

New maps of global geological provinces and tectonic plates

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Abstract

Accurate spatial models of tectonic plates and geological terranes are important for analyzing and interpreting a wide variety of geoscientific data and developing compositional and physical models of the lithosphere. We present a global compilation of active plate boundaries and geological provinces in a shapefile format with interpretive attributes (e.g., crust type, plate type, province type, last orogeny). The initial plate and province boundaries are constructed from a combination of published global and regional models that we refine using a variety of geoscientific constraints including, but not limited to, relative GPS motions, earthquakes, mapped faults, potential field characteristics, and geochronology. These new plate model show improved correlation to observed earthquake and volcano occurrences within deformation zones and microplates, compared to existing models, capturing 73 and 80% of these criteria, respectively. We estimate 57.7% of the Earth's surface is covered by oceanic crust, which is a slight increase relative to the most recent seafloor age model. The model of last orogenies agrees well with peaks in the globally summed geochronology data. There is room for improvement in future editions of our global plate and geologic provinces model where basins, ice, or lack of geological data fidelity obscure bedrock geology, particularly in the eastern Central Asian Orogenic Belt, much of Africa, East Antarctica, and eastern Australia. Additionally, some province types—ogens, shields, and cratons that are homogenized within our global scheme—can likely be partitioned into smaller terranes with more precise geodynamic attributes. Despite some of these shortcomings, the digital maps presented here form a self-consistent data standard for adding spatial metadata to geoscientific databases. The database is available on GitHub where the geoscience community can provide updates to improve the models and their contemporaneity as new knowledge is acquired. The files are also released in formats suitable for use in Generic Mapping Tools and GoogleEarth.

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Keywords: tectonic plate, tectonic province, orogenic system, orogeny, geodynamics, geospatial analysis

Contents

1	Introduction	4
2	Existing Global Models of Tectonic Plates and Provinces	4
3	Method of Construction	8
3.1	Plate Model	10
3.1.1	Plate Type	11
3.1.2	Plate Boundary Type	13
3.1.3	Ocean–Continent Boundary	14
3.1.4	Oceanic Domain	15
3.2	Global Geologic Province Model	16
3.2.1	Province Type	21
3.2.2	Last Orogeny	24
3.2.2.1	<i>Active Orogens</i>	
	· · · · ·	27
3.2.2.2	<i>Neoproterozoic to Mesozoic Orogens</i>	
	· · · · ·	30
3.2.2.3	<i>Paleo- to Mesoproterozoic Orogens</i>	
	· · · · ·	33
4	Model Evaluation	41
4.1	Ocean–Continent Boundary	41
4.2	Plate Boundaries	43
4.3	Plate Model	44
4.3.1	Comparison with Bird (2003)	44
4.3.2	Comparison with Tomography and Volcanism	46
4.3.3	Uncertain Plate Boundaries	49
4.4	Geologic Province Model	50
4.4.1	Comparison with Matthews et al. (2016) and Meredith et al. (2021) . . .	52
4.4.2	Last Orogeny	52
4.4.3	Uncertain Province Boundaries	55

4.5 Future Improvements	57
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5 Summary	58
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1 1. Introduction

2 The structural architecture, tectonic environment, and temporal evolution of rocks at the
3 surface of the Earth are frequently correlated with the chemical and physical
4 characteristics of the enclosing lithosphere (Gard et al., 2019b; Artemieva, 2019; Tang et al.,
5 2020; Tetley et al., 2020). As a result, it is useful to have spatially accurate maps of geologic
6 provinces and terrane boundaries that encompass a pragmatically uniform set of common geo-
7 logical characteristics for comparative global studies. Such maps also form the foundation for
8 accurate plate reconstructions (Merdith et al., 2021). While there are some regional models
9 of tectonic provinces that are digital (Artemieva, 2006; Laske et al., 2013), there are few ac-
10 curate global models easily accessible to the geoscience community built on a multiplicity of
11 comparative attributes that approach self-consistency.

12 In this paper, we present two basic models: (1) a global set of geologic provinces and
13 (2) a model for present-day plate boundaries. Both models are presented in a vector format
14 with accompanying metadata that can be used to improve and simplify the process of global
15 tectonic data analysis and/or modeling across a diverse range of geoscientific phenomena. These
16 models have been produced using a wide variety of geologic and geophysical data and have been
17 partially validated, wherever possible, using igneous and metamorphic age dates allied with
18 additional geophysical datasets. Our hope is these models can be used as a data standard for
19 common classification across the variety of geological databases that currently exist. The global
20 models presented below are freely available in open-source and form a basic digital architecture
21 that can be progressively updated as geological data and interpretations continue to improve.

22 2. Existing Global Models of Tectonic Plates and Provinces

23 Previous global plate and province models have been published that incorporate tectonic
24 setting, juvenile age, or thermotectonic age (e.g., Artemieva, 2006; Goutorbe et al., 2011; Laske
25 et al., 2013; Szwilus et al., 2019). Ideally, the digital nature of these maps makes them easy
26 to use and allows one to add desirable attributes to underlying datasets. However, the raster

format of these models is often an impediment to accurate spatial analysis at or near province boundaries due to their low resolution and pixelated nature. Even though some plate and province maps include age and province type, it may not be possible to separate individual terranes that potentially have distinctive chemical and physical characteristics. For example, a terrane-type map ([Laske et al., 2013](#), e.g.,) can be useful for identifying data within a volcanic arc setting, but it can be difficult to separate individual volcanic arcs to compare temporal and spatial geochemical patterns that illuminate the geodynamic character. Furthermore, while some of these maps have been included in peer-reviewed publications as part of global studies, it is unclear to what degree the maps themselves have been examined in detail because the underlying geological data used to construct the maps is not available to the geological community. In some cases, it is difficult to obtain digital versions from the authors, making it challenging to validate or improve the models as a geoscientific community and achieve widespread use. As a consequence, some of these maps of global geological provinces are essentially artistic and have opaque underlying rationale that cannot be interrogated.

A few global shapefiles of province polygons do exist ([Klett et al., 1997](#); [Torsvik and Cocks, 2016](#); [Matthews et al., 2016](#)), but they cannot be accurately matched to the province boundaries as they are identified in regional studies. For example, the [Klett et al. \(1997\)](#) proposed geologic provinces were developed to assess global hydrocarbon reserves. Therefore, a narrow perceived attribute rather than broad actual attributes were used to delineate geological provinces. The models by [Torsvik and Cocks \(2016\)](#), [Matthews et al. \(2016\)](#), and [Merdith et al. \(2021\)](#) are based on tectonic blocks as they were developed to perform plate reconstructions (Figure 1b), but they are typically shape-defined by contemporary global geography. Some terrane boundaries agree well with published models (e.g., Africa, [McCourt et al. \(2013\)](#); South America, [Ibañez-Mejia et al. \(2011\)](#)), while others appear greatly simplified and/or do not closely follow geophysical trends for reasons that are not given (e.g., western United States, [Hasterok and Chapman \(2007\)](#)). Furthermore, many models are time diluted. For example, models by [Matthews et al. \(2016\)](#) and [Merdith et al. \(2021\)](#) were developed for reconstructions to 400 Ma and 1000 Ma, respectively. As a result, most Mesoproterozoic and older regions are not divided into separate terranes except where they behave as separate entities during the plate model timeframe, despite a wealth of data that would otherwise allow the organization of these older participants to be

57 illuminated.

58 In addition to the lack of globally a comprehensive and self-consistent scheme for the de-
59 piction of geological provinces, there are still ambiguities and lack of self-consistency in con-
60 temporary plate tectonic maps. Bird (2003) released a widely used plate boundary shapefile
61 (Figure 1a). Since its release, several additional microplates have been proposed, generally
62 on the basis of GPS motion data. The new proposed microplates are the Adria Microplate
63 (Battaglia et al., 2004; Breton et al., 2017), the Danakil Microplate (Eagles et al., 2002; Mc-
64 Clusky et al., 2010), the Yakutat Microplate (Fletcher and Freymueller, 1999; Bruhn et al.,
65 2012), the Sierra Nevada Microplate (Dixon et al., 2000; Schweickert et al., 2004), several mi-
66 croplates around the Caribbean (DeMets and Wiggins-Grandison, 2007; Sun et al., 2020), and
67 few microplates on the Somali Plate (Saria et al., 2014; Stamps et al., 2021). The Bird model
68 is also missing a few microplates proposed at the time of publication including the Capricorn
69 Microplate and Indo-Australian Deformation zone (Royer and Gordon, 1997), the Coiba and
70 more recently associated Malpelo Microplates (Hardy, 1991; Zhang et al., 2017). The Bird
71 model included some plate boundary deformation zones, but many zones of known deformation
72 were excluded from the model (e.g., Gordon, 1998). Additional improvements can be gained by
73 incorporating more recent kinematic models of strain within plate boundaries to discriminate
74 between discrete and more continuously deforming zones (Kreemer et al., 2014).

75 Beyond the global models discussed above, there exists a plethora of regional geologic
76 province and terrane maps that are downloadable. Many of these models are more precise
77 because they may incorporate a variety of datasets as constraints (e.g., topography, bedrock
78 geology, seismic tomography, and crustal faults or shear zones), but they are poorly designed
79 for digital processing as they are only available as raster images in papers despite many be-
80 ing created using geographic software. While these depictions are sufficient for presentation
81 in publications, the rationale and metadata are generally not available for scrutiny. Conse-
82 quently, interpreted tectonic boundaries are often inconsistent between publications, especially
83 when boundaries are envisaged beneath cover. In many publications (e.g., Gee and Stephen-
84 son, 2006; Xu et al., 2016; He et al., 2018), a province or terrane map is often secondary to
85 the main point of the paper, and boundary definition may be schematic rather than rigorous
86 and is commonly simply adopted from pre-existing publications. An obvious example is the

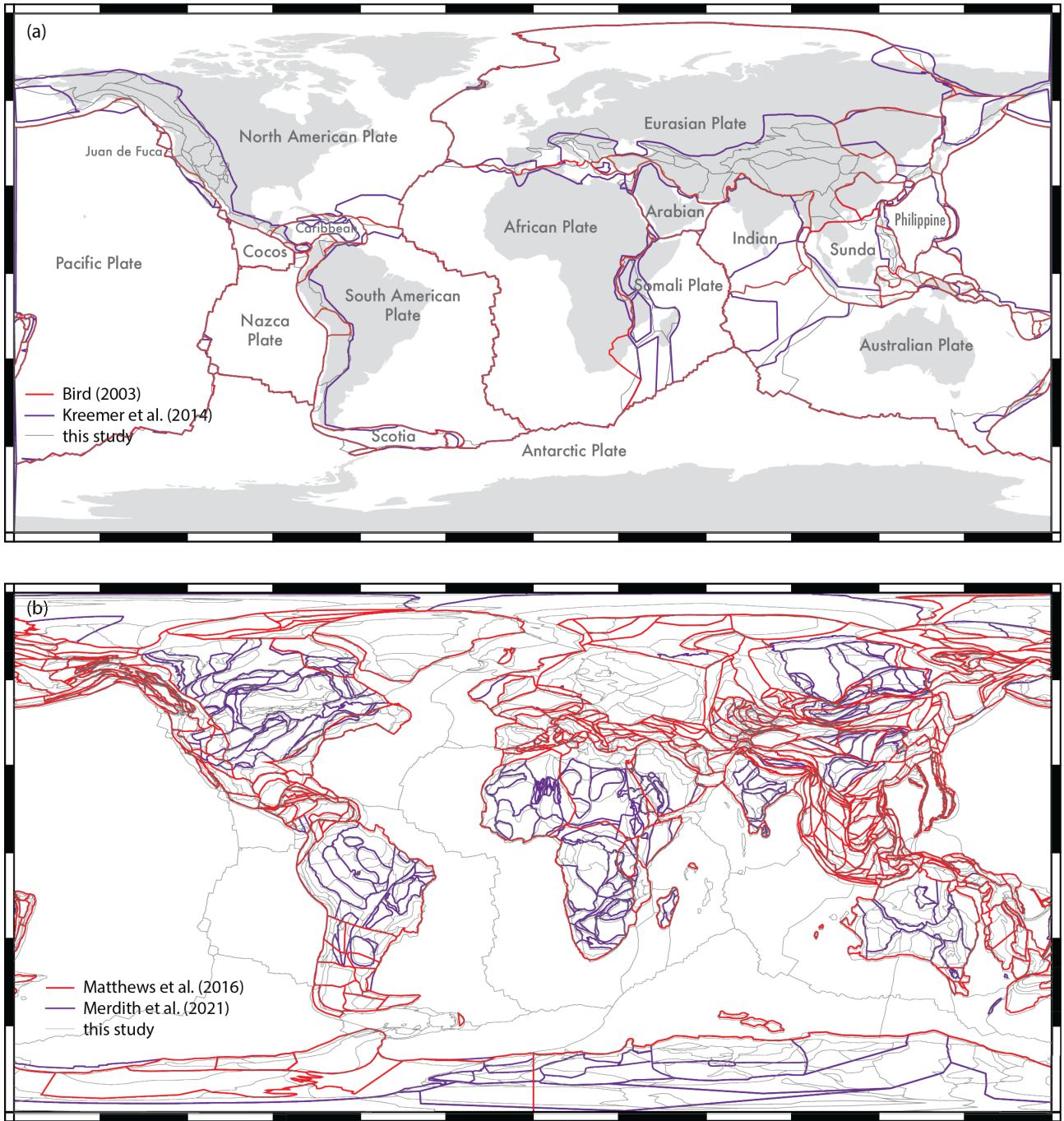


Figure 1: Tectonic plate and geologic province models. (a) A comparison between the [Bird \(2003\)](#) plate model, limit of modeled plate boundary zones ([Kreemer et al., 2014](#)) and the plate model from this study. Note that the limit of the deformation model domain is not an exact limit of the deformation, in many cases wider than the true deformation zone. (b) A comparison between the [Matthews et al. \(2016\)](#) and [Merdith et al. \(2021\)](#) plate models and province models from this study.

depiction of the Sahara Metacraton (Figure 14 and Figure 1 of Liégeois et al., 2013; Kwékam et al., 2020, respectively). Furthermore, the locations and level of detail lack continuity between publications, such that province boundaries cannot be easily matched up between maps of adjacent regions, even when apparently depicted in the same map projection scheme (e.g., wei Zhou and Murphy, 2005; Blayney et al., 2019).

The objective of the work we present here is to try and smooth deficiencies between existing models for tectonic province and plate boundaries by delineating a set boundaries in a vectorized format that incorporates attributes derived from geophysical, geochemical and geochronological data. Our models represent an attempt to extend tectonic boundaries back to ca. 2.3 Ga and possibly further in some cases, incorporating numerous existing geologic interpretations. Some of these interpretations are modified slightly to meet additional observational constraints from global geophysics, geologic maps and geochronology. Our goal is to create an adaptable and interactive environment that allows the geoscience community to improve delineation of geological provinces and the behaviors of their boundaries and interiors during Earth evolution. In particular, we hope creating an interactive data environment will bring more illumination to the Mesoproterozoic and older Earth, which hosts a significant fraction of metallic resources.

3. Method of Construction

The maps constructed in this study come from four separate shapefiles—also released in GMT and KML formats for use in Generic Mapping Tools and GoogleEarth, respectively. These files include the plate polygons, tectonic province polygons, the oceanic–continental crustal boundary, and plate boundary types. The ocean–continent boundary and plate boundary types are both developed in conjunction with the plate polygons. Each of these files contains a number of attributes that include a variety of contextual information. The metadata for each file are described in Table 1.

Vector format shapefiles have several advantages over the raster maps that dominate the existing literature. The polygons and lines created across the four shapefiles are seamless, i.e., they use common boundaries where geologically appropriate. Vectorized data permit multiple attributes to be assigned to each polygon, line, or point, which can be unique. Raster models allow only a single attribute per pixel, often requiring multiple maps (generally from separate

Table 1: A description of shapefile attributes for the plate and province files released in this study.

Field name	Field description
<i>Plate polygons</i> (plates.shp)	
id	unique polygon identifier
poly_name	unique polygon name
plate	major plate, includes microplates and deformation zones
plate_id	numeric subplate id
plate_code	subplate abbreviation
subplate	separates microplates and deformation zones from major plates
plate_type	rigid plate, microplate, or deformation zone
crust_type	continental or oceanic crust
sea_name	name of ocean or sea
domain	oceanic domain for geochemical grouping purposes
area	polygon area in square kilometers
plate_ref	reference for initial plate polygon
<i>Plate boundaries</i> (plate_boundaries.shp)	
feature_id	unique boundary segment identifier
feature	name of boundary segment
type	type of boundary segment
plate1 and plate2	subplates on either side of boundary segment
level	assigns an integer value, 1 = major plate boundaries and 2 = minor plate boundaries
comment	specific comments about a boundary segment
length	length of segment in kilometers
<i>Ocean–continent boundary</i> (oc_boundary.shp)	
id	line segment identification number
length	length of segment in kilometers
<i>Province polygons</i> (geologic_provinces.shp)	
id	unique polygon identifier
prov_name	unique province name
prov_type	dominant tectonic character of a tectonic province
prov_group	name for multiple polygons with a shared geological history, may contain multiple tectonic styles
lastorogen	most recent significant orogenic event
continent	continent name if on a continent
crust_type	continental, oceanic, or transitional
area	polygon area in square kilometers
comment	specific comments about a province
prov_ref	reference for initial province polygon

116 studies), which could lead to non-causal juxtaposition of such attributes. For example, slight
117 differences in the boundaries of different attributes from unrelated studies could result in oceanic
118 crust being incorporated into a continental orogenic belt or continental crust being excluded
119 from the same orogenic system. Vector format files are also typically more memory efficient
120 than raster images. Thus the seamless nature of the polygons in this project is a distinct
121 advantage when constructing physical and/or compositional models of the lithosphere.

122 We recognize the study presented here attempts to do on a global scale, what a large number
123 of studies have done at a variety of subordinate scales. Therefore we are aware of an element
124 of hypocrisy in being somewhat critical of continuity between existing studies. However our
125 goal isn't to be correct, it's to create an environment where decisions about tectonic provinces
126 and their boundaries are globally determined where possible, using the same style of data
127 sets and their interpretation. Where our study differs from predecessors, is the compiled data
128 used is freely available and adoptable by more informed practitioners because the models are
129 available on GitHub (https://github.com/dhasterok/global_tectonics) where community
130 users can correct errors and omission and propose refinements.

131 *3.1. Plate Model*

132 The plate model consists of two separate shapefiles including the plate polygons and the
133 boundary lines (Table 1; Figure 2). The plate shapefiles were created in QGIS using global
134 vector and raster datasets (Table 2). We used the widely-distributed model by Bird (2003)
135 as the initial plate boundaries to construct the shapefile (Figure 1a). Newer models that
136 incorporate proposed past plates and microplates for use with plate reconstruction software
137 such as GPlates were also used for reference (i.e., Zahirovic et al., 2014; Matthews et al., 2016).
138 Although Bird's model is an excellent framework, increased spatial coverage of GPS and an
139 additional 18 years of earthquake observations has improved our ability to recognize additional
140 microplates and identify presently deforming regions. The strain-rate model by Kreemer et al.
141 (2014) is also used as an initial constraint on the boundaries of microplates and deformation
142 zones and as a way to distinguish between the two (Figure 1). Digital elevation models and
143 global models of active faults have also improved since Bird's model (Amante and Eakins,
144 2009; Styron and Pagani, 2020), which allow for more accurate and precise positioning of the
145 boundaries.

