

16 **Abstract**

17 The lapse-rate feedback is the dominant driver of stronger warming in the Arctic than the
18 Antarctic in simulations with increased CO₂. While Antarctic surface elevation has been
19 implicated in promoting a weaker Antarctic lapse-rate feedback, the mechanisms in which
20 elevation impacts the lapse-rate feedback are still unclear. Here we suggest that weaker
21 Antarctic warming under CO₂ forcing stems from shallower, less intense climatological
22 inversions due to limited atmospheric heat transport above the ice sheet elevation and
23 elevation-induced katabatic winds. In slab ocean model experiments with flattened Antarctic
24 topography, stronger climatological inversions support a stronger lapse-rate feedback and
25 annual-mean Antarctic warming comparable to the Arctic under CO₂ doubling. Unlike the
26 Arctic, seasonality in warming over flat Antarctica is mainly driven by a negative shortwave
27 cloud feedback which exclusively dampens summer warming, with a smaller contribution from
28 the winter-enhanced lapse-rate feedback.

29 **Plain Language Summary**

30 Models project stronger surface warming in the Arctic than the Antarctic under climate change.
31 A climate feedback in which more warming occurs near the surface than at higher altitudes in
32 the atmosphere promotes this increased warming in the Arctic. Antarctica's surface elevation is
33 thought to weaken this feedback in comparison to the Arctic, but how this occurs is unclear.
34 Here we show that Antarctic elevation weakens surface warming by changing the base-state
35 vertical temperature structure. When Antarctic topography is flattened in model experiments,
36 Antarctica experiences more warming under climate change, resembling Arctic warming.
37 Similarly to the Arctic, flat Antarctica warms most during the winter, but this seasonality is
38 driven by different climate feedbacks in the Arctic versus Antarctic. These results indicate the
39 importance of base-state temperatures for warming under climate change, and suggest that
40 strong polar amplification is possible without local sea-ice loss.

41 **I Introduction**

42 The Arctic has warmed about twice the global average in recent decades in a pattern
43 known as Arctic amplification (Serreze et al., 2009; Screen and Simmonds, 2010a). In contrast,
44 Antarctic amplification is not observed in the same timeframe, and Antarctic warming is
45 dwarfed by Arctic warming in 21st century projections (Marshall et al., 2015; Smith et al., 2019).
46 While Antarctic warming is delayed in part by Southern Ocean upwelling and associated heat
47 uptake (Collins et al., 2013; Armour et al., 2016), global climate models also project weaker
48 equilibrium warming for the Antarctic than the Arctic (e.g., Danabasoglu and Gent, 2009). In an
49 evaluation of models in the fifth phase of the Coupled Model Intercomparison Project (CMIP5),

50 Goosse et al. (2018) find that the lapse-rate feedback is the single greatest factor contributing
51 to this hemispheric asymmetry in polar warming.

52 In a warming climate, a positive polar lapse-rate feedback results from stable
53 temperature inversions which contribute to stronger warming near the surface than aloft,
54 leading to inefficient longwave emission to space. In the Arctic, the ice-albedo feedback
55 promotes surface warming and thus contributes to a more-positive lapse-rate feedback
56 (Graversen et al., 2014). Dai et al. (2019) have more recently argued that Arctic sea-ice loss
57 enables the lapse-rate feedback via increased turbulent heat fluxes and upward longwave
58 radiation, enhancing lower-tropospheric warming over newly opened ocean. These results align
59 with evidence that sea-ice loss promotes seasonality in Arctic warming by enhancing winter
60 heat transfer from the ocean to the atmosphere and strengthening winter longwave radiative
61 feedbacks (e.g., Screen and Simmonds, 2010b; Bintanja and van der Linden, 2013). Given the
62 proposed dependence of Arctic lower-tropospheric warming and lapse-rate feedback on sea-ice
63 loss, a weaker Antarctic lapse-rate feedback may be driven by the persistence of the Antarctic
64 ice sheet.

65 Using fully-coupled Community Earth System Model (CESM) experiments, Salzmann
66 (2017) instead finds that Antarctic elevation drives the weaker Antarctic lapse-rate feedback. As
67 the sign and magnitude of the lapse-rate feedback depends on base-state static stability (Cronin
68 and Jansen, 2015; Payne et al., 2015), we expect Antarctic elevation to control the lapse-rate
69 feedback through impacts on base-state inversions. Confinement of the radiatively-active
70 atmospheric column over Antarctica to a shallower layer than in the Arctic may additionally
71 weaken the lapse-rate feedback.

72 With the mechanisms linking Antarctic elevation to a weaker lapse-rate feedback still
73 unclear, we investigate hemispheric asymmetry in the polar lapse-rate feedback using slab ocean
74 CESM experiments with present-day and flattened Antarctic topography under preindustrial
75 and doubled CO₂. We analyze Antarctic elevation impacts on climatological inversions, the
76 lapse-rate feedback, and polar amplification, in addition to investigating the seasonality of
77 warming in the flat Antarctic compared to the Arctic. These experiments provide insight into
78 hemispheric polar warming differences, with an ultimate goal of better understanding the
79 mechanisms behind polar amplification.

