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1 **Imaging the seismic velocity structure of the crust and upper mantle in the northern East**
2 **African Rift using Rayleigh wave tomography**

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14 Key Points:

- 15 • The Ethiopian Plateau has slow velocities in the crust that likely require melt. At mantle
16 depths, off rift velocities do not require melt.
- 17 • Beneath the rift, mantle anomalies are punctuated along the rift and offset from the
18 slowest crustal anomalies.
- 19 • This suggests that either melt migrates laterally and/ or is ephemeral in a dynamic
20 lithosphere-asthenosphere system.

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23 **Summary**

24 Within the northern East African Rift, multiple seismic models have been produced to
25 understand the evolution of magmatism, however variations in method, resolution, and scale
26 make direct comparisons challenging. The lack of instrumentation off rift further limits our

1 understanding of the spatial extent of tectonic and magmatic processes, which is crucial to
2 understanding magmatic continental rifting. In this paper, we jointly invert Rayleigh wave
3 dispersion curves from ambient noise and teleseisms to obtain absolute shear velocity maps at
4 10–150 km depth. This includes data from a new seismic network located on the Ethiopian
5 Plateau and enhanced resolution at Moho and upper mantle depths from the joint inversion. At
6 crustal depths, velocities are slowest beneath the Main Ethiopian Rift and the off rift Ethiopian
7 Plateau ($<3.00\text{--}3.75 \pm 0.04$ km/s, 10–40 km depth), which are slow enough to require ongoing
8 magmatic emplacement. At 60–80 km depth off rift, we observe a fast velocity lid (>0.1 km/s
9 faster than surroundings), which corresponds to previous estimates of the lithosphere-
10 asthenosphere-boundary. The fast lid is not observed within the rift in locations underlain by
11 asthenospheric slow velocity anomalies ($<4.05 \pm 0.04$ km/s at 60–120 km depth), suggesting
12 melt is infiltrating the lithosphere within the rift. Furthermore, the asthenospheric slow velocity
13 anomalies are segmented ($\sim 110 \times 80$ km wide), existing in areas that have not undergone
14 significant crustal and plate thinning, suggesting segmented melt supply starts prior to
15 significant plate deformation. Finally, the segmented asthenospheric slow velocity zones are
16 not directly located beneath melt-rich crustal regions particularly for those off rift, suggesting
17 mantle melt either migrates laterally during ascent, and/or that melt is ephemeral.

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19 **Key words:** Seismology - Joint inversion - Seismic tomography - Africa - Surface waves and
20 free oscillations

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1.0 Introduction

Magma injected into the continental lithospheric crust and mantle, and the subsequent weakening and thinning, has been proposed as a mechanism for continental rifting, especially in regions such as the northern East African Rift (EAR) (Buck, 2006). Models and observations of magmatic continental rifting suggest there is melt production and localised emplacement beneath the rift valley (Buck, 2006). However, the locus, depth and along rift variability of melt production and storage remains poorly understood, as does the possibility that magmatic processes occur over a broader region outside the rift valley. Understanding these processes is essential for a full understanding of magmatic continental rifting. The northern EAR is a magmatic rift system where the initial stages of rifting through to incipient seafloor spreading can be observed subaerially (Ebinger & Casey, 2001; Furman et al., 2006). This provides a unique opportunity to study how rifting develops and modifies plate structure through progressive rift sector development (Barberi et al., 1972). A major barrier to understanding magmatic processes subsurface has been that the majority of passive seismic stations were deployed near the rift valley only, preventing comparison to off rift structures. In addition, despite the numerous seismic images created, absolute seismic velocities for the crust and upper mantle are unconstrained beneath a broad region including multiple sectors of the rift and to significant distances off rift (Chambers et al., 2019; Keranen et al., 2009).

Seismic tomography provides constraints on the elastic properties of the Earth, which are used to understand tectonic and magmatic processes. The rift has been extensively imaged, creating multiple velocity models for smaller disparate sections of the region, which are not necessarily directly comparable due to variations in methodology, resolution, and scale. For example, ambient noise tomography (ANT) has been used to produce 3 models of the northern EAR that

1 are not directly comparable due to variations in the type of velocity (phase, group and shear
2 velocity) and regions covered (Afar, central Main Ethiopian Rift and broader rifted regions)
3 (Chambers et al., 2019; Kim et al., 2012; Korostelev et al., 2015). Comparisons between
4 models are important for determining the relationship between magmatic and tectonic
5 processes including variations in plate structure, melt generation and melt migration processes.
6 These comparisons are not possible without a single absolute velocity model that covers the
7 full region. In addition, the rift flanks have not been tomographically imaged at crustal and
8 upper mantle depths due to a lack of instrumentation.

9
10 To solve the challenges in comparing multiple models, we produce a comprehensive model by
11 performing a joint inversion between Rayleigh wave dispersion curves from ambient noise and
12 teleseisms to image the absolute seismic shear velocity structure of the nEAR. Our unified
13 model of absolute shear velocity from 10–150 km depth, allows us to determine variations in
14 plate structure both within and off rift. Data is obtained from seismometers present from 1999-
15 2017 located in Ethiopia, Yemen, Djibouti and Eritrea, allowing coverage of regions unaffected
16 by rifting (e.g. west of the Ethiopian Plateau) to areas at incipient seafloor spreading (e.g. Afar).
17 The joint inversion provides increased resolution at lower crustal and uppermost mantle depths
18 which would not be possible with ambient noise or teleseisms individually (Chambers et al.,
19 2019; Gallacher et al., 2016). Furthermore, the addition of the Ethiopian Plateau network
20 allows us to image the transition from unrifted lithosphere to the rift, and obtain a view of the
21 asthenospheric mantle dynamics of rifting at depth. In particular the larger area with increased
22 resolution and improved crustal resolution allows us to compare the presence/absence of slow
23 velocity anomalies likely related to melt in the asthenosphere to those in the crust in a consistent
24 manner to better constrain the pathways of melt and associated dynamics.

25

2.0 Geological Background and Previous Tomographic Studies

Precambrian basement, formed during the Neoproterozoic Pan-African orogeny, underlies much of Ethiopia and the surrounding area (Mège & Korme, 2004). Volcanism initiated 45 Ma and continues to the present day, with the first eruptions beneath the Amaro province of southern Ethiopia (Ebinger et al., 1993; Rooney, 2017; Rooney et al., 2014). The largest volcanic event was the emplacement of the Ethiopian flood basalt province, occurring 31-29 Ma (Hofmann, et al., 1997; Rooney, et al., 2012a; Ukstins et al., 2002) with sporadic volcanism continuing on the Ethiopian Plateau, including the emplacement of alkali shield volcanoes above the flood basalts (e.g. Choke and Guguftu 22 Ma (Kieffer et al., 2004)). Volcanic activity today is largely focused within ~15 km-wide and 60 km-long crustal magmatic segments at the rift axis which are oblique to the border faults and trend en-echelon along the rift (Ebinger et al., 2001; Wolfenden et al., 2004). Recent volcanism occurs off rift at Nabro volcano in Afar (Goitom et al., 2015), in the Yerer-Tullu Wellel Volcanotectonic Lineament (YTVL) (Rooney et al., 2014) and eruptive centres are present south of Lake Tana (Corti, 2009; Kieffer et al., 2004) (Figure 1). The volcanic segments in Afar and the Main Ethiopian Rift (MER) are the main focus of present day extension, while prior to the Pleistocene, extension was focussed at the border faults (Ebinger et al., 2001; Wolfenden et al., 2004, 2005).

Rifting in Afar initiated between 29 Ma (Wolfenden et al., 2004) and 26 Ma (Bosworth et al., 2005), just after the main flood basalt emplacement at 31-29 Ma (Hofmann et al., 1997; Pik et al., 1998; Wolfenden et al., 2004), while the MER started rifting later at 20 Ma in the south, and 11 Ma in the north (Kazmin et al., 1978; Wolfenden et al., 2004). Extension rates are variable with full spreading rates of ~6 mm/yr. for the MER (Birhanu et al., 2016; Jestin et al., 1994; Saria et al., 2014), 16 mm/yr. for the Gulf of Aden Rift (Jestin et al., 1994; Vigny et al., 2006) and 18 mm/yr. for the Red Sea Rift (McClusky et al., 2010; Vigny et al., 2006) (Figure

1 1a) The Tendaho Goba'ad discontinuity (TGD) separates the east-west directed extension in
2 the northern-most MER from the northeast-southwest directed extension in Afar (Tesfaye et
3 al., 2003).

4
5 Crustal thickness has been modified by rifting and volcanism, with the 35–40 km thick
6 Somalian Plateau (Mackenzie et al., 2005; Stuart et al., 2006) and western edge of the Ethiopian
7 Plateau (30–35 km thick, Ogden et al., 2019) considered representative of pre-rifted
8 unmodified crust (Mackenzie et al., 2005; Ogden et al., 2019; Stuart et al., 2006). In contrast
9 crustal thickness for the eastern part of the Ethiopian Plateau ranges from 40–45 km, and has
10 been significantly affected by several kilometre thick magmatic addition from the flood basalt
11 volcanism and likely associated lower crustal intrusions (Mackenzie et al., 2005; Ogden et al.,
12 2019; Stuart et al., 2006; Wang et al., 2021). The crust beneath the rift is predominantly thinner
13 than the plateaus at 25–35 km thick, with the exception of the southern MER, where crustal
14 thickness ranges from 30–38 km. There is also evidence of magmatic additions to the MER
15 crust (Hammond et al., 2011; Maguire et al., 2006; Stuart et al., 2006; Wang et al., 2021). In
16 Afar crustal thickness ranges from <16–26 km, with the Danakil depression having the thinnest
17 crust of the subaerial rift system (Dugda et al., 2005; Hammond et al., 2011; Lavayssière et al.,
18 2018; Makris & Ginzburg, 1987).

19
20 Lithosphere-asthenosphere-boundary (LAB) depths have been interpreted from S-to-P receiver
21 functions, surface wave tomography and a joint inversion between Rayleigh wave velocities
22 and receiver functions. The LAB is thought to lie at 60–80 km depth beneath the Ethiopian
23 Plateau; whereas within the rift the LAB is shallower, with seismic imaging suggesting that the
24 mantle lithosphere is either very thin or non-existent (Dugda et al., 2007; Lavayssière et al.,
25 2018; Rychert et al., 2012).

