

Revisiting the surface-energy-flux perspective on the sensitivity of global precipitation to climate change

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1 **Abstract** Climate models simulate an increase in global precipitation at a rate of approxi-
2 mately 1-3% per Kelvin of global surface warming. This change is often interpreted through
3 the lens of the atmospheric energy budget, in which the increase in global precipitation is
4 mostly offset by an increase in net radiative cooling. Other studies have provided differ-
5 ent interpretations from the perspective of the surface, where evaporation represents the
6 turbulent transfer of latent heat to the atmosphere. Expanding on this surface perspective,
7 here we derive a version of the Penman-Monteith equation that allows the change in ocean
8 evaporation to be partitioned into a thermodynamic response to surface warming, and ad-
9 ditional diagnostic contributions from changes in surface radiation, ocean heat uptake, and
10 boundary-layer dynamics/relative humidity. In this framework, temperature is found to be
11 the primary control on the rate of increase in global precipitation within model simulations
12 of greenhouse gas warming, while the contributions from changes in surface radiation and
13 ocean heat uptake are found to be secondary. The temperature contribution also dominates
14 the spatial pattern of global evaporation change, leading to the largest fractional increases
15 at high latitudes. In the surface energy budget, the thermodynamic increase in evaporation
16 comes at the expense of the sensible heat flux, while radiative changes cause the sensible
17 heat flux to increase. These tendencies on the sensible heat flux partly offset each other,
18 resulting in a relatively small change in the global mean, and contributing to an impression
19 that global precipitation is radiatively constrained.

20 **Keywords** Hydrologic Cycle · Global Warming

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21 1 Introduction

22 In Earth’s hydrologic cycle, water evaporates from the surface, condenses in the atmosphere,
 23 and returns to the surface as precipitation. It takes an average water molecule 8-10 days to
 24 complete this cycle (e.g., Van Der Ent and Tuinenburg, 2017). On timescales much longer
 25 than this, the rates of global-mean evaporation and precipitation are essentially equal. Be-
 26 cause of the latent heat absorbed and released during evaporation and condensation, the hy-
 27 drologic cycle plays an important role in the heat engine of the climate system, transferring
 28 energy from the warm surface, where most sunlight is absorbed, to the cooler atmosphere,
 29 where infrared radiation is emitted back to space.

30 In response to CO₂-induced warming, climate models predict that the intensity of the
 31 global hydrologic cycle (i.e., global-mean evaporation/precipitation) will increase by around
 32 1-3%/K, which is significantly less than the approximately 7%/K increase in atmospheric
 33 water vapor resulting from the Clausius-Clapeyron equation and near-constant relative hu-
 34 midity (e.g., Boer, 1993; Allen and Ingram, 2002). This disparity has been invoked to
 35 explain important aspects of the climate response to CO₂-induced warming, including a
 36 slowing down of the atmospheric circulation (Held and Soden, 2006), and an increase in
 37 the frequency and intensity of floods and droughts (Allen and Ingram, 2002; Trenberth,
 38 1999, 2011). Understandably, therefore, the relationship between surface temperature and
 39 the global hydrologic cycle has been a central topic in climate science for decades (e.g.,
 40 Manabe and Wetherald, 1975)

41 Many previous studies have investigated the change in global precipitation with warming
 42 from the perspective of the global-mean atmospheric energy budget,

$$L\bar{P} = \bar{R}_a - \bar{H}, \quad (1)$$

43 where L is the latent heat of vaporization, P is the rate of precipitation, R_a is the net heat
 44 lost through radiation, H is the sensible heat flux from the surface to the atmosphere, and
 45 $(\bar{\cdot})$ indicates the global mean of a quantity (e.g., Allen and Ingram, 2002; Stephens and
 46 Ellis, 2008; Previdi, 2010; Pendergrass and Hartmann, 2014; Fläschner et al, 2016). Eq.
 47 1 implies that any increase in $L\bar{P}$ must be offset by a decrease in \bar{H} and/or an increase
 48 in \bar{R}_a . In most simulations of CO₂-induced warming by GCMs, the change in \bar{H} tends to
 49 be small compared with changes in $L\bar{P}$ and \bar{R}_a , with the latter dominated by an increase in
 50 longwave emissions (e.g., Lambert and Webb, 2008; Pendergrass and Hartmann, 2014). This
 51 result has sometimes been interpreted as evidence that the change in global precipitation is
 52 primarily determined by the ability of the atmosphere to radiate more energy as it warms
 53 (e.g., Allen and Ingram, 2002). Yet in most cases, there is not a one-to-one trade-off between
 54 changes in \bar{R}_a and $L\bar{P}$, because changes in \bar{H} are often not negligible. For example, among
 55 GCMs participating in the most recent Coupled Model Intercomparison Project (CMIP5),
 56 Pendergrass and Hartmann (2014) found that the change in \bar{H} was small but significant, with
 57 a magnitude about a third as large as the change in $L\bar{P}$. Simulations spanning a wider range
 58 of climate states have also been found to exhibit large variability in the change in \bar{H} with
 59 global temperature (O’Gorman and Schneider, 2008; O’Gorman et al, 2012). On its own,
 60 however, the atmospheric energy budget provides limited insight into how the changes in \bar{H} ,
 61 $L\bar{P}$, and \bar{R}_a are partitioned.

62 An alternative to the atmospheric energy budget perspective is to treat the hydrologic
 63 cycle as a turbulence-driven process, in which evaporation, E , represents the turbulent flux
 64 of water vapor from the surface to the near-surface atmosphere (e.g., Penman, 1948; Priestly
 65 and Taylor, 1972; Monteith, 1981; Pierrehumbert, 2002; Richter and Xie, 2008; Lorenz et al,

2010; Pierrehumbert, 2010). Over the oceans, where 85% of global evaporation occurs (e.g.,
 66 Trenberth et al, 2007), the rate of evaporation is given approximately by the bulk transfer
 67 equation,
 68

$$E = [q^*(T_s) - rq^*(T_a)]\rho C_H u \quad (2)$$

69 where q^* is the specific humidity at saturation, T_s and T_a are the temperatures of the ocean
 70 surface and near-surface atmosphere, r is the near-surface relative humidity, ρ is the near-
 71 surface air density, C_H is the bulk transfer coefficient, and u is the near-surface wind speed.
 72 Similarly, the sensible heat flux is approximately given by

$$H = c_p(T_s - T_a)\rho C_H u, \quad (3)$$

73 where c_p is the specific heat capacity of air, and C_H is typically assumed to have the same
 74 value as in Eq. 2. Eqs. 2 and 3 are derived from Monin-Obukhov similarity theory (e.g.,
 75 Pierrehumbert, 2010), and they imply that the fluxes of latent and sensible heat are deter-
 76 mined by the speed, temperature, and relative humidity of the near-surface winds, and by
 77 the difference in temperature between the ocean surface and the near-surface atmosphere.¹
 78 While these variables are all strongly dependent on atmospheric physics, collectively they
 79 must satisfy the surface energy budget,

$$LE + H = R_s - G, \quad (4)$$

80 where L , E , and H are defined as in Eq. 1, R_s is the net downward radiation flux at the
 81 surface, and G is the rate of heat storage by the ocean plus—at high latitudes—the heat
 82 required to melt frozen hydrometeors that reach the surface. Unlike Eq. 1, Eqs. 2-4 are valid
 83 at a specific location, and not just in the global mean.

84 With three equations instead of one, the turbulent-flux perspective provides insight into
 85 the partitioning between H and LE that is not possible with the atmospheric energy budget
 86 alone. For example, Pierrehumbert (2002) has shown that one implication of Eqs. 2-4 is
 87 that $LE \ll H$ at very cold temperatures due to low values of q^* , while $LE \gg H$ at very
 88 warm temperatures due to high values of q^* , and correspondingly weak (or even negative)
 89 values of $T_s - T_a$, which are necessary to satisfy the surface energy budget (e.g., Le Hir et al,
 90 2009). Interpolating between these two extremes, it must be true that, in more moderate
 91 climates like that of the present day, warming will tend to cause LE to increase (via an
 92 increase in q^*), and H to decrease (via a decrease in $T_s - T_a$). However, while such changes
 93 are easily diagnosed in GCM simulations (e.g., Richter and Xie, 2008; Lorenz et al, 2010),
 94 they are difficult to quantify from first principles, since they also depend on changes in net
 95 surface radiation, ocean heat uptake, and boundary-layer dynamics, which are only weakly
 96 constrained by Eqs. 2-4.

