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Please, feel free to get in contact: [gino@ginodegelder.nl](mailto:gino@ginodegelder.nl)

1       **Quaternary E-W extension uplifts Kythira Island and segments the Hellenic Arc**

2       **G. de Gelder<sup>1,2</sup>, D. Fernández-Blanco<sup>2,3</sup>, N. Öğretmen<sup>4,5</sup>, S. Liakopoulos<sup>6</sup>, D.**  
3       **Papanastassiou<sup>6</sup>, C. Faranda<sup>5</sup>, R. Armijo<sup>2</sup>, and R. Lacassin<sup>2</sup>**

4       <sup>1</sup>ISTerre, Université Grenoble-Alpes, 1381 Rue de la Piscine, 38400 St. Martin d'Hères, France

5       <sup>2</sup>Université de Paris, Institut de physique du globe de Paris, CNRS, F-75005 Paris, France

6       <sup>3</sup>Barcelona Center for Subsurface Imaging, Passeig Marítim de Barceloneta 37-49, E-08003  
7       Barcelona, Spain

8       <sup>4</sup>Eurasia Institute of Earth Sciences, Istanbul Technical University, 34467, Istanbul, Turkey

9       <sup>5</sup>Dipartimento di Scienze, Università degli Studi Roma Tre, Largo San Leonardo Murialdo, 1,  
10       00146 Rome, Italy

11       <sup>6</sup>Institute of Geodynamics, National Observatory of Athens, GR-11810 Athens, Greece

12       Corresponding author: Gino de Gelder ([gino.de-gelder@univ-grenoble-alpes.fr](mailto:gino.de-gelder@univ-grenoble-alpes.fr))

13       **Key Points:**

- 14       • ~100 m of Tortonian-Pliocene subsidence of Kythira Island was followed by ~300-400 m  
15       of uplift since ~2.6 Ma.
- 16       • Regional E-W extension, caused by large-scale plate boundary changes ~1.5-0.7 Ma, is  
17       uplifting Kythira Island.
- 18       • We highlight the general importance of upper crustal faulting as a major driver of forearc  
19       uplift.  
20

## 21 **Abstract**

22 Several crustal and lithospheric mechanisms lead to deformation and vertical motion of the upper  
23 plate during subduction, but their relative contribution is often enigmatic. The Hellenic Forearc  
24 has been uplifting since Plio-Quaternary times, yet the spatiotemporal characteristics and sources  
25 of this uplift are poorly resolved. The remarkable geology and geomorphology of Kythira Island,  
26 in the southwestern Hellenic forearc, allow for a detailed tectonic reconstruction since the Late  
27 Miocene. We present a morphotectonic map of the island, together with new biostratigraphic  
28 dating and detailed analyses of active fault strikes and marine terraces. We find that the  
29 Tortonian-Pliocene stratigraphy in Kythira records ~100 m of subsidence, and a wide coastal  
30 *rasa* marks the ~2.8-2.4 Ma maximum transgression. Subsequent marine regression of ~300-400  
31 m and minor E-W tilt are recorded in ~12 marine terrace levels at uplift rates of ~0.2-0.4 mm/yr.  
32 Guided by simple landscape evolution models, we interpret the coastal morphology as the result  
33 of initial stability or of slow, gradual sea-level drop since ~2.8-2.4 Ma, followed by faster uplift  
34 since ~1.5-0.7 Ma caused by roughly N-S normal faulting. Our findings on- and offshore  
35 emphasize that E-W extension is the dominant mode of regional active upper crustal  
36 deformation, and N-S normal faults accommodate most, if not all of the uplift on Kythira. We  
37 interpret the initiation of E-W extension as the result of a change in plate boundary conditions, in  
38 response to either propagation of the North Anatolian Fault, incipient collision with the African  
39 plate, mantle dynamics or a combination thereof.

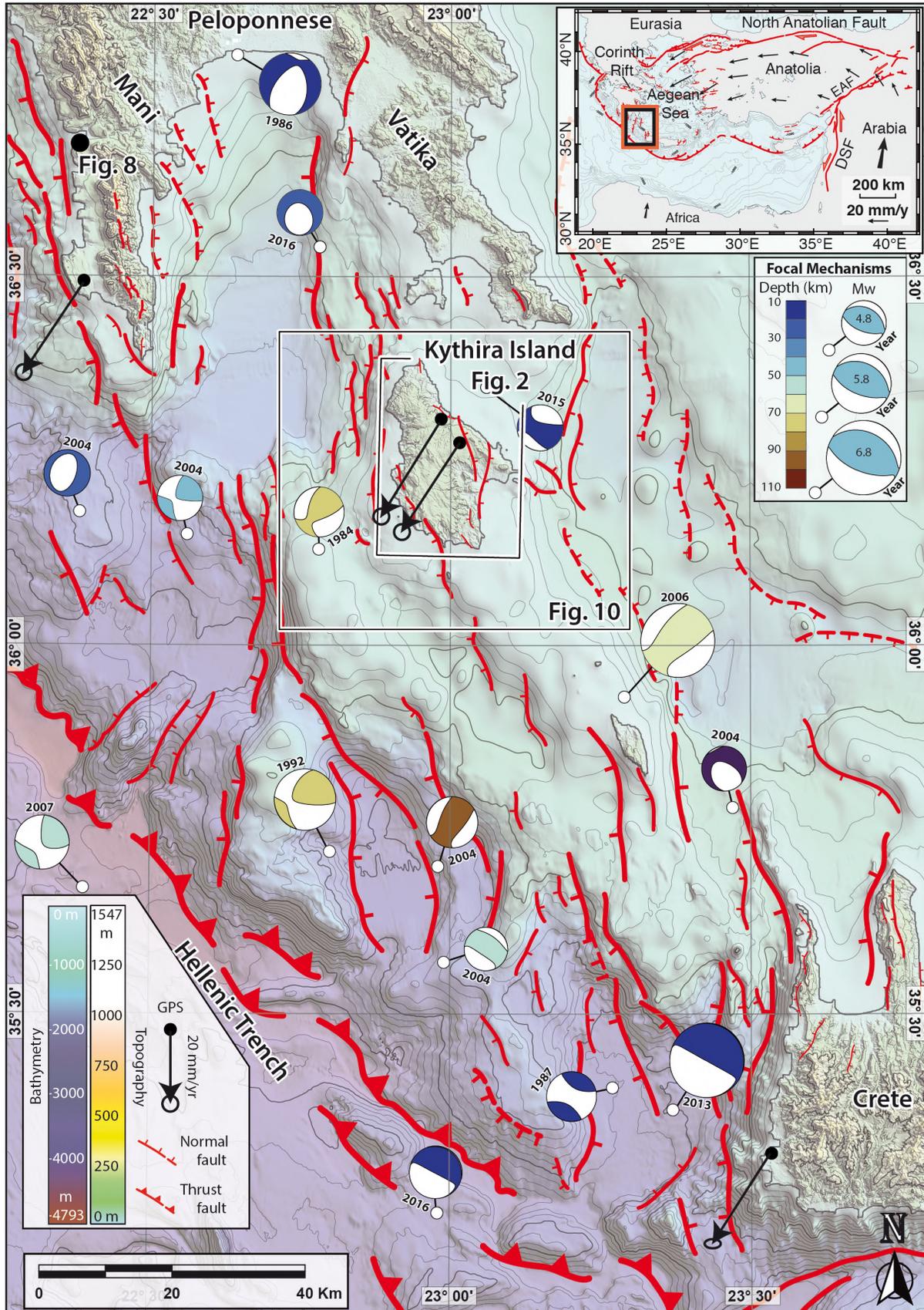
## 40 **1 Introduction**

41 The mode of deformation of the upper plate during subduction and the mechanisms of  
42 uplift at the front of the overriding plates are two fundamental questions largely addressed that  
43 remain unresolved (e.g. Davis et al., 1983; Willet et al., 1993; Larroque et al., 1995; Gutscher et  
44 al., 1996; Fuller et al., 2006; Gallen et al., 2014). The Hellenic subduction assumes the largest  
45 convergence velocity and holds the highest seismic activity among Mediterranean arcs  
46 (McClusky et al., 2000; Reilinger et al., 2006), and thus, its forearc is key in this debate. The  
47 steep normal faults that control the present-day structuration and seismicity of the Hellenic Arc  
48 are oblique to the intraplate contact and change in orientation along the arc, transecting former  
49 fold-and-thrust structures that are parallel to the trench (Lyon-Caen et al., 1988; Armijo et al.,  
50 1992; de Chabali er et al., 1992; Shaw and Jackson, 2010). Constraining the timing and

51 kinematics of these normal faults, and of the recent structuration of the arc between uplifted  
52 islands and offshore grabens, is of prime relevancy to better understand the mechanics of the  
53 Hellenic plate boundary.

54 Kythira is the largest island between SW Peloponnese and NE Crete (Fig. 1), and  
55 provides an exceptional but surprisingly unattended opportunity to understand the interaction  
56 between past and active tectonics in the Hellenic Arc. Kythira's sedimentary record and  
57 morphological characteristics allow a detailed reconstruction of its tectonic evolution since the  
58 Late Miocene. Whereas emerged marine sedimentary basins record the island vertical motions  
59 during the Neogene, its landscape archives a recent uplift phase as coastline abrasion surfaces,  
60 marine terraces and steep river gorges. Holocene fault scarps and numerous historical  
61 earthquakes (Papadopoulos and Vassilopoulou, 2001) evidence active faulting, but controversy  
62 remains on the processes that drove the initiation of the main fault systems, both on the island  
63 itself and within the regional context of the Hellenic Arc.

64 We reconstruct the Late Miocene-Recent tectonic evolution of Kythira through a multi-  
65 disciplinary approach that includes morphometric and basin analyses, fieldwork, biostratigraphic  
66 dating and numerical modeling. We developed a 2 m-resolution Digital Surface Model (DSM)  
67 from Pleiades satellite imagery that allows us to detect and map geologic and geomorphic  
68 features like faults, marine terraces, drainage systems and stratigraphic layering with  
69 unprecedented detail. We present a detailed geological, structural and geomorphological map of  
70 Kythira Island, and a comprehensive analysis of marine terraces based on a combination of  
71 fieldwork and the high resolution DSM. We constrain absolute ages and evaluate the rates of  
72 surface uplift dating marine deposits near the highest and most extensive marine abrasion surface  
73 of the island with microfossils. Combining these ages and the coastal morphology, we then use  
74 simple landscape evolution models to test different uplift scenarios. This allows us to put  
75 forward a detailed discussion on the tectonic changes that have affected the island, and discuss  
76 our findings in the context of the evolution of the Hellenic Arc.



78 **Figure 1. Active tectonics of the SW Hellenic Arc.** Topography is an ALOS Global Digital  
79 Surface Model (DSM), and bathymetry is from the European Marine Observation and Data  
80 Network (EMODNet) Digital Terrain Model (DTM). Offshore fault mapping is an interpretation  
81 of the bathymetry based on relief offset and morphology, and accounts for fault mapping by  
82 Lyberis et al. (1982) and Armijo et al. (1992). Focal mechanisms (>4.5 Mw since 1976, colored  
83 by depth and sized by magnitude) are build using both the main and the monthly-curated  
84 databases of the Global Centroid Moment Tensor (GCMT) catalog (<https://www.globalcmt.org/>;  
85 Ekström et al., 2012), as downloaded on the 14th December 2020. GPS vectors are relative to  
86 stable Eurasia, and taken from England et al. (2016), but we note that these would be similar in  
87 direction and magnitude if re-calculated with respect to stable Nubia (Reilinger et al., 2006;  
88 Shaw and Jackson, 2010). The location of figures 2, 8 and 10 is indicated.

## 89 **2 Tectonic, geologic, and geomorphic setting**

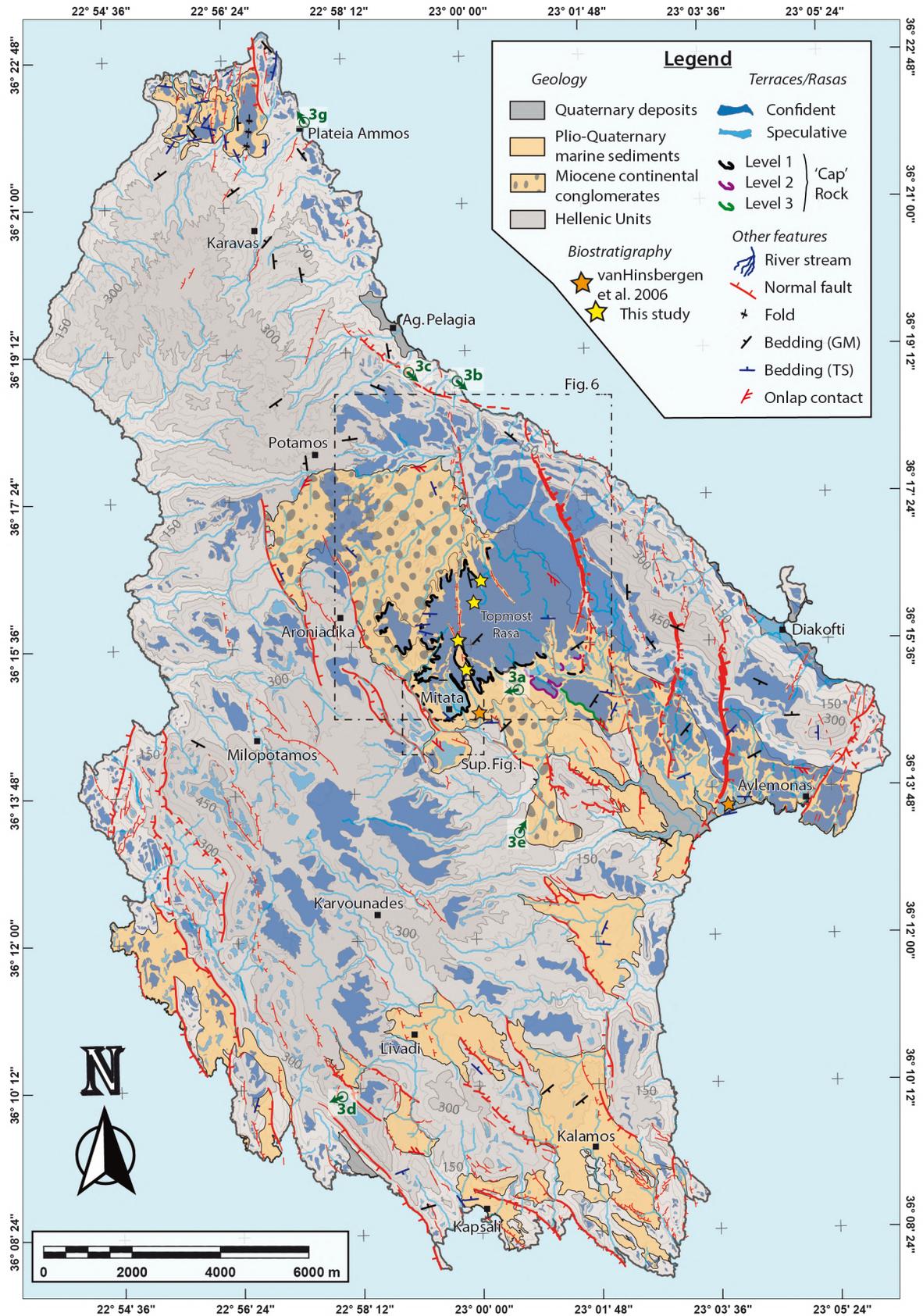
90 Subduction of the African plate below the Aegean has stacked the Hellenic fold-and-  
91 thrust belt nappes since the early Mesozoic (e.g. Aubouin, 1957; Jacobshagen, 1986; Facenna et  
92 al., 2003; van Hinsbergen et al., 2005; Jolivet and Brun, 2010). Three of those nappes are  
93 exposed on Kythira: the Arna Unit, the Tripolis Unit and the Pindos Unit (Manolessos, 1955;  
94 Papanikolaou and Danamos, 1991; Danamos, 1992). The HP/LT metamorphic Arna Unit  
95 (Gerolymatos, 1994; Marsellos and Kidd, 2008; Marsellos et al., 2014) is highly folded and  
96 crops out in the north of the island. Overthrusting the Arna Unit, the Tripolis Unit contains  
97 Jurassic-Eocene limestones, dolomites and flysch (Danamos, 1992). The Tripolis Unit crops out  
98 along large parts of the NE and SW coasts and in the center of the island, typically along NW-SE  
99 trending ridges parallel to the Hellenic trench. The Pindos Unit overthrusts both underlying units  
100 and contains Cretaceous-Early Cenozoic limestones and flysch (Danamos, 1992). It mainly crops  
101 out in the central and SE parts of the island, and has a NW-SE trend similar to the Tripolis Unit.  
102 Apatite and zircon fission track cooling ages for the metamorphic Arna Unit indicate exhumation  
103 during the Middle-Late Miocene (Marsellos and Kidd, 2008; Marsellos et al., 2014). The oldest  
104 preserved Neogene sediments on Kythira Island are of Tortonian age (Theodoropoulos, 1973;  
105 Meulenkamp et al., 1977; van Hinsbergen et al., 2006), and were probably deposited not long  
106 after exhumation ceased.

