

1            **Intermediate-magnitude postseismic slip follows**  
2            **intermediate-magnitude ( $M$  4 to 5) earthquakes in**  
3            **California**

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8            **Key Points:**

- 9            • We examine postseismic slip following M4-5 earthquakes using borehole strain data  
10           • The median estimated postseismic moment is 0.4 times the coseismic moment  
11           • The postseismic moment ratio is intermediate between estimates for small and large  
12           earthquakes

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13 **Abstract**

14 The magnitude of postseismic slip is useful for constraining physical models of fault  
 15 slip. Here we examine the postseismic slip following intermediate-magnitude ( $M$  4 to 5)  
 16 earthquakes by systematically analyzing data from borehole strainmeters in central and  
 17 northern California. We assess the noise in the data and identify 36 records from 12 earth-  
 18 quakes that can be interpreted. We estimate postseismic to coseismic moment ratios by  
 19 comparing the coseismic strain changes with strain changes induced by afterslip in the  
 20 following 1.5 days. The median estimated postseismic moment is 0.36 times the coseis-  
 21 mic moment, with a 90% confidence interval between 0.22 and 0.54. This postseismic  
 22 moment is slightly larger than typically observed following large ( $M > 6$ ) earthquakes  
 23 but smaller than observed following small ( $M$  2 to 4) earthquakes. The intermediate-  
 24 magnitude postseismic slip suggests a size dependence in the dynamics of earthquakes  
 25 or in the properties of fault areas that surround earthquakes.

26 **1 Introduction**

27 Deformation in the hours to years following earthquakes accumulates via a range  
 28 of processes, including afterslip, poroelastic flow, and viscoelastic deformation. After-  
 29 slip is usually the largest cause of deformation in the first few hours to months. The af-  
 30 terslip that accumulates following large ( $M > 6$ ) earthquakes typically has moment equal  
 31 to 10 to 30% of the coseismic moment (e.g., Cetin et al., 2012; D’Agostino, Cheloni, Fornaro,  
 32 Giuliani, & Reale, 2012; Donnellan & Lyzenga, 1998; Gahalaut et al., 2008; Gonzalez-  
 33 Ortega et al., 2014; Johanson & Bürgmann, 2010; Lin et al., 2013; Segall, Bürgmann,  
 34 & Matthews, 2000), though afterslip moments for individual earthquakes range from a  
 35 few percent to several hundred percent of the coseismic moment (see Figure 4, Bürgmann  
 36 et al. (2001); Dogan et al. (2014); Freed (2007); Langbein, Murray, and Snyder (2006);  
 37 Paul, Lowry, Bilham, Sen, and Smalley (2007)).

38 While afterslip moments vary, the postseismic to coseismic moment ratios estimated  
 39 for large earthquakes show no systematic trend with magnitude (see Lin et al. (2013),  
 40 Fattahi, Amelung, Chaussard, and Wdowinski (2015), and Figure 4 for summaries). The  
 41 inferred magnitude-independent afterslip is consistent with a self-similar model of earth-  
 42 quakes, where large events are scaled versions of smaller ones. However, large afterslip  
 43 following small earthquakes has been proposed to explain the long recurrence intervals  
 44 of small repeating earthquakes (Chen & Lapusta, 2009), and analysis of small ( $M$  1.9  
 45 to 3.5) earthquakes near San Juan Bautista, CA revealed that those small earthquakes  
 46 had afterslip with moment roughly equal to the coseismic moment, on average (Hawthorne,  
 47 Simons, & Ampuero, 2016). Those large afterslip moments could simply indicate that  
 48 the frictional properties of the San Andreas Fault near San Juan Bautista region are un-  
 49 usual, and more prone to large afterslip. Afterslip with moment 1.5 to 3 times the co-  
 50 seismic moment was identified following several larger earthquakes in the area: the 2004  
 51  $M$  6 Parkfield earthquake (Barbot, Fialko, & Bock, 2009; Freed, 2007; Langbein et al.,  
 52 2006), the 2007  $M$  5.4 Alum Rock earthquake (Murray-Moraleda & Simpson, 2009), and  
 53 the 1998  $M$  5.1 San Juan Bautista earthquake (Taira, Bürgmann, Nadeau, & Dreger,  
 54 2014).

55 But the large afterslip moments identified following  $M < 3.5$  earthquakes could  
 56 also indicate that small earthquakes generally behave differently—that the self-similar  
 57 scaling of postseismic slip breaks down as earthquakes become smaller. For example, the  
 58 afterslip moment could change as earthquake rupture extents become smaller than the  
 59 seismogenic zone width, so that most afterslip occurs within the seismogenic zone, rather  
 60 than above and below it. Or the afterslip moment could change as earthquake rupture  
 61 extents become similar to the minimum earthquake nucleation size, and thus become too  
 62 small to drive more rapid slip (Chen & Lapusta, 2009).

