

The Warnie Volcanic Province: A Jurassic Volcanic Province in Central Australia

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Abstract

The Cooper and Eromanga Basins of South Australia and Queensland are the largest onshore hydrocarbon producing region in Australia. Igneous rocks have been documented infrequently within end of well reports over the past 34 years, with a late Triassic to Jurassic age determined from well data. However, the areal extent and nature of these basaltic rocks were largely unclear. Here, we integrate seismic, well, gravity and magnetic data to clarify the extent and character of igneous rocks preserved within Eromanga Basin stratigraphy overlying the Nappamerri Trough of the Cooper Basin. We recognise mafic monogenetic volcanoes that extend into tabular basalt lava flows, igneous intrusions and, more locally, hydrothermally altered compound lava flows. The volcanic province covers ~7500 km² and is proposed to have been active between ~180-160 Ma. We deem this Jurassic volcanic province the Warnie Volcanic Province after the Warnie East 1 exploration well, drilled in 1985. The distribution of extrusive and intrusive igneous rocks is primarily controlled by basement structure, with extrusive and intrusive igneous rocks elongate in a NW-SE direction. Finally, we detail how the WVP fits into the record of Jurassic volcanism in eastern Australia. The WVP is interpreted as a product of extension and intraplate convective upwelling above the subducting Pacific Slab. The discovery of the WVP raises the possibility of other, yet unidentified, volcanic provinces worldwide.

Keywords

Intraplate; volcanism; monogenetic; Australia; Jurassic

1. Introduction

The Cooper and the overlying Eromanga Basin extend for 130,000 km² and 1,000,000 km² respectively across central, north and eastern Australia. Together, they represent the largest conventional onshore hydrocarbon-producing region in Australia (Hall, 2016). The first gas discovery was made in 1963 and since then, over 1,400 producing wells have been drilled, with 190 gas and 115 oil fields discovered throughout both basins (Fig. 1) (Mackie, 2015). The majority of discoveries within the Cooper Basin are situated within structural traps on regional highs, however, a number of companies began to pursue unconventional hydrocarbon

33 plays within the basin during the late 2000s, resulting in an increase in activity within the
34 Nappamerri Trough (Hillis *et al.*, 2001; Pitkin *et al.*, 2012; Khair *et al.*, 2013; Scott *et al.*, 2013;
35 Hall *et al.*, 2016). The Nappamerri Trough is the largest and deepest depocentre within the
36 Cooper Basin covering $\sim 10,000$ km² and reaching present day depths of ~ 4.5 km (Fig. 1A).
37 The acquisition of new well and seismic data due to the exploration activities has resulted in
38 the increased recognition of igneous rocks in the Cooper and Eromanga Basins.

39 Igneous rocks of suspected late Triassic-Jurassic age have been sporadically
40 encountered by drilling during the past 34 years (Short, 1984; Boothby, 1986; Bucknill, 1990;
41 Allen, 1998), though there has been no systematic study of the character, origin and
42 significance of these igneous rocks. Here, we combine extensive seismic, well and airborne
43 geophysical data to document the extent and character of Mesozoic volcanics recognised
44 within Eromanga Basin stratigraphy overlying the Nappamerri Trough of Cooper Basin age.
45 The Warnie Volcanic Province (WVP) is the suggested name for this $>7,500$ km² suite of
46 volcanics identified within the Jurassic succession of the Eromanga Basin (Fig. 1B). The
47 province is named after the Warnie East 1 exploration well, which was drilled in 1985 and
48 encountered 65 m of basalt. We provide an in-depth study of the regional implications of the
49 WVP for the Cooper and Eromanga Basins. We find that SE-NW intracratonic sag faulting
50 (Kulikowski *et al.*, 2015) controlled the morphology of igneous rocks whilst the basin structure
51 likely controlled the location of igneous rocks.

52 Finally, we discuss the origin of the WVP. Evidence for Triassic to mid-Cretaceous
53 volcanic activity is pervasive throughout the sedimentary basins of central Australia,
54 manifested by an influx of volcanic arc-derived sediment into the Eromanga Basin (Boult *et al.*,
55 1998), silicic tuffs in the Surat Basin (Wainman *et al.*, 2015) and widespread volcanogenic
56 zircons throughout the Great Artesian Basin (Bryan *et al.*, 1997). However, the origin of
57 intraplate volcanism is contentious and can be difficult to isolate from tectonic processes
58 occurring close to the region of volcanism (Conrad *et al.*, 2011; Zhou *et al.*, 2018). The WVP
59 was located >500 km away from a significant plate boundary during the Jurassic. It is therefore
60 important to consider the source of this volcanism and the tectonic implications for central
61 Australia. We theorise that the WVP's geodynamic origin is ultimately linked to subduction
62 at the eastern margin of Gondwana, through asthenospheric shear ~ 1500 km from the oceanic
63 trench, however, we also discuss other mechanisms that can produce intraplate volcanism.
64 Our discovery of a previously undescribed volcanic province in an area that has undergone
65 >50 years of extensive subsurface data collection due to continued hydrocarbon exploration,

66 raises the prospect of other undiscovered intraplate volcanic provinces both in Australia and
67 in other continental areas worldwide.

68 **2. Geological Overview of the Cooper and Eromanga Basins**

69 The geology of southwest Queensland and northeast South Australia, central Australia, is
70 defined by a series of intracratonic basins stacked on top of each other; the Warburton,
71 Cooper, Eromanga and Lake Eyre Basins (Fig. 2 & Fig. 3).

72 The Warburton Basin is a lower Palaeozoic sedimentary basin that hosted Silurian to
73 Carboniferous granite emplacement (Murray 1986; Boucher 1997; Alexander *et al.*, 1998).
74 The Cooper Basin unconformably overlies the Warburton Basin. The Cooper Basin is a
75 northeast to southwest trending intracratonic structural depression that accommodated the
76 deposition of sedimentary rocks from the late Carboniferous to the Middle Triassic (Laws &
77 Gravestock, 1998). Several tectonic origins for the Cooper Basin have been suggested,
78 including a mildly compressional structural depression (Apak, 1994 & Sun, 1997), a depression
79 created through wrenching (Kuang, 1985) and extension during post-compression flexural
80 relaxation (Kulikowski *et al.*, 2015). Deposition throughout the early Permian was dominated
81 by highly sinuous fluvial systems situated on a major floodplain with peat rich swamps and
82 lakes (Abul Khair *et al.*, 2015). Later Triassic deposition took place in a period of tectonic
83 quiescence (Gravestock *et al.*, 1998).

84 Deposition in the Cooper Basin was terminated by regional uplift, compressional folding and
85 erosion in the middle Triassic. The Cooper Basin is unconformably overlain by the Eromanga
86 Basin, which forms part of the Great Artesian Superbasin (GAB), a group of interconnected
87 basins that cover much of Queensland, South Australia and New South Wales. The Eromanga
88 Basin is an intracratonic sag basin whose subsidence has been attributed to dynamic
89 topography induced by subduction of the Pacific plate below Australia (Gallagher & Lambeck,
90 1989; Russell & Gurnis, 1994; Matthews *et al.*, 2011). Deposition within the Eromanga Basin
91 was controlled by subsidence rates and tectonic activity on the margins of the Australian Plate
92 that were attributable to subduction of the Pacific plate beneath the Australian plate
93 (Alexander *et al.* 1996; Draper, 2002).

94 **2.1 The Jurassic Succession of the Eromanga Basin**

95 The Jurassic succession of the Eromanga Basin is of stratigraphic importance as it is where the
96 extrusive igneous rocks documented within this study are believed to be located. Much of the

97 Jurassic strata in the Eromanga Basin consists of predominantly quartzose fluviatile sandstones
98 with subordinate shales and some coals (Senior *et al.*, 1978; Exon & Burger, 1981; Draper
99 2002). Sediments in the Eromanga Basin can be divided into three packages of decreasing age;
100 lower non-marine, marine and upper non-marine. The focus of this study is the lower non-
101 marine package that was deposited between the early Jurassic (~206 Ma) and the early
102 Cretaceous (~119 Ma) (Fig. 2). During this time, large sand-dominated, braided fluvial systems
103 drained into lowland lakes and swamps. Throughout deposition, it is believed that sediment
104 supply was a competing product of input from arc volcanism to the east, and sediment sourced
105 from stable cratonic regions to the south (Boult *et al.*, 1997). In the next part of this paper
106 we detail the Early to Late Jurassic stratigraphy in order to provide a palaeogeographic
107 framework for the study. The stratigraphic framework of the Jurassic succession has been
108 iterated considerably in the last 20 years, evidenced by the progression of Alexander &
109 Hiburt's stratigraphic ages (1996) to Reid *et al.*'s (2009), to those listed on the Australian
110 Stratigraphic Units Database (Geoscience Australia, 2017). For each formation we have used
111 the most up-to-date reference available.

112 2.1.1 Poolowanna Formation, Early Jurassic, ~200–178 Ma

113 In southwest Queensland, the Poolowanna Fm. sits unconformably on top of the Late Triassic
114 unconformity, which formed as a result of uplifting, tilting and erosion of the Cooper Basin at
115 the end of the Early to Middle Triassic (Fig. 5). The palaeotopography of this unconformity
116 strongly controlled deposition of the Poolowanna Formation and the subsequent Hutton
117 Sandstone (Hall *et al.*, 2015). Lithologies consist of coal and silt deposited in a highly sinuous
118 fluvial setting with minor coal swamps (Hall, 2015).

119 2.1.2 Hutton Sandstone, Early to Middle Jurassic, ~178–166 Ma

120 The Hutton Sandstone is a high energy braided fluvial Jurassic sandstone that either sits
121 conformably on top of the Poolowanna Fm. or, where the Poolowanna is not present,
122 unconformably above Triassic sediments (Kapel, 1966; Watts, 1987). The Hutton Sandstone
123 is a mineralogically mature fine to coarse-grained sandstone with minor siltstone interbeds
124 (Hall *et al.*, 2015). Throughout much of the Great Artesian Basin, the Hutton is a regional
125 aquifer extending over an area of 1.7 million km² (Hamilton *et al.* 1998). Similarly, the Hutton
126 is the main reservoir within the Jurassic succession of the Eromanga Basin.

127 2.1.3 Birkhead Formation, Middle to Upper Jurassic, ~ 166–160 Ma

128 The boundary between the Hutton and Birkhead formations is sharp to transitional across
129 the Eromanga Basin (Lanzilli, 1999) and is associated with a change in the depositional setting
130 from the high energy, braided fluvial setting of the Hutton Sandstone to a low energy, fluvial-
131 deltaic to deltaic and lacustrine setting in the Birkhead Formation (Watts, 1987; Lanzilli, 1999).
132 This decrease in depositional energy is also associated with a change in sediment provenance
133 from craton-derived sediment to lithic-rich, volcanogenic sediment from a proposed volcanic
134 arc situated to the east above the subducting Pacific plate (Lanzilli, 1999). The diachronous
135 influx of volcanogenic sediments can be traced across the Eromanga Basin (Boult, 1997). It is
136 commonly picked as the top of the porous unit in the Hutton Sandstone (Patton, 1986; Watts,
137 1987), however, it can also be identified by the influx of lithic rich sandstones (Watts, 1987).
138 There are several key features associated with the sediments shed from the volcanic arc
139 (Patton, 1986):

- 140 • 50% quartz with high amounts of potassium and plagioclase feldspar. Trace amounts
141 of tourmaline, pyroxene, mica and zircon were also noted (Watts, 1987; Alexander
142 and Sansome, 1996).
- 143 • Lithic fragments mostly consisting of labile minerals.
- 144 • Common altering of the above lithics to kaolinite and chlorite or replacement of
145 carbonate cements such as siderite or dolomite.
- 146 • Rare glassy fragments.

147 Based on the lithology of the volcanogenic sediments, it has been concluded that the volcanics
148 that formed them were of acid to intermediate affinity (Patton, 1986). The influx of
149 volcanogenic sediments is believed to have been controlled by the margin above the
150 subducting Pacific plate (Boult *et al.*, 1998). During the Jurassic it is suggested that subduction
151 of the Pacific plate beneath Australia caused extensive arc volcanism (Tucker *et al.*, 2016).
152 When uplift in this area was greatest, volcanogenic sediments were shed into the Eromanga
153 Basin (Boult *et al.*, 1997). Volcanogenic sediment in the Birkhead Fm. often forms a weak link
154 with lower sealing capacity within what is the main sealing lithology for the Hutton Sandstone
155 (Boult *et al.*, 1997).

156 2.1.4 Adori Sandstone, Upper Jurassic, ~ 160-155 Ma

157 The Birkhead Formation is capped by the Adori Sandstone, marked by a transition to clean,
158 non-volcanogenic sedimentation. The Adori Sandstone is a well sorted and typically cross
159 bedded, fine to coarse-grained sandstone (Hibburt, PGSA). It is interpreted to have been
160 deposited in a low sinuosity fluvial system consisting of channel, point bar and flood plain
161 deposits in the Central Eromanga Basin (Burger & Senior, 1975; Hibburt, PGSA). Due to the
162 absence of igneous rocks in the Mesozoic to Cenozoic succession younger than the Adori,
163 the succession is not discussed in detail here.

164 2.2 Structural Setting of the Study Area

165 Structurally, the Cooper Basin can be divided into southern and northern sections
166 (Gravestock & Jensen-Schmidt, 1998; Hall, 2016). To the north, depocentres are dominantly
167 Triassic, with Permian depocentres in the south (Hall *et al.*, 2018). The northern and southern
168 sections of the Cooper Basin are separated by the Jackson-Naccowlah-Pepita Trend (JNP
169 trend) (Fig. 1). In the south west, depocentres are generally thicker, reaching a maximum
170 thickness of 2400 m in the Nappamerri Trough (Hall *et al.*, 2015) and many of these
171 depocentres have experienced synclinal folding due to differential compaction (Apak, 1997).
172 The major depocentres are separated by northeast-southwest trending ridges (Gravestock &
173 Jensen Schmidt, 1998).

174 The area of focus in this study is the eastern Nappamerri Trough of north east South Australia
175 and south west Queensland. The Nappamerri Trough contains the deepest basement within
176 the Cooper Basin and hosts the Halifax well that intersected the deepest Permian-age
177 sedimentary rocks in the basin at 4,209 m (Beach Energy, 2013b). To the north, the
178 Nappamerri Trough is bounded by the Gidgealpa, Canway, Packsaddle and Innamincka ridges
179 (Fig. 1A). The southern extent of the Nappamerri Trough is bounded by the Della-
180 Nappacoongee, Dunoon and Murteree ridges (Fig. 1) (Hall, 2016). The ridges bounding the
181 northern and southern Nappamerri Trough are cored by deep thrust faults reactivated from
182 pre-existing structures within the Warburton Basin (Robert *et al.*, 1990). These were believed
183 to have been reactivated in the Permian and the Cenozoic, when the Cooper and Eromanga
184 Basins were subject to contractional deformation (Sun, 1997).

185 **3. Methodology**

186 **3.1 Identification of Volcanics Using Subsurface Data**

187 A number of volcanic features can be recognised using subsurface data. Here we describe the
188 main techniques that were used to investigate igneous rocks within the Nappamerri Trough.

189 *3.1.1 Seismic Data*

190 Seismic reflection is a geophysical technique that estimates properties of the earth's
191 subsurface through the measurement of reflected seismic waves. It can be used to provide an
192 estimate of the acoustic impedance of the boundaries to layers of rock within the subsurface
193 (the product of the acoustic velocity and density of a rock). Seismic reflection data have been
194 shown by numerous studies to be especially effective at imaging igneous rocks due to the high
195 acoustic impedance of igneous material when compared to surrounding sedimentary rocks,
196 and the distinctive morphology formed by lava flows and intrusions (Planke *et al.*, 2000). Here
197 we provide a brief introduction to the different types of igneous rock identified in the
198 subsurface on seismic reflection data

199 Vents identified within seismic reflection data are typically grouped into either hydrothermal
200 or volcanic vents (Reynolds *et al.*, 2017). The morphologies of hydrothermal vents and
201 volcanoes are similar as they both exhibit eye, dome or crater shaped morphologies in seismic
202 data (Fig. 4B)(Planke *et al.*, 2005; Grove, 2013; Reynolds *et al.*, 2017). This can make
203 distinguishing hydrothermal vents from volcanoes difficult when seismic reflection data are
204 the sole data source.

205 Volcanoes can be confidently identified using seismic reflection data where they extend into
206 lava flows (e.g. Fig. 4A-D). Lava flows are distinguishable as layer-parallel acoustically hard,
207 bright reflectors within the subsurface (Planke *et al.*, 2000; Schofield & Jolley, 2013; Hardman
208 *et al.*, 2018A) (Fig. 4B). Extrusive igneous rocks are distinguishable from intrusive igneous
209 rocks as they do not transect stratigraphy.

210 Igneous intrusions have been studied intensely using 3D seismic data (Bell & Butcher, 2002;
211 Thomson & Schofield, 2004; Archer *et al.*, 2005; Magee *et al.*, 2014). Igneous intrusions form
212 acoustically hard, bright reflectors that are layer parallel or transect stratigraphy. In places,
213 they can produce forced folding of the overburden with onlapping stratigraphy facilitating age-
214 dating of the intrusion (Trude *et al.*, 2003).