107 Neogene-Quaternary rocks unconformably overlie the Hellenic nappes and are scattered  
108 around the island, with the largest basin located in its central-eastern sector, between Potamos  
109 and Avlemonas (Fig. 2). The Neogene stratigraphy has previously been described as a Tortonian  
110 terrigenous-clastic succession at the bottom overlaid in angular unconformity by a Pliocene  
111 calcareous succession (Theodoropoulos, 1973; Meulenkamp et al., 1977). Paleobathymetry  
112 estimates from a Plio-Pleistocene calcareous section near the present-day coast at Avlemonas  
113 (orange star on Fig. 2) suggest deepening from ~300 to ~750 m depth between ~3.5 and ~2.5 Ma  
114 (van Hinsbergen et al., 2006).

115 The highest elevated sediments of the calcareous succession form part of a large marine  
116 *rasa* (Fig. 2), i.e. a wide coastal planation surface of polygenic origin formed by sea erosion  
117 during multiple sea-level stands (e.g. Guilcher, 1974; Pedoja et al., 2006; Regard et al., 2010;  
118 Dawson et al., 2013). The *rasa* in Kythira is carved into both basement nappes and Neogene

119 sediments (Fig. 2), is most extensive around the airport (NE of Mitata), and can be found all  
120 around the island between elevations of ~200 and ~400 m. It forms the highest marine abrasion  
121 surface in a flight of marine terraces that reaches the present-day coastline, which is best  
122 preserved along the east coast. Gaki-Papanastassiou et al. (2011) described 6-8 different terrace  
123 levels in the southern part of the island, with sea caves and notches at 0.4-0.6, 2, and 4 meters  
124 above sea level evidencing recent uplift.

125         The active faults on the island have been described as a combination of NW-SE and NE-  
126 SW trending normal faults related to trench-normal and trench-parallel extension respectively  
127 (Marsellos and Kidd, 2008), or NNW-SSE trending normal faults related to N-NE directed  
128 subduction (Gaki-Papanastassiou et al., 2011). Historical earthquakes in Kythira have been  
129 documented in 800 AD, 1750, 1798, 1866 and 1903 (Gaki-Papanastassiou et al., 2011), but are  
130 difficult to associate with specific faults. The 2006 magnitude 6.7 Mw earthquake that caused  
131 significant destruction to the town of Mitata (Fig. 2) had a focal depth of ~60 km (Fig. 1), and is  
132 thus probably unrelated to any of the faults observed on the island (Konstantinou et al., 2006).



134 **Figure 2. Active tectonics on Kythira.** Map with main geologic and geomorphic features, based  
135 on 2m-resolution Digital Surface Model developed from Pleiades satellite imagery and fieldwork  
136 (see text). Topographic contours are given at 50 m intervals, with color changes every 150 m.  
137 Active faults in red have different stroke thickness to show 5 levels of relative offset, and most  
138 present fault scarps as corroborated in situ during fieldwork. Marine terraces and the topmost  
139 rasa levels are shown in two tones of blue, with a lighter tone for more speculative surfaces (see  
140 main text). Bedding measurements are both from this study (TS) and the geological map (GM) of  
141 Danamos (1992). Observed onlap contacts between the largest uplifted basin and basement are  
142 also shown. Insets show locations of Fig. 6 and Sup. Fig. 1. Green symbols mark the field  
143 photographs of Fig. 3 with viewing direction marked with arrows. Yellow stars mark the  
144 locations of samples dated by biostratigraphy (see results on Fig. 6); orange stars, near Mitata  
145 and W of Avlemonas, show location of Van Hinsbergen (2006) sites.

### 146 **3 Methods**

147 3.1 Developing and analysing high-resolution DSMs derived from Pleiades imagery.

148 Tri-stereo Pleiades satellite images of 0.5 m resolution covering the whole island of Kythira were  
149 obtained through the ISIS and Tosca programs of the CNES. The open-source software MicMac  
150 (Rosu et al., 2015; Rupnik et al., 2016) was used to create tie-points, orientate the images and  
151 calculate 0.5 m-resolution DSMs, using ground control points at 0 m elevation for several  
152 locations along the coastline. To reduce the topographic effects of vegetation, crops and man-  
153 made structures, the DSMs were downsampled to 2 m resolution. Objects of ~50 cm height are  
154 easily detected, indicating a relative vertical accuracy of less than 1 m. The accuracy and  
155 refinement of our geologic/geomorphic map is relevantly improved thanks to usage of the  
156 Pleiades DSM in comparison to freely available digital elevation models (Supplementary Fig. 1).  
157 In addition, we also used stacked swaths of various sectors of the Pleiades DSM to gain a “2.5D”  
158 view of the topography. Stacked swaths (Armijo et al., 2015, Fernández-Blanco et al., 2019b)  
159 contain a significant number of parallel swath profiles derived from topographic data, which are  
160 plot together orthogonal to their strike as hairlines. By stacking swath profiles, the resulting  
161 profile highlights topographic coherence in depth (perpendicular to the viewpoint), allowing the  
162 distinction of structural and morphological features that are continuous over large scales. We

163 share a georeferenced hillshade image and slope map of the 2 m-resolution Digital Surface  
164 Model through these links: <https://doi.org/10.6084/m9.figshare.18715535.v1> (hillshade image)  
165 and <https://doi.org/10.6084/m9.figshare.18714914.v1> (slope map). The map of Figure 2 can be  
166 downloaded in georeferenced format (as Geospatial PDF) at  
167 <https://doi.org/10.6084/m9.figshare.18703496.v1>.

### 168 3.2 Active faults and sedimentary basins

169 We carried out a detailed, albeit preliminary, mapping of faults, unit contacts and marine  
170 terraces and rasas by analyzing a combination of satellite imagery and DSM-derived hillshade  
171 maps, slope maps and topographic profiles. We carefully mapped contacts between the basement  
172 nappes and sedimentary basins as well as contacts within the basins, refining them from previous  
173 maps (Geological Map of Kythira Island – IGME, 1966; Danamos, 1992). During this stage prior  
174 to fieldwork, we also mapped the terraces semi-automatically, using the slope and roughness of  
175 the topography as guidelines, but evaluating every mapped terrace manually (section 3.4).

176 Fieldwork was carried out to verify and improve the mapped structures, study the terrace  
177 and fault morphologies in detail, resolve the stratigraphic contacts of Neogene deposits with the  
178 basement rocks, and understand the overall tectonic architecture of the sedimentary basins. All  
179 preliminary mapped faults were checked in the field to corroborate its correct mapping, and most  
180 faults presented a (sometimes degraded) fault scarp. We focused our analysis of the stratigraphic  
181 succession to the largest basin between Potamos and Avlemonas. Bedding measurements were  
182 taken at representative locations and kinematic indicators on fault scarps were measured using  
183 standard structural methods.

184 To quantify the overall strike of the active faults on Kythira, we subdivided them in 200-  
185 m segments, roughly corresponding to the smallest mapped faults, and placed them in bins of 5°.  
186 The same was done for all faults mapped on- and off-shore to the NE of the Hellenic Trench  
187 (Fig. 2) with 4-km long fault segments. Fault dip directions should be perpendicular to the strike  
188 of active faults. As a coarser but more objective alternative to measure the dominant fault strike  
189 offshore, we also quantified the slope direction of the bathymetry NE of the Hellenic Trench,  
190 calculating the 3x3 pixel slope direction for every ~200x200 m pixel. To a first order, the slope  
191 direction of steep slopes should be perpendicular to the strike of active faults.

### 192 3.3 Biostratigraphic dating

193 Marine foraminifer and ostracod species typically live(d) within a restricted water depth  
194 range, and within restricted time-intervals throughout geological history. As such, their  
195 occurrence in marine sediments - if not re-worked - can provide constraints on the age and/or  
196 paleodepth of those sediments. We sampled the deposits immediately below the uppermost  
197 marine sediments at 5 different sites for microfossil analysis (planktic and benthic foraminifers  
198 and ostracods), to obtain an approximate age for the formation of the rasa in locations around the  
199 area NE of Mitata (yellow stars on Fig. 2). These uppermost sediments typically form a 2-to-5  
200 m-thick layer of coarse marine sandstones, rich in shells and oysters, acting as a caprock to the  
201 underlying finer sandstones that were sampled (Fig. 2).

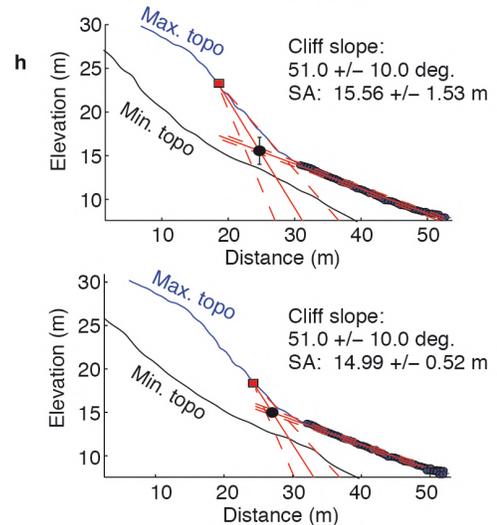
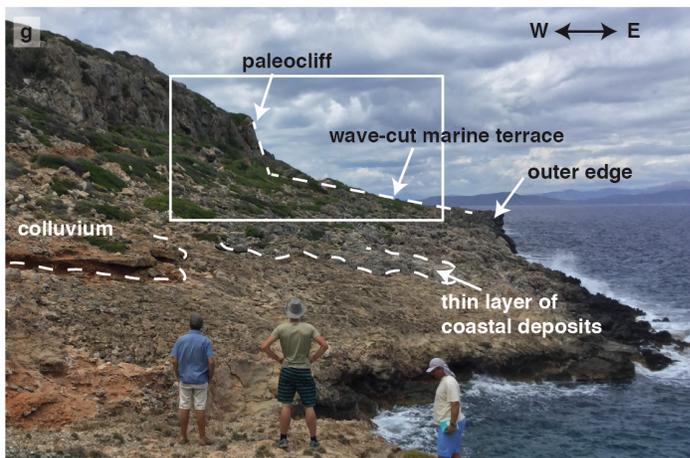
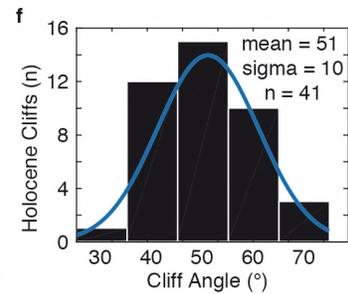
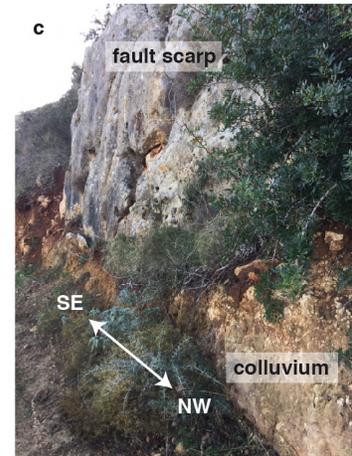
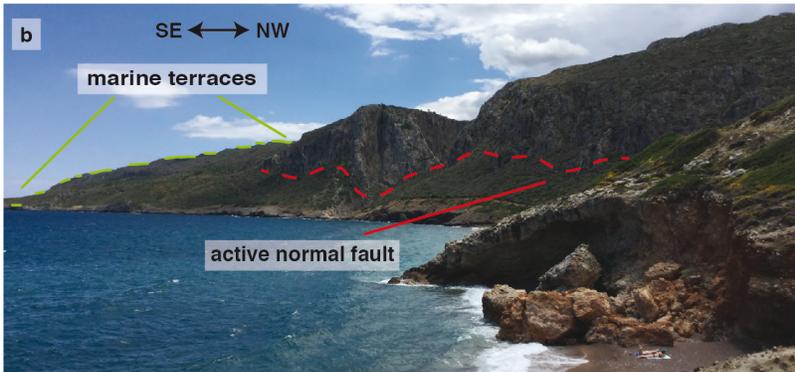
202 Samples were soaked in a H<sub>2</sub>O<sub>2</sub> 5% solution for 24-48 hours, sieved under tap water using  
203 0.066 and 0.125 mm mesh sieves, and dried in a 40°C oven. Whenever possible, up to 300  
204 individuals were hand-picked from the dry sieved samples under a stereomicroscope, in order to  
205 collect a significant sample of the thanatocoenosis. The taxonomic identification of planktic  
206 foraminifers was based on Parker (1962), Postuma (1971), Kennett and Srinivasan (1983),  
207 Iaccarino et al. (2007), and Lirer et al. (2019). Benthic foraminifer classification was based on  
208 AGIP (1982), Morkhoven et al., (1986), and Meriç et al. (2014). For further visual comparison,  
209 we used the foraminifer databases of WoRMS, Foraminifera.eu, and [www.microtax.org](http://www.microtax.org). The  
210 biostratigraphy of the study area was updated following the planktic foraminifer biozones of  
211 Lirer et al. (2019) and the Global Time Scale 2012 edited by Anthonissen and Ogg (2012). We  
212 took several pictures for identification under the Scanning Electron Microscope Philips XL30  
213 (Suppl. Fig. 2).

### 214 3.4 Marine terrace analysis

215 Marine terraces are geomorphic markers that record the former position of the sea-level,  
216 and can thus be used to derive relevant information about vertical tectonic movements (e.g.  
217 Dupré, 1984; Armijo et al., 1996; De Gelder et al., 2019). The marine terraces on Kythira are  
218 wave-cut terraces, which are typically formed by wave-abrasion during sea-level rise and  
219 highstands (Anderson et al., 1999) and expressed as smooth planar surfaces with slope angles of  
220 1°-15° (Scott and Pinter, 2003) that may be covered by a thin layer of sediments.

