

1 **Simultaneous classification and location of volcanic deformation** 2 **in SAR interferograms using a convolutional neural network**

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6 **Key Points:**

- 7 • The three channels of pretrained CNNs cannot be used with different InSAR data
8 (e.g. phase and DEM).
- 9 • Our VolcNet database of InSAR time series contains up to ~500,000 labelled inter-
10 ferograms.
- 11 • Our CNN uses unwrapped data to differentiate between deformation patterns, and
12 determines their size.

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Abstract

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With the evolution of InSAR into a tool for active hazard monitoring, new methods are sought to quickly and automatically interpret the large number of interferograms that are created. We present a convolutional neural network (CNN) that is able to both classify the type of deformation, and to locate the deformation within an interferogram in a single step. We achieve this through building a “two headed model”, which returns both outputs after one forward pass of an interferogram through the network. We train our model by first creating a dataset of synthetic interferograms, but find that our model’s performance is improved through the inclusion of real Sentinel-1 data. When building models of this type, it is common for some of the weights within the model to be transferred from other models designed for different problems. Consequently, we also investigate how to best organise interferograms such that the filters learned in other domains are sensitive to the signals in interferograms, but find that using different data in each of the three input channels degrades performance when compared to the simple case of repeating wrapped or unwrapped phase across each channel. We also release our labelled Sentinel-1 interferograms as a database named VolcNet, which consists of ~500,000 labelled interferograms. VolcNet comprises of time series of unwrapped phase and labels of the magnitude, location, and duration of deformation, which allows for the automatic creation of interferograms between any two acquisitions, and greatly increases the amount of data available compared to other labelling strategies.

1 Introduction

In recent years, work to extend volcano monitoring to all of the world’s ~1400 sub-aerial volcanoes has resulted in the application of several machine learning methods to ground deformation maps produced by interferometric synthetic aperture radar (InSAR). Convolutional neural networks (CNNs) have been used in *Anantrasirichai et al.* [2018, 2019a] and *Valade et al.* [2019] to determine if individual interferograms contain deformation. This approach has been extended, through using cumulative time series, to more subtle deformation signals that are not visible in a single short temporal baseline Sentinel-

44 1 interferogram [Anantrasirichai *et al.*, 2019b]. Time series have been used by Sun *et al.*
45 [2020] to detect subtle deformation, with independent component analysis (ICA) by Gad-
46 des *et al.* [2018] to detect signs of unrest relative to a baseline stage of a volcano's be-
47 haviour, and with the CUSUM algorithm by Albino *et al.* [2020] to detect signs of unrest.
48 However, in all of the examples detailed above, each algorithm demonstrates very limited
49 knowledge of the diverse types of deformation that may be measured at volcanoes. The
50 algorithm presented in Anantrasirichai *et al.* [2019a] assigns all data containing deforma-
51 tion to one label, whilst the algorithms presented in Gaddes *et al.* [2018] and Albino *et al.*
52 [2020] alerts users to changes in the signals present, but does not identify the type of de-
53 formation present. Consequently, we seek to improve upon these approaches by developing
54 a CNN that is able to differentiate between different types of deformation, and to detect
55 the spatial extent of it.

56 Figure 1A shows the hierarchy of computer vision object/signal identification meth-
57 ods. The algorithm presented in Anantrasirichai *et al.* [2018] contains a model that per-
58 forms classification and, by breaking larger images into smaller tiles that are each classi-
59 fied, the algorithm as a whole is able to perform localisation. This approach has the lim-
60 itation that the deep learning model used in this algorithm does not need to learn how to
61 determine the location or size of the object (or signal) of interest, and at a more funda-
62 mental level, remains a classification and not localisation model. However, in the field of
63 computer vision, CNNs have been developed that are able to perform both classification
64 and localisation on images that contain either single or multiple objects. The location of
65 an object is either indicated through encompassing it in a rectangle (e.g. localisation or
66 object detection, Simonyan and Zisserman [2014]; Redmon *et al.* [2016]) or, in more com-
67 plex algorithms, indicating the exact outline of an object by identifying which pixels com-
68 prise it (e.g. instance segmentation, He *et al.* [2017]). These approaches should provide
69 more detailed information on the spatial extent of a signal of interest than a classification
70 model that is repeatedly used on different areas of the representation. Consequently, we
71 endeavour to advance the state of the art through developing a CNN that is able to both
72 localise deformation within an interferogram, and to classify different types of deformation
73 (the hierarchy of which we show in Figure 1B).

74 When constructing a CNN to perform both classification and localisation with data
75 derived from SAR satellites, a new CNN could be designed before all the parameters
76 within it are trained. However, this approach has the risk of failing to utilise both the suc-

77 successful structures and the learned parameters of CNNs that have been successfully applied
78 to other computer vision problems (e.g. the classification of natural images in *Krizhevsky*
79 *et al.* [2012] and *Simonyan and Zisserman* [2014], the instance segmentation of biomed-
80 ical images in *Ronneberger et al.* [2015], or the detection of buildings in satellite imagery
81 in *Zhang et al.* [2016]). In order to describe how we can utilize these successes, we must
82 first introduce the structure of a CNN in more detail, which we do with the use of Fig-
83 ure 1C. In this figure, a CNN can be seen to comprise of a convolutional part, and a fully
84 connected part. The convolutional part comprises of filters that are convolved across an
85 image to extract deep representations, whilst downsampling is performed simultaneously to
86 reduce the spatial size of the features as their depth increases. In the case of the example
87 network shown in Figure 1C, a three channel (colour) image of size $(224 \times 224 \times 3)$ pixels
88 is transformed into a spatially smaller but deep $(7 \times 7 \times 512)$ representation by this process.
89 In the second part, this 3D representation is flattened into a vector (which in this example
90 would be of size $(7 \times 7 \times 512 = 25088)$), before a traditional neural network comprising
91 of interconnected neurons is used to create the desired model outputs. The size of the last
92 layer of this second part is dependent on features such as the number of different classes
93 present in the data and, in this example case with two neurons in the last layer, would be
94 used in a case in which there were only two different classes.

95 Consequently, when using an existing model on a new problem, any change in the
96 number or type of output classes will require changing the fully connected part of the net-
97 work. Therefore, it is common to retain the structure of the convolutional layers (i.e. part
98 one of the model) and design a new fully connected network (i.e. part two of the model)
99 that outputs the classes required by the new problem. However, this approach still requires
100 the training of a CNN that is likely to contain tens of millions of parameters, which will
101 be both computationally expensive, and require a large volume of training data. AlexNet,
102 a previously state-of-the-art image classification CNN (named after one of the design-
103 ers, Alex Krizhevsky), has 60 million parameters, was trained on 1.2 million images, and
104 even when implemented on GPUs took around one week to train [*Krizhevsky et al.*, 2012].
105 Therefore, a common approach termed transfer learning is to retain both the structure and
106 weights of the initial convolutional layers, and to train only the fully connected part of the
107 network.

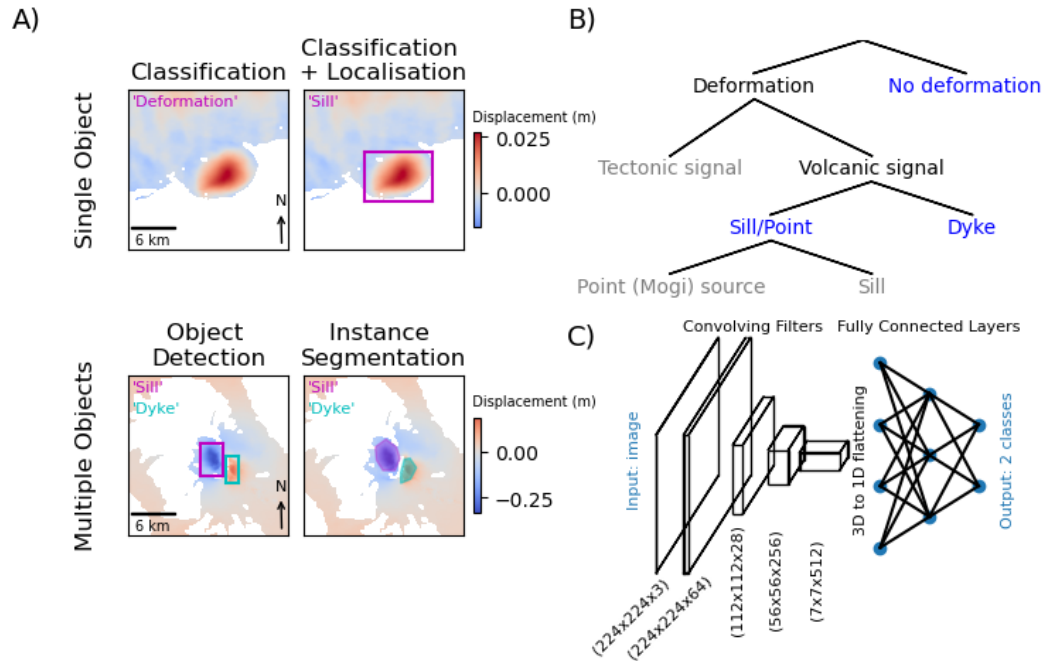
108 The weights learned in the convolutional filters of a CNN are of great importance
109 to a network's ability to detect features, as the filters must be sensitive to the patterns that

110 these features present in an image. As networks such as AlexNet [Krizhevsky *et al.*, 2012]
 111 and VGG16 (Simonyan and Zisserman [2014], named after the University of Oxford Vi-
 112 sual Geometry Group) were originally developed to compete in the ImageNet competitions
 113 [Deng *et al.*, 2009], the filters have been trained to detect the type of features present in
 114 natural images (e.g. photographs of a person, or car). When performing transfer learning,
 115 it is these filters that must be sensitive to the patterns presented in a deformation signal if
 116 the network is to correctly classify and locate it. However, as interferograms can be ex-
 117 pressed in differing formats we also seek to explore which of these formats allows for the
 118 filters in models trained on natural images to excel.

129 **2 Classification with different data formats**

130 As the most common CNNs for computer vision are trained on images comprising
 131 of a channel for each of the red, green, and blue values for each pixel, other data that are
 132 to be used with the network would also ideally be three channel. However, when consid-
 133 ering an image of interferometric phase, these images contain only a single value for each
 134 pixel, and so consist of only one channel. This difference in the number of channels can
 135 be circumvented through duplicating the one channel interferogram in each of the three in-
 136 put channels of a CNN, or by discarding parts of the filters of the first convolution (e.g. a
 137 filter of size $(5 \times 5 \times 3)$ be reduced to $(5 \times 5 \times 1)$). However, in this section of our study
 138 we wish to determine if this approach can be improved upon by utilising the three channel
 139 structure of many pre-trained CNNs to input more data to the model.

140 When two SAR images are combined to form a single interferogram, the result-
 141 ing image is a $2D$ array of complex numbers [Hanssen, 2001]. Whilst the magnitude of
 142 each of these complex numbers relates to the underlying brightness and coherence of a
 143 given pixel, it is common for only the argument to be displayed, as these phase values
 144 can be used to infer ground movement. However, the phase values of an interferogram
 145 are wrapped in the range $[-\pi, \pi]$ as only the fractional part of the phase value can be
 146 measured, but this ambiguity can be estimated to produce an unwrapped interferogram
 147 [Chen and Zebker, 2001]. We postulate that in addition to the use of either wrapped or
 148 unwrapped data duplicated to fill three channels, the original complex numbers of an inter-
 149 ferogram could be used in two channels, and so allow the network to use interferometric
 150 amplitude as an indicator of the reliability of the phase.



