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8   ***welcome feedback!***

9 **Interplay between geometry and brittle deformation of bedrock fault scarps**

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15 **Key Points:**

16 • Brittle deformation of the fault core accounts for a portion of an earthquake's energy

17 budget, but how this varies is poorly constrained.

18 • We study how faults and fractures vary along two fault scarps in Malta to quantify spatial-

19 temporal variations in brittle deformation.

20 • Fracture intensity and geometry correlate with fault-plane geometry, and cross-faults

21 rotate fractures and the fault scarp.

22 **Key words**

23 Earthquakes, Active faults, Fracture networks, Energy budget, Seismic Hazard, Fault core

24   **Abstract**

25   Energy built up during interseismic phases is released during earthquakes as seismic energy and

26   dissipated within fault zones. Seismological observations indicate spatial variations in earthquake

27   behavior across a fault (e.g., rupture speed, stress drop, wave directivity). However, how this

28   relates to energy dissipation through brittle fracturing in the fault zone is difficult to quantify

29   from seismological observations. Spatial patterns of on-fault fractures across two bedrock normal

30   fault scarps in Malta are investigated, with fracture network properties related to identified

31   cross-faults and changes in fault-plane geometry. Cross-faults (i.e., faults that offset the main

32   plane) are common across both faults, with throws from 0.2 to 14 m. On-fault fractures are

33   commonly en echelon and contribute to a connected fracture network. The orientation, intensity

34   and connectivity of on-fault fractures display spatial changes associated both with proximity to

35   cross-faults and changes in fault-plane geometry (e.g., corrugations). On-fault fracture and

36   corrugation trends are block rotated by 10° to 40° by cross-fault zones, indicating corrugations

37   and some on-fault fractures pre-date cross-faults. The elevated fracture intensity and

38   connectivity adjacent to cross-faults, and the sharp contrast in fracture properties across them,

39   indicate that on-fault fractures continued to develop during or after cross-fault formation. This

40   suggests cross-faults are an energy sink during earthquake slip, reducing radiation efficiency and

41   possibly promoting rupture arrest. This observation indicates interaction between the cross-

42   faults and the main fault at surface, and suggests cross-faults play a significant role in earthquake

43   energy dissipation in the shallow sub-surface.

44

45 **Plain Language Summary**

46 During an earthquake, the energy that has built up on a fault is released during an earthquake,

47 potentially leading to extensive damage and/or loss of life. Only part of this energy is radiated as

48 seismic waves with a substantial proportion dissipated within the fault zone through frictional

49 heating and the breakdown of rock to create subsidiary faults and fractures. How energy is

50 partitioned controls rupture behaviour and ground shaking intensity, however, the spatial

51 variability of energy dissipation remains poorly understood.

52 To investigate this, we mapped fault plane geometry, cross-cutting faults, and on-fault fractures

53 across two normal fault scarps in Malta. We find that on-fault damage is highly heterogeneous

54 with fracture intensity, length, and network connectivity vary systematically along strike,

55 increasing near bends in the fault surface. Cross-faults form through the progressive localization

56 of strain within the on-fault fracture network. Cross-faults rotate the fault plane and pre-existing

57 fractures and localise the development of new fractures, producing abrupt step-changes in

58 fracture properties. Our findings demonstrate that geometric heterogeneities and cross-faults

59 strongly effect on-fault damage, and that mapping these features prvides key insights into the

60 energy budget during shallow earthquake slip.

61

62

63   **1 Introduction**

64   The total elastic strain energy (E_{total}) accumulated during the interseismic phase caused

65   by tectonic loading is released during earthquakes and/or aseismic slip. However, only part of

66   this energy is converted into seismic radiation. The energy budget, which ultimately controls

67   rupture dynamics and the amount of radiated seismic energy, can be expressed as:

68        $$E_{total} = E_R + E_G + E_{Th}$$

69   where E_R is the radiated energy through seismic waves, E_G is the fracture energy required for

70   rupture propagation, leading to the production of fault rocks and fractures, and E_{Th} is the energy

71   dissipated within the fault zone through frictional heating (Ide, 2002; Kanamori & Rivera, 2006;

72   Nielsen et al., 2016). The ratio of the radiated energy E_R to that energy dissipated during rupture

73   propagation and within the fault zone ($E_G + E_{Th}$) is referred to as the radiation efficiency, which

74   is derived from seismological sources (Beeler et al., 2003; Savage & Wood, 1971; Shearer, 2019).

75   Earthquakes with low radiation efficiency, where a high proportion of energy is dissipated within

76   the fault zone, result in reduced ground shaking compared to earthquakes with high radiation

77   efficiency in faults of similar size and tectonic loading history (e.g., Kanamori, 2001). Therefore,

78   understanding how energy dissipation varies both between and across active faults is crucial for

79   the assessment of the seismic hazard they pose.

80   Seismological observations indicate large along-strike and down-dip variability in rupture

81   speed, stress drop, and radiation efficiency for a given earthquake (e.g., Fletcher & McGarr,

82   2006). This suggests that energy dissipation varies along a fault's length, making it important to

83   consider energy dissipation at a given point on the fault (Cocco et al., 2023; Nielsen et al., 2016).

84 Faults are not single planar surfaces but rather fault zones consisting of one or more fault cores,

85 where the majority of strain is accommodated, surrounded by a damage zone (Caine et al., 1996;

86 Choi et al., 2016; Kim et al., 2004). Energy dissipated within the fault zone may therefore either

87 be on-fault (i.e., within the fault core), or off-fault (e.g., creation of new fractures or shear along

88 existing structures such as antithetic or synthetic faults within the damage zone (Cocco et al.,

89 2023; Nielsen et al., 2016)). Coseismic off-fault damage can reduce the radiation efficiency of

90 earthquakes on a fault when compared to scenarios where slip is confined to the fault core

91 (Cocco et al., 2023; De Paola et al., 2025; Huang, 2018; Okubo et al., 2019; Thomas et al., 2017).

92 Non-planar fault geometries such as fault bends and fault segmentation, which are oten referred

93 to as geometrical complexities, have been shown to promote additional off-fault damage (Okubo

94 et al., 2019; Rodriguez Padilla et al., 2024). The role non-planar geometries and fault

95 segmentation has on the distribution of on-fault energy sinks is challenging to deduce due to

96 resolution limitations of seismological methods, making small-scale stress-drop variations along

97 a fault surface difficult to detect.

98 Fault segmentation and non-planar geometries that affect the energy budget can be

99 identified through field studies of exhumed faults (e.g., Barth et al., 2013; Chester & Chester,

100 2000; Chester et al., 2005; Faulkner et al., 2003; Shipton et al., 2006). Most of the faults described

101 in those papers have, however, undergone multiple deformation phases (e.g., reactivation an/or

102 uplift and exhumation), complicating the distinction between features that formed during

103 earthquakes and those resulting from exhumation processes. One way to overcome this is

104 through deep borehole studies (e.g., the San Andreas Fault Observatory at Depth (SAFOD)

105 borehole) which provide direct insights into fault structure at depth (Bradbury et al., 2011;

106   Holdsworth et al., 2011). Borehole studies are costly and only sample a very small section of the

107   fault, therefore, extrapolating these data carries considerable uncertainities given the significant

108   spatial heterogeneity observed across fault zones (De Rosa et al., 2018; L. McKay et al., 2021;

109   Lucy McKay et al., 2019; Torabi et al., 2019). It is therefore difficult to deduce how fault zone

110   structure may affect along-strike changes in radiation efficiency at depth from current datasets.

111   Fracture development is an important energy sink in the shallow crust (Nielsen et al.,

112   2016; Shearer, 2019), and can develop both on the fault plane and within a damage zone. On-

113   fault fractures (i.e., fractures on the slip-surface) are abundant across bedrock fault scarps of

114   active normal faults (Chester & Chester, 2000; Friedman & Logan, 1970; Hancock & Barka, 1987;

115   Lacazette, 2009; Smeraglia et al., 2017; Stewart & Hancock, 1990). Cross-faults (i.e., faults that

116   occur at a high angle to, and cross the trend of, principle faults in an actively deforming region)

117   have been identified as playing an important role in shallow earthquake slip. Examples include

118   the 1987 Supersitition Hills earthquake sequence in California (Hudnut et al., 1989; Larsen et al.,

119   1992), the 1992 Big Bear and Landers, California, earthquakes (Jones & Hough, 1995), the 2010

120   Darfield earthquake in New Zealand (Barnhart et al., 2011; Elliott et al., 2012), the 2010-2011

121   Rigan earthquake sequence in Iran (Rezapour & Mohsenpur, 2013; Walker et al., 2013), the 2012

122   Sumatra earthquake (Meng et al., 2012; Wei et al., 2013), and 2019 Ridgecrest, California

123   (Barnhart et al., 2019; Ponti et al., 2020). Cross-cutting pre-existing faults can act as barriers to

124   rupture propagation. For example, an inherited Miocene-Pliocene thrust ramp is suggested as a

125   barrier to the propagation of the first two earthquakes of the 2016 Central Italy Sequence (24th

126   August, Mw 6.0 and 26th October, Mw 5.9), leading to a stress concentration on the pre-existing

127 fault that acted as the nucleation zone of the third and most powerful earthquake in the

128 sequence (30th October, Mw 6.5) (Pizzi et al., 2017).

129   Bedrock fault scarps associated with active or recently active faults provide an ideal

130 setting for systematically studying earthquake energy dissipation through brittle fracturing,

131 because 1) they have not been exhumed from great depth since faulting, 2) can be exposed

132 across hundreds of meters, 3) display consistent fault zone structures, and 4) commonly cut by

133 abundant cross-faults and on-fault fracture networks. We identify and quantify the properties of

134 cross-faults and on-fault fractures to investigate consistent dissipative energy sinks and examine

135 how these vary spatially across a fault plane, while also testing whether fault plane geometry

136 influences the growth of on-fault fractures. To achieve this, we use drone-derived Digital

137 Elevation Models (DEMs) of two bedrock fault scarps in Malta (the Victoria Lines and Maghlaq

138 faults, Fig. 1) to document and compare variations in fault plane geometry with the presence of

139 cross-faults (i.e., faults that offset the fault plane). The use of both fault scarps enabled us to

140 compare cross-fault development between an isolated segment of the Victoria Lines Fault, and a

141 portion of the Maghlaq Fault that is proposed to have formed through the linkage of several

142 initially isolated segments. Furthermore, we compare changes in fault plane geometry along the

143 Maghlaq Fault with the distribution of on-fault fractures mapped at a scale of 1:50. Our findings

144 suggest that fault plane geometry influences on-fault fracture distributions, and that cross-faults

145 cause the block-rotation of the fault plane at shallow depths. On-fault fractures continue to

146 develop after cross-fault development, with jumps in fracture intensity observed either side of

147 cross-faults, and an increase in on-fault fracture intensity close to cross-faults as they develop a

148 damage zone. Our data suggests fault plane geometry and cross-faults influence energy

149   dissipation in the near surface, and represent an important contribution to the energy budget

150   within the shallow sub-surface.

151   **2. Geological setting**
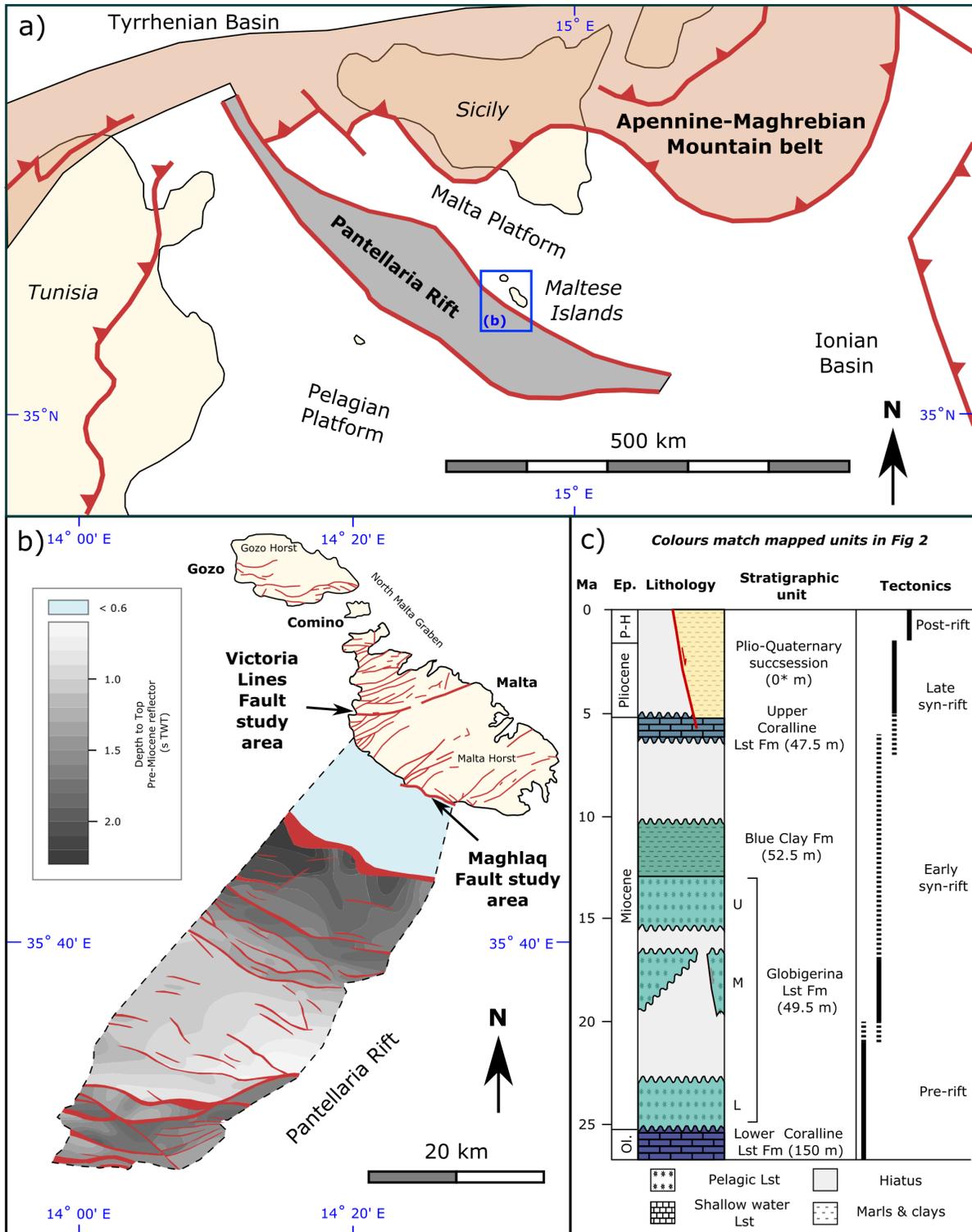
Figure 1: Geological setting of the Maltese Islands: a) Regional geology of the central

