

# The rupture extent of low frequency earthquakes near Parkfield, CA

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## Abstract

The low frequency earthquakes (LFEs) that constitute tectonic tremor are often inferred to be slow: to have durations of 0.2 to 0.5 s, a factor of 10 to 100 longer than those of typical  $M_W$  1-2 earthquakes. Here we examine LFEs near Parkfield, CA in order to assess several proposed explanations for LFEs' long durations. We determine LFE rupture areas and location distributions using a new approach, similar to directivity analysis, where we examine how signals coming from various locations within LFEs' finite rupture extents create differences in the apparent source time functions recorded at various stations. We use synthetic ruptures to determine how much the LFE signals recorded at each station would be modified by spatial variations of the source-station travel time within the rupture area given various possible rupture diameters, and then compare those synthetics with the data. Our synthetics show that the methodology can identify inter-station variations created by heterogeneous slip distributions or complex rupture edges, and thus lets us estimate LFE rupture extents for unilateral or bilateral ruptures. To obtain robust estimates of the sources' similarity across stations, we stack signals from thousands of LFEs, using an empirical Green's function approach to isolate the LFEs' apparent source time functions from the path effects. Our analysis of LFEs in Parkfield implies that LFEs' apparent source time functions are similar across stations at frequencies up to 8 to 16 Hz, depending on the family.

The inter-station coherence observed at these relatively high frequencies, or short wavelengths (down to 0.2 to 0.5 km), suggest that LFEs in each of the 7 families examined occur on asperities. They are clustered in patches with sub-1-km diameters. The individual LFEs' rupture diameters are estimated to be smaller than 1.1 km for all families, and smaller than 0.5 km and 1 km for the two shallowest families, which were previously found to have 0.2-s durations. Coupling the diameters with the durations suggests that it is possible to model these  $M_W$  1-2 LFEs with earthquake-like rupture speeds: around 70% of the shear wave speed. However, that rupture speed matches the data only at the edge of our uncertainty estimates for the family with highest coherence. The data for that family are better matched if LFEs have rupture velocities smaller than 40% of the shear wave speed, or if LFEs have different rupture dynamics. They could have long rise times, contain composite sub-ruptures, or have slip distributions that persist from event to event.

# 1 Introduction

Tectonic tremor is a long-duration seismic signal, best observed at frequencies between 1 and 10 Hz (e.g., *Obara, 2002; Rogers and Dragert, 2003; Payero et al., 2008; Peterson and Christensen, 2009; Rubinstein et al., 2009; Fry et al., 2011*). It is thought to consist of numerous small low frequency earthquakes, or LFEs (*Shelly et al., 2006, 2007; Wech and Creager, 2007; Brown et al., 2009*). LFEs are often inferred to have magnitudes between  $M_W$  1 and 2.5 but to have corner frequencies of a few Hz, a factor of 10 to 100 times smaller than corner frequencies observed for “normal”  $M_W$  1-2.5 earthquakes (*Fletcher and McGarr, 2011; Zhang et al., 2011; Bostock et al., 2017*). LFEs are found to have durations around 0.2 seconds in Parkfield (*Thomas et al., 2016*) and around 0.5 s in Cascadia (*Bostock et al., 2015*), which are a factor of 10 to 100 longer than “normal”  $M_W$  1-2.5 earthquakes.

## 1.1 Potential Causes of LFEs’ Long Durations

The durations of typical earthquakes are determined by their spatial extent: by how long it takes the rupture to progress across the earthquake area. Models and observations suggest that earthquake ruptures usually progress at speeds of 2 to 3 km/s, or 60 to 95% of the shear wave speed  $V_s$  (*Kanamori and Brodsky, 2004; McGuire, 2004; Madariaga, 2007; Seekins and Boatwright, 2010; Taira et al., 2015; Folesky et al., 2016; Ye et al., 2016; Melgar and Hayes, 2017; Chounet et al., 2018*). Earthquakes’ durations can thus be roughly estimated by dividing their rupture lengths by the shear wave speed. If LFEs, like normal earthquakes, rupture at speeds close to the shear wave speed, their long durations could indicate that LFEs have unusually large lengths given their moment: perhaps 0.7 to 1.5 km. In this scenario, LFEs would have lower stress drops than normal earthquakes—0.1 to 10 kPa, but they could otherwise be governed by the same physical processes. LFEs could be driven by unstable frictional sliding, and their slip speeds could be limited by the energy that they dissipate via seismic waves (e.g., *Rice, 1980; Kanamori and Brodsky, 2004*).

However, it is also possible that seismic wave generation has minimal impact on LFE dynamics, and that LFEs are governed by different fault zone processes. LFEs’ slip rates may be limited by a spatial constraint or by a speed-limiting frictional rheology (e.g., *Liu and Rice, 2005, 2007; Shibazaki and Shimamoto, 2007; Rubin, 2008; Segall et al., 2010; Skarbak et al., 2012; Fagereng et al., 2014; Yabe and Ide, 2017*). For instance, LFEs might occur on faults with a velocity-strengthening rheology, which inhibits increases in slip rate. The brief slip rate increases seen in LFEs could result from imposed local stress concentrations, perhaps created by the creep fronts of large slow slip events (e.g., *Perfettini and Ampuero, 2008; Rubin, 2009*). Alternatively, LFEs could occur on faults with a more complex rheology, which encourages initial increases in slip rate but inhibits slip rates higher than some cutoff speed. Such rheologies are commonly proposed for slow slip events and may be created by shear-induced dilatancy or by a minimum asperity size (e.g., *Shibazaki and Iio, 2003; Shibazaki and Shimamoto, 2007; Liu et al., 2010; Segall et al., 2010; Hawthorne and Rubin, 2013; Poulet et al., 2014*). The possibility that LFEs are small versions of slow slip events is intriguing because slip rates vary widely from slow slip to tremor (*Ide et al., 2007, 2008; Aguiar et al., 2009; Gao et al., 2012; Ide and Yabe, 2014; Hawthorne and Bartlow, 2018*). Several of the processes proposed to govern slow slip would have difficulty producing such a wide range of slip rates (e.g., *Liu and Rice, 2005, 2007; Shibazaki and Shimamoto, 2007; Hawthorne and Rubin, 2013; Fagereng et al., 2014; Veveakis et al., 2014*). If LFE slip rates are

79 limited primarily by frictional resistance to shear and not by seismic wave radiation, LFEs need not  
80 rupture across the fault at speeds close to the shear wave speed. They could rupture more slowly  
81 and have diameters far smaller than 1 km despite their 0.2-s durations.

82 LFEs could also have small rupture diameters if their 0.2-s durations and low corner frequencies  
83 are actually apparent values, not true ones. LFEs could be “normal”  $M_W$  1-2.5 earthquakes, with  
84 0.01-s durations and 10-m rupture diameters. They may appear to be dominated by low-frequency  
85 signals only because their high-frequency signals are attenuated when they pass through a highly  
86 damaged fault zone or through a region of high pore fluid pressure (*Gomberg et al.*, 2012; *Bostock*  
87 *et al.*, 2017). Regions of high pore pressure or increased attenuation are frequently identified near  
88 the slow slip region (*Audet et al.*, 2009; *Song et al.*, 2009; *van Avendonk et al.*, 2010; *Kato et al.*,  
89 2010; *Fagereng and Diener*, 2011; *Kitajima and Saffer*, 2012; *Nowack and Bostock*, 2013; *Yabe*  
90 *et al.*, 2014; *Saffer and Wallace*, 2015; *Audet and Schaeffer*, 2018), though we note that any regions  
91 with attenuation strong enough to produce tremor’s frequency content might have to be localized  
92 into patches. Earthquakes do occur below the tremor-generating region, and some of them show  
93 higher-frequency signals than tremor (*Seno and Yamasaki*, 2003; *Shelly et al.*, 2006; *Bell et al.*,  
94 2010; *Kato et al.*, 2010; *Gomberg et al.*, 2012; *Bostock et al.*, 2017).

## 95 **1.2 Potential Role of Tremor Asperities**

96 Tremor is often patchily distributed along the plate interface; it is densely concentrated in some  
97 regions but appears absent in others (e.g., *Payero et al.*, 2008; *Maeda and Obara*, 2009; *Walter*  
98 *et al.*, 2011; *Ghosh et al.*, 2012; *Armbruster et al.*, 2014). Some observations and models suggest  
99 that tremor occurs only on a set of tremor-generating asperities (e.g., *Ariyoshi et al.*, 2009; *Ando*  
100 *et al.*, 2010; *Shelly*, 2010b; *Nakata et al.*, 2011; *Ando et al.*, 2012; *Sweet et al.*, 2014; *Veedu and*  
101 *Barbot*, 2016; *Chestler and Creager*, 2017a,b; *Luo and Ampuero*, 2017). Such asperities may also  
102 be suggested by the success of template matching approaches to tremor identification, in which  
103 LFEs are detected and grouped into families according to waveform similarity. Each LFE family  
104 could reflect an individual tremor asperity (*Shelly et al.*, 2007; *Brown et al.*, 2008; *Bostock et al.*,  
105 2012; *Frank et al.*, 2013; *Kato*, 2017; *Shelly*, 2017). However, the family grouping could also result  
106 from more gradual variations in the path effects. LFEs located more than 1 or a few km away from  
107 each other may be grouped into distinct families simply because the path effects vary significantly  
108 on several-km length scales, so that well-separated LFEs give rise to distinct seismograms.

109 A few studies have provided further indications that LFE families reflect tremor clusters. *Sweet*  
110 *et al.* (2014) relocated LFEs within an isolated family in Cascadia and found that they clustered  
111 within a 1-km-wide patch. *Chestler and Creager* (2017b) relocated LFEs within around 20 families  
112 in Cascadia and found that LFEs cluster within 1 to 2-km-wide patches that are often separated by  
113 > 5-km-wide areas with few to no LFEs. Tremor-generating asperities are also suggested by the  
114 highly repetitive recurrence intervals of one isolated LFE family near Parkfield, CA, which suggest  
115 that the LFEs could be repeating similar ruptures of a particular asperity (*Shelly*, 2010b; *Veedu and*  
116 *Barbot*, 2016). Repetitive LFE rupture is also suggested by LFE moments and durations that vary  
117 little from event to event, creating exponential amplitude distributions (*Watanabe et al.*, 2007;  
118 *Shelly and Hardebeck*, 2010; *Chamberlain et al.*, 2014; *Sweet et al.*, 2014; *Bostock et al.*, 2015;  
119 *Chestler and Creager*, 2017a), though it is also possible that each LFE ruptures only a portion of a  
120 tremor-generating asperity. The total slip on an LFE patch could result from a range of ruptures of  
121 different types, as well as some aseismic slip (*Chestler and Creager*, 2017a).

### 1.3 Analysis to Be Presented

In this study, we further assess whether small asperities control tremor generation and whether LFEs are governed by earthquake-like or slow slip rheologies by determining the rupture extents of LFEs in seven families near Parkfield, CA. We will place upper bounds on the spatial distribution of LFEs in each family and on the average LFE rupture area. In order to obtain these bounds, we will introduce a new coherence-based approach, which can be thought of as a version of directivity analysis that we have modified so that we can combine data from thousands of LFEs which may rupture unilaterally or bilaterally (e.g. *Mueller, 1985; Mori and Frankel, 1990; Got and Fréchet, 1993; Velasco et al., 1994; Lengliné and Got, 2011; Wang and Rubin, 2011; Kane et al., 2013*). We examine how signals coming from various locations within LFEs' finite rupture areas can produce complex apparent source time functions (ASTFs) that vary from station to station. We quantify the ASTF variation as a function of frequency, or seismic wavelength, in order to determine the LFE rupture area.

