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3 **Carbonate polymorphism controlled by microbial iron redox**
4 **dynamics at a natural CO₂ leakage site (Crystal Geysir, Utah)**

5 JULIE COSMIDIS^{1,2}, SHANE O'REILLY^{3,4}, ERIC T. ELLISON¹, KATHERINE L. CRISPIN⁵,
6 DAVID DIERCKS⁶ and ALEXIS S. TEMPLETON¹

7 ¹*Department of Geological Sciences, University of Colorado, Boulder, CO 80309, USA*

8 ²*Department of Earth Sciences, University of Oxford, Oxford OX1 3AN, UK*

9 ³*School of Earth Sciences, University College Dublin, Belfield, Dublin 4, Ireland*

10 ⁴*Institute of Technology Sligo, Sligo F91 YW50, Ireland*

11 ⁵*Materials Research Institute, Pennsylvania State University, University Park, PA 16802, USA*

12 ⁶*Department of Metallurgical and Materials Engineering, Colorado School of Mines, Golden,*
13 *CO 80401, USA*

14 **Corresponding author:** Julie Cosmidis. Email: julie.cosmidis@earth.ox.ac.uk

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16 **ABSTRACT**

17 Crystal Geysir (Utah, USA) is a CO₂-rich low-temperature geysir that is studied as a natural
18 analog for CO₂ leakage from carbon capture and storage (CCS) sites. In order to better constrain
19 the biogeochemical processes influencing CaCO₃ precipitation at geological CO₂ escape sites,
20 we characterized fast-forming iron-rich calcium carbonate pisoids and travertines precipitating
21 from the fluids expelled by the geysir. The pisoids, located within a few meters from the vent,
22 are composed of concentric layers of aragonite and calcite. Calcite layers contain abundant

23 ferrihydrite shrubs in which iron is encasing bacterial forms. The aragonite layers contain less
24 abundant and finely dispersed iron, present either as iron-oxide microspherules or iron adsorbed
25 to organic matter dispersed within the carbonate matrix. We propose that carbonate
26 polymorphism in the pisoids is mostly controlled by local fluctuations of the iron redox state of
27 the fluids from which they form, caused by episodic blooms of iron-oxidizing bacteria. Indeed,
28 the waters expelled by Crystal Geyser contain $>200 \mu\text{M}$ dissolved iron (Fe^{2+}), a known inhibitor
29 of calcite growth. The calcite layers of the pisoids may record episodes of intense microbial iron
30 oxidation, consistent with observations of iron-oxide rich biofilms thriving in the rimstone pools
31 around the geyser and previous metagenomic analyses showing abundant neutrophilic,
32 microaerophilic iron-oxidizing bacteria in vent water. In turn, aragonite layers of the pisoids
33 likely precipitate from Fe^{2+} -rich waters, registering periods of less intense iron oxidation.
34 Separately, CaCO_3 polymorphism in the travertines, where calcite and aragonite precipitate
35 concurrently, is not controlled by iron dynamics, but may be locally influenced by the presence
36 of microbial biofilms. This study documents for the first time an influence of microbial iron
37 oxidation on CaCO_3 polymorphism in the environment, and informs our understanding of
38 carbonate formation at CO_2 leakage sites and in CCS contexts.

39 **Keywords**

40 Travertines, pisoids, carbonates, CCS, iron oxidation, frutexitis

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42 INTRODUCTION

43 Carbon Capture and Storage (CCS) are climate mitigation technologies whereby carbon dioxide
44 (CO₂) is captured at large emission sources (e.g., fossil fuel power plants or industrial sites), and
45 injected into deep sedimentary reservoirs for long-term storage. These negative carbon emission
46 strategies are essential tools to meet global climate goals (van Vuuren *et al.*, 2017) but several
47 challenges and uncertainties are still limiting their application (Kelemen *et al.*, 2019). Among
48 them, the potential for CO₂ stored in deep aquifers to leak to the surface along natural fractures
49 or through injection wells and boreholes is of particular concern (Alcalde *et al.*, 2018).

50 Naturally occurring as well as man-made CO₂ leakage sites have been studied to understand the
51 fate and impact of geological CO₂ in surface environments (e.g., Lewicki *et al.*, 2006; Roberts &
52 Stalker, 2017). Crystal Geysir (Utah, USA) is an abandoned petroleum exploration well, drilled
53 in 1935, from which CO₂ has been leaking to the surface for decades at a pace of ~12 kt/yr
54 (Gouveia *et al.*, 2005). The well bore reaches a deep CO₂-charged aquifer in the Navajo
55 Sandstone (Colorado Plateau), from which CO₂ has also been leaking along natural faults for
56 more than 400,000 years (Shipton *et al.*, 2004; Burnside *et al.*, 2013). CO₂ at Crystal Geysir is
57 escaping to the surface both as free gas and dissolved in brines expelled during eruptions
58 (Assayag *et al.*, 2009; Kampman *et al.*, 2014). It is estimated that 1 to 10% of the dissolved CO₂
59 is precipitated as calcium-carbonates (travertines), forming a mound around the geyser (Shipton
60 *et al.*, 2004; Burnside *et al.*, 2013).

61 Here, we performed a detailed characterization of travertines and pisoids forming at Crystal
62 Geysir, in order to identify chemical and biological processes influencing CaCO₃ precipitation,
63 with a particular focus on CaCO₃ polymorphism. Various efforts have been made to constrain

64 the physicochemical factors influencing the polymorphism of calcium carbonates precipitating
65 from CO₂-rich springs. Parameters such as temperature, pH, CO₂ content and degassing rate,
66 fluid ionic strength, and presence of sulfate, metals cations, and organics, determine whether
67 calcite, aragonite, or (more rarely) vaterite, may form in CO₂-rich spring systems (Chang *et al.*,
68 2017; Jones, 2017). In addition, biological factors such as the presence of microbial cells and
69 extracellular polymeric substances (EPS) may also influence CaCO₃ polymorphism in travertines
70 (Guo & Riding, 1992; Okumura *et al.*, 2013a; b; Peng & Jones, 2013). Parameters controlling
71 CaCO₃ polymorphism at CO₂ leakage sites need to be constrained to improve our understanding
72 of the fate and long-term stability of carbon trapped as carbonate minerals. Moreover,
73 understanding the processes influencing the precipitation, polymorphism, and physicochemical
74 properties of CaCO₃ formed in CO₂ storage settings participates in efforts aimed at valorizing
75 CaCO₃ as value-added materials for different industries (construction materials, food and
76 pharmaceuticals, paper, etc.), potentially offsetting the cost of CCS (Chang *et al.*, 2017).
77 This work combines mineralogical and chemical characterizations of the travertines and pisoids
78 at Crystal Geysers, geochemical analyses of the fluids from which they form, and investigations
79 of microbe-mineral interactions through microscopy and lipid biomarker analyses, in order to
80 better constrain physicochemical and biological processes impacting CaCO₃ mineralization and
81 polymorphism at CO₂ leakage sites. An influence of microbial iron oxidation on CaCO₃
82 polymorphism in a CO₂-rich environment is shown here for the first time, a process that may be
83 important in CCS contexts.

84 **BACKGROUND**

85 **Setting: Crystal Geysers**

86 Crystal Geyser is a cool CO₂-driven geyser located 14.5 kilometers southeast of the town Green
87 River in Utah (USA). Its hydrology has been described elsewhere (Shipton *et al.*, 2004; Gouveia
88 *et al.*, 2005; Gouveia & Friedmann, 2006; Assayag *et al.*, 2009; Heath *et al.*, 2009; Wilkinson *et*
89 *al.*, 2009; Han *et al.*, 2013; Kampman *et al.*, 2014), and a brief summary will be given here.
90 Crystal Geyser waters erupt from an abandoned exploration well drilled in 1935 (a pipe, standing
91 ~2 m above the ground, was added for safety in the 1990s). The well was drilled into 21.5 m of
92 pre-existing travertines, showing the existence of natural springs at this location prior to drilling
93 (Baer & Rigby, 1978). Although the drill-hole is ~800 m deep, erupted waters discharge
94 predominantly from the Jurassic Navajo Sandstone at ~200-350 m depths. Crystal Geyser erupts
95 mildly reducing (Eh ~ -5 mV), slightly acidic (pH ~6.5), cold waters (~18 °C), containing
96 mixtures of groundwater of meteoritic origin (80-90% of erupted waters) and brines emanating
97 from deeper (~1.5 km depth) carboniferous evaporite formations (~10-20 % of erupted waters)
98 (Wilkinson *et al.*, 2009; Kampman *et al.*, 2014). Crystal Geyser erupts CO₂ both as a free gas
99 (representing ~96% of the erupted gases; Kampman *et al.*, 2014) and dissolved in the waters
100 (which start degassing at ~120 m depths as the waters migrate vertically in the drill-hole;
101 Assayag *et al.*, 2009). The CO₂ emanates from carbonate dissolution by acidic groundwater in
102 the Navajo sandstone (Heath *et al.*, 2009) and deep supercritical CO₂ reservoirs migrating
103 upwards through a normal fault system (Shipton *et al.*, 2004; Gilfillan *et al.*, 2008).

104 The intensity of Crystal Geyser eruptions has been declining with time since the drilling of the
105 oil well at its origin. In 1973, the discharge from a single eruption was measured to be ~120 m³,
106 while in 2001 the discharge was only ~25 m³ (Waltham, 2001). The frequency and duration of
107 eruption is also evolving with time. Earlier publications have reported a bimodal eruption cycle,

108 with large eruptions lasting between 1 and 1.5h (Type B) and 5-7h (Type D), occurring every 7-
109 10 h or 20-30 h, respectively (Gouveia *et al.*, 2005; Gouveia & Friedmann, 2006; Han *et al.*,
110 2013). Between these large eruptions, smaller magnitude “bubbling events” (Type A and C
111 eruptions; Han *et al.*, 2013) would occur approximatively every 15 minutes. However, longer
112 duration (~24 h) and lower frequency (every ~70 h) eruptions have been described by Kampman
113 *et al.* (2014) who observed the geyser in 2012. For the present study, Crystal Geyser was visited
114 in 2014, and a short (1-2 hours) eruption event was observed followed by continuous “bubbling”
115 lasting approximatively 24 hours. After this, the geyser and surrounding pools were totally dry
116 for ~24 hours. During the following ~24 hours, the pool directly around the geyser vent was
117 observed to be slowly filling with water flowing from holes at the base of the pipe, after which
118 another eruption/bubbling event started. In 2015, Crystal Geyser was visited for a few hours on
119 two consecutive days. On the first day, the geyser was observed to be “bubbling”, while on the
120 following day no water was flowing from the geyser and its surroundings were dry.

121 During eruptions, Crystal Geyser waters flow towards Green River, located ~80 m downstream.
122 Along their flow path, the waters deposit terraced travertines which form a gently sloping mound
123 with a lateral extent of ~85 m (Barth & Chafetz, 2015). The travertines are bright red in color
124 directly around the vent of the geyser and become orange and yellow more distally. Brighter
125 colors are observed centrally respective to the water flow path, and paler colors laterally (Fig.
126 1A,G). Green mats, likely corresponding to algae and cyanobacteria (Takashima *et al.*, 2011b;
127 Barth & Chafetz, 2015), cover the central portion of the travertines in the warm season (Fig.
128 S1A,B).

129 During eruptions, pools around the geyser fill with water, which can be flowing or stagnant
130 depending on the timing and intensity of the eruptions. When waters are stagnant, rusty materials
131 (most likely iron-(oxyhydr)oxides) form at the surface of the pool directly surrounding the
132 geyser (Fig. 1D), while oily-looking films are observed at the surface of shallow rimstone pools
133 located slightly further downstream. Between eruptions, the area surrounding geyser dries up
134 completely (Fig. 1C).

135 The modern travertine (deposited since the drilling of the oil well bore in 1935) is estimated to
136 be ~1 m (Burnside *et al.*, 2013) to several meters (Barth & Chafetz, 2015) thick, corresponding
137 to sedimentation rates of more than one to several centimeters per year. Fast carbonate
138 precipitation at Crystal Geyser is also evidenced by the presence of recently encrusted plants and
139 objects across the field site (Figs. S1J,K).

140 **Crystal Geyser travertines and pisoids**

141 Crystal Geyser terraced travertines form a gently sloping mound around the geyser. Near Green
142 River, the slope abruptly becomes steeper and the travertines form staircase-like steps (Fig. 1G).
143 The travertines are often laminated (Fig. 1H), which was interpreted as resulting from daily
144 banding combined with the eruption pattern of the geyser (with dark, micritic bands, influenced
145 by photosynthetic organisms, forming when both sunlight and water are available, and lighter
146 bands with larger crystals developing at night when water is present) (Takashima *et al.*, 2011b).
147 Some travertines furthermore display microstromatolitic horizons, interpreted as resulting from
148 the encrustation of microbial mats (Barth & Chafetz, 2015). Travertines located downslope or
149 laterally respective to the water flow path are often more porous, and may contain fenestrae-like
150 holes and include calcified bubbles (Fig. 1I). Crystal Geyser travertines contain iron-

151 (oxyhydr)oxides which give them their red to orange colors (depending on iron abundances). The
152 iron-(oxyhydr)oxides sometimes form filaments and hollow tube-like structures, interpreted as
153 encrusted sheaths of iron-oxidizing bacteria (Barth & Chafetz, 2015). At the surface of the
154 travertines, calcium carbonates frequently form fan-shaped bundles of feather-like crystals,
155 eventually producing botryoidal structures, previously interpreted as being microbial in origin
156 (Parenteau & Cady, 2010; Barth & Chafetz, 2015). Large (> 1 cm) botryoidal, cauliflower-
157 shaped carbonate structures we also observed at the surface of the travertines in pools located
158 distally from the vent of the geyser (Figs. S1C-E). Other botryoidal structures shaped like toroids
159 were also found (Fig. S1F). They could possibly form around calcified gas bubbles which are
160 frequently observed in the rimstone pools (Fig. S1G-I). These botryoidal structures will not be
161 further described here.

162 Reddish pisoids (*i.e.*, rounded coated grains measuring less than a centimeter in diameter)
163 accumulate in rimstone pools mostly located proximal to the geyser (Fig. 1F). Small pisoids are
164 typically simple and well-rounded, while larger ones can be complex (composed of multiple
165 aggregated smaller coated grains) and irregularly shaped. The nuclei of the pisoids often contain
166 quartz grains cemented by calcium carbonates (Barth & Chafetz, 2015). Their cortices are
167 formed by alternating layers of calcium carbonate and iron. Iron-(oxyhydr)oxides frequently
168 form outwardly branching growth patterns, or “shrubs”, resembling frutexites, which are often
169 interpreted as fossil microbial structures (Takashima *et al.*, 2008; Jakubowicz *et al.*, 2014; Guido
170 *et al.*, 2016; Reitner *et al.*, 2017; Grădinaru *et al.*, 2020). Due to their proximity to Crystal
171 Geyser’s vent, Barth and Chafetz (2015) proposed that the pisoids may be formed within the

172 geyser's plumbing system and ejected during eruptions. Analyses were focused on these pisoids
173 due to their potential to probe CaCO_3 formation in the subsurface.

174 **MATERIALS AND METHODS**

175 **Sample collection**

176 Crystal Geyser was visited in April 2014 and February 2015. Samples of the pisoids (found
177 within ~10 m of the geyser) and two different types of travertine were collected: a visibly
178 laminated travertine (Fig. 1H), referred to hereafter as TL, and a non-laminated, porous
179 travertine containing calcified gas bubbles (Fig. 1I), referred to hereafter as TB. Both travertines
180 were collected distally from the geyser, in the area where travertines form steep slopes towards
181 the river (Fig. 1G). Water samples were also collected from Crystal Geyser's vent for
182 geochemical analyses during eruptions or "bubbling" events. Particles from erupted waters were
183 collected by filtration on 0.2 μm polycarbonate filters. The filters were rinsed immediately with
184 deionized water and air-dried, and preserved for later mineralogical analyses.

185 **Water chemistry**

186 Major cations (Mn^{2+} , Fe^{2+} , Mg^{2+} , Ca^{2+} , Al^{3+} , Sr^{2+} , Na^+ , K^+) and SiO_2 were analyzed on water
187 samples collected from Crystal Geyser's vent in 2014 and 2015. The water samples were filtered
188 and acidified in the field, and cations were analyzed using Inductively Coupled Plasma-Mass
189 Spectrometry (ICP-MS) on a ThermoScientific X Series 2. Major anions (F^- , Cl^- , Br^- , NO_2^- ,
190 NO_3^- , PO_4^{3-} , SO_4^{2-}) were analyzed on field-filtered samples collected in 2014 only. Anion
191 analyses were performed using Ion Chromatography (IC) on a Dionex IC25 with an IonPac
192 column and a 9 mM sodium carbonate eluent.

193 Profiles of dissolved Fe(II) and total Fe (measured by colorimetry) were obtained across the field
194 site using a portable Hach DR890 instrument in 2014. pH and temperature were measured with a
195 separate Hach multimeter with a pH probe.

196 **X-ray diffraction**

197 Samples of the pisoids and travertines TB and TL were finely ground with an agate mortar and
198 pestle, and the powders were analyzed using a Bruker D2 Phaser operated at 30 kV and 10 mA.
199 X-ray diffraction (XRD) spectra were obtained in the 2θ range $10\text{-}65^\circ$ using Cu $K\alpha$ radiation ($\lambda =$
200 1.5418\AA) and a Lynxeye 1D detector with a step size of 0.02° and collection time of 1 s per step.

201 **X-ray absorption spectroscopy**

202 X-ray absorption spectroscopy at the Fe K-edge was performed to determine iron speciation in
203 the travertines and pisoids. Samples were ground to a fine powder using an agate mortar and
204 pestle, and loaded into sample folders sealed with kapton tape. X-ray absorption spectroscopy
205 measurements were performed on beamline 4-1 of the Stanford Synchrotron Radiation
206 Lightsource (SSRL). X-ray absorption spectra were collected in fluorescence mode using a Si
207 $(220) \phi = 0$ monochromator and a Lytle detector Energy was calibrated by setting the first
208 inflection point of the Fe K-edge XANES spectrum of a reference Fe^0 foil to 7112 eV. Two to
209 four spectra were collected and averaged for each sample. X-ray absorption near edge structure
210 (XANES) spectra were background subtracted and normalized to unit step edge using the
211 SIXPack software package (Webb, 2005). Extended X-ray absorption fine structure (EXAFS)
212 spectra were extracted with SIXPack using a threshold energy of 7125 eV. Previously published
213 spectra of 2-Line ferrihydrite, hematite, goethite (Maillot *et al.*, 2011) and lepidocrocite (Pantke
214 *et al.*, 2012) were used for comparison with the samples.