Mediterranean showing the tectonic setting of the Maltese islands (after Dart et al., 1993). b)

155 ***Location of major normal faults in the Maltese islands and an offshore section of the Pantellaria***

156 ***Rift (after Bonson et al., 2007). c) Tectono-stratigraphic log of the Oligocene to Quaternary***

157 ***sediments of the Maltese archipelago (after Dart et al., 1993). The timings of the tectonic***

158 ***phases are taken from Dart et al, (1993) and Martinelli et al, (2019). The Maghlaq Fault***

159 ***represents the most north-easterly extent of the faults of the Pantellaria Rift, with faulting***

160 ***initiating at ~21 Ma and locally offsetting the Plio-Quaternary succession.***

161   The Victoria Lines and Maghlaq Faults outcrop across the center and along the SW coast

162 of Malta respectively (Fig 1a). The Maltese islands are located on the Pelagian Platform, which is

163 on the foreland of the Apennine-Sicilian-Maghrebian orogenic belt (Argnani, 1990a; Dart et al.,

164 1993; Hill & Hayward, 1988). The platform is cut by the ESE-WNW trending Pantelleria Rift

165 system, also known as the Strait of Sicily Rift, which started opening in the late Miocene (Cello et

166 al., 1985; Dart et al., 1993; Finetti, 1985; Reuther & Eisbacher, 1985). The rift system is composed

167 of three ~NW-SE trending depocenters; the Pantelleria, Linosa, and Malta troughs (Reuther &

168 Eisbacher, 1985). The late Oligocene to present tectonic evolution of the Pantelleria Rift resulted

169 from the northward migration of the Apennine-Sicilian-Maghrebian orogenic belt (Argnani,

170 1990b; Dart et al., 1993).

171   Rifting caused two main sets of normal faults, trending WNW-ESE and ENE-WSW

172 respectively, that offset the predominantly carbonate Oliogo-Miocene stratigraphic succession

173 (Dart et al., 1993; Illies, 1981; Reuther & Eisbacher, 1985). Whilst there is some debate upon the

174 exact timings and kinematic details surrounding the structural evolution (e.g., Martinelli et al.,

175 2019 *and references therein*), four tectono-sedimentary stages have been identified (Fig. 1b; Dart

176 et al., 1993): (i) pre-rift (>21 Ma); (ii) early syn-rift, extending from 21 to 6 Ma (Dart et al., 1993),

177 or 20 to 17 Ma (Martinelli et al., 2019), characterized by minor faulting and dyke emplacement;

178 (iii) late syn-rift, extending from ~5 to 1.5 Ma (Dart et al., 1993), or 7 to 1.5 Ma (Martinelli et al.,

179 2019), characterised by major fault growth, and (iv) post-rift (<1.5 Ma). Kinematic analysis of

180 faults on the Maltese Islands show that faults that formed during the early rift phase developed

181 under a different stress field than those that developed during the late syn-rift phase (Martinelli

182 et al., 2019). Early faults, which typically have <5 m throw, developed under oblate WNW-ESE

183 extension (Martinelli et al., 2019). Faults during the late syn-rift phase show throws of up to 200

184   m, and include the regional fault sets on the Maltese Islands (e.g., Victoria Lines and Maghlaq

185   faults) (Martinelli et al., 2019).
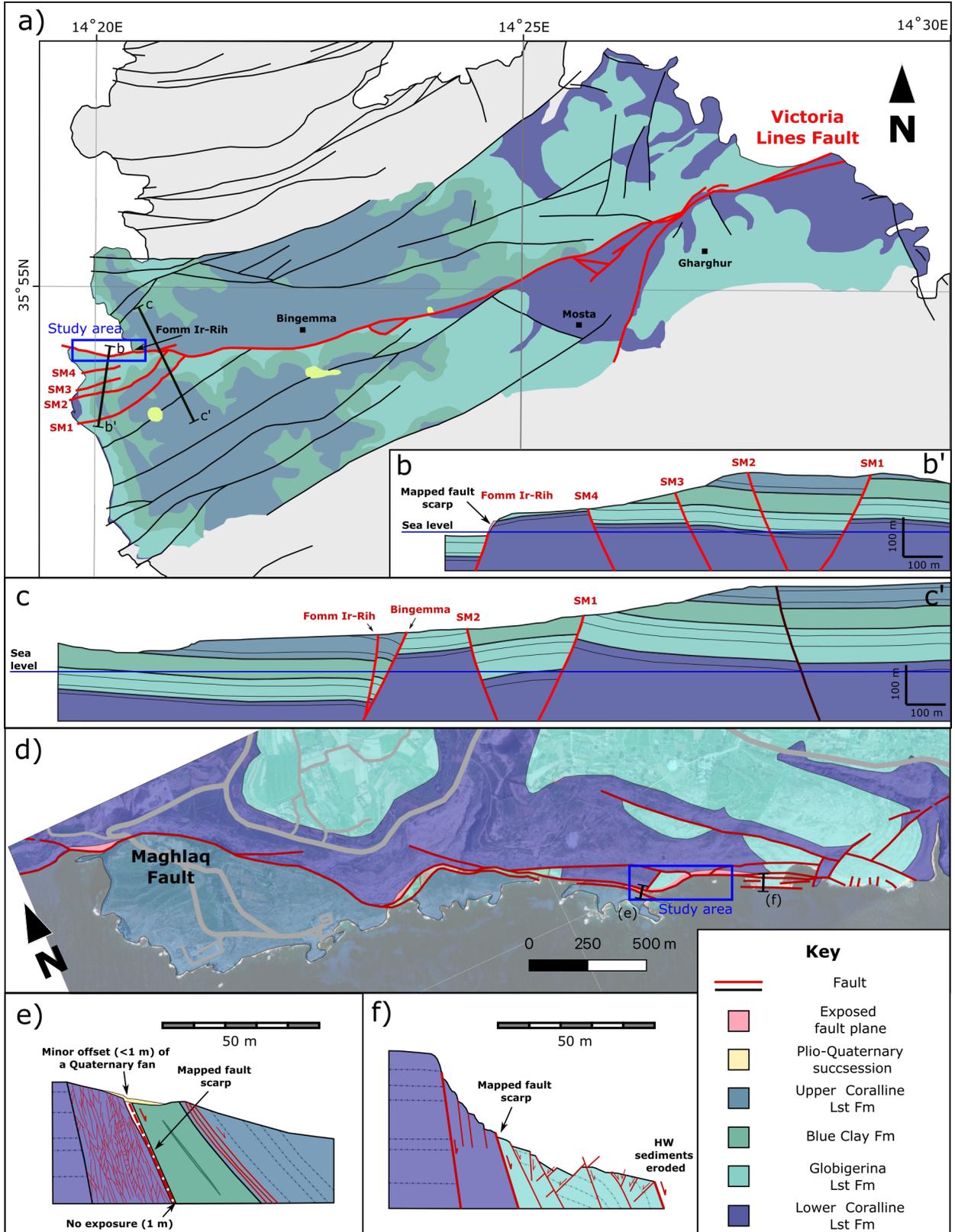
187   **Figure 2: Geological map and cross-sections of the Victoria Lines and Maghlaq Faults (after**

188   **Villani et al. (2018) and Bonson et al. (2007) respectively). a) Geological map of the area**

189   **surrounding the Victoria Lines Fault, showing the study area and location of cross-sections (b)**

190   **and (c). (b-c) Cross-sections showing the throw distribution across the Fomm Ir-Rih and St.**

191   **Martin fault segments of the Victoria Lines fault. d) geology of the area sorrounding the**

192   **Maghlaq Fault showing the location of the study area and panels (e) and (f). (e-f) Cross-sections**

193   **either side of the study area in the Maghlaq Fault.**

194   The Victoria Lines Fault, also called the Great Fault (Putz-Perrier & Sanderson, 2010),

195   extends for ~14 km across the center of Malta, creating a prominent escarpment that trends

196   between 250° in the east and 270° in the west (Fig 1, 2a). The Victoria Lines Fault separates the

197   North Malta Graben to the north from the Malta Horst to the south and has a maximum throw

198   of ~195 m (Putz-Perrier & Sanderson, 2010). There is no evidence of Quaternary activity along

199   the Victoria Lines Fault (Illies, 1981; Trechmann, 1938), with the last postulated movement based

200   on seismic tomography profiles across the western extent of the fault being during the Middle

201   Pleistocene (~0.6 Myr) (Villani et al., 2018). The Victoria Lines Fault is not a continuous fault, but

202   is composed of multiple sub-parallel and overlapping strands, with some localized splay faults

203   and branches in the westernmost and easternmost portions of the fault (Fig 2a-c) (Villani et al.,

204   2018). Displacement in the western portion of the fault is accommodated across a few segments

205   that include the 2.1 km-long Fomm Ir-Rih segment, four sub-ordinate NW- and SE- dipping splays

206   termed the St. Martin Segments, and the ~4.6 km long Bingemma segment (Fig 2 a-c). This study

207   focuses on the Fomm Ir-RiH segment, which displays a maximum throw of 119 m and decreases

208   to near zero at the east of the study area (Fig 2c; Villani et al., 2018). The study area is a 1.2 km

209 long continuous coastal section (sea cliff) of a northerly dipping normal fault plane (Fig 2a),

210 enabling along-strike patterns in cross-faults to be systematically studied across the fault

211 segment. Because of weathering of the Victoria Lines Fault scarp, it was not possible to reliably

212 map fractures at a consistent resolution across the study area. It was therefore decided that only

213 cross-faults will be studied for the Victoria Lines Fault.

214 The Maghlaq Fault is located along the south-west coast of Malta, with the footwall of

215 the fault exposed along a ~4 km continuous coastal section (Bonson et al., 2007). The trend of

216 the Maghlaq Fault is ESE-WNW, matching fault orientations observed in the Pantelleria Rift to

217 the south-west of Malta (Fig. 1) (Bonson et al., 2007; Pedley et al., 1976; Reuther & Eisbacher,

218 1985). Displacement probably started during the Messinian (7-5 Ma), with activity increasing

219 after 5 Ma (Dart et al., 1993; Martinelli et al., 2019). Quaternary fans are offset in places by the

220 Maghlaq Fault (Fig 2e), suggesting slip on the Maghlaq Fault continued for longer than the

221 Victoria Lines Fault (Bonson et al., 2007). Despite seismicity being recorded in the Northern

222 Pantelleria Rift, there is no direct evidence for neotectonic fault movements along the Maghlaq

223 Fault (Bonson et al., 2007; Reuther & Eisbacher, 1985). Faulting across the exposed scarp likely

224 occurred in to top few hundered meters, with the burial of the footwall sediments not exceeding

225 250 to 300 m (Bonson et al., 2007). The Maghlaq Fault consists of left-stepping, en echelon

226 normal fault array striking ESE-WNW striking, with 60˚ dipping segments 1-2 km long (Fig 1;

227 Bonson et al., 2007). Locally, these are joined by 10s to 100s m long fault segments that form

228 fault bends, including our study area (Figure 2d-f) (Bonson et al., 2007). In our study area, a

229 southerly dipping fault scarp is, similarly to the Victora Lines Fault, preserved over a continuous

230 ~430 m coastal section. Unlike the Victoria Lines Fault segment, the Maghlaq Fault segment has

231   an exceptionally well exposed slip surfaces (Fig. S1-2), enabling on-fault fractures to be mapped

232   across the whole study area.

233   **3 Materials and Methods**

234   3.1 Quantifying fault plane geometry and inferring slip-vector

235   An unmanned aerial vehicle (drone) was used to capture imagery across the Victoria Lines

236   and Maghlaq fault scarps, ensuring a minimum 70% overlap between photographs to facilitate

237   photogrammetric processing, which was undertaken using the software Agisoft MetaShape to

238   create virtual outcrop models. From the virtual outcrop models, digital elevation models (DEM)

239   and orthomosaics were extracted and loaded into QGIS for analysis. The images for the Maghlaq

240   Fault covered sections of the fault with exceptional exposures of the fault scarp, extending ~430

241   m along strike and between ~4 and ~25 m down-dip. The Victoria Lines Fault model covered the

242   full extent of the Fomm Ir-RiH segment, extending ~1250 m along strike and between 5 and 19

243   m down dip.

244   The fault plane geometry was quantified, and slip-vectors inferred, from the DEM. Dip

245   and dip-direction maps were generated using the slope and aspect raster tools in QGIS.

246   Corrugations, used to infer slip vector, were inferred through the visual analysis of the dip-

247   direction map to map the location of down-dip ridges and troughs. These features were mapped,

248   using a minimum length of 0.4 m to reduce the effect of noise. The trend of the corrugation was

249   calculated in QGIS, and the dip was determined from calculating the slope of the linear regression

250   of the depth values extracted along the mapped corrugation.

