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This manuscript has been submitted for publication in **Journal of Quaternary Science**. This pre-print has **not undergone peer-review** and subsequent versions of the manuscript may differ from this version. If accepted, the final version will be available via a DOI link on this page. Please contact the corresponding author by email with any queries – [pennos4@gmail.com](mailto:pennos4@gmail.com). Prepared for EarthArxiv on 17<sup>th</sup> December 2021.

2

3 **Deciphering late Quaternary climatic histories from the Hermes Cave**  
4 **speleothem record, Corinth Rift, Greece**

5

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14

15 **Keywords:** speleothem, paleoclimate, Late Quaternary, Corinth Gulf, Eastern Mediterranean

16

17 **Abstract**

18 The Greek peninsula is located at the crossroads of several major atmospheric circulation patterns  
19 and is consequently characterized by highly variable climatic conditions, making it an important  
20 location to examine past climate dynamics. Over the last decades, the focus of many studies in the  
21 region has been to unravel Holocene paleoclimatic oscillations and their impact on the  
22 development of ancient civilizations using terrestrial archives and especially speleothem records.  
23 In this study we contribute to the regional climate record over the early Holocene and late

24 Quaternary using a speleothem from the Hermes Cave located at the Corinth Rift shoulder in  
25 southern Greece. Our stalagmite grew over two distinct periods, from ~127 to 105 ka and from 18  
26 to 8 ka separated by a distinct hiatus. We have examined its growth history, stable isotope  
27 geochemistry and elemental composition. Higher growth rates are observed during the Eemian and  
28 the early Holocene and are attributed to high water recharge implying humid conditions. A gradual  
29 isotopic enrichment before the growth hiatus of the stalagmite suggests a gradual reduction in  
30 water infiltrating the system that may have been related to glacier advance. Our record correlates  
31 with other paleoclimate records from the broader area confirming and extending a pattern of  
32 coherent changes in paleoclimate across the Eastern Mediterranean basin.

33

## 34 **1. Introduction**

35 Climate in the Mediterranean Basin (MB) is a complex result of the conjunction of several  
36 atmospheric systems: westerlies from the North Atlantic Ocean, subtropical high-pressure systems  
37 originating over North Africa's arid zones, the Siberian High pressure system (SH), the North  
38 Atlantic Oscillation (NAO) (Lionello et al., 2006 ; Xoplaki et al., 2000) and the African and Asian  
39 Monsoons (Lionello and Galati, 2008). The NAO in particular strongly impacts winter  
40 atmospheric circulation patterns in the MB, with subsequent effects on river runoff (e.g. Tsimplis  
41 et al., 2006; Zerefos et al, 2011). Examining climatic patterns over the past 500 years, Luterbacher  
42 et al. (2006) concluded that a negative NAO index is related to wet and cool conditions in the MB,  
43 while a positive NAO index is associated with strong westerlies at high and mid latitudes, and dry  
44 and warm conditions in the MB.

45 However, the interplay of the atmospheric systems does not have the same intensity across the  
46 whole length of the MB. Luterbacher and Xoplaki (2003) suggest that there is a substantial  
47 differentiation between the Eastern Mediterranean basin (EMB) and the Western and Central  
48 Mediterranean. For example, winter air temperature in the Eastern part appears to be negatively  
49 correlated with the NAO index, in contrast to the Western and Central part where there seems to  
50 be a small positive correlation (Zerefos et al., 2011).

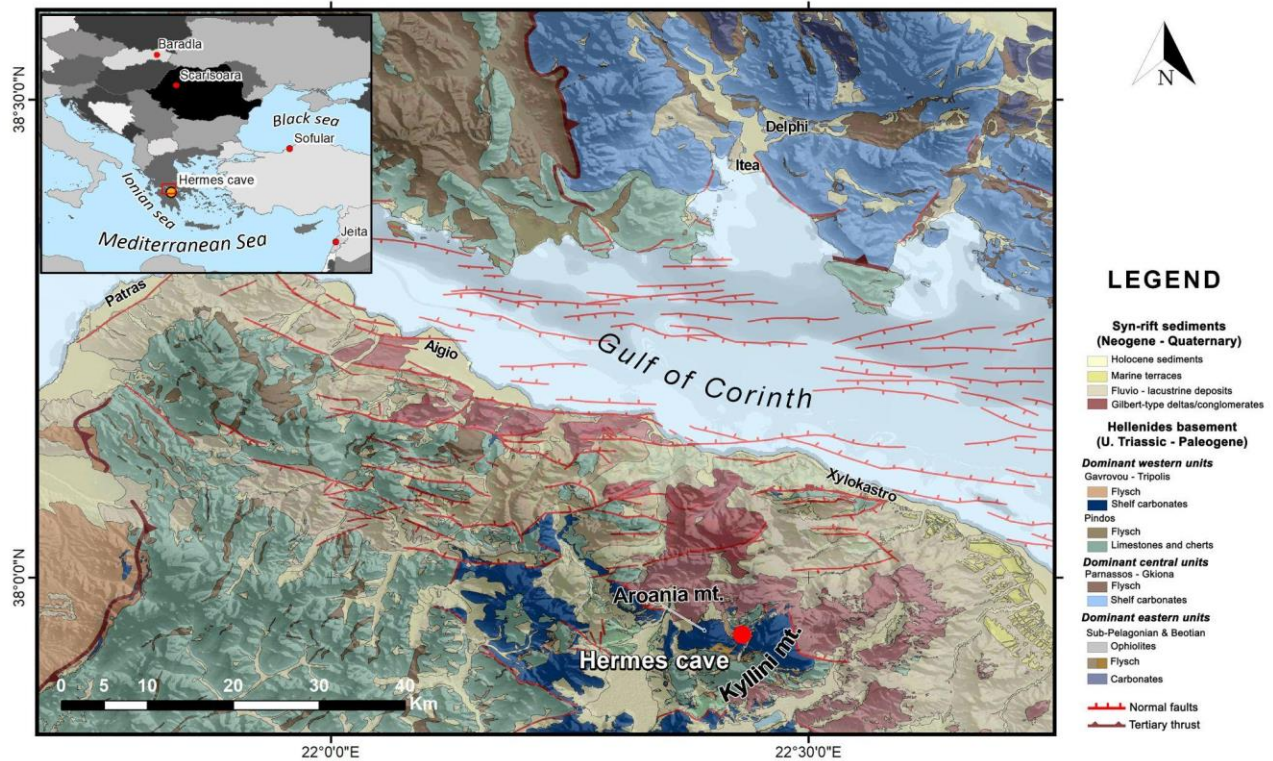
51 Climate research continues in the MB, utilizing many different archives from the area. Many are  
52 based on marine proxies, and aim to understand the paleoceanographic evolution in the EMB and  
53 the prevailing paleoclimatic conditions driving that evolution (e.g. Triantaphyllou et al., 2009;  
54 Koukousioura et al., 2012; Kouli et al., 2012; Rohling et al., 2015; Triantaphyllou et al., 2015;  
55 Gogou et al., 2016; Rohling et al., 2019). Terrestrial records can also provide insight, in some  
56 cases more directly, on the paleoclimatic evolution of the area. For the wider Mediterranean area,  
57 existing records are many and diverse, including pollen studies (e.g. (Tzedakis et al., 2002;  
58 Tzedakis et al., 2006; Tzedakis, 2010; Jones et al., 2012; Milner et al., 2013), speleothems (e.g.  
59 Bar-Matthews et al., 2000; Fleitmann et al., 2009; Psomiadis et al., 2009; Finné et al., 2014; Finné  
60 et al., 2015; Nehme et al., 2015; Nehme et al., 2018; Psomiadis et al., 2018; Regattieri et al., 2018;  
61 Peckover et al., 2019; Regattieri et al., 2020), geomorphic indexes (e.g. Styllas et al., 2018;  
62 Leontaritis et al., 2020) and clastic sedimentary sequences (e.g. Lespez et al., 2017; Styllas and  
63 Ghilardi, 2017; Katrantsiotis et al., 2019; McNeill et al., 2019; Pennos et al., 2021). Despite this  
64 plethora of publications, most of the studied records do not present continuous temporal coverage  
65 for the Quaternary, and/or focus mainly on periods relevant to archaeology, when ancient  
66 civilizations flourished.

67 Here we aim to investigate the paleoclimatic evolution of the area from a stalagmite collected from  
68 the Corinth Rift shoulder in southern Greece which covers parts of the late Quaternary and early  
69 Holocene. For these periods, this is the first published speleothem record for the Peloponnese/Gulf  
70 of Corinth region. We compare our results with other speleothem records from the broader area to  
71 understand regional scale climate dynamics, and to compliment the findings of the recent IODP  
72 Expedition 381 (McNeill et al., 2019) to decipher climate forcings affecting fluvial sediment fluxes  
73 in the Gulf of Corinth.