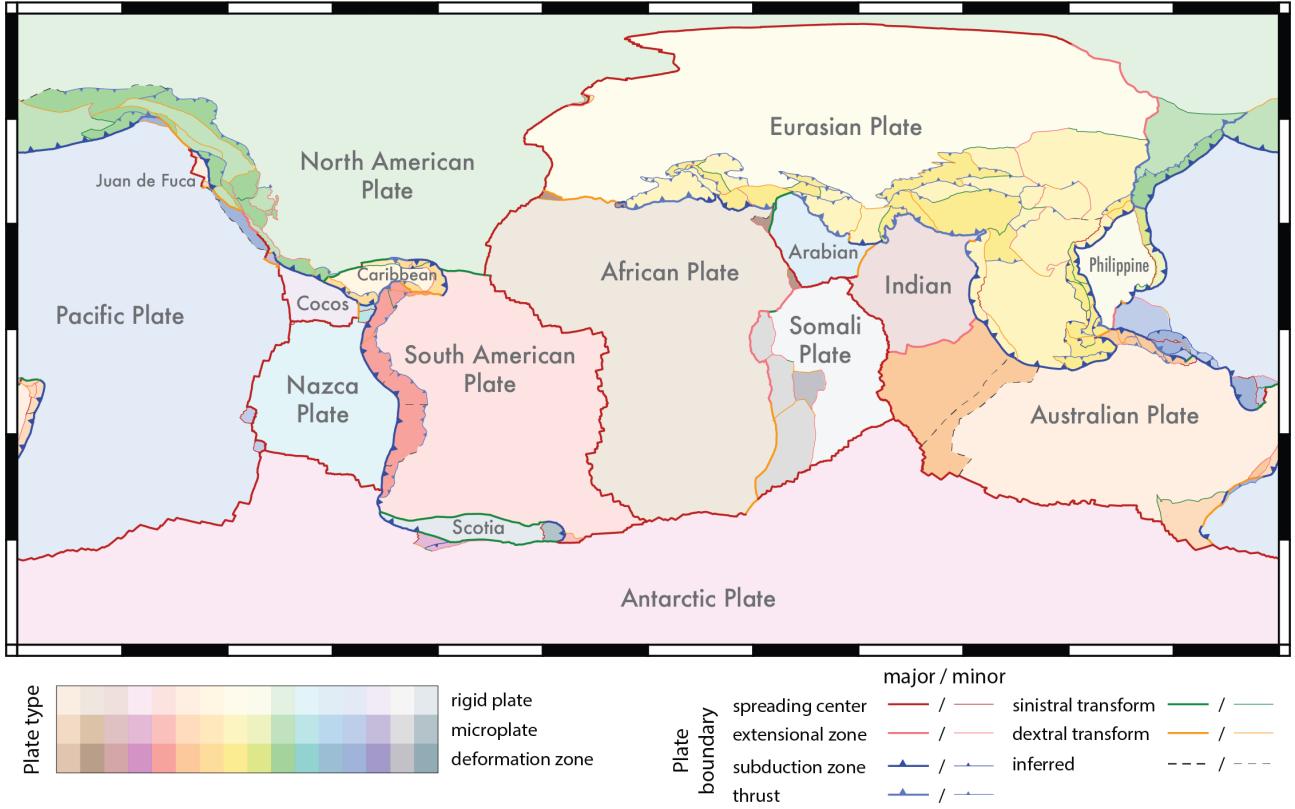


Figure 2: An updated plate model along with the plate boundary types. Microplates and deformation zones are illustrated as darker shades of the same hue as the major plate that they are most closely associated with.

146 Where there is ambiguity in the location of the plate boundary from topography and its
 147 gradient, the location was chosen to fit with the pattern of recent seismicity (Figure 3). We
 148 found it helpful to compute a spatial histogram of earthquakes because it is easier to identify
 149 zones of high seismicity relative to a simple scatter plot. To produce the histogram, we use global
 150 seismicity M5.5+ for 1970–1990 and M3.5+ for 1990–2020 extracted from the US Geological
 151 Survey’s Advanced National Seismic System (ANSS) global seismic catalog. The earthquake
 152 density accurately traces out much of the mid-ocean ridge systems (Figure 3).

153 3.1.1. Plate Type

154 We include metadata with the plate model, including a plate_type field. The following
 155 definitions are used to distinguish plate types:

156 **fixed plate**, a region with distinct plate boundaries generally defined by seismicity, little in-
 157 ternal deformation, distinct motion relative to several other plates, and generally large
 158 area;

159 **microplate**, a region with distinct plate boundaries generally defined by seismicity, little in-
 160 ternal deformation, motion controlled by surrounding plates, and generally small area;

Table 2: Datasets used to develop and evaluate the plate, plate boundary, and oceanic–continental boundary models.

Region	Data Type	Description	Resolution	Reference
global	plate model	shapefile		Bird (2003)
global	plate model	GPlates shapefile		Zahirovic et al. (2014)
global	plate model	GPlates shapefile		Matthews et al. (2016)
global	earthquakes	1990–2020, <30 km, M3–M5.5		ANSS (2020)
global	earthquakes	1970–2020, <30 km, M5.5+		ANSS (2020)
global	earthquakes	heat map (ANSS data above)	0.1 degree	
global	volcanic centers			Global Volcanism Program (2013)
global	topography	ETOPO1	1 arcmin	Amante and Eakins (2009)
global	topographic gradient ^a	derived from ETOPO1	1 arcmin	Seton et al. (2020)
global	seafloor age	EMC-3D2018_08Sv at 70 km depth	1 arcmin	Debayle et al. (2016), Debayle et al. (2019)
global	shear wave tomography	GEM GAF-DB	2 degree	Styron and Paganini (2020)
global	active faults	computed as a fixed-plate reference frame for each major plate	various	Kreemer et al. (2014)
global	GPS velocities			
global	strain rate		0.25 degree	Kreemer et al. (2014)
Azores Microplate	GPS velocities	global model		DeMets et al. (2010)
Adria Microplate	GPS velocities	case study		Breton et al. (2017)
Alaska and northwest Canada	GPS velocities	case study		Elliott and Freymueller (2020)
Danakil Microplate	GPS velocities	case study		McClusky et al. (2010)
Somali Plate	GPS velocities	case study		Saria et al. (2014), Stamps et al. (2021)
Greater Antilles	GPS velocities	case study		DeMets and Wiggins-Grandison (2007), Sun et al. (2020)
Lesser Antilles	magnetics and seismic	case study		Allen et al. (2019)
Coiba and Malpelo Microplates	GPS velocities	case study		Zhang et al. (2017)
New Zealand (North Island)	Seismicity	case study		Shi et al. (2019)
Sierra Nevada Microplate	GPS velocities	case study		Schweickert et al. (2004)
Philippines and East Indonesia	GPS velocities	case study		Zahirovic et al. (2014)
Yakutat Microplate	GPS velocities	case study		Bruhn et al. (2012)

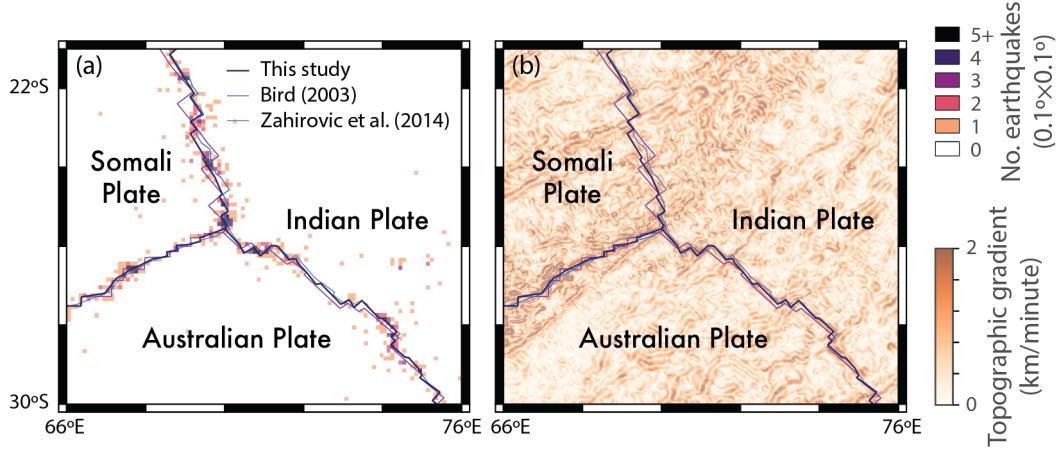


Figure 3: Refining locations of mid-ocean ridges relied heavily on the earthquake catalog, shown as histogram in (a), and topography along with its gradient (b).

and

deformation zone, a zone identified by GPS motions as distinct from, yet controlled by surrounding plates/microplates with significant internal deformation and seismicity.

The microplates and deformation zones are associated with a parent plate for instances where a simpler model is required (Figure 2).

3.1.2. Plate Boundary Type

Plate boundaries are frequently formed by a series of several fault zones that accommodate a portion of the plate motion rather than a single discrete structure. However, there is often one fault zone that accommodates the majority of the motion, which is chosen as the boundary for the plate and plate boundary shapefiles. This choice means that in some cases, the presently most-seismically active fault may not mark the plate boundary as defined here, as it may not have accumulated the greatest displacement. For some extensional plate boundaries, the deformation is distributed over hundreds of kilometers. In these cases, we chose the boundaries by the major structures that bound the extension.

There are several plate boundary types defined in our model:

subduction zone, convergent plate boundaries, kinematically active footwall, plate1 field is the upper plate;

thrust, compression-dominated plate boundaries, kinematically active hanging wall, plate1 field is the upper plate;

180 **ridge**, divergent plate boundaries, type reserved for mid-ocean spreading centers (includes
181 transform segments);

182 **extensional zone**, extension dominated plate boundaries, plate1 field is the upper plate;

183 **dextral transform**, right-lateral transform boundary;

184 **sinistral transform**, left-lateral transform boundary; and

185 **inferred**, unknown boundary types, or location.

186 Our definitions are simplified as they do not include transpressional or transtensional styles,
187 which are classified most commonly as transform boundaries, but these may be considered in
188 future versions.

189 To construct the plate boundary shapefile, we converted the plate polygons to lines and
190 removed duplicate lines. We then split the boundaries so that each line segment represents the
191 boundary between two plates, microplates, or deformation zones (Figure 2). Plate attributes
192 including the boundary type are then added and stylized in QGIS. In order to ensure that
193 thrusts and subduction zone boundaries had the barbs displayed on the upper plate, some lines
194 were topologically reversed in direction.

195 3.1.3. Ocean–Continent Boundary

196 We split the plate polygons into continental and oceanic parts because the ocean–continent
197 boundary is colocated with many plate boundaries. For example, subduction zones generally
198 form the boundary between plates as well as oceanic and continental regions. However, rifts and
199 transform boundaries may form boundaries between plates but rarely form ocean–continental
200 boundaries. Passive margins do not represent plate boundaries but do contain the ocean–
201 continent transition.

202 A single depth contour cannot be used to identify the ocean–continent transition because the
203 depth differs from region to region as a result of variations in sediment thickness and isostatic
204 state. To create the initial model for the ocean-continent boundary, we created a polygon that
205 defines the global distribution from the global seafloor age model (Seton et al., 2020). This
206 initial polygon was then modified to provide a better match to the base of steep topographic
207 gradients (continental slope) computed from ETOPO1 (Figure 4).

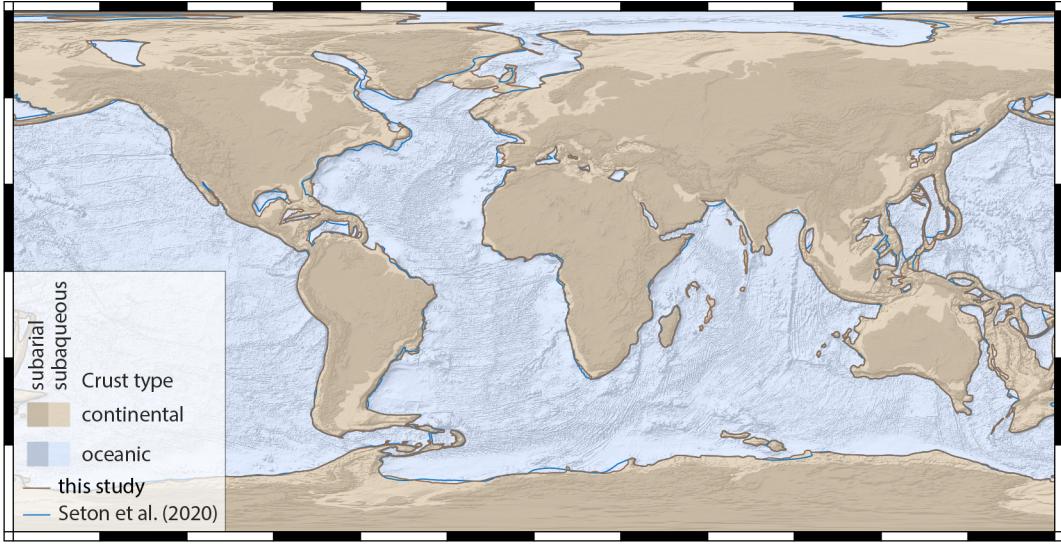


Figure 4: A model of continental and oceanic crustal domains. Plate boundaries as defined in Figure 2.

208 The ocean–continent boundary model could be improved with the addition of crustal thick-
 209 ness and/or seismic velocity estimates across the ocean–continent transition. Presently, global
 210 seismic models are of insufficient resolution to precisely identify the boundary. A compilation
 211 of seismic profiles is beyond the scope of the present work.

212 3.1.4. Oceanic Domain

213 There appear to be differences in chemistry between ocean basins (domains). For example,
 214 Brandl et al. (2013) documents a difference in basaltic geochemistry between the Atlantic and
 215 Pacific Oceans, possibly related to temperatures of melt generation. Back-arc basins, separated
 216 from the major oceans by continental ribbons behind subduction zones, tend to contain enriched
 217 basalts compared with mid-ocean ridge basalts (Langmuir et al., 2006). Thus, it is desirable to
 218 have a way to quickly divide geoscientific data into these separate domains.

219 For the oceanic-type crustal polygons a mantle chemical domain is included in the plates
 220 shapefile attributes table (Table 1). We have separated the oceans into nine separate domains
 221 (Figure 5). The domains are intended to make separation of data from the different ocean
 222 basins easier for geochemical and geophysical studies. However, the divisions are speculative
 223 rather than data driven, specifically the exact boundaries of these chemical domains. The only
 224 boundary that has been tested geochemically lies between the Indian and Pacific oceanic mantle
 225 (Pyle et al., 1992). This study used isotopic analyses of Sr, Nd, and Pb to place the Indian–
 226 Pacific boundary of Australia along the mid-ocean ridge at approximately 126°E. Similarly,
 227 there is also evidence for a more complex chemistry in the seafloor basalts of the Philippine

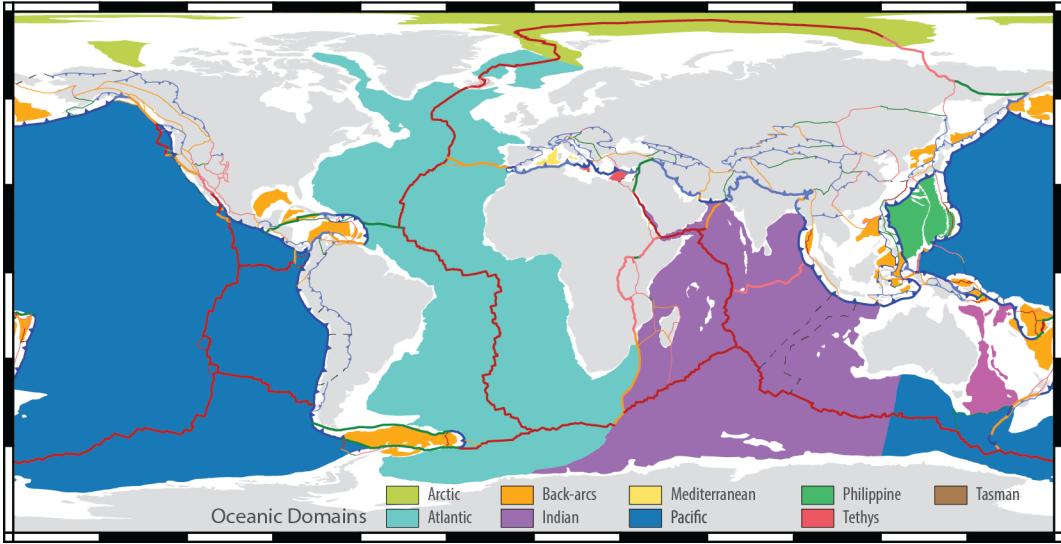


Figure 5: A hypothetical model of mantle chemical domains that source oceanic crust. At present these models are largely untested, but will be refined in future versions using global geochemical datasets.

228 plate compared to the rest of the Pacific ([Hickey-Vargas et al., 2006](#)).

229 *3.2. Global Geologic Province Model*

230 The initial province model was produced by creating a collage of overlapping geological
 231 maps from the published literature that varied from the regional to continental scale (Table
 232 [4](#)). Many of the images were georeferenced in GIS software and vectorized using polygon tools,
 233 while others were used as a visual reference. Province names are assigned based on commonly
 234 published terms, generally taken from the reference maps themselves (Table [4](#)).

Table 3: Geological and geophysical models used to develop province model.

Region	Data type	Description	Resolution	Reference
global	topography	ETOPO1	1 arcmin	Anante and Eakins (2009)
global	composite gravity	GGM2012	2 arc-min	Balmino et al. (2011)
global	composite magnetics	EMAG2-V3	2 arc-min	Meyer and Saltus (2016)
global	composite magnetics	WDMAMv2	5 km	Lesur et al. (2016)
global	seafloor age		2 arc-min	Seton et al. (2020)
global	volcanic centers			Global Volcanism Program (2013)
global	crustal thickness	Szwilus et al. (2019)	1 degree	
global	active faults	Styron and Pagani (2020)	various	
global	digital lithology	Hartmann and Moosdorf (2012b), Hartmann and Moosdorf (2012a)	0.5 degrees	
global	igneous dates	Gard et al. (2019a)	global geochemical database	Brown and Johnson (2018), updated
global	metamorphic dates	P-T database	DateView	Eglington (2004)
global	metamorphic dates	USGS WEP ^a	1:5,000,000	Persits et al. (1997a)
Africa	surface geology and geologic provinces	USGS WEP ^a	1:2,000,000	Pollastro et al. (1999)
Arabian Peninsula	surface geology and geologic provinces	USGS WEP ^a	1:5,000,000	Wandrey and Law (1998)
South Asia (India)	surface geology and geologic provinces	USGS WEP ^a	1:7,500,000	Schenk et al. (1999)
South America	surface geology and geologic provinces	USGS WEP ^a	1:5,000,000	Gómez-Tapias et al. (2019)
South America	surface geology	CGMW ^b	1:7,500,000	Persits et al. (1997b)
Former Soviet Union	surface geology and geologic provinces	USGS WEP ^a	1:7,500,000	Steinshöner et al. (1999)
China, Southeast Asia, and Australia	surface geology and geologic provinces	USGS WEP ^a	1:2,500,000	Pollastro et al. (1997)
Iran	surface geology and geologic provinces	USGS WEP ^a	1:3,000,000	Pawlewicz et al. (1997)
Europe	surface geology and geologic provinces	USGS	1:7,500,000	Garry and Soller (2009)
North America	surface geology provinces		250 m	Australia (2004)
Australia	magnetic anomaly		830 m	Wynne and Bacchin (2009)
Australia	gravity anomaly			

Table 3: continued.

Region	Data type	Description	Resolution	Reference
Antarctica	bedrock topography	BEDMAP2	1 km	Fretwell et al. (2013)
Antarctica	free-air and Bouguer gravity anomalies	ANTGG	1.0 km	Scheinert et al. (2016)
Antarctica	shear wave tomography	AN1	4 arcmin	An et al. (2015)
Antarctica	magnetic anomaly	ADMAP2	7 km	Golynsky et al. (2018)
Antarctica	mantle gravity anomaly		5 km	Baranov et al. (2017)
Antarctica	crustal thickness	gravity and seismic based	1 degree	Baranov et al. (2017)
Antarctica	crustal thickness	satellite gravity based	0.25 degrees	Llubes et al. (2018)
Antarctica	crustal thickness	seismic methods	4 arcmin	An et al. (2015)
India	Bouguer gravity map		3 arc-min	Geological Survey of India (2006)

^aUnited States Geological Survey (USGS) World Energy Project (WEP).

^bCommission for the Geological Map of the World (CGMW)

Table 4: Published province models used to construct the initial global model.

