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5 **New constraints from Central Chile on the origins of enriched continental compositions in thick-crust**  
6 **arc magmas**

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## 17 **Abstract**

18 Magmas from continental arcs built on thick crust have elevated incompatible element abundances and  
19 “enriched” radiogenic isotope ratios compared to volcanics erupted in island and continental arcs overlying  
20 thinner crust. The relative influence of the slab, mantle, and upper plate on this variability is heavily debated.  
21 The Andean Southern Volcanic Zone (SVZ; 33-46° S) is an ideal setting to investigate the production of  
22 enriched continental arc volcanics, because both crustal thickness and magma chemistry vary coherently  
23 along strike. However, the scarcity of primitive magmas in the thick-crust northern SVZ has hindered  
24 previous regional studies. To better address the origin of enriched continental compositions, we investigate  
25 the geochemistry (major and trace element abundances, <sup>87</sup>Sr/<sup>86</sup>Sr and <sup>143</sup>Nd/<sup>144</sup>Nd) of new mafic samples  
26 from Don Casimiro and Maipo volcanoes in Diamante-Maipo Caldera Complex of the northern SVZ. While  
27 evolved Diamante-Maipo samples show evidence for crustal assimilation, the trace element and isotopic  
28 enrichment of the most mafic samples cannot result from crustal processing, as no known regional or global  
29 basement lithologies are enriched in all of the necessary incompatible trace elements. Subduction erosion  
30 models similarly fail to account for the enriched isotopic and trace element signature of these samples.

31 Instead, we suggest that the enrichment of northern SVZ magmas is derived from an enriched ambient  
32 mantle component (similar to EM1-type ocean island basalts), superimposed on a northward decline in melt  
33 extent. A substantial, but nearly uniform contribution of melts from subducting sediment and altered  
34 oceanic crust are required at all latitudes. The EM1-like enrichment may arise from recycling of  
35 metasomatized subcontinental lithospheric mantle (M-SCLM), as the isotopic trajectory of primitive rear-arc  
36 monogenetic cones trend towards the compositions of SCLM melts sampled across South America. Isotopic  
37 data from spatially distributed rear-arc centres demonstrate that the arc-parallel variations in the degree of  
38 EM1-type enrichment observed in arc-front samples are also present up to 600 km behind the trench in the  
39 rear-arc. Rear-arc trace element systematics require significant but variable quantities of slab melts to be  
40 transported to the mantle wedge at these large trench distances. Overall, we show that a unified model  
41 incorporating variable mantle enrichment, slab additions, and melt extents can account for along and across-  
42 arc trends within the SVZ. The recognition that mantle enrichment plays a key role in the production of  
43 enriched continental compositions in the SVZ has important implications for our understanding of the  
44 chemical evolution of the Earth. If ambient mantle enrichment is not taken into account, petrogenetic  
45 models of evolved lavas may overestimate the role of crustal assimilation, which, in turn, may lead models of  
46 continental crust growth to overestimate the amount of continental material that has been recycled back  
47 into the mantle.

48

#### 49 **Key Words**

- 50 • Mantle heterogeneity
- 51 • Andean Southern Volcanic Zone

#### 52 **1. Introduction**

53 Magmas which erupt from continental arc-front stratovolcanoes are compositionally distinct from  
54 oceanic arc magmas. With the exception of volcanics from intra-arc rift zones (e.g., Conrey et al., 2004),  
55 continental magmas exhibit elevated incompatible element abundances, steeper incompatible trace  
56 element patterns (Fig. 1a), and often have isotopic signatures offset to higher  $^{87}\text{Sr}/^{86}\text{Sr}$  and lower  $^{143}\text{Nd}/^{144}\text{Nd}$   
57 than oceanic arc magmas (Leeman, 1983; Plank and Langmuir, 1988; Hildreth and Moorbath, 1988; Turner  
58 and Langmuir, 2015a; Farner and Lee, 2017). Many of these compositional characteristics correlate strongly  
59 with crustal thickness (Turner et al., 2015a), and have been used as proxies to estimate changes in crustal  
60 thickness and rates of surface uplift through time (Profeta et al., 2015; Chiaradia, 2015; Chapman et al., 2015;  
61 Scott et al., 2018). Some of the compositional offsets between thick and thin-crust arcs can be accounted  
62 for by higher extents of crystal fractionation and crustal assimilation in continental settings (e.g., Farner and  
63 Lee, 2017). However, other compositional differences persist in high Mg# arc lavas (Fig. 1a) that have not  
64 been extensively overprinted in the crust (e.g., Plank and Langmuir, 1988; Turner and Langmuir, 2015a;

65 Turner et al., 2017; Schmidt and Jagoutz, 2017), suggestive of a mantle origin. Compositional differences  
66 among parental magmas from continental vs. oceanic arc settings have been attributed to variations in slab  
67 temperature (Ruscitto et al., 2012; Turner and Langmuir, 2015b; Schmidt and Jagoutz, 2017), different  
68 extents of mantle melting (e.g., Plank and Langmuir, 1988; Tormey et al., 1991; Turner et al., 2016),  
69 subduction erosion (e.g., Stern, 1989), and heterogeneity of the “ambient” mantle (prior to the addition of  
70 subducted materials; e.g., Hickey et al., 1986; Ewart and Hawkesworth, 1987; Rogers and Hawkesworth,  
71 1989; Hochstaedter et al., 2001; Pearce et al., 2007; Turner et al., 2017). Furthering our understanding of the  
72 thermal structures of subducting plates, the growth of the continental crust, and the petrogenesis of evolved  
73 arc magmas requires constraints on the relative influence of each factor in producing the enriched  
74 compositions of continental arc magmas.

75

76 The Andean Southern Volcanic Zone (SVZ) is an excellent natural laboratory in which to investigate how the  
77 thick crusts and lithospheres of active compressional plate margins affect the compositions of magmas  
78 added to the continents. The SVZ (33-46° S) is one of four volcanically active segments within the Andean  
79 Cordillera, where the Nazca Plate subducts under South America. It is commonly subdivided into three sub-  
80 segments: the northern, transitional, and southern SVZ (NSVZ, TSVZ and SSVZ respectively; Dungan et al.,  
81 2001; Fig. 2a). From the easily-accessed segments of the SSVZ to the NSVZ (42° S to 33° S), crustal thickness  
82 increases from ~30 to 50 km (Fig. 2c), sub-arc slab depth increases from ~70 to 120 km (Tassara and  
83 Echaurren, 2012), and the lower plate age increases from 16 to 37 Ma (Völker et al., 2011), while trench  
84 sediment thickness decreases (Völker et al., 2013). In the NSVZ, seismic and heat flow data indicate that  
85 lithosphere thickness increases rapidly behind the arc, while behind the SSVZ, the lithosphere is thinner  
86 (Tassara and Echaurren, 2012; Valdenegro et al., 2019). As a result, the mantle wedge in the NSVZ is colder  
87 and deeper (Turner et al., 2016). There are also well-established geochemical gradients from the SSVZ to the  
88 NSVZ (Tormey et al., 1991; Hildreth and Moorbath, 1988; Hickey et al., 2016, Fig. 2d-e). The trace element  
89 signatures of mafic SSVZ volcanics overlap those of oceanic arcs (Fig. 1a). In contrast, the more mafic TSVZ  
90 lavas have substantially elevated trace element abundances. The limited amount of trace element data from  
91 mafic NSVZ lavas (Fig. 1a; Hickey et al., 1986) suggest that this segment is even more enriched than the TSVZ,  
92 with compositions similar to the bulk continental crust and mafic to intermediate volcanics from other thick-  
93 crusted arc segments such as the Northern Cascades (Fig. 1a). Thus, the chemical variability between the  
94 SSVZ and NSVZ is analogous to the chemical offsets between island arcs, continental arcs, and the bulk  
95 continental crust.

96

97 While many studies have investigated the origin of along-strike compositional variability of the SVZ arc-front  
98 volcanoes (Lopez-Escobar et al., 1977; Hickey et al., 1986; Hildreth and Moorbath, 1988; Tormey et al., 1991;

99 Jacques et al., 2014; Hickey et al., 2016), the role of enrichment and heterogeneity within the ambient  
100 asthenospheric mantle (prior to the addition of subducted material) has received relatively little attention  
101 (Hickey et al., 1986; Jacques et al., 2014). Though often overlooked, the extent to which the compositions of  
102 continental arc volcanics are inherited from the ambient mantle has important implications for models of  
103 crustal growth. If the trace element enrichment of continental arcs is inherited partially from the mantle,  
104 and not exclusively from extensive crystal fractionation and crustal remelting, a substantially smaller amount  
105 of material must be returned to the mantle to account for the continental mass balance (e.g. Sisson and  
106 Keleman, 2018). The amount of crustal recycling, in turn, has direct relevance for our understanding of the  
107 generation of large-scale mantle heterogeneity.

108  
109 Heterogeneity of the upper mantle is commonly inferred from trace element and isotopic compositions of  
110 mid-oceanic ridge basalts (MORBs) and ocean island basalts (OIBs; Zindler and Hart, 1986; Willbold and  
111 Stracke, 2010). OIB isotopic variability can generally be accounted for by a handful of isotopically distinct  
112 end-member components, such as the “enriched mantle” end-members (EM1, EM2; e.g., Zindler and Hart,  
113 1986). Isotopic end-members are often modelled as the products of recycled lithospheric mantle, sediment,  
114 and ancient oceanic/continental crust (Willbold and Stracke, 2010). There is also evidence for heterogeneity  
115 in the ambient mantle that feeds volcanic arcs, despite additional tectonic complexity. For example, a global  
116 compilation of samples from rear-arc provinces, filtered to minimize slab contributions (Turner and  
117 Langmuir, 2015b), form a linear array between Depleted MORB Mantle (DMM) and EM-1 type OIB  
118 compositions (Fig. 3a). As corner flow transports the mantle feeding rear-arc volcanism over hundreds of  
119 kilometres towards the arc-front (MacDougall et al., 2017), isotopic enrichment within rear-arc provinces  
120 indicates that ambient sub-arc mantle compositions also vary substantially (e.g., Woodhead et al., 2012).  
121 The Andean Southern Volcanic Zone (SVZ) rear arc exemplifies this enriched ambient mantle signature (Kay  
122 et al., 2013; Jacques et al. 2013, 2014; Søger et al., 2015a; Turner et al. 2017), with more enriched  $^{87}\text{Sr}/^{86}\text{Sr}$   
123 and  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios than any other rear-arc province globally (Fig. 3a).

124  
125 This study investigates the petrogenesis of mafic lava samples from the volcanoes Don Casimiro and Maipo  
126 (within the Diamante-Maipo Caldera complex; 34°S, Fig. 2b) in order to determine the geochemical character  
127 of high Mg# NSVZ magmas. These lavas from Don Casimiro and the basal portion of Maipo, which erupted  
128 immediately after the catastrophic formation (~150 ka) of the Diamante Caldera (Orozco et al., 2015), are  
129 among the most primitive ever sampled within the NSVZ (Fig. 4a; Hickey et al., 1986). Their primitive nature  
130 may indicate that the storage and eruption of these melts preceded the re-organization and maturation of  
131 crustal magma reservoirs beneath the Diamante caldera (analogous to post-collapse lavas at Campi Flegrei;  
132 Forni et al., 2018), or that ascending magmas utilized tension cracks formed during the collapse event



133 (Jacques et al., 2014). Regardless of the exact mechanism permitting the eruption of mafic lavas within the  
134 thick-crust NSVZ, these new samples provide an opportunity to characterise and explore the origins of the  
135 prominent chemical differences between the SSVZ and the NSVZ. As the chemical variability between the  
136 SSVZ and NSVZ is analogous to the chemical offsets between island arcs and continental arcs, exploring these  
137 regional trends provides constraints on origins of magmatic enrichment within thickened continental  
138 subduction zones globally.

139

140 This study also incorporates samples erupted within the last 1 Myr from seven Argentinean monogenetic  
141 cones from the Payenia Volcanic Province (34.33° to 35.38°S; Folguera et al., 2009). The origin of this rear-arc  
142 volcanism, located up to 350 km behind the NSVZ arc front, remains a topic of active research. Volcanism  
143 initiated in the earliest Miocene, associated with a period of shallow slab subduction, and continued into the  
144 Holocene (Kay and Copeland, 2006; Pallares et al., 2016). In the southernmost part of Payenia, it is generally  
145 accepted that the slab steepened towards the present-day dip of 33° (Tassara and Echarren, 2012) by 5 Ma  
146 (Kay et al., 2005). Folguera et al. (2009) suggest that volcanism propagated northwards following extensional  
147 relaxation of the San Rafael block, while Gudnason et al. (2012) attribute the northward propagation of  
148 volcanism within Payenia to a northward delay in slab rollback. Despite uncertainty in the chronology and  
149 mode of magmagenesis of the SVZ rear-arc, the wide spatial distribution of volcanic products, along with  
150 their relatively primitive chemistry, provide vital constraints on the composition of the ambient sub-arc  
151 mantle beneath the SVZ (Jacques et al., 2013; Turner et al., 2017).

## 152 **2. Samples and methods**

153 Six samples were collected from the basal sequence of Maipo Volcano, the active stratovolcano at the centre  
154 of the Diamante-Maipo caldera, and nine from Don Casimiro Volcano, a small eroded stratovolcano located  
155 ~10 km SW of Maipo with activity restricted to preglacial times (Charrier, 1979; Fig. 2b).  $^{40}\text{Ar}/^{39}\text{Ar}$  dates  
156 indicate that early activity at these centres was contemporaneous (Orozco et al., 2015). All lavas contain  
157 olivine and clinopyroxene phenocrysts within a glassy groundmass. The more evolved samples also contain  
158 plagioclase, oxides, and orthopyroxene (details in SIA2). Seven samples were collected from monogenetic  
159 scoria cones in the northern rear-arc (34.3 to 35.4 °S; Fig. 2a). Olivine, clinopyroxene, and oxide phenocrysts  
160 dominate in these scorias, with microcrysts of plagioclase and alkali feldspar, and minor apatite.

161

162 All 22 samples were powdered for whole-rock analysis with an agate ball mill at the University of Oxford.  
163 Major element analysis was conducted using a PANalytical Axios Advanced X-ray fluorescence spectrometer  
164 at the University of Leicester (Knott et al., 2016). Powders were digested in clean labs at the University of  
165 Oxford for ICP-MS analysis. Ba and Sr were measured at a 1,200,000x dilution using a Thermo Element 2 ICP-  
166 MS, while all other trace elements were measured at a 6000x dilution on a Perkin Elmer NexION 350D

167 quadrupole ICPMS (both at the University of Oxford). Four certified reference materials (CRMs; BCR-2,  
168 BHVO-2, AGV-2, and W-2a) were digested alongside these samples. Data was reduced using calibration  
169 curves generated from these 4 CRMs. Repeat solution analysis of one sample (n=19) and repeated digestions  
170 indicate an analytical uncertainty of <3% (<5% for Cs), and calculated values for the CRMs yield  
171 concentrations within  $\pm 5\%$  of the preferred values (Appendix C).

172

173  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  isotope ratios were analysed for eleven samples at the University of Leeds.  
174 Following Sr and Nd purification, samples were loaded onto outgassed W (Sr) and Re (Nd) filaments, and  
175 analysed on a Thermo Scientific Triton-series multicollector mass spectrometer. NIST SRM-987, La Jolla and  
176 BHVO-2 were analysed throughout the session to monitor accuracy and precision (Appendix C).

177

178 Olivine grains were analysed for major elements on a Cameca SXFive FEG-electron microprobe at the  
179 University of Oxford. St Johns Olivine and Fayalite were analysed as secondary standards to assess accuracy  
180 and precision. LA-ICP-MS olivine analysis was conducted at the University of Cambridge on a 193ESI Laser  
181 Ablation system coupled to a Nexion ICP-MS. Multiple secondary standards were run to monitor instrument  
182 drift, and accuracy and precision. Additional analytical details are provided in SIA3.

### 183 **3. Results**

184 The new Don Casimiro-Maipo samples are primitive calc-alkaline basaltic andesites to andesites. The whole-  
185 rock Mg#s (assuming  $\text{Fe}^{3+}/\text{Fe}_T=0.3$ ; Holm et al., 2016) of Don Casimiro-Maipo samples mostly range from 60-  
186 70, indicating that these samples have undergone only minor crystal fractionation from primary magmas  
187 (Fig. 4a). There is little variation in  $\text{SiO}_2$  (55.3–57.9 wt.%) within the sample set. Previous studies of the  
188 Diamante-Maipo Caldera (Hickey et al., 1986; Futa and Stern, 1988, Sruoga et al., 2005; Holm et al., 2014)  
189 have primarily sampled volcanics with significant europium anomalies ( $\text{Eu}^* = [\text{Sm}_N \times \text{Gd}_N]^{0.5}$ ; Fig. 4b),  
190 indicating fractionation of plagioclase or mixing between primitive and evolved magmas (e.g., Turner and  
191 Langmuir, 2015a). In contrast, our higher Mg# samples have  $\text{Eu}/\text{Eu}^* > 0.9$  (Fig. 4b). Olivine phenocryst  
192 compositions range from  $\text{Fo}_{83.6-75.7}$  at the arc-front, which lie below the olivine-whole rock equilibrium line,  
193 indicating that some olivine accumulation has occurred. The highest Mg# Maipo lavas (Mg#~70) have  
194 compositions that reflect the addition of up to 7 wt%  $\text{Fo}_{81}$  olivine to more evolved melts (Mg# ~60), and  
195 should not be regarded as “primitive” bulk compositions (Fig. 4a-b; Fig. SB2; Hickey et al., 2016). We regard  
196 the five Don Casimiro samples with the highest MgO,  $\text{Eu}/\text{Eu}^*$  and Mg#s (circled on Fig. 4; 4.85-5.63wt%  
197 MgO) to be representative of primitive magma compositions within the Diamante-Maipo caldera.

198

199 Like most arc magmas, mafic Don Casimiro-Maipo lavas have high abundances of Cs, Rb, Ba, U, Th, Pb, and  
200 Sr, and depletions in Nb and Ta relative to other elements of similar incompatibility during mantle melting

201 (Fig. 1b). Highly incompatible trace element abundances of these lavas are elevated relative to typical mafic  
202 to intermediate samples from oceanic arcs and the SSVZ. Compared to other primitive SVZ samples, Don  
203 Casimiro-Maipo lavas also exhibit lower Cs/Rb, smaller Zr-Hf depletions ( $\text{Hf}/\text{Sm} \sim 0.9$ ; Fig. SB3), and  
204 significantly more enriched isotope ratios (higher  $^{87}\text{Sr}/^{86}\text{Sr}$  and lower  $^{143}\text{Nd}/^{144}\text{Nd}$ ; Fig. 3b). The isotopic  
205 compositions of the most primitive Don Casimiro samples from this study have lower  $^{87}\text{Sr}/^{86}\text{Sr}$  and higher  
206  $^{143}\text{Nd}/^{144}\text{Nd}$  than our Maipo lavas, and literature samples from throughout the Diamante-Maipo caldera (Fig.  
207 3b).

208

209 Rear-arc scoria samples are basalts-trachybasalts (Fig. 3a; Mg#s of 58–71; if  $\text{Fe}^{3+}/\text{Fe}_T = 0.15$ ; Sjøager et al.,  
210 2015b). Olivine core compositions have higher forsterite contents than Don Casimiro-Maipo samples ( $\text{Fo}_{90.1-}$   
211  $80.6$  vs.  $\text{Fo}_{83.6-75.7}$ ), and lie close to the equilibrium field (Appendix C; Fig. SB2a). Trace element patterns in the  
212 rear-arc show prominent subduction signatures (e.g., Nb-Ta depletion, Pb enrichment, and Th/Nb up to  
213 0.72), and concentrations even higher than Don Casimiro-Maipo lavas (Fig. 1b). As noted by Holm et al.  
214 (2016), rear-arc samples have slight depletions in the high field strength elements Zr and Hf, though such  
215 depletions are comparable to those observed within the SVZ arc front ( $\text{Hf}/\text{Sm} \sim 0.6$ ; Fig. SB3).  $^{87}\text{Sr}/^{86}\text{Sr}$  and  
216  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios of the new rear-arc samples plot between Don Casimiro-Maipo and literature SVZ data,  
217 and are mostly offset from the filtered rear-arc literature data to higher  $^{87}\text{Sr}/^{86}\text{Sr}$  at a given  $^{143}\text{Nd}/^{144}\text{Nd}$  (Fig.  
218 3b).

219

220 The new whole rock trace element and isotopic data from Don Casimiro-Maipo and rear-arc samples  
221 collected in this study is presented in Table 1. Complete geochemical analyses (including olivine  
222 compositions) are presented as a spreadsheet in Appendix C.

## 223 4. Discussion

224 The northward increase in crustal thickness in the SVZ provides the ideal setting to understand the  
225 relationship between primitive magma compositions and crustal thickness. However, the scarcity of  
226 geochemical data for mafic lavas in the NSVZ has hindered assessments of the composition of parental  
227 magma compositions, and thus the origin of regional geochemical variations. The more primitive nature of  
228 Don Casimiro-Maipo samples investigated in this study provide a new opportunity to re-evaluate the relative  
229 contributions from crustal processing, mantle melting, slab fluxes, subduction erosion, and ambient mantle  
230 heterogeneity.

