

1 Time scales of shallow magma chamber
2 replenishment at Campi Flegrei caldera.

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8 **Abstract**

9 Ascent of primitive magmas from depth into shallow, partially de-
10 gassed reservoirs is commonly assumed to be a viable eruption trigger.
11 The resulting processes of convection and mixing have played an impor-
12 tant role both in pre- and syn-eruptive stages in many eruptions of differ-
13 ent sizes at the unrest Campi Flegrei caldera in Southern Italy.

14 We performed numerical simulations of magma chamber replenishment
15 referring to an archetypal case whereby a shallow, small magma cham-
16 ber containing degassed phonolite is invaded by volatile-rich shoshonitic
17 magma coming from a deeper, larger reservoir. The system evolution is
18 driven by buoyancy, as the magma entering the shallower chamber is less
19 dense than the degassed, resident phonolite.

20 The evolution in space and time of physical quantities such as pressure,
21 gas content and density is highly heterogeneous; nonetheless, an overall
22 decreasing exponential trend in time can be observed and characterizes the
23 efficiency of the whole process. The same exponentially decreasing trend
24 can be observed in the amplitude of the synthetic ground deformation
25 signals (seismicity over the whole frequency spectrum) calculated from the
26 results of the magmatic dynamics. Depending on the initial and boundary
27 conditions explored, such as chamber geometry or density contrast, the
28 time constant thus the inferred duration of the process can vary. An
29 initial vigorous phase of convection and mixing among the two magma
30 types reaches an asymptotic stage after a few hours to half a day.

31 Independently, the evolution of pressure in the magmatic system also
32 depends on the initial and boundary conditions, leading either to eruption-
33 favorable conditions or not. Relating the time scales for convective pro-
34 cesses to be effective with their outcomes in terms of stresses at the bound-
35 aries of the magmatic system can substantially improve our ability to
36 forecast the evolution of volcanic unrest crises worldwide.

37 **Keywords:** magma chamber, magma mixing, caldera unrest, Campi
38 Flegrei, magma dynamics

39 1 Introduction

40 The restless Campi Flegrei caldera, in southern Italy, is one of the highest-risk
41 volcanic areas in the world, due to the millions of people living in its surround-
42 ings, that include the city of Naples. Within a radius of around 7 km intra-
43 calderic craters and volcanic edifices testify the post-collapse eruptive history
44 of the volcano, that last erupted in 1538 AD (Di Vito et al., 1999; D’Oriano
45 et al., 2005; Di Vito et al., 2016). Nowadays, the system is characterized by
46 intense hydrothermal circulation with abundant fumaroles, episodes of acceler-
47 ating ground deformation, and sporadic seismic swarms (Chiodini et al., 2015).

48 Geochemical (Arienzo et al., 2009; Di Renzo et al., 2011) and geophysical
49 (Judenherc and Zollo, 2004; Zollo et al., 2008; De Siena et al., 2010) evidence
50 suggests that the plumbing system at Campi Flegrei is composed of a large, deep
51 (around 8 km), dominantly shoshonitic reservoir (Mangiacapra et al., 2008), and
52 shallower, smaller (less than 1 km³) reservoirs that discontinuously occur in a
53 depth range of 3 km to 6 km and host small pockets of trachytic to phonolitic
54 magmas (Arienzo et al., 2010).

55 The arrival of primitive, volatile-rich magma from depth into the shallow
56 reservoirs has been identified as a viable trigger mechanism for some past erup-
57 tions at Campi Flegrei (Tonarini et al., 2009). Numerical simulations of this
58 process highlight the convective patterns leading to efficient mingling between
59 the two magmas (Montagna et al., 2015). The resulting pressure oscillations gen-
60 erate ground displacement patterns in the ultra-long-period band (ULP, 10⁻⁴ Hz
61 to 10⁻² Hz), that can be recorded by strainmeters networks (Longo et al., 2012b;
62 Bagagli et al., 2017).

63 In this work we perform a through analysis of the time scales over which
64 the aforementioned replenishment processes are expected to be efficient. Those

65 same processes can be related to the stress evolution at the reservoir boundaries
66 in order to determine in which cases the conditions for eruption are met. The
67 method can provide estimates of mixing-to-eruption time scales, which can be
68 of great help for hazard assessment.

69 **2 Material and Methods**

70 **2.1 Numerical model**

71 Numerical simulations of magma chamber replenishment at Campi Flegrei were
72 performed using the finite-element C++ GALES code (Longo et al., 2012a).
73 The physico-mathematical model describes the time-dependent 2D dynamics of
74 a compressible-to-incompressible multicomponent mixture consisting of liquid
75 (or crystal suspension) in thermodynamic equilibrium with an $\text{H}_2\text{O}+\text{CO}_2$ gas
76 phase at the local conditions of pressure, temperature and composition. The
77 numerical algorithm used in the solution of the conservation equations is based
78 on an extension of the finite element formulation by Hauke and Hughes (1998)
79 to include multicomponent fluids (Longo et al., 2006), allowing the investigation
80 of processes involving mixing of fluids, chemical changes, and phase transitions.
81 The algorithm consists of a space-time discretization with Galerkin least-squares
82 and discontinuity capturing terms, with third order accuracy in time and space.
83 This method allows the simulation of both compressible and incompressible flow
84 dynamics (Shakib et al., 1991; Chalot and Hughes, 1994; Hauke and Hughes,
85 1998; Garg et al., 2018), and it is effective in stabilizing the numerical solution
86 without introducing excessive numerical diffusion. A large number of problems
87 can be solved, such as natural and forced convection, shock waves and their
88 interaction with contact discontinuities, evolution of internal interfaces in in-
89 compressible or compressible flows, and bubbly flows with evaporation or gas
90 dissolution. The conservation equations for the mass of single components and
91 momentum of the whole mixture, together with the gas-liquid thermodynamic
92 equilibrium model and the constitutive equations for mixture properties (den-
93 sity and viscosity), are discretized and solved for the primitive variables pressure,
94 velocity, temperature, and concentration of components. Magmas are described
95 as ideal mixtures with components that may be either in the liquid or gaseous
96 state, with instantaneous phase change according to the non-ideal multicom-
97 ponent $\text{H}_2\text{O}+\text{CO}_2$ saturation model of Papale et al. (2006). The solid phase
98 (suspended crystals) can be taken into account in the computation of mixture