Region	Reference
	<i>Global</i>
large igneous provinces	Johansson et al. (2018)
modern passive margins	Berndt et al. (2019)
	<i>Africa</i>
Africa	Begg et al. (2009), Hinsbergen et al. (2011)
West Africa	Ennih and Liégeois (2008)
Sahara Metacraton	Liégeois et al. (2013); Şengör et al. (2020)
Mozambique Belt	Chaúque et al. (2019), Goscombe et al. (2020)
Central Africa	Jelsma et al. (2018)
Southern Africa	McCourt et al. (2013), Hanson (2003)
Madagascar	Collins et al. (2003)
	<i>Antarctica</i>
Antarctica	Harley et al. (2013), Stål et al. (2019)
East Antarctica	Golynsky (2007), Harley and Kelly (2007), Elliot et al. (2015), Leitchenkov et al. (2016), Pierce et al. (2014), Pant and Dasgupta (2017), Mulder et al. (2019), Ruppel et al. (2020), Flowerdew et al. (2013), Aitken et al. (2014), Maritati et al. (2016), Maritati et al. (2019), Wang et al. (2020), Ebbing et al. (2021), Jacobs et al. (2015), Dunkley et al. (2020)
West Antarctica	Jordan et al. (2020)
	<i>Asia</i>
Siberian Craton	Tretiakova et al. (2017)
West Siberian Basin	Cherepanova et al. (2013)
Russia, far east	Isbell et al. (2016)
Central Asian Orogenic Belt	Xiao et al. (2015), Janoušek et al. (2018), Ivanov et al. (2014), Windley et al. (2007), Buslov et al. (2001)
North China Craton	Liu et al. (2017)
South China Craton	Wang et al. (2013)
Tian Shan Belts	Charvet et al. (2011)
Tibetan plateau	Wei Zhou and Murphy (2005), Blayney et al. (2019)
southeast Asia	Mitchell et al. (2012), Burrett et al. (2014), Zhang et al. (2019), Morley and Searle (2017), Dew et al. (2021)
	<i>Australia and Zealandia</i>
Australia	Foster and Goscombe (2013), Pilia et al. (2015), Abdullah and Rosenbaum (2018)
Zealandia	Stagg et al. (2002), Gallais et al. (2019)
New Zealand	Mortimer (2004)
	<i>Europe</i>
Baltic Shield	Bogdanova et al. (2015), Zhao et al. (2002)
Mediterranean Europe	Schmid et al. (2020)
western Europe	Topuz et al. (2020)
	<i>India and Middle East</i>

Table 4: continued.

Region	Reference
Arabian-Nubian Shield	Johnson (2014)
Iran	Naimi-Ghassabian et al. (2018)
Pakistan	Kazmi and Rana (1982)
India	French et al. (2008)
Sri Lanka	Cooray (1994)
<i>North America</i>	
United States and Canada	Whitmeyer and Karlstrom (2007) , Hasterok and Chapman (2007) , Lund et al. (2015) , Ontario Geological Survey (2011) , Berman et al. (2013b) , Fyffe et al. (2012) , Bjorkman (2017) , Linde et al. (2017)
Greenland	White et al. (2016)
Alaska and Canadian Cordillera	Colpron and Nelson (2011)
Mexico	Sedlock et al. (1993)
Caribbean and Gulf of Mexico	Allen et al. (2019); Davison et al. (2020)
<i>South America</i>	
South America	Ibañez-Mejia et al. (2011) , Chew et al. (2008) , Egydio-Silva et al. (2018)
western South America	Ramos and Aleman (2000) , Eude et al. (2015) , Charrier et al. (2014)

235 Each boundary was then adjusted for seamless fits in the global model using a combination
 236 of geologic maps, active fault databases, geochronology, topography, and geophysical anomalies/
 237 models such as gravity, magnetics, and crustal thickness (Table 3). Magnetic anomalies
 238 were the most useful geological dataset for developing the province model as the magnitude
 239 and visual character is often distinctive within a province (Figure 6). We used two global mag-
 240 netic anomaly models to aid with locating province boundaries. Both are constructed from a
 241 combination of airborne and shiptrack magnetics. Because the airborne and ship track data are
 242 recorded at different altitudes and line spacing, this leads to variations in resolution and gaps
 243 filled with very low-resolution satellite observations. Where the resolution is low, it is more
 244 difficult to precisely position the province boundaries (Figure 6), , requiring more emphasis be
 245 placed on the other datasets listed in Table 3.

246 Topography is helpful for identifying the boundaries of many provinces as fault and shear
 247 zones are often expressed topographically. Topographic features are most helpful in active ter-
 248 ranes, but many ancient terranes can also be delineated by the changes in morphology when
 249 the faults/shear zones are no longer immediately apparent. The relatively high resolution of
 250 ETOPO1 (~ 1.85 km at the equator) makes the positioning of boundaries reasonably precise.

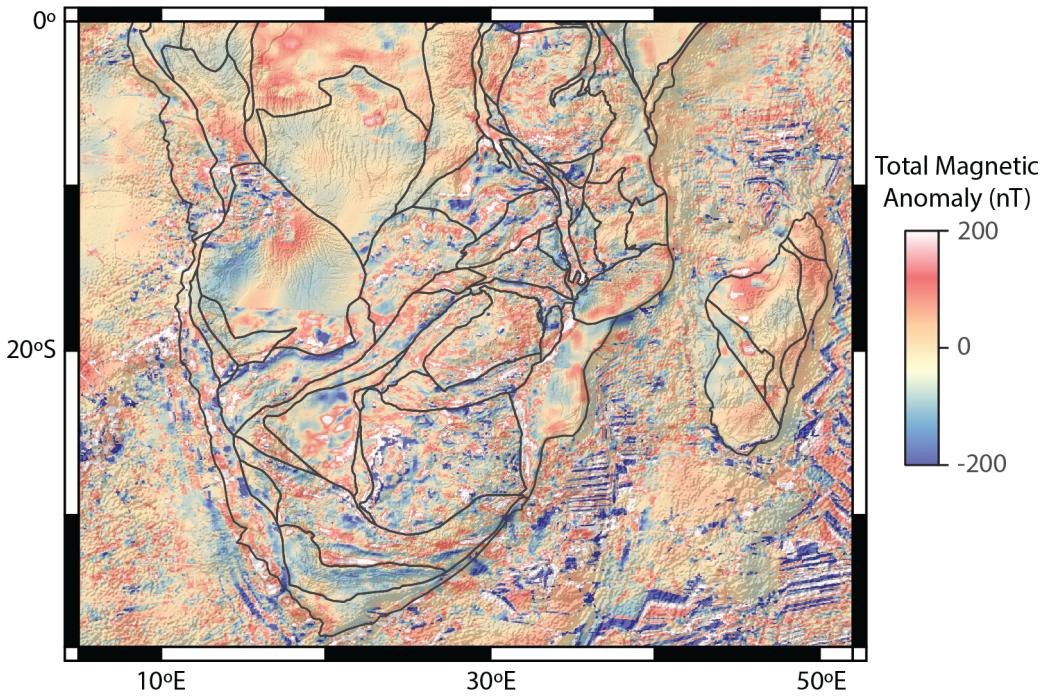


Figure 6: Magnetic anomaly map of southern Africa overlaying topographic relief. The WDMAM_V2 magnetic model has a pixel size of 1.5 minutes, which is approximately 2.8 km at the equator. The model is created from shiptrack, airborne and satellite datasets resulting in variable resolution that is evident in this image and affects the accuracy of province boundaries.

When there is a clear topographic expression at the boundary between provinces, the uncertainties are probably on the order of 3 to 5 pixels (~ 5 to 10 km).

We incorporate a number of attributes ascribed to the province polygons including a province name, province group, tectonic province type, and last significant orogeny (Table 1). The province names are generally The data standards used to define these additional attributes are described below.

3.2.1. Province Type

A geologic terrane captures a set of geologic units that describe a coherent block of crust with a shared geologic history, which can include tectonic setting, magmatic history, and/or metamorphic evolution. While a terrane is less fundamental than a geologic unit or suite, there are often similarities in the physical architecture and chemistry of terranes created in similar tectonic settings (e.g., Ducea et al., 2015; Ueki et al., 2018). However, there are also differences between terranes of similar types that can be uniquely expressed in the architecture, composition, and thermal history (e.g., Furman, 2007; Dilek and Furnes, 2014; Profeta et al., 2015).

To facilitate the analysis of terranes, we include a basic province type attribute with the

shapefile. The province definitions are based on the terrane type that covers the majority of a polygon. The focus of the geologic province model is on basement terranes, hence large igneous provinces (LIPs) are not included in this model as they are superimposed on the basement, and basins have been removed except where the character of the underlying basement is unknown (Figure 7). A good model for LIPs currently exists (Johansson et al., 2018) so there is no need to recreate one as a separate layer as part of this project. A reasonable resolution (5-arc-minute) sediment thickness model by Straume et al. (2019) covers the ocean basins. Australia is covered by a high-resolution, 15-arc-second basin model (Geognostics, 2021). A 30-arc-second sediment thickness model by Pelletier et al. (2016) covers all continents except Antarctica, but only provides values for regions with less than 50 m of sediment thickness. The only available global basin thickness model has a relatively low resolution of $1^\circ \times 1^\circ$ (Laske and Masters, 1997).

Most of the terrane definitions we have chosen distinct characteristics within a modern plate tectonic setting:

craton, predominantly Archean core, contains granite-greenstone belts and other undifferentiated terranes with relatively small area;

shield, similar to a craton, predominantly Meso- to Paleoproterozoic lithosphere, undifferentiated;

passive margin, sediment accumulation built on thinned continental crust facing a tectonically inactive continent–ocean margin;

accretionary complex, active/subduction margin consisting of sedimentary wedges built on oceanic or continental crust;

basin, intracontinental sedimentary cover built on preexisting continental crust with uncertain or unknown basement provenance;

foredeep basin, thick intracontinental sedimentary basin created during continent-continent collision, basement uncertain;

orogenic belt, fold and thrust belts created during accretionary, collisional and intracontinental settings that may incorporate a variety of preexisting terrane types, often commingled,

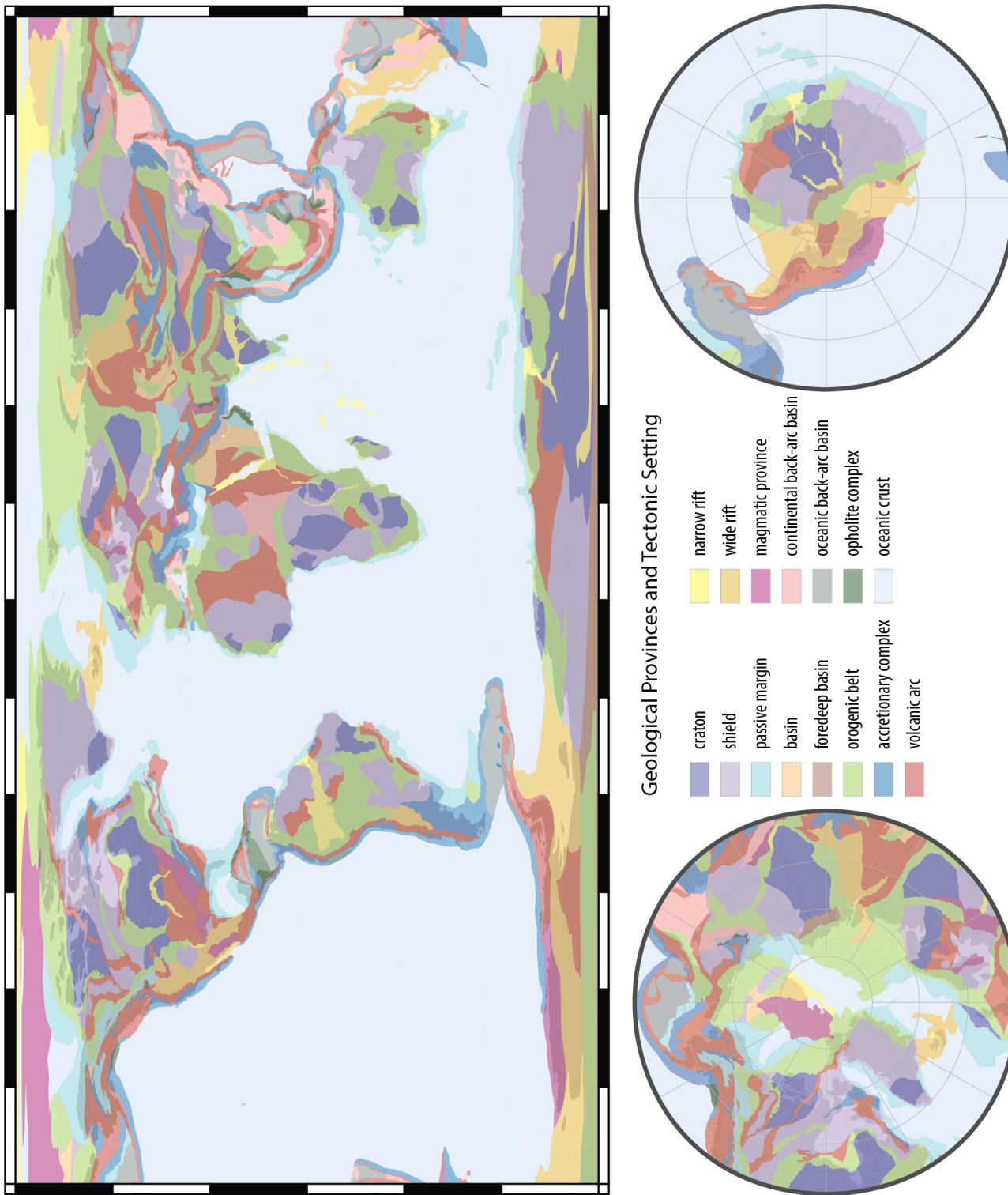


Figure 7: A global model of geological provinces with similar tectonic and compositional histories. Provinces are colored by their dominant tectonic setting. Though the setting may change with time, the provinces are defined by the environment which dominates the majority of rocks found within.

making them difficult to differentiate at the regional scale;

narrow rift, focused extensional terrane with continental basement;

wide rift, distributed extensional terrane with continental basement;

volcanic arc, predominantly magmatic arc crust related to subduction, but may contain crust predating the arc and/or interspersed accretionary material in island arcs and in seaward migrating arcs due to retreating trenches;

continental back-arc basin, a hyper-extended basin either transitional continental or oceanic crust created as a result of upper plate extension in response to subduction rollback;

oceanic back-arc basin, a back-arc basin where seafloor spreading has been sustained, creating enriched basaltic compositions relative to mid-ocean ridge basalt;

ophiolite complex, obducted oceanic crust of some variety, excluding volcanic arc-type, but including supra-subduction zone oceanic crust;

magmatic province, a large intraplate magmatic terrane not clearly associated with subduction or extension processes; and

oceanic crust, typical oceanic crust not created in a back-arc setting.

3.2.2. Last Orogeny

The most recent (last) high-temperature orogenic event to affect a province often has an enduring influence on the present day thermal and physical state of the lithosphere and therefore, its future potential to deform, metamorphose and melt (e.g., Sandiford et al., 2001; Fossen et al., 2017; Hyndman, 2019). Here we define the last orogeny as the most recent regional high-temperature thermotectonic event, excluding regions that may have experienced plume-related activity. The most recent thermotectonic event is generally correlated with elevated surface heat flow (Luazeau, 2019), and lithospheric buoyancy (Fischer, 2002), depending on its thermal intensity, it may reset high-temperature thermochronometers (Wan et al., 2011).

Orogenic systems often span a few hundred million years and comprise multiple, smaller orogenic events that exhibit significant regional variability (e.g., Ge et al., 2014; de Gromard et al., 2019). For example, orogenic activity may propagate along a system over time as

exhibited by protracted continent–continent collision as in the Alpine Himalayan Belt (Kuhnt et al., 2004; Dilek, 2006; Ustaszewski et al., 2010; Hu et al., 2016; Symeou et al., 2018; An et al., 2021), hence the last orogeny descriptor is not as finely resolved in age as the activity in any given region. Instead, we use the last orogeny term to represent long-lived tectonic/geodynamic systems. These descriptors are often related to consumption of ocean basins ± continental collision; though they can also apply to an intraplate orogeny. In the Phanerozoic, the divisions are generally well-defined, but in the Precambrian, the connection between orogenic systems and now-isolated provinces may be less certain (e.g., Li et al., 2008). In these cases, the names refer to periods of orogenic activity rather than discrete systems. While this somewhat blurs the meaning of the term, we prefer it over several colloquial orogenic names. It also represents a research opportunity for improving models of orogenies and more accurately capture multiple distinct systems that may overlap in age.

The last orogeny model is built from reviews and large-scale studies of orogenic systems and plate reconstructions, assigning a single orogeny to each province polygon (Figure 8). We then validate the orogeny model against databases of igneous and metamorphic dates. Recently, Condie et al. (2021) attempted to quantify orogens and link them to other global processes by using the ‘number of orogens’ as a measurable quantity. We suggest that this is too arbitrary a parameter and instead we have attempted an ‘orogenic province’ approach where we suggest linking orogens and orogenies based on their interpreted tectonic/geodynamic system. To take the modern Earth as an example, this approach then links the active circum-Pacific orogens into one orogenic province, but separates them from the Alpine-Himalayan system. Understandably, this gets more subjective in deep time, but it provides a framework for building geodynamic models and presents hypotheses that can be robustly tested by new observations.

In the Phanerozoic, the divisions are generally well-defined, but in the Precambrian, the connection between orogenic systems and now-isolated provinces is often less certain (e.g. Li et al., 2008). We have used the emerging full-plate tectonic reconstructions of Merdith et al. (2021) and Cao et al. (in prep.) to guide us here (Figure 9). While in many cases we believe these were coherent systems, the names may refer to periods of orogenic activity rather than discrete systems (e.g., Siberian Orogeny as defined below). While this somewhat blurs the meaning of the term, we prefer it over several discrete orogenies with colloquial names. It

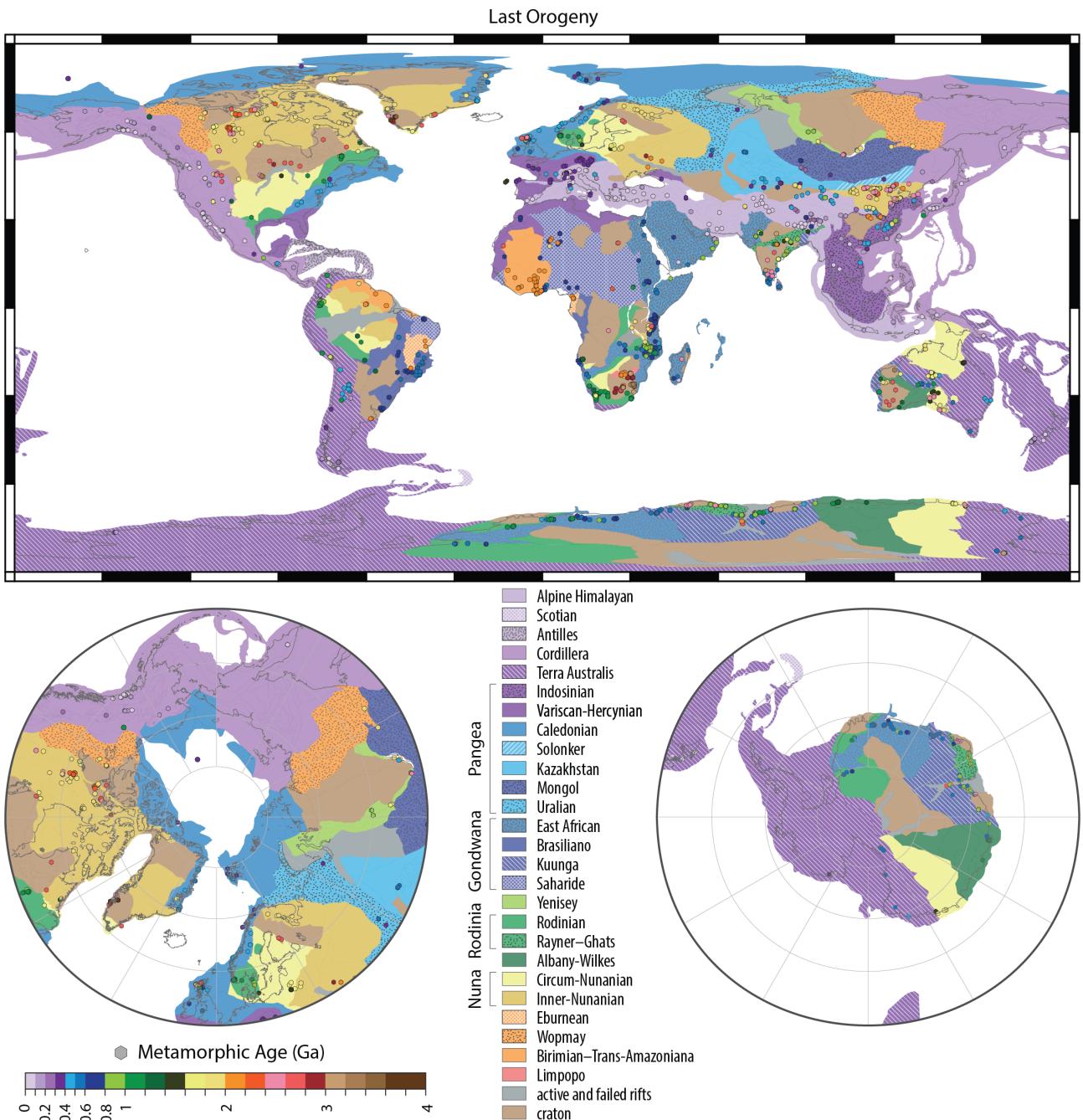


Figure 8: A map of the last orogenic event to affect a province. The colors are approximately related to the age of the orogenic event. The points show locations of observed and estimated metamorphic ages (Brown and Johnson (2018), updated; Eglington (2004)). Classification of the Phanerozoic orogens follows a systems approach as discussed in the main text.

