

State-dependence of Cenozoic thermal extremes

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State-dependence of Cenozoic thermal extremes

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1 **Oxygen isotopes in sediments reflect Earth’s past temperature, revealing a**
2 **cooling over the Cenozoic punctuated by multimillennial thermal extreme events.**
3 **These extremes are captured by the generalized extreme value distribution,**
4 **and the distribution’s shape changes with baseline temperature such that large**
5 **thermal extremes are more likely in warmer climates. Anthropogenic warm-**
6 **ing has the potential to return the baseline climate state to one where large**
7 **thermal extremes are more likely.**

8 Analysis of geochemical archives provides insight into Earth’s climate history through prox-
9 ies of paleoclimate conditions (1). Characterizing this history is critical for understanding
10 the evolution of modern Earth and for constraining possible future responses to anthropogenic
11 greenhouse gas emissions (2). Estimates of cumulative emissions so far, remaining fossil fuel
12 reservoirs and the long-term sensitivity of climate to cumulative emissions (3, 4) indicate that
13 humanity has the potential to perturb the climate system enough that the large changes in Earth’s
14 paleorecords (1) are relevant indicators of its potential response on millennial timescales. It is
15 thus particularly important to determine how paleoclimatic variations may depend on baseline
16 climate state, because this is directly linked to the risk of a large long-term Earth system re-
17 sponse to anthropogenic forcing. Variations in Cenozoic climate are studied using deep-sea

18 benthic foraminiferal $\delta^{18}\text{O}$, which relates approximately inversely to global temperature and lin-
19 early to global ice volume such that low $\delta^{18}\text{O}$ corresponds to warm climate states (5). Much of
20 the Cenozoic was a greenhouse climate state with minimal ice volume (1), and so $\delta^{18}\text{O}$ is used
21 as an inverse linear proxy for global temperature (6). Analogously, foraminiferal $\delta^{13}\text{C}$ records
22 past carbon cycle changes through isotopic fractionation during photosynthesis. Tremendous
23 scientific effort has gone into producing, refining, and interpreting these records; it is a marvel
24 that we can infer with some confidence so much about Earth's climate tens of millions of years
25 ago based on the isotopic composition of shells of protist algae that sink to and are preserved in
26 the seabed (7,8). Figure 1 shows the $\delta^{18}\text{O}$ record from (8) leveraging new methods and measure-
27 ments, which we focus on here. Four phenomena are evident: i) a long-term cooling trend, ii)
28 the emergence of periodic Pleistocene glacial-interglacial cycles at 2.6 million years ago (Ma),
29 iii) noisy sub-million-year fluctuations before then, and iv) punctuations of the record by large,
30 rapid, negative $\delta^{18}\text{O}$ excursions corresponding to multimillennial timescale warming events,
31 most notably the Paleocene-Eocene Thermal Maximum (PETM, 56Ma). The long term cooling
32 trend and Pleistocene glacial-interglacial cycles have been the subject of extensive study (1, 7),
33 and the sub-million year noise has recently been shown to be consistent with multiplicative
34 fluctuations (9), potentially due to metabolic temperature-sensitivity of the biosphere (10). The
35 tendency for large negative $\delta^{18}\text{O}$ excursions, perhaps the most concerning from a future climate
36 perspective, has been noted (9), and considerable investigation of individual events such as the
37 PETM shows promise for providing useful constraints on Earth's climate sensitivity (11). How-
38 ever, these thermal extreme events (iv) have not been studied quantitatively and collectively,
39 meaning a general explanation for these extremes and their magnitude is lacking, impairing our
40 ability to use these extremes to make inferences about future climate.

41 The generalized extreme value (GEV) distribution is widely used to study such extremes in
42 other settings (12). Analogously to how the ubiquity of normal and log-normal phenomena in