99 properties. Gas bubbles are assumed to be undeformable, a good approxima-
100 tion if the bubble size is smaller than $\sim 10\ \mu\text{m}$ (Marchetti et al., 2004). This
101 corresponds to a gas volume of 5 % to 50 % for bubble number densities in the
102 range $1 \times 10^{14}\ \text{m}^{-3}$ to $1 \times 10^{15}\ \text{m}^{-3}$. The role on the relevant properties (den-
103 sity and viscosity) of dissolved water and of dispersed gas and solid phases, and
104 the mutual roles of H_2O and CO_2 in affecting their saturation contents, are
105 accounted for. Mixture density is calculated using the Lange (1994) equation
106 of state for the liquid phase, real gas properties and standard mixture laws for
107 multiphase fluids. Mixture viscosity (under the assumption of Newtonian rheol-
108 ogy) is computed through standard rules of mixing (Reid et al., 1977) for one
109 phase mixtures and with a semi-empirical relation (Ishii and Zuber, 1979) in
110 order to account for the effect of non-deformable gas bubbles. Liquid viscosity
111 is modeled as in Giordano et al. (2008), and it depends on liquid composition
112 and dissolved water content.

113 2.2 Campi Flegrei system setup

114 In order to understand the dynamics of magmas beneath the Campi Flegrei
115 caldera, we simplify the magmatic system trying to retain its most peculiar
116 features. A schematic picture of the feeding system includes the deep magmatic
117 sill containing the less evolved, CO_2 -rich shoshonitic magma; and more or less
118 persistent features, such as dikes, fractures or conduits, linking the deep reservoir
119 to shallower, degassed batches of magma. Shallower reservoirs contain magmas
120 with compositions from trachyte to phonolite.

121 We model the injection of CO_2 -rich, shoshonitic magma coming from a deep
122 reservoir into a shallower, much smaller chamber, containing more evolved and
123 partially degassed phonolitic magma. The deep reservoir is schematized as a sill;
124 the geometry of the shallow chamber has been varied from oblate to prolate in
125 a set of simulations, as shown in Figure 1. Initial conditions for the simulated
126 cases are reported in Table 1.

127 The top of the deep reservoir is at 8 km depth, and its horizontal and vertical
128 semi-axes measure 4 km and 0.5 km, respectively. The shallow reservoir has
129 its top at 3 km below the surface; its surface area is kept to $0.25 \times 10^6\ \text{m}^2$,
130 representing a volume of magma of order a few tenths of km^3 . The connecting
131 dike is 20 m wide. The oxides composition of the two magmas is detailed in
132 Table 2.

133 At the initial time, the two magmas are placed in contact at the bottom of

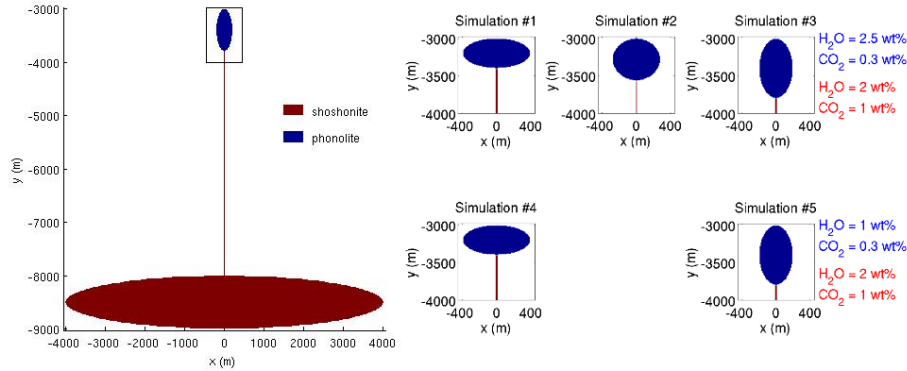


Figure 1: Summary of the five simulated settings. The geometry and volatile content of the shallow reservoir have been varied in a range of conditions relevant to Campi Flegrei.

134 the shallow chamber (top of the feeding dike: Figure 1). Because of the higher
 135 volatile content thus lower overall density of the deeper magma, the system
 136 is gravitationally unstable. Therefore, the deep magma tends to rise into the
 137 shallow chamber, pushing the resident magma down through the feeding dike.

