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Polar ice-cores unravel the formation of a UV window during magnetic field collapse ~ 42 ka BP

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Key Points

- Observational evidence of strong ozone layer depletion in the Southern Hemisphere during the Laschamp Geomagnetic Excursion
- Timing of UV window formation is constrained to within 100 years of magnetic field collapse using sulfur isotope anomalies detected in ice core sulfate
- Potential for hemispherical discrepancy exists in the realization of the immediate effect of increased UV influx at Earth's surface during geomagnetic excursions

Abstract

During geomagnetic excursions (GEs), compromised magnetic field and increased cosmic-ray bombardment can deplete the ozone layer forming ‘UV window(s)’ in the Earth’s atmosphere. Here, using triple sulfur-isotope systematics in polar ice-core sulfate record spanning 600 years of the Laschamp GE, we provide direct evidence of the formation of a UV window. Several events of UV-induced anomalous ice-core sulfate were detected under background conditions in Antarctica concomitant with the evolution of paleomagnetic intensity, with a distinct peak observed between ≈ 41.7 ka BP and ≈ 41.6 ka BP corresponding to the weakest magnetic field phase. The timing and amplitude of these events unequivocally confirm the existence of a UV window in the southern hemisphere during the LGE. However, such geochemical imprints were not detected in discrete subsampled northern polar ice-core record. We postulate that, among other possibilities, non-uniformity in excursion behaviour at the Earth’s surface could be modulating this geochemical imprint.

Plain Language Summary

The ozone layer in the upper atmosphere acts as a shield against harmful ultraviolet (UV) radiation, thereby protecting life on the planet. However, drastic changes occurring in the Earth’s magnetic field over geological timescales, termed geomagnetic excursions (GEs), can damage the ozone layer making it feasible for harmful UV radiation to penetrate deeper into the troposphere. This is here termed as the creation of ‘UV window(s)’ in the Earth’s atmosphere. Direct observational evidence of the timing of formation of such UV windows during the GE events has been lacking and is needed to better understand the climate/environmental and societal impact of GEs in the past (and potentially in the future). Here, employing a novel geochemical tracer of UV radiation in polar ice core records, we detect the formation of such a UV window in the Earth’s atmosphere ≈ 41.5 thousand years in the past during a period of severe magnetic field collapse which profoundly impacted life on the planet. We find the existence of this UV window had strongly materialized in the Southern Hemisphere. Overall, we

46 suggest that GEs can indeed create such UV windows, but their immediate effect might not be realized
47 uniformly across the globe.

49 **1. Introduction**

50 The Earth's magnetic field, virtue of its dipole moment, shields life on the planet by deflecting
51 much of the cosmic radiation reaching the Earth's surface (Merrillet et al 1983; Usoskin et al., 2008).
52 However, variations in this dipole moment and as such the magnetic field have been noted to occur on
53 geological timescales. Geomagnetic excursions are characterised by major changes in the Earth's
54 magnetic dipole moment, lasting a few hundred to a few thousand years, and have been observed in the
55 past (Herndon et al., 2020; Jackson et al., 2000; Simon et al., 2016; Valet et al., 2024). It is plausible
56 that a weakened magnetic field could lead to the depletion of the stratospheric ozone layer through
57 enhanced chemical effects in the upper atmosphere (Norval et al., 2007; Solomon et al., 1982, 1999;
58 Winkler et al., 2008). Consequently, such an event could lead to the formation of 'UV window(s)'
59 thereby allowing the harmful UV rays (Cooper et al., 2021; Dasari et al., 2022; Mukhopadhyay et al.,
60 2025; Solomon et al., 1999), which otherwise would have been absorbed by the ozone layer, to
61 penetrate deeper into the troposphere and, thus, towards the Earth's surface. Under such scenarios, the
62 formation and persistence of UV windows would therefore be detrimental to both the environment and
63 life on the plane (Cooper et al., 2021; Mukhopadhyay et al., 2025).

64 Studying past events of magnetic field collapse might hold the key to an improved
65 understanding of the formation of UV window(s) (Cooper et al., 2021; Channell et al., 2019; Dasari et
66 al., 2022; Valet et al., 2010). The period between ≈ 42 ka BP and ≈ 41 ka BP (thousand years before
67 present), wherein the intensity of the Earth's magnetic field decreased by more than 80 % and then
68 partially recovered (Valet et al., 2024), represents one of the most significant changes in the magnetic
69 dipole moment (Cooper et al., 2021; Channell et al., 2019; Simon et al., 2007; Valet et al., 2024) (Figure
70 1a). This event, known as the Laschamp Geomagnetic Excursion (LGE), was a global and synchronous

71 event observed in many lavas of similar ages in New Zealand and California (Nowaczyk et al., 2025;
72 Valet et al., 2024) and highlighted in marine stratigraphic records (Valet et al., 2010) and ice-core
73 records from Greenland and Antarctica (Raisbeck et al., 1997, 2017; Yiou et al., 1999). Moreover, this
74 event is particularly intriguing and important because it has been proposed to explain global climatic
75 shifts during this period that yield to major environmental changes and life extinction (Cooper et al.,
76 2021; Valet et al., 2010). For instance, the LGE has been attributed as a contributing factor for the
77 subsequent disappearance of the Neanderthals (Cooper et al., 2021; Valet et al., 2010) contemporaneous
78 with this period, despite the notion being extensively debated (Picin et al., 2021; Hawks 2021). As such,
79 the LGE event represents a compelling case study for the formation and existence of UV window(s) in
80 Earth's atmosphere.

81 Mostly climate proxies and modelling works have been used to study the instances of increase
82 in UV influx in connection to the thinning/depletion of the ozone layer in the Earth's past (Hodson et
83 al., 2005; Marshall et al., 2020). Recently, fingerprinting sulfur-isotope anomalies recorded in Antarctic
84 ice-core sulfate records has shown promise in being able to observationally trace ozone layer depletion
85 events in the past (Dasari et al., 2022). In general, such isotopic anomalies are only ever recorded in
86 polar ice-core records during stratospheric volcanic eruptions wherein the plumes reach altitudes above
87 the ozone layer (Baroni et al., 2007; Burke et al., 2019; Gautier et al., 2019; McConnell et al., 2017;
88 Savarino et al., 2003). Subsequently, through a series of photochemical reactions involving UV
89 wavelengths between 190 and 250 nm and oxidation kinetics of S-bearing species such as sulfur dioxide
90 (SO_2) or carbonyl disulfide (CS_2), the synthesized sulfur-mass independent fractionation (S-MIF)-
91 bearing sulfate aerosols get deposited in the polar ice (represented by ' $\Delta^{33}\text{S}$ ', also referred to as 'sulfur-
92 isotope anomaly', in ice-core sulfate isotope measurements; see Methods) (Farquhar et al., 2000;
93 Gautier et al., 2018; Hattori et al., 2013; Ono et al., 2017; Whitehill et al., 2015). However, it is logical
94 to envisage that in the event of a sustained depletion of the ozone layer occurring over hundreds of
95 years, such as during geomagnetic excursion events, the UV rays would penetrate towards Earth's

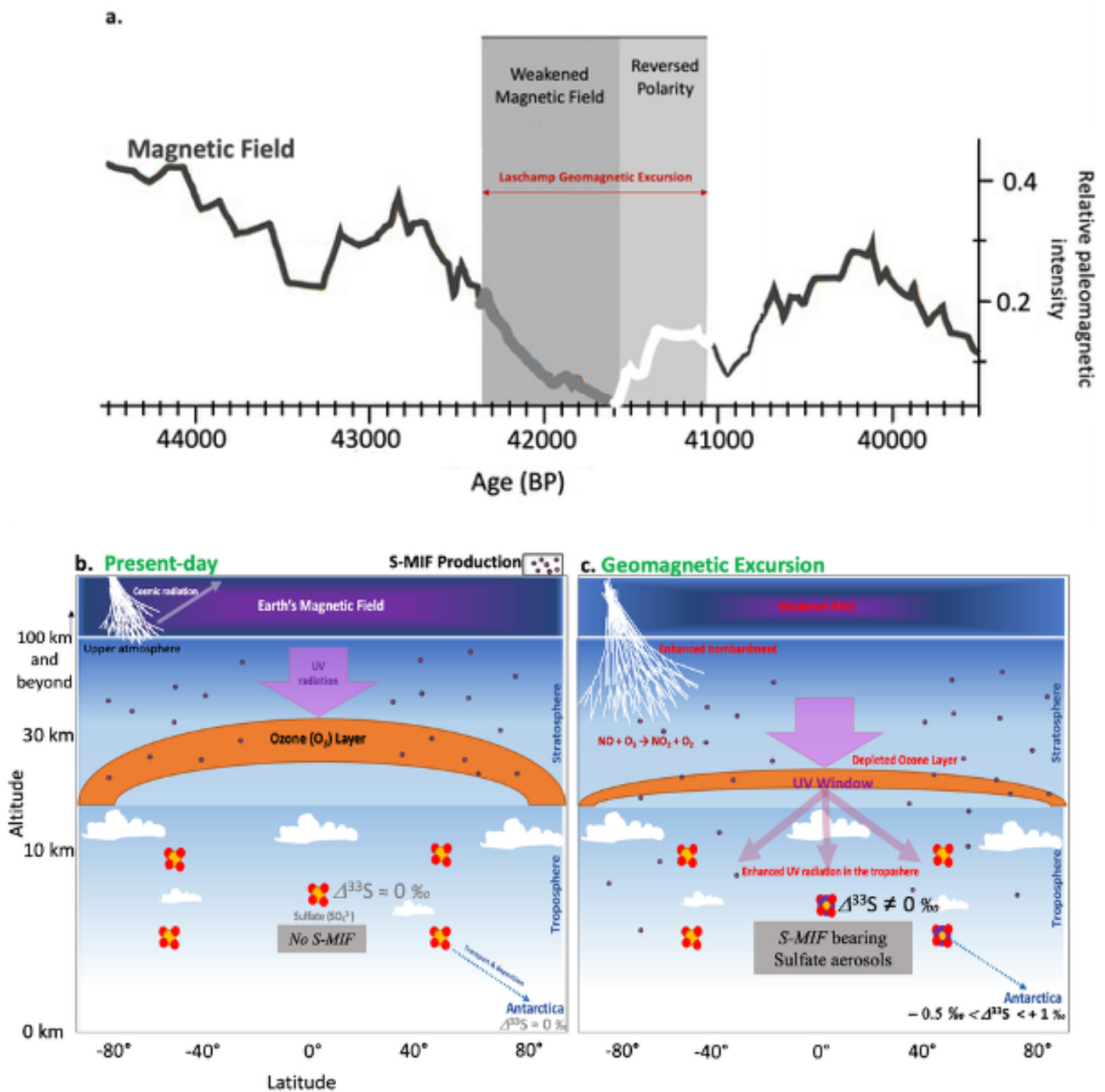
96 surface, thereby leading to the production of S-MIF-bearing sulfate aerosols in the troposphere during
97 such periods (Figure 1). Consequently, the geochemical imprint of UV window(s) from ozone layer
98 depletion could be found during such corresponding periods in polar ice-core records (Dasari et al.,
99 2022).

100 However, ice-core sulfur-isotope records had only ever been used to detect volcanic events in
101 the past (e.g., Baroni et al., 2007; Burke et al., 2019; Gautier et al., 2019; McConnell et al., 2017;
102 Savarino et al., 2003) limiting the possibility of such a pursuit to trace UV windows using this
103 geochemical tool. In this regard, our pilot study pioneering the application of sulfur isotope anomalies
104 in polar ice-core records for a renewed purpose of tracing UV window(s) (Dasari et al., 2022) had
105 recently investigated the period of ≈ 41.5 ka BP to ≈ 41.2 ka BP of the LGE event i.e., the reversed
106 polarity phase (Figure 1a). During this period the partial recovery of the magnetic field had commenced,
107 and the study found evidence for sulfur isotope anomalies in ice-core record not associated with any
108 stratospheric volcanic influence, hinting the possibility of the existence of a UV in the Earth's
109 atmosphere during the LGE event. However, direct evidence of the timing of the formation, the overall
110 period of existence and above all, the globality of the UV window(s) during the LGE event is still
111 lacking. These are crucial aspects which need investigating as the UV-influx from the LGE event may
112 have profoundly impacted life on the planet (Cooper et al., 2021; Channel et al., 2019; Valet et al.,
113 2010).

