

1                    **VESIcal Part II: A critical approach to volatile**  
2                    **solubility modelling using an open-source Python3**  
3                    **engine**

4                    **P. E. Wieser<sup>1,2\*</sup>, K. Iacovino<sup>3</sup>, S. Matthews<sup>4</sup>, G. Moore<sup>3</sup>, C. M. Allison<sup>5,6</sup>**

5                                    <sup>1</sup>Department of Earth Sciences, University of Cambridge, UK.

6                                    <sup>2</sup>College of Earth, Ocean and Atmospheric sciences, Oregon State University

7                                    <sup>3</sup>Jacobs, NASA Johnson Space Center, Houston, TX 77058, USA.

8                    <sup>4</sup>Johns Hopkins University, Department of Earth and Planetary Sciences, Baltimore, MD 21218, USA.

9                                    <sup>5</sup>Cornell University, Department of Earth and Atmospheric Sciences, Ithaca, NY 14853

10                                   <sup>6</sup>City College of New York, City University of New York, New York, NY 10031

11                    **Key Points:**

- 12                    • The Python3 tool VESIcal allows extensive comparisons to be drawn between  
13                    different H<sub>2</sub>O-CO<sub>2</sub> solubility models
- 14                    • Solubility models are not interchangeable - for a single magma composition,  
15                    different models can predict a wide range of solubility relationships
- 16                    • The P-T-X calibration range of each solubility model must be critically evalu-  
17                    ated before application to a specific volcanic system

---

Corresponding author: P. Wieser, [penny.wieser@gmail.com](mailto:penny.wieser@gmail.com)

**Abstract**

Accurate models of H<sub>2</sub>O and CO<sub>2</sub> solubility in silicate melts are vital for understanding volcanic plumbing systems. These models are used to estimate the depths of magma storage regions from melt inclusion volatile contents, investigate the role of volatile exsolution as a driver of volcanic eruptions, and track the degassing paths followed by magma ascending to the surface. However, despite the large increase in the number of experimental constraints over the last two decades, many recent studies still utilize the earlier generation of models, which were calibrated on experimental datasets with restricted compositional ranges. This may be because many of the available tools for more recent models require large numbers of input parameters to be hand-typed (e.g., temperature, concentrations of H<sub>2</sub>O, CO<sub>2</sub>, and 8–14 oxides), making them difficult to implement on large datasets. Here, we use a new open-source Python3 tool, VESIcal, to critically evaluate the behaviours and sensitivities of different solubility models for a range of melt compositions. Using literature datasets of andesitic-dacitic experimental products and melt inclusions as case studies, we illustrate the importance of evaluating the calibration dataset of each model. Finally, we highlight the limitations of particular data presentation methods such as isobar diagrams, and provide suggestions for alternatives, and best practices regarding the presentation and archiving of data. This review will aid the selection of the most applicable solubility model for different melt compositions, and identifies areas where additional experimental constraints on volatile solubility are required. (242/250 words)

**Plain Language Summary**

Being able to accurately model the solubility of H<sub>2</sub>O and CO<sub>2</sub> in magmas is very important for understanding a wide variety of volcanic processes, such as the depths at which magma is stored in the crust, the driving force behind volcanic eruptions, and the release of volatile elements into the atmosphere. However, there has been no easy way for volcanologists to perform calculations on large datasets, or to compare different models. This review uses a new, open-source tool called VESIcal written in the popular programming language Python3. This allows us to compare different models for a wide variety of melt compositions, temperatures and pressures, helping researchers to identify the most suitable model for their study.

50 We also suggest areas where further experimental constraints are required. Finally,  
51 we highlight the limitations of particular data presentation methods such as isobar  
52 diagrams, and provide suggestions for alternatives, and best practices regarding the  
53 presentation and archiving of data.

## 54 **1 Introduction**

55 The most abundant volatile components found in terrestrial magmatic systems  
56 are H<sub>2</sub>O and CO<sub>2</sub>. It has been known for nearly a century (Bowen, 1928; Tuttle &  
57 Bowen, 1958) that these volatile species have profound effects on the chemical and  
58 material properties of magmas (e.g., phase equilibria, melting temperatures, magma  
59 viscosity and density; Burnham, 1979; Husen et al., 2016; Burnham & Davis, 1974;  
60 Hess & Dingwell, 1996; Ochs & Lange, 1999), so significantly affect their geochem-  
61 ical and dynamical behavior (e.g., eruption and degassing style, erupted volume;  
62 Papale et al., 1999; Huppert & Woods, 2002; La Spina et al., 2021). Thus, it is vital  
63 to be able to predict how H<sub>2</sub>O and CO<sub>2</sub> solubilities change as a function of inten-  
64 sive variables such as pressure, temperature, melt and fluid composition in order to  
65 understand plutonic and volcanic systems.

66 The solubility of a volatile species is defined at a given pressure and temper-  
67 ature as the maximum concentration that can be dissolved within a silicate melt  
68 of a specified composition. Ignoring disequilibrium effects, if the volatile content of  
69 the system exceeds this solubility limit, a separate fluid/vapour phase will exsolve  
70 from the magma. In this review, we favour the term fluid because of the supercrit-  
71 ical nature of exsolved volatile phases at magmatic temperatures. In general terms,  
72 a magma is described as volatile undersaturated when there is no fluid phase, and  
73 volatile saturated once a fluid phase is present (also referred to as vapour under-  
74 saturated/saturated, or fluid undersaturated/saturated). In detail, different volatile  
75 species do not act as independent entities, but influence one another. For this rea-  
76 son, a magma may exsolve a mixed CO<sub>2</sub>-H<sub>2</sub>O fluid even if the dissolved concentra-  
77 tions of H<sub>2</sub>O and CO<sub>2</sub> do not exceed the pure solubility limit of each species.

78 Despite the obvious importance of accurate volatile solubility modelling, very  
79 few studies of volcanic systems have evaluated results using several different sol-  
80 ubility models to determine possible sources of systematic and random error, and

81 assess the suitability of each model for the conditions of interest (e.g. temperature,  
82 pressure, and melt composition). This lack of intercomparison likely results from the  
83 fact that it is extremely time consuming to perform the large numbers of calcula-  
84 tions necessary for thorough comparisons using available tools. For example, many  
85 solubility models were released as stand-alone Excel spreadsheets (e.g., G. Moore  
86 et al., 1998; Newman & Lowenstern, 2002; Allison et al., 2019) or web apps (e.g.,  
87 Iacono-Marziano et al., 2012; Ghiorso & Gualda, 2015), where saturation pressures,  
88 dissolved volatile contents, degassing paths, and isobars can only be calculated for  
89 one sample and set of conditions at a time. The more recent models which include  
90 several terms accounting for the effect of melt composition on volatile solubility  
91 require users to hand-type a large number of input parameters. For example, to  
92 calculate a saturation pressure in MagmaSat (Ghiorso & Gualda, 2015), users must  
93 hand-type 9–14 oxide concentrations in addition to entering H<sub>2</sub>O and CO<sub>2</sub> concen-  
94 trations, and a melt temperature. Similarly, the web app of Iacono-Marziano et al.  
95 (2012) requires users to input 8 major element oxide concentrations. Calculating  
96 isobars using these web apps is a particularly daunting task, as users must evaluate  
97 dissolved volatile contents at multiple fluid compositions, and then use curve fitting  
98 to produce a smooth isobar to display on plots. Other models were released with no  
99 calculator at all, requiring each user to correctly interpret and combine the relevant  
100 equations in the manuscript (Dixon, 1997; Shishkina et al., 2014).

101 Here, we take advantage of the recent release of VESIcal (Volatile Equilibria  
102 and Saturation Identification calculator; Iacovino et al., 2021), an open-source tool  
103 written in Python3. VESIcal contains functions to calculate saturation pressures,  
104 dissolved volatile contents, isobars, and degassing paths automatically for seven dif-  
105 ferent models. Calculations can be performed based on melt compositions provided  
106 in an Excel spreadsheet, and users can take full advantage of Python’s extensive  
107 flexibility to perform large numbers of calculations automatically (e.g., creating for  
108 loops to perform calculations across a range of pressures, temperatures, and fluid  
109 compositions). To our knowledge, the only other model with similar functionality to  
110 VESIcal is the Linux program Solwcad supplied by Papale et al. (2006), which per-  
111 forms calculations automatically on a user-supplied .txt file containing melt composi-  
112 tions, pressures and temperatures (<http://www.pi.ingv.it/progetti/eurovolc/>).

113 Solwcad was used alongside VESICAL in this review, through the Windows Subsystem  
114 for Linux (WSL2).

115 The overall aim of this review is to summarize the formulation, strengths and  
116 weaknesses of popular solubility models to inform users who wish to model volatile  
117 solubility in silicate melts, whether that be the calculation of melt inclusion saturation  
118 pressures, degassing paths, incorporating volatile exsolution in physical model of  
119 magma chambers (e.g. Huber et al., 2019), or calculating the dissolved volatile con-  
120 tents of fluid-saturated experimental products where the pressure, temperature and  
121 exsolved fluid composition are known (e.g. Waters & Lange, 2015). Specifically, we  
122 demonstrate in a number of ways how users investigating a specific subset of compo-  
123 sitional space (e.g., melt inclusions from a single volcano) can assess the similarities  
124 and differences between models, and evaluate these findings in the context of the  
125 calibration dataset and formulation of each model. We start by briefly summarizing  
126 the major results from volatile solubility experiments over the last century (section  
127 2), before describing nine of the most popular solubility models (section 3). We then  
128 compare the solubility of pure H<sub>2</sub>O, mixed H<sub>2</sub>O-CO<sub>2</sub>, and pure CO<sub>2</sub> predicted by  
129 different models for representative mafic and silicic compositions (section 4). We  
130 also explore the sensitivity of these models to parameters such as temperature and  
131 redox state, which are often poorly constrained in igneous systems (section 5). Fi-  
132 nally, we evaluate the suitability of these models for intermediate melt compositions,  
133 where experimental constraints are sparse relative to basaltic and rhyolitic melts  
134 (section 6). We conclude by discussing best practices for presenting and archiving  
135 data related to volatile components in igneous systems (section 7). Overall, these  
136 discussions demonstrate that there are large differences between model outputs,  
137 even in relatively "normal" melt compositions, so the choice of solubility model is a  
138 critical part of any study investigating magmatic volatiles (and needs to be justified  
139 in all cases). This manuscript will act as a guide to help users assess the suitability  
140 of each model for their specific application (supplemented by the python code pro-  
141 vided in the supporting information, which can be easily adapted to evaluate melt  
142 compositions and conditions relevant to a specific study).

## 2 Major findings from experimental studies investigating volatile solubility in magmas

One of the earliest volatile studies was that of Goranson (1931), who investigated the effect of pressure on the solubility of water in granitic melts. The classic treatise of Tuttle and Bowen (1958) investigated the impact of H<sub>2</sub>O on mineral phase equilibrium. This study led to a wider recognition of the importance of volatiles, and motivated the development of experimental and analytical approaches to determining volatile solubilities as a function of pressure, temperature, and melt composition. Hamilton et al. (1964) was one of the first to compare H<sub>2</sub>O solubilities for differing melt compositions (basalt and andesite), while also investigating the effect of dissolved H<sub>2</sub>O and oxidation state on the magmatic phase equilibria. These studies were followed by the fundamental experimental measurements of the Burnham group on the dissolution of H<sub>2</sub>O in albite melts (e.g., Burnham & Davis, 1971, 1974).

Further investigation of volatile solubility over the next four decades in natural samples and experimental products was aided significantly by analytical developments, allowing volatile contents in quenched glasses to be measured by techniques such as Fourier Transform infra-red spectroscopy (FTIR; e.g., Stolper, 1982; Fine & Stolper, 1986; Silver et al., 1990) and secondary ion mass spectrometry (SIMS; Hervig & Williams, 1988; Hauri, 2002). In particular, the high spatial resolution of FTIR and SIMS (a few tens of micrometers) meant that volatile concentrations could be measured within quenched pockets of melt trapped within crystals (termed melt inclusions). Unlike subaerially-erupted lavas which have degassed almost all their H<sub>2</sub>O and CO<sub>2</sub> following their ascent to shallow pressures, melt inclusions remain pressurized during ascent as they are trapped in relatively incompressible crystals, so retain high volatile contents (Roedder, 1979; Anderson, 1974).

Melt inclusion analyses have greatly advanced our understanding of the behavior of volatiles in volcanic systems (Lowenstern, 2003; Hauri et al., 2002; Roggensack, 2001; Wallace et al., 1995; Métrich & Wallace, 2008; Sides et al., 2014a). For example, melt inclusions provide insights into pre-eruptive volatile contents (e.g., Saal et al., 2002; Hervig et al., 1989), and links between melt volatile contents and eruption styles (Lucic et al., 2016). The strong pressure-dependence on volatile solubility means that H<sub>2</sub>O and CO<sub>2</sub> contents within melt inclusions trapped from a

176 volatile-saturated magma can be used to determine the pressure at which the inclu-  
177 sion was sealed off (termed the saturation pressure or entrapment pressure). In turn,  
178 the distribution of saturation pressures in a suite of melt inclusions can reveal the  
179 locations of the main regions of magma storage in a volcanic system. This explosion  
180 of new information from melt inclusions greatly increased the demand for flexible  
181 and accurate solubilities models that could be applied to a broad range of pressures,  
182 temperatures, and melt/fluid compositions (G. Moore, 2008; Ghiorso & Gualda,  
183 2015).

184 It has become increasingly apparent from solubility experiments that the  
185 solubility of H<sub>2</sub>O is relatively insensitive to melt composition (e.g., G. Moore &  
186 Carmichael, 1998; Shishkina et al., 2010), while CO<sub>2</sub> solubility is highly sensitive  
187 to melt composition, particularly in mafic melts where the carbonate ion is the  
188 dominant species (Dixon, 1997; Brooker et al., 2001a; Shishkina et al., 2010; Iacono-  
189 Marziano et al., 2012; Shishkina et al., 2014; Allison et al., 2019). This has lead to  
190 a great diversity in the way that various models treat the dependence of CO<sub>2</sub> sol-  
191 ubility on melt composition. In general, models have become more complex with  
192 time as the region of compositional space spanned by solubility experiments has in-  
193 creased to include more alkaline lavas. The individual role and relative importance  
194 of each cation species in the melt is still associated with a large degree of uncer-  
195 tainty (Allison et al., 2019), accounting for the larger discrepancies between different  
196 model predictions for CO<sub>2</sub> vs. H<sub>2</sub>O.

197 Experimental work has also highlighted the complexities of mixing between  
198 H<sub>2</sub>O and CO<sub>2</sub> in igneous systems. In the simplest case, the addition of one com-  
199 ponent in a melt-fluid system decreases the activity, and therefore the solubility of  
200 the other component in the melt (Lowenstern, 2001). This behavior is referred to as  
201 Henrian/ideal behavior. Henry's Law states that the amount of a volatile dissolved  
202 in a liquid is proportional to its partial pressure in the gas phase in equilibrium with  
203 that liquid. Neglecting the possible entropic effects of speciation, the addition of  
204 H<sub>2</sub>O to the fluid/gas phase acts to lower the partial pressure of CO<sub>2</sub>, and therefore  
205 lowers the solubility of CO<sub>2</sub> in the liquid. Similarly, addition of CO<sub>2</sub> to the fluid/gas  
206 phase causes the solubility of H<sub>2</sub>O in the melt to decrease.

207 Experimental studies have shown that Henry’s law is generally obeyed at low  
 208 pressures ( $<1$  kbar) in basaltic (Dixon et al., 1995) and rhyolitic melts (Blank et al.,  
 209 1993). However, at higher pressures, some experimental observations have shown  
 210 that the mixing behavior of  $\text{CO}_2$  and  $\text{H}_2\text{O}$  becomes strongly non-Henrian (Papale,  
 211 1999). For example, Eggler (1973), Mysen et al. (1976) and Mysen (1976) show that  
 212 the solubility of  $\text{CO}_2$  in albitic melts increases with the addition of  $\text{H}_2\text{O}$  at higher  
 213 pressures. This has been attributed to the fact that the addition of small amounts  
 214 of water as  $\text{OH}^-$  species decreases melt polymerization, and therefore enhances the  
 215 solubility of  $\text{CO}_2$  relative to an entirely anhydrous melt. After a certain amount of  
 216  $\text{H}_2\text{O}$  is added, solubility decreases once more because the addition of  $\text{H}_2\text{O}$  to the  
 217 fluid phase causes the fugacity of  $\text{CO}_2$  to decrease (Mysen, 1976; Dingwell, 1986;  
 218 King & Holloway, 2002). More recently, this behavior has been demonstrated for  
 219 dacitic and rhyolitic melts by Behrens et al. (2004) and Liu et al. (2005).

220 The non-ideal behavior of  $\text{H}_2\text{O}$  and  $\text{CO}_2$  in basaltic and andesitic melts at  
 221 higher pressures is less well constrained. King and Holloway (2002) show that at  
 222 1 kbar, andesitic melts ( $\text{SiO}_2=58.4$  wt%) exhibit a sharp increase in the solubility  
 223 of  $\text{CO}_3^{2-}$  species with increasing melt  $\text{H}_2\text{O}$  contents between 0–3.39 wt%. In con-  
 224 trast, Jakobsson (1997) show that the solubility of  $\text{CO}_2$  in an icelanditic melt (54.6  
 225 wt%  $\text{SiO}_2$ ) at 10 kbar is essentially constant, despite variation in melt water con-  
 226 tents between  $\sim 1$ –9 wt%. Similar behaviour to that observed by Jakobsson (1997)  
 227 was noted for andesitic magmas (57.4 wt%  $\text{SiO}_2$ ) at 5 kbar by Botcharnikov et al.  
 228 (2006), who suggest that differences between  $\text{H}_2\text{O}$ - $\text{CO}_2$  mixing in their experiments  
 229 and those of King and Holloway (2002) may result from differences in oxygen fu-  
 230 gacity of these experiments ( $\text{Fe}^{3+}/\text{Fe}_T=0.2$ – $0.6$  vs.  $\text{Fe}^{3+}/\text{Fe}_T=0.09$ – $0.2$ ). However,  
 231 as we discuss in section 4.1, the influence of oxygen fugacity on volatile solubility  
 232 is still poorly constrained. Recent basaltic  $\text{H}_2\text{O}$ - $\text{CO}_2$  experiments generally show a  
 233 relatively flat plateau for  $\text{CO}_2$  solubility with increasing melt  $\text{H}_2\text{O}$  contents between  
 234  $\sim 0$ –4 wt% (Iacono-Marziano et al., 2012; Shishkina et al., 2010, 2014), although  
 235 relatively large error bars on analyses of  $\text{CO}_2$  in experimental products make it dif-  
 236 ficult to determine whether this plateau is truly flat, or shows a slight positive or  
 237 negative gradient at low  $\text{H}_2\text{O}$  contents (see Fig. 7a of Iacono-Marziano et al., 2012).  
 238 The effect of  $\text{H}_2\text{O}$ - $\text{CO}_2$  mixing in the 9 solubility models evaluated here is discussed  
 239 further in sections 4.2.2 and 4.3.2.



### 3 Models discussed in this review

In this review, we focus on the seven models implemented in VESIcal (Iacovino et al., 2021):

1. The simplified parameterization of the Dixon (1997) model for H<sub>2</sub>O and CO<sub>2</sub>, which was implemented in the popular Excel tool VolatileCalc (Newman & Lowenstern, 2002), hereafter VolatileCalc-Basalt.
2. The H<sub>2</sub>O model of G. Moore et al. (1998), hereafter M-1998.
3. The H<sub>2</sub>O and CO<sub>2</sub> models of Liu et al. (2005), hereafter L-2005.
4. The H<sub>2</sub>O and CO<sub>2</sub> models of Iacono-Marziano et al. (2012), hereafter IM-2012.
5. The H<sub>2</sub>O and CO<sub>2</sub> models of Shishkina et al. (2014), hereafter S-2014.
6. The combined H<sub>2</sub>O and CO<sub>2</sub> model of Ghiorso and Gualda (2015), hereafter MagmaSat.
7. The CO<sub>2</sub> models of Allison et al. (2019), hereafter A-2019.

We also consider the two additional models, reflecting both their popularity and relative ease of calculation using previously published tools:

8. The combined H<sub>2</sub>O and CO<sub>2</sub> model of Papale et al. (2006), hereafter P-2006, accessed using the Linux program solwcad.
9. The Rhyolite functions in the VolatileCalc spreadsheet, hereafter VolatileCalc-Rhyolite

We do not consider the models of X. Duan (2014), Eguchi and Dasgupta (2018), or Burgisser et al. (2015) because no tool exists to automate the necessary calculations. We also do not discuss models with more limited pressure (P), temperature (T) or compositional (X) ranges.

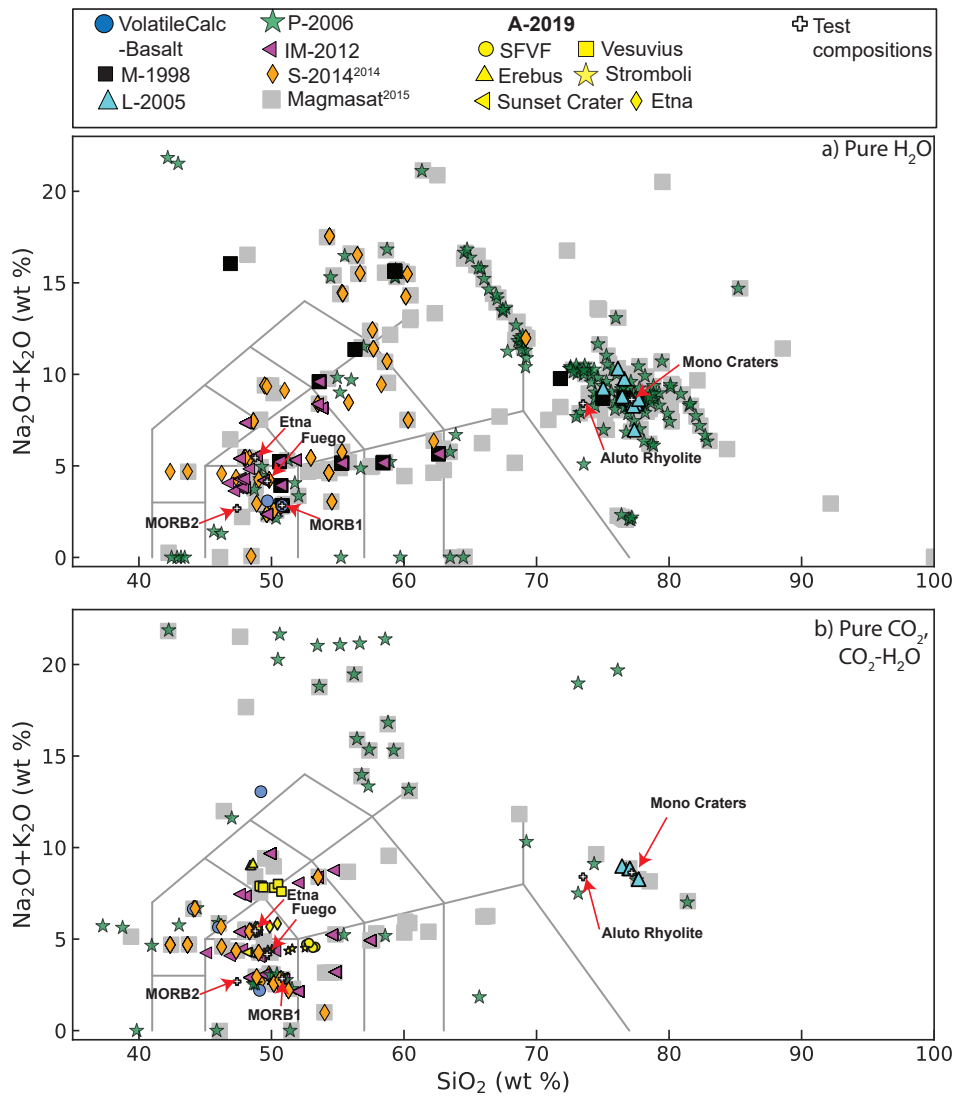
A summary of the P, T and X range covered by the calibration dataset of each of the 9 models evaluated, as well as available tools to perform calculations, is provided in Figure 1. The calibration dataset of each model is shown on a total alkali-silica diagram in Figure 2. Detailed descriptions of each model are provided below.

Publication	Volatile	Speciation	P (bar)	T (°C)	Notes
Moore et al., 1998	H <sub>2</sub> O	N/A	0–3000 <sup>1</sup>	700–1200 <sup>1</sup>	<sup>1</sup> Author-suggested calibration range. The calibration dataset spans 190 to 6067 bar, and 800–1200 °C
Newman and Lowenstern (2002) VolatileCalc–Basalt	H <sub>2</sub> O CO <sub>2</sub>	H <sub>2</sub> O as OH <sup>-</sup> and H <sub>2</sub> O <sub>mol</sub> CO <sub>2</sub> as CO <sub>3</sub> <sup>2-</sup>	0–5000 <sup>1</sup> 0–2000 <sup>2</sup> 0–1000 <sup>3</sup>	600–1500 <sup>1</sup> 1200 <sup>4</sup>	<sup>1</sup> Warnings implemented in VolatileCalc (Newman and Lowenstern, 2002). <sup>2</sup> Calibration range suggested by Lesne et al. (2011) <sup>3</sup> Calibration range suggested by Iacono–Marziano et al. (2012) <sup>4</sup> Calibration temperature of Dixon (1997) *if normalized (not recommended), different proportions of FeO and Fe <sub>2</sub> O <sub>3</sub> will slightly change the normalized SiO <sub>2</sub> content
Newman and Lowenstern (2002) VolatileCalc–Rhyolite	H <sub>2</sub> O CO <sub>2</sub>	H <sub>2</sub> O as OH <sup>-</sup> and H <sub>2</sub> O <sub>mol</sub>	0–5000 <sup>1</sup>	600–1500 <sup>1</sup>	<sup>1</sup> Calibration range warning implemented in VolatileCalc (Newman and Lowenstern, 2002).
Liu et al. 2005	H <sub>2</sub> O CO <sub>2</sub>	N/A	0–5000 <sup>1</sup>	700–1200 <sup>1</sup>	<sup>1</sup> Author-suggested calibration range for the mixed fluid model. The calibration dataset covers 750–5510 bar and 800–1150 °C for the Carbon model, and 1–5000 bar and 700–1200 °C for the water model
Papale et al. (2006)	H <sub>2</sub> O CO <sub>2</sub>	N/A	0–10,000 <sup>1</sup>	~630–1630 <sup>1</sup>	<sup>1</sup> Interception of H <sub>2</sub> O–CO <sub>2</sub> fields on Fig. 10 of Papale (1999)
Iacono–Marziano et al., 2012	H <sub>2</sub> O CO <sub>2</sub> H <sub>2</sub> O	N/A	100–10000 (mostly <5000) <sup>1</sup> 163–6067 <sup>1</sup>	1100–1400 (preferably 1200–1300) <sup>2</sup> 1000–1250 <sup>1</sup>	<sup>1</sup> Range of calibration dataset, as authors do not state a preferred range. We note that the vast majority of experiments were conducted at <5000 bar. <sup>2</sup> Authors state that most experiments were conducted between 1200–1300 °C (whole range 1100–1400 °C).
Shishkina et al. 2014	H <sub>2</sub> O <sup>1</sup> CO <sub>2</sub> <sup>1</sup>	N/A	0–5000 <sup>2</sup> 500–5000 <sup>2</sup>	1050–1400 (preferably 1150–1250) <sup>2,3</sup> 1200–1250 <sup>2,3</sup>	<sup>1</sup> Although their empirical expressions are for pure fluids, they were mostly calibrated on mixed CO <sub>2</sub> –H <sub>2</sub> O experiments. <sup>2</sup> Author-suggested range <sup>3</sup> Note, this model contains no temperature term.
Magmasat (Ghiorso and Gualda, 2015)	H <sub>2</sub> O CO <sub>2</sub> H <sub>2</sub> O–CO <sub>2</sub>	H <sub>2</sub> O: as OH <sup>-</sup> CO <sub>2</sub> : as CaCO <sub>3</sub> and CO <sub>2mol</sub>	0–20,000 <sup>1</sup> 0–30,000 <sup>1</sup> 0–10,000 <sup>1</sup>	550–1420 <sup>1</sup> 1139–1400 <sup>1</sup> 800–1400 <sup>1</sup>	<sup>1</sup> Ranges extracted from Fig. 2d of Ghiorso and Gualda, 2015
Allison et al., 2019	CO <sub>2</sub> <sup>1</sup>	N/A	0–7000 <sup>2</sup>	1200 <sup>3</sup> (~1000–1400)	<sup>1</sup> Although this model is for pure CO <sub>2</sub> , it was calibrated on mixed CO <sub>2</sub> –H <sub>2</sub> O experiments. <sup>2</sup> Author-suggested range. The calibration dataset spans: (SFVF-4133–6141 bar, Sunset Crater: 4071–6098 bar, Erebus: 4078–6175 bar, Vesuvius: 269–6175 bar, Etna=485–6199 bar, Stromboli=524–6080 bar). <sup>3</sup> Note, all calculations and experiments were performed at 1200 °C. Authors suggest applicable between 1000–1400 °C

Continued over page...

Publication	Compositional range	Formulation	Redox sensitive?	Available Tools
Moore et al., 1998	Broad compositional range: subalkaline basalts to rhyolites, alkaline trachybasalts–andesites, foidites, phonolites	Thermodynamic basis, empirical effect of melt composition	No (input FeO <sub>T</sub> )	<ul style="list-style-type: none"> <li>Macro-enabled Excel spreadsheet supplied by authors</li> <li>Python3 code (<b>VESICAL</b>; Iacovino et al. 2021)</li> </ul>
Newman and Lowenstern (2002) VolatileCalc–Basalt	Alkali basalts: 40–49 wt% SiO <sub>2</sub>	Thermodynamic basis, empirical effect of melt composition	No (input SiO <sub>2</sub> only)*	<ul style="list-style-type: none"> <li>Macro-enabled Excel spreadsheet (<b>VolatileCalc</b>; Newman and Lowenstern, 2002)</li> <li>Python2 code (Rasmussen et al. 2020)</li> <li>Python3 code and web application (<b>VESICAL</b>; Iacovino et al. 2021)</li> </ul>
Newman and Lowenstern (2002) VolatileCalc–Rhyolite	Rhyolites	Thermodynamic basis, no compositional term	No	<ul style="list-style-type: none"> <li>Macro-enabled Excel spreadsheet (<b>VolatileCalc</b>; Newman and Lowenstern, 2002)</li> </ul>
Liu et al. 2005	Haplogranites and rhyolites	Empirical basis, no compositional term	No	<ul style="list-style-type: none"> <li>Python3 code (<b>VESICAL</b>; Iacovino et al. 2021)</li> </ul>
Papale et al. (2006)	Broad compositional range (SiO <sub>2</sub> =37–85, Na <sub>2</sub> O+K <sub>2</sub> O=0–20) for H <sub>2</sub> O, more limited for CO <sub>2</sub> and CO <sub>2</sub> –H <sub>2</sub> O liquids. Poor coverage of intermediate compositions (SiO <sub>2</sub> =55–75 wt%).	Fully thermodynamic (formulation + effect of melt composition)	Yes (input FeO and Fe <sub>2</sub> O <sub>3</sub> )	<ul style="list-style-type: none"> <li>Web application, Linux application, and Fortran source code (SOLWCAD, hosted at <a href="http://www.pi.ingv.it/progetti/eurovolc/">http://www.pi.ingv.it/progetti/eurovolc/</a>)</li> </ul>
Iacono–Marziano et al., 2012	Predominantly mafic compositions: subalkaline and alkaline basalts–andesites	Thermodynamic basis, empirical effect of melt composition	No (input FeO <sub>T</sub> )	<ul style="list-style-type: none"> <li>Web application (<a href="http://calcul-isto.cnrs-orleans.fr/apbs/h2o-co2-systems/">http://calcul-isto.cnrs-orleans.fr/apbs/h2o-co2-systems/</a>)</li> <li>Python3 code (<b>VESICAL</b>; Iacovino et al. 2021)</li> </ul>
Shishkina et al. 2014	H <sub>2</sub> O: Mafic and intermediate compositions: Subalkaline basalts–basaltic andesites, alkali basanites–phonolites. SiO <sub>2</sub> >65 wt%. CO <sub>2</sub> : Predominantly mafic compositions: subalkaline basalts, alkaline basanites, trachybasalts	Fully empirical (formulation+ effect of melt composition)	Only for CO <sub>2</sub> (π* uses Fe <sup>2+</sup> ). Calibrated with Fe=Fe <sup>2+</sup>	<ul style="list-style-type: none"> <li>Python3 code (<b>VESICAL</b>; Iacovino et al. 2021), which uses Fe=Fe<sup>2+</sup></li> </ul>
Magmasat (Ghiorso and Gualda, 2015)	Very broad compositional range of natural silicate melt compositions: subalkaline picrobasalts–rhyolites, wide variety of mafic–silicic alkaline compositions	Fully thermodynamic (formulation + effect of melt composition)	Yes (input FeO and Fe <sub>2</sub> O <sub>3</sub> )	<ul style="list-style-type: none"> <li>Web application (<a href="http://melts.ofm-research.org/CORBA_CTserver/GG-H2O-CO2.html">http://melts.ofm-research.org/CORBA_CTserver/GG-H2O-CO2.html</a>)</li> <li>Mac App Store (MagmaSat)</li> <li>Python3 code (<b>VESICAL</b>; Iacovino et al. 2021)</li> </ul>
Allison et al., 2019	Alkali–rich mafic magmas from 6 volcanic fields (San Francisco Volcanic Field, Sunset Crater, Erebus, Vesuvius, Etna, Stromboli). Separate model coefficients for each composition.	Thermodynamic basis, separate parameters for each of 6 melt compositions.	No	<ul style="list-style-type: none"> <li>Excel spreadsheet supplied by authors</li> <li>Python3 code (<b>VESICAL</b>; Iacovino et al. 2021)</li> </ul>

**Figure 1.** Summary of the calibration range of each model, as well as available tools to perform calculations. Models are colored using the same palette used for comparison figures.



