

# **Broadband Ocean Bottom Seismometer Noise Properties**

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## Summary

We present a new compilation and analysis of broadband ocean bottom seismometer noise properties from 15 years of seismic deployments. We compile a comprehensive dataset of representative four-component (seismometer and pressure gauge) noise spectra and cross-spectral properties (coherence, phase, and admittance) for 551 unique stations spanning 18 US-led experiments. This is matched with a comprehensive compilation of metadata parameters related to instrumentation and environmental properties for each station. We systematically investigate the similarity of noise spectra by grouping them according to these metadata parameters to determine which factors are the most important in determining noise characteristics. We find evidence for improvements in similarity of noise properties when grouped across parameters, with groupings by seismometer type and deployment water depth yielding the most significant and interpretable results. Instrument design, that is the entire deployed package, also plays an important role, although it strongly covaries with seismometer and water depth. We assess the presence of traditional sources of tilt, compliance, and microseismic noise to characterize their relative role across a variety of commonly used seismic frequency bands. We find that the presence of tilt noise is primarily dependent on the type of seismometer used (covariant with a particular subset of instrument design), that compliance noise follows anticipated relationships with water depth, and that shallow, oceanic shelf environments have systematically different microseism noise properties (which are, in turn, different from instruments deployed in shallow lake environments). These observations have important implications for the viability of commonly used seismic analysis techniques. Finally, we compare spectra and coherences before and after vertical channel tilt and compliance noise removal to evaluate the efficacy and limitations of these now standard processing techniques. These findings may assist in future experiment planning and instrument development, and our newly compiled noise dataset serves as a building block for more targeted future investigations by the marine seismology community.

## Key Words

Seismic noise; Seismic instruments; Instrumental noise; Site effects; Pacific Ocean; Atlantic Ocean

### 1 1. Introduction

2 Over recent decades, the marine seismological community has made steady progress in the  
3 deployment of increasingly high-quality and large(r)-N broadband ocean-bottom seismometer  
4 (BBOBS) networks. It is approximately 30 years since the advent of modern-standard ocean-floor  
5 seismic instruments (Cox, et al., 1984; Montagner et al., 1994; Webb et al., 1994; Purdy & Orcutt,  
6 1995; Collins et al., 2001; Stephen et al., 2003) led to the formation of the Ocean Bottom Seismic  
7 Instrument Pool (OBSIP) in 1999 (Aderhold et al., 2019). It is approximately 10 years since the

8 conception and execution of one of the most ambitious community BBOBS deployments to date:  
9 the Cascadia Initiative (Toomey et al., 2014). Systematic archiving of seismic and pressure-gauge  
10 data at the Incorporated Research Institutions for Seismology (IRIS) Data Management Center  
11 (DMC), along with community tools for preprocessing BBOBS data (*e.g.*, ATaCR, Janiszewski et  
12 al., 2019; DLOPy, Doran & Laske, 2017) have substantially expanded the reach and salience of  
13 BBOBS data across the wider seismological community. With the recent reorganization of the  
14 national US instrument pool into the Ocean Bottom Seismic Instrument Center (OBSIC), the  
15 emergence of improving hardware, renewed planning towards long-term BBOBS observatories  
16 (Kohler et al., 2020), and the evolution of new seafloor technologies (Spica et al., 2020; Lior et  
17 al., 2021), this is an apposite juncture to assess the systematics of BBOBS noise traits.

18 Analyses of spectra (including both seismometer and pressure-gauge data) for individual  
19 deployments suggest significant seismic noise variations exist among deployed BBOBSs (Yang et  
20 al., 2012; Sumy et al., 2015; Barcheck et al., 2020; An, 2021). Direct comparisons of their noise  
21 have largely focused on single-instrument tests rather than arrays (Webb, 1998), or pilot studies  
22 that contrast different installation techniques (*e.g.*, Collins et al., 2001). These are limited in  
23 location and do not include all types of BBOBS design in the modern OBSIC fleet. To date, a  
24 systematic noise comparison across deployments that encompasses the full range of instrument  
25 designs, water depths, and site conditions does not exist. Recent analyses using the Cascadia  
26 Initiative dataset demonstrate variability as a function of instrument design and water depth (Bell  
27 et al., 2015; Hilmo & Wilcock, 2020), motivating expansion of such analyses across deployments.

