

# Lubrication of dislocation glide in forsterite by Mg vacancies: insights from Peierls-Nabarro modeling

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## 6 **Abstract**

7 Dislocation glide is an important contributor to the rheology of olivine under conditions of high stress  
8 and low to moderate temperature, such as occur in mantle wedges. Interactions between point defects  
9 and dislocation core may alter the Peierls stress,  $\sigma_p$ , and has been suggested that vacancy-related  
10 defects may selectively enhance glide on certain slip systems, changing the olivine deformation fabric.  
11 In this study, the Peierls-Nabarro model, parameterized by generalized stacking fault (GSF) energies  
12 calculated atomistically using empirical interatomic potentials, is used to determine the effect of bare  
13 Mg vacancies on the Peierls stresses of [100](010) and [001](010) dislocations in forsterite. Mg  
14 vacancies considerably reduce GSF energies and, consequently,  $\sigma_p$  for dislocations gliding on (010) in  
15 olivine. The magnitude of this decrease depends strongly on dislocation and the type of the lattice site,  
16 with vacant M2 sites producing the largest reduction of  $\sigma_p$ . The [001](010) slip system is found to be  
17 more sensitive than the [100](010) slip system to the presence of vacancies. Although, at ambient  
18 pressure,  $\sigma_p$  is lower for [100](010) than [001](010) edge dislocations,  $d\sigma_p/dP$  is greater for [100](010)  
19 dislocations, resulting in a change in the preferred slip system at 1.5 GPa. By preferentially lubricating  
20 [001](010) glide, Mg vacancies reduce the pressure at which this cross-over occurs. An M2 vacancy

21 concentration at the glide plane of 0.125 defects/site is sufficient to reduce cross-over to 0.7 GPa. This  
22 may account for the existence of the B-type olivine deformation fabric in the corners of mantle wedges.

23 **Keywords:** Forsterite; dislocation; Peierls stress; cation vacancies; atomic-scale modeling

## 24 **1. Introduction**

25 Forsterite-rich olivine is the dominant component of the Earth's upper mantle, composing ~60-70% of  
26 its bulk by volume. Olivine is also the weakest major phase in this region of the Earth's interior, and  
27 accordingly controls its rheology. Interpreting seismological models of the Earth's upper mantle in  
28 terms of the dynamics of plastic flow thus requires an intimate understanding of the atomic-scale  
29 mechanisms that contribute to the deformation of olivine. A number of mechanisms contribute to the  
30 rheology of olivine under mantle conditions, such as grain boundary diffusion (e.g. Mei and Kohlstedt  
31 2000); grain boundary sliding, accommodated by either elastic and diffusional relaxation of the grain  
32 boundary (Jackson et al. 2014) or dislocation-enabled deformation of individual grains (e.g. Hansen et  
33 al. 2011; Hansen et al. 2012ab); dislocation climb (e.g. Goetze and Kohlstedt 1973); and dislocation  
34 glide (e.g. Evans and Goetze 1979; Katayama and Karato 2008).

35 Dislocations play a particularly important role in the deformation of olivine under low temperature or  
36 moderate to high stress conditions, such as in the mantle wedge above a subducting slab. The strain rate  
37 in this creep regime is controlled by the rate of kink-pair nucleation, which depends on the elasticity of  
38 the crystal and the intrinsic lattice friction of the glide plane. This latter property is quantified by the  
39 Peierls stress,  $\sigma_p$ , the critical stress required to initiate free glide of a dislocation at 0 K. Defect  
40 chemistry can therefore exert an important influence if immobile impurities segregate strongly to  
41 dislocation cores and inhibit glide creep by pinning dislocations, a phenomenon called solute drag  
42 (Cottrell and Bilby 1949).

43 Olivine is a plastically anisotropic mineral, and develops a measurable lattice preferred orientation  
44 (LPO) when deformed in the dislocation glide-controlled creep regime (e.g. Nicolas and Christensen  
45 1987; Mainprice 2007; Long and Silver 2009; Long and Becker 2010). The dominant slip system for  
46 dislocation creep at low pressure is [100](010). Pressure changes the preferred dominant slip system of  
47 olivine, and much of the variation in seismic wave anisotropy in the upper mantle is explicable in terms  
48 of this pressure-induced transition (Mainprice et al. 2005; Ohuchi et al. 2011; Raterron et al. 2016). At  
49 high pressure, the dominant slip system is [001](010), as this slip system hardens less in response to  
50 increasing pressure than the [100](010) slip system (Raterron et al. 2011; Hilairet et al. 2012).

51 Experimental studies have reported Peierls stresses for olivine, ranging widely from as little as 3.8 GPa  
52 (Idrissi et al. 2016) to ~15 GPa (Demouchy et al. 2013), although typical values are in the range 5-10  
53 GPa (e.g. Evans and Goetze 1979; Kranjc et al. 2016; Proietti et al. 2016). The Peierls stress measured  
54 in these studies represents a weighted average of the Peierls stresses of the individual slip systems that  
55 contribute to the total strain. While  $\sigma_p$  has not been measured experimentally for any individual slip  
56 system in olivine, this information gap has been partially filled using atomic scale modeling.  
57 Mahendran et al. (2017) directly calculated  $\sigma_p$  for [001] and [100] screw dislocations gliding on several  
58 crystallographic planes. Consistent with experimental observations, they found that [100](010) glide is  
59 easier than [001](010) glide in forsterite, with Peierls stresses of 3.1 and 7.2 GPa, respectively, within  
60 the range of Peierls stresses reported from experiments.

61 Hydrogen, probably incorporated into the olivine lattice by protonation of cation vacancies (e.g. Martin  
62 and Donnay 1972; Bai and Kohlstedt 1993; Kohlstedt et al. 1996; Lemaire et al. 2004), can influence  
63 the rheology of olivine under conditions of low-temperature (<1273 K) and high-stress, reducing  $\sigma_p$  to  
64 1.6-2.9 GPa (Katayama and Karato 2008). Similarly, the incorporation of water as the  
65 "titanoclinohumite" defect, consisting of charge-coupled  $\{Ti_M\}''$  and  $\{2H_{Si}\}''$  defects apparently results  
66 in a systematic increase in strain rate (Berry et al. 2005; Faul et al. 2016).

67 Protonated cation vacancies also change the preferred slip system in olivine. Ohuchi et al. (2012)  
68 investigated the effect of increasing water content on the dominant slip system, and observed that the  
69 [001](010) slip was preferred over [100](010) slip at moderate water contents, with the transition  
70 occurring at ~650 ppm H/Si. Natural olivine crystals in subduction zone-derived peridotites show this  
71 B-type fabric (Mizukami et al. 2004), and 'wet' olivine fabrics have also been found in some peridotites  
72 sourced from the deep upper mantle (Katayama et al. 2005). Such a change in preferred slip system  
73 may explain the existence of trench-parallel orientation of 'fast' shear-wave polarizations in mantle  
74 wedges above subduction zones (Margheriti et al. 1996; Smith et al. 2001), which has previously been  
75 attributed to trench-parallel flow. A water-induced change of deformation fabric implies that protonated  
76 vacancies have a differential effect on the critical resolved shear stresses (CRSS) of the dominant slip  
77 systems, whether by stabilizing dislocation kinks or reducing the Peierls stress.

78 Atomic-scale calculations show that the energy required to create an Mg vacancy is considerably less  
79 than that required to create a Si vacancy (Wright and Catlow 1994; Brodholt 1997). Similarly,  
80 protonated Mg vacancies have lower energies than protonated Si vacancies (Walker et al. 2007). In the  
81 Earth's upper mantle, which is MgSiO<sub>3</sub>-saturated, M-site vacancies in olivine can thus be expected to  
82 be many orders of magnitude more abundant than protonated Si vacancies. The predominance of M site  
83 vacancies in mantle olivine could be reinforced in oxidized regions of the mantle, as the oxidation of  
84 Fe<sup>2+</sup> to Fe<sup>3+</sup> may be accompanied by the creation of M-site vacancies to maintain charge neutrality. The  
85 two mechanisms for creating mobile M-site vacancies may co-exist, particularly in mantle wedges  
86 above subduction zones, which are oxidized by fluid transfer from the subducting slab (Kelley and  
87 Cottrell 2009). High oxygen fugacity has been shown to enhance seismic attenuation in olivine (Cline  
88 et al. 2018). The similarly high oxygen fugacities present in the earlier deformation experiments of  
89 Faul et al. (2016), produced by the platinum sleeve used to wrap the specimen (Faul et al. 2017), may

90 indicate that oxidizing conditions may indicate that ferric-iron associated vacancies may have a  
91 comparable effect to water on the rheology of olivine.

92 In this study we use computational simulations to investigate the possibility that cation vacancies may  
93 be able to lubricate dislocation glide in forsterite by reducing the Peierls stress  $\sigma_p$ , providing a  
94 mechanism by which point defects may influence the rheology of this important mantle mineral in the  
95 glide-controlled creep regime. Core structures and Peierls stresses are calculated for important slip  
96 systems in forsterite using a semi-discrete Peierls-Nabarro (PN; Peierls 1940; Nabarro 1947) model. In  
97 the PN model, a dislocation is represented as a discrete array of partial dislocations distributed on the  
98 glide plane, with the balance between repulsive elastic forces and inelastic restoring forces determining  
99 the shape of the dislocation. The inelastic component of the dislocation energy is parametrized here  
100 using atomistic calculations of generalized stacking fault (GSF; Christian and Vitek 1970) energies on  
101 olivine slip planes, using empirical interatomic potentials to model interactions between ions. PN  
102 modeling has been used to investigate glide lubrication by interstitial hydrogen atoms (Lu et al. 2001)  
103 and vacant lattice sites (Lu and Kaxiras 2002) in fcc Al, and  $\{2H_{Mg}\}^x$  defects in MgO (Skelton and  
104 Walker 2018). We consider only cation vacancies on the M1 and M2 sub-lattices, represented using  
105 Kröger-Vink notation (Kröger and Vink 1956) as  $\{V_{M1}\}''$  as  $\{V_{M2}\}''$ , such as may be created in mantle  
106 olivine by oxidation of  $Fe^{2+}$  to  $Fe^{3+}$  at  $fO_2$  or through incorporation of water under hydrous conditions.

