

1 **How, where, and when do radial faults grow near salt diapirs?**

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12 **ABSTRACT**

13 We examine three-dimensional seismic data from the Santos Basin, offshore Brazil, to
14 determine how, where, and when radial faults grow near a salt diapir. We show roof stretching
15 alone cannot account for the large heights and lengths of the kilometer-scale radial faults,
16 suggesting stock widening ('stem push'), a mechanism implied in numerical models but not yet
17 recognized in natural examples, played a pivotal role in fault formation. We suggest that, when a
18 diapir is covered by a roof, radial faults form due to roof stretching, extending no further than the
19 limit of the drape folding. The roof may then be shouldered aside and the faults buried along the
20 stock flanks, exposing these strata to stem push-related stresses that may then re-activate
21 preexisting, or form new, radial faults. We suggest the causal mechanism for radial fault formation
22 will likely change as roof thickness varies during diapirism, with this reflecting the ratio between
23 sedimentation rate and salt volumetric flux.

24 **INTRODUCTION**

25 Sub-circular salt diapirs or ‘stocks’ are ubiquitous in salt-bearing sedimentary basins, and
26 are typically associated with complex fault networks in surrounding country rock. The most
27 common fault networks comprise ‘radial faults’ (i.e., normal faults that extend radially from a
28 stock into flanking strata). Radial faults may control the migration of crustal fluids (e.g., Davison
29 et al., 2000a), may compartmentalize hydrocarbon reservoirs (e.g., Carruthers et al., 2013), and
30 may provide a relatively high-fidelity record of the evolving near-salt stress conditions associated
31 with salt diapirism (cf. Quintà et al., 2012; Nikolinakou et al., 2014).

32 Despite being widespread, and geologically and economically important, the origin of
33 radial faults remains unclear. Radial faults in unpierced roofs above rising stocks are undoubtedly
34 related to outer-arc extension during active rise (‘roof stretching’; Fig. 1A). Roof stretching–
35 related radial faults may nucleate anywhere in, but not necessarily extend fully across, the arched
36 overburden (Withjack and Scheiner, 1982). As a stock pierces its overburden, roof radial faults
37 may be eroded or shouldered aside, and buried along the stock flanks (e.g., Carruthers et al., 2013).
38 Stretching and shouldering aside of the roof may occur during ‘drape folding’ during passive
39 diapirism (Giles and Rowan, 2012) or regional extension, or alternatively during active rise driven
40 by regional shortening (Dooley et al., 2009; Jackson and Hudec, 2017).

41 Radial faults may also form due to a widening stock pushing outward against its flanking
42 strata (e.g., Bishop, 1978; Nikolinakou et al., 2014, their figure 7), herein termed ‘stem push’ (Fig.
43 1B). Stem push–related radial faults form at the salt-sediment interface where circumferential
44 stretching is greatest and the horizontal stresses are anisotropic. Although numerical models
45 suggest stem push is a plausible mechanism to form radial faults (e.g., Nikolinakou et al., 2014),

46 observations from a natural salt stock have never critically tested this prediction, nor characterized
47 and quantified the associated strain.

48 Drape folding above a passively rising diapir is typically recorded by synkinematic
49 Composite Halokinetic Sequences (CHS; Giles and Rowan, 2012), with two end-members
50 recognized: tapered and tabular. Tapered CHS form when the salt is buried by a relatively thick
51 roof, as sedimentation rate outpaces the volumetric flux of salt; in this case, broad, kilometer-scale
52 drape folds form. Tabular CHS form when the salt is at, or very near, the surface and covered by
53 only a relatively thin roof, and the volumetric flux of salt is greater or equal to the sedimentation
54 rate (e.g., Giles and Rowan, 2012; Jackson and Hudec, 2017); in this case, narrow drape folds (<
55 200 m) form next to the salt-sediment interface. As the roof is pierced, strata containing drape
56 fold-related radial faults are either eroded or buried along with the contained faults along the flanks
57 (Fig. 1A). In tapered CHS, radial faults are expected to extend greater distances (<1000 m) from
58 the salt due to broader folding compared to tabular CHS (<200 m). It follows that radial faults
59 extending more than a few hundred meters laterally in tabular CHS must have formed due to stem
60 push rather than roof stretching alone, although this has never been tested.

61 Here, we test these hypotheses by identifying CHS and applying quantitative fault analysis
62 to infer where faults nucleated and how they grew around a salt stock imaged in three-dimensional
63 (3-D) seismic reflection data from the Santos Basin, offshore Brazil (Fig. 1C). Using this approach,
64 we: (1) link the genetic mechanism of radial fault formation to salt diapirism, and (2) for the first
65 time using a natural example, test the validity of the stem push model, using exceptionally well-
66 imaged radial faults flanking and overlying a salt stock. These data not only allow us to map radial
67 fault-diapir relationships in 3-D and constrain their kinematics, but also investigate when, during
68 diapirism, roof stretching and stem push may occur.

69 **DATASET AND METHODS**

70 We used 225 km² of a 850 km² Kirchhoff pre-stack time-migrated (PSTM), zero-phase
71 processed, 3-D seismic data set. Inline and crossline spacing are 14 m and 25 m, respectively. A
72 frequency of ~15–40 Hz and assumed average velocity of ~2000 m/s (after Jackson et al., 2014)
73 yield an estimated vertical resolution of ~12 m at shallow depths, decreasing to ~35 m toward the
74 base of supra-salt minibasins (see Appendix DR1 in the GSA Data Repository¹ for details). All
75 seismic data are displayed in milliseconds two-way time (ms TWT), but measurements are
76 converted from time to depth using an interval velocity of 2000 m/s. We first mapped three seismic
77 horizons (H1–H3) to constrain salt body geometry, and the 3-D distribution of throw on, and
78 kinematics of, individual faults (Appendix DR2). Quantitative fault analysis was not undertaken
79 for H1 because throw was at the limit of seismic resolution (i.e., <25 m; Appendix DR3). We then
80 identified nine Late Cretaceous–Tertiary stratigraphic units adjacent to the stock, assigning them
81 to the two end-member CHS styles of Giles and Rowan (2012) based on the width of folding and
82 thinning and the geometry (convergent or parallel) of the bounding unconformities. These CHS
83 allowed us to interpret periods when the rising diapir was buried by a thick (tapered CHS) or thin
84 (tabular CHS) roof (Fig. 2) (see Giles and Rowan [2012] for recognition criteria). We then grouped
85 the units into three packages based on CHS style, and whether the stock had pierced strata at the
86 level of observation. Package A consists of tabular CHS, whereas packages B and C contain
87 tapered CHS. Packages A and B have been pierced by the salt, whereas C has not. H1 lies in
88 package A, H2 at the boundary between packages A and B, and H3 in package C.

