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29 Chapter – **Feedbacks between internal and external Earth dynamics**

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32

33 **Abstract**

34 Countless continuously interacting processes determine the functioning and evolution of the
35 Earth. Even geodynamic and climate changes, which have been classically studied
36 independently because they pertain to different Earth ‘spheres’, are linked by mutual cause-
37 effect relationships that recent research has just started to recognize and quantify. Modeling, be
38 it analogue or numerical, is a trump card in this research for it allows rigorous integrations and
39 interpretations of multiple observations that report on processes with different characteristic
40 spatial and temporal scales and occurring at the Earth’s surface or deep within its interior. In
41 this solicited chapter, I let my academic journey thus far illustrate the challenges of the study
42 of the feedbacks between internal and external Earth dynamics and its relevance for the Earth
43 Sciences as well as for facing and mitigating ongoing fast and extreme global changes.

44 **Keywords:** Internal and external Earth dynamics; interactions between geodynamic and
45 climate changes; geological carbon cycle; geological observations and modeling.

46

47 **1. The ground up**

48 My creed for the continental drifting and plate tectonics theories grew firm from 2003 to
49 2008, when I was a lucky undergraduate student of Earth Sciences at the State University of
50 Milan. Knowing and believing, however, blend into some sort of flawed religious
51 understanding when one learns passively. I first realized this in 2007 when I met Laurent Jolivet

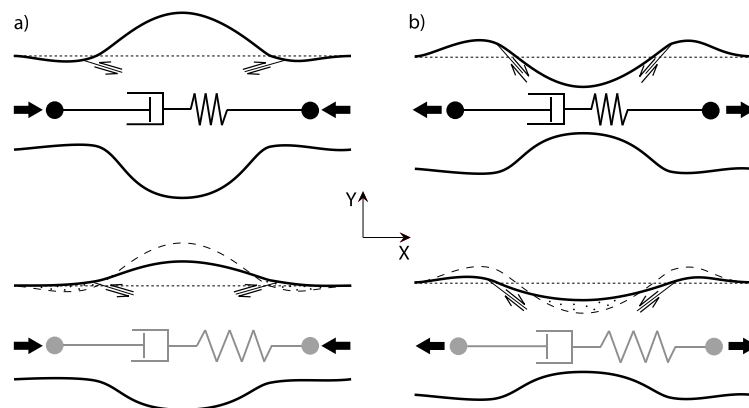
52 and Evgueni Burov in the frame of the Erasums Exchange at the Sorbonne Université and Ecole
53 Normale Supérieure in Paris. As a task for one of the exams, Laurent asked me to read the
54 Burov & Diament, 1995, and Jackson, 2002, papers and present my understanding of the
55 rheology of the continental lithosphere. The different perspectives involved by the ‘jelly
56 sandwich’ and ‘crème brûlée’ rheological models of the continental lithosphere and the
57 arguments in support of and against these models helped me realize about patently unsolved
58 and yet very fundamental issues in the plate tectonics and continental drift theories. During my
59 stay in Paris, Laurent and Evgueni further showed me a different Geology from the classical
60 observational and qualitative Science I was used to. They taught me the basis of a rigorous
61 Geology oriented to the understanding of processes and interactions between processes through
62 integration of quantitative analyses of geological observations.

63 The Deformation Mechanisms, Rheology and Tectonics (DRT) conference was held in
64 Milan in 2007. As a student from the organizing Department, I could help with the organization
65 of the meeting in exchange for free access to my first geological conference. The presentation
66 by Taras Gerya, who showed numerical models of subduction and continental collision,
67 impressed me particularly. Every frame of those physics-based ‘movies’ (literally the opposite
68 of ‘science fiction’) of fundamental geologic processes showed structures extraordinarily
69 similar to those I could observe in the European Alps, with the impressive addition of the stress
70 and thermal fields in which they originate, an information that none of the many structural
71 Alpine cross-sections I came across until then (models just as well) could provide. Once I
72 graduated a few months later, I applied for all open PhD positions that involved numerical
73 modeling of geologic processes. I was eventually offered a position by Sean Willett at ETH-
74 Zurich to work on a project (part of the TOPO-EUROPE program, funded by the European
75 Science Foundation) aimed at constraining the role of glaciation in affecting the Plio-
76 Quaternary topographic and tectonic history of the Alps. I was trained as a ‘classical’ structural

77 geologist and the focus on geomorphology was largely outside my comfort zone. I was also
78 underestimating the considerable (but worthy) effort required to learn the basics of numerical
79 modeling without an appropriate background. I learnt soon after that stepping out of your
80 comfort zone and engaging into new and demanding challenges is exactly what research is
81 about.

82 During the PhD, I was mainly supervised by Frédéric Herman who gave me what I now
83 acknowledge as a fairly balanced amount of constraints and liberty to allow for independent
84 and in depth learning. The learning process was facilitated by Matthew Fox, who came from
85 Oxford to begin his PhD at the same time as me, but with a much better training in any sort of
86 quantitative analyses. It did not take long to realize that surface processes - erosion and sediment
87 deposition but also the formation and melting of continental ice-caps and associated sea level
88 changes - are affected by but also affect the most relevant expressions of plate tectonics and,
89 thus, are critical in determining the evolution of the Earth system. Simple physical
90 considerations make this point easily understandable. The horizontal motion of tectonic plates
91 implies their shortening or stretching, both of which involve vertical deformation of the Earth's
92 surface plates. Thickening or thinning of tectonic plates 'floating' on the underlying mantle
93 imply a departure from the equipotential state and, thus, work against the horizontal motion of
94 plates (Fig. 1). The dismantling of uplifted terrains via erosion and the filling of subsiding
95 basins with sediments contribute to restore the equipotential state, thereby facilitating the
96 horizontal motion of plates and the associated strain. Within this frame, convergent mountain
97 belts have been considered as crustal scale accretionary wedges in which deformation is
98 described by a simple Coulomb behavior since the pioneering works by Davis, Dahlen, and
99 Suppe in the 1980s (e.g., Davis, et al., 1983; Dahlen, et al., 1988). Sandbox numerical as well
100 as analogue models show localized thrusting during convergence that controls the geometric,
101 topographic, and architectural evolution of the wedge. Basal décollements play a major role

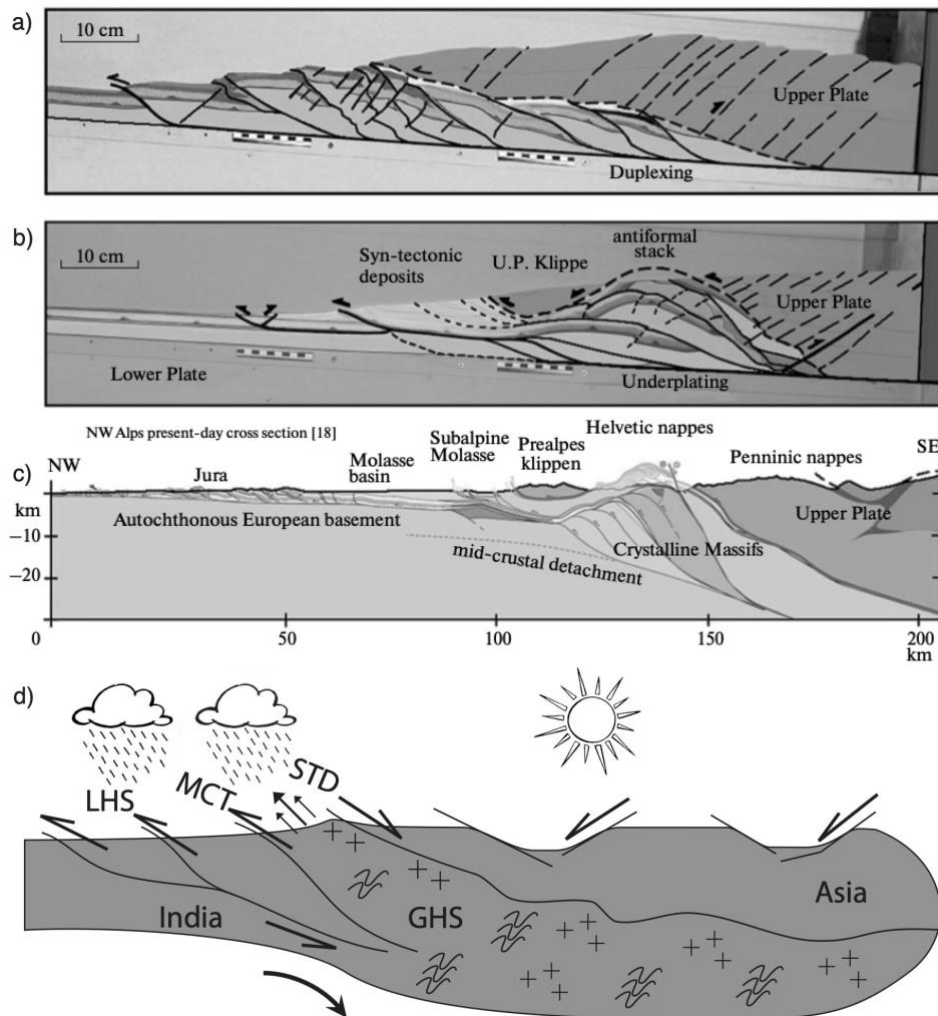
102 allowing for the formation of duplex structures and underplating at different structural levels
 103 within the wedge, whereas frontal accretion characterizes the deformation in the foreland. The
 104 strong localization of brittle deformation in thrust wedges exerts a major influence on
 105 topography and, thus, on erosion. In turn, erosion and sedimentation decrease the topographic
 106 slope thereby favoring a change from overcritical to stable to undercritical mechanical state of
 107 the wedge (e.g., Willett, 1999) and allowing for a feedback loop that plays a particularly
 108 important role for what concerns exhumation of deep crustal and metamorphic rocks.
 109



