

# 1 Lowermost mantle shear-velocity structure from hierarchical trans- 2 dimensional Bayesian tomography

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## 6 7 **Abstract**

8 The core-mantle boundary (CMB) is the most extreme boundary within the Earth where the liquid,  
9 iron-rich outer core interacts with the rocky, silicate mantle. The nature of the lowermost mantle atop  
10 the CMB, and its role in mantle dynamics, is not completely understood. Various regional studies  
11 have documented significant heterogeneities at different spatial scales. While there is a consensus on  
12 the long scale-length structure of the inferred S-wave speed tomograms, there are also notable  
13 differences stemming from different imaging methods and datasets. Here we aim to overcome over-  
14 smoothing and avoid over-fitting data for the case where the spatial coverage is sparse and the  
15 inverse problem ill-posed. Here we present an S-wave tomography model at global scale for the  
16 Lowermost Mantle (LM) using the Hierarchical Trans-dimensional Bayesian Inversion (HTDBI)  
17 framework, LM-HTDBI. Our HTDBI analysis of ScS-S travel times includes uncertainty, and the  
18 complexity of the model is deduced from the data itself through an implicit parameterization of the  
19 model space. Our comprehensive resolution estimates indicate that short-scale anomalies are  
20 significant and resolvable features of the lowermost mantle regardless of the chosen mantle-model  
21 reference to correct the travel times above the D'' layer. The recovered morphology of the Large-  
22 Low-Shear-wave Velocity Provinces (LLSVPs) is complex, featuring small high-velocity patches  
23 among low-velocity domains. Instead of two large, unified, and smooth LLSVPs, the newly obtained  
24 images suggest that their margins are not uniformly flat.

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## 29 **Plain Language Summary**

30 The lowermost mantle sits atop the core-mantle boundary - the most dramatic boundary within our  
31 planet, with contrasts in physical properties that exceed those that exist at the surface. Despite  
32 significant progress in recent years, this part of the Earth is not well understood, and various  
33 tomographic studies on global scale, along with regional studies that focus on seismic waveform  
34 modelling, pave the path towards higher resolution and new understanding. Important questions to  
35 answer are on the distribution, shape, size and composition of inhomogeneities in the lowermost  
36 mantle, and their critical role in mantle and core dynamics.  
37 While there is a general consensus on the long-scale length structures inferred from long-period  
38 shear waves, there are notable differences in details of the tomograms of the lowermost mantle,  
39 stemming from the use of different imaging methods and datasets. Here, we utilize a large traveltimes  
40 dataset of ScS and S waves with a significant addition of new measurements sensitive to the  
41 lowermost mantle to perform a probabilistic shear-wave tomography, and we retrieve a high-  
42 resolution image of the lowermost mantle. The new shear-wave speed tomogram and comprehensive  
43 resolution-estimations indicate that short and medium scale inhomogeneities are omnipresent  
44 features of the lowermost mantle.



## 45 **1. Introduction**

46 The core-mantle boundary (CMB) lies approximately halfway down to the Earth's center and is the  
47 boundary with the most extreme contrasts in seismic velocity, viscosity, and density, where the heat  
48 from the outer core is being transmitted to the mantle. This thermal boundary per se cannot explain  
49 the observation of isolated blobs of partial melt (Williams & Garnero, 1996; Li 2002; Suzuki et al.  
50 2020) and sharp changes in seismic velocity and anisotropy (e.g. Wenk et al., 2011; Nawacki &  
51 Woodi 2016; Creasy et al., 2017) in the bottom 300-km part of the mantle, also known as D''.  
52 Geodynamical simulations show that mantle convection sweeps intrinsically hotter material towards  
53 upwelling regions (e.g. McNamara & Zhong, 2005) and alternatively, pushes the subducted slabs to  
54 the base of mantle (e.g. Ricard et al., 1993; Fukao et al. 2001). Therefore, the subduction history can  
55 potentially lead to the development of thermochemical heterogeneities in D''.

56 The two dominant features at the top of the CMB identified by seismic imaging of long-period S-  
57 waves are the so-called Large-Low-Shear-wave Velocity Provinces (LLSVPs), and they are situated  
58 roughly beneath Africa and the Pacific (e.g. Masters et al., 2000; Garnero and McNamara,  
59 2008; Simmons et al., 2010) covering more than 25% of the CMB (e.g. Koelemeijer et al., 2017).  
60 The geographical position and overall shape of LLSVPs, as inferred from different tomographic  
61 models using various datasets and methods, are in a good general agreement, although differences  
62 are present with regards to the specific shape and lateral extension of LLSVP margins (eg. Garnero  
63 et al, 2016; Cottaar & Lekić, 2016). A recent geochemical analysis has shown that the two LLSVPs  
64 hold different evolutionary histories which are related to the assembly and breakup of the  
65 supercontinents, supporting a dynamic relationship between plate tectonics and the lowermost mantle  
66 structures (Doucet et al., 2020).

67 Apart from LLSVPs, past seismological studies that involve waveform modeling presented  
68 arguments for the existence of intermediate heterogeneity (e.g. Rost & Earle 2010; Wysession et al.,  
69 1994), sharp transitions (e.g. Ni and Helmberger, 2001), short-scale heterogeneity (e.g. Tkalčić and  
70 Romanowicz, 2002); as well as ultra-low-velocity zones (ULVZs) (e.g. Garnero and Helmberger,  
71 1995; Avants et al., 2005; Thorne et al., 2013; Pachhai et al., 2014; McNamara, 2019). Recent  
72 studies by employing P-wave travel times that focused either solely on the lowermost mantle (Young  
73 et al., 2014; Tkalčić et al., 2015; Muir and Tkalčić, 2020) or the whole mantle (Hosseini et al., 2019)  
74 have indicated that the D'' is far more complex than what global S-wave studies revealed (e.g.  
75 Grand, 2002; Simmons et al., 2010; Moulik & Ekström 2014), with heterogeneity scales comparable  
76 to those found in the Earth's crust.

77 Importantly, while various regional studies have argued for the existence of short-scale  
78 heterogeneities (e.g. He & Wen 2011; Takeuchi, et al., 2008; Takeuchi 2012; Frost et al, 2013;  
79 Hung et al., 2005; Suzuki et al., 2020; Li et al., 2021) the global seismic studies of shear-waves  
80 robustly revealed a dominant character of long-scale features (e.g. Houser et al., 2008; Durand et al.,  
81 2017). A recent mantle model using shear waves, developed by French and Romanowicz (2014)  
82 utilized waveform inversion of selected windows of body waves (300 – 32 s) and surface waves (400  
83 – 60 s) to provide a detailed image of the Earth’s mantle. This model also found evidence for short-  
84 scale heterogeneity at the top of the CMB.

85 The remarkable diversity of the short and medium scale anomalies suggests a significant uncertainty  
86 amongst models, even at the scales that they allegedly resolve (e.g. Moulik and Ekström 2014;  
87 Kustowski et al., 2008; Simmons et al., 2010). In our hypothesis, most discrepancies among various  
88 tomographic models stem from the adopted procedures. First, they are prone to subjective choices of  
89 model parameters and regularization, which, in turn, determines the level of model complexity. In  
90 model regularizations, smoothing and damping are usually applied to the entire model, and  
91 consequently, they mute the sharp edges that are likely present in Earth and transform them into  
92 pixelated gradual transitions. Second, building models based on explicit parameterizations such as  
93 regular cells which ignores the spatial variability in the data’s resolving power usually introduces  
94 small-scale artefacts in regions with less ray coverage. Third, the difficulty in quantifying data  
95 uncertainties takes its toll since the level of data noise is usually determined by the user, prior to the  
96 inversion, and it regulates the level of data fit.

97 To take a progressive approach and improve the imaging capability of structural features in the D’’  
98 by using travel times only, we believe it is essential to possess the following two ingredients: a)  
99 differential travel times of body waves, whose use suppresses the unwanted crustal and upper mantle  
100 effects near the source and receiver, thus capturing the contribution from the lowermost mantle, b) an  
101 improved theoretical and methodological framework to overcome the aforementioned limitations in  
102 various seismic-tomography inversion approaches. For this purpose, we use the Hierarchical Trans-  
103 dimensional Bayesian Inversion (HTDBI) framework to perform a tomographic inversion that  
104 circumvents subjective choices. In fact, the first global imaging application of the HTDBI method  
105 was to study the lowermost mantle (Young et al., 2013; Tkalčić et al. 2015), but the improvement  
106 here in comparison with the previous studies is that we now deploy truly-spherical Voronoi  
107 parameterization. This type of parameterization has recently been applied in an attenuation  
108 tomography (Pejić et al., 2019) and the P-wave velocity tomography studies of the upper inner core

109 (Burdick et al., 2019). The main limitation of our study (due to the computational expense  
110 consideration) is that we do not account for the finite frequency effects. In contrast, other global  
111 tomography studies have utilized more complete information by utilizing waveforms, a wide range  
112 of body waves and phases, and normal modes.

113 In subsequent Sections, we present a development of a new shear-wave velocity model of the D''. In  
114 Section 2, we explain the data that was used in the inversion. In Section 3, we discuss how the  
115 HTDBI method differs from the linearized travel-time tomography. Next, in Section 4, we present  
116 the resolution tests including an all-embracing test for an abstract, complex, D'' model and a test for  
117 a more realistic structure, which mimics the final pattern of anomalies in D''. In Section 5, we  
118 explain the inversion and data analysis that we carried out through HTDBI. In Section 6, we first  
119 describe our approach to correcting for mantle structure and then present the averaged shear-wave  
120 velocity model together with its standard deviation map. In Sections 7 and 8 we compare our model  
121 with other available models of D'' and discuss the advantages of using HTDBI in global imaging of  
122 D''.

## 123 **2. Data**

124 In this study, we utilize differential travel time of shear waves to image the heterogeneity pattern at  
125 the lowermost mantle. Differential travel times of body waves are particularly helpful: first, by  
126 reducing the unwanted crustal and upper mantle effects near the source and receiver, since the  
127 raypaths are close for the large epicentral distances in the crust and upper mantle; second, source  
128 mislocations are not projected into the model.

129 We used ScS-S differential travel times from Houser et al., (2008) which encompasses events from  
130 1976 – 2007, and augmented it by our newly measured data in the time window 2008 – 2018,  
131 focusing on the source-receiver geometries that fill the gaps in spatial coverage of the Houser et al  
132 (2008) dataset. We considered only earthquakes with simple source-time functions and magnitudes  
133  $5.5 < M < 7.5$ . To reduce the effect of upper mantle structure on our ScS-S data set, we imposed a  
134 relatively strict criterion and only selected the event-station pairs at epicentral distances between 60  
135 and 75 degrees (Fig. 1a). In these epicentral distances, the raypaths of ScS and S waves are more  
136 proximate to each other and thus their differential travel time is a better proxy for the travel time  
137 affected by D'' structure where they differ the most. The vertical separation between the bottoming  
138 points of S and ScS varies from 1450 km to 950 km for epicentral distances of 60 and 75 degrees.  
139 Our data-selection procedure resulted in 18,131 rays in total, including 16,344 rays out of the

140 original 41,000 rays (Houser et al., 2008) indicated by the red dashed line in Fig. 1b, and the newly-  
141 collected 1,787 rays, to augment the geometric coverage and enhance the imaging capacity of D''.  
142 The resulting coverage of rays in CMB is shown in the top panel of Fig. 2 and the Supporting  
143 Information (Fig. S1). A simple calculation for our dataset shows that over 70% of the CMB has a hit  
144 count of 5 or more if we divide the area of CMB to 4 ° by 4 ° cells. The distribution of the bounce  
145 points of ScS rays and their raypaths for both datasets at the CMB are shown in Fig. 2b & 2c. We  
146 compare the newly picked data and dataset from Houser et al., (2008) in the Supporting Information  
147 (Fig. S2 to Fig. S3).

148 To ensure the compatibility between the two datasets we followed the same measurement technique  
149 as Houser et al., (2008) and handpicked the differential travel times through waveform correlation of  
150 band-pass filtered waveform data at 10–50 s (Fig. 3). Correspondingly, the size of the first Fresnel  
151 zone for ScS in the lowermost mantle is near 400 km in diameter, which can be interpreted as a  
152 rough approximation of the potential resolution of the data in the areas with intermediate or low  
153 coverage and a criterion for the size of resolved short-scale heterogeneities.

154 The advantage of handpicking is that it allows us to align the onsets of S and ScS phases and discard  
155 noisy waveforms. The measurements were made on transverse components to suppress any  
156 interference from converted phases.

157 Given the fact that S and ScS raypaths are not identical in the lower mantle (albeit they are similar),  
158 to further reduce the effects of mantle heterogeneity, S and ScS travel times were corrected for 3D  
159 mantle structure using 1D ray paths down to the top of D'' (300 km above CMB) using five global  
160 mantle models: TX2011 (Grand, 2002), S362ANI+M (Moulik and Ekström, 2014), GyPSuM-S  
161 (Simmons et al, 2010), SEMUCB-WM1 (French and Romanowicz, 2014) and SAW642ANb  
162 (Panning et al., 2010). S and ScS traced in the mantle using “tau-p” (Buland and Chapman, 1983)  
163 through PREM and the extracted mantle velocity structure (based on the selected five global models  
164 and their reference models) along each ray path give the time corrections for S and ScS. While this  
165 does not account for ray bending that might occur due to 3D structure, the difference between  
166 velocities sampled along 1D and 3D ray paths is minimal given the long wavelength of features  
167 imaged in tomographic models. Mantle correction reasonably suppresses the projection of mantle  
168 heterogeneities into the inferred structure of D''. At these epicentral distances and in the crust, the  
169 raypaths of S and ScS are nearly identical, which means that the correction for the crust is not  
170 necessary. The corrected ScS-S data from all 5 models are then inverted for the velocity anomalies in  
171 D''.

172

### 173 **3. Method**

174 We employed the HTDBI method (Bodin et al., 2012; Young et al., 2013; Pejić et al., 2019) as our  
175 preferred inversion method to image the lowermost mantle. The HTDBI method includes the overall  
176 level of data noise as a free parameter in the inversion process, together with the number of model  
177 parameters. Consequently, large, small, smooth, and sharp discontinuous features can be resolved  
178 based on the trade-off between data uncertainty and resolution. In the non-HTDBI tomographic  
179 inversions, arbitrary choices, such as a single parameterization type, damping, and smoothing, must  
180 be made a priori. These choices can be especially problematic in global seismic tomography, where  
181 an uneven distribution of sources and receivers causes uneven ray coverage. Also, the use of a fixed  
182 regular grid of some chosen size typically introduces artifacts in regions of low ray coverage, where  
183 the particular property being investigated is poorly constrained.

184 Here, we assume a 300-km-thick layer atop CMB to perform the inversion. The choice of 300 km is  
185 based on the average thickness of D'' (350±50 km) presented by statistical analysis of Garcia et al.  
186 (2009) and empirical examination of different D'' thicknesses in Tkalčić et al. (2015). Given the  
187 trade-off between layer thickness and the amplitude of perturbations, further advances will include  
188 the HTDBI approaches where the thickness and the amplitude of the perturbation will be treated as  
189 variables.

190 The Trans-dimensional inverse problem in tomography may be solved with the reversible jump  
191 algorithm (Bodin et al., 2012) in which the solution is represented by multi-scale irregular ensemble  
192 of Voronoi cells with variable number and size (see Supporting information, Fig. S5 for example of a  
193 Voronoi model). For a review of the trans-dimensional Bayesian inference, see a review of  
194 Sambridge et al. (2013). In this approach, the Voronoi cells dynamically change their position,  
195 values (here velocity) and number (via death and birth of the cells) during the inversion, within the  
196 context of multi-dimensional probabilistic sampling.