1
2 Previous tomographic studies suggest the upper mantle beneath the northern East African Rift
3 is slower than expected for a mantle peridotite with $V_s = 3.80\text{-}4.25$ km/s (e.g. Chambers et al.,
4 2019; Gallacher et al., 2016) and V_p 4–10% slower (e.g. Bastow et al., 2008, 2005; Fishwick,
5 2010). These velocities are significantly slower than those observed at similar depths beneath
6 continental interiors, where shear velocities are typically >4.45 km/s (Kennett et al., 1995).
7 Elevated mantle potential temperatures of $100\text{--}170^\circ\text{C}$ are also interpreted beneath Ethiopia
8 from petrological modelling (Armitage et al., 2015; Ferguson et al., 2013; Rooney et al., 2012b)
9 and can fully account for off rift velocity variations between tomographic models and ak135
10 but not for those beneath the rift (e.g. Gallacher et al., 2016). Regional studies for both P-wave
11 and shear velocity, find significantly slow velocities beneath the rift valley, which require a
12 component of partial melt after accounting for variations in temperature and composition
13 (Bastow et al., 2005; Chambers et al., 2019; Gallacher et al., 2016; Hammond, et al., 2013).
14
15 At crustal depths within the rift, P-wave studies image fast velocities which have previously
16 been interpreted as solidified magmatic intrusions (Mackenzie et al., 2005; Maguire et al.,
17 2006). More recently S-wave and surface wave velocity studies, which are more sensitive to
18 fluids, observe slow velocities in the rift valley, suggesting a fluid component is present, in
19 contrast to P-wave studies (Chambers et al., 2019; Kim et al., 2012; Korostelev et al., 2015)
20 Consequently V_p/V_s ratios for the crust are high, and in places >2.0 (Hammond, et al., 2011).
21 Elevated V_p/V_s values in the presence of active volcanism have been interpreted as fluids and
22 melt beneath the rift (Dugda et al., 2005; Hammond, et al., 2011; Ogden et al., 2019; Stuart et
23 al., 2006). Magnetotelluric surveys have imaged high conductivity bodies within the crust
24 beneath the MER and Afar which provide further evidence for ongoing magmatic emplacement
25 beneath the rift (Samrock et al., 2018; Whaler & Hautot, 2006). Off rift, there is similar

1 evidence for ongoing lower crustal intrusions and magmatic activity. S-wave studies beneath
2 parts of the Ethiopian Plateau, YTVL and Nabro detect slow velocities (Chambers et al., 2019;
3 Kim et al., 2012) coupled with high conductivity bodies beneath the YTVL (Didana et al.,
4 2014; Samrock et al., 2018, 2015; Whaler & Hautot, 2006). In addition, Nabro volcano in
5 Eritrea erupted in 2011 (Goitom et al., 2015). While these studies hint at the presence of off
6 rift partial melt, lack of instrumentation off rift mean the interpretations for the crustal and
7 overall off rift plate structure remain unclear.

8

9 **3.0 Methods**

10 **3.1 Datasets**

11 We used data from 13 temporary seismic networks and 5 permanent stations installed between
12 1999 and 2017 (Figure 1a). We used the vertical component data from the broadband
13 seismometer stations. Fifty-eight stations from 3 networks were included (ARGOS XM 2012-
14 2014, Plateau YY 2014-2016 and Afar0911 2H 2012-2013), in addition to the 170 used in
15 Chambers et al. (2019) and 290 in Gallacher et al. (2016) (Figure 1a). The variation in the
16 number of stations between the two studies arise from the 3λ station separation requirement
17 for the ambient noise cross-correlations resulting in the removal of the RiftVolc (Y6 2016–
18 2017) network. Furthermore, networks and stations with short deployment durations such as
19 the EAGLE phase iii network, were not included in the ambient noise.

20

21 **3.2 Ambient Noise Phase Velocity**

22 **3.2.1 Data Processing**

23 The vertical component data was downsampled to 1Hz, normalised and whitened with a 4th
24 order Butterworth bandpass filter between 0.005–0.4 Hz following the method of Bensen et al.
25 (2007). The data were cross-correlated on the 24-hour long waveforms for each concurrently

1 running station pair. The cross-correlograms for every day and each station pair were then
2 stacked to improve the signal to noise ratio (SNR). Station pairs with less than 10 days of
3 continuous recording, 10 days' worth of stacked cross-correlation functions, interstation
4 distances $<3\lambda$ or a SNR <3 were removed (Bensen et al., 2007; Chambers et al., 2019; Harmon
5 et al., 2007) resulting in 6716 cross-correlation functions (Supplementary Figure S1). We
6 examined the stability of the cross-correlations through time, and found that typically >1 month
7 stacks produced waveforms with phases that were within 1-2 s of the long term stack. SNR
8 increased as expected with longer time period stacks and is consistent with numerous previous
9 studies (e.g. Bensen et al., 2007). Then the fundamental mode Rayleigh wave data were
10 windowed using a time variable filter (Landisman et al., 1969), and the Fourier amplitude and
11 phase calculated at each frequency of interest via a fast Fourier transform.

12 **3.2.2 1-D Phase Velocity**

13 The phase velocity dispersion across the region was estimated using a spatial domain technique
14 across the entire array. A zero order Bessel function of the first kind was fitted to the real part
15 of the Noise Correlation Function (NCF) in the Fourier domain by searching over phase
16 velocities from 2.5–5 km/s in 0.01 km/s steps for every period of interest. For each stacked
17 NCF the phase was measured at each period by unwrapping the phase, using the average phase
18 velocity curve at the longest periods, to resolve cycle ambiguity (Bensen et al., 2007; Harmon
19 et al., 2008).

20 **3.2.3 2-D Phase Velocity**

21 The phase velocity maps were then generated by inverting the phase data using the Born
22 approximation 2-D phase sensitivity kernels (Zhou et al., 2004) and an iterative damped least
23 squares approach (Tarantola & Valette, 1982). We used a regular $0.25^\circ \times 0.25^\circ$ grid of nodes
24 as our parameterization for the inversion (Figure 2) and averaged the sensitivity kernel between
25 each station pair onto the nodes (Harmon et al., 2013; Yang, Y. & Forsyth, 2006). The

1 sensitivity kernel is calculated at every period for each station pair on a densely sampled grid
2 ($0.1^\circ \times 0.1^\circ$) and then the Gaussian distance-weighted average value is taken to determine the
3 value at each node on the coarser grid with a Gaussian width (2-sigma) of 40 km. The inversion
4 estimates the average phase velocity at each node, then we “undo” the Gaussian weighted
5 average to recover a $0.1^\circ \times 0.1^\circ$ sampled grid by determining the Gaussian weighted
6 contribution of the nearest nodes to each pixel using the same 40 km Gaussian width. We
7 present the formal error from the last iteration of the inversion, which is propagated from the
8 nodal parameterization to the higher density grid using the Gaussian weights of each node at
9 each pixel using the full covariance matrix. We used an a priori damping parameter of 0.2 km/s
10 in the phase velocity inversion. This value choice stabilises the inversion but is not restrictive
11 as the value is much larger than the standard deviations from the mean velocity at each period
12 (Forsyth & Li, 2005). Choices smaller than 0.2 km/s resulted in damped velocity variations.
13 This resulted in well resolved phase velocity maps between 8-26 s varying between ± 0.02 –
14 0.07 km/s which are indicated in the average 1D profiles in Figure 3 and S3, with lower errors
15 in the rift and at shallower depths.

16 **3.3 Teleseismic Rayleigh wave phase velocity**

17 **3.3.1 Data Processing**

18 We extracted amplitude and phase information from vertical component seismograms for
19 earthquakes with >5.5 magnitude and epicentral distances of 25 - 150° (1053 events, Figure 1
20 inset and supplementary Figure S1). For each teleseismic event, data were processed in the
21 following way. The raw data were demeaned and detrended and their instrument response was
22 removed. The data was then bandpass filtered using a 4th order Butterworth filter between
23 0.005 – 0.4 Hz. Then the fundamental mode Rayleigh wave data were windowed using a time

1 variable filter (Landisman et al., 1969), and the Fourier amplitude and phase calculated at each
2 frequency of interest via a fast Fourier transform.

3 **3.3.2 1-D phase velocity inversion.**

4 We determined the average dispersion curve for the area using a 1-D version of the two-plane
5 wave inversion method (Forsyth et al., 2005). The inversion was completed in two steps, with
6 the first stage utilising a simulated annealing method to fit the two plane wave parameters for
7 each event, while trying a range of starting phase velocities for the model between 3.00-4.40
8 km/s (Press et al., 1992). This ensured a global starting model was found for input into the
9 second stage which utilised an iterative damped least squares inversion (Tarantola et al., 1982)
10 (Figure 3). The inversion simultaneously solves for the phase velocity, azimuthal anisotropy
11 and wave parameters for each event.

12 **3.3.3 2-D phase velocity inversion.**

13 We used the two plane wave method of Forsyth & Li (2005) at each period to invert for a phase
14 velocity map using the amplitude and phase measurements described above from the
15 teleseismic events, using 2-D finite frequency kernels (Yang & Forsyth 2006; Zhou et al.,
16 2004). We used the same nodal parameterization as used in the ambient noise tomography, i.e.,
17 a $0.25^\circ \times 0.25^\circ$ nodal grid with the outermost row and column spaced at 1° to absorb velocity
18 heterogeneities outside the target region (Figure 2). The average 1-D phase velocity described
19 in the previous paragraph is used as our starting model at each period. The phase velocity
20 inversion used 2-D finite frequency kernels (Forsyth et al., 2005; Nishida, 2011; Tromp et al.,
21 2010; Yang et al., 2006) and an iterative damped least squares approach (Tarantola & Valette
22 1982), which solves for phase velocities at each node, and the plane wave parameters for each
23 event. We use the same Gaussian averaging scheme described above to generate higher
24 resolution phase velocity grids at $0.1 \times 0.1^\circ$. The phase velocity at the nodes represent an

1 average phase over the smoothed area around the node, so the final phase velocity maps at 0.1
2 x 0.1° resolution are determined from the Gaussian distance weighted contributions of the
3 nearest nodes to each pixel. We present the formal error from the last iteration of the inversion,
4 which is propagated from the nodal parameterization to the higher density grid using the
5 Gaussian weights of each node at each pixel using the full covariance matrix. The inversion is
6 run twice. After the first set of inversions, events with phase misfits of >4 s are removed from
7 the starting dataset and this is the input used for the final set of inversions. The removal of
8 these poorly fit events is necessary as it removes waveforms with complicated source radiation
9 patterns and other effects not accounted for in the inversion.

10 **3.4 Joint Inversion for Shear Velocities**

11 For the shear velocity inversion, we inverted each pixel of the phase velocity maps across all
12 periods for a 1-D shear velocity structure at every node as a function of depth (Figure 3). The
13 combined 1-D shear velocities at every pixel collectively form the 3-D volume. We used the
14 phase velocity maps from the ambient noise for 8–26 s and the phase velocity maps from the
15 teleseismic results for 29–100 s. The transition at 26 s from one data-type to another is chosen
16 based on the relative amounts of data, i.e. where the teleseismic has a greater number of ray
17 paths. Where the ambient noise and teleseismic phase velocity maps overlap, they are within
18 error of each other, e.g. at 20–33 s (Figure S2).

