

1 **Unsupervised clustering of Southern Ocean Argo float**
2 **temperature profiles**

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6 **Key Points:**

- 7 • We apply Gaussian mixture modeling (GMM) to Southern Ocean temperature data
- 8 • GMM identifies spatially coherent profile types without using latitude or longitude
- 9 information
- 10 • GMM offers a complementary approach for objectively classifying temperature pro-
- 11 files

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Abstract

The Southern Ocean has a complex density structure characterized by sharp fronts, steeply tilted isopycnals, and deep seasonal mixed layers. Methods of defining Southern Ocean density structures traditionally rely on somewhat ad-hoc combinations of physical, chemical, and dynamic properties. As a step towards an alternative approach for defining water masses, here we apply an unsupervised classification technique (that is, Gaussian mixture modelling or GMM) to Southern Ocean Argo float temperature profiles. GMM, without using any latitude or longitude information, automatically identifies several spatially coherent circumpolar classes influenced by the Antarctic Circumpolar Current. In addition, GMM identifies classes that bear the imprint of mode/intermediate water formation and export, large-scale gyre circulation, and the Agulhas Current, among others. Because GMM is robust, standardized, and automated, it can potentially be used to identify structures (such as fronts) in both observational and model datasets, possibly making it a useful complement to existing classification techniques.

Plain Language Summary

The Southern Ocean is an important part of the climate system, in part because it absorbs a large fraction of the heat and carbon that is added to the atmosphere/ocean system by human-driven fossil fuel burning. In this work, we use a machine learning technique to automatically sort Southern Ocean temperature measurements into groups based on how those temperature measurements change with depth. Different groups have the fingerprints of different large-scale circulation patterns, such as the powerful Antarctic Circumpolar Current that flows around Antarctica. The groups that we identify are consistent with our understanding of the Southern Ocean, which gives us confidence that our machine learning technique may be useful for automatically grouping measurements and computer model data in the future. This matters because the climate science community needs a new set of tools, possibly including the machine learning technique that we use in this paper, to deal with a very large, ever-increasing volume of observational and computer model data.

1 Introduction

The Southern Ocean (SO) is a critical component of Earth's climate system, having thus far absorbed greater than 75% of the energy added via anthropogenic emissions and 50% of the excess carbon [Fletcher *et al.*, 2006; Frölicher *et al.*, 2015]. Its ability to absorb heat and carbon comes in part from its unique density structure and circulation, which features upwelling of cold, nutrient rich waters and regions of dense water formation [Lumpkin and Speer, 2007]. Characterizing and understanding the mean state and variability of SO density structure remains an important and climatically-relevant goal of modern oceanography.

Through decades of effort, the oceanographic community has converged on a description of ocean structure that uses temperature, salinity, dynamical, and biogeochemical patterns to define different water masses [Emery, 2003; Talley, 2013, and references therein]. For example, Herraiz-Borreguero and Rintoul [2011] use potential vorticity minima and particular neutral density surfaces to locate SO mode water pools. Such systematic approaches employ the understanding that water mass properties are "set" in their formation regions and modified by advection, mixing, and biogeochemical processes. Classification in latitude-longitude has traditionally been centered around several fronts of the Antarctic Circumpolar Current (ACC), defined by sharp transitions in sea surface height or neutral density [Kim and Orsi, 2014]. The classical southern boundary of the ACC (SBDY) marks the transition between subpolar, gyre-dominated circulations and lower-latitude, more circumpolar flow. The ACC itself features three circumpolar fronts, namely the southern ACC front (SACCF), the Polar Front (PF), and the Subantarctic Front (SAF) [Orsi *et al.*, 1995]. These three fronts separate the subpolar SO from the subtropical domain [Garabato *et al.*, 2011].

The modern, property-driven classification scheme is extremely useful and will continue to be useful well into the future, but it is not necessarily ideal for every application. Many of the temperature, salinity, and density values used to delimit one water mass from another are somewhat ad-hoc and very specific (e.g. boundaries between different types of mode water). These schemes are useful for observational data analysis but difficult to apply to numerical models of the ocean, which do not necessarily feature exactly the same structure as the observed ocean [Sallée *et al.*, 2013]. In addition, traditional classification approaches that define water masses by specific property ranges are limited by the fact that these properties may change over time (for example, the warming of Antarctic Bottom

71 Water observed by *Purkey and Johnson* [2010]). We suggest that it is prudent to develop and
72 test alternative methods for the classification of oceanic temperature, salinity, and density
73 structures, as a complement to existing expertise-driven methods.

74 *Maze et al.* [2017] have shown that Argo temperature profile data from the North At-
75 lantic Ocean can be usefully grouped into classes using Gaussian mixture modelling (GMM),
76 an unsupervised classification technique. GMM describes the spatial structure of Argo pro-
77 files as a collection of Gaussian modes whose means and standard deviations generally vary
78 with pressure. In this work, we apply GMM to Southern Ocean Argo temperature profiles
79 in the upper 1000 m of the water column. We find that GMM identifies several circumpolar
80 classes, gyres, the Agulhas current, and pathways broadly associated with the formation and
81 export of mode and intermediate waters. In section 2, we describe the Argo dataset and the
82 basics of GMM. In section 3, we present the results of applying GMM to Southern Ocean
83 Argo data, and in sections 4 and 5 we offer a brief discussion and summarize our conclu-
84 sions.

85 **2 Methods**

86 We applied an unsupervised classification method (i.e. Gaussian mixture modelling,
87 hereafter GMM) to Southern Ocean Argo float data. In this section, we briefly describe the
88 Argo dataset and the basics of GMM. We use the *scikit-learn* machine learning library for
89 Python (<http://scikit-learn.org/>), and the source code used for much of the analysis
90 in this paper is available via Github (<https://github.com/DanJonesOcean/OceanClustering>).
91 We refer the reader to *Maze et al.* [2017] for further detail on applying GMM to Argo float
92 data.

