

1 **Unsupervised clustering of Southern Ocean Argo float**
2 **temperature profiles**

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6 **Key Points:**

- 7 • We apply Gaussian mixture modeling (GMM) to Southern Ocean temperature data
- 8 • GMM identifies spatially coherent profile types without using latitude or longitude
- 9 information
- 10 • GMM offers a complementary approach for objectively classifying temperature pro-
- 11 files

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Abstract

The Southern Ocean has a complex density structure characterized by sharp fronts, steeply tilted isopycnals, and deep seasonal mixed layers. Methods of defining Southern Ocean density structures traditionally rely on somewhat ad-hoc combinations of physical, chemical, and dynamic properties. As a step towards an alternative approach for defining water masses, here we apply an unsupervised classification technique (that is, Gaussian mixture modelling or GMM) to Southern Ocean Argo float temperature profiles. GMM, without using any latitude or longitude information, automatically identifies several circumpolar classes influenced by the Antarctic Circumpolar Current. In addition, GMM identifies classes that bear the imprint of mode/intermediate water formation and export, large-scale gyre circulation, and the Agulhas Current, among others. Because GMM is robust, standardized, and automated, it can potentially be used to identify structures in both observational and model datasets, possibly making it a useful complement to existing classification techniques.

Plain Language Summary

The Southern Ocean is an important part of the climate system, in part because it absorbs a large fraction of the heat and carbon that is added to the atmosphere/ocean system by human-driven fossil fuel burning. In this work, we use a machine learning technique to automatically sort Southern Ocean temperature measurements into groups based on how those temperature measurements change with depth. Different groups have the fingerprints of different large-scale circulation patterns, such as the powerful Antarctic Circumpolar Current that flows around Antarctica. The groups that we identify are consistent with our understanding of the Southern Ocean, which gives us confidence that our machine learning technique may be useful for automatically grouping measurements and computer model data in the future. This matters because the climate science community needs a new set of tools, possibly including the machine learning technique that we use in this paper, to deal with a very large, ever-increasing volume of observational and computer model data.

1 Introduction

The Southern Ocean is a critical component of Earth's climate system, having thus far absorbed greater than 75% of the energy added via anthropogenic emissions and 50% of the excess carbon [Fletcher *et al.*; Frölicher *et al.*, 2015]. Its ability to absorb heat and carbon comes in part from its unique density structure and circulation, which features upwelling of cold, nutrient rich waters and regions of dense water formation [Lumpkin and Speer, 2007]. Characterizing and understanding the mean state and variability of Southern Ocean density structure remains an important and climatically-relevant goal of modern oceanography.

Through decades of effort, the oceanographic community has converged on a description of ocean structure that uses temperature, salinity, dynamical, and biogeochemical patterns to define different water masses (e.g. using potential vorticity minima to locate mode water pools) [Talley, 2013, and references therein]. These systematic approaches employ the understanding that water mass properties are "set" in their formation regions and modified by advection, mixing, and biogeochemical processes. This modern classification scheme is extremely useful and will continue to be useful well into the future, but it is not necessarily ideal for every application. Many of the temperature, salinity, and density values used to delimit one water mass from another are somewhat ad-hoc and very specific (e.g. boundaries between different types of mode water). These schemes are useful for observational data analysis but difficult to apply to numerical models of the ocean, which do not necessarily feature exactly the same structure as the observed ocean [Sallée *et al.*, 2013]. It is therefore prudent to develop and test alternative methods for the classification of different oceanic temperature, salinity, and density structures, as a complement to existing expertise-driven methods.

Maze *et al.* [2017] have shown that Argo temperature profile data from the North Atlantic Ocean can be usefully grouped into classes using Gaussian mixture modelling (GMM), an unsupervised classification technique. GMM describes the spatial structure of Argo profiles as a collection of Gaussian modes whose means and standard deviations generally vary with pressure. In this work, we apply GMM to Southern Ocean Argo temperature profiles in the upper 1000 m of the water column. We find that GMM identifies several circumpolar classes, gyres, the Agulhas current, and pathways broadly associated with the formation and export of mode and intermediate waters. In section 2, we describe the Argo dataset and the basics of GMM. In section 3, we present the results of applying GMM to Southern Ocean

70 Argo data, and in sections 4 and 5 we offer a brief discussion and summarize our conclu-
71 sions.

72 **2 Methods**

73 We applied an unsupervised classification method (i.e. Gaussian mixture modelling,
74 hereafter GMM) to Southern Ocean Argo float data. In this section, we briefly describe the
75 Argo dataset and the basics of GMM. We use the *scikit-learn* machine learning library for
76 Python (<http://scikit-learn.org/>), and the source code used for much of the analysis
77 in this paper is available via Github (<https://github.com/DanJonesOcean/OceanClustering>).
78 We refer the reader to *Maze et al.* [2017] for further detail on applying GMM to Argo float
79 data.