119 **Figure 1.** A) Introduction to the hierarchy of computer vision object/signal identification methods. The
 120 upper and lower rows show 12 day descending Sentinel-1 interferograms of Sierra Negra and Wolf volcano
 121 (Galapagos Archipelago, Ecuador), respectively. The Sierra Negra interferogram contains only one signal
 122 (an inflating sill), whilst the Wolf interferogram contains two signals (a deflating sill and an opening dyke).
 123 B) Proposed hierarchy for signals of interest in interferograms at volcanic centres. We propose a model that
 124 is able to classify interferograms into one of the three classes shown in blue: "no deformation", "Dyke", and
 125 "'Sill/Point". We envisage that future studies may add further classes which we mark in grey, such as those
 126 that differentiate between sills and point sources. C) Overview of a traditional convolutional neural network
 127 (CNN), showing how convolving filters and downsampling create a small but deep representation of an image
 128 ((224 × 224 × 3) to (7 × 7 × 512)), which is then flattened and passed through a traditional neural network.

151 However, we can also consider external data to feed into the CNN. When a human
152 observer interprets an interferogram, they are likely to use data such as a digital elevation
153 model (DEM) as this can be used to help determine if a signal is due to deformation, or
154 due to a topographically-correlated atmospheric phase screen. This problem is of partic-
155 ular importance at stratovolcanoes, as the cones typical of these volcanoes can be several
156 kilometres high, and therefore be capable of creating large and spatially stationary signals
157 in interferograms. The body of literature that covers the application of InSAR to volcanic
158 deformation is replete with studies that consider which of the two mechanisms are respon-
159 sible for the observed signals, and examples include *Beauducel et al.* [2000]; *Rémy et al.*
160 [2015]; *Yip et al.* [2019]. When considering previous attempts at the automatic detection
161 of deformation signals in Sentinel-1 interferograms, *Anantrasirichai et al.* [2019a] also re-
162 ported that many of the false positives recovered by their algorithm were caused by signals
163 correlated with topography. Consequently, we postulate that the inclusion of a DEM in the
164 inputs to our CNN will improve its ability to differentiate between deformation signals and
165 atmospheric signals that are correlated with topography, and therefore seek to investigate
166 its use as an input into a multichannel model.

167 To perform this analysis, we first synthesise a dataset of labelled interferograms.
168 To achieve this, we have created an open source Python3 package named *SyInterferoPy*,
169 which we make freely available to the community via GitHub: ([https://github.com/matthew-](https://github.com/matthew-gaddes/SyInterferoPy)
170 [gaddes/SyInterferoPy](https://github.com/matthew-gaddes/SyInterferoPy)). The collection of enough labelled data to train a CNN is com-
171 monly time consuming or expensive, and we find that the addition of localisation labels
172 to our data makes it more time consuming than in previous studies. Additionally, due to
173 the large number of data that are required to train CNNs and our expansion to classifi-
174 cation of different types of deformation, procuring enough real data to do this may not
175 be possible. Consequently, we perform this analysis using only synthetic data. Follow-
176 ing the hierarchy proposed in Figure 1B, we create interferograms that contain either no
177 deformation, deformation due to an opening dyke, or deformation due to a sill or point
178 source. These sources were chosen after reviewing the database of volcanic deformation
179 events measured using InSAR in *Biggs et al.* [2014] as we believe they cover the majority
180 of the observed signals that are of importance for volcano monitoring (i.e. we disregard
181 signals due to processes such as the cooling of lava flows). Model parameters were cho-
182 sen to be both physically realistic (e.g. dykes have near vertical dips), and for the resulting
183 deformation patterns to have absolute magnitudes in the range [0.05, 0.3] m which ensured

184 that the signals are visible over the synthetic atmospheric signals. We model the dykes
 185 as vertical dislocations with uniform opening in an elastic half space [Okada, 1985] with
 186 strikes in the range $[0, 359^\circ]$, dips in the range $[75, 90^\circ]$, openings in the range $[0.1, 0.7]$
 187 m, top depths in the range $[0, 2]$ km, bottom depths in the range $[0, 8]$ km, and lengths in
 188 the range $[0, 10]$ km. We model the sill/point sources as horizontal dislocations with uni-
 189 form opening in an elastic half space [Okada, 1985] with strikes in the range $[0, 359^\circ]$,
 190 dips in the range $[0, 5^\circ]$, openings in the range $[0.2, 1]$ m, depths in the range $[1.5, 3.5]$
 191 km, and widths and lengths in the range $[2, 6]$ km. It should be noted that our proposed
 192 hierarchy of volcanic deformation signals also includes processes that could be modelled
 193 as a point pressure source (commonly referred to as a ‘‘Mogi’’ source [Mogi, 1958]) within
 194 the sill/point category, but given that we do not envisage that a deep learning model us-
 195 ing satellite data from only one look angle (i.e. ascending or descending) would be able to
 196 differentiate between these two models, we generate our synthetic data using only one of
 197 them for simplicity.

198 These deformation patterns are then combined with a topographically correlated at-
 199 mospheric phase screen (APS), and a turbulent APS, which we discuss generating in more
 200 detail in *Gaddes et al.* [2018]. We calculate the topographically correlated APS using the
 201 Shuttle Radar Topography Mission (SRTM) 90m DEM [Farr et al., 2007], and use the
 202 coastline information contained within the product to mask areas of water. We also syn-
 203 thesise areas of incoherence within our interferograms, which we mask in order for our
 204 synthetic interferograms to be as similar as possible to the Sentinel-1 interferograms au-
 205 tomatically created by the LiCSAR processor [Lazec̆ky et al., 2020]. Figure 2 shows the
 206 results of mixing these different elements to create our synthetic interferograms, and the
 207 range of sizes of deforming regions that the different deformation model parameters pro-
 208 duce (e.g. Interferogram 2 versus Interferogram 3).

209 This process creates unwrapped data, which can be converted to wrapped data through
 210 finding modulo 2π of the unwrapped phase. However, to synthesise both the real and
 211 imaginary part of a complex interferogram requires knowledge of both the brightness of
 212 a pixel and its phase. To achieve this, we again use the SRTM DEM, and calculate the
 213 intensity of reflected electromagnetic radiation at the angles of incidence used by the
 214 Sentinel-1 satellites ($29.1 - 46.0^\circ$), before adding speckle noise, and calculating the in-
 215 terferometric amplitude between two images (i.e. the product of the two amplitudes). As
 216 inputs to CNNs that are to be trained using transfer learning must be rescaled to the in-

217 puts used in the original training data, we use only relative values in the range $[(-1), 1]$
 218 for the synthetic intensities. With knowledge of the modulus (relative intensity) and ar-
 219 gument (wrapped phase) of each pixel of our synthetic interferogram, the real/imaginary
 220 components are simply the products of the modulus and cosine/sine of the argument, re-
 221 spectively. Figure 3 shows five different ways we can represent an interferogram using the
 222 three channels available. Whilst this is not an exhaustive list of possible combinations or
 223 data sources, we believe that these five types are able to fully explore our hypothesis on
 224 the use of three channel data, yet are not so numerous as to be too computationally expen-
 225 sive to train.

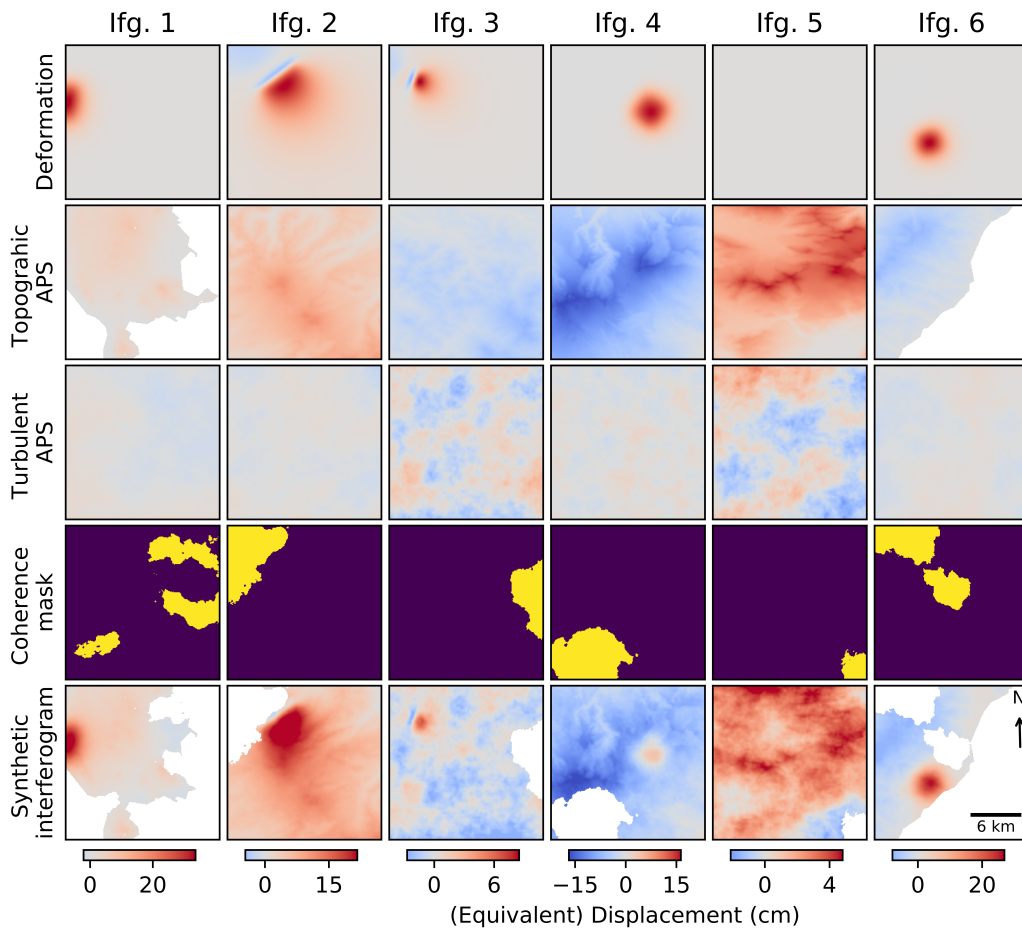
226 The CNN we build to classify the synthetic interferograms uses the five convolu-
 227 tional blocks of VGG16 [*Simonyan and Zisserman, 2014*], with our own fully connected
 228 network after this. This network was chosen as, when used in the field of computer vi-
 229 sion for classifying natural images, it outperformed older models such as AlexNet [*Si-*
 230 *mony and Zisserman, 2014*], which is used in the algorithm presented in *Anantrasirichai*
 231 *et al.* [2018], yet remains relatively simple to work with and train when compared to even
 232 newer models such as ResNet [*He et al., 2016*], and Inception [*Szegedy et al., 2015*]. Ad-
 233 ditionally, VGG16 was used by *Simonyan and Zisserman* [2014] to perform localisation of
 234 items it classifies, and therefore aligns with our goals. Figure 4B shows an overview of
 235 the model, in which interferograms of shape $(224 \times 224 \times 3)$ are passed through the five
 236 convolutional blocks of VGG16 to create a tensor of shape $(7 \times 7 \times 512)$. This is flattened to
 237 make a vector of size 25,088, before being passed through fully connected layers of size
 238 256, 128, and an output layer of size three (i.e., dyke, sill/point, or no deformation). The
 239 localisation output shown in the figure is not used in our preliminary exploration of which
 240 channel format to use (Section 2), but is used in Section 3. To produce a set of outputs
 241 that can be used as probabilities, we use a softmax activation for the last layer [*Bridle,*
 242 *1990*], but on the remaining layers we use rectified linear units (ReLus) to reduce compu-
 243 tation time [*Agostinelli et al., 2014*]. As our model seeks to solve a classification problem,
 244 we use categorical cross entropy for the loss function, which we seek to reduce using the
 245 Nadam optimizer as this does not require the choice of a learning rate [*Dozat, 2016*].