110
 111 **Figure 1: Schematic representation of the effect of surface processes on the stress state within a deforming continental**
 112 **lithosphere subject to horizontal compression (a) or shortening (b). Relaxed (gray) and stressed (black) visco-elastic**
 113 **systems are represented by an elastic spring and viscous damper. Surface processes relax deforming systems, thereby**
 114 **contributing to the restoration of the equipotential state lost due to horizontal and vertical tectonic strain.**

115
 116 For instance, analogue models allow relating many of the main structures within classical
 117 geological sections across the Swiss Alps to erosion and/or sedimentation during convergence
 118 (e.g., Malavieille & Konstantinovskaya, 2010; Fig. 2a-c). In response to shortening without
 119 surface processes, an analogue basement is commonly subject to initial thrusting and
 120 imbrication upon inherited structural and sedimentary weaknesses. Then, the homogenous part
 121 of the basement underthrusts and a high friction wedge is originated. With erosion and
 122 sedimentation, convergence leads to initial thrusting and frontal accretion in the foreland basin,
 123 followed by formation of an antiformal stack of duplexes in the internal part. Here, protracted
 124 strain localization, erosion and exhumation isolates a frontal synformal klippen of formerly

125 imbricated thrust units and the antiformal structure eventually outcrops as a tectonic window,
126 as observed in the natural case study (e.g., Burkhard & Sommaruga, 1998). The influence of
127 surface processes on orogenic dynamics in the European Alps is also expressed at short
128 timescales since at least ~50% of the geodetically measured present-day vertical displacements
129 are currently ascribed to the deglaciation of the Last Glacial Maximum ice-sheet and Plio-
130 Quaternary erosion of the belt (e.g., Sternai, et al., 2019, and references therein). Another
131 example is provided by exhumation of the high-grade metamorphic Greater Himalayan
132 Sequence along the Himalayan-Tibetan range, interpreted as the result of southward growth of
133 the Tibetan plateau due to ductile flow in the middle to lower crust (e.g., Burchfiel, et al., 1992;
134 Royden, 1996; Grujic, et al., 1996; Wu, et al., 1998). Thermo-mechanical numerical models,
135 which further allow to account for the ductile behavior of rocks during convergence and flexure
136 of the lithosphere due to the topographic growth and erosion or sediment deposition (e.g.,
137 Garcia-Castellanos, 2002), clearly indicate a dynamic link between the ductile flow and
138 exhumation of lower crustal rocks and localized surface erosion at the orogen front where
139 topographic slopes and orographic precipitations are particularly significant (e.g., Beaumont,
140 et al., 2001; Zeitler, et al., 2001; Fig. 2d).



141

142 **Figure 2:** (a) Analogue models of crustal convergence and orogenic wedge evolution without (a) and with (b) erosion
 143 and syn-tectonic sedimentation (Malavieille & Konstantinovskaya, 2010). Modeling results are compared to a classical
 144 cross section of the Swiss Alps (c) from Burkhard & Sommaruga, 1998. d) Crustal tectonic framework of the Himalaya
 145 and Southern Tibet where erosion at the mountain front contributes to localized exhumation of weak lower crustal
 146 material (e.g., Wu, et al., 1998; Burchfiel, et al., 1992; Beaumont, et al., 2001). LHS, Lesser Himalayan sequence; MCT,
 147 Main Central Thrust; STD, South Tibetan Detachment; GHS, Greater Himalayan Sequence.

148

149 Thanks to an enormous amount of research performed in less than one generation, it is
 150 now established that deformation, surface uplift or subsidence and erosion or sediment
 151 deposition (these latter may act in addition to ice-building/melting or sea level changes)
 152 comprise a system with feedbacks that links plate tectonics and continental drifting to the
 153 evolution of the surface topography. Being the main driver of both the motion of tectonic plates
 154 and the redistribution of the surface masses, the gravity acceleration is the fundamental engine
 155 of the feedbacks between internal and external dynamics. Climate, however, has a special role

156 to play because it determines the driving mechanisms of the redistribution of the Earth's surface
157 masses and, thus, the rate at which the lost equipotential state due to the tectonic strain may be
158 re-established (or re-approached to).

159

160 **2. The ground down**

161 In 2012, a few months before defending my PhD I contacted Laurent Jolivet to ask for
162 postdoc opportunities in his research group. He replied quickly and positively since his
163 RHEOLITH ERC Advanced Grant proposal had just been funded. The project involved a study
164 of the rheology of the continental lithosphere based on three main pillars: field observations,
165 experimental measurements, and numerical modeling. Laurent, who transitioned to the
166 University of Orléans while I was a graduate student at ETH-Zurich, offered me a postdoc
167 position to work on numerical modeling of subduction and collisional systems constrained by
168 available and newly produced observational and experimental data. Because numerical
169 modeling provides a means to test hypotheses formulated based on field observations, Laurent
170 brought me in continental Greece and the Cyclades, where large-scale GPS measurements,
171 seismic data and outcrop-scale strain structures allow observing continental chunks (e.g.,
172 Anatolia) being extruded out of a collisional system (e.g., the Bitlis-Zagros domain) into the
173 back-arc domain of an active subduction system (e.g., the Aegean). These observations raised
174 a number of considerations in my head. First, if continents extrude like toothpaste squeezed out
175 the tube, they must be somewhat soft. Second, convergent settings involve more than just
176 shortening, thickening and uplift of the surface topography. The assumption that a collisional
177 orogen is a broad uniformly and steadily uplifting area, which I came across and adopted myself
178 many times during my PhD, seemed a particularly coarse oversimplification ever since. Third,
179 lateral boundaries of convergent settings are particularly important for they determine whether
180 and where lateral escape of continental material can occur via transcurrent strain, thereby

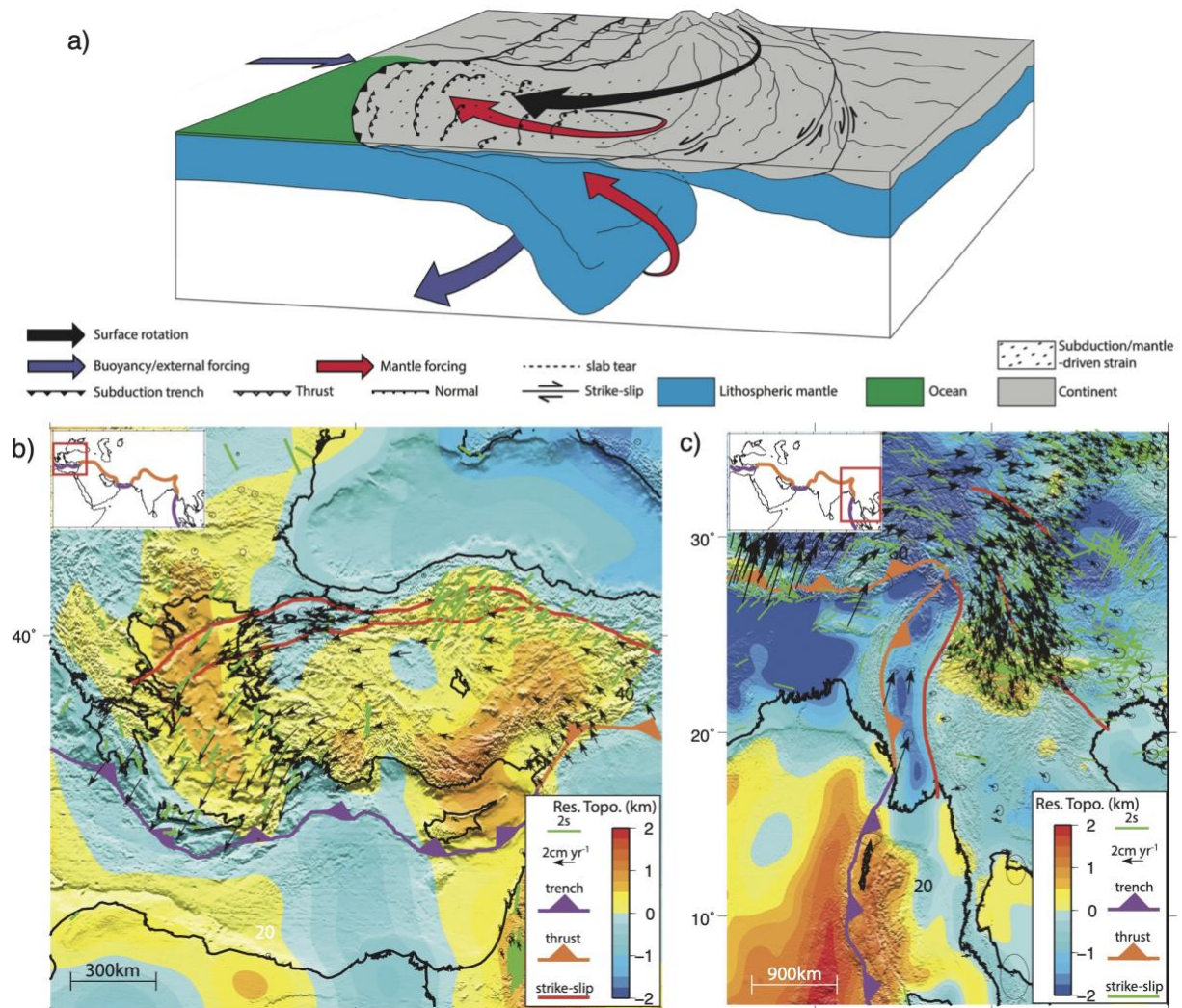
181 allowing the convergence to continue or forcing it to slow and eventually end. Indeed, in
182 reducing the work against gravity that horizontal tectonics produces, extrusion tectonics play a
183 very similar role to that of surface processes (Fig. 3a) and I refer again to the magnificent
184 Tibetan plateau and mountains of Asia to provide an example.

