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# Three-Station Interferometry and Tomography: Coda vs. Direct Waves

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## Summary

Traditional two-station ambient noise interferometry estimates the Green's function between a pair of synchronously deployed seismic stations. Three-station interferometry considers records observed three stations at a time, where two of the stations are considered receiver-stations and the third is a source-station. Cross-correlations between records at the source-station with each of the receiver-stations are correlated or convolved again to estimate the Green's function between the receiver-stations, which may be deployed asynchronously. We use data from the EarthScope USArray in the western US to compare Rayleigh wave dispersion obtained from two-station and three-station interferometry. Three three-station interferometric methods are distinguished by the data segment utilized (coda-wave or direct-wave) and whether the source-stations are constrained to lie in stationary phase zones approximately inline with the receiver-stations. The primary finding is that the three-station direct wave methods perform considerably better than the three-station coda-wave method and two-station ambient noise interferometry for obtaining surface wave dispersion measurements in terms of signal-to-noise ratio, bandwidth, and the number of measurements obtained, but possess small biases relative to two-station interferometry. We present a ray-theoretic correction method that largely removes the bias below 40 s period and reduces it at longer periods. Three-station direct-wave interferometry provides substantial value for imaging the crust and uppermost mantle, and its ability to bridge asynchronously deployed stations may impact the design of seismic networks in the future.

**Key words:** Seismic noise; Seismic interferometry; Seismic tomography; Surface waves and free oscillations; Coda waves.

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## 1 Introduction

Inter-station seismic interferometry is designed to extract an estimate of the Green’s function between pairs of seismic stations or receivers. Generally speaking, there are two established methods to perform this task, which we will call “*two-station interferometry*” and “*three-station interferometry*”. In this paper, we attempt to discuss and characterize important variants of three-station interferometry, and compare the characteristics amongst the variants and to two-station interferometry using data from the EarthScope Transportable Array (TA) in the US.

Two-station interferometry is the traditional method of “*ambient noise interferometry*” or “*ambient noise correlation*”. It is the more commonly applied method and is based on a single cross-correlation between ambient noise recorded at two stations. The cross-correlation can be converted to an estimate of the Green’s function of the medium if the time series is long enough (e.g., [Shapiro & Campillo, 2004](#)). In this case, one of the stations acts as a virtual source of the seismic energy and the other as the receiver. When many pairs of stations are considered, it is the basis for *ambient noise tomography* of surface waves, and many applications of this method have emerged since [Shapiro et al. \(2005\)](#); [Sabra et al. \(2005\)](#); [Yao et al. \(2006\)](#).

Three-station interferometry, in contrast, considers recordings from three seismic stations at a time. This method takes the cross-correlation between recordings of ambient noise at one station, which acts as a virtual source and which we call the “*source-station*”, with recordings from two other stations, which are called the “*receiver-stations*”. These two cross-correlations, or particular segments of them, are then cross-correlated again (or, as discussed further below, convolved). Stacking the resulting waveforms from many source-stations for the same pair of receiver-stations provides an estimate of the Green’s function between the two receiver-stations. This method, therefore, is based on cross-correlations performed three at a time, where the last one has been referred to as the “*correlation of correlations*” ([Stehly et al., 2008](#)) but in certain circumstances will be a convolution of correlations. We refer to this method generally speaking as “*three-station interferometry*”, to distinguish it from traditional two-station ambient noise methods. When the final cross-correlation is between the coda-wave parts of the first two correlations the method is commonly referred to as the “*correlation of the coda of correlations*” or  $C^3$  ([Stehly et al., 2008](#)).

60 **Fig. 1** illustrates some of the notation introduced in this paper. For two-station  
 61 interferometry, we denote the cross-correlation between a pair of seismograms observed  
 62 at stations  $r_i$  and  $r_j$  as  $C_2(r_i, r_j)$ . With an appropriate phase-shift,  $C_2(r_i, r_j)$  can be con-  
 63 verted to an estimate of the Green’s function between the two stations,  $\hat{G}_2(r_i, r_j)$ , where  
 64 we suppress the time-dependence of the correlations and the estimated Green’s function.  
 65 For three-station interferometry, cross-correlations between observations at a source-station,  
 66  $s_k$  ( $1 \leq k \leq N$ ), with the two receiver-stations,  $C_2(s_k, r_i)$  and  $C_2(s_k, r_j)$ , are corre-  
 67 lated again (or in some circumstances convolved). This produces the three-station “*source-*  
 68 *specific interferogram*”,  $C_3(r_i, r_j; s_k)$ , for source-station  $s_k$ . (The subscript “3” distin-  
 69 guishes the final cross-correlation or convolution from the first two correlations.) The  
 70 “*composite Green’s function*” for three-station interferometry is produced by taking a  
 71 weighted sum over the contributing source-specific interferograms from the  $N$  source-  
 72 stations:

$$\hat{G}_3(r_i, r_j) = \sum_{k=1}^N w_k C_3(r_i, r_j; s_k) \quad (1)$$

73 where  $w_k$  is a weight.  $\hat{G}_3$  provides information about the medium between the two receiver-  
 74 stations. For this equation to hold,  $C_3$  must have an appropriate phase-shift applied prior  
 75 to the summation.

76 The advantages of two-station interferometry include its simplicity and general ap-  
 77 plicability. The principal advantage of the three-station method over the two-station method  
 78 is that the two receiver-stations do not have to operate at the same time, although they  
 79 do have to operate synchronously with each source-station for some length of time. Thus,  
 80 three-station interferometry can be applied to asynchronously deployed stations (Ma &  
 81 Beroza, 2012; Curtis et al., 2012), which provides the opportunity for what Curtis et al.  
 82 (2012) call “retrospective seismology”. In terms of applications, the method will be most  
 83 impactful in settings where there is a long-term backbone seismic network to provide the  
 84 source-stations and shorter term deployments from which the receiver-stations are taken.

85 In practice, the data processing involves three noteworthy subtleties. (1) The cross-  
 86 correlations of seismic noise data that form the basis for both the two-station and three-  
 87 station methods involve refined data processing methods that aim to speed convergence  
 88 and reduce sensitivity to earthquakes and localized persistent noise sources (e.g., Ritz-  
 89 woller & Feng, 2019). We discuss the methods of data processing that we use in **sections**

90 **2 and 3** below, but we do not attempt to optimize data processing procedures for three-  
 91 station interferometry.

92 (2) We must specify which parts of the cross-correlations of seismic noise,  $C_2(s_k, r_i)$   
 93 and  $C_2(s_k, r_j)$ , that are correlated or convolved to produce the source-specific interfer-  
 94 ogram for source  $s_k$ ,  $C_3(r_i, r_j; s_k)$ . **Fig. 2** identifies the two parts of the cross-correlations  
 95 relevant to this study: the coda-wave (CW) and the direct-wave (DW) parts. If coda waves  
 96 are correlated, we refer to the method to produce an estimated Green’s function as “*coda-*  
 97 *wave interferometry*” and if direct waves are correlated or convolved we call it “*direct-*  
 98 *wave interferometry*”.

99 (3) Finally, it is important to specify how to determine the weights,  $w_k$ , that con-  
 100 vert individual source-specific interferograms to the estimated Green’s function. One as-  
 101 pect of the choice of weights is the geometrical relationship between the receiver-stations  
 102 and each source-station. For coda-wave interferometry there is no geometrical constraint  
 103 so that all source-stations are used for a given receiver-station pair irrespective of their  
 104 relative position; that is, the geometrical-weights are all unity (**Fig. 3a**). However, for  
 105 direct-wave interferometry we impose the constraint that the source-stations lie within  
 106 appropriately defined “stationary phase zones” so that sources outside those zones are  
 107 given zero geometrical-weight and sources inside the zones are given unit geometrical-  
 108 weight. The stationary phase zone is a Fresnel ellipse for source-stations between the receiver-  
 109 stations (**Fig. 3c**) or hyperbolae for source-stations not between the receiver-stations  
 110 (**Fig. 3b**), where the receiver-stations are the foci of both the ellipse and the hyperbo-  
 111 lae. Another aspect of these weights is based on a measure of the quality of each source-  
 112 specific interferogram,  $C_3(r_i, r_j, s_k)$ . Both aspects of assigning weights are discussed in  
 113 greater detail in **section 3.2**.

114 It is useful to define nomenclature to distinguish the interferometric methods con-  
 115 sidered here. Traditional two-station ambient noise (AN) interferometry is denoted:

$$\mathcal{I}_2^{AN},$$

116 where the “2” represents the number of stations used. Three-station methods require the  
 117 specification of two additional fields, “*type*” and “*geometry*”, so that three-station inter-  
 118 ferometric methods are denoted generally as:

$$geometry\mathcal{I}_3^{type}.$$

119 Here, “type” indicates either coda-wave (CW) or direct-wave (DW) interferometry, “ge-  
 120 ometry” represents the shape of the stationary phase zone, and the “3” indicates the num-  
 121 ber of stations used in the method prior to stacking over source-stations. Of course, in  
 122 the stacking of eq. (1) multiple source-stations will typically be used, but data analy-  
 123 sis is performed three stations at a time. There is no geometrical constraint for coda-  
 124 wave interferometry; thus this field is left blank in this case. For direct-wave interfer-  
 125 ometry the geometrical constraint is either an ellipse (ell) or a hyperbola (hyp).

126 Therefore, we identify three general methods of three-station interferometry to es-  
 127 timate Green’s functions. First, three-station coda-wave interferometry is denoted as

$$\mathcal{I}_3^{CW}.$$

128 Hence, there is the following relationship between our notation and earlier notation:  $\mathcal{I}_3^{CW} \equiv$   
 129  $C^3$ . Second, three-station direct-wave interferometry with sources in the elliptical sta-  
 130 tionary phase zone between the receiver-stations is represented as

$$ell\mathcal{I}_3^{DW}.$$

131 Finally, we indicate three-station direct-wave interferometry with sources in the hyper-  
 132 bolic stationary phase zones radially outside the receiver-stations as

$$hyp\mathcal{I}_3^{DW}.$$

133 When we refer to direct-wave interferometry generally without distinguishing between  
 134 the geometry of the stationary phase zones, we will use the symbol  $\mathcal{I}_3^{DW}$ , leaving the ge-  
 135 ometry field blank.

136 Three-station coda-wave interferometry ( $\mathcal{I}_3^{CW}$ ) was initiated by [Stehly et al. \(2008\)](#)  
 137 and has been fairly well studied ([Garnier & Papanicolaou, 2009](#); [Froment et al., 2011](#);  
 138 [Ma & Beroza, 2012](#); [Zhang & Yang, 2013](#); [Haendel et al., 2016](#); [Sheng et al., 2017, 2018](#);  
 139 [Spica et al., 2017](#); [Ansaripour et al., 2019](#)). Applications of  $\mathcal{I}_3^{CW}$  to surface wave tomog-  
 140 raphy or 3-D model construction remain rare, however, in particular at regional or con-  
 141 tinental scales. To the best of our knowledge, the principal exceptions are two studies  
 142 that combine group velocity measurements from  $\mathcal{I}_3^{CW}$  with traditional ambient noise in-  
 143 terferometry ( $\mathcal{I}_2^{AN}$ ) to improve 3-D models of Mexico and the southern US ([Spica et al.,](#)  
 144 [2016](#)), and of the Iranian Plateau ([Ansaripour et al., 2019](#)).

145 In comparison, three-station direct-wave interferometry ( $\mathcal{I}_3^{DW}$ ) has received much  
 146 less attention. [Froment et al. \(2011\)](#) discussed the possibility for using direct versus coda

147 waves, and differentiated between two types of correlations of correlations:  $C_{coda}^3$  and  
 148  $C_{all}^3$ , where  $C_{coda}^3$  denotes the correlation of the coda of correlations and  $C_{all}^3$  refers to  
 149 correlating the entirety of the correlations. Thus, as noted above, their  $C_{coda}^3$  is similar  
 150 to our  $\mathcal{I}_3^{CW}$  and because the direct-waves dominate the coda-waves in the correlations,  
 151 their  $C_{all}^3$  is in some ways similar to our  $\mathcal{I}_3^{DW}$ . They, however, do not discuss constrain-  
 152 ing the source-stations in direct-wave interferometry to lie in stationary phase zones, al-  
 153 though other studies do (Curtis & Halliday, 2010; Duguid et al., 2011; Curtis et al., 2012;  
 154 Entwistle et al., 2015). Moreover, the latter studies also recognize that for the ellipti-  
 155 cal stationary phase zone, when source-stations lie generally between the receiver-stations,  
 156 the original cross-correlations should be convolved with one another rather than cross-  
 157 correlated. Therefore, for  $^{hyp}\mathcal{I}_3^{DW}$  the three data operations are all cross-correlations, but  
 158 for  $^{ell}\mathcal{I}_3^{DW}$  the third data operation is a convolution. Discussion of the role of convolu-  
 159 tion in interferometry goes back at least to Slob and Wapenaar (2007). Entwistle et al.  
 160 (2015) applied aspects of direct-wave interferometry to data from the EarthScope Trans-  
 161 portable Array, but to the best of our knowledge  $\mathcal{I}_3^{DW}$  has not yet been applied tomo-  
 162 graphically or in the context of inversions for 3-D models and its properties remain poorly  
 163 understood.

164 The purpose of this paper is to determine and compare empirically the character-  
 165 istics of the three-station methods to each other and to two-station interferometry. In  
 166 particular, we focus on obtaining reliable surface wave dispersion measurements in the  
 167 context of tomography. From the outset, it is evident that coda-wave interferometry has  
 168 the advantage that any geometrical relationship can exist between the source-stations  
 169 and the receiver-stations, whereas for direct-wave interferometry only a small subset of  
 170 stations can be used as source-stations for each pair of receiver-stations. In coda-wave  
 171 interferometry, however, signals emerge very slowly with the addition of source-stations,  
 172 which means that many more source-stations are needed to recover reliable estimated  
 173 Green’s functions. Therefore, the relative merits of direct-wave interferometry and coda-  
 174 wave interferometry (which of the methods will be preferable, in what ways, and in which  
 175 settings) need to be determined empirically.

176 We address these questions by applying  $\mathcal{I}_2^{AN}$ ,  $\mathcal{I}_3^{CW}$ ,  $^{ell}\mathcal{I}_3^{DW}$ , and  $^{hyp}\mathcal{I}_3^{DW}$  across the  
 177 central and western US to all stations west of 95°W longitude from the EarthScope Trans-  
 178 portable Array to measure Rayleigh wave dispersion from 8 s to 80 s period and present  
 179 associated phase speed maps from 10 s to 60 s period. We pay particular attention to

180 the agreement between the three-station results and the two-station results, including  
 181 systematic differences (bias) and fluctuation, and to the distributions of measurements  
 182 as functions of signal-to-noise ratio (SNR), band-width, and the number of measurements  
 183 produced for asynchronously deployed receiver-stations.

## 184 2 Data

185 Three-station interferometry ( $\mathcal{I}_3$ ) is based on data output from two-station inter-  
 186 ferometry ( $\mathcal{I}_2$ ). As the basis for the three-station interferometry in this study, we use  
 187 the two-station database of ambient noise cross-correlations ( $C_2$ ) constructed by [Shen](#)  
 188 [and Ritzwoller \(2016\)](#). Stations in the database of [Shen and Ritzwoller \(2016\)](#) extend  
 189 across the contiguous US, but we use only a subset of them in the central and western  
 190 US (west of 95°W longitude), which defines our region of study ([Fig. 4](#)). We use all 1047  
 191 EarthScope USArray stations in this region deployed from 2005 to 2010, including 979  
 192 Transportable Array (\_US-TA) stations and 68 Reference Network (\_US-REF) stations.  
 193 We retain a two-station cross-correlation only if its signal-to-noise ratio (SNR) is greater  
 194 than 10, where SNR is defined as the ratio of the maximum amplitude of the waveform  
 195 in the time window of the direct fundamental Rayleigh wave to the root-mean square  
 196 of the waveform in the coda-wave window ([Fig. 2](#)). SNR defined in this way is indepen-  
 197 dent of frequency. Among the 547,581 possible combinations of pairs from the 1047 sta-  
 198 tions, 66% (364,103) operated synchronously so that two-station ambient noise interfer-  
 199 ometry could be employed. Of these, we retained 325,446 (89%) cross-correlations that  
 200 met the SNR criterion. In contrast, 34% (183,478) of the station-pairs were deployed asyn-  
 201 chronously.

202 The deployment of the Transportable Array started from the West Coast and rolled  
 203 eastward, with stations deployed temporarily for  $\sim 2$  years ([Fig. 4](#)). This rolling pat-  
 204 tern provides an ideal geometry for direct-wave interferometry with an elliptical station-  
 205 ary phase zone,  $^{ell}\mathcal{I}_3^{DW}$ , in which source-stations lie approximately between receiver-stations.  
 206 In contrast, the Reference Network was deployed permanently and was scattered across  
 207 the US with a station spacing of  $\sim 300$  km. This is a good geometry for coda-wave in-  
 208 terferometry,  $\mathcal{I}_3^{CW}$ , and direct-wave interferometry with a hyperbolic stationary phase  
 209 zones,  $^{hyp}\mathcal{I}_3^{DW}$ , in which source-stations lie approximately radially outward from receiver-  
 210 stations.