**Figure 2.** Total alkalis (Na<sub>2</sub>O+K<sub>2</sub>O) vs. SiO<sub>2</sub> (TAS) diagram showing the composition of melts in the calibration dataset of each model for a) pure H<sub>2</sub>O and b) pure CO<sub>2</sub> and mixed CO<sub>2</sub>-H<sub>2</sub>O. Pure CO<sub>2</sub> and CO<sub>2</sub>-H<sub>2</sub>O experiments are combined because pure CO<sub>2</sub> experiments are often contaminated by variable amounts of H<sub>2</sub>O due to exposure with the earth's atmosphere, and the high mobility of H<sup>+</sup> through experimental apparatus (Mangan et al., 2021). As dissolved H<sub>2</sub>O contents in glasses from pure CO<sub>2</sub> experiments are rarely reported, it is nontrivial to distinguish these from a mixed H<sub>2</sub>O-CO<sub>2</sub> experiment. For P-2006 and MagmaSat, points were extracted from the TAS diagrams shown in these papers using Web Plot Digitizer (Rohatgi, 2017). For the other models, the calibration dataset is provided in the supplementary information of Iacovino et al. (2021). TAS plot drawn using Python code from Stevenson (2015).

269 **3.1 VolatileCalc-Basalt: a simplification of Dixon (1997)**

270 The Dixon (1997) model calculates the solubility of H<sub>2</sub>O and CO<sub>2</sub> in basaltic  
 271 silicate melts, combining thermodynamic expressions as a function of pressure and  
 272 temperature described in Dixon et al. (1995) with empirical parameters accounting  
 273 for the effect of melt composition in terms of melt SiO<sub>2</sub> content from Dixon (1997).  
 274 The thermodynamic expressions are originally from Fine and Stolper (1986) for car-  
 275 bon, and Silver and Stolper (1989) for water. The Dixon (1997) model considers the  
 276 solubility of the carbonate ion (CO<sub>3</sub><sup>2-</sup>) for CO<sub>2</sub>, and both molecular water (H<sub>2</sub>O<sub>mol</sub>)  
 277 and hydroxyl groups (OH<sup>-</sup>) for H<sub>2</sub>O.

The solubility of molecular H<sub>2</sub>O is calculated using an adapted version of equa-  
 tion 3 of Dixon et al. (1995). In the original 1995 equation the X<sub>H<sub>2</sub>O<sub>mol</sub></sub><sup>m</sup>(P<sub>0</sub>, T<sub>0</sub>)  
 term representing the mole fraction of molecular H<sub>2</sub>O in equilibrium with fluid with  
 a fugacity of water specified by  $f_{H_2O}(P_0, T_0)$  at 1473.15 K and 1 bar was fixed at  
 $3.28 \times 10^{-5}$ . To account for the effect of melt composition on H<sub>2</sub>O solubility, Dixon  
 (1997) parametrize this term as a function of melt SiO<sub>2</sub> content:

$$X_{H_2O_{mol}}^m(P_0, T_0) = 3.04 \times 10^{-5} + 1.29 \times 10^{-6} [SiO_2]^{wt\%} \quad (1)$$

278 This relationship was derived from experimental observations of Cocheo and  
 279 Holloway (1993), and predicts that there is a ~30% increase in the solubility of H<sub>2</sub>O  
 280 with increasing SiO<sub>2</sub> between nephelinite (~40 wt% SiO<sub>2</sub>) and tholeiite (~49 wt%  
 281 SiO<sub>2</sub>) melt compositions. The concentration of OH<sup>-</sup> is then calculated as a function  
 282 of the mole fraction of molecular H<sub>2</sub>O in the melt using the solution model of Silver  
 283 and Stolper (1989) (see equation 4 of Dixon, 1997). Interestingly, this is the only  
 284 model discussed here which considers more than one species for dissolved H<sub>2</sub>O in the  
 285 melt.

For CO<sub>2</sub> solubility, Dixon (1997) adapted the model of Dixon et al. (1995) to  
 account for the effect of melt composition, based on observations from experiments  
 that CO<sub>2</sub> solubility increases from tholeiitic (49 wt% SiO<sub>2</sub>) to basaltic (46 wt%  
 SiO<sub>2</sub>) to leucitic (44.1 wt% SiO<sub>2</sub>) melts at 1200°C, 1 kbar. A linear regression with  
 CO<sub>2</sub> solubility was achieved using a composition parameter (II) expressed in terms

of the cation fractions,  $X_i$  (Dixon, 1997):

$$\Pi = -6.50(X_{\text{Si}^{4+}} + X_{\text{Al}^{3+}}) + 20.17(X_{\text{Ca}^{2+}} + 0.8X_{\text{K}^+} + 0.7X_{\text{Na}^+} + 0.4X_{\text{Mg}^{2+}} + 0.4X_{\text{Fe}^{2+}}) \quad (2)$$

However, based on the strong correlation between  $\Pi$  and  $\text{SiO}_2$  in a suite of lavas from the North Arch Volcanic Field, Dixon (1997) express the  $X_{\text{CO}_3}^m(P_0, T_0)$  term from equation 6 of Dixon et al. (1995) solely as a function of melt  $\text{SiO}_2$  content:

$$X_{\text{CO}_3}^m(P_0, T_0) = 8.70 \times 10^{-6} - 1.70 \times 10^{-7}[\text{SiO}_2]^{\text{wt}\%} \quad (3)$$

286 Where  $X_{\text{CO}_3}^m(P_0, T_0)$  is the mole fraction of carbonate in equilibrium with fluid  
 287 with a fugacity of carbon dioxide specified by  $f_{\text{CO}_2}(P_0, T_0)$  at 1473.15 K and 1 bar.  
 288 Fugacities are calculated using the Redlich-Kwong equation of state (Holloway,  
 289 1977), with the correction of Flowers (1979).

290 This simplified expression was designed to aid the investigations of volatile  
 291 solubility in the suite of lavas from the North Arch, where it effectively captures  
 292 the observed  $5\times$  decrease in  $\text{CO}_2$  solubility from 40 to 49 wt%  $\text{SiO}_2$ . However, this  
 293 simplified parameterization became very widely used in a wide variety of tectonic  
 294 settings following its implementation in the Excel-based tool VolatileCalc (New-  
 295 man and Lowenstern, 2002). Here, we refer to this model as VolatileCalc-Basalt, to  
 296 differentiate it from the full  $\Pi$  parameterization of Dixon (1997).

297 The advantage of the  $\Pi$ - $\text{SiO}_2$  simplification is that users only have to input the  
 298 concentration of one oxide component (melt  $\text{SiO}_2$ ) in addition to melt temperature  
 299 and melt volatile contents to calculate saturation pressures or degassing paths. The  
 300 limited number of inputs required by this model meant that users can calculate sat-  
 301 uration pressures for large numbers of melt inclusions relatively quickly compared  
 302 to more recent models such as Iacono-Marziano et al. (2012) and MagmaSat (which  
 303 require users to input 8–14 oxide concentrations).

304 However, extreme care must be taken when using this simplified model to  
 305 calculate  $\text{CO}_2$  solubility. Firstly, the North Arch lavas span  $\text{SiO}_2$  contents of only  
 306 40–49 wt%. Because of the rapid drop in  $\Pi$  with increasing  $\text{SiO}_2$ , extrapolation  
 307 beyond 51.2 wt%  $\text{SiO}_2$  yields a negative value for  $X_{\text{CO}_3}^m(P_0, T_0)$ , corresponding  
 308 to a negative amount of dissolved  $\text{CO}_2$ . To avoid this issue, VolatileCalc-Basalt re-  
 309 turns an error, and will not perform the calculation if users enter a  $\text{SiO}_2$  content

310 >49 wt%. Most studies extrapolate beyond this by simply entering  $\text{SiO}_2=49$  wt%  
311 into VolatileCalc-Basalt if their melts have higher silica contents (e.g. Sides et al.,  
312 2014a, 2014b; Tucker et al., 2019), and this approach is implemented in VESIcal for  
313 consistency. Newman and Lowenstern (2002) suggest that this extrapolation will be  
314 “generally applicable for other basaltic rocks with  $<52$  wt%  $\text{SiO}_2$ ”. However, if a  
315 large proportion of a sample suite has  $\text{SiO}_2$  contents with  $>49$  wt%  $\text{SiO}_2$ , the simpli-  
316 fied  $\text{II-SiO}_2$  parameterization treats all melts as if they have the same composition,  
317 neglecting variations in solubility that may exist within that suite (see Wieser et al.,  
318 2021). Additionally, even if samples have  $\text{SiO}_2$  contents between 40–49 wt%, this  
319 simplification can yield spurious results for melts which do not follow the same trend  
320 in  $\text{II-SiO}_2$  space to the North Arch lavas (see section 4.2.3). Thus, we suggest that  
321 any users wishing to apply VolatileCalc-Basalt to their system first check whether  
322 their melt compositions lie close to the trend defined by the North Arch lavas using  
323 the Jupyter notebook provided in the supplementary information of Iacovino et al.  
324 (2021).

325 It is also worth noting that, because VolatileCalc-Basalt parameterizes the  
326 effect of melt composition in terms of the absolute concentration of  $\text{SiO}_2$  (rather  
327 than other empirical models which use cation fractions), it is extremely sensitive to  
328 normalization. For example, consider the MORB2 composition in Table 1 which has  
329 a measured  $\text{SiO}_2$  content of 47.4 wt%, and an anhydrous total of 97.375 wt%. For  
330 1000 ppm  $\text{CO}_2$  and  $T=1200^\circ\text{C}$ , using raw data ( $\text{SiO}_2=47.4$  wt%) the calculated sat-  
331 uration pressure is 1206 bars. However, it is relatively common in the literature that  
332 major elements (excluding volatiles) are scaled to sum to 100%, while volatile con-  
333 centrations are left unchanged. This would cause the melt  $\text{SiO}_2$  content to increase  
334 to 48.68 wt%  $\text{SiO}_2$ , corresponding to a saturation pressure of 1765 bar respectively  
335 ( $1.5\times$  higher!). We encourage users not to normalize their data, as we note that the  
336  $\text{II-SiO}_2$  plot of North Arch Glasses in Dixon (1997) is best recreated using unnor-  
337 malized data (see Supporting Fig. S1). Unnormalized data is used throughout this  
338 review for all VolatileCalc-Basalt outputs, which gives results comparable to those  
339 produced in the VolatileCalc-Basalt spreadsheet when users enter the  $\text{SiO}_2$  contents  
340 given in Table 1.

341

### 3.2 VolatileCalc-Rhyolite: Newman and Lowenstern, 2002

342

343

344

345

346

347

348

349

350

351

352

353

354

355

356

357

358

359

360

361

In addition to the functionality for basalts described above, the VolatileCalc spreadsheet also allows users to calculate saturation pressures, degassing paths, isobars and isopleths for rhyolitic magmas (hereafter, VolatileCalc-Rhyolite). For CO<sub>2</sub>, VolatileCalc-Rhyolite uses the simple thermodynamic model from Stolper et al. (1987) that was later applied to rhyolitic melts by Fogel and Rutherford (1990). The molar enthalpy change for CO<sub>2</sub> dissolution in the melt is from Fogel and Rutherford (1990), the single-O melt mass from Silver et al. (1990), and the CO<sub>2</sub> molar volume and solubility at standard state from Blank et al. (1993). The thermodynamic basis for the H<sub>2</sub>O model is the same as that used in VolatileCalc-Basalt. The fitted parameters for H<sub>2</sub>O solubility in the standard state is from Silver (1988), and the single-O melt mass and molar enthalpy change for H<sub>2</sub>O dissolution in the melt from Silver et al. (1990). The partial molar volume of H<sub>2</sub>O was adjusted to 5 cm<sup>3</sup>/mol to provide a better fit to experimental data. There are two main differences of the rhyolite model relative to the basaltic model. First, while both the models use a temperature-dependent equation of state, VolatileCalc-Rhyolite model also contains a term for the heat of solution of volatile solubility, so is far more sensitive to temperature (see section 5). Second, unlike VolatileCalc-Basalt which require users to enter melt SiO<sub>2</sub> contents, VolatileCalc-Rhyolite is independent of melt composition. Thus, Newman and Lowenstern (2002) caution that this model may not be applicable for strongly peralkaline or peraluminous rhyolites.

362

### 3.3 M-1998 (Moore et al., 1998)

The Moore et al. (1998) model calculates the solubility of H<sub>2</sub>O for a wide range of silicate melt compositions using an empirical expression valid between 700–1200°C and 0–3000 bars:

$$2 \ln(X_{\text{H}_2\text{O}}^{\text{melt}}) = \frac{a}{T} + \sum_i b_i X_i \frac{P}{T} + c \ln(f_{\text{H}_2\text{O}}^{\text{fluid}}) + d \quad (4)$$

363

364

365

366

Where  $X_{\text{H}_2\text{O}}^{\text{melt}}$  is the mole fraction of H<sub>2</sub>O dissolved in the melt, T is the temperature in Kelvin, P is the pressure in bars, and  $X_i$  is the anhydrous molar fraction of each oxide component.  $f_{\text{H}_2\text{O}}^{\text{fluid}}$  is the fugacity of H<sub>2</sub>O in the fluid, calculated using the modified version of the Redlich-Kwong equation of state provided in the ap-



pendix of Holloway and Blank (1994). Equation 4 is associated with the following fit parameters ( $\pm$  standard error):

Coefficient	a	$b_{Al_2O_3}$	$b_{FeO_r}$	$b_{Na_2O}$	c	d
Value ( $\pm 1\sigma$ )	2565 ( $\pm 362$ )	-1.997 ( $\pm 0.706$ )	-0.9275 ( $\pm 0.394$ )	2.736 ( $\pm 0.871$ )	1.171 ( $\pm 0.069$ )	-14.21 ( $\pm 0.54$ )

As equation 4 includes a term for the fugacity of H<sub>2</sub>O in the fluid, this model can be integrated with CO<sub>2</sub> models implemented in VESIcal (e.g., the II-SiO<sub>2</sub> simplification of Dixon, 1997, Liu et al., 2005, Iacono-Marziano et al., 2012, Shishkina et al., 2014, Allison et al., 2019) to investigate mixed H<sub>2</sub>O-CO<sub>2</sub> fluids.

The model calibration dataset combines the authors' pure H<sub>2</sub>O experiments with literature data, spanning sub-alkaline basaltic to rhyolitic compositions, as well as some alkaline compositions (Fig. 2). As with other fully-empirical models implemented in VESIcal, or those including empirical expressions, extreme care must be taken when extrapolating this model outside of the calibration range. In particular, the authors warn against extrapolating this model to pressures exceeding 3 kbar, in part due to the complexities of the critical behavior of fluids at higher pressures.

### 3.4 L-2005 (Liu et al., 2005)

The Liu et al. (2005) model calculates the solubility of H<sub>2</sub>O and CO<sub>2</sub> in metaluminous, high-silica rhyolitic melts using empirical expressions, valid between 700–1200°C, and 0–5000 bars. The following expression is used to calculate CO<sub>2</sub> solubility:

$$[CO_2]^{ppm} = \frac{P_{CO_2}(b_1 + b_4 P_W)}{T} + P_{CO_2}(b_2 P_W^{0.5} + b_3 P_W^{1.5}) \quad (5)$$

T is temperature in Kelvin,  $b_1$ – $b_4$  are fit parameters, and the  $P_W$  and  $P_{CO_2}$  terms account for the partial pressures of each volatile species in the co-existing fluid, with:

$$P_{CO_2} = X_{CO_2}^f P \quad (6)$$

$$P_W = X_W^f P \quad (7)$$

Where P is pressure in MPa,  $X_{CO_2}^f$  is the mole fraction of CO<sub>2</sub> in the fluid, and  $X_W^f$  is the mole fraction of H<sub>2</sub>O in the fluid. These empirical terms mean that no equation of state is used (unlike M-1998 and VolatileCalc-Basalt). The fit parameters associated with equation 5 are shown below ( $\pm$  error):

Coefficient	$b_1$	$b_2$	$b_3$	$b_4$
Value (Error)	5668 ( $\pm 127$ )	0.4133 ( $\pm 0.0491$ )	$2.041 \times 10^{-3}$ ( $\pm 0.285 \times 10^{-3}$ )	-55.99 ( $\pm 8.36$ )

386

Similarly, they provide the following expression for H<sub>2</sub>O:

$$[\text{H}_2\text{O}_t]_{\text{wt}\%} = \frac{a_1 P_w^{0.5} + a_2 P_w + a_3 P_w^{1.5}}{T} + a_4 P_w^{1.5} + P_{\text{CO}_2} (a_5 P_w^{0.5} + a_6 P_w) \quad (8)$$

387

Using the following fit parameters:

Coefficient	$a_1$	$a_2$	$a_3$	$a_4$
Value (Error)	354.94 ( $\pm 4.55$ )	9.623 ( $\pm 0.923$ )	-1.5223 ( $\pm 0.0722$ )	0.0012439 ( $\pm 0.0000499$ )
	$a_5$		$a_6$	
	-1.084 $\times 10^{-4}$ ( $\pm 0.406 \times 10^{-4}$ )		-1.362 $\times 10^{-5}$ ( $\pm 0.352 \times 10^{-5}$ )	

388

389

390

391

392

393

394

395

The model calibration dataset combines pure H<sub>2</sub>O solubility experiments by the authors between  $\sim 1$  and 250 bars with literature experiments investigating the solubility of pure H<sub>2</sub>O, H<sub>2</sub>O-H<sub>2</sub>, CO<sub>2</sub> and CO<sub>2</sub>-H<sub>2</sub>O fluids, spanning significantly higher pressures (up to 5000 bars). Unlike the M-1998 model, their empirical expressions do not incorporate a term for melt composition, so care is needed when applying this model to melts with different major element compositions to the calibration dataset (Fig. 2).

396

### 3.5 P-2006 (Papale et al., 2006)

397

398

399

400

401

402

403

404

405

406

407

408

409

410

Papale et al. (2006) present a fully non-ideal thermodynamic model for H<sub>2</sub>O and CO<sub>2</sub> solubility, which is a recalibration of the earlier models of Papale (1999) and Papale (1997). This updated model capitalizes on the large amount of volatile solubility experiments performed between 1997 and 2005, which nearly doubled the size of the calibration dataset, and allowed experimental data on CO<sub>2</sub> solubility collected prior to 1980 to be discarded (removing systematic errors associated with different analytical techniques, see Papale, 1999). Unlike the models discussed above which are calibrated on a specific subregion of compositional space and use empirical parametrizations to account for the effect of melt composition, the models of Papale et al. (2006) and Papale (1999) treat the composition of the silicate liquid using a thermodynamic approach based on Ghiorso et al. (1983). Papale et al. (2006) note that this thermodynamic approach means that for any specific region of composition space (e.g., comparing model results to a specific experiment), the fit may not be as good as an empirical model tuned to that composition. However, carefully calibrated

411 thermodynamic models will be significantly more successful than empirical models  
 412 when applied to melts which are not represented in the calibration dataset.

P-2006 considers a silicate liquid in mechanical, thermal and chemical equilibrium with a fluid phase containing H<sub>2</sub>O and CO<sub>2</sub>. The model uses the modified Redlich-Kwong equation of state of Kerrick and Jacobs (1981) to describe the fluid phase, and considers only the dissolution of CO<sub>2</sub> and H<sub>2</sub>O in the melt (while natural silicate melts contain molecular CO<sub>2</sub> and CO<sub>3</sub><sup>2-</sup> species, and molecular H<sub>2</sub>O and OH<sup>-</sup> species). The model calculates the Gibbs free energy of mixing, considering 10 major oxide components in addition to CO<sub>2</sub> and H<sub>2</sub>O. Binary interaction coefficients, denoted by  $w_{ij}$ , account for the attractive-repulsive behaviour between an oxide and volatile component. For example,  $w_{CO_2, MgO}$  describes the interaction of MgO with CO<sub>2</sub>. Interaction coefficients for CO<sub>2</sub> are expressed as a function of pressure (relative to a reference pressure of 0.1 MPa) requiring two coefficients, while those for H<sub>2</sub>O are invariant of pressure (requiring 1 coefficient):

$$w_{CO_2,i} = w_{CO_2,i}^0 + w_{CO_2,i}^1 \ln \frac{P}{P_0} \quad (9)$$

$$w_{H_2O,i} = w_{H_2O,i}^0 \quad (10)$$

413 P-2006 uses interaction terms for SiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, MgO, CaO, Na<sub>2</sub>O, K<sub>2</sub>O, FeO and  
 414 Fe<sub>2</sub>O<sub>3</sub>. The presence of two Fe terms means that the model is sensitive to melt re-  
 415 dox. Papale et al. (2006) show that the inclusion of  $w$  terms for MnO and TiO<sub>2</sub> lead  
 416 to overfitting, so the effect of these oxides on model outputs is only through the di-  
 417 lution of the concentration of components allocated  $w$  terms. The values of the 24  $w$   
 418 terms, as well as 5 terms accounting for molar volumes and fugacities, are calculated  
 419 from a calibration dataset comprising  $\sim$  1100 solubility experiments with pure CO<sub>2</sub>,  
 420 pure H<sub>2</sub>O and mixed CO<sub>2</sub>-H<sub>2</sub>O fluids. While the calibration dataset contains well  
 421 populated clusters for basaltic and rhyolitic compositions, intermediate compositions  
 422 and basaltic melts with high alkali contents are poorly represented, particularly for  
 423 CO<sub>2</sub> (Fig. 2).

424 Papale et al. (2006) demonstrate that despite the addition of hundreds of new  
 425 experimental datapoints for H<sub>2</sub>O, there are no significant changes in coefficients  
 426 compared to those which were published with their 1999 model. The percent errors  
 427 on the H<sub>2</sub>O coefficients in the 2006 model are  $< 10\%$  for all species (defined as  $100^* \sigma / \text{coefficient}$ ).  
 428 In contrast, the addition of new CO<sub>2</sub> data to the calibration dataset

429 resulted in significant changes in coefficients, and the percentage errors on these  
 430 coefficients in the 2006 model remained large ( $\sim 800\%$  for FeO,  $\sim 150\%$  for Na<sub>2</sub>O,  
 431  $\sim 190\%$  for MgO, Fig. 3). Based on these large error values, these coefficients would  
 432 likely change again if this model was recalibrated to include all new CO<sub>2</sub> experi-  
 433 ments published since 2006.

434 The pressure-dependence of the CO<sub>2</sub> melt interaction terms, combined with  
 435 the fact that the  $w_o$  and  $w_1$  terms have different signs for all oxides except FeO,  
 436 means that a given change in melt chemistry may cause an increase in CO<sub>2</sub> solu-  
 437 bility at one pressure, but a decrease at another pressure (see Section 4.1). Ghiorso  
 438 and Gualda (2015) note that the coefficient for the compressibility of CO<sub>2</sub> in the  
 439 P-2006 model is negative, which is physically impossible (implying the volume of the  
 440 CO<sub>2</sub> fluid increases when pressure is increased), which they suggest may arise from  
 441 the inclusion of pressure-dependent  $w$ -terms.

### P-2006

Element	CO <sub>2</sub>				H <sub>2</sub> O	
	$w_o$	% error (1 $\sigma$ )	$w_1$	% error	$w_o$	% error
SiO <sub>2</sub>	-59,962	18	6049	19	-34,093	1.85
Al <sub>2</sub> O <sub>3</sub>	-590,957	32	41,395	47	-189,117	2.54
Fe <sub>2</sub> O <sub>3</sub>	4,469,623	9	-529,301	9	135,935	9.32
FeO	21,666	806	1214	1500	-195,751	3.13
MgO	52,866	189	-13,446	78	-86,418	7.06
CaO	-328,792	23	12,789	63	-209,997	1.67
Na <sub>2</sub> O	140,034	146	-35,213	60	-322,253	1.42
K <sub>2</sub> O	309,070	48	-58,010	27	-349,798	1.79

### MagmaSat

Element	CO <sub>2</sub> -CO <sub>3</sub> <sup>2-</sup>				H <sub>2</sub> O	
	W (CO <sub>2</sub> )	% error (1 $\sigma$ )	W (CaCO <sub>3</sub> )	% error	W	% error
SiO <sub>2</sub>			63.281	2.52	27.557	0.065
TiO <sub>2</sub>	-19.266	1.34	-79.203	0.46	88.199	2.87
Al <sub>2</sub> O <sub>3</sub>			46.716	2.52	11.768	21.8
Fe <sub>2</sub> O <sub>3</sub>	-3.187	4.3	65.509	0.26	50.105	17.0
Fe <sub>2</sub> SiO <sub>4</sub>	-32.465	0.95	-72.997	0.40	30.936	18.8
Mg <sub>2</sub> SiO <sub>4</sub>	-40.854	1.6	-24.873	4.17	20.910	21.5
CaSiO <sub>3</sub>	30.012	4.8	37.534	2.70	9.715	27.5
Na <sub>2</sub> SiO <sub>3</sub>			-311.011	0.24	-82.460	2.9
KAlSiO <sub>4</sub>			-27.865	8.21	1.057	112
Ca <sub>3</sub> (PO <sub>4</sub> ) <sub>2</sub>	-3.473	0.86	37.534	0.13	44.133	0.76

**Figure 3.** Interaction coefficients for P-2006 and MagmaSat. Percentage errors calculated as  $100 \times 1\sigma / \text{coefficient}$ . Error are colored green if they are  $< 10\%$ , light pink if  $10\text{--}25\%$ , and red if  $> 25\%$ .

442

### 3.6 IM-2012 (Iacono-Marziano et al. 2012)

The Iacono-Marziano et al. (2012) model expresses the solubility of H<sub>2</sub>O and CO<sub>2</sub> in mafic melts by combining simplified thermodynamic expressions for melt-fluid thermodynamics with empirical formulations accounting for melt composition. For CO<sub>2</sub>, they present the following expression:

443

444

$$\ln[\text{CO}_2]^{\text{ppm}} = X_{\text{H}_2\text{O}}d_{\text{H}_2\text{O}} + X_{\text{Al}}d_{\text{Al}} + X_{\text{FeO}+\text{MgO}}d_{\text{FeO}+\text{MgO}} + X_{\text{Na}_2\text{O}+\text{K}_2\text{O}}d_{\text{Na}_2\text{O}+\text{K}_2\text{O}} + a_{\text{CO}_2}\ln[\text{P}_{\text{CO}_2}] + b_{\text{CO}_2}\left[\frac{\text{NBO}}{\text{O}}\right] + B_{\text{CO}_2} + C_{\text{CO}_2}\frac{\text{P}}{\text{T}} \quad (11)$$

Where P is the pressure in bars, T is the temperature in Kelvin, X<sub>H<sub>2</sub>O</sub> is the molar fraction of H<sub>2</sub>O in the melt and P<sub>CO<sub>2</sub></sub> is the partial pressure of CO<sub>2</sub> in bars. The partial pressure of CO<sub>2</sub> is calculated from the pressure multiplied by the mole fraction of CO<sub>2</sub> in the fluid. This means that this model does not rely on an equation of state (as with the L-2005 model). The other terms account for the effect of melt composition using molar fractions calculated on a hydrous basis; X<sub>FeO+MgO</sub> is the sum of molar fractions of FeO<sub>t</sub> and MgO, X<sub>Na<sub>2</sub>O+K<sub>2</sub>O</sub> is the sum of the molar fractions of Na<sub>2</sub>O and K<sub>2</sub>O, and X<sub>Al</sub> is the aluminic index (AI):

$$X_{\text{AI}} = \frac{X_{\text{Al}_2\text{O}_3}}{X_{\text{CaO}} + X_{\text{K}_2\text{O}} + X_{\text{Na}_2\text{O}}} \quad (12)$$

The NBO/O term represents the number of non-bridging oxygens divided by oxygen, expressing the availability of oxygen to form carbonate groups within the melt. NBO/O can be calculated from mol fraction of different oxides, X<sub>i</sub>, on an anhydrous or hydrous basis:

$$\frac{\text{NBO}}{\text{O}}^{\text{Anhyd}} = \frac{2(X_{\text{K}_2\text{O}} + X_{\text{Na}_2\text{O}} + X_{\text{CaO}} + X_{\text{MgO}} + X_{\text{FeO}} - X_{\text{Al}_2\text{O}_3})}{2X_{\text{SiO}_2} + 2X_{\text{TiO}_2} + 3X_{\text{Al}_2\text{O}_3} + X_{\text{MgO}} + X_{\text{FeO}} + X_{\text{CaO}} + X_{\text{Na}_2\text{O}} + X_{\text{K}_2\text{O}}} \quad (13)$$

$$\frac{\text{NBO}}{\text{O}}^{\text{Hyd}} = \frac{2(X_{\text{H}_2\text{O}} + X_{\text{K}_2\text{O}} + X_{\text{Na}_2\text{O}} + X_{\text{CaO}} + X_{\text{MgO}} + X_{\text{FeO}} - X_{\text{Al}_2\text{O}_3})}{2X_{\text{SiO}_2} + 2X_{\text{TiO}_2} + 3X_{\text{Al}_2\text{O}_3} + X_{\text{MgO}} + X_{\text{FeO}} + X_{\text{CaO}} + X_{\text{Na}_2\text{O}} + X_{\text{K}_2\text{O}} + X_{\text{H}_2\text{O}}} \quad (14)$$

445

446

447

448

In both cases, mole fractions are calculated on a hydrous basis (Iacono-Marziano, written comms). Iacono-Marziano give coefficients for equation 11 for both cases ( $\pm 2\sigma$ ), leading to two forms of this model: IM-2012-A (anhydrous) and IM-2012-H (hydrous):

Coefficient	d <sub>H<sub>2</sub>O</sub>	d <sub>Al</sub>	d <sub>FeO+MgO</sub>	d <sub>Na<sub>2</sub>O+K<sub>2</sub>O</sub>	a <sub>CO<sub>2</sub></sub>	b <sub>CO<sub>2</sub></sub>	c <sub>CO<sub>2</sub></sub>	B <sub>CO<sub>2</sub></sub>
<b>Hydrous (±error)</b>	-16.4 (± 1.2)	4.4 (±0.4)	-17.1 (± 0.9)	22.8 (± 1.1)	1 (± 0.03)	17.3 (± 0.9)	0.12 (±0.02)	-6 (±0.4)
<b>Anhydrous (±error)</b>	2.3 (± 0.5)	3.8 (±0.4)	-16.3 (± 0.9)	20.1 (± 1.1)	1 (± 0.03)	15.8 (± 0.9)	0.14 (±0.02)	-5.3 (±0.4)

449

450 We note for completeness that in the original publication, equation 11 was  
 451 incorrectly expressed in terms of  $\ln[\text{CO}_3^{2-}]$  (Iacono-Marziano, written. comms).

452 For  $\text{H}_2\text{O}$ , the authors state that it is statistically unjustified to include  $d_i$   
 453 terms similar to those in the  $\text{CO}_2$  expression, due to the relatively small effect of  
 454 melt composition on  $\text{H}_2\text{O}$  solubility. The effect of melt composition is incorporated  
 455 only through the NBO/O term:

$$\ln [\text{H}_2\text{O}]^{\text{wt}\%} = a_{\text{H}_2\text{O}} \ln[\text{P}_{\text{H}_2\text{O}}] + b_{\text{H}_2\text{O}} \left[ \frac{\text{NBO}}{\text{O}} \right] + B_{\text{H}_2\text{O}} + C_{\text{H}_2\text{O}} \frac{\text{P}}{\text{T}} \quad (15)$$

456 Where  $\text{P}_{\text{H}_2\text{O}}$  is the partial pressure of  $\text{H}_2\text{O}$  in bars. As for  $\text{CO}_2$ , coefficients are  
 457 provided for NBO/O calculated on a hydrous and anhydrous basis. The hydrous co-  
 458 efficients in the published paper differ from those used in the web app over the last  
 459 decade (although a new web app using the published coefficients appeared briefly in  
 460 2021). VESICAL uses the web app hydrous coefficients by default, as recommended  
 461 by Iacono-Marziano (written. comms). The coefficients in the published paper were  
 462 from an older version of the model, and predict extremely high  $\text{H}_2\text{O}$  solubility at  
 463  $\sim 10$  kbar ( $>100$  wt%).

Coefficient	$a_{\text{H}_2\text{O}}$	$b_{\text{H}_2\text{O}}$	$B_{\text{H}_2\text{O}}$	$C_{\text{H}_2\text{O}}$
<b>Hydrous (<math>\pm</math>error)</b>	0.53 ( $\pm$ 0.02)	2.35 ( $\pm$ 0.28)	-3.37 ( $\pm$ 0.13)	-0.02 ( $\pm$ 0.02)
<b>Anhydrous (<math>\pm</math>error)</b>	0.54 ( $\pm$ 0.02)	1.24 ( $\pm$ 0.28)	-2.95 ( $\pm$ 0.17)	0.02 ( $\pm$ 0.02)
<b>Web App</b>	0.52096846	2.11575907	-3.24443335	0.02238884

465 The authors state that the differences between calculations performed with  
 466 NBO/O calculated on a hydrous and anhydrous basis are relatively small, but that a  
 467 slightly better fit to experimental data is obtained using the hydrous model (particu-  
 468 larly for  $\text{H}_2\text{O}$ -rich, and  $\text{CO}_2$ -poor melts). For completeness, we perform calculations  
 469 using both versions (referred to as IM-2012-H and IM-2012-A). Interestingly, we  
 470 show that the anhydrous version is most similar to other models for MORB-like  
 471 compositions.

472 The calibration dataset for  $\text{CO}_2$  combines the authors experiments with those  
 473 from a variety of literature studies for mixed  $\text{H}_2\text{O}$ - $\text{CO}_2$  fluids, spanning tempera-  
 474 tures between 1100 and 1400°C, and pressures between 100 and 10,000 bars (but  
 475 mostly  $<5000$  bars). The calibration dataset for  $\text{H}_2\text{O}$  incorporates pure  $\text{H}_2\text{O}$  exper-  
 476 iments from the literature (spanning 163–6067 bars, and 1000–1250°C), as well as

477 the experiments on H<sub>2</sub>O-CO<sub>2</sub> fluids used to calibrate the CO<sub>2</sub> expression. Melt com-  
 478 positions are predominantly mafic, spanning subalkaline-alkaline basalts to basaltic  
 479 andesites (45–57 wt% SiO<sub>2</sub> for mixed H<sub>2</sub>O-CO<sub>2</sub> experiments, and 46–63 wt% SiO<sub>2</sub>  
 480 for pure H<sub>2</sub>O; Fig. 2).

481 The empirical nature of the fitting terms incorporating melt composition, pres-  
 482 sure and temperature means that users should be cautious when extrapolating this  
 483 model to conditions lying outside the P-T-X range of the calibration dataset. In  
 484 particular, Iacono-Marziano et al. (2012) highlight five weaknesses of their model:

- 485 1. The effect of melt MgO and FeO contents on CO<sub>2</sub> solubility is poorly con-  
 486 strained, because of the small variation in the concentrations of these oxides  
 487 in the calibration database.
- 488 2. While their compositional terms for the effect of melt composition on CO<sub>2</sub>  
 489 solubility gives equal weight to Na<sub>2</sub>O and K<sub>2</sub>O, the calibration dataset only  
 490 includes K<sub>2</sub>O-rich melts with a range of pressures, so the effect of substituting  
 491 Na and K is poorly constrained.
- 492 3. The effect of temperature on the solubility of mixed H<sub>2</sub>O-CO<sub>2</sub> is poorly con-  
 493 strained because the majority of experiments in the calibration dataset were  
 494 performed at 1200–1300°C.
- 495 4. The relative role of molecular H<sub>2</sub>O vs. OH<sup>-</sup> on melt structure, which in turn  
 496 influences CO<sub>2</sub> solubility, needs to be evaluated further.
- 497 5. The model was calibrated assuming that all Fe was Fe<sup>2+</sup>, so calculated solu-  
 498 bilities are not sensitive to melt redox (unlike the model of P-2006).