251
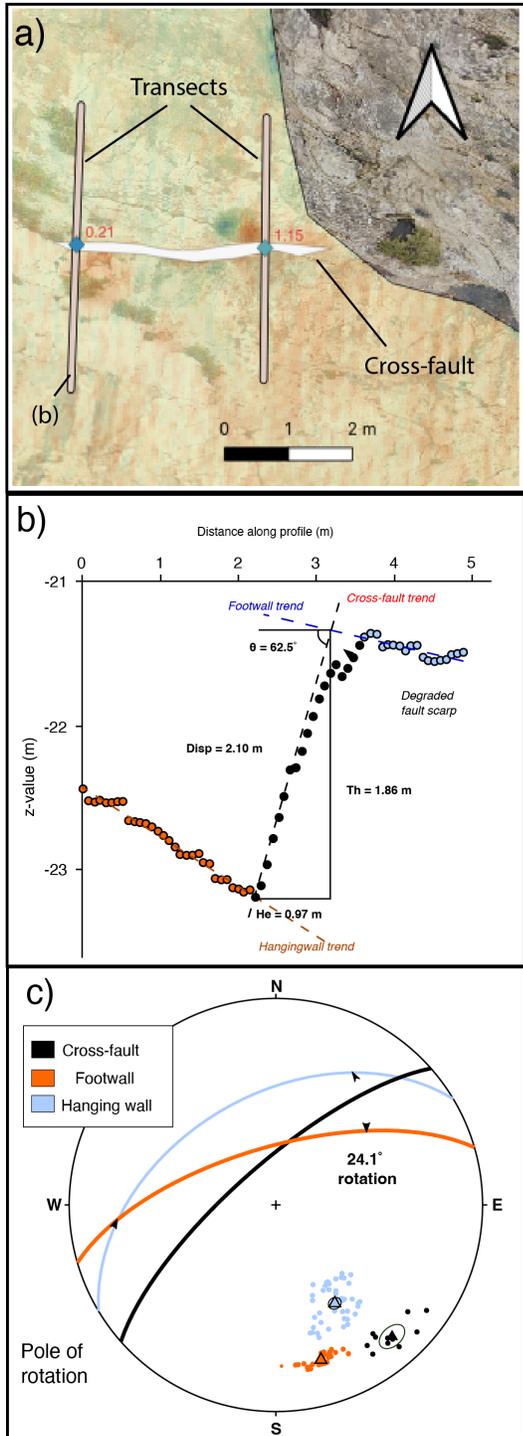
252 **Figure 3: Quantification of cross-fault throw and rotation. a) Example transects used to extract**

253 **cross-fault throw. b) Profile extracted from the DEM, highlighting how throw is measured; c)**

254 **Stereographic projection of fault plane geometry extracted from the transect that is used to**

255 **calculate scarp rotation using stereograph analysis.**

256 Several cross-faults are observed along both the Victoria Lines and Maghlaq faults. These

257 are mapped based on abrupt changes in dip-direction that coincide with a change in z-value on

258 the DEM. Once mapped, fault-perpendicular transects, with sufficient density to capture along-

259 stike changes in fault throw, were used to extract slope profiles across the fault. Profiles are then

260 visually split into footwall (FW), hanging wall (HW), and fault plane subsections to calculate fault

261 throw, heave, displacement and an estimation of fault dip (Fig. 3a). Additionally, these profiles

262 are used to determine the fault plane geometry from the dip and dip direction maps with the

263 mean orientation calculated for each subsection (HW, FW and Fault Plane) using Stereonet 10

264 (Fig 3b) (Allmendinger et al., 2013; Cardozo & Allmendinger, 2013). The angular rotation between

265 the footwall and hanging wall, as well as the line of intersection, is calculated for each transect

266 and compared to the extracted throw value (fig. 3b).

267 3.2 On-fault fracture mapping and fracture network analysis

268 Fracture networks were analyzed through lineament mapping in QGIS by a single

269 interpreter, maintaining a constant scale of 1:50, and with snapping enabled to reduce subjective

270 bias (Andrews et al., 2019). Features above a map-view length of ~0.05 to 0.1 m (~5-10 pixels)

271 were mappable at this scale. Fracture type could not be reliably discerned, leading to all

272 lineaments being mapped as a single layer. For lineaments wider than the pixel size (>0.02 m),

273   the fracture was digitized at its center, with connected fractures snapped onto the mapped

274   fracture. Although this method introduces a small uncertainty in the spatial location, the

275   magnitude of this uncertainty is low, thus not affecting comparisons between fracture properties

276   and fault plane geometry. Fracture aperture was not considered in this study, as the majority of

277 apertures fall below the pixel resolution of the orthomosaic and aperture may have been

278 widened by weathering and erosion.

279   Following digitization, the map view trace length (Tl_{map}) and orientation of fractures were

280 added to the attribute table in QGIS. An estimation of true trace length (Tl_{true}) is then calculated

281 by considering the best-fit slope along the transect similarly to the calculation of cross-fault dip.

282   On-fault fracture trend (i.e., fault-fracture intersection lineation) was calculated using the

283 start and end point of the polylines. Whilst this will induce small orientation errors for sinuous

284 fractures, most digitized fractures had low sinuosity (average = 1.005), so should not affect the

285 results presented here. For plotting purposes, the central points of fractures were chosen when

286 displaying spatial trends in fracture orientation. An adapted gridded circular scanline method

287 (after Mauldon et al., 2002) with a spacing of 0.2 m and a 1 m radius was used to investigate

288 spatial distributions of fault scarp geometry, fracture intensity and fracture connectivity (See

289 Supplementary 2 for further details). We then compared trends in fault scarp geometry, fracture

290 intensity, and fracture connectivity to the presence of mapped cross-faults.

291 **3 Results**

292 *3.1 Fault plane geometries and corrugation trends*

293
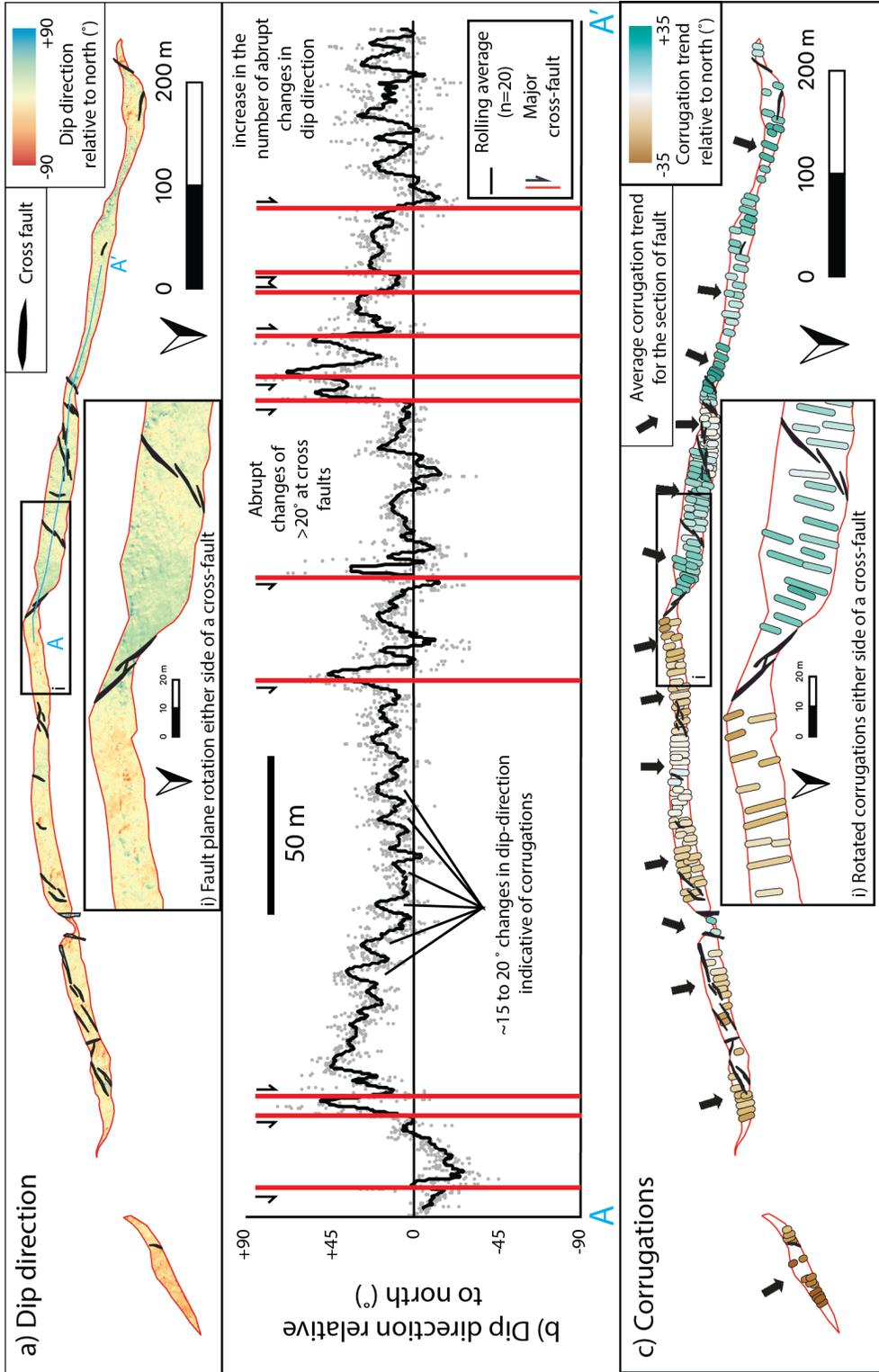
294   *Figure 4: Fault plane geometry and corrugation trends for the Victoria Lines Fault. a) Dip-*

295   *direction map highlighting the rotation of the fault plane either side of major cross-faults. b)*

296   *Along-strike variation in dip-direction along the profile A-A', highlighting the changes in dip-*

297   *direction used to map out corrugations and abrupt changes used to identify cross-faults. c) Map*

298   *showing the trend of mapped corrugations across the Victoria Lines Fault. Black arrows indicate*

299   *the average corrugation trend between prominent cross-faults. Note how corrugations*

300   *converge towards the center of the study area, and are rotated across major cross faults (see*

301   *sub-panel ci).*
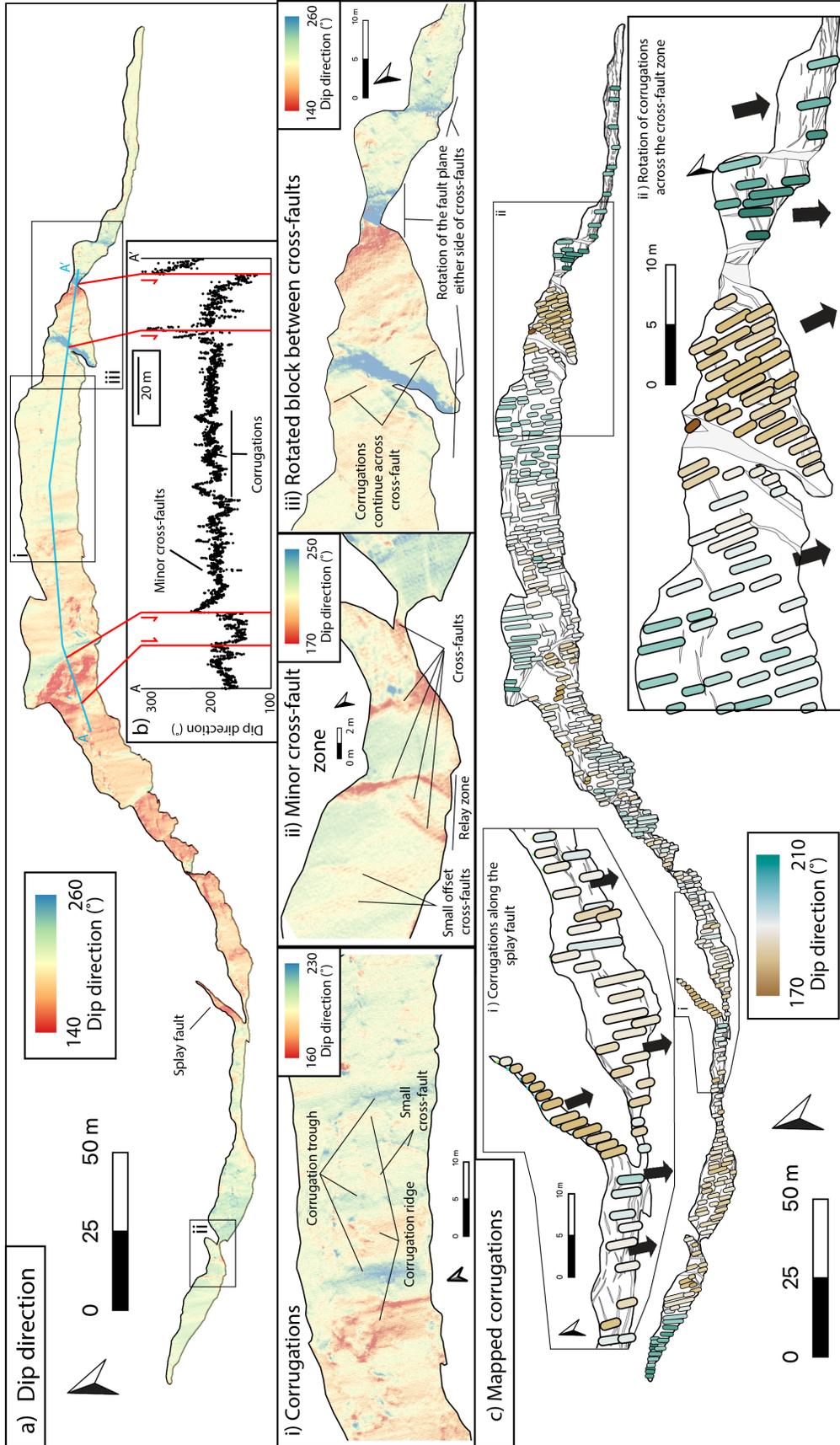
303   *Figure 5: Fault plane geometry and corrugation trends for the Maghlaq Fault. a) Dip-direction*

304   *map highlighting the rotation of the fault plane either side of major cross-faults and the*

305   *presence of corrugations with insets i-iii showing detailed sections. b) Along-strike variation in*

306   *dip-direction along the profile A-A', highlighting the changes in dip-direction used to map out*

307   *corrugations and abrupt changes used to identify cross-faults. c) Map showing the trend of*

308   *mapped corrugations (used to infer slip vectors) across the Maghlaq Fault. Black arrows*

309   *indicate the average corrugation trend. Close up sections are highlighted to show i) how*

310   *corrugation trend varies across the splay fault and ii) how corrugation trends are rotated by*

311   *cross-faults. Note how fault plane geometry varies along the length of the studied section and*