74

## 75 **2. Setting**

76 Hermes cave (HC) is located on Kyllini mountain at 1614 m a.m.s.l. close to the Ziria ski resort in  
77 the Peloponnese peninsula in southern Greece (Fig.1). It is named after the ancient god Hermes,  
78 who, according to Greek mythology, was born and raised inside the cave. The cave entrance is  
79 located on a cliff facing towards Flavouritsa valley and the Gulf of Corinth. The cave has been  
80 known since antiquity, and has been visited by numerous people throughout the years who caused  
81 extensive damage to the speleothems.



82

83 Figure 1. Regional geological map modified from Skourtsos et al. (2017). Inset shows the  
 84 location of caves presented in Figs. 6 & 7.

85

86 HC formed in upper Triassic-Paleogene limestones belonging to the Gavrovou-Tripoli geotectonic  
 87 zone of the Hellenides orogenic belt (e.g. Skourtsos et al., 2017; Gawthorpe et al., 2018). These  
 88 carbonates outcrop on the southern flank of the Corinth extensional rift, have undergone brittle  
 89 deformation and are densely fractured. The extensive fracturing has allowed surface water to  
 90 penetrate into the limestone and initiate cave formation. Hermes Cave is elongated in a NE-SW  
 91 direction and the cave floor dips steeply,  $> 45^\circ$ , toward the SW, following the bedding of the  
 92 limestones. Speleothem formations are extensive, with stalagmites intercepting rock debris and  
 93 forming small terraces in some places.

94 The modern climate in the broader area corresponds to “Mediterranean Climate” with 9.7°C  
95 mean annual temperature and 1296 mm mean annual precipitation (Mamara et al., 2017).  
96 Regional climate is further characterized by mild and wet winters (December – March) that  
97 contribute most of the annual precipitation, with dry, warm summers where occasional  
98 convective precipitation occurs, resulting in strong, stormy rainfall events (Xoplaki et al., 2000;  
99 Feidas et al., 2007). Xoplaki et al. (2004) concluded that although there is large spatio-temporal  
100 variability in the region’s winter precipitation, a significant portion (30%) is explained by large-  
101 scale atmospheric circulation. This pattern is clearly observed when winter NAO-driven  
102 depressions move northeast from the North Atlantic and release rainfall on western Greece  
103 (Styllas et al., 2015 and references therein).

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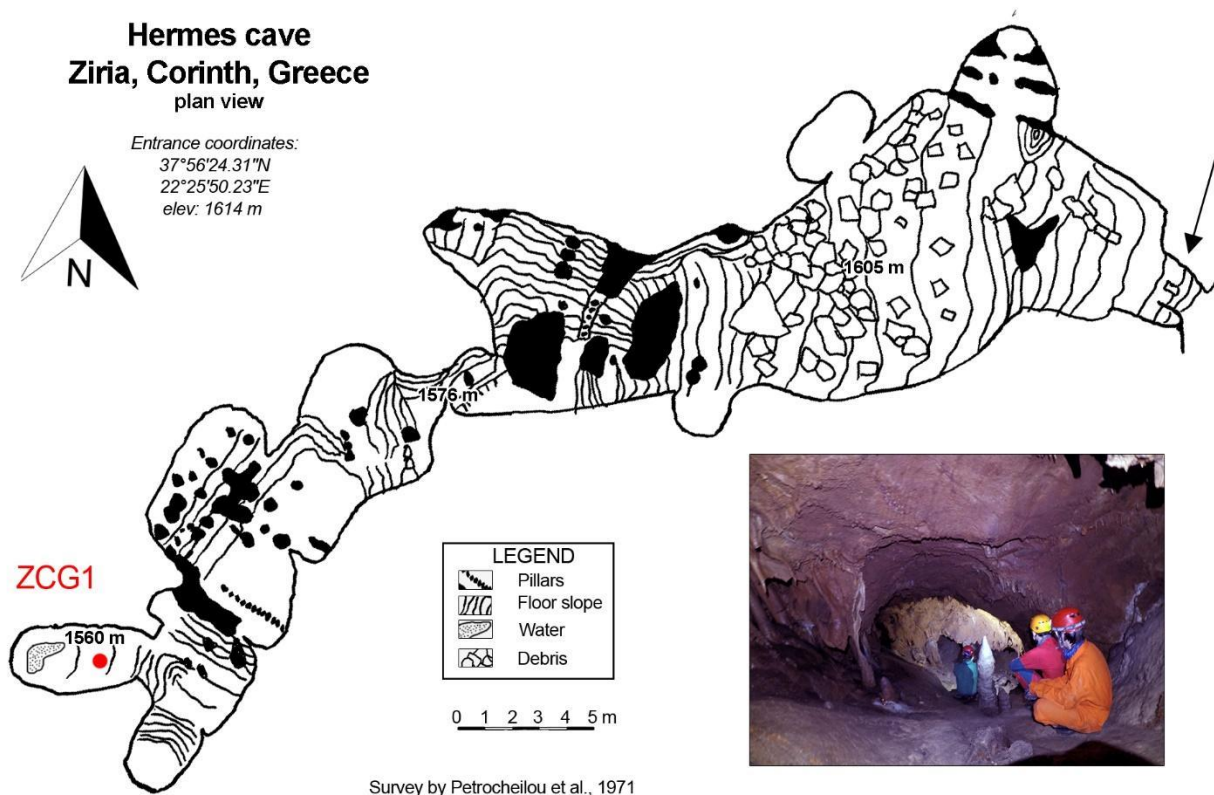
### 105 **3. Methods**

#### 106 **3.1 Speleothem sampling**

107 In order to collect a stalagmite that is active throughout the year, we visited the cave during the  
108 dry season (late August). We collected an active stalagmite (ZCG1) in situ from a relatively small  
109 chamber far from the entrance (Fig.2), where no air draft was evident, to exclude drip water  
110 evaporation. The sampling chamber is a small blind passage, 6 m long, 2.5 m high and 4 m wide,  
111 that dips steeply towards the main development axis of the cave. It is located 50 m from the  
112 entrance and ~ 60 m below the surface. The 30 cm tall stalagmite was extracted from the cave and  
113 was later cut in half along the growth axis. One part of the stalagmite was stored for reference, and  
114 from the other half a 2 cm thick slab was extracted. From macroscopic observation it is evident

115 that ZCG1 is a densely laminated stalagmite. Most of the laminae are opaque, with no visible signs  
116 of diagenesis of the calcite fabric (Fig. 3).

117

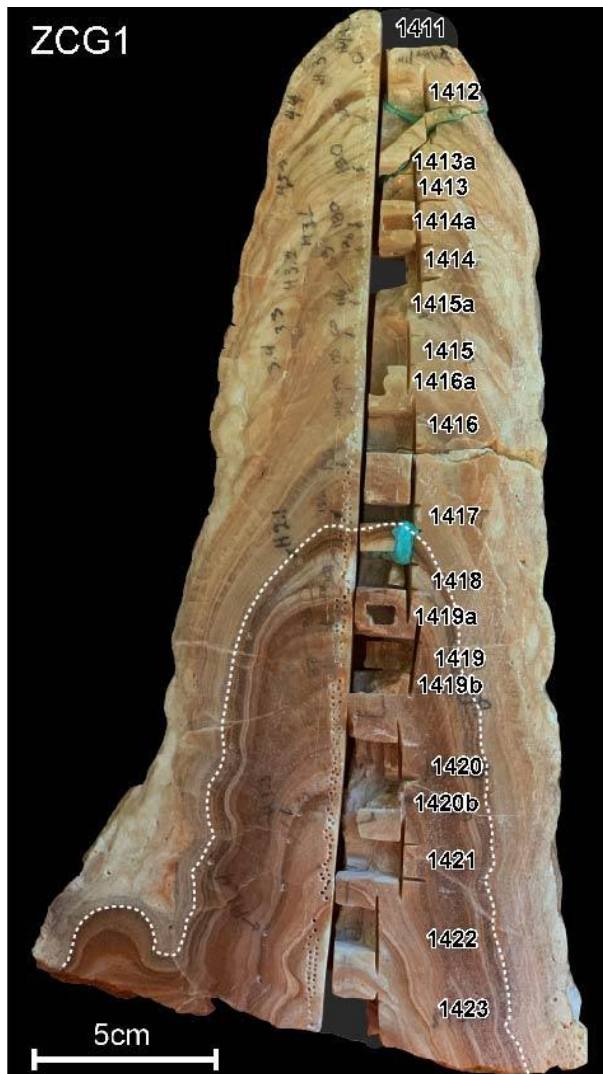


118

119 Figure 2. Plan view of Hermes Cave from Petrocheilou (1972). Red dot shows the position of the  
120 stalagmite. Inset photograph shows the chamber where stalagmite ZCG1 was formed.