19 We used an iterative damped least squares inversion (Tarantola & Valette 1982) and
20 parameterised the shear velocity every 5 km vertically with 0.1° x 0.1° pixel size. The partial
21 derivatives that relate variations in shear velocity to changes in phase velocity were calculated
22 using DISPER80 (Saito, 1988). We assigned a nominal a priori standard error of 0.2 km/s for
23 the shear velocity starting model and fixed the V_p/V_s ratio to 1.8, which is the crustal average
24 from receiver function analyses (Hammond, et al., 2011; Stuart et al., 2006) and also a typical

1 mantle value (Dziewonski, A. M. & Anderson, 1981). Variations in the choice of V_p/V_s (1.5–
2 2.1 the observed V_p/V_s ratios in this area) produced results within error and we present the
3 formal error from the inversion in Figure 3. Finally, we interpolated the velocity structure to 1
4 km depth for presentation purposes using a linear interpolation.

5 **4.0 Errors and resolution**

6 To examine the resolution of the phase velocity tomography, we produced checkerboard
7 resolution tests (Figure 4) and synthetic structure recovery tests (Figure 5). Checkerboard tests
8 were produced at lateral length scales of 70 km (periods 8–40 s), and 165 km ($\sim 1.5^\circ$, periods
9 8–100 s)(See supplementary S4 and S5 for further checkerboard tests at length scales of 110
10 km (1°) and 220 km (2°)). We show the checkerboards at the shortest and longest periods for
11 the ambient noise (8 and 26 s) and teleseisms (29 and 100 s). In the period range where the two
12 data sources overlap, the teleseisms offer improved resolution at periods longer than 29s,
13 without losing resolution for crustal structure provided by the ambient noise. We also mask
14 results outside the 2σ standard error contour from the phase velocity inversion. This is the
15 formal error of the inversion from the linearized iterative least squares at the last iteration.
16 Errors are propagated to the higher density grid in a similar fashion to the phase velocities.

17

18 The checkerboard tests indicate that 70 km length scale anomalies are well recovered for 8–40
19 s period, and 165 km length scale anomalies are resolved for all periods within the 2σ error
20 contour (Figure 4). Inside the rift, anomalies are resolved at 70 km length scales for 8–40 s
21 period, and 110 km anomaly length scale anomalies can be resolved using periods of less than
22 71s (Figure S4). These periods are reflective of crustal and mantle depths which are discussed
23 here down to ~ 120 km depth (Figure S4). Off rift, 70 km length scale anomalies are well
24 resolved at periods less than 17s for the ambient noise and at periods less than 40 s from the
25 teleseisms. The 110 km length scales anomalies off rift (Figure S4), can resolve anomalies at

1 periods less than 71s. The 165 km length scale anomalies for all periods, both on and off rift,
2 are well resolved, however amplitude decreases at the longer periods. For checker anomaly
3 length scales of 220 km, features are resolved at periods up to 100s (Figure S5). The decrease
4 in amplitude recovery is likely related to the broad sensitivity kernels for the longest periods
5 (first Fresnel zone ~500-1000 km wide for 100s period (Yoshizawa & Kennett, 2002)) and
6 fewer ray paths. In areas where ray coverage is sparse we have poorer resolution, such as the
7 Red Sea, Gulf of Aden and eastern part of Afar. There is also northeast-southwest smearing of
8 the checkerboards beneath the Red Sea Rift and northwest-southeast smearing beneath Yemen.
9 Consequently, we do not interpret these areas.

10

11 In the synthetic recovery tests, we input a slow velocity anomaly of similar magnitude to our
12 output models in the MER and beneath the Ethiopian Plateau. Broadly, based on depth
13 sensitivity kernels (Figure 3), features >70 km in length scale are well resolved in the upper 70
14 km and features >150 km in length are resolvable at all depths of the shear velocity model. We
15 show periods at the minimum and maximum period for the ANT (8 and 26s) and teleseisms
16 (29 and 100s) to show the full range of recovery (Figure 5). Within the rift, the recovery tests
17 indicate phase velocity variations are resolved in our models at all periods (Figure 5) though at
18 periods >71 s the anomaly beneath the rift is less elongate and focusses to the northwest of the
19 MER. Off rift synthetic recovery tests at periods <29s, indicate the two slow velocity features
20 are resolvable if they are >70 km wide. At longer periods the two anomalies are no longer
21 distinct. Synthetic recovery tests within the Red Sea and Gulf of Aden Rifts suggest there is
22 smearing in a northeast direction particularly for the Gulf of Aden at longer periods.

23 To examine vertical resolution, we perform a spike test (Backus & Gilbert, 1970) for a range
24 of model depths (Figure 6). These kernels show the recovery of a spike function based on the

1 diagonals of the formal resolution matrix. The formal resolution matrix is derived from the
2 damped least squares inversion where the Resolution:

3
$$\mathbf{R} = \left(\mathbf{G}^T \mathbf{C}_{nn}^{-1} \mathbf{G} + \mathbf{C}_{mm} \right)^{-1} \left(\mathbf{G}^T \mathbf{C}_{nn}^{-1} \mathbf{G} \right).$$
 \mathbf{G} is the matrix of partial derivatives from the
4 kernel at each node, \mathbf{C}_{nn} is the data covariance matrix and \mathbf{C}_{mm} is the model covariance matrix
5 (Saito, 1988). Resolution ranges between 0-1, where 1 indicates a completely resolved model
6 parameter (e.g. velocity of a layer) and 0 means not resolved. A value of 0.33 would require 3
7 adjacent layers to resolve 1 piece of independent information about velocity. The kernels
8 suggest shear velocities are resolvable down to 150 km depth (Figure 3 and Figure 6) after
9 which the resolution starts to deviate from the typical bell curve shape. At the shallowest depths
10 (7-22 km), depth slices averaged over ± 10 km are resolved, and for the deepest slices at 142-
11 162 km depth are resolved when averaged over ± 50 km velocity averages.

12

13 **5.0 Results**

14 **5.1 1-D dispersion curves and shear velocity model**

15 Average 1-D dispersion curves measured from the ambient noise and teleseismic data are
16 shown in Figure 3. Phase velocities range from 2.93 ± 0.02 km/s at 8s to 3.87 ± 0.03 km/s at
17 100s (Figure 3 and Figure 7). Where the phase velocities from the ambient noise and teleseisms
18 overlap (20-33 s) (Figure S2), the velocities are similar to one another and within the standard
19 error. The shear velocity structure is displayed in Figure 3b with shear velocities ranging from
20 3.10 ± 0.02 km/s at 5 km depth to 4.42 ± 0.03 km/s at 150 km depth. The average 1-D shear
21 velocities are slower than the input model (green line) at most depths except from 40-80 km
22 depth. In Figure 3c we present the sensitivity kernels, which indicate the depths of peak

1 sensitivity for the shear velocity inversion at 8, 15, 20, 26, 29, 40, 71 and 100 s period. The
2 kernels suggest 150 km depth is the limit of our sensitivity.

3

4 **5.2 2-D Phase Velocities**

5 We generate phase velocity maps from 8-100 s (Figure 7) and observe velocity variations that
6 correlate with geologic and tectonic features. Phase velocities are more variable at shorter
7 periods ranging from $2.85\text{-}3.45 \pm 0.04$ km/s at 10 s becoming less laterally variable at 100 s
8 period ($3.80\text{-}3.90 \pm 0.06$ km/s).

9

10 Within the rift system, the MER is the slowest region of our study for all periods (at 10s the
11 Red Sea is slower but due to ray path bias, we do not interpret this region). We observe
12 minimum velocities ranging from 2.90 ± 0.03 km/s at 10 s to 3.82 ± 0.04 km/s at 100 s period.
13 Beneath Afar, velocities are ~ 0.20 km/s faster than the MER at all periods. Phase velocity
14 ranges from 3.15 ± 0.03 km/s at 10 s to 3.90 ± 0.04 km/s at 100 s period. Within the rifts, slow
15 velocities are not laterally continuous and at periods < 71 s, display segmentation. The
16 anomalies are ~ 100 km in length and in most cases broadly correlate with Quaternary
17 volcanism or hydrothermal sites.

18

19 The velocity structure on the Ethiopian Plateau is variable, exhibiting both fast and slow
20 velocity regions across the period range of interest. The eastern part of the Ethiopian Plateau
21 is slow at short periods (from 8 to 26 s) with velocities of 3.00 ± 0.04 km/s at 10 s to 3.50 ± 0.05
22 km/s at 26 s period. At 20 to 26 s period, the shape of the two slow velocity regions connect
23 which may in part be due to smearing according to our resolution tests (Figure 5). At longer
24 periods (> 29 s period), phase velocities beneath the eastern part of the Ethiopian Plateau are
25 similar to background phase velocity values ($\sim 3.80\text{-}3.90 \pm 0.05$ km/s). In contrast, the western

1 part of the Ethiopian Plateau is the fastest area of our study at all periods with phase velocities
2 of 3.35 ± 0.04 km/s at 10 s period increasing to 3.92 ± 0.05 km/s at 100 s depth.

3

4 **5.3 3-D Shear Velocity Structure**

5 At lithospheric depths (10-80 km), we observe strong lateral variations in shear velocity, up to
6 0.85 km/s across our study region, which likely reflect a combination of significant changes in
7 crustal thickness and variability in mantle structure. Figure 8 and Figure 9 show depth slices
8 and cross-sections through the shear velocity model, respectively. The slowest velocities in the
9 region are beneath the MER from 10–50 km depth and the Danakil depression from 50-80 km
10 depth (due to variations in lithospheric thickness). For instance, at crustal depths (10 km) we
11 find $V_s = 3.00 \pm 0.03$ km/s and at mantle depths (60 km) we find $V_s = 4.05 \pm 0.03$ km/s (Figure
12 8 a-c). In profile A–A' (Figure 9a) south of the MER, there is a fast lid visible between 37–
13 38.5° E at 60–80 km depth. Within the rift, this fast lid is broken by slow velocity anomalies
14 before becoming prevalent again near the Arabian Peninsula.

15 At 20-40 km depth there is a slow velocity region centred ~100 km southeast of Lake Tana
16 beneath the eastern part of the Ethiopian Plateau (3.10 - 3.85 ± 0.04 km/s at depths of 10 to 40
17 km respectively), and another anomaly ~100 east of this anomaly centred beneath the border
18 fault region. These anomalies are not present below 60 km depth (Figure 8c).

19 The western Ethiopian Plateau is one of the fastest regions at 20-40 km depths with velocities
20 >3.80 km/s at 20 km and >4.15 km/s at 40 km (± 0.04 km/s), but is close to the average velocity
21 across the region at 60 km. Profile B-B' (Figure 9b) shows there is a high velocity lid (velocity
22 >4.15 km/s), from 60-80 km depth west of the rift ($<38.5^\circ$ E), and to the east of the rift ($>40^\circ$

1 E). We note that this study is the first to obtain shear wave velocities this far west on the
2 Ethiopian Plateau due to the addition of recent seismic networks (Figure 1).

3 Afar has some of the fastest velocities in the region at 20 km depth (~ 3.80 km/s, which likely
4 reflects the thinner crust in this region compared to the Plateau. At 40-60 km depth, the Afar
5 region is characterised by several punctuated slow velocity regions with velocities < 3.9 km/s,
6 which are coincident with regions of active hydrothermal sites (stars in Figure 8) or recent
7 volcanism (red polygons in Figure 8). Specifically, beneath the crustal magmatic segments and
8 along the TGD (shown in Figure 1 as red polygons and dashed line respectively) the velocities
9 are ~ 0.2 km/s slower than the rest of Afar at all lithospheric depths.