97 As we argue in this paper, however, the turbulent-flux perspective may have more ex-
 98 planatory power than previously assumed. By combining Eqs. 2-4 into a variant of the
 99 Penman-Monteith equation (Penman, 1948; Monteith, 1981), we show that the change in
 100 evaporation can be partitioned into four distinct terms, each with a straightforward physi-
 101 cal interpretation. The first term depends only on surface temperature, and represents the
 102 thermodynamic response to warming, which is independent of other changes in the climate
 103 system. Meanwhile, the other terms represent the change in evaporation due to changes in
 104 surface radiation, ocean heat uptake, and boundary-layer dynamics/relative humidity. These
 105 terms cannot be derived from first principles, and must therefore be diagnosed from GCM

¹ On long timescales, $T_s - T_a$ is generally positive over the ocean, implying a transfer of both sensible and latent heat from the surface to the atmosphere.

106 simulations. Nevertheless, if the spatial patterns of changes in surface radiation and ocean
 107 heat uptake are known, the Penman-Monteith equation provides a way to quantify the re-
 108 sulting spatial pattern of the change in evaporation. When extended to the global mean, the
 109 Penman-Monteith framework also sheds light on the partitioning between changes in \bar{H} and
 110 \overline{LE} in the surface energy budget (or LP in the atmospheric energy budget).

111 The paper is organized as follows. We begin in Section 2 by deriving a variant of the
 112 Penman-Monteith equation, which governs the rate of evaporation over the oceans subject
 113 to energetic constraints. In Section 3, we use this equation to diagnose the contributions to
 114 evaporation change in GCM simulations of global warming. In this diagnostic framework,
 115 we find that most of the increase in global evaporation is a direct consequence of warmer
 116 temperatures, while changes in surface radiation and ocean heat uptake play a secondary
 117 role. Based on these results, we then derive an approximation for evaporation change as a
 118 function of changes in temperature, surface radiation, and ocean heat uptake, and show that
 119 it accurately represents the fast and slow responses of evaporation to CO_2 forcing, while
 120 also shedding light on the partitioning between changes in the latent and sensible heat flux.
 121 In Section 4, we apply the Penman-Monteith framework to a series of idealized simulations
 122 run by O’Gorman and Schneider (2008), and find that thermodynamics alone can account
 123 for much of the changes in global precipitation among the simulations. We conclude with a
 124 brief summary and discussion in Section 5.

125 2 Derivation and interpretation of the Penman-Monteith equation.

126 We build upon a long history of research on the physics of ocean evaporation, beginning
 127 with the fundamental equations (2-4) that govern the exchange of energy between the ocean
 128 surface and the atmosphere (Penman, 1948; Monteith, 1981; Pierrehumbert, 2002). In par-
 129 ticular, we seek to solve this system of equations for E while eliminating $T_s - T_a$ and H . To
 130 do so, we first approximate $q^*(T)$ as the first-order Taylor expansion about the point $T = T_a$:

$$q^*(T) \approx q^*(T_a) + \frac{dq^*}{dT} [T - T_a]. \quad (5)$$

131 Given the Clausius-Clapeyron relation,

$$\frac{dq^*}{dT} = \alpha q^*, \quad (6)$$

132 Eq. 5 implies that

$$q^*(T_s) \approx q^*(T_a)(1 + \alpha[T_s - T_a]), \quad (7)$$

133 where

$$\alpha = \frac{L}{R_v T_a^2} \quad (8)$$

134 is the Clausius-Clapeyron scaling factor, with R_v representing the specific gas constant for
 135 water vapor. Given Eq. 7, we can express the air-sea moisture difference in Eq. 2 as a func-
 136 tion of r , T_a , and $T_s - T_a$:

$$E \approx q^*(T_a)(1 - r + \alpha[T_s - T_a])\rho C_H u. \quad (9)$$

137 This allows us to eliminate $T_s - T_a$ and H from Eqs. 2-4 to arrive at the following expression
 138 for the rate of evaporation from the ocean surface:

$$LE \approx \eta(R_s - G + \kappa), \quad (10)$$

139 where η and κ are defined as

$$\eta \equiv \frac{1}{1 + \beta_0}, \quad (11)$$

140

$$\kappa \equiv (1 - r)c_p\rho C_H u \alpha^{-1}, \quad (12)$$

141 with

$$\beta_0 = \frac{c_p}{\alpha L q^*(T_a)} \quad (13)$$

142 representing the Bowen ratio (H/LE) in the limit of 100% near-surface relative humidity
143 ($r = 1$).

144 Eq. 10 is equivalent to the Penman-Monteith equation for terrestrial evapotranspiration
145 from a saturated surface (i.e., where stomatal resistance is zero) (e.g., Scheff and Frierson,
146 2014). As written above, however, Eq. 10 is much easier to interpret than its more conven-
147 tional form: besides R_s and G , it comprises just two terms (η and κ), each of which has a
148 straightforward physical meaning.

149 The first term, η , is a function of T_a alone. In the context of climate change, therefore, η
150 captures the thermodynamic response of evaporation to warming, which is independent of
151 changes in other variables like wind speed, relative humidity, and net surface radiation. On
152 the other hand, η does implicitly involve $T_s - T_a$ and H , as the following thought experiment
153 illustrates.

154 Let us assume that T_s and T_a were to increase by the same amount, while all other
155 variables in Eqs. 2-4 were held constant. In this scenario, H would not change, while LE
156 would increase with T_a at the same rate as $q^*(T_a)$ —i.e., at the Clausius-Clapeyron rate of
157 $\alpha \approx 7\%/K$ (Eq. 9; Richter and Xie, 2008). But without compensating changes in $R_s - G$,
158 such an increase in LE would clearly violate the surface energy budget. In reality, therefore,
159 $T_s - T_a$ must decrease with warming, such that H decreases and LE increases at a rate less
160 than $7\%/K$. This constraint on $T_s - T_a$ has been described in similar terms by Pierrehumbert
161 (2002) and Lorenz et al (2010), but in Eq. 10, it is implicitly captured by a single variable,
162 η .

163 Meanwhile, κ represents the dependence of evaporation on relative humidity (r) and
164 boundary-layer dynamics (u and C_H), while its temperature-dependence (via α in its de-
165 nominator) is negligibly small. An increase in r will reduce the air-sea moisture difference,
166 thereby causing LE to decrease. Yet in the absence of changes in $R_s - G$, this decrease must
167 be accompanied by an increase in $T_s - T_a$ to ensure no change in $LE + H$. Likewise, an in-
168 crease in u or C_H will cause both LE and H to increase, and thus requires a simultaneous
169 decrease in $T_s - T_a$ to satisfy the surface energy budget.

170 These thought experiments illustrate the crucial role played by $T_s - T_a$ within the Penman-
171 Monteith framework (Eqs. 2-4). In response to any environmental change that, in isolation,
172 would violate the surface energy budget, $T_s - T_a$ must adjust to conserve energy and, in the
173 process, alter the partitioning between LE and H . Physically, this framework is consistent
174 with previous studies that have emphasized the role of $T_s - T_a$ in limiting the rate of increase
175 in global precipitation with surface warming (Pierrehumbert, 2002; Richter and Xie, 2008;
176 Lorenz et al, 2010). However, the advantage of the Penman-Monteith surface-energy per-
177 spective developed here is that it allows the change in evaporation with surface warming to
178 be partitioned into a predictable component due to changes in η , and a diagnostic component
179 due to changes in surface radiation, ocean heat uptake, and relative humidity/dynamics. In
180 the next section, we show that this partitioning provides insight into the factors controlling
181 the change in global and regional evaporation in response to CO_2 -induced warming.