114 In this study, to test the timing of the formation, period of existence and globality of the UV
115 window(s) during the LGE event, we provide a new record of sulfur-isotope anomalies ($\Delta^{33}\text{S}$) by
116 sampling ice-cores from both Antarctica (Figure 2; see also Table S1 in Supporting Information) and
117 Greenland (Figure 3; Tables S2-S3 in Supporting Information) with a focus on the weakened magnetic
118 field phase (≈ 42 ka BP and ≈ 41.5 ka BP; Figure 1a) of the LGE event. We couple this with the re-
119 analysed S-isotope record (see Methods) of the same Antarctic ice-cores from the pilot study (Dasari et
120 al., 2022) for the reversed polarity phase (≈ 41.5 ka BP to ≈ 41.3 ka BP; Figure 1a) to overarchingly

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explore the link between magnetic field collapse and recovery over the investigated 700-year period of the LGE event, and the concomitant changes in the stratospheric ozone layer in the Earth's past.



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Figure 1. S-MIF based investigation of a UV window during geomagnetic excursion events. (A) The evolution of the paleomagnetic intensity during the Laschmap geomagnetic excursion event is shown.

126 **(B)** Schematic of tropospheric S-isotope imprint without UV window. Most of the isotopic
127 fractionation mechanisms for sulfur fractionate isotopes are in proportion to their mass, according to
128 the so-called "mass dependent" fractionation laws (MDF). For MDF, one can write $\delta^{33}\text{S} \approx 0.515 \times \delta^{34}\text{S}$
129 (for its linear approximation), If $\Delta^{33}\text{S} \neq 0$, then it is referred to as "mass-independent" fractionation
130 (MIF) (Farquhar et al., 2000; Gautier et al., 2018; Hattori et al., 2013; Ono et al., 2017; Whitehill et al.,
131 2015). In the present atmosphere, the presence of the ozone layer along with an intact geomagnetic
132 field prevents the creation of such large-scale fractionation i.e., $\Delta^{33}\text{S} = 0 \text{ ‰}$ in modern day atmospheric
133 sulfate in polar ice. **(C)** Schematic of tropospheric S-isotope imprint without UV window. During the
134 Geomagnetic Excursion events creation of a UV 'window' could lead to deeper penetration of UV
135 radiation towards the Earth's surface. The UV-induced photochemistry can then generate measurable
136 S-MIF splits in tropospheric sulfate produced from S-containing compounds such as SO_2 , CS_2 , which
137 can be detected in polar ice-core sulfate records for the corresponding periods (Dasari et al., 2022).

138 139 **2. Materials and Methods**

140 **2.1 Ice-core sulfate sampling, isolation and isotope measurements**

141 High-resolution sampling of ice-cores was conducted by cutting ~ 3 cm slices along the wings of the
142 cores. Outside edges of the ice-cores were scraped clean, and ice-core slices were then decontaminated
143 by rinsing the ice with ultrapure water (Milli-Q, Millipore Corp.) and, finally, melted and filtered at
144 $0.45 \mu\text{m}$. Major water-soluble ions were measured by ion chromatography (Metrohm IC Professional
145 850), each slice requiring 2 ml of the sample. The quality assurance of the analytical data (within 5 %)
146 was ascertained by preparing standard stock solutions from analytical grade reagents and salts, and
147 reported concentrations were suitably blank corrected. Depending on the sulfate concentrations,
148 samples were pooled (providing ~ 10 -year time resolution) and sulfate was then isolated using ion
149 chromatography-based fraction separation at IGE as described elsewhere (Gautier et al., 2019).

150 The sampled Antarctic ice-cores (core: Vostok 3G; 78.46°S 106.83°E) were from 611-615 m
151 in depth (≈ 41.6 to ≈ 42 ka BP in age; AICC 2023 chronology (Bouchet et al., 2023); Table S1 in
152 Supporting Information) whereas for the discrete sub-sampled Greenland ice-cores (core: NGRIP;
153 75.10°N, 42.32°W) this was 2126–2129 m in depth (≈ 41.5 to ≈ 41.7 ka BP in age; AICC 2023
154 chronology; Table S2 in Supporting Information). The reason for the partial overlap was the limited
155 availability of sample as well as the intention to cover the overlapping period of the lowest magnetic
156 field intensity (Figure 1) in the dual-polar ice-core samples (see also age transfer function in Table S3
157 in Supporting Information). This new Antarctic ice-core (site: Vostok 3G) sulfate record was then
158 coupled with the data from the pilot study (for the ‘non-volcanic’ sulfate period alone; see Discussion)
159 wherein sampled cores were from 607-611 m in depth (≈ 41.3 to ≈ 41.6 ka BP in age; AICC 2023
160 chronology) (Dasari et al., 2022). The pilot study also analyzed the period 41.2 to 41.3 ka BP which
161 was dominated by ‘volcanic’ sulfate and was therefore not used in the merged record here.

162 The isolated sulfate extracts were then analyzed for S-isotopes at CRPG with the Neptune Plus
163 MC-ICP-MS (Thermo Fisher Scientific Neptune Plus) following well-established protocols for small
164 sample analysis (Paris et al., 2013, 2014, 2024). The MC-ICP-MS was coupled with a desolvation
165 membrane (Aridus II, Cetac) which reduces the interference of oxides and hydrides on the masses ^{32}S ,
166 ^{33}S , and ^{34}S and hence increases the sensitivity to the signal. The samples were analysed by standard-
167 sample bracketing using an in-house Na_2SO_4 standard to correct for instrumental isotopic fractionation
168 (Paris et al., 2024). Before introducing the sample to the instrument, the sulfate extracts were
169 evaporated to dryness, resuspended in 5 % HNO_3 , and NaOH was added to the sample to match the
170 concentration and matrix of the in-house Na_2SO_4 bracketing standard with a concentration of 40 $\mu\text{mol/L}$
171 (Dasari et al., 2022b). As part of quality control, the bracketing standard was measured as a sample
172 with decreasing sulfate concentration, showing no bias induced by sulfate concentration as long as Na
173 concentration matches that of the bracketing standard, in agreement with previous studies (Dasari et
174 al., 2022a, 2024). Each sample was run a minimum of four independent times, usually five, with each

175 run making fifty measurements and a sulfate concentration between 20 and 40 $\mu\text{mol/L}$. The reported
176 values (Tables S1 in Supporting Information) are given as an average of at least five independent
177 measurements on the Neptune for a given sample. This analytical error (usually $< 0.05 \text{ ‰}$, 1σ) is an
178 improvement from previous MC-ICP-MS-based studies (Dasari et al., 2022a, 2022b, 2024; Paris et al.,
179 2013, 2014, 2024) as samples are run (i) at a higher concentration (up to 40 $\mu\text{mol/l}$ instead of less than
180 20 $\mu\text{mol/L}$) and (ii) 5 times instead of 2, respectively. The error accounts for instrumental
181 reproducibility.

182 The triple sulfur isotopes (^{32}S , ^{33}S , ^{34}S) are reported relative to Vienna-Canyon Diablo Troilite
183 (V-CDT). $\Delta^{33}\text{S}$ is here defined as $\delta^{33}\text{S} - 0.515 \times \delta^{34}\text{S}$, where $\delta^{3x}\text{S} = \ln[(^{3x}\text{S}/^{32}\text{S})_{\text{sample}} / (^{3x}\text{S}/^{32}\text{S})_{\text{standard}}]$
184 (Hulston & Thode, 1965). The bracketing standard was calibrated against the international standard
185 IAEA-S1 assuming $\delta^{34}\text{S} = -0.3 \text{ ‰}$ and $\delta^{33}\text{S} = -0.061 \text{ ‰}$ (and thus $\Delta^{33}\text{S} = 0.094 \text{ ‰}$). Further quality
186 control was conducted using a secondary sulfur isotope reference material 'SMIF-1' which yielded a
187 $\delta^{34}\text{S}$ value of $10.06 \pm 0.14 \text{ ‰}$ and $\Delta^{33}\text{S}$ of $9.46 \pm 0.14 \text{ ‰}$ ($n = 5$) in agreement with the reported $\Delta^{33}\text{S} =$
188 $9.54 \pm 0.09 \text{ ‰}$ from five different laboratories worldwide (Geng et al., 2019). A seawater sample was
189 also run further to ensure the precision and accuracy of the measurements. It yielded $\delta^{34}\text{S}$ and $\Delta^{33}\text{S}$
190 values of $21.14 \pm 0.09 \text{ ‰}$ and $0.07 \pm 0.12 \text{ ‰}$ (2 s.d., $n=5$), in agreement with previously published
191 values (Paris et al., 2013, 2014, 2024). The seawater sample was purified at CRPG following a column
192 chemistry method (Paris et al., 2013, 2014, 2024). Procedural blanks were also run with each set of
193 samples. Overall, high-precision measurements of both $\delta^{34}\text{S}$ and $\Delta^{33}\text{S}$ were ensured during this
194 investigation (Tables S1-S2 in Supporting Information).

195

196 **2.2 Calculation of the non-sea salt (NSS)-sulfate isotopic signatures**

197 The polar ice-core sulfate is a mix of sea water sulfate and atmospheric sulfate (from a -variety of
198 sources). To isolate the atmospheric component, a sea-salt sulfate (SS) correction needs to be made

199 both for concentration and isotope ratios (e.g., Gautier et al., 2019). The concentration of SS present in
 200 aerosols is usually estimated according to the relation:

$$202 \quad [\text{SO}_4^{2-}]_{\text{SS}} = k [\text{Na}^+] \quad (1)$$

203
 204 A k value of 0.25 is often used for this calculation in seawater, but in Antarctica, due to temperature
 205 differences, the sea salt emitted from the sea ice surface is depleted in SO_4^{2-} relative to Na^+ . As such,
 206 lower values have been reported at various Antarctic sites (Alexander et al., 2003; Ishino et al., 2019),
 207 e.g., at Dome C (0.16) and DDU (0.13), respectively. Therefore, a mean k value of 0.14 has been used
 208 here.

209
 210 Quantity of non-sea-salt (NSS) sulfate is then determined using the relation as in previous studies
 211 (Gautier et al., 2019; Uemura et al., 2022):

$$213 \quad [\text{SO}_4^{2-}]_{\text{NSS}} = [\text{SO}_4^{2-}]_{\text{M}} - [\text{SO}_4^{2-}]_{\text{SS}} \quad (2)$$

214
 215 Where $[\text{SO}_4^{2-}]_{\text{M}}$ is the concentration of sulfate measured in the sampled ice-core

216
 217 The isotopic composition of the NSS is then approximated using the following mixing equation:

$$219 \quad \delta^{34}\text{S}_{\text{NSS}} \approx (\delta^{34}\text{S}_{\text{M}} - [1-X] \delta^{34}\text{S}_{\text{SS}}) / X \quad (3)$$

220
 221 where X is the proportion of NSS defined as $X = [\text{SO}_4^{2-}]_{\text{NSS}} / [\text{SO}_4^{2-}]_{\text{M}}$ and $\delta^{34}\text{S}_{\text{SS}}$ is the average isotopic
 222 composition measured in seawater. This equation (3) can also be used to express the $\delta^{33}\text{S}_{\text{NSS}}$.

224 Finally, the $\Delta^{33}\text{S}_{\text{NSS}}$ is expressed (in per mil) as follow :

$$225 \quad \Delta^{33}\text{S}_{\text{NSS}} = \delta^{33}\text{S}_{\text{NSS}} - \{[(\delta^{34}\text{S}_{\text{NSS}}) + 1]^{0.515} - 1\} \quad (4)$$

227
228 Note the data in the pilot study (concentrations and isotope values) were not corrected for contribution
229 from sea salt. As such, in the merged record here (Table S1 in Supporting Information) the data from
230 the pilot study was reanalyzed accounting for the sea salt correction as was done for the new record in
231 the present study.