89

90 **GEOLOGICAL SETTING**

91 The Santos Basin formed during Early Cretaceous rifting and initial opening of the South
92 Atlantic, during which time a thick Aptian salt layer was deposited (Ariri Formation) (Mohriak et
93 al., 2008; Contreras et al., 2010). Subsequent deposition of Albian (carbonate-dominated) and
94 Cenomanian-Holocene (siliciclastic-dominated) rocks, in addition to thin-skinned gravity-driven
95 extension, drove seaward salt flow and diapir growth (Demercian et al., 1993; Modica and Brush,
96 2004; Davison et al., 2012). We focus on a salt stock located (Fig. 1C) in the proximal, extensional
97 domain (after Davison et al., 2012), in an area unlikely to have undergone Albian shortening. Like
98 many salt structures in this area, the stock initiated as a reactive diapir, before undergoing passive
99 and active rise driven by sediment loading (Jackson et al., 2015). Here, we focus only on the latter
100 stages of diapirism once the stock had developed, where CHS and radial faults formed.

101

102 **SALT STOCK AND OVERBURDEN GEOMETRY, AND DIAPIRISM**

103 The salt stock is expressed in seismic data as a package of chaotic, low-amplitude
104 reflections. In cross section, the stock is ~4 km tall and has a ‘finger’ geometry, consisting of a
105 <2.3-km-wide smooth head and stem, and a <6-km-wide pedestal (Fig. 2). In plan view, the stock
106 is sub-circular at shallow depths (~2000 ms TWT) and oblate at greater depths (~4000 ms TWT),
107 with its long axis trending northeast. The presence of tabular CHS (with narrow drape folds <200
108 m from the salt) at deeper levels indicates that, following diapir initiation, the stock entered a
109 protracted phase of passive diapirism when the volumetric flux of salt equaled or exceeded the
110 background sedimentation rate (package A). Tapered CHS (with broad drape folds <1000 m from
111 the salt) dominate at shallower levels, suggesting that sedimentation rate outpaced the volumetric

112 flux of salt (packages B and C). This could reflect an increase in the regional sedimentation rate,
113 or a decreased volumetric flux of salt as the source layer thinned and ultimately welded.

114 Shortening has been documented regionally in the Santos Basin (e.g., Demercian et al.,
115 1993; Modica and Brush, 2004; Contreras et al., 2010). However, based on the stock's location in
116 the extensional domain (after Davison et al., 2012), the sub-circular map-view geometry (cf.
117 Jackson and Hudec, 2017), a lack of thrusts in the roof or flanking minibasins (cf. Davison et al.,
118 2000b; Dooley et al., 2009), and fault patterns unlike those expected during compression (cf.
119 Withjack and Scheiner, 1982), we interpret diapir growth was *not* driven by shortening. If
120 shortening has occurred, the strains associated with this must be minimal (e.g., Davison et al.,
121 2000a, 2000b).

122

123 **RADIAL FAULTS**

124 **Geometry and Distribution**

125 Radial faults are broadly linear in map view at all stratigraphic intervals (H1–H3), although
126 they vary in their distribution, density, and length (Fig. 3). They occur over an ~2.5 km depth range
127 (~1–3.5 km) within tapered and tabular CHS, although they tend to cluster around the stock head
128 in tapered CHS (H2-H3). It is possible that radial faults exist but are not imaged at greater depths
129 (>3.5 km). Faults are planar, 400–1400 m tall, have height-to-length aspect ratios of <2 (Appendix
130 DR4), dip 50–60°, and have throws of <80 m. Faults occur in vertically stacked tiers. Faults in
131 each tier have similar geometric characteristics; e.g., heights, lengths, and densities. Tier
132 boundaries are undeformed, or at least deformation is sub-seismic. Tall radial faults, which are
133 best-developed at shallower levels around the head of the stock in packages B and C, may cross-
134 cut several tier boundaries (Fig. 4; Appendix DR5).

135 **Throw Distribution**

136 We study the distribution of radial fault throw to determine where these structures
137 nucleated with respect to the stock, which may reveal their formation mechanism. Throw maxima
138 for faults offsetting H2 and H3 occur immediately at or some distance from the salt-sediment
139 interface (<3 km from the salt; white squares in Fig. 3; Appendix DR6). Faults typically have ‘C-
140 type’ throw-depth profiles (*sensu* Muraoka and Kamata, 1983), with a throw maximum near their
141 centers and very low gradients toward their tips (<0.1) (Fig. 4A; Appendix DR7). Some faults may
142 have several throw maxima separated by throw minima, and may offset presumably older,
143 neighboring faults (Fig. 4B). Faults are not associated with growth strata (expansion indices of ~1;
144 Fig. 4), suggesting they were blind.

145 **Kinematics and Origin**

146 Based on their geometry, stratigraphic occurrence in tapered and tabular CHS, and throw
147 distribution, we propose the radial faults have two origins. Radial faults developed in the roof and
148 are contained in tapered CHS, with throw maxima (i.e., nucleation points; Muraoka and Kamata,
149 1983; Baudon and Cartwright, 2008) located both above and outboard of the stock, and do not
150 intersect the salt, formed only due to roof stretching (H3 in Fig. 3; package C in Fig. 2).
151 Predominantly NNW-SSE–striking faults, whose throw maxima occur outboard of the drape
152 folding limit in H3 and do not encounter the salt, may reflect stresses related to diapir growth to
153 the NNW and SSE, or movement of deeper-lying salt structures (Fig. 1C).

154 In contrast, radial faults in tabular CHS, and which are in contact with and extend several
155 kilometers from the salt (that is, well beyond the limit of drape folding), formed due to stem push
156 (H1 in Fig. 3; package A in Fig. 2). These faults nucleated at the salt-sediment interface where the
157 circumferential extension is greatest (e.g., Nikolinakou et al., 2014; Jackson and Hudec, 2017). As

158 the stock was at or near the surface, with only a relatively thin roof during deposition of tabular
159 CHS, radial faults associated with roof stretching would be limited to the relatively narrow extent
160 of drape folding (<200 m), immediately adjacent to the stock. Roof stretching, therefore, cannot
161 be responsible for the formation of radial faults that extend several kilometers away from the salt,
162 now deeply buried in the stock flanks. Given that the majority of deep radial faults are not
163 physically connected to shallow radial faults associated with roof stretching (Appendix DR5), the
164 deep faults cannot be attributed to downward propagation of the shallower-level structures; they
165 must therefore reflect a mechanism other than drape folding. The vertical extent of some radial
166 faults was several kilometers (Fig. 4), again suggesting it is unlikely they formed due to roof
167 stretching alone, as passive diapirs cannot arch kilometer-thick roofs (e.g., Davison et al., 2000b;
168 Jackson and Hudec, 2017). Given the lack of evidence for regional shortening, a mechanism that
169 could have lifted kilometer-thick diapir roofs and generated kilometer-tall faults, we propose such
170 tall and laterally extensive faults grew due to stem push during passive diapirism.