231

### 232 4.1. Crustal processing at Don Casimiro-Maipo

233 Hildreth and Moorbath (1988) suggested that mantle-derived magmas in the SVZ are processed in a melting,  
234 assimilation, storage, and homogenization (MASH) zone in the lower crust. In their model, mantle melts

235 throughout the SVZ have similar compositions, and the enriched “baseline” composition of NSVZ magmas  
236 results from increased MASH processing in the north, where the crust is thickest. A compilation of samples  
237 from the entire Don Casimiro-Maipo caldera (this study and literature; Appendix D) indicates a role for  
238 crustal assimilation or mixing during fractional crystallization (AFC) in the petrogenesis of more evolved  
239 lavas. For example, the  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios of Diamante Caldera lavas correlate with MgO and  
240 Eu/Eu\* (Fig. SB1). These geochemical trajectories can be recreated with a simple model whereby the  
241 compositions of sampled crustal lithologies (Lucassen et al. 2001; Lucassen et al., 2004) are mixed in variable  
242 proportions with primitive Don Casimiro-Maipo lavas (Fig. 4c-d). More complex assimilation models were  
243 not utilized, due to the uncertainty regarding the thermal properties of such a wide variety of basement  
244 lithologies.

245

246 The MASH model proposed that the isotopic and trace element enrichment in even the most primitive NSVZ  
247 volcanics were caused by crustal assimilation and fractionation of parental magmas that are similar to SSVZ  
248 primary magmas. To test the power of the MASH model to explain the compositions of our new primitive  
249 NSVZ samples, we represent the composition of primitive SSVZ volcanics by an average composition of mafic  
250 samples from Villarrica. Villarrica erupts high-Mg# lavas and has a similar trace element and isotopic  
251 composition to the other volcanoes of the SSVZ (Fig. 1a, Fig. 3b; Turner et al., 2016). We mixed the crustal  
252 lithologies which recreate the geochemical trajectories *within* the Diamante-Maipo Caldera into the  
253 composition of mafic Villarrica lavas as an initial test of whether assimilation into a primitive SSVZ magma  
254 can recreate the primitive Don Casimiro-Maipo lava compositions. These mixing trends produce rapidly  
255 decreasing K/Rb ratios prior to producing adequate enrichment in  $^{87}\text{Sr}/^{86}\text{Sr}$ , and none reach high enough  
256 Rb/Y to replicate the primitive Don Casimiro-Maipo lava compositions (Fig. 4c-d). Thus, assimilation of these  
257 lithologies cannot bridge the compositional gap between the SSVZ and NSVZ.

258

259 To test whether *any* plausible Andean crustal assimilant can bridge the compositional gap between the NSVZ  
260 (Don Casimiro-Maipo) and SSVZ (Villarrica), the compositions of 348 basement outcrops and basement  
261 xenoliths from the SVZ and the surrounding area (22-46°S) were compiled (Lucassen et al. 2001; Lucassen et  
262 al., 2004 and others; Appendix E). As above, assimilation was modelled by mixing the average composition of  
263 mafic Villarrica samples with each basement lithology (for mixing proportions between 0–100%). Only 14%  
264 of potential assimilants produce mixed compositions within  $^{87}\text{Sr}/^{86}\text{Sr} \pm 0.0001$  and  $^{143}\text{Nd}/^{144}\text{Nd} \pm 0.00005$  of  
265 the primitive Don Casimiro-Maipo average (Fig. 5a). Of these isotopically plausible mixed compositions, only  
266 35% have Eu anomalies within the range of the primitive Don Casimiro-Maipo samples ( $\text{Eu}/\text{Eu}^* > 0.9$ ). None  
267 of this subset of mixed compositions have trace element signatures resembling Don Casimiro-Maipo (Fig.  
268 5b).

269

270 Incorporating crystal fractionation alongside these simple assimilation models also cannot account for these  
271 compositional offsets, because the resulting compositions deviate from primitive Don Casimiro-Maipo  
272 samples not only in overall elemental abundances, but also in a variety of incompatible element ratios that  
273 are not fractionated by early-crystallizing phases (e.g., Dy/Yb; Davidson et al., 2007). Additionally, while  
274 models of recharging, evacuating and fractionation (REFC) magma chambers (Lee et al., 2014) could elevate  
275 incompatible element concentrations, this would drive prominent Fe enrichment, which is at odds with the  
276 high Mg#s observed in Don-Casimiro-Maipo and rear-arc samples. Finally, REFC would most probably occur  
277 in the deep crust (based on thermal constraints and the absence of Eu/Eu\* anomalies in our samples). This  
278 would drive magma compositions towards peraluminous compositions, inconsistent with our major element  
279 data.

280

281 It remains possible that the required MASH assimilant has evaded sampling within the Southern Andes. To  
282 address this possibility, we assess the chemical characteristics that a hypothetical assimilant must possess to  
283 bridge the compositional gap between the NSVZ and SSVZ. Assuming a maximum of 20% assimilation, based  
284 on the relatively high Mg#s of primitive Don Casimiro-Maipo lavas (assimilation drives cooling and crystal  
285 fractionation; DePaolo, 1981), any suitable assimilant must have greater than ~450ppm Zr, ~1420ppm Sr,  
286 and ~0.63wt% P<sub>2</sub>O<sub>5</sub>. It is improbable that crustal lithologies will possess these characteristics, because these  
287 elements typically become depleted, rather than enriched, during late stage crystal fractionation or crustal  
288 melting (Turner and Langmuir, 2015a). To illustrate this point, a compilation of all continental granites,  
289 diorites, syenites, and monzonites in the GEOROC database (N~3000 with trace element data) was searched  
290 for suitable assimilants (considering only trace element abundances). Only 15 samples (0.5% of the entire  
291 compilation) possessed sufficient concentrations of Zr, Sr, and P<sub>2</sub>O<sub>5</sub>. Mixing of these lithologies with mafic  
292 Villarrica lavas produces compositions with erratic trace element patterns, testament to the complex  
293 petrological histories of these unusually enriched basement lithologies. None recreate the observed trace  
294 element pattern of Don Casimiro-Maipo, particularly the negative Nb-Ta anomaly (Fig. 5c). Thus, although  
295 crustal assimilation is common among more evolved samples from the NSVZ, the compositional offset  
296 between mafic NSVZ and SSVZ lavas is unlikely to have been generated by assimilation in the lower crust.  
297 The failure of the MASH model is most apparent when the full trace element signatures of regional crustal  
298 assimilants are considered, along with fundamental considerations regarding the production of enriched  
299 basement lithologies worldwide.

300

301 The <sup>87</sup>Sr/<sup>86</sup>Sr and <sup>143</sup>Nd/<sup>144</sup>Nd ratios of rear-arc samples presents further challenges to the MASH model's  
302 ability to explain the isotopic offsets between NSVZ and SSVZ samples. Interpreting the trace element and

303 isotopic compositions of rear-arc samples is complicated by the sporadic infiltration of slab melts to the rear-  
304 arc mantle (Jacques et al., 2013). However, by filtering a compilation of all published data from SVZ rear-arc  
305 monogenetic cones to remove samples with large slab additions ( $\text{Th/Nb} > 0.4$ ), it becomes apparent that rear-  
306 arc  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  data plot along a “mantle array” between DMM and EM1-type ocean island  
307 basalts (Fig. 3a; Sørensen et al., 2013; 2015a-b; Sørensen and Holm, 2013; Kay et al., 2013; Jacques et al., 2013;  
308 2014; Holm et al., 2014; Turner et al. 2017). This trajectory is not consistent with assimilation of available  
309 basement (e.g., Fig. 5b), because samples with high Mg#s are available at all points along the array.  
310 Additionally, unlike the arc-front samples, many of the primitive rear-arc volcanics bear highly forsteritic  
311 olivines, nearly in equilibrium with the mantle (Fig. SB2a). It is difficult to reconcile the presence of these  
312 primitive olivines with models invoking large amounts of crustal assimilation (e.g., 70% assimilation; Sørensen  
313 et al., 2013), as heat loss during assimilation would drive olivine compositions to lower forsterite contents.  
314 Additionally, rear-arc isotopic variability shows similar spatial trends to that observed at the arc-front,  
315 despite substantial longitudinal offsets (Fig. 6a-b). Crustal contamination cannot explain this similarity, as  
316 accretion and eastward migration of the arc-front since the Palaeozoic has produced a longitudinally  
317 fragmented basement, with different basement domains striking parallel to the coast (Kay et al., 2005).  
318 Assimilation of lithologies within these different basement domains would be expected to produce  
319 uncorrelated rear-arc and arc-front isotope systematics. In contrast, mantle flow paths *are* predicted to  
320 travel directly from the rear-arc to arc front, providing a coherent mechanism to account for the  
321 synchronized latitudinal changes in arc-front and rear-arc magmas. Thus, the coincidence of along-strike  
322 rear-arc and arc-front geochemical trends is one of the strongest pieces of evidence that the enriched  
323 compositions of NSVZ magmas are derived from the mantle rather than the crust. Additional consideration  
324 will be given to this observation in section 4.5.

325

#### 326 **4.2 Assessing the slab and mantle melting contributions to SVZ compositional variation**

327 As with other arcs, variable slab components and mantle melting processes may affect the chemistry of SVZ  
328 volcanics. The NSVZ mantle wedge is likely to have a lower maximum extent of melting than the SSVZ,  
329 because the thickened over-riding plate in the NSVZ forces high wedge temperatures to greater pressures  
330 (Turner et al., 2016). This should lead to enrichment of incompatible elements in the NSVZ (e.g. Tormey et  
331 al., 1991; Jacques et al., 2014; Turner et al., 2016). In addition to variability in mantle melt extent (F), the  
332 trace element compositions of primary arc magmas are influenced by slab materials. For example, Jacques et  
333 al. (2014) attribute compositional variability between 38-43° S to varying slab fluxes. To assess whether the  
334 compositional differences between the NSVZ and SSVZ can be accounted for by varying F and/or slab  
335 components, it is useful to consider the maximum solution space generated by varying these parameters in  
336 tandem.

337

338 For this exercise, trace element concentrations and isotopic ratios were calculated using an adapted version  
339 of the mantle melting model of Turner and Langmuir (2015a; see SIA4). The compositions of subducted  
340 sediment and AOC were taken from the average isotopic and trace element compositions of sediment core  
341 ODP1232 (Turner et al., 2017) and NMORB (Gale et al., 2013). As numerical models indicate that SVZ slab top  
342 temperatures are well above the wet pelite and AOC solidi (Syracuse et al., 2010; Hermann and Spandler,  
343 2007; Carter et al. 2015; Jacques et al., 2013; 2014), and only hydrous melt (as opposed to aqueous fluid)  
344 eclogite partition coefficients produce successful models of trace element abundances (Turner et al., 2017),  
345 it was assumed that slab melts (rather than fluids) were added to the mantle wedge. Slab melt compositions  
346 were calculated using the aggregate fractional melting equation and partition coefficients consistent with  
347 experiments (SIA4.1; Kessel et al., 2005; Hermann and Rubatto, 2009; Skora and Blundy, 2010). Slab melts  
348 were then mixed with a depleted upper mantle composition assigned the trace element abundances of  
349 Workman and Hart (2005), and the isotopic composition of the rear-arc sample 126171 (Baseline Andean  
350 Mantle - BAM; Figs 3a, 6a-b; Søager et al., 2013). This sample defines the low Sr and high Nd end-member of  
351 the rear-arc array (Fig. 3a) and has a trace element composition that appears free of slab material (e.g.,  
352  $\text{Th/Nb} < 0.06$ ,  $\text{Ce/Pb} > 25$ ). Finally, the compositions of mantle melts were estimated using the batch melting  
353 equation, with experimentally-derived hydrous peridotite partition coefficients and melt reaction  
354 coefficients (model details in SIA4.1).

355

356 Firstly, we consider the solution space for  $^{87}\text{Sr}/^{86}\text{Sr}$  vs.  $^{143}\text{Nd}/^{144}\text{Nd}$ , which is simplified by the fact that these  
357 trajectories are not noticeably fractionated by variations in slab temperature, and are independent of melt  
358 extent. The pink field on Fig. 7a shows the solution space produced by variations in slab temperature and the  
359 proportions of slab melts. The isotopic composition of Villarrica (and most SSVZ) volcanics is encompassed  
360 within this field. However, TSVZ and Don Casimiro samples plot significantly below this field. As variable slab  
361 contributions predominantly drive wedge compositions to higher  $^{87}\text{Sr}/^{86}\text{Sr}$  at near constant or increasing  
362  $^{143}\text{Nd}/^{144}\text{Nd}$ , another process is clearly needed to recreate the low  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios in TSVZ and NSVZ  
363 volcanics.

364

365 A similar solution space for trace element ratios is shown in Fig. 7b. The pink area encompasses the space  
366 made available by varying the proportions of a slab melts (with slab melting at  $\sim 800$  °C) and F within  
367 reasonable bounds.  $F=0.04$  was chosen as the lowest plausible mantle melting extent because F values  $< 0.04$   
368 tend to produce alkali-basalt or silica-undersaturated major element compositions (e.g., Baasner et al.,  
369 2016). Villarrica samples plot within this field, and a model fit can be found that reproduces the entire suite  
370 of Villarrica incompatible trace elements and radiogenic isotope ratios (Figs. 7-8). While this model fit does

371 not represent a unique solution for Villarrica, it does demonstrate that a feasible quantitative solution exists  
372 that is consistent with available experimental constraints. As in Fig. 7a, the compositions of most TSVZ and  
373 Don Casimiro-Maipo lavas fall outside of this solution space. The addition of sediment and AOC melt to the  
374 depleted mantle source can generate high La/Sm, but only at excessively high Sr/Nd (Fig. 7a). While reducing  
375 F can increase La/Sm, even at F=0.04 there is no plausible trace element match to the Don Casimiro-Maipo  
376 lavas within this parameter space.

377

378 Slab temperatures might also affect the trace element compositions of SVZ volcanics. There are systematic  
379 changes in the slab age and depth beneath the SVZ arc front, both of which may lead to variability in the  
380 temperature of the slab surface beneath the arc front (Syracuse et al., 2010). Certain aspects of the trace  
381 element differences between Don Casimiro-Maipo and Villarrica are indeed suggestive of slab temperature  
382 variations. Compared to Villarrica, Don Casimiro-Maipo has low Cs/Rb and high Hf/Sm ratios (Fig. 1b). These  
383 element ratios are not substantially fractionated from each other during melting, and may indicate higher  
384 slab temperatures in the NSVZ. For example, Cs is highly mobile in low temperature aqueous fluids, and may  
385 have been disproportionally lost during early slab dehydration in the NSVZ, where the sub-arc slab is deeper  
386 (Savov et al., 2007; Spandler et al., 2007). Additionally, a hotter slab temperature can destabilize zircon  
387 during slab melting, which may lead to the less pronounced Zr-Hf depletions at Don Casimiro-Maipo (e.g.,  
388 Hirai et al., 2018). The cyan field on Fig. 7b shows the expansion of the solution space for slab temperatures  
389 hotter than 900°C (calculated using the maximum measured experimental mobility of light rare earth  
390 elements in slab melts; see SIA4.1). While the solution space at very high slab temperatures may shift toward  
391 higher La/Sm ratios at a given Sr/Nd ratio, the mafic NSVZ volcanic rocks remain well outside of the realm of  
392 possible solutions. Overall, the trace element and isotopic differences between mafic SSVZ and NSVZ  
393 volcanics cannot be recreated by variable slab additions and melt extents, even in combination. This  
394 indicates that the sub-arc mantle within the SVZ must be variably enriched to account for the prominent  
395 latitudinal variations in geochemistry.

396

### 397 **4.3 Is there geochemical evidence for subduction erosion?**

398 SVZ rear-arc volcanics form a linear array between MORB-like compositions and EM1-type OIBs in  $^{87}\text{Sr}/^{86}\text{Sr}$   
399 vs.  $^{143}\text{Nd}/^{144}\text{Nd}$  space (Kay et al., 2013; Søger et al., 2015a; Fig. 3a), and both rear-arc and arc-front samples  
400 are characterized by similar along-strike variability in these isotope ratios (Fig. 6a-b). As the isotopic offset  
401 between the mafic NSVZ and SSVZ samples is inconsistent with crustal assimilation or variable slab additions,  
402 an alternative mechanism is required to introduce an EM1-like contaminant to the mantle source of the SVZ.  
403 The origins of EM1-type OIBs are debated, with suggestions including deep recycling of lower continental  
404 crust (LCC), pelagic sediment and oceanic crust, or metasomatized subcontinental lithospheric mantle



405 (Willbold and Stracke, 2010). The first suggestion is particularly apt, as it has been argued that the NSVZ  
406 endured extensive subduction erosion (the removal of upper plate material via abrasion and plucking by the  
407 lower plate) throughout the Cenozoic (Stern, 1989; Stern, 2011). Although the composition of the Andean  
408 LCC is relatively uncertain, an extensive compilation of regional basement compositions (Appendix E)  
409 constrains the likely compositional range of this reservoir. While these samples were collected from surface  
410 outcrops, Miocene episodes of tectonic thickening and under-thrusting have resulted in similar lithologies  
411 residing at deeper levels in the crust (Hildreth and Moorbath, 1991). Hence, this compilation can be used to  
412 assess the subduction erosion hypothesis, regardless of whether the UCC, LCC or the entire crustal column  
413 (Holm et al., 2014; 2016) on the leading edge of the South American plate undergoes recycling.

414  
415 We assess the ability of subduction erosion to reproduce along-arc geochemical trends by mixing the  
416 composition of basement lithologies (Appendix E) into the estimated composition of the metasomatized  
417 Villarrica mantle source. Melts of the resulting mixes were calculated using the model described above at  
418 conditions indicated by thermal models for Don Casimiro-Maipo ( $F=0.056$ ,  $P=33\text{kbar}$ ; Turner et al., 2016).  
419 Only 11% of the mixed compositions pass within  $^{87}\text{Sr}/^{86}\text{Sr}\pm 0.0001$  and  $^{143}\text{Nd}/^{144}\text{Nd}\pm 0.00005$  of the most  
420 primitive Don Casimiro-Maipo samples (Fig. 9a). Only 15% of these isotopically plausible mixed compositions  
421 have Eu anomalies within the range of the primitive Don Casimiro-Maipo samples ( $\text{Eu}/\text{Eu}^* > 0.9$ ), none of  
422 which recreate the concentrations of Y, Gd, Sm, Sr, Ba and  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios (Fig. 9b-c). This analysis  
423 demonstrates that no mixed composition with adequate trace element data plots within the range of the  
424 Don Casimiro-Maipo for  $^{87}\text{Sr}/^{86}\text{Sr}$ ,  $^{143}\text{Nd}/^{144}\text{Nd}$ , and trace element abundances.

425  
426 While some previous studies of the NSVZ (e.g., Holm et al., 2014; 2016) have concluded that eroded  
427 continental material is the most plausible candidate for the enriched mantle end-member in this region, our  
428 analysis does not support this conclusion. This discrepancy arises from the fact that the newly sampled  
429 primitive Don Casimiro-Maipo lavas have  $\text{Eu}/\text{Eu}^*$  close to 1. This implies that the low  $\text{Eu}/\text{Eu}^*$  ratios of  
430 previous studies are a secondary overprint from plagioclase fractionation, magma mixing, or crustal  
431 assimilation, rather than a mantle source characteristic.

432  
433 As with crustal assimilation during magmatic ascent, it is possible that the appropriate eroded crustal  
434 lithology has avoided sampling due to the restricted basement exposure in the SVZ. However, regional  
435 considerations indicate that LCC recycling is unlikely to produce the observed EM1-like enrichment. Willbold  
436 and Stracke (2010) present a model in which EM1-like isotopic compositions are produced by mixing LCC  
437 starting materials into normal MORB mantle. However, the success of this model requires that the recycled  
438 LCC domains have been extracted and isolated from the mantle for  $>4$  Ga, while the Chilenia and Cuyania

439 terranes comprising the lower crust of the NSVZ-TSVZ are only of Grenvillian age (~1Ga; Ramos, 2010). This  
440 relatively young age causes typical Southern Andean basement to have  $^{187}\text{Sr}/^{86}\text{Sr}$  ratios that are too high  
441 (relative to  $^{143}\text{Nd}/^{144}\text{Nd}$ ) to serve as an appropriate end-member (see mixed compositions in Fig. 9a).

442

443 Rear-arc geochemical variability is also inconsistent with the subduction erosion hypothesis, because rear-  
444 arc isotopic enrichment persists at distances exceeding 600 km from the trench. It is unlikely that material  
445 eroded at the trench could spread hundreds of kilometres towards the rear arc, in the opposite direction of  
446 corner flow (MacDougall et al., 2017). While eroded material may instead be carried down with the  
447 subducting slab and released alongside other slab inputs, in this case the enriched signature would co-vary  
448 with other slab components. Instead, at 550 km from the trench, rear-arc monogenetic samples have Ce/Pb  
449 ratios that range from 5 (similar to the arc front) to >20 (similar to MORBs; Fig. 10b), but retain  $^{143}\text{Nd}/^{144}\text{Nd}$   
450 ratios that are consistently enriched relative to DMM (Fig. 3a). Additionally, within a given latitude bracket,  
451 rear-arc  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios have limited variability, and do not correlate significantly with Ce/Pb or other  
452 indices of slab additions. Finally, isotopic enrichment in both the rear arc and arc front increases in a  
453 reasonably symmetrical, coherent fashion on either side of a pronounced minima at 39° S (Fig. 6a-b; Turner  
454 et al., 2017). Although it has been suggested that subduction erosion increases northwards (Stern, 1989),  
455 this phenomenon cannot account for the increase in enrichment south of Villarrica. The spatial systematics  
456 of rear-arc compositions are therefore also inconsistent with subduction erosion.

