

Detecting lightning infrasound using a high-altitude balloon

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Key Points:

- First lightning infrasound detected using free-flying balloon at stratospheric altitudes over Tasman Sea in May 2016.
- Infrasonic signals matched with a few lightning strokes within 100 km range of balloon as it flew over at least two thunderclouds.
- Strong atmospheric effects likely reduced the number of recorded signals generated by lightning strokes within range of the balloon.

Abstract

Acoustic waves with a wide range of frequencies are generated by lightning strokes during thunderstorms, including infrasonic waves (0.1 to 20 Hz). The source mechanism for these low frequency acoustic waves is still debated and studies have so far been limited to ground-based instruments. Here we report the first confirmed detection of lightning generated infrasound with acoustic instruments suspended at stratospheric altitudes using a free-flying balloon. We observe high-amplitude signals generated by lightning strokes located within 100 km of the balloon as it flew over the Tasman Sea on 17 May 2016. The signals share many characteristics with waveforms recorded previously by ground-based instruments near thunderstorms. The ability to measure lightning activity with high-altitude infrasound instruments has demonstrated the potential for using these platforms to image the full acoustic wavefield in the atmosphere.

Plain-language summary

Lightning generates sound waves across a wide range of frequencies, including below the threshold for human hearing at 20 Hz. How these waves at less than 20 Hz, also known as infrasound waves, are generated during a lightning stroke is currently an area for debate. So far, measurements of lightning infrasound waves have been limited to microphones fixed to the ground and models have shown that only a small section of sound waves actually reach the ground. Here we show lightning infrasound that has been detected using microphones suspended over a thunderstorm using a balloon flying at 32 km height. This opens up the possibility of using balloons in future studies to make better measurements of infrasound waves generated by lightning activity and in turn, give a better idea of how they are generated.

1 Introduction

Acoustic signals with frequencies between 0.02 to 20 Hz are classified as infrasound and are not audible to humans. A wide variety of sources have been found to generate infrasound, including: volcanoes, earthquakes, avalanches, tsunamis, meteors, aurora, thunderstorms, wind-mountain interactions, supersonic aircraft, rockets, and chemical and nuclear explosions [*Campus and Christie*, 2010]. Infrasonic signals can travel hundreds to thousands of kilometers through the atmosphere, sampling areas from the Earth's surface up to the thermosphere. A variety of institutions maintain arrays for monitoring purposes, such as volcano observatories [*Fee and Matoza*, 2013] or the International Monitoring System [*Christie and Campus*, 2010]. The vast majority of infrasound studies currently use ground-based instrument arrays and networks. Infrasound instruments deployed on the ground may be subject to high-levels of background noise which may obscure the signals of interest, or the signals may arrive distorted due to topographic or atmospheric propagation effects [e.g. *Lacanna and Ripepe*, 2013; *Kim and Lees*, 2014]. Furthermore, the intensity of the acoustic wavefield may be as much as 50% greater directly above the source compared that at a similar horizontal distance [*Blackstock*, 2000].

A series of studies have recently taken place to explore how to fill this gap in our ability to monitor the atmospheric acoustic wavefield. These experiments have tested the use of microphones suspended underneath free-floating balloons to record infrasound at high-altitude [e.g. *Bowman and Lees*, 2015, 2017; *Bowman and Albert*, 2018]. Balloon deployments conducted as part of the NASA High-Altitude Student Platform (HASP) program described evidence of the ocean microbarom as well as other signals of unknown provenance [*Bowman and Lees*, 2015]. A follow-up experiment showed that the ocean microbarom was often detectable in the stratosphere but not at ground-level, either due to low noise at the balloons, an elevated acoustic duct, or both [*Bowman and Lees*, 2017]. In 2017, four microphone-bearing solar balloons were launched concurrently and successfully detected and located a chemical explosion on the ground [*Bowman and Albert*, 2018].

64 So far, none of these experiments have confirmed the detection of other infrasonic sig-
65 nals from sources such as lightning storms.

