

Abstract

On November 12th 2017, a $M_W=7.3$ earthquake struck near the Iranian town of Ezgeleh, close to the Iran-Iraq border. This event was located within the Zagros fold and thrust belt which delimits the continental collision between the Arabian and Eurasian Plates. Despite a high seismic risk, the seismogenic behaviour of the complex network of active faults is not well documented in this area due to the long recurrence interval of large earthquakes. In this study, we jointly invert InSAR and near-field strong-motions to infer the geometry of a flat fault and a kinematic slip model of the rupture. The kinematic slip distribution reveals an impulsive seismic source with a strong southward rupture directivity, consistent with significant damage South of the epicenter. We also show that the slip direction does not match plate convergence, implying that some of the accumulated strain must be partitioned onto other faults.

Plain Language Summary

Iran is a very seismically active region. However, the 2017 Ezgeleh earthquake ($M_W=7.3$) occurred in a region where large earthquakes have not been documented for several centuries. Our knowledge of fault locations, geometry, and seismic behaviour is therefore limited in this region. We use near-field seismological and geodetic data to retrieve the spatial and temporal distribution of slip occurring on the fault during the Ezgeleh earthquake. We show that the high slip rate and Southward directivity of the rupture may have worsened damage South of the epicentre. We also observe that tectonic motion is partitioned between different type of faults. Although the Ezgeleh earthquake did release a significant part of that strain, other seismogenic faults in the region could represent an important hazard for nearby population.

1 Introduction

On November 12th, 2017, the Iranian province of Kermanshah and the Iraqi Kurdistan was shaken by a severe $M_W=7.3$ earthquake located underneath the border. It caused the death of ~ 630 people and considerable damage, in particular in the Iranian city of Sarpol-e Zahab (c.f. Figure 1). The earthquake triggered numerous landslides and rock falls, including a massive 4x1 km landslide in Kermanshah (Miyajima et al., 2018).

48 The hypocenter is located within the Zagros Mountains near the Iranian town of
49 Ezgeleh, a tectonically active region that accommodates crustal shortening (e.g., Berberian & King, 1981) resulting from the collision between the Arabian Plate and the Eurasian
50 Plate. About a third to a half of current convergence is accommodated within the Za-
51 gros belt (Vernant et al., 2004). The belt hosts many moderate earthquakes ($M=5-6$)
52 with depths ranging from 4 km to 20 km, although these values are debated (e.g., Ni-
53 aзи et al., 1978; Talebian & Jackson, 2004; Nissen et al., 2011). Our knowledge of the re-
54 gional seismo-tectonics is further complicated by the very rare occurrence of co-seismic
55 surface rupture (Talebian & Jackson, 2004; Walker et al., 2005).
56

57 The Ezgeleh earthquake occurred at the transition between the Lorestan Arc in the
58 South-East and the Kirkuk Embayment in the North-West (c.f. Figure 1). The area is
59 covered by a 8-13 km thick sedimentary cover heavily folded into numerous anticlines
60 (e.g., Falcon, 1969; Alavi, 2007). Sediments are crossed by many thrust faults that flat-
61 tens within the basement (Sadeghi & Yassaghi, 2016; Tavani et al., 2018). As expected
62 from the lack of surface ruptures and fault scarps, most of these faults are blind, hence
63 the difficulty in inferring their geometry. In this region, plate convergence is roughly North-
64 South (c.f., Figure 1) with a rate between 19 mm/yr (Kreemer et al., 2014) and 24 mm/yr
65 (DeMets et al., 2010). Slip is partitioned between thrust faults at the front of the belt,
66 such as the Mountain Front Fault, the High Zagros Fault and the Zagros Foredeep Fault,
67 and the Main Recent Fault, a right-lateral strike-slip fault located at the back of the belt
68 (c.f., Figure 1; Berberian, 1995). This part of the Zagros belt hosts moderate seismic-
69 ity, but the last significant earthquakes ($5.9 \lesssim M \lesssim 6.4$) to strike the area happened in
70 958 and 1150 (Ambraseys & Melville, 2005). Therefore, our understanding of the regional
71 seismo-tectonic setting is obscured by the lack of significant earthquakes and the absence
72 of ground geodesy. The 2017 Ezgeleh earthquake highlighted the seismic hazard in this
73 portion of the Zagros belt. Its analysis hence provides a unique opportunity to enrich
74 our understanding of the region and the associated seismic hazard. In addition, the avail-
75 ability of near-field strong-motion records offers the possibility to closely study the prop-
76 agation of the rupture on the fault and its interaction with the surrounding rheology.