457

458 Overall, when the full array of trace element abundances and isotopic ratios are considered, subduction  
459 erosion cannot recreate the notable geochemical trends in rear-arc and arc-front data. This may indicate  
460 that the extent of subduction erosion has been over-estimated for the SVZ. In some cases, erosion rates are  
461 calculated based on the assumption that the distance between the trench and the arc front remained  
462 constant throughout the Miocene (Stern, 1989). However, more recent geophysical models find that arc-  
463 front migration relative to trench location might be a common process (e.g., Karlstrom et al., 2014), and the  
464 locus of subduction-related volcanism has also likely been influenced by variability of the slab angle over this  
465 same interval (Kay et al., 2006). Alternatively, it is possible that subduction erosion did indeed influence the  
466 mantle wedge composition at some point in the Miocene, but corner flow towards the arc has since flushed  
467 the present-day mantle wedge of eroded material. Regardless of the explanation, the range of plausible  
468 geochemical signatures that would be produced by subduction erosion are inconsistent with the along-strike  
469 enrichment trend of the SVZ, and thus some other mechanism is required to explain SVZ geochemical  
470 variability.

471

#### 472 **4.4 EM1-type enrichment from a sub-continental lithospheric mantle source**

473 An alternative source for the enriched EM1-like mantle signature of the NSVZ is the recycling of  
474 Metasomatised Subcontinental Lithospheric Mantle (M-SCLM). The infiltration of low-degree mantle melts  
475 ( $F \sim 0.005$ ) into the base of the lithospheric mantle produces enriched incompatible element concentrations,  
476 which rapidly evolve to EM1-like isotopic compositions (McKenzie and O’Nions, 1995; Turner et al., 2017;  
477 Fig. 11). Unlike crustally-derived materials, M-SCLM provides a more homogenous end-member composition  
478 that is capable of recreating the remarkably coherent geochemical trends observed in the SVZ (Rogers and  
479 Hawkesworth, 1989; Turner et al, 2017). M-SCLM material may be stored for long periods in the lithospheric  
480 mantle until it is returned to the asthenosphere by delamination or erosion driven by corner flow.

481

482 Field evidence supports the hypothesis that the SCLM surrounding the SVZ has EM1-like isotopic affinities.  
483 Three geographically separated suites of isotopically enriched igneous rocks across South America have been  
484 interpreted as direct melts of the enriched SCLM (see Fig. 2a):

- 485 1) Alkaline igneous rocks erupted near the edges of the Rio Apa-Luis Alves and Sao Francisco Cratons  
486 (Gibson et al., 1995; 2005; Carlson et al., 1996).
- 487 2) Mesozoic metabasites from the Southern Rift (Lucassen et al., 2002).
- 488 3) Carboniferous granitoids from the Santo Domingo Complex of the Coastal Batholith (Parada et al.,  
489 1999).

490 The  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios of samples from these regions plot as a linear extension of the SVZ rear-  
491 arc array (Fig. 11), and cannot be accounted for by addition of variable slab components or assimilation of  
492 available basement. The remarkable alignment of these lithologies in isotopic space, despite being  
493 geographically separated by hundreds to thousands of kilometres, suggests that M-SCLM east and north of  
494 the SVZ has relatively homogeneous EM1 affinities. The isotopic trajectory of these SCLM melts is consistent  
495 with the expected isotopic evolution of SCLM enriched by low-degree mantle melts (Turner et al., 2017; Fig.  
496 11). In contrast, the field produced by addition of sampled basement lithologies to the mantle has higher  
497  $^{87}\text{Sr}/^{86}\text{Sr}$  at a given  $^{143}\text{Nd}/^{144}\text{Nd}$  than the array of SCLM-melts and rear-arc volcanics, and a significantly  
498 broader trajectory (grey field, Fig. 11).

499

500 Rather than attempt to estimate the full trace element composition of very low-degree metasomatising  
501 mantle melts, a proxy for the composition of mantle domains which has been enriched by M-SCLM melts  
502 was estimated by inverting for the mantle source composition of Gough Island EM1-type OIB volcanics  
503 (Turner et al., 2017), where enrichment has been specifically linked to deep recycling of SCLM (Gibson et al.,  
504 2005; Boyet et al., 2019). Trace element and isotopic ratios of the inverted source composition were  
505 subsequently adjusted within the limits of EM1-like ocean island basalts (Stracke et al. 2003; Willbold and

506 Stracke, 2010) to produce a generic EM1 source (SIA4.2). Including EM1-like mantle heterogeneity as an  
507 additional parameter sufficiently expands the solution spaces on Fig. 7a-b (grey fields) to overlap the  
508 composition of primitive Don Casimiro-Maipo lavas. Furthermore, a model fit incorporating this EM1-type  
509 mantle component recreates the isotopic composition and entire suite of incompatible trace element  
510 concentrations at Don Casimiro-Maipo (Figs. 7–8). As with the model fit for Villarrica, this does not represent  
511 a unique solution, but does suggest that the observed geochemical enrichment at Don Casimiro-Maipo is  
512 quantitatively consistent with all available constraints, unlike models relying upon MASH, LCC recycling, or  
513 variable slab components.

514

515 The addition of M-SCLM to the mantle by erosion or delamination at the base of continental cratons behind  
516 the volcanic arc also accounts for the spatial variability in rear-arc and arc-front geochemistry. Turner et al.  
517 (2017) proposed two explanations for the gradual decline in isotopic enrichment towards a distinct minimum  
518 at 39° S (Fig. 6a-b). Firstly, they note the presence of a Permian terrane suture at 39° S (Rapalini et al., 2010).  
519 It is possible that M-SCLM was lost during rifting or subsequent collision between cratonic blocks (Fig. 12a).  
520 Secondly, a vertical and horizontal slab tear has been inferred based on seismic imaging at ~39° S (Pesicek et  
521 al., 2012). If M-SCLM was entrained equally throughout the SVZ, upwelling of depleted Pacific mantle (which  
522 has not interacted with M-SCLM) through these features could dilute the EM1 signature. The upwelling,  
523 depleted mantle would then spread northwards and southwards, perhaps aided by toroidal flow (Zandt and  
524 Humphreys, 2008), resulting in a gradual northward and southward increase in enrichment (Fig. 12b). This  
525 upwelling material could also produce the magnetotellurically imaged "plume" in the rear-arc (Burd et al.,  
526 2014). Geochemical signatures associated with upwelling through a slab window in the Kula Volcanic Field  
527 (W. Anatolia, Turkey) are traced over >250 km (Klaver et al., 2016), similar to the distances observed here.

528

#### 529 **4.5 A unified model for the SVZ arc-front and rear-arc**

530

531 The preceding sections have shown that the first-order geochemical variability in the SVZ arc-front and rear-  
532 arc is consistent with the mixing of EM1-type and isotopically depleted mantle end-members. This is similar  
533 to the proposal of Jacques et al. (2013, 2014), though by including the influence of ambient mantle  
534 enrichment on trace element abundances, our model does not require substantial variations in the  
535 composition of slab components, or exceedingly low extents of melting to account for NSVZ arc-front  
536 compositions. In contrast to the unified models of arc-front and rear-arc geochemistry presented in Jacques  
537 et al. (2013, 2014) and this study, other models have proposed unique processes specific to particular  
538 regions of the rear-arc. For example, Søger et al. (2015a;b) invoke the presence of a distinct "Rio Colorado"  
539 mantle component in Southern Payenia, Søger et al. (2013;2015a) suggest that rear-arc basalts of the

540 Nevado volcanic field (~35-36°S) have assimilated up to 70% LCC, and Holm et al. (2016) conclude that both  
541 enriched mantle and eroded crust produce geochemical variation observed in the Northern Segment of the  
542 rear-arc (~34-35 °S). To assess whether these region-specific observations can be accounted for by our  
543 unified model, it is important to consider the characteristics of across-arc geochemical trends and temporal  
544 variability in SVZ subduction geometry.

545

546 The extent to which slab materials have infiltrated the rear-arc mantle can be investigated using trace  
547 element ratios sensitive to the addition of slab melts (e.g., Th/Nb, Th/La, and Ce/Pb; Johnson and Plank,  
548 1999; Plank, 2005; Fig. 10a). Remarkably, while these slab-sensitive ratios generally show declining slab  
549 fluxes with increasing trench distance, some cones retain elevated Th/La and Th/Nb and low Ce/Pb at nearly  
550 600km from the trench (Fig. 10b). Strong slab signals in older (>5 Ma) Andean rear-arc volcanics have been  
551 interpreted, alongside other evidence, to result from transient variations in the dip angle of the subducting  
552 slab (Kay et al., 2006). However, the rear-arc samples of this study were likely erupted between 0.01–0.7 Ma  
553 (Folguera et al., 2009), yet still exhibit trace element patterns that clearly indicate the addition of slab melts  
554 (Figure 1b). One possible explanation for this phenomenon is that this slab signal is still a remnant from the  
555 Miocene period of shallow subduction, and not related to the present-day slab geometry (Søager et al.,  
556 2013; Holm et al., 2016). However, it is difficult to reconcile this model with expectations from geodynamic  
557 models and the chronology of arc-front volcanism.

558

559 In Southern Payenia, it is generally accepted that the slab steepened towards its current configuration at ~5  
560 Ma (Kay et al., 2005). However, Gudnason et al. (2012) attribute the northward propagation of quaternary  
561 rear-arc volcanism to a significant delay in slab rollback in this region, while Folguera et al. (2009) suggest  
562 that the recent volcanism behind the NSVZ is due to extension collapse. Of these proposals, only the model  
563 of Folguera et al. accounts well for activity within the Diamante-Maipo caldera, which has been essentially  
564 uninterrupted since the Late Miocene (Sruoga et al., 2005). This near continuous activity indicates that the  
565 NSVZ slab geometry is unlikely to have shifted substantially over the past 5 Ma. Over 5 My, given a  
566 convergence rate of ~70 km/Ma, the mantle wedge should have turned over ~350 km in between the  
567 establishment of the current slab configuration and the eruption of the youngest rear-arc volcanics,  
568 effectively flushing older slab-derived materials from the wedge entirely. Thus, the slab signatures in recent  
569 rear-arc volcanics were most likely derived from the slab in its current geometry.

570

571 Variation in slab temperature should significantly fractionate many elements in slab melts (e.g. Hermann and  
572 Rubatto, 2009), so provide an additional set of constraints on the dynamics of the SVZ mantle wedge. If  
573 across-arc slab temperature variations are in accord with SVZ thermal models (Syracuse et al., 2010), then

574 Th/La in slab melts may increase by ~25% from the arc-front to the rear-arc, and Th/Nb by nearly an order of  
575 magnitude. This variability in the composition of slab melts should result in rear-arc and arc-front lavas  
576 following distinct mixing trajectories. Yet, rear-arc samples lie on mixing lines between the modelled arc-  
577 front slab melt composition (as used in the model in Fig. 8) and an ambient mantle with varying amounts of  
578 EM1-like enrichment (Fig. 10a). The high slab temperatures beneath rear-arc cones should also destabilize  
579 zircon (e.g., Hirai et al., 2018). However, rear-arc samples have Zr-Hf anomalies similar to TSVZ and SSVZ arc-  
580 front volcanoes (Fig. SB3). These observations imply that the slab components reaching the rear-arc and arc-  
581 front mantle sources were extracted from the slab at similar temperatures.

582

583 Thus, the slab component observed in northern rear-arc cones was likely derived from the slab in its current  
584 geometry, but is unlikely to have been extracted from the slab directly beneath the cones. Two-phase  
585 thermo-mechanical models provide an explanation for this phenomenon. The mantle source at the arc front  
586 may be influenced by slab materials that were extracted from the slab beyond the arc front, and then  
587 deflected towards the trench along compaction pressure gradients (Wilson et al., 2014). The rear-arc  
588 asthenosphere is plausibly influenced by slab melts extracted at shallower depths and then transported  
589 away from the trench in zones of low porosity (Cerpa et al., 2017; Fig. 14b). It is also possible that slab  
590 materials could be transported non-vertically by “mélange diapirs”. However, Cruz-Uribe et al. (2018) found  
591 that melts of such diapirs have >50 wt. % SiO<sub>2</sub>, >18 wt. % Al<sub>2</sub>O<sub>3</sub>, and <8 wt. % MgO, while the most primitive  
592 SVZ rear-arc basalts have >10 wt. % MgO, <48 wt. % SiO<sub>2</sub>, and <15 wt. % Al<sub>2</sub>O<sub>3</sub>, as is typical of arc alkali  
593 basalts generated by low degrees of melting from a lherzolite source (e.g., Baasner et al., 2016). The  
594 experimental melts also have positive Zr-Hf anomalies, unlike the negative Zr-Hf anomalies of SVZ rear-arc  
595 lavas (Fig. 1b). Overall, the geochemical systematics of SVZ rear-arc volcanics appear to be most consistent  
596 with a strong role for slab melting followed by advective transport, compaction channelling, and generally  
597 non-vertical transport of slab-derived materials through the wedge, which provides support for recent  
598 numerical models of two-phase flow (Cagnioncle et al., 2007; Wilson et al., 2014; Cerpa et al., 2017).

599

600 Given the general viability of our unified SVZ model, where geochemical trends in rear-arc and arc-front  
601 lavas can be accounted for by variations in the composition of the ambient mantle, slab additions and  
602 melting variations, it is worthwhile to consider whether this model can also account for the observations  
603 that have motivated region-specific models. Here, we limit this discussion to trends (Fig. 13a-c) that have  
604 been interpreted as evidence for extensive lower crustal assimilation (up to 70%; Søger et al., 2013; Søger  
605 et al., 2015a). While rear-arc compositions on these plots appear to trend toward the compositions of  
606 certain lower crustal xenoliths (Kay et al., 1996), calculated mixing lines with these xenoliths are strongly  
607 hyperbolic, and deviate substantially from the actual rear-arc compositions (Fig. SB5). While “adjusted”

608 lower crust compositions have been proposed to mediate such discrepancies (Søager et al., 2013), the  
609 resulting compositions resemble neither generic models of the LCC (e.g., Rudnick and Gao, 2003), nor the  
610 actual compositions of Andean basement samples (Fig. SB6). These observations, along with the general  
611 difficulty of maintaining high Mg# bulk compositions while enduring such large amounts of crustal  
612 assimilation, motivate a re-examination of the assimilation model.

613  
614 Our analysis suggests that the regional systematics identified by Søager et al. (2013; 2015a; 2015b) can be  
615 fully accounted for by the variations expected from slab fluxes, mantle melting extents, and the ambient  
616 mantle. On Figure 13a, both axes are controlled by mantle and slab components, while the y-axis (La/Sm) is  
617 also affected by variations in F. The arc-front volcanics have a relatively uniform slab contribution, so follow  
618 a trajectory determined by F and the ambient mantle (e.g., Turner et al., 2017). Many of the rear-arc  
619 samples overlap the arc-front data. However, when slab additions are diminished, the extent of mantle  
620 melting often drops to very low values (Jacques et al., 2013), which drives up La/Sm. This is apparent from  
621 the similarly high La/Sm compositions of basalts from the Gough Island chain, which are also alkali basalts  
622 produced by low-F melting of garnet lherzolite. In Gough, the very low degree melting occurs within a region  
623 of enriched ambient mantle, leading to lower  $^{143}\text{Nd}/^{144}\text{Nd}$  and even higher La/Sm than the SVZ rear-arc  
624 volcanics. Similar systematics can account for variability on Figure 13b, though the divergence between arc-  
625 front and rear-arc compositions is less apparent here, as both low-F melt generation and bulk slab additions  
626 skew compositions toward very high Th/Sm (much higher than seen in the Gough basalts). Thus, the  
627 variability on both Figs. 13a and 13b are easily accommodated without a crustal assimilation.

628  
629 Unlike Th/Sm and La/Sm, the highly-incompatible element ratios on Figure 13c are less affected by variations  
630 in melt extent, and can be accounted for by variations in slab flux and mantle enrichment alone. Ba/Th is  
631 particularly sensitive to small variations in the compositions of subducting slab materials and temperatures  
632 of slab melting (e.g., Carter et al., 2015; Patino et al., 2000). This is evident from the high Ba/Th mafic  
633 enclaves erupted at Longaví Volcano, which are thought to be representative of parental magmas at this  
634 centre (Rodríguez et al., 2007). Rodríguez et al. (2007) demonstrate that these lavas reflect input of a  
635 particularly high Ba/Th slab component to the mantle, and suggest this is caused by the proximal subduction  
636 of the Mocho Fracture Zone (Fig. 13d). Hydrothermal alteration, which was likely enhanced along the  
637 fracture zone, produces elevated Ba abundances in both ocean crust and sediment (Kelley et al., 2003; Plank  
638 et al., 2013). Alternatively, the fracture zone may introduce an additional high-Ba fluid component to the  
639 wedge. Subducting fracture zones have inevitably swept along the arc over time, causing transient  
640 geochemical signals among sensitive slab tracers in local volcanics (Fig. 13d). Given the large range in the  
641 Ba/Th ratio of the slab component, there is no need to invoke extremely large quantities of lower crustal

642 assimilation (>70%) to account for the regional systematics. While geochemical variability, when considered  
643 on a small enough scale, necessarily requires case-specific ad-hoc explanations, the regional trends of the  
644 SVZ appear consistent with the relatively simple three parameter model developed here (Fig. 7; Fig. 10; Fig.  
645 13) which incorporates processes seen in subduction zones worldwide.

646

## 647 **5. Conclusions**

648 The origins of trace element and isotopic enrichment in thick-crustal continental arcs have profound  
649 implications for the formation of the continental crust and the petrogenesis of evolved magmas. The new  
650 high-Mg# samples characterized here provide valuable constraints on the compositions of parental magmas  
651 from the thick-crustal Northern SVZ (NSVZ). While evolved NSVZ samples underwent crustal assimilation  
652 during fractional crystallization, the geochemical offsets between the most primitive Southern SVZ (SSVZ)  
653 and NSVZ volcanics cannot be produced assimilation of any basement lithologies either within the Southern  
654 Andes, or globally. Similarly, while variations in the extent of mantle melting and subtle differences in slab  
655 additions may arise due to the thicker lithosphere and deeper slab in the NSVZ, the solution space made  
656 available by these parameters does not encompass the trace element or isotopic compositions of mafic Don-  
657 Casimiro Maipo samples. Only models incorporating an enriched ambient mantle domain that is  
658 compositionally similar to the mantle source of EM1-type OIBs can recreate the observed isotopic and trace  
659 element enrichment in mafic samples from Don Casimiro-Maipo. Coincident arc-front and rear-arc regional  
660 variations in isotopic enrichment suggest that this signal arises from the ambient mantle. A forward model  
661 based on an inversion for an EM1-like mantle source successfully reproduces the isotopic composition and  
662 trace element composition of Don-Casimiro Maipo, demonstrating that this interpretation is consistent with  
663 available experimental constraints.

664

665 The isotopic trajectory of rear-arc lavas is co-linear with other South American lavas thought to represent  
666 melts of metasomatized sub-continental lithospheric mantle (M-SCLM). This supports a model in which the  
667 enriched mantle source is derived from the addition of M-SCLM material to a depleted mantle wedge by  
668 erosion or delamination from the cratonic lithosphere behind the arc. An alternative scenario in which the  
669 EM1 signal results from subduction erosion of continental material produces unsuitable isotopic systematics  
670 and erratic trace element patterns. Additionally, subduction erosion of fragmented basement domains fails  
671 to account for coherent along and across-arc geochemical variations. However, unlike the longitudinally  
672 fragmented and highly heterogeneous Andean basement, the composition of the M-SCLM appears to be  
673 relatively homogenous across a large area of South America. Thus, unlike subduction erosion or crustal  
674 assimilation, the trench-ward flow of variably enriched mantle domains by corner flow accounts for the  
675 coherent latitudinal variations in the isotopic ratios of arc-front and rear-arc lavas.



676

677 In addition to chemical diversity induced by variable M-SCLM addition, rear-arc cones receive highly variable  
678 slab inputs that are surprisingly similar in composition to the arc front. This, along with variable melt extents,  
679 accounts for the offset trajectories of rear-arc and arc-front lavas, and provides support for recent two-  
680 phase flow models indicating non-vertical transport of slab materials (Cagnioncle et al., 2007; Wilson et al.,  
681 2014; Cerpa et al., 2017). This unified model of rear-arc and arc-front geochemical variability is able to  
682 account for the geochemical variations previously attributed to a variety of region-specific processes.