232 233 234 **2.3 Source apportionment of polar ice-core sulfate**

235 We consider three possible sources of sulfate in the polar ice core samples: marine biogenic (*bio*),
236 volcanic (*vol*) and terrestrial (*terr*). Combining the source contributions is expected to yield the
237 observed $\delta^{34}\text{S}_{\text{NSS}}$ as per the following equations:

$$238 \quad f_{\text{bio}} \delta^{34}\text{S}_{\text{bio}} + f_{\text{vol}} \delta^{34}\text{S}_{\text{vol}} + f_{\text{terr}} \delta^{34}\text{S}_{\text{terr}} = \delta^{34}\text{S}_{\text{NSS}} \quad (5)$$

$$240 \quad f_{\text{bio}} + f_{\text{vol}} + f_{\text{terr}} = 1 \quad (6)$$

241
242 where $f_{\text{bio,vol,terr}}$ and $\delta^{34}\text{S}_{\text{bio,vol,terr}}$ are the fractions and isotopic signatures attributable to of sulfate from
243 the biogenic, volcanic and terrestrial sources, respectively. Here, we have used the values for $\delta^{34}\text{S}_{\text{bio}} =$
244 $18.8 \pm 0.3 \text{ ‰}$ and $\delta^{34}\text{S}_{\text{vol}} = 4.1 \pm 0.5 \text{ ‰}$ determined elsewhere (Kiyosu 1985; Jongebloed et al., 2023).
245 $\delta^{34}\text{S}_{\text{terr}}$ has been shown to vary regionally, as such we have used a global average of $4.4 \pm 4.5 \text{ ‰}$ (Burke
246 et al., 2018).
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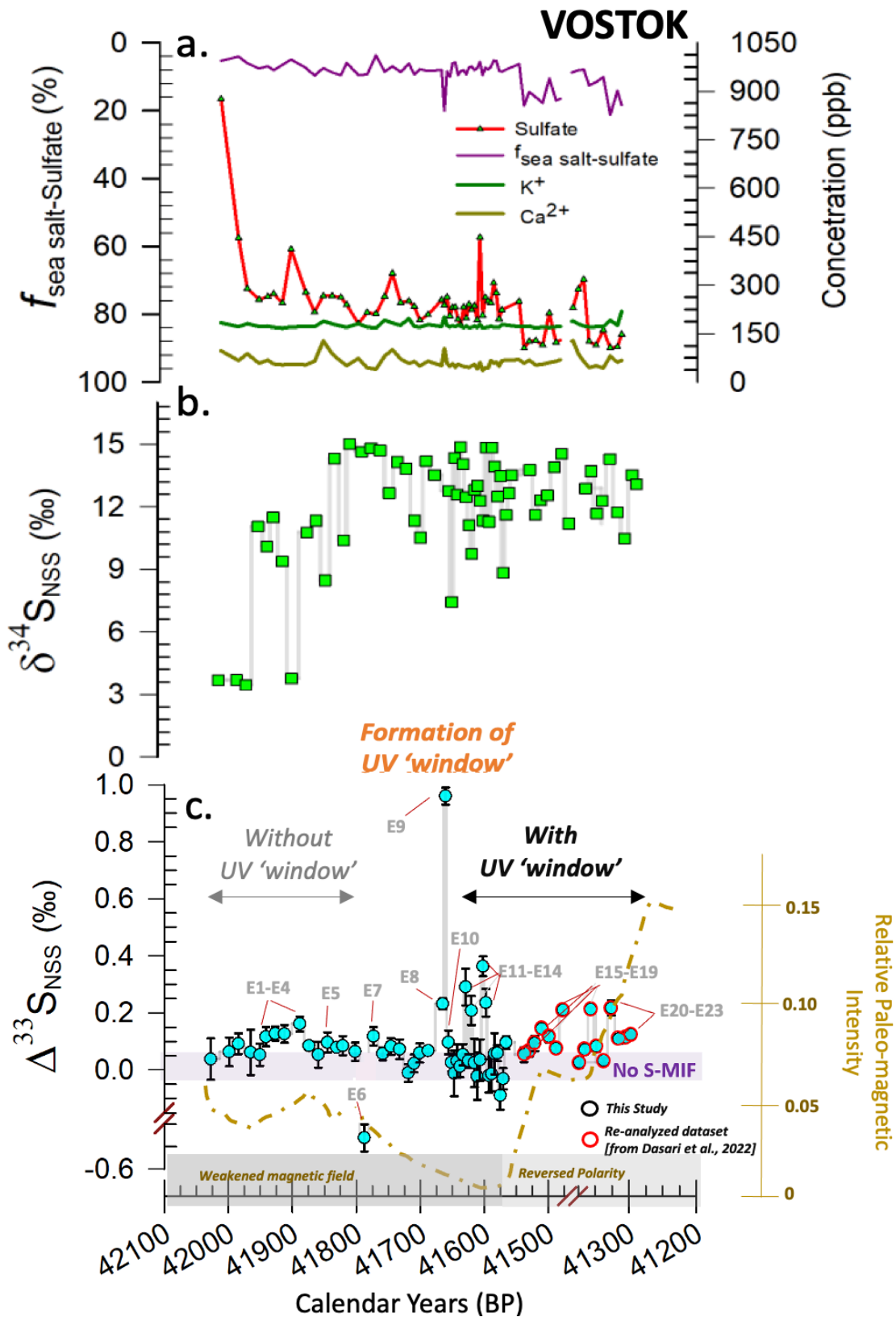
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A Bayesian mixing model was implemented in PyMC (Python) building on previous works (Dasari et al., 2020, 2021, 2022b, 2024) to quantify the proportional contributions of these three sulfur sources—'bio', 'vol' and 'terr'—to observed $\delta^{34}\text{S}$ values, where each source endmember was treated as a hierarchical random variable with a normal prior distribution defined by its mean and standard deviation to propagate endmember uncertainty through the model. For each sample, the source fractions (f_{bio} , f_{vol} , f_{terr}) were assigned a Dirichlet prior with $\alpha = [1,1,1]$, ensuring non-negativity and sum-to-one constraints while providing weakly informative regularization. The predicted $\delta^{34}\text{S}$ for each sample was calculated as the weighted sum of the three endmembers, and the likelihood was defined as a normal distribution centered on this predicted value with a standard deviation equal to the quadrature sum of the sample-specific analytical uncertainty (ranging from 0.02 to 0.28‰) and a 0.05‰ systematic uncertainty term to account for unmodeled processes. Posterior distributions were estimated using Markov Chain Monte Carlo (MCMC) simulation with four independent chains, each running 10,000 sampling iterations after 2000 warmup iterations, yielding 15,000 posterior draws per parameter. Convergence was assessed using the Gelman-Rubin R-hat statistic (<1.05 for all parameters) and effective sample size (>400), while posterior predictive checks confirmed adequate model fit. The output provides for each sample the mean source fractions, standard deviations, and 95% highest density intervals (HDI) for each source, allowing comprehensive uncertainty characterization that integrates analytical measurement error, endmember variability, and mixing model uncertainty into probabilistic source contribution estimates (Tables S4-S5 in Supporting Information). Further information on the model and the uncertainty estimation can be found elsewhere (Dasari et al., 2022b, 2024)

3. Results and Discussion

In the Antarctic ice-core sulfate record, considerable variability was found in the sulfate concentrations and the corresponding isotopic signals. The total measured sulfate concentrations, $\delta^{34}\text{S}$, and $\Delta^{33}\text{S}$ values varied from 100 to 900 ng g⁻¹, from 4 ‰ to 16 ‰, and from -0.1 ‰ to +0.5 ‰, respectively. The estimated average sea-salt sulfate fraction and sulfate flux were $10 \pm 6 \%$ and 3 ± 2 kg km⁻¹ yr⁻¹ over the sampled event, respectively. Upon correcting the record for the sea salt fraction, the non-sea salt sulfate concentrations ($[\text{SO}_4^{2-}]_{\text{NSS}}$), $\delta^{34}\text{S}_{\text{NSS}}$, and $\Delta^{33}\text{S}_{\text{NSS}}$ values varied from 80 to 830 ng g⁻¹, from 3.5 ‰ to 15 ‰, from -0.50 ‰ to +1 ‰, respectively (Figure 2; see also Table S1 in Supporting Information).

Here, the $\delta^{34}\text{S}_{\text{NSS}}$ values were found to be mostly stable with average value being $13 \pm 2 \%$, apart from the period between 42 ka BP and ≈ 41.9 ka BP when the average $\delta^{34}\text{S}_{\text{NSS}}$ was $7 \pm 4 \%$ (Figure 2). Changes to the $\delta^{34}\text{S}_{\text{NSS}}$ values were, however, concomitant with changes in the average $[\text{SO}_4^{2-}]_{\text{NSS}}$ which increased from 190 ± 70 ng/g (≈ 41.9 to 41.3 ka BP) to 320 ± 170 ng/g (≈ 42 and 41.9 ka BP), respectively (Figure 2). This contrasts with the $\Delta^{33}\text{S}_{\text{NSS}}$ values with maximum variability found in the period when both the $\delta^{34}\text{S}_{\text{NSS}}$ and $[\text{SO}_4^{2-}]_{\text{NSS}}$ were relatively stable (Figure 2). This amplitude of variations in the $\Delta^{33}\text{S}_{\text{NSS}}$ values in the weakened magnetic field phase (-0.53 ‰ to +0.96 ‰; new record) was found to be more pronounced than in the reversed polarity phase (+0.04 ‰ to +0.16 ‰; reanalysed pilot study record) (Figure 2). Overall, $72 \pm 19\%$ and $22 \pm 20\%$ of the non-sea salt sulfate in Antarctic ice-cores during this period was sourced from biogenic and volcanic emissions, respectively, with the remaining $6 \pm 4\%$ was sourced from terrestrial sources (see Methods; Table S4 in Supporting Information).



299
 300 **Figure 2. The evolution of Antarctic ice-core sulfate dynamics and isotopic signatures during the**
 301 **investigated 700-years of the Laschamp Geomagnetic Excursion event. (A) Sulfate along with other**
 302 **water-soluble ion concentrations are reported. The non-sea-salt sulfate fraction is shown. (B) $\delta^{34}\text{S}_{\text{NSS}}$**

303 values are reported. **(C)** $\Delta^{33}\text{S}_{\text{NSS}}$ values are reported from Antarctic ice-cores (see also Methods, Figure
304 S1 and Tables S2-S3 in Supporting Information) along with the periods of the formation and existence
305 of the UV window during the event. The relative paleomagnetic intensity (from Figure 1a) is shown
306 (dotted line) here for the analysed period. E1-E23 mark the events wherein anomalous sulfate
307 ($\Delta^{33}\text{S}\neq 0$) was detected during the investigated period. Note (in A - C) the data from this study (≈ 42 to
308 ≈ 41.5 ka BP) is merged with the reanalysed data from a previous study (Dasari et al., 2022) for the
309 periods ≈ 41.5 ka BP to ≈ 41.2 ka BP (see Methods) to create the 700-year record. The distinct phases
310 of the magnetic field evolution (see Figure 1) during this period are shown. The AICC2023 chronology
311 has been used (Bouchet et al., 2025).

312
313 In Greenland ice-core sulfate record, measured for the overlapping period of the magnetic field
314 collapse between ≈ 41.8 and ≈ 41.5 ka BP, the estimated average sea-salt sulfate fraction and sulfate
315 flux were 17 ± 3 % and 60 ± 20 kg km⁻¹ yr⁻¹ over the sampled event (Figure 3; see also Table S2 in
316 Supporting Information). Upon correcting the record for the sea salt fraction, the average non-sea salt
317 sulfate concentrations ($[\text{SO}_4^{2-}]_{\text{NSS}}$), $\delta^{34}\text{S}_{\text{NSS}}$, and $\Delta^{33}\text{S}_{\text{NSS}}$ values were 250 ± 90 ng/g, from 4.2 ± 1 ‰,
318 and $+0.08 \pm 0.01$ ‰, respectively (Figure 3). Overall, $14 \pm 9\%$ and $62 \pm 8\%$ of the non-sea salt sulfate
319 in Greenland ice-cores during this period was sourced from biogenic and volcanic emissions,
320 respectively, and the remaining $24 \pm 8\%$ was sourced from terrestrial sources (see Methods; Table S5
321 in Supporting Information). Taken together, the findings show considerable shifts in $\Delta^{33}\text{S}_{\text{NSS}}$ in the
322 southern polar ice-core record alone and a lack of similar response in the northern polar ice-core record.
323 This is a notable feature of the bi-polar ice-core sulfate record for the LGE event, and an aspect that
324 has not yet been found in other geochemical proxies from this period (Cooper et al., 2021; Reisbeck et
325 al., 2017).

