

1 **Is the Earth Lazy? A review of work minimization in fault evolution**

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10 **Keywords:** work minimization; fault propagation; energy budget; numerical model

11

11 **Abstract**

12 The principle of work minimization has been used in various forms to account for the
13 development of active fault systems within a wide range of tectonic settings. We review the
14 successes, challenges and implications learned from previous applications of work
15 minimization. Examination of the energy budget provides insight into the competing
16 influences of different processes within fault systems at a variety of scales. We present
17 each work budget component with considerations for numerical implementation.
18 Additionally, we demonstrate numerical implementation by solving for the energy budgets
19 and finding the most efficient propagation paths of a joint and several faults. Consideration
20 of the energy budget captures growth patterns that are consistent with both laboratory
21 observations and theoretical predictions using the energy release rate, G . Unlike using G ,
22 work minimization is not limited to collinear growth. In comparison with predictions using
23 Coulomb failure planes, work minimization eliminates uncertainty in predicting fault
24 propagation by (1) evaluating the trade-off between tensile and shear failure ahead of the
25 fault and (2) avoiding the ambiguity introduced by the two potential Coulomb failure
26 planes. Application of work minimization to faulting demonstrates the effectiveness of this
27 approach in predicting both the orientation and timing of fault propagation.

28 **1. Introduction**

29 Crustal deformation results from complex plate motions acting on a variety of time scales,
30 from the seconds required for earthquake ruptures to the millions of years required for
31 mountain building. Brittle failure plays a large role in this deformation, as faults slip and
32 grow to accommodate tectonic loading and dykes propagate due to internal magma
33 pressure, for example. Predicting the growth of brittle structures is critical to unraveling
34 many geophysical processes and is particularly important for understanding fault
35 evolution.

36 Inglis (1913) provided fundamental insights into the role of concentrated stresses
37 around flaws, and their role in flaw growth and the subsequent failure of the material.
38 Using these insights, the growth of fractures can be predicted if the stress concentration
39 factor at the fracture tip meets the strength of the material. This concept is fundamental to

40 theories of linear elastic fracture mechanics and has been successful in predicting
41 conditions for and paths of opening mode propagation that is characteristic of joints, veins
42 and dikes in rock materials (e.g. Pollard and Aydin, 1988). Analysis of the propagation of
43 faults using this approach has been more problematic. The challenges arise from the
44 distributed nature of damage preceding fault growth and the subsequent linkage of cracks
45 to form fault surfaces. The deformational processes around a fault tip do not match the in-
46 line growth pattern required for using stress concentration factors to predict the potential
47 for fault growth. In addition, while seismologists have been successful in using a local
48 energy balance at the fracture tip to model earthquake dynamics, strength heterogeneities
49 in the Earth make application of this approach over large areas and long time-scales
50 particularly difficult.

51 In contrast, work minimization provides a global approach to predicting failure by
52 postulating that the crust deforms in order to minimize the external tectonic work acting
53 on it. This deformation includes the propagation of faults and work minimization suggests
54 that fault propagation will occur in the direction that minimizes external energy. As a
55 consequence, a fault system may evolve over geologic time to efficiently accommodate
56 forcing from plate motions. A work minimization approach does not ignore fracture
57 mechanics, but rather drives mechanics based on a global energy budget, rather than solely
58 on a criterion local to the fault or fracture tip. Furthermore, fault growth along a
59 mechanically efficient path can still require that local conditions for failure are met at a
60 fault tip, though the timing and orientation of growth are based on the global energy
61 budget.

62 While the principle of work minimization has been applied in various forms in the
63 past, now is an ideal time to apply it to research on crustal deformation in general and fault
64 behavior in particular. Current computing capabilities allow fault mechanics to be
65 considered and the full energy budget of deformation to be determined. This eliminates
66 resurgent concerns regarding minimization approaches, such as that strain minimization
67 fails to capture the “complete energetics of the system” (Bird and Yuen, 1979).
68 Computational advances also allow for better constraints on and quantification of
69 individual work budget components, including the energy required in work against friction,

70 energy required for propagation, energy converted into seismic waves, work against
71 gravity, and the internal strain energy of the deforming system.

72 In this review paper, we present the individual components of the energy budget
73 along with guides for numerical implementation of the calculations. Different approaches
74 that previous investigators have taken to predict fault evolution using the principle of work
75 minimization are outlined, as well as the historical debates about the appropriateness of
76 work minimization in the study of fault network evolution. To demonstrate the benefits of
77 analyzing the complete work budget, we apply work minimization to the growth of a joint
78 and a fault, and contrast their energy budgets. The paper concludes with discussion of the
79 insights provided by using a work minimization approach to investigate tectonic
80 deformation in general and fault system evolution in particular, as well as on-going
81 considerations for doing so and current challenges.

82 **2. Energy budget for fault evolution**

83 Prior to reviewing the contributions of previous investigations of fault evolution using
84 work minimization, we outline the individual components of the energy budget. Over the
85 years, various workers have used slightly different definitions of these energy terms and
86 utilized different subsets of the complete energy budget. Here we present a set of
87 consistent formulations for a complete energy budget associated with fault growth. This
88 budget includes internal work of deformation around faults, W_{int} , work against gravity,
89 W_{grav} , work against friction along faults, W_{fric} , energy to create fault surfaces, W_{prop} , and
90 energy of ground shaking, W_{seis} . The sum of these terms equals the total work of the system,
91 which is the external work, W_{ext} , along the boundaries of the fault system (Fig. 1):

$$92 \quad W_{ext} = W_{int} + W_{grav} + W_{fric} + W_{prop} + W_{seis} \quad \text{Eq. 1}$$

93 The internal work and work against gravity often are considered together as mechanical
94 work. Within fault systems, gravitational force plays a different role within contractional
95 and extensional systems (Dempsey et al., 2012), so separating these two work components
96 provides relevant insights. All work terms except for the energy consumed by creating
97 fault surfaces (W_{prop}) can be derived from the product of a force and a displacement within
98 the fault system. In addition to describing the theoretical underpinning of each component

99 of the work budget, we note special considerations for the numerical calculation of each
100 component.

101 2.1 Internal work of deformation

102 Internal work is the work of deformation within the fault system. W_{int} can be
103 measured as strain energy density, which is the product of stress and strain (e.g.
104 Timoshenko and Goodier, 1951; Jaeger et al., 2007). The integral of the strain energy
105 density within the fault system provides the internal work of the system:

$$106 \quad W_{int} = \iiint \frac{1}{2} \sigma_{ij} \varepsilon_{ij} dV \quad \text{Eq. 2}$$

107 For a two-dimensional system in x-z space, Eq. 2 simplifies to:

$$108 \quad W_{int} = \frac{1}{2} \iint \sigma_{xx} \varepsilon_{xx} + \sigma_{zz} \varepsilon_{zz} + 2\sigma_{xz} \varepsilon_{xz} dx dz \quad \text{Eq.3}$$

109 Although the internal strain energy represents elastic recoverable strain, the internal work
110 term also incorporates energy available for consumption by inelastic processes. For
111 example, pervasive deformation mechanisms, such as calcite twinning, may reduce stresses
112 within deformed host rock between faults and subsequently consume internal work. The
113 role of distributed permanent deformation on the work budget has not yet been
114 investigated. The strain energy density around faults is often greatest near fault tips and
115 irregularities. Consequently, the internal work is a powerful control on fault propagation
116 and many studies have predicted fault growth using minimization of internal work (Melosh
117 and Williams, 1989; Du and Aydin, 1992; Du and Aydin, 1996; Okubo and Schulz, 2005).
118 W_{int} is reversible, meaning that it can either increase or decrease during fault system
119 evolution.

120 In the numerical implementation of W_{int} , the volume integral is calculated via
121 sampling points throughout the model. In the absence of faults, integrating a volume of
122 densely spaced sample points can provide accurate solutions for W_{int} ; however, the
123 discretization of the fault into elements introduces approximations to the stress field. Near
124 fault tips, where stresses vary a lot over short distances, these approximations can impede
125 accurate sampling of W_{int} . Of all of the work components, calculation of W_{int} has the greatest
126 errors. Additionally, we need to consider that the total W_{int} is sensitive to the size of the

127 fault system analyzed. For two models with identical fault length, but different model
128 dimensions, the larger model will have greater W_{int} . For this reason, the difference in W_{int}
129 before and after fault growth can be more informative than its total value.

130 2.2 Work against gravity

131 Gravitational work considers the upward displacement, d_z , against gravitational
132 force, g ,

$$133 \quad W_{grav} = \iiint \rho g d_z(z) dV \quad \text{Eq. 4}$$

134 where ρ is the density of the material. Like W_{int} , W_{grav} is reversible and can increase or
135 decrease during fault system evolution. Contractual fault systems will have an overall
136 upward displacement of points within the system that result in a positive W_{grav} while
137 extensional systems will have negative W_{grav} . Dempsey et al. (2012) show that a decrease
138 in W_{grav} can drive extensional faulting, even at the expense of an increase in W_{int} . This
139 complicates the elastic rebound paradigm of faulting, which was developed for strike slip
140 faults and assumes that slip events are associated with a decreased in stored elastic energy.
141 For all fault systems, decreases in the sum of W_{grav} and W_{int} , which is termed the
142 mechanical work, provide energy for fault slip, propagation and ground shaking.

143 The numerical calculation of W_{grav} requires sampling the displacement field
144 throughout the model. Unlike W_{int} , the displacements are not singular near the fault so
145 calculations of W_{grav} are not as sensitive to numerical discretization of the fault as
146 calculations of W_{int} . Both mechanical work terms are sensitive to the size of the model
147 domain.

148 2.3 Work against frictional resistance

149 The frictional work is the energy required to slide fault surfaces past each other.
150 W_{fric} is irreversible, because the heat created during frictional slip exits the system. The
151 onset of sliding involves processes related to the work required to generate new fault
152 surfaces and W_{fric} (Fig. 2A). The component attributed to work against frictional resistance
153 is calculated from the strength of the fault at stable sliding and the total slip, s . For a single

154 fault segment using a tension positive, compression negative sign convention, this
155 component is calculated as:

$$156 \quad W_{fric} = (c - \sigma_n \mu_d) sA = |\tau| sA \quad \text{Eq. 5}$$

157 where c is cohesion, σ_n is normal stress, μ_d is the dynamic friction coefficient during sliding,
158 τ is the shear strength during sliding, and A is the slipped area of the fault. When positive,
159 σ_n is tensile along the fault and allows opening, so W_{fric} is zero. The complete frictional
160 work for one increment of loading is integrated over the surface area, A , of the fault so that:

$$161 \quad W_{fric} = \iint [c - \sigma_n \mu_d] dA = \iint \tau dA \quad \text{Eq. 6}$$

162 Frictional work can be calculated as fault strength evolves over a loading increment (e.g.
163 Savage and Cooke, 2010); however, most quasi-static studies use the stress state at the end
164 of the model convergence to calculate W_{fric} (e.g. Hardy et al., 1998; Burbridge and Braun,
165 2002; Cooke and Murphy, 2004; Del Castello and Cooke, 2008). Because frictional work is
166 inelastic, numerical implementations often load the faults system with multiple monotonic
167 steps so that W_{fric} is calculated for each loading step and integrated over the applied
168 loading (e.g. Cooke and Murphy, 2004; Del Castello and Cooke, 2008)(Fig. 2B).

169 Observations generally support the inference that mature faults have lower sliding
170 friction than immature faults (e.g. Marone, 1998). Nevertheless, the formulation of W_{fric} is
171 the same for the development of new fault surfaces and slip along mature fault surfaces. In
172 addition, a new fault surface with high strength and low slip may produce identical
173 frictional heating to a mature fault surface with low strength and high slip.

174 Until recently, we have not been able to measure the frictional heat produced by
175 fault slip events. On-going projects that drill and sample active fault zones (e.g. Townend et
176 al., 2009) and particularly efforts to drill soon after earthquake events (e.g. Fulton et al.,
177 2013) are providing constraints on frictional heating so that we can validate the theoretical
178 formulations. Studies of heating along exhumed faults are further constraining values of
179 W_{fric} (Savage et al., 2014).