In this study, we further examine how afterslip moment varies with earthquake size by examining intermediate-magnitude ( $M$  4 to 5) earthquakes. Only a few afterslip moments have been estimated for  $M$  4 to 6 earthquakes, and all reported values are large. Afterslip moments between 1 and 5 times the coseismic moment were observed following  $M$  4.7 to 5.5 earthquakes on the Chaman fault (Furuya & Satyabala, 2008), on the Ghazaband fault (Fattahi et al., 2015), near Alum Rock, CA (Murray-Moraleda & Simpson, 2009), and near Mogul, NV (Bell, Amelung, & Henry, 2012). However, these large afterslip moments could reflect a reporting bias, as smaller amounts of afterslip would be difficult to observe.

Here we use high-precision borehole strain data to examine postseismic slip following  $M$  4 to 5 earthquakes in central and northern California. We systematically identify earthquakes with resolvable coseismic deformation at each of 12 strainmeters in the PBO and USGS networks. We are able to assess the afterslip moment with reasonable accuracy for 12 earthquakes. The median afterslip moment is roughly 0.4 times the coseismic moment, between the values obtained for smaller and larger earthquakes.

## 2 Available Data and Earthquakes

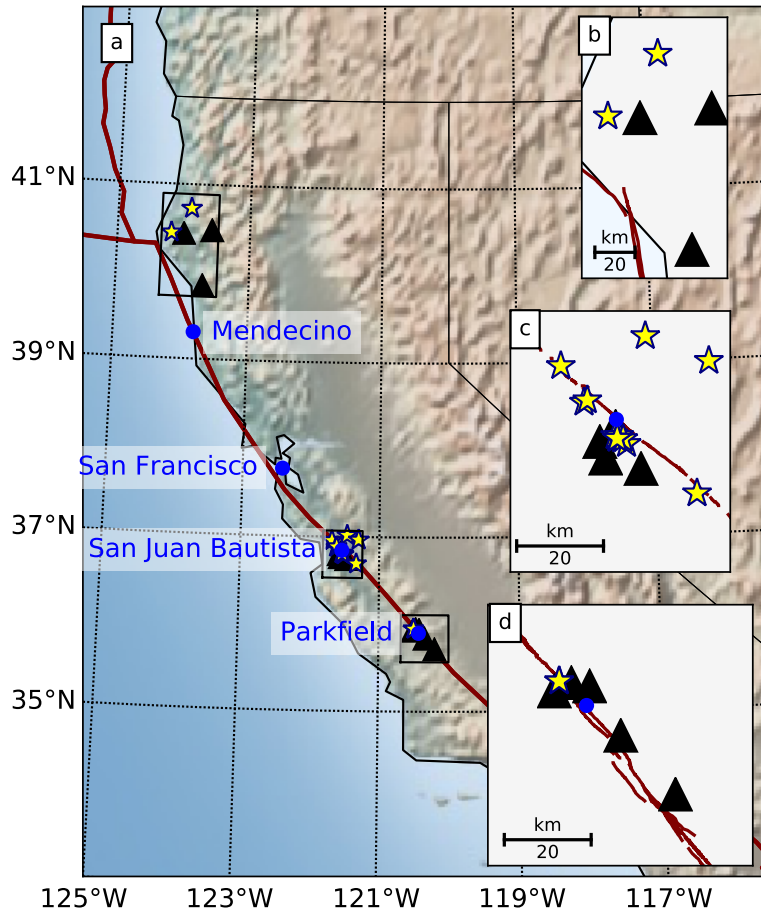
We examine data from 12 strainmeters located along the San Andreas Fault in California, shown in Figure 1. More than half of the high-quality earthquake records will come from strainmeter SJT, which was installed by the USGS in 1983 at the northern end of the central creeping section of the San Andreas Fault Gladwin, Gwyther, Hart, Francis, and Johnston (1987). The remaining records come from strainmeters that were installed by UNAVCO as part of the Plate Boundary Observatory (PBO). We consider data from 11 PBO strainmeters installed between 2006 and 2008. Strainmeters B073, B075, B076, B078, and B079 are located near Parkfield, at the southern edge of the central creeping section, while strainmeters B058, B065, and B067 are located near San Juan Bautista, at the northern edge of the central creeping section, and strainmeters B045, B934, and B935 are located close to the Mendocino triple junction, near another creeping section of the San Andreas.

Small earthquakes occur frequently along these creeping sections, mostly at depths of 4 to 15 km (e.g., Irwin & Barnes, 1975; Waldhauser & Schaff, 2008). We begin our analysis by identifying all  $M$  4 to 6 earthquakes that occurred within 30 km of the strainmeters while the strainmeters were operating, as recorded in the NCSN catalog. This identification recovers 112 potential earthquake-station pairs, or 112 potential earthquake records. But we will find that only 14 records, which cover 12 unique earthquakes, have low enough noise level that we can usefully assess the magnitude of afterslip.

