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On the modulation of kinetic energy transfer by internal gravity waves

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Key Points:

- Twin submesoscale-permitting simulations of the North Atlantic Ocean were run with and without tidal forcing.
- The tidally-forced run shows enhanced internal gravity wave (IGW) signal during summer while the non-tidally forced run during winter.
- Tidally-forced IGWs (i.e. internal tides) enhance forward cascade of kinetic energy particularly during summertime.

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21 **Abstract**

22 Understanding how kinetic energy (KE) is exchanged across scales and eventually dissipated
23 remains a key question in physical oceanography. Recent theoretical works suggests that
24 both quasi-balanced submesoscale motions and internal gravity waves (IGWs) could play
25 a role in fluxing KE towards dissipation. How these classes of motions actually provide a
26 route to dissipation in the ocean is still debated. This study investigates the impact of IGWs
27 generated by tidal motions on cross-scale KE exchanges at mid-latitude. Our analysis is
28 based on the output of two realistic submesoscale permitting ocean model simulations of the
29 North Atlantic Ocean, run respectively with and without tidal forcing. These twin experi-
30 ments permit investigation of how tidally-forced IGWs modify the KE variance, cross-scale
31 exchanges, and associated seasonality. Our results show that, in the presence of externally-
32 forced IGWs, KE transfer towards dissipative scales is enhanced in summertime both at the
33 surface and in the ocean interior.

34 **Plain Language Summary**

35 Energetic oceanic currents on the scales of tens of kilometers that emerge as a re-
36 sult of the chaotic nature of the ocean, known as (sub)mesoscale flows, have been of great
37 scientific interest to the oceanographic community. These currents are associated with
38 the time scales of roughly a day, which overlap with the astronomical tidal frequency of
39 the ocean. Due to their similar time scales, it has been argued that the (sub)mesoscale
40 flows and tides would interact with each other. Here, by running a twin numerical sim-
41 ulation of the North Atlantic Ocean, one with tides and the other without, we examine
42 the physical interaction between the two and show that the tides stimulate a loss of en-
43 ergy from the (sub)mesoscale flows particularly during the summer.

44 **1 Introduction**

45 The ocean's kinetic energy (KE) is mostly concentrated in motions close to geostrophic
46 balance, with frequencies lower than the Coriolis frequency (f) and spatial scales larger
47 than the first Rossby deformation radius (R_d ; Vallis, 2017). These balanced motions (BMs)
48 are largely energized through baroclinic instability which extracts energy from large scale
49 stratification. Balanced motions include large-scale motions (> 300 km), mesoscale motions
50 (50–300 km) and submesoscale balanced motions (< 50 km) (McWilliams, 2016). Balanced
51 motions are characterized by an inverse cascade of energy (Scott & Wang, 2005; Scott &
52 Arbic, 2007; Eden, 2007; Aluie et al., 2017), so they do not provide a route to dissipation by
53 themselves. Therefore, energy has to be transferred from balanced motions to high-frequency

54 unbalanced motion for dissipation to occur. To equilibrate the well-known inverse cascade of
55 energy, the ocean requires ageostrophic processes to extract energy from balanced motions.
56 Mechanisms that might trigger a forward transfer of energy from balanced motions down to
57 dissipate scale include (i) bottom boundary-layer turbulence (Wunsch & Ferrari, 2004; Sen
58 et al., 2008; Arbic et al., 2009), (ii) generation of lee waves by mesoscale eddies interacting
59 with topography (Nikurashin & Ferrari, 2010; Nikurashin et al., 2013; Trossman et al.,
60 2013, 2016), (iii) generation of internal waves by upper-ocean frontal instabilities (Danioux
61 et al., 2012; Shakespeare & Taylor, 2014) and (iv) direct cascade of energy by energetic
62 submesoscale motions (Capet et al., 2008; Ferrari & Wunsch, 2009; Molemaker et al., 2010;
63 McWilliams, 2016). These studies have shown that, at fine-scale, submesoscale motions and
64 internal gravity waves (IGWs) can provide efficient transfer of energy to dissipative scales.

65 Submesoscale motions are flow structures in the form of density fronts, filaments and
66 topographic wakes at the surface and throughout the interior at scale smaller than $O(100\text{ km})$
67 (McWilliams, 2016). IGWs are a particular class of fast propagating unbalanced motions
68 with frequencies equal to or higher than the Coriolis frequency f and spatial scale smaller
69 than $O(100\text{ km})$. IGWs include wind-induced near-inertial waves with frequency near the
70 Coriolis frequency and internal tides generated by large-scale barotropic tidal flow over
71 topographic features. Near-inertial waves are usually stronger in winter than in summer
72 because they are driven by surface winds (D’Asaro, 1985) while the signature of internal
73 tides (i.e. IGWs with tidal frequencies) is typically stronger in summer time (Gerkema et
74 al., 2004). Recent works have highlighted that these two classes of motions—submesoscale
75 motions and IGWs—are out of seasonally phase (Rocha et al., 2016). Submesoscale motions
76 tend to be stronger than IGWs in wintertime. The emergence of submesoscales is due
77 to winter favored mechanisms such as mixed layer instability and wind-induced frontal
78 instability (e.g. Qiu et al., 2014; Sasaki et al., 2014; Callies, Flierl, et al., 2015; Brannigan
79 et al., 2015; McWilliams, 2016; Uchida et al., 2019; Khatri et al., 2021). KE associated
80 with IGWs shows stronger amplitude in summertime. This is due to the intensification of
81 vertical normal modes and shallow mixed layer (Callies, Ferrari, et al., 2015; Rocha et al.,
82 2016).