215 **Raman spectromicroscopy**

216 Raman spectra and hyperspectral maps were acquired on thin sections of the pisoids and
217 travertines, as well as on particles from Crystal Geysir's vent waters collected on polycarbonate
218 filters. The analyses were performed using a Horiba LabRAM HR Evolution Raman
219 spectrometer using a 532 nm frequency-doubled Nd:YAG laser and a Si-based CCD detector
220 (1024 x 256 pixels). The laser beam was focused through a 10x or 50x objective lens, yielding a
221 spatial resolution of ~5 μm or ~2 μm respectively. Spectra were collected from 80 to 1800 cm^{-1} .
222 A spectral resolution full width at half maximum (FWHM) of 4.5–8.4 cm^{-1} was obtained using a
223 600 lines/mm grating and adjustable confocal pinhole (100–200 μm). Prior to analysis,
224 calibration of the spectrometer was performed using the 520 cm^{-1} Raman peak of Si. Spectral
225 data were corrected for instrumental artifacts and baseline-subtracted using a polynomial fitting
226 algorithm in LabSpec 6 (Horiba Scientific). Raman maps were used to visualize the distribution
227 of mineral species using least-squares fitting. Spectra were averaged over a small portion of the
228 map containing relatively pure Raman spectra in order to define end-members. These end-
229 members were then used to fit the full map dataset by classical least squared constrained by non-
230 negativity of the fit coefficients using LabSpec 6 (Horiba Scientific).

231 **Electron microprobe**

232 Electron probe microanalyses (EPMA) were performed on a polished thin section of a pisoid.
233 The EPMA maps were collected on a CAMECA SX-Five microprobe using a LaB6 source at
234 15kV, 15nA with the beam defocused to 5 μm diameter. The stitched images are a mosaic
235 composed of 3x3 individual maps, collected at 256x256 pixels, with a step size of 10 μm and
236 dwell time of 25 ms per pixel. All elements were collected using the $\text{K}\alpha$ x-ray line except $\text{L}\alpha$ for

237 Sr. A LTAP crystal was used to collect Mg, and Al. A PET crystal was used to collect Ca, S, Si,
238 K, and P. A LPET crystal was used to collect Sr. A LLIF crystal was used to collect Fe. Energy-
239 dispersive X-ray spectroscopy was used to collect C and O.

240 **Scanning Electron Microscopy**

241 Scanning Electron Microscopy (SEM) was performed on different types of pisoid samples: (i)
242 polished pisoid cuts, (ii) pisoid cuts etched with HCl, and (iii) thin sections. SEM analyses were
243 also performed on particles from Crystal Geyser vent waters collected on polycarbonate filters.
244 All samples were coated with gold prior to SEM. The analyses were conducted on a JSM-7401F
245 field emission SEM. Images were acquired in the secondary electron mode with the microscope
246 operating at 5 kV and a working distance of 6 mm, and in the backscattered electron mode at 10
247 kV and a working distance of 10 mm. Elemental analyses and maps were obtained using Energy-
248 dispersive X-ray spectroscopy (EDX), performed at 20 kV with a working distance of 8 mm.

249 **Focused Ion Beam**

250 Specimens from a pisoid were prepared for Transmission Electron Microscopy (TEM) in an FEI
251 Helios NanoLab 600i Ga focused ion beam / scanning electron microscope (FIB/SEM) using an
252 in-situ lift-out method. Electron beam assisted deposition of Pt from a
253 Trimethyl(methylcyclopentadienyl)platinum(IV) source was used as an initial protective layer
254 over the regions of interest. This was followed by ion beam assisted deposition from the same
255 source to a total of ~ 2 μm thickness. An ion beam accelerating voltage of 30 kV was used to
256 mill around the regions of interest and for thinning of the specimens after they had been
257 extracted to a TEM grid by a nanomanipulator. Final thinning of the specimens was performed

258 using a 2kV ion beam accelerating voltage. Locations of the two FIB sections performed in the
259 pisoid are depicted in Figures S2 and S3.

260 **Transmission Electron Microscopy**

261 TEM analysis of the FIB sections was performed on an FEI Talos F200X instrument using a 200
262 keV accelerating voltage. Analyses included bright field and dark field imaging, high angle
263 annular dark field scanning TEM (HAADF-STEM) imaging, selected area diffraction, and EDX
264 compositional mapping. The EDX data were processed using Bruker Esprit 1.9 software.

265 **Scanning Transmission X-ray Microscopy**

266 Scanning Transmission X-ray Microscopy (STXM) was performed on FIB sections 1 and 2 on
267 beamline 10ID-1 (SM) of the Canadian Light Source (Saskatoon, Canada) (Kaznatcheev *et al.*,
268 2007). Energy calibration was achieved using the 3p Rydberg peak of gaseous CO₂ at 294.96 eV.
269 Images, maps and image stacks were obtained at the C K-edge, the Ca L_{2,3}-edge, and Fe L_{2,3}-
270 edges, using a 25 nm zone plate. STXM data was processed using the aXis2000 software
271 (Hitchcock, 2012). Organic carbon maps were obtained by subtracting an image at 280 eV (pre-
272 edge) and converted into an optical density (OD) image, from an OD-converted image at 288.5
273 eV (energy of the 1s→π* electronic transition in carbonyl and carboxylic groups). Carbonate
274 maps were obtained by subtracting an OD-converted image at 280 eV (pre-edge) from an OD-
275 converted image at 290.2 eV (energy of the 1s→π* electronic transition in carbonate groups).
276 Calcium maps were obtained by subtracting an OD-converted image at 342 eV (pre-edge) from
277 an OD-converted image at 349.2 eV (energy of the Ca L₃-edge main peak). Iron maps were
278 obtained by subtracting an OD-converted image at 700 eV (pre-edge) from an OD-converted
279 image at 710.2 eV (energy of main absorption peak in Fe-(oxyhydr)oxides). Pixels with negative

280 values in resulting maps were removed using the Clip Signal tool of aXis2000. XANES spectra
281 where extracted from aligned image stacks as described in Cosmidis & Benzerara (2014). Linear
282 background corrections were applied to the spectra in the 270-282 eV energy range at the C K-
283 edge, in the in the 330-345 eV energy range at the Ca L_{2,3}-edges, and in the 690-704 eV energy
284 range at the Fe L_{2,3}-edges. For some image stacks, representative XANES spectra for major
285 components of the samples were extracted, and the relative contribution of these representative
286 spectra at each pixel was mapped using the Stack Fit tool of aXis2000.

287 **Lipid biomarkers**

288 Lipid biomarkers were extracted from pisoid and travertine samples and analyzed using gas
289 chromatography-mass spectrometry (GC-MS). 20.0 g of Crystal Geyser pisoids and travertines
290 (TL and TB) were powdered using a shatterbox and accurately weighed into 60 mL glass
291 centrifuge tubes. Each sample was spiked with 1 µg of nonadecan-1-ol internal standard.
292 Samples were extracted with organic solvent as follows: 2:1 (v/v) methanol/dichloromethane
293 (×3), followed by 9:1 (v/v) dichloromethane/methanol (×3). For each extraction, the tubes were
294 sonicated for 10 minutes in an ultrasonic bath at room temperature. Extracts were separated from
295 solid residues by centrifugation, and supernatants from each step were combined to give a total
296 free lipid fraction, comprising surface and intercrystalline lipids. This fraction is referred to as
297 the free lipid fraction. 10.0 g of the extracted residues were subsequently diluted in
298 dichloromethane-cleaned water and carefully dissolved in HCl. When there was no evidence of
299 remaining carbonate, lipids were extracted from the aqueous solutions using liquid-liquid
300 extraction (dichloromethane, ×3). Lipids adsorbed to remaining residue were also extracted using

301 sonication/centrifugation as described above, and combined to give a carbonate-bound
302 (intracrystalline) lipid fraction.

303 All extracts were concentrated to minimal volume under a gentle stream of high purity N₂. A
304 portion of each extract was then subjected to acid methanolysis (0.5 N methanolic HCl, 60°C
305 ~10 h), followed by silylation (BSTFA (+1% trimethylchlorosilane) in pyridine, 70°C, 2 h).
306 Derivatized samples were analyzed by gas chromatography-mass spectrometry (Agilent 5890
307 GC hyphenated to an Agilent 5975C Mass Selective Detector). The GC was equipped with a
308 Gertsel programmable temperature vaporizer (70°C ramped to 360°C at a rate of 720°C min⁻¹)
309 and a J&W 60 m capillary column (0.25 mm inner diameter, 250 μm film thickness). The GC
310 temperature program was: 70°C for 2 min, ramp at 10°C min⁻¹ to 130°C, followed by a ramp to
311 300°C at 4°C min⁻¹ and a final hold time of 20 min. The mass spectrometer was operated in
312 electron impact ionization mode (70 eV), with a mass scan range from m/z 50 to 600. All
313 solvents used were high-purity (OmniSolv) and all aqueous solutions were cleaned with
314 dichloromethane prior to use, and procedural blanks were run to monitor background
315 contamination. The peak areas of analytes were compared with peaks of the internal standard and
316 can be considered semi-quantitative.

317 **RESULTS**

318 **The geochemical environment at Crystal Geyser**

319 Table 1 shows concentrations of major elements measured in waters collected from the vent of
320 the geyser on two different days in April 2014, as well as in February 2015. Crystal Geyser
321 waters are rich in chloride, sodium, sulfate, calcium, potassium, and magnesium (in order of
322 decreasing abundance), consistent with a contribution from brines originating from deep

323 evaporite formations (Wilkinson *et al.*, 2009; Kampman *et al.*, 2014). Dissolved iron values
324 measured by ICP-MS were 6.14-10.61 mg.L⁻¹, falling between the range of values reported by
325 other authors (12.8-15.7 mg.L⁻¹ in Kampman *et al.* (2014) and ~3.4 mg.L⁻¹ in Heath *et al.*
326 (2009)). Most of this dissolved iron is present under a reduced form, since colorimetric
327 measurements performed on vent water during an eruption in 2014 showed 10.05 mg.L⁻¹ Fe²⁺
328 and only 10.85 mg.L⁻¹ total iron. Fe²⁺ in the geyser waters likely results from the reductive
329 dissolution of hematite minerals present in the Navajo sandstone reservoir by CO₂-charged
330 brines (Kampman *et al.*, 2014).

331 According to previous studies, Crystal Geyser waters are supersaturated with respect to both
332 aragonite and calcite (Heath *et al.*, 2009). However, SEM and Raman analyses of vent particles
333 collected on filters showed the presence of numerous aragonite blades (Fig. S4) and no calcite.
334 Fine-grained iron minerals were also present, but it is not clear whether they originated from the
335 geyser or if they precipitated on the filters from oxidation of Fe²⁺-rich vent water during sample
336 preparation.

337 Geochemical profiles of dissolved Fe(II), total Fe, pH, and temperature were obtained on-site in
338 April 2014, using measurements performed at different locations along the geyser's flow path,
339 from the pool around the vent to Green River (Fig. 2). The waters expelled by Crystal Geyser
340 have temperatures of ~17-18 °C and pH values of ~6.5. As they flow towards Green River, they
341 get progressively warmer (~24 °C at the river) and slightly more basic (pH ~7.8 at the river).
342 This pH increase is consistent with progressive CO₂ degassing. Dissolved Fe(II) (Fe²⁺) and total
343 Fe (Fe_{tot}) concentrations decrease with distance from the vent, showing the progressive oxidation
344 and precipitation of iron from the water.

345 **Mineralogical description of the Crystal Geysir travertines and pisoids**

346 *Bulk mineralogical composition*

347 Both aragonite and calcite were identified by XRD in the travertine samples TL and TB as well
348 as in the pisoids (Fig. 3A). Fe K-edge EXAFS was used to determine the mineralogical
349 composition of the iron-bearing phases giving the travertines and pisoids their orange color.
350 EXAFS spectra match that of a ferrihydrite reference (Fig. 3A) (with possibly a slightly more
351 disordered structure in the travertine samples as compared with the pisoids). The samples'
352 spectra were fitted with reference spectra of different iron-(oxyhydr)oxides using a linear
353 combination-least squares fitting approach. Two-line ferrihydrite provided the best fit for all
354 samples, and no additional mineral improved the quality of the fits.

355 *Mineralogical mapping*

356 Raman spectromicroscopy was used to map the distribution of different mineral phases in the
357 samples. In travertine TL, distinct microlaminae dominated by either aragonite or calcite can be
358 found, although both phases are present in all microlaminae (Fig. 4). Quartz grains were also
359 found in travertine TL; it is possible that these quartz grains were undetected by XRD due to
360 their low abundance relative to calcium carbonates. Travertine TB is composed of randomly
361 distributed aragonite and calcite grains, ranging from a few microns to ~100 μm in size (Fig. 5),
362 with distinct areas of the travertine displaying different relative proportions of aragonite and
363 calcite grains (compare Figs. 5D and 5G). The cortices of the pisoids are composed of concentric
364 alternating layers of calcite and aragonite, measuring a few hundred micrometers in thickness
365 (Fig. 6,S5,S6). Interestingly, the nuclei of all the pisoids analyzed ($n = 5$) contain aragonite as the
366 main CaCO_3 phase, along with quartz grains. Some quartz grains were also found in the cortices

367 of the pisoids. Ferrihydrite was not mapped with Raman, due to its typically weak signal, but Fe-
368 rich regions were visible as red or brown areas on micrographs acquired along the maps. In all
369 pisoids, ferrihydrite is mostly present in well-delimited layers spatially co-located with calcite
370 (whereas aragonite layers are ferrihydrite-free). This co-occurrence between iron minerals and
371 calcite was not observed in the travertine samples, where iron phases are more dispersed and
372 finely intermixed with the carbonates.

373 *Elemental mapping of a pisoid*

374 Using EPMA, chemical maps showing the distribution of different elements in one Crystal
375 Geyser pisoid were obtained (Fig. 6C-E, Fig. S7). Fe is mostly present in the cortex, as discrete
376 layers of ferrihydrite co-located with calcite. Mg is particularly abundant in the nucleus of the
377 pisoid, while in the cortex it appears only in calcite layers (Fig. 6D). Sr is distributed more
378 homogeneously through the nucleus and cortex, with a slight enrichment in aragonite phases
379 (Fig. S7). S is absent from the nucleus, and in the cortex it is preferentially associated with
380 calcite (Fig. 6E). Si is found in hot spots, particularly abundant in the nucleus, and co-located
381 with high abundances of oxygen (Fig. S7), consistent with the presence of quartz grains. Al, and
382 K, sometimes with Mg, co-occur with some Si hot spots, likely corresponding to clay minerals
383 (undetected with other methods) in both the nucleus and the cortex.

384 **Nano-scale mineralogical characterization of a pisoid**

385 *Mineralogical and chemical characterization*

386 Two FIB sections from a pisoid were analyzed using TEM and STXM (Figs. 7,8). The FIB
387 sections were performed in porous areas of the cortex (Figs. S2,S3), in an attempt to sample and
388 characterize trapped organic materials (see following section). These areas correspond to

389 aragonitic layers of the cortex, as confirmed by electron diffraction (data not shown) and
390 STXM/XANES analyses at the Ca L_{2,3}-edge (Fig. S8). Distinct aragonite crystals are visible on
391 STXM carbonate maps (Fig. 7H, 8F), with sizes in the micrite range or smaller (< 2 μm).
392 Although SEM and Raman analyses have shown that aragonitic layers of the pisoids contain only
393 minor amounts of Fe, iron was found in both FIB sections, predominantly in an oxidized form
394 (Fe(III)) based on XANES analyses at the Fe L_{2,3}-edge (Fig. S9). In FIB section 1, iron forms
395 dense spherules, ~1-2 μm in diameter (Fig. 7B,I). Several spherules form an aggregate near the
396 top of the central part of the section. Iron also forms a circular structure, creating a broken ring
397 around carbonates (near the top left corner of the section). High resolution TEM and SAED
398 analyses show that the spherules are composed of nanocrystalline, 2-line ferrihydrite (Fig. 7E,F).
399 Iron is present in a porous area of FIB section 2 (Fig. 8B,G), either as fine silicate particles (most
400 likely clays) or as coatings of quartz grains, visible as Si-rich particles on EDX maps.

401 *Organic matter distribution and characterization*

402 Organic matter, mapped using STXM analyses at the C K-edge, is found concentrated around
403 ferrihydrite spherules, or dispersed in the porosity of the aragonite matrix (Fig. 7G,8E). Organic
404 matter around the ferrihydrite spherules is visible on TEM images of FIB section 1 showing an
405 amorphous light material accumulating in the space between the spherules and the carbonate
406 matrix (Fig. 7D). This organic matter has a composition similar to that associated with the
407 protective platinum layer deposited on top of the FIB sections, suggesting that the organics
408 present in the spaces around the ferrihydrite spherules were most likely introduced during the
409 FIB milling process (“contamination” spectrum in Fig. 7M). C K-edge XANES spectra of this
410 contaminating organic matter, also present in porous areas of the aragonite matrix at the top of

411 the foils just below the platinum layer, are dominated by peaks at 284.9 eV ($1s \rightarrow \pi^*$ transitions in
412 aromatic C=C groups and unsaturated carbon), 288.6 eV ($1s \rightarrow \pi^*$ transitions in carboxylic
413 groups), 290.3 eV ($1s \rightarrow \pi^*$ transitions in carbonate groups) and 292.4 eV ($1s \rightarrow \sigma^*$ transitions in
414 aromatic C-C groups), and a small shoulder at 286.1 eV ($1s \rightarrow \pi^*$ transitions in carbonyls or
415 phenols) (Brandes *et al.*, 2004; Schumacher *et al.*, 2005; Lehmann *et al.*, 2009). This
416 composition matches that of organics present in the platinum covering the FIB foils (resulting
417 from the decomposition of an organometallic precursor), dominated by aromatic groups (Carlut
418 *et al.*, 2010; Cosmidis *et al.*, 2013), with the additional contribution of carbonates originating
419 from aragonite matrix. On the other hand, C K-edge XANES spectra obtained on organic matter
420 found dispersed in other areas of both FIB foils (Fig. 7L) have no or weak absorption signal from
421 aromatics, and sometimes display additional peaks at 287.4 eV ($1s \rightarrow 3p/\sigma^*$ transitions in
422 aliphatic carbon) (Brandes *et al.*, 2004; Lehmann *et al.*, 2009), not present in the spectra of the
423 platinum layer organics, and indicating that this organic matter is most likely endogenous. C K-
424 edge spectra of the aragonite grains are dominated by an intense peak at 290.3 eV (carbonate
425 groups), consistent with its mineralogical composition, but also present a peak at 288.6 eV,
426 which could be due to a change in bonding environment around the carbon atoms of the crystals
427 (Brandes *et al.*, 2010), or to the presence of finely intermixed organics dominated by carboxylic
428 groups.