## 3 Initial Data Processing

The deformation produced by co- and postseismic slip are recorded at the strainmeters via three to four horizontal extensometers, which are located at depths of 150 to 250 m and measure changes in the horizontal borehole width at various azimuths. SJT records deformation at 18-minute intervals, and we use 10-minute data from the PBO strainmeters. We convert the time-varying extensometer measurements to the three time-varying horizontal components of strain  $\varepsilon_{E+N}$ ,  $\varepsilon_{E-N}$ , and  $\varepsilon_{2EN}$  using the tidal calibrations derived by J. Langbein for strainmeter SJT and by Hodgkinson, Langbein, Henderson, Mencin, and Borsa (2013) for the PBO strainmeters.

We will directly analyze the time series of  $\varepsilon_{E+N}$ ,  $\varepsilon_{E-N}$ , and  $\varepsilon_{2EN}$  recorded at strainmeter SJT. But for the PBO strainmeters, we analyze different linear combinations of these strain components. All the strain components are recorded with high instrumental precision, less than 1 nanostrain. But the various components of strain have different sensitivity to atmospheric and hydrologic noise. So we follow the approach of Hawthorne,



**Figure 1.** (a) Map of the central San Andreas Fault system in California. In all plots, black triangles are strainmeter locations, yellow stars are the earthquakes for which we estimate post-seismic signals, blue circles are cities, and the red lines mark the plate boundary. (b), (c), and (d) Maps of the three clusters of strainmeters. The strainmeters in each panel are, from west to east, (b) B045, B934, and B935; (c) B058, B067, SJT, and B065; (d) B075, B073, B076, B078, and B079. Note that the marker for SJT is behind the central group of earthquakes in panel c.

112 Bostock, Royer, and Thomas (2016) to identify linear combinations of  $\varepsilon_{E+N}$ ,  $\varepsilon_{E-N}$ , and  
 113  $\varepsilon_{2EN}$  that are normally less noisy: the components that have minimal response to at-  
 114 mospheric pressure variations. We refer to the linear combinations that are closest to  
 115 the original components as  $\varepsilon_{E+N-na}$ ,  $\varepsilon_{E-N-na}$ , and  $\varepsilon_{2EN-na}$ , and we will analyze all  
 116 three time series at the PBO strainmeters, though we note that only two of these three  
 117 components are independent. The third component is a linear combination of the first  
 118 two.

119 Having isolated the strain components of interest, we estimate and remove several  
 120 non-tectonic signals from each time series, following, for example, Hart, Gladwin, Gwyther,  
 121 Agnew, and Wyatt (1996); Langbein (2010); Roeloffs (2010), and Hawthorne, Bostock,  
 122 et al. (2016). To remove borehole curing signals, we discard the first 18 months of data  
 123 at each station and then fit and remove a linear trend and a decaying exponential with  
 124 time constant around 1 year. Then we estimate and remove shorter-timescale nontec-  
 125 tonic variations. We compare the tidal model of Cartwright and Edden (1973) with the  
 126 data and identify tidal frequencies that are likely to have tidal signals with amplitudes  
 127 of at least 0.5 times the noise level. We estimate best-fitting sinusoids at those frequen-  
 128 cies and remove them. At SJT, we also estimate and remove a linear response to atmo-  
 129 spheric pressure and a periodic 3-hour signal that is likely instrumental noise, as iden-  
 130 tified by Hawthorne, Simons, and Ampuero (2016).

#### 131 4 Interpreting Example Strain Records

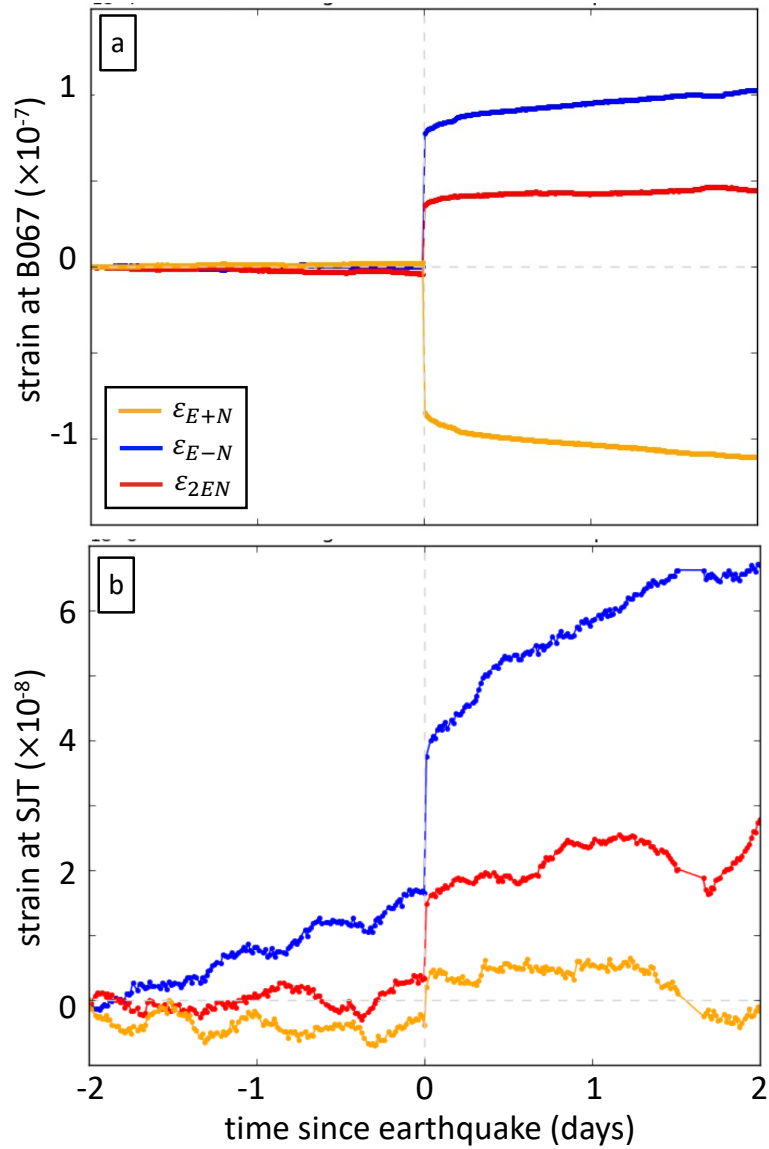
132 After removing these non-tectonic signals, we can analyze the earthquake- and afterslip-  
 133 induced strain. Several examples of the co- and postseismic strain are shown in Figures 2  
 134 and S1-S18. Figure 2a shows a high quality record of a  $M$  4.2 earthquake located about  
 135 5 km NE of strainmeter B067. An abrupt coseismic strain step is followed by the grad-  
 136 ual accumulation of postseismic strain over the two days shown. Figure 2b shows a sim-  
 137 ilar record of a  $M$  4.0 earthquake located about 5 km SE of strainmeter SJT, but here  
 138 the signal to noise ratio is lower.