221 To map marine terraces, we used the surface classification model (SCM) of Bowles and  
222 Cowgill (2012) as a guideline to select surfaces with a low slope and roughness. After smoothing  
223 the DSM with a 6x6 m moving window, a roughness threshold of 3.5 was used to incorporate  
224 90% of the data, and a slope threshold of  $6^\circ$  was used to avoid misinterpreting degraded paleo-  
225 cliffs, locally dipping as little as  $\sim 7^\circ$ - $9^\circ$ . After removing the SCM surfaces that were clearly not  
226 terraces (e.g. man-made surfaces), high-confidence terraces were separated from more  
227 speculative terraces (strong and light blue polygons in Fig. 2). The speculative surfaces are those  
228 that either; 1) have a parallel orientation to the local bedding, typically in the Neogene  
229 sediments; 2) are not part of a staircase sequence; 3) are dipping slightly landwards rather than  
230 seawards; 4) have a geometry/location that could indicate they are river terraces; 5) might have  
231 been affected by agricultural structures; or 6) have a slope higher than the SCM threshold, but  
232 geometry, position, relative slope and imagery suggest that they are indeed marine terraces.

233 The shoreline angle of marine terraces, at the intersection of paleo-cliffs and paleo-  
234 platforms, approximates the sea level during former highstands (Lajoie, 1986; Merritts and Bull,  
235 1989), but colluvial wedges caused by cliff degradation often obscure this angle (Hanks et al.,  
236 1984). For precise measurements of shoreline angles, we used TerraceM (Jara-Muñoz et al.,  
237 2016). Measurements were restricted to the eastern coast, where staircase sequences of terraces  
238 are least disturbed by cross-cutting faults and display a complete section between the coast and  
239 the upper rasa (Fig. 2). From field observations and DSM measurements, we estimated the slope  
240 of the present-day sea-cliff (Fig. 3) and used it as a proxy for the slope of eroded, older sea-cliffs.  
241 We obtained a range of possible shoreline angle elevations for every terrace in the selected area,  
242 by picking a most seaward/landward position of the paleo sea-cliff on the maximum elevations  
243 of cliff-perpendicular swath profiles and two points of the paleo-platform (Fig. 3; see De Gelder  
244 et al., 2015). We made a lateral correlation of terrace levels by comparing elevations of shoreline  
245 angles and their morphological characteristics (e.g., platform width, cliff height).



247 **Figure 3. Field photographs and shoreline angle analysis.** Locations are given in Fig. 2. (a)  
 248 *Section of the basin stratigraphy taken E of Mitata; dashed line outlines the topmost rasa; (b)*  
 249 *View of terraces and the fault south of Agia Pelagia along the NE-coast; (c) Detail of active fault*  
 250 *scarp near Agia Pelagia; (d) Example of active N-S striking normal fault near Livadi; (e)*  
 251 *Example of active NW-SE striking fault W of Avlemonas (f) Holocene cliff measurements along*  
 252 *NE-coast; (g) Detail of marine terrace along the NE-coast near Plateia Ammos; (h) Example of*  
 253 *shoreline angle analysis estimating a maximum (top) and minimum (bottom) position of the*  
 254 *shoreline angle (SA).*

## 255 **4 Results**

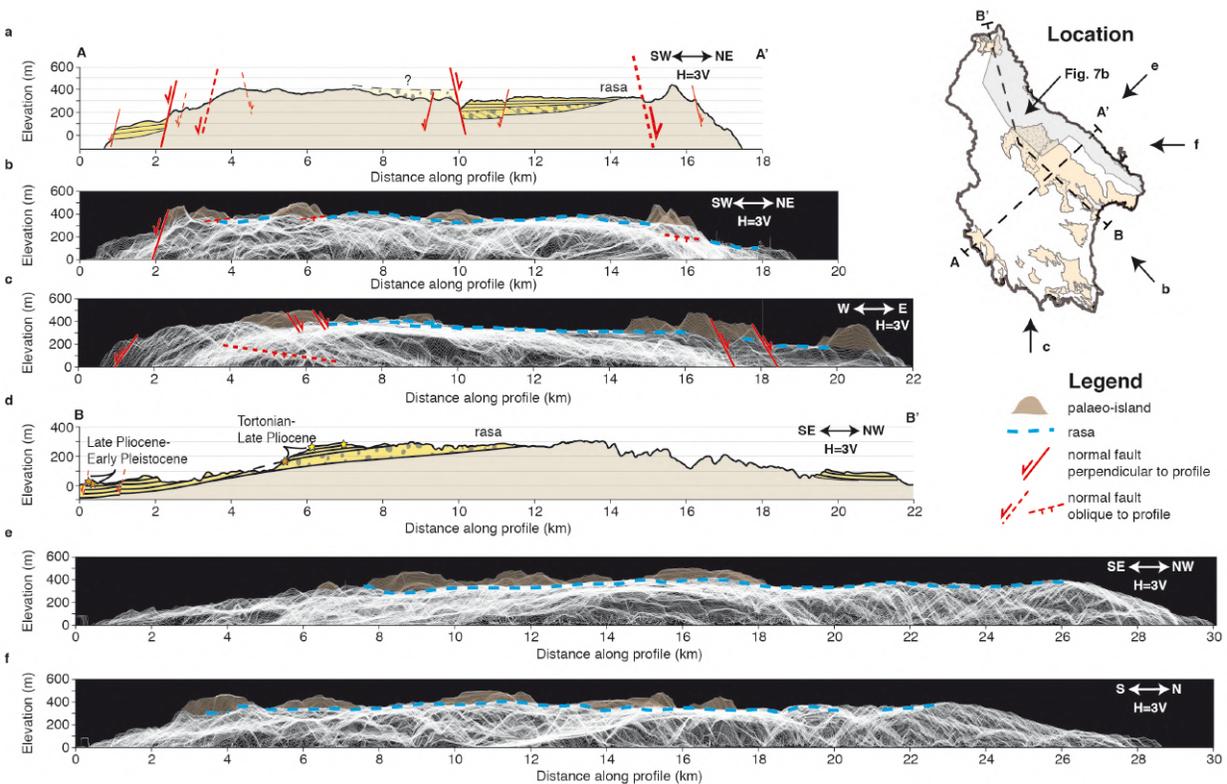
### 256 4.1 Active faults

257 The geometry of most Neogene-Quaternary sedimentary basins is controlled by NW-SE  
 258 trending normal faults. The largest basin, between Potamos and Avlemonas (Fig. 2; hereafter  
 259 Potamos-Avlemonas Basin), is a half-graben bounded by faults along its SW margin (Fig. 4a),  
 260 similar to the smaller basin near the coast east of Karvounades. The basins in the south of the  
 261 island, bounding the SW coast and around Kalamos, Livadi and Kapsali, are partly onlapping on  
 262 the basement, and partly bounded by NW-SE trending, and to a lesser extent N-S trending,  
 263 normal faults. The basin in the north of the island, W of Plateia Ammos, is mostly onlapping on  
 264 the basement, apart from its eastern margin, which is bounded by a N-S trending normal fault  
 265 dipping to the west. N-S trending faults are also cross-cutting the Potamos-Avlemonas Basin,  
 266 and the basement at several locations on the island. The area north of Potamos is largely devoid  
 267 of fault scarps, which are probably unpreserved in the metamorphic rocks of the Arna Unit.  
 268 Well-preserved fault scarps on both N-S and NW-SE trending faults indicate that both fault sets  
 269 are active (examples in Fig. 3b-e). For the fault scarps that contained kinematic indicators,  
 270 motion suggested nearly pure dip-slip motion along both sets of normal faults.

271 Taking the overall slope direction of the bathymetry, SW sloping directions towards the  
 272 Hellenic Trench are the most common (Fig. 5a). When filtering out for higher slopes, which  
 273 should be more representative of active faults, slope directions (dips) to the E and W are  
 274 dominant (Figs. 5b, c), suggesting that ~N-S striking faults are more active. The separation of  
 275 our mapped faults into distinct fault segments 200 m in length (Fig. 5d) indicates that N-S

276 striking faults are most dominant on the island, especially E-dipping ones. NW-SE trending  
 277 faults and trends in between N-S and NW-SE are less numerous, albeit only slightly. This result  
 278 is very similar to the faults we mapped offshore in the SW Hellenic Arc (Fig. 5e), which trend N-  
 279 S predominantly and NW-SE to a lesser extent.

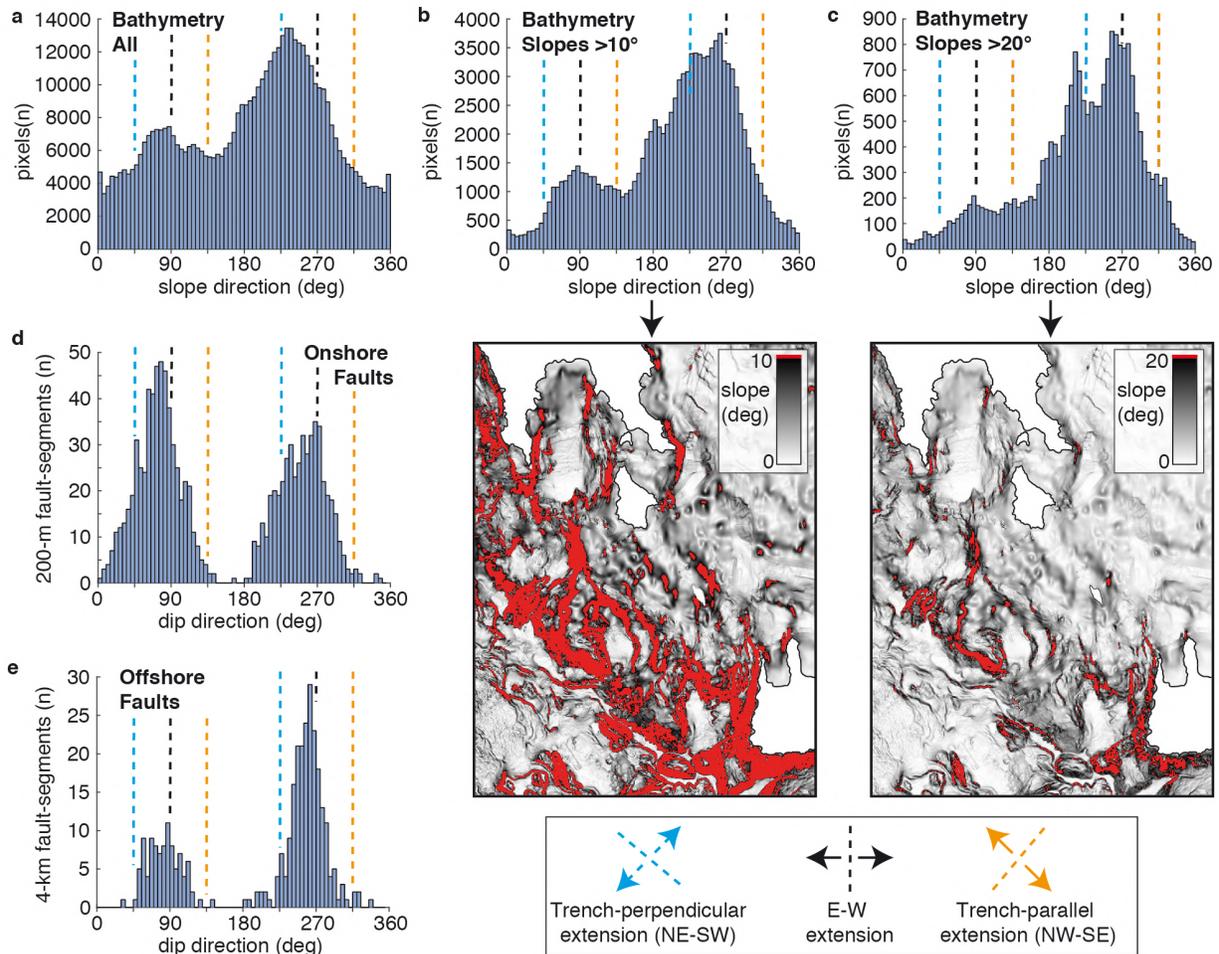
280 These findings collectively indicate that extension is mostly E-W directed, and thus  
 281 oblique to the trench direction, both on Kythira and on the scale of the SW Hellenic Arc. Trench-  
 282 perpendicular NE-SW extension has played a dominant role in the formation of Neogene-  
 283 Quaternary basins on Kythira bounded by NW-SE faults (Figs. 2, 5). It is still or again active, but  
 284 is less important in accommodating active deformation. Trench-parallel NW-SE extension is  
 285 minor or absent, both on the scale of Kythira and the scale of the SW Hellenic Arc.



286  
 287 **Figure 4. Cross-sections and island-scale stacked swath profiles.** Top-right map shows  
 288 locations of profiles in A-A' and B-B' on Kythira, as well as the viewing directions of the full-  
 289 island stacked swaths in panels b, c, e and f, and the stacked swath location of Fig. 7b (a)  
 290 NE-SW Cross-section perpendicular to major basin bounding faults. (b) Stacked swaths along a NE-  
 291 SW strike. (c) Stacked swaths along an E-W strike. (d) NW-SE Cross-section along-strike of the

292 major basin bounding faults. (e) Stacked swaths along a NW-SE strike (f) stacked swaths along a  
 293 N-S strike.

294



295

296 **Figure 5. Fault dip directions and slope directions.** (a) Slope direction for all bathymetric data  
 297 points NE of the Hellenic Trench in Fig. 2. To a first order, the slope direction of steep slopes  
 298 should be perpendicular to the strike of active faults. (b) Same as a, but filtering data points that  
 299 slope more than 10°, as plotted in map-view below. (c) Same as b, but for 20°. (d) Fault dip  
 300 directions of 200-m fault segments mapped on Kythira in Fig. 2. Fault dip directions should be  
 301 perpendicular to the strike of active normal faults (e) Fault dip directions of 4-km normal fault  
 302 segments mapped within the bathymetry of the SW-Hellenic Arc of Fig. 1.

