

A record of plume-induced plate rotation triggering subduction initiation

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23 **The formation of a global network of plate boundaries surrounding a mosaic of**
24 **lithospheric fragments was a key step in the emergence of Earth's plate tectonics. So far,**
25 **propositions for plate boundary formation are regional in nature; how plate boundaries**
26 **are created over thousands of kilometers in geologically short periods remains elusive.**
27 **Here we show from geological observations that a >12,000 km-long plate boundary formed**
28 **between the Indian and African plates around 105 Ma. This boundary comprised**
29 **subduction segments from the eastern Mediterranean region to a newly established India-**
30 **Africa rotation pole in the west Indian ocean, where it transitioned into a ridge between**
31 **India and Madagascar. We identify coeval mantle plume rise below Madagascar-India as**
32 **the only viable trigger of this plate rotation. For this, we provide a proof of concept by**
33 **torque balance modeling, which reveals that the Indian and African cratonic keels were**
34 **important in determining plate rotation and subduction initiation in response to the**
35 **spreading plume head. Our results show that plumes may provide a non-plate-tectonic**
36 **mechanism for large plate rotation, initiating divergent and convergent plate boundaries**
37 **far away from the plume head. We suggest that this mechanism may be an underlying**
38 **cause of the emergence of modern plate tectonics.**

39 The early establishment of plate tectonics on Earth was likely a gradual process that
40 evolved as the cooling planet's lithosphere broke into a mosaic of major fragments, separated by
41 a network of plate boundaries: spreading ridges, transform faults, and subduction zones¹. The
42 formation of spreading ridges and connecting transform faults is regarded as a passive process,
43 occasionally associated with rising mantle plumes². The formation of subduction zones is less
44 well understood. Explanations for subduction initiation often infer spontaneous gravitational
45 collapse of aging oceanic lithosphere², or relocations of subduction zones due to intraplate stress
46 changes in response to arrival of continents, oceanic plateaus, or volcanic arcs in trenches³.
47 Mantle plumes have also been suggested as drivers for regional subduction initiation, primarily
48 based on numerical modeling⁴⁻⁶. But while such processes may explain how plate tectonics
49 evolves on a regional scale, they do not provide insight into the geodynamic cause(s) for the
50 geologically sudden (<10 My) creation of often long (>1000 km) plate boundaries including new
51 subduction zones⁷. Demonstrating the causes of plate boundary formation involving subduction
52 initiation using the geological record is challenging and requires (i) establishing whether
53 subduction initiation was spontaneous or induced; (ii) if induced, constraining the timing and

54 direction of incipient plate convergence; (iii) reconstructing the entire plate boundary from triple
55 junction to triple junction, as well as the boundaries of neighboring plates, to identify collisions,
56 subduction terminations, or mantle plume arrival that may have caused stress changes driving
57 subduction initiation. In this paper, we provide such an analysis for an intra-oceanic subduction
58 zone that formed within the Neotethys Ocean around 105 Ma ago, to evaluate the driver of
59 subduction initiation and plate boundary formation.

60

61 **Induced subduction initiation across the Neotethys Ocean**

62 During induced subduction initiation, lower plate burial, dated through prograde mineral
63 growth in rocks of the incipient subduction plate contact, in so-called metamorphic soles⁸
64 predates upper plate extension that is inferred from spreading records in so-called supra-
65 subduction zone (SSZ) ophiolites^{8-10,11}. Such SSZ ophiolites have a chemical stratigraphy widely
66 interpreted as having formed at spreading ridges above a nascent subduction zone. Several SSZ
67 ophiolite belts exist in the Alpine-Himalayan mountain belt, which formed during the closure of
68 the Neotethys Ocean^{12,13} (Fig. 1A). One of these ophiolite belts formed in Cretaceous time and
69 runs from the eastern Mediterranean region, along northern Arabia, to Pakistan. Incipient lower
70 plate burial has been dated through Lu/Hf prograde garnet growth ages of ~104 Ma in
71 metamorphic soles in Oman as well as in the eastern Mediterranean region^{8,14}. Upper plate
72 extension and SSZ ophiolite spreading has been dated using magmatic zircon U/Pb ages and
73 synchronous metamorphic sole ⁴⁰Ar/³⁹Ar cooling ages and occurred at 96-95 Ma (Pakistan,
74 Oman)^{15,16} to 92-90 Ma (Iran, eastern Mediterranean region)¹⁷. The 8-14 Myr time delay
75 between initial lower plate burial and upper plate extension demonstrates that subduction
76 initiation was induced⁸.

77 An initial ~E-W convergence direction at this subduction zone was constrained through
78 paleomagnetic analysis and detailed kinematic reconstruction of post-subduction initiation
79 deformation of the eastern Mediterranean region, Oman, and Pakistan, and was accommodated at
80 ~N-S striking trench segments^{13,18-20}. This is surprising: for hundreds of Ma and throughout the
81 Tethyan realm, rifts and ridges accommodated the separation of continental fragments off
82 northern Gondwana in the south and their accretion to the southern Eurasian margin at
83 subduction zones in the north^{21,22}. The ~E-W convergence that triggered ~105 Ma subduction

84 initiation across the Neotethys ocean was thus near orthogonal to the long-standing plate
85 motions. To find the trigger inducing this subduction, we developed the first comprehensive
86 reconstruction of the entire ~12,000 km long plate boundary that formed at ~105 Ma and placed
87 this in context of reconstructions of collisions and mantle plumes of the Neotethyan realm (Fig.
88 1).

89

90 **Geological reconstruction of incipient plate boundary**

91 The SSZ ophiolites that formed at the juvenile Cretaceous intra-Neotethyan subduction
92 zone are now found as klippen on intensely deformed accretionary orogenic belts (Fig. 1A) that
93 formed when the continents of Greater Adria, Arabia, and India arrived in subduction zones. We
94 reconstructed these orogenic belts (Fig. 1) and restored these continents, and the Cretaceous
95 ophiolites that were thrust upon these, into their configuration at 105 Ma (Fig. 1C) (see
96 Methods).

97 The westernmost geological record of the Cretaceous intra-Neotethyan subduction zone
98 is found in eastern Greece and western Turkey, where it ended in a trench-trench-trench triple
99 junction with subduction zones along the southern Eurasian margin¹⁸. From there, east-dipping
100 (in the west) and west-dipping (in the east) subduction segments followed the saw-toothed shape
101 of the Greater Adriatic and Arabian continental margins (Fig. 1C) and initiated close to it: rocks
102 of these continental margins already underthrust the ophiolites within 5-15 My after SSZ
103 ophiolite spreading^{14,23,24}, and continent-derived zircons have been found in metamorphic sole
104 rocks²⁵. Subduction segments likely nucleated along ancient N-S and NE-SW trending fracture
105 zones and linked through highly oblique, north-dipping subduction zones that trended parallel to
106 and likely reactivated the pre-existing (hyper)extended passive margins (Fig. 1B, C)^{20,23}.
107 Subducted remnants of the Cretaceous intra-Neotethyan subduction are well-resolved in the
108 present-day mantle as slabs in the mid-mantle below the southeastern Mediterranean Sea, central
109 Arabia and the west Indian Ocean²⁶.