We qualitatively explain how the ASTFs' frequency-dependent variability should reflect LFEs' rupture extents in section 2. In section 3, we present our approach in more detail. We describe how we can isolate the ASTFs from observed seismograms using an empirical Green's function approach and then how we can quantify the ASTFs' coherence among LFEs and among stations. In sections 4 and 5, we analyze ASTF coherence for individual LFEs near Parkfield and then average over thousands of LFEs to obtain well-resolved estimates of inter-station coherence as a function of frequency. For comparison, we also compute ASTF coherence for a suite of synthetic LFEs with a range of diameters and rupture velocities (section 6). Finally, in sections 7 and 8, we compare the data with the synthetics to determine which rupture areas are plausible and which types of LFEs could match the observations.

## 2 Premise: Mapping Inter-Station Similarity to Rupture Area

In order to estimate LFE areas, we note that seismic waves generated at a range of locations throughout the source region require different amounts of time to travel to the various stations. For instance, in the rupture illustrated in Figure 1d, seismic waves generated by the high-slip asperity marked in red arrive earliest at the NW station (left) because the asperity is in the northwestern half of the rupture. But waves generated at the blue asperity, located farther SE, arrive first at the SE station. The time-shifted signals give rise to apparent source time functions (ASTFs) that differ among the recording stations, as seen in Figure 1a-c.

If we assume that Earth structure is relatively uniform within the source region, we may account for the travel time variations by modeling the observed seismograms  $d_k$  in terms of station-specific apparent source time functions  $s_k$ . At each station  $k$ ,

$$\hat{d}_k(\omega) = \hat{s}_k(\omega)\hat{g}_k(\omega). \quad (1)$$

Here  $g_k$  is an average Green's function for the source area, and  $\hat{d}_k$ ,  $\hat{s}_k$ , and  $\hat{g}_k$  are the Fourier coefficients of  $d_k$ ,  $s_k$ , and  $g_k$ , respectively.

If we define  $g_k(t)$  as the Green's function for a signal generated at a reference location  $x_0$ ,  $\Delta t_k(x)$  as the source-station travel time for a signal generated at location  $x$ , and  $\dot{\delta}(x, t)$  as the slip rate as a function of location  $x$  and time  $t$ , then the ASTF can be obtained by integrating over all

161 points  $x$  within the rupture area:

$$s_k(t) = \int_{\text{rupture area}} \dot{\delta}(x, t - \Delta t_k(x) + \Delta t_k(x_0)) dA. \quad (2)$$

162 The coloring in Figure 1 shows how three slip asperities shown contribute to ASTFs that differ  
 163 among stations located to the northwest, southeast, and above the earthquake. Note that the asper-  
 164 ities create differences at all three stations even though the earthquake ruptures radially out from  
 165 the center point.

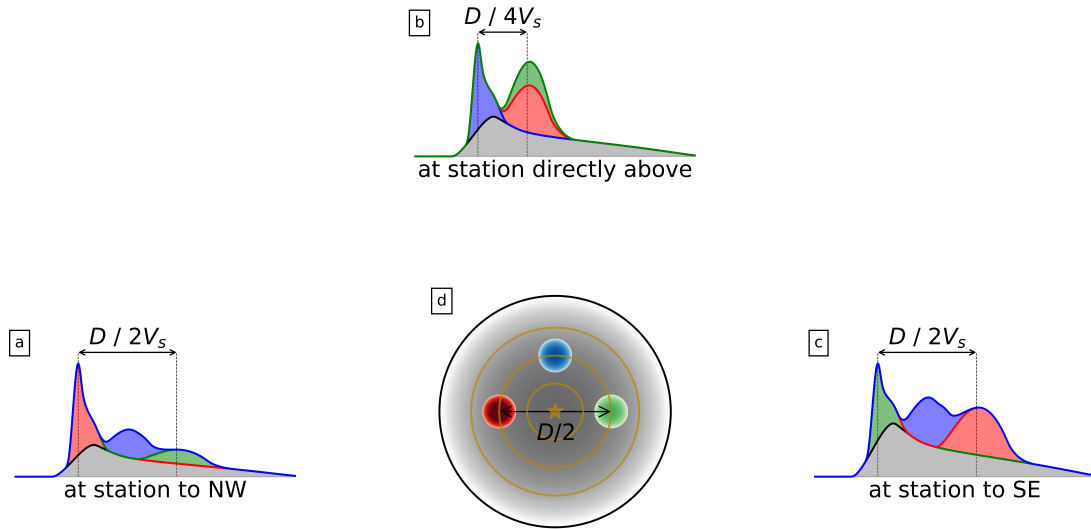


Figure 1: (a-c) ASTFs observed at 3 stations due to rupture of the slip distribution illustrated in panel (d). Rupture progresses outward from the center and moves through 3 high-slip asperities of varying magnitude, illustrated with colored circles. The asperities generate seismic waves which require different amounts of time to travel to the stations, giving rise to the various colored peaks in the ASTFs. Note that the timing of the asperity-created peaks varies among the stations by up to  $D/2V_s$ : half the rupture diameter divided by the shear wave speed.

166 There is, however, a limit to the ASTF differences. The spatially variable source-station travel  
 167 time may shift peaks in this earthquake's source time function by only a limited amount: up to  
 168  $D/V_s$ , the rupture diameter  $D$  divided by the seismic wavespeed  $V_s$ . Thus we can see differences  
 169 in the ASTFs only if we examine their short-period signal. If we examine ASTFs at periods much  
 170 longer than  $D/V_s$ , the travel time shifts will be a small fraction of the period, and the ASTFs will be  
 171 roughly the same at all stations. Synthetic rupture models described in section 6 show that ASTFs  
 172 are similar among stations at periods longer than 0.45 to 1.4 $D/V_s$ . Here the range of limiting  
 173 periods results from the earthquakes' other rupture parameters, but we note that periods at which  
 174 ASTFs are similar depend primarily on the diameter divided by seismic wave speed  $V_s$ , not on the  
 175 diameter divided by the LFEs' rupture speed  $V_r$ . We will thus be able to use the ASTFs' frequency-  
 176 dependent similarity to estimate LFE rupture extents without making restrictive assumptions about  
 177 LFE rupture dynamics.

### 3 Quantifying Coherence Across Events and Stations

#### 3.1 Removing the Path Effect

In order to examine ASTFs, we must first isolate them from the observed seismograms. To do so, we use an empirical Green's function approach similar to that of *Hawthorne and Ampuero* (2017) and compare each LFE's seismograms with a template event created via stacking (a variant on, e.g., *Mueller*, 1985; *Mori and Frankel*, 1990; *Velasco et al.*, 1994; *Hough*, 1997; *Prieto et al.*, 2004; *Baltay et al.*, 2010; *Kwiatek et al.*, 2011; *Uchide et al.*, 2014). The individual and template LFEs' seismograms  $d_{jk}$  and  $d_{tk}$  can be approximated as convolutions of ASTFs  $s_{jk}$  or  $s_{tk}$  and Green's functions  $g_k$ , so that

$$\hat{d}_{jk}(\omega) = \hat{s}_{jk}(\omega)\hat{g}_k(\omega). \quad (3)$$

To isolate the ASTFs from the Green's functions, we compute the normalized cross-spectrum  $\hat{x}_{jk}$  of the individual and template recordings:

$$\hat{x}_{jk} = \frac{\hat{d}_{jk}\hat{d}_{tk}^*}{|\hat{d}_{tk}^*|^2} = \frac{\hat{s}_{jk}\hat{s}_{tk}^*|\hat{g}_k|^2}{|\hat{s}_{tk}|^2|\hat{g}_k|^2} = \frac{\hat{s}_{jk}\hat{s}_{tk}^*}{|\hat{s}_{tk}|^2}, \quad (4)$$

where we have omitted the frequency indexing for readability. In the second equality, we have assumed that the template LFE has the same Green's functions as the individual event, so that the path effects cancel out, and we are left with a function that depends on the relative amplitudes and phases of the individual and template ASTFs. Note that we always normalize by the template amplitude, as this will allow us to stack ASTFs from thousands of LFEs. We will use the cross-spectra  $\hat{x}_{jk}$  to examine how ASTFs' amplitudes and phases vary among LFEs  $j$  and stations  $k$ .

#### 3.2 ASTF Energy: Direct and Inter-Station Coherence

As a first step, we ignore inter-station variations, and simply examine how much LFE source time functions vary from event to event. We assess the similarity between the individual and template ASTFs by computing the directly coherent power for each LFE  $j$

$$P_d = \frac{1}{N} \sum_{k=1}^N a_{jk}^2 [\text{Re}(\hat{x}_{jk})]^2 \text{sgn}[\text{Re}(\hat{x}_{jk})] \quad (5)$$

$$= \frac{1}{N} \sum_{k=1}^N a_{jk}^2 \left[ \text{Re} \frac{\hat{s}_{jk}\hat{s}_{tk}^*}{|\hat{s}_{tk}|^2} \right]^2 \text{sgn}[\text{Re}(\hat{s}_{jk}\hat{s}_{tk}^*)]. \quad (6)$$

Here the coefficients  $a_{jk}$  represent an optional weighting of the observed signals. The equality in equation (6) assumes that the individual LFE and the template have the same path effects. If the individual and template LFEs also have similar and well-aligned ASTFs  $\hat{s}_{jk}$  and  $\hat{s}_{tk}$ , the value  $\hat{s}_{jk}\hat{s}_{tk}^*$  in equation (6) will be real and positive. The directly coherent power  $P_d$  is thus also positive when the individual and templates ASTFs are the same. Its amplitude is determined by the relative power of the individual and template ASTFs.

The relative ASTF power also determines the amplitude of the inter-station coherent power  $P_c$ . With this power calculation, we seek to ignore ASTF variations across events, and instead assess

207 the ASTFs' similarity across stations. So we compute

$$P_c = \frac{2}{N(N-1)} \sum_{k=1}^N \sum_{l=k+1}^N a_{jk} a_{jl} \operatorname{Re}(\hat{x}_{jk} \hat{x}_{jl}^*) \quad (7)$$

$$= \frac{2}{N(N-1)} \sum_{k=1}^N \sum_{l=k+1}^N a_{jk} a_{jl} \operatorname{Re} \frac{(\hat{s}_{jk} \hat{s}_{jl}^*) (\hat{s}_{tk}^* \hat{s}_{tl})}{|\hat{s}_{tk}|^2 |\hat{s}_{tl}|^2}, \quad (8)$$

208 where the second equality again assumes common path effects. As noted in section 2, the ASTFs  
 209 are expected to be the same for all stations if the period being considered with these Fourier coeffi-  
 210 cients is long compared with  $D/V_s$ , the intra-source seismic wave travel time. If the ASTFs are the  
 211 same across stations at the period of interest, we will have  $\hat{s}_{jk} = \hat{s}_{jl}$  and  $\hat{s}_{tk} = \hat{s}_{tl}$ , so that  $\hat{s}_{jk} \hat{s}_{jl}^*$ ,  
 212  $\hat{s}_{tk}^* \hat{s}_{tl}$ , and finally  $P_c$  are all real and positive.

213  $P_d$  and  $P_c$  thus give us estimates of the direct or inter-station coherent power of an LFE, as  
 214 normalized by the template power. However, we can obtain a more interpretable normalization if  
 215 we also estimate the full template-normalized LFE power, including any incoherent components:

$$P_l = \frac{1}{N} \sum_{k=1}^N a_{jk}^2 |\hat{x}_{jk}|^2 \quad (9)$$

$$= \frac{1}{N} \sum_{k=1}^N a_{jk}^2 \frac{|\hat{s}_{jk}|^2}{|\hat{s}_{tk}|^2}. \quad (10)$$

216 We will use the LFE power  $P_l$  to normalize  $P_d$  and  $P_c$  and compute the fraction of the power that  
 217 is coherent across events and stations.