## 107 2. Methods

108 In the PN model, a dislocations with finite core-width is represented as continuous (e.g. Joós et al.  
109 1994) or discrete (e.g. Bulatov and Kaxiras 1997) distribution of dislocation density along the glide  
110 plane. Here we use an implementation of the discrete approach previously used for modeling  
111 lubrication of dislocation glide in MgO (Skelton and Walker 2018). The total energy of such a finite  
112 distribution of dislocations at a distance  $R$  from the dislocation line is

$$113 \quad E_{TOT}(R) = E_{INTERNAL} + E_{LONG-RANGED}(R) \quad (1)$$

114 where  $E_{LONG-RANGED}$  is the elastic strain energy of the dislocation contained within radius  $R$  and  $E_{INTERNAL}$   
 115 is the internal energy of the dislocation distribution. This can be further decomposed as

$$116 \quad E_{INTERNAL} = E_{ELASTIC} + E_{MISFIT} \quad (2)$$

117 where  $E_{ELASTIC}$  is the energy due to elastic interactions between the components of the dislocation  
 118 density and  $E_{MISFIT}$  is the inelastic energy due to the lattice discontinuity across the glide plane. If  $\mathbf{u}(x)$  is  
 119 the disregistry across the slip plane and  $\rho(x) = du_i(x)/dx$  is the associated dislocation density  
 120 distribution, then the elastic energy of the dislocation is the work required to insert this disregistry into  
 121 an infinite elastic medium:

$$122 \quad E_{ELASTIC}[\rho(x)] = -K \iint \rho(x') \rho(x) \ln|x - x'| dx' dx \quad , \quad (3)$$

123 where  $K$  is the elastic energy pre-factor, derived from the dislocation geometry and the elastic  
 124 constants, which are here calculated directly from the analytical derivatives of the total cell energy. In  
 125 the absence of a restoring force, the repelling elastic force between components of the dislocation  
 126 density distribution would cause the dislocation to have an infinitely wide core, with  $\rho$  zero  
 127 everywhere. In real crystals, it is the energy penalty associated with introducing misfit on either side of  
 128 the slip plane that provides this opposing force, and constrains dislocations to have finite width. For a  
 129 given disregistry profile  $\mathbf{u}(x)$ , the inelastic energy is

$$130 \quad E_{MISFIT} = \sum_n \gamma(u(na_p)) a_p \quad , \quad (4)$$

131 where  $a_p$  is the spacing between adjacent atomic planes and the function  $\gamma(\mathbf{u})$  (called the  $\gamma$ -line in 1D)  
 132 gives the energy required to displace one half of a crystal by  $\mathbf{u}$  (Christian and Vitek 1970). The  
 133 equilibrium dislocation core structure is calculated by minimizing equation (2), under the constraint  
 134 that the integral of the disregistry  $\mathbf{u}$  equals the Burgers vector.

135 The evolution of the disregistry profile under the action of an applied stress  $\sigma$  is computed by adding  
136  $E_{WORK} = -\sigma \int u(x) dx$  to the total internal energy (equation 2) of the dislocation and minimizing the  
137 energy functional as before. At the Peierls stress,  $\sigma_p$ , the energy barrier inhibiting free translation of the  
138 dislocation disappears, allowing it to glide without resistance. This is equivalent to searching for the  
139 stress  $\sigma$  at which the total energy function has no energy minimum, so that  $\sigma_p$  corresponds to the  
140 minimum stress for which the energy minimization step fails (Bulatov and Kaxiras 1997).

141 The  $\gamma$ -surface energy used to calculate the inelastic energy of a dislocation density distribution is  
142 constructed from generalized stacking fault (GSF) energies obtained from atomistic calculations. In an  
143 atomistic simulation, a GSF is inserted by cutting a simulation cell with the appropriate orientation, and  
144 displacing one half with respect to the other by  $\mathbf{u}$ , then allowing the atomic coordinates to relax to a  
145 local minimum energy configuration. For materials without rigid units, such as simple oxides, atomic  
146 positions are typically constrained to relax normal to the fault plane. However, olivine contains  $\text{SiO}_4$   
147 polyhedra, which accommodate shear strain primarily through rotation rather than deformation. To  
148 replicate this, in our simulations oxygen atoms have been permitted to relax freely while Mg and Si  
149 atoms can only move normal to the stacking fault plane. In all GSF calculations described here, a 15 Å  
150 thick vacuum layer was used to minimize interactions between the GSF and its periodic images. To  
151 ensure that the boundary conditions of the relaxation calculation match the bulk material, the  
152 coordinates of all atoms within 5 Å of the vacuum layer were held fixed. Calculated GSF energies  
153 depend on the thickness of the slab atoms used in the simulation. For stacking faults on (010), a slab  
154 thickness of  $8b$ , where  $b$  is the length of the [010] lattice vector, was found to be sufficient to converge  
155 the [100](010) and [001](010)  $\gamma$ -line maxima to  $< 5 \text{ meV}/\text{\AA}^2$ .

156 In this study, owing to the large number of atoms in the simulation cells, atomic scale interactions are  
157 modeled using empirical interatomic potentials. These potentials are taken from a widely used potential

158 model, which was parameterized by fitting to experimental data (Sanders et al. 1984; Lewis and Catlow  
159 1985), and reproduces the physical properties of forsterite reasonably well (Price et al. 1987). This  
160 potential, labeled THB1, has been widely used to model point and extended defects in forsterite,  
161 including Mg point defects (Wright and Catlow 1994), surface structures and energetics (de Leeuw et  
162 al. 2000), and screw dislocation core structures and energies (Walker et al. 2005). The THB1 model  
163 uses formal charges for the  $\text{Mg}^{2+}$  and  $\text{Si}^{4+}$  cations, while the polarizable oxygen anion is modeled as a  
164 positively charged core coupled by a harmonic potential to a negatively-charged massless shell (Dick  
165 and Overhauser 1958). Each cation-anion pair interacts through a Buckingham potential, while the  
166 rigidity of the  $(\text{SiO}_4)^{4-}$  tetrahedron is replicated using a short-ranged three-body harmonic potential. All  
167 atomistic calculations are performed using the molecular mechanics software GULP (Gale 1997; Gale  
168 and Rohl 2003).

169 The two defect species considered in this study,  $\{V_{M1}\}''$  and  $\{V_{M2}\}''$ , are charged, and inserting them  
170 into a simulation cell gives it a net charge. This is compensated for by applying a charge-neutralizing  
171 background. Creating a vacancy in the simulation cell is simple, and involves removing from the  
172 simulation cell a single atom of the desired type located a specified distance from the stacking fault  
173 plane. The slab thicknesses used to calculate GSF energies in the absence of Mg vacancies are  
174 sufficiently great to also ensure convergence of GSF energies with point defects present at the stacking  
175 fault plane. To minimize interactions between point defects and their periodic images, simulations are  
176 conducted by inserting a single Mg vacancy at the slip plane of a simulation cell with cross-section  
177  $2a \times 2c$ , which corresponds to a concentration of 0.125 defects/site at the stacking fault plane for both  
178 the M1 and M2 sub-lattices. While this would represent an unrealistically high defect concentration in  
179 the bulk lattice, the strain magnitude close to a dislocation core is substantial, so that the energy of an  
180 Mg vacancy on a core site may be considerably lower than in the bulk, and the defect concentration



181 correspondingly greater. In the case of MgO, fully atomistic calculations of Mg vacancy segregation to  
182  $1/2\langle 110 \rangle\{110\}$  edge dislocations give binding energy of up to 1.7 eV (Zhang et al. 2010).

### 183 **3. Results**

#### 184 *3.1 GSF energies in point defect free forsterite*

185 In materials with complex lattices, like olivine, there may be several inequivalent planes on which slip  
186 can occur. We find that the lowest energy (010) stacking fault plane is located at  $z = 0.25$  (or,  
187 equivalently,  $z = 0.75$ ), intersecting the sheet of  $M_2O_6$  octahedra (Fig. 1a). The THB1 potential model  
188 has previously been validated by Mahendran et al. (2017) at 0 GPa, who compared [100](010) and  
189 [001](010)  $\gamma$ -lines calculated with the interatomic potential with the earlier DFT calculations by  
190 Durinck et al. (2005). Although GSF energies calculated with THB1 are qualitatively consistent with  
191 the DFT calculations, the energies calculated with the interatomic potential are somewhat higher,  
192 particularly for displacement along [001](010). Indeed, the magnitude of the difference may actually be  
193 understated, as Durinck et al. (2005) used atomic slabs only a single unit cell thick along the stacking  
194 fault normal, so that their *ab initio* computed GSF energies are unlikely to be fully converged with  
195 respect to slab thickness.

196 The difference between GSF energies obtained in *ab initio* and force field calculations may be due to  
197 the use of formal charges in the THB1 model. The introduction of a GSF into a simulation cell often  
198 brings ions at or near the stacking fault plane into close proximity with ions of like charge, so that  
199 partially ionic models should in general predict lower energies than those that use formal charges. The  
200 ionicity  $\zeta$  of forsterite is calculated to be  $< 1$  from refinements of electron density distributions obtained  
201 using theoretical calculations (Liu et al. 2009) and synchrotron x-ray diffraction (Kirfel et al. 2005). It  
202 is thus unsurprising that Durinck et al. (2005) find lower GSF energies than calculations that use the  
203 THB1 potential.

204 In our calculations, we find that in point defect-free forsterite the maximum energy along the [100]  
205 (010)  $\gamma$ -line, which is the energy of a  $1/2[100](010)$  GSF, increases monotonically with pressure from  
206  $0.140 \text{ eV}/\text{\AA}^2$  at 0 GPa to  $0.203 \text{ eV}/\text{\AA}^2$  at 10 GPa (Fig. 2). By comparison, the energy of the  $1/2[001]$   
207 (010) GSF, which is the  $[001](010)$   $\gamma$ -line maximum, decreases modestly over the same pressure range,  
208 from  $0.163 \text{ eV}/\text{\AA}^2$  at 0 GPa to  $0.142 \text{ eV}/\text{\AA}^2$  at 10 GPa. This contrasts with DFT calculations (Durinck et  
209 al. 2005), which predict a pressure-independent  $[001](010)$   $\gamma$ -line maximum energy, although this is  
210 likely a consequence of the different simulation parameters, including simulation cell size and  
211 relaxation constraints. While, at ambient pressure, the energy of the  $[100](010)$   $\gamma$ -line maximum is  
212 lower than that of the  $[001](010)$   $\gamma$ -line maximum, the qualitatively different responses of the two  $\gamma$ -  
213 lines to applied pressure mean that the  $[001](010)$   $\gamma$ -line maximum is of lower energy above  $\sim 3$  GPa.  
214 The higher GSF energies reported here are partly due to the stricter relaxation constraints, particularly  
215 for Mg atoms, which were allowed to relax freely in previous studies (Durinck et al. 2005; Mahendran  
216 et al. 2017) but are here constrained to relax normal to the stacking fault plane.