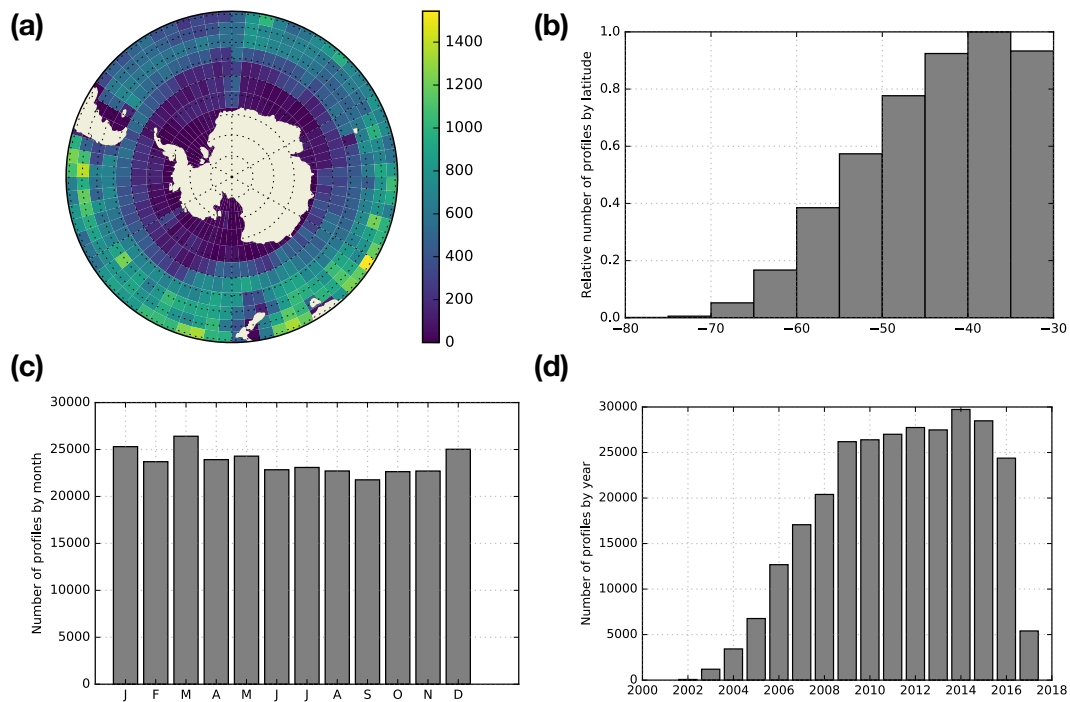
80 **2.1 Argo float dataset**

81 Argo floats are autonomous ocean instruments that measure, at minimum, the tempera-
82 ture and salinity of the ocean by periodically taking vertical profiles. Every 10 days, starting
83 at a “neutral” position of 1000 m, an Argo float dives down to 2000 m before rising to the
84 surface, taking a vertical profile of the water column along the way. The measurements are
85 transmitted via satellite and are ultimately made freely available via the Argo Global Data
86 Assembly Centers (GDACs) after some quality control checks. At the time of this writing,
87 over 3800 Argo floats are active in the global ocean, producing over 100,000 temperature and
88 salinity profiles per year with an average spacing of 3° (<http://www.argo.ucsd.edu/>).

89 For this study, we selected all available Argo profiles south of 30°S that have been
90 flagged by the GDACs as “observation good” (i.e. quality control flag = 1) covering the time
91 period from 2001 to early 2017. More specifically, we used a vertically interpolated product
92 with 400 pressure levels ranging from 0 to 2000 dbar. After discarding profiles with greater
93 than or equal to 6% NaN values (2% of the initial number of profiles) and discarding pres-
94 sure levels with greater than or equal to 3% NaN values, we were left with 284,427 profiles,
95 each with 192 pressure levels between 15 dbar and 980 dbar. We replaced all remaining NaN
96 values (\ll 1% of the total temperature measurements) with linearly interpolated estimates
97 using nearest neighbor values. We refer to the resulting dataset as the cleaned dataset.

98 Because of the autonomous and free-drifting nature of the floats, the profiles are not
99 distributed evenly in latitude/longitude (Figure 1). The profiles are more heavily concen-

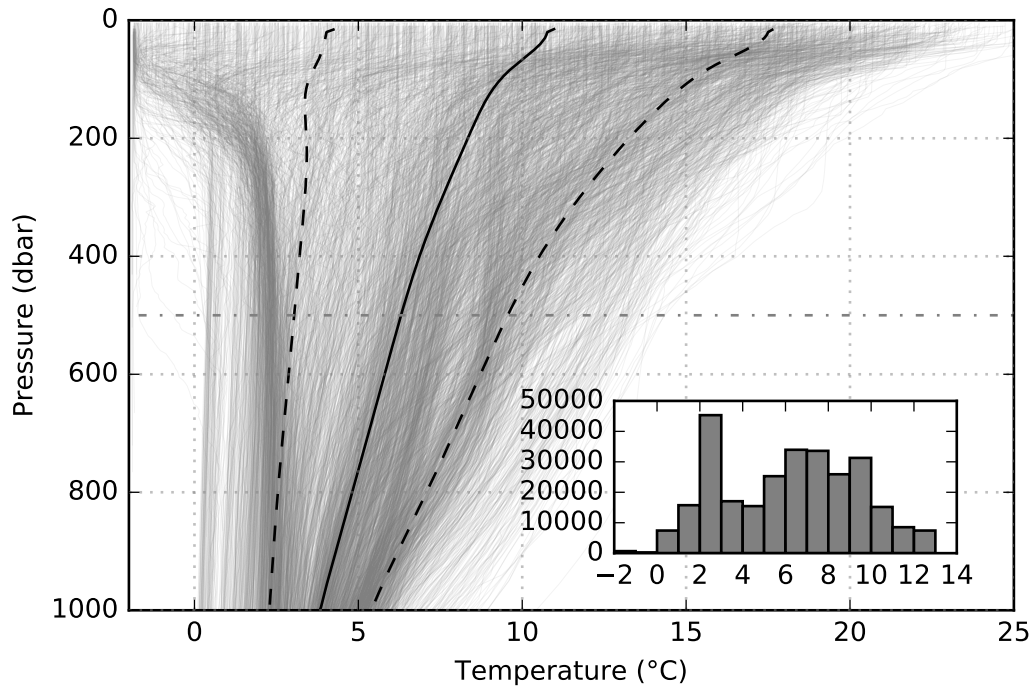
100 trated in the Pacific sector (roughly 890 profiles per degree longitude, totalling 47% of pro-
 101 files) and Indian sector (800 profiles per degree longitude, totalling 34% of profiles), with
 102 relatively fewer profiles in the Atlantic sector (610 profiles per degree longitude, 19% of
 103 total). When counted in equal-area bins and plotted by latitude, we see that the number or
 104 profiles decreases towards Antarctica (Figure 1(b)), which is partly due to challenging oper-
 105 ational conditions associated with seasonal sea ice, which can extend to just north of 60°S
 106 at maximum areal extent. The profiles are slightly over-represented in the Austral summer
 107 and autumn (DJF-MAM, 52% of profiles) and under-represented in the Austral winter and
 108 spring (JJA-SON, 48% of profiles), and the number of profiles increases until 2013 (Figure
 109 1(c,d)). Since we selected an Argo dataset that was created in early 2017, there are relatively
 110 few profiles from that year.