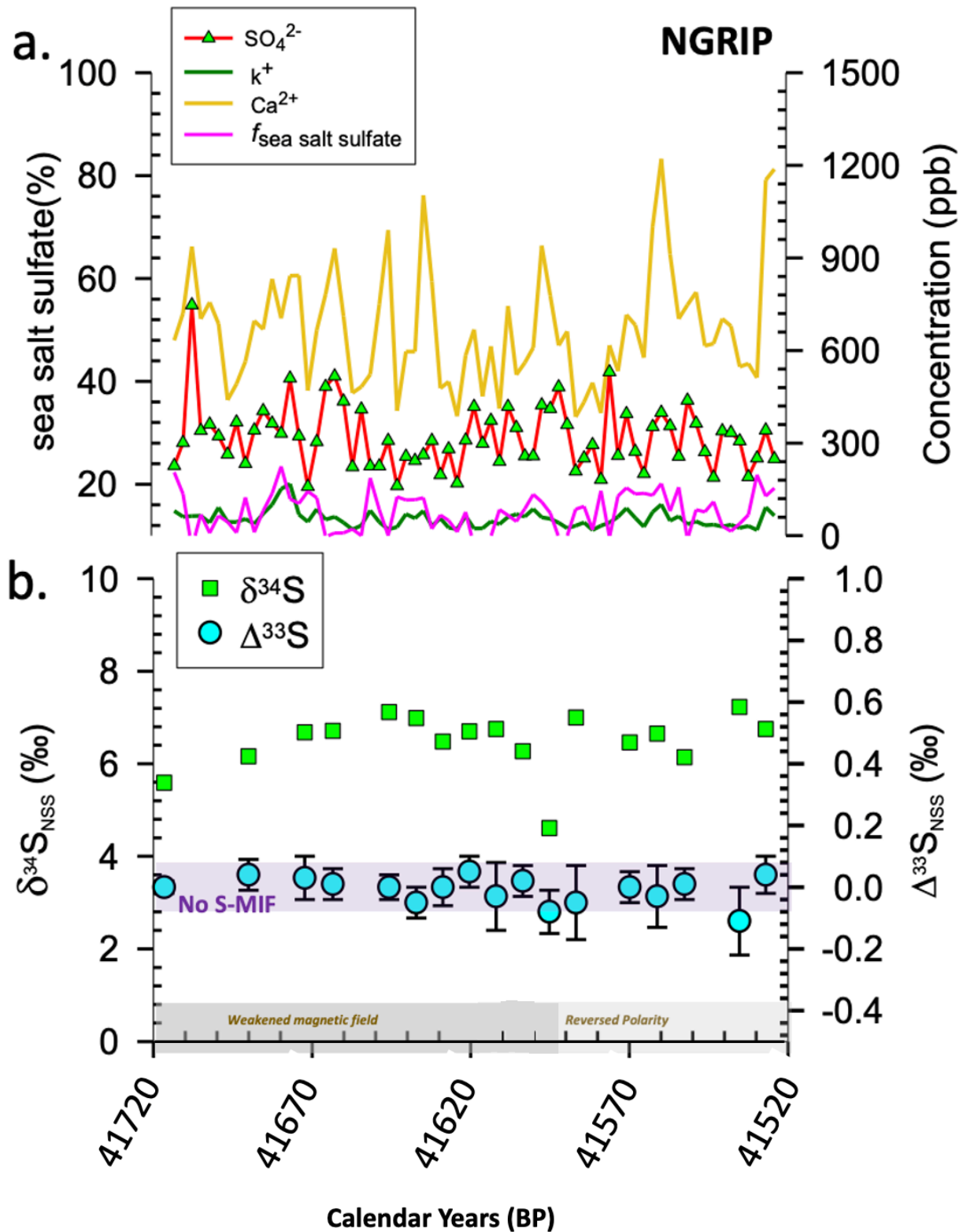


Figure 3. The evolution of North Greenland ice-core sulfate characteristics during the Laschamp Geomagnetic Excursion event for overlapping period of UV-window formation. (A) Sulfate along with other water-soluble ion concentrations are reported. The non-sea-salt sulfate fraction is shown.

(B) $\delta^{34}\text{S}_{\text{NSS}}$ values and $\Delta^{33}\text{S}_{\text{NSS}}$ values are reported from NGRIP ice-cores (see also Methods and Tables S2-S3 in Supporting Information). The distinct phases of the magnetic field evolution (see Figure 1) during this period are shown.

3.1 Investigating the origin of the observed S-MIF in the ice-core sulfate record

For the purpose of identifying the existence of a UV window, knowledge of the origin of S-MIF in ice-core sulfate—with the distinction of stratospheric volcanic sulfate—in the record is crucial (Dasari et al., 2022). During volcanic eruptions, the SO_2 oxidation kinetics imprinted in ice-core sulfate is such that the sulfur-isotope anomalies follow a cyclic pattern — changing from positive $\Delta^{33}\text{S}$ values at the beginning of the sulfate deposition to negative values at the end (Baroni et al., 2007; Burke et al., 2019; Gautier et al., 2018). The volcanic sulfate is superimposed on the marine-generated sulfate forming the continuous, low-concentration background (Baroni et al., 2008; Cole-Dai et al., 2021; Pearson et al., 2022). The volcanic deposition is often marked by an increase in sulfate concentrations and a corresponding decrease in $\delta^{34}\text{S}_{\text{NSS}}$ values in the ice-core record (Baroni et al., 2007; Burke et al., 2019; Baroni et al., 2008; Cole-Dai et al., 2021; Gautier et al., 2019; Pearson et al., 2022; Savarino et al., 2003).

Other than this aspect, the triple S-isotope signatures also offer the possibility to better understand the stratospheric volcanic origin of the observed S-MIF in ice-core sulfate record (Baroni et al., 2007; Gautier et al., 2019). As a consequence of the well-established stratospheric SO_2 oxidation kinetics and sedimentation dynamics, the $\Delta^{33}\text{S}_{\text{NSS}}$ versus $\delta^{34}\text{S}_{\text{NSS}}$ slope is constrained to be 0.09 ($\sigma = 0.02$) (Gautier et al., 2018) (Figure 5). Prior to this work, the validity of the slope had only been tested for ~2500 years of recorded polar ice-core S-MIF history (Gautier et al., 2019). As part of the pilot study covering the reversed polarity phase of the LGE event, a stratospheric volcanic eruption was recorded during ≈ 41.2 ka BP (Dasari et al., 2022). This allows for an opportunity to further test the validity of the slope for 41 millennia ago, which notably is found to be in alignment with the previously

347 established slope (Figure 5), thereby asserting the validity of the slope as a means to delineate
348 stratospheric volcanic influence on the observed S-MIF bearing sulfate in the ice-core record.

349 In the present study, the observed triple S-isotope systematics between 42 ka BP and \approx 41.9 ka
350 BP, accompanied with high sulfate concentration is suggestive of a stratospheric volcanic influence.
351 However, we note the time resolution of the ice-core samples was \sim 15 and 20 years, as such detecting
352 any cyclic S-MIF pattern from volcanic eruptions (as opposed to in studies with much higher time
353 resolution wherein such pattern was detected) was highly unlikely. Yet, S-MIF was detected in 3 out
354 of 8 samples during this period (and more in the remainder of the record as discussed later; SI Table
355 S1). Here, we refer to the samples with detectable S-MIF during the LGE as ‘Events (E; Figure 2)’ and
356 explore these further : the $\Delta^{33}\text{S}$ values in the 3 events (marked ‘E1-E3’; Figure 2) displayed an
357 enrichment in $\delta^{34}\text{S}_{\text{NSS}}$ concomitant to the sulfate becoming dominated with volcanic origin (Figure 4),
358 and were in alignment with observed volcanic slope (Figure 5). This is in line with the established
359 pattern of S-MIF from stratospheric volcanic eruptions and could rather be linked to extratropical
360 volcanoes (Burke et al., 2019). During such eruptions the initial transport of the volcanic plume is
361 expected to happen in the troposphere which could also lead to a depletion in the $\delta^{34}\text{S}_{\text{NSS}}$ and
362 concomitant increase in sulfate concentration in ice-core sulfate (as observed in the first 50 years of the
363 record; Figure 2), with the S-MIF bearing sulfate approaching later after the event (E1-E3; Figure 5).
364 However, such volcanic events last only a few years (Burke et al., 2019) but here the sulfur-isotope
365 anomalies in ‘E1-E3’ were observed on multi-decadal time resolution (spanning 60 years) which is not
366 expected from a ‘volcanic only’ process. As this period is concomitant with the abrupt changes in the
367 magnetic field (Figure 1a), it is therefore plausible that a secondary process could have been causing
368 this sustained detection of S-MIF bearing sulfate in the troposphere during this period. Further
369 investigation of such events in the remainder of the record could yield valuable insights into this
370 secondary process.

