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New maps of global geological provinces and tectonic plates

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Abstract

Accurate spatial models of tectonic plates and geological terranes are important for analyzing and interpreting a wide variety of geoscientific data and developing compositional and physical models of the lithosphere. We present a global compilation of active plate boundaries and geological provinces in a shapefile format with interpretive attributes (e.g., crust type, plate type, province type, last orogeny). The initial plate and province boundaries are constructed from a combination of published global and regional models that we refine using a variety of geoscientific constraints including, but not limited to, relative GPS motions, earthquakes, mapped faults, potential field characteristics, and geochronology. These new plate model show improved correlation to observed earthquake and volcano occurrences within deformation zones and microplates, compared to existing models, capturing 73 and 80% of these criteria, respectively. Deformation zones and microplates only account for 16% of Earth's surface area. We estimate 57.5% of the Earth's surface is covered by oceanic crust, which is a slight increase relative to the most recent seafloor age model. The model of last orogenies agrees well with peaks in the globally summed geochronology data. There is room for improvement in future editions of our global plate and geologic provinces model where basins, ice, or lack of geological data fidelity obscure bedrock geology, particularly in the eastern Central Asian Orogenic Belt, much of Africa, East Antarctica, and eastern Australia. Additionally, some province types—orogens, shields, and cratons that are homogenized within our global scheme—can likely be partitioned into smaller terranes with more precise geodynamic attributes. Despite some of these shortcomings, the digital maps presented here form a self-consistent data standard for adding spatial metadata to geoscientific databases. The database is available on GitHub where the geoscience community can provide updates to improve the models and their contempo-

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raneity as new knowledge is acquired. The files are also released in formats suitable for use in Generic Mapping Tools and GoogleEarth.

Keywords: tectonic plate, tectonic province, orogenic system, orogeny, geodynamics, geospatial analysis

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1. Introduction

The structural architecture, tectonic environment, and temporal evolution of rocks at the surface of the Earth is frequently correlated with the chemical and physical characteristics of the enclosing lithosphere (Gard et al., 2019b; Artemieva, 2019; Tang et al., 2020; Tetley et al., 2020). As a result, it is useful to have spatially accurate maps of geologic provinces and terrane boundaries that encompass a pragmatically uniform set of common geological characteristics for comparative global studies. Such maps also form the foundation for accurate plate reconstructions (Merdith et al., 2021). While there are some regional models of tectonic provinces that are digital (Artemieva, 2006; Laske et al., 2013), there are few accurate global models easily accessible to the geoscience community built on a multiplicity of comparative attributes that approach self-consistency.

In this paper, we present two basic models: (1) a global set of geologic provinces and (2) a model for present-day plate boundaries. Both models are presented in a vector format with accompanying metadata that can be used to improve and simplify the process of global tectonic data analysis and/or modeling across a diverse range of geoscientific phenomena. These models have been produced using a wide variety of geologic and geophysical data and have been partially validated, wherever possible, using igneous and metamorphic age dates allied with additional geophysical datasets. Our hope is these models can be used as a data standard for common classification across the variety of geological databases that currently exist. The global models presented below are freely available in open-source and form a basic digital architecture that can be progressively updated as geological data and interpretations continue to improve.

2. Existing Global Models of Tectonic Plates and Provinces

Previous global plate and province models have been published that incorporate tectonic setting, juvenile age, or thermotectonic age (e.g., Artemieva, 2006; Goutorbe et al., 2011; Laske

25 [et al., 2013](#); [Szwilius et al., 2019](#)). Ideally, the digital nature of these maps makes them easy
26 to use and allows one to add desirable attributes to underlying datasets. However, the raster
27 format of these models is often an impediment to accurate spatial analysis at or near province
28 boundaries due to their low resolution and pixelated nature. Even though some plate and
29 province maps include age and province type, it may not be possible to separate individual
30 terranes that potentially have distinctive chemical and physical characteristics. For example, a
31 terrane-type map (e.g., [Laske et al., 2013](#)) can be useful for identifying data within a volcanic
32 arc setting, but it can be difficult to separate individual volcanic arcs to compare temporal and
33 spatial geochemical patterns that illuminate the geodynamic character. Furthermore, while
34 some of these maps have been included in peer-reviewed publications as part of global studies,
35 it is unclear to what degree the maps themselves have been examined in detail because the un-
36 derlying geological data used to construct the maps is not available to the geological community.
37 In some cases, it is difficult to obtain digital versions from the authors, making it challenging
38 to validate or improve the models as a geoscientific community and achieve widespread use.
39 As a consequence, some of these maps of global geological provinces are essentially artistic and
40 have opaque underlying rationale that cannot be interrogated.

41 A few global shapefiles of province polygons do exist ([Klett et al., 1997](#); [Torsvik and Cocks,](#)
42 [2016](#); [Matthews et al., 2016](#)), but they cannot be accurately matched to the province boundaries
43 as they are identified in regional studies. For example, the [Klett et al. \(1997\)](#) proposed geologic
44 provinces were developed to assess global hydrocarbon reserves. Therefore, a narrow perceived
45 attribute rather than broad actual attributes were used to delineate geological provinces. The
46 models by [Torsvik and Cocks \(2016\)](#), [Matthews et al. \(2016\)](#), and [Merdith et al. \(2021\)](#) are based
47 on tectonic blocks as they were developed to perform plate reconstructions (Figure 1b), but they
48 are typically shape-defined by contemporary global geography. Some terrane boundaries agree
49 well with published models (e.g., Africa, [McCourt et al. \(2013\)](#); South America, [Ibañez-Mejia](#)
50 [et al. \(2011\)](#)), while others appear greatly simplified and/or do not closely follow geophysical
51 trends for reasons that are not given (e.g., western United States, [Hasterok and Chapman](#)
52 [\(2007\)](#)). Furthermore, many models are time diluted. For example, models by [Matthews et al.](#)
53 [\(2016\)](#) and [Merdith et al. \(2021\)](#) were developed for reconstructions to 400 Ma and 1000 Ma,
54 respectively. As a result, most Mesoproterozoic and older regions are not divided into separate

55 terranes except where they behave as separate entities during the plate model timeframe, despite
56 a wealth of data that would otherwise allow the organization of these older participants to be
57 illuminated.

58 In addition to the lack of globally a comprehensive and self-consistent scheme for the de-
59 pication of geological provinces, there are still ambiguities and lack of self-consistency in con-
60 temporary plate tectonic maps. Bird (2003) released a widely used plate boundary shapefile
61 (Figure 1a), which built upon the pioneering work by Minster and Jordan (1978); DeMets
62 et al. (1990). Since its release, several additional microplates have been proposed, generally
63 on the basis of GPS motion data. The new proposed microplates are the Adria Microplate
64 (Battaglia et al., 2004; Breton et al., 2017), Danakil Microplate (Eagles et al., 2002; McClusky
65 et al., 2010), Yakutat Microplate (Fletcher and Freymueller, 1999; Bruhn et al., 2012), Sierra
66 Nevada Microplate (Dixon et al., 2000; Schweickert et al., 2004), several microplates around
67 the Caribbean (DeMets and Wiggins-Grandison, 2007; Sun et al., 2020), and a few microplates
68 on the Somali Plate (Horner-Johnson et al., 2005, 2007; Saria et al., 2014; Stamps et al., 2021)
69 and African Plate (Njoroge et al., 2015; Wedmore et al., 2021). The Bird model is also missing
70 a few microplates proposed at the time of publication including the Capricorn Microplate and
71 Indo-Australian Deformation zone (Royer and Gordon, 1997), the Coiba and more recently as-
72 sociated Malpelo Microplates (Hardy, 1991; Zhang et al., 2017). The Bird model included some
73 plate boundary deformation zones, but many zones of known deformation were excluded from
74 the model (e.g., Gordon, 1998). Additional improvements can be gained by incorporating more
75 recent kinematic models of strain within plate boundaries to discriminate between discrete and
76 more continuously deforming zones (Kreemer et al., 2014).

77 Beyond the global models discussed above, there exists a plethora of regional geologic
78 province and terrane maps that are downloadable. Many of these models are more precise
79 because they may incorporate a variety of datasets as constraints (e.g., topography, bedrock
80 geology, seismic tomography, and crustal faults or shear zones), but they are poorly designed
81 for digital processing as they are only available as raster images in papers despite many be-
82 ing created using geographic software. While these depictions are sufficient for presentation
83 in publications, the rationale and metadata are generally not available for scrutiny. Conse-
84 quently, interpreted tectonic boundaries are often inconsistent between publications, especially

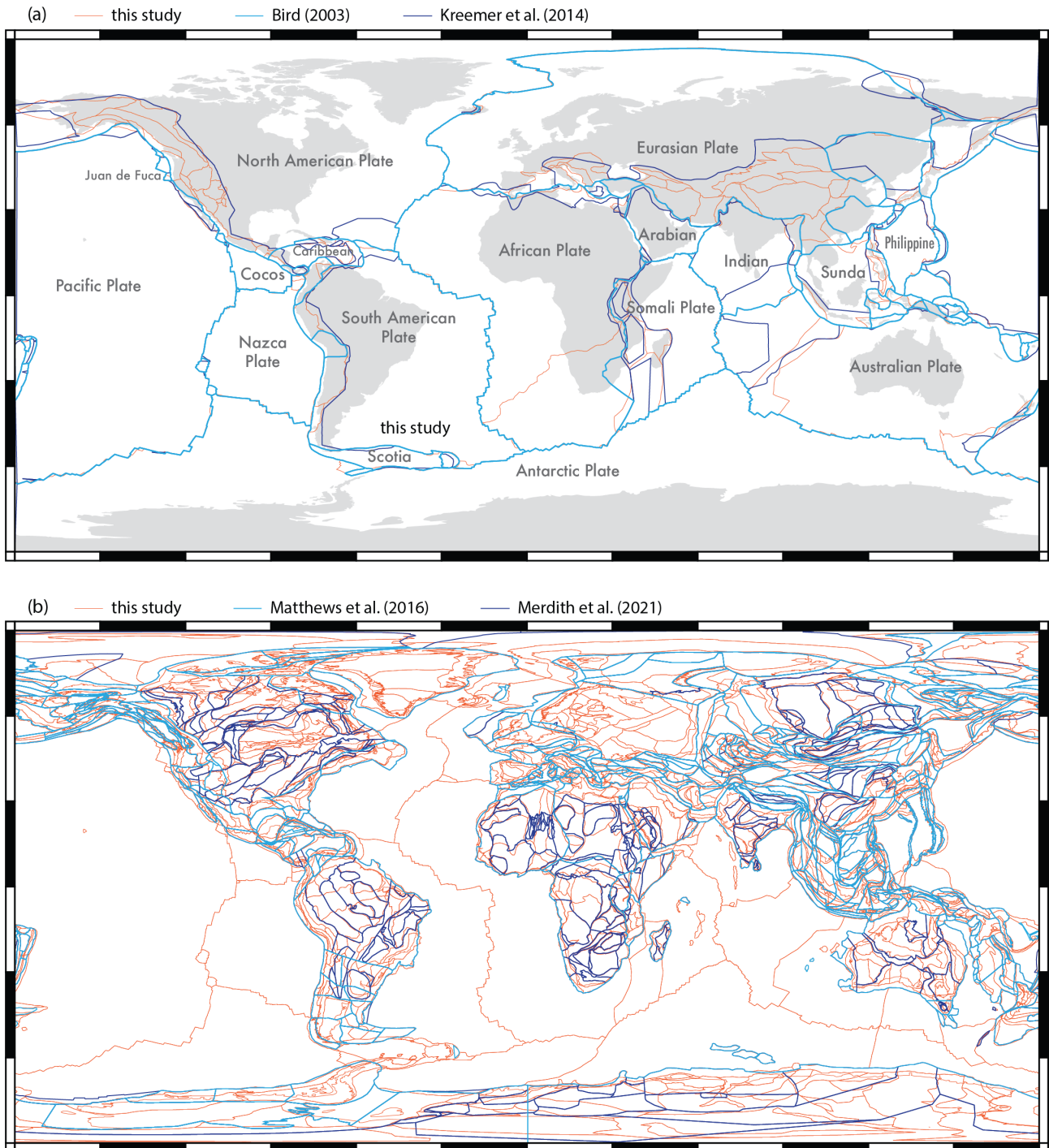


Figure 1: Tectonic plate and geologic province models. (a) A comparison between the Bird (2003) plate model, limit of modeled plate boundary zones (Kreemer et al., 2014) and the plate model from this study. Note that the limit of the deformation model domain is not an exact limit of the deformation, in many cases wider than the true deformation zone. (b) A comparison between the Matthews et al. (2016) and Merdith et al. (2021) plate models and province models from this study.

85 when boundaries are envisaged beneath cover. In many publications (e.g., [Gee and Stephen-](#)
86 [son, 2006](#); [Xu et al., 2016](#); [He et al., 2018](#)), a province or terrane map is often secondary to
87 the main point of the paper, and boundary definition may be schematic rather than rigorous
88 and is commonly simply adopted from pre-existing publications. An obvious example is the
89 depiction of the Sahara Metacraton (Figure 14 and Figure 1 of [Liégeois et al., 2013](#); [Kwékam](#)
90 [et al., 2020](#), respectively). Furthermore, the locations and level of detail lack continuity be-
91 tween publications, such that province boundaries cannot be easily matched up between maps
92 of adjacent regions, even when apparently depicted in the same map projection scheme (e.g.,
93 [wei Zhou and Murphy, 2005](#); [Blayney et al., 2019](#)).

94 The objective of the work we present here is to try and smooth deficiencies between existing
95 models for tectonic province and plate boundaries by delineating a set boundaries in a vectorized
96 format that incorporates attributes derived from geophysical, geochemical and geochronological
97 data. Our models represent an attempt to extend tectonic boundaries back to ca. 2.3 Ga
98 and possibly further in some cases, incorporating numerous existing geologic interpretations.
99 Some of these interpretations are modified slightly to meet additional observational constraints
100 from global geophysics, geologic maps and geochronology. Our goal is to create an adaptable
101 and interactive environment that allows the geoscience community to improve delineation of
102 geological provinces and the behaviors of their boundaries and interiors during Earth evolution.
103 In particular, we hope creating an interactive data environment will bring more illumination to
104 the Mesoproterozoic and older Earth, which hosts a significant fraction of metallic resources.

105 **3. Method of Construction**

106 The maps constructed in this study come from four separate shapefiles—also released in
107 GMT and KML formats for use in Generic Mapping Tools and GoogleEarth, respectively. These
108 files include the plate polygons, tectonic province polygons, the oceanic–continental crustal
109 boundary, and plate boundary types. The ocean–continent boundary and plate boundary
110 types are both developed in conjunction with the plate polygons. Each of these files contains a
111 number of attributes that include a variety of contextual information. The metadata for each
112 file are described in [Table 1](#).

113 Vector format shapefiles have several advantages over the raster maps that dominate the

Table 1: A description of shapefile attributes for the plate and province files released in this study.

Field name	Field description
<i>Plate polygons</i> (plates.shp)	
id	unique polygon identifier
poly_name	unique polygon name
plate	major plate, includes microplates and deformation zones
plate_id	numeric subplate id
plate_code	subplate abbreviation
subplate	separates microplates and deformation zones from major plates
plate_type	rigid plate, microplate, or deformation zone
crust_type	continental or oceanic crust
sea_name	name of ocean or sea
domain	oceanic domain for geochemical grouping purposes
area	polygon area in square kilometers
plate_ref	reference for initial plate polygon
<i>Plate boundaries</i> (plate_boundaries.shp)	
feature_id	unique boundary segment identifier
feature	name of boundary segment
type	type of boundary segment
plate1 and plate2	subplates on either side of boundary segment
level	assigns an integer value, 1 = major plate boundaries and 2 = minor plate boundaries
comment	specific comments about a boundary segment
length	length of segment in kilometers
<i>Ocean-continent boundary</i> (oc_boundary.shp)	
id	line segment identification number
length	length of segment in kilometers
<i>Province polygons</i> (geologic_provinces.shp)	
id	unique polygon identifier
prov_name	unique province name
prov_type	dominant tectonic character of a tectonic province
prov_group	name for multiple polygons with a shared geological history, may contain multiple tectonic styles
lastorogen	most recent significant orogenic event
continent	continent name if on a continent
crust_type	continental, oceanic, or transitional
area	polygon area in square kilometers
comment	specific comments about a province
prov_ref	reference for initial province polygon

114 existing literature. The polygons and lines created across the four shapefiles are seamless, i.e.,
115 they use common boundaries where geologically appropriate. Vectorized data permit multiple
116 attributes to be assigned to each polygon, line, or point, which can be unique. Raster models
117 allow only a single attribute per pixel, often requiring multiple maps (generally from separate
118 studies), which could lead to non-causal juxtaposition of such attributes. For example, slight
119 differences in the boundaries of different attributes from unrelated studies could result in oceanic
120 crust being incorporated into a continental orogenic belt or continental crust being excluded
121 from the same orogenic system. Vector format files are also typically more memory efficient
122 than raster images. Thus the seamless nature of the polygons in this project is a distinct
123 advantage when constructing physical and/or compositional models of the lithosphere.

124 We recognize the study presented here attempts to do on a global scale, what a large number
125 of studies have done at a variety of subordinate scales. Therefore we are aware of an element
126 of hypocrisy in being somewhat critical of continuity between existing studies. However our
127 goal isn't to be correct, it's to create an environment where decisions about tectonic provinces
128 and their boundaries are globally determined where possible, using the same style of data
129 sets and their interpretation. Where our study differs from predecessors, is the compiled data
130 used is freely available and adoptable by more informed practitioners because the models are
131 available on GitHub (https://github.com/dhasterok/global_tectonics) where community
132 users can correct errors and omission and propose refinements.

133 *3.1. Plate Model*

134 The plate model consists of two separate shapefiles including the plate polygons and the
135 boundary lines (Table 1; Figure 2). The plate shapefiles were created in QGIS using global
136 vector and raster datasets (Table 2). We used the widely-distributed model by Bird (2003)
137 as the initial plate boundaries to construct the shapefile (Figure 1a). Newer models that
138 incorporate proposed past plates and microplates for use with plate reconstruction software
139 such as GPlates were also used for reference (i.e., Zahirovic et al., 2014; Matthews et al., 2016).
140 Although Bird's model is an excellent framework, increased spatial coverage of GPS and an
141 additional 18 years of earthquake observations has improved our ability to recognize additional
142 microplates and identify presently deforming regions. Bird (2003) ignores deformation zones,
143 focusing on rigid plates and microplates. In contrast, our model includes these deformation

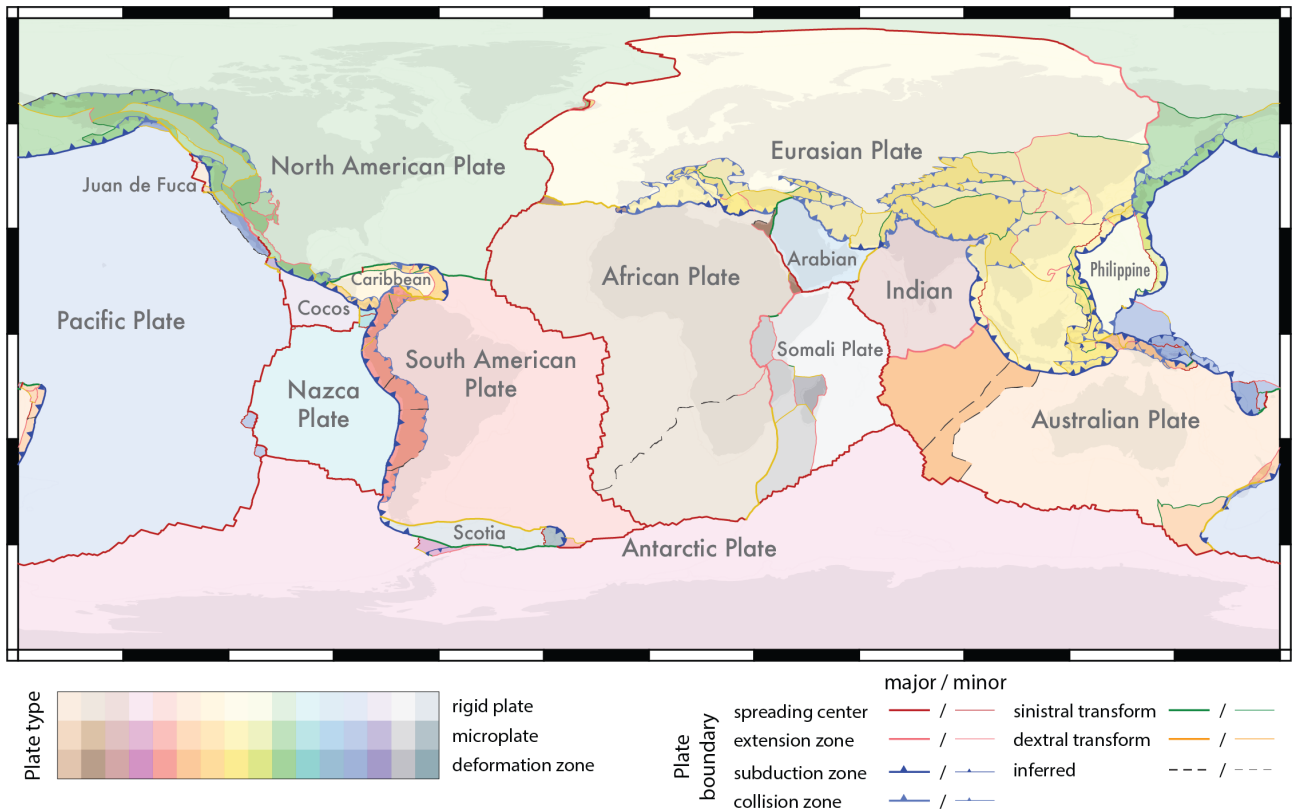


Figure 2: An updated plate model along with the plate boundary types. Microplates and deformation zones are illustrated as darker shades of the same hue as the major plate that they are most closely associated with. Note, geodetic studies often refer to the African Plate as the Nubian Plate, though we opt African as it appears to be more widespread. The case for a San Microplate (southern Africa) is weak, however, we include the proposed boundary in the shapefile (Section 4.3.3).

144 zones as part of the plate boundaries as they accommodate plate motion between the rigid
 145 plates (England and Molnar, 1997; Gordon, 1998; Freymueller, 2010; Kreemer et al., 2003).
 146 The strain-rate model by Kreemer et al. (2014) is also used as an initial constraint on the
 147 boundaries of microplates and deformation zones and as a way to distinguish between the two
 148 (Figure 1). Digital elevation models and global models of active faults have also improved
 149 since Bird’s model (Amante and Eakins, 2009; Styron and Pagani, 2020), which allow for more
 150 accurate and precise positioning of the boundaries.