111 **Figure 1.** Distribution of Argo temperature profiles from the cleaned dataset. (a) Number of profiles in
 112 $5^\circ \times 5^\circ$ bins. (b) Relative number of profiles by latitude, scaled by an area-weighting factor $\cos(\phi)$, where ϕ is
 113 latitude. The temporal distribution of profiles shown by (c) month and (d) year.

114 The profiles selected for this study display a large variety of vertical temperature struc-
 115 tures (Figure 2). The range of temperatures is wider in the surface and considerably narrower
 116 with pressure, in part reflecting the seasonal cycle in upper ocean temperatures. A large
 117 number of profiles feature colder temperatures near the surface and warmer temperatures

118 in the interior, a physical arrangement that would be unstable to convection without the com-
 119 pensating effect of salinity. Water masses around Antarctica tend to be fresher at the surface
 120 and saltier in the interior due to glacial melt, freshwater flux, and the balance of evapora-
 121 tion/precipitation. This arrangement of temperature and salinity can be stable to vertical mix-
 122 ing (called “salt stratification”). In addition, the thermocline, i.e. the region of the ocean that
 123 features a rapid change in temperature with pressure, is visible in some temperature profiles.



124 **Figure 2.** Plot of 10% of the Argo temperature profiles, chosen at random, in the upper 1000 dbar of the
 125 cleaned dataset, along with the mean (solid line) and the mean plus or minus one standard deviation (dashed
 126 lines) across the entire dataset. The inset figure is a histogram of temperatures at 500 dbar (marked by a
 127 dash-dot line on the main figure) with temperature on the horizontal axis and count on the vertical axis.

128 2.2 Gaussian mixture modeling

129 Gaussian mixture modeling (GMM) is a probabilistic approach for describing and clas-
 130 sifying data. It attempts to fit (or “model” in the statistical sense) the data as a linear com-
 131 bination of multi-dimensional Gaussian distributions with unknown means and unknown
 132 standard deviations. Let \mathbf{X} be the array of N vertical profiles, each with D pressure levels,
 133 and let $p(\mathbf{X})$ be the probability distribution function (PDF) representing the entire dataset.

134 GMM represents the PDF as a weighted sum of K Gaussian classes, indexed by k , i.e.:

$$p(\mathbf{X}) = \sum_{k=1}^K \lambda_k \mathcal{N}(\mathbf{X}; \mu_k, \Sigma_k). \quad (1)$$

135 Here, $\mathcal{N}(\mathbf{x}; \mu_k, \Sigma_k)$ is the multi-dimensional Gaussian PDF with a vector of means μ_k and
 136 covariance matrix Σ_k , i.e.:

$$\mathcal{N}(\mathbf{x}; \mu_k, \Sigma_k) = \frac{\exp\left[-\frac{1}{2}(\mathbf{x} - \mu_k)^T \Sigma_k^{-1} (\mathbf{x} - \mu_k)\right]}{\sqrt{(2\pi)^D |\Sigma_k|}}. \quad (2)$$

137 The probability associated with class k is $p(k) = \lambda_k$. The probability of profile \mathbf{x} being in
 138 class k is $p(k|\mathbf{x}) = \lambda_k \mathcal{N}(\mathbf{x}; \mu_k, \Sigma_k) / p(\mathbf{x})$, where the vector \mathbf{x} is a single profile taken from
 139 the complete array \mathbf{X} and $p(\mathbf{x})$ is equation (1) with a single profile \mathbf{x} as the argument, i.e. a
 140 normalizing factor. Both \mathbf{x} and μ_k are vectors of length D .

141 Starting with random initial guesses for the classes, GMM proceeds by iteratively ad-
 142 justing the means μ_k and standard deviations Σ_k (i.e. the ‘‘parameters’’) of the classes in or-
 143 der to maximize a logarithmic measure of likelihood, i.e.:

$$\log[p(\mathbf{X})] = \sum_{i=1}^N \log \left[\sum_{k=1}^K \lambda_k \mathcal{N}(\mathbf{X}_i; \mu_k, \Sigma_k) \right]. \quad (3)$$

144 GMM uses an expectation-maximization approach, described in *Maze et al.* [2017]. This
 145 algorithm monotonically converges on a local maximum. GMM is a generalization of k -
 146 means clustering, which only attempts to minimize in-group variance by shifting the means.
 147 By contrast, GMM attempts to identify means and standard deviations, allowing for some
 148 variation about the centres of the Gaussian distributions.

149 In our instance of GMM, each pressure level is treated as a ‘‘dimension’’, and the Gaus-
 150 sian parameters are associated with each pressure level. However, we may not need all of
 151 these pressure levels to accurately describe the dataset, as ocean temperature changes much
 152 more rapidly in the mixed layer and thermocline than in the interior. In order to reduce the
 153 computational complexity of the problem, we transform the profile data from pressure space
 154 to an alternative space using principal component analysis (PCA). Specifically, we calculate
 155 principal components that capture a desired fraction of the vertical variability of the dataset.
 156 Each eigenvector may be thought of as a ‘‘profile type’’ that describes a certain amount of
 157 variance in the data with pressure (note that this is not necessarily the same thing as a ‘‘typi-
 158 cal profile’’). We calculate J principal components via the transformation:

$$\mathbf{X}(z) = \sum_{j=1}^J \mathbf{P}(z, j) \mathbf{Y}(j), \quad (4)$$

159 where z is the pressure level, J is the total number of principal components (index j), and
 160 $\mathbf{P}(z, j)$ is the transformation matrix between pressure space and principal component space.
 161 This strategy is an example of “dimensionality reduction”, which is common in machine
 162 learning approaches. We find that $J = 6$ captures 99.9% of the variance in the vertical struc-
 163 ture, which greatly reduces the number of dimensions needed to describe the Argo profile
 164 data used here, i.e. from 194 pressure levels to 6 principal components.