77 In this study, we propose a stochastic analysis of the 2017 earthquake source pro-
78 cess. We use a Bayesian framework to infer a population of co-seismic slip models that
79 fit available observations. While currently available studies were either limited to the fi-
80 nal distribution of slip on the fault (He et al., 2018; Wanpeng et al., 2018; Barnhart et

81 al., 2018; Yang et al., 2018; Vajedian et al., 2018) or used far-field teleseismic data (Chen
82 et al., 2018), we jointly invert InSAR and near-field strong-motion data to propose a kine-
83 matic description of the earthquake source. Unlike these studies, we use a local layered
84 elastic model (Supplementary Table T1) to limit mismodelling.

85 **2 Inversion of co-seismic slip**

86 **2.1 Observations**

87 Due to the remote location of the event, the only available geodetic data come from
88 interferometric Synthetic Aperture Radar (InSAR). We use three SAR interferograms
89 computed from acquisition by the Sentinel-1 satellite, along two ascending and one de-
90 scending tracks (Figures 2a and S1-2). We use the ISCE software with precise orbits and
91 SRTM DEM to compute the co-seismic interferograms (Rosen et al., 2012). The coher-
92 ence of the radar phase is excellent, likely due to the arid conditions of this region. We
93 measure up to 80 cm of ground displacement toward the satellite in the ascending tracks,
94 suggesting uplift and/or displacement toward the South-West. The number of data points
95 in the unwrapped interferograms is reduced using a recursive quad-tree algorithm (cf.,
96 Fig.S1; Lohman & Simons, 2005). We estimate uncertainties due to tropospheric per-
97 turbations in the phase by estimating empirical covariance functions for each interfer-
98 ograms (Jolivet et al., 2014). Estimated covariance parameters are summarized in Ta-
99 ble T2.

100 We include near-field seismic waveforms recorded by 10 strong-motion accelerom-
101 eters from the Iran Strong Motion Network (ISMN) to constrain the temporal evolution
102 of slip during the earthquake rupture. Although located only on one side of the rupture,
103 all stations are within 102 km of the epicentre (c.f. Figure 2b). Details on strong mo-
104 tion data processing are given in Supplementary Text T1. The East component of the
105 two stations located South of the rupture (SPZ and GRS) was not used due to poor qual-
106 ity of the record. We integrate accelerometric data to recover ground velocity, downsam-
107 pled to 1 sps. Waveforms are bandpass filtered between 7 Hz and 50 Hz using a 4th or-
108 der Butterworth band-pass filter. Waveforms are then windowed around the first arrivals.

2.2 Estimation of the fault plane

The two nodal planes of the global CMT mechanism (Ekström et al., 2005) are either a shallow North-East dipping plane (351° strike and 11° dip) or a nearly vertical plane (121° strike and 83° dip). We conduct a grid-search on fault geometry parameters for each nodal plane. The goal is to discriminate between the two planes and to find the optimal fault geometry to limit forward modelling errors.

We grid-search the fault location and its strike and dip angles by inverting the InSAR displacement to find the geometry that better explains the observations. For each tested geometry, slip is inverted on 96 subfault patches using a simple least-square technique. More details on the method are given in Supplementary text T2. We find that even the best sub-vertical plane has a RMS six times larger than the shallow-dipping plane (c.f. Figures S4 and S5). Although the sub-vertical plane is compatible with a back-thrust fault that may exist in the region (Tavani et al., 2018) or with the reactivation of steep normal faults (Jackson, 1980), the shallow dipping plane is in better agreement with the tectonic setting (e.g. Berberian, 1995; Paul et al., 2010; Vergés et al., 2011). Our optimal plane (351° strike, 14° dip, 13 km depth) agrees well with other studies using a similar grid-search approach (Barnhart et al., 2018; Wanpeng et al., 2018). In the following, we will consider that the Ezgeleh earthquake occurred on our optimum shallow dipping plane.