429 **Evidence for microbe-mineral interactions in the pisoids**

430 *Morphological evidence for microbial influences in calcium carbonate and iron precipitation*

431 Ferrihydrite-rich layers of the pisoids display “shrub” textures, also called “frutexites” (Fig.
432 9A,B). Fractured pisoids etched with HCl were imaged with SEM (Fig. 9C-F). Etching dissolves

433 away the surface of the carbonate matrix, revealing the internal structure of the ferrihydrite
434 shrubs (Fig. 9C). Iron forms cavities with sub-spherical, rod-shaped or filamentous shapes (Fig.
435 9D-F). These cavities have internal diameters ranging from 0.5 to 2 μm and are interpreted as
436 casts of microbial cells. Similar looking iron-rich “honeycomb” textures were depicted in ancient
437 carbonate spring deposits and interpreted as microbial fossils (Potter-McIntyre *et al.*, 2017).

438 Preserved microbial shapes were also visible in the carbonate matrix of the pisoids, forming
439 either empty casts (Fig. 9G) or carbonate-filled molds (Fig. 9H,I), generally with larger
440 dimensions ($>5 \mu\text{m}$), and possibly corresponding to entombed microalgae.

441 *Lipid analyses*

442 A detailed description of free and carbonate-bound lipid biomarkers extracted from Crystal
443 Geysers’s pisoids and travertines can be found in the Supplementary Materials (Table S1,
444 Supplementary Text). Total lipid abundances are low overall, reflecting the low abundance of
445 organic carbon in the carbonates by mass. The carbonate-bound lipid fraction of the travertine
446 TB yields 4x lower lipid concentration than the pisoids and travertine TL. Fatty acids are the
447 major lipid components detected in all samples but significant variation between each sample
448 type and free and bound fractions are observed (Fig. 10). The ratio of bound: free fatty acids is
449 4.1:1 for the pisoids, 1.4:1 for travertine TL and 1:4.3 for travertine BL. This confirms the higher
450 abundance of microbial organic matter in the pisoids and travertine TL compared with travertine
451 TB. It is not possible to identify the source of all fatty acids as many, especially *n*-saturated and
452 monounsaturated, are common to bacteria and eukaryotes (Table S1). However, the presence of
453 *iso*- and *anteiso*- branched saturated and 3-hydroxy acids indicates a strong contribution from
454 anaerobic bacteria (Kaneda, 1991; O’Reilly *et al.*, 2017). *n*-alkanols (notably, the unusual

455 nonacosan-12-ol and hentriacontan-12-ol), β -sitosterol and stigmasterol (both C₂₉ sterols) and
456 odd-carbon-number long chain (>C₂₃) alkanes were also present in the free lipid fractions of all
457 samples, indicating the presence of vascular plants (most likely from eolian input) and
458 photosynthetic microalgae (forming visible mats the surface of the travertines) (O'Reilly *et al.*,
459 2017, and references therein). Archaeol occurs in high relative abundance in the free lipid
460 fraction of the pisoids, and is absent from the travertine samples. This confirms the presence of
461 archaea as major clades in microbial communities in the growth environment of the pisoids
462 (Kate, 1993), a finding that is consistent with the detection of abundant archaea in Crystal
463 Geyser vent waters as determined by metagenomic and lipid analyses (Probst *et al.*, 2014, 2017,
464 2020).

465 **DISCUSSION**

466 **Origin of Crystal Geyser pisoids**

467 Pisoids are coated grains larger than 2 mm displaying concentric internal layering in a cortex
468 growing around a central nucleus. Carbonate pisoids and smaller coated grains (ooids) are
469 frequent in marine environments, but have also been described in travertine settings (Kano *et al.*,
470 2019; Della Porta *et al.*, 2021). Although small and irregularly shaped ooids (<1 mm) have
471 recently been shown to form *in-situ* within microbial mats (Suarez-Gonzalez & Reitner, 2021),
472 coated grains usually thought to indicate growth while rolling in flowing water. At Crystal
473 Geyser, the pisoids may have formed in two distinct types of settings: either at the surface in
474 pools where water is flowing during large eruptions, or in the subsurface in the borehole where
475 water is turbulent during eruptions and “bubbling” events. The pisoids are mostly found in the
476 vicinity of Crystal Geyser’s vent, with larger pisoids occurring closer to the vent, and smaller

477 ones are found more distally. For these reasons, Barth and Chafetz (2015) proposed that the
478 pisoids may be formed within the plumbing system of the geyser, and ejected during eruptions.
479 The nuclei of the pisoids often contain abundant quartz grains, a major component of the
480 sandstone formation from which Crystal Geyser's waters originate, as well as cementing calcium
481 carbonates. In all pisoids analyzed with Raman, aragonite was the only CaCO_3 phase identified
482 in the nuclei (Figs. 6,S5,S6). Aragonite is also the carbonate structure found in all carbonate
483 particles filtered out of the waters expelled from Crystal Geyser's vent (Fig. S4), suggesting that
484 the nuclei of the pisoids were formed under similar geochemical conditions as these particles, i.e.
485 in the subsurface. The nuclei of the pisoids are furthermore relatively free of iron and sulfur (Fig.
486 6C,E), suggesting that they formed under reducing conditions. Indeed, in the presence of oxygen,
487 Fe^{2+} dissolved in the water would oxidize and precipitate as Fe(III) phases which would be
488 incorporated in the pisoids during growth (as observed in the cortices). Similarly, in the presence
489 of oxygen, sulfide (which is also present in the water as evidenced by the characteristic sulfide
490 smell at the vent of the geyser) would oxidize as sulfate which is readily incorporated into
491 carbonates. Thus, absence of S and Fe in the nuclei of the pisoids, along with the presence of
492 these elements in their cortices, demonstrates initial formation under reducing conditions
493 followed by further growth in more oxidizing conditions. The fact that at least some pisoid
494 growth occurs above ground is furthermore evidenced by the presence of lipid biomarkers for
495 higher plants and microalgae in the carbonate-bound fraction of pisoids. Overall, results thus
496 support an initial formation of the pisoids in the subsurface in Crystal Geyser's plumbing system,
497 with some further growth after ejection at the surface.

498 The precipitation of CaCO₃ minerals composing the pisoids is likely mostly driven by degassing
499 of CO₂-rich waters, either in the subsurface while fluids migrate vertically during eruptions and
500 “bubbling”, or in pools at the surface. Although a primarily abiotic process, microbial influences
501 on CaCO₃ mineralization in the pisoids are reflected in the high proportion of bound fatty acids
502 in these objects (Fig. 10), and the presence of microorganisms encased in the carbonate matrix
503 (Fig. 9G,I). Barth and Chafetz (2015) described carbonate spherulites, formed by aragonite
504 crystals radiating from central clumps of bacteria-shaped objects, in the nuclei of some Crystal
505 Geyser pisoids. These objects are usually interpreted as resulting from carbonate nucleation on
506 microbial cells and their EPS (Chafetz *et al.*, 2018). Influences on CaCO₃ precipitation by
507 microbial activity and EPS was documented in marine coated grains (Diaz *et al.*, 2015, 2017),
508 but the results obtained here on the Crystal Geyser pisoids do not allow to determine whether
509 similar mechanism are at play here.

510 **Calcium carbonate polymorphism and iron redox dynamics in the pisoids**

511 Carbonate polymorphism in CO₂-rich spring systems is a complex, multi-parameter problem,
512 and may be influenced by a great number of geochemical and biological factors including water
513 temperature, pH, CO₂ content and degassing rate, calcium carbonate saturation state, the
514 presence of sulfate, metals and divalent ions, organics substances, and microbial mats (Chang *et*
515 *al.*, 2017; Jones, 2017). The pisoids at Crystal Geyser are particularly interesting due to a clear
516 relationship between carbonate polymorphism and iron behavior. Indeed, their cortices are
517 composed of alternating layers of aragonite and calcite, suggestion shifting (bio)geochemical
518 conditions during pisoid growth. While calcite layers contain abundant iron, forming ferrihydrite
519 shrubs, aragonite layers contain only minor amounts of iron, present as ferrihydrite spherulites or

520 iron associated with clays, quartz grains, and organics (Figs. 6,7,8). Different hypotheses to
521 explain this correlation are discussed here.

522 *Iron control on CaCO₃ polymorphism in Crystal Geysir's pisoids*

523 A first hypothesis is that iron behavior exerts a direct control on CaCO₃ polymorphism in the
524 pisoids. Numerous experiments have shown that Fe²⁺ is an inhibitor of calcite growth (Meyer,
525 1984; Gutjahr *et al.*, 1996; de Leeuw, 2002; Mejri *et al.*, 2015), promoting the precipitation of
526 aragonite over calcite. In Crystal Geysir's plumbing system, where the pisoids are thought to
527 start forming, dissolved iron (Fe²⁺) is present at relatively high concentrations (values ranging
528 from 3.4 to 15.7 mg.L⁻¹ have been measured by ourselves and others at the vent; Table 1; Heath
529 *et al.*, 2009; Kampman *et al.*, 2014). The inhibitory effect of Fe²⁺ on calcite formation would
530 explain why the nuclei of the pisoids (as well as particles present in vent water) contain
531 exclusively aragonite. At the surface, the pisoids grow in rimstone pools proximal to the geyser,
532 where Fe²⁺ is still present (conditions similar to site 2 Fig. 2), favoring aragonite formation.
533 Ferrihydrite-rich layers in the cortices of the pisoids show that Fe²⁺ is episodically oxidized (a
534 process that may be biologically mediated – see next section), causing Fe(III) precipitation. The
535 resulting local decrease in dissolved Fe²⁺ in the pools would remove calcite inhibition and allow
536 the formation of the calcitic layers of the pisoids. In some layers, calcite seems to precipitate
537 before ferrihydrite starts to form (see for instance the outermost calcite layer on top of Fig. 6A).
538 However, it is possible that ferrihydrite incorporation in the growing pisoids occurs with a delay
539 compared to Fe²⁺ oxidation in solution. Indeed, observations of rusty materials as well as oily-
540 looking films (often attributed to iron-oxidizing bacteria; Dyer, 2003) at the surface of stagnant

541 pools around the geyser (Fig. 1D,E) indicate that following Fe^{2+} oxidation, Fe(III)-phases do not
542 immediately sink to the bottom of the pools.

543 *Competing hypotheses for CaCO_3 polymorphism in the pisoids*

544 Although the model depicted above, based on iron redox dynamics and calcite inhibition by Fe^{2+} ,
545 satisfactorily explain mineralogical observations reported in this study, competing hypotheses
546 need to be considered to account for the correlation between CaCO_3 polymorphism and iron
547 behavior in Crystal Geyser pisoids. It could be proposed that CaCO_3 polymorphs directly control
548 iron behavior. Experiments have shown that calcite has a catalytic effect on iron oxidation, with
549 Fe(II) adsorption at the surface of calcite grains accelerating the rate at which it is oxidized to
550 Fe(III) in the presence of oxygen (Mettler *et al.*, 2009). However, aragonite most likely has a
551 similar catalytic effect on iron oxidation. It is thus unlikely that CaCO_3 polymorphism directly
552 controls iron oxidation and distribution.

553 Since at least part of the pisoids growth occurs in the plumbing system of the geyser, it is
554 necessary to consider the potential impact of physicochemical fluctuations in the subsurface on
555 carbonate polymorphism. These fluctuations are mostly driven by the eruption cycle of the
556 geyser (Kampman *et al.*, 2014; Han *et al.*, 2017). The temperature and pH of the geyser waters
557 are relatively constant over time and through the eruption cycle, with temperature variations
558 smaller than 3.5 °C (ranging from 15.5 to 18.8 °C) and pH variations smaller than 1.5 units
559 (ranging from 6.2 to 7.6), as measured by several authors (Baer & Rigby, 1978; Shipton *et al.*,
560 2004; Assayag *et al.*, 2009; Heath *et al.*, 2009; Takashima *et al.*, 2011b; Kampman *et al.*, 2014;
561 Emerson *et al.*, 2016; Han *et al.*, 2017). Temperature favors aragonite precipitation at values
562 greater than 35 °C, and the influence of pH on CaCO_3 polymorphism is relatively insignificant

563 compared with other chemical parameters except at pH values higher than 10 (Chang *et al.*,
564 2017; Jones, 2017). It thus seems unlikely that variations in Crystal Geyser's water temperature
565 and pH may be driving CaCO₃ polymorphism in the pisoids. Similar to the effect of Fe²⁺, the
566 presence of magnesium (as Mg²⁺ ions) in solution inhibits calcite growth and promotes aragonite
567 precipitation. The Mg/Ca ratio appears to be particularly important for controlling CaCO₃
568 polymorphism (with higher ratios favoring the precipitation of aragonite over calcite) (Lin &
569 Singer, 2009). Mg/Ca molar ratios in Crystal Geyser's waters measured by ourselves (Table 1)
570 and others (Kampman *et al.*, 2014; Han *et al.*, 2017) are remarkably constant around 0.37 (±
571 0.01), and variations of this parameter over time are thus unlikely to be the cause of changes in
572 CaCO₃ polymorphism at this site. The presence of strontium is also a factor favoring aragonite
573 precipitation over calcite (Jones, 2017). Sr variations in Crystal Geyser's waters are relatively
574 wide (~127-220 μM; Kampman *et al.*, 2014; Han *et al.*, 2017). However, EPMA analyses show
575 constant concentrations of Sr across aragonite and calcite layers of a pisoid (Fig. S7), therefore
576 Sr probably do not control CaCO₃ polymorphism. The presence of sulfate in solution is known to
577 selectively inhibit calcite growth (Walter, 1985), especially in the presence of Mg²⁺ (Nielsen *et*
578 *al.*, 2016). However, Han *et al.* (2017) measured that SO₄²⁺ concentrations in Crstal Geyser vent
579 waters vary by less than 15% through several eruption periods, suggesting that sulfate fluctuation
580 in the subsurface probably do not affect CaCO₃ polymorphism in the pisoids.

581 CO₂ degassing rate is an important factor likely to vary dramatically over an eruption cycle of
582 the geyser. Aragonite precipitation is thought to be favored over calcite in waters with high CO₂
583 degassing rates (Holland *et al.*, 1964; Jones, 2017). This effect is consistent with the presence of
584 aragonite in the nuclei of the pisoids, forming within the plumbing system of the geyser, where

585 CO₂ content and degassing rates (due to turbulent mixing during the vertical migration of the
586 fluid) are high. However, a model where CO₂ degassing rate is the main driver of CaCO₃
587 polymorphism cannot account for calcite and Fe(III) co-precipitation in the pisoids. Indeed,
588 intense CO₂ degassing correlates with periods of water-air mixing during eruptions or
589 “bubbling”, i.e. turbulent events that are also likely to cause Fe²⁺ oxidation and Fe(III)
590 precipitation. Thus, if CO₂ degassing were the main driver for CaCO₃ polymorphism in the
591 pisoids, ferrihydrite would be mostly co-located with aragonite layers rather than calcite.

592 Physicochemical changes that may be occurring in the rimstone pools where (at least some of)
593 the pisoid growth is occurring should now be considered. Unfortunately, geochemical parameters
594 in pools were not measured in time series. However, an indication of the changes that may affect
595 Crystal Geysir water once at the surface is shown by the geochemical profile in Figure 2.
596 Depending on air temperature, the temperature of the water in the pools may increase with time,
597 but it is not likely to reach the values (well above 35 °C) where it affects CaCO₃ polymorphism.
598 Similarly, pH is unlikely to show dramatic changes after a slight increase (by less than 1.5 pH
599 units) due to CO₂ degassing. No data on the evolution of Mg, Sr, SO₄²⁻, or other species likely to
600 affect the structure of precipitating CaCO₃ in pools, has been acquired. However, the correlation
601 between CaCO₃ polymorphism and iron distribution in the pisoids suggest that what controls
602 shifts from aragonitic to calcitic conditions is probably a redox-active process. Iron is
603 experiencing dramatic changes at the surface due to oxidation (Fig. 2), as also shown by changes
604 in the visual aspect of the pools (Fig. 1D,E). Sulfate is also likely to improve with time in the
605 pools due to oxidation of reduced sulfur species. However, increases in SO₄²⁻ concentrations
606 should favor aragonite formation over calcite. If sulfur oxidation in surface pools was controlling

607 CaCO₃ polymorphism, periods of pool water oxygenation and consecutive Fe(III) formation
608 would correspond to the precipitation of aragonite layers. Since the opposite correlation is
609 observed, sulfate variations can be ruled out as an important factor in CaCO₃ polymorphism. It
610 can be concluded that CaCO₃ polymorphism in Crystal Geyser pisoids is mostly controlled by
611 iron redox dynamics, either in the subsurface or in surface pools where the pisoids are growing.