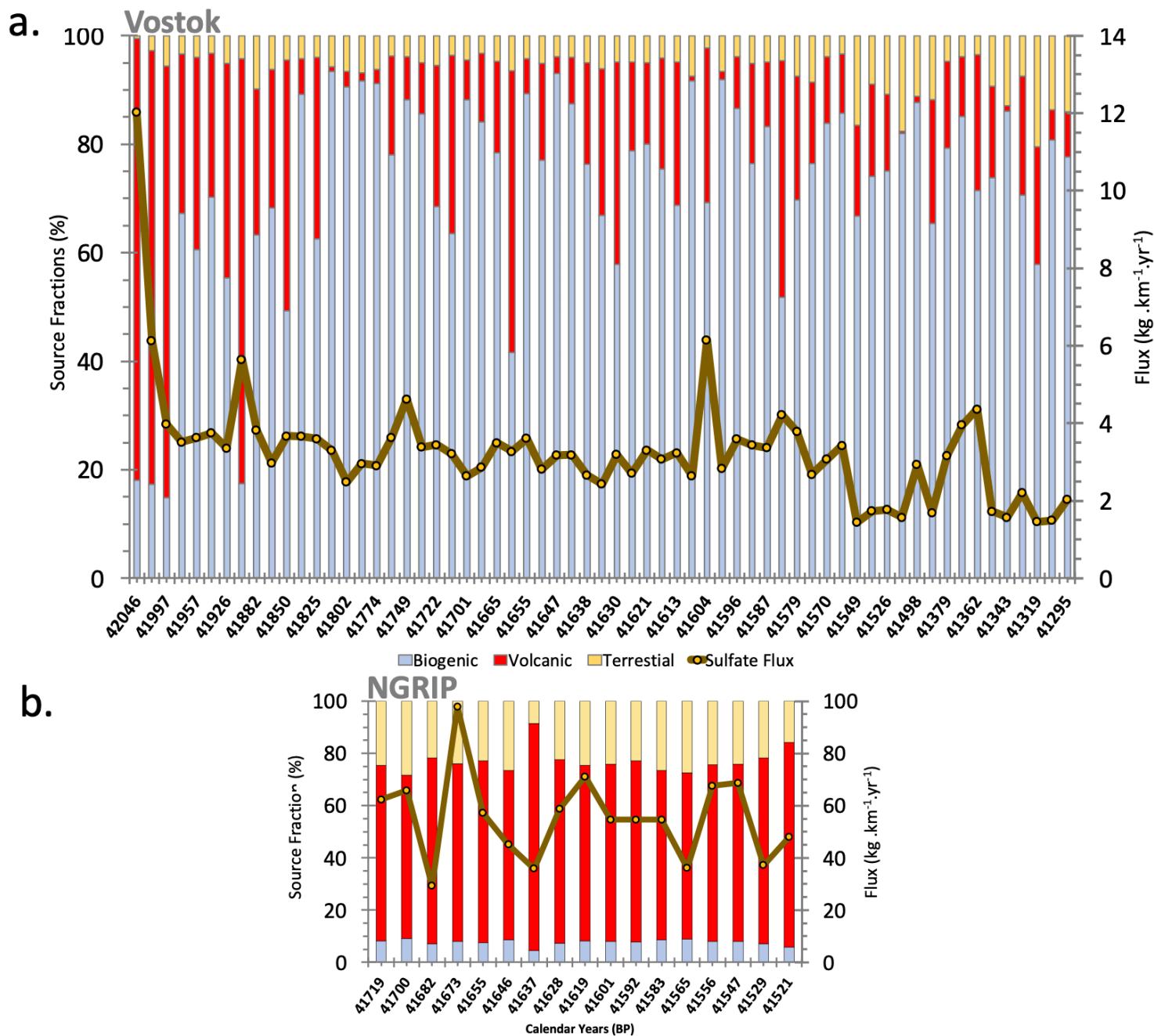
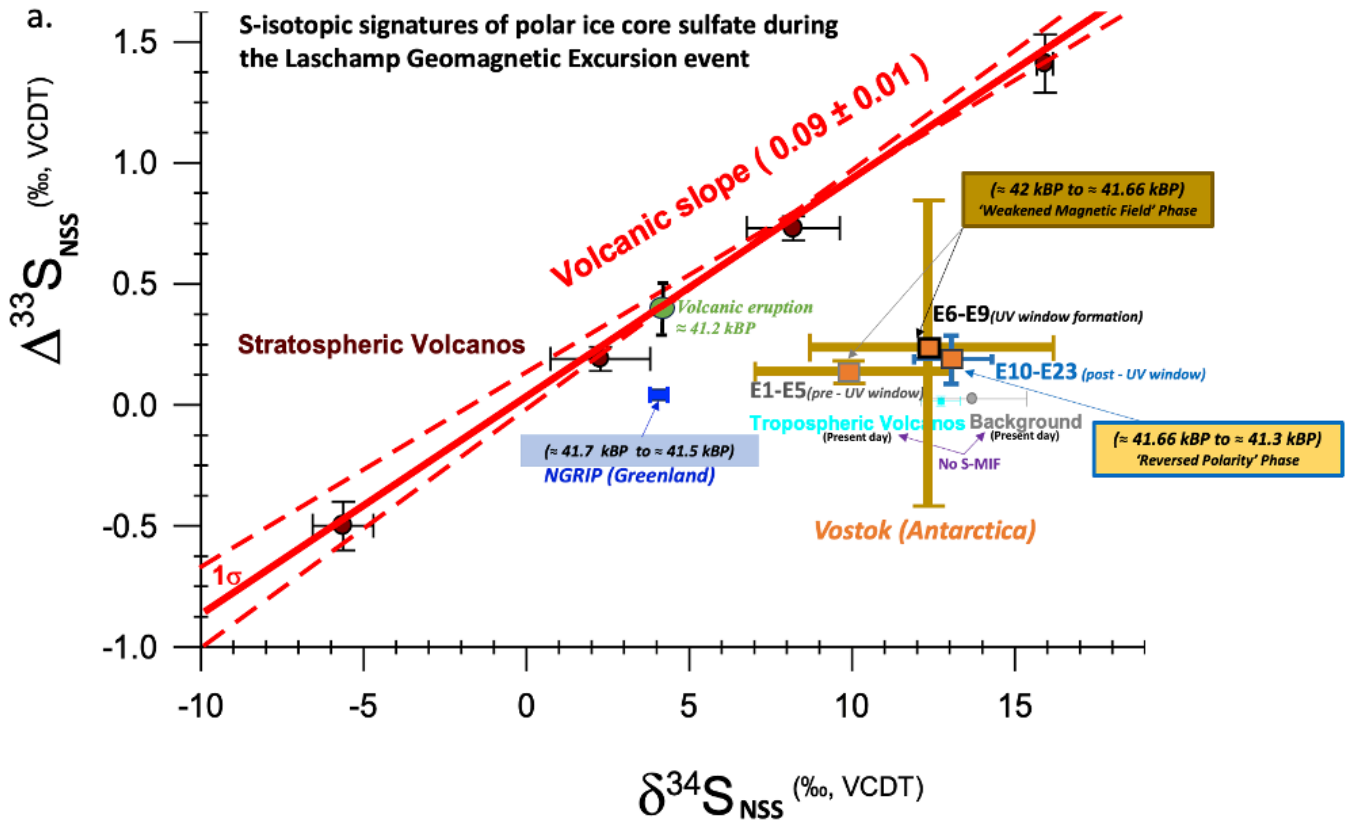


Figure 4. Source apportionment of polar ice core sulfate during the Laschamp Geomagnetic Excursion event. The observed $\delta^{34}\text{S}_{\text{NSS}}$ values are used to apportion the ice-core sulfate into marine biogenic, volcanic and terrestrial source fractions (see Methods) for Vostok (A) and NGRIP (B) samples used in this study (see Tables S1-S2 in Supporting Information). Sulfate flux computed for both these locations is also shown.

371

372 In fact, the triple S-isotope signatures of the remaining events ('E4-E16'; Figure 5) in the period
373 between ≈ 41.9 ka BP and ≈ 41.5 ka BP are not consistent with the volcanic slope even if we consider
374 uncertainties. This is also the case for the reanalysed $\Delta^{33}\text{S}$ values between ≈ 41.5 ka BP and ≈ 41.3 ka BP;
375 'E17-E23' (from reanalysed data from the pilot study; Figures 2 and 5). Furthermore, the sulfate concentrations
376 and $\delta^{34}\text{S}_{\text{NSS}}$ values for these periods (Figure 2) are comparable to that of the previously reported background
377 values across different sites in Antarctica, such as at Dome C and the South pole (concentration of 80 ± 20
378 ng/g, $\delta^{34}\text{S}_{\text{NSS}}$ of 14 ± 2 ‰) (Jourdin et al., 2002; Legrand et al., 2017). This indeed implies that S-MIF bearing
379 tropospheric sulfate being present in 'background' Antarctic atmosphere between ≈ 41.9 ka BP and ≈ 41.3 ka
380 BP, i.e., for the combined 600-year period of the LGE event.

381 The finding of S-MIF in tropospheric sulfate aerosols across a period of 600-years is intriguing. This
382 is because an estimated $\sim 80\%$ (20%) of the sulfate in 'E4-E23' was sourced from marine biogenic (volcanic)
383 emissions (Figure 4), which in general are not found to be associated with significant S-MIF ($\Delta^{33}\text{S}_{\text{biogenic}} \leq 0.05$
384 ‰; $\Delta^{33}\text{S}_{\text{tropospheric volcanoes}} \sim 0$ ‰) (Alexander et al., 2003; Baroni et al., 2008; Burke et al., 2019). This is also
385 evidenced in the year-round investigation of atmospheric sulfate aerosols in Antarctica, wherein the reported
386 $\Delta^{33}\text{S}$ was ≈ 0.01 ‰ on average for samples collected at both Dome C and Dumont d'Urville research stations
387 (Ishino et al., 2019). Likewise, sulfate formed in the troposphere and found in the snow or soil of Antarctica
388 has never shown any sulfur isotope anomaly either (Alexander et al., 2003). The amplitude of isotopic shifts
389 ($\Delta^{33}\text{S}$ varying between -0.5 ‰ to $+1$ ‰) for marine-biogenic and tropospheric volcanic emissions sourced
390 sulfate are therefore unique from the perspective of S-MIF based investigations in pristine environments. This
391 potentially implicates deeper penetration of UV radiation in the troposphere as the secondary process
392 generating S-MIF (other than stratospheric volcanic influence potentially for 'E1-E3'; Figure 5) during the
393 investigated period of the LGE event.



b. Deviation of the UV window S-isotopic signatures from the volcanic slope in ice-core sulfate

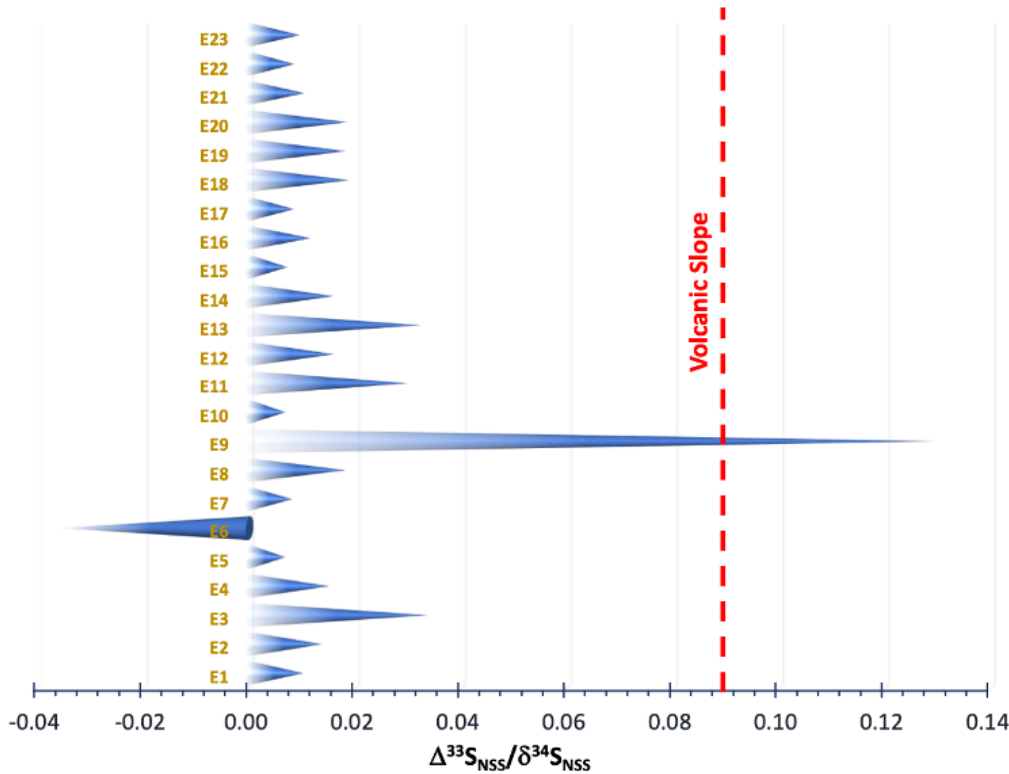


Figure 5. Triple-isotope-based deconvolution of the S-MIF signals during the Laschamp geomagnetic excursion event. (A) Five stratospheric volcanoes (circles), Kuwae, Samalas, 1809 UE, 1259 UE, and

398 Agung, are shown. Also shown is the 'volcanic' eruption detected during the Laschamp geomagnetic
399 excursion event in the pilot study¹¹ showing the validity of the volcanic slope during the corresponding
400 period as well. The isotopic signatures of previously reported tropospheric volcanoes and background
401 sulfate are shown. The averaged UV window isotopic signatures (E1-E23) in various phases are shown.
402 (B) Deviation of the events E1-E23 from volcanic slope is shown.

403
404 An increase in stratospheric tropospheric transport of airmasses is a possibility bringing in sulfate
405 bearing S-MIF and the existence linked to a DO event. Such an event could not have spanned the An
406 influx of UV radiation towards the Earth's surface is possible through the thinning/destruction of the
407 ozone layer during geomagnetic excursions as shown in Figure 1b, which is likely the case during the
408 investigated period. Taken together, the detection of S-MIF bearing sulfate in background Antarctic
409 conditions during the LGE event here unequivocally confirms the presence of a UV window during \approx
410 41.9 ka BP to \approx 41.3 ka BP period. We postulate that while the annually forming ozone hole over
411 Antarctica for short duration (few months) does not show any measurable S-MIF in ice-core sulfate,
412 geomagnetic excursions in the past can ideally be traced in Antarctic ice-core records. While mostly
413 positive sulfur-isotope anomalies values were observed during the LGE event, it is plausible that the
414 negative S-isotopic anomalies created in the process are stored in other S-reservoirs (e.g., 'COS', 'SO',
415 'CS₂', CH₃SO_{3H}) and as such remain undetected in the present ice-core 'sulfate' isotopic record
416 (Colman et al.,1996; Wanatabe et al., 2009; Shaheen et al., 2014). Some reports have also suggested
417 detection of only positive anomalies in ice-core sulfate during volcanic eruptions and the negative
418 anomalies were either lost, not detected in those ice-cores or remain less transported to Antarctica in
419 general (Baroni et al., 2007; Gautier et al., 2016. It is here speculated that this remains a possibility in
420 the present sampling scheme as well.