139 In our modelling, we will assume that the postseismic strain is created by after-  
 140 slip. If afterslip occurs in an area within 1 to 2 earthquake radii of the earthquake rup-  
 141 ture, as is usually observed following large earthquakes (D'Agostino et al., 2012; Ryder,  
 142 Parsons, Wright, & Funning, 2007), and if the earthquake radius is small relative to the  
 143 earthquake-strainmeter distance, then the co- and postseismic slip should appear co-located  
 144 from the perspective of the strain observations. In other words, the co- and postseismic  
 145 slip should have approximately the same Green's functions. Such similar Green's func-  
 146 tions are consistent with the data in Figure 2. The ratio of postseismic to coseismic strain  
 147 is similar on the three strain components, as expected if the strains are given by mul-  
 148 tiplying the co- and postseismic moments by the same Green's functions.

149 In this study, we are interested in the ratio of the postseismic moment to the co-  
 150 seismic moment. If the Green's functions are the same, the moment ratio can be obtained  
 151 simply by dividing the postseismic strain by the coseismic strain. In Figure 2a, for in-  
 152 stance, we may note that the postseismic strains accumulated within 2 days of the earth-  
 153 quake have magnitude about 20% of the amplitude of the coseismic strains. Such strain  
 154 ratios suggest that the afterslip moment accumulated within 2 days of the earthquake  
 155 is equal to about 20% of the coseismic moment.

#### 156 5 Identifying Interpretable Strain Records

157 Our search for earthquakes within 30 km of the strainmeters revealed 112 poten-  
 158 tial earthquake records. But most of these are not interpretable. First, many are too noisy.  
 159 In our initial culling of the data, we visually examine the 112 earthquake records and  
 160 retain only those where the coseismic offset is well resolved on at least one component.



**Figure 2.** Illustrative records of co- and postseismic strain for (a) a  $M$  4.2 earthquake on 20-Nov-2015, recorded at B067 and located 5 km NE of the strainmeter at 6 km depth and (b) a  $M$  4.0 earthquake on 11-Feb-2001, recorded at SJT and located 5 km SE of the strainmeter at 6 km depth.

161 This selection does not bias our analysis toward small postseismic to coseismic moment  
 162 ratios because the noise in the strain data has a random walk character (Hawthorne &  
 163 Rubin, 2013; Langbein & Johnson, 1997). The uncertainty in the postseismic strain ac-  
 164 cumulation, which occurs over a few days, is larger than the uncertainty in the coseis-  
 165 mic strain accumulation. For the postseismic strain to be resolvable when the coseismic  
 166 strain is not, it would have to be about 10 times larger than the coseismic strain.