66 Acoustic emissions from lightning, described as thunder, can produce a broadband
67 range of frequencies. The origin of the audible portion of the thunder (20-20,000 Hz) is
68 widely understood to come mostly from shock waves produced by the rapid heating and
69 expansion of a lightning channel due to current flow [Few *et al.*, 1967]. In addition, nu-
70 merous studies have observed infrasound generated by cloud-to-ground (CG) and intra-
71 cloud (IC) lightning flashes [e.g. Balachandran, 1983; Assink *et al.*, 2008; Farges and
72 Blanc, 2010; Arechiga *et al.*, 2014]. The lightning signal is often a discrete pulse char-
73 acterized by an initial compression followed by a rarefaction with maximum amplitudes
74 in the range of 0.05 to 5 Pa and a spectral peak in the range of 0.2 to 2 Hz [Dessler, 1973;
75 Bohannon *et al.*, 1977; Assink *et al.*, 2008; Campus and Christie, 2010]. Multiple pro-
76 duction mechanisms have been postulated for the infrasonic acoustics detected during
77 lightning storms. This includes rapid intensification of the electric field just prior to the
78 flash, ohmic heating of the air by charge flowing into the channel, and interaction be-
79 tween the positive and negative charge layers in the storm cloud [Dessler, 1973; Bohan-
80 non *et al.*, 1977; Pasko, 2009]. Furthermore, the acoustic wavefield generated by this mech-
81 anism was predicted to be orientated vertically, restricting the horizontal detection range
82 of lightning infrasound [Dessler, 1973; Pasko, 2009]. Validation of the production mech-
83 anism and acoustic wavefield has been confounded by the difficulty in locating the charge
84 layers in the storm cloud, as well as characterizing the structure of the parent lightning
85 flash. Advancements in location algorithms and instruments deployments such as the
86 Lightning Mapping Array (LMA) have produced observations that refute the previously
87 proposed production mechanisms [Arechiga *et al.*, 2014]. Instead, the observations sug-
88 gest that the infrasonic signals from lightning flashes may be produced by electrostatic
89 interaction of charge deposited in the streamer zone of a lightning channel [Arechiga *et al.*,
90 2014]. Lightning infrasound has also been detected at ranges of up to 150 km from the
91 source, contrary to previous predictions of a vertically orientated acoustic wavefield [e.g.
92 Farges and Blanc, 2010].

93 Here we report on detections of lightning infrasound recorded by a high-altitude
94 balloon flying over the Tasman Sea on 17 May 2016. We present evidence for signals recorded
95 from at least two groups of lightning flashes during a 6 hour period. Measurements sug-
96 gest that the detection of lightning infrasound was limited by the distance and atmo-
97 spheric conditions between source and receiver. A few example signals from lightning
98 strokes are isolated and their waveform characteristics are briefly discussed. These ob-
99 servations, the first of their kind reported, suggest that microphones deployed on high-
100 altitude balloons can offer additional insights into the production mechanisms of light-
101 ning infrasound.

102 2 Data

103 An acoustic sensor package was included as a piggyback payload on the NASA Ul-
104 tra Long Duration Balloon (ULDB) flight launched from Wanaka, New Zealand on 16
105 May 2016. The ULDB landed in Peru on 2 July 2016 for a total flight duration of 46 days.
106 The ULDB balloon position and height was recorded using an onboard GPS unit, and
107 records show the full flight included a full circumnavigation of the southern hemisphere
108 [Bowman *et al.*, 2017]. The acoustic sensor package recorded data for the first 20 days
109 of the flight, and the ocean microbarom was recorded throughout as well as other sig-
110 nals of unknown provenance [Bowman and Lees, 2018].

111 The sensor package contained three InfraBSU microphones [Marcillo *et al.*, 2012]:
112 one control and a pair with reversed polarities. The reversed polarity sensor was achieved
113 by placing the mechanical sensor on the opposite port. The reversed polarity microphone