312   *that fault geometry and corrugation trend rotates close to, or between closely spaced, cross*

313   *faults.*
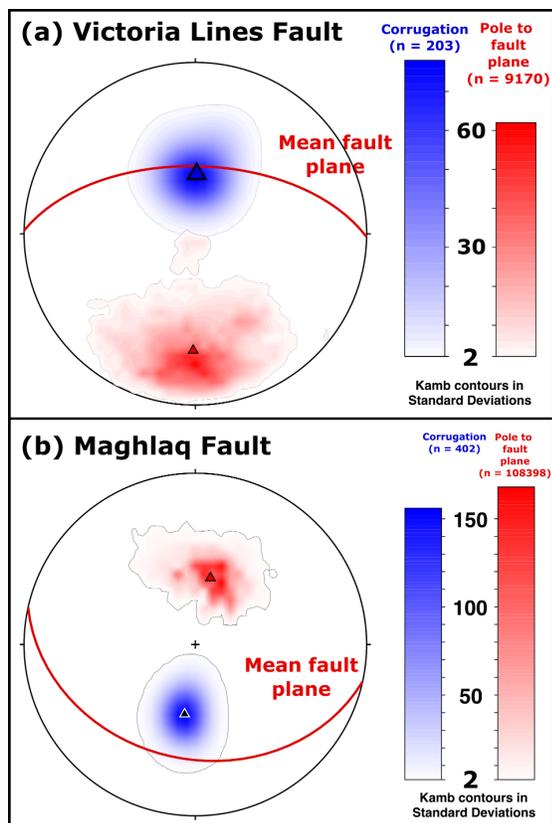
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315 ***Figure 6: Stereographic projections of the (a) Victoria Lines and (b) Maghlaq faults showing the***

316 ***poles to the fault plane (red) and trend and plunge of mapped corrugations (blue). Triangles***

317 ***show the mean values of orientations of pole to the fault plane and of corrugations. Note the***

318 ***high dispersion in both fault plane geometry and corrugation trend, in part caused by the***

319 ***presence of cross faults that rotate the fault plane.***

320 Along-strike variations in fault plane geometry are observed at multiple scale for both the Victoria

321 Lines and Maghlaq faults (Figs. 4-6). The Victoria Lines Fault dips to the north, with a mean

322 orientation of 001˚/58˚ (dip direction/dip), though dip and dip direction shows high variability as

323 shown by the contours in Fig. 6a. The Maghlaq Fault dips to the south-southwest, with a mean

324 orientation of 193˚/56˚, and exhibits similar dispersion in fault plane orientation as the Victoria

325 Lines Fault (Fig 6b). Additionally, a splay fault with an average orientation of (160˚/67˚) was

326 identified to the west of the study area (Fig 5a). Spatial analysis of dip-direction variations (Fig

327 4a,b, 5a) reveals that small-scale fault plane geometry changes occur at the meter to decimeter

328 scale, while larger variations appear either as step changes or over short distances of <10 m (Fig

329 4b, 5a). Along-strike transects across these small-scale variations (e.g., Fig 4b) show a sigmoidal

330 pattern in dip direction, with a change of dip direction of between 20˚ and 30˚ occurring over

331 along-strike distances of between 5 to 20 m. These observations combined with the down-dip

332 elongation of these features (Fig 5ai) has led to these being interpreted as corrugations.

333 Where abrupt changes in fault plane geometry occur, such as at ~245 m in Fig 4b and in the

334 center of the Maghlaq Fault (Fig 5a), the change in dip direction frequently exceeds 20˚ and, in

335 some locations, exceeds 45˚ over a distance of up to 15 m. These abrupt changes have been

336   interpreted as faults that intersect the main fault surface, hereafter termed cross-faults, which

337   offset the fault scarp. Cross-fault mapping was conducted using dip-direction and dip-attribute

338   maps, identifying 57 cross-faults on the Victoria Lines Fault and 535 on the Maghlaq Fault. The

339   overall patterns and parameters of cross-faults are further explored in section 3.2.

340   Visual mapping of corrugations enabled abundant corrugations to be identified across the

341   Victoria Lines (n = 203) and Maghlaq (n = 402) fault scarps (Fig 4c, 5b). The average corrugation

342   orientations (61˚ -> 001˚and 56˚-> 189˚ respectively) indicate that both faults display normal slip

343   vectors (Fig. 4). However, similar to the fault plane geometry, corrugation trends show high

344   dispersion, ranging from 327˚ to 030˚ for the Victoria Lines Fault and 280˚ to 135˚ for the

345   Maghlaq Fault (Fig 6). Spatial variations in corrugation trend reveal both similarities and

346   differences between the two fault scarps. For both faults, corrugations either side of major cross-

347   faults show a degree of rotation (Fig 4ci, 5bii). For the Victoria Lines Fault, corrugations exhibit

348   'convergent slip vectors' (cf. Roberts et al., 1996), with oblique corrugations towards the eastern

349   and western edges of the study area (arrows in Fig 4c). This suggests that the Fomm Ir-RiH

350   segment of the Victoria Lines Fault is behaving as an isolated fault, with its tips located near the

351   study area's boundary. Conversly, the Maghlaq Fault shows no systematic changes in corrugation

352   trend, other than in the vicinity of the splay fault, suggesting the study area is located in the

353   center of a larger fault segment.

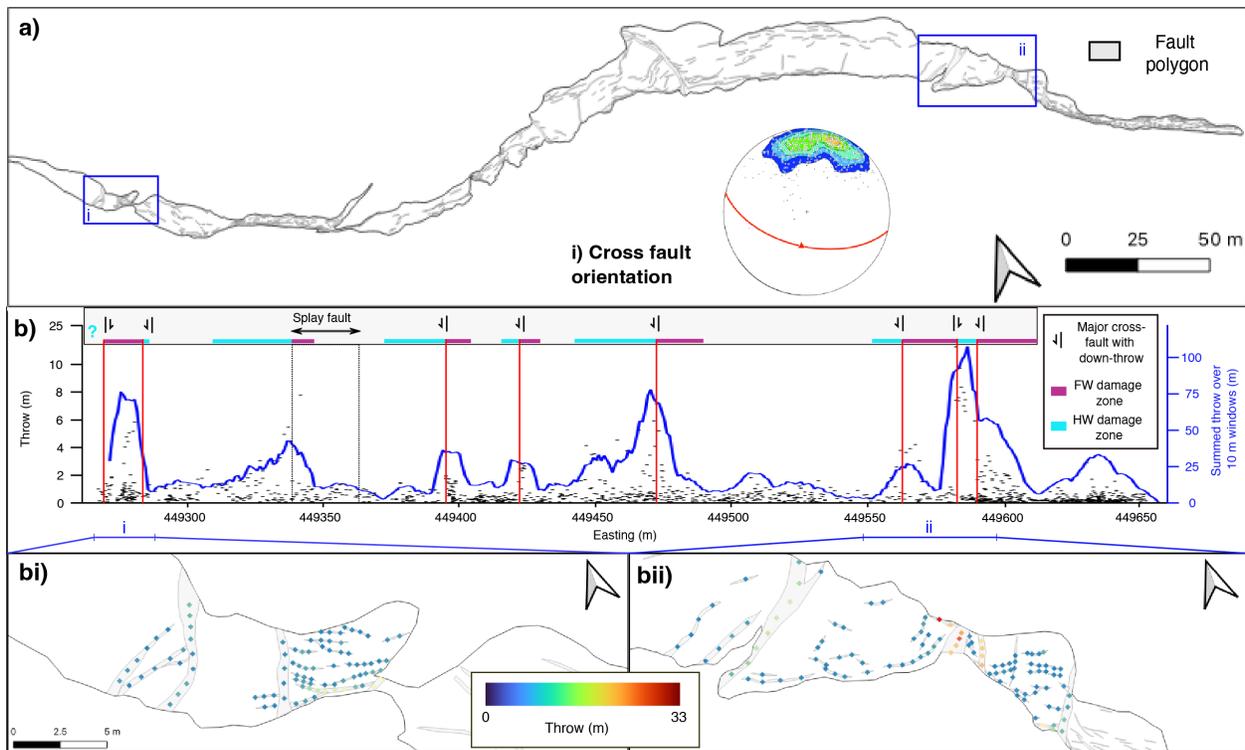354   ***3.2 Cross-faults properties***

355
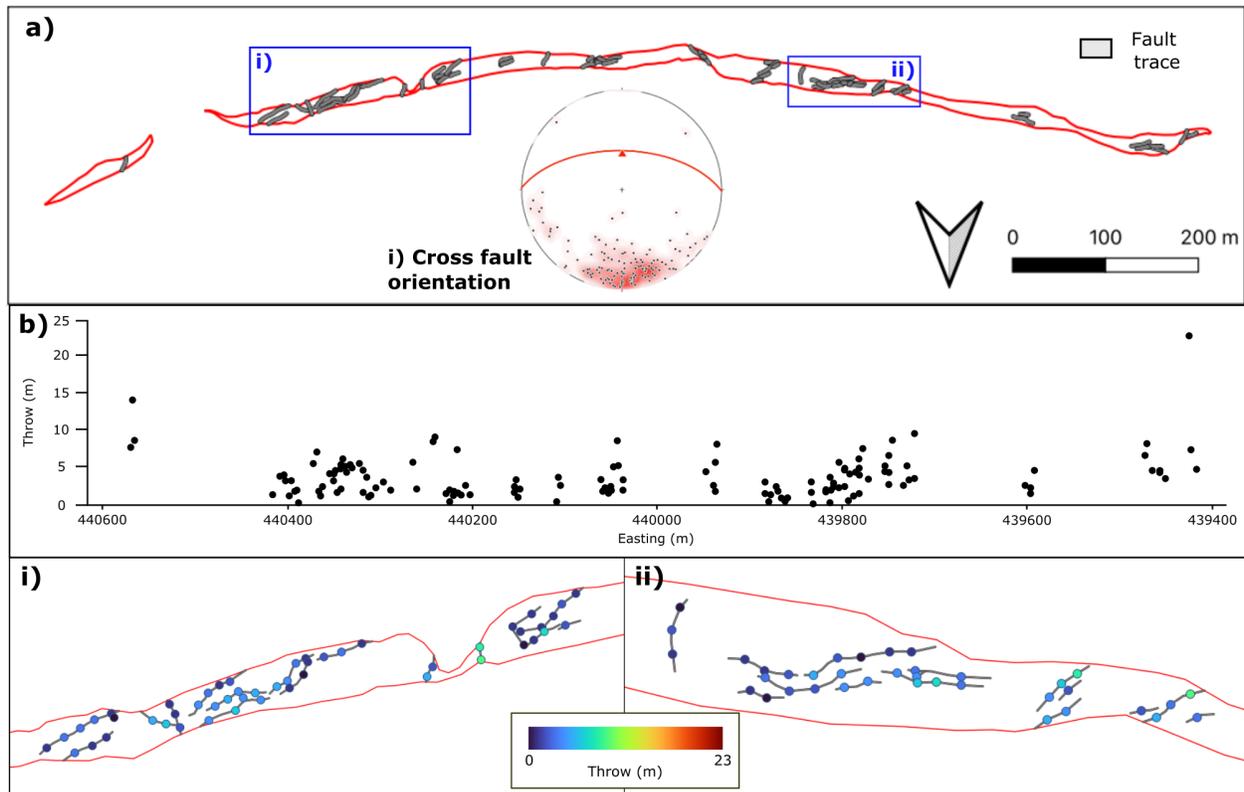
Figure 7: Cross-fault properties for the Victoria Lines Fault. a) Location of mapped cross-faults across the Victoria Lines Fault with an inset (i) highlighting their orientation. b) Along-strike distribution of cross-fault throw. How the throw distribution relates to the mapped cross-fault networks are shown in sub-panels bi and bii.

Cross-faults mapped using the dip- and dip-direction attribute maps reveal similarities and differences between the Victoria Lines and Maghlaq Faults (Fig 7, 8). Across the Victoria Lines Fault, cross-faults are observed across the whole study area, forming fault zones where the cross-fault perpendicular extent of mapped cross-faults ranges from 2 to 15 m (Fig 7a). Within these zones, faults that intersect the main fault plane at high angles tend to exhibit slightly greater throw than faults that intersect at a low angle (Figs 7b). This distribution pattern contrasts with that observed along the Maghlaq Fault, where cross-fault distribution varies across the study

367   area (Fig 8). Fewer cross-faults were mapped in the western portion of the Maghlaq Fault, and

368   the throw along these structures is typically below 2 m (Fig 8b). In contrast, the eastern portion

369   of the Maghlaq Fault contains a larger number of small-offset cross-faults, leading to an overall

370   increase in the rugosity of the main fault plane (Fig 5a, 8). The central portion of the Maghlaq

371   Fault is characterized by high (up to 6 m) offset cross-fault zones composed of tens of individual

372   cross-faults, which are separated by portions of the study area where less cross-faults are

373   observed (Fig 8b). Cross-fault zones, where total throw across mapped strands exceeds 10 m,

374   strongly influence fault plane geometry. These zones contribute to localized clockwise, or

375   anticlockwise rotation of the fault plane, development of relay zones, and the formation of

376   kinematically linked fault systems that span tens of meters perpendicular to cross-fault strike (Fig

377   8bi, bii). Both the Maghlaq and Victoria Lines Faullts show similar cross-fault patterns; however,

378   the Maghlaq Fault exhibits a higher frequency of small-throw faults compared to the Victoria

379   Lines Fault.

Figure 8) *Cross-fault properties for the Maghlaq Fault. a) Location of mapped cross-faults across the Maghlaq Fault with an inset ai highlighting the orientation of all mapped cross-faults. b) Along-strike distribution of cross-fault throw. Note that because of the high number of mappable cross-faults, a rolling summed throw over 10 m windows was also undertaken, enabling the identification of cross-fault damage zones. How the throw distribution relates to the mapped fault networks are shown in sub-panels bi and bii.*

Cross-faults along both faults display a variability in extracted cross-fault properties (**Fig 7, 8**). Cross-faults display a broad range of orientations, with distinct average trends and downthrow directions between study areas. Cross-faults across the Victoria Lines Fault have an average trend of 073° and most faults downthrowing to the north compared to 104° with downthrow directions