167 We do not analyze all records with well-resolved coseismic offsets, however. Some  
 168 are too complex to interpret because they occur within earthquake clusters or because  
 169 they trigger creep events. Creep events on the shallow San Andreas Fault are mostly hours-  
 170 to days-long intervals when part of the fault accelerates to rates of order mm to cm per  
 171 hour (Bilham, Suszek, & Pinkney, 2004; Gladwin, Gwyther, Hart, & Breckenbridge, 1994;  
 172 S. Schulz, Burford, & Mavko, 1983; S. S. Schulz, 1989; Thurber & Sessions, 1998). In  
 173 principle, creep events triggered by earthquakes can be classified as afterslip (e.g., Floyd  
 174 et al., 2016; Fukuda, Johnson, Larson, & Miyazaki, 2009; Langbein et al., 2006). We choose  
 175 to exclude creep events from our analysis for two reasons. First, it is unclear whether  
 176 the fault zone processes that create triggered creep events are the same as the processes  
 177 that usually create afterslip. The spontaneous occurrence of creep events suggest a par-  
 178 tially slip rate-weakening rheology that drives acceleration (e.g., Belardinelli, 1997; Wei,  
 179 Kaneko, Liu, & McGuire, 2013) while afterslip is often modeled with a slip rate-strengthening  
 180 rheology, so that increased slip rates are driven exclusively by the imposed coseismic stress  
 181 (Helmstetter & Shaw, 2009; Marone, Scholtz, & Bilham, 1991; Perfettini & Avouac, 2004).

182 Second, and more importantly, we exclude creep events because it is difficult to es-  
 183 timate their moments. As noted above, we can estimate the relative moment of after-  
 184 slip that is located close to the coseismic slip simply by computing the ratio of the post-  
 185 seismic to coseismic strain. But to estimate the relative moment of the triggered creep  
 186 events, we would need to account for the difference in Green’s functions between the co-  
 187 seismic rupture and the creep event slip. And we do not know the creep event locations  
 188 or the coseismic Green’s functions well enough to account for that difference. So we iden-  
 189 tify creep events by (1) looking at the nearby surface USGS creepmeter records and (2)  
 190 examining how the postseismic to coseismic strain ratio varies among the different com-  
 191 ponents of strain. The earthquakes excluded because of noise and nearby creep events  
 192 are listed in table S1.

## 193 **6 Estimates of Postseismic Strain**

194 After identifying earthquakes with well-resolved coseismic steps and excluding those  
 195 with visible creep events, we are left with 18 records, or 18 earthquake-station pairs, which  
 196 cover 14 independent earthquakes. We will estimate the coseismic and postseismic strain  
 197 of these earthquakes within a 1.5-day period. But first, we remove a linear trend that  
 198 may represent seasonal or hydrological variations. We estimate the trend from the strain  
 199 observations made in the two days before the earthquakes: by dividing the change in strain  
 200 between 2 days and 5 hours before the earthquake by that time interval (43 hours). We  
 201 extrapolate the estimated “long-term” strain rate through the co- and postseismic pe-  
 202 riod and subtract it from the strain time series. Note that the offset-over-time approach  
 203 to determining offsets and trends is appropriate for random walk noise, as appears to  
 204 characterize the strain data (Hawthorne & Rubin, 2013; Langbein & Johnson, 1997).

205 After removing the long-term trend, we estimate the coseismic and postseismic strain  
 206 changes for each of the three strain components. The coseismic strain change is defined  
 207 as the strain accumulated within the 40-minute interval centered on the earthquake time.  
 208 This 40-minute coseismic interval allows us to identify the entire coseismic strain step  
 209 from the 10- to 18-minute data. We interpolate between data points as necessary. The  
 210 postseismic strain change is defined as the strain that accumulates between 20 minutes  
 211 and 1.5 days after the earthquake. The 1.5-day interval is chosen as a compromise be-

212 tween signal and noise. It is long enough to allow significant strain to accumulate but  
 213 is short enough to avoid the large atmospheric and hydrological noise that can accumu-  
 214 late when longer time intervals are considered.

215 As noted earlier, we take the ratio of the postseismic to coseismic strains on each  
 216 component as an estimate of ratio of the postseismic to coseismic moments. To assess  
 217 how well resolved each postseismic moment ratio is, we examine how it would change  
 218 if noise were added. We randomly pick 3000 4-day-long intervals of the strain time se-  
 219 ries to use as 3000 realizations of the noise. We add each realization to the strain data  
 220 of interest, recompute and subtract a long-term trend, extract the coseismic and post-  
 221 seismic strain changes, and compute their ratio. With this approach, we obtain a prob-  
 222 ability distribution for the postseismic to coseismic moment ratios for each earthquake  
 223 and component. We identify and will examine the ratios from the earthquakes and com-  
 224 ponents that have reasonably low uncertainty ranges: those where 70% of the estimated  
 225 ratios fall within a range smaller than 2 (e.g., between 0 and 2, or between 0.5 and 2.5).  
 226 These ratios and their 70% confidence bounds are listed in table S2 and are plotted in  
 227 Figure 3a.

228 Most of the estimated postseismic to coseismic moment ratios cluster between 0  
 229 and 1. 64% of them are between 0.2 and 0.6. Note that two earthquakes, a  $M$  4.9 and  
 230 a  $M$  4.25, appear to have well-resolved but negative moment ratios. Such negative post-  
 231 seismic moment ratios would seem to imply that the fault is slipping backwards. How-  
 232 ever, the signal from the  $M$  4.9 may simply result from hydrological noise; there was sig-  
 233 nificant rainfall in January 1993, when the earthquake occurred (Figures S3 and S19).  
 234 The signal from the  $M$  4.25 is well resolved (Figure S10). But its postseismic to coseis-  
 235 mic moment ratios are different on the different components, suggesting that the post-  
 236 seismic strain may result from a triggered creep event, or from slip occurring in a dif-  
 237 ferent location than the coseismic slip. We choose to keep these two earthquakes in our  
 238 analysis because we had not excluded them before estimating the collection of moment  
 239 ratios.

