

1 **AGE, PETROGENESIS AND TECTONIC IMPLICATIONS OF THE LATE PERMIAN**  
2 **PERALUMINOUS AND METALUMINOUS MAGMATIC ROCKS IN THE MIDDLE**  
3 **GOBI VOLCANOPLUTONIC BELT, MONGOLIA**

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## 38 Abstract

39

40 The Mongol–Okhotsk Belt, the youngest segment of the Central Asian Orogenic Belt, formed by  
41 the evolution and closure of the Mongol–Okhotsk Ocean. The oceanic closure formed two  
42 volcanoplutonic belts: Selenge Belt in the north and Middle Gobi Belt in the south (in present  
43 day coordinates). However, the origin and tectonic evolution of the Mongol–Okhotsk Belt in  
44 general, the origin and formation age of the Middle Gobi Belt in particular, remain enigmatic. To  
45 better understand the history of the magmatic activity in the Middle Gobi Belt, we conducted  
46 geochemical, U–Pb geochronological, zircon Hf, whole-rock Nd isotopic analyses of volcanic  
47 and plutonic rocks of the Mandalgovi suite, the major component of the Middle Gobi Belt. Our  
48 results show that the Mandalgovi suite consists of (i)  $265 \pm 2$  Ma biotite-granite; (ii)  $250 \pm 3$  Ma  
49 hornblende-granitoids; (iii) their volcanic equivalents of both: and (iv) gabbro-diorites. The  
50 geochemical compositions indicate that their precursor magmas were derived from crustal  
51 source. The protoliths of the biotite and hornblende-granitoids were metagraywacke and  
52 metabasalt, respectively. They are characterized by positive whole-rock  $\epsilon_{Nd}(t)$  and zircon  $\epsilon_{Hf}(t)$   
53 values, indicating the molten protoliths were juvenile crust. The biotite-granites formed by  
54 remelting of fore-arc sediments by ridge subduction and later hornblende-granites were  
55 emplaced at an intra-oceanic arc by the subduction of the Mongol–Okhotsk Ocean. We conclude  
56 that the magmatic rocks of the Middle Gobi formed in an active continental margin and/or intra-  
57 oceanic arc setting.

58

## 59 1. Introduction

60

61 The growth of the continental crust has represented one of the most exciting debates in the  
62 Earth Sciences for at least the last three decades (e.g., [Taylor and McLennan, 1985](#); [Jahn et al.,](#)  
63 [2000](#); [Kemp et al., 2006](#); [Kröner et al., 2014, 2017](#); [Rino et al., 2008](#); [Belousova et al., 2010](#);  
64 [Ganbat et al., 2021a](#)). Pioneering research pointed out that the formation of continental crust was  
65 episodic and most of its volume formed in the Precambrian (e.g., [Armstrong, 1981](#)). However,  
66 more recent studies suggest that significant continental growth took place during the Phanerozoic  
67 ([Jahn, 2004](#); [Hawkesworth et al., 2010](#); [Safonova, 2017](#); [Rosebaum, 2018](#)). Understanding the  
68 continental growth during the Phanerozoic can provide crucial knowledge on the mechanisms of

69 magmatic differentiation from the mantle to the crust, as well as important perceptions about the  
70 compositions and evolution of the continental crust.

71 The Central Asian Orogenic Belt (CAOB) is the world largest Pacific-type accretionary orogeny.  
72 It is located between Siberia, Kazakhstan, Tarim, and North China, and has a long history from  
73 the Neoproterozoic to the late Paleozoic of continuous amalgamation of multiple microcontinents,  
74 island arcs, oceanic plateaus, seamounts, ophiolites, and accretionary complexes after suturing at  
75 least two oceans (Paleo-Asian and Mongol–Okhotsk Oceans) and possibly several seaways and  
76 back-arcs (e.g., [Zonenshain et al., 1990](#); [Dobretsov et al., 1995](#); [Buslov et al., 2001](#); [Xiao et al.,  
77 2003](#); [Windley et al., 2007](#); [Safonova et al., 2011](#), [Safonova, 2017](#); [Wilhem et al., 2012](#);  
78 [Yakubchuk, 2017](#)). The CAOB is also the locus of rapid and extensive juvenile granitic formation  
79 during the Phanerozoic, but the petrogenesis, geodynamic and relationships with CAOB’s crustal  
80 growth are controversial ([Jahn, 2004](#); [Kröner et al., 2014](#); [Safonova, 2017](#)). A classical view of the  
81 CAOB holds that nearly half of the granites of the CAOB formed after the successive accretion of  
82 juvenile arc complexes and subduction accretion during the Phanerozoic ([Sengör, 1990](#)). Other  
83 view proposes that the Phanerozoic CAOB granitoids were the product of basaltic underplating in  
84 a post-collisional setting ([Jahn et al., 2000](#); [Wu et al., 2011](#)). Nonetheless, the origins of the  
85 CAOB’s granitoids might be far more complex. Some authors pointed out that many key features  
86 are explicable by Paleozoic ridge subduction (e.g. [Windley and Xiao, 2018](#)), and this model  
87 recently supported by local cases in the southwestern and southeastern CAOB (e.g., [Tang et al.,  
88 2012](#); [Li et al., 2021](#)).

89 The Mongol-Okhotsk Ocean (MOO) was the youngest oceanic basin, which closure finalized the  
90 welding of CAOB terranes ([Kravchinsky et al., 2002](#); [Sorokin et al., 2020](#)). The suturing of the  
91 MOO closure formed two volcanoplutonic belts in the territory of Mongolia: Selenge and Middle  
92 Gobi, on its northern and southern sides (in actual coordinates), respectively ([Fig. 1](#); [Zorin, 1999](#);  
93 [Parfenov, 2001](#)). The Middle Gobi Belt is thought to represent the southern active margin of the  
94 MOO ([Fig. 1](#)), however it remains a deficiency in detailed structural studies and up-to-date  
95 radiogenic ages and geochemical data. Consequently, the geodynamic evolution of the southern  
96 margin of the MOO and the reliable tectonic setting of the Middle Gobi Belt remains unclear. In  
97 this study, we present new zircon U–Pb ages, zircon Hf isotope and whole-rock geochemical and  
98 Nd isotope data from igneous rocks of the Mandalgovi volcanoplutonic suite of the Middle Gobi  
99 Belt. We will discuss the ages and tectonic settings of their emplacement, and petrogenesis in order

100 to contribute to the better understanding of the crustal growth and tectonic evolution of the Middle  
101 Gobi Belt.

102

## 103 **2. Geological setting**

### 104 *2.1 Geological background of Mongolia*

105 Mongolia is located in the central CAOB (Fig. 1), and has recorded multiple episodes of orogeny  
106 during the Neoproterozoic to the Late Paleozoic (e.g., Badarch et al., 2002; Windley et al., 2007;  
107 Rojas-Agramonte et al., 2011). Geologically, Mongolia consists of a Northern Domain and a  
108 Southern Domain separated by the Main Mongolian Lineament (Tomurtogoo et al., 1998; Badarch  
109 et al., 2002; Fig. 1). The Northern Domain comprises Neoproterozoic–Early Paleozoic  
110 accretionary units and a collage of collided Precambrian blocks (Badarch et al., 2002; Xiao et al.,  
111 2003; Wilhem et al., 2012). The Southern Domain includes relicts of an extended long-lived  
112 subduction-accretionary system of the Paleo-Asian Ocean, which was active from the  
113 Neoproterozoic to the Permian (e.g., Demoux et al., 2009; Kröner et al., 2010; Batsaikhan et al.,  
114 2018). In Late Paleozoic to Mesozoic time, the Southern Domain approached and eventually  
115 docked with the Northern Domain to close MOO.

116

### 117 *2.2 The Mongol–Okhotsk Belt*

118 The Mongol–Okhotsk Belt extends from central Mongolia to the Okhotsk Sea over 3000 km. Its  
119 formation is linked with the evolution of the Mongol–Okhotsk Ocean (MOO), which existed  
120 between the Central Mongolia–Erguna block to the south, and the Siberian Craton to the north.  
121 The suture zone is marked by the ca. 325–314 Ma Adaatsag–Khukhu Davaa ophiolite belt  
122 (Tomurtogoo et al., 2005; Zhu et al., 2018) (Fig. 1). The MOO opened in late Ordovician–early  
123 Silurian times, possibly as a back-arc basin on the northern (in present coordinates) active margin  
124 of the Paleo-Asian Ocean near the Central Mongolia–Erguna block (Bussien et al., 2011; Miao et  
125 al., 2020; Wilfred et al., 2020). Other scientists suggested that the MOO opened much earlier in  
126 Neoproterozoic to Cambrian times (e.g., Zonenshain, 1990; Zorin, 1999).

127 The final closure of the MOO occurred in the Mesozoic (Jurassic to Cretaceous) as suggested by  
128 paleomagnetic data and the early Jurassic age of the youngest marine sediments (e.g., Cogné et al.,  
129 2005; Metelkin et al., 2010; Sorokin et al., 2020). So far, most scientists agree on the kinematic  
130 model of closure: a scissors-style, which is reflected by the younging trend of sediments and

131 intrusions along the suture zone from west to east (e.g., [Metelkin et al., 2010](#); [Donskaya et al.,](#)  
132 [2013](#)).

133 The Mongol–Okhotsk suture zone is bounded by the Khangay–Khentey accretionary terrane in the  
134 north-west ([Fig. 1](#); [Kelty et al., 2008](#); [Bussien et al., 2011](#); [Ruppen et al., 2014](#)). The Khangay–  
135 Khentey terrane hosts all typical units of Ocean Plate Stratigraphy (OPS; from bottom to top):  
136 oceanic basalt (MORB, OIB), radiolarian pelagic chert, hemipelagic siliceous mudstone, siltstone,  
137 thick trench turbidite, and sandstone (e.g., [Kelty et al., 2008](#); [Safonova et al., 2009](#); [Ruppen et al.,](#)  
138 [2014](#); [Dagva-Ochir et al., 2020](#)). The provenance analysis of sandstones and the U–Pb ages of  
139 detrital zircons show that the sediments of the northern part deposited from the Silurian to early  
140 Carboniferous times in a subduction-accretion setting, which apparently stopped in the  
141 Carboniferous. Later, subduction re-initiated during the Permian.

142 The Selenge and Middle Gobi belts are abutting with the Khangay–Khentey terrane ([Fig. 1](#)).  
143 Selenge Belt is believed to represent the northern, ([Fig. 1](#)), the Middle Gobi Belt is the southern  
144 active margin of the MOO ([Zorin, 1999](#); [Parfenov, 2001](#)). The southern part of the Khangay–  
145 Khentey terrane experienced a back-arc extension during the Early Paleozoic ([Bussien et al., 2011](#);  
146 [Wilnkler et. al., 2020](#)), and from the Devonian and to early Carboniferous it evolved as a passive  
147 margin. In the late Carboniferous, the subduction re-initiated resulting the transportation of detrital  
148 zircons from volcanic rocks and plutons of the Middle Gobi Belt until the Triassic ([Kelty et al.,](#)  
149 [2008](#); [Bussien et al., 2011](#)).

150

### 151 *2.3 The Middle Gobi Belt*

152 The Middle Gobi Belt is a magmatic domain located south of the Mongol–Okhotsk suture  
153 ([Fig. 1](#)) ([Badarch et al., 2002](#)) and contains Precambrian basement rocks as bulk of the crust. The  
154 Precambrian basement consists of Meso–Neoproterozoic quartz-feldspathoid gneisses,  
155 metamorphic schists, pelitic gneisses, and metasandstones and marbles ([Figs. 1, 2](#)) overlain by the  
156 Silurian and Devonian marine sedimentary and volcanic rocks (turbiditic and pelagic series),  
157 which are intruded and overlapped by Permian plutonic and volcanic rocks, respectively ([Badarch](#)  
158 [et al., 2002](#)).

159 The Middle Gobi volcanoplutonic belt is dominated by Permian calc-alkaline andesites, dacites,  
160 rhyolites, trachyrhyolites and subalkaline granites ([Khanchuk, 2015](#); [Gerel et al., 2019](#)). Several  
161 authors, according to shallow marine flysch deposit with volcanic horizons, considered the

162 northern margin of the Middle Gobi Belt as a separate belt—North Gobi, representing a fossil  
163 forearc basin (Khanchuk et al., 2015) (Fig. 1). The emplacement of both ~270 Ma peraluminous  
164 granites and their associated volcanics, and ~240 Ma metaluminous granites (Machiowak et al.,  
165 2012; Gerel et al., 2019) was related to the subduction of the MOO lithosphere (e.g., Zhao et al.,  
166 2017; 2020). The late Triassic (~210 Ma) bimodal volcanic suites and coeval A-type granites were  
167 interpreted to form in a back-arc setting due to rollback of the MOO slab (Zhu et al., 2016).

168 The Mandalgovi volcanoplutonic suite occurs in the western fringe of the Middle Gobi  
169 Belt and occupies a large area of  $250 \times 650 \text{ km}^2$  (Fig. 2; Gerel et al., 2019). Devonian granites and  
170 volcanics are overlain unconformably by Carboniferous to Permian volcanics, volcanoclastics,  
171 clastic rocks and gneissic migmatites, which occur with Mandalgovi gabbro, gabbro-diorites,  
172 hornblende-granite and biotite-granite (Figs. 1c and 2; Badarch et al., 2002). Mafic dykes are  
173 abundant and occur associated with the Mandalgovi suite and sedimentary strata (Fig. 3). Pre-  
174 Mesozoic rocks are unconformably covered by Jurassic to Cretaceous sedimentary rocks  
175 (conglomerate, sandstone, and gravelite) and volcanics basalts, basaltic andesites, andesites and  
176 rhyolites) (Fig. 2).