215 3.1.2 Seismic Attributes

216 3D seismic reflection data facilitates the analysis of the geophysical properties of igneous rocks
217 in 3D. Seismic attributes such as RMS amplitude and spectral decomposition have proven to
218 be effective tools in picking the extent and morphology of lava flows in seismic data (Fig.4C)
219 (Schofield & Jolley, 2013; Planke *et al.*, 2017; Hardman *et al.*, 2018A; Reynolds *et al.*, 2018).
220 Spectral decomposition involves filtering 3D seismic reflection wavelets to produce three
221 amplitude components at distinct frequencies that are displayed as separate colour channels
222 using the primary colours red, green and blue. These channels are then blended to produce
223 full colour spectrum images where the igneous rocks appear visibly different to the
224 surrounding sediments, due to their differing reflectiveness for particular frequencies (they
225 transmit acoustic energy efficiently at all frequencies). Furthermore, this technique can be
226 used to investigate the geometries of igneous rocks, and features such as inflation lobes (Fig.
227 8C) can help with distinguishing igneous intrusions from extrusive igneous rocks.

228 3.1.3 Well Data

229 In addition to the use of seismic data, the integrated analysis of borehole data including
230 wireline log responses, core and cuttings is essential to accurately identifying igneous rocks
231 within the subsurface (Nelson *et al.*, 2009; Rider & Kennedy, 2011; Watson *et al.*, 2017). The
232 onshore response of igneous rocks in outcrop has been linked to offshore observations
233 facilitating the identification of different volcanic facies through petrophysical and petrological
234 analysis (Nelson *et al.*, 2009; Millet *et al.*, 2016).

235 Here, we use a combination of petrophysical and petrological data in our description of the
236 WVP. Although cuttings and core were not examined within this study, descriptions were
237 available within well reports that were integrated with petrophysical data and seismic
238 reflection data to provide a comprehensive overview of the data available.

239 3.2 Description of Dataset

240 The Cooper and Eromanga Basins have been explored extensively over the past 60 years, so
241 are characterised by the largest collection of well and seismic reflection data for any onshore
242 sedimentary basin in Australia (Carr *et al.*, 2016). A large database of seismic reflection data
243 was available through Santos Limited and the Queensland Government's Department of
244 Natural Resources and Mines. 3D and 2D seismic reflection surveys throughout the South
245 Australian and Queensland sides of the basin were assessed for the presence of volcanic rocks

246 (the extent of the seismic data examined is visible on Fig. 1B). Notably, only four surveys
247 available to this study (the Winnie 3D, the Gallus 2D, the Challum 3D and the Snowball 3D
248 survey (provided by Santos Limited)), were interpreted to contain volcanic rocks. Unless
249 stated, seismic data were displayed using normal (American) polarity, whereby a downward
250 increase in acoustic impedance corresponds to a positive (blue) reflection and a downward
251 decrease a negative (red) reflection.

252 Over 2000 wells have been drilled in the Cooper and Eromanga Basins. Despite this, of the
253 wells available, only 3 (Lambda 1, Orientos 2 and Kappa 1) were identified as having
254 intersected extrusive volcanic rocks with only 1 well (Warnie East 1) encountering an
255 intrusive igneous rock. The Warnie East 1, Lambda 1 and Orientos 2 wells were key in
256 constraining volcanism above the central Nappamerri Trough (Fig. 1C). The Kappa 1 well was
257 adjacent to the Snowball 3D survey, aiding interpretation of volcanism in the south west (Fig.
258 1C). To estimate the relative age of the volcanic units, well data were tied, where possible,
259 to the seismic reflection surveys available. For constraining the age of the Eromanga
260 succession within the Central Nappamerri trough, the Halifax, Ety, Padme, Charal and Anakin
261 wells were key (location on Fig. 1C).

262 Here we present an overview of all the available datasets that constrain the distribution and
263 age of igneous rocks within the study area. First, we detail the main WVP, within which the
264 lithology and morphology of the igneous rocks appear broadly similar throughout the eastern
265 Nappamerri Trough. We then deal with the Snowball 3D survey and the Kappa 1 well that
266 image volcanic rocks that have a markedly different petrological and geophysical character
267 when compared to the rest of the WVP. It is important we recognise that whilst this is the
268 extent of the data that was available during this study, the authors are aware of other seismic
269 reflection surveys within the Nappamerri Trough that were not available to download digitally
270 and were not possible to acquire within the period of time during which this study was
271 conducted. These surveys may also image igneous rocks that could fuel further work on the
272 area.

273 **4. Description of Igneous Rocks in Seismic Data within the eastern Nappamerri** 274 **Trough**

275 **4.1 Winnie 3D**

276 The main 3D seismic reflection survey utilised within this study is the Winnie 3D survey.
277 Acquired in 2012 by Drillsearch Energy Ltd. The survey was shot above the eastern
278 Nappamerri Trough in southwest Queensland over an area of 2545 km² (Fig. 1). The main
279 focus when acquiring the survey was on the unconventional targets in the area, typically the
280 extensive Permian coal-rich source rocks (Hillis *et al.*, 2001; Schmidt, 2013). The dominant
281 frequency within the Jurassic succession is ~40 Hz, with volcanic rocks in the region having
282 an acoustic velocity of 4 to 6 kms⁻¹ (taken from the Lambda-1 exploration well). Using an
283 average velocity of 5 kms⁻¹ a vertical resolution of ~30 m and a detection limit of 4 m was
284 calculated.

285 *4.1.1 Extrusive Igneous Rocks*

286 We identified ~100, <= 4 km² cone shaped features that often express an eye-shaped
287 morphology (Fig. 2B). The cones within the survey are less than 2 km long and 750 m wide,
288 and often elongate in a NW-SE direction. We have interpreted these as monogenetic
289 volcanoes, small cumulative volume volcanic edifices built up by one continuous, or many
290 discontinuous, small eruptions over a short time scale (<= 10 years) (Nemeth & Kereszturi,
291 2015).

292 Twelve of the volcanoes are linked with what appear to be elongate, NW-SE oriented
293 lava flows that are conformable to stratigraphy and have dimensions of 2-7 km in length and
294 0.5-2.5 km in width (covering areas of 4-13.5 km²) (e.g. Fig. 6 & 7). The flows appear to be
295 strata-concordant and confined to the Hutton Sandstone or the base of the Birkhead
296 Formation which would place the flows between 160 and 178 Ma (Alexander & Hibburt, 1996)
297 (Fig. 7).

298 We have also interpreted one feature within the survey as a diatreme. Maar-diatreme
299 volcanism can occur where a volcanic pipe forms through gaseous explosions that cut into
300 the country rock (Fig. 4 E-H)(White & Ross, 2011). These explosions form a maar (the crater
301 and associated ejecta ring) and a diatreme (the root to the maar), consisting of a steep-sided
302 cone shaped structure filled with pyroclastic, volcanic and country rocks)(White & Ross,

2011). Unlike vents and associated lava flows, they consist of downward dipping reflectors that truncate the underlying stratigraphy with typically chaotic internal reflectors (Fig. 4F). Spectral decomposition proved to be the most powerful tool for identifying maar-diatremes above the Nappamerri Trough, as the circular crater morphology stood out against the surrounding sediments (Fig. 4G). One feature interpreted to be a diatreme was identified within the survey, displayed in detail in (Fig. 4). It is $\sim 2.25 \text{ km}^2$ and $\sim 120 \text{ m}$ deep (150 ms, TWT) and is found in the west of the Winnie 3D survey (Fig. 6).

4.1.2 Intrusive Igneous Rocks

Igneous intrusions within the Winnie 3D survey are identified as igneous rocks that cross-cut the Triassic to Jurassic strata; notably the Nappamerri, Hutton and Birkhead Fms (e.g. Fig. 8A). We identified and mapped 14 intrusions in the seismic data; these occur $\sim 100\text{--}200 \text{ ms}$ ($\sim 105\text{--}210 \text{ m}$) below the palaeosurface. Most of the sills identified are of a similar scale to the lava flows in the area (2 – 4.5 km long, 1.2 – 1.7 km wide),

The Winnie 3D survey also hosts a spectacular single intrusion, 14 km long and 8 km wide, by far the largest igneous feature in the WVP (Fig. 8A & B). Classical intrusion related features such as inflation lobes are observed within spectral decomposition (Fig. 8C). ~ 20 vents cross-cut this intrusion (see pockmarks on Fig. 8C), in places leading to extrusive lava flows in the overlying sediments (see surface flows on Top Birkhead overlying the intrusion in Fig. 8B). The presence of the pockmarks on the surface of the intrusion suggest emplacement of the sill predated a later stage of volcanism that occurred towards the top of the Birkhead Fm.

4.1.3 Timing of Volcanism and Intrusion

The timing of igneous activity within the Winnie 3D survey was constrained by the stratigraphic relationship between igneous activity and the Triassic to Jurassic stratigraphy in the study area. In places, intrusions are observed to have caused forced folding of the overburden, as evidenced by onlap of Birkhead Fm. sediments onto the deformed overburden overlying the intrusions (Fig. 8B). Onlap of sedimentary rocks onto these forced folds was used to age constrain the intrusion (Trude *et al.*, 2003), whilst the strata-concordant surface flows were assigned an approximate age based on their stratigraphic level that was tied to known stratigraphy within the wells. We estimate that volcanism lasted throughout deposition of the Hutton and Birkhead formations between ~ 178 and 160 Ma, with the youngest

334 volcanics identified as surface flows emplaced on the top Birkhead marker horizon (Alexander
335 & Hibburt, 1996).

336 *4.1.4 Basement Structure and Igneous Rock Distribution*

337 The basin structure of the eastern Nappamerri Trough is largely obscured by the thick
338 Cooper and Eromanga Basin fill (the deepest Permian-age sedimentary rocks drilled in the
339 basin are 4,209 m deep (Beach Energy, 2013b)) and the highly reflective coals of the Permian
340 Toolachee and Patchawarra Fms that make imaging of deeper sediments difficult. Of all the
341 3D seismic reflection surveys available for this study, only the Winnie 3D survey imaged
342 reflections ~2800 ms deep within the Nappamerri Trough (approximately 4.3 km subsurface
343 depth based on the Charal-I time-depth relationship). Notably, the exact lithology of this
344 reflection is unclear, largely due to the lack of any intersections in the Central Nappamerri
345 wells. Unlike the flat-lying Cooper and Eromanga stratigraphy, these deep reflections were
346 often inclined at an angle of $\sim 10^\circ$ (taken with the scale set so that 1s TWT depth = 1 km
347 distance laterally). Although not confidently picked throughout the whole of the Winnie 3D
348 survey, the reflections can be mapped over an area of $\sim 600 \text{ km}^2$ (Fig. 9C). We interpret these
349 to represent the top Basement in the Nappamerri Trough.

350 In places, the reflections are visibly offset on the order of ~ 100 ms (Fig. 9A). When mapped
351 and in plan view, it is clear that the most major offset (a ~ 400 ms step in the data visible in
352 Fig. 9A) trends ~ 15 km NW-SE across the Winnie 3D survey (Fig. 9C). The same 400 ms step
353 has been noted by previous mapping (NGMA, 2003) where it is visible on a top basement
354 map in the eastern Nappamerri Trough. We interpret offset of the top basement as faulting
355 and, so that it may be easily referred to throughout the text, name the major 400 ms offset
356 the Central Nappamerri Fault.

357 Beneath the vents and volcanoes of the Winnie 3D survey, the seismic reflection data
358 exhibited a velocity pull-up effect, where reflection events directly beneath the vents were
359 pulled up in two-way time relative to the surrounding area (Fig. 9A). This could either
360 represent shallow high velocity material (i.e., the vents or volcanoes themselves) or a high
361 velocity feeder zone representing a vertical column of igneous rock or hydrothermally altered
362 host rock. This effect was particularly obvious on the well-imaged Toolachee Fm. and, hence,
363 the Toolachee Fm. was mapped in order to better constrain the distribution of the pull-up
364 effects throughout the area (Fig. 9C). Here, the distorted zones are imaged as dark patches
365 that are clearly distinguished from the fluvial channels that make up the Toolachee Fm. (Fig.

366 9C). We have interpreted these as vents that are either hydrothermal vents or basaltic dykes.
367 The position of these vents was highlighted and superimposed on top of an image of the
368 basement reflections (Fig. 9C). The vents broadly fall into two groups, one to the SW of the
369 Central Nappamerri Fault and one to the NE of the Central Nappamerri Fault. These groups
370 fall adjacent to the Top Basement reflections or directly above visible offset in the basement,
371 strongly implying a pervasive structural control on the location of vents in the Winnie 3D
372 survey.

373 **4.2 Gallus 2D**

374 The Gallus 2D seismic reflection dataset was shot in 2013 by Beach Energy. The survey covers
375 420 km² of the Nappamerri Trough, to the north of the Winnie 3D survey (Fig. 11). It was
376 used to constrain the northern extent of the WVP. Although a dominant frequency of ~35
377 Hz and average velocity of 5 kms⁻¹ for the igneous rocks in the area suggests a vertical
378 resolution of ~40 m and a detection limit of 5.5 m, imaging was generally poorer than that of
379 the Winnie 3D survey. This was probably due to the inferior signal-to-noise ratio of stacked
380 2D seismic reflection data by comparison with that of an equivalent section line extracted
381 from stacked 3D data. In particular, the 2D nature of the survey and, hence, the inability to
382 conduct attribute analysis made the mapping of igneous rocks difficult.

383 A northwards continuation of the monogenetic vents, flows and intrusions that were detailed
384 in the Winnie 3D survey is imaged within the Gallus 2D survey (Fig. 11A-C). In order to
385 better constrain the distribution of the volcanics in the Gallus 2D survey, the top and bottom
386 of the igneous rocks were mapped. From these, we calculated the isochron thickness of the
387 igneous rocks in the area (Fig. 11A). Like the Winnie 3D survey, volcanism was more pervasive
388 towards the east, with the thickest volcanics (~300 ms, approximately 163 m) located at the
389 eastern edge of the survey. Whilst none of the volcanics were intersected by any wells, the
390 approximate stratigraphic age was constrained by the nearby Halifax and Ety wells (Fig. 11C),
391 placing the volcanics within the Hutton and Birkhead formations, as found for the Winnie 3D
392 survey.

393 **4.3 Madigan 3D**

394 The Madigan 3D seismic reflection survey was shot in 1996 by Santos Ltd. The Madigan 3D
395 survey covers an area of 584 km² over the Challum Field, which consists of gas reserves
396 hosted within the Toolachee Formation and pre-Permian carbonates (Yew & Mills, 1989)(Fig.

397 10). The survey is situated across the very north eastern edge of the Nappamerri Trough,
398 above the JNP trend, adjacent to the confluence point of the JNP trend and the Innamincka
399 Ridge. The dominant frequency within the Jurassic succession was ~ 35 Hz, and using an
400 average velocity of 5 km s^{-1} a vertical resolution of ~ 35 m and a detection limit of 5 m was
401 calculated.

402 Two vents were identified within the seismic data that consisted of a single positive amplitude
403 reflection (Fig. 10). The smaller vent ($750 \text{ m} \times 1.5 \text{ km}$) is situated directly above the JNP trend.
404 The second is ~ 5 km to the SW and larger, at $1.5 \text{ km} \times 1.5 \text{ km}$, with a central 500 m
405 depression interpreted to be a central crater (Fig. 10). RMS amplitude extraction reveals the
406 central crater of both vents (Fig. 10). As neither vent was penetrated by a well and no surface
407 flows were mapped away from the vents, it cannot be concluded whether they are volcanic
408 or hydrothermal. Stratigraphic data from the neighbouring Challum wells were used to
409 constrain a rough stratigraphic age for the vents, placing them towards the base of the Hutton
410 Sandstone (~ 178 - 200 Ma). It is unclear on the basis of the 3D seismic reflection data what the
411 age of these vents is. They appear to sit near the top of the Nappamerri Group, however, this
412 would place them in the middle Triassic ~ 240 Ma, ~ 60 Myrs older than the rest of the WVP.
413 An alternative explanation would be the emplacement of these vents at the very base of the
414 Jurassic in this region (a maximum age of ~ 200 Ma, Fig. 2), as they appear to cross-cut all of
415 the Triassic Nappamerri Group stratigraphy. This second explanation is preferable as it would
416 fit with the proposed age for the rest of the WVP.

417 **5. Description of Igneous Rocks in Well Data within the Eastern Nappamerri** 418 **Trough**

419 **5.1 Extrusive Igneous Rocks**

420 *5.1.1 Lambda I*

421 The Lambda I well was drilled in 1984 by Delhi Petroleum Pty. Ltd. At a depth of 1570.9 m,
422 283 m of igneous rocks were intersected, directly underlying the base of the Birkhead Fm.
423 (Fig. 12). The upper 33 m of basalt are heavily weathered, fractured and vesicular, as noted in
424 the Lambda well report (Short, 1984). This is represented by poor seismic imaging due to the
425 low density and acoustic velocity of the upper 33 m relative to the rest of the volcanic
426 succession (Fig. 12A). The remaining 250 m of basalt is fresh and crystalline. It typically has

427 low gamma ray values, with consistently high density and acoustic velocity (Fig. 12A). There
428 are localised areas where the density drops drastically, however, these are associated with
429 increases in the caliper and may therefore point to caving of the wellbore when it was drilled
430 or fractures within the basalt, rather than lithological variations. At the base of the basalt ~1.5
431 m of core was cut, within which the fine grained, crystalline nature of the basalt is apparent
432 (Fig. 12A, see picture).