### 217 *3.2 Influence of Mg vacancies on GSF energies*

218 In the mantle, where olivine coexists with  $(\text{Mg,Fe})\text{SiO}_3$  pyroxene, Mg vacancies are expected to be  
219 more abundant than Si vacancies. In the bulk lattice, M1 vacancies have lower energies than M2  
220 vacancies, and are thus more abundant (Brodholt 1997). Any change in the GSF energy will depend on  
221 the interaction between the Mg vacancy and the stacking fault, which can be expected to vary between  
222 the two sites. Considering first slip in the [100] direction, we find that the energy of the  $1/2[100](010)$   
223 GSF with a  $\{\text{V}_{\text{M1}}\}''$  defect adsorbed to the stacking fault plane is  $0.132 \text{ eV}/\text{\AA}^2$  at 0 GPa, 5.7% lower  
224 than the value computed without point defects. The  $1/2[100](010)$  GSF energy increases with pressure  
225 (Fig. 3), reaching  $0.198 \text{ eV}/\text{\AA}^2$  at 10 GPa, 2.5% less than the point defect-free value.  $\{\text{V}_{\text{M2}}\}''$  defects  
226 have a more substantial impact on the calculated  $1/2[100](010)$  GSF energy, reducing it to  $0.121 \text{ eV}/\text{\AA}^2$

227 at 0 GPa, 13.6% lower than the point defect-free value. Pressure decreases the relative reduction of the  
228  $1/2[100](010)$  GSF energy with adsorption of  $\{V_{M2}\}$ ; its value at 10 GPa ( $0.184 \text{ eV}/\text{\AA}^2$ ) corresponds to  
229 a reduction of only 9.5%.

230 The  $1/2[001](010)$  GSF energy calculated with a  $\{V_{M1}\}$  defect at the fault plane is  $0.168 \text{ eV}/\text{\AA}^2$  at  
231 ambient pressure, decreasing to  $0.145 \text{ eV}/\text{\AA}^2$  at 10 GPa. Comparing these with the values reported  
232 above for point defect-free case, we see that M1 vacancies do not lubricate  $[001](010)$  slip, and may  
233 actually inhibit it. In contrast,  $\{V_{M2}\}$  defects reduce the  $1/2[001](010)$  GSF energy to  $0.127 \text{ eV}/\text{\AA}^2$  at 0  
234 GPa, a decrease of -22.1%. At 10 GPa, the energy of this GSF is  $0.101 \text{ eV}/\text{\AA}^2$ , corresponding to a  
235 -28.9% change in the GSF energy relative to the point defect-free GSF case. Overall, (010) stacking  
236 fault energies are reduced much more by the presence of M2 than M1 vacancies, unsurprising given  
237 that the fault plane intersects the sheet of  $M2O_6$  octahedra. The lubrication effect is greatest for  $[001]$   
238 (010)  $\gamma$ -line energies, as the minimum distance between M2 sites across the fault plane is  $2.61 \text{ \AA}$  (at the  
239  $\gamma$ -line maximum), compared with  $3.89 \text{ \AA}$  for the  $[100](010)$   $\gamma$ -line. As the  $\{V_{M2}\}$  defect interacts more  
240 strongly than the  $\{V_{M1}\}$  defect with both  $1/2[100](010)$  and  $[001](010)$  stacking faults,  $\{V_{M2}\}$  defects  
241 can be expected to be relatively more abundant near the dislocation core than in the bulk lattice.

242 The markedly different sensitivities of the  $[100](010)$  and  $[001](010)$   $\gamma$ -lines to adsorption of  $\{V_{M1}\}$   
243 versus  $\{V_{M2}\}$  defects to the stacking fault are a direct consequence of the location of the slip plane in  
244 the unit cell. As the lowest energy slip plane lies across the center of the sheet of  $M2O_6$  octahedra that  
245 is parallel to the (010) plane, it does not intersect  $M1O_6$  octahedra or  $SiO_4$  tetrahedra (Fig. 1a), so that  
246 the creation of M1 vacancies cannot contribute to the total number of vacant lattice sites at the slip  
247 plane. Nevertheless,  $\{V_{M1}\}$  defects do still reduce the energy of the  $[100](010)$   $\gamma$ -line, although  $[001]$   
248 (010)  $\gamma$ -line energies are essentially unchanged. This is because displacement of the upper half the  
249 simulation cell by  $1/2[100]$  brings M2 and Si sites on opposite sites of the stacking plane into close  
250 proximity. Relaxing the atom in the M2 site away from the stacking fault reduces the energy of the

251 slab, which is easier when a void has been created above the M2 site by the insertion of a  $\{V_{M1}\}$   
252 defect. At 0 GPa, this increases the relaxed M2-Si distance from 2.939 Å to 3.021 Å. Furthermore, a  
253 sub-periodic modulation of the M1 site coordinates is introduced, as the Mg atoms in the layer of M1  
254 sites closest to the GSF stacking fault are displaced  $\sim 0.2$  Å along the fault plane normal. The sign of the  
255 displacement vector alternates along [001]. For the olivine crystal structure, [001](010) slip does not  
256 cause cations to approach as closely as does [100](010) slip, and there is thus less potential for vacant  
257 cation sites away from the slip plane to reduce GSF energies by accommodating relaxation away from  
258 the stacking fault.

### 259 *3.3 Peierls stresses of dislocations in point defect-free forsterite*

260 Core structures and Peierls stresses for edge and screw dislocations in forsterite were calculated using  
261 the PN model (equation 2), with the inelastic energy parameterized using the  $\gamma$ -lines calculated in the  
262 previous section. At 0 GPa, the Peierls stress calculated for the [100](010) edge dislocation in the point  
263 defect-free crystal is 2.4 GPa,  $\sim 25\%$  greater than the value computed by Durinck et al. (2007) using a  
264 DFT-parameterized continuous PN model. The Peierls stress increases with pressure (Fig. 4), reaching  
265 6.3 GPa at 10 GPa, much higher than the corresponding value of 2.2 GPa calculated by Durinck et al.  
266 The pressure derivative of the Peierls stress is thus  $d\sigma_p/dP = 0.39$ . For the [001](010) edge dislocation,  
267  $\sigma_p$  is calculated to be 3.4 GPa at ambient pressure, decreasing with increasing pressure to 1.4 GPa at 10  
268 GPa, in qualitative agreement with DFT-parameterized PN calculations by Durinck et al. (2007). This  
269 gives a pressure derivative for the Peierls stress of  $d\sigma_p/dP = -0.19$ . The disagreement between this  
270 calculation and that of Durinck et al. may be partially attributed to the present use of a discrete PN  
271 model, in contrast to the continuous formulation employed by Durinck et al. However, the principal  
272 disagreement concerns the pressure dependence of the  $\gamma$ -line energies. Whereas in this study the [001]  
273 (010)  $\gamma$ -line maximum decreases with pressure, the  $\gamma$ -line previously calculated is nearly pressure

274 independent (Durinck et al. 2005; Durinck et al. 2007). The cores of [100](010) and [001](010) edge  
275 dislocations are undissociated across the entire pressure range 0-10 GPa, consistent with previous PN  
276 calculations (Durinck et al. 2007).

277 A screw dislocation can glide on any lattice plane containing the dislocation line vector  $\xi$ . Each glide  
278 plane has a different structure and a correspondingly different lattice resistance, giving rise to preferred  
279 glide planes. Here, we calculate the Peierls stress for [100] and [001] glide on (010), corresponding to  
280 the screw dislocations that enable deformation of the [100](010) and [001](010) slip systems. The  
281 Peierls stress for glide of [100] on the (010) plane at ambient pressure (Fig. 5) is 6.8 GPa, comparable  
282 with the 6.3 GPa predicted by DFT-parametrized PN calculations (Durinck et al. 2007). However, we  
283 calculate  $d\sigma_p/dP = 0.47$ , slightly greater than the pressure derivative of  $d\sigma_p/dP = 0.39$  calculated by  
284 Durinck et al. (2007). This is somewhat lower than the pressure derivative  $d\sigma_p/dP \sim 0.67$  measured in  
285 high-stress deformation experiments for the average Peierls stress, which is presumably controlled by  
286 the easy [100](010) slip system (Proietti et al. 2016).

287 The [001] screw dislocation glides on the (010) plane with a Peierls stress  $\sigma_p = 5.4$  GPa at 0 GPa  
288 pressure (Fig. 5), comparable to values calculated in previous theoretical studies, which range between  
289 3.1 GPa (Durinck et al. 2007) and 7.2 GPa (Mahendran et al. 2017). As was found for the [001](010)  
290 edge dislocation, pressure enhances glide on this slip system by reducing its Peierls stress, with  $d\sigma_p/dP$   
291 = -0.21. Fully atomistic calculations show that the static [001] screw dislocation core has a labile non-  
292 planar structure and that the effect of applied stress is to transform into a high core energy-low Peierls  
293 stress planar configuration (Carrez et al. 2008). The critical stress for dislocation glide is thus the stress  
294 required to activate the locking-unlocking mechanism for this slip system (Mahendran et al. 2017).  
295 This locking-unlocking mechanism cannot be modeled within the PN formalism, and the  $\sigma_p$  reported in  
296 this study is for the glissile (i.e. unlocked) dislocation core gliding on the median plane of the  $M_2O_6$   
297 sheet.