151 Where there is ambiguity in the location of the plate boundary from topography and its
 152 gradient, the location was chosen to fit with the pattern of recent seismicity (Figure 3). We
 153 found it helpful to compute a spatial histogram of earthquakes because it is easier to identify
 154 zones of high seismicity relative to a simple scatter plot. To produce the histogram, we use global
 155 seismicity M5.5+ for 1970–1990 and M3.5+ for 1990–2020 extracted from the US Geological
 156 Survey’s Advanced National Seismic System (ANSS) global seismic catalog. The earthquake

Table 2: Datasets used to develop and evaluate the plate, plate boundary, and oceanic-continental boundary models.

Region	Data Type	Description	Resolution	Reference
global	plate model	shapefile		Bird (2003)
global	plate model	GPplates shapefile		Zahirovic et al. (2014)
global	plate model	GPplates shapefile		Matthews et al. (2016)
global	earthquakes	1990-2020, <30 km, M3-M5.5		ANSS (2020)
global	earthquakes	1970-2020, <30 km, M5.5+		ANSS (2020)
global	earthquakes	heat map (ANSS data above)	0.1 degree	Global Volcanism Program (2013)
global	volcanic centers			Amante and Eakins (2009)
global	topography	ETOPO1	1 arcmin	Seton et al. (2020)
global	topographic gradient ^a	derived from ETOPO1	1 arcmin	Debayle et al. (2016), Debayle et al. (2019)
global	seafloor age		1 arcmin	Styron and Pagani (2020)
global	shear wave tomography	EMC-3D2018_08Sv at 70 km depth	2 degree	Kreemer et al. (2014)
global	active faults	GEM GAF-DB		
global	GPS velocities	computed as a fixed-plate reference frame for each major plate	various	
global	strain rate	global model	0.25 degree	Kreemer et al. (2014)
Azores Microplate	GPS velocities	case study		DeMets et al. (2010)
Adria Microplate	GPS velocities	case study		Breton et al. (2017)
Alaska and northwest Canada	GPS velocities	case study		Elliott and Freymueller (2020)
Danakil Microplate	GPS velocities	case study		McClusky et al. (2010)
San Microplate	GPS velocities	case study		Njoroge et al. (2015), Wedmore et al. (2021)
Somali Plate	GPS velocities	case study		Saria et al. (2014), Stamps et al. (2021)
Greater Antilles	GPS velocities	case study		DeMets and Wiggins-Grandison (2007), Sun et al. (2020)
Lesser Antilles	magnetics and seismic	case study		Allen et al. (2019)
Coiba and Malpelo Microplates	GPS velocities	case study		Zhang et al. (2017)
New Zealand (North Island)	Seismicity	case study		Shi et al. (2019)
Sierra Nevada Microplate	GPS velocities	case study		Schweickert et al. (2004)
Philippines and East Indonesia	GPS velocities	case study		Zahirovic et al. (2014)
Yakutat Microplate	GPS velocities	case study		Bruhn et al. (2012)

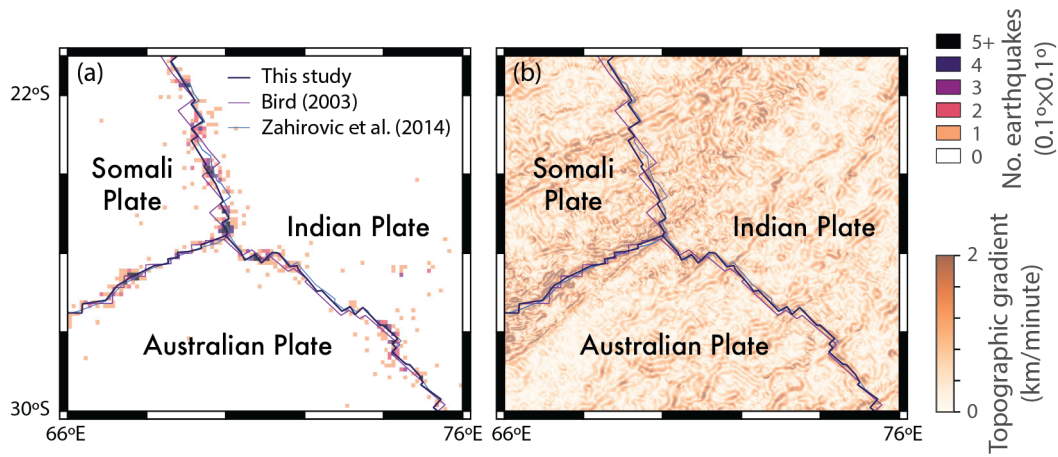


Figure 3: Refining locations of mid-ocean ridges relied heavily on the earthquake catalog, shown as histogram in (a), and topography along with its gradient (b).

157 density accurately traces out much of the mid-ocean ridge systems (Figure 3).

158 3.1.1. Plate Type

159 We include metadata with the plate model, including a `plate_type` field. The following
 160 definitions are used to distinguish plate types:

161 **rigid plate**, a region with distinct plate boundaries generally defined by seismicity, little in-
 162 ternal deformation, distinct motion relative to several other plates, and generally large
 163 area (median 14 million km²);

164 **microplate**, a region with distinct plate boundaries generally defined by seismicity, little in-
 165 ternal deformation, motion controlled by surrounding plates, and generally small area
 166 (median 0.28 million km²); and

167 **deformation zone**, a zone identified by GPS motions as distinct from, yet controlled by
 168 surrounding plates/microplates with significant internal deformation, seismicity, and gen-
 169 erally small area (median 0.25 million km²).

170 The microplates and deformation zones are associated with a parent plate for instances
 171 where a simpler model is required (Figure 2).

172 3.1.2. Plate Boundary Type

173 Plate boundaries are frequently formed by a series of several fault zones that accommodate
 174 a portion of the plate motion rather than a single discrete structure. However, there is often
 175 one fault zone that accommodates the majority of the motion, which is chosen as the boundary

176 for the plate and plate boundary shapefiles. This choice means that in some cases, the presently
177 most-seismically active fault may not mark the plate boundary as defined here, as it may not
178 have accumulated the greatest displacement. For some extension zones, the deformation is
179 distributed over hundreds of kilometers. In these cases, we chose the boundaries by the major
180 structures that bound the extension.

181 While most models of plate boundaries are limited to three types: convergent, divergent
182 and transform, we opt for six that provide greater contextual information. There are several
183 plate boundary types defined in our model:

184 **subduction zone**, convergent plate boundaries, kinematically active footwall, plate1 field is
185 the upper plate;

186 **collision zone**, convergent plate boundaries, kinematically active hanging wall, plate1 field is
187 the upper plate;

188 **spreading center**, divergent plate boundaries, type reserved for mid-ocean spreading centers
189 (includes transform segments);

190 **extension zone**, divergent plate boundaries, often associated with diffuse extension, plate1
191 field is the upper plate;

192 **dextral transform**, right-lateral transform boundary;

193 **sinistral transform**, left-lateral transform boundary; and

194 **inferred**, unknown boundary types, or location.

195 Our definitions are simplified as they do not include transpressional or transtensional styles,
196 which are classified most commonly as transform boundaries, but these may be considered in
197 future versions.

198 To construct the plate boundary shapefile, we converted the plate polygons to lines and
199 removed duplicate lines. We then split the boundaries so that each line segment represents the
200 boundary between two plates, microplates, or deformation zones (Figure 2). Plate attributes
201 including the boundary type are then added and stylized in QGIS. In order to ensure that
202 thrusts and subduction zone boundaries had the barbs displayed on the upper plate, some lines
203 were topologically reversed in direction.

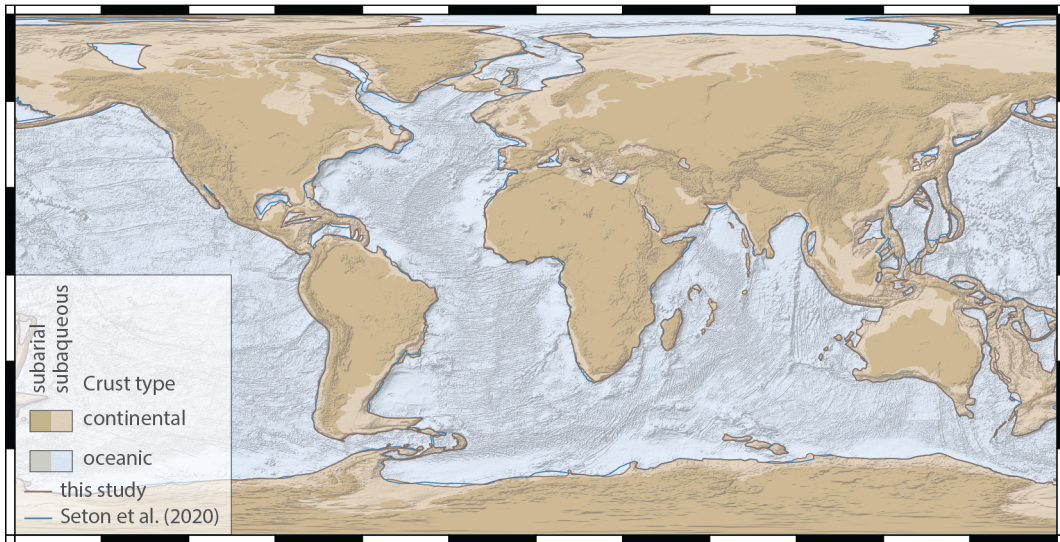


Figure 4: A model of continental and oceanic crustal domains. Plate boundaries as defined in Figure 2.

204 3.1.3. Ocean–Continent Boundary

205 We split the plate polygons into continental and oceanic parts because the ocean–continent
 206 boundary is colocated with many plate boundaries. For example, subduction zones generally
 207 form the boundary between plates as well as oceanic and continental regions. However, rifts and
 208 transform boundaries may form boundaries between plates but rarely form ocean–continental
 209 boundaries. Passive margins do not represent plate boundaries but do contain the ocean–
 210 continent transition.

211 A single depth contour cannot be used to identify the ocean–continent transition because the
 212 depth differs from region to region as a result of variations in sediment thickness and isostatic
 213 state. To create the initial model for the ocean–continent boundary, we created a polygon that
 214 defines the global distribution from the global seafloor age model (Seton et al., 2020). This
 215 initial polygon was then modified to provide a better match to the base of steep topographic
 216 gradients (continental slope) computed from ETOPO1 (Figure 4).

217 In reality, the ocean–continent boundary is rarely a sharp discontinuity, rather a transitional
 218 zone that can extend for 10’s to 100’s of kilometers. However, in many cases it is difficult to
 219 capture at the resolution of this model. Therefore, we place this line at the limit of the transition
 220 on the oceanic side. The province shapefile also includes an additional transitional crustal type
 221 where this boundary is sufficiently wide that it can be easily digitized, though we do not discuss
 222 it further.

223 Intra-oceanic volcanic arcs, such as Izu-Bonin and Marianas, are defined as continental

224 crust. Such arcs are often classified as oceanic crust because they are built upon crust created
225 by seafloor spreading. However, we have reserved the definition of oceanic crust for only
226 regions that have been created by seafloor spreading and not those chemically modified by arc
227 magmatism.

228 The ocean–continent boundary model could be improved with the addition of crustal thick-
229 ness and/or seismic velocity estimates across the ocean–continent transition. Presently, global
230 seismic models are of insufficient resolution to precisely identify the boundary. A compilation
231 of seismic profiles is beyond the scope of the present work.

232 *3.1.4. Oceanic Domain*

233 There appear to be differences in chemistry between ocean basins (domains). For example,
234 [Brandl et al. \(2013\)](#) documents a difference in basaltic geochemistry between the Atlantic
235 and Pacific Oceans, possibly related to temperatures of melt generation. Distinct chemical
236 signatures have long been recognized in the Indian Ocean ([Saha et al., 2020](#)). Back-arc basins,
237 separated from the major oceans by continental ribbons behind subduction zones, tend to
238 contain enriched basalts compared with mid-ocean ridge basalts ([Langmuir et al., 2006](#)). Thus,
239 it is desirable to have a way to quickly divide geoscientific data into these separate domains.

240 For the oceanic-type crustal polygons a mantle chemical domain is included in the plates
241 shapefile attributes table (Table 1). We have separated the oceans into nine separate domains
242 (Figure 5). The domains are intended to make separation of data from the different ocean
243 basins easier for geochemical and geophysical studies. However, the divisions are speculative
244 rather than data driven, specifically the exact boundaries of these chemical domains. The only
245 boundary that has been tested geochemically lies between the Indian and Pacific oceanic mantle
246 ([Pyle et al., 1992](#)). This study used isotopic analyses of Sr, Nd, and Pb to place the Indian–
247 Pacific boundary of Australia along the mid-ocean ridge at approximately 126°E. Similarly,
248 there is also evidence for a more complex chemistry in the seafloor basalts of the Philippine
249 plate compared to the rest of the Pacific ([Hickey-Vargas et al., 2006](#)).

250 *3.2. Global Geologic Province Model*

251 The initial province model was produced by creating a collage of overlapping geological
252 maps from the published literature that varied from the regional to continental scale (Table
253 4). Many of the images were georeferenced in GIS software and vectorized using polygon tools,

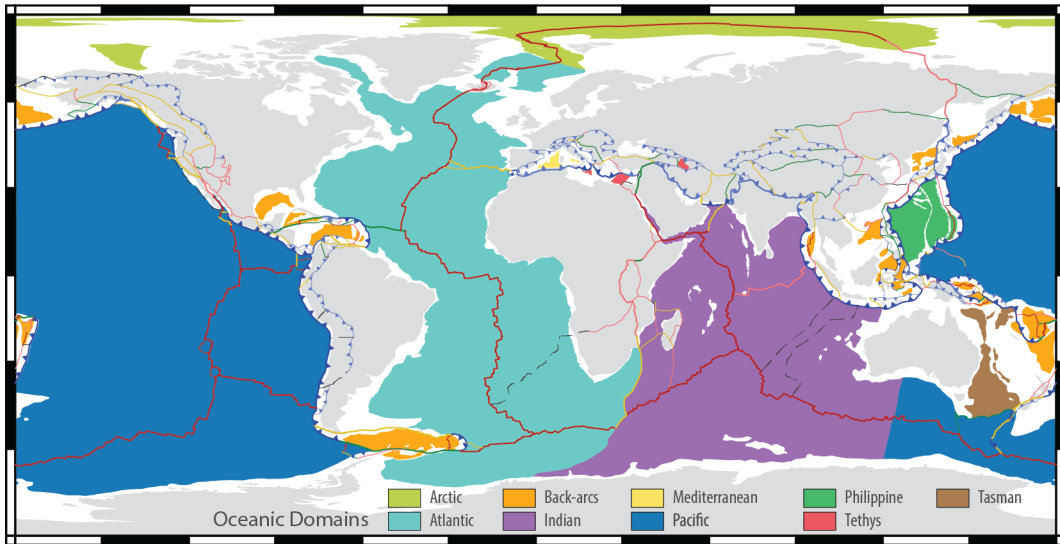


Figure 5: A hypothetical model of mantle chemical domains that source oceanic crust. At present these models are largely untested, but will be refined in future versions using global geochemical datasets.

254 while others were used as a visual reference. Province names are assigned based on commonly
 255 published terms, generally taken from the reference maps themselves (Table 4).

Table 3: Geological and geophysical models used to develop province model.

Region	Data type	Description	Resolution	Reference
global	topography	ETOPO1	1 arcmin	Amante and Eakins (2009)
global	composite gravity	GGM2012	2 arc-min	Balmino et al. (2011)
global	composite magnetics	EMAG2_V3	2 arc-min	Meyer and Salties (2016)
global	composite magnetics	WDMAMv2	5 km	Lesur et al. (2016)
global	seafloor age		2 arc-min	Seton et al. (2020)
global	volcanic centers			Global Volcanism Program (2013)
global	crustal thickness		1 degree	Szwilius et al. (2019)
global	active faults	GEM GAF-DB	various	Styron and Pagani (2020)
global	digital lithology	GLiM	0.5 degrees	Hartmann and Moosdorf (2012b), Hartmann and Moosdorf (2012a)
global	igneous dates	global geochemical database		Gard et al. (2019a)
global	metamorphic dates	metamorphic database	P-T	Brown and Johnson (2018), updated
global	metamorphic dates	DateView		Eglinton (2004)
Africa	surface geology and geologic provinces	USGS WEP ^a	1:5,000,000	Persits et al. (1997a)
Arabian Peninsula	surface geology and geologic provinces	USGS WEP ^a	1:2,000,000	Pollastro et al. (1999)
South Asia (India)	surface geology and geologic provinces	USGS WEP ^a	1:5,000,000	Wandrey and Law (1998)
South America	surface geology and geologic provinces	USGS WEP ^a	1:7,500,000	Schenk et al. (1999)
South America	surface geology	CGMW ^b	1:5,000,000	Gómez-Tapias et al. (2019)
Former Soviet Union	surface geology and geologic provinces	USGS WEP ^a	1:7,500,000	Persits et al. (1997b)
China, Southeast Asia, and Australia	surface geology and geologic provinces	USGS WEP ^a	1:7,500,000	Steinshouer et al. (1999)
Iran	surface geology and geologic provinces	USGS WEP ^a	1:2,500,000	Pollastro et al. (1997)
Europe	surface geology and geologic provinces	USGS WEP ^a	1:3,000,000	Pawlewicz et al. (1997)
North America	surface geology provinces	USGS	1:7,500,000	Garity and Soller (2009)
Australia	magnetic anomaly		250 m	Australia (2004)
Australia	gravity anomaly		830 m	Wynne and Bacchin (2009)

Table 3: continued.

Region	Data type	Description	Resolution	Reference
Antarctica	bedrock topography	BEDMAP2	1 km	Fretwell et al. (2013)
Antarctica	free-air and Bouguer gravity anomalies	ANTGG	10 km	Scheinert et al. (2016)
Antarctica	shear wave tomography	AN1	4 arcmin	An et al. (2015)
Antarctica	magnetic anomaly	ADMAP2	7 km	Golynsky et al. (2018)
Antarctica	mantle gravity anomaly		5 km	Baranov et al. (2017)
Antarctica	crustal thickness	gravity and seismic based	1 degree	Baranov et al. (2017)
Antarctica	crustal thickness	satellite gravity based	0.25 degrees	Llubes et al. (2018)
Antarctica	crustal thickness	seismic methods	4 arcmin	An et al. (2015)
India	Bouguer gravity map		3 arc-min	Geological Survey of India (2006)

^aUnited States Geological Survey (USGS) World Energy Project (WEP).

^bCommission for the Geological Map of the World (CGMW)

Table 4: Published province models used to construct the initial global model.

Region	Reference
	<i>Global</i>
large igneous provinces	Johansson et al. (2018)
modern passive margins	Berndt et al. (2019)
	<i>Africa</i>
Africa	Begg et al. (2009), Hinsbergen et al. (2011)
West Africa	Ennih and Liégeois (2008)
Sahara Metacraton	Liégeois et al. (2013); Şengör et al. (2020)
Mozambique Belt	Chauúque et al. (2019), Goscombe et al. (2020)
Central Africa	Jelsma et al. (2018)
Southern Africa	McCourt et al. (2013), Hanson (2003)
Madagascar	Collins et al. (2003)
	<i>Antarctica</i>
Antarctica	Harley et al. (2013), Stål et al. (2019)
East Antarctica	Golynsky (2007), Harley and Kelly (2007), Elliot et al. (2015), Leitchenkov et al. (2016), Pierce et al. (2014), Pant and Dasgupta (2017), Mulder et al. (2019), Ruppel et al. (2020), Flowerdew et al. (2013), Aitken et al. (2014), Maritati et al. (2016), Maritati et al. (2019), Wang et al. (2020), Ebbing et al. (2021), Jacobs et al. (2015), Dunkley et al. (2020)
West Antarctica	Jordan et al. (2020)
	<i>Asia</i>
Siberian Craton	Tretiakova et al. (2017)
West Siberian Basin	Cherepanova et al. (2013)
Russia, far east	Isbell et al. (2016)
Central Asian Orogenic Belt	Xiao et al. (2015), Janoušek et al. (2018), Ivanov et al. (2014), Windley et al. (2007), Buslov et al. (2001)
North China Craton	Liu et al. (2017)
South China Craton	Wang et al. (2013)
Tian Shan Belts	Charvet et al. (2011)
Tibetan plateau	wei Zhou and Murphy (2005), Blayney et al. (2019)
southeast Asia	Mitchell et al. (2012), Burrett et al. (2014), Zhang et al. (2019), Morley and Searle (2017), Dew et al. (2021)
	<i>Australia and Zealandia</i>
Australia	Foster and Goscombe (2013), Pilia et al. (2015), Abdullah and Rosenbaum (2018)
Zealandia	Stagg et al. (2002), Gallais et al. (2019)
New Zealand	Mortimer (2004)
	<i>Europe</i>
Baltic Shield	Bogdanova et al. (2015), Zhao et al. (2002)
Mediterranean Europe	Schmid et al. (2020)
western Europe	Topuz et al. (2020)
	<i>India and Middle East</i>

Table 4: continued.

Region	Reference
Arabian-Nubian Shield	Johnson (2014)
Iran	Naimi-Ghassabian et al. (2018)
Pakistan	Kazmi and Rana (1982)
India	French et al. (2008)
Sri Lanka	Cooray (1994)
	<i>North America</i>
United States and Canada	Whitmeyer and Karlstrom (2007), Hasterok and Chapman (2007), Lund et al. (2015), Ontario Geological Survey (2011), Berman et al. (2013b), Fyffe et al. (2012), Bjorkman (2017), Linde et al. (2017)
Greenland	White et al. (2016)
Alaska and Canadian Cordillera	Colpron and Nelson (2011)
Mexico	Sedlock et al. (1993)
Caribbean and Gulf of Mexico	Allen et al. (2019); Davison et al. (2020)
	<i>South America</i>
South America	Ibañez-Mejia et al. (2011), Chew et al. (2008), Egydio-Silva et al. (2018)
western South America	Ramos and Aleman (2000), Eude et al. (2015), Charrier et al. (2014)

Each boundary was then adjusted for seamless fits in the global model using a combination of geologic maps, active fault databases, geochronology, topography, and geophysical anomalies/models such as gravity, magnetics, and crustal thickness (Table 3). Magnetic anomalies were the most useful geological dataset for developing the province model as the magnitude and visual character is often distinctive within a province (Figure 6). We used two global magnetic anomaly models to aid with locating province boundaries. Both are constructed from a combination of airborne and shiptrack magnetics. Because the airborne and ship track data are recorded at different altitudes and line spacing, this leads to variations in resolution and gaps filled with very low-resolution satellite observations. Where the resolution is low, it is more difficult to precisely position the province boundaries (Figure 6), , requiring more emphasis be placed on the other datasets listed in Table 3.