114 pair were combined into a single channel via:

$$M = \frac{M_+ - M_-}{2} \quad (1)$$

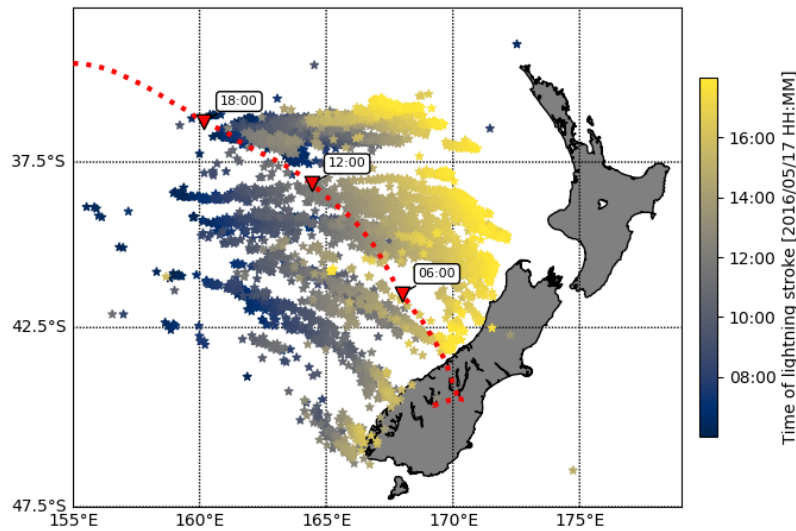
115 where M is the data analyzed in this article, and M_+ and M_- are the data from the mi-
 116 crophones with positive and negative polarities, respectively. The control sensor was a
 117 microphone that was disabled by removing the mechanical filter entirely. This acoustic
 118 sensor trio was designed to robustly distinguish between true pressure fluctuations and
 119 spurious signals, such as electronic interference [Bowman *et al.*, 2017]. Data was recorded
 120 at 200 samples per second at 64x gain using an Omnirecs Datacube digitizer. The mi-
 121 crophones were not calibrated to the pressure and temperature conditions experienced
 122 during the flight, but their primary effect should be to lower the corner period of the sen-
 123 sors [Bowman *et al.*, 2017]. The acoustic waveforms presented here are high-pass filtered
 124 at 0.6 Hz in order to remove high-amplitude signals contributed from the ocean micro-
 125 barom [Bowman and Lees, 2018], atmospheric gravity waves generated by thunder cloud
 126 convection [Blanc *et al.*, 2010], and balloon oscillations [Anderson and Taback, 1991].
 127 (Unfiltered signals recorded by the acoustic package can be seen in Figure S1 in supple-
 128 mentary information.) The microphones and digitizer were each powered by separate Lithium
 129 battery packs, and contained within high density styrofoam shipping boxes for thermal
 130 insulation. Internal temperatures within the digitizer ranged from -26 to 7 °C during the
 131 flight.

132 The lightning stroke detections and location data used in this article were detected
 133 and recorded by the World Wide Lightning Location Network (WWLLN). The WWLLN
 134 is an instrument network capable of locating and timing lightning strokes at long range
 135 (thousands of kilometers) to within <10 km and <10 μ s [Hutchins *et al.*, 2012]. The net-
 136 work uses very-low-frequency radio wave (3-30 kHz) receivers distributed around the globe
 137 to identify the time of group arrival for individual lightning waveforms, or sferics. The
 138 network is capable of detecting both CG and IC discharges, but the latter are typically
 139 underrepresented in detection databases as they produce weaker electromagnetic pulses
 140 [Behnke and McNutt, 2014]. As of 2010, the estimated detection efficiency for the net-
 141 work was \sim 11% for all strokes and >30% for more powerful strokes [Hutchins *et al.*, 2012].
 142 These values may seem low, but the WWLLN was not designed to detect all lightning
 143 strokes but instead to provide a global overview of lightning activity [Dowden *et al.*, 2008].

144 3 Observations

145 The ULDB was launched from Wanaka, New Zealand just before 0000 UTC on 17
 146 May 2016 started flying east as it ascended. Once the craft approached and breached
 147 30 km altitude, it turned to the west and flew out over the Tasman Sea and towards Aus-
 148 tralia (Fig. 1). During this period, the WWLLN detected intense lightning activity from
 149 multiple thunderstorms approaching New Zealand from the west (Fig. 1, Movie S1 in
 150 supporting information). From 0800 to 1400 UTC, 2994 strokes were detected and lo-
 151 cated by the WWLLN across the Tasman Sea, of which 2554 were located within 500
 152 km of the ULDB (Fig. 2a). At approximately 0945 and 1200 UTC the ULDB passed di-
 153 rectly over or near lightning activity which correlates with an increase in acoustic activ-
 154 ity recorded at the ULDB (Fig. 2a, b). Acoustic signals are recorded with peak-to-peak
 155 amplitudes of up to 0.05 Pa and a broadband range of frequencies from 0.6 to 20 Hz (Fig.
 156 2b, c).