177

### 178 **3. Analytical methods**

#### 179 *3.1 U–Pb geochronology*

180 Zircon crystals were separated in the Graduate School of Science, Tohoku University,  
181 using standard techniques including conventional rock-crushing, magnetic and heavy liquid  
182 separation, and handpicking under a binocular microscope. Then, zircon crystals of similar size  
183 were mounted in epoxy discs. Zonation of zircon interiors was documented using  
184 cathodoluminescence (CL) imaging with a Hitachi S-3400N scanning electron microscope,  
185 equipped with a Gatan MiniCL. In-situ zircon U–Pb dating was carried out in the Okayama  
186 University of Science using a Thermo Fisher Scientific iCAP-RQ single-collector quadrupole  
187 coupled to a Teledyne Cetac Technologies Analyte G2 ArF excimer laser ablation (LA) system  
188 equipped with a HelEx 2 volume sample chamber. The laser ablation was conducted at the laser  
189 spot size of  $25 \mu\text{m}$ , the fluence of  $1.8 \text{ J/cm}^2$  and the repetition rate of 5 Hz (for details see Aoki et  
190 al., 2019, 2020). Zircon standard Nancy 91500 (1065 Ma; Wiedenback et al., 2004) was used for  
191 age calibration, NIST SRM 612 standard (Jochum et al., 2011) for instrument optimization, and  
192 Plešovice zircons (337 Ma; Sláma et al., 2008) as secondary standards for quality control. U–Pb

193 ages and concordia diagrams were, respectively, calculated and plotted using the programs  
194 IsoplotR (ver. 3.75; [Vermeesch, 2018](#)); concordia age of each sample incorporates errors on the  
195 decay constants and includes evaluation of concordance of apparent ages. The concordia ages and  
196 errors are at two-sigma level.

197

### 198 *3.2 Whole-rock geochemistry*

199 Eighteen samples were selected for whole-rock analysis. Concentrations of major and trace  
200 elements were measured at Activation Laboratories Ltd., Canada, using Code 4Litho  
201 Litho geochemistry Package with fusion inductively coupled plasma optical emission spectrometry  
202 (FUS-ICPOES) and inductively coupled plasma mass spectroscopy (FUS-ICPMS), respectively.  
203 One more sample of was analyzed at the Analytical Center for Multi-Element and Isotope Studies  
204 of the Institute of Geology and Mineralogy, Novosibirsk, Russia. Major oxides were determined  
205 by the X-ray fluorescence (XRF) method using an Applied Research Laboratories ARL-9900-XP  
206 analyzer, following the standard procedure. Trace elements were analyzed by mass spectrometry  
207 with inductively coupled plasma (ICPMS) after fusion with LiBO<sub>2</sub>. Simultaneous determination  
208 of all elements was carried out to low, medium, and high resolution, on a Finnigan Element-II  
209 high-resolution mass spectrometer with external calibration using BHVO-1 reference samples and  
210 an internal standard. The method has been validated through the analysis of nine reference  
211 materials. Relative standard deviations for all elements were <10% within the determined  
212 concentration ranges.

213

### 214 *3.3 Zircon Hf isotopes*

215 Hf isotope analysis was carried out using a Thermo Fisher Scientific Neptune Plus multi-  
216 collector (MC)-ICPMS in combination with a Geolas 2005 excimer ArF laser ablation system (193  
217 nm) at the Institute of Geology and Geophysics, Chinese Academy of Science, Beijing. The  
218 analyses for zircon grains from the granites were conducted with a beam diameter of 63 μm, 6 Hz  
219 repetition rate, and energy of 15 mJ/cm<sup>2</sup>. This setting yielded a signal intensity of 10 V at <sup>180</sup>Hf  
220 for the standard zircon Nancy 91500. Typical ablation time was 26 s, resulting in pits 20–30 μm  
221 deep. The initial <sup>176</sup>Hf/<sup>177</sup>Hf ratios for the unknown samples were calculated to their initial value,  
222 using the measured <sup>176</sup>Lu/<sup>177</sup>Hf ratios, the apparent age of each zircon grain, and an <sup>176</sup>Lu decay  
223 constant of  $1.867 \times 10^{-11} \text{ yr}^{-1}$  ([Söderlund et al., 2004](#)). Epsilon Hf were calculated using a present-

224 day chondritic  $^{176}\text{Hf}/^{177}\text{Hf}$  value of 0.282785 and  $^{176}\text{Lu}/^{177}\text{Hf}$  of 0.0336 (Bouvier et al., 2008) and  
225 the present-day felsic crustal ratio of  $^{176}\text{Lu}/^{177}\text{Hf} = 0.015$ .

226

### 227 3.4. Sm–Nd isotopic analysis

228 Sm–Nd isotopic analyses were performed at the Institute of Geology and Geochronology,  
229 Russian Academy of Sciences, Saint-Petersburg. About 100 mg of whole-rock powder was  
230 dissolved in a mixture of HF, HNO<sub>3</sub>, and HClO<sub>4</sub>. A  $^{149}\text{Sm}$ – $^{150}\text{Nd}$  spike solution was added to all  
231 samples before dissolution. REEs were separated on BioRad AGW50-X8 200–400 mesh resin  
232 using conventional cation-exchange techniques. Sm and Nd were separated by extraction  
233 chromatography with a LN-Spec (100–150 mesh) resin. The total blank in the laboratory was  
234 0.1–0.2 ng for Sm and 0.1–0.5 ng for Nd. Isotopic compositions of Sm and Nd were determined  
235 on a TRITON TIMS mass-spectrometer. The precision ( $2\sigma$ ) of Sm and Nd contents and  
236  $^{147}\text{Sm}/^{144}\text{Nd}$  ratios were 0.5% and 0.005% for  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios.  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios were  
237 adjusted relative to a value of 0.512115 for the JNdi-1 standard. During the period of analysis, the  
238 weighted average of 10 JNdi-1 Nd standard runs yielded  $0.512108 \pm 7$  ( $2\sigma$ ) for  $^{143}\text{Nd}/^{144}\text{Nd}$ ,  
239 normalized against  $^{146}\text{Nd}/^{144}\text{Nd} = 0.7219$ . The  $\epsilon_{\text{Nd}}(t)$  values were calculated using the present-day  
240 values for a chondritic uniform reservoir (CHUR)  $^{143}\text{Nd}/^{144}\text{Nd} = 0.512638$  and  $^{147}\text{Sm}/^{144}\text{Nd} =$   
241  $0.1967$  (Jacobsen and Wasserburg, 1984). Whole-rock Nd model ages  $T_{\text{Nd(DM)}}$  were calculated  
242 using the model of Goldstein and Jacobsen (1988) according to which the Nd isotopic  
243 composition of the depleted mantle evolved linearly since 4.56 Ga ago and has a present-day  
244 value  $\epsilon_{\text{Nd}}(0)$  of +10 ( $^{143}\text{Nd}/^{144}\text{Nd} = 0.513151$  and  $^{147}\text{Sm}/^{144}\text{Nd} = 0.21365$ ). Two-stage (crustal)  
245 Nd model ages  $T_{\text{Nd(C)}}$  were calculated using a mean crustal ratio  $^{147}\text{Sm}/^{144}\text{Nd}$  of 0.12.

246

## 247 4. Results

248

### 249 4.1 Petrography

250 Gabbros (D0909) are coarse- to medium-grained gabbro-diorite with hypidiomorphic texture.  
251 The major minerals are plagioclase (50–55%), hornblende (25–30%), and biotite (15–20%) (Fig.  
252 4). Euhedral to subhedral plagioclase exhibits polysynthetic twin zoning, kinking, and is partially  
253 replaced by sericite (Fig. 4a). Accessory minerals are ilmenite, magnetite, and apatite. Biotite-  
254 grained granites (D0906) is from a porphyritic biotite-granite (Fig. 4c). The major minerals



255 include quartz (30–35%), plagioclase (30–35%), K-feldspar (20–25%), and biotite (15–20%).  
256 Plagioclase and K-feldspar occur as porphyritic minerals and biotite is common in fractures and  
257 along plagioclase boundaries (Fig. 4d). Accessory minerals are magnetite, apatite, and zircon.  
258 Hornblende-granites (D0914) are coarse- to fine-grained hypidiomorphic granodiorite consisting  
259 of subhedral to anhedral plagioclase (40–45%), K-feldspar (20–25%), subhedral quartz (15–  
260 20%), euhedral hornblende (15–20%), and biotite (5–10%) (Fig. 4e, f) plus accessory zircon and  
261 opaque minerals and secondary sericite and chlorite.

262

#### 263 4.2 Whole-rock major and trace element geochemistry

##### 264 *Gabbro-diorite*

265 This group of samples includes gabbro-diorite and diorite (Fig. 5a). They have relatively narrow  
266 spanned SiO<sub>2</sub> content (54–58 wt.%), high contents of MgO (4.08–4.34%), CaO (8.05–8.75%),  
267 FeO<sup>T</sup> (6.10–7.9%), and Al<sub>2</sub>O<sub>3</sub> (16.2–18.5%) and low total alkalis (Na<sub>2</sub>O + K<sub>2</sub>O = 3.23–3.92%)  
268 (Fig. 5c). The contents of TiO<sub>2</sub> and P<sub>2</sub>O<sub>5</sub> range from 0.77% to 0.79% and from 0.21% to 0.24%,  
269 respectively. The samples possess clear metaluminous features (Fig. 5d).

270 In terms of trace elements, the sample of gabbro-diorite display a comparably flat REE pattern  
271 [(La/Yb)<sub>N</sub> = 63–65] without significant Eu anomalies (Eu/Eu\* = 0.96–1.15) (Table 1; Fig. 6a),  
272 indicating no to weak fractionation and/or accumulation of plagioclase. They all have consistent  
273 primitive mantle-normalized patterns, characterized by lower Rb, Ba, Th, and U than the  
274 intermediate-felsic rocks of the same pluton, and characteristic troughs at Nb and Ta (Fig. 6b).

275

##### 276 *Biotite-granites*

277 Biotite-granite rock group consists of granites and rhyolites (Fig. 5a, b) and have relatively narrow  
278 ranges of SiO<sub>2</sub> (69.8–77.9 wt.%), total alkalis (Na<sub>2</sub>O + K<sub>2</sub>O = 6.7–8.21 wt.%) and Al<sub>2</sub>O<sub>3</sub> (11.2–  
279 15.3 wt.%). They are characterized by low TiO<sub>2</sub> (0.05–0.26 wt.%), moderate FeO<sup>T</sup> (0.84–0.96  
280 wt.%), and Mg# (12–44.6). The samples are high-K calc-alkaline granites (Fig. 5c). According to  
281 the A/CNK vs. A/NK diagram they can be classified as peraluminous granitoids (Fig. 6d). The  
282 chondrite normalized REE patterns (Fig. 6c) show enrichment in LREEs [(La/Yb)<sub>CN</sub> = 0.7–3.6],  
283 and weak negative Eu anomalies. The multi-element patterns show negative Nb, P, and Ti  
284 anomalies (Fig. 6d) and significant enrichment in K, Rb, Ba, Pb.

285

## 286 *Hornblende-granites*

287 This group of samples include granodiorite and dacite with lower SiO<sub>2</sub> (61–70 wt.%), K<sub>2</sub>O (1.91–  
288 3.99%), higher Al<sub>2</sub>O<sub>3</sub> (13.7–17.3%), Na<sub>2</sub>O (3.09–5.22%) and K<sub>2</sub>O + Na<sub>2</sub>O (5.15–8.2%) than the  
289 biotite-granitoids. The samples are characterized by high FeO<sup>T</sup> (3.09–6.08%), CaO (0.1–0.2%),  
290 TiO<sub>2</sub> (0.4–1%), and MgO (1.14–2.59%). Accordingly, they belong to the calc-alkaline to high  
291 calc-alkaline series (Fig. 6c; Table 1) and represent metaluminous granites (Fig. 5d).

292 In the CI-normalized REE diagrams (Fig. 6a), they exhibit various degrees of LREE enrichment  
293 with (La/Yb)<sub>N</sub> ratios of 3.5–25.4 and weak to zero negative anomalies at Eu (Eu/Eu\* = 0.52–1.03).

294 In the PM-normalized spidergrams (Fig. 6d), all samples display patterns with enrichment of Rb,  
295 Th, K, La, Ce, and Zr and significant troughs at P and Ti.

296

## 297 *4.3 .Geochronology*

298 *Biotite-granite (Sample D0906):* Zircons were colorless or light yellow and are stubby or euhedral  
299 to anhedral, about 100–300 μm in size, elongation ratio 1 to 3. CL images show fine-scaled  
300 oscillatory zoning and rare unzoned xenocrystic cores (Fig. 7a). Fourteen analyses show the  
301 contents of U and Th varying from 67 μg/g to 175 μg/g and from 57 μg/g to 204 μg/g (Table 2),  
302 respectively, with Th/U ratios ranging from 0.6 to 1.2, suggesting a magmatic origin for the zircon  
303 (e.g., Kelly and Harley, 2005; Yakymchuk et al., 2018). The U–Pb ratios give a concordia age of  
304 265 ± 2 Ma (MSWD = 0.19, Fig. 8a), which is interpreted as the crystallization age of biotite-  
305 granite.

306 *Granodiorite (Sample D0914):* Zircons are colorless or light brown, euhedral to subhedral, and  
307 stubby prisms up to 200 μm long with aspect ratios not exceeding 2 and exhibit fine oscillatory  
308 zoning (Fig. 7b). Sixteen analyses of zircons show the contents of U and Th varying from 143 μg/g  
309 to 236 μg/g and from 136 μg/g to 300 μg/g, respectively, and the ratios of Th/U ranging from 0.9  
310 to 1.3. All analyses plotted on or near the concordia and yielded an age of 250 ± 3 Ma (MSWD =  
311 1, Fig. 8b). This age is interpreted to represent the emplacement time of the granodiorite.