499 These weaknesses are explored in more detail in sections 5 and 6.

### 500 3.7 S-2014 (Shishkina et al., 2014)

The Shishkina et al. (2014) model calculates the solubility of H<sub>2</sub>O and CO<sub>2</sub> using fully-empirical expressions. Their expression for CO<sub>2</sub> solubility was calibrated on a dataset of mixed H<sub>2</sub>O-CO<sub>2</sub> experiments on predominantly mafic compositions between 1200–1300°C, and 500–5000 bars:

$$\ln[\text{CO}_2]^{\text{ppm}} = 1.150 \ln(P) + 6.71 \Pi^* - 1.345 \quad (16)$$

Where  $P$  is the pressure in MPa, and  $\Pi^*$  is a compositional parameter expressed in terms of the cation fractions of 7 species:

$$\Pi^* = \frac{X_{\text{Ca}^{2+}} + 0.8 X_{\text{K}^+} + 0.7 X_{\text{Na}^+} + 0.4 X_{\text{Mg}^{2+}} + 0.4 X_{\text{Fe}^{2+}}}{X_{\text{Si}^{4+}} + X_{\text{Al}^{3+}}} \quad (17)$$

501 We note for completeness that the expression provided in Shishkina et al. (2014)  
502 incorrectly states that  $\text{CO}_2$  in equation 16 was in wt%, rather than ppm.

Their expression for  $\text{H}_2\text{O}$  solubility was calibrated on a dataset of pure  $\text{H}_2\text{O}$  experiments with mafic to intermediate compositions between 1200–1250°C, and 485–5009 bars. It incorporates a composition parameter expressed in terms of the anhydrous cation fractions of Na and K:

$$[\text{H}_2\text{O}]^{\text{wt}\%} = (3.36 \times 10^{-7} P^3 - 2.33 \times 10^{-4} P^2 - 0.0711 P - 1.1309)(X_{\text{Na}} + X_{\text{K}}) - 1.2 \times 10^{-5} P^2 + 0.0196 P + 1.1297 \quad (18)$$

503 In general, the compositional range of the Shishkina et al. (2014) dataset includes  
504 a larger variety of mafic compositions than that of Iacono-Marziano et al. (2012),  
505 particularly with respect to alkali-rich, or highly depolymerized melts (Fig. 2). How-  
506 ever, as was true for IM-2012, the empirical nature of the compositional term means  
507 that extreme care is needed when extrapolating this model beyond the composi-  
508 tional range of the calibration dataset (see section 6).

509 One caveat of the implementation of this model in VESICAL is the treatment of  
510 mixing between  $\text{CO}_2$  and  $\text{H}_2\text{O}$ . Shishkina et al. (2014) note that their experimental  
511 data shows evidence for significant non-ideality, with isobars remaining almost hori-  
512 zontal between 0–4 wt%  $\text{H}_2\text{O}$  (see their Fig. 6). However, the isobars shown on their  
513 plots are fitted to experimental data, rather than derived from their equations for  
514  $\text{CO}_2$  and  $\text{H}_2\text{O}$  solubility. These fits cannot be applied to melts with different com-  
515 positions, and the authors give no guidance as to how to combine their equations for  
516 pure  $\text{CO}_2$  and pure  $\text{H}_2\text{O}$  to reproduce this non-ideal behaviour for any given melt  
517 composition. Thus, due to an absence of other information, VESICAL treats mixing  
518 between  $\text{H}_2\text{O}$  and  $\text{CO}_2$  as ideal in this model. To emphasize this assumption, the  
519  $\text{H}_2\text{O}$ - $\text{CO}_2$  model called ShishkinaIdealMixing in VESICAL.

### 520 **3.8 MagmaSat: Ghiorso and Gualda (2015)**

521 Ghiorso and Gualda (2015) present a comprehensive thermodynamic model  
522 (MagmaSat) for mixed  $\text{H}_2\text{O}$ - $\text{CO}_2$  solubility, calibrated on the most chemically-



523 diverse set of natural silicate melt compositions of all the models discussed thus far  
 524 (Fig. 2). For this reason, it is the default model in VESIcal. MagmaSat uses the  
 525 equation of state of Z. Duan and Zhang (2006) for the CO<sub>2</sub>-H<sub>2</sub>O fluid, and is the  
 526 volatile solubility model implemented in rhyolite-MELTS v.1.2 (Gualda et al., 2012).  
 527 Thus, it is currently the only model which can be directly integrated with phase  
 528 equilibrium calculations (e.g., to track microlite growth during degassing upon as-  
 529 cent, or post-entrapment modification to melt inclusions). Like P-2006, the model  
 530 considers the Gibbs free energy of solution using interaction parameters (denoted  
 531 with a capital  $W$  in this model), although, unlike in P-2006, these terms are in-  
 532 dependent of pressure (as well as temperature). The exact choice of components  
 533 differs from that of P-2006, as MagmaSat adapts the formulation from Rhyolite-  
 534 MELTS (e.g., Mg is considered as Mg<sub>2</sub>SiO<sub>4</sub>, Ghiorso & Sack, 1995). Unlike P-2006,  
 535 MagmaSat considers the dissolution of CO<sub>2</sub> as both molecular CO<sub>2</sub> and carbonate  
 536 species (as CaCO<sub>3</sub>), with a set of  $W$  coefficients for each. MagmaSat assumes that  
 537 water dissolves entirely as a hydroxyl species, rather than considering both hydroxyl  
 538 and molecular species. This helps to reduce the number of interaction parameters  
 539 for volatile-melt species, and seems to be a justified simplification based on available  
 540 experimental data (see Ghiorso and Gualda, 2015 for a more detailed discussion).

541 The calibration dataset for H<sub>2</sub>O spans 550–1420°C, and pressures of 0–20,000  
 542 bars, and for CO<sub>2</sub> spans 1140–1400 °C and 0–30,000 bars. However, as discussed in  
 543 detail later, care is required when interpreting H<sub>2</sub>O solubility at pressures >10 kbar.  
 544 Importantly, unlike P-2006, MagmaSat is not calibrated for synthetic liquids (e.g.,  
 545 compositions only containing a small number of oxide species like albite), so should  
 546 only be applied to natural silicate liquid compositions.

### 547 **3.9 A-2019: Allison et al. (2019)**

548 Allison et al. (2019) present thermodynamic models to calculate CO<sub>2</sub> solubility  
 549 for six different basaltic compositions from Stromboli (alkali basalt), Etna (trachy-  
 550 basalt), Vesuvius (phonotephrite), Erebus (phonotephrite), Sunset Crater (alkali  
 551 basalt), and the San Francisco Volcanic Field (basaltic andesite, Fig. 2). Specifi-  
 552 cally, they performed experiments at 1200°C, and ~4000–6000 bars to address the  
 553 paucity of experiments examining CO<sub>2</sub> solubility in alkali systems at mid crustal  
 554 pressures. In addition to these experiments, their models for Vesuvius, Etna and

555 Stromboli incorporate experiments from the literature, extending the calibration  
 556 range of these three models to upper crustal pressures (see Fig. 1). Unlike models  
 557 which incorporate the effect of changing melt composition empirically (e.g., Newman  
 558 & Lowenstern, 2002; Iacono-Marziano et al., 2012; Shishkina et al., 2014), Allison  
 559 et al. (2019) determine the parameters  $\Delta V_r^{0,m}$  (the molar volume change of the  
 560 condensed components of the reaction) and  $K_0(P_0, T_0)$  (the equilibrium constant  
 561 at the reference pressure and temperature) within their thermodynamic equation  
 562 empirically for each of the six compositions they examine, and create 6 separate  
 563 models (each of which contains no compositional dependence). The A-2019 models  
 564 incorporate the modified Redlich-Kwong equation of state provided in the appendix  
 565 of Holloway and Blank (1994).

566 Interestingly, Allison et al. (2019) show that  $\text{CO}_2$  solubility does not sim-  
 567 ply scale with total alkali contents. Erebus melts have  $\text{Na}_2\text{O}+\text{K}_2\text{O}=8.8$  wt%, but  
 568 dissolve less  $\text{CO}_2$  than Etna and Vesuvius melts ( $\text{Na}_2\text{O}+\text{K}_2\text{O}=5.2$  and 7.8 wt% re-  
 569 spectively). They suggest that CaO, MgO and  $\text{Al}_2\text{O}_3$  may play a role in the lower  
 570 solubility of Erebus compared to Etna, but the fact that 5 of the 7 major elements  
 571 they examine show notable differences between these melt compositions make it  
 572 difficult to conclusively determine the origin of solubility variations. Ideally, users  
 573 would apply the A-2019 Etna model to lavas erupted at Etna, the Stromboli model  
 574 to lavas erupted at Stromboli and so on. The absence of an empirical term for melt  
 575 composition means that extreme care should be taken when applying these equa-  
 576 tions to alkaline lavas with different major element contents to those used in the  
 577 experiments of Allison et al. (2019), even if the lavas originated from one of the 6  
 578 volcanoes they examine.

579 Allison et al. (2019) only present equations for  $\text{CO}_2$  solubility, as their experi-  
 580 ments weren't designed to have a wide range of  $\text{H}_2\text{O}$  contents at different pressures,  
 581 and their high  $X_{\text{CO}_2}^f$  values mean that errors in their fluid fraction measurements  
 582 propagate into large errors for  $\text{H}_2\text{O}$  fugacity (relative to the insignificant errors for  
 583  $\text{CO}_2$  fugacity). In their supplementary spreadsheet, they integrate their  $\text{CO}_2$  sol-  
 584 ubility models with an power law fit for water solubility developed specifically for  
 585 Etna (Equation 2 of Lesne, Scaillet, Pichavant, Iacono-Marziano, & Beny, 2011).  
 586 In VESical, users can combine any of the A-2019 carbon models with  $\text{H}_2\text{O}$  models  
 587 from M-1998, IM-2012 and S-2014, or write their own.

## 4 Model Comparisons

To aid comparisons between models, we use a number of silicate melt compositions spanning a range of typical end-member compositions found in nature (Table 1, Fig. 2) to examine the relationship between volatile solubility and pressure, the treatment of mixing between H<sub>2</sub>O and CO<sub>2</sub> (manifested in the shapes of isobars), as well as sensitivity to parameters such as temperature and oxygen fugacity. For basalts, we consider two mid-ocean ridge basalts (MORB1, 50.8 wt% SiO<sub>2</sub>, (Dixon et al., 1995), MORB2, 47.4 wt% SiO<sub>2</sub>, (Allan et al., 1989), one alkali basalt from Etna (48.8 wt% SiO<sub>2</sub>, (Allison et al., 2019) and an arc basalt from Fuego (49.7 wt% SiO<sub>2</sub>, (Lloyd et al., 2013). Comparisons between the MORB and Etna compositions were performed at 1200°C, while those for Fuego were performed at 1000°C to reflect the lower temperatures typical of more H<sub>2</sub>O-rich basalts.

For rhyolitic magmas, we perform calculations at 800°C for a high-Si rhyolite from Mono Craters (Eastern California) with 77.19 wt% SiO<sub>2</sub> from Liu et al. (2005), and a peralkaline rhyolitic melt inclusion from Aluto Volcano with 73.5 wt% SiO<sub>2</sub> measured by Iddon and Edmonds (2020). The Aluto rhyolite has much lower Al<sub>2</sub>O<sub>3</sub> and higher FeO contents than the Mono Craters Rhyolite (Table 1).

**Table 1.** Representative compositions used for comparisons. MORB1 is a Mid-Oceanic Ridge Basalt tholeiite from Dixon et al. (1995). MORB2 is the MORB composition given in Table 3 of Ghiorso and Gualda (2015), originally from Allan et al. (1989). Etna is sample ET-8 from the supplementary information of Allison et al. (2019). Fuego is the composition of a melt inclusion from Lloyd et al. (2013). Mono Craters is from Table 1 of Liu et al. (2005). Aluto rhyolite is the composition of a quartz-hosted melt inclusion from the East African Rift (MI70 from sample MER055A; Iddon & Edmonds, 2020).

Name	MORB1	MORB2	Etna	Fuego	Mono Craters	Aluto Pumice
SiO <sub>2</sub>	50.8	47.4	48.77	49.67	77.19	73.51
TiO <sub>2</sub>	1.84	1.01	1.79	1.17	0.06	0.23
Al <sub>2</sub> O <sub>3</sub>	13.7	17.64	16.98	16.50	12.8	9.18
Fe <sub>2</sub> O <sub>3</sub>	2.1	0.89	2.51	1.65	0.26	1.41
FeO	10.5	7.18	6.44	8.43	0.71	3.81
MnO	0	0	0.18	0.19	0	0.25
MgO	6.67	7.63	6.33	4.38	0.03	0
CaO	11.5	12.44	11.26	7.90	0.53	0.2
Na <sub>2</sub> O	2.68	2.65	3.65	3.37	3.98	4.18
K <sub>2</sub> O	0.15	0.03	1.79	0.79	4.65	4.22
P <sub>2</sub> O <sub>5</sub>	0.19	0.08	0.53	0.22	0	0

#### 4.1 Redox sensitivity

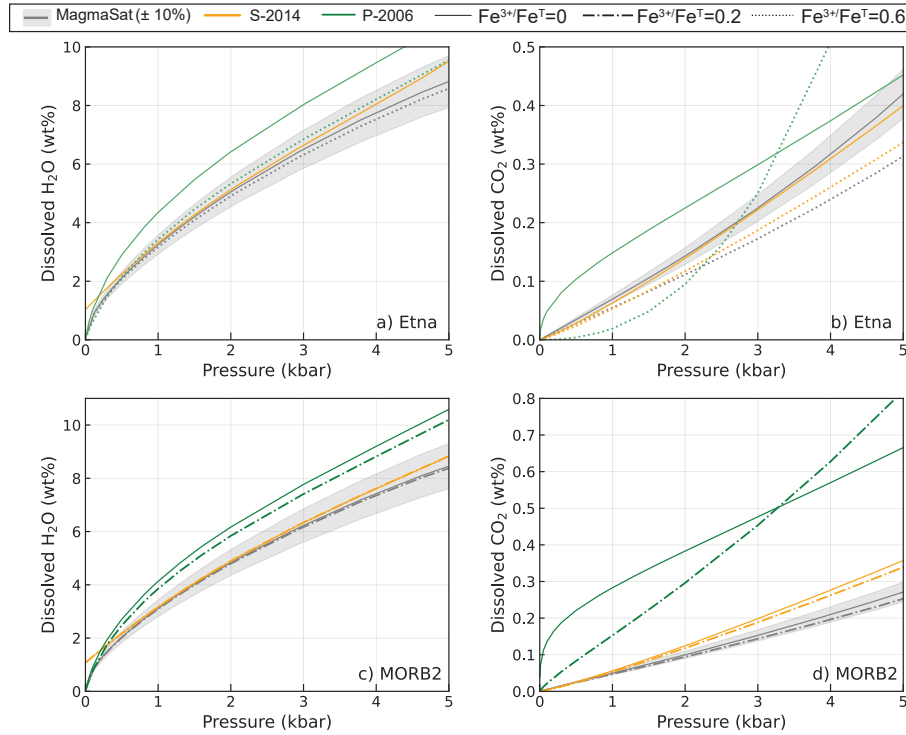
Before proceeding with these comparisons, it is worth noting that the vast majority of studies report whole-rock, melt inclusion and matrix glass compositions in terms of  $\text{FeO}_t$ , because the proportions of  $\text{FeO}$  vs.  $\text{Fe}_2\text{O}_3$  are difficult to determine precisely using common analytical techniques such as electron probe microanalysis (EPMA) and x-ray fluorescence (XRF). VolatileCalc-Rhyolite and L-2005 have no compositional terms, and VolatileCalc-Basalt is only parametrized in terms of the melt  $\text{SiO}_2$  content, so these 3 models are not sensitive to the choice of  $\text{FeO}$  vs.  $\text{Fe}_2\text{O}_3$  for the representative compositions in Table 1. Similarly, IM-2012 and M-1998 are parameterized using an  $\text{FeO}_t$  term, so also show no sensitivity to melt redox. S-2014 is technically slightly redox-sensitive for  $\text{CO}_2$ , because the  $\Pi^*$  term is expressed in terms of  $\text{Fe}^{2+}$  species (equation 17). However, as the model was calibrated assuming  $\text{Fe}^{2+} = \text{Fe}_T$ , any sensitivity to redox is likely spurious. Thus, VESical calculates  $\Pi$  with  $\text{Fe}_T$  by default for consistency with their calibration. In Figure 4, we perform calculations for different  $\text{Fe}^{3+}$  ratios for completeness, but the rest of the figures in the manuscript for S-2014 are calculated using  $\text{Fe}_T$ .  $\text{H}_2\text{O}$  solubility in S-2014 is not redox sensitive, because the effect of melt composition is only parametrized in terms of cation fractions of Na and K (equation 18). Both P-2006 and MagmaSat have interaction parameters for  $\text{Fe}^{2+}$  and  $\text{Fe}^{3+}$ -bearing species, so are redox sensitive for both  $\text{CO}_2$  and  $\text{H}_2\text{O}$  solubility.

We examine the sensitivity of calculations using S-2014, MagmaSat and P-2006 to melt redox by performing calculations for 0, 10 and 20%  $\text{Fe}^{3+}$  for MORB2, and 0, 30 and 60%  $\text{Fe}^{3+}$  for Etna (the higher redox accounting for the highly oxidising conditions of experiments on Etna melts, e.g., Lesne, Scaillet, Pichavant, & Beny, 2011).

For Etna, pure  $\text{H}_2\text{O}$  solubility in MagmaSat is relatively insensitive to redox, predicting variations in dissolved  $\text{H}_2\text{O}$  which are well within model error (Fig. 4a). Pure  $\text{CO}_2$  solubility in MagmaSat is more redox sensitive than  $\text{H}_2\text{O}$ , predicting  $\sim 1.2\text{--}1.3\times$  more  $\text{CO}_2$  for 0%  $\text{Fe}^{3+}$  vs. 60%  $\text{Fe}^{3+}$  (Fig. 4b). Pure  $\text{H}_2\text{O}$  solubility in P-2006 shows the same directionality as MagmaSat, but is more sensitive to redox ( $1.8\times$  more  $\text{H}_2\text{O}$  dissolves at 0.1 kbar for 0%  $\text{Fe}^{3+}$  vs. 60%  $\text{Fe}^{3+}$ , dropping to  $1.2\times$  at  $> 2$  kbar). Pure  $\text{CO}_2$  solubility in P-2006 is extremely redox-sensitive, with melts

637 with 0%  $\text{Fe}^{3+}$  vs. 60%  $\text{Fe}^{3+}$  dissolving  $25\times$  more  $\text{CO}_2$  at 0.5 kbar, but  $0.5\times$  less  
 638  $\text{CO}_2$  at 5 kbar. S-2014 is slightly less redox sensitive than MagmaSat for  $\text{CO}_2$ .

639 Varying  $\text{Fe}^{3+}$  proportions between 0–20% for MORB2 produces similar pat-  
 640 terns as Etna, with changes lying within model error for MagmaSat and S-2014,  
 641 but showing significant differences for P-2006 (Fig. 4c-d). Isobars for different redox  
 642 states for this composition can be found in Supporting Figure S2. For the smaller  
 643 changes in  $\text{Fe}^{3+}$  proportions considered for MORB2, changes in dissolved  $\text{H}_2\text{O}$  and  
 644  $\text{CO}_2$  contents for MagmaSat and S-2014 are well within model uncertainty (gener-  
 645 ally stated as 10–20%). In contrast, P-2006 shows changes in dissolved  $\text{CO}_2$  which  
 646 are significantly larger than quoted errors on solubility models.



**Figure 4.** Relationship between volatile solubility and the proportion of  $\text{Fe}^{3+}$  for: a-b) the Etna composition at  $1200^\circ\text{C}$ ; c-d) the MORB2 composition at  $1200^\circ\text{C}$ . VESICAL uses  $\text{FeO}_t$  in S-2014 to calculate  $\Pi^*$  for consistency with the calibration of this model. Here, dashed and dotted lines show the results from calculations where  $\text{FeO}_t$  is multiplied by  $\text{Fe}^{2+}/\text{Fe}_T$ , to show the results that would be obtained if VESICAL calculated  $\Pi^*$  using only  $\text{Fe}^{2+}$ . Isobars for MORB2 and lines for intermediate  $\text{Fe}^{3+}$  ratios for each composition are shown in Supporting Fig. S2.

647 The cause of the extreme sensitivity of CO<sub>2</sub> in P-2006 to redox relative to  
 648 MagmaSat is apparent from examining the interaction coefficients in Table 3. In  
 649 MagmaSat, the  $W_{Fe_2O_3, H_2O}$  coefficient is only  $1.6\times$  bigger than  $W_{Fe_2SiO_4, H_2O}$  ( $\sim 50$   
 650 vs. 31), and these two coefficients overlap within  $\pm 1.5\sigma$  of the uncertainty of these  
 651 coefficients. This accounts for the relatively weak effect of redox on calculated H<sub>2</sub>O  
 652 solubility. For the CaCO<sub>3</sub> component representing the carbonate ion, the coefficients  
 653 have similar magnitudes, but opposite signs ( $W_{Fe_2O_3, CaCO_3} \sim 66$ ,  $W_{Fe_2SiO_4, CaCO_3}$   
 654  $\sim -73$ ), and this difference is much larger than the error on the coefficients (account-  
 655 ing for the stronger effect of melt redox on CO<sub>2</sub> solubility compared with H<sub>2</sub>O). The  
 656 Fe<sub>2</sub>O<sub>3</sub> and Fe<sub>2</sub>SiO<sub>4</sub> coefficients for the CO<sub>2</sub> component (which becomes more domi-  
 657 nant in more evolved compositions) are also significantly different outside the quoted  
 658 error but have the same sign ( $W_{Fe_2O_3, CO_2} \sim -32$ ,  $W_{Fe_2SiO_4, CO_2} \sim -3$ ).

659 In P-2006, the  $w_{H_2O-FeO}^0$  coefficient is of similar magnitude, but opposite sign  
 660 to  $w_{H_2O-Fe_2O_3}^0$  ( $1.4\times 10^5$  vs.  $-2\times 10^5$ ), and clearly distinct outside the error on  
 661 each coefficient. This accounts for the slightly stronger sensitivity of H<sub>2</sub>O in P-2006  
 662 to redox compared with MagmaSat. In stark contrast to all the comparisons thus  
 663 far, the  $w_{CO_2-Fe_2O_3}^0$  coefficient is  $>200\times$  larger than  $w_{CO_2-FeO}^0$  for P-2006. In fact,  
 664  $w_{CO_2-Fe_2O_3}^0$  is  $\sim 8\times$  higher than the next largest coefficient, suggesting that for a  
 665 given mole fraction in the melt, it has the largest effect on carbonate ion solubil-  
 666 ity. The  $w_{CO_2-Fe_2O_3}^1$  coefficient in P-2006 model, which becomes more dominant  
 667 at higher pressures, has the opposite sign to that for  $w_0$ . This accounts for the fact  
 668 that at low pressures ( $<3$  kbar), increasing proportions of Fe<sup>3+</sup> cause CO<sub>2</sub> solubility  
 669 to decrease, while at higher pressures, increasing proportions of Fe<sup>3+</sup> cause CO<sub>2</sub>  
 670 solubility to increase.

671 It is difficult to trust the extreme sensitivity of CO<sub>2</sub> in P-2006 to redox given  
 672 the large uncertainty associated with the proportions of Fe<sup>3+</sup> in volatile solubility  
 673 experiments. For example, S-2014 note that only 7 of the 48 experiments in their  
 674 calibration dataset contain non-zero values of Fe<sub>2</sub>O<sub>3</sub>, which is why they choose to  
 675 calibrate the model using FeO<sub>t</sub>. In the P-2006 dataset, only 6 studies used in the  
 676 calibration directly determined the proportion of Fe<sup>3+</sup>, and a further 9 reported  
 677 the experimental oxygen fugacity. Thus, for the vast majority of their experimental  
 678 calibration dataset, Papale et al. (2006) calculate the proportion of Fe<sup>3+</sup> assuming  
 679 the oxygen fugacity is controlled by the H<sub>2</sub>O-H<sub>2</sub> equilibrium at the stated experi-

680 mental conditions. However, this method requires accurate measurements of fluid  
 681 composition, is affected by Fe and H<sup>+</sup>-loss during experiments, and it is unclear  
 682 how applicable this method is for mixed H<sub>2</sub>O-CO<sub>2</sub> experiments (Botcharnikov et al.,  
 683 2006). The P-2006 calibration dataset contains some very surprising values: in the  
 684 calibration dataset for pure CO<sub>2</sub> experiments, the experiments of Fogel and Ruther-  
 685 ford (1990) have been allocated Fe<sup>3+</sup>/Fe<sub>T</sub> ratios of ~0.9–1, despite the authors  
 686 debating whether their experiments were actually reducing enough to stabilize a CO  
 687 species. Overall, it is worth noting that the coefficients for CO<sub>2</sub> are relatively under-  
 688 constrained. While there are 10 coefficients for H<sub>2</sub>O calibrated using 865 datapoints  
 689 of pure-H<sub>2</sub>O solubility, there are 20 coefficients for CO<sub>2</sub>, and only 173 datapoints for  
 690 pure-CO<sub>2</sub> solubility and 84 for mixed fluids. In particular, the highly oxidising  
 691 Fe<sup>3+</sup>/Fe<sub>T</sub> ratios calculated assuming H<sub>2</sub>O-H<sub>2</sub> equilibrium are more concentrated  
 692 in experiments with low CO<sub>2</sub> contents and pressures, making it difficult to decon-  
 693 volve the differential effects of these parameters in a model with a large number of  
 694 coefficients being calibrated on a relatively small calibration dataset.

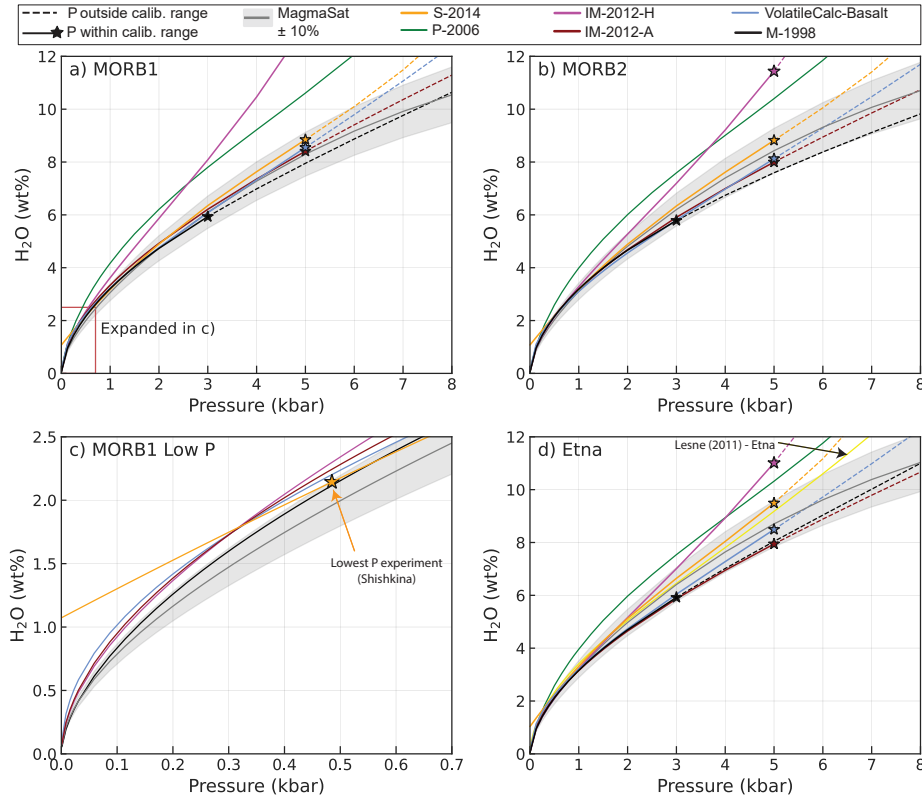
695 In sections 4.2 to 5, we show calculations using the Fe<sup>3+</sup>/Fe<sub>T</sub> proportions  
 696 in Table 1, as these best-estimates for each center are representative of what a  
 697 user would select when calculating melt inclusion saturation pressures, dissolved  
 698 volatile contents etc. For MORB1, MORB2, and Fuego, these proportions are from  
 699 the original publications. For Etna, Fe<sub>2</sub>O<sub>3</sub> was calculated from FeO<sub>t</sub> assuming  
 700 Fe<sup>3+</sup>/Fe<sub>T</sub>=0.26 after Gaborieau et al. (2020). For Mono Craters and the Aluto  
 701 pumice, Fe<sup>3+</sup>/Fe<sub>T</sub> was set at 0.25 based on available data on other rhyolites (e.g.,  
 702 Ghiorso and Gualda, 2015), and modelling studies of the fractional crystallization  
 703 path at Aluto (Gleeson et al., 2017).

## 704 **4.2 Mafic Compositions**

### 705 **4.2.1 Pure H<sub>2</sub>O**

706 The 7 models applicable to H<sub>2</sub>O in basaltic systems predict a sharp rise in pure  
 707 H<sub>2</sub>O solubility with increasing pressure (Fig. 5). For all three melt compositions,  
 708 S-2014, IM-2012-A, VolatileCalc-Basalt, and M-1998 predict H<sub>2</sub>O concentrations  
 709 within ±10% of MagmaSat (grey envelope) between ~1–5 kbar. For MORB1 and  
 710 MORB2, IM-2012-H begins to deviate to higher H<sub>2</sub>O contents than MagmaSat at

711 >1 kbar. For Etna, IM-2012-H follows a similar trajectory between 0–3 kbar to the  
 712 solubility model of Lesne, Scaillet, Pichavant, Iacono-Marziano, and Beny (2011)  
 713 developed specifically for Etna melts (yellow line, Fig. 5d). In contrast, P-2006 plots  
 714 to substantially higher H<sub>2</sub>O solubilities compared to all other models at >0.5 kbar  
 715 (although P-2006 and IM-2012-H intercept at higher pressures).



**Figure 5.** Relationship between pure H<sub>2</sub>O solubility and pressure for MORB1, MORB2 and Etna melts at 1200°C. The grey field shows a  $\pm 10\%$  error window around MagmaSat. Model lines are dashed when extrapolated above the recommended pressure range. The low pressure region of a) is expanded in c), emphasizing the non-zero solubility of H<sub>2</sub>O at P=0 bar in the S-2014 model.

716 The fact that IM-2012-H predicts higher H<sub>2</sub>O solubility relative to the cluster  
 717 of other models lying within the error window of MagmaSat is an interesting obser-  
 718 vation. Iacono-Marziano et al. (2012) favour their hydrous model, particularly for  
 719 CO<sub>2</sub>-poor, H<sub>2</sub>O-rich melts, based on regressions between predicted and measured  
 720 H<sub>2</sub>O contents, and by comparing the two models to experiments conducted between



721 1 and 4 kbar on Etna melts (their Figure 7a). However, their Figure 8c, which com-  
 722 pares dissolved volatile contents calculated by the model for the entire calibration  
 723 dataset, shows that predicted H<sub>2</sub>O contents using the hydrous version are overesti-  
 724 mates for experimental products with >6–7 wt% H<sub>2</sub>O (although these predictions  
 725 still lie within the ~17% error associated with their H<sub>2</sub>O model). Our comparisons  
 726 suggest that the anhydrous model is most similar to other models, so should not  
 727 automatically be discounted in favour of the hydrous model.

Another notable oddity is the nearly linear trajectory of H<sub>2</sub>O vs. P in S–2014  
 at <0.5 kbar, causing this model to predict a non-zero solubility of H<sub>2</sub>O at 0 bar  
 (Fig. 5c). This contrasts with the power-law shapes followed by the other models  
 which intercept very close to the origin. This anomalous behaviour is because the  
 S–2014 equation for H<sub>2</sub>O solubility (equation 18) simplifies at P=0 to:

$$[\text{H}_2\text{O}]^{\text{wt}\%} = -1.1309(\text{X}_{\text{Na}} + \text{X}_{\text{K}}) + 1.1297 \quad (19)$$

728 In the S–2014 calibration dataset,  $\text{X}_{\text{Na}} + \text{X}_{\text{K}}$  varies from 0.05 to 0.25, which corre-  
 729 sponds to solubilities of 0.85–1.07 wt% H<sub>2</sub>O at 0 bar. This demonstrates the issue  
 730 with extrapolating empirical expressions beyond the calibration range (the lowest  
 731 pressure experiment in the calibration dataset of S–2014 was conducted at 485 bar).  
 732 When combined with the assumption of ideal mixing used in VESIcal, this non-zero  
 733 solubility of H<sub>2</sub>O at 0 bar results in S–2014 predicting unusual degassing paths and  
 734 isobar shapes relative to other models. For example, if a melt has <1 wt% H<sub>2</sub>O, S–  
 735 2014 predicts that the co-existing fluid contains no H<sub>2</sub>O, despite abundant evidence  
 736 that volcanic plumes in low H<sub>2</sub>O systems such as Hawai'i are dominated by H<sub>2</sub>O at  
 737 low pressures (Gerlach, 1986). It also causes isobars to be entirely flat at low H<sub>2</sub>O  
 738 contents (see section 4.2.2, Fig. 6).

739 Overall, excluding P–2006 and IM–2012–H based on their higher predictions  
 740 of H<sub>2</sub>O solubility, and S–2014 based on anomalous behavior at low H<sub>2</sub>O contents,  
 741 the remaining four solubility models predict dissolved H<sub>2</sub>O concentrations within  
 742 error of one another at pressures lower than the upper calibration limit. This likely  
 743 reflects the relatively small effect of melt composition of H<sub>2</sub>O solubility, meaning  
 744 that more recent models calibrated on a wider compositional range display similar  
 745 behavior to older models (G. Moore & Carmichael, 1998; Papale et al., 2006). The  
 746 larger deviation between models at higher pressures reflect the fact that very few

747 pure-H<sub>2</sub>O solubility experiments have been performed at > 5 kbar (Table 1). One  
 748 reason for this shortage of higher pressure experimental data results from the fact  
 749 that it is very difficult to quench silicate melts with >9 wt% H<sub>2</sub>O to a glass phase  
 750 which can be analysed by FTIR or SIMS (Gavrilenko et al., 2019; Mitchell et al.,  
 751 2017).

