

1 Upper mantle mush zones beneath low 2 melt flux ocean island volcanoes: 3 insights from Isla Floreana, Galápagos

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9 **ABSTRACT**

10 The physicochemical characteristics of sub-volcanic magma storage regions have important
11 implications for magma system dynamics and pre-eruptive behaviour. The architecture of magma
12 storage regions located directly above high buoyancy flux mantle plumes (such as Kīlauea, Hawai'i
13 and Fernandina, Galápagos) are relatively well understood. However, far fewer constraints exist on
14 the nature of magma storage beneath ocean island volcanoes that are distal to the main zone of mantle
15 upwelling or above low buoyancy flux plumes, despite these systems representing a substantial
16 proportion of ocean island volcanism globally. To address this, we present a detailed petrological
17 study of Isla Floreana in the Galápagos Archipelago, which lies at the periphery of the upwelling
18 mantle plume and is thus characterised by an extremely low flux of magma into the lithosphere.
19 Detailed *in situ* major and trace element analyses of crystal phases within exhumed cumulate
20 xenoliths, lavas and scoria deposits, indicate that the erupted crystal cargo is dominated by
21 disaggregated crystal-rich material (i.e., mush or wall rock). Trace element disequilibria between
22 cumulus phases and erupted melts, as well as trace element zoning within the xenolithic
23 clinopyroxenes, reveals that reactive porous flow (previously identified beneath mid-ocean ridges) is
24 an important process of melt transport within crystal-rich magma storage regions. In addition,
25 application of three petrological barometers reveal that the Floreana mush zones are located in the
26 upper mantle, at a depth of 23.7 ± 5.1 km. Our barometric results are compared to recent studies of
27 high melt flux volcanoes in the western Galápagos, and other ocean island volcanoes worldwide, and

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28 demonstrate that the flux of magma from the underlying mantle source represents a first-order control
29 on the depth and physical characteristics of magma storage.

30 **KEY WORDS**

31 Galápagos; magma storage; reactive porous flow; barometry.

32 **1 INTRODUCTION**

33 The physicochemical characteristics (such as size, pressure, volatile content and geochemical
34 heterogeneity) of magma storage at volcanic centres located directly above high buoyancy flux mantle
35 plumes (e.g. Kīlauea, Hawai'i and Isabela, Galápagos) have been subject to intense study over the
36 past few decades (Bagnardi et al., 2013; Bernard et al., 2019; Clague and Denlinger, 1994; Geist et
37 al., 1998; Naumann and Geist, 1999; Neal et al., 2019; Park et al., 2007; Pietruszka et al., 2015;
38 Poland et al., 2015; Sides et al., 2014; Stock et al., 2018; Wieser et al., 2020, 2019). Systems such as
39 Kīlauea are characterised by frequent volcanic activity, and geophysical (seismicity, ground
40 deformation) and geochemical (gas emissions) monitoring is prevalent. Monitoring data, combined
41 with petrological and geochemical analysis of erupted products (mineral textures, deformation
42 characteristics and chemistry) provide important insights into the architecture and dynamics of their
43 sub-volcanic plumbing systems (Amelung et al., 2000; Davidge et al., 2017; Geist et al., 2014;
44 Hartley et al., 2018; McCormick Kilbride et al., 2016). However, these systems (which we term 'high
45 melt flux') represent only one endmember of global plume-derived volcanism. Low melt flux
46 systems, either above low buoyancy flux plumes (e.g. Canary Islands; Longpre et al., 2014) or at
47 volcanic systems distal to the centre of mantle melting at high buoyancy flux mantle plumes (e.g.
48 eastern and south-eastern Galápagos; Harpp and Geist, 2018), are the other endmember.

49 While a substantial number of hotspot-related volcanic systems that have been active during the
50 Holocene are located in regions characterised by a relatively low flux of magma into the lithosphere
51 (i.e., regions distal to the main zone of plume upwelling or above low buoyancy flux plumes; Samoa,
52 Canary Islands, Cape Verde; Global Volcanism Program, 2013), only a small number of eruptions
53 have been observed (and recorded) at these systems since the advent of modern volcano monitoring

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54 techniques. As a result, few constraints exist on the conditions of magma storage in regions
55 characterised by a low flux of magma into the lithosphere, relative to volcanic centres located above
56 the centre of mantle plumes with a large buoyancy flux (and thus generating a large flux of magma
57 and more frequent eruptions).

58 The flux of mantle-derived magma into the lithosphere is thought to impart a first-order control on the
59 evolution of ocean island volcanoes and the homogeneity of erupted liquids (Geist et al., 2014).

60 Therefore, placing constraints on the physicochemical characteristics of magma storage at low melt
61 flux ocean island volcanoes is essential for determining the influence of mantle dynamics and melt
62 generation processes on the structure and physical characteristics of sub-volcanic magma plumbing
63 systems. In the absence of detailed monitoring data, petrological and geochemical analyses of
64 volcanic products from past eruptions represent the only available tools for determining the structure
65 and processes operating within these systems, as well as possible eruption precursors.

66 Isla Floreana in the south-eastern Galápagos is currently located ~100 km downstream from where the
67 centre of the Galápagos plume impacts on the base of the lithosphere beneath the island of Isabela in
68 the western archipelago (Fig. 1; Villagómez et al., 2014). Hence, although the Galápagos plume has a
69 relatively high buoyancy flux compared to regions such as the Canary Islands (Jackson et al., 2017),
70 Floreana's location relative to the main zone of mantle plume upwelling results in an extremely low
71 flux of magma entering the lithosphere and, consequently, very infrequent volcanic activity (Harpp et
72 al., 2014a; Harpp and Geist, 2018). Floreana is considered to be an infrequently active volcanic
73 centre, rather than extinct, owing to the persistent volcanic activity since ~1 Ma and the long-lived
74 nature of volcanism on the eastern Galapagos islands that lie >100 km 'downstream' of the Galapagos
75 mantle plume (e.g. the youngest lavas on San Cristobal are ~9 ka; Mahr et al., 2016).

76 In this paper, we present a thorough petrological study of scoria, lava and xenolith samples from
77 Floreana and place constraints of the structure, depth and crystallinity of magmatic systems beneath
78 this low melt flux ocean island volcano. We compare our results with more frequently active volcanic
79 centres in the western Galápagos (near the centre of plume upwelling; Geist et al., 1998; Naumann
80 and Geist, 1999; Stock et al., 2018), as well as other ocean island volcanoes worldwide (Hammer et

81 al., 2016; Hartley et al., 2018; Poland et al., 2015), to investigate how the flux of magma into the
82 lithosphere influences the depth and crystallinity of sub-volcanic magma storage regions.

83 **2 GEOLOGICAL BACKGROUND**

84 The Galápagos Archipelago in the eastern equatorial Pacific is one of the most volcanically active
85 regions on Earth, with eruptions typically occurring every 2–3 years (Global Volcanism Program,
86 2013). Although most historic Galápagos eruptions have taken place on the two westernmost islands
87 of Isabela and Fernandina (Fig. 1), infrequent volcanic activity has also occurred on several islands in
88 the eastern and south-eastern Galápagos (for example, Santiago in 1906 and Marchena in 1991;
89 Global Volcanism Program, 2013). In fact, volcanic activity in the eastern Galapagos, >100 km
90 ‘downstream’ of the postulated position of the plume stem, has been shown to be long-lived, with
91 volcanic activity on San Cristobal extending over 2 Myrs with the most recent lavas erupted at ~9 ka
92 (Mahr et al., 2016).

93 Volcanoes in the western Galápagos likely emerged within the last 500 kyr (Naumann and Geist,
94 2000), whereas those in the eastern and south-eastern Galápagos are considerably older (eruption ages
95 up to 2.3 Ma and 3.2 Ma have been measured on San Cristobal and Espanola, respectively; Bailey,
96 1976; Geist et al., 1986). In addition, substantial differences in geomorphology and the style of
97 volcanic activity are observed across the archipelago (Geist et al., 1995; Harpp et al., 2014a; Harpp
98 and Geist, 2018). For example, volcanoes in the western archipelago are typified by large summit
99 calderas (<700m deep), which are not present on the eastern islands (Chadwick and Howard, 1991;
100 Cleary et al., 2020; Harpp and Geist, 2018).

101 Geochemical distinctions between the western and eastern/south-eastern Galápagos islands are also
102 observed, which are primarily related to variations in the composition of the underlying mantle source
103 (Geist et al., 1988; Gibson and Geist, 2010; Gleeson et al., 2020; Harpp and White, 2001; White et al.,
104 1993) or the volume flux of mantle-derived magma that ascends into the lithosphere (Geist et al.,
105 1995, 2014; Gibson et al., 2016; Harpp and Geist, 2018). For example, variations in the flux of mantle
106 derived magma are hypothesised to influence the geochemical heterogeneity of erupted basalts at each

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107 island: volcanoes in the western archipelago typically erupt a very narrow range of basaltic
108 compositions over hundreds of millennia during their main shield building phase, whereas basalts
109 erupted from a single island in the eastern and/or south-eastern archipelago, such as Floreana, tend to
110 display far greater compositional heterogeneity (Geist et al., 2014; Harpp and Geist, 2018).

111 Floreana is characterised by numerous scoria cones and blocky, heavily vegetated lava flows that can
112 typically be traced to the cone from which they originated (Bow and Geist, 1992; Harpp et al., 2014a).

113 The crustal thickness beneath Floreana is ~16 km, similar to that observed in the western Galápagos
114 10–18 km (Feighner and Richards, 1994), and the lithospheric thickness beneath the western and
115 south-eastern Galápagos is very similar (~50–60 km; Gibson and Geist, 2010). Recent work has
116 shown that the average volumetric eruption rate on Floreana over the past 1–1.5 Myrs is 1–10 m³·yr⁻¹,
117 6 orders of magnitude lower than the current volcanic productivity at volcanoes in the western
118 Galápagos (cf. ~4.4·10⁶ m³·yr⁻¹ at Fernandina; Harpp et al., 2014a; Kurz et al., 2014). This difference
119 in volumetric eruption rate likely reflects a substantially lower flux of magma into the lithosphere
120 beneath Floreana than beneath each volcanic centre in the western Galápagos.

121 Despite the relatively low volcanic productivity of Floreana over the last ~1–1.5 Myrs, the erupted
122 products have several important characteristics that provide insights into the nature of the underlying
123 magmatic system. For example, Floreana has a high proportion of pyroclastic deposits compared to
124 the other Galápagos islands (Harpp et al., 2014a) and eruption deposits typically contain a large
125 number of cumulate xenoliths (Bow and Geist, 1992; Lyons et al., 2007), which have been interpreted
126 as evidence for very high magma ascent rates (Harpp et al., 2014a).

127 Floreana is the only Galápagos island that displays evidence for multiple stages in its volcanic
128 evolution. Submarine parts of the island have isotopic and trace element characteristics that are
129 similar to those measured in recent basalts erupted on southern Isabela, near the centre of plume
130 upwelling (e.g. Sierra Negra and Cerro Azul, Fig. 1A), whereas the subaerial material is isotopically
131 distinct (high ²⁰⁶Pb/²⁰⁴Pb and ⁸⁷Sr/⁸⁶Sr ratios; Harpp et al., 2014). The trace element and isotopic
132 differences between the erupted basalts is mirrored in xenoliths found in the Floreana lava and scoria
133 deposits: gabbroic xenoliths have radiogenic isotope ratios that are similar to modern Isabela basalts,

134 whereas wehrlitic xenoliths have trace element and isotopic compositions that resemble recent
135 Floreana subaerial basalts (Lyons et al., 2007). Differences in the isotopic characteristics of the
136 Floreana lavas (submarine vs subaerial) and xenoliths (gabbros vs wehrlites) are thought to indicate a
137 change in the mean composition of magma produced by mantle melting beneath the island at ~1–1.5
138 Ma (Harpp et al., 2014a). In this study, we focus on constraining the depth and physical
139 characteristics of magma storage during the most recent period of volcanic activity on Isla Floreana
140 (<1–1.5 Ma) using chemical and textural analysis of crystal phases in lava flows and xenolithic
141 nodules.

142 **3 SAMPLES AND PETROGRAPHY**

143 The Floreana samples analysed in this study consist of lavas (27 samples), scoria (2 samples) and
144 xenoliths (4 wehrlite, 3 dunite and 2 gabbro samples; Fig. 1B). Most lava samples were collected from
145 the unaltered, low vesicularity cores of blocky flows or glassy flow fronts. All samples form part of
146 the Main Series of Floreana lavas identified by Bow and Geist (1992). The scoria samples were
147 collected from two separate deposits and comprise rapidly cooled scoria lapilli (~0.5–2 cm across;
148 17MMSG16) and bombs (~10 cm across; 17MMSG20). Xenolithic fragments (3–15cm across) were
149 sampled from two different scoria cones on the north-east coast of Floreana; similar xenoliths are also
150 found within most lava flows across the entire island (Supplementary Information).

151 **3.1 Lavas and scoria**

152 The lava and scoria samples analysed in this study are typically olivine phyric with minor anhedral
153 clinopyroxene and very rare orthopyroxene. Except for small plagioclase laths in the microcrystalline
154 groundmass, plagioclase crystals are extremely rare in the Floreana lavas. Plagioclase macrocrysts are
155 only present in one of our lava samples (17MMSG29) where they occur as isolated phenocrysts and in
156 plagioclase-olivine crystal clots (Table S.1). The abundance of olivine and absence of plagioclase in
157 the Floreana lavas and scoria contrasts with basalts in the central, northern, and western parts of the
158 Galápagos Archipelago, where plagioclase-phyric and ultraphyric basalts are common (Geist et al.,
159 2002; Gibson et al., 2012; Harpp et al., 2014b).

160 Despite their relatively simple mineralogy, Floreana lava and scoria samples contain texturally diverse
161 olivine crystals which can be divided into five distinct groups (Fig. 2):

162 - Group 1 olivines are present in all lava and scoria samples and are the most abundant type of
163 olivine (~60-70% of all crystals). They are characterised by homogeneous cores, with respect
164 to major elements, and narrow normally zoned rims (Fig. 2A). Group 1 olivines are generally
165 subhedral to euhedral.

166 - Group 2 olivines are the second most abundant group (~20-30%) and display reverse zoning.
167 They are typically euhedral, with occasional small embayments (Fig. 2B).

168 - Group 3 olivines are also reversely zoned, but are distinguished by skeletal overgrowths,
169 indicating significant undercooling of the host magma and rapid crystal growth (Fig. 2C;
170 Donaldson, 1976; Welsch et al., 2014). Group 3 olivines are less abundant than Groups 1 and
171 2 olivines (<10%).

172 - Group 4 olivines are present in low abundance (<10%). They have homogeneous cores and
173 reverse-zoned rims (up to ~300 μm thick). The rims have sieve textures, potentially
174 suggesting resorption and chemical disequilibrium with their carrier melts (Fig. 2D).

175 - Group 5 olivines are only found in a minority of samples and are characterised by the
176 presence of 4 compositional zones with alternating high and low forsterite contents (visible in
177 back-scattered electron images; Fig. 2E and F).

178 3.2 Xenoliths

179 3.2.1 Gabbroic xenoliths

180 Floreana gabbroic xenoliths predominantly comprise plagioclase (33–66 vol.%), clinopyroxene (28–
181 46 vol.%) and orthopyroxene (5–15 vol.%), with little or no olivine (Table S.2). Plagioclase and
182 pyroxene crystals are typically >500 μm and grain sizes are relatively uniform within a single xenolith
183 sample. Where three plagioclase grains meet at a triple junction, 120° grain boundaries indicate a high
184 degree of textural equilibrium (Fig. 3A; Holness et al., 2005). Some of the gabbros have
185 clinopyroxene-rich and plagioclase-rich layers of 2–5 mm thickness.