211 Shen and Ritzwoller (2016) used a common method of ambient noise data process-  
 212 ing (Bensen et al., 2007). Briefly, continuous records of vertical component seismograms  
 213 are cut to day-long segments and downsampled from 40 Hz to 1 Hz. Then the instru-  
 214 ment response, mean and trend are removed. To minimize the effects of strong directional  
 215 sources (in particular earthquakes) and to broaden the usable bandwidth, temporal nor-  
 216 malization and spectral whitening are applied. The temporal normalization uses a 80 s  
 217 running time window, which strongly attenuates signals with periods above 80 s. For this  
 218 reason we will focus our interpretation on measurements only up to 80 s period and show  
 219 tomographic results only up to 60 s period.

220 After pre-processing, daily seismograms from all available combinations of station-  
 221 pairs  $(r_i, r_j)$  are cross-correlated to produce  $C_2(r_i, r_j)$ , between correlation lag times of  
 222  $\pm 3000$  s. Daily correlations are then stacked to generate two-station estimated Green’s  
 223 functions between each pair of stations ( $\hat{G}_2(r_i, r_j)$ ). Finally, we compute the so-called  
 224 “symmetric component” of the estimated Green’s function by averaging the estimated  
 225 Green’s function at positive and negative correlation lags for simplicity. We will also re-  
 226 fer to this symmetric component estimated Green’s function as  $\hat{G}_2(r_i, r_j)$ , even though  
 227 it is defined only for positive lag. This database of symmetric component estimated Green’s  
 228 functions is the basis for the three-station analysis (**section 3**).

### 229 **3 Data Processing for Three-Station Interferometry**

230 The input for three-station interferometry are the two-station symmetric compo-  
 231 nent cross-correlations (or estimated Green’s functions) taken from the database of Shen  
 232 and Ritzwoller (2016) with SNR  $> 10$ . As inter-station cross-correlations, these func-  
 233 tions are denoted by  $C_2$  and as estimated Green’s functions by  $\hat{G}_2$ . Three-station source-  
 234 specific interferograms ( $C_3$ ) are cross-correlations of the coda-wave parts of the inter-  
 235 station cross-correlations, or cross-correlations or convolutions of the direct-wave parts  
 236 of the inter-station cross-correlations. Three-station data processing aims to compute  
 237 the composite Green’s function between pairs of receiver-stations by stacking the three-  
 238 station interferograms over contributions from various source-stations.

239 For concreteness, consider a receiver-station pair  $(r_i, r_j)$  and a set of source-stations,  
 240  $\{s_k\}_{k=1}^N$ , that operate synchronously with both  $r_i$  and  $r_j$  at least for some time. **Fig.**  
 241 **1b** depicts this situation, where one source-station is shown. Let the coda-wave parts

242 of the two-station cross-correlations be denoted  $C_2^{CW}(s_k, r_i)$  and  $C_2^{CW}(s_k, r_j)$ , and the  
 243 direct-wave parts be written  $C_2^{DW}(s_k, r_i)$  and  $C_2^{DW}(s_k, r_j)$ , where the coda-wave and  
 244 direct-wave segments are defined in **Fig. 2**. The three-station data processing procedure  
 245 breaks into three principal steps (**sections 3.1 - 3.3**).

### 246 3.1 Constructing Source-Specific Interferograms

247 The first step in three-station data processing is devoted to cross-correlating or con-  
 248 volving segments of the two-station cross-correlations. It is broken into three categories  
 249 depending on whether one considers the direct- or coda-wave segments of the two-station  
 250 cross-correlations and the geometrical relationship between the receiver-station pair and  
 251 each source-station. For direct-waves, the geometrical relationship is summarized in terms  
 252 of hyperbolic or elliptical stationary phase zones (**Fig. 3b,c**).

253 (1) The first category is, for each source-station, to compute the three-station source-  
 254 specific interferograms based on the coda-waves in the two-station cross-correlations. That  
 255 is, correlate  $C_2^{CW}(s_k, r_i)$  and  $C_2^{CW}(s_k, r_j)$  for all  $s_k$  to produce  $C_3^{CW}(r_i, r_j; s_k)$  for  $1 \leq$   
 256  $k \leq N$ . An example record-section containing three-station coda-wave source-specific  
 257 interferograms is presented in **Fig. 5a**, where each trace is for a separate source-station.

258 (2) The second category is to compute the three-station source-specific interfero-  
 259 grams based on the direct-waves in the two-station cross-correlations for the source-stations  
 260 in the hyperbolic stationary phase zones. For each source-station  $s_k$  in the stationary-  
 261 phase hyperbolae for the receiver-station-pair, cross-correlate  $C_2^{DW}(s_k, r_i)$  and  $C_2^{DW}(s_k, r_j)$   
 262 to produce  $^{hyp}C_3^{DW}(r_i, r_j; s_k)$ . An example record-section for three-station direct-wave  
 263 source-specific interferograms computed by cross-correlation is shown in **Fig. 5b**, where  
 264 each trace is for a separate source-station. For this record-section, cross-correlations are  
 265 computed based on source-stations irrespective of whether they lie in the stationary-phase  
 266 hyperbolae. However, the green-shaded regions identify the stationary phase zones.

267 (3) The third category is similar to the second, but we compute the three-station  
 268 source-specific interferograms based on the direct-waves in the two-station cross-correlations  
 269 for the source-stations in the elliptical stationary phase zone. For each source-station  $s_k$   
 270 in the stationary-phase ellipse for this receiver-station-pair, convolve  $C_2^{DW}(s_k, r_i)$  and  
 271  $C_2^{DW}(s_k, r_j)$  to produce  $^{ell}C_3^{DW}(r_i, r_j; s_k)$ . An example record-section for three-station  
 272 source-specific direct-wave interferograms computed by convolution is shown in **Fig. 5c**,

273 where each trace is for a separate source-station. As in **Fig. 5b**, convolutions are pre-  
 274 sented irrespective of whether the source-station lies in the stationary-phase ellipse, but  
 275 the green-shaded region identifies the stationary phase zone.

276 Convolution of the direct-wave parts of the two-station records when source-stations  
 277 lie in the elliptical stationary phase zone has been formally justified by other studies (Halliday  
 278 & Curtis, 2009; Curtis & Halliday, 2010). We provide a heuristic argument for illumi-  
 279 nation. When a source-station lies radially outward from a pair of receiver-stations, it  
 280 is the time-difference between the travel times from the source-station to the two receiver-  
 281 stations that approximates the travel time between the two receiver-stations. Cross-correlation  
 282 of two records finds the time-difference between them, therefore when source-stations lie  
 283 outside the receiver-stations it is the appropriate method to apply. In contrast, convo-  
 284 lutions find the sum of the times. When a source-station lies between two receiver-stations,  
 285 we wish to find the sum of the times from the source-station to each receiver-station, so  
 286 that convolution is the appropriate method to apply in this case.

287 We define the hyperbolic and elliptical stationary phase zones in a straightforward  
 288 and simplified manner. An ellipse is defined as the locus of points where the sum of the  
 289 distances to the foci is constant. Let  $d_{ij}$  be the great-circle distance between the two receiver-  
 290 stations,  $d_{ki}$  be the distance between a point  $s_k$  on the ellipse and receiver-station  $r_i$ , and  
 291  $d_{kj}$  be the distance between  $s_k$  and receiver-station  $r_j$ . Then we define the elliptical sta-  
 292 tionary phase zone for method  $^{ell}\mathcal{I}_3^{DW}$  as

$$d_{ki} + d_{kj} \leq (1 + \alpha)d_{ij}, \quad (2)$$

293 where  $\alpha \geq 0$  and we choose  $\alpha = 10^{-2}$ . Thus, if source-station  $s_k$  lies within the ellip-  
 294 tical stationary phase zone, the sum of distances from  $s_k$  to  $r_i$  and to  $r_j$  is less than 1%  
 295 longer than the distance between the receiver-stations.

296 Similarly, a hyperbola is defined as the locus of points where the difference of the  
 297 distances to the foci is constant. We therefore define the hyperbolic stationary phase zones  
 298 for method  $^{hyp}\mathcal{I}_3^{DW}$  as

$$|d_{ki} - d_{kj}| \geq (1 - \alpha)d_{ij}, \quad (3)$$

299 where  $\alpha \in [0, 1]$  and again we choose  $\alpha = 10^{-2}$ . This means that if source-station  $s_k$   
 300 lies within the hyperbolic stationary phase zone, the difference of distances from  $s_k$  to  
 301  $r_i$  and to  $r_j$  is greater than 99% of the distance between the receiver-stations. On a sphere,

302 the locus of points where the difference of the distances to the foci is constant, however,  
 303 approximates a hyperbola only near the foci.

304 The stationary phase zones can be defined alternatively using azimuthal angle  $\theta$   
 305 (**Fig. 3**) instead of  $\alpha$ . For the methods  $\mathcal{I}_3^{CW}$  and  $^{hyp}\mathcal{I}_3^{DW}$ ,  $\theta$  is the angle from the source-  
 306 station to the mid-point between the receiver-stations (**Fig. 3a,b**), which defines the slopes  
 307 of the asymptotes of a hyperbola. It is related to  $\alpha$  by  $\cos \theta = 1 - \alpha$ , where  $\theta \in [0, 2\pi]$ .  
 308 The definition of angle  $\theta$  for a given source-station for method  $^{ell}\mathcal{I}_3^{DW}$  is motivated by  
 309 the symmetry in eqs. (4) and (5) below. To do so, first identify the ellipse on which the  
 310 source-station lies with the two receiver-stations as foci. Then find the intersection point  
 311 between the ellipse and the perpendicular bisector of the line segment linking the two  
 312 receiver-stations. Angle  $\theta$  is the angle between a receiver-station and this intersection  
 313 point. **Fig. 3c** shows an example of this intersection point, but does not identify the lo-  
 314 cation of the source-station or the ellipse on which it lies. In this case,  $\theta$  is related to  $\alpha$   
 315 by  $\cos \theta = 1/(1+\alpha)$ , where  $\theta \in [0, \frac{\pi}{2}]$ . For the same  $\alpha$ ,  $\theta$  is generally larger for  $^{hyp}\mathcal{I}_3^{DW}$  than  
 316 for  $^{ell}\mathcal{I}_3^{DW}$ . Our choice of  $\alpha = 10^{-2}$  corresponds to a maximum  $\theta \approx 8^\circ$  for both  $^{hyp}\mathcal{I}_3^{DW}$  and  
 317  $^{ell}\mathcal{I}_3^{DW}$ .

318 We use eqs. (2) and (3) with  $\alpha = 10^{-2}$  to define the stationary phase zones in this  
 319 paper for methods  $^{ell}\mathcal{I}_3^{DW}$  and  $^{hyp}\mathcal{I}_3^{DW}$ , respectively. These definitions are chosen for sim-  
 320 plicity and because they appear to provide reliable results in the applications we con-  
 321 sider. However, the choice of the value of  $\alpha$  is ad-hoc as is its frequency-independence.  
 322 More elaborate, perhaps frequency-dependent, definitions may prove to be preferable.

323 The approximate arrival time,  $\delta t$ , for method  $^{hyp}\mathcal{I}_3^{DW}$  is known (Tsai, 2009; Yao  
 324 & van der Hilst, 2009):

$$\delta t = \frac{d_{ij}}{v} \cos \theta, \quad (4)$$

325 for a plane-wave in a medium with constant wave speed  $v$ , where  $d_{ij}$  is the inter-receiver-  
 326 station distance and  $\theta$  is shown in **Fig. 3b**. The grey line plotted in **Fig. 5b** is for this  
 327 formula. Analogously, the approximate arrival time  $t_{sum}$  for method  $^{ell}\mathcal{I}_3^{DW}$  is:

$$t_{sum} = \frac{d_{ij}}{v} \sec \theta, \quad (5)$$

328 for  $\theta$  shown in **Fig. 3c**. The grey line plotted in **Fig. 5c** is for this formula.

### 3.2 Stacking Weights

Appropriate stacking weights  $w_k$  must be applied for each source-station  $s_k$  to compute the composite Green's functions for each of the three-station methods ( $\mathcal{I}_3^{CW}$ ,  $hyp\mathcal{I}_3^{DW}$  and  $ell\mathcal{I}_3^{DW}$ ). The principal weight that we use is to set  $w_k$  equal to the reciprocal of the root-mean-square (rms) of the noise in the coda-wave part of each source-specific interferogram,  $C_3(r_i, r_j; s_k)$  for receiver-stations  $r_i$  and  $r_j$ . Defined in this way, we down-weight each contributing cross-correlogram by the rms of trailing noise. We do not, however, normalize the amplitude of the cross-correlograms. Therefore, down-weighting by the rms of trailing noise is approximately equivalent to normalizing the amplitudes of the cross-correlograms then weighting by peak signal-to-rms trailing noise ratio (SNR). Because the peak signal grows approximately linearly with the time series length of the records used to compute the cross-correlations, and rms trailing noise grows approximately as the square root of the time series length, SNR grows approximately as the square root of time series length (Snieder, 2004; Bensen et al., 2007). Thus, the use of this weighting scheme tends to accentuate the contribution from longer cross-correlations, but less strongly than if we had not normalized by peak amplitude and inversely by the rms of the trailing noise.

There are three other aspects of the data processing that can be considered to be stacking weights. First, for the direct-wave three-station methods, we only include a source-station in the stack if it lies within an appropriately defined stationary phase zone, which is referred to as geometrical-weighting in the Introduction. This choice can be thought of as applying binary weights to source-stations depending on their position relative to the receiver-stations. Second, also as mentioned above, unless the two constituent two-station interferograms,  $C_2(s_k, r_i)$  and  $C_2(s_k, r_j)$ , both have  $SNR \geq 10$ , the weight of the corresponding three-station interferogram,  $C_3(r_i, r_j; s_k)$ , is set to zero; otherwise it is unity. Third, to include signals for the longest paths ( $> 3000$  km) in the coda-wave three-station method, a source-station is excluded if the length of either  $C_2^{CW}(s_k, r_i)$  or  $C_2^{CW}(s_k, r_j)$  is less than 1500 s.

### 3.3 Estimating Composite Green's Functions

To compute the composite Green's function,  $\hat{G}(r_i, r_j)$ , for each of the three-station methods we apply the weighted sum given by eq. (1) based on the stacking weights (sec-

tion 3.2). **Fig. 6** provides some examples using the same pair of receiver-stations used in the record-sections of **Fig. 5**.

**Fig. 6a** presents an example composite Green’s function for three-station coda-wave interferometry ( $\mathcal{I}_3^{CW}$ ). For this method, no stationary phase zone is needed, so contributions from all source-stations are included in the stack. This is the black line in **Fig. 6a**, labelled “Stack all”, which is compared to the two-station ambient noise cross-correlation plotted as the red line and labelled  $\mathcal{I}_2^{AN}$ . Two observations of noteworthy: First, one of the features of coda-wave interferometry is the tendency for the composite Green’s functions to be more symmetric than for two-station ambient noise methods (e.g., [Stehly et al., 2008](#), and many others), and this is also observed in this example. We found it, however, to be an artifact due to the use of symmetric components ([Sheng et al., 2018](#)). Second, the SNR of the three-station coda-wave composite Green’s function is lower than for the two-station record, even though in this case 510 source-stations contribute to the three-station interferogram. This highlights another aspect of coda-wave interferometry, i.e., signals emerge from noise very slowly as source-stations are introduced. And, as can be seen in **Fig. 5a**, constituent source-specific three-station interferograms are typically very noisy so that signals cannot be discerned in any of them. This implies that the presence of many long duration source-stations may be necessary for coda-wave interferometry to play a useful role in ambient noise interferometry, unless more sophisticated data processing procedures are applied (**section 7.3**). For comparison, we also plot in **Fig. 6a** the recovered composite Green’s function based on source-stations that lie exclusively in the hyperbolic stationary phase zone. The choice of source-stations in this zone further degrades the SNR of the composite Green’s function, indicating that there is no geometrical advantage to choosing source-stations in the end-fire directions in coda-wave interferometry.

**Fig. 6b** shows an example composite Green’s function for three-station direct-wave interferometry where the source-stations lie in the hyperbolic stationary phase zone ( $^{hyp}\mathcal{I}_3^{DW}$ ). In this case, the green line, which is the stack for source-stations only in the hyperbolic stationary phase zones, is the Green’s function estimate, and there are 25 source-stations. Retaining source-stations at all azimuths (black line) degrades the result by adding precursory noise. Two comments are worthy of note in comparing the three-station composite Green’s function (green line) with two-station Green’s function (red line). First, the relative amplitudes for the different correlation lags are more similar than for coda-

393 waves. Second, precursory noise is lower for the three-station estimate. These are both  
 394 common characteristics when comparing two-station to three-station Green’s functions.