#### 752 *4.2.2 Mixed H<sub>2</sub>O-CO<sub>2</sub>*

753 The majority of experiments used to calibrate expressions for pure-CO<sub>2</sub> solu-  
 754 bility contained dissolved H<sub>2</sub>O and CO<sub>2</sub> (e.g., Iacono-Marziano et al., 2012; Shishk-  
 755 ina et al., 2014; Allison et al., 2019), requiring authors to assess H<sub>2</sub>O-CO<sub>2</sub> mixing  
 756 behaviour to determine the solubility of pure-CO<sub>2</sub> fluids. Thus, it makes sense to  
 757 consider the treatment of mixing between CO<sub>2</sub> and H<sub>2</sub>O species in each model  
 758 before considering predictions of pure CO<sub>2</sub> solubility which are affected by these  
 759 assumptions. The treatment of H<sub>2</sub>O-CO<sub>2</sub> mixing is best demonstrated using isobar  
 760 diagrams, which show the solubility of H<sub>2</sub>O and CO<sub>2</sub> in a given silicate melt com-  
 761 position at a given pressure for proportions of  $X_{H_2O}^f$  in the co-existing fluid ranging  
 762 from 0 (interception with the y axis) to 1 (interception with the x axis). The treat-  
 763 ment of mixed fluids differs quite considerably in each model.

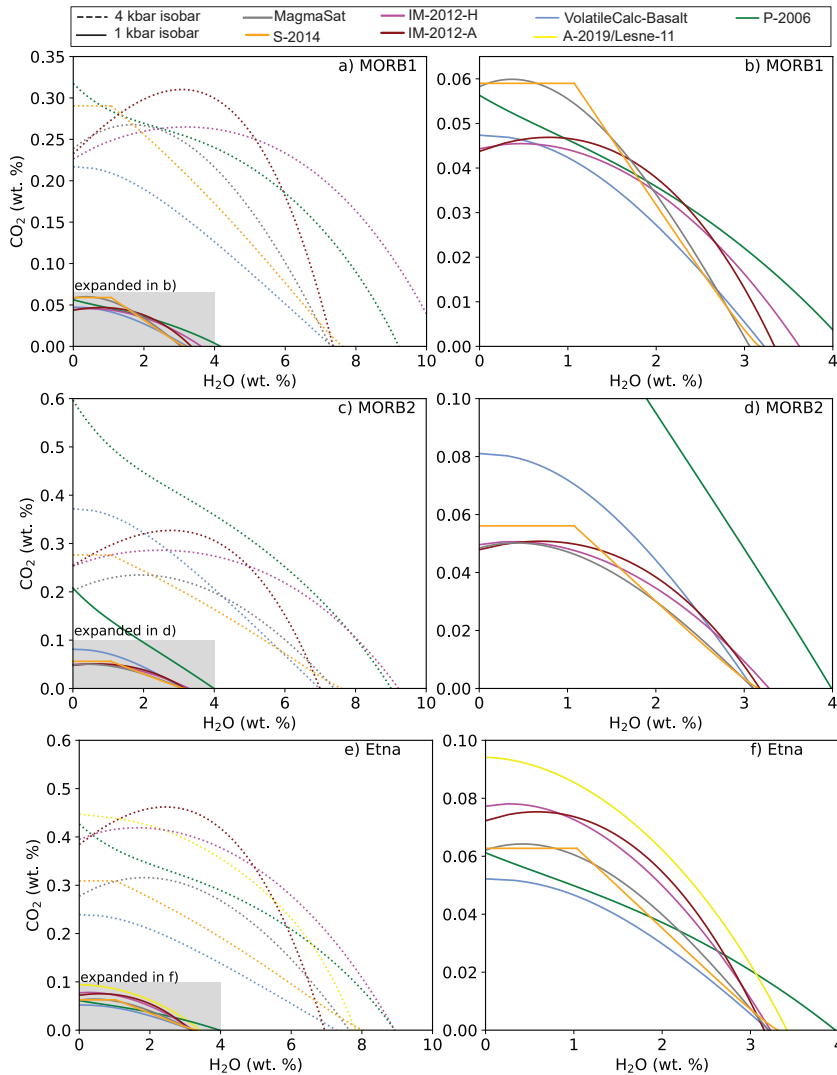
764 VolatileCalc-Basalt models mixed fluids under the assumption of Henrian  
 765 (ideal) mixing in the fluid and melt phase. Thus, the addition of H<sub>2</sub>O always causes  
 766 the solubility of CO<sub>2</sub> to decrease (and vice versa), and isobars possess a nega-  
 767 tive gradient, with a slightly sloping plateau at low H<sub>2</sub>O contents merging into a  
 768 concave-down shape (Fig. 6).

769 S-2014 does not provide an equation for the treatment of non-ideal mixing,  
 770 despite their experiments showing that increasing H<sub>2</sub>O contents at high pressure  
 771 cause almost no change in CO<sub>2</sub> solubility. Using the assumption of ideal mixing in  
 772 VESIcal, S-2014 isobars exhibit a flat plateau at low H<sub>2</sub>O contents, merging into  
 773 a negative slope at higher H<sub>2</sub>O contents. This flat plateau results from the fact  
 774 that there are no partial pressures at which S-2014 yields H<sub>2</sub>O < 1 wt%, so the y  
 775 co-ordinate for lower H<sub>2</sub>O contents is equal to the solubility of pure CO<sub>2</sub>.

776 P-2006 is fully non-ideal, which causes isobars to have complex shapes, ex-  
 777 hibiting both positive and negative gradients. In detail, the shape of isobars calcu-

778 lated using P-2006 differ as a function of both melt composition and temperature  
779 (see Fig. 12 of Papale, 1999). For the basaltic compositions considered here, isobars  
780 show a sharp increase to higher CO<sub>2</sub> contents at very low H<sub>2</sub>O contents (Fig. 6a,  
781 c), because the model predicts a decrease in CO<sub>2</sub> solubility following the addition of  
782 small quantities of H<sub>2</sub>O.

783 IM-2012-A and IM-2012-H incorporate empirical representations of non-  
784 ideality through the inclusion of a term for the molar fraction of H<sub>2</sub>O in the melt in  
785 their expression for CO<sub>2</sub> solubility (equation 11). This means that these models pre-  
786 dict that maximum CO<sub>2</sub> solubility occurs at non-zero H<sub>2</sub>O concentrations, causing  
787 isobars to display prominent domed shapes (Fig. 6). Isobars calculated using IM-  
788 2012-A show a more extreme peak than IM-2012-H, because of the difference in the  
789 sign and magnitude of the  $d_{H_2O}$  coefficient combined with the differences between  
790 NBO/O calculated on a hydrous and anhydrous basis (see Supporting Information  
791 for further detail).



**Figure 6.** 1 and 4 kbar isobars for MORB1 (a-b) and MORB2 (c-d) and Etna (e-f) at 1200°C. The 1 kbar isobar is expanded in b), d) and e). The distinctive shapes of isobars from different models reflects different treatment of H<sub>2</sub>O-CO<sub>2</sub> mixing. This is most apparent at higher pressures. The prominent plateau at <1 wt% H<sub>2</sub>O for S-2014 at all pressures results from the non-zero solubility of H<sub>2</sub>O at low pressures.

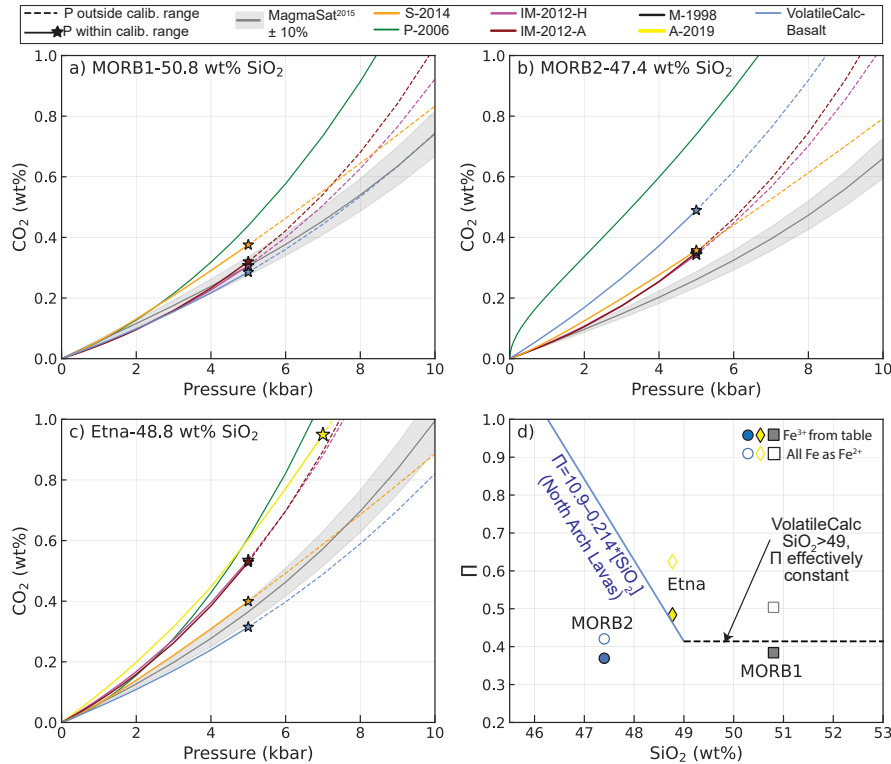
792 Like P-2006, MagmaSat is fully non-ideal. However, unlike P-2006, the treat-  
 793 ment of non-ideality in MagmaSat predicts that the addition of small amounts of  
 794 H<sub>2</sub>O always causes the solubility of CO<sub>2</sub> to increase (so isobars peak at non-zero  
 795 H<sub>2</sub>O concentrations; Fig. 6, c). This peak becomes more pronounced at higher pres-  
 796 sures, but is generally smaller than that predicted by IM-2012.

797 These different mixing assumptions result in large discrepancies between the  
 798 predicted volatile solubilities for melts in equilibrium with H<sub>2</sub>O-CO<sub>2</sub> fluids, par-  
 799 ticularly at higher pressures where non-ideal behaviour is more pronounced. For  
 800 example, while IM-2012-A predicts similar pure CO<sub>2</sub> and pure H<sub>2</sub>O solubilities to  
 801 VolatileCalc-Basalt and MagmaSat for MORB1 at 4 kbar (interception with x and  
 802 y axis on Fig. 6a), IM-2012-A predicts that melts with ~ 4 wt% H<sub>2</sub>O can dissolve  
 803 more than twice as much CO<sub>2</sub> as that predicted by VolatileCalc-Basalt.

#### 804 **4.2.3 Pure CO<sub>2</sub>**

805 All basaltic compositions and models show a large increase in the solubility of  
 806 pure CO<sub>2</sub> with increasing pressure (Fig. 7). The solubility of pure CO<sub>2</sub> is approx-  
 807 imately an order of magnitude lower than for H<sub>2</sub>O (compare Fig. 7 with Fig. 5).  
 808 This solubility difference accounts for the fact that Mid Oceanic Ridge (MOR) mag-  
 809 mas, which have similar concentrations of H<sub>2</sub>O and CO<sub>2</sub> (~ 0.07 wt% H<sub>2</sub>O, 0.1–0.2  
 810 wt% CO<sub>2</sub>; Le Voyer et al., 2019), are almost always CO<sub>2</sub> saturated during crustal  
 811 storage (Saal et al., 2002) but only exsolve measurable quantities of H<sub>2</sub>O if erupted  
 812 at very low pressures (Le Voyer et al., 2019).

813 For MORB1, IM-2012-A and H, and VolatileCalc-Basalt lie within, or close to  
 814 the ± 10% error window on MagmaSat at <5 kbar, and S-2014 lie within ± 20%.  
 815 The deviation at higher pressures is expected, because only P-2006 and MagmaSat  
 816 are calibrated on large numbers of experiments performed at > 5 kbar (Fig. 1). For  
 817 example, the relationship between  $\Pi$  and CO<sub>2</sub> solubility of Dixon (1997) used in  
 818 VolatileCalc-Basalt was based on experiments at 1 kbar, and Newman and Lowen-  
 819 stern (2002) suggest that it should not be extrapolated above 5 kbar. Similarly, only  
 820 the experiments of Jakobsson (1997) in the IM-2012 database were conducted at >5  
 821 kbar, and there are no experiments in the calibration dataset of S-2014 performed at  
 822 >5 kbar. Unlike for pure H<sub>2</sub>O, IM-2012-A and H predict very similar pure CO<sub>2</sub> sol-  
 823 ubilities to one another. This reflects the fact the coefficients for CO<sub>2</sub> between these  
 824 2 model versions are very similar apart from the  $d_{H_2O}$  term, which is multiplied by a  
 825 zero when calculating pure CO<sub>2</sub> solubility. In contrast, P-2006 plots to significantly  
 826 higher pressures than the other models (~ 2× higher at ~8 kbar).



**Figure 7.** Relationship between pure CO<sub>2</sub> solubility and pressure for three mafic melts at 1200°C: a) MORB1 from Dixon et al. (1995), b) MORB2 from Ghiorso and Gualda (2015), and c) Etna from Allison et al. (2019). Models extrapolated beyond their calibrated pressure range are shown as dashed lines, with the colored star marking the recommended upper calibration limit. d) The relationship between  $\Pi$  and SiO<sub>2</sub> defined by the North Arch lavas is shown in blue (Dixon, 1997). Generally, VolatileCalc-Basalt is applied to melts with >49 wt% SiO<sub>2</sub> by setting SiO<sub>2</sub>=49 wt%; the black dashed line represents this extrapolation. These simplified relationships incorporated into VolatileCalc-Basalt (blue and black lines) underestimate the true  $\Pi$  value for Etna, and overestimate it for MORB2.

827 MORB2 shows a significantly larger discrepancy between models than MORB1  
 828 (Fig. 7b), While S-2014, IM-2012-H and -A follow very similar trajectories at  
 829 <5 kbar, P-2006 predicts that MORB2 dissolves ~3370 ppm CO<sub>2</sub> at 2 kbar while  
 830 MagmaSat predicts only ~950 ppm (factor of 3.5×), and VolatileCalc-Basalt  
 831 predicts higher CO<sub>2</sub> solubility relative to MagmaSat by a factor of 1.8× at 2 kbar.

832 The high CO<sub>2</sub> solubility predictions by P-2006 relative to other models and ex-  
 833 perimental measurements has also been noted by Shishkina et al. (2010), Shishkina

834 et al. (2014) and Mangan et al. (2021). This may result from the fact that the P-  
 835 2006 uses a negative compressibility for the CO<sub>2</sub> fluid (Ghiorso & Gualda, 2015).  
 836 Alternatively, it may reflect the large errors on the CO<sub>2</sub>  $w$  coefficients, meaning that  
 837 the effect of melt composition is uncertain, accounting for the larger discrepancy  
 838 for MORB1 vs. MORB2 compared to other models. Finally, Shishkina et al. (2010)  
 839 suggest that the overprediction of CO<sub>2</sub> solubility by P-2006 may result from the  
 840 inclusion of anomalously high CO<sub>2</sub> contents from the experiments of Freise (2004)  
 841 in the calibration dataset of P-2006 (these values have now been revised to lower  
 842 numbers, as the original FTIR thickness correction factor is thought to have been  
 843 incorrect).

844 The fact that VolatileCalc-Basalt plots close to other models for MORB1 but  
 845 not MORB2 is a good example of the main caveat of the  $\Pi$ -SiO<sub>2</sub> simplification  
 846 used to account for the effect of melt composition on CO<sub>2</sub> solubility. For melts with  
 847 40–49 wt% SiO<sub>2</sub>, VolatileCalc-Basalt assumes that the relationship between CO<sub>2</sub>  
 848 solubility and SiO<sub>2</sub> is identical to that defined by the North Arch lavas, shown in  $\Pi$   
 849 vs. SiO<sub>2</sub> space as a blue line in Fig. 7d. This is a reasonable approximation for the  
 850 MORB1 composition, which has a  $\Pi$  value similar to North Arch Lavas with 49 wt%  
 851 SiO<sub>2</sub>. However, the MORB2 composition lies significantly below the line defined by  
 852 North Arch lavas, so has a lower  $\Pi$  value, and therefore a lower CO<sub>2</sub> solubility at a  
 853 given SiO<sub>2</sub> content compared to the North Arch Lavas. Thus, VolatileCalc-Basalt  
 854 likely overpredicts the solubility of CO<sub>2</sub> in this melt composition.

855 Furthermore, VolatileCalc-Basalt predicts that MORB2 dissolves  $\sim 1.7$  times  
 856 more CO<sub>2</sub> at a given pressure than MORB1. This is because MORB2 has 3.4 wt%  
 857 less SiO<sub>2</sub> than MORB1, and VolatileCalc-Basalt predicts that CO<sub>2</sub> solubility in-  
 858 creases drastically with decreasing SiO<sub>2</sub>. However, if the full  $\Pi$  expression of Dixon  
 859 (1997) were used, MORB1 and MORB2 would have very similar CO<sub>2</sub> solubility,  
 860 as they have similar  $\Pi$  values despite different SiO<sub>2</sub> contents. P-2006 also predicts  
 861 that MORB2 dissolves  $5\text{--}6\times$  more CO<sub>2</sub> at 0.4 kbar, and  $1.9\times$  more at 4 kbar than  
 862 MORB1. In contrast, the models of S-2014, IM-2012, and MagmaSat predict that  
 863 MORB1 and MORB2 dissolve similar amounts of CO<sub>2</sub> (MORB2/MORB1= $\sim 0.99\times$ ,  
 864  $\sim 1.125\times$  and  $\sim 0.81\text{--}0.89\times$  respectively). These three more recent models utilize  
 865 significantly larger basaltic calibration datasets to parametrize the effect of multiple  
 866 oxide species melt on CO<sub>2</sub> solubility (Fig. 2), so likely predict more realistic solubil-

867 ity relationships than VolatileCalc-Basalt and P-2006. In summary, CO<sub>2</sub> solubility  
868 in melt compositions that do not follow a similar trajectory in II-SiO<sub>2</sub> space as the  
869 North Arch Lavas (Fig. 7d) is unlikely to be accurately predicted by VolatileCalc-  
870 Basalt.

871 There is also significant deviation between different models for Etna melts  
872 (Fig. 7c), which is far greater than that observed for H<sub>2</sub>O (Fig. 5). The A-2019  
873 model, developed specifically for the composition of Etna magmas, predicts much  
874 higher CO<sub>2</sub> solubility at a given pressure than VolatileCalc-Basalt, S-2014, and  
875 MagmaSat, while P-2006 and IM-2012-H and -A follow similar trajectories to  
876 A-2019. The success of both IM-2012 models likely reflects the large number  
877 of alkaline compositions in their calibration dataset, including some from Etna.  
878 VolatileCalc-Basalt predicts the lowest CO<sub>2</sub> solubility (factor of 0.5–0.6× that of  
879 A-2019). The calculated II value for Etna lies significantly above the line defined  
880 by North Arch lavas (so VolatileCalc-Basalt predicts lower CO<sub>2</sub> solubility; Fig. 7d).  
881 However, even the full II expression of Dixon (1997) is unlikely to be successful, be-  
882 cause alkaline magmas show considerable variation in CO<sub>2</sub> solubility at a given II  
883 value (Allison et al., 2019). As S-2014’s expression for CO<sub>2</sub> solubility incorporates  
884 a II\* term very similar to the II term of Dixon (1997), the deviation of this model  
885 from that of A-2019 (0.6–0.7×) may also result from variations in CO<sub>2</sub> solubility  
886 that are not incorporated by this simplified melt composition parameter (Allison et  
887 al., 2019).

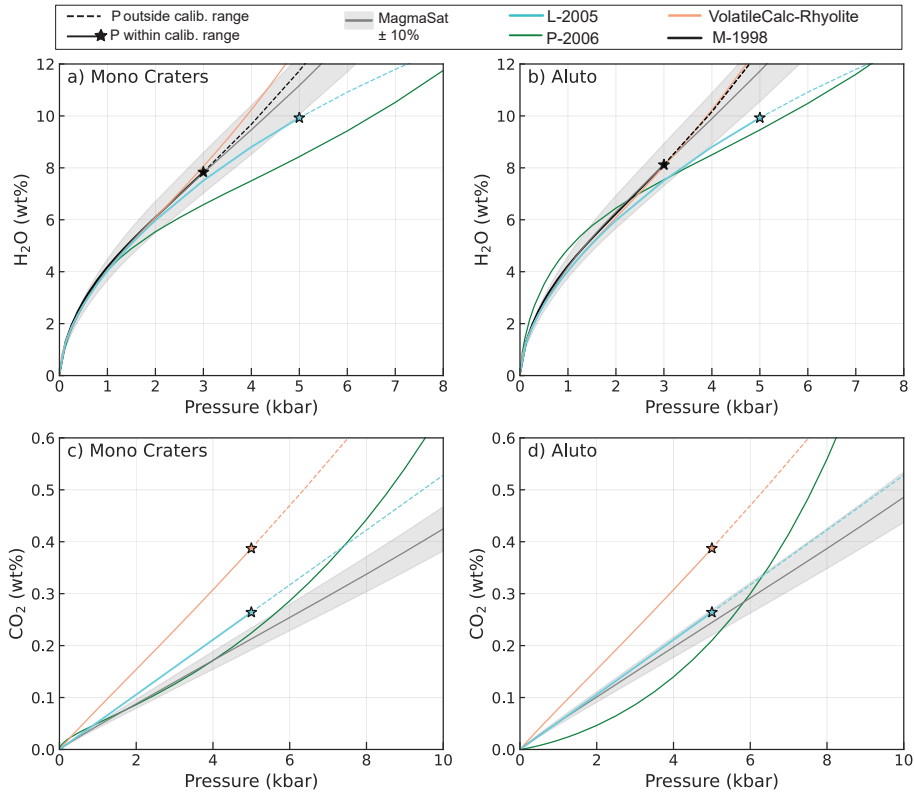
888 Interestingly, MagmaSat also underpredicts CO<sub>2</sub> concentrations at a given  
889 pressure relative to A-2019 by a factor of 0.6–0.7×, despite incorporating CO<sub>2</sub> ex-  
890 periments on Etna basalts from Lesne, Scaillet, Pichavant, and Beny (2011) and  
891 Iacono-Marziano et al. (2012) in its calibration dataset. This is a good example of  
892 the main pitfall of comprehensive models such as MagmaSat and P-2006 which can  
893 predict volatile solubility across the entire range of natural silicate melt composi-  
894 tions (Papale et al., 2006). For any specific melt composition, the model is highly  
895 unlikely to be as well tuned as models calibrated on melts from a specific volcanic  
896 center (e.g., Allison et al., 2019) or heavily weighted towards a specific region of  
897 compositional space (e.g., Iacono-Marziano et al., 2012, for alkaline basalts). Tuning  
898 MagmaSat to provide a better fit to Etna would almost certainly cause this model to  
899 show larger discrepancies for experiments conducted on different melt compositions.



### 900 **4.3 Silicic Compositions**

#### 901 **4.3.1 Pure $H_2O$**

902 All five  $H_2O$  models calibrated for silicic magmas (MagmaSat, P-2006, L-2005,  
903 VolatileCalc-Rhyolite and M-1998) predict very similar  $H_2O$  concentrations at <1-  
904 1.5 kbar for the Mono Craters rhyolite composition (Fig. 8a, Table 1). At higher  
905 pressures, P-2006, and to a much lesser extent L-2015, show a smaller increase in  
906  $H_2O$  solubility with pressure compared to MagmaSat, M-1998 and VolatileCalc-  
907 Rhyolite (the difference in  $H_2O$  solubility between models reaches  $\sim 4$  wt% at 5  
908 kbar). As discussed for basalt, the large discrepancies at higher pressures and  $H_2O$   
909 contents likely results from an absence of experimental constraints because of chal-  
910 lenges related to quenching melts with >6-9 wt%  $H_2O$  (Gavrilenko et al., 2019). The  
911 trajectories of models with compositional terms (M-1998, P-2006, MagmaSat) are  
912 very similar for Aluto.



**Figure 8.** Relationship between pure H<sub>2</sub>O (a-b) and pure CO<sub>2</sub> (c-d) solubility and pressure at 800°C for the Mono Lake rhyolite from Liu et al. (2005), and a peralkaline Rhyolite from Aluto in the East African Rift (Iddon and Edmonds, 2020). Models extrapolated beyond their calibrated pressure range are shown as dashed lines, with the colored star marking the recommended upper calibration limit.

913

#### 4.3.2 Mixed H<sub>2</sub>O-CO<sub>2</sub>

914

915

916

917

918

919

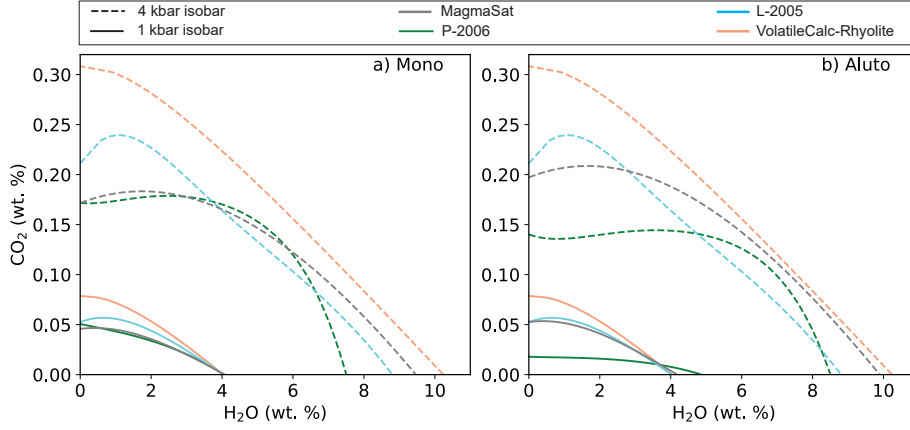
920

921

922

Differences in the treatment of H<sub>2</sub>O-CO<sub>2</sub> mixing for rhyolitic melts are more subtle than for basaltic compositions (Fig. 9). Unlike for basalts, the differences in isobar positions mostly result from large differences between the pure CO<sub>2</sub> solubility predicted by different models rather than different formulations of mixing. Only VolatileCalc-Rhyolite assumes ideal mixing of H<sub>2</sub>O-CO<sub>2</sub>, causing isobars to have a negative gradient at all pressures. L-2005 accounts for non-ideal mixing through the inclusion of a term for the mole fraction of H<sub>2</sub>O in the fluid in their expression for CO<sub>2</sub> solubility (equation 5). This empirical representation of non-ideality causes isobars to exhibit a prominent peak at low H<sub>2</sub>O contents (Fig. 9). MagmaSat and

923 P-2006, which include a fully non-ideal treatment of mixing, show a far less promi-  
 924 nent peak than L-2005. The slight up-tick in the P-2006 isobars at very low H<sub>2</sub>O  
 925 contents is much smaller than for basaltic compositions (e.g., Fig. 6).



**Figure 9.** 1 and 4 kbar isobars for Mono Craters (a) and the Aluto rhyolite (b) at 800°C.

### 926 4.3.3 Pure CO<sub>2</sub>

927 Unlike the relatively good correspondence between rhyolite H<sub>2</sub>O models (par-  
 928 ticularly at low pressures), there is substantial divergence between CO<sub>2</sub> models at all  
 929 pressures (Fig. 8c, d). For the Mono Craters rhyolite, VolatileCalc-Rhyolite predicts  
 930  $\sim 1.8\times$  more dissolved CO<sub>2</sub> than MagmaSat at 2 kbar, while Liu, P-2006 and Mag-  
 931 maSat plot relatively close to each other at  $<5$  kbar. As VolatileCalc-Rhyolite and  
 932 L-2005 have no compositional dependence, the model lines are identical for Mono  
 933 Craters and Aluto. MagmaSat predicts that the Aluto composition has slightly  
 934 higher CO<sub>2</sub> solubility at a given pressure compared to the Mono Lake composition  
 935 (factor of  $\sim 1.2\times$ ), so the discrepancy between MagmaSat, VolatileCalc-Rhyolite and  
 936 L-2005 is smaller for Aluto than Mono Craters.

937 The P-2006 model shows a substantially different trajectory for CO<sub>2</sub> vs. pres-  
 938 sure compared to the other three solubility models for both rhyolite compositions,  
 939 showing a strongly concave-up shape compared to the near linear trajectory of L-  
 940 2005 and VolatileCalc-Rhyolite, and the slightly concave-up shape of MagmaSat  
 941 (Fig. 8c-d). For Mono Craters, P-2006 predicts similar CO<sub>2</sub> solubility to MagmaSat  
 942 at  $<4$  kbar, but rapidly rises to higher CO<sub>2</sub> contents at higher pressures, predict-

943 ing almost as much dissolved CO<sub>2</sub> as VolatileCalc-Rhyolite at ~12 kbar (Fig. 8c).  
 944 For Aluto, the curvature of the P-2006 model is even more prominent, predicting  
 945 drastically lower CO<sub>2</sub> contents than all other models at <6 kbar, and then rapidly  
 946 rising, predicting higher CO<sub>2</sub> solubility than even VolatileCalc-Rhyolite at >9 kbar  
 947 (Fig. 8d). These large deviations between models, as well as the large errors on the  
 948 interaction terms for CO<sub>2</sub> solubility in MagmaSat (Table 3) demonstrate that while  
 949 H<sub>2</sub>O solubility in rhyolites is well constrained, more work is required to determine  
 950 the effect of melt composition on CO<sub>2</sub> solubility at a range of pressures and temper-  
 951 atures.

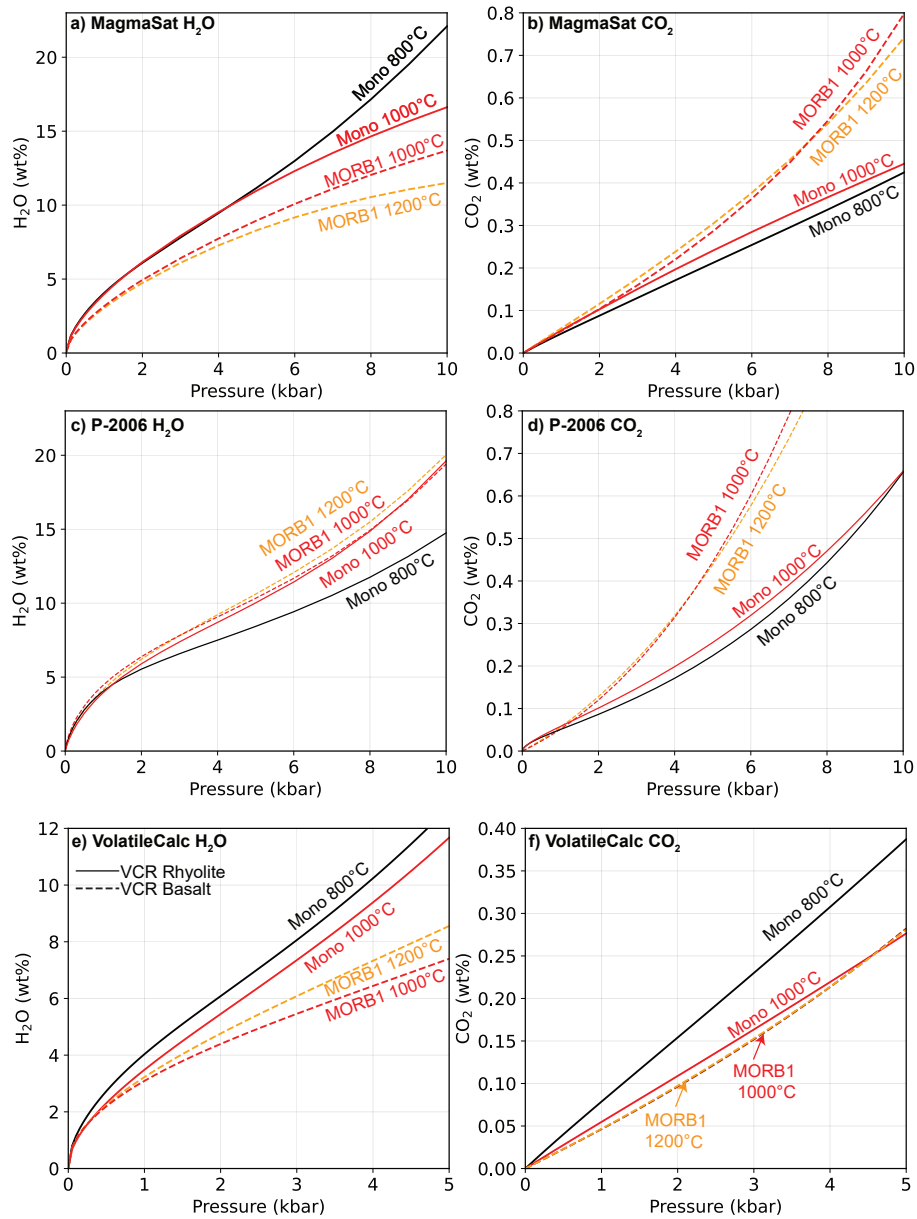
#### 952 **4.4 Comparisons between Basalts and Rhyolites**

953 In this section, we briefly discuss the differences in solubility between basalts  
 954 (using the MORB1 composition) and rhyolites (using the Mono Craters composi-  
 955 tion). To differentiate the effect of melt composition from temperature (because  
 956 basaltic melts tend to be hotter), we perform calculations at 800 and 1000°C for  
 957 Mono Craters, and 1000°C and 1200°C for MORB1.

958 When all solubility models are compared (4 applicable to rhyolites, 6 to  
 959 basalts), there is substantial overlap between curves calculated for MORB1 at  
 960 1200°C and Mono Craters at 800°C (compare Fig. 11a vs. Fig. 8a). To get around  
 961 this problem of large differences between models, we compare the predictions from  
 962 the three models which can be applied to both Rhyolites and Basalts: MagmaSat  
 963 (Fig. 10a-b), P-2006 (Fig. 10c-d) and VolatileCalc-Basalt and -Rhyolite (Fig. 10e-f).