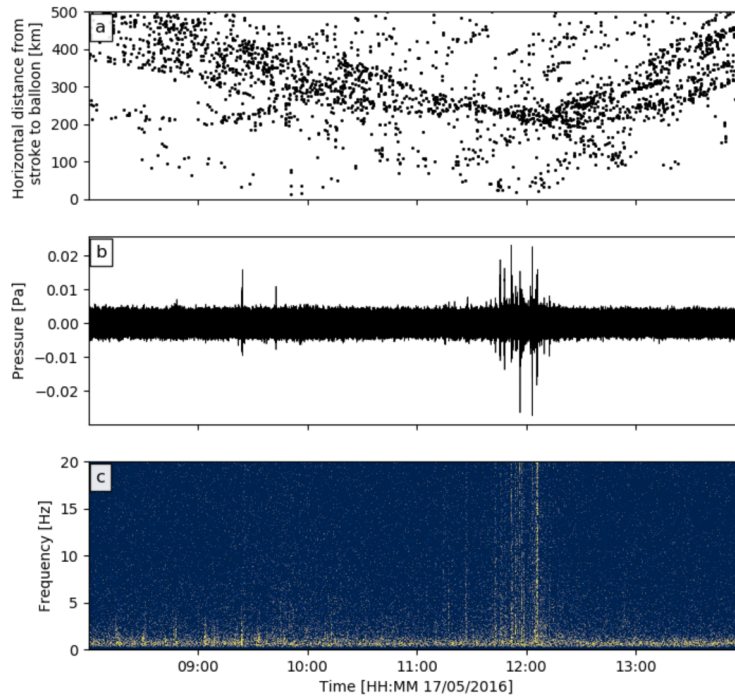
166 As the acoustic sensor package on the ULDB was fundamentally a single element
 167 station, back-azimuths and slowness vectors cannot be calculated to locate sources of de-
 168 tected signals. To estimate detection ranges, ray tracing was used to model infrasonic
 169 propagation paths between lightning and the ULDB. Ray tracing was performed using
 170 classical geometric acoustics techniques and a plane wave assumption, calculated within
 171 the open source GeoAc ray tracing software [Blom and Waxler, 2012]. Rays were launched



157 **Figure 1.** Map of the Tasman Sea with the locations of lightning detected by the WWLLN
 158 from 0600 to 1800 UTC on 17 May 2016, where color represents the progression of time (see col-
 159 orbar). Also plotted is the path of the ULDB balloon after it was launched from Wanaka, New
 160 Zealand (red dotted line), and its location at 0600, 1200 and 1800 UTC on 17 May 2016 (red
 161 triangles). (For an animated version of this figure, see Movie S1 in supporting information.)

172 at intervals of 1° from a point source at a height of 4 km. This source height was based
 173 on previously used heights for modeling lightning infrasound [*Pasko, 2009; Farges and*
 174 *Blanc, 2010*]. Atmospheric profiles were derived from the 12z Global Forecast System (GFS)
 175 analysis model run, located at the latitude/longitude coordinates for the ULDB at 1200
 176 UTC on 17 May 2016 (Fig. S2 in supplementary information). For a source at 4 km height
 177 and a receiver at 32 km height, direct arrivals from the source should only be expected
 178 <110 km horizontal distance in all directions (Fig. 3). At this distance, 34 lightning strokes
 179 were recorded when the ULDB flew near a storm at approximately 1200 UTC (Fig. 4a).
 180 10 strokes were recorded within 100 km during the earlier storm at 0945 UTC (Fig. S3
 181 in supplementary information).