93 **2.1 Argo float dataset**

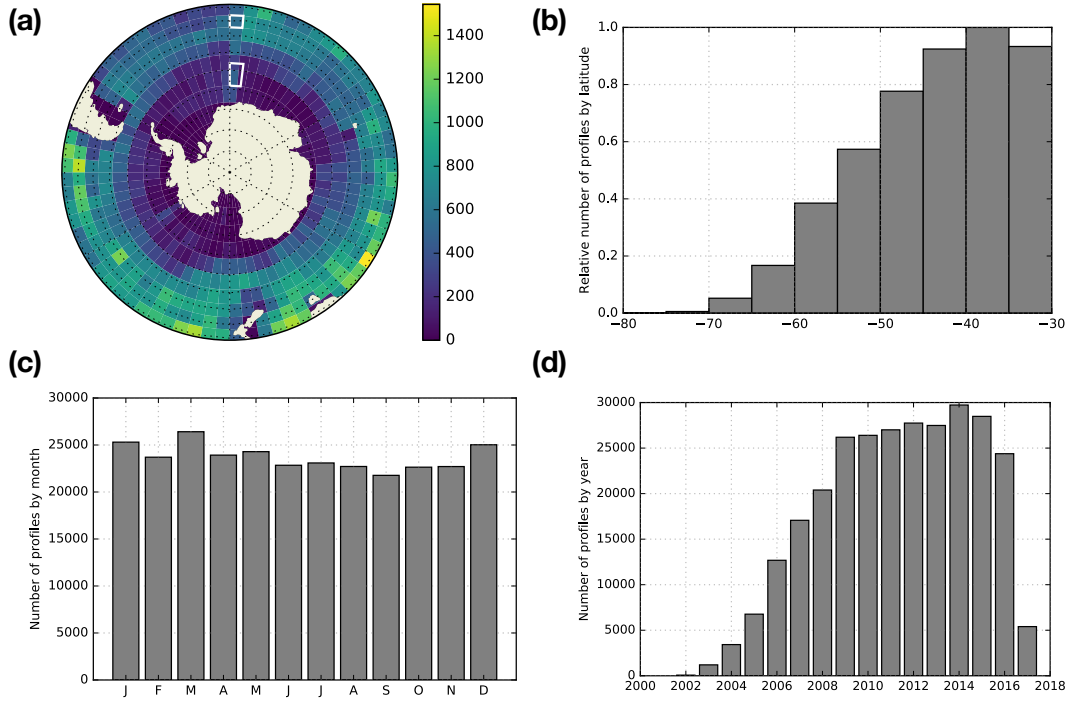
94 Argo floats are autonomous ocean instruments that measure, at minimum, the tempera-
95 ture and salinity of the ocean by periodically taking vertical profiles. Every 10 days, starting
96 at a “neutral” position of 1000 m, an Argo float dives down to 2000 m before rising to the
97 surface, taking a vertical profile of the water column along the way. The measurements are
98 transmitted via satellite and are ultimately made freely available via the Argo Global Data
99 Assembly Centers (GDACs) after some quality control checks. At the time of this writing,

100 over 3800 Argo floats are active in the global ocean, producing over 100,000 temperature and
101 salinity profiles per year with an average spacing of 3° (<http://www.argo.ucsd.edu/>).

102 For this study, we selected all available Argo profiles south of 30°S that have been
103 flagged by the GDACs as “observation good” (i.e. quality control flag = 1) covering the time
104 period from 2001 to early 2017. More specifically, we used a vertically interpolated product
105 with 400 equally spaced pressure levels ranging from 5 to 2000 dbar in 5 dbar increments.
106 After discarding profiles with greater than or equal to 6% NaN values (2% of the initial num-
107 ber of profiles) and discarding pressure levels with greater than or equal to 3% NaN values,
108 we were left with 284,427 profiles, each with 192 pressure levels between 15 dbar and 980
109 dbar. Most of these initially removed NaN values came from interpolation below roughly
110 1000 dbar, as opposed to gaps in the original dataset. We selected our NaN cut-off values
111 based on the relatively large increase in the number of NaN values below 1000 dbar. We re-
112 placed all remaining NaN values ($\ll 1\%$ of the total temperature measurements) with lin-
113 earlyly interpolated estimates using nearest neighbor values with respect to pressure. We refer
114 to the resulting dataset as the cleaned dataset.

115 Because of the autonomous and free-drifting nature of the floats, the profiles are not
116 distributed evenly in latitude/longitude (Figure 1). The profiles are more heavily concen-
117 trated in the Pacific sector (roughly 890 profiles per degree longitude, totalling 47% of pro-
118 files) and Indian sector (800 profiles per degree longitude, totalling 34% of profiles), with
119 relatively fewer profiles in the Atlantic sector (610 profiles per degree longitude, 19% of
120 total). When counted in equal-area bins and plotted by latitude, we see that the number of
121 profiles decreases towards Antarctica (Figure 1(b)), which is partly due to challenging oper-
122 ational conditions associated with seasonal sea ice, which can extend to just north of 60°S
123 at maximum areal extent. The profiles are slightly over-represented in the Austral summer
124 and autumn (DJF-MAM, 52% of profiles) and under-represented in the Austral winter and
125 spring (JJA-SON, 48% of profiles), and the number of profiles increases until 2013 (Figure
126 1(c,d)). Since we selected an Argo dataset that was created in early 2017, there are relatively
127 few profiles from that year.

132 The profiles selected for this study display a large variety of vertical temperature struc-
133 tures (Figure 2). The range of temperatures is wider in the surface and considerably narrower
134 with pressure, in part reflecting the seasonal cycle in upper ocean temperatures. A large
135 number of profiles feature colder temperatures near the surface and warmer temperatures