121





122

123 Figure 3. Polished section of the ZCG1 stalagmite. Small pits indicate stable isotope sampling  
124 position. Rectangular trenches indicate the position of the Th/U samples with their corresponding  
125 laboratory identification numbers. White dashed line indicates the position of the growth hiatus.

126

### 127 3.2 Th/U dating

128 To construct the age-depth model, we extracted seventeen samples along the growth axis either by  
129 drilling or milling. Samples were heated at 650°C for 4 hours to remove organic material. Chemical

130 separation of U and Th was adapted from Edwards (1988). The samples were spiked with a mixed  
131  $^{233}\text{U}$ - $^{236}\text{U}$ - $^{229}\text{Th}$  solution (calibrated against Harwell Uraninite (HU-1) solution considered at  
132 secular equilibrium) and dissolved in concentrated  $\text{HNO}_3$ . Column chemistry cleaned Fe solution  
133 was added, and Fe precipitates were formed by dropwise addition of  $\text{NH}_4\text{OH}$ . Fe precipitates were  
134 rinsed with 18.2 M $\Omega$  deionized water, re-dissolved with 6 M HCl and loaded onto AG1X8 resin  
135 for U-Th separation. Uranium and Th were further purified through consecutive passes onto  
136 U/TEVA and AG1X8 resins, respectively. Th and U isotopes were analyzed on a Nu Plasma II  
137 MC-ICP-MS. Mass bias was corrected by standard-sample bracketing using HU-1 solution. Blank  
138 concentrations were  $^{238}\text{U} < 0.2 \text{ ng}$ ;  $^{234}\text{U} < 30 \text{ fg}$ ;  $^{232}\text{Th} < 11 \text{ pg}$ ;  $^{230}\text{Th}$  was below the detection limit.  
139 Activity ratios were calculated using decay constant values from Bourdon et al. (2003). Ages were  
140 calculated using Isoplot 3.75 (Ludwig, 2003) without decay constant uncertainties. Long term  
141 analytical reproducibility of the HU-1 solution (n = 28, measured over 15 months) is  $\text{AR}(^{234}\text{U}/^{238}\text{U})$   
142  $1.002 \pm 0.001$  and  $\text{AR}(^{230}\text{Th}/^{238}\text{U}) 1.003 \pm 0.002$  (2 SD). All U-series data reported in tables  
143 are presented with  $\pm 2\sigma$  uncertainty, propagated to include analytical and spike calibration  
144 uncertainties, unless otherwise indicated.

145

### 146 **3.3 Stable isotopes**

147 Sampling for stable isotope analyses was performed by milling with a 0.7 mm diameter bit along  
148 the growth axis, with 1 mm step, resulting in a total of 218 samples.  $\delta^{18}\text{O}$  and  $\delta^{13}\text{C}$  analyses were  
149 conducted at the University of Bergen (FARLAB) using a MAT 253 mass spectrometer coupled  
150 to an automated Kiel IV preparation device. Approximately 50 ( $\pm 20$ )  $\mu\text{g}$  of sample powder was  
151 reacted with concentrated ortho-phosphoric acid ( $\text{H}_3\text{PO}_4$ ) at a constant 70 °C. Isotope values are  
152 reported on the Vienna Pee Dee Belemnite (VPDB) scale calibrated using the scale reference

153 standard NBS 19 (value 1.95‰ and 2.2‰ for  $\delta^{13}\text{C}$  and  $\delta^{18}\text{O}$ , respectively) together with NBS 18  
154 (-5.01‰ and -23.2‰ for  $\delta^{13}\text{C}$  and  $\delta^{18}\text{O}$ , respectively; Friendman et al., 1982; Hut, 1987; Stiltcher,  
155 1993; Coplen et al., 2006). Analytical reproducibility (1s), based on replicate measurements of the  
156 in-house carrara marble standard CM12, and spanning the same mass range and run over the same  
157 analysis period (n=128), was 0.06 and 0.03 for  $\delta^{18}\text{O}$  and  $\delta^{13}\text{C}$ , respectively. Finally, we performed  
158 Hendy's tests (Hendy, 1971) within 4 distinct laminae as a first check for correlation between  $\delta^{18}\text{O}$   
159 and  $\delta^{13}\text{C}$ , providing information about possible kinetic effects during precipitation.

### 160 **3.4 $\mu\text{-XRF}$**

161 Relative elemental composition was determined by x-ray fluorescence using an Itrax core scanner.  
162 Core scanning was conducted on the 2 cm thick stalagmite slab, along the growth axis at 1 mm  
163 intervals using a Mo x-ray tube (Croudace et al., 2006). Exposure time was 10 s, power supply  
164 was 30 kV/55 mA. The output was later processed using Q-spec software. Following the calcite  
165 growth modelling approach of Wong et al. (2011), we interpret high Sr/Ca values as representing  
166 summer season speleothem growth and low values as winter growth, respectively.

167

## 168 **4. Results**

### 169 **4.1 Age Model and growth rate**

170 The Th/U analysis produced ages ranging from  $7.0 \pm 4.2$  ka to  $127.9 \pm 52.5$  ka (Table 1). To  
171 generate a growth model we used the Mod-Age software (Hercman and Pawlak, 2012) that  
172 employs a weighted scatterplot smoothing (LOESS) interpolation to build the chronological  
173 model. Three Th/U dates were indicated by the software to be outliers (see Fig. 4b) and were not  
174 used for the model. The growth period covers the period 133.2 to 7.4 ka but is interrupted by a

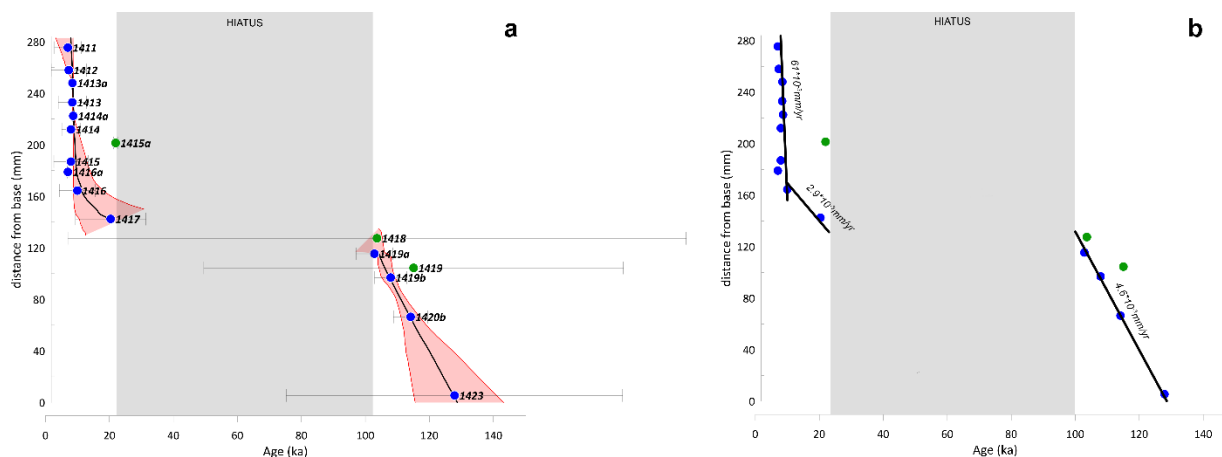
175 hiatus in growth at ~ 135 mm from the base. This hiatus extends between approximately 105 to 23  
 176 ka (Fig. 4a).