## 303 4.2 Sedimentary basins

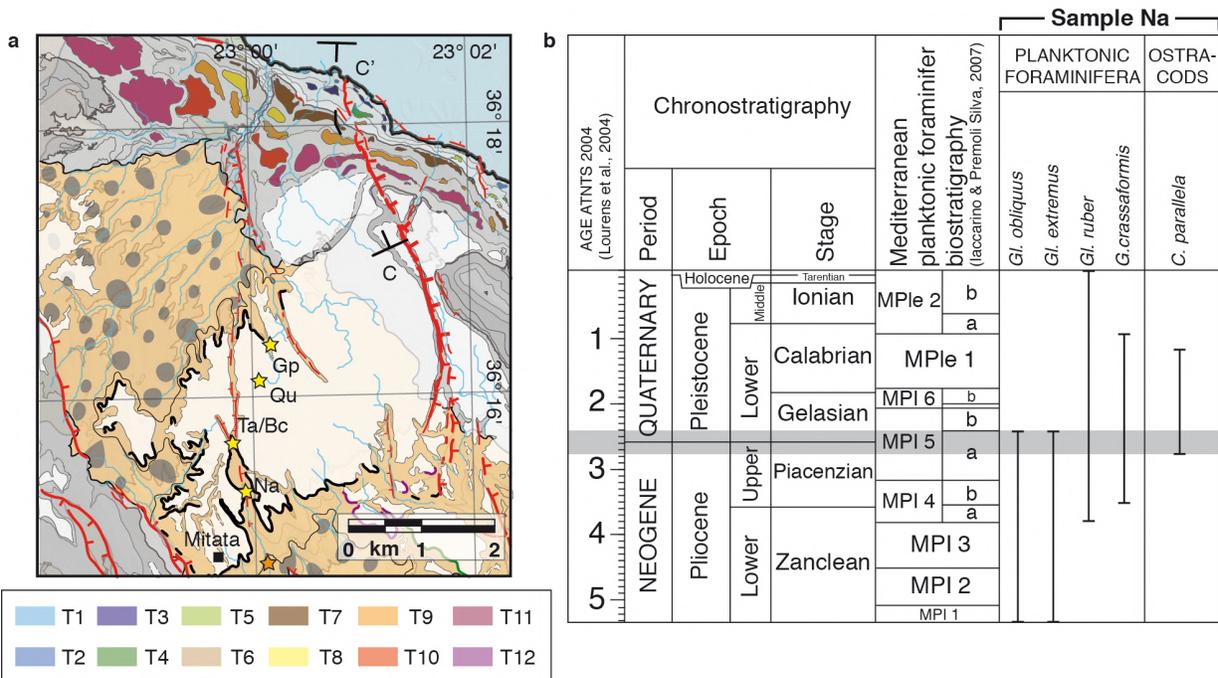
304 The stratigraphically lowest sedimentary infill in central Kythira consists of continental  
305 clastic deposits of variable grain size, including conglomerates, sandstones, siltstones and clays,  
306 and reaching a maximum thickness of ~100 m SE of Potamos (Figs. 2, 4a, d). The conglomerates  
307 that dominate the base of this section become gradually sparser up-section. These conglomerates  
308 are clast supported, and contain sub-rounded, polymict, and poorly sorted clasts of the three  
309 basement units, locally carving the basement with channel geometries. Imbrications and dip of  
310 infill deposits in those channels indicate transport towards the NE, similar to the present-day  
311 river reaching the coast near Agia Pelagia (Fig. 2). East of Mitata, the overlying marine deposits  
312 are deposited conformably on top of the continental deposits (Figs. 3a and 4a), in a gradual up-  
313 sequence transition from continental to coastal to shallow marine. Along the NE-margin of the  
314 Potamos-Avlemonas Basin, marine deposits are deposited directly on top of the basement (Fig.  
315 4a). The marine deposits around Mitata and the rasa consist of limestones, sandstones and marls  
316 and are rich in *Pectinidae*, oysters and other macro-fossils, indicating a shallow marine  
317 environment. The marine deposits close to Avlemonas city consist mostly of laminated and  
318 homogeneous marls intercalated with coarse sandy limestone beds, and are poor in macro-fossils,  
319 possibly indicating a deeper marine environment.

320 The marine deposits are generally dipping ~0-10° to the SE (Fig. 4), and combined with  
321 the geometry of the continental-marine transition, indicate that the Potamos-Avlemonas basin is  
322 consistently 100-200 m thick, with larger content of continental deposits in its NW margin than  
323 towards its SE margin, where marine deposits dominate (Fig. 4d).

## 324 4.3 Micropaleontological analyses and biostratigraphy

325 Planktic and benthic foraminifers and ostracods were analyzed from five sites (Fig. 6a;  
326 Supplementary Table 1). Scanning electron microscope images of selected foraminifers and  
327 ostracods are presented in Supplementary Fig. 2. Ostracods were recovered in all the collected  
328 samples, while benthic foraminifers were absent in sample Qu. In general, they were well  
329 preserved and abundant, albeit not taxonomically diverse. Benthic foraminiferal assemblages  
330 consisted of a few reworked (Miocene) species, some eurybathic species and some other species  
331 typical of the infralittoral environment (approximately 0-50 m depth), such as *Elphidium* spp. (*E.*  
332 *advenum*, *E. complanatum*, *E. crispum*, *E. macellum*), *Asterigerinata planorbis*, *Aubignyana*

333 *perlucida*, *Lobatula lobatula*, and *Cancris oblongus* (Morkhoven et al., 1986; Sgarella and  
 334 Montcharmont Zei, 1993; Di Bella, 2010). Ostracod assemblages, mainly consisting of *Aurila*  
 335 spp., *Callistocythere* spp., and *Semicytherura* spp., confirm a shallow infralittoral environment  
 336 with vegetated bottom. Conversely, planktic foraminifers were collected only in sample Na,  
 337 where several reworked species were identified together with some long-ranging (Miocene-  
 338 Recent) species, and a few markers, such as *Globorotalia crassaformis*, *Globigerinoides*  
 339 *extremus*, *Globigerinoides obliquus* and *Globigerinoides ruber*. On the base of their stratigraphic  
 340 distribution (Lirer et al., 2019), the co-occurrence of these species in sample Na and the presence  
 341 of the ostracod *Callistocythere parallela* in samples NA, Bc and Ta (Ruggieri and D'Arpa, 1993;  
 342 D'Arpa and Ruggieri, 2004) constrain the age of the deposits immediately below the coastal rasa  
 343 to the late Piacenzian-early Gelasian (2.8-2.4 Ma; Fig. 6b). Given its elevation of ~325 m above  
 344 present sea level, this gives an average long-term uplift rate of ~0.13 mm/yr. This value is  
 345 averaged over the last ~2.6 Ma, and should be taken as a minimum value.



346  
 347 **Figure 6. Biostratigraphic dating results.** (a) Zoom-in of Fig. 2 with location of samples used  
 348 for biostratigraphic dating, as well as map of marine terraces for which shoreline angles were  
 349 determined (Fig. 7a), and the location of profile C-C' (Fig. 8c) (b) Mediterranean planktic  
 350 foraminifer biostratigraphic scheme of Pliocene-Quaternary, with the distribution of selected  
 351 species of planktic foraminifers and ostracods recovered in sample Na. The grey band

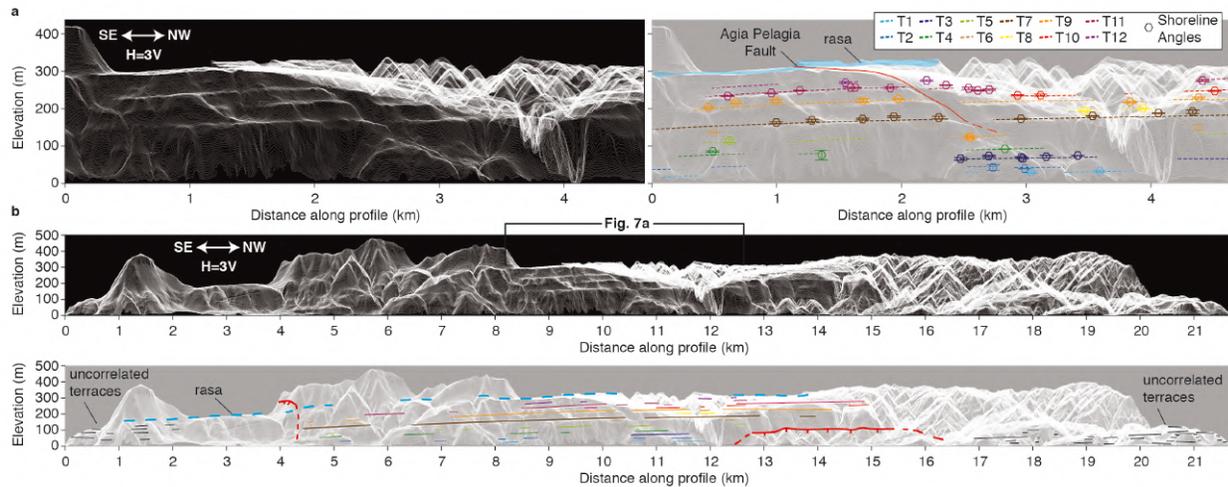
352 *corresponds to the constrained age of the deposits immediately below the rasa. Timescale is*  
353 *after ATNTS 2004 (Lourens et al., 2004), and biostratigraphy after Lirer et al. (2019).*

#### 354 4.4 Marine terraces and rasa

355 The wide top-most rasa covers a large part of the island (Figs 2, 4), but is not preserved  
356 above the metamorphic Arna Unit in the northern part of the island. The very few topographic  
357 highs elevated above the rasa would have been paleo-islands when sea-level was around rasa  
358 elevations. These highs roughly follow the NW-SE trend of the basement nappes (Figs. 2, 4). As  
359 the rasa is only a few meters above the dated shallow marine deposits, it must be of younger age,  
360  $\leq 2.8\text{-}2.4$  Ma (see above).

361 As for the rasa, the marine terraces are preserved within the Pindos and Tripolis basement  
362 units as well as in the sedimentary basins, but not within the metamorphic Arna Unit. Apart from  
363 the rasa caprock mentioned in section 3.3, similar caprocks of coastal deposits appear at the  
364 highest terrace levels within the Potamos-Avlemonas Basin (Fig. 2), in a low-angle unconformity  
365 with underlying marine basin deposits. Along the NE coast a thin layer of coastal deposits  
366 overlying the wave-cut platform exists at some locations (e.g. Fig. 3c), well cemented and  
367 containing sparse shell fragments. This layer typically dips seaward  $<5^\circ$ , which is commonly  
368 much less than the bedding of the deformed basement.

369 Our detailed analysis of shoreline angles (Supplementary Fig. 4) reveals that the  
370 terraces along the NE coast are slightly tilted along the strike of the coastline over a distance of  
371  $\sim 10$  km (Fig. 7). The best-preserved and most continuous terraces are T7, T9 and T11 (Figs. 6b,  
372 7), and decrease in elevation by  $\sim 25\%$  over this distance. Other terraces are more scattered but  
373 can be correlated laterally following a similar trend. The continuity of marine terraces within the  
374 footwall of the Agia Pelagia Fault (Fig. 7a) indicates that it does not affect footwall uplift  
375 significantly. When looking at the scale of the whole island along four different projections the  
376 rasa is tilted eastward within the northwards view (Fig. 4c), varying from  $\sim 400$  m elevation in  
377 the west to  $\sim 200$  m in the east over a distance of  $\sim 15$  km, and at a relatively constant  $\sim 400$  m  
378 elevation in the westwards view (Fig. 4f). The trends in rasa elevation in the other viewing  
379 directions are more complicated (Figs. 4b, 4e). The above supports an overall island uplift with  
380 an approximately E-ward tilt. This is best compatible with a N-S striking source of uplift to the  
381 west of Kythira, which can also explain the along-coast tilt we see in the studied terraces (Fig. 7).



382  
 383 **Figure 7. Terraces along the NE coast.** (a) Detail of stacked swath profile showing area for  
 384 which shoreline angles were determined. Thin white lines represent 300 parallel swath profiles  
 385 of 10 m width, with arrows indicating orientation of profiles. Shoreline angles and error bars of  
 386 the twelve terrace levels (T1 to T12) are derived similarly to the example given in Fig. 3c, with  
 387 details provided in Supplementary Fig. 4 (b) Stacked swath profile of whole NE-coastline (top  
 388 plot; location in Fig. 4), including trends of marine terraces and rasa (bottom plot).

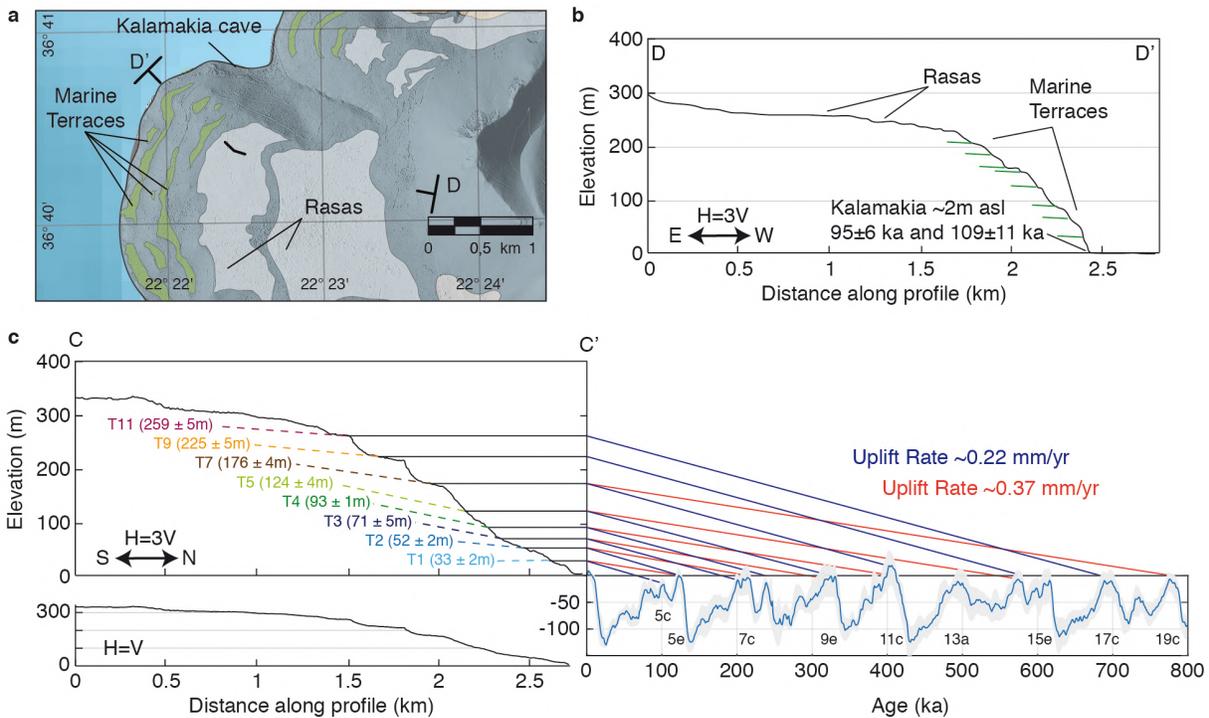
## 389 5 Discussion

### 390 5.1 Terrace correlation

391 Marine terraces are mainly formed during major interglacial sea-level highstands (e.g.  
 392 Lajoie, 1986; Anderson et al., 1999). The most commonly recorded highstand within marine  
 393 terrace sequences worldwide is the interglacial Marine Isotope Stage (MIS) 5e (Kopp et al.,  
 394 2009, Pedroja et al., 2011; 2014). The MIS 5e peaked around  $124 \pm 5$  ka (e.g. Stirling et al., 1998;  
 395 Masson-Delmotte et al., 2010) with eustatic sea-levels around  $6 \pm 4$  m higher than today (Murray-  
 396 Wallace and Woodroffe, 2014). Given this, we expect one of the terraces of the sequence on  
 397 Kythira to have been formed during MIS 5e. However, reliable ages of marine terraces on  
 398 Kythira are lacking, making terrace attribution difficult. We thus estimate possible correlations  
 399 of marine terraces to highstands based on the similar morphostratigraphy in the S-Peloponnese  
 400 (Fig. 8).