180 *2.4 Seismic radiated energy*

181 The strength of a fault evolves during slip, from the initial strength (τ_0) to the
182 weakened sliding strength (τ) (Fig. 2A). This change in shear stress with slip produces

183 energy that is available for both fault propagation, W_{prop} , and ground shaking, W_{seis} , which
 184 are irreversible and lost to the fault system. The portion of work allocated to each process
 185 depends on the amount of fault slip relative to the slip required to bring the fault to its
 186 sliding strength, referred to as the slip weakening distance, L (Fig. 2). All energy associated
 187 with the shear stress drop ($\Delta\tau = \tau_0 - \tau$) when s exceeds L is available for ground shaking. In
 188 other words, the energy required to generate the new fault surface is expended while $s < L$,
 189 whereas work associated with $s > L$ is available for ground shaking. If $L = 0$, W_{seis} for one
 190 increment of loading is:

$$191 \quad W_{seis} = \iint \int \Delta\tau ds dA \quad \text{Eq. 7}$$

192 Numerical models typically do not consider the evolution of shear stress with slip and only
 193 calculate the change in shear stress, $\Delta\tau$, from comparing model results before and after
 194 fault growth. In this case, the work need not be integrated over slip and Eq. 7 simplifies to:

$$195 \quad W_{seis} = \iint \frac{1}{2} \Delta\tau s dA \quad \text{Eq. 8}$$

196 In the event of multiple loading increments, W_{seis} is calculated from the change in shear
 197 stress and the change in slip from one loading step to the next (Fig. 2B). The calculation of
 198 the W_{seis} is not largely affected by discretization.

199 The formulation of W_{seis} is the same for both established and new fault surfaces, but
 200 the values of τ and τ_0 may differ. For example, intact rock often requires greater shear
 201 stress to reach failure than an established fault surface. While quasi-static models cannot
 202 be used to calculate shaking directly, they can be used to examine $\Delta\tau$ and estimate the
 203 energy available for shaking.

204 *2.5 Energy of fault propagation/growth*

205 The energy of fault propagation, W_{prop} , is the energy required to create
 206 discontinuities within a material. This energy is irreversible within most natural
 207 conditions. Although Obriemoff (1930) was able to reverse the growth of a crack in mica by
 208 growing the crack within a vacuum, under atmospheric conditions, available cations
 209 bonded to new crack surfaces and impeded healing.

210 Griffith (1921) demonstrated that the energy release rate, G , required to grow a
 211 crack can be calculated directly from the stress intensity factors on the crack, K_i , where i
 212 refers to growth by modes I, II or III:

$$213 \quad G = \frac{1 - \nu^2}{E} \left[K_I^2 + K_{II}^2 + \frac{K_{III}^2}{(1 - \nu)} \right] \quad \text{Eq. 9}$$

214 Here, ν is Poisson's ratio and E is Young's Modulus. However, this formulation presumes
 215 co-linear crack propagation of a single discontinuity, so Eq. (8) works well for opening-
 216 mode propagation of joints and veins in rock, and is applicable for controlled experiments
 217 that result in co-linear growth. However, this formulation does not apply universally to
 218 faults, which often propagate out of plane. Consequently, predicting fault growth using G is
 219 problematic.

220 Taking an observational approach, some workers have estimated W_{prop} from
 221 measurements of damage around faults. Using the measured energy release rate of
 222 opening mode cracks ($G \cong 1 \text{ J/m}^2$), the total energy consumed in fault growth can be
 223 estimated by summing the damage area around faults (e.g. Wong, 1982, 1986; Cox &
 224 Scholz, 1988). For cases of new faulting within intact rock, these values are 10^5 - 10^6 J/m^2
 225 (Wilson et al., 2005; Pittarello et al., 2008). In contrast, values from laboratory samples
 226 range only up to 10^4 J/m^2 (Wong, 1982, 1986; Cox & Scholz, 1988; Lockner et al., 1992).
 227 These different ranges from laboratory and field measurements highlight the need to
 228 develop a better understanding of W_{prop} . In section 4, we present numerical models that
 229 simulate laboratory experiments of faults loaded at values critical for growth. In addition
 230 to demonstrating the work budget of fault growth, these experiments allow for calculation
 231 of W_{prop} . For co-linear propagation, this value can be validated against calculations of G .

232 A formula for W_{prop} that does not require co-linear growth is related to that for W_{seis} .
 233 As mentioned in the previous section, the portion of work allocated to each of these
 234 processes depends on the amount of fault slip, s , relative to the slip weakening distance, L
 235 (Fig. 2A). The energy required to generate the new fault surface is expended while $s < L$,
 236 whereas work associated with $s > L$ is available for ground shaking. For loading steps
 237 where the total slip remains less than L , this is:

$$238 \quad W_{prop} = \iint \frac{1}{2} \Delta\tau \Delta s dA \quad \text{Eq. 10}$$

239 where Δs is the change in slip associated with that loading step and W_{seis} equals zero. For
 240 loading steps that produce slip that spans the slip weakening distance, such as step 3 on
 241 Fig. 2B, this is:

$$242 \quad W_{prop} = \iint \frac{1}{2} \Delta \tau s_L dA \quad \text{Eq. 11a}$$

$$243 \quad W_{seis} = \iint \frac{1}{2} \Delta \tau (\Delta s - s_L) dA \quad \text{Eq. 11b}$$

244 Here, s_L represents slip during this loading step when $s < L$. The total W_{prop} and W_{seis} for the
 245 system are summed at each loading step.

246 2.6 External work

247 The external work reflects the overall mechanical efficiency of the fault system, such
 248 that an efficient system will require less W_{ext} to accommodate the same tectonic
 249 displacement than an inefficient system. W_{ext} is the product of the shear and normal
 250 tractions, τ and σ_n , and shear and normal displacements, u_s and u_n , along the model
 251 boundaries, B (Fig. 1):

$$252 \quad W_{ext} = \iint (\tau u_s + \sigma_n u_n) dB \quad \text{Eq. 12}$$

253 If multiple loading steps are applied, the calculation of W_{ext} should be integrated over the
 254 loading steps. The value of W_{ext} also equals the sum of the energy budget components in Eq.
 255 (1). This relationship provides an independent check on the calculations of individual
 256 energy terms within the budget. Considerations of fault system evolution benefit from
 257 attention to changes in external work, ΔW_{ext} , between stages of fault growth. To meet
 258 requirements of energy conservation, we can consider that ΔW_{ext} equals the sum of the
 259 associated changes in all work budget components.

260 This energy budget differs from the energy considerations single earthquakes,
 261 where it is assumed that no change in remote loading or strain occurs during the short
 262 time-span of fault slip (e.g. Scholz, 2002). Under this assumption, which follows that
 263 presented by Griffith (1921), the change in mechanical work, $\Delta(W_{int} + W_{grav})$, equals the
 264 sum of the changes in W_{fric} , W_{prop} and W_{seis} . During an earthquake, the decrease in
 265 mechanical work provides the energy that is consumed in frictional heating, generation of
 266 damage, and ground shaking. The assumption that $\Delta W_{ext} = 0$ is not appropriate for studies

267 of fault evolution, however, which consider fault development over multiple earthquake
268 cycles due to tectonic loading. Consequently, we consider ΔW_{ext} as the fault system evolves,
269 following the work of Del Castello and Cooke (2007), which shows that a fault system
270 prefers to propagate in the direction that maximizes ΔW_{ext} .

271 The approach of analyzing the increasing efficiency of evolving fault systems in
272 numerical models depends on the method of applied deformation, i.e. applied tractions or
273 displacements. For fault systems that are loaded with displacement along the model
274 boundaries, fault growth will increase the compliance of the system, so that the
275 corresponding tractions along the model boundaries decrease. In this case, W_{ext} will
276 decrease with the growth of more efficient faults within the system. In contrast, if the
277 system is loaded with tractions along the model boundaries, then the increase in
278 compliance due to fault growth will increase the displacement of the model boundaries. In
279 this case, W_{ext} will increase with the growth of more efficient faults within the system.
280 Either displacement or traction loading is suitable for a work budget analysis, but we
281 advise against using mixed non-zero boundary conditions. In the cases of both
282 displacement only and traction only loading, the most efficient fault growth path is that
283 which maximizes the change in external work, ΔW_{ext} , from before to after fault growth. We
284 demonstrate both types of loading in section 4, but continue to refer to growth that
285 optimizes energy efficiency as “work minimization”.

286 **3. Work minimization in structural geology**

287 For over a century, engineers have used energetic principles to understand material
288 deformation and failure. Griffith (1921, 1924) used the energy budget associated with
289 opening mode fracture growth to determine the equations governing fracture propagation,
290 which launched the field of fracture mechanics. In an interesting revisitation of this classic
291 approach, Engelder and Fisher (1996) execute Griffith’s conceptual model within the
292 laboratory and discuss implications for the loading of joints in the crust. Here, we review
293 past applications of energy minimization to assess fault growth in detail. These studies
294 present many interesting insights, as well as challenges that remain to be resolved.

295 Though we do not review them in detail, work minimization approaches have been

296 used in other geoscience disciplines as well. For example, within metamorphic petrology,
297 the numerical code ELLE predicts microstructure development based on work
298 minimization of forces and chemical reactions within the system (e.g. Jessell et al, 2001).
299 Modern fluvial geomorphology is founded on the work of Langbein and Leopold (1964),
300 which analyzes the energy budget of evolving river channels. Another formulation of work
301 minimization is maximum entropy. Rather than assessing the energy available within the
302 system, the maximum entropy approach assesses the irreversible energy that is no longer
303 available to the system. An example of this application is atmospheric circulation models
304 that utilize maximum entropy to predict flow (e.g. Kleidon et al., 2003).

305

306 *3.1 Minimum viscous dissipation*

307 For several decades, the minimization of work has been utilized to predict the
308 development of fault systems. The earliest approaches minimized the viscous dissipation
309 within ocean ridge and transform fault systems. In this application, analytical solutions
310 characterizing the system were solved to find the fault network that uses the least force to
311 accommodate the prescribed plate boundary displacements via viscous flow within the
312 fault zones. Lachenbruch and Thompson (1972) showed that minimum viscous dissipation
313 predicts the orthogonal orientation of ridges and transform faults. This study was followed
314 by several others that augmented this analysis to successfully account for various
315 observations of ocean ridge geometry (Fig. 3) (e.g. Froidevaux, 1973; Stein, 1978; Kleinrock
316 and Morgan, 1988). In 1979, Bird and Yuen published a critique of the approximations used
317 in the analytical implementation of the minimum viscous dissipation approach. They
318 pointed out that the approximations are avoided with numerical implementations that
319 directly solve the governing equations and consider more complex rheology and coupled
320 processes within the crust (e.g. Regenauer-Lieb et al., 2006; Holtzman et al., 2005). This
321 critique was followed by a comment in support of the original methodology by Sleep et al.
322 (1979).

323 A similar approach to minimizing viscous shear is minimizing on-fault shear stress,
324 which has been used to predict the orientation of frictionally sliding faults. Reches (1983)

325 implemented a minimum dissipation approach for frictionally sliding faults, which showed
326 that sets of orthorhombic faults require the least shear stresses to accommodate three-
327 dimensional states of strain. In this same study, Reches (1983) demonstrated that the fault
328 networks that minimize shear stress also minimize the internal work in the rock between
329 the sliding surfaces.

330 *3.2. Minimization of internal work*

331 More efficient fault systems allow for more fault slip, so that the surrounding rock is
332 relatively less strained and subject to less stress. Consequently, the minimization of
333 internal work, which integrates the product of the stress and strain fields Eq. (2), can
334 achieve similar predictions of fault geometry as methods that minimize the forces required
335 to accommodate prescribed deformation. Melosh and Williams (1989) demonstrated that
336 antithetic pairs of normal faults result in lower internal work than synthetic pairs of
337 normal faults. The greater efficiency of antithetic normal faults may account for the
338 frequent occurrence of antithetic normal faults within grabens (Melosh and Williams,
339 1989). Hexagonal cracks within uniformly expanding materials, such as mud cracks, are
340 another common observation that has been explained by the minimization of strain energy
341 per unit of new crack area (Lachenbruch, 1962).

342 Within elastic systems, the spatial variations of internal work throughout a fault
343 system can be mapped with the Strain Energy Density, SED, so that SED is the internal
344 work at a point within the system. The distribution of SED has been used to assess the
345 growth and propagation faults. This approach utilizes the premise that faults develop in
346 regions of high stress and strain, so that the growth of faults should minimize the total SED,
347 i.e. the internal work (e.g. Du and Aydin, 1993; Du and Aydin, 1996; Okubo and Schulz,
348 2005; Olson and Cooke, 2005). Following this approach, Olson and Cooke (2005) used
349 three-dimensional models to compare fault orientations predicted by the locations of high
350 SED with those predicted by Coulomb failure theory, and showed that locales of high SED
351 better predict the interpreted evolution of imbricate thrust faults within the Los Angeles
352 basin (Fig. 4). Griffith and Cooke (2004) used SED patterns around alternative three-
353 dimensional configurations for the intersection of the Whittier and Puente Hills faults to

354 show that the configuration that minimizes internal work also reproduces the observed
355 slip rate.