177 Table 1. Activity ratios and age calculations from ZCG1 stalagmite.

Sample ID	Depth (mm)	238U $\mu\text{g/g}$	2 $\sigma$ e	(230Th/238U)	2 $\sigma$ e	(234U/238U)	2 $\sigma$ e	Age uncr (ka)	2 $\sigma$ e	(232Th/238U)	2 $\sigma$ e	230Th/232Th	2 $\sigma$ e	Age cr (ka)	2 $\sigma$ e (ka)
1411	0.85	222	7	0.0683	0.0338	1.078	0.074	7.1	4.2	0.0040	0.0020	17.2	8.5	7.0	4.2
1412	2.6	197	7	0.0727	0.0451	1.099	0.093	7.4	5.5	0.0053	0.0033	13.7	8.5	7.3	5.5
1413a*	3.6	197	1	0.0818	0.0005	1.079	0.006	8.6	0.1	0.0042	0.0003	19.6	0.1	8.5	0.3
1413	5.1	196	7	0.0841	0.0352	1.103	0.078	8.6	4.4	0.0078	0.0034	10.8	4.5	8.4	4.4
1414a*	6.15	219	1	0.0866	0.0004	1.109	0.005	8.8	0.1	0.0037	0.0003	23.1	0.1	8.7	0.3
1414	7.2	219	7	0.0789	0.0222	1.095	0.064	8.1	2.9	0.0056	0.0017	14.1	4.0	8.0	2.9
1415a*	8.25	229	1	0.2021	0.0012	1.095	0.006	22.1	0.2	0.0172	0.0035	11.7	0.1	21.9	0.6
1415	9.7	186	6	0.0795	0.0447	1.109	0.078	8.1	5.4	0.0040	0.0023	20.1	11.3	8.0	5.4
1416a*	10.5	344	1	0.0702	0.0004	1.112	0.006	7.1	0.1	0.0023	0.0002	30.4	0.2	7.0	0.2
1416	11.95	176	6	0.1033	0.0431	1.123	0.094	10.5	5.6	0.0190	0.0084	5.4	2.3	10.0	5.5
1417	14.15	165	8	0.2129	0.0688	1.125	0.118	22.7	11.2	0.0915	0.0392	2.3	0.8	20.4	11.0
1418	15.65	77	8	0.7299	0.1407	1.154	0.183	105.6	97.0	0.0813	0.0723	9.0	1.6	103.6	96.6
1419a*	16.85	117	0	0.7040	0.0040	1.125	0.007	104.1	1.5	0.0694	0.0492	10.1	0.1	102.8	5.8
1419	17.95	107	5	0.7381	0.1202	1.101	0.096	117.7	65.8	0.1053	0.0924	7.0	1.3	115.1	65.6
1419b*	18.7	126	1	0.7063	0.0045	1.112	0.007	107.0	1.7	0.0505	0.0359	14.0	0.1	107.9	5.0
1420b*	21.75	158	1	0.7541	0.0043	1.140	0.006	114.1	1.7	0.0972	0.0737	7.8	0.0	114.1	5.4
1423	27.85	243	8	0.7381	0.0879	1.057	0.067	128.3	52.5	0.0151	0.0126	48.7	5.9	127.9	52.5

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179



180

181 Figure 4. a) Age depth model (blue dots = dates, horizontal error bars = 2 $\sigma$  uncertainty; red  
 182 shading = age model 2 $\sigma$  uncertainty). Outliers removed from the model shown in green. b)

183 Growth rates of the stalagmite indicated per major intervals.

184

185 To estimate the growth rate for each interval, we employed a linear regression approach. Three  
 186 different linear regression lines with high coefficient ( $R^2 > 0.85$ ) were generated (Fig. 4b). The

187 oldest part of the stalagmite developed between 127.9 ka to ca 105 ka at a rate of  $4.6 \times 10^{-3}$  mm/yr.  
188 Following the growth hiatus, the rate between 20.4 ka to 10 ka B.P was lower at  $2.9 \times 10^{-3}$  mm/yr.  
189 The youngest part of the stalagmite, 10 ka to 7 ka, formed at a much higher growth rate of  $6.1 \times$   
190  $10^{-3}$  mm/yr (Fig. 4b).

191 The age estimate for the oldest sample (1423, Table 1, Fig. 3) has relatively poor precision. This  
192 age estimate is critical for constraining the age of the oldest part of the record, thus the  
193 measurement was repeated with a new portion of stalagmite material. Unfortunately, the result of  
194 this repeat measurement was not an improvement over the original, likely due to detrital  
195 contamination. To compensate for the resulting age model uncertainty on ages older than  $\sim 115$   
196 ka, we compare the stable isotope results grouped over the Last Interglacial (116 – 129 ka;  
197 Tzedakis et al., 2018) and Holocene (to  $\sim 12.5$  ka; Styllas et al., 2018) according to the median age  
198 model dates, and compare to the same data but given ages at the maximum and minimum values  
199 of the age model error envelope (see discussion).

200

#### 201 **4.2 $\delta^{18}\text{O}$ and $\delta^{13}\text{C}$ record**

202 The results from the Hendy tests (Hendy, 1971) suggest isotopic equilibrium conditions during the  
203 majority of ZCG1 deposition (see supplement). It is evident that the variations of  $\delta^{18}\text{O}$  and  $\delta^{13}\text{C}$   
204 show no positive correlation along each of the tested layers, and there is no enrichment in  $^{18}\text{O}$   
205 towards the external part of the stalagmite. The exception to this is near the termination of growth  
206 before the hiatus, which is marked by an isotopic enrichment (Fig. 5) that might be indicative of  
207 kinetic effects, possibly due to a change in climatic conditions.

208  $\delta^{18}\text{O}$  values range from -8.8 to -6.2 ‰ (Fig. 5). For the first 10 ka of stalagmite development the  
209 values show minor variations from -7.7 to -6.7 ‰ until ca 123 ka. This period is followed by a  
210 period of lower values (down to -8.7 ‰) lasting nearly 3000 a. An increase in  $\delta^{18}\text{O}$  values follows,  
211 which peaks at 118 ka B.P (maxima -6.8 ‰). This period is trailed by a decrease of  $\delta^{18}\text{O}$  values  
212 until ca 115 ka where the values oscillate between -7.9 and -7.3 ‰ until 112 ka. From 106 ka to ~  
213 105 ka, when the growth of the stalagmite halted,  $\delta^{18}\text{O}$  values increase again (maxima -6.2 ‰)  
214 but, as stated above, this section of the sample might be affected by kinetic fractionation, thus it  
215 will not be interpreted in terms of climate.

216 At ~ 20 ka, stalagmite growth resumed; concurrently, the  $\delta^{18}\text{O}$  record exhibits a slight increase,  
217 from -8.3 to -7.9 ‰, for the succeeding 3.5 ka. This period is followed by an abrupt decrease in  
218  $\delta^{18}\text{O}$  to -8.5 ‰, followed by an interval of minor oscillations (between -8.5 to -8 ‰) until ca 11  
219 ka. Low values of  $\delta^{18}\text{O}$  are recorded during the next growth interval until ca  $10.7 \pm 3.5$  ka, followed  
220 by a general increase in  $\delta^{18}\text{O}$  to -7.4 ‰ that peaks at ca 8.5 ka. This increase in the  $\delta^{18}\text{O}$  values is  
221 interrupted by two prominent negative peaks at  $9.5 \pm 3$  and  $9.1 \pm 3$  ka ( $\delta^{18}\text{O}$  values -8.72 and -8.7  
222 ‰ respectively). Finally, at the youngest part of the stalagmite,  $\delta^{18}\text{O}$  values oscillate from -8.4 to  
223 -7.8 ‰ when the stalagmite growth halts at ca. 8 ka.

224  $\delta^{13}\text{C}$  values range between -10.4 to -3.5 ‰ VPDB (Fig.5). From 126 ka to 118 ka,  $\delta^{13}\text{C}$  values  
225 show small variations from -8.7 to -8.2 ‰. From 118 to 113 ka, the carbon isotope record exhibits  
226 quite constant values around -6.4 ‰. After a short and sharp  $^{13}\text{C}$  enrichment at 113 ka, a period of  
227 ca 4 millennia with minor oscillations in  $\delta^{13}\text{C}$  values follows. During the next evolutionary stage  
228 of the stalagmite, a strong increase in  $\delta^{13}\text{C}$  to around -3.7 ‰ is observed and lasts for ~ 2 millennia.  
229 The following period begins with a sharp decrease of  $\delta^{13}\text{C}$  at ca 111 ka BP and continues with  
230 slight variations overprinted on this long term  $^{13}\text{C}$ -depletion until the stalagmite growth is