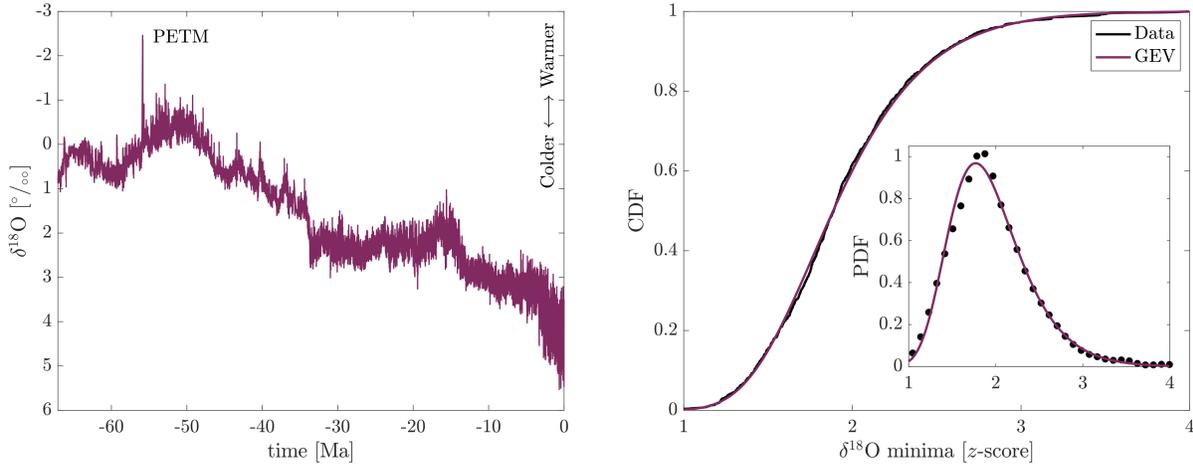


Figure 1: Left: $\delta^{18}\text{O}$ over the Cenozoic (66Ma-present), from (8). Right: Cumulative distribution functions for standardized $\delta^{18}\text{O}$ block minima and generalized extreme value distribution with maximum-likelihood-estimated parameters. Inset: corresponding probability density functions.

43 nature is explained by the central limit theorem (13), the maxima of many natural phenomena
 44 tend to be GEV-distributed, which is explained by the extreme value theorem (Methods). The
 45 GEV distribution has three parameters μ , σ , and ξ , the last of which controls the weight of its
 46 upper tail (12) (Methods). We show that the GEV distribution describes thermal extremes (i.e.
 47 $\delta^{18}\text{O}$ minima) in the Cenozoic excellently, then utilize it to study how the magnitude of these
 48 extremes depends on baseline climate state, allowing us to project the increased likelihood of
 49 large (>3 standard deviations above baseline) thermal extremes as a function of cumulative
 50 emissions.

51 The distribution of thermal extremes, as captured by standard (z -) scores of $\delta^{18}\text{O}$ minima
 52 in blocks of consecutive $\delta^{18}\text{O}$ values, is well-characterized as GEV-distributed (Figure 1). The
 53 Kolmogorov-Smirnov statistic D quantifies the deviation between the theoretical and empirical
 54 distributions; here $D = 0.0213$, well below the threshold $D_{5\%} = 0.0389$ for significance at the
 55 5% level for this sample size (a smaller D -value indicates a better correspondence between the
 56 null hypothesis of GEV distribution, and a D -value below $D_{5\%}$ indicates a failure to reject the

57 GEV distribution at the 5% significance level; Methods). The GEV distribution also applies
58 for $\delta^{18}\text{O}$ maxima (i.e. thermal minima, $D = 0.0142$), $\delta^{13}\text{C}$ maxima ($D = 0.0145$), and $\delta^{13}\text{C}$
59 minima ($D = 0.0203$). This result is also robust to choice of block size (Methods). This
60 excellent agreement suggests we can utilize the GEV distribution to characterize the rarity of
61 individual events in terms of return levels and return periods, but more importantly motivates
62 the use of the GEV to investigate the possible dependency of extremes on baseline climate state.

63 Through this lens of the GEV distribution we investigate whether the magnitude of ther-
64 mal extremes changes with baseline climate state. We fit the GEV distribution to ‘metablocks’
65 of standardized $\delta^{18}\text{O}$ minima grouped according to their associated mean $\delta^{18}\text{O}$ values. Figure
66 2A shows that the shape parameter ξ decreases monotonically as baseline $\delta^{18}\text{O}$ increases, from
67 $\xi = +0.01 \pm 0.03$ when $\delta^{18}\text{O} = 0 \pm 0.5\text{‰}$, to $\xi = -0.32 \pm 0.08$ when $\delta^{18}\text{O} = 4 \pm 0.5\text{‰}$. The im-
68 plication of this ξ -change is shown in Figure 2B, which plots the GEV distribution with best-fit
69 parameters for $\delta^{18}\text{O} = 0 \pm 0.5$ and $\delta^{18}\text{O} = 4 \pm 0.5$. The relative likelihood of an $\delta^{18}\text{O}$ minimum
70 $> z$ standard deviations below the mean for a given z -score is captured by the ratio of these dis-
71 tributions’ complementary cumulative distribution functions (CCDFs). When $\delta^{18}\text{O} \sim 4$ as over
72 much of the past $\sim 3.5\text{Ma}$, $\delta^{18}\text{O}$ minima with z -scores > 3 are virtually impossible/nonexistent,
73 whereas when $\delta^{18}\text{O} \sim 0$ as at the boundary between the Paleocene and the Eocene, such large
74 excursions still had some probability of occurring. We found no other significant or system-
75 atic changes in any other parameters (μ , σ , ξ) of extremes’ (maxima/minima of $\delta^{18}\text{O}$ or $\delta^{13}\text{C}$)
76 distributions as a function of baseline climate or carbon cycle state (mean $\delta^{18}\text{O}$ or $\delta^{13}\text{C}$), indi-
77 cating this phenomenon is restricted to the potential for large thermal maxima depending on the
78 baseline climate state.