395 Finally, **Fig. 6c** presents an example composite Green’s function for three-station  
 396 direct-wave interferometry where the source-stations lie in the elliptical stationary phase  
 397 zone ( $^{ell}\mathcal{I}_3^{DW}$ ). The green line, which is the stack for source-stations only in the ellip-  
 398 tical stationary phase zones, is the composite Green’s function estimate, and there are  
 399 7 source-stations. As with the hyperbolic stationary phase zone, retaining source-stations  
 400 at all azimuths (black line) degrades the result but in this case adds both precursory and  
 401 trailing noise, especially the trailing noise. In this case, too, there is lower precursory noise  
 402 for the three-station estimate than for  $\mathcal{I}_2^{AN}$ .

## 403 4 Dispersion Measurements

### 404 4.1 Frequency-Time Analysis

405 To measure frequency dependent phase speed, we apply frequency-time analysis  
 406 (FTAN; [Dziewonski et al., 1969](#); [Levshin & Ritzwoller, 2001](#); [Bensen et al., 2007](#)). We  
 407 assume that the measured phase of a two-station interferogram ( $\mathcal{I}_2^{AN}$ ) at frequency  $\omega$   
 408 in the frequency domain for receiver-stations  $r_i$  and  $r_j$  is approximately ([Lin et al., 2008](#))

$$409 \phi_{ij}^{AN}(\omega) = \frac{\omega}{c_{ij}} d_{ij} + \frac{\pi}{4} + \phi_s + 2N\pi, \quad N \in \mathbb{Z}, \quad (6)$$

410 where  $d_{ij}$  is the distance between the two receiver-stations,  $\pi/4$  is from the far-field or  
 411 high-frequency asymptotic approximation of the Bessel function,  $\phi_s$  is an initial phase  
 412 term, and  $c_{ij}$  is the frequency-dependent phase speed, which is what we aim to measure.

413 For two-station ambient noise interferometry ( $\mathcal{I}_2$ ),  $\phi_s \approx 0$  has been shown the-  
 414 oretically ([Snieder, 2004](#)) and empirically ([Yao et al., 2006](#); [Lin et al., 2008](#)). For three-  
 415 station coda-wave interferometry ( $\mathcal{I}_3^{CW}$ ),  $\phi_s$  should also be approximately 0. However,  
 416 for three-station direct-wave interferometry ( $\mathcal{I}_3^{DW}$ ),  $\phi_s$  will differ from 0, and this ini-  
 417 tial phase must be taken into account when measuring phase speed.

418 For  $^{hyp}\mathcal{I}_3^{DW}$ , let the source-station  $s_k$  lie outside the two receiver-stations at dis-  
 419 tances  $d_{ki}$  from  $r_i$  and  $d_{kj}$  from  $r_j$  (**Fig. 7b**). Because correlation of two interferograms  
 420 will determine the difference of the phases in the frequency domain, the phase of  $^{hyp}C_3(r_i, r_j; s_k)$

421 is

$${}^{hyp}\phi_{ij;k} = \phi_{ki}^{AN} - \phi_{kj}^{AN} = \omega \left( \frac{d_{ki}}{c_{ki}} - \frac{d_{kj}}{c_{kj}} \right) + 2N\pi. \quad (7)$$

422 From straight-ray and far-field assumptions (Tsai, 2009), we have

$$\frac{d_{ki}}{c_{ki}} - \frac{d_{kj}}{c_{kj}} \approx \frac{d_{ki} - d_{kj}}{c_{ij}}, \quad (8)$$

423 thus

$${}^{hyp}\phi_{ij;k} = \frac{\omega}{c_{ij}} (d_{ij} + {}^{hyp}\delta d_{ij;k}) + 2N\pi, \quad (9)$$

424 where

$${}^{hyp}\delta d_{ij;k} = d_{ki} - d_{kj} - d_{ij}. \quad (10)$$

425 For  ${}^{ell}\mathcal{I}_3^{DW}$ , source-stations lie generally between the two receiver-stations (**Fig.**  
426 **7a**). Because convolution of two interferograms will determine the sum of the phases in  
427 the frequency domain, the phase of  ${}^{ell}C_3(r_i, r_j; s_k)$  is

$${}^{ell}\phi_{ij;k} = \phi_{ki}^{AN} + \phi_{kj}^{AN} = \omega \left( \frac{d_{ki}}{c_{ki}} + \frac{d_{kj}}{c_{kj}} \right) + \frac{\pi}{2} + 2N\pi. \quad (11)$$

428 Based on approximations similar to  ${}^{hyp}\mathcal{I}_3^{DW}$ , we find

$${}^{ell}\phi_{ij;k} = \frac{\omega}{c_{ij}} (d_{ij} + {}^{ell}\delta d_{ij;k}) + \frac{\pi}{2} + 2N\pi, \quad (12)$$

429 where

$${}^{ell}\delta d_{ij;k} = d_{ki} + d_{kj} - d_{ij}. \quad (13)$$

430 Assuming  ${}^{hyp}\delta d = 0$  gives  ${}^{hyp}\phi_s = -\pi/4$  by comparing eqs. (6) and (9). Simi-  
431 larly, assuming  ${}^{ell}\delta d = 0$  yields  ${}^{ell}\phi_s = \pi/4$  by comparing eqs. (6) and (12). The as-  
432 sumption that  $\delta d = 0$  will lead to biased measurements for the three-station direct-wave  
433 methods and its correction is discussed in **section 5**.

434 **Fig. 8** compares example frequency-time (FTAN) diagrams for the two-station method  
435 and the three-station methods, for the two receiver-stations M07A and M15A. The four  
436 diagrams are similar at short periods but the diagrams for the two direct-wave meth-  
437 ods show larger relative amplitudes at longer periods. For the coda-wave diagram, longer  
438 periods are too noisy to measure and the 26 s stripe correspond to a spatially localized  
439 microseism source. The effects of the 26 s microseism are discussed in **section 7.2**.

440 We apply two additional quality control criteria to the dispersion measurements.  
441 First, for a dispersion measurement to be retained, we apply a spectral SNR (Bensen et

442 [al., 2007](#)) criterion to the composite Green’s function, where again SNR is defined as the  
 443 peak amplitude in the direct-wave window divided by the rms of the waveform in the  
 444 coda-wave window. That is, at a given period the composite Green’s function must have  
 445 a  $\text{SNR} \geq 10$  otherwise the dispersion measurement at that period is discarded. Second,  
 446 the distance between the two receiver-stations must be greater than three wavelengths  
 447 ([Lin et al., 2008](#)) for the dispersion measurement to be retained. For example, if the phase  
 448 speed is 4 km/s, at 20 s period the receiver-stations must be separated by more than 240  
 449 km. This criterion becomes more restrictive as period increases.

## 450 4.2 General Characteristics

451 **Fig. 9a** summarizes the spectral signal-to-noise ratio (SNR) of each of the four in-  
 452 terferometric methods, averaging over the entire data set of dispersion measurements.  
 453 Generally speaking, SNR decreases with period and the trends are similar between  $\mathcal{I}_3$   
 454 and  $\mathcal{I}_2^{AN}$ . The peaks near 16 s and 8 s periods correspond to the primary and secondary  
 455 microseisms, respectively, while the dip near 26 s period corresponds to the existence of  
 456 a spatially localized microseismic source (e.g., [Shapiro et al., 2006](#); [Xia et al., 2013](#)). **Fig.**  
 457 **9b** presents the SNR results relative to the SNR for  $\mathcal{I}_2^{AN}$ . The SNR for  $^{ell}\mathcal{I}_3^{DW}$  is slightly  
 458 larger than for  $^{hyp}\mathcal{I}_3^{DW}$ , while both have a SNR more than twice that of  $\mathcal{I}_2^{AN}$  across a  
 459 broad bandwidth. In contrast,  $\mathcal{I}_3^{CW}$  has a much lower median SNR ( $< 10$ ) across all pe-  
 460 riods.

461 Because SNR plays a significant role in the quality control of dispersion measure-  
 462 ments, the number of accepted  $\mathcal{I}_2$  and  $\mathcal{I}_3$  measurements varies with period similar to SNR  
 463 (**Fig. 10a**). The number of accepted  $\mathcal{I}_3$  measurements can be divided into three cat-  
 464 egories depending on whether the two receiver-stations operated at the same time (syn-  
 465 chronously) and whether an  $\mathcal{I}_2$  measurement exists for the path so that the  $\mathcal{I}_3$  measure-  
 466 ment is new or repeated. These three categories of  $\mathcal{I}_3$  measurements are referred to as  
 467 “Synchronous-Repeated” (receiver-stations deployed synchronously, with both an  $\mathcal{I}_3$  and  
 468 an  $\mathcal{I}_2$  measurement), “Synchronous-New” (receiver-stations deployed synchronously, with  
 469 an  $\mathcal{I}_3$  but not an  $\mathcal{I}_2$  measurement), and “Asynchronous-New” (receiver-stations deployed  
 470 asynchronously, with only an  $\mathcal{I}_3$  measurement). In the Synchronous-New case, the receiver-  
 471 stations produced an  $\mathcal{I}_2$  measurement but it was rejected, usually because it did not meet  
 472 the SNR requirement. The numbers of  $\mathcal{I}_3$  measurements that derive from these three  
 473 categories are shown in **Fig. 10b-d**. In all categories,  $^{ell}\mathcal{I}_3^{DW}$  measurements somewhat

474 outnumber the  $hyp\mathcal{I}_3^{DW}$  measurements, and both outnumber the  $\mathcal{I}_2^{AN}$  measurements (in  
 475 cases where they exist) and greatly outnumber the  $\mathcal{I}_3^{CW}$  measurements.

476 **Fig. 10b** is for the Synchronous-Repeated category of  $\mathcal{I}_3$  measurements. By def-  
 477 inition, the number of  $\mathcal{I}_3$  measurements will be no larger than the number of  $\mathcal{I}_2$  mea-  
 478 surements. Nearly every existing  $\mathcal{I}_2^{AN}$  measurement is accompanied by an  $\mathcal{I}_3^{DW}$  measurement,  
 479 but the number of  $\mathcal{I}_3^{CW}$  measurements is considerably smaller. The number of these mea-  
 480 surements generally decreases with period after maximizing between 20 and 30 s, although  
 481 the  $\mathcal{I}_3^{CW}$  measurement maximizes nearer to 15 s period and decays very rapidly at longer  
 482 periods.

483 **Fig. 10c** is for the Synchronous-New category of  $\mathcal{I}_3$  measurements, and illustrates  
 484 that many new longer periods measurements emerge from the  $\mathcal{I}_3^{DW}$  method. Above about  
 485 50 s period,  $\mathcal{I}_3^{DW}$  nearly doubles the number of measurements between synchronously de-  
 486 ployed stations. Although a principal attraction of the three-station methods is the abil-  
 487 ity to obtain measurements between asynchronously deployed stations, but many new  
 488 measurements result from the  $\mathcal{I}_3^{DW}$  methods even for synchronously deployed stations  
 489 particularly at long periods. There are essentially no new measurements from  $\mathcal{I}_3^{CW}$  in  
 490 this category.

491 **Fig. 10d** is for the Asynchronous-New category of  $\mathcal{I}_3$  measurements, measurements  
 492 from the  $\mathcal{I}_3$  methods that are inherently non-existent for  $\mathcal{I}_2$ . Relative to the number of  
 493 measurements delivered by  $\mathcal{I}_2^{AN}$ , the greatest impact of the  $\mathcal{I}_3$  methods is at the longer  
 494 periods of the bandwidth considered. The vast majority of the measurements for  $\mathcal{I}_3^{CW}$  are  
 495 from synchronously deployed stations (**Fig. 10b**), indicating that it is difficult for  $\mathcal{I}_3^{CW}$  to  
 496 bridge asynchronous stations.

## 497 5 Correcting the Bias in Three-Station Direct-Wave Interferometry ( $\mathcal{I}_3^{DW}$ )

498 As described above, the three-station methods are based on measuring the phase  
 499 speed of the composite Green's function (eq. (1)),  $\hat{G}_3(r_i, r_j)$ , between a pair of receiver-  
 500 stations  $(r_i, r_j)$ , which is a stack of source-specific interferograms,  $C_3(r_i, r_j; s_k)$ , that emerge  
 501 from particular source-stations  $s_k$ . In direct-wave methods  $hyp\mathcal{I}_3^{DW}$  and  $ell\mathcal{I}_3^{DW}$ , the phase  
 502 speed,  $c_{ij}$ , is measured using the composite Green's function based on eqs. (9) and (12),  
 503 respectively, under the assumption that  $\delta d = 0$ . It is this assumption for the compos-  
 504 ite Green's function that can produce the systematic bias in the three-station direct-wave

505 methods. Because  $^{ell}\delta d$  is always positive, assuming  $^{ell}\delta d = 0$  will result in a phase speed  
 506 that is biased slow for  $^{ell}\mathcal{I}_3^{DW}$ . In contrast, because  $^{hyp}\delta d$  is always negative, assuming  
 507  $^{hyp}\delta d = 0$  will result in a phase speed that is biased fast for  $^{hyp}\mathcal{I}_3^{DW}$ .

508 Therefore, the correct distance to be used in measuring phase speed will depend  
 509 on the specific location of each source-station. The direct use of the composite Green’s  
 510 function invariably will yield a biased phase speed measurement. To “de-bias” the phase  
 511 speed measurements, we abandon the composite Green’s function and measure a phase  
 512 speed curve for each source-specific interferogram ( $C_3(r_i, r_j; s_k)$ ) independently based  
 513 on corrections from the more accurate ray-theoretic distance,  $^{ell}\delta d_{ij;k}$  or  $^{hyp}\delta d_{ij;k}$ , and  
 514 then average the resulting phase speed curves.

515 **Fig. 11** presents an example of the set of source-specific phase speed curves that  
 516 have been de-biased by using the source-specific ray-theoretic distances. At each period  
 517 we reject a source-specific measurement if its SNR  $< 10$  or either of the source-receiver  
 518 distances is  $< 2\lambda$ . We do not, however, apply the wavelength criterion to the two source-  
 519 receiver distances in constructing the composite Green’s function before the de-biasing  
 520 correction (**section 3.3**) because that would require stacking over different source-stations  
 521 at different periods. Then we reject the 10% of measurements most different from the  
 522 mean. Finally, we calculate the standard deviation ( $\sigma$ ) and discard the mean measure-  
 523 ment altogether if  $\sigma > 60$  m/s.

524 **Fig. 12a** shows the correction averaged over the entire data set for the two three-  
 525 station direct-wave methods. Our definition of stationary phase zones that  $\alpha = 1\%$  (eqs.  
 526 (2) and (3)) provides an upper limit on the bias as 1%. The absolute mean correction  
 527 is about 10 m/s at all periods for both methods, which is around 0.3%, and thus con-  
 528 sistent with the definition of stationary phase zones. The average standard deviation amongst  
 529 the constituent source-specific curves over the entire data set is presented in **Fig. 12b**.  
 530 The standard deviations for the  $^{ell}\mathcal{I}_3^{DW}$  method are generally smaller for the  $^{hyp}\mathcal{I}_3^{DW}$  method,  
 531 consistent with the latter having larger and more complex sensitivity zones (**section 7.1**).  
 532 These standard deviations may serve in the future as uncertainty estimates for the re-  
 533 sulting dispersion measurements.

## 6 Validate Three-Station ( $\mathcal{I}_3$ ) against Two-Station ( $\mathcal{I}_2$ ) Interferometry

To test if three-station methods are consistent with two-station interferometry, and if the de-biasing correction for  $\mathcal{I}_3^{DW}$  presented in **section 5** is effective, we statistically compare the differences in Rayleigh wave phase speed measurements and also the associated phase speed maps from the methods.

### 6.1 Phase Speed Measurements

**Fig. 13** and **Table 1** present comparisons of Rayleigh wave phase speed measurements derived from the three-station methods to two-station interferometry for common receiver-station pairs.

**Fig. 13c** and **Table 1** (column 2) show that the mean difference between the two-station Green's functions and the three-station composite Green's functions based on coda-waves is negligible ( $< 2$  m/s, on average), from which we infer that the three-station method based on coda-waves is unbiased. The standard deviation of the difference decreases with period to achieve a minimum around 15 s, but then increases rapidly with period although results extend only up to 30 s period.

In contrast, **Fig. 13a,b** and **Table 1** (columns 4 & 6) show the existence of a non-zero systematic difference or bias between each of the three-station direct-wave methods with two station interferometry before correction. For  $^{ell}\mathcal{I}_3^{DW}$ , the bias is always negative and the absolute bias increases with period. For  $^{hyp}\mathcal{I}_3^{DW}$ , the bias is positive and the absolute bias generally decreases with period.