Topography is helpful for identifying the boundaries of many provinces as fault and shear zones are often expressed topographically. Topographic features are most helpful in active terranes, but many ancient terranes can also be delineated by the changes in morphology when the faults/shear zones are no longer immediately apparent. The relatively high resolution of ETOPO1 (~ 1.85 km at the equator) makes the positioning of boundaries reasonably precise.

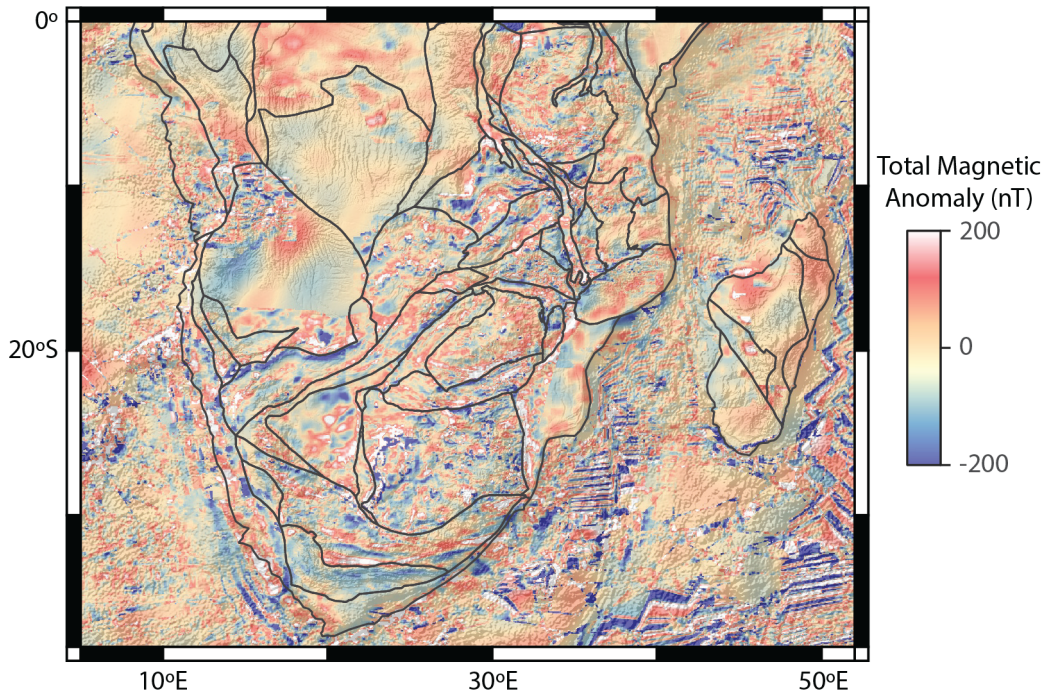


Figure 6: Magnetic anomaly map of southern Africa overlaying topographic relief. The WDMAM_V2 magnetic model has a pixel size of 1.5 minutes, which is approximately 2.8 km at the equator. The model is created from shiptrack, airborne and satellite datasets resulting in variable resolution that is evident in this image and affects the accuracy of province boundaries.

272 When there is a clear topographic expression at the boundary between provinces, the uncer-
 273 tainties are probably on the order of 3 to 5 pixels (~ 5 to 10 km).

274 We incorporate a number of attributes ascribed to the province polygons including a province
 275 name, province group, tectonic province type, and last significant orogeny (Table 1). The data
 276 standards used to define these additional attributes are described below.

277 3.2.1. Province Type

278 A geologic terrane captures a set of geologic units that describe a coherent block of crust
 279 with a shared geologic history, which can include tectonic setting, magmatic history, and/or
 280 metamorphic evolution. While a terrane is less fundamental than a geologic unit or suite,
 281 there are often similarities in the physical architecture and chemistry of terranes created in
 282 similar tectonic settings (e.g., [Ducea et al., 2015](#); [Ueki et al., 2018](#)). However, there are also
 283 differences between terranes of similar types that can be uniquely expressed in the architecture,
 284 composition, and thermal history (e.g., [Furman, 2007](#); [Dilek and Furnes, 2014](#); [Profeta et al.,](#)
 285 [2015](#)).

286 To facilitate the analysis of terranes, we include a basic province type attribute with the
 287 shapefile. The province definitions are based on the terrane type that covers the majority of

288 a polygon. Most of the terrane definitions we have chosen distinct characteristics within a
289 modern plate tectonic setting:

290 **craton**, predominantly Archean core, contains granite-greenstone belts and other undifferen-
291 tiated terranes with relatively small area;

292 **shield**, similar to a craton, predominantly Meso- to Paleoproterozoic lithosphere, undifferen-
293 tiated;

294 **passive margin**, sediment accumulation built on transitional crust between continental and
295 oceanic crust marking half of a tectonically inactive fossil rift;

296 **accretionary complex**, active/subduction margin consisting of sedimentary wedges built on
297 oceanic or continental crust;

298 **basin**, intracontinental sedimentary cover built on preexisting continental crust with uncertain
299 or unknown basement provenance;

300 **foredeep basin**, (foreland basin) thick intracontinental sedimentary basin created during continent-
301 continent collision, basement uncertain;

302 **orogenic belt**, fold and thrust belts created during accretionary, collisional and intracontinen-
303 tal settings that may incorporate a variety of preexisting terrane types, often commingled,
304 making them difficult to differentiate at the regional scale;

305 **narrow rift**, focused extensional terrane with continental basement;

306 **wide rift**, distributed extensional terrane with continental basement;

307 **volcanic arc**, predominantly magmatic arc crust related to subduction, but may contain crust
308 predating the arc and/or interspersed accretionary material in island arcs and in seaward
309 migrating arcs due to retreating trenches;

310 **continental back-arc basin**, a hyper-extended basin either transitional continental or oceanic
311 crust created as a result of upper plate extension in response to subduction rollback;

312 **oceanic back-arc basin**, a back-arc basin where seafloor spreading has been sustained, cre-
313 ating enriched basaltic compositions relative to mid-ocean ridge basalt;

314 **ophiolite complex**, obducted oceanic crust of some variety, excluding volcanic arc-type, but
315 including supra-subduction zone oceanic crust;

316 **magmatic province**, a large intraplate magmatic terrane not clearly associated with subduc-
317 tion or extension processes; and

318 **oceanic crust**, typical oceanic crust not created in a back-arc setting.

319 The focus of the geologic province model is on basement tectonic terranes, hence large
320 igneous provinces (LIPs) are not included in this model as they are superimposed on the
321 basement. LIPs are volcanic features often associated with mantle plumes rather than tectonics.
322 Our model is a tectonic model. We do not discount LIPs as an important aspect of crustal
323 and mantle evolution, but they can obscure the underlying tectonic terranes. In fact, many
324 old LIPs whose volcanics have long been eroded away are recognized only by dike swarms that
325 occupy a fraction of the original surface area (e.g., [Ernst and Bleeker, 2010](#); [Ciborowski et al.,](#)
326 [2015](#)). These older terranes retain much of their prior tectonic character beneath the volcanics.
327 Since a good model for LIPs currently exists ([Johansson et al., 2018](#)) so there is no need to
328 recreate one as a separate layer as part of this project. Users who wish to include LIPs can
329 easily incorporate them any spatial analysis.

330 Basins have been ignored except where the character of the underlying basement is unknown
331 (Figure 7). A reasonable resolution (5-arc-minute) sediment thickness model by [Straume et al.](#)
332 ([2019](#)) covers the ocean basins. Australia is covered by a high-resolution, 15-arc-second basin
333 model ([Geognostics, 2021](#)). A 30-arc-second sediment thickness model by [Pelletier et al. \(2016\)](#)
334 covers all continents except Antarctica, but only provides values for regions with less than 50
335 m of sediment thickness. The only available global basin thickness model has a relatively low
336 resolution of $1^\circ \times 1^\circ$ ([Laske and Masters, 1997](#)).

337 3.2.2. Last Orogeny

338 The most recent (last) high-temperature orogenic event to affect a province often has an en-
339 during influence on the present day thermal and physical state of the lithosphere and therefore,
340 its future potential to deform, metamorphose and melt (e.g., [Sandiford et al., 2001](#); [Fossen](#)
341 [et al., 2017](#); [Hyndman, 2019](#)). Here we define the last orogeny as the most recent regional
342 high-temperature thermotectonic event, excluding regions that may have experienced plume-

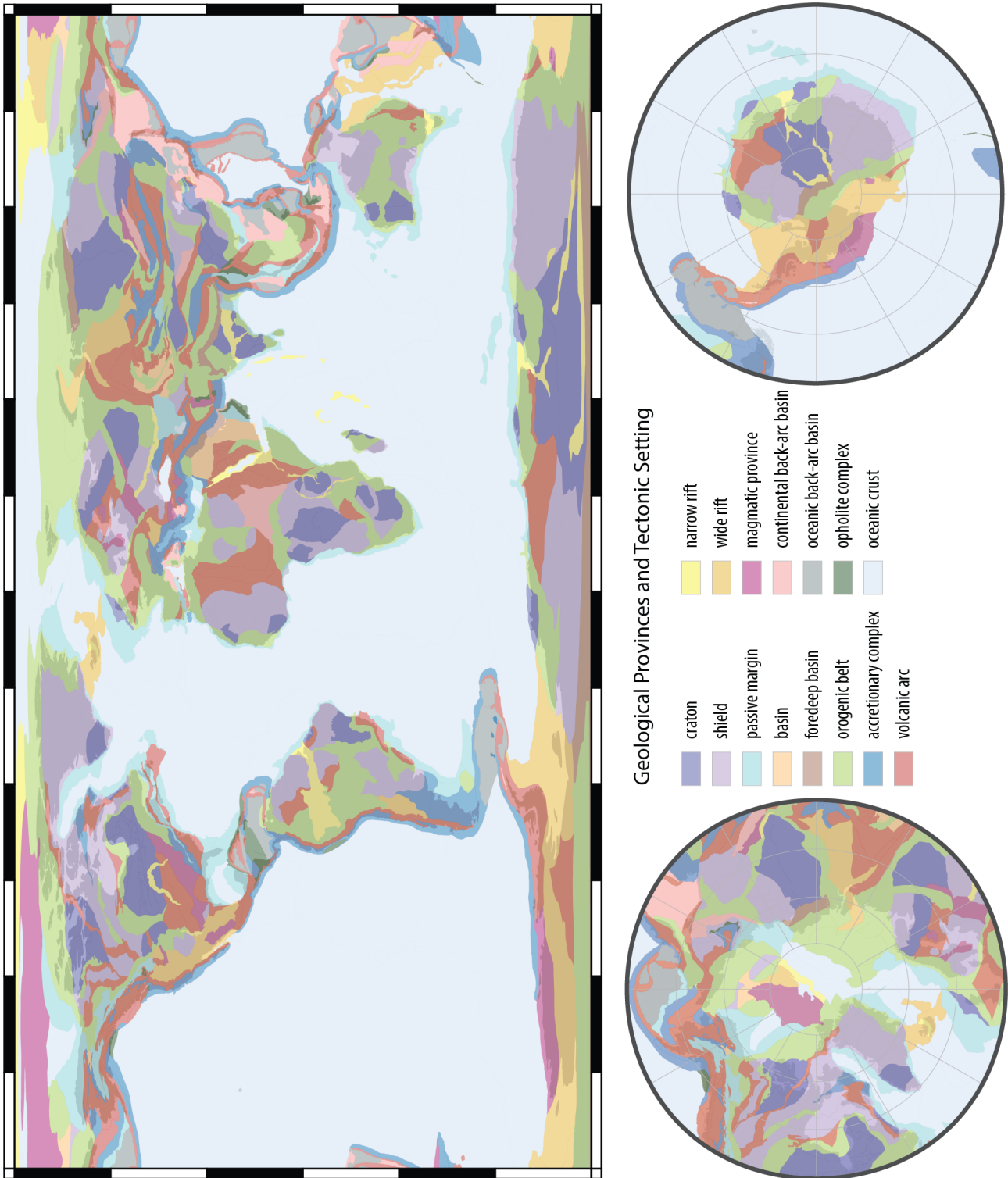


Figure 7: A global model of geological provinces with similar tectonic and compositional histories. Provinces are colored by their dominant tectonic setting. Though the setting may change with time, the provinces are defined by the environment which dominates the majority of rocks found within.

343 related activity. The most recent thermotectonic event is generally correlated with elevated
344 surface heat flow (Lucazeau, 2019), and lithospheric buoyancy (Fischer, 2002), depending on
345 its thermal intensity, it may reset high-temperature thermochronometers (Wan et al., 2011).

346 Orogenic systems often span a few hundred million years and comprise multiple, smaller
347 orogenic events that exhibit significant regional variability (e.g., Ge et al., 2014; de Gromard
348 et al., 2019). For example, orogenic activity may propagate along a system over time as
349 exhibited by protracted continent–continent collision as in the Alpine Himalayan Belt (Kuhnt
350 et al., 2004; Dilek, 2006; Ustaszewski et al., 2010; Hu et al., 2016; Symeou et al., 2018; An et al.,
351 2021), hence the last orogeny descriptor is not as finely resolved in age as the activity in any
352 given region. Instead, we use the last orogeny term to represent long-lived tectonic/geodynamic
353 systems. These descriptors are often related to consumption of ocean basins \pm continental
354 collision; though they can also apply to an intraplate orogeny. In the Phanerozoic, the divisions
355 are generally well-defined, but in the Precambrian, the connection between orogenic systems
356 and now-isolated provinces may be less certain (e.g., Li et al., 2008). In these cases, the names
357 refer to periods of orogenic activity rather than discrete systems. While this somewhat blurs
358 the meaning of the term, we prefer it over several colloquial orogenic names. It also represents
359 a research opportunity for improving models of orogenies and more accurately capture multiple
360 distinct systems that may overlap in age.

361 The last orogeny model is built from reviews and large-scale studies of orogenic systems and
362 plate reconstructions, assigning a single orogeny to each province polygon (Figure 8). We then
363 validate the orogeny model against databases of igneous and metamorphic dates. Recently,
364 Condie et al. (2021) attempted to quantify orogens and link them to other global processes by
365 using the ‘number of orogens’ as a measurable quantity. We suggest that this is too arbitrary
366 a parameter and instead we have attempted an ‘orogenic province’ approach where we suggest
367 linking orogens and orogenies based on their interpreted tectonic/geodynamic system. To take
368 the modern Earth as an example, this approach then links the active circum-Pacific orogens into
369 one orogenic province, but separates them from the Alpine-Himalayan system. Understandably,
370 this gets more subjective in deep time, but it provides a framework for building geodynamic
371 models and presents hypotheses that can be robustly tested by new observations.

372 In the Phanerozoic, the divisions are generally well-defined, but in the Precambrian, the

Last Orogeny

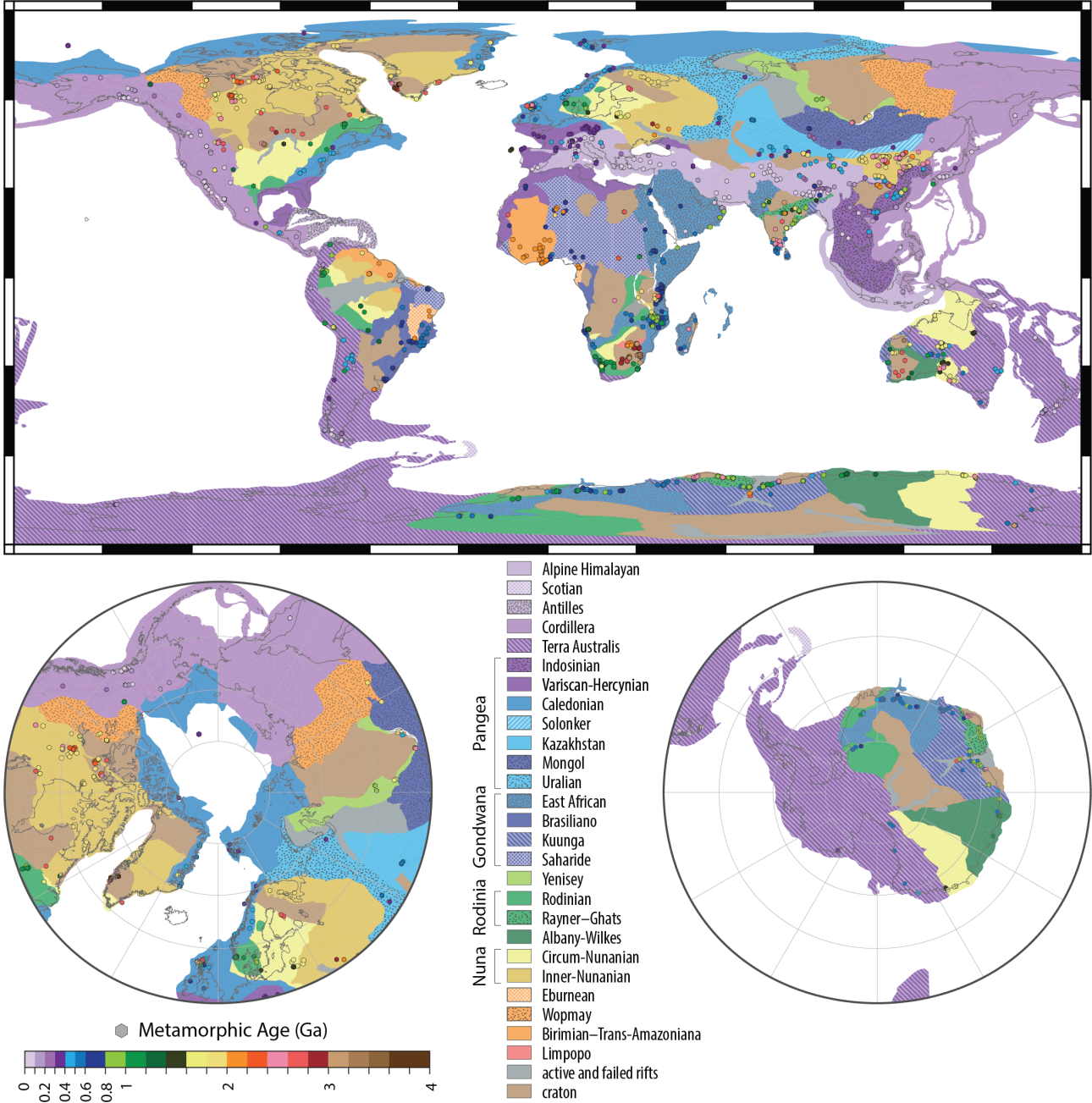


Figure 8: A map of the last orogenic event to affect a province. The colors are approximately related to the age of the orogenic event. The points show locations of observed and estimated metamorphic ages (Brown and Johnson (2018), updated; Eglington (2004)). Classification of the Phanerozoic orogens follows a systems approach as discussed in the main text.

373 connection between orogenic systems and now-isolated provinces is often less certain (e.g. [Li](#)
374 [et al., 2008](#)). We have used the emerging full-plate tectonic reconstructions of [Merdith et al.](#)
375 [\(2021\)](#) and [Cao et al. \(in prep.\)](#) to guide us here (Figure 9). While in many cases we believe
376 these were coherent systems, the names may refer to periods of orogenic activity rather than
377 discrete systems (e.g., Siberian Orogeny as defined below). While this somewhat blurs the
378 meaning of the term, we prefer it over several discrete orogenies with colloquial names. It
379 also represents a research opportunity for improving models of orogenies and more accurately
380 capture multiple distinct systems that may overlap in age. We define each of the orogens below.

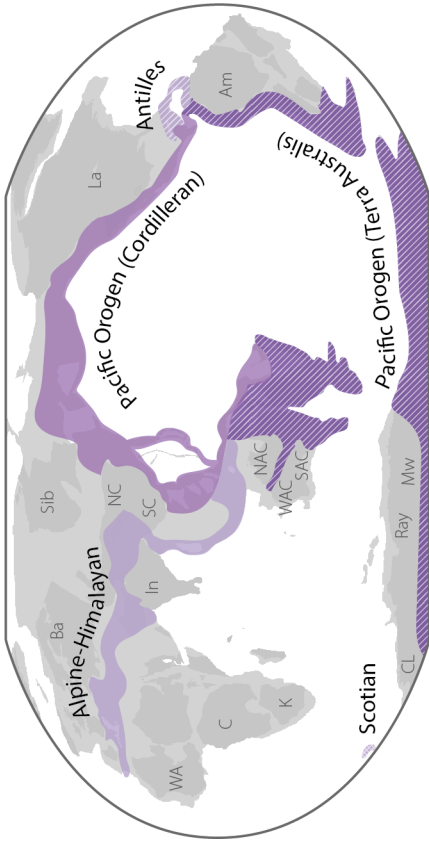
381 *3.2.2.1. Active Orogens*

382 . The Alpine-Himalayan Orogeny includes the collision between several plates with Eurasia,
383 which began ca. 65 Ma. The orogeny was initially driven by the subduction of the Tethys
384 Ocean beneath Eurasia but has continued even as the ocean has closed in many regions. The
385 continent–continent collisions with Eurasia were heterogeneous in time, beginning with the
386 collision of Apulia with Europe ca. 65 Ma ([Ustaszewski et al., 2010](#)), India with Tibet ca. 61–
387 59 Ma (earliest suggested timing; [Hu et al., 2016](#); [An et al., 2021](#)), Australia with Indonesia ca.
388 25 Ma ([Kuhnt et al., 2004](#)), and Arabia with Iran ca. 25 Ma ([McQuarrie and van Hinsbergen,](#)
389 [2013](#); [Gaina et al., 2015](#)).

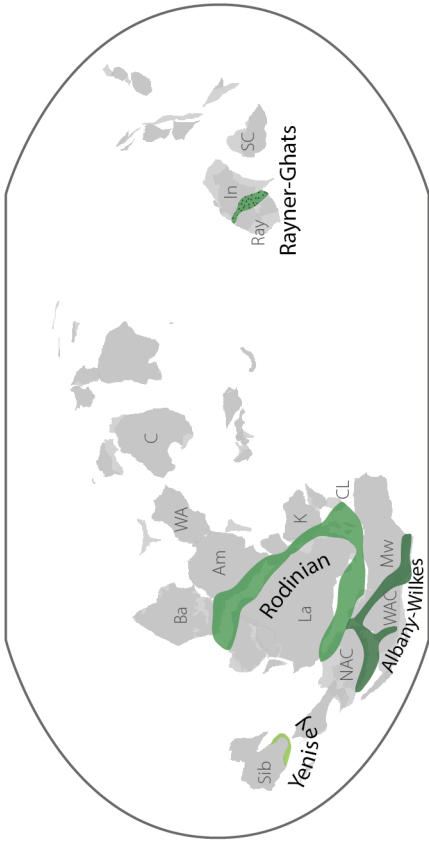
390 The Scotian Orogeny began with the initiation of subduction along the Antarctic and South
391 American Plate margins ca. 80 Ma and continues to the present day beneath the Sandwich
392 Islands ([Eagles, 2016](#); [van de Lagemaat et al., 2021](#)). The Scotian Orogeny (Figure 8 and 9a),
393 like the Antilles Orogeny (below), is geodynamically governed by a retreating subduction zone
394 consuming the southwest Atlantic, which has created the Scotia Plate and Sandwich Microplate
395 in its wake.