2.3 Co-seismic slip modelling

We use fault parameters inferred in part 2.2 to construct a planar fault and divide it in 96 subfault patches, each with a dimension of 7x7 km². Source model parameters include total final slip, rupture velocity, and rise time for each patch along with hypocenter location. We define \mathbf{m}_S the vector including the two components of static slip (i.e. final integrated slip), and \mathbf{m}_K the vector of kinematic parameters describing the temporal evolution of slip.

We solve the problem in a Bayesian framework using AlTar, an Markov Chain Monte Carlo algorithm based on the algorithm described by Minson et al. (2013). It samples the full posterior probability distribution of the models that fit observations and that are consistent with prior information. The strength of our solution is that it does not rely on any spatial smoothing and provides accurate estimates of the posterior slip un-

140 certainty. We sample the posterior probability density $p(\mathbf{m}_S, \mathbf{m}_K | \mathbf{d}_S, \mathbf{d}_K)$ given by

$$141 \quad p(\mathbf{m}_S, \mathbf{m}_K | \mathbf{d}_S, \mathbf{d}_K) \propto p(\mathbf{m}_K) p(\mathbf{m}_S) p(\mathbf{d}_S | \mathbf{m}_S) p(\mathbf{d}_K | \mathbf{m}_S, \mathbf{m}_K) \quad (1)$$

142 where \mathbf{d}_S and \mathbf{d}_K are the InSAR and strong-motion observations, respectively. The prior
 143 PDFs $p(\mathbf{m}_S)$ and $p(\mathbf{m}_K)$ are mostly uniform distributions designed to prevent some model
 144 features such as back-slip. They are described in details in Table T3. For further details
 145 on the method, the reader can refer to Supplementary text T3, Minson et al. (2013) and
 146 Gombert et al. (2018).

147 **3 Results**

148 In the first seconds following the hypocentral time, slip propagates in every direc-
 149 tion around the hypocentre (c.f. Fig 3 and supplementary movie M1). Approximately
 150 5 seconds after, the rupture almost only propagates toward the South. The largest slip
 151 rate occurs roughly after 6 seconds, 20 km South of the epicentre. We observe a strong
 152 directivity toward the South, consistent with observations of large ground velocities recorded
 153 on the North-South component of stations SPZ and GRS (c.f., Fig 2 and S3). In addi-
 154 tion, we infer a large slip rate on the fault. As shown in Figures 4d-e and S3, slip rate
 155 increases up to more than 3 m/s where the slip is maximum. The slip rate functions of
 156 two fault patches presented here show the fast increase in slip rate associated with a short
 157 rise time of ~ 5 s, defining a sharp slip pulse (Heaton, 1990). Although larger than the
 158 values usually reported in kinematic slip models (usually ranging from 0.1 m/s to 1 m/s),
 159 our slip rate estimates for this event are compatible with well documented earthquakes
 160 (e.g., Minson et al., 2014; Cirella et al., 2012) and numerical models (e.g., Kaneko et al.,
 161 2008). The fast slip rate of the fault is reflected on the total moment rate function (Fig.
 162 3c), which shows that 90% of the moment was released within the first 14 seconds of the
 163 rupture, depicting an overall impulsive earthquake.

164 The posterior mean model of the final cumulative slip is shown in Figure 4a. At
 165 first order, this solution is in agreement with previously published static models (Barn-
 166 hart et al., 2018; Wanpeng et al., 2018). We infer a ~ 50 km long and ~ 30 km wide rup-
 167 ture, with a peak slip of $5.5 \text{ m} \pm 0.5 \text{ m}$. One difference arises as previous models proposed
 168 that two distinct asperities ruptured during the earthquake. Our posterior mean model
 169 does not show a clearly distinct rupture area in the North, closer to the hypocenter. How-
 170 ever, roughly 20% of the models in our solution present such a feature (see Supplemen-

171 tary Movie M2). This indicates that it is in the realm of possibilities but available ob-
 172 servations cannot entirely resolve it. The slip direction is constant along most of the fault,
 173 with a $131.5^\circ \pm 0.8^\circ$ rake corresponding to a motion toward the South-West. The inferred
 174 focal mechanism is therefore consistent with long-period moment tensor inversions.