## 218 4 Calculating Powers of Parkfield LFEs

219 When we extract the coherent and incoherent powers of LFEs near Parkfield, we will also have to  
 220 estimate and remove the noise, and we will have to average over thousands of LFEs to obtain well-  
 221 resolved powers. To begin, we describe the LFE catalog and seismic data (section 4.1) and create  
 222 templates for seven LFE families (section 4.2). Then we demonstrate our approach by estimating  
 223 template-normalized powers for an individual LFE (section 4.3). Finally, we average the powers  
 224 over the LFEs in each family (section 5).

### 225 4.1 Data and LFE Families

226 The LFEs considered here occurred between 2006 and 2015 at depths of 16 to 23 km near Parkfield,  
 227 CA (see Figure 2). They were identified via cross-correlation by *Shelly* (2017) as part of his 15-  
 228 year tremor catalog and are grouped into seven families numbered 37140, 37102, 70316, 27270,  
 229 45688, 77401, and 9707, with 2500 to 8300 LFEs in each family (see also *Shelly et al.* (2009);  
 230 *Shelly and Hardebeck* (2010)). LFEs in families 37140 and 37102 were examined by *Thomas*  
 231 *et al.* (2016) and found to have best-fitting source durations of 0.19 and 0.22 s, respectively. We  
 232 use LFE seismograms from 17 borehole seismic stations in the Berkeley HRSN (High Resolution

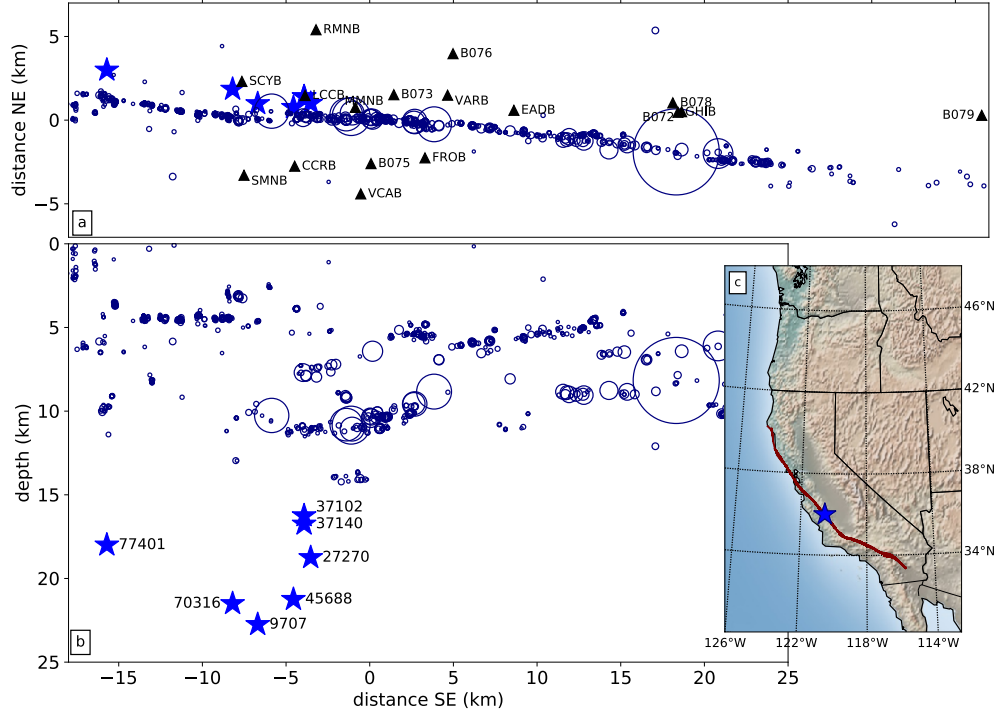


Figure 2: (a) Map view and (b) depth section of the LFE families (blue stars), local  $M > 2.5$  earthquakes (circles), and the HRSN and PBO seismic stations used (triangles). Earthquake sizes are scaled to the radii expected for 3-MPa stress drops, and locations are taken from the NCEDC catalog and the relocations of *Waldhauser (2009)*.

233 Seismic Network) and in the PBO (Plate Boundary Observatory) network. Since this analysis  
 234 relies on high-quality recordings of small LFEs, we correct the data for some errors identified by  
 235 *Shelly (2017)*. We have also gone through the data from each station and channel and discarded  
 236 weeks- to years-long intervals where the LFE amplitudes are much more scattered than usual, as  
 237 these intervals likely have larger-than-average noise.

## 238 4.2 Stacked LFE Templates

239 For each LFE family, we create a low-noise template by averaging the LFE records for each chan-  
 240 nel. We bandpass filter the LFE seismograms from 2 to 30 Hz, normalize them by their maximum  
 241 values, and then average, weighting each record by the station-averaged cross-correlation coeffi-  
 242 cient obtained by *Shelly (2017)*. Then we rescale these normalized stacks so that their amplitudes  
 243 match the amplitudes of individual records, as described in section S2. We iterate the stack a few  
 244 times to slightly improve the signal to noise ratio, each time discarding records with very small or  
 245 unusual amplitudes (for details see section S2).

246 We estimate the signal to noise ratio of the stacks using a 3-second window starting just before  
 247 the S arrival. We keep only the stacks which have average amplitude spectra at least 3 times larger  
 248 than the noise in the 2 to 10 Hz band. The procedure leaves us with 16 to 29 well-resolved template  
 249 seismograms for each LFE family, observed on the two horizontal components of 9 to 16 stations.  
 250 Some are shown in Figure 3a, as well as in Figures S1 to S7.



### 4.3 Coherent and Total Powers for One LFE

We will use the obtained templates to remove the Green’s functions from individual LFE records, so that we can probe the LFEs’ ASTFs. To prepare, we realign each LFE’s origin time to better match the template, as poor alignment could reduce the direct coherence  $P_d$ . We bandpass filter to 2 to 5 Hz, cross-correlate to obtain a preferred shift at each station, and then shift the seismograms of all stations by the median shift.

Next, we remove the path effects for the power calculations. We extract 3-second-long segments of the template seismograms, starting just before the S arrival, and cross-correlate the segments with the individual LFE records. The individual LFE records are truncated 0.2 seconds before the S arrival to reduce contamination by the P arrival, but they are not truncated after the S wave. We average the cross-correlations over the available channels at each station.

Cross-correlations obtained for one LFE are illustrated in Figure 3b. The cross-correlations are often roughly but not entirely symmetric, suggesting that the individual and template LFEs have slightly different source time functions. The asymmetry is also apparent in the non-zero phases of the cross-correlations’ Fourier coefficients, which are equal to the phases of the normalized cross-spectra  $\hat{x}_{jk}$  (equation (4), Figure 3c). To estimate the  $\hat{x}_{jk}$ , we first extract a 6-second portion of the cross-correlations, multiply by a Slepian taper concentrated at frequencies lower than 0.4 Hz, and compute the Fourier transform (Thomson, 1982). Then we normalize; we divide by the Fourier transform of the template seismograms’ autocorrelation, computed via the same procedure.

We use the cross-spectra to compute the power that is directly coherent ( $P_d$ , equation (5)) and coherent among stations ( $P_c$ , equation (7)) and plot them in yellow and red in Figure 3d. The total power  $P_t$  in the template-normalized cross-correlation is also computed, following equation (9), and is plotted in green. However, a significant fraction of this total power comes from noise, not from the LFE signal. To estimate the noise contribution, we cross-correlate the template seismograms with data from noise intervals starting 8 seconds before the S arrivals. We compute the power ( $P_n$ ) in those noise correlations, again following equation (9), and plot it in gray in Figure 3d. Finally, we subtract the noise power  $P_n$  from the total power  $P_t$  to determine the power contributed by the LFE ( $P_l$ , blue in Figure 3d).

In all the power calculations, we use weightings  $a_{jk}$  equal to one over the standard deviation of the 2 to 30-Hz filtered waveform, as computed in the four seconds ending 0.5 s before the LFE S arrival. This weighting allows us to downweight records with large noise, but it does not bias our results because all of the power in  $P_t$ ,  $P_c$ , and  $P_d$  comes from after 0.2 s before the S arrival and because almost all of the subtracted noise power  $P_n$  comes from more than 5 seconds before the S arrival.

In an ideal scenario, we would now interpret the estimated powers, and compare the coherent powers  $P_d$  and  $P_c$  with the LFE power  $P_l$ . However, for this and other individual LFEs, the powers are too poorly resolved to allow direct interpretation. In Figure 3d, the ratios  $P_d/P_l$  and  $P_c/P_l$  vary by tens of percent among the frequencies but show no systematic trend, and there is further variation if we use different subsets of the stations. So in the next section, we will average the powers over several thousand LFEs to obtain well-resolved and stable coherent power fractions.

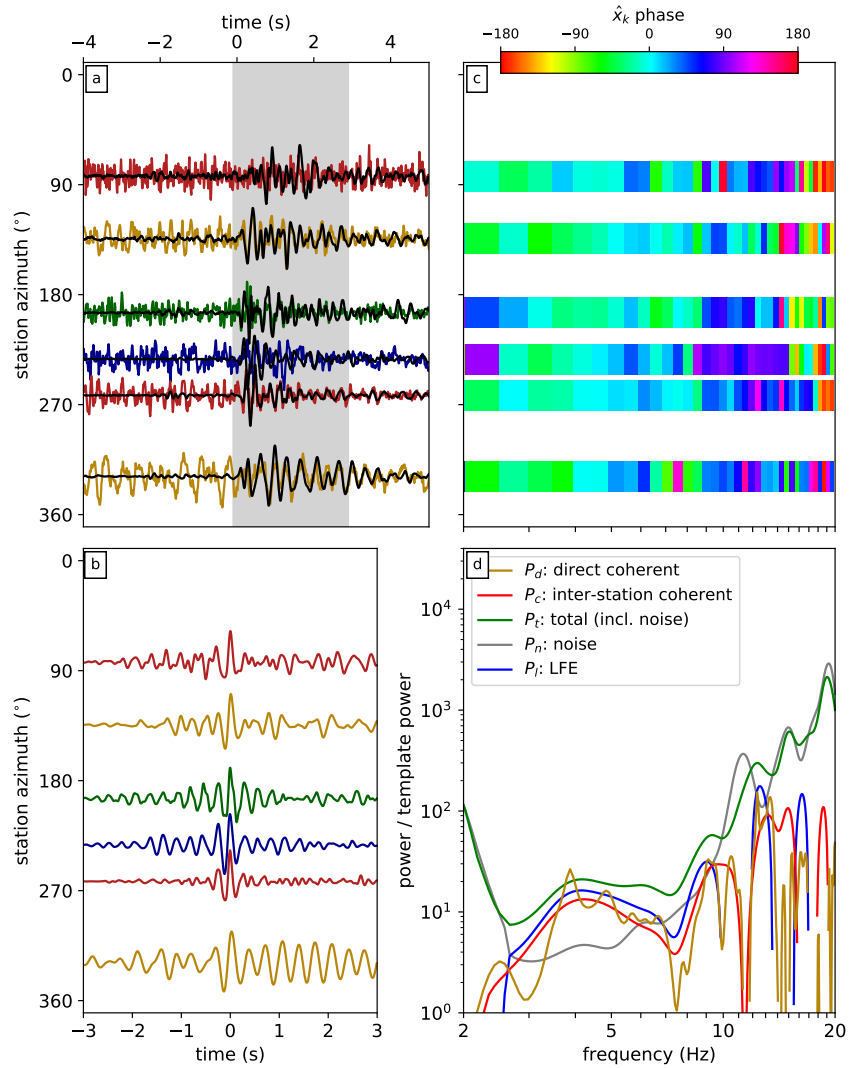


Figure 3: (a) Some of the template seismograms (black) for family 37102 along with seismograms observed for one LFE (color). Traces are organized according to the station’s azimuth relative to the LFE and are scaled to their maximum value. The gray shading indicates the portion of the template that is correlated with the individual observations. (b) Cross-correlations of the observed seismograms with the template. (c) Phase of the cross-spectra: of the Fourier coefficients of the cross-correlations in panel b. (d) Yellow, red, and green curves:  $P_d$ ,  $P_c$ , and  $P_t$ —the coherent and total template-normalized powers from the LFE interval. Gray:  $P_n$ —the noise power, computed in an interval without the LFE. Blue:  $P_l = P_t - P_n$ —the power likely contributed by the LFE. Note that with just this one LFE, it is not practical to interpret the relative values of the coherent and total powers.