683

684 Overall, we demonstrate that the enriched trace element and isotopic compositions of primitive lavas in the  
685 thick-crust NSVZ are produced by low-degree melting of an enriched ambient mantle that is  
686 metasomatized by slab melts (Turner et al., 2017), rather than crystal fractionation and crustal assimilation.  
687 The similarities between the trace element signatures of primitive NSVZ lavas and the bulk continental crust  
688 (Fig. 1) implies that ambient mantle enrichments may play a major role in the origin of enriched continental  
689 compositions globally. Reevaluation of enriched signatures in primitive lavas of other thick-crust arcs (e.g.,  
690 Cascades, Mexico, Guatemala, Colombia) will allow assessment of the global importance of ambient mantle  
691 enrichment. If similarities with the SVZ are found, models of continental crust generation and elemental  
692 cycling within the silicate earth will require re-evaluation.

693

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698 isotopic analysis in Leeds.

699

700

#### 701 **Figure Captions**

702

703 Fig. 1– Incompatible trace element “spider diagrams”, normalized to D-MORB (Gale et al., 2013). a)  
704 Literature data for Don Casimiro-Maipo (Hickey et al., 1986) has incompatible trace element abundances  
705 comparable to bulk continental crust (Rudnick and Gao, 2003) and the Northern Cascades (only samples with  
706  $Mg\# > 0.64$  for  $Fe^{3+}/Fe_T = 0.8$ , and  $Eu/Eu^* > 0.9$  are shown for Mt Ranier and Mt Hood; Sisson et al., 2014;  
707 Baggerman et al. 2011). Don Casimiro-Maipo shows significantly more enriched trace element abundances  
708 than the TSVZ, SSVZ, and the majority of island arcs (GEOROC compilation). Villarrica, which has a similar  
709 trace element signature to other SSVZ centres (Llaima, Puyehue and Osorno shown here; Turner et al. 2016),

710 is used as a representative SSVZ end-member in figures 3, 4, 5, 7, 8 and 9. b) Trace element data collected in  
711 this study for 7 rear-arc cones, and the five most primitive samples from Don Casimiro-Maipo (circled on Fig.  
712 4). The observed HREE depletion in Don Casimiro-Maipo samples indicates that garnet is residual in the  
713 mantle source during mantle melting. Rear-arc samples show distinctive arc-like signatures, and even  
714 greater trace element enrichment than Don Casimiro-Maipo samples.

715

716 Fig. 2-Geographical context and sample locations. a) Map of the SVZ, showing the segment boundaries  
717 defined by Dungan et al. (2001). The locations of arc-front and rear-arc samples investigated in this study are  
718 shown, along with the locations of arc-front volcanoes passing the filters of Turner et al. (2016). The  
719 geographical regions containing inferred M-SCLM melts (see Fig. 11) are overlain. Base map from GEOMAP  
720 APP. b) Sample locations of Don Casimiro and Basal Maipo lavas within the Diamante-Maipo Caldera  
721 Complex. Imagery from Google Earth. c) Moho depth (from the model of Tassara and Echaurren, 2012) vs.  
722 volcano latitudes projected along the vector of convergence to the trench. d-e) Incompatible major and  
723 trace element 6-values (Turner et al. 2016, and this study) increase northwards, mirroring changes in crustal  
724 thickness.

725

726 Fig. 3- a) SVZ rear-arc samples which are not dominated by slab inputs ( $Th/Nb < 0.4$ ) plot along the “mantle  
727 array”, between DMM and EM1-type ocean island basalts. These rear-arc samples greatly extend the  
728 isotopic range observed among average rear-arc compositions from other arcs (black dots; Turner et al.,  
729 2017). Literature values from Don Casimiro-Maipo lie at the enriched end of the global rear-arc array (Hickey  
730 et al., 1986). The pink and blue fields envelop all EM1 and EM2-type ocean island basalts from Stracke et al.  
731 (2003). The enriched isotopic end-member used for quantitative modelling is from Pitcairn (57DS9; Stracke  
732 et al., 2003). The isotopic composition of the Baseline Andean Mantle (BAM) end member was taken from a  
733 SVZ rear-arc sample with no trace of a slab contribution (e.g.  $Th/Nb < 0.06$ ,  $Ce/Pb > 25$ ; Søger et al., 2013;  
734 sample 126171). The Nazca Plate sediment composition is the average of ODP1232 (Turner et al., 2017). b)  
735 Don Casimiro and Maipo and rear-arc isotopic data from this study. Literature data from Villarrica and the  
736 rest of SVZ that pass filters indicating they are not substantially affected by crustal contamination or mixing  
737 (Turner et al., 2016) are overlain. The filtered rear-arc samples shown in a) are also shown in b), along with  
738 unfiltered data from the entire Diamante Caldera.

739

740 Fig. 4- a) Don Casimiro-Maipo samples from this study are high Mg# basaltic andesites. Samples circled in red  
741 have major element characteristics indicative of olivine accumulation (Fig. SB2). The five most primitive  
742 samples used to characterize the composition of primitive magmas from Don Casimiro-Maipo are circled in  
743 grey. Rear-arc samples are primitive basalts-trachybasalts. Mg#s were calculated assuming  $Fe^{3+}/Fe_T = 0.3$  for

744 the arc-front (Holm et al., 2016), and  $Fe^{3+}/Fe_T = 0.15-0.3$  for the rear-arc (Søager et al., 2015b, cyan and blue  
745 triangles respectively). b) Eu anomalies are negligible in the most primitive samples of this study ( $>0.9$ ),  
746 whereas the majority of literature samples from the Diamante-Maipo caldera show substantial negative Eu  
747 anomalies.  $Eu/Eu^*$  was calculated with  $Eu^* = (Sm_N \times Gd_N)^{1/2}$  or with  $Eu^{**} = (Sm_N^2 \times Tb_N)^{1/3}$  where no Gd data  
748 was reported. In both cases, element concentrations were normalized to CI chondrite (McDonough and Sun,  
749 1995). c-d) Mixing arrays between eight SVZ basement samples and primitive Don Casimiro-Maipo samples  
750 encompass the chemical diversity within the Diamante-Maipo Caldera (from Lucassen et al. 2001; # 84-31-1,  
751 84-31-2, 84-42-1, 84-42-11, and Lucassen et al. 2004; # 00-77, 00-55, 00-58 and 00-05). However, mixing  
752 between the same 8 samples and Villarrica (grey lines) do not pass through primitive Don Casimiro-Maipo  
753 samples. Models originate from the primitive average of Don Casimiro-Maipo and Villarrica samples. Error  
754 bars in c) show  $\pm 1\sigma$  of the variation among primitive samples.

755

756 Fig. 5-Crustal assimilation models. a) Mixing lines between 348 crustal lithologies thought to be  
757 representative of SVZ basement (compiled from GEOROC and other sources) and the average composition of  
758 primitive Villarrica samples. Only a small proportion pass near Don Casimiro-Maipo. b) Trace element  
759 patterns of the mixed compositions which have isotopic ratios close to Don Casimiro-Maipo (red lines in a),  
760 as well as  $Eu/Eu^* > 0.9$ . Mixtures providing a reasonable match to the La-Ce concentrations (cyan lines) have  
761 trace element patterns that are completely distinct from that of Don Casimiro-Maipo. c) Trace element  
762 signatures of mixed compositions for the small subset of the  $\sim 3000$  plutonic rocks from GEOROC that have  
763 sufficient concentrations of Zr, Sr and  $P_2O_5$  to recreate the enrichment of these elements at Don Casimiro-  
764 Maipo. Mixed compositions were calculated for the % mixing required to recreate Zr concentrations. All  
765 produce erratic trace element patterns that are dissimilar to those observed at Don Casimiro-Maipo.

766

767 Fig. 6 - Across and along-arc variations in rear-arc isotopes. a-b) Rear-arc isotopic ratios form latitudinally  
768 coherent trends, with statistically significant trend lines ( $p < 0.005$ ). A pronounced minimum in  $^{87}Sr/^{86}Sr$ , and a  
769 maximum in  $^{143}Nd/^{144}Nd$  is observed at  $\sim 39^\circ S$ , with a reasonably linear and symmetric decrease to the north  
770 and south that can be explained by variable EM1 enrichment (pink arrows). Scatter about these first order  
771 trends within the rear-arc results from variable slab fluxes (and possible local heterogeneity). The blue bar  
772 shows the range of bulk slab compositions for a mix of 20% sediment melt, 80% AOC melt, incorporating the  
773  $\pm 1\sigma$  variation of the sediment composition from ODP1232 (Turner et al., 2017). Rear-arc samples from this  
774 study fit within the regional trends. Arc-front centres lie close to the rear-arc best fit line, showing that these  
775 spatial trends in enrichment are also present at the arc-front. The deviation in  $^{143}Nd/^{144}Nd$  between the rear-  
776 arc and the arc-front south of  $39^\circ S$  likely results from the sparse sampling of rear-arc centres, making it

777 difficult to resolve regional trends from local variations in heterogeneity. This offset may also reflect the  
778 much smaller slab input to the southern rear-arc compared to the southern arc-front

779

780 Fig. 7 – Trace element and isotopic solution space produced by varying the composition of slab melts, the  
781 extent of melting, and the composition of the ambient mantle. a) The pink field represents the maximum  
782 isotopic solution space available by the addition of varying slab melt additions (up to 2% sediment melt, 10%  
783 AOC melt) to a depleted mantle (“baseline ambient mantle”; BAM), with a trace element composition of  
784 DMM, and isotopic composition of Sample 126171 (Søager et al., 2013). Variable slab temperatures have  
785 very little effect on the size of this isotopic solution space, so are not shown. The grey field represents the  
786 additional solution space made available by the addition of an EM1-type mantle domain to the composition  
787 of BAM. b) The pink field represents the maximum trace element solution space available by varying slab  
788 additions as in a), in addition to variations in F between 0.04 and 0.3. The teal field represents the additional  
789 solution space made available by an increase in slab temperature (with a change in the partition coefficients  
790 of La, Sm, and Nd; see SIA4.1). as in a), the grey field represents variable enrichment by an EM1-type mantle  
791 source. In both figures, it is clear that only the addition of an EM1-type mantle source to the sub-arc mantle  
792 can recreate the composition at Don Casimiro-Maipo. The diamonds represent the best model fits for  
793 Villarrica (F=0.1, 7.5% slab addition consisting of 17% sediment, with no EM1 enrichment) and Don Casimiro-  
794 Maipo (F=0.056, 7.5% slab addition consisting of 23% sediment, with 80% EM1 enrichment). The full trace  
795 element signatures of these models are shown in Fig. 8.

796

797 Fig. 8 – Trace element modelling results. The trace element composition of Villarrica can be well matched at  
798 F=0.1 with addition of 7.5% slab melt consisting of 17% sediment. The extreme trace element enrichment at  
799 Don Casimiro-Maipo is best recreated by addition of 80% enriched source, and slight changes in the  
800 composition of the slab addition (23% sediment vs. 17% at Villarrica). The mantle melting model includes a  
801 parameterization for changing garnet proportions based on melt depths ( $Gt=0.0067*GPa+0.0017$ ); we  
802 assume that melt production at Don Casimiro-Maipo and Villarrica occur at 3.3 GPa and 2.2 GPa respectively  
803 (Turner et al., 2016). The solid black lines in both figures represents the best model fit for each centre at the  
804 enrichment and proportions of AOC:SED mentioned above using low temperature slab partition coefficients.  
805 This provides the best match for Villarrica. The dashed grey lines represent the model fit for each centre at  
806 increased slab temperatures, assuming 70% loss of Cs in the forearc. This provides the best model fit for Don  
807 Casimiro-Maipo. Model envelopes were produced by adding and subtracting  $1\sigma$  of 33 measurements of  
808 ODP1232 (Turner et al., 2017). Melting conditions were generated directly from the physical modelling  
809 results of Turner et al. (2016).

810

811 Fig. 9– Modelling contamination of the mantle source by subduction erosion. a) Mixing lines produced by the  
812 contamination of the Villarrica mantle source by SVZ crustal material introduced by subduction erosion. The  
813 isotopic composition of the Villarrica mantle source was taken from the average composition of primitive  
814 Villarrica samples, and the trace element composition of the Villarrica mantle source was calculated from  
815 the best model fit in Fig. 8. Only a small percentage of mixing lines pass near Don Casimiro-Maipo. b-c) Trace  
816 element and isotopic signatures following melting at  $F=0.056$ ,  $P=33\text{kbar}$  of the mixed mantle compositions  
817 that produce isotopic ratios close to Don Casimiro-Maipo (red lines in a), and have  $\text{Eu}/\text{Eu}^* > 0.9$ . None of  
818 these melts recreate Sr, Ba and  $^{87}\text{Sr}/^{86}\text{Sr}$  systematics at Don Casimiro-Maipo. Error bars show  $1\sigma$  of primitive  
819 samples at Don Casimiro-Maipo and Villarrica based on 5 and 3 analyses respectively.

820

821 Fig. 10- Variation in rear-arc slab inputs. a) Mixing lines for slab sensitive ratios following the progressive  
822 addition of slab melts (20% sediment, 80% AOC) to a depleted and enriched (80% EM1 enrichment) rear-arc  
823 mantle source ( $F=0.02$ , melting at 3.3 Gpa). The orange wedge reflects progressive mixing of 10% slab (with  
824  $\text{ODP1232} \pm 1\sigma$ ) into a mantle source with 0% EM1 enrichment, and the green wedge reflects mixing of 17%  
825 slab (with  $\text{ODP1232} \pm 1\sigma$ ) into a mantle source with 80% EM1 enrichment. Rear-arc trace element ratios  
826 follow these mixing lines remarkably closely. The small amount of scatter away from these modelled  
827 trajectories likely reflects variable conditions of melting along the rear-arc. b) Trace element ratios sensitive  
828 to slab additions versus trench distance. MORB-like ratios (blue histogram; Gale et al., 2013) start to appear  
829 at distances greater than 425 km. Remarkably high, arc-like ratios (black histogram shows primitive samples  
830 from the SVZ arc-front; Turner et al. 2016) are seen up to 600 km from the trench, with large scatter in ratios  
831 at large trench distances, contrary to the narrow range of observed isotopes.

832

833 Fig. 11 –Isotopic compositions of inferred SCLM melts across South America (approximate locations on Fig.  
834 2a). These clearly extend the array that passes from DMM/DAM through the arc-front and rear-arc, to Don  
835 Casimiro-Maipo. The field produced by contaminating the Villarrica mantle source with crustal lithologies  
836 (Fig. 9a) is also overlain (grey). This follows a much shallower trajectory than SCLM melts, and is significantly  
837 broader than the SCLM field. Mobile belt M-SCLM melts represent alkaline volcanics erupted on the edges of  
838 the Sao Fransisco craton and the Rio Apa-Luis Alves craton on the east coast of South America (Gibson et al.,  
839 1995; 2005; Carlson et al., 1996). Southern Rift melts were erupted just north of the NSVZ (Lucassen et al.,  
840 2002). Carboniferous granitoids are from the Santo Domingo Complex of the Coastal Batholith (Parada et al.,  
841 1999). All are suggested to have sampled M-SCLM. The green and blue lines show a recycling model in which  
842 low-degree melts are extracted from the depleted mantle (Salters and Stracke, 2004) and then aged for up  
843 to 2 Ga using the parameters from Stracke et al. (2003; see Turner et al., 2017). A model with  $F=0.003-0.008$   
844 encapsulates the range of isotopic compositions observed in SCLM melts.

845

846 Fig. 12- Schematic diagrams showing two possible explanations for the isotopic enrichment minimum and its  
847 relationship to tectonic features. a) A Permian suture zone corresponds with an absence of M-SCLM. b) A  
848 vertical and horizontal slab tear at  $\sim 39^\circ\text{S}$  may allow depleted Pacific MORB mantle to upwell, and dilute the  
849 EM1-type enrichment in the sub-arc mantle located above the tear. As the proportion of depleted Pacific  
850 MORB mantle declines northwards and southwards, isotopic signatures become more enriched. Upper plate  
851 geometry adapted from box model of S ager et al. (2015b), pp. 1514.

852

853 Fig. 13 –A unified model for the SVZ can account for the different trajectories of arc-front and rear-arc  
854 samples. a-c) Trace element and isotopic plots showing rear-arc, and filtered SVZ arc-front samples  
855 ( $\text{Eu}/\text{Eu}^* < 0.9$ ), along with the composition of the two mantle components used in this study (BAM, EM1), and  
856 the compositional field of MORB (Gale et al., 2013). The composition of mafic enclaves from Longav  are  
857 overlain to demonstrate the amount of variability that can be generated from the subducting slab (see d).  
858 The unit vectors in a)-c) indicate how varying ambient mantle enrichment,  $F$ , and slab components will alter  
859 a the composition of the Baseline Andean Mantle (BAM). Melting trajectories from the model described in  
860 section 4.2 are shown in a and b) to demonstrate the fractionation of certain trace element ratios by the low  
861 melt extents that likely dominate the rear-arc. S ager et al. (2013; 2015a) suggest that the offset trajectory  
862 of the SVZ arc-front and rear-arc in the cross plots shown in a-c are produced by LCC assimilation. However,  
863 varying degrees of EM1 enrichment, variable slab fluxes and variable melt extents can easily account for the  
864 offsets between these trajectories. d) Large variations in Ba/Th within arc-front and rear-arc lavas can be  
865 accounted for by variable slab additions. For example, mafic enclaves from Longav  volcano show Ba/Th  
866 ratios up to  $\sim 500$ , which have been attributed to the subduction of the Mocho fracture zone beneath this  
867 centre (Rodr guez et al., 2007).

868

869 Fig. 14 – Schematic diagrams of rear-arc slab supply. a) Traditional model of slab supply, with buoyant rise of  
870 slab fluids to the rear-arc from a slab melting at higher pressure and temperature than at the arc-front. The  
871 composition of these fluids should change with increasing trench distance, as the depth to the slab  
872 increases. Additionally, at large trench distances, very little to no slab supply would be expected, due to the  
873 exhaustion of hydrous phases. b) Proposed model of this study, inspired by Cagniocle et al. (2007), Wilson et  
874 al. (2014) and Cerpa et al. (2017). Down dragging of fluids released at arc-front depths explains the similar  
875 composition of arc-front and rear-arc slab supply.