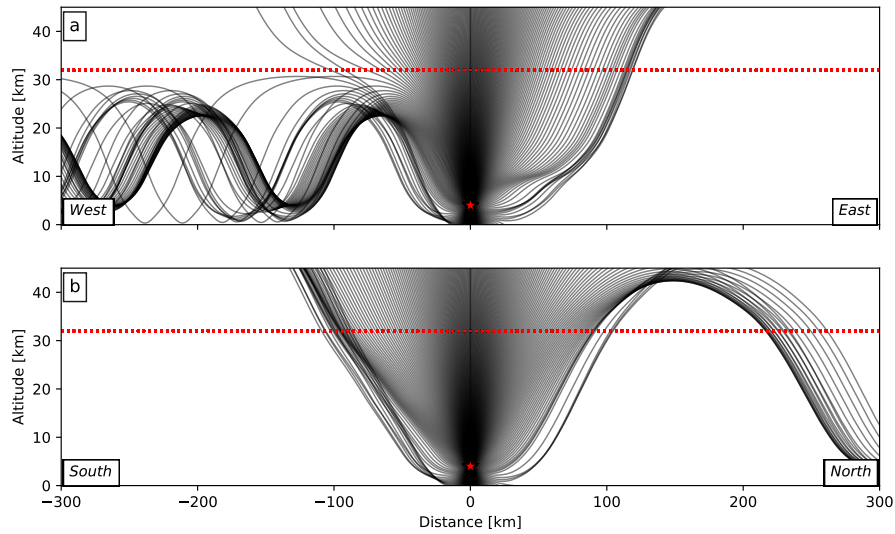
186 To match infrasonic signal peaks and specific lightning strokes, we compute the time
 187 needed for waveforms from each stroke within a limited distance to arrive at the ULDB.
 188 We take a simplified approach and assume that the atmosphere can be approximated
 189 with a bulk acoustic wave speed of 300 ms^{-1} . Furthermore, all acoustic waveforms ar-
 190 riving at the ULDB platform are assumed to be direct arrivals from the source. For the
 191 34 lightning strokes within 100 km of the balloon between 1130 and 1230 UTC, only a
 192 few events appear to match directly with peaks in recorded acoustic amplitudes (Fig.
 193 4b). Three matches occur for signals recorded at the ULDB at 1145, 1148, and 1152 UTC
 194 and are plotted in Fig. 4c, d, and e, with their respective spectrograms (Fig. 4f, g, h).
 195 None of the detected strokes directly match with infrasonic signals during the earlier storm
 196 at 0945 UTC (Fig. S3 in supplementary information). The apparent low matching rate
 197 between lightning strokes and acoustic signals indicates that recording infrasonic thun-
 198 der signals at stratospheric altitudes was dependent on source-receiver distance, light-
 199 ning stroke energy, and the atmospheric conditions.



162 **Figure 2.** (a) Horizontal distances for each stroke within 500 km of the balloon’s location at
 163 the time of the stroke, from 0800 to 1400 UTC on 17 May 2016. (b) High-pass filtered (0.6 Hz)
 164 infrasound over the same time period as recorded at the ULDB. (c) Frequency spectrogram of the
 165 waveform plotted in (b).

206 4 Discussion and Conclusions

207 Here we have presented evidence that lightning infrasound was observable by acous-
 208 tic instruments suspended at stratospheric altitudes by free-flying balloons. Several wave-
 209 forms were matched with detected lightning strokes through a simple time delay approach
 210 (Fig. 4c, d, e). Here, we have assumed that the signals represent direct arrivals between
 211 the source and receiver. To test this assumption, we searched for eigenray solutions us-
 212 ing the GeoAc software package. Direct arrivals for all three waveforms are found for the
 213 distances and azimuths to their associated lightning strokes (Fig. S4 in supplementary
 214 info). However, the arrival times using the eigenray paths described earlier to calculate
 215 their arrival times does not readily match with the recorded arrival times of signals (Fig.
 216 S4 in supplementary information). It is worth noting that there were a number of sig-
 217 nals recorded that do not readily match with any lightning strokes detected by the WWLLN,
 218 and *vice versa* (Fig. 4a, b). If the estimated detection rates of the WWLLN are correct
 219 [11-30%; *Hutchins et al.*, 2012], then there may have been as many as 100-300 lightning
 220 strokes within 100 km of the ULDB. This number of high-amplitude signals was not recorded
 221 at the ULDB (Fig. 4b), therefore the total number signals recorded at the ULDB un-
 222 derrepresents the true total of lightning strokes that occurred within range of the bal-
 223 loon. This is similar to detection rates of lightning infrasound by ground-based instru-
 224 ments [e.g. *Farges and Blanc*, 2010]. This may be attributed to very low signal-to-noise
 225 ratios, especially for smaller lightning strokes or those located further from the instru-