171 Having considered radial faults in tabular CHS, we now explore which mechanism likely
172 produced radial faults in package B (pierced tapered CHS). Radial faults in package B have their
173 throw maxima either outboard of the salt or at the salt-sediment interface (H2 in Fig. 3). The former
174 suggests roof stretching must have occurred over a broad region with discontinuous faulting;
175 however, the latter could feasibly be explained by either: (1) stem push, or (2) roof stretching, and
176 subsequent diapiric piercement of the overburden. In the first case, radial faults nucleate where
177 circumferential extension is greatest due to stem push at the salt-sediment interface (Fig. 1B). In
178 the second case, piercement of the overburden removes sections of the roof and portions of radial
179 faults formed by roof stretching, thus truncating the original throw distribution. Throw maxima
180 could therefore be only coincidentally located at the salt-sediment interface. Because the radial

181 faults were blind and were not associated with growth strata, we are unable to identify whether
182 stem push re-activated preexisting roof stretching faults as the strata became buried (e.g., package
183 B). Irrespective of their origin, radial faults grew, dip-linked, and offset preexisting radial faults
184 beside the stock (Fig. 4B) (cf. Muraoka and Kamata, 1983; Baudon and Cartwright, 2008).

185

186 **DISCUSSION AND IMPLICATIONS**

187 By undertaking detailed mapping of 3-D seismic reflection data, we are able to not only
188 better determine the full, 3-D geometry of *in situ* radial fault networks, but also constrain their
189 kinematics. Based on our observations from the Santos Basin, we offer a genetic model that may
190 be broadly applicable to other diapirs. We propose that, as a salt stock grows and roof thickness
191 varies with changes in the volumetric flux of salt and/or sedimentation rate, it is likely that the
192 mechanism responsible for forming radial faults will vary. Such changes in the relative balance of
193 salt flux and sedimentation rate may, for example, reflect progressive welding of supra-salt
194 minibasins, or changes in regional sedimentation rate.

195 Once passive diapirism occurs and a stock starts to grow, the volumetric flux of salt may
196 outpace the background sedimentation rate, meaning the stock will be at or near the depositional
197 surface, covered only by a relatively thin roof (i.e., tabular CHS). As this thin roof is arched and
198 is shouldered aside by the rising diapir, roof stretching–related radial faults will be buried adjacent
199 to (<200 m) the salt-sediment interface. As the source layer thins and the volumetric flux of salt
200 decreases, the stock may be buried by a relatively thick roof (i.e., tapered CHS). Subsequent rise
201 of the diapir generates stretching-related radial faults in the aggrading overburden, over a relatively
202 broad area (<1000 m). Shouldering-aside and burial of the roof along the flanks (regardless of the
203 CHS type) may expose these strata to stem push–related stresses, re-activating preexisting, or

204 forming new, radial faults. Faults in the vicinity of the salt may continue to grow throughout
205 diapirism, becoming taller and propagating laterally. Stem push–related re-activation of old faults
206 and the formation of new faults will likely be concentrated toward the upper parts of stocks where
207 the greatest stress perturbations occur (e.g., Nikolinakou et al., 2014, their figures 8 and 9). Finally,
208 as the salt supply is exhausted and minibasins weld, sedimentation rate may outpace the volumetric
209 flux of salt, causing stock burial (cf. Giles and Rowan, 2012; Jackson and Hudec, 2017). Once
210 diapirism ceases, no further radial faults form unless latter extension or shortening occurs.

211 As the genetic mechanism for forming radial faults likely changes during diapirism, the
212 geometry and kinematics of those faults will likely change, especially where they have interacted
213 to create complex fault geometries. This could prove problematic when inverting fault network
214 geometry for paleostress conditions (cf. Quintà et al., 2012; Carruthers et al., 2013), leading to
215 questionable interpretations of salt diapir–related stresses, and the mode and distribution of
216 fractures around salt stocks. In addition, we highlight the structural variability and potential
217 reservoir compartmentalization that may occur around salt stocks, providing insights into areas
218 where radial faults are not exposed or are poorly imaged (e.g., Jones and Davison, 2014).

219 Finally, we note the Santos Basin radial faults are shorter (<3 km versus <6 km) than those
220 suggested by the strain field in the numerical models of Nikolinakou et al. (2014, see their figure
221 7). These differences may reflect variations in the country rock rheology and salt geometry through
222 time, and, in particular, the diapir radius which dictates fault length.

223

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296 316.

297 **FIGURE CAPTIONS**

298 Figure 1. Radial fault formation via roof stretching (A) and stem push (B) with idealized fault
299 throw-length plots. Radial fault Y throw increases toward the salt. Radial fault X does not
300 encounter the salt. Red and blue units are tapered and tabular Composite Halokinetic Sequences
301 (CHS; Giles and Rowan, 2012), respectively. C: Geographic context and variance (i.e., trace-to-
302 trace variability in acoustic impedance) slice at 1500 ms two-way time (TWT).

303

304 Figure 2. Seismic section showing the salt stock and stratigraphic position of horizons H1–H3 and
305 packages A–C. Interpreted tapered Composite Halokinetic Sequences (CHS) in red, and tabular
306 CHS in blue. CHS may exhibit different degrees of upturn next to the salt, forming cusps (inset).
307 For location, see Figure 1C.

308

309 Figure 3. Variance map for horizons H1–H3, delineating radial faults and the salt. Throw maxima
310 (white squares) for individual radial faults, throw-length plots for radial faults 1 and 2 (right), and
311 the location of Figure 4 are also shown. For location, see Figure 1C. Throw maxima are absent for
312 H1, as measured throw is at the limit of seismic resolution.

313

314 Figure 4. Seismic sections showing radial faults 3 (A) and 4 (B) with throw-depth profiles and
315 expansion indices. Solid line is throw-depth, dashed line is expansion index. White circles are
316 vertical fault tips, white squares are throw maxima. Radial fault 3 has a simple throw-depth profile
317 with a single throw maximum. Radial fault 4 offsets older faults and has two throw maxima,
318 indicative of dip linkage. See Figure 3 for location.