392   to the south to southwest for the Maghlaq Fault (**Fig 7a, b**). The map-view length and throw of

393   cross-fault also show considerable differences between the two faults, with cross-faults across

394   the Maghlaq Fault typically being shorter and having a greater range, and lower average, throw

395   (**Fig 8c, d**). This difference results in a higher proportion of censored cross-faults along the

396   Victoria Lines Fault (37%) compared to the Maghlaq Fault (18%). The differences in throw and

397   length leads to higher throw-to-length ratios along the Maghlaq Fault relative to the Victoria

398   Lines Fault.
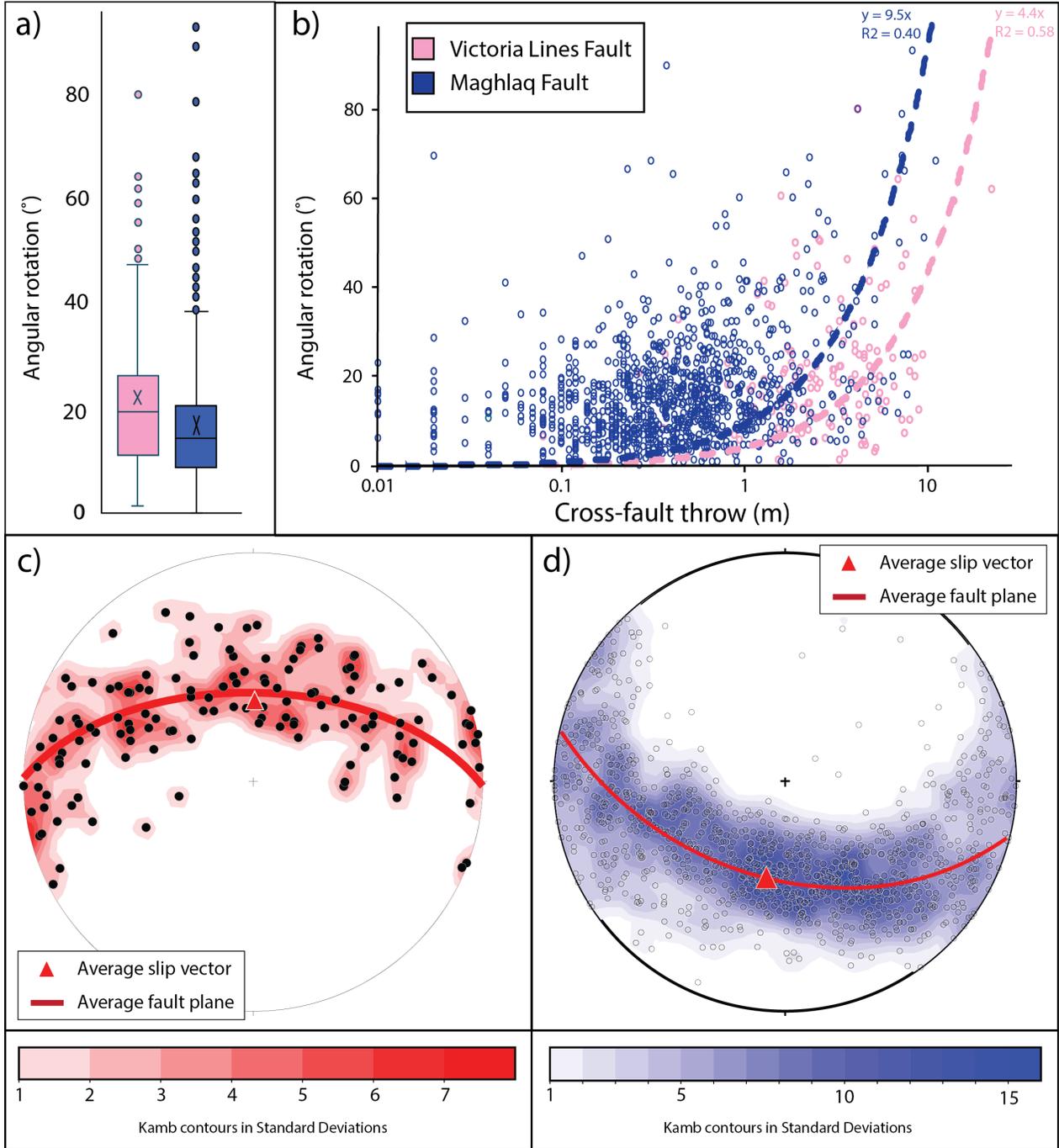
Figure 9: Angular rotation of main fault plane caused by cross-faults: a) Box and whisker plot of angular rotations caused by cross-faults for the Victoria Lines and Maghlaq faults. X refers to the mean angular rotation. b) Relationship between cross-fault throw and magnitude of

403   **angular rotation. Stereographic projections of the rotation poles for the c) Victoria Lines Fault**

404   **and d) Maghlaq Faults.**

405   Either side of cross-faults, measurable rotations of the main fault plane orientation are observed

406   (Fig 4a, 5a, 9). These rotations were quantified by analyzing cross-fault transects and extracting

407   fault plane orientations from the footwall, hanging wall and fault plane of cross-faults (Fig 3b).

408   Angular rotations of up to 55˚ were recorded, with median values of 20° and 17° for the Victoria

409   Lines and Maghlaq Faults, respectively (Fig. 9a). A weak positive correlation is observed between

410   angular rotation and throw (Fig 9b), though the relationship varies across different segments of

411   the fault (Figs. 4a, 5a). Rotation axes cluster around the mean fault plane orientation, with a wide

412   spread occurring along this plane (Fig 9 c, d). There is a slight concentration of rotation planes

413   close to the mean slip-vectors of both faults, but no clear patterns can be observed from this

414   data. Fault plane rotation is further evidenced by changes in corrugation trends near major cross-

415   faults. Across both faults, corrugation orientations are rotated either side of cross-faults (Figs. 4c,

416   5b). For example, in the central portion of the Victoria Lines Fault, a southeast-trending cross-

417   fault induces an 18° rotation in corrugations, shifting their orientation from 62° → 344° to 57° →

418   017° (Fig. 4ci). Similarly, along the Maghlaq Fault, multiple locations exhibit block rotation of

419   corrugations adjacent to cross-fault zones that exceed 10 m throw (Figs. 5dii). The magnitude of

420   these corrugation rotations is comparable to the angular rotation observed in overall fault plane

421   geometry (Fig. 4b, 5a). This pattern suggests that fault corrugations developed prior to cross-fault

422   formation in response to the early stages of normal faulting. Due to the resolution of the imagery,

423   it is not possible to test whether any other forms of kinematic marker (e.g., striations) are rotated

424   by cross-faults.

425   *3.3 On-fault fractures on the Maghlaq Fault*
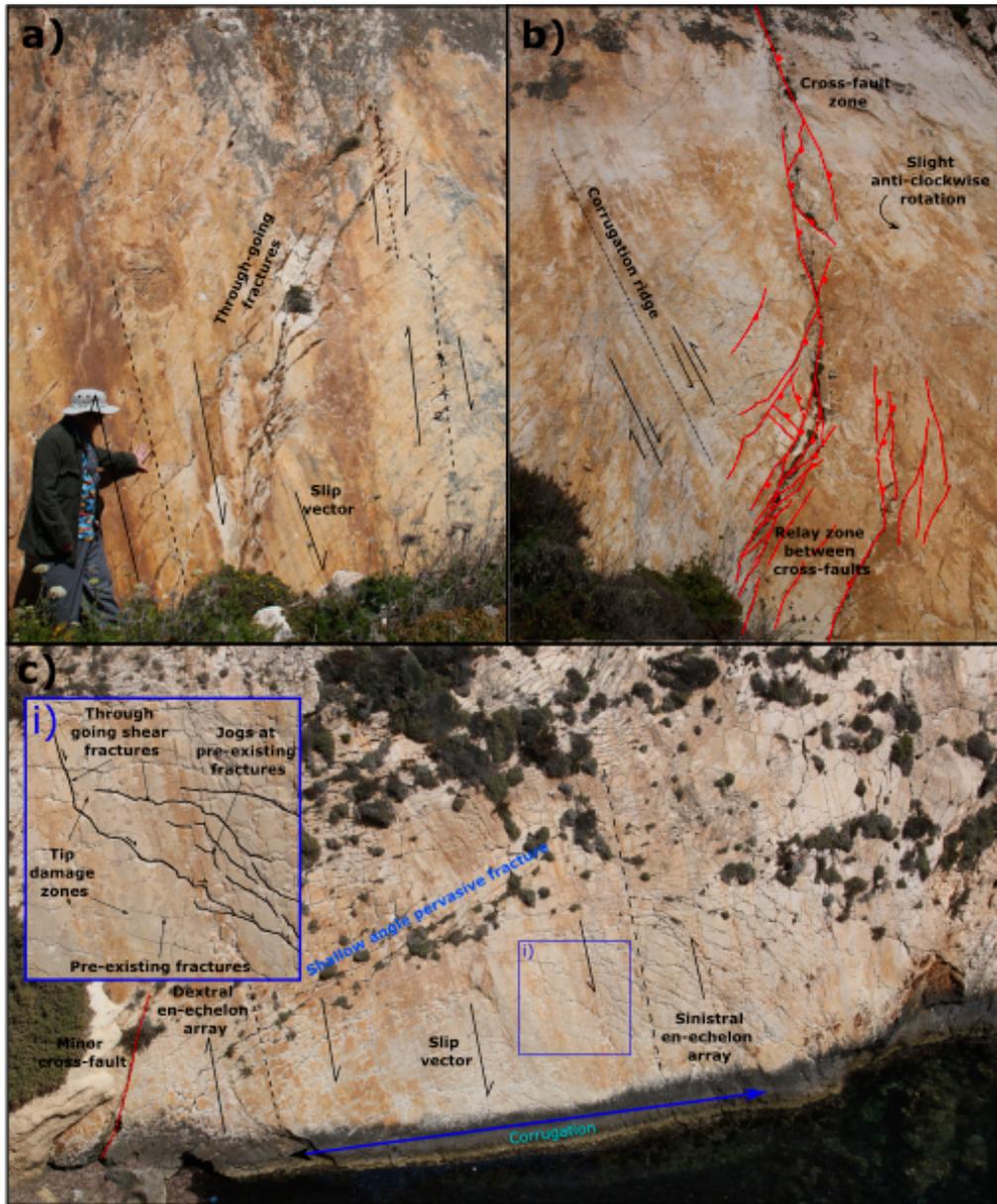
426

427   **Figure 10: Field photographs of on-fault fractures on the Maghlaq Fault showing a) through-**

428   **going fractures and en-echelon arrays cutting across a minor corrugation; b) minor cross-fault**

429   **zone showing how cross-faults interact with on-fault fractures; c) en-echelon arrays either side**

430   **of a corrugation, with the dashed lines representing the corrugation ridges where the arrays**

431 **nucleate, with the corrugation trough occurring where slip-vector is marked. The inset (i)**

432 **highlights how through going shear fractures interact with pre-existing on-fault fractures.**

433 The well-exposed slip surface of the Maghlaq Fault exhibits an extensive fracture network that

434 provides an exceptional opportunity to analyze fracture patterns and along-strike trends in on-

435 fault fracture properties, and infer their development (Fig 10-12). On-fault fractures occur as

436 either opening-mode fractures with no disernable offset, or shear fractures with cm offsets, with

437 no evidence of mineralization on either fracture type (Fig 10). Fractures are commonly en

438 echelon, with the trend of the arrays aligned within 20° of the observed slip vector, as defined by

439 striations and corrugations on the fault surface (Fig 10). These arrays can be either sinistral or

440 dextral, with the sense of shear reversing across corrugation ridges and troughs (Figs 10b, c).

441 Arrays occur either as isolated features, where fractures of similar length terminate at i-nodes

442 (Fig 10a), or as linked structures that form larger connected fractures (i.e., through-going

443 fractures), with trace lengths ranging from tens of centimeters to over a meter (Fig 10c). Where

444 through-going fractures (i.e., fractures that cut across and utilize pre-existing fractures in the

445 network) develop, tip-damage zones (cf. Kim et al., 2004) occur between fracture tips, and jogs

446 are observed where a through-going fracture steps between previously isolated fractures (Fig

447 10ci). This suggests that there is a degree of shear across these fractures as strain becomes

448 progressively localized onto larger fractures, which evolve into faults, while smaller en-echelon