80 **2 Data and Methods**

81 2.1 Model Experiments

82 To investigate hemispheric asymmetry in the lapse-rate feedback, we use CESM (Hurrell
83 et al., 2013) version 1.2.2 with the Community Atmosphere Model version 4 (CAM4; Neale et
84 al., 2013), with a horizontal resolution of $0.9^\circ \times 1.25^\circ$ and 26 vertical levels. For all experiments,
85 CAM4 is coupled to a slab ocean forced with a spatially heterogeneous monthly climatology of
86 ocean heat flux convergence derived from a fully-coupled preindustrial control simulation (Bitz
87 et al., 2012). Differing from Salzmann (2017) in the use of a slab ocean rather than fully-coupled
88 model, these experiments isolate the role of Antarctic elevation in determining the equilibrium
89 climate response to CO₂ forcing and exclude potential effects of ocean heat uptake changes on
90 the lapse-rate feedback (e.g., Po-Chedley et al., 2018; Singh et al., 2018). CESM also uses the
91 Community Land Model version 4 (CLM4; Oleson et al., 2010) and the Los Alamos Sea Ice
92 Model version 4 (CICE4; Hunke and Lipscomb, 2008).

93 We perform CESM experiments with present-day topography, referred to as the
94 control Antarctic, or with the elevation of Antarctica flattened to 0 m above sea level, referred

95 to as the flat Antarctic. Branching from preindustrial runs for the control and flat Antarctic,
96 CO₂ is abruptly doubled from 285 to 570 ppm. We run each simulation for 50 years, and
97 calculate climatologies over the final 30 years. Net top-of-atmosphere (TOA) energy imbalances
98 of -0.09 W/m² (control, preindustrial), -0.06 W/m² (control, doubled CO₂), -0.14 W/m² (flat,
99 preindustrial), and -0.09 W/m² (flat, doubled CO₂) indicate near equilibrium for all experiments
100 in this period.

101 2.2 Radiative Feedbacks

102 We calculate the lapse-rate, Planck, surface-albedo, water-vapor, and cloud radiative
103 feedbacks using the radiative kernel method (Shell et al., 2008; Soden et al., 2008). This method
104 calculates radiative feedbacks as the product of (a) the change in radiative forcing per unit
105 change in a given climate variable, termed the radiative kernel, and (b) the modeled change in
106 this climate variable normalized by the surface temperature change. While the Planck feedback
107 is defined by propagating the surface temperature change through the entire troposphere, the
108 lapse-rate feedback calculates the effect of departures from this vertically uniform temperature
109 change. To calculate cloud feedbacks, the kernel method is used to determine the effect of
110 noncloud variables (temperature, water vapor, and surface albedo) on the change in cloud
111 radiative forcing (ΔCRF), and this cloud masking effect is subtracted from the total ΔCRF (Shell
112 et al., 2008; Soden et al., 2008). We calculate the residual term by subtracting the kernel-
113 estimated TOA radiation change from the modeled TOA radiation change, normalizing by the
114 surface temperature change.

115 Following Goosse et al. (2018), we also calculate feedback contributions to polar
116 warming (ΔTS) by dividing the energetic contribution of each feedback ($\lambda_i\Delta\text{TS}$), CO₂ forcing (F),
117 change in atmospheric heat transport convergence (ΔAHT), and residual term (R ΔTS) by the

118 magnitude of the Planck response in either the control Antarctic (Figure 3b) or control Arctic
 119 (Figure 3c) ($\lambda_{p,c}$), where $\lambda'_p = \lambda_p - \lambda_{p,c}$ is the difference between the local Planck response
 120 (λ_p) and $\lambda_{p,c}$:

$$121 \quad \Delta TS = -F/\lambda_{p,c} - \lambda'_p \Delta TS/\lambda_{p,c} - \sum_i \lambda_i \Delta TS/\lambda_{p,c} - \Delta AHT/\lambda_{p,c} - R\Delta TS/\lambda_{p,c}. \quad (1)$$

122 For this study, changes in climate variables are obtained from the doubled CO₂ minus
 123 preindustrial simulations, and radiative kernels are taken from Shell et al. (2008), calculated with
 124 an offline radiative transfer version of the Community Atmospheric Model, version 3 (CAM3).
 125 An alternative estimation of the albedo and shortwave cloud feedbacks using model-specific
 126 albedo kernels following Donohoe et al. (2020b) produces similar results (not shown) to those
 127 shown here using CAM3 albedo kernels. Feedbacks are calculated with respect to local rather
 128 than global surface temperature change (Armour et al., 2013; Feldl & Roe, 2013), and Arctic
 129 and Antarctic regional averages are defined for 70 to 90 °N and °S, respectively. For the flat
 130 Antarctica experiment, we apply zonally averaged kernels from the Arctic between 60 and
 131 90°N to the flat Antarctic between 60 and 90°S, shifted by six months. As the standard kernels
 132 do not exist below the Antarctic surface elevation, this allows for feedback calculation with the
 133 assumption that climate variable effects on TOA radiation for the flat Antarctic are similar to
 134 the Arctic. The kernel-estimated change in TOA radiation approximates the modeled TOA
 135 radiation change for the flat Antarctic almost as well as the control Antarctic experiments
 136 (Figure S1), supporting this flipped kernel method.