## 5 Results: Event-Averaged Coherent and Incoherent Powers

To estimate  $P_c$ ,  $P_d$ ,  $P_t$ , and  $P_n$  for a given family of LFEs, we compute the powers for individual events and then average. However, some LFE records have exceptionally large noise, so we check the signals' amplitudes before the calculation and discard records when the S arrival or the preceding noise interval has standard deviation that differs by more than a factor of 10 from that channel's median. This record selection, coupled with data availability, leaves us with 860 to 4220 LFEs per family which have template-normalized powers computed from at least 5 stations.

Figure 4a shows the summed coherent and total powers obtained from 2000 LFEs in family 37140, one of the two families with duration estimates from *Thomas et al.* (2016). The shading indicates 95% uncertainty ranges on the powers, obtained by bootstrapping the LFEs included in the summation. All of the template-normalized powers increase with frequency, suggesting that the high-frequency template power is damped relative to a typical LFE. The stacks' high-frequency signal may be averaged out by stacking if LFEs are more different at higher frequencies or if the LFE timing is not accurate enough to allow coherent stacks at higher frequencies. The stacking effectively creates a template LFE which has slightly broader and simpler ASTFs (*Royer and Bostock*, 2014). Note that this ASTF modification may reduce the direct coherence between the template and the individual LFEs  $P_d$ , but it should not affect the inter-station coherence  $P_c$ , as  $P_c$  is independent of inter-event ASTF differences.

We compute the coherent power fractions  $P_d/P_l$  and  $P_c/P_l$  for all 7 families and plot the results in Figure 4b-h. For family 37140 (panel b), the direct coherence  $P_d/P_l$  is larger than 0.8 at frequencies of 2 to 4 Hz, suggesting that most 0.2-second-long LFE source time functions are similar when viewed at these frequencies. We should note, however, that  $P_d/P_l$  may be slightly higher than its true value in this range because we allowed for an LFE origin time shift using data in the 2 to 5-Hz range.  $P_d/P_l$  decreases at higher frequencies, falling below 0.6 at a frequency of 5 Hz. The decrease in direct coherence could imply (1) that the LFE source time functions are more different at higher frequencies, (2) that the LFEs are too poorly aligned to show direct coherence at high frequencies, or (3) that the stacking has modified the source time functions being compared. We have tried improving the alignment by using higher-frequency signals in the alignment cross-correlation, outside the 2 to 5-Hz range. We find that the high-frequency signals does result in large  $P_d/P_l$  out to higher frequencies, but we choose not to use it here because some of the increase in  $P_d/P_l$  could come from the alignment of high-frequency noise.

Family 37140's inter-station coherent power  $P_c/P_l$  is insensitive to the alignment, and it remains coherent over a wider frequency range.  $P_c/P_l$  is above 0.8 at frequencies up to 12 or 15 Hz and falls below 0.6 only at 16.5 Hz. The persistence of high  $P_c/P_l$  out to frequencies  $>15$  Hz suggests that the ASTFs vary little among stations at  $>0.07$ -second periods. We will use synthetic rupture calculations to interpret this high-frequency coherence in terms of LFE rupture area in section 7.

The other six LFE families show similar or slightly lower coherence, as seen in Figure 4c-h and in Figures S8 - S14. Family 37102, the other family with an estimated duration (*Thomas et al.*, 2016), displays gradually decaying  $P_d/P_l$  and  $P_c/P_l$  (Figures 4b and S9). Its  $P_d/P_l$  falls below 0.6 at 4 Hz, and its  $P_c/P_l$  stays above or hovers near 0.6 until 9 Hz. For the remaining families, the direct coherence  $P_d/P_l$  remains above 0.6 out to 4 to 5 Hz. The inter-station coherence  $P_c/P_l$  remains above 0.6 out to 8 to 13 Hz: to 8, 9, 11, 12, and 13 Hz.

These high-coherence frequency limits are likely lower bounds on the true high-coherence fre-

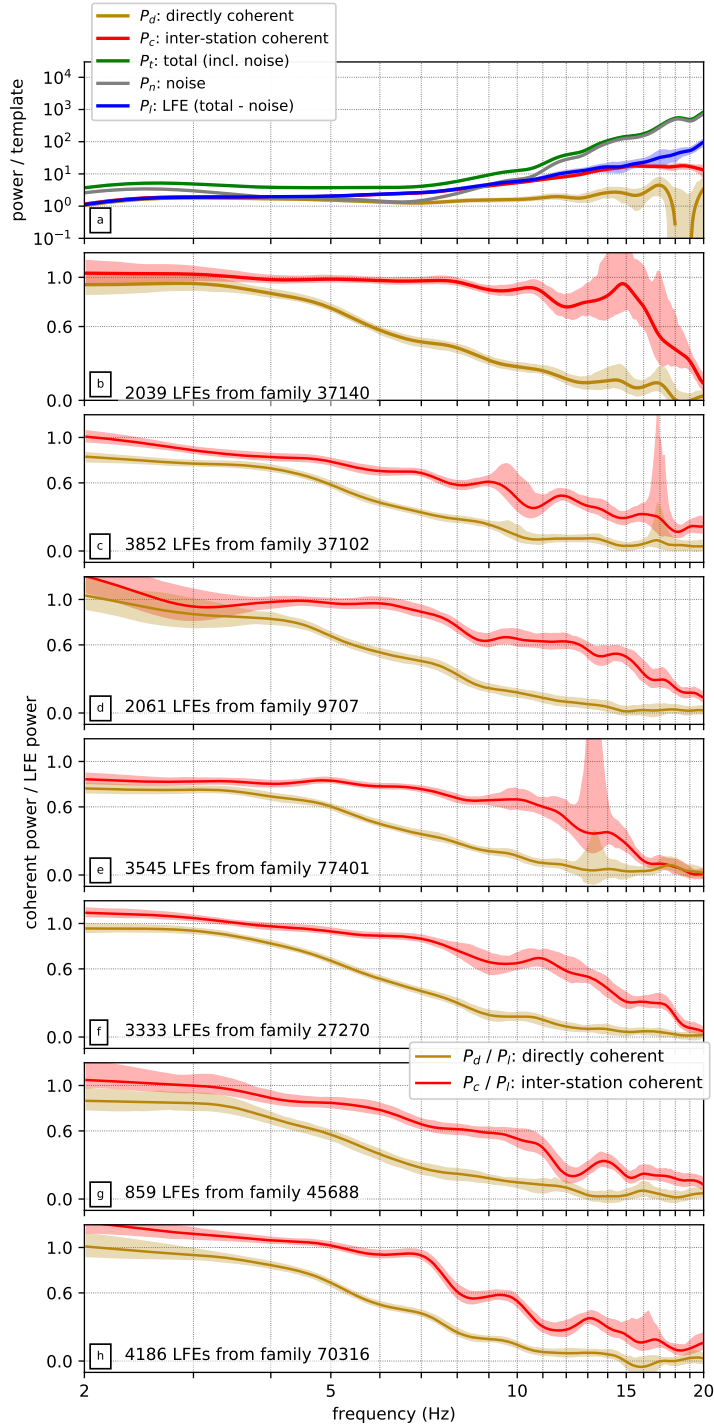


Figure 4: (a) Coherent and incoherent powers, as in Figure 3d, but averaged over 2000 LFEs from family 37140. Color indicates the power of interest. In all panels, the line indicates the value obtained with all allowable LFEs, and the shaded region delimits 95% confidence intervals obtained by bootstrapping the included events. (b-h) Ratios of the direct and inter-station coherence:  $P_c/P_l$  (yellow) and  $P_d/P_l$  (red). Each panel is computed for a different LFE family, as indicated by the text in the bottom left.

335 quencies. Our coherence estimates could be affected by several factors, including LFE clustering,  
 336 data selection, LFE origin time alignment, and template accuracy. We describe the uncertainties  
 337 in Appendix A1 and note that only the LFE origin time alignment is likely to give artificially high  
 338 coherence, and it affects only  $P_d/P_l$ , not  $P_c/P_l$ . The remaining factors would result in our un-  
 339 derestimating the true  $P_d/P_l$  and  $P_c/P_l$ . In section 7, we will therefore interpret our coherence  
 340 estimates as lower limits when we consider their implications for LFE rupture areas and location  
 341 distributions.