110 East of Arabia, we trace the intra-oceanic plate boundary to a NE-SW striking, NW-
111 dipping subduction zone between the Kabul Block and the west Indian passive margin. The 96
112 Ma Waziristan ophiolites of Pakistan formed above this subduction zone, perhaps by inverting an
113 Early Cretaceous spreading ridge between the Kabul Block and India¹³ and were thrust eastward

114 onto the Indian margin^{13,16} (Fig. 1B, C). The Cretaceous intra-Neotethyan plate boundary may
115 have been convergent to the Amirante Ridge in the west Indian Ocean¹³, from where it became
116 extensional instead and developed a rift, and later a spreading ridge, in the Mascarene Basin that
117 accommodated separation of India from Madagascar^{13,27,28} (Fig. 1B). The plate boundary ended
118 in a ridge-ridge-ridge triple junction in the south Indian Ocean^{13,28} (Fig. 1B).

119 The newly formed Cretaceous plate boundary essentially temporarily merged a large part
120 of Neotethyan oceanic lithosphere between Arabia and Eurasia to the Indian plate. This plate was
121 >12,000 km long from triple junction to triple junction, and reached from 45°S to 45°N, with
122 4500 km of rift/ridge in the southeast and 7500 km of subduction zone in the northwest and with
123 a transition between the convergent and divergent segments, representing the India-Africa Euler
124 pole¹³, in the west Indian Ocean, at a latitude between Pakistan and the Amirante Ridge (Fig.
125 1B). Marine geophysical constraints show a ~4° counterclockwise rotation of India relative to
126 Africa about the west Indian Ocean Euler pole during rifting preceding the ~83 Ma onset of
127 oceanic spreading in the Mascarene Basin²⁷⁻²⁹, associated with up to hundreds of km of ~E-W
128 convergence across the Neotethys (Fig. 1D).

129 The neighboring plates of the intra-Neotethyan subduction zone at 105 Ma were thus
130 Africa and India. The African plate was mostly surrounded by ridges and had a complex
131 subduction plate boundary in the Mediterranean region³⁰. The Indian plate was surrounded by
132 ridge-transform systems in the south and east and by subduction in the north, and may have
133 contained rifts and ridges between the Indian continent and Eurasia^{13,28}. The Neotethys
134 lithosphere between Arabia-Greater Adria and Eurasia continued unbroken to the north-dipping
135 subduction zone that had already existed along the southern Eurasian margin since the
136 Jurassic^{31,32}: the spreading ridges that existed during Neotethys Ocean opening in the Permian-
137 Triassic (north of Arabia)³³, and Triassic-Jurassic (eastern Mediterranean region)²³ had already
138 subducted below Eurasia before 105 Ma^{19,33} (Fig. 1B, C).

139

140 **Identifying potential drivers of subduction initiation**

141 Candidate processes to trigger the reconstructed plate boundary formation at 105 Ma are
142 terminations of existing subduction zones by arrival of buoyant lithosphere or the rise of mantle
143 plumes. Southern Eurasia contains relics of many microcontinents that accreted at or clogged

144 subduction zones since the Paleozoic, but none of these events started or ended around 105
145 Ma^{13,21-23,33-35}. Continental subduction and collision was ongoing in the central Mediterranean
146 region²³, but it is not evident how this or any other changes in subduction dynamics along the E-
147 W trending southern Eurasian margin would lead to E-W convergence in the Neotethys Ocean.
148 In the eastern Neotethys, a mid-Cretaceous collision of the intra-oceanic Woyla Arc with the
149 Sundaland continental margin led to a subduction polarity reversal initiating eastward subduction
150 below Sundaland³⁶, which is recorded in ophiolites on the Andaman Islands. There, metamorphic
151 sole rocks with ⁴⁰Ar/³⁹Ar hornblende cooling ages of 105-106 Ma, and likely coeval SSZ
152 ophiolite spreading ages³⁷ reveal that this subduction zone may have developed slab pull around
153 the same time as the Indian Ocean-western Neotethys plate boundary formed (Fig 1C). However,
154 eastward slab pull below Sundaland cannot drive E-W convergence in the Neotethys to the west,
155 and Andaman SSZ extension may well be an expression rather than the trigger of Indian plate
156 rotation. We find no viable plate tectonics-related driver of the ~105 Ma plate boundary
157 formation that we reconstructed here.

158 A key role, however, is possible for the only remaining geodynamic, non-plate-tectonic,
159 plate-motion driver in the region: a mantle plume. India-Madagascar continental breakup is
160 widely viewed^{13,27,37} as related to the ~94 Ma and younger formation of the Morondava Large
161 Igneous Province (LIP) on Madagascar³⁸ and southwest India³⁹. This LIP, however, started
162 forming ~10 Ma after initial plate boundary formation. To understand whether the plume may be
163 responsible for both LIP emplacement and plate boundary formation, we explore existing
164 numerical models of plume-plate interaction and conduct explorative torque-balance simulations
165 of plume-lithosphere interaction.

166

167 **Mantle plumes driving subduction initiation**

168 Numerical simulations of plume-lithosphere interaction have already identified that
169 plume head spreading below the lithosphere leads to horizontal asthenospheric flow that exerts a
170 ‘plume push’ force on the base of the lithosphere, particularly in the presence of a cratonic
171 keel^{5,40,41}. Plume push may accelerate plates by several cm/yr⁴¹ and has been proposed as a
172 potential driver of subduction initiation⁵.

173 In many cases, including in the case of the Morondava LIP, LIP eruption and
174 emplacement shortly preceded continental breakup, but pre-break up rifting preceded LIP
175 emplacement by 10-15 Myr²⁷. This early rifting typically is interpreted to indicate that the plume
176 migrated along the base of the lithosphere into a pre-existing rift that formed independently of
177 plume rise²⁷. However, in numerical simulations dynamic uplift⁴² and plume push⁴¹ already start
178 to accelerate plates 10-15 Myr before the plume head reaches the base of the lithosphere and
179 emplaces the LIP. Numerical simulations thus predict the observed delay between plume push,
180 as a driver for early rifting and subduction initiation, and LIP eruption and emplacement.

181 Here, we add to these plume-lithosphere coupling experiments by conducting proof-of-
182 concept torque-balance simulations particularly exploring why the observed India-Africa Euler
183 pole is so close to the plume head such that the associated plate rotation between Africa and
184 India caused E-W convergence in the Neotethys. We performed semi-analytical computations,
185 including both the Indian and African plates at ~105 Ma, and assess the influence of cratonic
186 keels on the position of the India-Africa Euler pole (Fig. 2, see Methods).