396 The Antilles Orogeny is a young (ca. 118 Ma to present), active orogenic system in the
397 Caribbean that began in the mid-Cretaceous and created the Caribbean Plate as a result of
398 rapid trench retreat ([García Casco et al., 2006](#)). An alternative explanation by [Whattam and](#)
399 [Stern \(2015\)](#) suggests plume-induced thinning of the upper plate promoted subduction initiation
400 in the Caribbean ca. 95 to 85 Ma. The orogen is responsible for the creation of three separate
401 arc systems: the Greater Antilles, Lesser Antilles and Aves Ridge (Figure 8 [García Casco et al.,](#)
402 [2006](#); [Neill et al., 2011](#); [Allen et al., 2019](#)).

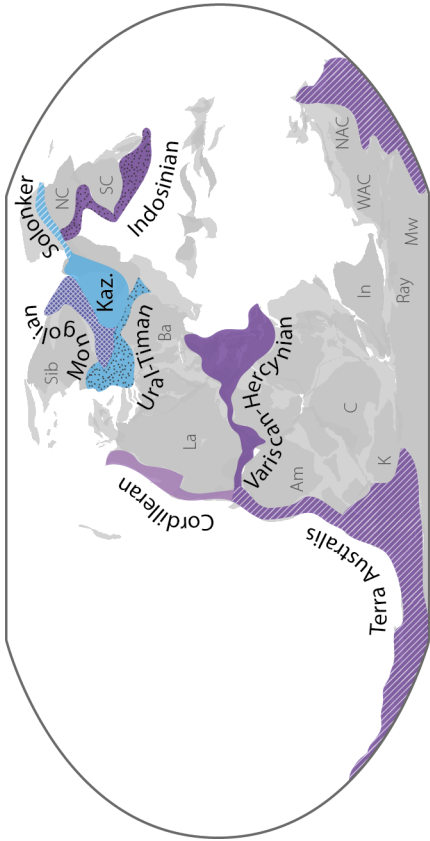
(a) Present Day



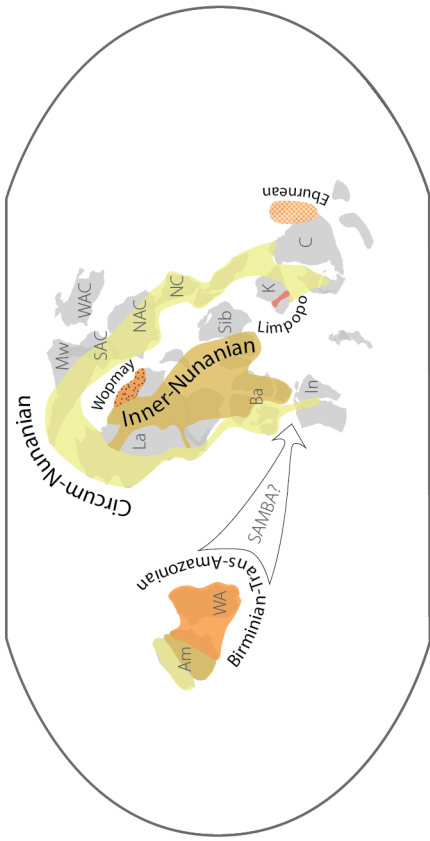
(d) 900 Ma



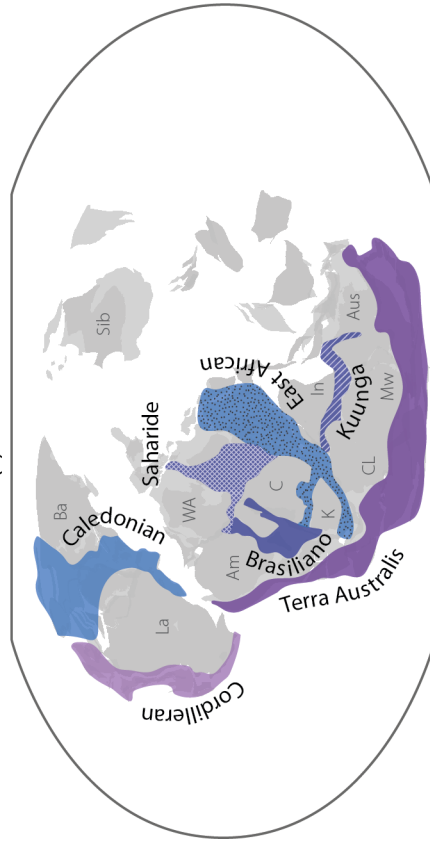
(b) 210 Ma



(e) 1450 Ma



(c) 400 Ma



(f) 2410 Ma

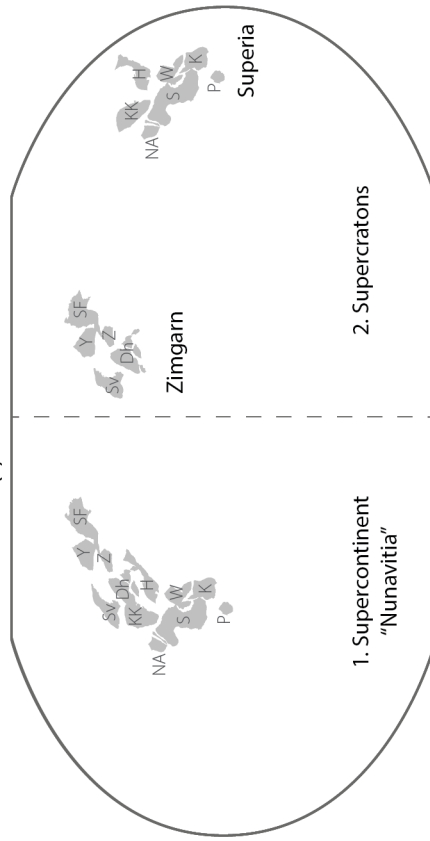


Figure 9: Reconstructions of orogenic systems at key dates. (a) Present-day shows the five active orogens. (b) Post-Pangea assembly, 210 Ma, keeping the position of Antarctica fixed. (c) Post-Gondwana assembly, 400 Ma, keeping the position of Amazonia fixed. (d) Post-Rodinia assembly, keeping Amazonia fixed. (e) Post-Nuna/Columbia assembly, 1450 Ma. Arrow indicates the alternate location of Amazonia and the West African Craton in the SAMBA reconstructions (see Section 3.2.2.3 for discussion). (f) Supercontinent model for Nunavutia at 2410 Ma on the left and supercratons model on the right. Reconstructions (a-d) from [Merdith et al. \(2021\)](#), (e) from [Cao et al. \(in prep.\)](#), and (f) from [Li et al. \(2021\)](#) with terranes that were connected at 2.3 Ga: Gr nehogna Craton with the Kaapvaal Craton, Penokean Orogen, North Atlantic Craton and the Nain Province with the Superior Craton; Napier Complex with the East Dharwar Craton; and the Mantiqera Province with the Sa  Francisco Craton. Abbreviations are as follows: Ba, Baltica; C, Congo Craton; CL, Coats Land; Dh, Dharwar Craton, H, Hearne Province; In, India; K, Kaapvaal/Kalahari Craton; KK, Kola-Karelia; La, Laurentia; Mw, Mawson Craton; NA, North Atlantic Craton; NAC, North Australian Craton; NC, North China; P, Pilbara Craton; Ray, Rayner Complex; S, Superior; SAC, South Australian Craton; SC, South China; SF, Sa  Francisco Craton, Sib, Siberia; Sv, Slave Craton; VU, Volgo-Ukranian Shield; W, Wyoming; WA, West Africa; WAC, West Australian Craton; Y, Yilgarn Craton; and Z, Zimbabwe Craton.

403 The Pacific Orogeny is defined by a set of circum-Pacific subduction zones associated with
404 the destruction of the Pacific, Philippine Sea, Cocos and Nazca Plates beneath the South
405 American, North American, Eurasian, and Australian Plates (Figure 9a to c). In the modern
406 lexicon, this system is often referred to as the Cordilleran Orogeny ([Dickinson, 2004](#)); however,
407 this is too simplistic as the Cordilleran Orogeny either merged with or grew out of the Terra
408 Australis Orogeny ([Muttoni et al., 2003](#)), which is defined by the subduction of the Panthalassic
409 Ocean beneath Gondwana ([Cawood, 2005](#)). We have kept these two orogens separate to preserve
410 the association of Terra Australis with Gondwana, but acknowledge there is little reason to
411 distinguish them geodynamically.

412 The Terra Australis Orogeny began ca. 530 to 520 Ma in response to subduction of the
413 Pacific along the Gondwana margin (Figure 9c; [Cawood, 2005](#); [Chew et al., 2007](#); [Paulsen et al.,](#)
414 [2020](#)), which continues through to the present day ([Glen et al., 2016](#)). The orogen includes a
415 series of alternating subduction-related back-arc extensional and collisional events that built
416 eastern Australia, Zealandia, and much of the Transantarctic Mountains ([Fergusson and Hen-](#)
417 [derson, 2015](#)). Because the Terra Australis Orogen involved phases of significant extension
418 ([Gaina et al., 1998](#); [Abdullah and Rosenbaum, 2018](#); [Jessop et al., 2019](#)), parts of the oro-
419 gen are no longer active and/or have been separated by seafloor spreading (i.e., Delamerian,
420 Thompson and New England orogens of Australia, Patagonia in South America, Ross Orogen
421 in Antarctica, and the Cape Fold Belt in South Africa). Some authors consider parts of the
422 Terra Australis Orogen to be part of the Cordilleran (Ross Orogen in Antarctica and Southern
423 Alpine Orogen in New Zealand, [Tagami and Hasebe, 1999](#); [Dickinson, 2004](#)) or the Alpine-
424 Himalayan Belt (Zealandia, [Lister et al., 2001](#)). However, we associate these regions with the

425 Terra Australis Orogeny due to (1) their former common Gondwana association and (2) the
426 orogenic system largely predates both the Alpine-Himalayan and Cordilleran Orogenies. In the
427 eastern South Pacific, South America is moving westward over the young oceanic lithosphere
428 (Schepers et al., 2017). The modern Andes are ca. 66 Ma old (Capitanio et al., 2011), reach-
429 ing their current heights ca. 14 Ma (Evenstar et al., 2015); however, some estimates suggest
430 subduction formed a volcanic arc in South America by 530 Ma (Chew et al., 2007).

431 The Cordilleran Orogeny from Canada to Mexico is long-lived, starting in the late Devonian
432 (ca. 370 Ma) with the collision of the Antler Orogeny and subsequently the Sonoma Orogeny in
433 the Triassic along the western North American margin (Figure 9c; Dickinson, 2004). At the end
434 of the Sonoma Orogeny, a continental magmatic arc system extended along the western North
435 American boundary (Dickinson, 2004), some of the remnants of which are still active today in
436 the Cascades, Trans-Mexican Volcanic Belt and Middle America Arc. Subduction has all but
437 ceased along the Canada to Mexico margin as a transform margin developed ca. 50 Ma (Queen
438 Charlotte Fault predecessor, Rusmore et al., 2010) and ca. 28 Ma (San Andreas Fault, Atwater
439 and Stock, 1998) leading to gravitationally relaxation of the orogen (Liu and Shen, 1998). In the
440 western Pacific, subduction zones are generally in retreat, causing significant upper-plate exten-
441 sion in the back-arcs (Vaes et al., 2019). We also include the Verkhoyansk-Kolyma Orogeny of
442 far east Russia in our Cordilleran classification (Figure 8), which experienced intracontinental
443 deformation during the mid-Cretaceous in response to compressional forces applied by relative
444 Siberian and Alaskan convergence, but is no longer active (Oxman, 2003; Filatova and Khain,
445 2008). The eastern half of the North China Craton also experienced widespread volcanism in
446 the Mesozoic and Cenozoic related to Pacific subduction (Wu et al., 2019). While there is no
447 metamorphic evidence for resetting of high-temperature thermochronometers, there is ample
448 evidence for significant modification of the lithosphere (Kusky et al., 2007; Yang et al., 2018;
449 Li et al., 2019; Dong et al., 2021).

450 3.2.2.2. Neoproterozoic to Mesozoic Orogens

451 . The Indosinian Orogeny, ca. 310 to 200 Ma, resulted from the closure of the Paleo-Tethys
452 and Paleo-Pacific oceans in the late Paleozoic and early Mesozoic and led to the formation of
453 much of East Asia (Lepvrier et al., 2004; Morley et al., 2013; Arboit et al., 2016; Gao et al.,
454 2017; Dew et al., 2021). Specifically, the orogeny involves the collision of South China with

455 North China, Indochina with South China, and the Sibumasu Terrane with Indochina and the
456 intervening Sukhothai arc terrane (Arboit et al., 2016; Gao et al., 2017; Dew et al., 2021). It
457 is responsible for the accretion and amalgamation of much of East Asia (Figure 9b).

458 The Variscan-Hercynian Orogeny is defined by the closure of the Rheic Ocean ca. 290 Ma
459 (Matte, 2001; Nance et al., 2012). The orogeny spans the period from ca. 360 to 280 Ma (Edel
460 et al., 2014; Žák et al., 2014) and extended from north and western Mexico to Florida in North
461 America, the Iberian Peninsula to the Tornquist zone in Europe and included parts of north
462 and west Africa (Figure 9b; Catalán et al., 2021). The orogeny concluded with the collision of
463 Gondwana with the Carolina Arcs, Meguma, Amorica and Avalonian terranes with Laurussia
464 (Stampfli et al., 2013). In Europe, the orogeny involved several additional microcontinents
465 that comprise western Europe (i.e., Iberia and Cadomia). Further east, the orogen merges into
466 the early Alpine-Himalayan Orogen and the late Central Asian Orogenic Belt and Indosinian
467 orogens with the closure of the various strands of the Paleotethys and Paleo-Asian Oceans
468 (Sengör and Natal'in, 1996; Robertson et al., 2004; Xiao et al., 2015; Gardiner et al., 2016).

469 The Caledonian Orogeny comprises a series of orogenic events spanning the period ca. 540 to
470 350 Ma resulting from the closure of the Iapetus Ocean (Figure 9c; McKerrow et al., 2000; Weller
471 et al., 2021). The orogen is typically associated with deformation from northeastern Greenland,
472 Svalbard and western Baltica, extending through the British Isles and continuing south in the
473 Appalachian Mountains (Weller et al., 2021). In eastern North America and Baltica, this
474 orogeny ended with the collision of Avalonia with Laurussia (Nance et al., 2012). We have
475 also included the Arctic Innuitian (Ellesmerian) Orogeny in our Caledonian Orogeny definition
476 due to its proximity in space and time to the Caledonian system sensu stricto (Barnes et al.,
477 2020), though the Innuitian occurred near the end of the Caledonian and is nearly orthogonal
478 in strike. While portions of the Innuitian Orogeny were overprinted in the early Tertiary by
479 the Eurekan Orogeny, this intracontinental deformation was relatively minor and is difficult to
480 distinguish from the earlier Innuitian Orogeny (Gion et al., 2017, and references therein).

481 The Central Asian Orogenic Belt (CAOB; Windley et al., 2007) has been divided into the
482 three separate orogenic systems following the model presented by Xiao et al. (2015), which
483 involves subduction and eventual closure of distinct regions of the Paleo-Asian Ocean from the
484 Tonian until the Triassic. The larger system is divided into (1) the Kazakhstan tectonic collage

485 and orocline; (2) the Mongolia tectonic collage; and (3) the Tarim-North China system, which
486 we refer to as the Solonker Orogeny (Figures 8 and 9b), that overlaps in age with the more
487 southern Indosinian Orogeny (Xiao et al., 2015; Song et al., 2018). Accretion of the Kazakhstan
488 and Mongolian tectonic collages continued through the Neoproterozoic and Paleozoic and were
489 terminated by the Solonker Orogeny at ca. 270 to 235 Ma that marks the final closure of
490 the Paleo-Asian Ocean (Eizenhöfer et al., 2014; Song et al., 2018, 2021). The CAOBS was
491 subsequently reactivated during the Meso-Cenozoic in response to distant events related to the
492 progressive consumption of the Tethys and Mongol-Okhotsk Oceans (e.g., Glorie and Grave,
493 2016). The Bureja-Jziamusy Terrane also experienced deformation as part of the Solonker
494 Orogen, but was subsequently overprinted by Cordilleran deformation as recently as ca. 95 to
495 90 Ma (Derbeko, 2013).

496 The Uralian-Timan Orogeny is the result of several subduction-related magmatic periods
497 spanning the period ca. 610 to 250 Ma (Figure 9b Fershtater, 2012; Pease, 2021). Oceanic arc
498 volcanics associated with the subduction of the Paleo-Asian orogen are recorded in the Pechora
499 arc (ca. 560 Dovzhikova et al., 2004) and blueschists and eclogites in the Urals (ca. 530 Willner
500 et al., 2019). The latter set of magmatic events is related to the subduction and closure of the
501 Paleo-Asian ocean and collision with the Kazakhstan Orocline (Xiao et al., 2015). The Uralian
502 Orogeny extends into the Taimyr Fold belt where compression continued to 220 Ma, after the
503 orogen had ceased elsewhere (Torsvik and Andersen, 2002).

504 The Neoproterozoic to Cambrian Gondwana-forming orogenies are a set of generally con-
505 temporaneous orogens—though they span nearly 700 Ma (1200 to 500 Ma)—culminating in
506 the amalgamation of Gondwana (Figure 9c; Collins and Pisarevsky, 2005; Meert and Lieber-
507 man, 2008; da Silva Schmitt et al., 2018; Goscombe et al., 2020; Şengör et al., 2020; Collins
508 et al., 2021b). We have separated these orogenies into the East African, Saharide, Kuunga, and
509 Brasiliano orogenies in much the same way as the Central Asian Orogenic belt was separated
510 into distinct systems. The East African Orogen (Stern, 1994; Collins and Windley, 2002; John-
511 son et al., 2011; Fritz et al., 2013; Collins et al., 2021b) runs through Arabia, eastern Africa,
512 Madagascar, southern India and into the Lützow-Holm Bay area of Antarctica. Whereas, we
513 have located the Kuunga Orogen (Meert and Voo, 1997) as being the orogen that separated
514 Neoproterozoic India from Australia/Mawson (Collins and Pisarevsky, 2005), running through

515 NE India, SW Australia and into Antarctica. This is broadly the trace of the Pinjarra Orogen
516 of [Fitzsimons \(2003\)](#) and Prydz-Denman-Darling Orogen of [Collins and Pisarevsky \(2005\)](#), but
517 may include the Mirny Fault ([Daczko et al., 2018](#)) and Gamburtsev suture ([Ferraccioli et al.,](#)
518 [2011](#)) as Neoproterozoic plate boundaries between Indo-Antarctica and Australo-Antarctica
519 ([Mulder et al., 2019](#)). [Şengör et al. \(2020\)](#) recently suggested the Tuareg Shield, Arabian-
520 Nubian Shield, and portions of the Saharan Metacraton constitute a single volcanic arc system
521 that was segmented and recombined in a fashion very similar to the Kazakhstan and Mongolian
522 oroclinal. We tentatively accept this model for the Tuareg Shield and Saharan Metacraton,
523 but include the younger Arabian Nubian Shield in the East African Orogen (Figure 8), whilst
524 appreciating a likely continuity of orogenesis from one to the other ([Blades et al., 2021](#); [Collins](#)
525 [et al., 2021a](#)). Our use of Brasiliano Orogen encompasses all the South American Gondwana-
526 forming orogens, as well as orogens along the west coast of Africa that correlate with them
527 (including the West Congo Orogen, the Rokelides, the Gariep Belt and the Kaoko Belt).

528 In the early Neoproterozoic, ca. 880 to 500 Ma, a small subduction-related orogen occurred
529 along the present-day eastern and southern margin of the Siberian Craton ([Vernikovsky et al.,](#)
530 [2003](#); [Kuzmichev and Sklyarov, 2016](#)), which we refer to as the Yenisey Orogeny (Figures 8 and
531 9). The orogeny coincided with the accretion of the Angara terrane at ca. 870 Ma, ([Vernikovsky](#)
532 [et al., 2007](#); [Gladkochub et al., 2010](#)). Numerous A-type magmatic dates, ca. 880 to 720 Ma,
533 are interpreted as part of a back-arc basin system ([Kozlov et al., 2012](#); [Kuzmichev and Sklyarov,](#)
534 [2016](#)) and is consistent with metamorphism recorded during this interval ([Gladkochub et al.,](#)
535 [2010](#)). The system transitioned to seafloor spreading creating the Isakovka Terrane, an arc
536 ophiolite, ca. 700 to 635 Ma ([Vernikovsky et al., 2003](#); [Kuzmichev and Sklyarov, 2016](#)), which
537 later accreted to the continent associated with a 500 to 470 Ma high-temperature metamorphic
538 event in the middle of the orogen ([Gladkochub et al., 2010](#), and references therein).

539 3.2.2.3. *Paleo- to Mesoproterozoic Orogens*

540 . Today, the Mesoproterozoic and Paleoproterozoic orogens are fragmented and scattered across
541 multiple continents (Figure 8), and in many places, reworked by more recent events ([Phillips](#)
542 [et al., 2009](#)). This dispersion and tectonic overprinting obscures the orogenic systems with
543 time and makes it more difficult to associate terranes with individual orogens. As a result, we
544 recognise that our orogenic-systems approach becomes more subjective. To retain an orogenic-

545 systems approach as much as possible, we have used paleomagnetic-based (e.g., [Condie et al.,](#)
546 [2021](#)) and full-plate tectonic ([Merdith et al., 2021](#); [Cao et al., in prep.](#)) reconstructions (Fig-
547 ure 9d to f), while recognising an inevitable shift to a more temporal-based scheme for the
548 pre-Neoproterozoic.

549 The late Mesoproterozoic to Tonian orogenies include the orogens that assembled Rodinia
550 (Figure 9d), which are now widely dispersed across the globe (Figure 8; [Li et al., 2008](#)). Many
551 studies refer to orogenesis during the period 1.3 to 0.9 Ga as Grenvillian-aged (e.g., [Tohver](#)
552 [et al., 2006](#); [Sheppard et al., 2007](#); [Goodge et al., 2010](#); [Chattopadhyay et al., 2015](#)). However,
553 the conflation of orogen names to mean stretches of time has caused considerable confusion. For
554 example, the term ‘Pan-African’ has been used to mean any orogen that occurred between ca.
555 800 and 400 Ma ([Kröner, 1980](#)), whereas orogens of this age form a number of discrete orogenic
556 systems (see discussion of Gondwana above). Similarly with the term ‘Grenvillian’ (e.g., [Kra-](#)
557 [nendonk and Kirkland, 2013](#)). [Fitzsimons \(2000\)](#) pointed out that the late Mesoproterozoic to
558 Tonian orogens that appear to surround Antarctica, in fact, fall into discrete time brackets that
559 relate to three different orogenic systems. Using the reconstruction by [Merdith et al. \(2021\)](#) as
560 a guide, we have separated the orogens of this period into the Rodinian, Rayner-Ghats, and
561 Albany-Wilkes Orogenies as separate systems active between 1.3 to 0.9 Ga.