10 At asthenospheric depths (> 60 km in Afar and > 80 km in the MER and Plateau), we also
11 observe several punctuated slow velocity regions. Specifically, in the 80-120 km depth range,
12 the region beneath the MER is the slowest in the region with velocities < 4.15 km/s (Figure 8
13 and S6). The slow velocity anomaly is not centred beneath the rift but is offset towards the
14 west, beneath Addis Ababa and straddling the rift. There are two other slow velocity regions
15 with velocities < 4.15 km/s located near the Red Sea and the Gulf of Aden. Within Afar, in this
16 depth range, the slow velocities are not as strong.

17 The slow velocity anomalies within the rift appear to systematically extend to shallower depths
18 going from the MER northwards. Profile A-A', along the MER rift axis (Figure 9a), extending
19 into Afar, shows the relationship between the slow velocity anomalies. Near the MER, the slow
20 velocities are visible beneath the fast lid, extending from 80-120 km depth using the 4.05 km/s
21 contour (Figure 9a). Going northwards, the slow velocity anomalies are centred at shallower
22 depths going to 75 km and then to 65 km depth beneath Afar. The base of the anomalies appears

1 fairly constant at ~120 km depth. There does not appear to be much variation in structure at
2 greater depths in our models though this is close to the limits of our depth resolution.

3 The strongest anomalies at crustal depths within the MER and the Ethiopian Plateau, are
4 displaced from the strongest anomalies in the asthenosphere, while in Afar the anomalies in
5 the asthenosphere are close to regions of geologically recent volcanism. In the MER the slow
6 velocities at 20 km depth are located ~100 km southwest of the slowest velocities in the
7 asthenosphere. Beneath the Ethiopian Plateau, velocities at asthenospheric depths are faster
8 than those within the MER and the slowest asthenospheric velocities that may link to those
9 within the crust off rift, are again located within the asthenosphere beneath the MER. Beneath
10 Afar, the slow velocity regions are predominantly located beneath the locations of active
11 hydrothermal sites (stars Figure 8) and recent volcanos (red triangles Figure 8).

12 **6.0 Interpretation**

13 Our results are the first absolute seismic velocity model for the crust and upper mantle beneath
14 both the rift and the Ethiopian Plateau. The primary new components are the crustal structure
15 beneath the Ethiopian Plateau and significant improvement in the resolution of upper mantle
16 depths from the joint inversion. The crustal structure beneath the rift is broadly consistent to
17 that described in Chambers et al. (2019) due to similar methods and networks used. The
18 addition of the Ethiopian Plateau network has provided additional data that have been used to
19 provide fundamental new constraints of the absolute shear velocity structure of the crust and
20 upper mantle off rift beneath the Ethiopian Plateau and beneath parts of the Somalian Plateau.
21 Below we summarize the previous findings for the crustal and uppermost mantle structure from
22 Chambers et al. (2019) and then focus on the new results off rift beneath the Ethiopian Plateau.

1 We then discuss mantle depths beneath the area, where we have significantly improved
2 resolution from the joint inversion.

3 **6.1 Crustal Structure**

4 At shallow depths (10-40 km) we observe the largest range in shear velocity. This has been
5 observed in previous tomographic studies which attributed the range to changes in crustal
6 thickness, melt and temperature (Chambers et al., 2019; Hammond, 2014; Kim et al., 2012;
7 Korostelev et al., 2015). Within the MER observed velocities are slow and have been attributed
8 to higher temperatures and/or the presence of partial melt in the crust. Beneath Afar, where
9 slow velocities are observed beneath crustal magmatic segments in our model (>0.2 km/s
10 slower than surroundings, $V_s = 3.50 \pm 0.02$ km/s), a similar interpretation has been made for
11 the origin of the slow velocities (e.g. Chambers et al., 2019; Rooney, 2020b) (Figure 8).
12 Temperatures in the crust would be too high to account for the observed velocity variations
13 making partial melt the most likely explanation. The amount of partial melt required to explain
14 the slowest velocities ranges from 0.5 to 4% (Chambers et al., 2019). Variations in crustal
15 thickness are also thought to have a control on the observed velocities as thicker crust can
16 increase melt residence times resulting in more felsic compositions and a deep crustal melt
17 zone (Annen et al., 2006; Karakas & Dufek, 2015; Siegburg et al., 2018). The faster velocities
18 in Afar at 20–40 km depth ($3.80\text{--}4.05 \pm 0.02$ km/s), are consistent with slower than average
19 mantle velocities constrained in previous studies ($\sim 3.8\text{--}4.1$ km/s) (e.g. Bastow et al., 2008;
20 Gallacher et al., 2016).

21 In contrast to other studies, we provide the first shear velocity model for the Ethiopian Plateau
22 allowing the first seismic interpretation of the velocity structure away from the rift. We observe
23 slow anomalies off rift southeast of Lake Tana on the eastern part of the Ethiopian Plateau
24 ($3.10\text{--}3.85 \pm 0.04$ km/s at depths of 10 to 40 km). At mid to lower crustal depths, and periods
25 >15 s, synthetic recovery tests suggest there is smearing between the two slow velocity

1 anomalies, and with those in the MER (Figure 5). We do not observe a slow velocity anomaly
2 beneath the Ethiopian Plateau below 40 km depth which could be a result of features being too
3 small to be resolved, are isolated to the crust, or finally, are at the limit of our well resolved
4 region off rift. The slow velocity anomalies are a similar magnitude to those within the MER,
5 and broadly correlate in map view to known geothermal activity and past volcanism (Figure
6 8). The last major volcanic event on the Ethiopian Plateau was >21 Myrs (Rooney, 2017)
7 though there is evidence for limited quaternary activity along the YTVL and southwest of lake
8 Tana (Kieffer et al., 2004; Meshesha & Shinjo, 2007). It is therefore unlikely that a remnant
9 thermal anomaly could explain our observations (Figure S7), as simple conductive cooling
10 calculations indicate that a 1300°C thermal anomaly would dissipate in <10 Myr. We suggest
11 there has been further recent magmatic emplacement.

12 The western part of the Ethiopian Plateau is one of the seismically fastest areas of this study
13 (in contrast to the East), and is located 250-500 kms from the rift, providing us with information
14 about pre-rift structure. The distance from the rift coupled with geological studies (e.g. Jones,
15 (1976); Rooney, (2019)) suggests the western part of the Ethiopian Plateau has been minimally
16 impacted by rifting processes. In addition there is little evidence for significant flood basalt
17 magmatism (Mège et al., 2004). We interpret the velocity observations as being most similar
18 to original plate structure before rifting, with our inferred crustal thickness between 20–30 km
19 thick (using the 3.60 km/s contour as used in Chambers et al., (2019)). Ogden et al. (2019),
20 performed receiver function analysis in the same area, finding crustal thicknesses of ~30 km at
21 the western edge of the Ethiopian Plateau and interpreted the area as being unaffected by lower

1 crustal intrusion from the Oligocene flood basalts or younger rifting, consistent with our
2 observations.

3 **6.2 Lithospheric Mantle Structure**

4 In the upper mantle we observe a fast lid in 1-D profile (Figure 3) and cross-sections (Figure
5 9) that are >0.1 km/s faster than the surroundings. We note that while fast, the velocities are
6 slower than lithospheric mantle observed in other continental settings (maximum of 4.30 ± 0.05
7 km/s for our model in comparison to 4.45 km/s for ak135) (Kennett et al., 1995). This feature
8 is broadly at 60–80 km depth off rift, but is absent beneath Afar. In the MER the velocities we
9 observe at similar depths alternate between slow and fast, part way between fast velocities
10 observed off rift and the absence in Afar. The base of the fast lid is commonly associated with
11 the lithosphere-asthenosphere-boundary (LAB) (Fishwick, 2010; Lavayssière et al., 2018;
12 Rychert et al., 2012, 2005) and we therefore plot the S-to-P LAB results of Lavayssière et al.
13 (2018) on our cross-sections (Figure 9 red diamonds in lower panels). We find good agreement
14 between the locations of our fastest velocities ($V_s > 4.15$ km/s) and the existence of a strong
15 LAB phase from the S-to-P results suggesting we are imaging a similar feature. The fast lid is
16 most prominent off rift beneath the rift flanks and Ethiopian Plateau in Profile B-B' (Figure
17 9b), in regions less affected by rifting. In the rift Profile A-A' which extends from the southern
18 MER to the Arabian Peninsula (Figure 9a), a weaker fast velocity zone ($V_s = 4.1-4.15$ km/s)
19 is intermittently visible within the rift and is underlain by the slowest asthenospheric velocities
20 (discussed in the next section). The fast lid becomes prevalent again near the Arabian
21 Peninsula. Our result, with no discernible fast lid in Afar, is consistent with S-to-P results which
22 also did not find a strong, significant LAB phase beneath the majority of the rift (Lavayssière
23 et al., 2018; Rychert et al., 2012). The presence of the slowest asthenospheric velocities ($3.98-$
24 4.06 ± 0.03 km/s) beneath areas where receiver functions do not detect a significant LAB phase

1 (Lavayssière et al., 2018; Rychert et al., 2012), is consistent with a lack of a fast lid in these
2 regions, possibly caused by partial melt at shallow depths (10-60 km depth).

3 **6.3 Asthenospheric Anomalies**

4 At asthenospheric depths of ~80 km beneath the Ethiopian Plateau we observe velocities of
5 4.20 ± 0.05 km/s, which is lower than the global average for continents at that depth (4.45 km/s)
6 using ak135 (Gallacher et al., 2016; Kennett et al., 1995) (Figure 3b and Figure 8). Similar
7 velocity structure imaged in the asthenosphere beneath the Ethiopian Plateau has been
8 explained with elevated temperatures of 100–170°C (e.g. Gallacher et al., 2016), consistent
9 with petrological constraints of the regional mantle potential temperature (Armitage et al.,
10 2015; Ferguson et al., 2013; Rooney et al., 2012b). These velocities provide the regional
11 background velocity for the asthenosphere without melt for comparison to the sub-rift
12 structure. Melt may not necessarily be required for this modest velocity anomaly at least over
13 the broad depth ranges (~40 km) and lateral areas (>160 km) resolvable by surface waves in
14 this region. However, receiver functions image a strong, sharp discontinuity beneath the
15 Ethiopian Plateau (Lavayssière et al., 2018; Rychert et al., 2012) that likely requires a small
16 amount of partial melt, indicating it may exist in a more limited depth and/or lateral area.