## 342 6 Frequencies With Coherent Power: Synthetics

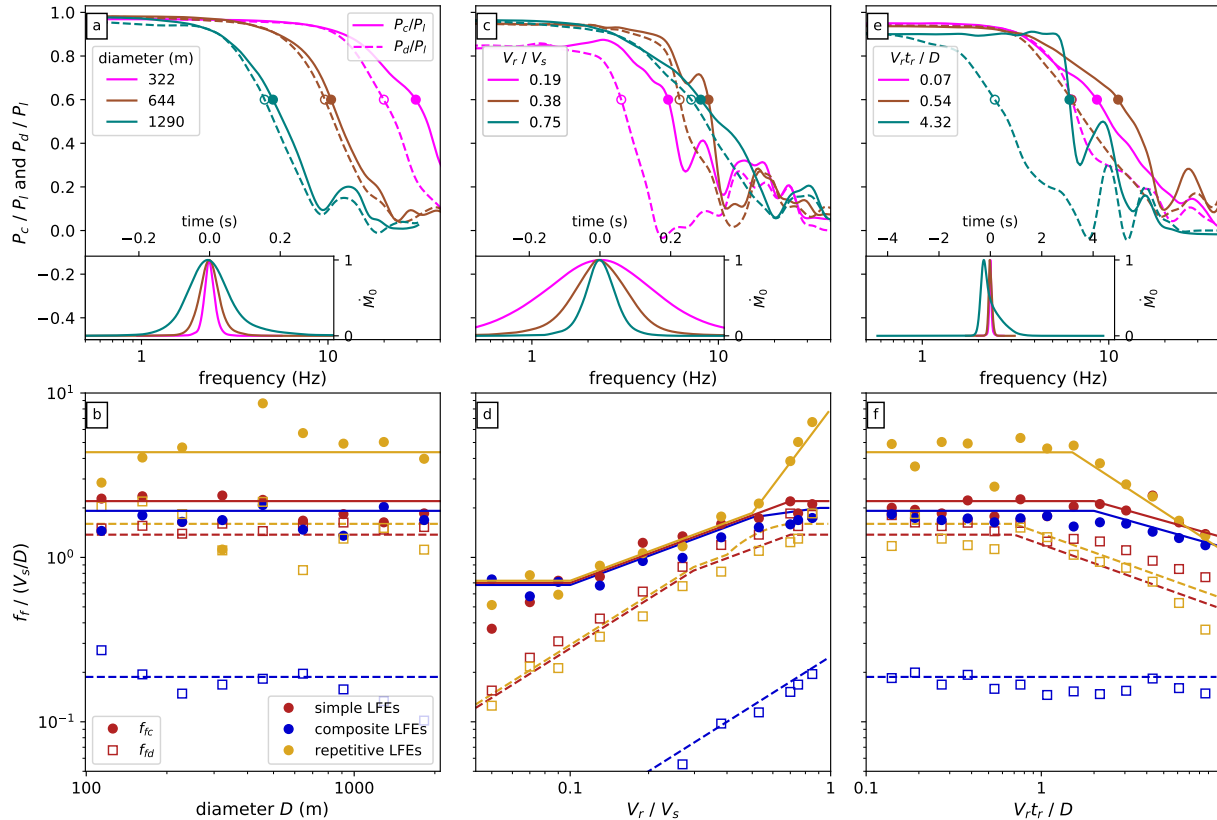


Figure 5: (a, c, e) Coherent power fractions  $P_c/P_l$  (solid lines) and  $P_d/P_l$  (dashed lines) as a function of frequency for various groups of synthetic LFEs. Circles mark the coherence falloff frequencies: when  $P_c/P_l$  or  $P_d/P_l$  falls below 0.6. Inset panels show the moment rate functions averaged over LFEs in each group. Color indicates diameter (panel a), rupture velocity (panel c), and rise time (panel e). (b, d, f) Normalized coherence falloff frequencies  $f_{fc}/(V_s/D)$  (filled circles) and  $f_{fd}/(V_s/D)$  (open squares) as a function of the LFE properties. Color indicates the type of LFE rupture. Solid and dashed lines indicate approximations of the numerically identified  $f_{fc}$  and  $f_{fd}$  to be used in our interpretations. In panels a, b, c, and d,  $t_r = 0.27D/V_r$ . In panels a, b, e, and f,  $V_r = 0.75V_s$ . In panels c and e,  $D = 456$  m. In panels d and f, the values plotted are medians taken from synthetics with 7 different diameters.

343 To consider the coherence’s implications for LFE rupture areas, we need to know how  $P_a/P_l$   
344 and  $P_c/P_l$  depend on LFE rupture properties. So we generate and analyze groups of synthetic LFEs  
345 with various diameters  $D$ , rupture velocities  $V_r$ , and rise times  $t_r$ . We create synthetic ruptures for  
346 three types of LFEs (section 6.1), analyze their waveforms (section 6.2), and examine the coherent  
347 frequencies as a function of the LFE properties (section 6.3).

## 348 **6.1 Synthetic LFEs Models**

349 We create and analyze groups of 100 LFEs. The individual events are assigned diameters  $D$ ,  
350 rupture velocities  $V_r$ , and rise times  $t_r$  that cluster around specified mean values. The diameters,  
351 rupture velocities, and rise times are chosen from lognormal distributions with factor of 1.3, 1.1,  
352 and 1.3 standard deviations, respectively. Moments are chosen from lognormal distributions with  
353 factor of 1.5 standard deviation and assigned with no consideration of the radii.

354 In the simplest version of our LFEs, each event is assigned a random heterogeneous slip dis-  
355 tribution within a roughly circular area, as detailed in section S4 and motivated by inferences of  
356 fractal earthquake slip distributions (*Frankel, 1991; Herrero and Bernard, 1994; Mai and Beroza,*  
357 *2002*). Rupture initiates at a random location within  $0.4D$  of the center and spreads radially at rate  
358  $V_r$ . Once a location starts slipping, slip accumulates following a regularized Yoffe function with  
359 duration  $t_r$  (*Tinti et al., 2005*).

360 We also construct groups of LFEs with more repetitive rupture patterns, as it is possible that  
361 LFEs within a given family recur not just on the same patch, but with similar rupture patterns within  
362 that patch (e.g., *Ariyoshi et al., 2009; Ando et al., 2010; Sweet et al., 2014; Chestler and Creager,*  
363 *2017b*). In our repetitive LFEs, slip is the sum of two heterogeneous distributions: one that varies  
364 randomly from event to event and one that is the same from event to event. The distributions  
365 are scaled so that the repetitive component contributes twice as much moment, and slip always  
366 nucleates within  $0.1D$  of the LFE center points.

367 Finally, we construct groups of composite LFEs, as it is possible that individual LFEs comprise  
368 a series of small ruptures of the complex fault zone at depth (*Fagereng et al., 2014; Hayman and*  
369 *Lavier, 2014; Chestler and Creager, 2017b; Rubin and Bostock, 2017*). Each of our relatively  
370 crude composite LFE contains five simple ruptures whose rupture velocities, diameters, and slip  
371 distributed are chosen from the lognormal and heterogeneous distributions described above. The  
372 five sub-ruptures begin at random times within a  $2.5D/V_r$  interval.

## 373 **6.2 Computing and Analyzing LFE Waveforms**

374 Having defined the location and timing of slip in the LFEs, we compute ASTFs for nearby stations.  
375 We assume that the synthetic LFEs are in the location of family 37140 and calculate ASTFs for the  
376 12 stations used in its analysis, as shown in Figures 2 and S1. To calculate ASTFs, we integrate  
377 the slip rate over the slipping area at each time step, but shift the signals’ arrival times to account  
378 for the travel time from each point in the source region to the observing stations, as in equation (2).  
379 To calculate seismograms, we convolve these ASTFs with fake Green’s functions, which are taken  
380 to be white noise tapered by an exponential with a 3-s decay constant.

381 We may now process the synthetic seismograms. As with the real data, we create templates for  
382 each LFE group, normalizing the synthetic seismograms by their maximum values and stacking.  
383 We iterate this stack three times. Each time, we cross-correlate the template seismograms with

384 the individual LFEs’ waveforms. We identify a station-averaged time shift for each LFE, realign  
 385 according to those shifts, and stack.

386 Next, we use the templates to compute the cross-spectrum  $\hat{x}_{jk}$  for each synthetic LFE record  
 387 (equation (4)). As with the real data, we compute the cross-spectra from the tapered cross-  
 388 correlations, but we adjust the taper duration to ensure that it is always significantly longer than  
 389 the LFEs’ durations. Finally, we compute the LFEs’ template-normalized powers  $P_c$ ,  $P_d$ , and  $P_l$   
 390 (equations (5), (7), and (9)). Figure 5a, c, and e shows the coherent power fractions  $P_d/P_l$  and  
 391  $P_c/P_l$  obtained for simple LFEs with various diameters, rupture velocities, and rise times.

## 392 **6.3 Coherence Falloff Frequencies as a Function of $D$ , $V_r$ , and $t_r$**

### 393 **6.3.1 Coherence Falloff with Diameter**

394 As anticipated in section 2, both  $P_d/P_l$  and  $P_c/P_l$  decrease at lower frequencies (longer periods)  
 395 when the LFE diameters are larger (panel a).  $P_d/P_l$  falls off earlier when diameters are larger  
 396 because larger diameters imply longer ruptures, which allow for complexity and inter-LFE vari-  
 397 ability at lower frequencies.  $P_c/P_l$  falls off earlier because larger diameters imply larger shifts  
 398 in the source-station travel time within the rupture area, and thus allow for inter-station ASTF  
 399 variability at lower frequencies. To examine the coherence falloff systematically, we identify the  
 400 frequencies at which  $P_d/P_l$  and  $P_c/P_l$  first fall below 0.6. These falloff frequencies  $f_{fd}$  and  $f_{fc}$  are  
 401 normalized by  $V_s/D$  and plotted as a function of LFE diameter  $D$  in Figure 5b. In the simple LFE  
 402 simulations in Figure 5b, which have  $V_r/V_s = 0.75$  and  $t_r = 0.27R/V_r$ ,  $f_{fd}$  is roughly  $1.4V_s/D$   
 403 (open red squares and dashed red line), and  $f_{fc}$  is roughly  $2.2V_s/D$  (filled red circles and solid red  
 404 line).

405 Note that these ratios  $f_{fd}/(V_s/D)$  and  $f_{fc}/(V_s/D)$  could change slightly if we assumed a  
 406 different distribution of stations, as different takeoff angles and azimuths could change the apparent  
 407 source durations and travel time shifts. Here we have chosen a station distribution consistent with  
 408 the stations used in analyzing the Parkfield data.

### 409 **6.3.2 Coherence Falloff with Rupture Velocity**

410 The direct coherence falloff frequency  $f_{fd}$  decreases relative to  $V_s/D$  if LFE rupture velocities  
 411 are reduced, as shown Figure 5c and d. Note that when we plot  $f_{fd}/(V_s/D)$  and  $f_{fc}/(V_s/D)$  in  
 412 Figure 5d and f, we take the median of estimates computed for 7 groups of LFEs, with different  
 413 diameters, in order to reduce the scatter. The decrease of  $f_{fd}/(V_s/D)$  with decreasing rupture  
 414 velocities arises because lower rupture velocities allow for longer ruptures and therefore more  
 415 complexity and inter-event variability at lower frequencies. The LFEs’ heterogeneous slip distri-  
 416 butions give rise to source time functions that differ among events at all frequencies shorter than  
 417 the rupture duration, which scales as  $D/V_r$  in simulations of simple LFEs. The direct coherence  
 418 falloff frequency  $f_{fd}$  thus scales inversely with the durations of these ruptures, with value around  
 419  $2.8V_r/D$  when  $V_r < 0.4V_s$ , though it decreases relative to  $V_r/D$  for rupture velocities larger than  
 420  $0.8V_s$  (red dashed line in Figure 5d).

421 The inter-station coherence falloff frequency  $f_{fc}$  depends more weakly on rupture velocity  $V_r$ .  
 422  $f_{fc}$  increases from  $0.7$  to  $2.2V_s/D$  as  $V_r$  increases from  $0.05$  to  $1V_s$  (filled red circles and solid  
 423 red line in Figure 5d).  $P_c/P_l$  depends only weakly on  $V_r$  because  $P_c/P_l$  measures how much

424 the ASTFs vary among stations, not among events. The inter-station ASTF variability depends  
425 primarily on the S-wave travel time across the source region, which scales with  $D/V_s$ , not  $D/V_r$ .  
426 The  $V_r$  dependence that does exist likely results from the simpler ASTF pulses associated with  
427 higher rupture velocities. As  $V_r$  approaches  $V_s$ , the ASTFs tend toward single pulses, and inter-  
428 station complexity is harder to distinguish.

### 429 **6.3.3 Coherence Falloff With Rise Time**

430 Both  $f_{fd}$  and  $f_{fc}$  vary minimally in response to modest changes in the duration  $t_r$  of slip at each  
431 point in the rupture, especially when the rise time  $t_r$  is less than  $D/V_r$  (Figure 5c and f). In  
432 our implementation, we have assumed a spatially uniform rise time for each LFE. As a result,  
433 changing the rise time is roughly equivalent to convolving all of an LFE's ASTFs by a single  
434 function, and such a convolution has little effect on the inter-ASTF coherence. We do allow roughly  
435 10% variability in rise time and rupture velocity among the LFEs in each group. These rise time  
436 differences, coupled with the increased complexity visible in longer-duration ruptures, are likely  
437 responsible for the reduced coherence falloff frequencies that become apparent once  $t_r$  exceeds 1  
438 to  $2D/V_r$ . (red symbols and lines in Figure 5d).