After the de-biasing correction, the mean and standard deviation of the difference between the  $\mathcal{I}_3^{DW}$  and  $\mathcal{I}_2^{AN}$  measurements are shown in **Fig. 13d,e** and **Table 2**, which should be contrasted with **Fig. 13a,b** and **Table 1** that contains the same statistics without the de-biasing. The correction decreases the absolute mean difference between the  $\mathcal{I}_3^{DW}$  and  $\mathcal{I}_2^{AN}$  measurements at most periods. If we consider the mean difference to be a measure of residual bias, then the bias of the corrected measurements is relatively small ( $< 5$  m/s) for both  $\mathcal{I}_3^{DW}$  methods at periods  $< 40$  s. However, the residual bias generally increases at longer periods for both  $\mathcal{I}_3^{DW}$  methods. Potential causes of and corrections for the residual bias are discussed in **section 7.1**.

563 In contrast with the bias, generally the standard deviations of the differences be-  
 564 tween the dispersion measurements from the  $\mathcal{I}_2^{AN}$  method to both  $\mathcal{I}_3^{DW}$  methods grow  
 565 with period (**Fig. 13a,b** and **Table 1 (columns 5 & 7)**). Partly, this is due to the  
 566 decrease in signal-to-noise ratio (SNR) in both the three-station and two-station inter-  
 567 ferograms at longer periods (**Fig. 10a**). However, irrespective of SNR, we do not ex-  
 568 pect the dispersion measurements from the three-station methods to agree with those  
 569 from the two-station method as well at longer periods. The reason is that the Fresnel  
 570 Zone or sensitivity kernel for the three-station methods is not identical to the sensitiv-  
 571 ity kernel for the two-station method and the differences in sensitivity grow with period  
 572 (**section 7.1**).

## 573 6.2 Eikonal Tomography

574 To further validate and compare the three-station methods we report results from  
 575 surface wave tomography based on them. To perform tomography, we apply the eikonal  
 576 tomography method ([Lin et al., 2009](#)) to Rayleigh wave phase speed measurements ob-  
 577 tained from the two-station and three-station methods. We employ the eikonal tomog-  
 578 raphy method rather than traditional tomographic methods that minimize a penalty func-  
 579 tional (e.g., [Barmin et al., 2001](#)) because eikonal tomography applies no ad-hoc regular-  
 580 ization that depends on data coverage. This simplifies comparison of results from dif-  
 581 ferent datasets because they are less affected by differences in the number and distribu-  
 582 tion of wave paths. In this section, we consider  $\mathcal{I}_3^{DW}$  only after the de-biasing correction.

583 The Rayleigh wave phase speed maps produced by the three-station ( $\mathcal{I}_3$ ) and two-  
 584 station ( $\mathcal{I}_2^{AN}$ ) methods are generally quite similar, as displayed at periods of 10 s, 20 s,  
 585 40 s, and 60 s in **Figs 14 - 17**. The touchstone is the  $\mathcal{I}_2^{AN}$  map, and at each period there  
 586 is substantial agreement between the  $\mathcal{I}_3$  maps with the  $\mathcal{I}_2^{AN}$  map. However, we do not  
 587 show the three-station coda-wave ( $\mathcal{I}_3^{CW}$ ) maps at periods of 40 s and 60 s because the  
 588  $\mathcal{I}_3^{CW}$  method does not provide enough measurements to perform tomography reliably at  
 589 periods above 30 s. Presumably, this is because the coda is enriched at the shorter pe-  
 590 riods ([Spica et al., 2016](#); [Ansaripour et al., 2019](#)).

591 A more careful comparison of the tomographic maps requires detailed inspection  
 592 of the differences between the maps. Let us assume that we have two dispersion maps  
 593 on the same grid of longitudes ( $x_i$ ) and latitudes ( $y_j$ ):  $c_{ij}^{(1)} = c^{(1)}(x_i, y_j)$  and  $c_{ij}^{(2)} =$

594  $c^{(2)}(x_i, y_j)$ . Let  $\Delta_{ij}$  be the difference between these maps:

$$\Delta_{ij} = c_{ij}^{(1)} - c_{ij}^{(2)}, \quad (14)$$

595 whose mean over  $(x_i, y_j)$  is denoted as  $\bar{\Delta}$  and standard deviation as  $\sigma_{\Delta}$ . **Figs 18 - 20**  
 596 display such differences between the three-station methods with two-station interferom-  
 597 etry in map form and **Table 3** summarizes the differences, tabulating  $\bar{\Delta}$  and  $\sigma_{\Delta}$ .

598 **Fig. 18** (and **Table 3**, column 2) shows the difference between the Rayleigh wave  
 599 phase speed maps at periods of 10 s and 20 s from three-station coda-wave interferom-  
 600 etry ( $\mathcal{I}_3^{CW}$ ) and two-station interferometry ( $\mathcal{I}_2^{AN}$ ). There is a small systematic differ-  
 601 ence between the maps ( $\bar{\Delta} \approx 7$  m/s) and the standard deviation of the differences is  
 602 also small ( $\sigma_{\Delta} < 15$  m/s). Unfortunately, we are unable to produce meaningful tomo-  
 603 graphic maps from  $\mathcal{I}_3^{CW}$  at longer periods, while it may be more feasible to push  $\mathcal{I}_3^{CW}$  towards  
 604 shorter periods than what can be produced by  $\mathcal{I}_2^{AN}$  (**section 7.3**).

605 **Fig. 19** presents difference maps at periods from 10 s to 60 s for the three-station  
 606 direct-wave method  $^{ell}\mathcal{I}_3^{DW}$  relative to  $\mathcal{I}_2^{AN}$ . **Table 3**, columns 4-5, summarizes the mean  
 607 and standard deviation of the difference over the maps. The standard deviation of the  
 608 differences generally grow with period because the  $\mathcal{I}_3^{DW}$  methods increasingly sample the  
 609 earth differently than the ( $\mathcal{I}_2^{AN}$ ) method at longer periods (**section 7.1**). Larger dis-  
 610 crepancies are observed near the peripheries of the maps, where both methods have larger  
 611 uncertainties. However, the maps are reasonably consistent ( $\sigma_{\Delta} < 25$  m/s) across all  
 612 periods.

613 **Fig. 20** presents difference maps at periods from 10 s to 60 s for the three-station  
 614 direct-wave method  $^{hyp}\mathcal{I}_3^{DW}$  relative to  $\mathcal{I}_2^{AN}$  whose mean and standard deviation are sum-  
 615 marized in **Table 3**, columns 6-7. Similar patterns are observed as in  $^{ell}\mathcal{I}_3^{DW}$ .

## 616 7 Discussion

### 617 7.1 Residual Bias of Three-Station Interferometry

618 Our de-biasing correction methods are based on straight-ray theory. As shown in  
 619 **section 6**, some residual bias exists between three-station direct-wave interferometry  
 620 ( $\mathcal{I}_3^{DW}$ ) and two-station interferometry ( $\mathcal{I}_2^{AN}$ ) even after the correction, especially at the  
 621 longer periods. We believe this is due to deviation from ray theory. In particular, we dis-

622 cuss here the finite frequency effects and the differences in the Fresnel Zones or sensi-  
 623 tivity kernels between the methods.

624 **Fig. 21** schematically depicts the difference in sensitivity for the three-station direct-  
 625 wave measurements and the two-station measurement, in which we approximate the Fres-  
 626 nel Zone for the two-station method as an ellipse, shown with dashed lines, with the two  
 627 receiver-stations at the ellipse’s foci. The Fresnel Zone for the method  $^{ell}\mathcal{I}_3^{DW}$  is approx-  
 628 imately the sum of the two Fresnel zones for each of the constituent waves that emanate  
 629 from the source-station (red dot in **Fig. 21a**) which lies between the receiver-stations  
 630 for this method. The sensitivity zone for  $^{ell}\mathcal{I}_3^{DW}$  is smaller than for  $\mathcal{I}_2$ , on average, and  
 631 we therefore expect that the method  $^{ell}\mathcal{I}_3^{DW}$  will have a higher resolution than  $\mathcal{I}_2$ , ev-  
 632 erything else being equal. In contrast, the Fresnel Zone for the method  $^{hyp}\mathcal{I}_3^{DW}$  is ap-  
 633 proximately the difference of the two Fresnel zones for each of the constituent waves that  
 634 emanate from the source-stations (red dots in **Fig. 21b**), which lie outside the receiver-  
 635 stations. This sensitivity zone for  $^{hyp}\mathcal{I}_3^{DW}$  is larger and considerably more complicated  
 636 than for  $\mathcal{I}_2$ , on average. We, therefore, expect that the method  $^{hyp}\mathcal{I}_3^{DW}$  will have a lower  
 637 resolution than  $\mathcal{I}_2$ , everything else being equal.

638 The Fresnel zones for the  $\mathcal{I}_2^{AN}$  method widen with period, as will those for the  $\mathcal{I}_3^{DW}$  methods.  
 639 Therefore, differences between the Fresnel zones of the  $\mathcal{I}_3^{DW}$  methods compared with the  
 640 Fresnel zone of the  $\mathcal{I}_2^{AN}$  method will increase with period, too, as the various methods  
 641 sample the earth between and around the pair of receiver-stations increasingly differently.  
 642 We believe this is the source of the increase in the standard deviations of the differences  
 643 between the phase speed measurements and maps for the various methods (e.g., **Fig. 13**).

644 The analysis of Fresnel Zones presented here is schematic and illustrative. The Fres-  
 645 nel Zones have internal structure that will produce details in the sums and differences  
 646 presented in **Fig. 21**. General conclusions about the nature of the differences between  
 647 the various Fresnel Zones are robust, but to use this information quantitatively to im-  
 648 prove images in the future will require much more careful computation of the Fresnel zones  
 649 (e.g., de Vos et al., 2013).

## 650 7.2 Effects of the 26 s Microseism

651 A noteworthy observation is that few reliable measurements exist for three-station  
 652 coda-wave interferometry ( $\mathcal{I}_3^{CW}$ ) beyond 40 s (**Fig. 10a**). The degradation of quality

653 with period for  $\mathcal{I}_3^{CW}$  is also observed in [Spica et al. \(2016\)](#) and [Ansaripour et al. \(2019\)](#).  
 654 To understand its cause, we compare the spectra for all methods (**Fig. 22**). Specifically,  
 655 we randomly choose 10,000 interferograms from each method and calculate their ampli-  
 656 tude spectra. Then the amplitude spectra are normalized and stacked to form mean am-  
 657 plitude spectra curves with standard deviations for each method.

658 As shown in **Fig. 22**, the 26 s spatially localized microseism source (e.g., [Shapiro](#)  
 659 [et al., 2006](#); [Xia et al., 2013](#)) leaves an imprint on the spectra for all methods although  
 660 is somewhat stronger for  $\mathcal{I}_3^{CW}$ , which is also indicated in the example FTAN diagrams  
 661 (**Fig. 8**). We also compare spectra of  $\mathcal{I}_3^{CW}$  with and without spectral whitening (**Fig.**  
 662 **22de**). The whitening makes the spectra flatter at short periods but does not substan-  
 663 tially remove the effects of the 26 s microseism.

664 The spectra of  $\mathcal{I}_3^{CW}$  also show much stronger variability at long periods than other  
 665 methods. Spectral whitening does not help reduce its variability. Thus, although the 26 s  
 666 microseism has a stronger effect on  $\mathcal{I}_3^{CW}$  than other methods, we believe the lack of sig-  
 667 nals at long periods for  $\mathcal{I}_3^{CW}$  is largely due to the nature of the coda in two-station in-  
 668 terferometry ( $\mathcal{I}_2^{AN}$ ).

### 669 7.3 Potential for Further Refinement

670 We have chosen many of the characteristics of the two-station and three-station  
 671 interferometric methods in a reasoned but largely ad-hoc way. Thus, all of the proce-  
 672 dures we describe above may be refined to improve some aspect of the results. Such re-  
 673 finements could be made (1) to the data processing procedures, (2) to the definition of  
 674 the stationary phase zones for the direct-wave methods, (3) to the de-biasing procedure  
 675 applied to the direct-wave methods, and (4) to the use of the results from the different  
 676 methods in concert with one another.

677 (1) Data processing procedures include the definition of both the direct-wave and  
 678 coda-wave windows, the wavelength criterion for the minimum inter-receiver-station dis-  
 679 tance, the chosen values of the stacking weights  $w_j$ , and the use of only the symmetric  
 680 component of the two-station ambient noise interferograms as the basis for all of data  
 681 processing. In addition, the two-station data processing procedures of [Shen and Ritz-](#)  
 682 [woller \(2016\)](#) underlie our results, including the use of an 80 s moving average time-domain  
 683 normalization window and spectral whitening. All of these choices may be revised in the

684 future to optimize the result of data processing. For three-station coda-wave interfer-  
 685 ometry ( $\mathcal{I}_3^{CW}$ ), performing interferometry on hourly or daily  $\mathcal{I}_2$  and then stack hourly  
 686 or daily  $\mathcal{I}_3^{CW}$  (Zhang & Yang, 2013; Haendel et al., 2016) may greatly increase SNR be-  
 687 cause cross-talk between incoherent asynchronous signals are avoided (Sheng et al., 2018).  
 688 Despite its inapplicability to asynchronous pairs, this pre-stacking scheme may be promis-  
 689 ing for extraction of short-period information (Sheng et al., 2018).

690 (2) Another important characteristic of the three-station direct-wave methods is  
 691 the definition of the stationary-phase zones. We choose  $\alpha = 10^{-2}$  in eqs. (2) and (3)  
 692 to be period-independent, which produces a maximum angle of both the elliptical and  
 693 hyperbolic stationary phase zones of about  $\theta = 8^\circ$ . An optimal period-dependent pa-  
 694 rameterization of the stationary phase zones may be possible. Moreover, because increas-  
 695 ing  $\alpha$  should increase the bias of the three-station methods, in station-rich settings  $\alpha$  may  
 696 be reduced and in station-poor regions it may be increased, although at the expense of  
 697 increasing bias.

698 (3) The de-biasing method outlined in **section 5** applies corrections to dispersion  
 699 curves before they are statistically summarized for each path, based on a great-circle ray-  
 700 theoretic procedure. This method could be improved by correcting additional errors from  
 701 off-great-circle propagation (Yao et al., 2006; Foster et al., 2014), non-plane waves (Pedersen,  
 702 2006) and finite-frequency effects (Yao et al., 2010; de Vos et al., 2013). Alternately, a  
 703 completely different approach may be possible, which applies phase corrections to source-  
 704 specific interferograms ( $C_3^{DW}$ ) and then makes the dispersion measurements on the com-  
 705 posite Green’s function ( $\hat{G}_3$ ). The correction is a (frequency-dependent) phase shift to  
 706 each of the source-specific interferograms prior to stacking.

707 (4) Because the three-station methods ( $\mathcal{I}_3$ ) are consistent with the two-station method  
 708 ( $\mathcal{I}_2$ ), measurements from all methods can be combined simultaneously. It might be par-  
 709 ticularly advantageous to combine measurements from methods  $^{ell}\mathcal{I}_3^{DW}$  and  $^{hyp}\mathcal{I}_3^{DW}$  because  
 710 they are oppositely biased, and their biases may cancel approximately without an ex-  
 711 plicit bias correction.

#### 712 **7.4 Connections to Other Methods**

713 We now discuss how three-station interferometry ( $\mathcal{I}_3$ ) methods connect to and dif-  
 714 fer from other interferometric methods, and how the methods can gain insight from each

715 other. In particular, we discuss the earthquake two-station method (Sato, 1955), source-  
 716 receiver interferometry (Curtis & Halliday, 2010), and generalized interferometry not based  
 717 on Green’s function retrieval (Fichtner et al., 2017).

718 In the earthquake two-station method, for an earthquake lying approximately in-  
 719 line with two receivers, seismograms recorded at two receivers are correlated to extract  
 720 information about the inter-receiver medium (e.g., Landisman et al., 1969). Thus, if the  
 721 earthquake is replaced by a station and the earthquake seismograms are replaced by inter-  
 722 station noise correlations, then the configuration of the earthquake two-station method  
 723 is somewhat similar to three-station direct-wave interferometry method  $^{hyp}\mathcal{I}_3^{DW}$ . In this  
 724 study, source-receiver distances are similar in scale to inter-receiver distances while global  
 725 earthquakes are often used for regional studies in the earthquake two-station method (e.g.,  
 726 Yao et al., 2006). Thus, global source-stations may also be used in  $\mathcal{I}_3^{DW}$  which may pro-  
 727 vide longer period information in the future.