964 MagmaSat and VolatileCalc (Rhyolite vs. Basalt) predict that Mono Craters  
 965 dissolves more H<sub>2</sub>O than MORB1 at realistic temperatures (800°C vs. 1200°C and  
 966 at a fixed temperature (1000°C). In contrast, P-2006 predicts that MORB1 dis-  
 967 solves more H<sub>2</sub>O than Mono at realistic temperatures (although their solubilities are  
 968 nearly identical when compared at 1000°C). For CO<sub>2</sub>, MagmaSat and P-2006 pre-  
 969 dict higher solubilities in MORB1, with the difference becoming more pronounced at  
 970 higher pressures, while VolatileCalc predicts that at realistic temperature, Mono  
 971 Craters has higher CO<sub>2</sub> solubility. When compared at the same temperature,  
 972 VolatileCalc predicts very similar CO<sub>2</sub> solubility for Mono Craters and MORB1.  
 973 These comparisons demonstrate that at <5 kbar, the difference in solubility between

974 basalts and rhyolites in each model easily overwhelmed by differences in predictions  
975 from different solubility models. These discrepancies are enhanced by the different  
976 sensitivities of these models to temperature (see section 5.3).



**Figure 10.** Comparison of solubility of basalt (represented by MORB1) to rhyolite (represented by Mono Craters) using MagmaSat (a-b), P-2006 (c-d) and VolatileCalc-Basalt and Rhyolite (d-e). Solubility curves are calculated for 1200°C and 1000°C for basalt, and 1000°C and 800°C for rhyolite.

## 5 Model Sensitivities

In this section, we explore the sensitivity of the different models to parameters such as temperature and variable proportions of H<sub>2</sub>O and CO<sub>2</sub>. Specifically, we consider how these inputs affect calculations of the pressure at which a melt inclusion was trapped (termed the saturation pressure). To calculate saturation pressures, the initial concentration of major and volatile elements as well as the temperature must be estimated at the time of melt inclusion entrapment. However, a number of processes, such as crystallization of the host mineral on the wall of the inclusion (termed post-entrapment crystallization, or PEC), growth of a vapour bubble or daughter phases within the inclusion, and diffusive re-equilibration with a changing carrier liquid composition can make it difficult to reconstruct initial major element and volatile contents (Lowenstern, 1995). Similarly, diffusive re-equilibration of the major elements in the melt inclusion and host mineral, as well as the errors associated with mineral-melt and melt-only thermometers, can lead to uncertainties in the entrapment temperature, which propagates into the saturation pressure. By investigating the effect of varying these parameters within realistic limits, insight can be gained into the uncertainties associated with estimating magma storage depths using melt inclusions.

### 5.1 Relationship between saturation pressure and dissolved H<sub>2</sub>O content

Melt inclusion H<sub>2</sub>O contents are vulnerable to diffusional re-equilibration with the melt surrounding the crystal (here termed the carrier melt), because of the fast diffusion rate of H<sup>+</sup> through silicate minerals (Portnyagin et al., 2008). H<sup>+</sup> diffusion is particularly fast in olivine (Gaetani et al., 2012), with melt inclusions losing significant amounts of water in hours to days (Bucholz et al., 2013). Thus, this discussion focuses on mafic compositions, where olivine-hosted melt inclusions are frequently analysed.

In relatively H<sub>2</sub>O-poor mafic systems such as MORs and ocean islands (e.g., Hawai'i), diffusive re-equilibration can increase melt inclusion H<sub>2</sub>O contents if crystals are mixed into more H<sub>2</sub>O-rich carrier melts (Hartley et al., 2015), or, more commonly, cause melt inclusion H<sub>2</sub>O contents to drop if the crystal is in contact

1008 with a carrier melt that has degassed its H<sub>2</sub>O upon eruption (Gaetani et al., 2012;  
 1009 Bucholz et al., 2013). To assess how uncertainty in initial H<sub>2</sub>O contents translates  
 1010 into errors on saturation pressures, we calculate saturation pressures for the MORB1  
 1011 melt composition with 200, 1000, and 3000 ppm CO<sub>2</sub> (representing melt inclusions  
 1012 trapped at low, medium and high pressures) for H<sub>2</sub>O contents between 0–1.5 wt%  
 1013 (Fig. 11a-c). CO<sub>2</sub> contents are held constant while H<sub>2</sub>O contents are varied, sim-  
 1014 ulating the changes undergone by melt inclusions during diffusive re-equilibration  
 1015 (which strongly affects H<sub>2</sub>O contents in the inclusion, but does not change the total  
 1016 CO<sub>2</sub> budget of the inclusion).

1017 The relationship between saturation pressure and dissolved H<sub>2</sub>O predicted by  
 1018 each solubility model is strongly dependent on the amount of CO<sub>2</sub> in the melt, and  
 1019 therefore the pressure. To quantify model sensitivity to H<sub>2</sub>O, saturation pressures  
 1020 calculated at H<sub>2</sub>O=1.5 wt% are divided by the saturation pressure calculated at  
 1021 H<sub>2</sub>O=0 wt%, representing the possible discrepancy between the calculated satura-  
 1022 tion pressure and the real saturation pressure for melt inclusions which have under-  
 1023 gone complete H<sup>+</sup> re-equilibration with a fully degassed erupted melt at 0 bar. At  
 1024 low pressures (200 ppm CO<sub>2</sub>), all models show a decrease in calculated saturation  
 1025 pressure with decreasing H<sub>2</sub>O contents, with entrapment pressures being 1.2–1.8×  
 1026 higher before complete H<sub>2</sub>O-loss (Fig. 11a). MagmaSat shows the strongest sensitiv-  
 1027 ity to H<sub>2</sub>O content, and both IM-2012 models the weakest.

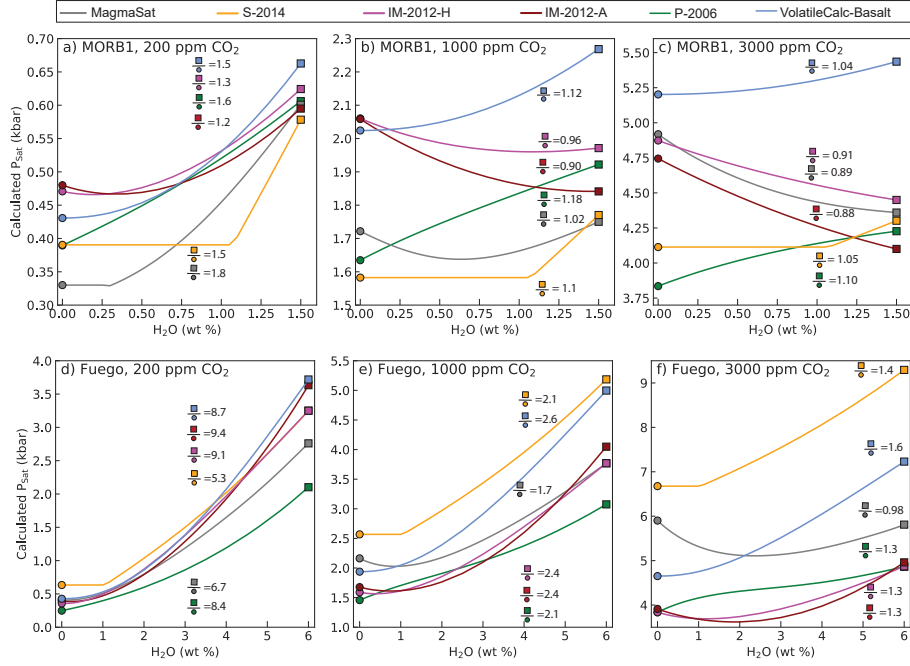
1028 At moderate pressures (1000 ppm CO<sub>2</sub>), loss of H<sub>2</sub>O causes a significantly  
 1029 smaller decrease in saturation pressure for VolatileCalc-Basalt, P-2006 and S-2014  
 1030 compared to the 200 ppm CO<sub>2</sub> scenario (Fig. 11b, 1.1–1.2×). Saturation pressures  
 1031 for 1000 ppm CO<sub>2</sub> calculated using MagmaSat and IM-2012-H first decrease, then  
 1032 increase with H<sub>2</sub>O loss. This is because these models predict that the maximum  
 1033 CO<sub>2</sub> solubility occurs at H<sub>2</sub>O contents at ~0.5–1.25 wt% (see Fig. 6).

1034 At higher pressures (3000 ppm CO<sub>2</sub>), saturation pressures from VolatileCalc-  
 1035 Basalt, P-2006 and S-2014 only drop by ~5-10% with progressive H<sub>2</sub>O-loss, while  
 1036 saturation pressures continually increase with progressive H<sub>2</sub>O-loss for IM-2012-H  
 1037 and -A and MagmaSat (because these models predict that maximum CO<sub>2</sub> solubility  
 1038 is found at H<sub>2</sub>O contents >1.5 wt% at these pressures; Fig. 6).



1039           Within a given suite of MORB or OIB melt inclusions, the range of measured  
1040 H<sub>2</sub>O contents, and the uncertainty involved in reconstructing initial H<sub>2</sub>O contents  
1041 following diffusional re-equilibration, is likely significantly smaller than the 1.5 wt%  
1042 H<sub>2</sub>O considered here (Koleszar et al., 2009; Sides et al., 2014a; Wieser et al., 2021).  
1043 Thus, except at low pressures (<1 kbar), uncertainties in saturation pressures due to  
1044 diffusive re-equilibration of H<sub>2</sub>O in relatively anhydrous systems are likely compara-  
1045 ble to the analytical errors associated with the measurements of volatile species by  
1046 FTIR or SIMS ( $\pm 5$ –10%), errors on each solubility model ( $\sim 10$ –20%), and signifi-  
1047 cantly smaller than the differences between solubility models.

1048           The higher H<sub>2</sub>O contents of melt inclusions from subduction zones ( $\sim 2$ –6  
1049 wt%; Plank et al., 2013) mean that substantially more H<sub>2</sub>O can be lost following  
1050 diffusive re-equilibration with a degassed carrier melt upon eruption. Additionally,  
1051 arc melt inclusions are vulnerable to diffusive re-equilibration during crustal stor-  
1052 age. This is because these relatively hydrous magmas saturate in a H<sub>2</sub>O-rich fluid  
1053 at high pressures in the crust. Thus, as a melt and its crystal cargo ascends from  
1054 a deeper storage reservoir to a shallower storage reservoir, significant quantities of  
1055 H<sub>2</sub>O will be degassed and the H<sub>2</sub>O contents of melt inclusions will rapidly diffusively  
1056 re-equilibrate with the new carrier melt composition (Gaetani et al., 2012). Even  
1057 if samples are rapidly quenched upon eruption (preventing syn-eruptive H<sub>2</sub>O diffu-  
1058 sion), only the volatile contents of melt inclusions trapped in the shallowest storage  
1059 reservoir can be reliably converted into saturation pressures (Gaetani et al., 2012).  
1060 This contrasts with more H<sub>2</sub>O-poor systems such as mid-oceanic ridges and oceanic  
1061 islands, where H<sub>2</sub>O only degasses in the upper few hundred metres of the crust, so  
1062 ascent to a shallower reservoir is not accompanied by a drop in melt H<sub>2</sub>O contents  
1063 (although diffusive re-equilibration could occur if the resident melts in the shallower  
1064 reservoir have different H<sub>2</sub>O contents; Hartley et al., 2015).



**Figure 11.** a-c) Relationship between saturation pressure and melt H<sub>2</sub>O content for H<sub>2</sub>O-poor melts (using the MORB1 composition at 1200°C). Three different melt CO<sub>2</sub> contents (200, 1000, 3000 ppm) are shown in part a, b and c respectively. The numbers on the graphs show the saturation pressure at 1.5 wt% H<sub>2</sub>O (square symbol) divided by the saturation pressure at 0 wt% H<sub>2</sub>O for each model. d-f) Sensitivity tests using the composition of a Fuego melt inclusion (Table 1) at 1000°C, and H<sub>2</sub>O contents between 0–6 wt%. Note that the y scale for parts a-c is significantly smaller than parts d-f.

1065 To investigate the effect of H<sub>2</sub>O re-equilibration on melt inclusion saturation  
 1066 pressures in arcs, we repeat the sensitivity test described above, using the major  
 1067 element composition of a Fuego melt inclusion with 49.7 wt% SiO<sub>2</sub> from Lloyd et  
 1068 al. (2013, Table 1) and H<sub>2</sub>O contents between 0–6 wt% (Fig. 11d-f). For melt in-  
 1069 clusions with 200 ppm CO<sub>2</sub>, complete diffusive loss of H<sub>2</sub>O may result in saturation  
 1070 pressures being underestimated by a factor of ~5–10×. Even for melts with 3000  
 1071 ppm CO<sub>2</sub> (the highest pressure regime examined, and thus the best case scenario),  
 1072 diffusive loss can affect saturation pressures by factors of 0.98–1.6× (similar in mag-  
 1073 nitude to the sensitivity displayed by H<sub>2</sub>O-poor melts at the lowest pressures; Fig.  
 1074 11a vs. f). Only saturation pressures calculated in MagmaSat for the most CO<sub>2</sub>-rich  
 1075 melts display variations with variable H<sub>2</sub>O-loss similar in magnitude to analytical

1076 errors. Thus, it is extremely important to determine whether melt inclusions have  
1077 undergone H<sub>2</sub>O-loss during ascent to a shallower reservoir or syn-eruptive degassing  
1078 before using saturation pressures to deduce magma storage depths in H<sub>2</sub>O-rich vol-  
1079 canic systems.

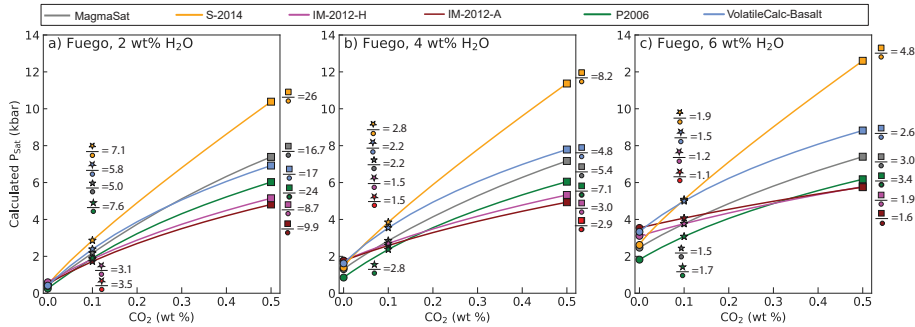
## 1080 **5.2 Relationship between saturation pressure and dissolved CO<sub>2</sub>** 1081 **content**

1082 Estimating the initial CO<sub>2</sub> contents of melt inclusions is also challeng-  
1083 ing. While the total CO<sub>2</sub> content of the inclusion is not affected by diffusive re-  
1084 equilibration, CO<sub>2</sub> may be partitioned from the melt phase into a vapour bubble.  
1085 Cooling following melt inclusion entrapment is accompanied by the formation of a  
1086 denser mineral phase from a less dense silicate melt, and differential thermal con-  
1087 traction of the melt and crystal. These processes cause the internal pressure of the  
1088 inclusion to drop (Steele-Macinnis et al., 2011; Maclennan, 2017), driving the nu-  
1089 cleation and growth of a vapour bubble. This may be enhanced by the diffusive  
1090 loss of H<sub>2</sub>O, which also causes a pressure drop in the melt inclusion because of the  
1091 high molar volume but low molecular weight of H<sub>2</sub>O (Gaetani et al., 2012; Aster et  
1092 al., 2016). A drop in pressure, combined with a decrease in the solubility of CO<sub>2</sub>  
1093 in the melt phase because of changes to the major element composition accompa-  
1094 nying post-entrapment crystallization, causes CO<sub>2</sub> to partition strongly into the  
1095 vapour bubble (L. R. Moore et al., 2015; Steele-Macinnis et al., 2011; Maclennan,  
1096 2017; Wieser et al., 2021). A number of recent studies have quantified the amount of  
1097 CO<sub>2</sub> in vapour bubbles using Raman Spectroscopy, and demonstrated that between  
1098 15–99% of the total CO<sub>2</sub> budget of the inclusion may be held within the vapour  
1099 bubble (Hartley et al., 2014; L. R. Moore et al., 2015; Wieser et al., 2021; Allison  
1100 et al., 2021). This means that a large proportion of literature melt inclusion data,  
1101 which only measured the CO<sub>2</sub> content of the glass phase, may have significantly  
1102 underestimated initial CO<sub>2</sub> contents (and therefore saturation pressures).

1103 In relatively H<sub>2</sub>O-poor systems such as Hawai'i and Iceland, where melt in-  
1104 clusion CO<sub>2</sub> contents have the dominant control on saturation pressures (shown by  
1105 the near horizontal slopes of most model isobars at low H<sub>2</sub>O contents; Fig. 6), it is  
1106 readily apparent that saturation pressures will be significantly underestimated if a  
1107 CO<sub>2</sub>-rich vapour bubble is not measured. In arcs, H<sub>2</sub>O contents inferred from melt

1108 inclusions or mineral hygrometers are sometimes used to place first order constraints  
 1109 on saturation pressures (e.g. Plank et al., 2013; Blundy & Cashman, 2005; Goltz  
 1110 et al., 2020). However, even in very H<sub>2</sub>O-rich melts, the non-vertical orientation of  
 1111 isobars at high H<sub>2</sub>O contents indicates that CO<sub>2</sub> contents still have an important  
 1112 role in determining the saturation pressure (Fig. 6). Additionally, only a very small  
 1113 number of studies have measured CO<sub>2</sub> in melt inclusion vapour bubbles from arc  
 1114 systems (L. R. Moore et al., 2015; Aster et al., 2016; Venugopal et al., 2020; Mironov  
 1115 et al., 2020). Thus, it is vital to determine the effect of CO<sub>2</sub> on saturation pressures  
 1116 in H<sub>2</sub>O-rich systems.

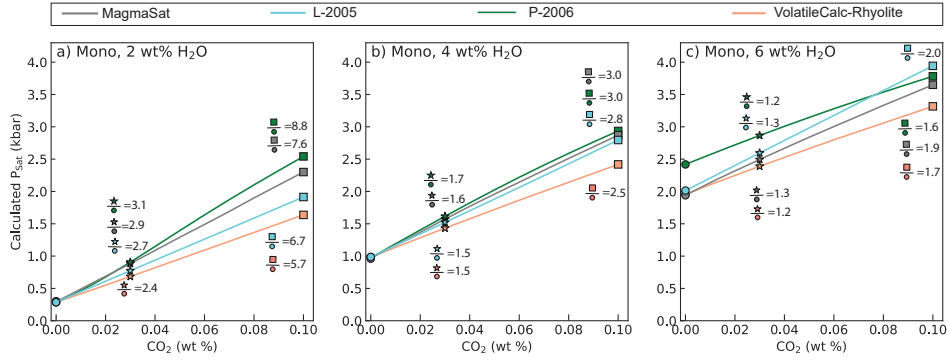
1117 Using a similar method to that for H<sub>2</sub>O discussed above, we calculate satura-  
 1118 tion pressures for the composition of a Fuego melt inclusion from Lloyd et al. (2013)  
 1119 with varying CO<sub>2</sub> and H<sub>2</sub>O contents. The mean melt inclusion glass CO<sub>2</sub> content  
 1120 from this melt inclusion suite was 340 ppm (range of 59–786 ppm). However, Raman  
 1121 analyses of vapour bubbles in the same sample set by L. R. Moore et al. (2015) re-  
 1122 veals that 993–4776 ppm of CO<sub>2</sub> has migrated from the glass phase into the vapour  
 1123 bubble following melt inclusion entrapment. Thus, we calculate saturation pressures  
 1124 for CO<sub>2</sub> contents between 0–5000 ppm at 1000°C for 2, 4 and 6 wt% H<sub>2</sub>O respec-  
 1125 tively (after Plank et al., 2013).



**Figure 12.** Relationship between saturation pressures and melt CO<sub>2</sub> contents for H<sub>2</sub>O-rich melts (using the composition of a Fuego melt inclusion at 1000°C; Table 1). Three different melt H<sub>2</sub>O contents (2, 4 and 6 wt%) are shown in part a, b and c) respectively. The numbers on the graphs show the saturation pressure at 5000 ppm CO<sub>2</sub> (square symbol) divided by the saturation pressure at 0 ppm CO<sub>2</sub> (circle), and the saturation pressure at 1000 ppm (star symbol) divided by the saturation pressure at 0 ppm CO<sub>2</sub> (circle) for each model.

1126 S-2014 is most sensitive to CO<sub>2</sub> content, and IM-2012-H and -A the least  
1127 sensitive. With increasing H<sub>2</sub>O, the change in saturation pressure with increasing  
1128 CO<sub>2</sub> becomes smaller, but is still significant (Fig. 12). For example, calculating a  
1129 H<sub>2</sub>O-only saturation pressure for a melt inclusion with H<sub>2</sub>O=6 wt% in MagmaSat  
1130 would underestimate magma storage depths by a factor of 1.5 if the melt inclusion  
1131 had 1000 ppm CO<sub>2</sub>, and a factor of 3 if the inclusion had 5000 ppm CO<sub>2</sub>. For a melt  
1132 inclusion with 4 wt% H<sub>2</sub>O, H<sub>2</sub>O-only saturation pressures underestimate by a factor  
1133 of ~2.2 for 1000 ppm CO<sub>2</sub>, and 5.4 for 5000 pm CO<sub>2</sub>. These variations in satura-  
1134 tion pressure overwhelm the other errors associated with melt inclusion barometry  
1135 (e.g., uncertainty in crustal density profiles, analytical errors associated with volatile  
1136 measurements by FTIR or SIMS, differences between solubility models). Further  
1137 investigation of the prevalence of CO<sub>2</sub>-rich vapour bubbles in arc lavas is clearly  
1138 required to have confidence in published barometric estimates in studies which did  
1139 not measure the vapour bubbles, or used mineral hygrometers.

1140 Saturation pressures in rhyolitic magmas are also very sensitive to melt CO<sub>2</sub>  
1141 contents (Fig. 13). For example, saturation pressures calculated for 1000 ppm CO<sub>2</sub>  
1142 vs. 0 ppm CO<sub>2</sub> differ by factors of 5.7–8.8× for 2 wt% H<sub>2</sub>O, and 1.6–2× for 6 wt%  
1143 H<sub>2</sub>O. Even saturation pressures calculated for 300 ppm CO<sub>2</sub> (0.03 wt%) vs 0 ppm  
1144 CO<sub>2</sub> are a factor of ~2–3× higher for 2 wt% H<sub>2</sub>O, and 1.2–1.3× higher for 6 wt%  
1145 H<sub>2</sub>O. The strong effect of CO<sub>2</sub> on saturation pressure is important to recognise  
1146 when calculating saturation pressures using only melt H<sub>2</sub>O contents, such as studies  
1147 using mineral-melt hygrometers (Waters & Lange, 2013), or volatiles-by-difference  
1148 methods to estimate H<sub>2</sub>O contents of melt inclusions. It is also interesting to note  
1149 that, to our knowledge, there are no published Raman measurements of CO<sub>2</sub> in  
1150 vapour bubbles which grew after melt inclusion entrapment in dacitic-rhyolitic melt  
1151 compositions. While the extremely low CO<sub>2</sub> contents of many rhyolitic melt inclu-  
1152 sions are commonly interpreted to result from shallow crustal storage, it is becoming  
1153 increasingly recognised that mafic melt inclusions with CO<sub>2</sub> below detection limit  
1154 contain large quantities of CO<sub>2</sub> in the vapour bubble (Wieser et al., 2021). Thus, ex-  
1155 amination of vapour bubbles in melt inclusions from more silicic systems (e.g., Fig. 1  
1156 of Lowenstern, 2001) is likely warranted, to rule out the possibility that these melts  
1157 crystallized at greater depths than calculated using measurements of CO<sub>2</sub> hosted in  
1158 just the glass phase.



**Figure 13.** Sensitivity of saturation pressures to melt  $\text{CO}_2$  contents at three different  $\text{H}_2\text{O}$  contents (2, 4 and 6 wt%) for the Mono Craters rhyolite. Ratios of saturation pressures at 0.03 wt%  $\text{CO}_2$  (300 ppm; star symbol) vs. 0 wt% (circle), and 0.1 wt% (1000 ppm; square symbol) vs. 0 wt%  $\text{CO}_2$  are shown on the figure.

### 1159 5.3 Sensitivity to Temperature

1160 The temperature of the melt at the time of entrapment is another source of un-  
 1161 certainty when calculating saturation pressures, as melt or mineral-melt thermome-  
 1162 ters are relatively imprecise. For example, the liquid MgO thermometer presented  
 1163 in eq. 13 of Putirka (2008) has a standard error of  $\pm 71^\circ\text{C}$ , while the clinopyroxene-  
 1164 liquid thermometer presented in their equation 28b has a standard error of  $\pm 48^\circ\text{C}$ .  
 1165 Changes in the major element compositions of the melt inclusion during PEC and  
 1166 diffusive  $\text{H}_2\text{O}$ -loss can also introduce errors when estimating entrapment tempera-  
 1167 tures (as most thermometers are highly sensitive to the MgO and  $\text{H}_2\text{O}$  content of  
 1168 the melt). Additionally, almost all solubility experiments are performed at supra-  
 1169 liquidus conditions, while melt inclusion formation must take place at sub-liquidus  
 1170 conditions, so extrapolation to lower temperatures is an unfortunate necessity.

1171 To investigate the sensitivity of different models to temperature, we calculate  
 1172 the 0.5 and 2 kbar isobars for the MORB1 composition at  $1000^\circ\text{C}$  and  $1400^\circ\text{C}$ . Only  
 1173 S-2014 shows no temperature dependency, because there is neither a temperature  
 1174 or fugacity term in their equations. Interestingly, there is considerable disagree-  
 1175 ment between the other models as to whether a hotter magma dissolves more or less  
 1176 volatiles. MagmaSat and P-2006 predict an increase in pure  $\text{CO}_2$  solubility with  
 1177 increasing temperature, while VolatileCalc-Basalt and IM-2012-A and -H predict a

1178 much smaller decrease (Fig. 14a-c). In all models but IM-2012-A, isobars calculated  
 1179 for lower temperatures intercept the x axis ( $\text{CO}_2=0$  wt%) at higher  $\text{H}_2\text{O}$  contents,  
 1180 so the temperature dependency of  $\text{H}_2\text{O}$  solubility is opposite to that for  $\text{CO}_2$  solu-  
 1181 bility. To visualize the effect of these trends on calculated saturation pressures for  
 1182 the MORB1 composition, the calculated saturation pressures for melts with volatile  
 1183 contents represented by the yellow and cyan stars on Figure 14b are plotted against  
 1184 temperature (between 1000–1400°C; Fig. 14c-d). MagmaSat and P-2006 show the  
 1185 strongest temperature sensitivity, with a slope opposite to that of the more subtle  
 1186 changes predicted by VolatileCalc-Basalt and IM-2012.

1187 The lack of consensus as to whether increasing temperature increases or de-  
 1188 creases the solubility of  $\text{H}_2\text{O}$  and  $\text{CO}_2$  indicates that this effect is relatively subtle,  
 1189 and overwhelmed by analytical errors associated with measuring experimental prod-  
 1190 ucts (and other sources of experimental scatter; e.g., Fig. 16a-b). This makes it  
 1191 very difficult for empirical models to fully constrain the temperature sensitivity,  
 1192 particularly given that the experiments conducted by any given study are usually  
 1193 performed at a single temperature. For example, all the experiments used to cali-  
 1194 brate the VolatileCalc-Basalt model were conducted at 1200°C, so the temperature  
 1195 dependency of this model results from the fugacity function, and  $1/T$  terms from  
 1196 Dixon et al. (1995), rather than experimental observations.

1197 IM-2012 is calibrated on experiments mostly performed at 1200–1300°C (with  
 1198 a few spanning 1100–1400°C). Their empirical expressions contain  $C_{\text{H}_2\text{O}} \times \frac{P}{T}$  and  
 1199  $C_{\text{CO}_2} \times \frac{P}{T}$  terms in their expressions for  $\text{H}_2\text{O}$  and  $\text{CO}_2$  solubility respectively (equa-  
 1200 tion 11 and 15), where  $C_{\text{H}_2\text{O}}$  and  $C_{\text{CO}_2}$  are empirically-derived constants, and P  
 1201 and T are pressure and temperature. In the hydrous model,  $C_{\text{H}_2\text{O}}$  is negative (-  
 1202  $0.02 \pm 0.02$ ), so  $\text{H}_2\text{O}$  solubility increases with increasing temperature, while in the  
 1203 anhydrous model  $C_{\text{H}_2\text{O}}$  is positive ( $0.02 \pm 0.02$ ), so  $\text{H}_2\text{O}$  solubility decreases with  
 1204 increasing temperature (Fig. 14). As the magnitude of these coefficients is small, the  
 1205 temperature effect on  $\text{H}_2\text{O}$  solubility is small, and only visible at higher pressures  
 1206 (because of the P part of these terms; Fig. 14a vs. c). It is also worth noting that  
 1207 these coefficients in both models are within error of zero, showing that the experi-  
 1208 mental dataset used to calibrate this model showed very little evidence for a change  
 1209 in  $\text{H}_2\text{O}$  solubility with temperature. In both the hydrous and anhydrous models,

1210  $C_{CO_2}$  is positive ( $0.12\pm 0.02$  and  $0.14\pm 0.02$  respectively) and larger in magnitude  
 1211 than  $C_{H_2O}$ , so  $CO_2$  solubility decreases with increasing temperature (see Fig. 14c).

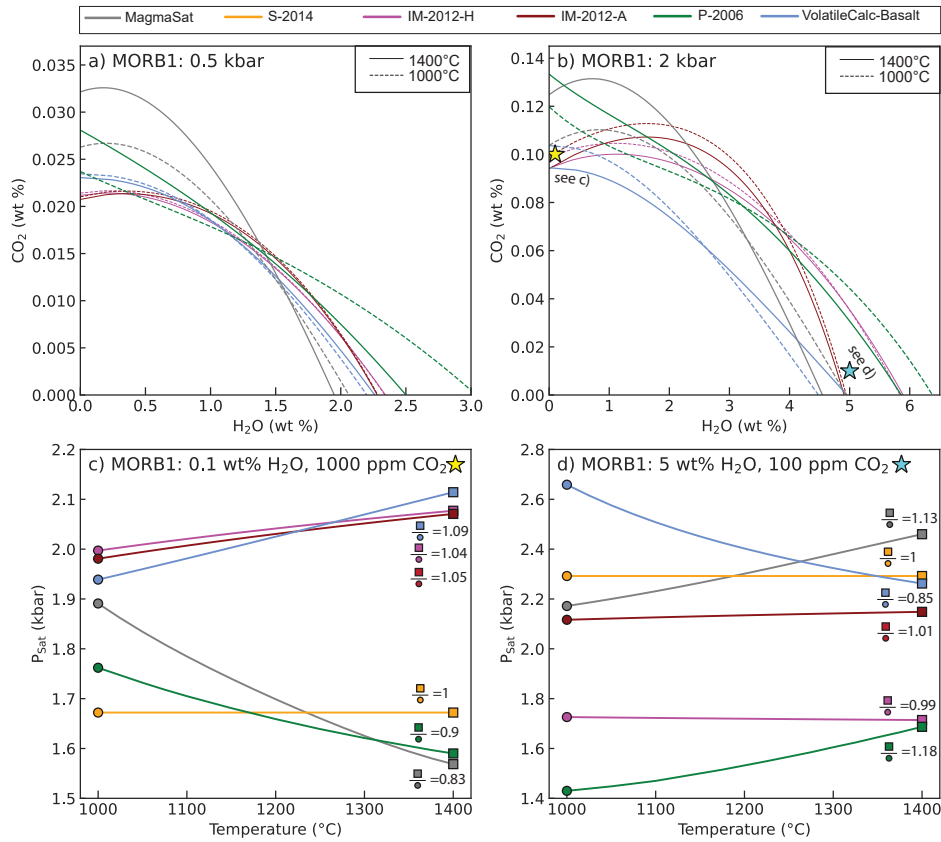
1212 The approach taken by S-2014 and A-2019 is an interesting alternative when  
 1213 constructing solubility models. While S-2014 is calibrated on experiments conducted  
 1214 between 1200–1300°C for  $CO_2$ , and 1200–1250°C for  $H_2O$ , their solubility equa-  
 1215 tions contain no temperature or fugacity term. Instead, these authors suggest that  
 1216 the  $H_2O$  model should ideally be used between 1150–1250°C. They perform addi-  
 1217 tional tests on experiments between 1050–1400°C not used in the calibration, and  
 1218 show that their model predicts  $H_2O$  solubility within  $\pm 10\%$  for 78% of experiments  
 1219 for this wider temperature range. The S-2014 testing dataset for  $CO_2$  only has a  
 1220 slightly different temperature range than the calibration dataset (1170–1250°C vs.  
 1221 1200–1250°C), so they do not suggest an expanded temperature range for  $CO_2$ . Sim-  
 1222 ilarly, the spreadsheet for A-2019 (and the implementation of this model in VESI-  
 1223 cal) performs all calculations at 1200°C, regardless of the user-input temperature.  
 1224 Allison et al. (2019) suggest that this approach is likely valid between 1000–1400°C.

1225 Unlike empirical models, the temperature sensitivity of P-2006 and MagmaSat  
 1226 arises from the entropy differences between melt and fluid components. Given the  
 1227 limited experimental evidence for changes in solubility with temperature, the direc-  
 1228 tionality inferred by thermodynamical models is more likely to be correct, as the  
 1229 relative entropy differences between components are easier to constrain than decon-  
 1230 voluting subtle differences between the solubility of  $CO_2$  and  $H_2O$  in experiments  
 1231 run at different temperatures, and because all empirical models are being extrapo-  
 1232 lated to lower temperatures than the supra-liquidus experiments used to calibrate  
 1233 them. However, the fact that P-2006 doesn't account for volatile speciation for ei-  
 1234 ther  $H_2O$  or  $CO_2$ , and MagmaSat doesn't account for the two  $H_2O$  species in the  
 1235 melt means that these thermodynamic temperature sensitivities may also need further  
 1236 interrogation.