449 fractures are gradually abandoned as the network evolves (Figs. 10a, 10ci, e.g., Cowie, 1998).
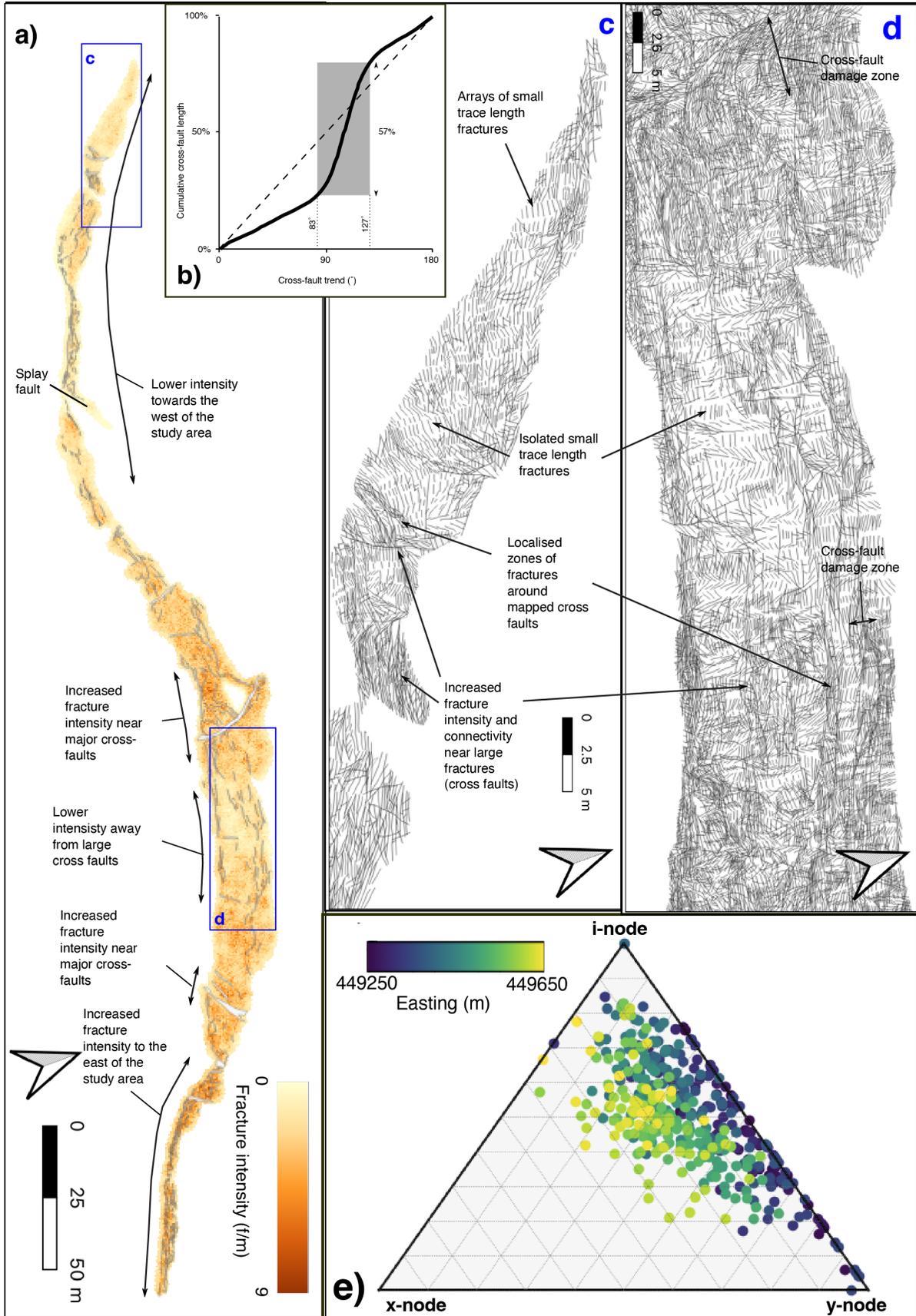
450

451 In some areas, through-going fractures exhibit cm-scale offsets (Fig 10ci), and in localized zones,

452 these features evolved into the aforementioned cross-faults (Fig 10b). This likely occured through

453   the progressive linkage of shear fractures, such as those in Figure 10a. These zones can extend

454   several meters across the fault scarp, creating highly connected networks that displace the fault

455   plane (Fig 10b). More developed cross-fault zones display centimeter- to decimeter-scale out-of-

456   plane offsets, resulting in a non-planar fault surface with a well-developed damage zone and

457   relay ramps between segments (Fig 10b). The throw observed across a single cross-fault mapped

458   using the dip- and dip-direction maps therefore probably represents throw accommodated by

459   several smaller structures and localized fault-plane rotation. These structures, as well as the

460   larger-scale cross-fault zones across the Victoria Lines and Maghlaq fault scarps (Fig 6), indicate

461   a systematic progression of deformation from isolated en-echelon fractures to cross-faults.

462

463   Systematic mapping of 27,459 on-fault fractures on the Maghlaq Fault at a scale of 1:50 highlights

464   the spatial variations in fracture intensity, connectivity, and trace length (Fig 11). Fracture

465   intensity ranges from 0.04 f/m to 7.52 f/m (mean = 2.55 f/m) while fracture lengths vary between

466   0.05 m and 18.44 m (average = 0.94 m). Smaller fractures typically form arrays, whereas larger

467   fractures exhibit greater connectivity and are more commonly located near cross-faults (Figs

468   10c,11c, d). The highest fracture intensities coincide with areas containing a greater proportion

469   of through-going fractures (i.e., fractures that utilize pre-existing fractures in the network) and

470   cross-faults (Fig 11a). Network connectivity is also high, with a percentage of connected branches

471   (Pc) of 85%, with the majority of nodes classified as Y-nodes (50.3%), followed by I-nodes (34.9%)

472   and X-nodes (14.7%) (Fig. 11d). Connectivity, however, displays significant variations across the

473   study area (Fig 11d). In general, fracture intensity and connectivity are higher towards the eastern

474   portion of the study area, where small offset cross-faults are more frequent, while the western

475   portion exhibits lower intensities, except in the vicinity of large offset cross-faults. These trends

476   indicate a strong spatial correlation between fracture network properties and cross-faults.

477

a)

Splay
fault

Lower intensity
towards the
west of the
study area

Increased
fracture
intensity near
major cross-
faults

Lower
intensisty away
from large
cross faults

Increased
fracture
intensity near
major cross-
faults

Increased
fracture
intensity to the
east of the
study area

0
25
50 m

Fracture intensity (f/m)

b)

Arrays of small
trace length
fractures

Isolated small
trace length
fractures

Localised
zones of
fractures
around
mapped cross
faults

Increased
fracture
intensity and
connectivity
near large
fractures
(cross faults)

c

d

Cross-fault
damage zone

Cross-fault
damage zone

0
2.5
5 m

e)

i-node

449250 449650

Easting (m)

x-node

y-node

478

479 **Figure 11: Fracture properties across the Maghlaq Fault. a) Fracture intensity (P10) map across**

480 **the MF. b) Trend of all mapped on-fault fractures. c-d) Insets of the digitized network**

481 **highlighting key features discussed in the main text. e) Node triangle showing the distribution**

482 **of node classifications across the study area.**

483 On-fault fracture properties display both gradual variations over distances of tens of meters and

484 abrupt step-changes at cross-faults with > 10 m cumulative throw (**Fig 12**). Gradual changes

485 correspond to changes in fault plane geometry at corrugations (**Fig 12**). This relationship is

486 observed at feature 'CA' in **Figure 12**, where a corrugation with a 30˚ change in dip direction over

487 a 40 m along-strike distance coincides with a localized increase with fault dip. On-fault fractures

488 within the corrugation show a change in orientation from ESE to SE before returning to ESE,

489 fracture intensity decreases from 2 f/m to 1 f/m, while network connectivity increases at the

490 boundary of the corrugation, before decreasing in the trough (**Fig 12**). Similar trends are observed

491 across the study area, and in the detailed field photographs (**Fig 10c**). Additionally, the effect of

492 fault plane geometry on fracture network properties is evidenced where the splay fault intersects

493 the fault scarp (Fig 5a, 12a), where a bimodal distribution of on-fault fracture orientation

494 (Fig.12d), and an increase in fracture intensity and connectivity (Fig. 12e-f), is observed as the

495 fault segments link (**Fig 12**). Overall, corrugation ridges appear to cause an increase in fracture

496 intensity and connectivity when compared to corrugation troughs, and the orientation of on-fault

497 fractures is related to the orientation of the fault plane (**Fig 10c, 12a-d**).
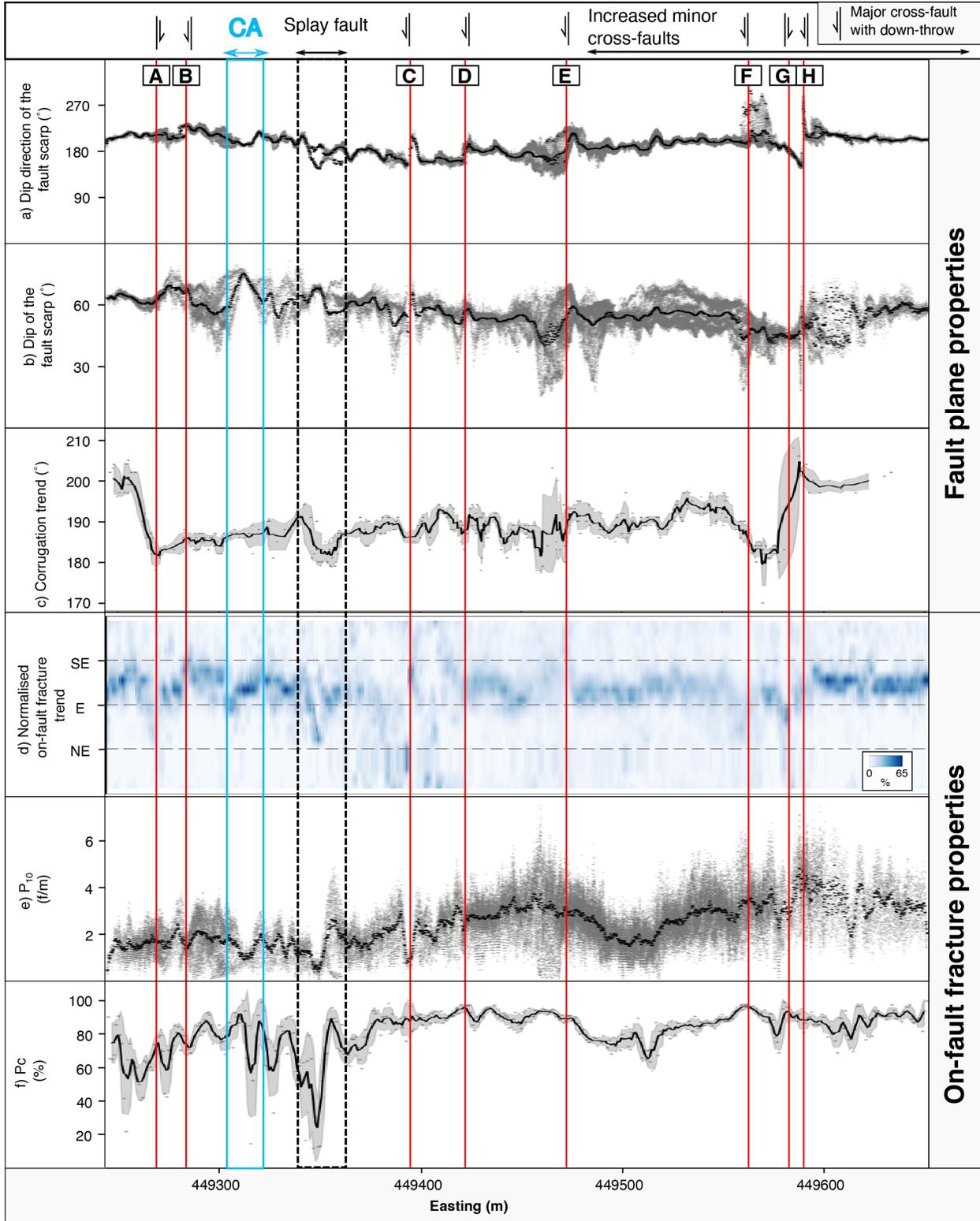
499 **Figure 12: Along-strike trends in fault plane and on fault fracture properties across the Maghlaq**

500 **Fault, with the location of major cross-faults indicated with red vertical lines. a) Dip-direction**

501 **of the Maghlaq Fault. b) Dip- of the Maghlaq Fault. For panels (a) and (b), the grey dots show**

502 **individual sampled values, whilst the black line shows the average for a given easting value. c)**

503 **Corrugation trend of mapped corrugations, showing all mapped corrugations (grey dots),**

504 **rolling average (n = 5), and standard deviation (grey shade). d) Heat map showing the**

505 **normalized on-fault fracture trend by 5 m along-strike bins. e) Fracture intensity (P10), showing**

506 **each datapoint (grey) and a 5 m wide rolling average (black). f) Percentage of connected nodes**

507 **showing all sample circles (grey dots), a rolling average (n = 5), and standard deviation (grey**

508 **shade).**

509 Properties of on-fault fractures are strongly influenced by the presence of cross-faults. Mapping

510 of cross-faults identified a series of cross-fault zones (A-H), which correspond to changes in on-

511 fault fracture properties (Figs. 6d-f, 12). Step-changes in fracture trend, intensity, and

512 connectivity are frequently observed at these locations (Fig. 12d-f). For example, on-fault

513 fracture trend changes from 116˚ in the footwall of cross-fault zone H to 096˚ in the footwall of

514 cross-fault zone G (Fig. 12d). This aligns closely with the 23˚ rotation in fault-plane orientation

515 across the same faults (Fig. 12a), suggesting that some on-fault fractures must predate cross-

516 faults and have been block-rotated during faulting. Compared to on-fault fracture orientation,

517 fracture intensity (Fig. 12e) and connectivity (Fig. 12f) exhibit more variable responses to cross-

518 faults. In some cases, step changes occur immediately adjacent to cross-fault zones (e.g., cross-

519 fault zones G, H), whereas in others, a more gradual increase is observed approaching the fault

520 zones (cross-fault zones A, D, E). A decrease in fracture intensity and connectivity is recorded

521   across cross-fault zones B and C, but this appears to be caused by preferential erosion of the fault

522   plane at these locations, which reduces the image resolution close to those cross-fault zones.

523   Fracture intensity and connectivity are low in the western portion of the study area, where cross-

524   fault density and throw are relatively low (Fig11e-f). In the eastern portion, where small offset

525   cross-faults are abundant, fracture intensity and connectivity are high and show a greater

526   variability compared to elsewhere in the study area (Fig 11e-f). These observations suggest that

527   while cross-faults post-date some on-fault fractures, as evidenced by the rotation of the fracture

528   orientations, fractures continued to form and the fracture network developed after the cross-

529   faults developed, as shown by cross-faults modifying intensity (Fig 12e) and connectivity (Fig 12f)

530   patterns. This implies a progressive deformation history for the Maghlaq Fault Scarp, which is

531   explored in the discussion.