17 We observe slow velocity anomalies ($V_s < 4.05$ km/s) at asthenospheric depths of > 60 km
18 within the rifts, which are punctuated segments that are ~110x80 km wide, spaced ~70 km
19 apart, and have a base at ~120 km depth. The segments we describe do not correlate to the
20 surface expressions of magmatic segments described in crustal studies. The slow velocity
21 anomalies get progressively shallower, slower and broader in lateral extent northwards,
22 towards areas at more advanced stages of rifting (Figure 9a). The slowest and largest anomaly
23 with the shallowest starting depth (60 km) exists beneath Afar, the region in the latest stage of
24 rifting (9% velocity decrease when comparing the segments to the rift flanks), while for the
25 MER the shallowest starting depth for the slow velocity bodies are ~80 km depth. We note that

1 while the anomalies themselves are well-resolved, absolute depths could vary by 20-30 km
2 given our sensitivity at these depths. Mantle potential temperatures in the rift are only
3 moderately elevated, $\sim 1450^{\circ}\text{C}$ (Ferguson et al., 2013; Petersen et al., 2015; Rooney, Herzberg,
4 et al., 2012) based on geochemical observations. With this temperature, we would expect
5 velocities to be reduced by $\sim 3\%$ using a Burgers model (Jackson & Faul, 2010) for a peridotite
6 mantle and an estimated geotherm for the MER (Chambers et al., 2019). Composition and
7 temperature are therefore not sufficient to account for the velocity reduction, which suggests a
8 fluid component is required in the asthenosphere beneath the rift. We postulate the fluid
9 component is most likely partial melt, from the abundance of Quaternary volcanism at the
10 surface, which is in agreement with previous tomographic studies (Bastow et al., 2008;
11 Chambers et al., 2019; Gallacher et al., 2016). The anomalies occur at depths consistent with
12 previous geochemical estimates of melt generation (53–120 km depth) (Ferguson et al., 2013;
13 Rooney et al., 2005).

14 The slow velocity segments in the asthenosphere (Figure 9a), are present at the earliest stages
15 of rifting beneath areas where the crust and lithosphere is not significantly thinner than beneath
16 the Ethiopian Plateau (40 km thick crust and 80 km thick lithosphere (Lavyssi re et al., 2018;
17 Stuart et al., 2006)). Furthermore the segments persist into areas of later stage rifting in the
18 northern section of A-A' in Afar (Figure 9a) where the crust is thinner (crustal thicknesses of
19 22 km (Hammond, et al., 2011)). These observations suggest segmented melt supply starts
20 prior to significant crustal and plate thinning.

21 Segmentation beneath rifts is not isolated to this study and has been observed beneath more
22 mature rifts and mid ocean ridges (e.g. Gulf of California, Red Sea Rift, the Mid Atlantic Ridge
23 (Harmon et al., 2020; Lekic et al., 2011; Ligi et al., 2012; Wang et al., 2009) and also beneath
24 the EAR in the same region as our study (Civiero et al., 2015, 2019; Gallacher et al., 2016).
25 The processes generating segmentation beneath the EAR are debated (e.g. Bastow et al., 2008;

1 Civiero et al., 2015; Gallacher et al., 2016; Ligi et al., 2012). The 2 main hypotheses for
2 segmentation are decompression melting with buoyancy driven upwelling from the release of
3 melt (e.g. Gallacher et al., 2016) or the presence of mantle plumelets (e.g. Civiero et al., 2019).
4 The spatial correlation of our anomalies within the rift coupled with larger, slower and
5 shallowing of the top of the slow velocity anomalies towards more advanced stages of rifting,
6 favours rift related buoyancy driven upwelling.

7 Off rift beneath the Ethiopian Plateau at crustal depths, we observed velocities slow enough to
8 contain melt (Figure 9). However velocities in the asthenosphere can be accounted for by
9 temperature alone at scales resolvable by surface waves (e.g. Gallacher et al., 2016). To explain
10 the disconnect between slow velocities off rift in the crust, and limited evidence for melt
11 production in the asthenosphere off rift we propose 3 possibilities. 1. Melt is present off rift
12 (Civiero et al., 2015) but is localised in pockets that are below our resolving power, 2. Melt is
13 ephemeral and was previously located off rift but has now been drained (Sim et al., 2020) and
14 3. Melt has laterally migrated from the rift axis along permeability boundaries (Sparks &
15 Parmentier, 1991). The first option is suggestive of ponding beneath a permeability barrier,
16 and therefore likely also includes migration and/or ephemeral character, and therefore we
17 proceed discussing the latter two options. Similar observations of disconnected slow velocity
18 anomalies have been observed in previous studies at more developed rifts or ridges, which
19 hypothesise lateral melt migration is occurring along a permeability boundary (Braun & Sohn,
20 2003; Ghods & Arkani-Hamed, 2000; Harmon et al., 2020; Holtzman & Kendall, 2010;
21 Rychert et al., 2020; Varga et al., 2008; Wang et al., 2020) and/or melt is ephemeral (Harmon
22 et al., 2020; Rychert et al., 2020; Wang et al., 2020). Within the MER we also observe the
23 slowest velocity anomalies at asthenospheric depths in our model, are not located directly
24 beneath the slowest crustal velocity anomalies (Figure 9). In line with previous studies (Braun
25 et al., 2003; Gallacher et al., 2016; Ghods et al., 2000; Holtzman & Kendall, 2010) we suggest

1 the most likely explanation for the disconnect between melt in the asthenosphere and melt in
2 the crust within the rift, is likely due to lateral melt migration along permeability boundaries in
3 the lithosphere. Overall, this study suggests melt generation and migration are dynamic
4 processes, which require further study to fully understand rifting processes and the factors that
5 dictate the locations of active volcanic/hydrothermal regions.

6

7 **7.0 Conclusions**

8 We present the results from a joint inversion of ambient noise and teleseismic Rayleigh waves
9 to produce a 3-D absolute shear velocity map from 10-150 km depth, for the northern East
10 African Rift. At crustal depths we observe significant lateral velocity variations which can
11 partly be explained by variations in crustal thickness. At crustal depths velocities are slowest
12 beneath the Main Ethiopian Rift and the eastern part of the Ethiopian Plateau and are slow
13 enough to require a component of partial melt. A fast lid, consistent with previous
14 measurements for the Lithosphere-Asthenosphere-Boundary, is observed at 60–80 km depth
15 off rift (>0.1 km/s faster than surroundings). However, within the rift, the fast lid is obscured
16 where it is underlain by asthenospheric segmented slow velocity anomalies, which we interpret
17 as melt infiltration to shallow depths. At asthenospheric depths beneath the Ethiopian Plateau
18 we observe velocities only slightly slower (4.30 ± 0.05 km/s) than global models, that can be
19 explained by elevated temperatures. The rift is significantly slower than off rift at
20 asthenospheric depths, and we observe segmented slow velocity anomalies at 60-120 km depth,
21 including in areas that have not undergone significant crustal and plate thinning. The
22 asthenospheric slow velocity segments are interpreted as areas of partial melt and suggest
23 segmented melt supply starts prior to significant crustal and plate thinning. Furthermore,
24 asthenospheric anomalies are not directly beneath the melt-rich crustal regions, including those

1 off rift, suggesting melt laterally migrates within the mantle, and/or melt is ephemeral and the
2 mantle source for the anomalies observed at crustal depths, have been drained.

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9 al. 2013).

10

11 **Data Availability Statement**

12 All data used to generate the models are freely available from the IRIS Data Management Center
13 (IRISDMC; <https://service.iris.edu/fdsnws/dataselect/1/>) funded through SAGE Proposal of the NSF
14 under Cooperative Agreement EAR-126168. The networks used are:
15 RiftVolc Y6 (https://doi.org/10.7914/SN/Y6_2016), ARGOS XM
16 (https://doi.org/10.7914/SN/XM_2012), Plateau YY (https://doi.org/10.7914/SN/YY_2013), EKBSE
17 XI (https://doi.org/10.7914/SN/XI_2000), YOCMAL XW(
18 https://doi.org/10.7914/SN/XW_2009), EAGLE geophysical XJ
19 (https://www.fdsn.org/networks/detail/XJ_2002/), EAGLE geoscientific YJ
20 (https://www.fdsn.org/networks/detail/YJ_2001/), Boina YZ
21 (https://www.fdsn.org/networks/detail/YZ_2005/), AFAR07 ZE
22 (https://doi.org/10.7914/SN/ZE_2007), Afar depression ZK (https://doi.org/10.7914/SN/ZK_2009),
23 Africa Array AF (<https://doi.org/10.7914/SN/AF>), IU (<https://doi.org/10.7914/SN/IU>), GEOFON GE
24 (<https://doi.org/10.14470/TR560404>), Eritrea YR/5H (https://doi.org/10.7914/SN/5H_2011), Danakil
25 depression ZR/2H (https://doi.org/10.7914/SN/2H_2009), GEOSCOPE G
26 (<https://doi.org/10.18715/GEOSCOPE.G>).