FIGURE 1

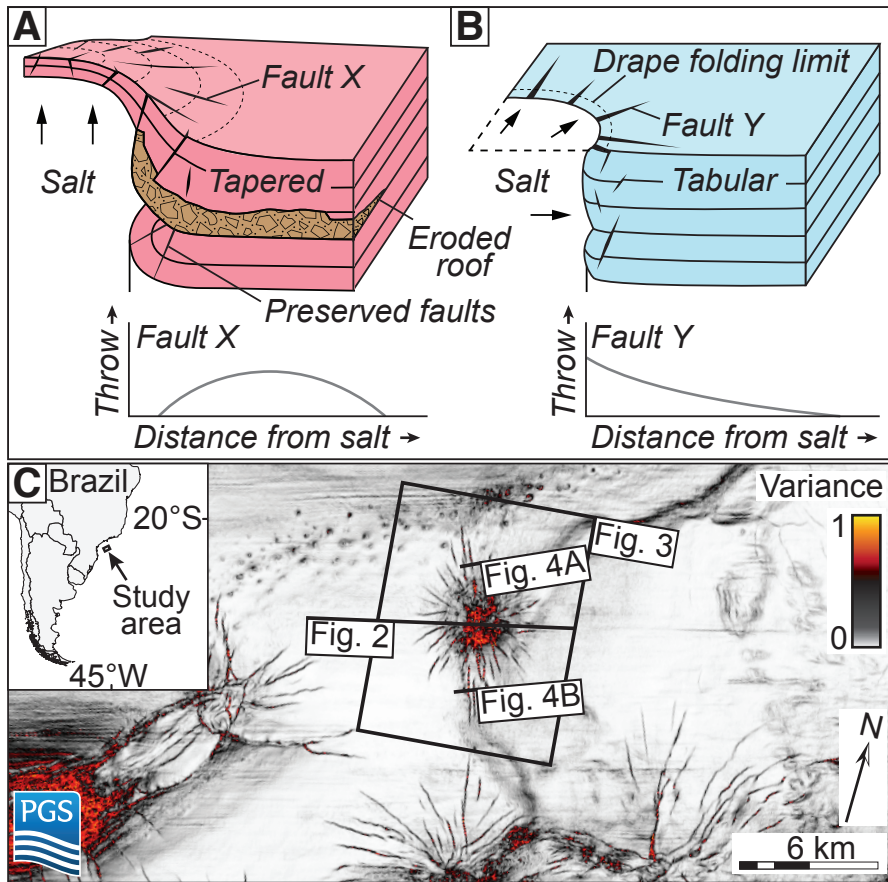


FIGURE 2

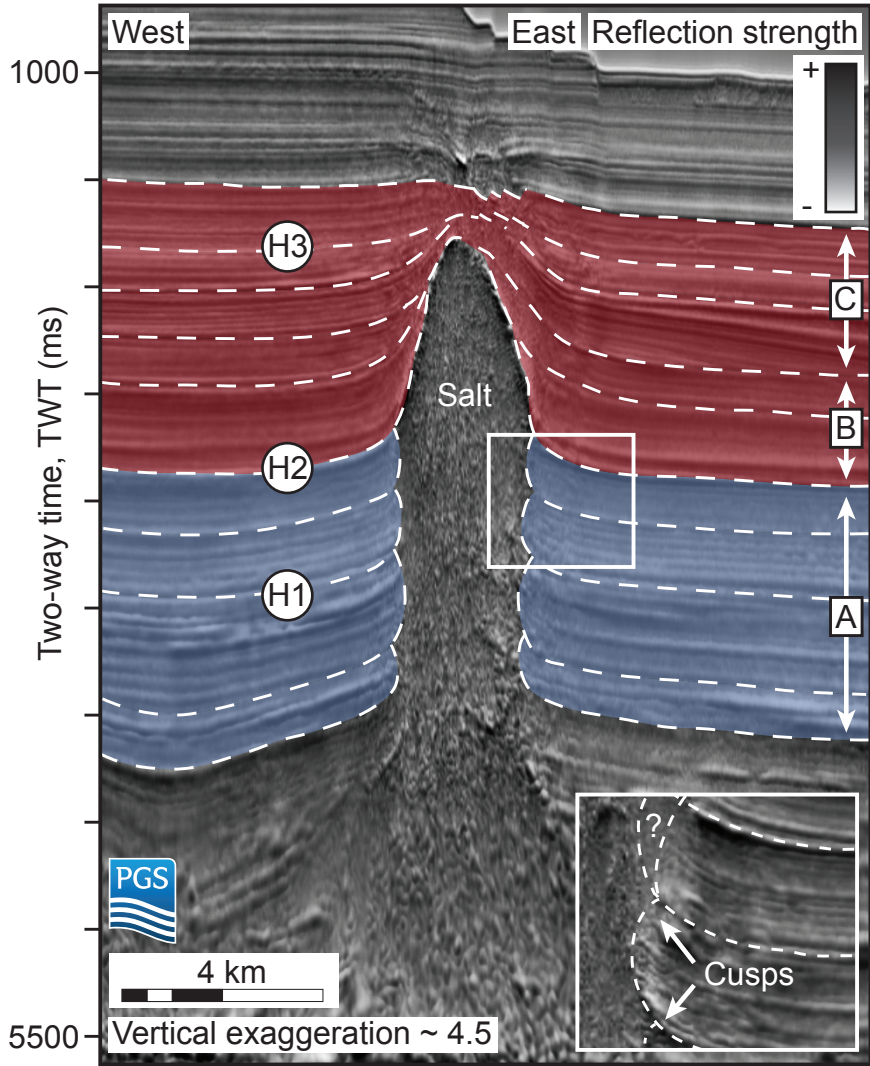


FIGURE 3