312

## 313 *4.4 Nd–Hf isotopes*

314 Seven analytical spots on the zircon grains/domains from the biotite-granite (Sample D0906)  
315 yielded initial ε<sub>Hf</sub>(*t*) values and two-stage model (T<sub>DM2</sub>) ages ranging from + 5.07 to + 7.54 and  
316 from 775 to 898 Ma, respectively (Table 3). Eight analytical spots on zircon grains/domains from

317 the granodiorite (Sample D0914) have  $\epsilon_{\text{Hf}}(t)$  values from + 4.47 to + 6.23 and  $T_{\text{DM2}}$  between 639  
318 and 734 Ma. The whole-rock Nd isotope ratios yielded the values of  $\epsilon_{\text{Nd}}(t)$  at 0.4, which  
319 corresponds to a  $T_{\text{DM2}}$  of 877 Ma.

320

## 321 **5. Discussion**

322

### 323 *5.1 Timing of magmatism*

324 Based on lithostratigraphic relationships and K–Ar ages, previous researchers considered  
325 that the Mandalgorvi volcanoplutonic suite of the Middle Gobi Belt consists of Neoproterozoic  
326 (Zaytsev and Luchitsky, 1979), early Cambrian, Devonian or early Permian (Gerel et al., 2012,  
327 2019) formations. The state geological map shows the Mandalgorvi biotite-granites as early  
328 Cambrian in age (Zaytsev and Luchitsky, 1979). Our new zircon U–Pb data demonstrate that these  
329 granitoids crystallized/emplaced during the Permian at ~265 Ma (Fig. 8). The designation of  
330 Precambrian age for the basement metamorphic rocks (Fig. 2) was mainly based on the intrusive  
331 relation between the so-called “early Cambrian” granites, where no precise radiogenic dating were  
332 performed previously. Therefore, reevaluating the age of the allegedly “Precambrian basement”  
333 metamorphic rocks in the future is needed.

334

### 335 *5.2 Petrogenesis and magma sources*

336

#### 337 *5.2.1 Biotite-granites*

338 The Mandalgorvi biotite-granites are characterized by high  $\text{SiO}_2$ , K, Rb, Pb, and low  $\text{Na}_2\text{O}$ , REE,  
339 and Sr, and high A/CNK (Fig. 5), and in the  $\text{CaO}/\text{Na}_2\text{O}$  vs.  $\text{Al}_2\text{O}_3/\text{TiO}_2$  plot they fall in the field  
340 of peraluminous granites (Fig. 5d). Peraluminous, Si-rich granites can be produced by (1) partial  
341 melting of Al-rich metapelite and metagraywacke (Patiño Douce, 1999; Eyal et al., 2004; Healy  
342 et al., 2004); (2) partial melting of tonalite or granodiorite at pressures  $\geq 8$  kbar with restitic  
343 clinopyroxene (Patiño Douce, 1999); (3) mixing of the basaltic magma with terrigenous sediments.  
344 We did not observe wide range values of  $\epsilon_{\text{Hf}}(t)$  and  $\epsilon_{\text{Nd}}(t)$  nor abundance of the mafic microgranular  
345 enclaves within biotite-bearing granites, both being considered as an indicators of magma mixing  
346 and source heterogeneity, precluding the mixing model. The depletion of the granites in Sr, Nb, P,  
347 and Eu do suggest fractional crystallization (Fig. 6d). The negative Nb–Ti and P anomalies can be

348 related to the fractionation of Ti-phases (ilmenite, titanite, etc.) and apatite, respectively. The Eu  
349 troughs in the REE spectra (Fig. 9c) and the Sr negative trend in the Ba vs. Sr and Sr vs. Eu  
350 diagrams (Fig. 9) indicate fractionation of plagioclase, i.e., shallow depth of melting.  
351 Compositions of biotite-granites are comparable with melts produced by dehydration-melting of  
352 various kinds of metasediments at low pressure (Fig. 10a), suggesting the derivation of their  
353 precursor magma from shallow crustal levels. This is consistent with the compositions of these  
354 granitoids which are close to the minimum point in the normative Qz–Ab–Or phase diagram ( $P <$   
355  $5$  kbar) (Fig. 10b). This condition precludes the partial melting of the tonalite at the pressures  $P \geq$   
356  $8$  kbar. In the diagrams of  $(\text{Na}_2\text{O} + \text{K}_2\text{O})/(\text{FeO}^{\text{T}} + \text{TiO}_2)$  vs.  $(\text{Na}_2\text{O} + \text{K}_2\text{O} + \text{FeO}^{\text{T}} + \text{MgO} + \text{TiO}_2)$   
357 and  $\text{Al}_2\text{O}_3/(\text{MgO} + \text{FeO}^{\text{T}})$  vs.  $\text{CaO}/(\text{MgO} + \text{FeO}^{\text{T}})$  diagrams indicate partial melting of a  
358 metagraywacke source for generating the biotite-granitic magma (Fig. 11). The Rb/Ba vs. Rb/Sr  
359 ratios indicate that the source magma was from clay-poor melts in a shallow depths (Fig. 11c).

360

### 361 5.2.2 Hornblende-granites

362 The Mandalgovi hornblende-granites are classified as metaluminous granitoids (Fig. 5). In general,  
363 calc-alkaline metaluminous granites form by assimilation and fractional crystallization of mantle-  
364 derived basaltic sources, mixing of mafic magma with crust-derived felsic magma, or partial  
365 melting of juvenile mafic crust (e.g., Chappell and White, 2001; Kemp et al., 2007; Clemens et al.,  
366 2011; Moyen et al., 2017). Metaluminous granites generated by magma mixing typically carry  
367 abundant mafic enclaves and have a variable chemical composition (e.g., Chappell, 1996), which  
368 is not observed in our samples. Na-rich ( $\text{Na}_2\text{O}/\text{K}_2\text{O} > 1$ ), high- $\text{Al}_2\text{O}_3$  melts with intermediate to  
369 silicic compositions may be produced by  $\sim 20$ – $40\%$  dehydration melting of metabasalts at  
370 temperatures between  $1050$  and  $1100$  °C, leaving a granulite residue at  $8$ – $12$  kbar and garnet  
371 granulite to eclogite residues at  $12$ – $32$  kbar (Rapp and Watson, 1995; Rapp, 1995). However,  
372 unlike the biotite-granites, show no notable trends of fractionation (Fig. 9a, b), thus excluding the  
373 partial melting scenario. The Mandalgovi hornblende-granitoids have low K and Th; and  
374 dominated by intermediate compositions, which are consistent with those of experimental melts  
375 derived by the dehydration melting of metabasalts or amphibolite (Figs. 11a–c). Therefore, we  
376 think that the Mandalgovi hornblende-granites formed by the partial melting of mafic protoliths.

377

### 378 5.2.3 Gabbro-diorites

379 Gabbro-diorites are mostly tholeiitic and characterized by low SiO<sub>2</sub> and total alkali, but high MgO,  
380 FeO<sup>T</sup> and CaO but enriched in LREE and LILE and depleted in HREEs and HFSEs. Those  
381 geochemical features are typical of arc magmas. In addition, those features can result from crustal  
382 contamination, since the average lithospheric mantle and continental crust are both depleted in Nb  
383 and Ta relative to neighboring elements (Th, La) (Rudnick and Gao, 2003). However, significant  
384 crustal contamination leads not only to Nb–Ta depletion but also increases concentrations of Zr  
385 and Hf, giving rise to negative Nb–Ta anomalies and positive Zr–Hf anomalies in mantle-  
386 normalized trace element spidergrams (Fig. 6b) (Zhao and Zhou, 2009). The Mandalgovi gabbro-  
387 diorites display troughs at Nb–Ta, and Zr, and peaks at Pb, which cannot be completely attributed  
388 to crustal contamination, but could indicate arc-derived magmatism.  
389 Nb is commonly depleted in the lithospheric mantle relative to La, so low Nb/La ratios (<0.5) for  
390 mafic magmas suggest a lithospheric mantle source and higher ratios (>1) indicate an OIB-like  
391 asthenospheric mantle source (Smith et al., 1999). The Nb/La ratios for the gabbro-diorites are  
392 between 0.1 and 0.3, i.e. closer to the lithospheric mantle source.

393

### 394 *5.3 Tectonic implications*

395 The Mandalgovi suite of the Middle Gobi Belt hosts spatially and temporally associated  
396 peraluminous biotite-granite, metaluminous hornblende-granite igneous rocks and gabbro-diorites.  
397 The formation of peraluminous granites is generally attributed to collisional and/or post-collisional  
398 settings (Barbarin 1996, 1999; Chappell et al., 2011). A normal crustal thickness of 15–20 km was  
399 suggested in the Central Mongolia in the Paleozoic by Zorin et al. (1993), which is distinct from  
400 collisional orogens with overthickened crust. Additionally, the small amount of peraluminous  
401 granites and absence of detachment fault zones, gneiss dome and metamorphic core complexes,  
402 usually associated to the post-collisional granites are indicate that they were not formed in a post-  
403 collisional setting (Fig. 2). Moreover, studies showed that continuous subduction-related  
404 widespread magmatism in the Central Mongolia from early Permian and to the late Triassic (e.g.,  
405 Ganbat et al., 2021b), and possible post-collisional granites appeared after ca. 220 Ma (Fig. 12;  
406 Zhu et al., 2016)

407

408 Normal oceanic crust subduction may hardly provide sufficient heat for partial melting of  
409 metasedimentary rocks in a shallow depth, and their generation requires specific geothermal  
410 condition (e.g., [Zhang et al., 2004](#)).

411 More recent studies proposed that peraluminous granites could also form in other geodynamic  
412 scenarios that provide an enhanced thermal gradient to the crust facilitating crustal melting  
413 including: ridge subduction ([Cai et al., 2011](#); [Kong et al., 2019](#)), mantle plumes ([Li et al., 2011](#)).  
414 Alternative studies proposed back-arc extension model to facilitate the generation of S-type  
415 granites in the New England orogen of circum-Pacific ([Collins and Richards, 2008](#); [Collins et al.,](#)  
416 [2020](#)). This model suggests that back-arc extension is induced by repeated, long-term subduction  
417 retreat and the sediment-dominated back-arc basin is triggered to melt when hot basaltic magmas  
418 intruded into the thinned back-arc crust once slab retreat is re-established. Such a setting requires  
419 a long history of subduction before the extension to provide crustal thickening before the formation  
420 of peraluminous granites, flat subduction initiation, and subsequent rollback ([Collins and Richards,](#)  
421 [2008](#)). We decline such a scenario as the peraluminous granites of the Middle Gobi Belt emplaced  
422 in the fore-arc setting ([Fig. 1](#)). The positive zircon  $\varepsilon_{\text{Hf}}(t)$  values and whole-rock  $\varepsilon_{\text{Nd}}(t)$  isotopic  
423 values observed in peraluminous samples make them distinct from classical peraluminous granites  
424 ([Fig. 12](#)), and are generally interpreted to require an extreme heat source ([Appleby et al., 2010](#);  
425 [Liu and Zhao, 2018](#)). These characters also are typical of greywackes formed by destruction of  
426 juvenile igneous rocks formed at intra-oceanic arcs ([Safonova et al., 2017, 2021](#)). Therefore, we  
427 suggest that the peraluminous granites may have formed by melting of the sediments in a fore-arc  
428 setting of an intraoceanic arc.

429 After ~250 Ma, peraluminous granitoids were supplanted by calc-alkaline metaluminous granites  
430 and gabbro-diorites, where they took place when by parental magmas being derived from the  
431 melting of the lithospheric mantle or juvenile mafic lower crustal sources, which led by downgoing  
432 oceanic slab.

433 Medium to high-K calc-alkaline metaluminous granites can form in two tectonic settings: (1) post-  
434 collisional settings similar to the Caledonides (e.g., [Litvinovsky et al., 2021](#)), (2) continental  
435 arcs similar to the Andes ([Roberts and Clemens, 1993](#)), and (3) intraoceanic arcs similar to the  
436 Japanese and Aleutian islands ([Kay et al., 1990](#)). We preclude the post-collisional setting because  
437 there is no evidence for the Permian intra-plate magmatism and post-collisional detachment fault  
438 in the study area. The ~250 Ma igneous association lacks abundant felsic magmas (rhyolites)

439 which occur rather as pyroclastic flow material derived from a continental margin (Kay et al.,  
440 1990; Wilson, 2007) (Fig. 2). Additionally, they have lower concentrations of K<sub>2</sub>O, Pb, Sr, Rb, Ba,  
441 Th, and U than those formed in an Andean-type active continental margin (Fig. 6b).

442 Their neutral to positive  $\epsilon_{\text{Hf}}(t)$  (+0.4 to +7.5) and  $\epsilon_{\text{Nd}}(t)$  (0.4) characteristics and the  $T_{\text{DM2}}$   
443 of 775 to 1019 Ma (Table 3) suggest the generation of primary magmas by the partial melting of  
444 a relatively depleted juvenile lower crust or depleted mantle source. Detrital zircon  $\epsilon_{\text{Hf}}(t)$  value  
445 from the northern margin of the Middle Gobi Belt also exhibits mostly positive  $\epsilon_{\text{Hf}}(t)$  (-7 to +14)  
446 (Bussien et al., 2011; Wilfred et al., 2020), with slightly negative values, implying that the crustal  
447 development occurred on the mature arc. Based on the above evidence, we conclude that the  
448 ~250 Ma granitoids and gabbro-diorites of the Mandalgovi pluton formed in an intra-oceanic  
449 subduction setting.

450 The abundance of hornblende and the incompatible trace element composition of the  
451 Mandalgovi gabbro-diorites suggest that they were generated in a subduction-related tectonic  
452 setting. In the tectonic discrimination diagrams for mafic rocks gabbro-diorites plot in the field of  
453 island arc (Fig. 13a, b). Their association with hornblende-granitoids (Fig. 2) and linear trend from  
454 tholeiitic to high-K calc-alkaline series suggest close ages of the two units. The emplacement of  
455 the peraluminous granites and coeval volcanic rocks was short after the peak of voluminous  
456 subduction-related metaluminous granites (after ~15 Ma; Fig. 12).

457 Normal oceanic crust subduction may hardly provide sufficient heat for partial melting of  
458 metasedimentary rocks in a shallow depth, and their generation requires specific geothermal  
459 condition. Mafic-intermediate dykes, and the presence of migmatites (Fig. 3) were associated  
460 which peraluminous granites and sedimentary strata also may require extreme heat source.