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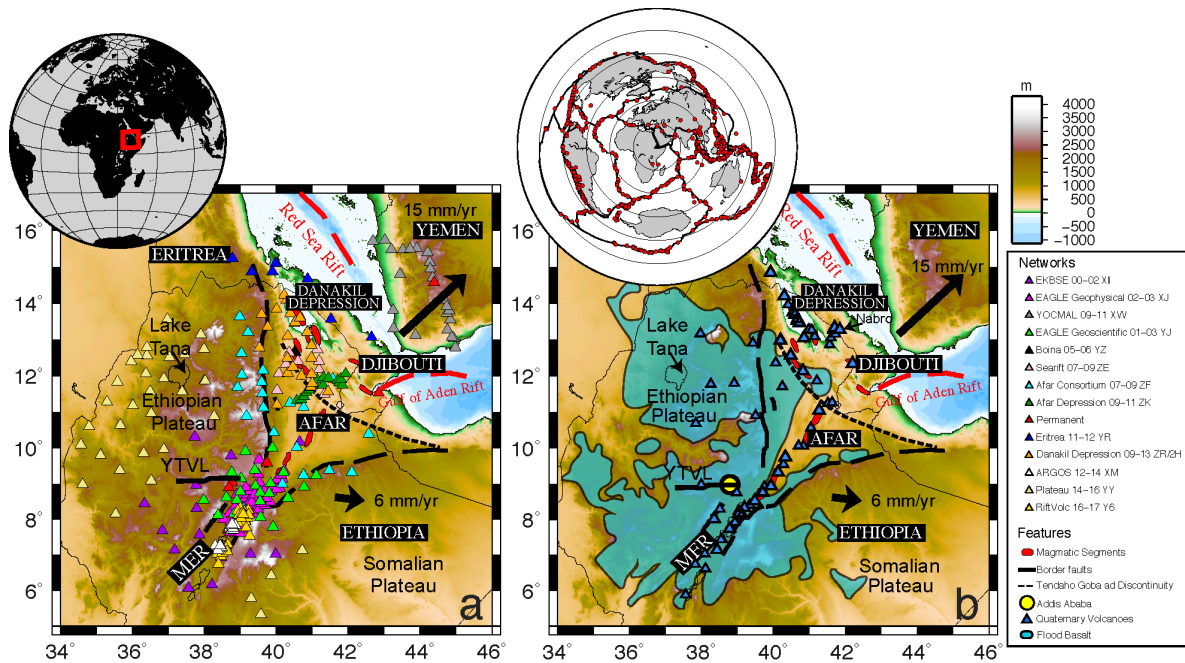
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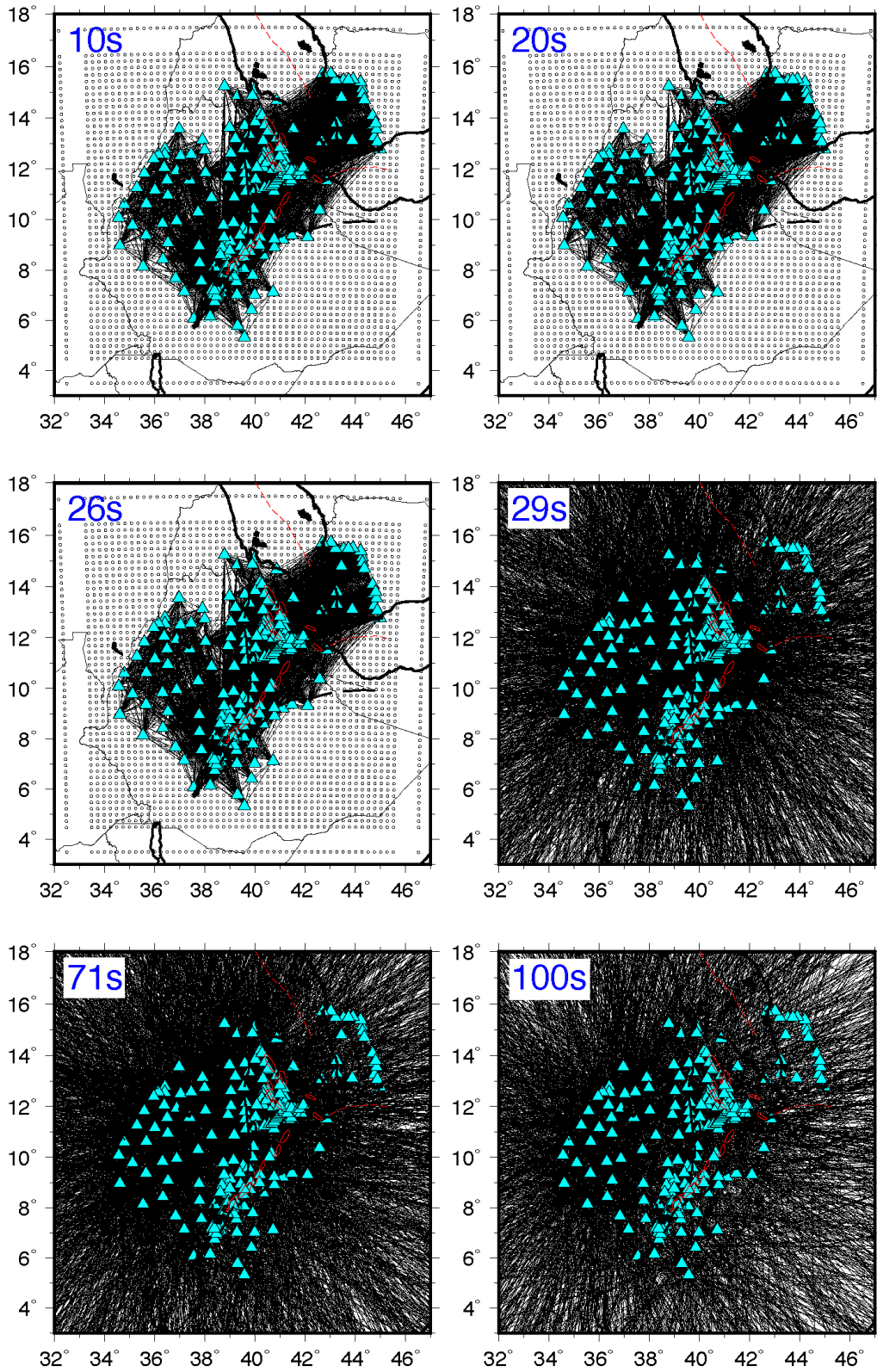
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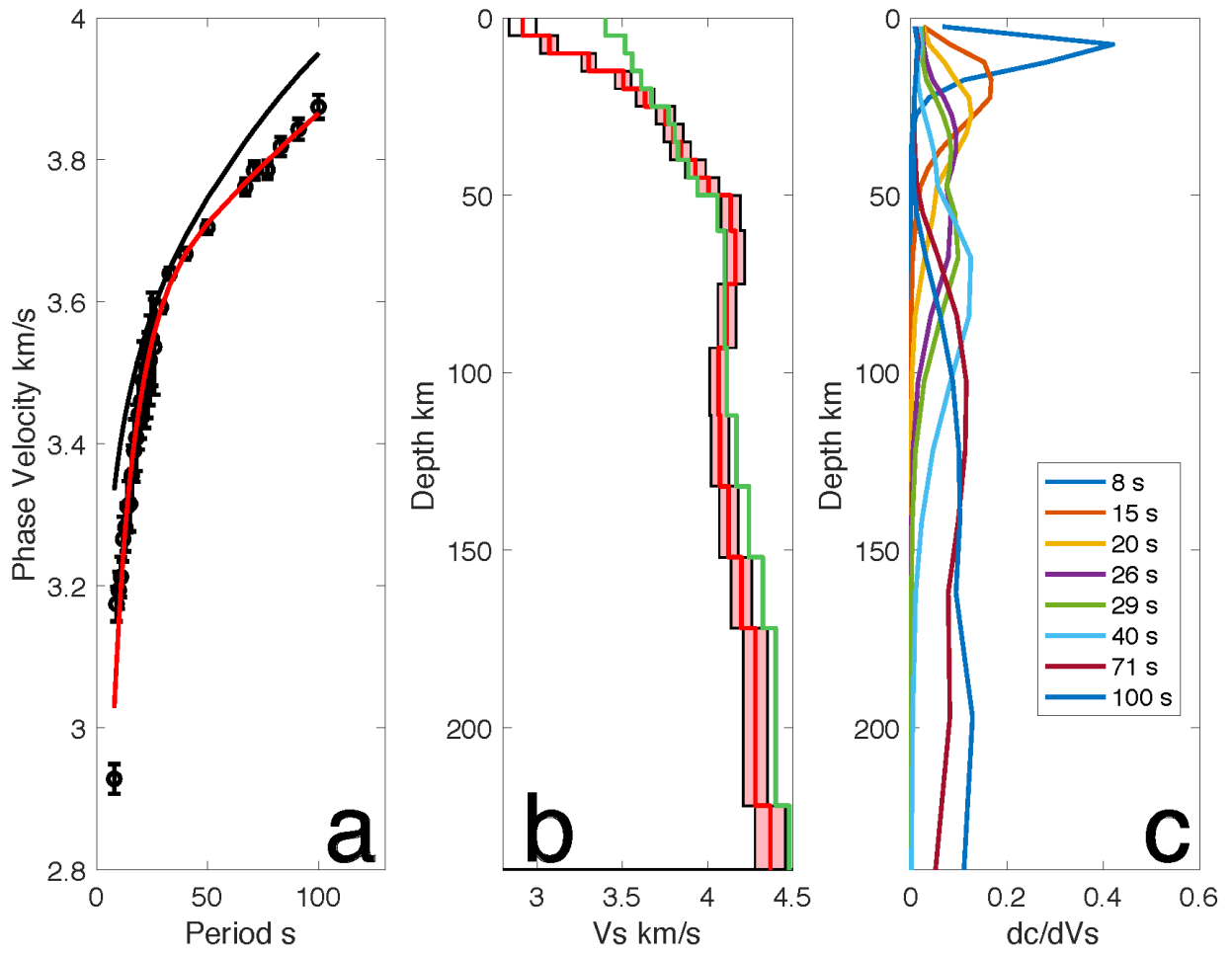
21



1
 2 Figure 1a) Seismic station map of the northern East African Rift. Thick black lines show border faults, red polygons crustal
 3 magmatic segments, and dashed lines the Tendaho-Goba'ad discontinuity (TGD). Stations are triangles coloured to their
 4 project deployment with the final 4 networks (Danakil depression–RiftVolc) new networks not used for ambient noise or
 5 teleseismic tomography. Arrows indicate extension rates relative to the stationary Nubian plate (Jestin et al., 1994; Saria et
 6 al., 2014; Vigny et al., 2006). (b) Geological map. Volcanoes are represented by blue triangles and crustal magmatic segments
 7 red polygons. Flood basalt provinces are shown in blue. Addis Ababa is marked by yellow circle. Inset figure red dots show
 8 locations of the 1053 earthquakes used in this study. Arrows again indicate extension rates relative to the stationary Nubian
 9 plate (Jestin et al., 1994; Saria et al., 2014; Vigny et al., 2006).



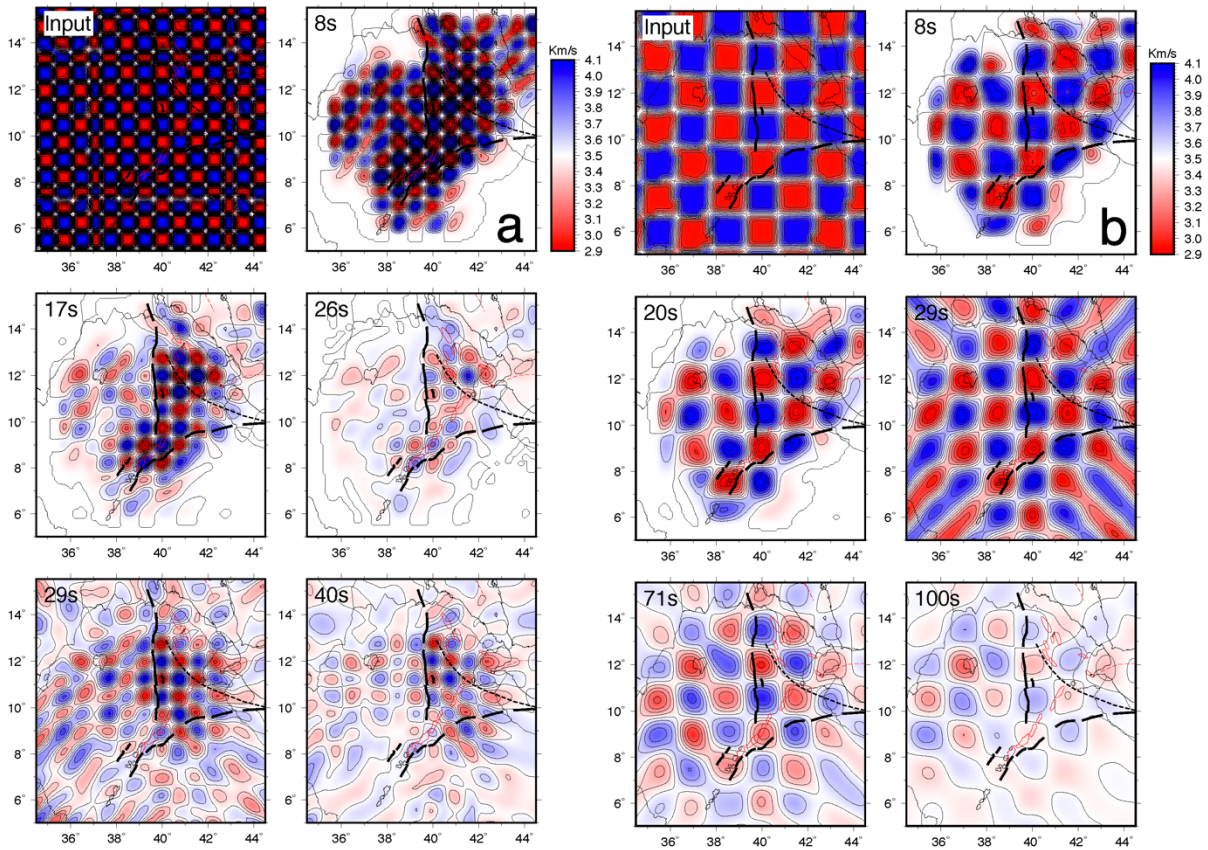
1
 2 Figure 2 Nodal grid at 0.25 ° spacing with ray paths for ambient noise (a–c, 8, 20 and 26s) and teleseisms (d–f, 29, 71 and
 3 100 s) overlain as black lines. Blue triangles indicate stations and red polygons the crustal magmatic segments.