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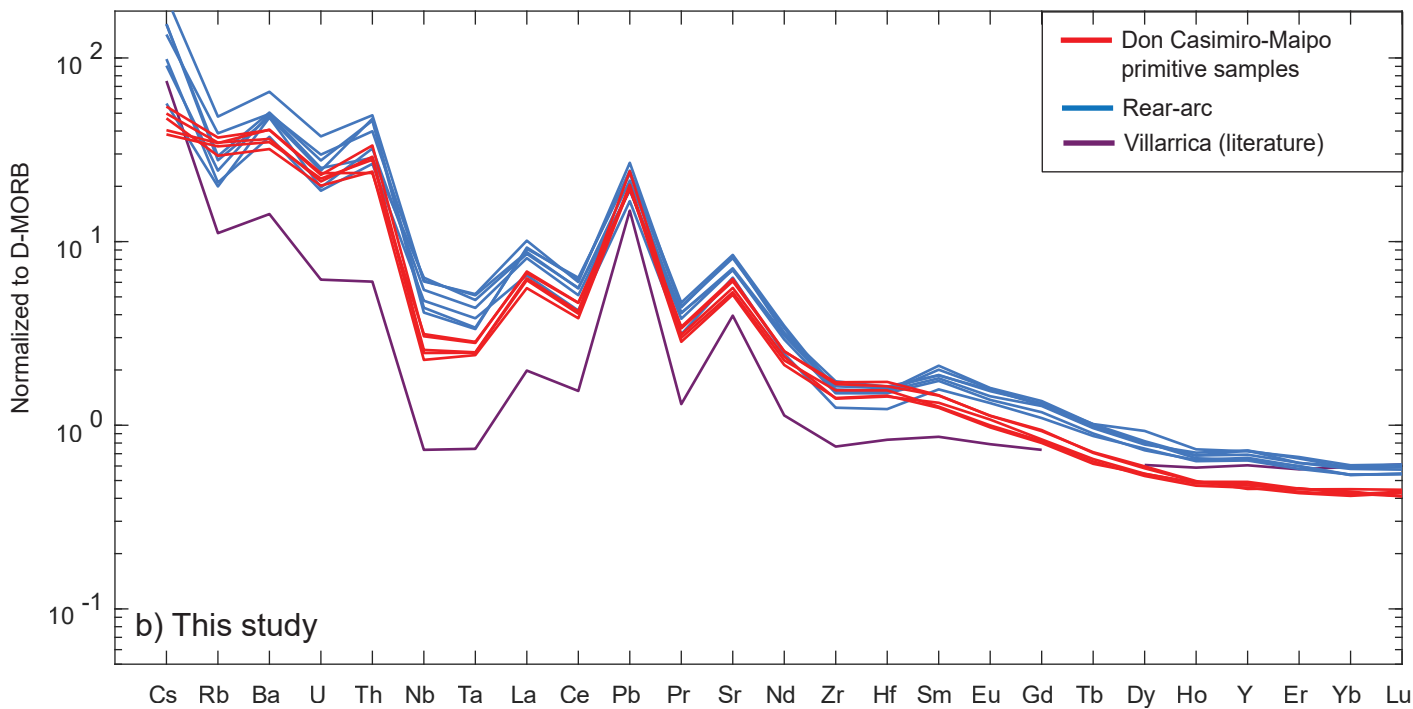
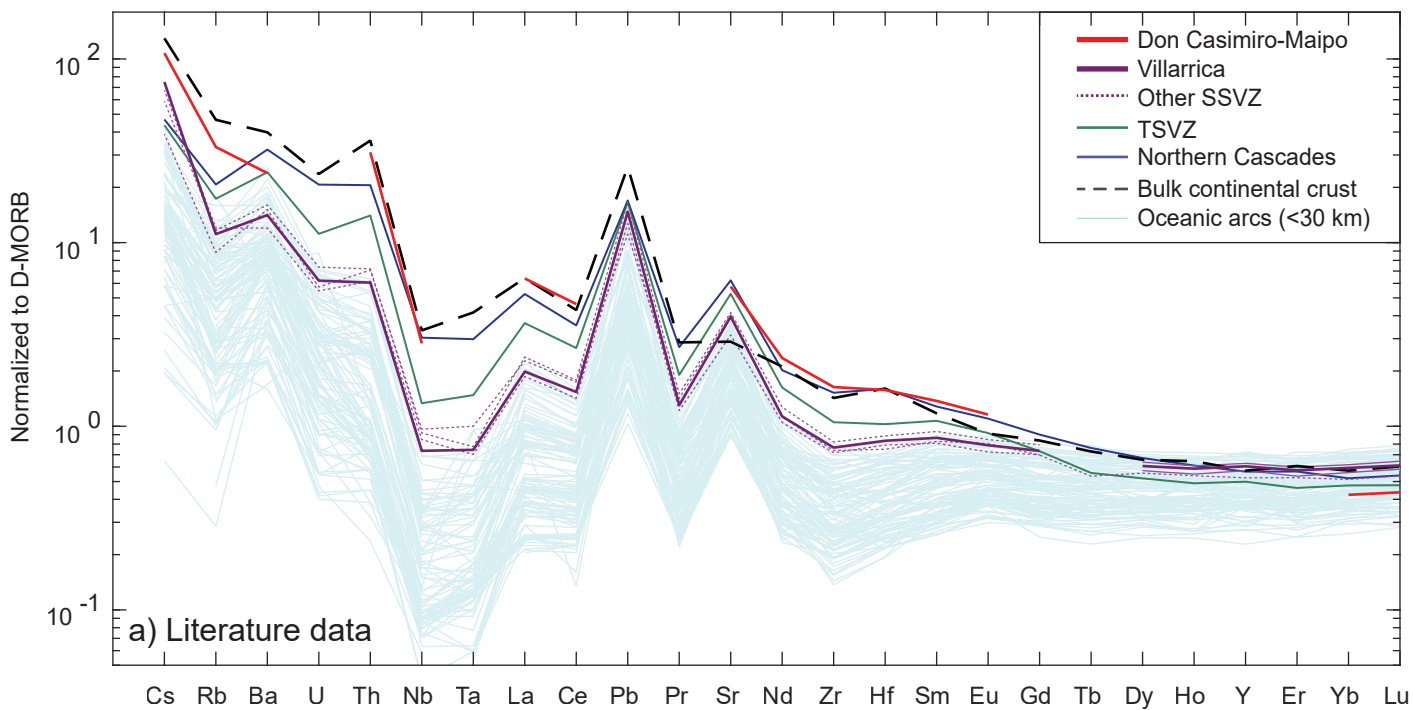
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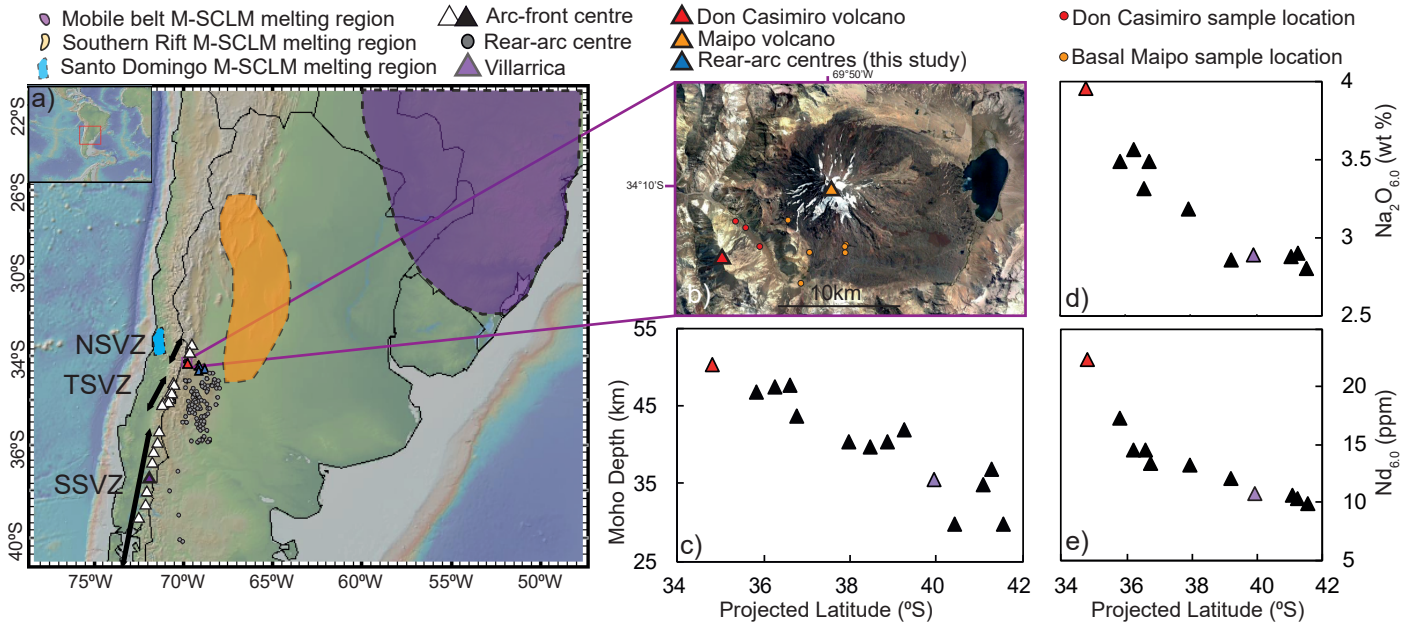
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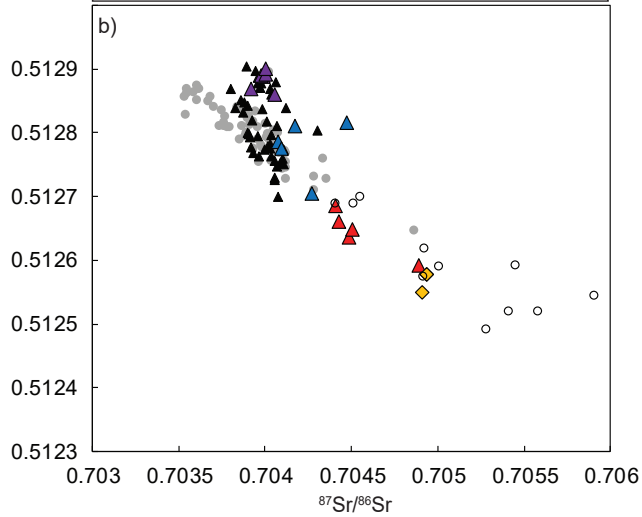
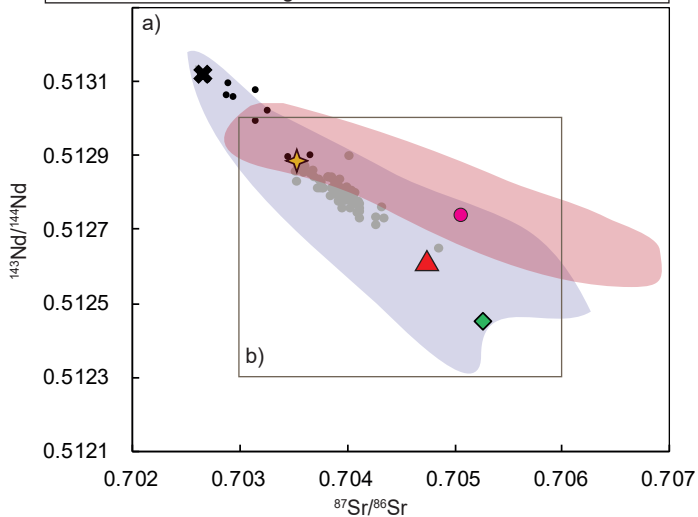
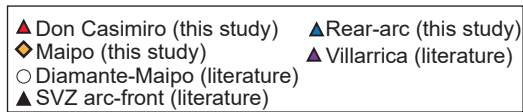
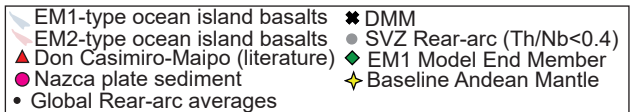
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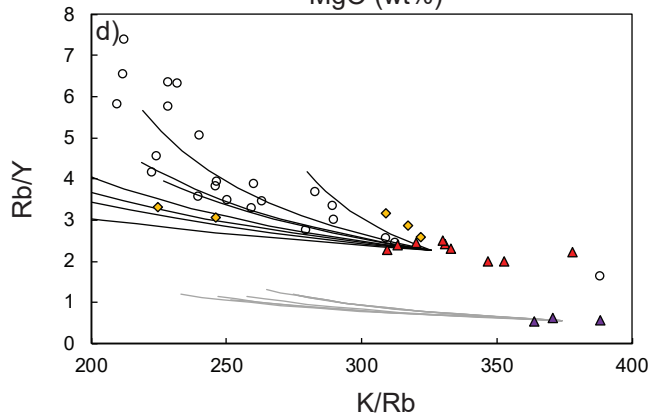
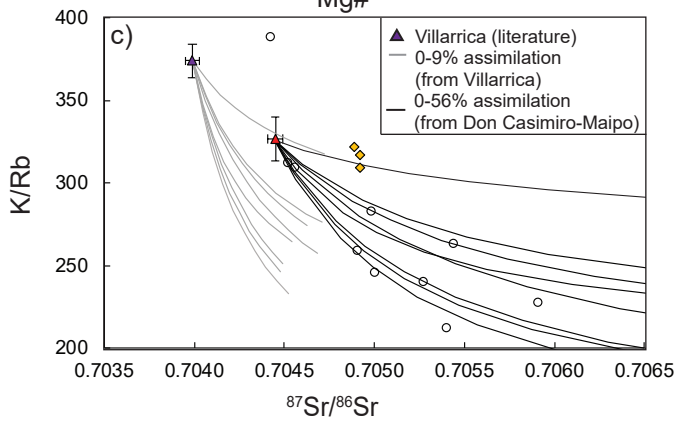
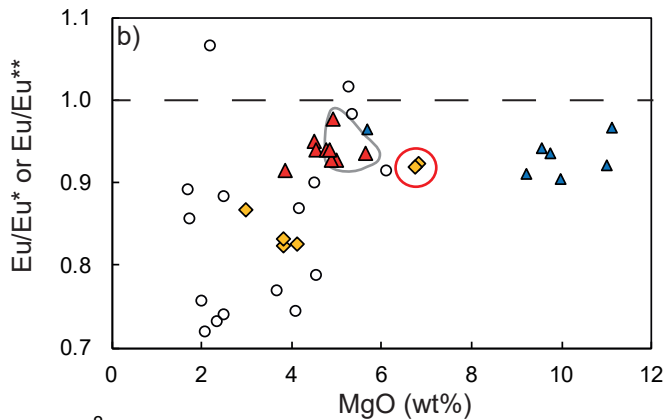
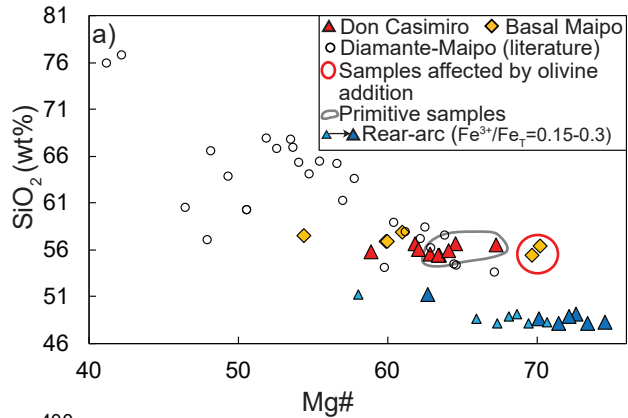
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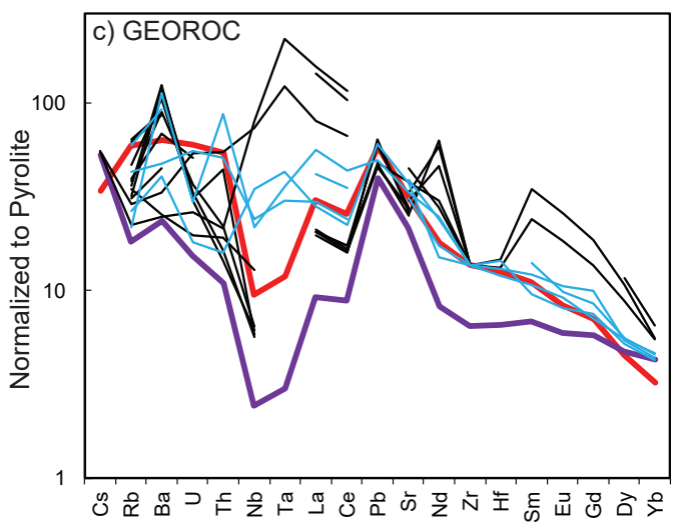
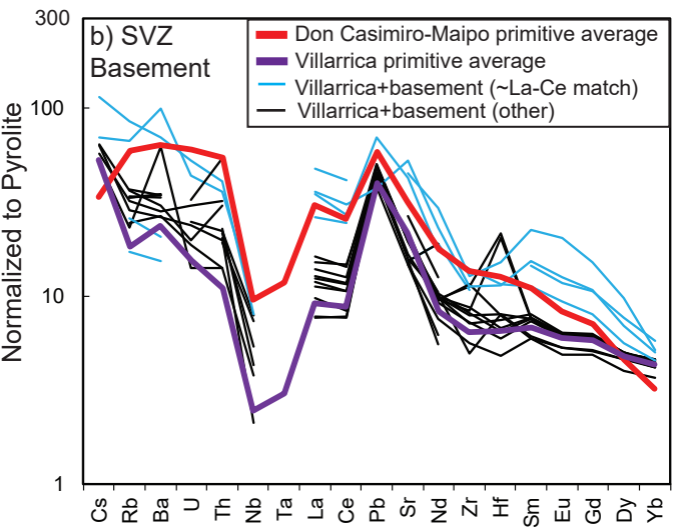
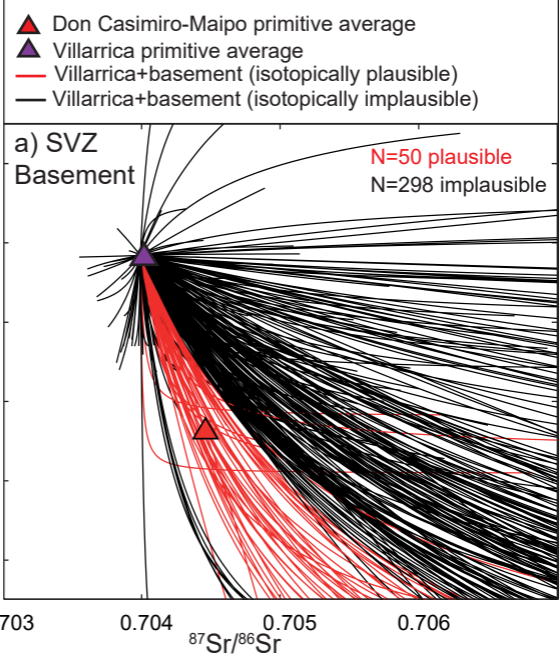
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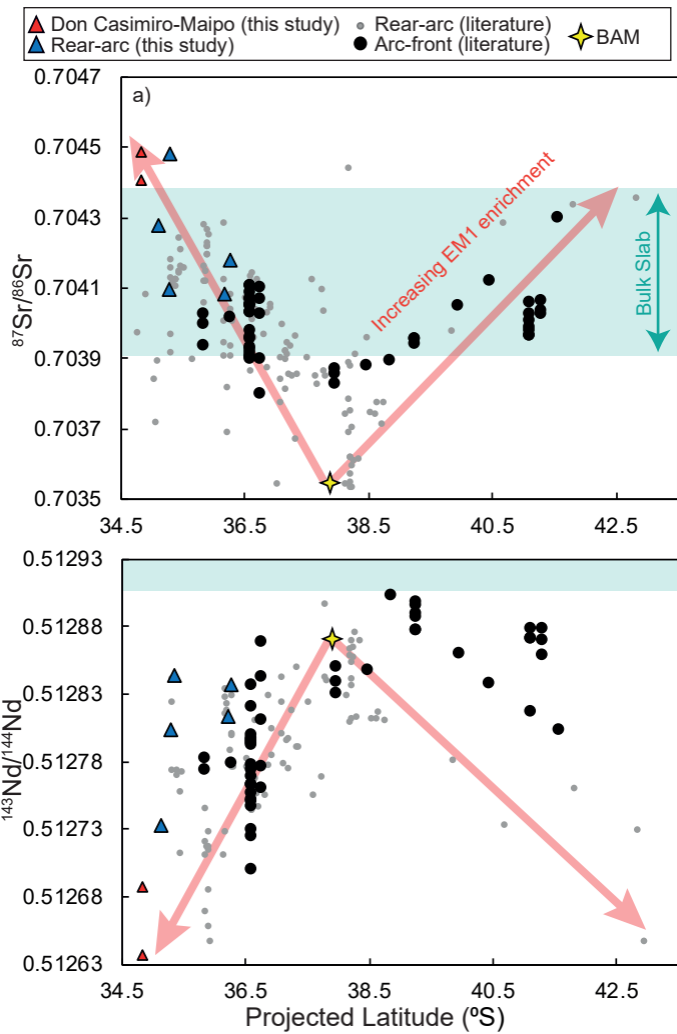