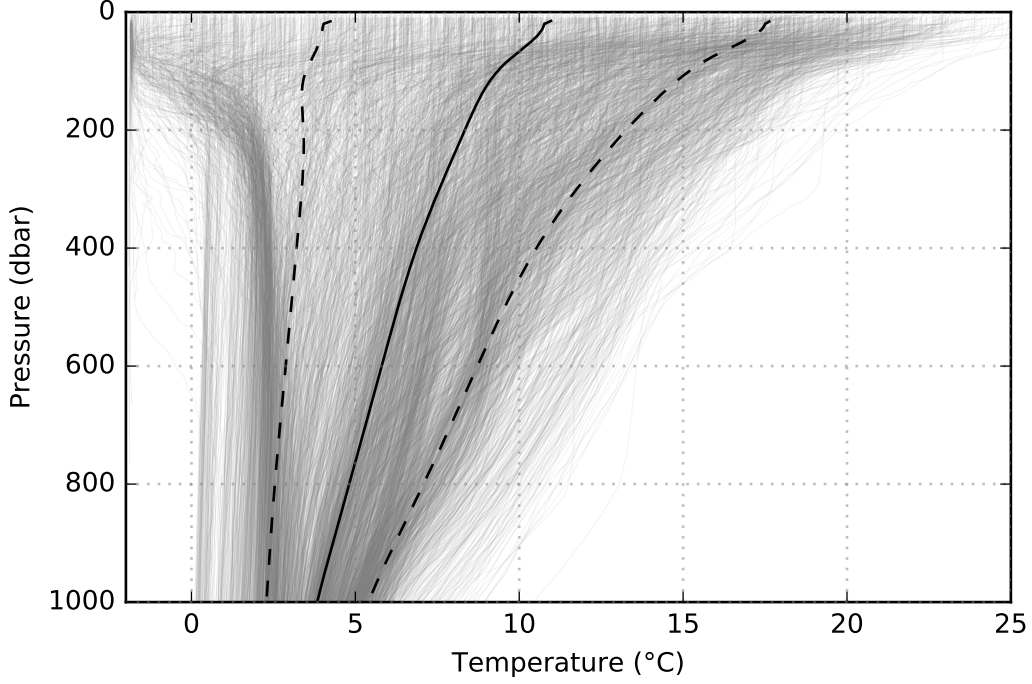


128 **Figure 1.** Distribution of Argo temperature profiles from the cleaned dataset. (a) Number of profiles in
 129 $5^\circ \times 5^\circ$ bins. Two equal-area boxes are shown for reference (solid white lines). (b) Relative number of pro-
 130 files by latitude, scaled by an area-weighting factor $\cos(\phi)$, where ϕ is latitude. The temporal distribution of
 131 profiles shown by (c) month and (d) year.

136 in the interior, a physical arrangement that would be unstable to convection without the com-
 137 pensating effect of salinity. Water masses around Antarctica tend to be fresher at the surface
 138 and saltier in the interior due to glacial melt, freshwater flux, and the balance of evapora-
 139 tion/precipitation. This arrangement of temperature and salinity can be stable to vertical mix-
 140 ing (called “salt stratification”). In addition, the thermocline, i.e. the region of the ocean that
 141 features a rapid change in temperature with pressure, is visible in some temperature profiles.

145 **2.2 Gaussian mixture modeling**

146 Gaussian mixture modeling (GMM) is a probabilistic approach for describing and clas-
 147 sifying data. It attempts to fit (or “model” in the statistical sense) the data as a linear com-
 148 bination of multi-dimensional Gaussian distributions with unknown means and unknown
 149 standard deviations. Let \mathbf{X} be the array of N vertical profiles, each with D pressure levels,
 150 and let $p(\mathbf{X})$ be the probability distribution function (PDF) representing the entire dataset.



142 **Figure 2.** Plot of 10% of the Argo temperature profiles, chosen at random, in the upper 1000 dbar of the
 143 cleaned dataset, along with the mean (solid line) and the mean plus or minus one standard deviation (dashed
 144 lines) across the entire dataset.

151 GMM represents the PDF as a weighted sum of K Gaussian classes, indexed by k , i.e.:

$$p(\mathbf{X}) = \sum_{k=1}^K \lambda_k \mathcal{N}(\mathbf{X}; \mu_k, \Sigma_k). \quad (1)$$

152 Here, $\mathcal{N}(\mathbf{x}; \mu_k, \Sigma_k)$ is the multi-dimensional Gaussian PDF with a vector of means μ_k and
 153 covariance matrix Σ_k , i.e.:

$$\mathcal{N}(\mathbf{x}; \mu_k, \Sigma_k) = \frac{\exp\left[-\frac{1}{2}(\mathbf{x} - \mu_k)^T \Sigma_k^{-1} (\mathbf{x} - \mu_k)\right]}{\sqrt{(2\pi)^D |\Sigma_k|}}. \quad (2)$$

154 The probability associated with class k is $p(k) = \lambda_k$. The probability of profile \mathbf{x} being in
 155 class k is $p(k|\mathbf{x}) = \lambda_k \mathcal{N}(\mathbf{x}; \mu_k, \Sigma_k) / p(\mathbf{x})$, where the vector \mathbf{x} is a single profile taken from
 156 the complete array \mathbf{X} and $p(\mathbf{x})$ is equation (1) with a single profile \mathbf{x} as the argument, i.e. a
 157 normalizing factor. Both \mathbf{x} and μ_k are vectors of length D , and Σ_k is a matrix of size $D \times D$.

158 Starting with random initial guesses for the classes, GMM proceeds by iteratively ad-
 159 justing the means μ_k and standard deviations Σ_k (i.e. the “parameters”) of the classes in or-

160 der to maximize a logarithmic measure of likelihood, i.e.:

$$\log[p(\mathbf{X})] = \sum_{i=1}^N \log \left[\sum_{k=1}^K \lambda_k \mathcal{N}(\mathbf{X}; \mu_k, \Sigma_k) \right]. \quad (3)$$

161 GMM uses an expectation-maximization approach, described in *Maze et al.* [2017]. This
 162 algorithm monotonically converges on a local maximum. GMM is a generalization of k -
 163 means clustering, which only attempts to minimize in-group variance by shifting the means.
 164 By contrast, GMM attempts to identify means and standard deviations, allowing for some
 165 variation about the centres of the Gaussian distributions.

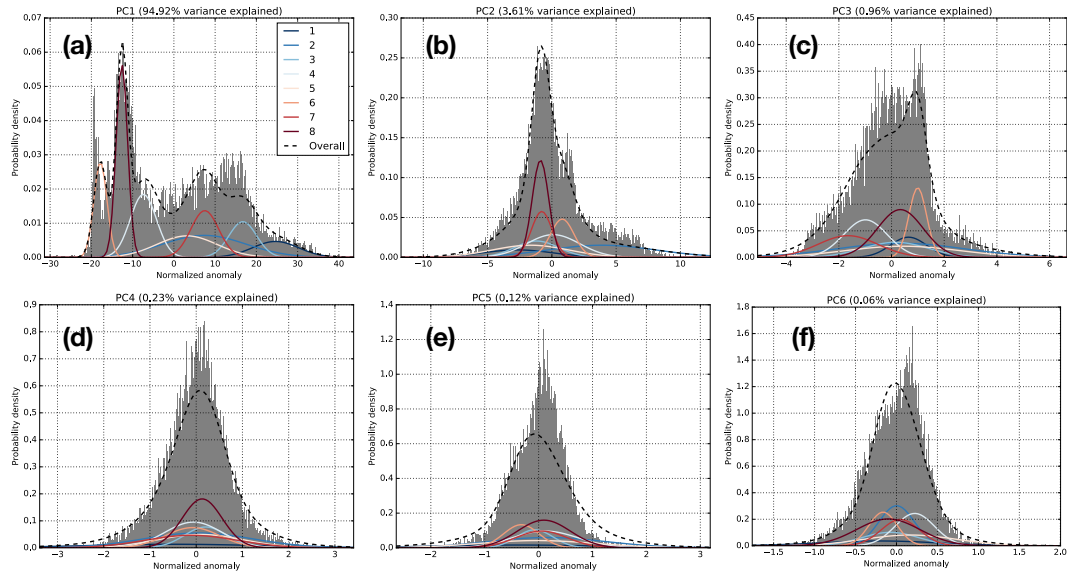
166 In our instance of GMM, each pressure level is treated as a “dimension”, and the Gaus-
 167 sian parameters are associated with each pressure level. However, we may not need all of
 168 these pressure levels to accurately describe the dataset, as ocean temperature changes much
 169 more rapidly in the mixed layer and thermocline than in the interior. In order to reduce the
 170 computational complexity of the problem, we transform the profile data from pressure space
 171 to an alternative space using principal component analysis (PCA). Specifically, we calculate
 172 principal components that capture a desired fraction of the vertical variability of the dataset.
 173 Each eigenvector may be thought of as a “profile type” that describes a certain amount of
 174 variance in the data with pressure (note that this is not necessarily the same thing as a “typi-
 175 cal profile”). We calculate J principal components via the transformation:

$$\mathbf{X}(z) = \sum_{j=1}^J \mathbf{P}(z, j) \mathbf{Y}(j), \quad (4)$$

176 where z is the pressure level, J is the total number of principal components (index j), and
 177 $\mathbf{P}(z, j)$ is the transformation matrix between pressure space and principal component space.
 178 This strategy is an example of “dimensionality reduction”, which is common in machine
 179 learning approaches.

180 We find that $J = 6$ captures 99.9% of the variance in the vertical structure, which
 181 greatly reduces the number of dimensions needed to describe the Argo profile data used here,
 182 i.e. from 194 pressure levels to 6 principal components (PCs). We refer to this dataset as
 183 the “cleaned, compressed” dataset. Nearly 95% of the variance is explained by the first PC
 184 (i.e. PC1), and the Gaussian functions associated with PC1 are relatively distinct, captur-
 185 ing the broad shape of the temperature distribution (Figure 3). For higher indexed PCs, the
 186 Gaussians overlap more, but their sum still makes up a representation of the temperature dis-
 187 tribution that is sufficiently accurate for our purposes. The fact that we only need six PCs
 188 to capture 99.9% of the variance is consistent with the strong vertical coherence found in

189 the Southern Ocean, which is well-described by an equivalent barotropic model [Karsten
 190 and Marshall, 2002]. For more information on the principal components that we used in this
 191 work, see the supporting information (Figure S1 and S2).



192 **Figure 3.** Probability density functions for the (dimensionless) principal component amplitude coefficients
 193 associated with each profile, along with the Gaussian functions generated by GMM with $K = 8$ classes.

194 We used a “training” dataset, a subset of the cleaned, compressed dataset, to estimate
 195 the parameters (i.e the means and standard deviations) of the GMM classes. To generate the
 196 GMM training set, we randomly selected a single profile from each $1^\circ \times 1^\circ$ bin. Each training
 197 dataset contains 12,286 profiles (roughly 4% of the cleaned, compressed dataset), distributed
 198 evenly in latitude/longitude space. Note that this sub-selection is not related cross-validation
 199 analysis, in which there are “training” and “test” datasets [Maze *et al.*, 2017]. Instead, we use
 200 a random sub-selection that is roughly uniform in latitude-longitude as our test dataset, and
 201 then we apply the GMM model to the entire cleaned, compressed dataset. As discussed in
 202 the supporting information, our results are not sensitive to our choice of test dataset.

203 Once we have our test dataset and calculate the optimized parameters (that is, the means
 204 and standard deviations of the Gaussians), we then statistically represent (i.e. ‘model’) the
 205 entire cleaned, compressed dataset with the fitted Gaussian model using optimized parame-
 206 ters. The end result is a probabilistic description of the cleaned, compressed Argo tempera-
 207 ture profile dataset in terms of a linear combination of Gaussian distributions that vary with
 208 pressure. Each profile then has a probability distribution across the classes, and the profile is

209 assigned to the class with the highest probability. Our results are not sensitive to our choice
 210 of training dataset (see supporting information, Table S1).

211 **2.2.1 Selecting the number of classes**

212 GMM does have one free parameter, i.e. the maximum number of classes K . In order
 213 to determine the most appropriate value for K , we applied a statistical test, namely a
 214 Bayesian Information Criterion (BIC). BIC uses an empirically formulated cost function that
 215 rewards likelihood and penalizes the number of classes K :

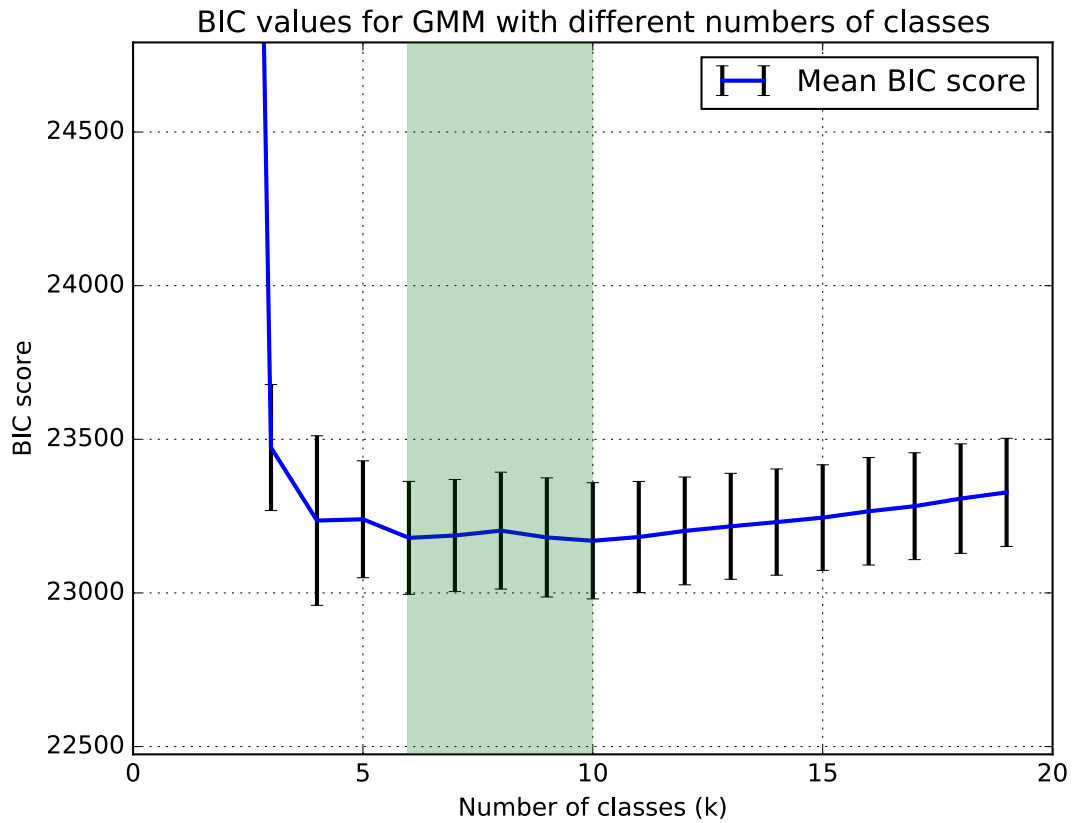
$$BIC(K) = -2\mathcal{L}(K) + N_f(K) \log(n), \quad (5)$$