3.2 Testing the globality of the UV window

The geographical representativeness of the sampled ice-cores can be used to investigate the globality of the UV window. While the sampled Greenland ice-core sulfate record here is limited to 200-years, this overlapping period (≈ 41.7 ka BP and ≈ 41.5 ka BP; SI Table S3) with the Antarctic ice-core records maximizes the possibility for ascertaining the globality of the UV window (given the concomitance with the lowest relative paleomagnetic intensity; see Figure 1a). Long-term volcanic records have shown the detection of ‘global’ events in bi-polar ice-core sulfate records (Gautier et al., 2019; Lin et al., 2022; Sigl et al., 2013; Svensson et al., 2020). In general, a global volcanic event is often found to be synchronous in bi-polar ice-core sulfate records. In the present Antarctic record here, initially between ≈ 42 and 41.9 ka BP we find fewer and sporadic events (E1-E3) of S-MIF bearing sulfate in Antarctica (Figure 2). Immediately following this period, striking positive (E6) and negative peaks (E9) in sulfur-isotope anomalies are detected between $\approx 41,787$ BP and $\approx 41,659$ BP (Figure 2), which are synchronous with the weakest magnetic field intensity (Figure 2) and suggests the possibility of the UV window being formed in the Earth’s atmosphere during these ~ 100 years of the LGE event. Beyond 41,659 BP and into the reversed polarity phase of the LGE event, the sporadicity in detecting sulfur-isotope anomalies in Antarctic ice-core record also decreases and more events are detected (E10-E14 still in the weakened magnetic field phase and E15-E23 in the reversed polarity phase; Figure 2) implying the sustained presence of the UV window in the Earth’s atmosphere. Further investigation into background sulfate ≈ 41 ka BP in the pilot study revealed no detectable S-MIF implying that the UV window (Dasari et al., 2022a), and thereby the ozone layer, could have plausibly recovered by that period, in line with the regaining of the magnetic field intensity (Figure 1a). This is therefore potentially new evidence regarding the ‘timing’ of formation of a UV window in the Earth’s atmosphere during the LGE event.

Unlike the Antarctic record, the observed ice-core sulfate dynamics in Greenland during the sampled overlapping period of the LGE event showed no measurable S-MIF in background sulfate in

447 the north (Figure 3). The absence of the discernable S-MIF in northern polar ice-core records can be
448 linked to the several putative possibilities which we explore in detail:

449 (i) A higher snow accumulation rate, and thereby the sulfate flux (Figure 4), in Greenland
450 compared to Antarctica could lead to dilution of the S-MIF signal (Leonard et al., 2004; Neilsen
451 et al., 2014). This is possible only under the assumption that most of the sulfate transported to
452 Greenland was predominately formed from mass-dependent fractionation (MDF) pathways. But
453 events E5-E23 show that ‘background’ tropospheric sulfate (detected across the 600-year
454 period) was S-MIF bearing during this period. There is no clear way to apportion the fraction
455 of MIF-bearing and MDF-bearing sulfate in the sulfate flux.

456 (ii) Alternatively, the $\delta^{34}\text{S}_{\text{NSS}}$ values of Greenland ice-core sulfate suggest that sulfate could have
457 been sourced largely from continental origin (Jonegbløed et al., 2023) relative to the marine
458 biogenic emissions (Figure 4). Conflicting reports have emerged regarding S-MIF in continental
459 sulfate with recent studies suggesting such possibility exists in urban locations perhaps
460 speculatively linked to the role of organics whereas chamber studies in the past have advocated
461 against this process (au Yang et al., 2019; Dasari et al., 2022b, 2024). Moreover, combustion-
462 associated sulfate aerosols in biomass burning have also displayed S-MIF. Sulfate from biomass
463 burning has been mostly associated with positive $\Delta^{33}\text{S}$ anomalies (Dasari et al., 2023) whereas
464 from biofuel (e.g., wood) burning with negative $\Delta^{33}\text{S}$ anomalies (Han et al., 2017; Pei et al.,
465 2026). The $\delta^{34}\text{S}$ value of such sources are rather similar to the ones found in the Greenland ice-
466 core sulfate record, yet, lacking evidence of any S-MIF imprint. Other than this source, the
467 remaining sulfate was primarily sourced from a mixture of tropospheric volcanic and marine
468 biogenic emissions (Table S2 in Supporting Information and Figure 4). As found in the
469 Antarctic record (Figure 5), sulfate from a mix of these sources did display S-MIF during the
470 investigated period of 600 years of the LGE event linked to the presence of a UV window.

471 (iii) While accounting for the role of atmospheric deposition and dynamics during the LGE event
472 remains beyond the scope of the present study, it is plausible that the S-MIF bearing sulfate
473 could arrive at the NGRIP site later than the period investigated here (Gautier et al., 2016).
474 However, given the synchronicity between the collapse of the magnetic field and the events E6-
475 E9 (Figure 2), we argue that such a geochemical imprint should have been witnessed in the
476 northern ice core record either instantaneously or within the ~100-year period of covered by E6-
477 E9, as supported by bi-polar records of other cosmogenic tracers during this period (Reisbeck
478 et al., 2007, 2017) . Furthermore, during events of climate transitions such as the between
479 Greenland Interstadial and stadial periods, an increase in stratospheric tropospheric transport
480 (STT) has been reported (Geng et al., 2017). Here, the subsampled NGRIP ice core sulfate
481 covers such a transitional period (Cooper et al., 2021) yet no evidence of S-MIF bearing sulfate
482 is found. While further tracer-based fingerprinting (e.g. $\Delta^{17}\text{O}$ of ice core (NO_3^-)) is warranted
483 to better understand any potential influence of STT processes on this geochemical imprint, here
484 accounting for (i- iii), we argue that there exists another possibility of a geographically varying
485 UV-influx during the investigated phases of the LGE event (Mukhopadhyay et al., 2025), which
486 would make it feasible for sulfate sourced from similar sources to exhibit discernible anomalous
487 characteristic in different hemispheres.

488

489 **4. Conclusion and Implications**

490 Our study provides first direct evidence of the formation of a UV ‘window’ during the LGE
491 event. Despite the possibility of dilution effects impacting the S-MIF signal, the bi-polar ice-core sulfate
492 record alternatively supports the notion of a geographically varying influence of the UV window
493 formed from magnetic field collapse. Data-based global paleomagnetic field models could provide
494 novel insights into the formation of the geomagnetic excursion linked UV window. Quantities, such as
495 field morphology at the core-mantle boundary or individual spherical harmonic degree power variations

496 need to be accounted with caution (Korte et al., 2019). Modeling works relying on these parameters
497 suggest that the excursion process during the Laschamp was mainly governed by axial dipole decay
498 and recovery, without a significant influence from the equatorial dipole or non-dipole fields (Korte et
499 al., 2019; Leonhardt et al., 2009; Laj et al., 2006; Panovska et al., 2017). This results in an excursionsal
500 field behavior seen globally, but ‘non-uniformly’ at the Earth’s surface and is supportive of the notion
501 that the ensuing magnitude of the ozone depletion was also non-uniform fields (Korte et al., 2019; Laj
502 et al., 2006). Studies of the effects of cosmic radiation during paleo-magnetospheric polarity transitions
503 reveal that, with a reduction in dipole moment strength, lower latitudes including the tropical zone
504 become accessible to high fluxes of lower-energy particles, ushering in zones of impact for diverse
505 magnetospheric configurations (Mukhopadhyay et al., 2025). In the period of the overlapping bi-polar
506 ice-core record, the excursion was strongest in the Indian Ocean and the south Atlantic Ocean as well
507 as the southern Oceanic regions yet in the proximity of Antarctica (Korte et al., 2019; Mukhopadhyay
508 et al., 2025). While in the Northern hemisphere, the expanded auroral zones towards the equator suggest
509 such impact zones for high energy cosmic rays were further way from Greenland (Mukhopadhyay et
510 al., 2025). Therefore, it is possible that the effect of the UV window (detected ≈ 41.6 ka BP) strongly
511 materialized in the Southern Hemisphere than in the Northern hemisphere during the investigated
512 period of the LGE event. However, further research is warranted to better understand the globality of
513 the UV window during the entirety of the LGE event.

514 We also note that the atmospheric modeling of the NO_x -ozone chemistry is at odds with this
515 notion (Cooper et al., 2021; Suter et al., 2014). Atmospheric modelling studies have provided
516 quantitative estimates for ozone depletion during such events (Cooper et al., 2021; Suter et al., 2014;
517 Norval et al., 2007; Winkler et al., 2008). For instance, during the LGE event, such modeling works
518 estimated only ~ 5 % ozone loss in the lower stratosphere over Antarctica with a similar increase in
519 tropospheric ozone mixing ratios and the opposite for the Arctic region (Cooper et al., 2021). Such a
520 scenario would impede S-MIF-generating UV-induced photochemical processes. Furthermore, a

521 relatively higher ozone loss over the Arctic than over Antarctica should have led to the detection of the
522 UV window in the northern hemisphere with effect strongly materialized in the northern polar ice-cores
523 which is not the case. As such, considering the ‘non-uniform’ evolution of the magnetic field could be
524 valuable in accurately modeling the ozone depletion during such events. The morphological alterations
525 experienced by Earth’s magnetosphere and its auroral zones together with the geochemical imprint
526 shown in this study support the argument for further research on the evolving geomagnetic shielding
527 behavior during geomagnetic excursions to better understand the environmental, atmospheric and
528 climatic effects during such events.

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546 Guillaume Paris, Julien Charreau, Joel Savarino **Writing – review & editing:** Sanjeev Dasari,

547 Guillaume Paris, Julien Charreau, Alexis Lamothe, Joel Savarino

548
549 **Data Availability Statement**

550 All data is included in the manuscript and/or supporting information.

551
552 **Conflict of Interest**

553 The authors declare no conflict of interest relevant to this study.

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892 [16-1565-2020](https://doi.org/10.5194/cp-16-1565-2020)

899 **Supporting Information: Polar ice-cores unravel the formation of a**
900 **UV window during magnetic field collapse ~ 42 ka BP**

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908
909 **Contents of this file**

910 1. Tables S1 to S5

911 **Introduction** This document is used to provide additional tables to supplement specific details mentioned in
912 the manuscript.

913 **Table S1. S-isotope values for ice-core sulfate samples from Antarctica (site: Vostok).** The mean and
 914 standard deviations of the ensemble of the measurements are shown. Sampling resolution was increased for
 915 ice-core VK-611 and VK-612. This data is not corrected for contribution from sea salt. NA refers to data not
 916 available. The AICC2023 age chronology is followed here.