3 Results from comprehensive GCMs

In this section, we use the Penman-Monteith framework (Eq. 10) to quantify the factors contributing to the change in ocean evaporation in GCM simulations of CO₂-induced warming. We examine both the equilibrium response to an abrupt doubling of atmospheric CO₂, as simulated by an ensemble of atmosphere-only GCMs with slab oceans, and the transient response a century after an abrupt quadrupling of atmospheric CO₂, as simulated by an ensemble of coupled atmosphere-ocean GCMs. The transient simulations were performed as part of the most recent Coupled Model Intercomparison Project (CMIP5; Taylor et al, 2012), while the equilibrium simulations were performed with a previous generation of climate models included in CMIP3² (Meehl et al, 2007). In the equilibrium (slab ocean) case, we focus on the change in ocean evaporation between the last five years of the pre-industrial control simulations and years 21-25 after CO₂ doubling. The transient results reflect the change in evaporation between the pre-industrial control simulations and years 96-100 after CO₂ quadrupling.

Tables 1 and 2 list the names of the 10 CMIP3 models and 12 CMIP5 models included in our analysis. Some models were excluded because the required variables were not readily available, or in the CMIP3 case, because the simulations were not in radiative equilibrium. For consistency with our later analysis in Section 3.1, we further restrict the CMIP5 models to those that also performed CO₂ quadrupling simulations with prescribed climatological sea-surface temperatures (SSTs). The second column of each table gives the percent rate of change in global evaporation per Kelvin of global warming, while the third column gives the equivalent rate over the oceans (i.e., using ocean-mean instead of global-mean values of ΔE , E , and ΔT). Even though the rate of change in ocean evaporation exceeds the global change in all models, the two values are highly correlated across models ($r = 0.93$ and 0.81 for CMIP3 and CMIP5, respectively). This high correlation is not surprising, considering that the oceans account for 85% of global evaporation in the current climate (Trenberth et al, 2007), and an even larger share of the increase in evaporation under global warming (e.g., Fu and Feng, 2014). To first order, therefore, we can understand the change in global evaporation (and thus precipitation) by focusing on the ocean-mean changes.

Fig. 1 shows the fractional change in ocean evaporation per Kelvin of local warming in the ensemble mean of the equilibrium (left column) and transient (right column) simulations, along with the individual contributions to evaporation change due to changes in η , R_s , G , and κ (rows 2-5). We have calculated these contributions using the discrete form of the fractional derivative of Eq. 10 with respect to T_a :

$$\frac{1}{\Delta T_a} \left[\frac{\overbrace{\Delta E}^1}{E} \approx \frac{\overbrace{\Delta \eta}^2}{\eta} + \frac{\eta}{LE} (\overbrace{\Delta R_s}^3 - \overbrace{\Delta G}^4 + \overbrace{\Delta \kappa}^5) \right], \quad (14)$$

where Δ indicates the change in a variable between the control and warmed climate. The number above each term in Eq. 14 indicates its order, from top to bottom, in Fig. 1. Each contribution was calculated using monthly-mean model output (see Appendix for details). The ocean-mean of each contribution is given in the top left corner of each panel. Columns 4-7 of Tables 1 and 2 give the equivalent values for each ensemble member, along with the standard deviation of each contribution across the ensemble. In this Section we discuss only the ensemble-mean results (Fig. 1), but will address the inter-model variability in Section 5.

² Equilibrium simulations were not part of CMIP5.

223 In both the equilibrium and transient cases, we find that $\Delta\eta$ accounts for the largest
 224 contribution to evaporation change, increasing from about 1.2%/K at the equator to more
 225 than 4%/K at higher latitudes. To understand this meridional structure—which is also re-
 226 flected in the pattern of total evaporation change—consider the analytic expression for the
 227 $\Delta\eta$ contribution, which can be derived from Eqs. 6, 8, 11, and 13:

$$\frac{d\ln\eta}{dT_a} = \frac{\beta_0}{1+\beta_0} \left(\alpha - \frac{2}{T_a} \right). \quad (15)$$

228 Because β_0 decreases almost exponentially with increasing temperature, $d\ln\eta/dT_a$ is largest
 229 at cold temperatures where it approaches $(\alpha - 2/T_a)$ —not much less than the Clausius-
 230 Clapeyron scaling of atmospheric water vapor under fixed relative humidity. As T_a increases,
 231 however, the gap between α and $d\ln\eta/dT_a$ grows ever larger. The increase in $d\ln\eta/dT_a$
 232 with latitude therefore reflects the meridional structure of T_a in the control climate. In the
 233 global mean, $\Delta\eta$ causes an increase in evaporation of about 1.5% per Kelvin of global
 234 warming in both the equilibrium and transient cases, accounting for the largest fraction of
 235 the total increase.³

236 Compared with $\Delta\eta$, the contributions from the other terms in Eq. 14 are generally
 237 smaller in magnitude. ΔR_s accounts for the second largest contribution to evaporation change
 238 (Fig. 1, row 3), increasing evaporation by about 1%/K globally in both the equilibrium and
 239 transient cases due to an increase in net surface radiation in the tropics and midlatitudes.⁴
 240 In contrast, the contribution from ΔG is negligible in the equilibrium (slab-ocean) simula-
 241 tions, but significantly negative in the transient simulations due to ocean heat uptake in the
 242 subpolar North Atlantic Ocean and Southern Ocean (row 4). Finally, the contribution from
 243 $\Delta\kappa$ tends to be negative but quite weak (row 5), implying that changes in boundary-layer
 244 dynamics and relative humidity play a minor role, particularly in determining the spatial
 245 pattern of evaporation change.

246 Fig. 1 reveals two key points about how ocean evaporation responds to CO₂-induced
 247 warming in GCMs. First, the largest contribution to evaporation change comes from $\Delta\eta$,
 248 which is a direct consequence of warmer temperatures (Eq. 15). This means that much of
 249 the increase in ocean evaporation with warming—both globally and regionally—is inde-
 250 pendent of the nature of both the forcing that drove the temperature change (e.g., aerosols
 251 vs. greenhouse gases) and how certain physical processes are parameterized within a model
 252 (e.g., boundary-layer dynamics, convection, clouds, etc.).

253 Second, the small contribution from $\Delta\kappa$ suggests that changes in boundary layer dynam-
 254 ics and relative humidity play a limited role in determining the change in ocean evaporation
 255 in response to CO₂-induced warming, particularly at regional scales. We know of no *a pri-*
 256 *ori* reason why this should be the case, since in principle the boundary layer could adjust in
 257 any number of ways that would alter the surface energy balance (e.g., Pierrehumbert, 2002).
 258 However, if we take as given that the contribution from $\Delta\kappa$ is small, the total change in
 259 evaporation can then be approximated as

$$L\Delta E \approx LE \frac{d\ln\eta}{dT_a} \Delta T_a + \eta \Delta(R_s - G), \quad (16)$$

³ The global-mean contributions represent the average of the fractional changes (Fig. 1), weighted by the product of mean-state evaporation and local temperature change. Further details are provided in the Appendix.

⁴ The contribution from ΔR_s is negative at high latitudes in the Southern Hemisphere, reflecting a decrease in shortwave absorption as a result of increased cloud cover.

260 where the first term represents the thermodynamic response to warming (Eq. 15), and the
 261 second term represents the diagnostic component of evaporation change due to the combined
 262 changes in surface radiation and ocean heat uptake.

263 3.1 Testing Eq. 16 on the fast and slow responses of evaporation to CO₂ forcing.

264 Although $\Delta\kappa$ has little impact on evaporation in Fig. 1, it is reasonable to question the
 265 generality of this result. For example, it could be that $\Delta\kappa$ is not physically independent
 266 of the other terms in Eq. 14, in which case its small contribution in Fig. 1 might reflect a
 267 cancellation between larger competing effects driven by changes in η and $R_s - G$. If true,
 268 this would imply that Eq. 16 does not generally hold.