185 The Tibetan-Himalayan range has been long ascribed to lithospheric shortening and
186 thickening along the India-Eurasia margin (e.g., Argand, 1924; Molnar & Tapponnier, 1975).
187 Unlike the abrupt Himalayan front, the gentler but still impressive topography along the eastern
188 margin of Tibet developed in a predominantly trans-tensional tectonic regime (e.g., Leloup, et
189 al., 1995; Wang, et al., 1998; Hall & Morley, 2004). These fundamental observations and the
190 geophysical evidence suggesting the presence of a weak lower crust below Tibet (e.g., Nelson,
191 et al., 1996; Xu, et al., 2007) triggered a debate about the partitioning between clock-wise rigid
192 rotation (e.g., Tapponnier, et al., 1981; Armijo, et al., 1986; Avouac & Tapponnier, 1993;
193 Leloup, et al., 1995; Meade, 2007) or viscous eastward evacuation of the Asian crust and
194 lithosphere, possibly involving crustal channel flow (e.g., England & Houseman, 1986;
195 Royden, 1996; Clark & Royden, 2000; Clark, et al., 2006; Copley & McKenzie, 2007). These
196 proposals put different emphasis on strain localization, vertical gradients of strain due to depth-
197 dependent rheologies, the role of gravitational body forces and tractions at the base of the
198 lithosphere, and the influence of plate boundary dynamics, but all agree that extrusion tectonics
199 permitted protracted and ongoing convergence of India into Eurasia. Similarly, common to all
200 proposed models are a focus on crustal dynamics and, regarding the effects of plate boundary
201 dynamics, the assumption that subduction of oceanic lithosphere along the Sunda and western
202 Pacific margins created the accommodation space for unconstrained continental extrusion. I
203 stress here that these considerations are not unique to the Asian mountains, but apply to many
204 places along the Alpine-Himalayan orogenic belt, which is a continuous alternation between
205 collisional (the Alpine-Dinarid, Bitlis-Zagros and Tibetan-Himalayan mountains) and

206 subduction zones (Carpathian-Pannonian, Ionian-Aegean, Makran and Sunda subduction
207 systems), toward which continental extrusion systematically takes place (Fig. 3b,c).

208 The necessary requirement of an ‘open’ boundary to allow for extrusion, be it by rigid
209 rotation of upper crustal continental blocks or by viscous flow of lower crustal material, and
210 protracted horizontal convergence of plates puzzled me particularly. Widespread extension
211 across the Anatolian extruded continental domains as it transitions into the back-arc regions of
212 the bounding rollback Aegean subduction zone suggests that a ‘suction’ toward the
213 subduction zone facilitates extrusion of continental slivers and, thus, horizontal convergence in
214 the collisional domain. This hypothesis was already proposed for the eastern Tibetan margin
215 (e.g., Burchfiel & Royden, 1985; Jolivet, et al., 1990; Northrup, et al., 1995; Schellart & Lister,
216 2005; Royden, et al., 2008). With the I3ELVIS geodynamic model provided by Taras Gerya,
217 we could validate its physical and geological feasibility building a setup that accounts for a
218 collisional and an ocean-continent subduction domain (Sternai, et al., 2014; Sternai, et al.,
219 2016b; Menant, et al., 2016; Roche, et al., 2018). In these experiments, a major bended strike-
220 slip surface structure on the upper plate, geometrically and kinematically similar to the North
221 Anatolian or Altyn-Tagh fault zones, joints compressive structures in the collisional zone to
222 extensional structures in the back arc domain of the subduction system. The strike-slip structure,
223 however, only develops when the subducting slab is subject to rollback during convergence,
224 whereas compression occurs across the entire upper plate otherwise. The numerical models also
225 provide access to deep levels, where basal lithospheric shearing due to the asthenospheric return
226 flow associated with slab rollback facilitates the upper plate extensional strain in the back-arc
227 of the subduction zone. These results allowed us to conclude that rollback subduction zones
228 at the sides of a collisional domain not only provide ‘free space’ but also actively drive
229 continental extrusion thereby promoting the horizontal convergence between the colliding
230 continental plates, a conclusion that may apply (or may have applied in the geological past) to

231 multiple extruded domains along the Neo-Tethyan margin, Fig. 3a (e.g., Sternai, et al., 2014;
 232 Sternai, et al., 2016b).



233
 234 **Figure 3: (a) Schematic representation of the dynamic interactions between continental collision, oceanic subduction,**
 235 **mantle flow and surface deformation (modified after Sternai et al., 2014). (b) Maps of the residual topography, $z_{res} =$**
 236 **$z_{obs} - z_{iso}$. The surface elevation of a lithospheric column with respect to a reference elevation, H, is determined by**
 237 **$z_{iso} = ((\rho_a - \rho_c)/\rho_a l_c + (\rho_a - \rho_m)/\rho_a l_m) \cdot H$, where l_c is the crustal thickness, ρ_c is the average crustal density, l_m is**
 238 **the lithospheric mantle thickness, ρ_m is the average mantle lithosphere density, and ρ_a is the asthenosphere average**
 239 **density. Maps were generated based on the CRUST1.0 model and assuming a lithospheric thickness of 100 km,**
 240 **uniform crustal and mantle densities and H equal to the average mid-oceanic ridge elevation (e.g., Lachenbruch &**
 241 **Morgan, 1990). More details on the methodology can be found in Faccenna, et al., 2014; Sternai, et al., 2016b; Sternai,**
 242 **et al., 2019. Positive values across extruded continental domains suggest that differential along-strike kinematics at the**
 243 **transition between subduction and collision zone along the Neo-Tethyan margin may produce vigorous asthenospheric**
 244 **flow that contributes to the surface strain and topographic uplift. GPS velocity vectors are from Zhang, et al., 2004;**
 245 **Gan, et al., 2007; Reilinger, et al., 2010. Seismic anisotropies (green markers) are from Wüstefeld, et al., 2009, and**
 246 **Biryol, et al., 2010.**

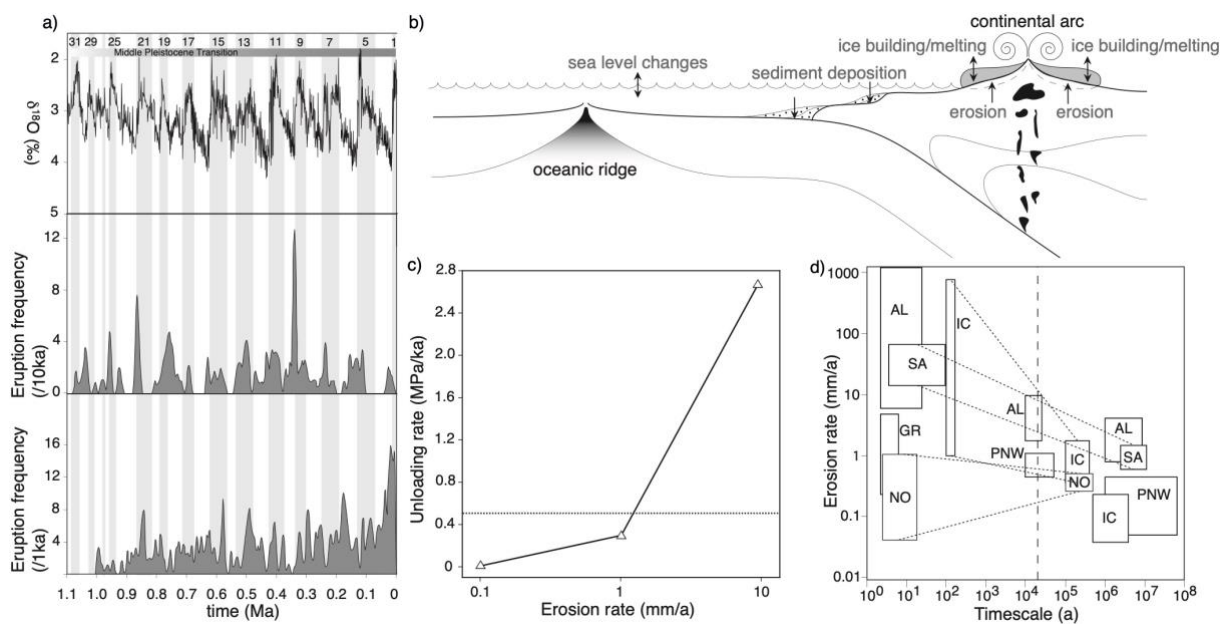
247
 248 In 2013, Jean Philippe Avouac and his students from Caltech joined one of the fieldtrips to
 249 Greece and I had the chance to show the recently obtained numerical results and discuss our
 250 preliminary interpretation. Jean Philippe suggested to extract the vertical component of the