1

2 *Figure 3a) Average 1-D phase velocity for the study area with 3σ error bars (circles), starting model (black line) with our*
 3 *best fit shear velocity model dispersion overlain (red line). (b) Best fit shear velocity model for the study area (red line) and*
 4 *formal 2σ error bounds (thin black lines and shaded area). Green line is initial starting model using the average shear velocity*
 5 *from Chambers et al., (2019) and Gallacher et al., (2016). (c) Sensitivity kernels for Rayleigh waves at selected periods.*

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Figure 4 a) Checkerboard tests at $70 \text{ km} \times 70 \text{ km}$ for phase velocities from ambient noise (8–26s) and teleseisms (29 and 40

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s) which are sensitive to crustal and upper most mantle depths (left 2 panels). Within the rift and eastern part of the Ethiopian

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Plateau we can resolve the checkerboards with confidence. Outside these areas features should not be interpreted. Right 2

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panels: b) Checkerboard tests at $165 \text{ km} \times 165 \text{ km}$ ($\sim 1.5^\circ \times 1.5^\circ$) for phase velocities from ambient noise (8–20 s) and

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Teleseisms (29–100 s) for the minimum and maximum periods. Within the rift and eastern part of the Ethiopian Plateau we

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can resolve the checkerboards with confidence to 100 s period, although the magnitude of the anomaly weakens with

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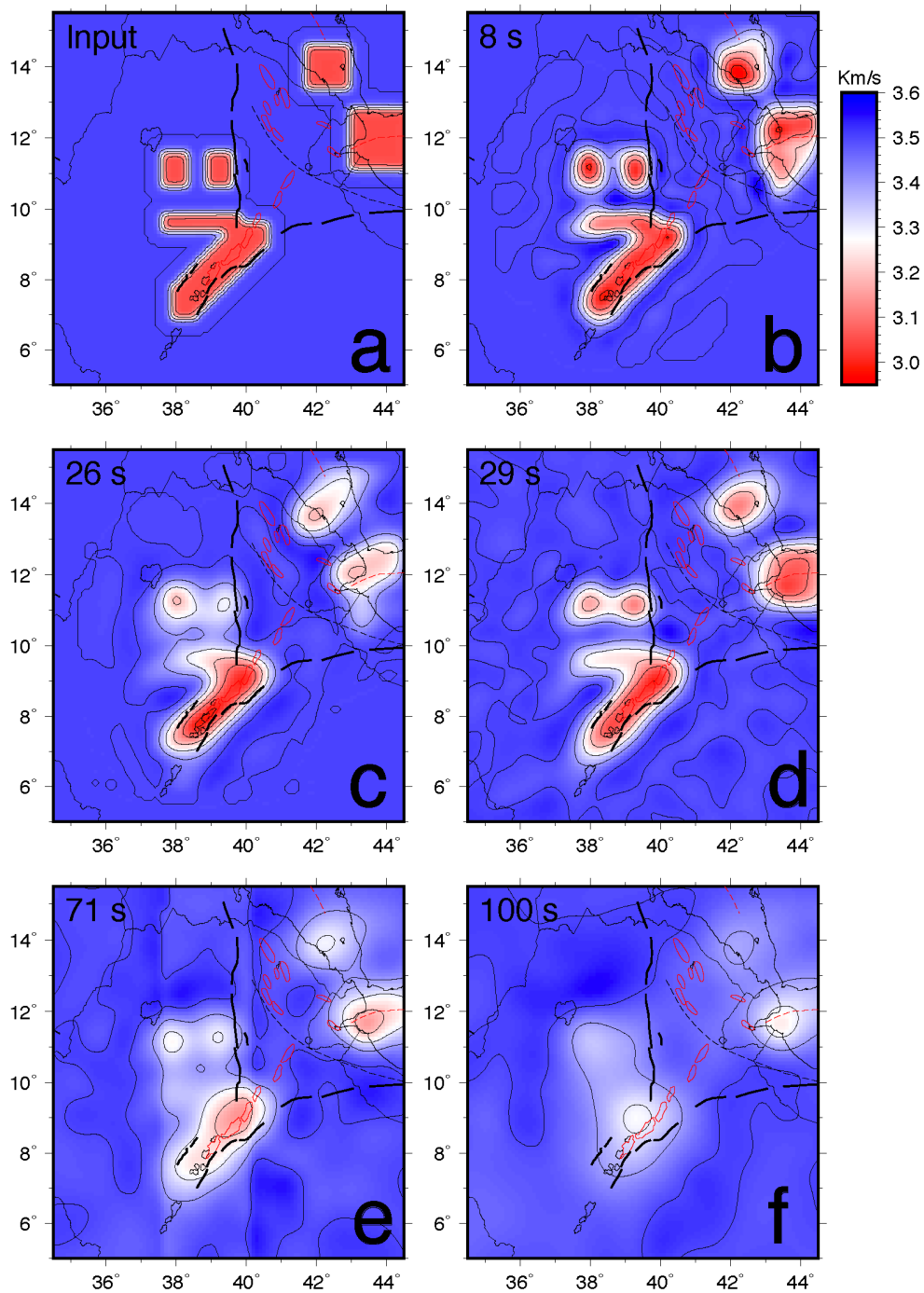
increasing period. Thick black lines show border faults, red polygons crustal magmatic segments, and dashed lines the

9

Tendaho-Goba'ad discontinuity (TGD). For further checker board tests at 110 km and 220 km spacings see supplementary

10

Figures S4 and S5.



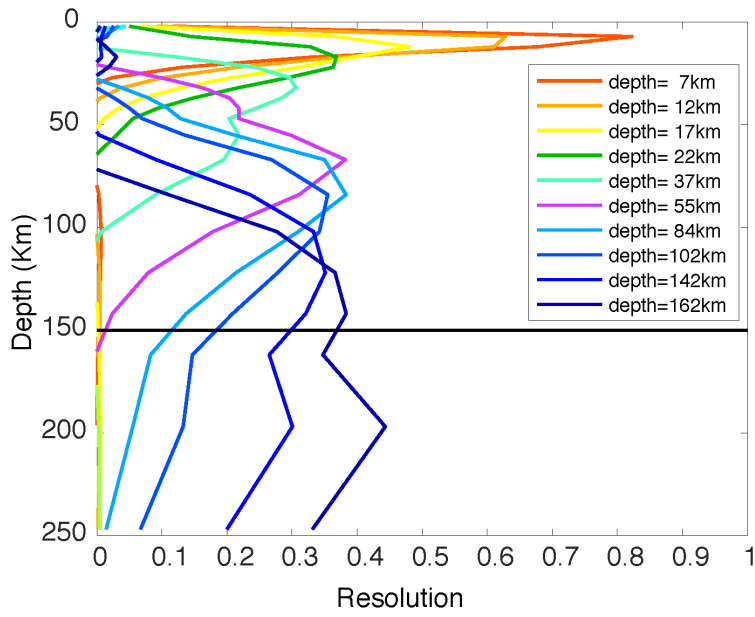
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3 *Figure 5 Synthetic recovery tests inputting slow velocity anomalies beneath the Main Ethiopian Rift, YTVL, Red Sea Rift, Gulf*
 4 *of Aden Rift and beneath the eastern part of the Ethiopian Plateau at 8, 26, 29, 71 and 100 s period. Thick black lines show*
 5 *border faults, red polygons crustal magmatic segments, and dashed lines the Tendaho-Goba'ad discontinuity (TGD). Off rift*
 6 *features are resolvable at short periods equivalent to crustal depths though smear together at longer periods. The MER and*
 7 *Gulf of Aden anomalies are visible at all periods with decreasing amplitude at longer periods.*

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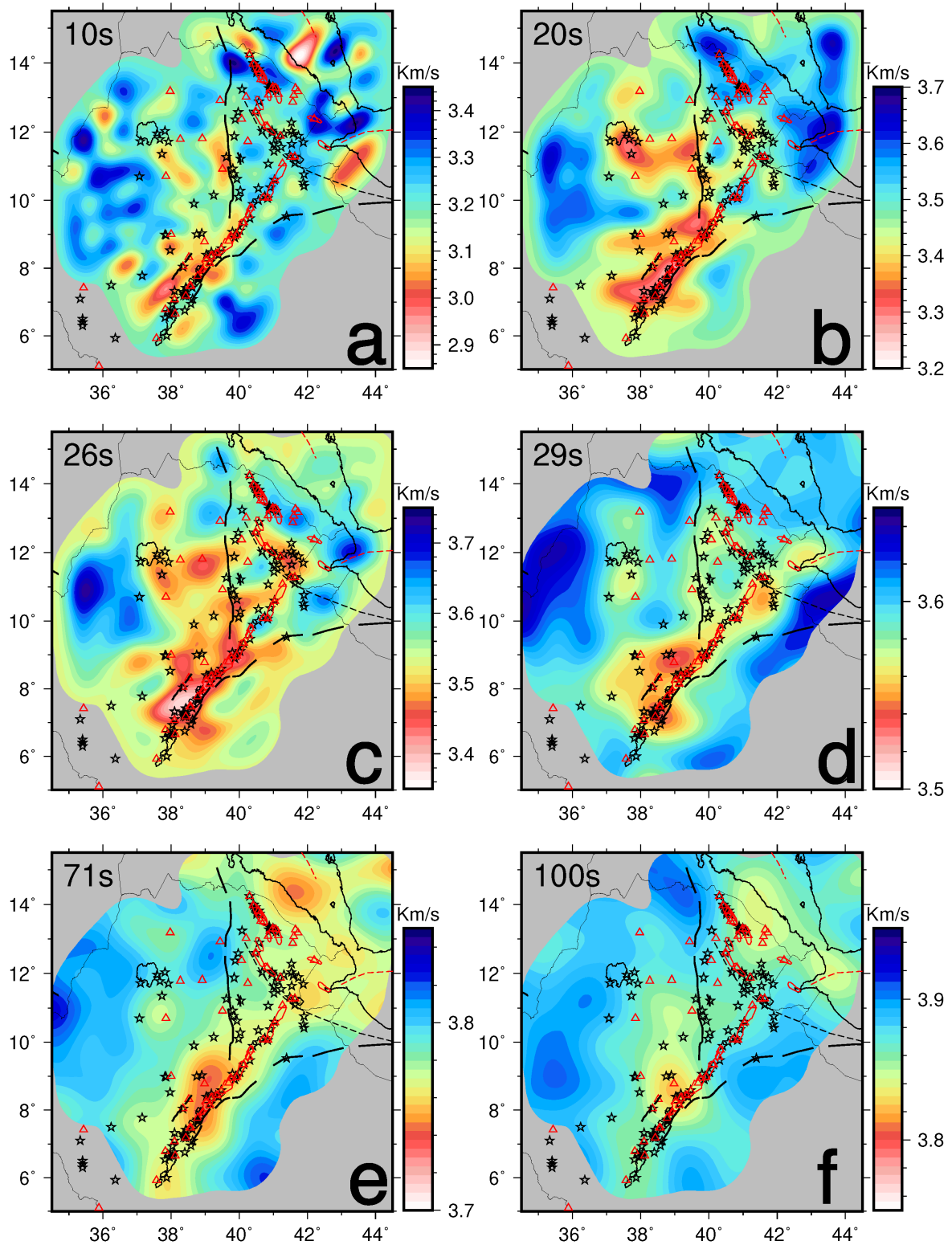