186 3.2.2 *Dunitic xenoliths*

187 Olivine dominates the dunitic Floreana xenoliths (>90 vol.%). The olivine crystals are subhedral to
188 euhedral and may have undergone partial textural re-equilibration, with some olivine triple junctions
189 approaching 120° grain boundaries. Minor intercumulus clinopyroxene is present along grain
190 boundaries and between pre-existing olivine crystals (Fig. 3B).

191 3.2.3 *Wehrlitic xenoliths*

192 Floreana wehrlitic xenoliths contain olivine (>50 vol.%), clinopyroxene (20–40 vol.%),
193 orthopyroxene (~0–7 vol.%) and minor spinel (<1 vol.%; Table S.2). Clinopyroxene typically occurs
194 as large (<5 mm) oikocrysts, which enclose rounded olivine chadacrysts <500 µm in diameter (Fig.
195 3C and D). Fine-scale orthopyroxene exsolution lamellae (<2 µm) is observed in the clinopyroxenes
196 of a single wehrlite (17MMSG03a; Fig. 3E and F). Olivine grains that are not enclosed by
197 clinopyroxene are typically larger (>1 mm) and more euhedral than the chadacrysts. In some samples,
198 the boundary between clinopyroxene and olivine crystals is characterised by a thin (<20–30 µm) layer
199 of glass and very fine-grained microcrysts. Excluding the rare exsolution lamellae in sample
200 17MMSG03a, orthopyroxene crystals are typically anhedral, infilling the space between earlier
201 formed clinopyroxene and olivine grains. Our observations of dunitic and wehrlitic xenoliths (which
202 have the isotopic signatures of modern day Floreana basalts; Lyons et al., 2007) indicate that the
203 typical order of crystallisation beneath Floreana is olivine, followed by clinopyroxene, with little to no
204 crystallisation of plagioclase.

205 **4 ANALYTICAL METHODOLOGY**

206 **4.1 Electron microprobe analysis**

207 Glass chips, olivine and clinopyroxene crystals were hand-picked from crushed scoria and lava
208 samples, mounted in epoxy or indium, and then ground and polished prior to analysis (crystals
209 mounted in indium were polished individually prior to mounting). Xenolithic crystals were analysed
210 as individual crystals mounted in indium or *in situ* in petrographic thin sections. The major and minor
211 element concentrations of olivine, clinopyroxene and glass were measured using a Cameca SX100

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212 electron microprobe in the Department of Earth Sciences, University of Cambridge. Calibrations were
213 made using mineral and metal standards prior to each analytical session (see Gleeson and Gibson,
214 2019 for details). Glasses were analysed using a 6 nA, 15 kV, defocused (5 μm) beam for most
215 elements. Na and K were analysed first (10 s peak count time) to avoid alkali migration. Other
216 elements were analysed with peak count times of 10 s (Si), 20 s (Fe), 30 s (Al, P, Ca, Mg), 40 s (Mn),
217 or 60 s (Ti). Sulphur was analysed last using a 20 nA beam current and a 60 s peak count time.

218 Pyroxene compositions were determined by spot analyses using a 20 nA, 15 kV, focused ($\sim 1 \mu\text{m}$)
219 beam, with Na, K and Si analysed first (10 s). Element maps of Cr, Ti, and Al in key clinopyroxene
220 crystals from the Floreana xenoliths were created using a 60 nA, 15 kV, focused ($\sim 1 \mu\text{m}$) beam, with
221 a dwell time of 150 ms. Cr counts were collected on a PET and a LIF crystal, Al counts were
222 collected on two TAP crystals, and Ti counts were collected on a PET crystal. Olivine electron
223 microprobe analysis was carried out using the method outlined in Gleeson and Gibson (2019).

224 Analytical uncertainties were tracked through analysis of appropriate Smithsonian Microbeam
225 Standards (Jarosewich et al., 1980). Accuracy was typically between 98 and 102% for all phases. 2σ
226 analytical precision of clinopyroxene and olivine analyses are typically better than 2–3% for major
227 elements ($>1 \text{ wt}\%$) and typically $\sim 5\text{--}10\%$ for minor elements ($<1 \text{ wt}\%$). Similarly, the 2σ precision for
228 glass analysis was typically $<3\%$ for major elements, $\sim 5\%$ for Na, and $\sim 10\%$ for K (see
229 Supplementary Information).

230 **4.2 Laser ablation Inductively Coupled Plasma Mass Spectrometry**

231 Trace element concentrations were measured in the apparent cores (i.e. as exposed in the 2D plane) of
232 clinopyroxene crystals from scoria and xenolith samples using an ESI193 laser coupled to a Nexion
233 350D inductively coupled plasma mass spectrometer in the Department of Earth Sciences, University
234 of Cambridge. Analyses were collected in spot mode using a 20 Hz laser repetition rate, 4 J/cm^2
235 fluence and $80 \mu\text{m}$ spot size, or in transect mode using a 10 Hz repetition rate, 3.5 J/cm^2 fluence and
236 $30 \mu\text{m}$ spot size. For transects, individual spots were offset into two (alternating) lines to increase the
237 spatial resolution. Laser ablation inductively coupled plasma mass spectrometry (LA-ICP-MS) data

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238 reduction was carried out in Iolite[®], with NIST 612SRM as the standard reference material (Hinton,
239 1999) and ⁴³Ca (from electron microprobe analysis) as the internal reference standard. Analytical
240 accuracy was tracked using a USGS glass standard (Jochum et al., 2016) and was between 95% and
241 105% for most elements (See Supplementary File). 2 σ analytical precision of spot analyses were
242 monitored through analysis of an in-house clinopyroxene standard and was 5–10% for the light rare-
243 earth elements (LREE), Y, Sr, and Zr and 10–20% for the heavy rare earth element (HREE). 2 σ
244 analytical precision was ~10% for all elements of interest in transect analyses (Ce, Y).

245 **5 GLASS AND MINERAL CHEMISTRY**

246 **5.1 Matrix glass compositions**

247 The matrix glass compositions measured in one scoria (17MMSG16) and two glassy lava samples
248 (17MMSG12 and 17MMSG27) from Floreana have very similar MgO concentrations (mean
249 concentrations of 6.06–6.67 wt%) but exhibit differences in the concentrations of other elements (Fig.
250 4). For example, sample 17MMSG12 has consistently lower TiO₂ and Al₂O₃ concentrations than
251 17MMSG16, which must either reflect heterogeneity in the composition of primary mantle melts or
252 variations in crustal processing (e.g. the extent of plagioclase or clinopyroxene crystallisation; Fig.
253 4A,D). The largest variation in the glass major element composition, however, is seen in sample
254 17MMSG27 where, at a near constant MgO contents, the CaO, Na₂O, and K₂O contents vary by ~6
255 wt%, 4.5 wt%, and 1.5 wt%, respectively (Fig. 4B,C). Differences between our matrix glass major
256 element analyses and previously-published whole-rock data from Floreana (Harpp et al., 2014a) are
257 primarily due to olivine accumulation in the whole-rock samples (additional accumulation of
258 clinopyroxene may explain the high CaO content of some whole-rock samples; Fig. 4C).

259 Our Floreana matrix glass analyses have higher Al₂O₃ concentrations, at a given MgO content, than
260 basaltic glass and whole-rock measurements from the western Galápagos Archipelago (excluding
261 whole-rock samples with accumulated plagioclase, Fig. 4D; Geist et al., 2002). This indicates
262 substantially lower extents of plagioclase fractionation in the Floreana magmatic system and is
263 consistent with the scarcity of plagioclase phenocrysts in erupted Floreana lavas. Reduced plagioclase

264 crystallisation could be due to the major element composition or H₂O content of primary mantle melts
265 and/or increased pressure of magma storage (Asimow and Langmuir, 2003; Neave et al., 2019;
266 Thompson, 1987; Winpenny and Maclennan, 2011).

267 **5.2 Olivine compositions**

268 Olivine crystals in our Floreana lava and scoria samples show large variations in their forsterite
269 contents (Fo = 70–92, where Fo=(Mg/(Mg+Fe²⁺) molar) with histograms showing a primary density
270 peak at Fo~85 (Fig. 5), more primitive than the olivine compositions in equilibrium with basaltic
271 glasses from Floreana (K_D=0.27-0.34, assuming a Fe³⁺/Fe_{tot} ratio of 0.15; Matzen et al., 2011;
272 Peterson et al., 2015; Roeder and Emslie, 1970). Although there is no clear correlation between Fo
273 and Ca concentration in these crystals, the most forsteritic olivines (Fo>83) have extremely diverse Ca
274 contents (~250 to ~2600 ppm; Fig. 5), whereas crystals with lower Fo contents (<83) have
275 ubiquitously low Ca concentrations (<1500 ppm). All crystals with >1500 ppm Ca are classified as
276 Group 1 olivines. Crystals with <1500 ppm Ca, however, may be classified in any of the 5 olivine
277 groups, with the most evolved crystals (Fo ~70) the only to display the sieve-like rim texture of Group
278 4 olivines.

279 Floreana olivines separated from the lava and scoria samples also have a large range of Ni
280 concentrations (~700 to ~3200 ppm), consistent with crystallisation from a peridotite-derived melt
281 (Fig. 5; Gleeson and Gibson, 2019; Herzberg, 2011; Matzen et al., 2017b, 2017a). All olivine crystals
282 analysed in the wehrlite and dunite xenoliths have a narrow range of Fo contents (83-87) and,
283 crucially, have uniformly low Ca concentrations (<1000 ppm) and moderately high Ni contents
284 (~2000 ppm). Core-rim profiles were performed on a small number of the wehrlitic olivines, revealing
285 typically flat profiles in Fo, Ni, and Ca, with one profile displaying evidence for an increase in Ca in
286 the outer 10 – 20 µm (see Supplementary Information). The range in Ca contents of the Floreana
287 olivines contrasts with isolated olivine crystals from other Galápagos islands where Ca is typically
288 >1000 ppm (Gleeson and Gibson, 2019; Vidito et al., 2013). The Ca and Ni contents of our Floreana
289 olivines are inversely related at a set Fo content (Fig. 5B).

290 **5.3 Clinopyroxene compositions**

291 *5.3.1 Major elements*

292 The Floreana clinopyroxenes separated from the scoria and xenolith samples are augitic, and have a
293 relatively high Mg# (0.85–0.90; Mg# = Mg/(Mg+Fe_i) molar). The clinopyroxenes reach high Na
294 concentrations up to 1.58 wt% Na₂O (0.11 Na atoms per formula unit; Fig. 6A) and, correspondingly,
295 up to 10% of the jadeite component. In general, clinopyroxene separates from scoria samples display
296 a wide range of Na concentrations, although some xenolithic clinopyroxenes extend to higher Na
297 contents (Fig. 6A). The Floreana clinopyroxenes display a large range of Cr contents, ranging from
298 <0.05 wt% in the most evolved crystals to ~1.72 wt% in the more primitive crystals (Fig. 6B).

299 Clinopyroxene analyses are typically taken from the cores of crystals, but zoning was characterised by
300 a series of transects (and maps) on the xenolithic clinopyroxenes (see Supplementary Information).
301 Results indicate that the clinopyroxenes are unzoned with respect to most element (including MgO
302 and FeO), but some zoning is present in their Al₂O₃, TiO₂ and Cr₂O₃ contents. Specifically,
303 clinopyroxene crystal rims are typically characterised by elevated TiO₂, but lower Cr₂O₃ contents.

304 *5.3.2 Trace elements*

305 Our Floreana clinopyroxenes display a wide range of geochemical enrichment, with LREE to MREE
306 ratios varying from [La/Sm]_n ~0.2 to [La/Sm]_n ~3.1 (where n represents normalisation to the
307 primitive mantle composition of Sun and McDonough, 1989). Wehrlitic clinopyroxenes typically
308 have more enriched trace element ratios (such as [La/Sm]_n or [Ce/Y]_n) than clinopyroxenes from the
309 scoria samples (Fig. 7). Furthermore, melt [La/Sm]_n ratios calculated to be in equilibrium with
310 clinopyroxenes from the scoria and xenolith samples range from ~1 to ~15 (calculated using the major
311 element composition of the clinopyroxene at the location of LA-ICP-MS analysis and the elastic
312 strain model of Wood and Blundy (1997) at 1225°C and 700 MPa; Fig. 7B), significantly greater than
313 the range measured in Floreana whole-rock samples (~2–5, with a small number of outliers up to
314 ~7.5; Harpp et al. 2014a). Almost all of the xenolithic crystals, and a large proportion of the
315 clinopyroxenes separated from scoria deposits, are too enriched to be in equilibrium with the typical
316 composition of melts erupted on Floreana (Harpp et al., 2014a). In addition, there is a strong

317 correlation between the Na concentrations and highly/moderately incompatible trace element ratios of
318 the Floreana clinopyroxenes ($p < 10^{-3}$), such that crystals with enriched trace element signatures
319 typically contain a high jadeite component (Fig. 7A). Finally, the Floreana clinopyroxenes all contain
320 Eu anomalies ($Eu^* = Eu_n / \sqrt{Sm_n \times Gd_n}$) within analytical uncertainty of 1, and Sr anomalies ($Sr^* =$
321 $Sr_n / \sqrt{Pr_n \times Nd_n}$) that have a similar range to that observed in the erupted basalts (Harpp et al.,
322 2014a).

323 **6 MAGMA SYSTEM ARCHITECTURE BENEATH** 324 **ISLA FLOREANA**

325 **6.1 Mush crystallisation and textural equilibration**

326 Based on pyroxene trace element and radiogenic isotope ratios, Lyons et al. (2007) hypothesised that
327 gabbroic xenoliths in the Floreana lava and scoria deposits formed in an ancient (>1 -1.5 Ma)
328 magmatic system, compositionally similar to those currently beneath Cerro Azul and Sierra Negra
329 volcanoes in the western Galápagos. In contrast, wehrlite xenoliths preserve isotopic ratios similar to
330 more recent subaerial lavas on Floreana, suggesting that they are fragments of the present-day
331 magmatic system (Lyons et al., 2007).

332 Our wehrlitic xenoliths preserve an original poikilitic igneous texture (clinopyroxene oikocrysts
333 surrounding olivine chadacrysts) and display no evidence for the metamorphic breakdown of
334 plagioclase (e.g. pseudomorphs or relict cores) as hypothesised by Lyons et al. (2007). We suggest
335 that the clinopyroxene crystals in our wehrlitic xenoliths grew within an olivine-dominated cumulate
336 mush (i.e. interstitial growth of clinopyroxene oikocrysts; Wager et al., 1960). If clinopyroxene
337 growth is principally within a crystal-rich (i.e. relatively viscous and immobile) mush zone, this could
338 explain its relatively low abundance in Floreana lava and scoria deposits ($<5\%$ of separated crystals).
339 In contrast, plagioclase triple junctions in the gabbroic xenoliths have $\sim 120^\circ$ dihedral angles (Fig. 3A),
340 indicating a high degree of textural equilibration (Holness et al., 2019, 2005). Plagioclase textural
341 equilibrium, along with the two-pyroxene phase assemblage, suggests that the gabbroic xenoliths

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342 represent magmatic cumulates which were stored at high temperatures (>900°C) on long timescales
343 (Holness et al., 2006). These petrographic observations are consistent with the gabbroic nodules
344 sampling an ancient magmatic system beneath Floreana (>1 Ma; Lyons et al., 2007).

