

A Quality-Control Procedure for Bio-Optical Applications of Hyperspectral Radiometric Upwelling Radiance and Downwelling Irradiance Profiles Measured by BioGeoChemical-Argo Floats.

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ABSTRACT

24

25 Autonomous in-situ radiometric observations are increasingly used to constrain bio-optical
26 processes and validate satellite ocean-color products, such as remote sensing reflectance and
27 diffuse attenuation coefficients. Because these observations are collected independently of
28 weather and sea-state conditions, their application critically depends on robust quality control.

29 Starting in 2012, the BioGeoChemical-Argo (BGC-Argo) program has measured
30 downwelling irradiance (E_d) at three wavelengths on autonomous floats. Since 2022, a pilot array
31 of 12 BGC-Argo floats equipped with TriOS-RAMSES hyperspectral radiometers measuring E_d
32 and upwelling radiance (L_u) has been deployed across open-ocean regions with diverse bio-
33 optical properties. To date, these floats have acquired hundreds of hyperspectral profiles from 0–
34 300 m at ~10-day intervals near local noon.

35 This study presents an automated Quality Control (QC) method for hyperspectral E_d and L_u
36 profiles measured by BGC-Argo floats, building upon previous QC procedures designed for
37 multispectral radiometry. The method flags perturbations in the light field caused by self-
38 shading, large tilt angles, passing clouds, wave focusing, spikes, and corrects for dark current
39 signals. The QC is first applied at five key wavelengths (380, 443, 490, 555, and 620 nm) to
40 generate wavelength-specific flags along each vertical profile, which are then combined into a
41 final global classification for each spectral profile as *Good*, *Questionable*, or *Bad*.

42 This paper, along with its Python code and data files, provides the community with a robust
43 and computationally efficient approach for assessing hyperspectral BGC-Argo data quality,
44 preparing it for further bio-optical applications.

45 **1. Introduction**

46 Since its inception in the early 2000s, the Argo program has maintained approximately 4,000
47 autonomous floats across the world's oceans. These floats measure pressure, temperature, and
48 salinity, profiling from a depth of 2,000 meters to the surface every 10 days. After each cycle, the
49 collected data are transmitted by satellite to land before the float begins a new sampling cycle.
50 This unprecedented collaboration among 30 countries resulted in the collection of over 2 million
51 profiles of ocean parameters, revolutionizing the scientific community's understanding of climate
52 systems, weather forecasting, circulation, and more (Roemmich et al., 2019; Wong et al., 2020).

53 In 2016, the Argo program was expanded to include measurements of biogeochemical variables
54 (Bittig et al., 2019; Johnson & Claustre, 2016). The BioGeoChemical-Argo (BGC-Argo) initiative
55 aims to maintain a network of ~1,000 floats that, in addition to the core Argo parameters, measure
56 dissolved oxygen, nitrate, pH, phytoplankton chlorophyll-a, optical backscattering coefficient,
57 photosynthetically available radiation (PAR), and downwelling irradiance (E_d), providing critical
58 insights into the ocean's biological and chemical processes (Claustre et al., 2020). These BGC-
59 Argo floats follow the same sampling cycle as Argo floats, sampling every 10 days from 2,000
60 meters up to the surface, except the radiometric parameters (E_d and PAR) that are collected in the
61 upper 300 meters.

62 Starting in 2022, as part of the ERC-funded REFINE (Robots Explore plankton-driven Fluxes in
63 the marine twilight zoNE) project, a subset of these floats has been equipped with ruggedized
64 hyperspectral TriOS-RAMSES G2 radiometers to measure profiles of both E_d and upwelling
65 radiance (L_u) at at least 70 wavelengths. These hyperspectral BGC-Argo floats will hereafter be
66 referred to as *hyperspectral floats*. Twelve of these floats are operational and acquire data in near
67 real time across open-ocean regions representative of the ocean's diverse bio-optical conditions.
68 This hyperspectral fleet, programmed to reach the surface at local noon, provides a new in situ
69 data source for bio-optical applications and validation of hyperspectral satellite ocean color
70 missions such as NASA's newly launched Plankton, Aerosol, Cloud, Ocean Ecosystem (PACE)
71 satellite (Werdell et al., 2019).

72 Raw radiometric data need to be processed to identify perturbations in the light field (e.g., self-
73 shading, float tilt, wave focusing, passing clouds, etc.) that limit their utility for bio-optical
74 applications. Quality control (QC) procedure for radiometric quantities measured by autonomous
75 floats has already been designed for E_d acquired by SeaBird-OCR radiometers mounted on floats
76 (Organelli et al., 2016; Wojtasiewicz et al., 2018). Earlier QC procedures were developed for
77 multispectral radiometers measuring only three discrete bands, typically selected from 380, 412,
78 443, 490, and 555 nm. Because these instruments did not sample wavelengths longer than 555 nm,
79 the original QC framework was neither designed nor evaluated for the yellow–red region of the
80 spectrum. In contrast, hyperspectral radiometers cover a much broader spectral range, extending
81 approximately from 320 to 780 nm. This spectral region includes key optical features such as
82 chlorophyll fluorescence and Raman scattering (Bartlett et al., 1998; Bricaud et al., 2004;
83 Desiderio, 2000), as well as strong absorption by water, complicating quality assessment.
84 Consequently, an updated QC procedure is required to reliably detect artifacts—such as wave-
85 focusing effects—at these longer wavelengths. This updated method needs to be computationally
86 efficient, applicable to both E_d and L_u , and relies solely on these 2 variables to be effective across
87 the entire visible spectrum and the entire hyperspectral fleet.

88 This study, therefore, aims to build upon the previous QC procedures developed for BGC-Argo
89 multispectral radiometry and adapt them to hyperspectral TriOS-RAMSES G2 measurements.

90 It functions by first assessing individual data points along depth within a profile at five key
91 wavelengths (380, 443, 490, 555, and 620 nm). For the rationale of the key wavelength selection,
92 please see Section 3.a. Based on these QC by wavelengths, the whole spectrum is then
93 characterized and given one overall flag, but there is no data correction or removal.

94 This QC procedure has been developed and tested on 899 profiles available from 12 BGC-Argo
95 floats, profiling in various open ocean regions in the period since June 2022. It is important to note
96 that this procedure is not designed to be implemented in the BGC-Argo Global Data Assembly
97 Center (GDAC). Indeed, data distributed from the GDAC (both in digital counts and in scientific
98 units) only aim to be free of sensor issues. In contrast, the QC presented here is an additional
99 procedure that relies on data delivered by GDAC in order to detect light field disturbances in the

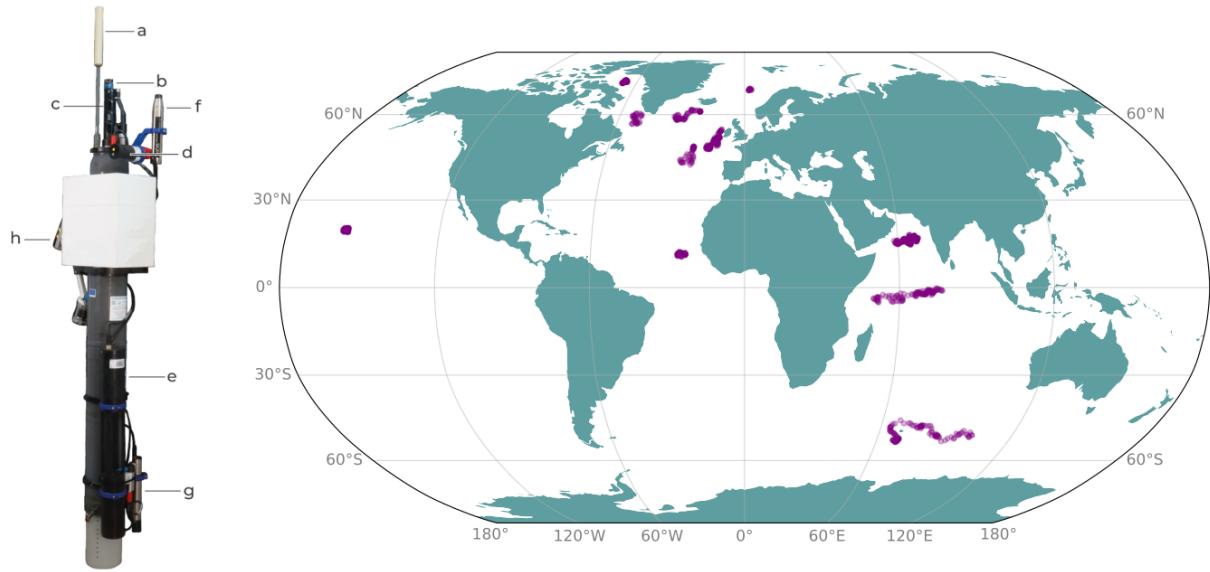
100 signal (e.g., passing clouds). This enables users to efficiently quality control spectra that can be
101 used for bio-optical applications such as ocean color radiometry validation or diffuse light
102 attenuation coefficient (K_d/K_L) computation. Therefore, there is a fundamental difference between
103 data post-GDAC quality control (free of sensor issues but not necessarily usable for bio-optical
104 applications) and the output from the quality control presented in this paper (free of sensor issues
105 and usable for bio-optical applications).

106 **2. Instruments and data**

107 *a. Hyperspectral data*

108 A new generation of BGC-Argo floats has been designed in the context of the ERC-REFINE
109 project. The platform (NKE Provor CTS5 Jumbo, Fig. 1) enables the integration of two
110 hyperspectral radiometers (TriOS RAMSES G2), one measuring E_d and the other L_u , along with
111 various optical sensors such as a fluorometer, a backscattering sensor, a transmissometer, and an
112 Underwater-Vision-Profiler (UVP6). To reduce shading effects, the E_d and L_u sensors are mounted
113 respectively at the top of the float and at the very bottom (each sensor collector is respectively 0.50
114 m above the waterline and 1.76 m under the waterline when the float is at the surface). However,
115 at specific sun azimuth angles (saa), L_u measurements can still be shaded by the float body and
116 thus require a specific processing (see details in Section 3.c).

117 This study uses 899 profiles collected by 12 floats between June 2022 and June 2025 in various
118 oceanic areas (Fig. 1). Raw data are publicly available online at
119 <ftp://ftp.ifremer.fr/ifremer/argo/aux/coriolis/>.