269 To test the robustness of the approximation in Eq. 16, therefore, we revisit the coupled
 270 CMIP5 simulations analyzed previously, but now with the change in evaporation partitioned
 271 into “fast” and “slow” components that represent, respectively, the direct response of evap-
 272 oration to CO₂ quadrupling with fixed SSTs, and the more gradual changes that occur as
 273 the climate warms. Following previous studies (e.g., Lambert and Faull, 2007; Lambert and
 274 Webb, 2008; Lambert et al, 2009; Andrews et al, 2009, 2010; Andrews and Forster, 2010;
 275 Frieler et al, 2011; Samset et al, 2016; Fläschner et al, 2016), we define the fast response as
 276 the change that occurs when CO₂ is quadrupled and SSTs are fixed at pre-industrial values
 277 (i.e., the difference between sstClim4xCO2 and sstClim experiments in CMIP5 parlance),
 278 and the slow response as the difference between the coupled 4xCO2 simulations and the
 279 fixed-SST 4xCO2 simulations. Defined in this way, the sum of the slow and fast responses
 280 gives the total change in the top right panel of Fig. 1.

281 In the fast response to CO₂ forcing, an increase in the longwave optical depth of the
 282 atmosphere will cause an increase in net surface radiation (e.g., Allen and Ingram, 2002;
 283 McInerney and Moyer, 2012) and a decrease in outgoing longwave radiation. The latter ef-
 284 fect adds heat to the climate system, which is mostly absorbed by the ocean (thus increasing
 285 G). Meanwhile, little warming occurs over the oceans because T_s is held constant, imply-
 286 ing that the first term on the RHS of Eq. 16 will be small. As a result, the fast change in
 287 evaporation is approximately given by

$$L\Delta E_{\text{fast}} \approx \eta \Delta(R_s - G)_{\text{fast}}, \quad (17)$$

288 where $\Delta(R_s - G)_{\text{fast}}$ represents the change in $R_s - G$ that is directly caused by CO₂ forcing,
 289 independent of surface warming.

290 On longer timescales, T_a will gradually rise, impacting ocean evaporation in two ways.
 291 First, there will be a thermodynamic response represented by the first term on the RHS of Eq.
 292 16. Second, R_s and G will also change as the atmosphere warms and the earth approaches
 293 top-of-atmosphere radiative balance (e.g., Pierrehumbert, 1999). Applying the regression
 294 method of Gregory et al (2004) at each grid point, we find that the slow change in $R_s - G$ over
 295 most of the global oceans is well approximated in all models as a linear surface-temperature
 296 feedback,

$$\Delta(R_s - G)_{\text{slow}} \approx \lambda \Delta T_a, \quad (18)$$

297 where λ represents the slope of the regression. The slow response of evaporation can then
 298 be expressed as

$$L\Delta E_{\text{slow}} \approx \left(LE \frac{d \ln \eta}{dT_a} + \eta \lambda \right) \Delta T_a. \quad (19)$$

299 Fig. 2 shows the approximations in Eqs. 17 and 19 (top row) along with the actual fast
 300 and slow changes in evaporation from the ensemble mean of the CMIP5 simulations (bottom
 301 row). Comparing the two rows, we find that the approximations capture the spatial pattern of
 302 evaporation change remarkably well, with spatial correlations of $r > 0.99$ in both cases. The
 303 approximations are also quite accurate at the global scale, deviating from the actual change
 304 in ocean-mean evaporation by 16% and 4% in the fast and slow cases, respectively.⁵ These
 305 results show that the contribution from $\Delta \kappa$ is small even when $\Delta(R_s - G)$ is independent of
 306 surface warming, suggesting that Eq. 16 is robust within GCM simulations of CO₂-induced
 307 warming.

308 3.2 Implications of the Penman-Monteith perspective for changes in the sensible heat flux.

309 Having demonstrated the approximate validity of Eq. 16, let us now consider its implications
 310 for the change in sensible heat flux (ΔH). When combined with the surface energy budget
 311 (Eq. 4), Eq. 16 implies that

$$\Delta H \approx -LE \frac{d \ln \eta}{dT_a} \Delta T_a + (1 - \eta) \Delta(R_s - G), \quad (20)$$

312 which can be further decomposed into fast and slow components, following the same line of
 313 reasoning used to derive Eqs. 17 and 19:

$$\Delta H_{\text{fast}} \approx (1 - \eta) \Delta(R_s - G)_{\text{fast}}, \quad (21)$$

314

$$\Delta H_{\text{slow}} \approx \left(-LE \frac{d \ln \eta}{dT_a} + (1 - \eta) \lambda \right) \Delta T_a. \quad (22)$$

315 These equations provide insight into how $L\Delta E$ and ΔH are partitioned in the fast and
 316 slow responses to CO₂ warming. In the fast case, Eqs. 17 and 21 imply that

$$\frac{\Delta H_{\text{fast}}}{L\Delta E_{\text{fast}}} \approx \frac{1 - \eta}{\eta} = \beta_0, \quad (23)$$

317 which is a function of T_a alone (Eq. 13). In the current climate, β_0 ranges from around 0.25
 318 in the tropical warm pool to more than 1.5 at high latitudes, with an ocean-mean value of
 319 around 0.5. Even though this result is specific to the oceans, it helps explain why in the fast
 320 response to CO₂ forcing, $L\Delta \bar{E}$ significantly exceeds $\Delta \bar{H}$ in the global mean (e.g., Bala et al,
 321 2008).

322 In the slow case, $L\Delta E$ (Eq. 19) and ΔH (Eq. 22) both consist of two terms: one stemming
 323 from a change in η , and the other from the slow change in $R_s - G$ (i.e., $\lambda \Delta T_a$). Fig. 3
 324 shows the contributions from these terms in the ensemble mean of the CMIP5 simulations.
 325 The $\Delta \eta$ terms (top row) are equal and opposite to each other, representing an increase in
 326 LE at the expense of H . In contrast, both $\lambda \Delta T_a$ terms (second row) are generally positive,
 327 with magnitudes that differ by a factor of β_0 , mirroring the response to $\Delta(R_s - G)_{\text{fast}}$ (Eq.
 328 23). When these contributions are combined (third row), the positive tendencies on $L\Delta E$
 329 reinforce each other, while the opposing tendencies on ΔH mostly cancel. This explains
 330 why ocean-mean ΔH is close to zero in the slow response to CO₂ forcing (bottom row;

⁵ These percentages are based on a comparison of the ocean-mean values that appear in the top left of each panel in Fig. 2.

Andrews et al, 2009), and why it is significantly smaller in magnitude than ocean-mean $L\Delta E$ in the total (fast + slow) response, where $\Delta(R_s - G)$ is also generally positive.

This result points to an important conceptual difference between the Penman-Monteith perspective and other interpretations of evaporation change based on diagnostic assessments of the surface or atmospheric energy budgets. In the surface energy budget, the relatively small decrease in H over the global oceans implies that the increase in LE is mostly offset by an increase in $R_s - G$. Similarly, in the global-mean atmospheric energy budget, the small magnitude of $\Delta\bar{H}$ means that $L\Delta\bar{P}$ is mostly offset by $\Delta\bar{R}_a$ (Eq. 1).

Within the Penman-Monteith framework, however, Fig. 1 shows that $\Delta(R_s - G)$ accounts for just 45% and 41% of the change in ocean-mean evaporation in the equilibrium and transient simulations, respectively—a very different diagnosis than suggested by the surface energy budget alone. This difference stems from the fact that, in addition to the surface energy budget, the Penman-Monteith equation also incorporates other physical constraints related to the turbulent transfer of latent and sensible heat from the surface (Eqs. 2-3). In particular, η represents a thermodynamic constraint on the partitioning between LE and H , as noted in Section 2. When this constraint is combined with energy conservation, ΔH and $L\Delta E$ are found to be closely related, each responding to ΔT_a and $\Delta(R_s - G)$ according to Eqs. 16 and 20. In GCM simulations, ΔT_a and $\Delta(R_s - G)$ conspire to make ΔH relatively small over most of the oceans, but this is by no means a general result. In a much cooler climate, for example, Eq. 20 indicates that ΔH would be significantly larger in response to a similar amount of warming.