216 where \mathcal{L} is a measure of likelihood, n is the number of profiles used in the BIC test, and N_f
 217 is the number of independent parameters to be estimated:

$$N_f(K) = K - 1 + KD + \frac{KD(D-1)}{2}. \quad (6)$$

218 In this framework, the optimum value of K minimizes the BIC score. We perform a number
 219 of BIC tests, using different subsets of the data and different values of K , to estimate the dis-
 220 tribution and variability of BIC. Using the roughly 300 km decorrelation scale of the South-
 221 ern Ocean as guidance [Ninove *et al.*, 2016], we randomly select a profile from each $4^\circ \times 4^\circ$
 222 grid cell, returning 884 random profiles for each BIC test. We calculate BIC scores for each
 223 set of 884 random profiles (in principal component space) using a range of classes K from
 224 1 to 19 (Figure 4). For each value of K , we repeat the random selection and BIC process 50
 225 times. BIC analysis does not feature a clear minimum, but instead it suggests that the opti-
 226 mum value of K lies between 6 and 10.

231 It may seem counterintuitive that BIC does not return a single optimum value for K ,
 232 but this is consistent with the nature of K as a weakly constrained free parameter that con-
 233 trols the level of complexity of the statistical description of the dataset. Oceanography has
 234 a rich history of expertise-driven clustering using physical and biogeochemical criteria (e.g.
 235 PV minima, oxygen minima) and the fingerprints of various processes (e.g. gyre circula-
 236 tion). These descriptions can be arranged into hierarchies, from coarse/simple (e.g. two-layer
 237 quasi-geostrophic models) to rich and complex (e.g. the descriptions found in Talley [2013]).
 238 The level of detail required in the description depends on the application at hand. For exam-
 239 ple, a simple β -plane model is sufficient to explain the existence of gyres and western bound-
 240 ary currents; it constitutes a first-order description of gyres. Algorithmic clustering offers a

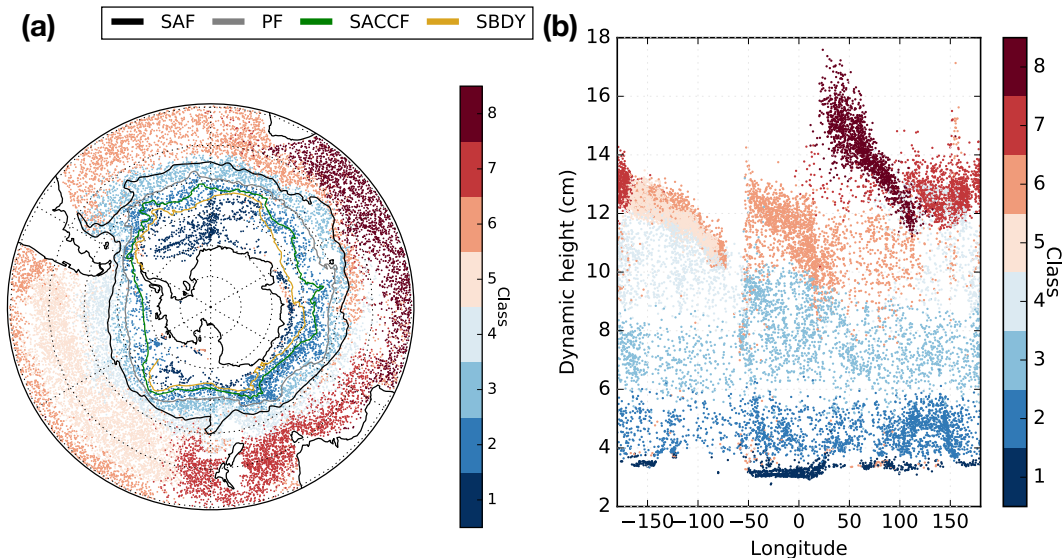


227 **Figure 4.** Bayesian Information Criteria (BIC) scores versus the specified number of classes K . For each
 228 K , we calculate the BIC score 50 times using randomly selected profiles as discussed in the text. The means
 229 (solid blue line) and standard deviations (error bars) are shown for each K . The range of the smallest mean K
 230 values is indicated by green shading.

241 robust way to traverse this hierarchy using a range of K values. Although statistical tests can
 242 be used as rough guides for choosing the number of classes, there is not necessarily a single
 243 ideal value for K . We explore the impact of K on our results in section 4.

244 3 Results

245 In order to identify patterns in the temperature structure of the Southern Ocean, we de-
 246 scribe the cleaned, compressed Argo temperature profile dataset as a linear combination of
 247 multi-dimensional Gaussian functions that vary with pressure, using $K = 8$ different classes.
 248 Despite the fact that GMM does not have access to the longitudes and latitudes of the pro-
 249 files, it identifies spatially coherent structures, some of which are roughly demarcated by the
 250 fronts of the ACC as defined by *Kim and Orsi [2014]* (Figure 5). For ease of interpretation,
 251 we sorted the classes by mean temperature (Table 1).