197 Bayes' theorem is applied to solve the inverse problem in which the solution is represented as a  
198 posterior probability distribution. The Bayes' theorem for a trans-dimensional model can be  
199 represented as:

$$200 \quad p(m, k|d) \propto p(d|m, k) \times p(m|k)p(k) \quad (1)$$

201 where  $m$  is the model and  $d$  is the observed data,  $p(d|m, k)$  is the likelihood term,  $p(m|k)$  is the  
 202 prior Probability Distribution Function (PDF) for model  $m$  given  $k$ ,  $p(k)$  is the prior on  $k$ , the  
 203 number of Voronoi cells, typically chosen to be uniform between chosen limits.

204 The Likelihood term includes the forward modeling, where we make predictions of travel times for a  
 205 given Earth model. Here, the forward model involves computation of a travel time prediction along  
 206 each ray path in D'' using the tau-p algorithm through the PREM Earth model (Dziewonski and  
 207 Anderson, 1981). The accrued time differences between the 1D and 2D ray-tracing are on the same  
 208 order of magnitude as the uncertainty introduced by ray tracing through various mantle tomographic  
 209 models (up to several seconds for most extreme velocity perturbations). Still, moving from 1D to 2D  
 210 or complete 3D ray tracing is an ambitious computational goal within the HTDBI framework, which  
 211 hopefully the next generation of models will achieve.

212 The downward and upward ray-legs of ScS in D'' layer with 300 km thickness are projected along  
 213 with the longitude and latitude coordinates on the surface of a sphere. The predicted travel times of  
 214 ScS are attributed to the surface projection of ScS ray path on the D'' spherical shell. The prediction  
 215 is then compared to the observation. We assume travel time errors are independently and normally  
 216 distributed which leads to a Gaussian likelihood of the following form:

$$217 \quad L = p(d|m) = \frac{1}{\sqrt{2\pi} \prod_{i=1}^N \sigma_i} e^{-\sum_{i=1}^N \frac{(G(m)_i - d_i)^2}{2\sigma_i^2}} \quad (2)$$

218 where  $G(m)$  is data predicted from model  $m$ ,  $d_i$  is observation,  $N$  is the total number of observations  
 219 and  $\sigma_i$  is the error of the observation. The negative log-likelihood,  $-\log(L)$ , is a measure of data  
 220 misfit used by the algorithm. Minimizing the negative log-likelihood is equivalent to maximizing the  
 221 posterior probability.

222 In the hierarchical approach, the data noise,  $\sigma_i$ , is assumed a fixed parameter and a scaling  
 223 hierarchical parameter,  $\lambda$ , is allowed to vary. Together, they define the noise parameter,  $\sigma_{N,i}$ :

$$224 \quad \sigma_{N,i} = \sigma_i \cdot \lambda \quad (3)$$

225 A Bayesian setting is adopted to solve the inverse problem in which the solution is represented as a  
 226 posterior probability distribution over a multi-dimensional Earth model,  $m$  (the first term in eq. 1).  
 227 An ensemble of models generated with the Birth-Death Markov chain Monte Carlo, MCMC, (Geyer

228 and Moyer, 1994), which is a special case of the generalized reversible jump algorithm (Green,  
229 1995).

230 The McMC algorithm probabilistically samples Earth models, with variable numbers of unknowns,  
231 in proportion to their support as expressed by the combination of the likelihood and prior. The  
232 property known as natural parsimony ensures more complex, higher  $k$  models, which, if optimized,  
233 would fit the data better, but are not necessarily favored over less complex, lower  $k$ , models  
234 (Mackay, 2003). In the McMC algorithm, new models are proposed according to some chosen  
235 proposal distribution, and these get accepted or rejected probabilistically according to an acceptance  
236 ratio. Typically a Gaussian proposal PDF width for velocity perturbations is adjusted so that actual  
237 acceptance levels are between 30 – 50% (Tarantola, 2005). If acceptance levels are much lower than  
238 this then the algorithm spends a large proportion of time-solving the forward problem for models that  
239 are ultimately rejected. This is often the case when the proposal step size is too large. If acceptance  
240 levels much higher than this range are obtained, the algorithm is likely inefficiently walking through  
241 model space taking small steps and thereby sampling many similar models. The acceptance rate  
242 between 30 – 50% is where the random walk establishes a balance between large and small steps.

243 By definition, each new model (or step) along a Markov chain is dependent on the previous step.  
244 Since each chain is typically initiated at an arbitrary point in model space, which may be poorly  
245 supported by the data, (i.e. having low Likelihood) then early steps along the chain are discarded as  
246 ‘burn-in’. The number of models discarded is usually decided to be once the Likelihood reaches  
247 some stabilized level, or when models provide a satisfactory fit to data given estimated noise levels.  
248 In addition, the chain is ‘thinned’ in order to reduce the correlation of nearby samples along the  
249 chain. In this procedure, only every  $n^{th}$  model is taken into account. The collected output samples  
250 from all chains after burn-in and thinning is referred to as the ‘ensemble’. Statistical properties of  
251 this ensemble, such as mean and standard deviation, can then be calculated to aid the interpretation  
252 of the results.

253 In this study, we use a recently developed extension in which the model space is defined by a  
254 network of spherical Voronoi cells forming an irregular mesh over the 2D sphere (Pejić et al., 2019;  
255 Burdick et al., 2019). Spherical Voronoi cells directly partition a spherical surface, which is critical  
256 for global tomographic studies. We view this as an improvement over previously employed Cartesian  
257 Voronoi cells parameterizations (e.g. Young et al, 2013, Tkalčić et al., 2015) defined through a  
258 Cartesian projection over the sphere. The implementation of spherical Voronoi cells removes

259 distortion at the poles inherent with the Cartesian approach, and also naturally accommodates  
260 cyclicity over the globe.



#### 261 **4. Resolution tests**

262 We carried out two all-inclusive synthetic tests (Fig. 4), two synthetic tests with simpler shapes and a  
263 synthetic test including a model with smooth variations (Supporting Information, Fig. S4) to assess  
264 the resolving power of our ray coverage (Fig. 2) and the capabilities of the HTDBI method. We used  
265 the data predicted by the two models featured in Fig. 4 and added Gaussian noise with the standard  
266 deviation of  $\sigma_N=0.5$  s to the travel times. Velocity perturbations, relative to an average PREM  
267 velocity in the lowermost 300 km of the mantle, were allowed to range in the interval  $\pm 4\%$ .

268 Our synthetic tests are improvements over the checkerboard pattern tests as they check the recovery  
269 of coexisting heterogeneities of different sizes and shapes. They examine retrieval of both simple and  
270 more complex heterogeneous structures with sharp boundaries as well as varying strengths in S-wave  
271 speed in D''. The first synthetic test (Fig. 4a) examines recovery of a pseudo, complex heterogeneity  
272 in D'' dominated by the spherical degree 2 anomalies (LLSVPs) as revealed in numerous long-  
273 period S-wave tomographic studies, but assumes other, smaller-scale features of varying amplitudes  
274 superimposed on the main signature. For example, it includes fast and slow velocities outside of the  
275 LLSVPs, beneath Central America, Alaska, Greenland, East Asia, South-East Asia and Australia.

276 The second synthetic test (Fig. 4b&c) includes the maximum power amplitudes and explores the  
277 recovery of several multi-scale features of different shapes (circle, rhombus, square and a small  
278 circle) that are superimposed on top of a predominantly hemispherical structure with sharp  
279 boundaries (Young et al., 2013; Tkalčić et al., 2015). The size of the overlaying anomalies onto the  
280 hemispherical structures varies from 5000 km (the diameter of the big circles) to 300 km (the narrow  
281 parts of the squares). We show the mean and standard deviation of the posterior ensemble in the  
282 second and third row of Fig. 4, respectively. The recovery of both long- and short-scale features in  
283 the synthetic tests highlights the benefits of the HTDBI approach. As expected, the best model-  
284 recovery is achieved in regions with a dense ray coverage, in particular North Africa, North Atlantic,  
285 Asia, and Australia.

286 In our synthetic tests, we used 6 Markov chains with 3 million iterations in each chain. The first  
287 200,000 iterations were discarded as burn-in period and every 250<sup>th</sup> model was selected during the  
288 thinning procedure. We let the number of cells vary from 50 to 2000. Since the number of the model  
289 parameters can change, the required number of cells to fit the data can be expressed as a posterior  
290 probability distribution function. The histograms for the number of cells for both synthetic tests are  
291 shown in Fig. 5. In these synthetic examples, the true noise is known and there is no need to use

292 hierarchical sampling to estimate the character and level of the noise. However, adding the noise as a  
293 free parameter gives us confidence that we are able to recover noisy data in the inversion (Bodin et  
294 al., 2012). Therefore, as in a hierarchical approach, we allowed the scaling factor,  $\lambda$ , to be estimated  
295 through the inversion in all of our synthetic tests. As we added Gaussian random noise to the  
296 synthetic travel times and introduced the same value for  $\sigma$  with a standard deviation of 0.5, we  
297 expect  $\lambda$  to arrive at a value around 1 (equation 3). The histogram of  $\lambda$  (Fig. 5b & 5d) confirms that  
298 the data noise is properly resolved.

299 From synthetic test #1, it is evident that the method is able to retrieve the pattern and strength of the  
300 model parameters with a good resolution in areas with reasonable ray coverage. The ensemble mean  
301 of the retrieved models from inversion of the synthetic travel times illustrates that the combination of  
302 our ray coverage and this inversion scheme is successful in recovering both long- and short-scale  
303 features. According to the synthetic tests, small and large structures are well resolved at most parts of  
304 the northern hemisphere and Australia. We should expect smearing and distortions to happen in the  
305 east Pacific and parts of the southern hemisphere which is also evident in the standard deviation  
306 maps (Fig. 4g, 4h & 4i), and this is in agreement with the ray coverage (Fig. 2). The higher standard  
307 deviation occurs in the poorly sampled regions and at the sharp boundaries of positive and negative  
308 anomalies, which has been shown to be associated with uncertainty in boundary locations (Hawkins  
309 et al., 2019). The northern hemisphere is characterized by a higher ray coverage which results in a  
310 lower standard deviation and a good capacity to recover the details of S-wave speed variations. In  
311 contrast, the poorly sampled areas or the sharp velocity transitions go through more versatility in  
312 parameterization (mobile geometry of the Voronoi cells) and hence have a higher standard deviation.

## 313 **5. Inversion**

314 We used our high-quality, hand-picked measurements collected through waveform correlation,  
315 corrected the differential travel times for the effect of mantle structure and simulated the shear wave  
316 structure in the D''. In the HTDBI approach often a number of chains are utilized to thoroughly  
317 sample the model space. Here, at each chain, the initial 200,000 burn-in iterations were discarded  
318 and every 250<sup>th</sup> model was selected for further processing (similar to the synthetic tests). Note that  
319 these chains start at a different random point and sample the model space independently. We allowed  
320 between 50 and 2000 cells with a uniform prior PDF over the globe, and set the shear wave velocity  
321 perturbation with a uniform prior in the interval  $\langle -7.5, 5\% \rangle$  relative to an average PREM shear-wave  
322 speed in D''. A uniform prior PDF for the model parameters means that the possible outcome could  
323 be with equal probability within the prior bound but not outside. The hierarchical parameter,  $\lambda$ , was

324 set over the range  $<1.0e-9, 1.0e1>$  and is sampled from the Jeffreys prior distribution. Bounds of the  
325 priors are set to include all possible values that a model parameter can take.

326 The determining factors for the resulting number of Voronoi cells and their spatial distribution are  
327 the ray distribution and the noise amplitude. Larger Voronoi cells form in the regions of scattered  
328 sampling while much smaller ones form in the regions of denser sampling (Bodin and Sambridge,  
329 2009). We performed the inversion hierarchically, meaning that the noise is treated as a free  
330 parameter in the inversion (i.e. the inversion attempts to explore the trade-off between the model  
331 complexity (the number of Voronoi cells) and the observational noise (hierarchical error)).

332 Our LM-HTDBI approach generates a total of 54 million models. Given our burn-in period and the  
333 thinning stage, the posterior inference ends up with an ensemble of 180,000 models. Each model of  
334 the ensemble is a crude and implausible representation of the  $D''$ , however, the mean of many such  
335 crude models represents a more interpretable field of shear-wave velocity variation in  $D''$ .

## 336 **6. Results**

337 The mean and standard deviation of the posterior ensemble solution can be expressed visually by  
338 plotting them on a pixelated 2-D velocity map. To suppress the effect of the mantle on our differential  
339 travel times, we have tested a number of mantle models and analyzed the corresponding corrections  
340 for 3-D mantle structure. The structure of the mantle has been imaged by different research groups  
341 and some differences can be observed among them, which makes the decision of choosing a mantle  
342 model non-trivial. Here we have taken a more general approach, explained below.

### 343 **6.1 Mantle correction**

344 Given the variability among global mantle models and to avoid possible bias towards a specific one,  
345 we corrected our differential travel times using five different mantle models including TX2011  
346 (Grand, 2002), GyPSuM (Simmons et al., 2010), SAW642ANb (Panning et al., 2010), S362ani+M  
347 (Moulik & Ekström 2014) and SEMUCB-WM1 (French & Romanowicz 2014). We ran nine Markov  
348 chains for each corrected dataset, resulting in 45 chains with 1.2 million model-iterations performed  
349 under each chain. The resulting  $D''$  models inferred from data corrected for these five models are  
350 presented in Fig. 6. The patchy distribution of high- and low- velocity structure is independent of our  
351 choice of mantle model and shows more or less a consistent image of  $D''$  (Fig. 6a to 6e). The short-  
352 scale heterogeneities exist across all five models obtained from corrected data sets, which suggests

353 that these patchy anomalies are a feature of the Earth's structure at D'' rather than a bias towards a  
354 specific mantle model.

355 We show the chain history of negative log-likelihood as a function of iterations for 45 Markov chains  
356 in Fig. 7. The chain history of negative log-likelihood demonstrates that all Markov chains stabilize  
357 around a similar value (left panel in Fig. 7). A significant overlap can be observed in the histograms  
358 of different chains, which demonstrates that all chains converged to a similar set of models (right  
359 panel in Fig. 7).

360 Posterior probability plots of the data noise and the number of cells show the estimated noise in the  
361 dataset and the number of Voronoi cells to which the inversion converges (Fig. 8a & 8b). The  
362 frequency histogram of the number of cells shows a Gaussian distribution with a maximum centered  
363 on 600 cells. Fig. 8c shows three distinctive groups representing the trans-D ensembles recovered  
364 from the five different data corrections used. They are not due to different chains converging to  
365 different parts of the model space. Hence in this figure, we see the effect of the change in data on the  
366 ensemble. The key things to note are that the effect of the change in model correction on the  
367 distribution of cell numbers is not significant since all are centered around 600-650 cells, and most  
368 have similar spreads. The hierarchical noise parameter,  $\lambda$ , sees the most change, which again is  
369 understandable since each correction results in different levels of scatter in the observations and  
370 hence differing levels of data consistency (or noise). A sampling of the  $\lambda$  parameter responds  
371 precisely to this effect. It will be higher when the (uncorrelated) residual scatter is larger and smaller  
372 otherwise. Also, the slope of each cluster in terms of the  $\lambda$  vs. the number of cells is also as expected,  
373 i.e., when  $\lambda$  becomes larger (data noise is assumed larger), then fewer numbers of cells tend to be  
374 introduced, and vice versa. The Natural Parsimony principle can be seen in action on the distribution  
375 of cells as fewer models with a higher number of cells are sampled. The natural parsimony is also  
376 evident in Fig. 8c, since there is a negative correlation between data noise level and numbers of cells  
377 introduced by the algorithm. This is consistent with the fact that more cells, and hence structure, are  
378 required when there is more apparent signal in the data (noise levels are lower). Given that the  
379 majority of the models recovered in the inversion have around 600 cells, this signifies a well-  
380 sampled parameter space. Also, given that the inversion was performed with data noise as a free  
381 parameter, we show the frequency histogram for the scaling factor of the hierarchical parameter,  $\lambda$ .  
382 Since we assumed a  $\sigma = 5\%$ , and given the peak value of 0.27 in the distribution of  $\lambda$ , the estimated  
383 noise in HTDBI is about 1.35% ( $\sigma_N$ , in equation 3). Given this estimation for data noise, and

384 assuming a typical travel time of 170 seconds for ScS within D'', we arrive at a representative  
385 traveltimes uncertainty of about 2.3 s for our dataset.