Average Depth (m)	Average Age (AICC2023; BP)	$\delta^{34}\text{S}$ (‰)		$\Delta^{33}\text{S}$ (‰)		Sulfate Flux kg km ⁻¹ yr ⁻¹
		Mean	s.d.	Mean	s.d.	
607.07±0.04	41295±4	14.54	0.02	0.05	0.02	2
607.19±0.04	41307±4	14.58	0.02	0.06	0.02	1
607.31±0.04	41319±4	12.72	0.08	0.03	0.04	1
607.43±0.04	41331±4	12.68	0.09	0.17	0.05	2
607.55±0.04	41343±4	15.06	0.05	0.00	0.02	2
607.66±0.03	41353±3	13.39	0.04	0.04	0.04	2
607.75±0.03	41362±3	12.41	0.12	0.17	0.02	4
607.84±0.03	41371±3	14.30	0.10	0.04	0.03	4
607.93±0.03	41379±3	13.59	0.15	0.00	0.04	3
609.07±0.04	41483±4	12.86	0.12	0.13	0.03	2
609.19±0.04	41498±4	15.22	0.03	0.04	0.02	3

609.31±0.04	41511±4	15.16	0.03	0.05	0.04	2
609.43±0.04	41526±4	13.91	0.06	0.08	0.03	2
609.54±0.03	41538±3	13.58	0.10	0.04	0.05	2
609.63±0.03	41549±2	13.34	0.07	0.00	0.04	1
609.73±0.04	41561±4	14.21	0.12	0.03	0.05	3
610.04	41570.00	14.15	0.14	0.05	0.04	3
610.09	41574.50	13.34	0.12	0.01	0.07	3
610.13	41579.00	12.11	0.08	-0.08	0.10	4
610.18	41583.50	9.48	0.05	0.04	0.06	4
610.22	41587.00	14.00	0.06	0.03	0.14	3
610.27	41591.50	13.16	0.07	-0.03	0.12	3
610.31	41596.00	14.45	0.04	-0.02	0.12	4
610.36	41600.50	15.42	0.06	0.09	0.10	3
610.40	41604.00	11.82	0.11	0.41	0.07	6
610.45	41608.50	15.29	0.06	0.01	0.14	3
610.49	41613.00	12.12	0.04	-0.03	0.17	3
610.54	41617.50	12.91	0.06	0.03	0.17	3
610.58	41621.00	13.59	0.11	0.17	0.10	3
610.63	41625.50	13.57	0.07	0.00	0.05	3
610.67	41630.00	10.66	0.04	0.25	0.13	3

610.72	41634.50	11.94	0.14	0.03	0.07	2
610.76	41638.00	13.27	0.02	-0.02	0.07	3
610.81	41642.50	14.45	0.10	0.02	0.12	3
610.85	41647.00	15.23	0.10	0.01	0.16	3
610.90	41651.50	13.43	0.05	0.01	0.06	3
610.94	41655.00	14.90	0.05	0.06	0.08	4
610.99	41659.50	14.38	0.03	0.03	0.06	3
611.07±0.03	41665±3	13.42	0.13	0.19	0.04	3
611.18±0.04	41678±4					
611.31±0.04	41690±4	14.15	0.04	0.04	0.01	3
611.43±0.03	41701±3	14.72	0.11	0.04	0.07	3
611.52±0.04	41711±3	11.50	0.05	0.00	0.06	3
611.63±0.04	41722±4	11.93	0.08	-0.02	0.06	3
611.77±0.04	41736±3	14.45	0.15	0.04	0.07	3
611.92±0.07	41749±3	14.62	0.09	0.06	0.06	5
612.09±0.04	41762±5	13.37	0.07	0.03	0.05	4
612.22±0.04	41774±5	14.94	0.08	0.10	0.06	3
612.36±0.04	41787±5	15.39	0.07	0.04	0.07	3
612.51±0.04	41802±5	15.23	0.10	0.03	0.06	2
612.65±0.03	41814±3	15.37	0.09	0.06	0.07	3

612.76±0.04	41825±5	11.40	0.10	0.04	0.03	4
612.87±0.04	41838±5	14.90	0.06	0.07	0.07	4
612.96±0.04	41850±6	9.40	0.13	0.03	0.09	4
613.13±0.04	41865±5	12.27	0.09	0.02	0.03	3
613.27±0.04	41882±5	11.53	0.09	0.13	0.05	4
613.38±0.03	41896±3					
613.49±0.04	41910±5	4.62	0.05	0.11	0.06	6
613.63±0.04	41926±5	10.14	0.08	0.07	0.05	3
613.74±0.03	41940±3	12.26	0.08	0.09	0.07	4
614.10±0.05	41957±5	10.84	0.15	0.18	0.08	4
614.27±0.06	41977±6	11.81	0.08	0.08	0.16	4
614.45±0.06	41997±5	4.51	0.09	0.07	0.07	4
614.62±0.06	42017±6	4.40	0.14	0.05	0.10	6
614.77±0.03	42034±3					
614.92±0.03	42048±3	4.61	0.15	0.02	0.14	12

917 **Table S2. S-isotope values for ice-core sulfate samples from Greenland (site: NGRIP).** The mean and
 918 standard deviations of the ensemble of the measurements are shown. This data is not corrected for contribution
 919 from sea salt. NA refers to data not available. The AICC2023 age chronology is followed here.

Average Depth (m)	Average Age (AICC2023; BP)	$\delta^{34}\text{S}$ (‰)		$\Delta^{33}\text{S}$ (‰)		Sulfate Flux kg km ⁻¹ yr ⁻¹
		Mean	s.d.	Mean	s.d.	
2126.34±0.04	41521±3	6.75	0.12	0.04	0.06	48
2126.46±0.04	41529±3	7.23	0.13	-0.11	0.11	37
2126.58±0.04	41538±3	NA				
2126.70±0.04	41547±3	6.14	0.02	0.01	0.05	69
2126.82±0.04	41556±3	6.65	0.08	-0.03	0.10	68
2126.94±0.04	41565±3	6.46	0.11	0.00	0.05	36
2127.06±0.04	41574±3	NA				
2127.18±0.04	41583±3	7.00	0.03	-0.05	0.12	55
2127.30±0.04	41592±3	4.61	0.07	-0.08	0.07	54
2127.42±0.04	41601±3	6.27	0.06	0.02	0.05	78
2127.54±0.04	41619±3	6.75	0.14	-0.03	0.11	71
2127.66±0.04	41628±3	6.70	0.05	0.05	0.05	59
2127.78±0.04	41637±3	6.48	0.28	0.00	0.06	36
2127.90±0.04	41646±3	6.99	0.05	-0.05	0.05	45

2128.02±0.04	41655±3	7.12	0.07	0.00	0.04	57
2128.14±0.04	41664±3					
2128.26±0.04	41673±3	6.71	0.17	0.01	0.05	98
2128.38±0.04	41682±3	6.68	0.15	0.03	0.07	29
2128.50±0.04	41691±3	NA				
2128.62±0.04	41700±3	6.16	0.07	0.04	0.05	66
2128.74±0.04	41709±3	NA				
2129.00±0.05	41719±4	5.59	0.07	0.00	0.03	62

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921 **Table S3. Age transfer function for identifying corresponding periods of the Laschamp geomagnetic**
 922 **excursion on the NGRIP ice-cores.** The periods are corresponding to the periods where the maximum isotopic
 923 anomalies were detected in the Vostok ice-cores. Bi-polar mp = bi-polar volcanic match point from Table 1
 924 and Table S2 from Svensson et al. 2020. For the bi-polar matching we first convert AICC2023 ages (41503-
 925 41680) to AICC2012 ages (Veres et al., 2013; Bazin et al., 2013).

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Event	AICC2012age (BP 1950)	Vostok depth (m)	EDML depth (m)	NGRIP depth (m)	GICC05 age (b2k)
Bi-polar mp			1378.51	2123.98	
Interval start	41540	609.71	1382.10	2126.52	41623
Interval end	41710	611.53	1385.42	2128.87	41790
Bi-polar mp			1387.31	2130.21	

Table S4. Bayesian statistical source apportionment results for VOSTOK ice core sulfate samples.

$\delta^{34}\text{S}$ (‰)	Source	Source Fraction	SD	95% CI Lower	95% CI Upper	CV (%)
14.54 ± 0.02	Biogenic	0.78	0.04	0.65	0.84	6
	Volcanic	0.08	0.03	0.07	0.20	24
	Terrestrial	0.14	0.02	0.06	0.16	24
14.58 ± 0.02	Biogenic	0.81	0.04	0.65	0.84	6
	Volcanic	0.06	0.03	0.07	0.20	24
	Terrestrial	0.14	0.02	0.06	0.16	24
12.72 ± 0.08	Biogenic	0.58	0.05	0.61	0.80	7
	Volcanic	0.22	0.03	0.10	0.25	20
	Terrestrial	0.20	0.02	0.06	0.16	25
12.68 ± 0.09	Biogenic	0.71	0.05	0.61	0.80	7
	Volcanic	0.22	0.03	0.10	0.25	20
	Terrestrial	0.07	0.02	0.06	0.16	25
15.06 ± 0.05	Biogenic	0.86	0.04	0.66	0.85	6
	Volcanic	0.01	0.03	0.06	0.19	25
	Terrestrial	0.13	0.02	0.06	0.16	24
13.39 ± 0.04	Biogenic	0.74	0.04	0.63	0.82	6.7
	Volcanic	0.17	0.03	0.09	0.23	22
	Terrestrial	0.09	0.02	0.06	0.16	25
12.41 ± 0.12	Biogenic	0.72	0.05	0.60	0.8	7
	Volcanic	0.25	0.03	0.11	0.26	20
	Terrestrial	0.03	0.02	0.06	0.16	25
14.30 ± 0.10	Biogenic	0.85	0.04	0.65	0.84	6
	Volcanic	0.11	0.03	0.07	0.21	23
	Terrestrial	0.04	0.02	0.06	0.16	24

	Biogenic	0.79	0.04	0.63	0.83	6
13.59 ± 0.15	Volcanic	0.16	0.03	0.08	0.22	22
	Terrestrial	0.05	0.02	0.06	0.16	25
	Biogenic	0.65	0.05	0.61	0.81	7
12.86 ± 0.12	Volcanic	0.23	0.03	0.10	0.24	21
	Terrestrial	0.12	0.02	0.06	0.16	25
	Biogenic	0.88	0.04	0.67	0.85	6
15.22 ± 0.03	Volcanic	0.01	0.03	0.06	0.19	25
	Terrestrial	0.11	0.02	0.06	0.16	24
	Biogenic	0.82	0.04	0.67	0.85	6
15.16 ± 0.03	Volcanic	0.00	0.03	0.06	0.19	26
	Terrestrial	0.18	0.02	0.06	0.16	24
	Biogenic	0.75	0.04	0.64	0.83	6
13.91 ± 0.06	Volcanic	0.14	0.03	0.08	0.21	23
	Terrestrial	0.11	0.02	0.06	0.16	25
	Biogenic	0.74	0.04	0.63	0.83	6.7
13.58 ± 0.10	Volcanic	0.17	0.03	0.08	0.22	22
	Terrestrial	0.09	0.02	0.06	0.16	25
	Biogenic	0.67	0.04	0.63	0.82	6
13.34 ± 0.07	Volcanic	0.17	0.03	0.09	0.23	22
	Terrestrial	0.16	0.02	0.06	0.16	25
	Biogenic	0.86	0.04	0.65	0.83	6.5
14.21 ± 0.12	Volcanic	0.11	0.03	0.07	0.21	23
	Terrestrial	0.03	0.02	0.06	0.16	24
	Biogenic	0.84	0.04	0.64	0.83	6.5
14.15 ± 0.14	Volcanic	0.12	0.03	0.08	0.21	23

	Terrestrial	0.04	0.02	0.06	0.16	24
	Biogenic	0.76	0.04	0.63	0.82	6
13.34 ± 0.12	Volcanic	0.15	0.03	0.09	0.23	22
	Terrestrial	0.09	0.02	0.06	0.16	25
	Biogenic	0.70	0.05	0.59	0.79	7.3
12.11 ± 0.08	Volcanic	0.23	0.03	0.12	0.27	20
	Terrestrial	0.07	0.02	0.06	0.16	25
	Biogenic	0.52	0.05	0.51	0.72	8
9.48 ± 0.05	Volcanic	0.44	0.04	0.18	0.351	16
	Terrestrial	0.05	0.02	0.06	0.174	24
	Biogenic	0.83	0.04	0.64	0.83	6.7
14.00 ± 0.06	Volcanic	0.12	0.03	0.08	0.22	23
	Terrestrial	0.05	0.02	0.06	0.16	25
	Biogenic	0.76	0.04	0.62	0.81	6
13.16 ± 0.07	Volcanic	0.18	0.03	0.09	0.23	21
	Terrestrial	0.05	0.02	0.06	0.16	25
	Biogenic	0.87	0.04	0.65	0.84	6
14.45 ± 0.04	Volcanic	0.09	0.03	0.07	0.20	24
	Terrestrial	0.04	0.02	0.06	0.16	24
	Biogenic	0.92	0.04	0.67	0.85	6
15.42 ± 0.06	Volcanic	0.01	0.03	0.06	0.18	26
	Terrestrial	0.07	0.02	0.06	0.16	24
	Biogenic	0.69	0.05	0.58	0.78	7
11.82 ± 0.11	Volcanic	0.29	0.03	0.12	0.28	19
	Terrestrial	0.02	0.02	0.06	0.16	25
15.29 ± 0.06	Biogenic	0.92	0.04	0.67	0.85	6