433 Interval velocity histograms were created for the Lambda I volcanics. They were then binned
434 with a bin spacing of 0.125 m order to create a velocity histogram (Fig. 12C). These were
435 superimposed on the velocity histogram fields of Nelson *et al.*, (2009) that were obtained
436 from boreholes on the Faroe Islands in the North Atlantic in order to compare the acoustic
437 velocities of igneous rocks in the Nappamerri Trough with values from elsewhere. It can be
438 seen in Fig. 12C that the Lambda I volcanics show a very strong similarity to the tabular basalt
439 analysed by Nelson *et al.* (2009).

440 Additionally, vitrinite reflectance data (Boothby, 1986) were available for sediments
441 deposited on top of the Lambda-I volcanics (Fig. 12B). Notably, in the sediment contact with
442 the volcanics, there is no deviation towards high %RoMax, suggesting that the Lambda I
443 volcanics had little thermal effect on the overlying organic matter. If this were an intrusion, it
444 is expected that the VR values would be elevated (Stewart *et al.*, 2005) and the absence of
445 such elevated values implies that the basalt is extrusive.

446 K-Ar dating was conducted on a bag of drill cuttings taken from a depth of 1658 m. Although
447 the samples were too altered to be used for total rock analysis, fresh plagioclase within the
448 cuttings was separated for analysis and used to determine an age of 227 +/- 3 Ma, suggesting
449 emplacement of the basalt during the early Triassic (Murray, 1994). However, in Lambda-I,
450 the basalt is situated between the Jurassic Birkhead Fm. and the Triassic Nappamerri Gp.,
451 supporting a Late Triassic to Jurassic age (Draper, 2002). K-Ar dating of basaltic volcanic
452 rocks is often unreliable (Schofield *et al.*, 2015), due to unquantified argon loss (Kelley, 2002).
453 It is noted within the well report that the loss of Argon due to subsequent thermal events or
454 contamination of drill cuttings is not accounted for in the quoted error (Murray, 1994). In
455 combination with the extensive seismic evidence for Middle to Late Jurassic age volcanics, we
456 deem a Jurassic age of eruption most likely.

457 *2D Seismic across the Lambda I well*

458 A single 2D seismic line intersecting the Lambda I well was available (Fig. 13). Within the 2D
459 line, the tabular basalt intersected by the Lambda I well corresponds to a high amplitude
460 reflection most pronounced at the location of the well but also continuing ~3 km to the west
461 of the survey. The centre of the survey (and the location of the well) corresponds with a ~1.5
462 km wide anticline that exhibits doming on the order of ~150 ms.
463 Below the anticline and the location of the volcanics, the 2D seismic reflections are visibly
464 distorted, being most pronounced beneath the location of the well where a velocity pull-up
465 effect, similar to those observed below the igneous rocks in the Winnie 3D survey, is visible.
466 We interpret the Lambda I volcanics to be a central vent drilled by the well, with a lava flow
467 extending 3 km to the west of the survey.

468 Stratigraphically, the Birkhead Fm. extends across the whole survey, however, the Hutton
469 Sandstone and equivalent sediments (not penetrated by the Lambda I well) are seen to onlap
470 the Lambda I volcanics. The lava flow extending away from Lambda I sits atop Triassic
471 Nappamerri Group sediments of the Cooper Basin whereas the volcanics penetrated by
472 Lambda I appear to cross cut the Triassic stratigraphy. In combination with the well data, we
473 believe a Hutton to Birkhead Fm. age (178 – 163 Ma) for the volcanics to be most likely,
474 although we cannot discount that some of the volcanics may be Late Triassic in age.

475 *5.1.2 Orientos 2*

476 Similarly to Lambda I, Orientos 2 drilled through volcanics situated at 1555 m, beneath the
477 Jurassic age Adori Sandstone (Fig. 14). The well drilled through 34 m of volcanics before
478 drilling was ceased, before reaching the base of the volcanics. The volcanics are described as
479 dark green, grey black with quartz inclusions and common amphiboles. Similar to the Lambda
480 I well, Orientos 2 also encountered ~14 m of chloritic, weathered basalt identifiable by low
481 density and sonic values at the top of the volcanic succession and by the cuttings described in
482 the end-of-well report (Bucknill, 1990). Draper (2002) noted that Orientos 2 was drilled <1
483 km from Orientos 1, which did not intersect basalt, suggesting that the basalt in Orientos 2
484 might be intrusive in origin. However, we note that the Padme, Anakin and Charal wells were
485 all drilled less than 2-5 km from what appear to be extrusive volcanics identified in the Winnie
486 3D survey, but due to the localised nature of the volcanics, they are not noted in any of the
487 well logs. As such, we do not take the absence of volcanics in Orientos 1 to be indicative of
488 an intrusive origin for the Orientos 2 volcanics. Due to the similarity in their log signature and

489 the lithological information recorded in the end-of-well report (Bucknill, 1990), we believe
490 Orientos 2 drilled through a series of Jurassic-aged extrusive basalts.

491 *2D Seismic across the Orientos 2 well*

492 A single 2D seismic reflection survey line running south to north across the Orientos 2 well
493 path was available. High amplitude reflections running across the survey appear similar to the
494 extrusive volcanics interpreted in the Winnie 3D survey, with a ~ 1 km reflection penetrated
495 by Orientos 2 and a ~ 4 km reflection ~100 m to the north of the Orientos 2 reflection. Like
496 the Lambda I volcanics, we interpreted these to be two lava flows, although vents visible
497 below the flows suggest these flows may be made up of multiple eruptions.

498 Stratigraphically the flows sit on top of the Triassic Nappamerri Group and are onlapped and
499 age equivalent to the Birkhead and Hutton Sandstone formations (178 – 160 ma). However,
500 the northernmost vent and the Orientos 2 vent are situated directly below the Adori
501 Sandstone raising the possibility of some volcanic rocks in the area that are younger than 160
502 Ma.

503 **5.2 Intrusive Igneous Rocks**

504 *5.2.1 Warnie East I*

505 Igneous rocks intersected within the Warnie East I well are lithologically unique amongst the
506 igneous rocks penetrated by wells in the WVP. Located at 2103 m depth within the Permian
507 Toolachee Fm, the 65 m thick basalt is fine grained and dominated by plagioclase laths
508 intergrown with augite (Fig. 16). The upper 12 m of the volcanics appear altered and
509 calcareous, containing phenocrysts of augite with minor pale green talc and minor calcite
510 veining (Boothby, 1986).

511 The age of the igneous rocks is unclear based on available data. Palynological data indicates
512 that sedimentary rocks immediately overlying igneous rocks in the Warnie East I well are
513 Permian in age (belonging to the Toolachee Fm.). Furthermore, sedimentary rocks below the
514 basalt to 2228 m are barren of palynological data. Like Lambda-I, K-Ar dating was conducted
515 on a sample of drill core from 2163 m (Murray, 1994). Again, the rock was too altered to be
516 used for whole rock K-Ar and, instead, unaltered pyroxene was used for dating. A middle
517 Cretaceous age of 100 ± 9 Ma was calculated (Murray, 1994), significantly younger than the
518 rest of the WVP and the Permian host rock within which the intrusion is located. Draper

519 (2002) found the age dating to be 'equivocal' due to the large difference in age for the Lambda-
520 I extrusive basalt ($227 \pm 3\text{Ma}$) and the Warnie East intrusive basalt ($100 \pm 9\text{Ma}$)(ages from
521 Murray, 1994). It is therefore important consider the K-Ar ages and the stratigraphical
522 position of the igneous rocks with respect to other data.

523 Vitrinite reflectance data were acquired within the Warnie East I well. (Fig. 16B). Samples of
524 vitrain were obtained and mean maximum reflectance of vitrinite calculated (Smith, 1987).
525 Vitrinite reflectance samples adjacent (directly above and below) to the basalt show a marked
526 increase to 4.5% relative to surrounding Toolachee sediments that are typically 1.5-1.75% (Fig.
527 16). Whereas extrusive volcanics appear to have little thermal effect on the surrounding
528 sediments (e.g. Grove, 2014), igneous intrusions typically show uniform heating on either side
529 of the volcanics, as is observed here. Coupled with its stratigraphic position within the
530 Toolachee Fm., where no other volcanics have been documented within the Cooper Basin,
531 we conclude that Warnie East I contains the sole identified intersection of an intrusion in the
532 WVP.

533 *2D Seismic across the Warnie East I well*

534 As with the previous two wells, a 2D seismic reflection survey line runs across the Warnie
535 East well (Fig. 17). Unlike the Lambda I and Orientos 2 volcanics, the igneous rocks within
536 the Warnie East I well sit within the Toolachee Fm., a highly reflective sequence of
537 heterogeneous lithologies consisting of sandstones, siltstones and coals. As such, the igneous
538 rocks are difficult to distinguish on seismic reflection survey data, although a shallow saucer-
539 shaped reflection intersects the well path of the Warnie East I well at the depth that igneous
540 rocks are located. The reflection is $\sim 1.5\text{ km}$ wide and transects the Toolachee Fm. over a
541 depth range of 50 ms. The shallow saucer shape that the reflection event exhibits is common
542 among igneous intrusions within sedimentary basins, corroborating the well data that suggests
543 the igneous rocks are intrusive in nature.

544

545 **6. Description of Anomalous Igneous Rocks within the southern Nappamerri** 546 **Trough**

547 This study has so far dealt with igneous rocks of a very consistent lithology preserved within
548 or above the Central Nappamerri Trough. However, the Kappa I well and Snowball 3D

549 seismic reflection survey, located on the very southern edge of the mapped extent of the
550 WVP, preserve igneous rocks of a very different character (Figs. 18 -20).

551 **6.1 Kappa I Well**

552 The Kappa I well was drilled in 1997 on behalf of Santos Limited (Allen, 1998). The aim of
553 the well was to evaluate the Toolachee, Patchawarra and Epsilon formations (Allen, 1998).
554 Kappa I is located above a large anticline and is the southernmost of a string of prospects
555 located above the Kinta structural trend, which deepens northwards into the Nappamerri
556 Trough (Fig. 1)(Allen, 1998). Between 1895 and 2115 m, the Kappa well intersected a 120 m
557 thick succession of volcanics below a thinned 21 m succession of Hutton Sandstone (Fig. 18).
558 Within the well report, the volcanics are ascribed a late Permian to Middle Triassic age (Allen,
559 1998).

560 The primary lithological information for the volcanics comes from assorted rock chips and
561 cuttings. Due to the large amount of alteration, the volcanics appear as a chloritized basalt
562 with abundant plagioclase, chlorite, carbonates and quartz (Allen, 1998). The rock is mottled
563 and light green to a reddish brown colour (Allen, 1998). However, some variation was noted.
564 Firstly, a vesicular microporphyritic basalt was identified that consisted of olivine and
565 pyroxene phenocrysts contained within a fine grained or glassy groundmass (Allen, 1998).
566 Secondly, less abundant coarse grained holocrystalline basalt that was composed of randomly
567 oriented plagioclase laths, ferromagnesian crystals and accessory magnetite plus ilmenite
568 (Allen, 1998). Importantly, and pervasively throughout the succession, a fibrous chlorite has
569 replaced much of the ferromagnesian minerals and groundmass as well rimming and filling
570 vesicles within the basalt (Allen, 1998).

571 The Igneous succession in Kappa I also has an unusual petrophysical expression. In our
572 interpretation, we have divided the volcanics into two facies; in situ basalt and a volcanoclastic
573 breccia (Fig. 18). The in situ basalt (e.g. the section between 1900 m and 1960 m (Fig. 18)) has
574 a 'saw-tooth' response with a relatively low gamma response of ~40. Acoustic velocities from
575 the sonic velocity log (~4 kms⁻¹) are consistently lower than those observed in the Lambda I
576 and Orientos 2 wells (~ 5.5-6 kms⁻¹). The chaotic, saw-tooth response is indicative of the
577 highly altered nature of the volcanic succession in Kappa I and the compound nature of the
578 lava flows (Millet *et al.*, 2015; Hardman *et al.*, 2018a) (Fig. 18). The resistivity throughout the
579 section interpreted to be in situ volcanics consistently measures 15 Ωm. As with the Lambda

580 well, we constructed a velocity histogram for the igneous rocks in the Kappa I well. When
581 compared to Nelson *et al.*'s velocity histograms (2009) the best match was with that of
582 Compound-braided flows (Fig. 18).

583 The volcanoclastic breccia typically consists of poorly sorted, subangular, volcanoclastic igneous
584 rock of very fine sand to granule size (Allen, 1998). The log response for the volcanoclastic
585 breccia is characterised by a jagged petrophysical response when compared to the in situ
586 basalt. Although the gamma ray values (~40 api) are similar to those of the in situ basalt, the
587 acoustic velocities (between 2800 and 4 kms⁻¹) are consistently lower. Furthermore, the
588 volcanoclastic breccia is water saturated with a resistivity of 2 – 3 Ωm, although this may also
589 be a reflection of the magnetite and ilmenite within the basalt. Finally, below 2050 m to the
590 base of the volcanic succession, the Caliper tool widened significantly suggesting the
591 volcanoclastics in this part of the succession are considerably unconsolidated and/or fractured.

592 **6.2 Snowball 3D Seismic Survey**

593 The Snowball 3D seismic reflection survey is situated ~ 5 km south of the southwest tip of
594 the Winnie 3D survey, on the border of Queensland and South Australia. Unfortunately, the
595 Kappa I well is located ~2 km outside of the Snowball 3D survey data that was available to
596 used within this study. Furthermore, no sonic velocity data were acquired for the Kappa I
597 well. In order to tie this well to the Snowball 3D survey, a pseudo TWT/depth plot was
598 created using the velocity data synthesised for the wells drilling within the central Nappamerri
599 Trough (Fig. 19)(Halifax, Ety, Anakin, Padme, Charal).

600 The Snowball 3D seismic reflection data are displayed such that a downward increase in
601 acoustic impedance corresponded to a red, negative reflection (European polarity). The
602 dominant frequency within the Jurassic succession was ~35 Hz, however, the average velocity
603 of the igneous rocks in the Kappa I survey was 4 kms⁻¹, lower than the other igneous rocks
604 in the WVP. Hence, a higher than average vertical resolution (lower resolvable thickness) of
605 ~29 m and a detection limit of 4 m were calculated.

606 The top of the Kappa I volcanics is interpreted to correlate with a bright, laterally continuous
607 red reflection that was present across the whole of the Snowball 3D survey (Figs. 19 & 20).
608 Above the top Kappa I volcanics a second reflector, deemed here the 'Top Volcanic Mounds,'
609 was mapped in parts of the Snowball 3D survey adjacent to a series of mound-like structures,
610 appearing in places to downlap onto the top Kappa I volcanics (Fig. 20). The centres of these-

611 mound shaped structures are marked by a notable brightening of the reflector that is underlain
612 by vertical vents similar to those seen in the other surveys. We thus interpret these mound
613 structures to be a series of volcanic edifices.

614 The volcanics interpreted to be present in the Snowball 3D survey are notably different to
615 those preserved in the Nappamerri Trough (e.g. Figs 7-9). Rather than isolated, monogenetic
616 vents or flows, they are interpreted to represent a thick package (up to ~150 ms) of mixed
617 volcanics and volcanoclastic breccias (based on evidence from Kappa-I) that thin towards the
618 southwest of the survey (Fig. 20A) (i.e. the southern edge of the Nappamerri Trough).
619 Although the true extent of them is not imaged within the survey (they are interpreted to
620 extend beyond the survey limits), laterally they extend over 6 km. Spectral decomposition
621 was used to gain a better understanding of these volcanics in 3D (Fig. 20C). We found darker
622 colours surrounding the vents with these colours interpreted to represent the extent of the
623 lava flow fields.

624 **6.3 Interpretation of Snowball Volcanics**

625 The volcanic succession is described as pyroclastic in the Kappa-I well report. Whilst this
626 interpretation can account for some of the altered volcanoclastics, it does not explain the
627 extensive sequence of altered basalt. An alternative explanation for the extensive sequence
628 of altered basalt is differing sedimentation and eruption rates within the Eromanga succession
629 deposited above the Nappamerri Trough. However, Mid-Late Jurassic sedimentation rates at
630 Kappa-I of 2.4 m.my^{-1} (calculated by dividing the thickness of the Hutton to Adori succession
631 by a duration of 23 Myrs) do not appear materially different to those in the Warnie or Lambda
632 wells (2.8 and 2.5 m.my^{-1} respectively). Hence, alteration does not appear to be a consequence
633 of extensive surface exposure.

634 A more compelling explanation for the Kappa I volcanics is hydrothermal alteration. In
635 volcanic regions, many lakes are typically fed by hot springs. In the East African Rift, discharge
636 of hydrothermal water can produce authigenic chlorite with calcite and quartz forming pore
637 filling cements (Renaut *et al.*, 2002). These minerals, particularly chlorite, are noted extensively
638 within the petrological description of the Kappa I well. Additionally, the Kappa I well lies
639 above the Kinta structural trend which could have provided a pathway for hydrothermal fluids
640 into the Eromanga succession (Kulikowski *et al.*, 2015). Basin-wide hydrothermal fluid
641 circulation has been noted within the Cooper Basin (Zwingmann *et al.*, 2001; Middleton *et al.*,

642 2014, 2015). In the Nappamerri Trough, these fluids are believed to have formed due to rifting
643 of the eastern Australian margin during the mid-Cretaceous (Middleton, 2015). We propose
644 that the Kappa I well represents a succession of volcanoclastic and primary volcanic rocks that
645 was highly altered by hydrothermal fluids.