## 386 **6.2 Shear-wave velocity in D''**

387 We combine the models with different mantle corrections (Fig. 6a-e) to obtain the mean and standard  
388 deviation over all models with the assumption that they all have equal weight. Therefore, our S-wave  
389 velocity variations in the D'' is the average of five models. This average velocity model and its  
390 associated standard deviation map are shown in Fig. 9. We compare the travel-time residual of the  
391 mean model (Fig. 9 a) with an individual Voronoi-cell model (Supporting Information, Fig. S5) and  
392 the initial travel-time residuals in the Supporting Information, Fig. S6. The histogram shows that the  
393 mean of the posterior ensembles can reduce travel time residuals compared to a single Voronoi-cell  
394 model shown in (Fig. S5) and the average velocity in PREM.

395 Following Simons et al (2006), we filtered the tomographic image of this study along with other  
396 tomographic models: SGLOBE-rani (Chang et al., 2015), S362ANI+M (Moulik and Ekström, 2014),  
397 SP12RTS (Koelemeijer et al., 2016), GyPSuM (Simmons et al., 2010), S40RTS (Ritsema et al.  
398 2011), SEMUCB-WM1 (French & Romanowicz, 2014) and HMSL (Houser et al., 2008) to the low  
399 spherical harmonic degree,  $l_{max}=6$  (Fig. 10). The long-scale patterns in LM-HTDBI are in general  
400 agreement about low velocities beneath Africa and the Pacific and fast in most parts of Asia and  
401 Antarctica. However, there is a striking difference across the existing global models, even for long-  
402 scale features in the shape and intensity of the LLSVPs (Fig. 10). The critical characteristic of our  
403 tomographic model is the appearance of short-scale features at the base of the mantle exterior to the  
404 LLSVPs, which our resolution analysis confirms are not artifacts of the inversion. In the Supporting  
405 Information (Fig. S7), we examine the heterogeneity-spectrum or the power-spectrum of the models.  
406 Spectral analyses reveal the long-scale content of our model is similar to the previous models as  
407 shown in the filtered plots (Supporting Information, Fig.10 and Fig. S7); however, its power of  
408 heterogeneity in higher degrees is larger than previous estimates (Fig. S7).

409 The main advantage of our approach over previous approaches is that as no explicit subjective  
410 regularization or parameterization are applied, the size of velocity heterogeneities is directly  
411 controlled by the data itself and the short-scale features are preserved. In terms of a comparison with  
412 other lowermost mantle models, we confirm that our model correlates well with SEMUCB-WM1 in  
413 terms of short-scale heterogeneity content, but it has more intense variation in velocity in comparison  
414 with the latter. The areas with the lowest and the highest shear-wave speed are scattered throughout

415 the lowermost mantle. Four specific examples of regions with high-velocity are mid-Indian ocean,  
416 west Australia, North Atlantic and North Chile. From Fig. 8b, it is obvious that the areas with higher  
417 standard deviation correspond to either the poorly sampled parts of D'' or to the areas that have gone  
418 through more velocity variations during the inversion (e.g. closer to the sharp contrasts in elastic  
419 properties).

## 420 **7. Discussions**

421 The chief goal of this study was to invert for the shear-wave velocity heterogeneities atop the CMB  
422 using ScS and S travel times. Our study has a couple of limitations. The first limitation is that we  
423 assume a fixed thickness of 300 km for the D'' layer. In addition, our forward method does not  
424 include the finite frequency effects. However, ray-theory based tomography at the frequencies we  
425 use is a reasonable approximation to interpret finite frequency travel times. For example, Hung et al.  
426 (2005) and Castle et al. (2000) used ScS-S data at the periods of 20 s to perform inversions with and  
427 without the sensitivity kernels. They showed that the tomography based on the sensitivity kernels  
428 yields similar results using infinite frequency ray paths. Hung et al. (2005) showed that the recovered  
429 velocity perturbations using finite frequency kernels (in the frequency band similar to our data) are  
430 approximately 1–2 times larger than those from the ray-based kernels. In our study, given the  
431 average number of 600 Voronoi cells for the area covering the CMB (Fig. 8a), the average diameter  
432 of a simplified Voronoi cell is  $\sim 500$  km, which is slightly larger than the diameter of the Fresnel  
433 zone (400 km). The structures we interpret are also larger than the size of the Fresnel zone. Hence,  
434 not accounting for finite frequency effects might not be a too severe limitation.

435 Another limitation of our approach is that the inversion scheme we deploy does not assess an  
436 independent impact of anisotropy on S and ScS data, and in turn, on our tomograms. Within the  
437 Bayesian framework, not accounting for anisotropy is considered a theory error. Accounting for both  
438 data noise and theory errors is at the leading edge of geophysical inference studies. Accurate  
439 mapping seismic anisotropy in the lowermost mantle using tomographic methods is challenging and  
440 still an open question (Romanowicz & Wenk, 2017). Many studies used differential measurements of  
441 pair phases such as ScS and S to calculate anisotropy in D'' (e.g., Wookey et al., 2005a; Nowacki et  
442 al., 2010; Lay, 2015). Shear wave splitting measurements of D'' show that horizontally polarized S-  
443 wave velocities are larger than vertically polarized ones by  $\sim 1\%$  (Panning & Romanowicz, 2004).  
444 The residual splitting in ScS, which is attributed to the anisotropy of D'' shows lag times between  
445 1.0 and 3.9 s for epicentral distances  $60^\circ$  to  $85^\circ$  beneath the north Pacific (Wookey et al., 2005).

446 Our tomographic approach in a combination with a high-quality dataset enables a multi-scale  
447 imaging of the lowermost mantle. Our model (LM-HTDBI) reveals multi-scale structures ranging in  
448 wavelengths from hundreds to thousands of kilometers. The deduced shear-wave tomogram of D''  
449 also reveals that the amplitudes of the short-scale low-velocity anomalies outside the LLSVPs are  
450 comparable to those observed inside them. Also, the morphology of the LLSVPs imaged by our  
451 method features small high-velocity patches creating separate regions of low-velocity anomalies,  
452 rather than only two large, unified and smooth LLSVPs. In addition, the appearance of the short-  
453 scale structures distorts the general shape of the LLSVPs as their margins do not look uniformly flat.

454 The results of this study are in general agreement with the D'' P-wave tomography of Tkalčić et al.  
455 (2015) using the HTDBI method in that both models contain short-scale and medium-scale structures  
456 superimposed on top of the long-scale structures. Furthermore, Muir and Tkalčić (2020) have  
457 recently obtained a similar result using Hamiltonian Monte Carlo sampling and low-degree spherical  
458 harmonics expansion with implicit regularization by basis truncation. They showed that short-scale  
459 perturbations are required to account for the observed travel-time residuals.

460 The complex pattern of short-scale heterogeneities that we obtained suggests that LLSVPs are not  
461 necessarily the only contributing factors to mantle instabilities. Given the existence of short-scale  
462 heterogeneities, we believe that it is feasible that they play a significant role in mantle dynamics.  
463 Several regional studies found evidence for locally sharp variation (up to or more than  $\pm 4\%$ ) in  
464 regions like eastern Eurasia (He & Wen, 2011), western Pacific (e.g. He et al., 2006; Takeuchi 2012;  
465 Suzuki et al., 2020) and the Caribbean (Hung et al., 2005). The recent study of Suzuki et al., (2020)  
466 illustrated several strong short-scale, low shear-wave velocities in the western Pacific and northern  
467 Australia. For instance, in ULVZs, shear-wave velocity reduction can reach 30 – 45% (e.g. Cottaar &  
468 Romanowicz 2012; Yu and Garnero 2018; Jenkins et al., 2021) and many of them show complex  
469 internal radial structures (Pachhai et al., 2014; 2015). We cannot expect to reveal such a level of  
470 details in our tomograms as the thickness of ULVZs varies most likely between 5 and 10 km  
471 (Garnero et al., 2016) or on the order of 20 – 30 km in the case of the Hawaiian mega-ULVZ  
472 (Cottaar and Romanowicz 2012; Jenkins et al., 2021). Our model presents an averaged image of D''  
473 with approximately 300 km thickness. Hence, some theory limitations inevitably contribute to the  
474 noise and are captured by our estimates of observational uncertainty through the hierarchical noise  
475 estimation. A recent study by Kim et al. (2020) found evidence of pervasive scattering (Sdiff  
476 postcursors present on  $\sim 50\%$  of paths) across the Pacific without a clear relationship between the  
477 postcursors' amplitudes and delay-times, suggesting either distributed heterogeneities or interaction

478 with boundaries of a larger structure. We speculate that the significant amplitude  $V_s$  variation  
479 obtained in this study is probably significant enough to produce the postcursors detected by Kim et  
480 al. (2020).

481 Our study reveals short-scale heterogeneities on top of the long-scale structures while most global  
482 models imaged predominantly long-scale features in  $D''$ . The method adopted here alleviates some  
483 of the shortcomings of using classical inversion methods including block parameterization,  
484 smoothing and damping regularization. Importantly, inverting independently for ScS-S and  
485 correcting for the mantle heterogeneities enables us to isolate the lowermost mantle from the  
486 heterogeneities in the rest of the mantle. To further investigate these differences, we compared seven  
487 different global maps of shear wave variations at  $D''$  including SGLOBE-rani (Chang et al., 2015),  
488 S362ANI+M (Moulik and Ekström 2014), SP12RTS (Koelemeijer et al. 2016), GyPSuM (Simmons  
489 et al., 2010), TX2011 (Grand 2002), SEMUCB-WM1 (French & Romanowicz 2014) and HMSL  
490 (Houser et al., 2008) with our model (Fig. 11). More broadly, these models were derived either from  
491 different datasets or by using different methods of inversion.

492 Fig. 10 and Fig. 11 reveal that there are similarities among different models for long-scale features:  
493 they agree on the existence and extent of large low-velocity patterns, the well-known LLSVPs, under  
494 the Pacific and Africa. However, the margins, shapes, strength of heterogeneity and gradients across  
495 the  $D''$  are significantly different. Although the data coverage is usually better in the northern  
496 hemisphere, most of the models in  $D''$  do not capture small-scale heterogeneities, which we  
497 speculate is most likely due to applying uniform damping and smoothing. For example, an  
498 interesting observation from our comparison with other global models (Fig. 11) is a clear difference  
499 in the recovery of short-scale anomalies between the HMSL model and this study. We employed a  
500 subset of data used in the HMSL model, which only includes data from larger epicentral distances.  
501 Although the outline of the LLSVPs is similar, the presence and lack of short-scale features represent  
502 a striking difference between these two models. This is attributable to the inversion strategy  
503 deployed here where no fixed model parameterization, damping and smoothing are required, and  
504 noise is treated as a free parameter, which is achieved at a significantly higher computational cost.

505 Another clear difference among the models selected for comparison is the amplitudes of  $V_s$   
506 variations. This is not surprising as different modelers use different data and methods, and they  
507 achieve different resolutions in different regions. We compared the amplitudes of LM-HTDBI with  
508 the previous models in the Supporting Information (Fig. S8), as in Burdick & Lekić (2017). The tilt  
509 of the ellipses in Fig. S8 shows that velocity variations in our model have higher amplitude relative



510 to the previous studies. Among all models presented in Fig. 11, SEMUCB-WM1 is generally  
511 consistent with our model in places where they both exhibit larger amplitudes of S-wave velocity  
512 heterogeneity, as well as in the features not seen in earlier generations of global D'' models.  
513 However, these two models certainly differ in detail, especially in terms of the location and  
514 distribution of short-scale heterogeneities. The existence of some short-scale features in SEMUCB-  
515 WM1 is likely due to the usage of short-period waveform data.

516 The only low-velocity anomaly outside of the LLSVPs that emerges in the majority of global models  
517 is the 'Perm anomaly' (e.g. Lekić et al., 2012), north of the Caspian Sea. The region around Perm  
518 anomaly is imaged as two smaller low-velocity anomalies in our model and this region is well  
519 constrained in our synthetic test (Fig. 4 d).

520 A possible origin and composition of the LLSVPs presented in global tomographic images of D'' is  
521 still enigmatic (e.g., Garnero et al., 2016; Doucet et al., 2020), yet our tomographic model confirms a  
522 significantly more complex image of D''. The base of the mantle outside of the LLSVPs has long  
523 been considered as the site of accumulation of cold subducted slabs (Grand, 2002; Jones et al., 2020;  
524 Suzuki et al., 2020; Li 2020). This is a possible explanation for the existence of high shear-wave  
525 velocity regions. A recent study by Li (2020) showed that the scattered low-velocity anomalies  
526 outside of the LLSVPs could have a hot thermal origin that are detached from the thermochemical  
527 piles. Along the same line, the existence of fast seismic velocity anomalies within LLSVPs might  
528 suggest principally thermal-heterogeneity characterized by the clusters of plumes (Davies et al.,  
529 2015; Tkalčić et al., 2015, Muir & Tkalčić, 2020, Davaille & Romanowicz 2020).

530

## 531 **8. Conclusions**

532 This study presents a new model of shear-velocity variations in the D'' using high quality,  
533 handpicked dataset of S and ScS differential travel times through waveform correlation. Inverting for  
534 the compilation of 18,131 ScS-S differential travel times in conjunction with careful consideration of  
535 global mantle models enabled us to "isolate" the lowermost mantle from the heterogeneities present  
536 in the rest of the mantle. In this paper, we assumed that the D'' could be modeled as a single layer.  
537 We utilized recent improvements in the inversion technique through the use of the HTDBI  
538 framework which is based on implicit parametrization (spherical Voronoi cells) and treats the model  
539 complexity and the data noise as free parameters, while a uniform application of damping and  
540 smoothing to the model space can be avoided. We found that our model is in agreement with most of

541 the recent global studies in terms of imaging the long-scale features at the base of the mantle.  
542 However, short-scale heterogeneity, as seen in our D'' tomogram, is far more omnipresent than in  
543 spherical-harmonic, degree-2-dominated images for shear-wave speed of the lowermost mantle  
544 obtained in most global models. This is the most striking difference between our model and the  
545 previous model obtained using the same subset of data (Houser et al., 2008). We demonstrate,  
546 however, that our method is able to recover much simpler, degree-2 images of the lowermost mantle  
547 if the lowermost mantle is indeed void of short- and medium-scale heterogeneity and only  
548 characterized by long-scale features. We are thus confident that the tomogram of the lowermost  
549 mantle we obtain is a realistic representation of the seismic structures. Significantly, the recovered  
550 LLSVP morphology of is complex, consisting of smaller high-velocity patches among low-velocity  
551 domains. The newly obtained image suggests that the LLSVPs' margins are not uniformly flat. We  
552 further argue that the methodology we adopt is a significant step forward in imaging the complicated  
553 structure of the D'' layer on a global scale, and provides an important bridge between long-scale  
554 features at a global scale and short-scale features of regional models. Future work may explore  
555 combining additional datasets of travel times to increase spatial coverage and resolution of the  
556 lowermost mantle, adopting similar principles in full-waveform modeling and novel methodological  
557 approaches to inversion.

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559 We thank Guy Masters for sharing his dataset with us, and his insights into the interpretation of our  
560 results. The authors would like to thank the associate editor Ved Lekić, Ed Garnero, and an  
561 anonymous reviewer for constructive discussions and feedback. We used GMT (Wessel et al., 1998)  
562 to plot the figures. We downloaded the data from IRIS DMC via the Wilber 3 tool  
563 (<http://ds.iris.edu/wilber3>). Calculations were performed on the Terrawulf cluster, a computational  
564 facility supported through the AuScope Ltd. AuScope is funded under the National Collaborative  
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567 NCI National Facility in Canberra, Australia, which is supported by the Australian Commonwealth  
568 Government. Generic Trans-dimensional Voronoi Cell Code for surface spherical problems is  
569 available through <https://zenodo.org/record/4771010#.YKRI7Kgzabg>.