401 Sequences of marine terraces on the Mani and Vatika peninsulas (Fig. 1) are also capped  
 402 by wide Plio-Pleistocene rasas (e.g. Kelletat et al., 1976; Dufaure, 1977; Kleman et al., 2016) at

403 comparable elevations (250-300 m asl) to the rasa in Kythira Island. The only radiometric dates  
404 in the region were taken at the Kalamakia cave on the Mani Peninsula (Fig. 8a). There, marine  
405 deposits ~2 m above present sea level are tentatively dated with U-Th as MIS 5c (~100 ka) (De  
406 Lumley et al., 1994). Assuming that this age is accurate, as well as constant uplift rates and a  
407 eustatic sea-level during MIS 5c of 5-40 m below present sea level (Fig. 8c; Spratt and Lisiecki,  
408 2016), MIS 5e should be at an elevation of 8-54 m above present sea level. On the Vatika  
409 Peninsula, biostratigraphic evidence for Tyrrhenian deposits (~80-130 ka) is found on marine  
410 terraces at different elevations up to ~20 m above the present-day sea-level (Dufaure, 1970;  
411 Kelletat et al., 1976; Kowalczyk et al., 1992). The highest and oldest of those deposits (named  
412 Euthyrenian; Dufaure, 1970) would correspond to MIS 5e, now at ~20 m elevation in this  
413 peninsula. Assuming an uplift rate of the same order of magnitude for Kythira, we estimate the  
414 lowest terrace (T1, 31-35 m) or T2 (50-54 m) to correspond to MIS 5e. Taking into account MIS  
415 5e elevation and age uncertainties mentioned above, this would imply uplift rates of  $\sim 0.22 \pm 0.04$   
416 mm/yr or  $\sim 0.37 \pm 0.04$  mm/yr, or approximately 0.2-0.4 mm/yr. We extrapolated 0.22 and 0.37  
417 mm/yr uplift rates to correlate marine terraces at higher elevations to older marine sea-level  
418 highstands up to 800 ka (Fig. 8c). If uplift rates have indeed been constant between the T1/T2  
419 terraces and the highest well expressed marine terrace (T11), the latter would have an age of  
420  $\sim 1.2$ - $0.7$  Ma. If uplift rates have been constant between the T1/T2 terraces and the rasa, the latter  
421 would have started emerging from the sea  $\sim 1.5$ - $0.9$  Ma.



422

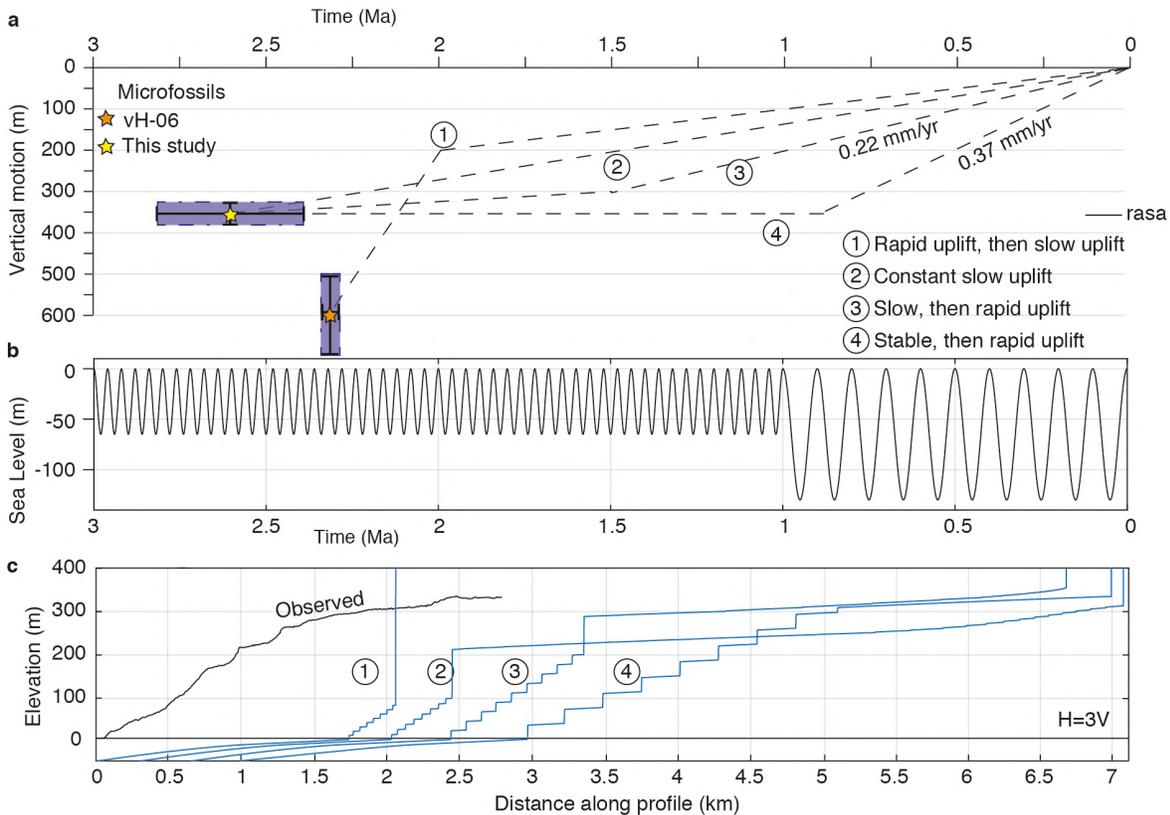
423 **Figure 8. Proposed correlation of marine terraces to sea-level highstands.** (a) Map of  
 424 terraces/rasas around the Kalamakia cave on the Mani Peninsula (location in Fig. 1) (b)  
 425 Topographic profile D-D' of Kalamakia terrace sequence, derived from a Pleiades-DSM (as  
 426 used for Kythira) with radiometric ages. (c) Correlation of main terrace levels to sea-level  
 427 highstands, using the eustatic sea-level curve of Spratt and Lisiecki (2016) and the graphical  
 428 method of Bloom and Yonekura (1990) adopting the shoreline angle elevations of profiles 31, 42  
 429 (T1), 18, 42 (T2), 15, 17, 19, 29, 30 (T3), 19 (T4), 16, 17 (T5), 14, 28 (T7), 10, 11 (T9), 11 and  
 430 22 (T11) (see Supplementary Fig. 4). The location of profile C-C' is specified in Fig. 6a.

## 431 5.2 Uplift history

432 Whereas the age and elevation of the rasa indicate an uplift rate of ~0.12 mm/yr averaged  
 433 over ~2.6 Ma for the center of the island (section 4.3), the Late Quaternary uplift rate estimates  
 434 based on marine terraces are distinctly higher (0.2-0.4 mm/yr). Previous paleobathymetry  
 435 estimates based on microfossils suggested another uplift history: initial rapid uplift of 2 mm/yr  
 436 after ~2.3 Ma, followed by slower uplift of ~0.1 mm/yr since ~2 Ma (Van Hinsbergen et al.,  
 437 2006). To discriminate between these potential scenarios of surface uplift, we used a simple  
 438 landscape evolution model. We test 4 different scenarios (Fig. 9a): initial very rapid uplift (2

439 mm/yr) followed by slow uplift (0.1 mm/yr); constant slow uplift (0.12 mm/yr); initial slow  
440 uplift (0.05 mm/yr) followed by faster uplift (0.22 mm/yr); and stable conditions followed by  
441 rapid uplift (0.37 mm/yr). We use the cliff erosion model in TerraceM (Jara et al., 2019; adapted  
442 from Anderson et al., 1999), with standard parameter values for wave height (4 m), erosion rate  
443 (0.5 m/yr) and an initial slope based on the approximate slope of the sequence (12°). For  
444 simplicity, and to avoid biased results by sea-level noise inherent to long-term (>1 Ma) sea-level  
445 curves (De Gelder et al., 2020), we approximate Quaternary sea-level by combining a 40 ky-  
446 period, 65 m-amplitude sine function (2.6 - 1 Ma) and a 100 ky-period, 130-m amplitude sine  
447 function (Fig. 9b). Additional tests with different parameter values and published Plio-  
448 Quaternary sea-level curves are given in Supplementary Fig. 3, and do not change the main  
449 modeling outcomes.

450         The resulting first order morphology can be reproduced by the third and fourth scenarios.  
451 The first uplift scenario does not match the observed terrace sequence, as the decelerating uplift  
452 rate leads to the formation of a high cliff, since increased erosion due to re-occupation removes  
453 the older terraces. Slower uplift rates imply that terrace carving is distributed over a narrower  
454 vertical range, a process that was recently illustrated by Malatesta et al. (2021). The second  
455 scenario does not match the observed profile either. The change from ~40 ky period low-  
456 amplitude sea-level oscillations to ~100 ky period high-amplitude sea-level oscillations around  
457 the Mid-Pleistocene Transition (~1 Ma; Clark et al., 2006) should lead to a major change in  
458 morphology, which is not recorded by the marine terrace sequence in Kythira. In published Plio-  
459 Quaternary sea-level curves (Bintanja and V.d. Wal, 2008; De Boer et al., 2010; Bates et al.,  
460 2014 and Rohling et al., 2014) the change in sea-level oscillations is less abrupt, but also using  
461 those sea-level curves instead of the simplified one, big cliffs are at least twice as high as any  
462 cliff observed in the coastal morphology of Kythira (Supplementary Fig. 3). The third and the  
463 fourth scenarios are more compatible with the observed morphology of a continuous marine  
464 terrace sequence culminating in a wide *rasa* around 300 m elevation. Accounting for this, and for  
465 the ~0.22 and ~0.37 mm/yr uplift rates that are compatible with the marine terrace elevations, we  
466 favour an uplift history with initial stable conditions or slow uplift of Kythira from the sea,  
467 followed by faster uplift since ~1.5-0.7 Ma.



468

469 **Figure 9. Landscape evolution modeling under different uplift scenarios.** (a) Different vertical  
 470 motion scenarios, with scenario 1 assuming 400 m of rapid uplift followed by 200 m of slow  
 471 uplift, as suggested by Van Hinsbergen et al. (2006), and scenarios 2, 3 and 4 assuming the  
 472 dated samples of this study are equivalent to maximum transgression at ~2.6 Ma, and either  
 473 steady slow uplift since then (scenario 2), a slow uplift followed by a rapid uplift (scenario 3 -  
 474 uplift rate 0.22 mm/yr as estimated in Fig. 8) or stable conditions followed by a rapid uplift  
 475 (scenario 4 - uplift rate 0.37 mm/yr as estimated in Fig. 8). (b) Extremely simplified sea-level  
 476 curve used for the modeling (following De Gelder et al., 2020). (c) Modeling results for the  
 477 different scenarios, compared to the terraced topography as given in Fig. 8. Parameter values  
 478 are given in Supplementary Table 2.

479

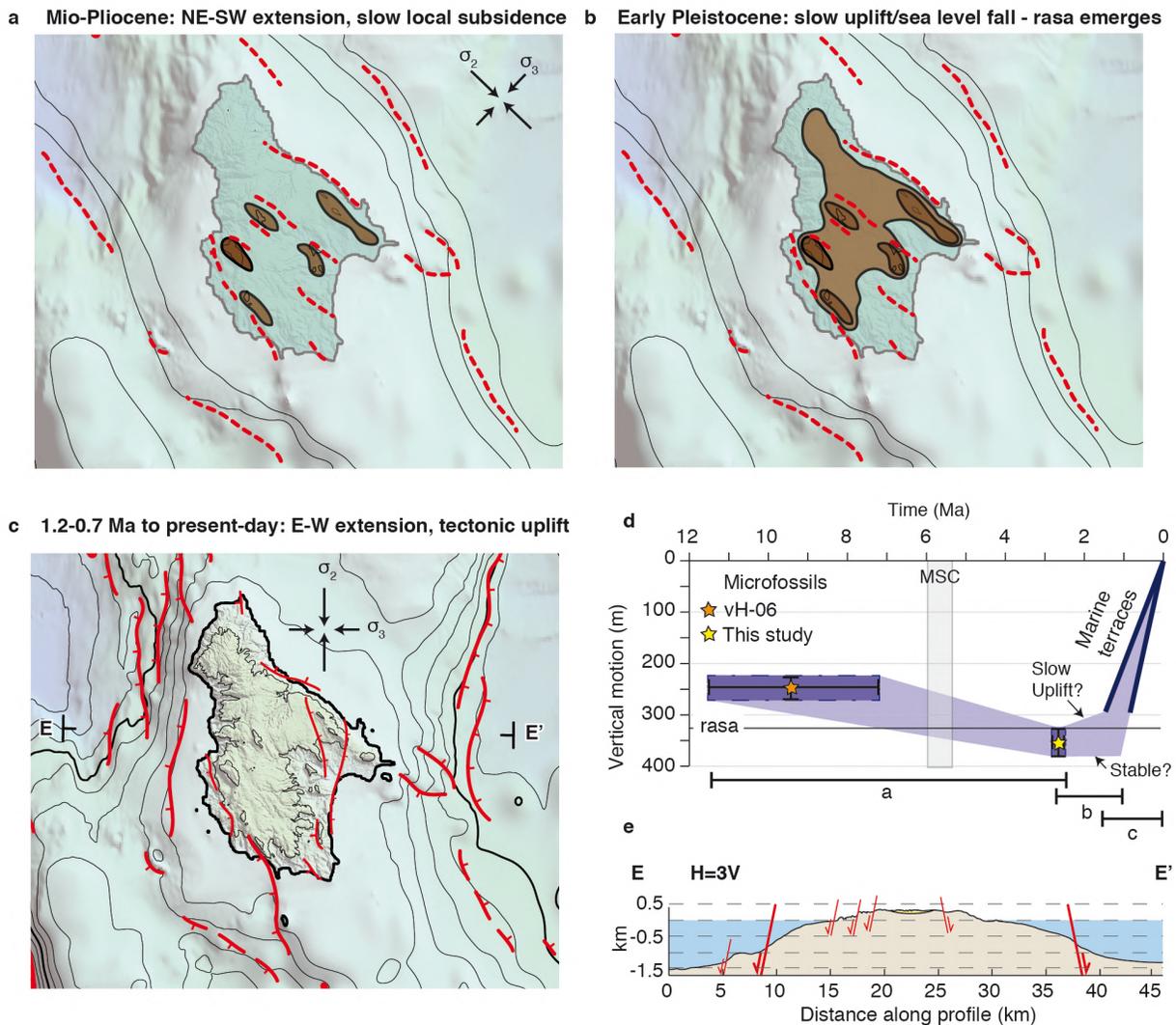
### 5.3 Miocene-Recent tectonics of Kythira Island

480

481 The sedimentary infill in the Potamos-Avlemonas Basin either infilled a former  
 482 paleotopography after nappe stacking, or deposited in relation with the NW-SE striking normal  
 faults that currently bound the basin. We favor the latter, fault-controlled sedimentation, on the

483 basis of the lack of sedimentary deposits within the footwall of the basin-bounding normal faults  
484 (Fig. 4). This in turn supports that the ~100 m of relative subsidence between the Tortonian  
485 shallow marine deposits close to Mitata (Van Hinsbergen et al., 2006), and the shallow marine  
486 Plio-Pleistocene deposits dated in this study a few kms to the north (Figs. 2, 10), are largely of  
487 local origin and due to NE-SW extension. The geometry, depositional environment and ages of  
488 the sediments indicate that the coastline gradually shifted towards the NW with increasing  
489 subsidence, while the ~200 m thick basin transitioned laterally from continental conditions in the  
490 NW to marine conditions in the SE (Fig. 4). Within this framework, the Plio-Pleistocene section  
491 near Avlemonas (Fig. 2) (Van Hinsbergen et al., 2006) should have been deposited in a deeper  
492 marine environment than our dated deposits immediately below the *rasa*. Their vertical motion  
493 trend is very similar to ours, but when compared to our data, their paleo-depth estimate suggests  
494 much larger vertical motions (~400 m subsidence, ~800 m uplift). We did not find clear geologic  
495 or geomorphic evidence of marine conditions higher than the topmost *rasa*, and therefore suspect  
496 that the paleo-depth range of the fossils they considered is biased, possibly as a result of local  
497 upwelling. We did not observe the angular unconformity between continental and marine  
498 deposits described by Meulenkamp et al. (1977). As an explanation, we speculate that the  
499 Mediterranean sea-level drop of tens to hundreds of meters during the Messinian Salinity Crisis  
500 (MSC; 5.97-5.33 Ma) (Roveri et al., 2014) may have locally resulted in erosive/unconform  
501 sedimentary contacts near former river valleys, as described elsewhere in the Mediterranean  
502 (Bache et al., 2012).