120

121 **Fig. 1.** On the left panel: the PROVOR CTS-5 profiling float equipped with (a) Iridium antenna; (b) oxygen sensor;
 122 (c) conductivity-temperature-pressure sensor; (d) sensor measuring chlorophyll-a fluorescence, dissolved organic
 123 matter fluorescence, and light backscattering; (e) transmissometer; (f) hyperspectral downwelling irradiance
 124 radiometer (E_d); (g) hyperspectral upwelling radiometer (L_u); (h) Underwater Vision Profiler (UVP). Photo
 125 by Thomas Jessin and Thomas Boniface. On the right panel: map of the profiles used in this study. Each circle
 126 corresponds to a location where hyperspectral radiometric profiles of E_d and L_u were measured, for a total of 899 float
 127 profiles.

128

129 *b. Technical specs: TriOS hyperspectral radiometers*

130 The E_d radiometers (RAMSES G2-ACC-VIS-Ti-2000m) have a cosine collector of 3.5 mm in
 131 radius. L_u radiometers (RAMSES-G2-ARC-VIS-Ti-2000m) have a 7° full-angle Field Of View
 132 (FOV) in air. Both have a 256-channel (or -pixel) detector with a Full Width Half-Max of 9.5 nm,
 133 including ~17 darkened pixels, and spectral sensitivity coefficients were determined by the
 134 manufacturer on individual sensors pre-deployment (TriOS, 2024). Immersion coefficients were
 135 laboratory-determined by the manufacturer for each ACC-type sensor and theoretically for ARC-
 136 type sensors (Ohde & Siegel, 2003). Following the TriOS user manual, the integration time is
 137 automatically set between 4 ms and 4096 ms: a first scan is performed at 4096 ms; if any of the
 138 pixels is saturated, a second scan is performed with a new integration time equal to half of the
 139 previous one (here: $4096/2 = 2048$ ms), and so on until one of the pixels is saturated. Detailed
 140 sensor specifications can be found in Vabson et al. (2024). TriOS sensors are equipped with an

141 internal tilt sensor that takes a measurement before and after each spectrum. In this study, the post-
142 tilt measurement is used as the tilt value because it is the closest in time to the spectral acquisition.

143 Measurements are collected during the upward casts, programmed every 10 days, with the first
144 spectrum taken at the parking depth (1,000 m). During the 4 months of the initial acquisition phase
145 of the PACE satellite, the measurements were acquired every 5 days. Radiometric data are
146 recorded by the float from ~300 m depth to the surface with a vertical resolution rising from 10 to
147 0.3 m as the float ascends: 4 points per meter between the surface and 20 m, 1 point per meter
148 between 20 to 100 m, 1 point every 2 meters from 100 to 200 m and 1 point every 5 meters from
149 200 to 300 m. The floats therefore have a relatively low vertical resolution - less than 4 points per
150 meter compared to ~15 points per meter for other known systems (HyperPro, C-OPS, HyperNav,
151 Zibordi et al., 2019). Combined with the fact that there is only a single upward profile per cycle,
152 mitigation strategies such as multicasting cannot be used to reduce wave perturbations (Zibordi et
153 al., 2004). Thus, BGC-Argo data requires a specific QC capable of detecting data affected by wave
154 focusing. Concurrent dark measurements are collected on ~17 “black” detector channels consisting
155 of purposely masked detector pixels.

156 All hyperspectral BGC-Argo floats are programmed to reach the surface at local noon. For the
157 PROVOR floats, the hyperspectral measurements are transmitted from 320 to 780 nm at the
158 original 3.3 nm spectral sampling or binned at 6.6 nm (a two-by-two channels average), depending
159 on the float’s configuration (TriOS, 2025). The central wavelength associated with each pixel
160 slightly varies (<4 nm) between sensors, meaning that all hyperspectral radiometers from the fleet
161 do not have the same distribution of spectral bands.

162

163 *c. Processing of raw data.*

164 Electronic raw counts are transmitted to land through a two-way Iridium satellite communication,
165 which also enables the remote adjustment of mission and sensor parameters (sampling frequency,
166 pixel binning, drift duration, drift depth, detector channel start, detector channel stop). Radiometric

167 quantities are then retrieved following the TriOS recommendations (TriOS, 2025), detailed in
168 Appendix A.

169 E_d and L_u data in physical units (respectively $W\ m^{-2}\ nm^{-1}$ and $W\ m^{-2}\ nm^{-1}\ sr^{-1}$) are available at
170 <ftp://ftp.ifremer.fr/ifremer/argo/>.

171

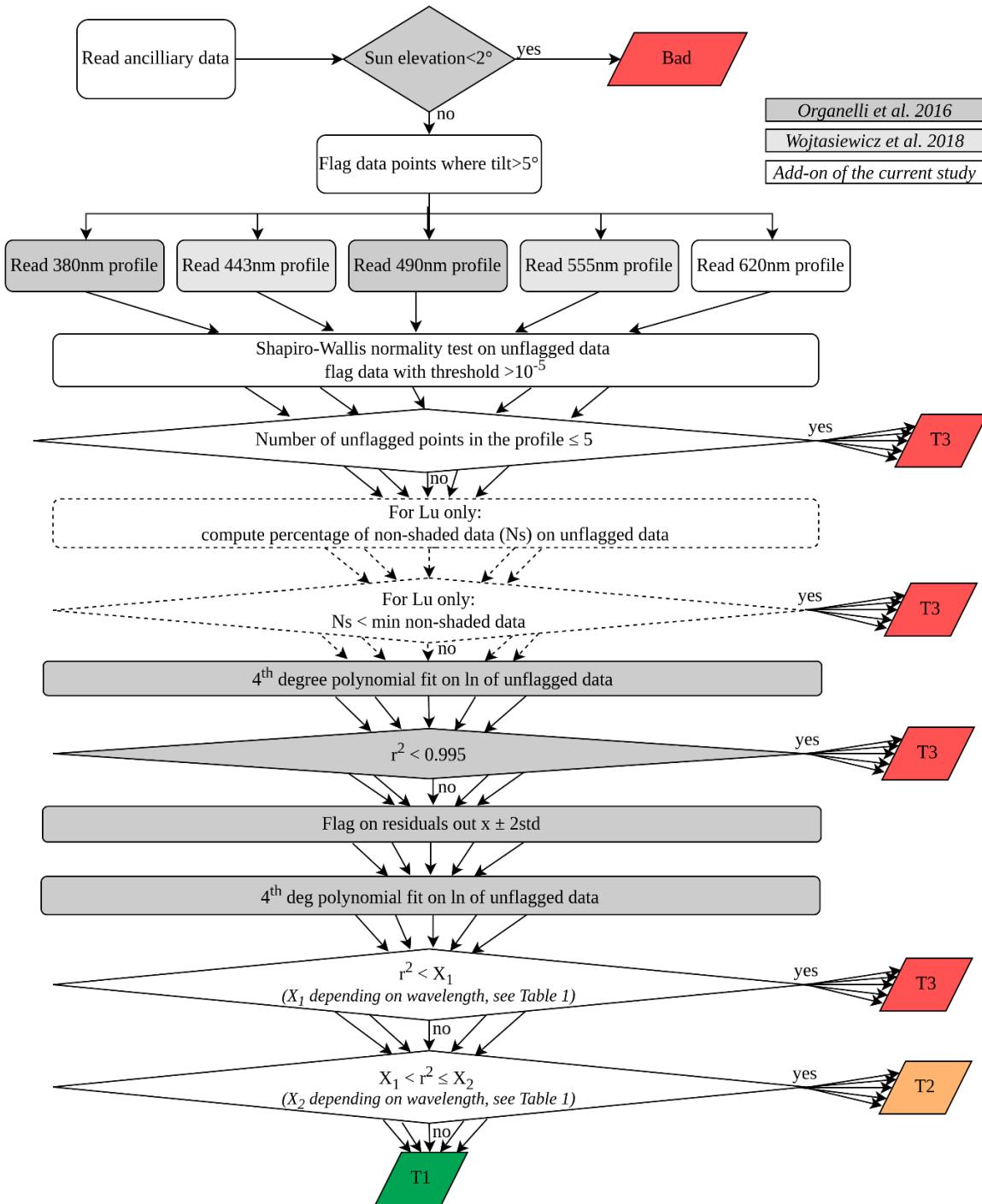
172 **3. Quality control protocol**

173 The following section details the different steps of the QC protocol specifically adapted for
174 hyperspectral floats. It should be noted that this QC procedure and flags are not equivalent to those
175 applied as part of the standard Argo data quality framework. Argo real-time QC applies a suite of
176 automatic tests to core physical and biogeochemical variables (e.g., global range check, pressure
177 gradient, spikes), aiming at identifying only if the instrument performance is satisfactory (Poteau
178 et al., 2019). Thus, multispectral E_d data are not subject to screening for light perturbations such
179 as wave focusing. Instead, this present QC evaluates the light environment at the time of the
180 measurements, which can affect the subsequent derivation of optical quantities (e.g., remote
181 sensing reflectances and diffuse attenuation coefficients), a need previously identified (Claustre et
182 al., 2020). Building on previously published QC approaches for multispectral radiometry
183 (Organelli et al., 2016; Wojtasiewicz et al., 2018), this method is designed for hyperspectral
184 radiometers to detect the occurrence of passing clouds, self-shading, large tilt angles, wave-
185 focusing effects, and spike events along a profile, without attempting to attribute these
186 perturbations to their physical causes.