612 *Microbial iron oxidation in the pisoids formation environment*

613 As noted by Kappler *et al.* (2021), in the presence of oxygen, iron oxidation by biotic and abiotic
614 pathways typically occur in parallel, making it challenging to identify the occurrence and
615 quantitative contribution of iron-oxidizing bacteria to overall Fe(II) oxidation. Shiraishi *et al.*,
616 (2018) showed that both microbial and abiotic iron oxidation may account for ferrihydrite
617 deposition in spring systems, depending on fluctuations of the water oxygen concentration.
618 However, several lines of evidence suggest that ferrihydrite in Crystal Geyser's pisoids is at least
619 in part a product of microbial iron oxidation, either at the subsurface in pools, or in the
620 subsurface in the plumbing system of the geyser. Oily-looking, iridescent films developing in
621 stagnant pools after eruptions (Fig. 1E) are a typical signature of microaerophilic iron-oxidizing
622 bacteria (Dyer, 2003; Fru *et al.*, 2012). Neutrophilic microaerophilic iron-oxidizing bacteria
623 dominate microbial diversity in Crystal Geyser's vent water (Emerson *et al.*, 2016), including
624 *Mariprofundus* (Zetaproteobacteria), and several members of the Gallionellales
625 (Betaproteobacteria), such as *Gallionella* and *Sideroxydans*. Iron oxidizers of the genera
626 *Mariprofundus* and *Gallionella* produce recognizable extracellular structures forming ribbon-like,
627 twisted stalks composed of Fe-(oxyhydr)oxides and organic polymers (Chan *et al.*, 2011, 2016),
628 which were not found in the pisoids. However, ferrihydrite in the iron-rich layers of the pisoids

629 forms “honeycomb” microtextures, encasing sub-spherical, rod-shaped or filamentous shapes
630 with sizes ranging from 0.5-2 μm (Fig. 9D-F), interpreted as iron-encrusted microbial cells
631 (Potter-McIntyre *et al.*, 2017). The encrusted cells may correspond to iron-oxidizers of the genus
632 *Sideroxydans* or other members of the *Gallionellales* which precipitate extracellular Fe-
633 (oxyhydr)oxides not associated with any stalks or other recognizable extracellular structures
634 (Emerson & Moyer, 1997; Weiss *et al.*, 2007; Fleming *et al.*, 2014). The overall texture of the
635 ferrihydrite layers of the pisoids furthermore corresponds to what has been described as iron
636 shrubs (Chafetz *et al.*, 1998; Chafetz & Guidry, 1999; Takashima *et al.*, 2008; Parenteau &
637 Cady, 2010) or frutexites (Jakubowicz *et al.*, 2014; Guido *et al.*, 2016; Reitner *et al.*, 2017;
638 Grădinaru *et al.*, 2020) (Fig. 9A,B), and which are commonly interpreted as microbial in origin.
639 Of particular relevance here, upward-branching iron shrubs described by Takashima *et al.* (2008)
640 in laminated travertines forming at the Shionoha hot spring (Japan) are composed of ferrihydrites
641 encrusting rod-shaped structures produced by microaerophilic iron-oxidizers of the genus
642 *Siderooxidans*. Shrub-like dendritic iron-oxide structures associated with Gallionellaceae were
643 also described in carbonates forming from CO_2 - and iron-rich circumneutral hot springs at
644 Okuoku-hachikurou Onsen (Japan) by Ward *et al.* (2017). In other hot spring laminated
645 travertines (Ilia Hot Spring, Greece), iron shrubs are associated with iron-oxidizing
646 Zetaproteobacteria (Kanellopoulos *et al.*, 2019). Iron shrubs can also be formed by
647 microorganisms other than microaerophilic iron-oxidizers. For instance, in microbial mats
648 forming in iron-rich hot springs (Chocolate Pots, Yellowstone National Park), iron shrubs are
649 produced by cyanobacteria such as *Oscillatoria*, *Synechococcus*, and *Cyanothece* encrusted with
650 ferrihydrite (Trouwborst *et al.*, 2007; Parenteau & Cady, 2010).

651 The iron-rich layers in Crystal Geyser pisoids may thus record changes in the abundance and
652 activity of neutrophilic microaerophilic iron-oxidizers, episodically precipitating Fe(III).
653 Abundances of these microorganisms can fluctuate in the environment due to a number of
654 physicochemical factors that may include availability of complex organic carbon, iron
655 abundance, and the steepness of the redoxcline (Fleming *et al.*, 2014; Blackwell *et al.*, 2019). At
656 Crystal Geyser, the eruption cycle is likely to be an important factor controlling variations in the
657 abundance and activity of iron-oxidizing bacteria. At the surface, iron-oxidizers may bloom after
658 each eruption of the geyser, introducing reduced iron from the subsurface. In the subsurface,
659 iron-oxidizers may be active during eruptions of bubbling events when turbulent mixing
660 introduces oxygen in water.

661 In iron-poor (aragonitic) layers of the pisoids, iron is present as ferrihydrite spherules (Fig. 7),
662 Fe(III) in clays and coatings of quartz grains, but also associated with organic matter found in the
663 porosity of the carbonate matrix (Fig. 8). Fe(III) has a strong affinity for organic matter,
664 adsorbing on negatively charged functional groups such as carboxylates or phosphorylates
665 (González *et al.*, 2014), and frequently forms organo-ferric colloids in the environment (Ilina *et*
666 *al.*, 2016; Liao *et al.*, 2017). The origin of the ferrihydrite microspherules is more enigmatic.
667 Since they are included in the aragonite matrix, they are likely to have formed in solution prior to
668 CaCO₃ formation. The presence of some space between the spherules and the carbonates (where
669 contaminating organic matter could accumulate during the FIB milling process; Fig. 7D,K,M),
670 suggests some shrinking after their incorporation within the aragonite matrix. Spheroidal
671 ferrihydrite particles were observed to form aggregates around bacteria in iron-rich laminated

672 carbonate spring deposits (Takashima *et al.*, 2011a) but an abiotic origin for the ferrihydrite
673 microspherules in the Crystal Geyser pisoids cannot be discounted.

674 **Microbial influences on CaCO₃ precipitation and polymorphism in Crystal Geyser**
675 **travertines**

676 *Evidence for microbial influence on travertine formation*

677 Travertine formation at Crystal Geyser is most probably a primarily abiotic process resulting
678 from CO₂ degassing. It is unclear what impact microbial activities may have on the intensity of
679 CaCO₃ precipitation in such CO₂-rich environment. However, microbial influences on CaCO₃
680 mineralization have been documented for many travertine systems (Shiraishi *et al.*, 2008; Perri *et*
681 *al.*, 2012; Okumura *et al.*, 2013a; Kano *et al.*, 2019; Della Porta *et al.*, 2021), producing
682 recognizable sedimentary fabrics and textures (Guo & Riding, 1992; Kano *et al.*, 2019). In
683 Crystal Geyser travertines, such microbial influences are thought to be responsible for
684 lamination, microstromatolitic horizons, and other features such as botryoidal carbonate textures
685 (Takashima *et al.*, 2011b; Barth & Chafetz, 2015). Lipid analyses have shown abundant bound
686 fatty acids in the laminated travertines (TL), while most lipids in the non-laminated travertine
687 (TB) were free (Fig. 10). This difference may indicate more important contributions of
688 microorganisms to CaCO₃ precipitation in travertine TL, as compared with travertine TB, in
689 agreement with the absence of lamination in the latter sample. However, the presence of
690 fenestrae and calcified bubbles in travertine TB (Fig. 5A) indicates gas formation concurrent
691 with CaCO₃ precipitation, possibly resulting from microbial activity (e.g., O₂ production by
692 aerobic phototrophs; Bosak *et al.*, 2010; Della Porta *et al.*, 2021).

693 *Origin of CaCO₃ polymorphism in the travertines*

694 Iron oxidation is not likely to be important factor controlling CaCO_3 polymorphism in the
695 travertines, which are formed at a distance from the geyser, where the well-oxygenated waters
696 are Fe^{2+} -poor (conditions similar to site 3 in Fig. 2). Ferrihydrite is present at relatively low
697 abundances (compared with the pisoids) and there is no correlation between iron distribution and
698 CaCO_3 polymorphism in the travertines. Aragonite and calcite are concentrated respectively in
699 different microlaminae of travertine TL, but both phases can be found together in each
700 microlaminae. In travertine TB, grains of both aragonite and calcite are found together in the
701 same areas (Figs. 4,5). Small-scale variations of CaCO_3 polymorphism in the travertines
702 probably results from localized physicochemical changes occurring in micro-environments at the
703 surface of the growing travertines, most likely under the influence of microorganisms and their
704 EPS. Peng & Jones, (2013) described the co-precipitation of calcite, aragonite and amorphous
705 calcium carbonate within distances of a few microns in hot spring deposits. They proposed that
706 microbial biofilms growing on the carbonates were forming microdomains, within which specific
707 physicochemical conditions developed as a result of microbial activity, influencing CaCO_3
708 precipitation and polymorphism. Biofilms contain abundant EPS, which influence CaCO_3
709 precipitation principally due the presence of negatively charged, Ca^{2+} -binding organic functional
710 groups (Dupraz *et al.*, 2009), and have been shown exert a control on CaCO_3 polymorphism in
711 laboratory experiments (Tourney & Ngwenya, 2008, 2009). Microlaminations formed by
712 alternating layers of aragonite and calcite in hot spring travertines in Japan were interpreted as
713 resulting from diurnal cycles affecting microbial EPS (Okumura *et al.*, 2013b): during the day,
714 EPS built by photosynthetic organisms would bind Ca^{2+} , reducing supersaturation and promoting
715 calcite formation, while at night, decomposition of EPS by heterotrophs would release Ca^{2+} and

716 lead to aragonite precipitation. Similarly, CaCO₃ polymorphism in Italian hot spring travertine
717 deposits was interpreted to reflect microbial diurnal control (Guo & Riding, 1992), but there
718 aragonite was thought to grow during the day. Different types of microbial metabolisms,
719 including oxygenic photosynthesis (performed by cyanobacteria and microalgae visibly thriving
720 at the surface of Crystal Geyser travertines in warm months; Fig. S1A,B) may cause local
721 increases in alkalinity and saturation with respect to CaCO₃ phases (Dupraz *et al.*, 2009), also
722 possibly affecting CaCO₃ polymorphism. CaCO₃ polymorphism in microbial precipitation
723 experiments was recently shown to be affected by bacterial metabolism in a strain-specific way,
724 which was interpreted as differential precipitation kinetics resulting from different levels of
725 enzymatic activities (Clarà Saracho *et al.*, 2020). It can be concluded that different types of
726 microbial processes may be responsible for the small-scale variations in CaCO₃ polymorphism in
727 Crystal Geyser travertines, but that, as opposed to the pisoids, iron redox dynamics is not an
728 important factor. Barth & Chafetz (2015) reported the presence of iron-rich tube-like structures
729 morphologically similar to the iron sheaths produced by microaerophilic iron-oxidizing bacteria
730 such as *Leptothrix ochracea* (Fleming *et al.*, 2011) in Crystal Geyser's travertines. However, the
731 absence of well-defined ferrihydrite layers suggests that the travertine-formation area does not
732 experience intense blooms of iron oxidizers. Moreover, iron oxidizers were not detected in
733 bacterial 16s rDNA phylotype analyses performed on a Crystal Geyser travertine sample by
734 Takashima *et al.* (2011), showing that they are not dominant members of the bacterial
735 community thriving at the surface of these travertines.

736 CONCLUSIONS

737 Pisoids and travertines formed at Crystal Geysers, a natural analogue for CO₂ leakage at CCS
738 sites, were characterized. Microbially driven iron oxidation was shown to exert a strong
739 influence on CaCO₃ polymorphism, as recorded in the pisoids. In the travertines, microbial
740 activity may produce small-scale variations in CaCO₃ polymorphism, and textural features such
741 as laminations and calcified gas bubbles. A control on CaCO₃ polymorphism by iron redox
742 dynamics was shown here for the first time in a natural environment. Microbial iron oxidation
743 may play an important role in controlling polymorphism of the carbonate products of CO₂ escape
744 from geological storage, and may also be relevant to subsurface carbonation at CCS sites.
745 Indeed, blooms of iron-oxidizing Betaproteobacteria have been occurring following CO₂
746 injections at a geological CO₂ storage site (Trias *et al.*, 2017). Fractures can introduce oxygen in
747 deep groundwater, creating iron-oxidizer hotspots in the subsurface (Bochet *et al.*, 2020). Future
748 studies will have to determine whether intense microbial iron oxidation may influence subsurface
749 carbonate formation and stability, impacting the efficacy of geological CO₂ storage.

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772 **DATA AVAILABILITY STATEMENT**

773 The data supporting the findings of this study are openly available in the supplied supplementary
774 figures and tables.

775

776 **TABLE**

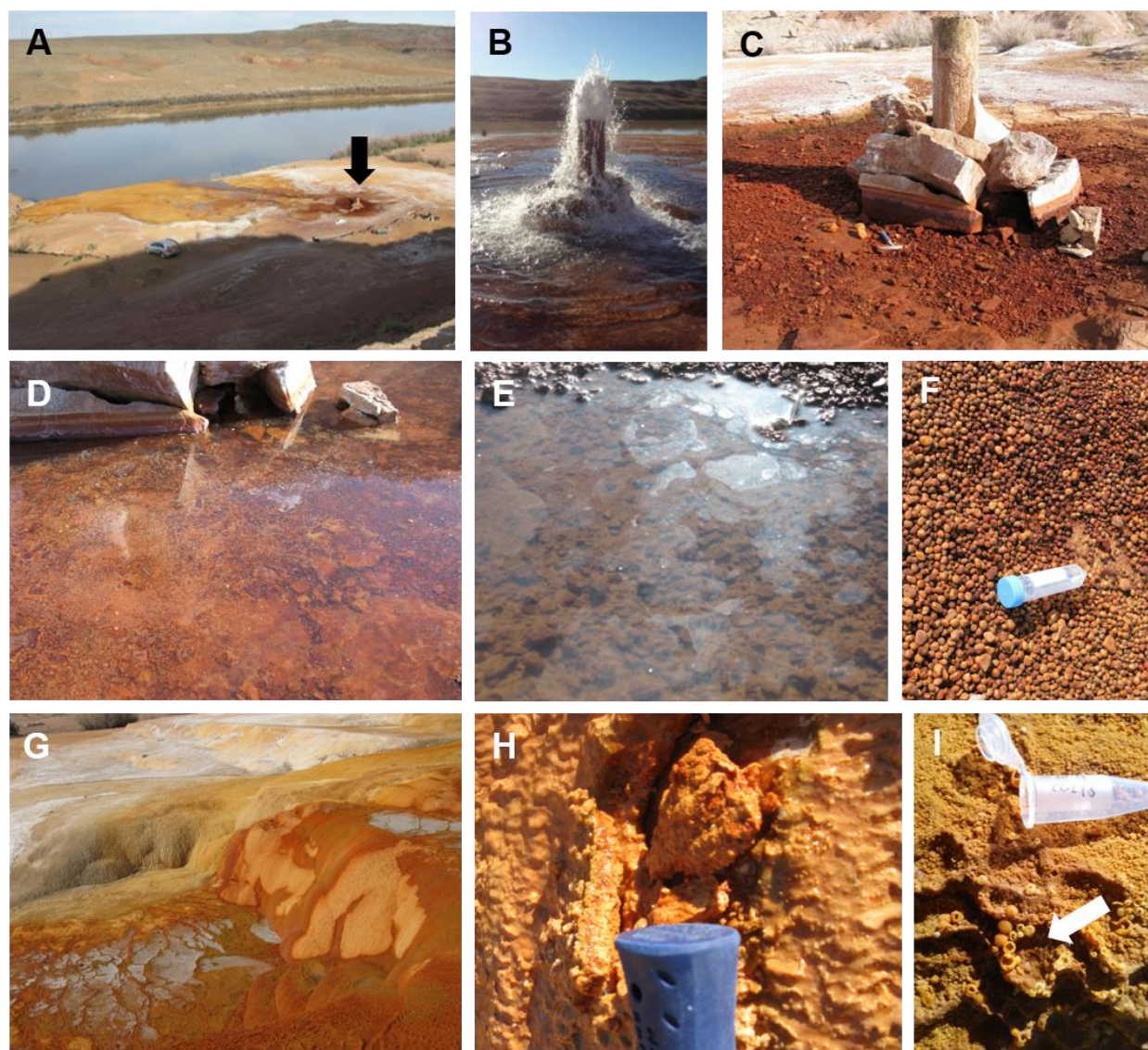
SiO₂ and cations	Si ppm	Mn ppm	Fe ppm	Mg ppm	Ca ppm	Al ppm	Sr ppm	Na ppm	K ppm
<i>Detection limit</i>	0.026	0.001	0.003	0.002	0.003	0.002	0.001	0.020	0.029
2014-1	8.12	1.86	10.61	249.6	1089	<DL	15.4	4231.5	351.9
2014-2	8.14	1.84	10.56	248.2	1092	<DL	15.3	4303.9	370.2
2015	6.6	1.37	6.14	221.3	955	<DL		3616	324

Anions	F ppm	Cl ppm	NO ₂ ppm	Br ppm	NO ₃ ppm	PO ₄ ppm	SO ₄ ppm
<i>Detection limit</i>	0.02	0.02	0.03	0.05	0.05	0.2	0.2
2014-1	<DL	4817	<DL	1.29	<DL	<DL	2433.9
2014-2	<DL	4830	<DL	1.38	<DL	<DL	2432.4

777 **Table 1.** Major elements in Crystal Geysir vent waters collected in April 2014 (on two different
778 days) and in February 2015. <DL: below detection limit. Note that Sr and anions were not
779 measured in 2015.