646 **7. Airborne Geophysical Surveys of the Nappamerri Trough**

647 **7.1 Gravity Data**

648 Alongside seismic and well data, airborne geophysical surveys were examined in order
649 to further delineate the regional distribution of the WVP (Fig. 21A). The gravity data in this
650 study were adopted from Tracey & Bacchin (2008). In southern Queensland, the Nappamerri
651 Trough is typically characterized by a broad gravity low. However, a pronounced ~50 x 50
652 km gravity high is apparent in the northeast Nappamerri Trough (Fig. 1B). Gravity modelling
653 of the Cooper Basin suggests that localised gravity lows adjacent to the 50 x 50 km high within
654 the Nappamerri Trough are the result of low density granites and granodiorites within the
655 basement (Meixner *et al.*, 2012). Furthermore, the quantity of mapped intrusive and extrusive
656 volcanics increases towards the centre of the 50 x 50 km gravity anomaly (Fig. 1B). We
657 interpret the 50 x 50 km gravity anomaly to represent an expression of the WVP, as it
658 highlights an increase in the volume of igneous rocks towards the centre of the Nappamerri
659 Trough and perhaps also a weak point between granitic bodies that facilitated the transport
660 of magma into the Eromanga succession. Our lack of constraint on the exact size of the WVP
661 means we cannot confirm that the thickness and areal extent of the igneous rocks are
662 sufficient to produce the size of the gravity anomaly observed in the Nappamerri Trough.

663 **7.2 Magnetic Data**

664 In addition to gravity data, a total magnetic intensity survey has been analysed (Fig. 22).
665 Previous investigations of magnetic anomalies in the Cooper Basin by Meixner *et al.* (2000)
666 have concluded that the sources of magnetic anomalies all occur at or below the base of the
667 Cooper Basin succession, due to the absence of any significant magnetic sources within the
668 sedimentary sequences of the Cooper and Eromanga basins. The Nappamerri Trough is
669 generally marked by a low in the magnetic intensity whilst the bounding ridges are magnetic
670 highs. However, a ~75 x 40 km magnetic intensity high runs from the Innamincka Ridge across
671 the northern extent of the WVP in a NW-SE direction, with many smaller (<8 km² anomalies)
672 scattered throughout the Nappamerri Trough (Fig. 22A-D). Basalt is highly susceptible to

673 being magnetised (Tarling, 1966), which has facilitated the recognition of buried volcanic rocks
674 elsewhere in the world (Segawa *et al.*, 1975). Notably, many of the smaller magnetic anomalies
675 do crudely correspond to the location of lava flows in the Winnie 3D survey (Fig. 22D).
676 However, the large 75 x 40 km anomaly does not correspond with extrusive igneous rocks,
677 otherwise this would have been imaged in the northern Gallus 2D lines. Without good seismic
678 imaging below the Cooper Basin sediments, or boreholes that drill down through the entire
679 Cooper Basin succession, it is difficult to confidently identify what this magnetic anomaly
680 represents. We postulate that this could be an expression of the plumbing system of the
681 WVP, although this is highly speculative. If true, it could point to an extensive network of
682 igneous intrusions that are not imaged by available data.

683 **8. Discussion**

684 **8.1 The Age of the Warnie Volcanic Province**

685 Within this manuscript we have detailed the rough age of igneous rocks within the Cooper
686 and Eromanga basins. Here we synthesise the different methods used to define an age range
687 for the eruption of the igneous rocks.

688 K-Ar dating was conducted on both the Lambda I extrusive basalt and the Warnie East
689 I intrusive basalt deducing an age of 227 ± 3 and 100 ± 9 Ma, respectively (Murray 1994).
690 However, the large differences in age for these igneous rocks led Draper (2002) to question
691 the validity of the results suggesting they were “equivocal.” Furthermore, Murray (1994)
692 noted that loss of argon due to subsequent thermal events was not accounted for in the
693 quoted error of the ages. K-Ar dating has been shown to be unreliable due to unquantified
694 argon loss (Kelley, 2002; Schofield *et al.*, 2015) and in combination with the altered lithology
695 of many of the igneous rocks within the Warnie Volcanic Province (Murray, 1994) we cannot
696 consider the K-Ar dating to be reliable in the absence of further evidence for the age of these
697 rocks.

698 Stratigraphically, all the extrusive igneous rocks analysed within well data are situated between
699 Jurassic sedimentary rocks (the Hutton and Birkhead fms.) and the Triassic Nappamerri
700 Group. which supports a late Triassic to Jurassic age (Draper, 2002). The corroboration of
701 well data with 3D seismic reflection data has allowed us to define further intrusive and
702 extrusive volcanic events that are strata-concordant with the Birkhead Fm. and Hutton
703 Sandstone, placing them between ~178 and 160 Ma. The oldest extrusive volcanics identified

704 using 3D seismic reflection data are the vents within the northern Challum 3D survey that
705 appear to sit near the top of the Nappamerri Group below the Hutton Sandstone. This could
706 extend the age of the Warnie Volcanic Province into the Triassic which would agree with the
707 K-Ar dating for the Lambda 1 well. Equally, the vents could be interpreted as cross-cutting
708 the Triassic strata which would give them an Early Jurassic age. In the absence of further data,
709 their exact age is unclear.

710 We conclude that the vast majority of igneous rocks identified using 3D seismic reflection
711 data and well data throughout this study have been constrained to a broad age range between
712 ~178 and 160 Ma spanning the Middle Jurassic, however, further work is encouraged to better
713 define the age of the Warnie Volcanic Province.

714 **8.2 Structural Controls on the Emplacement of the Warnie Volcanic Province**

715 From the late Triassic to the early Jurassic, the Eromanga Basin was subject to an east-west
716 compressional event named the Hunter-Bowen event (Mayromatidis, 2005; Kulikowski &
717 Amrouch, 2017), which uplifted the basin and inverted the major highs. Subsequently, the
718 basin underwent post-compressional flexural relaxation with the resultant accommodation
719 space infilled by the Middle to Late Jurassic succession discussed within this manuscript (Lowe-
720 Young, 1997; Mayromatidis, 2005). During this time, a strike-slip extensional regime is thought
721 to have been present (Kulikowski, 2018). In this context, it is interesting to consider whether
722 the stress regime of the Cooper Basin helped to control the location and morphology of
723 igneous rocks within the WVP and, more broadly, the location of the WVP within central
724 Australia.

725 **8.2.1 The Influence of Basement Structure on the Location of Igneous Rocks**

726 Within other basins, such as those in the North Atlantic, it has been noted how the location
727 of igneous intrusions and volcanism is controlled by the basin structure, in particular the
728 location of major faults (Schofield *et al.*, 2015; McClean *et al.*, 2017; Hardman *et al.*, 2018a).
729 To investigate whether basement structure and faulting had a direct control on volcanism
730 erupted during the Jurassic, we have utilised the 3D seismic reflection datasets available.

731 We have already considered the Top Basement structure within the Winnie 3D survey and
732 interpreted faulting of the Top Basement, in particular the 25 km long Central Nappamerri
733 Fault. We superimposed the location of vents within the Winnie 3D survey on top of a TWT

734 map of the Top Basement to see if there was any relationship between basement structure
735 and the location of the vents (Fig. 9C). On a broad scale, two vents are located directly above
736 the Central Nappamerri Fault (CNF) with the others situated adjacent to the mapped Top
737 Basement Horizon (i.e. where it deepens to > 3600 ms and is no longer mappable). On a
738 broad scale, the location of vents does appear to be related to the basement structure.

739 We also superimposed the location of the Madigan 3D volcanoes on top of the NGMA (2003)
740 basement map (Fig. 25). Like the igneous rocks in the Winnie 3D survey, the two volcanoes
741 imaged in the Madigan 3D survey sit atop the tip of the 25 km Challum Fault and atop a
742 second, ~5 km fault northwest of the main Challum Fault. This reinforces the idea that the
743 location of igneous rocks was largely controlled by the basin structure.

744 *8.2.2 The Influence of Discontinuities and Faults within the Eromanga Succession on the* 745 *Morphology of Igneous Rocks*

746 Here, we consider how basin structure controlled the morphology of igneous rocks within
747 the Nappamerri Trough. Using the Winnie 3D survey, the length, width and elongate direction
748 of the volcanoes, lava flows and intrusions was picked (Fig. 23). The high quality of the Winnie
749 3D survey facilitated confident mapping of the outline of individual volcanic events, however,
750 the Gallus 2D and Snowball 3D surveys were not considered as the outline of individual
751 volcanics could not be picked confidently. The data shows that igneous rocks in the area are
752 roughly/typically twice as long as they are wide (Fig. 23). Furthermore, almost all the igneous
753 rocks are elongated in a NW-SE direction, closely matching the strike of the faults within the
754 Nappamerri Trough. It is clear that during eruption and emplacement of the igneous rocks,
755 the basin structure exerted a strong control on the morphology of lava flows and igneous
756 intrusions.

757 However, we have also considered fracturing and faulting within the Eromanga succession and
758 its control on the morphology of lava flows and igneous intrusions in the Nappamerri Trough.
759 In vertical seismic reflection sections, faulting of the Eromanga Basin stratigraphy is not clearly
760 imaged within the Winnie 3D survey, largely due to the considerable noise within the section.
761 However, spectral decomposition produced for the Top Volcanics horizon images a number
762 of discontinuities (shown as black lines) within the colour blend (Fig. 24A). These
763 discontinuities are not thought to be a product of the acquisition and processing of the seismic
764 data, due to their non-linear nature and their coincidence with other geological features such

765 as the igneous rocks and the CNF. When mapped, the features form a series of largely NE-
766 SW trending lineaments (Fig. 24B). Although imaging of the discontinuities is best in the SW
767 of the survey, where there are fewer igneous rocks, it is evident that many of the igneous
768 rocks are aligned with these features suggesting they must have exerted some topographical
769 or structural control when the igneous rocks were emplaced or erupted over the Nappamerri
770 Trough (Fig. 24). As they do not resemble any sedimentary features such as valleys or fluvial
771 systems, we interpret these features to be faulting or fracturing of the Eromanga succession
772 above the Nappamerri Trough.

773 Determining the tectonic activity responsible for the faulting within the Nappamerri Trough
774 and, hence, the location and morphology of igneous rocks within the WVP is difficult due to
775 the complex structural and stress history of the Cooper Basin. Throughout the basin, a
776 conjugate set of large NNE-SSW and field scale SE-NW striking dextral strike-slip faults
777 developed under a SSE-NNW strike-slip stress regime (Kulikowski *et al.*, 2018). This is
778 pertinent when applied to the central Nappamerri Trough, as Kulikowski *et al.* (2018)
779 identified a series of vertical SE-NW structural lineaments with normal displacement or no
780 displacement. Due to their vertical nature, it is thought that these faults developed as strike
781 slip faults during successive periods of flexural relaxation or sag. The faulting identified within
782 this study is also determined to consist of little offset with an orientation that matches that
783 of the strike-slip faulting determined by Kulikowski *et al.*, (2008) (Fig. 24A-D). We surmise
784 that sag within the basin produced a series of faults and fractures within the basin that
785 controlled the morphology of igneous rocks in the WVP.

786 **8.3 The Plumbing System of the WVP**

787 Due to the risks associated with drilling igneous rocks within the Warnie Volcanic Province,
788 being able to identify the presence of them is especially important. Whilst this manuscript has
789 dealt with surface vents, flows and shallowly emplaced igneous intrusions (~100-200 ms below
790 the paleo land surface, here believed to be ~Top Birkhead), we have only briefly discussed
791 the plumbing system of the Warnie Volcanic Province and the lack of constraint on the
792 presence of dykes and deeper sills

793 Within this study we have already noted that vents are observed to extend through the
794 Cooper Basin succession, below the extrusive and intrusive igneous rocks identified in the
795 Eromanga Basin succession above (e.g. Fig. 9). Vents are positioned directly above visible offset

796 in the basement or adjacent basement highs (Fig. 9C). Furthermore, when viewed in a TWT
797 cross section, the vents are all vertical, with no sills observed >200 ms below the palaeo-land
798 surface (taken here to be Top Birkhead). This suggests that the magmatic plumbing system of
799 the WVP is dominantly vertical throughout the Nappamerri Trough.

800 However, whilst the intrusions identified within the Winnie 3D survey are emplaced at
801 shallow depths (~100-200 ms in Fig. 8) below the palaeo-land surface, the intersection of an
802 intrusion within the Permian succession of the Warnie East well (410 m below the Birkhead
803 Fm., Fig. 14) suggests that the plumbing system of the WVP is not well understood and may
804 be subject to lateral transport of magma, in addition to the predominantly vertical vents
805 observed in this manuscript. When detailing the magnetic data we discussed the importance
806 of a ~75 x 40 km magnetic intensity high that runs from the Innamincka Ridge across the
807 northern extent of the WVP in a NW-SE direction (Fig. 22B). Without good seismic imaging
808 of the stratigraphy below the Permian Cooper Basin succession, it is difficult to confidently
809 identify what this represents. Notably, it does not correspond to an increase in the
810 concentration of extrusive volcanics, otherwise this would have been imaged in the Gallus 2D
811 survey (Fig. 11A). This 75 x 40 km magnetic anomaly could be an expression of an extensive
812 network of igneous intrusions that are not imaged within the available data, however, a
813 confident interpretation would require high quality 3D seismic reflection survey data covering
814 the area north of the Winnie 3D survey.

815 **8.4 The Broader Record of Jurassic Volcanism in Eastern Australia**

816 We believe it is important to consider the WVP in the context of the broader record of
817 Jurassic volcanic activity in eastern Australia. Coeval with the eruption of the WVP, the
818 Eromanga Basin was undergoing post-compressional flexural relaxation with thermally
819 controlled subsidence in the absence of significant fault control (Gallagher & Lambeck, 1989
820 Mayromatidis, 2005). Subsidence has been attributed to subduction of the Pacific plate
821 beneath southern and eastern Australia, during which time rifting also initiated between
822 Australia and Antarctica (Griffiths, 1971; Johnstone *et al.*, 1973).

823 Evidence for Jurassic volcanic activity is pervasive throughout the sedimentary basins of
824 eastern Australia, manifested by volcanic arc-derived sediment within the Eromanga Basin
825 (Boult *et al.*, 1998), silicic tuffs in the Surat Basin (Wainman *et al.*, 2015) and widespread
826 volcanogenic zircons within the Eromanga Basin (Bryan *et al.*, 1997). This volcanoclastic

827 material is thought to be derived from an acid to intermediate volcanic arc positioned off the
828 coast of Queensland during the Jurassic however, the remnants of this arc volcanism are yet
829 to be identified (Patton, 1986, Boulton *et al.*, 1997). Jurassic alkali basalt has been intersected on
830 the Marion Plateau (~400 km east of Townsville in Queensland), although this may also be
831 related to rifting in the region during the Jurassic (Isern *et al.*, 2002).

832 ~1500 km south of the Marion Plateau, mafic and ultramafic breccia pipes and igneous
833 intrusions are noted in the Eastern Highlands and Sydney Basin with K-Ar dating suggesting
834 ages of between 193 and 163 Ma (McDougall and Wellman, 1976). To the north west of the
835 Sydney Basin, in New South Wales, lower Jurassic intrusions and volcanics of the Garawilla
836 alkali basalts and alkaline Mittagong complexes are noted in the Gunnedah Basin (195-177 Ma)
837 (Sutherland, 1973, Gurba & Weber, 2001). Further south of the Sydney and Gunnedah basins,
838 Jurassic volcanics are scattered throughout Victoria and Tasmania (Veevers, 1984). These
839 include voluminous tholeiitic intrusions in Tasmania (165 m.y.) that are thought to have been
840 buried under a thick pile of since eroded extrusive volcanics (Sutherland, 1973; McDougall &
841 Wellman, 1976).

842 ~1000 km to the north-west, the Wisanger Basalt is a ~170 Ma, 20 km fragmented lava flow
843 emplaced on top of Permian fluvial sediments on Kangaroo Island (McDougall & Wellman,
844 1976). Geochemically, it is similar to the Tasmanian and Antarctic tholeiitic dolerite and
845 basalts that are usually attributed to the Ferrar Large Igneous Province and extension that latterly
846 resulted in the separation of Australia and Antarctica (Compston *et al.*, 1968; Milnes *et al.*,
847 1982). Of note in South Australia, Jurassic kimberlites record the extension of Mesozoic
848 kimberlites found along the margin of southern Gondwana, above the subducting Pacific plate
849 (Tappert *et al.*, 2009), differing petrologically from the predominantly basaltic volcanism found
850 elsewhere in eastern Australia.

851 Plate reconstructions indicate that the Eromanga Basin was located at significant distances
852 (~750 km) from the oceanic trench of the subducting Pacific plate (Veevers, 1984) (Fig. 27).
853 Importantly, alongside the WVP, the South Australian Kimberlites are the only other Jurassic
854 volcanics documented in eastern Australia that occur away (>500 km) from a significant plate
855 boundary (here we have ascribed the Wisanger Basalts to the zone of separation between
856 Australia and Antarctica and, hence, a 'significant' plate boundary). We therefore believe the
857 Warnie Volcanic Province to be an intraplate volcanic province and, as the origin of intraplate

858 volcanism can often be contentious (Conrad *et al.*, 2011), it is important to consider the
859 source of this volcanism and implications for the formation of intraplate volcanism.