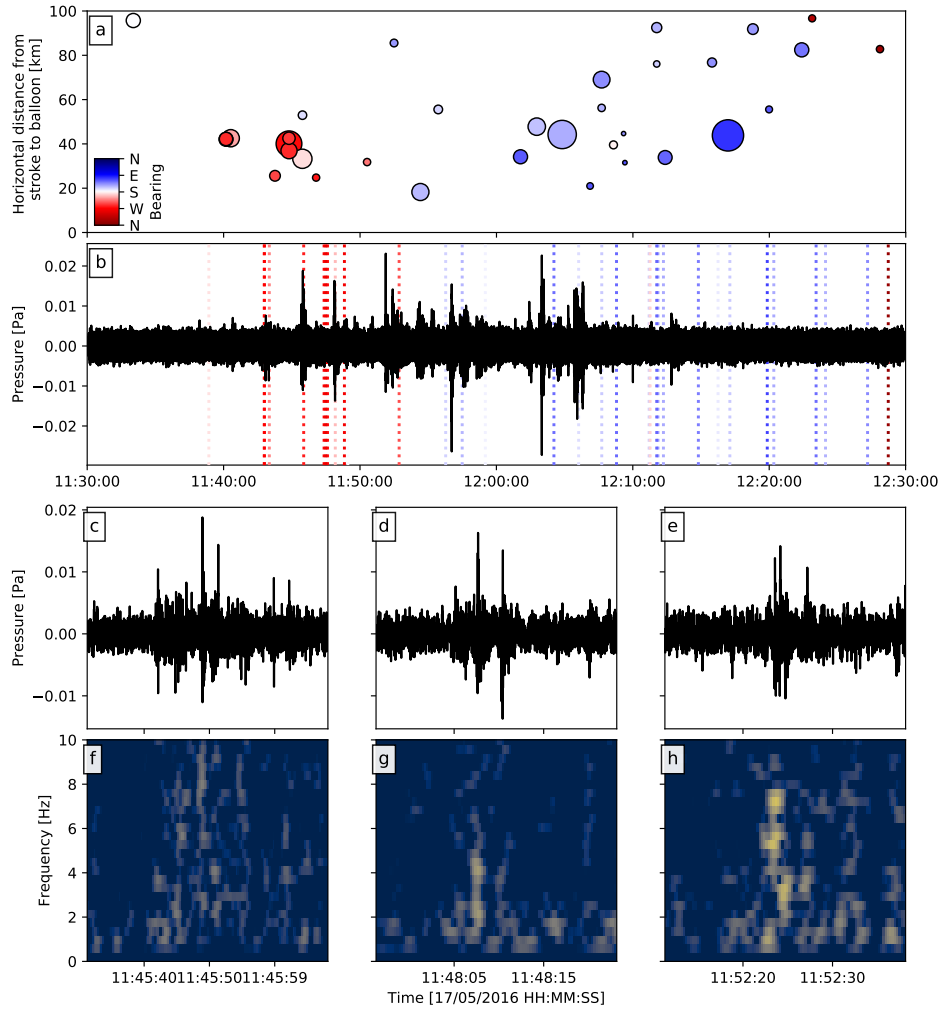


182 **Figure 3.** Ray-tracing propagation for an acoustic source (red star) at 4 km height along
 183 East-West (a) and North-South (b) profiles using realistic atmospheric conditions derived from
 184 the Global Forecast System. The red dotted line at 32 km indicates the approximate height of
 185 the ULDB balloon on 17 May 2016.

226 ment than the rest of the thunderstorm cloud [Farges and Blanc, 2010]. Complex atmo-
 227 spheric conditions likely refract the generated acoustic waves away from the receiver [Jones
 228 and Bedard, 2015]. Additionally, it is possible that not all lightning strokes generate mea-
 229 surable infrasound.