79 As the state-dependency seen in Figure 2A is restricted to thermal maxima and does not
80 materialize in the $\delta^{13}\text{C}$ record, we interpret it to be driven by state-dependency of the physical
81 climate system, rather than the carbon cycle. Thermal extremes have been interpreted as be-

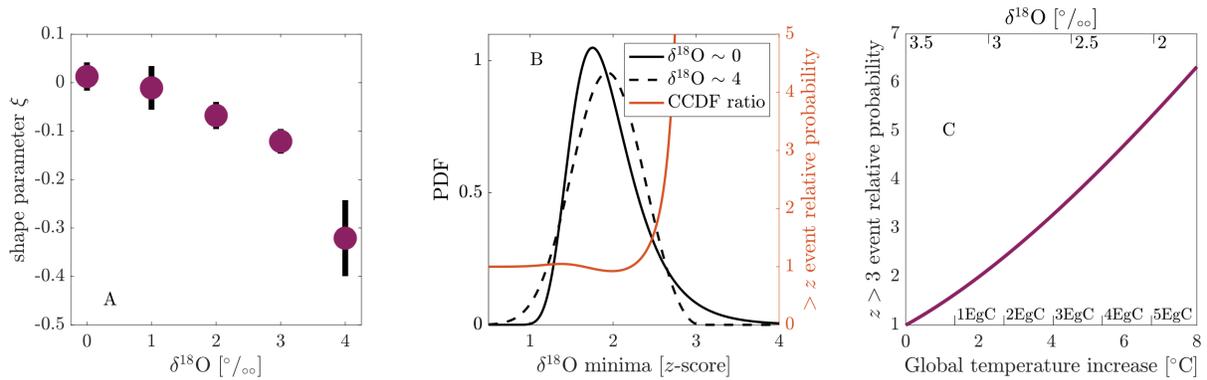


Figure 2: (A) Generalized extreme value (GEV) distribution’s shape parameter ξ as a function of baseline climate state. Compare with analogous figures including $\delta^{13}\text{C}$ and maxima in $\delta^{18}\text{O}$ in Fig-S1. (B) GEV probability density function with the parameters from mean $\delta^{18}\text{O} = 0 \pm 0.5$ (solid black line) and 4 ± 0.5 (dashed black line) from Figure 2A. Orange line is the ratio of these two distributions’ complementary cumulative density functions, indicating e.g. that $\delta^{18}\text{O}$ extremes $>2\frac{2}{3}$ standard deviations below the mean are $>3x$ more likely when $\delta^{18}\text{O} \sim 0$ than when $\delta^{18}\text{O} \sim 4$. (C) Relative likelihood of $\delta^{18}\text{O}$ extremes $>3z$ (three standard deviations below the mean) for different mean $\delta^{18}\text{O}$ values (upper x -axis) compared to present (mean $\delta^{18}\text{O} = 3.5 \pm 0.25$). For instance, such extremes are $\sim 5x$ more likely when $\delta^{18}\text{O} = 2 \pm 0.25$.