175 Our Bayesian framework allows us to directly infer the posterior uncertainties as-
 176 sociated with the model parameters. Slip uncertainties are represented on Figure 4a by
 177 the 95% confidence ellipses. In addition, posterior marginal distributions after the static
 178 and kinematic inversions of the along-rake slip of two fault patches are shown in 4b-c.
 179 Unsurprisingly, the inclusion of kinematic observations reduces the posterior uncertain-
 180 ties of those parameters. On the highest slipping patch for instance, the $1-\sigma$ posterior
 181 uncertainty decreases from 0.82 m to 0.52 m. Over the fault, we observe a rather low pos-
 182 terior uncertainty at shallow and intermediate depths, where slip is located. At depths
 183 larger than 15 km, uncertainties become more significant. However, the inspection of each
 184 model composing the solution reveals a good consistency in the slip distribution, with
 185 nonetheless a larger variability in the northern part of the rupture (c.f., supplementary
 186 movie M2).

187 As shown in Figures S1, S2 and S6, the model predictions of our solution strongly
 188 fit the Sentinel-1A observations. Residuals are particularly small for the ascending track,
 189 which has the lowest observational errors and the narrowest time window around the main-
 190 shock (see Table T2). Stochastic model predictions of the strong-motion data are shown
 191 in Figure 2 and S3. Overall, our solution can explain the observations with a great ac-
 192 curacy. Posterior model predictions of stations KAT, SNI and MHD suffers from a larger
 193 uncertainty, likely explained by the larger distance separating them from the hypocen-
 194 ter.

195 **4 Discussion**

196 As suggested by previous studies (Barnhart et al., 2018; He et al., 2018; Wanpeng
 197 et al., 2018), the Ezgeleh earthquake likely occurred on the Mountain Flexure Fault (some-
 198 times referred as Main Front Fault, noted MFF in Figure 1). Along the major part of
 199 the Zagros belt, the MFF follows a NW-SE axis with a $\sim 120^\circ$ azimuth and is aligned with
 200 many topographic features (visible on the DEM presented in Figure 4). However, the
 201 strike of the fault differs by about 50° with the topography orientation at the location

202 of the earthquake. This discrepancy is explained by a major bend in the MFF at this
203 location as it transitions between the Lurestan Arc (LA) in the South and the Kirkuk
204 Embayment (KE) in the North (e.g., Koshnaw et al., 2017; Vergés et al., 2011). Inter-
205 estingly, the fault bend between the LA and KE corresponds to the northern bound of
206 the rupture (Fig. 3). This geometry change possibly stopped the rupture propagation,
207 as suggested by numerical models (Aochi et al., 2000). The rupture may also have been
208 halted by the 8 km to 10 km thick sediment cover, whose depth roughly corresponds to
209 the updip limit of slip.

210 These sediments are heavily folded in the forearc basin and hosts many large an-
211 ticlinal folds (e.g., Kent, 2010; Casciello et al., 2009). These folds are evidence for thin-skin
212 shortening occurring within the belt (Koshnaw et al., 2017; Tavani et al., 2018). How-
213 ever, the slip of the 2017 earthquake occurred at larger depth, between 10 km and 15 km.
214 This deeper co-seismic deformation suggests that thick-skin shortening is also happen-
215 ing in this part of the Zagros range (Nissen et al., 2011; Vergés et al., 2011). The slip
216 direction of the Ezgeleh earthquake on the MFF is nearly perpendicular to the alignment
217 of the topographic features mentioned above (cf., Fig. 4a), creating a maximum 65 cm
218 of uplift and 33 cm of subsidence across the belt (c.f., Figure S8). Despite the relatively
219 large depth of the Ezgeleh earthquake, such co-seismic deformation may thus contribute
220 to the growth of the Zagros topography. Afterslip might also contribute although it seems
221 to occur on a shallow dipping decollement at the front of the mountain range (Barnhart
222 et al., 2018).