28 In this study, we present a comprehensive dataset describing the last 15 years of US-funded  
29 BBOBS array deployments (Figure 1). We compute representative multi-component noise spectra  
30 from stations deployed in a variety of environments, water depths, and using several different

31 instrument designs. We calculate cross-channel coherences, upper and lower bounds on typical  
32 noise, and investigate systematics of noise within a variety of frequency bands spanning from  
33 0.001 to 1 Hz. Using this dataset, we offer a comprehensive and quantitative review of the  
34 character and sources of noise on BBOBS instruments.

35

## 36 **2. Background**

### 37 **2.1 Noise Sources**

38 The noise power spectrum from 0.001–1 Hz on BBOBS instruments is influenced by the presence  
39 and strength of microseism noise, infragravity waves, and tilt or bottom current noise, as well as  
40 complications due to instrument response and shear-mode wave propagation at higher frequencies  
41 (Figure 2). Microseism noise is the broad, high-amplitude peak between 0.5-0.05 Hz that  
42 dominates the ambient seismic energy field world-wide (Peterson, 1993; McNamara & Buland,  
43 2004). Its prevalence has resulted in the traditional distinction between “high frequency” ( $> 1$  Hz)  
44 and “low frequency” ( $< 0.05$  Hz) low-noise observational seismic bands. To first order, seafloor  
45 observations of microseismic noise are consistent with the long history of observations on land.  
46 The noise spectrum in this band is typically divided into two peaks - the secondary microseism,  
47 with multiple sub-peaks at frequencies within 0.1 - 0.5 Hz (Stephen et al., 2003) and a dominant  
48 peak located at  $\sim 0.14$  Hz, and the primary microseism, which peaks at  $\sim 0.07$  Hz (Webb, 1998).

49 Acoustic waves produced by the interaction of wind-generated ocean waves with the seafloor  
50 constitute the secondary microseism (Longuet-Higgins, 1950). Several dominant mechanisms  
51 generate these interacting waves, including storm-generated swell, coastline reflected waves, and  
52 interactions of waves generated by multiple storms (Bromirski et al., 2005; Ardhuin et al., 2011).

53 Broadly, the amplitudes of higher frequency energy within the secondary microseism correlate  
54 with the local sea state, while waves generated from distant storms and their coastal reflections  
55 play a more important role at longer periods within this band (Babcock et al., 1994; Stephen et al.,  
56 2003; Bromirski et al., 2005). These properties can lead to systematic differences between ocean  
57 basins. The Pacific Ocean appears to propagate energy from larger, more distant storms with  
58 higher sustained wind speeds, leading to a secondary microseism peak that extends to lower  
59 frequencies than in the North Atlantic Ocean (Babcock et al., 1994; Webb, 1998), although only  
60 limited numbers of instruments were used for these early measurements. More recently, Yang et  
61 al. (2012) observed systematic differences in long period microseismic energy between BBOBSs  
62 deployed in the South Pacific and the Tasman Sea off opposite coasts of New Zealand, with  
63 instruments in the marginal sea relatively deficient in longer period energy. Additionally, the  
64 secondary microseism peak may shift to higher frequencies in lake environments (Xu et al., 2017;  
65 Smalls et al., 2019).

66 The primary microseism peak is generated by direct interaction (shoaling) of ocean waves with  
67 the shoreline and rough seafloor topography (Hasselmann, 1963; Arduin, 2018). In the deep  
68 ocean, the primary microseism peak is weaker than the secondary microseism (Arduin et al.,  
69 2015). At longer periods than the primary microseism is a noise spectral-amplitude minimum  
70 termed the “noise notch” (Webb, 1998). Significant differences exist in the microseism properties  
71 between deep and shallow water; in shallow water, the primary microseism has higher amplitudes  
72 than the secondary microseism due to the direct coupling between the ocean swell with the seafloor  
73 (Webb & Crawford, 2010; Hilmo & Wilcock, 2020). This also reduces or removes the noise notch  
74 at shallow BBOBS (Hilmo & Wilcock, 2020).

