

1 **A convolutional neural network approach to estimate earthquake kinematic parameters from back-**
2 **projection images**

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24 **Abstract**

25

26 The retrieval of earthquake finite-fault kinematic parameters after the occurrence of an earthquake is a crucial
27 task in observational seismology. Routinely-used source inversion techniques are challenged by limited data
28 coverage and computational effort, and are subject to a variety of assumptions and constraints that restrict the
29 range of possible solutions. Back-projection (BP) imaging techniques do not need prior knowledge of the
30 rupture extent and propagation, and can track the high-frequency (HF) radiation emitted during the rupture
31 process. While classic source inversion methods work at lower frequencies and return an image of the slip over
32 the fault, the BP method underlines fault areas radiating HF seismic energy. HF radiation is attributed to the
33 spatial and temporal complexity of the rupture process (e.g., slip heterogeneities, changes in rupture speed and
34 in slip velocity). However, the quantitative link between the BP image of an earthquake and its rupture
35 kinematics remains unclear. Our work aims at reducing the gap between the theoretical studies on the
36 generation of HF radiation due to earthquake complexity and the observation of HF emissions in BP images.
37 To do so, we proceed in two stages, in each case analyzing synthetic rupture scenarios where the rupture
38 process is fully known. We first investigate the influence that spatial heterogeneities in slip and rupture velocity
39 have on the rupture process and its radiated wave field using the BP technique. We simulate different rupture
40 processes using a 1D line source model. For each rupture model, we calculate synthetic seismograms at three
41 teleseismic arrays and apply the BP technique to identify the sources of HF radiation. This procedure allows
42 us to compare the BP images with the causative rupture, and thus to interpret HF emissions in terms of along-
43 fault variation of the three kinematic parameters controlling the synthetic model: rise time, final slip, rupture
44 velocity. Our results show that the HF peaks retrieved from BP analysis are better associated with space-time
45 heterogeneities of slip acceleration. We then build on these findings by testing whether one can retrieve the
46 kinematic rupture parameters along the fault using information from the BP image alone. We apply a machine
47 learning, convolutional neural network (CNN) approach to the BP images of a large set of simulated 1D rupture
48 processes to assess the ability of the network to retrieve from the progression of HF emissions in space and
49 time the kinematic parameters of the rupture. These rupture simulations include along-strike heterogeneities
50 whose size is variable and within which the parameters of rise-time, final slip, and rupture velocity change
51 from the surrounding rupture. We show that the CNN trained on 40,000 pairs of BP images and kinematic
52 parameters returns excellent predictions of the rise time and the rupture velocity along the fault, as well as

53 good predictions of the central location and length of the heterogeneous segment. Our results also show that
54 the network is insensitive towards the final slip value, as expected from a theoretical standpoint.

55

56 **Keywords:** Computational seismology, Earthquake source observations, Neural networks, Numerical
57 modelling

58

59 **1. Introduction**

60

61 Characterizing earthquake finite-fault parameters, such as the rupture extent, rupture velocity, and the spatio-
62 temporal distribution of the slip along the fault, are all fundamental to achieving a better understanding of
63 earthquake dynamics. Earthquake parameters are often estimated through slip inversion techniques that rely
64 on a-priori assumptions on the fault geometry and the rupture mechanism. Such techniques have succeeded to
65 image the finite-fault parameters in a relatively low-frequency range ($f < 2.0 - 3.0$ Hz) (e.g., Zeng et al. 1993;
66 Mai et al, 2016). However, if we are interested in uncovering fine-scale, detailed structure in the rupture
67 process, high-frequency (HF) waveform data must be taken into account (e.g. Mendoza and Hartzell 1988, Ide
68 1999).

69

70 The complexity of the source manifests itself in terms of short-period seismic waves, at frequencies higher
71 than the corner frequency f_c (e.g. Madariaga 1977, Spudich and Frazer 1984, Ruiz et al. 2011) controlled by
72 the source duration. At these higher frequencies, the classic inversion techniques are no longer adequate, both
73 because of computational limitations and by our lack of knowledge of the Earth's structure at those frequencies.
74 Hence, high-resolution imaging techniques, such as the back-projection (BP, Ishii et al. 2005), have recently
75 become a prominent means to unraveling aspects of the rupture complementary to the ones supplied by classic
76 inversion. When applied to teleseismic body-waves, the BP method takes advantage of the coherence of HF
77 waveforms among traces recorded at nearby stations. This technique requires minor a priori constraints and
78 bypasses the procedure of inverting for Earth structure (e.g., Kiser and Ishii 2017). BP has become a very
79 popular method and several observational studies in recent years have highlighted the ability of such a
80 technique to illuminate the HF emission sources excited during earthquake rupturing (e.g., Walker and Shearer
81 2009, Xu et al. 2009, Zhang and Ge 2010, Meng et al. 2011, Koper et al. 2011, Lay et al. 2012, Satriano et al.
82 2014, Vallée and Satriano 2014, Grandin et al. 2015). From a theoretical perspective, studies have long
83 attributed the generation of HF radiation during earthquake faulting to small-scale variations or roughness in
84 final slip, slip velocity, or rupture velocity —acceleration and deceleration— over the fault plane (Madariaga
85 1977, Andrews 1981, Herrero and Bernard 1994, Somerville et al. 1999, Ruiz et al. 2011). HF radiation
86 enlightened by BP images of large earthquakes could, therefore, help constrain the variability of parameters
87 controlling the rupture process. However, unanimous consent on how the BP image relates to earthquake

88 parameters has not yet been achieved. Specifically, Ishii et al. (2005) suggested that the BP image of the
89 earthquake is related to the radiated seismic energy. Both works from Yao et al. (2012) and Fukahata et al.
90 (2014) point out an intrinsic ambiguity in BP image towards resolving slip velocity or slip acceleration.

91

92 Our work has, therefore, a dual purpose. First, we aim to investigate the link between coherent images of the
93 rupture process and the mechanism of high-frequency (HF) generation and, therefore, our understanding of
94 what the BP image truly represents. Second, we work toward extracting information on the earthquake
95 kinematic properties from the HF radiation enhanced by the BP analysis of earthquakes. To accomplish the
96 objectives of our study, we take a data-driven approach, where for a Mw 7.5 strike-slip event we simulated
97 40,000 rupture scenarios with different characterizing traits. We parameterize the system as a 1D line source
98 model (Lancieri and Zollo 2009). This choice allows us to ease the computational effort and describe the
99 complex earthquake rupture process in a simplified manner that seeks to represent its key elements by three
100 kinematic parameters: the rise time, the final slip, and the rupture velocity. We apply the back-projection
101 technique and compare the resulting images with the originating rupture model. Doing so enables us to
102 understand not only the link between the generation of HF radiation and the complexity of the slip rate function,
103 but also how different arrays can enlighten different aspects of the same rupture process. Using a 1D source
104 allows us to represent the BP image of the rupture as a 2D map where we can easily follow the HF progression
105 in space and time, drawing the attention to the mechanism behind the generation of high-frequency radiation
106 from spatial heterogeneities in the kinematic rupture parameters.

107

108 After developing a conceptual understanding of the relation between the BP image and the variability of the
109 parameters controlling the rupture process, we move on to adopting a convolutional neural network (CNN)
110 approach to look for the statistical link between the BP images and the rupture kinematic parameters, exploring
111 the role of input and target parameters on the accuracy of the CNN predictions. In particular, we aim to assess
112 which kinematic characteristics of the rupture process, as well as the location and the extent of a spatial
113 heterogeneity —when present— can be reliably determined from the BP image of the seismic event.

114 **2. Methodology**

115

116 **2.1 *The line source model***

117 The line source is an intermediate approach between the point source and the extended source model (Lancieri
118 and Zollo 2009). Its advantage lies in the ability to reproduce the typical 2D-source directivity effects on the
119 signal while requiring a lower computational effort. Use of this model is most appropriate for earthquakes with
120 aspect ratios (fault length divided by width) well in excess of one, such as for large strike slip earthquakes and
121 great subduction zone earthquakes. Our source model is built by placing a series of equally spaced point
122 sources along a line (with total length equal to the fault length). These points are set at the hypocentral depth
123 of the event and are distributed along the strike direction of the fault. The point sources begin to slip at different
124 subsequent activation times, related to the rupture propagation velocity along the line. Once crossed by the
125 rupture front, each point source slips following a ramp function whose duration is the rise time. We describe
126 the line source model with the following parameters:

127

- 128 • the hypocentral coordinates of the seismic event;
- 129 • the strike direction of the fault;
- 130 • the rupture length L ;
- 131 • the rupture velocity V_r along the line;
- 132 • the discretization along the line ΔL ;
- 133 • the rise time t_r , or the duration of the dislocation, for each point source;
- 134 • the final slip value s_f , reached by each point source at the end of the dislocation.