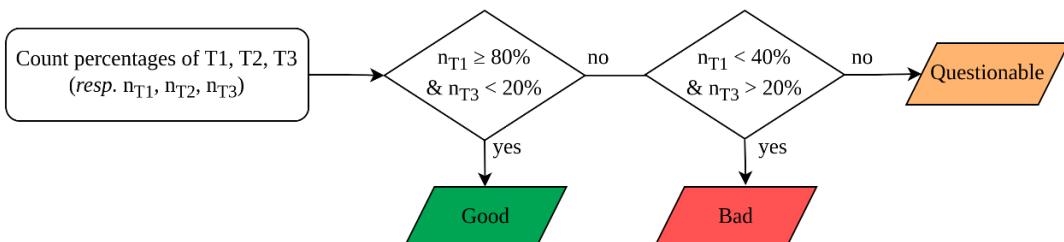
187 The first QC step is to determine whether the profile is acquired at daylight and flag data points
188 where the tilt angle of the float remains below 5° (Zibordi et al., 2019). Then, the method
189 distinguishes the section of the profile containing a radiometric signal from the section dominated
190 by instrument noise (i.e., dark current). Quality control will only be applied on the “signal layer”
191 (i.e., the layer where the signal is significantly greater than zero), which varies over wavelengths,
192 time of day, and water composition. As a result, the depth at which the QC is performed will differ
193 across profiles and wavelengths. The QC protocol begins with a check of individual data points,

194 followed by verification of the vertical profile shape at five “reference” wavelengths—380, 443,
195 490, 555, and 620 nm (see Section 3.b for justification), and each obtains a wavelength-specific
196 quality flag after two successive 4th degree polynomial fits and an optional shaded data filter on
197 L_u measurements only. The first polynomial fit detects passing clouds and large spikes, whereas
198 the second one, applied on data points without residuals from the fit, flags weaker perturbations
199 such as wave focusing (Organelli et al., 2016). The last step is to assign a final flag to the full
200 hyperspectral profile.

201 The overall procedure is summarized in Fig. 2. Following the convention defined in Organelli et
202 al. 2016, for each wavelength, three profile types are introduced. “type 1” qualifies good profiles,
203 that can readily be used for bio-optical applications, “type 3” qualifies bad profiles, affected by
204 environmental disturbances, while “type 2” corresponds to “questionable” profiles, with
205 potentially usable measurements for bio-optical application but who could require some case-by-
206 case additional processing/data manipulation (see examples in Fig. 3).

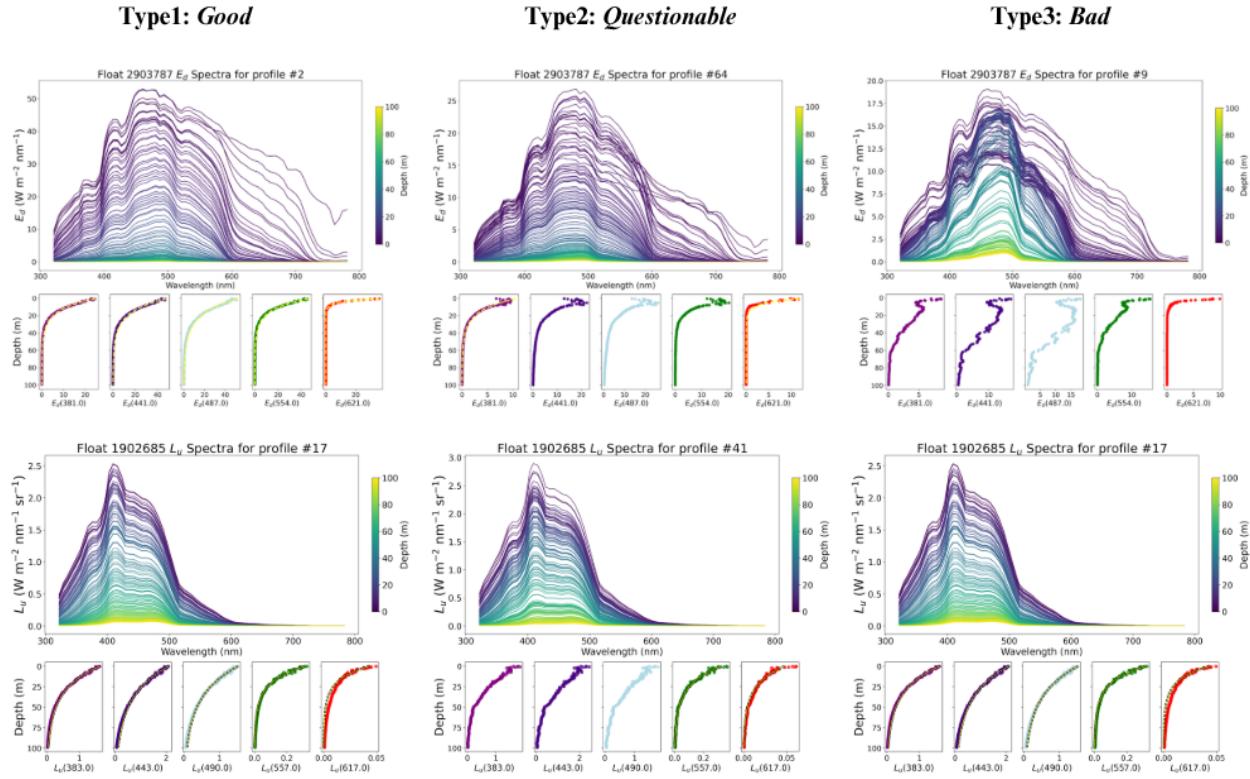


Overall quality of the spectrum



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209
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Fig 2: Flowchart of the QC procedure for hyperspectral data with the same conventions as in (Organelli et al., 2016): “x” is the mean of residuals, and T1, T2, T3 refer to the profile “type”. For the last step, “ n_{Ti} ” refers to the number of wavelengths with a given type. The colors of the boxes refer to the method from which the step is taken: Organelli et al., 2016 (dark gray), Wojtasiewicz et al., 2018 (light gray), and the QC procedure described in this paper (white).



213

214 **Fig 3:** Top row: Example of the three different E_d profile types acquired by float WMO 2903787. On the left-hand
215 side at the top, a type 1 “Good” profile. Type 2 profile is in the middle, with some wave-focusing at the intermediate
216 wavelengths. On the right side, a type 3 “Bad” profile, likely affected by passing clouds. In dashed yellow, the 4th
217 degree polynomial fit to each wavelength’s E_d profile, on a linear scale, from the computed K_d (only computed when
218 the wavelength is classified as type 1 or type 2). On the bottom row: three different L_u profile types acquired by float
219 WMO 1902685. Same order as for the E_d profiles, with, from left to right, type 1 to type 3 classification.

220

221 a. Selection of the different wavelengths

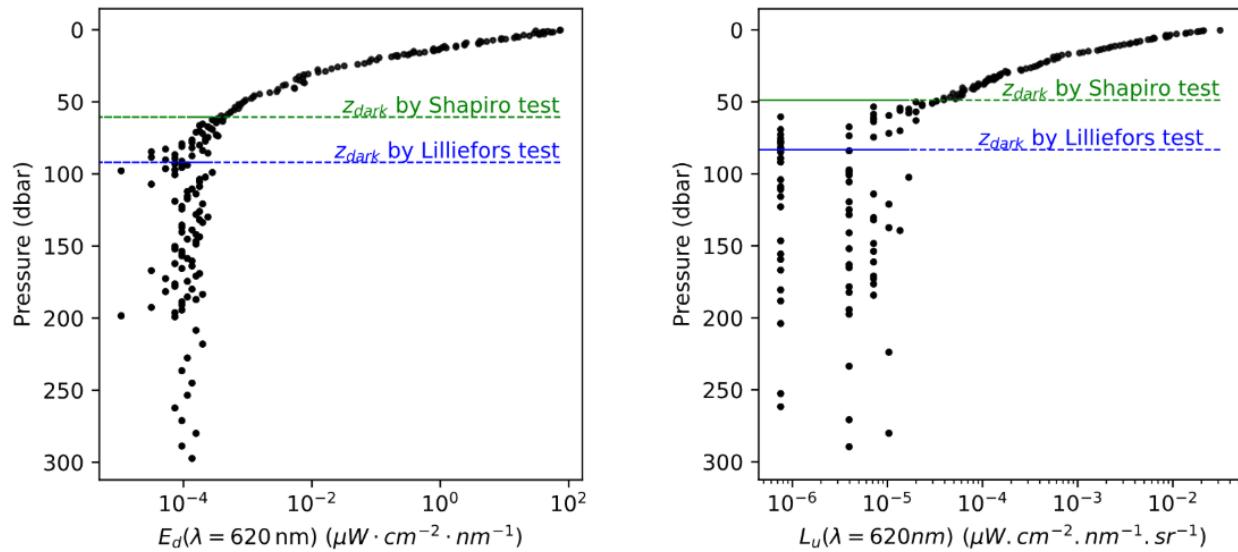
222 Although the present QC procedure is designed for hyperspectral data, it is performed on a subset
223 of wavelengths, ensuring robust quality control while achieving computational efficiency. The
224 comparison with a full hyperspectral QC approach is discussed in Section 4.e.
225 The reference wavelengths were selected based on those adopted by the BGC-Argo community

226 for multispectral float measurements and those typically used by satellite ocean-color missions
227 (e.g., Sentinel-3 OLCI, MODIS-Aqua, VIIRS). The 380 nm and 490 nm wavelengths are used for
228 the QC in Organelli et al. (2016) and correspond to two of the channels on the multispectral E_d
229 sensor from the BGC-Argo program (OCR-504, Sea-Bird Scientific). 443 nm and 555 nm are used
230 in the QC of Wojtasiewicz et al. (2018) for both L_u and E_d and will be on the new version of
231 multispectral E_d sensors adopted by the BGC-Argo community. The last wavelength (620 nm) is
232 used to evaluate the red part of the visible spectrum. The specific 620 nm value matches the
233 Sentinel-3 OLCI red band (Oa7, Donlon et al., 2012). In practice, since the wavelengths can differ
234 by about 4 nm between radiometers, the sensor's wavelength closest to each reference wavelength
235 is selected.