860 **8.5 The Origin of the Warnie Volcanic Province**

861 The WVP was located far from active plate boundaries during the Jurassic (Fig. 27A). When
862 considering the source of intraplate volcanism, several different mechanisms must be
863 considered:

- 864 • The presence of a mantle plume
- 865 • Asthenospheric upwelling due to extension
- 866 • Local scale mantle convection

867 Mantle plumes are typically associated with the eruption of voluminous flood basalt provinces
868 that often mark the earliest volcanic activity of major hot spots (Richard *et al.*, 1989).
869 Furthermore, mantle plumes cause dynamic uplift of the land's surface of up to several
870 hundred kms followed by surface subsidence due to the withdrawal of mantle plume material
871 and loading of the crust with the volcanic sequence (Nadin *et al.*, 1997; Dam *et al.*, 1998; Leng
872 & Zhong, 2010; Hartley *et al.*, 2011; Hardman *et al.*, 2018b). No evidence of major surface
873 uplift in the Cooper and Eromanga Basins is coeval with the emplacement and eruption of the
874 WVP (Hall *et al.*, 2015). Furthermore, subsidence and sag throughout the emplacement of the
875 WVP is minor, with no large changes in sedimentary facies noted throughout the Jurassic
876 lower non-marine Eromanga succession (Alexander & Hibburt, 1996). No major crustal
877 and/or lithospheric extension is evidenced by the lack of major Jurassic faults, during a time
878 in which the basin underwent post-compressional flexural relaxation in a strike slip
879 extensional regime (Lowe-Young, 1997; Mayromatidis, 2005; Kulikowski, 2018). Furthermore,
880 the small spatial area covered by the Warnie Volcanic Province (~7500 km² within the
881 Nappamerri Trough area) and, by extension, the volume of magma emplaced is too low to be
882 related to a mantle plume source (Conrad *et al.*, 2011).

883 Low-effusive volcanism occurring within tectonic plates has been also attributed to
884 several locally operating processes, such as minor upwelling plumes (Courtilot *et al.*, 2003;
885 Conrad *et al.*, 2011), downwelling drops (Elkins-Tanton *et al.*, 2006) and sub-lithospheric
886 (Ballmer *et al.*, 2009) or edge-driven convection (Conrad *et al.*, 2010). In the case of edge-
887 driven convection, mantle flow can induce upwelling and volcanism through interaction with
888 lithospheric or asthenospheric heterogeneities (Conrad *et al.*, 2011; Davies & Rawlinson,

889 2014). In the Newer Volcanic Province in South Australia, <5 kyr volcanism forming as a
890 product of edge driven convection is noted to consist of basaltic monogenetic volcanoes <4
891 km² in size covering an area of ~20,000 km² (Demidjuk *et al.*, 2007; Davies & Rawlinson, 2014),
892 bearing a striking similarity to the WVP (the volcanism is basaltic and individual volcanoes like
893 Mount Gambier are of a similar scale (~1.5 x 1.5 km) to the volcanoes imaged in the
894 Nappamerri Trough area). Although not much is known about the crustal structure below
895 the eastern Nappamerri Trough, seismic reflection profiles have revealed that Devonian
896 troughs to the east of the study area are associated with crustal thinning of 7 – 10 km (Mathur,
897 1983). Furthermore, on a continental scale, the Moho is observed to shallow to ~20 km below
898 the Cooper Basin (and more broadly, Central Australia) (Kennett *et al.*, 2011). Therefore,
899 there is observable lithospheric thinning below the basins in Central Australia that could have
900 produced edge driven convection.

901 However, because Central Australia was relatively inactive during the Jurassic, we
902 must consider driving mechanisms for asthenospheric flow beneath the continent during the
903 eruption and emplacement of the WVP. As we have already detailed, the Pacific plate was
904 subducting beneath eastern Australia during the Jurassic. In intraplate volcanic provinces
905 adjacent to subduction zones (e.g. Western North America, China), there is a causal link
906 between subduction and intraplate volcanism, in that subduction below the continental crust
907 acts as a driving force for asthenospheric shear and, hence, the mantle upwelling that produces
908 intraplate volcanism (Conrad *et al.*, 2011; Tang *et al.*, 2014; Zhou *et al.*, 2018). If the extension
909 in the area during the Jurassic is not great enough to stimulate volcanism, then an alternative
910 model may involve rapid asthenospheric shear produced by the subducting Pacific Plate,
911 localised beneath SW Queensland due to edge driven convection.

912 We propose a model for the Warnie Volcanic Province based on our understanding of the
913 lithospheric and geodynamic state of Central Australia in the Middle to Late Jurassic. We
914 propose that asthenospheric shear above the subducting Pacific plate stimulated mantle flow
915 below Australia. Localisation of mantle flow occurred beneath southwest Queensland because
916 of edge driven convection leading to emplacement of the Warnie Volcanic Province above
917 the Nappamerri Trough. Despite our proposed model, we strongly believe that more work
918 needs to be conducted on the area before it can be concluded what the source of the WVP
919 is. In the Newer Volcanic Province of southeastern Australia, geochemistry, in particular major
920 and trace element analysis, provided insights into why mantle melting occurred in the absence
921 of extension or a mantle plume (Demidjuk *et al.*, 2007). As such, we would strongly

922 recommend further studies to be conducted on the geochemical signature of the basalts of
923 the WVP.

924 Due to its low volume and small areal extent, we do not believe that the WVP was
925 the source of the volcanically derived sediment that was distributed throughout much of
926 Australia during the Mesozoic (Boult *et al.*, 1998; MacDonald *et al.*, 2013; Barham *et al.*, 2016).
927 However, our findings raise the possibility that other, yet unidentified, intra-basinal volcanic
928 sources may contribute to Mesozoic volcanogenic sedimentation in eastern Australia.

929 Finally, it must be considered why the WVP has not been documented until now
930 despite the presence of volcanic rocks being known for over three decades. By our estimates,
931 only 0.13% of the wells drilled in the Cooper Basin between 1959 and 2015 drilled the WVP
932 (based on Hall *et al.*, 2016's estimate of the number of wells), with 35 years of exploration
933 since the first well that drilled the volcanics. This is despite the WVP occupying ~6% of the
934 geographic extent of the Cooper Basin. Largely, this is due to the lack of data and attention
935 given to the Nappamerri Trough region until recently. Acquisition of the high-quality Winnie
936 3D seismic reflection survey facilitated confident delineation of the WVP. This study also
937 underlines the importance for collaboration between industry and academia. Whilst volcanics
938 have been described within industry reports and imaged within data acquired, they have been
939 overlooked within the literature. The obscurity of the WVP in an area of such intense
940 exploration points to the probability of other undiscovered volcanic provinces globally. It is
941 therefore an important analogue in the search for undiscovered intraplate volcanism that may
942 inform our understanding of the mantle processes occurring beneath continental interiors.

943

944 **9. Conclusions**

945 We have integrated seismic, well, gravity and magnetic data and clarified the extent and
946 character of igneous rocks emplaced above the Nappamerri Trough of the Cooper Basin
947 within Eromanga Basin stratigraphy. Monogenetic volcanoes, igneous intrusions and compound
948 lava flows extending over ~7500 km² are proposed to have been active between ~180 – 160
949 Ma forming part of the proposed 'Warnie Volcanic Province.' Regionally, the distribution of
950 igneous rocks is controlled by basement structure. On a continental scale, we interpret the
951 Warnie Volcanic Province to be a product of intraplate convective upwelling above the
952 subducting Pacific slab.

953

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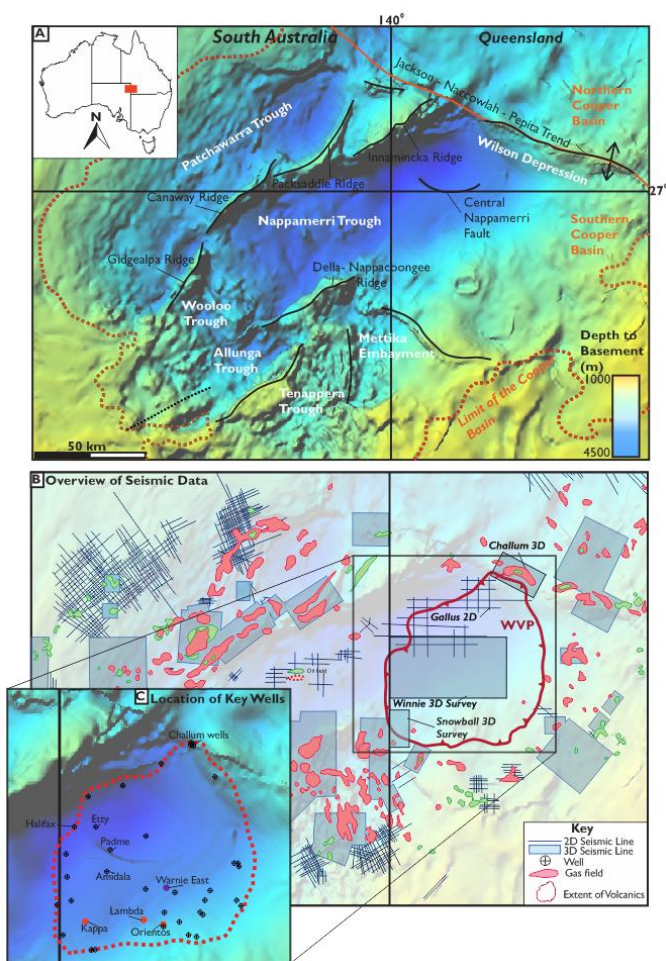
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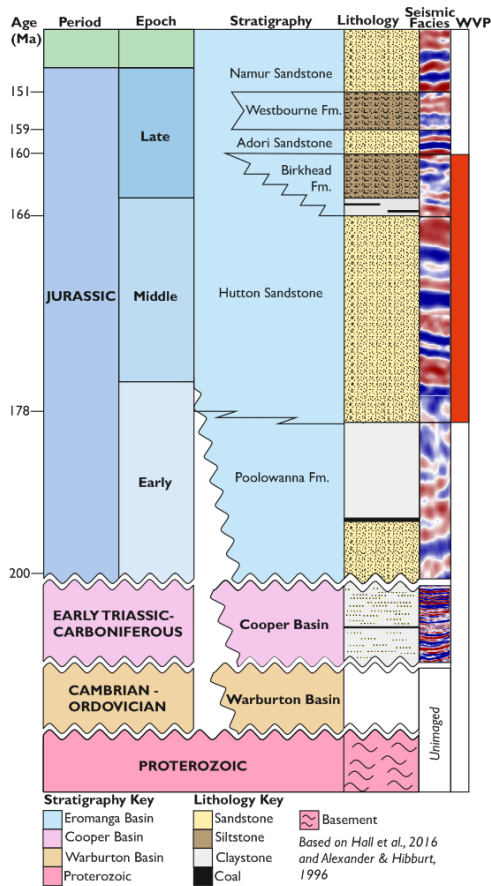
1262 **12. Figures**



1263

1264 **Figure 1.** Location map of the location of the Warne Volcanic Province and the data available. **A** Top
 1265 basement map adapted from the NGMA (2003). Key structures of the southern Cooper Basin
 1266 highlighted. **B** Location map highlighting the location of the Warne Volcanic Province within the
 1267 Cooper Basin. Top basement map superimposed with the location of 3D and 2D seismic surveys
 1268 utilised in this study and the location of key exploration wells.

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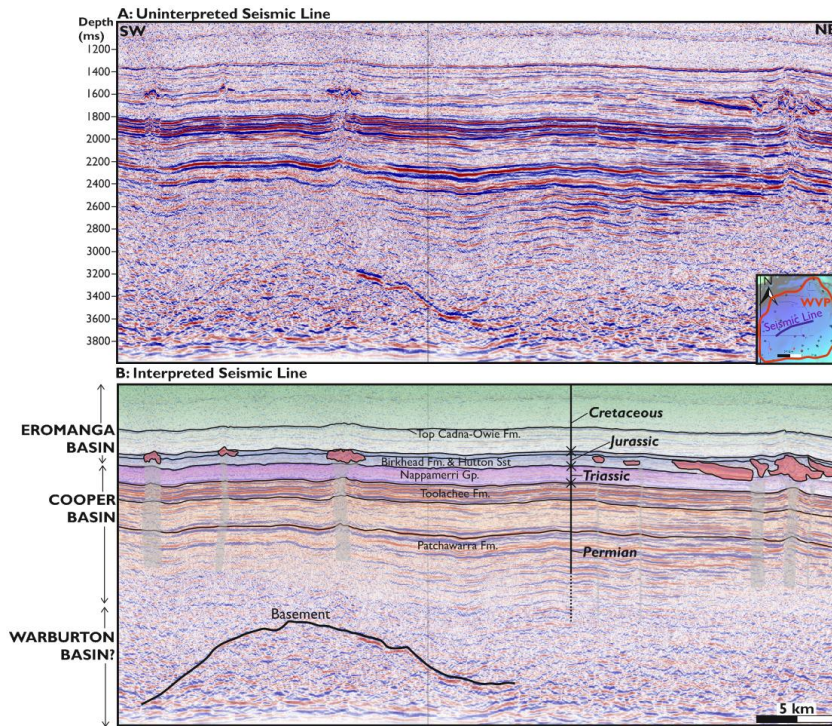


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1272 **Figure 2.** Stratigraphic column for the stratigraphy discussed within the paper, based on the work
 1273 of Alexander & Hibburt (1996), Hall et al. (2016) & Reid et al. (2009). Seismic facies for each section
 1274 is taken from the Winnie 3D survey of the eastern Nappamerri Trough. Rough stratigraphic location
 1275 of the WVP highlighted.

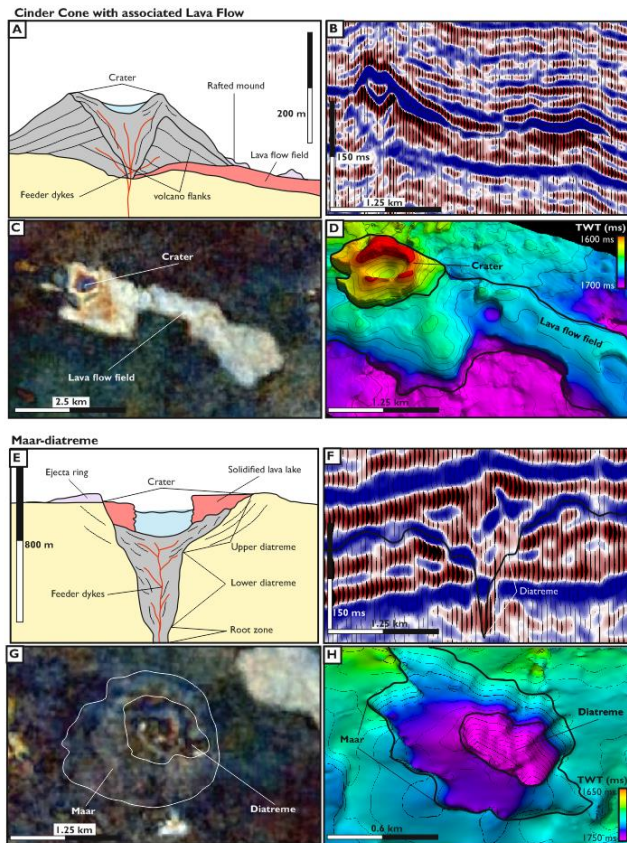
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1278 **Figure 3.** Seismic line from the Winnie 3D survey of the Nappamerri Trough highlighting the broad
1279 stratigraphy of the study area. **A** Uninterpreted line. **B** Interpreted line. Note the stacked nature of
1280 the Eromanga, Cooper and Warburton Basins and the stratigraphic location of the Warne Volcanic
1281 Province.

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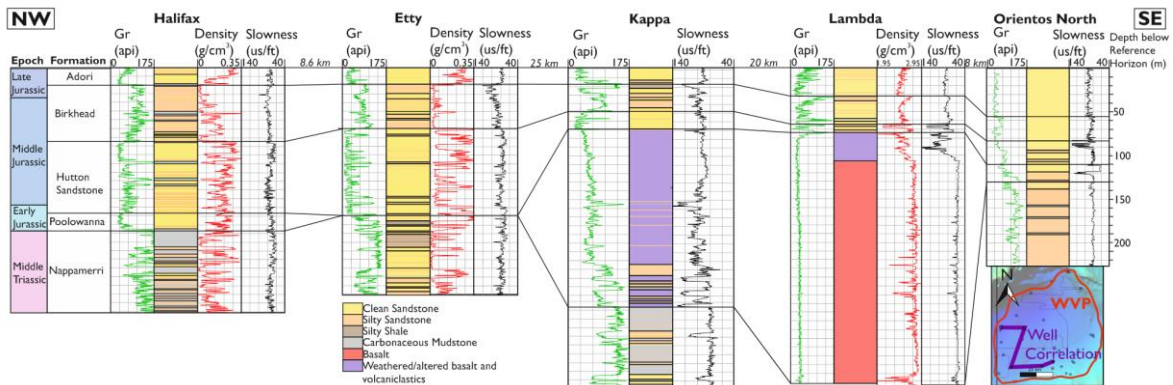
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1284 **Figure 4.** Examples of different monogenetic vents. **A** Cross section through a cinder cone, adapted
 1285 from Nemeth & Kereszturi (2015). **B** Seismic line across a cinder cone and lava flow, taken from the
 1286 Winnie 3D survey, highlighting the general morphology, seismic response. Note the characteristic eye
 1287 shape within the centre of the vent. **C** Plan view spectral decomposition of the cinder cone in **B**
 1288 highlighting how different it is from the surrounding sediments. **D** Oblique view of the same cinder
 1289 cone shown in TWT. **E** Cross section through a Maar-diatreme, adapted from Nemeth & Kereszturi
 1290 (2015). **F** Seismic line across a proposed Maar-diatreme from the Winnie 3D survey. Note how it cuts
 1291 down into the subsurface with an internally chaotic seismic response. **G** Plan view spectral
 1292 decomposition of a Maar-diatreme. **H** Oblique, TWT view of the proposed Maar-diatreme.