135

136 Under the line source approximation, each point has the same focal mechanism (slip vector and fault
137 orientation) and the total seismic moment of the event is given by the sum of the seismic moments of each
138 point source. An estimation of the fault length (L) and width (W) is derived from the event magnitude using,
139 e.g., the Wells and Coppersmith (1994) empirical relationships. To avoid space-time aliasing, the contributions
140 emitted by each elementary source must overlap in time; i.e., the duration of each elementary source (the rise
141 time) must be greater than the time the rupture takes to propagate to the next source (here called τ), as discussed

142 in Lancieri and Zollo (2009). To properly set the point source spacing along the line, different simulations were
 143 performed, leading to the optimal sampling assigned in **table 1**.

144

Table 1: Hypocentral coordinates and focal mechanism of the $M_w = 7.8$, 2001 Kunlun earthquake, used as a frame for our synthetic study. Rupture length L and width W derived from Wells and Coppersmith empirical relations. Space ΔL and time Δt sampling chosen in order to avoid space-time aliasing, as discussed by Lancieri and Zollo (2009). P - and S - wave velocities and quality factor used for the Green's functions.

Hypocentral Coordinates (°)	Focal Mechanism (°)
Lon = 90.59, Lat = 35.93	Strike = 78, Dip = 61, Rake = -12
Hypocentral depth (km)	Rupture Geometry (km)
$Z = 13$	$L = 100, W = 15$
P-wave velocity (km/s)	Space sampling (km)
$\alpha = 5.8$	$\Delta L = 0.4$ km
S-wave velocity (km/s)	Time sampling (s)
$\beta = 3.46$	$\Delta t = 0.02$ s
Attenuation	
$Q = 730$	

145

146

147 In the present study, we locate the earthquake in Northwestern China, and we use the geometry of the $M_w =$
 148 7.5, 2001 Kunlun earthquake (Klinger et al. 2005, Vallée et al. 2008). The generated source models do not aim
 149 at reproducing the exact rupture process of the Kunlun event, but they rather explore different rupture
 150 scenarios. However, our synthetic model does share the same hypocentral coordinates, seismic moment, focal
 151 mechanism and rupture propagation direction of the $M_w 7.5$, Kunlun earthquake, as summarized in **table 1**.
 152 Hypocentral coordinates and focal mechanism are obtained from the Global CMT Moment Tensor Solution
 153 (Dziewonski et al, 1981). We simulate 40,000 rupture scenarios of the $M_w 7.5$ strike-slip event adopting the
 154 1D line source model with large range of complexity. To capture a range of different scenarios, we vary the
 155 kinematic parameters (rise time, final slip and rupture velocity) for each simulation. Each kinematic parameter
 156 is randomly selected within a range of values shown in **table 2** according to the empirical relationships
 157 proposed by Wells and Coppersmith (1994) and Geller (1976). In addition, we introduce a heterogeneous
 158 segment whose length L_H ranges from one space sample, that is 400 m, to 40 km. Within the heterogeneous

159 segment, the rise time, the final slip and the rupture velocity assume a different value from the surrounding
160 length.

161

Table 2: Ranges of kinematic parameters for rupture process simulation. We call t_r the rise time, s_f the final slip, V_r the rupture velocity, L_H the length of the heterogeneous segment and x_c its central position. Kinematic parameters are uniformly distributed within the range listed below.

tr (s)	Sf (m)	V_r(m/s)	LH (km)	Xc (km)
1-8	1-10	2500-3460	0.4-40	20-80

162

163

164 **2.2 Modeling body waves at teleseismic distance**

165 For each rupture scenario generated in the previous section, we compute synthetic seismograms at three arrays
166 of seismic stations at teleseismic distances (between 50 and 90 degrees): Alaska (AK), Europe (EU), and
167 Australia (AU). Each array is comparable to an antenna that tracks in space and time the strongest coherent
168 sources of HF seismic energy. The arrays are located at three complementary locations with respect to the
169 epicenter of the event, allowing us to obtain three different viewpoints of the same rupture process. Each array
170 is composed of 55 stations (black triangles in **Fig.1**). The interest of working at teleseismic distances lies in
171 the possibility of investigating high-frequency emissions using only the far-field term of the Green's function
172 in the representation theorem (Aki and Richards, 1980). This approximation leads to many simplifications of
173 the calculations and simple physical models of the generation of high-frequency waves. In the current study,
174 the Green's functions are calculated in a spherically averaged Earth's structure AK135 (Kennett et al., 1995).
175 The ground motion associated with the displacement on the fault plane is computed by considering the
176 representation integral in the frequency domain. Here, in particular, teleseismic body waves are computed by
177 taking into account only the direct P arrival and associated depth phases, pP and sP. We compute the
178 displacement associated with a teleseismic P-wave in a geometrical ray solution following the formula of Okal
179 (1992). However, in our study, we neglect the geometrical spreading and the response of both the receiving
180 site and the recording instrument by normalizing the individual traces. To compute the reflection coefficients
181 for pP and sP phases into the synthetic seismograms, we use the calculations made by Aki and Richards (1980).
182 Values for P- and S- wave velocity are shown in **table 1**.

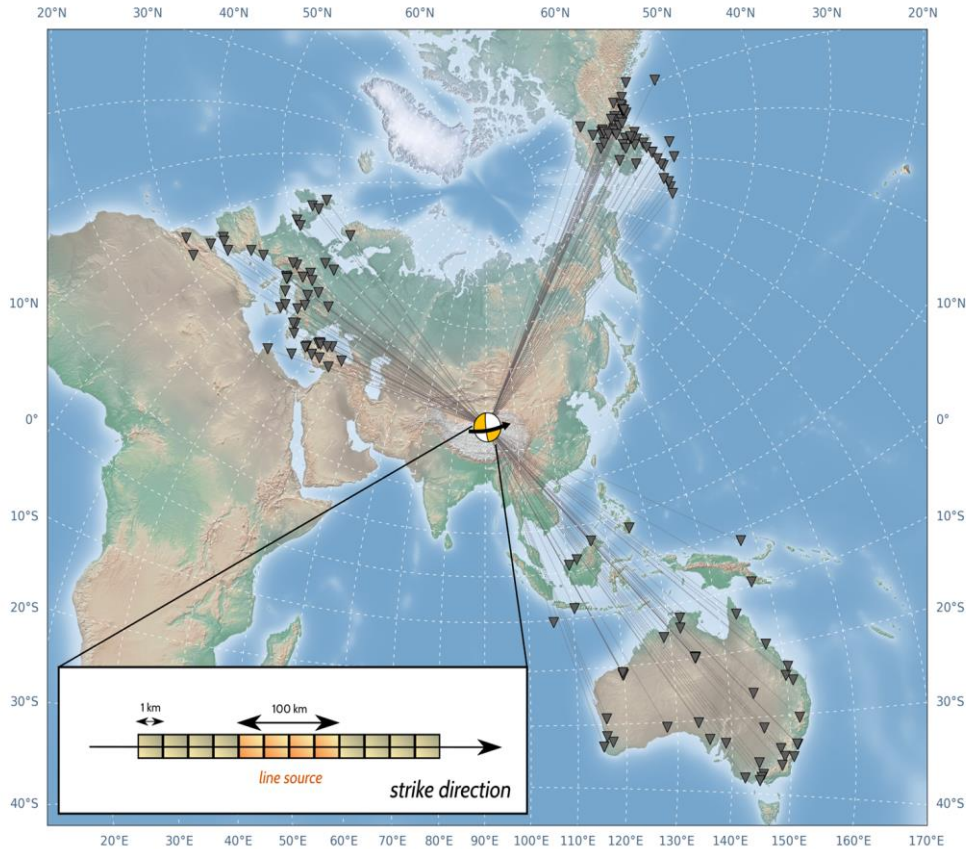


Figure 1: Geographical setting of the present study. Focal mechanism and epicentral location of the synthetic test are placed in Northwestern China. Alaskan (AK), Australian (AU) and European (EU) arrays are composed each of 55 seismic stations (black triangles). The bottom window shows a zoom over the 300 km long one-dimensional BP grid that surrounds the line source.

183

184 **2.3 Description of the synthetic data-set**

185 Our data-set is composed of 55 Z-component synthetic displacement traces for each array in **Fig.1**. The AK
 186 and AU arrays are located almost in the directive position respect to the rupture propagation, whereas the EU
 187 array is located almost in the antidirective position. To identify high-frequency pulses, the synthetic
 188 displacement traces are then differentiated in time and band-pass filtered between 0.5 and 4 Hz. The filter is a
 189 zero-phase 4-pole Butterworth, to obtain an acausal filtered signal where the arrival time of the signal's peak
 190 is respected.