298 3.4 *Effect of vacancies on the Peierls stress*

299 For [100](010) edge dislocations, the adsorption of a  $\{V_{M1}\}$  defect to the stacking fault plane reduces  
300 Peierls stress by -16% at ambient pressure, to 2.0 GPa. The pressure dependence of  $\sigma_p$  is not  
301 significantly changed by the presence of the defect as  $d\sigma_p/dP = 0.37$ , compared with 0.39 calculated  
302 above for the point defect-free dislocation, and at 10 GPa  $\sigma_p = 5.7$  GPa. The effect of the  $\{V_{M2}\}$  defect  
303 is more substantial, and for this defect  $\sigma_p = 1.2$  GPa at ambient pressure, slightly more than half the  
304 Peierls stress calculated for a dislocation without adsorbed vacancies. Moreover, unlike the  $\{V_{M1}\}$   
305 defect,  $\{V_{M2}\}$  defects reduce the pressure sensitivity of this slip system, as  $d\sigma_p/dP = 0.31$ . At 10 GPa,  
306  $\sigma_p = 4.3$  GPa, -31% lower than the value calculated above.

307 As can be seen in Fig. 4, the [001](010) edge dislocation responds quite differently from [100](010)  
308 edge dislocation to the presence of Mg vacancies. The adsorption of a  $\{V_{M1}\}$  defect to the stacking  
309 fault plane increases [001](010)  $\gamma$ -line energies, leading to considerably higher Peierls stresses for the  
310 [001](010) edge dislocation. At ambient pressure,  $\sigma_p = 3.7$  GPa, nearly 10% greater than the Peierls  
311 stress without point defects. The Peierls stress decreases with pressure at a rate  $d\sigma_p/dP = -0.10$ .  
312 Compared with the  $d\sigma_p/dP$  calculated for this dislocation without adsorbed point defects,  $\{V_{M1}\}$   
313 reduces the pressure sensitivity of [001](010) slip by almost a factor of 2. Unlike  $\{V_{M1}\}$  defects,  
314  $\{V_{M2}\}$  defects reduce the [001](010)  $\gamma$ -line energy. Correspondingly, as can be seen in Fig. 4c, the  
315 Peierls stress of the [001](010) edge dislocation is considerably reduced by  $\{V_{M2}\}$ , to 1.6 GPa at 0  
316 GPa applied pressure, decreasing to only 150 MPa at 10 GPa, giving  $d\sigma_p/dP = -0.14$ . Due to the  
317 relative softness of this slip system, the relative Peierls stress reduction actually increases with  
318 pressure, from -52% at 0 GPa to -90% at 10 GPa.

319 For [100](010) screw dislocations (Fig. 5), the presence of Mg vacancies at the glide plane reduces the  
320 Peierls stress. If the Mg vacancy is located on an M2 site,  $\sigma_p$  is calculated to be 6.1 GPa, 10.2% lower  
321 than in point defect-free forsterite, and increases with pressure to 10.7 GPa at 10 GPa. The pressure  
322 sensitivity of the [100] screw dislocation is unaffected by  $\{V_{M1}\}$  defects as  $d\sigma_p/dP = 0.46$ , nearly  
323 identical to the value found previously for this dislocation in the absence of adsorbed vacancies. For an  
324 Mg vacancy located on an M1 site close to (but not on) the slip plane, calculated values for  $\sigma_p$  at these  
325 pressures are comparable, 6.3 GPa and 10.8 GPa at 0 and 10 GPa, respectively, as is the pressure  
326 derivative  $d\sigma_p/dP$ , which is 0.45.

327 The Peierls stresses of [001] screw dislocations gliding on (010) are similarly reduced by the presence  
328 of M2 vacancies at the slip plane, to 3.0 GPa at ambient pressure and 0.8 GPa at 10 GPa (Fig. 5). The  
329 pressure derivative  $d\sigma_p/dP = -0.21$ , identical to the value above for point defect-free forsterite. Unlike  
330 the [100](010) screw dislocation, the relative Peierls stress reduction induced by the presence of M2  
331 vacancies actually increases with pressure, from -44% at ambient pressure to -75% at 10 GPa.  
332 However, just as in the case of [001](010) edge dislocations, glide of [001] screw dislocations on (010)  
333 is modestly inhibited by the presence of M1 vacancies. At ambient pressure,  $\sigma_p = 5.7$  GPa, decreasing  
334 to 4.3 GPa at 10 GPa pressure. The magnitude of  $d\sigma_p/dP$  is reduced by  $\sim 1/3$ , to -0.14, relative to the  
335 point defect-free value.

#### 336 4. Discussion

337 In vacancy-free forsterite, the ratio  $\sigma_{p,[100]}/\sigma_{p,[001]}$  increases with  $P$  for both the edge and screw  
338 components of both slip systems. For edge dislocations,  $\sigma_{p,[100]}/\sigma_{p,[001]}$  varies from 0.71 at ambient  
339 pressure, to 4.4 at 10 GPa (Fig. 6a). The addition of  $\{V_{M1}\}$  vacancies changes this ratio, decreasing it  
340 to 0.54 at ambient pressure and reducing the pressure sensitivity so that, at 10 GPa, it is just 2.16. In  
341 contrast,  $\{V_{M2}\}$  defects increase the ratio, to 0.77 at 0 GPa and 28.51 at 10 GPa. In each case, the ratio

342  $\sigma_{p,[100]}/\sigma_{p,[001]} = 1$  corresponds to the pressure at which [100](010) glide becomes harder than [001](010)  
343 glide. This occurs at 1.5 GPa with no point defects present, increasing to 3.6 GPa when  $\{V_{M1}\}$  defects  
344 are adsorbed to the stacking fault and decreases to 0.7 GPa upon adsorption of  $\{V_{M2}\}$  defects to the  
345 stacking fault plane. In the case of screw dislocations,  $\sigma_{p,[100]}/\sigma_{p,[001]}$  is strictly  $> 1$ . However,  $\{V_{M1}\}$   
346 defects reduce it from 1.26 to 1.09 at ambient pressure, and from 3.49 to 2.50 at 10 GPa. The  $\{V_{M2}\}$   
347 defect has the reverse effect on the two slip systems, increasing  $\sigma_{p,[100]}/\sigma_{p,[001]}$  to 2.06 at ambient pressure  
348 and 13.10 and 10 GPa. From this, we conclude that, while  $\{V_{M1}\}$  defects cause the relative activities  
349 of the two slip systems to converge at constant  $P$ ,  $\{V_{M2}\}$  defects enhance slip on [001](010).

350 As qualitatively predicted by our models, experimental studies have reported a pressure-induced  
351 change in the deformation fabric of anhydrous olivine, caused by a change of the preferred slip system  
352 from [100](010) to [001](010) (Couvry et al. 2004; Raterron et al. 2016). This transition may be  
353 observable in seismological models for the Earth's upper mantle. The LPO of Fe-bearing  $Fo_{90}$  olivine  
354 suggests a change from dominant [100](010) to [001](010) slip at  $\sim 3$  GPa, which corresponds to a  
355 mantle depth of approximately 70-80 km (Jung et al. 2009). At 0 GPa, the [100](010) edge dislocation  
356 has the lowest Peierls stress and so should be the dominant slip system during glide deformation.  
357 However, because pressure is calculated to suppress the activity of this slip system, the Peierls stress of  
358 the [001](010) edge dislocation is lower at high pressure. The  $P$ - $\sigma_p$  curves of the two slip systems in  
359 point defect-free forsterite intersect at 1.5 GPa (Fig. 4a), slightly lower than the pressures reported in  
360 experiments.

361 Deformation experiments produce different olivine textures under dry and wet conditions. One  
362 plausible reason for this is that hydrated vacancies lubricate [001](010) glide more than [100](010)  
363 glide (Katayama et al. 2004), leading to the development of the flow normal B-type fabric associated  
364 with the mantle wedge. Ohuchi et al. (2012) have found that the transition between the two fabrics  
365 occurs at  $\sim 650$  ppm H/Si. Consistent with this, we find that the Peierls stresses of [100](010) edge and

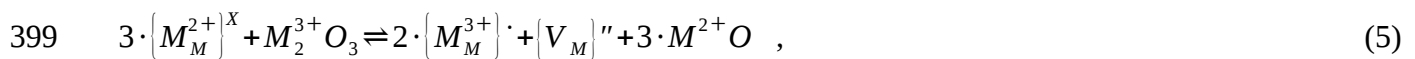


366 screw dislocations are reduced less by the presence of M-site vacancies than [001](010) edge and screw  
367 dislocations. Realistic vacancy-related defect concentrations are insufficient to change the weakest slip  
368 system at ambient pressure. However, the differential effect on the Peierls stresses of the [100](010)  
369 and [001](010) edge dislocations is enough to reduce the pressure at which the Peierls stresses of the  
370 two slip systems cross over. When the inelastic restoring force is parametrized using the  $\gamma$ -lines  
371 calculated with vacancies adsorbed to the slip plane,  $P_{\text{cross-over}}$  decreases from 1.5 GPa to 0.7 GPa.  
372 Higher defect concentrations at the dislocation core will cause the slip system cross-over to be  
373 displaced to even lower pressures. Similarly, differences between concentrations of vacancy-related  
374 defects at the glide planes of the two slip systems will displace the cross-over pressure. Increasing the  
375 concentration  $\{V_{M2}\}$  defects adsorbed to the [100](010) dislocation or decreasing the concentration of  
376 defects adsorbed to the [001](010) dislocation will reduce  $P_{\text{cross-over}}$ . Changing either of the  
377 concentrations in the opposite direction increases  $P_{\text{cross-over}}$ .

378 The calculated influence of vacancy-related defects on the Peierls stress of olivine may be of  
379 significance for development of LPO during deformation. It has also been suggested that a distinct LPO  
380 may form in olivine when it is deformed via diffusion creep (Miyazaki et al. 2013). Similarly, a  
381 transition to a grain boundary sliding deformation mechanism may explain the presence of the B-type  
382 deformation fabric, rather than water-induced changes to the relative strengths of olivine slip systems.  
383 However, as these calculations show, the adsorption of vacancies, and by implication vacancy-related  
384 defects, is sufficient to induce a change in the deformation fabric of olivine similar to that observed by  
385 Katayama and Karato (2008) under hydrous conditions. In tomographic images of shear wave  
386 anisotropy, this might show up as a reduction of the depth at which the deformation fabric changes  
387 from A-type (i.e. [100](010)) to B-type (i.e. [001](010)). Alternatively, if water is present only at  
388 shallow depths, the shear wave anisotropy may match the B-type fabric at shallow depths, transitioning  
389 to the A-type deformation fabric as the modal abundance of water decreases. The B-type deformation

390 fabric then re-emerges at high pressure as the relative activity of the [001](010) slip system increases.  
 391 These conclusions carry over to the dislocation-accommodated grain-boundary sliding regime. This  
 392 deformation mechanism leads to the development of a pronounced LPO that relates to the slip system  
 393 accommodating grain boundary sliding (Hansen et al. 2011; Hansen et al. 2012ab), which will be  
 394 altered by the presence of hydrous vacancies.