461 A ridge subduction is probably viable scenario for the tectonic setting of the Mandalgovi suite.  
462 Ridge subduction impacts strongly on magmatic activity, metamorphism and mineralization near  
463 convergent plate margins (Thorkelson, 1996; Sisson, 2003). When the ridge intersects with the  
464 subduction zone, a “slab window” will form between the subducted parts of the diverging oceanic  
465 plates (Dickinson and Snyder, 1979; Thorkelson, 1996). The upwelling of asthenospheric mantle  
466 through the slab window provides high heat flow that can induce partial melting of the slab edge,  
467 overlying mantle wedge and/or upwelling asthenospheric mantle and crustal rocks, and can  
468 produce a wide variety of magmas (e.g., Sisson et al., 2003). Ridge subduction in the CAOB was  
469 introduced by Windley et al. (2007) and it was one of the most remarkable tectonic setting that

470 responsible for the many key features of the CAO, as the modern circum-Pacific accretionary  
471 orogens (Liu et al., 2020; Ganbat et al., 2021c).

472 Thus, we suggest that the Permian oceanic ridge subduction caused anomalies of thermal and  
473 physical conditions in the surrounding mantle and upwelling of asthenosphere, triggered partial  
474 melting metasedimentary rocks which was derived from pre-Permian mature arc (Miao et al.,  
475 2020; Wilfred et al., 2020) in the forearc setting in a shallow depth. Intrusion of mafic–intermediate  
476 dykes into the accretionary prism and supports the ridge subduction and slab window model  
477 (Sisson et al., 2003; Windley et al., 2007). The mafic crust was the source for metaluminous  
478 granitoids. Although our study was confined to a relatively small area in the central part of the  
479 CAO (Fig. 1), it supports the view that ridge subduction was common during formation of the  
480 entire belt.

481 In the early-Carboniferous (ca. 325 Ma), the southward subduction of the MOO lithosphere started  
482 to form the Adaatsag and Khuhu Davaa ophiolites (Fig. 1). Two subduction zones were proposed  
483 at the southern margin of the MOO, which generated two arcs: intra-oceanic arc (Onon) and  
484 continental arc (Ereendavaa, Middle Gobi) (e.g., Zhu et al., 2018; Miao et al., 2020). However,  
485 there is no evidence of coeval or younger accretionary wedge, which should be an integral  
486 constituent of an intra-oceanic arc system and direct evidence for an arc except for the late  
487 Carboniferous detrital zircons. The positive  $\varepsilon_{\text{Hf}}(t)$  values of the detrital zircons may suggest  
488 tectonic erosion of former magmatic arc rocks, as those in NE Japan (Safonova et al., 2015; Pastor-  
489 Galán et al., 2021). The Middle Gobi Belt has been previously considered as a Permian–Triassic  
490 Andean-type active continental margin terrane (Parfenov, 2001; Tomurtogoo et al. 2005; Zhu et  
491 al., 2016). Our results show that the tectonic history of the Middle Gobi Belt is more complex than  
492 previously thought and that the Mandalgovi suite could have formed in an intra-oceanic arc setting  
493 with ridge subduction, but we still cannot exclude a continental margin setting for the whole belt.  
494 The northern margin of the Middle Gobi Belt, the North Gobi, could have formed in a fore-arc of  
495 the MOO, as well (Khanchuk et al., 2015).

496 During the ~265 Ma, the Mongol–Okhotsk Oceanic ridge subducted beneath the arc. As  
497 the slab window opened, the upwelling asthenosphere heated and then melted the overlying fore-  
498 arc sediments and induced to generate peraluminous granites. The mafic dykes in the Mandalgovi  
499 batholiths probably formed from the partial melting of a refractory mantle with previous magma  
500 extraction. This is also supported by migmatite occurrence adjacent to the batholith (Fig. 3). Later,



501 at ~250 Ma, metaluminous granitoids and gabbro-diorites of the Mandalgovi pluton formed in an  
502 intra-oceanic subduction of the Mongol–Okhotsk Ocean.

503 Previous studies show that the CAOBS may contain many examples of ridge subduction and slab  
504 windows, for example in the Junggar and the Chinese Altai, and in East Junggar, Beishan, the  
505 Tianshan, the Russian Altai, and in Inner Mongolia (Windley and Xiao, 2018). Ridge subduction  
506 may have caused extensive underplating/intraplating of hot mantle-derived basaltic magma as a  
507 result of asthenosphere upwelling, which could have provided not only sufficient heat for the  
508 widespread crustal melting, but also significant juvenile materials for the crustal growth in the  
509 CAOBS.

510

## 511 6. Conclusions

512 Based on the obtained result we draw the following conclusions:

513 1. The Mandalgovi suite consists of peraluminous and metaluminous granitoids, their  
514 volcanic counterparts, and gabbro-diorites.

515 2. The Mandalgovi peraluminous granites emplaced at ~265 Ma and derived by partial  
516 melting from metagraywacke in a fore-arc setting just above the slab window  
517 produced after the mid-ocean ridge subduction.

518 3. The Mandalgovi metaluminous granitoids were emplaced at ~250 Ma and resulted  
519 from the partial melting of a metabasaltic source in an intra-oceanic arc setting.

520 4. The Mandalgovi gabbro-diorites were formed by partial melting of a metasomatized  
521 lithospheric mantle source in a supra-subduction setting.

522 5. The northern margin of the Middle Gobi Belt may include both intra-oceanic setting  
523 above the subducted ridge and active continental margin terranes of the Mongol–  
524 Okhotsk Ocean.

525

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## 824 CAPTIONS

- 825 Figure 1. Sketch map illustrating the major tectonic units along the Mongol-Okhotsk Belt (Modified after  
 826 [Zorin, 1999](#); [Khanchuk et al., 2015](#); [Miao et al., 2020](#)). Inset map showing the location of the  
 827 CAOB (Modified after [Safonova et al., 2017](#)).
- 828 Figure 2. Simplified geological map of the Middle Gobi Belt (modified after the 1:200 000 State Geological  
 829 Map), showing sample locations.
- 830 Figure 3. Representative field photos of the Mandalgovi pluton
- 831 Figure 4. Photomicrographs of cross-polarized light view showing textures and mineral assemblage of the  
 832 studied samples from the Middle Mandalgovi pluton, Middle Gobi Belt. (a, b) Gabbrodiorite  
 833 (Sample D0909); (c, d) Biotite-granite (sample D0906); (e, f) Granodiorite (sample D0914); Bt—  
 834 biotite; Hbl—hornblende; Kfs—K-feldspar; Pl—plagioclase; Qz—quartz.
- 835 Figure 5. Major element discrimination diagrams showing the compositions and characteristics of the  
 836 studied samples from Middle Gobi Belt. (a) SiO<sub>2</sub> versus (Na<sub>2</sub>O + K<sub>2</sub>O) total alkali-silica (TAS)  
 837 diagram for plutonic rocks (after [Irvine and Baragar, 1971](#)), (b) SiO<sub>2</sub> versus (Na<sub>2</sub>O + K<sub>2</sub>O) total  
 838 alkali-silica (TAS) diagram for volcanic rocks (after [Le Bas et al., 1986](#)), (c) K<sub>2</sub>O wt% versus  
 839 SiO<sub>2</sub> wt% plot ([Peccerillo and Taylor, 1976](#)), (d) A/CNK [molar Al<sub>2</sub>O<sub>3</sub>/(CaO × Na<sub>2</sub>O × K<sub>2</sub>O)]  
 840 versus A/NK [molar Al<sub>2</sub>O<sub>3</sub>/(Na<sub>2</sub>O × K<sub>2</sub>O)] diagram, the boundary line is from [Maniar and Piccoli](#)  
 841 (1989).
- 842 Figure 6. CI-chondrite-normalized REE patterns and primitive-mantle-normalized trace element  
 843 spidergrams for the studied samples from the Middle Gobi Belt. Both chondrite and primitive-  
 844 mantle normalized values are from [Sun and McDonough \(1989\)](#). Green shaded area show island  
 845 arc derived rocks, pink shaded area shows active continental margin. Data from GEOROC.  
 846 (<http://www.earthchem.org/>).
- 847 Figure 7. Cathodoluminescence (CL) images of representative zircon crystals from the studied samples  
 848 from the Middle Gobi Belt. White circles show individual analysis spots, corresponding Pb–Pb  
 849 ages and red circles show an individual spot of Lu–Hf isotope and their ε<sub>Hf</sub>(t) values.
- 850 Figure 8. Concordia diagrams of zircons for samples from the Middle Gobi Belt, showing U–Pb isotope  
 851 ratios. Light grayish ellipses indicates discordant data excluded from the calculation.
- 852 Figure 9. (a) Ba versus Sr diagram, (b) Sr versus Eu diagram showing mineral fractionation.
- 853 Figure 10. (a) CaO/Al<sub>2</sub>O<sub>3</sub> versus CaO + Al<sub>2</sub>O<sub>3</sub> wt% plot; (b) The Ab–Or–Qz normative diagram of the of  
 854 the studied samples from Middle Gobi Belt. Solidus curves from Manning (1981).
- 855 Figure 11. (a) (Na<sub>2</sub>O+K<sub>2</sub>O)/(FeO<sup>T</sup>+TiO<sub>2</sub>) versus Na<sub>2</sub>O+K<sub>2</sub>O+FeO<sup>T</sup>+MgO+TiO<sub>2</sub>, wt% diagram, (b) molar  
 856 Al<sub>2</sub>O<sub>3</sub>/(Mg+FeO<sup>T</sup>) versus molar CaO/(MgO+FeO<sup>T</sup>) for source determination, (c) Plots of Rb/Ba  
 857 versus Rb/Sr, (d) CaO/Na<sub>2</sub>O versus Al<sub>2</sub>O<sub>3</sub>/TiO<sub>2</sub> plots for source characteristics ([Patiño Douce and](#)  
 858 [Johnston, 1991](#); [Patiño Douce and Harris, 1998](#)).
- 859 Figure 12. Correlations between whole-rock ε<sub>Nd</sub>(t) and zircon concordant ages. (b) Correlations between  
 860 ε<sub>Hf</sub>(t) and U–Pb ages of zircons for the studied samples from the Middle Gobi Belt. Compared  
 861 data from [Zhu et al., 2016](#); [Zhao et al., 2017](#).
- 862 Figure 13. (a) MgO–Al<sub>2</sub>O<sub>3</sub>–FeO<sup>T</sup> ternary diagram for tectonic discrimination of mafic rocks from the  
 863 Middle Gobi Belt ([Pearce et al. 1977](#)); (b) Ba versus Zr tectonic discrimination diagram of mafic  
 864 rocks from the Middle Gobi Belt ([Shervais, 1982](#)).
- 865 Table 1. Major (wt%) and trace (µg/g) element compositions including sample location and rock type of  
 866 the studied samples from the Middle Gobi Belt.
- 867 Table 2. LA-ICPMS U–Th–Pb analytical data for zircons of the studied samples from the Middle Gobi  
 868 Belt. \* Discordant data excluded from calculation.
- 869 Table 3. Zircon Lu–Hf and whole-rock Sm–Nd isotope data of the studied samples from the Middle Gobi