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577 **Figure Captions**

578

579 **Fig. 1.** a) Diagram of S and ScS raypaths from hypocenter (star) to receivers (reversed triangles) at  
580 epicentral distances of 45° and 75°. b) Histogram of data distribution (40,843 rays) based on  
581 epicentral distances (Houser et al., 2008). The red dashed line indicates the portion of data (16,144  
582 rays) that was selected for this study.

583

584 **Fig. 2.** The D'' sampling by the ray coverage in this study. The green rays are a subsample of  
585 Houser et al., (2008) for epicentral distances of 60-75 deg. The blue rays are the newly-picked data  
586 which include events from 2008 to 2018 (1,787 rays). b) Distribution of ScS bounce points of the  
587 newly picked dataset on the CMB. Observed ScS-S travel time residuals relative to PREM projected  
588 at the bouncing points of ScS. c) Distribution of ScS bounce points of Houser et al., (2008) on the  
589 CMB. The gray lines beneath the bouncing points represent ray path at D''.

590

591 **Fig. 3.** An example of S and ScS alignment to measure the differential travel time of S and ScS for  
592 two events 2017-02-24T17:28:44 and 2008-02-25T21:02:19 recorded at stations GIRL, BEAR and  
593 CMSA. The measurement technique is manual fitting of the S and ScS waveforms concentrating on  
594 the first half cycle. The y-axis is normalized to the maximum amplitude.

595

596 **Fig. 4.** Two synthetic tests designed to show the resolving power of ScS in the lowermost mantle  
597 based on the ray coverage shown in Fig. 1. Top row is the input models, second one is the recovered  
598 mean model and third row is the estimated standard deviation. a) The synthetic test that mimics the  
599 D'' pattern obtained in most studies. b) A hemispherical multi-scale synthetic test centered at 0°. c)  
600 Same as b) but centered at 180°. The positive anomalies are 4% faster than the PREM average  
601 velocity and the negative anomalies are 4% slower. d), e) and f) are the recovered mean models. g),  
602 h) and i) are the standard deviation of the posterior ensemble.

603

604 **Fig. 5.** The posterior probability distributions of the number of cells and the hierarchical parameter:  
605 a) and b) for the synthetic test #1. c) and d) for the synthetic test #2.

606

607 **Fig. 6.** S-wave velocity perturbations from the PREM of D'' using mantle corrections according to  
608 the following models: a) TX2011 (Grand, 2002), b) SEMUCB-WM1 (French and Romanowicz,  
609 2014), c) GyPSuM (Simmons et al., 2010), d) SAW642ANb (Panning et al., 2010) and e)  
610 S362ANI+M (Moulik and Ekström, 2014).

611

612 **Fig. 7.** Chain history of 45 chains from five models each including 9 chains that presented in Fig. 6.  
613 Each color line represents a chain. The vertical axis is Negative-log likelihood from equation 2, and  
614 the iterations are shown in the horizontal axis. The right panel shows the histograms of the Negative-  
615 log likelihood functions across the chains. The overlapping histograms indicates that all the chains  
616 converged to a similar set of models and different corrections lead to chains with similar likelihoods.

617

618 **Fig. 8.** a) Posterior probability distributions of the number of Voronoi cells used to describe LM-  
619 HTDBI. b) Posterior probability distributions of the hierarchical parameter,  $\lambda$ , of the data noise. c)  
620 Distribution of hierarchical parameter (estimated noise in the data) versus numbers of cells. Each  
621 color represents a different chain. The trade-off is, as expected, where higher data noise ( $\lambda$ ) results in  
622 fewer unknowns, and hence less structure is introduced. This property is often termed ‘natural  
623 parsimony’.

624

625 **Fig. 9.** The average shear-wave velocity model of D’’ (LM-HTDBI) calculated by averaging over  
626 the five models presented in Fig. 6. a) S-wave velocity perturbations from the PREM model and b)  
627 the corresponding standard deviation. This average model and standard deviation is determined  
628 across multiple MCMC chains in Fig. 6, using the individual means and standard deviations in S-  
629 wave speed of each chain. By using some standard algebraic results this becomes equivalent to an  
630 equally weighted combination of the ensemble of models from all chains and all cases in Fig.6.

631

632 **Fig. 10.** Global maps of shear wave velocity variations at the core-mantle boundary for the eight  
633 models including SGLOBE-rani (Chang et al., 2015), S362ANI+M (Moulik and Ekström, 2014),  
634 SP12RTS (Koelemeijer et al., 2016), GyPSuM (Simmons et al., 2010), TX2011 (Grand, 2002),  
635 SEMUCB-WM1 (French & Romanowicz, 2014), HMSL (Houser et al., 2008) and this study (LM-  
636 HTDBI) filtered to the spherical-harmonic degree  $l_{max} = 6$ . Each model extreme value annotated on  
637 the top left side of their map. All scale bars are in  $\% \delta V_s$  and set to the  $\pm 5\%$  to show the amplitude  
638 diversity between models.

639

640 **Fig. 11.** Comparison of the global shear-wave velocity models for the lowermost mantle. All scale  
641 bars are in  $\% \delta V_s$  and set to the  $\pm 5\%$  to show the amplitude diversity between models. The  
642 tomographic models for eight models shown in Fig. 10. As in Fig. 10, the numbers in the corner of  
643 each panel show maximum variation for each filtered model and scale bars are set to  $\pm 5\%$ .

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Figure 1.

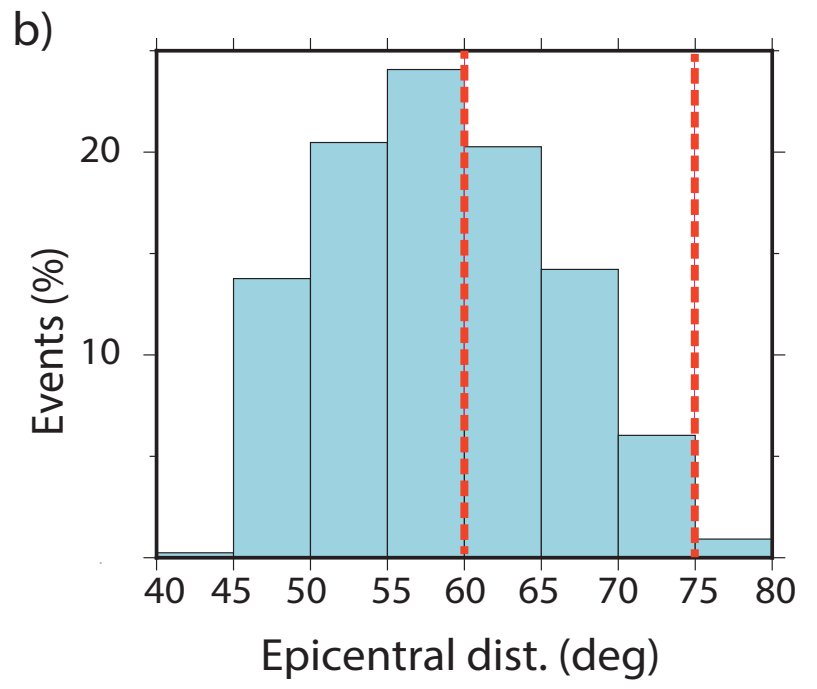
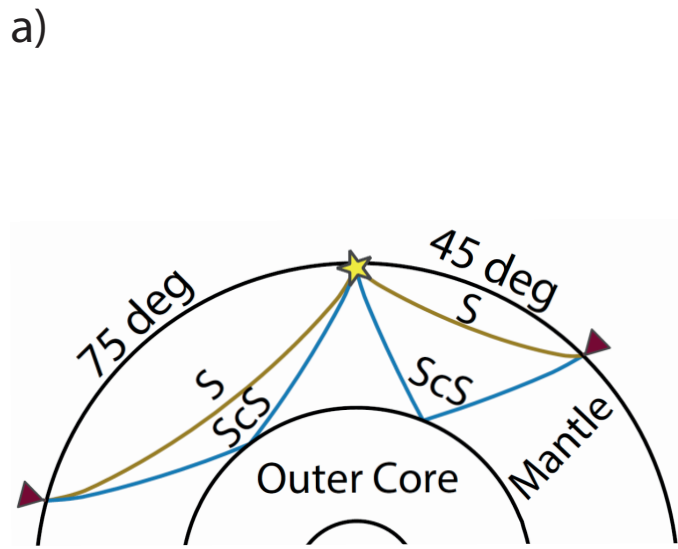


Figure 2.