83 A forward cascade is seen in spectral KE fluxes computed from gridded satellite altimeter
84 products (Scott & Wang, 2005) at scales smaller than $O(100\text{ km})$. Later work (Arbic
85 et al., 2013) argued that while this forward cascade could be physically based, and related
86 to energy transfer between mesoscale eddies and IGWs, it could also be due to the spatial
87 and temporal smoothing inherent in the construction of gridded altimeter products. Due to
88 high-frequency winds and tidal motions, energy extraction via IGWs seems to be a highly
89 probable mechanism of KE sinks for balanced motions (J. Thomas & Daniel, 2021). There

90 have so far been two well documented energy transfer mechanisms from balanced motions to
 91 IGWs away from topography, namely the *stimulated generation* of a forward cascade of KE
 92 by near-inertial waves from balanced flows (Gertz & Straub, 2009; Rocha et al., 2018; Barkan
 93 et al., 2021), and *spontaneous generation* of near-inertial waves from balanced flows (Nagai
 94 et al., 2015; Shakespeare & Hogg, 2017). In stimulated generation, near-inertial waves are
 95 first introduced by external forcing (e.g. wind) and then grow by extracting energy from
 96 the balanced flow (Gertz & Straub, 2009; Barkan et al., 2017; L. N. Thomas, 2017) while
 97 spontaneous generation is the emission of waves by unbalanced, large Rossby number flow
 98 at density fronts without external forcing. These waves then radiate vertically downwards
 99 into the interior and amplify by extracting energy from deep balanced flow (Shakespeare
 100 & Hogg, 2017). Spontaneous generation is localized at sharp submesoscale fronts and is
 101 inefficient at small Rossby numbers (Danioux et al., 2012; Nagai et al., 2015; Shakespeare &
 102 Hogg, 2017). On the other hand, stimulated generation is efficient at small Rossby numbers
 103 provided that the waves are forced externally.

104 As listed above, several studies have focused on the impact of wind-generated near-
 105 inertial waves on energy dissipation (e.g. Barkan et al., 2017; J. Thomas & Arun, 2020). In
 106 contrast, little is known as to the role of internal tides on KE exchanges (cf. Barkan et al.,
 107 2021). We know that internal tides contribute to the building up of IGW continuum (Garrett
 108 & Munk, 1975) and that they eventually contribute to diapycnal mixing in the ocean interior
 109 (Arbic et al., 2018). Whether internal tides could play a significant role in the down-scale
 110 transfer of mesoscale eddy KE is yet to be fully explored. In this context, we investigate
 111 the role of internal tides on cross-scale KE exchanges in a regime with active submesoscale
 112 motions by using a twin submesoscale resolving numerical simulations (with/without tides)
 113 of the North Atlantic Ocean with a horizontal resolution of $1/60^\circ$. This permits us to
 114 investigate how IGWs affect KE exchanges in the presence of active submesoscales motions.

115 This paper is organized as follows: in the next section, we describe the numerical
 116 simulations. Section 3 presents the seasonality of BMs and IGWs characterized by the KE
 117 frequency-wavenumber spectral density. The contribution of BMs and IGWs to the KE
 118 transfer is presented in section 4. We discuss the impacts of this observed seasonality on
 119 the KE spectral flux in section 5, and we identify two different mechanisms that support a
 120 direct cascade of energy in a dynamical regime with/without tidal motions.

121 2 North Atlantic Ocean Simulation

122 In this study, we use numerical outputs from a NEMO-based submesoscale eddy-permitting
 123 simulations of the North Atlantic with a horizontal resolution of $1/60^\circ$ (eNATL60; Brodeau
 124 et al., 2020). eNATL60 is a spatially-extended version of NATL60 (Ajayi et al., 2020,