75 At lower frequencies still, noise from infragravity-waves and bottom-currents is prevalent in  
76 BBOBS data. These signals are largely absent or strongly diminished at onshore sites. Infragravity  
77 waves are long period ( $< \sim 0.03$  Hz) ocean waves generated in coastal regions. Typically, these  
78 have maximum amplitudes along the continental shelves, but a small amount of infragravity wave  
79 energy may reach and subsequently propagate efficiently across the open ocean (Webb et al., 1991;  
80 Uchiyama & McWilliams, 2008; Ardhuin et al., 2014). Propagation into the deep ocean depends  
81 on coastal morphology (Aucan & Ardhuin, 2013; Crawford et al., 2015; Bogiatzis et al., 2020).  
82 Infragravity waves in the deep ocean perturb the seafloor at long ( $> 40$  km) wavelength, such that  
83 coherent signals are observed on the seismic and pressure channels of BBOBSs (Crawford et al.,  
84 1991). On the seismometers, this compliance noise is strongest on the vertical component.  
85 Compliance noise also affects horizontal components, but is typically obscured by other noise  
86 sources, chiefly the effects of seafloor currents (Webb et al., 1991; Doran & Laske, 2016). The  
87 pressure perturbations associated with infragravity waves have a frequency-dependent decay with  
88 depth in the water column. As a result, the maximum frequency at which seafloor compliance  
89 affects BBOBSs decreases in deeper water (Crawford & Webb, 2000; Bell et al., 2015), and the  
90 minimum frequency extends beyond the low-frequency end of the sensitivity of modern  
91 instruments (Figure 2).

92 Bottom-current noise is a consequence of seafloor currents directly buffeting the instrument  
93 (Webb, 1998; Collins et al., 2001). It affects the entire seismic band (Webb, 1998), but is strongest  
94 at frequencies  $< 0.1$  Hz, and leads to substantially higher noise levels on the horizontal  
95 components. In the case that the sensor is not perfectly level, this bottom-current noise can also  
96 couple into the vertical component, resulting in tilt noise (Crawford & Webb, 2000). These follow  
97 a power-law increase with decreasing frequency. Current noise is analogous to wind-driven noise

98 in land stations, but generally much stronger. On horizontal components, this noise largely  
99 eliminates the traditional low-noise observational band at frequencies below the microseism;  
100 analyses of earthquake signals on these components are generally limited to high-amplitude  
101 recordings (*i.e.*, large magnitude and/or nearby events).

## 102 **2.2 Noise Corrections**

103 Tilt and compliance noise imply predictable relationships between the vertical and horizontal, and  
104 vertical and pressure time series, respectively, at an individual BBOBS. Transfer functions  
105 quantify the admittance, coherence, and phase relationships (in frequency space) between these  
106 components (see Crawford & Webb, 2000; Bell et al., 2015). Vertical seismic components can be  
107 corrected for tilt and compliance noise using the appropriate transfer functions (Crawford & Webb,  
108 2000), leading to a reduction of noise levels on this component. The approach relies on the  
109 statistical property of signal stationarity: temporally consistent transfer functions can be obtained  
110 by averaging frequency-domain relationships between components across multiple time windows.  
111 Typical approaches for calculating transfer functions include averaging over long time periods,  
112 such that transient signals occupy a relatively small percentage of time windows (Yang et al.,  
113 2012), or removing transient signals prior to processing (Bell et al. 2015; Janiszewski et al., 2019).