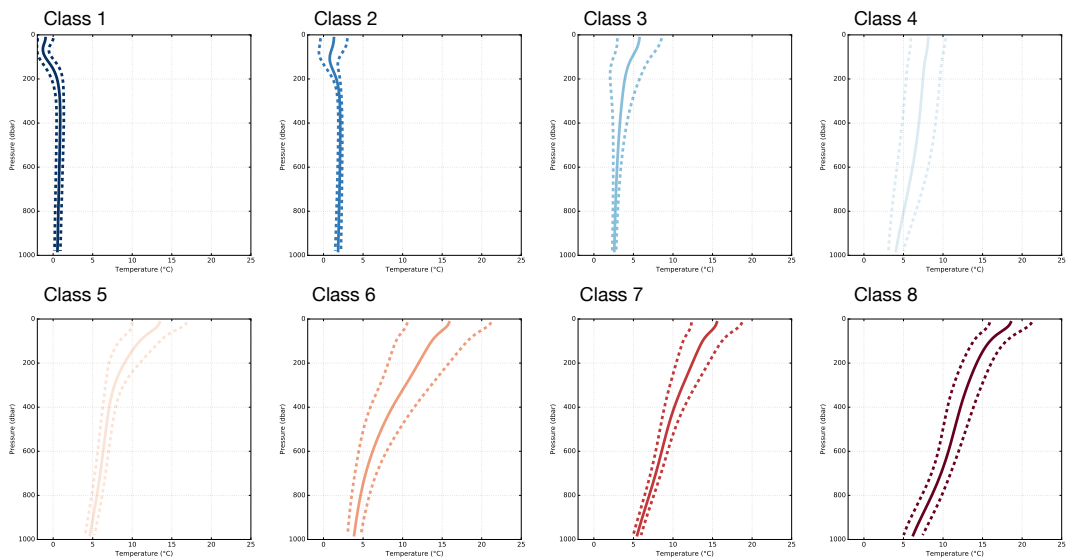
252 **Figure 5.** (a) GMM-derived class distribution for $K = 8$, shown with four fronts of the Antarctic Circum-
 253 polar Current, i.e. the Subantarctic Front (SAF), Polar Front (PF), Southern ACC Front (SACCF), and the
 254 Southern Boundary (SBDY) [Kim and Orsi, 2014]. (b) Class distribution shown in dynamic height space
 255 ($\phi_{1500\text{dbar}}^{300\text{dbar}}$). Note that only points with posterior probability ≥ 0.9 are shown. The classes are sorted by mean
 256 temperature, from coldest ($k = 1$) to warmest ($k = 8$).

260 The class nearest Antarctica (class 1) extends throughout the Weddell Gyre and coastal
 261 Antarctica (Figure 5a). The mean temperature profile in this region is inverted, that is, it is
 262 colder near the surface and warmer in the interior (Figure 6). This near-Antarctic class coin-
 263 cides with regions of Antarctic Bottom Water (AABW) export [Orsi *et al.*, 1999; Ohshima
 264 *et al.*, 2013], the subpolar Weddell and Ross gyres, and its northern extent approximately cor-
 265 responds with the classical Southern Boundary (SBDY) of the ACC [Kim and Orsi, 2014].
 266 This class occupies a narrow range in dynamic height space, with a class mean and standard
 267 deviation of 3.3 ± 0.2 cm ($\phi_{1500\text{dbar}}^{300\text{dbar}}$, Figure 5b); it is fairly distinct from the other classes, that
 268 is, class 1 profiles are rarely found north of the SBDY. For reference, Kim and Orsi [2014]
 269 associate the SBDY with the 3.1 cm dynamic height contour ($\phi_{1500\text{dbar}}^{500\text{dbar}}$). As their limits of
 270 integration over pressure are different than ours, this value of dynamic height is not directly
 271 applicable to our data, but it is roughly consistent with the gap between classes 1 and 2 in
 272 our analysis (Figure 5b). Assuming that the data features sufficiently uniform spatial cov-
 273 erage, gaps in dynamic height space may be indicative of fronts, as they may suggest sharp
 274 gradients in dynamic height over relatively short physical distances. We do not pursue this

Class	Number of profiles	Mean	Std. dev.	Min.	Max.
1	10680	0.48	0.81	-2.11	2.52
2	33031	1.83	0.72	-1.87	8.89
3	40268	3.38	1.50	-1.82	19.70
4	39619	6.36	2.24	-1.85	17.17
5	48252	7.32	2.56	2.76	25.37
6	48770	8.22	4.49	-1.88	27.56
7	38682	9.70	3.07	3.25	27.11
8	25130	11.57	3.43	3.56	28.08

257 **Table 1.** Temperature statistics for each class, using values from every pressure level. All temperature statis-
 258 tics are shown in °C. The classes have been sorted by mean temperature, calculated using values from all
 259 pressure levels.

275 analysis further here. For an in-depth analysis of SO front positions, see *Sokolov and Rintoul*
 276 [2009], for example.



277 **Figure 6.** Temperature profile statistics, separated by class, as functions of pressure. Shown are the mean
 278 (solid lines) and the mean plus or minus one standard deviation (dashed lines) for all profiles in the indicated
 279 class.

280 The second coldest class (class 2) is a circumpolar class with profiles that sit north of
 281 the SBDY and south of the Polar Front (PF) across all longitudes; it is the dominant class in
 282 the dynamic height range 4-6 cm, with a class mean value of 4.8 ± 0.7 cm ($\phi_{1500dbar}^{300dbar}$, Fig-
 283 ure 5). Its mean profile is also inverted, though not as sharply as the mean profile of class
 284 1 (Figure 6). A second circumpolar class (class 3) sits roughly north of the PF and south of
 285 the Subantarctic Front (SAF). In dynamic height space, class 3 is found between roughly 6-8
 286 cm, except in the Atlantic sector, where it extends to roughly 10 cm. For reference, *Kim and*
 287 *Orsi* [2014] associate the PF with the 5.0 cm dynamic height contour and the SAF with the
 288 7.0 cm dynamic height contour ($\phi_{1500dbar}^{500dbar}$). These values are roughly consistent with (but not
 289 directly comparable to) the gap positions in our data. Unlike the first two classes, the mean
 290 profile of class 3 is not inverted, that is, it gets colder with pressure. The presence of these
 291 two circumpolar classes is consistent with the homogenizing influence of the ACC, which
 292 typically encourages mixing along the strong jets associated with fronts and suppresses mix-
 293 ing across them [*Ferrari and Nikurashin*, 2010].

294 The profiles assigned to class 4 are mostly located north of the SAF in the Pacific and
 295 Indian sectors, roughly coinciding with regions of Subantarctic Mode Water (SAMW) and
 296 Antarctic Intermediate Water (AAIW) formation in the Pacific Ocean and south of Australia
 297 [*Sallée et al.*, 2010]. Despite its relatively narrow range in latitude, class 4 profiles occupy
 298 a broad, distinct range in dynamic height space in the Pacific Sector, with a class mean of
 299 11 ± 1.5 cm. The mean vertical profile associated with class 4 changes relatively gently with
 300 pressure, with no clear thermocline and a relatively large standard deviation across all pres-
 301 sures.