236 *b. Wavelength-specific signal layer detection*

237 The depth range over which QC is applied varies by wavelength, as light attenuates differently at
238 different wavelengths. To ensure consistency, QC is performed over the full depth range over
239 which a light signal is detected for each wavelength. To differentiate the environmental dark signal
240 from the light signal, a normality test was performed on the data, as dark values are normally
241 distributed as a function of depth, whereas light measurements are not (Organelli et al., 2016). We
242 define z_{dark} as the threshold depth, corresponding to the level below which light is effectively
243 negligible. During the gradual shift from very low light levels to darkness, the light is not
244 necessarily completely null around z_{dark} . Assuming z_{dark} as the dark threshold potentially leads to
245 the exclusion of a small portion of the profile, although this portion has a negligible amount of
246 light. Following La Forgia & Organelli (2025), the Shapiro-Wilk test (Shapiro & Wilk, 1965) of
247 normality was applied to all the selected bands with a p-value threshold of 10^{-5} , as the Lilliefors
248 test (Thode, 2002) adopted by Organelli et al. (2016) was not sensitive enough for the 620 nm
249 band, resulting in lower sections of the profile flagged as light when they corresponded to a dark
250 signal (Fig. 4). The Shapiro–Wilk test evaluates whether a set of values is drawn from a normally
251 distributed population. It is based on comparing the ordered sample values to the expected values
252 of a normal distribution using a correlation-type statistic (the p-value). A high p-value, therefore,
253 indicates that the distribution of values is statistically indistinguishable from Gaussian noise, i.e.,

254 the depth range where the signal becomes dominated by instrumental dark current. In practice, the
 255 Shapiro-Wilk test was processed iteratively: for each pressure level P^* associated with a spectrum,
 256 the p-value was derived from all the data points associated with a pressure higher than P^* (data
 257 from greater depth where less light is expected). This profile-by-profile approach has the
 258 advantage of being dependent on data distribution only, avoiding noise equivalent irradiance (NEI)
 259 estimation or sensor sensitivity drift issues.



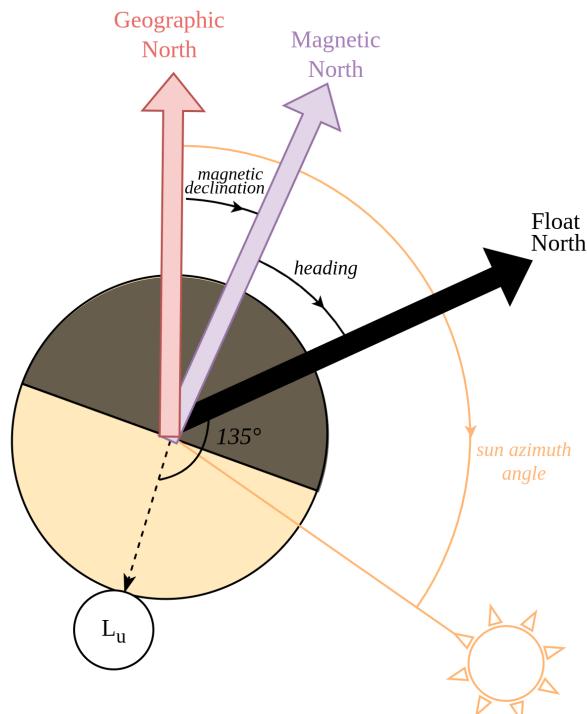
260
 261 **Fig. 4.** Comparison of the two dark detection methods on the E_d profile N° 58 at 620 nm of the float WMO 1902637
 262 (black dots) (left panel) and on the L_u profile N° 10 at 620 nm of the float WMO 1902637 (right panel). The dashed
 263 lines indicate the limit depth (z_{dark}) between the “signal” layer and the dark layer. Both z_{dark} s are derived from statistical
 264 tests for normality of data: Lilliefors test (Thode, 2002) in blue and Shapiro-Wilk (Shapiro & Wilk, 1965) in green.

265 *c. Detection of shaded points (L_u specific steps)*

266 Due to the float geometry, L_u measurements are strongly dependent on solar azimuth angle.
 267 Ensuring data quality therefore requires the ability to identify shaded measurements. For this
 268 reason, all hyperspectral floats are equipped with an inertial measurement unit (IMU) that includes
 269 a compass providing the heading, defined as the angle between float north and the magnetic north
 270 (resp. black and purple arrows, Fig. 5). Float north is defined by the manufacturer and aligned with
 271 the UVP sensor. Compass measurements are recorded as digital counts and stored alongside the

272 radiometric data, along with the equations needed to convert them to physical units and correct for
 273 magnetic deviation (i.e. error of the compass caused by local magnetic field created by the float,
 274 see details in Appendix B). The Python scripts accompanying this paper compute the solar azimuth
 275 angle relative to geographic north. To align the reference frames, the magnetic declination—
 276 defined as the angle between geographic and magnetic north (resp. red and purple arrows, Fig.
 277 5)—is derived using the World Magnetic Model 2020 (NOAA NCEI, 2024) and applied to convert
 278 the measured heading to a geographic north-referenced orientation. Finally, using the known angle
 279 between the radiance sensor and the float north provided by the manufacturer (135° , Fig. 5), data
 280 points for which the radiometer is on the sunny side of the float can be identified following the
 281 equation below:

282 $\text{sun azimuth angle} - (\text{heading} + \text{magnetic declination}) \in [45^\circ; 225^\circ]$



283

284 Fig. 5. View from above a hyperspectral float. When the Sun lies within the yellow sector of the circle (representing
 285 the float body), the measurements are considered non-shaded. When the relative orientation of the float and the Sun
 286 places the Sun within the dark sector, the corresponding L_u data are flagged as shaded by the float's body. The Sun's
 287 position relative to the float's North (black arrow) is calculated as the solar azimuth angle (saa) minus the sum of the

288 float heading (i.e. the angle between magnetic and float north, which relies on the compass measurement) and the
289 magnetic declination (i.e. the angle between geographic and magnetic north).

290

291 Although the presented method flags shaded data points and allows users to specify a minimum
292 acceptable number of non-shaded measurements throughout the profile (below which a profile is
293 classified as Bad), the results presented hereafter in this paper follow the recommendations of
294 Gerbi et al. (2016) by retaining all data points, both shaded and non-shaded, to preserve sufficient
295 data density.

296 *d. Revision of the coefficients for profile type assignment*

297 Once z_{dark} has been estimated by the Shapiro-Wilk method, all data below this depth are flagged
298 and will be ignored further in the QC processing. The same method developed by Organelli et al.
299 (2016) is applied to identify moving clouds, wave focusing, and spikes: a 4th-degree polynomial
300 fit on the Napierian logarithm (\ln) of E_d and L_u , followed by the analysis of r^2 and the residuals of
301 the difference between the polynomial fit and the data points (Fig. 2 and Fig. 3).

302 The threshold values for r^2 and residuals to determine the flag of each datapoint are wavelength-
303 dependent, but apply to both E_d and L_u . The original QC threshold values for the 380, 443, 490,
304 and 555 nm wavelengths are preserved (Organelli et al., 2016; Wojtasiewicz et al., 2018). At 620
305 nm, we slightly relaxed the criteria by choosing a threshold of 0.995 for the lower limit and 0.998
306 for the upper limit of type 2. This change accounts for a less robust polynomial fit in the red
307 wavelengths because of the presence of inelastic scattering and the reduced light signal due to a
308 stronger attenuation by water, which lowers the signal-to-noise ratio (see Section 4.c.). Thresholds
309 are summarized in Table 1.

310

| Band / Flag | X_1 | X_2 |
|-------------|-------|-------|
| 380 nm | 0.997 | 0.999 |

| | | |
|--------|-------|-------|
| 443 nm | 0.996 | 0.998 |
| 490 nm | 0.996 | 0.998 |
| 555 nm | 0.996 | 0.998 |
| 620 nm | 0.995 | 0.998 |

311 **Table 1.** Thresholds of r^2 used for the profile type identification for each wavelength. X1 and X2 denote the lower
 312 and upper limits of r^2 , respectively (see Fig. 2).

313 *e. Overall quality of the hyperspectral profile*

314 Once each of the 5 wavelengths is assigned to a “type”, the overall quality of the spectrum is
 315 determined based on the frequency of each of the three “types”. Each criterion is detailed below
 316 and in the last step of the flowchart (Fig. 2), and an example of spectra corresponding to each type
 317 is available in Fig. 3. To summarize:

318 ● If 80% or more of the reference wavelengths’ profiles are categorized as “type 1” and less
 319 than 10% as “type 3”, the radiometric profile (the hyperspectral E_d or L_u profile) is labeled
 320 as “Good”.

321 ● If more than 20% of the reference wavelengths’ profiles are categorized as “type 3” and
 322 less than 40% as “type 3”, the radiometric profile is flagged as “type 3” (Bad).

323 ● For all other cases, type 2 is used by default, i.e. the data quality is considered
 324 “Questionable”.

325

326 *f. An extension to a hyperspectral method*

327 To analyze whether this 5-wavelengths approach (called 5-QC hereon) is sufficient to spectrally
 328 qualify a given radiometric profile from 320 to 780 nm, we compared the results of the proposed

329 QC method with a hyperspectral version of it (called Hyper-QC hereon). The Hyper-QC
330 characterizes profiles for every single wavelength available. The above-mentioned criteria were
331 thus adapted to obtain the Hyper-QC approach. Correlation coefficient thresholds depend on
332 whether wavelengths are shorter or longer than 600 nm. Details are presented in Table 2.

| λ / Flag | X ₁ | X ₂ |
|-----------------------|----------------|----------------|
| $\lambda < 600$ nm | 0.997 | 0.999 |
| $\lambda \geq 600$ nm | 0.995 | 0.998 |

333 **Table 2.** Thresholds of r^2 used in profile type identification for the Hyper-QC approach. X₁ and X₂ denote the lower
334 and upper limits of r^2 , respectively.

335 Once all wavelengths for a given radiometric profile are quality controlled, an overall quality flag
336 is assigned to the whole profile, following the same criteria as the 5-QC described above, but only
337 applied to wavelengths < 650 nm. Note that the number of wavelengths can vary between
338 hyperspectral floats depending on the sampling channels' settings.

339 The overall Hyper-QC procedure follows the same steps as the 5-QC procedure, presented in Fig. 2,
340 with differences only in the threshold values.