	Volcanic	0.01	0.03	0.06	0.18	25
	Terrestrial	0.07	0.02	0.06	0.16	24
	Biogenic	0.69	0.05	0.59	0.79	7
12.12 ± 0.04	Volcanic	0.26	0.03	0.12	0.27	20
	Terrestrial	0.05	0.02	0.06	0.16	25
	Biogenic	0.75	0.05	0.61	0.81	7
12.91 ± 0.06	Volcanic	0.20	0.03	0.10	0.24	21
	Terrestrial	0.04	0.02	0.06	0.16	25
	Biogenic	0.80	0.04	0.63	0.83	6
13.59 ± 0.11	Volcanic	0.15	0.03	0.08	0.22	22
	Terrestrial	0.05	0.02	0.06	0.16	25
	Biogenic	0.79	0.04	0.63	0.82	6
13.57 ± 0.07	Volcanic	0.16	0.03	0.08	0.22	22
	Terrestrial	0.05	0.02	0.06	0.16	25
	Biogenic	0.58	0.05	0.55	0.75	7
10.66 ± 0.04	Volcanic	0.37	0.04	0.15	0.31	17
	Terrestrial	0.05	0.02	0.06	0.17	25
	Biogenic	0.67	0.05	0.58	0.78	7.4
11.94 ± 0.14	Volcanic	0.27	0.03	0.12	0.27	19
	Terrestrial	0.06	0.02	0.06	0.16	25
	Biogenic	0.76	0.04	0.62	0.82	6.8
13.27 ± 0.02	Volcanic	0.19	0.03	0.09	0.23	22
	Terrestrial	0.05	0.02	0.06	0.16	25
	Biogenic	0.87	0.04	0.65	0.84	6.4
14.45 ± 0.10	Volcanic	0.09	0.03	0.07	0.20	24
	Terrestrial	0.04	0.02	0.06	0.16	24
15.23 ± 0.10	Biogenic	0.93	0.04	0.67	0.85	6

	Volcanic	0.03	0.03	0.06	0.19	25
	Terrestrial	0.04	0.02	0.06	0.16	24
	Biogenic	0.77	0.04	0.63	0.82	6
13.43 ± 0.05	Volcanic	0.18	0.03	0.09	0.22	22
	Terrestrial	0.05	0.02	0.06	0.16	25
	Biogenic	0.89	0.04	0.66	0.84	6
14.90 ± 0.05	Volcanic	0.06	0.03	0.07	0.20	24
	Terrestrial	0.04	0.02	0.06	0.16	24
	Biogenic	0.42	0.04	0.65	0.84	6.4
14.38 ± 0.03	Volcanic	0.52	0.03	0.07	0.21	23
	Terrestrial	0.06	0.02	0.06	0.16	24
	Biogenic	0.78	0.04	0.63	0.82	6.7
13.42 ± 0.13	Volcanic	0.17	0.03	0.09	0.22	22
	Terrestrial	0.05	0.02	0.06	0.16	25
	Biogenic	0.84	0.04	0.64	0.83	6.5
14.15 ± 0.04	Volcanic	0.13	0.03	0.08	0.21	23
	Terrestrial	0.03	0.02	0.06	0.16	24
	Biogenic	0.88	0.04	0.66	0.84	6
14.72 ± 0.11	Volcanic	0.07	0.03	0.07	0.20	23
	Terrestrial	0.05	0.02	0.06	0.16	24
	Biogenic	0.64	0.05	0.57	0.77	7
11.50 ± 0.05	Volcanic	0.33	0.04	0.13	0.29	19
	Terrestrial	0.04	0.02	0.06	0.16	25
	Biogenic	0.69	0.05	0.58	0.78	7
11.93 ± 0.08	Volcanic	0.26	0.03	0.12	0.27	19
	Terrestrial	0.05	0.02	0.06	0.16	25

	Biogenic	0.86	0.04	0.65	0.84	6
14.45 ± 0.15	Volcanic	0.09	0.03	0.07	0.20	24
	Terrestrial	0.05	0.02	0.06	0.16	24
	Biogenic	0.88	0.04	0.65	0.84	6
14.62 ± 0.09	Volcanic	0.08	0.03	0.07	0.20	24
	Terrestrial	0.04	0.02	0.06	0.16	24
	Biogenic	0.78	0.04	0.63	0.82	6
13.37 ± 0.07	Volcanic	0.18	0.03	0.09	0.23	22
	Terrestrial	0.04	0.02	0.06	0.16	25
	Biogenic	0.91	0.04	0.66	0.84	6
14.94 ± 0.08	Volcanic	0.03	0.03	0.07	0.20	24
	Terrestrial	0.06	0.02	0.06	0.16	24
	Biogenic	0.92	0.04	0.67	0.85	6
15.39 ± 0.07	Volcanic	0.01	0.03	0.06	0.18	26
	Terrestrial	0.07	0.02	0.06	0.16	24
	Biogenic	0.91	0.04	0.67	0.85	6
15.23 ± 0.10	Volcanic	0.03	0.03	0.06	0.19	25
	Terrestrial	0.07	0.02	0.06	0.16	24
	Biogenic	0.93	0.04	0.67	0.85	6
15.37 ± 0.09	Volcanic	0.01	0.03	0.06	0.18	26
	Terrestrial	0.06	0.02	0.06	0.16	24
	Biogenic	0.63	0.05	0.57	0.77	7.6
11.40 ± 0.10	Volcanic	0.33	0.04	0.13	0.29	18
	Terrestrial	0.04	0.02	0.06	0.16	25
	Biogenic	0.89	0.04	0.66	0.84	6
14.90 ± 0.06	Volcanic	0.07	0.03	0.07	0.20	24

	Terrestrial	0.04	0.02	0.06	0.16	24
	Biogenic	0.49	0.05	0.51	0.71	8
9.40 ± 0.13	Volcanic	0.46	0.04	0.18	0.35	16
	Terrestrial	0.04	0.02	0.06	0.17	24
	Biogenic	0.68	0.05	0.60	0.79	7
12.27 ± 0.09	Volcanic	0.26	0.03	0.11	0.26	20
	Terrestrial	0.06	0.02	0.06	0.16	25
	Biogenic	0.63	0.05	0.57	0.77	7
11.53 ± 0.09	Volcanic	0.27	0.04	0.13	0.29	19
	Terrestrial	0.10	0.02	0.06	0.16	25
	Biogenic	0.17	0.05	0.38	0.58	10
4.62 ± 0.05	Volcanic	0.78	0.04	0.32	0.50	11
	Terrestrial	0.78	0.02	0.05	0.15	24
	Biogenic	0.55	0.05	0.53	0.74	8
10.14 ± 0.08	Volcanic	0.40	0.04	0.16	0.32	17
	Terrestrial	0.05	0.02	0.06	0.17	25
	Biogenic	0.70	0.05	0.60	0.79	7
12.26 ± 0.08	Volcanic	0.26	0.03	0.11	0.26	20
	Terrestrial	0.03	0.02	0.06	0.16	25
	Biogenic	0.61	0.05	0.55	0.76	7
10.84 ± 0.15	Volcanic	0.36	0.04	0.14	0.30	18
	Terrestrial	0.04	0.02	0.06	0.17	25
	Biogenic	0.67	0.05	0.58	0.78	7
11.81 ± 0.08	Volcanic	0.29	0.03	0.12	0.28	19
	Terrestrial	0.03	0.02	0.06	0.16	25
4.51 ± 0.09	Biogenic	0.15	0.05	0.38	0.57	10

	Volcanic	0.80	0.04	0.32	0.50	11
	Terrestrial	0.06	0.02	0.05	0.15	24
	Biogenic	0.17	0.05	0.37	0.57	10
4.40 ± 0.14	Volcanic	0.80	0.04	0.33	0.51	11
	Terrestrial	0.03	0.02	0.05	0.15	24
	Biogenic	0.18	0.05	0.38	0.58	10
4.61 ± 0.15	Volcanic	0.81	0.04	0.32	0.50	11
	Terrestrial	0.01	0.02	0.05	0.15	24

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946 **Table S5. Bayesian statistical source apportionment results for NGRIP ice core sulfate samples.**

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$\delta^{34}\text{S}$ (‰)	Source	Source Fraction	SD	95% CI Lower	95% CI Upper	CV (%)
6.75 ± 0.12	Biogenic	0.08	0.02	0.03	0.14	31
	Volcanic	0.67	0.05	0.56	0.77	7
	Terrestrial	0.24	0.04	0.17	0.33	17
7.23 ± 0.13	Biogenic	0.09	0.02	0.04	0.15	31
	Volcanic	0.62	0.05	0.52	0.72	8
	Terrestrial	0.28	0.04	0.20	0.38	15
6.14 ± 0.02	Biogenic	0.07	0.02	0.03	0.12	33
	Volcanic	0.71	0.04	0.61	0.80	6
	Terrestrial	0.21	0.03	0.14	0.30	17
6.65 ± 0.08	Biogenic	0.07	0.02	0.03	0.13	32
	Volcanic	0.68	0.05	0.57	0.77	7
	Terrestrial	0.24	0.04	0.16	0.33	16
6.46 ± 0.11	Biogenic	0.07	0.02	0.03	0.13	32
	Volcanic	0.69	0.04	0.59	0.79	7
	Terrestrial	0.23	0.04	0.15	0.31	17
7.00 ± 0.03	Biogenic	0.08	0.02	0.04	0.15	31
	Volcanic	0.64	0.05	0.54	0.74	8
	Terrestrial	0.26	0.04	0.19	0.36	16
4.61 ± 0.07	Biogenic	0.04	0.01	0.01	0.08	39
	Volcanic	0.87	0.03	0.79	0.92	3
	Terrestrial	0.08	0.02	0.05	0.13	23
6.27 ± 0.06	Biogenic	0.07	0.02	0.03	0.13	32
	Volcanic	0.70	0.04	0.60	0.79	6

	Terrestrial	0.22	0.03	0.15	0.31	17
	Biogenic	0.08	0.02	0.03	0.14	32
6.75 ± 0.14	Volcanic	0.67	0.05	0.56	0.77	7
	Terrestrial	0.24	0.04	0.17	0.33	17
	Biogenic	0.07	0.02	0.03	0.14	31
6.70 ± 0.05	Volcanic	0.67	0.05	0.57	0.77	7
	Terrestrial	0.24	0.04	0.16	0.33	16
	Biogenic	0.07	0.02	0.03	0.13	33
6.48 ± 0.28	Volcanic	0.69	0.05	0.58	0.78	7
	Terrestrial	0.23	0.04	0.15	0.31	17
	Biogenic	0.08	0.02	0.04	0.15	31
6.99 ± 0.05	Volcanic	0.64	0.05	0.54	0.74	8
	Terrestrial	0.26	0.04	0.19	0.36	16
	Biogenic	0.08	0.02	0.04	0.15	30
7.12 ± 0.07	Volcanic	0.63	0.05	0.52	0.73	8
	Terrestrial	0.27	0.04	0.20	0.37	16
	Biogenic	0.07	0.02	0.03	0.14	32
6.71 ± 0.17	Volcanic	0.67	0.05	0.56	0.77	7
	Terrestrial	0.24	0.04	0.16	0.33	17
	Biogenic	0.07	0.02	0.03	0.14	31
6.68 ± 0.15	Volcanic	0.67	0.05	0.57	0.77	7
	Terrestrial	0.24	0.04	0.16	0.33	17
	Biogenic	0.07	0.02	0.03	0.12	33
6.16 ± 0.07	Volcanic	0.71	0.04	0.60	0.80	6
	Terrestrial	0.21	0.03	0.14	0.30	17
	Biogenic	0.05	0.02	0.02	0.10	34
5.59 ± 0.07	Volcanic	0.78	0.04	0.69	0.85	5

Terrestrial	0.16	0.02	0.10	0.21	17
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