165 2.2.1 Selecting the number of classes

166 GMM does have one free parameter, i.e. the maximum number of classes K . In or-
 167 der to determine the most appropriate value for K , we applied a statistical test, namely a
 168 Bayesian Information Criterion (BIC). BIC uses an empirically formulated cost function that
 169 rewards likelihood and penalizes the number of classes K :

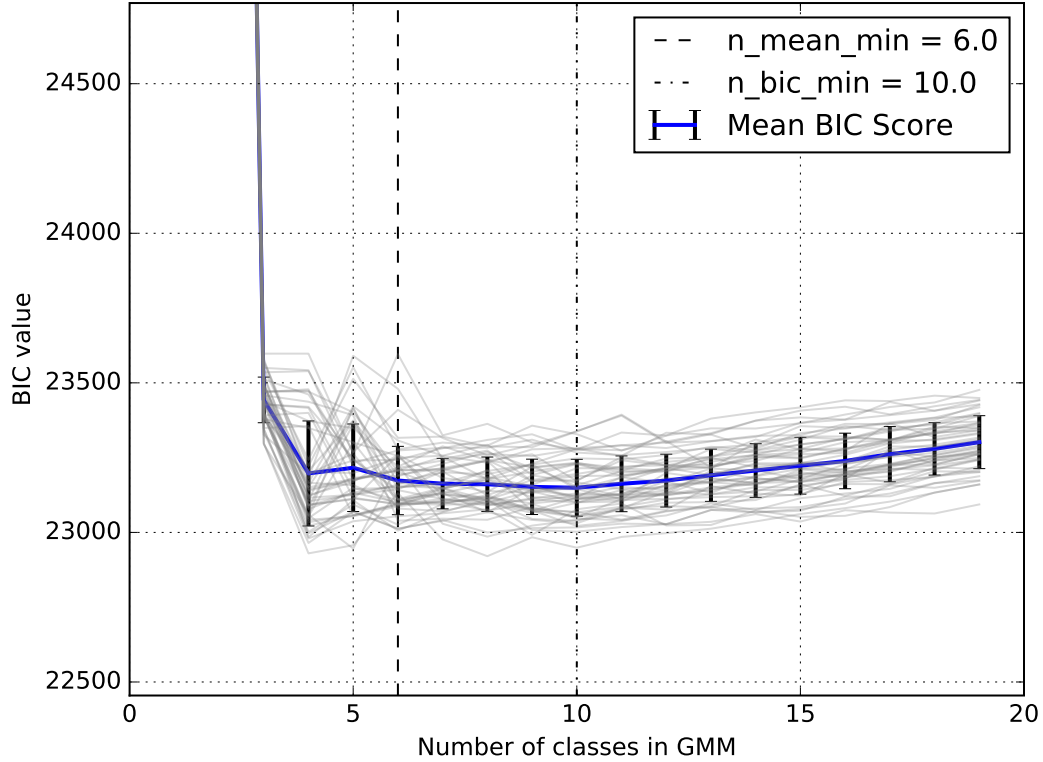
$$BIC(K) = -2\mathcal{L}(K) + N_f(K) \log(n), \quad (5)$$

170 where \mathcal{L} is a measure of likelihood, n is the number of profiles used in the BIC test, and N_f
 171 is the number of independent parameters to be estimated:

$$N_f(K) = K - 1 + KD + \frac{KD(D - 1)}{2}. \quad (6)$$

172 In this framework, the optimum value of K minimizes the BIC score. We perform a number
 173 of BIC tests, using different subsets of the data and different values of K , to estimate the dis-
 174 tribution and variability of BIC. Using the roughly 300 km decorrelation scale of the South-
 175 ern Ocean as guidance [Ninove *et al.*, 2016], we randomly select a profile from each $4^\circ \times 4^\circ$
 176 grid cell, returning 884 random profiles for each BIC test. We calculate BIC scores for each
 177 set of 884 random profiles (in principal component space) using a range of classes K from 1
 178 to 19 (Figure 3). BIC analysis does not feature a clear minimum, but instead it suggests that
 179 the optimum value of K lies between 6 and 10.

185 It may seem counterintuitive that BIC does not return a single optimum value for K ,
 186 but this is consistent with the nature of K as a weakly constrained free parameter that con-
 187 trols the level of complexity of the statistical description of the dataset. Oceanography has
 188 a rich history of expertise-driven clustering using physical and biogeochemical criteria (e.g.
 189 PV minima, oxygen minima) and the fingerprints of various processes (e.g. gyre circula-
 190 tion). These descriptions can be arranged into hierarchies, from coarse/simple (e.g. two-layer
 191 quasi-geostrophic models) to rich and complex (e.g. the descriptions found in Talley [2013]).



180 **Figure 3.** Bayesian Information Criteria (BIC) scores versus the specified number of classes K . Shown are
 181 the individual trials for different subsets of the temperature profile datasets (gray lines), the mean (blue line),
 182 and standard deviations computed from the profiles (error bars). The dashed vertical line represents the average
 183 of the minimums from each profile and the dash-dot vertical line represents the minimum of the average
 184 of the profiles. These two minima indicate a range of suitable values for the maximum number of classes K .