341 **4. Results and Discussion**

342
343 The development of hyperspectral sensors will significantly enhance our ability to detect bio-
344 optical features, such as chlorophyll absorption bands, fluorescence peaks, and Raman scattering,
345 that multispectral instruments cannot resolve (Jemai et al., 2021; Organelli et al., 2021). Ensuring
346 the quality of these full-spectrum measurements is critical for validating upcoming hyperspectral
347 satellite missions (e.g., PACE) and for deriving spectral products such as hyperspectral Kd,
348 spectral reflectance, and pigment-specific absorption metrics. Therefore, a QC procedure
349 encompassing the full hyperspectral range enables robust identification of wavelength-dependent

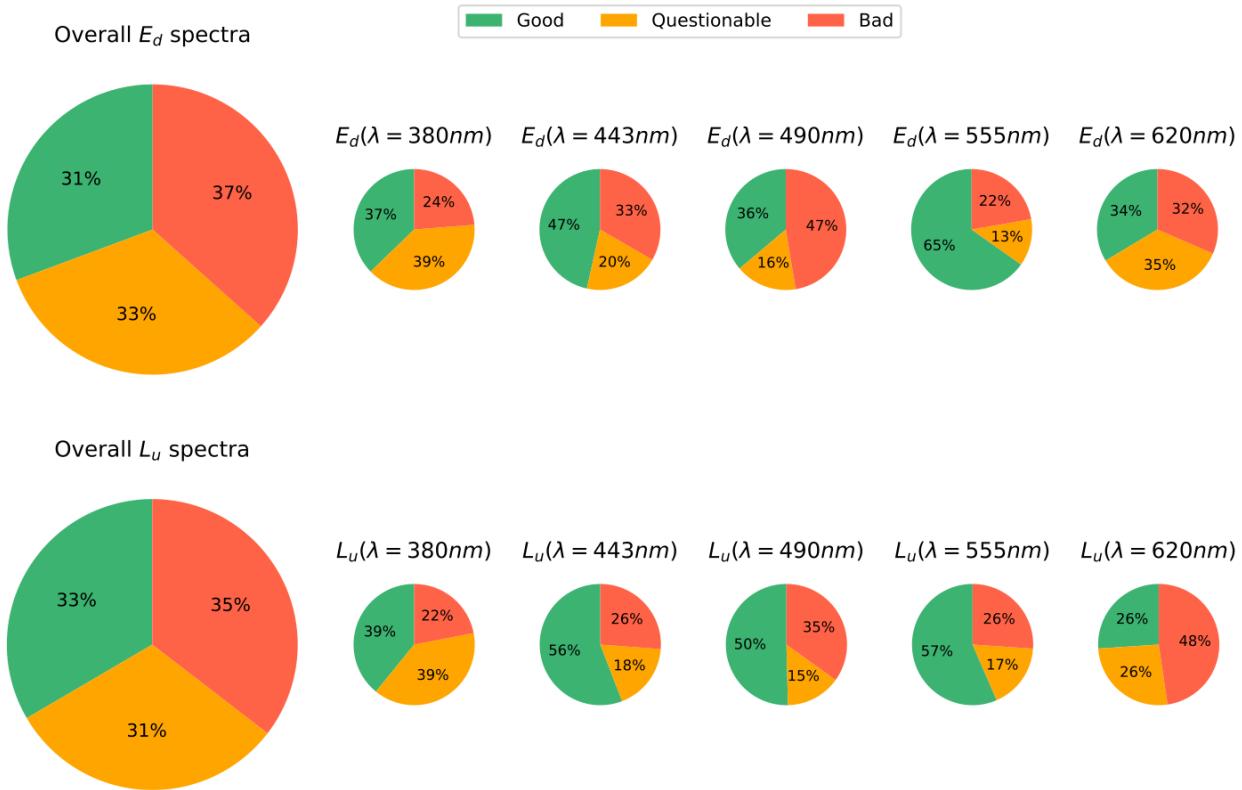
350 perturbations in the light field. This, in turn, increases confidence in applications that leverage the
351 hyperspectral capability of the measurements.

352 We elected to use the same curve-fitting QC-based methodology for both E_d and L_u . Indeed,
353 applying a QC originally designed for E_d to L_u is justified given that perturbations in the light field
354 affect both E_d and L_u (albeit differently). Additionally, polynomial and regression-based screening
355 of hyperspectral L_u profiles is already used in autonomous systems supporting satellite calibration
356 (e.g., HyperNav, Barnard et al., 2024). A further discussion on the difference in QC result between
357 the two variables can be found below in Section 4.c.

358

359 *a. Quality of the profiles with the 5-QC*

360 A total of 899 radiometric profiles from a large diversity of ocean regions and weather conditions
361 (Fig.1) were QC-ed with the methods presented in section 3: for the E_d spectra, ~31% of the
362 profiles were labeled as type 1 (ie: “Good”), 33% of the spectra are considered “Questionable”
363 and ~37% are qualified as “Bad” (Fig. 6). For L_u , 33% of hyperspectral profiles were type 1, 31%
364 type 2, and 35% were identified as type 3.



365

366 **Fig. 6.** Distribution of E_d profile (upper panel) and L_u profile (lower panel) types for each wavelength and for the
367 overall hyperspectral profile.

368 There are two ways in which a profile can show variability in the 5-QC method: between the
369 different wavelengths of the same radiometric profile (inter-wavelength variability, discussed in
370 4.b) or between E_d and L_u profiles that are measured simultaneously (discussed in 4.c).

371 *b. Inter-wavelength variability*

372 Significant variability in 5-QC classification exists across wavelengths of a given measurement
373 (Fig. 6). Importantly, this variability does not indicate that certain spectral regions are unsuitable
374 for the 4th-order polynomial test. Rather, it reflects wavelength-dependent differences in the
375 physical and biological processes shaping the radiometric profiles, which motivates the use of
376 wavelength-dependent r^2 thresholds when discriminating between quality levels. This can be
377 explained by the predominantly oligotrophic nature of the waters sampled by the floats, with a

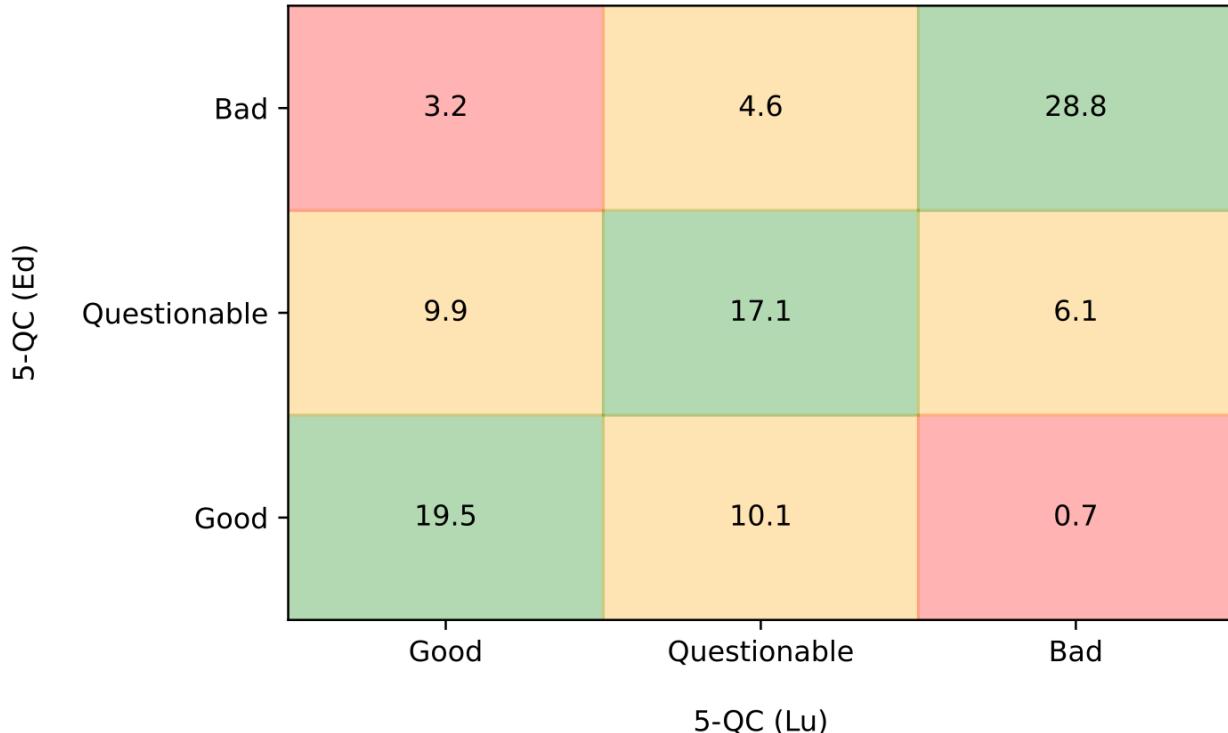
378 strong signal in the blue-green region (443, 490, and 555 nm in this study), penetrating deep into
379 the water column (Morel et al., 2007; Pope & Fry, 1997). Consequently, at those wavelengths, the
380 QC is often performed over greater depths, increasing the number of available data points for the
381 polynomial fitting and effectively reducing the impact of outliers near the surface on the fits
382 performance. On the other hand, the float spends more time in the “signal” layer (as opposed to
383 the dark layer where the QC is not performed), increasing the likelihood of weather perturbations
384 (e.g., passing clouds).

385 Phytoplankton absorbs significantly more light at 490 nm than at longer wavelengths like 555 nm
386 (Bricaud et al., 1995; Mobley, 2022). E_d profiles at 490 nm are therefore more susceptible to
387 variations caused by the float crossing the Deep Chlorophyll Maximum (DCM), leading to changes
388 in K_d . These changes can introduce an uneven decrease of E_d with depth as it moves across water
389 masses with different phytoplankton concentrations, ultimately degrading the fit quality.
390 For L_u , the largest number of type 3 profiles is for the red wavelength. In the red wavelengths,
391 absorption of light by pure water is strong, and most of the light disappears in the first few meters.
392 The resulting shape of the L_u profile is therefore different than at other wavelengths, resulting in
393 lower performance of the polynomial fit, regardless of biology and instrument performance. As
394 phytoplankton naturally fluoresces with a relative maximum at ~685 nm (Gordon, 1979), the L_u
395 profile shape can deviate from log-normal linearity for the wavelengths around, such as our
396 reference wavelength at 620 nm, which contributes to explain why polynomial fit performances
397 can be affected. Lastly, the red wavelengths are affected by Raman scattering, an inelastic
398 scattering process where the incident blue light excites water molecules. As the water molecule
399 decays post-excitation, it emits light at a longer wavelength (lower energy), i.e., in the visible
400 spectrum's red wavelengths (Mobley, 2022). Therefore, the overall L_u/E_d profile is affected not
401 only by constituents present in the water but also by a combination of pure water absorption,
402 phytoplankton fluorescence, and Raman scattering. Those additional processes do not follow the
403 same exponential decrease with depth, therefore the red wavelength can not be described by the
404 polynomial fit as accurately as other wavelengths.