Fig. 1

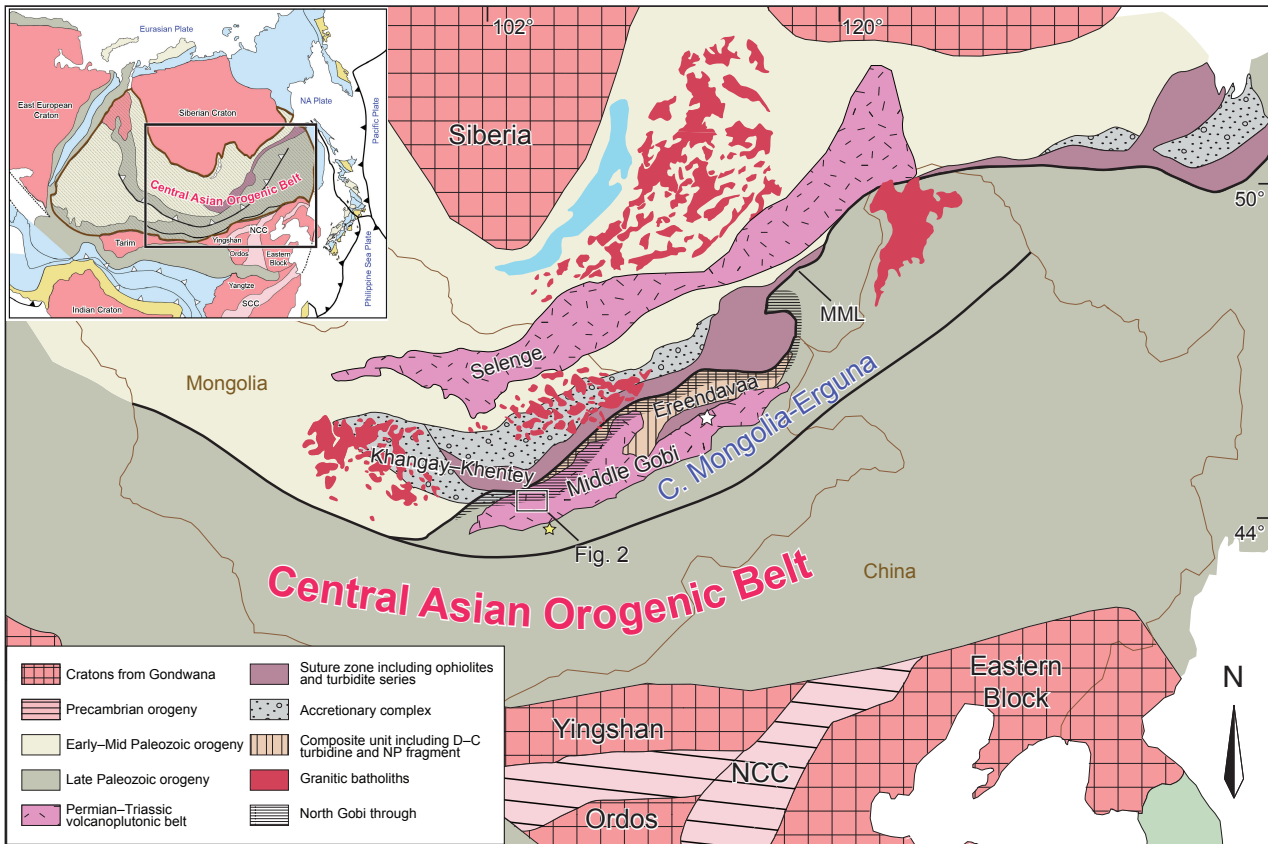


Fig. 2

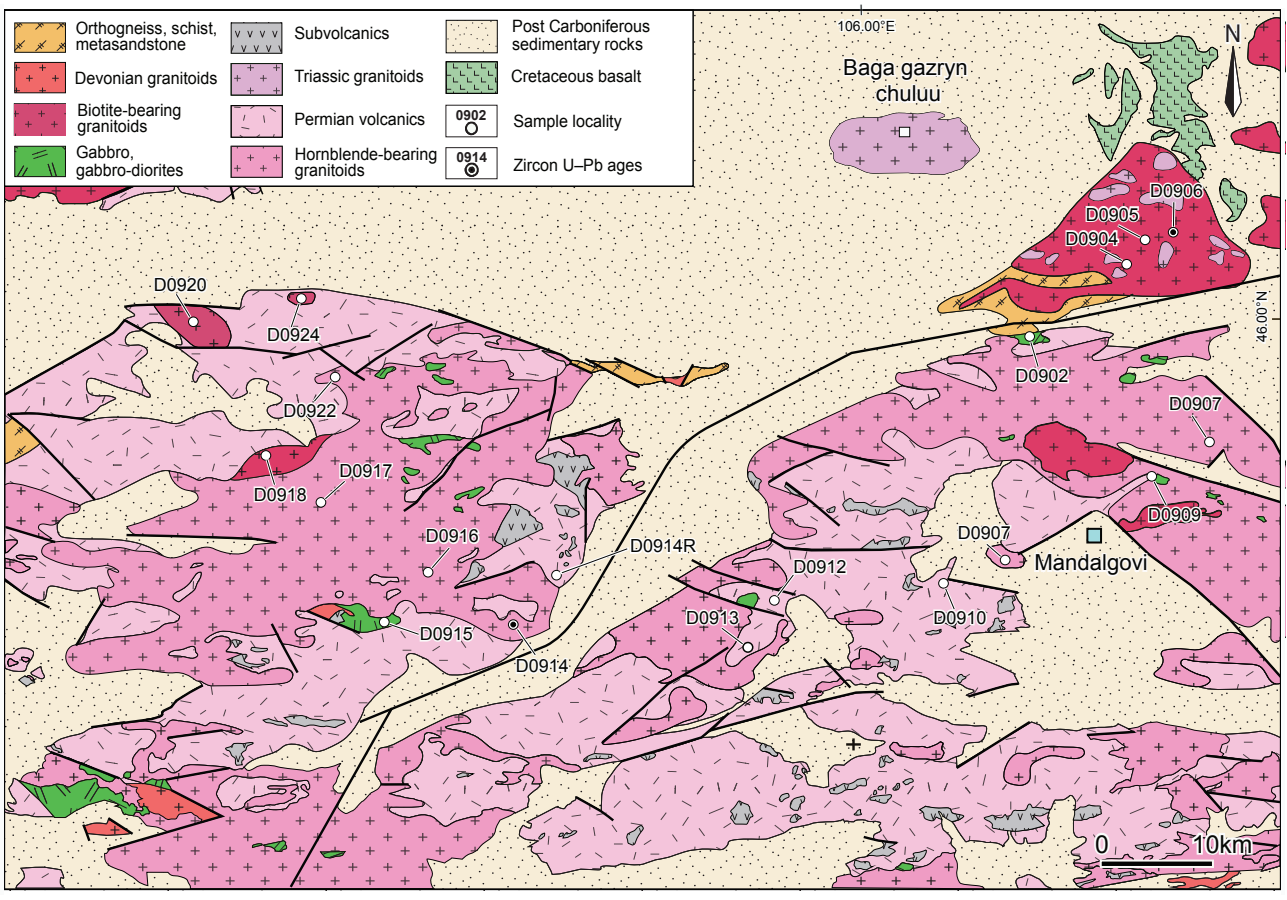


Fig. 3

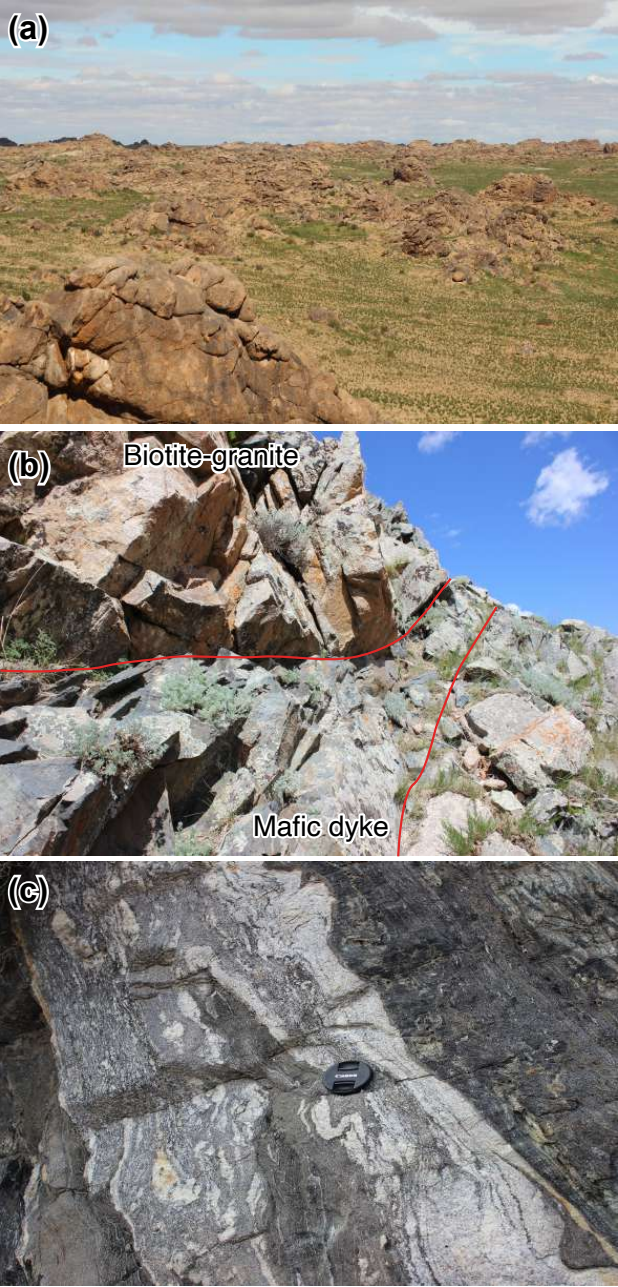


Fig. 4

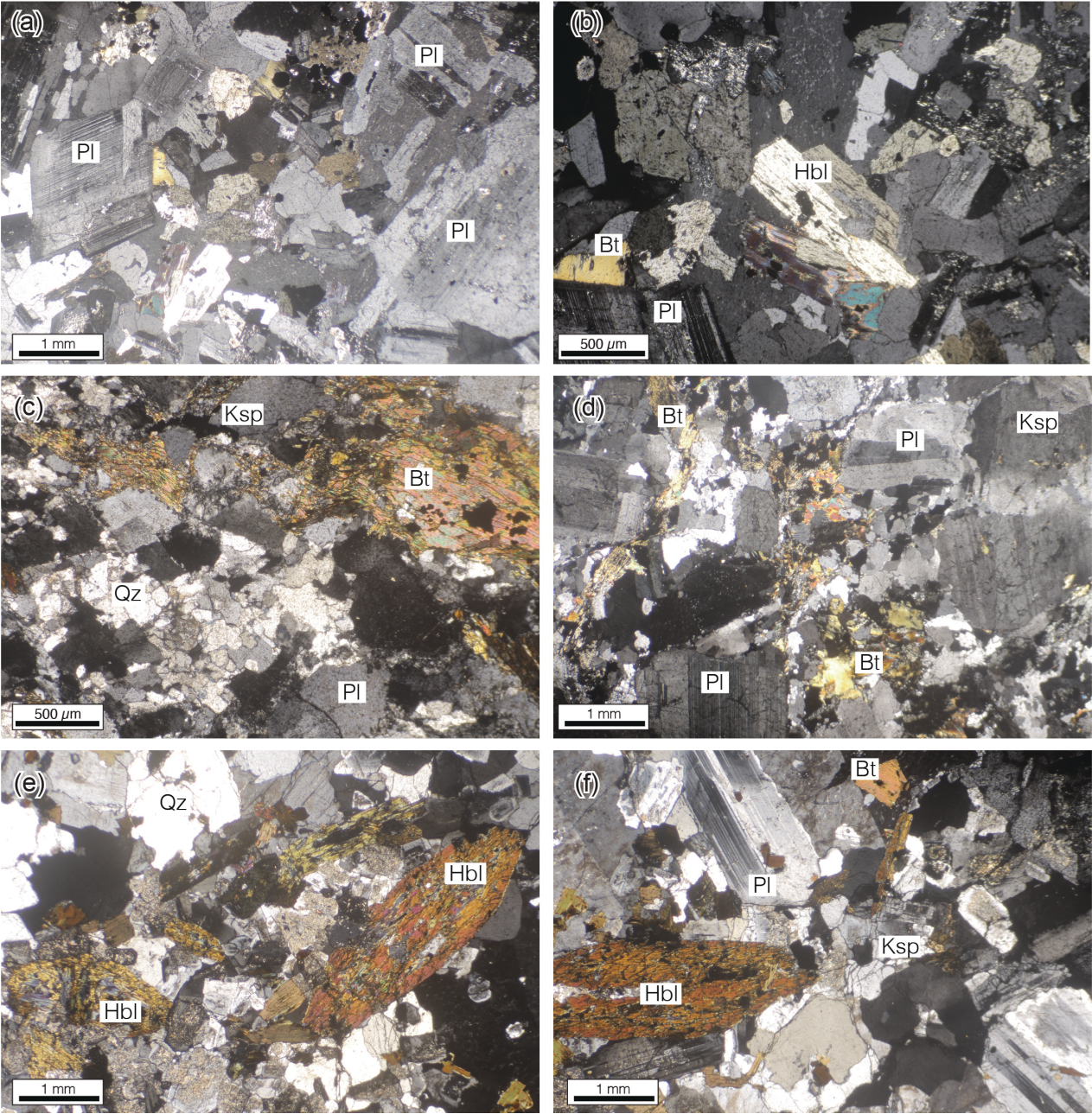


Fig. 5

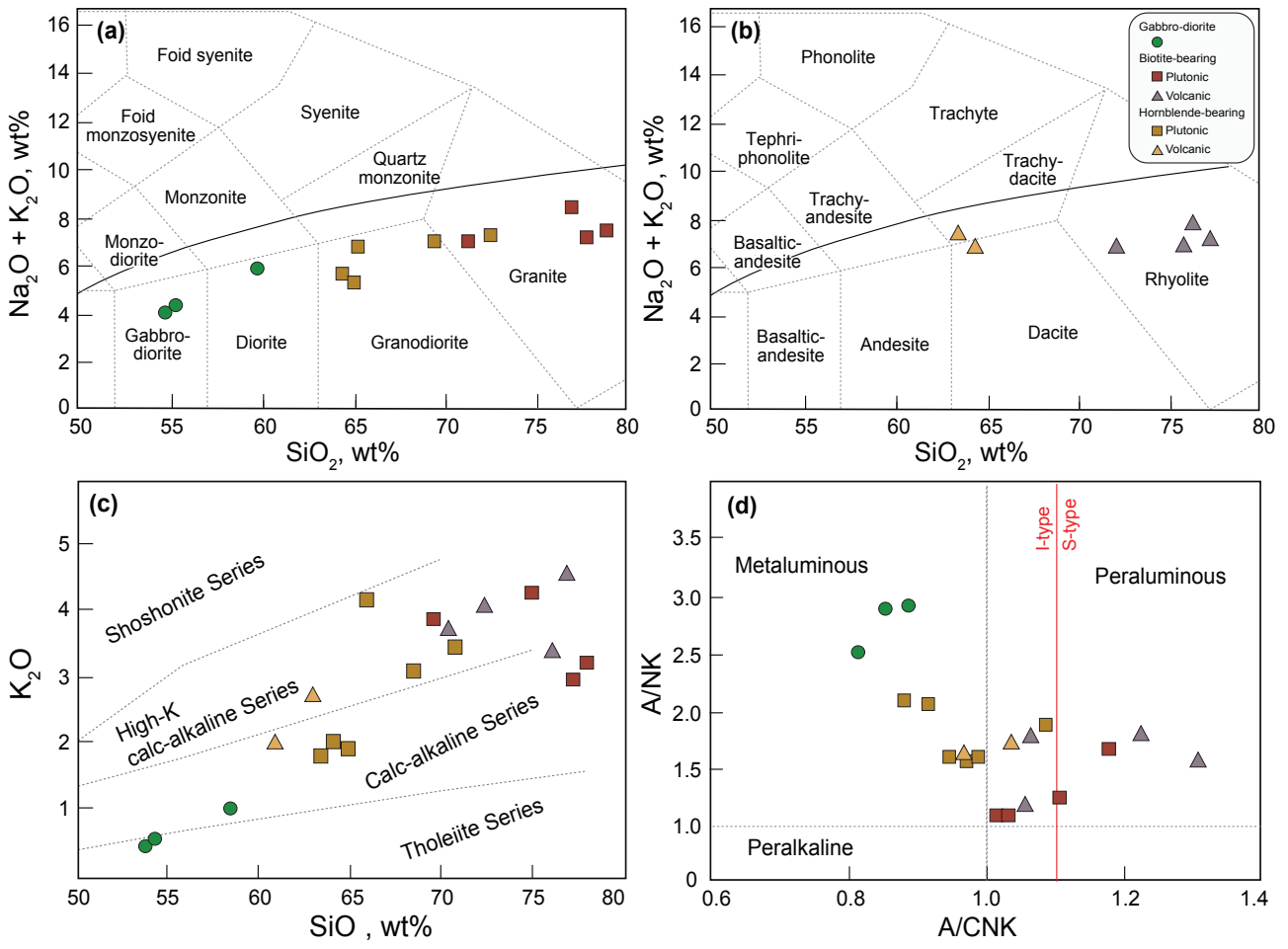


Fig. 6

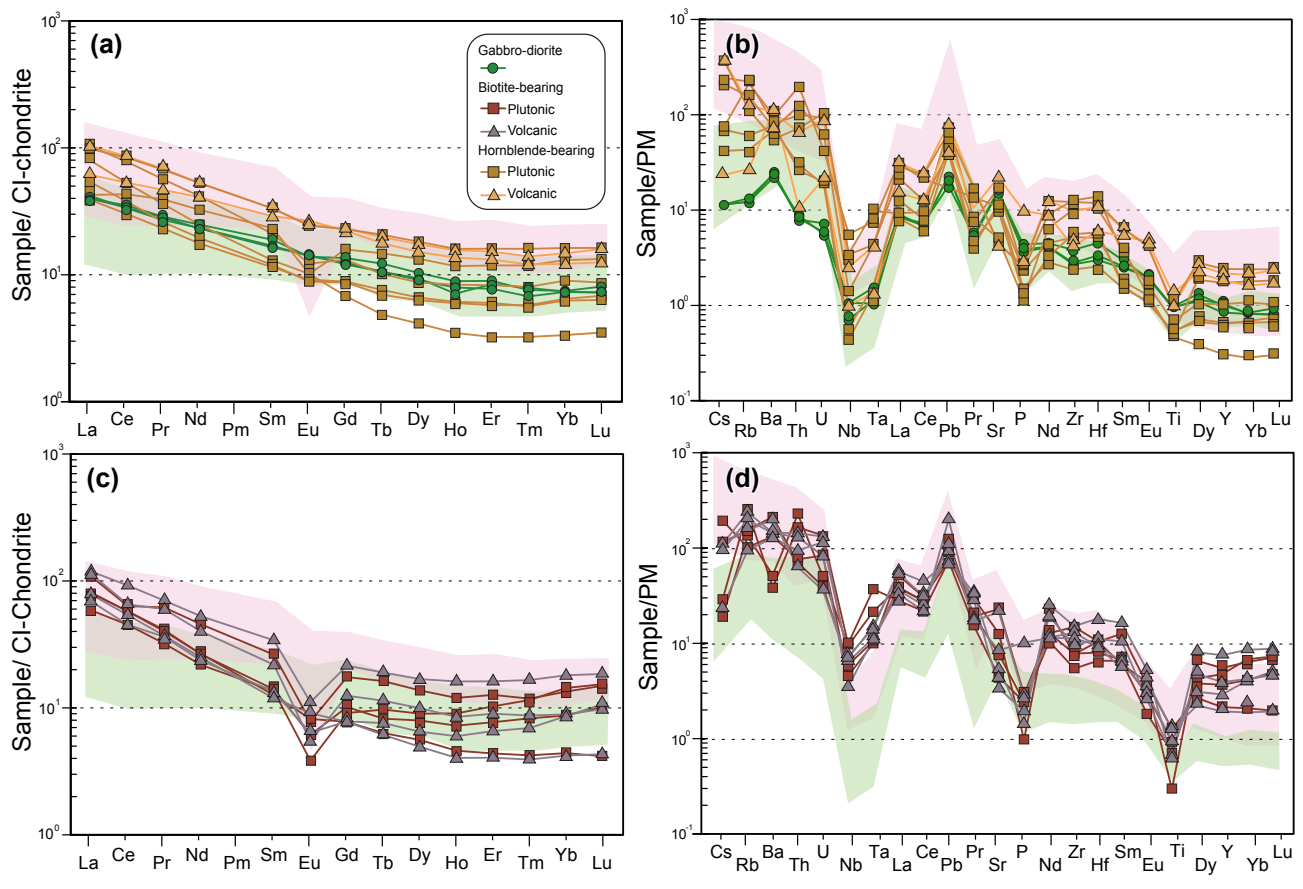


Fig. 7

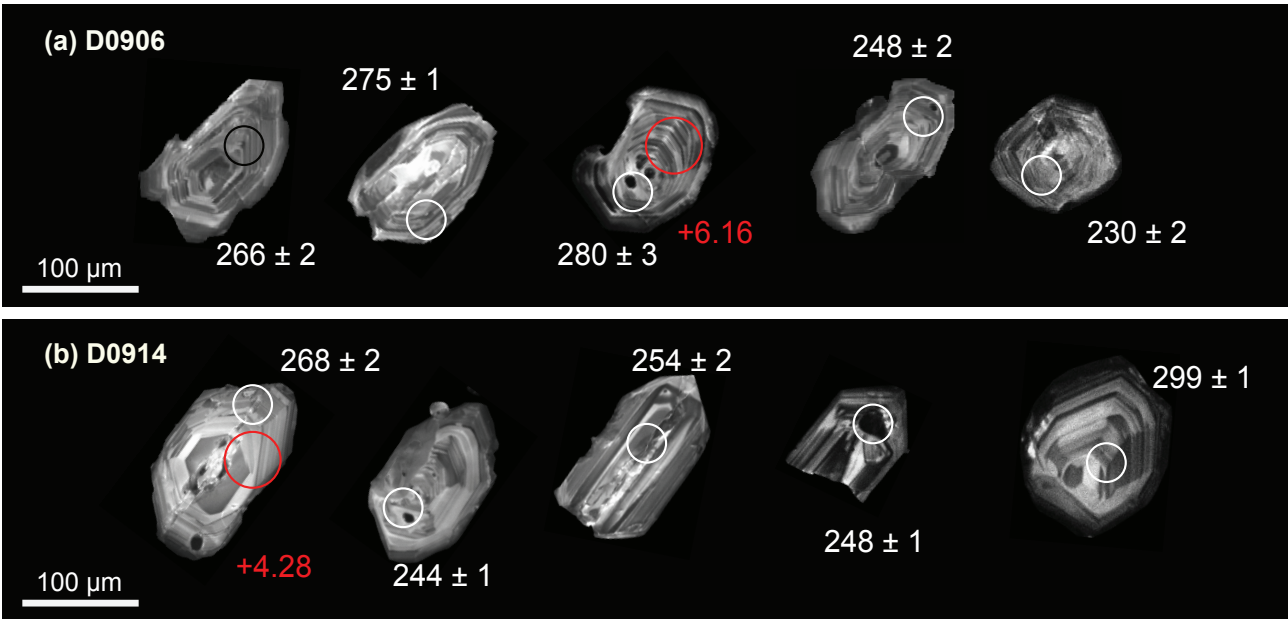


Fig. 8

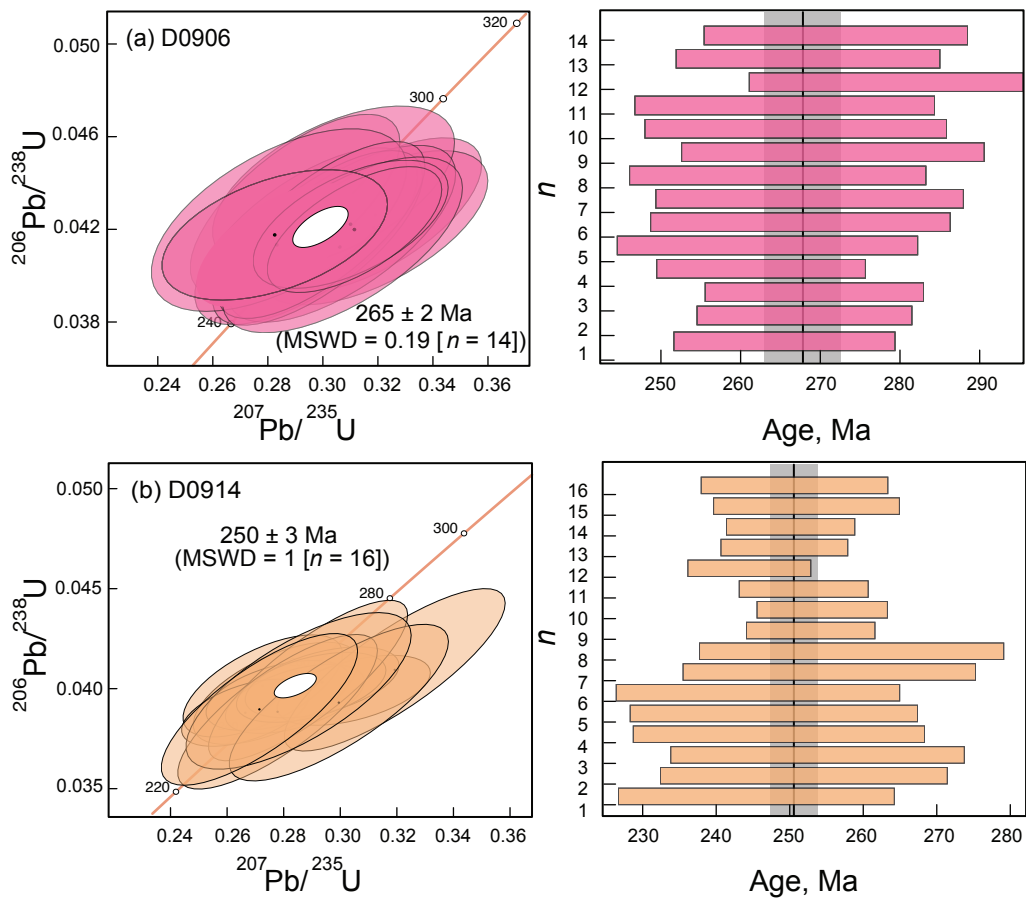




Fig. 9

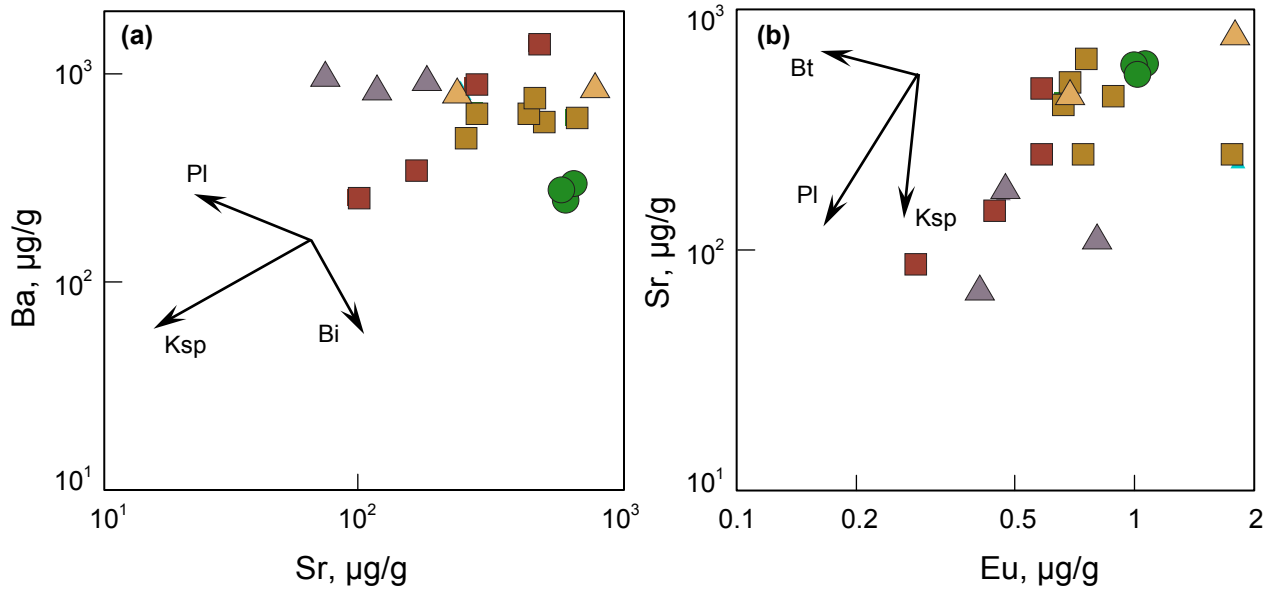


Fig. 10

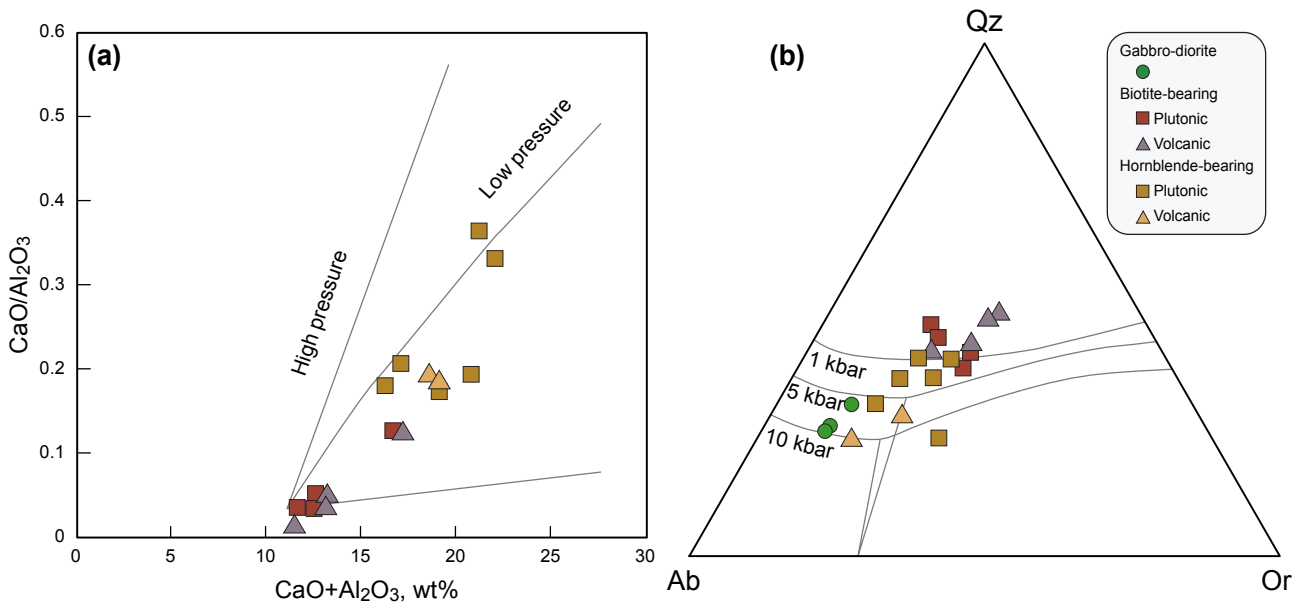


Fig. 11

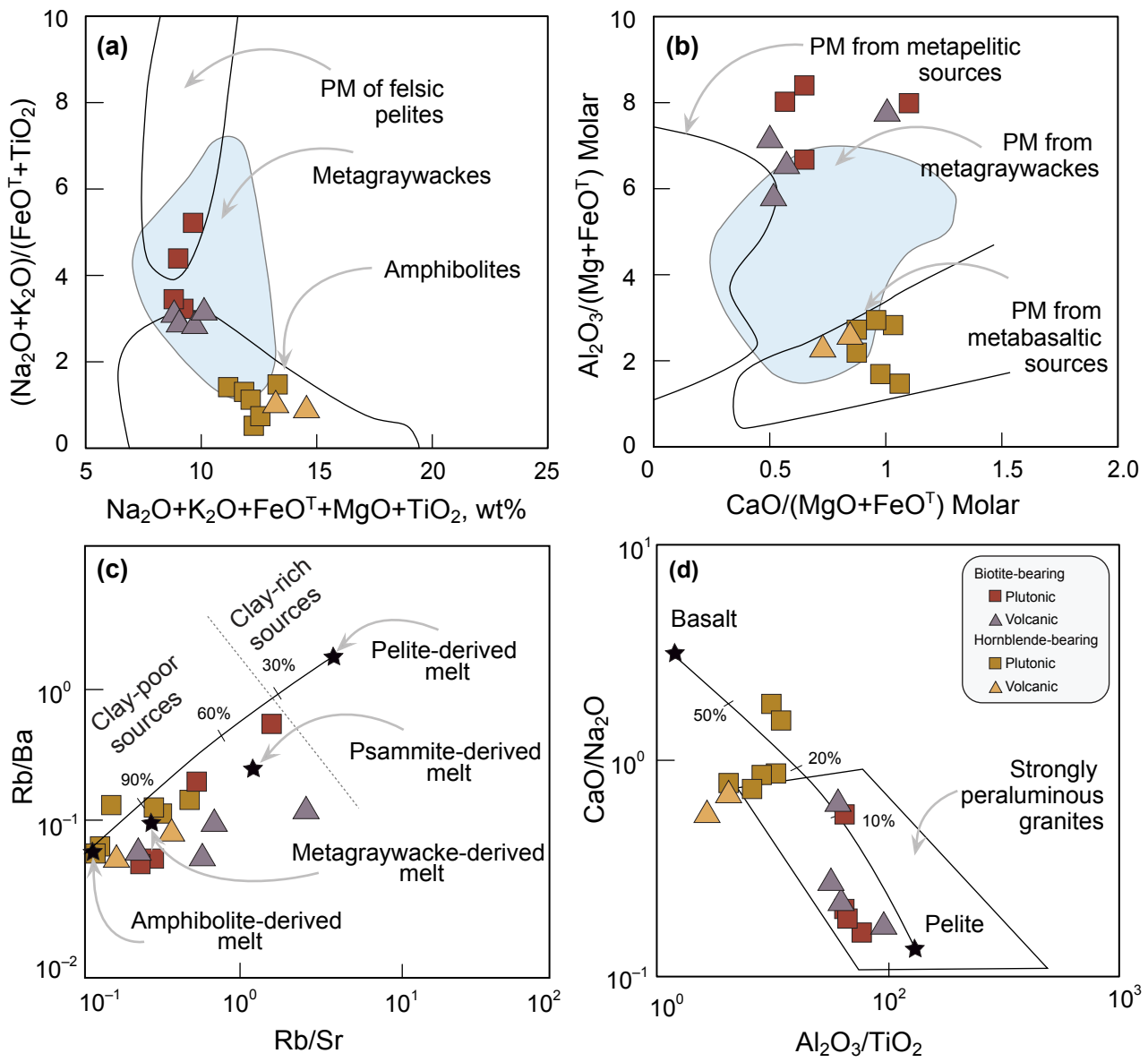


Fig. 12

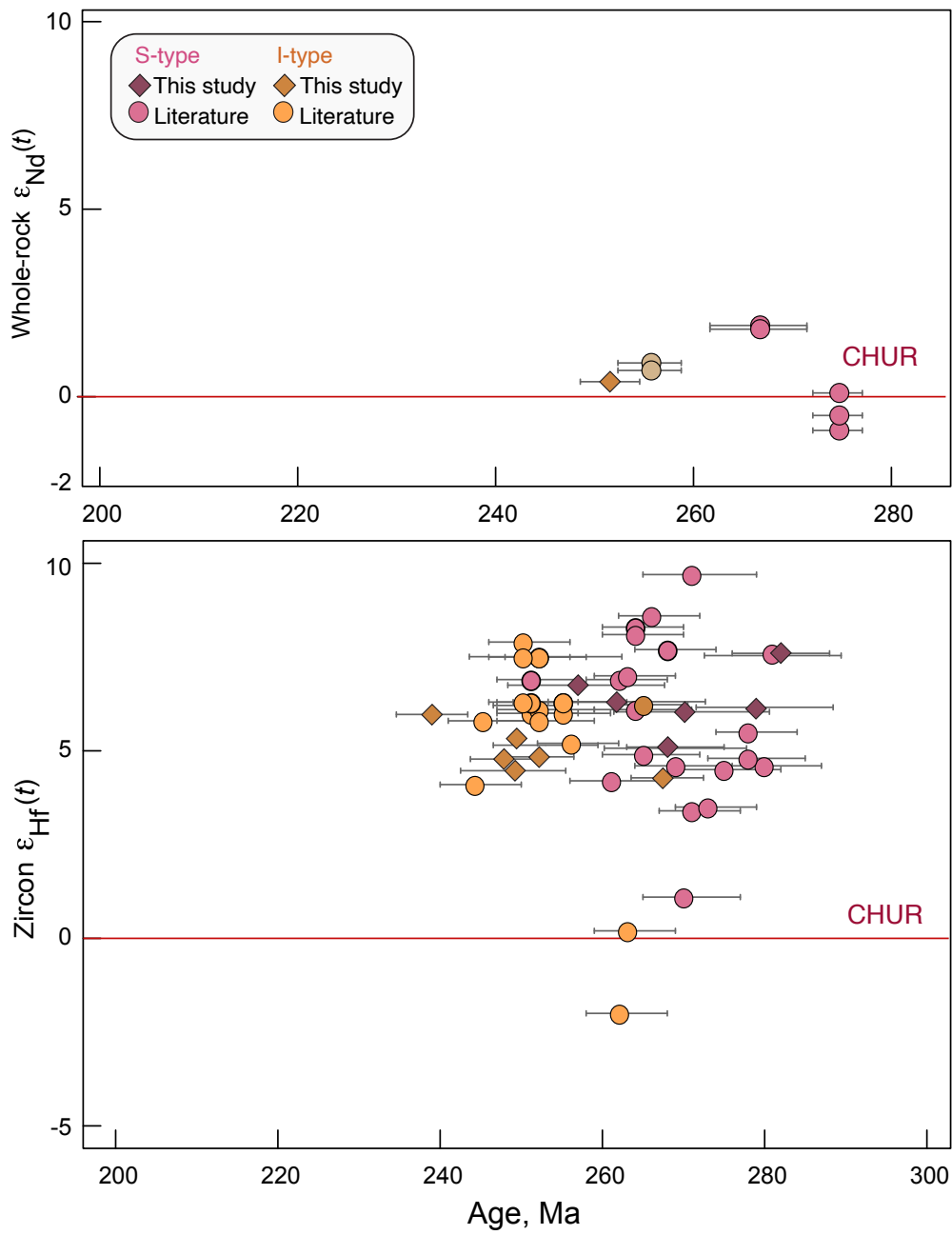


Fig. 13

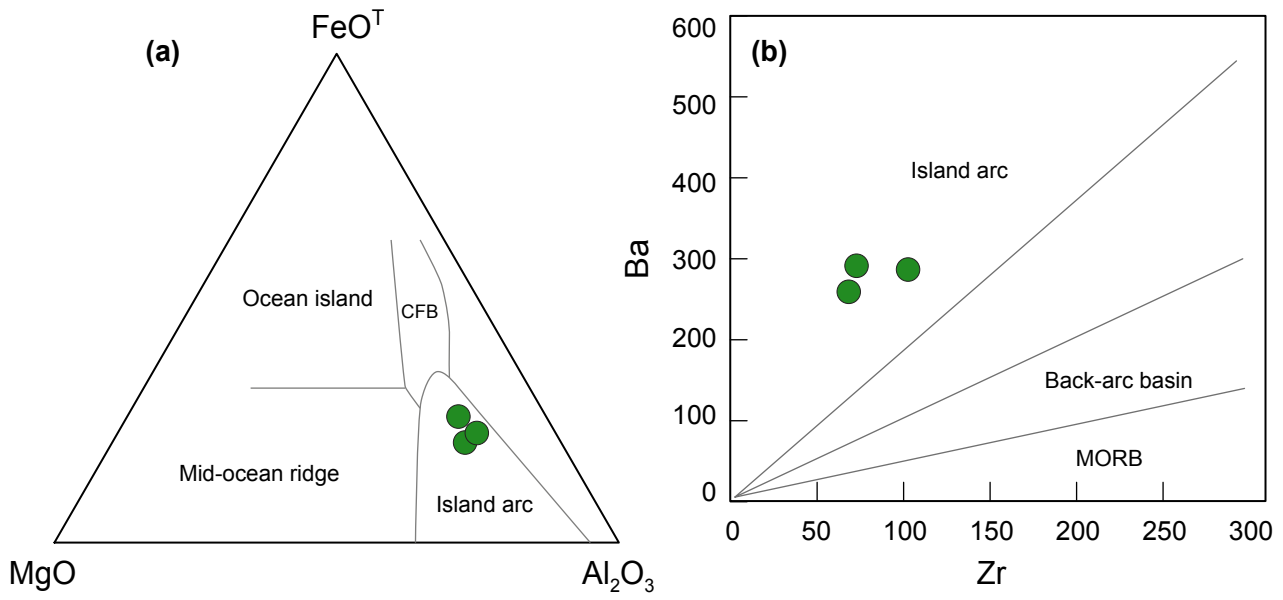


Table 1

Sample	D0902	D0904	D0905	D0907	D0909	D0910	D80914	D0915	D0917	D0920	D0912	D0913	D0909-2	D0914-R	D0918-1	D0918-2	D0922-1	D0922-2	D0924
SiO <sub>2</sub>	54.06	77.39	74.75	70.75	63.01	72.53	65.23	58.74	68.59	63.53	61.1	76.93	54.12	76.22	69.82	70.48	64.19	64.83	0.059
TiO <sub>2</sub>	0.792	0.188	0.146	0.471	0.796	0.191	0.61	0.774	0.464	0.506	1.006	0.136	0.793	0.124	0.255	0.264	0.514	0.834	11.25
Al <sub>2</sub> O <sub>3</sub>	18.52	12	12.13	13.77	16.15	12.68	16.14	16.24	14.11	16	15.69	11.44	17.67	12.38	15.05	15.31	15.32	17.36	1.69
FeO	8.19	1.84	1.42	3.43	5.48	2.27	4.25	6.78	4.09	4.95	6.76	2.67	8.78	2.46	1.69	1.75	5.12	4.79	1.52
MnO	0.128	0.035	0.024	0.055	0.104	0.029	0.09	0.16	0.045	0.085	0.126	0.024	0.138	0.028	0.023	0.028	0.083	0.107	0.036
MgO	4.08	0.25	0.17	1.16	1.18	0.1	1.15	2.89	1.37	2.59	1.14	0.09	4.34	0.13	0.31	0.32	2.57	1.19	0.08
CaO	8.05	0.6	0.45	2.49	2.97	0.44	2.93	5.75	2.93	5.31	2.97	0.1	8.51	0.57	1.88	1.88	5.49	3.36	0.4
Na <sub>2</sub> O	3.26	3.98	3.89	3.44	4.33	2.72	4.22	3.98	3.79	3.45	5.22	2.62	3.15	4.31	3.14	3.06	3.09	4.64	4.08
K <sub>2</sub> O	0.77	2.99	4.32	3.44	2.77	4.07	3.99	2.79	3.07	1.91	2.02	4.53	0.79	3.36	3.85	3.72	2.06	2.01	3.25
P <sub>2</sub> O <sub>5</sub>	0.21	0.05	0.05	0.09	0.17	0.21	0.15	0.15	0.15	0.11	0.44	0.04	0.24	0.03	0.06	0.06	0.1	0.16	0.02
LOI	2.42	1.43	1.23	1.47	1.36	3.4	1.02	2.42	1.22	1.62	3.99	2.02	2.25	0.72	3.61	3.5	1.61	1.52	1.28
Total	100.5	100.8	98.58	100.6	98.33	98.64	99.94	100.7	99.81	100.1	100.4	100.6	100.8	100.3	99.7	100.4	100.2	100.8	100.1
Rb	15	80	150	128	85	118	124	105	98	38	27	142	16	57	93	93	49	77	61
Sr	611	153	88	258	228	168	455	642	635	478	779	67	591	107	455	440	427	238	253
Y	15.3	15.8	20.4	26	30.8	12.5	17.8	25.6	7.3	12.8	27.2	16.9	17.4	32.3	9.2	8.8	12.4	33.4	23.7
Cs	0.5	0.4	0.6	4.8	6.6	2	2	1.9	4.3	1.3	1.2	2.3	0.5	0.5	2.4	2.4	1.9	6.5	4.1
Ba	256	344	255	665	804	948	759	484	614	592	852	955	271	844	1400	1339	643	505	875