223 An interesting feature of the Ezgeleh earthquake is the discrepancy between the
224 co-seismic slip direction and the current plate motion. Both the GSRM v2.1 model (Kreemer
225 et al., 2014) and the MORVEL model (DeMets et al., 2010) predict a nearly N-S plate
226 convergence (see Fig. 1) while the overall co-seismic slip vector is oriented on a S 30° W
227 axis (see Fig. 4). This axis difference suggests that strain partitioning is occurring in this
228 part of the Zagros belt, with a partial decoupling between the thrust and right-lateral
229 strike-slip motion (Platt, 1993; McCaffrey, 1992). Strain partitioning in the Lurestan Arc
230 and the Kirkuk Embayment has been proposed before based on the analysis of regional
231 focal mechanisms (Talebian & Jackson, 2004). The Main Recent Fault (MRF; see Fig-
232 ure 1) is a major NW-SE 800 km long right-lateral strike-slip fault which accommodates
233 some of the strain (Tchalenko & Braud, 1974). It hosted several large earthquakes and
234 has a ~ 50 km horizontal offset (Talebian & Jackson, 2002). However, other structures

235 may be accommodating the strike-slip component of the convergence. Between July and
236 November 2018, three significant aftershocks with respective magnitudes of $M_W=5.8$,
237 $M_W=6.0$, and $M_W=6.2$ occurred south of the mainshock epicenter (c.f. Figure 1b). These
238 events present a right-lateral strike-slip focal mechanism, but are located more than 100 km
239 West of the MRF. They could be located on the Khanaqin fault, a N-S strike-slip struc-
240 ture marking the boundary between the Lurestan Arc and the Kirkuk Embayment (e.g.,
241 Blanc et al., 2003; Hessami et al., 2001; Berberian, 1995). However, there is very lim-
242 ited evidence that the Khanaqin fault is actually a strike-slip fault. As a matter of fact,
243 a recent study by Tavani et al. (2018) using reconstruction of seismic profile proposed
244 that the Khanaqin fault is back-thrust structure accommodating the SW-NE motion.
245 Therefore, undetected strike-slip faults may be accommodating some of the strike-slip
246 deformation closer to the forearc than the MRF. These faults represent a major seismic
247 risk for population of nearby cities and villages, both in Iran and in Iraq.

248 The good spatial and temporal resolution of our kinematic slip model reveals in-
249 teresting features. Fig. 3 and S7 shows that the rupture starts as a growing crack that
250 rapidly transition into a pulse with a rise time of about 4 sec. This crack-pulse transi-
251 tion occurs within the first four seconds and less than 7 km from the hypocenter (Fig
252 S7), therefore away from the rupture boundaries. This pulse-like behaviour is therefore
253 unlikely to result from healing phases emanating from the along-dip finiteness of the fault
254 (Day, 1982). A rapid crack-pulse transition is in agreement with early observations by
255 Heaton (1990) and later studies (e.g., Beroza & Mikumo, 1996; Meier et al., 2016). Such
256 self-healing pulse may result from a number of mechanisms such as frictional self-healing,
257 fault strength or stress heterogeneities, bimaterial effects and wave reflections within low-
258 velocity fault zones (e.g., Perrin et al., 1995; Andrews & Ben-Zion, 1997; Huang & Am-
259 puero, 2011). After this early transition from a growing crack, the rupture continues its
260 journey along-strike as a decaying pulse toward the North, and a strong growing pulse
261 toward the South.

262 This strong southward propagating pulse seems to have a significant impact in the
263 distribution of damage and landslides triggered by the earthquake. The Ezgeleh earth-
264 quake induced extensive destructions of dwellings in Iraqi Kurdistan, but mostly in the
265 Iranian province of Kermanshah. Figure 1b) shows the intensity of damage created by
266 the mainshock. It is obtained from field observations conducted by the International In-
267 stitute of Earthquake Engineering and Seismology of Iran (IIEES). Damage intensity roughly

268 follows the surface projection of the slip distribution, but larger damage was reported
269 in the South. In addition to building damage, many rockfalls and landslides occurred
270 south of the rupture and up to 125 km from the centroid, including a large 4 km long
271 and 1 km wide landslide (Miyajima et al., 2018). Many different factors can largely in-
272 fluence the aftermath of an earthquake, like soil nature or mountain slopes. In addition
273 to rupture directivity, studies have suggested that the strong impulsiveness of the source
274 can intensify ground shaking (Melgar & Hayes, 2017). The large slip-rate and short rise-
275 time of the southward propagating pulse may therefore have aggravated the damage ob-
276 served South-West of the Ezgeleh earthquake.