3

4 *Figure 6 Spike test for depths of 7, 12, 17, 22, 37, 55, 84, 102, 142 and 162 km depth. Vertical resolution ranges from ± 10 km*

5 *at the shallowest depths (7-22 km) and ± 50 km for the deepest slices at 142-162 km depth.*

6

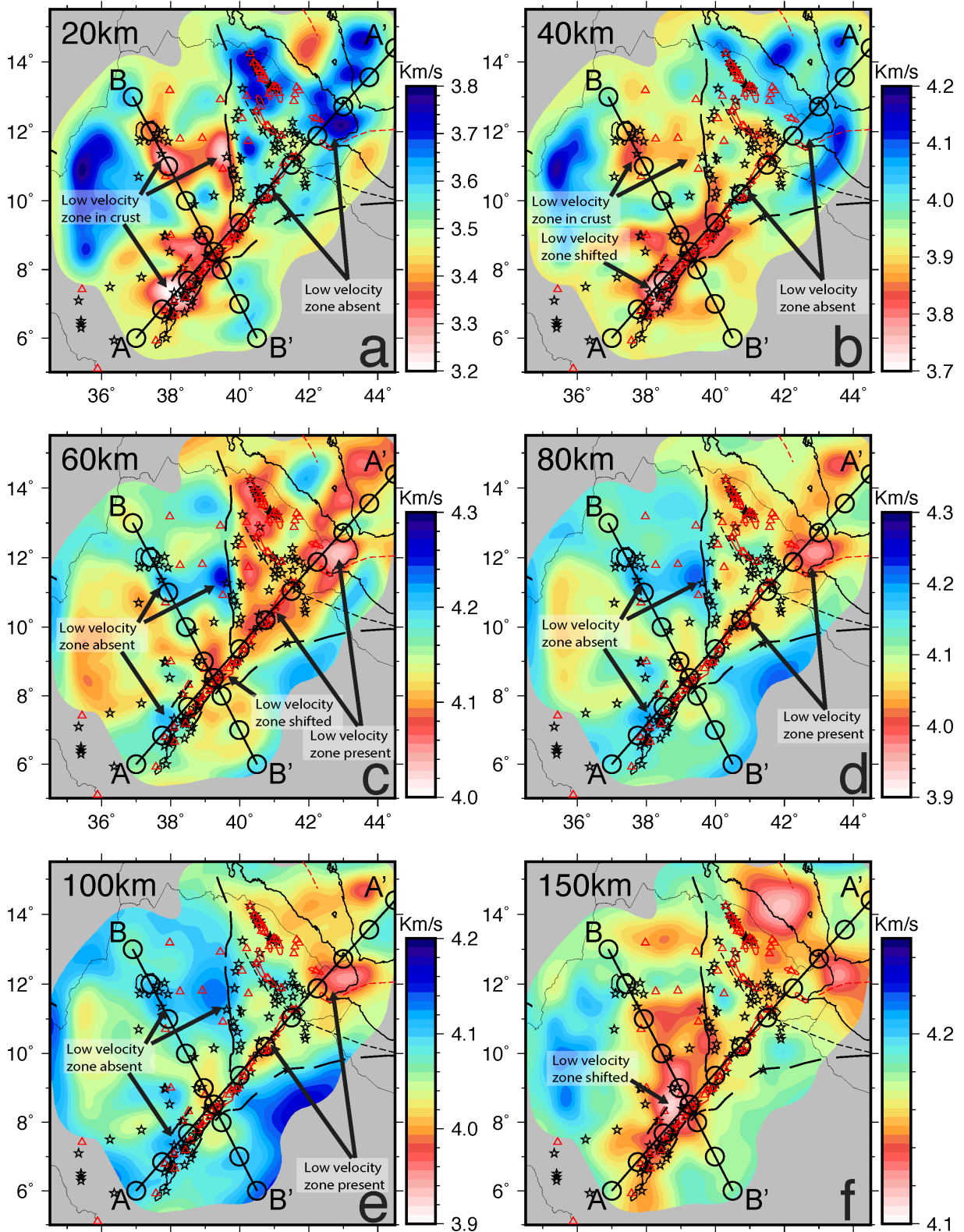


1

2 *Figure 7 Phase velocity maps generated by tomographically inverting dispersion curves from ambient noise (10, 20, 26s) and*

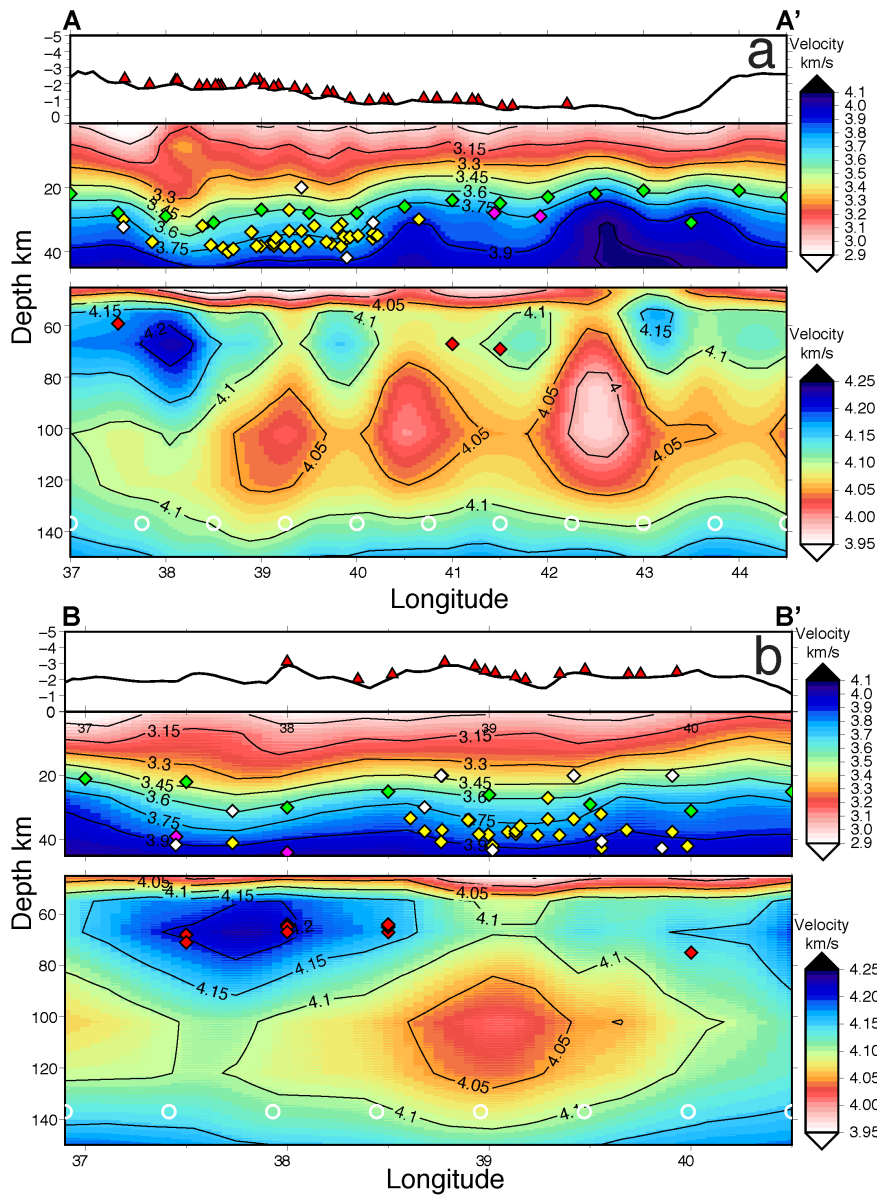
3 *teleseisms (29, 71 and 100 s). Models have been cropped to the 2σ standard error contour. Pink and red colours are slower*

- 1 velocities and blue faster velocities. Thick black lines indicate border faults, red polygons crustal magmatic segments, dashed
- 2 lines the Tendaho-Goba'ad discontinuity (TGD), red triangles volcanoes, and black stars geothermal activity.



3

1 Figure 8 Interpolated absolute shear wave velocity at 20, 40, 60, 80, 100 and 150 km. Models have been cropped to the
 2 standard error contour. Pink and red colours indicates slower velocities and blue faster velocities. Thick black lines indicate
 3 border faults. 2 profiles (thin black lines with black rings) are the cross-section locations for Figure 9. Red polygons indicate
 4 crustal magmatic segments, red triangles volcanoes, and black stars geothermal activity. Arrows indicate areas discussed in
 5 the text. See supplementary Figure S6 for slices at 10 km depth intervals.



6
 7 Figure 9 Cross-sections through the interpolated absolute shear velocity depth slices a) along the rift and b) across the rift
 8 (See Figure 8 for locations). The cross-sections have been split into crustal section (0-45 km depth, top panel) and mantle
 9 (45-150 km depth, bottom panel) for display purposes. Red and pink colours indicate slower velocities, and blue faster
 10 velocities. Red triangles above section indicate quaternary volcanoes with topography as black line. Thin lines are velocity
 11 contours and diamonds represent previous receiver function results for the Moho in top sections (green (Lavayssière et al.,

- 1 2018), *magenta* (Hammond, et al., 2011), *yellow* (Stuart et al., 2006), *white* (Ogden et al., 2019)) and *LAB* in bottom sections
- 2 (*red* (Lavayssière et al., 2018)). *White rings are the same as black rings in Figure 8, for location reference.*