82 ing caused by the release of isotopically depleted organic carbon into the surface environment,
 83 such as methane hydrates, permafrost, or dissolved organic carbon. Many of these thermal ex-
 84 tremes have been shown to be accompanied by extremes in $\delta^{13}\text{C}$ (9). The lack of ξ -changes in
 85 $\delta^{13}\text{C}$ is consistent with a temperature-dependent climate feedback; when the climate is warmer,
 86 the same input of carbon produces a larger temperature change (11). Temperature-dependent
 87 climate feedbacks occur in most Earth System Models, most notably due to the water vapor
 88 feedback (14) though also possibly due to e.g. ice-albedo or cloud feedbacks. While we cannot
 89 exclude the possibility that this dependency is due to carbon cycle perturbations that are bal-
 90 anced in their effect on organic and inorganic carbon and therefore not observed in the $\delta^{13}\text{C}$
 91 record, or an external aspect of the Earth system that co-varies with the baseline climate state
 92 such as silicate weathering, these are less parsimonious explanations given the lack of any re-
 93 lationship involving $\delta^{13}\text{C}$ extremes and baseline $\delta^{18}\text{O}$ or vice versa. Additionally, while by any

94 analysis the PETM is an outlier in the $\delta^{18}\text{O}$ and $\delta^{13}\text{C}$ record, our results help contextualize it
95 statistically; such a large outlier was far more probable during such a warm climate state, due
96 to the far heavier tail of the thermal extreme distribution.

97 We can utilize the trend in Figure 2A to estimate Earth's increased susceptibility to large
98 ($>3\sigma$) multimillennial thermal extremes resulting from potential human emissions. As a 0.22‰
99 change is associated with a $\sim 1^\circ\text{C}$ temperature change (15), cumulative carbon emissions so far
100 plus remaining fossil fuel carbon resources are on the order of 5 EgC (= 5 TtC = 5000 PgC =
101 5000 GtC) (3), and 1 EgC cumulative emissions is associated with $\sim 1.35^\circ\text{C}$ long-term warm-
102 ing (4), we focus on the $\delta^{18}\text{O}$ range $3.5(\pm 0.25)\text{‰}$ - $2(\pm 0.25)\text{‰}$, and estimate the ξ change over the
103 equivalent ranges 3.5‰ - 1.75‰ $\delta^{18}\text{O}$, $0\text{--}8^\circ\text{C}$ temperature anomaly, and $0\text{--}6$ EgC emissions. The
104 probability of large multimillennial thermal extremes increases with background warming, dou-
105 bling at approximately 2°C warming, quadrupling at approximately 5°C warming, and sextu-
106 pling at approximately 7.5°C warming. (These relationships are approximate; this extrapolation
107 should be taken illustratively/qualitatively.)

108 Altogether our results suggest that thermal extremes over the Cenozoic are more likely to
109 be large when the baseline climate state is warmer. As similar behavior is not seen in carbon
110 cycle extremes, this dependency is most plausibly due to the temperature-sensitivity of physical
111 climate feedbacks. The probability of large multimillennial thermal extremes (superimposed on
112 anthropogenic warming) may considerably increase if a substantial portion of remaining fossil
113 fuel reserves are combusted.

114 (Word count including abstract, references and figure legends: 1446)

115

116 **References and Notes**

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164 github.com/bbcael and given a DOI via Zenodo should this manuscript be accepted for publica-
165 tion. Reprints and permissions information will be available at www.nature.com/reprints should
166 this manuscript be accepted for publication. Correspondence and requests for materials should
167 be addressed to B. B. Cael (cael@noc.ac.uk).

168 **Online Methods**

169 $\delta^{18}\text{O}$ records were taken from (8) (Figure 1), along with associated $\delta^{13}\text{C}$ records; these variables
170 and their relationship to temperature and other aspects of the Earth system are described exten-
171 sively elsewhere. For blocks of consecutive values the mean, standard deviation, and minima
172 were calculated to determine the standard deviations below the mean (z -score) of the minimum
173 $\delta^{18}\text{O}$ value for that block. The distribution of minima's z -scores is then fit by a generalized
174 extreme value (GEV) distribution via maximum likelihood estimation (Matlab's `mle` function).
175 The extreme value theorem states that the GEV distribution is the only possible limit distribu-
176 tion of properly normalized maxima of a sequence of independent and identically distributed
177 (i.i.d.) random variables. Natural phenomena are rarely if ever truly i.i.d., but the GEV dis-
178 tribution holds and is applied broadly nonetheless (12), analogous to the central limit theorem
179 holding quite accurately for only a handful of summed or multiplied random variables (13). The
180 GEV distribution has the form:

$$f(x; \mu, \sigma, \xi) = \frac{1}{\sigma} t(x)^{\xi+1} e^{-t(x)}$$

181 where $f(\cdot)$ is the probability density function and

$$t(x) = \begin{cases} (1 + \xi(\frac{x-\mu}{\sigma}))^{-1/\xi} & \text{if } \xi \neq 0 \\ e^{-(x-\mu)/\sigma} & \text{if } \xi = 0 \end{cases}$$