187 In our computations without cratonic keels, plume push under Madagascar/India caused
188 counterclockwise rotation of India versus Africa, but about an Euler pole situated far north of
189 Arabia, (Fig. 2A) without inducing significant E-W convergence within the Neotethys. However,
190 in experiments that include keels of the Indian and African cratonic lithosphere, which are
191 strongly coupled to the sub-asthenospheric mantle, the computed Euler pole location is shifted
192 southward towards the Indian continent, inducing E-W convergence along a larger part of the
193 plate boundary within the Neotethys Ocean (Fig. 2B).

194 Convergence of up to several hundreds of km, sufficient to induce self-sustaining
195 subduction²⁷, is obtained if plume material is fed into – and induced flow is confined to – a 200
196 km thick weak asthenospheric layer. The thinner this layer is, the further the plume head spreads,
197 and pushes the plate. The modern Indian cratonic root used in our computations has likely eroded
198 considerably during interaction with the ~70-65 Ma Deccan plume⁴³. India may have had a
199 thicker and/or laterally more extensive cratonic root at ~105 Ma than modeled here which would
200 further enhance coupling of the lithosphere and the sub-asthenospheric mantle. Furthermore, an
201 Euler pole close to India and a long convergent boundary to the north requires much weaker
202 coupling in the northern (oceanic) part of the India plate (Fig. 2). In this case, results remain

203 similar as long as the plume impinges near the southern part of the western boundary of
204 continental India.

205 An order of magnitude estimate of the maximum plume-induced stresses, assuming no
206 frictional resistance at other plate boundaries, is obtained from the rising force of $\sim 1.5 \cdot 10^{20}$ N of
207 a plume head with 1000 km diameter and density contrast 30 kg/m^3 . If half of this force acts on
208 the India plate and with a lever arm of 4000 km, this corresponds to a torque of $3 \cdot 10^{26}$ Nm. Once,
209 at the onset of rifting, ridge push is established as an additional force in the vicinity of the plume,
210 we estimate that this number may increase by up to a few tens of per cent. This torque can be
211 balanced at the convergent boundary (length ~ 5000 km, plate thickness ~ 100 km) involving
212 stresses of ~ 240 MPa, much larger than estimates of frictional resistance between subducting and
213 overriding plates that are only of the order of tens of MPa⁴⁴. For this estimate, we neglect any
214 frictional resistance at the base of the plate and at any other plate boundary – essentially
215 considering the plate as freely rotating above a pinning point. This is another endmember
216 scenario, as opposed to our above convergence estimate, where we had considered friction at the
217 plate base but neglected it at all plate boundaries. Therefore, the estimate of 240 MPa may be
218 considered as an upper bound but being compressive and oriented in the right direction it shows
219 the possibility of subduction initiation as has occurred in reality along the likely weakened
220 passive margin region of Arabia and Greater Adria. Moreover, the plume-induced compressive
221 stresses may have added to pre-existing compressive stresses, in particular due to ridge-push
222 around the African and Indian plates. Such additional compressive stresses may contribute to
223 shifting the Euler pole further south, closer to the position reconstructed in Fig. 1.

224 Subduction became self-sustained ~ 8 -12 Ma after its initiation, as marked by the 96-92
225 Ma age of SSZ spreading^{15,17}: inception of this spreading shows that subduction rates exceeded
226 convergence rates, and reconstructed SSZ spreading rates were an order of magnitude higher¹⁵
227 than Africa-Arabia or Indian absolute plate motions^{41,45} signaling slab roll-back, i.e. self-
228 sustained subduction^{20,46}. Numerical models suggest that self-sustained subduction may start
229 after ~ 50 -100 km of induced convergence⁷, corresponding to $\sim 1^\circ$ of India-Africa rotation
230 between ~ 105 and ~ 96 -92 Ma. Subsequent east and west-dipping subduction segments (Fig. 1)
231 may have contributed to and accelerated the India-Africa/Arabia rotation, driving the
232 propagation of the Euler pole farther to the south (compare Fig. 2A, C).

233

234 **Mantle plumes as an initiator of plate tectonics?**

235 Previously, numerical modeling has shown that mantle plumes may trigger circular
236 subduction initiation around a plume head⁴, where local plume-related convection may drive
237 subduction of thermally weakened lithosphere. This subduction would propagate through slab
238 roll-back and may have started the first subduction features on Earth⁴. 3D convective models do
239 produce a global network of plate boundaries^{47,48} but the role of plumes in initiating new
240 subduction zones within this network is unclear. Here, we have provided the first evidence that
241 plume rise formed a >12,000 km long plate boundary composed of both convergent and
242 divergent segments. Our documented example is Cretaceous in age but geological observations
243 showing a general temporal overlap between LIP emplacement and formation of SSZ ophiolite
244 belts over more than a billion years⁴⁹ suggest that plume rise is a key driving factor in the
245 formation of subduction plate boundaries. Because mantle plumes are thought to be also
246 common features on planets without plate tectonics, such as Mars and Venus⁵⁰, they may have
247 played a vital role in the emergence of modern style plate tectonics on Earth. That plumes may
248 have been key for the evolution of plate tectonics on Earth, as we suggest, but apparently
249 insufficient on Mars and Venus, provides a new outlook on understanding the different planetary
250 evolutions.

251

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448 PJmcP, CGa, ELA and RLMV developed the kinematic reconstruction; BS performed
449 modelling; DJJvH, BS, CGu, WS wrote the paper, all authors made corrections and edits.

450

451 **Competing interests:** All authors declare no competing interests.

452

453 **Fig. 1. Plate kinematic reconstructions of the Neotethys Ocean and surrounding continents.**

454 a) the present-day; b) 70 Ma; c) 105 Ma, corresponding to the timing of intra-Neotethyan
455 subduction initiation and d) 110 Ma, just before intra-Neotethyan subduction initiation. See
456 Methods for the plate reconstruction approach and sources of detailed restorations.

457 Reconstructions show in a mantle reference frame⁴⁵. AR = Amirante Ridge; Emed = Eastern
458 Mediterranean Region; Ir = Iran; LIP = Large Igneous Province; Mad = Madagascar; Mas =
459 Mascarene Basin; Pak = Pakistan, Tur = Turkey; Waz = Waziristan Ophiolite.

460

461 **Fig. 2. Torque balance modeling results of plumes affecting plates similar to India and**
462 **Africa with, and without cratonic keels.** The computed total displacement, induced by the
463 Morondava plume (pink circle) for the restored ~105 Ma plate configuration (Fig. 1c) for plates
464 without (a, b) and with (c, d) African and Indian cratonic keels, in an Africa-fixed (a, c), or
465 mantle reference frame⁴⁵ (b, d) (see Methods). Ten degree grid spacing; locations of plates,
466 lithosphere thickness and the plume are reconstructed in a slab-fitted mantle reference frame⁴⁵.