503 We interpret the elevations and overall concave morphology of the terraces and overlying  
504 *rasa* as indicative of land motions at pace with sea-level variations or as a slow apparent uplift  
505 followed by faster uplift around ~1.5-0.7 Ma. The age of the *rasa* is coeval to that of the Mid-  
506 Pliocene Warm Period (~3.3-2.9 Ma; Roveri et al., 2014), after which eustatic sea-levels may  
507 have dropped by a few tens of meters (Raymo et al., 2011; Dutton et al., 2015). Therefore, an  
508 initial apparent uplift of Kythira could have been an effect of eustatic sea-level fall, and hence of  
509 climatic rather than tectonic origin. The rate and magnitude of the second uplift phase cannot be  
510 accounted for by eustatic sea-level changes alone, and implies a recent major change in tectonic  
511 conditions.



512

513 **Figure 10. Schematic reconstruction of Miocene-Recent tectonic history of Kythira. (a)**514 *Reconstruction for the phase of NE-SW extension during the Tortonian-Pliocene, resulting in*515 *NW-SE oriented topographic ridges and a large, coastal rasa in the latest Pliocene. (b)*516 *Reconstruction for the Early Pleistocene, with slow uplift resulting in emergence of the rasa*517 *from the sea (c) Reconstruction for the present-day setting, with dominantly E-W extension and*518 *uplift of the island as the results of offshore N-S trending normal faults. (d) Vertical motion,*519 *assuming a paleo-depth of our samples and the shallow marine Tortonian samples of Van*520 *Hinsbergen et al. (2006), of 0-50 m, and a marine terrace uplift rate of 0.22-0.37 mm/yr.*

## 521           5.4 Forearc uplift and regional implications

522           The N-S trending normal faults cutting through the topmost rasa and the Tortonian-  
523 Pliocene basin, and the associated surface uplift, post-date ~2.8-2.4 Ma. This supports that the  
524 uplift and overall eastward tilt of the rasa is due to the large N-S trending normal faults offshore  
525 (Fig. 10), and the uplift of Kythira as an asymmetric horst. These offshore faults would thus be  
526 of similar age as the N-S trending faults onshore but with larger vertical throws, >1 km (Fig.  
527 10c). Given uplift to subsidence ratios of ~1:1-2.5 found for normal faults (King et al., 1988;  
528 McNeill et al., 2005, De Gelder et al., 2019), these N-S offshore faults can account for the ~200-  
529 400 m of observed uplift. Whereas we find an extreme lower bound for the onset of E-W  
530 extension and uplift of ~2.8-2.4 Ma, we deem it more likely that these events started as recently  
531 as ~1.5-0.7 Ma.

532           Our results on Kythira and in the offshore SW Hellenic Arc indicate that active extension  
533 is dominantly E-W oriented at a regional scale, at an oblique angle to the Hellenic Trench and  
534 the preceding phase of NE-SW extension. As shown in both analogue (Henza et al., 2011) and  
535 numerical models (e.g. Deng et al., 2018), preceding normal fault systems can be re-activated by  
536 a new phase of extension, even if that new phase has a different extension direction. The overall  
537 zigzag-pattern of the two active fault systems in the W Hellenic Forearc suggests that the NW-  
538 SE striking fault system was moderately well-developed (Henza et al., 2011) when the new  
539 phase of E-W extension initiated.

540           Several authors have interpreted the offshore faults in the SW Hellenic Arc as dominantly  
541 trench-normal and trench-perpendicular normal faults resulting from trench migration and slab-  
542 rollback (e.g. Angelier et al., 1982; van Hinsbergen and Schmidt, 2012; Gallen et al., 2014). E-W  
543 extension is difficult to reconcile with slab-rollback as a driving mechanism, and instead  
544 suggests that another mechanism must have induced the present-day deformation field.  
545 Sedimentary underplating has been highlighted by other studies as a potentially important driver  
546 of vertical motions around the Hellenic Forearc (Gallen et al., 2014; Menant et al., 2020).  
547 Although this can explain the cyclic pattern of Miocene-Pliocene subsidence followed by  
548 Pleistocene uplift on Kythira, there is no obvious explanation to link trench-perpendicular  
549 underplating with trench-oblique E-W extension. As such, we do not consider sedimentary  
550 underplating as a primary mechanism to explain tectonic changes in the SW Hellenic Arc.

551 A process that may explain Pleistocene changes around the forearc is the occurrence of  
552 slab tearing, which has been proposed for both the eastern (e.g. Özbakir et al., 2013) and western  
553 (e.g. Royden and Papanikolaou, 2011; Jolivet et al., 2013) limits of the Hellenic subduction zone.  
554 Given the location of Kythira, especially the W-limit of the Hellenic Arc is potentially relevant.  
555 Analogue models of slab-tearing, treating the Kefalonia Fault Zone (Fig. 11) as the strike-slip  
556 surface expression of the slab tear, suggest toroidal mantle flow would lead to ~NNE-SSW  
557 directed horizontal lithospheric strain in the Hellenic Forearc (Fig. 11; Guillaume et al., 2013).  
558 Although this does not perfectly match E-W extension, strain orientations are close enough that  
559 we cannot exclude the possibility that slab tearing plays a role in the SW Hellenic Arc  
560 deformation.

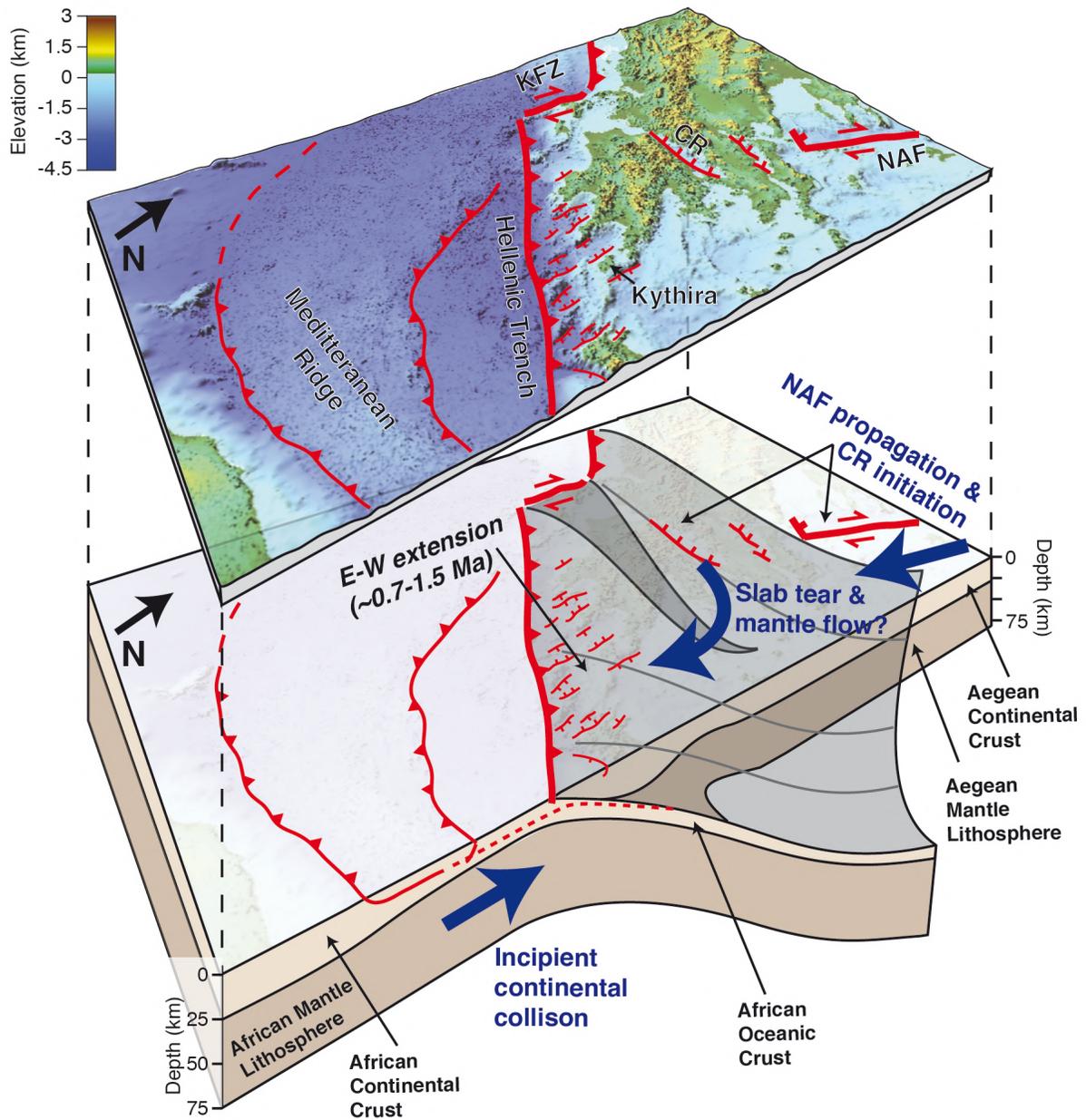
561 As proposed by Lyon-Caen et al. (1988) and Armijo et al. (1992), the change of  
562 extensional regime implies a change in boundary conditions, probably in relation to the nearby  
563 subduction zone. They suggested it could be due to incipient collision of the arc with the  
564 increasingly buoyant crust of the subducting African margin (Fig. 11). A similar tectonic-scale  
565 change in kinematics is supported by a wealth of evidence across the East Mediterranean (e.g.,  
566 Mascle and Chaumillon, 1997; Schattner, 2010; Aksu et al., 2021), and in agreement with a  
567 growing body of observations. These include seismic and tomographic images, local bathymetric  
568 mapping and marine basins uplifted on land, all reporting changes in the stress and deformation  
569 fields throughout the Mediterranean in the last ~5 Ma (e.g., Prada et al., 2020; Zitellini et al.,  
570 2020; Gómez de la Peña et al., 2021).

571 A third possible process to explain tectonic changes in the forearc is the extrusion of the  
572 Anatolian plate towards the Aegean domain, controlled by the westward propagation of the  
573 North Anatolian Fault (NAF; Fig. 11; e.g. Armijo et al., 1996, 1999; Flerit et al, 2004). This  
574 process has been proposed as the cause of localized extension in the Corinth Rift (Fig. 11; e.g.  
575 Taymaz et al., 1991; Armijo et al., 1996; Taylor et al., 2011), which is currently extending at  
576 plate-boundary rates of 1-2 cm/yr (Briole et al., 2000). The initiation of localized rapid extension  
577 in the Corinth Rift around ~2-0.6 Ma (Armijo et al., 1996, Nixon et al., 2016, Gawthorpe et al.,  
578 2017, Fernández-Blanco et al., 2019a), is approximately coeval with the initiation of N-S fault-  
579 driven uplift we propose on Kythira, and we note that river profile inversion on Crete also  
580 suggests accelerated uplift over the past ~2 Ma (Roberts et al., 2013). The coeval timing of these  
581 events support that the recent growth into the Aegean of the propagating NAF changed the

582 convergence rate on the Hellenic Trench, possibly doubling it within the Pleistocene (as  
583 proposed by Flerit et al. 2004). This could have resulted in stronger mechanical coupling on the  
584 deep subduction interface and an increase of the N-S directed horizontal forces, explaining the  
585 recent change in extensional fault geometry in the upper crust.

586 The three processes mentioned above, i.e. incipient collision, NAF propagation and slab  
587 tearing, can account for large-scale tectonic changes in the SW Hellenic Arc individually, but we  
588 emphasize that these processes are non-exclusive. As such, a combination of two or three of  
589 these processes, possibly inter-related, is also a feasible scenario.

590 In a more general sense, our study of Kythira Island highlights the importance of upper  
591 crustal faulting in forearc uplift. As mentioned above, the N-S striking offshore normal faults  
592 explain most, if not all of the observed uplift on Kythira. Deeper seated processes like  
593 underplating (e.g., Menant et al., 2020), dynamic topography (e.g., Conrad and Husson, 2009),  
594 lower-crustal mantle flow (e.g., Fernández-Blanco, 2014; Fernández-Blanco et al., 2020) and  
595 mantle flow in relation to roll-back and/or slab tearing (e.g., Guillaume et al., 2013) are likely to  
596 play a role in any subduction zone around the world, but their relative contributions are difficult  
597 to quantify if the forearc is crosscut with upper crustal faults. This is the case further east in the  
598 Hellenic Forearc (Gallen et al., 2014; Howell et al., 2017) as well as in many other subduction  
599 zones, like the Caribbean (e.g. Leclerc et al., 2014), Japan (e.g. Matsu'ura et al., 2009) and New  
600 Zealand (e.g., Clark et al., 2010). A key factor to distinguish between uplift sources that we can  
601 underline from our study is the deformation pattern and wavelength. We know, from well-  
602 resolved examples like the Corinth Rift, that the deformation wavelength of upper crustal normal  
603 faults is around ~15-20 km (De Gelder et al., 2019). Whereas this is compatible with km-scale  
604 tilting on Kythira, the wavelength of lithospheric/mantle scale processes would be much larger,  
605 possibly by an order of magnitude. Continuous deformation markers such as marine terraces and  
606 rasas can be especially valuable in that sense, and are thus of primary importance in resolving  
607 uplift mechanisms at the front of overthrusting upper plates.



608

609 **Figure 11. Possible regional drivers of E-W extension.** Top plot shows topo-bathymetry with  
 610 main tectonic features, using a Global Mapper (version 15.0) 3D view and a 3x vertical  
 611 exaggeration. Topography is an ALOS Global Digital Surface Model (DSM), and bathymetry is  
 612 from the European Marine Observation and Data Network (EMODNet) Digital Terrain Model  
 613 (DTM). KFZ = Kefalonia Fault Zone, CR = Corinth Rift and NAF = North Anatolian Fault.  
 614 Bottom plot shows the same area but with underlying African and Aegean lithosphere. Marked in  
 615 blue are 3 possible mechanisms, not mutually exclusive, that could have led to E-W extension

616 *since ~1.5-0.7 Ma. This plot is approximately to scale, with African lithospheric structure S of*  
617 *the Hellenic Trench based on Meier et al. (2004), Aegean lithospheric structure based on Jolivet*  
618 *et al. (2013) and schematic geometry of a possible slab tear based on Guillaume et al. (2013).*  
619 *There is some debate on the location of the main plate boundary (e.g. discussion in Shaw and*  
620 *Jackson, 2010), with estimates ranging from the Hellenic Trench to S of the Mediterranean*  
621 *Ridge, but this does not affect our main inferences.*

## 622 **6 Conclusions**

623 Our study of the main tectono-stratigraphic and morphological features of Kythira allows  
624 us to draw the following conclusions:

625 1) Slow local subsidence of ~100 m within the largest sedimentary basin on Kythira  
626 occurred between the Tortonian and Latest Pliocene, possibly largely accommodated by a NW-  
627 SE trending fault system forming a half-graben.