532   **4 Discussion**

533   4.1 Evolution of brittle deformation across the Victoria Lines and Maghlaq faults

534   We deduced a four-stage structural evolution of the studied fault scarps from the detailed

535   analysis of fault plane geometry, on-fault fracture mapping, and the identification of cross-faults

536   (Fig 13). The proposed structural evolution serves a as basis for the discussion on the spatial and

537   temporal evolution of energy dissipation through brittle deformation in Section 4.2
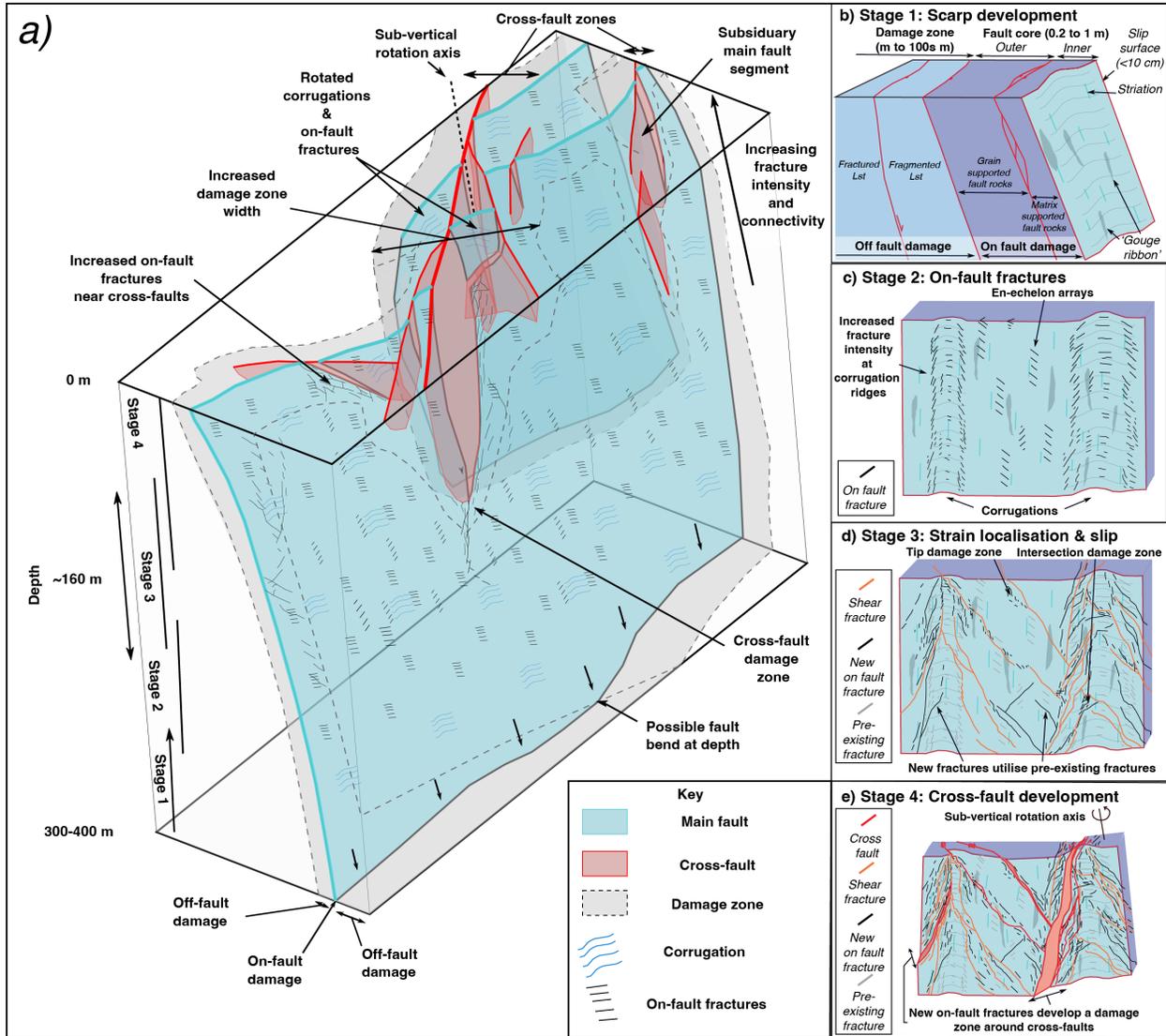
538

539   **Fig 13: Conceptual model outlining the key stages of brittle deformation: a) conceptual model**

540   **of deformation at depth based on the observed cross-cutting relationships; b) conceptual**

541   **model of limestone fault scarps after (Ferraro et al., 2018); c-e) conceptual evolutionary model**

542   **highlighting the progressive on-fault deformation of non-planar fault scarps during footwall**

543   **uplift, see text for discussion.**

544   **Stage 1: Initial fault scarp development and corrugation formation (Fig 13b).** Faulting

545   initiated on the Victoria Lines and Maghlaq faults during the late syn-rift phase (~7 to 5 Ma), with

546   the maximum burial of footwall sediments not exceeding 250 to 300 m (Bonson et al., 2007; Dart

547   et al., 1993; Martinelli et al., 2019; Villani et al., 2018). The Victoria Lines and Maghlaq faults have

548   well-developed bedrock fault scarps that preserve corrugations (Fig 5). As normal faults accrue

549   throw and deformation of the fault scarp, the footwall uplifts relative to the hanging wall

550   (Doutsos & Koukouvelas, 1998; Quye-Sawyer et al., 2021; Schlische & Anders, 1996; Stein et al.,

551   1988; Stein & Barrientos, 1985), with uplift to subsidence ratios for the Mediterranean commonly

552   1:1.6 to 1:2.3 (Doutsos & Koukouvelas, 1998; Quye-Sawyer et al., 2021; Stein et al., 1988).

553   Therefore, the fault scarp and footwall were progressively exhumed over time. Based on this

554   burial and tectonic history, fault scarp development and subsequent brittle deformation must

555   have occurred in the top 200 to 300 m of the crust.

556   Corrugation trends across the Victoria Lines Fault show convergent slip vectors, where

557   slip is nearly dip-slip in the faults center and becomes more oblique towards the tips (Fig 5). Such

558   patterns, recognized across many active normal faults (e.g., Jackson et al., 1982; Roberts 1996;

559   Roberts and Ganas, 2000; Roberts and Michetti, 2004; Michetti et al., 2000; Maerten, 2000), and

560   are used to distinguish kinematically isolated verses connected fault segments. Where not

561   influenced by cross-fault rotations, corrugation trends along the Maghlaq Fault remain close to

562   dip-parallel (i.e. , normal fault slip) with little to no evidence of convergence (Fig. 5b), implying

563   that the fault tips lie outside the study area. This aligns with Bonson et al. (2007), who interpret

564   the Maghlaq Fault as initially comprising WNW-ESE trending isolated segments that were later

565   linked by NE-SW trending branch faults, likely forming at a depth below 160 m. Corrugations on

566   the splay fault (Fig 5ai), and evidence of SSW-directed slip prior to cross-fault rotation, support a

567   protracted linkage history on the Maghlaq Fault. In contrast, corrugations across the Victoria

568   Lines Fault are consistent with the study area being an individual fault segment. This is supported

569   by throw distributions (fig 7b), which decrease eastward, and the segment tipping out onshore

570   (Fig 2b, c; Villani et al., 2018). Thus, although the faults share some features and both develop

571   cross-faults, the Maghlaq Fault has a more protracted linkage history, providing insights into the

572   role of this on subsequent brittle deformation.

573   **Stage 2: Early on-fault fracture formation (Fig 13c).** On-fault fractures across the Maghlaq

574   Fault cross-cut the fault scarp, so must post-date the formation of the fault scarp and

575   corrugations. Most small (<2 m) trace length on-fault fractures mapped across the Maghlaq Fault

576   have isolated nodes and form sinistral or dextral arrays, with the array trends close to the

577   observed corrugation trend (Fig 10, 11ai, aii). Abundant aligned en echelon fractures are

578   indicative of pure shear dominated transtension (Dewey et al., 1998; Fossen and Tickoff, 1998;

579   Sanderson and Marchini, 1984). In the case of the Maghlaq Fault, shear occurs down the

580   corrugation trend and a degree of extension occurs across the fault surface.

581   In addition to overall trends, changes in fracture orientation and array shear sense, coupled with

582   an increase in fracture intensity, are observed around geometrical complexities (**Fig 10c)** related

583   to corrugation ridges and troughs (Fig 12). Geometrical complexities are known to influence the

584   near-fault stress field (Chester and Chester, 2000) and the type and distribution of fault rocks

585   across the fault surface, and often refer to as "asperities" (Scholz, 2002; Chambon et al., 2006).

586   Large asperities, whether geometrical or rheological, commonly act as areas of increased stress

587   accumulation and causes earthquakes to nucleate at this location (Aki, 1984). Smaller

588   geometrical heterogeneities can directly influence the velocity and behavior of an earthquake

589   rupture, reducing an earthquakes rupture velocity and acting as a barrier to slip, limiting the size

590   and extent of ruptures (Li et al., 2023; Wesnousky, 2006; Zhang et al., 2023). This reduction in

591   rupture velocity can explain the local change in dominant array type across corrugation ridges

592   and troughs, whereby the bulk velocity decreases locally around the asperity relative to the

593   overall trends observed across the fault surface. Our observations suggest small geometrical

594   variations across the fault surface are sufficient to increase the degree of on-fault fracturing and

595   may locally perturb the velocity field across the fault surface during earthquakes.

596   **Stage 3: Strain localization and slip along on-fault fractures (Fig 13d).** Cross-cutting and abutting

597   relationships reveal that as a sub-section of the network grow, tip damage zones develop as on-

598   fault fractures from different en-echelon fractures interact (Fig 10ci). Tip damage zones (e.g., Kim

599   et al., 2004) suggest a minor amount of shear occurs along the fractures. These fractures form a

600   throughgoing shear fracture that displays abundant jogs at the intersection with pre-existing

601   fractures (Fig 10ci). As fracture networks evolve, strain commonly becomes localized onto a

602   subset of fractures, preserving blocks of early fractures that no longer accrue strain surrounded

603   by larger fractures that do (Cowie, 1998; Cowie et al., 2005; Segall & Pollard, 1983). The

604   orientation and spacing of the pre-existing fractures within a network is known to strongly

605   influence the properties (e.g., length, connectivity, orientation) of subsequent fractures (e.g.,

606 (Cruikshank et al., 1991; Peacock et al., 2018, 2023; Peacock & Sanderson, 2018). Across the

607 Maghlaq Fault, longer (>2 m) through-going fractures show this clearly, with rapid changes in

608 orientation present where through-going fractures intersect with isolated en-echelon fractures

609 (Fig 10a, b), causing localized (up to 2 m wide) increases in the length, connectivity, and intensity

610 of fractures across the Maghlaq Fault Scarp (Fig 11a, ai, aii). Similarly to isolated on-fault

611 fractures, the location of these localized zones of increased fracturing correlate with small

612 wavelength changes in fault plane geometry (Fig 12), suggesting geometrical variations also

613 control the location of strain localization across the fault scarp. Additionally, as fractures display

614 evidence of progressive growth over multiple earthquake cycles, it suggests geometrical

615 variations continue to act as asperities as the footwall is uplifted during shallow earthquake slip.

616 **Stage 4: Cross-fault development and rotation of the fault scarp (Fig 13e).** Numerous cross-

617 faults are observed across both studied fault scarps, and post-date and rotate the fault planes

618 and corrugations (Fig, 7, 8), as well as a subset of the on-fault fractures on the Maghlaq Fault

619 (Figs 12). Cross-faults are typically longer across the Victoria Lines Fault and have greater throw

620 than those across the Maghlaq Fault (Figure 7). This may be partly caused by the difference in

621 resolution between the DEMs of the Victoria Lines (pixel size = 0.13 m) and Maghlaq (pixel size =

622 0.06 m), leading to the smallest cross-faults across the Victoria Lines Fault to be below the

623 resolution of the dataset. Despite this, cross-faults across both datasets show high throw values

624   for their respective lengths and are overdisplaced when compared to global datasets of normal

625   faults (e.g., (Lathrop et al., 2022).

626   Suggestions for the formation of cross-faults include: i) strain compatibility associated

627   with relay ramp development (Walsh and Watterson, 1991; Long and Imber, 2021); ii) the

628   presence of weak fault rock along non-oriented fault segments (ref); iii) the presence of pre-

629   existing faults and/or structural inheritance (Phillips et al., XXX; Pizzi and Galadini, 2009). Options

630   (i) and (ii) require the presence of faults that are separate, and at a high angle to, the main fault.

631   Cross-faults in our dataset typically occur within 25˚ of the main faults strike and show a

632   progressive evolution after the main fault slipped in a different direction (Fig 4, 5), therefore this

633   observation implies (i) and (iii) are unlikely. The only remaining viable explanation, where cross-

634   faults develop because of strain incompatability during relay ramp development, could match

635   previous models of fault growth for the Maghlaq Fault. Using detailed fault zone mapping,

636   Bonson et al. (2007) suggest the Maghlaq Fault displays evidence of progressive smoothing and

637   linkage during the development of a breached relay ramp. The linkage of fault strands, as

638   evidenced by the presence of the splay fault (Fig. 5), may contribute to the increased cross-fault

639   intensity across the Maghlaq Fault compared to the Victoria Lines Fault. Contrary to the findings

640   of Bonson et al. (2007), however, we find that cross-faults act to rotate the fault scarp and

641   therefore increase the rugosity of the overall slip surface. Additionally, the evidence of

642   convergent slip vectors across the Victoria Lines Fault suggests that this fault slipped as an

643   isolated segment (e.g., Roberts et al., 1996), but still displays similar cross-fault patterns and

644   orientations to the Maghlaq Fault. Cross-faults rotating the fault plane and the elevated fracture

645   intensity around cross-faults therefore indicate that cross-faults begin to form at geometrical

646   heterogeneities, and the presence of cross-faults across both studied faults, indicates that cross-

647   faults likely form in response to stress rotations and irregular slip patterns that occur around

648   geometrical heterogeneity during shallow earthquakes (e.g., Marshall and Morris, 2012; Resor

649   and Meer, 2009).