138 We conceived different setups in order to explore a parameters' range suit-
 139 able for the Campi Flegrei feeding system. Keeping the upper chamber top at
 140 a depth of 3 km and its 2D cross-sectional area, we changed its geometry from
 141 a sill to a circle and to a vertically elongated, dike-like ellipsoid. For the two
 142 ellipsoidal chambers, we also varied the volatile content of the shallow phonolitic
 143 magma as detailed in Table 1, to represent larger or smaller amount of degassing.
 144 The deep shoshonitic magma contains in all cases 2 wt% H_2O and 1 wt% CO_2
 145 (Table 1). These quantities represent the total volatiles' amount, which is dis-
 146 tributed between liquid and gas phases according to the local conditions and
 147 the saturation model of Papale et al. (2006). The simulated systems are closed,
 148 with fixed boundaries. The effect of wall-rock elasticity on the chamber dynam-
 149 ics is neglected, as it is assumed to cause small pressure changes with respect
 150 to those originating from the magma dynamics (see also section 4).

151 The setup, albeit being a simplification, aims at representing complex, in-
 152 terconnected reservoirs that are believed to exist at many active volcanoes, in-
 153 cluding Campi Flegrei (Amoruso et al., 2014); its basic features are chosen as
 154 to avoid arbitrariness whenever possible.

155 **2.3 Shallow reservoir dynamics**

156 This work focuses on the evolution of the shallow chamber as a consequence of
 157 the replenishment process (Montagna et al., 2015; Papale et al., 2017), which has
 158 been testified in a variety of settings, in both volcanic and plutonic environments
 159 (Perugini and Poli, 2012). The onset and development of convection and mixing
 160 following the arrival of magma from depth into a shallower reservoir have been
 161 found to depend on system geometry (Turner and Campbell, 1986) as well as on
 162 the physical properties of the magmas involved, notably density and viscosity
 163 (Bachmann and Bergantz, 2006).

164 Significant physical quantities including pressure, density and gas fraction
 165 can be described using a lumped approach (Degruyter and Huber, 2014; Papale
 166 et al., 2017), whereby the spatial heterogeneities are averaged out in order to
 167 obtain a larger-scale description of the reservoir dynamics. The lumped value
 168 of the variable y in the region \mathcal{D} is given by

$$y_L = \int_{\mathcal{D}} y(\mathbf{x}) dx. \quad (1)$$

169 Figure 2 shows the heterogeneous space-time evolution resulting from the
 170 simulations alongside the corresponding lumped time series, obtained by numerical
 171 integration of the simulation results in space.

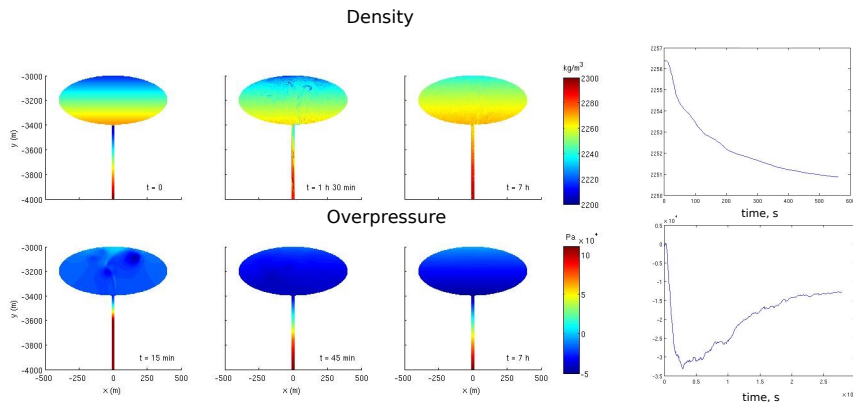


Figure 2: Space-time evolution of density and overpressure in the shallow reservoir for simulation 1, and corresponding lumped time series.

172 The upper reservoir is an open system, exchanging mass with the surround-
 173 ings through the feeder dike. A measure of the mass exchange between the two

174 magmas in a certain region of the domain is given by the convection efficiency
175 η_C :

$$\eta_C = \frac{|m_P(t) - m_P(0)|}{m_P(0)}. \quad (2)$$

176 η_C represents the relative variation of the mass of the phonolitic magma m_P
177 with time t (Montagna et al., 2015); it is a lumped variable by definition.

178 The efficiency of mass exchange depends on the dike width as well, because
179 larger dikes will allow for easier flow and enhanced mass exchange. A dike width
180 of 20 m has been chosen for consistency with exposed mafic dikes and to avoid
181 too rapid cooling near the walls (Costa and Macedonio, 2002).

182 **3 Results**

183 **3.1 Magma Dynamics**

184 In all the simulated scenarios, plumes of more primitive magma rise into the
185 shallow chamber and originate convective patterns enhancing magma mixing,
186 while a portion of the degassed phonolitic magma initially hosted in the chamber
187 sinks into the feeder dike (Figure 3).

188 Soon after the start of the simulation, disruption of the initial interface
189 by buoyancy forces produces a series of discrete plumes of light magma rising
190 through the shallow chamber and reaching its top after having developed com-
191 plex velocity fields. At the same time, part of the dense magma originally hosted
192 in the shallow chamber sinks into the feeding dike. Intense mixing originates
193 at the dike level, so that no pure shoshonitic end-member can be found in the
194 shallow chamber; the rising plumes are rather made of a mixture with 30 wt%
195 to 50 wt% deep component. In our framework, only mechanical mixing is con-
196 sidered, as chemical reactions apart from volatile exsolution/dissolution are not
197 implemented in the numerical scheme. Therefore, mixing at the scale of the sim-
198 ulation, with maximum resolution of order 1 m, refers to mechanical mingling
199 rather than chemical mixing. The simulated dynamics suggests a complex pat-
200 tern of convection and mixing (or mingling) whereby the original dense magma
201 rapidly mixes up with the volatile-rich, lighter magma coming from depth at
202 chamber bottom or dike level. The original phonolite carried down into the
203 dike mixes up with the shoshonite; the mixed magma is overall lighter than the
204 magma above it or in its immediate surroundings, therefore part of the initially