2021). The simulation spans the North Atlantic Ocean from about 6°N up to the polar circle and has a horizontal grid spacing ranging from 1.6 km at 6°N to 0.9 km at 65°N. The model has 300 vertical levels with a resolution of 1 m at the top-most layers to better resolve a realistic surface boundary layer. In practice, the model effective resolution is about 10–15 km in wavelength, the same as the resolution of the anticipated Surface Water and Ocean Topography (SWOT) altimetry mission (Fu & Ubelmann, 2014). The initial and open boundary conditions are based on GLORYS2v3 ocean reanalysis with a relaxation zone at the northern boundary for sea-ice concentration and thickness. The atmospheric forcing is based on ERA-Interim (ECMWF; Dee et al., 2011), the grid and bathymetry follow (Ducousso et al., 2017). A third-order upwind advection scheme is used for both momentum and tracers in the model simulation to implicitly adapt lateral viscosity and diffusivity to flow properties. The model is spun up for 18 months, and a one-year simulation output from July 2009 to June 2010 is used in this study. eNATL60 has two identical runs (i) eNATL60 with tidal forcing herein referred to as eN60-WT and (ii) eNATL60 with no tidal forcing eN60-NT. The two simulations have perfectly the same configuration except for the inclusion of tidal motions in eN60-WT. In the rest of this article, we use eNATL60 to refer to the two simulations while individual runs are addressed as eN60-WT (with tides) or eN60-NT (no tides). The inclusion of tidal forcing in eN60-WT run provides conversion of tidal energy into the internal wave field through, both, flow-topography interactions and wave-balanced motions interactions (Arbic et al., 2008, 2018).

To investigate cross-scale energy exchanges between different scales of motions, we estimate KE spectral density in frequency-wavenumber space as a proxy to understand the energetic nature of balanced/unbalanced motions in regimes with/without tidal motions. Also, we estimate the rate at which nonlinear mechanisms exchange energy across temporal and spatial scales in the two scenarios. In what follows, our analysis of KE density and transfer is based on the hourly output of horizontal total velocity field and are computed using the following equation;

$$\frac{\partial \widehat{KE}}{\partial t} = T_{KE} + \widehat{\mathbf{u}}^* \cdot OT \quad (1)$$

$$\widehat{KE} = \frac{1}{2} \widehat{\mathbf{u}}^* \cdot \widehat{\mathbf{u}} \quad (2)$$

$$T_{KE} = -\widehat{\mathbf{u}}^* \cdot [\mathbf{u} \cdot \widehat{\nabla} \mathbf{u}] \quad (3)$$

Equations (1)–(3) are derived from the Fourier transform of momentum equation multiplied by horizontal velocity field (Scott & Wang, 2005; Capet et al., 2008; Müller et al., 2015). In the momentum equation (Equation 1), KE and T_{KE} represents the KE density and transfer respectively while OT stands “Other Terms”. $\widehat{}$ refers to Fourier transform and $*$ represents the complex conjugate. Before performing spectral analysis the 2D time series were de-

150 trended and windowed in space and time consistent with standard procedures previously
 151 used in Rocha et al. (2016), Müller et al. (2015) and Torres et al. (2018).

152 The eNATL60 simulation resolves well to a reasonable extent, mesoscale motions, sub-
 153 mesoscale motions, and IGWs (see Figure 1 in SI). The comparison of eNATL60 sea surface
 154 height (SSH) spectral density with SARAL AltiKa (Figure 1a) shows that the predicted
 155 SSH variance by the model compares well with the satellite observation for scales > 100 km.
 156 There are differences at scales < 100 km that are due to the satellite instrument noise level.
 157 There seems to be quite a robust agreement between the two runs of eNATL60 simulations in
 158 wintertime. However, of particular interest is the difference between the runs in summertime
 159 where variance at fine-scales is of higher magnitude in eN60-WT than to eN60-NT.

160 A similar analysis of the KE spectral density in the same region (Figure 1b), shows that
 161 the variance associated with fine-scale motions smaller than 100 km is higher in the eN60-
 162 WT compared to eN60-NT. So what are the mechanism/dynamics at fine-scales in eN60-
 163 WT that could be responsible for this higher variance? A possible answer to this is that the
 164 inclusion of tidal motion in eN60-WT simulation is responsible for enhanced wave activity,
 165 and this is why we see higher variance at fine scales in the SSH and KE spectra density plot.
 166 To qualitatively investigate this, we separate the flow into its rotational and divergent part,
 167 which represents the balanced and the unbalanced wave motions, respectively. Figure 1c
 168 presents the spectral density for these two components. The spectra of the rotational part
 169 for the two runs are almost indistinguishable, indicating that both simulations have nearly
 170 equal energy levels in geostrophically balanced motions. However, the divergence spectra
 171 of the KE is very different between the two simulations. This difference is obvious at scales
 172 less than 500 km, and indeed, the divergent motions are more energetic in eN60-WT by a
 173 factor of 2 with two interesting peaks. We can conclude that the higher variance in eN60-
 174 WT at fine-scales compared to eN60-NT is primarily due to stronger divergent motions in
 175 eN60-WT, which is caused by the inclusion of tidal forcing in this simulation.