356 Further insights can be gained by decomposing SED into its two deformational
357 components, the dilatational and distortional strain energy densities (e.g. Jaeger et al.,
358 2007). Because faults accommodate distortion, the pattern of faulting may be predicted by
359 the maximization of distortional SED (e.g. Du and Aydin, 1992; Du and Aydin, 1996; Okubo
360 and Schulz, 2005). Limiting analysis to the distortional SED is akin to using a maximum
361 shear stress or second invariant of shear stress failure criterion, such as von Mises's
362 criterion. While this approach does not consider the role of the normal stress, as Coulomb
363 failure does, the predicted fault propagation patterns resemble general observations of
364 fault step-overs (Du and Aydin, 1993) and the propagation path of deformation bands
365 (Okubo and Schulz, 2005), as well as the development of the Eastern California Shear Zone
366 north of a bend in the San Andreas fault in California, USA (Du and Aydin, 1996),.
367 Furthermore, Okubo and Schulz (2005) used dilatational SED to predict fault initiation
368 locations, which may be related to dilatational failure, in contrast to distortion dependent
369 propagation processes. These studies demonstrate that faults grow into regions of high
370 internal work, which consequently can reduce the W_{int} of the system.

371 Another application is in crustal models, where minimum internal work is used to
372 interpolate deformation in regions where geologic data are scarce (e.g. Saucier and
373 Humphries, 1993; Peltzer and Saucier; 1996). These models use geodetic velocities and
374 geologic slip rates as input where they are available. Minimizing internal work provides a
375 way to resolve slip along portions of the faults, while satisfying data constraints. The
376 widespread use of this approach supports the premise that faults develop to minimize
377 internal work.

378 A corollary to the minimization of internal work is the maximization of slip. The
379 theory here is that the fault system with the least internal work will display the largest
380 amount of fault slip. DeBremaecker and Ferris (2004) show that maximum slip predicts the
381 orientation of wing-crack development at fault tips. Using analog experiments of
382 restraining bends, Cooke et al. (2013) showed that the percentage of slip accommodated by
383 fault systems increases with the propagation of new faults as the system becomes more

384 efficient. This analog modeling study further supports the premise that faults evolve to
385 greater efficiency by maximizing fault slip and minimizing internal work.

386 *3.3 Minimization of strain*

387 The calculation of either internal work or viscous dissipation requires solving for
388 both the displacement (or velocity) and the stress (or stress rate) fields. An analytically
389 more simple and computationally faster approach is to minimize only the strains or
390 displacements. By assuming that locations of high stress are also locations of high strain,
391 minimizing strain should have the same effect as minimizing internal work. The
392 computational benefit is that the kinematic solutions can be used to find the strain field,
393 without having to solve for stress. This approach has been used to assess the ratios of pure
394 shear to simple shear during non-steady state deformation (Fossen and Tikoff, 1997).
395 Following publication of that paper, Jiang (1998) criticized the minimum strain path
396 approach, because it did not consider the full energetics of the system. In their reply,
397 Fossen and Tikoff (1998) note that, due to the complexities of the system, it is not known if
398 minimizing work gives a different result from the kinematic solutions found by minimizing
399 strain. In many ways, this discussion and reply echo the discussion two decades prior
400 between Bird and Yuen (1979) and Sleep et al. (1979).

401 Minimization of strain also is used in several applications to determine fault slip,
402 when the system is under-constrained. Strain minimization can be used in a similar way to
403 the minimization of internal work, to resolve unconstrained fault slip rates within crustal
404 deformation models (e.g. Fleck et al., 2001). Similarly, several three-dimensional structural
405 restoration algorithms minimize strain to determine slip along faults within a system (e.g.
406 Plesch et al., 2007; Shackleton et al., 2009).

407 *3.4 Limit analysis*

408 Limit analysis provides an alternative approach for finding the most efficient failure
409 surface using analytical solutions. Like other methods, it draws from engineering
410 mechanics and has a long history of success. For example, Bishop (1955) developed a
411 standard method to investigate slope stability by delineating the most likely landslide

412 surface as that which fails at the least applied stress. Similar methodology has been used to
413 find the fault surface within an accretionary wedge that fails under the lowest force (e.g.
414 Maillot and Leroy, 2003; Maillot and Leroy, 2006; Cubas et al., 2008; Souloumiac et al.,
415 2009; Souloumiac et al., 2010; Mary et al., 2013). The limit analysis requires analysis of
416 many potential surfaces to find the one that requires the minimum force to deform. This
417 surface is at its failure limit, so it requires the 'least upper bound of applied force', a phrase
418 often used in this literature. Lesser force can be applied, but the upper bound of force is
419 that required for failure of the system, i.e. the development of faults. One of the great
420 benefits of the limit analysis approach is that it is analytical, so that the search for the
421 preferred surface completes within seconds. Consequently, limit analysis can be used to
422 assess sequential failure of multiple surfaces and shows many of the same features as
423 accretionary wedges (Fig. 5) (e.g. Cubas et al., 2008, Mary et al., 2013).

424 *3.5. Minimization of total and external work*

425 Many workers have recognized that internal work does not capture the complete
426 system and so have sought to find the fault geometry that minimizes the system's total
427 work. This total work represents different combinations of work budget components in
428 different applications, including: 1) work against gravity, which is sometimes integrated
429 with internal work, 2) work against friction, 3) energy required to propagate faults, and 4)
430 seismic radiated energy (Fig. 1) (e.g. Cooke and Murphy, 2004). Jones and Wesnousky
431 (1992) considered internal work and work against friction to demonstrate that under
432 transpressional loading, slip partitioning between two parallel fault planes, one with dip
433 slip and another with strike slip, minimizes the total work of the deforming system. Cooke
434 and Kameda (2002) showed that, for a variety of proposed cross-sections of the Los
435 Angeles basin, the system that requires the least total work to accommodate tectonic
436 loading is that which best matches the available fault slip rates. In a study of the Lake
437 Meade Fault zone, Marshall et al. (2010) assessed the relative efficiency of different fault
438 interpretations. The configuration that minimizes work is consistent with independent
439 observations and provides inside into interaction among faults. These studies lend support
440 to the postulation that fault systems evolve to minimize total work.

441 By far, the bulk of investigations of fault evolution using work minimization with a
442 full work budget analysis examine accretionary wedge systems. Both analog models and
443 crustal wedges provide excellent laboratories for studying fault growth, because the
444 material is relatively undeformed prior to incorporation into the wedge, while within many
445 other deformational regimes, the rock mass may have inherited fault structures that
446 complicate a minimum work analysis. Dahlen et al. (1984) is a watershed study that
447 introduced the concept of critically tapered wedges. A few years later, Mitra and Boyer
448 (1986) framed the development of duplexes, which can be viewed as the cellular unit of an
449 accretionary wedge, in terms of the system's work budget. Dahlen et al. (1989), followed up
450 with a work budget for the entire accretionary wedge that has been augmented by many
451 workers (e.g. Gutcher et al., 1998; Hardy et al., 1998; Burbidge and Braun, 2002).

452 Several studies assess fault propagation within accretionary systems through
453 minimization of the energy required to overcome friction and gravity (Gutscher et al., 1998;
454 Hardy et al., 1998; Burbidge and Braun, 2002). Gutscher et al. (1998) showed analytically
455 that minimization of frictional and gravitational work accurately predicts the length of a
456 new accretionary forethrust. Burbidge and Braun (2002) extended this analysis to
457 determine the friction required for frontal and back accretion, while Hardy et al. (1998)
458 used a combined Eulerian-Lagrangian scheme within numerical models to minimize
459 frictional and gravitational work to predict the progressive sequence of accretion. Masek
460 and Duncan (1998) expanded these analyses by including internal work, showing that this
461 component has a large influence on fault geometry. As with previous applications of work
462 minimization, Masek and Duncan (1998) evoked a discussion and reply (DeBremaecker,
463 1999; Masek and Duncan, 1999). However, this discussion focuses on the effect of
464 numerical discretization on the work budget, rather than on the premise of a fault growing
465 to minimize work.

466 In a recent study of accretion, Del Castello and Cooke (2007) used a numerical
467 implementation of work minimization to simulate analog experiments of accretion and
468 present the evolving work budget during the process of new forethrust development (Fig.
469 6). In these finite models, the external work acting on the system represents the system's
470 total work, and equals the sum of the five work budget components. A new forethrust
471 reduces the external work required to deform the system due to a reduction in frictional

472 work, even though the new fault increases the system's internal work (Fig. 6B).
473 Additionally, the investigation of potential position and vergence of thrusts shows that the
474 thrust in the numerical model that minimizes W_{ext} matches the forethrust that develops in
475 the experimental sandbox. Del Castello and Cooke (2007) further found that propagation
476 of the forethrust occurs when the decrease in external work due to adding the fault (ΔW_{ext})
477 exceeds the work required to create the new fault surface ($W_{prop} + W_{seis}$; Fig. 6C). This
478 outcome highlights the critical role of the work of fault propagation, a work term not
479 previously considered. This study suggests that consideration of external work and the
480 work of fault propagation together allow for prediction of not only the path of fault
481 propagation, but also the timing of fault propagation.

482 **4. Work minimization predictions of fault propagation**

483 To demonstrate the power of using a work minimization approach to study fault
484 growth, we present the propagation paths and work budgets for co-linear growth of a joint
485 and faults both with and without slip weakening. Following investigations by Del Castello
486 and Cooke (2007), the most efficient growth paths are found by maximizing ΔW_{ext} , the
487 difference in W_{ext} before and after propagation. Potential growth elements, or *pupative*
488 *elements*, are radial to the fault tip at a prescribed range of angles to the fault. The growth
489 of the fault via each pupative element is assessed separately and comparison of the results
490 reveals the most efficient radial propagation path. The stresses along some pupative
491 elements will not meet the tensile or shear failure criterion, so the fault will not be able to
492 grow in these directions and W_{ext} remains unaltered (i.e. $\Delta W_{ext} = 0$). Among the pupative
493 elements that do fail, the element that maximizes ΔW_{ext} is added to the initial structure.
494 Though only one growth iteration is presented here, this sequence can be repeated for
495 subsequent pupative elements.

496 We simulate joint and fault propagation under the conditions of laboratory
497 experiments on fault growth by Bobet and Einstein (1998). Both the joint and the fault are
498 12.7 mm long and are comprised of 100 boundary elements of 0.127 mm embedded within
499 a slab of gypsum (Fig. 7A, 7B). The pupative elements have the same length as elements
500 along the pre-existing fault. The orientations of pupative elements are considered at 5°

501 angle increments between 45° and 315° clockwise from the joint or fault tip (Fig. 7B inset).
502 We use the material properties for gypsum from Bobet and Einstein (1998) with a water to
503 gypsum ratio of 0.4, tensile strength of 3.2 MPa, an average Poisson's ratio of 0.15, and an
504 average Young's modulus of 5.96 GPa.

505 We utilize the two-dimensional boundary element method (BEM) program Fric2D
506 (Cooke and Pollard, 1997). Dislocation surfaces made up of a series of elements of equal
507 length are free to open or slip, but not to interpenetrate, in response to the tractions or
508 displacement applied on model boundaries and interactions with other elements. Fric2D
509 solves the quasi-static equations of deformation on all elements. We take a compression
510 negative, tension positive sign convention. Elements fail in tension when the normal stress,
511 σ_n , exceeds the prescribed tensile strength, T :

$$512 \quad T \geq \sigma_n. \quad \text{Eq. 13}$$

513 Shear failure is governed by a frictional failure criterion, so that an element along an
514 established fault slips when the shear stress, τ , exceeds the frictional strength, which is the
515 sum of cohesion, c , and the product of the coefficient of static friction, μ_s , and σ_n :

$$516 \quad |\tau| \geq c - \mu_s \sigma_n \quad \text{Eq. 14}$$

517 Fric2D also captures slip-weakening behavior (Savage and Cooke, 2010). When slip on a
518 fault element exceeds a prescribed weakening distance, L , the friction evolves linearly from
519 its static value, μ_s , to a dynamic value, μ_d .

520 The criterion governing the shear failure of intact rock has the same form as Eq.
521 (14), but c is replaced by the rock's inherent shear strength, S_o , and μ_s is replaced by the
522 coefficient of internal friction, μ_o :

$$523 \quad |\tau_o| \geq S_o - \mu_o \sigma_n \quad \text{Eq. 15}$$

524 Following onset of failure in intact rock, laboratory tests have shown that fault strength
525 decreases as a fault surface develops during slip (e.g. Handin and Hager, 1957; Hazzard and
526 Young, 2000; Mitchell and Lockner, 2008). In Fric2D, when a pupative element fails, its
527 strength decreases so that shear stress drops from τ_o to τ as its internal friction value, μ_o ,
528 drops to μ_d . The inelastic behavior of frictional slip is solved iteratively until the model
529 converges (Cooke and Pollard, 1997).