467

468 **Methods: Kinematic reconstruction** – The kinematic restoration of Neotethyan intra-
469 oceanic subduction was made in GPlates plate reconstruction software (www.gplates.org)⁵¹.
470 First, we systematically restored stable plates using marine geophysical data from the Atlantic
471 and Indian Ocean, and then restored continental margin deformation that occurred following the
472 arrival of continental lithosphere below the oceanic lithosphere preserved as ophiolites. These
473 restorations are based on a systematic reconstruction protocol, based on magnetic anomalies and
474 fracture zones of present-day sea floor and geophysical constraints on pre-drift extension in
475 adjacent passive continental margins²³, followed by kinematic restoration of post-obduction
476 orogenic deformation using structural geological constraints on continental extension, strike-slip
477 deformation, and shortening, and paleomagnetic constraints on vertical axis rotations. We then
478 restored pre-emplacement vertical axis microplate rotations^{52,53}, as well as paleo-orientations of
479 the SSZ spreading ridges at which the ophiolitic crust formed¹⁸⁻²⁰. The reconstruction shown in
480 Fig. 1B compiles kinematic restorations for the eastern Mediterranean region²³, Iran⁵⁴, Oman²⁰,
481 Pakistan¹³, and the Himalaya³⁴. Ophiolites interpreted to be part of the Cretaceous subduction
482 system include the 96-90 Ma, Cretaceous ophiolites exposed in SE Greece, Anatolia, Cyprus,
483 Syria, and Iraq, the Neyriz ophiolite of Iran, the Semail ophiolite in Oman, and the Waziristan-
484 Khost ophiolite in Pakistan and Afghanistan^{15-17,55}. The Jurassic ophiolite belts of northern
485 Turkey and Armenia⁵⁶⁻⁵⁸ and the late Cretaceous (<80 Ma) Kermanshah ophiolite of Iran⁵⁹ are
486 not included and are instead interpreted to have formed along the southern Eurasian margin²³.
487 The Masirah Ophiolite of East Oman⁶⁰ and the uppermost Cretaceous Bela, Muslim Bagh, and
488 Kabul-Altimur ophiolites of Pakistan and Afghanistan^{61,62} are interpreted to reflect oblique latest
489 Cretaceous to Paleogene India-Arabia convergence¹³ and are also unrelated to the event studied

490 here. Restoration of intra-oceanic subduction prior to the arrival of the continental margins used
491 paleomagnetic data from the ophiolites of Oman, Syria, Cyprus, and Turkey that constrain
492 vertical axis rotations, as well as the orientation of sheeted dyke following cooling after
493 intrusion^{18-20,52,53} as proxy for original ridge and intra-oceanic trench orientations. These
494 paleomagnetic data systematically revealed N-S to NW-SE primary sheeted dyke orientations¹⁸⁻
495 ^{20,52,53}. Because the ages of the SSZ ophiolites in the Neotethyan belt do not laterally progress,
496 spreading must have occurred near-orthogonal to the associated trench, which must thus also
497 have been striking N-S to NE-SW, as shown in the reconstruction of Fig. 1.

498 How far the Indian plate continued northwards around 105 Ma is subject to ongoing
499 debate. On the one hand, the northern Indian continental margin has been proposed to have rifted
500 off India sometime in the Cretaceous^{34,63}, but recent paleomagnetic data suggest that this process
501 occurred in the late Cretaceous, well after 100 Ma⁶⁴. Others inferred that the north Indian
502 continent had a passive margin contiguous with oceanic Neotethyan lithosphere since the middle
503 Jurassic or before and continued to a subduction zone below the SSZ ophiolites found in the
504 Himalayan suture zone and the Kohistan arc^{35,65,66}. Sedimentary and paleomagnetic data
505 demonstrate that these ophiolites formed adjacent to the Eurasian margin in the Early
506 Cretaceous⁶⁷, although they may have migrated southward during slab roll-back in the Late
507 Cretaceous³⁵. Recent paleomagnetic data have shown that a subduction zone may have existed
508 within the Neotethys to the west of the Andaman Islands, above which the West Burma Block
509 would have been located (Figure 1)⁶⁸. Our reconstruction of the eastern Neotethys may thus be
510 oversimplified. However, the geological record of the West Burma Block shows that this
511 subduction zone already existed as early as 130 Ma, and E-W trending until well into the
512 Cenozoic⁶⁸, and we see no reason to infer that changes in the eastern Neotethys contributed to
513 the plate boundary formation discussed here. Some have speculated that the West Burma
514 subduction zone would have been connected to a long-lived, equatorial subduction zone within
515 the Neotethys all along the Indian segment that would already have existed in the Early
516 Cretaceous⁶⁹: this scenario remains unconstrained by paleomagnetic data, and is inconsistent
517 with sediment provenance data from the Himalaya and overlying ophiolites³⁵. In summary, the
518 Indian plate around 105 Ma continued far into the Neotethyan realm, and the India-Africa
519 rotation is a likely driver of E-W convergence sparking subduction initiation close to the
520 northern Gondwana margin purported in Figure 1.

521 *Torque balance modeling* – Forces considered here include (i) the push due to plume-
522 induced flow in the asthenosphere and (ii) the drag due to shear flow between the moving plate
523 and a deeper mantle at rest (Fig. S1). In the first case, we disregard any lateral variations. Plume-
524 induced flow is treated as Poiseuille flow, i.e. with parabolic flow profile, in an asthenospheric
525 channel of thickness h_c , radially away from the plume stem. Since at greater distance plume-
526 induced flow will eventually not remain confined to the asthenosphere, we only consider it to a
527 distance 2400 km, in accord with numerical results⁴¹, and consistent with the finding that there is
528 a transition from dominantly pressure-driven Poiseuille flow at shorter wavelengths to
529 dominantly shear-driven Couette flow at length scales approximately exceeding mantle
530 depth^{70,71}. With v_0 the velocity in the center of the channel at a distance d from the plume stem
531 the total volume flux rate is $2/3 \cdot v_0 \cdot 2\pi d \cdot h_c$ (here neglecting the curvature of the Earth surface
532 for simplicity). Its time integral is equal to the volume of the plume head with radius estimated⁷²
533 to be about $r_p=500$ km, with considerable uncertainty. That is, integration is done over a time
534 interval until the entire plume head volume has flown into the asthenospheric channel. Hence the
535 corresponding displacement vector in the center of the channel is

$$\mathbf{x}_{plu} = \int_{\Delta t} v_0 dt \cdot \mathbf{e}_r = \frac{r_p^3}{d \cdot h_c} \cdot \mathbf{e}_r$$

536
537 where \mathbf{e}_r is the unit vector radially away from the plume (red arrows in Extended DataFig. 1).
538 Because of the parabolic flow profile, the vertical displacement gradient at the top of the channel
539 is

$$2 \cdot \frac{\mathbf{x}_{plu}}{0.5 \cdot h_c} = 2 \cdot \int_{\Delta t} v_0 dt \cdot \frac{1}{0.5 \cdot h_c} \cdot \mathbf{e}_r = \frac{4r_p^3}{d \cdot h_c^2} \cdot \mathbf{e}_r.$$