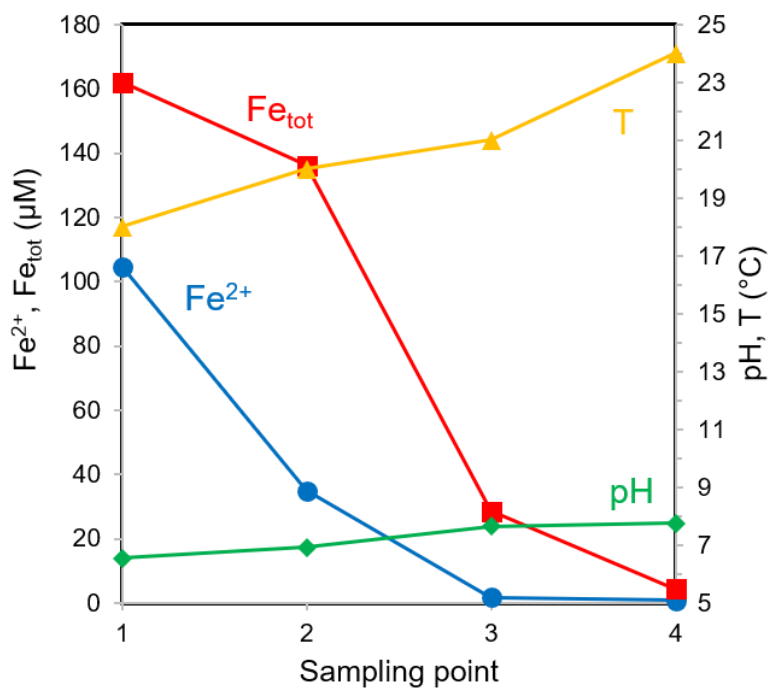
780

781

782 **FIGURES**

783
 784 **Figure 1.** Photographs of the Crystal Geyser field site, pools, pisoids and travertines. (A) Aerial
 785 view of Crystal Geyser (black arrow) between two eruptions. The reddish pool around the geyser
 786 is visible, as well as the orange travertines deposited downstream along the flow path of the
 787 geyser waters to Green River. (B) Crystal Geyser during an eruption. (C) The geyser and its
 788 surrounding pools completely dry up between eruptions. (D) Pool formed around the geyser after

789 an eruption. Iron-(oxyhydr)oxides forming at the surface of the pool are visible. (E) Oily-looking
790 films formed at the surface of wet pools in the pisoid-formation area. (F) Pisoids found near the
791 geyser. 50 mL falcon tube shown for scale. (G) Terraced travertines found downstream. (H)
792 Close-up on a section of the “layered” travertine (TL). Rock hammer grip shown for scale. (I)
793 Calcified bubbles (white arrow) at the surface of the “bubbly” travertine (TB). 1.5 mL
794 microcentrifuge tube shown for scale. Photographs were taken in April 2014 (A, C-H) and
795 February 2015 (B,I).
796



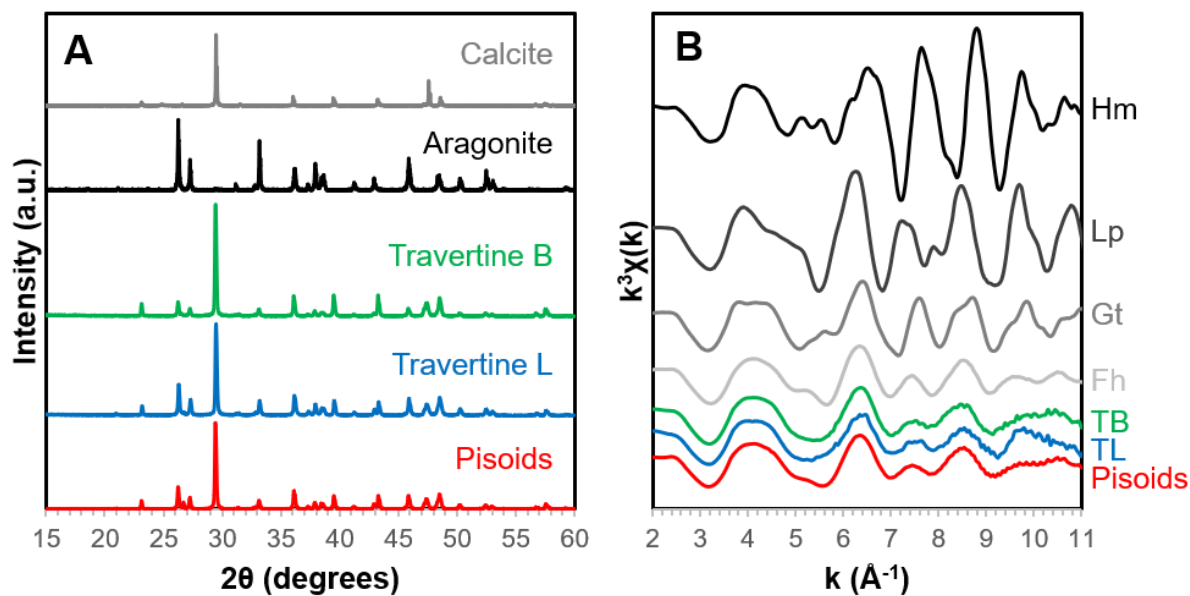
797

798 **Figure 2.** Geochemical profiles along the flow path of Crystal Geyser waters. Sampling point #1

799 corresponds to the pool directly surrounding the geyser, and sampling point #4 corresponds to

800 the water entering Green River. Blue circles: concentration of dissolved Fe(II) (Fe²⁺). Red801 squares: concentration of total dissolved Fe (Fe_{tot}). Yellow triangles: temperature (T). Green

802 diamonds: pH.



803

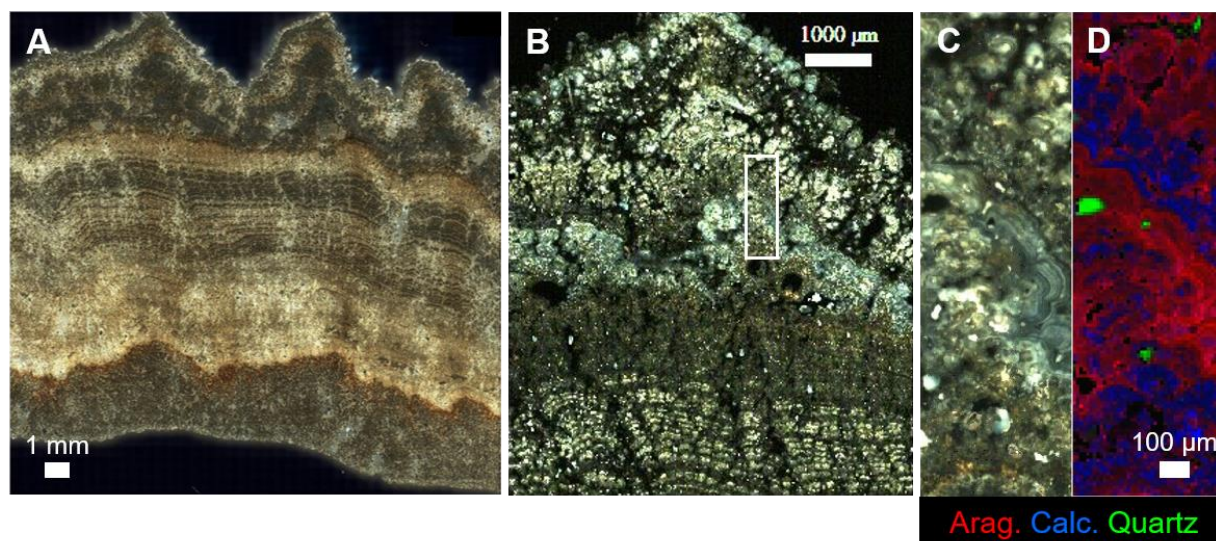
804 **Figure 3.** Mineralogy of the Crystal Geyser travertines (TL and TB) and pisoids. (A) XRD

805 spectra. (B) Fe K-edge EXAFS spectra. Reference spectra for calcite and aragonite (A) and

806 ferrihydrite (Fh), Goethite (Gt), Lepidocrocite (Lp) and Hematite (Hm) (B) shown for

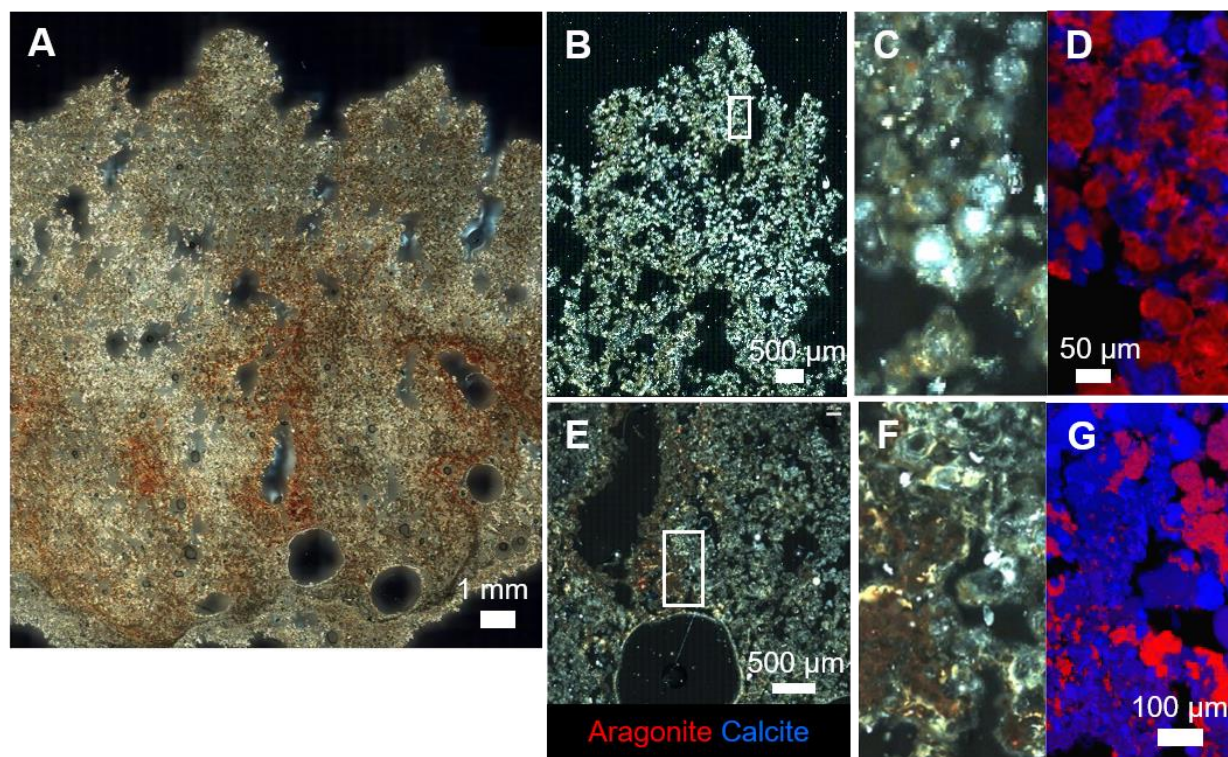
807 comparison.

808

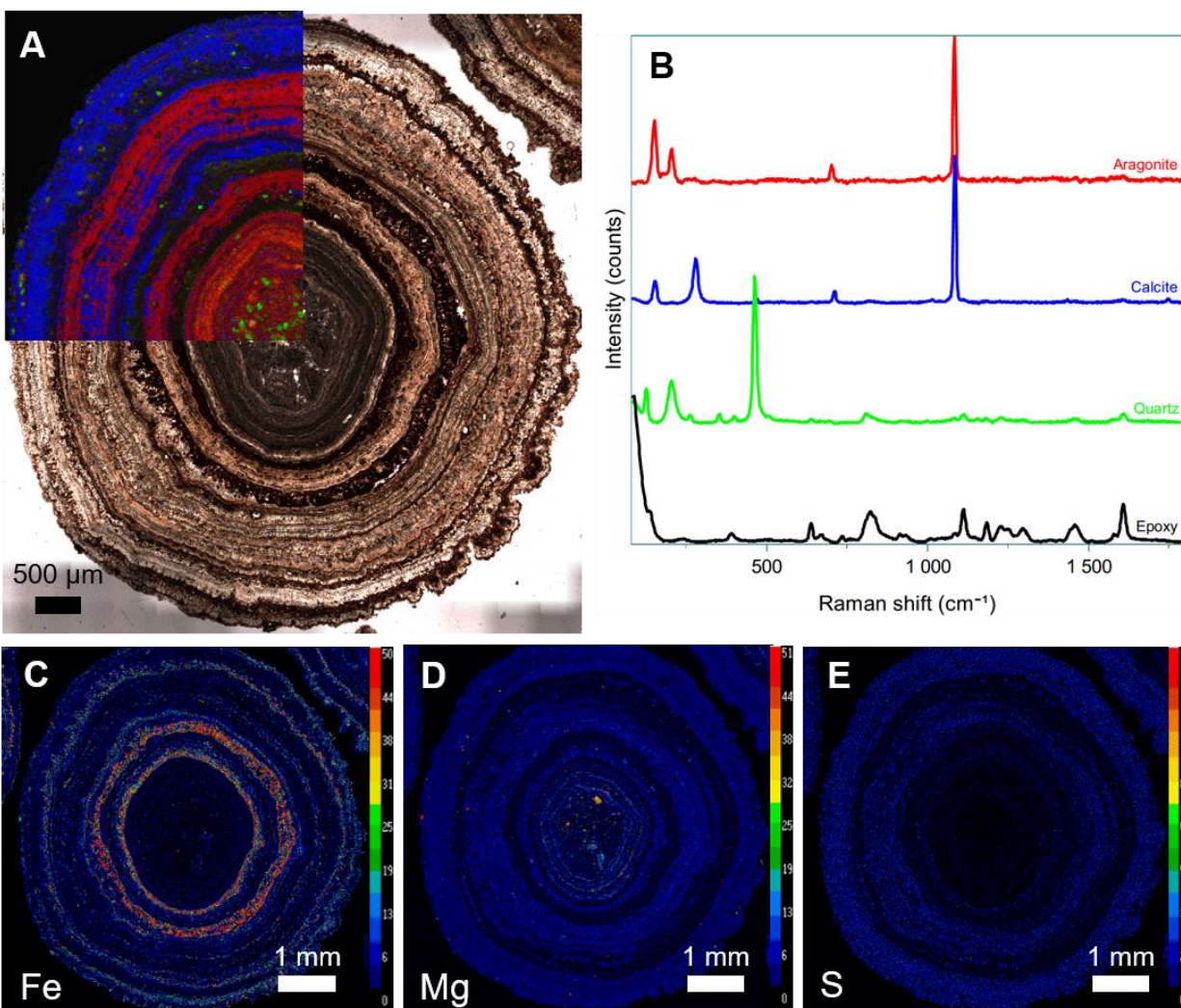


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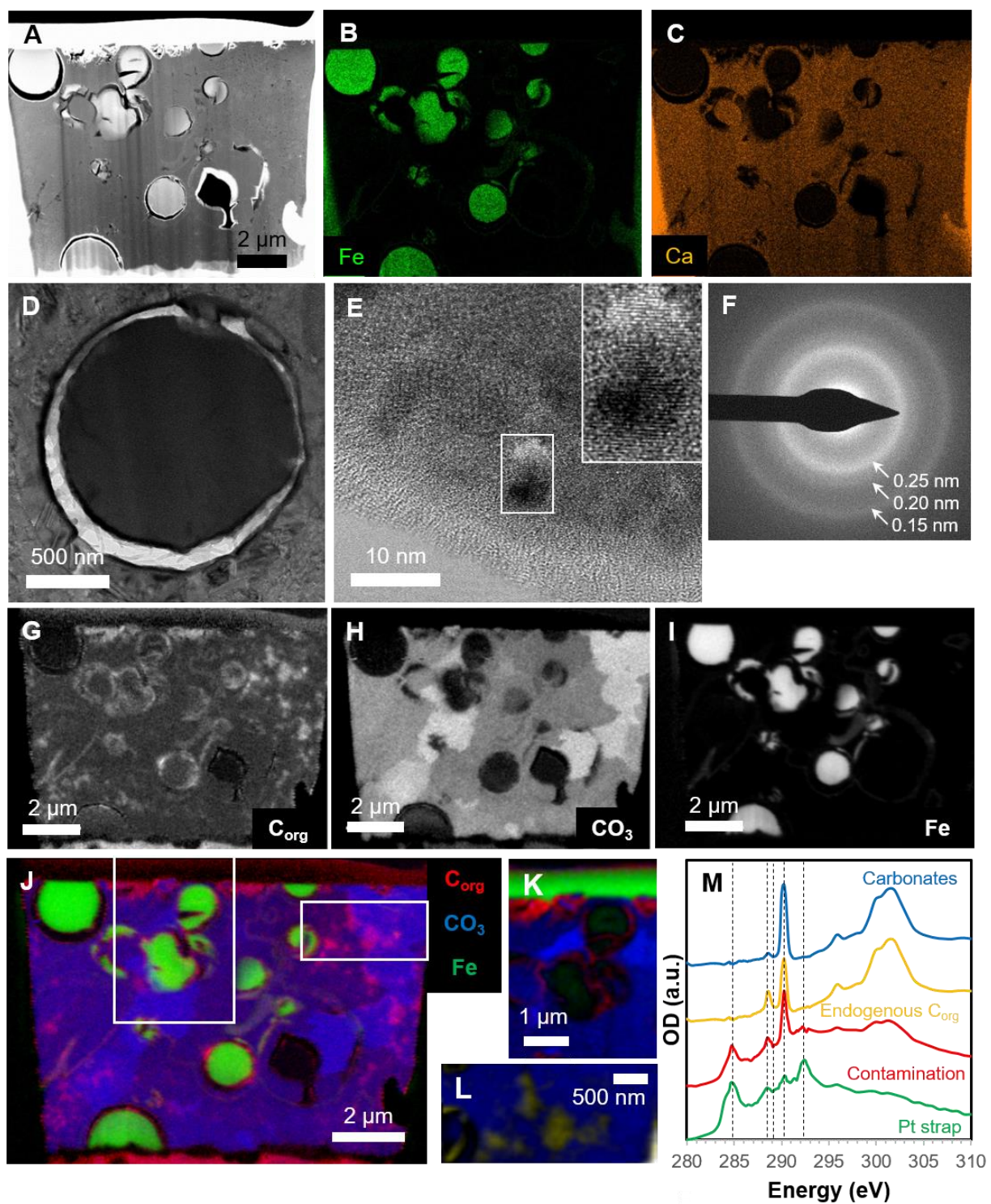
810 **Figure 4.** Micrographs and Raman map of the “layered” travertine (TB). (A) Large composite
811 micrograph in cross-polarized transmitted light. (B) Close-up in cross-polarized light showing
812 the location of the Raman map (white rectangle). (C) Cross-polarized image corresponding to the
813 area mapped with Raman. (D) Raman map showing the distribution of aragonite (red), calcite
814 (blue) and quartz (green).