246 To train the model using the five different types of synthetic data, we perform what
 247 is termed “bottleneck learning” in machine learning literature (e.g. *Yu and Seltzer* [2011]).
 248 This method of training a CNN is used when only the weights within the fully connected
 249 layer are updated (i.e. transfer learning is being performed on the convolutional filters),

250 and comprises of first computing the results from passing our entire dataset through the
251 first five blocks of VGG16, before then training only the fully connected parts of our net-
252 work (i.e. the classification output). When a three channel image is passed through the
253 first five blocks of VGG16, a tensor of shape $(7 \times 7 \times 512)$ and termed a bottleneck fea-
254 ture is created, which we illustrate in Figure 4A. This method is highly efficient as we
255 do not generally wish to update the weights in the convolutional blocks of VGG16, yet
256 passing the data through these blocks is computationally expensive. By passing the data
257 through the convolutional blocks just once, we can then repeat only the relatively inexpen-
258 sive passes of the data through the fully connected parts of our network as we update the
259 weights contained within these layers. This method is of particular use for practitioners
260 who do not have access to high power computing facilities or GPUs.

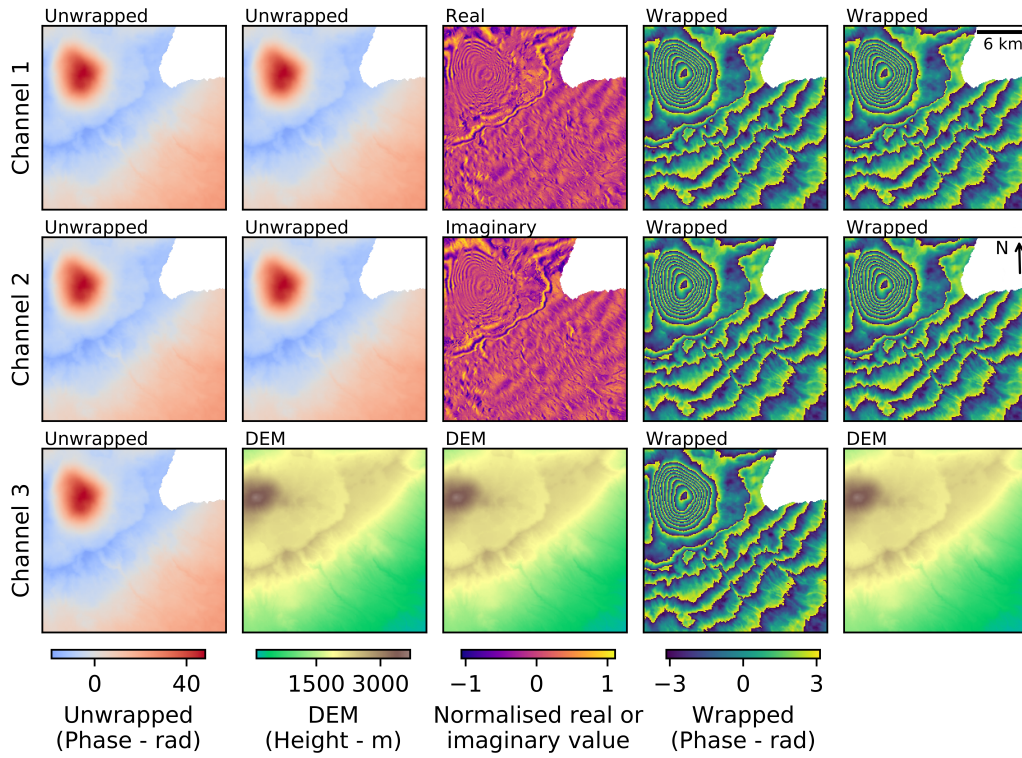
261 A common problem of CNNs that are used for classification can be overfitting of
262 the training data, which results in a model that generalises to new data poorly [Krizhevsky
263 *et al.*, 2012]. Overfitting is commonly caused by insufficient training data, but can also be
264 caused by issues such as using a model with too much complexity for the desired task,
265 or training a model for too many epochs [Chollet, 2017]. We endeavour to limit overfit-
266 ting through the use of dropout [Srivastava *et al.*, 2014] before both the 256 and 128 neu-
267 ron layers, as through randomly removing some connections during each pass of the data
268 through our model, this method aims to ensure that our model is forced to learn more ro-
269 bust representations of the training data. As we use synthetic data, we are not limited by
270 the usual cost of collecting labelled data, and therefore are able to generate 20000 unique
271 interferograms that are evenly distributed between classes without the use of data augmen-
272 tation.

273 Figure 5 shows the results of training five models with each of the data formats pre-
274 viously discussed. The highest classification accuracy achieved is ~ 0.95 , which is achieved
275 when the models are trained with either wrapped or unwrapped data repeated across the
276 three input channels. However, it should be noted that the accuracy of the unwrapped
277 phase model takes the full 20 epochs to achieve this performance, which contrasts with
278 the wrapped phase model which shows little change after the eighth epoch. Inclusion of
279 the DEM as the third channel appears to reduce classification accuracy, whilst very low
280 accuracies are achieved in the real and imaginary channel case. We discuss these results
281 in more detail in Section 4, but for the remainder of the paper we choose to work with
282 data that is unwrapped and repeated across the three input channels. We choose this ap-

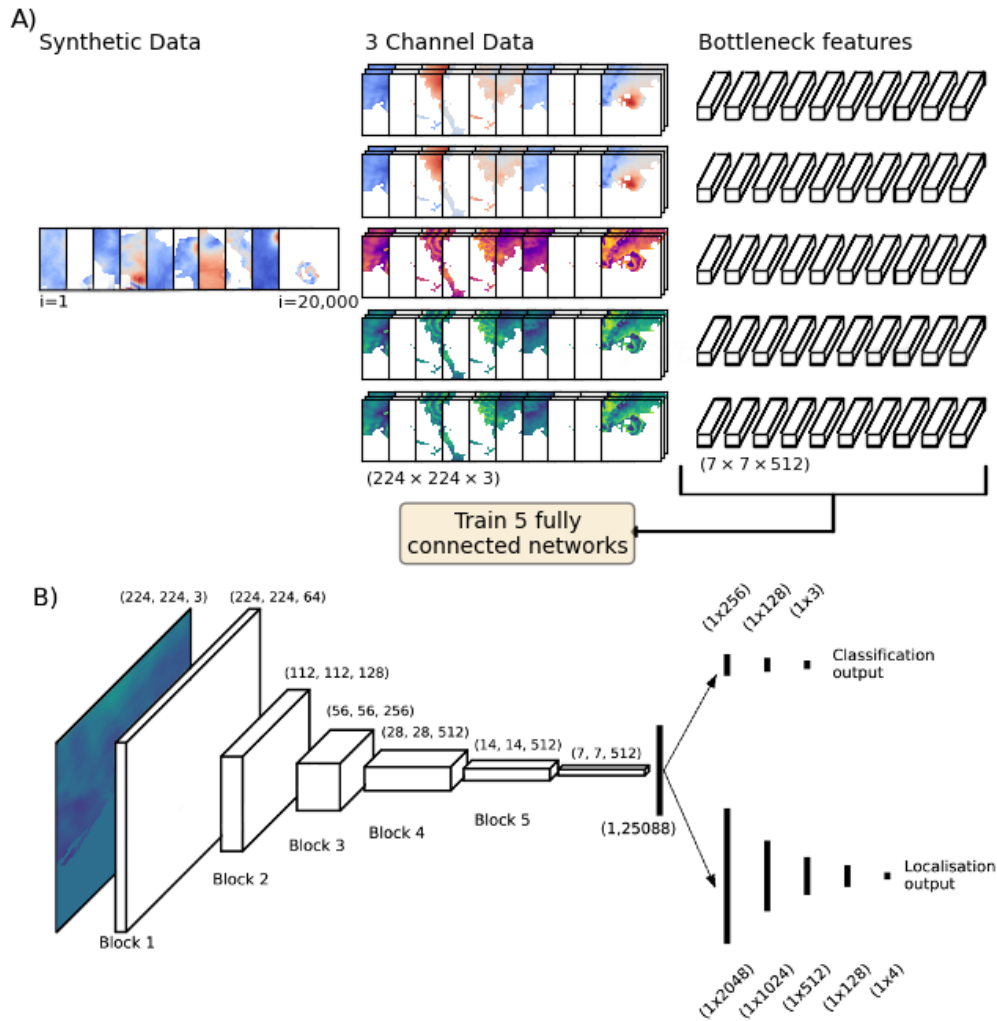


289 **Figure 2.** An example of the constituent parts of seven synthetic interferograms. Interferogram 5 does
 290 not feature deformation, interferograms 1, 4, and 6 feature deformation due to an sill/point source, and in-
 291 terferograms 2 – 3 feature deformation due to an opening dyke. These signals are geocoded and areas of
 292 water masked, before being combined with a topographically correlated APS, and a turbulent APS. Areas of
 293 incoherence are also synthesised, and these are used to mask the combination of the three signals to create the
 294 final synthetic interferograms.

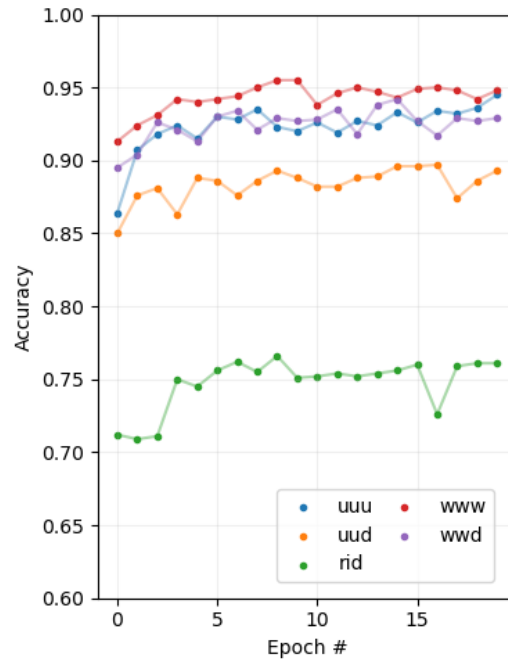
283 proach as no significant differences are seen between the classification accuracy ultimately
 284 achieved with either wrapped or unwrapped data, but the use of unwrapped data may al-
 285 low for a model to be used with unwrapped time series, and so detect subtle signals pro-
 286 duced by low strain rate processes. Additionally, a model that works with unwrapped data
 287 may also provide the opportunity to be expanded to locate and classify unwrapping errors
 288 automatically.