628 2) The highest elevated marine deposits below a wide paleo-coastal rasa constrain  
629 maximum transgression to ~2.8-2.4 Ma, and give a maximum age for the onset of recent uplift.

630 3) Marine regression following rasa-formation occurred in two steps. The first phase  
631 occurs with stable conditions or slow uplift, possibly as an effect of eustatic sea-level drop  
632 instead of a tectonic origin. This phase was followed by a faster uplift phase of ~0.2-0.4 mm/yr  
633 that initiated around 1.5-0.7 Ma.

634 Our evidence supports that this last and ongoing phase of faster uplift is largely driven by  
635 N-S trending normal faults that are part of a regionally dominant mode of active E-W extension.  
636 We attribute the change in tectonic regime and vertical motion around 1.5-0.7 Ma to changing  
637 boundary conditions at the subduction zone in relation to incipient collision with the African  
638 plate, Aegean-Anatolian extrusion and propagation of the North Anatolian Fault, and/or slab  
639 tearing. In all those cases, increased N-S directed horizontal forces could have resulted in a new  
640 direction of extension and active uplift of horsts like Kythira. In general, we emphasize the  
641 importance of upper crustal faulting in dictating forearc uplift patterns.

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#### 654 **Availability Statement**

655 The Pleiades satellite imagery was obtained through the ISIS and Tosca programs of the  
656 Centre National d'Etudes Spatiales (CNES, France) under an academic license and is not for  
657 open distribution. On request, we'll provide the DSM calculated from this imagery to any  
658 academic researcher who gets approval from CNES (contact [isis-pleiades@cnes.fr](mailto:isis-pleiades@cnes.fr) for quoting  
659 this paper, and with [lacassin@ipgp.fr](mailto:lacassin@ipgp.fr) in copy). We do share a georeferenced hillshade image and  
660 slope map of the 2 m-resolution Digital Surface Model that was developed from Pleiades satellite  
661 imagery. This image can be retrieved with these links:

662 <https://doi.org/10.6084/m9.figshare.18715535.v1> (hillshade image)

663 <https://doi.org/10.6084/m9.figshare.18714914.v1> (slope map).

664 The map of Figure 2 can be downloaded in georeferenced format (as Geospatial PDF) at  
665 <https://doi.org/10.6084/m9.figshare.18703496.v1>. The other data that support the findings of this  
666 study are available within the publication, referenced studies and/or from the corresponding  
667 author on request.

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**Quaternary E-W extension uplifts Kythira Island and segments the Hellenic Arc**

G. de Gelder<sup>1,2</sup>, D. Fernández-Blanco<sup>2,3</sup>, N. Öğretmen<sup>4,5</sup>, S. Liakopoulos<sup>6</sup>, D. Papanastassiou<sup>6</sup>, C. Faranda<sup>5</sup>, R. Armijo<sup>2</sup>, and R. Lacassin<sup>2</sup>

<sup>1</sup>ISTerre, Université Grenoble-Alpes, 1381 Rue de la Piscine, 38400 St. Martin d'Hères, France

<sup>2</sup>Université de Paris, Institut de physique du globe de Paris, CNRS, F-75005 Paris, France

<sup>3</sup>Barcelona Center for Subsurface Imaging, Passeig Marítim de Barceloneta 37-49, E-08003 Barcelona, Spain

<sup>4</sup>Eurasia Institute of Earth Sciences, Istanbul Technical University, 34467, Istanbul, Turkey

<sup>5</sup>Dipartimento di Scienze, Università degli Studi Roma Tre, Largo San Leonardo Murialdo, 1, 00146 Rome, Italy

<sup>6</sup>Institute of Geodynamics, National Observatory of Athens, GR-11810 Athens, Greece

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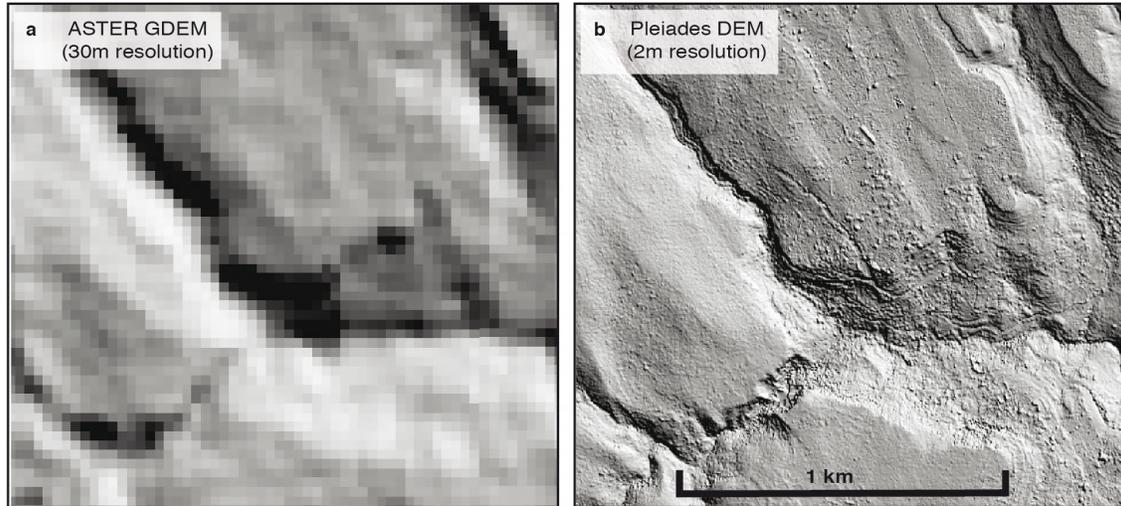
Tables S1 to S2

**Additional Supporting Information (Files uploaded separately)**

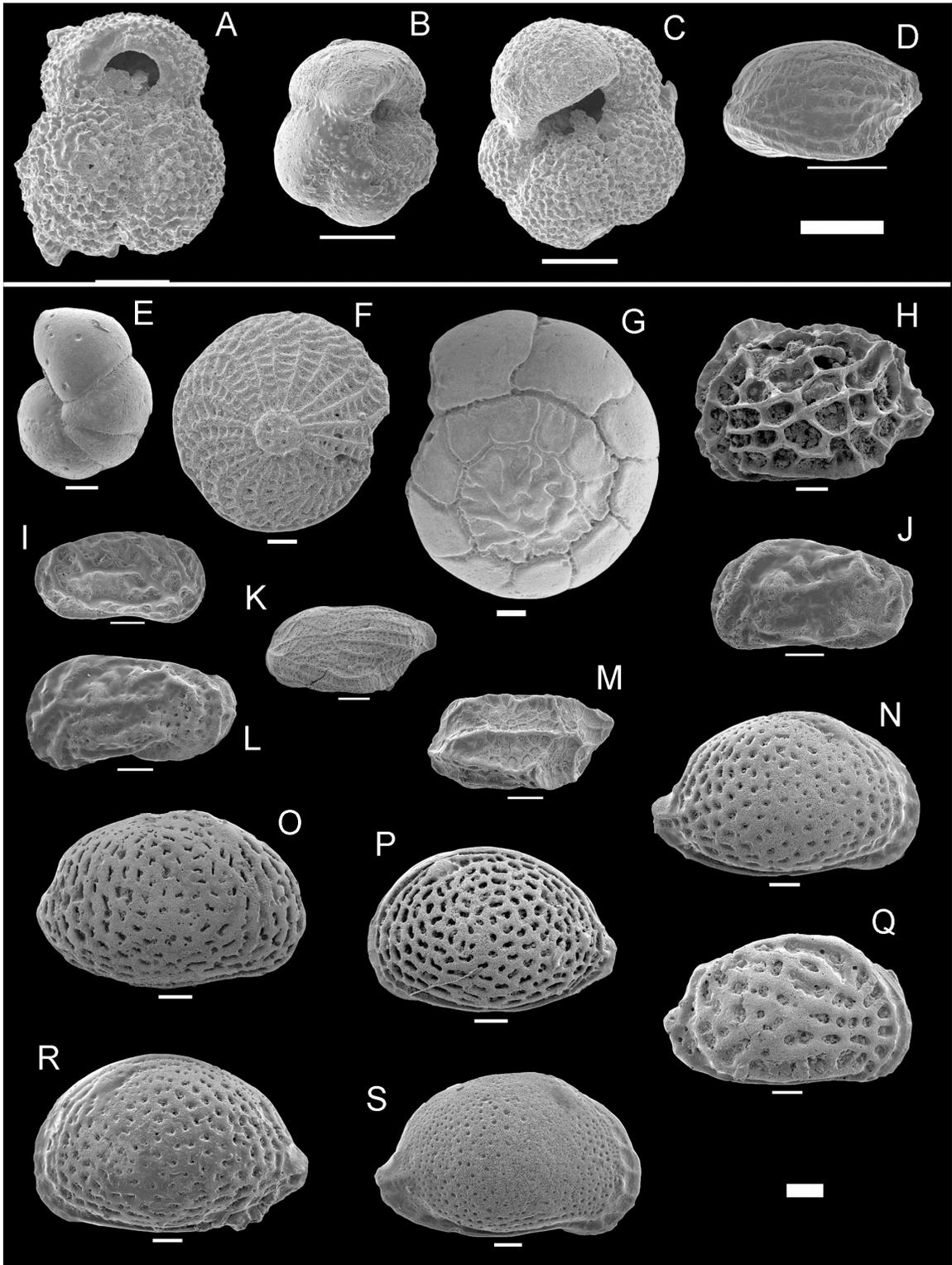
We share a georeferenced hillshade image and slope map of the 2 m-resolution Digital Surface Model through these links: <https://doi.org/10.6084/m9.figshare.18715535.v1> (hillshade image) and <https://doi.org/10.6084/m9.figshare.18714914.v1> (slope map). The map of Figure 2 can be downloaded in georeferenced format (as Geospatial PDF) at <https://doi.org/10.6084/m9.figshare.18703496.v1>.

**Introduction**

Figure S1 is a comparison between Digital Surface Models produced with ASTER or Pleiades imagery (as in this study). Figure S2 shows some Scanning Electron Microscope images for some of the microfossils used in the biostratigraphic dating section. Figure S3 shows sensitivity tests of the landscape evolution modeling. Figure S4 are detailed topographic cross-sections used to determine shoreline angles of marine terraces. Table S1 lists all microfossils found in the samples, and Table S2 the model parameters used for the landscape evolution modeling.

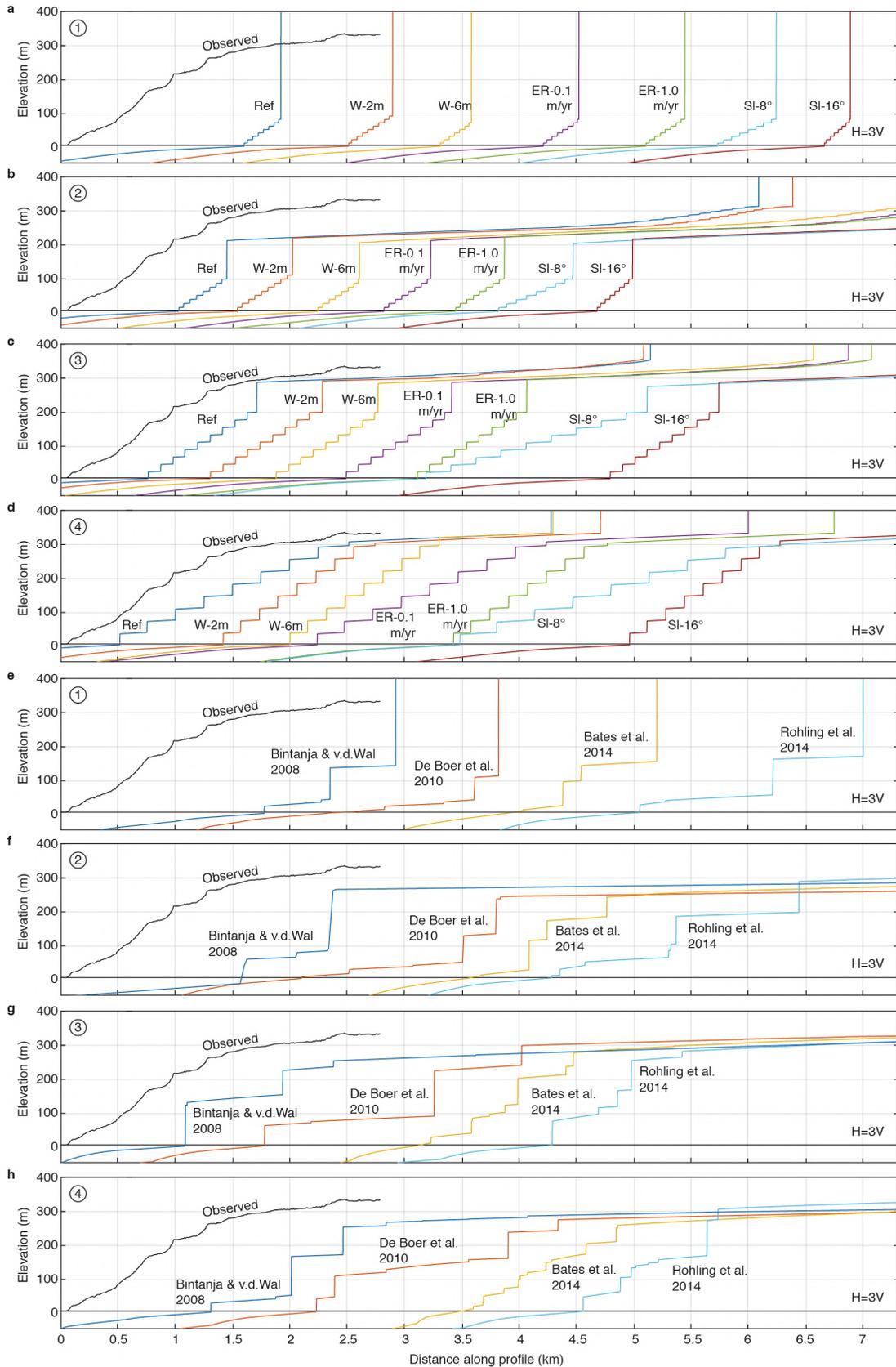


**Figure S1. Topographic data comparison.** Comparison between the standard Aster GDEM (30m horizontal resolution) and the DSM (2m horizontal resolution) we produced with images of the Pleiades satellite; location given by inset in Fig. 2.