352 also represents a research opportunity for improving models of orogenies and more accurately
353 capture multiple distinct systems that may overlap in age. We define each of the orogens below.

354 *3.2.2.1. Active Orogen*

355 . The Alpine-Himalayan Orogeny includes the collision between several plates with Eurasia,
356 which began ca. 65 Ma. The orogeny was initially driven by the subduction of the Tethys
357 Ocean beneath Eurasia but has continued even as the ocean has closed in many regions. The
358 continent–continent collisions with Eurasia were heterogeneous in time, beginning with the
359 collision of Apulia with Europe ca. 65 Ma ([Ustaszewski et al., 2010](#)), India with Tibet ca.
360 61–59 Ma (earliest suggested timing; [Hu et al., 2016](#); [An et al., 2021](#)), Australia with Indonesia
361 ca. 25 Ma ([Kuhnt et al., 2004](#)), and Arabia with Iran ca. 13 Ma ([Dilek, 2006](#); [Symeou et al.,](#)
362 [2018](#)).

363 The Scotian Orogeny began with the initiation of subduction along the Antarctic and South
364 American Plate margins ca. 80 Ma and continues to the present day beneath the Sandwich
365 Islands ([Eagles, 2016](#); [van de Lagemaat et al., 2021](#)). The Scotian Orogeny (Figure 8 and 9a),
366 like the Antilles Orogeny (below), is geodynamically governed by a retreating subduction zone
367 consuming the southwest Atlantic, which has created the Scotia Plate and Sandwich Microplate
368 in its wake.

369 The Antilles Orogeny is a young (ca. 118 Ma to present), active orogenic system in the
370 Caribbean that began in the mid-Cretaceous and created the Caribbean Plate as a result of
371 rapid trench retreat ([García Casco et al., 2006](#)). The orogen is responsible for the creation
372 of three separate arc systems: the Greater Antilles, Lesser Antilles and Aves Ridge (Figure 8
373 [García Casco et al., 2006](#); [Neill et al., 2011](#); [Allen et al., 2019](#)).

374 The Pacific Orogeny is defined by a set of circum-Pacific subduction zones associated with
375 the destruction of the Pacific, Philippine Sea, Cocos and Nazca Plates beneath the South
376 American, North American, Eurasian, and Australian Plates (Figure 9a to c). In the modern
377 lexicon, this system is often referred to as the Cordilleran Orogeny ([Dickinson, 2004](#)); however,
378 this is too simplistic as the Cordilleran Orogeny either merged with or grew out of the Terra
379 Australis Orogeny ([Muttoni et al., 2003](#)), which is defined by the subduction of the Panthalassic
380 Ocean beneath Gondwana ([Cawood, 2005](#)). We have kept these two orogens separate to preserve
381 the association of Terra Australis with Gondwana, but acknowledge there is little reason to

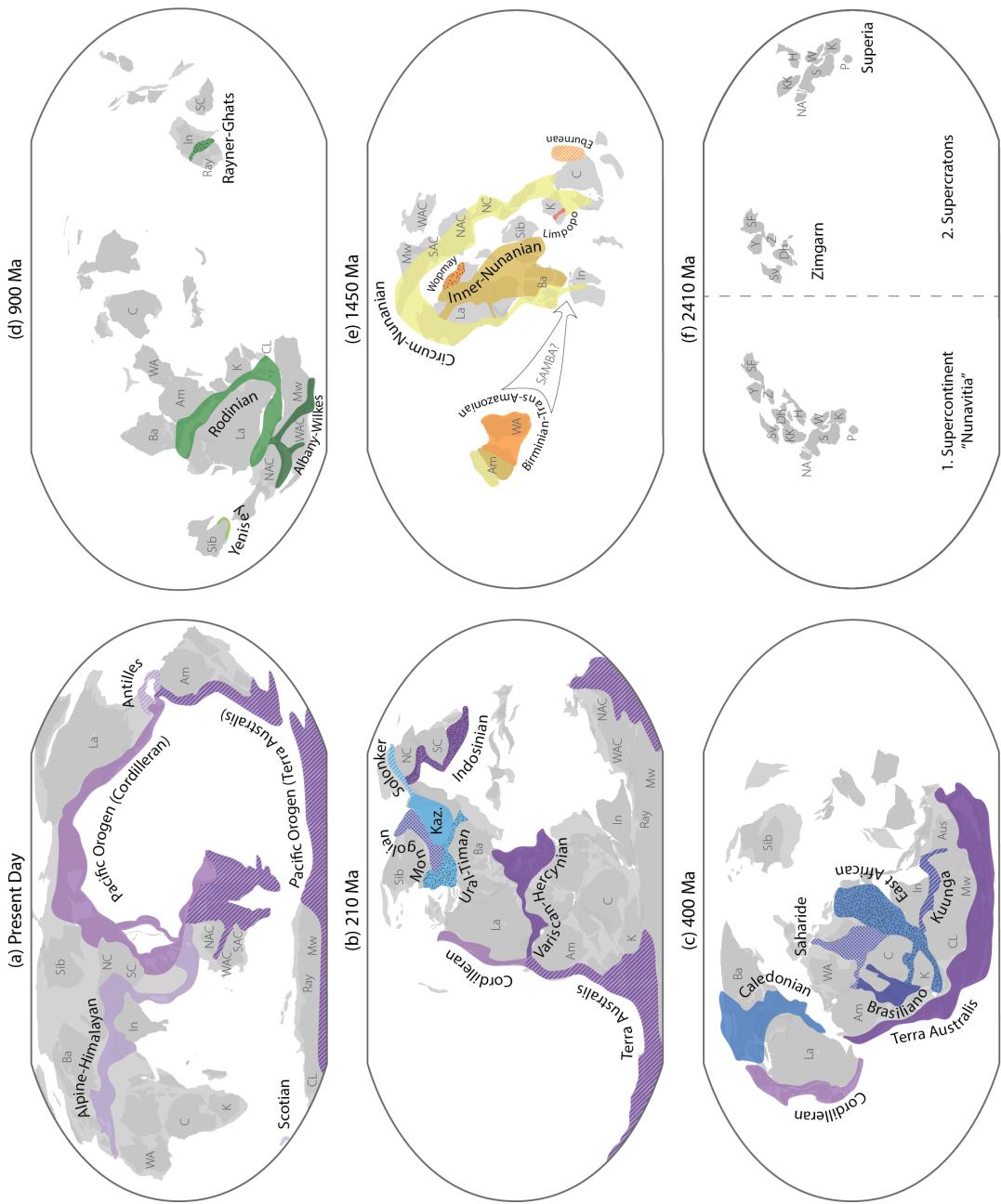


Figure 9: Reconstructions of orogenic systems at key dates. (a) Present-day shows the five active orogens. (b) Post-Pangea assembly, 210 Ma, keeping the position of Antarctica fixed. (c) Post-Gondwana assembly, 400 Ma, keeping the position of Amazonia fixed. (d) Post-Rodinia assembly, keeping Amazonia fixed. (e) Post-Nuna/Columbia assembly, 1450 Ma. Arrow indicates the alternate location of Amazonia and the West African Craton in the SAMBA reconstructions (see Section 3.2.2.3 for discussion). (f) Supercontinent model for Nunavutia at 2410 Ma on the left and supercratons model on the right. Reconstructions (a-d) from Merdith et al. (2021), (e) from Cao et al. (in prep.), and (f) from Li et al. (2021) with terranes that were connected at 2.3 Ga: Grünehogna Craton with the Kaapvaal Craton, Penokean Orogen, North Atlantic Craton and the Nain Province with the Superior Craton; Napier Complex with the East Dharwar Craton; and the Mantiquera Province with the Saô Francisco Craton. Abbreviations are as follows: Ba, Baltica; C, Congo Craton; CL, Coats Land; Dh, Dharwar Craton; H, Hearne Province; In, India; K, Kaapvaal/Kalahari Craton; KK, Kola-Karelia; La, Laurentia; Mw, Mawson Craton; NA, North Atlantic Craton; NAC, North Australian Craton; NC, North China; P, Pilbara Craton; Ray, Rayner Complex; S, Superior; SAC, South Australian Craton; SC, South China; SF, Saô Francisco Craton; Sib, Siberia; Sv, Slave Craton; VU, Volgo-Ukrainian Shield; W, Wyoming; WA, West Africa; WAC, West Australian Craton; Y, Yilgarn Craton; and Z, Zimbabwe Craton.

382 distinguish them geodynamically.

383 The Terra Australis Orogeny began ca. 530 to 520 Ma in response to subduction of the
384 Pacific along the Gondwana margin (Figure 9c; Cawood, 2005; Chew et al., 2007; Paulsen et al.,
385 2020), which continues through to the present day (Glen et al., 2016). The orogen includes a
386 series of alternating subduction-related back-arc extensional and collisional events that built
387 eastern Australia, Zealandia, and much of the Transantarctic Mountains (Fergusson and Hen-
388 derson, 2015). Because the Terra Australis Orogen involved phases of significant extension
389 (Gaina et al., 1998; Abdullah and Rosenbaum, 2018; Jessop et al., 2019), parts of the oro-
390 gen are no longer active and/or have been separated by seafloor spreading (i.e., Delamerian,
391 Thompson and New England orogens of Australia, Patagonia in South America, Ross Orogen
392 in Antarctica, and the Cape Fold Belt in South Africa). Some authors consider parts of the
393 Terra Australis Orogen to be part of the Cordilleran (Ross Orogen in Antarctica and Southern
394 Alpine Orogen in New Zealand, Tagami and Hasebe, 1999; Dickinson, 2004) or the Alpine-
395 Himalayan Belt (Zealandia, Lister et al., 2001). However, we associate these regions with the
396 Terra Australis Orogeny due to (1) their former common Gondwana association and (2) the
397 orogenic system largely predates both the Alpine-Himalayan and Cordilleran Orogenies. In the
398 eastern South Pacific, South America is moving westward over the young oceanic lithosphere
399 (Schepers et al., 2017). The modern Andes are ca. 66 Ma old (Capitanio et al., 2011), reach-
400 ing their current heights ca. 14 Ma (Evenstar et al., 2015); however, some estimates suggest
401 subduction formed a volcanic arc in South America by 530 Ma (Chew et al., 2007).

402 The Cordilleran Orogeny from Canada to Mexico is long-lived, starting in the late Devonian
403 (ca. 370 Ma) with the collision of the Antler Orogeny and subsequently the Sonoma Orogeny in
404 the Triassic along the western North American margin (Figure 9c; Dickinson, 2004). At the end
405 of the Sonoma Orogeny, a continental magmatic arc system extended along the western North
406 American boundary (Dickinson, 2004), some of the remnants of which are still active today in
407 the Cascades, Trans-Mexican Volcanic Belt and Middle America Arc. Subduction has all but
408 ceased along the Canada to Mexico margin as a transform margin developed ca. 50 Ma (Queen
409 Charlotte Fault predecessor, Rusmore et al., 2010) and ca. 28 Ma (San Andreas Fault, Atwater
410 and Stock, 1998) leading to gravitationally relaxation of the orogen (Liu and Shen, 1998). In the
411 western Pacific, subduction zones are generally in retreat, causing significant upper-plate exten-

412 sion in the back-arcs (Vaes et al., 2019). We also include the Verkhoyansk-Kolyma Orogeny of
413 far east Russia in our Cordilleran classification (Figure 8), which experienced intracontinental
414 deformation during the mid-Cretaceous in response to compressional forces applied by relative
415 Siberian and Alaskan convergence, but is no longer active (Oxman, 2003; Filatova and Khain,
416 2008). The eastern half of the North China Craton also experienced widespread volcanism in
417 the Mesozoic and Cenozoic related to Pacific subduction (Wu et al., 2019). While there is no
418 metamorphic evidence for resetting of high-temperature thermochronometers, there is ample
419 evidence for significant modification of the lithosphere (Kusky et al., 2007; Yang et al., 2018;
420 Li et al., 2019; Dong et al., 2021).

421 *3.2.2.2. Neoproterozoic to Mesozoic Orogens*

422 . The Indosinian Orogeny, ca. 310 to 200 Ma, resulted from the closure of the Paleo-Tethys
423 and Paleo-Pacific oceans in the late Paleozoic and early Mesozoic and led to the formation of
424 much of East Asia (Lepvrier et al., 2004; Morley et al., 2013; Arboit et al., 2016; Gao et al.,
425 2017; Dew et al., 2021). Specifically, the orogeny involves the collision of South China with
426 North China, Indochina with South China, and the Sibumasu Terrane with Indochina and the
427 intervening Sukhothai arc terrane (Arboit et al., 2016; Gao et al., 2017; Dew et al., 2021). It
428 is responsible for the accretion and amalgamation of much of East Asia (Figure 9b).

429 The Variscan-Hercynian Orogeny is defined by the closure of the Rheic Ocean ca. 290 Ma
430 (Matte, 2001; Nance et al., 2012). The orogeny spans the period from ca. 360 to 280 Ma (Edel
431 et al., 2014; Žák et al., 2014) and extended from north and western Mexico to Florida in North
432 America, the Iberian Peninsula to the Tornquist zone in Europe and included parts of north
433 and west Africa (Figure 9b; Catalán et al., 2021). The orogeny concluded with the collision of
434 Gondwana with the Carolina Arcs, Meguma, Amorica and Avalonian terranes with Laurussia
435 (Stampfli et al., 2013). In Europe, the orogeny involved several additional microcontinents
436 that comprise western Europe (i.e., Iberia and Cadomia). Further east, the orogen merges into
437 the early Alpine-Himalayan Orogen and the late Central Asian Orogenic Belt and Indosinian
438 orogens with the closure of the various strands of the Paleotethys and Paleo-Asian Oceans
439 (Sengör and Natal'in, 1996; Robertson et al., 2004; Xiao et al., 2015; Gardiner et al., 2016).

440 The Caledonian Orogeny comprises a series of orogenic events spanning the period ca. 540 to
441 350 Ma resulting from the closure of the Iapetus Ocean (Figure 9c; McKerrow et al., 2000; Weller

442 et al., 2021). The orogen is typically associated with deformation from northeastern Greenland,
443 Svalbard and western Baltica, extending through the British Isles and continuing south in the
444 Appalachian Mountains (Weller et al., 2021). In eastern North America and Baltica, this
445 orogeny ended with the collision of Avalonia with Laurussia (Nance et al., 2012). We have
446 also included the Arctic Innuitian (Ellesmerian) Orogeny in our Caledonian Orogeny definition
447 due to its proximity in space and time to the Caledonian system sensu stricto (Barnes et al.,
448 2020), though the Innuitian occurred near the end of the Caledonian and is nearly orthogonal
449 in strike. While portions of the Innuitian Orogeny were overprinted in the early Tertiary by
450 the Eurekan Orogeny, this intracontinental deformation was relatively minor and is difficult to
451 distinguish from the earlier Innuitian Orogeny (Gion et al., 2017, and references therein).

452 The Central Asian Orogenic Belt (CAOB; Windley et al., 2007) has been divided into the
453 three separate orogenic systems following the model presented by Xiao et al. (2015), which
454 involves subduction and eventual closure of distinct regions of the Paleo-Asian Ocean from the
455 Tonian until the Triassic. The larger system is divided into (1) the Kazakhstan tectonic collage
456 and orocline; (2) the Mongolia tectonic collage; and (3) the Tarim-North China system, which
457 we refer to as the Solonker Orogeny (Figures 8 and 9b), that overlaps in age with the more
458 southern Indosinian Orogeny (Xiao et al., 2015; Song et al., 2018). Accretion of the Kazakhstan
459 and Mongolian tectonic collages continued through the Neoproterozoic and Paleozoic and were
460 terminated by the Solonker Orogeny at ca. 270 to 235 Ma that marks the final closure of
461 the Paleo-Asian Ocean (Eizenhöfer et al., 2014; Song et al., 2018, 2021). The CAOB was
462 subsequently reactivated during the Meso-Cenozoic in response to distant events related to the
463 progressive consumption of the Tethys and Mongol-Okhotsk Oceans (e.g., Glorie and Grave,
464 2016). The Bureja-Jziamusy Terrane also experienced deformation as part of the Solonker
465 Orogen, but was subsequently overprinted by Cordilleran deformation as recently as ca. 95 to
466 90 Ma (Derbeko, 2013).

467 The Uralian-Timan Orogeny is the result of several subduction-related magmatic periods
468 spanning the period ca. 610 to 250 Ma (Figure 9b Fershtater, 2012; Pease, 2021). Oceanic arc
469 volcanics associated with the subduction of the Paleo-Asian orogen are recorded in the Pechora
470 arc (ca. 560 Dovzhikova et al., 2004) and blueschists and eclogites in the Urals (ca. 530 Willner
471 et al., 2019). The latter set of magmatic events is related to the subduction and closure of the

⁴⁷² Paleo-Asian ocean and collision with the Kazakhstan Orocline (Xiao et al., 2015). The Uralian
⁴⁷³ Orogeny extends into the Taimyr Fold belt where compression continued to 220 Ma, after the
⁴⁷⁴ orogen had ceased elsewhere (Torsvik and Andersen, 2002).