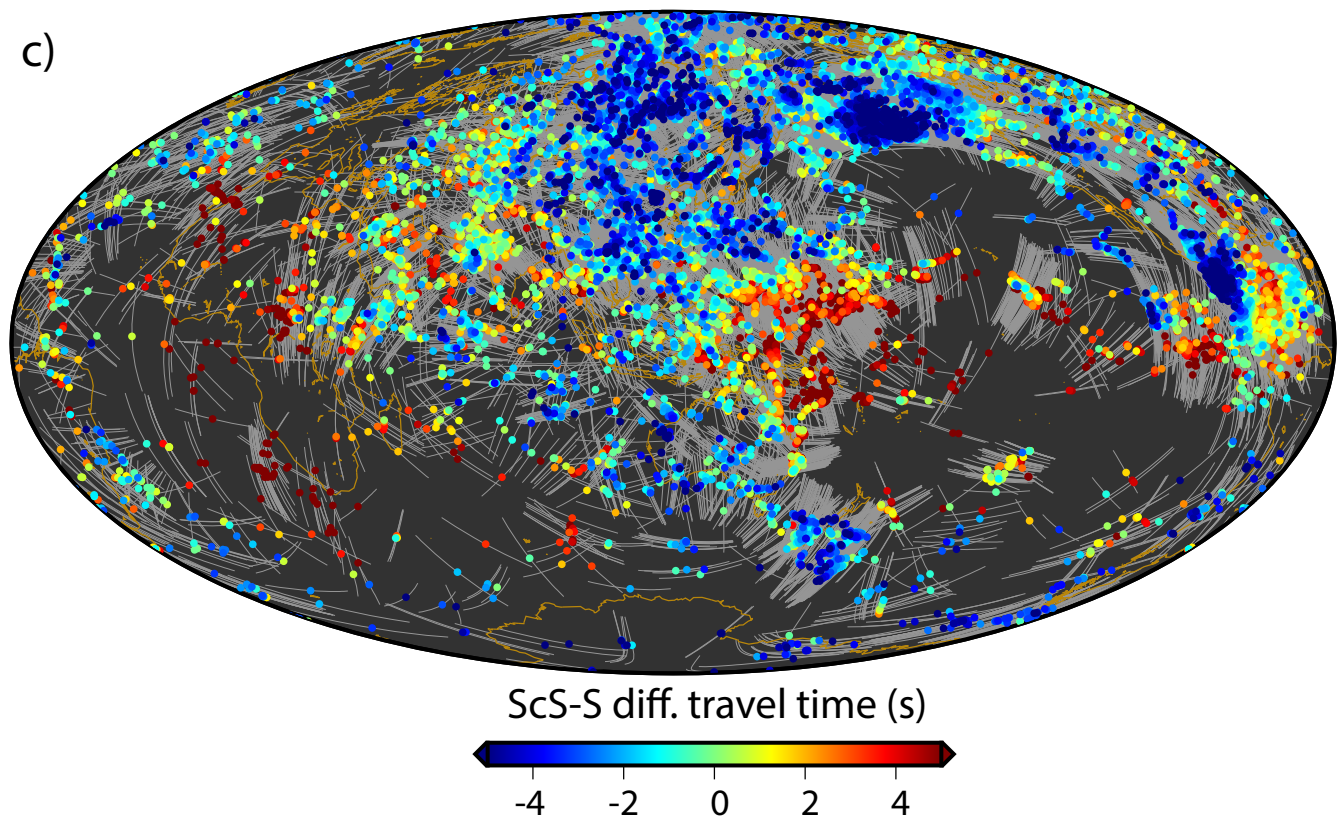
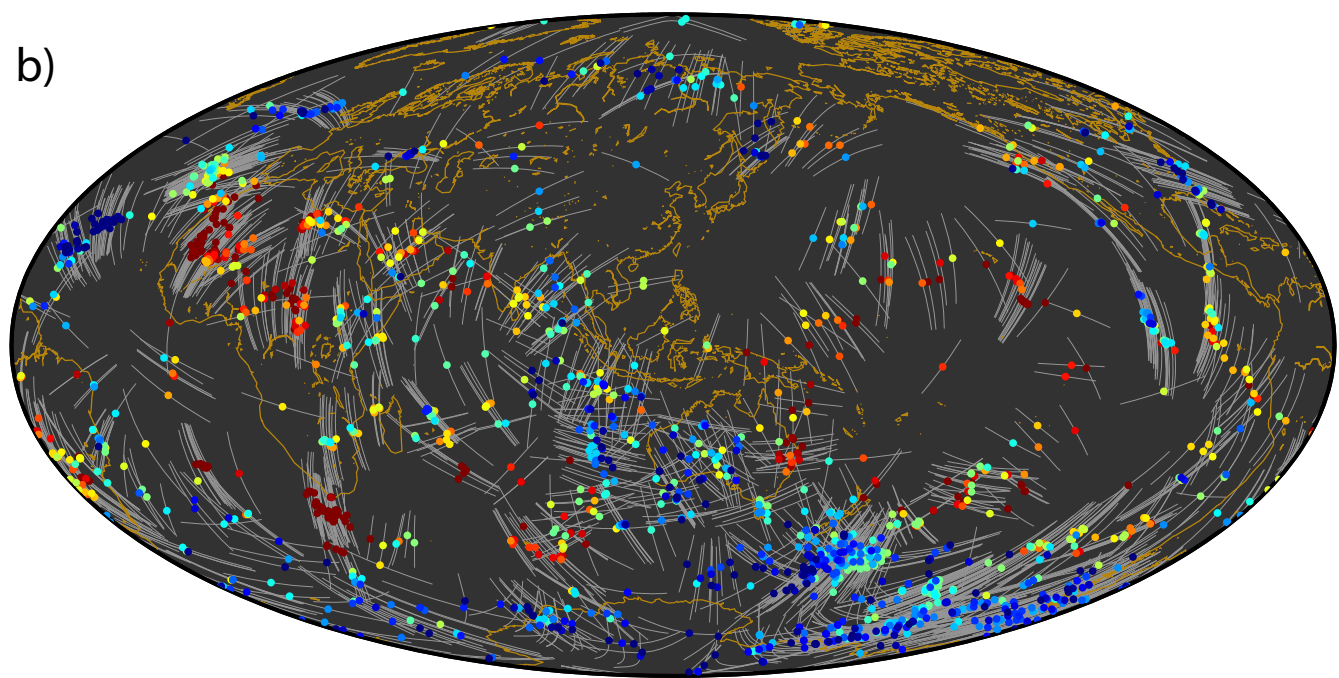
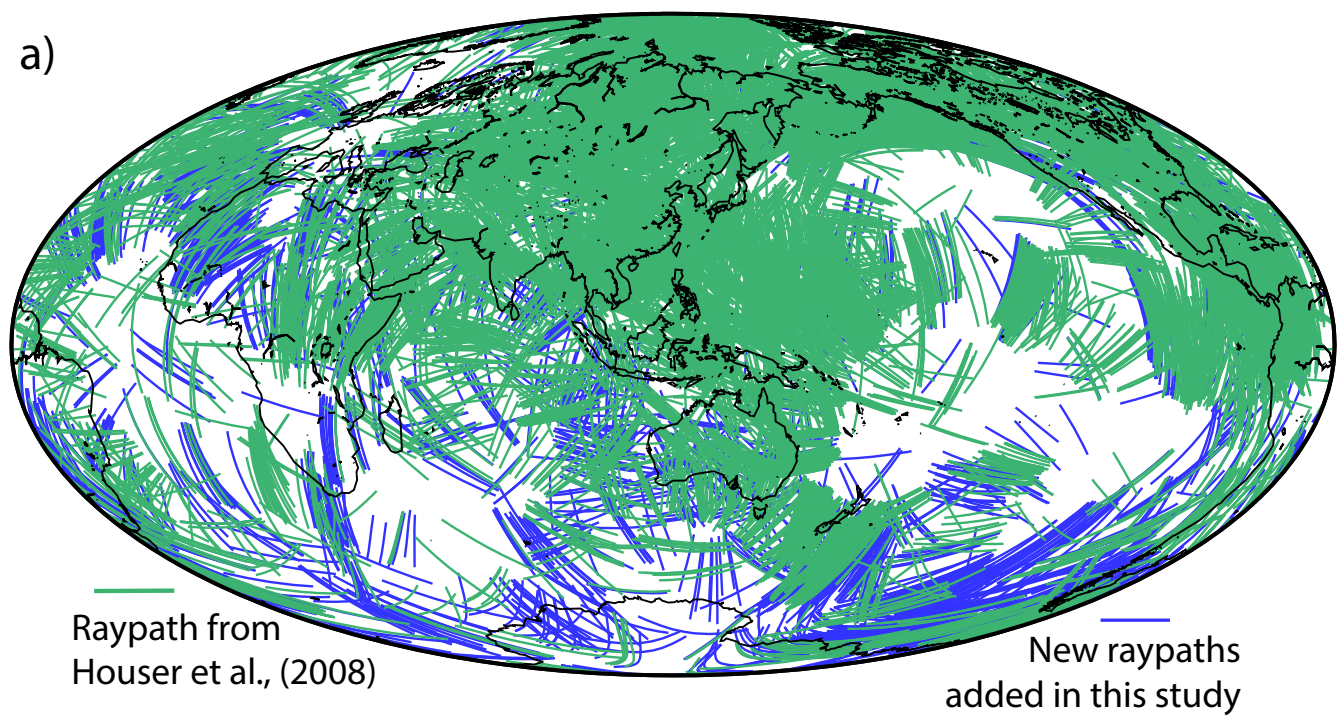
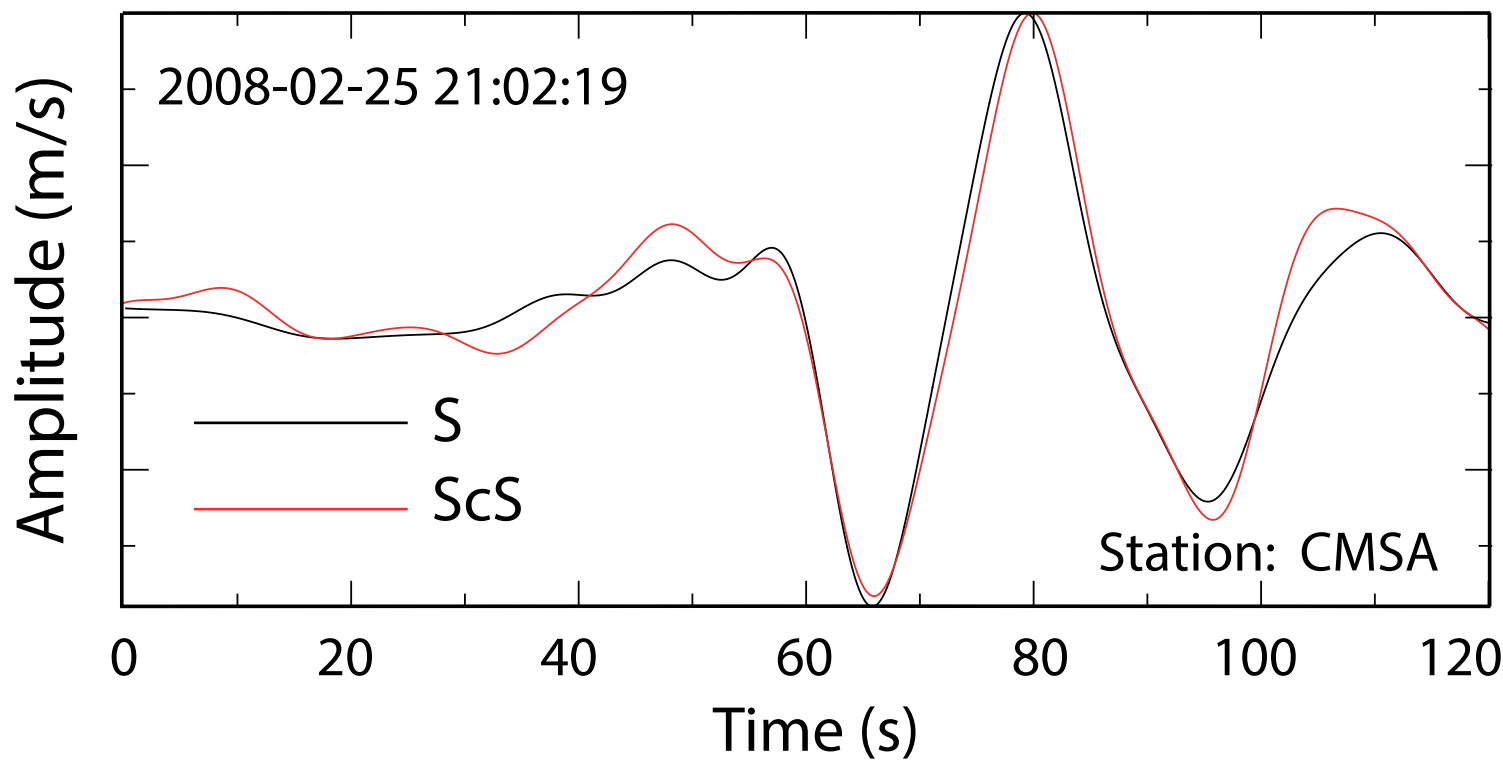
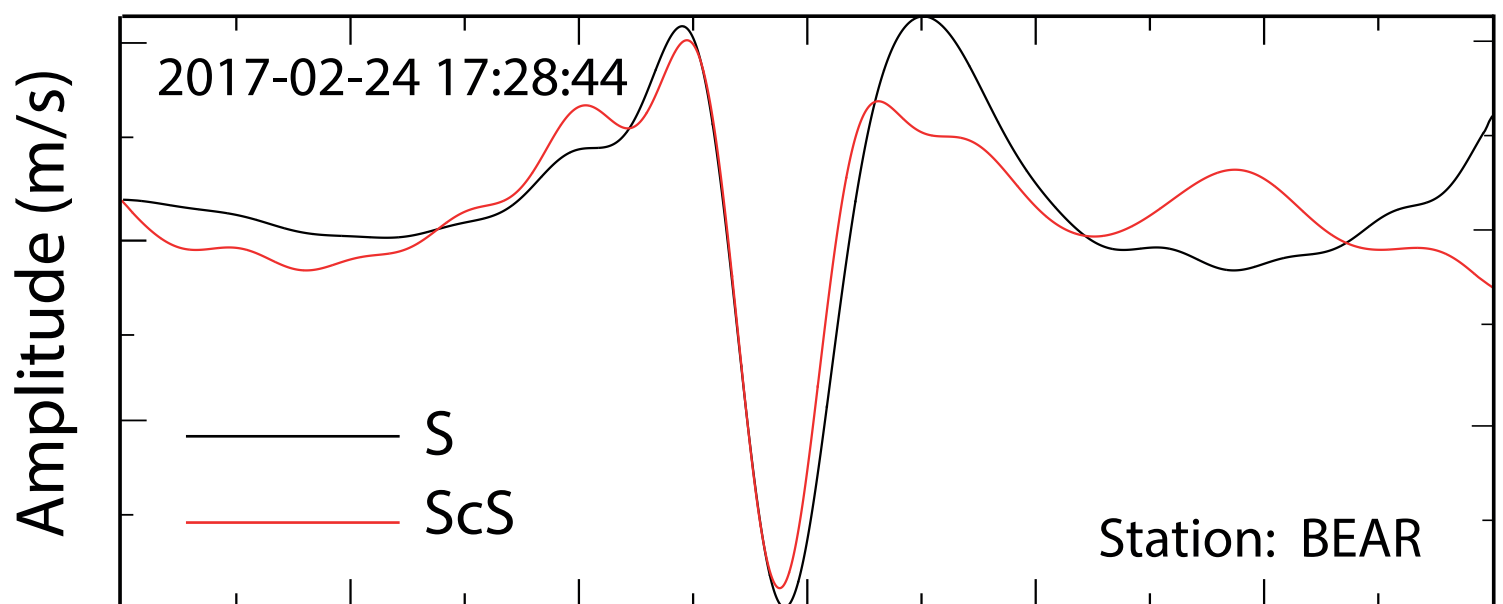
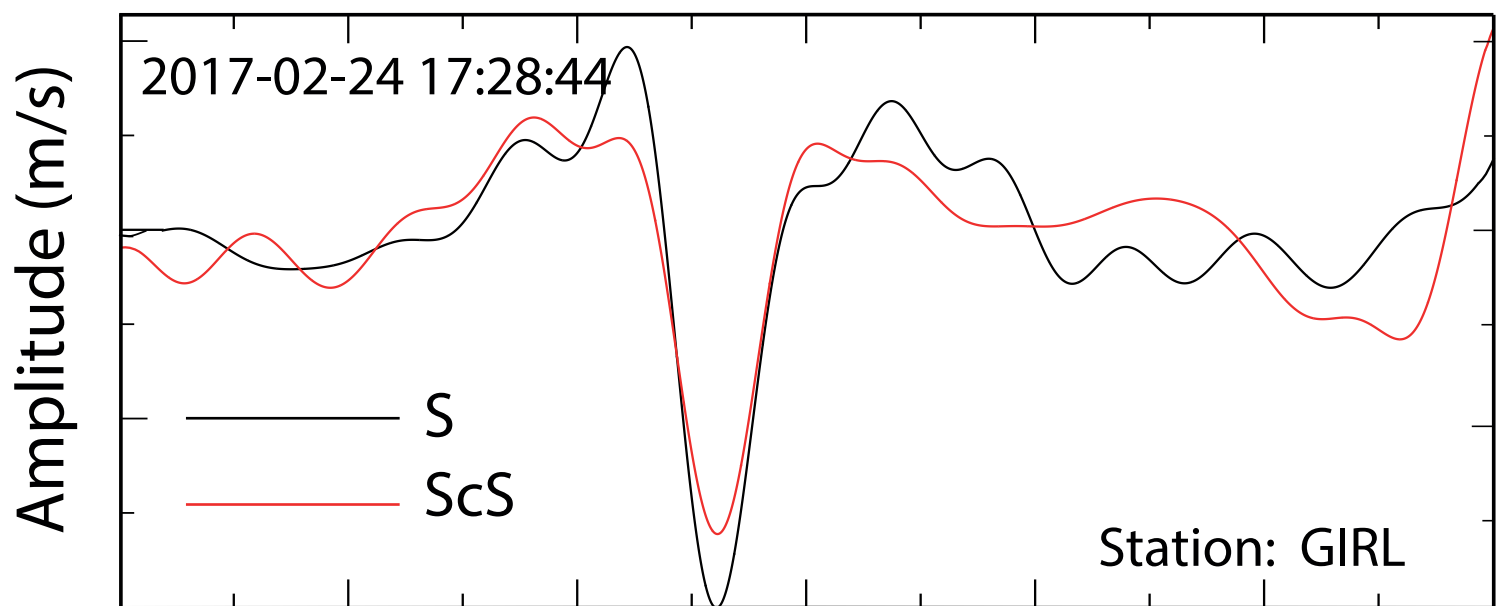


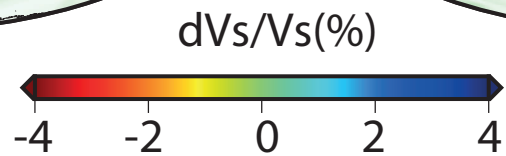
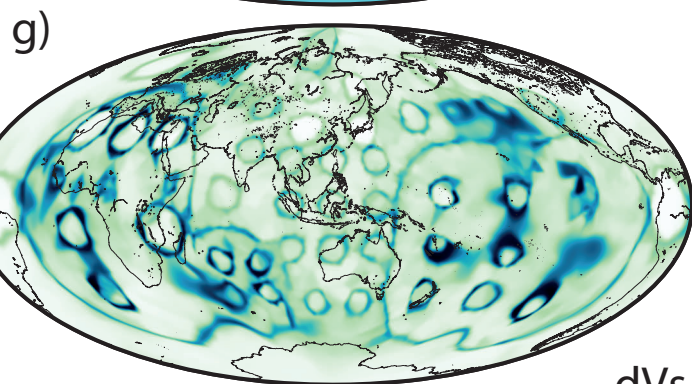
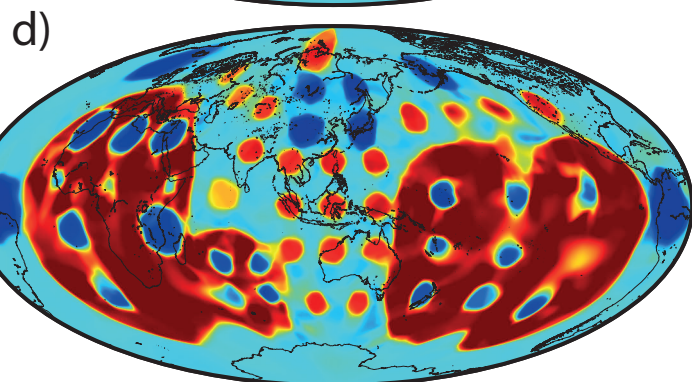
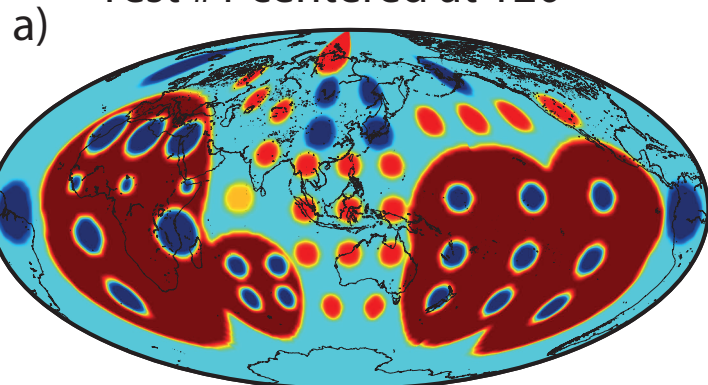
Figure 3.