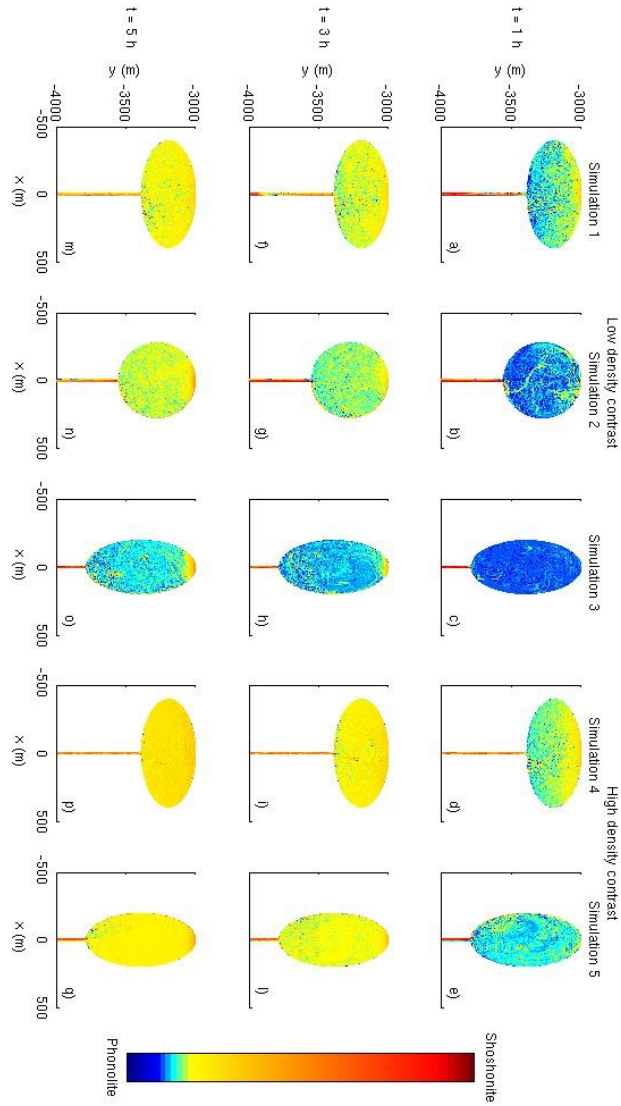


Figure 3: Space-time evolution of composition in the upper region of the simulated domain, for the different scenarios. Rows represent different times, columns represent different simulations.

205 sunk magma is carried up again into the shallow chamber, while other portions
 206 continue to sink down into the dike. This complex process originates a composi-
 207 tional, density, and gas volume stratification inside the chamber (Figure 2). The

208 original density contrast at chamber bottom is smoothed, so that the convective
209 process tends to slow down in time.

210 3.2 Exponential Trends and Time Scales

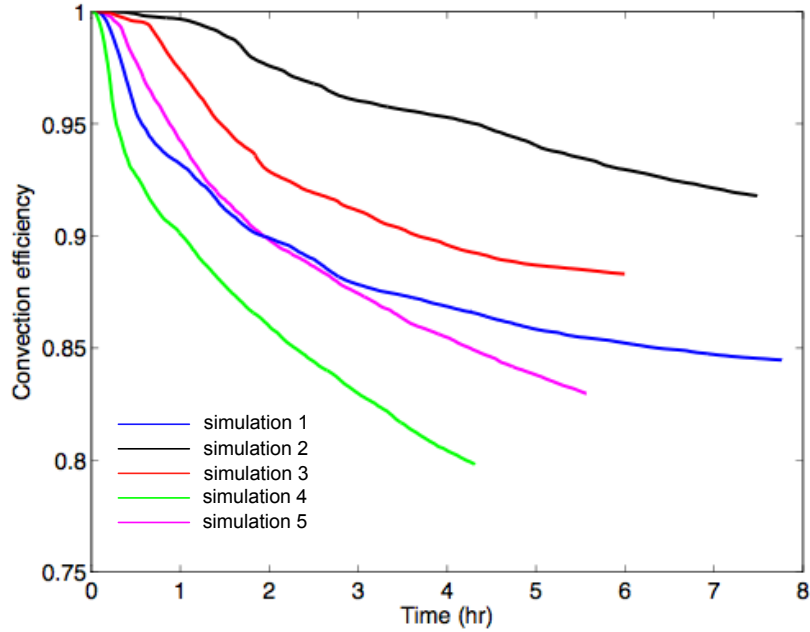


Figure 4: Time evolution of convection efficiency $\eta_C(t)$ for the different simulated scenarios.

211 Figure 4 shows the time evolution of the convection efficiency for all the sim-
212 ulations in Table 1. Convection and mixing slow down with time, as revealed
213 by the progressive decrease of the slopes. An initial vigorous phase leads to an
214 asymptotically decreasing efficiency of the mingling process. The density con-
215 trast among the entering and the resident magmatic mixtures becomes smaller
216 as the dynamics proceeds and mixing becomes intense both in the chamber and
217 in the feeding dike. More than 80% of the dense, degassed magma initially
218 residing in the shallow chamber remains at chamber level after waning of the
219 convective process. The complex patterns of convection and mixing are such
220 that most of the degassed magma is not carried down to larger depths when an
221 asymptotic stage has been reached. While shallow chamber degassing creates

222 unstable conditions by increasing the density of the magma at shallow levels,
 223 the ingress of small amounts of volatile-rich magma rapidly re-establishes a
 224 quasi-equilibrium dynamical condition, whereby small batches of mixed magma
 225 still enter the chamber while other batches sink into the dike (Figure 3).