4 Thermodynamic constraints on global precipitation over a wide range of climates.

In the previous Section, we showed that the change in ocean evaporation in simulations of CO₂-induced warming can be separated into a thermodynamic component that is a direct consequence of surface warming, and a diagnostic component that represents the effects of changes in net surface radiation, ocean heat uptake, and boundary-layer dynamics/relative humidity. Of these, the thermodynamic component was found to account for 2/3 of the total increase in ocean evaporation in equilibrium slab-ocean simulations, and for an even larger share in transient coupled simulations. The thermodynamic component is also strongly dependent on mean-state temperature, explaining why the largest fractional increase in evaporation tends to occur at high latitudes in Fig. 1. In light of these results, it is worth considering how thermodynamics might influence global precipitation over a wide range of climates much warmer and cooler than our own.

For this portion of our analysis, we revisit a series of idealized gray-radiation simulations run by O’Gorman and Schneider (2008), which were designed to exhibit a wide range of global-mean surface temperatures in response to imposed changes in atmospheric longwave optical depth, albeit with no representation of the radiative effects of clouds and atmospheric water vapor. Since the output from these simulations is not publicly available, we are unable to diagnose the factors contributing to their differences in global-mean precipitation, as we did for the CMIP ensembles. However, we can estimate the thermodynamic contribution to these differences by assuming that η varies with global-mean surface temperature (Eq. 11), and that all other variables are approximately constant ($R_s - G + \kappa = 197 \text{ Wm}^{-2}$; see Appendix). This gives the following approximation for global precipitation as a function of global-mean surface temperature alone (Eq. 10):

$$L\bar{P} \approx \eta(\bar{T}_a) \times 197 \text{ Wm}^{-2}. \quad (24)$$

375 Fig. 4 shows the estimated global precipitation based on Eq. 24 (gray line) compared
 376 with the actual results of O’Gorman and Schneider (2008) (black x’s). Despite some dis-
 377 crepancies at low temperatures, the agreement is good overall, suggesting that most of the
 378 change in global precipitation among their idealized simulations can be explained as a direct
 379 consequence of the change in global-mean surface temperature.

380 Of course, the real world is more complicated than the idealized, gray-radiation GCM
 381 used by O’Gorman and Schneider (2008). This might explain why changes in surface radi-
 382 ation appear to be smaller in these simulations than in the more realistic CMIP simulations
 383 discussed previously.

384 Yet even if surface radiation is not constant, it is still instructive to consider how global
 385 precipitation varies with surface temperature as a consequence of thermodynamics alone.
 386 Taking the derivative of η with respect to T_a , we find that the slope of the \bar{P} -vs.- \bar{T}_a curve in
 387 Fig. 4 is proportional to

$$\frac{d\bar{P}}{dT_a} \propto \frac{\beta_0}{(1+\beta_0)^2} \left(\alpha - \frac{2}{T_a} \right). \quad (25)$$

388 Importantly, this equation encapsulates what Pierrehumbert (2002) has identified as distinct
 389 constraints on global precipitation operating in different temperature regimes. At the cold
 390 extreme, $\beta_0 \gg 1$ as q^* approaches 0. While this results in a large *fractional* increase in
 391 global precipitation with warming (Eq. 15), the actual increase is small (Eq. 25), reflect-
 392 ing the atmosphere’s limited capacity at cool temperatures to maintain water vapor against
 393 condensation. Conversely, at very warm temperatures, $\beta_0 \ll 1$ due to high values of q^* . In
 394 this regime, atmospheric water vapor is plentiful, but the change in global precipitation with
 395 warming is limited by net radiation at the surface (Eq. 4). It is therefore between these lim-
 396 its, where β_0 is $O(1)$ (i.e., $T_a \approx 280$ K), that global precipitation is thermodynamically most
 397 sensitive to changes in global mean surface temperature. While changes in surface radiation
 398 will affect the upper bound on global precipitation in the warm limit, the broad shape of the
 399 \bar{P} -vs.- \bar{T}_a curve is guaranteed by thermodynamic constraints inherent in η .

400 Finally, it is interesting to consider how the primacy of η in GCM simulations of CO_2 -
 401 induced warming relates to two other ideas for how global precipitation changes with warm-
 402 ing. The first, put forth by Kleidon and Renner (2013a,b), is that the hydrologic-cycle heat
 403 engine operates near the thermodynamic limit of maximum power. Using an idealized en-
 404 ergy balance model, Kleidon and Renner (2013b) explore the implications of this assump-
 405 tion for how global evaporation/precipitation scales with warming, assuming no change in
 406 net surface radiation. Consistent with Eq. 24, they find that global evaporation is propor-
 407 tional to a term that can be shown to be *identical* to η , except that \bar{T}_a in their model rep-
 408 represents the average temperature at which water vapor condenses in the atmosphere. This
 409 suggests that Kleidon and Renner’s foundational assumption—that atmospheric convection
 410 approaches the thermodynamic limit of maximum power—may indeed have some relevance
 411 to GCM simulations, and perhaps also to the real atmosphere.

412 A second, similarly idealized conceptualization of the hydrologic cycle was proposed by
 413 Takahashi (2009), who argued that the change in precipitation with global warming is con-
 414 trolled by radiative cooling from the free troposphere rather than the full atmosphere. This
 415 idea—which represents a variant of the atmospheric energy budget perspective—is based
 416 on the principle that little of the surface sensible heat flux makes it out of the marine bound-
 417 ary layer. Using a 1-dimensional radiative-convective-equilibrium model constrained by this
 418 and a few other *a priori* assumptions, Takahashi found changes in precipitation with warm-
 419 ing that were similarly consistent with the idealized results of O’Gorman and Schneider in
 420 Fig. 4 (O’Gorman et al, 2012). While the idealized models of Takahashi (2009) and Kleidon

421 and Renner (2013a,b) do not appear to be incompatible, it remains unclear how they might
 422 relate to each other, or to the Penman-Monteith framework presented in this paper. We hope
 423 that future research will shed light on this important question.

424 5 Discussion

425 In this paper, we have shown that the Penman-Monteith equation, as expressed in Eq. 10,
 426 allows any change in ocean evaporation to be partitioned into distinct contributions from
 427 changes in surface temperature, net surface radiation, ocean heat uptake, and boundary layer
 428 dynamics/relative humidity. In GCM simulations of CO₂-induced warming, we find that the
 429 majority of the change in ocean evaporation is a direct consequence of warming, represented
 430 by $\Delta\eta$ in Eq. 14. This component of evaporation change derives from fundamental thermo-
 431 dynamics, and therefore does not depend on the specific nature of the radiative forcing or
 432 on the model physics. Physically, this term represents a change in the partitioning between
 433 latent and sensible heat fluxes due to an increase in the surface-air moisture gradient (re-
 434 quired by the Clausius-Clapeyron equation), and a corresponding decrease in the surface-air
 435 temperature gradient (required by energy conservation). In fractional terms, the change in
 436 evaporation due to this effect diminishes with warming, explaining why the largest frac-
 437 tional changes in ocean evaporation tend to occur at high latitudes in GCM simulations
 438 (Fig. 1). Compared with this thermodynamic effect, the contribution to evaporation change
 439 from changes in net surface radiation (ΔR_s) and ocean heat uptake (ΔG) were found to
 440 be secondary but still significant, while changes in boundary-layer dynamics and relative
 441 humidity ($\Delta\kappa$) were found to be less important, particularly at regional scales.