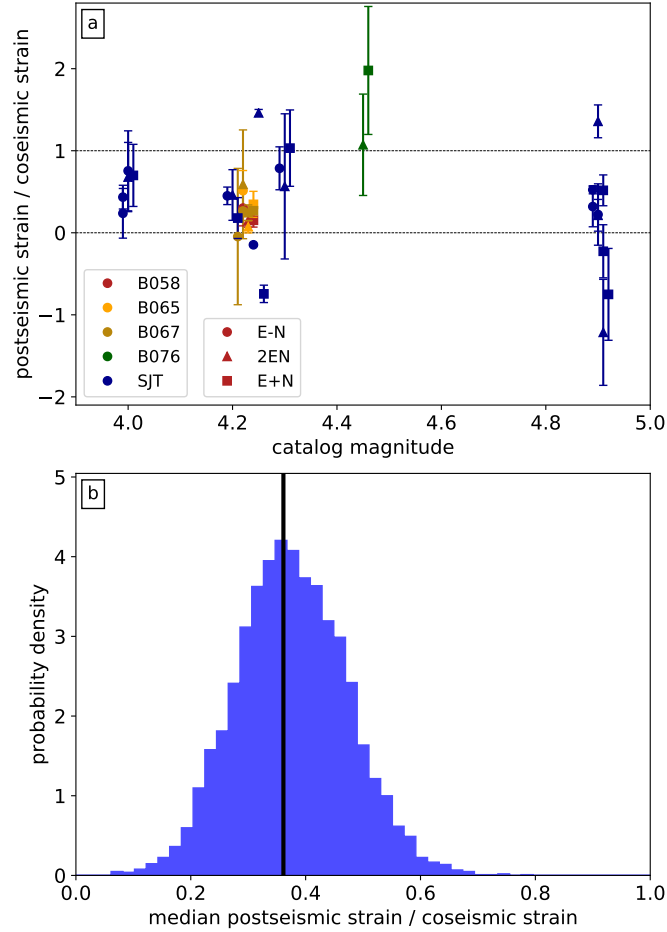
240 Our goal here is not understand individual events, but to determine the typical af-  
 241 terslip moments of the available  $M$  4 to 5 earthquakes. The median of the estimated post-  
 242 seismic to coseismic moment ratios is 0.31. But it may be more appropriate to take the  
 243 median over earthquakes, to avoid overweighting a few events that have more than one  
 244 observation. So we group the estimated moment ratios by earthquake, take the median  
 245 for each group, and then take the median among the 12 unique earthquakes. The me-  
 246 dian earthquake-grouped moment ratio is 0.36.

247 To determine the uncertainty on the median moment ratio, we randomly pick sets  
 248 of values from the probability distributions for each earthquake and component, which  
 249 were obtained above by adding various realizations of the noise. We pick 4000 sets of ra-  
 250 tios and estimate their earthquake-grouped medians. Figure 3b shows a histogram of these  
 251 medians, which represents the probability distribution of the median postseismic to co-  
 252 seismic moment ratio. The distribution implies that the median postseismic ratio is be-  
 253 tween 0.28 and 0.48 with 70% probability and between 0.22 and 0.54 with 90% proba-  
 254 bility.