⁴⁷⁵ The Neoproterozoic to Cambrian Gondwana-forming orogenies are a set of generally con-
⁴⁷⁶ temporaneous orogens—though they span nearly 700 Ma (1200 to 500 Ma)—culminating in
⁴⁷⁷ the amalgamation of Gondwana (Figure 9c; Collins and Pisarevsky, 2005; Meert and Lieber-
⁴⁷⁸ man, 2008; da Silva Schmitt et al., 2018; Goscombe et al., 2020; Sengör et al., 2020; Collins
⁴⁷⁹ et al., 2021b). We have separated these orogenies into the East African, Saharide, Kuunga, and
⁴⁸⁰ Brasiliano orogenies in much the same way as the Central Asian Orogenic belt was separated
⁴⁸¹ into distinct systems. The East African Orogen (Stern, 1994; Collins and Windley, 2002; John-
⁴⁸² son et al., 2011; Fritz et al., 2013; Collins et al., 2021b) runs through Arabia, eastern Africa,
⁴⁸³ Madagascar, southern India and into the Lützow-Holm Bay area of Antarctica. Whereas, we
⁴⁸⁴ have located the Kuunga Orogen (Meert and Voo, 1997) as being the orogen that separated
⁴⁸⁵ Neoproterozoic India from Australia/Mawson (Collins and Pisarevsky, 2005), running through
⁴⁸⁶ NE India, SW Australia and into Antarctica. This is broadly the trace of the Pinjarra Orogen
⁴⁸⁷ of Fitzsimons (2003) and Prydz-Denman-Darling Orogen of Collins and Pisarevsky (2005), but
⁴⁸⁸ may include the Mirny Fault (Daczko et al., 2018) and Gamburtsev suture (Ferraccioli et al.,
⁴⁸⁹ 2011) as Neoproterozoic plate boundaries between Indo-Antarctica and Australo-Antarctica
⁴⁹⁰ (Mulder et al., 2019). Sengör et al. (2020) recently suggested the Tuareg Shield, Arabian-
⁴⁹¹ Nubian Shield, and portions of the Saharan Metacraton constitute a single volcanic arc system
⁴⁹² that was segmented and recombined in a fashion very similar to the Kazakhstan and Mongolian
⁴⁹³ oroclines. We tentatively accept this model for the Tuareg Shield and Saharan Metacraton,
⁴⁹⁴ but include the younger Arabian Nubian Shield in the East African Orogen (Figure 8), whilst
⁴⁹⁵ appreciating a likely continuity of orogenesis from one to the other (Blades et al., 2021; Collins
⁴⁹⁶ et al., 2021a). Our use of Brasiliano Orogen encompasses all the South American Gondwana-
⁴⁹⁷ forming orogens, as well as orogens along the west coast of Africa that correlate with them
⁴⁹⁸ (including the West Congo Orogen, the Rokelides, the Gariep Belt and the Kaoko Belt).

⁴⁹⁹ In the early Neoproterozoic, ca. 880 to 500 Ma, a small subduction-related orogen occurred
⁵⁰⁰ along the present-day eastern and southern margin of the Siberian Craton (Vernikovsky et al.,
⁵⁰¹ 2003; Kuzmichev and Sklyarov, 2016), which we refer to as the Yenisey Orogeny (Figures 8 and

502 9). The orogeny coincided with the accretion of the Angara terrane at ca. 870 Ma, (Vernikovsky
503 et al., 2007; Gladkochub et al., 2010). Numerous A-type magmatic dates, ca. 880 to 720 Ma,
504 are interpreted as part of a back-arc basin system (Kozlov et al., 2012; Kuzmichev and Sklyarov,
505 2016) and is consistent with metamorphism recorded during this interval (Gladkochub et al.,
506 2010). The system transitioned to seafloor spreading creating the Isakovka Terrane, an arc
507 ophiolite, ca. 700 to 635 Ma (Vernikovsky et al., 2003; Kuzmichev and Sklyarov, 2016), which
508 later accreted to the continent associated with a 500 to 470 Ma high-temperature metamorphic
509 event in the middle of the orogen (Gladkochub et al., 2010, and references therein).

510 *3.2.2.3. Paleo- to Mesoproterozoic Orogens*

511 . Today, the Mesoproterozoic and Paleoproterozoic orogens are fragmented and scattered across
512 multiple continents (Figure 8), and in many places, reworked by more recent events (Phillips
513 et al., 2009). This dispersion and tectonic overprinting obscures the orogenic systems with
514 time and makes it more difficult to associate terranes with individual orogens. As a result, we
515 recognise that our orogenic-systems approach becomes more subjective. To retain an orogenic-
516 systems approach as much as possible, we have used paleomagnetic-based (e.g., Condie et al.,
517 2021) and full-plate tectonic (Merdith et al., 2021; Cao et al., in prep.) reconstructions (Fig-
518 ure 9d to f), while recognising an inevitable shift to a more temporal-based scheme for the
519 pre-Neoproterozoic.

520 The late Mesoproterozoic to Tonian orogenies include the orogens that assembled Rodinia
521 (Figure 9d), which are now widely dispersed across the globe (Figure 8; Li et al., 2008). Many
522 studies refer to orogenesis during the period 1.3 to 0.9 Ga as Grenvillian-aged (e.g., Tohver
523 et al., 2006; Sheppard et al., 2007; Goodge et al., 2010; Chattopadhyay et al., 2015). However,
524 the conflation of orogen names to mean stretches of time has caused considerable confusion. For
525 example, the term ‘Pan-African’ has been used to mean any orogen that occurred between ca.
526 800 and 400 Ma (Kröner, 1980), whereas orogens of this age form a number of discrete orogenic
527 systems (see discussion of Gondwana above). Similarly with the term ‘Grenvillian’ (e.g., Kra-
528 nendonk and Kirkland, 2013). Fitzsimons (2000) pointed out that the late Mesoproterozoic to
529 Tonian orogens that appear to surround Antarctica, in fact, fall into discrete time brackets that
530 relate to three different orogenic systems. Using the reconstruction by Merdith et al. (2021) as
531 a guide, we have separated the orogens of this period into the Rodinian, Rayner-Ghats, and

532 Albany-Wilkes Orogenies as separate systems active between 1.3 to 0.9 Ga.
533 The Rayner-Ghats Orogeny, ca. 1.1-0.9 Ga, has been interpreted as a distinct orogen outside
534 of Rodinia in recent global plate models (Figure 9d; [Merdith et al., 2021](#)). This orogen includes
535 the Rayner Complex in Antarctica ([Halpin et al., 2013](#); [Liu et al., 2014](#); [Morrissey et al., 2015](#)),
536 conjugate terranes in India (the Eastern Ghats; [Korhonen et al., 2011](#)) and the Central Indian
537 Tectonic Zone ([Bhowmik, 2019](#)).

538 The 1.3 to 1.0 Ga Rodinian Orogeny (Figures 8 and 9d) includes the Grenville and Llano
539 provinces in North America ([Whitmeyer and Karlstrom, 2007](#); [Johansson et al., 2022](#)), which,
540 during Rodinia assembly, we link with the 1.1 to 1.0 Ga Namaqua-Natal Belt in southern Africa,
541 the Maud Belt and Coats Land Block in Antarctica, and possibly the eastern South Tasman
542 Rise ([Mulder et al., 2018](#)). We also include the ca. 1.1 to 0.9 Ga Laurentia-Australia transform
543 as part of the Rodinian Orogeny that [Mulder et al. \(2018\)](#) connects to Rodinian subduction
544 zones. In South America, the Rondonia-Juruena Province and its continuation underneath the
545 Llanos Basin had a long-lived history of deformation spanning ca. 1.32 to 0.96 Ga ([Tohver
546 et al., 2006](#)) with late magmatic activity in the Sunsás Orogen (1.17 to 1.08 Ga; [Santos et al.,
547 2008](#); [Nedel et al., 2020](#); [Johansson et al., 2022](#)).

548 The Albany-Wilkes Orogeny, ca. 1.38 to 1.13 Ga, resulted from the collision of the West
549 Australian Craton with the North and South Australian Cratons and the Mawson Craton of
550 Antarctica (Figure 9d; [Maritati et al., 2019](#); [Pawley et al., 2020](#)). [Mulder et al. \(2018\)](#) suggests
551 that the Albany-Wilkes Orogen is likely a continuation of the Grenville-Maud system (Rodinian
552 in our lexicon). However, the Albany-Fraser and Wilkes Orogenies may have started on the
553 same broad margin of Nuna/Columbia during break-up through to the assembly of Rodinia
554 ([Pisarevsky et al., 2003, 2014](#); [Yang et al., 2020](#); [Kirscher et al., 2020](#)). Most of the Rodinian
555 deformation occurs on the opposite side of Laurentia from East Antarctica/Australia and starts
556 at a later time; hence we consider the Albany-Wilkes to be a separate system. Recent dating of
557 1.38 to 1.275 Ga metamorphism in the Rudall Province and Western Musgravites, respectively
558 (the Parnngurr and Mount West Orogenies; [Howard et al., 2015](#); [Payne et al., 2021](#)) extends
559 the early Albany-Fraser orogenesis north and documents collision between the West Australian
560 Craton and the combined South and North Australian Cratons. However, in places these were
561 subsequently overprinted during the Miles (ca. 650 to 625 Ma) and Paterson-Petermann (ca.

562 580 to 530 Ma) orogenies and thus related to the Kuunga Orogeny (Figure 9c). Two meta-
563 morphic events have been dated in the Albany Fraser Orogen, ca. 1.345 to 1.260 Ga and 1.215
564 to 1.140 Ga, resulting from the geometry of the collision between the West Australian Cra-
565 tons and the rest of Proterozoic Australia, with post-Parnngurr orogenic rotation and collision
566 of the West and South Australian Cratons (Clark et al., 2000). The Albany-Fraser Orogen,
567 Coompana and Madura provinces of Australia (e.g., Kirkland et al., 2017; Spaggiari et al.,
568 2018; Pawley et al., 2020) have been linked to their Antarctic conjugates from detrital zircon
569 spectra in offshore sediments and onshore geophysical characteristics (Maritati et al., 2019).
570 The Wilkes Orogen records amphibolite facies metamorphism at ca. 1.305 Ga and granulite
571 facies overprinting associated with charnockite intrusions, ca. 1.20 to 1.16 Ga (Morrissey et al.,
572 2017). The system also includes the 1.380 to 1.275 Ga Parnngurr and Mount West Orogenies
573 (Payne et al., 2021), an intracontinental contractional orogeny between the Yilgarn and Pilbara
574 Cratons. Tectonothermal events are also present in the central Australian Arunta Block (ca.
575 1.13 Ga; Scrimgeour et al., 2005; Morrissey et al., 2011; Wong et al., 2015) and in the Western
576 Australian Capricorn Orogen between the Yilgarn and Pilbara cratons (the 1.321–1.171 Ga
577 Mutherbukin Tectonic Event and the 1.026–0.954 Ma Edmudian Orogeny), which occur as
578 intracontinental far-field orogenesis to the amalgamation of Proterozoic Australia. It is unclear
579 how far the orogen extends into Antarctica due to the extensive ice cover.

580 Between ca. 2.1 and 1.45 Ga two major orogenic systems are associated with Nuna (Fig-
581 ure 9e Pisarevsky et al., 2014; Condie et al., 2021; Cao et al., in prep.): the Inner-Nunianian
582 Orogeny (ca. 2.1 to 1.7 Ga), which formed the core of the supercontinent; and the accre-
583 tionary Circum-Nunianian Orogeny (ca. 1.85 to 1.45 Ga) that is driven by a subduction girdle
584 surrounding the core. There are several reconstructions for Nuna (Bispo-Santos et al., 2008;
585 Elming et al., 2009; Johansson, 2009; Zhang et al., 2012; Pisarevsky et al., 2014; D’Agrella-
586 Filho and Cordani, 2016; Meert and Santosh, 2017; Cawood et al., 2020; Elming et al., 2021;
587 Cao et al., in prep.), and while most are sufficiently similar to yield little difference in the last
588 orogeny designation, there are competing models for the likely participation of Amazonia that
589 will affect our model. In the SAMBA models, first proposed by Johansson (2009), Amazonia
590 (Central Amazonian and Ventuari-Tapajós Belts) is contiguous with the Baltic Shield, ca. 2.1
591 to 1.8 Ga, both of which experienced significant intrusive magmatism during the Nunianian Oro-

592 genies (e.g., Almeida et al., 2007; Bogdanova et al., 2015; Juliani et al., 2021). In the alternative
593 configuration by Pisarevsky et al. (2014), Amazonia and the West African Craton form a lesser
594 continent separate from Nuna (Figure 9e Cao et al., in prep.), deforming as part of a separate
595 accretionary margin. We have included the Amazonian Belts as part of the Nunanian Orogenies
596 (Figure 8 and 9e), but acknowledge that they may have evolved as a separate system.

597 The Inner-Nunanian Orogeny, ca. 2.1 to 1.76 Ga, was a major global event related to the
598 closure of the Manikewan Ocean and assembly of the Nuna/Columbia supercontinent (Corrigan
599 et al., 2009; Weller et al., 2021). In North America, the Superior Craton collided with the
600 Reindeer Zone and Sask Craton, in a collision that has been compared to India colliding with
601 southern Asia (St-Onge et al., 2006; Darbyshire et al., 2017; Weller and St-Onge, 2017). On
602 the opposing side of the Reindeer Zone, the core of the orogen, contains a number of terranes,
603 grouped into the Hearne and Rae Cratons, which are sutured together by the Snowbird Tectonic
604 Zone (Thiessen et al., 2018). Deformation extended across to the Taltson-Thelon Arc (ca. 1.87
605 to 1.84 Ga) between the Rae and Slave Cratons (Chacko et al., 2000; Whalen et al., 2018).
606 Baltica also experienced widespread orogenic activity as part of the Inner-Nunanian system that
607 is recorded in igneous and metamorphic activity that affected the Kola Block (e.g., Mikkola
608 et al., 2018; Daly et al., 2001; Tuisku and Huhma, 2006; Makkonen et al., 2020), which at the
609 time, the present-day northern margin of the Kola Peninsula was adjacent to the eastern margin
610 of northern Greenland (Evans and Mitchell, 2011). Several reconstructions place the southern
611 margin of the Siberian Craton against the northern Margin of Laurentia at this time (Condie
612 and Rosen, 1994; Sears and Price, 2003) and Baltica to its east (Evans and Mitchell, 2011).
613 However, more recent models suggest it collided post 1.9 to 1.84 Ga after the Anabar and Aldan
614 terranes had accreted to the eastern margin of the Siberian Craton based on geochronology of
615 mafic dike swarms, and post-collisional granitoids (Donskaya et al., 2009; Ernst et al., 2016).
616 The Inner-Nunanian Orogeny also saw the amalgamation of Volgo-Uralia and Sarmatia, ca.
617 2.1 to 2.0 Ga (Savko et al., 2015; Baltybaev et al., 2017), and the subsequent collision with
618 Baltica at ca. 1.82 to 1.80 Ga (Bogdanova et al., 2015). This collision is distinguished from
619 the Circum-Nunanian belts in Baltica that run orthogonal to the Volgo-Sarmatia collision.
620 This collision is distinguished from the Circum-Nunanian belts in Baltica that run orthogonal
621 to the Volgo-Sarmatia collision. The spatial complexity of the Inner-Nunanian Orogeny may

622 result from multiple systems just as the inner Gondwana orogenies, but uncertainties in the
623 geographic positions of many key crustal elements make it difficult to divide further at present.

624 The next few paragraphs discuss the Circum-Nunanian Orogeny and some of the variations
625 in configurations. Some of these variations are slight while others are quite dramatic, but
626 despite these differences the orogens all appear to occur along the exterior of Nuna's core as
627 terranes were accreted. Thus regardless of the accuracy of the geologic connections between
628 terranes, the last orogeny classification remains the same.

629 Wyoming was likely the first accretionary terrane added to Laurentia during the Circum-
630 Nunanian Orogeny, colliding with the Medicine Hat Terrane causing deformation in the Great
631 Falls Tectonic Zone (ca. 1.86 to 1.73 Ga; [Gifford et al., 2018](#)). Also during the first phase
632 of the Circum-Nunanian Orogeny, ca. 1.85 to 1.75 Ga, the Penokean Orogeny was a small
633 deformation event on the southern margin of the Superior craton recorded in metamorphism
634 and accompanying magmatism ([Holm et al., 2007](#); [Vallini et al., 2007](#); [Klier, 2019](#); [Zi et al.,](#)
635 [2021](#)). On the southern margin of Laurentia, a series of exotic terranes, Yavapai (ca., 1.80 to
636 1.70 Ga), Mazatzal (1.70 to 1.65 Ga), and Granite-Rhyolite terranes (1.50 to 1.45 Ga), were
637 accreted over the course of approximately 300 Ma ([Karlstrom et al., 2001](#); [Whitmeyer and](#)
638 [Karlstrom, 2007](#); [Amato et al., 2008](#); [Mako et al., 2015](#)). The Yavapai and Mazatzal terranes
639 include juvenile arcs, ophiolites and metasediments that were accreted ca. 1.71 to 1.68 Ga
640 and 1.646 to 1.633 Ga during the Yavapai and Mazatzal Orogenies, respectively ([Whitmeyer](#)
641 [and Karlstrom, 2007](#); [Amato et al., 2008](#)). The ca. 1.49–1.45 Ga Picuris Orogeny occurred
642 during a rare period of orogenic preservation between supercontinent cycles and is relatively
643 limited geographically to southern Laurentia which included parts of Precambrian Australia at
644 the time. The orogeny is identified in the southwestern United States where it deforms older
645 crust and appears not to have juvenile magmatism associated with it ([Daniel et al., 2013](#); [Mako](#)
646 [et al., 2015](#); [Aronoff et al., 2016](#)). The orogen extends into the northeastern United States
647 ([Medaris et al., 2021](#)), where it is progressively overprinted by the Grenvillian Orogeny in the
648 east. On the present-day eastern margin of Greenland, several exotic terranes were accreted to
649 the Kola-Karelia Craton, including Bergslagen-Livonia (ca. 1.89 and 1.84 Ga) and Amberland
650 (ca. 1.84 and 1.83 Ga; [Bogdanova et al., 2015](#)).

651 In Australia, extensive plate-margin orogenesis (ca. 1.82 to 1.55 Ga), similar to that inter-

652 preted for SW Laurentia, occurs throughout the South Australian Craton (Kimbang and Kararan
653 orogenies; [Hand et al., 2007](#)) and North Australian Craton (Yambah-Strangways-Leibig oroge-
654 nies). These likely formed a continuous accretionary system ([Payne et al., 2009](#); [Betts et al.,](#)
655 [2008](#); [Betts and Giles, 2006](#)). Extensive intracontinental orogenesis within Western Australia
656 is marked by the 1.82 to 1.77 Ga Capricorn Orogeny ([Johnson et al., 2013](#)) and the 1.68 to 1.62
657 Ga Mangaroon Orogeny ([Sheppard et al., 2005](#)). The Isan Orogeny (including the Chewings
658 and Olary orogenies) spanned the Paleo-Mesoproterozoic boundary (ca. 1.65-1.55 Ga). The
659 effects of this orogeny dominate the eastern entirety of pre-Phanerozoic Australia ([Morrissey](#)
660 [et al., 2011](#); [Tiddy et al., 2020](#); [Volante et al., 2020](#)), and extend into the Gawler Craton in
661 southern Australia ([Cutts et al., 2011](#)) and into the central North Australian craton ([Anderson](#)
662 [et al., 2013](#)). This orogeny is envisaged to have occurred as a consequence of collision between
663 Paleoproterozoic Australia (then consisting of the North Australia Craton and the South Aus-
664 tralian Craton with the North China Craton and a large piece of East Antarctica) and Laurentia
665 ([Pourteau et al., 2018](#)), and is recorded by the Rackett Orogeny in NW Laurentia ([Furlanetto](#)
666 [et al., 2013](#)). Orogenic activity coeval with younger Picuris Orogeny are found in the Gawler
667 Craton of Australia ([Hall et al., 2018](#); [Morrissey et al., 2019](#)), and in the Mount Isa region of
668 NE Australia ([Cave et al., 2022](#)).

669 In Antarctica, ca. 1.7 Ga orogenesis recorded along the coast of the Mawson Craton ([Peucat](#)
670 [et al., 1999](#)), as well as in the central Transantarctic Mountains ([Goodge et al., 2001](#); [Brown](#)
671 [et al., 2021](#)), although the extent of this orogenic activity into the interior of Antarctica is
672 unknown, it appears to be an extension of deformation in the Gawler Craton. There is some
673 ambiguity in the connections between Australia, Antarctica and western Laurentia ([Wingate](#)
674 [et al., 2002](#)), however, the connection of these three bodies is established from paleomagnetics
675 and geologic observations (aforementioned Nuna reconstructions; [Whitmeyer and Karlstrom,](#)
676 [2007](#)). The difficulty in precisely resolving the connections results from reworking of western
677 North America during the Cordilleran Orogeny and the uncertainties in paleomagnetic poles.