442 Because $\Delta\kappa$ is small, the Penman-Monteith framework allows the change in evaporation
 443 to be estimated from the spatial pattern of ΔT_a , ΔR_s , and ΔG alone (Eq. 16). For example,
 444 in the fast response to CO₂ forcing, $R_s - G$ decreases due to significant ocean heat uptake,
 445 causing a decrease in global evaporation (and thus precipitation). Because SSTs are fixed,
 446 this decrease in evaporation is well approximated as $\eta\Delta(R_s - G)$. On longer timescales, sur-
 447 face temperatures rise and evaporation increases, in part due to thermodynamics, and in part
 448 because ocean heat uptake declines as the climate system returns to radiative equilibrium.
 449 Combined with the surface energy budget, Eq. 16 also leads to an equally accurate approx-
 450 imation of the change in sensible heat flux, H (Eq. 20). Thus, from the Penman-Monteith
 451 perspective, LE and H represent two sides of the same coin, each responding to ΔT_a , ΔR_s ,
 452 and ΔG according to Eqs. 16 and 20.

453 This interpretation of global hydrologic change is somewhat different from those based
 454 on the atmospheric energy budget, in which the change in global precipitation ($L\Delta\bar{P}$) is
 455 offset by $\Delta\bar{H}$ and a change in net atmospheric radiative cooling ($\Delta\bar{R}_a$). The energy-budget
 456 perspective provides little insight into $\Delta\bar{H}$, but has much to say about the physics behind
 457 $\Delta\bar{R}_a$ (e.g., Lambert and Webb, 2008; Stephens and Ellis, 2008; Previdi, 2010; Pendergrass
 458 and Hartmann, 2014; DeAngelis et al, 2015; Fläschner et al, 2016). In contrast, the Penman-
 459 Monteith equation provides new insight into the partitioning between $L\Delta E$ and ΔH over
 460 the oceans (Eqs. 16 and 20), but only if ΔT_a and $\Delta(R_s - G)$ (which is equal to ΔR_a in the
 461 global mean) are already known. This shows that the energy-budget and Penman-Monteith
 462 perspectives are fully complementary, and together provide a more complete understanding
 463 of evaporation change than either can provide by itself.

464 The Penman-Monteith perspective may also shed light on the response of the global
 465 hydrologic cycle to a change in the solar constant (e.g., Wetherald et al, 1975; Andrews
 466 et al, 2009), to changes in radiation due to solar geoengineering (Bala et al, 2008), or to

467 non-greenhouse forcings like a volcanic eruption (Trenberth and Dai, 2007). Of course, the
 468 accuracy of Eq. 16 is contingent on the forcing having little impact on the dynamics or
 469 relative humidity of the atmospheric boundary layer. In certain scenarios—e.g., a change
 470 in the concentration of absorbing aerosols in the boundary layer (Ming et al, 2010; Samset
 471 et al, 2016)—this condition might not be met. Yet even in these cases, the Penman-Monteith
 472 framework could prove to be a powerful tool for diagnosing the various contributions to
 473 changes in ocean evaporation at both global and regional scales.

474 Finally, it is important to note that while changes in surface radiation are of secondary
 475 importance to the overall change in ocean evaporation, they account for most of the inter-
 476 model spread, as evidenced by the standard deviations in the bottom row of Tables 1 and
 477 2. Relative to the other terms, the standard deviation of the ΔR_s contribution is roughly 2-3
 478 times larger across both the equilibrium and transient ensembles. This is not surprising given
 479 that $\Delta \bar{R}_s$ is closely tied to $\Delta \bar{R}_a$ (Eqs. 1 and 4), which depends on several model variables
 480 that are not well constrained, including clouds, tropospheric humidity, and the radiative
 481 transfer parameterization for calculating shortwave absorption by water vapor (DeAngelis
 482 et al, 2015; Fläschner et al, 2016). In contrast, the $\Delta \eta$ contribution is more consistent due
 483 to broad model agreement in the spatial patterns of warming, mean-state temperature, and
 484 mean-state evaporation. Altogether, these results suggest that thermodynamics alone will
 485 contribute to an increase in global precipitation with surface warming at a rate of about
 486 1.5%/K; whether global precipitation increases at a rate closer to 1 or 3%/K will largely
 487 depend on radiative changes.

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490 A Calculating the contributions to evaporation change in Eq. 14.

491 The terms in Eq. 14 were calculated as follows: LE and R_s were taken directly from model output; G was
 492 determined from R_s , LE , and H based on the surface energy budget (Eq. 4); η was calculated from the two-
 493 meter air temperature using Eqs. 11 and 13. Finally, given LE , η , R_s , and G , we then solved for κ in Eq.
 494 10. The contributions were calculated from ensemble-mean output over the last five years of the simulation
 495 period. In the equilibrium warming simulations, this was typically 21-25 years after CO_2 doubling. In the
 496 transient warming simulations, we used years 96-100 after CO_2 quadrupling. The contributions were first
 497 calculated for each month, and then the monthly contributions were averaged to arrive at an annual-mean
 498 value. However, the results were essentially unchanged when the contributions were calculated from annual-
 499 mean output.

500 To understand the *global* impact of the fractional contributions in Fig. 1, we must account for spatial
 501 variability in the magnitude of the mean-state evaporation and surface-air warming. To do so, we multiply
 502 each term in Eq. 14 by the following (dimensionless) weighting function,

$$w = \frac{E \Delta \bar{T}_a}{\bar{E} \Delta \bar{T}_a}, \quad (26)$$

503 where the overbars in the denominator indicate the ocean-mean values of each variable. These results are then
 504 averaged in space, yielding the ocean-mean contributions given in the top left of each panel in Fig. 1.

505 B Estimating $R_s - G + \kappa$ in the idealized simulations of O’Gorman and Schneider 506 (2008).

507 To estimate the value of $R_s - G + \kappa$ in O’Gorman and Schneider’s (2008) simulations, we use the fact that their
 508 control climate exhibits a global-mean surface-air temperature of $\bar{T}_a = 288$ K, and a global-mean precipitation

509 of 4.3 mm/day, which equates to $L\bar{E} = 124 \text{ Wm}^{-2}$. Given $\eta \approx 0.63$ at $T = 288 \text{ K}$, this implies a combined
510 value of $R_s - G + \kappa = 197 \text{ Wm}^{-2}$. If we assume that this sum is constant, global precipitation is directly
511 proportional to η , resulting in the gray curve in Fig. 4.

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Table 1 First column: The names of the CMIP3 (slab-ocean) models included in our analysis. Second column: The rate of increase in global-mean evaporation in response to a doubling of atmospheric CO₂, after reaching radiative equilibrium. Third column: The rate of increase in ocean-mean evaporation. Columns 4-7: The individual contributions to changes in ocean-mean evaporation from $\Delta\eta$, ΔR_s , ΔG , and $\Delta\kappa$, according to Eq. 14. The second row from the bottom gives the ensemble-mean rates (Fig. 1). These were calculated from ensemble-mean variables, and therefore differ slightly from the average rates of the individual models. The bottom row gives the standard deviation across models. In the second (third) column, global-mean (ocean-mean) rates were calculated using the global-mean (ocean-mean) values of ΔE , E , and ΔT . Global-mean and ocean-mean rates are highly correlated at $r = 0.93$, indicating the dominant influence of the ocean on global evaporation.