176 **3 Seasonality of BMs and IGWs**

177 This section presents the different classes of motions and their seasonality based on
 178 frequency-wavenumber (ω - K) spectral density. This diagnostic will help us better under-
 179 stand how the difference in wave activity between the two simulations affects oceanic mo-
 180 tions' spectral signature across different temporal and spatial scales. Following Torres et
 181 al. (2018), we begin by presenting a schematic (Figure 2a) showing the different observable
 182 dynamical regimes in the ocean as a function of their temporal and spatial scale. These
 183 classes of motions starting with low-frequency, low-wavenumber motions to high-frequency,
 184 high-wavenumber motions are Rossby waves (RW), mesoscale balanced motions (MBM),

185 submesoscale balanced motions (SBM), unbalanced submesoscale motions (USM) and in-
 186 ternal gravity waves (IGW). Due to the computational cost of this diagnostic tool, we
 187 perform the ω - K spectral analysis in a $5^\circ \times 5^\circ$ (-40° to -35° , 40° to 45°) box located
 188 inside the previous large box (see Figure 1 of SI).

189 We also show, in Figure 2, the winter and summer averages of surface KE ω - K spectra
 190 for the two runs. The classes of motions described previously in Figure 2a are identifiable in
 191 the figure. The winter-summer contrast shows a strong seasonality in the spectral density
 192 of SBMs and IGWs. In wintertime, for eN60-WT, energy is mostly concentrated in BMs,
 193 near-inertial waves, and the dispersion curve of IGWs (Figure 2b), while in summertime,
 194 energy is mostly concentrated in the MBMs, near-inertial waves, and internal tides (Fig-
 195 ure 2d). In particular, the variance associated with SBMs is stronger in winter, while that
 196 of IGWs is stronger in summer. Our understanding is that IGWs are stronger in summer
 197 due to shallow mixed layer depth and vertical normal mode intensification. Simultaneously,
 198 SBMs are stronger in winter because they are driven by winter-favored mechanisms such as
 199 frontogenesis, wind-induced frontal instabilities, and mixed layer instability, among other
 200 processes (Qiu et al., 2014; Sasaki et al., 2014; Callies, Flierl, et al., 2015; Brannigan et
 201 al., 2015; McWilliams, 2016; Uchida et al., 2019; Khatri et al., 2021). This out of phase
 202 seasonality of SBMs and IGWs is consistent with the findings of Rocha et al. (2016) and
 203 Torres et al. (2018).

204 Similarly, eN60-NT resolves the same classes of motion as eN60-WT except that in-
 205 ternal tides are absent and supertidal IGWs are less energetic. In wintertime, energy is
 206 mostly concentrated in BMs, near-inertial waves, and along the dispersion curve of IGWs
 207 (Figure 2a). This is consistent with the winter dynamics in eN60-WT. In summertime,
 208 energy is concentrated in MBMs and near-inertial waves (Figure 2c). Unlike eN60-WT, the
 209 seasonality observed in eN60-NT is associated with stronger SBMs and IGWs in winter.
 210 The seasonality of IGWs is reversed in eN60-NT when compared to eN60-WT. How can
 211 this be? We know that the classical paradigm for the generation of the supertidal IGW
 212 continuum is that winds produce near-inertial waves, barotropic tidal flow over topographic
 213 features creates internal tides, and the energy along the IGW dispersion curve is due to
 214 nonlinear interactions. Both simulations are forced with realistic high-frequency winds with
 215 3-hourly outputs (Brodeau et al., 2020). These winds are stronger in winter, hence there is
 216 a well-resolved near-inertial wave and IGWs dispersion curve in winter. The dynamics in
 217 summertime are different between the two runs. For the eN60-WT simulation, internal tides
 218 generated by tidal motions are amplified by a shallow mixed layer in summertime, and non-
 219 linearity produces energy in the IGW dispersion curve. Thus for eN60-NT, the mechanism
 220 for generating waves in summertime is relatively weak; no tidal forcing and weaker winds,

221 hence relatively weak wave motions in summer. The difference in the spectral energy density
 222 of different dynamical regimes between the two simulations extends below the surface (see
 223 Figure 3 of SI).

224 To obtain frequency spectra, we integrate the ω - K spectra over all wavenumbers for the
 225 two runs (Figure 3). In summertime, the variance at high frequencies (M_2 and supertidal
 226 frequencies) is higher in eN60-WT compared to eN60-NT. We believe that this is likely due
 227 to the amplification of IGWs by tidal motions. eN60-WT spectra approximately follow the
 228 estimated Garrett-Munk spectra (Garrett & Munk, 1975; Cairns & Williams, 1976; Müller
 229 et al., 2015) in summertime. Visible in eN60-WT spectra are the peaks at the inertia
 230 frequency f and the M_2 tidal frequency. In contrast, only the near-inertial peak is visible in
 231 eN60-NT. To a large extent, we now understand the dynamics responsible for the differences
 232 in KE density that we see in the two simulations. In the following sections, we shall discuss
 233 how the enhanced IGWs arising from tidal forcing affect the redistribution of energy in
 234 different dynamical regimes.

235 **4 Modulation of KE forward flux by IGWs**

236 In this section, we will discuss the impact of resolving internal tides on the magnitude
 237 and direction of KE cascade at fine-scales in the wavenumber and frequency-wavenumber
 238 domain by comparing the twin eNATL60 runs.