277 **5 Conclusion**

278 The 2017 Ezgeleh earthquake breaks a long hiatus on strong events affecting the
279 Zagros thrust and fold belt in the Kermanshah province. The joint inversion of InSAR
280 and near-field strong-motion observations reveals a predominantly thrust motion on a
281 near-horizontal blind crustal fault. We also infer a highly impulsive source propagating
282 toward the South. These kinematic properties may have play a role in the numerous slope
283 instabilities and in the important damage that affected Iranian cities.

284 Furthermore, the misalignment between the plate convergence and the slip direc-
285 tion provide additional evidences for a strain partitioning in this part of the Zagros belt
286 between thrust motion on flat crustal faults and right-lateral strike-slip. As suggested
287 by late aftershocks, unmapped dextral faults could be accommodating part of that shear
288 strain, and therefore represent an important seismic risk for nearby populations.

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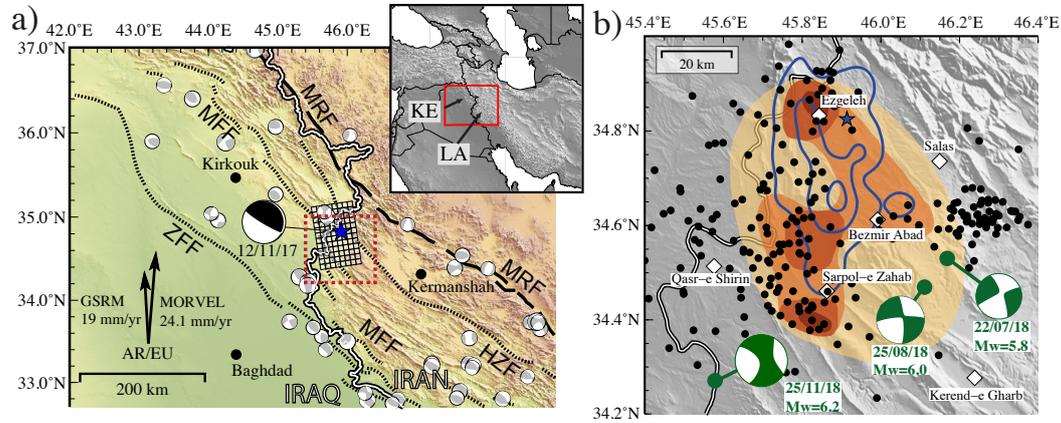
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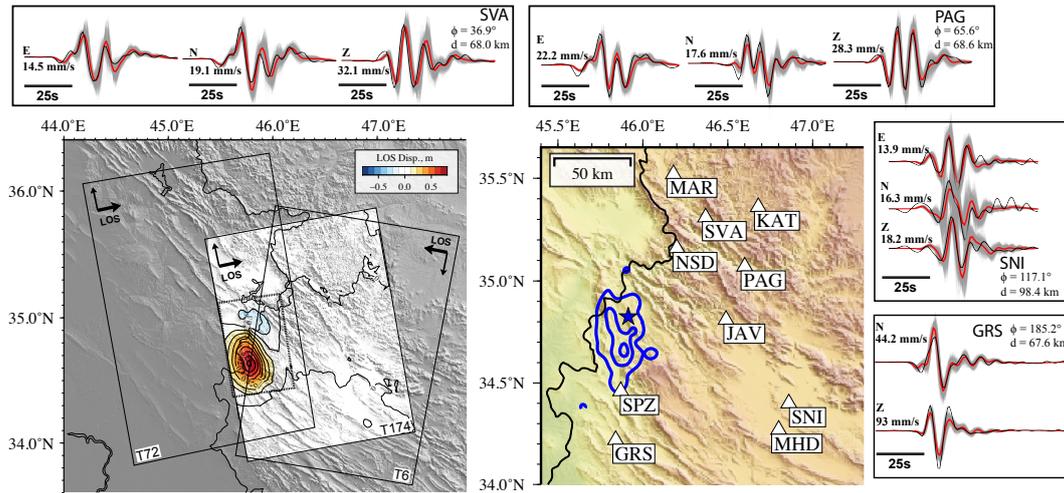
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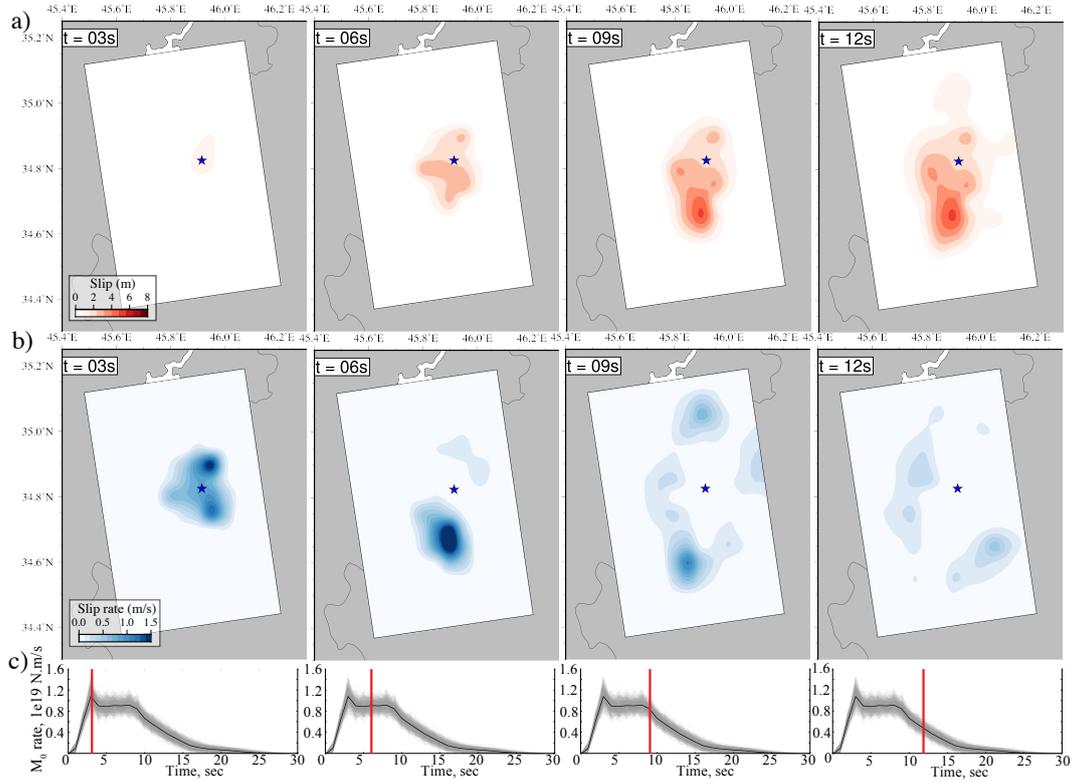
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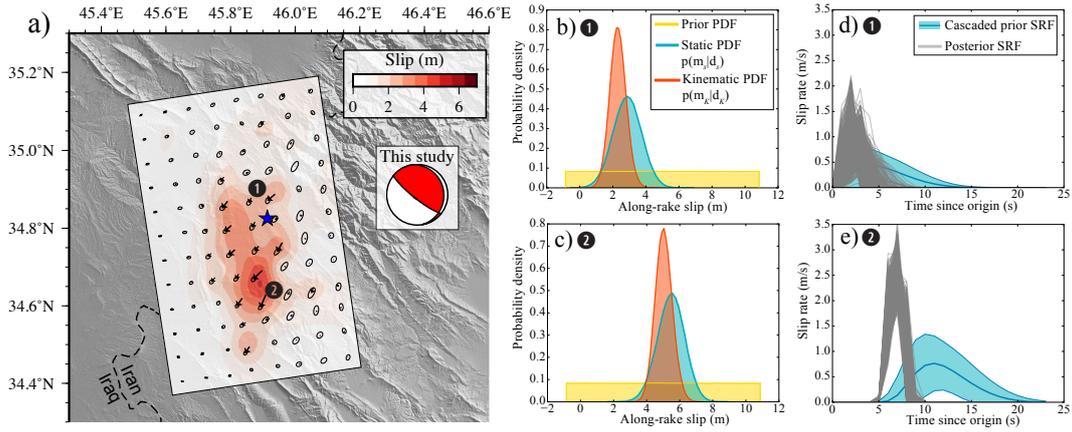