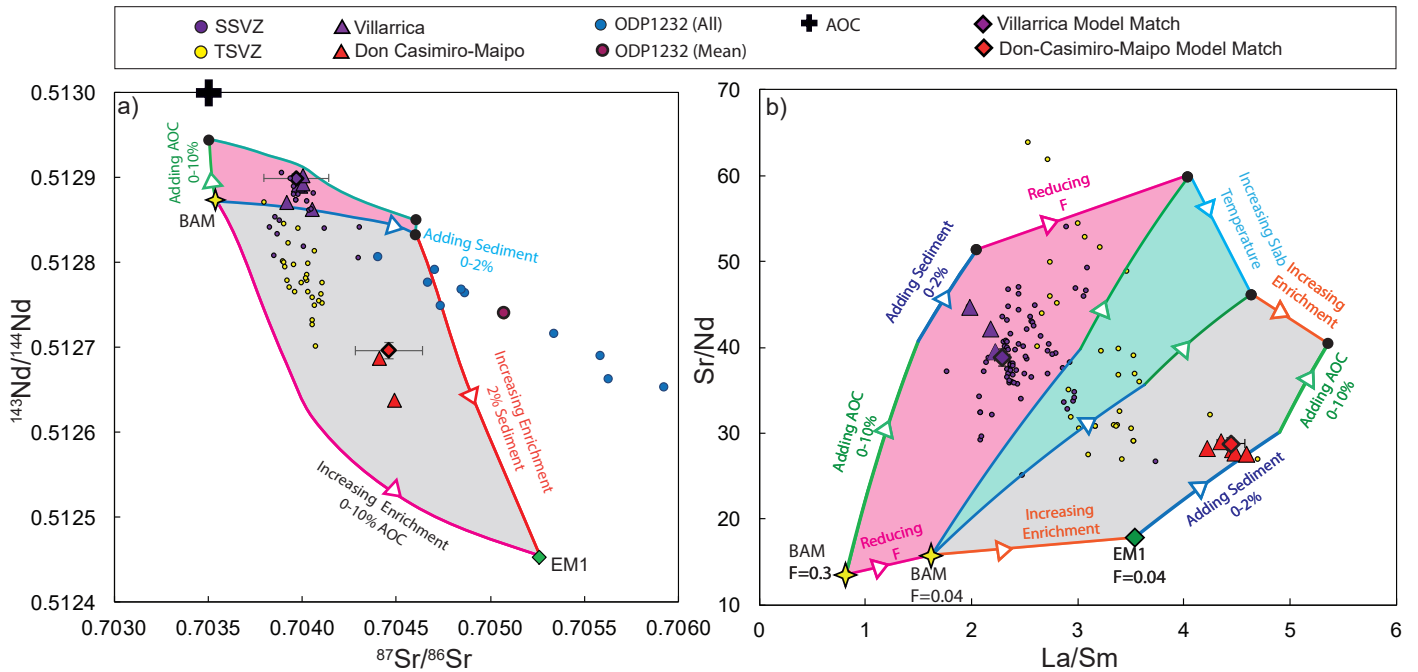




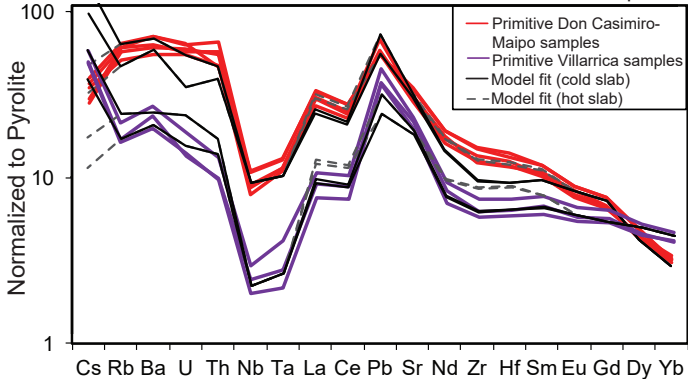




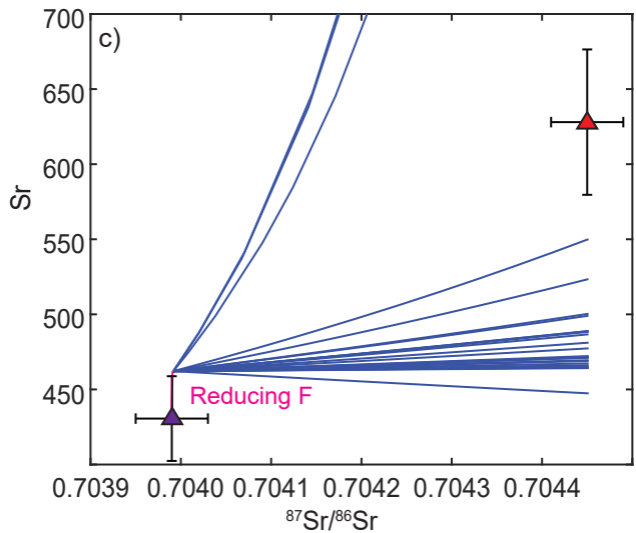
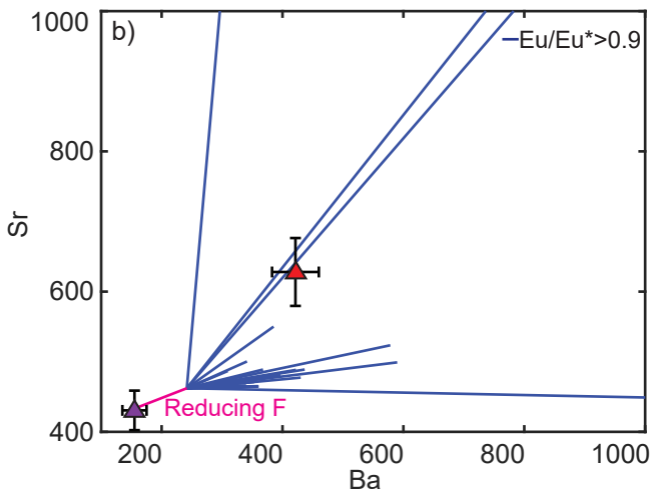
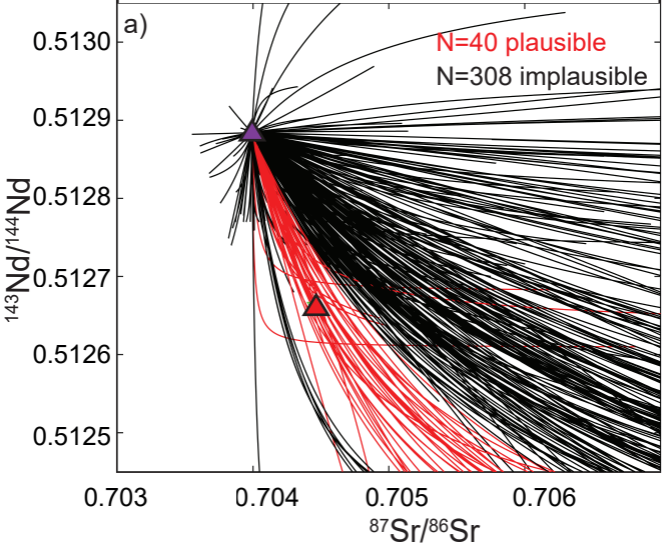


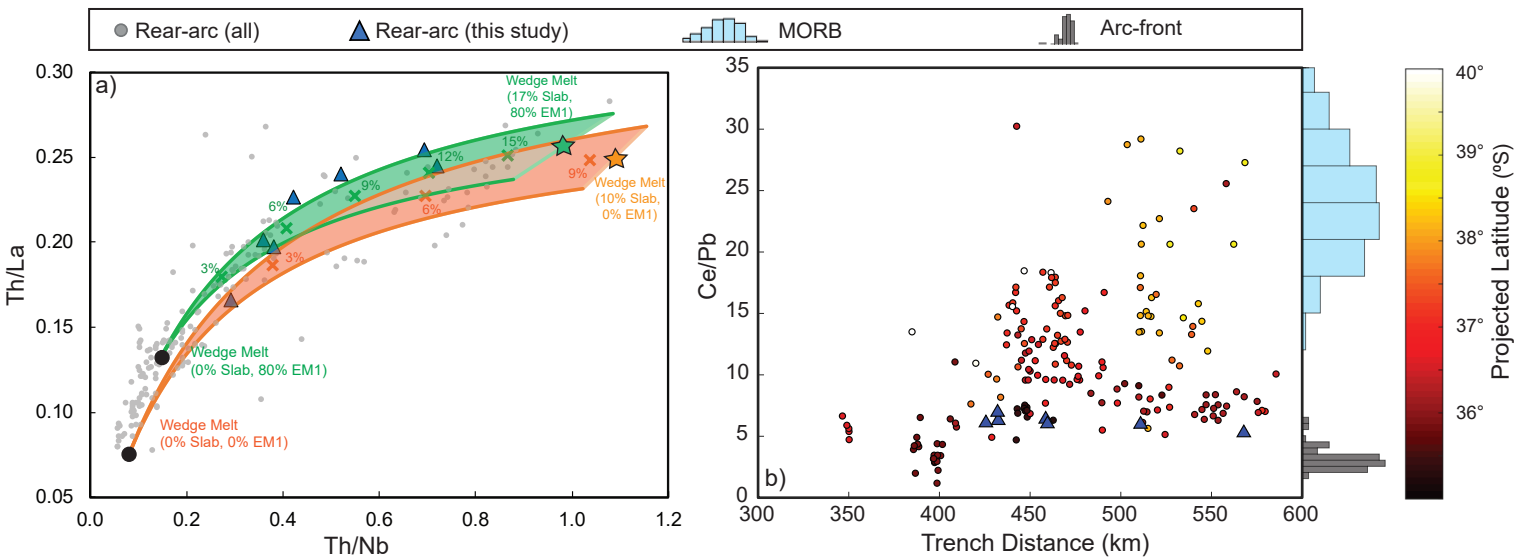


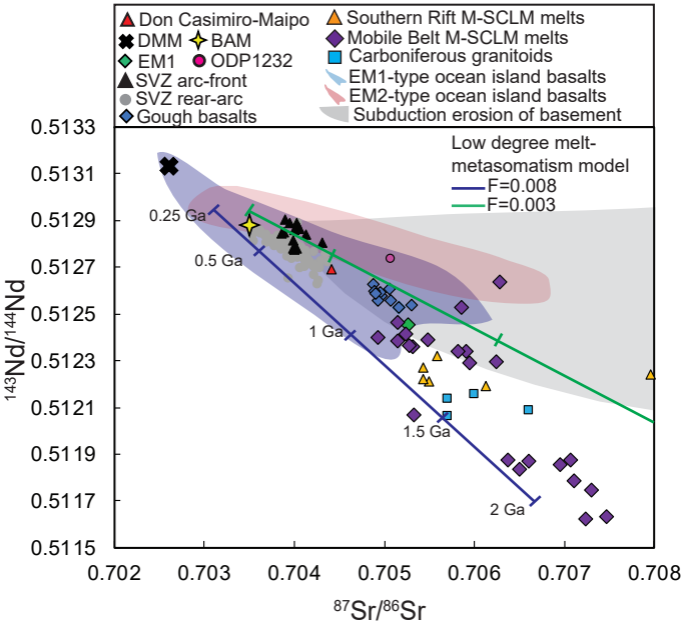
Data vs. model results for Villarrica and Don Casimiro-Maipo

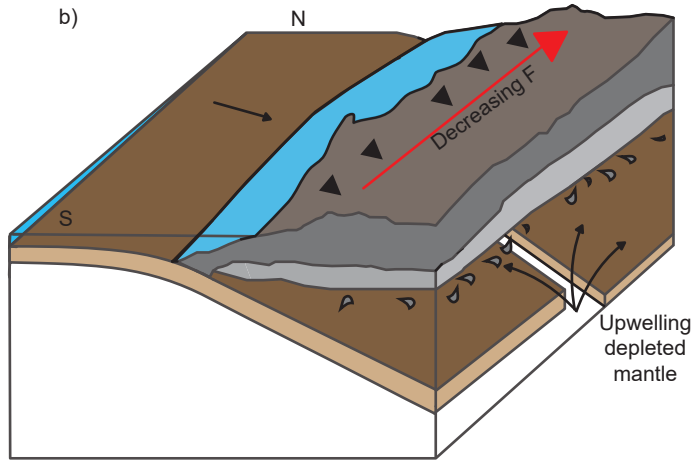
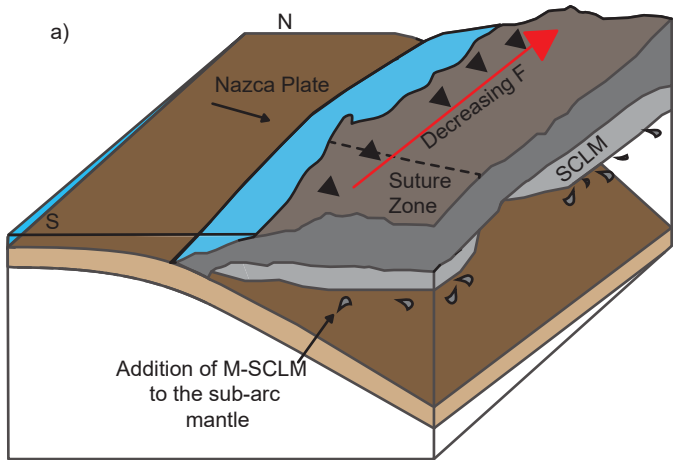


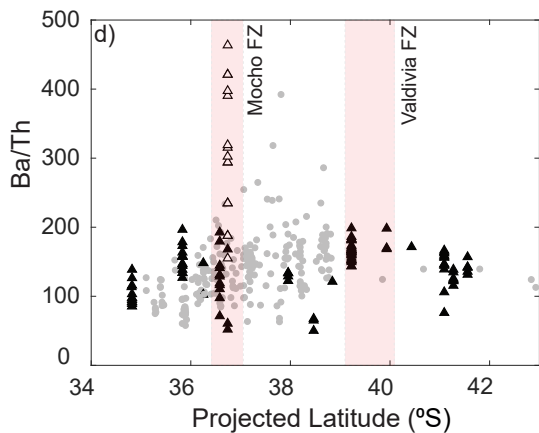
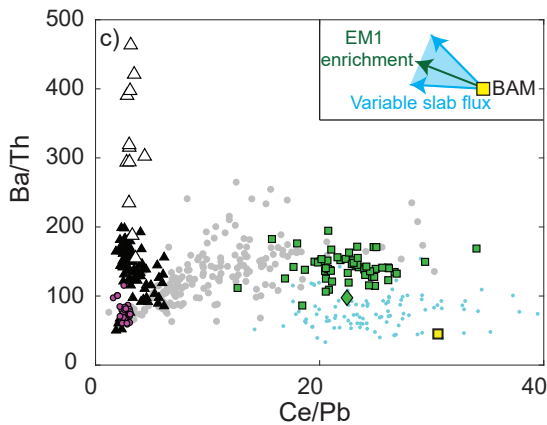
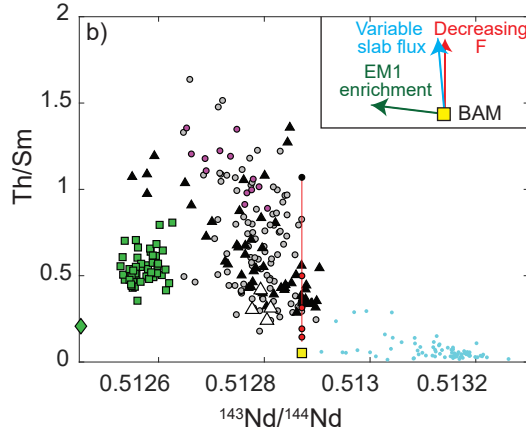
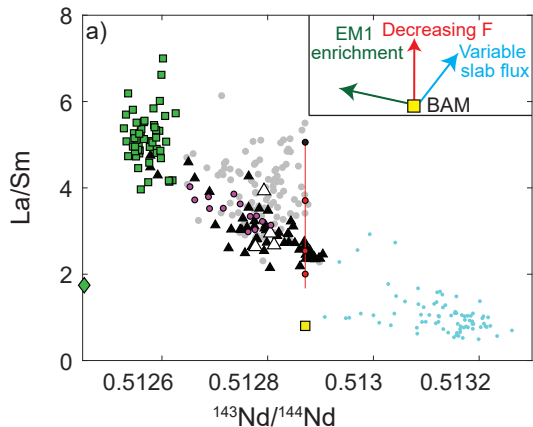
- ▲ Don Casimiro-Maipo primitive average
- ▲ Villarrica primitive average
- Isotopically plausible
- Isotopically implausible



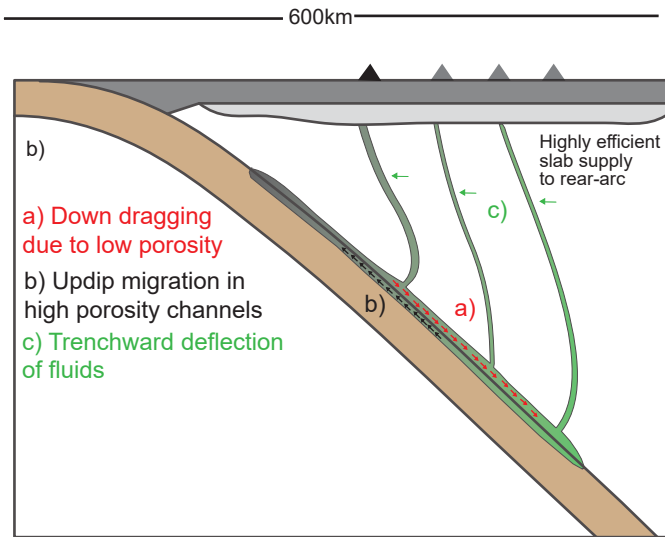
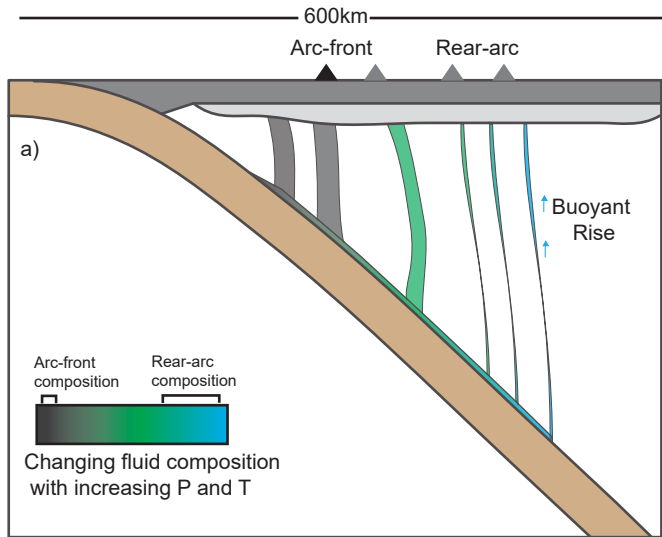












## Appendix 1: Methods and Sample Details

### SIA1 Data and Databases

1 Whole-rock and olivine analysis from this study is shown in Appendix C. Compiled rear-arc  
 2 and Diamante-Maipo literature data used in this study is in Appendix D. The compilation  
 3 of SVZ basement compositions is in Appendix E.

### 4 SIA2 Sample Characteristics

5 Six 100 $\mu$ m sections were made for Don Casimiro-Maipo samples spanning a range of MgO  
 6 contents. These were scanned using a 5MP slide scanner. Point counting was conducted in  
 7 JMicroVision by overlaying a 1000 point grid on these scans. Microscope observations were  
 8 made on any crystals that couldn't be identified with certainty in scans. Crystals smaller  
 9 than  $\sim 0.05$ mm were classified as matrix. Amorphous or crystalline material sufficiently  
 10 distinct from the surrounding matrix that could not be identified as a mineral phase was  
 classified as alteration.

**Table SA1:** Point counting results for Don Casimiro-Maipo samples (ordered by decreasing wt% MgO).

Sample	MgO (wt%)	Crystals (%)	Matrix (%)	Vesicles (%)	Alteration (%)
G0020315-1	6.82	11	89	0	0.1
G0020315-2	6.73	10	90	0	0.3
G0270215-1D	5.63	17	81	0	0.3
G0270215-1C	4.85	14	85	0	0.2
G0030315-4	3.87	11	72	16	0.0
G0030315-1	2.99	42	48	0	10

**Table SA2:** Percentage of each phenocryst phase in the crystal fraction. \*The fine grained nature of these samples meant that phenocryst phases were harder to identify, so proportions are not as reliable as for other samples.

Sample	Olivine (%)	Clinopyroxene (%)	Plagioclase (%)	Oxide (%)	Orthopyroxene (%)
G0020315-1	66	34	0	0	0
G0020315-2	66	34	0	0	0
G0270215-1D*	6	15	46	33	0
G0270215-1C*	8	20	32	41	0
G0030315-4	7	5	89	0	0
G0030315-1	3	13	82	0	3

**Table SA3:** Description of Don Casimiro-Maipo samples in thin section.

Sample	Olivine	Plagioclase	Clinopyroxene	Matrix Textures
G0020315-1	Predominantly subhedral-anhedral 2-0.5mm grains, and smaller, <0.5mm anhedral grains. Both types form clots, and show association with clinopyroxene. Most olivines contain oxide inclusions.	No phenocrysts.	Subhedral-anhedral 2-0.5mm grains, often forming clots. Smaller, <0.5mm anhedral grains within the matrix. Some show breakdown in their cores.	Dark, fine grained matrix with white streaks of amorphous material. Needle-shaped matrix plagioclase aligns with these streaks. A few 1-2mm patches of alteration are associated with reacted olivine cores.
G0020315-2	Euhedral 1-3mm grains, subhedral 0.5-2mm clots associated with clinopyroxene, and smaller isolated <1mm anhedral grains.	No phenocrysts.	Subhedral 1-2mm grains, often associated with olivine clots. As above, some show broken down cores.	Dark, fine grained matrix, with white amorphous material in streaks. Small needle-shaped plagioclase shows alignment around large crystals. Abundant anhedral oxide grains.
G0270215-1D	Anhedral <0.2mm grains with abraded edges. These are associated with patches of small red amorphous material.	Blocky to needle-shaped plagioclase. No alignment.	Sub-anhedral, 0.3-1mm grains. These are more commonly associated with the red material than olivines.	Dark, fine grained matrix, with poorly aligned, needle-shaped plagioclase, and dispersed anhedral oxides. Patches of white, amorphous material.
G0270215-1C	Anhedral, <0.1mm grains.	Elongated needles, 20:1 aspect ratio. Aligned in matrix.	Sub-anhedral 0.3-1mm grains, associated with red amorphous material and oxide growth.	Dominated by small, aligned, needle-shaped plagioclase, with euhedral-anhedral opaque oxides.
G0030315-4	Anhedral, <1mm grains, with highly abraded edges, likely these were originally clots.	Large, 2mm+, blocky plagioclase, showing sieve textured cores (indicating rapid growth/cooling/heating).	Euhedral, 2-4mm grains, and smaller, <1mm highly abraded anhedral grains.	Highly vesicular texture. Matrix dominated by glass.
G0030315-1	A few subhedral, 1-2mm olivines. Some associated with large, 1-2mm cubic oxide phases.	Several populations of plagioclase. 1) blocky 2-4mm rectangular macrocrysts, with large fluid inclusions. 2) lath shaped 1-3mm grains, with inclusion-free cores. 3) needle shaped (<1mm) grains, with sieve textured cores.	Euhedral-subhedral, 1-2mm grains with oxide inclusions. Some are also associated with very large, 1-3mm oxide grains. Additionally, one 4mm clot of several grains and oxides is observed. In some places, clinopyroxene forms anhedral intergrowths with plagioclase, and orthopyroxene.	Extremely crystal-rich. Large areas of amorphous material within the slide are associated with large blocky plagioclase macrocrysts. Between phenocrysts, the matrix is dark and amorphous.