678 The accretionary orogenesis at this time on both the Dharwar-Bastar Cratons (South In-
679 dia) and the Bundelkhand Craton (northern India) is recorded in the Krishna Orogeny of the
680 Ongole Domain (1.68-1.60 Ma; [Henderson et al., 2014](#)) and in the Central Indian Tectonic Zone
681 ([Bhowmik, 2019](#)), respectively. The two halves of Peninsula India were likely separate conti-

682 nents before the Neoproterozoic. Paleomagnetic reconstructions have placed Southern India
683 conjugate to Antarctica or NE Australia ([Zhang et al., 2012](#)); however, more recent models
684 place the Indian continent adjacent to Baltica ([Pisarevsky et al., 2003; Cawood et al., 2020](#)).
685 Regardless, the Eastern Ghats appear to have been part of the active accretionary margin of
686 Nuna from 1.85 to 1.60 Ga or possibly as late as 1.45 Ga based on dating of tectonomagmatic
687 activity and an accreted ophiolite terrane in the Krishna Province ([Dasgupta et al., 2013](#)).

688 Prior to 1.95 Ga, the North China Craton was a set of microcontinents separated by ocean
689 basins that closed during the same period as both major Nunanian orogenies, completing by
690 1.85 Ga ([Zhao et al., 2012](#)), which is recorded in magmatism and widespread metamorphism
691 (e.g., [Yin et al., 2014; Cai et al., 2015; Wu et al., 2016; Liu et al., 2017](#)). However, the location of
692 the North China Craton has been subject to extensive debate, with various hypothesis ranging
693 from the northern margin of Siberia ([Halls et al., 2000](#)), paired with the Kola-Karelia Craton
694 and ([Wilde et al., 2002](#)), positioned between Baltica and Amazonia ([Pesonen et al., 2012](#)), or
695 joined with India outboard of the Nuna accretionary margin ([Zhao et al., 2011](#)). However,
696 we prefer more recent models that suggest a long-lived connection with the North Australian
697 Craton on the basis of more extensive paleomagnetic and geologic correlations ([Wang et al.,
698 2019; Nixon et al., 2022; Zhang et al., 2022](#)). Much of the North China Craton has similar
699 ages to the Inner-Nunanian Orogen, about half the craton experience widespread magmatism
700 as part of the Cordilleran Orogeny ([Wu et al., 2019](#)).

701 Most models for Nuna do not include the Kalahari Craton; however, the two models that
702 do, place it in opposite hemispheres ([Djeutchou et al., 2021; Cao et al., in prep.](#)). A recent
703 paleomagnetic reconstruction by [Djeutchou et al. \(2021\)](#) suggests the Zimbabwe Craton (the
704 northern part of the Paleoproterozoic Kalahari Craton) was juxtaposed against the southern
705 margin of the Superior Craton at 1.88 Ga. Additionally, their model calls for a pair of sub-
706duction zones along the south and western margins of the Kaapvaal Craton (the southern part
707 of the Kalahari Craton), reworking the Magondi and Kheis Belts and accreting the Rehoboth
708 Block at this time ([Kleinhanns et al., 2013](#)). However, such placement would leave little time for
709 the Kalahari Cratons to rift away from Superior, prior to the arrival of the Mazatzal Orogeny
710 ca. 1650 Ma along a margin that experienced the accretion of at least three major terranes ca.
711 1850 to 1450 Ma (Yavapai, Mazatzal and the Granite-Rhyolite terranes). The paleomagnetic

712 data and timing dates of deformed terranes are also consistent with a collision between Congo-
713 Tanzanian Craton with the south African cratons, which is consistent with ca. 2.0 Ga, resulting
714 in exhumation of eclogites in the Ubendian-Usagaran Belts (Collins et al., 2004; Tamblyn et al.,
715 2021). Regardless, both interpretations place these deformed and accretionary terranes within
716 the Circum-Nunanian Orogeny.

717 The Wopmay Orogeny, ca. 1.95 to 1.84 Ga, was a small subduction-related event on the
718 western margin of the present-day Slave Craton in northern Canada (Figure 8 Bowring and
719 Podosek, 1989). East-dipping subduction on the western margin accreted three separate arc
720 terranes, Great Bear, Hottah, and Fort Simpson, to the craton at ca. 1.88 Ga in the short-lived
721 Calderian Orogeny (Hildebrand et al., 2009; Cook, 2011).

722 The Eburnean Orogeny records the collision between the São Francisco Craton and the
723 Gabon Belt on the eastern Congo Craton margin, (ca. 2.12 to 2.0 Ga; Weber et al., 2016).
724 The orogeny is recorded in a set of tectonomagmatic events (Doumbia et al., 1998; Barbosa
725 et al., 2008; Peucat et al., 2011; Loose and Schenk, 2018; de Carvalho Filgueiras et al., 2020).
726 Although the Eburnean Orogeny was accretionary at the margins of Nuna (Figure 8e), it was
727 geographically isolated and earlier than the majority of Circum-Nunanian orogenesis.

728 Early models of the Limpopo Orogeny suggested it was active between 2.7 and 2.65 Ga
729 as the result of the collision between the Kaapvaal and Zimbabwe Cratons (Barton and van
730 Reenen, 1992), which is based on the age of granitoids contained within the thrust sheets.
731 However, metamorphic ages and more recent interpretations suggest it was active ca. 2.0 Ga
732 (Yin et al., 2019). The Limpopo Orogeny was probably intracontinental due to a lack of arc or
733 accretionary sediments of appropriate age (Yin et al., 2019).

734 The Birimian–Trans-Amazonian Orogeny occurred during the early stages or just prior to
735 the assembly of Nuna and may have been multiple spatially discrete events between ca 2.3 and
736 1.9 Ga. During this period, the West African Craton records considerable tectonomagmatic
737 activity in sedimentary deposits (Grenholm, 2019; Grenholm et al., 2019) as it collided with
738 the present-day northeastern Amazonian Craton (Grenholm, 2019). Both magmatic and meta-
739 morphic events are recorded in Amazonia during this period (De Roever et al., 2003; Savko
740 et al., 2015; Baltybaev et al., 2017; Klaver et al., 2015; da Rosa-Costa et al., 2008).

741 In the early Paleoproterozoic (ca. 2.5 to 2.3 Ga), a purported drop in magmatism corre-

sponded with several cratonic regions that experienced high-temperature, often contractional, metamorphic events (Pehrsson et al., 2013, 2014). Pehrsson et al. (2013) hypothesized these events are related to a formation of a supercontinent (Nunavutia), however, a recent paleomagnetic reconstruction suggested two separate supercratons were also consistent with pole determinations and patterns of dike swarms (Figure 9f; Liu et al., 2021). The reconstructions of the Siderian are based on paleomagnetic poles from 11 terranes (Liu et al., 2021; Salminen et al., 2021); however, less than one-quarter of blocks with igneous dates older than 2.3 Ga (Figure 10) are included in the reconstructions. Igneous activity in the interval 2.5 to 2.3 Ga is nearly ubiquitous across Archean terranes with the exception of regions covered by ice, sediments, and or Phanerozoic large igneous provinces. As a result, the classification as a super-continent/craton may be premature without the reconstructed positions of additional Archean terranes. Metamorphic dates in this time interval are sparse (Figure 10), but the metamorphic database is incomplete, so it is difficult to make clear inferences at this point. We refer to the regions with significant 2.4 Ga tectonothermal activity collectively as the Siderian Metamorphic Event, referring to the geologic time period rather than a coherent orogen due to the present uncertainty in reconstructions. The ca. 2.4 Ga affected terranes are now scattered across all seven continents and many have been reworked (Figure 9f and 10). Some of the terranes where the Siderian Metamorphic Event has been documented include the Mawson Craton (Duclaux et al., 2008), the Sask Craton (Chiarenzelli et al., 1998), Arrowsmith Orogen in Northern Canada (Hartlaub et al., 2007; Schultz et al., 2007; Berman et al., 2013a), the East Dharwar in India (Clark et al., 2009; Li et al., 2018), and the Sleaford Complex in south Australia (Halpin and Reid, 2016) and the North China Craton (Liu et al., 2017).

4. Model Evaluation

4.1. Ocean–Continent Boundary

Not counting for topographic relief, we estimate 57.7% of the Earth’s surface is covered by oceanic crust and 42.3% is covered by continental crust (Figure 4). The seafloor age model by Seton et al. (2020) covers a slightly smaller proportion of the Earth’s surface with seafloor ages, 57.4%. However, there are a few significant differences between the models.

In general, the edges of the seafloor age model by Seton et al. (2020) are easily correlated with high bathymetric gradients and deep water. However, there are a few regions where lo-

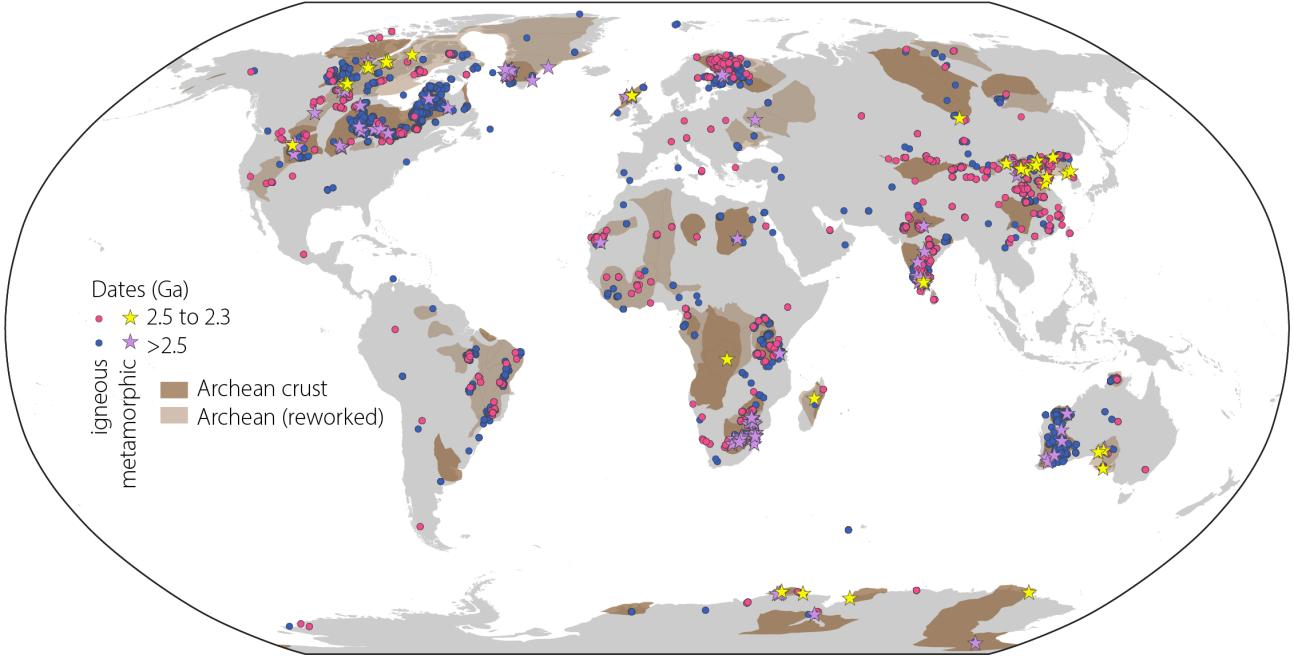


Figure 10: Regions with crust >2.3 Ga with superimposed locations of dated igneous and metamorphic activity. The observed dates are divided into the period 2.5 to 2.3 Ga, and older than 2.5 Ga. Reworked Archean crust is displayed in a lighter shade. Province boundaries from Figure 7 and a few additional Archean basement provinces in North America by Lund et al. (2015). Igneous ages extracted from Gard et al. (2019a) and Puetz (2018); metamorphic ages from DateView (Eglington, 2004) and the expanded metamorphic database by Brown and Johnson (2018).

772 cations differ significantly between their model and our ocean–continent boundary. Some of
 773 the differences may be due to the quality of magnetic data near the continents where remnant
 774 magnetization may be reset by high temperatures beneath insulating sediments or where mag-
 775 netic data is of insufficient quality and/or density to resolve seafloor ages. A few of the larger
 776 differences include the Greenland-Iceland-Faroe Ridge (GIFR), Blake Plateau, Gulf of Mexico,
 777 and some microcontinents.

778 Perhaps the most obvious difference between our model and Seton et al. (2020) is the
 779 inclusion of the GIFR as a region of continental crust. In a recent comprehensive paper by
 780 Foulger et al. (2020), the authors make a compelling case that the GIFR is a peculiar region with
 781 variably extended continental crust rather than the product of anomalous oceanic volcanism.
 782 The nearby Jan Mayen microcontinent has been recognized as a microcontinents for decades
 783 (Peron-Pinvidic et al., 2012, and references therein), lending credence to the model. The total
 784 thickness of the crust approaches 40 km thick beneath the Iceland microcontinent—a value
 785 more typical of continental than oceanic crust. The symmetric, linear magnetic anomalies
 786 characteristic of seafloor spreading are muddled in this region, possibly indicating a complex
 787 history of rifting and volcanism. The upper crust in the region exhibits seismic properties and

788 layer thicknesses typical of oceanic crust (3–10 km), but the middle and lower crustal seismic
789 velocities and densities are better explained by continental material (Foulger et al., 2020).
790 Therefore, we favor their interpretation in our model.

791 There are a number of other microcontinents that we have included in our model that are
792 not found in the Seton et al. (2020) model. We have included the Hovgaard Ridge and the East
793 Greenland Ridge in the Arctic Ocean (Funck et al., 2016). Several microcontinents lie in the
794 Indian Ocean including the Mascarene Plateau, Chagos–Laccadive Ridge, Gulden Draak, and
795 Batavia Knoll (Torsvik et al., 2013; Gardner et al., 2015; Halpin et al., 2017), which are formerly
796 pieces of Madagascar and India, respectively. We have also included the Bollons Seamount east
797 of the Campbell Plateau in the Pacific (Davy, 2006).

798 The Blake Plateau on the Atlantic side of Florida sits at a bathymetric depth of \sim 1000 m
799 and is underlain by transitional crust (Dillon et al., 1988). In the Gulf of Mexico, the ocean–
800 continent transition is obscured by sedimentary cover making magnetic data the most useful for
801 identifying the boundary, but the global models are relatively low resolution. However, there
802 are industry magnetic datasets that we do not have access to that yield a much clearer view of
803 the Gulf of Mexico.

804 In the future, incorporation of seismic reflection profiles taken across the ocean–continent
805 transition and improvements in the magnetic datasets may warrant alterations to the ocean–
806 continent boundary.

807 *4.2. Plate Boundaries*

808 There are significant differences in the lengths of plate boundaries by type (Table 5). The
809 total length of all major plate boundaries is nearly five times the Earth’s circumference. Among
810 major boundaries, the total length of divergent boundaries is significantly greater than con-
811 vergent boundaries. The mid-ocean ridge system and associated transforms is \sim 1.7 times the
812 length of subduction zones. Transform boundaries account for the least length of major bound-
813 aries, but this does not include the transforms between mid-ocean ridge segments.

814 Including minor deformation zone and microplate boundaries more than doubles the total
815 length of boundaries (Table 5). The latter number may be a bit misleading as the minor bound-
816 aries generally bracket the deformation, while in reality multiple structures may accommodate

Table 5: Lengths of plate boundaries. Major and minor boundaries are identified in Figure 2.

Boundary Type	Major (km)	Minor (km)
spreading center	87,001	12,680
extensional zone	13,702	40,134
subduction zone	52,966	9,688
thrust	17,614	65,077
dextral transform	15,476	41,723
sinistral transform	13,276	24,099
inferred	0	19,177

the motion. Among minor boundaries, convergent and transform boundaries are longer than divergent boundaries. Also in contrast to major boundaries, spreading ridges and subduction zones are less common than extensional zones and thrusts.

4.3. Plate Model

4.3.1. Comparison with Bird (2003)

The major differences between the [Bird \(2003\)](#) plate model and our model are the microplates and deformation zones (Figure 1a). Most major plate boundaries are very similar to [Bird \(2003\)](#) with some refinement. The [Bird \(2003\)](#) model contains 52 polygons, whereas ours contains 121 regions comprising rigid plates, microplates and deformation zones. Most of these added regions are located in the Alpine–Himalayan Belt and the North American Cordillera. These changes are driven by a significant improvement in the number and quality of land-based GPS coverage; however, several of our additions have long been recognized as plate boundary zones (e.g., [Gordon, 1998](#); [Lowman and Yates, 2002](#); [Kreemer et al., 2014](#)). Many of these regions have also been previously identified as microplates (case studies in Table 2). In total, we estimate 16% of the Earth’s surface is covered by microplates and deformation zones, divided roughly equally between them whereas the Bird model only includes 6%.

The southern Somali–African boundary has changed considerably with respect to the [Bird \(2003\)](#) model. The [Bird \(2003\)](#) model connects several discrete regions of seismicity in southern Mozambique and central South Africa. However, GPS motions suggest there is little deformation on either side of this boundary ([Kreemer et al., 2014](#)). From GPS data, the boundary appears to be further east, near the southern Mozambique coast ([Stamps et al., 2021](#)).

Another region of significant refinement relative to Bird’s model among the microplates and deformation zones of the Philippines and eastern Indonesia. Our model more closely follows

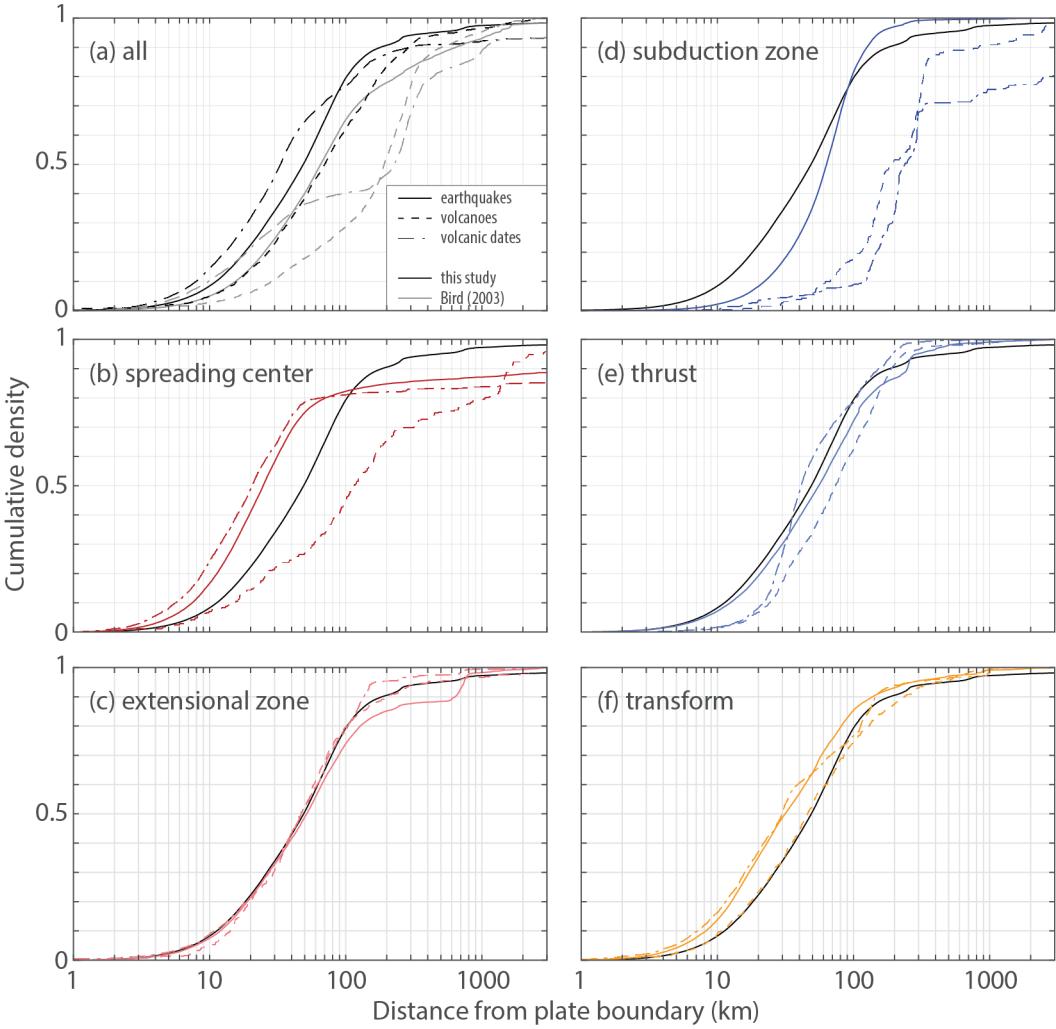


Figure 11: Cumulative distributions of earthquakes and volcanoes by distance to plate boundaries. Each subplot contains CDFs from three datasets: (solid) earthquakes from the ANSS catalog, magnitude 3.0 to 5.5, 1990 to 2020 and magnitude 5.5+, 1970 to 2020 ([ANSS, 2020](#)); (dashed) Quaternary volcanoes ([Global Volcanism Program, 2013](#)); and (dashed-dotted) dates of volcanic samples limited to Quaternary samples [Gard et al. \(2019a\)](#) with duplicate sample locations removed to limit oversampling. (a) Computed using all data in each respective dataset, black lines using our model and grey lines using [Bird \(2003\)](#). (b-f) Computed for listed plate boundary types as classified in Figure 2. The black line on each plot is earthquakes from all plate boundaries as a reference.

several boundaries identified by [Zahirovic et al. \(2014\)](#), which shows better correlations with seismicity and active fault models ([ANSS, 2020; Styron and Pagani, 2020](#)).