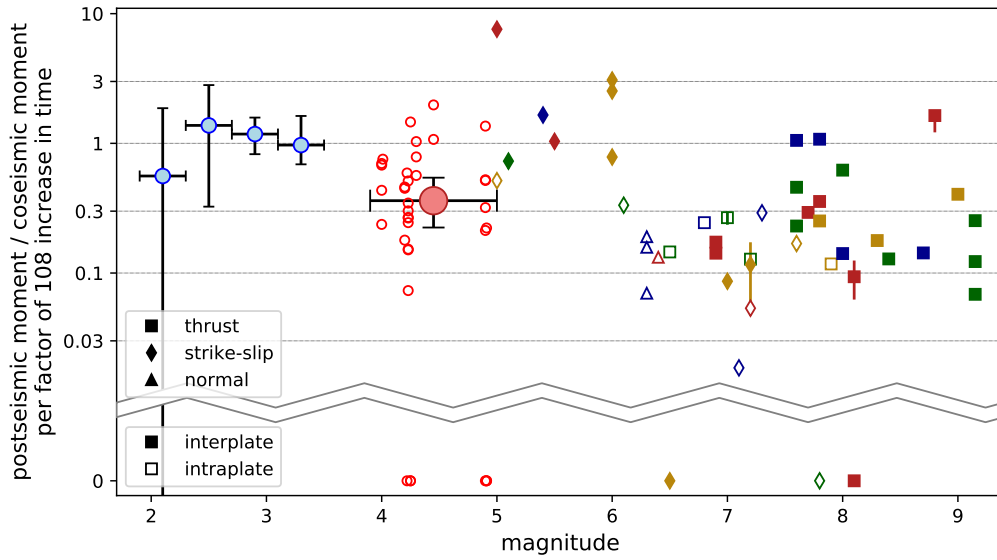
## 255 7 Discussion and Conclusion

256 The postseismic to coseismic moment ratios estimated here are plotted along with  
 257 the ratios inferred for smaller and larger earthquakes in Figures 4 and S20. The post-  
 258 seismic moments for smaller ( $M < 3.5$ ) earthquakes were obtained over the same time  
 259 interval considered here: from 20 minutes to 1.5 days after the earthquakes (Hawthorne,  
 260 Simons, & Ampuero, 2016). The postseismic observations of larger earthquakes were made  
 261 over a range of timescales, from days to years after the earthquakes. Postseismic moment





**Figure 3.** (a) Observed ratios of the postseismic strain changes, from 20 minutes to 1.5 days after the earthquakes, to the coseismic strain changes, from 20 minutes before to 20 minutes after the earthquakes. Each measurement comes from one component at one strainmeter, as indicated by color and symbol. Error bars indicate 70% uncertainty ranges. Note that we randomly shift the magnitudes by up to 0.015 to avoid plotting points on top of each other. (b) Vertical black line: median postseismic to coseismic strain ratio, obtained by taking the median over earthquakes. Probability distribution of the median postseismic to coseismic strain ratio, obtained by considering various realizations of the noise.



**Figure 4.** Time-normalized postseismic to coseismic moment ratios for earthquakes with a range of magnitudes. Note that the y-axis is on a log scale but that zero values are plotted below the break. The small red open circles are the individual measurements from this study, and the large red circle with error bars is the event-averaged median with 90% uncertainty ranges vertically and the range of magnitudes horizontally. The blue circles with error bars on the left are the moment ratios obtained by Hawthorne, Simons, and Ampuero (2016) for small earthquakes near San Juan Bautista, again with 90% uncertainty ranges vertically and the range of magnitudes horizontally. The points on the right come from a range of studies of intermediate and large-magnitude earthquakes, as listed in the text.

262 is often found to accumulate as the logarithm of time after the earthquake, so for a sim-  
 263 ple comparison in Figure 4, we normalize the larger earthquake observations to repre-  
 264 sent the moment accumulation expected for factor of 108 (=1.5 days / 15 minutes) in-  
 265 crease in time after the earthquake, assuming a logarithmic moment accumulation.

266 The postseismic to coseismic moment ratios we observe, with median 0.36, is slightly  
 267 larger than the moment typically seen after large ( $M > 6$ ) earthquakes. Most (though  
 268 not all) large earthquakes show postseismic moments smaller than 0.3 times the coseis-  
 269 mic moment (Amoruso & Crescentini, 2009; Barbot, Hamiel, & Fialko, 2008; Bürgmann  
 270 et al., 2001; Cetin et al., 2012; Cheloni et al., 2010; Chlieh et al., 2007; D’Agostino et  
 271 al., 2012; Diao, Wang, Wang, Xiong, & Walter, 2018; Dogan et al., 2014; Floyd et al.,  
 272 2016; Freed, 2007; Gonzalez-Ortega et al., 2014; Heki, Miyazaki, & Tsuji, 1997; Hobbs,  
 273 Kyriakopoulos, Newman, Protti, & Yao, 2017; Hsu et al., 2006; Jacobs, Sandwell, Fialko,  
 274 & Sichoix, 2002; Johanson & Bürgmann, 2010; Jónsson, 2008; Langbein et al., 2006; Lin  
 275 et al., 2013; Mahsas et al., 2008; Malservisi et al., 2015; Melbourne, Webb, Stock, & Reig-  
 276 ber, 2002; Miura, Suwa, Hasegawa, & Nishimura, 2004; Podgorski et al., 2007; Pritchard  
 277 & Simons, 2006; Rolandone et al., 2018; Ryder, Bürgmann, & Sun, 2010; Ryder et al.,  
 278 2007; Savage & Svarc, 1997; Segall et al., 2000; Shrivastava et al., 2016; Sreejith et al.,  
 279 2016; Subarya et al., 2006; Wen, Li, Xu, Ryder, & Bürgmann, 2012). Our  $M$  4 to 5 mo-  
 280 ment ratios are smaller than the reported moments for  $M$  5 to 6 earthquakes (Barbot  
 281 et al., 2009; Fattahi et al., 2015; Freed, 2007; Furuya & Satyabala, 2008; Langbein et al.,  
 282 2006; Murray-Moraleda & Simpson, 2009; Taira et al., 2014). But the high values for  $M$  5  
 283 to 6 earthquakes could result from observational bias; smaller postseismic moments may  
 284 not be reported because they would be harder to observe. More interestingly, then, we  
 285 note that our  $M$  4 to 5 moment ratios are also smaller than the roughly one to one ra-  
 286 tios observed for  $M < 3.5$  earthquakes Hawthorne, Simons, and Ampuero (2016).