562 The Rayner-Ghats Orogeny, ca. 1.1-0.9 Ga, has been interpreted as a distinct orogen outside
563 of Rodinia in recent global plate models (Figure 9d; [Merdith et al., 2021](#)). This orogen includes
564 the Rayner Complex in Antarctica ([Halpin et al., 2013](#); [Liu et al., 2014](#); [Morrissey et al., 2015](#)),
565 conjugate terranes in India (the Eastern Ghats; [Korhonen et al., 2011](#)) and the Central Indian
566 Tectonic Zone ([Bhowmik, 2019](#)).

567 The 1.3 to 1.0 Ga Rodinian Orogeny (Figures 8 and 9d) includes the Grenville and Llano
568 provinces in North America ([Whitmeyer and Karlstrom, 2007](#); [Johansson et al., 2022](#)), which,
569 during Rodinia assembly, we link with the 1.1 to 1.0 Ga Namaqua-Natal Belt in southern Africa,
570 the Maud Belt and Coats Land Block in Antarctica, and possibly the eastern South Tasman
571 Rise ([Mulder et al., 2018](#)). We also include the ca. 1.1 to 0.9 Ga Laurentia-Australia transform
572 as part of the Rodinian Orogeny that [Mulder et al. \(2018\)](#) connects to Rodinian subduction
573 zones. In South America, the Rondonia-Juruena Province and its continuation underneath the
574 Llanos Basin had a long-lived history of deformation spanning ca. 1.32 to 0.96 Ga ([Tohver](#)

575 [et al., 2006](#)) with late magmatic activity in the Sunsás Orogen (1.17 to 1.08 Ga; [Santos et al.,](#)
576 [2008](#); [Nedel et al., 2020](#); [Johansson et al., 2022](#)).

577 The Albany-Wilkes Orogeny, ca. 1.38 to 1.13 Ga, resulted from the collision of the West
578 Australian Craton with the North and South Australian Cratons and the Mawson Craton of
579 Antarctica (Figure 9d; [Maritati et al., 2019](#); [Pawley et al., 2020](#)). [Mulder et al. \(2018\)](#) suggests
580 that the Albany-Wilkes Orogen is likely a continuation of the Grenville-Maud system (Rodinian
581 in our lexicon). However, the Albany-Fraser and Wilkes Orogenies may have started on the
582 same broad margin of Nuna/Columbia during break-up through to the assembly of Rodinia
583 ([Pisarevsky et al., 2003, 2014](#); [Yang et al., 2020](#); [Kirscher et al., 2020](#)). Most of the Rodinian
584 deformation occurs on the opposite side of Laurentia from East Antarctica/Australia and starts
585 at a later time; hence we consider the Albany-Wilkes to be a separate system. Recent dating of
586 1.38 to 1.275 Ga metamorphism in the Rudall Province and Western Musgraves, respectively
587 (the Parnngurr and Mount West Orogenies; [Howard et al., 2015](#); [Payne et al., 2021](#)) extends
588 the early Albany-Fraser orogenesis north and documents collision between the West Australian
589 Craton and the combined South and North Australian Cratons. However, in places these were
590 subsequently overprinted during the Miles (ca. 650 to 625 Ma) and Paterson-Petermann (ca.
591 580 to 530 Ma) orogenies and thus related to the Kuunga Orogeny (Figure 9c). Two meta-
592 morphic events have been dated in the Albany Fraser Orogen, ca. 1.345 to 1.260 Ga and 1.215
593 to 1.140 Ga, resulting from the geometry of the collision between the West Australian Cra-
594 tons and the rest of Proterozoic Australia, with post-Parnngurr orogenic rotation and collision
595 of the West and South Australian Cratons ([Clark et al., 2000](#)). The Albany-Fraser Orogen,
596 Coompana and Madura provinces of Australia (e.g., [Kirkland et al., 2017](#); [Spaggiari et al.,](#)
597 [2018](#); [Pawley et al., 2020](#)) have been linked to their Antarctic conjugates from detrital zircon
598 spectra in offshore sediments and onshore geophysical characteristics ([Maritati et al., 2019](#)).
599 The Wilkes Orogen records amphibolite facies metamorphism at ca. 1.305 Ga and granulite
600 facies overprinting associated with charnockite intrusions, ca. 1.20 to 1.16 Ga ([Morrissey et al.,](#)
601 [2017](#)). The system also includes the 1.380 to 1.275 Ga Parnngurr and Mount West Orogenies
602 ([Payne et al., 2021](#)), an intracontinental contractional orogeny between the Yilgarn and Pilbara
603 Cratons. Tectonothermal events are also present in the central Australian Arunta Block (ca.
604 1.13 Ga; [Scrimgeour et al., 2005](#); [Morrissey et al., 2011](#); [Wong et al., 2015](#)) and in the Western

605 Australian Capricorn Orogen between the Yilgarn and Pilbara cratons (the 1.321–1.171 Ga
606 Mutherbukin Tectonic Event and the 1.026–0.954 Ma Edmondian Orogeny), which occur as
607 intracontinental far-field orogenesis to the amalgamation of Proterozoic Australia. It is unclear
608 how far the orogen extends into Antarctica due to the extensive ice cover.

609 Between ca. 2.1 and 1.45 Ga two major orogenic systems are associated with Nuna (Fig-
610 ure 9e [Pisarevsky et al., 2014](#); [Condie et al., 2021](#); [Cao et al., in prep.](#)): the Inner-Nunianian
611 Orogeny (ca. 2.1 to 1.7 Ga), which formed the core of the supercontinent; and the accre-
612 tionary Circum-Nunianian Orogeny (ca. 1.85 to 1.45 Ga) that is driven by a subduction girdle
613 surrounding the core. There are several reconstructions for Nuna ([Bispo-Santos et al., 2008](#);
614 [Elming et al., 2009](#); [Johansson, 2009](#); [Zhang et al., 2012](#); [Pisarevsky et al., 2014](#); [D’Agrella-
615 Filho and Cordani, 2016](#); [Meert and Santosh, 2017](#); [Cawood et al., 2020](#); [Elming et al., 2021](#);
616 [Cao et al., in prep.](#)), and while most are sufficiently similar to yield little difference in the last
617 orogeny designation, there are competing models for the likely participation of Amazonia that
618 will affect our model. In the SAMBA models, first proposed by [Johansson \(2009\)](#), Amazonia
619 (Central Amazonian and Ventuari-Tapajós Belts) is contiguous with the Baltic Shield, ca. 2.1
620 to 1.8 Ga, both which experienced significant intrusive magmatism during the Nunianian Oro-
621 genes (e.g., [Almeida et al., 2007](#); [Bogdanova et al., 2015](#); [Juliani et al., 2021](#)). In the alternative
622 configuration by [Pisarevsky et al. \(2014\)](#), Amazonia and the West African Craton form a lesser
623 continent separate from Nuna (Figure 9e [Cao et al., in prep.](#)), deforming as part of a separate
624 accretionary margin. We have included the Amazonian Belts as part of the Nunianian Orogenies
625 (Figure 8 and 9e), but acknowledge that they may have evolved as a separate system.

626 The Inner-Nunianian Orogeny, ca. 2.1 to 1.76 Ga, was a major global event related to the
627 closure of the Manikewan Ocean and assembly of the Nuna/Columbia supercontinent ([Corrigan
628 et al., 2009](#); [Weller et al., 2021](#)). In North America, the Superior Craton collided with the
629 Reindeer Zone and Sask Craton, in a collision that has been compared to India colliding with
630 southern Asia ([St-Onge et al., 2006](#); [Darbyshire et al., 2017](#); [Weller and St-Onge, 2017](#)). On
631 the opposing side of the Reindeer Zone, the core of the orogen, contains a number of terranes,
632 grouped into the Hearne and Rae Cratons, which are sutured together by the Snowbird Tectonic
633 Zone ([Thiessen et al., 2018](#)). Deformation extended across to the Taltson-Thelon Arc (ca. 1.87
634 to 1.84 Ga) between the Rae and Slave Cratons ([Chacko et al., 2000](#); [Whalen et al., 2018](#)).

635 Baltica also experienced widespread orogenic activity as part of the Inner-Nunanian system that
636 is recorded in igneous and metamorphic activity that affected the Kola Block (e.g., [MIkkola](#)
637 [et al., 2018](#); [Daly et al., 2001](#); [Tuisku and Huhma, 2006](#); [Makkonen et al., 2020](#)), which at the
638 time, the present-day northern margin of the Kola Peninsula was adjacent to the eastern margin
639 of northern Greenland (Evans and Mitchell, 2011). Several reconstructions place the southern
640 margin of the Siberian Craton against the northern Margin of Laurentia at this time ([Condie](#)
641 [and Rosen, 1994](#); [Sears and Price, 2003](#)) and Baltica to its east (Evans and Mitchell, 2011).
642 However, more recent models suggest it collided post 1.9 to 1.84 Ga after the Anabar and Aldan
643 terranes had accreted to the eastern margin of the Siberian Craton based on geochronology of
644 mafic dike swarms, and post-collisional granitoids ([Donskaya et al., 2009](#); [Ernst et al., 2016](#)).
645 The Inner-Nunanian Orogeny also saw the amalgamation of Volgo-Uralia and Sarmatia, ca.
646 2.1 to 2.0 Ga ([Savko et al., 2015](#); [Baltybaev et al., 2017](#)), and the subsequent collision with
647 Baltica at ca. 1.82 to 1.80 Ga ([Bogdanova et al., 2015](#)). This collision is distinguished from
648 the Circum-Nunanian belts in Baltica that run orthogonal to the Volgo-Sarmatia collision.
649 This collision is distinguished from the Circum-Nunanian belts in Baltica that run orthogonal
650 to the Volgo-Sarmatia collision. The spatial complexity of the Inner-Nunanian Orogeny may
651 result from multiple systems just as the inner Gondwana orogenies, but uncertainties in the
652 geographic positions of many key crustal elements make it difficult to divide further at present.

653 The next few paragraphs discuss the Circum-Nunanian Orogeny and some of the variations
654 in configurations. Some of these variations are slight while others are quite dramatic, but
655 despite these differences the orogens all appear to occur along the exterior of Nuna's core as
656 terranes were accreted. Thus regardless of the accuracy of the geologic connections between
657 terranes, the last orogeny classification remains the same.

658 Wyoming was likely the first accretionary terrane added to Laurentia during the Circum-
659 Nunanian Orogeny, colliding with the Medicine Hat Terrane causing deformation in the Great
660 Falls Tectonic Zone (ca. 1.86 to 1.73 Ga; [Gifford et al., 2018](#)). Also during the first phase
661 of the Circum-Nunanian Orogeny, ca. 1.85 to 1.75 Ga, the Penokean Orogeny was a small
662 deformation event on the southern margin of the Superior craton recorded in metamorphism
663 and accompanying magmatism ([Holm et al., 2007](#); [Vallini et al., 2007](#); [Klier, 2019](#); [Zi et al.,](#)
664 [2021](#)). On the southern margin of Laurentia, a series of exotic terranes, Yavapai (ca., 1.80 to

665 1.70 Ga), Mazatzal (1.70 to 1.65 Ga), and Granite-Rhyolite terranes (1.50 to 1.45 Ga), were
666 accreted over the course of approximately 300 Ma (Karlstrom et al., 2001; Whitmeyer and
667 Karlstrom, 2007; Amato et al., 2008; Mako et al., 2015). The Yavapai and Mazatzal terranes
668 include juvenile arcs, ophiolites and metasediments that were accreted ca. 1.71 to 1.68 Ga
669 and 1.646 to 1.633 Ga during the Yavapai and Mazatzal Orogenies, respectively (Whitmeyer
670 and Karlstrom, 2007; Amato et al., 2008). The ca. 1.49–1.45 Ga Picuris Orogeny occurred
671 during a rare period of orogenic preservation between supercontinent cycles and is relatively
672 limited geographically to southern Laurentia which included parts of Precambrian Australia at
673 the time. The orogeny is identified in the southwestern United States where it deforms older
674 crust and appears not to have juvenile magmatism associated with it (Daniel et al., 2013; Mako
675 et al., 2015; Aronoff et al., 2016). The orogen extends into the northeastern United States
676 (Medaris et al., 2021), where it is progressively overprinted by the Grenvillian Orogeny in the
677 east. On the present-day eastern margin of Greenland, several exotic terranes were accreted to
678 the Kola-Karelia Craton, including Bergslagen-Livonia (ca. 1.89 and 1.84 Ga) and Amberland
679 (ca. 1.84 and 1.83 Ga; Bogdanova et al., 2015).

680 In Australia, extensive plate-margin orogenesis (ca. 1.82 to 1.55 Ga), similar to that inter-
681 preted for SW Laurentia, occurs throughout the South Australian Craton (Kimban and Kararan
682 orogenies; Hand et al., 2007) and North Australian Craton (Yambah-Strangways-Leibig oroge-
683 nies). These likely formed a continuous accretionary system (Payne et al., 2009; Betts et al.,
684 2008; Betts and Giles, 2006). Extensive intracontinental orogenesis within Western Australia
685 is marked by the 1.82 to 1.77 Ga Capricorn Orogeny (Johnson et al., 2013) and the 1.68 to 1.62
686 Ga Mangaroon Orogeny (Sheppard et al., 2005). The Isan Orogeny (including the Chewings
687 and Olary orogenies) spanned the Paleo-Mesoproterozoic boundary (ca. 1.65-1.55 Ga). The
688 effects of this orogeny dominate the eastern entirety of pre-Phanerozoic Australia (Morrissey
689 et al., 2011; Tidby et al., 2020; Volante et al., 2020), and extend into the Gawler Craton in
690 southern Australia (Cutts et al., 2011) and into the central North Australian craton (Anderson
691 et al., 2013). This orogeny is envisaged to have occurred as a consequence of collision between
692 Paleoproterozoic Australia (then consisting of the North Australia Craton and the South Aus-
693 tralian Craton with the North China Craton and a large piece of East Antarctica) and Laurentia
694 (Pourteau et al., 2018), and is recorded by the Racklan Orogeny in NW Laurentia (Furlanetto

695 [et al., 2013](#)). Orogenic activity coeval with younger Picuris Orogeny are found in the Gawler
696 Craton of Australia ([Hall et al., 2018](#); [Morrissey et al., 2019](#)), and in the Mount Isa region of
697 NE Australia ([Cave et al., 2022](#)).

698 In Antarctica, ca. 1.7 Ga orogenesis recorded along the coast of the Mawson Craton ([Peucat
699 et al., 1999](#)), as well as in the central Transantarctic Mountains ([Goodge et al., 2001](#); [Brown
700 et al., 2021](#)), although the extent of this orogenic activity into the interior of Antarctica is
701 unknown, it appears to be an extension of deformation in the Gawler Craton. There is some
702 ambiguity in the connections between Australia, Antarctica and western Laurentia ([Wingate
703 et al., 2002](#)), however, the connection of these three bodies is established from paleomagnetism
704 and geologic observations (aforementioned Nuna reconstructions; [Whitmeyer and Karlstrom,
705 2007](#)). The difficulty in precisely resolving the connections results from reworking of western
706 North America during the Cordilleran Orogeny and the uncertainties in paleomagnetic poles.

707 The accretionary orogenesis at this time on both the Dharwar-Bastar Cratons (South In-
708 dia) and the Bundelkhand Craton (northern India) is recorded in the Krishna Orogeny of the
709 Ongole Domain (1.68-1.60 Ma; [Henderson et al., 2014](#)) and in the Central Indian Tectonic Zone
710 ([Bhowmik, 2019](#)), respectively. The two halves of Peninsula India were likely separate conti-
711 nents before the Neoproterozoic. Paleomagnetic reconstructions have placed Southern India
712 conjugate to Antarctica or NE Australia ([Zhang et al., 2012](#)); however, more recent models
713 place the Indian continent adjacent to Baltica ([Pisarevsky et al., 2003](#); [Cawood et al., 2020](#)).
714 Regardless, the Eastern Ghats appear to have been part of the active accretionary margin of
715 Nuna from 1.85 to 1.60 Ga or possibly as late as 1.45 Ga based on dating of tectonomagmatic
716 activity and an accreted ophiolite terrane in the Krishna Province ([Dasgupta et al., 2013](#)).

717 Prior to 1.95 Ga, the North China Craton was a set of microcontinents separated by ocean
718 basins that closed during the same period as both major Nunanian orogenies, completing by
719 1.85 Ga ([Zhao et al., 2012](#)), which is recorded in magmatism and widespread metamorphism
720 (e.g., [Yin et al., 2014](#); [Cai et al., 2015](#); [Wu et al., 2016](#); [Liu et al., 2017](#)). However, the location of
721 the North China Craton has been subject to extensive debate, with various hypothesis ranging
722 from the northern margin of Siberia ([Halls et al., 2000](#)), paired with the Kola-Karelia Craton
723 and ([Wilde et al., 2002](#)), positioned between Baltica and Amazonia ([Pesonen et al., 2012](#)), or
724 joined with India outboard of the Nuna accretionary margin ([Zhao et al., 2011](#)). However,

725 we prefer more recent models that suggest a long-lived connection with the North Australian
726 Craton on the basis of more extensive paleomagnetic and geologic correlations (Wang et al.,
727 2019; Nixon et al., 2022; Zhang et al., 2022). Much of the North China Craton has similar
728 ages to the Inner-Nunanian Orogen, about half the craton experience widespread magmatism
729 as part of the Cordilleran Orogeny (Wu et al., 2019).

730 Most models for Nuna do not include the Kalahari Craton; however, the two models that
731 do, place it in opposite hemispheres (Djeutchou et al., 2021; Cao et al., in prep.). A recent
732 paleomagnetic reconstruction by Djeutchou et al. (2021) suggests the Zimbabwe Craton (the
733 northern part of the Paleoproterozoic Kalahari Craton) was juxtaposed against the southern
734 margin of the Superior Craton at 1.88 Ga. Additionally, their model calls for a pair of sub-
735 duction zones along the south and western margins of the Kaapvaal Craton (the southern part
736 of the Kalahari Craton), reworking the Magondi and Kheis Belts and accreting the Rehoboth
737 Block at this time (Kleinhanns et al., 2013). However, such placement would leave little time for
738 the Kalahari Cratons to rift away from Superior, prior to the arrival of the Mazatzal Orogeny
739 ca. 1650 Ma along a margin that experienced the accretion of at least three major terranes ca.
740 1850 to 1450 Ma (Yavapai, Mazatzal and the Granite-Rhyolite terranes). The paleomagnetic
741 data and timing dates of deformed terranes are also consistent with a collision between Congo-
742 Tanzanian Craton with the south African cratons, which is consistent with ca. 2.0 Ga, resulting
743 in exhumation of eclogites in the Ubendian-Usagaran Belts (Collins et al., 2004; Tamblyn et al.,
744 2021). Regardless, both interpretations place these deformed and accretionary terranes within
745 the Circum-Nunanian Orogeny.

746 The Wopmay Orogeny, ca. 1.95 to 1.84 Ga, was a small subduction-related event on the
747 western margin of the present-day Slave Craton in northern Canada (Figure 8 Bowring and
748 Podosek, 1989). East-dipping subduction on the western margin accreted three separate arc
749 terranes, Great Bear, Hottah, and Fort Simpson, to the craton at ca. 1.88 Ga in the short-lived
750 Calderian Orogeny (Hildebrand et al., 2009; Cook, 2011).

751 The Eburnean Orogeny records the collision between the Saõ Francisco Craton and the
752 Gabon Belt on the eastern Congo Craton margin, (ca. 2.12 to 2.0 Ga; Weber et al., 2016).
753 The orogeny is recorded in a set of tectonomagmatic events (Doumbia et al., 1998; Barbosa
754 et al., 2008; Peucat et al., 2011; Loose and Schenk, 2018; de Carvalho Filgueiras et al., 2020).

755 Although the Eburnean Orogeny was accretionary at the margins of Nuna (Figure 8e), it was
756 geographically isolated and earlier than the majority of Circum-Nunanian orogenesis.

757 Early models of the Limpopo Orogeny suggested it was active between 2.7 and 2.65 Ga
758 as the result of the collision between the Kaapvaal and Zimbabwe Cratons (Barton and van
759 Reenen, 1992), which is based on the age of granitoids contained within the thrust sheets.
760 However, metamorphic ages and more recent interpretations suggest it was active ca. 2.0 Ga
761 (Yin et al., 2019). The Limpopo Orogeny was probably intracontinental due to a lack of arc or
762 accretionary sediments of appropriate age (Yin et al., 2019).

763 The Birimian–Trans-Amazonian Orogeny occurred during the early stages or just prior to
764 the assembly of Nuna and may have been multiple spatially discrete events between ca 2.3 and
765 1.9 Ga. During this period, the West African Craton records considerable tectonomagmatic
766 activity in sedimentary deposits (Grenholm, 2019; Grenholm et al., 2019) as it collided with
767 the present-day northeastern Amazonian Craton (Grenholm, 2019). Both magmatic and meta-
768 morphic events are recorded in Amazonia during this period (De Roever et al., 2003; Savko
769 et al., 2015; Baltybaev et al., 2017; Klaver et al., 2015; da Rosa-Costa et al., 2008).