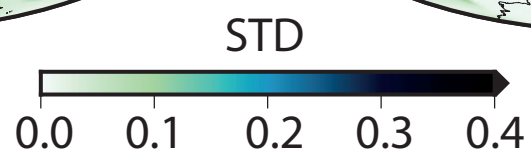
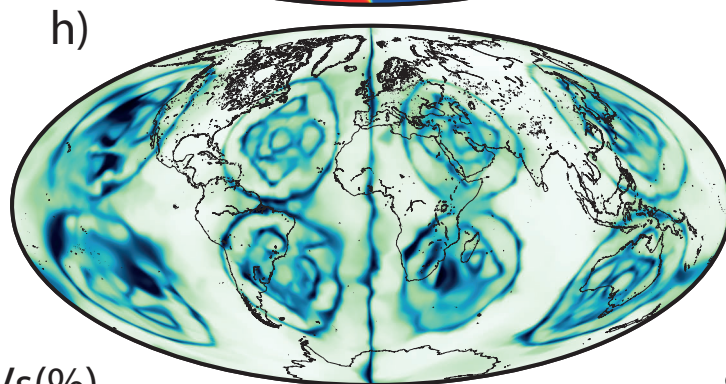
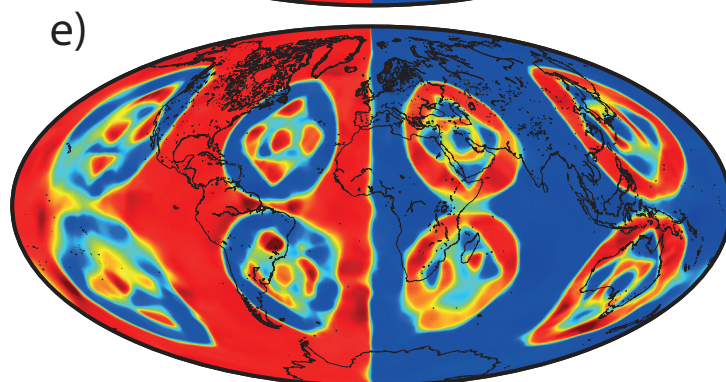
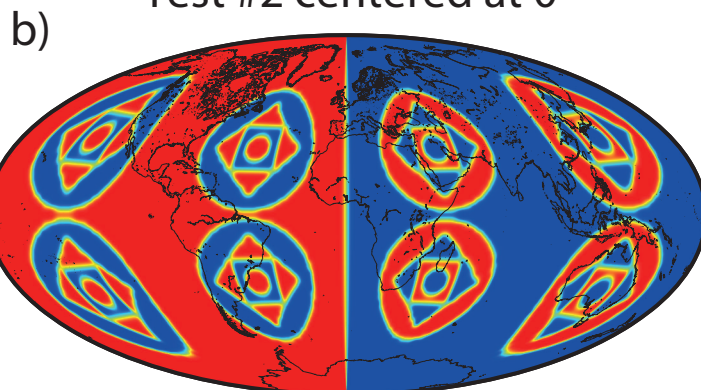
**Figure 4.**



Test #1 centered at 120°



Test #2 centered at 0°



Test #2 centered at 180°

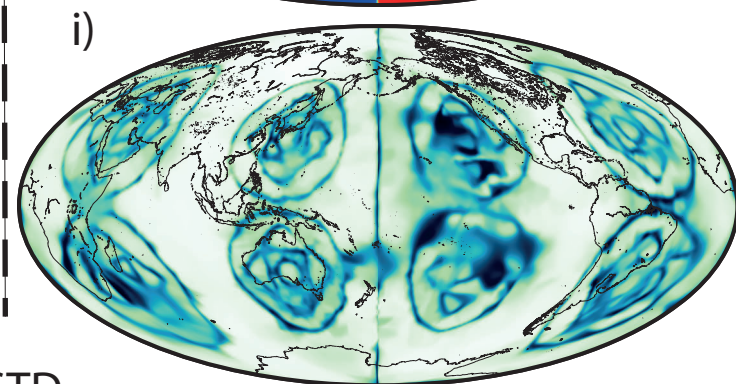
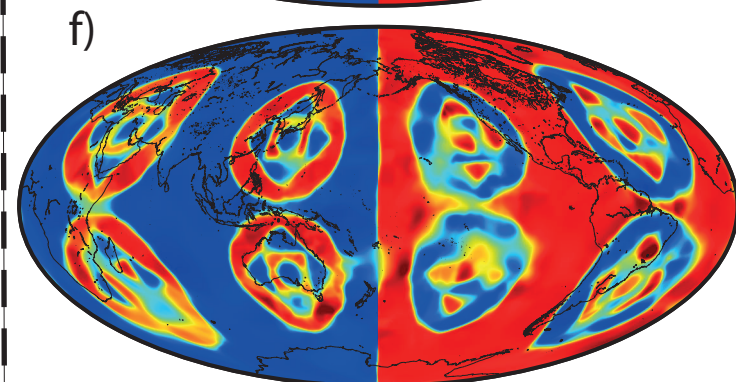
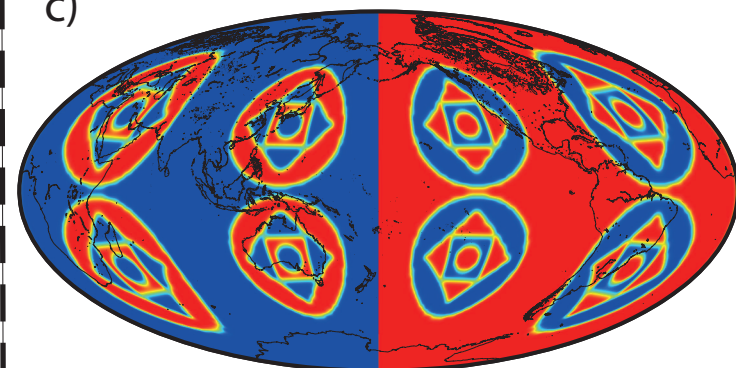


Figure 5.

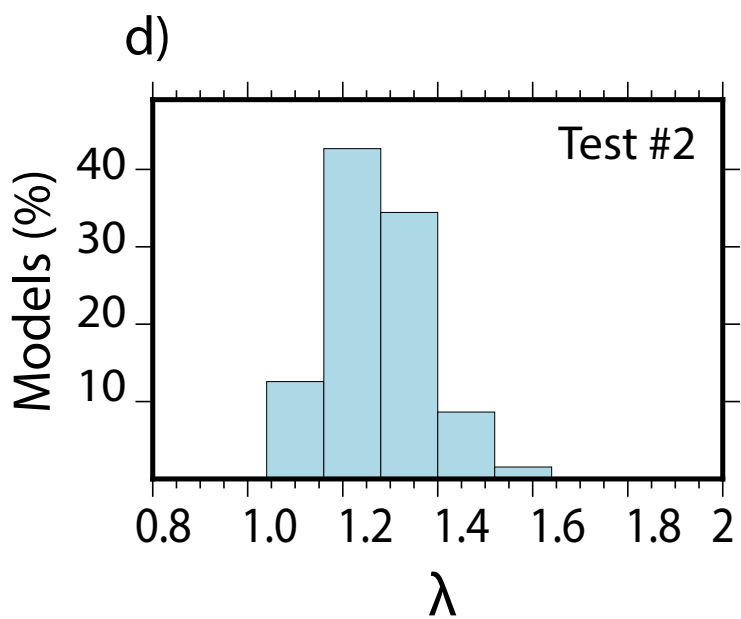
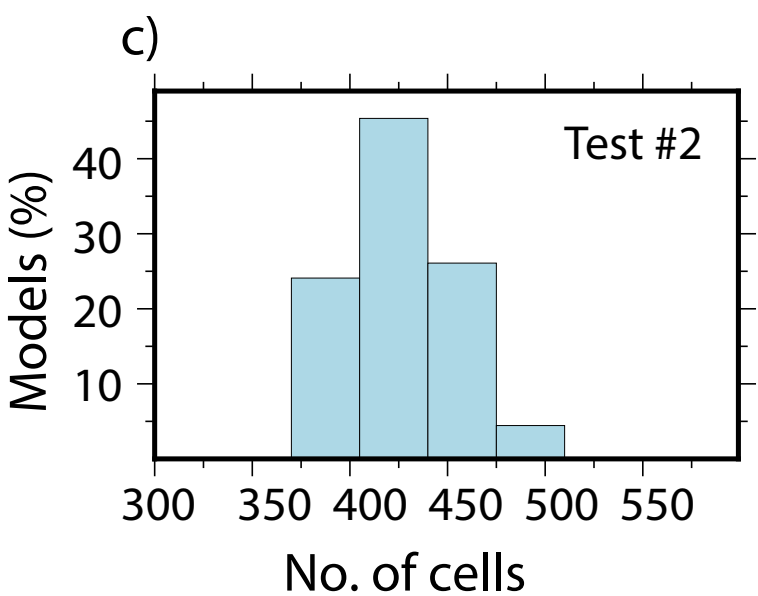
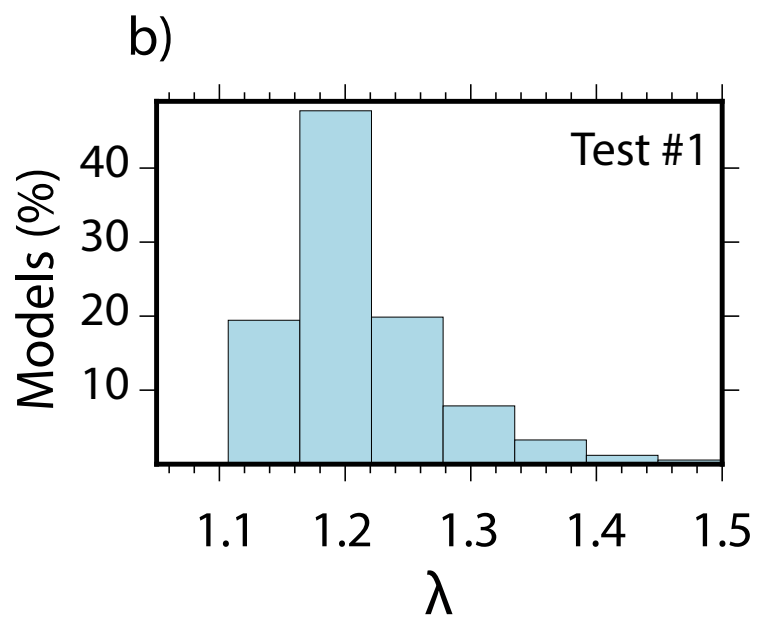
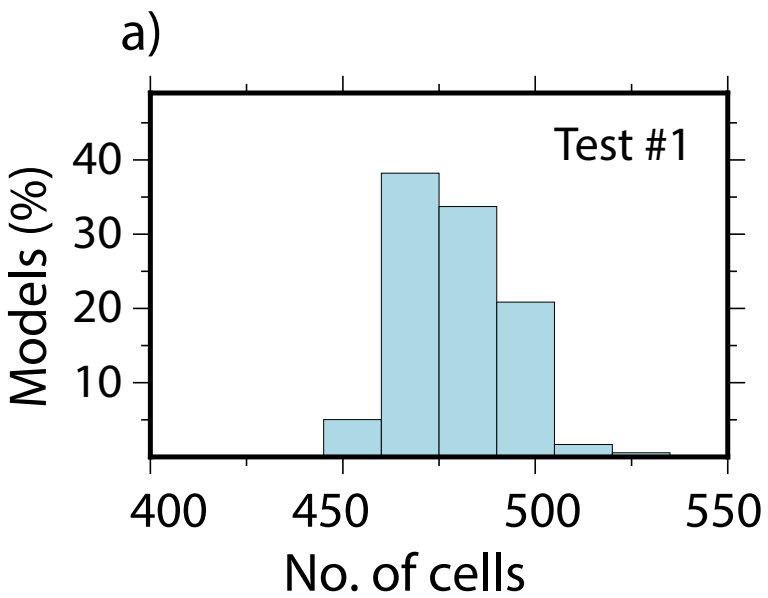
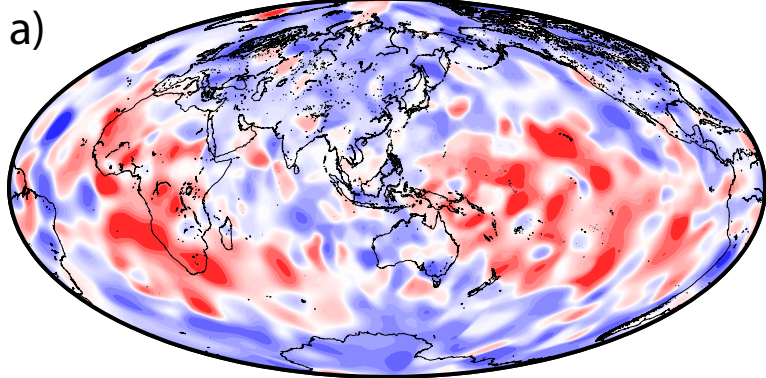


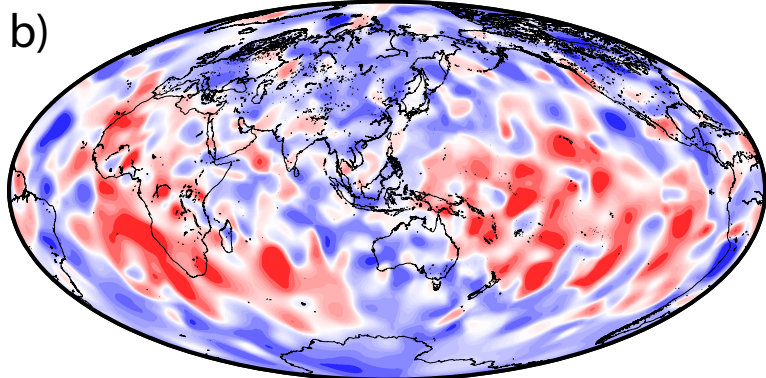
Figure 6.



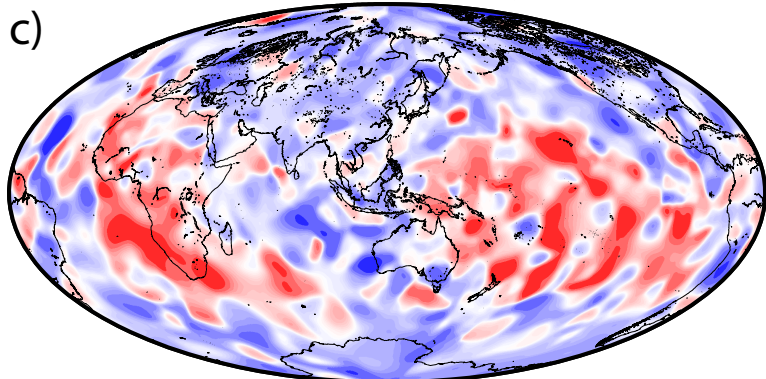
Mantle corrected using TX2011



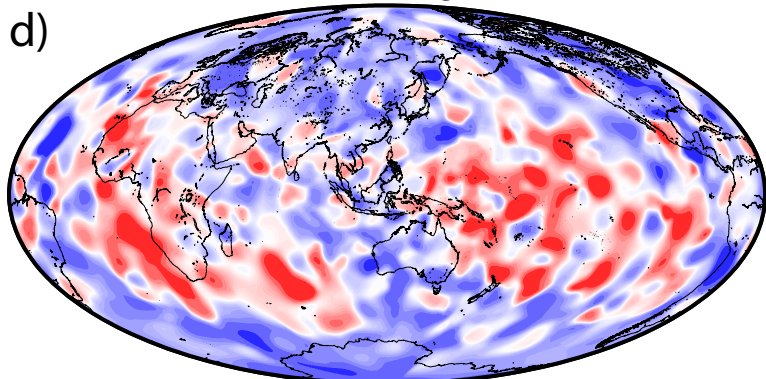
Mantle corrected using SEMUCB-WM1



Mantle corrected using GyPSuM



Mantle corrected using SAW642ANb



Mantle corrected using S362ANI+M

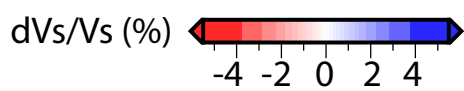
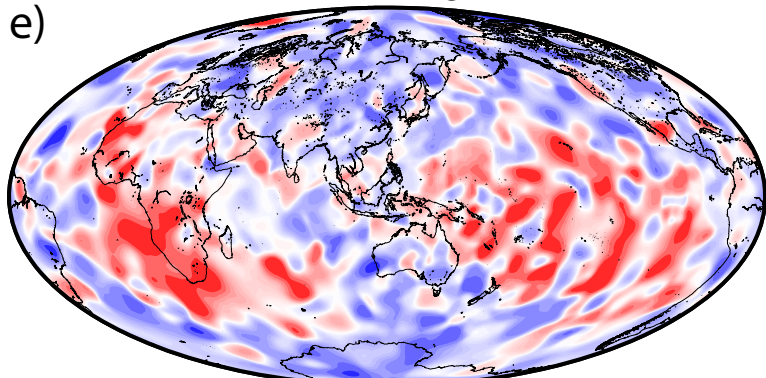


Figure 7.