137 2.3 Atmospheric Heat Transport

138 We calculate AHT across each latitude using the poleward integral of the difference
 139 between TOA and surface energy fluxes (and seasonal atmospheric energy storage for seasonal
 140 averages) from monthly model output (e.g., Kay et al., 2012; Donohoe et al., 2020a). We

141 calculate latent heat transport as the poleward integral of evaporation minus precipitation at a
 142 given latitude, multiplied by the latent heat of vaporization, and dry static AHT is the residual
 143 between total AHT and latent AHT. As in Kay et al. (2012), we estimate a feedback associated
 144 with poleward AHT by dividing the change in AHT at the edge of polar regions by the surface
 145 area of these regions, normalized by the polar surface temperature change.

146 We also determine the vertical structure of AHT due to stationary eddies (SE) and
 147 transient eddies (TE) from CESM monthly output at 60°N and 60°S. For the vertical structure
 148 of TE transport, we neglect the potential energy term in Donohoe et al. (2020a), which is small
 149 at high latitudes. Vertical structures of SE and TE are shown in units of PW per 1000 hPa by
 150 zonally and vertically integrating the heat transport at each level as if it applied to the entire
 151 vertical column, yielding for each pressure level and latitude θ :

$$152 \quad SE = \frac{2\pi a \cos(\theta)(100,000 \text{ Pa})}{g} [\bar{V}^* \overline{MSE}^*] \quad (2)$$

153 and

$$154 \quad TE = \frac{2\pi a \cos(\theta)(100,000 \text{ Pa})}{g} [c_p(\overline{VT} - \bar{V}\bar{T}) + L(\overline{VQ} - \bar{V}\bar{Q})], \quad (3)$$

155 where a is the radius of the Earth, g is acceleration due to gravity, V is meridional velocity, MSE
 156 is moist static energy, c_p is the specific heat of air, L is the latent heat of vaporization of water, T
 157 is atmospheric temperature, and Q is specific humidity. Square brackets indicate zonal averages,
 158 overbars indicate monthly means, and asterisks denote departures from the zonal mean.

159 For comparison with CESM, we include vertical profiles of AHT and temperature from
 160 the ERA-Interim (Dee et al., 2011) and NCEP-NCAR (Kalnay et al. 1996) reanalyses, using 6-
 161 hourly fields for 1979-2018 to calculate SE and TE following Donohoe et al. (2020a).

162 2.4 Inversion Depth and Intensity

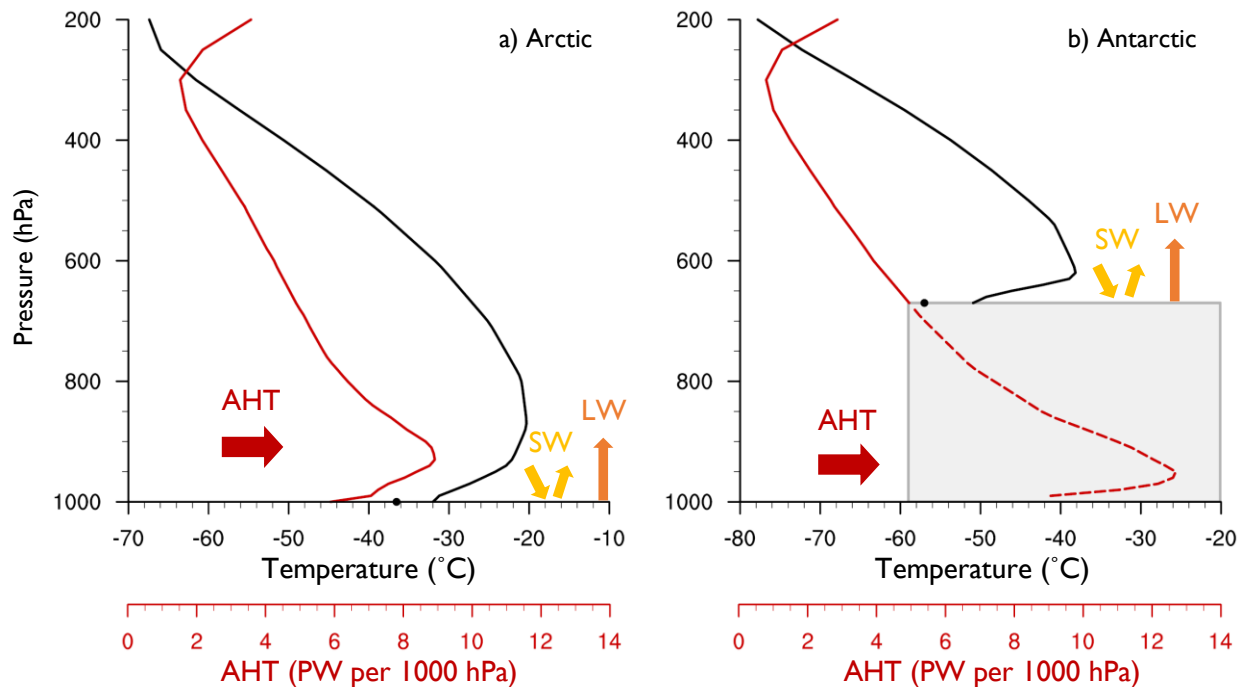
163 Following Zhang et al. (2011), we define surface-based inversion depth as the
164 geopotential thickness between the surface pressure and the first pressure level above which
165 temperature decreases with height, and inversion intensity as the difference in temperature
166 between these two levels.