251 asthenospheric return flow due to slab rollback and evaluate under which conditions the
252 resulting stresses provide significant support to the topography of the extruded continental
253 domains. A few months later he offered me a postdoc position to be held at Caltech and the
254 University of Cambridge to study the relationships between basal lithospheric traction due to
255 the asthenospheric flow and the surface evolution with a focus on Southeast Asia. The objective
256 was to assess whether the asthenospheric return flow owing to protracted northward migration
257 of the Indian indenter during rollback (late-Eocene to middle-Miocene) or stable (middle-
258 Miocene to present) subduction along the Sunda and western Pacific margins (e.g., Tapponnier,
259 et al., 1986; van der Hilst & Seno, 1993; Hall & Morley, 2004; Sibuet, et al., 2004; Honza &
260 Tokuyama, 2004; Royden, et al., 2008; Replumaz, et al., 2013) could provide active support to
261 the topography. Working with Jean Philippe helped me realize that, because complex 3D
262 numerical geodynamic models can account for a multitude of processes and interaction between
263 processes, the interpretation of the numerical results in terms of driving forces and resulting
264 effects requires careful analyses. An analytical investigation of the numerical results in terms
265 of Gravity Potential Energy (GPE) and depth-integration of strain and stress variations led to
266 recognition that the asthenospheric flow due to differential along-strike slab kinematics may be
267 vigorous enough to contribute to the surface strain and elevations at collision-subduction
268 transition zones (Sternai, et al., 2016b; Fig. 3b,c). If large-scale mantle convection drives the
269 overall India-Eurasia convergence (e.g., Alvarez, 2010; Becker & Faccenna, 2011), the more
270 local asthenospheric return flow related to relatively shallow dynamics contributes to the
271 topographic evolution, which indicates even tighter, smaller-scale and higher-order
272 relationships between internal and external dynamics than previously recognized.

273

274 **3. Merging concepts toward an integrative understanding of the Earth system**

275 At the beginning of 2015, I have been invited by Sébastien Castelltort to give a seminar at
 276 the Department of Earth Sciences of the University of Geneva. At the end of the seminar, Luca
 277 Caricchi pointed out that, if erosion can change the stress field at depth so to modulate the strain
 278 pattern, it may also affect the magma production because pressure is key to partial rock melting
 279 in many geodynamic contexts. The following discussions with Luca and Sébastien led to
 280 recognition of a poorly explored and very promising link between internal and external
 281 dynamics: the coupling between surface processes and magmatism. I was approaching the end
 282 of my second postdoc and it was the right time to try my hand at writing a research proposal
 283 which, about a year later, the Swiss National Science Foundation funded in the frame of the
 284 ‘Ambizione’ program.



285
 286 **Figure 4:** (a) Global $\delta^{18}O$ curve (Zachos, et al., 2001), tephra frequency at the Izu Bonin Mariana arc (central panel,
 287 using 10 ka binning after (Schindlbeck, et al., 2018)) and the Pacific Ring of Fire (lower panel, using 1 ka binning after
 288 (Kutterolf, et al., 2013)). The Middle Pleistocene Transition leads from dominant ~ 40 ka periodicity of climate
 289 oscillations to dominant ~ 100 ka cycles. Vertical gray bars mark marine isotope stages (MIS) after Railsback, et al.,
 290 2015. Note the cyclicity of the tephra record consistent with glacial-interglacial cycling. (b) Schematic representation of
 291 the relationships between surface processes (gray) and magmatic sources (black). Erosion, sediment deposition, ice
 292 building-melting and sea level changes can affect processes responsible for the production, transfer and eruption of
 293 magma by modulating the stress state at depth. (c) Estimated continental unloading rate owing to constant melting of
 294 1km of ice (dotted line for reference) and erosion (assuming a surface rock density of $2,700 \text{ kg/m}^3$) throughout the last
 295 interglacial (Sternai, et al., 2016a). (d) Boxes represent ranges of erosion rates from glaciated catchments or proximal
 296 basins including errors in estimations (vertical) and resolved timescale (horizontal). The data is from Brandon, et al.,
 297 1988; Hallet, et al., 1996; Sheaf, et al., 2003; Reiners, et al., 2003; Koppes & Hallet, 2006; Hebbeln, et al., 2007;
 298 Geirsdottir, et al., 2007; Berger & Spotila, 2008; Koppes, et al., 2009; Cowton, et al., 2012; Herman, et al., 2013;
 299 Herman & Brandon, 2015. AL, Alaska; SA, Southern Andes; GR, Greenland; NO, Norway; PNW, Pacific Northwest;
 300 IC, Iceland. The vertical dashed line represents the approximate time since the LGM (Lisiecky & Raymo, 2007). At
 301 such timescale, erosion rates between 1-10 mm/a are commonly observed.

302
303 Ice building/melting and associated sea level variations during glacial-interglacial cycles
304 affect the magmatic activity (Fig. 4a) modulating decompression partial melting (e.g., Crowley,
305 et al., 2015; Jull & McKenzie, 1996), generating new fractures that facilitate the magma
306 transport through the lithosphere (e.g., Rubin, 1993; Maccaferri, et al., 2011), and enhancing
307 gas exsolution within volatile saturated magmas, thereby leading to overpressure and increasing
308 the probability and magnitude of eruptive events (e.g., Jellinek & De Paolo, 2003). However,
309 the link between erosional unloading of continent and partial rock melting or the transfer
310 through the lithosphere and eruption of magma was not envisaged, although the density of rocks
311 and sediments is about three times the density of water or ice and deep fjords, glacial over-
312 deepened valleys, and the ubiquitous low-stand fluvial incisions on shelves testify for intense
313 erosion associated with glaciation. Pulses of erosion at regional or local scales are commonly
314 observed in the fluvial and glacial sedimentary records (e.g., Boulton, et al., 1988; Mullins &
315 Hinchey, 1989; Bjornsson, 1996; Singer, et al., 1997; Brown & Kennett, 1998; Koppes &
316 Hallet, 2006; Geirsdottir, et al., 2007) and numerical experiments (besides simple physical
317 considerations) indicate that erosion rate in the order of the mm/a or higher during deglaciation
318 unload continents by a similar or greater amount than the melting of large ice sheets (Sternai,
319 et al., 2016a; Fig. 4b,c). Long-term proxies of global sediment efflux from mountainous regions
320 show just such erosion rate variability and, at secular to millennial timescales, erosional fluxes
321 may be even higher and subject to strong variations due to modifications of the subglacial
322 hydraulic system and water supply (e.g., Koppes & Montgomery, 2009; Herman, et al., 2011).
323 The abrupt and high-magnitude magmatic pulses involved by such erosional changes are likely
324 to force centennial- to millennial-scale variations of atmospheric greenhouse gases seemingly
325 unrelated to ocean dynamics (e.g., Monnin, et al., 2004; Marcott, et al., 2014). In fact, on a
326 global scale, sea level rise following continental ice melting seems able to reduce the magma
327 productivity of mid-oceanic ridges (e.g., Crowley, et al., 2015), in turn potentially buffering the

328 increased subaerial volcanic activity and associated degassing owing to continental unloading
329 by the deglaciation. Such a buffering effect, however, does not apply to continental unloading
330 by erosion because sediment deposition in the ocean is unlikely to occur atop of oceanic ridges,
331 which are, for the vast majority, several hundreds of kilometers away from continental shelves
332 and stand up to a few thousand meters higher than abyssal plains or subduction trenches to
333 where the eroded sediments are transported (Fig. 4b). Therefore, continental erosion may have
334 greater net effects than ice melting on the CO₂ outflux and magma productivity from the solid
335 Earth. Earth system models that do not account for erosion may lead to significant
336 underestimations of the increase in atmospheric CO₂ concentration during interglacials
337 (Sternai, et al., 2016a).