530 4.1 Joint propagation

531 In the joint model, we apply horizontal and vertical displacement boundary
532 conditions, respectively u_h and u_v , on the edges of the gypsum block (Fig. 7A). Keeping
533 $u_h=0$, we adjust u_v to its critical value, when the stress intensity factor, K_I , at the joint tip
534 equals gypsum's mode I fracture toughness, K_{Ic} , of $0.15 \text{ MPa}\sqrt{\text{m}}$. This value of K_{Ic} value is
535 from the curve reported by Chen et al. (2006, Fig.2), extrapolated to the 0.4 water to
536 gypsum ratio used by Bobet and Einstein (1998). The failure criterion is met at $u_v=2.71e^{-5}$
537 m, which produces $\sim 3.1 \text{ MPa}$ of axial tension on the model boundaries.

538 Assessment of pupative elements around the joint tip reveals that the most efficient
539 orientation for growth, the orientation that minimizes ΔW_{ext} , is at 180° (Fig. 8a). This co-
540 linear growth is commonly observed for joints under perpendicular extension (e.g. Pollard
541 and Aydin, 1988). W_{ext} is 1.1135 J before growth and drops to 1.1131 J after growth,
542 resulting in ΔW_{ext} per pupative element area of -3.63 J/m^2 . Note that, though the model is
543 two-dimensional plane strain, all elements have a width of $w = 1 \text{ m}$ out of the plane under
544 consideration. The components of the joint's work budget are summarized in Table 1 and
545 Fig. 9. For this growth scenario, W_{seis} , W_{fric} and W_{grav} are 0, and $\Delta W_{int} = \Delta W_{ext} = W_{prop}$. Using
546 Eq. (10), we find that $W_{prop} = 3.62 \text{ J/m}^2$, which balances with ΔW_{ext} , indicating that the
547 change in external and internal work is a result of energy spent on propagating the joint.

548 Because the joint growth is co-linear, we can compare this estimate of W_{prop} to that
549 estimated by the critical energy release rate, G_c . Using $K_{Ic} = 0.15 \text{ MPa}\sqrt{\text{m}}$, we find from Eq.
550 (9) that $G_c = 3.71 \text{ J/m}^2$. Thus, $\Delta W_{ext} \cong W_{prop} \cong G_c$, demonstrating consistency between
551 opening-mode joint growth by work minimization and the theory outlined by Griffith
552 (1921) for co-linearly propagating cracks.

553 4.2 Fault Propagation (Fault 1)

554 For the more complex scenario of fault growth, we began with two widely spaced,
555 60° dipping faults in gypsum that demonstrate co-linear growth under biaxial loading
556 (scenario 60° -2a-4a from Bobet and Einstein (1998)). While Bobet and Einstein (1998)
557 suggest that faults with internal tips at least 3 half-lengths (19.05 mm) from one another
558 behave as isolated faults, the numerical simulations show some interaction between the

559 stress fields surrounding these faults, even at 28.40 mm of internal tip spacing. To remove
560 the effects of any interaction, we include only one fault in these simulations of fault growth
561 (Fig. 7B). While this geometry results in a slight overestimate of work budget components,
562 it provides a more robust demonstration of isolated fault growth patterns under biaxial
563 loading.

564 The numerical model for Fault 1 contains a pre-existing fault of 12.7 mm dipping
565 60° , with zero cohesion and equal static and sliding friction coefficients ($\mu_s = \mu_d = 0.3$)
566 (Bobet, 2000). The coefficient of internal friction on potential growth elements, μ_o , also is
567 0.3. The inherent shear strength, $S_o = 78.7$ MPa, for the growth elements is found by solving
568 Eq. (15) using the critical shear and normal stresses on the element oriented at 180° from
569 the fault tip at the onset of fault propagation (Fig. 7B) (Bobet, 1997, Table 4.2; Bobet and
570 Einstein, 1998). We apply traction boundary conditions on the horizontal and vertical
571 edges of the gypsum block of $\sigma_h = -2.5$ MPa and $\sigma_v = -29$ MPa, respectively (Bobet and
572 Einstein, 1998).

573 ΔW_{ext} is largest for the pupative growth element oriented at 180° from the fault tip
574 (Fig. 8b), consistent with the experimental observations of co-linear growth. The complete
575 work budget for Fault 1 is compared with that for the joint in Fig. 9 and Table 1. For growth
576 of 1 element length of 0.127 mm, W_{ext} increases from 878.5522 J to 878.5947 J. In contrast
577 to the joint model, which used displacement boundary conditions so that a decrease in W_{ext}
578 indicates an increase in efficiency, in this model an increase in W_{ext} indicates an increase in
579 efficiency (see discussion in Section 2.6). This results in a ΔW_{ext} per pupative element area
580 of 333.83 J/m². For this scenario, W_{seis} and W_{grav} are zero and $\Delta W_{ext} = \Delta W_{int} + \Delta W_{fric} = W_{prop}$
581 $+ \Delta W_{fric}$. Using Eq. (6) and Eq. (10), we find that $W_{fric} = 79.78$ J/m² and $W_{prop} = 260.77$ J/m².
582 The sum of W_{fric} and W_{prop} is 340.54 J/m², which balances with ΔW_{ext} , indicating that the
583 change in external work results from energy spent on fault propagation and the increase in
584 frictional sliding that fault growth allows.

585 As for the joint, this co-linear propagation allows for a comparison of the fault's
586 W_{prop} with that estimated by the critical energy release rate, G_c . Using the fracture
587 toughness reported by Fric2D at the fault tip at critical loading ($K_{Ic} = -1.27$ MPa \sqrt{m}), we

588 find that the critical energy release rate is $G_c = 263.1 \text{ J/m}^2$. Thus, G_c is consistent with the
589 work required for fault growth calculated using a work budget approach.

590 This model demonstrates the robustness of using a work minimization approach to
591 model fault growth. Because this approach considers the global ΔW_{ext} instead of the local
592 K_{IIc} , it can be applied to more complex fault propagation situations and is not restricted to
593 co-linear growth. The work minimization approach also provides an estimate of the energy
594 required for breaking intact rock to form new fault surfaces. In this case, we find that W_{prop}
595 for fault growth in gypsum is 260.77 J/m^2 . This value is reasonable, at two orders of
596 magnitude larger than W_{prop} for joint growth in gypsum and several orders of magnitude
597 smaller than the work inferred to grow faults in rock (10^4 - 10^6 J/m^2) (Wong, 1982, 1986;
598 Cox & Scholz, 1988; Lockner et al., 1992; Wilson et al., 2005; Pittarello et al., 2008).

599 *4.3 Propagation of slip-weakening faults (Fault 2, Fault 3)*

600 We consider slip-weakening along pupative growth elements in two additional
601 scenarios, Fault 2 and Fault 3. We prescribe $\mu_o = 0.6$ and $\mu_d = 0.3$ along both faults, but vary
602 the weakening distance, L , over which the drop in friction from μ_o to μ_d occurs. $L = 4 \times 10^{-6} \text{ m}$
603 for Fault 2 and $L = 1 \times 10^{-6} \text{ m}$ for Fault 3. The shear stress drop resulting from slip
604 weakening has the potential to produce seismic radiated energy, W_{seis} in addition to W_{prop}
605 (Fig. 2).

606 As shown in Table 1 and Fig. 9, varying L does not change the angle of the pupative
607 element that maximizes efficiency, which remains at 180° for both Fault 2 and Fault 3, nor
608 does it change ΔW_{int} or ΔW_{fric} . However, L does affect the partitioning of work between fault
609 propagation, W_{prop} , and the production of seismic waves, W_{seis} (Fig. 2). Shorter L (Fault 3)
610 means that less energy is required to break the intact rock ahead of the fault, leaving more
611 energy for ground shaking.

612 *4.4 Comparison of work minimization and the Coulomb criterion*

613 Correspondence between the maximum mechanical efficiency prediction of a fault's
614 propagation path with that observed in laboratory experiments supports the use of work
615 minimization as a fault growth criterion. This implementation of the principle of work

616 minimization using numerical models utilizes Coulomb theory as a failure criterion for both
617 the potential growth elements extending from the fault tip and those along the pre-existing
618 fault surface. However, this approach differs from Coulomb stress analyses that determine
619 a fault's propagation path from planes of maximum Coulomb shear stress out ahead of a
620 fault tip. We compare the two approaches here.

621 The Coulomb shear stress is found on a potential failure surface by rearranging Eq.
622 (15) to isolate S_o , replacing it with the Coulomb stress, σ_c , and replacing the frictional shear
623 strength, τ_o , with the shear stress resolved on the surface, τ :

$$624 \quad \sigma_c = |\tau| + \mu_o \sigma_n \quad \text{Eq. 16}$$

625 Positive σ_c along a surface indicates that it has the potential for failure, while negative σ_c
626 indicates that failure is inhibited. σ_c is useful for finding the relative failure potentials of
627 different fault surfaces.

628 The potential failure planes that maximize σ_c are at angles of $\pm\varphi$ to the maximum
629 compressive stress, σ_3 . φ depends upon the inherent friction coefficient, μ_o :

$$630 \quad \pm\varphi = \frac{1}{2} \text{atan}(1/\mu_o) \quad \text{Eq. 17}$$

631 The \pm indicates that Coulomb theory predicts two potential planes, on either side of σ_3 , that
632 carry the same shear stress magnitude and therefore the same failure potential. Selecting
633 between these planes is not trivial.

634 For this analysis, we consider both Coulomb and tensile failure near the fault tip. Fig.
635 10A shows the orientations of planes carrying the maximum tension and Coulomb stress, .
636 Figs. 10B and 10C display the contours of the maximum tensile stress and Coulomb stress,
637 respectively, following fault slip and prior to fault propagation. In Fig. 10B, tensile stress is
638 largest below and to the right of the fault tip, within the tensile quadrant of this left-lateral
639 fault. Tension values exceed the tensile strength of 3.2 MPa in much of this region. Fig. 10C
640 shows that the maximum Coulomb stress is located out ahead of the fault tip. The Coulomb
641 stress fails to exceed the inherent strength of the gypsum ($S_o = 78.7$ MPa) in much of this
642 region, except for directly in front of the fault. If we focus in the vicinity of growth, at half of
643 1 element length (0.0635 mm) from the fault tip (Fig. 10D), the Coulomb stress reaches
644 87.1 MPa and exceeds the shear strength of the rock. In addition, $\sigma_1 = 72.3$ MPa at this
645 location, exceeding the tensile strength of the gypsum ($T = 3.2$ MPa). T is exceeded at the
646 two locations to the right of the fault tip as well, where $\sigma_1 = 82.2$ MPa and 75.4 MPa.

647 Predicting the propagation path from these results is problematic. First, tensile
648 stresses are large and exceed the tensile strength of the rock in multiple locations. This
649 suggests that tensile failure in the form of a wing crack off the tip of the fault would occur
650 prior to any failure in shear. Second, ignoring the potential for tensile failure, both Coulomb
651 planes at the location where the shear strength is exceeded out ahead of the fault tip have
652 equal potential for failure. The plane with left-lateral slip aligns with the pre-existing fault
653 and may be the better candidate for fault growth in this scenario, but such a determination
654 is difficult to automate, as it may not be possible in other scenarios. Furthermore, selecting
655 incorrectly will greatly alter the prediction of any subsequent growth.

656 Using a work minimization approach has the advantage of evaluating both tensile
657 and shear failure by another metric, in this case the energy efficiency of the system, and
658 consequently avoids this fault plane ambiguity. The tensile failure of wing cracks is
659 considered explicitly within the work minimization analysis, but under the loading used
660 here, wing cracks are not as efficient as in-plane fault propagation. Considering the tensile
661 failure and Coulomb failure independently at points ahead of the fault does not provide
662 insights in the competition between these two mechanisms. In addition, within the work
663 minimization approach, a pupative growth element is evaluated as a continuation of the
664 pre-existing fault and experiences the slip and stress associated with that configuration.
665 For many scenarios, this provides an advantage over evaluating slip at a point ahead of the
666 fault tip. While the process zone around the fault tip no doubt includes microcracks that
667 are not radial to the fault tip, consideration of tip-radial propagation conveniently
668 considers the overall fault propagation path.

669 **5. Implications**

670 Examination of the complete energy budget provides insight into the competing
671 influences of different processes within deforming fault systems. In an innovative study of
672 multilayer folding, Ismat (2008) uses work minimization to explore the competition
673 between different flexural mechanisms within the competent and incompetent layers of an
674 evolving fold. Because gravity controls the growth of accretionary wedges, we might
675 expect gravitational work to be the largest part of the accretionary experiment energy

676 budget. DelCastello and Cooke (2007) showed that, within sandbox models of accretion, the
677 greatest energy is consumed by the work of against friction. The frictional work arises from
678 the weight of the overlying material, so gravity also plays a role in this scenario. The
679 growth of new faults in front of the wedge serves to decrease frictional work, at the
680 expense of an increase in internal work. The full work-budget analysis undertaken by
681 DelCastello and Cooke (2007) revealed this complex trade-off between different
682 components of the work budget. At a much larger scale, Meade (2013) showed that, within
683 crustal scale models, internal work is the dominant energy sink. Thus, different systems
684 require attention to different components of the work budget.