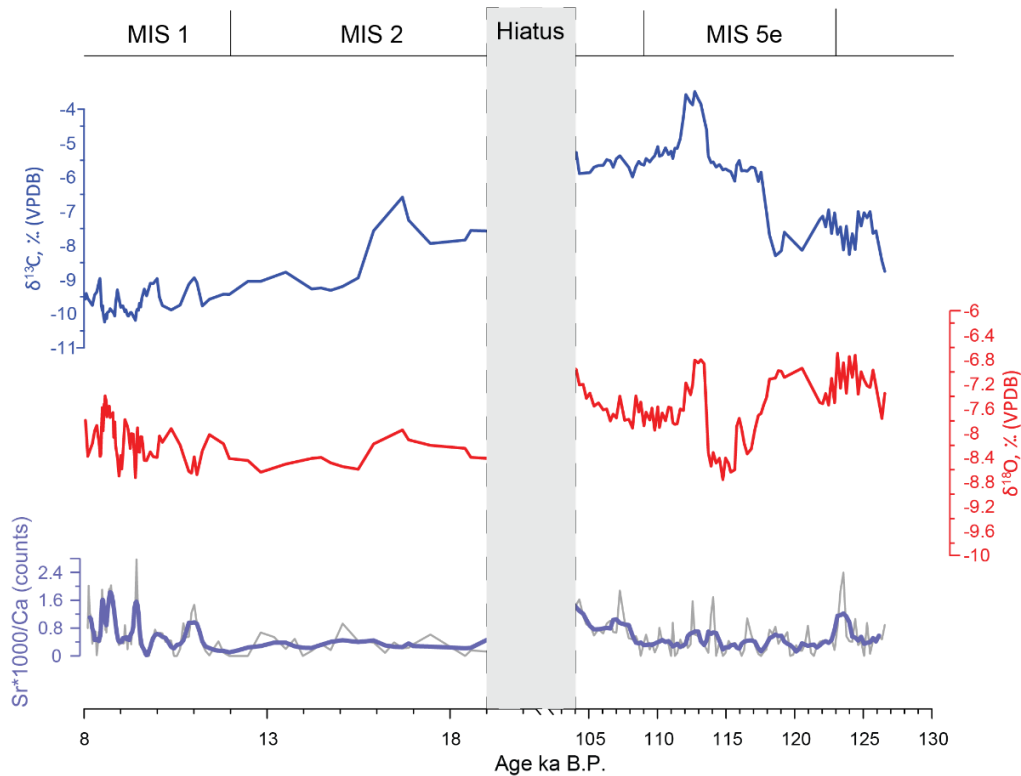
231 interrupted at ca 105 ka. Past the growth hiatus, the  $\delta^{13}\text{C}$  record presents constant values for 2 ka,  
232 until ca. 17 ka, and then a positive shift for another ~ 2 ka (until ca. 15.9 ka). From 15.9 to 11.2 ka,  
233  $\delta^{13}\text{C}$  decreases from -7.6 to -9.7 ‰. At the youngest part of the stalagmite the  $\delta^{13}\text{C}$  values are  
234 marginally decreasing, with minor fluctuations on the general trend interrupted by four prominent  
235 peaks at 10.5, 9.5, 8.9 and 8.4 ka.

236

### 237 **4.3 $\mu$ -XRF**

238 The XRF measurements generated a detailed trace element profile of the stalagmite. The x-ray  
239 fluorescence signal in dense material such as speleothems is strong; consequently, the scan  
240 returned good counting results for numerous elements (e.g. Al, Si, P, Sr, Ti, Fe). Here we focus  
241 mainly on the Sr/Ca ratio since it can be used as a paleoclimate proxy (e.g. Kluge et al.; Fairchild  
242 and Treble, 2009; Wong et al., 2011; Fairchild and Baker, 2012). Six distinct intervals exhibit an  
243 excess in Sr compared to Ca at 123, 120-117, 114, 113, 108-106.5. and ca 105 to 104 ka (Fig. 5).  
244 After the depositional hiatus and up until 11 ka BP, Sr/Ca shows minor fluctuations. From 11 ka  
245 until ca 9.8 ka, there are two intervals where again an excess in Sr is observed (at 11 and 9.8  
246 respectively). From 9.8 ka until the end of the record, there is a general increase in the Sr/Ca ratio  
247 with four peaks at 9.5, 8.9, 8.5 and 8.1 ka.

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249

250 Figure 5. Carbon isotope (blue line), oxygen isotope (red line) and Sr/Ca  $\mu$ -XRF (black and  
 251 purple lines) profiles along ZCG1 stalagmite.

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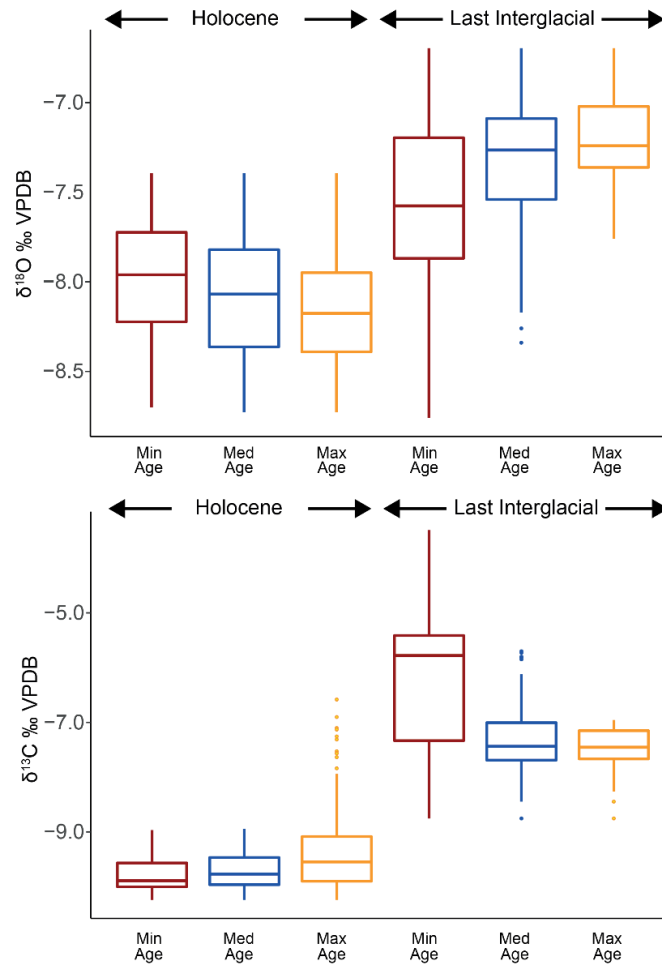
253 **5. Discussion**

254 **5.1 Major climate patterns**

255 ZCG1 stalagmite formation initiated around 127 ka, indicating the establishment of conditions that  
 256 favored the speleothem deposition in this karstic setting. This period correlates with the beginning  
 257 of the maximum interglacial conditions during MIS5 that enabled high growth rates recorded in  
 258 stalagmites across Europe (Drysdale et al., 2009; Genty et al., 2013; Demeny et al., 2017). In the  
 259 absence of monitoring data, we base our proxy interpretation on similar records (in terms of  
 260 recharge, speleothem growth period, climatic conditions) from the wider region. We interpret the



261 carbon isotope signal from the ZCG1 stalagmite as reflecting soil activity, which itself is dependent  
262 on both temperature and precipitation regime (e.g. Genty et al., 2001). The  $\delta^{18}\text{O}$  of calcite is more  
263 difficult to interpret in terms of both specific season it refers to and whether it is indicating climatic  
264 conditions or source  $\delta^{18}\text{O}$  (precipitation and/or moisture source). Interpretation of the ZCG1 is not  
265 simple since it can be influenced by various climate variables, such as variations in surface and  
266 cave air temperatures, seasonality of precipitation, storm tracks and ice volume (e.g. Fleitmann et  
267 al., 2009 and references within). The extent these variables controlled our record is not fully clear.  
268 However, the similarity our record with other records in the area suggests that changes in the  
269 oxygen values reflect the combined effect of temperature and moisture source (e.g. Dansgaard,  
270 1964; McDemott, 2004; Nehme et al., 2015 and references within) and in the absence of  
271 quantification, we only marginally discuss it in our interpretation.



272

273 Figure 6. Box plots comparing Ziria speleothem stable isotope compositions from the Holocene  
 274 (to 11.7 ka) against the Last Interglacial (127 to 116 ka) by possible ages (med age = mean age,  
 275 also age model used; min age = minimum age at error envelope limit; max age = maximum age  
 276 at error envelope limit).