345 **6.2 Mush disaggregation prior to eruption**

346 *6.2.1 Insights from olivine compositional heterogeneity*

347 Olivine crystals separated from the Floreana lava and scoria samples have a wide range of zoning
348 patterns, morphologies, and compositions (Fig. 2). The five olivine groups identified in the Floreana
349 samples have distinct morphologies and zoning patterns (see Section 2 above), suggesting chemically
350 heterogeneous magma storage (Holness et al., 2019). In particular, the most evolved crystals (Group
351 4; Fo~70–75) are in equilibrium with melts that are more evolved than the Floreana erupted basalts
352 (likely basaltic andesites). This is consistent with a recent study which identified highly evolved
353 (andesitic - dacitic) magmas beneath basaltic volcanoes in the western Galápagos Archipelago (Stock
354 et al., 2020).

355 As Fe-Mg interdiffusion in olivine is geologically fast (Chakraborty, 2010; Costa et al., 2020),
356 preservation of forsterite zoning in the Floreana olivine crystals suggests that multiple magma batches
357 interacted on relatively short pre-eruptive timescales. In Group 5 olivines, for example, four
358 compositional zones are preserved over ~100-200 µm (Fig. 2E and F). Whilst we do not have enough
359 Group 5 olivine crystals to calculate statistically robust timescales of pre-eruptive magma interactions
360 using diffusion chronometry, complex forsterite zoning over ~100 µm is estimated to last <3 yrs at the
361 approximate temperature of basaltic magma storage (~1225°C; using diffusion coefficients from
362 Chakraborty, 2010). Therefore, we suggest that the range of crystal morphologies and major element
363 compositions displayed by the Floreana olivines in lava and scoria deposits indicates mixing of
364 chemically heterogeneous magma storage regions over relatively short timescales prior to eruption.

365 The minor element chemistry of the olivine crystals allows us to investigate the crystallinity of these
366 chemically diverse magma storage regions. Olivine crystals in our Floreana lava and scoria deposits
367 have an unusually low, and large range of Ca concentrations (~250–2600 ppm compared with ~1500-
368 3000 ppm in the eastern Galápagos; Gleeson and Gibson, 2019; Fig. 5A). The lower end of the range

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369 in Ca concentrations measured in the Floreana lava and scoria deposits overlaps with those observed
370 in cumulate xenoliths (wehrlites) and thus are unlikely to represent mantle olivines (Thompson and
371 Gibson, 2000). Previous studies have shown that the Ca concentration of magmatic olivine is sensitive
372 to several parameters, including: (i) the major element composition of the co-existing melt phase
373 (Herzberg, 2011); (ii) the H₂O content of the co-existing melt phase (Gavrilenko et al., 2016); and (iii)
374 the temperature of the system (Adams and Bishop, 1982; Köhler and Brey, 1990; Shejwalkar and
375 Coogan, 2013).

376 Variations in the Ca content of primary mantle melts are commonly hypothesised to result from the
377 presence of a lithologically heterogeneous mantle source (Herzberg, 2011; Sobolev et al., 2007;
378 Vidito et al., 2013). Specifically, melting of a pyroxene-rich mantle lithology is expected to produce
379 Ca-poor and Ni-rich melts which would go on to form Ca-poor, but Ni-rich, olivines. While the low
380 Ca contents of many of the Floreana olivines would therefore appear to indicate derivation from melts
381 of a pyroxenitic source lithology, the olivine Ni contents are relatively low (<3000 ppm), inconsistent
382 with this hypothesis (Fig. 5B; Gleeson and Gibson, 2019). As a result, if the low Ca contents of the
383 Floreana olivines represent equilibrium with low Ca melts, we require a different process to generate
384 these compositions. This process must reduce the Ca concentration of the melt phase (and co-existing
385 olivines), without simultaneously reducing the melt Mg# (as low Ca concentrations are observed
386 across the entire range of forsterite contents in the Floreana olivines; Fig. 5A).

387 Evidence for the origin of the low-Ca contents in the Floreana olivines from lava and scoria deposits
388 is present in the texture and composition of the wehrlitic xenoliths, which contain uniformly low-Ca
389 olivine crystals (<1000 ppm; Fig. 5A). The petrography of the wehrlitic xenoliths attests to
390 clinopyroxene growth within olivine-dominated mush regions. Clinopyroxene crystallisation within
391 this mush would extract CaO and MgO from the residual melt. However, in an olivine-rich mush, the
392 large reservoir of MgO contained within the cumulus olivine grains would buffer the residual melt at
393 a near-constant Mg# during clinopyroxene crystallisation (Meyer et al., 1989). In contrast, the CaO
394 concentration of the melt is not buffered and decreasing melt CaO contents, due to clinopyroxene
395 crystallisation, will cause the CaO concentration of cumulus olivine grains to decrease (as a result of

396 diffusive re-equilibration). Support for this interpretation comes from the anomalously low CaO
397 concentrations in the matrix glass of sample 17MMSG27.

398 Intercumulus clinopyroxene growth would also increase the H₂O concentration of the residual melt
399 phase, decreasing the partition coefficient of Ca into olivine (Gavrilenko et al., 2016). Therefore,
400 variable amounts of clinopyroxene crystallisation within an olivine-dominated mush, and subsequent
401 disaggregation of this mush by an ascending melt, could explain the range of Ca contents measured in
402 the Floreana olivines derived from lava and scoria deposits. Specifically, we suggest that: (i)
403 forsteritic olivine crystals (Fo>83) with Ca contents >1500 ppm are consistent with those expected
404 from fractional crystallisation of mantle-derived melts in a liquid-rich magma storage region (Fig. 5A;
405 Gleeson and Gibson, 2019); (ii) olivines with Ca contents <1000 ppm, overlapping with the wehrlitic
406 olivines, reflect equilibrium with Ca-poor interstitial melts; and (iii) olivine crystals with intermediate
407 Ca concentrations (1000 – 1500 ppm) are sourced from regions where clinopyroxene growth is less
408 extensive, or ongoing at the time of eruption. Thus, the olivine crystal cargo of the Floreana magmas
409 is predominantly derived from crystal-rich domains with only a small number of olivine crystals
410 displaying compositions that are consistent with fractional crystallisation in liquid-rich storage regions
411 (Ca >1500 ppm; Gleeson and Gibson, 2019).

412 It is an important to note, however, that the partitioning of Ca between co-existing olivine and
413 clinopyroxene has been hypothesised to be sensitive to temperature (such that less Ca enters the
414 olivine structure at lower temperature; Shejwalkar and Coogan, 2013). As a result, the heterogeneity
415 in the Ca content of the Floreana olivines could instead represent disaggregation of xenolithic material
416 that has undergone variable amounts of cooling. Both hypotheses presented here can recreate the
417 range of Ca contents observed in the Floreana olivines and, as the majority of olivine analyses from
418 the Floreana lava and scoria deposits have low Ca concentrations (<1000 ppm) that overlap with those
419 in xenolithic nodules, indicate that a large proportion of the erupted crystal cargo derives from
420 disaggregated, highly crystalline magma storage regions.

421 6.2.2 *Insights from clinopyroxene major element compositions*

422 The compositions of clinopyroxene crystals from the Floreana scoria also overlap with those in our
423 xenolith samples, supporting the hypothesis that some of the erupted crystals are derived from
424 disaggregated sub-volcanic mush (Fig. 6). We used hierarchical cluster analysis to subdivide our 567
425 clinopyroxene major element analyses from the Floreana scoria and xenolith samples and determine
426 the proportion of material that is derived from each xenolith lithology in the erupted crystal cargo
427 (cluster analysis was performed in Python 3.8 using the scikit-learn package of Pedregosa et al. 2011).
428 We find that our clinopyroxene analyses form three distinct clusters (Fig. 8):

- 429 - Cluster 1 clinopyroxenes are predominantly from the wehrlite and dunite xenoliths and
430 include 90% of our analyses from these samples. 39% of clinopyroxenes analysed from the
431 scoria samples also fall into this cluster.
- 432 - Cluster 2 clinopyroxenes include all analyses from the gabbroic xenoliths, and ~10% of
433 analyses from crystals separated from the scoria samples.
- 434 - Cluster 3 clinopyroxenes are dominated by analyses of the scoria derived clinopyroxenes
435 (~50% of analyses from the scoria separates). However, 10% of analyses from the wehrlite
436 and dunite xenoliths also fall into this cluster.

437 Of our 248 clinopyroxenes analyses from the Floreana scoria, approximately half are classified as
438 Cluster 3 and thus have major element compositions that do not show a clear chemical affinity to
439 either the wehrlite/dunite or gabbroic cumulates. Therefore, these crystals may represent autocrysts
440 (defined here as crystals that are genetically related to primary mantle melts beneath Floreana, but are
441 not influenced by secondary cumulate processes) that grew in liquid-rich magma storage regions. The
442 remainder of clinopyroxene analyses from the scoria are either compositionally analogous to those in
443 the wehrlite and dunite xenoliths (Cluster 1; 39%) or the gabbroic xenoliths (Cluster 2; 11%); we
444 interpret these as representing disaggregated sub-volcanic mush or wall rock. The high proportion of
445 the clinopyroxene crystal cargo that is derived from highly crystalline storage regions beneath
446 Floreana is consistent with our interpretation of olivine minor element concentrations.

447 **6.3 Reactive Porous Flow within a cumulate mush**

448 Whilst the olivine and clinopyroxene major and minor element concentrations show that a large
449 proportion of the erupted crystal cargo is derived from highly crystalline magma storage regions,
450 clinopyroxene trace element concentrations (and zoning) reveal the magmatic processes that operate
451 within these crystal-rich domains. The trace element composition of melts in equilibrium with our
452 clinopyroxene crystals are calculated using the model of Wood and Blundy (1997). The results
453 indicate that many of our clinopyroxene analyses have incompatible trace element ratios (e.g.
454 $[Ce/Y]_n$) which are more enriched than any erupted basalt from Floreana (Harpp et al., 2014a). In fact,
455 nearly all clinopyroxene analyses from our xenolith samples, and ~50% of clinopyroxene analyses
456 from the scoria samples, record trace element disequilibrium with the typical composition of the
457 Floreana basalts (Fig. 7 and 9). Over-enriched trace element signatures are characteristic of Cluster 1
458 clinopyroxenes (i.e. chemical affinity to the wehrlitic or dunitic xenoliths), whereas crystals that are
459 near trace element equilibrium with Floreana basalts typically have Cluster 3 major element
460 compositions (i.e. the autocryst cluster).

461 Petrographic observations and olivine minor element data indicate that the Floreana sub-volcanic
462 system is characterised by clinopyroxene crystallisation within an olivine-dominated mush. If the
463 clinopyroxene grew from trapped melt within an olivine-dominated mush, progressive crystallisation
464 would increase the concentration of highly incompatible trace elements (e.g. Ba, La, Ce) relative to
465 less incompatible trace-elements (e.g. Sm, Y) in the residual melt. A simple fractional crystallisation
466 model indicates that ~70% crystallisation is required to generate melt $[Ce/Y]_n$ ratios that are in
467 equilibrium with enriched clinopyroxenes from the scoria samples and even greater extents of
468 crystallisation (~80%) would be required to generate the extremely high $[Ce/Y]_n$ ratios in some of the
469 xenolithic clinopyroxenes (Fig. 9).

470 Such extensive fractional crystallisation would be expected to result in the saturation and
471 crystallisation of plagioclase and other accessory phases (e.g. apatite, magnetite/ilmenite, quartz),
472 which are observed in more evolved xenoliths from Rabida island in the central Galápagos (Holness
473 et al., 2019). However, these phases are absent in the Floreana xenoliths, indicating that either

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474 infiltration of melts from a highly enriched mantle source component or a different magmatic process
475 within the cumulate mush is responsible for generating the anomalous trace element signatures of the
476 Floreana clinopyroxenes. Infiltration of mantle-derived melts that are more enriched than anything
477 observed in erupted basalts on Floreana is considered unlikely owing to the overlapping isotopic
478 composition of xenoliths and erupted lavas (Lyons et al., 2007), and the absence of a significant
479 correlation between trace element enrichment and radiogenic isotope compositions in the erupted
480 Floreana basalts (see Supplementary Information; Harpp et al., 2014a). As a result, a process other
481 than source heterogeneity must be responsible for generating the trace element variation in both the
482 Floreana xenoliths and erupted basalts.

483 One alternative mechanism that might be able to generate the observed trace element over-enrichment
484 is reactive porous flow. In this scenario, clinopyroxene is precipitated from ascending clinopyroxene-
485 saturated melts that continuously react with the existing, olivine-dominated crystal framework as the
486 system approaches equilibrium, and thus deviates from a simple fractional crystallisation trajectory
487 (Lissenberg and MacLeod, 2016). As a result, reactive porous flow in the olivine-dominated mush
488 beneath Floreana will likely result in substantial crystallisation of clinopyroxene, possibly at the
489 expense of pre-existing crystal phases, with little to no formation of olivine or saturation of minor
490 phases. Consequently, reactive porous flow can lead to enrichment of highly- to moderately-
491 incompatible trace elements in the resulting melt (Coogan et al., 2000; Gao et al., 2007; Lissenberg et
492 al., 2013; Lissenberg and MacLeod, 2016), and is consistent with the petrography of the Floreana
493 xenoliths. For example, major element maps of clinopyroxene crystals in the Floreana wehrlites show
494 that they are zoned, with Ti-rich rims (Fig. 10); equivalent zoning patterns have been attributed to
495 reactive porous flow in plutonic clinopyroxenes from the oceanic crust (e.g. Hess Deep; Lissenberg
496 and MacLeod, 2016). In addition, if pre-existing Cr-spinel was dissolved by the reacting melt, then
497 reactive porous flow could also explain the high Cr contents of clinopyroxene in our wehrlitic
498 xenoliths (Fig. 6; Lissenberg and MacLeod, 2016).

499 To test whether reactive porous flow of clinopyroxene saturated melts through an olivine-dominated
500 mush is consistent with the trace element compositions of melts calculated to be in equilibrium with
501 our Floreana clinopyroxenes, we use the zone refining model of Harris (1957; Fig. 9):

$$\frac{C_l}{C_l^0} = \frac{1}{D} - \left(\frac{1}{D} - 1\right)^{-DI}$$

503 (eq. 1)

504 where D is the bulk partition coefficient; C_l^0 and C_l are the initial and final concentration of that
505 element in the melt phase, respectively; and I is the 'equivalent volumes of solid processed by the
506 liquid' (Lissenberg and MacLeod, 2016). The model assumes continuous reaction of the percolating
507 melt front with the existing crystal framework, and has previously been employed to investigate
508 geochemical signatures in oceanic gabbros (Lissenberg and MacLeod, 2016). Results indicate that
509 reactive porous flow can produce melts with trace element compositions that are comparable with
510 those in equilibrium with our Floreana clinopyroxenes (i.e. $[\text{Ce}/\text{Y}]_n \sim 8-13$) using I values that are
511 similar to those invoked in other magmatic settings worldwide ($\sim 2-5$ compared to $\sim 4-8$ for the Hess
512 Deep; Lissenberg and MacLeod, 2016). Hence, reactive porous flow represents a realistic mechanism
513 for generating the geochemical diversity of Floreana clinopyroxenes, including the trace element
514 enriched crystals analysed in the wehrlitic nodules (Fig. 9).