In our plate model, $\sim 80\%$ of earthquakes occur within 100 km of a plate boundary and $\sim 91\%$ within 200 km (Figure 11a). Approximately 73% of earthquakes occur within deformation zones and microplates. For the [Bird \(2003\)](#) model, these percentages are significantly lower. Less than 27% of earthquakes lie within Bird's microplates and 65% of earthquakes lie within 100 km of plate boundaries. Therefore, we suggest our new model provides a more accurate representation of the actively deforming crustal regions.

The pattern of distance of earthquakes from plate boundaries varies depending on the type

of plate boundary. Earthquakes are centered close to spreading centers (Figure 11b), but are more diffuse around transforms (Figure 11f), which occurs because transform plate motion is frequently accommodated by multiple faults rather than a single structure (e.g., Pacific–North American boundary in California, Hauksson et al., 2013; DeMets et al., 2014). In contrast to spreading centers and transforms that identify the centers of deformation, extensional zones and thrusts identify the boundaries of internal deformation so it makes sense that earthquakes are distributed at a greater distance from these boundaries (Figure 11c, e). Subduction zone earthquakes are distributions furthest from their associated plate boundaries (Figure 11d). The more rapid increase in cumulative density of subduction zone earthquakes could be due to our limiting earthquakes to 30 km depth and/or the earthquakes extend far enough into the deformation zone that the opposite boundary becomes the closest (Figure 2).

4.3.2. Comparison with Tomography and Volcanism

Our plate boundary model correlates well with slow seismic velocity slices from 40 to 90 km depth, with 70 km displaying the most similarity to the plate boundary zones (Figure 12). Greater than 90 km depth, the mantle beneath the oceans has significantly larger negative velocity anomalies. Greater than approximately 125 km depth, some of the continental plate boundary zones begin to lose their negative velocity anomalies. The correlation between plate boundary zones and shallow mantle shear wave velocity anomalies are the result of thinner and warmer lithosphere in actively deforming regions with respect to cold, thick, rigid plate interiors.

Only a few plate boundaries and deformation zones are not clearly associated with negative seismic velocity anomalies. The Cordilleran Frontal Thrust, Lesser Antilles Arc and the Lwandle Plate have positive seismic anomalies. Oceanic intraplate deformation zones do not show clear correlations with negative tomographic anomalies. For example, the Capricorn region between the Indian and Australian Plates, the Macquarie Microplate, and the boundary between North and South America are not clearly delineated by tomography.

Volcanoes also show high correlation with the plate boundary zones as ~80% lie within deformation zones and microplates (Figure 12). Like earthquakes, there are differences between the distribution of distance from plate boundaries by type (Figure 11). Volcanoes are furthest from subduction zones, with the majority ranging between 130 to 330 km. For transforms

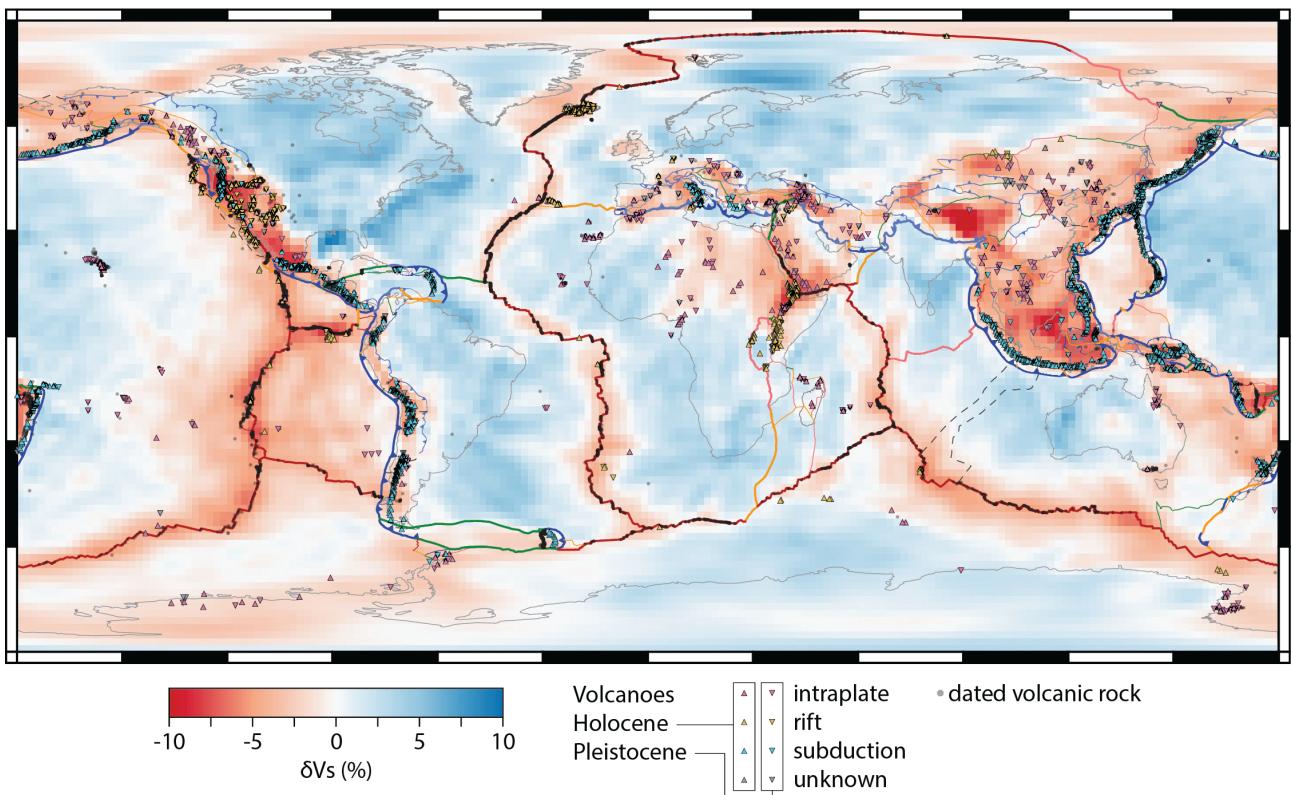


Figure 12: Shear wave tomography at 70 km depth demonstrates a high spatial correlation between negative velocity perturbations and plate boundaries and deformation zones. The shear wave velocity model is by [Debayle et al. \(2016\)](#). Nearly 80% of recently active, Quaternary volcanoes lie within microplates and deformation zones. Volcano locations from [Global Volcanism Program \(2013\)](#). Pleistocene and younger volcanics from [Gard et al. \(2019a\)](#); originally from EarthChem.org affiliated databases.

and extension zones, >75% lie within 100 km of the boundary. The distribution of volcanoes from spreading centers is more complex because most of the volcanoes identified near these environments are seamounts associated with hotspots rather than flows at ridges. If a full accounting of flows could be made along ridges, it is likely the distribution would indicate most are much closer to the drawn boundaries [Rubin \(2016\)](#). We try to account for this by using dates from a global geochemical dataset ([Gard et al., 2019a](#), and Figure 12), which results in a similar distribution of distances as the volcanic eruption database (Figure 11). The only exception is the distribution of dates from recent volcanic samples, which fall considerably closer to spreading centers than the eruption database similar to earthquakes (Figure 11b).

The correlation between volcanoes and seismic velocities may imply partial melt is common beneath most plate boundary zones. However, there are a few that have minimal volcanism but high seismicity, including Tibet and the Tien Shan Mountains. The Amur, Yangtze and Okhotsk regions also have very little volcanism outside the volcanic arcs on their margins. These regions also have little seismicity in their interiors.

There are a few regions with negative seismic anomalies and/or recently active volcanic centers that do not correspond with plate boundaries. Several of these regions are hotspots associated with mantle plumes (e.g., Hawaii, Reunion, Cape Verde). These regions do not have clear negative velocity anomalies at 70 km depth within the scale of the tomography model (Figure 12). There is a negative shear wave anomaly in the Arctic without volcanics, which may be related to prior rifting and the Eureka Orogeny ([Darbyshire, 2005](#)). The Saharan Metacraton has both low shear velocities and volcanism, but little seismicity. There are currently no GPS data from western Egypt, Chad, Sudan or Libya above the seismic anomaly, but GPS data in Nigeria do not indicate active deformation above the negative velocity region. The West Antarctic Rift contains both volcanoes and a negative shear wave anomaly related to renewed extension in the Cenozoic ([Winberry and Anandakrishnan, 2004](#); [Gupta et al., 2009](#); [O'Donnell et al., 2019](#)). While there is active volcanism associated with the Marie Byrd and Erebus hotspots, seafloor magnetic anomalies suggest the rift was active as recently as 11 Ma ([Granot and Dymant, 2018](#)) and minimal seismicity suggests the region is inactive at present.

907 4.3.3. *Uncertain Plate Boundaries*

908 Many oceanic deformation zones and microplates have relatively uncertain boundaries due to
909 a lack of GPS constraints, significant seismicity and/or distinctive bathymetric features akin to
910 boundaries. The Azores deformation zone, for example, has a well-defined boundary along the
911 northern margin by seismicity. GPS data on individual islands suggest internal deformation
912 ([Fernandes et al., 2006](#); [Marques et al., 2013](#)), but the southern boundary is not clear from
913 seismicity. Our best estimate places the southern boundary on the East Azores Fracture Zone
914 due to its clear topographic expression.

915 The Capricorn Plate is a region of diffuse extension between the Indian and Australian
916 Plates. The region is large, nearly the same size as the Indian Plate, but the eastern and
917 southeastern boundaries are very uncertain (Figure 13). Magnetic anomaly maps of the plate
918 are sparse on the southern boundary (i.e., Wharton Basin). Bathymetry is complicated by
919 several features including the Ninetyeast Ridge, Diamantina Escarpment, Roo Rise, Vening
920 Meinesz Seamounts, and the Raitt Rise. Earthquakes, while indicating extension, are not of
921 sufficient density to clearly delineate the boundaries (Figure 13). Furthermore, the Indian
922 and Australian Plates have relatively high rates of diffuse intraplate seismicity, without clear
923 clustering indicative of rigid plates boundaries.

924 The eastern, western, and northern boundary of the Lwandle Microplate are poorly con-
925 strained. Most authors have drawn the northern boundary as an extension of the Quasama
926 Seismic Axis on the southern end of the Rovuma Microplate to Madagascar and then traversing
927 Madagascar along a constant line of latitude before running along the eastern margin of the
928 island (e.g., [Saria et al., 2014](#)). However, there is little to no observed seismicity to constrain
929 this model. A more recent model suggests the northern boundary is constrained by a dextral
930 transform that runs along the Comoros archipelago which is consistent with seismicity and GPS
931 velocities in northern Madagascar ([King et al., 2017](#); [Famin et al., 2020](#); [Stamps et al., 2018,](#)
932 [2021](#)). For this reason, we prefer the northernmost boundary in our model, and have separated
933 the northern Lwandle region into the Comoros deformation zone.

934 The boundary between the oceanic portion of the North and South American Plates east of
935 the Lesser Antilles Trench is poorly defined by limited diffuse seismicity. In fact, the low level
936 of seismicity is similar to many intraplate regions not considered deformation zones. The lack

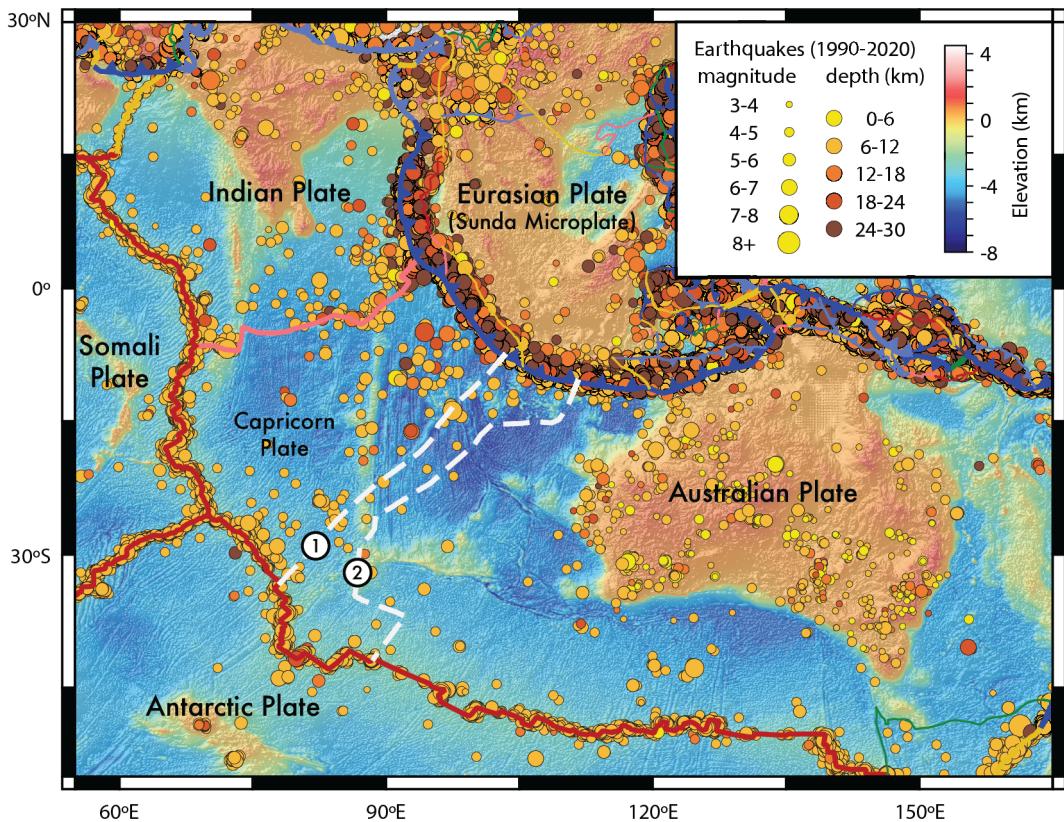


Figure 13: Models of the Capricorn Plate and associated Mid-Indian deformation zone. The white boundaries are a couple of the proposed southern boundaries: 1, (Rathnayake et al., 2019); 2, (Royer and Gordon, 1997).

of GPS stations in both regions makes it difficult to define the edges of the deformation zone reliably. Given the lack of a clear diffuse region, we draw the boundary as the most prominent feature, the Fifteen-Twenty Fracture Zone, which is consistent with prior interpretations (Roest and Collette, 1986; Dixon and Mao, 1997).

4.4. Geologic Province Model

The geological province model includes 918 polygons, of which 790 (86%) are continental. Most continental terranes are linear belts with areas that range between 10^4 and 10^6 km 2 . The median province area is 175,000 km 2 , but varies between different continental regions from 10^5 and 3×10^5 km 2 (Figure 14a). The distribution of province sizes is largest in South America and Africa. Part of the reason may be due to the inability to pick out smaller terranes beneath thick sedimentary cover and relatively few studies with high resolution geophysical data. It is also due to the number of composite type terranes such as cratons, shields and orogens that are built from ophiolites, accretionary complexes, and volcanic arcs (Figure 14b). However, some provinces are naturally larger due to the processes involved in their formation (i.e., passive margins and wide rifts).

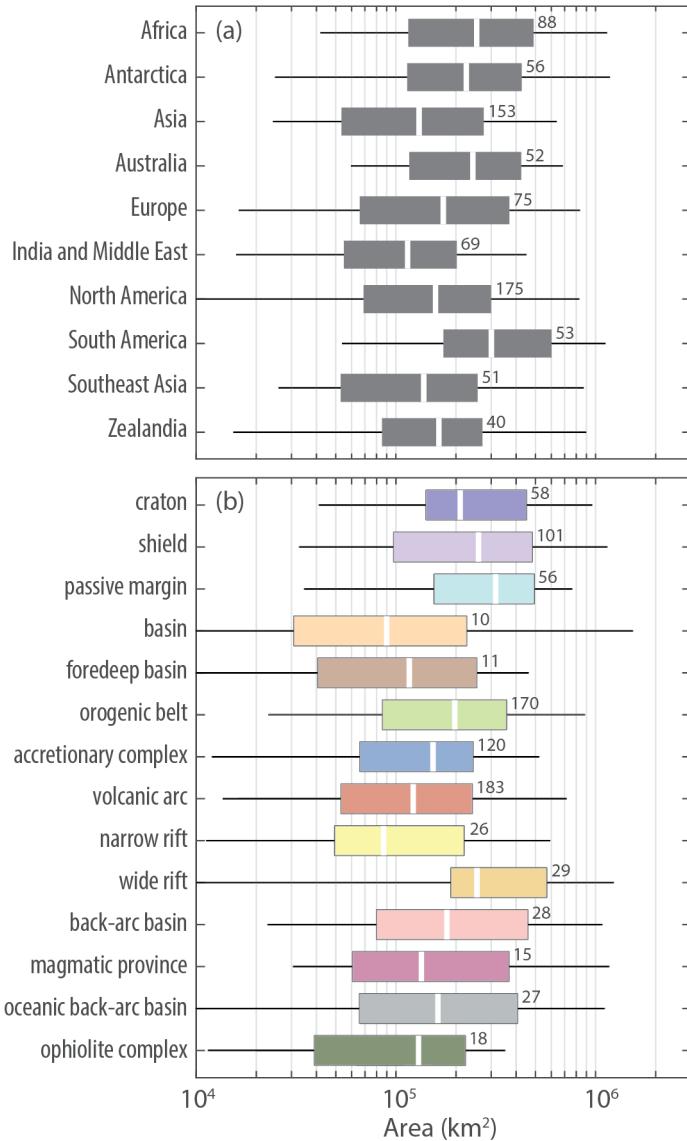


Figure 14: Area of geological provinces for (a) continental regions and by (b) province type. The white bar indicates the median, box enclose the 0.25 to 0.75 quantiles and whiskers extend to the 0.05 and 0.95 quantiles. The number beside the boxes indicates the number of provinces for each region.

952 The most common province types are orogenic belts, volcanic arcs and accretionary com-
953 plexes (Figure 14b). There are many fewer extended terranes, however, passive margins are
954 generally rifted margins with thick sediment on top. Many ancient rifts are incorporated into
955 orogenic belts. Where orogenic belts have been separated into their individual constituents,
956 there is often an accretionary complex and associated arc. However, in many cases only a vol-
957 canic arc is identified, which may be because the related accretionary complex is commingled
958 with the volcanic arc and difficult to separate or was destroyed during orogenesis.

959 *4.4.1. Comparison with Matthews et al. (2016) and Merdith et al. (2021)*

960 The model by [Matthews et al. \(2016\)](#) was constructed to model plate motions over the past
961 400 Ma. Given their model timeframe, continental provinces that moved relative to one another
962 prior to 400 Ma are not typically subdivided. As a result, our model contains a significantly
963 greater number of terrane divisions. Many of the divisions in the [Matthews et al. \(2016\)](#) model
964 are also simpler than ours. This simplicity requires fewer polygon vertices, which may make
965 computation of plate rotations more rapid, but it compromises accuracy of filtering geologic
966 data for analysis.