770 In the early Paleoproterozoic (ca. 2.5 to 2.3 Ga), a purported drop in magmatism corre-
771 sponded with several cratonic regions that experienced high-temperature, often contractional,
772 metamorphic events (Pehrsson et al., 2013, 2014). Pehrsson et al. (2013) hypothesized these
773 events are related to a formation of a supercontinent (Nunavutia), however, a recent paleo-
774 magnetic reconstruction suggested two separate supercratons were also consistent with pole
775 determinations and patterns of dike swarms (Figure 9f; Liu et al., 2021). The reconstructions
776 of the Siderian are based on paleomagnetic poles from 11 terranes (Liu et al., 2021; Salmi-
777 nen et al., 2021); however, less than one-quarter of blocks with igneous dates older than 2.3
778 Ga (Figure 10) are included in the reconstructions. Igneous activity in the interval 2.5 to
779 2.3 Ga is nearly ubiquitous across Archean terranes with the exception of regions covered by
780 ice, sediments, and or Phanerozoic large igneous provinces. As a result, the classification as
781 a super-continent/craton may be premature without the reconstructed positions of additional
782 Archean terranes. Metamorphic dates in this time interval are sparse (Figure 10), but the
783 metamorphic database is incomplete, so it is difficult to make clear inferences at this point. We
784 refer to the regions with significant 2.4 Ga tectonothermal activity collectively as the Siderian

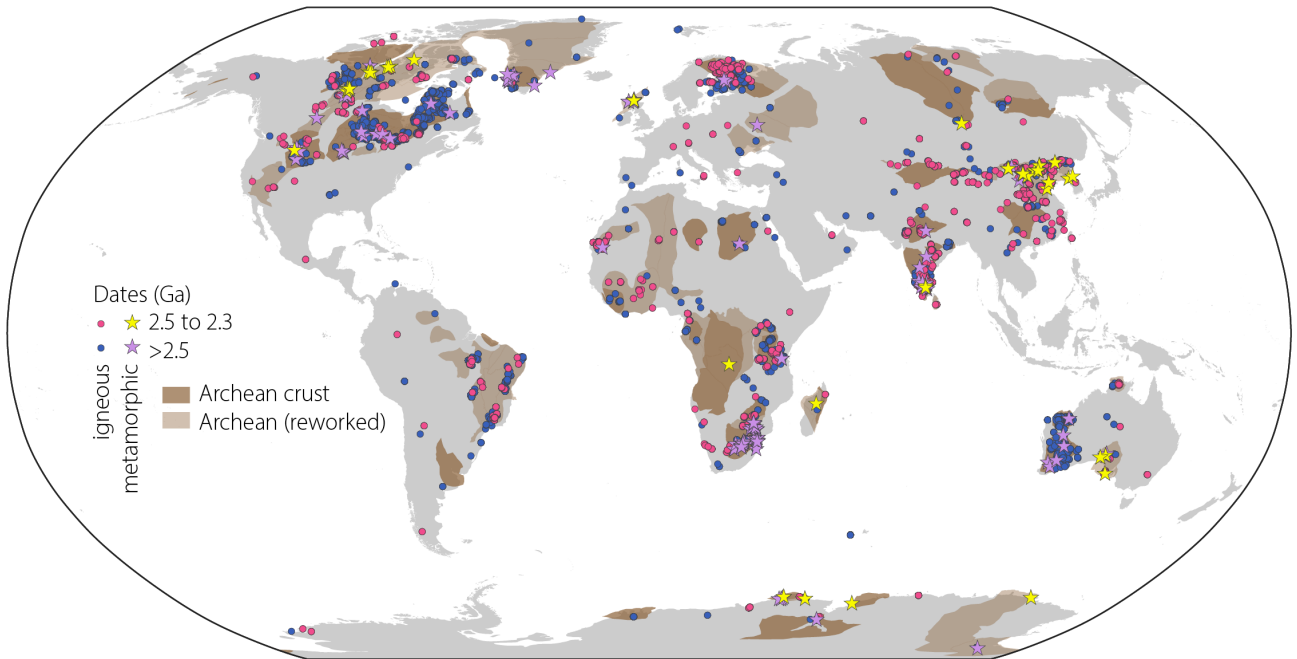


Figure 10: Regions with crust >2.3 Ga with superimposed locations of dated igneous and metamorphic activity. The observed dates are divided into the period 2.5 to 2.3 Ga, and older than 2.5 Ga. Reworked Archean crust is displayed in a lighter shade. Province boundaries from Figure 7 and a few additional Archean basement provinces in North America by [Lund et al. \(2015\)](#). Igneous ages extracted from [Gard et al. \(2019a\)](#) and [Puetz \(2018\)](#); metamorphic ages from DateView ([Eglington, 2004](#)) and the expanded metamorphic database by [Brown and Johnson \(2018\)](#).

785 Metamorphic Event, referring to the geologic time period rather than a coherent orogen due
 786 to the present uncertainty in reconstructions. The ca. 2.4 Ga affected terranes are now scat-
 787 tered across all seven continents and many have been reworked (Figure 9f and 10). Some of
 788 the terranes where the Siderian Metamorphic Event has been documented include the Mawson
 789 Craton ([Duclaux et al., 2008](#)), the Sask Craton ([Chiarenzelli et al., 1998](#)), Arrowsmith Orogen
 790 in Northern Canada ([Hartlaub et al., 2007](#); [Schultz et al., 2007](#); [Berman et al., 2013a](#)), the
 791 East Dharwar in India ([Clark et al., 2009](#); [Li et al., 2018](#)), and the Sleaford Complex in south
 792 Australia ([Halpin and Reid, 2016](#)) and the North China Craton ([Liu et al., 2017](#)).

793 4. Model Evaluation

794 4.1. Ocean–Continent Boundary

795 Not counting for topographic relief, we estimate 57.5% of the Earth’s surface is covered by
 796 oceanic crust and 42.5% is covered by continental crust (Figure 4). The seafloor age model by
 797 [Seton et al. \(2020\)](#) covers a slightly smaller proportion of the Earth’s surface with seafloor ages,
 798 57.3%. However, there are a few significant differences between the models.

799 In general, the edges of the seafloor age model by [Seton et al. \(2020\)](#) are easily correlated

800 with high bathymetric gradients and deep water. However, there are a few regions where lo-
801 cations differ significantly between their model and our ocean–continent boundary. Some of
802 the differences may be due to the quality of magnetic data near the continents where remnant
803 magnetization may be reset by high temperatures beneath insulating sediments or where mag-
804 netic data is of insufficient quality and/or density to resolve seafloor ages. A few of the larger
805 differences include the Greenland-Iceland-Faroe Ridge (GIFR), Blake Plateau, Gulf of Mexico,
806 and some microcontinents.

807 Perhaps the most obvious difference between our model and [Seton et al. \(2020\)](#) is the
808 inclusion of the GIFR as a region of continental crust. In a recent comprehensive paper by
809 [Foulger et al. \(2020\)](#), the authors make a compelling case that the GIFR is a peculiar region with
810 variably extended continental crust rather than the product of anomalous oceanic volcanism.
811 The nearby Jan Mayen microcontinent has been recognized as a microcontinents for decades
812 ([Peron-Pinvidic et al., 2012](#), and references therein), lending credence to the model. The total
813 thickness of the crust approaches 40 km thick beneath the Iceland microcontinent—a value
814 more typical of continental than oceanic crust. The symmetric, linear magnetic anomalies
815 characteristic of seafloor spreading are muddled in this region, possibly indicating a complex
816 history of rifting and volcanism. The upper crust in the region exhibits seismic properties and
817 layer thicknesses typical of oceanic crust (3–10 km), but the middle and lower crustal seismic
818 velocities and densities are better explained by continental material ([Foulger et al., 2020](#)).
819 Therefore, we favor their interpretation in our model.

820 There are a number of other microcontinents that we have included in our model that are
821 not found in the [Seton et al. \(2020\)](#) model. We have included the Hovgaard Ridge and the East
822 Greenland Ridge in the Arctic Ocean ([Funck et al., 2016](#)). Several microcontinents lie in the
823 Indian Ocean including the Mascarene Plateau, Chagos–Laccadive Ridge, Gulden Draak, and
824 Batavia Knoll ([Torsvik et al., 2013](#); [Gardner et al., 2015](#); [Halpin et al., 2017](#)), which are formerly
825 pieces of Madagascar and India, respectively. We have also included the Bollons Seamount east
826 of the Campbell Plateau in the Pacific ([Davy, 2006](#)).

827 The Blake Plateau on the Atlantic side of Florida sits at a bathymetric depth of ~ 1000 m
828 and is underlain by transitional crust ([Dillon et al., 1988](#)). In the Gulf of Mexico, the ocean–
829 continent transition is obscured by sedimentary cover making magnetic data the most useful for

830 identifying the boundary, but the global models are relatively low resolution. However, there
 831 are industry magnetic datasets that we do not have access to that yield a much clearer view of
 832 the Gulf of Mexico.

833 In the future, incorporation of seismic reflection profiles taken across the ocean–continent
 834 transition and improvements in the magnetic datasets may warrant alterations to the ocean–
 835 continent boundary.

836 *4.2. Plate Boundaries*

837 There are significant differences in the lengths of plate boundaries by type (Table 5). The
 838 total length of all major plate boundaries is nearly five times the Earth’s circumference. Among
 839 major boundaries, the total length of divergent boundaries is significantly greater than con-
 840 vergent boundaries. The mid-ocean ridge system and associated transforms is ~ 1.6 times the
 841 length of subduction zones. Transform boundaries account for the least length of major bound-
 842 aries, but this does not include the transforms between mid-ocean ridge segments.

Table 5: Lengths of plate boundaries. Major and minor boundaries are identified in Figure 2.

Boundary Type	Major (km)	Minor (km)
spreading center	87,001	12,663
extension zone	16,608	41,663
subduction zone	52,994	9,688
collision zone	18,295	65,113
dextral transform	20,856	41,663
sinistral transform	13,276	24,099
inferred	0	24,923

843 Including minor deformation zone and microplate boundaries more than doubles the total
 844 length of boundaries (Table 5). The latter number may be a bit misleading as the minor bound-
 845 aries generally bracket the deformation, while in reality multiple structures may accommodate
 846 the motion. Among minor boundaries, convergent and transform boundaries are longer than
 847 divergent boundaries. Also in contrast to major boundaries, spreading ridges and subduction
 848 zones are less common than extensional zones and thrusts.

849 *4.3. Plate Model*

850 The plate model consists of 16 rigid plates, 54 microplates and 73 deformation zones. The
 851 areas of rigid plates span (Figure 11 just over three orders of magnitude from the smallest

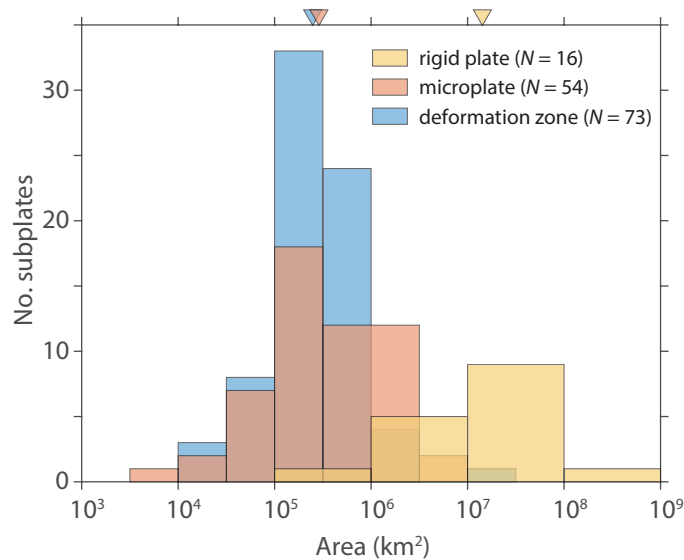


Figure 11: Areas of subplate polygons by type. The triangles are centered about the median area for each type, respectively.

852 (Juan de Fuca, $0.15 \times 10^6 \text{ km}^2$) to the largest (Pacific, $103 \times 10^6 \text{ km}^2$). The areas of microplates
 853 and deformation zones have similar ranges to rigid plates, but the median sizes are nearly
 854 two orders of magnitude lower than rigid plates. The subplates are further split into smaller
 855 polygons (208) to split oceanic from continental regions.

856 4.3.1. Comparison with Bird (2003)

857 The major differences between the Bird (2003) plate model and our model are the mi-
 858 croplates and deformation zones (Figure 1a). Most major plate boundaries are very similar to
 859 Bird (2003) with some refinement. The Bird (2003) model contains 52 polygons, whereas ours
 860 contains 121 regions comprising rigid plates, microplates and deformation zones. Most of these
 861 added regions are located in the Alpine–Himalayan Belt and the North American Cordillera.
 862 These changes are driven by a significant improvement in the number and quality of land-based
 863 GPS coverage; however, several of our additions have long been recognized as plate bound-
 864 ary zones (e.g., Gordon, 1998; Lowman and Yates, 2002; Freymueller, 2010; Kreemer et al.,
 865 2014). Many of these regions have also been previously identified as microplates (case studies
 866 in Table 2). In total, we estimate 16% of the Earth’s surface is covered by microplates and
 867 deformation zones, divided roughly equally between them whereas the Bird model only includes
 868 6%.

869 The southern Somali–African boundary has changed considerably with respect to the Bird
 870 (2003) model. The Bird (2003) model connects several discrete regions of seismicity in southern

871 Mozambique and central South Africa. However, GPS motions suggest there is little deforma-
872 tion on either side of this boundary (Kreemer et al., 2014). From GPS data, the boundary
873 appears to be further east, near the southern Mozambique coast (Stamps et al., 2021).

874 Another region of significant refinement relative to Bird’s model among the microplates and
875 deformation zones of the Philippines and eastern Indonesia. Our model more closely follows
876 several boundaries identified by Zahirovic et al. (2014), which shows better correlations with
877 seismicity and active fault models (ANSS, 2020; Styron and Pagani, 2020).

878 In our plate model, $\sim 80\%$ of earthquakes occur within 100 km of a plate boundary and $\sim 91\%$
879 within 200 km (Figure 12a). Approximately 73% of earthquakes occur within deformation zones
880 and microplates. For the Bird (2003) model, these percentages are significantly lower. Less than
881 27% of earthquakes lie within Bird’s microplates and 65% of earthquakes lie within 100 km of
882 plate boundaries. Therefore, we suggest our new model provides a more accurate representation
883 of the actively deforming crustal regions.

884 The pattern of distance of earthquakes from plate boundaries varies depending on the type
885 of plate boundary. Earthquakes are centered close to spreading centers (Figure 12b), but are
886 more diffuse around transforms (Figure 12f), which occurs because transform plate motion is
887 frequently accommodated by multiple faults rather than a single structure (e.g., Pacific–North
888 American boundary in California, Hauksson et al., 2013; DeMets et al., 2014). In contrast to
889 spreading centers and transforms that identify the centers of deformation, extensional zones
890 and thrusts identify the boundaries of internal deformation so it makes sense that earthquakes
891 are distributed at a greater distance from these boundaries (Figure 12c, e). Subduction zone
892 earthquakes are distributions furthest from their associated plate boundaries (Figure 12d).
893 The more rapid increase in cumulative density of subduction zone earthquakes could be due to
894 our limiting earthquakes to 30 km depth and/or the earthquakes extend far enough into the
895 deformation zone that the opposite boundary becomes the closest (Figure 2).

896 4.3.2. Comparison with Tomography and Volcanism

897 Our plate boundary model correlates well with slow seismic shear-wave velocity slices from
898 40 to 90 km depth, with 70 km displaying the most similarity to the plate boundary zones
899 (Figure 13). Greater than 90 km depth, the mantle beneath the oceans has significantly larger
900 negative velocity anomalies. Greater than approximately 125 km depth, some of the continental

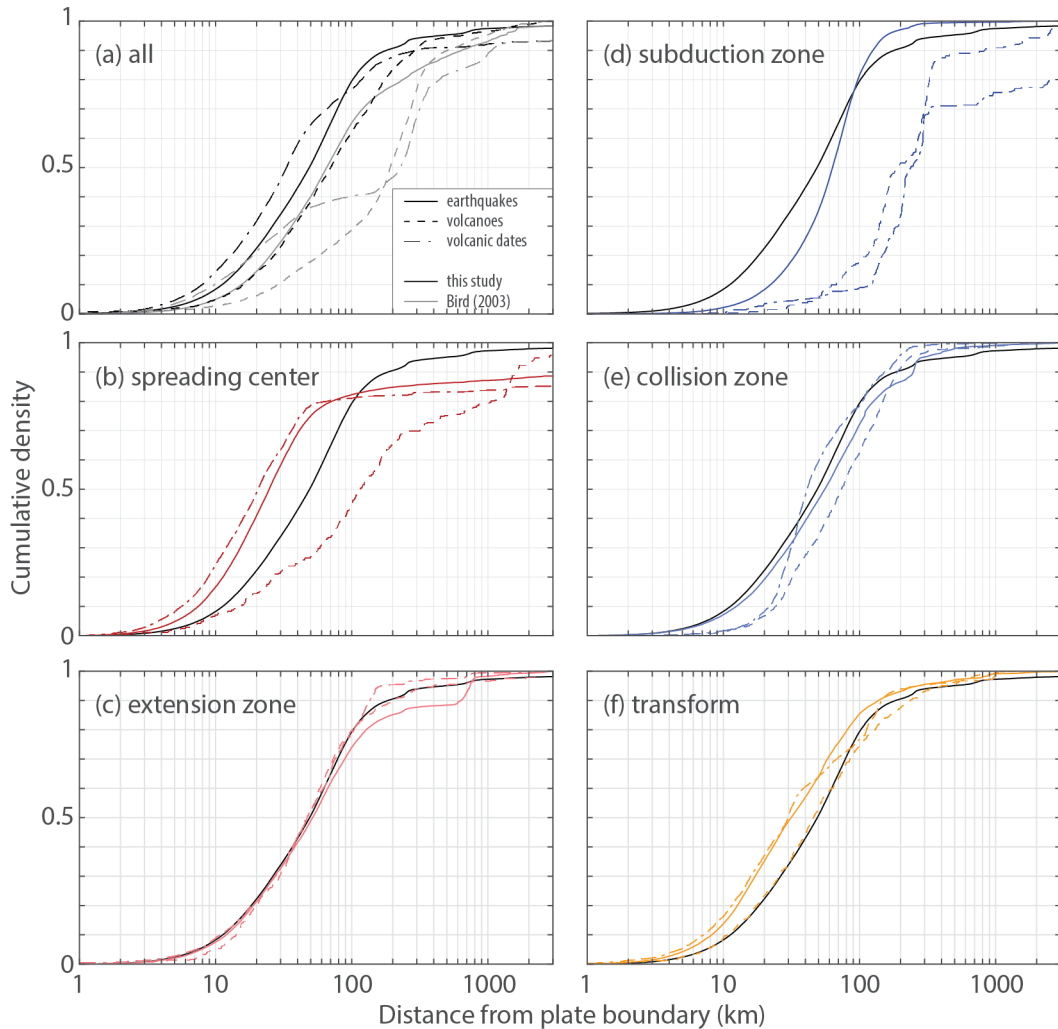


Figure 12: Cumulative distributions of earthquakes and volcanoes by distance to plate boundaries. Each subplot contains CDFs from three datasets: (solid) earthquakes from the ANSS catalog, magnitude 3.0 to 5.5, 1990 to 2020 and magnitude 5.5+, 1970 to 2020 (ANSS, 2020); (dashed) Quaternary volcanoes (Global Volcanism Program, 2013); and (dashed-dotted) dates of volcanic samples limited to Quaternary samples Gard et al. (2019a) with duplicate sample locations removed to limit oversampling. (a) Computed using all data in each respective dataset, black lines using our model and grey lines using Bird (2003). (b-f) Computed for listed plate boundary types as classified in Figure 2. The black line on each plot is earthquakes from all plate boundaries as a reference.

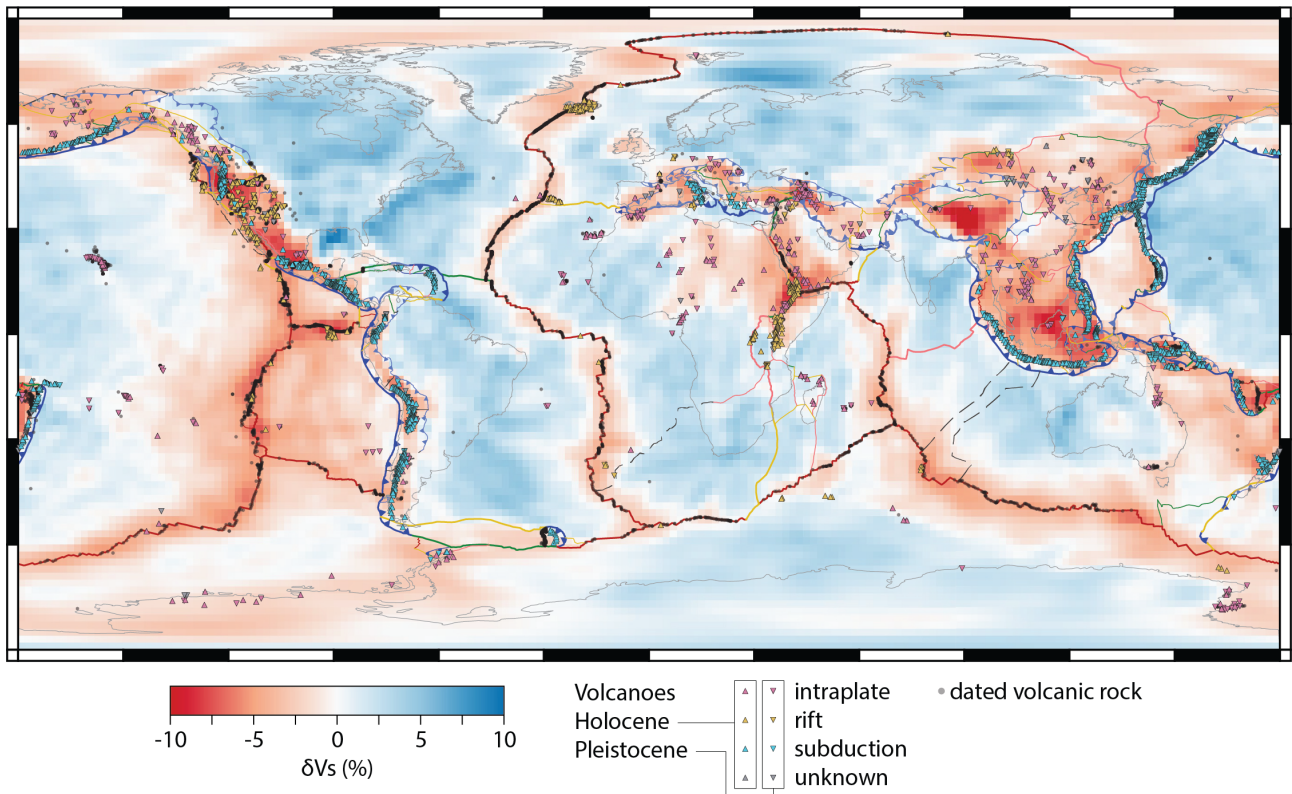


Figure 13: Shear wave tomography at 70 km depth demonstrates a high spatial correlation between negative velocity perturbations and plate boundaries and deformation zones. The shear wave velocity model is by [Debayle et al. \(2016\)](#). Nearly 80% of recently active, Quaternary volcanoes lie within microplates and deformation zones. Volcano locations from [Global Volcanism Program \(2013\)](#). Pleistocene and younger volcanics from [Gard et al. \(2019a\)](#); originally from EarthChem.org affiliated databases.