1237 Overall, although the differences between models is interesting, and important  
 1238 to recognise when extrapolating beyond the range of the calibration dataset, the  
 1239 uncertainty that temperature variations introduce to saturation pressure calcula-  
 1240 tions are relatively small for mafic melts. For example, an uncertainty in the initial  
 1241 entrapment temperature of  $\pm 100^\circ C$  introduces an uncertainty similar to that asso-



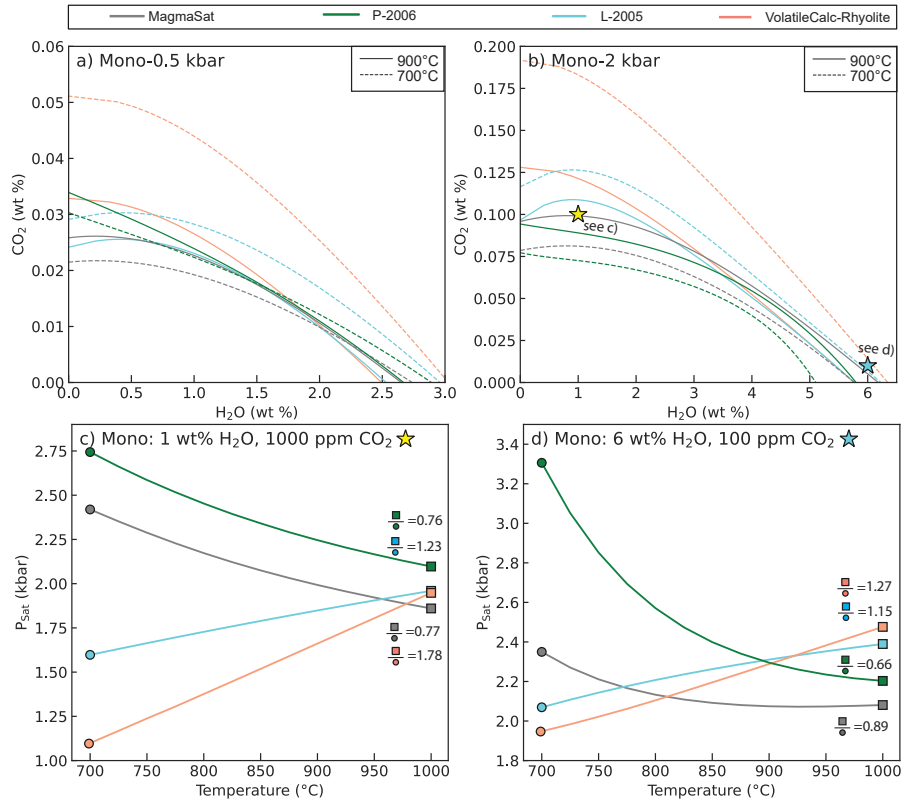
1242 ciated with in-situ measurements of melt inclusion volatile contents ( $\pm 5\text{--}10\%$ ; Fig.  
 1243 14).



**Figure 14.** Evaluating model sensitivity to temperature (using the MORB1 composition).  
 a-b) Isobars evaluated at 1000 and 1400°C and 0.5 and 2 kbar. c) Relationship between saturation  
 pressure and temperature for a melt with 1000 ppm CO<sub>2</sub>, 0.1 wt% H<sub>2</sub>O (yellow star on a),  
 d) 100 ppm CO<sub>2</sub>, 5 wt% H<sub>2</sub>O (cyan star on b). Ratios of saturation pressures at 1400°C (square  
 symbol) to 1000°C (circle symbol) are shown on the figure.

1244 Similarly, temperature sensitivity in rhyolitic melts was evaluated by calcu-  
 1245 lating isobars at 0.5 and 2 kbar for 700 and 900°C using the Mono Craters rhyolite  
 1246 composition. As for the basaltic example, the directionality and magnitude of effect  
 1247 of temperature on saturation pressures for melts with volatile contents indicated by  
 1248 the colored stars is shown in Fig. 15c-d for temperatures between 700 and 1000°C.  
 1249 VolatileCalc-Rhyolite shows the strongest temperature sensitivity, predicting that  
 1250 the solubility of CO<sub>2</sub> decreases with increasing temperature. L-2005 also predicts  
 1251 decreasing CO<sub>2</sub> solubility with increasing temperature, although this effect is smaller

1252 than in VolatileCalc-Rhyolite. Decreasing solubility of molecular CO<sub>2</sub> with increas-  
 1253 ing temperature was demonstrated experimentally by Fogel and Rutherford (1990).  
 1254 In contrast, the two thermodynamic models, P-2006 and MagmaSat, predict that  
 1255 pure CO<sub>2</sub> solubility increases with increasing temperature. VolatileCalc-Rhyolite  
 1256 also predicts that H<sub>2</sub>O solubility decreases with increasing temperature, but the  
 1257 effect is smaller than for CO<sub>2</sub>.



**Figure 15.** Sensitivity of saturation pressures for the Mono Craters rhyolite to temperature. a-b) Isobars calculated for different solubility models at 700 and 900°C and 0.5 and 2 kbar. c) Relationship between saturation pressure and temperature for a melt with 1000 ppm CO<sub>2</sub>, 1 wt% H<sub>2</sub>O (yellow star in b), d) 100 ppm CO<sub>2</sub>, 6 wt% H<sub>2</sub>O (cyan star in b). Ratios of saturation pressures at 900°C (square symbol) to 700°C (circle symbol) are shown on the figure.

1258 It is noteworthy that the temperature sensitivity of CO<sub>2</sub> solubility predicted  
 1259 by L-2005 and VolatileCalc-Rhyolite is much greater than that shown by any of the  
 1260 basaltic models (Fig. 14d-e vs. Fig 15d-e), and significant considering other sources  
 1261 of error associated with saturation pressure calculations. MagmaSat and P-2006 also

1262 show a far greater sensitivity to H<sub>2</sub>O solubility between 700–850°C in rhyolites than  
1263 any of the basaltic models between 1000–1400°C (Fig. 14f vs. Fig 15f), although the  
1264 sensitivity decreases between 800–900°C.

1265 Given the contrasting behaviour of empirical and thermodynamic models,  
1266 and the relatively strong effect of temperature in rhyolitic melts, we suggest that  
1267 users proceed with caution when extrapolating empirical models to temperatures  
1268 significantly lower or higher than the calibration temperature of each model. It may  
1269 be best to use empirical models at the calibration temperature (e.g., 1200°C for  
1270 VolatileCalc-Basalt, 850°C for VolatileCalc-Rhyolite), which is the approach used by  
1271 in the models of S-2014 and A-2019, rather than introduce a temperature sensitivity  
1272 with the wrong sign. This is discussed in further detail for VolatileCalc-Rhyolite in  
1273 section 6.

## 1274 **6 Intermediate Compositions**

1275 In this section, we compare the predictions of different solubility models for in-  
1276 termediate melt compositions (andesites to dacites). Lavas with these compositions  
1277 are dominant within subduction zones, and volcanoes erupting these compositions  
1278 are extremely hazardous. Yet, there is a notable paucity of solubility experiments  
1279 for andesitic and dacitic compositions relative to basalts and rhyolites (Fig. 2; King  
1280 & Holloway, 2002; Botcharnikov et al., 2006). This section builds on the sensitiv-  
1281 ity tests performed in section 5 to evaluate possible discrepancies between model  
1282 outputs and experimental constraints.

1283 The calibration dataset of MagmaSat has the broadest coverage of andesitic-  
1284 dacitic compositions of all the models described here (although it is far from exten-  
1285 sive). While there are a number of pure H<sub>2</sub>O experiments, MagmaSat only includes  
1286 one pure CO<sub>2</sub> experiment on an andesitic melt (conducted at 1 GPa; King & Hol-  
1287 loway, 2002), and no pure CO<sub>2</sub> experiments on dacitic melts. Similarly for mixed  
1288 H<sub>2</sub>O-CO<sub>2</sub>, the calibration dataset for andesitic melts includes only four experiments  
1289 from King and Holloway (2002), 21 from Botcharnikov et al. (2006) and three from  
1290 Botcharnikov et al. (2007). Dacitic liquids are represented by the 12 experiments  
1291 on mixed H<sub>2</sub>O-CO<sub>2</sub> solubility by Behrens et al. (2004). As the P-2006 model had a  
1292 decade fewer experimental constraints available for calibration, it only includes the

1293 one pure CO<sub>2</sub> and four mixed H<sub>2</sub>O-CO<sub>2</sub> andesitic experiments of King and Holloway  
 1294 (2002). The IM-2012 model includes two pure H<sub>2</sub>O experiments that lie within the  
 1295 andesite field on a TAS diagram, but no H<sub>2</sub>O-CO<sub>2</sub> experiments, and no experiments  
 1296 in the dacitic field. None of the other models contain any andesitic or dacitic melts  
 1297 in their calibration datasets.

### 1298 **6.1 Comparing solubility models to experimental products**

1299 The suitability of different solubility models for andesitic-dacitic compositions  
 1300 can be evaluated by calculating isobars using the melt compositions, pressures and  
 1301 temperatures of different experimental studies, and comparing these isobars to mea-  
 1302 surements of dissolved volatile contents in experimental products (similar to the  
 1303 method used in the supplementary material of Ghiorso & Gualda, 2015). The 2  
 1304 and 5 kbar andesitic experiments of Botcharnikov et al. (2006) are shown in Fig.  
 1305 16a-b, the 1, 2, and 5 kbar dacitic experiments of Behrens et al. (2004) are shown in  
 1306 Fig. 16c-e, and the 10 kbar andesite experiments of King and Holloway (2002) are  
 1307 shown in Fig. 16f. Additional isobar diagrams for the 3–12 kbar basaltic-andesite  
 1308 experiments of Mangan et al. (2021) are shown in the Supporting Information.

1309 Isobar diagrams show that S-2014 significantly underpredicts CO<sub>2</sub> for all ex-  
 1310 periments except the most H<sub>2</sub>O-poor composition of King and Holloway (2002).  
 1311 L-2005, VolatileCalc-Basalt and -Rhyolite also mostly plot to lower CO<sub>2</sub> contents  
 1312 than experimental products. IM-2012-H and -A do a reasonable job of recreating  
 1313 the most H<sub>2</sub>O-poor experiments at <5 kbar, but curve rapidly down to intercept the  
 1314 x axis at lower H<sub>2</sub>O contents than experimental products and other models. Mag-  
 1315 maSat is a good match to experimental data in Fig. 16d, e and f, but plots to lower  
 1316 CO<sub>2</sub> contents than experiments in Fig. 16 a, b and c. Using Fe<sup>3+</sup> proportions best  
 1317 representing the experimental conditions, P-2006 only passes through experimental  
 1318 data on Figs. 16a-b, and e, and plots to significantly lower CO<sub>2</sub> contents than exper-  
 1319 iments (lower than MagmaSat) on Fig. 16c and d. P-2006 is a better match to most  
 1320 experiments if Fe<sup>3+</sup>/Fe<sub>T</sub>=0, but overpredicts CO<sub>2</sub> solubility at 10 kbar for all redox  
 1321 states. Interestingly, none of the available models recreate the near-flat trajectory of  
 1322 dissolved CO<sub>2</sub> contents with increasing H<sub>2</sub>O from Botcharnikov et al. (2006).

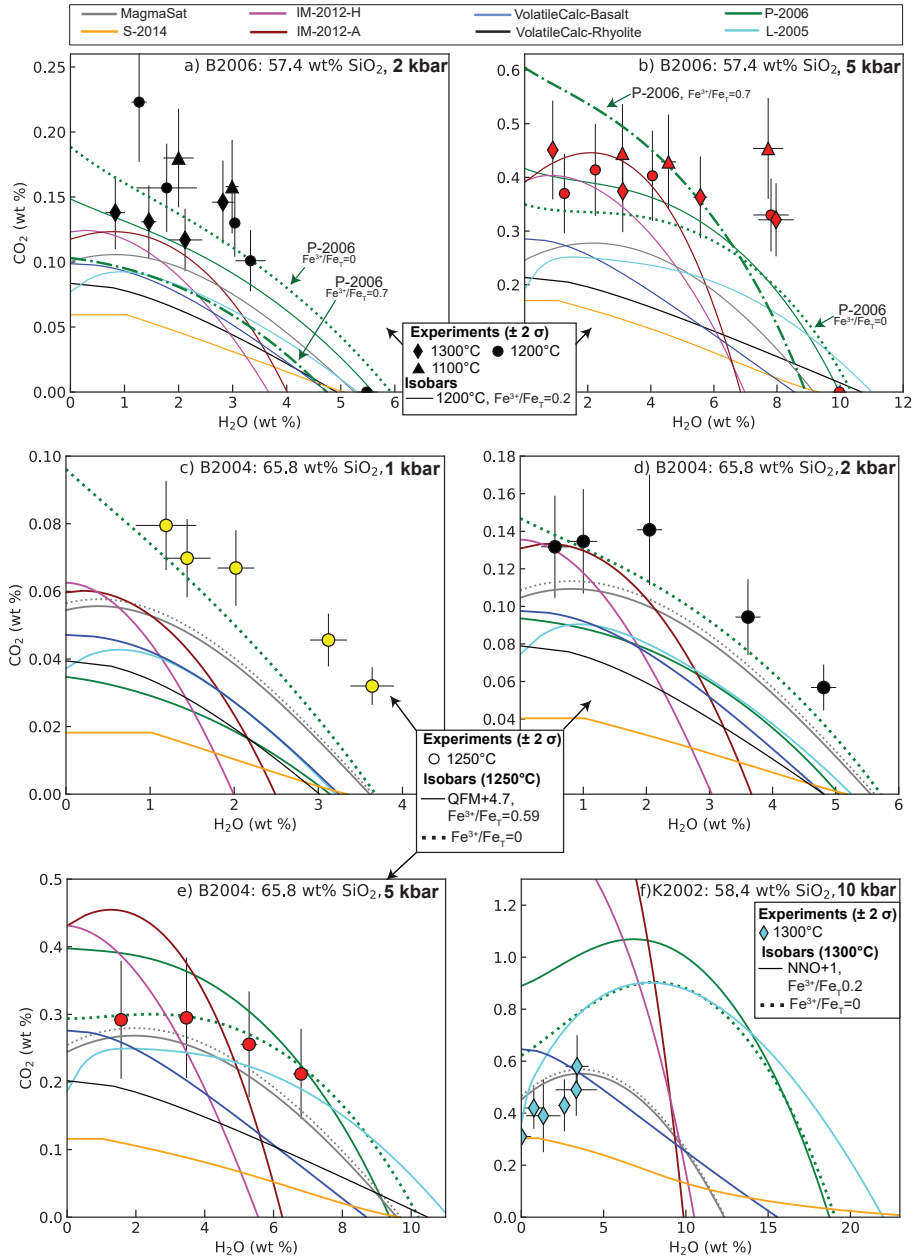
1323 The underprediction of CO<sub>2</sub> solubility by S-2014 is a good example of the  
1324 dangers of extrapolating models accounting for the effect of melt composition using  
1325 empirical expressions beyond the compositional range of the calibration dataset. The  
1326 S-2014 model expresses CO<sub>2</sub> solubility as a function of the composition parameter,  
1327  $\Pi^*$ , with CO<sub>2</sub> solubility increasing as an exponential function of  $\Pi^*$  at a given pres-  
1328 sure (equation 16). The melt compositions for the three sets of experimental studies  
1329 shown in Fig. 16 all plot to much lower  $\Pi^*$  values than any of the melts in the cal-  
1330 ibration dataset (orange diamonds; Fig. 17a). These low  $\Pi^*$  values mean that the  
1331 S-2014 model predicts very low CO<sub>2</sub> solubilities. However, CO<sub>2</sub> solubility for melts  
1332 with  $\Pi^*$  values outside the range of the calibration dataset may not follow the same  
1333 exponential function of this parameter as melt compositions within the calibration  
1334 range. Additionally, the exponential dependency of CO<sub>2</sub> solubility on  $\Pi^*$  incorpo-  
1335 rated by S-2014 likely breaks down in more evolved melts, because  $\Pi^*$  represents the  
1336 ability of the melt to form carbonate-bearing species, while more evolved melts con-  
1337 tain increasing proportions of molecular CO<sub>2</sub> (Shishkina et al., 2014). For example,  
1338 the proportion of molecular CO<sub>2</sub> to carbonate species varies between 0–4 wt% in the  
1339 experiments of Botcharnikov et al. (2006), and 3–30 wt% in the dacitic experiments  
1340 of Behrens et al. (2004).

1341 VolatileCalc-Basalt and VolatileCalc-Rhyolite underpredict CO<sub>2</sub> solubility  
1342 for all intermediate experiments, with VolatileCalc-Rhyolite predicting lower pure  
1343 CO<sub>2</sub> solubility than VolatileCalc-Basalt. This is noteworthy, because many publi-  
1344 cations have calculated saturation pressures for andesitic and dacitic melts using  
1345 VolatileCalc-Rhyolite (e.g., Blundy et al., 2006; Atlas et al., 2006; Cassidy et al.,  
1346 2015; Koleszar et al., 2012). In the original publication, Newman and Lowenstern  
1347 (2002) state: “because many andesites contain rhyolitic interstitial melt, Volatile-  
1348 Calc may also be applicable to these intermediate compositions.” However, this  
1349 should not be taken to suggest that VolatileCalc-Rhyolite is safely extrapolated to  
1350 andesitic-dacitic melts. Instead, this statement is referring to the fact that many  
1351 lavas with andesitic bulk compositions have rhyolitic groundmass/melt inclusion  
1352 compositions (e.g., Tamura & Tatsumi, 2002; Reubi & Blundy, 2009), where  
1353 VolatileCalc-Rhyolite may be applicable. Interestingly, VolatileCalc-Rhyolite iso-  
1354 bars calculated for the reference temperature of this model (850°C) rather than  
1355 the experimental temperature are a good match to experimental data at < 5 kbar

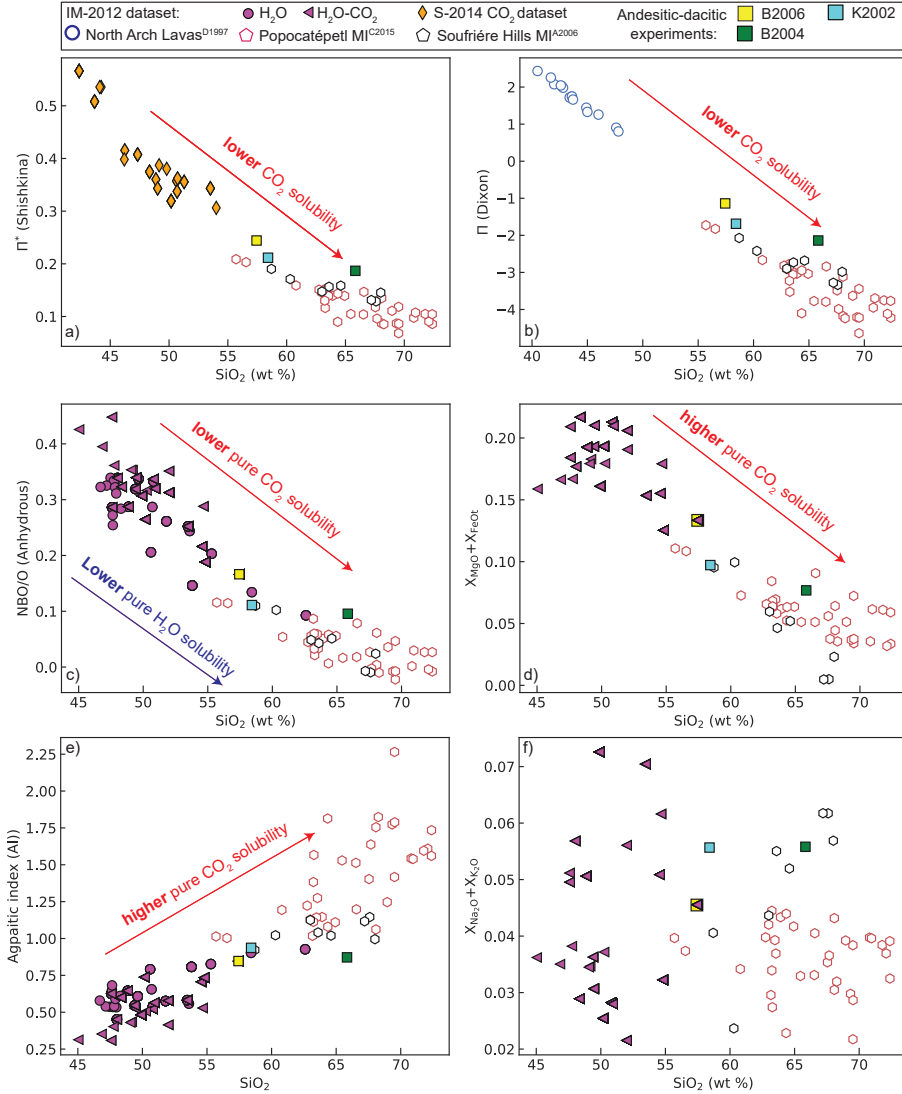
1356 (see Supporting Fig. S3). Thus, the main failure of this model in intermediate com-  
1357 positions appears to result from the fact that this model is extremely sensitive to  
1358 temperature, and these melts have much higher temperatures than the experiments  
1359 used to calibrate this model. This supports our suggestion in section 4.2.1 that it  
1360 may be better run models at their reference temperature, rather than extrapolate  
1361 beyond the calibration range.

1362 It is also worth noting that all these experimental products have negative val-  
1363 ues of the Dixon  $\Pi$  parameter (equation 2). As highlighted by G. Moore (2008),  
1364 the full  $\Pi$  expression of Dixon (1997) cannot be used to calculate  $\text{CO}_2$  solubility in  
1365 calc-alkaline lavas, because  $\Pi$  is negative (yielding a negative solubility of  $\text{CO}_2$ ; Fig.  
1366 17).

1367 Differences between experimental data and isobars from IM-2012-H and -A are  
1368 more complicated, because the discrepancies between models and experiments are  
1369 very dependent on the pressure. For example, at 1 kbar, IM-2012-H and -A under-  
1370 predict  $\text{H}_2\text{O}$  and  $\text{CO}_2$  solubility relative to the experiments of Behrens et al. (2004),  
1371 while at 5 kbar, they overpredict  $\text{CO}_2$  solubility, but underpredict  $\text{H}_2\text{O}$  solubility  
1372 (Fig. 16c vs.e). These discrepancies likely reflect this model being extrapolated  
1373 towards the limits of its calibration dataset in terms of both pressure (most exper-  
1374 iments were conducted at  $<5$  kbar) and melt composition (Fig. 17, see the next  
1375 section for more discussion).



**Figure 16.** Assessing model fits to experimental data for andesitic and dacitic compositions a-b) Experiments from Botcharnikov et al. (2006). Isobars were calculated for 1200°C and Fe<sup>3+</sup>/Fe<sub>T</sub>=0.2 (~QFM+1.5). P-2006 isobars are also shown for Fe<sup>3+</sup>/Fe<sub>T</sub>=0.7 (the upper estimate of Fe<sup>3+</sup>/Fe<sub>T</sub> in experimental products; dash-dotted line), and Fe<sup>3+</sup>/Fe<sub>T</sub>=0 (dotted line) c-e) Experiments from Behrens et al. (2004). Isobars are shown for Fe<sup>3+</sup>/Fe<sub>T</sub>=0.59 (*f*O<sub>2</sub> of QFM+4.7). For P-2006 and MagmaSat, isobars are also shown for Fe<sup>3+</sup>/Fe<sub>T</sub>=0. f) Experiments from King and Holloway (2002). Isobars are shown for Fe<sup>3+</sup>/Fe<sub>T</sub>=0.2. For P-2006 and MagmaSat, isobars are also shown for Fe<sup>3+</sup>/Fe<sub>T</sub>=0. VolatileCalc-Rhyolite isobars are not shown, as the spreadsheet doesn't calculate isobars above 5 kbar. Error bars on all plots shows the 2σ uncertainties from measurements of volatile contents in experimental products. Fe<sup>3+</sup>/Fe<sub>T</sub> ratios were calculated from author-stated buffers using MELTS for Excel (Gualda & Ghiorso, 2015).



**Figure 17.** Comparing the calibration datasets of S–2014 and IM–2012 to the andesitic experiments of Botcharnikov et al. (2006), (B2006), Behrens et al. (2004), (B2004) and King and Holloway (2002), (K2002) shown in Fig. 16, and the melt inclusions from Popocatepetl (Atlas et al., 2006) and Soufrière Hills (Cassidy et al., 2015) shown in Fig. 18. Anhydrous molar fractions are used to calculate compositional parameters in parts c-f, because when accounting for discrepancies between isobars (e.g., on Fig. 18, the H<sub>2</sub>O content and therefore hydrous cation fractions vary as a function of the pressure).



1376 The fact that no model passes through all available experiments demonstrates  
 1377 that further investigation of solubility in andesitic to dacitic melts is warranted.  
 1378 Using representative experimental  $\text{Fe}^{3+}/\text{Fe}_T$  ratios, MagmaSat is the most accurate  
 1379 model, predicting dissolved volatile contents within  $\sim 20\%$  of experimental products  
 1380 (considering reported error bars on dissolved volatile contents).

1381 The extreme sensitivity to the  $\text{Fe}^{3+}/\text{Fe}_T$  ratio makes it very difficult to assess  
 1382 the accuracy of the P-2006 model (particularly given the relatively large uncertain-  
 1383 ties in the oxygen fugacity of experimental run products; King & Holloway, 2002;  
 1384 Botcharnikov et al., 2006). In all of the examples shown, P-2006 isobars calculated  
 1385 for  $\text{Fe}^{3+}/\text{Fe}_T=0$  are a better fit to the experimental data than isobars calculated us-  
 1386 ing estimates of the  $\text{Fe}^{3+}/\text{Fe}_T$  ratio of experiments. This suggests that, in relatively  
 1387 oxidising intermediate melts, P-2006 is overestimating the effect of  $\text{Fe}^{3+}$  species on  
 1388 volatile solubility. It is noteworthy that P-2006 is a particularly poor match to the  
 1389 high pressure andesitic experiments of King and Holloway (2002), despite the fact  
 1390 that these are the only intermediate experiments in the calibration dataset of this  
 1391 model.

1392 As all the experiments shown were used to calibrate MagmaSat, similar anal-  
 1393 ysis applied to new experimental data on andesitic compositions when it becomes  
 1394 available will provide further constraints on the accuracy of this model. Isobars for  
 1395 recent solubility experiments on a basaltic-andesite at 4–12 kbar by Mangan et al.  
 1396 (2021) are shown in Supporting Fig. S4 and S5. MagmaSat is a much better fit  
 1397 to this new data than P-2006, with experiments performed 400–815 MPa plotting  
 1398 within error of calculated MagmaSat isobars. In contrast, P-2006 overpredicts  $\text{CO}_2$   
 1399 solubility (regardless of  $\text{Fe}^{3+}/\text{Fe}_T$ ) at  $<600$  MPa.

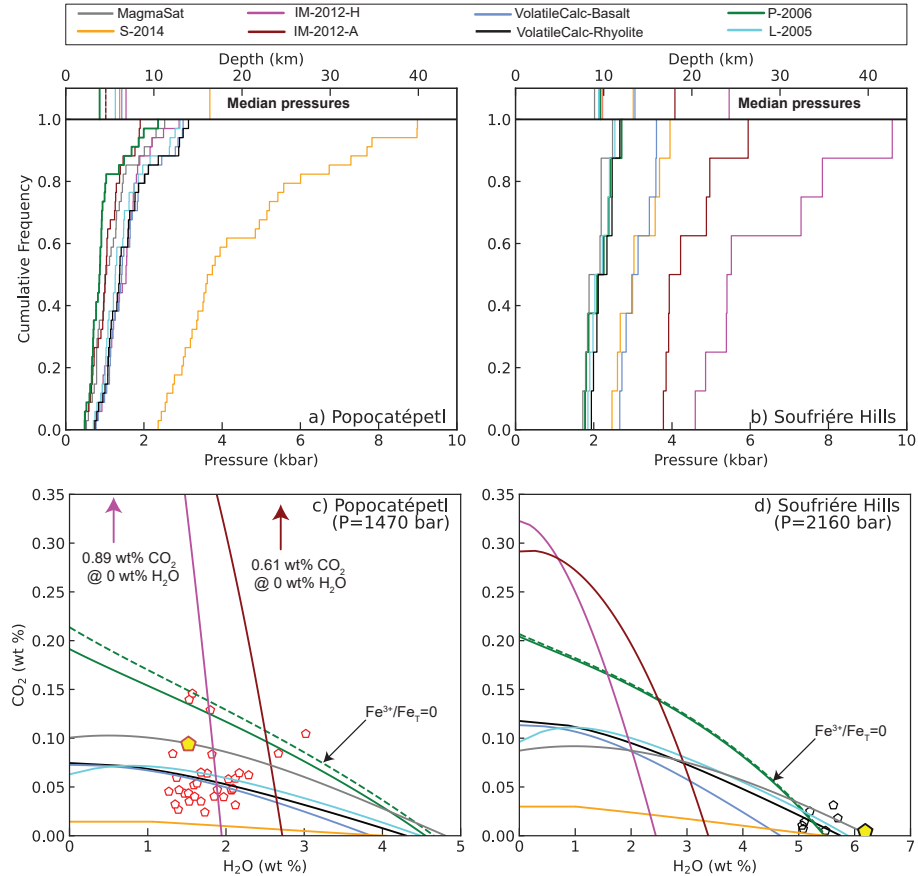
## 1400 **6.2 Case Study: Intermediate melt inclusions**

1401 To assess the impact of these model differences on the depths of magma stor-  
 1402 age reservoirs inferred from melt inclusions in volcanic arcs, we calculate saturation  
 1403 pressures using a variety of models for two suites of melt inclusions with andesitic-  
 1404 dacitic liquid compositions: 1) 34 melt inclusions from Volcán Popocatepetl, Mexico  
 1405 with 55.7–73.4 wt%  $\text{SiO}_2$  (Figs. 2, 18a; Atlas et al., 2006), and 2) 8 melt inclusions  
 1406 from Soufrière Hills Volcano, Montserrat with 58.7–68.0 wt%  $\text{SiO}_2$  (Figs. 2, 18b;

1407 Cassidy et al., 2015). Both studies calculated saturation pressures (and therefore  
1408 magma storage depths) using VolatileCalc-Rhyolite.

1409 Cumulative frequency distributions for the Popocatepetl melt inclusions (Fig.  
1410 18a) shows that P-2006 predicts the lowest saturation pressures, and S-2014 the  
1411 highest, with MagmaSat, IM-2012-A, VolatileCalc-Basalt, VolatileCalc-Rhyolite,  
1412 and IM-2012-H lying in-between these model extremes. Based on our analysis in  
1413 the previous section suggesting that MagmaSat is the best calibrated model for  
1414 intermediate melt compositions, we ratio saturation pressures from each model to  
1415 those determined using MagmaSat (allowing model differences to be quantified).  
1416 Additionally, because MagmaSat is a thermodynamic model that has been shown to  
1417 work well for basaltic and rhyolitic compositions, it is effectively being interpolated  
1418 to andesitic-dacitic compositions which are not represented in its calibration dataset  
1419 (i.e., these melt inclusion compositions), rather than extrapolated (as for empiri-  
1420 cal models such as S-2014, VolatileCalc-Basalt, and IM-2012 which are primarily  
1421 calibrated on more mafic melt compositions; Fig. 2).

1422 The median saturation pressure for Popocatepetl melt inclusions calculated  
1423 for VolatileCalc-Rhyolite is  $1.26\times$  higher than for MagmaSat. The median for Mag-  
1424 maSat is  $\sim 1.2\times$  higher than for P-2006. The deviation relative to MagmaSat in-  
1425 creases with  $\text{SiO}_2$  content for VolatileCalc-Rhyolite ( $\sim 1.1$  to  $1.5\times$ ), VolatileCalc-  
1426 Basalt ( $\sim 1$  to  $1.35\times$ ), and L-2005 ( $\sim 1$  to  $1.4\times$ ). The deviation between S-2014  
1427 and MagmaSat is very large, increasing from a factor of 2 at  $\sim 55$  wt%  $\text{SiO}_2$  to  
1428  $\sim 5.5\times$  at 72.5 wt%  $\text{SiO}_2$ . The deviation between IM-2012-H and -A and P-2006  
1429 vs. MagmaSat shows no clear correlation with  $\text{SiO}_2$  content (Supporting Fig. S6).  
1430 The presence of discrepancies which correlate with melt composition is particularly  
1431 concerning, because it means that choice of an inappropriate solubility model may  
1432 introduce systematic error into a dataset as a function of melt inclusion composition.



**Figure 18.** a-b) Cumulative distribution functions of saturation pressures from different models for melt inclusions from a) Popocatépetl (Atlas et al., 2006) and b) Soufrière Hills (Cassidy et al., 2015).  $\text{Fe}^{3+}/\text{Fe}_T$  was set to 0.15. c) Isobars from different models calculated at the saturation pressure from MagmaSat (1470 bar) for the Popocatépetl melt inclusion shown with a yellow pentagon (at 1050°C following Atlas et al., 2006). The scale is trimmed to emphasize the differences between models at lower CO<sub>2</sub> contents, with the interception of the IM-2012 isobars with the y axis labelled on the plot. d) Isobars from different models calculated at the saturation pressure from MagmaSat (2160 bar) for the Soufrière Hills melt inclusion shown with a yellow pentagon (at 1000°C following Cassidy et al., 2015). Isobars are also shown for P-2006 with  $\text{Fe}^{3+}/\text{Fe}_T=0$  (dotted line)

1433 MagmaSat predicts the lowest saturation pressures for the Soufrière Hills melt  
 1434 inclusions, with L-2005, P-2006, and VolatileCalc-Rhyolite predicting reasonably  
 1435 similar pressures. VolatileCalc-Basalt, S-2014 and IM-2012-A and -H are offset

1436 to higher pressures. Similar to the results for Popocatépetl, the ratio of saturation  
 1437 pressures for Soufrière Hills melt inclusions relative to MagmaSat for L-2005 ( $\sim 1$   
 1438 to  $1.1\times$ ), VolatileCalc-Basalt ( $\sim 1.25$  to  $1.35\times$ ) and VolatileCalc-Rhyolite ( $\sim 1.05$   
 1439 to  $1.17\times$ ) increase with increasing  $\text{SiO}_2$  (Supporting Fig. S7). Unlike Popocatépetl  
 1440 melt inclusions, the discrepancy between IM-2012 and MagmaSat increases dra-  
 1441 matically with increasing  $\text{SiO}_2$  (from  $\sim 1.5$  to  $2.5\times$  for anhydrous, and 2 to  $4.5\times$  for  
 1442 hydrous), while that for S-2014 shows a weak negative correlation with  $\text{SiO}_2$  (from  
 1443  $\sim 1.7$ – $1.35\times$ ).

1444 The large discrepancies shown by IM-2012 (for Soufrière Hills) and S-2014 (for  
 1445 Popocatépetl) are best understood by examining an isobar calculated for the inclu-  
 1446 sion showing the largest deviation relative to MagmaSat in each dataset ( $\text{SiO}_2=72.4$   
 1447 wt% for Popocatépetl, and 67.6 wt% for Soufrière Hills; yellow pentagon on Fig.  
 1448 18c-d) at the saturation pressure obtained from MagmaSat. For both sets of inclu-  
 1449 sions, S-2014 isobars intercept the x axis at similar  $\text{H}_2\text{O}$  contents to other models,  
 1450 but intercept the y axis at drastically lower  $\text{CO}_2$  contents. This underestimation of  
 1451  $\text{CO}_2$  solubility likely results from the fact that the  $\Pi^*$  values of these inclusions lie  
 1452 well below that of the calibration dataset (as discussed for andesitic experiments;  
 1453 Fig. 17b).