## 12 SIA3 Methodologies

### 13 SIA3.1 Whole Rock Chemistry

14 Sixteen samples from Don Casimiro-Maipo and seven rear-arc samples were prepared for whole-rock analysis  
 15 at the University of Oxford. Arc-front lava samples were cleaned in an ultrasonic bath until the water ran  
 16 clear. These were then split into 2-3cm chunks using a pneumatic rock splitter. Rear-arc samples were wet  
 17 sieved, and unaltered clasts from the coarsest fraction (>2mm) were hand picked for whole rock analysis.  
 18 Both sample types were crushed in a steel jaw crusher, and then powdered in an agate ball mill.

19  
 20 **Major Element Analysis:** Major element analyses was conducted by using a PANalytical Axios Advanced  
 21 X-ray fluorescence spectrometer at the University of Leicester, UK, following the methods of Knott et al. (2016).

22  
 23 **Trace Element Analysis:** ~ 100 mg of each powder as well as reference materials BCR-2a, BHVO-2,  
 24 W-2a and AGV-2, were digested in ~4ml of 7.5M HNO<sub>3</sub> and ~1ml of 24M HF. The samples were then  
 25 evaporated and redissolved twice in ~4ml 7.5M HNO<sub>3</sub> to eradicate fluoride precipitates. At this stage, a  
 26 sample split was made for isotopic analysis. Trace element analyses were conducted using a Perkin Elmer  
 27 NexION 350D quadrupole ICP-MS for the majority of elements. An additional 1200x dilution and addition  
 28 of an internal standard containing Rh, Re, and In was achieved using the ESI prepFAST system producing  
 29 a 6000x solution in which to measure lower concentration elements. All prior dilutions were performed  
 30 gravimetrically. Sample G0030315-4 was run every four samples, and the entire run was duplicated in reverse  
 31 order to monitor drift. Drift correction and data reduction were carried out following Turner et al. (2013).  
 32 Due to malfunctioning of the Prepfast system, a planned run at 60,000x dilution for higher concentration  
 33 elements using the same method failed. Thus, Sr and Ba concentrations were measured using a Thermo  
 34 Element 2 ICP-MS (also at the University of Oxford) at a 1,200,000x dilution (performed gravimetrically).

35 The digestion procedure, and sample homogeneity can be assessed by calculating the % difference between  
 36 the measured element concentrations of two duplicates produced from the powders of this study, and two  
 37 from an additional set of samples digested alongside these. Most measurements lie within 5%, suggesting  
 38 that the digestion process, and samples, were reasonably homogeneous (table SA4).

#### 39 Isotopic Analysis

40 Strontium (Sr) and neodymium (Nd) isotope analyses were performed on a Thermo Scientific Triton-series  
 41 multicollector mass spectrometer at the School of Earth and Environment, University of Leeds. Eleven  
 42 whole-rock powders were digested alongside those for trace element analysis, followed by subsequent drying  
 43 and pick up in concentrated ultrapure HNO<sub>3</sub> and HCl acids. The final dried samples were diluted and  
 44 centrifuged for 15 minutes at 2000 rpm. Sr and Nd were extracted by using Sr-Spec and TRU-spec resins and  
 45 conventional ion-exchange chromatographic techniques. <sup>87</sup>Sr/<sup>86</sup>Sr and <sup>143</sup>Nd/<sup>144</sup>Nd ratios were normalized for  
 46 mass fractionation to <sup>86</sup>Sr/<sup>88</sup>Sr=0.1194 and <sup>146</sup>Nd/<sup>144</sup>Nd=0.7219. Samples for Sr were loaded onto previously  
 47 outgassed W filaments and samples for Nd were loaded onto previously outgassed Re filaments. Sr and Nd  
 48 were analysed while the current was maintained between 3-5 V (for <sup>88</sup>Sr) and 0.4-0.8 V (for <sup>144</sup>Nd).

49 The average <sup>87</sup>Sr/<sup>86</sup>Sr obtained from repeated measurements of NIST SRM-987 during the course of this  
 50 study was 0.710271. Similarly, we measured <sup>143</sup>Nd/<sup>144</sup>Nd=0.511837 for the LaJolla standard. Two samples  
 51 (G0020315-2 and G0270215-1C) were analyzed for <sup>143</sup>Nd/<sup>144</sup>Nd in a separate run, in which we measured  
 52 <sup>143</sup>Nd/<sup>144</sup>Nd=0.511850 for the LaJolla standard. We also successfully reproduced the published USGS  
 53 standard BHVO-2 values for <sup>87</sup>Sr/<sup>86</sup>Sr=0.703468 and <sup>143</sup>Nd/<sup>144</sup>Nd=0.512995 (Weis et al., 2005). Details  
 54 on the sample preparation and analytical protocols are described in Crummy et al. (2014).

### 55 SIA3.2 Correction of Isotopic data of this study and Literature 56 Data

57 Sr and Nd isotopes from the literature and this study were corrected to be in line with the analytical procedure  
 58 used for SVZ data presented in Turner et al. (2016). Various standards were corrected to be equivalent  
 59 to La Jolla=0.511858, NBS 987=0.710235 based on session averages.

**Table SA4:** % deviation of four duplicates ran during analysis on the "quad" (except starred elements, where % deviation is reported from the "element"). (4) represents a duplicate of G0030315-4, (1) represents a duplicate of MDZ1, 079a and B1 are samples measured in the same run, but for use in a study on the Main Ethiopian Rift. % deviation is defined as (Measurement 1-Measurement 2)/(Average of measurement 1 and 2) x 100.

	% difference (4)	% difference (1)	% difference (079)	% difference (B1)
Li	4.13	1.74	-0.97	0.35
Be	-4.52	60.2	-12.4	1.29
Sc	5.89	-27.0	-20.5	-0.72
V	2.02	-1.95	-0.65	0.15
Cr	3.02	0.41	1.14	-0.99
Ni	3.53	0.74	0.86	1.47
Cu	4.71	-3.94	-0.26	1.76
Rb	3.64	0.28	-9.66	0.30
Sr*	5.54	1.74	0.35	-5.67
Y	2.99	-3.08	-6.18	-0.94
Zr	2.12	-0.29	-0.42	0.32
Nb	3.09	0.21	-0.30	1.16
Cs	8.90	-5.49	-4.43	9.11
Ba*	6.08	3.17	3.65	-2.89
La	3.52	0.58	1.10	0.82
Ce	3.31	-1.09	0.94	0.70
Pr	4.57	0.71	2.05	0.80
Nd	3.29	3.63	-0.88	-0.53
Sm	0.00	2.17	1.35	2.23
Eu	2.21	2.20	1.55	-1.12
Gd	3.09	-2.92	-2.87	1.04
Tb	3.26	-2.17	2.29	0.91
Dy	3.83	-0.60	2.67	0.78
Ho	4.64	-2.73	-0.32	3.88
Er	2.99	4.01	-2.53	3.70
Tm	2.81	-5.67	-3.75	0.51
Yb	2.67	1.49	-0.23	1.69
Lu	1.28	1.41	-3.58	1.15
Hf	3.34	0.93	-0.87	1.40
Ta	-2.98	-0.57	1.60	-0.94
Pb	3.55	-2.14	-9.60	1.72
Th	3.48	-2.05	-17.6	-1.10
U	3.33	-0.09	0.56	-0.11

### 60 SIA3.3 Olivine Chemistry

61 Olivines were picked from the 1000-250 $\mu$ m grain fractions, and mounted in epoxy rings. These  
62 mounts were polished and carbon coated for EMPA analysis. Analysis was conducted with  
63 an accelerating voltage of 15KeV, a current of 100nA and a 10 $\mu$ m spot size. St John's Island  
64 and Fayalite were run as secondary standards, with % recoveries reported in table SA5.

65 LA-ICP-MS analysis was conducted at the University of Cambridge. Three instrument  
66 runs were performed over two days. The run conditions were 20Hz, 80 $\mu$ m, 9.5J/cm<sup>2</sup>, 0.8L/min  
67 of gas flow for He and Ar. Where possible, three spots were placed in the centre of the olivine,  
68 with no overlap. For very small grains, only 2 spots were placed, as it is preferable to have  
69 fewer spots nearer the centre than more spots towards the grain edges, due to the reasonably  
70 high diffusivity of elements of interest to this study. To allow drift correction, 3 spots on each of

71 NISTSRM612, NISTSRM610, 06SGOL03 (inhouse olivine standard) and BD4074OL-7 (a San  
 72 Carlos olivine) were run every 30 analyses.  $^{29}\text{Si}$  was used as an internal standard, with values  
 73 obtained from EMPA grain averages. Signals were selected using Glitter Processing software,  
 74 with normalization to NISTSRM612 via a quadratic drift correction. This was effective at  
 75 removing drift, as determined by observing the stability of NISTSRM610, 06SGOL03 and  
 76 BD4074OL-7 following this correction. Multiple secondary standards were run to assess the  
 77 accuracy of the measurement for elements of interest (NISTSRM610, NISTSRM614, BCR-2G).  
 78 Measurement of inhouse olivines O6SGO103 and BD407OL-7 provided an excellent assessment  
 79 of the success of the drift correction performed on glass standards. Elements of interest in this  
 80 study drifted by less than 7%. Although Mn and Ni recovery on glass standards were good,  
 81 these elements were recovered more poorly on BCR-2G. This, and the fact that previous  
 82 studies in this area have used EMPA Ni and Mn data, resulted in us choosing to use EMPA  
 83 data for consistency, as most of the peridotite-pyroxenite ratios involve normalization to Fe  
 84 (which is reported from EMPA). Detection limits, and 1 sigma uncertainties calculated from  
 85 Glitter for a representative sample (MDZ2) are shown in table SA6.

**Table SA5:** Peak count times, and calibration materials used. % recovery calculated below is defined as Measured Average/Published Average x 100. Where no published value exists for a standard, - is shown. Detection limits are reported from 9 spots on MDZ8. These are assumed to be representative of the analysis as a whole.

	Mn	Fe	Mg	Si	Ca	Ni
Peak count times (s)	20	30	20	20	70	70
Calibration material	Mn metal	Andradite	MgO	Albite	Wollastonite	Ni metal
% Rec Fayalite	97.7	94.2	-	99.6	97.0	-
% Rec St Johns Island Olivine	106.1	96.1	99.6	99.2	-	90.1
Detection limit (ppm)	173	197	173	98	74	137

**Table SA6:** Minimum (99% Confidence) detection limits (ppm), and 1 sigma uncertainties (ppm) for all spots on MDZ2, taken as representative for the analysis as a whole. The mass in brackets represents the isotope measured on the ICP-MS.

Element	MDL	1 $\sigma$ uncertainty
Li(7)	0.115	0.151
Na(23)	1.121	6.92
Al(27)	0.318	39.3
P(31)	13.76	47.3
Ca(43)	133	192
Sc(45)	0.062	0.412
Ti(47)	0.551	18.4
V(51)	0.043	0.504
Cr(53)	0.578	56.0
Mn(55)	0.136	88.4
Co(59)	0.049	7.91
Ni(60)	0.171	191
Cu(65)	0.351	0.416
Zn(66)	0.373	10.1
Ga(69)	0.042	0.049
Ge(72)	0.370	0.208
Y(89)	0.037	0.025
Zr(90)	0.033	0.052
Nb(93)	0.035	0.021

**Table SA7:** % recovery of NISTSRM610 (average of 3 spots). (x,y) refers to when this secondary standard was run, x referring to the instrument run (1-3), and y to the order of measurement within a given run. These values were used to check the success of the quadratic drift correction. Across the three instrument runs, most elements drift by less than 10%.

Element	(1,1)	(1,2)	(1,3)	(1,4)	(1,5)	(2,1)	(2,2)	(2,3)	(2,4)	(3,1)	(3,2)	(3,3)	(3,4)	AVERAGE
Li(7)	98.0	105.5	103.7	106.7	105.9	107.3	104.3	101.1	103.6	108.0	99.6	96.8	98.2	103.0
Na(23)	98.2	105.7	107.0	106.1	105.9	105.9	104.9	101.4	103.9	105.0	103.0	101.1	102.1	103.9
Al(27)	102.3	105.5	105.9	106.2	107.8	108.3	104.2	104.9	107.3	105.6	107.2	103.5	104.9	105.7
P(31)	107.8	105.2	99.7	130.9	194.5	100.5	103.9	101.0	100.5	125.0	113.1	117.6	133.9	118.0
Ca(43)	95.9	103.7	100.6	106.7	107.3	105.2	102.5	99.2	102.2	105.2	100.1	99.5	104.5	102.5
Sc(45)	98.7	109.1	108.0	109.8	108.8	110.2	110.9	105.6	108.9	108.5	106.7	108.0	109.6	107.9
Ti(47)	115.2	120.8	116.9	120.0	119.0	125.1	128.7	124.5	125.5	125.3	125.9	122.8	125.4	122.7
V(51)	95.7	101.9	100.1	105.2	105.8	104.8	102.6	100.1	105.1	105.9	101.7	99.9	103.7	102.5
Cr(53)	103.4	107.7	109.4	113.6	119.5	120.7	115.9	112.1	115.1	114.1	111.4	112.7	116.7	113.2
Mn(55)	94.6	100.0	100.0	103.7	102.4	105.1	100.2	99.0	103.5	104.9	97.1	96.1	97.7	100.3
Co(59)	95.5	103.0	101.7	103.7	100.8	104.2	102.0	98.9	104.3	103.6	101.0	100.3	103.4	101.7
Ni(60)	92.1	100.9	99.8	103.7	99.9	105.6	100.1	98.2	100.0	104.3	101.8	100.5	102.2	100.7
Cu(65)	89.9	98.8	97.3	100.1	97.5	100.5	98.4	94.9	99.0	100.1	96.0	95.9	97.6	97.4
Zn(66)	98.5	104.1	102.3	108.8	110.2	111.1	104.4	100.1	106.6	109.9	106.4	104.1	102.9	105.3
Ga(69)	92.2	97.7	96.5	101.1	101.5	101.7	99.0	97.2	99.4	101.4	98.9	98.3	100.8	98.9
Ge(72)	81.5	86.7	85.7	89.1	88.8	92.4	89.0	85.6	86.6	92.0	87.5	87.8	89.8	87.9
Y(89)	93.1	102.1	99.8	102.8	102.1	104.0	102.9	100.5	100.3	102.5	101.4	101.8	103.8	101.3
Zr(90)	89.8	95.9	94.6	97.9	98.1	98.6	98.5	94.6	95.4	97.5	96.8	96.0	98.7	96.3
Nb(93)	92.8	99.9	98.9	102.4	101.7	102.2	102.2	98.1	100.2	102.9	99.9	99.2	101.1	100.1

**Table SA8:** % recovery of BCR-2G, using averages from 3 spots, and % recovery of NISTSRM614. The % recovery for NISTSRM614 is not shown for the first run. These were the first spots ran by the laser, and it has been shown that the system takes a few spots to warm up and settle after switching from solution mode to laser mode.

Element	(1,1)	(2,1)	(2,2)	(3,1)	(3,2)	BCR-2G Average	(2,1)	(3,1)	NISTSRM614 Average
Li(7)	100.2	104.5	101.7	105.8	94.6	101.4	84.4	103.7	94.1
Na(23)	96.2	99.1	97.0	101.4	96.9	98.1	105.5	103.1	104.3
Al(27)	103.8	103.4	99.9	104.6	103.3	103.0	106.6	106.3	106.4
P(31)	101.3	84.9	86.8	108.4	120.1	100.3	268.2	260.0	264.1
Ca(43)	99.0	97.9	98.9	101.1	99.1	99.2	106.2	103.7	105.0
Sc(45)	97.5	102.1	96.4	100.8	99.9	99.3	217.1	185.9	201.5
Ti(47)	107.3	108.4	105.2	110.9	106.3	107.6	122.1	101.3	111.7
V(51)	97.6	100.1	101.1	102.8	99.6	100.2	109.7	99.9	104.8
Cr(53)	93.2	104.4	92.9	97.1	101.1	97.7	94.8	113.0	103.9
Mn(55)	92.4	96.6	94.9	98.5	90.0	94.5	105.6	89.7	97.7
Co(59)	99.0	97.8	97.0	97.9	96.2	97.6	92.6	95.0	93.8
Ni(60)	89.4	86.9	83.8	90.2	86.5	87.4	107.9	71.8	89.8
Cu(65)	71.0	72.5	67.5	71.1	69.6	70.3	88.8	109.0	98.9
Zn(66)	121.5	134.0	127.4	134.9	120.7	127.7	91.4	89.2	90.3
Ga(69)	111.6	117.1	109.2	118.4	114.5	114.2	96.9	93.9	95.4
Ge(72)	99.6	112.9	95.6	101.1	102.4	102.3	86.3	110.8	98.5
Y(89)	87.3	90.0	81.9	90.6	89.1	87.8	102.0	99.4	100.7
Zr(90)	84.5	87.0	78.3	87.7	86.3	84.7	90.3	91.6	90.9
Nb(93)	88.6	91.6	86.6	91.0	90.2	89.6	93.7	99.3	96.5

## 86 SIA4 Mantle Melting Model Details

87 As discussed in the text, this study uses an updated version of the model of Turner and  
88 Langmuir (2015b). Below, the justification of various inputs to the model are discussed. The  
89 updated Matlab scripts ADD\_SLAB.m, ADD\_SLAB\_HOTP.m and ModeMelt\_Hydrous.m  
90 have been uploaded into the supplementary information.

### 91 SIA4.1 Model Partition Coefficients

92 **Hydrous Partition Coefficients for Mantle Melting:** This study utilizes hydrous  
93 partition coefficients and mantle modes from Turner and Langmuir (2015b; detail in their  
94 supplement). However, we slightly updated the partitioning behaviour of several elements:  
95

96 1) The bulk partition coefficient of Eu is obtained by a log extrapolation from those of Sm and  
97 Gd:

$$98 \quad Kd_{Eu} = (Kd_{Sm} * Kd_{Gd})^{0.5}$$

99 correcting an error in Turner and Langmuir (2015b).

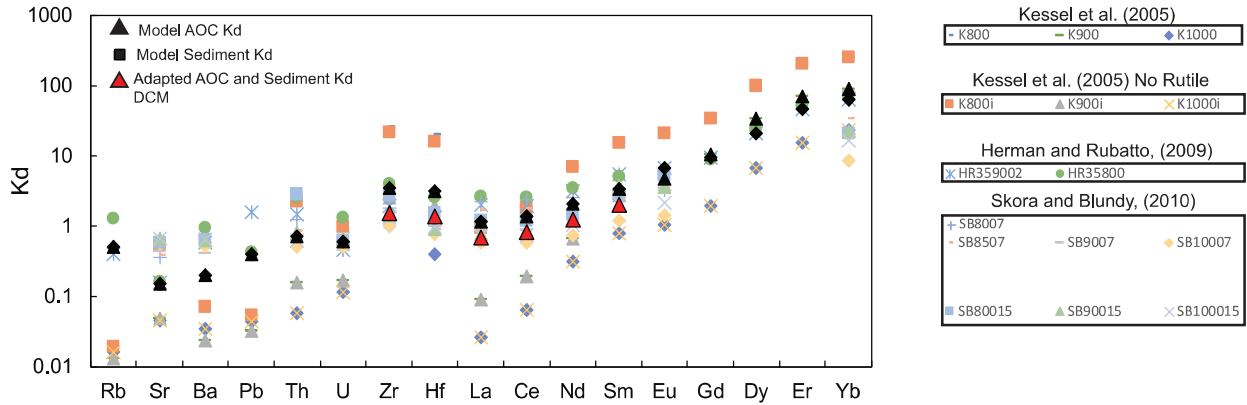
100 2) The partition coefficient of Hf is assumed to be equal to that of Sm.

101 3) The partition coefficient of Zr is approximated as:

$$102 \quad Kd_{Zr} = (Kd_{Nd} * Kd_{Sm})^{0.5}$$



103 **Hydrous Partition Coefficients for slab melting:** As above, this study uses slightly  
 104 adapted partition coefficients for slab melting to those in Turner and Langmuir (2015b).  
 105 These lie within the limits of experimental values (figure SA1). The partition coefficients for  
 106 the LREE (La, Ce, Nd, and Sm) were reduced by 40% to reflect the hotter slab underneath  
 107 Don Casimiro-Maipo (inline with experimental high temperature partitioning results; figure  
 108 SA1). Additionally, the partition coefficients of Zr and Hf were reduced, and 70% of Cs  
 109 was assumed to be lost to the forearc.



**Figure SA1:** Slab melting partition coefficients lie within experimental values for 800-1000°C.

## 110 SIA4.2 Model End-Members

111 **"Ambient Andean Mantle" (Depleted Mantle End-Member):** The isotopic composition  
 112 of the depleted mantle end member was set to  $^{87}\text{Sr}/^{86}\text{Sr}=0.70355$  and  $^{143}\text{Nd}/^{144}\text{Nd}=0.512871$ ,  
 113 the isotopic composition of sample 126171 of Soager et al. (2013) following the isotopic  
 114 correction discussed above. This sample lies at the depleted end of the rear-arc array,  
 115 and thus reflects the composition of the Andean mantle prior to the addition of EM1  
 116 enrichment. For trace element modelling, the depleted mantle end-member is DMM of  
 117 Workman and Hart (2005).

118 **AOC:** The subducting Nazca plate has Pacific affinities. Trace element concentrations  
 119 were obtained from melting of N-MORB (Gale et al., 2013) at 876°C,  $F=0.3$ . The isotopic  
 120 composition of AOC was set to  $^{87}\text{Sr}/^{86}\text{Sr} = 0.7035$ ,  $^{143}\text{Nd}/^{144}\text{Nd} = 0.5130$ .