967 The [Merdith et al. \(2021\)](#) model extends the [Matthews et al. \(2016\)](#) plate motions back
968 to 1000 Ma, which required the addition of many more province boundaries. The models
969 for southern Africa, the northwest Cordillera, and the Mongol Orocline are similar to ours.
970 However, there are significant differences in the number of provinces in Australia, Antarctica,
971 eastern Europe and the Superior Craton. Many of these differences occur where we include a
972 finer resolution of terrane boundaries with tectonic histories that converge prior to 1000 Ma.
973 In light of the agreement with previous reconstruction polygons, the model presented in this
974 study could be used in the future for plate reconstructions into the Archean.

975 *4.4.2. Last Orogeny*

976 To ensure that our orogenic model is reasonable (Figures 8 and 9), we have computed kernel
977 density estimates (KDE) for both magmatic crystallization dates and metamorphic dates for
978 each orogen (Figure 15). The dates are mostly derived from U–Pb, ^{207}Pb – ^{206}Pb , and U–Th–Pb
979 analyses of zircon and monazite. The zircon dates can provide constraints for both igneous and
980 metamorphic events whereas most monazite dates record metamorphic events in the datasets
981 we used ([Eglington, 2004](#); [Puetz, 2018](#); [Brown and Johnson, 2018](#); [Gard et al., 2019a](#)). Other

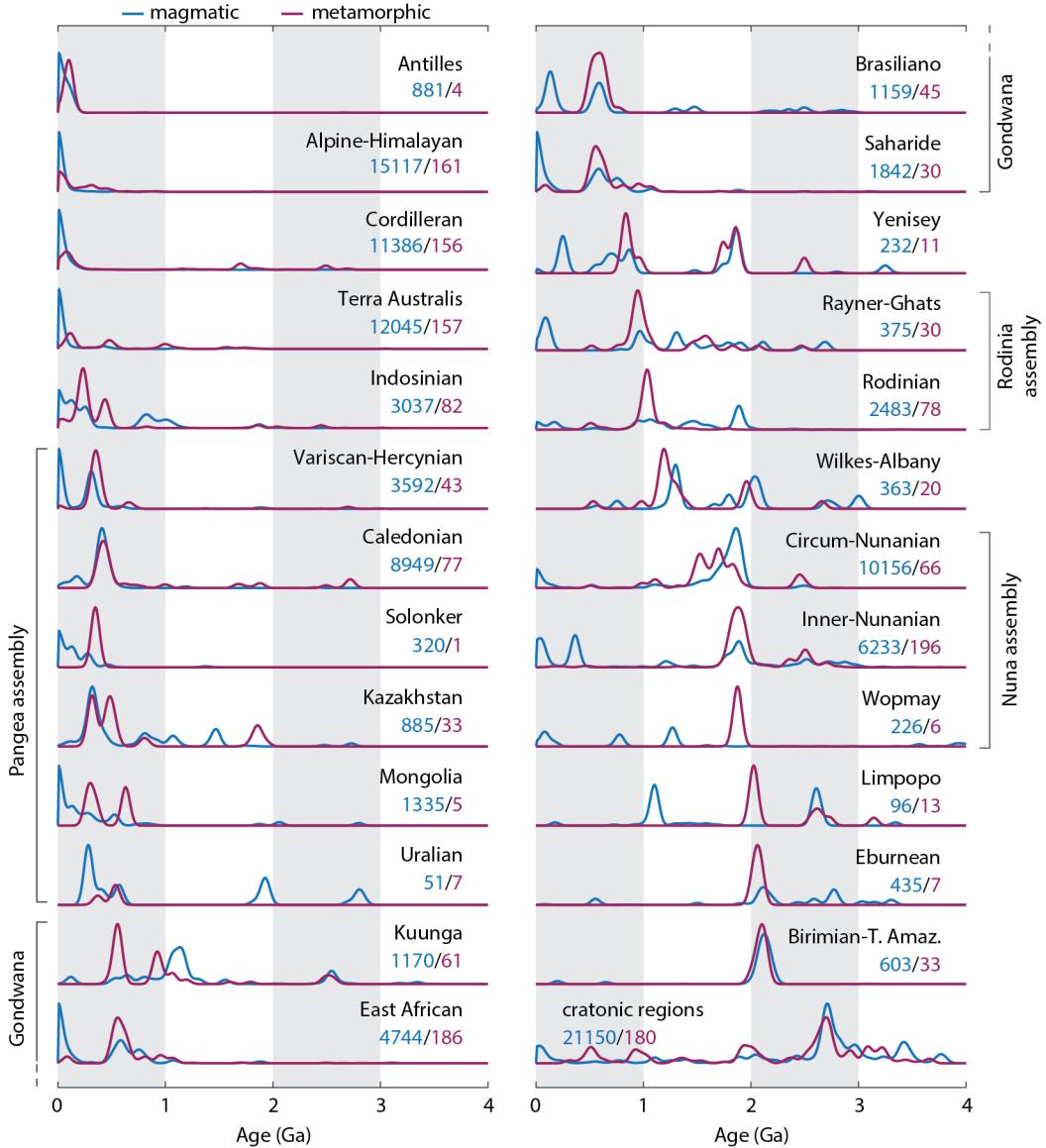


Figure 15: Kernel density estimates for magmatic crystallization dates (blue, Gard et al. (2019a)) and metamorphic dates (violet, Brown and Johnson (2018) updated; Eglington (2004)). A list of metamorphic dates are given in the Supplemental Material. No dates for the Scotian Orogeny are available in the database.

isotopic dating systems (e.g., ^{40}Ar - ^{39}Ar , K-Ar, and Rb-Sr) may record lower closure temperatures and can correspond to less significant heating events and are therefore generally excluded from the dataset except where the dates match the regional higher temperature metamorphism or in younger systems where U-Th-Pb lose precision.

The largest age peak for metamorphic KDEs always falls within the range of dates attributed to an orogen as discussed in Section 3.2.2. It is important to note that these dates were not used to define the orogen, but undoubtedly many of these orogen models are based in part on geochronology. In Figure 8, one can see the spatial correspondence between metamorphic dates and our orogen interpretation. In a few cases, there are smaller-magnitude metamorphic date peaks that record P-T conditions of previous orogenic events. For example, in the Canadian

992 Cordillera Frontal Thrust Belt a few dates are associated with the Wyoming Craton (ca. 1780
993 Ma; [Cheney et al., 2004b,a](#)) and not the more recent Cordilleran deformation (Figure 8). These
994 older metamorphic events are still evident because Cordilleran deformation reaches relatively
995 shallow levels in these regions and has not reset the isotopic systems associated with the older
996 events.

997 Very few orogens have significant age peaks younger than the assigned orogen. The excep-
998 tions include the Rodinian Orogeny which in some cases includes some Gondwana-associated
999 dates, particularly in Antarctica, and the Circum-Nunalian terranes, which include some
1000 Rodinia-associated overprinting (Figure 15)—specifically in North America where metamorphic
1001 conditions are poorly sampled (Figure 8). These discrepancies between the orogenic model and
1002 observed metamorphic conditions may indicate some refinement of the province model may be
1003 necessary.

1004 The magmatic dates are noisier, often indicating several episodes of magmatism, which
1005 attests to the multi-generational history to the growth of many provinces (Figure 15). There
1006 are a surprising number of provinces which have magmatic peaks near the present day that lie in
1007 generally much older orogens. In the East African Orogen, the younger ages are associated with
1008 the East African Rift. In the Saharides and Trans Hudson Orogen, these younger magmatic
1009 ages are the result of Cenozoic intraplate magmatism, which does not appear to be associated
1010 with active tectonics. In many regions with younger volcanism, the geodynamic process that
1011 has created the melts does not appear to have reset metamorphic conditions or alternatively
1012 the depths at which metamorphism has occurred have not yet been exhumed.

1013 While magmatic dates appear to be less reliable for identifying the last orogen event, there
1014 is often a peak in the KDE that corresponds with the metamorphic peak. The Limpopo Orogen
1015 is a clear exception where no magmatic peak is observed during the orogenic event. The missing
1016 magmatic peak at ca. 2.0 Ga is the reason the Limpopo orogen was originally interpreted as a
1017 ca. 2.7 Ga event, rather than a younger intracontinental orogen ([Barton and van Reenen, 1992](#);
1018 [Yin et al., 2019](#)). The only other orogen with missing magmatic dates is the Wopmay Orogen
1019 in northwest Canada, however, this orogen is poorly exposed so there are very few data from
1020 the terranes that comprise it.

1021 The KDEs for the cratonic regions display significant magmatic and metamorphic activity

1022 prior to 2.5 Ga, with a few minor metamorphic and igneous peaks. Many of these small
1023 younger (<2.5 Ga) metamorphic peaks do not appear to correspond with magmatism, which
1024 may indicate minor degrees of intracratonic deformation as orogens occur around them. We
1025 suggest this observation is consistent with a persistent mechanically strong lithosphere able
1026 to resist deformation in most cases. The young magmatic peak within cratonic regions is
1027 common across most orogens. As most of these regions are intraplate several questions arise
1028 regarding the lack of widespread intraplate volcanism in the past. Are intraplate volcanics
1029 easily eroded and therefore poorly preserved? If not, is there a sampling bias, i.e., petrologists
1030 and geochemists vastly oversample present-day intraplate volcanism, or are they generally too
1031 small and distributed in the past to easily recognize and thus ignored?

1032 4.4.3. Uncertain Province Boundaries

1033 Perhaps the poorest geologically characterized region is Antarctica due to the lack of out-
1034 crops. Most geological province information about Antarctica comes from regions near the
1035 coast typically combined with constraints from conjugate terranes on now-distant continents
1036 (e.g., [Boger, 2011](#); [Flowerdew et al., 2013](#); [Goodge et al., 2017](#); [Maritati et al., 2019](#); [Rup-
1037 pel et al., 2020](#)). Due to the nature of these peripheral constraints, province boundaries are
1038 more uncertain into the subglacial interior and are typically guided by geophysical fields such
1039 as gravity and magnetics following linear trends ([Aitken et al., 2014](#); [Maritati et al., 2016](#)).
1040 Because of the uncertainties, it has resulted in a myriad of province models that can change
1041 significantly as additional geological data are added and higher resolution geophysical models
1042 are produced. Very few models include interior provinces that do not reach the coast because of
1043 the lack of geological control (cf., [Ferraccioli et al., 2011](#)), however new multivariate approaches
1044 to mapping lithospheric boundaries suggest significant interior complexity characterizes East
1045 Antarctica ([Stål et al., 2019](#)).

1046 Our Antarctic province model is derived from several geophysical constraints (Table 3). As
1047 with other continents, magnetic anomalies are often the best constraint on province bound-
1048 aries in the upper crust (Figure 16a). Geochronology is important for defining provinces and
1049 identifying boundaries in areas of outcrop, which are mostly concentrated near the coast. The
1050 PetroChron Antarctica database is a compilation of geochronology and other geological data of
1051 Antarctica ([Sanchez et al., 2021](#)), that will surely continue to improve our insights into the pre-

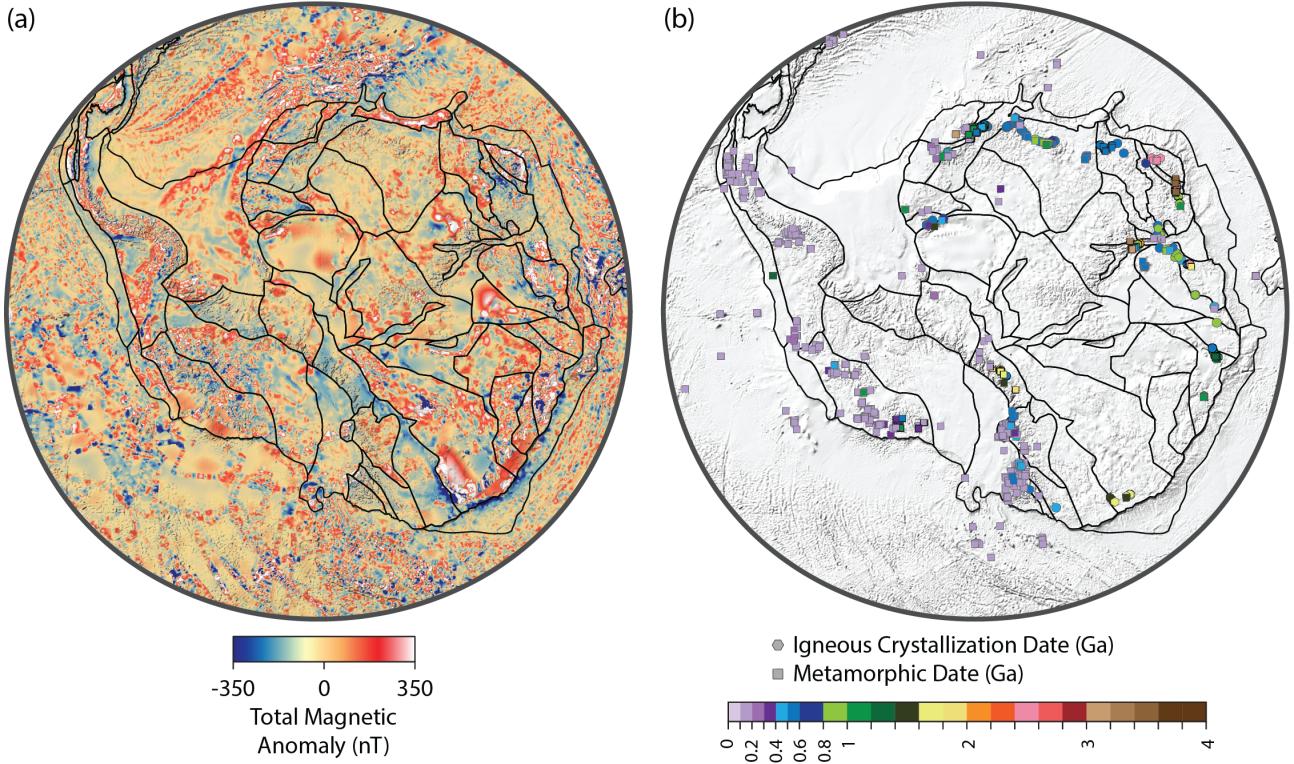


Figure 16: Magnetic (a) and geochronologic (b) constraints on the Antarctic province model. Magnetic data from ADMAP2 (Golynsky et al., 2018) and EMAG2_V3 (Meyer and Saltus, 2016) to fill gaps. Igneous crystallization dates from Gard et al. (2019a) and Puetz (2018) and metamorphic dates from Brown and Johnson (2018), updated, and DateView, (Eglinton, 2004).

1052 cise positioning of Antarctic province boundaries. Our model includes a few interior provinces
 1053 that we interpret as cratonic or shield terranes due to the lithospheric characteristics, though
 1054 they are relatively large. These terranes are ringed by orogens at the edges of the continent,
 1055 similar to many other continental shields (e.g., Kaapvaal, Congo, and Siberian Cratons).

1056 Most other poorly resolved regions include those covered by thick sedimentary cover: Patag-
 1057 onia, Parana Basin, Saudi Arabian Platform, North Africa, and West Siberian Basin. In these
 1058 regions the provinces tend to be much larger than the average and there are multiple in-
 1059 terpretations for their divisions. These models are difficult to independently assess because
 1060 the geophysical data in many of these regions is relatively poor. For example, the Saharan
 1061 Metacraton and Tuareg Shield have considerably different interpretations for their evolution.

1062 The Tuareg Shield is constructed of a set of imbricated arc terranes. Most models show lit-
 1063 tle to no association with the Saharan Metacraton, which is interpreted as a set of Archean
 1064 cores with reworked subcontinental lithosphere (Liégeois et al., 2013; Sobh et al., 2020). This
 1065 interpretation is based on thermophysical modeling using multiple geophysical data, but the
 1066 data are relatively low resolution. An alternative explanation developed using geochronological

1067 and geochemical data suggests the Tuareg Shield, Arabian–Nubian Shield and portions of the
1068 Saharan Metacraton were a single volcanic arc system that was broken in multiple segments,
1069 reorganized and then accreted to the core of the metacraton, deforming it in the process in the
1070 Neoproterozoic ([Liégeois, 2018](#); [Sengör et al., 2020](#); [Blades et al., 2021](#)).

1071 *4.5. Future Improvements*

1072 There are a number of improvements that can be made to the province definitions in future
1073 versions. Higher resolution geologic maps could be incorporated that have better lithological
1074 data and finer age resolution in the Precambrian, especially the Archean. Likewise, higher
1075 resolution magnetic data will be very useful for improving terrane boundaries, especially when
1076 the resolution is \sim 10 km pixel spacing. Most global geophysical datasets are available at too
1077 low a resolution to precisely resolve many terrane boundaries.

1078 Province types that do not provide information about the tectonic construction of the crust
1079 (basin, shield, and craton), should be replaced in future versions. Doing so may require im-
1080 proved geologic models of Precambrian regions, which may require the addition of new tectonic
1081 settings that are unique to the early Earth such as granite-greenstone belts. Likewise, orogenic
1082 belts could also be deconstructed into the types of terranes that build them, i.e., separating out
1083 arcs and accretionary margins could help improve models of crustal growth and composition.

1084 Although we currently only ascribe the last orogenic event to provinces, many regions have
1085 experienced multiple orogenic events throughout their history. Adding this history to the at-
1086 tribute tables would provide a more complete view of an individual province’s chronological his-
1087 tory, which could be paired with information about the changing tectonic environments. Such
1088 a model would require a more sophisticated analysis of global geochemical and geochronological
1089 data. This process would be tedious to perform on a province by province scale so an automated
1090 approach will be warranted and could include recent machine-learning based approaches to pre-
1091 dicting tectonic environments (e.g., [Tang et al., 2020](#); [Tetley et al., 2020](#)). These improvements
1092 would place valuable constraints on plate reconstruction models.

1093 Constructing the province models can include a community-driven approach that is updated
1094 by individuals or small groups with expertise in specific regions. To that end, we have set
1095 up a GitHub project page where individuals can download the shapefiles, update them, and
1096 upload their changes along with a description, references, and rationale for the changes. The

1097 changes can then be evaluated and incorporated into the model if deemed credible. Thus, one
1098 would not need to wait for a formally published reference before using the improved maps
1099 (https://github.com/dhasterok/global_tectonics).

1100 **5. Summary**

1101 We have produced a set of plate and geologic province maps that can be used to improve spa-
1102 tial analysis of geoscientific data. The plate model boundaries are validated against earthquake
1103 locations, active fault traces, and GPS motions and shows good correlation with shear wave
1104 velocities at 70 km depth and active volcanism. The geological province model is constructed
1105 from a collage of published models and refined using a wide-variety geophysical and geological
1106 data. The most useful data were found to be aeromagnetic anomalies when the models are high
1107 resolution. The province model has not been independently validated at this point, but relies
1108 on the accuracy of the original studies. However, the last orogeny is validated using metamor-
1109 phic and igneous dates. The plate and province polygons are drawn so that the boundaries are
1110 seamless between the files and additionally include an ocean–continent boundary and a plate
1111 boundary type.

1112 The maps are available in a shapefile format that can be easily interpreted by many modern
1113 computer languages and have advantages over raster maps for geographically selecting data
1114 for various types of analysis. The maps can also be used as a data standard for prescribing
1115 spatial metadata in global databases. Because the models are available on GitHub, the geologic
1116 community can submit updates and fixes to improve their accuracy.

1117 **Data Availability**

1118 Plate and province models produced in this study are available at the GitHub repository,
1119 https://github.com/dhasterok/global_tectonics. The models can be found in shapefile
1120 format suitable for GIS programs, KML for programs such as GoogleEarth, and GMT format for
1121 Generic Mapping Tools. The models are also available in the global tectonics library on Zenodo,
1122 <http://doi.org/10.5281/zenodo.5093930>, which includes additional global geophysical and
1123 geochronological datasets that are useful for research and educational applications.

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