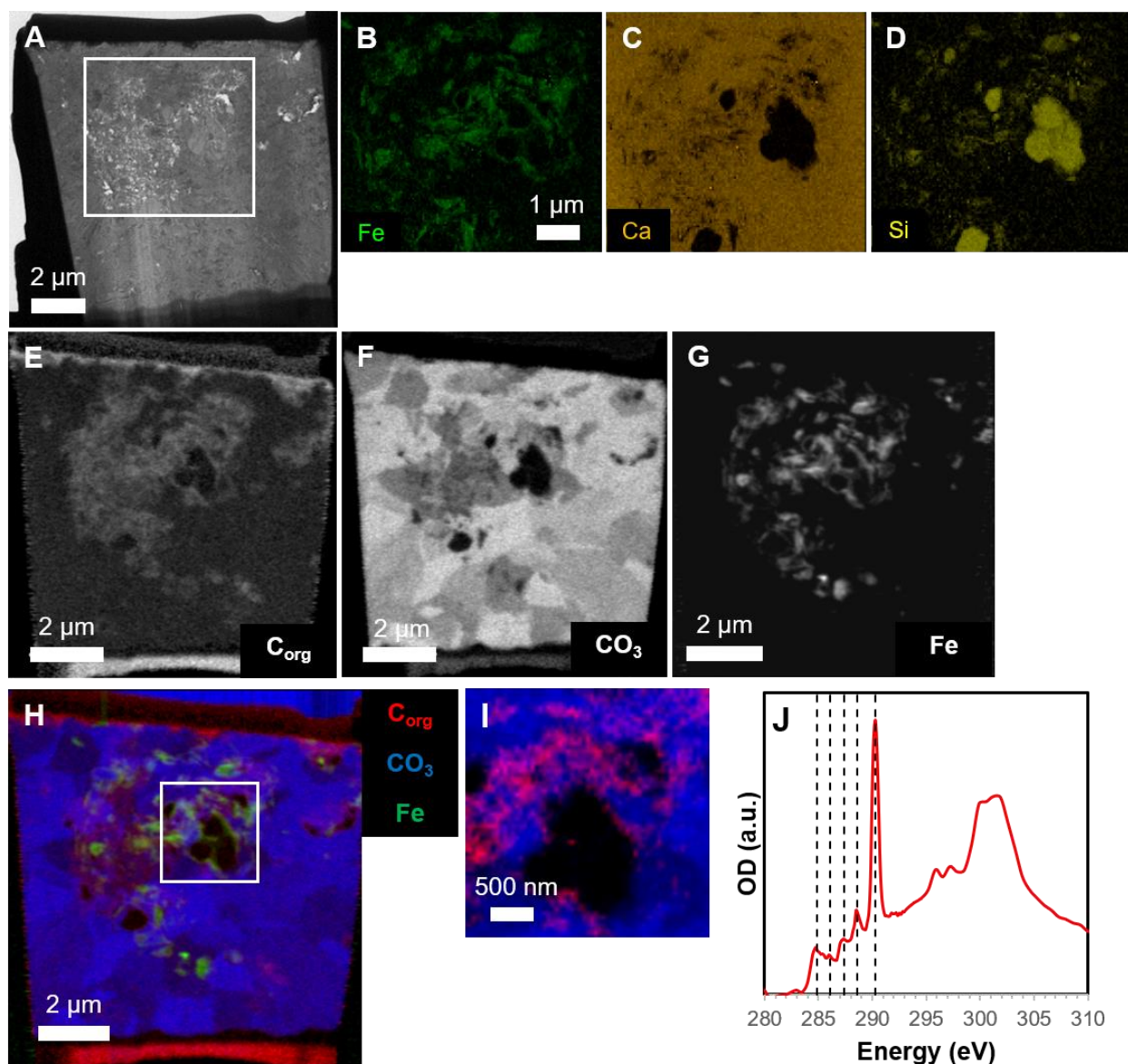
815
 816 **Figure 5.** Micrographs and Raman maps of the “bubbly” travertine (TB). (A) Large composite
 817 micrograph in cross-polarized transmitted light. (B,E) Close-ups in cross-polarized light showing
 818 the location of the Raman maps (white rectangles). (C,F) Cross-polarized images corresponding
 819 to the areas mapped with Raman. (D,G) Raman maps showing the distribution of aragonite (red)
 820 and calcite (blue).
 821



822
 823 **Figure 6.** Mineralogical and chemical mapping of a simple pisoid. (A) Montage of transmitted
 824 light micrographs and overlaid Raman map of a pisoid, showing the distribution of aragonite
 825 (red) and calcite (blue), as well as quartz (green). (B) Representative end-member Raman spectra
 826 of aragonite, calcite, quartz, and embedding epoxy resin, used to fit the map datasets. (C-E)
 827 Elemental maps of Fe, Mg and S acquired on the same pisoid using EPMA. The color scales
 828 represent relative abundances of the different elements in the sample, but the number indicated
 829 on the scales are arbitrary (no calibration has been performed).



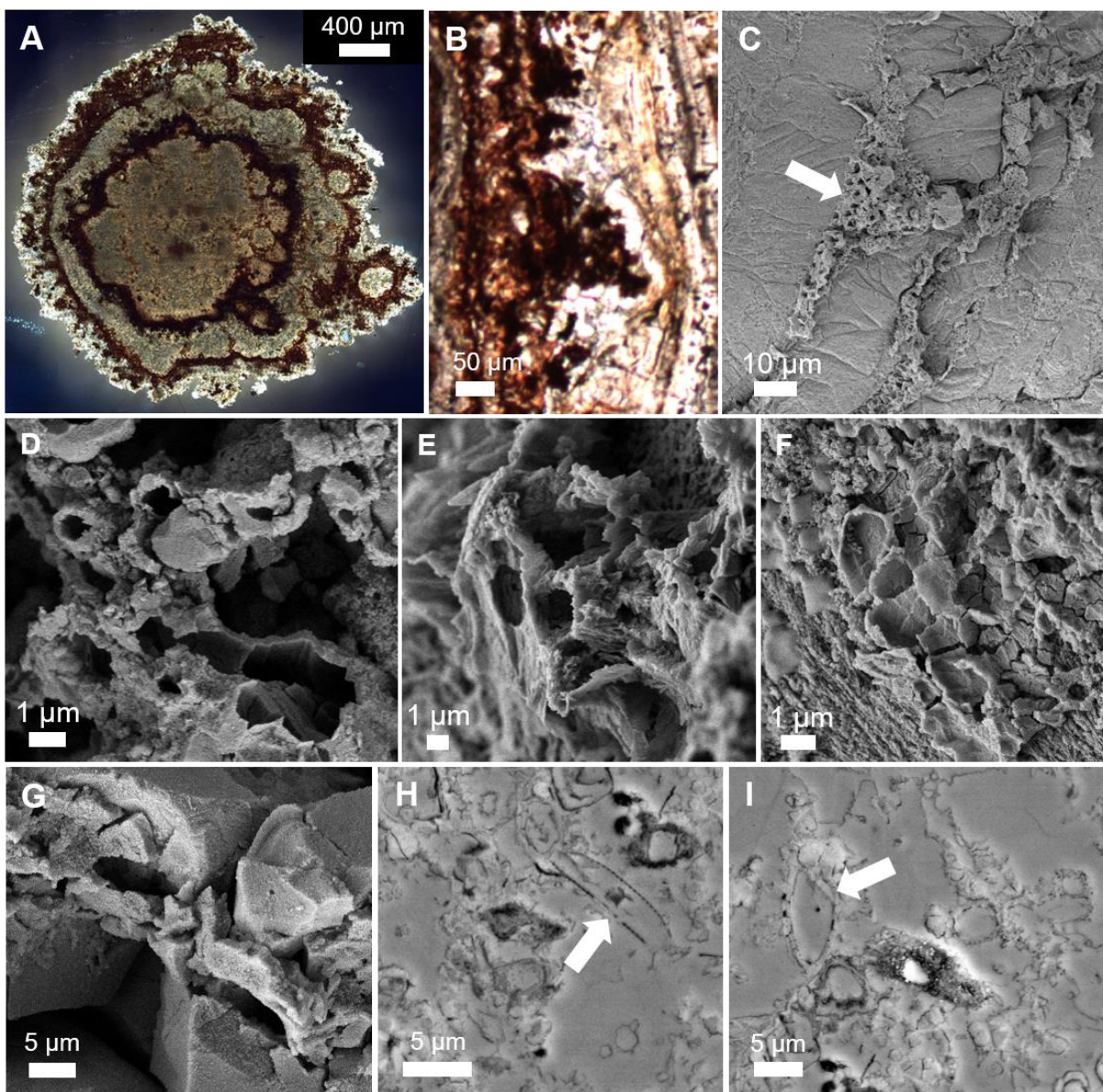
831 **Figure 7.** TEM (A-F) and STXM (G-M) analyses of FIB section 1 (pisoid). (A) STEM high
832 angle annular dark field image. (B) Fe map. (C) Ca map. (D) Bright field TEM close-up on an
833 iron-rich spherule. (E) High-resolution TEM image obtained on the border of a spherule. Small
834 (<10 nm) crystalline domains are visible. A close-up on a crystalline domain shows lattice
835 fringes with a spacing of ~0.2 nm. (F) SAED pattern obtained on the spherule border, showing
836 diffuse rings, and confirming the nanocrystalline structure of the spherules. d-spacing values,
837 corresponding to ferrihydrite, are indicated. (G) Organic carbon map. (H) Carbonate maps. (I)
838 Iron map. (J) Composite STXM map showing the distribution of organic carbon (red),
839 carbonates (blue) and iron (green). The white boxes show the areas corresponding to maps and
840 spectra shown in (K-M). (K) Map showing the protective platinum layer at the top of the FIB
841 section 1 (green), contaminating organics (red) and carbonate minerals (blue). (L) Map showing
842 organic-rich areas (yellow) and carbonate minerals (blue). (M) Representative C K-edge XANES
843 spectra for the carbonate matrix, the platinum layer, and organic contaminants in (K) and the
844 organic-rich areas in (L). Vertical lines show the positions of absorption peaks at 284.9 eV,
845 286.1 eV, 287.4 eV, 288.6 eV, 290.3 eV and 292.4 eV.



846

847 **Figure 8.** TEM (A-D) and STXM (E-J) analyses of FIB section 2 (pisoid). (A) TEM bright field
 848 image of the whole section. The white rectangle shows the areas mapped in (B-D). (B) Fe map.
 849 (C) Ca map. (D) Si map showing the presence of quartz grains. (E) Organic carbon map. (F)
 850 Carbonate map. (G) Iron map. (H) Composite STXM map showing the distribution of organic
 851 carbon (red), carbonates (blue) and iron (green). The white box shows the areas mapped in (I).

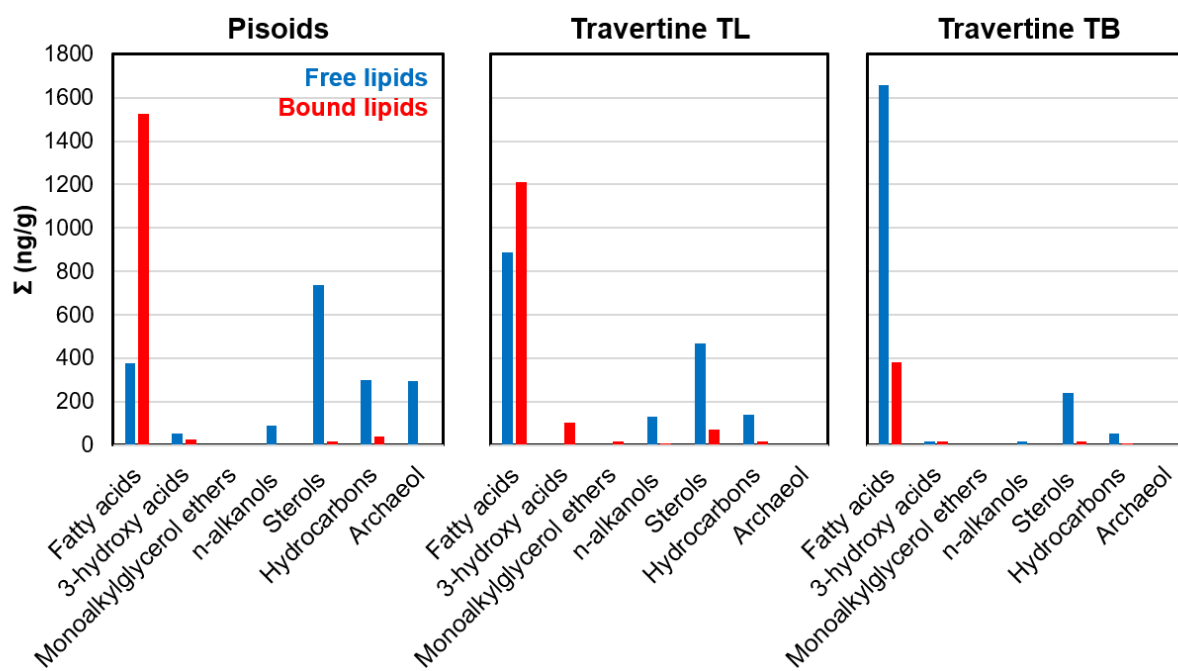
852 (I) Map showing the distribution of organic matter (red) and carbonate minerals (blue). (J) C K-
853 edge XANES spectrum representative of the organic matter mapped in red in (I). Vertical lines
854 show the positions of absorption peaks at 284.9 eV, 286.1 eV, 287.4 eV, 288.6 eV, and 290.3
855 eV.
856



857
 858 **Figure 9.** Optical micrographs and SEM of pisoids. (A) Montage of light micrographs showing a
 859 pisoid. (B) Close-up on iron shrubs. (C-F) SEM images of HCl-etched pisoids. (C) Low
 860 magnification image showing the partially dissolved carbonate matrix revealing the structure of
 861 an Fe-rich layer (arrow). (D-F) Close-ups showing microbial iron casts. The images were
 862 obtained on three different pisoids. (G-I) SEM images of microbial shapes in carbonate layers of

863 the pisoids. (G) Empty rod-shaped casts. (H,I) Possible microalgae entombed in the carbonates
864 (arrows). (G) was acquired on a HCl-etched sample. (H,I) were acquired on a polished thin
865 section.

866



867

868 **Figure 10.** Free and carbonate-bound lipid components in Crystal Geyser pisoids and travertines

869 (TL and TB).

870

871

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Supplementary data for:

Carbonate polymorphism controlled by microbial iron redox dynamics at a natural CO₂ leakage site (Crystal Geysers, Utah)

Julie Cosmidis, Shane O'Reilly, Eric T. Ellison, Katherine L. Crispin, David Diercks and Alexis S. Templeton

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Supplementary Table

Table S1. Free and carbonate-bound lipids in Crystal Geysers pisoids and travertines TL and TB.

Supplementary Text:

Detailed interpretation of lipid analyses

Supplementary Figures

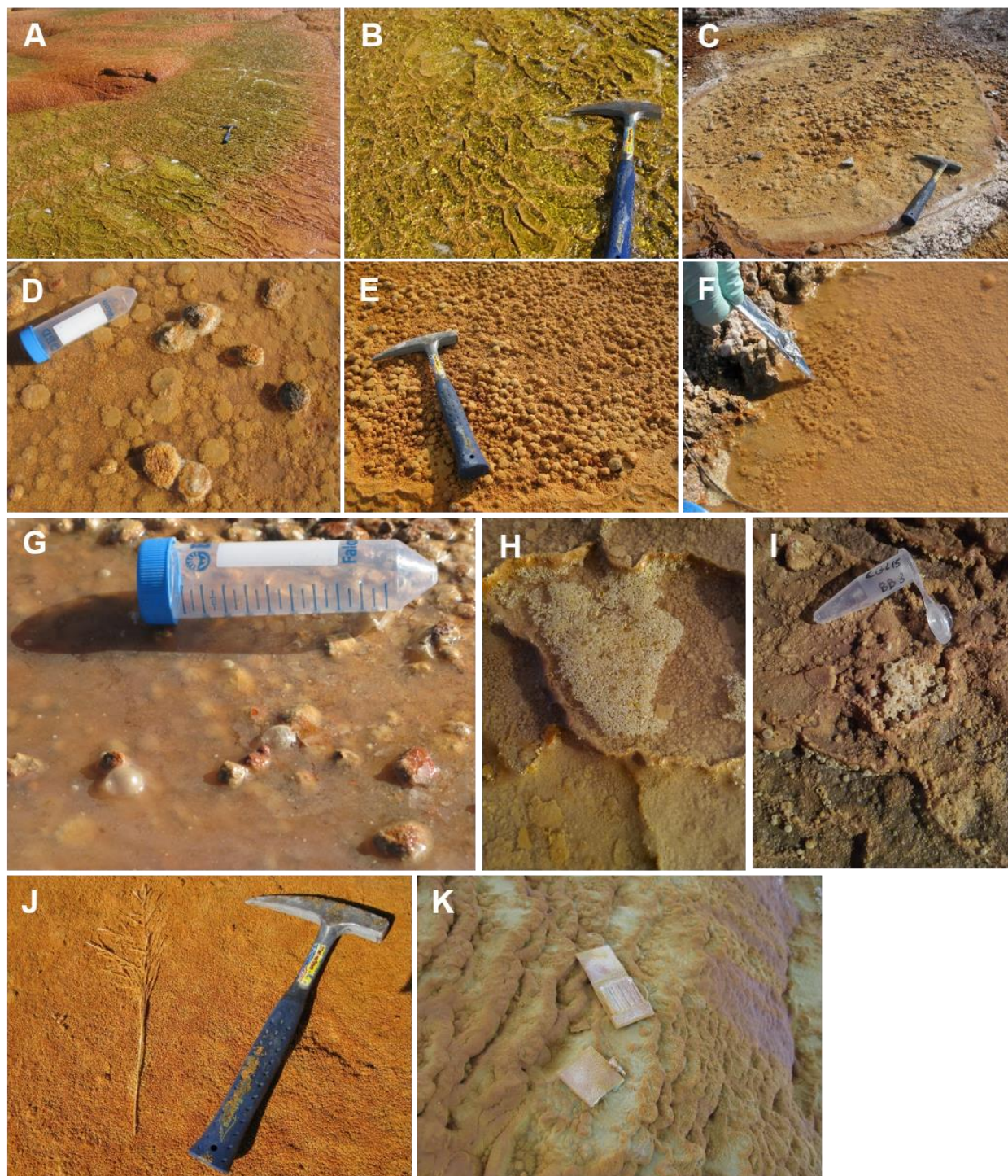


Figure S1. Additional pictures of the Crystal Geyser field site. (A,B) Green microbial mats at the surface of the travertine. (C-E) Cauliflower-shaped carbonate structures. (F) Carbonate toroids. (G-I) Calcified bubbles. (J,K) Plants and objects encrusted by rapidly precipitating carbonates.