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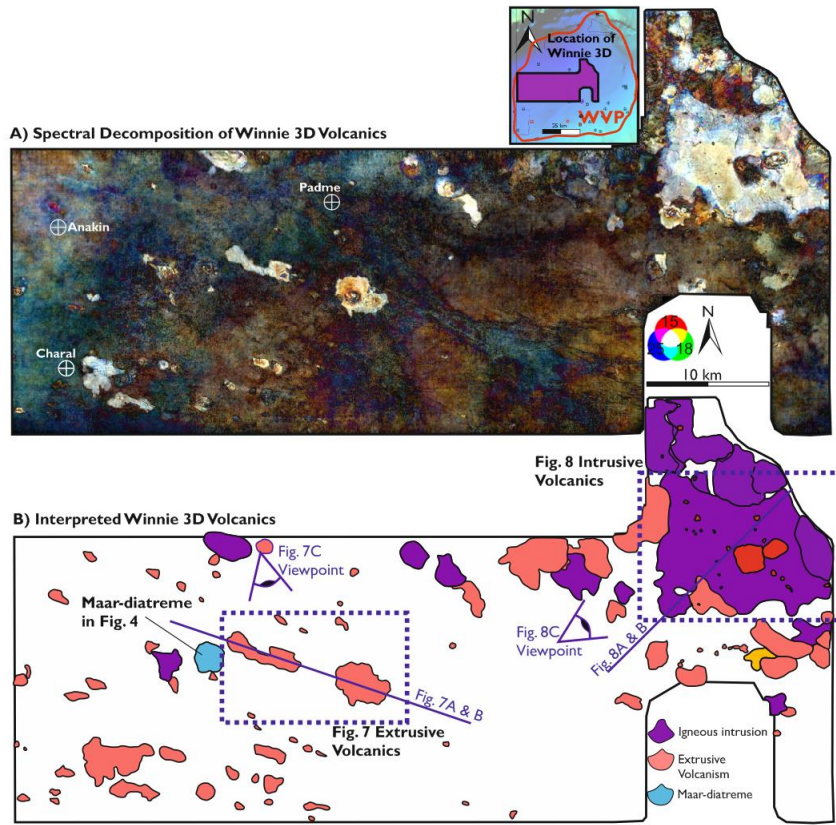


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1297 **Figure 5.** Well correlation across the eastern Nappamerri Trough showing the changes in facies
 1298 across the area and the thickness of the volcanics intersected. Note in areas of thickened volcanics
 1299 (i.e. the Kappa and Lambda wells) the Hutton sandstone and Birkhead fm. are thinned considerably
 1300 suggesting sedimentary systems were diverted away from areas of eruption. As noted in the text, the
 1301 change from the clean sandstones of the Hutton Sandstone to silty sandstones in the Birkhead fm. is
 1302 associated with an influx of volcanic arc derived sediment from eastern Australia.

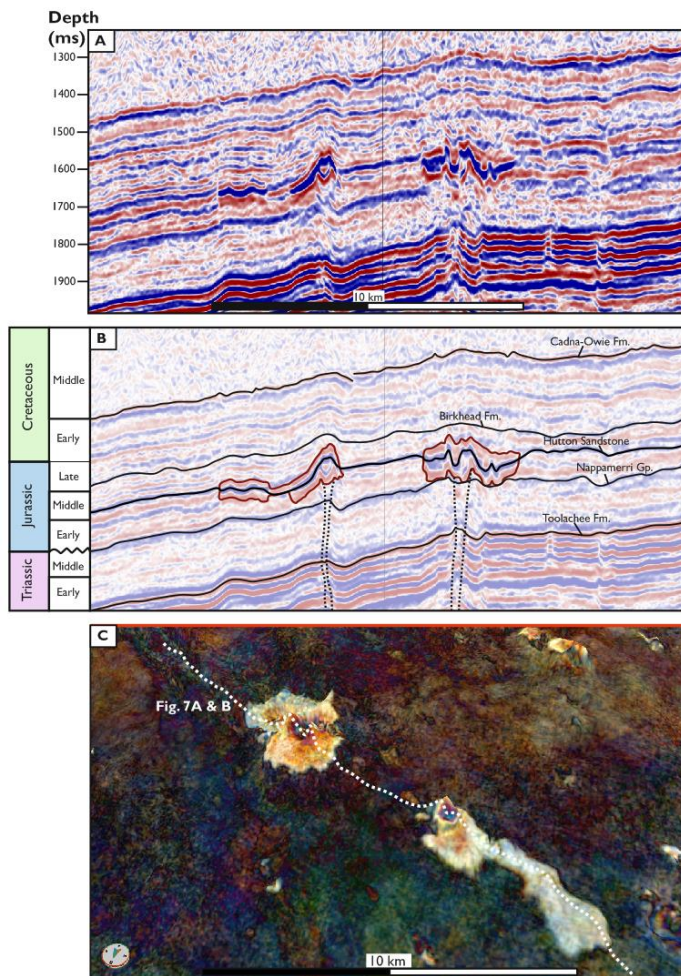
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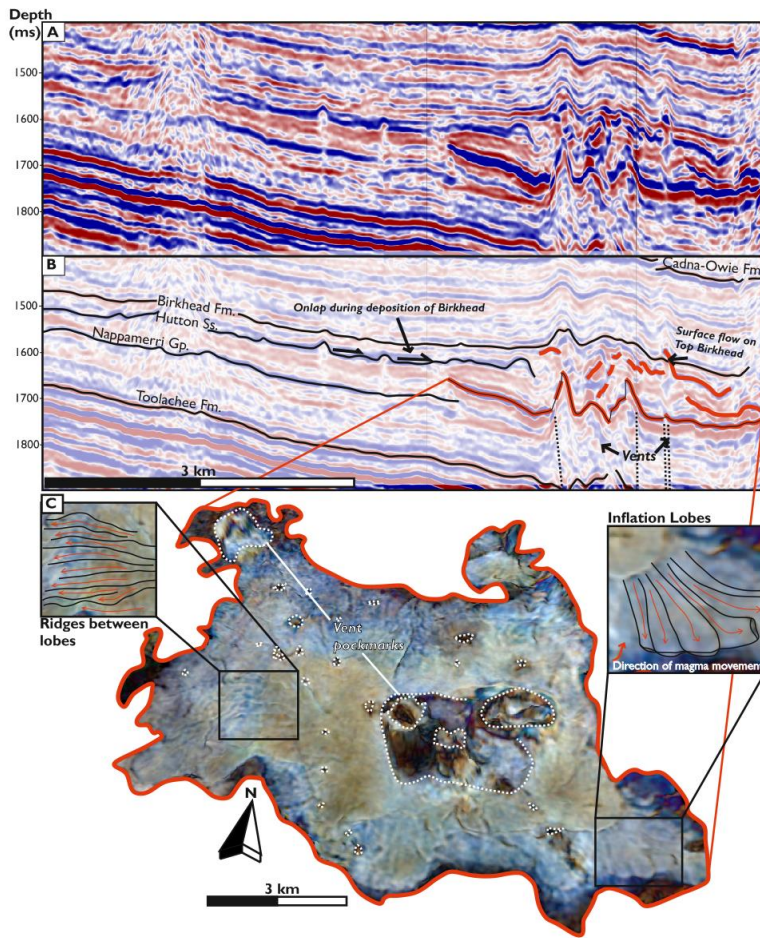
1305

1306 **Figure 6.** Spectral decomposition of a horizon mapped across the volcanics of the Winnie 3D survey.
1307 **A** Uninterpreted spectral decomposition. **B** Interpreted spectral decomposition highlight extrusive
1308 volcanics and the location of igneous intrusions. Note the increased number of intrusions towards the
1309 north east of the Winnie 3D survey.



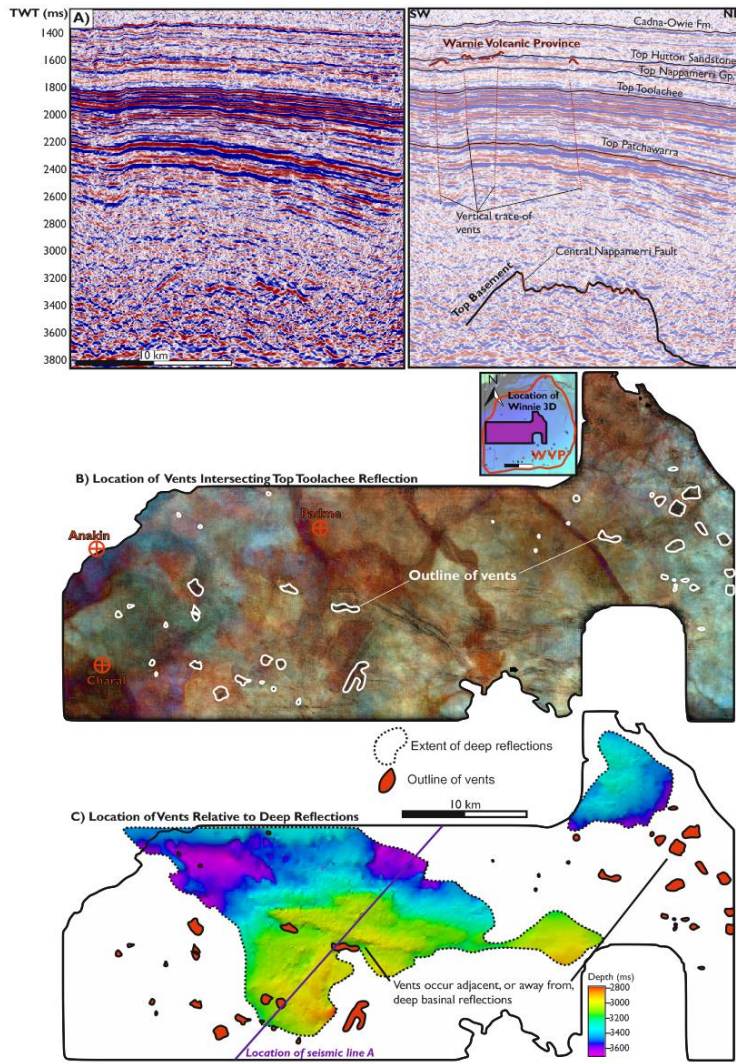
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1311 **Figure 7.** 3D Seismic across two monogenetic vents that extend into surface flows, from the Winnie
 1312 3D survey. **A** Uninterpreted line across the vents. **B** Interpreted line across the vents. Note the vertical
 1313 zone of disrupted seismic directly below the vents, indicative of underlying plumbing system that fed
 1314 the flows. **C** Spectral decomposition of a horizon mapped across the vents, highlighting their 3D
 1315 morphology.



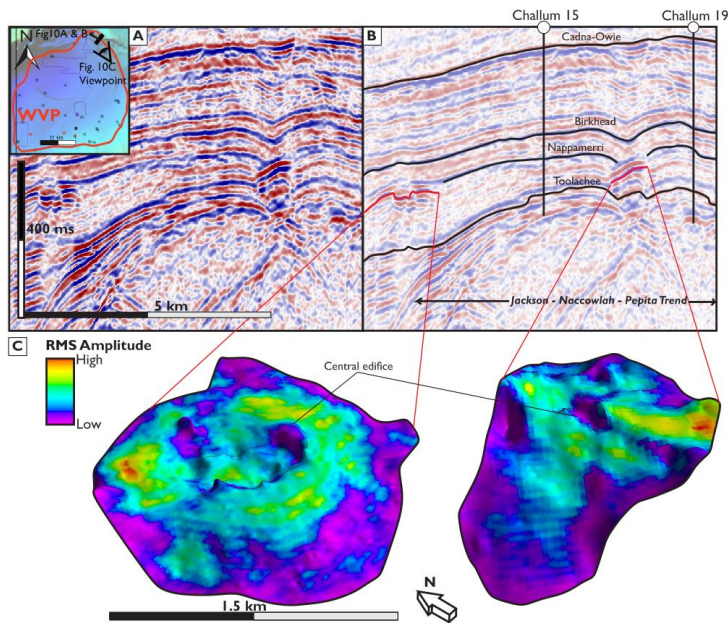
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1317 **Figure 8.** Seismic across the large intrusion in the Winnie 3D survey. **A** Uninterpreted seismic line.
1318 **B** Interpreted seismic line, illustrating the uplift of the overburden and the numerous vents and
1319 overlying volcanics associated with the intrusion. **C** Spectral decomposition of the intrusion, highlight
1320 the pockmarks of vents that pass through the intrusion and morphological characteristics associated
1321 with the intrusion.



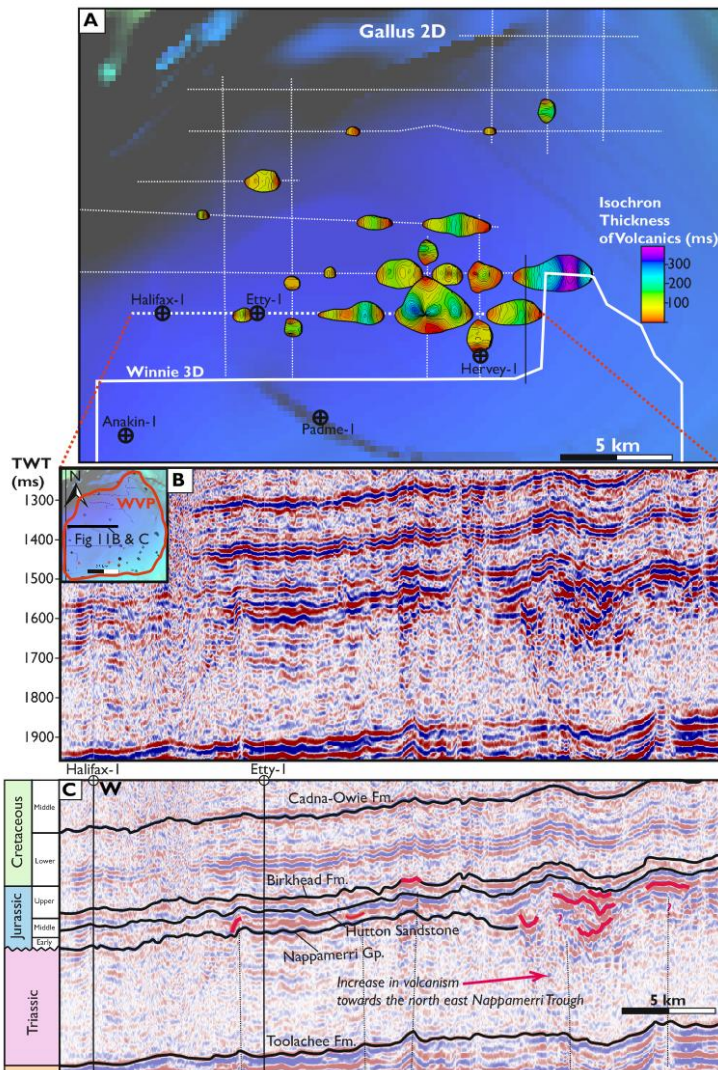
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1323 **Figure 9.** The basement structure and it's relation to vents within the Winnie 3D survey. **A**
1324 Uninterpreted and interpreted seismic line across the Winnie 3D survey, highlighting the deep
1325 basement reflections and their relation to vents feeding the WVP. **B** Spectral decomposition of the
1326 Top Toolachee horizon, highlighting the location of vents underlying the WVP. **C** Location of vents
1327 mapped using the Top Toolachee horizon superimposed on a TWT map of the top basement horizon.



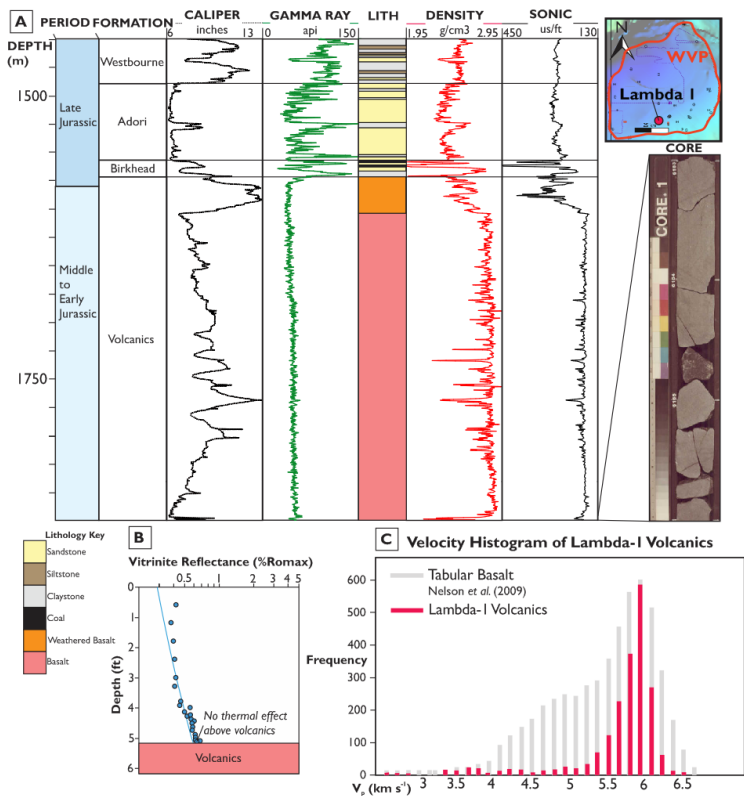
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1329 **Figure 10.** Seismic cross volcanics within the Madigan 3D survey. **A** Uninterpreted seismic line taken
 1330 across the two volcanic features. **B** Interpreted seismic horizon, showing the location of two Challum
 1331 wells and the Challum vents. **C** 3D view of the Challum vents, with an RMS amplitude extraction
 1332 draped over the horizons.