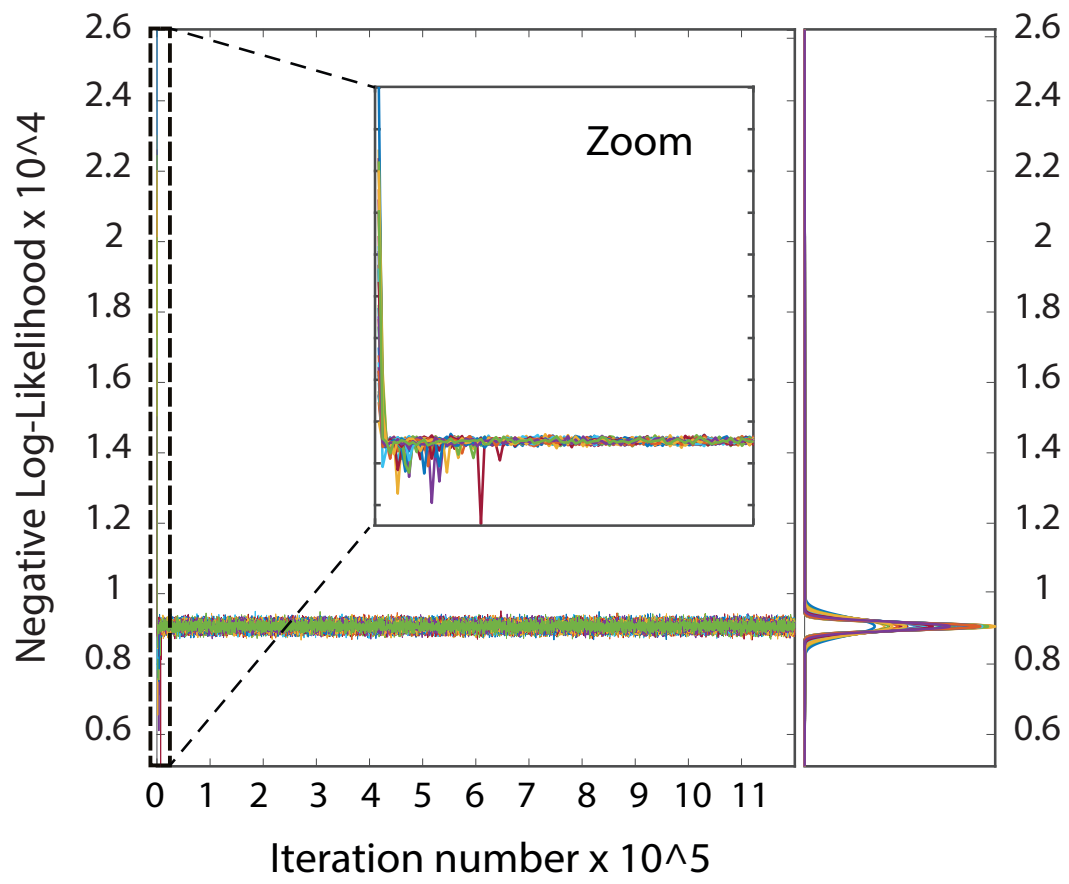
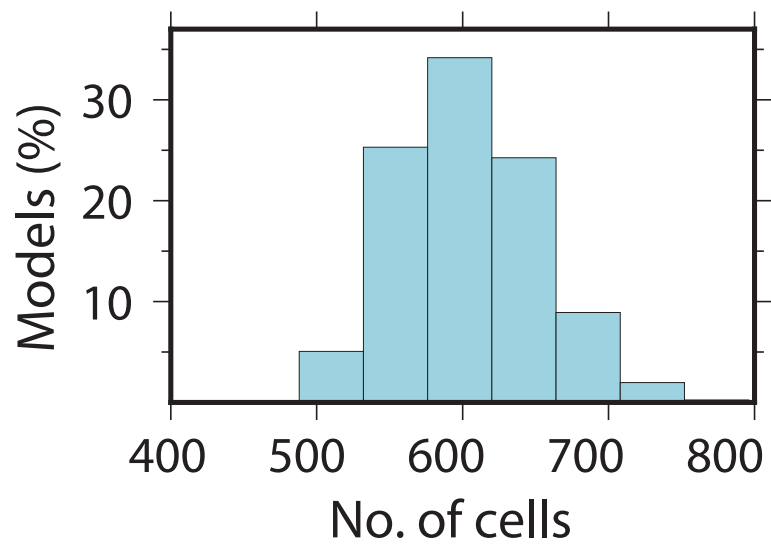
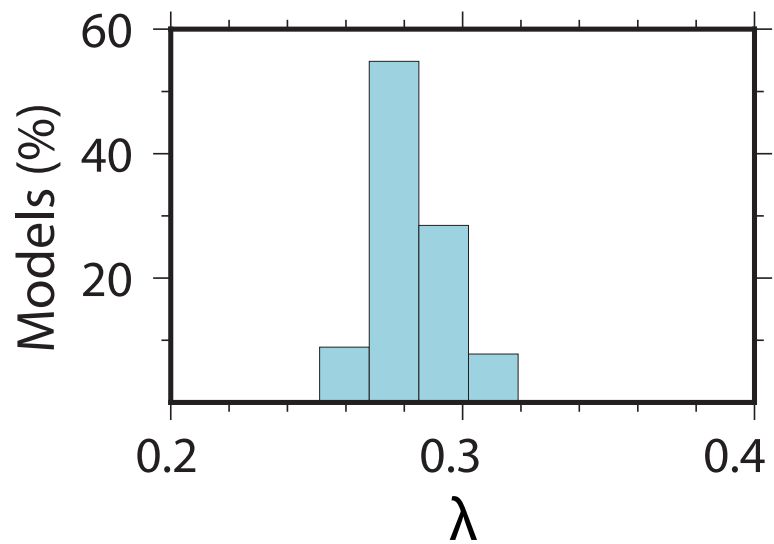


Figure 8.

a)



b)



c)

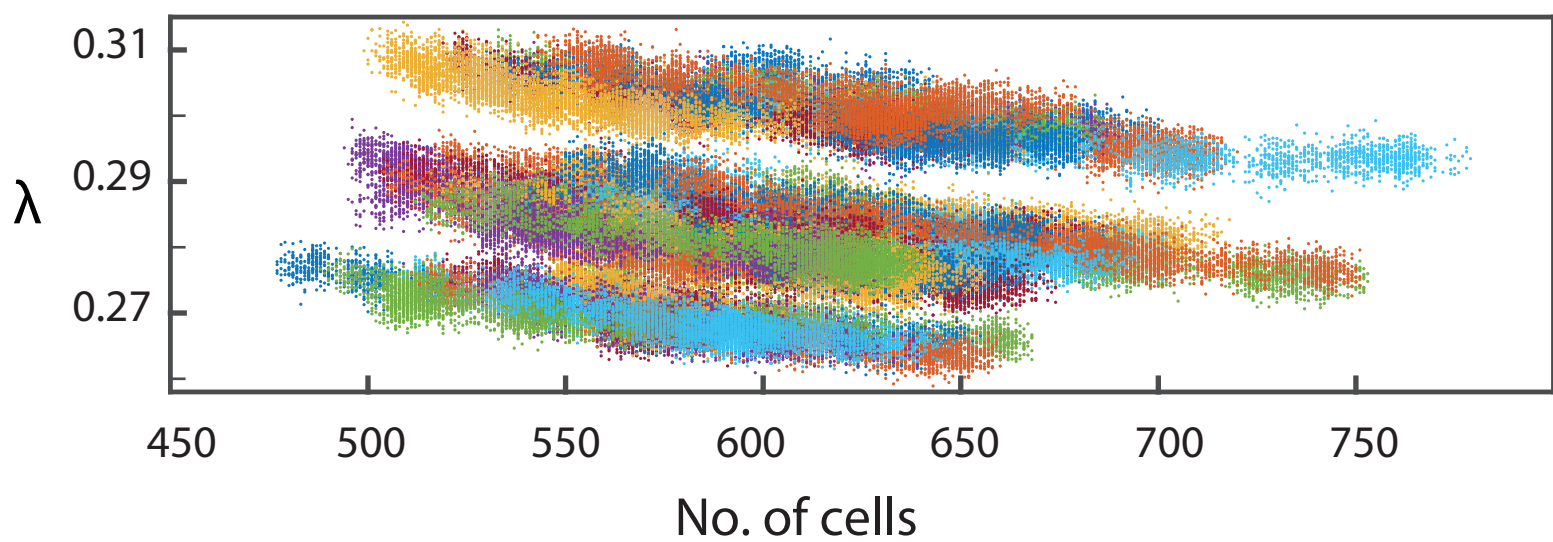


Figure 9.



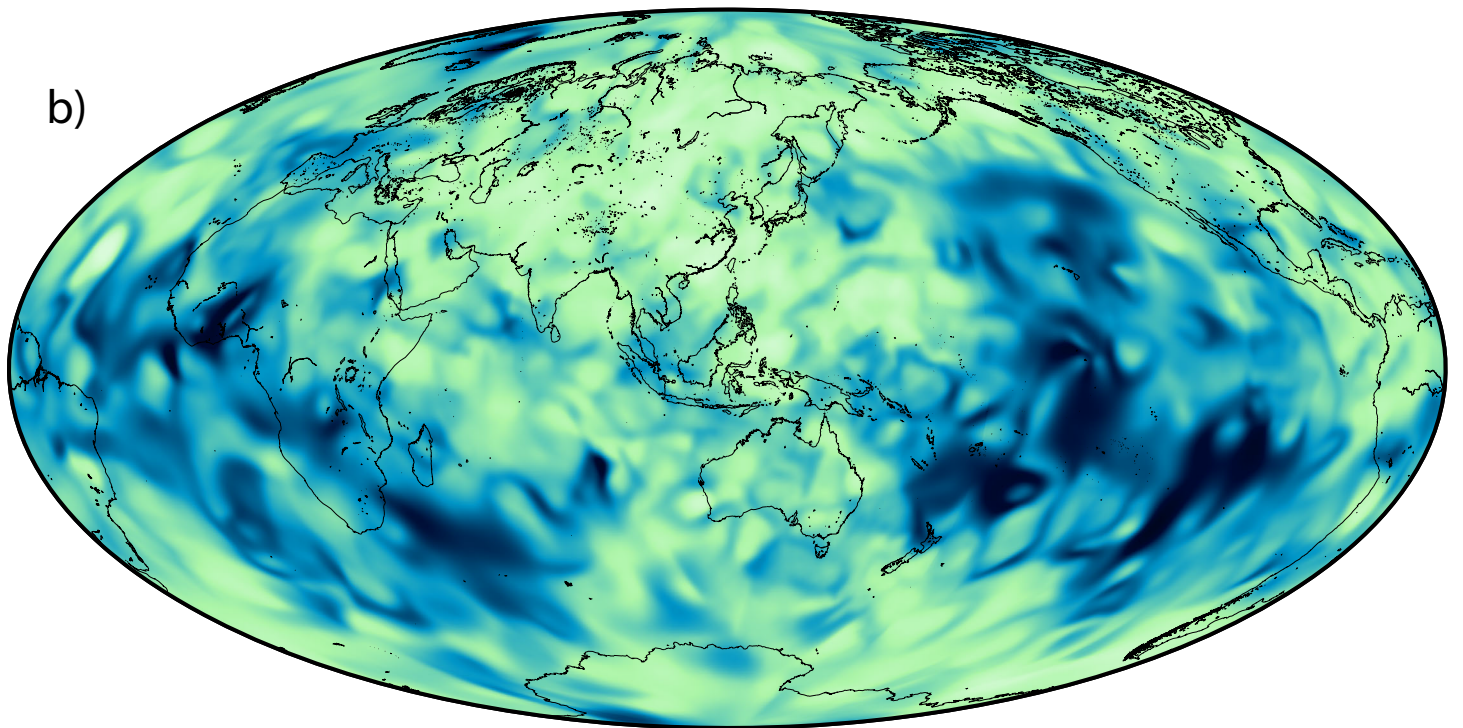
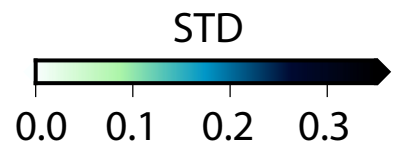
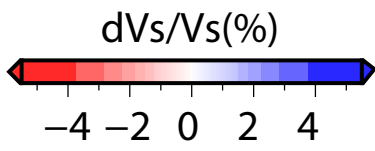
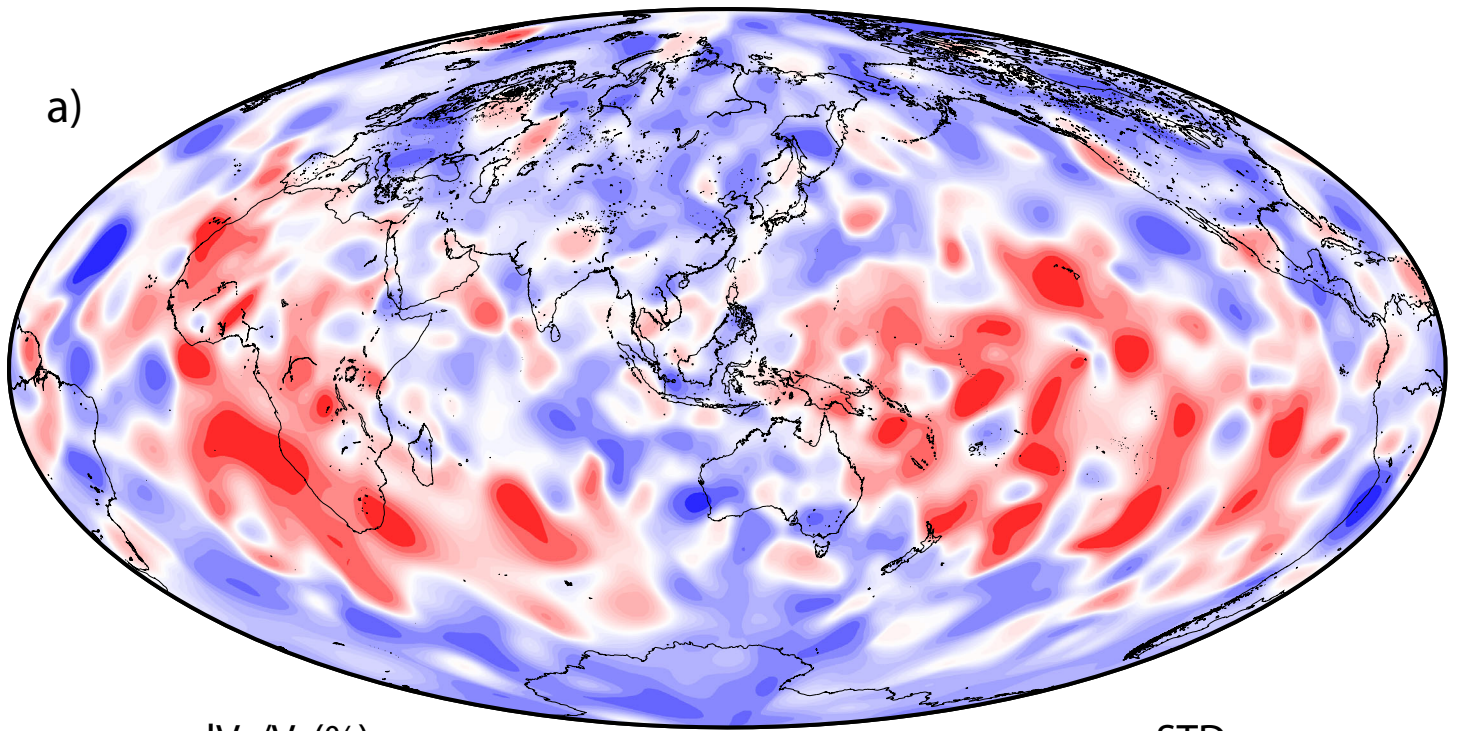
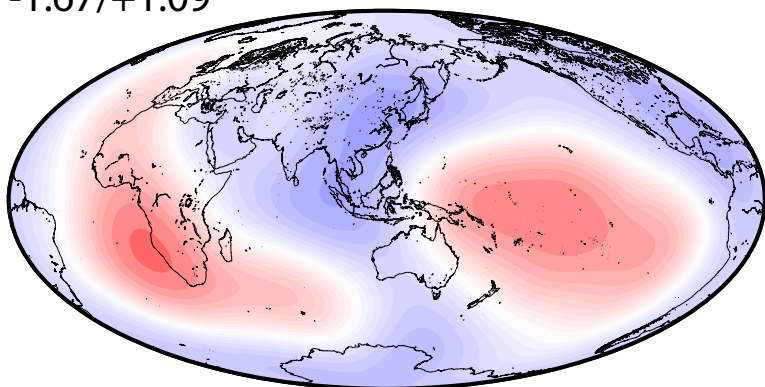


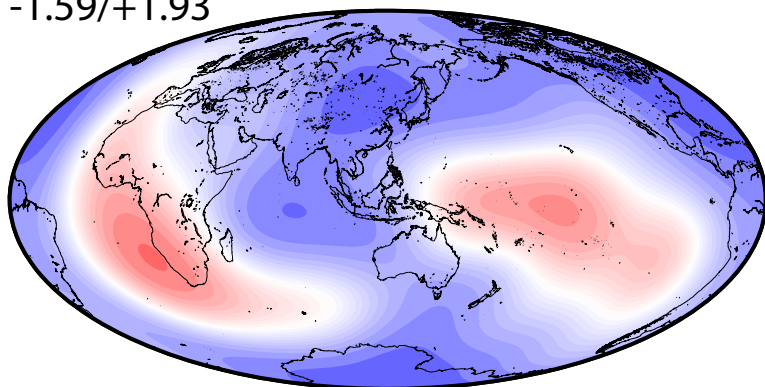
Figure 10.