302 Profiles assigned to class 5 are mostly found in the Pacific Sector, in a region associ-
 303 ated with the export of SAMW and AAIW from the surface ocean into the interior thermo-
 304 cline [*Iudicone et al.*, 2007; *Jones et al.*, 2016]. In contrast with class 4, class 5 occupies a
 305 relatively large range in latitude and a relatively small range in dynamic height, with a mean
 306 and standard deviation of 12 ± 0.7 cm. The mean vertical profile has a clear thermocline over
 307 the upper 400 dbar of the ocean, with a standard deviation that narrows considerably with
 308 pressure. This class spatially coincides with the southern part of the South Pacific gyre, sug-
 309 gesting that gyre circulation tends to homogenize properties in this region.

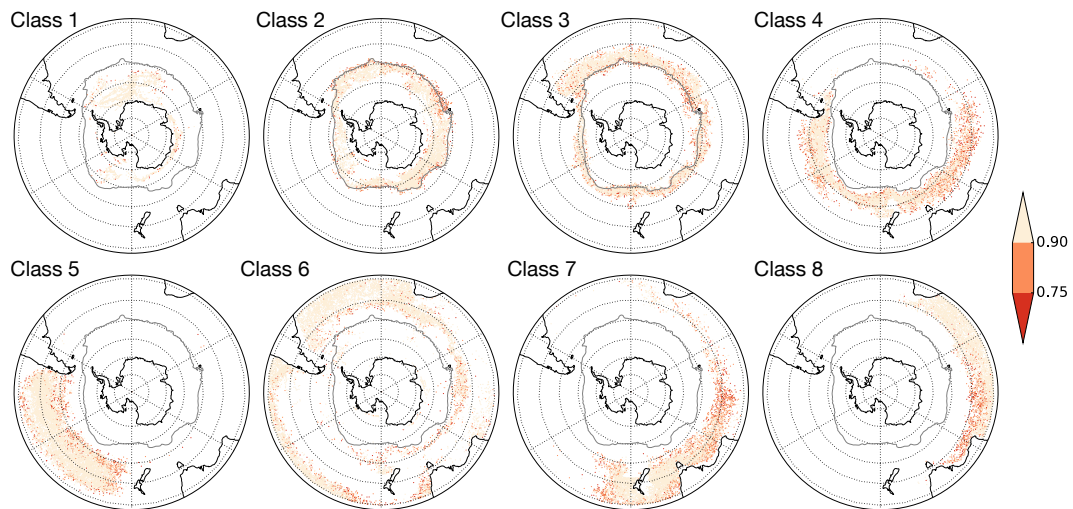
310 Class 6 highlights warmer subtropical waters and is mostly found in the Atlantic and
 311 Pacific sectors; it partially extends into the Indian sector, where it sits just north of the SAF.

312 From the surface to well into the interior, class 6 features some of the largest standard devia-
313 tions of any class, suggesting that class 6 consists of a wide variety of profiles; it can poten-
314 tially be split into a number of smaller classes. Classes 7 and 8 are also warmer subtropical
315 classes, with class 7 found mostly near Australia and New Zealand and class 8 found almost
316 exclusively in the Indian sector. Much of class 8 spatially coincides with the Indian Ocean
317 gyre. The spatial extent of class 8 near South Africa suggests that the Agulhas current influ-
318 ences the temperature structure in that region. The mean vertical profiles of classes 7 and 8
319 are similar, although class 7 features higher variability near the surface and class 8 features
320 slightly warmer surface temperatures. The higher variability in class 7 may be due to the
321 overlap of profiles in this class with a wider range of surface current features (e.g. boundary
322 currents around Australia and New Zealand, whereas class 8 largely overlaps with the Indian
323 Ocean gyre.

324 For a selected temperature profile, GMM predicts the probability distribution across
325 all K classes. That is, it calculates the probabilities that the profile belongs to each class k .
326 Next, the algorithm assigns the profile to the class with the highest probability. Note that
327 the sum of the posterior probabilities across all classes is one. Since these probabilities are
328 calculated with the full data set available, they are referred to as posterior probabilities. The
329 posterior probabilities are useful in their own right, as measures of confidence in GMM's
330 assignment of a profile to a particular class.

331 For our implementation of GMM on Argo temperature data, over 86% of the class as-
332 signments have posterior probabilities greater than 0.75, and over 74% of all class assign-
333 ments have posterior probabilities greater than 0.9 (Table 2). Class 1 features an especially
334 high percentage of very high posterior probabilities; over 90% of assignments into class
335 1 have posterior probabilities greater than or equal to 0.9. Outside of the Weddell Gyre,
336 we find the lowest posterior values in the Ross Sea and a few near-coastal areas (Figure 7).
337 The low posterior values could possibly be due to seasonal variability that is not well repre-
338 sented by a single class. Classes 2 and 3 also feature high posterior probabilities, for which
339 over 70% of assignments have values greater than or equal to 0.9. For both of these classes,
340 we find relatively low posterior probabilities upstream of Kerguelen Island (KI), clustered
341 around the PF. The area around KI is affected by upwelling, mixing, and the confluence of
342 the Agulhas Retroflexion and the ACC [Sallée *et al.*, 2010], and it also features relatively
343 high eddy diffusivities [Klocker and Abernathy, 2014]. The profiles in that area are influ-

344 enced by a number of competing processes and may be difficult to unambiguously separate
 345 into clear groups when using a value of K appropriate for the entire Southern Ocean.



346 **Figure 7.** Posterior probabilities for each class assignment, given the full cleaned, compressed dataset,
 347 shown together with the PF for reference [Kim and Orsi, 2014].

350 Although over 60% class 4 profiles have posterior values greater than or equal to 0.9,
 351 class 4 features some relatively low posterior values compared with the other classes, es-
 352 pecially in the Indian sector north of the SAF. In the Pacific sector, we find relatively low
 353 posteriors along the boundary between classes 4 and 5. Class 5 has a core of profiles with
 354 posterior values greater than or equal to 0.9, with relatively lower values all along its bound-
 355 ary. We find similar patterns for classes 6-8, except in the Indian sector between 60-120° E,
 356 north of the SAF. This region, which is downstream of Kerguelen Plateau, is characterized
 357 by relatively low posterior values for classes 4, 7, and 8. In general, although GMM performs
 358 well in all ocean basins, in terms of clear class separation with high posterior probabilities,
 359 its performance is somewhat weaker in the Indian sector.

360 4 Discussion

361 Here we explore the sensitivity of our results to the maximum number of classes K .
 362 We also explore a possible alternative to PCA that may be useful for incorporating salinity
 363 into our analysis, namely functional PCA.