167 **3 Results**

168 3.1 Climatological polar inversion asymmetry

169 Antarctic surface elevation drives modeled differences in base-state surface inversion
170 depth and intensity between the poles. Using the CESM preindustrial control simulation during
171 Arctic winter, Figure 1a shows a schematic of radiative-advective equilibrium, which controls
172 base-state inversions in polar regions (Cronin and Jansen, 2015; Payne et al., 2015). In this
173 framework, LW cooling and weak surface solar absorption promote cold near-surface
174 temperatures, while poleward AHT (shown at 60°N) maintains warmer temperatures aloft. As
175 is also seen in the NCEP and ERA-Interim reanalyses (Figure S2), AHT supporting these
176 inversions maximizes in the lower troposphere (around 900 hPa), near the tropospheric
177 temperature maximum. We propose that surface elevations above this level of maximum AHT
178 promote shallower, radiatively-driven winter inversions in the Antarctic (Figure 1b).

179 This preindustrial temperature profile at 90°S demonstrates CESM's ability to resolve
180 shallow surface inversions (here 50 hPa deep) matching radiosonde observations over the
181 Antarctic Plateau, although the model underestimates the intensity of observed mean
182 wintertime inversions which exceed 20 K at South Pole Station (Hudson & Brandt, 2005).
183 Weaker winter inversions over coastal slopes in CESM (Figure 2a; Figure S3c) align well with
184 coastal Antarctic radiosonde observations, which indicate typical depths shallower than 300 m
185 and intensities less than 5 K (Zhang et al., 2011).

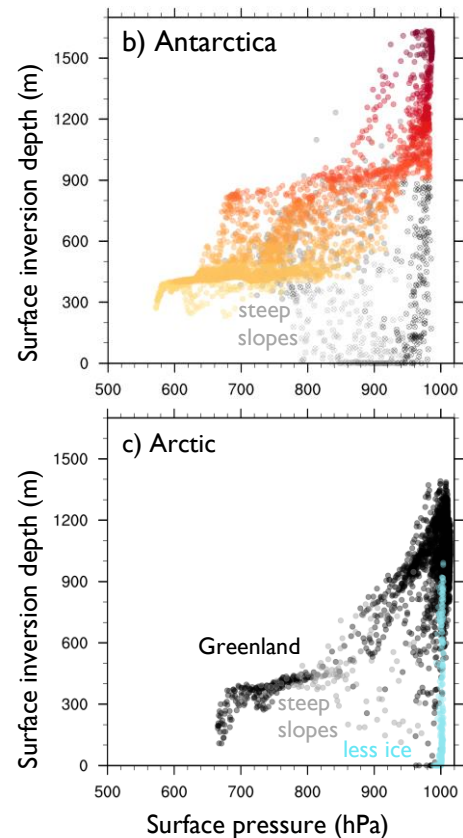
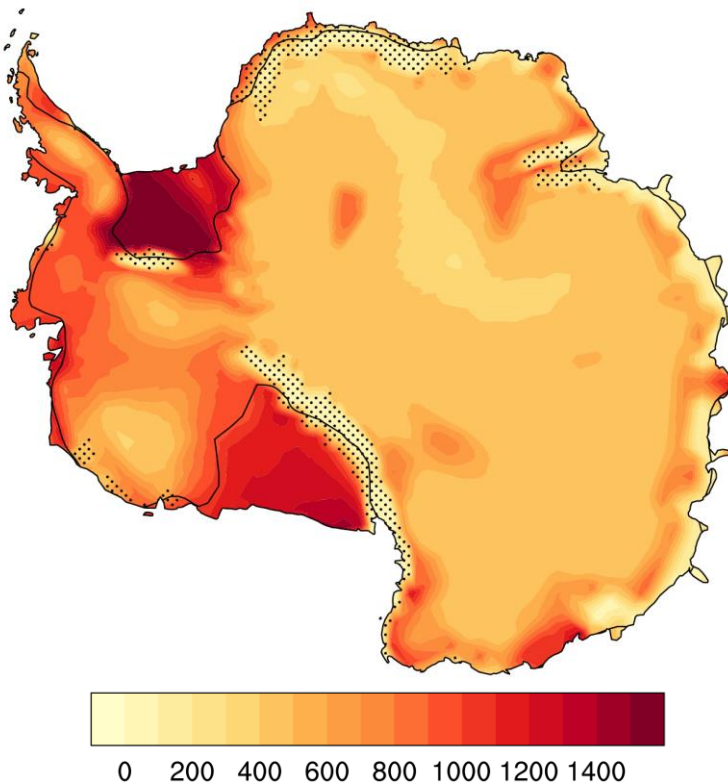


186 **Figure 1.** Schematic of radiative-advective equilibrium temperature based on the CESM preindustrial
 187 control topography experiment. Vertical profiles of winter poleward AHT due to transient and
 188 stationary eddies at (a) 60°N and (b) 60°S (red; PW per 1000 hPa) and winter temperature at (a) 90°N
 189 and (b) 90°S (black; dot for surface temperature; °C).