**Figure S2. SEM pictures of selected planktic and benthic foraminifers and ostracods collected in the shallow marine deposits in Kythira. A. *Globigerinoides ruber*, sample Na; B.**

*Globorotalia crassaformis*, sample Na; C. *Neogloboquadrina atlantica atlantica*, sample Na; D. *Semicytherura inversa*, sample Na; E. *Lobatula lobatula*, sample Gp; F. *Elphidium crispum*, sample Ta; G. *Aubignyana perlucida*, sample Ta; H. *Mutilus elegantulus*, sample Ta; I. *Callistocythere flavidofusca*, sample Na; J. *Callistocythere parallela*, sample Ta; K., *Semicytherura velata*, sample Ta; L. *Callistocythere intricatoides*, sample Ta; M. *Semicytherura* sp., sample Ta; N. *Aurila punctata*, sample Ta; O. *Aurila cephalonica*, sample Gp; P. *Aurila* sp., sample Ta; Q. *Cimbourila cimbaeformis*, sample Ta; R. *Aurila anguisfoveata*, sample Ta; S. *Aurila hesperiae*, sample Ta. White bars correspond to 0.1 mm.



**Figure S3. Sensitivity tests for LEM modeling.** (a-d) Testing the influence of wave depth, erosion rate and initial slope within reasonable ranges on the 4 tested uplift rate scenarios of Fig. 9. (e-h) Using the Bintanja and van der Wal (2008), De Boer et al. (2010), Bates et al. (2014) and Rohling et al. (2014) sea-level curves (as in De Gelder et al., 2020) on the 4 tested uplift rate scenarios, with the same reference parameter values as in Fig. 9 (also see Supplementary Table 2). As the Bintanja and van der Wal (2008) only spans 3 Ma, values for the 4-3 Ma period have been copied from the De Boer et al. (2010) sea-level curve, which was derived with a similar methodology.

Figure S4 is 15 pages and can be found through:  
<https://doi.org/10.6084/m9.figshare.18986210.v1>

**Figure S4. Swath profiles used to determine shoreline angles.** Map gives locations and numbering of swath profiles on a slope-map of the Pleiades DSM, and the average determined shoreline angles (dots). Color-code for the marine terraces is the same as in Fig. 8. Profiles show maximum and minimum topography, selected points on the terrace and paleo-cliff, and estimated shoreline angle elevation for two end-member scenarios: a most landward and a most seaward position of the hypothetical paleo sea-cliff.

<p><b>Sample BC (2.87 – 2.59 Ma)</b></p>	<p><b>GP Ostracods</b>  <i>Aurila anguisfoveata</i> (Uliczny, 1969)  <i>Bairdia</i> sp.  <i>Callistocythere flavidofusca</i> (Ruggieri, 1950)  <i>Carinocythereis whitei</i> (Baird, 1850)  <i>Caudites calceolatus</i> (Costa, 1853)  <i>Celtia quadridentata cephalonica</i> (Uliczny, 1969)  <i>Cystacythereis</i> sp.  <i>Cytheretta subradiosa</i> (Roemer, 1838)  <i>Eucytherura patercoli</i> (Mistretta, 1967)  <i>Hemicytherura gracilicosta</i> (Ruggieri, 1953)  <i>Loxococoncha ovulata</i> (Costa, 1863)  <i>Pontocythere turbida</i> (Müller, 1894)  <i>Ruggieria</i> sp.  <i>Sagmatocythere versicolor</i> (Müller, 1894)  <i>Semicytherura velata</i> (Ciampo, 1985)  <i>Tenedocythere exornata</i> (Terquem, 1878)    <i>Urocythereis exedata</i> (Uliczny, 1969)  <i>Urocythereis</i> sp.</p>
<p><b>BC Benthic foraminifera</b>  <i>Asterigerinata planorbis</i> (d'Orbigny, 1826)  <i>Astrononion stelligerum</i> (d'Orbigny, 1839)  <i>Cibicides refulgens</i> (de Montfort, 1808)  <i>Cibicidoides pseudungerianus</i> (Cushman, 1922)  <i>Elphidium complanatum</i> (d'Orbigny, 1839)  <i>Elphidium crispum</i> (Linnaeus, 1758)  <i>Elphidium macellum</i> (Fichtel &amp; Moll, 1798)  <i>Lobatula lobatula</i> (Walker &amp; Jacob, 1798)  <i>Melonis soldanii</i> (d'Orbigny, 1846)  <i>Uvigerina pygmaea</i> (d'Orbigny, 1826)</p>	<p><b>BC Ostracods</b>  <i>Aurila anguisfoveata</i> (Uliczny, 1969)  <i>Aurila cephalonica</i> (Mostafawi &amp; Matze-Kasarsz, 2006)  <i>Aurila convexa</i> (Baird, 1850)  <i>Bairdia</i> sp.  <i>Buntonia conularis</i> (Terquem, 1878)  <i>Callistocythere intricatoides</i> (Ruggieri, 1953)  <i>Callistocythere parallela</i> (Aruta, 1986)  <i>Carinocythereis whitei</i> (Baird, 1850)  <i>Costa edwardsii</i> (Roemer, 1838)  <i>Cystacythereis rubra</i> (Müller, 1894)  <i>Cytherella</i> sp.  <i>Cytheretta subradiosa</i> (Roemer, 1838)    <i>Cytheropteron</i> sp.  <i>Echinocythereis pustulata</i> (Namias, 1900)  <i>Eucytherura patercoli</i> (Mistretta, 1967)  <i>Graptocythere intricata</i> (Terquem, 1878)  <i>Hemicytherura gracilicosta</i> (Ruggieri, 1953)  <i>Loxococoncha ovulata</i> (Costa, 1863)    <i>Mutilus elegantulus</i> (Ruggieri &amp; Sylvester-Bradley, 1875)  <i>Paracytheridea</i> sp.  <i>Ruggieria</i> sp.  <i>Semicytherura inversa</i> (Seguenza, 1880)  <i>Semicytherura velata</i> (Ciampo, 1985)  <i>Semicytherura</i> sp. 1  <i>Semicytherura</i> sp. 2  <i>Semicytherura</i> sp. 3  <i>Urocythereis exedata</i> (Uliczny, 1969)  <i>Xestoleberis communis</i> (Müller, 1891)</p>
<p><b>Sample GP (3.81 – 2.59 Ma)</b></p>	<p><b>Sample NA (2.87 – 2.42 Ma)</b></p>
<p><b>GP Benthic foraminifera</b>  <i>Asterigerinata planorbis</i> (d'Orbigny, 1826)  <i>Cibicides refulgens</i> (de Montfort, 1808)</p>	<p><b>NA Planktonic foraminifera</b>  <i>Globigerina bulloides</i> (d'Orbigny, 1826)  <i>Globigerina falconensis</i> (Blow, 1959)  <i>Globigerinella pseudobesa</i> (Salvatorini, 1956)  <i>Globigerinita parkerae</i> (Loeblich &amp; Tappan, 1957)  <i>Globigerinoides elongatus</i> (d'Orbigny, 1839)  <i>Gobigerinoides extremus</i> (Bolli, 1957)  <i>Globigerinoides obliquus</i> (Bolli, 1957)  <i>Globigerinoides ruber</i> (d'Orbigny, 1839)  <i>Globigerinoides trilobus</i> (Reuss, 1850)  <i>Globorotalia crassaformis</i> (Galloway &amp; Wissler, 1927)  <i>Neogloboquadrina atlantica atlantica</i> (Berggren, 1972)  <i>Neogloboquadrina</i> sp.</p>
<p><b>GP Benthic foraminifera</b>  <i>Asterigerinata planorbis</i> (d'Orbigny, 1826)  <i>Cibicides refulgens</i> (de Montfort, 1808)</p>	<p><b>NA Benthic foraminifera</b>  <i>Amphycorina scalaris</i> (Batsch, 1791)  <i>Asterigerinata planorbis</i> (d'Orbigny, 1826)  <i>Brizalina catanensis</i> (Seguenza, 1862)  <i>Brizalina dilatata</i> (Reuss, 1850)  <i>Brizalina</i> sp.  <i>Cancris oblongus</i> (Williamson, 1858)  <i>Cibicidoides pseudungerianus</i> (Cushman, 1922)  <i>Elphidium aculeatum</i> (d'Orbigny, 1846)  <i>Elphidium advenum</i> (Cushman, 1922)  <i>Elphidium macellum</i> (Fichtel &amp; Moll, 1798)  <i>Lenticulina orbicularis</i> (d'Orbigny, 1826)  <i>Lobatula lobatula</i> (Walker &amp; Jacob, 1798)</p>

<i>Elphidium crispum</i> (Linnaeus, 1758) <i>Lobatula lobatula</i> (Walker & Jacob, 1798)	<i>Melonis soldanii</i> (d'Orbigny, 1846) <i>Planulina ariminensis</i> (d'Orbigny, 1826) <i>Pleurostomella alternans</i> (Schwager, 1866) <i>Sphaeroidina bulloides</i> (d'Orbigny, 1826)
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<i>Uvigerina peregrina</i> (Cushman, 1923) <i>Uvigerina proboscidea</i> (Schwager, 1866)	<i>Semicytherura inversa</i> (Seguenza, 1880) <i>Semicytherura velata</i> (Ciampo, 1985) <i>Tenedocythere exornata</i> (Terquem, 1878) <i>Urocythereis exedata</i> (Uliczny, 1969) <i>Xestoleberis communis</i> Müller, 1891
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<b>NA Ostracods</b> <i>Aurila hesperiae</i> (Ruggieri, 1975) <i>Aurila punctata</i> (von Munster, 1830) <i>Bairdia</i> sp. <i>Callistocythere flavidofusca</i> (Ruggieri, 1950) <i>Callistocythere parallela</i> (Aruta, 1986) <i>Carinocythereis whitei</i> (Baird, 1850)  <i>Costa</i> sp. <i>Cystacythereis</i> sp. <i>Cytherella</i> sp. <i>Echinocythereis pustulata</i> (Namias, 1900) <i>Eucytherura patercoli</i> (Mistretta, 1967) <i>Graptocythere intricata</i> (Terquem, 1878) <i>Hemicytherura gracilicosta</i> (Ruggieri, 1953) <i>Loxoconcha ovulata</i> (Costa, 1863) <i>Mutilus elegantulus</i> (Ruggieri & Sylvester-Bradley, 1875) <i>Parakrithe</i> sp. <i>Puloniella hellenica</i> (Mostafawi, 1989) <i>Pontocythere turbida</i> (Müller, 1894) <i>Ruggieria</i> sp. <i>Semicytherura inversa</i> (Seguenza, 1880) <i>Semicytherura velata</i> (Ciampo, 1985) <i>Semicytherura</i> sp. 1 <i>Tenedocythere</i> sp. <i>Urocythereis</i> sp. <i>Xestoleberis communis</i> (Müller, 1891)	<b>Sample TA (2.87 – 2.59 Ma)</b>
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	<b>TA Benthic foraminifera</b> <i>Aubignyna perlucida</i> (Heron-Allen & Earland, 1913) <i>Brizalina arta</i> (MacFadyen, 1931) <i>Cibicidoides pseudungerianus</i> (Cushman, 1922) <i>Elphidium crispum</i> (Linnaeus, 1758) <i>Lobatula lobatula</i> (Walker & Jacob, 1798) <i>Melonis soldanii</i> (d'Orbigny, 1846) <i>Neoconorbina terquemi</i> (Rzehak, 1888) <i>Pleurostomella alternans</i> <i>Uvigerina peregrina</i> Cushman, 1923
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	<b>TA Ostracods</b> <i>Aurila anguisfoveata</i> Uliczny, 1969 <i>Aurila hesperiae</i> Ruggieri, 1975 <i>Aurila punctata</i> (von Munster, 1830) <i>Aurila</i> sp. 1 <i>Bairdia</i> sp. <i>Buntonia conularis</i> (Terquem, 1878) <i>Callistocythere intricatoides</i> (Ruggieri, 1953) <i>Callistocythere parallela</i> Aruta, 1986 <i>Callistocythere</i> sp. <i>Carinocythereis whitei</i> (Baird, 1850) <i>Cimbourila cimbaeformis</i> (Seguenza, 1883) <i>Cymbaurila venetiensis</i> (Uliczny, 1969) <i>Costa punctatissima</i> Ruggieri, 1962 <i>Cystacythereis rubra</i> (Müller, 1894)  <i>Cytherella</i> sp. <i>Echinocythereis pustulata</i> (Namias, 1900) <i>Loxoconcha ovulata</i> (Costa, 1863) <i>Mutilus elegantulus</i> Ruggieri & Sylvester-Bradley, 1875 <i>Puloniella hellenica</i> Mostafawi, 1989 <i>Sagmatocythere</i> cf. <i>S. napoliana</i> (Puri, 1963) <i>Semicytherura inversa</i> (Seguenza, 1880) <i>Semicytherura velata</i> Ciampo, 1985 <i>Semicytherura</i> sp. 1 <i>Semicytherura</i> sp. 2
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<b>Sample QU (3.81 – 2.59 Ma)</b>	
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<b>QU Ostracods</b> <i>Aurila cephalonica</i> (Mostafawi & Matze-Kasarsz, 2006) <i>Aurila convexa</i> (Baird, 1850) <i>Aurila hesperiae</i> (Ruggieri, 1975) <i>Bairdia</i> sp. <i>Callistocythere intricatoides</i> (Ruggieri, 1953)  <i>Callistocythere flavidofusca</i> (Ruggieri, 1950) <i>Carinocythereis whitei</i> (Baird, 1850) <i>Cystacythereis rubra</i> (Müller, 1894) <i>Cytheretta subradiosa</i> (Roemer, 1838) <i>Hemicytherura gracilicosta</i> (Ruggieri, 1953) <i>Loxoconcha ovulata</i> (Costa, 1863)	
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<i>Mutilus elegantulus</i> (Ruggieri & Sylvester-Bradley, 1875)	<i>Tenedocythere exornata</i> (Terquem, 1878)
<i>Paracytheridea</i> sp.	<i>Urocythereis exedata</i> Uliczny, 1969
<i>Pontocythere turbida</i> (Müller, 1894)	<i>Urocythereis</i> sp.
<i>Sagmatocythere versicolor</i> (Müller, 1894)	<i>Xestoleberis communis</i> Müller, 1891

**Table S1. Micropaleontology.** List of all the identified species in the analyzed samples from Kythira

Parameter	Value			
Dx (m)	2			
Dt (yr)	50			
Wave Height (m)	4			
Erosion Rate (m/yr)	0.5			
Slope (degree)	12			
Scenario 1 Age (ka)	2300	2005	2000	0
Scenario 1 UR (mm/yr)	2.0	2.0	0.1	0.1
Scenario 2 Age (ka)	4000	2605	2600	0
Scenario 2 UR (mm/yr)	-0.1	0	0.12	0.12
Scenario 3 Age (ka)	4000	1505	1500	0
Scenario 3 UR (mm/yr)	-0.05	0.05	0.22	0.22
Scenario 4 Age (ka)	4000	905	900	0
Scenario 4 UR (mm/yr)	0	0	0.37	0.37

**Table S2. Model parameters.** Values used for the landscape evolution modeling analysis presented in Fig. 9 and described in section 5.2. Landscape evolution modeling was done using TerraceM-2 (Jara-Muñoz et al., 2019). Sensitivity tests with wave heights of 2 and 6 m, erosion rates of 0.1 and 1.0 m/yr, and slopes of 6° and 18° are given in Supplementary Figure 3