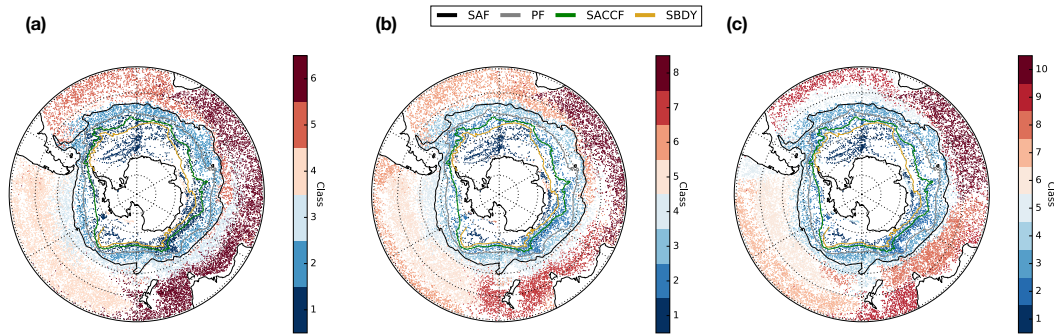
Class	[0.0, 0.50)	[0.50, 0.75)	[0.75, 0.9)	[0.9, 1.0]
1	<1	4	4	91
2	<1	11	11	77
3	<1	14	16	70
4	1	18	20	61
5	<1	7	8	84
6	<1	9	8	82
7	<1	19	17	64
8	<1	13	12	75

348 **Table 2.** Posterior probabilities for each class, divided into four unequal intervals. Each row shows the
349 percentage of profiles assigned to that class with posterior probabilities in the indicated range.

364 4.1 Sensitivity to number of classes K

365 In section 2, we estimated that the optimum number of classes K lies between 6 and
366 10. The weak constraint suggested by BIC allows for some tuning depending on the desired
367 level of complexity in the description of the dataset. Using $K = 6$ classes is sufficient to cap-
368 ture most of the large-scale structures identified in the $K = 8$ case, but there are some signif-
369 icant differences (Figure 8a,b). Specifically, there is one fewer circumpolar class, as classes
370 1-3 are reduced to classes 1-2 that sit roughly on either side of the PF. In the Pacific sector,
371 classes 4 and 5 merge into the new class 4. In the Indian sector, classes 7-8 merge into the
372 new class 6 that sits north of the SAF and south of Australia. We see that the overall descrip-
373 tion of ocean structure is simpler with $K = 6$; it is still a physically reasonable description of
374 ocean temperature structure, with circumpolar classes and clusters that span the major basins,
375 but it lacks some of the subtleties found in the $K = 8$ map.

378 As expected, the $K = 10$ case features more structure than the $K = 8$ case, and it is still
379 a physically reasonable distribution (Figure 8b,c). Classes 1-3 are still near-Antarctic or cir-
380 cumpolar classes; the additional structure all appears north of the SAF. In the Pacific basin,
381 the boundary between the $K = 8$ classes 5 and 6 and the $K = 10$ classes 6 and 7 is shifted
382 polewards, and a new class 5 is found along the Eastern Pacific, along the South American
383 coast. The $K = 10$ class 8 is found south of Australia, which in the $K = 8$ class is not a
384 distinct class. Interestingly, in the $K = 10$ case we find more profiles above 0.9 posterior



376 **Figure 8.** Comparison of GMM-derived classes, shown for (a) 6 classes, (b) 8 classes, and (c) 10 classes,
 377 along with fronts of the ACC [Kim and Orsi, 2014].

385 probability in the Indian sector, specifically in the region north of the SAF and between the
 386 longitudes of 60-120°E. Increasing K allowed for a more likely set of class assignments in
 387 this previously troublesome region. So, regions of low posterior probabilities may suggest a
 388 need for a higher value of K .

389 4.2 Functional PCA

390 In this work, we used PCA to reduce the dimensionality of our Argo temperature pro-
 391 file dataset. An alternative approach is to use functional principal component analysis (fPCA),
 392 in which PCA is performed on functions instead of the original data. In Pauthenet *et al.*
 393 [2017], the authors represent vertical temperature and salinity profiles from the Southern
 394 Ocean State Estimate [Mazloff *et al.*, 2010] as linear combinations of B-spline basis func-
 395 tions and apply fPCA to the resulting spline functions. They use the principal components to
 396 examine large-scale structures such as fronts in the Southern Ocean. Their approach offers
 397 another objective way to define water mass boundaries and could be used in concert with the
 398 GMM approach outlined in this work. This could offer a useful way to introduce salinity into
 399 the GMM analysis, which is especially relevant for stratification south of the PF [Pollard,
 400 2002].

401 5 Conclusions

402 We applied GMM, an unsupervised classification scheme, to Southern Ocean Argo
 403 temperature data above 1000 dbar. Without using longitude or latitude information, GMM
 404 identified spatially coherent patterns in the vertical temperature structure. The GMM-derived

405 classes broadly coincide with large-scale circulation and stratification features, including re-
406 gions of AABW formation and upwelling (i.e. adjacent to Antarctica), the ACC, formation
407 and export pathways of SAMW and AAIW, subtropical gyre circulation, and the Agulhas
408 Current and associated retroflection. We may say that GMM identifies *domains* in oceanographic
409 data, including gyre-dominated domains and circumpolar domains, among others.
410 GMM can be used to define these domains in a method that respects the structure of the
411 data, as opposed the simpler but physically unrealistic process of defining domains by simply
412 drawing rectangular boxes in latitude-longitude space. GMM also makes use of the interior
413 structure of the data, as opposed to only using surface variables like SSH. The class bound-
414 aries broadly coincide with several classically-defined fronts of the ACC, and the circumpo-
415 lar classes mostly occupy distinct regions in dynamic height space, indicating that GMM has
416 identified physically distinct profile types using only vertical temperature data. High poste-
417 rior probability distributions indicate regions where the classes are distinct and statistically
418 separate, whereas regions with low posterior probability indicate boundaries between classes
419 and/or regions of mixing influenced by a number of different temperature structures. GMM
420 may offer an alternative, complementary method for classification of Southern Ocean density
421 structures, and it is potentially useful for objectively and automatically comparing structures
422 across different observational and modeling datasets.

423 **Acronyms**

424 **AABW** Antarctic Bottom Water
 425 **AAIW** Antarctic Intermediate Water
 426 **ACC** Antarctic Circumpolar Current
 427 **BIC** Bayesian Information Criterion
 428 **fPCA** Functional principal component analysis
 429 **GDAC** Global Data Assembly Center
 430 **GMM** Gaussian mixture modeling
 431 **PC** Principal component
 432 **PCA** Principal component analysis
 433 **PDF** Probability distribution function
 434 **SAMW** Subantarctic Mode Water

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