462 **Figure 1. Regional seismotectonic context and damage associated with the 2017**
 463 **Ezgeleh earthquake. a)** Blue star marks the epicentre location, and the squares represent
 464 the fault parametrisation. Grey moment tensors are from the Global CMT catalogue (Ekström
 465 et al., 2012). Dashed black line is the Main Recent Fault (MRF) and dotted lines are supposed
 466 location of regional blind faults (MFF: Mountain Flexure Fault; HZF: High Zagros Fault; ZFF:
 467 Zagros Foredeep Fault; Berberian, 1995). Arrows indicate the convergence of the Arabian plate
 468 (AR) with respect to stable Eurasia (EU) from the GSRM v2.1 (Kreemer et al., 2014) and
 469 MORVEL (DeMets et al., 2010) models, computed with the UNAVCO Plate Motion Calculator.
 470 LA: Lorestan Arc. KE: Kirkuk Embayment. Red dashed rectangle indicates position of b). **b)**
 471 Black dots are aftershocks located by the International Institute of Earthquake Engineering and
 472 Seismology of Iran (IIEES). Focal mechanisms from the Global CMT catalogue of three large
 473 aftershocks are shown in green. Brown colours indicate the level of damage based on a compila-
 474 tion of destruction rate and landslide activity interpolated from field surveys conducted by the
 475 Geological Survey of Iran (GSI, 2017). The darker the colour, the more intense the damage. Blue
 476 lines are the 1.5 m co-seismic slip contour.