515 In addition, detailed LA-ICP-MS transects of two clinopyroxene grains from the most enriched
516 wehrlitic xenolith analysed in this study (17MMSG02c) show clear trace element zoning (Fig. 11).
517 The core of the larger clinopyroxene crystal has low $[\text{Ce}]$ and $[\text{Ce}/\text{Y}]_n$ contents that are approximately
518 in equilibrium with Floreana basalts (Harpp et al. 2014a; Fig. 11a), whereas the mantle and rim of the
519 crystal is characterised by increasing $[\text{Ce}]$ and $[\text{Ce}/\text{Y}]_n$ contents. We interpret this as core
520 crystallisation from a melt with a trace element signature similar to that of erupted Floreana basalts
521 (Harpp et al., 2014a), followed by growth from a melt which became progressively enriched during
522 reactive porous flow (Fig. 11).

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523 The mantle of the second, smaller xenolithic clinopyroxene shows a similar rim-ward increase in [Ce]
524 (interpreted as progressive melt enrichment during reactive porous flow). However, the [Ce] and
525 [Ce/Y]_n values of the crystal core are too high to be in equilibrium with erupted Floreana basalts (Fig.
526 11C). This is consistent with our spot analyses of clinopyroxene cores in other crystals and samples.
527 The high [Ce] and [Ce/Y]_n values measured in crystal cores cannot be explained by inward diffusion
528 of Ce, owing to significant differences in the diffusivities of Ce and Y and similar [Ce] and [Y]
529 zoning patterns in our two crystal transects (Fig. 11; Van Orman, 2001). Instead, we suggest that the
530 high apparent core [Ce] and [Ce/Y]_n contents in many of the Floreana clinopyroxenes record
531 crystallisation from melts that had already undergone geochemical enrichment via reactive porous
532 flow. However, we cannot discount that our apparent clinopyroxene cores are fragments of larger
533 oikocrysts that have been broken during mush disaggregation or sample crushing and, as a result, do
534 not represent the true core compositions of each crystal.

535 Nevertheless, our clinopyroxene major and trace element data, as well as petrographic observations of
536 the wehrlitic xenoliths, provide substantial evidence that reactive porous flow is an important
537 mechanism of melt migration and melt differentiation in highly crystalline magma storage regions
538 beneath Floreana. Although reactive porous flow has been identified as an important process in MOR
539 gabbros, this is the first study to identify reactive porous flow in an ocean island setting.

540 **6.4 Petrographic estimates of magma storage pressures**

541 Petrological and geophysical constraints on magma storage depths exist for several recently active
542 volcanoes in the western Galápagos Archipelago (Bagnardi et al., 2013; Case et al., 1973; Geist et al.,
543 1998; Stock et al., 2018; Vigouroux et al., 2008). However, in the absence of geophysical data (owing
544 to a paucity of recent eruptions), there are far fewer constraints on the structure of magma storage
545 regions in the eastern and south-eastern archipelago. To date, the only investigation of magma storage
546 depths beneath these volcanoes is by Geist et al. (1998), who undertook a visual comparison between
547 whole-rock lava compositions and the MORB olivine + plagioclase + augite + melt pseudoinvariant
548 point, parameterised by Grove et al. (1992). This approach is subject to substantial uncertainty, but the

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549 authors suggest that the Floreana magmas consistently equilibrate at >5 kbar (typically >7 kbar) at a
550 depth >16 km, within the upper mantle.

551 We used three petrological barometers to provide improved constraints on magma storage depths
552 beneath Floreana. First, we applied the clinopyroxene-only barometer and thermometer of Putirka
553 (2008; equations 32b and 32d, respectively), in which pressure and temperature are solved iteratively
554 based solely on the clinopyroxene major element composition (primarily the jadeite component;
555 standard error of estimate [SEE] ± 310 MPa). Second, we applied the clinopyroxene-melt barometer of
556 Neave and Putirka (2017; equation 1), which uses the composition of a co-existing melt phase and the
557 proportion of the Jadeite component in clinopyroxene to calculate the pressure of crystallisation (SEE
558 = ± 140 MPa; pressure is solved iteratively with temperature using the clinopyroxene-melt
559 thermometer of Putirka, 2008; equation 33). Third, for the xenolithic nodules, we estimate the final
560 pressure and temperature of storage using the two-pyroxene thermobarometer of Putirka (2008;
561 equations 36 and 39; SEE = ± 280 MPa).

562 Taken at face value, initial application of the clinopyroxene-only barometer to all clinopyroxene
563 analyses from the scoria and xenolith samples gives a range of pressure estimates between ~ 450 MPa
564 and ~ 1800 MPa. However, reactive porous flow has a substantial influence on the compositions of the
565 Cluster 1 (and Cluster 2) clinopyroxenes, which may influence the barometric results. Specifically,
566 crystals that show evidence for reactive porous flow also have elevated Na concentrations, leading to
567 an anomalously high jadeite component and thus calculated pressure. This likely originates from the
568 presence of unusual melt compositions that fall outside the calibration range of the clinopyroxene-
569 only barometer owing to reactive melt migration, which is supported by the unusual glass
570 compositions in sample 17MMSG27 (Fig. 4). Therefore, we filter our dataset to remove crystals that
571 show a chemical signature indicative of reactive porous flow and only use Cluster 3 clinopyroxenes
572 that have trace element compositions in equilibrium with the Floreana basalts (using the whole-rock
573 data from Harpp et al. 2014) in our barometric calculations (n=78). Barometric results from this
574 filtered dataset indicate that crystallisation beneath Floreana occurs at a pressure of 766 ± 322 MPa

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575 (2σ of calculated pressures), which equates to a depth of 25.2 ± 9.9 km (using the crustal density
576 estimate of Putirka (1997) and a mantle density estimate of 3300 kg/m^3 ; Fig. 12).

577 Application of the Neave and Putirka (2017) clinopyroxene-melt barometer requires identification of
578 equilibrium clinopyroxene-liquid pairs. We achieve this using an automated melt-matching algorithm
579 (as in Winpenny and MacLennan, 2011, Neave and Putirka, 2017, Stock et al. 2018), with $K_D(\text{Fe-Mg})$,
580 diopside-hedenbergite, enstatite-ferrosillite and calcium Tschermak's equilibrium tests ($K_D(\text{Fe-Mg})$
581 within ± 0.03 other components within 2 SEE; Putirka, 1999, Putirka, 2008, Mollo et al., 2013). We
582 used the whole-rock data of Harpp et al. (2014a) and basaltic glass analyses from this study as
583 potential equilibrium liquids. Input crystal compositions were again filtered to remove analyses that
584 showed evidence of reactive porous flow (i.e. only Cluster 3 clinopyroxenes in trace element
585 equilibrium with the Floreana whole-rock were used). In total, 70 of the 78 input clinopyroxene
586 analyses returned at least one equilibrium match to either the basaltic glass or whole-rock
587 compositions. Where clinopyroxene compositions produced an equilibrium match with more than one
588 melt composition, an average melt composition was used in the barometric model. Results from this
589 barometer, coupled to the thermometer of Putirka (2008), indicate that magma crystallisation occurred
590 at 717 ± 165 MPa (23.7 ± 5.1 km) and $1224 \pm 33^\circ\text{C}$ (Fig. 12).

591 Clinopyroxene-orthopyroxene thermobarometry records the final storage conditions of the cumulate
592 xenoliths, rather than the crystallisation conditions of clinopyroxene autocrysts (orthopyroxene is only
593 found as an intercumulus phase). Temperature and pressure estimates were only calculated from
594 orthopyroxene-clinopyroxene pairs in wehrlite and dunite xenoliths that passed the $K_D(\text{Fe-Mg})$
595 equilibrium test of Putirka (2008; 1.09 ± 0.14). Results suggest that the cumulates were stored at
596 $\sim 975\text{--}1100^\circ\text{C}$ and $600\text{--}900$ MPa, with a mean storage pressure of 712 ± 200 MPa (23.7 ± 6.4 km; Fig.
597 12).

598 The depths of magma storage calculated from our three petrological barometers show an excellent
599 agreement within the model uncertainties. These new data provide robust evidence that magma
600 storage beneath Floreana occurs below the Moho (~ 16 km; Feighner and Richards, 1994), in the upper
601 mantle.

602 **7 IMPLICATIONS FOR MAGMATIC PLUMBING** 603 **SYSTEMS BENEATH LOW MELT FLUX OCEAN** 604 **ISLAND VOLCANOES**

605 Our new petrological and geochemical data show that magma storage beneath Floreana occurs in
606 mush-dominated regions of the upper mantle (Fig. 13). Mineral chemistry (such as low olivine Ca
607 concentrations and clinopyroxene major elements) reveal that a substantial portion of the erupted
608 crystal cargo is derived from disaggregated mush and wall rock material which has been entrained
609 into the ascending magmas. During ascent, magmas may entrain coherent nodules (xenoliths) as well
610 as disaggregated mush (Fig. 13). Coherent nodules represent areas of the magmatic system beneath
611 Floreana that have undergone cooling to temperatures $<1100^{\circ}\text{C}$ (compared to the clinopyroxene
612 crystallisation temperatures of $\sim 1225^{\circ}\text{C}$) and may represent material from the border of the active
613 mush zone or older, almost completely solidified magma storage regions that are intersected during
614 magma ascent (Fig. 13).

615 Petrographic observations and clinopyroxene trace element chemistry from both the xenolith and
616 scoria samples reveal that clinopyroxene growth occurs via reactive porous flow in the mush-
617 dominated areas beneath Floreana. Reactive porous flow causes distinct trace element enrichment in
618 the percolating melt phase and crystallising clinopyroxene, which can explain the trace element
619 disequilibrium between the erupted Floreana basalts and their clinopyroxene cargo. Nevertheless, the
620 presence of some clinopyroxene crystals with major and trace element compositions in equilibrium
621 with erupted Floreana basalts indicates that at least some crystallisation occurs in liquid-rich sub-
622 volcanic storage regions, likely situated as localised melt pockets within the larger mush zone (Fig.
623 13).

624 Transport of melts modified by reactive porous flow into these melt pockets could impact the LREE
625 enriched signature of the resultant hybridised melts. This could help to explain the unique, concave up
626 REE signature of the Floreana basalts, which is not seen in other regions of the Galápagos
627 Archipelago (Harpp et al., 2014a). However, it is important to note that similar trace element
628 signatures are not observed in other, low melt-flux regions of the eastern Galapagos (such as San

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629 Cristobal; Geist et al., 1986). As a result, we hypothesise that the unique REE pattern of the Floreana
630 basalts is primarily a source signature, likely associated with the highly radiogenic Pb and Sr isotope
631 signatures that characterise the Floreana basalts (Harpp et al., 2014a; Harpp and White, 2001).

632 Nevertheless, a few basalts on Floreana have trace element signatures that are far more enriched than
633 the majority of erupted basalts ($[La/Sm]_n$ up to 7.5), but are isotopically indistinguishable (Harpp et
634 al., 2014a); we therefore suggest that these basalts contain an anomalously large contribution of melts
635 that have undergone geochemical modification due to reactive porous flow.

636 Our results indicate substantial differences in the architecture of the magmatic systems beneath
637 Floreana and the frequently active shield volcanoes in the western Galápagos Archipelago. For
638 example, previous petrological and geophysical studies have identified that western Galápagos
639 magmatic systems are characterised by crustal magma storage, often with a large storage region in the
640 mid-to-lower crust (~7 km depth) and a smaller storage region at shallow levels, within the volcanic
641 edifice (~1 km depth; Geist et al. 1998; Bagnardi et al. 2013; Bagnardi and Hooper, 2018; Stock et al.,
642 2018; Fig 12). In contrast, our barometric data indicate that magmas beneath Floreana ascend directly
643 from the upper mantle and undergo no detectable crustal storage. In addition, although mush-rich
644 regions have been inferred beneath the western Galápagos shield volcanoes (based on whole-rock data
645 and the presence of gabbroic glomerocrysts; Chadwick et al., 2011; Geist et al., 1995, 2014; Stock et
646 al., 2018), magmatic differentiation appears to be driven by simple fractional crystallisation and
647 mixing of chemically diverse magmas (Geist et al., 1995; Naumann and Geist, 1999; Stock et al.,
648 2020).

649 One major factor that differentiates Floreana from shield volcanoes in the western archipelago is the
650 flux of magma into the lithosphere, which is evidenced by the large variations in the volumetric
651 eruption rate on Floreana and the western shields (Harpp et al., 2014a; Harpp and Geist, 2018; Kurz et
652 al., 2014). Hence, we suggest that the greater pressure of magma storage and prevalence of reactive
653 porous flow beneath Floreana, relative to volcanoes in the western archipelago, are related to the
654 substantially lower flux of magma into the lithosphere from the underlying mantle source (and thus
655 the thermal structure of the lithosphere). For example, the magma flux entering the lithosphere

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656 beneath Wolf volcano (northern Isabela) has been substantially greater than that beneath Floreana for
657 several 100,000s of years (Geist et al., 2005). The high magma flux beneath Wolf maintains the
658 average temperature of the mid-to-lower crust at $\sim 1125^{\circ}\text{C}$ ($\Delta T \sim 22^{\circ}\text{C}$), with only small-scale thermal
659 and compositional heterogeneities present in the sub-volcanic plumbing system (Geist et al., 2014,
660 2005; Stock et al., 2020, 2018). In contrast, the flux of magma entering the magmatic system beneath
661 Floreana is much lower and the temperature of the mid-crust is likely to be significantly cooler than
662 the lowest temperature recorded by the Floreana xenoliths (that is, $\ll 900^{\circ}\text{C}$; Fig. 12). As the flux of
663 magma (and heat) from the mantle is insufficient to maintain an elevated crustal geotherm beneath
664 Floreana, magmas that stall in the crust are likely to rapidly crystallise, increase their viscosity, and
665 become uneruptable. Therefore, eruptions must be fed by melts ascending from much deeper storage
666 regions ($\sim 700\text{-}750$ MPa) where melts can persist over long time periods.

667 Our results have global implications for the architecture and dynamics of magma storage regions
668 beneath ocean island volcanoes worldwide. The observed difference in magma storage depths beneath
669 high and low melt flux volcanic systems in the Galapagos Archipelago is mirrored in a global
670 compilation of barometric data from ocean island volcanoes (Famin et al., 2009; Geist et al., 1998;
671 Hammer et al., 2016; Hartley et al., 2018; Klügel et al., 2015; Poland et al., 2015; Stock et al., 2018;
672 Zanon et al., 2020; Zanon and Pimentel, 2015). Using the average repose period between eruptions at
673 a particular basaltic volcanic centre as a proxy for the flux of magma entering the lithosphere from the
674 underlying mantle (Global Volcanism Program, 2013), Figure 14 shows that the most frequently
675 active volcanic centres (such as, Kīlauea, Hawai'i, and Piton de la Fournaise, Réunion) are
676 characterised by persistent magma storage in the mid to upper crust (Famin et al., 2009; Poland et al.,
677 2015). In contrast, less active centres located above low buoyancy flux plumes (e.g. El Hierro, Canary
678 Islands) and/or peripheral to the main zone of plume upwelling (e.g. Haleakalā, Hawai'i) are
679 characterised by longer repose periods and correspondingly greater magma storage pressures
680 (Hammer et al., 2016; Klügel et al., 2015; Zanon and Pimentel, 2015). In fact, although secondary
681 crustal magma staging can occur (Klügel et al., 2015), the main zone of magma storage beneath ocean
682 island volcanoes with repose periods >50 years is typically in the lithospheric mantle, below the base

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683 of the crust (Longpre et al., 2014; Taracsák et al., 2019; Zanon et al., 2020). Hence, we speculate that
684 the flux of magma from the underlying mantle source has a first-order control on the depth of magma
685 storage beneath ocean island volcanoes and, correspondingly, high-pressure magma storage – as
686 observed beneath Floreana – is characteristic of low melt flux ocean island volcanoes globally.