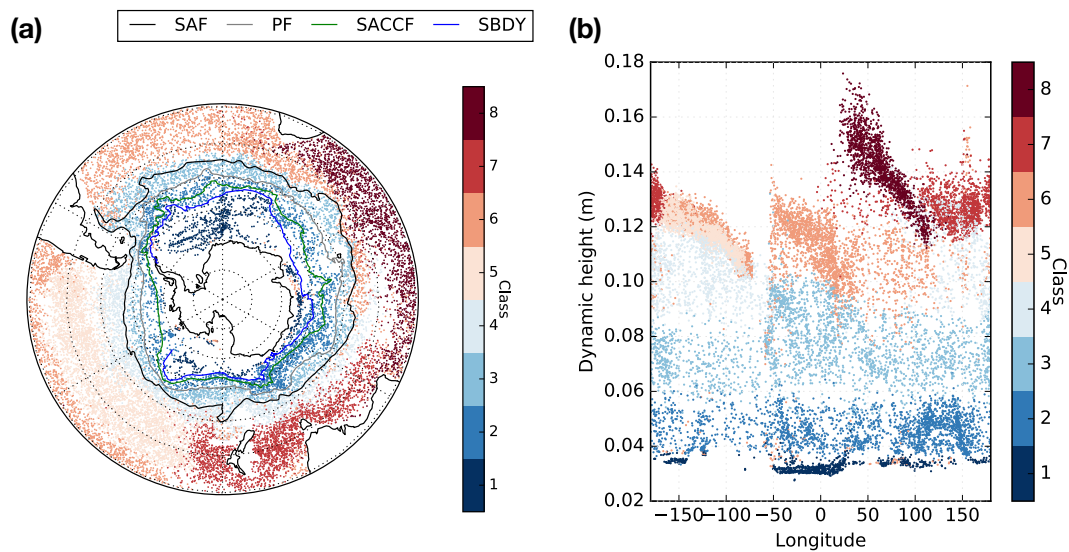
192 The level of detail required in the description depends on the application at hand. For exam-
 193 ple, a simple β -plane model is sufficient to explain the existence of gyres and western bound-
 194 ary currents; it constitutes a first-order description of gyres. Algorithmic clustering offers a
 195 robust way to traverse this hierarchy using a range of K values. Although statistical tests can
 196 be used as rough guides for choosing the number of classes, there is not necessarily a single
 197 ideal value for K . We explore the impact of K on our results in section 4.

198 We used a “training” dataset, a subset of the cleaned dataset, to estimate the param-
 199 eters (i.e the means and standard deviations) of the GMM classes. To generate the GMM
 200 training set, we randomly selected a single profile from each $1^\circ \times 1^\circ$ bin. Each training
 201 dataset contains 12,286 profiles (roughly 4% of the cleaned dataset), distributed evenly in lat-

202 itude/longitude space. We then statistically represent (i.e. ‘model’) the entire cleaned dataset
 203 with the fitted Gaussian model using optimized parameters. The end result is a probabilistic
 204 description of the cleaned Argo temperature profile dataset in terms of a linear combination
 205 of Gaussian distributions that vary with pressure. Each profile then has a probability distri-
 206 bution across the classes, and the profile is assigned to the class with the highest probability.

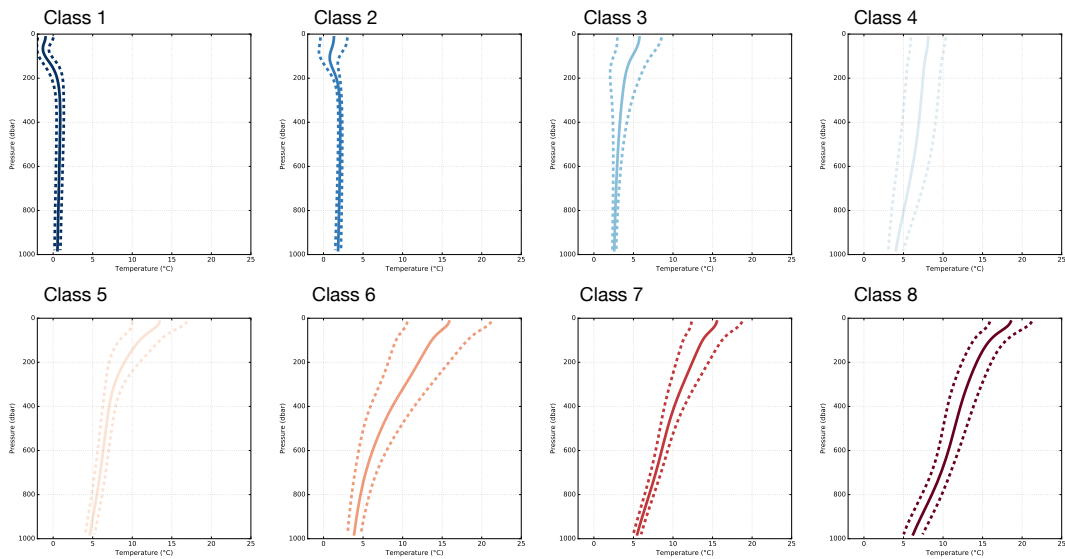
207 3 Results

208 In order to identify patterns in the temperature structure of the Southern Ocean, we de-
 209 scribe the cleaned Argo temperature profile dataset as a linear combination of multi-dimensional
 210 Gaussian functions that vary with pressure, using $K = 8$ different classes. Despite the fact
 211 that GMM does not have access to the longitudes and latitudes of the profiles, it identifies
 212 spatially coherent structures, some of which are roughly demarcated by the fronts of the
 213 ACC as defined by *Kim and Orsi* [2014] (Figure 4). For ease of interpretation, we sorted the
 214 classes by mean temperature (Table 1).



215 **Figure 4.** (a) GMM-derived class distribution for $K = 8$, shown with four fronts of the Antarctic Circum-
 216 polar Current, i.e. the Subantarctic Front (SAF), Polar Front (PF), Southern ACC Front (SACCF), and the
 217 Southern Boundary (SBDY) [*Kim and Orsi*, 2014]. (b) Class distribution shown in dynamic height space.
 218 Note that only points with posterior probability ≥ 0.9 are shown. The classes are sorted by mean temperature,
 219 from coldest ($k = 1$) to warmest ($k = 8$).

220 The class nearest Antarctica (class 1) extends throughout the Weddell Gyre and coastal
 221 Antarctica (Figure 4a). The mean temperature profile in this region is inverted, that is, it is
 222 colder near the surface and warmer in the interior (Figure 5). This near-Antarctic class co-
 223 coincides with regions of Antarctic Bottom Water (AABW) export [Orsi *et al.*, 1999], and its
 224 northern extent approximately corresponds with the classical Southern Boundary (SBDY) of
 225 the ACC [Kim and Orsi, 2014]. This class occupies a narrow range in dynamic height space,
 226 with a class mean and standard deviation of 3.3 ± 0.2 cm (Figure 4b); it is fairly distinct from
 227 the other classes, that is, class 1 profiles are rarely found north of the SBDY.