239 **4.1 Wavenumber KE flux**

240 We start by examining the KE flux in the wavenumber domain. We do this by esti-
 241 mating the net energy (spectral flux) passing through individual wavenumbers in spectral
 242 space. The spectral flux is obtained by integrating the energy transfer (Equation 3) from a
 243 particular wavenumber K to K_0 (the wavenumber corresponding to the box size).

244 We present in Figure 4a,d, the winter and summer averages of KE spectral flux for both
 245 simulations. In wintertime and in the two simulations (Figure 4a), the flux is nearly identical
 246 across all wavenumbers. The forward cascade starts at around 25 km and extends down to
 247 a kilometric scale. In summertime (Figure 4d), the magnitude of the forward cascade at
 248 high wavenumbers differs significantly between the two runs. The inclusion of tidal forcing
 249 in eN60-WT yields a forward flux at high wavenumbers that is a factor of 4 higher than
 250 the forward high-wavenumber flux in eN60-NT. This difference in cascade highlights how
 251 internal tides enhance the forward cascade of KE at high wavenumbers.

252 Thus far, the KE cascade has been estimated only at the surface. Considering we
 253 have three-dimensional information of the ocean, as opposed to satellite observations, it

254 is of great interest to understand the nature of the KE cascade in the ocean's interior.
 255 In Figure 4b,c,e,f, we present the spectral flux computed at 32 different vertical levels in
 256 the upper 1000 m of the water column. In winter and summer, at lower wavenumbers,
 257 the average KE flux in the two simulations is characterized by a net inverse cascade that
 258 extends down to around 700 m in the interior. In wintertime, the forward cascade at higher
 259 wavenumbers in eN60-WT (Figure 4c) is strong both at the surface and in the interior. In
 260 contrast, in eN60-NT (Figure 4b), the forward cascade is confined mostly to the surface. In
 261 summertime, the forward cascade in eN60-WT (Figure 4f) span the upper ocean but with
 262 a gradual decrease in magnitude farther down the water column. In contrast, in eN60-NT
 263 (Figure 4e), the forward cascade is nearly absent throughout the upper ocean. A stronger
 264 forward cascade (in summertime for eN60-WT) in the interior is an indication that internal
 265 tides transfer substantial amounts of KE to dissipative scales throughout the upper ocean
 266 water column.

267 4.2 Frequency-wavenumber KE transfer

268 We have demonstrated that internal tides enhance the supertidal IGW continuum and,
 269 in particular, enhance the forward cascade of energy in summertime. To better explain how
 270 internal tides modify cross-scale energy exchanges among the different classes of motions, we
 271 present in Figure 5 the winter and summer averages of KE spectral transfer in frequency-
 272 wavenumber space. In the ω - K spectra, negative values of spectra transfer imply that
 273 non-linearity extracts energy from these regions to feed other regions with positive values.
 274 In other words, sinks of energy are characterized by positive values, while sources of energy
 275 have negative values.

276 We start by discussing the spectral transfer in wintertime (cf. Equation 3; Figures 5a,c).
 277 In eN60-NT, the spectral transfer's positive values (blue) show that balanced motions serve
 278 as a source of energy for other motions, namely energy is being extracted from the balanced
 279 motions. In contrast, near-inertial motions and motions with scales less than 10 km are
 280 the major sinks of KE (red). The rate of non-linear exchanges in eN60-WT is similar to
 281 eN60-NT except for the mild intensification of energy gained by IGWs in eN60-WT. To
 282 summarize, submesoscale motions and internal gravity waves are sinks of KE in wintertime,
 283 with the former playing the major role.

284 The summer spectra (Figure 5b,d) differ significantly between the two runs as expected.
 285 In eN60-NT, BMs represent the major source of energy for other motions, while energy is
 286 gained mostly by near-inertial motions. The transfer at high frequencies and wavenumbers
 287 is very small. We can interpret this to signify that high-frequency motions and submesoscale
 288 motions are less energetic in eN60-NT in summertime. This result is consistent with results

289 in the KE spectra density. However in eN60-WT, the situation is different. The major
 290 energy source for other motions is the MBMs and the semi-diurnal tides (blue). Near-
 291 inertial waves, SBMs and the supertidal IGW continuum spectrum are seen in Figure 5
 292 to be gaining energy. The extraction of semi-diurnal tidal energy, and gain of energy in
 293 the IGW continuum, due to nonlinear interactions was also seen in (Müller et al., 2015)
 294 but is more clear here, perhaps because of the finer vertical and horizontal grid spacing in
 295 eN60-WT.