190 Considering CESM preindustrial winter inversion depth over the entire control
 191 Antarctic ice sheet, the deepest inversions are found at surface elevations below the level of
 192 maximum poleward AHT (black contour, Figure 2a). Shallower winter inversions exist over the
 193 Antarctic Plateau, and inversion depth generally increases with increasing surface pressure over
 194 Antarctica (Figure 2b). Inversions are shallower over steep slopes (grey circles, Figure 2b),
 195 where inversions may be disturbed by mixing due to katabatic winds (Vihma et al., 2011).
 196 Stippled in Figure 2a, points with shallow inversions for a given surface pressure (hollow grey,
 197 black circles, Figure 2b) are located either on or at the bottom of steep slopes (Figure S3a).
 198 In contrast to Antarctica, most points in the Arctic are at higher surface pressures,
 199 allowing AHT to support deep inversions (Figure 2c, S3b). Inversions are shallower over the
 200 Greenland ice sheet, particularly over steep slopes, and in regions with low sea-ice fraction.

201 While intense, shallow inversions (Figure S3c) exist at high Antarctic elevations, many areas
 202 show weaker inversion intensity than the Arctic (Figure S3d), especially over steep Antarctic
 203 slopes. These inversions in CESM are consistent with radiosonde observations of generally
 204 shallower, weaker inversions in the Antarctic compared to the Arctic (Zhang et al., 2011).

a) Surface inversion depth (m)



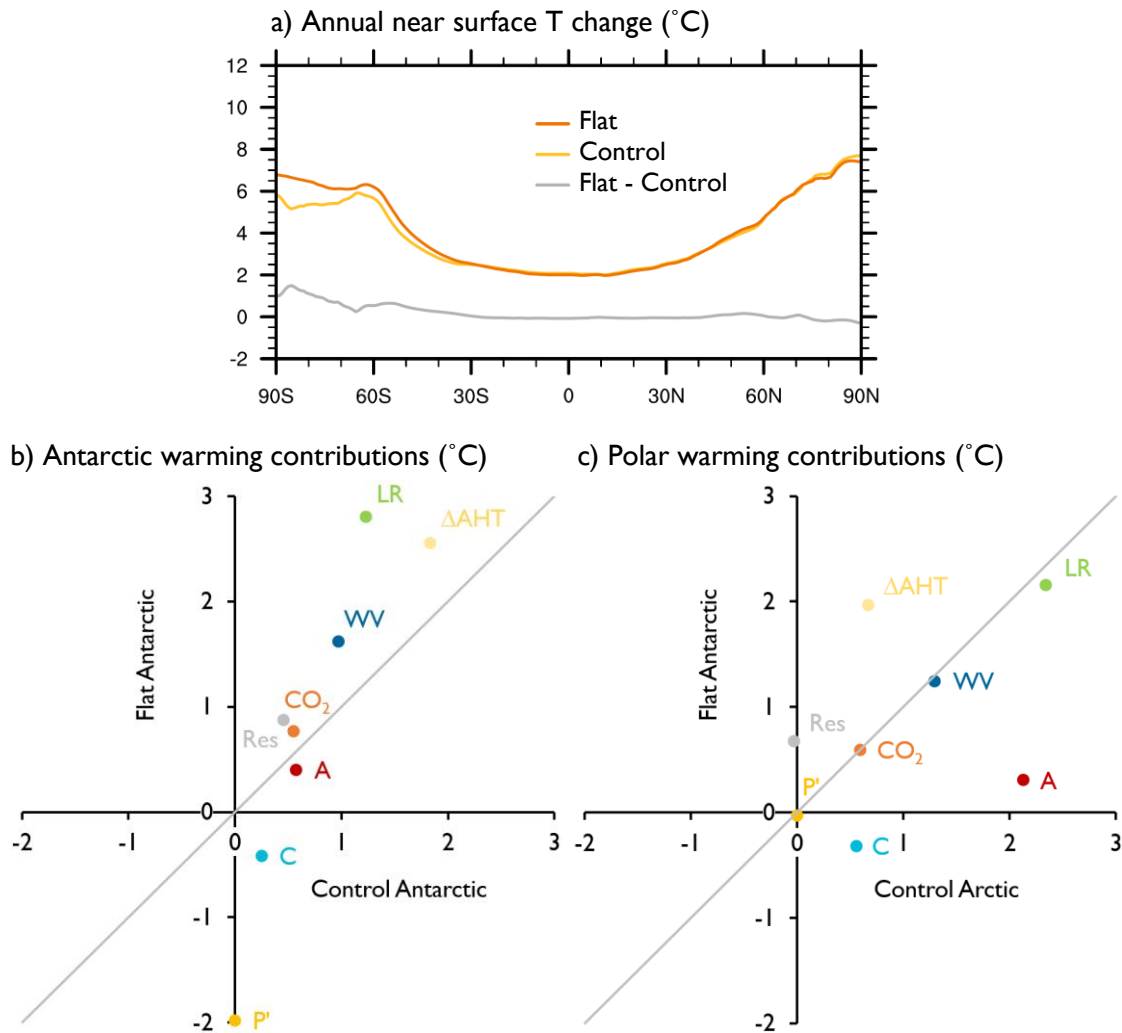
205 **Figure 2.** Winter surface-based inversion depth (m) for (a) Antarctica, (b) Antarctica from 70 to 90°S,
 206 and (c) the Arctic from 70 to 90°N from the CESM preindustrial control topography experiment. Black
 207 contour in (a) shows where the winter surface pressure equals 950 hPa: the level of the maximum AHT
 208 due to transient and stationary eddies during winter at 60°S. Colors in (b) match contours in (a), grey
 209 dots in (b,c) indicate surface slopes of at least .007, crosshatched grey and black dots in (b) correspond
 210 to stippled regions in (a), indicating shallow inversions for a given surface pressure on or at the bottom
 211 of steep slopes, and blue dots in (c) indicate winter sea ice fraction less than .97.