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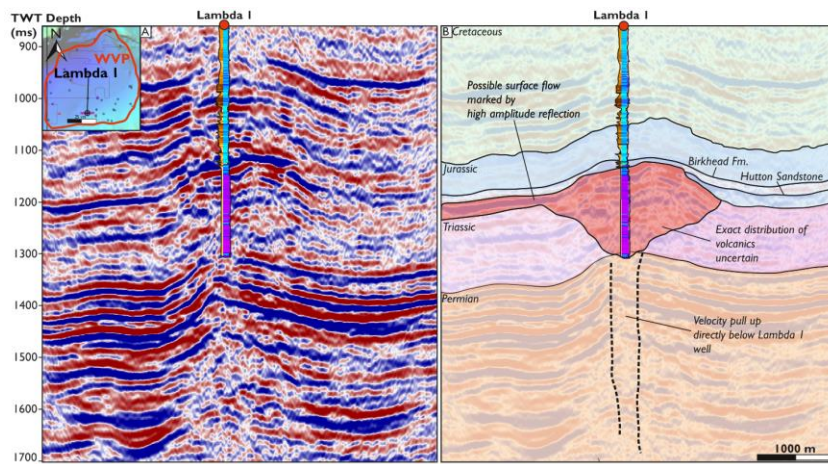
1334 **Figure 11.** Seismic images of the Gallus 2D survey. **A** Isochron thickness map of the volcanics
1335 identified throughout the Gallus 2D survey. **B** Uninterpreted seismic line from the Gallus 2D survey.
1336 **C** Interpreted seismic line from the Gallus 2D survey. An increase in the volume of volcanics towards
1337 the east of the survey is apparent.



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1339 **Figure 12.** Well data for the volcanics from the Lambda-I well. **A** Composite log of the volcanic
 1340 succession in the Lambda-I well. Includes an image of the core taken from the base of the volcanic
 1341 succession. **B** Vitrinite reflectance curve for the succession above the volcanics taken from Boothby
 1342 (1986). Little thermal effect is observed. **C** Velocity histogram of the Lambda-I volcanics
 1343 superimposed on Nelson *et al.*'s (2009) velocity histogram of Tabular basalt.

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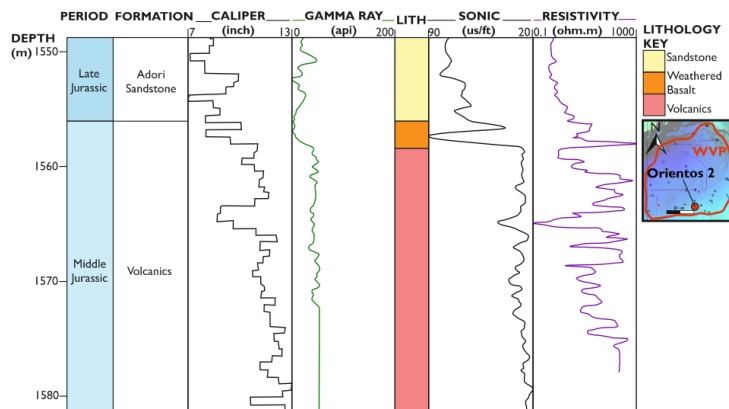
1345

1346 **Figure 13.** West to east seismic line running across the Lambda I well. The extent of the Lambda I
 1347 vent has been inferred base on the thickness of volcanics within the well.

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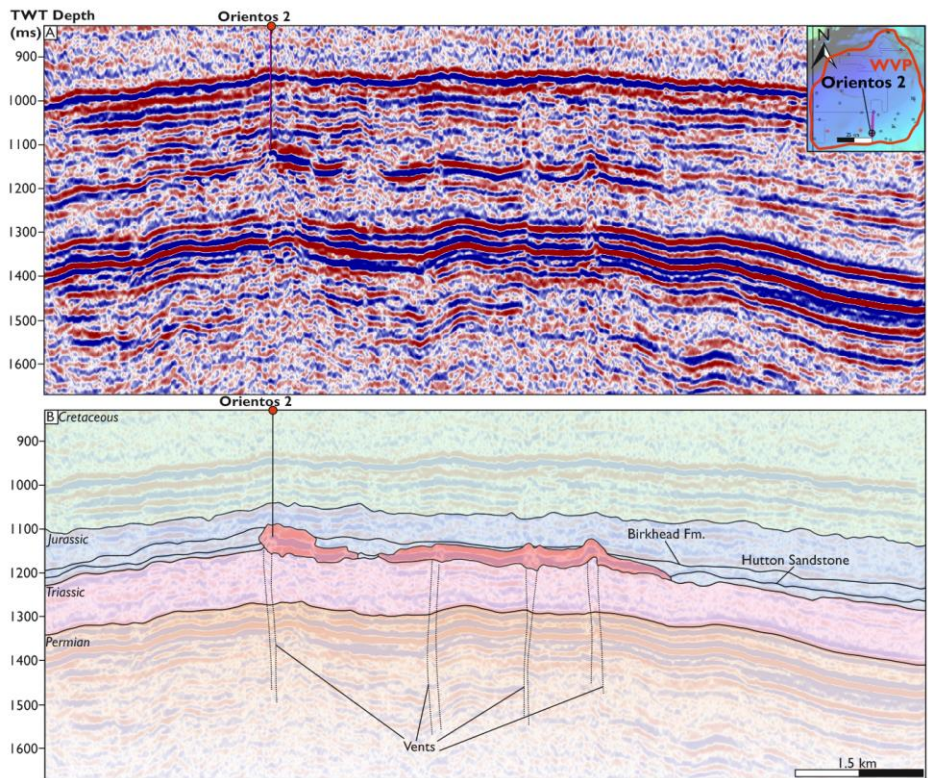
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1352 **Figure 14.** Digitised composite log for the volcanics within the Orientos 2 well.

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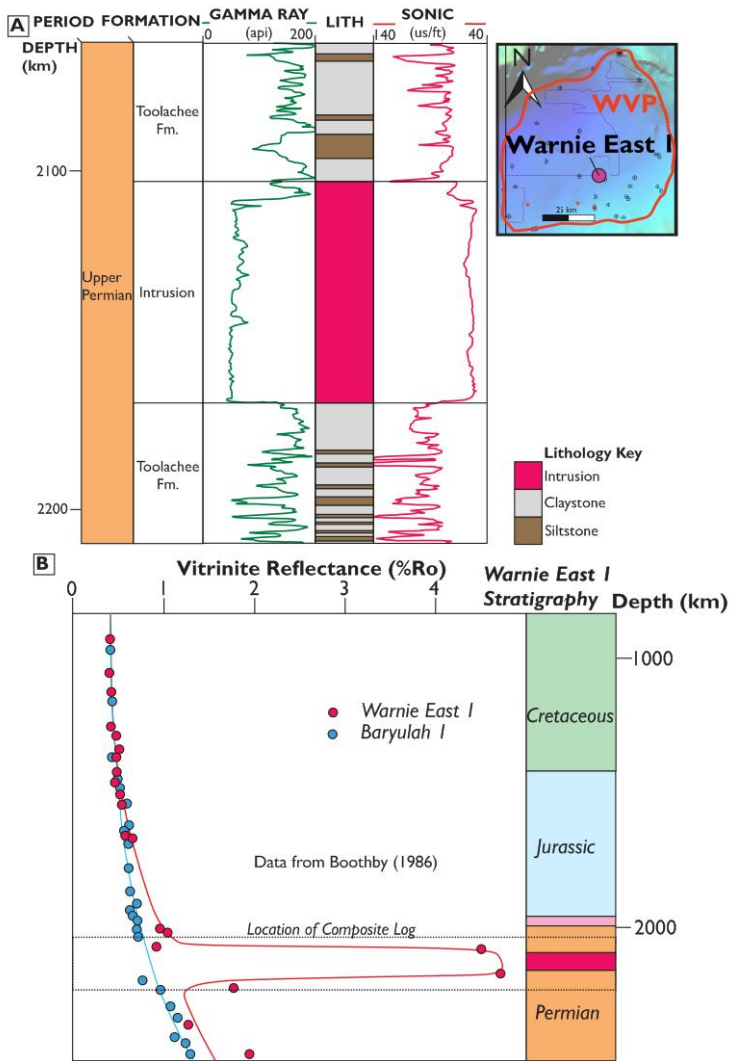
1354

1355 **Figure 15.** South to North 2D seismic line running across the location of the Orientos 2 well. A
 1356 series of tabular basalt lava flows emplaced on top of the Triassic Nappamerri group are visible, with
 1357 Orientos 2 intersecting the top of the southernmost flow.

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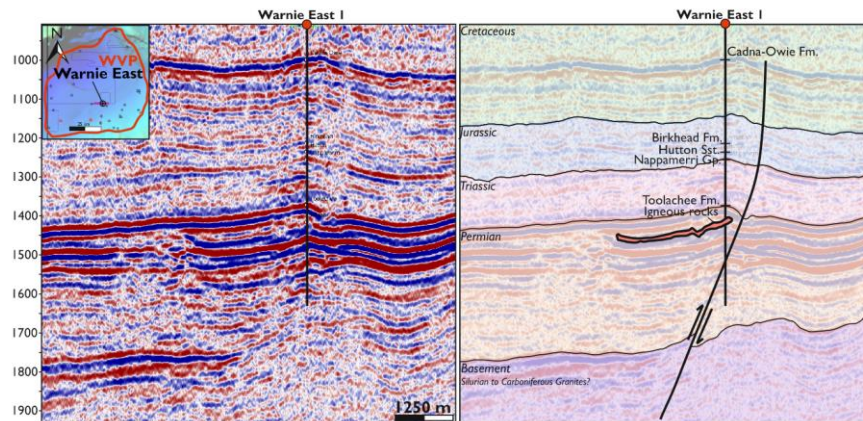
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1361

1362 **Figure 16.** Well data for intrusive volcanic rocks within the Warnie East I well. **A** Composite log
1363 across the intrusion in the Warnie East well. **B** Vitrinite reflectance profile across the intrusion in the
1364 Warnie East well taken from Boothby (1986). Thermal effect both above and below the well is noted.

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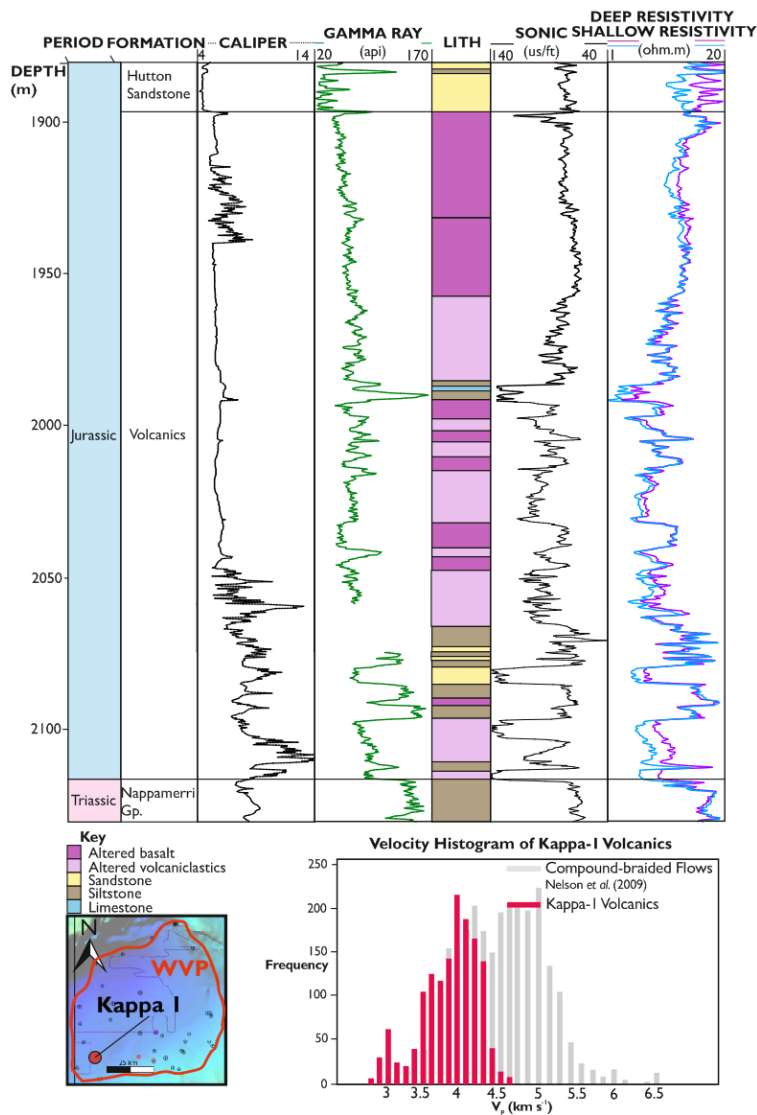
1366

1367 **Figure 17.** 2D seismic line running from west to east across the Warnie East I well. The exact
 1368 location of the igneous rocks is unclear on the line, however, a shallow saucer-shaped reflection with
 1369 the Toolachee Fm. intersects Warnie East I at the depth igneous rocks are found. Also notable is the
 1370 presence of a rollover anticline adjacent to a growth fault.

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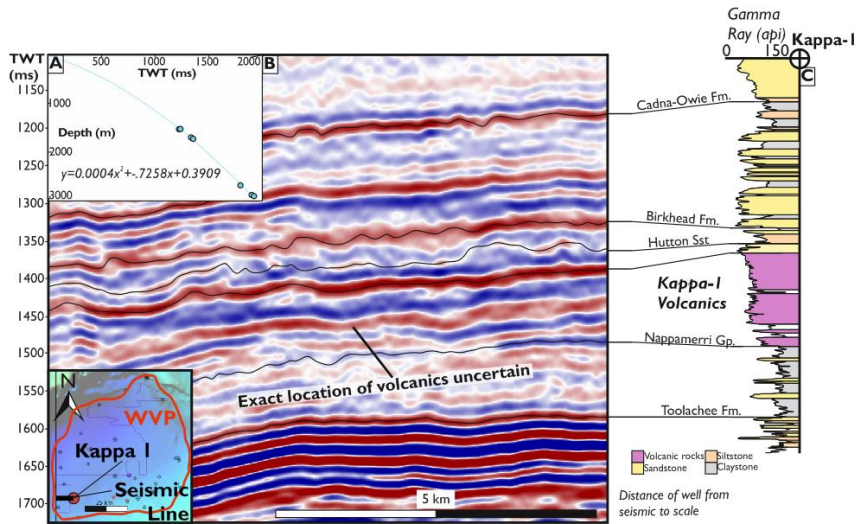
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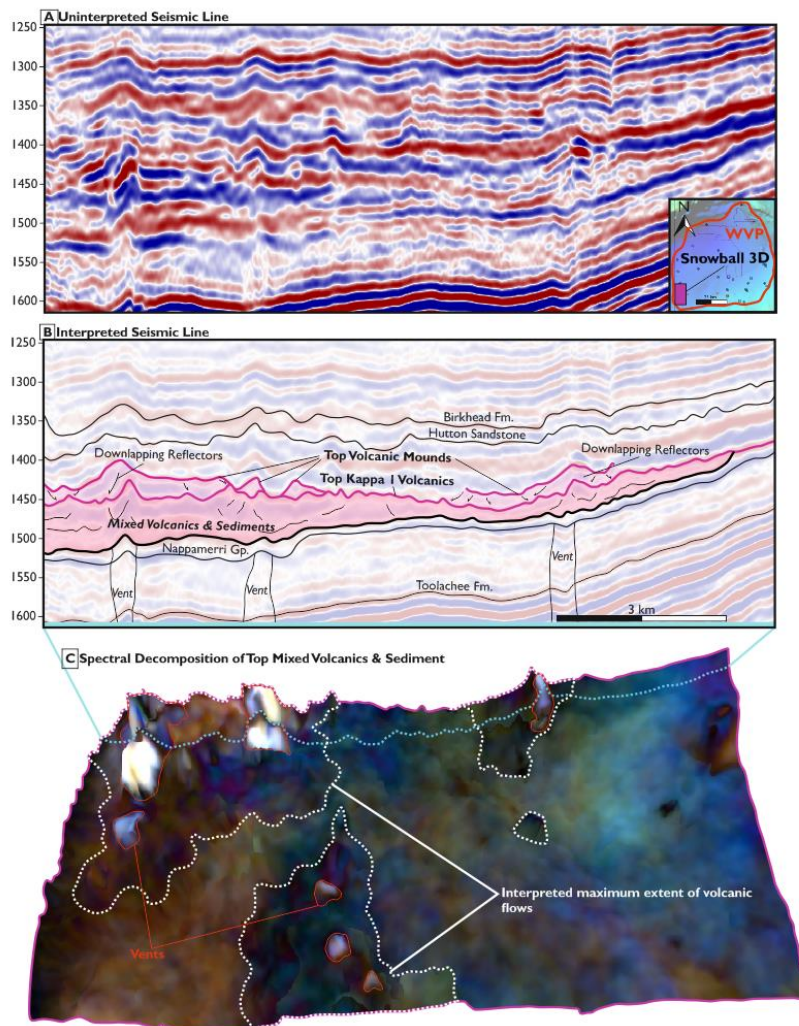
1374

1375 **Figure 18.** Well data for the Kappa-I well. **A** Composite log for the Kappa-I well volcanics. **B**
1376 Velocity histogram for the Kappa-I well volcanics, superimposed on top of Nelson *et al.*'s (2009)
1377 velocity histogram for compound-braided volcanics.