295 **Figure 3.** Organisation of an interferogram into three channel form. Columns one and two feature un-
 296 wrapped data that is repeated, and in column two the DEM is included as the third channel. In column three
 297 the real and imaginary elements of the complex values of each pixel of an interferogram occupy channels
 298 one and two, whilst the DEM is included in the third. Columns three and four feature wrapped data that is
 299 repeated, and in column five the DEM is included as the third channel.



300 **Figure 4.** A) Overview of our approach to creating a dataset of synthetic interferograms, arranging these
 301 into the five different three channel formats, computing the bottleneck features for each piece of data, and
 302 training the fully connected layers of a CNN B)Structure of our classification and localisation CNN. Input in-
 303 terferograms are first passed through the first five convolutional blocks of VGG16 to transform them from size
 304 $(224 \times 224 \times 3)$ to size (7×512) . These are flattened to create a large fully connected layer featuring 25088
 305 neurons, which is connected to both the upper branch/head, which performs classification, and the lower
 306 branch/head, which performs localisation. We find the localisation problem more complex than classification,
 307 and consequentially our localisation branch/head features more layers, each with more neurons. The output of
 308 the localisation head is a vector of four values determining the position and size of the deformation, whilst the
 309 output of the classification head is a vector of three values that indicate the probability for each class, and sum
 310 to one.



311 **Figure 5.** Accuracy of classifying validation data (10% of the total) during training using three channel
 312 data arranged in different formats. “u”: unwrapped data, “w”: wrapped data, “d”: DEM, “r” real compo-
 313 nent of interferogram, “i”: imaginary component of interferogram. Low accuracy is seen for the “rid” data,
 314 and in both the wrapped and unwrapped cases inclusion of the DEM in the third channel is seen to degrade
 315 classification accuracy. At the end of the 20 epochs of training, only a small difference is seen in accuracy
 316 between wrapped and unwrapped data, with both classifying ~95% of the validation data correctly, though the
 317 wrapped phase model is seen to achieve this level of accuracy more quickly (requiring only eight epochs of
 318 training). Whilst we see slight changes in the accuracy at the end of each of the latter epochs, we interpret the
 319 lines as having broadly plateaued and conclude that 20 epochs were sufficient for training these models.

3 Classification and localisation

3.1 Using synthetic data

In the previous section, we demonstrated that, when using VGG16 with convolutional weights learned on ImageNet data, roughly optimal performance for classifying synthetic interferograms is achieved when either the wrapped or unwrapped phase is repeated across the three input channels. We choose to progress with only the unwrapped phase model, as the computational cost of unwrapping is often already met by automatic processing systems (e.g. LiCSAR, *Lazeky et al.* [2020]), and the development of models that use unwrapped phase may lead to benefits such as the ability to classify and locate unwrapping errors. In this section, we build on the model used to perform classification by adding localisation output. We also endeavour to ascertain if the expense of collecting labelled data can be avoided entirely through the continued use of synthetic data when training our model.

We achieve both classification and localisation through dividing the fully connected section of our model to produce two distinct outputs. One output returns the class of the input data in the manner described in Section 2, whilst the second returns the location and size of any deformation within the scene. In machine learning parlance, models of this type are termed double headed, and we subsequently refer to either of the outputs and their corresponding preceding layers as either the classification head or localisation head. Figure 4B shows the structure of the two heads, and how they diverge after the output of the fifth block of VGG16 has been flattened. The localisation head is structured in a similar manner to the model described in *Simonyan and Zisserman* [2014], in which the model conveys the location of any deformation through outputting a column vector containing four values. Two of these values determine the centre of the deformation pattern and two display its horizontal and vertical extent. Together, these four values can be used to construct a box encompassing a deformation pattern. However, we find that an acceptable level of localisation performance cannot be achieved with a fully connected network with the same complexity as the classification head, and were required to increase both the number and size of layers in the localisation head's fully connected network. A simple network architecture search finds that the simplest model capable of achieving good performance has five layers consisting of 2048, 1024, 512, 128, and 4 neurons.

351 We use the mean squared error between the predicted location vector and the la-
352 belled location vector as our localisation loss function, which we seek to minimise. When
353 using three arc second pixels ($\sim 90\text{m}$) with a loss function of this type, a mean square er-
354 ror of 400 pixels would correspond to the localisation being incorrect by around $\sqrt{400} =$
355 20 pixels, or $\sim 2\text{km}$. However, when using a double headed network, training is compli-
356 cated by the fact that the model’s overall loss is now a combination of the classification
357 and localisation loss, which must be balanced using a hyperparameter commonly termed
358 loss weighting [Chollet, 2017]. In contrast to the localisation loss, we use categorical
359 cross-entropy for the classification loss and, as the value produced by this is generally sev-
360 eral orders of magnitude lower than the localisation loss, we find that a weighting of 1000
361 for the classification loss and 1 for the localisation loss produces a model which trains
362 well as the losses are approximately balanced.

363 To increase the performance of our classification and localisation model, we train it
364 using a two step process. In the first, we train it in a similar manner to that described in
365 the previous section, and update only the parameters within the fully connected network.
366 In the second step, we unfreeze the parameters in the fifth block (i.e. the last convolu-
367 tional filters of VGG16), and continue to train both these parameters and those contained
368 within the fully connected network. As the second step starts with parameters that are al-
369 ready approximately correct, optimizers that adaptively change the learning rate cannot be
370 used, as any initial large updates can destroy a model’s performance. Instead, we use the
371 “Adam” optimizer [Kingma and Ba, 2014] and, after experimentation, find a learning rate
372 of 1×10^{-5} neither destroys previous model performance, nor is too slow to train. As the
373 updates to the fifth block performed in the second step of our training preclude the use of
374 bottleneck features, we instead train our classification and localisation model on a Nvidia
375 GTX 1070 GPU.

376 **3.2 Application to real data: the VolcNet database**

377 Whilst the model described in the previous section achieved good performance when
378 classifying and locating deformation in synthetic interferograms, for use in automatic de-
379 tection algorithms we require our CNN to work with Sentinel-1 data. These data are of
380 particular importance for volcano monitoring, as the European Space Agency’s data pol-
381 icy ensures that Sentinel-1 data are available quickly and at no cost, whilst the low revisit
382 times ensure that the majority of sub-aerial volcanoes are imaged at least every 12 days.

383 We therefore create a database of labelled Sentinel-1 interferograms, which we term Volc-
384 Net and make freely available via GitHub: <https://github.com/matthew-gaddes/VolcNet>

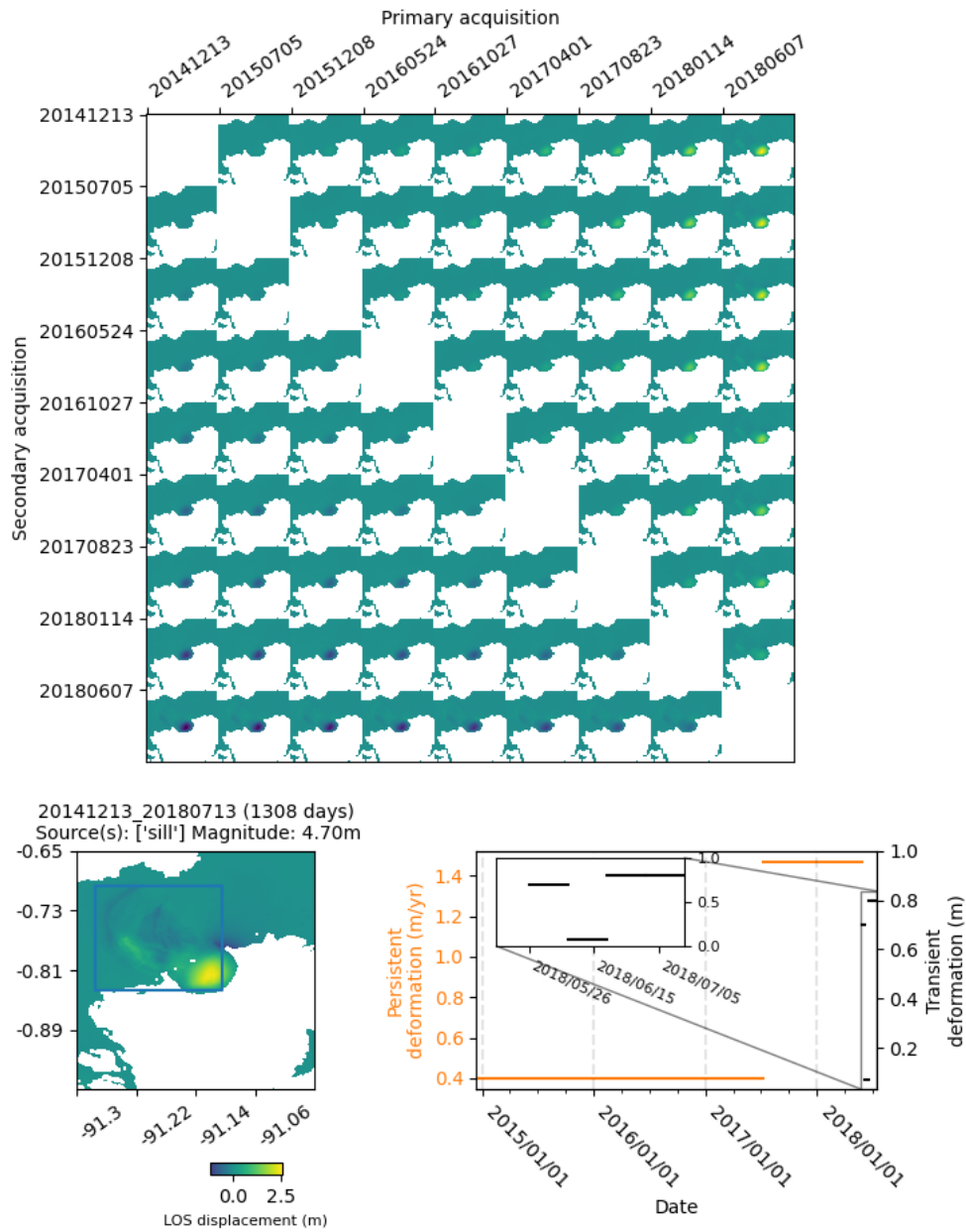
385 To populate our database, we chose a selection of volcanoes for which deformation
386 is known, and which the LiCSAR automatic interferogram processor ([https://comet.nerc.ac.uk/COMET-](https://comet.nerc.ac.uk/COMET-LiCS-portal/)
387 [LiCS-portal/](https://comet.nerc.ac.uk/COMET-LiCS-portal/)) had created networks of interferograms with no gaps and of long tempo-
388 ral duration (e.g. multiple years). This resulted in our use of Campi Flegrei, Vesuvius,
389 Agung, Wolf, Sierra Negra, Cerro Azul, Erta Ale, La Palma, and Domuyo. We filtered
390 the interferograms with a Goldstein filter [*Goldstein and Werner, 1998*], unwrapped using
391 SNAPHU [*Chen and Zebker, 2001*], and masked pixels with an average coherence below
392 0.7, before creating time series using LiCSBAS [*Morishita et al., 2020*].