338 The possibility of a positive feedback between factors internal to the climate system such
339 as erosion, subaerial volcano-magmatic CO₂ emissions, climate warming and deglaciation is
340 particularly worth of investigations. In a recent publication (Sternai, et al., 2020), we suggest
341 that such positive feedback may explain the “sawtooth” asymmetry (i.e., faster transitions to
342 warmer conditions than cooling trends) of Plio-Pleistocene glacial cycles (e.g., Lisiecky &
343 Raymo, 2007), which is not found in any orbital or insolation curve (e.g., Broecker & Donk,
344 1970). The main logic is that inhibition of subaerial eruptions during glacial periods forces
345 accumulation of gasses in magmatic reservoirs, which are then released over a few thousands
346 of years during the early interglacials (e.g., Jellinek, et al., 2004). Assuming that the long-term
347 weathering CO₂ sink is at equilibrium with the steady state volcanic CO₂ outgassing, the
348 weathering carbon sink slightly dominates over inhibited volcanic carbon emissions when ice
349 sheets grow, leading to a temporary reduction of atmospheric CO₂, which sustains climate
350 cooling. As soon as the orbital forcing of solar radiation overtakes the threshold to trigger the
351 deglaciation, enhanced volcanic carbon outgassing dominates over the weathering carbon sink,
352 in turn fostering climate warming and bringing the overall atmospheric CO₂ budget back to

353 equilibrium. The concept may be expressed with an example based on the early Pleistocene 40
354 ka and late Pleistocene 100 ka glacial-interglacial cyclicality. If the phase of enhanced outgassing
355 during early interglacials is extinguished in ~10 ka, then the cooling phase is forced to ~30 ka
356 and ~90 ka for early Pleistocene and late Pleistocene cycles respectively. The ~10 ka time
357 window is chosen arbitrarily, but the duration of warming of late Pleistocene climate
358 oscillations and the expected response time of magmatic systems to surface load changes
359 constrain this value (e.g., Jellinek, et al., 2004; Lisiecky & Raymo, 2007). Because both
360 constraints are largely independent on the period of climate oscillations (i.e., 40 or 100 ka), the
361 duration of the phase of enhanced outgassing determines the asymmetry of climate oscillations.
362 That is, the longer the phase of gas accumulation in magmatic reservoirs (glacials), the larger
363 the outgassing rate during the phase of enhanced emissions (early interglacials), and the more
364 pronounced the asymmetry between warming and cooling trends (Sternai, et al., 2020). This is
365 also consistent with increasing asymmetry of climate oscillations throughout the mid-
366 Pleistocene transition from dominant 40 to 100 ka cycles (e.g., Lisiecky & Raymo, 2007). Of
367 course, this simplistic analysis should be investigated further and possibly validated via
368 observations and more thorough modeling.

369

370 **4. A long way to go**

371 Since at least the '80, the recognition that major long term climate changes are related
372 to geodynamic events boosted research on the feedbacks between internal and external
373 dynamics and led to cross-disciplinary research efforts that built up most of our current
374 knowledge on the topic. Increasing awareness about the current climate crisis since the early
375 2000s helped maintaining a considerable interest on the subject. Indeed, knowledge about the
376 natural variability of climate and the interactions between geological processes behind them is
377 a fundamental prerequisite to quantify the characteristic magnitudes and rates of anthropogenic

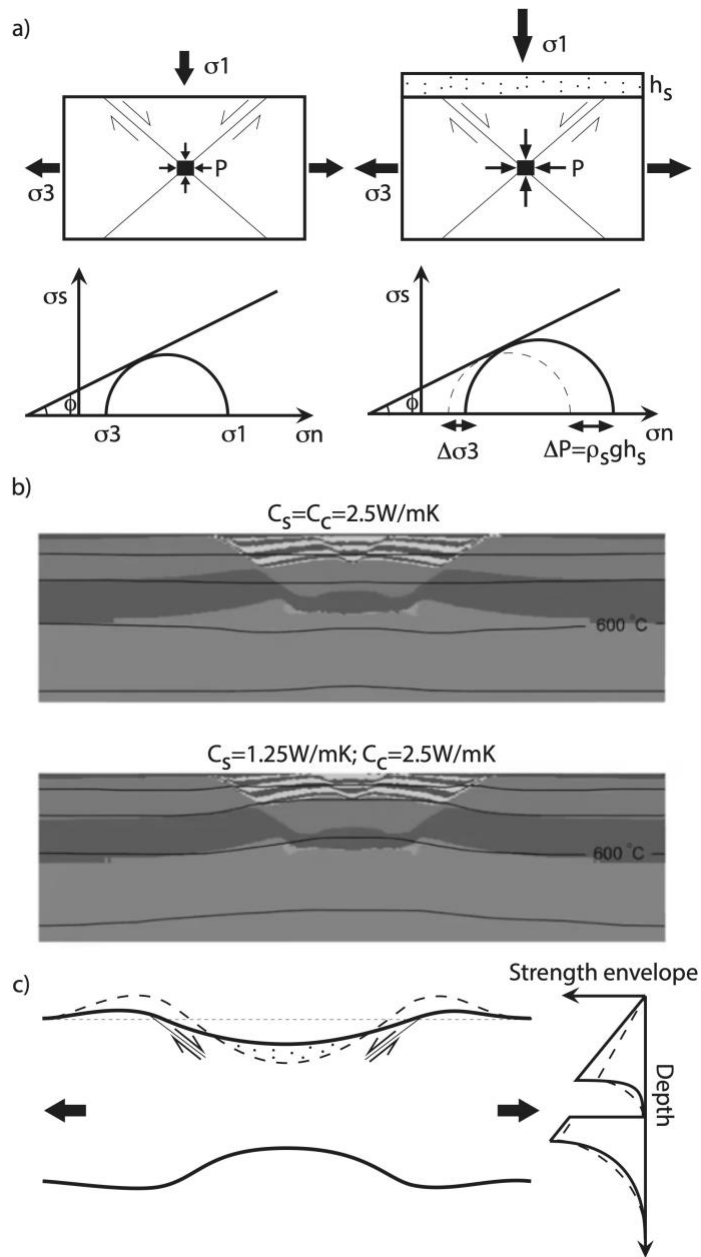
378 climate and environmental changes. Although much has been done, there remain many poorly
379 explored and likely fertile areas in this research field. I outline hereafter a few examples.

380

381 **4.1. Feedbacks between internal and external dynamics in extensional settings**

382 Much of what we know about the feedbacks between internal and external Earth
383 processes and resulting climate-tectonics interactions comes from the study of mountain
384 building in convergent tectonic settings. However, prominent topographic ridges and basins
385 generated by dominant extensional tectonics make divergent settings valuable contexts too
386 (e.g., Armijo, et al., 1996; Petit, et al., 2007; Sembroni, et al., 2016). In a recent online seminar
387 ([MCS RCN organized by the Community Surface Dynamics Modeling Systems](#)), Susanne
388 Buitter points out correctly that narrow and wide continental margins are commonly sediment
389 starved and rich respectively, which indicates that surface processes, particularly sediment
390 deposition, and the lithospheric extensional strain are related. The relationships that link surface
391 processes and lithospheric strain in extensional settings were studied through numerical models
392 that assume dominant hillslope-controlled erosion/deposition rates, \dot{e}_{hill} , commonly assessed
393 via a linear diffusion law such that, $\dot{e}_{hill} = k\nabla^2 z$, where k is a scaling coefficient (also referred
394 to as diffusivity) and z is the surface elevation (e.g., Burov & Cloetingh, 1997; Burov &
395 Poliakov, 2001; Buitter, et al., 2009; Sternai, 2020). Numerical models indicate particularly that
396 syn-extensional sediment deposition within rift basins produced opposite mechanical and
397 thermal effects. The increase in vertical stress involved by the deposition of sediments within
398 rifts basins enhances the lithostatic pressure and, thus, the brittle strength of crustal and mantle
399 rocks. On the other hand, thermal blanketing by sediment deposition prevents crustal rocks to
400 lose heat, thereby enhancing the viscous strain of the lithosphere (Fig. 5). By inhibiting
401 localized brittle strain and favouring distributed ductile flow of viscous rocks, sediment
402 deposition above a stretching lithosphere favours lateral migration of the extensional strain in

403 turn allowing for prolonged stretching and delayed continental lithospheric breakup (e.g.,
 404 Burov & Cloetingh, 1997; Buitter, et al., 2009; Sternai, 2020).



405

406 **Figure 5: (a) Effect of surface loading by deposition of a layer of sediments, h_s , on the brittle strain during extension**
 407 **(σ_1 : vertical stress; σ_3 : horizontal stress; σ_n : normal stress; σ_s : shear stress; ϕ : rock friction angle; ρ_s : density of**
 408 **sediments; g : gravity acceleration). Enhanced lithostatic pressure, P , due to sediment deposition within rifts basins**
 409 **enhances the brittle strength of crustal and mantle rocks and a greater σ_3 is required to rupture brittle rocks. (b)**
 410 **Comparison between two numerical models from Buitter, et al., 2009, to which the reader is referred for more detail.**
 411 **The simulations account for the same thermal conductivity of crustal rocks, C_c , and sediments filling the rift basin, C_s ,**
 412 **(upper panel) and lower thermal conductivity of the latter with respect to the former (lower panel), all other settings**
 413 **being the same. The thermal blanketing effect due to the low-conductivity sediment fill fosters the viscous strain of rocks**
 414 **(e.g., Buitter, et al., 2009; Sternai, 2020). (c) Schematic representation of the joint effects of surface processes on the**
 415 **brittle and ductile strain of lithospheric rocks during extension. Solid and dashed lines show the case with efficient and**
 416 **inefficient surface processes, respectively.**