395 Incorporation of water is not the only mechanism by which vacancies can be generated in olivine  
 396 minerals. Trivalent cations such as  $\text{Al}^{3+}$ ,  $\text{Sc}^{3+}$  and  $\text{Fe}^{3+}$  can substitute for  $\text{Mg}^{2+}$  (or  $\text{Fe}^{2+}$ ) in the olivine  
 397 lattice, charge-balanced either the replacement of  $\text{Si}^{4+}$  by a trivalent cation or, more commonly, the  
 398 creation of M-site vacancies (Colson et al. 1989). The latter substitution mechanism can be written as



400 where  $M^{2+}$  is typically  $\text{Mg}^{2+}$  or, in iron-bearing olivine,  $\text{Fe}^{2+}$  and  $M^{3+}$  is the substituting trivalent cation.  
 401 These vacancies can influence the mechanical properties of olivine, for instance by enhancing seismic  
 402 attenuation due to grain boundary mechanisms (Cline et al. 2018). The solubility of common trivalent  
 403 elements in olivine is sufficiently high that vacancies created by the substitution reaction (5) can greatly  
 404 outnumber intrinsic vacancies, such as those associated with Schottky defects (van Orman et al. 2009).  
 405 In  $(\text{Mg,Fe})_2\text{SiO}_4$ , M-site vacancies are produced by oxidation of iron from  $\text{Fe}^{2+}$  to  $\text{Fe}^{3+}$ , and at high  $f\text{O}_2$   
 406  $\left\{ \text{Fe}_M^{3+} \right\}^{\cdot}$  and  $\left\{ V_M \right\}''$  are the most abundant defects (Stocker 1978). These defects may be unassociated at  
 407 high temperature and pressure, as they are in the common oxide  $(\text{Mg,Fe})\text{O}$  (Otsuka et al. 2010), and we  
 408 expect such vacancies to alter the Peierls stress of dislocations in olivine.

## 409 5. Conclusions

410 In this study, we used atomistic calculations of generalized stacking fault energies in forsterite, the Mg  
 411 end-member of olivine, to parameterize Peierls-Nabarro models of pure edge and screw dislocation. To  
 412 probe the possible influence of vacancy-related defects on the Peierls stress, energies were calculated

413 for GSFs in point defect-free forsterite, and with  $\{V_{M1}\}''$  and  $\{V_{M2}\}''$  defects segregated to lattice sites  
414 in close proximity to the stacking fault plane. It was found that vacancies are capable of reducing  $\gamma$ -line  
415 energies and, correspondingly, Peierls stresses for many of the major slip systems in olivine. However,  
416 the magnitude of this decrease depends strongly on both the slip system and the site on which the  
417 vacancy is located.  $\{V_{M2}\}''$  defects at lattice sites adjacent to the glide plane have the greatest effect on  
418 the Peierls stress of both slip systems, while the magnitude of the  $\sigma_p$  reduction is greatest for the [001]  
419 (010) slip system. In the case of edge dislocations, this reduces the pressure at which the hardening of  
420 [100](010) glide cause [001](010) to become the dominant slip system.

421 Cation vacancies in olivine are associated with intrinsic defect mechanisms (Schottky and Frenkel  
422 defects). However, in natural olivine, vacancy concentrations are, in most cases, probably controlled by  
423 extrinsic defect mechanisms, such as the creation of protonated vacancies in hydrated systems, the  
424 oxidation of  $Fe^{2+}$  to  $Fe^{3+}$ , or the charge-balanced substitution of heterovalent impurities, such  $Al^{3+}$  or  
425  $Cr^{3+}$ . Each of these mechanisms will be operative in at least some regions of the Earth's upper mantle.  
426 Hydration and oxidation are especially relevant in mantle wedges above subduction zones. There,  
427 concentrations of vacancy related defects are likely to be elevated, due to the combination of oxidizing  
428 conditions and the availability of liquid water from dehydrating phases in the subducting slab. Our  
429 models suggest that these intrinsic point defects can reduce the Peierls stress of dislocations in olivine  
430 in mantle wedges, potentially linking crystal chemistry to the observed variations in seismic anisotropy  
431 in these regions.

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#### 441 **References**

- 442 Bai, Q., Kohlstedt, D.L., 1993. Effects of chemical environment on the solubility and incorporation  
443 mechanism for hydrogen in olivine. *Phys Chem Minerals* 19, 460–471.  
444 <https://doi.org/10.1007/BF00203186>
- 445 Berry, A.J., Hermann, J., O'Neill, H.S.C., Foran, G.J., 2005. Fingerprinting the water site in mantle  
446 olivine. *Geology* 33, 869–872. <https://doi.org/10.1130/G21759.1>
- 447 Braithwaite, J.S., Sushko, P.V., Wright, K., Catlow, C.R.A., 2002. Hydrogen defects in Forsterite: A test  
448 case for the embedded cluster method. *The Journal of Chemical Physics* 116, 2628–2635.  
449 <https://doi.org/10.1063/1.1433465>
- 450 Brodholt, J., 1997. Ab initio calculations on point defects in forsterite (Mg<sub>2</sub>SiO<sub>4</sub>) and implications for  
451 diffusion and creep. *American Mineralogist* 82, 1049–1053. <https://doi.org/10.2138/am-1997-11-1201>
- 452 Christian, J.W., Vitek, V., 1970. Dislocations and stacking faults. *Rep. Prog. Phys.* 33, 307.  
453 <https://doi.org/10.1088/0034-4885/33/1/307>
- 454 Bulatov, V.V., Kaxiras, E., 1997. Semidiscrete Variational Peierls Framework for Dislocation Core  
455 Properties. *Phys. Rev. Lett.* 78, 4221–4224. <https://doi.org/10.1103/PhysRevLett.78.4221>
- 456 Carrez, P., Walker, A.M., Metsue, A., Cordier, P., 2008. Evidence from numerical modelling for 3D  
457 spreading of [001] screw dislocations in Mg<sub>2</sub>SiO<sub>4</sub> forsterite. *Philosophical Magazine* 88, 2477–2485.  
458 <https://doi.org/10.1080/14786430802363804>
- 459 Cline II, C.J., Faul, U.H., David, E.C., Berry, A.J., Jackson, I., 2018. Redox-influenced seismic  
460 properties of upper-mantle olivine. *Nature* 555, 355–358. <https://doi.org/10.1038/nature25764>
- 461 Colson, R.O., McKay, G.A., Taylor, L.A., 1989. Charge balancing of trivalent trace elements in olivine  
462 and low-Ca pyroxene: A test using experimental partitioning data. *Geochimica et Cosmochimica Acta*  
463 53, 643–648. [https://doi.org/10.1016/0016-7037\(89\)90007-0](https://doi.org/10.1016/0016-7037(89)90007-0)
- 464 Cottrell, A.H., Bilby, B.A., 1949. Dislocation Theory of Yielding and Strain Ageing of Iron. *Proc. Phys.*  
465 *Soc. A* 62, 49. <https://doi.org/10.1088/0370-1298/62/1/308>

466 Couvy, H., Frost, D.J., Heidelbach, F., Nyilas, K., Ungár, T., Mackwell, S., Cordier, P., 2004. Shear  
467 deformation experiments of forsterite at 11 GPa - 1400°C in the multianvil apparatus. *European Journal*  
468 *of Mineralogy* 16, 877–889. <https://doi.org/10.1127/0935-1221/2004/0016-0877>

469 Demouchy, S., Tommasi, A., Boffa Ballaran, T., Cordier, P., 2013. Low strength of Earth's uppermost  
470 mantle inferred from tri-axial deformation experiments on dry olivine crystals. *Physics of the Earth and*  
471 *Planetary Interiors* 220, 37–49. <https://doi.org/10.1016/j.pepi.2013.04.008>

472 Dick, B.G., Overhauser, A.W., 1958. Theory of the Dielectric Constants of Alkali Halide Crystals.  
473 *Phys. Rev.* 112, 90–103. <https://doi.org/10.1103/PhysRev.112.90>

474 Durinck, J., Legris, A., Cordier, P., 2005. Pressure sensitivity of olivine slip systems: first-principle  
475 calculations of generalised stacking faults. *Phys Chem Minerals* 32, 646–654.  
476 <https://doi.org/10.1007/s00269-005-0041-2>

477 Durinck, J., Carrez, P., Cordier, P., 2007. Application of the Peierls-Nabarro model to dislocations in  
478 forsterite. *European Journal of Mineralogy* 19, 631–639. <https://doi.org/10.1127/0935-1221/2007/0019-1757>

480 Evans, B., Goetze, C., 1979. The temperature variation of hardness of olivine and its implication for  
481 polycrystalline yield stress. *J. Geophys. Res.* 84, 5505–5524. <https://doi.org/10.1029/JB084iB10p05505>

482 Faul, U.H., Cline II, C.J., David, E.C., Berry, A.J., Jackson, I., 2016. Titanium-hydroxyl defect-  
483 controlled rheology of the Earth's upper mantle. *Earth and Planetary Science Letters* 452, 227–237.  
484 <https://doi.org/10.1016/j.epsl.2016.07.016>

485 Faul, U.H., Ii, C.J.C., Berry, A., Jackson, I., Garapić, G., 2017. Constraints on oxygen fugacity within  
486 metal capsules. *Phys Chem Minerals* 1–13. <https://doi.org/10.1007/s00269-017-0937-7>

487 Gale, J.D., 1997. GULP: A computer program for the symmetry-adapted simulation of solids. *J. Chem.*  
488 *Soc., Faraday Trans.* 93, 629–637. <https://doi.org/10.1039/A606455H>

489 Gale, J.D., Rohl, A.L., 2003. The General Utility Lattice Program (GULP). *Molecular Simulation* 29,  
490 291–341. <https://doi.org/10.1080/0892702031000104887>