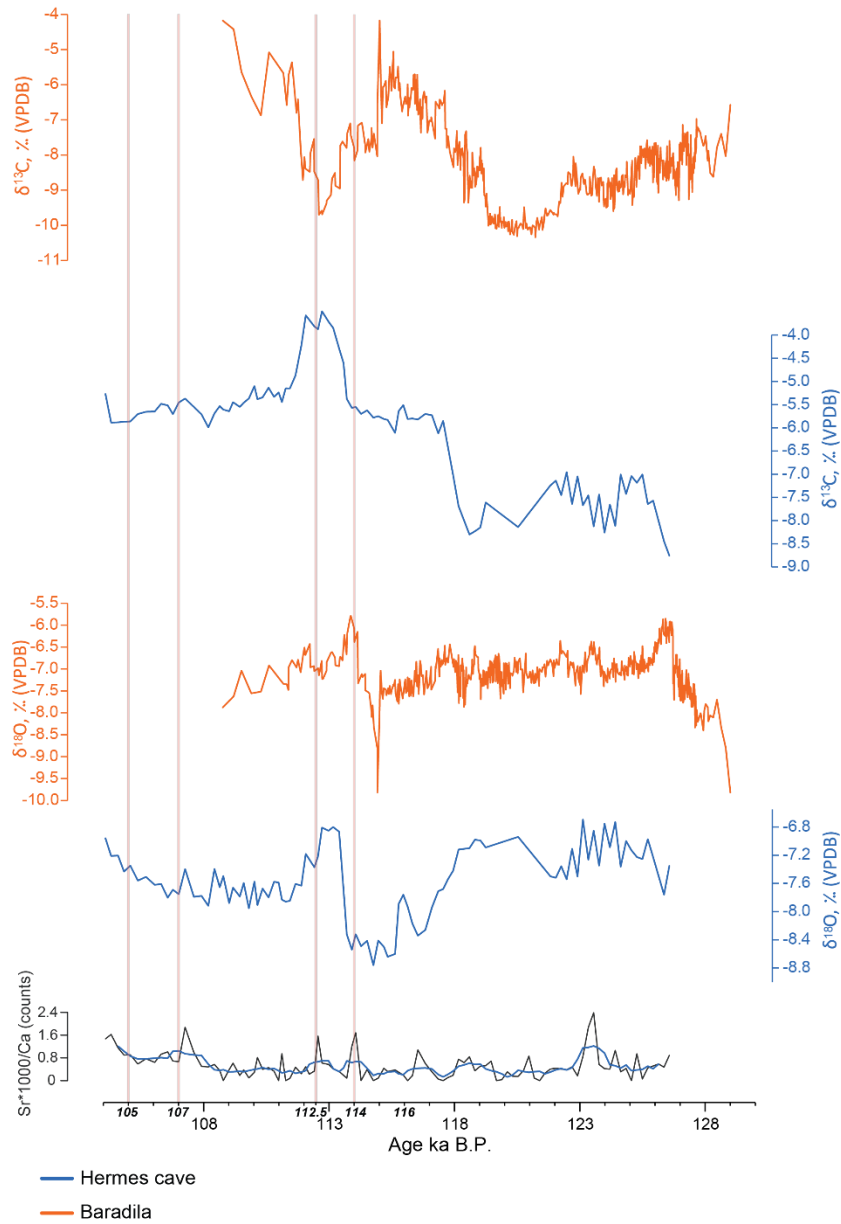
277

278 Fig. 6 compares the stable isotope data from the Holocene (to 11.7 ka) with those from the Last  
 279 Interglacial (116 to 127 ka), with the data grouped based on different options in the age model  
 280 (mean age, minimum age according to the error envelope, maximum age according to the error  
 281 envelope). As the oldest Th/U date has a large associated uncertainty, whereas it is required in

282 order to provide an age model to the base of the speleothem, this comparison aims to examine how  
283 the relationship between the range of isotopic compositions for each period changes with age  
284 model assignment. With the exception of  $\delta^{18}\text{O}$  compositions using minimum age values, all  
285 versions of the data are significantly different between the two time periods; each set exhibits  
286 similar differences in isotopic composition between the Holocene and Last Interglacial. This  
287 indicates that, at minimum, the different age possibilities for the oldest part of the Ziria stalagmite  
288 do not affect broad interpretations of climate for the Last Interglacial period, where the age model  
289 has the largest uncertainty. This test therefore demonstrates higher variability in stable isotope  
290 compositions, and thus likely climate, during the Last Interglacial compared to the Holocene,  
291 regardless of the selection of age model. A similar pattern has been observed in records from the  
292 North Atlantic and Southern European regions, e.g., speleothem records from Italy, marine cores  
293 and pollen records (Tzedakis et al., 2018).

294 Alongside the onset of initial stalagmite growth,  $\delta^{13}\text{C}$  values are increasing and  $\delta^{18}\text{O}$  values are  
295 decreasing (Fig. 7), suggesting that during this part of the interglacial cycle the climate was colder  
296 and drier in the region and had not yet reached stable interglacial conditions (Tzedakis et al., 2018).  
297 Following this period of instability and up to 117 ka,  $\delta^{13}\text{C}$  values exhibit two minima (at 122 and  
298 117 ka respectively) suggesting high soil activity that can be attributed to forest expansion due to  
299 humid conditions (Weiberg et al., 2016). The period between these two excursions exhibits higher  
300 values, possibly corresponding to a dryer period in the region. This is in contrast to Central  
301 European stalagmite records (e.g., Baradla cave; Demeny et al., 2017; Fig. 7) which indicate this  
302 period is considered the optimum within the interglacial (Govin et al., 2015) with high soil activity.  
303 The  $\delta^{18}\text{O}$  record for the same period exhibits low variability, suggesting no major changes in the  
304 source or amount of precipitation. The  $\delta^{18}\text{O}$  values during this period reach a minimum at 115 ka,

305 while concurrently the  $\delta^{13}\text{C}$  values increase, possibly implying low soil productivity and  
306 establishment of cold and dry conditions in the area towards the onset of the glacial period. The  
307 same trend for both  $\delta^{13}\text{C}$  and  $\delta^{18}\text{O}$  values continues until circa 112 ka with an interruption towards  
308 114 ka where the oxygen record exhibits a positive peak and the carbon record a negative peak.  
309 We suggest that these changes in the overall trend could be attributed to a period where cyclonic  
310 depressions forming over Africa are controlling the climate, similar to those argued to be  
311 responsible for the transport of Saharan dust to the area around this time (e.g. Stuut et al., 2009;  
312 Philandras et al., 2011; Remoundaki et al., 2011). The reduction in soil activity during this period  
313 is most likely related to a change in vegetation, as it corresponds chronologically to vegetation  
314 changes which occurred in the Balkan peninsula when low vegetation replaced thick interglacial  
315 forests (Tzedakis et al., 2004). This change in the type of vegetation slightly enhanced soil activity,  
316 likely resulting in the negative peak on the  $\delta^{13}\text{C}$  values at ~ 112 ka. Finally, the growth of the  
317 stalagmite continues up to 105 ka until the hiatus, albeit with slightly lower  $\delta^{13}\text{C}$  values (relative  
318 to the 112 ka peak), suggesting that the area was then experiencing a period with lower soil activity.  
319 The  $\delta^{18}\text{O}$  record during this final interval before the hiatus shows low variability and is similar to  
320 other records from central Europe (Kern et al., 2019).



321

322 Figure 7. Comparison of ZCG1 last interglacial part with Baradla cave record (Demeny et al.,

323 2017).

324

325 As warm and humid conditions are generally viewed as prerequisites for speleothem growth

326 (Dreybrodt, 1988; Baker and Smart, 1995; Genty et al., 2006; Nehme et al., 2020), inferences

327 about climate can potentially be made from the presence of speleothem growth hiatuses. Although  
328 the stalagmite growth hiatus might have occurred for different reasons, we suggest that in this case,  
329 the hiatus was the result of decreasing amounts of water infiltrating into the cave possibly due to  
330 the growth of small glaciers. The ZCG1 hiatus (~ 105 to 23 ka B.P; Fig. 3) spans approximately  
331 MIS4 to the LGM; local glacier expansion during this period is evidenced in deposits on the nearby  
332 Mount Chelmos, where glacial erratics have been observed, with emplacement dated throughout  
333 MIS 3 and the LGM (Leontaritis et al., 2020). The peak of Mt. Chelmos (also known as Aroania)  
334 is of a similar elevation to Mt. Kyllini, at 2355 m and 2376 m, respectively; the peaks are  
335 essentially next to each other, with Kyllini located about 15 km east of Chelmos. Speleothem  
336 growth interruptions may also be due to a change in the fluid pathways reaching the cave; while  
337 this cannot be ruled out, the fact that ZCG1 recovered and grew throughout the Holocene makes  
338 this option less likely. MIS3 and MIS4 are thought to generally have been wetter and colder in the  
339 Eastern Mediterranean (Bar-Matthews et al., 2003, 2019), indicating conditions more suited to the  
340 development of alpine glaciers, compared to the warmer temperatures of the preceding climate  
341 phases. In particular, the Eemian (MIS5e) is estimated at 9 – 11°C warmer in the eastern  
342 Mediterranean (e.g. Nehme et al., 2020); temperatures this much higher which would have driven  
343 the snowline to a much higher altitude compared to the cooler conditions of later  
344 glacial/interglacial cycles. These cold conditions likely extended to the LGM (Styllas et al., 2015;  
345 2018; Leontaritis et al., 2020 and references within), until at the LGM termination, the temperature  
346 and precipitation regime no longer fulfilled the requirements for glacier preservation. After local  
347 glaciers either shrank or disappeared, water was able to once again infiltrate Hermes Cave and  
348 ZCG1 growth resumed.

349 The post-LGM part of the ZCG1 record covers the interval from 18 to 8 ka. Our record exhibits  
350 higher temporal resolution towards the onset of the Holocene as a result of a higher growth rate  
351 that is most likely related to the increase of precipitation amount after ca 11.7 ka cal BP. The  $\delta^{18}\text{O}$   
352 record exhibits low variability between 18 and 11 ka, but after the onset of the Holocene, variability  
353 increases on an overall rising trend (likely linked to increasing air temperatures).