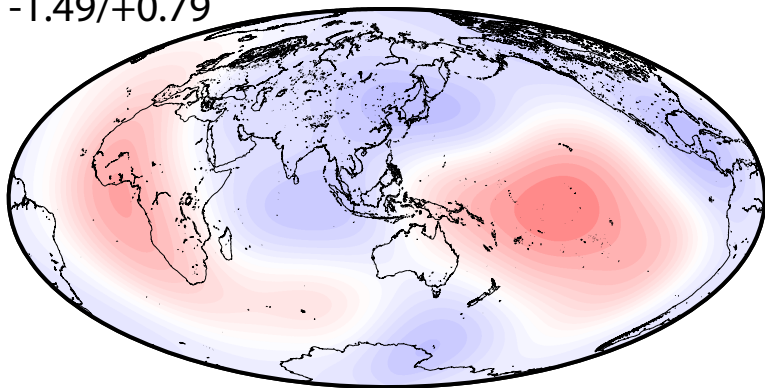
SGLOBE-rani

 $-1.67/+1.09$ 

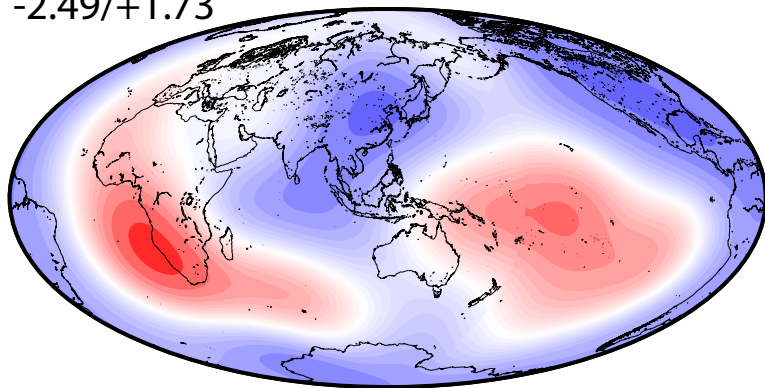
S362ANI+M

 $-1.59/+1.93$ 

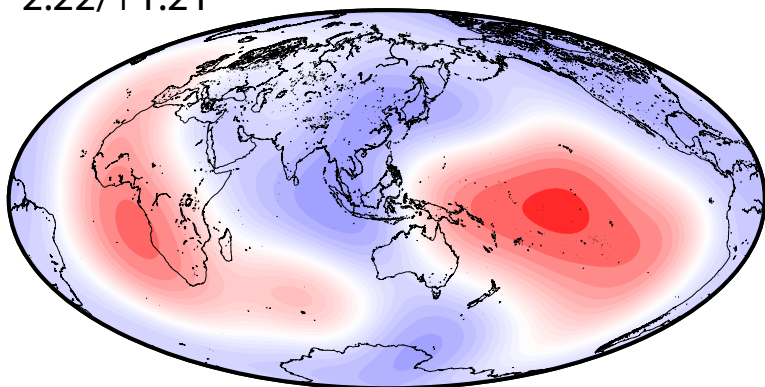
SP12RTS

 $-1.49/+0.79$ 

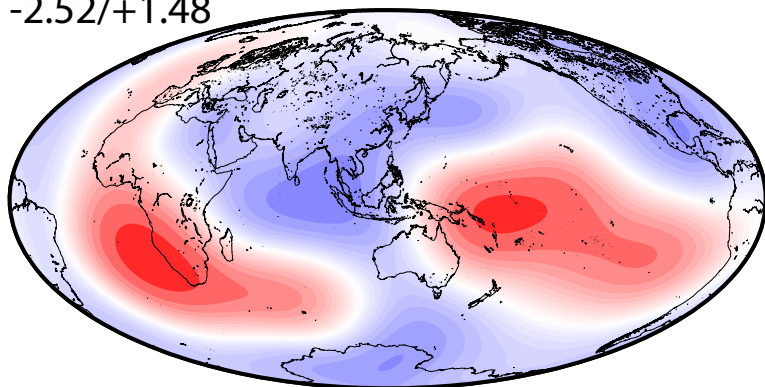
GyPSuM

 $-2.49/+1.73$ 

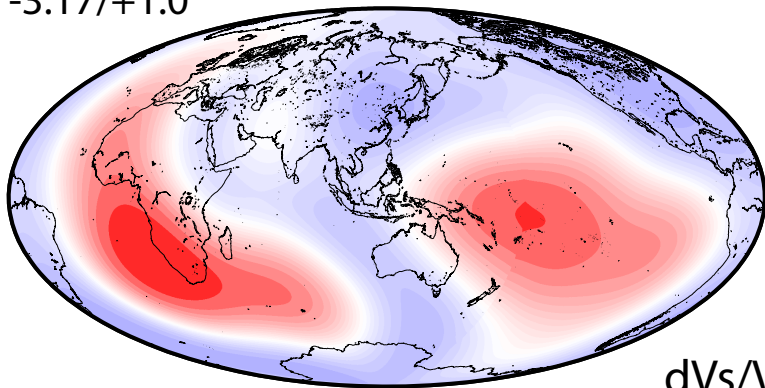
S40RTS

 $-2.22/+1.21$ 

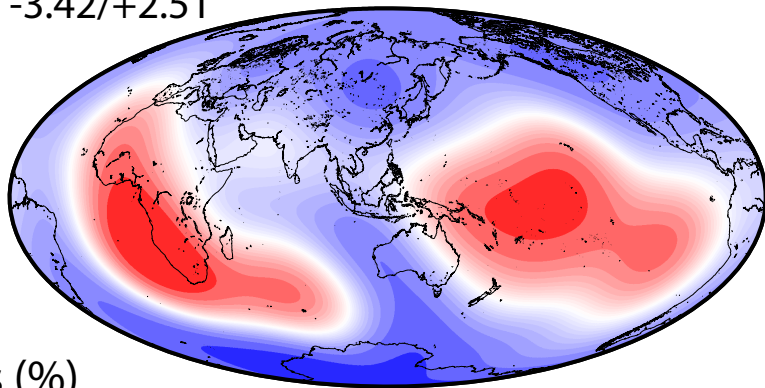
SEMUCB\_WM1

 $-2.52/+1.48$ 

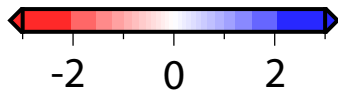
HMSL

 $-3.17/+1.0$ 

This Study

 $-3.42/+2.51$ 

dVs/Vs (%)



-2

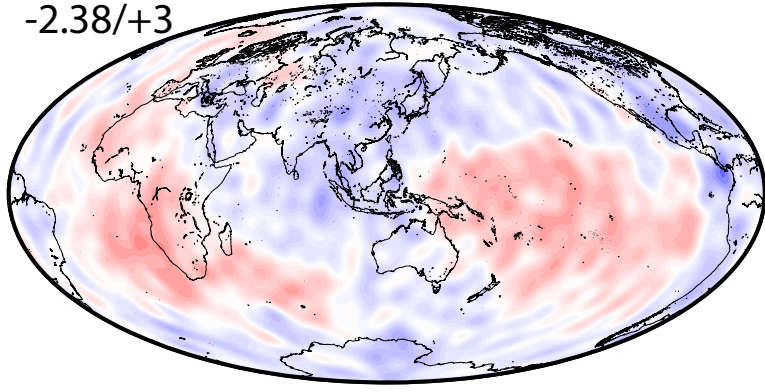
0

2

Figure 11.

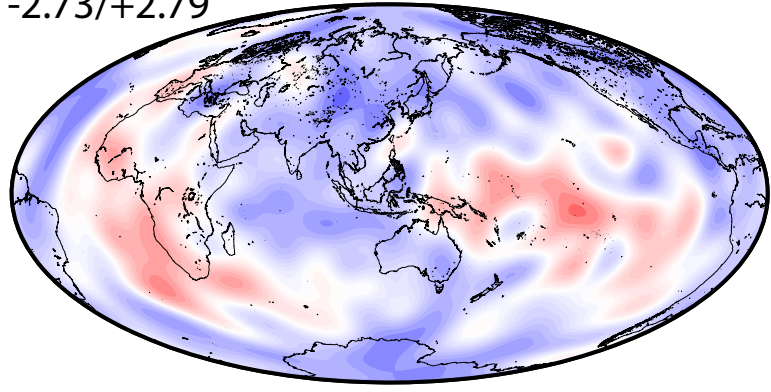
SGOLBERani

-2.38/+3



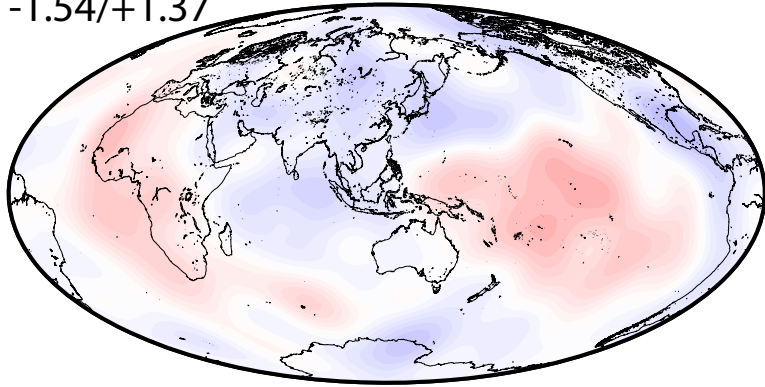
S362ANI+M

-2.73/+2.79



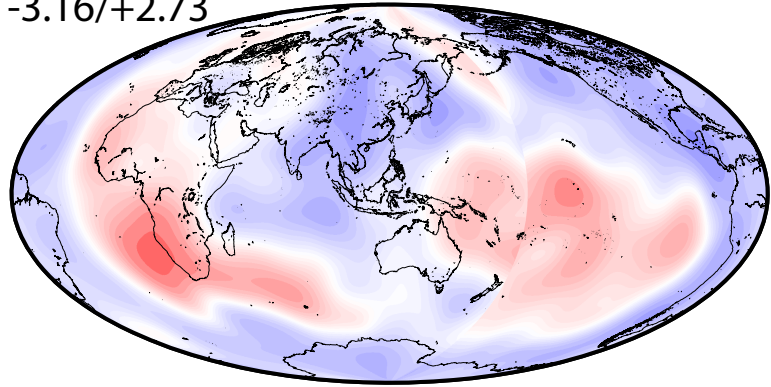
SP12RTS

-1.54/+1.37



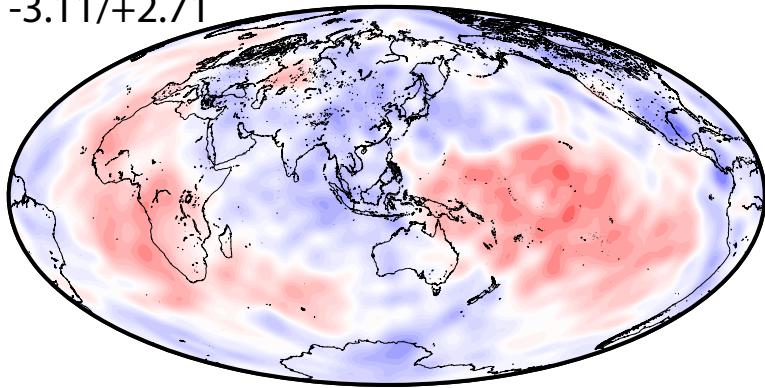
GyPSuM

-3.16/+2.73



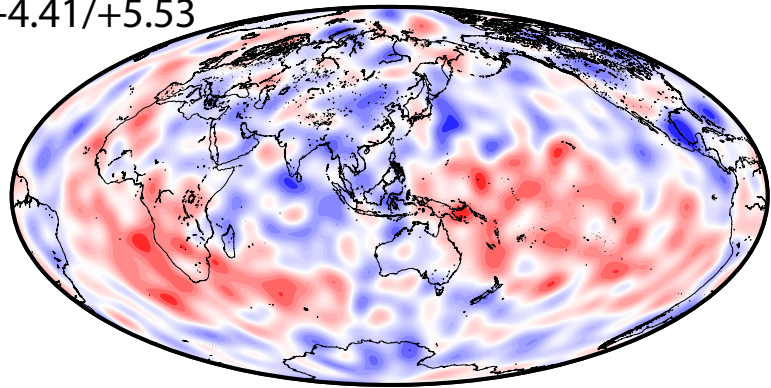
S40RTS

-3.11/+2.71



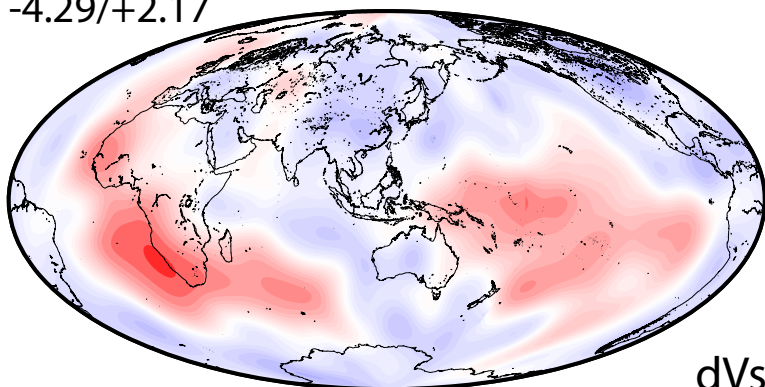
SEMUCB-WM1

-4.41/+5.53



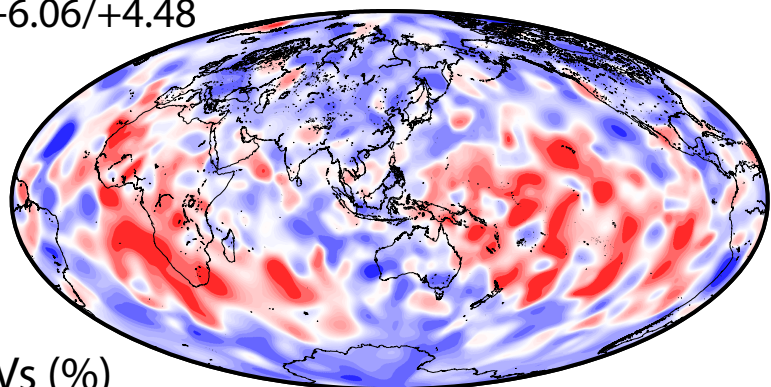
HMSL

-4.29/+2.17



This Study

-6.06/+4.48



dVs/Vs (%)