685 One challenge for any work budget investigation of fault systems is the appropriate
686 bounds for the system. Consider a complex network of active faults, such as at the
687 boundary of the Pacific and North American plates in southern California. Perhaps this
688 system has evolved to minimize work at the plate boundary scale, but can we extract
689 portions of this system (e.g. the Los Angeles basin) and reliably assume that this portion
690 should also minimize work? It could be that local mechanical inefficiencies develop within
691 a system that is efficient overall. If the system under consideration within the model is too
692 constrained in scale, then we may not be allowing for local inefficiencies to arise. Another
693 way to think of this is that we do not yet know if work minimization predictions of fault
694 evolution are independent of scale.

695 Another challenge in work minimization analyses is that active fault systems do not
696 always realize the most efficient configuration. We see this within sandbox models of
697 accretion, where underthrusting persists until a new fault can grow, even though a new
698 fault would be more efficient much earlier in the deformation (Fig. 6) (DelCastello and
699 Cooke, 2007). However, growth of this new thrust fault is not possible until the system has
700 enough energy to form the new fault surface. This energy requirement suggests that the
701 degree of inefficiency that a fault system can tolerate depends on the strength of the crust,
702 as well as on the details of the fault network geometry. It may be that inefficient fault
703 configurations persist longer in regions of thick-skinned tectonics, where crustal strength
704 is high, than in accretionary systems, where material is weaker.

705 We now have the tools and methods to make great advances in the study of fault
706 system evolution via work minimization. In the application demonstrated here using

707 numerical simulations of laboratory experiments, minimum external work determines the
708 direction of the most efficient fault propagation path, whether the pupative element fails by
709 in tension or shear. Elements that do not fail by one of these criteria do not alter W_{ext} , so the
710 efficiency of the system is unchanged. Thus, by utilizing these failure criteria along
711 pupative elements, the principle of work minimization is applied and theories governing
712 fault mechanics are honored. This approach has the advantage of accommodating both in-
713 line growth and growth at angles to pre-existing faults without invoking separate tensile
714 and shear failure criteria that require parallel analyses. The work minimization approach
715 also proves to be a viable alternative to using Coulomb theory, which empirically captures
716 the physics of faulting, but is limited in its ability to predict propagation at fault tips. The
717 main limitation is in generation of two equally viable potential failure planes. While one of
718 the failure planes may match the observations of fault growth, Coulomb theory does not
719 provide a means for objectively picking one plane over the other. While Coulomb theory is
720 remarkably helpful for predicting failure of pre-existing structures, or comparing failure
721 potential between multiple fault surfaces, it may not be so for predicting fault propagation
722 and evolution.

723 We see two main applications of the principle of work minimization to analyses of
724 geologic structure. The first is to use the full energy budget, or its individual components,
725 to gain insight into processes of deformation. For example, recent work by Savage et al.
726 (2014) constrained the energy required for fault frictional heating by estimating
727 temperatures from biomarkers near an exhumed fault. Work against friction also was
728 estimated recently by Fulton et al. (2013) from in-situ measurements of a fault that failed
729 in the 2011 Tohoku earthquake. Similarly, the work budget within accretionary systems
730 has shed light on the tradeoffs of underthrusting and accretionary phases in wedge
731 development (e.g. Gutscher et al., 1998; Hardy et al., 1998; Burbidge and Braun, 2002; Del
732 Castello and Cooke, 2007; Mary et al., 2013). Other studies that could benefit from similar
733 analyses include the partitioning of off-fault deformation between different faults, the
734 trade-off between various processes and uplift against gravity, and how different faulting
735 scenarios affect estimates of the energy available for seismic shaking.

736 The second application of work minimization to structural geology is the prediction
737 of fault propagation paths. This analysis has been done by comparing the total or external

738 work of different interpretations of fault configurations along the southern San Andreas fault
739 over the past 500,000 years (Cooke and Dair, 2011). Alternatively, work minimization can
740 be used in conjunction with numerical modeling tools to predict propagation paths over
741 multiple growth cycles, as presented here. The development of software to automate this
742 process would provide powerful tools for predicting fault growth.

743 **Acknowledgements**

744 Jess McBeck provided invaluable software support for the project. The manuscript was
745 greatly improved by reviews by Bill Dunne, Tom Blenkinsop and an anonymous reviewer.
746 This work was supported by NSF EAR grant 1219919.

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930

930 **Figures**

931 Figure 1. Schematic of fault deformation shows the tectonic work (W_{ext}), internal work of
932 deformation (W_{int}), work of uplift against gravity (W_{grav}), the work against frictional slip
933 (W_{fric}), the energy required to grow a fault (W_{prop}) and the radiated seismic energy (W_{seis}).
934 Taken from Cooke and Murphy, 2004.

935 Figure 2. Schematic of the partitioning of frictional heating, the energy required for fault
936 growth and energy available for seismic shaking. Work is done when fault slips and the
937 shear stress drops from the initial value, τ_0 , to the value at stable sliding, τ . One can
938 delineate the associated work terms in different ways. In most studies, W_{fric} is defined as
939 the energy due to slip under sliding shear stress, consequently W_{fric} depends on absolute
940 stress. The two energy terms that are consumed by fault growth (W_{seis} and W_{fric}) depend on
941 the shear stress drop during the slip event. Prior to the slip reaching the slip weakening
942 distance, L , work is primary consumed by the generation of new fault surface area (W_{prop})
943 and W_{fric} . Slip beyond the slip weakening distance produces W_{fric} and seismic radiated
944 energy (W_{seis}). While the sketch in (A) shows this relationship for a single increment of
945 loading the hypothetical sketch in (B) shows how W_{fric} , W_{prop} and W_{seis} are calculated over
946 multiple loading increments, which are typically implemented within numerical models.

947 Figure 3. Minimum viscous dissipation demonstrates that of the possible ocean ridge (red)
948 and transform (blue) intersections, several of which are show here, geometry (C) is the
949 most efficient. Modified from Lachenbruch and Thompson, 1972.

950 Figure 4. A) Strain energy density (SED) and B) Coulomb stress for the interpreted Puente
951 Hills fault system prior to the growth of the Coyote Hills fault (CH) in southern California.
952 The region of high SED envelops the location of the Coyote Hills fault. If we use high SED
953 and extreme values of Coulomb stress to predict the development of the next fault, the SED
954 pattern gives a better correlation with the location of the observed fault. The Coulomb
955 stress is less conclusive than the SED and suggests that the next faults may develop either
956 northeast or southwest of the Whittier fault. Taken from Olson and Cooke, 2005.

957 Figure 5. (A) Limit analysis can be used to predict the sequence of thrusting. (B) The
958 predicted evolution of the tectonic force shows drops in force with the growth of new

959 faults. Taken from Cubas et al. (2008).

960 Figure 6. DelCastello and Cooke (2007) explored the evolving work budget of forethrust
961 growth within the analog experiments of Adam (2005). (A) The numerical models simulate
962 lengthening of the thrust sheet prior to forethrust development. (B) With lengthening of
963 the thrust sheet length, W_{fric} increases as the fault normal compression increases due to the
964 thickening wedge. Consequently, the fault system becomes increasingly inefficient until a
965 new forethrust grows. This fault growth reduces the W_{ext} by decreasing W_{fric} along the fault
966 by an amount greater than the increase in W_{int} associated with the addition of the new fault.
967 (C) If the forethrust were not to grow, the W_{ext} would continue to increase. The timing of
968 fault growth depends on the work required to grow the new fault surface. The new fault
969 grows at 10.5 cm thrust fault length rather than sooner in the experiment because the total
970 energy reduction due to including the fault within the model (prefault W_{ext} – post fault
971 W_{ext}) exceeds the energy consumed in the growth of the fault (W_{prop} and W_{seis}). Taken from
972 DelCastello and Cooke, 2007.

973 Figure 7. Two models based on laboratory experiments on faults within gypsum slabs by
974 Bobet and Einstein (1998). (A) Set-up for the numerical experiment of a 12.7 mm joint in a
975 160 x 80 mm slab of gypsum. The block is subject to 27.1×10^{-6} m of uniaxial extension (u_v). E
976 is Young's modulus and ν is Poisson's ratio. (B) Set-up for the numerical experiment on a
977 12.7 mm fault within gypsum that dips 60° from the vertical compressive stress, $\sigma_v = -29$
978 MPa. A confining pressure of $\sigma_h = -2.5$ MPa also is applied. Box shows location of (C). Box at
979 fault tip is the location of the inset, showing pupative elements radial to the fault tip.

980 Figure 8. ΔW_{ext} per pupative element area from before growth to after growth of (A) the
981 joint and (B) the fault models. Red markers are pupative element orientations where
982 neither the tensile nor the shear failure criterion is met, so no growth can occur and there
983 is no change in work. Green markers are orientations where the tensile strength is
984 exceeded and failure can occur by opening. Blue markers are orientations where shear
985 failure can occur. Tensile strength, T , is 3.2 MPa. Shear strength, S_o , is 78.7 MPa.

986 Figure 9. Work budgets for joint and fault propagation by one pupative element in the
987 orientation that maximizes ΔW_{ext} . The energy to grow faults ($W_{prop} + W_{seis}$) comes from

988 ΔW_{int} . Growth elements for Fault 1 are not slip-weakening and do not experience a drop in
989 strength during sliding. The inherent friction, $\mu_o = 0.3$, is equal to the dynamic friction, μ_d .
990 Growth elements for Fault 2 are slip-weakening, with $\mu_o = 0.6$, $\mu_d = 0.3$, and $L = 4 \times 10^{-6}$ m.
991 Growth elements for Fault 3 also are slip weakening, with $\mu_o = 0.6$, $\mu_d = 0.3$, and $L = 1 \times 10^{-6}$
992 m. The most efficient propagation path for both the modeled joint and for all three faults is
993 collinear at 180° .

994 Figure 10. (A) Planes carrying maximum tensile stress are perpendicular to the maximum
995 tensile stress, σ_1 (top, green), while the two planes carrying the maximum Coulomb shear
996 stress are oriented at angle φ to the maximum compressive strength, σ_3 (bottom, red). (B)
997 Contours of the maximum tensile stress. (C) Contours of the maximum Coulomb shear
998 stress, σ_c , on optimally oriented planes with $\mu_o = 0.3$. (D) Tensile and Coulomb failure
999 planes at 5 points located at one half the element length from the fault tip (0.0635 mm).
1000 Tensile strength, T , is 3.2 MPa. Shear strength, S_o , is 78.7 MPa. Strength must be exceeded
1001 for failure to occur (solid lines). Tensile and shear failure can occur at a variety of points
1002 around the fault tip, but this approach does not permit prediction of the favored
1003 propagation path for the fault.