540
541 Viscosity is defined such that the force per area is equal to viscosity times the radial gradient of
542 horizontal velocity. Hence the time integral of torque on the plate is

$$\mathbf{T}_{plu} = \frac{4\eta_0}{h_c} \int_A \mathbf{r} \times \mathbf{x}_{plu} dA = \frac{4\eta_0 r_p^3}{d \cdot h_c^2} \int_A \mathbf{r} \times \mathbf{e}_r dA$$

543

544 where η_0 is viscosity in the channel and \mathbf{r} is the position vector. \mathbf{T}_{plu} is balanced by the time-
 545 integrated torque \mathbf{T}_{pla} of the plate rotating an angle $\boldsymbol{\omega}$ over the underlying mantle. With plate
 546 displacement vectors $\mathbf{x}_{pla} = \boldsymbol{\omega} \times \mathbf{r}$ (black arrows in Fig. S1) we obtain

$$\mathbf{T}_{pla} = -\frac{\eta_0}{h_s} \int_A \mathbf{r} \times \mathbf{x}_{pla} dA = -\frac{\eta_0}{h_s} \int_A \mathbf{r} \times (\boldsymbol{\omega} \times \mathbf{r}) dA$$

547
 548 Here h_s is an effective thickness of the layer over which shearing occurs, which is calculated
 549 below for a stratified viscosity structure, i.e. laterally homogeneous coupling of plate and mantle
 550 and which we will set equal to h_c for simplicity. Specifically, with \mathbf{T}_x being the time-integrated
 551 torque acting on a plate rotating an angle ω_0 around the x-axis

$$\mathbf{T}_x = -\frac{\omega_0 \eta_0}{h_s} \int_A \mathbf{r} \times (\mathbf{e}_x \times \mathbf{r}) dA,$$

552
 553 and \mathbf{T}_y and \mathbf{T}_z defined in analogy, the torque balance equation can be written

$$\mathbf{T}_{plu} = \frac{\omega_x}{\omega_0} \cdot \mathbf{T}_x + \frac{\omega_y}{\omega_0} \cdot \mathbf{T}_y + \frac{\omega_z}{\omega_0} \cdot \mathbf{T}_z$$

554
 555 ω_0 cancels out when \mathbf{T}_x , \mathbf{T}_y and \mathbf{T}_z are inserted. Integrals used to compute these torques only
 556 depend on plate geometry, η_0 cancels out in the torque balance, and we can solve for the rotation
 557 angle vector $\boldsymbol{\omega}$ simply by a 3 x 3 matrix inversion. In the more general case, where we do not set
 558 h_s and h_c equal, $\boldsymbol{\omega}$ is scaled by a factor h_s/h_c .

559 If a plate moves over a mantle where viscosity varies with depth, then the force per area
 560 F/A should be the same at all depths, and the radial gradient of horizontal velocity $dv/dz = F/A \cdot$
 561 $1/\eta(z)$. If we assume that the deep mantle is at rest (i.e. it moves slowly compared to plate
 562 motions), we further find that plate motion is

$$v_0 = \int_{z_0}^{z(\eta_{\max})} \frac{dv}{dz} dz = \frac{F}{A} \int_{z_0}^{z(\eta_{\max})} \frac{1}{\eta(z)} dz =: \frac{F}{A} \frac{h_s}{\eta_0} \quad (1)$$

563
 564 The integration is done from the base of the lithosphere z_0 to the depth where the approximation
 565 of the “mantle at rest” is probably the most closely matched, i.e. we choose the viscosity

566 maximum. The last equality is according to the definition of the effective layer thickness,
567 whereby η_0 is the viscosity just below the lithosphere. Solving this equation for h_s for the
568 viscosity structure in Extended DataFig. 2 and a 100 km thick lithosphere gives $h_s=203.37$ km.

569 The plume location at 27.1°E, 40.4° S, is obtained by rotating the center of the
570 corresponding LIP at 46° E, 26° S and an age 87 Ma (adopted from Doubrovine et al.⁷³) in the
571 slab-fitted mantle reference frame⁴⁵, in which also the plate geometries at 105 Ma are
572 reconstructed.

573 Results for this case (Fig. 2A) show that a plume pushing one part of a plate may induce
574 a rotation of that plate, such that other parts of that plate may move in the opposite direction. A
575 simple analog is a sheet of paper pushed, near its bottom left corner, to the right: Then, near the
576 top left corner, the sheet will move to the left. With two sheets (plates) on either side, local
577 divergence near the bottom (near the plume) may turn into convergence near the top (at the part
578 of the plate boundary furthest away from the plume). The length of that part of the plate
579 boundary, where convergence is induced may increase, if one plate is nearly “pinned” at a hinge
580 point slightly NE of the plume, perhaps due to much stronger coupling between plate and mantle.
581 At the times considered here ~105 My ago, the Indian continent, where coupling was presumably
582 stronger, was in the southern part of the Indian plate, whereas in its north, there was a large
583 oceanic part, with presumably weaker coupling. Hence the geometry was indeed such that
584 convergence could be induced along a longer part of the plate boundary.

585 In the second case, we therefore consider lateral variations in the coupling between plate
586 and mantle, corresponding to variations in lithosphere thickness and/or asthenosphere viscosity,
587 by multiplying the drag force (from the first case) at each location with a resistance factor. This
588 factor is a function of lithosphere thickness reconstructed at 105 Ma. On continents, thickness
589 derived from tomography⁷⁴ with slabs removed⁷⁵ is simply backward-rotated. In the oceans, we
590 use thickness [km] = 10 · (age [Ma] - 105)^{0.5} with ages from present-day Earthbyte age grid
591 version 3.6, i.e. accounting for the younger age and reduced thickness at 105 Ma, besides
592 backward-rotating. To determine the appropriate rotation, the lithosphere (in present-day
593 location) is divided up into India, Africa, Arabia, Somalia and Madagascar (paleo-)plates and
594 respective 105 Ma finite rotations from van der Meer et al.⁴⁵ are applied. For the parts of the
595 reconstructed plates where thickness could not be reconstructed in this way – often, because this

596 part of the plate has been subducted – we first extrapolate thickness up to a distance $\sim 2.3^\circ$, and
597 set the thickness to a default value of 80 km for the remaining part. Reconstructed thickness is
598 shown in Extended DataFig. 4. For the resistance factor as a function of lithosphere thickness we
599 use two models: Firstly, we use a continuous curve (Extended DataFig. 3) according to eq. (1)

$$\frac{F}{A} = \frac{v_0}{\int_{z_0}^{z(\eta_{\max})} \frac{1}{\eta(z)} dz}. \quad (2)$$

600
601 with the mantle viscosity model in Extended DataFig. 2 combined with variable lithosphere
602 thickness z_0 . However, this causes only a minor change in the plate rotations (Extended DataFig.
603 4 compared to Fig. 2B). Hence, we also use a stronger variation, further explained in the caption
604 of Fig 2 and with results shown in Fig. 2C and D.

605

606 **Data availability**

607 GPLates files with reconstructions used to draft Figure 1 are provided at
608 https://figshare.com/articles/dataset/van_Hinsbergen_NatureGeo_2021_GPlates_zip/13516727.