477 **Figure 2. Observations used in the inversion.** a) Unwrapped Sentinel-1A interfer-
 478 ograms showing surface displacement in LOS direction (Track 174). The footprint of one
 479 additional ascending and one descending tracks are also shown. Data, predictions and model
 480 performance of the 3 interferograms is available in Figs. S1-2. b) Location of strong-motion
 481 records (white triangles). c-f) Waveforms of four selected station around the epicenter. For each
 482 waveform, the bold number indicates its maximum amplitude. Φ and d are station azimuth
 483 and distance to epicentre, respectively. The black line is the recorded waveform, grey lines are
 484 stochastic predictions for our posterior model, and the red line is the mean of stochastic predic-
 485 tions. Remaining waveforms are shown in Fig. S3



486 **Figure 3. Temporal evolution of co-seismic slip.** a) Cumulative slip on the fault 3 s, 6 s,
 487 9 s, and 12 s after the origin time. The red colour-scale indicates slip amplitude. b) Evolution of
 488 slip rate on the fault. c) Source time function (STF) of the event. Grey lines are stochastic STFs
 489 inferred from our model population while the black curve represents the posterior mean STF.
 490 Vertical red lines indicate the temporal position of each one of the snapshots



491 **Figure 4. Final co-seismic slip distribution** a) Colour and arrows on the fault plane
 492 indicate amplitude and direction of slip, respectively. Ellipses represents the 95% posterior un-
 493 certainty. Results presented in subfigures b-e) are obtained for patches labelled 1 and 2. The
 494 background topography comes from the Shuttle Radar Topography Mission (SRTM; Farr et
 495 al., 2007). **b-c)** Prior, posterior static PDF, and posterior kinematic PDF of along-rake slip in
 496 patches 1 and 2. **d-e)** Slip rate evolution in patches 1 and 2. Blue line is the mean prior Slip
 497 Rate Function (SRF) used in the sampling, surrounded by 1- σ uncertainties. Posterior SRFs in
 498 grey are from 1000 thousands models randomly selected from our solution.