212 In comparison to the control Antarctic, inversions in the flat Antarctic preindustrial
213 experiment are intensified and deepened, particularly over steep slopes (Figure S4a,b). An
214 average inversion depth of 583 m and intensity of 7.5 K over the control Antarctic increases to
215 846 m and 11.9 K over the flat Antarctic, more comparable with the average inversion depth of
216 853 m and intensity of 9.0 K in the Arctic. We next investigate the extent to which these
217 stronger, deeper climatological inversions in the flat Antarctic simulation, which more closely
218 resemble Arctic inversions, may promote an Arctic-like lapse-rate feedback and stronger
219 Antarctic amplification.

220 3.2 Polar amplification and lapse-rate feedback asymmetry

221 In contrast to control Antarctic warming (5.4 K) under doubled CO₂, flat Antarctic
222 warming (6.3 K) is more comparable to control Arctic warming (6.7 K; Figure 3a) in the annual
223 mean. Applying Eq. (1), we find that the main driver of stronger amplification in the flat
224 Antarctic is the lapse-rate feedback (Figure 3b). A more-positive water-vapor feedback and
225 Δ AHT also support enhanced flat Antarctic warming, while more-negative Planck and cloud
226 feedbacks and a less-positive albedo feedback oppose flat Antarctic warming. The LW residual
227 term is similar for the control and flat Antarctic, although the SW residual term is slightly larger
228 in the flat Antarctic (Figure S5a). Flat Antarctica experiences lapse-rate, water-vapor, and Planck
229 warming contributions nearly equivalent to the Arctic, while a larger Δ AHT and residual term
230 in the flat Antarctic partly balance more-positive albedo and cloud feedbacks in the Arctic
231 (Figure 3c).

232 Figures S5b-g show the response to CO₂ doubling for flat minus control Antarctic
233 topography experiments for various climate variables relevant for feedbacks (left) and the
234 difference for the radiative kernels (right), where T is atmospheric temperature, TS is surface



235 **Figure 3.** (a) Zonal- and annual-mean near-surface temperature change ($^{\circ}\text{C}$) under CO_2 doubling in the
 236 control (yellow) and flat (orange) Antarctic experiments and their difference (grey); (b,c) Contributions
 237 of each feedback and atmospheric forcing to warming ($^{\circ}\text{C}$) for the flat Antarctic compared to (b) the
 238 control Antarctic and (c) the control Arctic for the lapse rate (LR), surface albedo (A), water vapor
 239 (WV), and cloud (C) feedbacks, the variation in the Planck response from its value in (b) the control
 240 Antarctic and (c) the control Arctic (P'), CO_2 forcing (CO_2), change in AHT convergence (ΔAHT), and
 241 residual term (Res).

242 temperature, q is specific humidity, and α is albedo. The strengthened lapse-rate feedback in the
 243 flat Antarctic is supported both by more surface-trapped warming in response to CO_2 doubling
 244 (more-negative $\Delta(T-T_S)$), and by a more-negative T kernel due to deepening and moistening the
 245 atmospheric column (Figure S5b,c). While it is difficult to quantify exactly how much of the
 246 increased lapse-rate feedback in the flat versus control Antarctic stems from the deeper

247 troposphere alone, we find that only 26% of this increase can be obtained from simply
248 extending the control Antarctic $\Delta(T-TS)$ from the lowest elevations above the ice sheet to all
249 pressure levels below the ice sheet and multiplying by the flat Antarctic temperature kernel.
250 This suggests that the increased lapse-rate feedback over flat Antarctica is strongly driven by
251 enhanced surface-trapped warming, rather than simply a deeper atmospheric column.

252 Surface-trapped warming enhancement for the flat Antarctic appears largest over the
253 Transantarctic Mountains and coastal slopes, where steep slopes promote shallow, weak
254 inversions in the control topography experiment (Figure S5b). Particularly for these regions,
255 increased preindustrial inversion depth and intensity in the flat experiment support stronger
256 low-level warming (Figure S4) and an enhanced lapse-rate feedback.

257 Both the stronger water-vapor feedback and weaker surface-albedo feedback in the flat
258 Antarctic are supported by deepening and moistening the atmospheric column. This amplifies
259 the greenhouse effect of water vapor by increasing the column-integrated specific humidity and
260 strengthening the water-vapor kernel (Figure S5d,e). The deeper atmospheric column over flat
261 Antarctica also dampens surface albedo impacts on TOA radiation, and this weaker (less-
262 negative) albedo kernel (Figure S5g) drives a weakened albedo feedback in the flat Antarctic.
263 Since feedbacks are normalized by local surface temperature change, the more-negative flat
264 Antarctic Planck feedback shown in Figure S5a can be attributed to a stronger atmospheric
265 temperature kernel due to a warmer, deeper emitting column.