| CMIP3 model | $\frac{\Delta E}{EAT}$ (globe) | $\frac{\Delta E}{EAT}$ (oceans) | $\Delta\eta$ | ΔR_s | ΔG | $\Delta\kappa$ |
|---------------------------|--------------------------------|---------------------------------|--------------|--------------|------------|----------------|
| Can-CGCM3.1 (T47) | 2.02 | 2.14 | 1.37 | 0.96 | -0.01 | -0.22 |
| Can-CGCM3.1 (T63) | 2.18 | 2.28 | 1.37 | 0.95 | 0.07 | -0.16 |
| CSIRO-Mk3.0 | 2.26 | 2.67 | 1.56 | 1.20 | -0.01 | -0.13 |
| GFDL-CM2 | 1.35 | 1.76 | 1.53 | 0.59 | 0.06 | -0.43 |
| HadGEM1 | 1.81 | 2.18 | 1.48 | 0.89 | 0.09 | -0.34 |
| INM-CM3 | 1.58 | 1.85 | 1.56 | 0.71 | -0.06 | -0.37 |
| MIROC3.2 (hires) | 1.92 | 2.13 | 1.47 | 1.05 | 0.02 | -0.46 |
| MIROC3.2 (medres) | 2.18 | 2.37 | 1.51 | 1.23 | -0.01 | -0.39 |
| MPI-OM | 1.95 | 2.23 | 1.46 | 1.07 | -0.10 | -0.23 |
| MRI-CGCM2 | 2.28 | 2.48 | 1.52 | 1.16 | 0.01 | -0.24 |
| Ensemble mean | 1.97 | 2.22 | 1.49 | 1.00 | 0.01 | -0.30 |
| Standard deviation | 0.30 | 0.27 | 0.07 | 0.21 | 0.06 | 0.12 |

Table 2 As in Table 1, but for the CMIP5 (coupled) simulations. The rates of change in evaporation are based on years 96-100 after CO₂ quadrupling. The global-mean and ocean-mean rates are correlated at $r = 0.81$.

| CMIP5 model | $\frac{\Delta E}{EAT}$ (globe) | $\frac{\Delta E}{EAT}$ (oceans) | $\Delta\eta$ | ΔR_s | ΔG | $\Delta\kappa$ |
|---------------------------|--------------------------------|---------------------------------|--------------|--------------|------------|----------------|
| BCC-CSM1.1 | 1.65 | 1.98 | 1.41 | 1.00 | -0.21 | -0.26 |
| CanESM2 | 1.41 | 1.85 | 1.44 | 0.74 | -0.19 | -0.17 |
| CCSM4 | 1.46 | 1.99 | 1.42 | 1.07 | -0.22 | -0.34 |
| CSIRO-Mk3.6.0 | 1.85 | 2.29 | 1.53 | 1.29 | -0.27 | -0.35 |
| HadGEM2-ES | 1.23 | 1.78 | 1.43 | 0.92 | -0.22 | -0.42 |
| INM-CM4 | 1.35 | 1.58 | 1.54 | 0.79 | -0.46 | -0.31 |
| IPSL-CM5A-LR | 2.02 | 2.64 | 1.71 | 1.53 | -0.27 | -0.38 |
| MIROC5 | 1.46 | 1.64 | 1.46 | 1.02 | -0.34 | -0.59 |
| MPI-ESM-LR | 1.63 | 2.19 | 1.52 | 1.20 | -0.34 | -0.24 |
| MPI-ESM-MR | 1.77 | 2.26 | 1.49 | 1.21 | -0.27 | -0.23 |
| MRI-CGCM3 | 2.27 | 2.27 | 1.43 | 1.28 | -0.29 | -0.23 |
| NorESM1-M | 1.43 | 1.96 | 1.52 | 1.12 | -0.32 | -0.40 |
| Ensemble mean | 1.62 | 2.04 | 1.51 | 1.10 | -0.28 | -0.33 |
| Standard deviation | 0.30 | 0.31 | 0.08 | 0.22 | 0.08 | 0.11 |

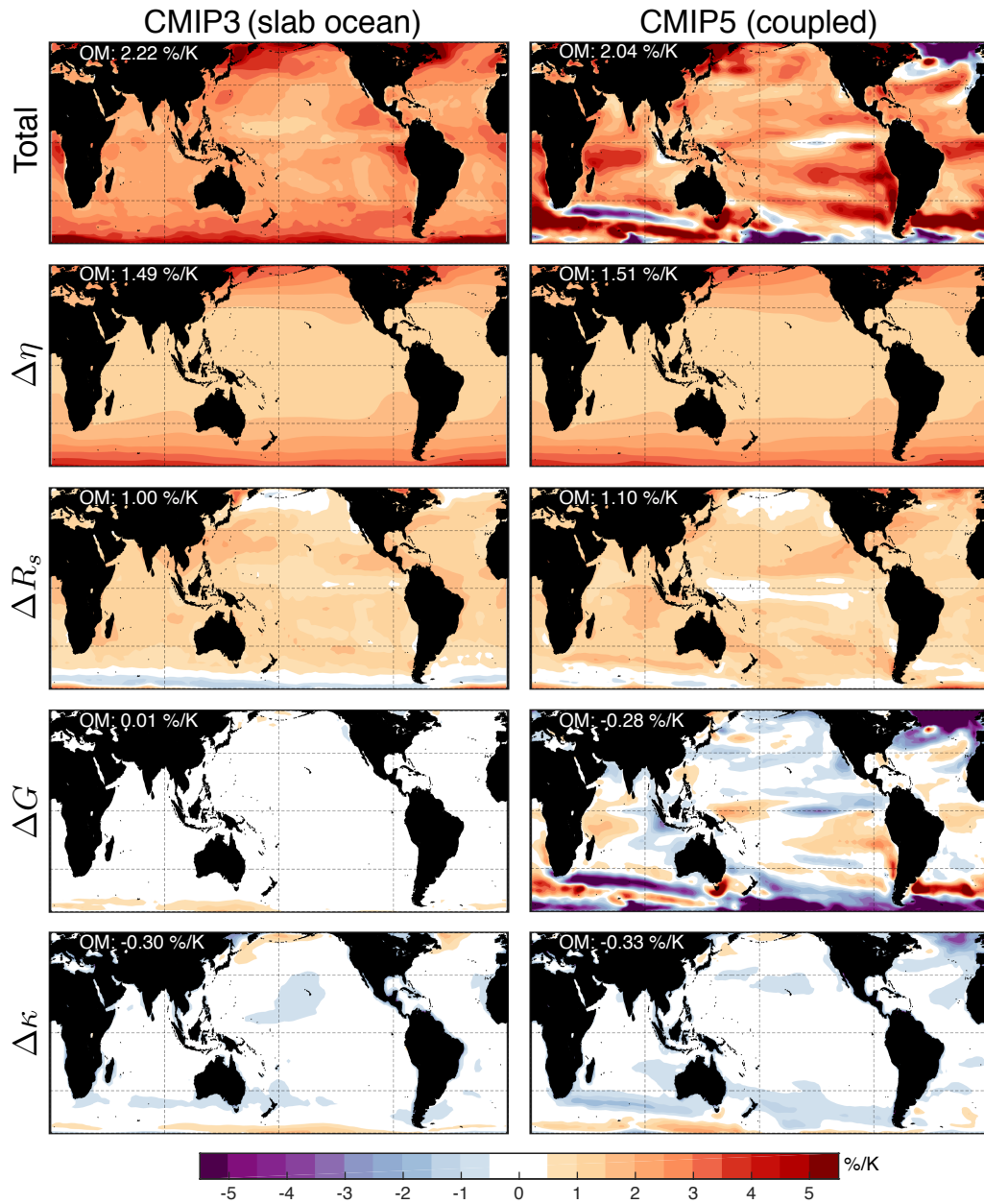


Fig. 1 The percent change in evaporation over the oceans in response to CO_2 -induced warming (top row), and the individual contributions from changes in η (second row), R_s (third row), G (fourth row), and κ (fifth row), in the equilibrium (left column) and transient (right column) simulations. Each contribution was calculated from ensemble-mean output according to Eq. 14 (see Appendix), and represents the change per Kelvin of global warming over the oceans, which is equal to 3.15 K in the $2\times\text{CO}_2$ equilibrium simulations, and 4.13 K in the $4\times\text{CO}_2$ transient simulations. The top left corner of each panel gives the ocean-mean (OM) values of each contribution. The results are broadly similar between the equilibrium and transient ensembles, with the exception of the contribution from ocean-heat uptake (ΔG), which is negligible in the equilibrium simulations due to the absence of a dynamical ocean.