728 Source-receiver interferometry (SRI; Curtis & Halliday, 2010) presents three types  
 729 of geometries where one can extract the Green’s function between a source and a receiver  
 730 without direct observation: correlation-correlation SRI, correlation-convolution SRI, and  
 731 convolution-convolution SRI. The geometries of three-station direct-wave interferome-  
 732 try methods  $^{hyp}\mathcal{I}_3^{DW}$  and  $^{ell}\mathcal{I}_3^{DW}$  are similar to the correlation-correlation SRI and correlation-  
 733 convolution SRI, respectively, with a station serving as a virtual source. A critical dif-  
 734 ference between three-station interferometry and SRI is that our goal is to obtain reli-  
 735 able dispersion measurements from the direct Rayleigh waves, which requires us to re-  
 736 solve the source phase  $\phi_s$  (section 4.1). The tapering of stationary phase zones and the  
 737 area weights of source-stations in Entwistle et al. (2015) will affect  $\phi_s$  in a complicated  
 738 way, so that the tapering and the area weights are not used here. When amplitude in-  
 739 formation can be reliably interpreted from two-station interferograms ( $\mathcal{I}_2$ ), the taper-  
 740 ing and the area weights may provide a closer approximation to the theoretical integral  
 741 of Green’s function retrieval and thus might also benefit phase measurements.

742 Finally, the three-station direct-wave methods ( $\mathcal{I}_3^{DW}$ ) could work optimally for new  
 743 generalized interferometric methods not based on estimating Green’s functions. New meth-  
 744 ods of interferometry are being developed that attempt to extract information about the  
 745 sources and propagating medium jointly irrespective of the relative position of the sources  
 746 and receiver-stations (e.g., Tromp et al., 2010; Hanasoge, 2014; Fichtner et al., 2017; Sager

747 *et al.*, 2018).  $\mathcal{I}_3^{DW}$ , where the location of source-stations is known exactly, may provide  
 748 an ideal application for these methods.

## 749 8 Conclusions

750 Our principal finding is that the three-station direct-wave interferometry methods  
 751  $ell\mathcal{I}_3^{DW}$  and  $hyp\mathcal{I}_3^{DW}$  generally outperform three-station coda-wave interferometry  $\mathcal{I}_3^{CW}$  for  
 752 obtaining Rayleigh wave dispersion measurements, even though direct-wave interferom-  
 753 etry has been largely ignored as an imaging tool to date. This outperformance includes  
 754 such metrics as signal-to-noise ratio, the number of measurements returned, and most  
 755 notably the band-width of the measurements because  $\mathcal{I}_3^{CW}$  is primarily confined to pro-  
 756 viding measurements below 25 s period. In addition, the direct-wave methods also out-  
 757 perform two-station interferometry in these metrics.

758 There are two primary caveats concerning the performance of the three-station direct-  
 759 wave methods. First, the  $ell\mathcal{I}_3^{DW}$  and  $hyp\mathcal{I}_3^{DW}$  methods are slightly biased relative to two-  
 760 station interferometry,  $\mathcal{I}_2^{AN}$ . However, we present a ray-theoretic de-biasing procedure  
 761 that nearly eliminates the bias at and below about 40 s period, where ray-theory is ex-  
 762 pected to work best, and substantially reduces bias at longer periods. Second, the sen-  
 763 sitivity kernels for the three-station direct-wave methods are more complicated than both  
 764 two-station interferometry and three-station coda-wave interferometry and remain poorly  
 765 understood. Research is needed to understand the nature of the sensitivity kernels for  
 766 the three-station direct-wave methods and how they compare to two-station interferom-  
 767 etry.

768 The tests presented here use data from the EarthScope Transportable Array (TA),  
 769 but the relative merits of the various methods tested may vary in different settings where  
 770 station coverage and geometries will differ. Indeed, the three-station methods that we  
 771 test here may be least needed in the contiguous US due to the outstanding data cover-  
 772 age provided by the TA. Tests in different regions (e.g., Antarctica, Tibet, Europe, Alaska,  
 773 the Juan de Fuca Plate, etc.) are needed to determine how the methods will perform in  
 774 a variety of settings.

775 Irrespective of these caveats, we believe that three-station direct-wave interferom-  
 776 etry promises to provide a substantial new tool to the toolbox of standard methods for  
 777 imaging the structure of the crust and uppermost mantle. We encourage seismologists

778 to bear in mind its ability to bridge asynchronously deployed stations in designing new  
779 seismic networks.

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**Table 1.** Differences (m/s) of Rayleigh wave phase speed measurements from the  $\mathcal{I}_3$  methods compared to  $\mathcal{I}_2$  before the de-biasing correction.

Period (s)	$\mathcal{I}_3^{CW}$		$ell\mathcal{I}_3^{DW}$		$hyp\mathcal{I}_3^{DW}$	
	Mean	SD	Mean	SD	Mean	SD
10	0.1	7.3	-9.3	13.5	11.5	14.1
20	-0.0	17.0	-9.9	13.8	13.3	12.9
30	-1.4	43.2	-9.9	18.4	14.0	18.2
40	-	-	-8.1	24.1	10.8	23.6
50	-	-	-11.4	27.2	10.4	27.5
60	-	-	-12.8	27.0	8.3	28.8
70	-	-	-14.7	25.8	5.0	28.4
80	-	-	-15.0	24.2	0.6	28.8

**Table 2.** Differences (m/s) of Rayleigh wave phase speed measurements from the direct-wave  $\mathcal{I}_3$  methods compared to  $\mathcal{I}_2$  after the de-biasing correction.

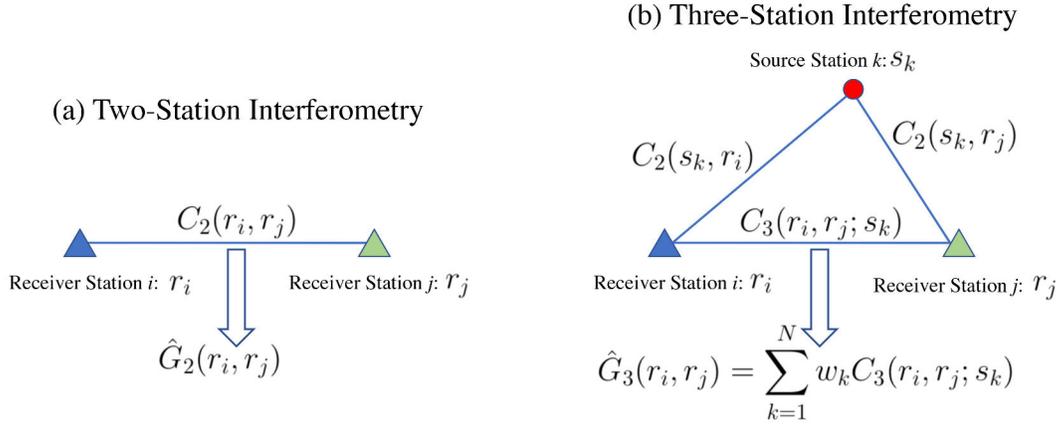
Period (s)	$ell\mathcal{I}_3^{DW}$		$hyp\mathcal{I}_3^{DW}$	
	Mean	SD	Mean	SD
10	0.6	5.5	-1.3	16.8
20	0.3	7.2	1.7	14.0
30	-0.9	11.5	3.4	19.4
40	-1.7	17.7	3.8	26.7
50	-7.6	24.2	8.1	31.7
60	-5.9	28.0	6.4	34.3
70	-6.7	28.4	4.5	35.2
80	-4.7	27.4	2.2	34.4

**Table 3.** Differences (m/s) of Rayleigh wave phase speed maps from the  $\mathcal{I}_3$  methods compared to  $\mathcal{I}_2$  after the de-biasing correction.

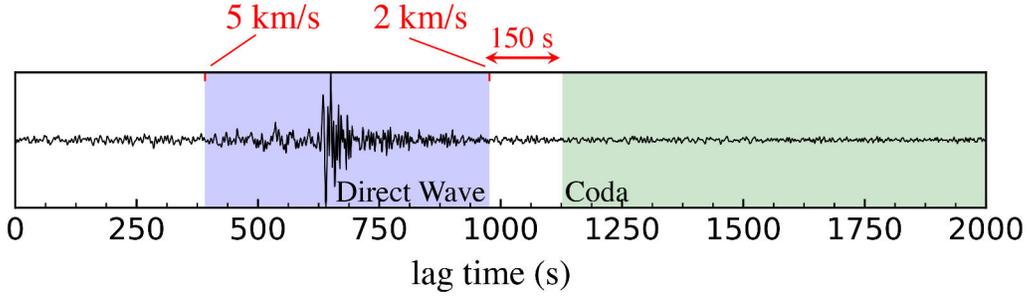
Period (s)	$\mathcal{I}_3^{CW}$		$ell\mathcal{I}_3^{DW}$		$hyp\mathcal{I}_3^{DW}$	
	Mean	SD	Mean	SD	Mean	SD
10	7.0	11.8	-0.8	8.1	-2.9	8.7
20	7.4	13.8	-1.8	5.2	-0.5	5.6
30	-	-	1.9	8.8	1.0	7.0
40	-	-	-0.4	12.3	0.8	15.4
50	-	-	9.9	16.2	2.1	17.9
60	-	-	1.2	24.1	6.5	23.3

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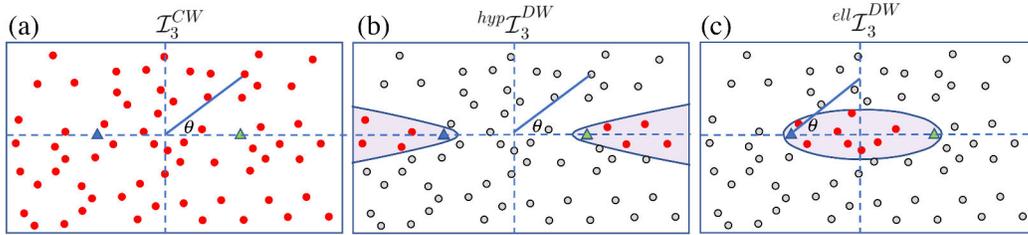
**Figures:**



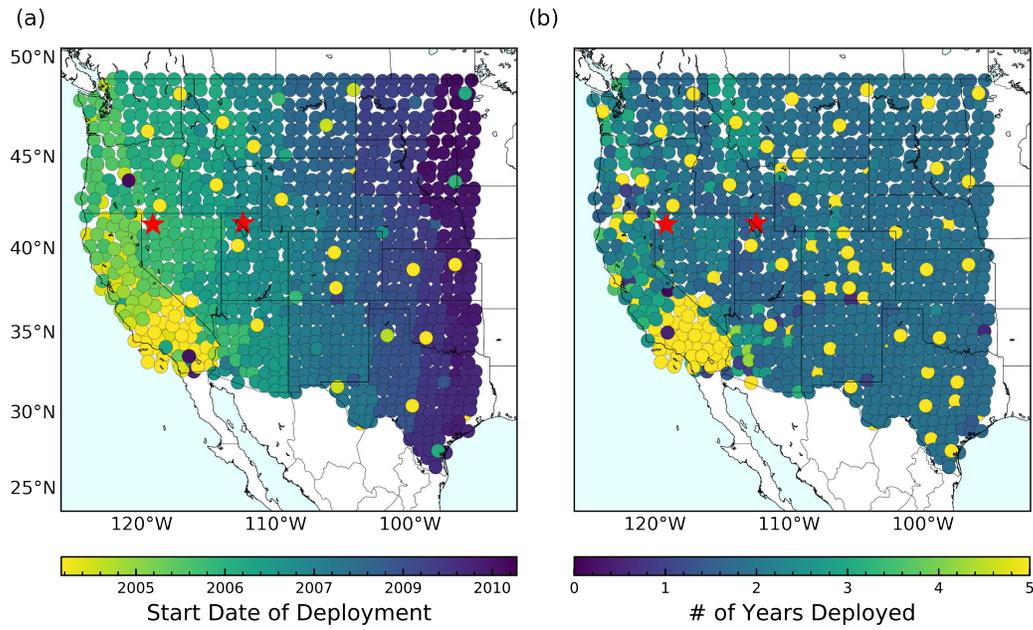
**Figure 1.** Notation for interferometry. (a) Two-station interferometry.  $C_2(r_i, r_j)$  is the cross-correlation between processed seismograms recorded at receiver-stations  $r_i$  and  $r_j$ . The two-station estimated Green’s function,  $\hat{G}_2(r_i, r_j)$ , can be determined from  $C_2$  after applying an appropriate phase shift. Receiver-stations  $r_i$  and  $r_j$  must operate synchronously. (b) Three-station interferometry. Cross-correlations between seismograms recorded at each source-station,  $s_k$ , with records at receiver-stations,  $r_i$  and  $r_j$ , are denoted  $C_2(s_k, r_i)$  and  $C_2(s_k, r_j)$ . Direct-wave or coda-wave parts of these records are cross-correlated or convolved to measure the source-specific interferogram,  $C_3(r_i, r_j; s_k)$ , which can be summed over contributions from many source-stations to produce the three-station composite Green’s function,  $\hat{G}_3(r_i, r_j)$ , between the receiver-stations. Receiver-stations  $r_i$  and  $r_j$  need not operate synchronously with one another, but both must overlap the operation of each source-station.



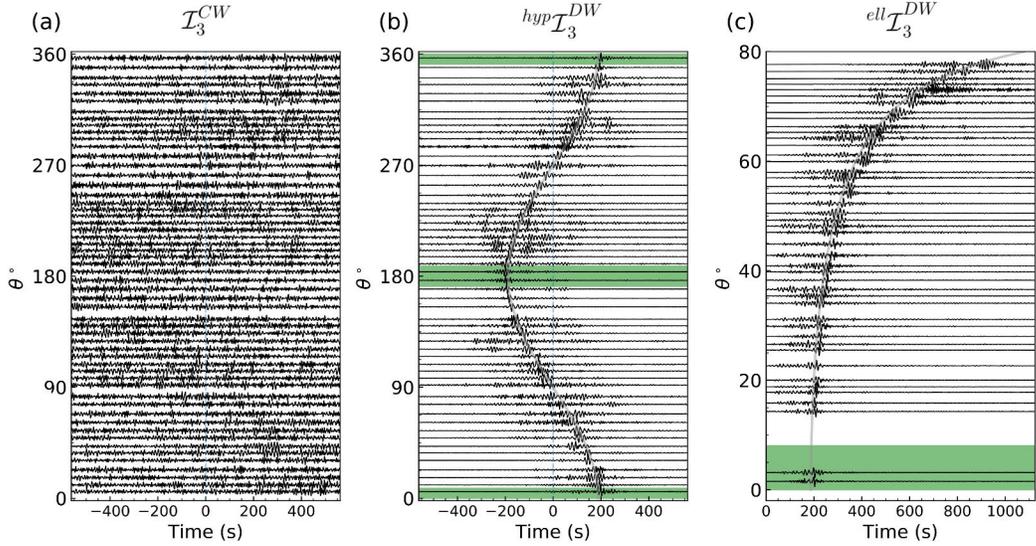
**Figure 2.** Example of the definition of the direct-wave and coda-wave segments of a two-station cross-correlation of ambient noise,  $C_2$ , for stations ANMO (Albuquerque, NM) and M47A (Cromwell, IN), at an inter-station distance of  $\sim 1950$  km. The direct-wave is the segment of the record between times corresponding to group speeds of 2 and 5 km/s. The coda-wave segment starts 150 s after the end of the direct-wave, and extends to the end of 3000 s. The symmetric component of the cross-correlation is shown (average of positive and negative correlation lags).



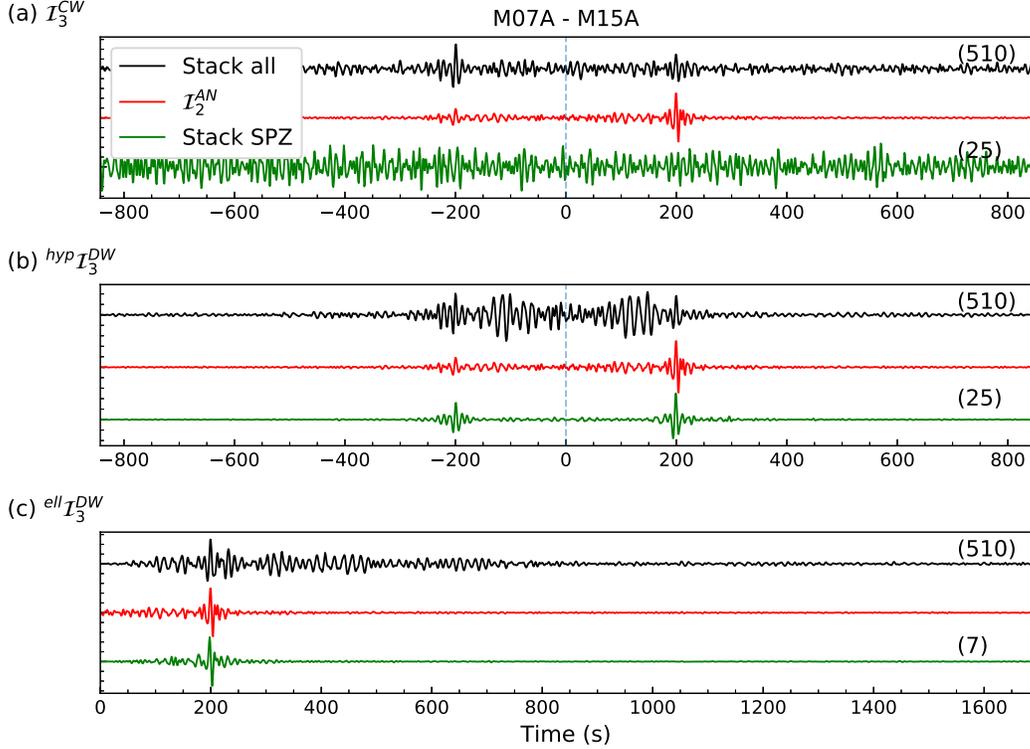
**Figure 3.** Schematic illustration of the geometrical constraints on source-stations for different methods of three-station interferometry. The two receiver-stations are shown with the blue and green triangles, and the circles are locations of other stations that may act as source-stations. Those stations that can act as source-stations are shown with red circles and those cannot with grey circles. (a) For three-station coda-wave interferometry,  $\mathcal{I}_3^{CW}$ , all stations whose operation overlaps the two receiver-stations can act as a source-station. (b) For three-station direct-wave interferometry with source-stations radially outside the receiver-stations,  $^{hyp}\mathcal{I}_3^{DW}$ , source-stations must lie in stationary phase hyperbolae (purple shading). (c) For three-station direct-wave interferometry with source-stations between the receiver-stations,  $^{ell}\mathcal{I}_3^{DW}$ , source-stations must lie in the stationary phase ellipse (purple shading). The angle  $\theta$  in each case is defined in [section 3.1](#) and used in [Fig. 5](#).