296 In summary, there are two mechanisms of energy extraction in summertime: (i) without
 297 tidal forcing, near-inertial waves gain energy from BMs, and (ii) with tidal forcing, near-
 298 inertial waves and the supertidal IGW continuum gain energy from internal tides and MBMs.
 299 In summary, the forward cascade in eN60-WT, is associated with the transfer of energy
 300 by nonlinearity from balanced motions and internal tides to near-inertial waves and the
 301 supertidal IGW continuum. This transfer is possible due to the intensification of IGWs in
 302 the presence of tidal motions. The lack of tidal motions in eN60-NT (compared to eN60-
 303 WT) shows how effective internal tides are in providing a route to energy dissipation in
 304 summertime, both at the surface and in the interior. This result strongly emphasizes the
 305 need to include tidal forcing in ocean model simulations to accurately predict cross-scale
 306 energy exchanges at fine-scales.

307 5 Summary and conclusion

308 The role of internal tides in the seasonality of KE density and spectral KE transfer
 309 was investigated in this study. Our analysis was based on the output of a realistic NEMO
 310 simulation of the North Atlantic Ocean with a horizontal resolution of $1/60^\circ$. We used
 311 two outputs of this numerical experiment; one with, and one without, tidal forcing. These
 312 twin experiments permit investigation of how IGWs generated by tidal forcing modify KE
 313 variance, cross-scale exchanges, and associated seasonality. In the simulation without tides,
 314 IGWs are stronger in winter, whereas, in simulation with tides, they are stronger in summer.
 315 The latter condition is consistent with the findings of Rocha et al. (2016) and Torres et al.
 316 (2018).

317 Our results also show that resolving internal tides in the presence of energetic subme-
 318 soscale motions has a strong impact on KE transfer in summertime (cf. Barkan et al., 2021,
 319 their Figure 3d,f). The magnitude of the estimated forward cascade at high wavenumbers
 320 (both at the surface and at depth) in the simulation with tidal forcing is a factor 4 higher
 321 than in the simulation without tidal forcing (Figure 4). Overall, we identified that two
 322 mechanisms supporting the KE forward cascade; (i) forward cascade due to energetic sub-
 323 mesoscale motions in wintertime and (ii) forward cascade due to IGWs (enhanced by tidal

324 forcing) in summertime (Figure 5). Our results underscore that at fine-scales, internal tides
325 can provide an effective route to KE dissipation.

326 In light of the SWOT altimeter mission (Morrow et al., 2019; Torres et al., 2019), the
327 difference between the runs with tidal forcing and without (e.g. Figures 1 and 3) highlight
328 the importance of tidal forcing in emulating the upcoming altimeter observations (Arbic et
329 al., 2018; Barkan et al., 2021; Yu et al., 2021).

330

331 **Open research**

332 The twin eNATL60 data are permanently stored at the CINES supercomputing center
333 in Montpellier, France. The data are available via OPeNDAP (doi:10.5281/zenodo.5910038).

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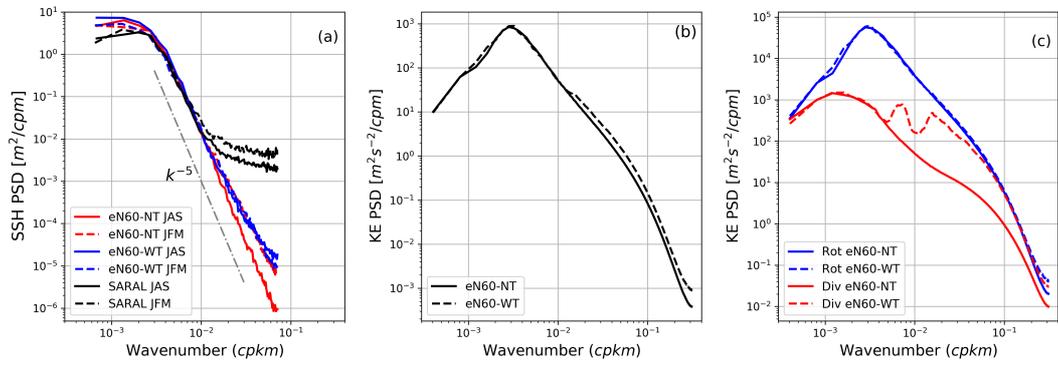


Figure 1. (a) Comparison of SSH wavenumber spectra between eNATL60 and SARAL AltiKa satellite. (b) One year average of surface KE wavenumber spectral density computed from hourly outputs of eN60-NT (no tides) and eN60-WT (with tides). (c) Helmholtz decomposition of KE into rotational (ζ) and divergent (δ) spectral components. Thick curves represent the simulation without tides (eN60-NT) and dashed curves represent the simulation with tides (eN60-WT).