687 **8 CONCLUSIONS**

688 Petrographic and geochemical analyses of lava, scoria and xenolith samples from Floreana in the
689 south-eastern Galápagos Archipelago provide new insights into the architecture and dynamics of
690 magma storage beneath low melt flux ocean island volcanoes. Comparison of olivine and
691 clinopyroxene major, minor and trace element contents between our different sample types reveal that
692 a substantial portion of the erupted crystal cargo is derived from mush-dominated magma storage
693 regions beneath Floreana. Mineral textures, highly enriched clinopyroxene trace element signatures
694 and trace element zoning in the xenoliths reveals that reactive porous flow is an important process of
695 chemical differentiation and melt transport within these mush-dominated regions. Mixing between
696 melts that have been geochemically enriched by reactive porous flow and those in overlying liquid-
697 rich storage regions could contribute to the anomalous, concave-up REE signature of the Floreana
698 basalts, which is absent in other parts of the Galápagos Archipelago where reactive porous flow has
699 not been identified.

700 Application of independent petrological barometers to crystals in Floreana scoria and xenolith
701 samples indicates that magmas are stored in the upper mantle ($\sim 23.7 \pm 5.1$ km). Floreana is in a distal
702 location to the Galápagos plume where the melt flux entering the lithosphere is low; the depth of
703 magma storage beneath Floreana contrasts with more proximal, higher melt flux volcanoes in the
704 western archipelago where magmas are stored in the crust (Geist et al., 1998; Stock et al., 2018).
705 Comparing our new data with ocean island volcanoes globally (e.g. Hawai'i, Iceland and the Canary
706 Islands) reveals that the Galápagos is not unique and that magma storage is ubiquitously shallower in
707 proximal magmatic systems above high buoyancy flux plumes than in off-axis systems, or above low
708 buoyancy flux plumes. We therefore suggest that the flux of mantle-derived magma entering the

709 lithosphere imparts a first-order control on the depth of magma storage beneath ocean island
710 volcanoes.

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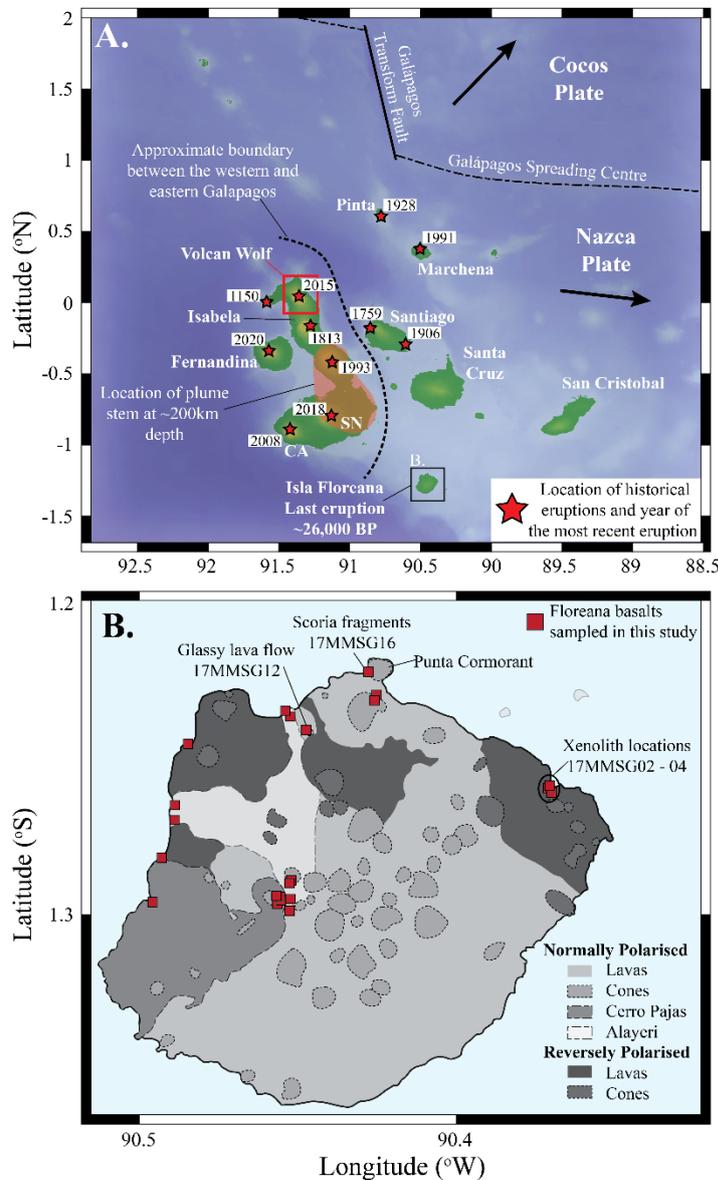
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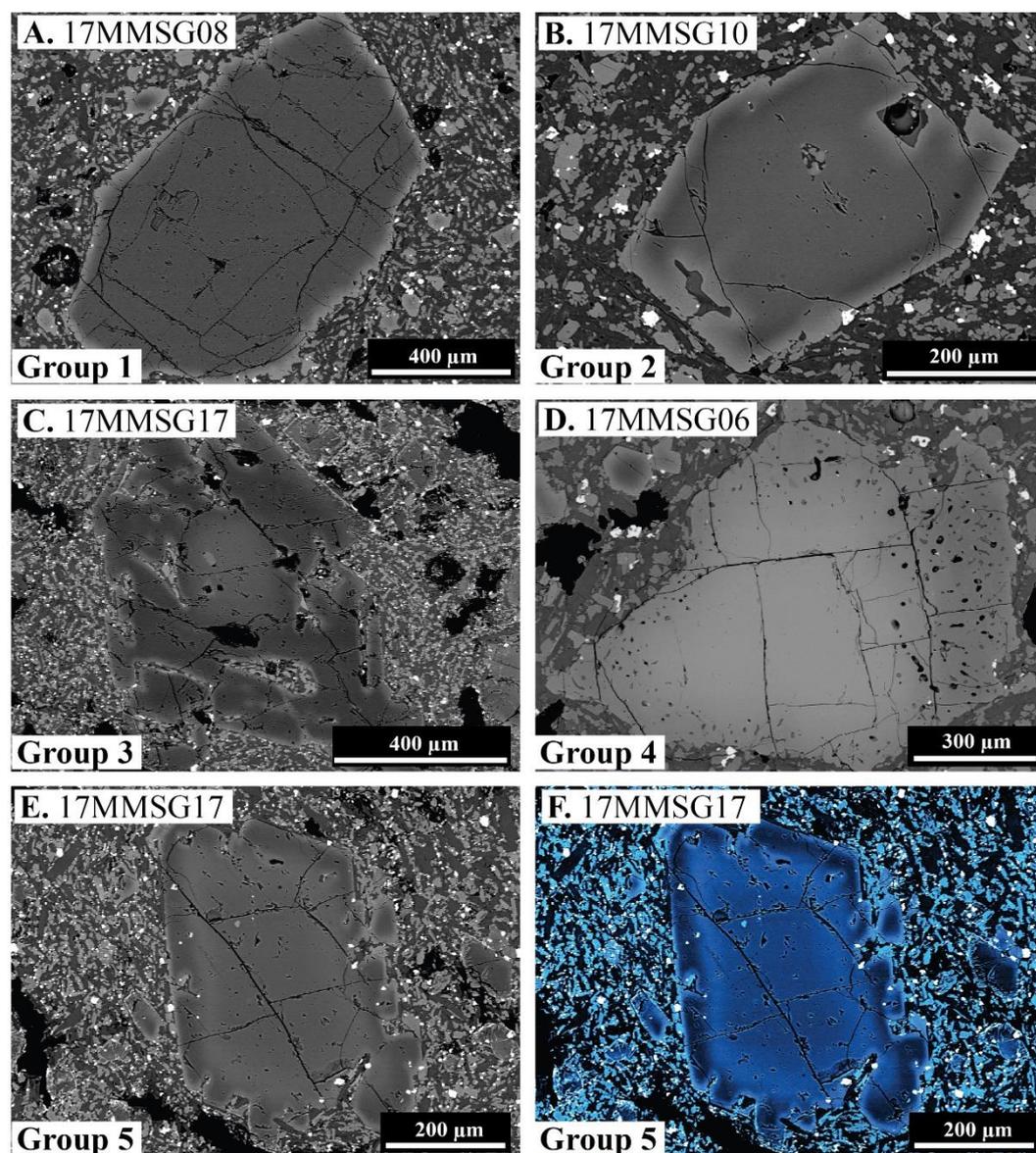
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FIGURES



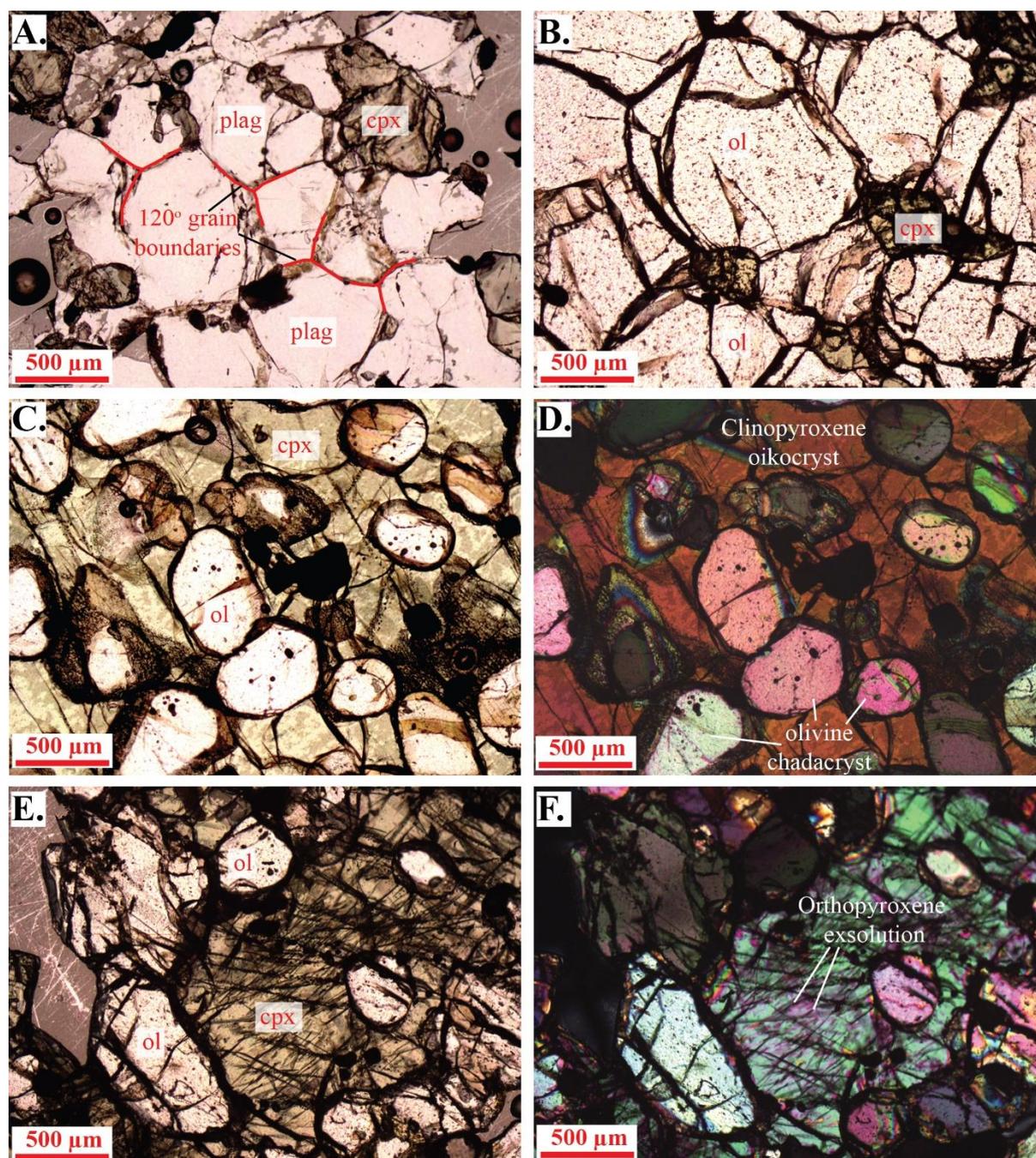
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1031 **Figure 1 - A.** Regional map of the Galápagos Archipelago highlighting the location of Isla Floreana,
1032 Cerro Azul (CA), Sierra Negra (SN) and Wolf volcanoes. Dates show the most recent eruptions at
1033 historically active volcanic centres. Black arrows show the direction of plate motion for the Nazca and
1034 Cocos tectonic plates. **B.** Geological map of Floreana adapted from Harpp et al. (2014a). Dashed lines
1035 delineate monogenetic scoria cones. Normally and reversely polarised lava flows are shown along
1036 with the largest (Cerro Pajas) and most recent (Alayeri; ~26,000 years) eruptions on the island.