La	12.4	24.8	18.2	13.2	32.3	21.4	32	36.2	24.8	12.1	19	35.1	12.3	36.8	25.1	24.2	16.7	30.5	33.6
Ce	27.3	46.7	36.5	34.8	69.2	36.5	63	76.6	42.8	24.4	43	52.8	28.5	74.5	46.5	43.6	30.6	67.6	51.3
Pr	3.46	5.02	3.92	4.8	8.38	4.43	6.9	8.68	4.41	2.83	5.69	7.31	3.5	8.54	4.89	4.52	3.18	8.22	7.54
Nd	13.9	15.9	13.1	19.5	31.7	14.2	25	30.6	14.9	10.6	24.5	24.1	14.9	31.7	16	15	11.7	31.7	27.4
Sm	3.2	2.73	2.65	4.54	6.51	2.62	4.1	5.93	2.53	2.23	5.64	4.27	3.68	6.64	2.46	2.32	2.33	6.49	5.12
Eu	1.05	0.45	0.282	0.75	1.82	0.468	0.9	1.03	0.748	0.673	1.96	0.402	1.02	0.815	0.602	0.668	0.649	1.79	0.575
Gd	3.13	2.63	2.36	4.1	5.99	2.03	3.4	4.84	1.76	2.29	5.76	3.23	3.51	5.57	2.01	2.01	2.26	5.94	4.53
Tb	0.5	0.39	0.46	0.7	0.94	0.36	0.48	0.78	0.23	0.36	0.84	0.54	0.58	0.9	0.3	0.29	0.33	0.98	0.77
Dy	2.92	2.56	2.9	4.31	5.59	2.09	2.8	4.67	1.33	2.13	4.87	3.2	3.33	5.43	1.8	1.59	2.03	5.86	4.42
Ho	0.57	0.52	0.65	0.84	1.12	0.43	0.6	0.92	0.25	0.44	0.98	0.61	0.64	1.16	0.33	0.29	0.43	1.15	0.86
Er	1.63	1.62	2.16	2.49	3.2	1.38	1.74	2.61	0.68	1.26	2.77	1.89	1.88	3.4	0.92	0.85	1.21	3.37	2.66
Tm	0.219	0.268	0.374	0.384	0.452	0.225	0.26	0.415	0.105	0.183	0.391	0.282	0.246	0.537	0.137	0.127	0.187	0.519	0.383
Yb	1.53	1.82	3	2.73	3.08	1.8	1.88	2.85	0.7	1.31	2.55	1.86	1.56	3.78	0.92	0.88	1.35	3.4	2.81
Lu	0.233	0.329	0.495	0.429	0.515	0.345	0.28	0.419	0.113	0.204	0.404	0.315	0.26	0.596	0.134	0.139	0.22	0.524	0.481
Pb	6	13	11	14	15	30		18	9	9	11	12	5	11	18	17	10	13	11
Th	1.56	18.1	13.2	9.42	6.91	10.4	15.7	13.6	11.1	3.87	1.79	11.5	1.45	7.58	5.59	5.16	3.49	7.46	6.1
U	0.28	1	2.71	2.24	2.19	1.65	1.25	1.75	1.68	0.68	0.76	2.63	0.34	2.3	0.79	0.74	0.71	2.43	1.67
Zr	73	82	57	227	223	119	268	75	139	70	118	100	111	159	156	142	147	261	81
Nb	2	6.2	6.5	6.5	5.1	2.3	9.2	8.4	3.3	1.4	2.6	4.6	2.7	4.8	3.9	4.4	1.7	6.4	3
Hf	1.9	2.7	2.2	7.1	5.5	3.3	6.4	2.4	3.8	1.9	3.4	3.2	3.1	4.9	3.8	3.9	3.8	6.7	2.7
Ta	0.16	0.8	1.37	0.78	0.42	0.41	0.65	0.6	0.43	0.18	0.18	0.56	0.15	0.53	0.39	0.44	0.21	0.43	0.37
Sc	22	4	4	8	14	2		17	5	14	21	2	24	4	2	2	14	15	3

Sample	D0902	D0904	D0905	D0907	D0909	D0910	D80914	D0915	D0917	D0920	D0912	D0913	D0909-2	D0914-R	D0918-1	D0918-2	D0922-1	D0922-2	D0924
Cr	40	<20	<20	30	<20	<20		30	50	50	<20	30	40	<20	<20	<20	40	<20	20
Co	22	2	1	7	7	1		16	8	14	7	<1	21	<1	1	<1	15	7	1
Ni	<20	<20	<20	<20	<20	<20		<20	<20	20	<20	<20	<20	<20	<20	<20	<20	<20	20
Zn	80	<30	<30	40	70	30		90	30	50	110	<30	80	40	40	<30	50	70	30
Ga	19	12	13	15	21	14		20	18	16	18	13	19	16	19	19	16	23	10
Tl	0.05	0.19	0.52	0.47	0.48	0.46		0.41	0.54	0.21	0.2	0.54	0.1	0.73	0.42	0.49	0.15	0.33	0.41

Table 2

Sample name	Isotope ratios						Age							