191

192 **2.4 Back Projection imaging**

193 We follow the conventional back projection (BP) method for source imaging first proposed by Ishii et al.
 194 (2005), where the body waves arriving at an array of recording stations are projected back to a (generally 2D)

195 source grid under the ray-theory asymptotic approximation. To apply the back-projection method, we define
196 the coordinates of the source grid points and assume a velocity model to calculate theoretical travel times. We
197 use a 1D grid of potential source locations placed over the line source. In particular, the grid is 300 km long
198 with a grid-step of 1 km (**Fig.1**) and all grid points are placed at the hypocentral depth. This is a simplification
199 of the standard approach in BP analysis where a two-dimensional grid is used. The BP image is then constructed
200 through the following steps:

201

- 202 1. For each grid-point, theoretical P-wave travel times are calculated at each station of the array in the AK135
203 velocity model (Kennett 2005);
- 204 2. Filtered seismograms are reverse time shifted according to the values of travel times;
- 205 3. Shifted traces are stacked, returning a signal related to the grid point. If the shifted traces sum constructively,
206 the stack is high, meaning that the grid point is a plausible source of coherent HF radiation (from the array
207 point of view). If the stack is low, in contrast, the grid point likely did not contribute coherent HF energy during
208 the rupture process.

209

210 Several methods exist for aligning and stacking seismic traces, for instance, the Nth root stacking (Muirhead
211 1968) and the F-ratio (Melton and Bailey 1957). In our study, traces are combined via the shift-and-sum
212 approach, because this linear approach is the simplest technique for combining the station traces, it requires
213 the least a priori assumptions and it does not deform the amplitude of the signals and. Mathematically, the
214 linear stack $\mathbf{s}_i(\mathbf{t})$ at the i -th potential grid point can be expressed as follows:

215

$$216 \quad \mathbf{s}_i(\mathbf{t}) = \sum_j \mathbf{u}_j(\mathbf{t} - \mathbf{t}_{ij}^P). \quad (1)$$

217

218 Where \mathbf{u}_j is the velocity trace at the j -th station and \mathbf{t}_{ij}^P is the theoretical P-wave travel time from the i -th source
219 to the j -th station. In the calculation of the BP image, we replaced the stacked signal by its squared value, the
220 beam power. The beam can be highly noisy especially for HF filtered signals, where further peaks appear
221 because of the filter. To further reduce the noise, the beam is smoothed using a zero-phase Gaussian filter,
222 parametrized by its standard deviation σ . We tested different values of Gaussian smoothing windows (σ), and

223 chose the value of $\sigma = 0.4$ s, which allows us to only focus on the main energy bursts neglecting the smaller
224 peaks coming from the filtering and stacking procedures.

225

226 The resolution of BP images mostly depends on two factors: the frequency content of the data and the geometry
227 of the array. In general, the higher the frequency we are looking at, the more detailed the BP image will be
228 (Schweitzer et al. 2011). Likewise, both the aperture and the position of the array with respect to the rupture
229 direction control the resolution of the BP image. Specifically, good resolution in BP images is achieved when
230 using large-aperture arrays (Xu et al. 2009). Low-frequency ($f < 1$ Hz) P-waves are less affected by scattering
231 and smaller-scale heterogeneity, thus their stack generally returns a high coherency. However, they do not
232 provide a good degree of detail on BP images. On the contrary, HF ($f > 1$ Hz) P-waves provide a high resolution
233 on BP images, but at the same time their waveforms are more easily distorted by small-scale heterogeneity and
234 scattering (e.g., Frankel et al, 1986, Takemura et al., 2013), often leading to a less coherent stack. In addition,
235 similarity among traces breaks down as interstation distance increases (Xu et al. 2009). The frequency range
236 0.5-4 Hz is indeed typical in BP analysis at teleseismic distance, because it is a good balance between the
237 resolution of the BP image and the waveform coherence (Xu et al. 2009).

238

239

240 ***2.5 Reducing artifacts in the BP image***

241 In BP analysis there are several potential artifacts that one needs to be aware of. These include “smearing” and
242 the “walking” effects (e.g., Meng et al, 2012, Xu et al. 2009, Walker and Shearer 2009), and the presence of
243 HF emissions carried out by later arrivals, namely the depth phases (Yagi et al, 2012). The smearing effect
244 makes the beam power signal of the true HF emission source to be blurred in space, whereas the walking effect
245 leads to having a false energy radiation appearing before and after the true emission time. In addition, the
246 presence of HF emissions carried out by depth phases can obscure some of the first-order features of the rupture
247 process. Okuwaki et al. (2018) show that BP has a depth-dependent bias proportional to the amplitude of the
248 Green’s function. At shallow depths (e.g., focal depths less than 10 km) these artifacts are caused by non-
249 physical interactions of the back-projected P wave and depth phases. For deep earthquakes, on the contrary,
250 depth phases are separated enough from the direct P-wave and they can be used to constrain depth (Kiser et al,

251 2011). However, in this case they might still complicate the interpretation of the BP image, as further HF peaks
 252 appear in the image, but they are not generated by the complexities of the source.

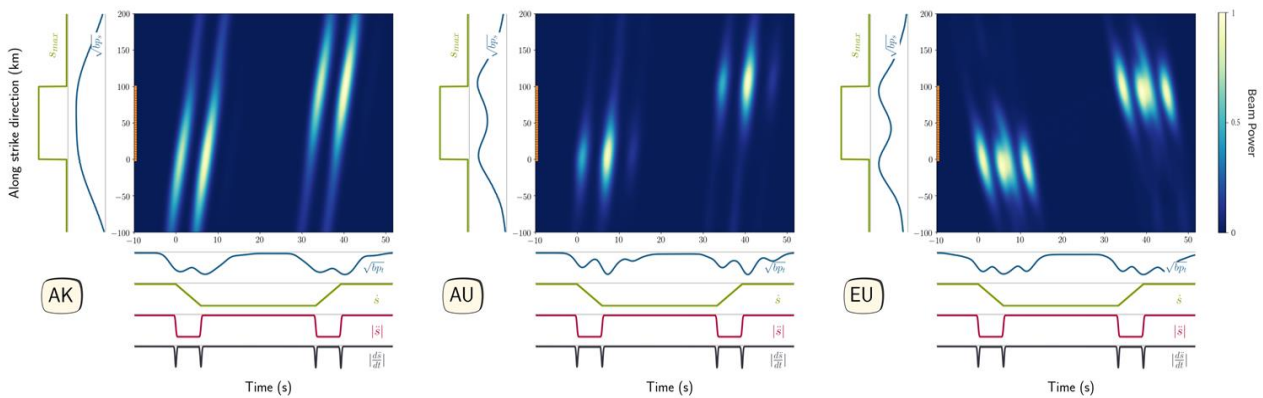


Figure 2: BP images of a homogeneous rupture process calculated at the three arrays. Time is plotted along the x-axis and the along strike direction is plotted along the y-axis. In the left subplots we show the squared root of the beam power integrated in time (in blue) and the maximum value of the slip-rate function (in green). In the bottom subplots we show the squared root of the beam power integrated in space (in blue), the slip-rate function (in green) and the absolute values of the slip-acceleration function (in red) and its time derivative (in black). The blurred and tilted patches correspond to the BP reconstruction of the HF emissions related to the source complexity and the contamination of the depth phases.

253

254 A current issue in BP analysis is therefore how one can remove the smearing effect and the nonphysical
 255 contribution of the depth phases that are incorrectly back-projected by the direct P wave travel times to obtain
 256 only the true location of the HF radiation in space and time. In **Fig.2** we show by way of example the BP
 257 images of a homogeneous rupture process calculated at the AK, AU and EU arrays using the conventional
 258 approach described in Ishii et al. (2005). The figure shows that a few pulses of HF emission retrieved by the
 259 BP analysis are found in correspondence with the nucleation and the stopping phases of the rupture. This
 260 happens because in a homogeneous rupture process all kinematic parameters are constant along the line fault,
 261 and only the nucleation and the stopping phases of the rupture generate abrupt changes in the slip-rate function
 262 (see the slip-rate function in the bottom subplot), that in turn produce seismic waves. Hence, the two bursts of
 263 HF radiation at the nucleation and at the end of the simulation are an effect of the finite duration of the rise-
 264 time. The presence of several pulses of HF emissions both at the nucleation and at the stopping of the rupture
 265 is due to the contamination of depth-phases (see in particular the BP image obtained at the EU array). In
 266 addition, the HF emissions are also tilted in time depending on the relative position between the array and the
 267 rupture direction. Several approaches have been proposed in the attempt to improve resolution and reduce the

268 artifacts in BP images (e.g., Lay et al. 2010, Wang et al. 2012, Haney 2014, He et al. 2015, Nakahara and
 269 Haney 2015, Wang and Mori 2016). In the first part of our study, we use the method described in Wang et al.
 270 (2016), hereinafter W2016. In their study, the authors propose a two-step procedure, first to correct the time
 271 tilt of HF emission patches and second to reduce the smear around the true HF energy peaks. The first step of
 272 the approach proposed by W2016 is performed by selecting a reference station lying in a central position within
 273 the array. In the conventional BP technique, the signal is shifted by the theoretical travel time t_{ij}^P between the
 274 grid point i and the station j . In the approach suggested by W2016, the signal recorded at a station j is shifted
 275 by the difference between the travel time at the station j and the travel time at the reference station:

276

$$277 \quad s_i(\mathbf{t}) = \sum_j \mathbf{u}_j (\mathbf{t} - \boldsymbol{\tau}_{ij}^P). \quad (2)$$

$$278 \quad \text{and } \boldsymbol{\tau}_{ij}^P = \mathbf{t}_{ij}^P - \mathbf{t}_{ij}^P. \quad (3)$$

279

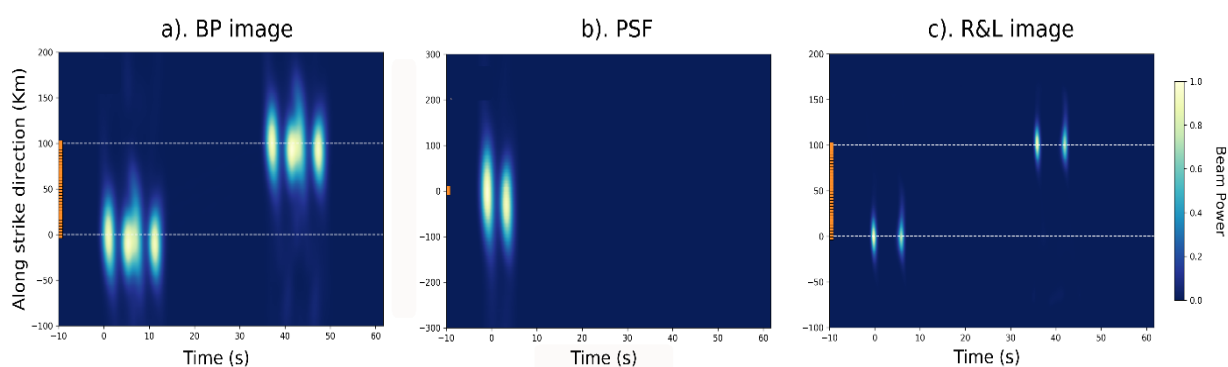
280 where \mathbf{u}_j is the velocity trace at the j -th station, $\boldsymbol{\tau}_{ij}^P$ is the i -th source travel-time difference between the j -th
 281 station and the reference station J . This procedure returns a BP image of the rupture as if it were seen by the
 282 reference station, where the HF emissions are no longer tilted in time. However, by following the ‘time
 283 correction’ of the approach proposed by W2016, we obtain a BP image where the time axis is no longer the
 284 absolute time at the source, but it rather corresponds to the apparent time at the reference station. It is instructive
 285 to notice that, in this new reference system, the directivity effect of the source appears in the BP images in
 286 terms of time stretching or compression of the HF emissions (see **Supplementary Material**).

287

288 Several authors have interpreted the smear around the true HF peaks on BP images as an artifact created by
 289 the array, referred to as array response (Xu et al. 2009, Meng et al. 2012, W2016). In particular, in their second
 290 step of the approach, W2016 suggest looking at the BP image of an earthquake as a convolution between the
 291 BP image of the true HF peaks and the BP image of the array response. The array response can be evaluated
 292 in different ways: theoretically, with synthetic Green’s functions (e.g., Rost and Thomas, 2002), or with
 293 empirical Green’s functions. The latter is usually performed by applying the BP analysis to a smaller
 294 earthquake that can be assumed to be a point source. The procedure proposed by W2016 can also help remove
 295 the contamination of depth phases into the BP images. In their study, W2016 use an aftershock as an effective

296 point source and a non-negative least squares (NNLS) algorithm to perform the deconvolution. In our study
297 we use a synthetic point source activating at the hypocenter as array response and perform the deconvolution
298 using the Richardson and Lucy restoration algorithm (Richardson 1972, Lucy 1974), hereinafter R&L. In **Fig.3**
299 we show an example of the R&L restoration algorithm applied to the BP image of a homogeneous rupture
300 process calculated at the EU array. Once the HF peaks have been extracted via the R&L algorithm (**Fig.3c**),
301 we restore the time axis in the BP images from the apparent time of the reference station to the absolute time
302 at the source by inversion of the W2016 time correction equation.

303



304

Figure 3: example of the Richardson and Lucy restoration algorithm applied to the BP image of a homogeneous rupture process calculated at the EU array. (a) Original BP image; (b) PSF obtained as the BP image of a point source activating at the hypocenter; (c) Restoration result after the R&L algorithm.

305 **3. Relation between BP and fault slip**

306

307 To better understand the relation between the BP image and the fault slip, we select two representative cases
308 among the 40,000 rupture scenarios:

- 309 1) a homogeneous rupture model, where the kinematic parameters are constant along the line;
- 310 2) a heterogeneous rupture model, where a segment of 30 km length is introduced in the middle of the fault;
311 within this segment, the rise-time value assumes a different value respect to the surrounding rupture.

312

313 **3.1 Homogeneous rupture model**

314 In a homogeneous rupture model, the kinematic parameters are constant along the line: the rupture propagates
315 along the line fault with constant velocity and the source points have the same value of rise time and final slip.
316 Specifically, the case we selected presents the following values: rupture velocity $V_r = 3$ km/s; rise time $t_r = 6$
317 s; final slip $s_f = 8$ m.

318

319 In a homogeneous rupture process, the shape of the slip-rate function is extremely simple, and the unique
320 abrupt changes are the slopes associated with the nucleation and stopping of the rupture. An example of
321 teleseismic synthetic traces generated at the three arrays is shown in **Fig.4a**, where the signal is characterized
322 by two abrupt changes, corresponding to the initiation and the stopping of the rupture, whereas the remaining
323 portion of the signal is flat, indicating a constant rupture process. In **Fig.4a** we notice that the HF band-pass
324 filter behaves like a time-derivative on the velocity trace as, more specifically, the pulses on the HF-filtered
325 signal highlight the discontinuities of the velocity trace. In particular, the effect of the finite duration of the
326 rise-time determines a two-pulse HF radiation both for the nucleation and the stopping of the signal. The effect
327 of the depth phases is almost imperceptible on the AK array, because of the radiation pattern. Here, four HF
328 pulses are visible in association with the discontinuities in the slip-rate function due to the nucleation and the
329 stopping phases of the rupture. In contrast, the waveforms at the AU and EU arrays are measurably perturbed
330 by the depth phases. Here, synthetics show at least three pulses for the AU array and four pulses for the EU
331 array.

332



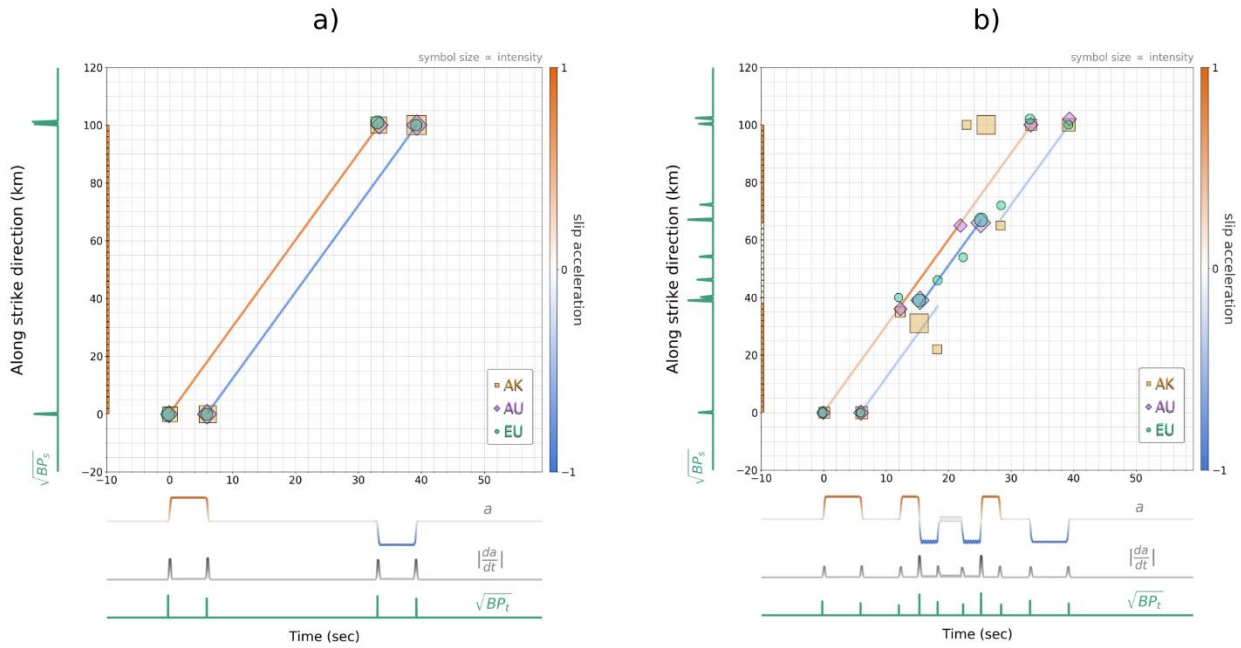
Figure 4: Example of a teleseismic synthetic trace generated at AU, AK, and EU array for a homogeneous (a) and a heterogeneous (b) rupture process. In each plot, we show the displacement trace on the top, the velocity trace in the middle, and the velocity trace band-pass filtered between 0.5-4 Hz at the bottom.