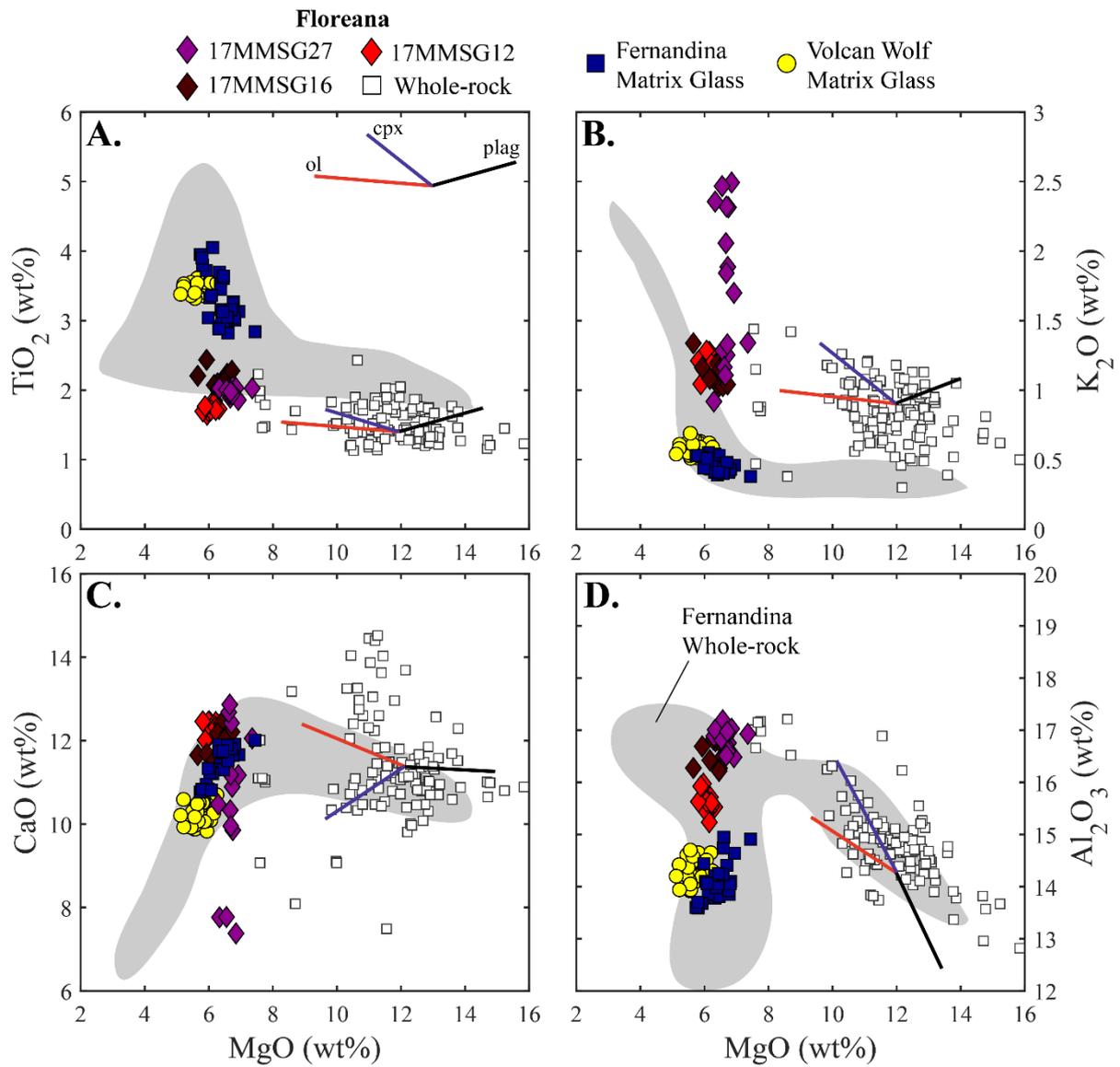
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1038 **Figure 2** – Backscatter Electron images. **A.** Group 1 olivines – euhedral to subhedral crystal
1039 morphologies with large, unzoned, crystal cores and narrow, normally-zoned rims. **B.** Group 2
1040 olivines – subhedral to euhedral crystals with clear, reverse-zoning profiles. **C.** Group 3 olivines –
1041 skeletal crystals with high forsterite overgrowths on low forsterite cores. **D.** Group 4 olivines –
1042 anhedral crystals with sieved textured, reverse zoned rims. **E.** (greyscale) and **F.** (false colour) Group
1043 5 olivines – crystals preserve at least 4 composition zones over ~100-200 μm. False colour image (**F.**)
1044 is used to highlight the compositional zoning of the Group 5 olivine, with the intensity of the blue
1045 colour associated with the Fo composition of the crystal (darker = higher Fo).



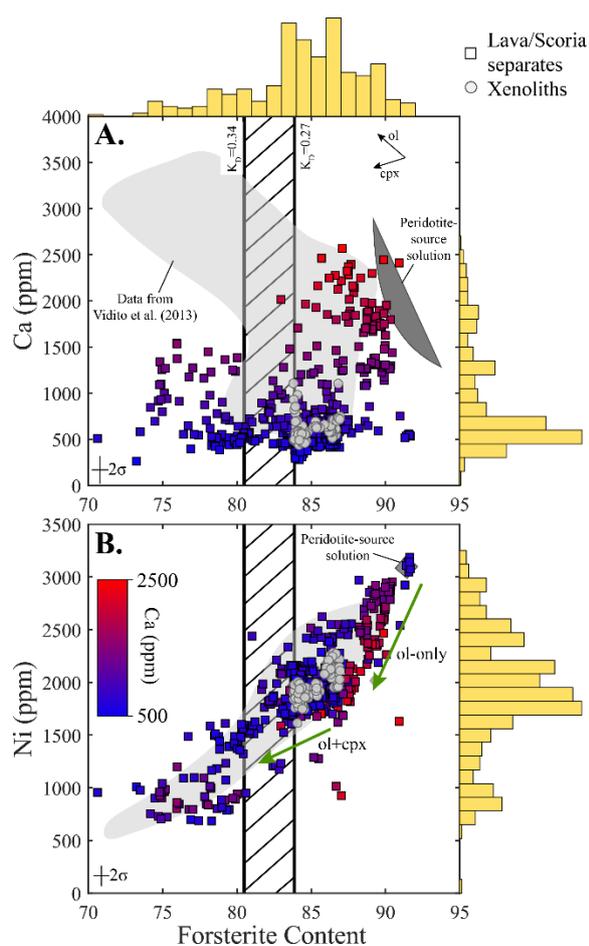
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1047 **Figure 3** - Plane Polarised Light (A. – C. and E.) and Crossed Polarised Light (D., F.) images of
1048 Floreana xenoliths. **A.** – gabbroic xenolith (17MMSG04b), highlighting near 120° grain boundaries at
1049 monomineralic plagioclase (plag) triple junctions. **B.** – dunitic xenolith (17MMSG04c) with
1050 intercumulus clinopyroxene (cpx). **C.** and **D.** – wehrlitic xenolith (17MMSG02c) showing a large
1051 clinopyroxene oikocryst surrounding olivine (ol) chadacrysts. **E.** and **F.** – wehrlitic xenolith (sample
1052 17MMSG03a) showing olivine chadacrysts within a clinopyroxene oikocryst. Orthopyroxene
1053 exsolution lamellae are visible within the clinopyroxene.



1054

1055 **Figure 4** – Major element compositions of matrix glasses (this study) and whole-rocks (Harpp et al.,
1056 2014a) from Floreana, as well as glasses from Fernandina (Peterson et al., 2017) and Wolf volcano
1057 (Stock et al., 2018) in the western Galápagos Archipelago. Lines show trajectories of liquid
1058 compositional evolution for olivine (ol; red), clinopyroxene (cpx; blue) and plagioclase (plag; black)
1059 crystallisation. The grey field shows whole-rock data from Isla Fernandina in the western Galápagos
1060 (Allan and Simkin, 2000; Geist et al., 2006). The 2 σ precision of our matrix glass analyses is smaller
1061 than the symbol size.



1062

1063 **Figure 5** – Major and minor element compositions of olivine crystals from the Isla Floreana basalts.

1064 **A.** Fo vs. Ca and **B** Fo vs. Ni in Galápagos olivine crystals displaying analyses from our lava/scoria

1065 separates and xenolith samples, as well as a compilation of available olivine data from Floreana

1066 (Vidito et al. 2013). Our lava/scoria analyses are coloured according to their Ca concentration. The

1067 histograms above and to the right of the plots show the data distributions (excluding *in situ* analyses

1068 of xenolithic olivines). Peridotite source solutions are taken from Herzberg (2011) and Matzen et al.

1069 (2017a). Black arrows in **A.** show the trajectory of crystal compositional evolution during olivine (ol)

1070 and clinopyroxene (cpx) crystallisation (taken from Gleeson and Gibson, 2019). The green lines in **B.**

1071 show the trajectories of crystal compositional evolution during olivine only and olivine +

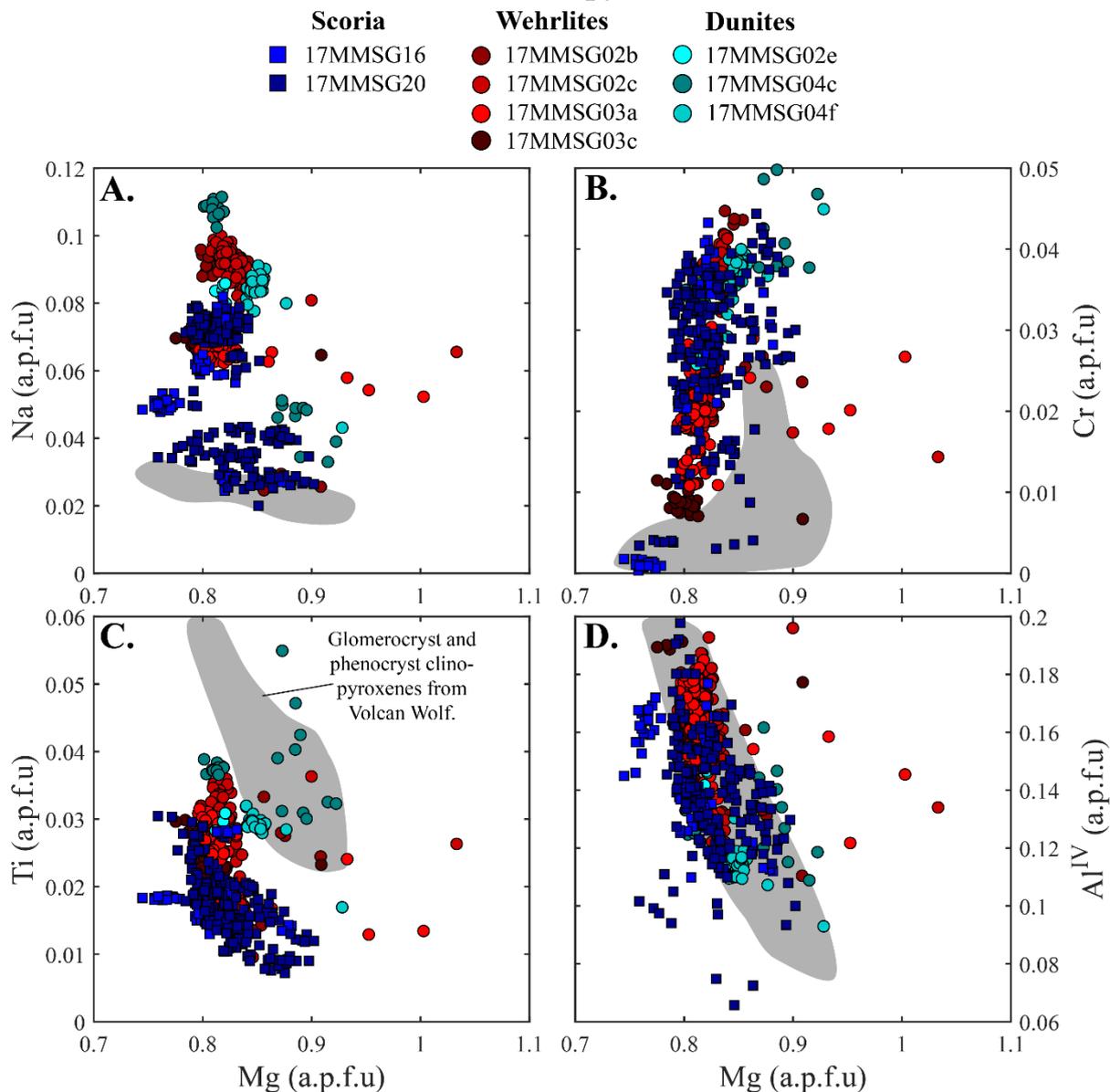
1072 clinopyroxene fractional crystallisation (from Gleeson and Gibson, 2019). The vertical black lines

1073 show the forsterite compositions of crystals calculated to be in equilibrium with the matrix glass

1074 composition of tephra sample 17MMSG16 ($K_d = 0.27 - 0.34$ after Matzen et al. 2011, and Roeder and

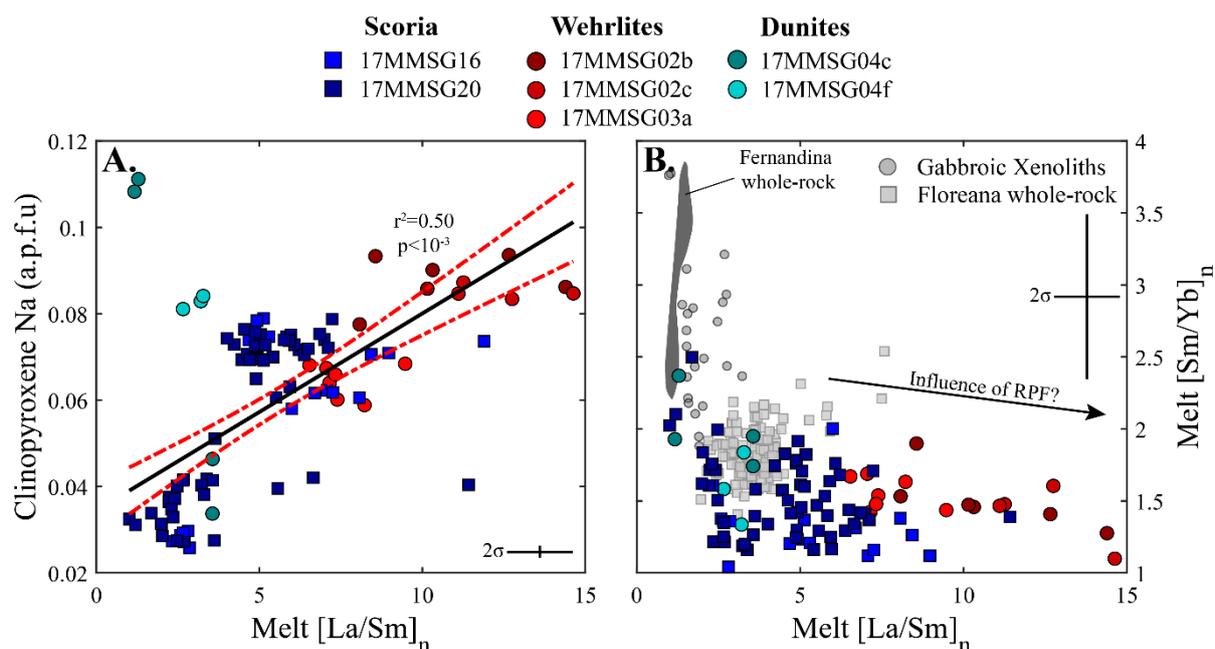
1075 Emslie 1970).