	$^{207}\text{Pb}/^{235}\text{U}$	$2\sigma$	$^{206}\text{Pb}/^{238}\text{U}$	$2\sigma$	$^{207}\text{Pb}/^{206}\text{Pb}$	$2\sigma$	$^{235}\text{U}-^{207}\text{Pb}$	$2\sigma$	$^{238}\text{U}-^{206}\text{Pb}$	$2\sigma$	$^{232}\text{Th}-^{208}\text{Pb}$	$2\sigma$	$^{207}\text{Pb}-^{206}\text{Pb}$	$2\sigma$
D0906-2	0.31903	0.01579	0.04262	0.00136	0.054288793	0.002052018	281.2	12.2	269.05	8.41	382	85	269.6	8.43
D0906-6	0.31848	0.01694	0.04207	0.00136	0.054905578	0.002324074	280.7	13	265.65	8.41	407.4	94.7	266.11	8.44
D0906-7	0.30387	0.01797	0.04382	0.00143	0.050288525	0.002482098	269.4	14	276.47	8.83	207	114	275.93	8.77
D0906-8	0.29395	0.01654	0.04170	0.00156	0.051124201	0.002148972	261.7	13	263.36	9.65	245.4	96.8	263.18	9.59
D0906-9	0.30686	0.01691	0.04181	0.00156	0.053230771	0.002156122	271.7	13.1	264.04	9.65	337.6	91.8	264.58	9.66
D0906-10	0.28987	0.01557	0.04305	0.00160	0.04883147	0.001892638	258.5	12.3	271.71	9.89	138.7	91	269.23	9.69
D0906-12	0.29598	0.01560	0.04151	0.00154	0.051715178	0.001932478	263.3	12.2	262.19	9.53	271.8	85.7	262.31	9.49
D0906-14	0.28405	0.01896	0.04240	0.00161	0.048585739	0.002664285	253.9	15	267.69	9.96	127	129	266.31	9.85
D0906-18	0.29943	0.01538	0.04197	0.00156	0.051744941	0.001838237	266	12	265.03	9.65	273.1	81.4	265.15	9.59
D0906-20	0.30614	0.01782	0.04124	0.00155	0.053836313	0.002393685	271.2	13.9	260.51	9.6	363	100	261.02	9.62
D0906-24	0.28327	0.01194	0.04133	0.00109	0.049713351	0.001635493	253.25	9.45	261.07	6.75	180.6	76.7	260.22	6.68
D0906-26	0.31008	0.01448	0.04222	0.00113	0.053261221	0.002040742	274.2	11.2	266.58	6.99	338.9	86.8	266.91	6.99
D0906-27	0.31148	0.01296	0.04199	0.00111	0.053794915	0.001731426	275.3	10	265.16	6.87	361.5	72.6	265.67	6.88
D0906-28	0.28250	0.01675	0.04176	0.00115	0.049061792	0.002574262	252.6	13.3	263.73	7.12	150	123	263.15	7.08
D0914-1	0.27719	0.01192	0.03968	0.00105	0.050659272	0.001719559	248.43	9.48	250.85	6.51	224.3	78.5	250.65	6.48
D0914-5	0.27261	0.01152	0.04005	0.00105	0.049371632	0.001630586	244.78	9.19	253.14	6.51	164.5	77.2	252.28	6.45
D0914-7	0.27792	0.01037	0.03957	0.00072	0.050933554	0.00166242	249.01	8.24	250.17	4.46	236.8	75.3	250.12	4.45
D0914-8	0.28394	0.01004	0.03940	0.00071	0.052260997	0.001592805	253.78	7.94	249.11	4.4	295.8	69.6	249.27	4.4
D0914-9	0.26643	0.00920	0.03870	0.00069	0.049936533	0.001477581	239.84	7.38	244.77	4.28	191	68.8	244.51	4.27
D0914-10	0.30477	0.01106	0.03983	0.00072	0.055494657	0.001744588	270.12	8.61	251.78	4.46	431.2	70.1	251.9	4.48
D0914-14	0.29357	0.01111	0.04023	0.00073	0.052924213	0.00175574	261.37	8.72	254.26	4.52	324.5	75.3	254.42	4.52
D0914-15	0.28971	0.01080	0.03998	0.00072	0.052559257	0.001711483	258.33	8.5	252.71	4.46	308.8	74.1	252.84	4.46