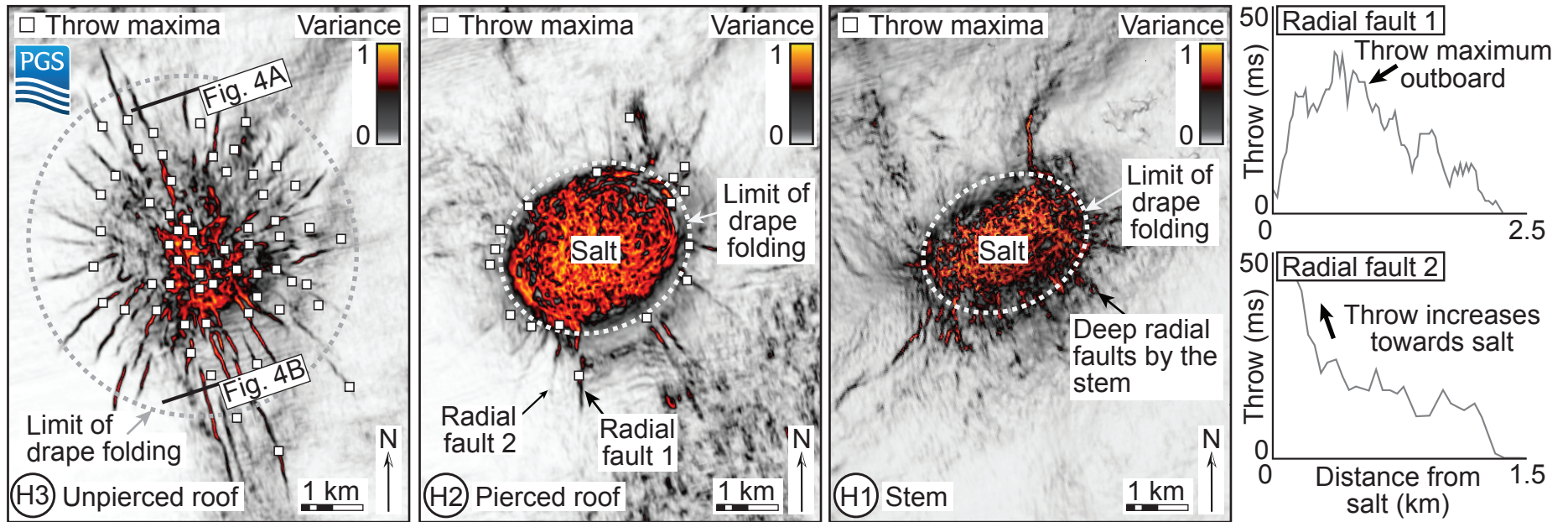
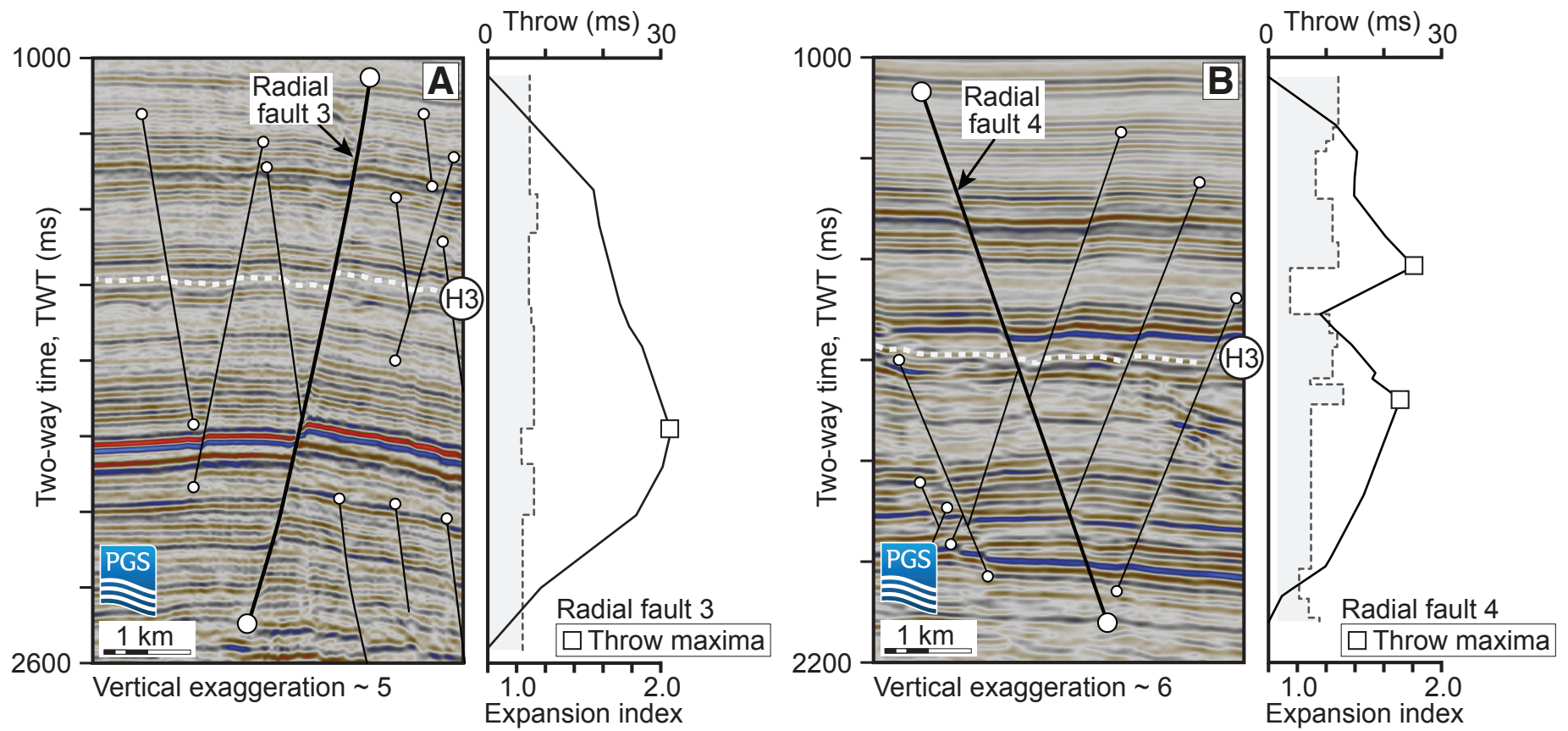
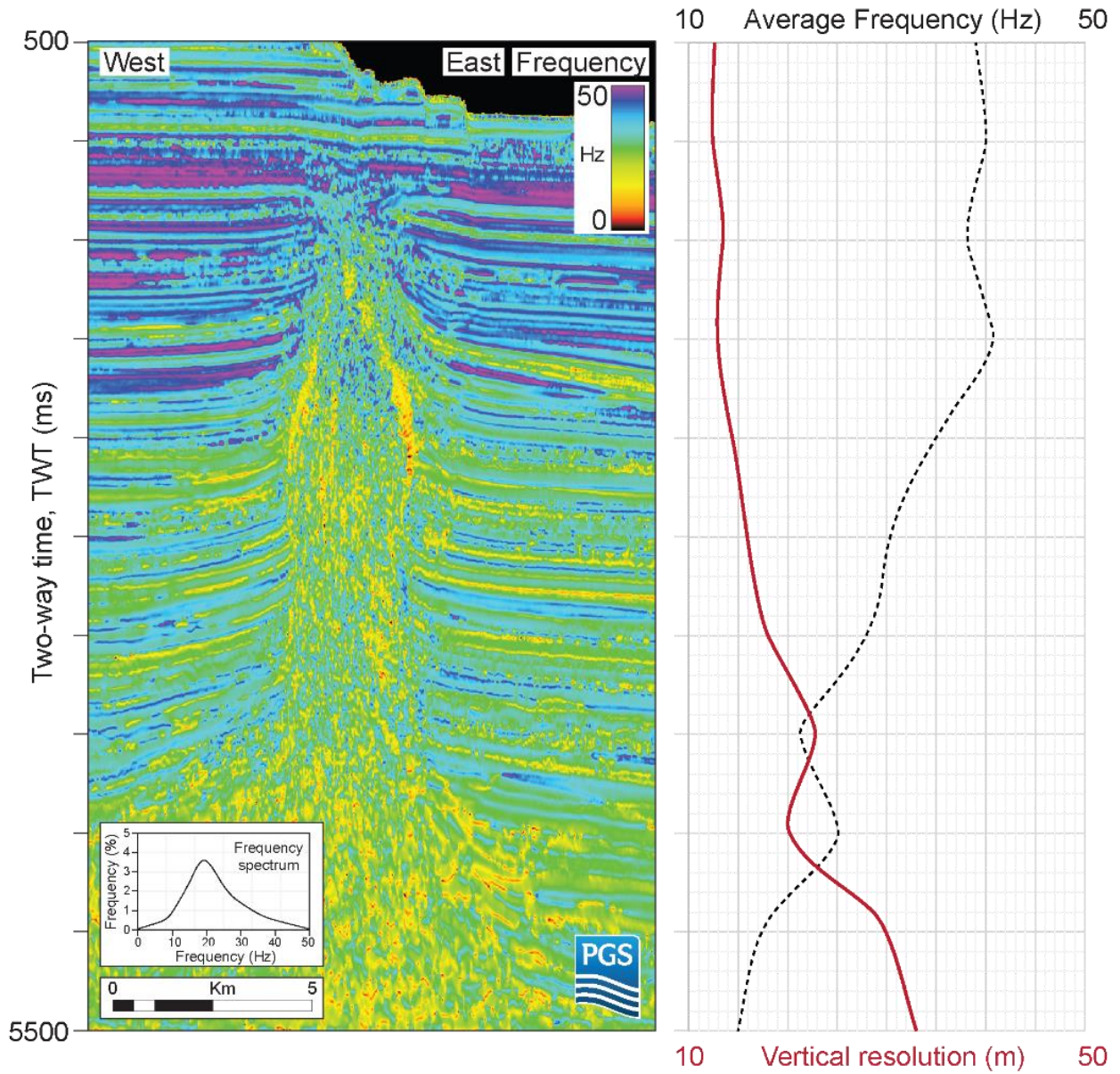