491 Goetze, C., Kohlstedt, D.L., 1973. Laboratory study of dislocation climb and diffusion in olivine. *J.*  
492 *Geophys. Res.* 78, 5961–5971. <https://doi.org/10.1029/JB078i026p05961>

493 Hansen, L.N., Zimmerman, M.E., Kohlstedt, D.L., 2011. Grain boundary sliding in San Carlos olivine:  
494 Flow law parameters and crystallographic-preferred orientation. *J. Geophys. Res.* 116, B08201.  
495 <https://doi.org/10.1029/2011JB008220>

496 Hansen, L.N., Zimmerman, M.E., Kohlstedt, D.L., 2012a. The influence of microstructure on  
497 deformation of olivine in the grain-boundary sliding regime. *J. Geophys. Res.* 117, B09201.  
498 <https://doi.org/10.1029/2012JB009305>

499 Hansen, L.N., Zimmerman, M.E., Kohlstedt, D.L., 2012b. Laboratory measurements of the viscous  
500 anisotropy of olivine aggregates. *Nature* 492, 415. <https://doi.org/10.1038/nature11671>

501 Hilairet, N., Wang, Y., Sanehira, T., Merkel, S., Mei, S., 2012. Deformation of olivine under mantle  
502 conditions: An in situ high-pressure, high-temperature study using monochromatic synchrotron  
503 radiation. *J. Geophys. Res.* 117, B01203. <https://doi.org/10.1029/2011JB008498>

504 Idrissi, H., Bollinger, C., Boioli, F., Schryvers, D., Cordier, P., 2016. Low-temperature plasticity of  
505 olivine revisited with in situ TEM nanomechanical testing. *Science Advances* 2, e1501671.  
506 <https://doi.org/10.1126/sciadv.1501671>

507 Jackson, I., Faul, U.H., Skelton, R., 2014. Elastically accommodated grain-boundary sliding: New  
508 insights from experiment and modeling. *Physics of the Earth and Planetary Interiors, High-Pressure*  
509 *Research in Earth Science: Crust, Mantle, and Core* 228, 203–210.  
510 <https://doi.org/10.1016/j.pepi.2013.11.014>

511 Joós, B., Ren, Q., Duesbery, M.S., 1994. Peierls-Nabarro model of dislocations in silicon with  
512 generalized stacking-fault restoring forces. *Phys. Rev. B* 50, 5890–5898.  
513 <https://doi.org/10.1103/PhysRevB.50.5890>

514 Jung, H., Mo, W., Green, H.W., 2009. Upper mantle seismic anisotropy resulting from pressure-  
515 induced slip transition in olivine. *Nature Geosci* 2, 73–77. <https://doi.org/10.1038/ngeo389>

516 Katayama, I., Karato, S., 2008. Low-temperature, high-stress deformation of olivine under water-  
517 saturated conditions. *Physics of the Earth and Planetary Interiors* 168, 125–133.  
518 <https://doi.org/10.1016/j.pepi.2008.05.019>

519 Katayama, I., Jung, H., Karato, S., 2004. New type of olivine fabric from deformation experiments at  
520 modest water content and low stress. *Geology* 32, 1045–1048. <https://doi.org/10.1130/G20805.1>

521 Katayama, I., Karato, S., Brandon, M., 2005. Evidence of high water content in the deep upper mantle  
522 inferred from deformation microstructures. *Geology* 33, 613–616. <https://doi.org/10.1130/G21332.1>

523 Kelley, K.A., Cottrell, E., 2009. Water and the Oxidation State of Subduction Zone Magmas. *Science*  
524 325, 605–607. <https://doi.org/10.1126/science.1174156>

525 Kirfel, A., Lippmann, T., Blaha, P., Schwarz, K., Cox, D.F., Rosso, K.M., Gibbs, G.V., 2005. Electron  
526 density distribution and bond critical point properties for forsterite, Mg<sub>2</sub>SiO<sub>4</sub>, determined with  
527 synchrotron single crystal X-ray diffraction data. *Phys Chem Minerals* 32, 301–313.  
528 <https://doi.org/10.1007/s00269-005-0468-5>

529 Kohlstedt, D.L., Keppler, H., Rubie, D.C., 1996. Solubility of water in the  $\alpha$ ,  $\beta$  and  $\gamma$  phases of  
530 (Mg,Fe)<sub>2</sub>SiO<sub>4</sub>. *Contrib Mineral Petrol* 123, 345–357. <https://doi.org/10.1007/s004100050161>

531 Kranjc, K., Rouse, Z., Flores, K.M., Skemer, P., 2016. Low-temperature plastic rheology of olivine  
532 determined by nanoindentation. *Geophys. Res. Lett.* 43, 2015GL065837.  
533 <https://doi.org/10.1002/2015GL065837>

534 Kröger, F.A., Vink, H.J., 1956. Relations between the Concentrations of Imperfections in Crystalline  
535 Solids. *Solid State Physics* 3, 307–435. [https://doi.org/10.1016/S0081-1947\(08\)60135-6](https://doi.org/10.1016/S0081-1947(08)60135-6)

536 de Leeuw, N.H., Parker, S.C., Catlow, C.R.A., Price, G.D., 2000. Modelling the effect of water on the  
537 surface structure and stability of forsterite. *Phys Chem Min* 27, 332–341.  
538 <https://doi.org/10.1007/s002690050262>

539 Lemaire, C., Kohn, S.C., Brooker, R.A., 2004. The effect of silica activity on the incorporation  
540 mechanisms of water in synthetic forsterite: a polarised infrared spectroscopic study. *Contrib Mineral  
541 Petrol* 147, 48–57. <https://doi.org/10.1007/s00410-003-0539-x>

542 Lewis, G.V., Catlow, C.R.A., 1985. Potential models for ionic oxides. *J. Phys. C: Solid State Phys.* 18,  
543 1149. <https://doi.org/10.1088/0022-3719/18/6/010>

544 Liu, L., Du, J., Zhao, J., Liu, H., Gao, H., Chen, Y., 2009. Elastic properties of hydrous forsterites under  
545 high pressure: First-principle calculations. *Physics of the Earth and Planetary Interiors* 176, 89–97.  
546 <https://doi.org/10.1016/j.pepi.2009.04.004>

547 Long, M.D., Becker, T.W., 2010. Mantle dynamics and seismic anisotropy. *Earth and Planetary Science  
548 Letters* 297, 341–354. <https://doi.org/10.1016/j.epsl.2010.06.036>

549 Long, M.D., Silver, P.G., 2009. Shear Wave Splitting and Mantle Anisotropy: Measurements,  
550 Interpretations, and New Directions. *Surv Geophys* 30, 407–461. <https://doi.org/10.1007/s10712-009-9075-1>

551

552 Lu, G., Kaxiras, E., 2002. Can Vacancies Lubricate Dislocation Motion in Aluminum? *Phys. Rev. Lett.*  
553 89, 105501. <https://doi.org/10.1103/PhysRevLett.89.105501>

554 Lu, G., Zhang, Q., Kioussis, N., Kaxiras, E., 2001. Hydrogen-Enhanced Local Plasticity in Aluminum:  
555 An Ab Initio Study. *Phys. Rev. Lett.* 87, 095501. <https://doi.org/10.1103/PhysRevLett.87.095501>

556 Mahendran, S., Carrez, P., Groh, S., Cordier, P., 2017. Dislocation modelling in Mg<sub>2</sub>SiO<sub>4</sub> forsterite:  
557 an atomic-scale study based on the THB1 potential. *Modelling Simul. Mater. Sci. Eng.* 25, 054002.  
558 <https://doi.org/10.1088/1361-651X/aa6efa>

559 Mainprice, D., 2007. Seismic Anisotropy of the Deep Earth from a Mineral and Rock Physics  
560 Perspective, in: *Treatise of Geophysics*, Vol.2. Elsevier, pp. 437–491.

561 Mainprice, D., Tommasi, A., Couvy, H., Cordier, P., Frost, D.J., 2005. Pressure sensitivity of olivine  
562 slip systems and seismic anisotropy of Earth's upper mantle. *Nature* 433, 731.  
563 <https://doi.org/10.1038/nature03266>

564 Margheriti, L., Nostro, C., Cocco, M., Amato, A., 1996. Seismic anisotropy beneath the Northern  
565 Apennines (Italy) and its tectonic implications. *Geophys. Res. Lett.* 23, 2721–2724.  
566 <https://doi.org/10.1029/96GL02519>

567 Martin, R.F., Donnay, G., 1972. Hydroxyl in the mantle. *American Mineralogist* 57, 554–570.

568 Miyazaki, T., Sueyoshi, K., Hiraga, T., 2013. Olivine crystals align during diffusion creep of Earth's  
569 upper mantle. *Nature* 502, 321–326. <https://doi.org/10.1038/nature12570>

570 Mizukami, T., Wallis, S.R., Yamamoto, J., 2004. Natural examples of olivine lattice preferred  
571 orientation patterns with a flow-normal a-axis maximum. *Nature* 427, 432–436.  
572 <https://doi.org/10.1038/nature02179>

573 Mei, S., Kohlstedt, D.L., 2000. Influence of water on plastic deformation of olivine aggregates: 1.  
574 Diffusion creep regime. *J. Geophys. Res.* 105, 21457–21469. <https://doi.org/10.1029/2000JB900179>

575 Nabarro, F.R.N., 1947. Dislocations in a simple cubic lattice. *Proc. Phys. Soc.* 59, 256.  
576 <https://doi.org/10.1088/0959-5309/59/2/309>

577 Nicolas, A., Christensen, N.I., 1987. Formation of Anisotropy in Upper Mantle Peridotites - A Review,  
578 in: Fuchs, K., Froidevaux, C. (Eds.), *Composition, Structure and Dynamics of the Lithosphere-*  
579 *Asthenosphere System*. American Geophysical Union, pp. 111–123.  
580 <https://doi.org/10.1029/GD016p0111>

581 Ohuchi, T., Kawazoe, T., Nishihara, Y., Nishiyama, N., Irifune, T., 2011. High pressure and temperature  
582 fabric transitions in olivine and variations in upper mantle seismic anisotropy. *Earth and Planetary*  
583 *Science Letters* 304, 55–63. <https://doi.org/10.1016/j.epsl.2011.01.015>