287 There are several possible explanations for the observed variation in postseismic  
 288 moment with magnitude. First, the varying postseismic moments could reflect fault prop-  
 289 erties. Smaller earthquakes may be more likely to occur on creeping sections of faults,  
 290 perhaps on asperities surrounded by velocity-strengthening fault sections that are more  
 291 prone to large postseismic slip (e.g., Rolandone et al., 2018; Vaca, Vallée, Nocquet, Battaglia,  
 292 & Régnier, 2018). The postseismic moment estimates for  $M < 3.5$  earthquakes all come  
 293 from a single 20-km-wide fault segment near San Juan Bautista, CA, which could have  
 294 particular properties. But most of the earthquakes investigated here come from that same  
 295 fault segment, and half are obtained from measurements on the same strainmeter, SJT  
 296 (see Figures S1 to S10 for the time series).

297 It seems unlikely that other physical processes create some of the postseismic de-  
 298 formation we observe. Significant viscoelastic deformation is unlikely to accumulate on  
 299 the brief, 2-day timescale examined here (e.g., Bruhat, Barbot, & Avouac, 2011; John-  
 300 son, Bürgmann, & Freymueller, 2009; Pollitz, Banerjee, Burgmann, Hashimoto, & Choosakul,  
 301 2006). Poroelastic deformation can accumulate more quickly, but it typically has smaller  
 302 magnitude, just few percent of the coseismic deformation (Jónsson, Segall, Pedersen, &  
 303 Björnsson, 2003; Peltzer, Rosen, Rogez, & Hudnut, 1996, 1998) unless there is a nonlin-  
 304 ear near-surface response (e.g., Chia, Wang, Chiu, & Liu, 2001; Manga & Wang, 2007;  
 305 Quilty & Roeloffs, 1997; Wang, Wang, & Manga, 2004) or near-borehole deformation due  
 306 to shaking (Barbour, Agnew, & Wyatt, 2015), and Hawthorne, Simons, and Ampuero  
 307 (2016) identified no strong near-surface response to passing seismic waves or to creep events  
 308 in the San Juan Bautista region or at strainmeter SJT.

309 Assuming, then, that the postseismic deformation reflects afterslip, the magnitude-  
 310 dependent moment ratios could reflect the time intervals in which we observe that afterslip.  
 311 Postseismic moment is often observed to accumulate as log of the time  $t$  since  
 312 the earthquake, or at a rate of  $1/t$ . But at short times  $t$  after the earthquake, the mo-  
 313 ment rate may be slower than would be predicted by a  $1/t$  extrapolation, perhaps be-  
 314 cause the slipping region is growing outward from the coseismic rupture (Ariyoshi et al.,

2009; Dublanchet, Bernard, & Favreau, 2013a, 2013b; Lui & Lapusta, 2016; Perfettini & Ampuero, 2008) or because the fault takes time to accelerate in response to the coseismic stress increase (Marone et al., 1991; Montési, 2004; Perfettini & Avouac, 2004; Savage, 2007).

The coseismic rupture geometry can also influence the magnitude of postseismic slip. Small earthquakes tend to be more circular (e.g., Abercrombie, 1995; Gomberg, Wech, Creager, Obara, & Agnew, 2016; Scholz, 1982; Shaw, 2013), and thus may have a larger perimeter-to-area ratio and a larger region close to the coseismic rupture that can experience and respond to strong coseismic stress changes (Hawthorne, Simons, & Ampuero, 2016). However, the transition from circular to rectangular ruptures is typically inferred to occur when ruptures first start to span the seismogenic zone, at a magnitude around 6 or 7. We observe a change in postseismic moment at a magnitude of 4 to 5.

Alternatively, the magnitude-dependent postseismic moments could reflect a more fundamental property of earthquake dynamics. For instance, Chen and Lapusta (2009) identified large postseismic slip in rate and state friction models of earthquakes occurring on small asperities, on patches that were not much wider than the earthquake nucleation size. The large afterslip arose because portions of the potentially unstable asperities did not rupture in the earthquakes, and instead slipped via aseismic afterslip.

As observations of postseismic slip continue to accumulate, the ratio of postseismic to coseismic moment may become an important constraint on physical models of earthquake rupture. The postseismic moment ratios will complement observations of coseismic stress drops, which are usually found to be magnitude independent, suggesting that earthquakes are self-similar: that large earthquakes are scaled-up small earthquakes. In this study, we have made observations that appear to contradict self-similarity. The median postseismic moment estimated for the 12 well-resolved  $M$  4 to 5 earthquakes is 0.36 (0.22 to 0.54 with 90% probability). This afterslip moment of these intermediate-magnitude moments is intermediate relative to previous observations; it is slightly larger than is typical of  $M > 6$  earthquakes and smaller than observed for  $M < 3.5$  earthquakes.

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