226 Mixing efficiency in magmas has been hypothesized to follow and exponen-
 227 tially decreasing trend in time, when measured through the degree of homo-
 228 geneization (Section 3.2.1; Morgavi et al. (2013)). The asymptotically decreasing
 229 trend of convection efficiency in Figure 4 can be fitted well by an exponentially
 230 decaying function in time, given by

$$\eta_C = \eta_0 \exp(-t/\tau) + \eta_1. \quad (3)$$

231 The time evolution of the lumped quantities density and gas volume is also
 232 found to follow exponential trends for all the simulated scenarios, given by

$$\begin{aligned} \rho &= \rho_0 \exp(-t/\tau) + \rho_1, \\ \phi &= \phi_0 [1 - \exp(-t/\tau)]. \end{aligned} \quad (4)$$

233 In Equations (3) and (4) above, t is time and τ is the time constant; ρ and
 234 ϕ are density and gas volume fractions, respectively, and subscripts 0 and 1
 235 indicate constants. Figure 5 shows the best exponential fits for simulation 1;
 236 Table 3 lists the time constants obtained by fitting the data from the different
 237 simulations. The goodness of fit, measured by the coefficient of determination
 238 R^2 , is very good, as all fits have $R^2 > 0.99$. Every setup is characterized by a
 239 a single time constant, that is obtained from separately fitting all the relevant
 240 quantities (Figure 5).

241 The decay times obtained by the exponential fits range from 2 hours to
 242 13 hours. Sill-like reservoirs, as compared to more vertically elongated, dike-like
 243 geometries, reach the asymptotic state first, thus are characterized by shorter
 244 time constants; the density contrast plays a less relevant role in determining the
 245 replenishment time scale (Appendix A).

246 **3.2.1 Concentration variance decay**

247 The degree of homogenization of fluid mixtures can be measured by the single-
248 component concentration variance σ (Liu and Haller, 2004), defined by

$$\sigma^2(t) = \frac{\sum_{j=1}^N (C_j(t) - \mu)^2}{N}. \quad (5)$$

249 In the above, $C_j(t)$ is the time-varying concentration of the component con-
250 sidered, μ its mean concentration and N is the number of samples. As the
251 system approaches homogeneity with mixing, this quantity decreases towards
252 zero. Laboratory experiments of magma mixing have shown an exponential de-
253 crease of concentration variance in time (Morgavi et al., 2013). We evaluated
254 concentration variance and its time evolution for the 5 different simulation set-
255 ups by choosing a number of sample points $N = 155$ within the upper chamber.
256 We performed calculations for the composition of the shoshonitic end-member
257 X_S , as the composition of the phonolitic magma is defined as $X_P = 1 - X_S$. The
258 mean composition is thus $\mu_S = 0.875$, the average composition over the whole
259 domain (Figure 1). Figure 6 shows the results of concentration variance decay
260 calculations, alongside the fits to exponentially decaying functions. Concen-
261 tration variance decays exponentially in our numerical experiments; the decay
262 times are the same observed for the physical quantities described in Section 3.2
263 - mixing efficiency, density and gas volume fraction.

264 **3.3 Ground Deformation**

265 The stress variations at the boundaries of the fluid system due to magmatic
266 convection and mixing can be calculated from the results of the numerical sim-
267 ulations. These perturbations to the original stress state of the crust are prop-
268 agated to the Earth's surface, and can be recorded as ground deformation. To
269 evaluate the observable effects in terms of ground deformation due to the simu-
270 lated magmatic dynamics, as a first approximation we considered the crust as
271 an homogeneous, infinite medium, and calculated seismic wave propagation by
272 Green's functions integration.

273 Wave propagation is calculated by assuming continuity of pressure and stress
274 at the magma-rock interface. One-way coupling between fluid and confining rock
275 has proved a necessary approximation in order to limit computational needs;
276 full coupling could bring about resonance effects (e.g., Chouet, 1986) that are
277 expected to be less energetic than the modes related to the dynamics. The

278 assumption of homogeneous and infinite rock medium is expected not to play
279 a major role as the most energy associated to magma convection and mixing
280 is in the frequency band 10^{-4} - 10^{-3} Hz (Longo et al., 2012b; Bagagli et al.,
281 2017), corresponding to wavelengths of thousands of km. We are thus dealing
282 with quasi-static ground deformation that is not severely affected by medium
283 heterogeneities and finiteness.

284 Figure 7 shows the amplitude of the synthetic seismic signal relative to simu-
285 lation 1 at a station right above the top of the magmatic system, from which the
286 continuous component has been removed. The amplitude shows a decreasing
287 trend in time, that can be related to the waning magmatic dynamics under-
288 ground. The seismic amplitude decreases following very closely the convection
289 efficiency trend, thus with the same exponential behaviour described in Section
290 3.2.

291 This well-defined decay in seismic energy in the ULP frequency band can
292 be detected from ground deformation measurements obtained by tiltmeters and
293 strainmeters, providing a means to discriminate waning magma dynamics un-
294 derground and contributing to hazard assessment.