182 so μ , and σ are the location and scale parameters and ξ is the parameter that controls the
 183 shape of the distribution. Whether the empirical distribution of maxima deviates significantly
 184 is then determined by calculating the Kolmogorov-Smirnov statistic D , which is the maxi-
 185 mum difference between the hypothesized and empirical cumulative distribution functions, and
 186 comparing it to a critical value at the 5% significance level, $D_{5\%}$ (16); the difference is not
 187 significant if $D < D_{5\%}$. Figure 1 uses the minimum block size of 20 from (17), for which
 188 $D = 0.0213 < D_{5\%} = 0.0389$ and the median z -score is 1.91. Here we focus on the minimum
 189 block size because maximizing the number of blocks is useful to assess changes in the distri-
 190 bution's shape as a function of baseline climate state, as this requires grouping sets of blocks
 191 into 'metablocks.' In general, the larger the block size, the larger the minima's z -scores will be,
 192 and also the larger $D_{5\%}$ will be due to a smaller sample size of maxima. For instance, using
 193 a block size of 67 (the largest prime factor of the length of the $\delta^{18}\text{O}$ record, 24321) yielded a
 194 $D = 0.0279 < 0.0708 = D_{5\%}$ and a median z -score of 2.46, while a block size of 33 (another
 195 factor of 24321) yielded $D = 0.0202 < 0.0498 = D_{5\%}$ and a median z -score of 2.17. D -values
 196 for $\delta^{18}\text{O}$ maxima and $\delta^{13}\text{C}$ maxima and minima reported in the main text are for the same block
 197 size of 20, and are also significant for larger block sizes.

198 Figure 2A was generated by repeating this process on metablocks of block minima, where
 199 blocks were grouped by their mean $\delta^{18}\text{O}$ values into the bins $(0,1,2,3,4)\pm 0.5\%$. Uncertain-
 200 ties (shown using the robust metric of median absolute deviation) were estimated by bootstrap
 201 resampling the distribution of maxima and re-fitting the GEV distribution. We use 10,000 boot-

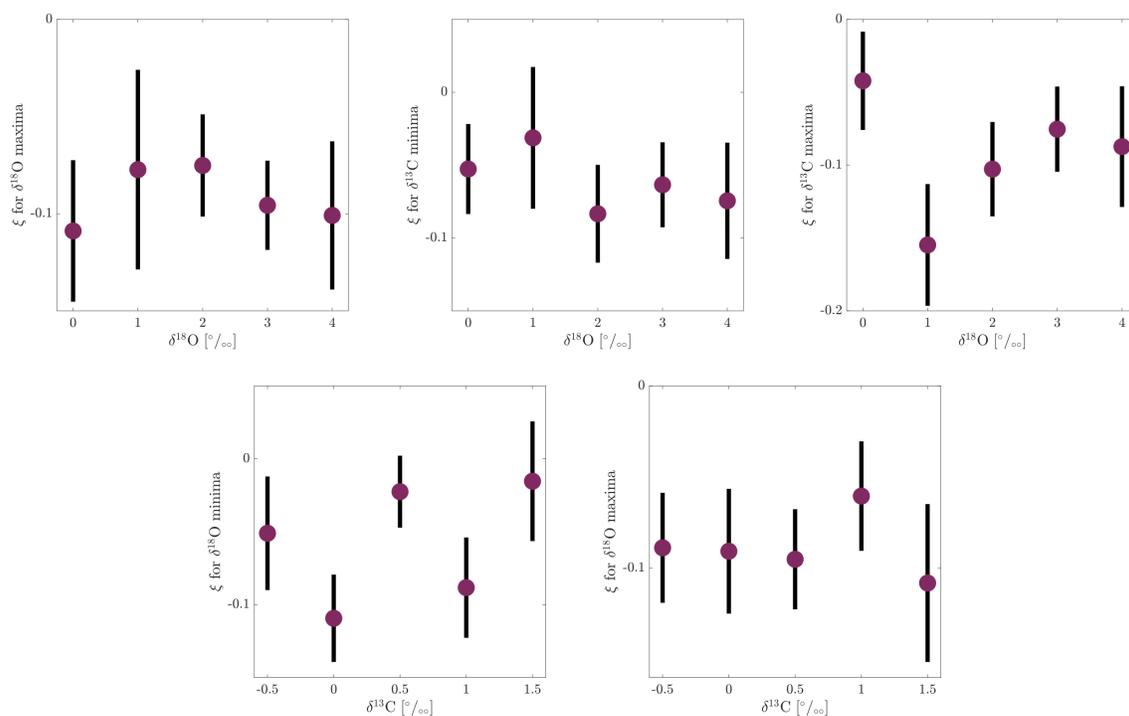
202 strap iterations, which we find to be more than sufficient as ten 1,000-member subsets were
203 negligibly different. The other GEV distribution parameters (location μ and scale σ) vary negli-
204 gibly, neither systematically nor significantly ($p \geq 0.33$ for the block and bin sizes in Figure 2A;
205 this also holds for the bin sizes in Figure 2C), with baseline climate state (i.e. across metablocks).
206 The decreasing trend of ξ with mean $\delta^{18}\text{O}$ holds for larger block sizes (e.g. 33 from above) or
207 narrower bin widths (e.g. ± 0.25 from Figure 2C. In Figure 2B, the complementary cumula-
208 tive distribution function of a probability distribution is one minus its cumulative distribution
209 function. For Figure 2C, we repeat the procedure to estimate the $\delta^{18}\text{O}$ -dependence of ξ (with
210 uncertainties) using the bins $(2,2.5,3,3.5)\pm 0.25$. We then perform a weighted regression of ξ
211 vs. mean $\delta^{18}\text{O}$ to estimate $\xi(\delta^{18}\text{O})$ over this range, yielding an estimate of the GEV distribution
212 for any given $\delta^{18}\text{O}$ value between 1.75-3.5‰. (Note again that other GEV parameters do not
213 change systematically or significantly over this range, or over the entire $\delta^{18}\text{O}$ range.) This is
214 then used to calculate the probability density >3 z -scores, which is shown in Figure 2C relative
215 to the probability density >3 z -scores estimated for present-day $\delta^{18}\text{O} = 3.5$ ‰. We underscore
216 that this subfigure, which includes assumed proportionalities between $\delta^{18}\text{O}$, global temperature
217 change, and cumulative emissions, should be interpreted as illustrative and qualitative.