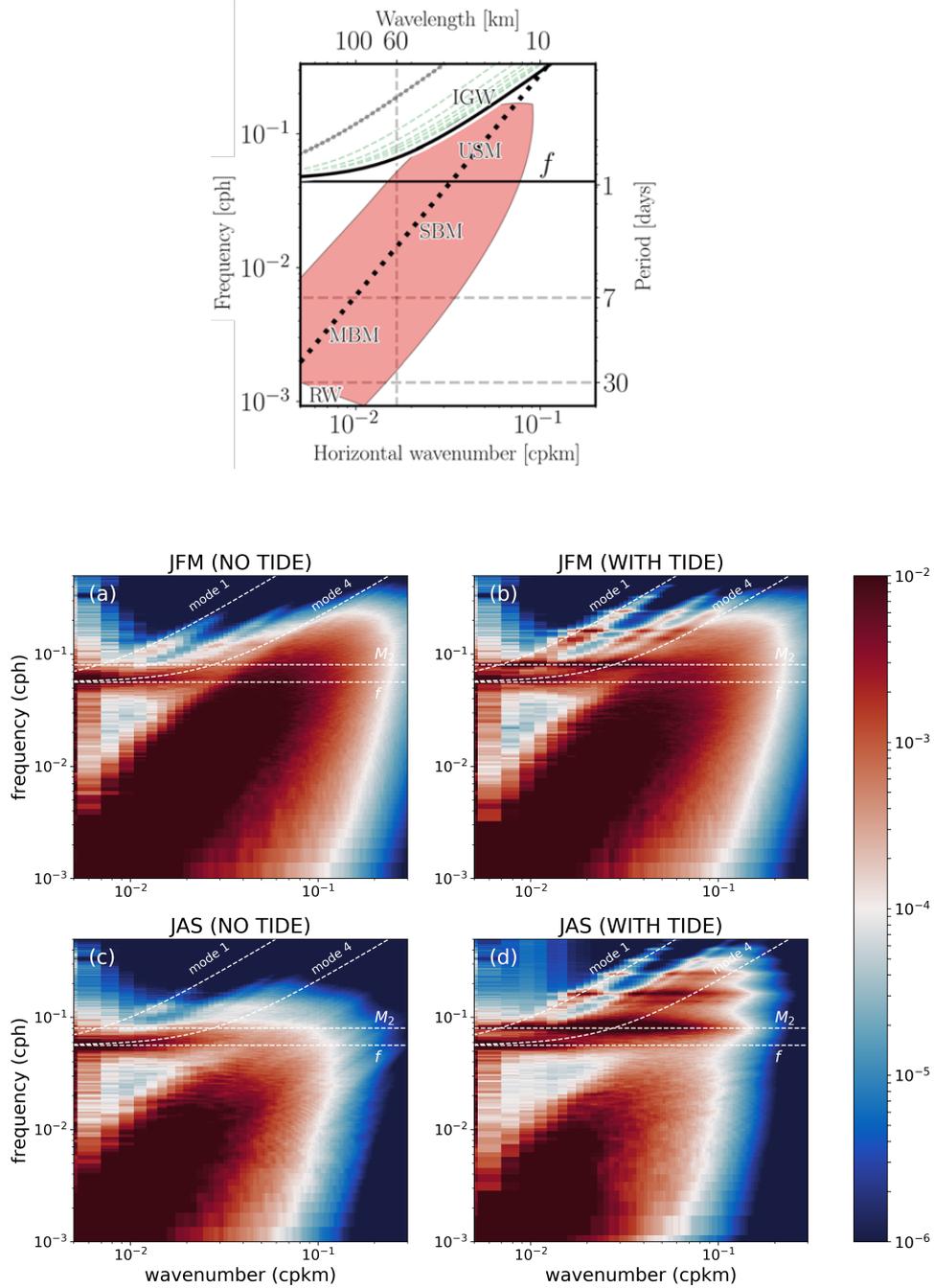


Figure 2. Top panel: A schematic of the observable dynamical regimes with different classes of motions in the ocean (adapted from Torres et al., 2018). These classes of motions starting with low frequency, small wavenumber motions to high frequency frequency, high wavenumber motions are Rossby waves (RW), mesoscale balanced motions (MBM), submesoscale balanced motions (SBM), unbalanced submesoscale motions (USM) and internal gravity waves (IGW). Bottom four plots: Surface KE frequency-wavenumber spectra computed from hourly outputs of eN60-NT (no tides) and eN60-WT (with tides) for winter (JFM) and summer (JAS) time.