### 439 **6.3.4 LFE Durations**

440 Increasing the rise time does increase LFE durations. To estimate durations for each LFE group,  
441 we stack the events' source time functions using the same alignment used in the stack creation.  
442 Then we identify the time interval containing the central 70% of the LFE moment. In our simple  
443 LFEs, these intervals have durations that increase from between  $0.29$  and  $0.31D/V_r$  when  $t_r$  is  
444  $0.27D/V_r$  to roughly  $0.28t_r$  as  $t_r$  gets longer than  $D/V_r$ . LFE durations are shorter when we  
445 require that LFEs nucleate near the rupture centers. For our modeled repetitive LFEs, which we  
446 assume nucleate within  $0.1D$  of their center points, durations are  $0.25$  to  $0.28D/V_r$  when  $t_r$  is  
447  $0.27D/V_r$ . If we instead specify nucleation locations within  $0.1D$  of the rupture edge, durations  
448 are  $0.35$  to  $0.37D/V_r$ . The durations of composite LFEs are determined by the number and timing  
449 of subevents. The presented LFEs, containing 5 subevents, have durations between  $3$  and  $3.3D/V_r$ .

### 450 **6.3.5 Composite LFEs**

451 The composite LFEs, with their long, complex ruptures, have reduced direct coherence. The direct  
452 coherence falloff frequency  $f_{fd}$  is around  $0.25V_r/D$  for all simulated events (open blue squares  
453 and dashed lines in Figure 5b, d, and f). On the other hand, the inter-station falloff frequencies  
454  $f_{fc}$  are similar for simple and composite LFEs (filled blue circles and solid blue line). Here again  
455  $P_c/P_l$  depends primarily on  $D/V_s$ : on how much the source-station travel time can shift peaks in  
456 the source time functions.

### 457 **6.3.6 Repetitive LFEs**

458 The coherent power fractions  $P_d/P_l$  and  $P_c/P_l$  can be significantly higher for repetitive LFEs,  
459 at least when the rupture velocity is larger than about  $0.5V_s$ . As described in section 6.1, the  
460 repetitive LFEs in each group have similar slip distributions, and they all nucleate near the rupture  
461 center, so they have similar ASTFs and similar waveforms. This similarity explains the increase in



462  $P_d/P_l$ , but the increase in  $P_c/P_l$  is surprising at first glance, as  $P_c/P_l$  measures similarity across  
463 stations, not across events. However, the cross-spectra calculation that goes into  $P_c$  (equation (4))  
464 is designed to remove complexity associated with the path effects, and it identifies as “path effect”  
465 any component of the source-path convolution (equation (3)) that is common to all events. If  
466 the ASTFs are the same for all events, the  $P_c$  calculation cannot distinguish inter-station ASTF  
467 variations from station-dependent Green’s functions, ASTF variations are thus attributed to path  
468 effects, and  $P_c/P_l$  is high when LFEs are highly repetitive. The falloff frequencies  $f_{fc}$  can increase  
469 by a factor of 6 when  $V_r > 0.8V_s$ .

470 We note, however, that this factor of 6 increase in  $f_{fc}$  is just one plausible value. Here we have  
471 assumed that two-thirds of the LFE moment came from a repetitive component of the rupture, but  
472 higher or lower coherence could be achieved by assuming that more or less of the moment came  
473 from the repetitive component. We also note that the high coherence arises only when the rupture  
474 nucleation location is consistent from event to event. The falloff frequencies  $f_{fc}$  remain low if only  
475 75% of the repetitive LFEs nucleate at the SE rupture edge and the other 25% nucleate on the NW  
476 edge (Figure S22).

## 477 **7 Interpretation of LFE Coherence**

478 We may now use our synthetic results to interpret the coherence obtained for the Parkfield LFE  
479 families, which show direct coherence  $P_d/P_l > 0.5$  out to 4 to 5 Hz and inter-station coherence  
480  $P_c/P_l > 0.5$  out to 8 to 16.5 Hz.

### 481 **7.1 LFE Location Distribution**

482 First, we note that the observed high-frequency coherence implies that LFEs within each family are  
483 strongly clustered in space. If LFEs were distributed over a wide range of locations, travel times  
484 from the LFE centroids to the recording stations would vary widely from event to event. But in our  
485 analysis, we allow only the origin time to be realigned from event to event. Any inter-station time  
486 shifts produced by varying LFE locations should show up in our results as a decrease in coherence.

487 To determine the maximum location variation allowed by the observations, we recompute co-  
488 herence values after artificially shifting the LFE locations by various amounts. We pick location  
489 shifts for each LFE in family 37140, drawing from bivariate normal distributions with 100-m to  
490 1-km standard deviations along strike and depth. We use the IASP91 velocity model and TauP  
491 to compute the arrival time change for the stations observing each LFE (*Kennett and Engdahl,*  
492 *1991; Crotwell et al., 1999*). We subtract the median arrival time change from these values, shift  
493 the seismograms by the station-dependent remainders, and compute the coherent power fractions.  
494 The family-averaged results are shown in Figures 6 and S15-S17. We find that the inter-station  
495 coherent fraction  $P_c/P_l$  obtained at 11 Hz is reduced by 40% even for location shifts with just  
496 250-m standard deviation (Figure 6). The  $> 0.6$  11-Hz coherence values obtained for the median  
497 family thus imply that LFEs in each family are strongly clustered, with standard deviation in their  
498 locations typically smaller than 250 m.

499 Note that the distribution of LFE locations within a family, when coupled with noise, is one  
500 way to explain all of the incoherence observed at higher frequencies in the data. It is possible that

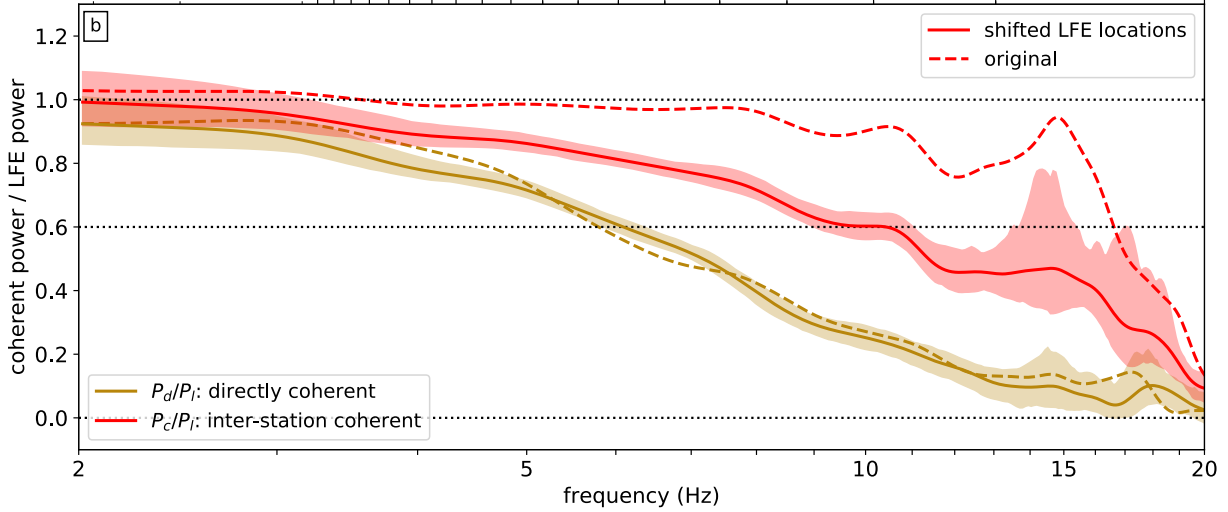


Figure 6: Solid lines and shading: coherent power fractions for family 37140, as in Figure 4b, but computed after shifting the LFE locations by random amounts with 250-m standard deviations along strike and along depth. Dashed lines: original  $P_d/P_l$  and  $P_c/P_l$ , without location shifts, reproduced from Figure 4b.

501 the individual LFEs appear to be point sources—that each LFE ruptures a tiny patch within the  
 502 sub-1-km asperity (Chestler and Creager, 2017a).

## 503 7.2 Matching $f_{fc}$ , $f_{fd}$ , and Duration With Simple Ruptures: Results

504 However, it is also possible that the finite rupture areas of individual LFEs contribute to the de-  
 505 crease in coherence at high frequencies. To determine the maximum rupture areas and rupture  
 506 velocities allowed by the data, we compare the observed coherence falloff frequencies and dura-  
 507 tions with those obtained from synthetics of simple, non-repetitive ruptures.

508 First, we note that the inter-station coherence  $P_c/P_l$  remains higher than 0.6 out to 8 to 16.5 Hz  
 509 for the various families. The median  $P_c/P_l$  falloff frequency  $f_{fc}$  is 11 Hz, and families 37102  
 510 and 37140, which we will discuss in more detail, have  $f_{fc}$  of 9 and 16.5 Hz, respectively. In the  
 511 synthetics,  $f_{fc}$  is 0.7 to  $2.2V_s/D$  for rupture velocities  $V_r$  between 0.05 and  $1V_s$  (red solid line  
 512 in Figure 5d). If the shear wave velocity  $V_s$  is around 4 km/s in the LFE area (Lin et al., 2010),  
 513 family 37102’s 9-Hz  $f_{fc}$  implies an average diameter smaller than 300 to 1000 m, with smaller  
 514 allowable diameters for slower rupture velocities. In Figure 7a, this range of allowable diameters  
 515 is marked with blue diagonal hatching. The thicker blue line presents alternative upper bound on  
 516 the diameter, ranging from 180 to 550 m. It is appropriate for a 16-Hz  $f_{fc}$ : the bootstrap-estimated  
 517 lower bound on  $f_{fc}$  obtained for family 37140.

518 The yellow diagonal hatching in Figure 7a illustrates a further, albeit weaker, constraint on the  
 519 LFEs’ diameters and rupture velocities: those obtained from the direct coherence  $P_d/P_l$ .  $P_d/P_l$   
 520 is higher than 0.6 out to 4 to 5 Hz for all seven LFE families, though it could be biased high or  
 521 low by uncertainties in the LFE origin time alignment (see Appendix A1). In the synthetics, the  
 522  $P_d/P_l$  falloff frequency  $f_{fd}$  scales roughly with 1 over duration. It ranges from 1.4 to  $2.8V_r/D$ ,

523 or from  $0.15$  to  $1.4V_s/D$  (blue dashed line in Figure 5d). Coupling the synthetics with a 5-Hz  $f_{fd}$   
524 constrains the LFE diameters to be less than 1100 m.

525 More important constraints on the LFE properties come from the LFE durations estimated  
526 by *Thomas et al. (2016)*. *Thomas et al. (2016)* compared LFE stacks with nearby earthquakes’  
527 waveforms and obtained best-fitting durations of 0.19 and 0.22s for LFEs in families 37140 and  
528 37102, respectively. To get a sense of the uncertainty, we note that their best fits come from  
529 averaging over comparisons with 12 or 17 different local earthquakes, but they also present the  
530 durations obtained by the individual earthquake comparisons. Only one earthquake comparison  
531 gives a family 37140 duration smaller than 0.15 or larger than 0.22, and only one comparison  
532 gives a family 37102 duration smaller than 0.15 or larger than 0.3, so we use these values as  
533 uncertainty bounds. To compare the durations to our synthetics, we note that 70% of the moment  
534 in the stacked synthetic LFEs accumulates within  $0.29$  to  $0.31V_r/D$ . *Thomas et al. (2016)* modeled  
535 the LFE waveforms with a Hann-like source time function, which accumulates 70% of its moment  
536 within 40% of length, so the 70% durations for families 37140 and 37102 are 0.060 to 0.087 and  
537 0.060 to 0.12 s, respectively. We multiply these 70% durations by 1.4 to  $2.8V_r$  to estimate LFE  
538 diameters and plot the results with red vertical hatching in Figure 7a. The lower and upper thick  
539 red lines mark the diameters expected for the best-fitting durations for families 37140 and 37102,  
540 respectively.