393 To label our database, we develop an approach in which we create generic labels
394 that describe the duration, magnitude, and spatial extent of deformation for each volcano.
395 In contrast to traditional labelling approaches that assign a label to individual interfero-
396 grams (e.g. for InSAR data, *Anantrasirichai et al. [2018]* and *Bountos et al. [2022]*), our
397 approach allows us to create labelled interferograms between any two Sentinel-1 acqui-
398 sitions. Consequently, with relatively few labels, time series with N acquisitions can be
399 quickly converted into sets of $N^2 - N$ labelled interferograms. We define two types of
400 deformation label: transient deformation, which is relatively short lived and would be im-
401 aged by a syneruptive interferogram, and persistent deformation, which is generally of low
402 rate but spans multiple acquisitions. A choice of threshold is also required for the defor-
403 mation predicted by the label to be considered as visible in an interferogram, as in the
404 cases of persistent deformation of low-rate, we do not want our short temporal baseline
405 interferograms (e.g. 12 days) to be labelled as containing deformation. Figure 6 shows
406 the VolcNet data and label for Sierra Negra, as this contains both persistent deformation
407 (inflation prior to the 2018 eruption), and transient deformation (the 2018 eruption).

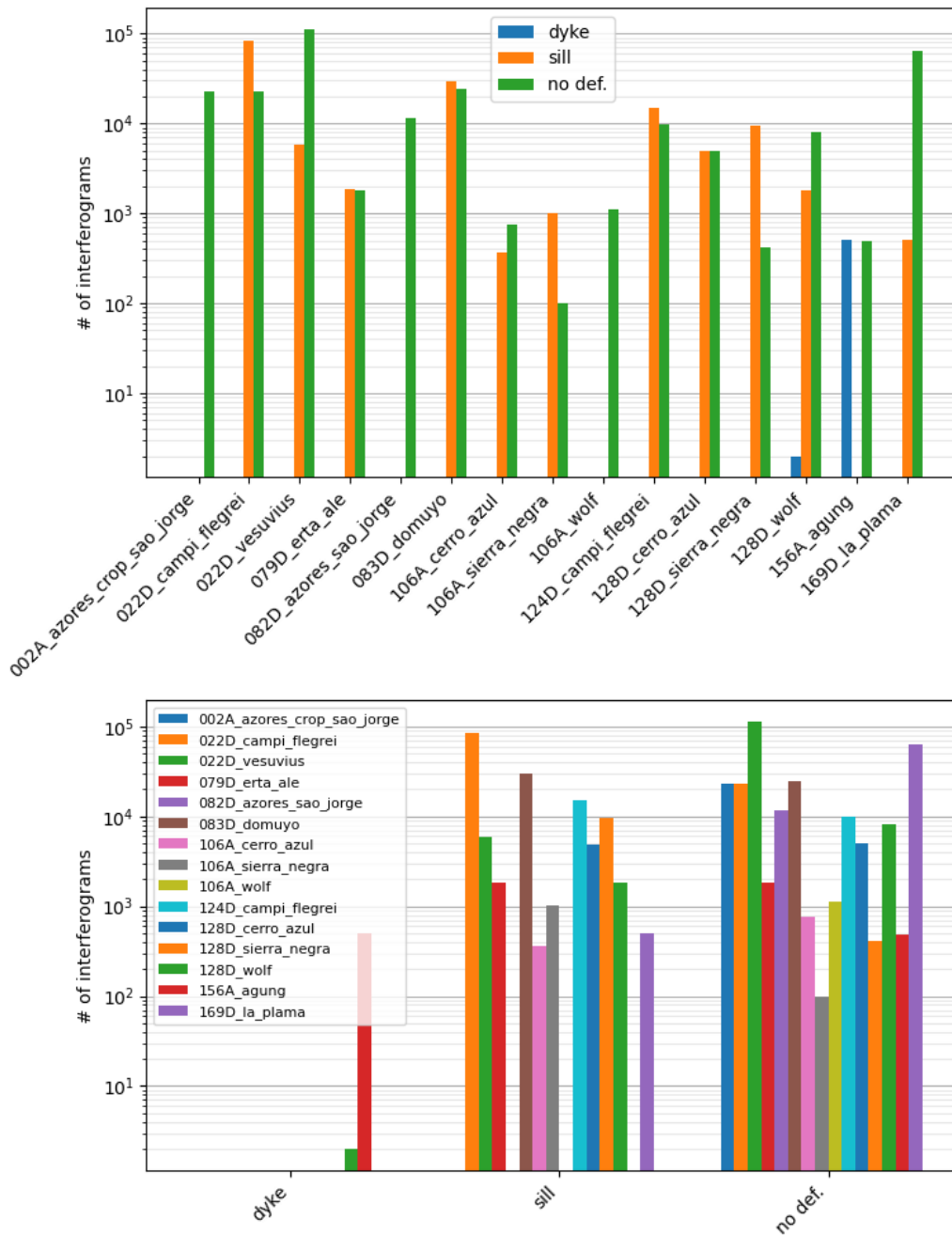
408 For the vast majority of time series in the collection, labelling was performed by
409 drawing on the results of previous studies in which inversions had been performed to fit
410 the signals observed in the interferograms, using *Albino et al. [2019]* for Agung, *Xu et al.*
411 *[2016]* for Wolf, *Gaddes et al. [2018]* for Sierra Negra, *Moore et al. [2019]* for Erta Ale,
412 and *Galetto et al. [2019]* for Cerro Azul. For the remaining time series, labelling was per-
413 formed through inspection of the signals present. Additionally, several of the studies from
414 which labels were created contain independent validation data in the form of ground truth

415 measurements made using global navigation satellite systems (e.g. Global Positioning Sys-
416 tem time series are available at Sierra Negra). These data ensure that signals present in
417 time series that are interpreted as being due to physical processes (such as the inflation of
418 a sill or point source) are not actually of atmospheric origin, and are in fact due to defor-
419 mation of the volcano. However, in some examples assigning a single class to a complex
420 deformation pattern is difficult, and we instead assign what we deem the dominant class to
421 be, whilst expecting that the network should assign some probability to other classes. This
422 is most evident at Wolf, in which signals were attributed to both the deflation of a sill and
423 the opening of a dyke [Novellis *et al.*, 2017; Xu *et al.*, 2016]. Figure 7 details the results of
424 labelling each of these time series, and then creating all possible interferograms between
425 all Sentinel-1 acquisitions.

442 Figure 8 shows the results of applying our trained classification and localisation
443 model to a quasi-random selection of Sentinel-1 interferograms from the VolcNet database
444 that we define as testing data (i.e. not be used when training models). Interferograms
445 such as Interferogram 8 show a very clear inflation signal at Sierra Negra, and are cor-
446 rectly classified by the CNN (“sill/point”), whilst the localisation is broadly correct. Other
447 promising results include the labelling of the two Wolf co-ruptive interferograms (inter-
448 ferograms ten and eleven) as containing a dyke (“sill/point”), which is also localised well.
449 However, some interferograms are wrongly classified, such as the subtle signal seen at
450 Vesuvius (interferogram zero), and the strong atmospheric signals at La Palma (interfero-
451 gram four), and Campi Flegrei (interferogram two). At Vesuvius, the deformation signal
452 is both small, and surrounded by incoherence and atmospheric signals, and is therefore
453 unlikely to be labelled by a human observer as deformation without inspection of the com-
454 plete time series. At Campi Flegrei and La Palma, the strong atmospheric signals juxta-
455 pose positive and negative signals in a manner somewhat similar to a dyke (our model’s
456 label), and this misclassification is likely to be due to our synthetic atmospheric signals
457 not being complex enough to allow our CNN to learn to differentiate between them and
458 deformation. The divergent nature of our CNN’s two heads also leads to outputs that show
459 disagreement between them. Interferogram six demonstrates this, in which deformation at
460 Erta Ale is localised approximately but the label is incorrect, although “dyke” has been
461 assigned a probability of 0.48. We again attribute this misclassification to a lack of com-
462 plexity in our synthetic data limiting what our CNN can learn, as the synthetic dykes we



426 **Figure 6.** Demonstration of VolcNet labelling for the time series that images Sierra Negra prior to and
 427 during the 2018 eruption. Upper: A subset (every 12th) interferogram that can be made between all possible
 428 acquisitions, showing the increasing deformation in longer temporal baseline interferograms. Interferograms
 429 along the diagonal are omitted as they contain only zeros. Lower left: Example of the longest temporal
 430 baseline interferogram that can be created, which features both inflation of the caldera floor (persistent de-
 431 formation), and complex syneruptive deformation propagating to the north west (transient deformation), for
 432 which a single bounding box is automatically created. Lower right: Graphical representation of the labelling,
 433 which shows an approximation of the increase in inflation rate that was observed approximately one year
 434 before the eruption as an increase in the height of the orange line, and the large but short-lived syneruptive
 435 signals in blue.



436 **Figure 7.** Summary of the VolcNet database. Top: Number of interferograms that can be created at each
 437 volcano divided into label type (sill/dyke/no deformation), showing the scarcity of time series that contain
 438 deformation attributed to dykes (Agung and Wolf, in blue). Many volcanoes are imaged in both ascending
 439 and descending orbits (e.g 128D and 106A for Sierra Negra), and some volcanoes feature in two frames (e.g.
 440 124D and 022D for Campi Flegrei). Bottom: Number of interferograms that can be created of each label type,
 441 showing the scarcity of interferograms that contain deformation attributed to a dyke.

463 use for training (e.g. Figure 2, interferograms two and three) are generally more elongate
464 and less complex than the signal seen at Erta Ale.