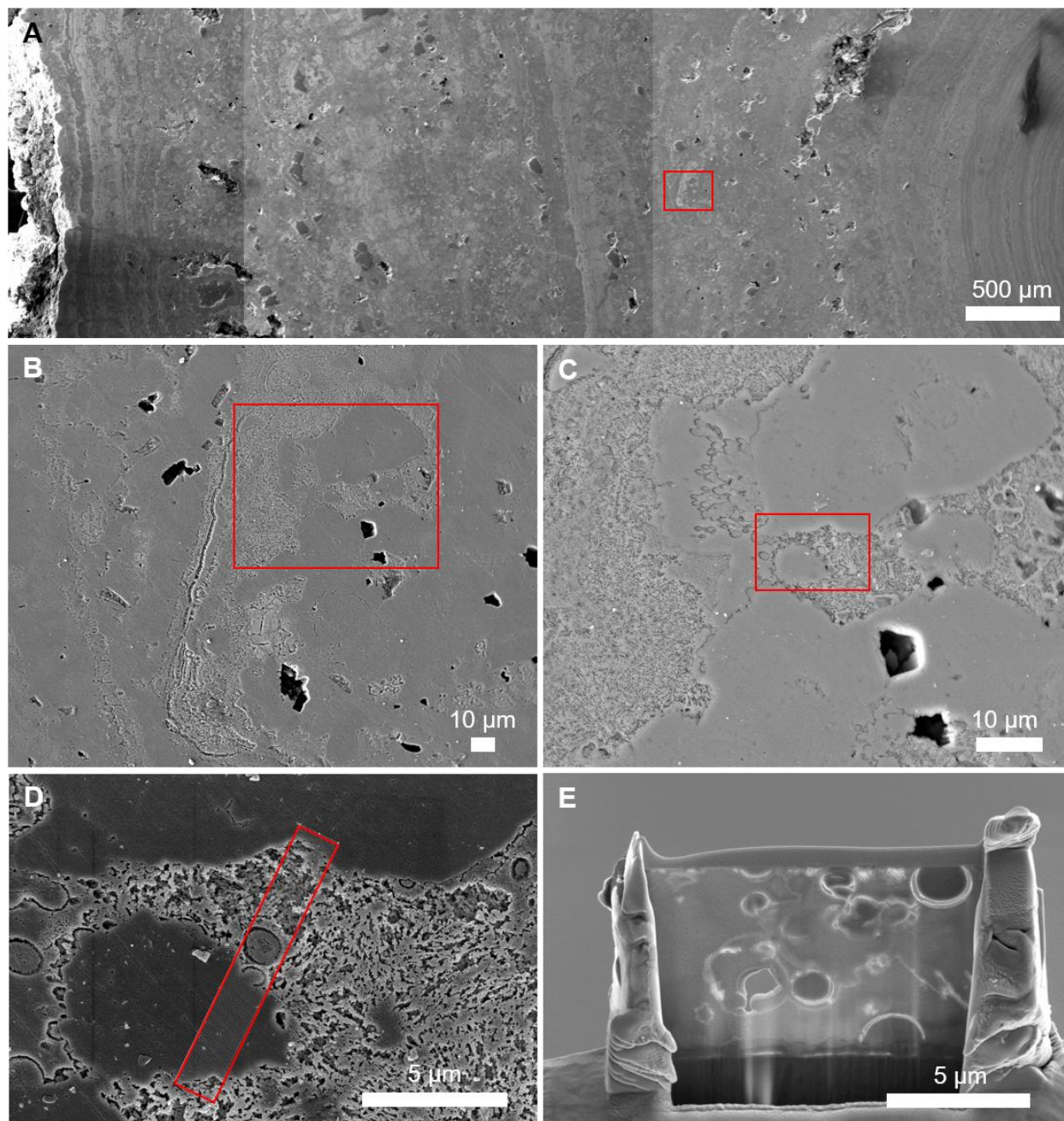


Figure S2. Location of FIB section 1. (A) Low-magnification SEM montage. The border of the pisoid is visible on the left side of the image. The red rectangles on (A), (B) and (C) show the locations of close-ups shown in the subsequent images. The rectangle in (D) indicates where FIB section 1 was milled. (E) SEM image of FIB section 1 after thinning.

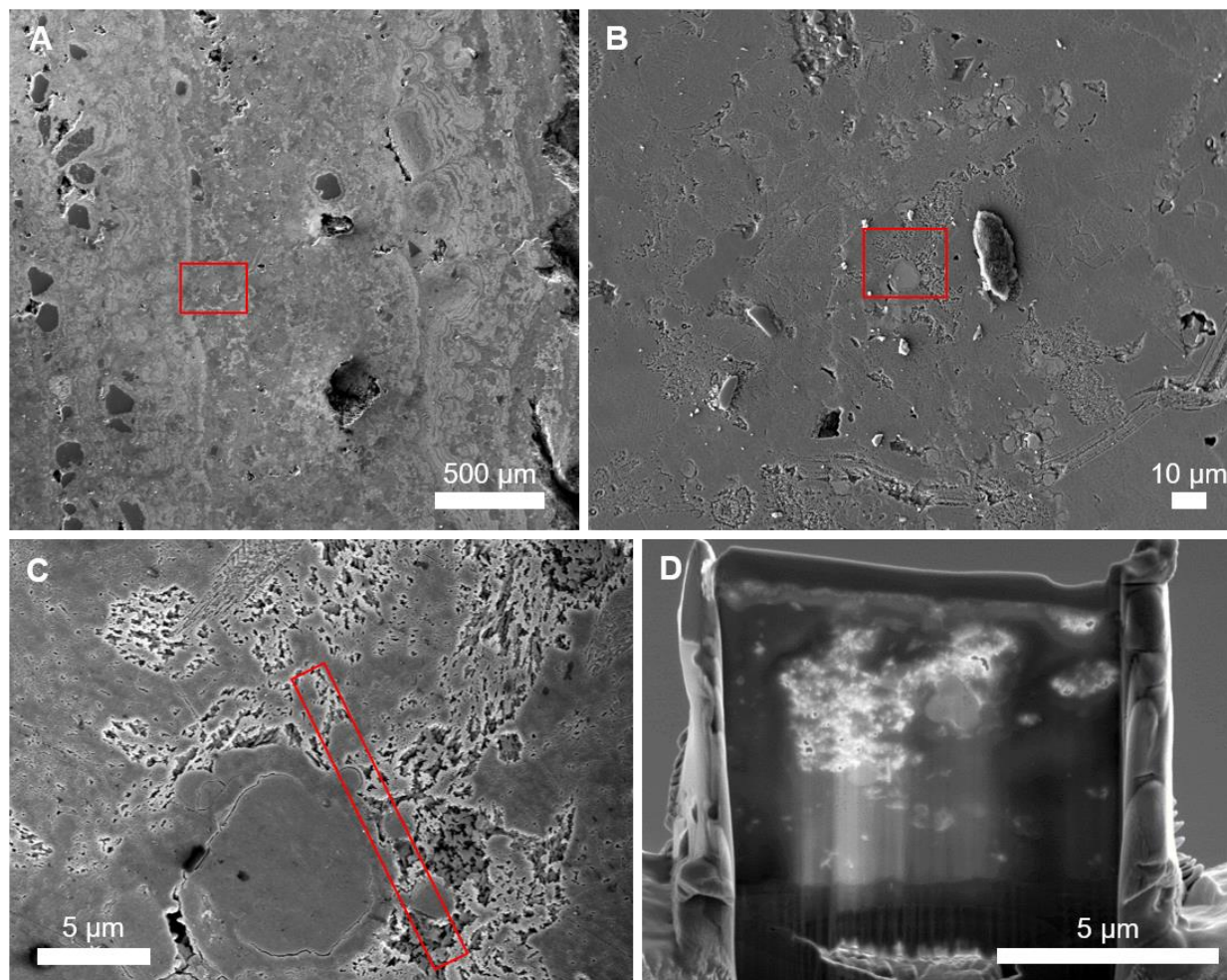


Figure S3. Location of FIB section 2. (A) Low-magnification SEM image. The border of the pisoid is visible on the right side of the image. The red rectangles in (A) and (B) correspond close-ups shown in (B) and (C), respectively. The rectangle in (C) indicates where FIB section 2 was milled. (D) SEM image of FIB section 2 after thinning.

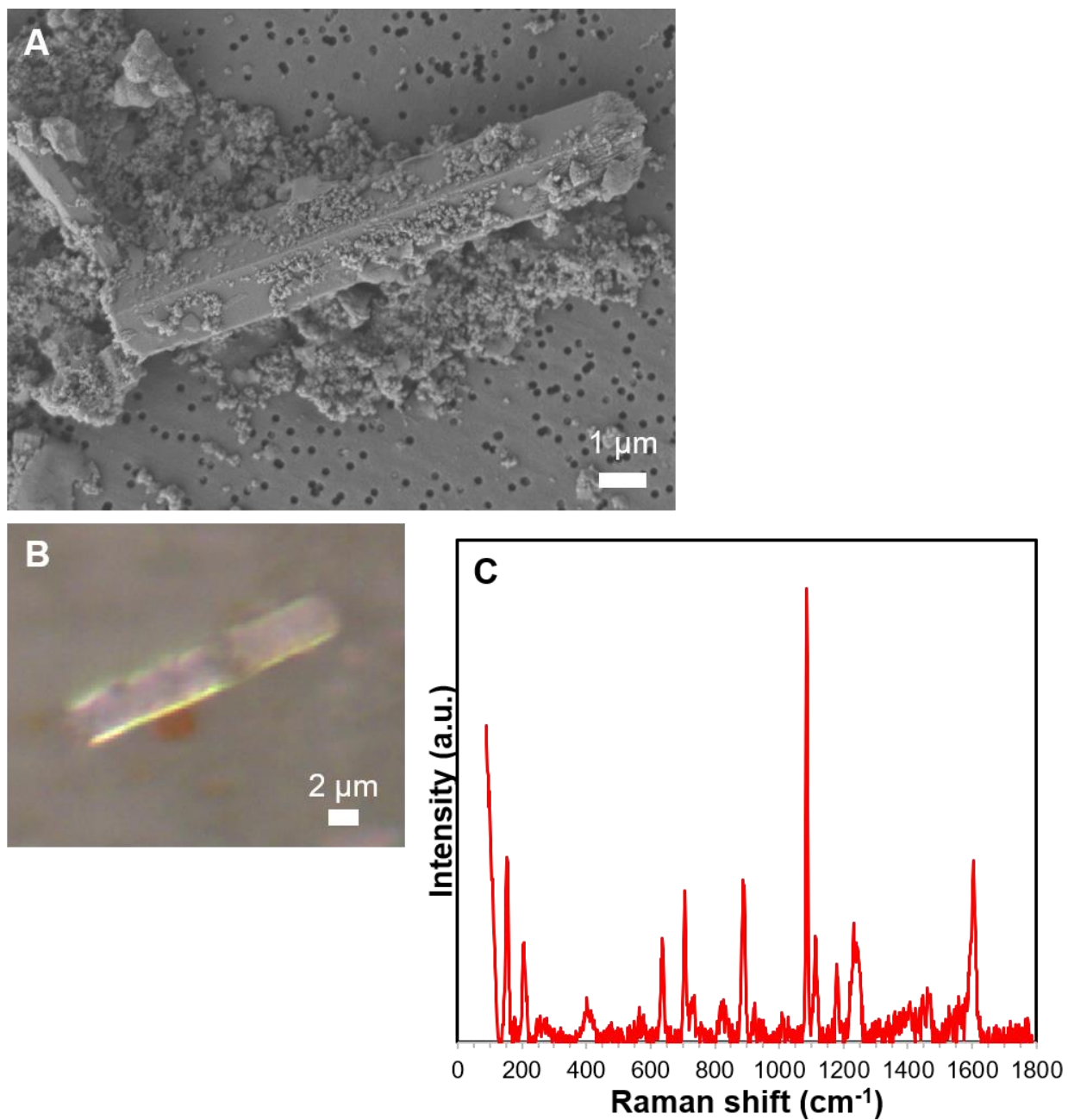


Figure S4. Aragonite particles samples from Crystal Geysers vent. (A) SEM image. Fine-grained iron minerals are also visible on and around the aragonite blade. (B) Light micrograph. (C) Raman spectrum (corresponding to aragonite).

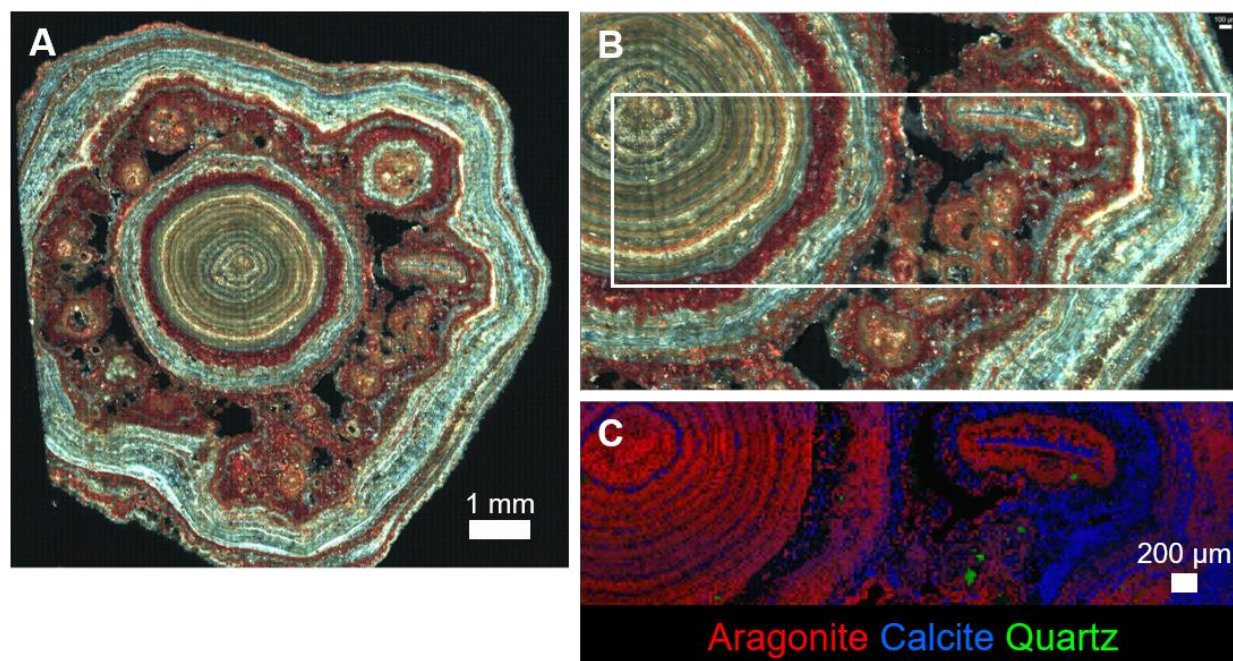


Figure S5. Micrographs and Raman map of a complex pisoid. (A) Montage of cross-polarized light micrographs. (B) Close-up showing the location of the Raman map (white rectangle). (C) Raman map showing the distribution of aragonite (red), calcite (blue) and quartz (green).

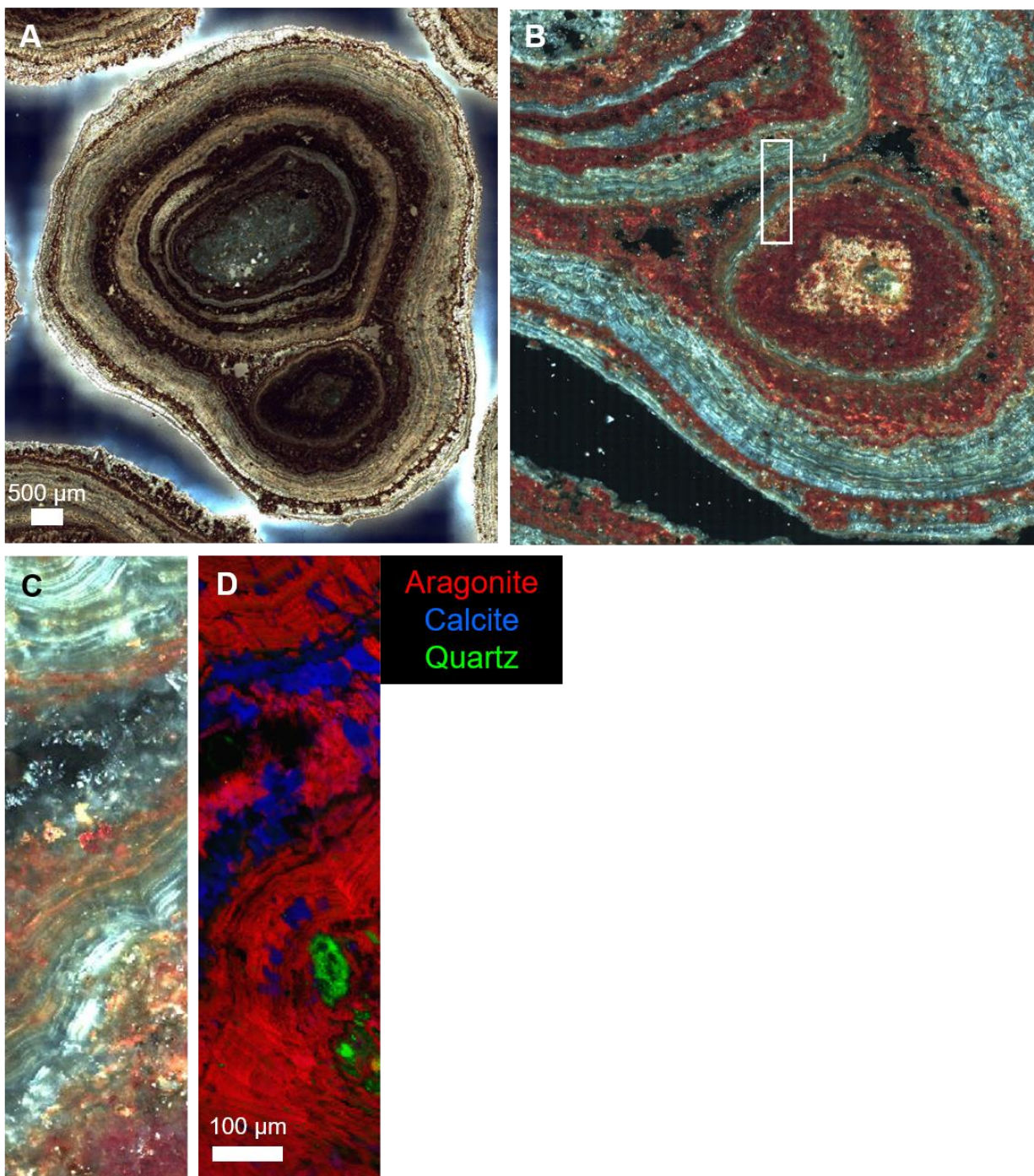


Figure S6. Micrographs and Raman map of a pisoid. (A) Montage of cross-polarized light micrographs. (B) Close-up in cross-polarized light showing the location of the Raman map (white rectangle). (C) Cross-polarized micrograph corresponding to the Raman map location. (D) Raman map showing the distribution of aragonite (red), calcite (blue) and quartz (green).