295 4 Discussion

296 Analysis of the dynamics of shallow magma chamber replenishment at Campi
297 Flegrei caldera shows that convection and mixing processes within magmatic
298 systems can be described by exponentially decreasing trends in time. An initial
299 vigorous phase of mass exchange among the incoming and the resident magmas is
300 followed by slowing interactions that asymptotically tend to the homogenization
301 of the system. The initial density contrast, which is the driving force behind the
302 overall dynamics, diminishes with increasing mixing time, causing the observed
303 behaviour.

304 The mathematical formulation of Equations (3) and (4) allows to define
305 precise time scales over which magma chamber replenishment processes are ex-
306 pected to be effective. Every simulated setup is thus characterized by a well-
307 defined and consistent time scale over which convective mingling processes are
308 effective. At Campi Flegrei, shallow chamber replenishment as depicted above
309 is thus characterized by typical mixing time scales in the range of some hours.
310 Following a similar approach, analysis of concentration variance decay in exper-
311 iments of magma mingling yields mingling-to-eruption time scales in the range

312 of tens of minutes (Perugini et al., 2015). Our results are in broad agreement
313 because explosive eruptions as those used as basis for the experimental setup are
314 expected to correspond to large ascent rates and more efficient mixing dynamics.

315 Eruptions can be triggered by chamber replenishment if the process is accom-
316 panied by a pressure increase in the magmatic system, that can cause rupture
317 of the host rock and initiate magma ascent towards the surface. The space-time
318 evolution of pressure in the case of buoyancy-driven replenishment analyzed here
319 is highly heterogeneous (Papale et al., 2017) and strongly depends on the initial
320 conditions of the system as shown in Figure 8, that reports lumped pressure
321 variations obtained from the model in the shallow reservoir. A pressure increase
322 is favoured by smaller density contrasts and less efficient dynamics; these cases
323 are also characterized by longer mixing time scales (Table 3). Albeit pressure
324 variations at the fluid boundary in the conditions explored in this work are
325 rather small, on the order of 10^{-1} MPa, they might be sufficient to initiate a
326 small-scale eruption in the fractured, low-tensile-strength shallow rock system
327 of Campi Flegrei (Krumbholz et al., 2014; Giudicepietro et al., 2017).

328 The results described above have necessarily been obtained under some as-
329 sumptions, mostly because of the large computational costs associated to the
330 detailed description of magmatic mixtures and their dynamical interactions de-
331 tailed in Section 2.1.

332 Extension to fully three-dimensional systems is expected not to change dras-
333 tically the time scales over which convection is efficient, given that the dynamical
334 interactions among the two magmas are mostly concentrated far from the bound-
335 aries (Bain et al., 2013). Assumed rigidity of the fluid boundary has implications
336 on the space-time evolution of pressure within the domain. A visco-elastic re-
337 sponse of the walls would have a buffering effect on the pressure variations, thus
338 on volatiles' exsolution in the chamber. Given the small pressure variations in
339 the magmatic system as a consequence of replenishment, though, that are in
340 the range of 0.1 Pa (Figure 8), the effect on chamber volume is expected to be
341 negligible. The assumptions of isothermal, crystal-free magmas stem from the
342 specific setting of Campi Flegrei: magmatic compositions are relatively mafic
343 and not too different among the two end-members (Table 2). Products with
344 the compositions used in our modeling have been found to have very little phe-
345 nocrysts, and the temperature differences inferred are negligible (Mangiacapra
346 et al., 2008). Solving the dynamics for an isothermal system, on the other
347 hand, greatly reduces computational needs as the energy conservation equation
348 does not need to be solved. The effect of wall cooling can be neglected on

349 the very short time scales analyzed here. Our model is thus meaningful for
350 melt-dominated, crystal-poor magma reservoirs, which is expected to be the
351 case for Campi Flegrei (Di Renzo et al., 2011). Long-lived magma chambers
352 can slowly solidify and originate mush zones, that, on the contrary, are domi-
353 nated by crystals (Pistone et al., 2017). Convection and mixing are common
354 in mushy reservoirs, too (Schleicher et al., 2016), but their features can depart
355 significantly from our results (Parmigiani et al., 2014).

356 Notwithstanding the enormous progress made towards reliable eruption fore-
357 casting, that is nowadays sometimes successful (Sparks, 2003), there is still no
358 widely accepted means of discriminating the precursor of an impending eruption,
359 especially at long-dormant volcanoes (Druitt et al., 2012). Our results offer in-
360 sights on the time scales of mixing among different magma types, that has been
361 invoked as eruption trigger at many volcanoes worldwide (e.g., Wark et al.,
362 2007; Druitt et al., 2012). Application to Campi Flegrei, one of the highest-risk
363 volcanic areas in the world, yields relatively short time scales from magma ar-
364 rival from depth to waning of shallow dynamics, which not necessarily leads to
365 eruption. As ground deformation signals related to magmatic convection can be
366 identified in strainmeter records (Bagagli et al., 2017), their duration and ampli-
367 tude modulation in time can be used for improved hazard assessment, possibly
368 discriminating between unrest phases leading or not to eruptive activity.

369 **Appendix A: Convective Overturn in Boundary** 370 **Layer Theory**

371 A rough estimate of the time scales over which convection in magma chambers
372 as described above is efficient can be obtained by evaluating the overturn time
373 predicted by boundary layer theory (e.g., Acheson, 1990).