333

334

335 In Fig.5a we show the location in space and time of the HF peaks obtained via the R&L restoration algorithm
 336 to eliminate the contamination of depth phases in the BP images. The three arrays are in agreement in terms of
 337 the location of the peaks in space and time and their intensity. Four HF peaks are observed: the first two,

338 generated at the hypocenter, corresponding to the rupture nucleation; the other two, generated at the end of the
 339 ruptured segment, corresponding to the rupture stopping. In particular, if we compare the HF peaks with the
 340 subplots in time axis, we notice that the HF peaks appear at the discontinuities of the slip acceleration function
 341 and are therefore better associated with the absolute value of the time derivative of the slip-acceleration
 342 function.
 343



344

Figure 5: HF peaks for a homogeneous (a) and a heterogeneous (b) rupture process retrieved at AK, AU and EU array. HF peaks size is proportional to the intensity of the beam power. In the main plot we show the normalized slip-acceleration function. The along strike direction is on the y-axis (in orange we plot the line source) and the time is on the x-axis. The marginal on the left is the squared root of the beam power signal for the EU array. The marginal on the bottom are: the total slip-acceleration function a ; its time derivative in absolute value; the squared root of the beam power signal for the EU array.

345

346

347 3.2 Heterogeneous rupture model

348 In this heterogeneous case, a 30 km long segment is placed in the middle of the line fault and we change the
 349 rise time value for the source points within it. The rupture propagates with constant rupture velocity ($V_r = 3$
 350 km/s) and the source points reach the same final slip value ($s_f = 8$ m). The rise time, on the contrary, has an
 351 abrupt change from 6 s to 3 s, determining a faster activation of the ‘heterogeneous’ points. An example of

352 teleseismic synthetic traces for this heterogeneous rupture process generated at the three arrays is shown in
353 **Fig.4b**, where the signals show a higher complexity beyond the nucleation and the stopping of the rupture. In
354 particular, the decrease in the rise time value within the heterogeneous segment produces additional elastic
355 waves, since the slip rate function experiences a transient acceleration when encountering the heterogeneity.
356 Multiple pulses in fact appear on the HF-filtered traces **Fig.4b**: they mark not only the nucleation and the
357 stopping phases, but also the major discontinuities in the slip-rate function due to the location of the
358 heterogeneity.

359

360 In figure **Fig.5b**, the three arrays are not always in agreement in terms of the location of the peaks in space and
361 time, neither do they retrieve the same number of peaks. As previously seen in the homogeneous rupture
362 process, HF peaks are again observed at the hypocenter, corresponding to the rupture nucleation, and at the
363 end of the ruptured segment, corresponding to the rupture stopping. Additional HF peaks appear at the edges
364 of the heterogeneous segment in space, and they match the discontinuities of the slip acceleration function in
365 time. The result of the peak extraction following the R&L restoration algorithm is not optimal for the AK array,
366 whereas it works better for the AU and EU arrays. For the EU array in particular, the beam power closely
367 matches the absolute value of the time derivative of the slip-acceleration function. Along the strike direction,
368 the localization of the peaks is good for the hypocenter and the end of the ruptured segment, whereas the
369 localization of the peaks generated by the presence of the heterogeneous segment is not always clear. The rise
370 time, on the contrary, has an abrupt change from 6 s to 3 s, determining a faster activation of the ‘heterogeneous’
371 points.

372

373 4. A Convolutional Neural Network approach

374

375 The second step of this work is to understand the statistical link between the BP image of an earthquake and
376 the kinematic parameters controlling the rupture process. Here we apply a CNN approach to the 40,000 BP
377 images of the rupture scenarios previously generated. To take full advantage of the features contained in the
378 BP image, whether they be artifacts or true source emissions, for each rupture process we calculate the BP
379 image using the conventional approach by Ishii et al. (2005). In the calculation of the BP image, we replace
380 the stacked signal by its squared value, the beam power. However, in this second part of our study, we want to
381 apply a minimal pre-processing and allow the CNN to find the best link between the BP image and the
382 kinematic parameters. Hence, hereafter the beam is not smoothed with a Gaussian filter and, most importantly,
383 no peak extraction or other manipulation is made on the BP images. Finally, we apply a CNN approach to
384 exploit the information carried by the HF radiation in the BP images and to understand the relation between
385 the features and the kinematic parameters.

386

387 4.2 CNN architecture

388 The CNN is a regression model that is trained in a supervised way and therefore needs an input and a target.
389 In our study, the input is a 2D matrix containing the pixels of the BP images of the simulated rupture processes,
390 and the target is a 1D vector containing the kinematic parameters corresponding to those simulations.
391 Specifically, the target vector d is defined as:

392

$$393 \quad \mathbf{d} = [t_r, s_f, V_r, H_{tr}, H_{sf}, H_{Vr}, x_c, L_H]. \quad (4)$$

394

395 where t_r is the rise time, s_f is the final slip, V_r is the rupture velocity; H_{tr} , H_{sf} and H_{Vr} are the heterogeneities
396 respectively in the rise time, in the final slip and in the rupture velocity; x_c and L_H are the central position and
397 the length of the heterogeneous segment. The objective of the CNN algorithm is to learn from data a functional
398 mapping between the input matrix and target vector (LeCun et al., 1998, Goodfellow et al. 2016).

399

400 We tested various network architectures in our study. Our preferred one has the structure sketched in **Fig.6**,
401 because it required a smaller number of learnable parameters (improving its generalization performance) and

402 was computationally efficient in training. In this architecture, the BP image is first downsampled with a max-
403 pooling operation, with equal pooling lengths of $[3] w \times [3] h$ along the image dimensions. Once downsampled,
404 the BP image is passed through two convolutional layers, each with a convolution, a max-pooling, and an
405 activation function. The convolution is performed with 5 and 10 filters, in the first and second layers,
406 respectively. The kernel size for both layers is $[5] w \times [5] h$. The max-pooling operation is performed with
407 pooling lengths of $[4] w \times [4] h$ in the first layer and of $[3] w \times [3] h$ in the second layer of the loop. A rectified
408 linear unit (ReLU) (Ramachandran et al. 2017) is used as an activation function. Two fully-connected layers
409 link these convolutional layers with the output target (**Fig.6**). We refrain from using a large number of
410 convolution and pooling layers, as is common in CNN applications, because we want to preserve a substantial
411 portion of the space-time axes of the image, since we infer that space-time resolution is important in back-
412 projection analysis to back-out the relevant kinematic parameters. Network weights are updated during the
413 learning to minimize the loss function, in our case, we use the mean squared error (MSE):

414

$$MSE = \frac{1}{M} \sum_{n=1}^M (d_n - y_n)^2. \quad (5)$$

415

416 where M is the number of samples in the dataset, d_n is the normalized target vector containing the true
417 kinematic parameters for the sample n , y_n is the normalized output vector containing the prediction of the
418 kinematic parameters for the sample n . The kinematic parameters used in the simulations are uniformly
419 distributed, hence we adopt a min-max normalization on target data to map them to the range 0 to 1:

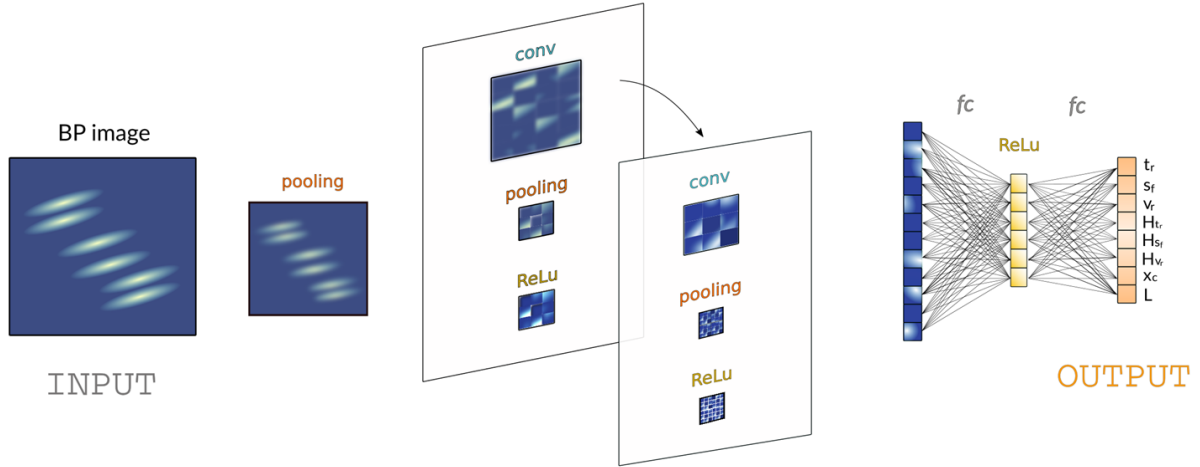
420

$$\tilde{d}_n = \frac{d_n - \min(d)}{\max(x) - \min(x)}. \quad (6)$$

421

422 where \tilde{d}_n is the normalized target. During the training, the updates to the model weights are controlled by the
423 learning rate, which quantifies how fast the model adapts to the problem. In our study, we initialize the weights
424 using the Glorot initialization scheme and set the learning rate equal to 7.5×10^{-5} . The Adam algorithm
425 (Kingma and Ba 2014), applied in the PyTorch framework (Paszke et al., 2017), is used to train the networks.
426 The original dataset composed of 40,000 simulations of rupture processes is divided into three subsets: training
427 (70%), validation (20%) and testing (10%). The network is trained for up to 500 iterations over the training