228 **Figure 5.** Temperature profile statistics, separated by class, as functions of pressure. Shown are the mean
 229 (solid lines) and the mean plus or minus one standard deviation (dashed lines) for all profiles in the indicated
 230 class.

231 The second coldest class (class 2) is a circumpolar class with profiles that sit north of
 232 the SBDY and south of the Polar Front (PF) across all longitudes; it is the dominant class in
 233 the dynamic height range 4-6 cm, with a class mean value of 4.8 ± 0.7 cm (Figure 4). Its
 234 mean profile is also inverted, though not as sharply as the mean profile of class 1 (Figure 5).
 235 A second circumpolar class (class 3) sits roughly north of the PF and south of the Subantarctic
 236 Front (SAF). In dynamic height space, class 3 is found between roughly 6-8 cm, except in
 237 the Atlantic sector, where it extends to roughly 10 cm. Unlike the first two classes, the mean
 238 profile of class 3 is not inverted, that is, it gets colder with pressure. The presence of these
 239 two circumpolar classes is consistent with the homogenizing influence of the ACC, which

240 typically encourages mixing along the strong jets associated with fronts and suppresses mix-
241 ing across them [*Ferrari and Nikurashin, 2010*].

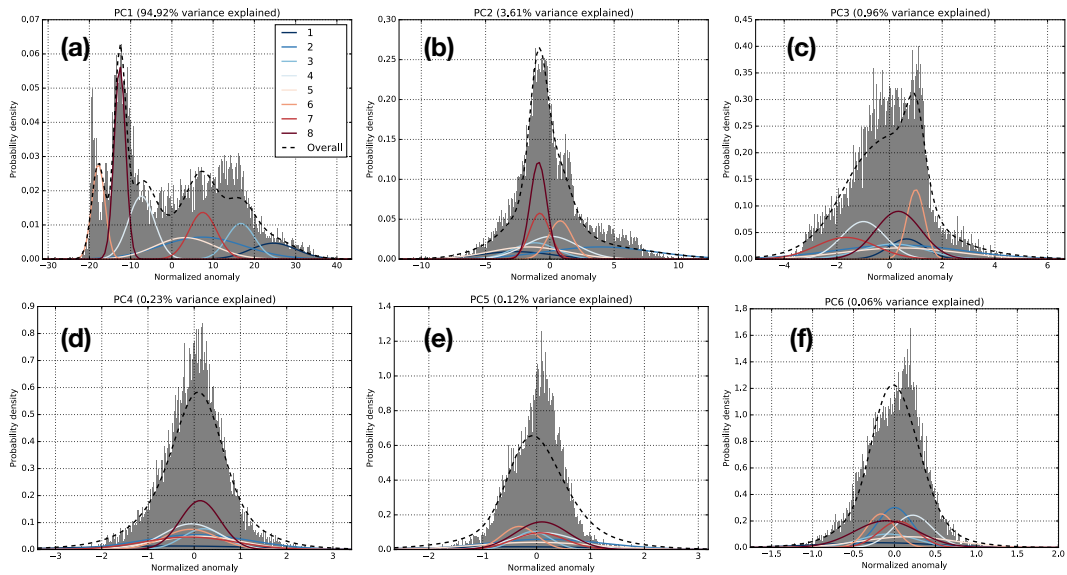
242 The profiles assigned to class 4 are mostly located north of the SAF in the Pacific and
243 Indian sectors, roughly coinciding with regions of Subantarctic Mode Water (SAMW) and
244 Antarctic Intermediate Water (AAIW) formation in the Pacific Ocean and south of Australia
245 [*Sallée et al., 2010*]. Despite its relatively narrow range in latitude, class 4 profiles occupy
246 a broad, distinct range in dynamic height space in the Pacific Sector, with a class mean of
247 11 ± 1.5 cm. The mean vertical profile associated with class 4 changes relatively gently with
248 pressure, with no clear thermocline and a relatively large standard deviation across all pres-
249 sures.

250 Profiles assigned to class 5 are mostly found in the Pacific Sector, in a region associ-
251 ated with the export of SAMW and AAIW from the surface ocean into the interior thermo-
252 cline [*Iudicone et al., 2007; Jones et al., 2016*]. In contrast with class 4, class 5 occupies a
253 relatively large range in latitude and a relatively small range in dynamic height, with a mean
254 and standard deviation of 12 ± 0.7 cm. The mean vertical profile has a clear thermocline over
255 the upper 400 dbar of the ocean, with a standard deviation that narrows considerably with
256 pressure.

257 Class 6 highlights warmer subtropical waters and is mostly found in the Atlantic and
258 Pacific sectors; it partially extends into the Indian sector, where it sits just north of the SAF.
259 From the surface to well into the interior, class 6 features some of the largest standard devia-
260 tions of any class, suggesting that class 6 consists of a wide variety of profiles; it can poten-
261 tially be split into a number of smaller classes. Classes 7 and 8 are also warmer subtropical
262 classes, with class 7 found mostly near Australia and New Zealand and class 8 found almost
263 exclusively in the Indian sector. The spatial extent of class 8 near South Africa suggests that
264 the Agulhas current influences the temperature structure in that region. The mean vertical
265 profiles of classes 7 and 8 are similar, although class 7 features higher variability near the
266 surface and class 8 features slightly warmer surface temperatures.

267 As discussed in section 2, in order to make the GMM algorithm more efficient, we
268 used PCA to reduce the number of variables required to represent the vertical structure of
269 the temperature profiles, from over 190 pressure levels down to six principal components.
270 Each PC is a vertical profile that can “explain”, in the statistical sense of being correlated
271 with, a fraction of the variance in temperature with pressure. Nearly 95% of the variance

272 is explained by the first PC (i.e. PC1), and the Gaussian functions associated with PC1 are
 273 relatively distinct, capturing the broad shape of the temperature distribution (Figure 6). For
 274 higher indexed PCs, the Gaussians overlap more, but their sum still makes up a representa-
 275 tion of the temperature distribution that is sufficiently accurate for our purposes.