Floreana Clinopyroxenes



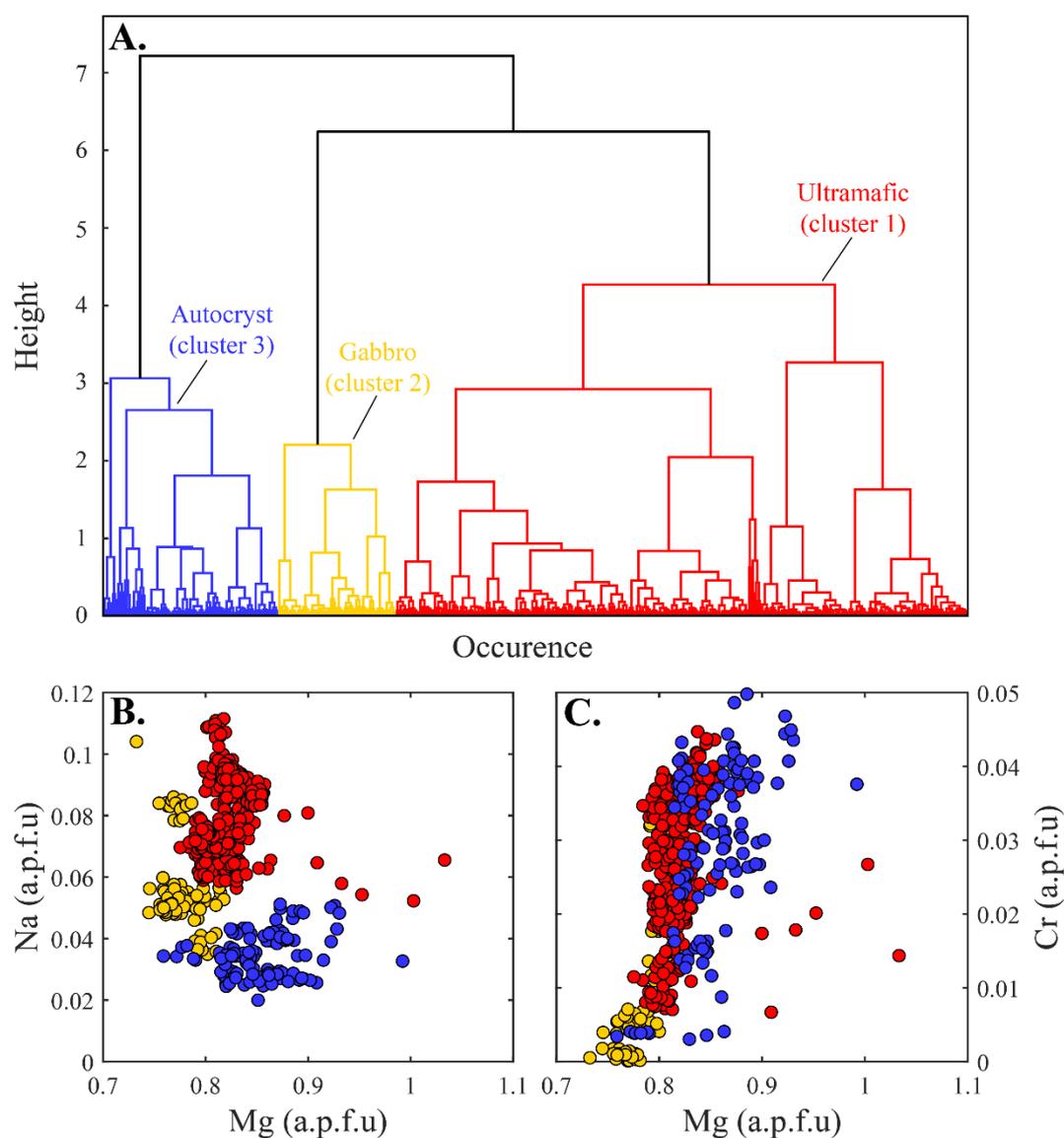
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1077 **Figure 6** – Major element composition of Floreana clinopyroxenes from our scoria, wehrlite and
1078 dunite samples. Data is shown as atoms per formula unit (a.p.f.u) on the basis of 6 oxygens. The grey
1079 field shows the compositions of clinopyroxenes from Wolf volcano in the western Galápagos
1080 Archipelago (from Stock et al. 2018). The 2σ precision of our clinopyroxene analyses is smaller than
1081 the symbol size.



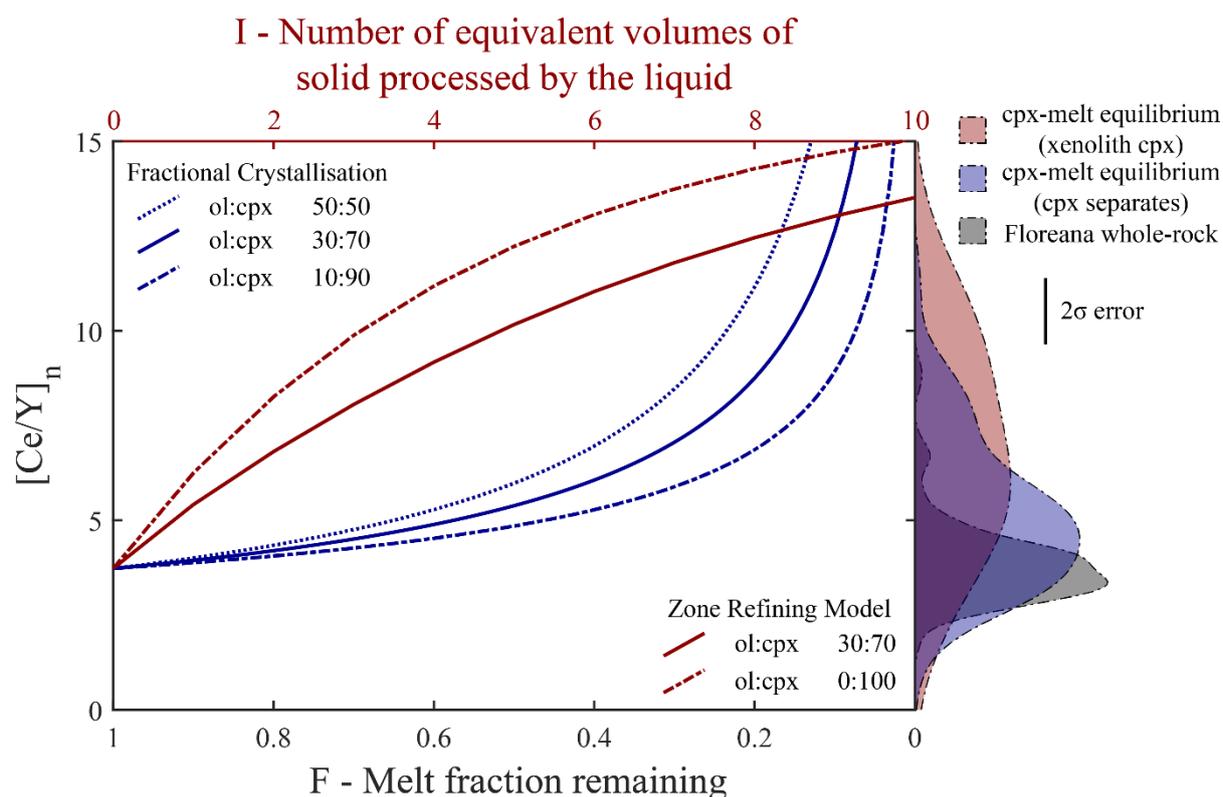
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1083 **Figure 7** – Trace element composition of the Floreana clinopyroxenes and their relation to major
 1084 element systematics. **A.** [La/Sm]_n vs. Na in clinopyroxenes from our scoria samples and wehrlite and
 1085 dunite xenoliths ([La/Sm]_n is shown as the composition of melt in equilibrium with each
 1086 clinopyroxene). The black line shows a regression through the data ($r^2 = 0.50$) and the red dashed
 1087 lines show the 95% confidence limits on the regression. The correlation indicates that trace element
 1088 enrichment in the Floreana clinopyroxenes is associated with anomalously high Na contents. **B.**
 1089 [La/Sm]_n vs. [Sm/Yb]_n of melts calculated to be in equilibrium with our Floreana clinopyroxenes
 1090 using the model of Wood and Blundy (1997). The black arrow shows the approximate trend of crystal
 1091 compositional evolution hypothesised to occur as a result of clinopyroxene crystallisation during
 1092 reactive porous flow within the cumulate mush (RPF). The grey field shows whole-rock compositions
 1093 from Fernandina (Geist et al., 2006; White et al., 1993). **B.** additionally shows the whole-rock
 1094 compositions of erupted Floreana lavas (Harpp et al., 2014a) and analyses of the gabbroic xenoliths
 1095 from Floreana (this study). Error bars show the fully propagated 2σ precision of our analyses.



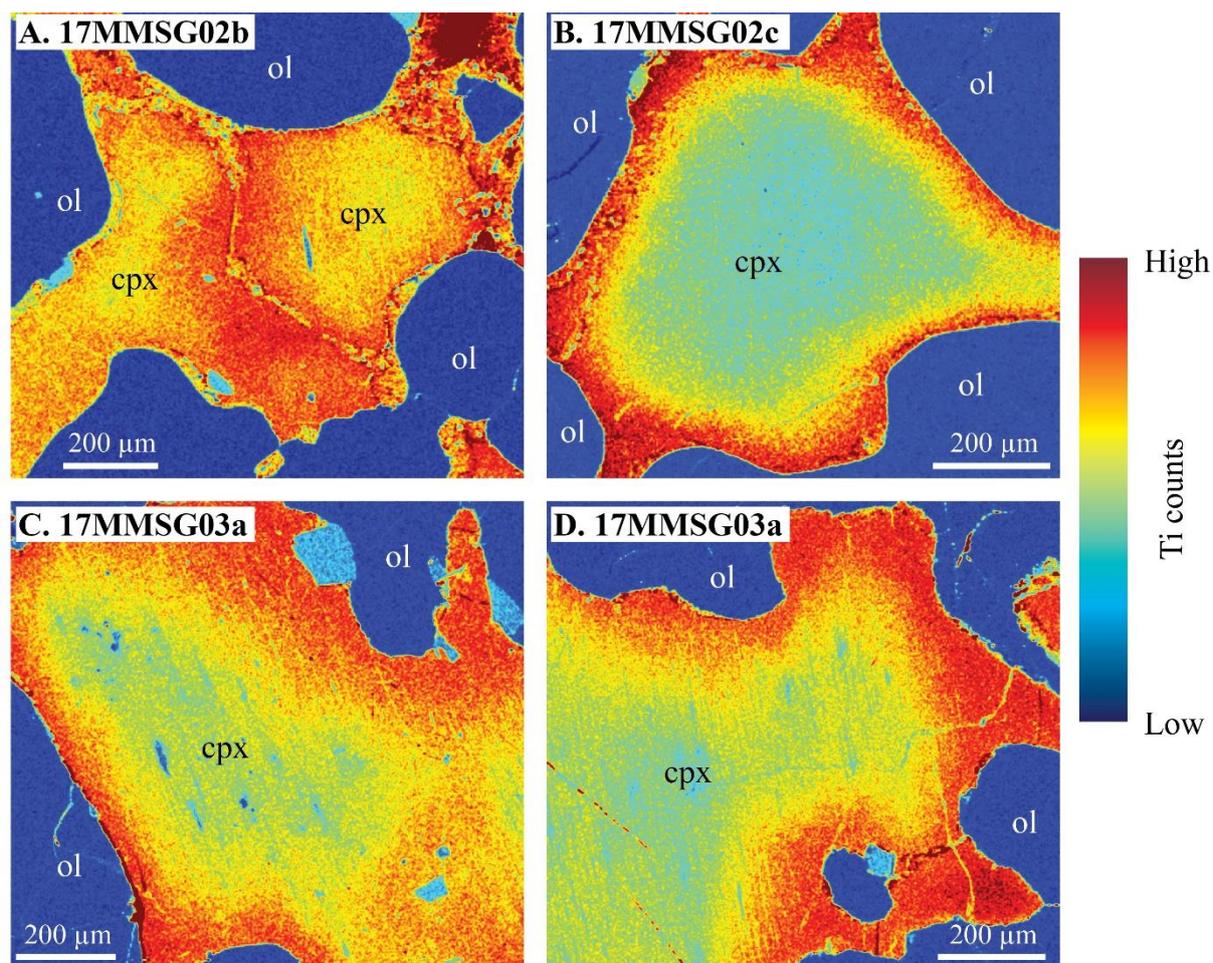
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1097 **Figure 8 – A.** Hierarchical cluster analysis of our clinopyroxene major element analyses. Cluster
1098 analysis was performed using Ward's method, which is built into the scikit-learn package in Python
1099 (Pedregosa et al., 2011; Ward, 1963). Height above the x-axis is a measure of the distance of
1100 separation of two clusters (i.e., the higher the join the more chemically distinct two clusters are).
1101 Colours show the high-level division of crystal compositions into three clusters: Cluster 1 is
1102 predominantly comprised of crystals from wehrlite and dunite xenoliths (red), Cluster 2 is composed
1103 of crystals from gabbroic xenoliths (yellow) and Cluster 3 (blue) is dominated by crystals separated
1104 from scoria samples. **B.** Na vs Mg and **C.** Cr vs Mg in our clinopyroxene analyses from the scoria and
1105 xenoliths, coloured by their cluster.



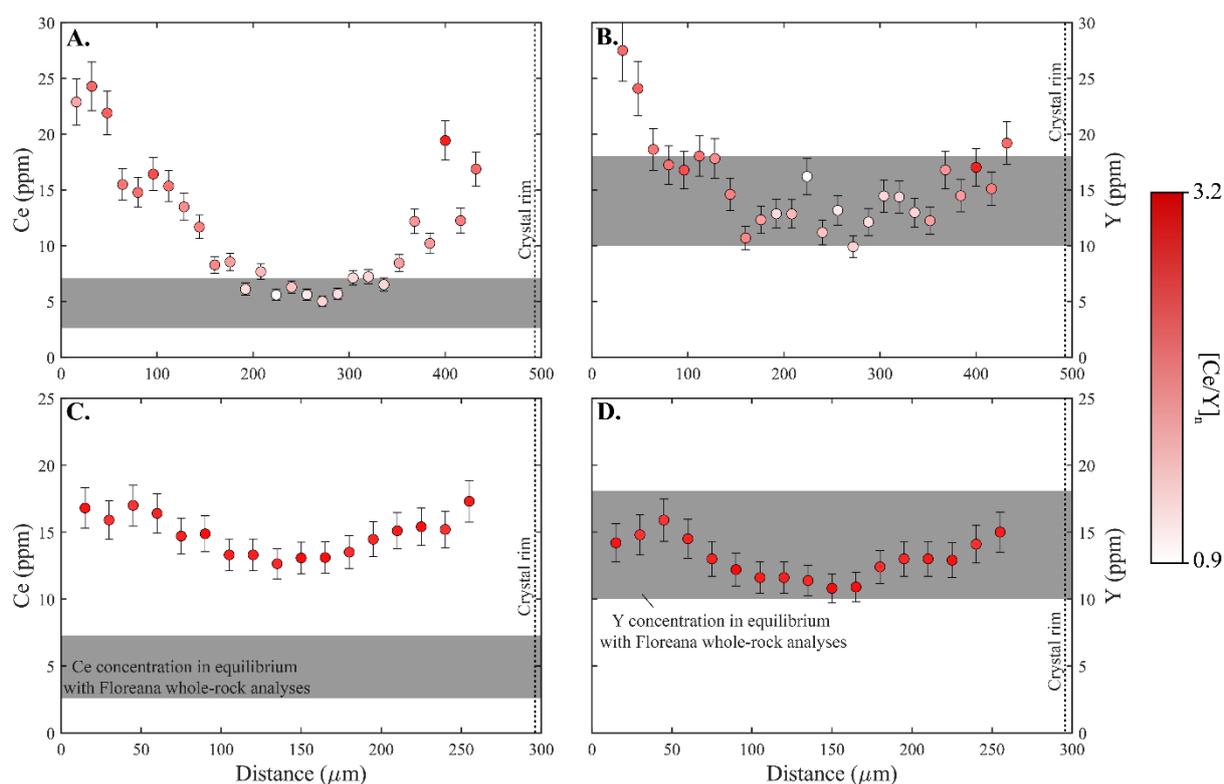
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1107 **Figure 9** – Models showing the evolution of melt $[Ce/Y]_n$ ratios as a function of the melt fraction
 1108 remaining (F) during fractional crystallisation and as a function of the number of equivalent volumes
 1109 of solid processed by the liquid (I) during reactive porous flow (red; using the zone-refining model of
 1110 Harris, 1957). Line types show how models vary with different ratios of crystallising olivine (ol) and
 1111 clinopyroxene (cpx). The kernel density plots show the $[Ce/Y]_n$ density distributions of liquids
 1112 calculated to be in equilibrium with clinopyroxene crystals in our wehrlite and dunite xenoliths (red),
 1113 scoria samples (blue; 2σ error for the clinopyroxene analyses is shown by the black bar), and whole-
 1114 rock analyses of erupted Floreana basalts (grey; Harpp et al., 2014a). Clinopyroxene partition
 1115 coefficients were calculated using the method of Wood and Blundy (1997), the mean major element
 1116 composition of the Cluster 3 clinopyroxenes from this study, a temperature of 1225°C, and a pressure
 1117 of 700 MPa. Initial Ce and Y concentrations for both models were taken as the mean values from the
 1118 whole-rock dataset of Harpp et al. (2014a). Calculation increments are 0.01 in F and 0.1 in I.