218 We repeated these calculations for block maxima of $\delta^{18}\text{O}$ and for block maxima and minima
219 of $\delta^{13}\text{C}$. All of these were well-characterized by GEV distributions ($D < D_{5\%}$ in each case),
220 but we found no evidence for any state-dependence other than that reported in the main text.
221 In other words only the shape parameter ξ for $\delta^{18}\text{O}$ minima was dependent on mean $\delta^{18}\text{O}$, and
222 no other GEV distribution parameter of any other maxima or minima was dependent on mean
223 $\delta^{18}\text{O}$ or $\delta^{13}\text{C}$.

224 The glacial-interglacial cycles of the Quaternary period (2.6Ma–present) are recognized
225 not to follow the same sort of fluctuation characteristics as the rest of the Cenozoic, which
226 must be accounted for in any analysis of extremes. Figures 1 and 2 exclude the last 2Ma;

227 neither increasing this to excluding the entire Quaternary period (2.6Ma) nor decreasing this to
228 excluding only after the mid-Pleistocene transition (1.25Ma) affects the results or conclusions.
229 Additionally, these are robust to including the Quaternary period and filtering out the glacial-
230 interglacial cycles via robust locally estimated scatterplot smoothing (R-LOESS) with a window
231 size of 10. Finally we note that our interpretation of $\delta^{18}\text{O}$ minima as thermal maxima is robust to
232 effects of ice volume on $\delta^{18}\text{O}$ because ice sheets primarily act to change the slope and intercept
233 of the linear temperature- $\delta^{18}\text{O}$ relationship $T \approx \alpha - \beta\delta^{18}\text{O}$, with $\alpha, \beta > 0$ approximately
234 constant over the timescales of the extremes considered here. Finally we note that excluding
235 the PETM did not affect our results (as would be expected, as this is only one thermal maximum,
236 and we are analyzing distributions of many thermal maxima) and therefore our inferences about
237 PETM likelihood are not confounded by including it in our analyses.

238 **Supplemental Figure**



239

240 **Fig-S1.** As Figure 2A but for ξ of maxima in $\delta^{18}\text{O}$ vs. baseline $\delta^{18}\text{O}$, ξ of maxima and minima
 241 in $\delta^{13}\text{C}$ vs. baseline $\delta^{18}\text{O}$, and ξ of maxima and minima in $\delta^{18}\text{O}$ vs. baseline $\delta^{13}\text{C}$.