374 The boundary layer theory postulates that the changes in composition that
375 cause the unstable density profile are confined at the bottom of the shallow
376 chamber, until the gravitational instability grows disrupting the system and
377 causing overturn of less dense layers. The criterion for the onset of instability
378 is that the Rayleigh number Ra be larger than 10^3 . The Rayleigh number is a
379 non-dimensional quantity indicating the ratio of buoyancy and viscosity forces
380 multiplied by the ratio of momentum and chemical diffusivities. For the setup

381 used in our simulations, the Rayleigh number can be defined as

$$\text{Ra} = \frac{g\Delta\rho L^3}{\mu\kappa}, \quad (6)$$

382 where g is gravity acceleration, $\Delta\rho$ is the density difference at the interface
383 among the two magmas, L is the vertical length scale of the reservoir, μ is
384 viscosity and κ is chemical diffusion coefficient.

385 A local Rayleigh number can be defined, in which the length scale is that of
386 the boundary layer δ , over which most of the compositional change occurs:

$$\text{Ra}_\delta = \frac{g\Delta\rho\delta^3}{\mu\kappa}. \quad (7)$$

387 The boundary layer thickness increases in time due to chemical diffusion:

$$\delta = \sqrt{\kappa t}, \quad (8)$$

388 thus the local Rayleigh number increases in time as well. Convective overturn
389 is predicted as the critical Rayleigh number $\text{Ra}_\delta^C \sim 10^3$ is reached, thus at a
390 time t_O given by

$$t_O \sim (\text{Ra}_\delta^C)^{2/3} \frac{\mu^{2/3}}{g^{2/3}\Delta\rho^{2/3}\kappa^{1/3}}. \quad (9)$$

391 Substituting the appropriate quantities for our simulations, $\mu \sim 500 \text{ Pa}\cdot\text{s}$,
392 $\Delta\rho \sim 100 \text{ kg/m}^3$ and $\kappa \sim 10^{-5} \text{ m}^2\text{s}^{-1}$, yields an overturn time of the order of
393 1 h.

394 This first-order estimate is of the same order of magnitude of the time scales
395 obtained by the exponential fits above; nonetheless, the decaying trend of mix-
396 ing efficiency cannot be incorporated in this theoretical framework. This same
397 theory also confirms that longer homogenization times are related to lower initial
398 density contrasts, as expected for buoyancy-only driven dynamics.

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Simulation no	Shoshonite		Phonolite		Shallow chamber geometry	Density contrast at initial interface (kg/m^3)
	CO_2 , (wt%)	H_2O , (wt%)	CO_2 , (wt%)	H_2O , (wt%)		
1	1	2	0.3	2.5	Oblate	30
2	1	2	0.3	2.5	Circular	25
3	1	2	0.3	2.5	Prolate	20
4	1	2	0.1	1	Oblate	120
5	1	2	0.1	1	Prolate	140

Table 1: Simulations performed and corresponding initial conditions.

	SiO ₂	TiO ₂	AlO ₂	Fe ₂ O ₃	FeO	MnO	MgO	CaO	Na ₂ O	K ₂ O
shoshonite	0.5247	0.0085	0.1760	0.0188	0.0574	0.0012	0.0360	0.0793	0.0343	0.0428
phonolite	0.5352	0.0060	0.1984	0.0160	0.0320	0.0014	0.0176	0.0676	0.0466	0.0791

Table 2: Oxides composition for the two magmas.

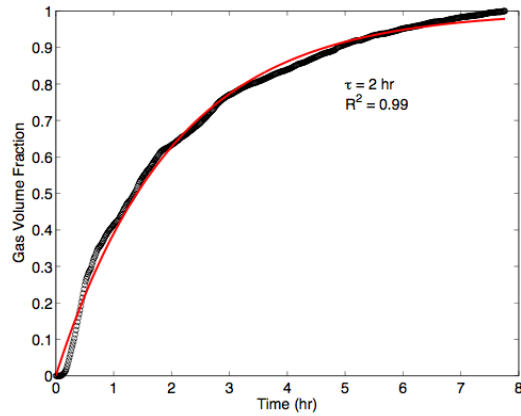
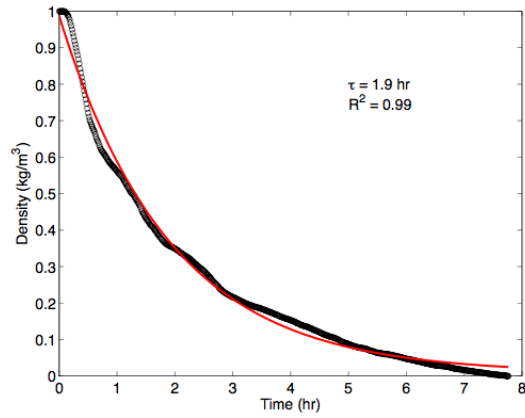
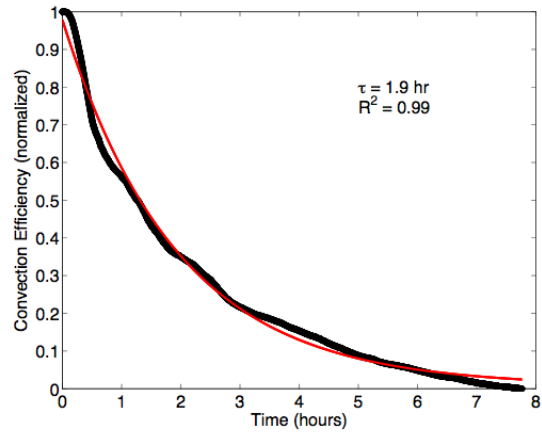


Figure 5: Exponential fits for a) convection efficiency, b) density and c) gas volume fraction for simulation 1.