1004

Figure 1

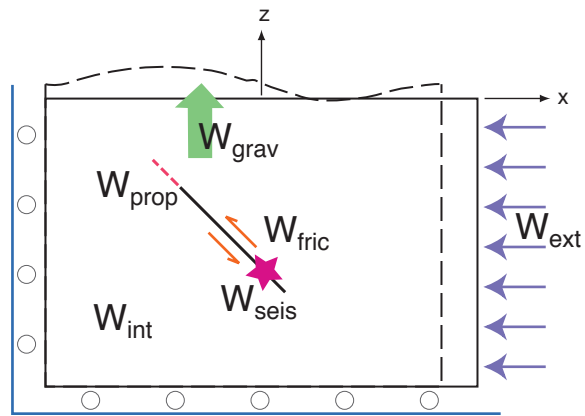


Figure 2

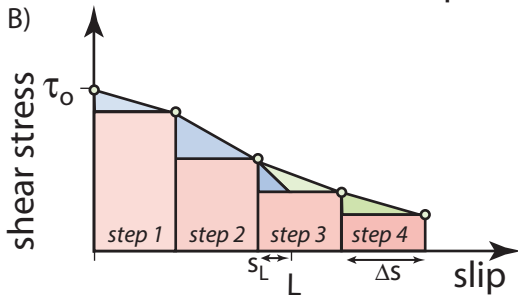
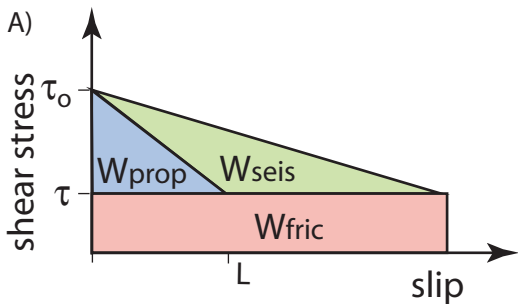


Figure3

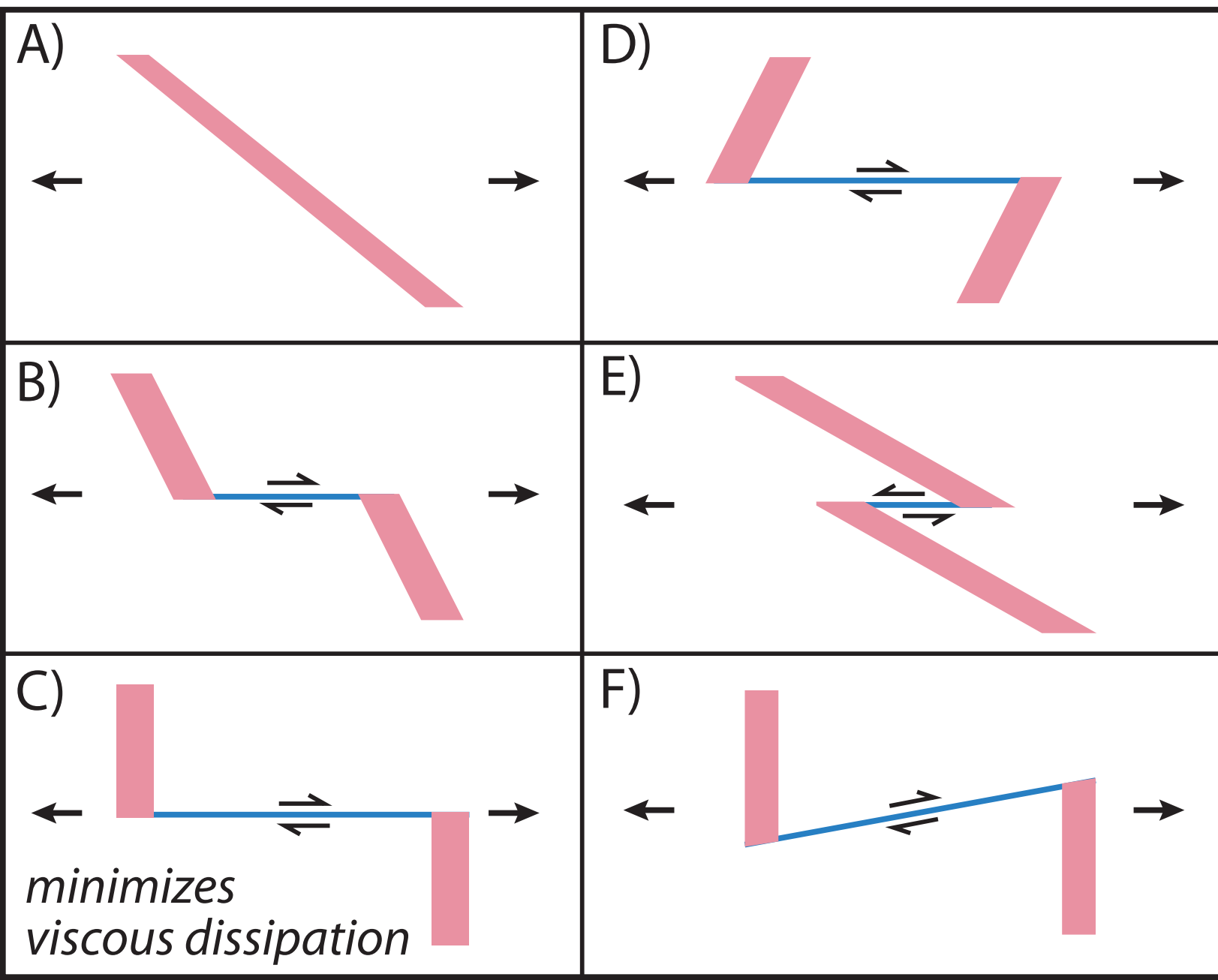
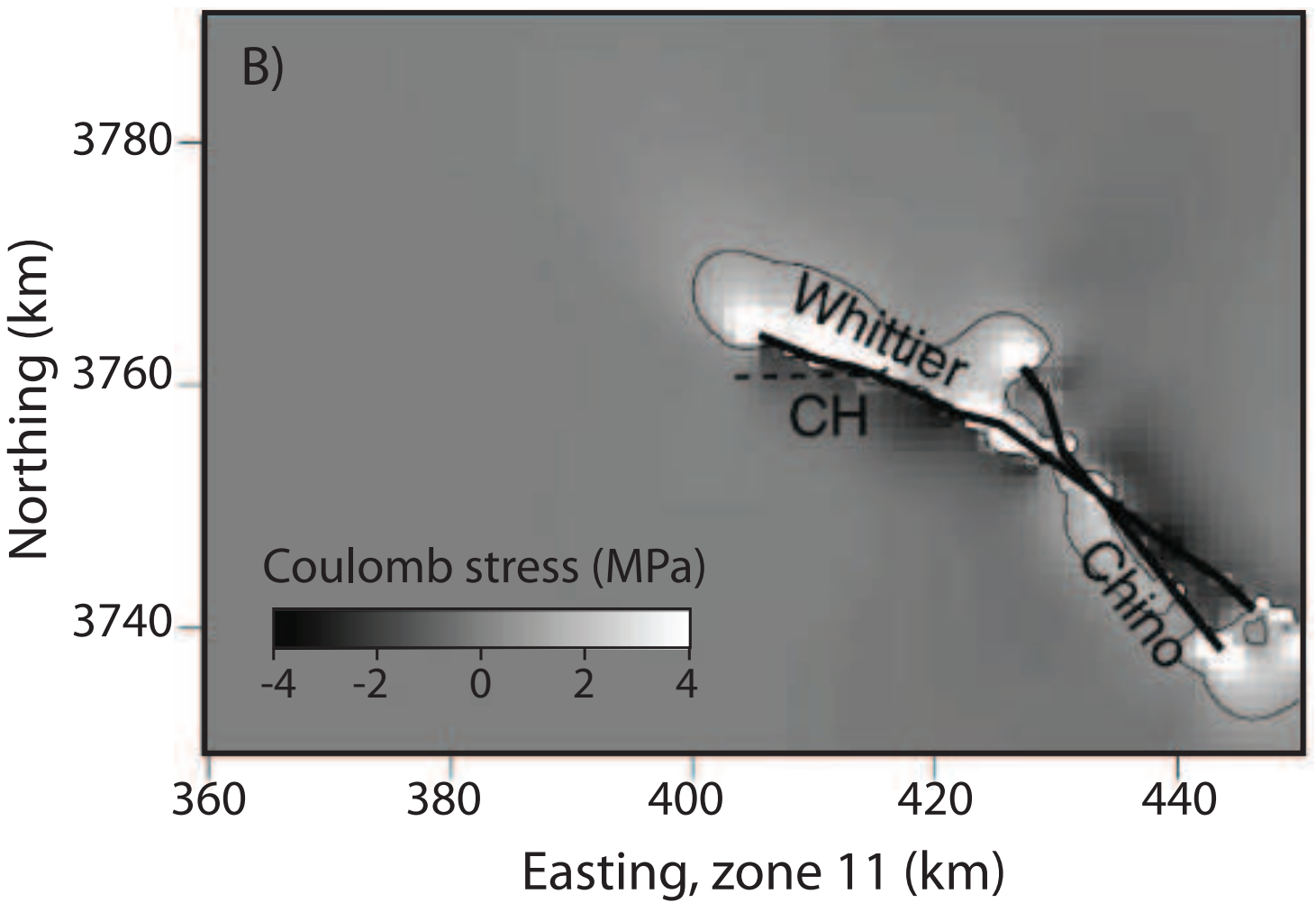
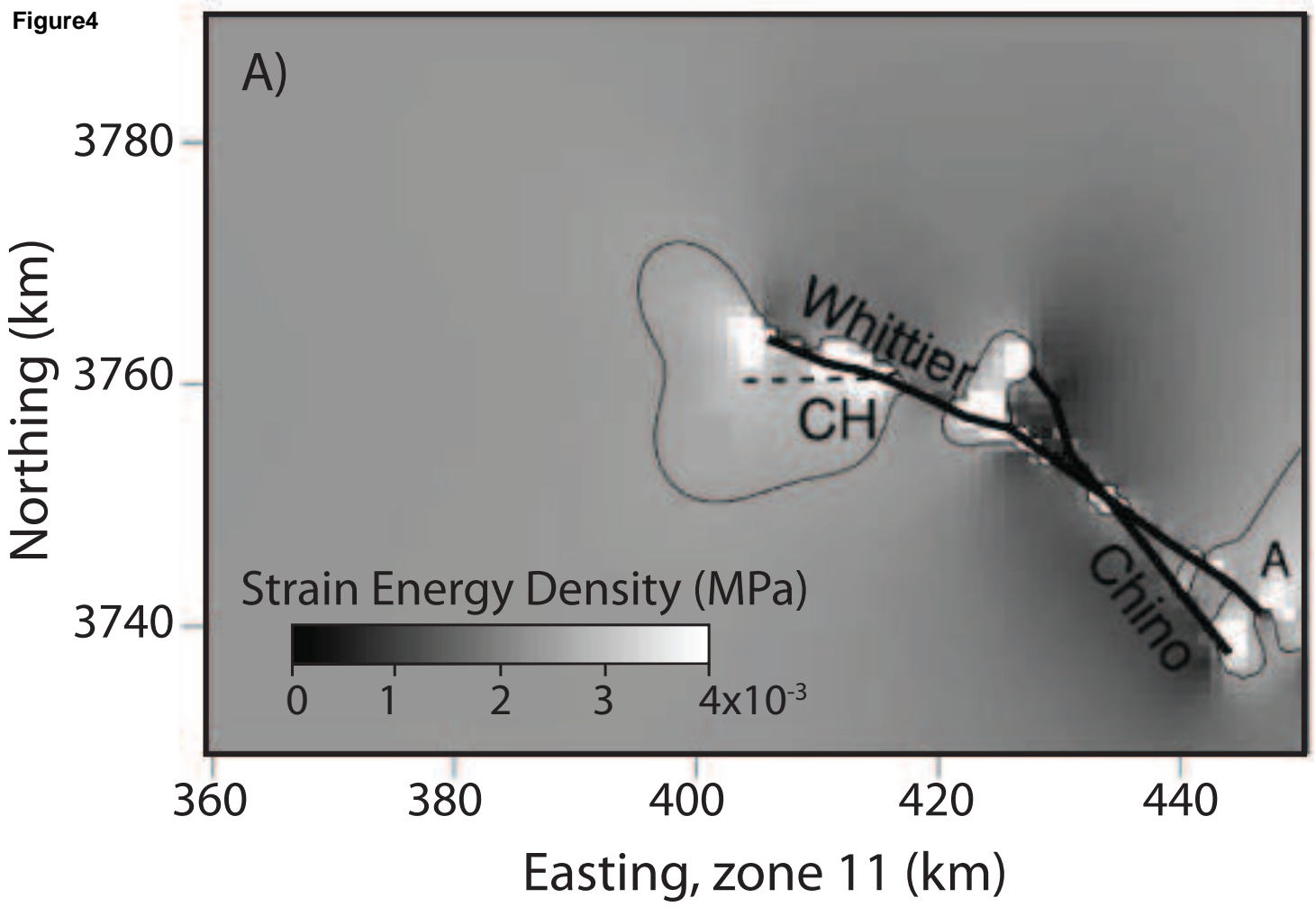
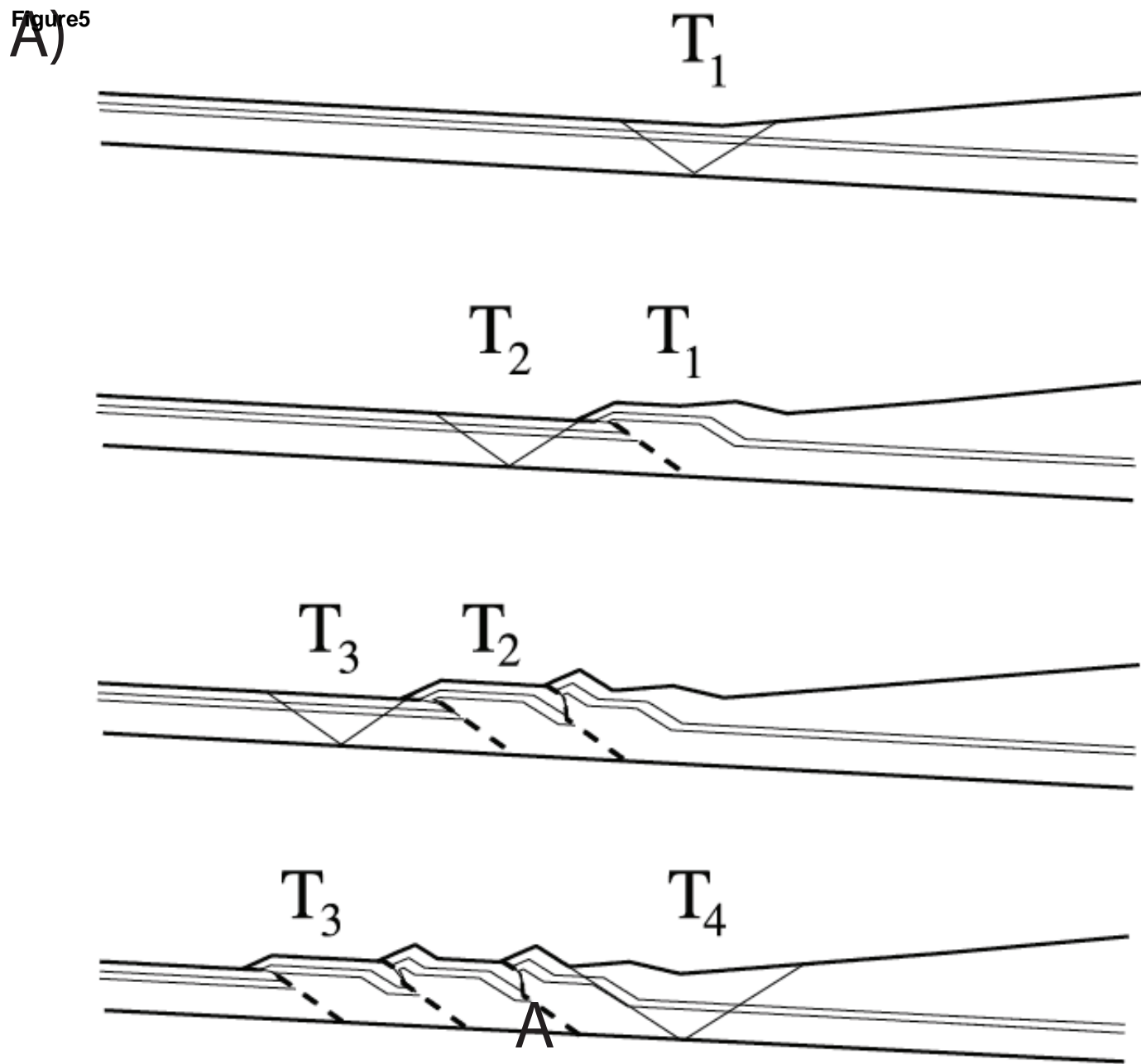