584 Ohuchi, T., Kawazoe, T., Nishihara, Y., Irifune, T., 2012. Change of olivine a-axis alignment induced  
585 by water: Origin of seismic anisotropy in subduction zones. *Earth and Planetary Science Letters* 317–  
586 318, 111–119. <https://doi.org/10.1016/j.epsl.2011.11.022>

587 van Orman, J.A., Fei, Y., Hauri, E.H., Wang, J., 2003. Diffusion in MgO at high pressures: Constraints  
588 on deformation mechanisms and chemical transport at the core-mantle boundary. *Geophys. Res. Lett.*  
589 30, 1056. <https://doi.org/10.1029/2002GL016343>

590 Otsuka, K., McCammon, C.A., Karato, S., 2010. Tetrahedral occupancy of ferric iron in (Mg,Fe)O:  
591 Implications for point defects in the Earth's lower mantle. *Physics of the Earth and Planetary Interiors*,  
592 *Transport properties of the lower mantle* 180, 179–188. <https://doi.org/10.1016/j.pepi.2009.10.005>

593 Peierls, R., 1940. The size of a dislocation. *Proc. Phys. Soc.* 52, 34. [https://doi.org/10.1088/0959-](https://doi.org/10.1088/0959-5309/52/1/305)  
594 [5309/52/1/305](https://doi.org/10.1088/0959-5309/52/1/305)

595 Price, G.D., Parker, S.C., Leslie, M., 1987. The lattice dynamics and thermodynamics of the Mg<sub>2</sub>SiO<sub>4</sub>  
596 polymorphs. *Phys Chem Minerals* 15, 181–190. <https://doi.org/10.1007/BF00308782>



597 Proietti, A., Bystricky, M., Guignard, J., Bějina, F., Crichton, W., 2016. Effect of pressure on the  
598 strength of olivine at room temperature. *Physics of the Earth and Planetary Interiors* 259, 34–44.  
599 <https://doi.org/10.1016/j.pepi.2016.08.004>

600 Raterron, P., Chen, J., Geenen, T., Girard, J., 2011. Pressure effect on forsterite dislocation slip systems:  
601 Implications for upper-mantle LPO and low viscosity zone. *Physics of the Earth and Planetary Interiors*  
602 188, 26–36. <https://doi.org/10.1016/j.pepi.2011.06.009>

603 Raterron, P., Chen, J., Li, L., Weidner, D., Cordier, P., 2016. Pressure-induced slip-system transition in  
604 forsterite: Single-crystal rheological properties at mantle pressure and temperature. *American*  
605 *Mineralogist* 92, 1436–1445. <https://doi.org/10.2138/am.2007.2474>

606 Sanders, M.J., Leslie, M., Catlow, C.R.A., 1984. Interatomic potentials for SiO<sub>2</sub>. *J. Chem. Soc., Chem.*  
607 *Commun.* 1271–1273. <https://doi.org/10.1039/C39840001271>

608 Skelton, R., Walker, A.M., 2018. Lubrication of dislocation glide in MgO by hydrous defects. *Phys*  
609 *Chem Minerals* 1–14. <https://doi.org/10.1007/s00269-018-0957-y>

610 Smith, G.P., Wiens, D.A., Fischer, K.M., Dorman, L.M., Webb, S.C., Hildebrand, J.A., 2001. A  
611 Complex Pattern of Mantle Flow in the Lau Backarc. *Science* 292, 713–716.  
612 <https://doi.org/10.1126/science.1058763>

613 Stocker, R.L., 1978. Influence of oxygen pressure on defect concentrations in olivine with a fixed  
614 cationic ratio. *Physics of the Earth and Planetary Interiors* 17, 118–129. [https://doi.org/10.1016/0031-](https://doi.org/10.1016/0031-9201(78)90053-5)  
615 [9201\(78\)90053-5](https://doi.org/10.1016/0031-9201(78)90053-5)

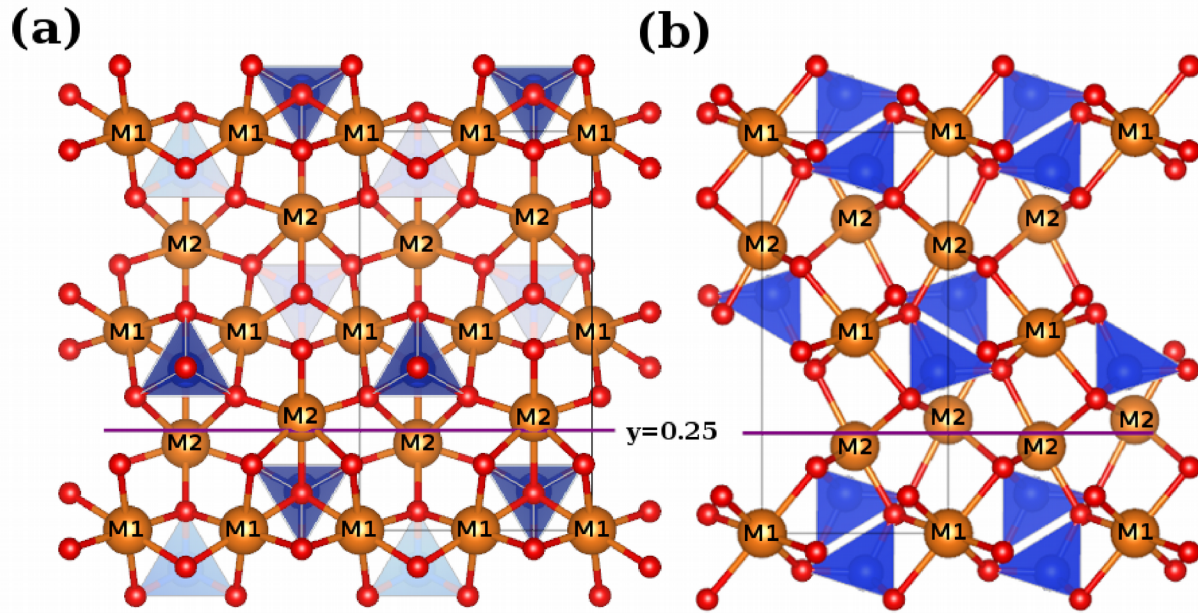
616 Walker, A.M., Gale, J.D., Slater, B., Wright, K., 2005. Atomic scale modelling of the cores of  
617 dislocations in complex materials part 2: applications. *Phys. Chem. Chem. Phys.* 7, 3235–3242.  
618 <https://doi.org/10.1039/B505716G>

619 Walker, A.M., Hermann, J., Berry, A.J., O’Neill, H.S.C., 2007. Three water sites in upper mantle  
620 olivine and the role of titanium in the water weakening mechanism. *J. Geophys. Res.* 112, B05211.  
621 <https://doi.org/10.1029/2006JB004620>

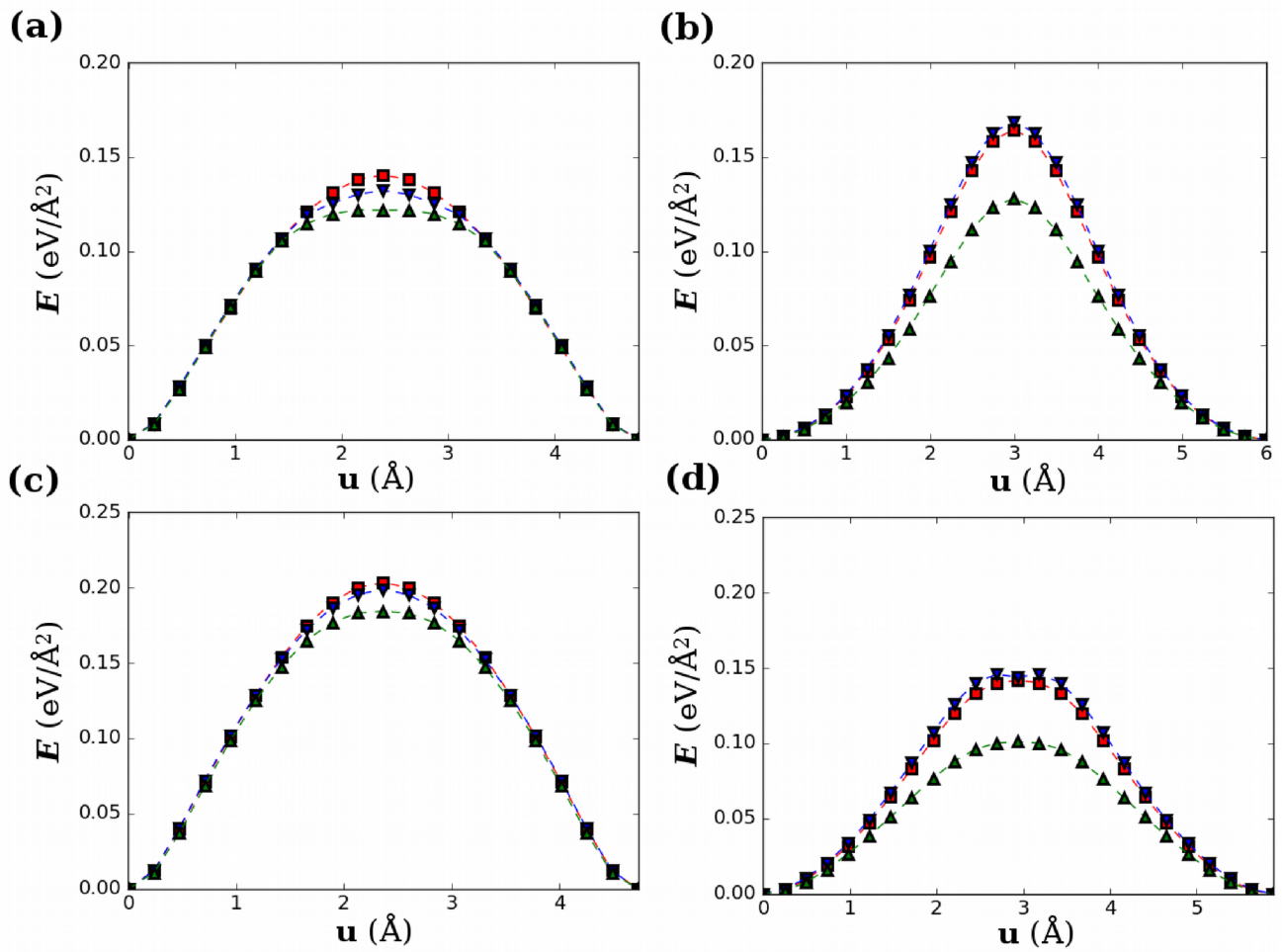
622 Wright, K., Catlow, C.R.A., 1994. A computer simulation study of (OH) defects in olivine. *Phys Chem*  
623 *Minerals* 20, 515–518. <https://doi.org/10.1007/BF00203222>