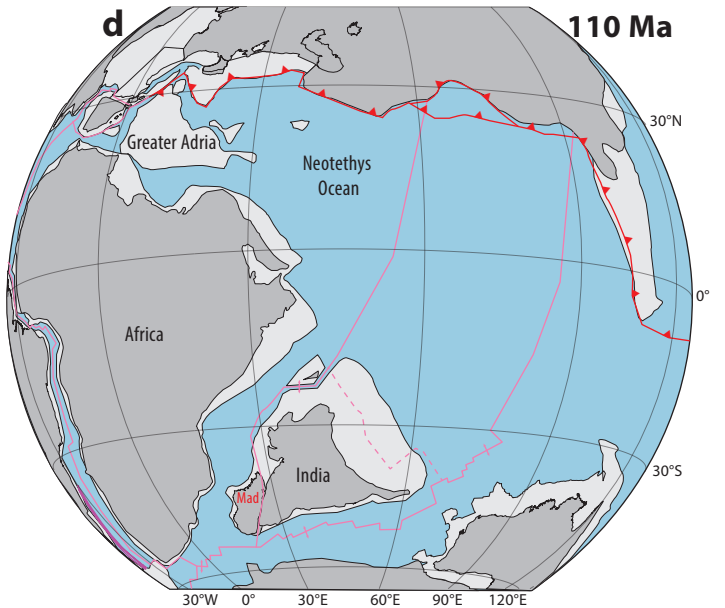
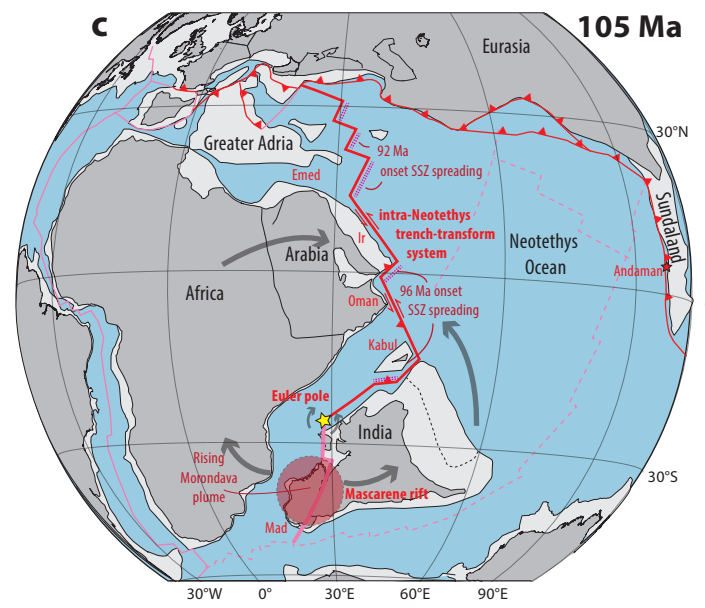
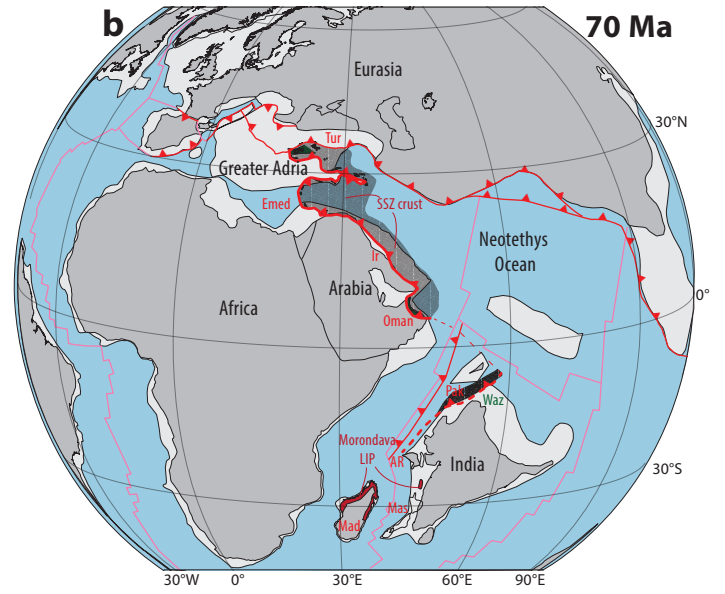
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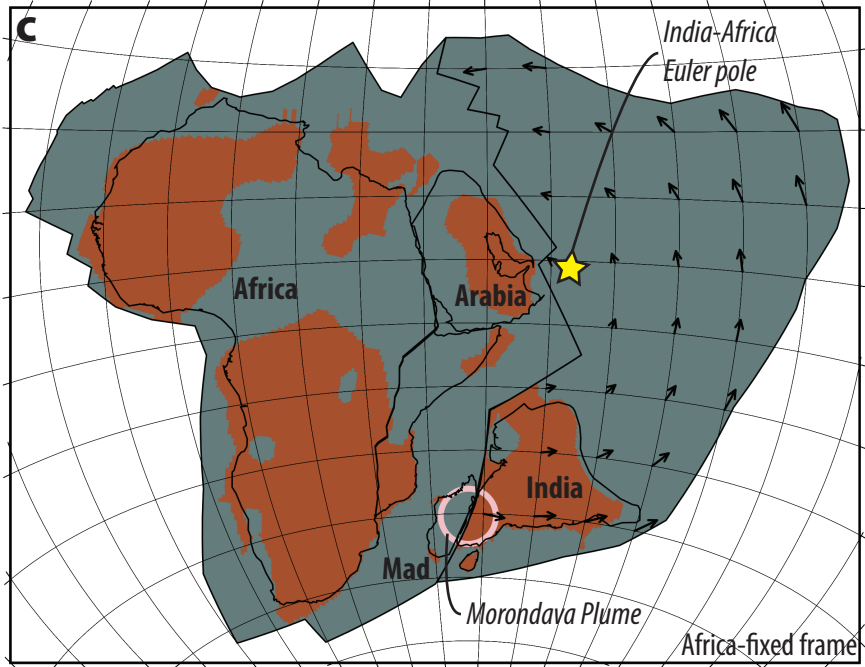
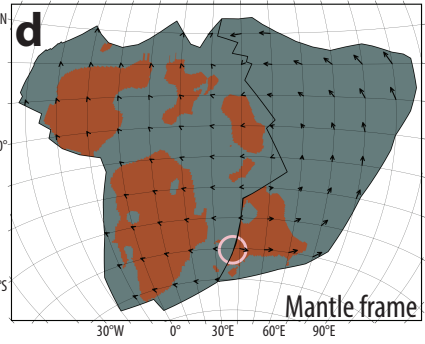
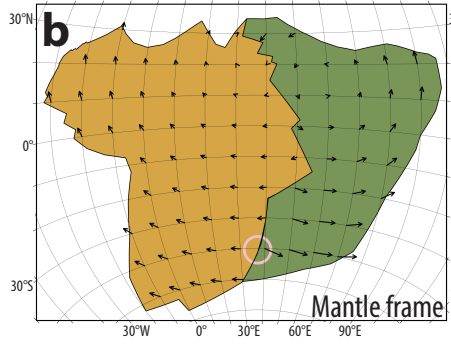
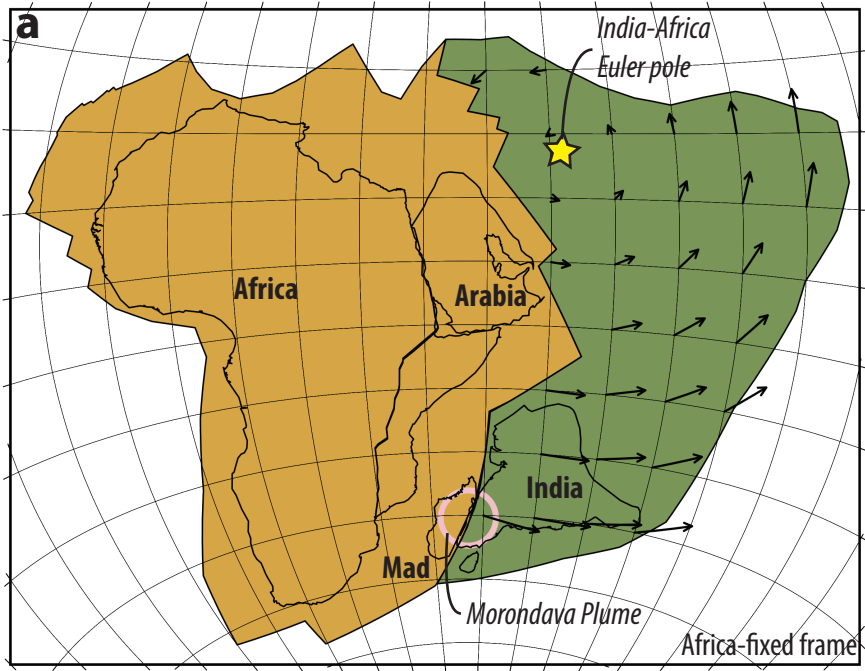
610 **Code availability**