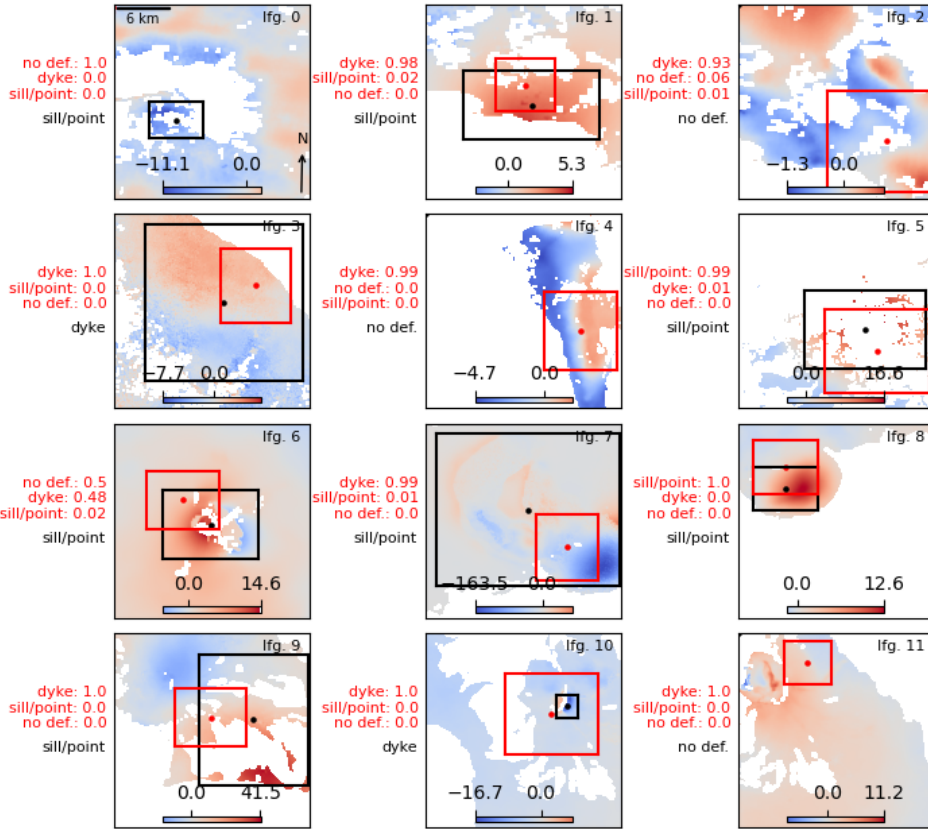
474 **3.3 Augmentation of training data with the VolcNet database of Sentinel-1 data**

475 To increase the performance of our model further, we seek to incorporate real Sentinel-
476 1 data from our VolcNet database into the training. Figure 7 details the distribution of
477 labels amongst the fully labelled database, from which the relative scarcity of those la-
478 belled as “dyke” can be seen ~ 500 , compared to $\sim 1 \times 10^5$ for “sill” and “no deformation”).
479 This class imbalance in the raw data requires preprocessing to ensure our real training
480 data is as balanced as our synthetic data (i.e. equally), which we achieve through select-
481 ing only a random subset of the “sill” and “no deformation” interferograms, resulting in
482 ~ 1500 labelled interferograms for training and validation use. However, 20000 synthetic
483 interferograms were used to train the previous model, and the inclusion of ~ 1500 new in-
484 terferograms is unlikely to impact the model significantly as these could still be classified
485 poorly with minimal increase in the loss function. We therefore apply data augmentation,
486 which involves creating random flips, rotations, and translations of the interferograms to
487 extend our set of real training data to feature 20000 unique, though often highly corre-
488 lated, Sentinel-1 interferograms. With the exception of including real data, we train our
489 model in the same manner as described in the previous section.

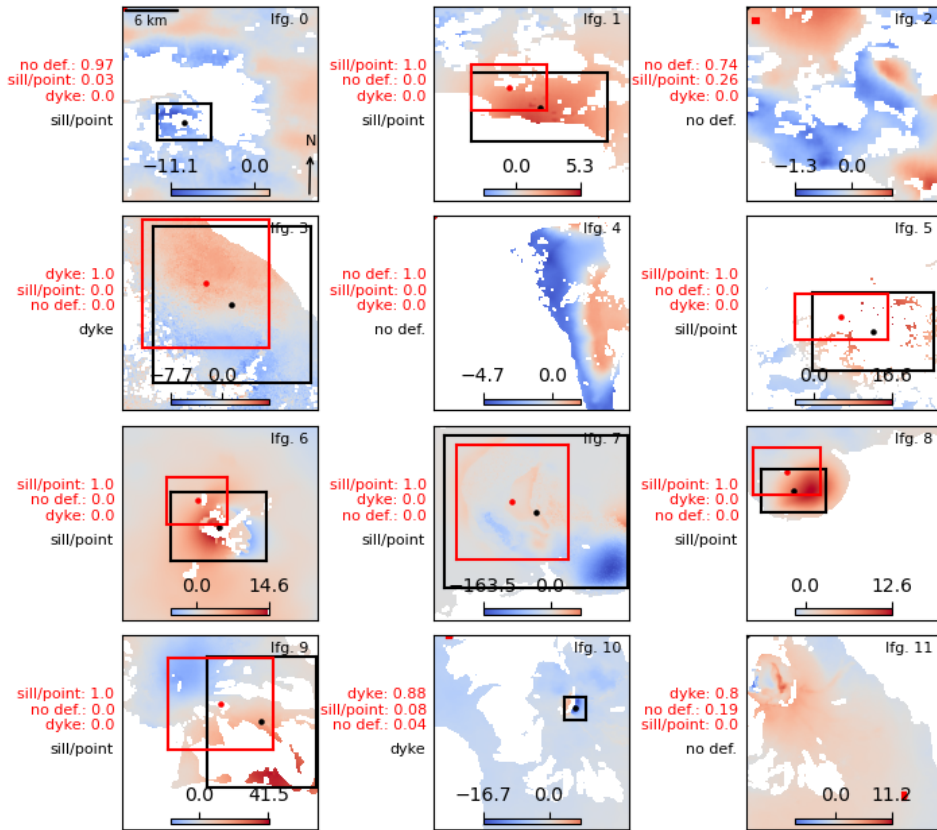
490 Figure 9 shows the results of applying our CNN to the same test set of real Sentinel-
491 1 VolcNet test interferograms used in Section 3.2. Inspection shows that our model is now
492 better able to handle interferograms with strong atmospheric signals, with interferograms
493 two and four now correctly classified as “no deformation”. Localisation is also improved,
494 with visibly smaller errors for interferograms three, and seven. Figure 10 compares the
495 results from the two models across the complete set of VolcNet test data (1000 interfero-
496 grams), and in all classes both the localisation loss and classification accuracy can be seen
497 to be improved through the incorporation of the real data.

507 **4 Discussion**

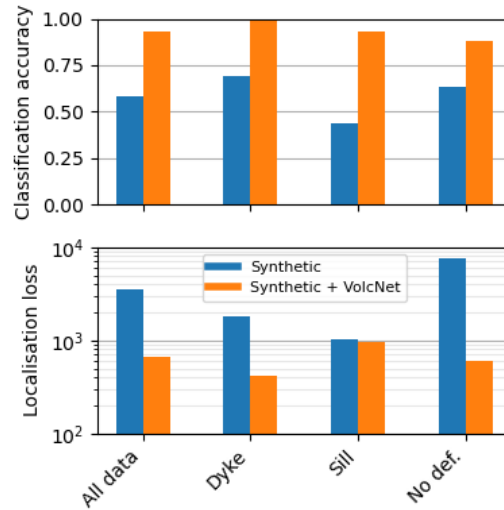
508 From the analysis performed in Section 2 we conclude that the incorporation of a
509 DEM into our CNN could not be achieved through the relatively simple step of using it as
510 one channel in multichannel data. This is likely because the weights in the first five con-



465 **Figure 8.** Results of our classification and localisation CNN on our testing set of Sentinel-1 interfero-
 466 grams when the CNN has been trained on synthetic data only. Model predictions are shown in red (including
 467 classification probabilities as decimals), and VolcNet labels are shown in black, with deformation shown in
 468 centimetres. Interferograms 0 : Vesuvius, 1 – 2 : Campi Flegrei, 3 : Agung, 4 : La Palma, 5 : Domuyo, 6 :
 469 Erta Ale, 7 – 8 : Sierra Negra, 9 : Cerro Azul, 10 – 11 : Wolf. Interferograms two and four feature strong at-
 470 mospheric signals which are misclassified as deformation, and the subtle deformation in zero is misclassified
 471 as no deformation. However, in the remaining cases both the classification and localisation is broadly correct,
 472 and in 11 the model classification and localisation outperforms the automatic labelling of the VolcNet data as
 473 subtle deformation is visible that falls below the threshold for being labelled.



498 **Figure 9.** Results of our classification and localisation CNN on our testing set of Sentinel-1 interferograms
 499 after incorporating real data into the training. The labelling convention and interferograms are as per Figure
 500 8. This model can be seen to outperform the CNN trained only on synthetic data, with improved classification
 501 and localisation in cases such as two and four where strong atmospheric signals are now classed as “no defor-
 502 mation”, and seven where the large deformation signal is localised correctly. However, several errors remain,
 503 such as the incorrect localisation of a co-ruptive Wolf interferogram (10).



504 **Figure 10.** Comparison of the VolcNet test data when evaluated with the model trained with only synthetic
 505 data (blue), and with both synthetic and real data (orange). In all cases, the synthetic and real model has both
 506 higher classification accuracies, and lower localisation losses than the synthetic only model.

511 volutional blocks of our model were transferred from VGG16 and, as VGG16 was trained
 512 using natural images, inputs which are broadly similar across all three channels are re-
 513 quired. It should be noted that we rescaled our training data to lie in the same range as
 514 the data that VGG16 was trained on (described further in Section 2), and therefore the
 515 lack of similarity across channels we refer to is not due to different magnitudes, but rather,
 516 different spatial patterns. However, an approach where the weights within the convolu-
 517 tional blocks of a classification and localisation model were trained from scratch may eas-
 518 ily allow for the incorporation of extra data in the different input channels. Whilst this
 519 approach was considered during the design of this study, we do not expect that training a
 520 CNN from scratch (i.e. training both the convolutional filters and the fully connected net-
 521 work) is feasible with only ~1500 real Sentinel-1 interferograms (i.e. the subset of data
 522 in which we balance our three data classes), and we did not have the resources available
 523 to create a larger database of labelled data. The results presented in *Sumbul et al.* [2019]
 524 explore this theme further, and they find that when using ~ 600,000 labelled Sentinel-2
 525 images they are able to train a shallow CNN with a channel for each of Sentinel-2's 13
 526 spectral bands that outperforms a deeper model that was pre-trained using ~ 1.2 million
 527 ImageNet images and used only the three visible Sentinel-2 spectral bands. Therefore, we
 528 expect that it is likely that through developing the VolcNet database that we introduce in

529 this work (and make freely available to the community), models that are able to use dis-
530 parate data in the different channels may be trainable, with resulting increases in perfor-
531 mance over the model presented here.

532 However, should the development of a larger training database continue to be prob-
533 lematic, information such as the DEM may be best incorporated through the use of a two
534 input model, in which one set of convolutional filters are applied to the phase information,
535 whilst a second is applied to the DEM. These two networks could then be merged at the
536 fully connected stage, in much the same way as our fully connected model diverges into
537 two outputs. Should this be successful, it may also provide a method to add further inputs
538 to a model, such as those outputted by a weather model, which may reduce false positives
539 due to occurrences such as a strong topographically correlated APS. However, training the
540 weights of a model from scratch and exploring more complex multi-input model architec-
541 tures remains beyond the remit of this study.

542 The results presented in Figure 8 show that a model trained only with synthetic data
543 is able to classify and locate deformation signals in Sentinel-1 data. However, it is only
544 successful in cases with particularly clear deformation patterns, and in cases with more
545 subtle signals generally erroneously resorts to labelling these as not containing deforma-
546 tion. Additionally, strong atmospheric signals are often misclassified as deformation. It is
547 possible that these limitations may be overcome through the use of more realistic synthetic
548 data, as our result suggests that our current methodology does not describe processes well
549 enough to be used without real data. The generation of more realistic deformation patterns
550 may be achieved through steps such as more intelligent sampling of the parameters used
551 in the forward models used to generate the deformation patterns, the use of different types
552 of deformation models such as penny-shaped cracks [Fialko *et al.*, 2001] or point/Mogi
553 sources [Mogi, 1958], and the superposition of multiple deformation patterns in a single
554 interferogram such as was observed prior to the 2005 eruption of Sierra Negra [Jónsson,
555 2009]. The generation of more realistic atmospheric signals could be achieved through
556 increasing the complexity of synthetic data, such as through the use of phase-elevation ra-
557 tios that are non-linear or spatially variable, or through using data from different sources.
558 Interferograms that image regions with little deformation could be used to increase the
559 complexity of the set of “no deformation” data, or combined with synthetic deformation
560 patterns to produce more complex semi-synthetic data.