1378

1379 **Figure 19.** Position of the Kappa-I well in relation to the Snowball 3D survey. **A** TWT to depth
 1380 relationship used to convert the Kappa-I well data from depth to time in order to display it adjacent
 1381 to the Snowball 3D survey. **B** Line across the Snowball 3D survey with the position of formation tops
 1382 based on the Kappa-I well displayed. **C** Gamma ray curve from the Kappa-I well with formation tops
 1383 marked.

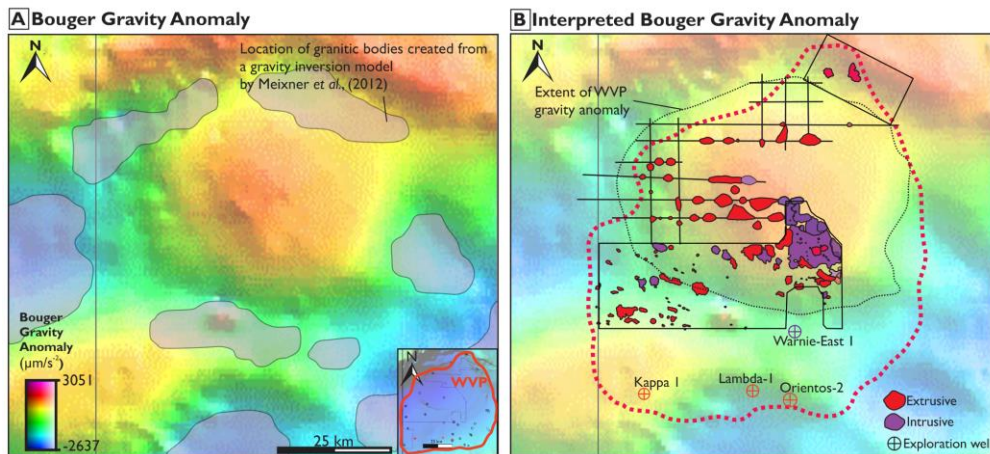


1384

1385 **Figure 20.** Seismic across volcanics within the Snowball 3D survey. **A** Uninterpreted seismic line. **B**
1386 Interpreted seismic line. Thick package of mixed volcanics & sediments based off of the Kappa-I well
1387 volcanics. **C** Spectral decomposition of the top mixed volcanics & sediment horizon. Location of vents
1388 noted by bright white colour with interpreted flow fields based on dark colours adjacent to vents.

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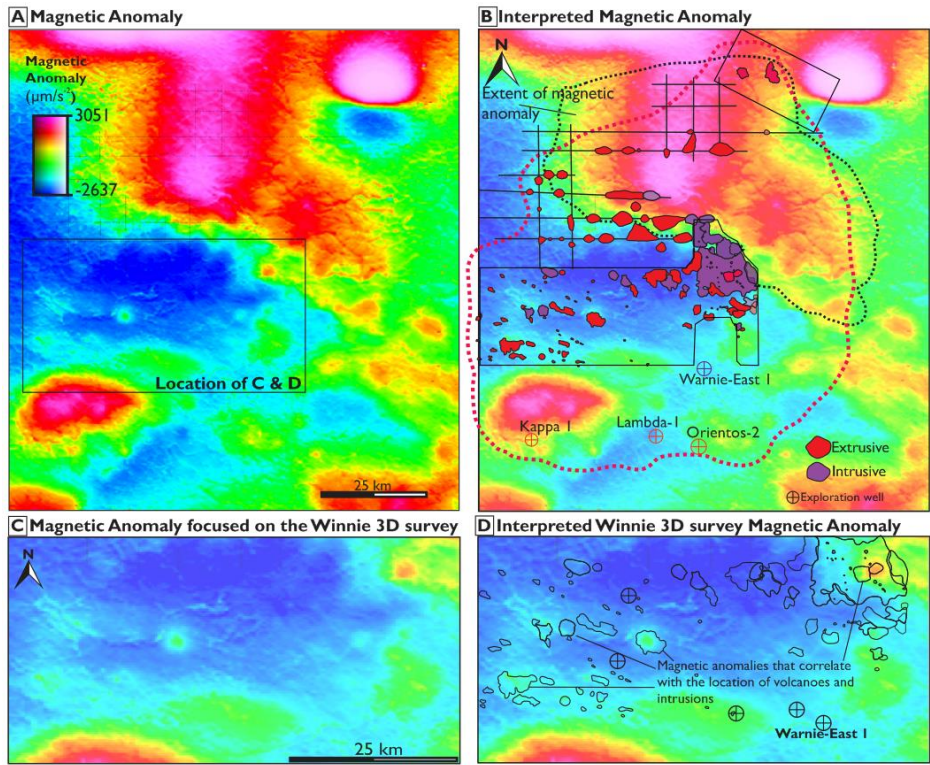
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1392 **Figure 21.** Bouguer gravity anomaly adopted from Tracey & Bacchin (2008). The gravity data shows
1393 spherical cap-corrected Bouguer anomalies with a high pass filter with a cut-off wavelength of 500 m.
1394 Spacing coverage is typically 11 km, with more closely spaced grids (2 km) in areas of economic interest
1395 (Tracey & Bacchin, 2008). **A.** Uninterpreted bouguer gravity anomaly overlain with the location of
1396 granitic bodies created from Meixner *et al.*'s (2002). Granitic bodies are expressed as lower values in
1397 the gravity anomaly, located adjacent to the 50 x 50 km gravity high below the Warnie Volcanic
1398 Province. **B.** Bouguer gravity anomaly superimposed with the location of the igneous rocks within the
1399 Warnie Volcanic Province.

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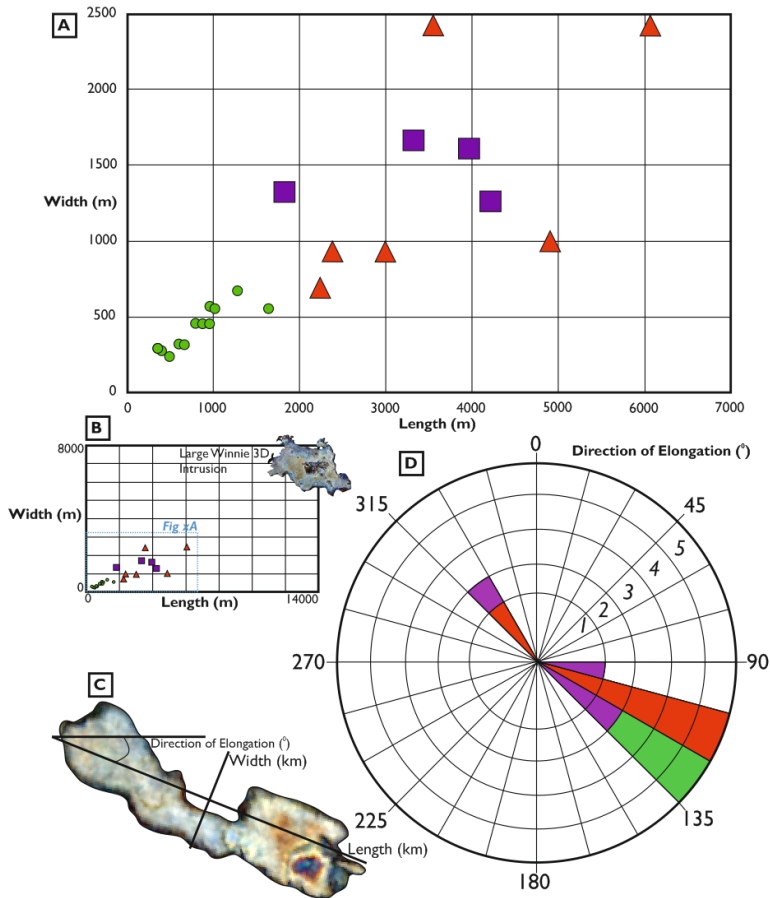


1401

1402 **Figure 22.** Total magnetic intensity across the Warnie Volcanic Province. This data was obtained for
1403 study from the Queensland Department of Natural Resources and Mines database. The survey was
1404 flown in an east-west direction at a height of 60 m with a distance of 400 m between the traverse
1405 lines. **A.** Unannotated total magnetic intensity over the eastern Nappamerri Trough. Location of C
1406 and D highlighted. **B.** Total magnetic intensity superimposed with the location of igneous rocks
1407 interpreted within the Warnie Volcanic Province. **C** Unannotated total magnetic intensity for the area
1408 covered by the Winnie 3D survey. **D** Total magnetic intensity superimposed with the location of
1409 igneous rocks interpreted within the Winnie 3D survey. Note the close correspondence between
1410 some magnetic anomalies and the location of igneous rocks.

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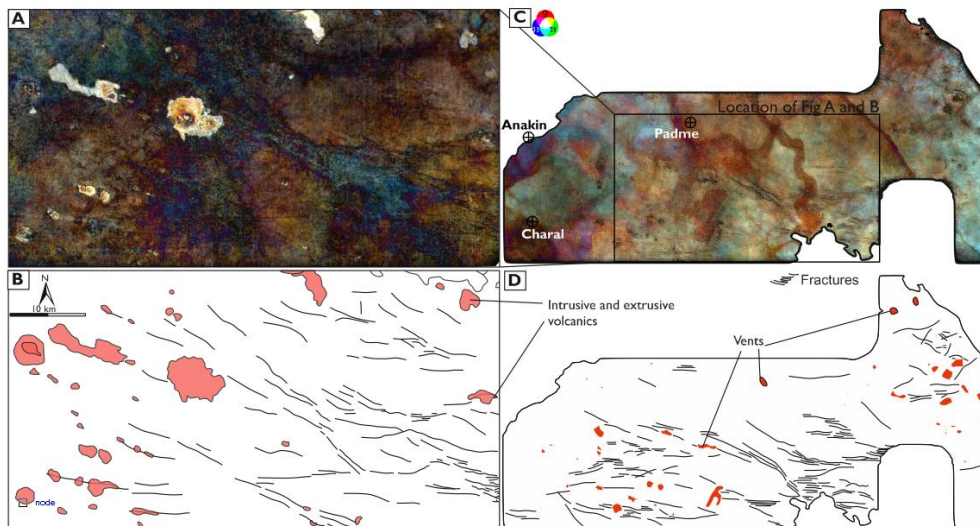
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1414 **Figure 23:** Size and orientation of volcanics within the Winnie 3D and Madigan 3D surveys. **A** Size
1415 of vents, intrusions and volcanic flows. **B** Figure A, with the singular large intrusion superimposed on
1416 it, highlighting the shear the scale of this volcanic features. **C** Measurements taken for each individual
1417 volcanic feature. **D** Direction of elongation for the volcanics, highlighting that they are elongate in a
1418 SE-NW direction.

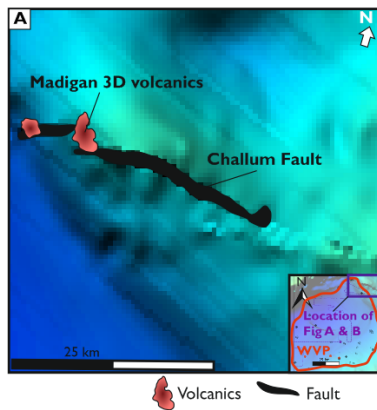
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1421 **Figure 24:** Mapping of minor faults within the Nappamerri trough. **A** Spectral decomposition of the
 1422 Top Volcanics horizon from the Winnie 3D survey. **B** Interpretation of A with discontinuities and
 1423 the location of volcanics highlighted. **C** Spectral decomposition of the Top Toolachee horizon. **D**
 1424 Interpretation of discontinuities within the Toolachee Horizon and the location of volcanics
 1425 highlighted.

1426



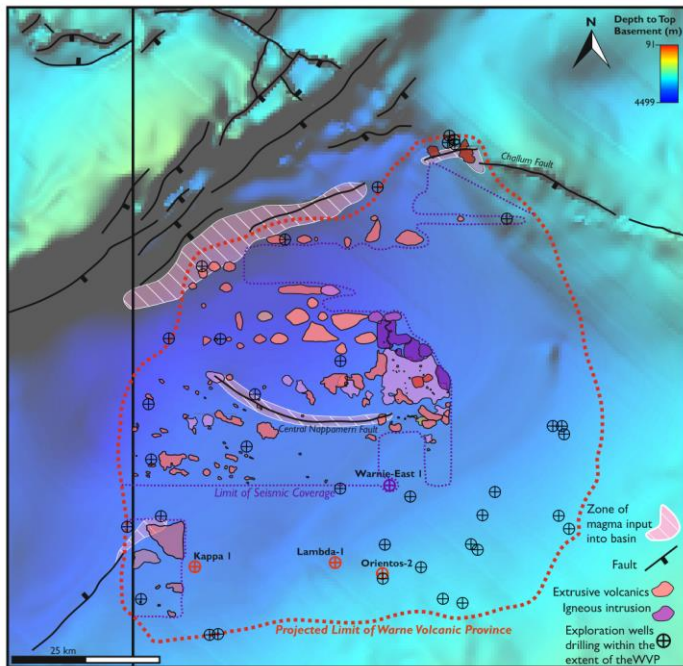
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1428 **Figure 25: A** Location of the Challum Fault and Madigan 3D volcanics superimposed on a Depth to
1429 top basement map adapted from NGMA (2003). Madigan 3D volcanics located above the North
1430 West tip of the Challum Fault.

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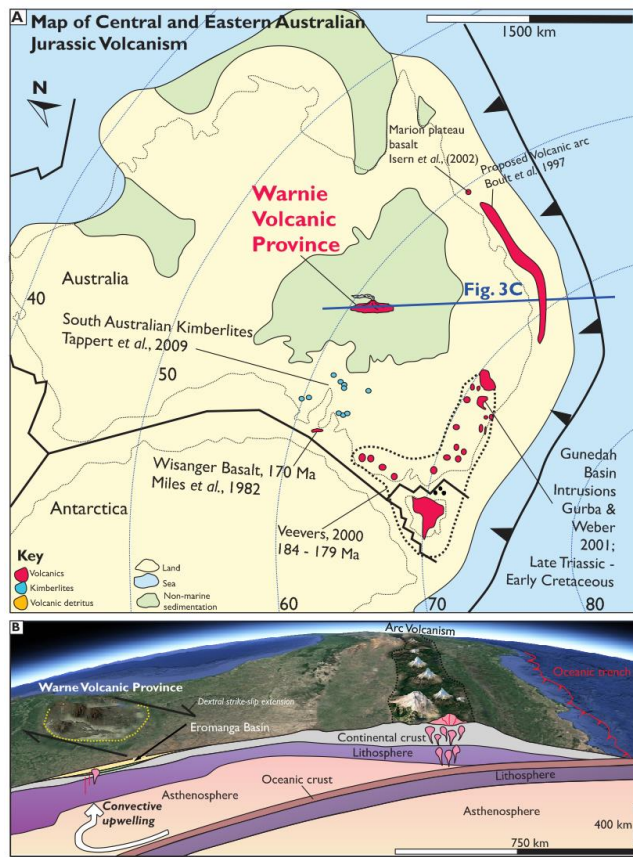
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1435 **Figure 26:** Regional map of the Warnie Volcanic Province. Depth to top basement map adapted
1436 from NGMA (2003). Location of extrusive volcanics and igneous intrusions based on seismic data.
1437 The volcanics, however, are thought to extend beyond the limit of seismic coverage towards the
1438 projected limit of the Warnie Volcanic Province. Zones of magma input into the basin are based off
1439 of the main basement bounding faults identified within the study, although many more minor input
1440 points are expected to occur within the Nappamerri trough from faults not imaged within this study.

1441



1442

1443 **Figure 27: A** Map of Australia, superimposed with the location of middle to late Jurassic volcanics
 1444 in eastern Australia, described in the text. Base map is a magnetics map used to highlight the Great
 1445 Artesian Basin. **B** Palaeogeographic map of Australia in the middle Jurassic (Oxfordian, 160 Ma),
 1446 highlighting the tectonic nature of the different volcanics. **C** Oblique view of Australia, highlighting
 1447 the subducting pacific slab and associated convective upwelling. Steps in lithospheric thickness have
 1448 adapted from Fishwick et al., (2008).

1449