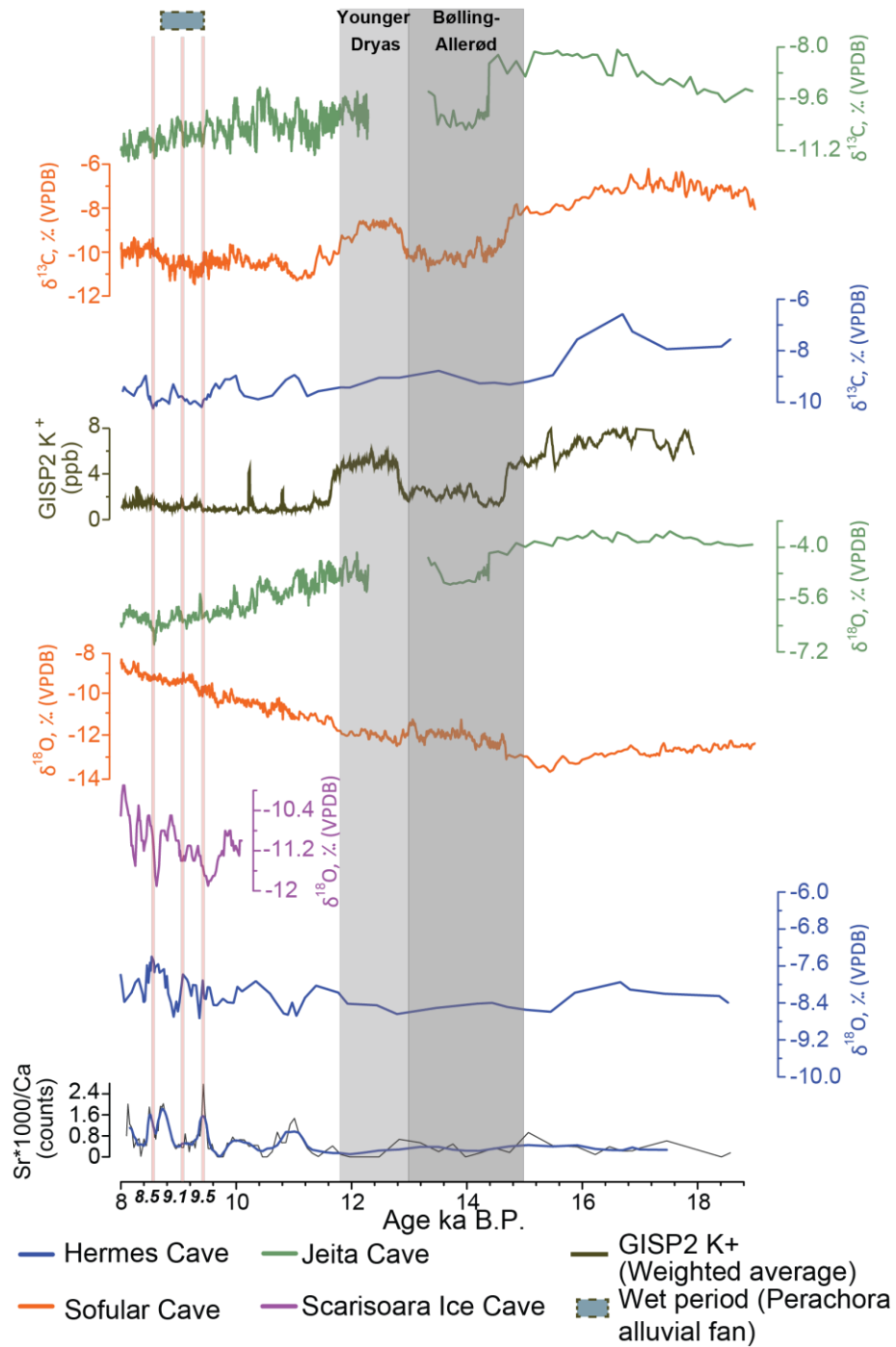
354 The overall trend of the  $\delta^{13}\text{C}$  records broadly follows that of similar records from regions with  
355 comparable climatic conditions, such as Sofular Cave in Northern Turkey (Fleitmann et al., 2009)  
356 and Jeita cave in Lebanon (Cheng et al. 2015). We have specifically chosen these records for  
357 comparison with our data as they both record winter climate variability, similar to our  
358 interpretation of  $\delta^{13}\text{C}$  and  $\delta^{18}\text{O}$  variability in the ZCG1 speleothem. While several other records  
359 exist in the area (referred to in Kern et al., 2019), these do not specifically mention what seasonally  
360 relevant climate variable is recorded so we restrict our discussion to proxies recording winter  
361 climate, only. In all these records, including the ZCG1 stalagmite,  $\delta^{13}\text{C}$  values peak at around 17  
362 ka and drop between 14.7 and 11.3 ka, resuming an increasing trend after the onset of the Holocene  
363 (11.7 ka). The low temporal resolution of our record does not allow us to go into detailed  
364 discussions, however, the overall similarity of these widely-spaced winter climate reconstructions  
365 suggest climatic processes acting on a regional scale. We suggest two climatic mechanisms at play  
366 during the B-A and the YD to be further tested by subsequent studies: 1) increased soil productivity  
367 linked to post-glacial forest expansion during the warm BA (Feurdean et al., 2014) and 2) increased  
368 precipitation delivered by strengthened westerlies linked to the southward displacement of the  
369 polar front during the YD (Lane et al., 2013). It is conceivable that while these moisture tracks  
370 reached both Greece and the Levant (as seen in the respective  $\delta^{13}\text{C}$  records), possibly also picking  
371 up moisture from the Eastern Mediterranean, they did not reach the Black Sea coast in Turkey,

372 explaining the drier conditions observed there, as indicated by the high  $\delta^{13}\text{C}$  values during the YD  
373 in the Sofular record (Fleitmann et al., 2009). Similarly high  $\delta^{13}\text{C}$  values were recorded in  
374 speleothem P10 in SW Romania (Constantin et al., 2010) at the end of the YD, suggesting drier  
375 conditions in the northern half of the region. Collectively, these observations point towards an  
376 important differentiation of climatic conditions between SE Europe and the Eastern Mediterranean  
377 region during the YD, with a band of cold and dry climatic conditions stretching from Central  
378 Europe across the northernmost Balkan Peninsula towards the northern coast of Turkey, and  
379 somewhat wetter conditions in southern Greece and the Levant, likely brought about by moisture  
380 delivered by southerly displaced westerlies. Additionally, we suggest that the complex interactions  
381 of climatic factors in SE Europe might have led to mutual cancellation of their effects, thus  
382 resulting in the poorly expressed climatic differentiation of the Bolling-Allerod and Younger  
383 Dryas. This hypothesis, presented as such, nevertheless needs to be tested with better temporally  
384 constrained records and greater spatial coverage.

385 The onset of the Holocene is marked by increased variability in the  $\delta^{13}\text{C}$  and  $\delta^{18}\text{O}$  records, similar  
386 to those seen in the Jeita and Sofular records. In the Hermes Cave, several potentially warm  
387 (recorded by  $\delta^{18}\text{O}$ ) and wet (recorded by  $\delta^{13}\text{C}$ ) periods punctuate the early Holocene, broadly  
388 centered at 9.5, 9.1 and 8.5 ka (Fig. 8); their timing correlates within error with a wet period  
389 recorded in alluvial fan formations on the Perachora peninsula (Peckover et al., 2019). These warm  
390 events (suggested, in the absence of monitoring data, by the corresponding peaks in the nearby  
391 Jeita and Sofular records and our knowledge of modern climatic variability in the region; Xoplaki  
392 et al., 2004) are coincident with cooling events in the Carpathian Mts, as recorded by the  $\delta^{18}\text{O}$  of  
393 cave ice (Perşoiu et al., 2017). These periods are further coincident with high  $\text{K}^+$  values measured  
394 in the GISP2 record (Mayewski et al., 1997; 2004), which indicate a stronger than usual Siberian



395 High. The Siberian High is a semi-permanent high-pressure cell located over Eurasia which affects  
396 European and Asian climate in winter (Cohen et al., 2001). Strengthening of the Siberian High  
397 results in atmospheric blocking that leads to cold air advection towards northern and central  
398 Europe and contrasting warming in southern Europe and the Levant with stronger than usual  
399 clockwise winds across SE Europe and enhanced cyclogenesis over the Central Mediterranean  
400 (Perşoiu et al., 2019). These conditions result in increased precipitation delivered to mainland  
401 Greece leading to wet conditions and thus potentially explaining the excursions towards negative  
402  $\delta^{13}\text{C}$  excursions in the ZCG1 speleothem. Alternatively, the increase in easterly anticyclonic winds  
403 could have increased the fraction of moisture sourced from the Aegean Sea, located east of our  
404 study site. Thus, the high  $\delta^{18}\text{O}$  values of ZCG1 speleothem calcite may have been the result of  
405 moisture from a warmer Aegean Sea (e.g. Marino et al., 2009). Younger speleothem  $\delta^{18}\text{O}$  records  
406 from the Peloponnese show a contrasting response to changes in precipitation, with the Kapsia  
407 record (Finné et al., 2014) indicating high  $\delta^{18}\text{O}$  values associated with dry conditions while the  
408 Alepotrypa record (Boyd, 2015) suggests high  $\delta^{18}\text{O}$  values are indicators of wet conditions (similar  
409 to ZCG1). Under these contrasting observations, it is likely that local infiltration conditions may  
410 have an outsized effect on the  $\delta^{18}\text{O}$ -climate relationship in all studied speleothems, which would  
411 further complicate their interpretation as past climate archives.