266 The more-negative cloud feedback in the flat versus control Antarctic is explained by
267 more-negative SW cloud forcing (Figure S6a). The flat Antarctic experiences a larger increase in
268 cloud cover and cloud water path due to increased low-level liquid-bearing clouds under CO₂
269 doubling (Figure S6d,e), supporting a stronger negative SW cloud feedback. Preindustrial cloud

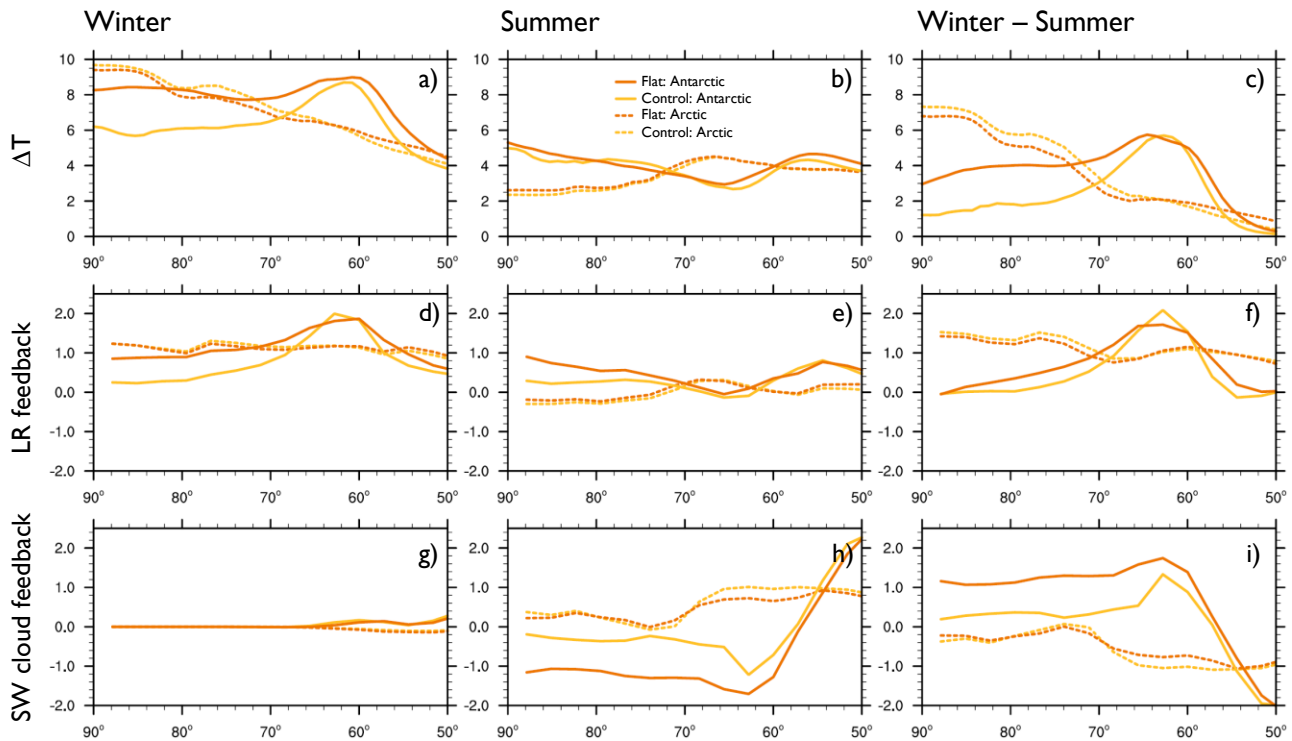
270 water path (not shown) is also larger for the flat Antarctic, likely due to an extended lower
271 atmosphere, enhanced water-vapor transport, and increased stability. Areas of stronger sea-ice
272 loss (Figure S6c) for the flat Antarctic also correspond to increased cloud water path and SW
273 forcing.

274 Increased preindustrial AHT toward the Antarctic balances stronger cooling to space
275 over flat Antarctica (Figure S7a), consistent with Singh et al. (2016) and Salzmann (2017). Under
276 doubled CO₂, AHT to the flat Antarctic also increases more, largely due to enhanced
277 southward latent AHT (Figure S7b). Δ AHT and feedback calculations for these slab ocean
278 experiments using the kernel method produce similar results to Salzmann (2017), who employ
279 partial radiative perturbation (PRP) feedback computations for the transient response to CO₂
280 doubling in fully-coupled runs: flat Antarctica experiences stronger lapse-rate and water-vapor
281 feedbacks and increased poleward Δ AHT, opposed by more-negative Planck and cloud
282 feedbacks and a less-positive albedo feedback.

283 3.3 Radiative feedback contributions to seasonality in Arctic and Antarctic amplification

284 Enhanced warming in the flat Antarctic compared with the control topography
285 experiment occurs predominantly during the winter season (Figure 4a-c), with similar summer
286 warming for flat and control experiments. This produces a greater difference between winter
287 and summer warming (4.0 K) in the flat Antarctic compared to the elevated Antarctic (2.1 K),
288 more comparable with control Arctic warming seasonality (5.3 K). To investigate what causes
289 enhanced seasonality in warming in the flat Antarctic experiment, we compare winter and
290 summer climate feedbacks. Feedbacks shown are normalized by the annual-mean local warming,
291 although we find similar results for normalizing by seasonal warming. Figures 4d-i highlight

292 seasonal lapse-rate and SW cloud feedbacks, with seasonality for all feedbacks shown in Figure
 293 S8.