D0914-20	0.31925	0.01595	0.04086	0.00167	0.056665203	0.001623141	281.3	12.3	258.2	10.3	477.5	63.3	258.4	10.6
D0914-22	0.28861	0.01438	0.04032	0.00165	0.051910446	0.001478487	257.5	11.3	254.8	10.2	280.4	65.2	255.4	10.2
D0914-24	0.27784	0.01442	0.03875	0.00159	0.051998939	0.001655886	248.9	11.5	245.08	9.87	284.3	72.8	245.68	9.84
D0914-25	0.28179	0.01524	0.03910	0.00161	0.052264812	0.001834673	252.1	12.1	247.25	9.99	296	80.1	247.84	9.97
D0914-29	0.29948	0.01583	0.03921	0.00161	0.055399705	0.001841826	266	12.4	247.93	9.99	427.4	74.1	248.5	10.1
D0914-33	0.32260	0.01746	0.04028	0.00162	0.058082313	0.002102738	283.9	13.4	254.6	10	531.9	79.3	253.8	10.2
D0914-34	0.28526	0.01626	0.03980	0.00161	0.051988824	0.002093598	254.8	12.8	251.59	9.98	283.9	92.1	251.93	9.95
D0914-35	0.27142	0.01418	0.03887	0.00156	0.050638598	0.001697749	243.8	11.3	245.82	9.68	223.4	77.5	245.46	9.57



Table 3

Sample No.	Age, Ma	$^{176}\text{Yb}/^{177}\text{Hf}$	$^{176}\text{Lu}/^{177}\text{Hf}$	$^{176}\text{Hf}/^{177}\text{Hf}$	2s	$\varepsilon_{\text{Hf}}(t)$	$T_{\text{DM}}$	$T_{\text{DMC Lu/Hf}=0.015}$	$(^{176}\text{Hf}/^{177}\text{Hf})_i$	$^{176}\text{Hf}/^{177}\text{Hf}(\text{DM})_i$	$f_{\text{Lu/Hf}}$	2s
D0914												
7.1	249	0.020205	0.000720	0.282747	0.000020	4.47	711	994	0.282743	0.283070	-0.98	0.6904
8.1	253	0.028939	0.001043	0.282768	0.000020	5.15	687	950	0.282763	0.283071	-0.97	0.6957
10.1	251	0.018934	0.000679	0.282795	0.000020	6.21	642	883	0.282792	0.283070	-0.98	0.7096
11.1	239	0.023660	0.000881	0.282796	0.000026	5.97	644	889	0.282792	0.283078	-0.97	0.9200
24.1	248	0.021468	0.000783	0.282758	0.000023	4.78	695	970	0.282755	0.283074	-0.98	0.7974
25.1	252	0.028522	0.001020	0.282760	0.000023	4.84	697	968	0.282755	0.283073	-0.97	0.8254
29.1	266	0.017429	0.000632	0.282797	0.000023	6.23	639	880	0.282794	0.283072	-0.98	0.7990
31.1	268	0.020348	0.000745	0.282730	0.000022	4.28	734	1020	0.282727	0.283057	-0.98	0.7759
D0906												
1.1	281	0.039100	0.001305	0.282853	0.000025	7.54	570	775	0.282848	0.283091	-0.96	0.8850
6.1	280	0.021466	0.000780	0.282785	0.000023	6.16	658	898	0.282781	0.283059	-0.98	0.7979
7.1	269	0.030064	0.001016	0.282750	0.000024	5.07	711	974	0.282745	0.283053	-0.97	0.8604
9.1	271	0.024690	0.000834	0.282783	0.000024	6.04	662	905	0.282779	0.283060	-0.97	0.8364
10.1	258	0.022115	0.000793	0.282800	0.000022	6.75	637	863	0.282796	0.283057	-0.98	0.7653
11.1	263	0.026817	0.000898	0.282820	0.000023	6.31	611	850	0.282816	0.283095	-0.97	0.8222
13.1	253	0.030760	0.001074	0.282825	0.000025	7.51	607	812	0.282819	0.283059	-0.97	0.8887