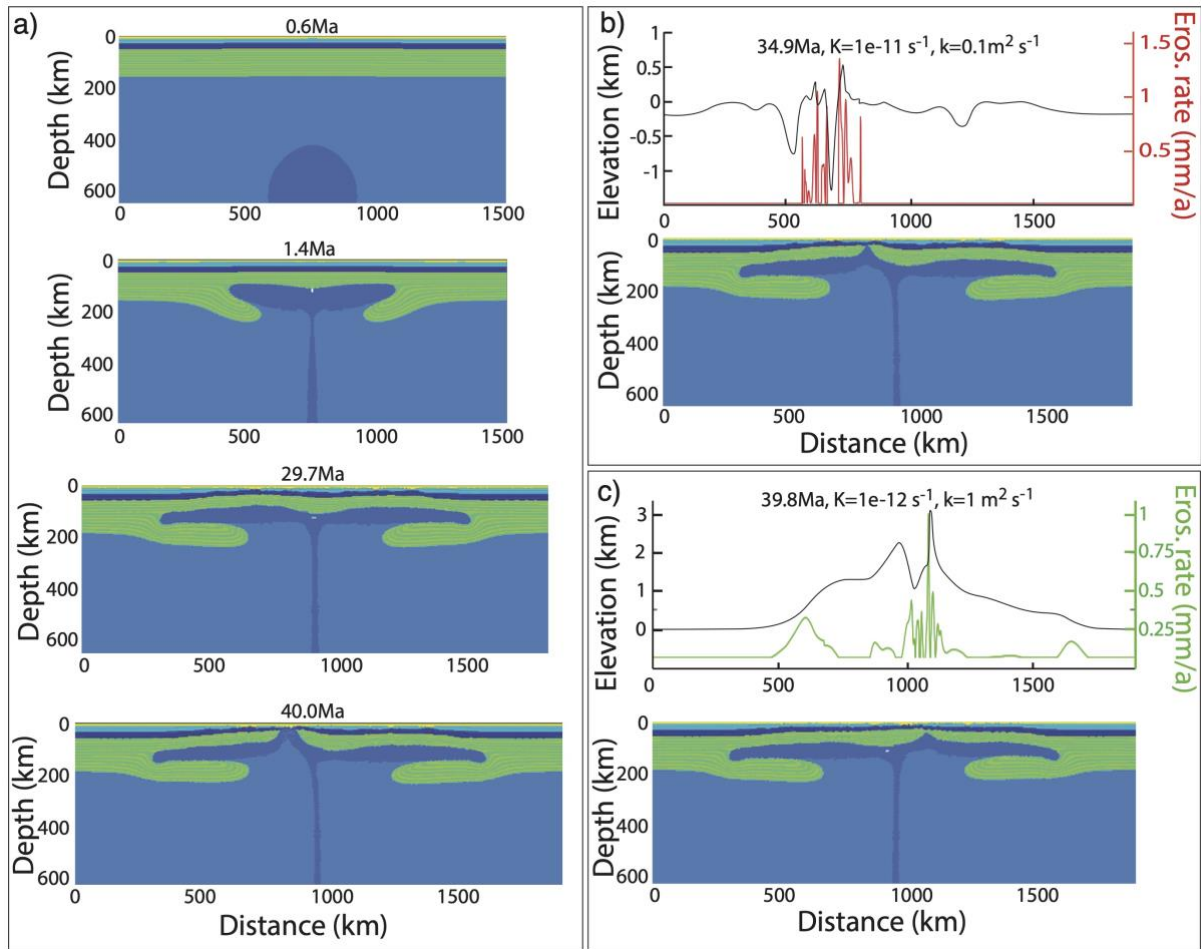
417

418 Syn-extensional and post-breakup magmatic units are nearly ubiquitous in extensional
419 settings and in some cases particularly voluminous (e.g., White & McKenzie, 1989; Franke,
420 2013), which makes extensional settings even more appealing for the study of the feedbacks
421 between internal and external dynamics. Magmatism likely provides a substantial contribution
422 to lithospheric rupturing (e.g., Kendall, et al., 2005; Lavecchia, et al., 2016) because extensional
423 stresses alone are estimated to be at most just enough to rupture the continental lithosphere
424 (e.g., Bott, 1991; Buck, 2004). The magma supply within fracture zones increases the pore fluid
425 pressure, thereby lowering the plastic yield strength of fractured rocks and further localising
426 the strain and topographic uplift or subsidence along weakening fault zones (e.g., Turcotte,
427 1982; Spence, et al., 1987; Connolly & Podladchikov, 1998; Gerya & Yuen, 2007; Katz, 2008;
428 Sternai, 2020). In this frame, peaks of igneous activity due to enhanced mantle decompression
429 melting have been ascribed to surface unloading by the deglaciation (e.g., Jull & McKenzie,
430 1996; Singer, et al., 1997) and/or erosion (Sternai, et al., 2016a) or sea level lowering (e.g.,
431 Crowley, et al., 2015; Sternai, et al., 2017). However, the sensitivity of extensional systems to
432 surface processes and the mechanisms that allow these latter to affect the production, transfer
433 and emplacement or eruption of magma are poorly constrained. Numerical modeling suggests
434 that flexural bending of the Moho due to efficient sediment delivery into a rift basin is an
435 efficient mechanism to enhance crustal melting in a stretching and warming lithosphere
436 (Sternai, 2020). Surface loading due to efficient filling of the rift basin, however, dampens
437 decompression partial melting of the asthenosphere by an amount proportional to the rate of
438 basin deepening/filling, the sediment density, and the surface-to-depth stress change transfer of
439 the rift system. For a given erosion/deposition rate, the modulation by surface processes to rock
440 melting in natural rift settings is inversely correlated to the extensional velocity, mantle
441 potential temperature (*sensu* McKenzie & Bickle, 1988) and initial Moho depth. In the extreme
442 case where surface processes redistribute the surface masses so efficiently to reset the

443 topography through time (i.e., the rate of erosion/deposition equals the rock uplift/subsidence
444 rate) the amount of crustal and mantle melts may respectively double and be reduced by ~50%
445 (Sternai, 2020). These modeling results does not account for brittle-plastic strain localisation
446 due to melts-rocks interactions (besides many other factors), so these estimates are likely
447 conservative. Increasing observational evidence corroborates these modeling results showing
448 that surface load changes in the order of the tens of MPa due to sea level changes during glacial
449 interglacial cycles can modulate the extensional magmatism (e.g., Crowley, et al., 2015;
450 Schindlbeck, et al., 2018; Kutterolf, et al., 2019; Satow, et al., 2021). Indeed, it seems somewhat
451 intuitive that a stretching and thinning lithosphere transmits surface stress changes at depth
452 more easily than a shortening and thickening lithosphere, implying that extensional systems
453 should be even more sensitive to the surface mass redistribution by surface processes than
454 convergent settings.

455 Bedrock fluvial incision plays a primary role in creating local relief in uplifting areas,
456 and this is also true for extensional settings. Geomorphologists have long proposed that the
457 bedrock fluvial erosion rate, \dot{e}_{fluv} , is proportional to the river stream power so that $\dot{e}_{fluv} =$
458 $KA^m \left| \frac{dz}{dx} \right|^n$, where $\frac{dz}{dx}$ is the local slope, A is the river drainage area, and K is the rock
459 erodibility, itself a function of precipitation and rock type (e.g., Gilbert, 1877). m and n are
460 arbitrary exponents usually determined by fitting models to river longitudinal profiles (e.g.,
461 Whipple & Tucker, 1999; Sternai, et al., 2012). Although simplistic, the definitions of \dot{e}_{hill} and
462 \dot{e}_{fluv} capture the main physics behind landscape evolution under hillslope- and fluvial-
463 dominated conditions and, most important for this chapter, they provide a straightforward
464 means to investigate feedbacks between the surface mass redistribution and lithospheric to sub-
465 lithospheric dynamics. The hillslope diffusion law has the clear advantage to reproduce jointly
466 erosion and sediment deposition and, thus, to better account for conservation of the surface
467 masses. The classical stream power erosion law, instead, does not account for sediment

468 deposition (i.e., no mass conservation) (e.g., Whipple & Tucker, 1999). Fig. 6 shows a rather
469 crude comparison between the resulting extensional continental lithospheric and sub-
470 lithospheric strain when hillslope or fluvial processes assessed via the diffusion and stream
471 power erosion laws are dominant. The numerical setup is similar to that described in Sternai,
472 2020, and Sternai, et al., 2021, to which the reader is referred for more information. This
473 comparison should be taken carefully because formulations of the stream power law that
474 account for sediment deposition do exist (e.g., Yuan, et al., 2019) and they should be
475 preferentially used for most applications. In addition, fluvial erosion leads to the formation of
476 hillslopes along river channels and, thus, $\dot{\epsilon}_{hill}$ and $\dot{\epsilon}_{fluv}$ are interdependent and should be
477 modeled jointly rather than alternatively or independently. However, the comparison shows
478 that different surface mass redistribution histories imply different effects on the deep dynamics
479 and, thus, informs us about the relevance of the physical treatment of surface processes in
480 geodynamic modeling. The classical stream power law is more effective in localizing the cross-
481 lithospheric strain, and even leads to earlier lithospheric breakup. However, this result is largely
482 due to the lack of a sediment deposition term and, thus, is reliable only if the natural sediment
483 routing brings the eroded material far out the extensional system. The main point is that typical
484 changes in the temporal and spatial pattern of surface mass redistribution may result in
485 protracted variations of the strain rate by up to a few orders of magnitude that reach down to
486 sub-lithospheric levels. Such variations imply changes in the location and amount of partial
487 rock melting, with further feedbacks on the strain pattern and topographic evolution. The
488 resulting chain of cause-effect relationships between surface, lithosphere and asthenosphere
489 dynamics are extremely complex and highly non-linear. I thus anticipate that unraveling the
490 substantial modifications to the architectural evolution of extensional systems involved by ever-
491 changing surface processes will be a tough but highly rewarding challenge for future research.