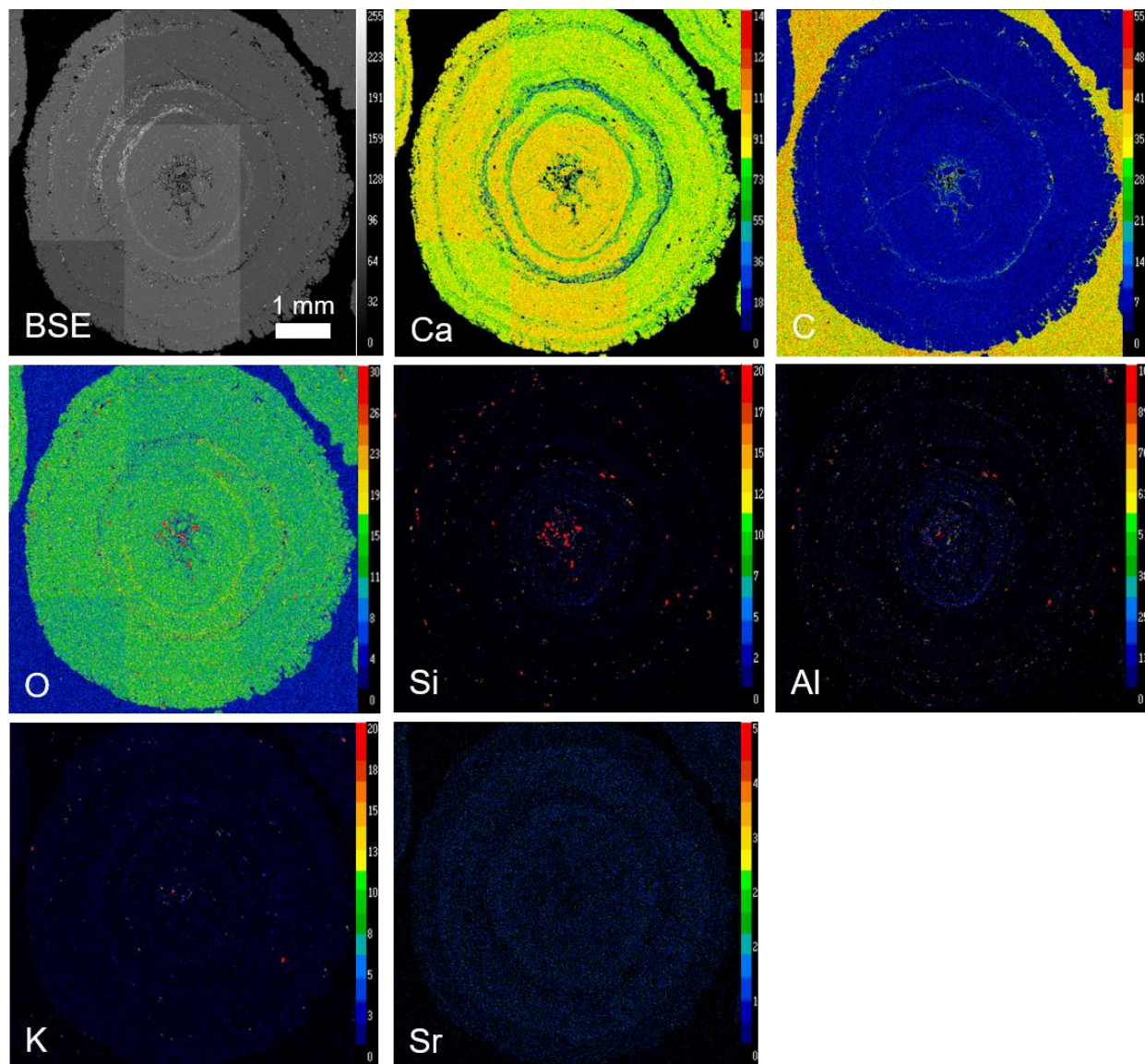


Figure S7. Montages of backscattered electron (BSE) images and additional EPMA maps of a pisoid. See also Figure 4C-E for maps of Fe, Mg and S. The color scales represent relative abundances of the different elements in the sample, but the number indicated on the scales are arbitrary (no calibration has been performed). High abundances of C and O in the background, as well as in porous areas of the pisoid, are due to the embedding epoxy resin used to prepare the thin section.

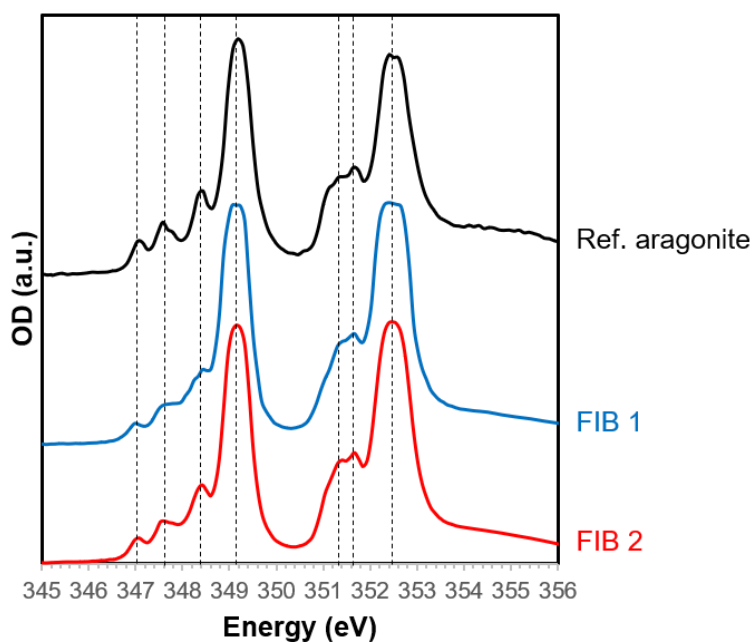


Figure S8. Ca L_{2,3}-edge XANES spectra acquired in STXM on the calcium carbonate phase of FIB sections 1 and 2. Dashed vertical lines show the positions of the main peaks and shoulders in the spectrum of a reference aragonite shown for comparison.

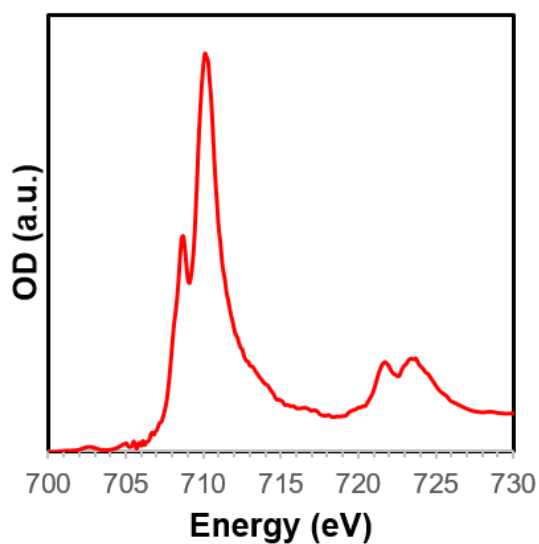


Figure S9. Fe L_{2,3}-edge XANES spectrum representative of Fe-rich areas in FIB sections 1 and 2, showing that Fe is present as Fe(III).

Supplementary Table

	Pisoid		Travertine TL		Travertine TB		Source
	Free	Bound	Free	Bound	Free	Bound	
Fatty acids (%)							
14		2.8		4.5	0.7	1.8	Mixed
i15		0.6		1.8	0.1	1.6	Bacteria
a15		0.2		0.7	0.1	1.3	Bacteria
15	3.2	0.6		0.5	0.2	0.6	Mixed
i16		0.2		0.6		0.7	Bacteria
16:1 ω 9		0.8		1.4	0.8	5.0	Mixed
16:1 ω 7		0.4					Mixed
16	69.3	52.1	9.4	42.4	7.5	40.0	Mixed
me-16							Bacteria
i17		0.2		0.7	0.1	0.5	Bacteria
a17		0.2		0.5	0.1	0.3	Bacteria
17	1.1	0.5		0.6	0.2	0.6	Bacteria
phytanic		0.2		0.2			Chlorophyll degradation product
18:1 ω 11c					2.6		
18:1 ω 9c		0.8	43.6	2.4	34.7		Green algae, other microalgae
18:1 ω 7c					17.7		Bacteria
18:1 ω 9t		1.2	23.7	3.1	13.9	1.2	
18:1 ω 7t					1.5		
18	13.5	30.7	5.7	18.1	7.4	26.9	Mixed
19					0.1		Mixed
20:1 ω 9c					7.1		Microeukaryotes
20:1 ω 9t					2.5		Microeukaryotes
20	3.7	0.8	1.7	2.0	0.7	1.3	
21		0.2		0.8	0.1	0.5	
22:1 ω 9c			2.1		0.8		Microeukaryotes
22	5.2	0.8	4.5	4.5	0.2	1.4	Algae, higher plants
23		0.2	0.3	0.6		0.6	Algae, higher plants
24	10.8	3.7	5.5	6.8	0.7	12.1	Algae, higher plants
25		0.3	0.3	0.5		0.4	Algae, higher plants
26	2.8	0.7	2.4	4.0	0.2	2.5	Higher plants, algae
27		0.2		0.3		0.2	Higher plants, algae
28	1.2	0.7	0.8	1.7		0.4	Higher plants, algae
29		0.1		0.2			Higher plants, algae
30		0.8		1.2			Higher plants, algae
Σ (ng g ⁻¹)	375	1525	887	1210	1657	383	
(Algae+Plant)/Bacteria	5.5	4.0		4.6	2.5	3.4	
3-hydroxy acids							

12		5		3		17	Bacteria (possibly anaerobic)
14		22		20		17	Bacteria (possibly anaerobic)
i15		7		10		8	Bacteria (possibly anaerobic)
a15		2		3			Bacteria (possibly anaerobic)
15		1		1			Bacteria (possibly anaerobic)
i16	18	8		1		7	Bacteria (possibly anaerobic)
a16				5			Bacteria (possibly anaerobic)
16	20	23		17		16	Bacteria (possibly anaerobic)
i17	39	23		21	43	18	Bacteria (possibly anaerobic)
a17	23	6		6	14	7	Bacteria (possibly anaerobic)
17		1		1			Bacteria (possibly anaerobic)
i18				1			Bacteria (possibly anaerobic)
18		7		14	42	9	Bacteria (possibly anaerobic)
Σ (ng g ⁻¹)	52	25	<i>n.d.</i>	104	15	16	
Monoalkylglycerol ethers (MGM)							
i15				15			Bacteria (likely aquatic, anaerobic)
a15				1			Bacteria (likely aquatic, anaerobic)
15				2			Bacteria (likely aquatic, anaerobic)
i16				3			Bacteria (likely aquatic, anaerobic)
16:1				3			Bacteria (likely aquatic, anaerobic)
16				24	100		Bacteria (likely aquatic, anaerobic)
me-16				13			Bacteria (likely aquatic, anaerobic)
me-16				10			Bacteria (likely aquatic, anaerobic)
i17				2			Bacteria (likely aquatic, anaerobic)
a17				3			Bacteria (likely aquatic, anaerobic)
17				2			Bacteria (likely aquatic, anaerobic)
me-17				2			Bacteria (likely aquatic, anaerobic)
i18							Bacteria (likely aquatic, anaerobic)
a18							Bacteria (likely aquatic, anaerobic)
18:1				2			Bacteria (likely aquatic, anaerobic)
18				7			Bacteria (likely aquatic, anaerobic)
me-18							Bacteria (likely aquatic, anaerobic)
i19				7			Bacteria (likely aquatic, anaerobic)
19:1				3			Bacteria (likely aquatic, anaerobic)
19							Bacteria (likely aquatic, anaerobic)
Σ (ng g ⁻¹)				19	2		
n-alkanols (%)							
16				1.6	13.3		Microalgae
17					0.4		Microalgae
18	1.4			2.4	21.4	76.8	Microalgae
20	0.6	4.1		0.7	1.7		Microalgae
21							Microalgae
22	1.1	3.1	0.5	0.9	1.9	23.2	Microalgae

24	1.8	6.7	1.0	1.5	3.0		Microalgae, higher plants
26	2.4	10.1	1.1	3.1	1.8		Microalgae, higher plants
Nonacosan-12-ol	24.2		22.9	11.3	40.5		Higher plants (related to <i>Tamarix</i> sp.)
28	2.6	7.7	1.3	1.8	1.3		Higher plants
Hentriacontan-12-ol	63.4	61.9	73.1	75.7	14.2		Higher plants (related to <i>Tamarix</i> sp.)
30	2.4	6.5		1.1	0.5		Higher plants
Σ (ng g ⁻¹)	91	<1	130	10	19	<1	
Higher Plant/Algae	31	12	185	17	2	0	
Ratio Alkan-12-ols	2.6		3.2	6.7	0.4		Different plant input, or diagenetic stability of C31 vs C29
Sterols (%)							
lanosterol					2.3		
cholesterol	36.2	33.6	12.0	5.4	6.8		Microeukaryotes, microalgae
cholestanol	3.2		5.2	8.7	6.9		degradation product of cholesterol
campesterol	3.6		9.1	6.5	1.1		Algae (diatoms, green algae), higher plants
stigmasterol	24.0	33.4	58.4	61.8	68.0	100.0	Algae (diatoms, green algae), higher plants
22-ethyl-coprostenol					11.5		degradation product of stigmasterol?
β -sitosterol	30.3	33.0	12.9	13.3	4.4		Algae (diatoms, green algae), higher plants
sitostanol	2.8		2.4	4.4	1.3		degradation product of β -sitosterol
Σ (ng g ⁻¹)	737	16	468	73	242	19	
stanol/stenol	0.1	0.0	0.3	0.7	0.7		more reducing environmental, higher anaerobic activity?
cholest/stigmast	1.5	1.0	0.2	0.1	0.1		greater eukaryote heterotrophic activity
stigmasterol/sitosterol	0.8	1.0	4.5	4.7	15.5		differences in photosynthetic community
Hydrocarbons							
2,2-dime-C16	0	0	0	0	32	0	Cyanobacteria?
23	6	0	0	0	13	0	Algae, higher plants
24	4	4	0	5	6	0	Algae, higher plants
25	9	8	10	9	9	17	Higher plants, algae
26	6	13	5	7	3	0	Higher plants
27	12	16	22	14	8	31	Higher plants
28	7	15	6	8	2	0	Higher plants
29	25	18	32	31	15	36	Higher plants
30	7	11	5	7	6	0	Higher plants
31	19	10	20	16	6	16	Higher plants
32	5	5	0	3	1	0	Higher plants
Σ (ng g ⁻¹)	298	38	141	17	51	6	
Higher Plant/Algae	8	21		17	1		
Other (ng g⁻¹)							
Archaeol	294						Archaea
Total							
Σ lipids (ng g ⁻¹)	1847	1605	1627	1432	1988	424	

Table S1. Free and carbonate-bound lipids in Crystal Geyser pisoids and travertines.

Supplementary Text: Detailed interpretation of lipid analyses

Fatty acids, ranged from 14 to 30 carbon chain lengths, were a major lipid component in all samples. Monounsaturated and methyl-branched (typically *iso* and *anteiso* chain positions) were present as well as *n*-saturated fatty acids. The highest abundances of fatty acids were observed in the bound lipid pools of the pisoids and travertine TL. Some of these fatty acids (especially *n*-saturated and monounsaturated) are common to both bacteria and eukaryote domains, so there is some uncertainty in source assignment. Methyl-branched fatty acids are primarily sourced from bacteria (Kaneda, 1991). 3-hydroxy acids were also identified in relatively minor amounts in all samples apart from the free lipid pool of travertine TL. Carbon chain lengths ranged from 12 to 18, and methyl-branched isomers were a major component, particularly *iso*- and *anteiso*-heptadecanoic acid. These are derived from gram negative bacteria, as components of lipopolysaccharides, and are often associated with anaerobic heterotrophic bacteria (Rietschel, 1976; Wang *et al.*, 2016).

Monoalkyl glycerol monoethers (MGM) were present in minor amounts in the bound lipid fraction of travertine TL and in the free lipid fraction of travertine TB. These are bacterial lipids, and based on current evidence are particularly common in aquatic extremophiles and heterotrophic mesophiles engaged in sulfur cycling (particularly sulfate reduction) (Rütters *et al.*, 2001; Hernandez-Sanchez *et al.*, 2014). MGM are rarely reported in non-aquatic settings, and a large contribution from methyl-branched members is indicative of sedimentary (probably sub-oxic/anoxic) bacteria rather than aerobic aquatic bacteria. The presence of MGM in the travertines and their absence in the pisoids suggest that the microbial communities present during precipitation of pisoids and travertine minerals are different.

n-alkan-1-ols between 12 and up to 24 carbon chain lengths are majorly associated with photosynthetic microalgae, while *n*-alkan-1-ols between about 24 and 32 carbon chain lengths are majorly associated with higher plants (Pancost & Boot, 2004; Volkman, 2006). Long chain *n*-alkan-12-ols are rarely reported and appear to be quite restricted to cuticular wax lipids from *Tamarix*-type species (Basas-Jaumandreu *et al.*, 2014), which are abundant in the region of Crystal Geyser. As such these likely reflect aeolian-sourced inputs. Minor amounts of alkan-12-ols were identified in all free and bound lipid fractions, although the abundance in the bound lipid fractions was generally over an order of magnitude lower. Interestingly the ratio of

hentriacontan-12-ol to nonacosan-12-ol was quite varied (0.4 to 6.7, Table S1). Assuming, similar degradation rates, this indicates distinct differences in the input of different vascular plant species and/or pathways between samples.

Sterols were abundant in all free lipid fractions, and are diagnostic for higher plants and microalgae (Volkman *et al.*, 1998; Volkman, 2006). These likely reflect aeolian input from regional vascular plants and photosynthetic microalgal biomass. Cholesterol is the sole sterol in animals and heterotrophic microeukaryotes, as well as a relatively minor sterol in certain photosynthetic microalgae. The presence of cholesterol amounts similar to C₂₉ sterols likely reflects input from heterotrophic microeukaryotes such as protists. Cholesterol was a major sterol in pisoids samples and likely reflect aquatic protists. High concentrations of stanols relative to sterols is generally indicative of reducing environments and significantly anaerobic bacterial hydrogenation reactions (Wakeham, 1989). Cholestanol/cholesterol ratios for the pisoids are <0.1, 0.5 for TB FL fraction, and > 1.0 for TL BL and TB FL. Thus, while the absence of sterols in bound lipid fractions may be related to low original concentrations, it may also reflect microbial activity under reducing conditions during mineral precipitation.

Hydrocarbons were identified in the free lipid fractions of all samples, and were dominated by long chain odd carbon number *n*-alkanes, particularly heptacosane, nonacosane and hentricontane. This is strong evidence of vascular plant wax lipid input (Eglinton & Hamilton, 1967; Bush & McInerney, 2013). One methyl-branched alkane was found in the free lipid pool of travertine TB. This is tentatively identified as 2,2-dimethyl-hexadecane and indicates the presence of cyanobacteria (Gomes *et al.*, 2020). Cyanobacterial hydrocarbons were not identified in any other sample, suggesting they play a minor role at Crystal Geyser, apart from a possibly role in mineral precipitation in travertine TB.

Archaeol (di-O-phytanylglycerol) is a membrane lipid found in certain archaea (Kate, 1993). It was found in relatively high abundances in the pisoid free lipid fraction. This most likely reflects aquatic, possibly extremophilic, archaea living in the water close to close to the geyser. The fact that it was not detected in the bound lipid fraction of the pisoids indicates that these archaea are associated with the aqueous phase and as detritus on mineral surfaces but are not closely associated with mineral precipitation in the Crystal Geyser pisoids.

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