FIGURE 4



APPENDIX 1. AVERAGE VERTICAL SEISMIC RESOLUTION

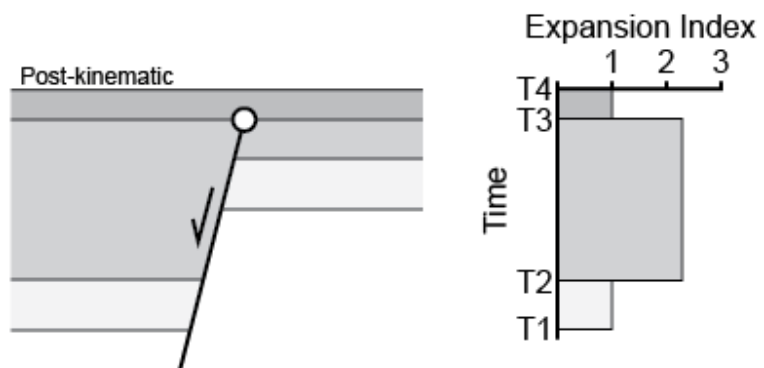
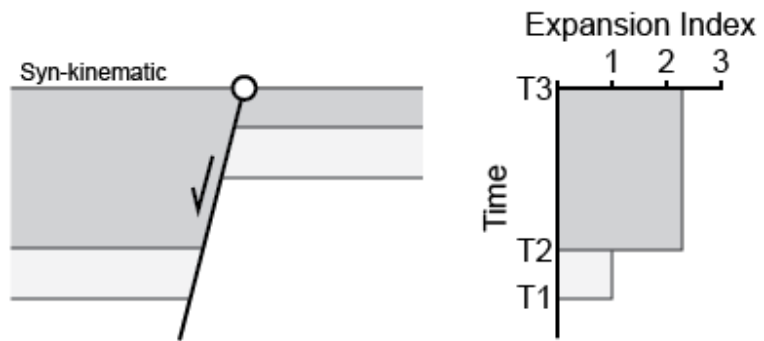
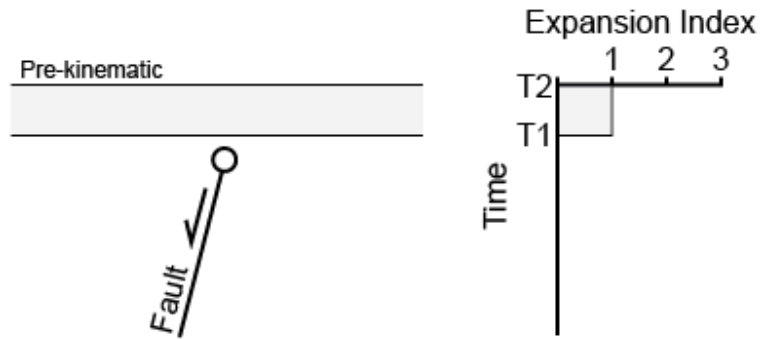
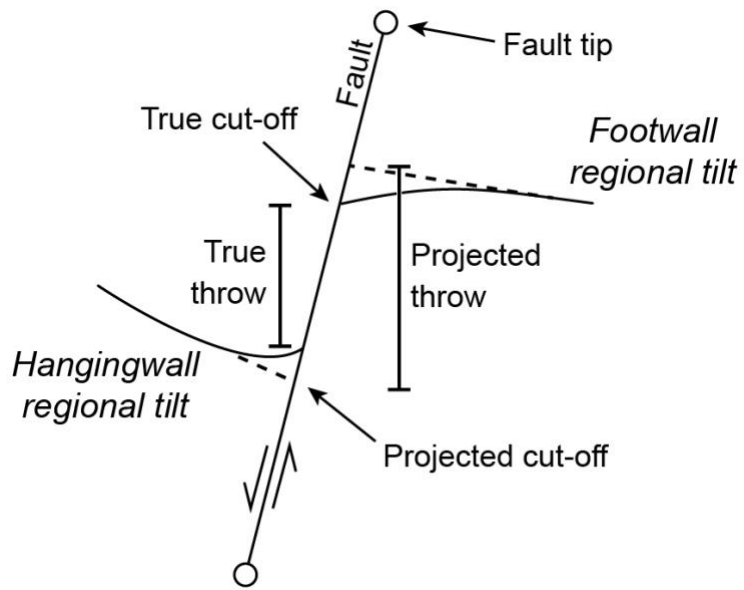
Average vertical seismic resolution (red solid line) with depth using a velocity of ~ 2 km/s (after Jackson et al., 2014) and the frequency (black dashed line). The average vertical seismic resolution was calculated using the frequency and velocity. The instantaneous frequency is shown (left). Vertical exaggeration ~ 4.5.



APPENDIX 2. QUANTITATIVE THROW ANALYSIS METHOD

Fault throw was measured perpendicular to radial fault strike every c. 50 – 100 m along the length of individual radial faults using horizon cut-offs (e.g. Muraoka and Kamata, 1983; Baudon and Cartwright, 2008). Cut-offs were defined using an extrapolated line that follows the regional trend of the chosen horizon prior to folding (Wilson et al., 2013), removing the effect of fault-parallel folding (Walsh et al., 1996). Therefore, total strain across the fault is accommodated, whether accommodated by ductile (continuous) or brittle (discontinuous) deformation (e.g. Long and Imber, 2010). The throw maxima was then identified on each radial fault, and plotted as white squares on Fig. 3. Fault throw was also measured with depth (T-z plots) using the aforementioned cut-offs, and throw maxima marked by white squares on Fig. 4.

Expansion indices illustrate variations in sediment thickness adjacent to fault systems, revealing the kinematics of bounding faults (e.g. Thorsen, 1963; Tvedt et al., 2013; Jackson et al., 2017). Expansion indices were calculated by dividing the hangingwall thickness of a stratal units by its corresponding footwall thickness and plotting these against geological time. An expansion index of 1 suggests no across-fault thickening, and a lack of syndepositional fault activity. An index of >1 suggests across-fault thickening and syndepositional fault activity. An index of <1 suggests stratal thinning from the footwall to the hangingwall, and may reflect difficulties in accurately measuring stratal thicknesses adjacent to a fault. Expansion indices near vertical fault tips may be slightly above and below one (± 0.1) due to ductile deformation (e.g. Barnett et al., 1987). T1 – T4 represent horizon tops. The white circle represents the vertical fault tip.