428 dataset. The model's parameters for which the MSE reaches its minimum during the validation step are then
429 selected as the best model's parameters which will be used for the testing.
430



431

Figure 6: Sketch illustrating the design of the CNN used in the present study. The input is the BP image which undergoes a max pooling operation first. Then, once downsampled, the BP image is passed through two convolutional layers, each composed of a convolution (conv), pooling and activation (ReLU) operation. The network ends with a two fully-connected layers (fc). The output of the network is the vector containing the kinematic parameters of the rupture process.

432

433 5. Statistical link between the BP image and the target parameters

434

435 5.1 Effect of target on CNN predictions

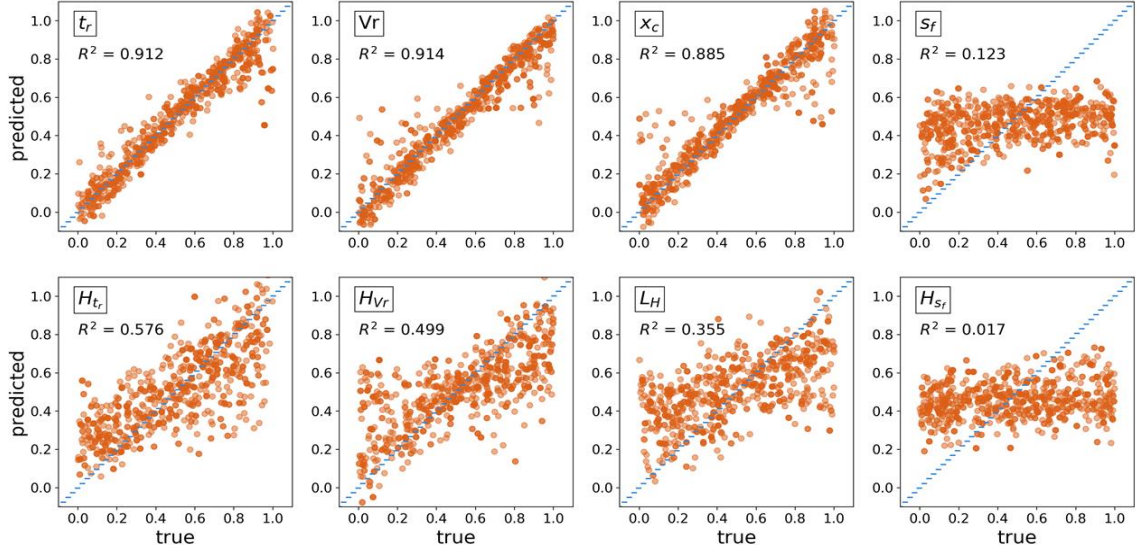
436 In Fig.7 we present the results obtained with the design of the CNN sketched in Fig.6 and the target vector
437 defined in (4). We include outputs of both the stacked version of the method (sum of the BP images obtained
438 from the three arrays), and the predictions for using each array separately. The plots show the output of the
439 CNN versus the target values, which are the predictions of the CNN versus the true values of kinematic
440 parameters, for the testing dataset. To summarize the prediction accuracies, we calculate the regression score
441 function, R^2 , for each of the different components of the target vector. This score is defined as:

442

$$443 \quad R^2 = 1 - \frac{\sum_{i=1}^n (d_i - y_i)^2}{\sum_{i=1}^n (d_i - \hat{d})^2}. \quad (7)$$

444 where d is the target vector, y is the CNN prediction and \hat{d} is the mean value of the target. This score captures
445 how well the model is able to return a prediction close to the target parameter in a linear regression setting. Its
446 range of possible values is $(-1, 1]$, where the best possible score is 1.0 and a score of 0.0 implies the model
447 predictions are no better than a simple guess based on the average value. A negative value for R^2 implies the
448 model predicts worse than the simple guess. The highest values of R^2 are found for the predictions of the rise
449 time t_r ($R^2 = 0.912$) the rupture velocity V_r ($R^2 = 0.914$) and the central position of the heterogeneous segment
450 x_c ($R^2 = 0.885$). Accurate predictions are also obtained for the heterogeneous values in rise time and rupture
451 velocity, H_{tr} ($R^2 = 0.576$) and H_{vr} ($R^2 = 0.499$). However, the length of the heterogeneous segment L_H is not
452 well predicted ($R^2 = 0.355$) and poor predictions are obtained for the final slip s_f ($R^2 = 0.123$) and its
453 heterogeneous values H_{sf} ($R^2 = 0.017$). We attribute this shortcoming to the inherent insensitivity of the BP
454 approach to image the low-frequency aspects of the rupture process.

455



456

Figure 7: Predictions of the CNN versus the true values of kinematic parameters, for the testing dataset. The regression score function R^2 is shown in each subplot. The input BP image is obtained as the sum of the BP images calculated at the three arrays AK, AU and EU. We adopt a min-max normalization on target data to map them to the range 0 to 1.

457

458

459 Difficult-to-predict parameters in the target vector slow down the loss function's convergence towards its
 460 minimum. Therefore, we define a different target vector, where we remove the final slip value and its
 461 heterogeneity and keep the following kinematic parameters:

462

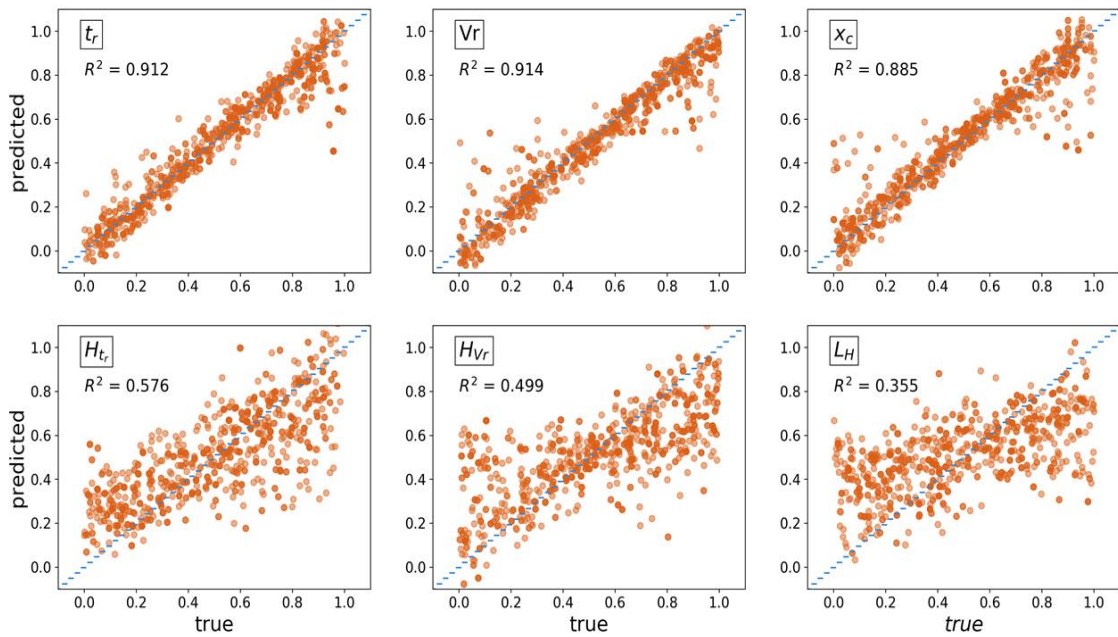
463

$$d = [t_r, V_r, H_{tr}, H_{Vr}, x_c, L_H]. \quad (8)$$

464

465 In **Fig.8** we show the prediction versus the true values of the reduced target vector in (8). The weights for each
 466 output channel are independent, therefore the performance on the remaining parameters does not change
 467 significantly with a different choice of target vector.