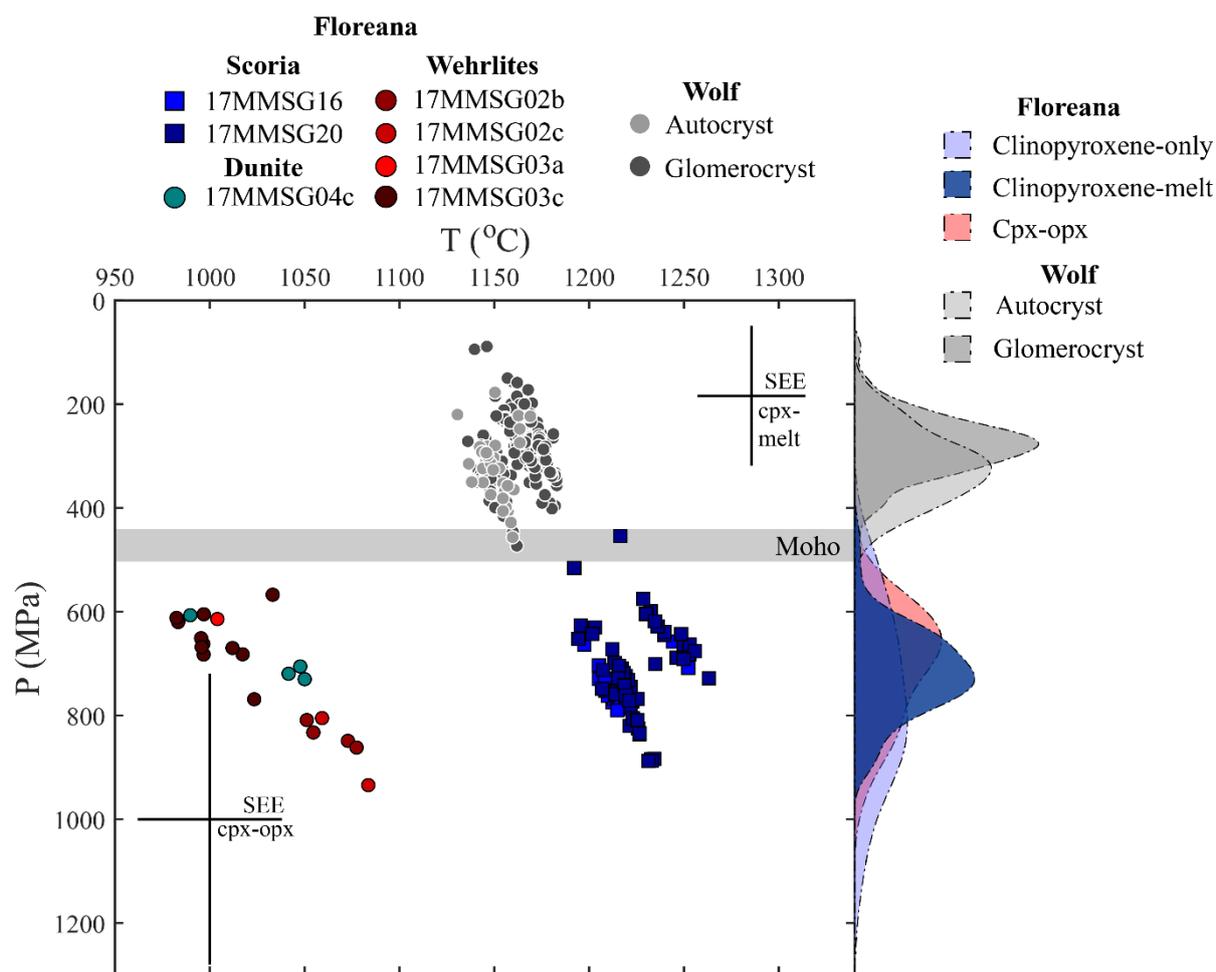
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1120 **Figure 10** – WDS maps of Ti counts in key clinopyroxene crystals from the Floreana wehrlite
1121 xenoliths. These maps display clear zoning in the xenolithic clinopyroxenes with Ti-poor cores and
1122 Ti-rich rims. EPMA transects across some of the wehrlitic clinopyroxenes indicates that the Ti
1123 concentrations in these crystals may vary from ~0.6 wt% to ~1.3 wt% TiO₂ (see Supplementary
1124 Information). The dark blue regions surrounding clinopyroxene grains are olivine crystals.



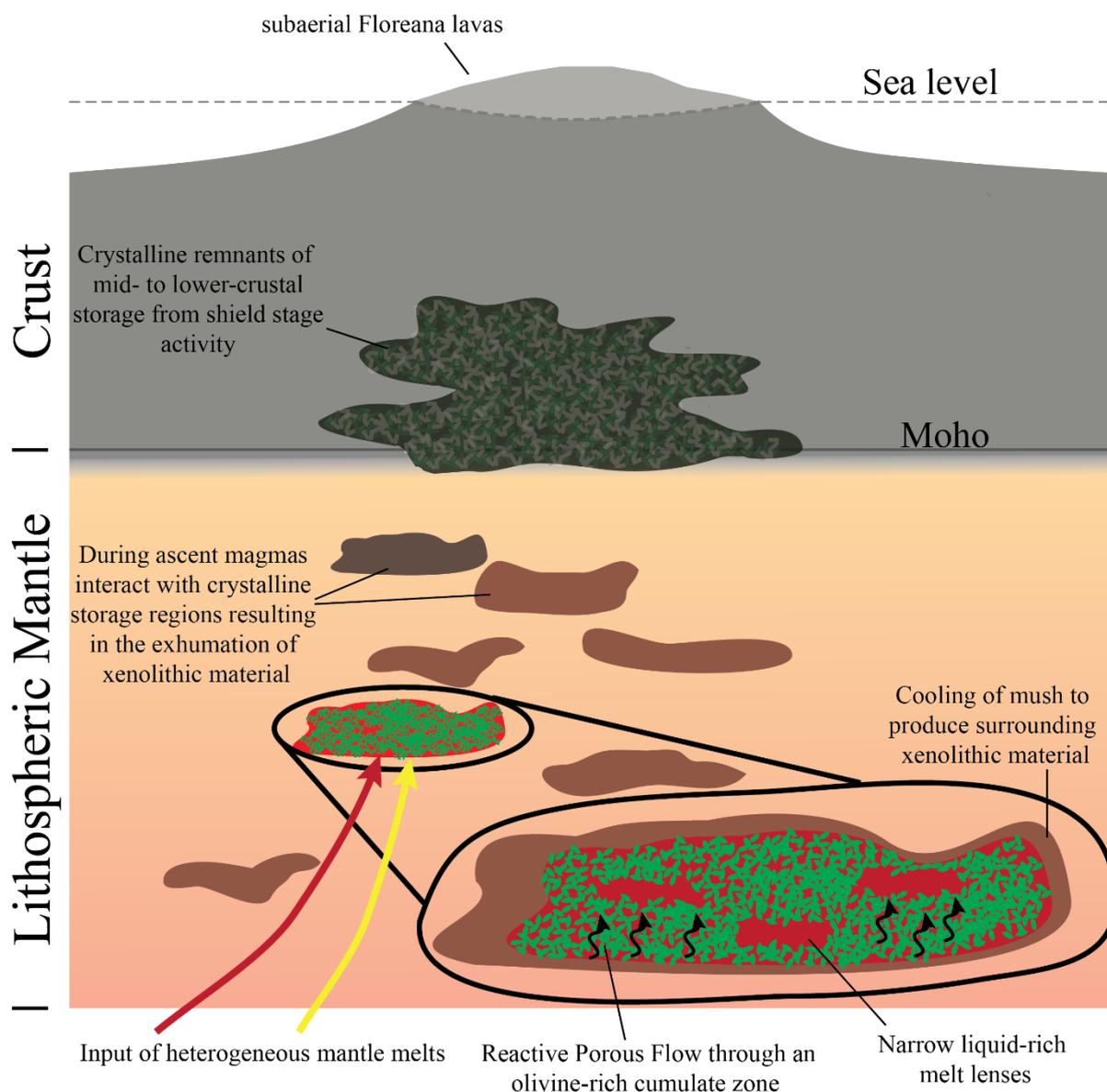
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1126 **Figure 11** – LA-ICP-MS transects showing Ce (**A, C**) and Y (**B, D**) zoning across two clinopyroxene
1127 grains in a wehrlitic xenolith (17MMSG02c). **A.** and **B.** show a transect across the core of a large
1128 clinopyroxene oikocryst. **C.** and **D.** show a transect across the apparent core of a smaller
1129 clinopyroxene oikocryst. Points are coloured according to their $[\text{Ce}/\text{Y}]_n$ ratio (see colour scale). The
1130 grey bars show the crystal compositions calculated to be in equilibrium with whole-rock analyses of
1131 erupted Floreana basalts (Harpp et al., 2014a).



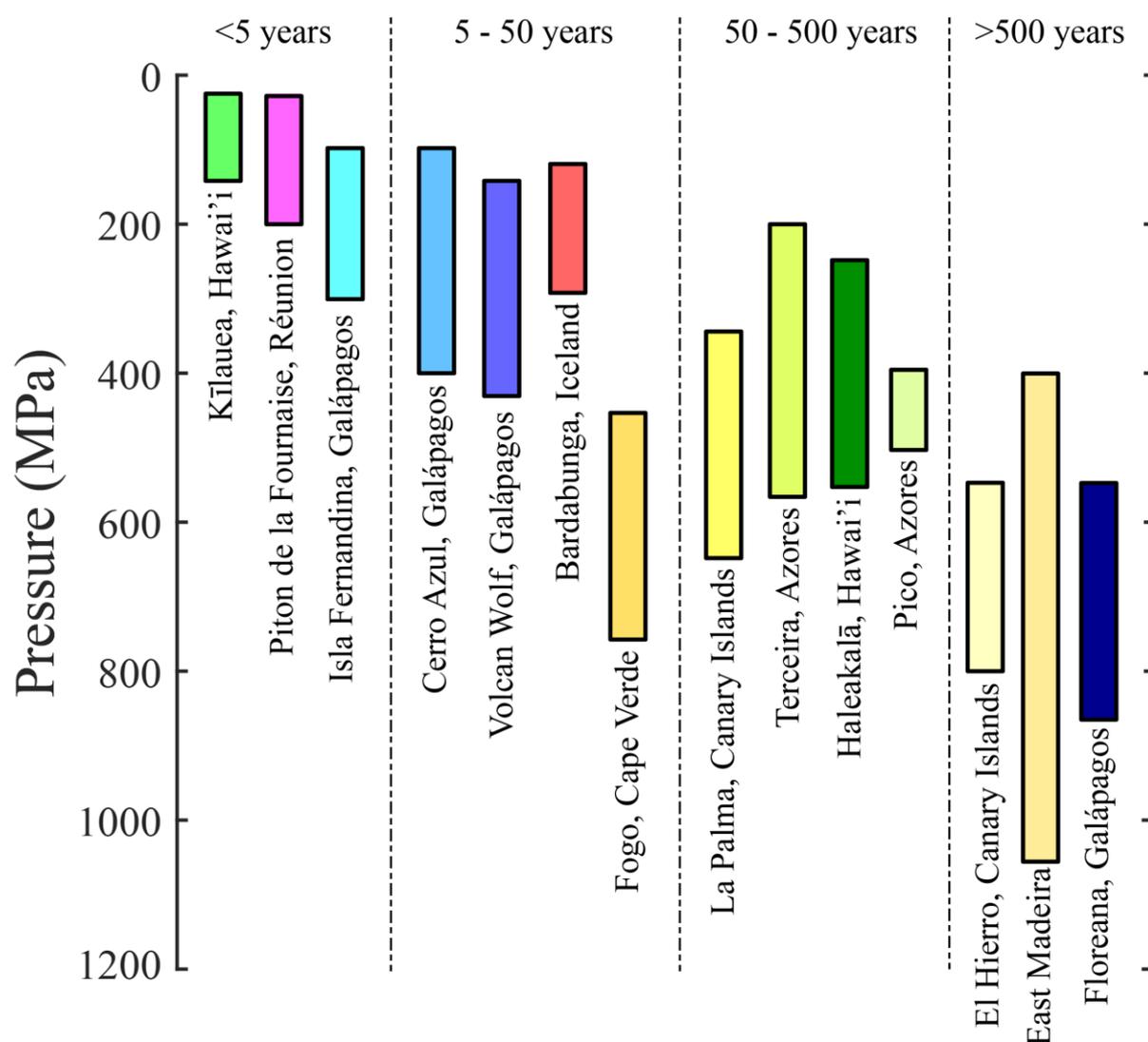
1132

1133 **Figure 12** – Petrological thermobarometry results. The kernel density plots to the right show the
 1134 density distributions of barometric results from different models (light blue – clinopyroxene-only,
 1135 Putirka (2008); dark blue – clinopyroxene-melt, Neave and Putirka (2017); red – clinopyroxene-
 1136 orthopyroxene Putirka (2008)). The grey bar shows the Moho depth beneath Fernandina (from
 1137 Feighner and Richards, 1994) and the grey points and kernel density estimates show clinopyroxene-
 1138 melt thermobarometric results for autocrysts and glomerocrysts from Wolf volcano (from Stock et al.,
 1139 2018). The Standard Estimated Error (SEE) of the clinopyroxene-melt and orthopyroxene-
 1140 clinopyroxene thermobarometers are given.



1141

1142 **Figure 13** – Schematic illustration of the magma plumbing system beneath Floreana. No magma
1143 storage is identified within the crust. Instead, our barometric results indicate that Floreana magmas
1144 ascend directly from the upper mantle, where they are stored at a depth of $\sim 23.7 \pm 5.1$ km. Floreana
1145 magma storage regions are dominated by crystal-rich domains (i.e. mush). Reactive porous flow is
1146 identified as an important mechanism of melt migration and magma differentiation in the crystal-rich
1147 storage regions, although our results shows that some crystallisation occurs within liquid-rich
1148 domains.



1149

1150 **Figure 14** - Compilation of barometric estimates for the primary magma storage region beneath
 1151 various ocean island volcanoes worldwide. Data is divided according to the approximate repose
 1152 period between eruptions at each volcano (estimated using data from Global Volcanism Program,
 1153 2013), which is used as a proxy for the flux of magma entering the lithosphere. A general trend to
 1154 greater magma storage pressures is observed with increasing repose period, indicating that the flux of
 1155 magma from the mantle has a first order control on the depth of magma storage. Data from (Famin et
 1156 al., 2009; Geist et al., 1998; Hammer et al., 2016; Hartley et al., 2018; Klügel et al., 2015; Poland et
 1157 al., 2015; Stock et al., 2018; Zanon et al., 2020; Zanon and Pimentel, 2015).