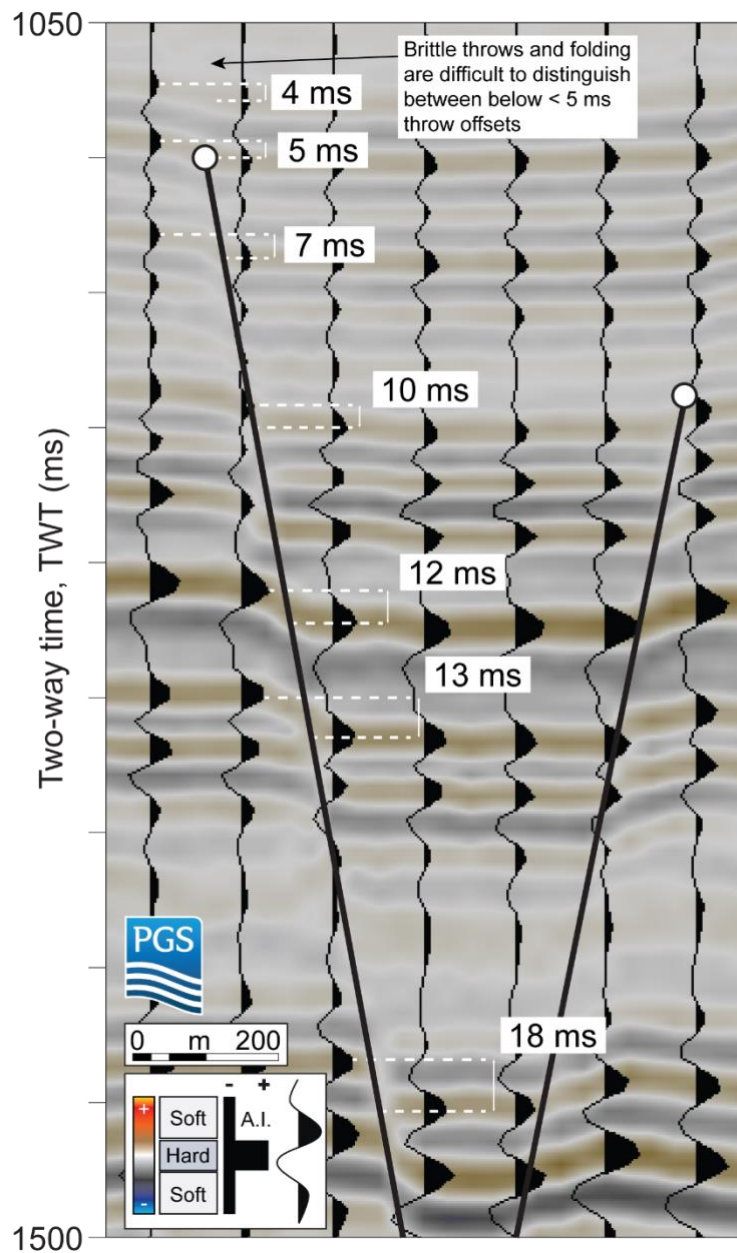
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APPENDIX 3. FAULT THROW RESOLUTION

Fault throw resolution for an example radial fault at Santos. Although vertical resolution may decrease with depth, the vertical offset between amplitude peaks between adjacent seismic traces permits fault throw to be measured to c. 5ms at shallow depths (< 3000 ms TWT). However, at greater depths (>3000 ms TWT), the peaks of individual traces become increasingly smeared as the vertical resolution decreases, and as such, vertical offsets are less distinct and measurement becomes increasingly difficult.



APPENDIX 4. ASPECT RATIO FOR SANTOS BASIN RADIAL FAULTS

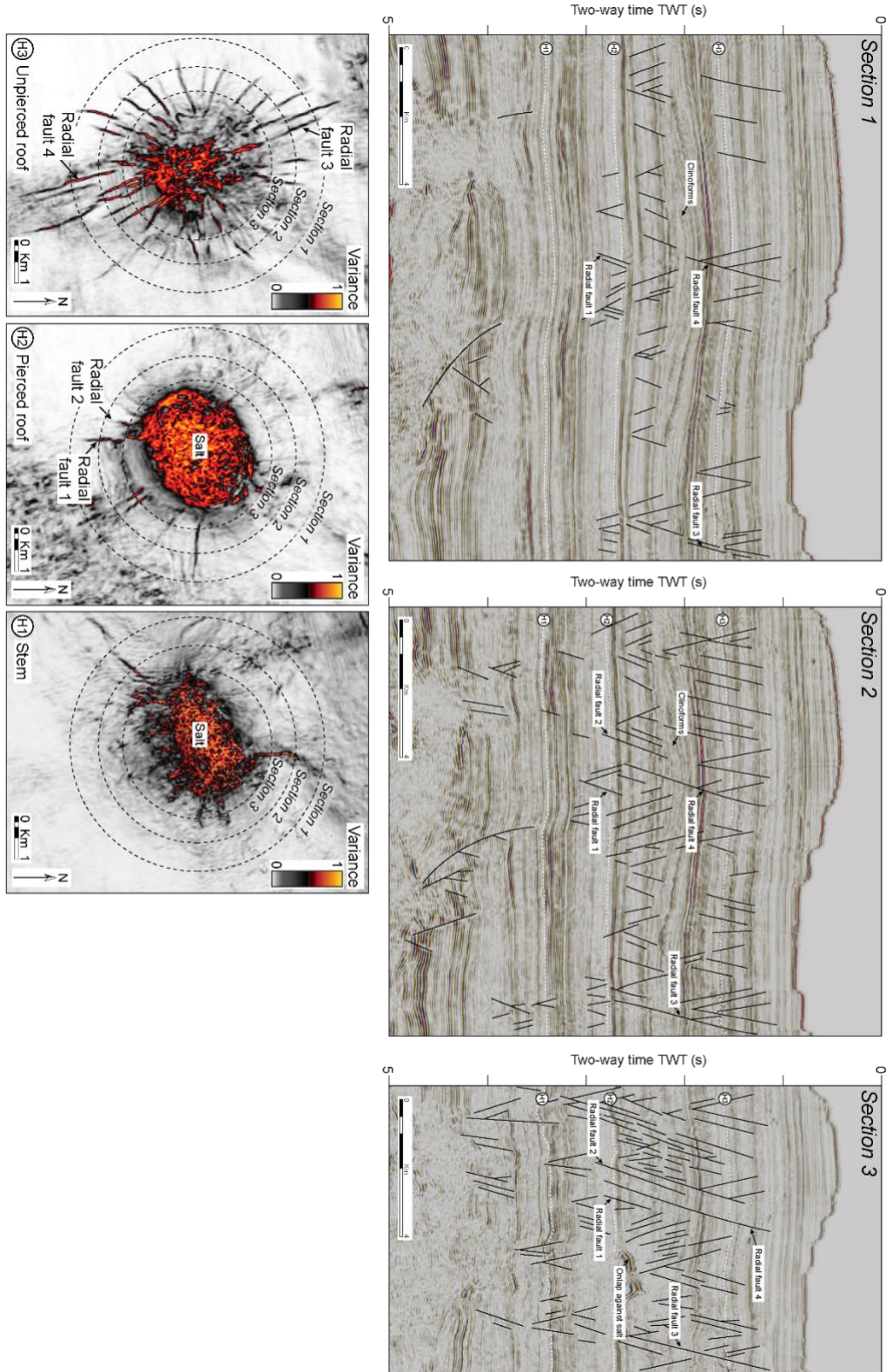
Aspect ratios for Santos Basin radial faults. Velocity ~ 2km/s after Jackson et al. (2014).