405

406 *c. E_d versus L_u variability in QC type*

407 Although E_d and L_u are measured almost simultaneously but with two different sensors, we
408 recommend performing independent QC on each hyperspectral profile. As a result, the same float
409 radiometric cast can be assigned two different types, depending on whether looking at the L_u or
410 the E_d sensor data. The new QC analysis resulted in around 65% of profiles being assigned to the
411 same type for the E_d and L_u (green boxes in Fig. 7) and 35% of differences, including 3.9% of
412 these profiles exhibiting a type 3/type 1 discrepancy (red boxes in Fig. 7). These differences,
413 depending on which sensor is QC-ed, are primarily due to two reasons. First, E_d corresponds to
414 the light going downward, hence the sensor points upward, toward the sky. When the float is close
415 to the surface, the E_d sensor is more sensitive to wave focusing (Gege & Pinnel, 2011) than the L_u
416 sensor (which points downward). This can affect the r^2 coefficient of the polynomial fit, thus
417 degrading the profile “type” of E_d as compared to L_u , especially in the blue (443 and 490 nm
418 channels). As such, the percentages of type 1 profiles for those wavelengths are lower for E_d than
419 L_u (47% vs. 56% and 36% vs. 50% for 443 and 490 nm, respectively -Fig. 6).



420

Fig. 7. Comparison matrix between QC classification of E_d profiles (on the y-axis) and QC classification of L_u profiles (on the x-axis). Green cells indicate the percentage of profiles for which the overall flag is the same for L_u and E_d ; orange cells correspond to the percentage of profiles with a difference of only one QC level between E_d and L_u (e.g. Questionable/Bad or Good/Questionable); red cells represent the percentage of bad E_d profiles for good L_u profiles or vice versa.

426 Secondly, given that L_u refers to light going upward from the water, and that water itself strongly
 427 absorbs at larger wavelengths (>500 nm; Pope & Fry, 1997), the amount of upward light at those
 428 wavelengths is significantly lower than downwelling light, resulting in $E_d(\lambda > 500) \gg L_u(\lambda > 500)$
 429 (Mobley, 2022). A small amount of light is not necessarily indicative of a lower signal; the TriOS
 430 radiometers modulate the integration time based on the amount of light, with a longer integration
 431 time leading to a stronger signal. This process is performed based on the channel that saturates
 432 first (usually the blue wavelength) (see details in Section 2.b; TriOS, 2025 and Mobley, 2022). As
 433 light penetrates deeper in the water column, given that red wavelengths attenuate faster than blue
 434 ones, the blue/red light ratio increases. Therefore, despite an increased integration time at depth
 435 (to keep the blue signal saturated), the red signal is lower, due to the faster decrease with depth.

436 A low signal, especially in the red, makes the fit performance worse (lower signal-to-noise ratio)
437 and thus partly explains the larger amount of 'Bad' L_u profiles at 620 nm (Fig 6; Fig 7) than E_d
438 profiles at 620 nm.

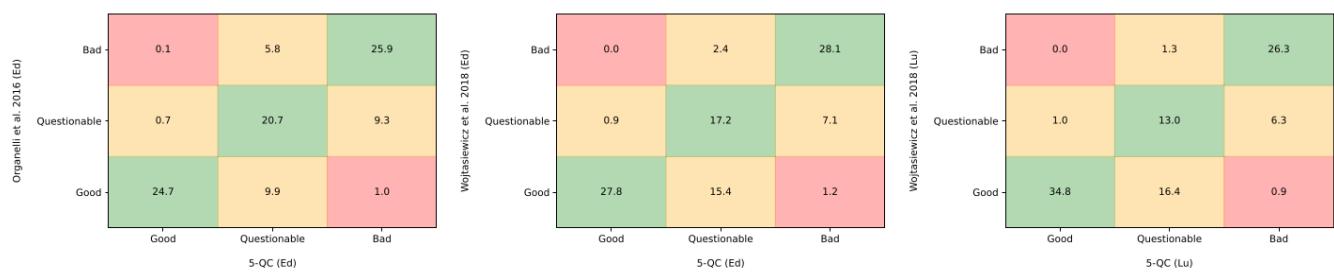
439 *d. Comparison with established methods*

440 To enable a more robust comparison of method performance, we applied all three classification
441 approaches (the 3-wavelengths by (Organelli et al., 2016), 4-wavelengths by Wojtasiewicz et al.,
442 2018, and the new 5-wavelengths methods) to the hyperspectral TriOS dataset. This is possible
443 because the TriOS-RAMSES radiometers include all the relevant bands from both versions of the
444 OCR-504 sensor. To directly compare the three method outputs, we computed a global flag after
445 performing the 3-wavelengths and 4-wavelengths QC based on the proportion of
446 Good/Questionable/Bad individual flags, following the percentages described in Section 3.d.

447 Overall, the QC distribution is consistent across the three methods (Fig. 8). The 5-QC classification
448 aligns with the 3-QC 71.3% of the time and with the 4-QC at rates of 73.1% for E_d and 74.1% for
449 L_u . Major differences in QC classification (type 1 vs. type 3) happened for 1.1%, 1.2%, and 0.9%
450 of profiles, respectively. Most of the differences arise from the previous QC methods classifying
451 profiles one level "better" than the 5-QC. For example, a profile categorized as "bad" by the 5-QC
452 is often labeled as "questionable" by the 3-QC and 4-QC, while a "questionable" profile in our
453 classification is frequently marked as "good" by the other two methods. The strong correlation
454 between methods for the same profiles is also consistent with the RAMSES radiometer design,
455 which relies on a single diffuser and thus can yield similar spectral responses (aside from effects
456 related to wavelength-dependent light penetration). While the three quality control methods are
457 largely similar, differing mainly in dark-signal and tilt detection and a slight variation in the r^2
458 cutoff value, we conclude that the primary source of their differences lies in the choice of
459 wavelengths.

460 As discussed in Section 4.b. and 4.c, red wavelength profiles are unlikely to be impacted by
461 weather phenomena when the float is at deeper depths (Fig. C1). However, they exhibit poor
462 performance particularly for L_u (Fig. 6), owing to their lower signal-to-noise ratio compared to

463 other parts of the spectrum, and the presence of inelastic scattering. These factors help explain part
 464 of the discrepancy in the overall quality assigned by the 5-QC compared to the 3-QC and the 4-
 465 QC approaches (Fig. 8). This highlights the importance of hyperspectral quality control to account
 466 for variability across the entire visible spectrum, as bio-physical disturbances do not have a
 467 uniform impact across wavelengths.



468

469 **Fig. 8.** Comparison matrix of 5-QC procedure with Organelli et al. 2016 method (3-QC, left panel)
 470 and Wojtasiewicz et al. 2018 method (4-QC, center panel) on E_d profiles. The right panel shows the
 471 comparison matrix of our QC procedure and Wojtasiewicz et al. (2018) method on L_u profiles.

472

473 *e. Multispectral QC method for hyperspectral profiles*

474 A comparison between the 5-QC and the Hyper-QC shows 69.1% and 70.5% of correspondence
 475 for E_d and L_u respectively. The significant correlation between the two methods shows that the 5-
 476 wavelengths chosen are representative of the full spectrum profile (Fig. 9). This strong correlation
 477 arises because the RAMSES radiometer is built with a single optical collector with an unique field
 478 of view. As a result, any wave-focusing effects—most pronounced in the upper layers—impact
 479 the signal across all wavelengths in a similar way, producing consistent responses throughout the
 480 spectrum and thus increasing the wavelength-to-wavelength correlation. Visualisations of the 3
 481 profiles where the 5-QC and the Hyper-QC disagree are provided in Appendix C.

482 The main justification behind implementing the 5-QC method to QC hyperspectral data was to
 483 reduce computing costs and the effect of subjective threshold choices. Performing the QC over 5
 484 wavelengths rather than 70 (or 140, depending on pixel settings) results in a computing time $\sim 16x$

485 faster than doing it on the full set of wavelengths. Over the 899 profiles of the BGC-Argo dataset,
486 this results in a ~2.5h run time for the 70-QC versus 12 minutes for the 5-QC (see Table 3).