1454 In contrast, IM-2012 predicts very high pure  $\text{CO}_2$  solubility, and low pure  $\text{H}_2\text{O}$   
 1455 solubility relative to the other models. Both IM-2012 models express the composi-  
 1456 tional dependence of  $\text{H}_2\text{O}$  solubility in terms of the parameter NBO/O, with melts  
 1457 with higher NBO/O values having higher  $\text{H}_2\text{O}$  solubility (equation 15). Both sets  
 1458 of melt inclusions possess much lower NBO/O values than the calibration dataset  
 1459 of IM-2012 (Fig. 17c-d). This is problematic, because the empirical relationship  
 1460 between NBO/O and pure  $\text{H}_2\text{O}$  solubility incorporated by IM-2012 has not been  
 1461 validated for these melt compositions. For example, Shishkina et al. (2014) show  
 1462 that IM-2012 drastically overestimates  $\text{H}_2\text{O}$  solubility in their basanite and nepheli-  
 1463 nite melt compositions. They point out that while IM-2012 conclude that there  
 1464 is only a small effect of melt composition on  $\text{H}_2\text{O}$  solubility, the IM-2012 model  
 1465 ends up showing a strong sensitivity to melt composition when extrapolated to the  
 1466 high NBO/O ratio in their basanite and nephelinite melts (and we invoke a similar  
 1467 explanation for the lower NBO/O ratios in melt inclusions discussed here).

1468 The IM-2012 expression for CO<sub>2</sub> solubility is more complicated, contain-  
 1469 ing terms for the Agpaitic index (AI), NBO/O, X<sub>Na<sub>2</sub>O+K<sub>2</sub>O</sub>, X<sub>FeO+MgO</sub> and X<sub>H<sub>2</sub>O</sub>  
 1470 (equation 11). These two sets of melt inclusions have higher AI, lower NBO/O ra-  
 1471 tios, similar values of X<sub>Na<sub>2</sub>O+K<sub>2</sub>O</sub>, and lower X<sub>FeO+MgO</sub> values than the calibration  
 1472 dataset (Fig. 17d-f). While the effect of NBO/O is more convoluted because it also  
 1473 affects the solubility of H<sub>2</sub>O (which feeds back into the expression for CO<sub>2</sub>), it is  
 1474 readily apparent that the positive coefficient attached to the AI term combined with  
 1475 the negative coefficient attached to the MgO+FeO term causes this model to predict  
 1476 higher CO<sub>2</sub> solubilities than the calibration dataset for the andesitic-dacitic melt  
 1477 inclusions considered here.

1478 The discrepancy between isobars for S-2014 and IM-2012 relative to Mag-  
 1479 maSat are relatively similar for the Popocatepetl and Soufrière Hills melt compo-  
 1480 sitions, while discrepancies for saturation pressures differ markedly (Fig. 18a-b vs.  
 1481 c-d). This is because the volatile contents of Popocatepetl melt inclusions are sig-  
 1482 nificantly more CO<sub>2</sub>-rich (~0.02–0.15 wt% and higher), and H<sub>2</sub>O-poor (~1–3 wt%)  
 1483 than Soufrière Hills melt inclusions (<0.04 wt% CO<sub>2</sub> and 5–6 wt% H<sub>2</sub>O). For this  
 1484 reason, calculated saturation pressures for Popocatepetl melt inclusions are sensitive  
 1485 to the treatment of both CO<sub>2</sub> and H<sub>2</sub>O in solubility models (Fig. 18b), while those  
 1486 for Soufrière Hills melt inclusions are mostly sensitive to pure H<sub>2</sub>O solubility. Thus,  
 1487 S-2014 overestimates saturation pressures for Popocatepetl melt inclusions because  
 1488 this model drastically underestimates the solubility of pure CO<sub>2</sub>. In contrast, S-2014  
 1489 only slightly underestimates H<sub>2</sub>O solubility relative to MagmaSat, so only slightly  
 1490 overpredicts saturation pressures for H<sub>2</sub>O-rich Soufrière Hills melt inclusions. The  
 1491 discrepancy between IM-2012 and MagmaSat is much smaller for Popocatepetl rel-  
 1492 ative to Soufrière Hills because, fortuitously, the IM-2012-H isobar intercepts the  
 1493 MagmaSat isobar at H<sub>2</sub>O contents similar to these Popocatepetl melt inclusions.  
 1494 IM-2012-H and -A drastically underestimates the solubility of pure H<sub>2</sub>O, so overes-  
 1495 timates saturation pressures for the H<sub>2</sub>O-rich Soufrière Hills melt inclusions relative  
 1496 to other models.

1497 It is worth noting that Iacono-Marziano never intended their model to be ap-  
 1498 plied to andesites, and when discussing the limitations of their model, they explicitly  
 1499 warn that their empirical expressions poorly incorporates the effect of MgO and FeO  
 1500 on CO<sub>2</sub> solubility because of the restricted range of these oxides in the calibration

1501 dataset. We have included this discussion as an extreme example of the danger of  
1502 extrapolating empirical models beyond their calibration range. However, the sen-  
1503 sitivity of this model to the FeO and MgO content of the melt also presents issues  
1504 when applied to high MgO basaltic liquids. For example, Wieser et al. (2021) show  
1505 that IM-2012 predicts higher saturation pressures for highly primitive (high MgO)  
1506 melt inclusions from Kilauea Volcano relative to S-2014 and MagmaSat. This likely  
1507 reflects the higher values of  $X_{\text{FeO}+\text{MgO}}$  in these melt inclusions relative to the cali-  
1508 bration dataset, which causes IM-2012 to predict lower  $\text{CO}_2$  solubility (the opposite  
1509 directionality to that seen for the intermediate melt inclusions discussed here).

1510 The  $\text{H}_2\text{O}$ -rich nature of Soufrière Hills melt inclusions means that VolatileCalc-  
1511 Rhyolite predicts much more similar saturation pressures to MagmaSat ( $1.07\text{--}1.15\times$ ;  
1512 Supporting Fig. S6) than for Popocatepetl ( $1.1\text{--}1.5\times$  higher; Supporting Fig. S5),  
1513 because the main failure of VolatileCalc-Rhyolite for intermediate compositions at  
1514 moderate to high temperatures ( $>850$  °C) is its prediction of pure  $\text{CO}_2$  solubility  
1515 (Figs. 16, 18c-d). The discrepancy for both VolatileCalc models and L-2005 relative  
1516 to MagmaSat is significantly smaller than for S-2014 and IM-2012. This is because  
1517 the solubility differences during evolution for basaltic to rhyolitic compositions are  
1518 relatively small (30–40%) compared to the error associated with the extrapolating  
1519 an empirical model far beyond its compositional range. Overall, this case study  
1520 shows the importance of checking that the calibration dataset of a model contains  
1521 melts similar to those in the sample set of interest, particularly if the effect of melt  
1522 composition is parameterized empirically.

## 1523 **7 Best practices for data visualization and curation**

### 1524 **7.1 Isobar Diagrams: limitations and alternatives**

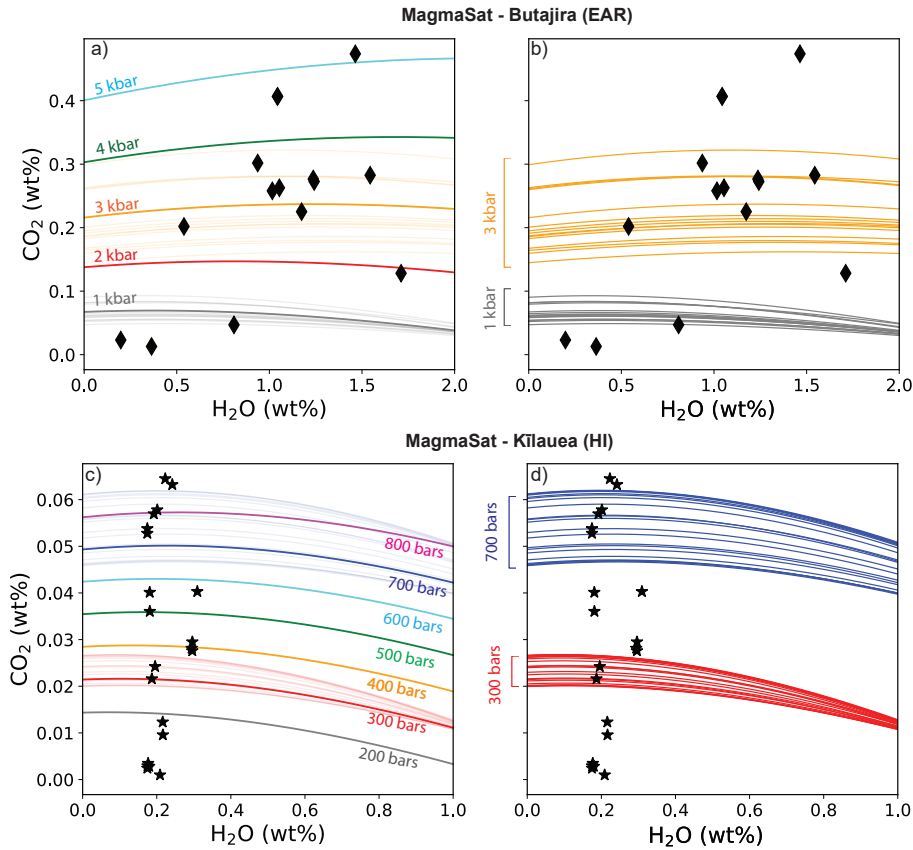
1525 As demonstrated by the preceding section, isobar diagrams are a useful tool to  
1526 visualize variations in volatile solubility for a specific melt composition. However,  
1527 many suites of melt inclusions have considerable major element variability, which  
1528 translates into differing solubilities of  $\text{H}_2\text{O}\text{--}\text{CO}_2$  at a given pressure, and different  
1529 isobar shapes at a given pressure (Wieser et al., 2021; Roggensack, 2001; Iacovino et  
1530 al., 2021). We use two suites of melt inclusions to demonstrate this point.

1531 First, we consider a suite of basaltic melt inclusions from Butajira volcanoes in  
1532 the Main Ethiopian Rift from Iddon and Edmonds (2020). We calculate isobars for a  
1533 representative inclusion composition (BJ08\_7; Fig. 19a), and then we compare these  
1534 to the isobars calculated for each individual melt inclusion composition at 1 and 3  
1535 kbar (Fig. 19a-b). 3 kbar isobars calculated from the composition of each individual  
1536 melt inclusion from Butajira (Fig. 19b) cover the entire region of H<sub>2</sub>O-CO<sub>2</sub> space  
1537 that is bracketed by the 2 and 4 kbar isobars calculated in MagmaSat for the repre-  
1538 sentative melt inclusion composition (Fig. 19b vs. a). In fact, saturation pressures  
1539 calculated using any given melt inclusion composition vs. the major element compo-  
1540 sition of the melt inclusion in question can affect the saturation pressure by almost a  
1541 factor of 2. By extension, visual inspection of melt inclusion volatile concentrations  
1542 plotted on isobar diagrams may be associated with an error of this magnitude).

1543 Second, we calculate isobars at 300 and 700 bars for a representative subset  
1544 of melt inclusions from the 2018 eruption of Kīlauea with PEC-corrected MgO con-  
1545 tents between 6.4 to 13.5 wt% (Wieser et al., 2021). Although less extreme than the  
1546 Butajira example, the spread of isobars calculated for a single sample at 700 bars  
1547 for multiple melt inclusions is wider than the distance between isobars calculated  
1548 100 bar apart for a specific melt inclusion. Similar isobar spreads are seen for both  
1549 Kīlauea and Butajira using IM-2012 (Supporting Fig. S8 and S9).

1550 VESICAL solves the problem of potentially misleading isobar diagrams by  
1551 facilitating rapid calculations of saturation pressures for large suites of melt in-  
1552 clusions. For example, calculating saturation pressures for the 33 melt inclusions  
1553 in the dataset of Iddon and Edmonds (2020) using VolatileCalc-Basalt, S-2014,  
1554 and IM-2012 takes only 4.2 seconds if VESICAL is run on the ENKI server. Mag-  
1555 maSat is slightly slower, taking 31.5 seconds (still <1 s per sample). MagmaSat  
1556 calculations may run faster if ThermoEngine is installed locally (calculations take  
1557 26.3 s using a Dell Inspiron laptop with 16 GB RAM and an Intel-i7 processor, see  
1558 <https://gitlab.com/ENKI-portal/ThermoEngine> for installation help).

1559 Once users have calculated saturation pressures for each inclusion, a number of  
1560 different x-y plots will provide more information than isobar diagrams. For example,  
1561 G. Moore (2008) suggests that users could plot two graphs, one showing melt inclu-  
1562 sion CO<sub>2</sub> content against saturation pressure and one showing melt inclusion H<sub>2</sub>O



**Figure 19.** a-b) Isobars for olivine-hosted melt inclusions from Iddon and Edmonds (2020) calculated at 1170°C. a) MagmaSat isobars for melt inclusion BJ08.7 from Butajira with melt inclusion volatile data overlain. b) 1 and 3 kbar isobars calculated for the major element composition of each individual melt inclusion from Butajira (underlain as light lines in a). c-d) As for a-b, but using a representative subset of melt inclusions from Kilauea (Wieser et al., 2021). To save computation time as this study analyses >100 melt inclusions, PEC-corrected compositions were sorted by MgO content, and every 5th inclusion composition was used to calculate an isobar. Isobars are calculated for a temperature calculated from the melt inclusion MgO content using the thermometer of Helz and Thornber (1987).

1563 contents against saturation pressure. This allows clustering of saturation pressures  
 1564 to be observed, and determination of the importance of each volatile species when  
 1565 calculating saturation pressure. Alternatively, saturation pressures could be plotted  
 1566 as histograms, cumulative density functions, or box/violin plots to assess clustering.  
 1567 Additionally, the ability to easily filter calculated saturation pressures by additional



1568 variables in python (e.g., host mineral content, amount of post-entrapment crystal-  
1569 lization, sample location, stratigraphic height) will help to elucidate the record of  
1570 the magmatic plumbing system preserved in different crystal populations. For ex-  
1571 ample, Wieser et al. (2021) show that melt inclusions hosted in low forsterite olivine  
1572 crystals from Kīlauea Volcano crystallized at  $\sim 1\text{--}2$  km depth, while melt inclusions  
1573 hosted within higher forsterite olivines crystallized at  $\sim 3\text{--}5$  km depth.

## 1574 **7.2 Assessing Errors**

1575 The ability to calculate saturation pressures using a number of different models  
1576 in VESICAL is advantageous, because it can provide assessment of the systematic  
1577 errors associated with model choice. If different solubility models produce satura-  
1578 tion pressures which are statistically distinguishable using tests such as ANOVA or  
1579 the Kolmogorov-Smirnov test, or differences between models exceed 10–20% (the  
1580 approximate quoted error on most models), users need to evaluate their melt compo-  
1581 sitions, pressures and temperatures in the context of the calibration dataset of each  
1582 solubility model. Several Jupyter notebooks aiding these comparisons are provided  
1583 alongside VESICAL part I (Iacovino et al., 2021). Additionally, comparisons between  
1584 any available experimental data for relevant melt compositions and different solu-  
1585 bility models using a workflow similar to that used here for andesites (e.g., isobar  
1586 diagrams as in Fig. 16, plots of melt composition vs. calibration datasets) will help  
1587 users select a suitable model. As well as examining melt compositions, users should  
1588 also evaluate whether they are extrapolating temperature-sensitive models beyond  
1589 the calibration range (as discussed here for VolatileCalc-Rhyolite).

1590 In general, if a natural silicate melt composition is poorly represented by ex-  
1591 perimental data, MagmaSat is probably the best model to use, as its thermodynamic  
1592 nature is more suitable to extrapolation to melt compositions not represented in the  
1593 calibration dataset than empirical models such as IM-2012, S-2014, or VolatileCalc-  
1594 Basalt. However, the comprehensive nature of MagmaSat means that the fit to  
1595 experimental data from any specific region of major element space is compromised  
1596 by the fact the model is optimizing the overall fit to many different major element  
1597 compositions. Thus, where present, models developed for specific volcanic centres,  
1598 or highly weighted towards specific melt compositions (e.g, A-2019 for the 6 cen-  
1599 ters they investigate, or IM-2012 for alkaline compositions) may return a better fit.

1600 While these composition-specific models may be well calibrated in terms of melt  
1601 composition, users must also check that they are applying the model within the  
1602 recommended pressure and temperature range.

1603 VESICAL also opens up the capability to handle errors arising from volatile  
1604 solubility modelling using the advanced functionality of Python3 packages such as  
1605 Numpy, SciPy and PyMC to perform Bayesian statistical techniques (e.g., Markov  
1606 Chain Monte Carlo methods). This means the uncertainty in all input parameters  
1607 (e.g., temperature, analytical uncertainties in volatile and major element contents,  
1608 uncertainty arising from post-entrapment crystallization corrections) can be prop-  
1609 agated into a resulting error distribution for each melt inclusion. These techniques  
1610 are increasingly being utilized by igneous petrologists, with recent applications in-  
1611 cluding calculating error distributions for diffusion timescales (Mutch et al., 2019),  
1612 the contribution of melts from distinct mantle sources (Gleeson et al., 2020), and  
1613 propagating uncertainties in vapour bubble growth models (Rasmussen et al., 2020).

### 1614 **7.3 Data Curation**

1615 Now that VESICAL makes it possible to calculate saturation pressures for large  
1616 melt inclusion datasets in short amounts of time, it is vital that data is published  
1617 in a way that allows such calculations to be performed retrospectively (e.g., recal-  
1618 culating literature saturation pressures from a given volcanic center/region to use a  
1619 single solubility model). The concentration of major elements and volatile elements  
1620 should be provided within a single spreadsheet or database, ideally alongside the  
1621 composition of the mineral host. Additionally, if melt inclusions are corrected for  
1622 post-entrapment crystallization (PEC), both raw and corrected major and volatile  
1623 element concentrations should be published. Particularly for more chemically-  
1624 complex host minerals like pyroxene and plagioclase, it is highly likely that com-  
1625 munity standards on the best PEC procedure will change with time. Finally, given  
1626 that numerous recent studies have shown that bubble CO<sub>2</sub> contents can change  
1627 calculated saturation pressures by up to an order of magnitude (section 5), even if  
1628 authors were unable to perform Raman analyses, it should be noted whether each  
1629 melt inclusion contained a vapour bubble (and ideally an estimate of the volume  
1630 proportion of the vapour bubble). This information may allow the CO<sub>2</sub> contents in  
1631 vapour bubbles to be reconstructed theoretically by future studies, given the recent

1632 proliferation of vapour bubble growth models (e.g. Aster et al., 2016; Rasmussen et  
1633 al., 2020; MacLennan, 2017).

1634 Proper data curation is particularly vital in the world of volatile solubility  
1635 modelling because it is very likely that many more experiments will be published  
1636 over the next decade, allowing the creation of new solubility models that are cali-  
1637 brated over an even wider region of P-T and compositional space. Many publica-  
1638 tions could not be used for comparisons in this study, because there was simply no  
1639 way to combine volatile element concentrations and major element concentrations  
1640 (which were often reported in different tables with non-unique or non-matching  
1641 sample names), or not reported at all. It would be a great shame if published melt  
1642 inclusion datasets could not be input into these new models to recalculate saturation  
1643 pressures, and gain a greater understanding of magma storage in the Earth's crust.

## 1644 **8 Future work**

1645 The comparisons drawn in this review highlight several research areas where  
1646 further experimental work is required to be able to distinguish which model be-  
1647 haviours are accurate. First, significantly more experiments are needed on andesitic-  
1648 dacitic melt compositions. Figure 16 shows that it is currently impossible to differ-  
1649 entiate a potential failure in any given solubility model from anomalies in any given  
1650 set of experiments (e.g., the differential effect of addition of H<sub>2</sub>O on CO<sub>2</sub> solubility  
1651 in different experiments; Fig. 16b. vs f).

1652 One of the challenges when assessing CO<sub>2</sub> solubility in andesitic-dacitic melts  
1653 is the fact that CO<sub>2</sub> is present as both carbonate and molecular CO<sub>2</sub>. Carbon  
1654 species do appear separately in FTIR spectra, but the accuracy of FTIR-derived  
1655 volatile concentrations can be affected by peak overlap (e.g., Brooker et al., 2001b),  
1656 as well as uncertainty in peak baseline and absorption coefficients (Botcharnikov et  
1657 al., 2006; Mangan et al., 2021). SIMS cannot distinguish different carbon species  
1658 (only yields total carbon), but may help to resolve issues with FTIR as a result of  
1659 increased understanding of the optimal analysis conditions for volatiles in silicate  
1660 glasses of the last few decades. However, in addition to its substantial cost relative  
1661 to FTIR, accurate SIMS measurements are reliant on having a suite of standards  
1662 with similar major element compositions and a range of volatile contents (and these

1663 standards are often characterized by FTIR, so are subject to the caveats mentioned  
1664 above).

1665         Second, the effect of redox on volatile solubility across the range encountered  
1666 in terrestrial magmas is still poorly constrained (section 5). This discrepancy largely  
1667 reflects the fact that the redox conditions at which many experiments in the litera-  
1668 ture were conducted are uncertain and/or highly variable (e.g., Botcharnikov et al.,  
1669 2006). Because of this uncertainty, many calibration datasets are built without being  
1670 able to constrain the quantities of  $\text{Fe}_2\text{O}_3$  and  $\text{FeO}$  for each experimental run. While  
1671 our investigation of intermediate melts indicates that the strong sensitivity of  $\text{CO}_2$   
1672 solubility to melt redox shown by P-2006 is likely anomalous because of the presence  
1673 of experiments with extremely high calculated  $\text{Fe}^{3+}/\text{Fe}_T$  ratios in the calibration  
1674 dataset, further experiments where  $\text{Fe}_2\text{O}_3$  and  $\text{FeO}$  proportions are accurately mea-  
1675 sured are needed to be certain that this behaviour is not real. It is also noteworthy  
1676 that almost all the andesitic experiments were performed at higher oxygen fugacities  
1677 than lavas erupted at volcanic arcs. When the calibration dataset for a given set of  
1678 compositions is so small, this makes it difficult to deconvolve changes in volatile sol-  
1679 ubility with melt composition compared to redox. Recent advances in measurements  
1680 of  $\text{Fe}^{3+}/\text{Fe}_T$  using Fe K-edge micro-X-ray absorption near-edge structure (XANES)  
1681 spectroscopy in hydrous glasses (Cottrell et al., 2018) could provide an avenue to  
1682 better constrain this parameter in future (and past) experimental products.

1683         It is also worth noting that all the models discussed here only consider the  
1684 effect of redox through terms for  $\text{Fe}^{2+}$  and  $\text{Fe}^{3+}$  in the melt, constraining their ap-  
1685 plicability to melts more oxidising than the IW buffer. In more reducing conditions,  
1686 the co-existing  $\text{CO}_2$ -rich phase may be graphite or diamond rather than a  $\text{CO}_2$ -rich  
1687 vapour phase (Eguchi & Dasgupta, 2018), and the dissolved volatile species may  
1688 be  $\text{CO}$ ,  $\text{CH}_4$  and  $\text{H}_2$  (Mysen et al., 2009). This means that extreme caution is re-  
1689 quired when applying these solubility models to highly reducing conditions such as  
1690 those found on other planetary bodies (e.g., the Moon, Mars and Mercury; Li et al.,  
1691 2017).

1692         Third, there is still significant uncertainty regarding the exact nature of  $\text{H}_2\text{O}$   
1693 and  $\text{CO}_2$  mixing at higher pressure. This reflects the difficulty in measuring mixed  
1694  $\text{H}_2\text{O}$ - $\text{CO}_2$  fluids that were in equilibrium with the melt during the experiment. If

1695 measured at all, methods in the literature span from puncture and weight loss of  
1696 frozen capsules (i.e., when frozen the CO<sub>2</sub> is released, but not the H<sub>2</sub>O; Shishkina  
1697 et al., 2010) to more sophisticated and accurate vacuum line manometry (G. Moore  
1698 et al., 2008; Iacovino et al., 2013; Allison et al., 2019). New infinite path laser spec-  
1699 trometry technology may offer potential improvement of this critical measurement,  
1700 but challenges associated with small sample sizes remain. More work determining  
1701 the pure CO<sub>2</sub> solubility as a function of pressure and temperature would also be of  
1702 great benefit in constraining the behavior of H<sub>2</sub>O-poor fluids.

1703 Fourth, we show that the sensitivity of dissolved volatile contents to temper-  
1704 ature is highly model-specific. Given the difficulties with constraining temperature  
1705 sensitivity experimentally (Iacono-Marziano et al., 2012), we suggest that it may be  
1706 best to parametrize future empirical models at a single temperature (e.g., A-2019,  
1707 S-2014), or incorporate the temperature sensitivity predicted by thermodynami-  
1708 cal models, rather than introduce a spurious temperature dependency which is not  
1709 founded in experimental data, nor consistent with the relative entropy of melt and  
1710 fluid terms.

1711 Finally, it is worth noting that all of the solubility models discussed only con-  
1712 sider H<sub>2</sub>O-CO<sub>2</sub> in the vapour phase, while in reality, natural fluids in volcanic sys-  
1713 tems may contain relatively large proportions of F, Cl, and S, as well as a separate  
1714 brine phase (Botcharnikov et al., 2007). Additionally, at higher pressures and tem-  
1715 peratures, significant quantities of major element species will dissolve into a H<sub>2</sub>O-  
1716 rich fluid (e.g., Si, Na, K), with silicate melt and hydrous fluids becoming completely  
1717 miscible above a critical temperature (Bureau & Keppler, 1999). This causes a pure  
1718 H<sub>2</sub>O model to underestimate the true solubility of H<sub>2</sub>O at these conditions. Com-  
1719 bined with the fact that it is near-impossible to quench silicate melts with >9±1  
1720 wt% H<sub>2</sub>O to a glass phase which can be analysed by SIMS or FTIR (Gavrilenko  
1721 et al., 2019; Mitchell et al., 2017), quantifying the solubility of H<sub>2</sub>O at condition  
1722 relevant to lower crustal magma storage will require experimental innovations (e.g.  
1723 Makhluף et al., 2020; Mitchell et al., 2017). Models will also need to be developed  
1724 which are capable of calculating equilibria between a silicic melt phase and a com-  
1725 plex aqueous fluid (Ghiorso & Sverjensky, 2016; Sverjensky et al., 2014; Huang &  
1726 Sverjensky, 2019). Recently, a new equation of state for estimating the dielectric  
1727 constant of water and new equation of state parameters for major solute species has

1728 enabled thermodynamic mass transfer calculations for fluid-rock systems up to 6  
1729 GPa and 1000 ° C (Facq et al., 2014; Debret & Sverjensky, 2017; Huang & Sverjen-  
1730 sky, 2019; Sverjensky et al., 2014), significantly advancing our ability to model fluid  
1731 behavior in volcanically relevant systems (e.g., Iacovino et al., 2020).

## 1732 **9 Conclusion**

1733 This review uses the new open-source Python3 tool VESIcal (Iacovino et  
1734 al., 2021), in addition to VolatileCalc (Newman & Lowenstern, 2002) and Solw-  
1735 cad (Papale et al., 2006) to draw extensive comparisons between the behaviour of  
1736 9 different solubility models for a range of melt compositions. We show that these  
1737 models predict surprisingly different volatile solubilities, particularly for pure CO<sub>2</sub>  
1738 or mixed CO<sub>2</sub>-H<sub>2</sub>O fluids. Even for melt compositions that are well represented in  
1739 the calibration datasets of multiple models (e.g., MORBs), calculated solubilities for  
1740 pure CO<sub>2</sub> can deviate from one another by factors of ~2. Differential treatment of  
1741 H<sub>2</sub>O-CO<sub>2</sub> mixing enhances these differences when calculating volatile solubility for  
1742 melts containing both volatile species. The solubility of CO<sub>2</sub> predicted by different  
1743 rhyolitic models also differs substantially, overwhelming other sources of uncertainty  
1744 such as analytical errors on measurements of volatile contents or uncertainties in  
1745 crustal density profiles. Differences are most pronounced for peralkaline rhyolites  
1746 where there are fewer experimental constraints.

1747 Overall, these comparisons demonstrate that it is vital to pick a model which is  
1748 calibrated for the pressure, temperature, and melt composition of interest. Choice of  
1749 a poorly calibrated model could introduce a systematic error of a factor of 2 or more  
1750 in estimates of saturation pressures. This has widespread implications for published  
1751 estimates of magma storage depths within volcanic systems, and indicates that re-  
1752 evaluation of published magma storage depths calculated using older models may be  
1753 warranted.

1754 We also investigate the sensitivity of different models to variation in parame-  
1755 ters such as H<sub>2</sub>O content (with relevance to diffusive re-equilibration), CO<sub>2</sub> content  
1756 (with relevance to melt inclusion vapour bubble growth), temperature and oxygen  
1757 fugacity. We suggest that by performing similar sensitivity tests in the future, the  
1758 uncertainties affecting calculations of volatile solubility in magmatic systems (and

1759 therefore the limitations of each study) can be quantified. We also demonstrate that  
1760 isobar diagrams are a poor visualization method for determining magma storage  
1761 depths in systems where melt inclusions possess diverse melt inclusion chemistry, so  
1762 encourage users to take advantage of the ease and speed of calculations in VESIcal  
1763 to determine the saturation pressure for each melt inclusion of interest. Saturation  
1764 pressures can then be visualized as various cumulative frequency distributions, his-  
1765 tograms, or violin plots, and plotted against melt inclusion H<sub>2</sub>O and CO<sub>2</sub> contents,  
1766 or parameters relating to host crystal chemistry, to gain greater insight into the  
1767 factors controlling magma storage depths within volcanic systems.

1768 Finally, we identify that further experimental constraints are required to ac-  
1769 curately estimate volatile solubility in andesitic-dacitic melts, and that further work  
1770 is needed to understand the effect of temperature, redox, and non-ideal mixing be-  
1771 tween H<sub>2</sub>O-CO<sub>2</sub> on volatile solubility.

1772

### 1773 **Acknowledgments**

1774 PW acknowledges funding from a NERC DTP studentship (NE/L002507/1) and  
1775 a National Science Foundation grant (1948862). KI and GMM were supported by  
1776 the NASA Jacobs JETS Contract (NNJ13HA01C). We acknowledge support from  
1777 the National Science Foundation, ICER-20-26904, granted to OFM Research (PI.  
1778 Ghiorso) for maintaining the ENKI server used to perform VESIcal calculations. We  
1779 thank Jackie Dixon, Giada Iacono-Marziano, Paolo Papale, and Mark Ghiorso for  
1780 help tracking down and using existing calculation tools, and for help constructing  
1781 model calibration datasets. We also thank Mark Ghiorso and Paolo Papale for help-  
1782 ful reviews which improved the clarity and robustness of the first version, and David  
1783 Neaveon the secon Version. We greatly appreciate the late Peter Fox for his editorial  
1784 handling of the first draft of this manuscript, and his overall support for the VESIcal  
1785 project.

1786 **Data Availability** The Jupyter notebooks and associated Excel spread-  
1787 sheets used to create the figures in this manuscript are currently hosted at  
1788 [https://github.com/PennyWieser/VESIcal\\_II](https://github.com/PennyWieser/VESIcal_II), and will be added to the  
1789 main VESIcal github upon acceptance of this article. Videos showing how to

1790 use VESICAL are hosted on YouTube [https://www.youtube.com/channel/](https://www.youtube.com/channel/UCpvCCs5KMXz0xXWm0seF8Qw)  
 1791 UCpvCCs5KMXz0xXWm0seF8Qw.