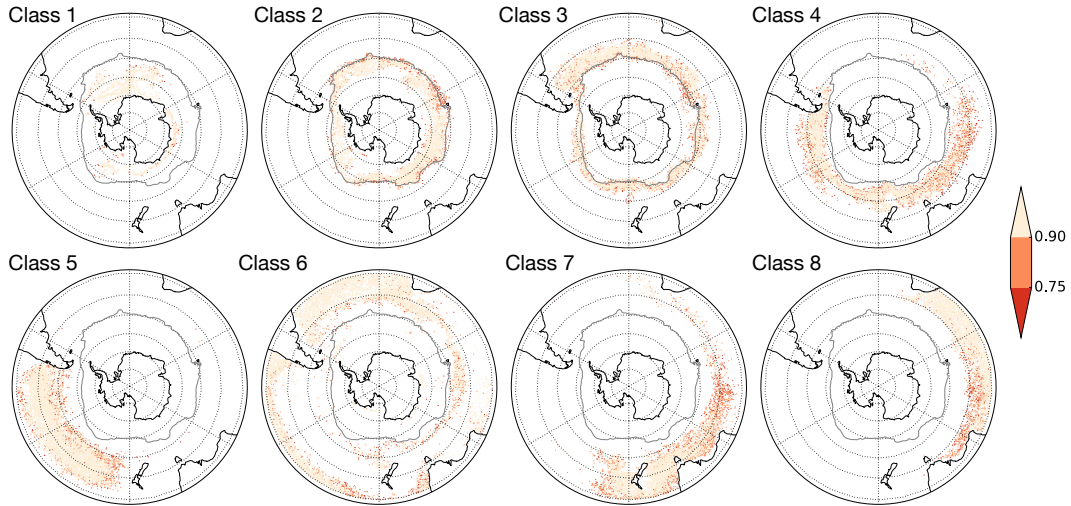


276 **Figure 6.** Probability density functions for the principal component amplitude coefficients associated with
 277 each profile, along with the Gaussian functions generated by GMM with $K = 8$ classes.

278 For a selected temperature profile, GMM predicts the probability distribution across
 279 all K classes. That is, it calculates the probabilities that the profile belongs to each class k .
 280 Next, the algorithm assigns the profile to the class with the highest probability. Since these
 281 probabilities are calculated with the full data set available, they are referred to as posterior
 282 probabilities. The posterior probabilities are useful in their own right, as measures of confi-
 283 dence in GMM's assignment of a profile to a particular class.

284 For our implementation of GMM on Argo temperature data, over 86% of the class as-
 285 signments have posterior probabilities greater than 0.75, and over 74% of all class assign-
 286 ments have posterior probabilities greater than 0.9 (Table 2). Class 1 features an especially
 287 high percentage of very high posterior probabilities; over 90% of assignments into class
 288 1 have posterior probabilities greater than or equal to 0.9. Outside of the Weddell Gyre,
 289 we find the lowest posterior values in the Ross Sea and a few near-coastal areas (Figure 7).
 290 Classes 2 and 3 also feature high posterior probabilities, for which over 70% of assignments

291 have values greater than or equal to 0.9. For both of these classes, we find relatively low pos-
 292 terior probabilities upstream of Kerguelen Island (KI), clustered around the PF.



293 **Figure 7.** Posterior probabilities for each class assignment, given the full cleaned dataset, shown together
 294 with the PF for reference [Kim and Orsi, 2014].

295 Although over 60% class 4 profiles have posterior values greater than or equal to 0.9,
 296 class 4 features some relatively low posterior values compared with the other classes, es-
 297 pecially in the Indian sector north of the SAF. In the Pacific sector, we find relatively low
 298 posteriors along the boundary between classes 4 and 5. Class 5 has a core of profiles with
 299 posterior values greater than or equal to 0.9, with relatively lower values all along its bound-
 300 ary. We find similar patterns for classes 6-8, except in the Indian sector between 60-120° E,
 301 north of the SAF. This region, which is downstream of Kerguelen Plateau, is characterized by
 302 relatively low posterior values for classes 4, 7, and 8. The area around KI is affected by up-
 303 welling, mixing, and the confluence of the Agulhas Retroflection and the ACC [Sallée et al.,
 304 2010], and it also features relatively high eddy diffusivities [Klokker and Abernathy, 2014].
 305 The profiles in that area are influenced by a number of competing processes and may be dif-
 306 ficult to unambiguously separate into clear groups when using a value of K appropriate for
 307 the entire Southern Ocean. In general, although GMM performs well in all ocean basins, in
 308 terms of clear class separation with high posterior probabilities, its performance is somewhat
 309 weaker in the Indian sector.

4 Discussion

4.1 Sensitivity to number of classes K

In section 2, we estimated that the optimum number of classes K lies between 6 and 10. The weak constraint suggested by BIC allows for some tuning depending on the desired level of complexity in the description of the dataset. Using $K = 6$ classes is sufficient to capture most of the large-scale structures identified in the $K = 8$ case, but there are some significant differences (Figure 8). Specifically, there is one fewer circumpolar class, as classes 1-3 are reduced to classes 1-2 that sit roughly on either side of the PF. In the Pacific sector, classes 4 and 5 merge into the new class 4. In the Indian sector, classes 7-8 merge into the new class 6 that sits north of the SAF and south of Australia. We see that the overall description of ocean structure is simpler with $K = 6$; it is still a physically reasonable description of ocean temperature structure, with circumpolar classes and clusters that span the major basins, but it lacks some of the subtleties found in the $K = 8$ map.