Fault #	Max Length (m)	Height (m)	Aspect Ratio
1	3727	871	4.28
2	704	596	1.18
3	639	639	0.69
4	2021	811	2.49
5	1340	809	1.66
6	1131	655	1.73
7	601	361	1.66
8	2909	1521	1.91
9	1075	434	2.48
10	1732	650	2.66
11	644	557	1.16
12	1536	923	1.66
13	833	833	0.72
14	1542	620	2.49
15	1826	683	2.67
16	1100	208	5.29
17	1742	666	2.62
18	1322	500	2.64
19	2001	736	2.72
20	809	545	1.48
21	579	579	0.78
22	1246	275	4.53
23	1969	501	3.93
24	947	323	2.93
25	2585	676	3.82
26	882	735	1.20
27	1726	669	2.58
28	1490	731	2.04
29	1713	368	4.65
30	600	580	1.03
31	1146	663	1.73
32	644	579	1.11
33	1471	782	1.88
34	1004	350	2.87
35	412	412	0.90
36	754	496	1.52
37	903	670	1.35
38	2510	401	6.26
39	1457	693	2.10
40	464	464	0.82
41	726	726	0.52

42	540	540	0.92
43	955	685	1.39
44	856	657	1.30
45	1328	493	2.69
46	1072	734	1.46
47	791	726	1.09
48	1619	566	2.86
49	1580	551	2.87
50	1276	707	1.80
51	1288	463	2.78
52	859	701	1.23
53	1764	699	2.52
54	1090	480	2.27
55	964	596	1.62
56	754	754	0.88

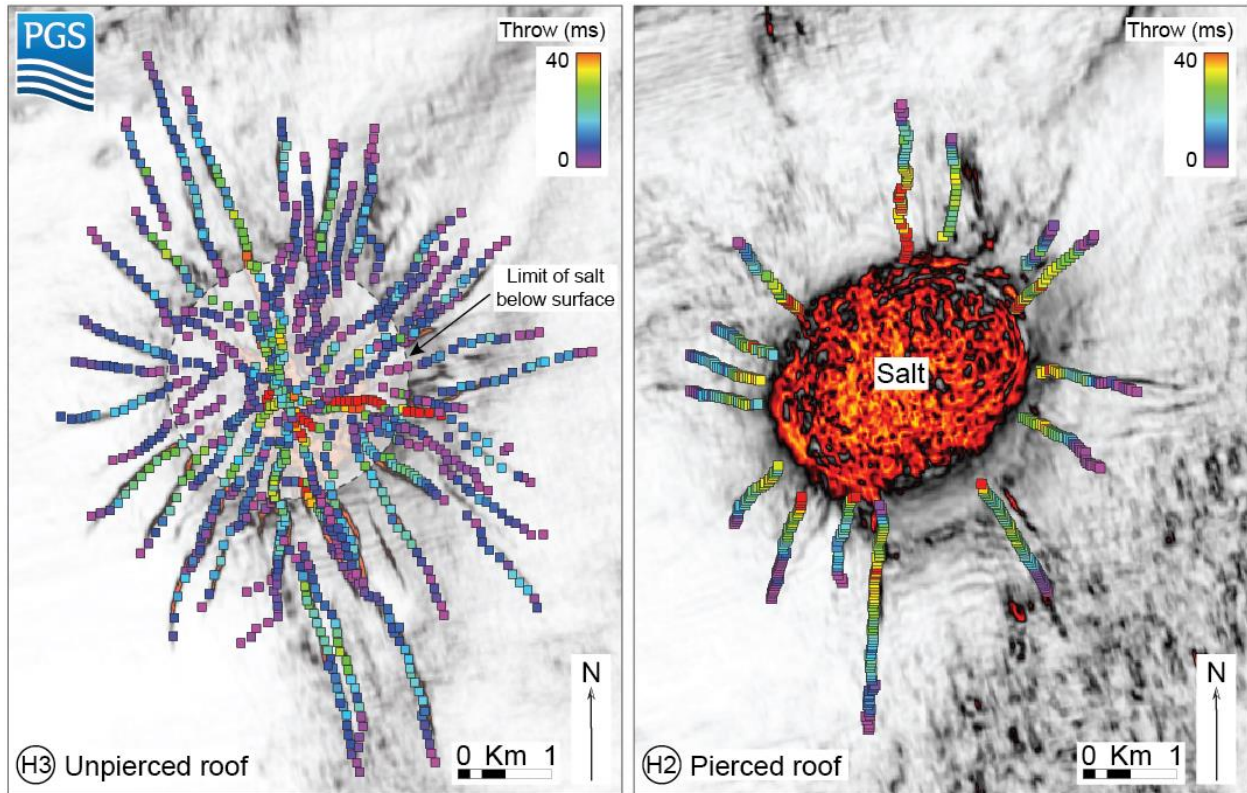
APPENDIX 5. CIRCUMFERENTIAL SEISMIC SECTIONS

Circumferential seismic sections parallel to the salt-sediment interface documenting the different vertical tiers of radial faults around the isolated salt stock. H1 – 3 and Faults 1 - 4 are also shown.



APPENDIX 6. RADIAL FAULT THROW FOR H2-3

Throw-distance on Fig. 3 used to determine the position of throw maxima along-strike for H2 - 3. Radial fault throw (i.e. strain) generally increases towards the diapir in H1 and H2.



APPENDIX 7. UPPER THROW TIP GRADIENTS

Upper throw tip gradients for the Santos Basin radial faults.

Fault	Throw (m)	Upper tip radius (m)	Vertical tip throw gradient
1	47	494	0.09
2	32	308	0.10
3	25	350	0.07
4	25	281	0.09
5	25	202	0.12
6	20	139	0.14
7	26	197	0.13
8	23	499	0.05
9	23	233	0.10
10	21	303	0.07
11	36	322	0.11
12	37	507	0.07
13	21	123	0.17
14	16	151	0.11
15	15	66	0.23
16	29	254	0.11
17	38	368	0.10
18	30	307	0.10
19	24	126	0.19
20	20	338	0.06
21	20	147	0.14
22	20	305	0.07
23	20	469	0.04
24	20	454	0.04

25	66	515	0.13
26	60	499	0.12
27	22	125	0.18
28	22	206	0.11
29	23	310	0.07
30	22	427	0.05
31	21	312	0.07
32	19	231	0.08
33	19	194	0.10
34	22	396	0.06
35	24	250	0.09
36	21	335	0.06
37	14	284	0.05
38	18	247	0.07
39	21	248	0.08
40	18	337	0.05