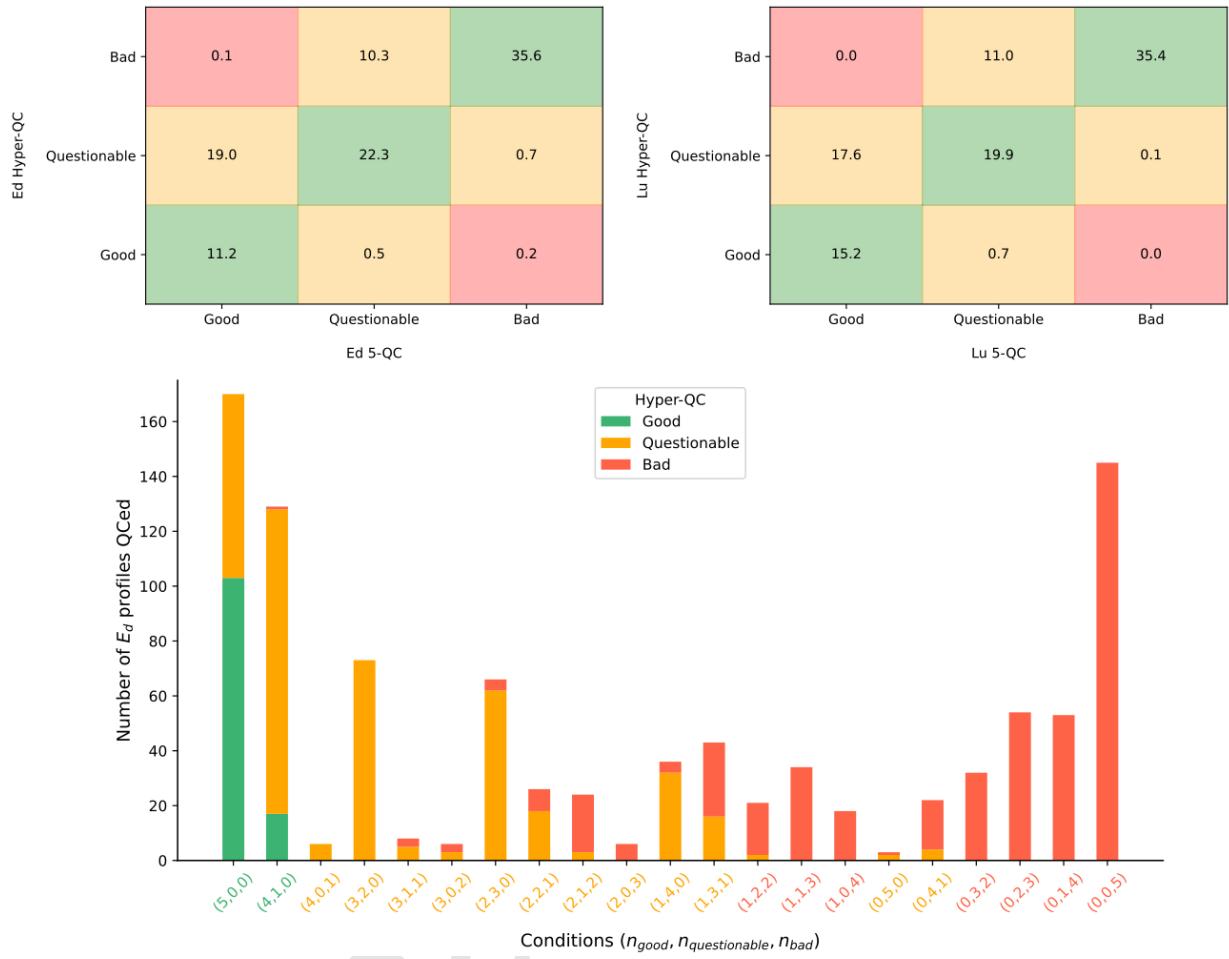
487

| Number of wavelengths in QC | Run time for 1 $E_d + L_u$ pair (min) | Run time for 899 $E_d + L_u$ pairs (min) |
|-----------------------------|---------------------------------------|--|
| 70 | 0.171 | 153 |
| 5 | 0.013 | 12 |
| 3 | 0.007 | 65 |

488 **Table 3.** Speed test of the computation of a wavelength-specific QC and the generation of a global spectrum flag for
489 a single hyperspectral BGC-Argo float. The test was performed on a M1 Macbook, 16GB RAM. Each row refers to
490 the amount of wavelength used to generate a global flag.

491

492 Although we have shown here that using the 5-QC represents a significant gain of time while
493 providing thorough QC, for some applications interested in specific wavelengths of the spectra
494 (e.g. chlorophyll- α natural fluorescence relative maximum at 690 nm, or hyperspectral K_d), it
495 would be more appropriate to do a QC on the specific wavelengths of interest. The associated
496 Python code was designed to allow for the user to easily modify both the number of wavelengths
497 and the thresholds used.

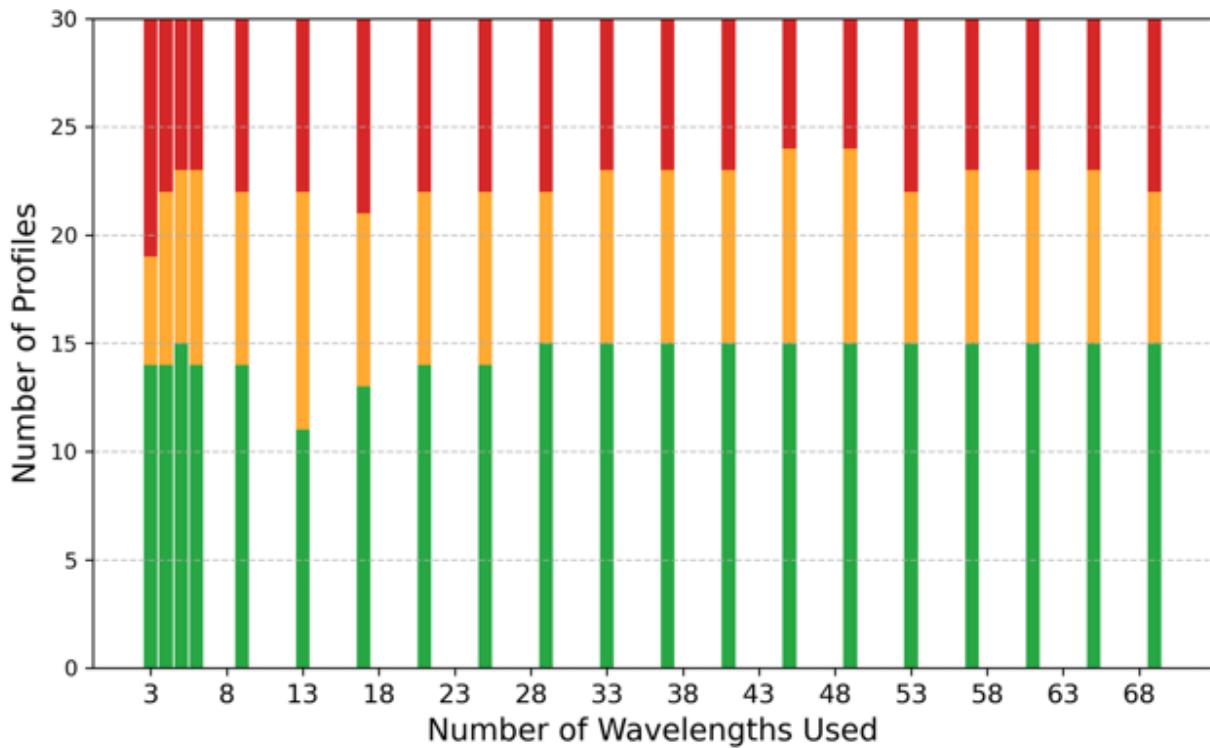


498

499 **Fig. 9:** At the top, a comparison table between the 5-QC and the Hyper-QC by overall type for both E_d (top left panel)
500 and L_u profiles (top right panel). At the bottom, histogram of the distribution of the overall QC by conditions for the
501 5 wavelengths, colored by the hyperspectral QC result. The x-label colors correspond to the 5 wavelengths QC result.
502 For example, the first x-tick is associated with the condition “(5,0,0)” which means: “5 wavelengths good, 0
503 wavelengths questionable and 0 wavelengths bad”. It is written in green as it corresponds to a “Good” spectra for the
504 5-QC method. Among the ~170 profiles that meet this condition, ~100 of them are also flagged as “Good” by the 70-
505 QC approach (green part of the bar) while ~70 of them are flagged as “Questionable” (orange part of the bar) and no
506 “Bad” (red part of the bar).

507

508 When comparing the overall distribution of Good, Questionable, and Bad profiles depending on the amount
509 of wavelengths used to perform the QC, we notice few differences in the overall distribution (Fig. 10), with
510 the 5 wavelengths being slightly more stringent between Questionable and Bad.



511

512 **Fig. 10.** The QC was performed on a varying number of wavelengths, ranging from 3 to 69, on 30 random profiles
 513 from float 2903787. The overall flag (Good, Questionable, Bad) was determined for a given profile based on the
 514 individual wavelengths' flags following the criteria described in section 3.e.

515

516 *f. Profile depth qualification alternatives*

517 In the method discussed in this paper, the quality control of radiometric profiles is performed over
 518 a different depth for each of the 5 reference wavelengths, depending on the amount of light that
 519 penetrates the water column (See Section 3.b). This allows the QC of the full profile, which is
 520 interesting for diffuse attenuation coefficient derivation and other variables related with
 521 biogeochemical properties (e.g. euphotic depth, PAR, isolume depth). However, this raises an
 522 issue about the consistency of weather conditions along the spectra. As the blue radiations
 523 penetrate the water deeper than the red ones, for example, a cloud passing when the float is between
 524 150 and 100 m depth could affect the radiometric profile at 443 nm, that still can measure light at
 525 this depth, but not the 620 nm band, which has no light at this time and depth. For near-surface

526 applications, it might be relevant to instead choose a constant depth for each wavelength. This
527 threshold could be chosen as the minimum of the five wavelengths' penetration depths. The
528 associated Python code allows for such modifications.

529

530 **5. Conclusion**

531 The BGC-Argo program has revolutionized our ability to observe biogeochemical processes on a
532 global scale, providing unprecedented spatiotemporal coverage of essential ocean variables in
533 regions that are otherwise undersampled (open-ocean gyres or southern ocean), overcoming some
534 of the obstacles associated with traditional oceanographic cruises, such as bad weather conditions
535 and high cost (Stoer et al., 2023). Since 2012, multispectral E_d measurements from Argo floats
536 have contributed to populate a dataset of around ~60,000 profiles, spanning multiple oceanic
537 regions and supporting a wide range of bio-optical studies (Organelli et al., 2017; Organelli &
538 Claustre, 2019; Uitz et al., 2023). Notable applications include the SOCA (Satellite Ocean Color
539 merged with Argo data to infer bio-optical properties to depth) methodology, which is based on
540 an artificial neural network trained with E_d profiles acquired from BGC-Argo floats (Renosh et al.,
541 2023) and the validation of remote sensing products such as remote sensing reflectance (Gerbi et
542 al., 2016; Organelli, Barbeau, et al., 2017; Wojtasiewicz et al., 2018), K_d (Begouen Demeaux &
543 Boss, 2022; Xing & Boss, 2021) and Chlorophyll concentration (Begouen Demeaux et al., 2025;
544 Xing et al., 2011), demonstrating the value of these measurements for characterizing ocean optical
545 properties (Jemai et al., 2021; Pitarch et al., 2025).

546 The quality of the BGC-Argo data for such applications relies on a dedicated QC methodology
547 that can provide qualified profiles without operator control and independently of the weather
548 conditions. As hyperspectral radiometers are now being integrated into the BGC-Argo fleet,
549 ensuring the quality of these measurements is essential to fully realize their scientific potential.
550 This study presents an automated, globally applicable quality-control methodology for
551 hyperspectral E_d and L_u profiles that is independent of absolute light levels. Adapted from
552 established multispectral approaches, the method reliably identifies sensor-related artifacts and

553 unstable light conditions attributable to passing clouds, self-shading, large tilt angles, spikes, and
554 wave focusing, enabling the robust use of hyperspectral radiometry for bio-optical applications.
555 The associated Python implementation is openly available (<https://gitlab.com/published-work-on-hyperspectral-bgc-argo/hyperspectral-5qc/>), and hyperspectral K_d products derived using this
556 framework have been released through SeaBASS (Haëntjens, 2022).