## 1792 References

- 1793 Allan, J. F., BATIZA, R., PERFIT, M. R., FORNARI, D. J., & SACK, R. O.  
 1794 (1989). Petrology of lavas from the lamont seamount chain and adjacent  
 1795 east pacific rise, 10 n. *Journal of Petrology*, 30(5), 1245–1298.
- 1796 Allison, C. M., Roggensack, K., & Clark, A. (2021). Highly explosive basaltic erup-  
 1797 tions driven by co<sub>2</sub> exsolution. *Nature Communications*.
- 1798 Allison, C. M., Roggensack, K., & Clarke, A. B. (2019). H<sub>2</sub>O–CO<sub>2</sub> solubility in  
 1799 alkali-rich mafic magmas: new experiments at mid-crustal pressures. *Contribu-  
 1800 tions to Mineralogy and Petrology*, 174(7), 58.
- 1801 Anderson. (1974). Evidence for a picritic, volatile-rich magma beneath mt. shasta,  
 1802 california. *Journal of Petrology*, 15(2), 243–267.
- 1803 Aster, E. M., Wallace, P. J., Moore, L. R., Watkins, J., Gazel, E., & Bodnar, R. J.  
 1804 (2016). Reconstructing CO<sub>2</sub> concentrations in basaltic melt inclusions using  
 1805 raman analysis of vapor bubbles. *Journal of Volcanology and Geothermal  
 1806 Research*, 323, 148–162.
- 1807 Atlas, Z. D., Dixon, J. E., Sen, G., Finny, M., & Martin-Del Pozzo, A. L. (2006).  
 1808 Melt inclusions from volcán popocatepetl and volcán de colima, mexico: melt  
 1809 evolution due to vapor-saturated crystallization during ascent. *Journal of  
 1810 Volcanology and Geothermal Research*, 153(3-4), 221–240.
- 1811 Behrens, H., Ohlhorst, S., Holtz, F., & Champenois, M. (2004). CO<sub>2</sub> solubility in  
 1812 dacitic melts equilibrated with H<sub>2</sub>O–CO<sub>2</sub> fluids: Implications for modeling the  
 1813 solubility of CO<sub>2</sub> in silicic melts. *Geochimica et Cosmochimica Acta*, 68(22),  
 1814 4687–4703.
- 1815 Blank, J., Stolper, E., & Carroll, M. (1993). Solubilities of carbon dioxide and water  
 1816 in rhyolitic melt at 850 c and 750 bars. *Earth and Planetary Science Letters*,  
 1817 119(1-2), 27–36.
- 1818 Blundy, J., & Cashman, K. (2005). Rapid decompression-driven crystallization  
 1819 recorded by melt inclusions from mount st. helens volcano. *Geology*, 33(10),  
 1820 793–796.
- 1821 Blundy, J., Cashman, K., & Humphreys, M. (2006). Magma heating by



- 1822 decompression-driven crystallization beneath andesite volcanoes. *Nature*,  
1823 *443*(7107), 76–80.
- 1824 Botcharnikov, R. E., Behrens, H., & Holtz, F. (2006). Solubility and speciation of  
1825 c–o–h fluids in andesitic melt at t= 1100–1300 c and p= 200 and 500 mpa.  
1826 *Chemical Geology*, *229*(1-3), 125–143.
- 1827 Botcharnikov, R. E., Holtz, F., & Behrens, H. (2007). The effect of co<sub>2</sub> on the solu-  
1828 bility of h<sub>2</sub>o-cl fluids in andesitic melt. *European Journal of Mineralogy*, *19*(5),  
1829 671–680.
- 1830 Bowen, N. (1928). The evolution of the igneous rocks, princeton, univ. *Press*,  
1831 *Princeton, New Jersey*.
- 1832 Brooker, R., Kohn, S., Holloway, J., & McMillan, P. (2001a). Structural controls on  
1833 the solubility of co<sub>2</sub> in silicate melts: part i: bulk solubility data. *Chemical Ge-*  
1834 *ology*, *174*(1-3), 225–239.
- 1835 Brooker, R., Kohn, S., Holloway, J., & McMillan, P. (2001b). Structural controls  
1836 on the solubility of co<sub>2</sub> in silicate melts: part ii: Ir characteristics of carbonate  
1837 groups in silicate glasses. *Chemical Geology*, *174*(1-3), 241–254.
- 1838 Bucholz, C. E., Gaetani, G. A., Behn, M. D., & Shimizu, N. (2013). Post-  
1839 entrapment modification of volatiles and oxygen fugacity in olivine-hosted  
1840 melt inclusions. *Earth and Planetary Science Letters*, *374*, 145–155.
- 1841 Bureau, H., & Keppler, H. (1999). Complete miscibility between silicate melts and  
1842 hydrous fluids in the upper mantle: experimental evidence and geochemical  
1843 implications. *Earth and Planetary Science Letters*, *165*(2), 187–196.
- 1844 Burgisser, A., Alletti, M., & Scaillet, B. (2015). Simulating the behavior of volatiles  
1845 belonging to the c–o–h–s system in silicate melts under magmatic conditions  
1846 with the software d-compress. *Computers & Geosciences*, *79*, 1–14.
- 1847 Burnham, C. W. (1979). The importance of volatile constituents. *The evolution of*  
1848 *the igneous rocks*, 439–482.
- 1849 Burnham, C. W., & Davis, N. (1971). The role of h<sub>2</sub>o in silicate melts; i, pvt rela-  
1850 tions in the system naalsi<sub>3</sub>o<sub>8</sub>-h<sub>2</sub>o to 10 kilobars and 1000 degrees c. *Amer-*  
1851 *ican Journal of Science*, *270*(1), 54–79.
- 1852 Burnham, C. W., & Davis, N. (1974). The role of h<sub>2</sub>o in silicate melts; ii, ther-  
1853 modynamic and phase relations in the system naalsi<sub>3</sub>o<sub>8</sub>-h<sub>2</sub>o to 10 kilobars,  
1854 700 degrees to 1100 degrees c. *American Journal of Science*, *274*(8), 902–940.

- 1855 Cassidy, M., Edmonds, M., Watt, S. F., Palmer, M. R., & Gernon, T. M. (2015).  
1856 Origin of basalts by hybridization in andesite-dominated arcs. *Journal of*  
1857 *Petrology*, *56*(2), 325–346.
- 1858 Cocheo, P., & Holloway, J. (1993). The solubility of h<sub>2</sub>O in basanitic melts at low  
1859 pressure. *EOS Transactions of the American Geophysical Union* *74*.
- 1860 Cottrell, E., Lanzirrotti, A., Mysen, B., Birner, S., Kelley, K. A., Botcharnikov, R.,  
1861 ... Newville, M. (2018). A mössbauer-based xanes calibration for hydrous  
1862 basalt glasses reveals radiation-induced oxidation of Fe. *American Mineralogist:*  
1863 *Journal of Earth and Planetary Materials*, *103*(4), 489–501.
- 1864 Debret, B., & Sverjensky, D. (2017). Highly oxidising fluids generated during serpen-  
1865 tinite breakdown in subduction zones. *Scientific reports*, *7*(1), 1–6.
- 1866 Dingwell, D. B. (1986). Volatile solubilities in silicate melts.
- 1867 Dixon, J. E. (1997). Degassing of alkalic basalts. *American Mineralogist*, *82*(3-4),  
1868 368–378.
- 1869 Dixon, J. E., Stolper, E. M., & Holloway, J. R. (1995). An experimental study of  
1870 water and carbon dioxide solubilities in mid-ocean ridge basaltic liquids. part i:  
1871 calibration and solubility models. *Journal of Petrology*, *36*(6), 1607–1631.
- 1872 Duan, X. (2014). A general model for predicting the solubility behavior of h<sub>2</sub>O–CO<sub>2</sub>  
1873 fluids in silicate melts over a wide range of pressure, temperature and composi-  
1874 tions. *Geochimica et Cosmochimica Acta*, *125*, 582–609.
- 1875 Duan, Z., & Zhang, Z. (2006). Equation of state of the h<sub>2</sub>O, CO<sub>2</sub>, and h<sub>2</sub>O–CO<sub>2</sub> sys-  
1876 tems up to 10 gpa and 2573.15 K: Molecular dynamics simulations with ab  
1877 initio potential surface. *Geochimica et cosmochimica acta*, *70*(9), 2311–2324.
- 1878 Eggler, D. (1973). Role of CO<sub>2</sub> in melting processes in the mantle. *Carnegie Inst.*  
1879 *Wash. Yearb*, *72*, 457–467.
- 1880 Eguchi, J., & Dasgupta, R. (2018). A CO<sub>2</sub> solubility model for silicate melts from  
1881 fluid saturation to graphite or diamond saturation. *Chemical Geology*, *487*,  
1882 23–38.
- 1883 Facq, S., Daniel, I., Montagnac, G., Cardon, H., & Sverjensky, D. A. (2014). In situ  
1884 raman study and thermodynamic model of aqueous carbonate speciation in  
1885 equilibrium with aragonite under subduction zone conditions. *Geochimica et*  
1886 *Cosmochimica Acta*, *132*, 375–390.
- 1887 Fine, G., & Stolper, E. (1986). Dissolved carbon dioxide in basaltic glasses: con-

- 1888 concentrations and speciation. *Earth and Planetary Science Letters*, *76*(3-4), 263–  
1889 278.
- 1890 Flowers, G. C. (1979). Correction of holloway’s (1977) adaptation of the modified  
1891 redlich-kwong equation of state for calculation of the fugacities of molecular  
1892 species in supercritical fluids of geologic interest. *Contributions to Mineralogy  
1893 and Petrology*, *69*(3), 315–318.
- 1894 Fogel, R. A., & Rutherford, M. J. (1990). The solubility of carbon dioxide in rhy-  
1895 olitic melts; a quantitative ftir study. *American Mineralogist*, *75*(11-12), 1311–  
1896 1326.
- 1897 Freise, M. (2004). *Differenzierung von basalten einer “large igneous province”  
1898 am beispiel des kerguelen plateaus. eine experimentelle studie* (Unpublished  
1899 doctoral dissertation). PhD thesis, University of Hannover.
- 1900 Gaborieau, M., Laubier, M., Bolfan-Casanova, N., Mccammon, C., Vantelon, D.,  
1901 Chumakov, A., . . . Venugopal, S. (2020). Determination of  $Fe^{3+}/\sigma_{Fe}$  of olivine-  
1902 hosted melt inclusions using mössbauer and xanes spectroscopy. *Chemical  
1903 Geology*, 119646.
- 1904 Gaetani, G. A., O’Leary, J. A., Shimizu, N., Bucholz, C. E., & Newville, M. (2012).  
1905 Rapid reequilibration of  $H_2O$  and oxygen fugacity in olivine-hosted melt inclu-  
1906 sions. *Geology*, *40*(10), 915–918.
- 1907 Gavrilenko, M., Krawczynski, M., Ruprecht, P., Li, W., & Catalano, J. G. (2019).  
1908 The quench control of water estimates in convergent margin magmas. *Ameri-  
1909 can Mineralogist: Journal of Earth and Planetary Materials*, *104*(7), 936–948.
- 1910 Gerlach, T. M. (1986). Exsolution of  $H_2O$ ,  $CO_2$ , and S during eruptive episodes at ki-  
1911 lauea volcano, hawaii. *Journal of Geophysical Research: Solid Earth*, *91*(B12),  
1912 12177–12185.
- 1913 Ghiorso, M. S., Carmichael, I. S., Rivers, M. L., & Sack, R. O. (1983). The gibbs  
1914 free energy of mixing of natural silicate liquids; an expanded regular solution  
1915 approximation for the calculation of magmatic intensive variables. *Contribu-  
1916 tions to Mineralogy and Petrology*, *84*(2), 107–145.
- 1917 Ghiorso, M. S., & Gualda, G. A. (2015). An  $H_2O$ – $CO_2$  mixed fluid saturation model  
1918 compatible with rhyolite-melts. *Contributions to Mineralogy and Petrology*,  
1919 *169*(6), 1–30.
- 1920 Ghiorso, M. S., & Sack, R. O. (1995). Chemical mass transfer in magmatic pro-

- 1921           cesses iv. a revised and internally consistent thermodynamic model for the  
1922           interpolation and extrapolation of liquid-solid equilibria in magmatic systems  
1923           at elevated temperatures and pressures.           *Contributions to Mineralogy and*  
1924           *Petrology*, 119(2-3), 197–212.
- 1925   Ghiorso, M. S., & Sverjensky, D. A. (2016). The melts-dew connection: Integration  
1926           of thermodynamic models for magmatic systems and aqueous fluids at elevated  
1927           temperatures and pressures.           In *Agu fall meeting abstracts* (Vol. 2016, pp.  
1928           V33H–04).
- 1929   Gleeson, M. L., Gibson, S. A., & Williams, H. M. (2020). Novel insights from fe-  
1930           isotopes into the lithological heterogeneity of ocean island basalts and plume-  
1931           influenced morbs. *Earth and Planetary Science Letters*, 535, 116114.
- 1932   Gleeson, M. L., Stock, M. J., Pyle, D. M., Mather, T. A., Hutchison, W., Yirgu,  
1933           G., & Wade, J. (2017). Constraining magma storage conditions at a restless  
1934           volcano in the main ethiopian rift using phase equilibria models.           *Journal of*  
1935           *Volcanology and Geothermal Research*, 337, 44–61.
- 1936   Goltz, A. E., Krawczynski, M. J., Gavrilenko, M., Gorbach, N. V., & Ruprecht, P.  
1937           (2020). Evidence for superhydrous primitive arc magmas from mafic enclaves  
1938           at shiveluch volcano, kamchatka.           *Contributions to Mineralogy and Petrology*,  
1939           175(12), 1–26.
- 1940   Goranson, R. W. (1931). Solubility of water in granite magmas. *Eos, Transactions*  
1941           *American Geophysical Union*, 12(1), 183–183.
- 1942   Gualda, G. A., & Ghiorso, M. S. (2015). Melts \_ e xcel: Am icrosoft e xcel-based  
1943           melts interface for research and teaching of magma properties and evolution.  
1944           *Geochemistry, Geophysics, Geosystems*, 16(1), 315–324.
- 1945   Gualda, G. A., Ghiorso, M. S., Lemons, R. V., & Carley, T. L. (2012). Rhyolite-  
1946           melts: a modified calibration of melts optimized for silica-rich, fluid-bearing  
1947           magmatic systems. *Journal of Petrology*, 53(5), 875–890.
- 1948   Hamilton, D., Burnham, C. W., & Osborn, E. (1964). The solubility of water and ef-  
1949           fects of oxygen fugacity and water content on crystallization in mafic magmas.  
1950           *Journal of Petrology*, 5(1), 21–39.
- 1951   Hartley, M. E., Maclennan, J., Edmonds, M., & Thordarson, T. (2014). Reconstruct-  
1952           ing the deep co<sub>2</sub> degassing behaviour of large basaltic fissure eruptions. *Earth*  
1953           *and Planetary Science Letters*, 393, 120–131.

- 1954 Hartley, M. E., Neave, D. A., MacLennan, J., Edmonds, M., & Thordarson, T.  
1955 (2015). Diffusive over-hydration of olivine-hosted melt inclusions. *Earth*  
1956 *and Planetary Science Letters*, *425*, 168–178.
- 1957 Hauri, E. (2002). Sims analysis of volatiles in silicate glasses, 2: isotopes and abun-  
1958 dances in hawaiian melt inclusions. *Chemical Geology*, *183*(1-4), 115–141.
- 1959 Hauri, E., Kent, A. J., & Arndt, N. (2002). Melt inclusions at the millennium: to-  
1960 ward a deeper understanding of magmatic processes. *ChGeo*, *183*(1-4), 1–3.
- 1961 Helz, R., & Thornber, C. R. (1987). Geothermometry of kilauea iki lava lake, hawaii.  
1962 *Bulletin of Volcanology*, *49*(5), 651–668.
- 1963 Hervig, R., Dunbar, N., Westrich, H. R., & Kyle, P. R. (1989). Pre-eruptive water  
1964 content of rhyolitic magmas as determined by ion microprobe analyses of melt  
1965 inclusions in phenocrysts. *Journal of Volcanology and Geothermal Research*,  
1966 *36*(4), 293–302.
- 1967 Hervig, R., & Williams, P. (1988). Sims microanalysis of minerals and glasses for h  
1968 and d. *SIMS VI Proceedings*, 961–964.
- 1969 Hess, K., & Dingwell, D. (1996). Viscosities of hydrous leucogranitic melts: A non-  
1970 arrhenian model. *American Mineralogist*, *81*(9-10), 1297–1300.
- 1971 Holloway, J. R. (1977). Fugacity and activity of molecular species in supercritical  
1972 fluids. In *Thermodynamics in geology* (pp. 161–181). Springer.
- 1973 Holloway, J. R., & Blank, J. G. (1994). Application of experimental results to coh  
1974 species in natural melts. *Reviews in mineralogy*, *30*, 187–187.
- 1975 Huang, F., & Sverjensky, D. A. (2019). Extended deep earth water model for pre-  
1976 dicting major element mantle metasomatism. *Geochimica et Cosmochimica*  
1977 *Acta*, *254*, 192–230.
- 1978 Huber, C., Townsend, M., Degruyter, W., & Bachmann, O. (2019). Optimal depth  
1979 of subvolcanic magma chamber growth controlled by volatiles and crust rheol-  
1980 ogy. *Nature Geoscience*, *12*(9), 762–768.
- 1981 Huppert, H. E., & Woods, A. W. (2002). The role of volatiles in magma chamber  
1982 dynamics. *Nature*, *420*(6915), 493–495.
- 1983 Husen, A., Almeev, R. R., & Holtz, F. (2016). The effect of h<sub>2</sub>o and pressure on  
1984 multiple saturation and liquid lines of descent in basalt from the shatsky rise.  
1985 *Journal of Petrology*, *57*(2), 309–344.
- 1986 Iacono-Marziano, G., Morizet, Y., Le Trong, E., & Gaillard, F. (2012). New experi-

- 1987            mental data and semi-empirical parameterization of h<sub>2</sub>o-co<sub>2</sub> solubility in mafic  
1988            melts. *Geochimica et Cosmochimica Acta*, *97*, 1–23.
- 1989    Iacovino, K., Guild, M. R., & Till, C. B. (2020). Aqueous fluids are effective oxidiz-  
1990            ing agents of the mantle in subduction zones. *Contributions to Mineralogy and*  
1991            *Petrology*, *175*(4), 1–21.
- 1992    Iacovino, K., Matthews, S., Wieser, P. E., Moore, G., & Begue, F. (2021). Vesi-  
1993            cal part i: An open source thermodynamic model engine for mixed volatile  
1994            solubility in silicate melts. *EarthArxiv* -<https://doi.org/10.31223/X5D606>.
- 1995    Iacovino, K., Moore, G., Roggensack, K., Oppenheimer, C., & Kyle, P. (2013). H<sub>2</sub>  
1996            o-co<sub>2</sub> solubility in mafic alkaline magma: applications to volatile sources and  
1997            degassing behavior at erebus volcano, antarctica. *Contributions to Mineralogy*  
1998            *and Petrology*, *166*(3), 845–860.
- 1999    Iddon, F., & Edmonds, M. (2020). Volatile-rich magmas distributed through the  
2000            upper crust in the main ethiopian rift. *Geochemistry, Geophysics, Geosystems*,  
2001            *21*(6), e2019GC008904.
- 2002    Jakobsson, S. (1997). Solubility of water and carbon dioxide in an icelandite at 1400  
2003            c and 10 kilobars. *Contributions to Mineralogy and Petrology*, *127*(1-2), 129–  
2004            135.
- 2005    Kerrick, D., & Jacobs, G. (1981). A modified redlich-kwong equation for h<sub>2</sub>o, co  
2006            2, and h<sub>2</sub>o-co<sub>2</sub> mixtures at elevated pressures and temperatures. *American*  
2007            *Journal of Science*, *281*(6), 735–767.
- 2008    King, P., & Holloway, J. (2002). Co<sub>2</sub> solubility and speciation in intermediate (an-  
2009            desitic) melts: the role of h<sub>2</sub>o and composition. *Geochimica et Cosmochimica*  
2010            *Acta*, *66*(9), 1627–1640.
- 2011    Koleszar, A., Kent, A. J., Wallace, P. J., & Scott, W. E. (2012). Controls on long-  
2012            term low explosivity at andesitic arc volcanoes: Insights from mount hood,  
2013            oregon. *Journal of Volcanology and Geothermal Research*, *219*, 1–14.
- 2014    Koleszar, A., Saal, A., Hauri, E., Nagle, A., Liang, Y., & Kurz, M. (2009). The  
2015            volatile contents of the galapagos plume; evidence for h<sub>2</sub>o and f open system  
2016            behavior in melt inclusions. *Earth and Planetary Science Letters*, *287*(3-4),  
2017            442–452.
- 2018    La Spina, G., Arzilli, F., Llewellyn, E., Burton, M., Clarke, A. B., Vitturi, M. d.,  
2019            ... Mader, H. (2021). Explosivity of basaltic lava fountains is controlled by

- 2020 magma rheology, ascent rate and outgassing. *Earth and Planetary Science*  
2021 *Letters*, 553, 116658.
- 2022 Lesne, P., Scaillet, B., Pichavant, M., & Beny, J.-M. (2011). The carbon dioxide  
2023 solubility in alkali basalts: an experimental study. *Contributions to Mineralogy*  
2024 *and Petrology*, 162(1), 153–168.
- 2025 Lesne, P., Scaillet, B., Pichavant, M., Iacono-Marziano, G., & Beny, J.-M. (2011).  
2026 The h<sub>2</sub>o solubility of alkali basaltic melts: an experimental study. *Contribu-*  
2027 *tions to Mineralogy and Petrology*, 162(1), 133–151.
- 2028 Le Voyer, M., Hauri, E., Cottrell, E., Kelley, K. A., Salters, V. J., Langmuir, C. H.,  
2029 ... Füre, E. (2019). Carbon fluxes and primary magma co<sub>2</sub> contents along the  
2030 global mid-ocean ridge system. *Geochemistry, Geophysics, Geosystems*, 20(3),  
2031 1387–1424.
- 2032 Li, Y., Dasgupta, R., & Tsuno, K. (2017). Carbon contents in reduced basalts at  
2033 graphite saturation: Implications for the degassing of mars, mercury, and the  
2034 moon. *Journal of Geophysical Research: Planets*, 122(6), 1300–1320.
- 2035 Liu, Y., Zhang, Y., & Behrens, H. (2005). Solubility of h<sub>2</sub>o in rhyolitic melts at low  
2036 pressures and a new empirical model for mixed h<sub>2</sub>o–co<sub>2</sub> solubility in rhyolitic  
2037 melts. *Journal of Volcanology and Geothermal Research*, 143(1-3), 219–235.
- 2038 Lloyd, A. S., Plank, T., Ruprecht, P., Hauri, E., & Rose, W. (2013). Volatile loss  
2039 from melt inclusions in pyroclasts of differing sizes. *Contributions to Mineral-*  
2040 *ogy and Petrology*, 165(1), 129–153.
- 2041 Lowenstern, J. B. (1995). Applications of silicate-melt inclusions to the study of  
2042 magmatic volatiles. *Magmas, fluids and ore deposits*, 23, 71–99.
- 2043 Lowenstern, J. B. (2001). Carbon dioxide in magmas and implications for hydrother-  
2044 mal systems. *Mineralium Deposita*, 36(6), 490–502.
- 2045 Lowenstern, J. B. (2003). Melt inclusions come of age: volatiles, volcanoes, and  
2046 sorby’s legacy. In *Developments in volcanology* (Vol. 5, pp. 1–21). Elsevier.
- 2047 Lucic, G., Berg, A.-S., & Stix, J. (2016). Water-rich and volatile-undersaturated  
2048 magmas at hekla volcano, iceland. *Geochemistry, Geophysics, Geosystems*,  
2049 17(8), 3111–3130.
- 2050 MacLennan, J. (2017). Bubble formation and decrepitation control the co<sub>2</sub> content of  
2051 olivine-hosted melt inclusions. *Geochemistry, Geophysics, Geosystems*, 18(2),  
2052 597–616.

- 2053 Makhluף, A. R., Newton, R., & Manning, C. (2020). Experimental investigation of  
2054 phase relations in the system  $\text{NaAlSi}_3\text{O}_8\text{-H}_2\text{O}$  at high temperatures and pres-  
2055 sures: liquidus relations, liquid–vapor mixing, and critical phenomena at deep  
2056 crust–upper mantle conditions. *Contributions to Mineralogy and Petrology*,  
2057 *175*(8), 1–20.
- 2058 Mangan, M. T., Sisson, T. W., Hankins, W. B., Shimizu, N., & Vennemann, T.  
2059 (2021). Constraints on deep,  $\text{CO}_2$ -rich degassing at arc volcanoes from solubility  
2060 experiments on hydrous basaltic andesite of Pavlof volcano, Alaska peninsula,  
2061 at 300 to 1200 MPa. *American Mineralogist: Journal of Earth and Planetary*  
2062 *Materials*, *106*(5), 762–773.
- 2063 Métrich, N., & Wallace, P. J. (2008). Volatile abundances in basaltic magmas and  
2064 their degassing paths tracked by melt inclusions. *Reviews in mineralogy and*  
2065 *geochemistry*, *69*(1), 363–402.
- 2066 Mironov, N., Tobelko, D., Smirnov, S., Portnyagin, M. V., & Krasheninnikov, S.  
2067 (2020). Estimation of  $\text{CO}_2$  content in the gas phase of melt inclusions using  
2068 Raman spectroscopy: Case study of inclusions in olivine from the Karymsky  
2069 volcano (Kamchatka). *Russian Geology and Geophysics*, *61*(5-6), 600–610.
- 2070 Mitchell, A. L., Gaetani, G. A., O’leary, J. A., & Hauri, E. H. (2017).  $\text{H}_2\text{O}$  sol-  
2071 ubility in basalt at upper mantle conditions. *Contributions to Mineralogy and*  
2072 *Petrology*, *172*(10), 1–16.
- 2073 Moore, G. (2008). Interpreting  $\text{H}_2\text{O}$  and  $\text{CO}_2$  contents in melt inclusions: constraints  
2074 from solubility experiments and modeling. *Reviews in Mineralogy and Geo-*  
2075 *chemistry*, *69*(1), 333–362.
- 2076 Moore, G., & Carmichael, I. (1998). The hydrous phase equilibria (to 3 kbar) of  
2077 an andesite and basaltic andesite from western Mexico: constraints on water  
2078 content and conditions of phenocryst growth. *Contributions to Mineralogy and*  
2079 *Petrology*, *130*(3-4), 304–319.
- 2080 Moore, G., Roggensack, K., & Klonowski, S. (2008). A low-pressure–high-  
2081 temperature technique for the piston-cylinder. *American Mineralogist*, *93*(1),  
2082 48–52.
- 2083 Moore, G., Vennemann, T., & Carmichael, I. (1998). An empirical model for the sol-  
2084 ubility of  $\text{H}_2\text{O}$  in magmas to 3 kilobars. *American Mineralogist*, *83*(1), 36–42.
- 2085 Moore, L. R., Gazel, E., Tuohy, R., Lloyd, A. S., Esposito, R., Steele-MacInnis, M.,



- 2086 ... Bodnar, R. J. (2015). Bubbles matter: An assessment of the contribution  
2087 of vapor bubbles to melt inclusion volatile budgets. *American Mineralogist*,  
2088 *100*(4), 806–823.
- 2089 Mutch, E. J., Maclennan, J., Shorttle, O., Edmonds, M., & Rudge, J. F. (2019).  
2090 Rapid transcrustal magma movement under iceland. *Nature Geoscience*, *12*(7),  
2091 569–574.
- 2092 Mysen, B. O. (1976). The role of volatiles in silicate melts; solubility of carbon  
2093 dioxide and water in feldspar, pyroxene, and feldspathoid melts to 30 kb and  
2094 1625 degrees c. *American Journal of Science*, *276*(8), 969–996.
- 2095 Mysen, B. O., Egger, D. H., Seitz, M., & Holloway, J. R. (1976). Carbon dioxide in  
2096 silicate melts and crystals; part i, solubility measurements. *American Journal*  
2097 *of Science*, *276*(4), 455–479.
- 2098 Mysen, B. O., Fogel, M. L., Morrill, P. L., & Cody, G. D. (2009). Solution behavior  
2099 of reduced coh volatiles in silicate melts at high pressure and temperature.  
2100 *Geochimica et Cosmochimica Acta*, *73*(6), 1696–1710.
- 2101 Newman, S., & Lowenstern, J. B. (2002). Volatilecalc: a silicate melt–h<sub>2</sub>o–co<sub>2</sub> solu-  
2102 tion model written in visual basic for excel. *Computers & Geosciences*, *28*(5),  
2103 597–604.
- 2104 Ochs, F. A., & Lange, R. A. (1999). The density of hydrous magmatic liquids. *Sci-*  
2105 *ence*, *283*(5406), 1314–1317.
- 2106 Papale, P. (1997). Modeling of the solubility of a one-component h<sub>2</sub>o or co<sub>2</sub>  
2107 fluid in silicate liquids. *Contributions to Mineralogy and Petrology*, *126*(3),  
2108 237–251.
- 2109 Papale, P. (1999). Modeling of the solubility of a two-component h<sub>2</sub>o+ co<sub>2</sub> fluid in  
2110 silicate liquids. *American Mineralogist*, *84*(4), 477–492.
- 2111 Papale, P., Moretti, R., & Barbato, D. (2006). The compositional dependence of  
2112 the saturation surface of h<sub>2</sub>o+ co<sub>2</sub> fluids in silicate melts. *Chemical Geology*,  
2113 *229*(1-3), 78–95.
- 2114 Papale, P., Neri, A., & Macedonio, G. (1999). The role of water content and magma  
2115 composition on explosive eruption dynamics. *Physics and Chemistry of the*  
2116 *Earth, Part A: Solid Earth and Geodesy*, *24*(11-12), 969–975.
- 2117 Plank, T., Kelley, K. A., Zimmer, M. M., Hauri, E., & Wallace, P. J. (2013). Why  
2118 do mafic arc magmas contain 4 wt% water on average? *Earth and Planetary*

- 2119 *Science Letters*, 364, 168–179.
- 2120 Portnyagin, M., Almeev, R., Matveev, S., & Holtz, F. (2008). Experimental evidence  
2121 for rapid water exchange between melt inclusions in olivine and host magma.  
2122 *Earth and Planetary Science Letters*, 272(3-4), 541–552.
- 2123 Putirka, K. D. (2008). Thermometers and barometers for volcanic systems. *Reviews*  
2124 *in Mineralogy and Geochemistry*, 69(1), 61–120.
- 2125 Rasmussen, D. J., Plank, T. A., Wallace, P. J., Newcombe, M. E., & Lowenstern,  
2126 J. B. (2020). Vapor-bubble growth in olivine-hosted melt inclusions. *American*  
2127 *Mineralogist: Journal of Earth and Planetary Materials*, 105(12), 1898–1919.
- 2128 Reubi, O., & Blundy, J. (2009). A dearth of intermediate melts at subduction zone  
2129 volcanoes and the petrogenesis of arc andesites. *Nature*, 461(7268), 1269–  
2130 1273.
- 2131 Roedder, E. (1979). Origin and significance of magmatic inclusions. *Bulletin de Min-*  
2132 *eralogie*, 102(5), 487–510.
- 2133 Roggensack, K. (2001). Unraveling the 1974 eruption of fuego volcano (guatemala)  
2134 with small crystals and their young melt inclusions. *Geology*, 29(10), 911–914.
- 2135 Rohatgi, A. (2017). *Webplotdigitizer*. Austin, Texas, USA.
- 2136 Saal, A. E., Hauri, E., Langmuir, C. H., & Perfit, M. R. (2002). Vapour undersat-  
2137 uration in primitive mid-ocean-ridge basalt and the volatile content of earth’s  
2138 upper mantle. *Nature*, 419(6906), 451–455.
- 2139 Shishkina, T., Botcharnikov, R. E., Holtz, F., Almeev, R., & Portnyagin, M. V.  
2140 (2010). Solubility of h<sub>2</sub>o-and co<sub>2</sub>-bearing fluids in tholeiitic basalts at pressures  
2141 up to 500 mpa. *Chemical geology*, 277(1-2), 115–125.
- 2142 Shishkina, T., Botcharnikov, R. E., Holtz, F., Almeev, R. R., Jazwa, A. M., & Jaku-  
2143 biak, A. A. (2014). Compositional and pressure effects on the solubility of h<sub>2</sub>o  
2144 and co<sub>2</sub> in mafic melts. *Chemical Geology*, 388, 112–129.
- 2145 Sides, I., Edmonds, M., MacLennan, J., Swanson, D., & Houghton, B. (2014a). Erup-  
2146 tion style at kilauea volcano in hawai ‘i linked to primary melt composition.  
2147 *Nature Geoscience*, 7(6), 464.
- 2148 Sides, I., Edmonds, M., MacLennan, J., Swanson, D., & Houghton, B. (2014b).  
2149 Magma mixing and high fountaining during the 1959 kilauea iki eruption,  
2150 hawai ‘i. *Earth and Planetary Science Letters*, 400, 102–112.
- 2151 Silver, L. A. (1988). *Water in silicate glasses* (Unpublished doctoral dissertation).

- 2152 California Institute of Technology.
- 2153 Silver, L. A., Ihinger, P. D., & Stolper, E. (1990). The influence of bulk composition  
2154 on the speciation of water in silicate glasses. *Contributions to Mineralogy and*  
2155 *Petrology*, *104*(2), 142–162.
- 2156 Silver, L. A., & Stolper, E. (1989). Water in albitic glasses. *Journal of petrology*,  
2157 *30*(3), 667–709.
- 2158 Steele-Macinnis, M., Esposito, R., & Bodnar, R. J. (2011). Thermodynamic model  
2159 for the effect of post-entrapment crystallization on the  $\text{H}_2\text{O}$ - $\text{CO}_2$  systematics  
2160 of vapor-saturated, silicate melt inclusions. *Journal of Petrology*, *52*(12),  
2161 2461–2482.
- 2162 Stevenson, J. (2015). *Tasplot*. [https://bitbucket.org/jsteven5/tasplot/src/](https://bitbucket.org/jsteven5/tasplot/src/master/)  
2163 [master/](https://bitbucket.org/jsteven5/tasplot/src/master/). bitbucket.
- 2164 Stolper, E. (1982). Water in silicate glasses: an infrared spectroscopic study. *Contri-*  
2165 *butions to Mineralogy and Petrology*, *81*(1), 1–17.
- 2166 Stolper, E., Fine, G., Johnson, T., & Newman, S. (1987). Solubility of carbon diox-  
2167 ide in albitic melt. *American Mineralogist*, *72*(11-12), 1071–1085.
- 2168 Sverjensky, D. A., Harrison, B., & Azzolini, D. (2014). Water in the deep earth: the  
2169 dielectric constant and the solubilities of quartz and corundum to 60 kb and  
2170 1200 c. *Geochimica et Cosmochimica Acta*, *129*, 125–145.
- 2171 Tamura, Y., & Tatsumi, Y. (2002). Remelting of an andesitic crust as a possible  
2172 origin for rhyolitic magma in oceanic arcs: an example from the Izu–Bonin arc.  
2173 *Journal of Petrology*, *43*(6), 1029–1047.
- 2174 Tucker, J. M., Hauri, E., Pietruszka, A. J., Garcia, M. O., Marske, J. P., & Trusdell,  
2175 F. A. (2019). A high carbon content of the Hawaiian mantle from olivine-  
2176 hosted melt inclusions. *Geochimica et Cosmochimica Acta*, *254*, 156–172.
- 2177 Tuttle, O. F., & Bowen, N. L. (1958). *Origin of granite in the light of experimental*  
2178 *studies in the system NaAlSi<sub>3</sub>O<sub>8</sub>-KAlSi<sub>3</sub>O<sub>8</sub>-SiO<sub>2</sub>-H<sub>2</sub>O* (Vol. 74). Geological Society  
2179 of America.
- 2180 Venugopal, S., Schiavi, F., Moune, S., Bolfan-Casanova, N., Druitt, T., & Williams-  
2181 Jones, G. (2020). Melt inclusion vapour bubbles: the hidden reservoir for  
2182 major and volatile elements. *Scientific Reports*, *10*(1), 1–14.
- 2183 Wallace, P. J., Anderson, A. T., & Davis, A. M. (1995). Quantification of pre-  
2184 eruptive exsolved gas contents in silicic magmas. *Nature*, *377*(6550), 612–616.

- 2185 Waters, L. E., & Lange, R. A. (2013). Crystal-poor, multiply saturated rhyolites  
2186 (obsidians) from the cascade and mexican arcs: evidence of degassing-induced  
2187 crystallization of phenocrysts. *Contributions to Mineralogy and Petrology*,  
2188 *166*(3), 731–754.
- 2189 Waters, L. E., & Lange, R. A. (2015). An updated calibration of the plagioclase-  
2190 liquid hygrometer-thermometer applicable to basalts through rhyolites. *Ameri-  
2191 can Mineralogist*, *100*(10), 2172–2184.
- 2192 Wieser, P. E., Lamadrid, H., Maclennan, J., Edmonds, M., Matthews, S., Iacovino,  
2193 K., . . . others (2021). Reconstructing magma storage depths for the 2018  
2194 kilauean eruption from melt inclusion co2 contents: the importance of vapor  
2195 bubbles. *Geochemistry, Geophysics, Geosystems*, *22*(2), e2020GC009364.