121 **Sediment:** The trace element composition of the sediment end-member was obtained from 33  
 122 samples of ODP1232 (Turner et al., 2017). To produce an envelope of model results for trace  
 123 elements and isotopes, one standard deviation of measured samples were added and subtracted  
 124 to the mean. Three outliers were discarded based on their anomalously high Sr isotope ratios.

125 **Enriched EM1 end-member:** Enrichment at Gough represents a mix of M-SCLM, and  
 126 ambient mantle. The isotopic signature of recycled M-SCLM depends on the storage time  
 127 (Turner et al., 2017). As several samples within this study have signatures more enriched  
 128 than Gough, we use the isotopic composition of a sample from Pitcairn (57DS9; Stracke et  
 129 al., 2003) as the EM1-like isotopic end member. This lies at the enriched end of the field

130 enclosing EM1-type ocean island basalts. The trace element concentration was obtained  
 131 from the inversion of Turner et al. (2017), with a few changes:

132 1) the bulk Eu partition coefficient was set at:

$$133 \quad Kd_{Eu} = (Kd_{Sm} * Kd_{Gd})^{0.5}$$

134 correcting an error in Turner et al. (2017) where it was set at 0.

135 2) The Zr concentration in Gough primary melts was adjusted to the standard EM1 value  
 136 Zr/Hf = 45.

137 3) Zr partition coefficients were extrapolated from Salters and Longhi (1999).

138 4) Nb and Ta partition coefficients were taken as  $Kd_{La}/3$ , based on ratios from Workman and  
 139 Hart (2005).

**Table SA9:** Gough mantle source composition used in this study compared to that of Turner et al. (2017; T2017).

	Eu	Zr	Nb	Ta
This study	0.130	6.70	0.403	0.024
T2017	0.0234	11.47	0.472	0.028

## 140 SIA5 References

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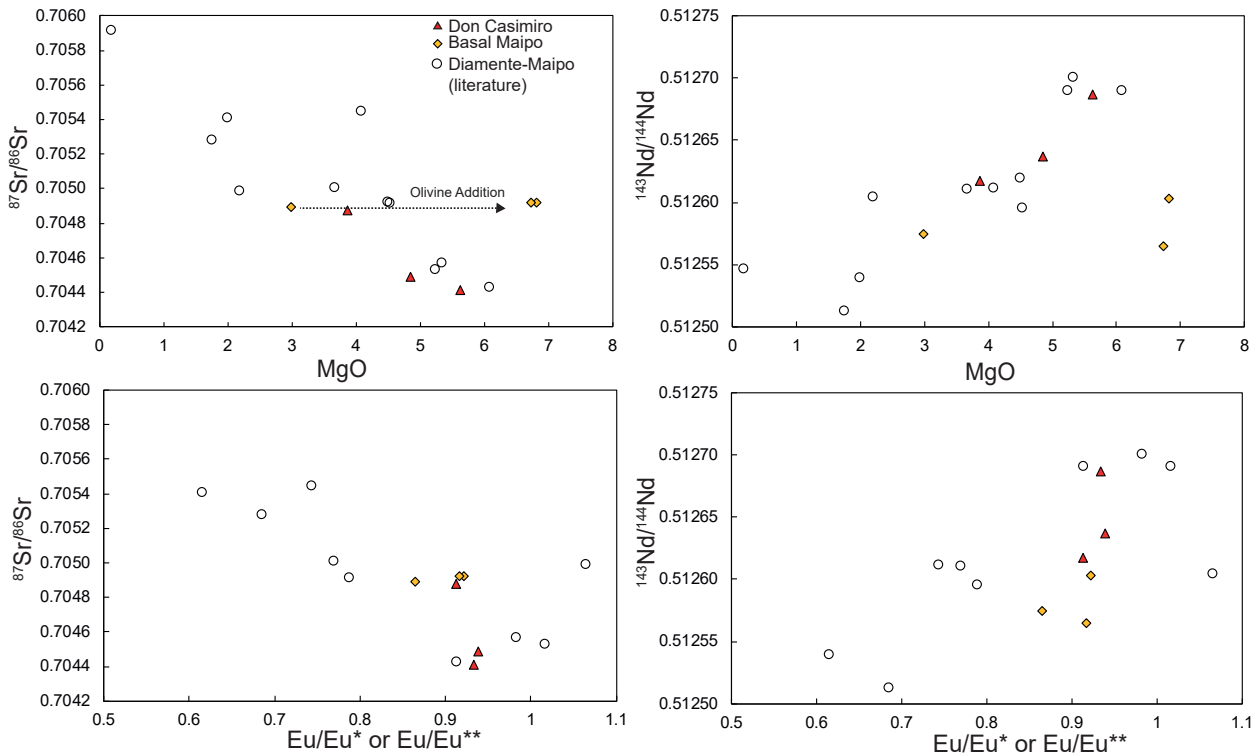
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## Appendix 2

### SIB1 Evidence for AFC processing

As mentioned briefly in the main text, when samples from the entire Diamante-Maipo caldera complex are considered, the correlations between  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  and indices for fractionation and assimilation (MgO and  $\text{Eu}/\text{Eu}^*$ ) provide strong evidence for AFC (figure SB1)

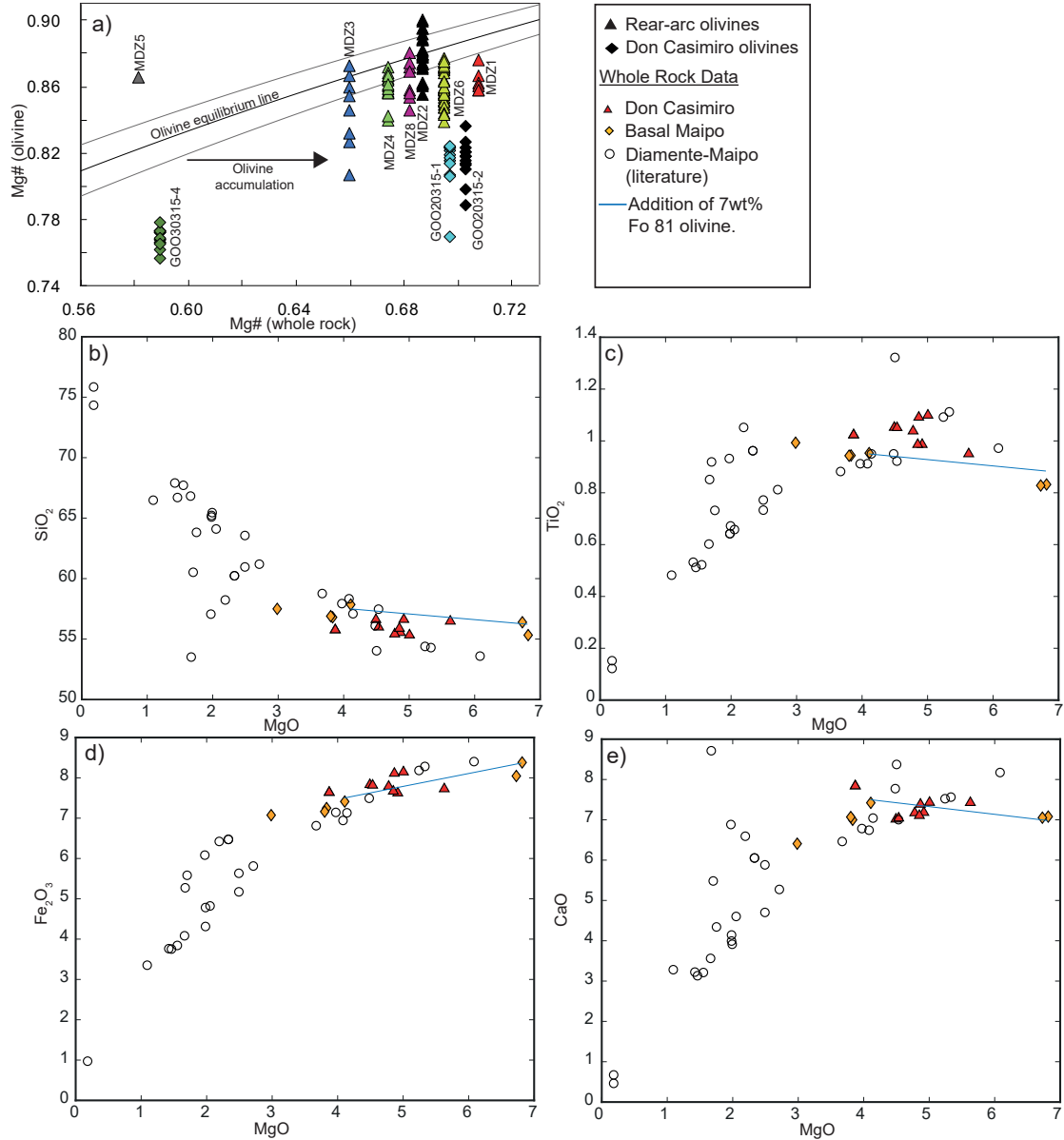


**Figure SB1:** Within the Diamante-Maipo Caldera, isotopic enrichment increases with assimilation and fractionation. The signal of olivine addition for the two most primitive Maipo samples is also clear.

### SIB2 Evidence for Olivine Addition

EMPA analysis indicates that Don Casimiro-Maipo olivines lie significantly below the olivine-whole rock equilibrium line (Figure SB2a), indicating excess olivine accumulation. The two most primitive samples, which are from the basal section of Maipo volcano, can be related to the otherwise coherent liquid line of descent among the Diamante Caldera samples by the addition of 7 wt. % Fo 81 olivine (figure SB2b-d.). All samples have phenocrysts of olivine and clinopyroxene, so this trend is not consistent with fractionation of these two phases together. Olivine may have been accumulated from a mush pile produced by fractionation before clinopyroxene appears on the liquidus. This serves as an excellent example of the importance of utilizing individual mineral analyses in conjunction with whole-rock data when interpreting subtle geochemical variations within individual volcanic groups. While it is tempting to represent the parental composition of a volcanic group using the most

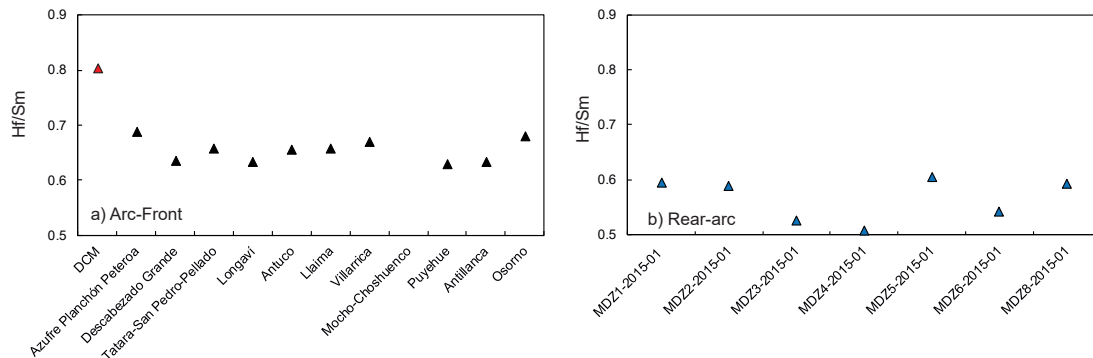
olivine-phyric samples (6-7vol% olivine, in this case), such samples can often have more complicated petrogenetic histories.



**Figure SB2:** a) Don Casimiro-Maipo samples show strong evidence for olivine addition, with olivine compositions plotting below an equilibrium line calculated using Roeder and Emslie (1970),  $K_d=0.3\pm 0.03$ , assuming whole-rock  $Fe^{3+}/Fe_T=0.3$  (Holm et al., 2016). Most rear-arc olivines plot near to or below the equilibrium line, suggesting minor olivine addition (assuming whole-rock  $Fe^{3+}/Fe_T=0.15$ , Søger et al., 2015b). b-d) Addition of 7 wt% Fo 81 olivine from more evolved Maipo samples produces a good match to the most primitive samples, showing that their high MgO contents are spurious. Diamante-Maipo whole-rock data literature data used in this study is from Hickey et al. (1986), Futa and Stern, (1988), Sruoga et al. (2005), Holm et al. (2014) and Hickey et al. (2016).

### SIB3 Changes in Slab Conditions Inferred from Zr-Hf Anomalies

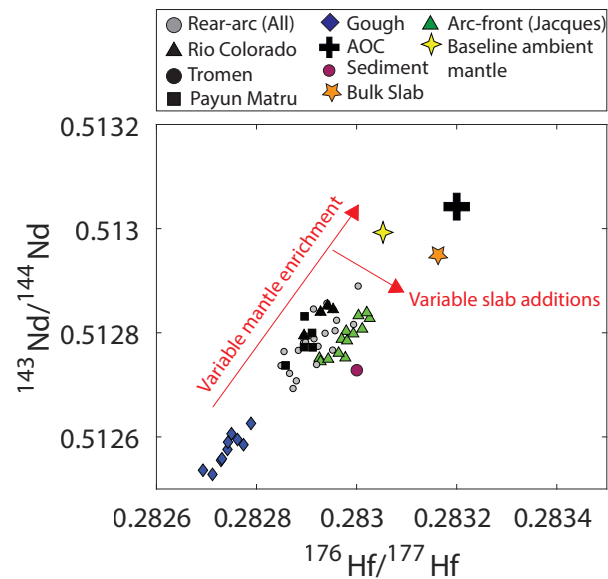
Using Hf/Sm as a proxy, we find that rear-arc and most arc-front volcanics have a similar range of Zr-Hf values, but Don Casimiro-Maipo has a less pronounced Zr-Hf anomaly (figure SB3). The smaller Hf/Sm (and thus Zr-Hf anomaly) at Don Casimiro-Maipo may reflect the fact that the slab is slightly hotter, which causes residual slab zircon to become more soluble in the slab melt (Hirai et al., 2018).



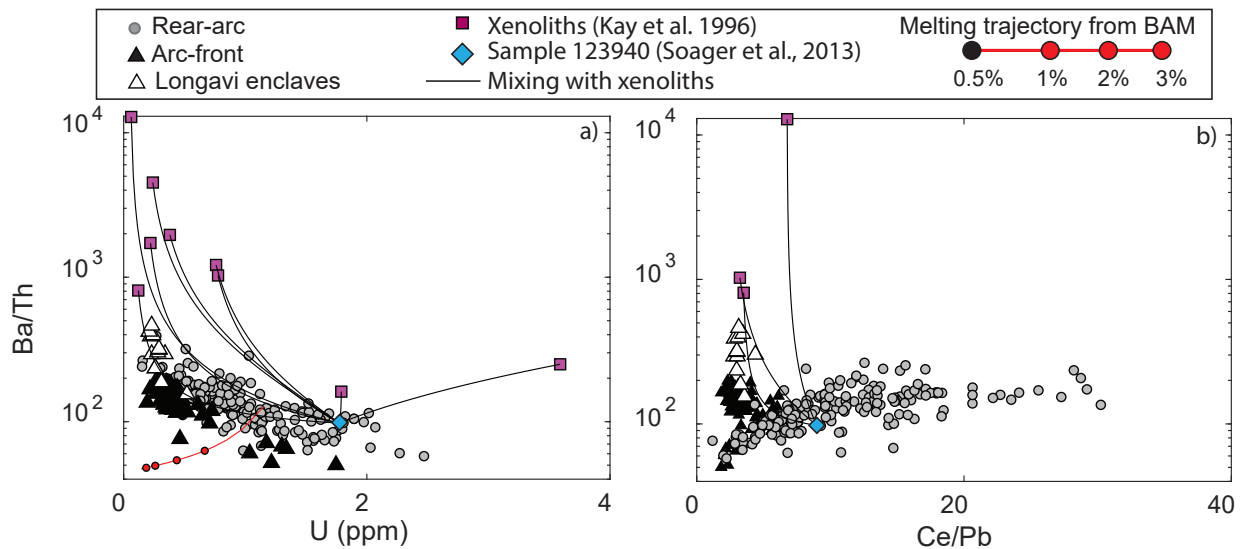
**Figure SB3:** a) Don Casimiro-Maipo has a significantly different Hf/Sm signature to the rest of the SVZ (data for other centers from Turner et al., 2016). b) Rear-arc centers of this study have similar Hf/Sm ratios to the SVZ arc-front as a whole.

### SIB4 Assessing the evidence for LCC assimilation in the SVZ rear-arc

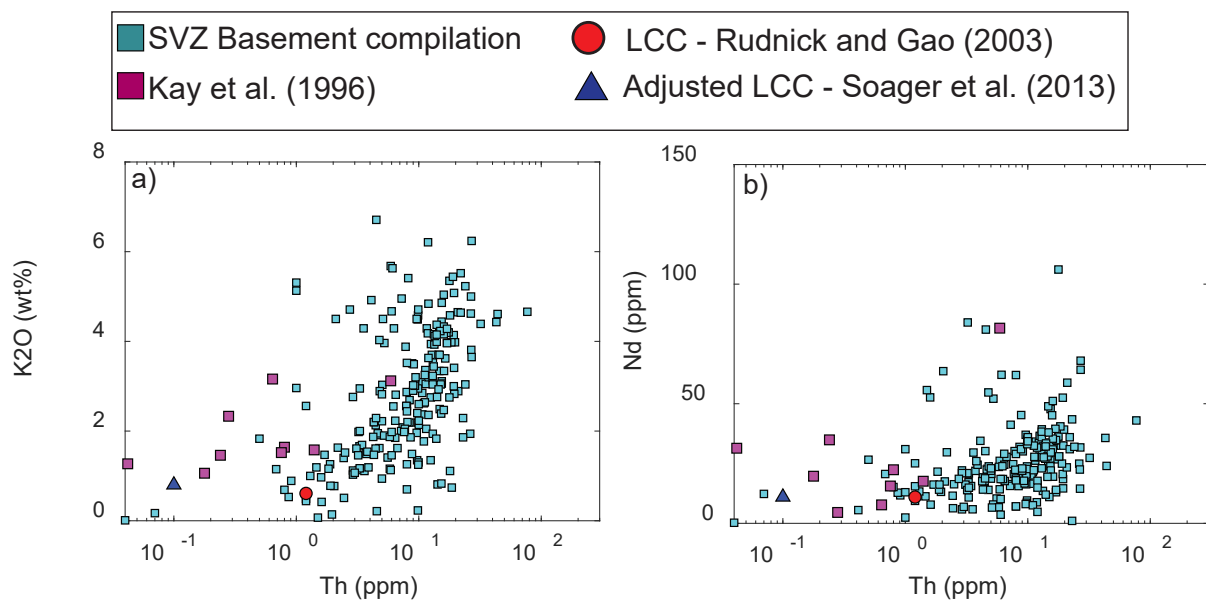
As discussed in section 4.5 of the main text, the offset trajectories of rear-arc and arc-front lavas in trace element and isotope space have been explained by various regional models. We show that these offsets can instead be accounted for a model incorporating variable amounts of EM1 enrichment, combined with variable slab fluxes and melt extents (figure 13 of the main text). This three-parameter model can also account for the offset trajectories in Nd-Hf isotope space (figure SB4). Arc-front and rear-arc lavas both spread along a mixing line between depleted and enriched mantle sources. Arc-front lavas are shifted to the right of the trajectory defined by rear-arc samples due to higher slab fluxes. Additionally, while rear-arc lavas appear to trend towards the composition of LCC xenoliths (Kay et al., 1996) in Ba/Th vs U or Ce/Pb space, calculated mixing lines are hyperbolic, and do not reproduce the trajectory of the rear-arc array (Figure SB5). To accommodate the mis-match between the LCC xenolith mixing model and the rear-arc data, Soager et al. (2013) instead adapt a model LCC composition that has been adjusted for several elements. These adjustments, however, diverge substantially from both global models of the LCC composition as well as the compositional field defined by the larger compilation of Andean basement lithologies (figure SB6).



**Figure SB4:** The offset trends of rear-arc and arc-front lavas in Hf-Nd isotope space can be accounted for a unified model with variable EM1-type enrichment combined with variable slab inputs (see Turner et al., 2017).



**Figure SB5:** Soager et al. (2013) suggest that the trajectory of rear-arc lavas trend towards the composition of xenoliths sampled by Kay et al. (1996) in  $\text{Ba}/\text{Th}$  vs.  $\text{U}$  and  $\text{Ce}/\text{Pb}$  space. However, calculated mixing trajectories between sample 123940 (used in the modelling of Soager et al., 2013) and these xenoliths fail to recreate the rear-arc trends on both plots.



**Figure SB6:** Comparison of SVZ basement compositions (cyan squares) to the composition of the xenoliths of Kay et al., (1996), the LCC (Rudnick and Gao, 2003), and the basement composition used in the modelling of Soager et al. (2013). It is clear that neither the xenoliths nor the LCC composition of Rudnick and Gao (2003) is particularly representative of the composition of SVZ basement lithologies. Given that the LCC composition of Rudnick and Gao (2003) already deviates substantially from observed SVZ basement lithologies, the adjustments made to the global LCC composition by Soager et al. (2013) are difficult to justify.

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