541 The diameters implied by the observed durations match those implied by family 37102’s  $>$   
542 9 Hz  $f_{fc}$  for a wide range of rupture velocities. The two sets of constraints overlap at least partially  
543 for all plotted  $V_r/V_s$ , and the inter-station coherence constraint matches the median duration when  
544  $V_r < V_s$ . According to these results, LFEs in family 37102 could be slow ruptures, with 200-  
545 m diameters and  $V_r = 0.2V_s$ . Or they could be relatively “normal” earthquakes, with 800-m  
546 diameters and  $V_r = 0.8V_s$ . Note that changing the assumed shear wave velocity  $V_s$  would change  
547 the estimated diameters in Figure 7, but not the  $V_r/V_s$  intersection ranges, as all of the plotted  
548 diameter constraints scale with  $1/V_s$ .

549 Given the uncertainties in the data, the constraints on LFEs in family 37140 could also be  
550 matched with a range of rupture speeds. This family’s  $f_{fc} \gtrsim 16$  Hz constraint (below the solid blue  
551 line) starts to intersect the duration constraints when  $V_r < 0.7V_s$ . However, we should note that the  
552 plotted 16-Hz constraint is already the 95% lower bound on  $f_{fc}$ , obtained from bootstrapping. The  
553 best-fitting  $f_{fc}$  is 16.5 Hz. Further, the  $f_{fc} \gtrsim 16$  Hz constraint intersects the best-fitting duration  
554 only when  $V_r < 0.4V_s$ . Family 37140’s data are thus best matched when  $V_r < 0.4V_s$ .

### 555 **7.3 Matching $f_{fc}$ , $f_{fd}$ , and Duration With Simple Ruptures: Uncertainties**

556 To further assess whether earthquake-like rupture velocities  $V_r$  of 0.7 to  $0.9V_s$  are plausible, not  
557 just possible, for family 37140, we consider additional sources of uncertainty in the duration- and  
558 coherence-derived diameters. The diameters implied by the durations would decrease slightly if all  
559 ruptures began at the asperity edge. Groups of synthetic ruptures starting within  $0.1D$  of the LFE  
560 edge have durations of  $0.35$  to  $0.37D/V_r$ , rather than the  $0.29$  to  $0.31D/V_r$  values estimated for  
561 events starting within  $0.4D$  of the center. However, synthetic ruptures starting from the edge also  
562 give  $f_{fc}$  values about 20% smaller than those starting closer to the center (Figure S19). Changing  
563 both constraints leaves the range of allowable rupture velocities almost unchanged.

564 Other minor modifications to the rupture parameters appear to affect the  $f_{fc}$  constraints min-  
565 imally. For instance, we observe little change in  $f_{fc}$  if we add a smooth tapered component to

566 the heterogeneous slip distributions (Figure S21) or if we limit the range of diameters within each  
567 group to a factor of 1.1 standard deviation (Figure S23). However, we have not explored the  
568 entire range of rupture parameters. Perhaps we would obtain higher coherence if we made the  
569 slip distribution and temporal evolution smoother or slightly more repetitive, more similar to the  
570 repeater-like LFEs discussed in sections 6.1 and 7.4.

571 Another scenario that seems unlikely but possible is that the 16.5-Hz  $f_{fc}$  obtained for family  
572 37140 reflects random variability. This  $f_{fc}$  is significantly larger than the median  $f_{fc}$  for the seven  
573 families, which is just 11-Hz, and the synthetics in Figure 5b do show tens of percent variability in  
574  $f_{fc}$  among LFE groups, simply as a result of random variations in the slip distributions. However,  
575 those synthetics use only 100 LFEs. Using several thousand should reduce the uncertainty. Further,  
576 bootstrapping events within each synthetic group gives a reasonable estimate of the variability  
577 among the groups. Bootstrapping the data in family 37140 gives 95% probability that  $f_{fc} > 16$  Hz.

578 The other uncertainties in the data, along with potential variation in LFE location, would imply  
579 that the estimated 16.5-Hz  $f_{fc}$  is a lower bound on the true value, as discussed in section 5 and  
580 appendix A1. Given these uncertainties, we cannot exclude the possibility that these LFEs are  
581 simple ruptures with “typical” earthquake rupture speeds around  $0.7V_s$ . But we consider it more  
582 likely that the rupture velocities are lower than  $0.7V_s$ . The data are best matched by simple LFEs  
583 when rupture velocities are less than  $0.4V_s$ .

## 584 7.4 Matching the Data With Modified LFE Ruptures

585 It is also possible to match the data if we modify the LFE dynamics significantly: if LFEs are  
586 composite ruptures, ruptures with long rise times, or repetitive ruptures, as described in section 6.1.  
587 Figure 7b-d illustrate the constraints obtained for some plausible rupture parameters.

588 Figure 7b illustrates the constraints on diameters and rupture velocity if LFEs are composed  
589 of 5 sub-ruptures distributed over an interval with duration  $2.5D/V_r$ . Here the inter-station coher-  
590 ence constraints (blue) are essentially unchanged, but the direct coherence and duration constraints  
591 imply smaller diameters.

592 Figure 7c illustrates the constraints if LFEs have rise times equal to  $5D/V_r$ . In these LFEs,  
593 rupture would progress to the asperity edge, and then the whole patch would continue slipping  
594 together.

595 Finally, Figure 7d illustrates the constraints on  $D$  and  $V_r/V_s$  if LFEs are repetitive ruptures,  
596 which persistently nucleate in the same region, and which have two-thirds of their moment is  
597 associated with a slip distribution that is consistent from event to event. With these repetitive  
598 ruptures, the 16-Hz  $f_{fc}$  of family 37140 can be matched even if the rupture diameters are larger.

599 A wide range of parameters could also match the data if LFE durations are actually reflections  
600 of local attenuation, not the LFE source dynamics (*Gomberg et al., 2012; Bostock et al., 2017*). In  
601 this case, the diameters estimated from the durations (red lines) are upper bounds, and the data can  
602 be matched by any combination of rupture velocity and diameter that plots below those bounds  
603 and within the  $f_{fc}$  (blue) and  $f_{fd}$  (red) constraints.

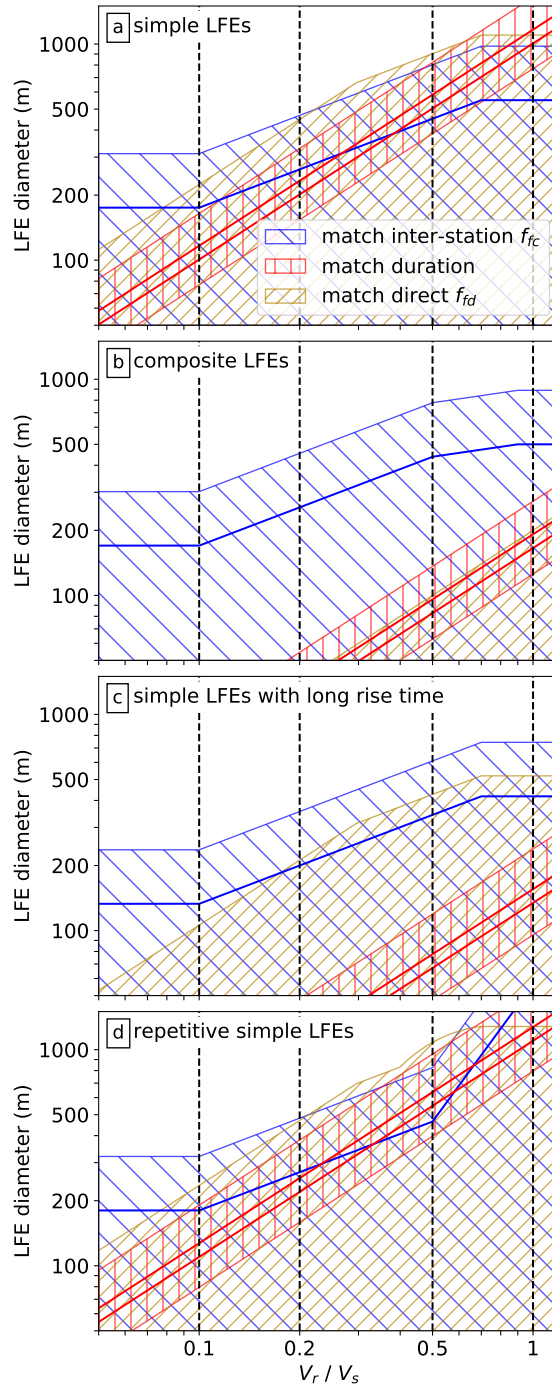


Figure 7: Hatched regions mark diameters (y-axis) and rupture velocities (x-axis) that match each of the 3 observations:  $f_{fc}$  (blue diagonal lines),  $f_{fd}$  (yellow diagonal lines), and the durations of *Thomas et al. (2016)* (red vertical lines). The four panels are for four approaches to constructing the LFEs, as indicated by the text in the upper left.

## 8 Discussion

### 8.1 Implications for Tremor Asperities

Regardless of the individual LFE rupture dynamics, our observations of high-frequency coherence suggest that LFEs are clustered in patches less than 1 km across. As noted in the introduction, such clustering has also been inferred by careful analysis of LFE families in Cascadia (*Sweet et al.*, 2014; *Chestler and Creager*, 2017a) and may be suggested by highly periodic LFE ruptures in Parkfield (*Shelly*, 2010b). The clustering may suggest a role for material heterogeneity in controlling the occurrence of tremor. It is consistent with proposals that tremor’s LFEs rupture a collection of unstable asperities embedded in a larger, more stable region (*Ando et al.*, 2010; *Nakata et al.*, 2011; *Ando et al.*, 2012; *Ariyoshi et al.*, 2012; *Veedu and Barbot*, 2016; *Luo and Ampuero*, 2017). Larger asperities may also exist, as patches of tremor are observed on scales of a few to tens of km. The larger tremor patches could represent groups of tremor asperities or regions more prone to distributed rapid slip (*Shelly*, 2010b; *Ghosh et al.*, 2012; *Armbruster et al.*, 2014; *Yabe and Ide*, 2014; *Savard and Bostock*, 2015; *Annoura et al.*, 2016; *Kano et al.*, 2018). Alternatively, the large and small tremor patches could represent persistent slip patterns that have arisen on a simple, homogeneous fault. Such patterns are sometimes seen in models that lack heterogeneity in material properties (*Horowitz and Ruina*, 1989; *Langer et al.*, 1996; *Shaw and Rice*, 2000), though it remains to be assessed whether these models can produce clusters of tremor that persist over many slow slip cycles, as we observe in Parkfield.

The family-based clustering implied by our coherence estimates and by others’ LFE relocations (*Sweet et al.*, 2014; *Chestler and Creager*, 2017a) suggests that cross-correlation based LFE families are more an observational convenience (*Shelly et al.*, 2007; *Brown et al.*, 2008; *Bostock et al.*, 2012; *Frank et al.*, 2013; *Kato*, 2017; *Shelly*, 2017). The analyzed families show sub-km LFE clustering even though some are separated from neighboring families by a few to 5 km. The LFEs’ tendency to occur on these asperities lends further confidence to studies that have interpreted LFE repeat rates as indicators of the slip rate in a creeping area surrounding the more unstable LFE patches (*Rubin and Armbruster*, 2013; *Royer et al.*, 2015; *Lengliné et al.*, 2017; *Thomas et al.*, 2018).