412

413

Figure 8. Comparison of ZCG1 post-LGM record with other regional records (see text for details).

414

415

416 The  $\mu$ -XRF results exhibit a relatively high variability in the Sr/Ca ratio which in combination  
417 with the stable isotope compositions can be used as indicators of seasonal environmental changes.  
418 The curve presents a prominent peak at 123-124 ka that correlates with lower values in the carbon  
419 record and a period of high oxygen values. The Sr and carbon isotope values suggest wet and warm  
420 summers linked to enhanced soil production in the area, while the oxygen isotope compositions  
421 did not significantly decrease, suggesting these changes were not particularly intense. Again, an  
422 increase in local, high  $\delta^{18}\text{O}$ , moisture sources at this time would also help to explain the persisting  
423 high oxygen isotope values. The period from 115 ka until the halt in stalagmite growth implies  
424 different climatic conditions relative to the older part of the stalagmite. In this younger period there  
425 are four prominent positive peaks at  $\sim$  114, 112.5, 107 and 105 ka that correlate with higher peaks  
426 in oxygen and lower peaks in the carbon record. This implies that the area was experiencing more  
427 precipitation during summer months, but the soil production was low since there is already a  
428 transition towards the last glacial period while mean annual temperature was decreasing. Finally,  
429 in the early Holocene section of the stalagmite, we observe four positive peaks in the Sr/Ca record  
430 centered at 9.5 and 8.5 ka that correlate with negative peaks in the carbon record and higher values  
431 in the oxygen record that suggest wet conditions during these periods that favoured soil formation  
432 (Fig. 8). Consequently, for the early Holocene, we suggest that easterly winds driven by a  
433 strengthened Siberian High resulted in moisture transfer from the warm surface waters of the  
434 Aegean Sea and subsequent increased precipitation in mainland Greece.

435

## 436 **5.2 Climatic impact on sediment delivery**

437 Recent results from the IODP Expedition 381 into the Gulf of Corinth (Shillington et al., 2019)  
438 show a profound variability in sediment accumulation rates in response to climate changes during

439 the Quaternary. In particular, relatively low sedimentation rates (0.3 to 0.7 mm/yr) are inferred for  
440 the last interglacial period (i.e. MIS 5) in contrast to the higher sedimentation rates that characterize  
441 the last glacial (~2.5 mm/yr) and the Holocene (2.5-3 mm/yr). McNeill et al. (2019) argue that the  
442 observed variability in sediment flux is due to changes in the type of vegetation cover, while the  
443 increased sedimentation rates during the Holocene are attributed to human deforestation in the area  
444 from 6 ka onward. Our results from the lower part of the ZCG1 stalagmite suggest high soil  
445 productivity, indicative of expanded vegetation cover in the area during the oldest parts of MIS 5.  
446 During this period, erosion was likely restricted by vegetation and consequently the fluvial  
447 network delivered less sediment into the Gulf of Corinth. After the demise of peak (penultimate)  
448 interglacial conditions, over the earliest part of the ZCG1 record, our data show a decline in soil  
449 productivity inferred by higher values of  $\delta^{13}\text{C}$ . This is consistent with results from other studies  
450 that imply the development of weak vegetation cover in the Balkan peninsula (e.g. Tzedakis et al.,  
451 2004, Weiberg et al., 2016). Although stalagmite growth halts during the last glacial period  
452 possibly due to a shortage of water infiltrating Hermes Cave as a result of the expansion of high-  
453 altitude small glaciers, at lower elevations there was probably an abundance of meltwater feeding  
454 the fluvial network. This combination of weak vegetation cover and higher water supply during  
455 the last glacial period could explain the high sedimentation rates observed in the Gulf of Corinth  
456 over this interval (McNeill et al., 2019). The younger part of the stalagmite (post growth hiatus)  
457 suggests a warmer and wetter period. However, during this period vegetation likely did not recover  
458 to pre-glacial levels as most of the soil cover was depleted due to erosion, enabling only low, open  
459 vegetation to grow (Weiberg et al., 2016). This, in combination with glacier melting, explains the  
460 higher erosion rates observed on the flanks of the Corinth Rift and increased sediment

461 accumulation into the gulf during the onset of Holocene (McNeill et al., 2019 and references  
462 therein).

463

## 464 **6. Conclusions**

465 This study contributes to the paleoclimatic reconstruction of the Eastern Mediterranean over the  
466 late Quaternary and early Holocene using a speleothem (ZCG1) from the Hermes Cave, which is  
467 located at the southern flanks of the Corinth Rift, central Greece. Our particular findings are the  
468 following:

- 469 ● The ZCG1 formation started at ~127 ka marking the establishment of climatic conditions  
470 that favored the speleothem deposition. This period coincides with the beginning of the  
471 maximum interglacial conditions which enabled the high growth rates observed in  
472 stalagmites in Europe.
- 473 ● We show that a low soil productivity period prevailed in the area from 122 to ca 117 ka  
474 implying overall dryer climatic conditions. This finding contrasts with the wetter climatic  
475 conditions observed in central Europe during the so-called 'last Interglacial optimum'. Dry  
476 conditions settle in the area later on until the beginning of MIS4, when the stalagmite stops  
477 growing at ca 105 ka.
- 478 ● The growth hiatus of ZCG1 from 105 to 23 ka occurred most probably due to low amounts  
479 of water infiltrating into the cave, likely caused by small glaciers which expanded in the  
480 area during the intervening glacial period.

- 481 ● Our record exhibits higher temporal resolution towards the onset of the Holocene due to a  
482 higher growth rate; we attribute this to increased soil productivity under warmer and wetter  
483 climatic conditions.
- 484 ● The ZCG1 record presents similarities with other records in the broader area of the Eastern  
485 Mediterranean. Specifically, during the Bølling-Allerød (BA) and Younger Dryas (YD)  
486 our data suggest enhanced soil productivity linked to post-glacial vegetation expansion  
487 during the warm BA and increased precipitation delivered by strengthened westerlies due  
488 to southward displacement of the polar front during the YD. Our data show a clear  
489 differentiation of climatic conditions during the YD between SE Europe and the Eastern  
490 Mediterranean region. During the Early Holocene three warm and wet periods, at ~ 8.5,  
491 9.4 and 10.1 ka, mark the overall warming trend. These warm periods may have resulted  
492 from cold air outbursts associated with a strengthened Siberian High which were  
493 nonetheless not strong enough to generate cold conditions over mainland Greece. These  
494 clockwise moving winds may also have absorbed moisture from the Aegean Sea and  
495 caused higher than usual precipitation over the area.
- 496 ● Our record shows high variability in soil productivity and precipitation. The combined  
497 effect of these two parameters is the controlling factor on catchment averaged erosion rates  
498 with implications for sediment delivery into the Gulf of Corinth.

## 499 **7. Acknowledgements**

500 The authors would like to express their gratitude to cavers Yorgos Sotiriadis, Charikleia  
501 Gkarlaoui, Christina Gkarlaoui and Nikolaos Kortimanitsis for their invaluable help during our  
502 visit to Hermes Cave. The Ephorate of Palaeoanthropology and Speleology of the Hellenic  
503 Ministry of Culture is thanked for granting permission to work inside the cave

504 (ΥΠΠΟΑ/ΓΔΑΠΚ/ΕΠΣ/ΤΑΕΜΓΠ/87570/59775/1006/40). We also thank Anna Kieu-Diem Tran  
505 for assistance with the stable isotope analyses (University of Bergen). RG acknowledges the award  
506 of the VISTA Professorship from the Norwegian Academy of Science and Letters. SP and some  
507 of the analytical work was also funded by the VISTA Professorship award. AP was supported by  
508 the Romanian Ministry of Education and Research, CNCS - UEFISCDI, project number PN-III-  
509 P4-ID-PCE-2020-2723, within PNCDI III. SEL and AP acknowledge the KARTSHIVES 2 project  
510 (EEA grants). This work was inspired by the conversations we had with late Prof. Patience Cowie,  
511 a great colleague and mentor.

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