650   **4.2: The effect of brittle deformation on energy dissipation during shallow earthquakes**
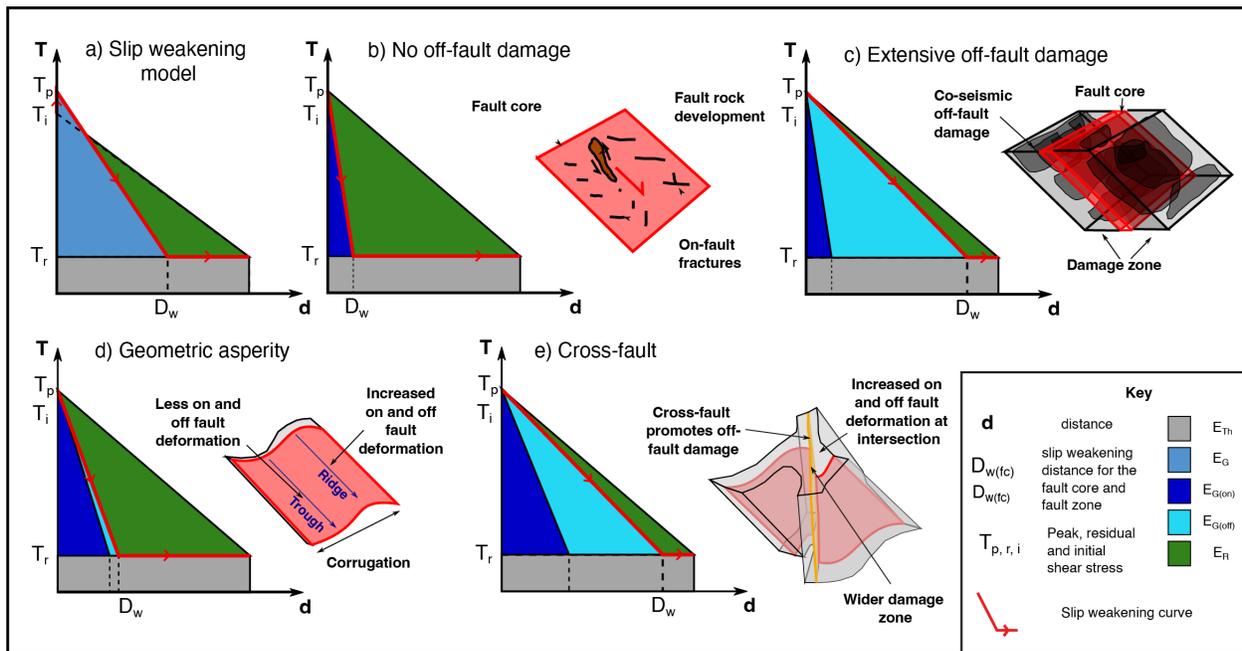
651

652   **Figure 14: Graphical representations of the earthquake energy budget for a range of on-fault**

653   **and near-fault deformation scenarios within a slip-weakening framework. Energy partitioning**

654   **is illustrated for: a) a case in which breakdown work consists only of E_G required for rupture**

655   **propagation (Kanamori and Rivera, 2006); b) deformation confined to the fault core, with on-**

656   **fault damage and negligible off-fault damage, as often observed in high velocity rock mechanics**

657   **experiments (De Paola et al., 2025); c) energy is dissipated within a faults damage zone (De**

658   **Paola et al., 2025); d) minor geometrical asperities, where damage is constrained to the fault**

659 **core prior to the development of throughgoing fractures and cross-faults; e) the vicinity of**

660 **cross-faults, where the magnitude of on- and near-fault damage increases. Where E_{Th} is the**

661 **energy dissipated within the fault zone through frictional heating, E_G is the fracture energy**

662 **required for rupture propagation, and can be split into energy dissipated within the fault core**

663 **(E_{Gon}) and energy dissipated within a faults damage zone (E_{Goff}), and E_R is the radiated energy**

664 **through seismic waves.**

665 Faults are heterogeneous, displaying geometrical, structural and lithological variations across the

666 fault surface that can directly affect earthquake rupture mechanics (e.g., Abercombie et al., 2006;

667 Faulkner et al., 2003; Mashoch et al., 2021; Coco et al., 2023; McKay et al., 2020). Our study

668 highlights that in addition to spatial heterogeneities, spatial-temporal variations occur in the type

669 and distribution of on-fault deformation, and therefore on-fault energy dissipation, in the top

670 few hundred meters of the crust across multiple earthquake cycles. The main components of an

671 earthquakes energy budget can be graphically represented on a slip-weakening curve (Fig 14),

672 whereby the total fracture energy (E_G) dissipated during earthquake rupture is equal to the work

673 done by the earthquake within the fault core (i.e., on-fault, **E_{Gon}**) and off-fault fracturing in the

674 surrounding damage zone (**E_{Goff}**). The remaining energy is radiated as seismic waves (**E_R**), or

675 through thermal heating (**E_H**) (Cocco et al., 2023; Kanamori & Rivera, 2006; Nielsen et al., 2016).

676 Our results show that on-fault damage varies spatially across the fault surface, as evidenced by

677 the range in observed fracture intensity (0.2 to >4 f/m, Fig. 12). The slip-weakening distance (Dw)

678 (i.e., the amount of slip required for the shear stress to decrease from the peak strength to the

679 residual dynamic strength) and total dissipated fracture energy (**E_G**) will also vary (**Fig 15ii**).

680   Fracture intensity decreases where the fault plane is relatively planar (e.g., between cross-faults

681   E and F in Fig 12). This suggests low on-fault energy dissipation (Fig 1b), enabling an increase in

682   radiated energy, surface ruptures, and surface shaking. In our data, fracture intensity was

683   observed to increase around geometrical heterogeneities, in particular at the edge of asperities

684   (e.g., ridges of corrugations), suggesting increased on-fault energy dissipation at these locations

685   (Fig 9, 11). Additionally, the prevalence of fractures that are large enough to propagate away

686   from the fault surface propably promote an increase in off-fault damage, although this was not

687   possible to quantify with the available datasets. In this case, Dw will be greater than in the case

688   of a planar fault surface, and a lower proportion of the energy will be radiated as seismic waves

689   (Fig 15iv).

690   The structural model (Fig. 14) suggests that on-fault fractures progressively develop during upper

691   few 100 metres of the crust during earthquake slip (Figure 12a), with the observed network

692   providing a final snapshot of on-fault deformation across multiple seismic cycles. To gain a better

693   understanding of the evolution of energy dissipation through time, and to gain insights into

694   different energy sinks during shallow earthquake slip, findings need to be considered in relation

695   to the presented structural model. Prior to fracture development (Stage 1), on-fault energy

696   dissipation occurred through the development of fault rocks that form a corrugated fault core.

697   Following this (Stage 2), during early fracture development, short trace length (<1 m) on-fault

698   fractures formed as arrays with their orientation influenced by minor asperities on the fault

699   surface. As these fractures are small, their contribution to the overall earthquake energy budget

700   will be limited, providing a minor energy sink across the whole fault surface. As strain localized

701   onto a sub-set of the fractures (Stage 3), the length of on-fault fractures increased, and new on-

702   fault fractures evolved as damage zones around the sub-set of fractures. This increase in fracture

703   length will increase the amount of fracture energy dissipated within the fault core and immediate

704   damage zone. Additionally, the energy sink caused by these fracture zones is localized at

705   geometrical heterogeneities, such as corrugation ridges, increasing the likelihood that these

706   features influenced earthquake rupture patterns. As cross-faults develop (Stage 4), aseismic slip

707   will be be transferred from the main fault into the damage zone, increasing off-fault deformation

708   (Fig 8, 11). Both these factors will increase the slip-weakening distance (Dw). Additionally, a

709   portion of earthquake energy will contribute to the rotation of the main fault surface, which will

710   also alter the relative angle of the main fault relative to the slip-direction creating a geometrical

711   barrier to further slip.
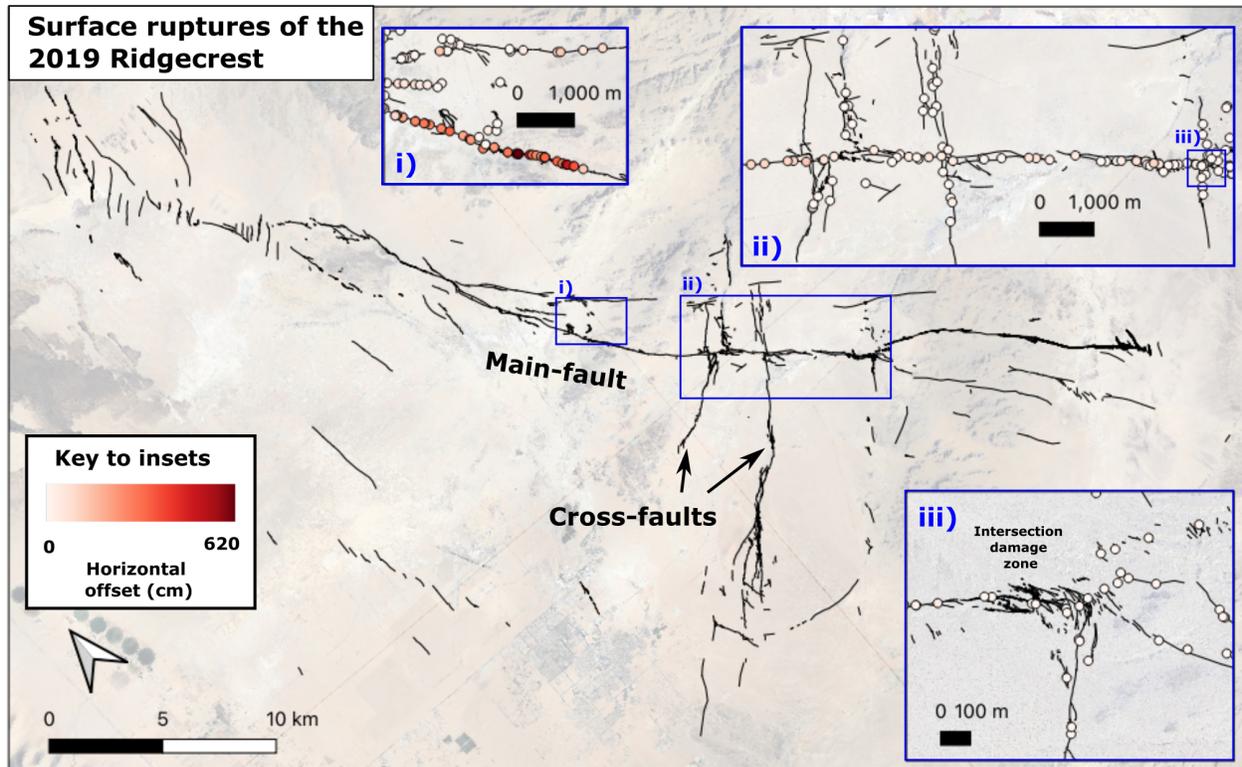
712

713   **Figure 15: Surface rupture and horizontal offset of the Ridgecrest Earthquake. Insets highlight**

714   **how the horizontal offset varies when segments of the fault with no surface ruptures (i) are**

715   **compared to where cross-faults cross (ii) or intersect with (iii) the main fault strand.**

716   Geometrical and/or rheological asperities strongly influence earthquake slip distributions (e.g.,

717   Lay et al., 1982). Faults are known to smooth as continued slip erodes asperities and slip becomes

718   more energy efficient (Sagy & Brodsky, 2009; Kirkpatrick and Brodsky, 2014; Hayward et al., 2019;

719   Imber et al., 2008). Most of these studies focus, however, on faults exhumed from depth and

720   give little information about slip in the shallow surface. Our work has shown that on-fault

721   fractures and cross-faults act to increase the geometrical complexity of the main fault through

722   time, with cross-faults being a type of geometric barrier, which can lead to slip being partitioned

723   into the damage zone. For example, during the 2019 Ridgecrest Earthquake (REF), cross-faults

724   caused an increase in the intensity of surface ruptures at the intersection of the cross-fault and

725   main fault (Fig 16ii, iii), as well as these areas having lower slip values compared to portions of

726   the fault where no cross-faults are present (Fig 16i). These are similar observations to the

727   patterns of cross faults observed in our study area, suggesting cross-faults, and their associated

728   damage zones, can have a significant impact on damaging surface rupturing earthquakes. Our

729   work highlights that the distribution of on-fault energy sinks in the near surface are spatially and

730   temporally variable, with cross-faults having a strong effect on the distribution of surface

731   ruptures and increased off-fault damage in the shallow sub-surface (Fig 15c). We therefore

732   suggest that the further study of on-fault fracture networks in natural and laboratory examples

733   can provide valuable insights into the shallow earthquake energy budget, how this evolves

734   through space and time, and factors that influence the distribution of surface ruptures. Studying

735   brittle surface deformation around faults provides useful constraints for seismic hazard

736   assessment. Variations in off-fault damage and fault-zone structure affects the energy available

737   for rupture propagation, thereby influences the maximum earthquake size and the likelihood of

738   multi-segment ruptures. Moreover, heterogeneous slip distributions and earthquake sized are

739   linked to more irregulat recurrence intervals (e.g. Rodriguez Piceda et al, 2025a,b). Finally, more

740   accurate ground motion preduction equations could account for fault-zone properties, rather

741 than relying solely on magnitude, distance and soil properties as is commonly done (e.g. (Bommer

742 et al., 2010) Bommer et al., 2010).

743 **5 Conclusions**

744 Our study highlights the importance of progressive brittle deformation in controlling the

745 dissipation of energy along fault planes during shallow earthquake slip. We have identified a four-

746 stage deformation model that explains the progressive development of brittle structures in the

747 upper few hundred meters of the crust by mapping fault plane geometry, on-fault fracture

748 networks, and cross-faults along the Maghlaq and Victoria Lines faults. Our findings demonstrate

749 that on-fault fractures form early but are subsequently modified by cross-faulting, with increased

750 fracture intensity and connectivity near major cross-faults. This suggests that energy dissipation

751 is spatially heterogeneous, with structures such as corrugations and cross-faults causing an

752 increase in the intensity and connectivity of on-fault fractures, and causing the partitioning the

753 seismic energy at different locations on the fault surface over scales of tens of meters. Our results

754 provide insights into how energy dissipation evolves within active fault zones, which has direct

755 implications for seismic hazard assessment.

756 **Acknowledgments**

758 **Open Research**

759 Spatial analysis was undertaken using Version 3.22 (2022) of QGIS, available from

760 https://qgis.org/download/

761 **Conflict of Interest Disclosure**

762 "The authors declare there are no conflicts of interest for this manuscript."

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