901 plate boundary zones begin to lose their negative velocity anomalies. The correlation between
 902 plate boundary zones and shallow mantle shear wave velocity anomalies are the result of thinner
 903 and warmer lithosphere in actively deforming regions with respect to cold, thick, rigid plate
 904 interiors.

905 Only a few plate boundaries and deformation zones are not clearly associated with negative
 906 shear-wave velocity anomalies. The Cordilleran Frontal Thrust, Lesser Antilles Arc, Lwandle
 907 Microplate and San Microplate have non-negative seismic anomalies. In the case of Lwandle
 908 and San Microplates, the lack of a clear velocity anomaly may be due to the very slow relative
 909 velocities between the Microplate and African Plate ([Wedmore et al., 2021](#)). Oceanic intraplate
 910 deformation zones do not show clear correlations with negative tomographic anomalies. For
 911 example, the Capricorn region between the Indian and Australian Plates, the Macquarie Mi-
 912 croplate, and the boundary between North and South America are not clearly delineated by
 913 tomography.

914 Volcanoes also show high correlation with the plate boundary zones as $\sim 80\%$ lie within

915 deformation zones and microplates (Figure 13). Like earthquakes, there are differences between
916 the distribution of distance from plate boundaries by type (Figure 12). Volcanoes are furthest
917 from subduction zones, with the majority ranging between 130 to 330 km. For transforms
918 and extension zones, >75% lie within 100 km of the boundary. The distribution of volcanoes
919 from spreading centers is more complex because most of the volcanoes identified near these
920 environments are seamounts associated with hotspots rather than flows at ridges. If a full
921 accounting of flows could be made along ridges, it is likely the distribution would indicate most
922 are much closer to the drawn boundaries Rubin (2016). We try to account for this by using
923 dates from a global geochemical dataset (Gard et al., 2019a, and Figure 13), which results in
924 a similar distribution of distances as the volcanic eruption database (Figure 12). The only
925 exception is the distribution of dates from recent volcanic samples, which fall considerably
926 closer to spreading centers than the eruption database similar to earthquakes (Figure 12b).

927 The correlation between volcanoes and seismic velocities may imply partial melt is common
928 beneath most plate boundary zones. However, there are a few that have minimal volcanism
929 but high seismicity, including Tibet and the Tien Shan Mountains. The Amur, Yangtze and
930 Okhotsk regions also have very little volcanism outside the volcanic arcs on their margins.
931 These regions also have little seismicity in their interiors.

932 There are a few regions with negative seismic anomalies and/or recently active volcanic
933 centers that do not correspond with plate boundaries. Several of these regions are hotspots
934 associated with mantle plumes (e.g., Hawaii, Reunion, Cape Verde). These regions do not have
935 clear negative velocity anomalies at 70 km depth within the scale of the tomography model
936 (Figure 13). There is a negative shear wave anomaly in the Arctic without volcanics, which may
937 be related to prior rifting and the Eureka Orogeny (Darbyshire, 2005). The Saharan Metacraton
938 has both low shear velocities and volcanism, but little seismicity. There are currently no GPS
939 data from western Egypt, Chad, Sudan or Libya above the seismic anomaly, but GPS data
940 in Nigeria do not indicate active deformation above the negative velocity region. The West
941 Antarctic Rift contains both volcanoes and a negative shear wave anomaly related to renewed
942 extension in the Cenozoic (Winberry and Anandkrishnan, 2004; Gupta et al., 2009; O'Donnell
943 et al., 2019). While there is active volcanism associated with the Marie Byrd and Erebus
944 hotspots, seafloor magnetic anomalies suggest the rift was active as recently as 11 Ma (Granot

945 [and Dymant, 2018](#)) and minimal seismicity suggests the region is inactive at present.

946 *4.3.3. Uncertain Plate Boundaries*

947 Many oceanic deformation zones and microplates have relatively uncertain boundaries due to
948 a lack of GPS constraints, significant seismicity and/or distinctive bathymetric features akin to
949 boundaries. The Azores deformation zone, for example, has a well-defined boundary along the
950 northern margin by seismicity. GPS data on individual islands suggest internal deformation
951 ([Fernandes et al., 2006](#); [Marques et al., 2013](#)), but the southern boundary is not clear from
952 seismicity. Our best estimate places the southern boundary on the East Azores Fracture Zone
953 due to its clear topographic expression.

954 The Capricorn Plate is a region of diffuse extension between the Indian and Australian
955 Plates ([Wiens et al., 1985](#); [Gordon et al., 1990](#)). The region is large, nearly the same size as
956 the Indian Plate, but the eastern and southeastern boundaries are very uncertain (Figure 14).
957 Magnetic anomaly maps of the plate are sparse on the southern boundary (i.e., Wharton Basin).
958 Bathymetry is complicated by several features including the Ninetyeast Ridge, Diamantina
959 Escarpment, Roo Rise, Vening Meinesz Seamounts, and the Raitt Rise. Earthquakes, while
960 indicating extension, are not of sufficient density to clearly delineate the boundaries (Figure 14).
961 Furthermore, the Indian and Australian Plates have relatively high rates of diffuse intraplate
962 seismicity, without clear clustering indicative of rigid plates boundaries.

963 The eastern, western, and northern boundary of the Lwandle Microplate are poorly con-
964 strained. Most authors have drawn the northern boundary as an extension of the Quasama
965 Seismic Axis on the southern end of the Rovuma Microplate to Madagascar and then traversing
966 Madagascar along a constant line of latitude before running along the eastern margin of the
967 island (e.g., [Saria et al., 2014](#)). However, there is little to no observed seismicity to constrain
968 this model. A more recent model suggests the northern boundary is constrained by a dextral
969 transform that runs along the Comoros archipelago which is consistent with seismicity and GPS
970 velocities in northern Madagascar ([King et al., 2017](#); [Famin et al., 2020](#); [Stamps et al., 2018,](#)
971 [2021](#)). For this reason, we prefer the northernmost boundary in our model, and have separated
972 the northern Lwandle region into the Comoros deformation zone.

973 The San Microplate was recently identified from geodetic studies of relative motions in Africa
974 ([Njoroge et al., 2015](#); [Wedmore et al., 2021](#)); however, the evidence is currently insufficient to

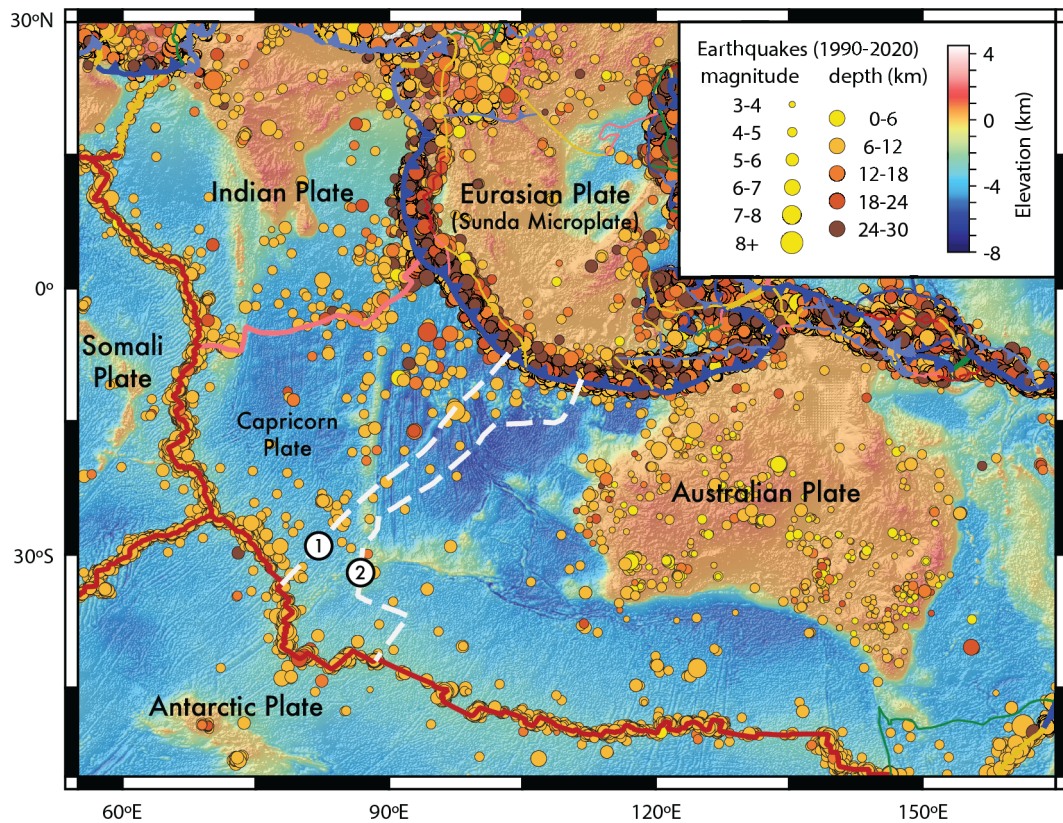


Figure 14: Models of the Capricorn Plate and associated Mid-Indian deformation zone. The white boundaries are a couple of the proposed southern boundaries: 1, (Rathnayake et al., 2019); 2, (Royer and Gordon, 1997).

975 justify its existence. The microplate shares its eastern boundary with the Lwandle Microplate.
 976 The northeastern boundary follows the southwest East African Rift which can be identified
 977 by mapped faults and seismicity (Poggi et al., 2017). However, the exact location of the
 978 boundary in southwestern Africa is difficult to trace as there is no Quaternary faults, seismicity
 979 is not above the intraplate background (Poggi et al., 2017) and there is no clear topographic
 980 expression. In the south Atlantic Ocean, Wedmore et al. (2021) infers the Walvis Ridge—a
 981 hotspot track (Sager et al., 2021)—marks the boundary. The lack of a clear shear-wave anomaly
 982 associated with the northern San Microplate boundary further weakens the case for a microplate
 983 boundary (Figure 13). Lastly, relative motions between the San Microplate and the African
 984 Plate are $<1 \text{ mm yr}^{-1}$ (Njoroge et al. (2015); Wedmore et al. (2021)). Such low velocities could
 985 result from a far-field long-wavelength glacial isostatic adjustment. As the case is currently
 986 weak, we include the San Microplate as part of the African Plate. We retain the polygons
 987 should the case be strengthened by future studies and indicate the boundary as inferred.

988 The boundary between the oceanic portion of the North and South American Plates east of
 989 the Lesser Antilles Trench is poorly defined by limited diffuse seismicity. In fact, the low level

990 of seismicity is similar to many intraplate regions not considered deformation zones. The lack
991 of GPS stations in both regions makes it difficult to define the edges of the deformation zone
992 reliably. Given the lack of a clear diffuse region, we draw the boundary as the most prominent
993 feature, the Fifteen-Twenty Fracture Zone, which is consistent with prior interpretations (Roest
994 and Collette, 1986; Dixon and Mao, 1997; DeMets et al., 2010).

995 4.4. Geologic Province Model

996 The geological province model includes 918 polygons, of which 790 (86%) are continental.
997 Most continental terranes are linear belts with areas that range between 10^4 and 10^6 km². The
998 median province area is 175,000 km², but varies between different continental regions from 10^5
999 and 3×10^5 km² (Figure 15a). The distribution of province sizes is largest in South America
1000 and Africa. Part of the reason may be due to the inability to pick out smaller terranes beneath
1001 thick sedimentary cover and relatively few studies with high resolution geophysical data. It is
1002 also due to the number of composite type terranes such as cratons, shields and orogens that are
1003 built from ophiolites, accretionary complexes, and volcanic arcs (Figure 15b). However, some
1004 provinces are naturally larger due to the processes involved in their formation (i.e., passive
1005 margins and wide rifts).

1006 The most common province types are orogenic belts, volcanic arcs and accretionary com-
1007 plexes (Figure 15b). There are many fewer extended terranes, however, passive margins are
1008 generally rifted margins with thick sediment on top. Many ancient rifts are incorporated into
1009 orogenic belts. Where orogenic belts have been separated into their individual constituents,
1010 there is often an accretionary complex and associated arc. However, in many cases only a vol-
1011 canic arc is identified, which may be because the related accretionary complex is commingled
1012 with the volcanic arc and difficult to separate or was destroyed during orogenesis.

1013 4.4.1. Comparison with Matthews et al. (2016) and Meredith et al. (2021)

1014 The model by Matthews et al. (2016) was constructed to model plate motions over the past
1015 400 Ma. Given their model timeframe, continental provinces that moved relative to one another
1016 prior to 400 Ma are not typically subdivided. As a result, our model contains a significantly
1017 greater number of terrane divisions. Many of the divisions in the Matthews et al. (2016) model
1018 are also simpler than ours. This simplicity requires fewer polygon vertices, which may make

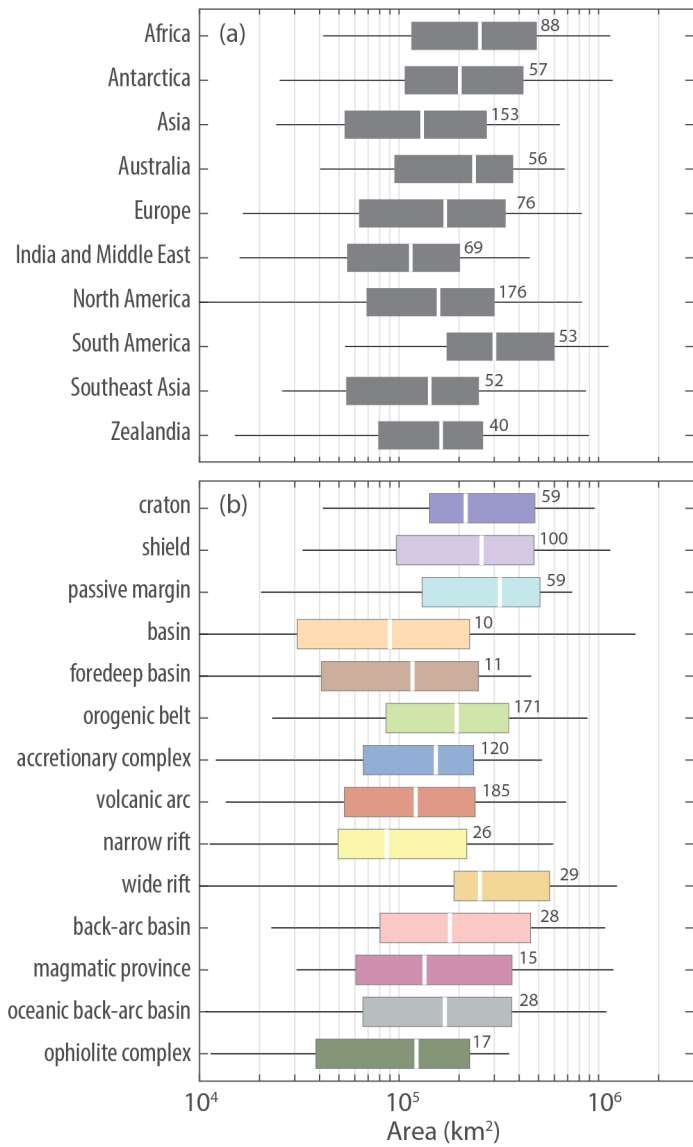


Figure 15: Area of geological provinces for (a) continental regions and by (b) province type. The white bar indicates the median, box enclose the 0.25 to 0.75 quantiles and whiskers extend to the 0.05 and 0.95 quantiles. The number beside the boxes indicates the number of provinces for each region.

1019 computation of plate rotations more rapid, but it compromises accuracy of filtering geologic
1020 data for analysis.

1021 The [Merdith et al. \(2021\)](#) model extends the [Matthews et al. \(2016\)](#) plate motions back
1022 to 1000 Ma, which required the addition of many more province boundaries. The models
1023 for southern Africa, the northwest Cordillera, and the Mongol Orocline are similar to ours.
1024 However, there are significant differences in the number of provinces in Australia, Antarctica,
1025 eastern Europe and the Superior Craton. Many of these differences occur where we include a
1026 finer resolution of terrane boundaries with tectonic histories that converge prior to 1000 Ma.
1027 In light of the agreement with previous reconstruction polygons, the model presented in this
1028 study could be used in the future for plate reconstructions into the Archean.

1029 *4.4.2. Comparison with CrustX models*

1030 We compare our province model with the model by [Laske et al. \(2013\)](#) (Table 6 and Fig-
1031 ure 16). Crust1.0 contains a model of crustal types updated from previous Crust5.1 and
1032 Crust2.0 ([Mooney et al., 1998](#); [Bassin et al., 2000](#); [Artemieva and Mooney, 2001](#); [Mooney,](#)
1033 [2007](#); [Laske et al., 2013](#)). We refer to the multiple generations of the USGS crustal models
1034 as CrustX. Crustal types are defined in [Mooney et al. \(1998\)](#). Most defined crustal types are
1035 similar between the two models. Oceanic crust, and forearc—accretionary complex in this
1036 study—are the only crustal types that have exact equivalencies. However, there are some key
1037 differences in definitions or grouping, with a few types exclusive to one province model or the
1038 other (Table 6).

1039 First, the slight lexical differences (Table 6). [Mooney et al. \(1998\)](#) define Precambrian
1040 regions as shields (exposed) and platforms (sediment covered). The update by [Laske et al.](#)
1041 [\(2013\)](#) almost entirely redefines the previous Crust5.1 platform regions as shield. We define
1042 cratons as regions with Archean cores and shields as regions with Paleo-Mesoproterozoic belts.
1043 In the thirty-some years since the first iteration of CrustX models, our understanding of Neo-
1044 proterozoic geology has improved sufficiently to classify many of these regions by their tectonic
1045 affiliation. [Mooney et al. \(1998\)](#) defines three types of extended crust: rifts, thinned continent
1046 and extended crust. It appears that only active rifts are defined as rifts. The precise differences
1047 between extended crust and thinned continent are not clear, but thinned continent has been
1048 almost completely removed in Crust1.0. In contrast to CrustX, our model separates rifts into

Table 6: Primary crustal types defined in Crust5.1 (Mooney et al., 1998) and the similar types from this study.

Crust5.1	This study	Comment
oceanic crust	oceanic crust	created by seafloor-spreading
anomalous o.c.	–	LIPs and hotspots; we suggest superimposing Johansson et al. (2018) if desired
–	ophiolite	generally below the resolution of Crust5.1
shield	craton or shield	Crust5.1 splits into Archean, early-middle Proterozoic, and late Proterozoic; late Proterozoic is classified as orogen or other types in this study
Phanerozoic rift	–	unknown type of Phanerozoic age
thinned continent	wide or narrow rift	thinned continent has been effectively removed from Crust1.0
extended crust	wide or narrow rift	unclear how this differs from thinned continent
transition	passive margin	Crust5.1 tile that covers oceanic and continental crust
shelf transition	passive margin	Crust5.1 tile that covers oceanic and continental crust
continental shelf orogen	passive margin or back-arc basin orogenic belt	orogens in this study are divided into other types if possible
platform	craton, shield, basin (rarely) or orogen	generally sediment covered craton/shield in crust5.1, applied inconsistently; Crust1.0 only contains platform in the West Siberian Basin and Arabia
–	foredeep basin	
continental plateau	–	not clear how this is defined, these are mixed crustal types
–	magmatic province	
continental arcs	volcanic arc	most arcs are below the resolution of Crust5.1
island arcs	volcanic arc	most arcs are below the resolution of Crust5.1
forearc	accretionary complex	not displayed in Mooney et al. (1998) Figure 1A

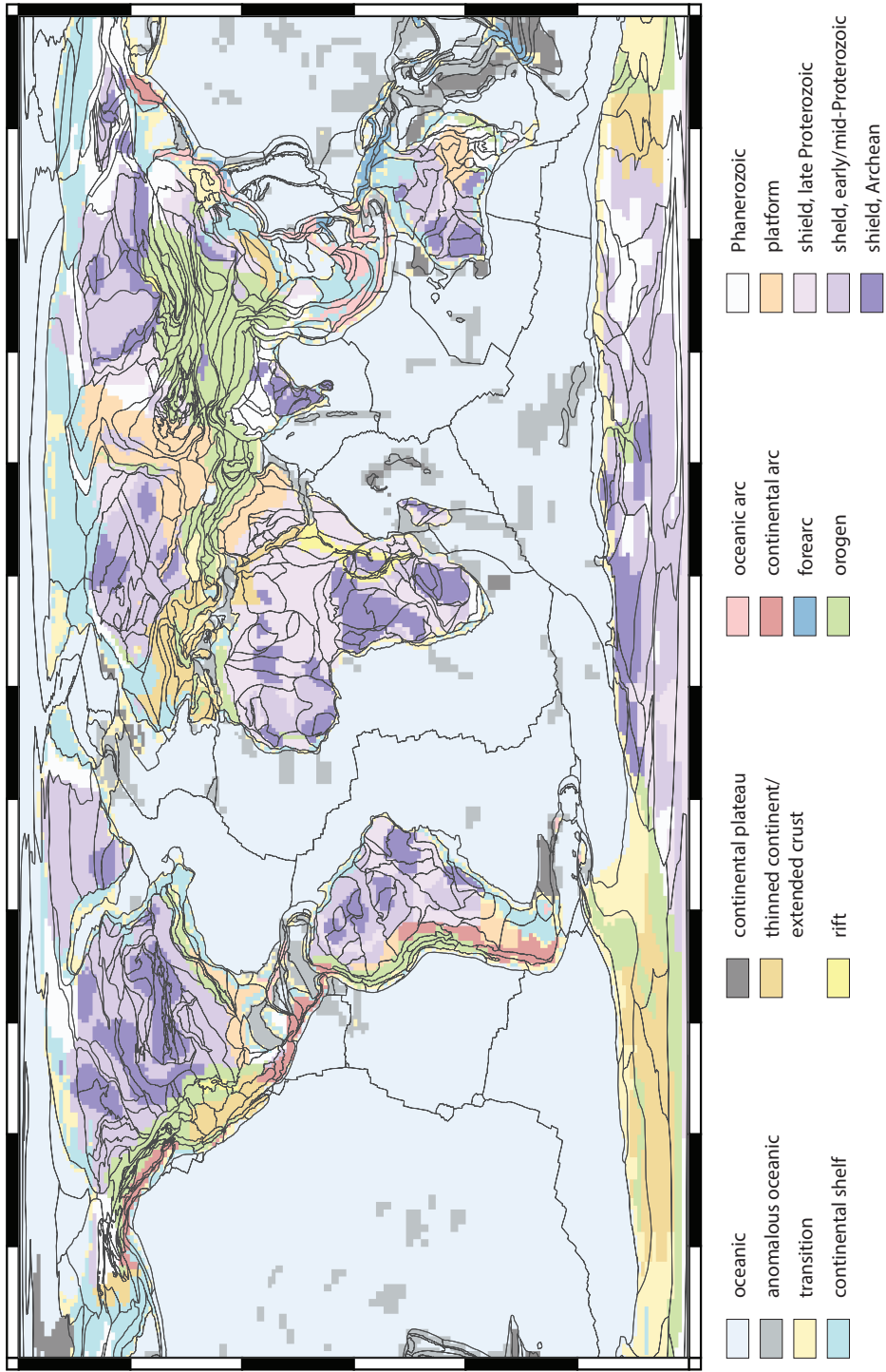


Figure 16: Crust1.0 model from [Laske et al. \(2013\)](#), with crustal types defined by [Mooney et al. \(1998\)](#). Geological province boundaries from this study are superimposed.