Figure4





B)

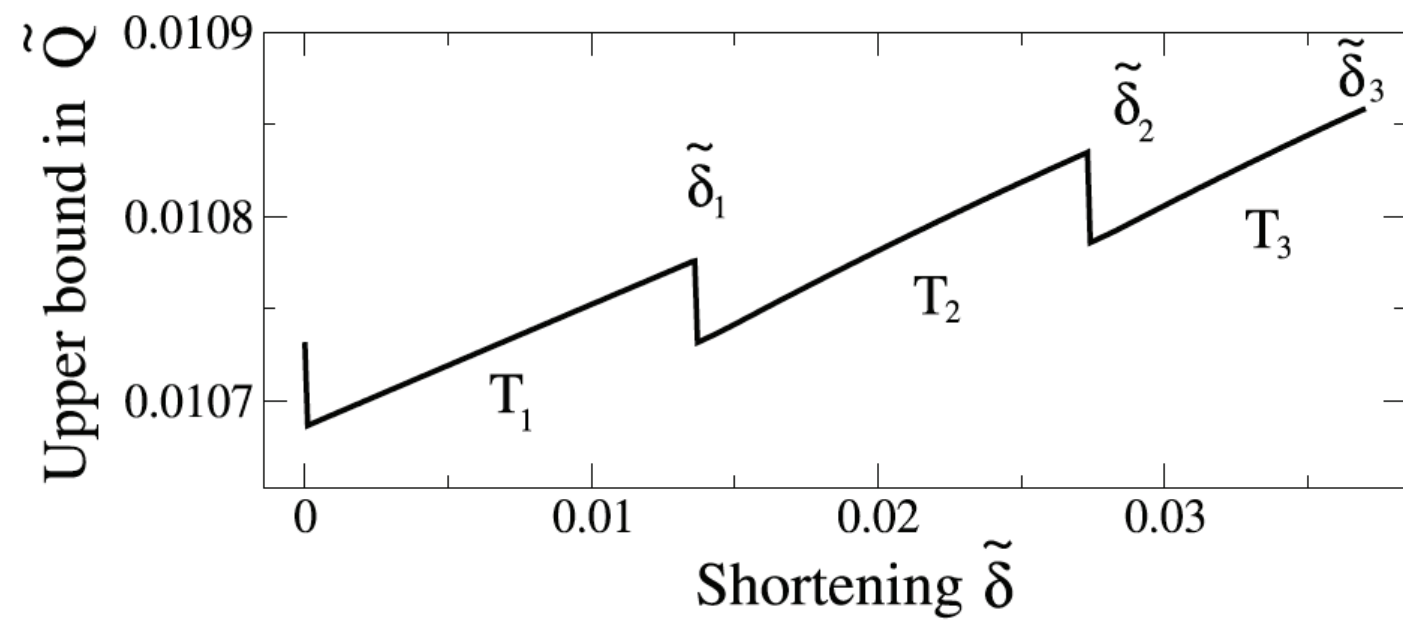
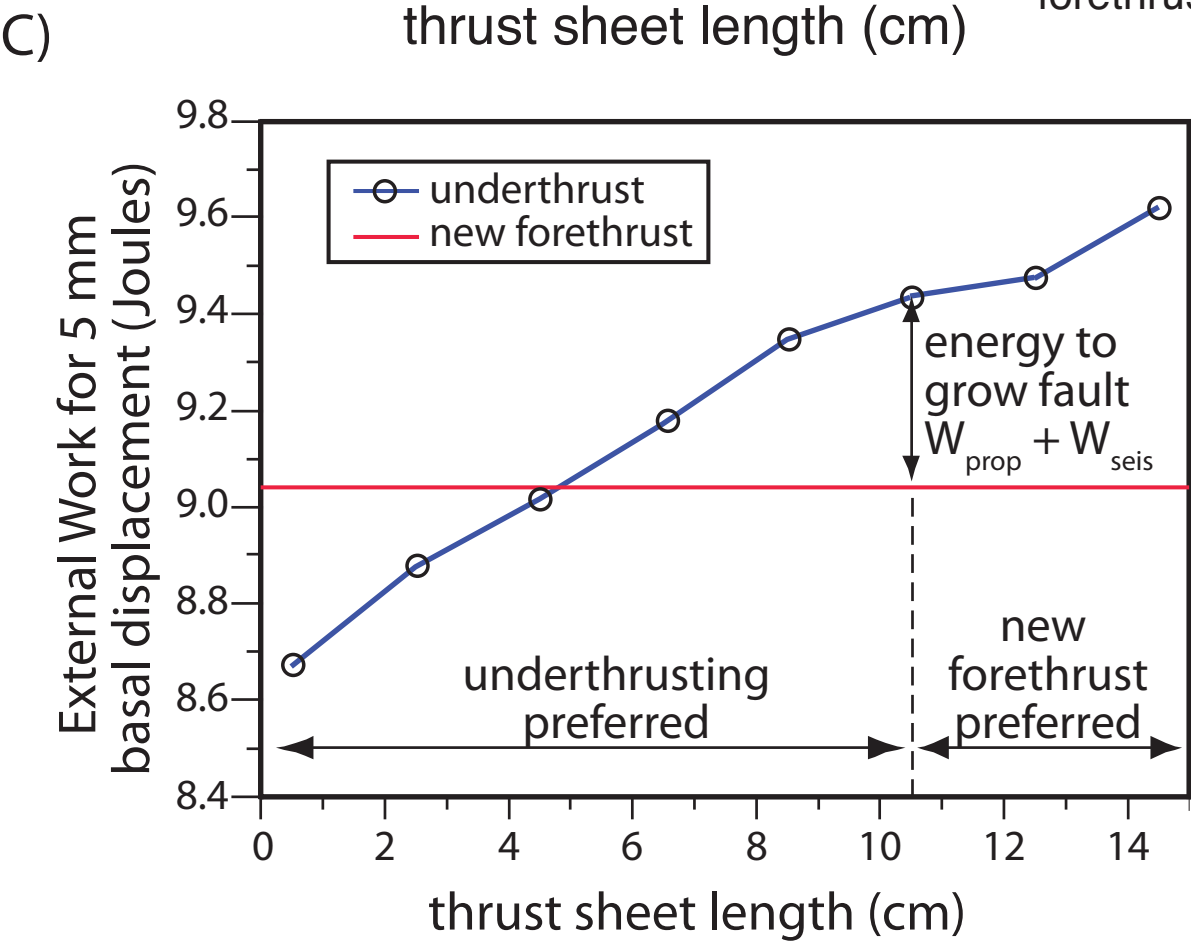
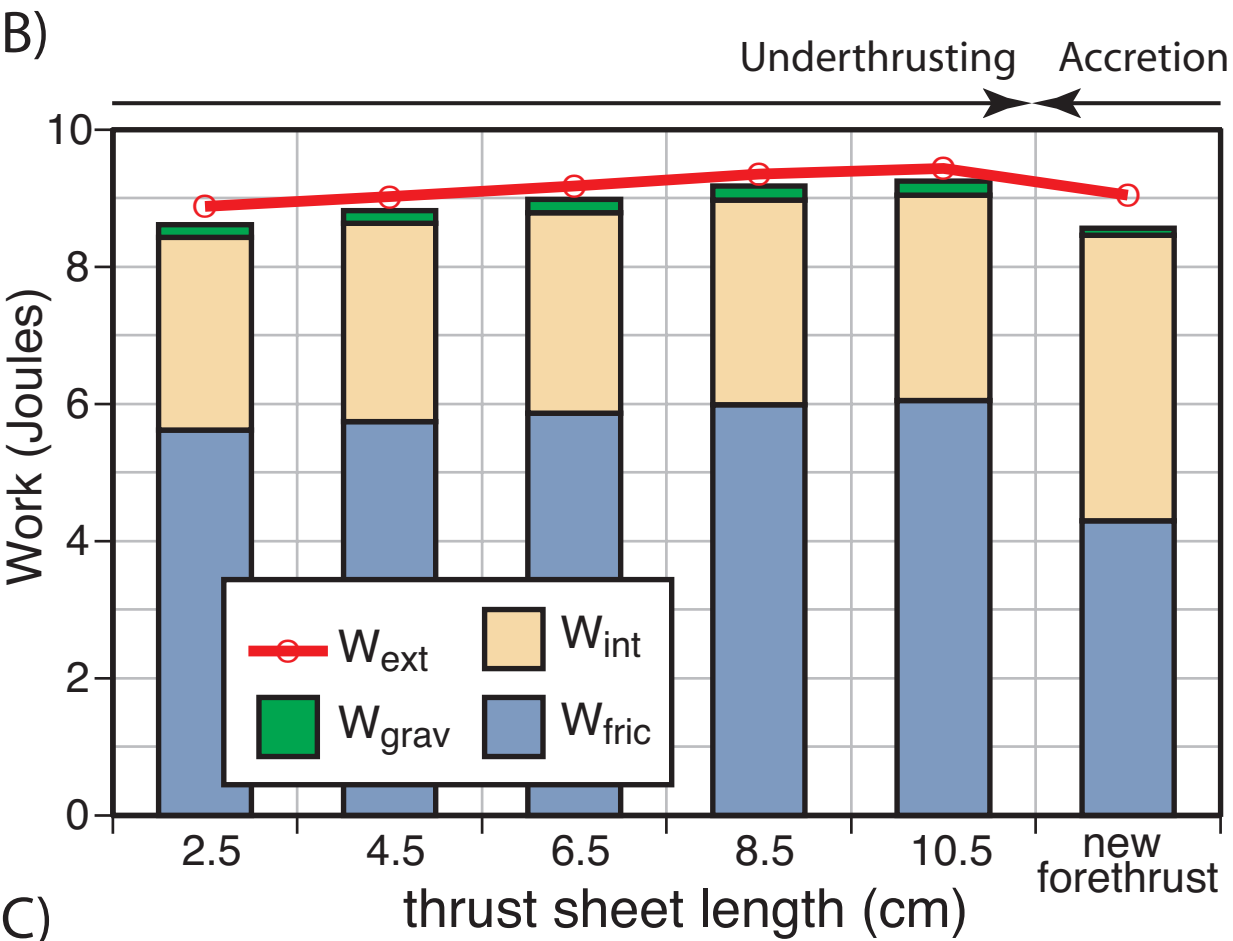
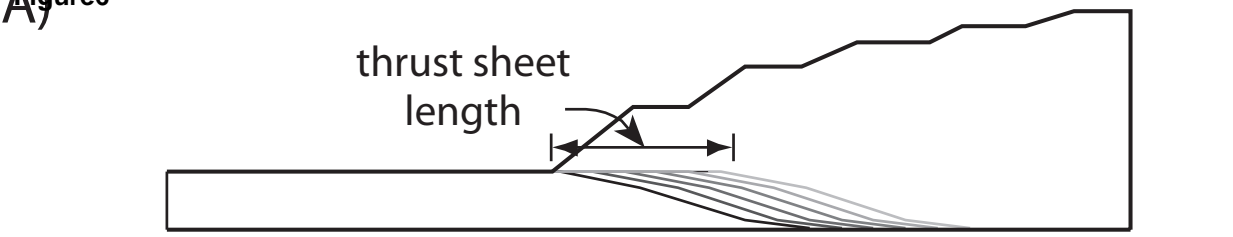
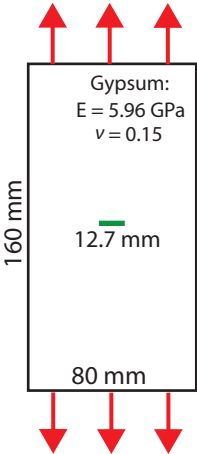