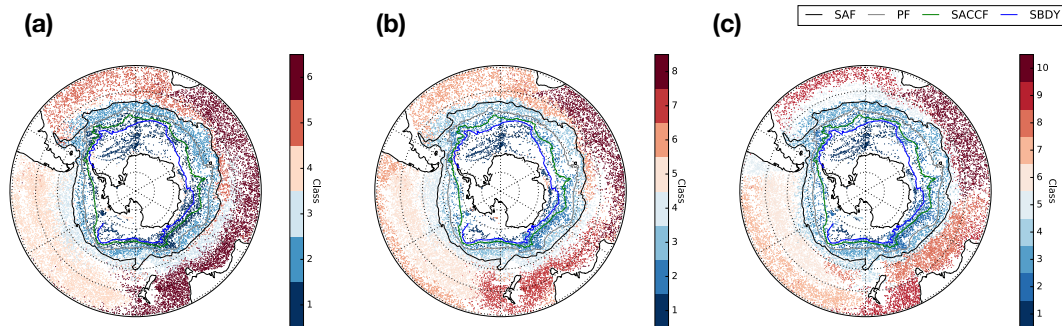


Figure 8. Comparison of GMM-derived classes, shown for (a) 6 classes, (b) 8 classes, and (c) 10 classes, along with fronts of the ACC [Kim and Orsi, 2014].

As expected, the $K = 10$ case features more structure than the $K = 8$ case, and it is still a physically reasonable distribution. Classes 1-3 are still near-Antarctic or circumpolar classes; the additional structure all appears north of the SAF. In the Pacific basin, the boundary between the $K = 8$ classes 5 and 6 and the $K = 10$ classes 6 and 7 is shifted polewards, and a new class 5 is found along the Eastern Pacific, along the South American coast. The $K = 10$ class 8 is found south of Australia, which in the $K = 8$ class is not a distinct class. Interestingly, in the $K = 10$ case we find more profiles above 0.9 posterior probability in the Indian sector, specifically in the region north of the SAF and between the longitudes of 60-120°E. Increasing K allowed for a more likely set of class assignments in this previ-

334 ously troublesome region. So, regions of low posterior probabilities may suggest a need for a
335 higher value of K .

336 4.2 Functional PCA

337 In this work, we used PCA to reduce the dimensionality of our Argo temperature pro-
338 file dataset. An alternative approach is to use functional principal component analysis (fPCA),
339 in which PCA is performed on functions instead of the original data. In *Pauthenet et al.*
340 [2017], the authors represent vertical temperature and salinity profiles from the Southern
341 Ocean State Estimate [*Mazloff et al.*, 2010] as linear combinations of B-spline basis func-
342 tions and apply fPCA to the resulting spline functions. They use the principal components
343 to examine large-scale structures such as fronts in the Southern Ocean. Their approach of-
344 fers another objective way to define water mass boundaries and could be used in concert with
345 the GMM approach outlined in this work. This could offer a useful way to introduce salinity
346 into the GMM analysis, which is especially relevant for stratification south of the PF [*Pollard*
347 *et al.*, 2002].

348 5 Conclusions

349 We applied Gaussian mixture modeling (GMM), an unsupervised classification scheme,
350 to Southern Ocean Argo temperature data above 1000 dbar. Without using longitude or lat-
351 itude information, GMM identified spatially coherent patterns in the vertical temperature
352 structure. The GMM-derived classes broadly coincide with large-scale circulation and strat-
353 ification features, including regions of AABW formation and upwelling (i.e. adjacent to
354 Antarctica), the ACC, formation and export pathways of SAMW and AAIW, subtropical
355 gyre circulation, and the Agulhas Current and associated retroflexion. The class bound-
356 aries broadly coincide with several classically-defined fronts of the ACC, and the circum-
357 polar classes mostly occupy distinct regions in dynamic height space, indicating that GMM
358 has identified physically distinct profile types using only vertical temperature data. Posterior
359 probability distributions indicate regions where the classes are distinct and statistically sep-
360 arate, whereas regions with low posterior probability indicate boundaries between classes
361 and/or regions of mixing influenced by a number of different temperature structures. GMM
362 may offer an alternative, complementary method for classification of Southern Ocean density
363 structures, and it is potentially useful for objectively and automatically comparing structures
364 across different observational and modeling datasets.

| Class | Number of profiles | Mean | Std. dev. | Min. | Max. |
|-------|--------------------|-------|-----------|-------|-------|
| 1 | 10680 | 0.48 | 0.81 | -2.11 | 2.52 |
| 2 | 33031 | 1.83 | 0.72 | -1.87 | 8.89 |
| 3 | 40268 | 3.38 | 1.50 | -1.82 | 19.70 |
| 4 | 39619 | 6.36 | 2.24 | -1.85 | 17.17 |
| 5 | 48252 | 7.32 | 2.56 | 2.76 | 25.37 |
| 6 | 48770 | 8.22 | 4.49 | -1.88 | 27.56 |
| 7 | 38682 | 9.70 | 3.07 | 3.25 | 27.11 |
| 8 | 25130 | 11.57 | 3.43 | 3.56 | 28.08 |

365 **Table 1.** Temperature statistics for each class, using values from every pressure level. All temperature statis-
366 tics are shown in °C. The classes have been sorted by mean temperature, calculated using values from all
367 pressure levels.

| Class | [0.0, 0.50) | [0.50, 0.75) | [0.75, 0.9) | [0.9, 1.0] |
|-------|-------------|--------------|-------------|------------|
| 1 | <1 | 4 | 4 | 91 |
| 2 | <1 | 11 | 11 | 77 |
| 3 | <1 | 14 | 16 | 70 |
| 4 | 1 | 18 | 20 | 61 |
| 5 | <1 | 7 | 8 | 84 |
| 6 | <1 | 9 | 8 | 82 |
| 7 | <1 | 19 | 17 | 64 |
| 8 | <1 | 13 | 12 | 75 |

368 **Table 2.** Posterior probabilities for each class, divided into four unequal intervals. Each row shows the
369 percentage of profiles assigned to that class with posterior probabilities in the indicated range.

370 Acronyms

371 **AABW** Antarctic Bottom Water

372 **AAIW** Antarctic Intermediate Water

373 **ACC** Antarctic Circumpolar Current

374 **ARGO** Array for Real-time Geostrophic Oceanography

375 **BIC** Bayesian Information Criterion

376 **fPCA** Functional principal component analysis

377 **GDAC** Global Data Assembly Center

378 **GMM** Gaussian mixture modeling

379 **PC** Principal component

380 **PCA** Principal component analysis

381 **PDF** Probability distribution function

382 **SAMW** Subantarctic Mode Water

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