Simulation	Time constant, hours
1	2.0
2	2.6
3	13
4	2.2
5	2.4

Table 3: Exponential decay times for the five simulated setups.

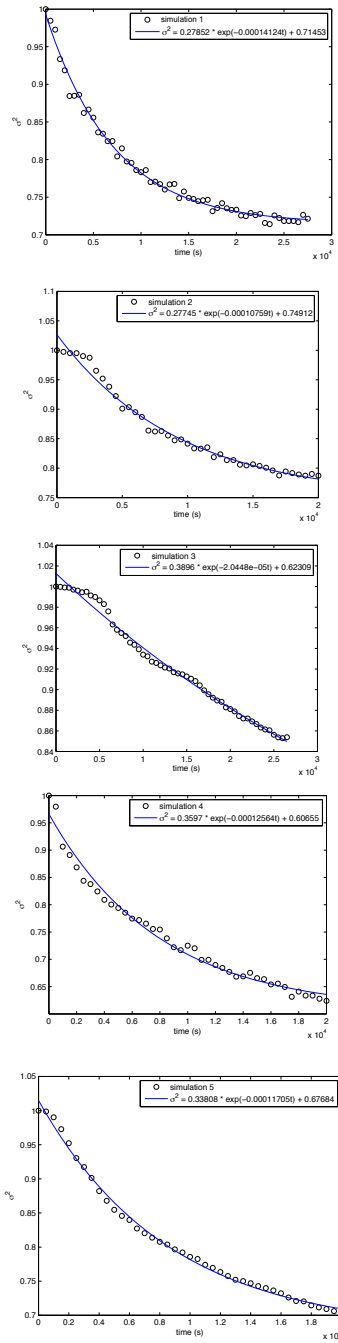


Figure 6: Concentration variance decay with time for the five simulated setups, alongside exponential fits.

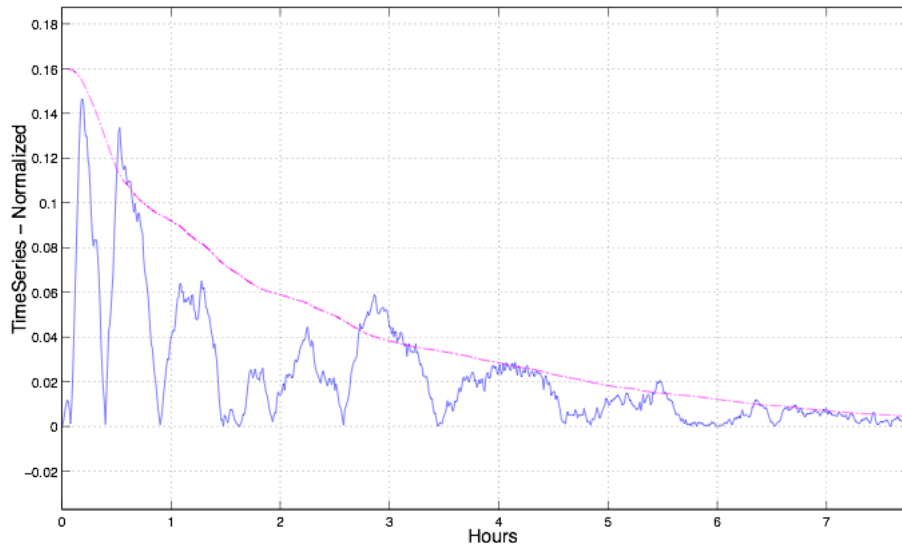


Figure 7: For simulation 1, amplitude of ground deformation at a synthetic station atop the magmatic system, in blue; superimposed is the convection efficiency, in violet.

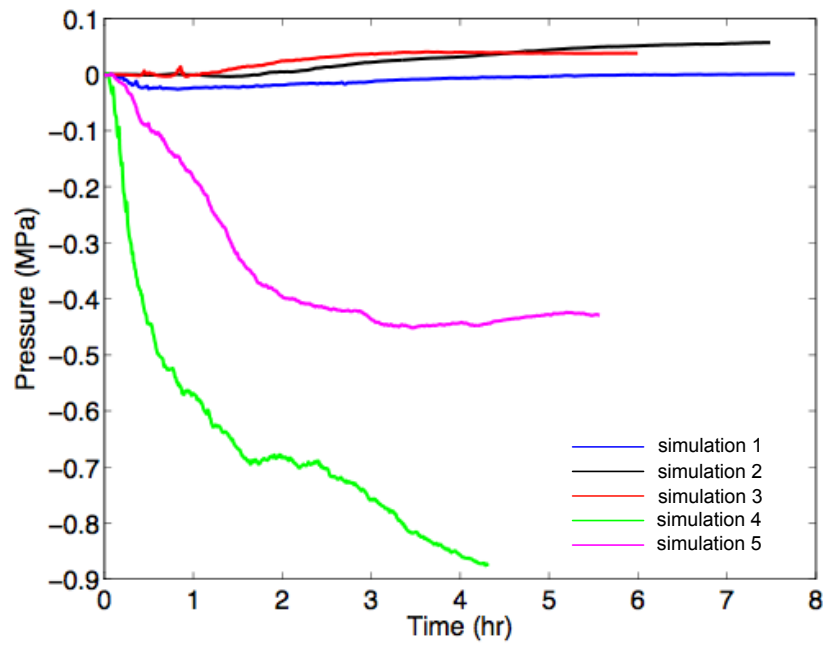


Figure 8: Pressure evolution at the top of the magmatic system for the 5 simulated scenarios.