294 **Figure 4.** Zonal-mean (a-c) near surface temperature change ($^{\circ}\text{C}$), (d-f) lapse rate feedback ($\text{W}/\text{m}^2/\text{K}$),
 295 and (g-i) shortwave (SW) cloud feedback ($\text{W}/\text{m}^2/\text{K}$) under CO_2 doubling in the control (yellow) and flat
 296 (orange) Antarctic experiments in the Antarctic (solid) and Arctic (dashed) for (a,d,g) winter, (b,e,h)
 297 summer, and (c,f,i) winter minus summer.

298 In the Arctic, the lapse-rate feedback strongly promotes greater warming in winter than
 299 summer (Figure 4f), consistent with Pithan and Mauritsen (2014). This enhanced winter lapse-
 300 rate feedback is driven by stronger base-state inversions in winter compared to summer, when
 301 sea-ice melting keeps surface temperatures near the freezing point. While strong lapse-rate
 302 feedback seasonality occurs over Southern Ocean sea ice, seasonality in this feedback is weaker
 303 over Antarctica for both the control and flat experiments. Flat Antarctica experiences stronger
 304 seasonality in warming than control Antarctica due in part to strengthened winter inversions,
 305 but the primary feedback enhancing flat Antarctic warming seasonality is the SW cloud
 306 feedback.

307 Due to polar night during winter, enhanced SW cloud cooling in the flat versus control
308 Antarctic occurs exclusively during summer (Figure 4h). While the lapse-rate and water-vapor
309 feedbacks promote greater warming for the flat Antarctic throughout the year, greater
310 warming seasonality over flat Antarctica results largely from the SW cloud feedback damping
311 only summertime warming. About two thirds of the difference in SW cloud feedback
312 seasonality between control and flat Antarctica arises from cloud radiative forcing, with the
313 other third coming from differences in the cloud masking term. A weaker summer surface
314 albedo feedback over flat Antarctica also contributes slightly to stronger winter versus summer
315 warming, with stronger albedo feedback seasonality changes over sea ice (Figure S8d).

316 **4 Conclusions**

317 With a goal of understanding the mechanisms driving lapse-rate feedback differences
318 between the Arctic and Antarctic, we compare CESM slab ocean experiments with control and
319 flattened Antarctic topography under preindustrial and doubled CO₂ forcing. We find
320 climatological differences in CESM preindustrial Arctic and Antarctic inversions, supported by
321 radiosonde observations. Limited poleward atmospheric heat transport above ice sheet
322 elevations, in addition to mixing due to katabatic winds on steep slopes, drives shallower,
323 weaker inversions for the Antarctic than the Arctic. Combined with weaker emission from the
324 relatively shallow and dry Antarctic atmospheric column, these weaker Antarctic inversions
325 prohibit the strong positive lapse-rate feedback seen in the Arctic. In contrast, over the flat
326 Antarctic, stronger climatological inversions support a lapse-rate feedback that is comparable to
327 that in the Arctic, with nearly equivalent degrees of polar amplification in each hemisphere in
328 these slab ocean experiments. While the seasonality of warming in the flat Antarctic is also
329 more comparable to the Arctic, this seasonality is supported by different climate feedbacks at

330 each pole: lapse-rate feedback seasonality contributes strongly to enhanced winter warming in
331 the Arctic, while a negative SW cloud feedback that only applies to non-winter months
332 contributes more to flat Antarctic seasonality.

333 As noted by Salzmann (2017), enhanced Antarctic amplification in flat Antarctica
334 experiments suggests that reduced Antarctic surface elevation due to mass loss would
335 accelerate Antarctic amplification under climate change. In the context of previous studies
336 considering the relative roles of sea ice and the lapse-rate feedback for polar amplification (e.g.,
337 Graversen et al., 2014; Dai et al., 2019), investigating Antarctic amplification can also provide
338 insight into mechanisms supporting Arctic amplification. Flat Antarctic experiments here and in
339 Salzmann (2017) demonstrate that strong, Arctic-like polar amplification is possible without
340 local sea-ice loss, although non-local sea-ice loss in the Southern Ocean likely contributes to
341 Antarctic amplification through changes in AHT. Our experiments additionally indicate that the
342 strongest seasonality in the lapse-rate feedback occurs in regions with sea ice: even over the
343 flattened Antarctic continent, lapse-rate feedback seasonality pales in comparison to lapse-rate
344 feedback seasonality over the Southern Ocean and Arctic. In line with Cronin and Jansen (2015)
345 and Payne et al. (2015), enhanced flat Antarctic warming supports the dependence of the lapse-
346 rate feedback on base-state inversions, which themselves depend on sea-ice concentration and
347 surface albedo in the Arctic. Further work to disentangle the lapse-rate feedback from sea-ice
348 effects may clarify how they will change separately and in tandem under climate change.

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356 model output relevant for study figures is available at <https://doi.org/10.5281/zenodo.3735441>
 357 (Hahn et al., 2020). ERA-Interim data was provided by the ECMWF Data Archive at
 358 <https://apps.ecmwf.int/datasets/data/interim-full-daily/levtype=pl/>, and NCEP data was provided
 359 at <https://www.esrl.noaa.gov/psd/> by the NOAA/OAR/ESRL PSD, Boulder, Colorado.

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