611 All codes used in the geodynamic modeling in this study are available at
612 https://figshare.com/articles/software/van_Hinsbergen_et_al_NatureGeo_2021_geodynamics_package/13635089.

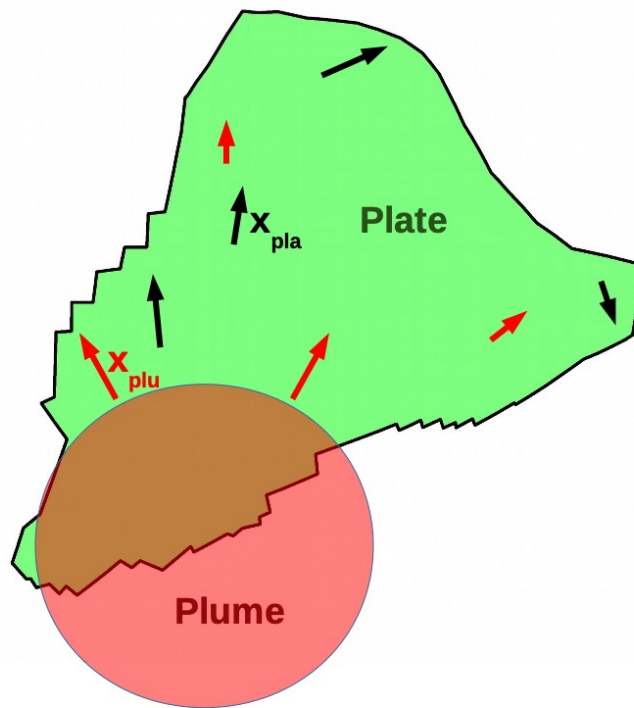
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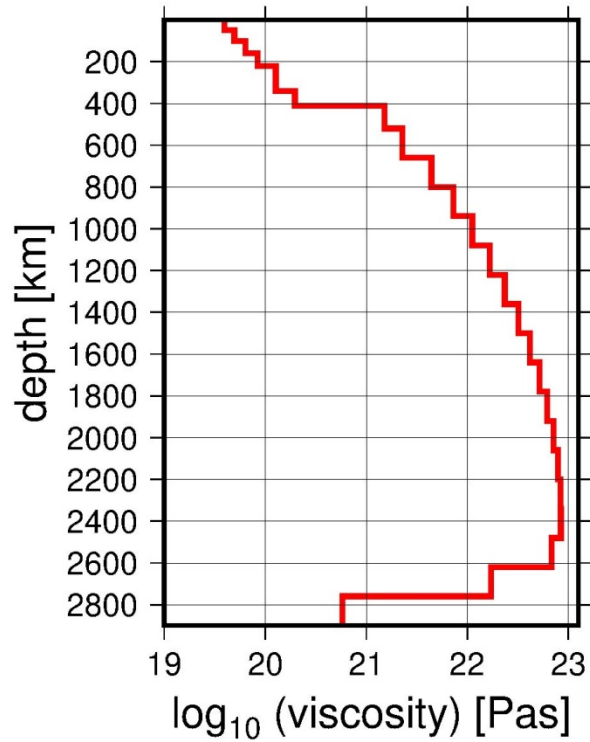