468



469

Figure 8: Predictions of the CNN versus the true values of kinematic parameters of the reduced target vector for the testing dataset. The input BP image is obtained as the sum of the BP images calculated at the three arrays AK, AU and EU. Target data have been scaled to 0-1 via min-max normalization.

470

471

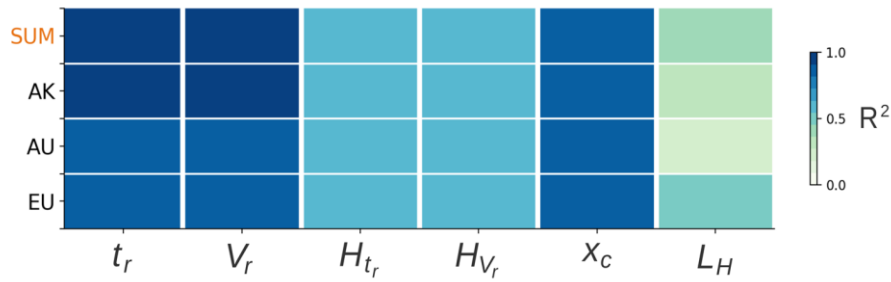
472 5.2 Effect of input on CNN predictions

473 As complementary information, we test the effect of the input parameter on the CNN results. In the previous
 474 cases, the input of the CNN was a matrix containing the BP images of the simulated rupture processes, obtained
 475 as the sum of the BP images calculated at the three arrays AK, AU and EU. Here, we test whether the
 476 combination of the three arrays or the employment of only one array at a time in the CNN approach could help
 477 us achieve better results.

478

479 In **Fig.9** we present the comparison between the results of the CNN approach applied to the sum of the BP
 480 images and to the single-array BP image using the reduced target vector. It is instructive to notice that not all
 481 arrays perform the same. In particular, slightly better predictions of the rise time and its heterogeneous value
 482 can be found when using the BP image of the AK array. The comparison also shows an improvement in the
 483 predictions of the length of the heterogeneous segment when we use an input BP image calculated at the EU
 484 array. On the contrary, when using the AK or the AU array in isolation, the predictions of the reduced target

485 vector do not improve in comparison to the initial test where the input is the sum of the three BP images. In
 486 these cases, the smearing or the walking effect on BP images may hamper the accuracy of the CNN predictions.
 487 In our case, the EU and AU array response functions are quite focused along the space direction, whereas the
 488 AK array shows higher smearing. The EU array shows the strongest walking effect, because of its large back-
 489 azimuth, whereas the HF emissions on AK and AU BP images are not very tilted in time (see as an example
 490 the **Fig.2**). However, only the EU array allows the CNN to return the best predictions of the length of the
 491 heterogeneous segment, even better than the outputs obtained with the summed BP image. On the contrary,
 492 both AK and AU arrays perform worse than the summed BP image for the length of the heterogeneous segment.
 493 Even though the smearing effect could encumber the ability of the network to extract information on the BP
 494 images, we can't attribute to it the shortcomings of the CNN on AK and AU arrays.
 495



496

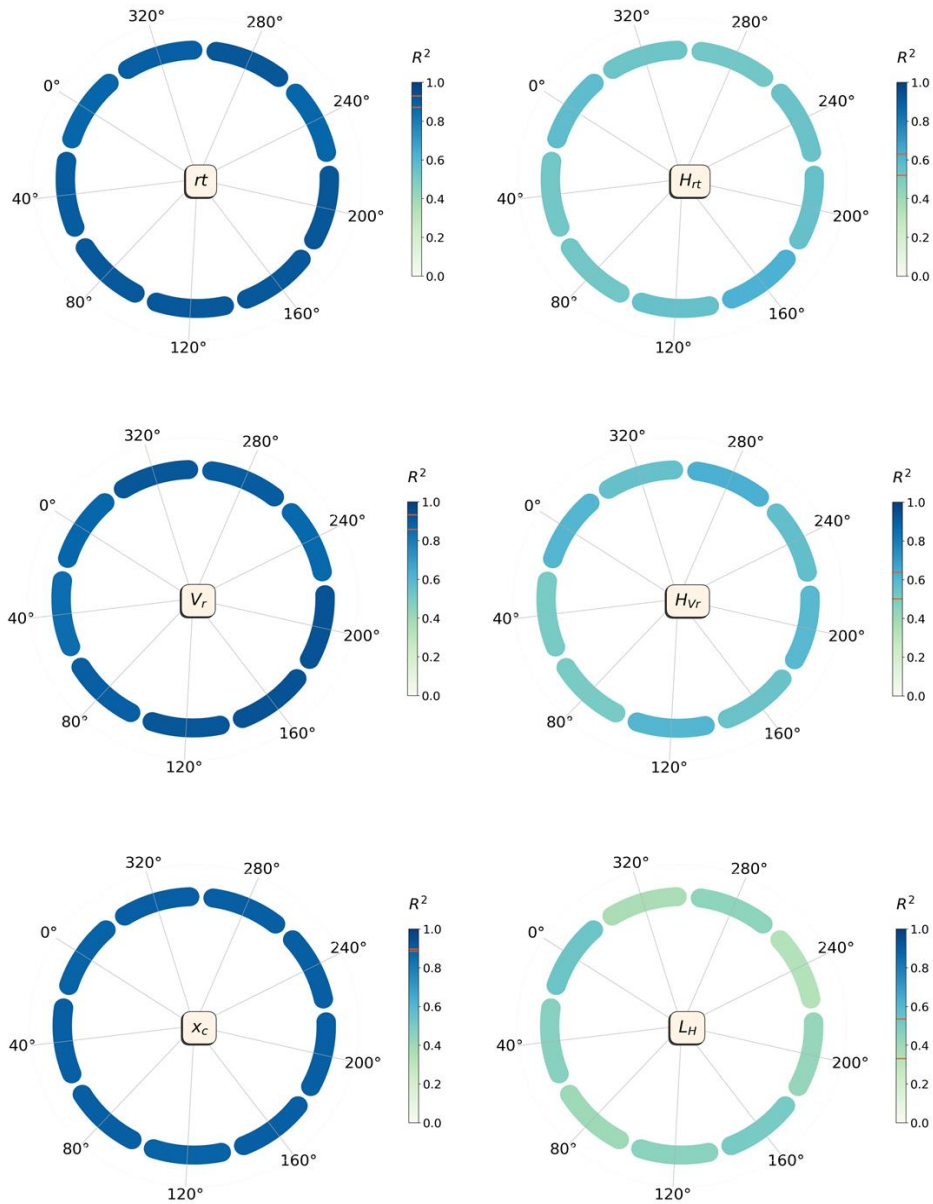
Figure 9: Regression score function (R^2) for the CNN predictions of the kinematic parameters in the reduced target vector when using the sum of the three BP images (on the top) or the single-array BP image as input for the network. The score represents the quality of the model indicating how close the predictions are to the target parameters in a linear regression setting.

497

498 We question, therefore, whether the different performances of the CNN on individual arrays could be
 499 attributable to the back-azimuth of the array relative to the rupture. In the conventional BP analysis, the delay-
 500 and-sum approach does not carry directivity effects in the duration of the rupture, nor in the time separations
 501 between HF pulses in the BP image. Hence, the time axis of the BP image is not affected by the relative position
 502 between the rupture direction and the array position. Nevertheless, different tilt angles for the smeared HF
 503 radiations are seen in the BP images depending on the back-azimuth. Because of the array configuration and
 504 position respect to the rupture direction, coherence among traces can vary from one array to another,
 505 determining a different pattern of HF peaks in the BP image among arrays at complementary azimuths [e.g.,

506 Xu et al. (2009)]. As a further analysis, we investigate whether the back-azimuth and thus the different HF
507 emission pattern and tilt in the BP image, can favor or restrain the quality of the CNN predictions. We focus
508 the analysis on the EU array because of its strong walking effect. We rotate the previously used line source at
509 nine different azimuths ranging from 0° , being the azimuth of the EU array reference station, to 320° . For each
510 azimuth, we generate 40,000 rupture scenarios and apply the CNN method on them. In **Fig.10** we show the
511 prediction of the CNN for the parameters in (6) and compare the quality of the CNN prediction obtained at
512 nine different azimuths. For each kinematic rupture parameter, we highlight in orange the variability of R^2 as
513 a function of azimuth in the corresponding regression score color-bar. From this analysis, we can notice that
514 the central position of the heterogeneous segment is the best retrieved parameter for all azimuths, whereas the
515 length of the heterogeneous segment shows the widest range of regression scores. But even this range is rather
516 modest [0.35 to 0.50], suggesting that azimuthal changes have a minimal effect on the quality of the CNN
517 predictions.

518



519

Figure 10: Azimuthal dependence of CNN predictions for the kinematic parameters in the reduced target vector. The variability of the regression score is represented with an orange interval. For this test the EU array is used and 0° is the back-azimuth of the EU array respect to the line fault strike.