561 The results presented in Figure 9 show the benefit of incorporating real data. How-
562 ever, much scope for improvement remains, with two classification and several localisa-
563 tion errors visible in this figure. The classification error at Vesuvius (interferogram zero)
564 relates to the subtle subsidence signal located near the summit of the volcano, and is so
565 unclear that a human expert would be unlikely to label this as deformation without further
566 analysis (e.g. inspecting the complete time series). The classification error at Wolf (inter-
567 ferogram 11) is more complex, and on inspection suggests that the error is caused by the
568 incorrect labelling of the interferogram during its construction from the VolcNet database,
569 and that our CNN is actually correct. This is likely to be a consequence of how we de-
570 termine a minimum threshold of deformation in an interferogram for it to be labelled as
571 “deformation”, and serves to illustrate a potential disadvantage of our automatic labelling
572 approach. The majority of the localisation errors are in the form of inaccuracy relating to
573 the centre of the deformation or its spatial size, and from this we conclude that our local-
574 isation head may not be complex enough to capture the large variety possible in both the
575 location and spatial extend of a signal. Further refinement of this part of the model lies
576 outside the scope of this paper, but is not likely to be addressed through incremental im-
577 provements to the fully connected head, but rather through complete replacement with a
578 more complex model such as R-CNN [Girshick *et al.*, 2013]. The training of more com-
579 plex models is likely to require more real data from the VolcNet database, which may be
580 addressed through incorporating more time series, and by addressing the large disparity
581 in the number of data per class (i.e. the scarcity of dykes) that limits the number of other
582 interferograms that we are able to use in this study.

583 The divergent nature of the two heads (classification and localisation) of our net-
584 work also allows for discrepancies between their outputs. This is seen in interferogram
585 10 of Figure 9, in which the localisation head produces a broadly correct output, but the
586 signal is incorrectly labelled as “no deformation”, although with a relatively low confi-
587 dence. However, we postulate that it may be possible to avoid errors of this type by using
588 more complex model architectures. Models such as YOLO [Redmon *et al.*, 2016] produce
589 bounding boxes and classifications in one step, and have the added bonus of being able to
590 work with images that contain multiple signals. If successfully applied to interferograms,
591 a model of this complexity may avoid the discrepancy errors we encounter, and be able to
592 handle interferograms that contain multiple deformation patterns. In the case that multiple
593 signals do exist in a single interferogram, we do not envisage these to be difficult to label

594 as it is likely that these would be considered interesting events by the scientific community
595 and therefore be the subject of detailed study (e.g. the multi-signal interferograms used in
596 this study are analysed in detail in [Xu *et al.*, 2016]).

597 Our approach to localisation avoids the need for repeated classification using a slid-
598 ing window approach, and allows for our network to reason using the entire image. Whilst
599 this approach is beneficial in terms of advancing the state-of-the-art towards that of a hu-
600 man interpreter, one caveat remains in that building a network that is able to utilise large
601 interferograms can be complex. In our model, we use pixels of three arc second size and,
602 with an input size of 224×224 , the resulting model is able to “see” an approximately
603 20km square around a volcano. If we wish to proceed at this resolution, our model’s vi-
604 sual field could be increased through changing the input size to around 400×400 which
605 would not impact our ability to use VGG16’s filters (or convolutional blocks), but would
606 increase the size of the first layer of the fully connected part of our network.

607 At present, an input with side length 224 is reduced to a feature map with side length
608 7 (shown in Figure 4) which, combined with a depth of 512, produces a flattened layer of
609 size $7 \times 7 \times 512 = 25088$. However, doubling the input side length would double the fea-
610 ture map side length, increasing the flattened layer to a size of $14 \times 14 \times 512 = 100352$.
611 Whilst our model contains millions of free parameters, connecting this layer to a subse-
612 quent layer would produce a significant increase in the total, and is likely to require either
613 more ingenuity or more data to be trained successfully. Analysis of the offsets of defor-
614 mation patterns at volcanic centres by *Ebmeier et al.* [2018] finds that 8% of signals are
615 located more than 10km from a volcanic edifice, and would therefore be missed by our
616 current model. Future models that wish to perform localisation using a global approach
617 may therefore require slight increases in size in order to capture all signals of interest. Al-
618 ternatively, as per the approach of [Anantrasirichai *et al.*, 2018], CNNs can themselves be
619 convolved across larger images (such as those routinely captured by the TOPSAR mode
620 of the Sentinel-1 satellites) to create repeat classifications, and this may provide a way to
621 apply our current model to images for which the number of parameters in the first layer
622 of the fully connected network is prohibitive. However, for application to large scenes that
623 capture non-volcanic deformation, a network similar to the fully convolutional network
624 (FCN) presented in *Rouet-Leduc et al.* [2020] may be more suitable, as this contains no
625 fully connected network and so can be applied to an input of near arbitrary size.

626 The smallest deformation signals that our model can accurately label are of approx-
627 imately 5 cm in magnitude, which is a product of us choosing this threshold as the min-
628 imum deformation required for a VolcNet interferogram to be considered as containing
629 deformation. A benefit of our novel labelling approach is that through decreasing this pa-
630 rameter, we can produce single interferograms that span persistent deformation that is not
631 visible to the human observer (e.g. 1 cm of deformation is not likely to be visible through
632 the atmospheric noise of several to tens of centimetres commonly encountered at volca-
633 noes). Relabelling the VolcNet database could therefore be done at increasingly lower de-
634 formation thresholds, and provide a route to train deep learning models that outperform
635 human domain experts. Through computing cumulative displacements in the manner de-
636 scribed in *Anantrasirichai et al.* [2019b], our existing method could also be extended to
637 extremely low rate signals, providing a long enough time series is present.

638 In addition to making our VolcNet database available via GitHub, we make all the
639 code for training our two deep learning models (VUDL-NET-21: “Volcanic Unrest De-
640 tection and Localisation NET, 2021”) available on GitHub: ([https://github.com/matthew-
641 gaddes/VUDLNet_21](https://github.com/matthew-gaddes/VUDLNet_21))

642 **5 Conclusion**

643 We find that either wrapped or unwrapped data are approximately equally suited for
644 use with the weights of VGG16’s filters trained on ImageNet data. We also find that in-
645 corporating extra information that a human interpreter may use (such as a DEM) in the
646 two otherwise unused channels of a model trained in this way acts to degrade model per-
647 formance, and we postulate that this is a result of the disparate nature of the signals con-
648 tained within a DEM and the phase of an interferogram. However, we expect this will not
649 be the case if the weights within VGG16’s filters are trained from scratch, as additional
650 data such as a DEM should help to separate deformation from noise.

651 We combine the five convolutional blocks of VGG16 with two fully connected net-
652 works to perform both classification and localisation of deformation, which allows our net-
653 work to reason using the whole interferogram (i.e. avoiding a sliding window approach),
654 and therefore move a step closer to interpreting InSAR data in a manner similar to a hu-
655 man expert. Additionally, our network is able to differentiate between several different
656 forms of deformation, and advances the state-of-the-art. We expect that further work may

657 build on the results presented in this manuscript and use the same method to increase
658 the number of deformation signals that a model is able to identify. For use with volcano
659 monitoring, this may include models that are able to classify signals such as those due
660 to cooling lava flows, or those due to unstable volcano flanks. For use in the broader re-
661 mote sensing community, this three class model could be adapted to perform tasks such as
662 differentiating between strike-slip, thrust, and normal fault earthquakes in single interfero-
663 grams.

664 As Sentinel-1 interferograms are being automatically created for the majority of the
665 world's subaerial volcanoes every 6 or 12 days, our algorithm provides a method to search
666 through this vast and regularly changing database to search for signs of deformation that
667 may indicate that a volcano has entered a period of unrest. Through doing this, the algo-
668 rithm could facilitate monitoring of many currently unmonitored volcanoes. Additionally,
669 as our model is able to localise any deformation it does encounter, this allows the model
670 to determine the spatial extent of a signal (i.e. the area of the bounding box it creates),
671 and so provide information that is likely to be useful when determining how urgently in-
672 terferograms that it flags should be inspected by a human expert.

673 To minimise the costly nature of labelling data, we initially train our model using
674 only synthetic data. We find that our model generalises well to some cases of Sentinel-1
675 data, but errors remain in cases such as subtle deformation signals, or unusual atmospheric
676 signals. We alleviate this problem through building a database of Sentinel-1 data, which
677 we term VolcNet, that uses a novel labelling strategy to create ~500,000 labelled inter-
678 ferograms from relatively few labels. The inclusion of a small amount of this real data
679 during the training phase improves model performance drastically, and we present a model
680 that is able to both classify and locate deformation within interferograms of ~20km side
681 length. For other practitioners seeking to train similar models, we make both our code for
682 generating synthetic interferograms (SyInterferoPy), our labelled Sentinel-1 data (VolcNet),
683 and our two models (VUDL-NET21) freely available via GitHub.

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 692 study from the Alaska Satellite Facility archive (<https://www.asf.alaska.edu/>). Figures were
 693 prepared in Matplotlib [Hunter, 2007], and all CNN work was carried out in KERAS us-
 694 ing the TensorFlow backend.

695 References

- 696 Agostinelli, F., M. Hoffman, P. Sadowski, and P. Baldi (2014), Learning activation func-
 697 tions to improve deep neural networks, *arXiv preprint arXiv:1412.6830*.
- 698 Albino, F., J. Biggs, and D. K. Syahbana (2019), Dyke intrusion between neighbouring arc
 699 volcanoes responsible for 2017 pre-eruptive seismic swarm at agung, *Nature communi-*
 700 *cations*, *10*(1), 1–11.
- 701 Albino, F., J. Biggs, C. Yu, and Z. Li (2020), Automated methods for detecting vol-
 702 canic deformation using sentinel-1 insar time series illustrated by the 2017–2018
 703 unrest at agung, indonesia, *Journal of Geophysical Research: Solid Earth*, *125*(2),
 704 e2019JB017,908.
- 705 Anantrasirichai, N., J. Biggs, F. Albino, P. Hill, and D. Bull (2018), Application of
 706 Machine Learning to Classification of Volcanic Deformation in Routinely Gener-
 707 ated InSAR Data, *Journal of Geophysical Research : Solid Earth*, pp. 1–15, doi:
 708 10.1029/2018JB015911.
- 709 Anantrasirichai, N., J. Biggs, F. Albino, and D. Bull (2019a), A deep learning approach to
 710 detecting volcano deformation from satellite imagery using synthetic datasets, *RSE*, pp.
 711 1–15, doi:10.1029/2018JB015911.
- 712 Anantrasirichai, N., J. Biggs, F. Albino, and D. Bull (2019b), The application of convolu-
 713 tional neural networks to detect slow, sustained deformation in insar time series, *Geo-*
 714 *physical Research Letters*, *46*(21), 11,850–11,858.
- 715 Beauducel, F., P. Briole, and J.-L. Froger (2000), Volcano-wide fringes in ers synthetic
 716 aperture radar interferograms of etna (1992–1998): Deformation or tropospheric effect?,
 717 *Journal of Geophysical Research: Solid Earth*, *105*(B7), 16,391–16,402.
- 718 Biggs, J., S. K. Ebmeier, W. P. Aspinall, Z. Lu, M. E. Pritchard, R. S. J. Sparks, and T. a.
 719 Mather (2014), Global link between deformation and volcanic eruption quantified by
 720 satellite imagery., *Nature communications*, *5*, 3471, doi:10.1038/ncomms4471.