### 8.2 Implications for Tremor Physics

Given our observations and synthetics of LFE coherence as a function of rupture diameter, there are still several ways to explain the long, 0.2-s durations of Parkfield LFEs. First, it is possible that families 37102 and 37140’s LFEs are normal earthquakes with near-shear-wave rupture speeds. A  $0.7V_s$  rupture speed is at the edge of the constraints for family 37140, but such rapid ruptures could match the data better if the LFEs are somewhat repetitive, with nucleation locations and slip distributions that persist from event to event. And a wide range of high rupture speeds could match the data if the 0.2-s durations we use are overestimates of the true durations, despite *Thomas et al.* (2016)’s careful empirical Green’s function analysis. The durations could be overestimated if a highly attenuating region is localized around the LFE patches, so that attenuation removes the high-frequency components of the LFE seismograms but has little effect on the seismograms of the reference earthquakes, which are located a few km away.

If LFEs do have durations of 0.2 s and rupture speeds up to  $0.7V_s$ , they could have diameters

645 up to 800 m. Uniform stress drop  $M_w$  1 to 2 earthquakes with 800-m diameters would have stress  
646 drops of 0.3 to 9 kPa and average slips of 0.002 to 0.06 mm (Eshelby, 1957; Shearer, 2009). These  
647 moment and slip estimates are imprecise, and difficult to estimate because LFE locations are offset  
648 from local earthquakes, but we note that if the larger slip estimates are representative, almost all of  
649 the slip on the LFE patch could be seismic. Even 800-m-wide LFEs could accommodate most of  
650 the long-term slip on the LFE patch, which Thomas *et al.* (2016) estimated to be around 0.05 mm  
651 per event.

652 But while LFEs from both families can be matched by rupture velocities up to  $0.7V_s$ , the data  
653 from family 37140 are better matched by LFEs with slower rupture speeds ( $< 0.4V_s$ ), long rise  
654 times, or a composite of subevents. Any of these scenarios would have interesting implications  
655 for the physics of LFE ruptures. For instance, rupture speeds around  $0.4V_s$ , which can match the  
656 data for both families, would suggest that the LFEs' radiation efficiency is around 0.5: that about  
657 half of the energy in LFEs is released via seismic wave generation, with the rest expended as  
658 fracture energy (e.g., Kostrov, 1966; Eshelby, 1969; Fossum and Freund, 1975; Venkataraman and  
659 Kanamori, 2004; Kanamori and Rivera, 2006). Such low but significant radiation efficiency *could*  
660 mean that LFEs are exceptionally weak but otherwise normal earthquakes. LFEs may be driven  
661 by unstable frictional sliding, and their slip rates may be limited by the dissipation of energy via  
662 seismic waves. Although  $0.4V_s$  is lower than typical earthquake rupture speeds (McGuire, 2004;  
663 Seekins and Boatwright, 2010; Folesky *et al.*, 2016; Ye *et al.*, 2016; Melgar and Hayes, 2017;  
664 Chouinet *et al.*, 2018), such speeds are sometimes observed in earthquakes, especially in shallow  
665 tsunami earthquakes (e.g., Ide *et al.*, 1993; Ihmlé *et al.*, 1998; Venkataraman and Kanamori, 2004;  
666 Bilek and Engdahl, 2007; Polet and Kanamori, 2009; Cesca *et al.*, 2011).

667 It is thus possible that LFEs are simply earthquakes driven by a frictional weakening process  
668 that is for some reason smaller in magnitude than the processes driving normal earthquakes. LFEs  
669 might nucleate “earlier” than most earthquakes, at times when there is only a modest stress drop  
670 available to drive rupture. Or LFEs could nucleate on small unstable patches but then move quickly  
671 into regions that resist high slip speeds, perhaps because they are velocity-strengthening or allow  
672 for large off-fault deformation. Such acceleration-resisting regions have been suggested to limit the  
673 rupture velocities of tsunami earthquakes (e.g., Bilek and Lay, 2002; Faulkner *et al.*, 2011a; Ma,  
674 2012). Off-fault deformation seems an appealing process to invoke for tremor because complex  
675 brittle and ductile deformation is observed at relevant depths (Fusseis *et al.*, 2006; Handy *et al.*,  
676 2007; Collettini *et al.*, 2011; Fagereng *et al.*, 2014; Hayman and Lavier, 2014; Angiboust *et al.*,  
677 2015; Behr *et al.*, 2018; Webber *et al.*, 2018). It is even possible that each LFE is a collection  
678 of small brittle failures, rupturing small faults or veins (Fagereng *et al.*, 2014; Ujiie *et al.*, 2018).  
679 However, it remains unclear how or if that distributed ductile deformation would limit the rupture  
680 speeds of LFEs. Off-fault ductile deformation can also accumulate in large earthquakes, with  
681 near-shear-wave rupture speeds (DeDontney *et al.*, 2011; Dunham *et al.*, 2011; Roten *et al.*, 2017).

682 Another possibility is that LFEs do rupture at near-shear-wave speeds, but that the shear wave  
683 speed is significantly reduced in the LFE area because of lithological variations, fault zone damage,  
684 or high pore pressures (Audet *et al.*, 2009; Song *et al.*, 2009; Kato *et al.*, 2010; Fagereng and Di-  
685 ener, 2011; Stefano *et al.*, 2011; Huang *et al.*, 2014). Fault damage zones are frequently observed  
686 at a range of depths (Shipton and Cowie, 2001; Rowe *et al.*, 2009; Faulkner *et al.*, 2011b; Rempé  
687 *et al.*, 2013; Leclère *et al.*, 2015), and they sometimes show 30 to 50% reductions in wavespeed, at  
688 least in shallow regions (Ben-Zion *et al.*, 2003; Cochran *et al.*, 2009; Lewis and Ben-Zion, 2010;  
689 Yang *et al.*, 2014; Li *et al.*, 2016). It is difficult to fully assess a low-wavespeed region's implica-

690 tions for our observations. The inter-station coherence we observe depends on the seismic waves’  
691 source-station travel times, and those times depend on which source-station paths are traveled. But  
692 in the simplest case, where LFE signals begin by traveling horizontally away from the fault, so  
693 that they move outside the fault zone before continuing to the surface, the travel time variation  
694 we probe with inter-station coherence would depend primarily on the higher wavespeed outside  
695 the fault zone. The higher wavespeeds could allow for the high-frequency inter-station coherence  
696 we observe even though the lower wave speed inside the fault zone limits the rupture velocity and  
697 produces long-duration events.

698 On the other hand, it is possible that LFE rupture velocities are not limited by seismic wave  
699 radiation at all, but by a different fault zone rheology. We note that the results from family 37140  
700 are best fit by simple LFE ruptures with  $V_r < 0.4V_s$ , and because of noise in the data, all of our  
701 coherence-constrained diameters and rupture speeds are upper bounds on the true values. So LFE  
702 rupture speeds could be much smaller:  $0.2V_s$ , for example. Such slowly rupturing LFEs would  
703 release more than 80% of their energy via fracture energy, making it unlikely that the energy dis-  
704 sipated via seismic wave radiation could limit the slip speeds. The low rupture velocities inferred  
705 for family 37140 could be telling us that LFE rupture dynamics are controlled by a different defor-  
706 mation mechanism than normal earthquakes—perhaps by the same speed-limiting rheology that  
707 controls slow slip events (e.g., *Ide et al.*, 2007; *Shibazaki and Iio*, 2003; *Shibazaki and Shimamoto*,  
708 2007; *Ide et al.*, 2008; *Aguiar et al.*, 2009; *Liu et al.*, 2010; *Segall et al.*, 2010; *Gao et al.*, 2012;  
709 *Hawthorne and Rubin*, 2013; *Ide and Yabe*, 2014; *Hawthorne and Bartlow*, 2018).

## 710 9 Conclusions

711 We have analyzed inter-station and inter-event coherence between LFEs in seven families near  
712 Parkfield, CA. Our synthetic analysis shows that we can use inter-station ASTF variations to es-  
713 timate LFE or earthquake rupture areas. Our observations of LFE coherence imply that LFEs in  
714 each family are strongly clustered, with standard deviation in their locations smaller than 250 m.  
715 Comparing the observed coherence with that of synthetic LFE ruptures implies that the LFE diam-  
716 eters are smaller than 500 to 1100 m, depending on the family. Coupling the coherence constraints  
717 with the LFE durations estimated by *Thomas et al.* (2016) suggests that we could match the data  
718 for LFEs in family 37102 with a wide range of rupture models, including earthquake-like rup-  
719 tures with rupture velocities  $V_r$  of 0.7 to 0.9 times the shear wave speed  $V_s$ . For family 37140,  
720  $V_r = 0.7V_s$  can match the data, but only on the edge of the constraints, and the data are better  
721 matched with  $V_r < 0.4V_s$ . Such low rupture speeds may indicate that LFEs are governed by a slow  
722 slip rheology, not by standard unstable frictional sliding, but we note that data from both families  
723 of LFEs could also be matched if LFEs are repetitive fast ruptures, composite ruptures, or ruptures  
724 with long rise times. Our synthetics illustrate how the coherence and durations might differ among  
725 these rupture types, and thus how we might probe the physics of LFEs with future observations.

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732 Figure 2 were obtained from the USGS and California Geological Survey fault and fold database,  
733 accessed from <http://earthquake.usgs.gov/hazards/qfaults> in 2016.

## 734 **A1 Decoherence from Noise**

735 Our coherence frequencies should probably be interpreted as lower bounds, as several sources of  
736 noise could reduce the observed  $P_d/P_l$  and  $P_c/P_l$  from their true values. First, decreased  $P_d/P_l$   
737 and  $P_c/P_l$  could arise if a significant portion of the “noise” comes from LFEs that are nearby but  
738 not in the family of interest. LFEs are clustered in space and time (e.g., *Shelly*, 2010a; *Bostock*  
739 *et al.*, 2015) so the noise from other LFEs may be higher during the LFE window than during the  
740 noise window before it. We estimate the noise power  $P_n$  in a window that starts just 8 s before the  
741 LFE S arrival to minimize the potential difference, but we cannot account for sub-8 s clustering.  
742 Note that in principle our noise window could include some of the P arrival. However, we find  
743 the P arrival is too late and too small to significantly affect the  $P_n$  estimates. Truncating the noise  
744 waveforms before the P arrivals and reprocessing changes our results negligibly.

745 Decreased  $P_d/P_l$  and  $P_c/P_l$  could also result from noise in the template LFEs. The template  
746 signals start to become poorly resolved at frequencies higher than 15 Hz, so it is difficult to calcu-  
747 late robust powers at those frequencies.

748 Finally, decreased or increased  $P_d/P_l$  could result from uncertainty in the LFE origin time. To  
749 accurately calculate direct coherence at high frequencies, we need well aligned waveforms, so we  
750 re-compute LFE origin times using 0.01-s precision. The realignment affects  $P_c/P_l$  negligibly but  
751 increases the frequencies with  $P_d/P_l > 0.5$  by several Hz relative to results without recomputed  
752 origin time. One might worry that the increase in coherence comes from aligning the template  
753 with coherent noise rather than with LFE signal. However, we require at least 5 stations for the  
754 power estimates for each LFE, and we allow only one origin time shift per LFE. Assuming noise  
755 is random among stations, realigning with noise should increase  $P_d/P_l$  by less than 0.2.

756 The LFE detection approach of *Shelly* (2017) could also result in slightly increased coherence  
757 if noise contributes a part of the identified coherent signals. Finally, slightly increased coherence  
758 could result from our exclusion of signals with especially high noise. Note that the detected-  
759 facilitated increases in coherence are most likely to occur at low frequencies, around a few Hz, as  
760 these frequencies contribute most of the seismogram power involved in LFE selection and align-  
761 ment.

762 There are no other obvious sources of artificially high coherence. Applying our processing to  
763 noise intervals rather than LFEs gives  $P_c/P_l$  and  $P_d/P_l$  of 0.01 or less.

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