520 **6. Discussion and Conclusion**

521 The objective of this study was to investigate the link between BP images and the kinematic parameters of the
522 rupture, and, in particular, slip motion on the fault (velocity or acceleration). For this purpose, we studied a
523 large data set of synthetic line rupture processes, characterized by a heterogeneous segment (in terms of final
524 slip, rise time or slip velocity) with variable length and position. Synthetic traces, filtered between 0.5 and 4
525 Hz, are back projected following the approach of Wang et al. (2016) and the HF peaks are retrieved through
526 deconvolution between the BP image and the array response function (ARF) for the given frequency range.
527 For the particular horizontal line-source configuration chosen here, depth phases strongly contaminate BP
528 images, introducing, for each HF peak, two "ghost" peaks associated with pP and sP phases. Deconvolution
529 with an ARF that includes depth phases restores the original number of peaks with good time accuracy, but
530 sometimes shifts the peaks in space. The role played by depth phases is amplified by the simple horizontal line
531 source geometry chosen here, where all the fault points lay at the same depth. For a more realistic 2D fault,
532 the recording stations will "see" at a given time the energy emitted from different points along a fault isochrone
533 (Bernard and Madariaga, 1984; Spudich and Frazer, 1984) which are generally incoherent and lay at different
534 depths (the isochrone for a line source is a single point). Moreover, the spatial extension of the ARF introduces
535 a further averaging scale which should reduce the coherency of depth phases. The resolution of BP becomes
536 very poor at depth, hence improved BP techniques such as the hybrid BP proposed by Yagi et al. (2012) could
537 be useful in mitigating the effect of depth phases on BP images of 2D synthetic fault models. An important
538 question is whether BP images are associated with slip velocity or acceleration (Fukahata et al., 2014). Here
539 we show that filtering plays an important role, since the 0.5-4 Hz band-pass filter, typically used in BP analyses
540 (Xu et al., 2009) behaves like a time derivative for the seismograms. Comparison between HF peaks extracted
541 from BP images with the slip rate and slip acceleration function, shows that the beam power is more likely to
542 be related to the absolute value of the time derivative of the slip acceleration function, when narrow-band
543 filtering is used. The ability of BP images to retrieve the rupture kinematic parameters was tested using a CNN
544 approach on BP images. CNN are data-driven predictive models, whose performance depends on the definition
545 of input and target parameters. We found that the CNN is able to predict the rise time, the rupture velocity, the
546 heterogeneous values in rise time and rupture velocity, the length and the central position of the heterogeneous
547 segment. However, the CNN fails at predicting the final slip and its heterogeneous value. We attribute this
548 shortcoming to the inherent insensitivity of the BP approach to the low-frequency aspects of the rupture

549 process. We also tested whether the information coming from one single array could be thorough for the
550 network or, on the contrary, if the combination of the information coming from the three arrays could provide
551 us with better predictions. In particular, we tested whether the back-azimuth of the array relative to the rupture
552 could influence the quality of the CNN predictions. Our analysis shows no strong azimuthal dependence in the
553 quality of the CNN predictions depending on the relative position between the source and the array. Thus, from
554 a CNN perspective, stacking multiple arrays may not always provide the best outcome, in contrast to what is
555 more common place in teleseismic BP analysis. It is worth noting that we trained our CNN on a simplistic case
556 of a line source, with the objective to assess the resolving power of BP in a controlled test. Generalization of
557 our CNN approach will require training the CNN on more realistic 2D source models, e.g., using fractal slip
558 distribution (Ruiz et al., 2011), dynamic modeling or real earthquakes.

559

560 Even though our study does not fully address the question of the generalizability of the CNN method, it does
561 demonstrate the potential upsides of machine learning approaches in providing reasonably accurate predictions
562 for the other kinematic parameters of the rupture process, which can open a field for its use. Further analysis
563 on the waveform content of synthetic data, as well as a careful analysis on the similarity between the BP images
564 of real data and synthetic data in a fixed frequency band, would enrich our study, potentially making it a
565 suitable approach for real data too, as long as waveforms are carefully pre-processed.

566

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754

755 ***Author contribution statement:***

756 M.C. and C.S. conceived the study. M.C. performed the computations and processed the data. M.C. C.S. and
757 P.B. interpreted and discussed the link between the BP image and the slip function. M.C., I.M. and D.T.T.
758 conceived the computational framework of the CNN analysis. I.M. designed the CNN architecture. M.C., I.M.
759 and D.T.T. interpreted the data and discussed the results of the CNN analysis. M.C. drafted the manuscript. All
760 of the authors provided critical feedback on the results and discussion and helped shape the manuscript. P.B.
761 and P.A.J. supervised the project.

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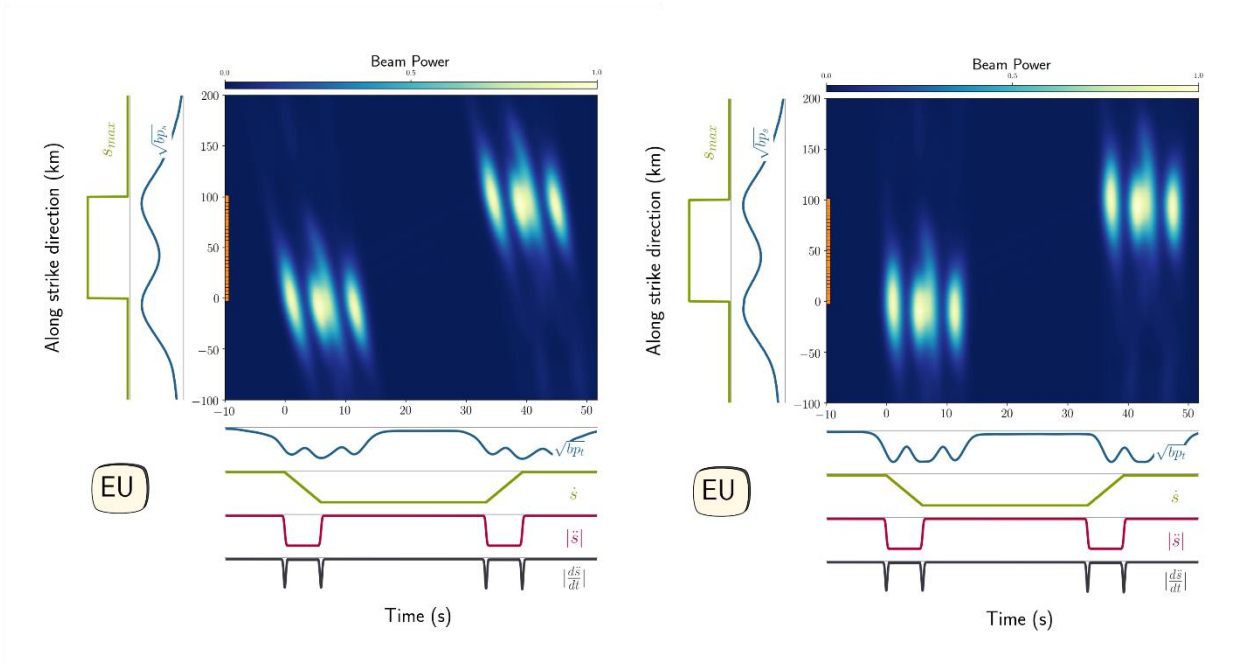
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776 **Supplementary Material**

777 The procedure suggested by W2016 transforms the BP image conventionally calculated in the time at the
778 source in a BP image in the apparent time of the reference station. In this new reference system, the HF
779 emissions do no longer suffer from the ‘walking effect’, however they are stretched or contracted along the
780 time axis because of the directivity effect. In **Fig.S1** we show, by way of example, the comparison between the
781 BP image of a homogeneous rupture process calculated at the EU array following the conventional approach
782 of Ishii et al. (2005) and the BP image of the same rupture process calculated using the W2016 reference station

783 correction. We notice that, in the reference station system, the HF radiation is stretched along the time axis.
 784 This happens because in this example the EU array is located in the anti-directive position respect to the
 785 direction of propagation of the rupture.
 786



787

788 **Figure S1:** HF emissions radiated by a homogeneous rupture process retrieved by BP analysis using
 789 the conventional approach (on the top) and the W2016 approach (at the bottom). Time is on the x-
 790 axis, the along strike direction is on the y-axis. The squared root of the beam power integrated in time
 791 (in blue) and the maximum value of the slip-rate function (in green) are plotted in the left subplots.
 792 The squared root of the beam power integrated in space (in blue), the slip-rate function (in green)
 793 and the absolute values of the slip-acceleration function (in red) and its time derivative (in black)
 794 are plotted in the bottom subplots. In the conventional BP analysis (BP image on the top), the HF
 795 emissions are tilted along the time axis. In the W2016 approach (BP image at the bottom), the HF
 796 emissions are stretched in time because of the directivity effect.