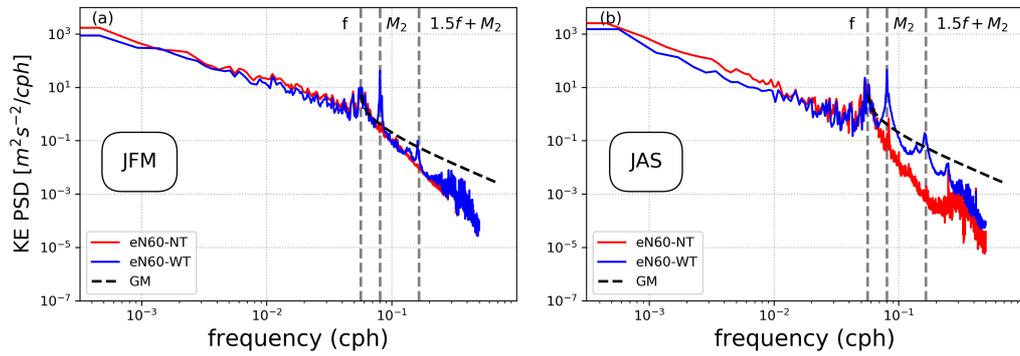


Figure 3. Surface KE frequency spectral density, obtained by integrating the ω - K spectra over all wavenumbers for eN60-NT (no tides) and eN60-WT (with tides). The three grey dash lines represent the inertia frequency f , the M_2 tidal frequency, and $1.5f + M_2$. The dashed black line represents the estimate of the Garrett-Munk spectra computed with reference values of total energy of the internal wavefield and stratification set to $E_0 = 6.3e^{-5} \text{m}^2 \text{s}^{-2}$ and $N_0 = 5.2e^{-3} \text{s}^{-1}$, respectively (Garrett & Munk, 1975; Cairns & Williams, 1976; Müller et al., 2015). (a) Winter (JFM) and (b) Summer (JAS).

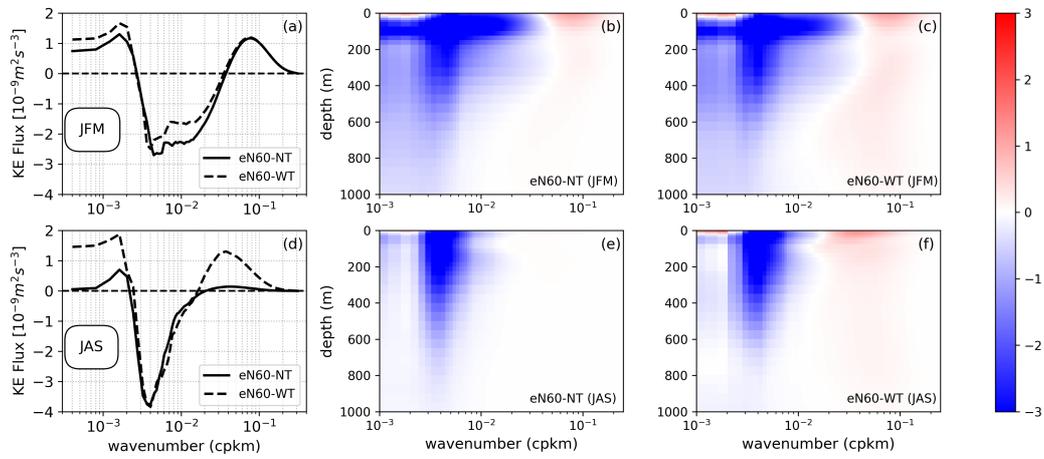


Figure 4. Surface KE spectral flux computed from hourly outputs of eN60-NT (no tides) and eN60-WT (with tides). Summer : July, August and September. Winter : January, February and March. (b,c,e,f) Winter and summer averages of KE spectral flux as a function of depth for eN60-NT and eN60-WT.

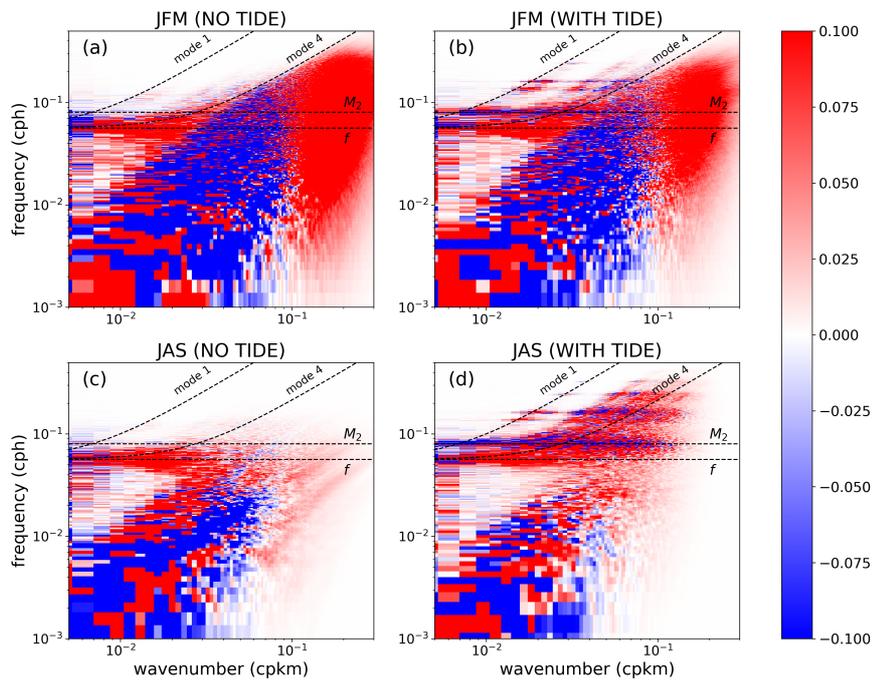


Figure 5. Surface KE transfer in frequency-wavenumber space computed from hourly outputs of eN60-NT (no tides) and eN60-WT (with tides) for winter (JFM) and summer (JAS) time.