492

493 **Figure 6:** (a) Example of thermo-mechanical model of lithospheric extension above a mantle plume (the reader is
 494 referred to Sternai, 2020, and Sternai, et al., 2021, for more details on the numerical model). (b, c) Selected timesteps of
 495 simulations that account for different stream power or diffusion erosion rates (scaled based on K and k , as described in
 496 the definitions of \dot{e}_{fluv} and \dot{e}_{diff} reported in the text), all other settings being equal. The different redistribution of the
 497 surface masses through time due to different surface processes lead to major changes in, e.g., the location of lithospheric
 498 rupturing and, thus, the rift architecture and topographic evolution.

499

500 4.2. The geological carbon cycle

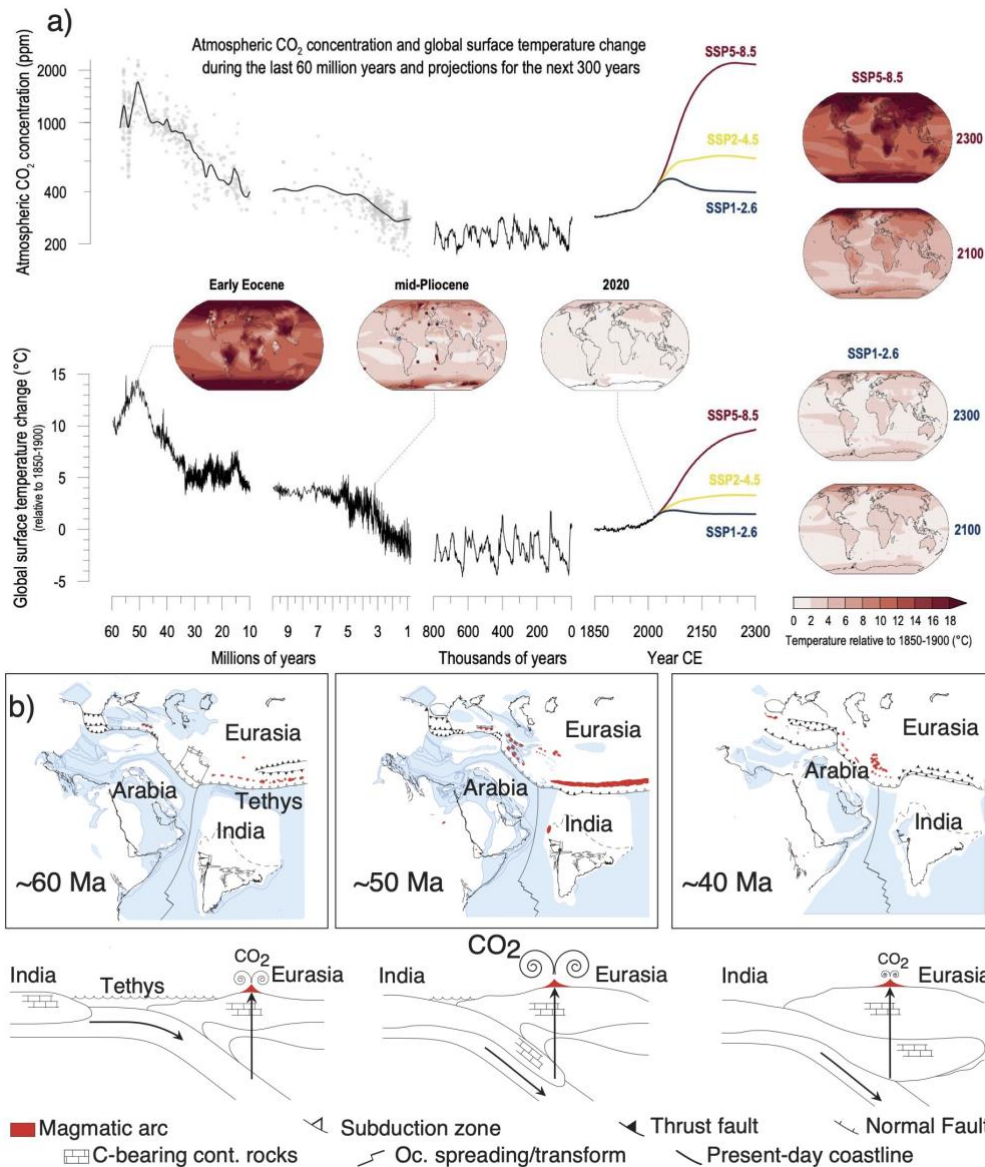
501 At timescales of millions to tens of millions of years, the Earth's carbon flows between
 502 the atmosphere, lithosphere, and mantle in an exchange called the geological carbon cycle (e.g.,
 503 Berner, 2003; Broecker, 2018). Carbon emissions from volcanic arcs above subduction zones
 504 are a critical input of carbon into the atmosphere, whereas the chemical weathering of silicate
 505 rocks exhumed at the Earth's surface through erosion of tectonically uplifted terrains consumes
 506 the atmospheric carbon (e.g., Lee & Lackey, 2015; Kelemen & Manning, 2015; Mason, et al.,
 507 2017). By preventing all the Earth's carbon from entering the oceans and atmosphere or being

508 stored within rocks, the geological carbon cycle keeps the Earth's temperature relatively stable
509 on the long term, like a global thermostat, linking the evolution of life and climate to plate
510 tectonics. The geological carbon cycle, thus, embodies the essence of the coupling between
511 tectonic and climatic changes.

512 Analyses of the sedimentary archives allow assessing the carbon outflux from the
513 atmosphere through erosion and chemical weathering. Instead, the uncertainty regarding the
514 amounts and driving mechanisms of carbon recycling and emissions from the Earth's interior
515 stands out as one of the most vexing problems facing us in understanding the geological carbon
516 cycle (e.g., Berner & Lasaga, 1989; Dasgupta, et al., 2007; Burton, et al., 2013). Available
517 estimates of current carbon fluxes between the Earth's deep and surface reservoirs, for instance,
518 vary by several orders of magnitude (e.g., Dasgupta, et al., 2007; Kelemen & Manning, 2015),
519 which is indicative of how little we know about this branch of the carbon exchange cycle.
520 Multidisciplinary integrations of geological data and modeling to quantify variations in global
521 carbon emissions due to fundamental geodynamic events throughout the Earth's history and
522 their critical effects on climate change represent a top-priority challenge for future research on
523 the natural carbon cycle, the quantitative understanding of which is also a fundamental
524 objective of the International Panel on Climate Change - IPCC (www.ipcc.ch).

525 The challenge is harsh but not impossible, especially if focused on geological times for
526 which there are well preserved observable records. I see in the Cenozoic an ideal time window
527 for this research because the global tectonic and climatic framework are relatively well
528 established (Fig. 7). The progressive closure of the Neo-Tethys and the rise of the Alpine-
529 Himalayan range perturbed the global atmospheric and ocean circulation, leading to climate
530 warming between ~60-50 Ma followed by unsteady cooling ever since (e.g., Ruddiman &
531 Raymo, 1988; Ruddiman & Kutzbach, 1989; Le Houedec, et al., 2012). The development of
532 high topography and closure of marine gateways alone are insufficient to cause the observed

533 long term Cenozoic climate trends, and changes in the concentration of CO₂ and other radiative
534 important gasses are required (e.g., Raymo & Ruddiman, 1992). Two end-member hypotheses
535 exist: the overall changes in atmospheric carbon budget and associated Cenozoic climate trends
536 are driven primarily by variations of global chemical weathering, the main atmospheric carbon
537 output term, or by variations in global degassing by magmatic and/or metamorphic processes,
538 the main input term. A currently established paradigm states that the India-Eurasia collision
539 and uplift of Tibet led to intense monsoonal circulation, increased rainfall on the Himalayas,
540 greater rates of rock weathering and, ultimately, lower atmospheric CO₂ concentrations and
541 long-term Cenozoic climate cooling (e.g., Raymo & Ruddiman, 1992). This view is often
542 associated with the proposal that Cenozoic climate cooling led to Plio-Quaternary increase in
543 global mountain elevation and sedimentation/erosion (Molnar & England, 1990). This proposal
544 embodies a true chicken or egg problem because erosion and weathering consume atmospheric
545 carbon and thus would be the cause, rather than the effect, of climate cooling.



546

547 **Figure 7:** (a) Changes in atmospheric CO₂ and global surface temperature (relative to 1850-1900) throughout the
 548 Cenozoic and for the next 300 years (modified after [IPCC 2021 report](#), to which the reader is referred for details on
 549 sources and data analyses). (b) Paleotectonic maps and cross-sections of the Neo-Tethyan margin during the lower
 550 Cenozoic (modified after the paleotectonic reconstructions by the DARIUS program, 2018,
 551 <http://istep.dgs.jussieu.fr/darius/maps.html> and Sternai et al., 2020).