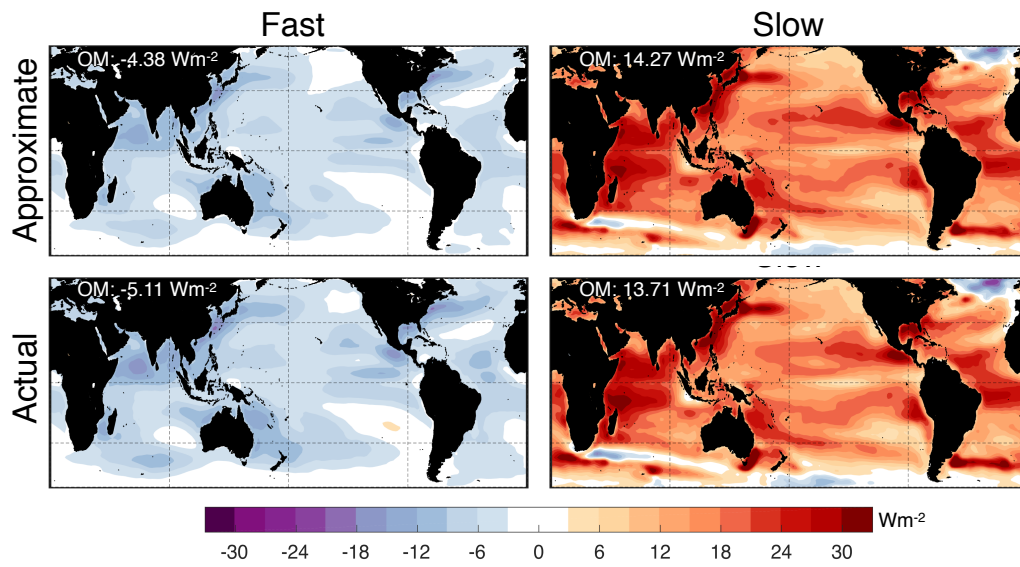


Fig. 2 The “fast” (left column) and “slow” (right column) changes in ocean evaporation in GCM simulations of an abrupt quadrupling of atmospheric CO_2 (in Wm^{-2}). The fast component represents the direct response of evaporation to CO_2 quadrupling with fixed SSTs, while the slow component represents the gradual changes that occur as the climate warms. Top row: the approximate changes calculated from Eqs. 17 and 19, using ensemble-mean values of F , S , E , T_a , and ΔT_a . Bottom row: the actual changes in the ensemble mean of CMIP5 simulations. Ocean-mean (OM) values are given in the top left corner of each panel.

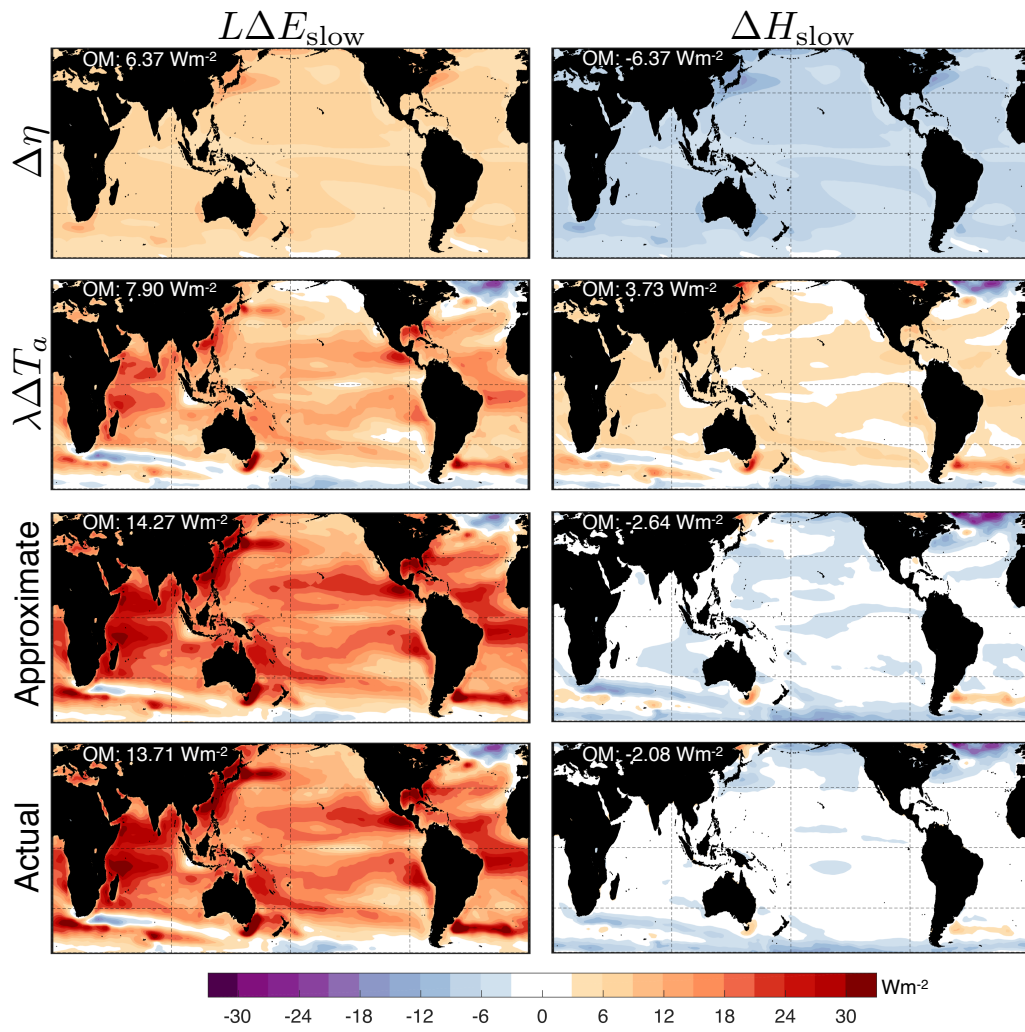


Fig. 3 The slow changes in LE (left column) and H (right column) in GCM simulations of an abrupt quadrupling of atmospheric CO_2 (in Wm^{-2}). Top row: the contribution from the $\Delta\eta$ term in Eqs. 16 and 20. Second row: the contribution from the $\lambda\Delta T_a$ term in Eqs. 16 and 20. Third row: the full approximation given by Eqs. 16 and 20. Fourth row: the actual change in the ensemble mean of CMIP5 simulations. Note that the two panels on the bottom left match those in the right column of Fig. 2.

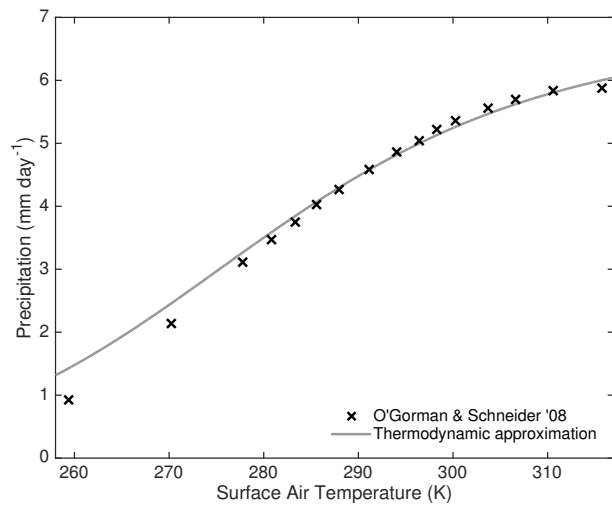


Fig. 4 Global mean precipitation (P) as a function of global-mean surface-air temperature ($\Delta\bar{T}_a$), according to the idealized simulations of O'Gorman and Schneider (2008) (black x's), and the thermodynamic approximation assuming that η varies with global-mean surface temperature and $R_s - G + \kappa$ is held fixed at $197 \text{ Wm}^{-2}\text{K}^{-1}$ (gray line).