1049 narrow and wide rifts. Narrow rifts tend to occur in cold/strong Precambrian lithosphere (e.g.,
1050 East African Rift, Baikal Rift) whereas wide-rifts tend to occur in weak/warm lithosphere at
1051 the time of rifting (e.g., North American Basin and Range). Orogen and orogenic belt are
1052 defined similarly, though we break orogenic belts into other crustal types (e.g., ophiolite, ac-
1053 cretionary complex and volcanic arc) if a province largely retains its distinctive characteristics.
1054 [Mooney et al. \(1998\)](#) separates volcanic arcs into continental and island whereas we do not.
1055 Trace element patterns are very similar between oceanic and continental arcs, with continental
1056 arcs more fractionated and consequently more enriched in light trace elements ([Kelemen et al.,](#)
1057 [2007](#)).

1058 There are a number of larger differences between our model and CrustX (Table 6). Transi-
1059 tional crust and continental shelf in the CrustX model are generally defined as passive margins
1060 in our study, but the CrustX definitions also includes regions we define as continental back-arcs
1061 in our model (Figures 7 and 16). We also define oceanic back-arcs as regions with subduction
1062 modified spreading that occurs in the hinterland of a volcanic arc that are not included in the
1063 CrustX. Curiously, the CrustX models includes oceanic LIPs (anomalous oceanic crust) but not
1064 continental LIPs (Columbia River Basalts, Deccan Traps, Parana-Entedeka Basalts, Siberian
1065 Traps, etc.) except for a few that sit below sea-level (e.g., Campbell Plateau, Ninetyeast Ridge,
1066 and Seychelles). As previously discussed in Section 3.2.1, we do not include LIPs as part of our
1067 model. As CrustX was developed for making crustal corrections to tomography it makes sense
1068 that CrustX includes LIPs whereas our tectonic model does not. However, ophiolite complexes
1069 are not included in CrustX, but are within our model when they are large enough to be resolved
1070 (Figure 7). Magmatic provinces, also excluded from CrustX, are a bit of an enigma. They are a
1071 few regions of predominantly magmatic orogen, not classified as LIPs, and which do not appear
1072 to be related to volcanic arcs. While there are some significant differences, a CrustX-like model
1073 can be reproduced by combining elements of our geologic province model, the LIP model by
1074 [Johansson et al. \(2018\)](#), and a sediment cover by [Laske and Masters \(1997\)](#) or [Straume et al.](#)
1075 [\(2019\)](#).

1076 The visual differences between our models are considerable, though there is some broad
1077 agreement (Figures 7 and 16). Some of these differences are related to the precision of province
1078 boundaries, which are often one to two pixels inside or outside of Crust1.0 crustal types. Some of

1079 the other clear differences are the classification of late Proterozoic regions (shields in Crust1.0)
1080 classified by their tectonic affiliation. Likewise, many orogens are also classified by more specific
1081 tectonic settings where possible. Although our model is likely a more accurate representation of
1082 many province types since it incorporates many recent studies, we anticipate it will be improved
1083 as new geochronological sampling and geoscientific studies expand our understanding.

1084 *4.4.3. Last Orogeny*

1085 To ensure that our orogenic model is reasonable (Figures 8 and 9), we have computed kernel
1086 density estimates (KDE) for both magmatic crystallization dates and metamorphic dates for
1087 each orogen (Figure 17). The dates are mostly derived from U–Pb, ^{207}Pb – ^{206}Pb , and U–Th–Pb
1088 analyses of zircon and monazite. The zircon dates can provide constraints for both igneous and
1089 metamorphic events whereas most monazite dates record metamorphic events in the datasets
1090 we used (Eglington, 2004; Puetz, 2018; Brown and Johnson, 2018; Gard et al., 2019a). Other
1091 isotopic dating systems (e.g., ^{40}Ar – ^{39}Ar , K–Ar, and Rb–Sr) may record lower closure tempera-
1092 tures and can correspond to less significant heating events and are therefore generally excluded
1093 from the dataset except where the dates match the regional higher temperature metamorphism
1094 or in younger systems where U–Th–Pb lose precision.

1095 The largest age peak for metamorphic KDEs always falls within the range of dates attributed
1096 to an orogen as discussed in Section 3.2.2. It is important to note that these dates were not
1097 used to define the orogen, but undoubtedly many of these orogen models are based in part on
1098 geochronology. In Figure 8, one can see the spatial correspondence between metamorphic dates
1099 and our orogen interpretation. In a few cases, there are smaller-magnitude metamorphic date
1100 peaks that record P-T conditions of previous orogenic events. For example, in the Canadian
1101 Cordillera Frontal Thrust Belt a few dates are associated with the Wyoming Craton (ca. 1780
1102 Ma; Cheney et al., 2004b,a) and not the more recent Cordilleran deformation (Figure 8). These
1103 older metamorphic events are still evident because Cordilleran deformation reaches relatively
1104 shallow levels in these regions and has not reset the isotopic systems associated with the older
1105 events.

1106 Very few orogens have significant age peaks younger than the assigned orogen. The excep-
1107 tions include the Rodinian Orogeny which in some cases includes some Gondwana-associated
1108 dates, particularly in Antarctica, and the Circum-Nunanian terranes, which include some

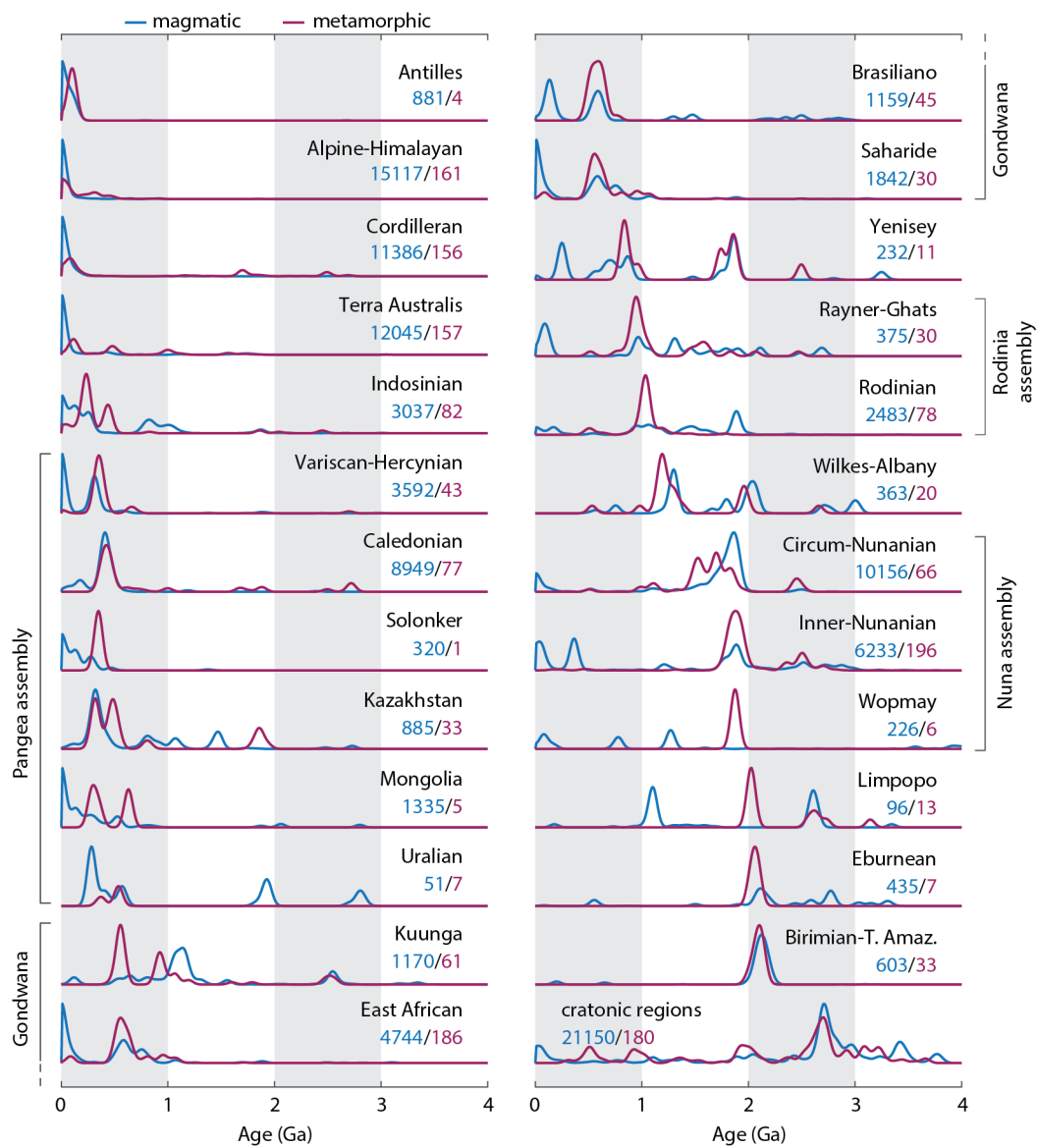


Figure 17: Kernel density estimates for magmatic crystallization dates (blue, Gard et al. (2019a)) and metamorphic dates (violet, Brown and Johnson (2018) updated; Eglington (2004)). A list of metamorphic dates are given in the Supplemental Material. No dates for the Scotian Orogeny are available in the database.

1109 Rodinia-associated overprinting (Figure 17)—specifically in North America where metamorphic
1110 conditions are poorly sampled (Figure 8). These discrepancies between the orogenic model and
1111 observed metamorphic conditions may indicate some refinement of the province model may be
1112 necessary.

1113 The magmatic dates are noisier, often indicating several episodes of magmatism, which
1114 attests to the multi-generational history to the growth of many provinces (Figure 17). There
1115 are a surprising number of provinces which have magmatic peaks near the present day that lie in
1116 generally much older orogens. In the East African Orogen, the younger ages are associated with
1117 the East African Rift. In the Saharides and Trans Hudson Orogen, these younger magmatic
1118 ages are the result of Cenozoic intraplate magmatism, which does not appear to be associated
1119 with active tectonics. In many regions with younger volcanism, the geodynamic process that
1120 has created the melts does not appear to have reset metamorphic conditions or alternatively
1121 the depths at which metamorphism has occurred have not yet been exhumed.

1122 While magmatic dates appear to be less reliable for identifying the last orogen event, there
1123 is often a peak in the KDE that corresponds with the metamorphic peak. The Limpopo Orogen
1124 is a clear exception where no magmatic peak is observed during the orogenic event. The missing
1125 magmatic peak at ca. 2.0 Ga is the reason the Limpopo orogen was originally interpreted as a
1126 ca. 2.7 Ga event, rather than a younger intracontinental orogen (Barton and van Reenen, 1992;
1127 Yin et al., 2019). The only other orogen with missing magmatic dates is the Wopmay Orogen
1128 in northwest Canada, however, this orogen is poorly exposed so there are very few data from
1129 the terranes that comprise it.

1130 The KDEs for the cratonic regions display significant magmatic and metamorphic activity
1131 prior to 2.5 Ga, with a few minor metamorphic and igneous peaks. Many of these small
1132 younger (<2.5 Ga) metamorphic peaks do not appear to correspond with magmatism, which
1133 may indicate minor degrees of intracratonic deformation as orogens occur around them. We
1134 suggest this observation is consistent with a persistent mechanically strong lithosphere able
1135 to resist deformation in most cases. The young magmatic peak within cratonic regions is
1136 common across most orogens. As most of these regions are intraplate several questions arise
1137 regarding the lack of widespread intraplate volcanism in the past. Are intraplate volcanics
1138 easily eroded and therefore poorly preserved? If not, is there a sampling bias, i.e., petrologists

1139 and geochemists vastly oversample present-day intraplate volcanism, or are they generally too
1140 small and distributed in the past to easily recognize and thus ignored?

1141 4.4.4. *Uncertain Province Boundaries*

1142 Perhaps the poorest geologically characterized region is Antarctica due to the lack of out-
1143 crops. Most geological province information about Antarctica comes from regions near the
1144 coast typically combined with constraints from conjugate terranes on now-distant continents
1145 (e.g., [Boger, 2011](#); [Flowerdew et al., 2013](#); [Goodge et al., 2017](#); [Maritati et al., 2019](#); [Rup-
1146 pel et al., 2020](#)). Due to the nature of these peripheral constraints, province boundaries are
1147 more uncertain into the subglacial interior and are typically guided by geophysical fields such
1148 as gravity and magnetics following linear trends ([Aitken et al., 2014](#); [Maritati et al., 2016](#)).
1149 Because of the uncertainties, it has resulted in a myriad of province models that can change
1150 significantly as additional geological data are added and higher resolution geophysical models
1151 are produced. Very few models include interior provinces that do not reach the coast because of
1152 the lack of geological control (cf., [Ferraccioli et al., 2011](#)), however new multivariate approaches
1153 to mapping lithospheric boundaries suggest significant interior complexity characterizes East
1154 Antarctica ([Stål et al., 2019](#)).

1155 Our Antarctic province model is derived from several geophysical constraints (Table 3). As
1156 with other continents, magnetic anomalies are often the best constraint on province bound-
1157 aries in the upper crust (Figure 18a). Geochronology is important for defining provinces and
1158 identifying boundaries in areas of outcrop, which are mostly concentrated near the coast. The
1159 PetroChron Antarctica database is a compilation of geochronology and other geological data of
1160 Antarctica ([Sanchez et al., 2021](#)), that will surely continue to improve our insights into the pre-
1161 cise positioning of Antarctic province boundaries. Our model includes a few interior provinces
1162 that we interpret as cratonic or shield terranes due to the lithospheric characteristics, though
1163 they are relatively large. These terranes are ringed by orogens at the edges of the continent,
1164 similar to many other continental shields (e.g., Kaapvaal, Congo, and Siberian Cratons).

1165 Most other poorly resolved regions include those covered by thick sedimentary cover: Patag-
1166 onia, Parana Basin, Saudi Arabian Platform, North Africa, and West Siberian Basin. In these
1167 regions the provinces tend to be much larger than the average and there are multiple in-
1168 terpretations for their divisions. These models are difficult to independently assess because

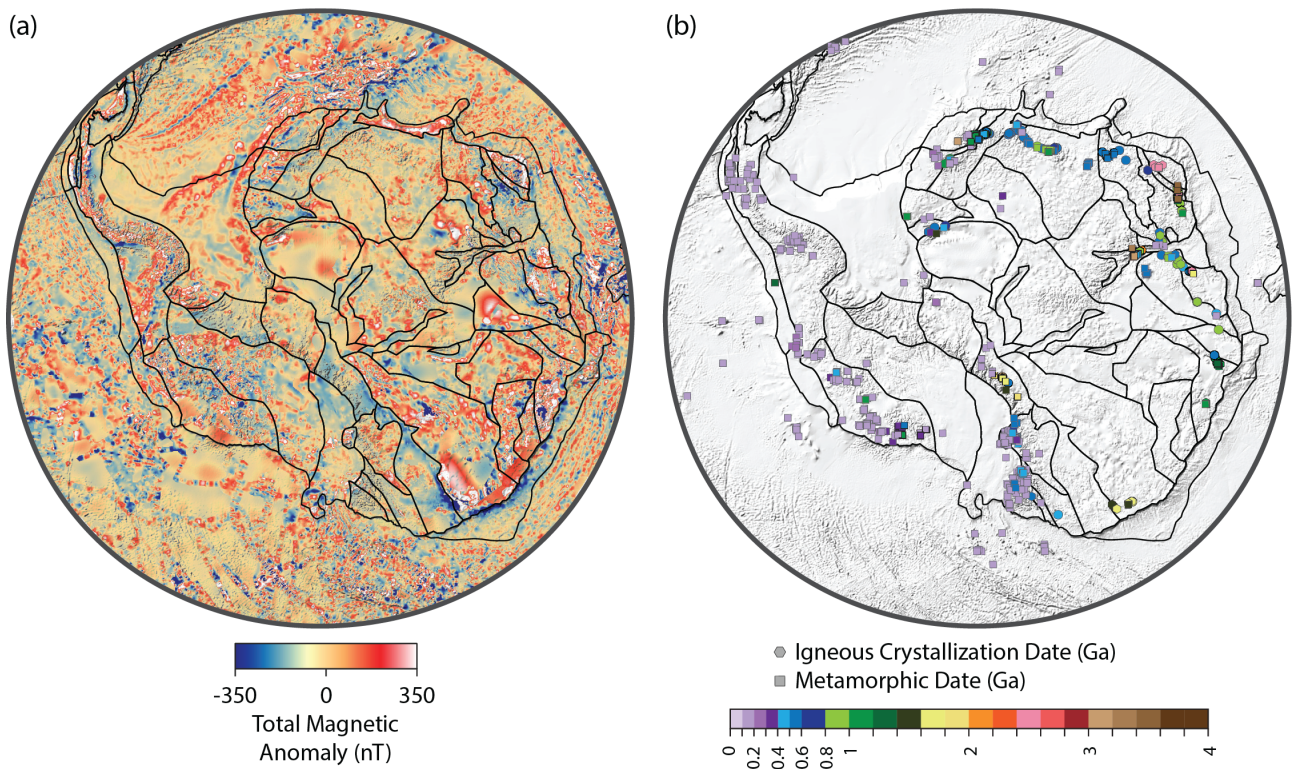


Figure 18: Magnetic (a) and geochronologic (b) constraints on the Antarctic province model. Magnetic data from ADMAP2 (Golynsky et al., 2018) and EMAG2_V3 (Meyer and Saltus, 2016) to fill gaps. Igneous crystallization dates from Gard et al. (2019a) and Puetz (2018) and metamorphic dates from Brown and Johnson (2018), updated, and DateView, (Eglington, 2004).

1169 the geophysical data in many of these regions is relatively poor. For example, the Saharan
 1170 Metacraton and Tuareg Shield have considerably different interpretations for their evolution.
 1171 The Tuareg Shield is constructed of a set of imbricated arc terranes. Most models show lit-
 1172 tle to no association with the Saharan Metacraton, which is interpreted as a set of Archean
 1173 cores with reworked subcontinental lithosphere (Liégeois et al., 2013; Sobh et al., 2020). This
 1174 interpretation is based on thermophysical modeling using multiple geophysical data, but the
 1175 data are relatively low resolution. An alternative explanation developed using geochronological
 1176 and geochemical data suggests the Tuareg Shield, Arabian–Nubian Shield and portions of the
 1177 Saharan Metacraton were a single volcanic arc system that was broken in multiple segments,
 1178 reorganized and then accreted to the core of the metacraton, deforming it in the process in the
 1179 Neoproterozoic (Liégeois, 2018; Şengör et al., 2020; Blades et al., 2021).

1180 4.5. Future Improvements

1181 There are a number of improvements that can be made to the province definitions in future
 1182 versions. Higher resolution geologic maps could be incorporated that have better lithological
 1183 data and finer age resolution in the Precambrian, especially the Archean. Likewise, higher

1184 resolution magnetic data will be very useful for improving terrane boundaries, especially when
1185 the resolution is ~ 10 km pixel spacing. Most global geophysical datasets are available at too
1186 low a resolution to precisely resolve many terrane boundaries.

1187 Province types that do not provide information about the tectonic construction of the crust
1188 (basin, shield, and craton), should be replaced in future versions. Doing so may require im-
1189 proved geologic models of Precambrian regions, which may require the addition of new tectonic
1190 settings that are unique to the early Earth such as granite-greenstone belts. Likewise, orogenic
1191 belts could also be deconstructed into the types of terranes that build them, i.e., separating out
1192 arcs and accretionary margins could help improve models of crustal growth and composition.

1193 Although we currently only ascribe the last orogenic event to provinces, many regions have
1194 experienced multiple orogenic events throughout their history. Adding this history to the at-
1195 tribute tables would provide a more complete view of an individual province's chronological his-
1196 tory, which could be paired with information about the changing tectonic environments. Such
1197 a model would require a more sophisticated analysis of global geochemical and geochronological
1198 data. This process would be tedious to perform on a province by province scale so an automated
1199 approach will be warranted and could include recent machine-learning based approaches to pre-
1200 dicting tectonic environments (e.g., [Tang et al., 2020](#); [Tetley et al., 2020](#)). These improvements
1201 would place valuable constraints on plate reconstruction models.

1202 Constructing the province models can include a community-driven approach that is updated
1203 by individuals or small groups with expertise in specific regions. To that end, we have set
1204 up a GitHub project page where individuals can download the shapefiles, update them, and
1205 upload their changes along with a description, references, and rationale for the changes. The
1206 changes can then be evaluated and incorporated into the model if deemed credible. Thus, one
1207 would not need to wait for a formally published reference before using the improved maps
1208 (https://github.com/dhasterok/global_tectonics).

1209 **5. Summary**

1210 We have produced a set of plate and geologic province maps that can be used to improve spa-
1211 tial analysis of geoscientific data. The plate model boundaries are validated against earthquake
1212 locations, active fault traces, and GPS motions and shows good correlation with shear wave
1213 velocities at 70 km depth and active volcanism. The geological province model is constructed

1214 from a collage of published models and refined using a wide-variety geophysical and geological
1215 data. The most useful data were found to be aeromagnetic anomalies when the models are high
1216 resolution. The province model has not been independently validated at this point, but relies
1217 on the accuracy of the original studies. However, the last orogeny is validated using metamor-
1218 phic and igneous dates. The plate and province polygons are drawn so that the boundaries are
1219 seamless between the files and additionally include an ocean–continent boundary and a plate
1220 boundary type.

1221 The maps are available in a shapefile format that can be easily interpreted by many modern
1222 computer languages and have advantages over raster maps for geographically selecting data
1223 for various types of analysis. The maps can also be used as a data standard for prescribing
1224 spatial metadata in global databases. Because the models are available on GitHub, the geologic
1225 community can submit updates and fixes to improve their accuracy.

1226 **Data Availability**

1227 Plate and province models produced in this study are available at the GitHub repository,
1228 https://github.com/dhasterok/global_tectonics. The models can be found in shapefile
1229 format suitable for GIS programs, KML for programs such as GoogleEarth, and GMT format for
1230 Generic Mapping Tools. The models are also available in the global tectonics library on Zenodo,
1231 <http://doi.org/10.5281/zenodo.5093930>, which includes additional global geophysical and
1232 geochronological datasets that are useful for research and educational applications.

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