552 On the one hand, the increase in global weathering paradigm and the chicken or egg
 553 controversy motivated a huge amount of research that greatly advanced our knowledge on the
 554 tectonics-climate interactions from the weathering perspective (i.e., atmospheric carbon sink).
 555 On the other hand, the role of changes in continental degassing or organic carbon erosion and
 556 oxidization from marine sediments on the passive and active Neo-Tethyan margins during the
 557 early Indian-Asian collision (i.e., the atmospheric carbon source perspective) have been

558 significantly overlooked (e.g., Beck, et al., 1995; Sternai, et al., 2020) and major climatic trends
559 such as the early Eocene climate optimum (EECO, i.e., the ~10 Ma climate warming preceding
560 post-50 Ma cooling) remain currently unexplained. I recently happened to assist a Professorship
561 promotion where eminent geologists did not account that the India-Asia collision since the early
562 Cenozoic involved the progressive demise of several thousands of kilometers of volcanic arcs
563 which were continuously and likely unsteadily emitting carbon into the ocean and atmosphere
564 since at least the Cretaceous. The entire Andean or Indonesian arcs today, which account for
565 some of the greatest CO₂ outgassing volcanoes worldwide (e.g., Burton, et al., 2013), are
566 modern analogues of these former arcs. I find it hard to believe that extinguishing emissions of
567 greenhouse gasses from these arcs would leave climate unaltered. Recent studies further suggest
568 that global Cenozoic silicate weathering fluxes weaker than previously thought (e.g., Tipper, et
569 al., 2020) are, at least to some extent, compensated by weathering of accessory carbonate and
570 sulfide minerals, a geologically relevant source of CO₂ (e.g., Torres, et al., 2017; Komar &
571 Zeebe, 2021), suggesting that other factors must be at play. Plausible changes in CO₂ emissions
572 from the Neo-Tethyan magmatic arcs during convergence are strikingly consistent with both
573 the ~60-50 Ma climate warming and post-50 Ma cooling (Fig. 7), and there is a general
574 agreement that Neo-Tethyan margin is key to global Cenozoic climate change (see e.g., Sternai,
575 et al., 2020, for a review on the topic). However, other major geodynamic events involving
576 likely significant climatic effects occurred worldwide during the early Cenozoic, for instance
577 the propagation of Atlantic Ocean spreading along the Reykjanes Ridge or of East African Rift
578 System (e.g., Seton, et al., 2012; Ebinger, 2005). I anticipate that integrated investigations
579 accounting for magmatism as well will considerably broaden our knowledge about how plate
580 tectonics affected the climate evolution. To this aim, quantification of the CO₂ content and
581 petrogenic characterization of preserved magmatic products as well as integration between
582 numerical geodynamic and thermodynamic models to explore how dynamic processes affect

583 the transformations of carbon-bearing rocks in different geological settings can provide direct
584 and quantitative constraints on CO₂ outgassing throughout the Cenozoic or longer timescales.

585

586 **4.3. Feedbacks between internal and external dynamics and effects on the evolution of life**

587 Several elements and nutrients useful to life (besides carbon) are cyclically transferred
588 through surface and deep reservoirs during geological timescales. This implies that geodynamic
589 events are intrinsically linked not only to the evolution of climate and the landscape, but also
590 to the biosphere. Increasing evidence indicates that the establishment of modern-style plate
591 tectonics contributed to the development of complex life on our Planet (e.g., DePaolo, et al.,
592 2008; Sobolev, et al., 2011; Stern, 2016; Zaffos, et al., 2017; Lee, et al., 2018). A global
593 continuously evolving mosaic of lithospheric plates, likely established gradually during the
594 geological past (e.g., Gerya, 2014; Sobolev & Brown, 2019; Gerya, 2019), supplies and
595 withdrawals nutrients and produces variations of environmental conditions that foster genetical
596 modifications and thus the evolution of life (e.g., Zerkle, 2018; Descombes, et al., 2018). The
597 redistribution of continents, the growth of mountains, the rise and demise of volcanic and
598 magmatic arcs, the opening and closing of marine gateways also produce moderate
599 environmental ‘stress’ that stimulates populations to adapt and evolve (e.g., Stern, 2016).
600 Indeed, some of the characteristic timescales of biological evolution are comparable to those at
601 which geodynamic reorganizations occur, which may be taken as further evidence of the
602 coupling between internal and external dynamics (e.g., DePaolo, et al., 2008). Research efforts
603 to quantitatively assess the mechanisms through which geodynamic processes influence the
604 evolution and diversification of species are a natural and logical step forth to enhance
605 knowledge on the feedbacks between internal and external dynamics and their multiple
606 implications. To this aim, geodynamicists, geologists, (geo)biologists, ecologists, geochemists,
607 palaeontologists, geomorphologists, and climate experts must cooperate to integrate/interpret

608 biological data with geodynamic, landscape evolution, carbon cycle and climate (*sensu lato*)
609 modeling. I anticipate that the necessary major cross-community efforts will lead to highly
610 rewarding discoveries about our history and that of our Planet.

611

612 **5. Closing Remarks**

613 Recognizing and quantifying how and to what extent internal and external dynamics are
614 linked has long been one of the grand challenges in the Earth Sciences (e.g., DePaolo, et al.,
615 2008; Huntington, et al., 2017). Overtaking this challenge is more than ever critical now that
616 warming temperatures, melting ice caps, rising sea levels, alteration of natural environments
617 and biodiversity, increasing catastrophic events, and pandemics have become undeniable
618 effects of the current anthropogenic climate crisis ([IPCC 2021 report](#)). It is essential to realize
619 that our ability to face the current climate emergency and the threatening global socio-economic
620 consequences entailed (Fig. 7) depends heavily on our understanding of the functioning of the
621 Earth system. Monitoring ongoing climate change will help improving warning systems and
622 the management and use of natural resources, but it will not improve our ability to anticipate
623 cascading effects and act consciously and effectively to mitigate future events and re-establish
624 the natural variability of climate. Only the study of the geological archives through ever-
625 improving observations and modeling tools can elucidate on the causes and effects of the
626 natural climate variability, thereby providing the baseline for understanding the current climate
627 state and forecasting future changes. I thus believe that research on the feedbacks between
628 internal and external dynamics will provide a critical help for decision makers for finding
629 optimal policies to lead us out the ongoing emergency.

630 It is increasingly clear that in-depth understanding of the functioning of our Planet can
631 only arise from the study of the atmosphere and oceans, but also of the landscape, the
632 lithosphere, and the mantle not to mention our societies and social behaviour, which delve into

633 it deeply. In fact, no problem of any significance can be perceived from within a single
634 compartmentalised discipline, but also the smallest topic, however seemingly minute, can only
635 be understood within and through its wider context. Complex systems commonly enhance and
636 inhibit qualities specific to individual components and, thus, are at the same time more and less
637 than the sum of their parts (Aristoteles, *Metafisica*). The significance of the study of individual
638 parts is inversely proportional to the complexity of the system if one aims at understanding the
639 functioning of the complex system itself. Trans-disciplinarity is an essential pre-requisite to
640 understanding Nature (i.e., the Earth system). Although this may seem a trivial statement, a
641 compartmentalized view of the Earth Sciences, where a scientist's work is meant to study the
642 Discipline rather than Nature, is still very much present and severely limits new relevant
643 scientific discoveries. I encourage any Earth Scientist concerned with the functioning of our
644 Planet to engage into a 'complex' thinking that involves as many disciplines and, thus,
645 collaborations as possible.

646 Since the advent of significant computing power, another just as harmful methodological
647 distinction between 'modelers' and 'observational' Earth Scientists added up to the classical
648 separation of the Earth Sciences into several compartmentalized disciplines. This distinction is
649 pure nonsense. When a complex system such as the Earth is investigated, observations and
650 models need to mutually draw meaning from each other. It is rarely straightforward to grasp
651 useful information from geological observations because the functioning of the Earth, which is
652 what generates those observables, is not so either. Most often, geological observables contain
653 mixed information about multiple processes, and we need models to unlock this information.
654 For instance, deformation structures into rocks contain information about the temporal and
655 spatial evolution of the stress and thermal fields responsible for those structures, but obtaining
656 such information requires a rheological model for rocks to say the least. Paleo-shorelines can
657 contain information about the postglacial rebound, but obtaining this information requires a

658 rheological model for the lithosphere and mantle as well as a model of the glacial ice sheet and
659 sea level variations during the time window of interest. Sedimentary structures can contain
660 information about eustatic changes, but obtaining this information requires a model for the
661 sediment production, routing system and deposition, which itself may depend on a tectonic
662 model that describes the evolution of the sediment source and sink areas. Just like doors need
663 the right key to be opened, the information within geological observables cannot be accessed
664 without the right model. Even trickier, the information delivered by geological observables can
665 be very different, even wrong, depending on which model we use to access it.

666 Modern computers allow us to generate and handle extremely complex numerical models,
667 which means we have increased potential to obtain useful information from observables. We
668 can even generate synthetic ‘observables’ that inform us about inaccessible parts of the Earth
669 or theoretical scenarios involving processes and interactions of processes that we cannot
670 observe directly. If used correctly, then recent numerical models can be for Earth Scientists
671 what a wind tunnel is for Engineers, that is a tool to test to test the response of a specific system
672 to given forcings (i.e., the ‘performance of a prototype’, in engineering terms) as a term of
673 comparison for real observations, or as synthetic data in absence of observations. Whereas there
674 is usually no aversion toward simple models, complex models are often regarded to as science-
675 fiction. Clearly, any scientist loves the sight of a straight line of data points because it can be
676 fit by a linear (i.e., simple) model. However, observations that report on the functioning of the
677 Earth system are practically never aligned, nor they should be because they arise from complex
678 interactions between complex processes. Interpreting these data with too simple a model (e.g.,
679 linear regression) may bring wrong information leading to flawed conclusions just as much as
680 interpreting the data with too complex a model that includes multiple poorly constrained
681 parameters. Being aware of the flaws and limitations of our data and models (e.g., van Zelst, et

682 al., 2021), rather than dismissing one or the other, should be the main concern of Earth
683 Scientists, regardless of their observational or modeling background.

684

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691

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