- 721 Bountos, N. I., I. Papoutsis, D. Michail, A. Karavias, P. Elias, and I. Parcharidis (2022),
 722 Hephaestus: A large scale multitask dataset towards insar understanding, in *Proceedings*
 723 *of the IEEE/CVF Conference on Computer Vision and Pattern Recognition*, pp. 1453–
 724 1462.
- 725 Bridle, J. S. (1990), Probabilistic interpretation of feedforward classification network out-
 726 puts, with relationships to statistical pattern recognition, in *Neurocomputing*, pp. 227–
 727 236, Springer.
- 728 Chen, C. W., and H. A. Zebker (2001), Two-dimensional phase unwrapping with use of
 729 statistical models for cost functions in nonlinear optimization, *Journal of the Optical*
 730 *Society of America*, 18(2), 338–351.
- 731 Chollet, F. (2017), *Deep learning with Python*, Simon and Schuster.
- 732 Deng, J., W. Dong, R. Socher, L.-J. Li, K. Li, and L. Fei-Fei (2009), Imagenet: A large-
 733 scale hierarchical image database, in *Computer Vision and Pattern Recognition, 2009.*
 734 *CVPR 2009. IEEE Conference on*, pp. 248–255, Ieee.
- 735 Dozat, T. (2016), Incorporating nesterov momentum into adam.
- 736 Ebmeier, S. K., B. J. Andrews, M. C. Araya, D. W. D. Arnold, J. Biggs, C. Cooper,
 737 E. Cottrell, M. Furtney, J. Hickey, J. Jay, R. Lloyd, A. L. Parker, M. E. Pritchard,
 738 E. Robertson, E. Venzke, and J. L. Williamson (2018), Synthesis of global satellite ob-
 739 servations of magmatic and volcanic deformation : implications for volcano monitoring
 740 & the lateral extent of magmatic domains, *Journal of Applied Volcanology*, pp. 1–26,
 741 doi:10.1186/s13617-018-0071-3.
- 742 Farr, T., P. Rosen, E. Caro, and R. Crippen (2007), The Shuttle Radar Topography Mis-
 743 sion, *Reviews of Geophysics*, 45(2005), 1–33, doi:10.1029/2005RG000183.1.
- 744 Fialko, Y., Y. Khazan, and M. Simons (2001), Deformation due to a pressurized horizontal
 745 circular craze in an elastic half-space, with application to volcano geodesy, *Geophys. J.*
 746 *Int.*, 146, 181–190.
- 747 Gaddes, M., A. Hooper, M. Bagnardi, H. Inman, and F. Albino (2018), Blind signal sep-
 748 aration methods for insar: The potential to automatically detect and monitor signals of
 749 volcanic deformation, *Journal of Geophysical Research: Solid Earth*.
- 750 Galetto, F., M. Bagnardi, V. Acocella, and A. Hooper (2019), Noneruptive unrest at the
 751 caldera of alcedo volcano (galápagos islands) revealed by insar data and geodetic mod-
 752 eling, *Journal of Geophysical Research: Solid Earth*, 124(4), 3365–3381.

- 753 Girshick, R. B., J. Donahue, T. Darrell, and J. Malik (2013), Rich feature hierarchies for
754 accurate object detection and semantic segmentation, *CoRR*, *abs/1311.2524*.
- 755 Goldstein, R. M., and C. L. Werner (1998), Radar interferogram filtering for geophysical
756 applications, *Geophysical Research Letters*, *25*(21), 4035, doi:10.1029/1998GL900033.
- 757 Hanssen, R. F. (2001), *Radar interferometry: data interpretation and error analysis*, vol. 2,
758 Springer Science & Business Media.
- 759 He, K., X. Zhang, S. Ren, and J. Sun (2016), Deep residual learning for image recogni-
760 tion, in *Proceedings of the IEEE conference on computer vision and pattern recognition*,
761 pp. 770–778.
- 762 He, K., G. Gkioxari, P. Dollár, and R. Girshick (2017), Mask r-cnn, in *Proceedings of the*
763 *IEEE international conference on computer vision*, pp. 2961–2969.
- 764 Hunter, J. D. (2007), Matplotlib: A 2d graphics environment, *Computing In Science &*
765 *Engineering*, *9*(3), 90–95.
- 766 Jónsson, S. (2009), Stress interaction between magma accumulation and trapdoor fault-
767 ing on Sierra Negra volcano, Galápagos, *Tectonophysics*, *471*(1-2), 36–44, doi:
768 10.1016/j.tecto.2008.08.005.
- 769 Kingma, D. P., and J. Ba (2014), Adam: A method for stochastic optimization, *arXiv*
770 *preprint arXiv:1412.6980*.
- 771 Krizhevsky, A., I. Sutskever, and G. E. Hinton (2012), Imagenet classification with deep
772 convolutional neural networks, in *Advances in neural information processing systems*, pp.
773 1097–1105.
- 774 Lazecký, M., K. Spaans, P. J. González, Y. Maghsoudi, Y. Morishita, F. Albino, J. Elliott,
775 N. Greenall, E. Hatton, A. Hooper, et al. (2020), Licsar: An automatic insar tool for
776 measuring and monitoring tectonic and volcanic activity, *Remote Sensing*, *12*(15), 2430.
- 777 Mogi, K. (1958), Relations between the eruptions of various volcanoes and the deforma-
778 tions of the ground surfaces around them, *Bulletin of the Earthquake Research Institute*,
779 *36*, 99–134, doi:10.1016/j.epsl.2004.04.016.
- 780 Moore, C., T. Wright, A. Hooper, and J. Biggs (2019), The 2017 eruption of erta’ale vol-
781 cano, ethiopia: Insights into the shallow axial plumbing system of an incipient mid-
782 ocean ridge, *Geochemistry, Geophysics, Geosystems*, *20*(12), 5727–5743.
- 783 Morishita, Y., M. Lazecký, T. J. Wright, J. R. Weiss, J. R. Elliott, and A. Hooper (2020),
784 Licsbas: an open-source insar time series analysis package integrated with the licsar
785 automated sentinel-1 insar processor, *Remote Sensing*, *12*(3), 424.

- 786 Novellis, V. D., R. Castaldo, C. D. Luca, S. Pepe, I. Zinno, F. Casu, R. Lanari, and
787 G. Solaro (2017), Source modelling of the 2015 Wolf volcano (Galápagos) eruption
788 inferred from Sentinel 1-A DInSAR deformation maps and pre-eruptive ENVISAT
789 time series, *Journal of Volcanology and Geothermal Research*, 344, 246–256, doi:
790 10.1016/j.jvolgeores.2017.05.013.
- 791 Okada (1985), Surface deformation due to shear and tensile faults in a half-space, *Internation-*
792 *ational Journal of Rock Mechanics and Mining Sciences Geomechanics Abstracts*, 75(4),
793 1135–1154, doi:10.1016/0148-9062(86)90674-1.
- 794 Redmon, J., S. Divvala, R. Girshick, and A. Farhadi (2016), You only look once: Unified,
795 real-time object detection, in *Proceedings of the IEEE conference on computer vision and*
796 *pattern recognition*, pp. 779–788.
- 797 Rémy, D., Y. Chen, J.-L. Froger, S. Bonvalot, L. Cordoba, and J. Fustos (2015), Revised
798 interpretation of recent insar signals observed at llaima volcano (chile), *Geophysical*
799 *Research Letters*, 42(10), 3870–3879.
- 800 Ronneberger, O., P. Fischer, and T. Brox (2015), U-net: Convolutional networks for
801 biomedical image segmentation, in *International Conference on Medical image comput-*
802 *ing and computer-assisted intervention*, pp. 234–241, Springer.
- 803 Rouet-Leduc, B., R. Jolivet, M. Dalaison, P. A. Johnson, and C. Hulbert (2020), Au-
804 tonomous extraction of millimeter-scale deformation in insar time series using deep
805 learning, *arXiv preprint arXiv:2012.13849*.
- 806 Simonyan, K., and A. Zisserman (2014), Very deep convolutional networks for large-scale
807 image recognition, *arXiv preprint arXiv:1409.1556*.
- 808 Srivastava, N., G. Hinton, A. Krizhevsky, I. Sutskever, and R. Salakhutdinov (2014),
809 Dropout: a simple way to prevent neural networks from overfitting, *The Journal of Ma-*
810 *chine Learning Research*, 15(1), 1929–1958.
- 811 Sumbul, G., M. Charfuelan, B. Demir, and V. Markl (2019), Bigearthnet: A large-
812 scale benchmark archive for remote sensing image understanding, *arXiv preprint*
813 *arXiv:1902.06148*.
- 814 Sun, J., C. Wauthier, K. Stephens, M. Gervais, G. Cervone, P. La Femina, and M. Hig-
815 gins (2020), Automatic detection of volcanic surface deformation using deep learning,
816 *Journal of Geophysical Research: Solid Earth*, 125(9), e2020JB019,840.
- 817 Szegedy, C., W. Liu, Y. Jia, P. Sermanet, S. Reed, D. Anguelov, D. Erhan, V. Vanhoucke,
818 and A. Rabinovich (2015), Going deeper with convolutions, in *Proceedings of the IEEE*

- 819 *conference on computer vision and pattern recognition*, pp. 1–9.
- 820 Valade, S., A. Ley, F. Massimetti, O. DâĂŽHondt, M. Laiolo, D. Coppola, D. Loibl,
821 O. Hellwich, and T. R. Walter (2019), Towards global volcano monitoring using mul-
822 tisensor sentinel missions and artificial intelligence: The mounts monitoring system,
823 *Remote Sensing*, *11*(13), 1528.
- 824 Xu, W., S. Jónsson, J. Ruch, and Y. Aoki (2016), The 2015 Wolf volcano (Galápagos)
825 eruption studied using Sentinel-1 and ALOS-2 data, *Geophysical Research Letters*, pp.
826 9573–9580, doi:10.1002/2016GL069820.
- 827 Yip, S. T. H., J. Biggs, and F. Albino (2019), Reevaluating volcanic deformation using
828 atmospheric corrections: Implications for the magmatic system of agung volcano, in-
829 donesia, *Geophysical Research Letters*, *46*(23), 13,704–13,711.
- 830 Yu, D., and M. L. Seltzer (2011), Improved bottleneck features using pretrained deep neu-
831 ral networks, in *Twelfth annual conference of the international speech communication*
832 *association*.
- 833 Zhang, Q., Y. Wang, Q. Liu, X. Liu, and W. Wang (2016), Cnn based suburban building
834 detection using monocular high resolution google earth images, in *2016 IEEE Interna-*
835 *tional Geoscience and Remote Sensing Symposium (IGARSS)*, pp. 661–664, IEEE.