558 While additional factors such as sensor drift and biofouling, have not been addressed in this study,
559 they remain critical for assessing the quality of radiometric measurements. Efforts have already
560 been made to characterize them on multispectral sensors (Antoine et al., 2008; Jutard et al., 2021)
561 and extending such analyses to hyperspectral instruments represents a valuable next step. In
562 addition, wavelengths beyond 650 nm, which were not addressed in this paper, could be of
563 particular interest for specific applications such as chlorophyll-*a* natural fluorescence or European
564 Space Agency FLuorescence EXplorer satellite mission validation (Vicent et al., 2016). For these
565 applications, adapted versions of the proposed quality-control approach could then be devised,
566 replacing for instance, the polynomial fit with an approach that uses the chlorophyll-*a*
567 concentration and Raman scattering estimations. Such developments would naturally benefit from
568 a more detailed analysis of wavelength-to-wavelength correlations across the hyperspectral range
569 following (Tan et al., 2024). In support of ocean-color satellite validation, particularly for new
570 hyperspectral missions such as NASA's PACE, future work will focus on developing and releasing
571 fully QCed datasets of L_w and R_{rs} derived from BGC-Argo hyperspectral floats.

572

573

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601 **Data Availability Statement.**

602 The associated Python script (<https://gitlab.com/published-work-on-hyperspectral-bgc-argo/hyperspectral-5qc/>) is freely available and customizable. The BGC-Argo hyperspectral

604 profiles types (from 5QC method) until June 2025 for both Ed and Lu are made available in the
605 Gitlab. A hyperspectral K_d product derived from E_d is downloadable in SEABASS (Haëntjens,
606 2022). E_d and L_u measurements by BGC-Argo floats in digital counts and in physical units are also
607 available online at <ftp://ftp.ifremer.fr/ifremer/argo/>.

608

APPENDIX

609 *Appendix A: Processing hyperspectral raw data for TriOS RAMSES mounted on Provor*
610 *BGC-Argo floats*611 Raw data in counts (x) are converted into scientific units (y) using calibration coefficients (B0 and
612 B1), integration time (t), and the calibration equations.613 1. First, the background spectrum y_{dark} is calculated using dark calibration coefficients
614 ($B0_{dark}$ and $B1_{dark}$ calculated as the average of the values measured by TriOS on the Dark
615 pixels) also referred to as background coefficients in the manufacturer documentation.
616 Then, we subtract the background from the normalized raw spectrum to obtain background-
617 corrected measurement (y_{bc}).

618
$$y_{dark} = B0_{dark} + B1_{dark} * \frac{t}{8192}$$

619
$$y_{bc} = \frac{x}{65535} - y_{dark}$$

620 2. The dark signal (x_{dark}) -defined as the average of 17 black detector channels (TriOS,
621 2025)- is then retrieved from y_{bc} . Finally y_{bc} is adjusted for the sensor's calibration factor
622 specific to seawater (S) defined by the manufacturer, and scaled to $\mu\text{W}.\text{cm}^{-2}.\text{nm}^{-1}$ or
623 $\mu\text{W}.\text{cm}^{-2}.\text{nm}^{-1}.\text{sr}^{-1}$ to obtain E_d or L_u respectively in physical units (y_{phys}). No time
624 interpolation is needed for the dark correction, given that the black channels are measured
625 simultaneously as the rest of the spectrum.

626
$$y_{phys} = \frac{y_{bc} - \frac{x_{dark}}{S}}{t} * \frac{8192}{t}$$

627

628

629 *Appendix B: Deriving heading from IMU raw data on Provor BGC-Argo floats.*

630

631 To derive heading angle from IMU raw data, we apply the following equations, assuming the
 632 verticality of the float, using magnetometer raw data in digital counts (Mag_count_x,
 633 Mag_count_y, Mag_count_z), magnetometer calibration coefficients (m_{x0} , m_{y0} and m_{z0}) and
 634 compass calibration coefficients (h_1 , h_2 si_{11} , si_{12} , si_{21} and si_{22}).

635 # 1. Orientation and simple calibration

636 PhyMag_x=Mag_count_x+m_{x0}

637 PhyMagy=Mag_count_z+m_{z0}

638 PhyMag_z=Mag_count_y+m_{y0}

639 Note that the y axis of the IMU corresponds to the vertical axis of the float.

640

641 # 2. Compass calibration

642 PhyMag_x=PhyMag_x+hi₁

643 PhyMag_y=PhyMag_y+hi₂

644

645 F_PhyMag_x=PhyMag_x*si₁₁ + PhyMag_y*si₁₂

$$646 \quad F_{\text{PhyMag}_y} = \text{PhyMag}_x * si_{21} + \text{PhyMag}_y * si_{22}$$

647

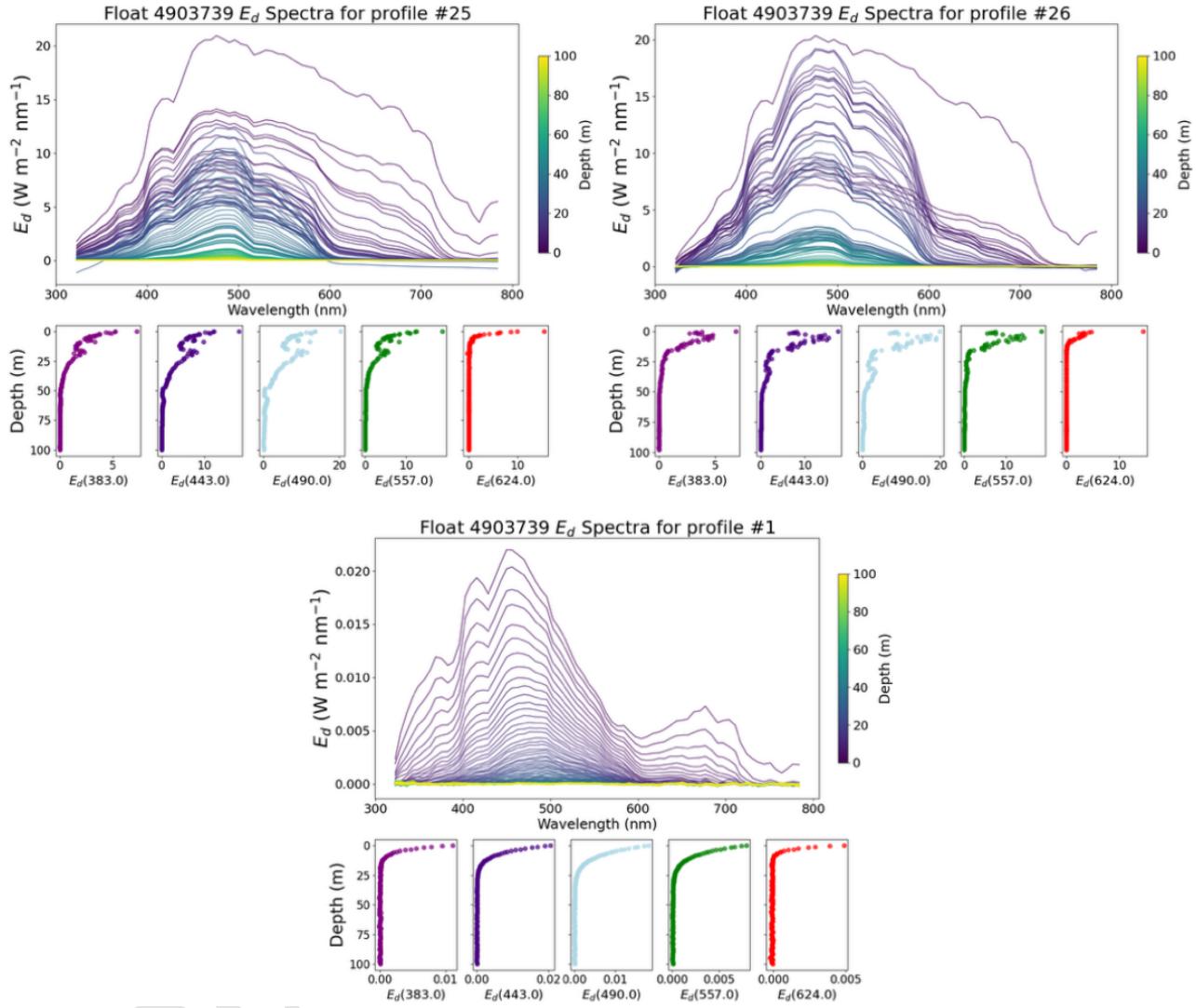
648 # 3. Heading computation

649 heading = arctan(F_PhyMag_y, F_PhyMag_x) *180.0 / π

650

651

Appendix C: Examples of spectra with different results depending on QC method chosen



652

653 **Figure C1.** Examples of spectra for which the former QC did not result in the same result as the newly-developed 5-
 654 QC. In both examples above, profiles at the 441nm and the 487nm did not pass quality control. However, higher
 655 wavelengths (554 and 621 nm) do pass quality control, resulting in a different overall flag depending on which
 656 wavelengths are considered.

657

658 In this example, profiles 25 and 26 from float 4903739 failed the 5-QC but passed the Hyper-QC.
659 Profile 25 is primarily affected by a pronounced spike at around 20 m depth, whereas profile 26
660 shows signatures consistent with wave focusing and the likely presence of a passing cloud. These
661 perturbations mainly influence the blue-green wavelengths, thereby affecting most of the
662 wavelengths used in the 5-QC. In contrast, they represent only a limited portion of the full
663 hyperspectral range (320–700 nm) considered in the Hyper-QC. Moreover, relatively deep
664 perturbations, such as the spike observed in profile 25, may be inconsequential for wavelengths
665 that penetrate less deeply into the water column, as the disturbance occurs below the wavelength-
666 dependent signal layer.

667 Conversely, profile 1 from float 4903739 is the only case in which the E_d profile passed the 5-QC
668 but failed the Hyper-QC. In this instance, very low ambient light levels lead to a highly sensitive
669 polynomial fit across all wavelengths, particularly in the near-UV and near-IR bands where the
670 signal is weaker than in the blue-green bands. As a result, the profile marginally satisfies the 5-QC
671 criteria, which emphasize blue-green wavelengths but fails the Hyper-QC which includes
672 numerous wavelengths for which the signal is already strongly absorbed by water, yielding
673 insufficient signal levels under these low-light conditions.

674 REFERENCES

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