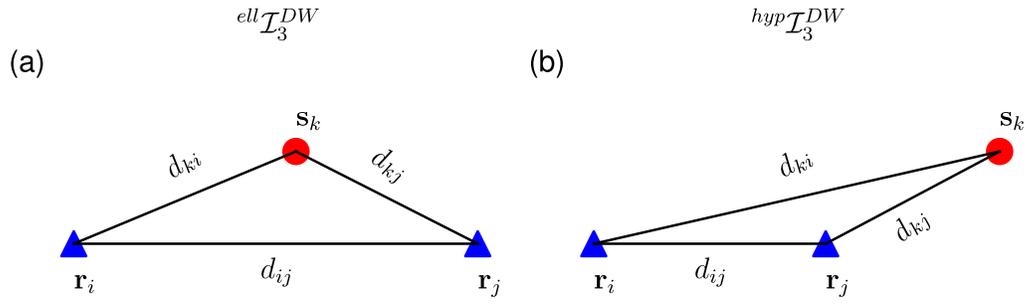
**Figure 4.** Map of stations used in this study. Red stars mark stations used in **Figs 5, 6 and 8**: M07A (Soldier Meadow, NV) and M15A (Promontory, UT). (a) The start dates for each station are color-coded, showing a rolling pattern from west to east. (b) Duration of deployment is color-coded. Most stations are deployed around two years with a few much longer from the USArray Reference Network (`_US-REF`) and the Southern California Seismic Network (CI).



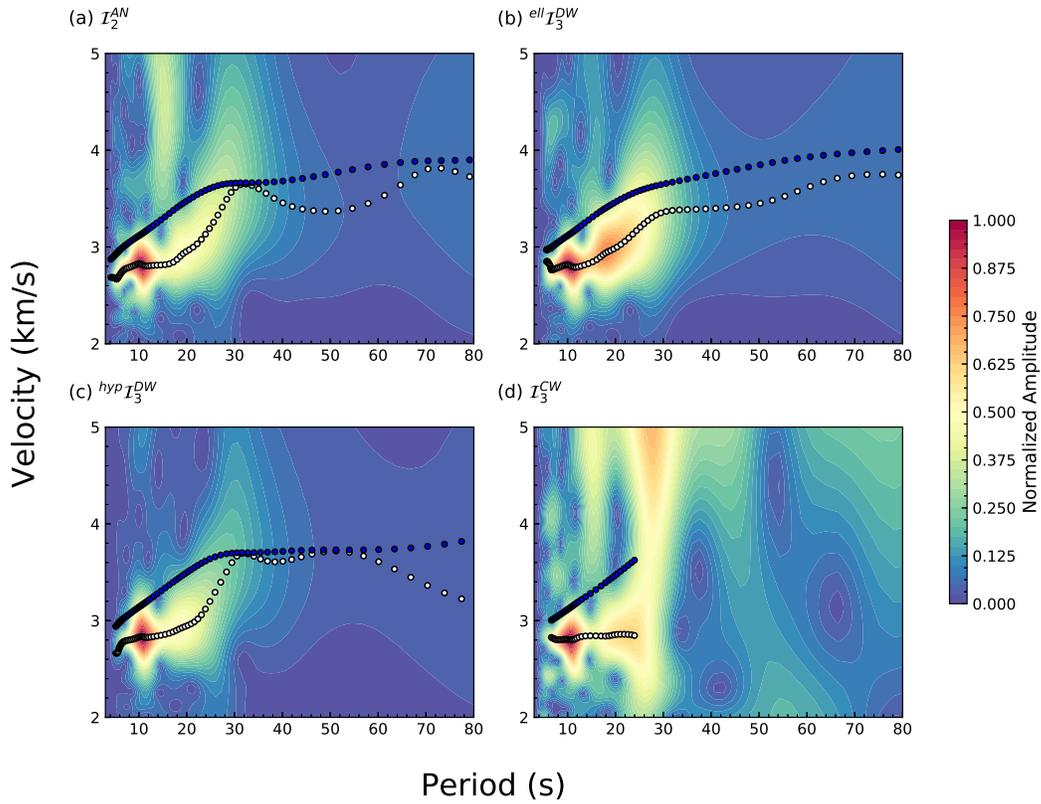
**Figure 5.** Example record sections of three-station interferograms for the receiver-station pair M07A-M15A, whose locations are shown in **Fig. 4**. (a) Coda-wave correlations ( $C_3^{CW}$ ) for different source-stations plotted at the azimuth angle  $\theta$  shown in **Fig. 3a**. (b) Direct-wave correlations ( $C_3^{DW}$ ) plotted for source-stations at the azimuth angle shown in **Fig. 3b**. The green regions are the hyperbolic stationary-phase zones for  $^{hyp}I_3^{DW}$ . (c) Direct-wave convolutions ( $C_3^{DW}$ ) plotted for source-stations at the azimuth angle shown in **Fig. 3c**. Only positive time lags are defined. The green region is the elliptical stationary-phase zone for  $^{ell}I_3^{DW}$ . Grey curves in (b) and (c) are predictions from eqs. (4) and (5), respectively, with  $c = 3$  km/s. Only selected three-station interferograms are shown to ease visualization.



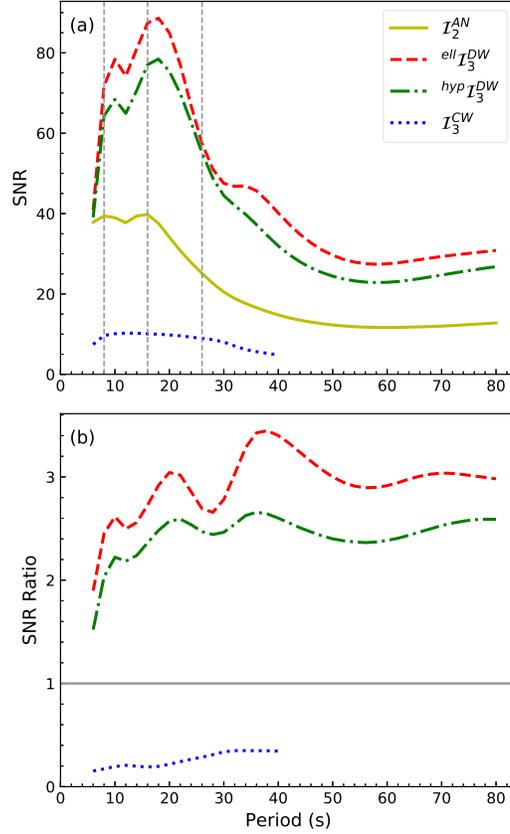
**Figure 6.** Examples of stacks of three-station interferograms for the receiver-station pair M07A-M15A of **Fig. 5**. In each panel the two-station estimated Green's function ( $\mathcal{I}_2^{AN}$ ) is plotted for reference (red). The number of source-stations for each stack is shown in parentheses above the stacked trace. (a) Method  $\mathcal{I}_3^{CW}$ . Two stacks of coda-wave interferograms are shown: (black line) stack of the interferograms from all source-stations irrespective of the azimuthal angle  $\theta$  (defined in **Fig. 3a**) and (green line) stack of the coda-wave interferograms for sources in the hyperbolic stationary phase zone. For  $\mathcal{I}_3^{CW}$ , the black line is the composite Green's function. (b) Method  $^{hyp}\mathcal{I}_3^{DW}$ . Black and green lines have similar meanings to those in (a), but here the direct-wave interferograms are stacked. For  $^{hyp}\mathcal{I}_3^{DW}$ , the green line is the composite Green's function. (c) Method  $^{ell}\mathcal{I}_3^{DW}$ . Black line is the same as in (b), but the green line is the stack of direct-wave interferograms in the elliptical stationary phase zone. For  $^{ell}\mathcal{I}_3^{DW}$ , the green line is the composite Green's function and only positive time lags are defined.



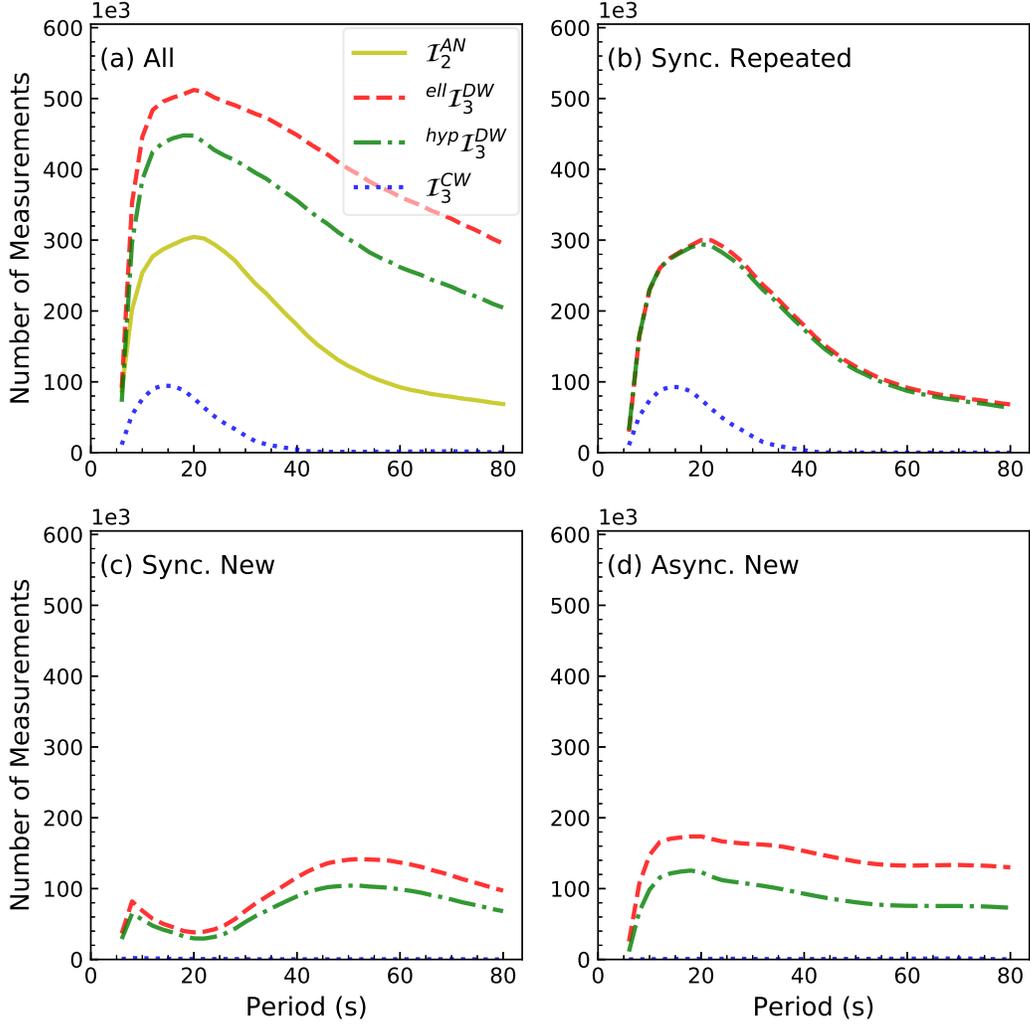
**Figure 7.** Geometry of the source-station ( $s_k$ ) and receiver-stations ( $r_i, r_j$ ) used to determine the phase for the three-station direct-wave methods: (a)  $ell\mathcal{I}_3^{DW}$  and (b)  $hyp\mathcal{I}_3^{DW}$ . Great circle distances between two stations are denoted as  $d$  with appropriate subscripts.



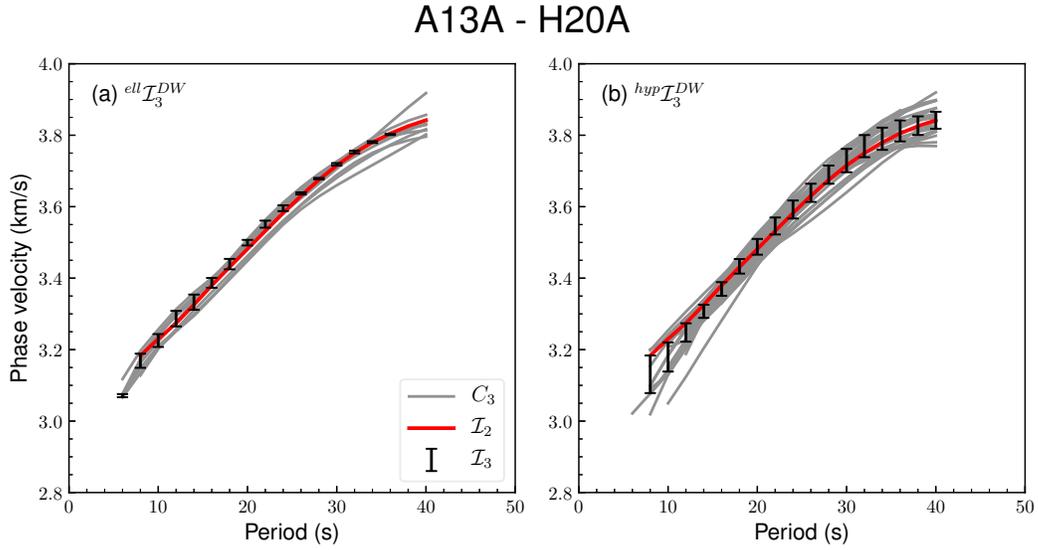
**Figure 8.** Frequency-time analysis (FTAN) diagrams for the receiver pair M07A-M15A using the waveforms from **Fig. 6**: (a)  $\mathcal{I}_2^{AN}$ , (b)  $\mathcal{I}_3^{CW}$ , (c)  $hyp\mathcal{I}_3^{DW}$ , and (d)  $ell\mathcal{I}_3^{DW}$ . White and blue circles are group and phase speed measurements, respectively.



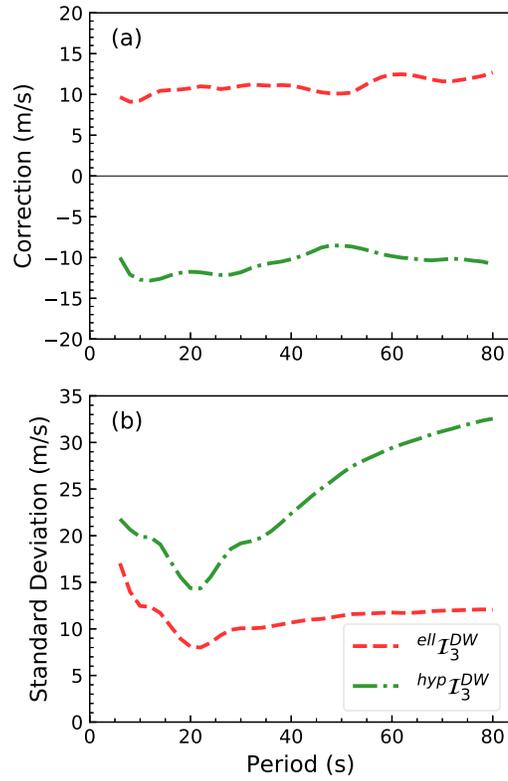
**Figure 9.** Signal-to-noise ratio (SNR) of estimated Green’s functions for the different interferometric methods (see legend) plotted versus period. (a) Median of the SNR for each method taken over all measurements at each period. SNR generally decreases with period for all methods, but the highest SNR is from the three-station direct-wave method with an elliptical stationary phase zone ( $ell\mathcal{I}_3^{DW}$ ) and the lowest is from the three-station coda-wave method ( $\mathcal{I}_3^{CW}$ ). (b) Paths common to two-station and three-station interferometry in (a) are selected such that the ratio of the median SNR for each three-station method to that for the two-station method is shown. The direct-wave methods increase SNR relative to  $\mathcal{I}_2^{AN}$  by a factor ranging from about 1.5 to 3 which grows with period, whereas the coda-wave method reduces SNR by a factor of 3-5.