230 For the ray propagation and eigenray modeling we have assumed a point source
 231 for the lightning infrasound at 4 km altitude. Infrasound sources from lightning have been
 232 mapped as high as 12 km altitude in the thundercloud [Anderson et al., 2014; Arechiga
 233 et al., 2014]. Furthermore, the mapped current flow within lightning strokes suggests the
 234 source geometry can resemble complex, dendritic structures [Anderson et al., 2014]. To
 235 test whether our source shape and height assumption was viable, we have repeated the
 236 ray tracing modeling but with a source at 32 km height instead, the height of the ULDB.
 237 In reverse, this can be seen as all possible locations for sources whose acoustic waveforms
 238 will be recorded at the receiver. Results suggest that a receiver at 32 km should record
 239 signals from heights up to 12 km and take-off angles between 45-90° within a 100 km
 240 horizontal range in all directions (Fig. S5 in supplementary information). Therefore, the
 241 source configuration used for the raypath modeling was reasonable for the lightning in-
 242 frasound described here.

243 A key assumption here was that the atmospheric conditions during the generation
 244 of the lightning infrasound was relatively simple and stratified. Thunderstorms require
 245 unstable air to form and persist, and often include significant vertical wind shear due
 246 to thermal plumes. The wind and temperature profiles used for ray path modeling here
 247 was likely a highly simplified representation of reality. Ray tracing from sources directly
 248 below thermal plumes such as those found within thunderstorm clouds have shown how
 249 they may act as vertical waveguides and thus can greatly distort the acoustic wavefield
 250 [Jones and Bedard, 2015]. Ray paths are found to either converge or diverge at the top
 251 of the thermal plume dependent on the height and width of the plume [Jones and Be-



200 **Figure 4.** (a) Horizontal distances for each stroke within 100 km of the balloon from 1130 to
 201 1230 UTC on 17 May 2016. each stroke is sized by the stroke energy, and colored by the bear-
 202 ing from the balloon to the stroke location. (b) High-pass (0.6 Hz) filtered acoustic waveform as
 203 recorded at the ULDB over the same time period. Vertical dotted lines indicate calculated time
 204 of arrivals for strokes in panel (a), colored by the bearing. (c,d,e) Example acoustic signals from
 205 lightning strokes and their respective spectrograms (f,g,h).

252 *ard*, 2015]. Therefore the ray-paths and eigenrays presented here are likely an oversim-
 253 plified representation of the true paths taken by the acoustic waves before their detec-
 254 tion at the ULDB. Future studies of acoustic wavefields generated by lightning infrasound
 255 must take into account the complex refraction patterns induced by vertical columns of
 256 wind shear within thunderclouds.

257 The measured waveforms do not display the compression-rarefaction-compression
 258 shape that had been modeled as generated by a rapid intensification and discharge of
 259 the electric field in the thundercloud [Pasko, 2009]. It must be noted that the waveforms
 260 modeled by Pasko [2009] are of 0.1-1 Hz frequency and the acoustic wavefield was strongly
 261 oriented in the vertical direction [Dessler, 1973]. The detection of lightning infrasound
 262 from strokes at more than several tens of kilometers from the ULDB does not support

263 the theory that acoustic wavefields generated by lightning infrasound are strictly orien-
 264 tated vertically. Instead, the waveforms here share amplitude, frequency and range de-
 265 tection characteristics with previously recorded lightning infrasound signals which were
 266 attributed to charge deposition in the lightning channels [e.g. *Farges and Blanc*, 2010;
 267 *Arechiga et al.*, 2014].

268 The observations here of lightning infrasound recorded by a high-altitude balloon
 269 over the Tasman Sea in May 2016 are fortuitous. Yet, they also demonstrate the poten-
 270 tial for using these platforms to fill gaps on our monitoring capabilities of infrasound gen-
 271 erating sources such as lightning storms. Future high-altitude balloon deployments com-
 272 bined with ground-based instruments will be required to image the full infrasonic wave-
 273 field generated by thunderstorms which, in turn, will provide a better understanding of
 274 the source mechanics of lightning infrasound. Furthermore, acoustic source localization
 275 is possible by the simultaneous deployment of multiple balloons to form a high-altitude,
 276 free-flying acoustic network [e.g. *Bowman and Albert*, 2018].

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