624 Zhang, F., Walker, A.M., Wright, K., Gale, J.D., 2010. Defects and dislocations in MgO: atomic scale  
625 models of impurity segregation and fast pipe diffusion. *J. Mater. Chem.* 20, 10445–10451.  
626 <https://doi.org/10.1039/C0JM01550D>

627 **Figure captions**



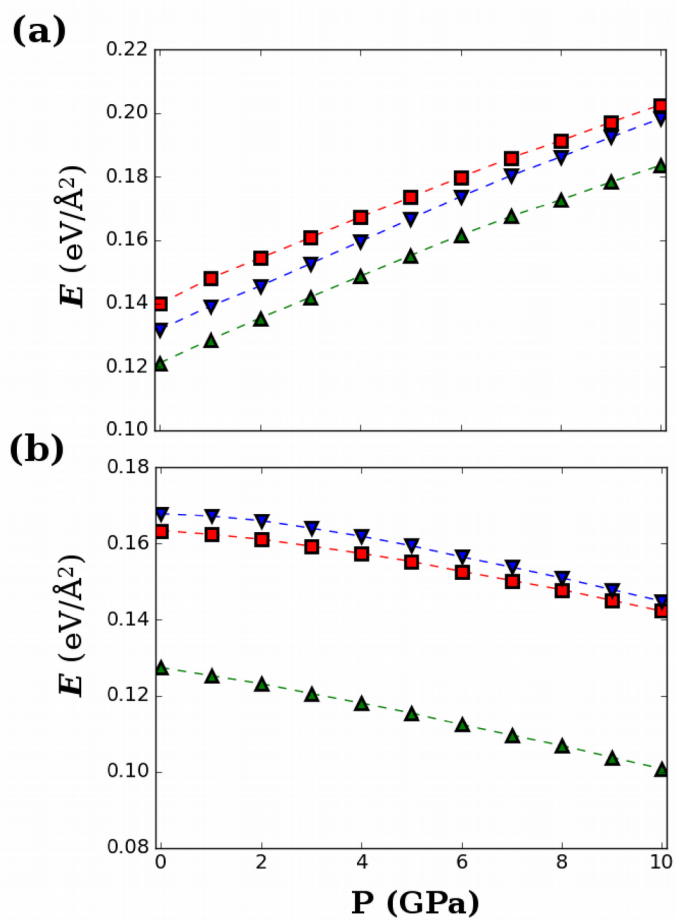
628 **Fig. 1** The olivine unit cell, viewed down the (a) [100] and (b) [001] cell directions, with the M1 and M2 sites  
629 labeled. The easy glide plane for dislocations gliding on (010) is marked. Visualization produced using VESTA 3  
630 (Momma and Izumi 2011).



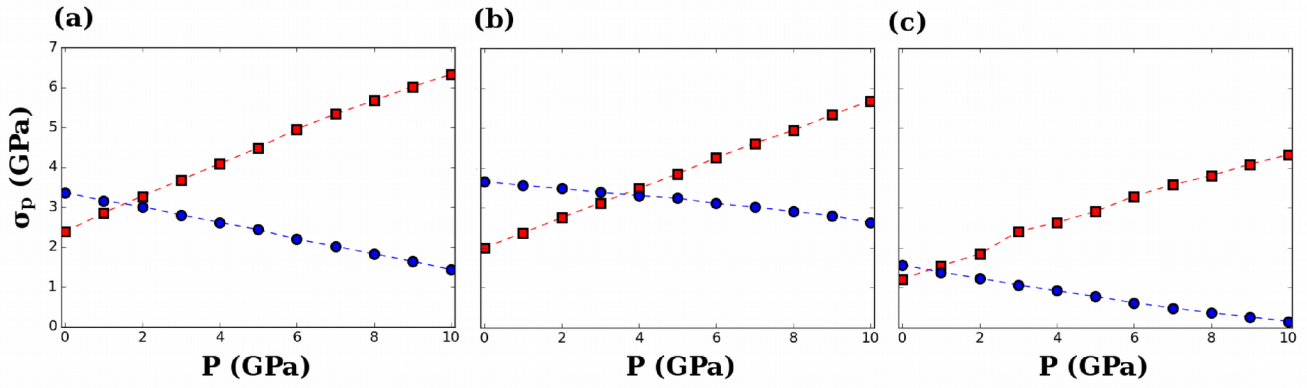
631 **Fig. 2** (a)  $[100](010)$  and (b)  $[001](010)$   $\gamma$ -lines at 0 GPa. (c)  $[100](010)$  and (d)  $[001](010)$   $\gamma$ -lines at 10 GPa.

632 Square, triangle, and inverted triangle symbols correspond to GSFs without point defects, with  $\{V_{M2}\}$  defects,

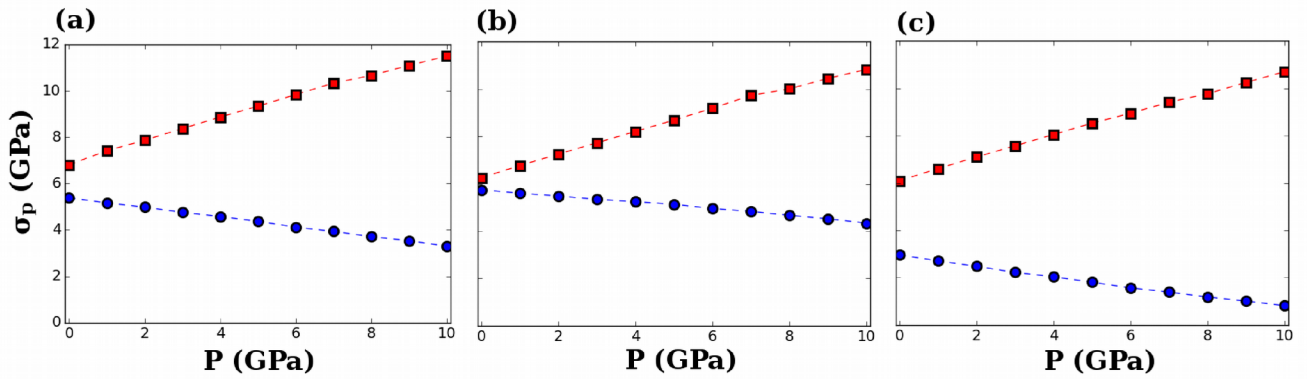
633 and with  $\{V_{M1}\}$  defects.



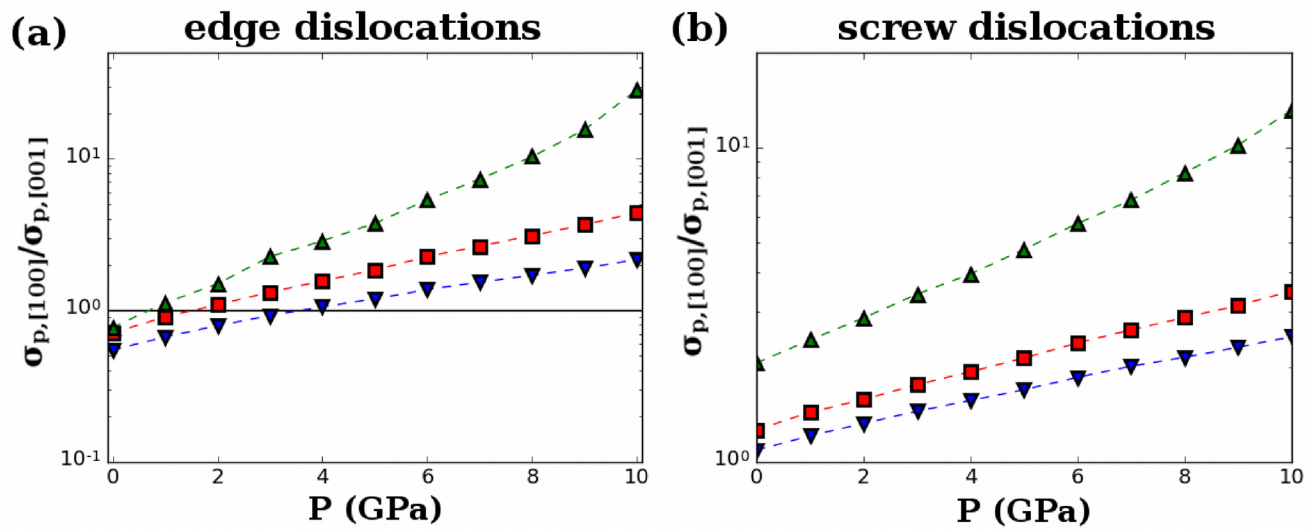
634 **Fig. 3** Maximum energies along the (a) [100](010) and (b) [001](010)  $\gamma$ -lines. Symbols have the same meaning  
 635 as in Fig. 2.



636 **Fig. 4** Pressure evolution of  $\sigma_p$  for (squares)  $[100](010)$  and (circles)  $[001](010)$  edge dislocations with (a) no  
 637 point defects, (b) a  $\{V_{M1}\}$  defect, and (c) a  $\{V_{M2}\}$  defect at the glide plane. All Peierls stresses are plotted using  
 638 the same scale. Symbols have the same meaning as in Fig. 2.



639 **Fig. 5** Pressure evolution of  $\sigma_p$  for (squares)  $[100](010)$  and (circles)  $[001](010)$  screw dislocations with (a) no  
 640 point defects, (b) a  $\{V_{M1}\}$  defect, and (c) a  $\{V_{M2}\}$  defect at the glide plane. All Peierls stresses are plotted using  
 641 the same scale.



642 **Fig. 6**  $\sigma_{p,[100]}/\sigma_{p,[001]}$  ratios for (a) edge, and (b) screw dislocations. Symbols have the same meaning as in Fig. 2.