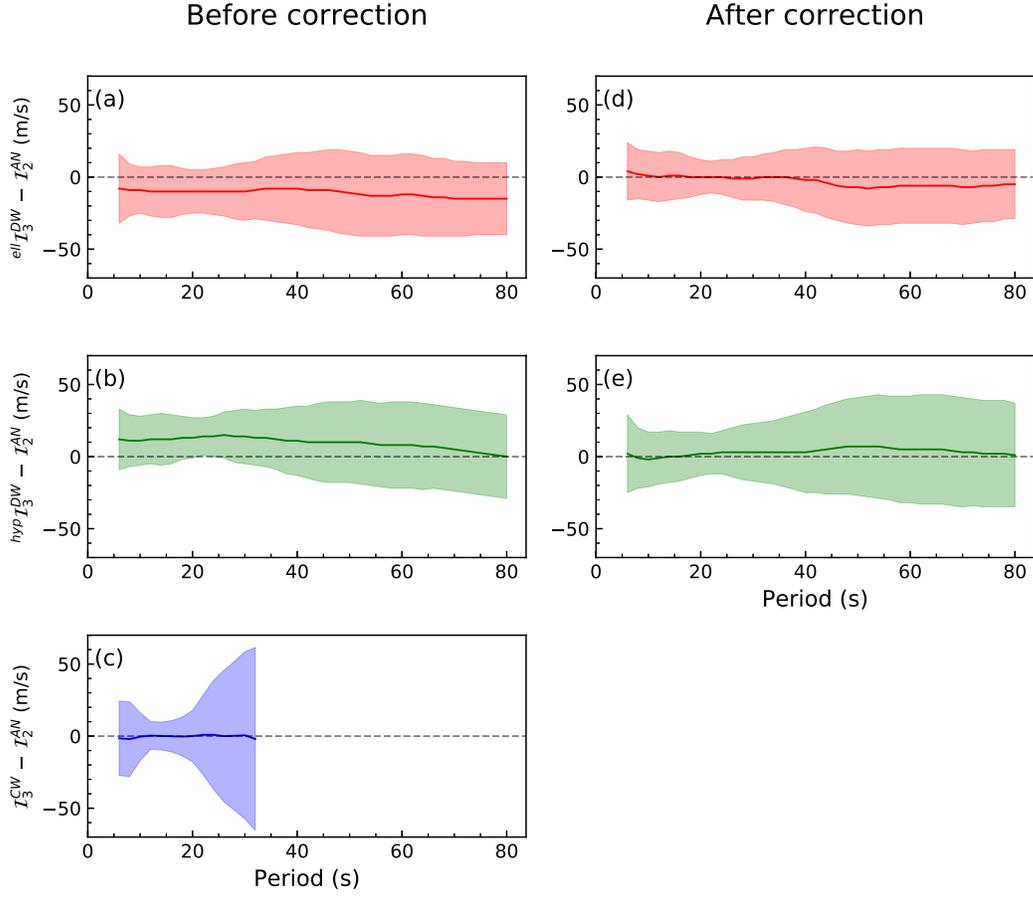
**Figure 10.** Number of resulting measurements (in thousands) versus period. (a) Number of accepted Rayleigh wave phase speed measurements plotted versus period for the different interferometric methods (see legend). The largest number of measurements is from the three-station direct-wave method with an elliptical stationary phase zone ( $^{ell}\mathcal{I}_3^{DW}$ ) and the smallest number is from the three-station coda-wave method ( $\mathcal{I}_3^{CW}$ ). The total number of measurements can be broken into three parts, as shown in (b)-(d). (b) Number of measurements from  $\mathcal{I}_3$  that exist for  $\mathcal{I}_2^{AN}$ . (c) Number of synchronous measurements from three-station interferometry methods ( $\mathcal{I}_3$ ) that are non-existent for two-station interferometry  $\mathcal{I}_2^{AN}$  (because of low SNR). (d) Number of asynchronous measurements from  $\mathcal{I}_3$  that are non-existent for  $\mathcal{I}_2^{AN}$  (because of asynchrony).



**Figure 11.** Examples of the de-biased Rayleigh wave phase speed curves for the receiver-station pair A13A (Polebridge, MT) and H20A (Greybull, WY) for the two three-station direct-wave methods: (a)  $^{ell}\mathcal{I}_3^{DW}$  and (b)  $^{hyp}\mathcal{I}_3^{DW}$ . Each gray curve is measured for a single source-specific interferogram ( $C_3$ ), where there are 9 source-stations for  $^{ell}\mathcal{I}_3^{DW}$  and 32 source-stations for  $^{hyp}\mathcal{I}_3^{DW}$ . The mean and standard deviation of the constituent curves are plotted with the black error bars. The two-station ambient noise ( $\mathcal{I}_2$ ) dispersion curve is shown in red.

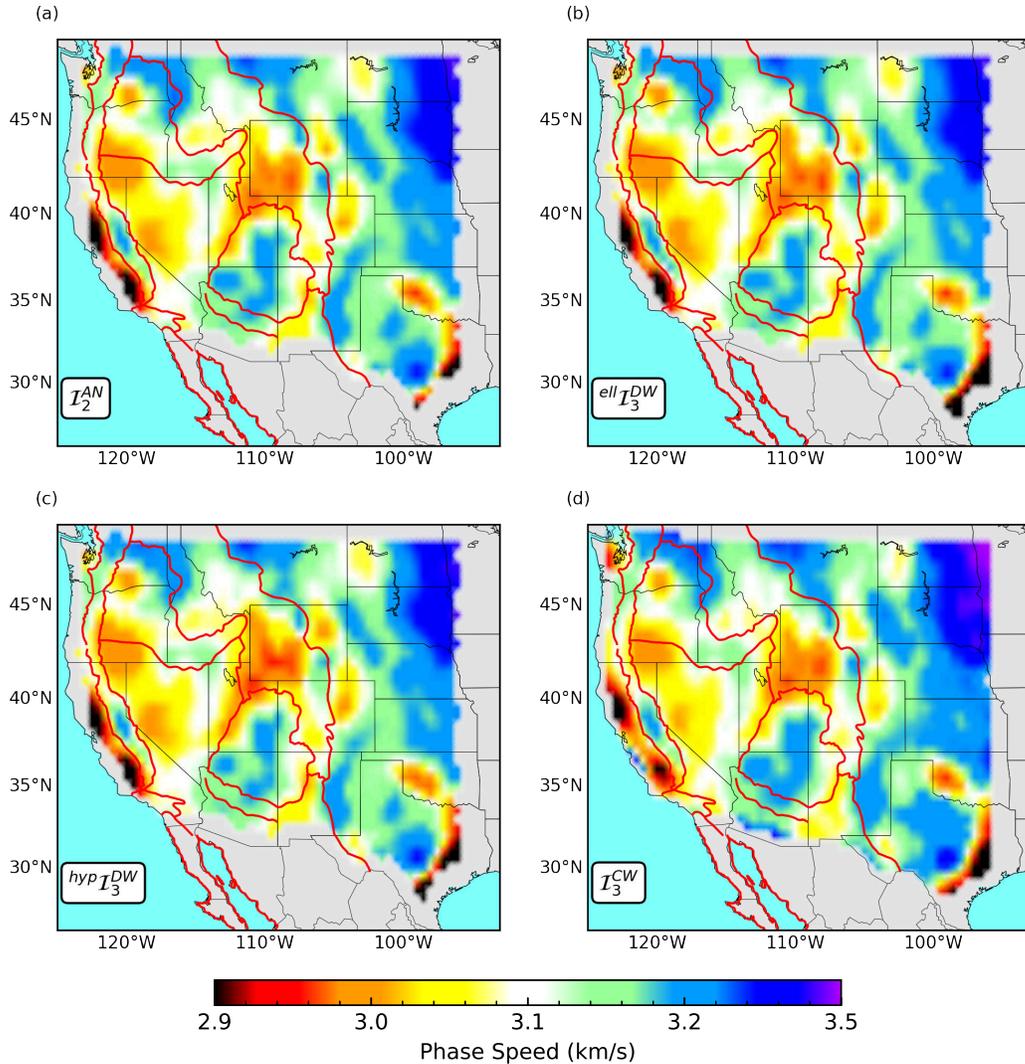


**Figure 12.** (a) Mean de-biasing correction averaged over all receiver-station pairs in the data set for  $ell \mathcal{I}_3^{DW}$  (red line) and  $hyp \mathcal{I}_3^{DW}$  (green line). (b) Standard deviation of the de-biased dispersion curves averaged over all receiver-station pairs in the data set.



**Figure 13.** Mean and standard deviation of the difference between Rayleigh wave phase speed measurements from the three-station methods ( $\mathcal{I}_3$ ) and the two-station ( $\mathcal{I}_2^{AN}$ ) method. (a)-(c) No bias correction has been applied. Measurements from the direct-wave three-station methods ( $\mathcal{I}_3^{DW}$ ) are systematically shifted from the  $\mathcal{I}_2^{AN}$  measurements, albeit with different signs, whereas the coda-wave measurements ( $\mathcal{I}_3^{CW}$ ) are not shifted relative to those from  $\mathcal{I}_2^{AN}$ . The standard deviation of the differences between the three-station and two-station measurements grow with period generally, but minimize around 20 s. (d)-(e) Similar to (a)-(b), but the  $\mathcal{I}_3^{DW}$  methods have been de-biased based on ray-theory. Systematic differences in Rayleigh wave phase speed measurements compared to the  $\mathcal{I}_2^{AN}$  method are largely removed at periods below 40 s, and are reduced at longer periods compared to the uncorrected values. The statistics are tabulated in **Tables 1, 2**.

10 s



**Figure 14.** Rayleigh wave phase speed maps constructed with eikonal tomography at 10 s period using four different interferometric methods: (a) traditional two-station ambient noise interferometry ( $\mathcal{I}_2^{AN}$ ), (b) three-station direct-wave interferometry with elliptical stationary phase zone ( $ell \mathcal{I}_3^{DW}$ ), (c) three-station direct-wave interferometry with hyperbolic stationary phase zone ( $hyp \mathcal{I}_3^{DW}$ ), and (d) three-station coda-wave interferometry ( $\mathcal{I}_3^{CW}$ ). Red lines depict geological provinces (Fenneman & Johnson, 1946).

20 s

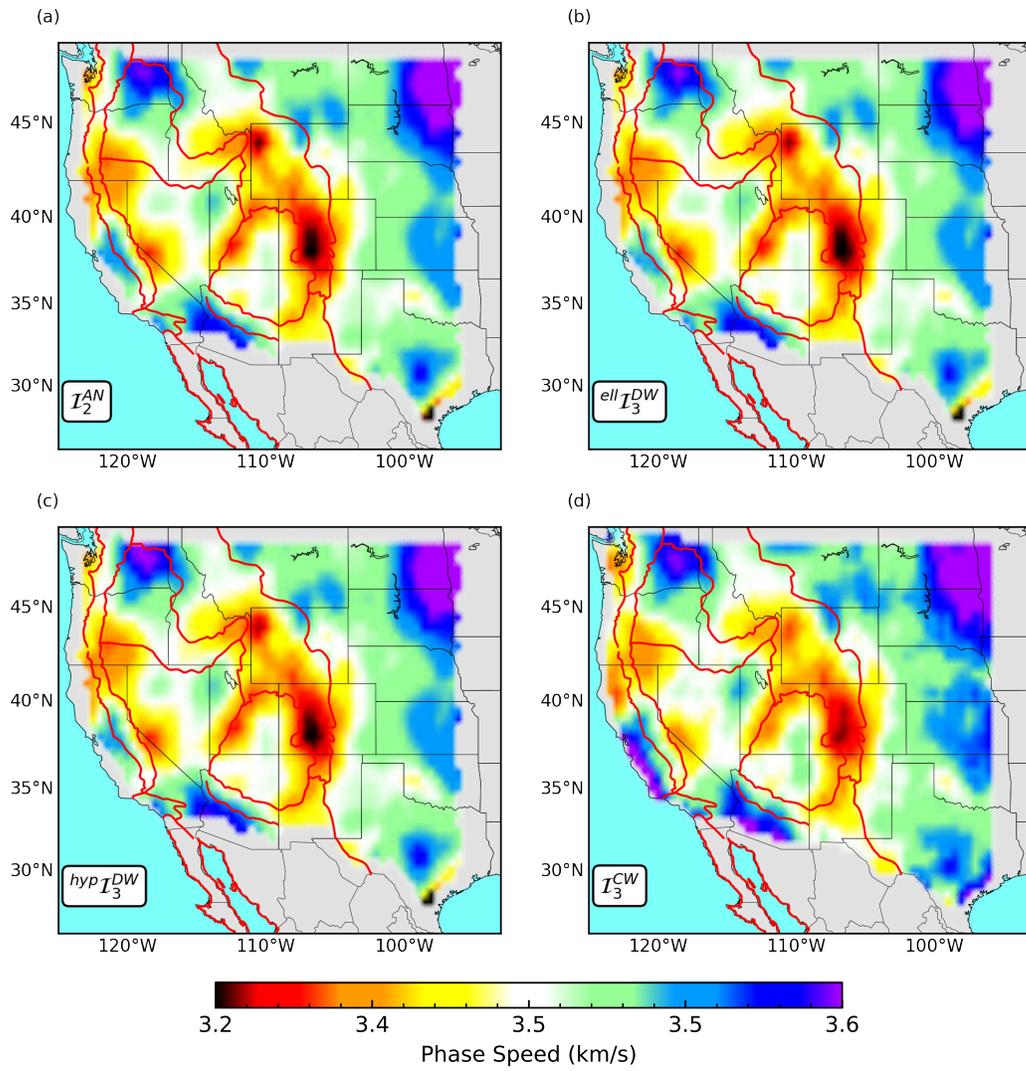
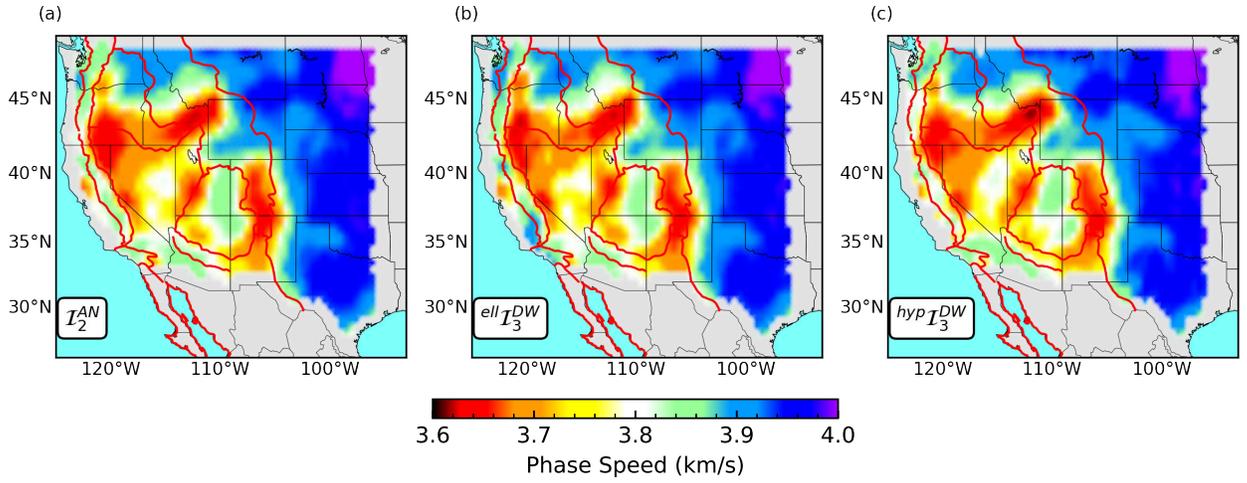


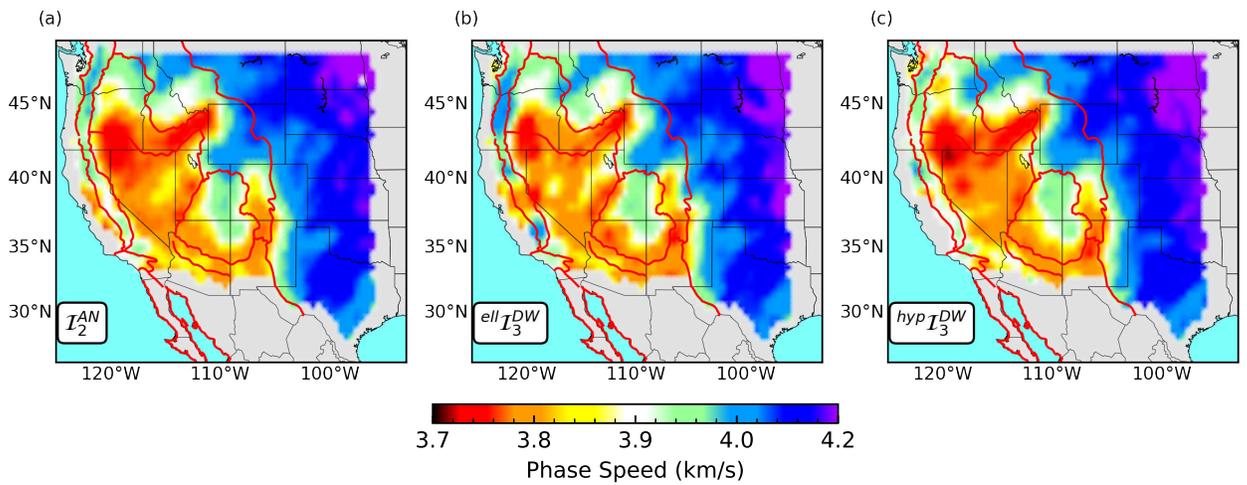
Figure 15. Similar to Fig. 14, but at 20 s period.