Figure 6



A) Figure 7



B)

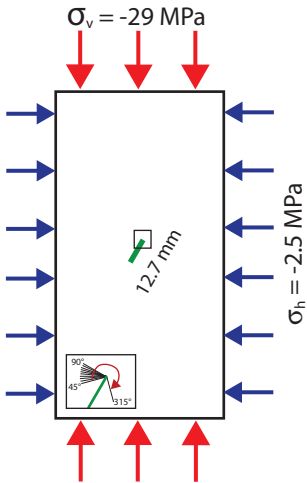


Figure 8

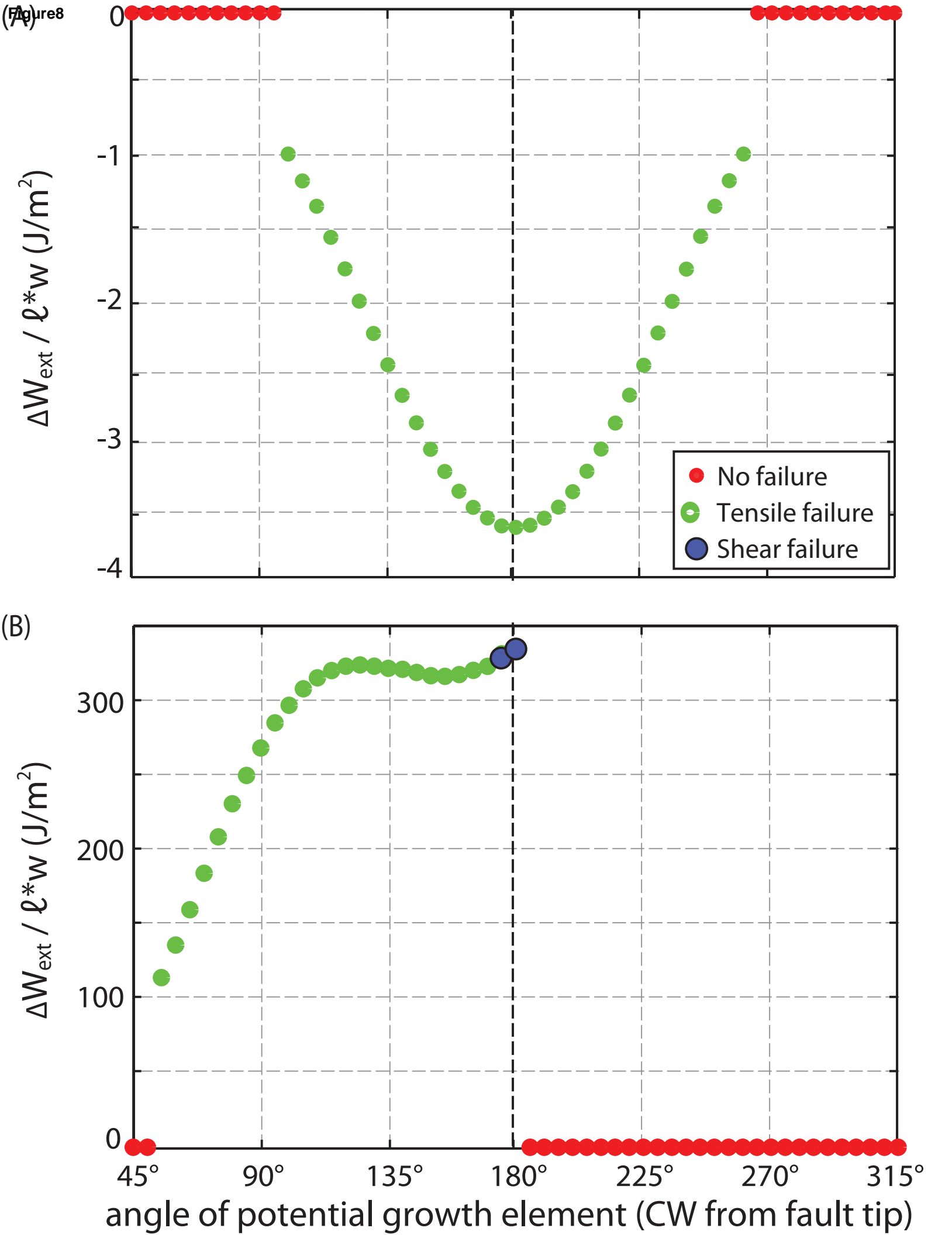
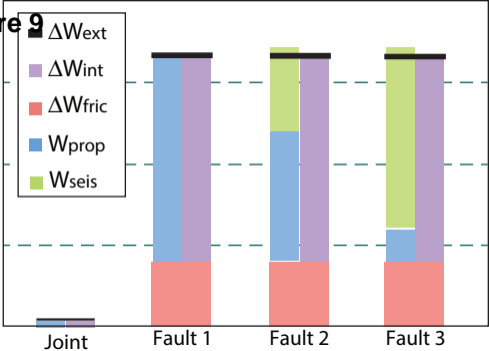
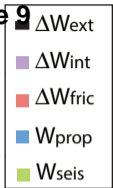


Figure 9 $\Delta\text{Work} / \ell^*w$ (J/m^2)

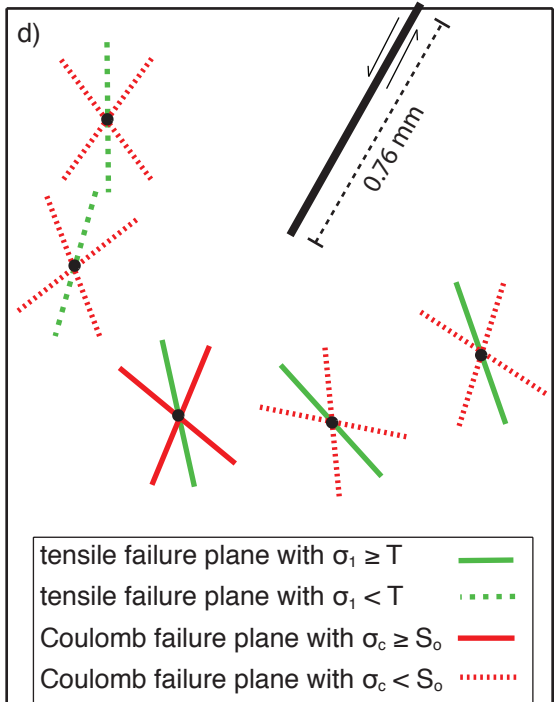
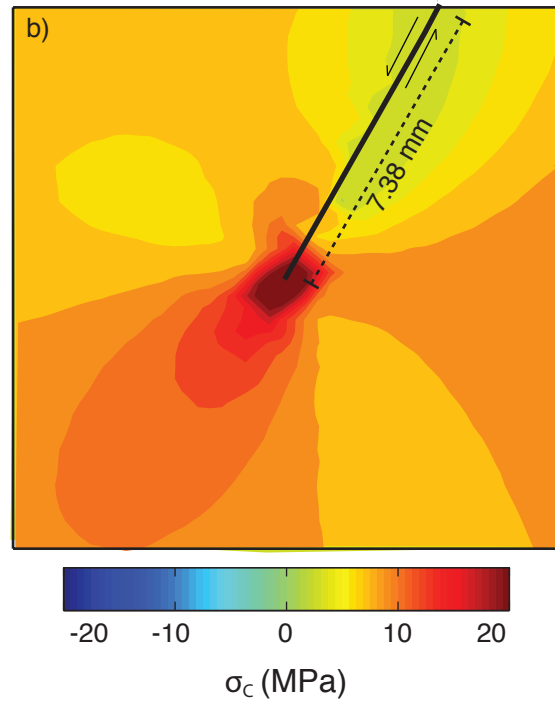
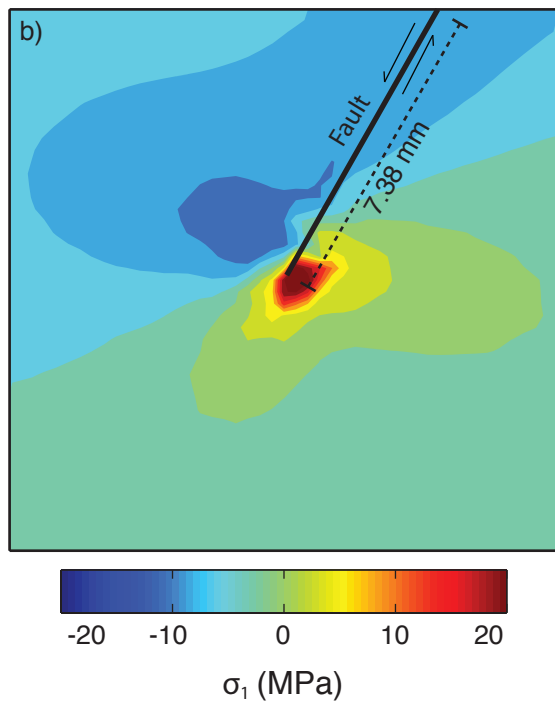
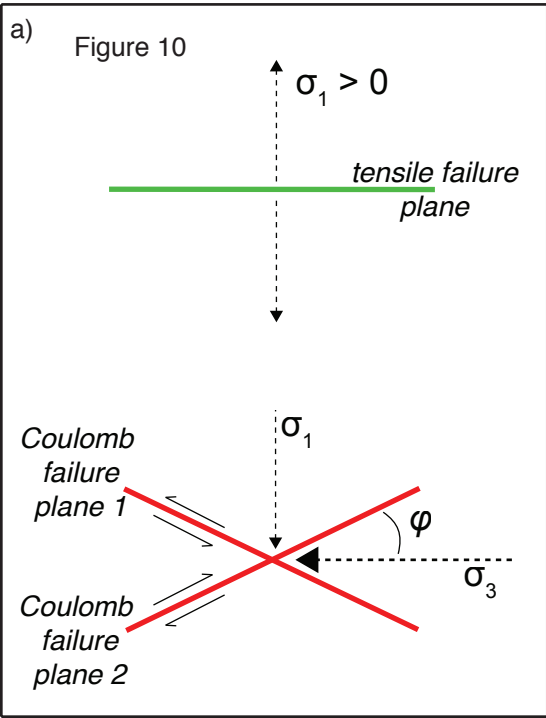


Table 1: Properties and work budgets for growing joint and fault models

Scenario	μ_s	μ_d	L (m)	W_{prop} (J/m ²)	W_{seis} (J/m ²)	ΔW_{fric} (J/m ²)	$W_{prop} + W_{seis} + \Delta W_{fric}$ (J/m ²)	ΔW_{int} (J/m ²)	ΔW_{ext} (J/m ²)
Joint	0	0	0	3.62	0.00	0.00	3.62	3.63	3.63
Fault 1	0.3	0.3	0	260.77	0.00	79.78	340.54	254.05	333.83
Fault 2	0.6	0.3	1e ⁻⁶	158.60	102.14	79.45	340.19	254.38	333.83
Fault 3	0.6	0.3	4e ⁻⁶	39.70	221.11	79.45	340.26	254.38	333.83