Extended Data

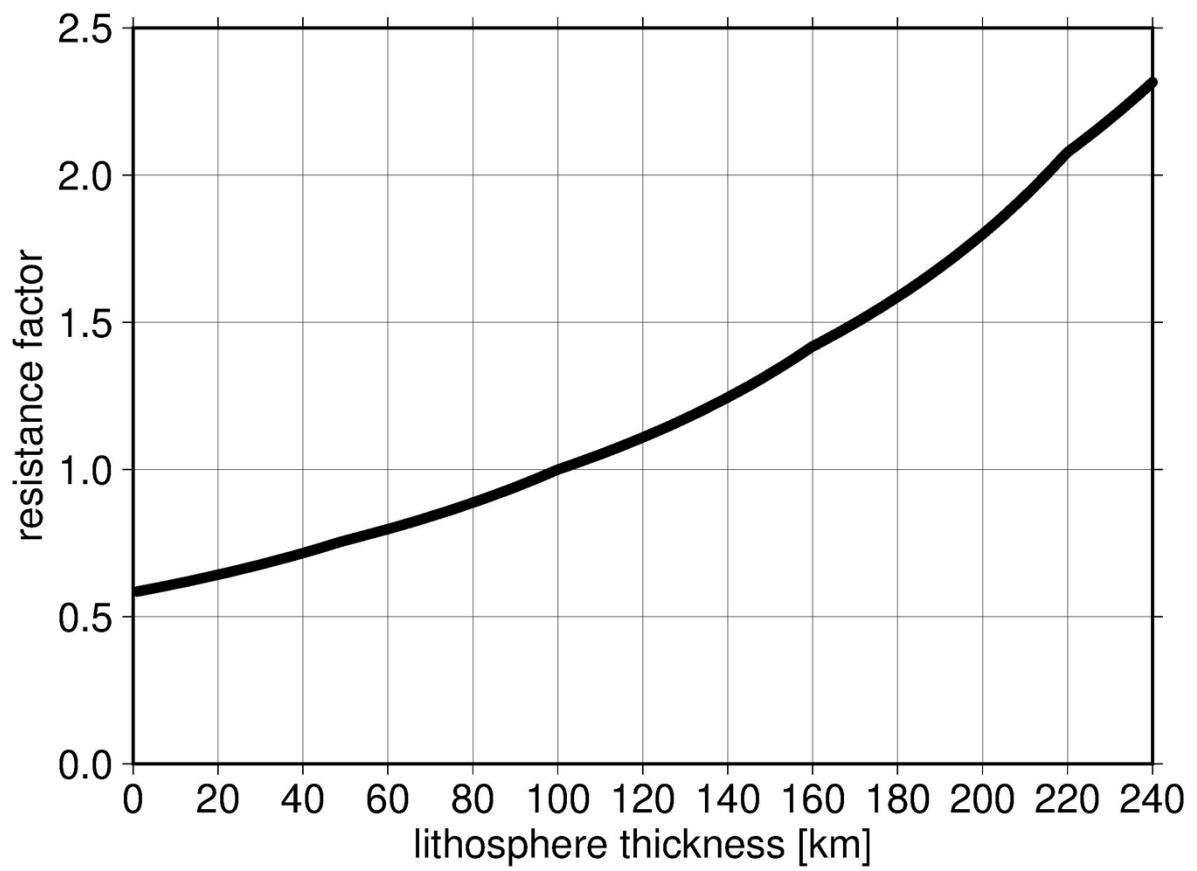


Extended Data Fig. 1: Sketch illustrating the geometry of a plume head (pink; not drawn to scale) hitting the boundary of a plate (green). x_{plu} (red arrows) are the (maximum) displacement vectors in the asthenosphere caused by emplacement of the plume. Motion vectors of the plate x_{pla} (black arrows) correspond to the plate rotation ω that is caused. Reversal of direction from left to right indicates the rotational component of motion induced by the plume push. Note that, since plume push is modelled as Poiseuille flow, the red arrows

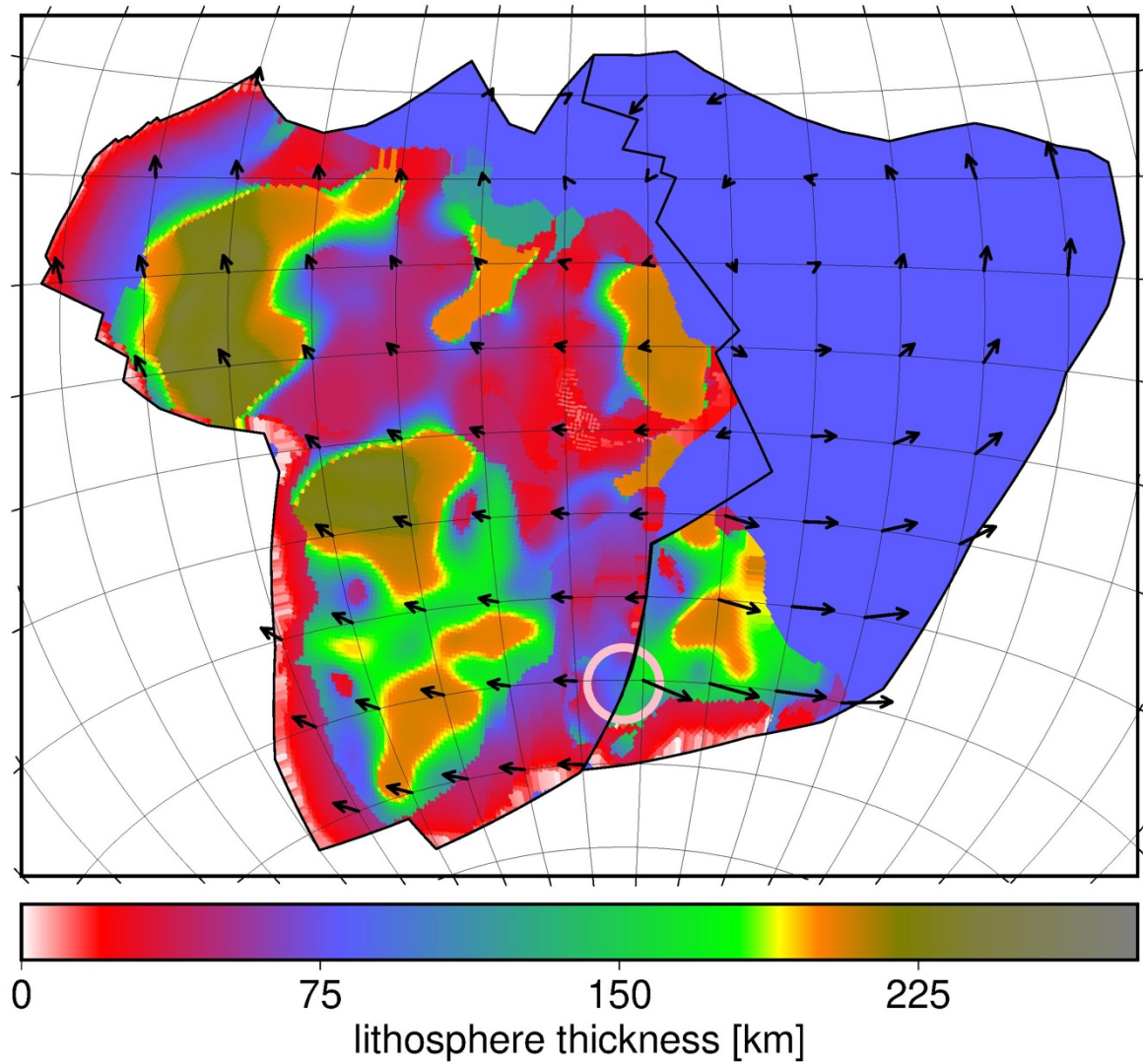
correspond to flow in the mid-asthenosphere, whereas plate motions induce Couette-type flow, therefore are shown at lithosphere depth.



Extended Data Fig. 2: Viscosity structure used, similar to Steinberger⁶⁶but without lithosphere. This is being combined with a lithosphere of constant or variable thickness.



Extended Data Fig. 3: Resistance factor to account for laterally variable coupling between lithosphere and mantle as a function of lithosphere thickness, for sub-lithospheric viscosity as in Fig. S2, according to eq. (2).



Extended Data Fig. 4: Computed total amount of displacement induced by the Morondava plume (pink circle), considering lithosphere thickness variations: It is assumed that, compared to a case with no lateral variations, the drag force due to the plate moving over the mantle is multiplied at each location with a resistance factor according to Fig. S3. Plates, plume and lithosphere thickness are reconstructed in the slab-fitted mantle reference frame⁴⁶.