40 s

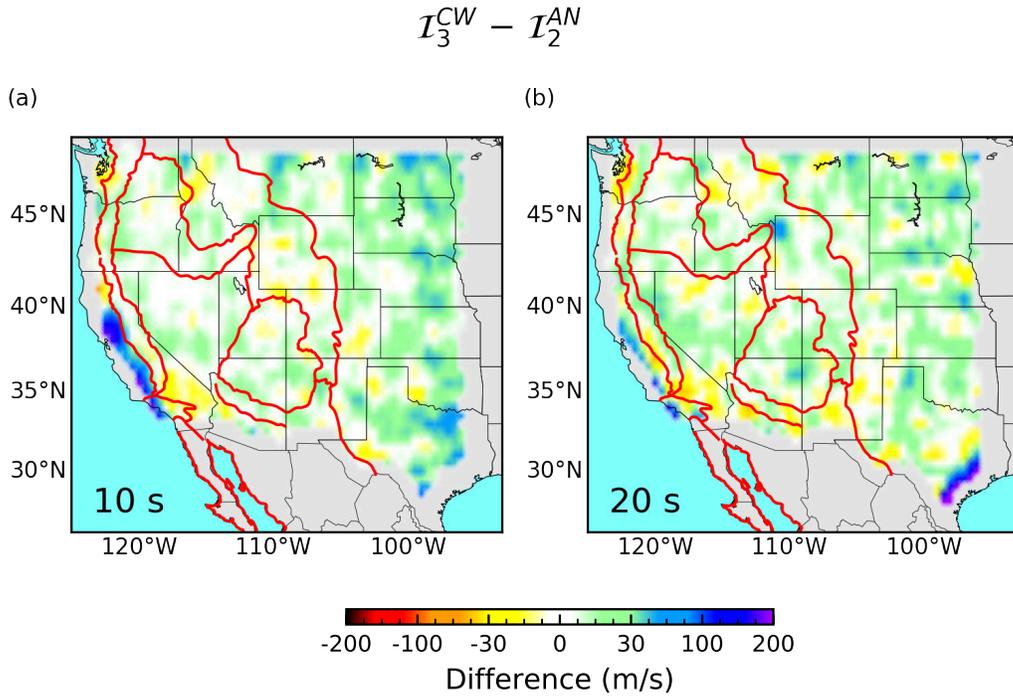


**Figure 16.** Similar to Fig. 14, but at a period of 40 s.  $I_3^{CW}$  yielded too few measurements to produce a tomographic map.

60 s

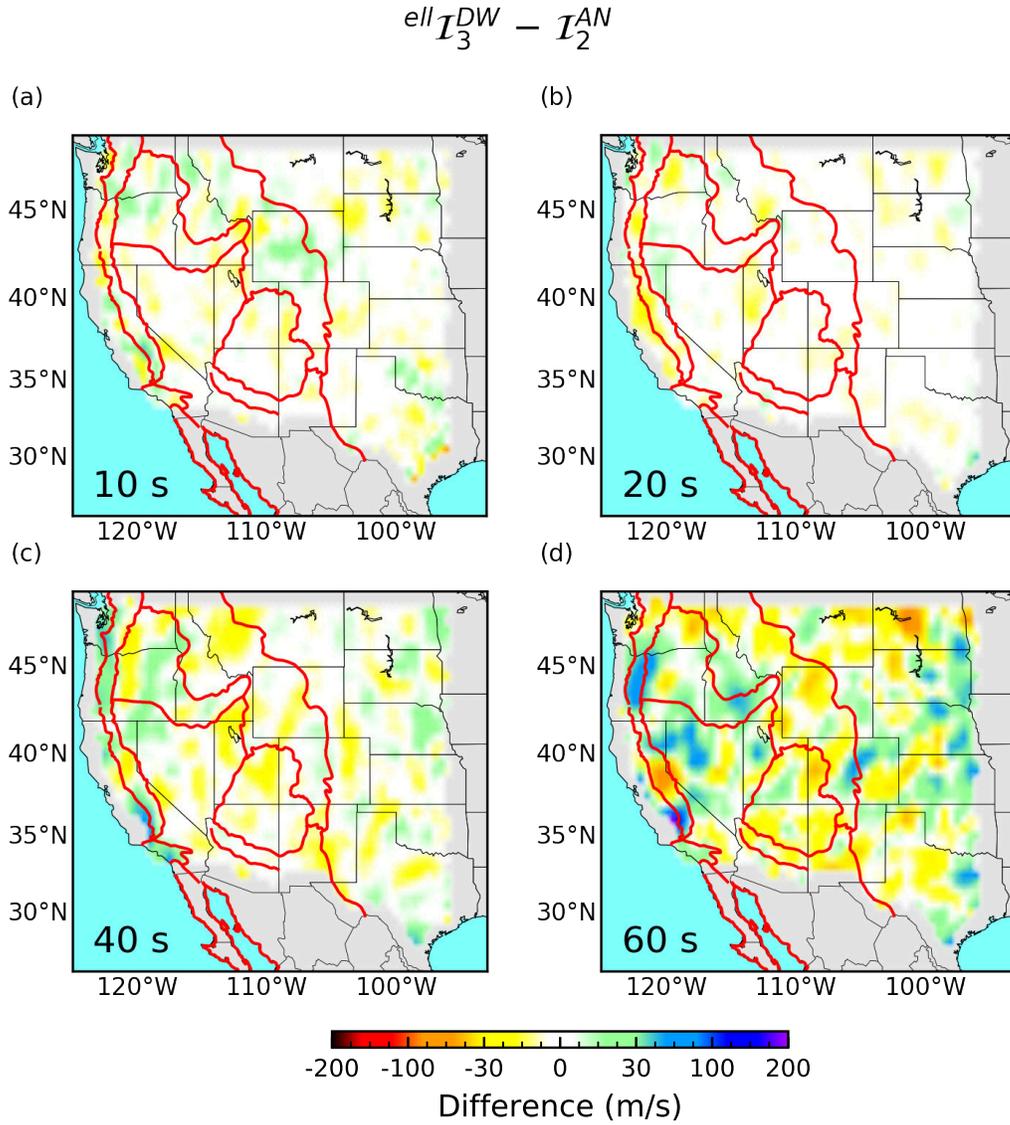


**Figure 17.** Similar to Fig. 16, but at a period of 60 s.

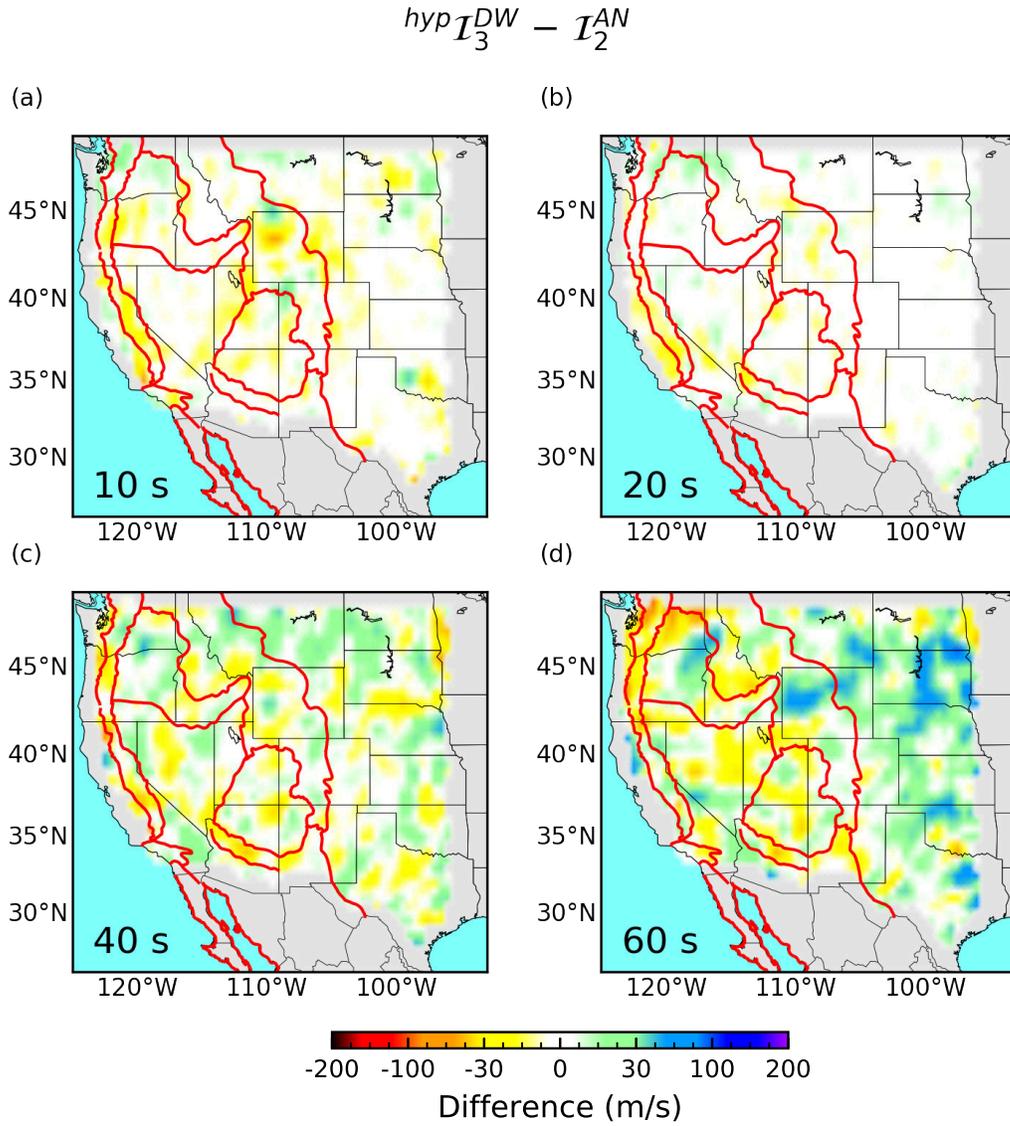


**Figure 18.** Differences in Rayleigh wave phase speed maps (Figs 14 and 15) between three-station coda-wave interferometry ( $I_3^{CW}$ ) and two-station ambient noise interferometry ( $I_2^{AN}$ ).

$I_3^{CW}$  yields too few measurements to produce tomographic maps at longer periods.

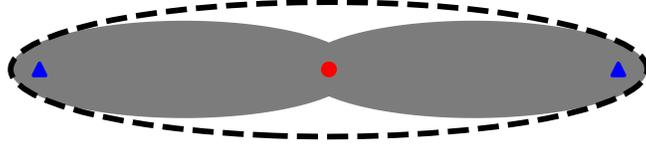


**Figure 19.** Similar to **Fig. 18** except differences are between three-station direct-wave interferometry  $ell \mathcal{I}_3^{DW}$  and  $\mathcal{I}_2^{AN}$  (**Figs 14 - 16**), and results are presented at four periods: 10 s, 20 s, 40 s, and 60 s.

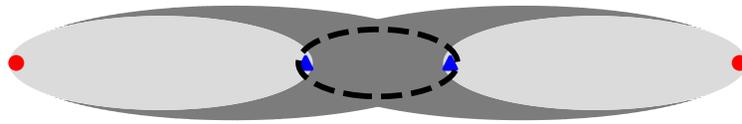


**Figure 20.** Similar to **Fig. 19** except between  $hyp I_3^{DW}$  and  $I_2^{AN}$ .

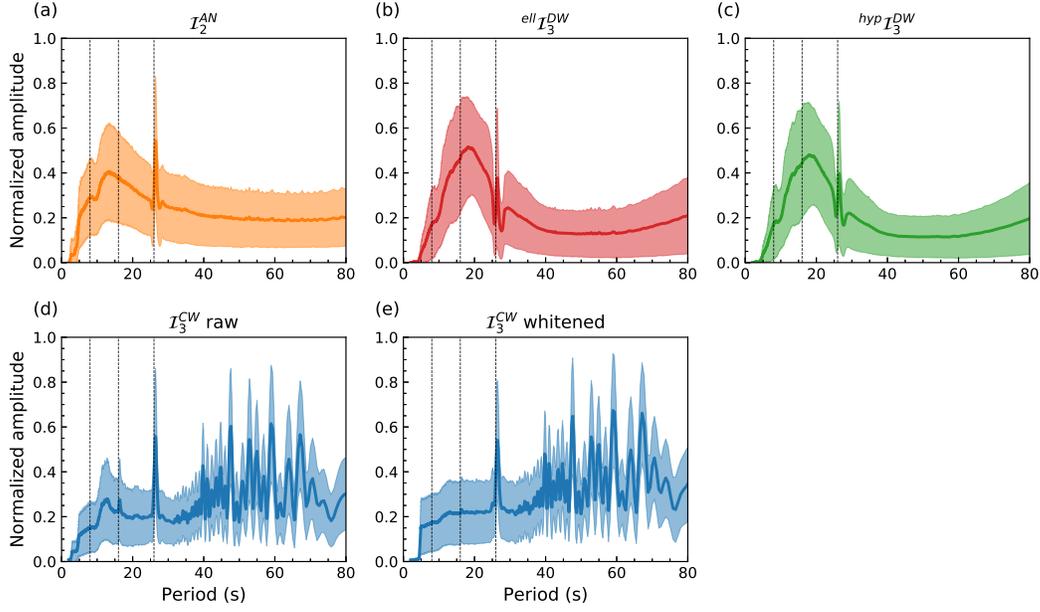
(a)  $ell\mathcal{I}_3^{DW}$



(b)  $hyp\mathcal{I}_3^{DW}$



**Figure 21.** Schematic illustration contrasting the sensitivity kernels for  $ell\mathcal{I}_3^{DW}$  and  $hyp\mathcal{I}_3^{DW}$  with that for  $\mathcal{I}_2$  which is shown as a Fresnel ellipse encompassing the two receiver-stations (blue triangles) and is depicted with the dashed lines. (a) The sensitivity kernel for  $ell\mathcal{I}_3^{DW}$  is a superposition of the two elliptical Fresnel zones where the source-station (red dot) is at one focus and each of the receiver-stations are at the other foci. The resulting sensitivity kernel for  $ell\mathcal{I}_3^{DW}$  (grey region) is smaller than the kernel for  $\mathcal{I}_2$  (zone encompassed by the dashed line). (b) The sensitivity kernel for  $hyp\mathcal{I}_3^{DW}$  is the difference of the two elliptical Fresnel zones where each source-station (red dots) is at one focus and each of the receiver-stations is at the other focus. The resulting sensitivity kernel for  $hyp\mathcal{I}_3^{DW}$  (grey region) is more complicated and larger than the kernel for  $\mathcal{I}_2$  (zone encompassed by the dashed line).



**Figure 22.** Mean (lines) and standard deviation (shaded areas) of spectra for (a) traditional two-station ambient noise interferometry ( $\mathcal{I}_2^{AN}$ ), (b) three-station direct-wave interferometry with elliptical stationary phase zone ( $^{ell}\mathcal{I}_3^{DW}$ ), (c) three-station direct-wave interferometry with hyperbolic stationary phase zone ( $^{hyp}\mathcal{I}_3^{DW}$ ), (d) three-station coda-wave interferometry ( $\mathcal{I}_3^{CW}$ ) without spectral whitening, and (e)  $\mathcal{I}_3^{CW}$  with spectral whitening. Dashed lines mark the secondary microseism peak (8 s), the primary microseism peak (16 s), and the 26 s microseism. The 26 s microseism shows a peak across all methods. Whitening of  $\mathcal{I}_3^{CW}$  only makes the spectra flatter at short periods but does not eliminate the 26 s peak. The spectra of  $\mathcal{I}_3^{CW}$  show strong variability at long periods which is not observed in other methods. Whitening of  $\mathcal{I}_3^{CW}$  does not reduce the variability at long periods.