114 Seafloor compliance is theoretically a time-invariant property describing the response of the local  
115 subsurface to infragravity waves, and its admittance spectrum can be inverted for shallow shear  
116 velocity structure (*e.g.*, Crawford et al., 1991; Ruan et al., 2014; Doran & Laske, 2019; Mosher et  
117 al., 2021). This implies that an effective compliance correction can be obtained from a small  
118 number of time-averaged transfer functions. However, tilt can vary with time, as instruments settle  
119 in soft sediment and the degree and azimuth of non-verticality change. Some instruments also

120 perform gyroscopic re-leveling cycles, with varying periodicity (Bell et al., 2015). To address these  
121 issues, tilt corrections derived from shorter-duration (*e.g.*, daily) transfer functions may prove  
122 more effective (*e.g.*, Bell et al., 2015).

123

### 124 **3. Methods**

#### 125 **3.1 Inclusion Criteria**

126 Our study includes BBOBSs deployed as part of experiments facilitated by OBSIP or OBSIC from  
127 2005 to the present. Each BBOBS in our dataset satisfies the following criteria. (1) It contains a 3-  
128 component, wideband or broadband seismometer (*i.e.*, with flat instrument response between  
129  $\sim 0.01$  and  $\sim 10$  Hz). We restrict our analysis to BBOBS designs with seismometers that are still  
130 actively used in the OBSIC fleet, which includes Guralp CMG-3T (CMG-3T), Nanometrics  
131 Trillium Compact (T-Compact), and Nanometrics Trillium 240 (T-240) instruments. (2) It includes  
132 a wide-band pressure sensor: either a differential pressure gauge (DPG) or an absolute pressure  
133 gauge (APG). (3) All four components of the BBOBS recorded data at a sample rate of at least 5  
134 samples-per-second (sps). Our study does not constitute a quantification of overall data quality;  
135 we do not account for station dropouts, broken channels, or instrument return rate. We focus on  
136 data that are considered “good” to offer an analysis of noise properties that are representative of  
137 normal BBOBS operations.

#### 138 **3.2 Data Selection and Processing**

139 To investigate the relationship of the noise characteristics of BBOBSs with deployment and  
140 instrument properties, we select a subset of data at each station from which to calculate power

141 spectra, cross-component coherence, admittance, and phase spectra, which make up the transfer  
142 functions used for noise corrections (Bell et al., 2015; Crawford & Webb, 2000). We examine 25  
143 days of data that are not significantly contaminated by earthquake signals, instrument glitches, or  
144 other transient signals at each BBOBS using the ATaCR code package (Janiszewski et al., 2019).  
145 These days are randomly distributed throughout each deployment to average across any long term  
146 drift, instrument releveling, or seasonal variability (Bell et al., 2015; Stähler et al., 2016). For all  
147 seismometer and DPG channels, we remove the instrument response using a high pass filter with  
148 a corner frequency of 1000 s. The response is not removed from APG channels, but we filter the  
149 data using the same procedure. All data are downsampled to 5 sps using an anti-alias filter with a  
150 corner frequency of 1.25 Hz.

151 We window each day of data into sixteen, 7200-second segments, overlapping by 30%, and apply  
152 a flat-Hanning taper to the windows. We calculate the auto- and cross-power spectral density  
153 functions from the finite Fourier transforms of the time series (Bell et al., 2015; Bendat & Piersol,  
154 2010) for each of the 16 windows. Any windows that contain transient signals identified via quality  
155 control procedures (see Janiszewski et al., 2019 for details) are discarded; if more than 6 windows  
156 are discarded, the entire day is rejected and not counted towards the 25-day sample. The windows  
157 are subsequently averaged to calculate spectral density functions for each day of data.

158 We then calculate deployment-average spectral functions for each station by averaging over all  
159 windows. A second quality-control step discards individual days that significantly (at 95%  
160 confidence level) increase the standard deviation of the noise properties (Janiszewski et al., 2019).  
161 This avoids the inclusion of days that are dominated by anomalous signals unrepresentative of  
162 normal station noise. While this processing procedure may not capture the full range of variability  
163 and discards malfunctioning data segments, it is appropriate for examining systematic trends at

164 functioning BBOBSs. For analysis and comparison, we take full-octave averages of the spectra in  
165  $\frac{1}{8}$  octave intervals following the procedure of McNamara & Buland (2004). We visually inspect  
166 all averaged spectra and discard any that contain data dropouts, flatlined or obviously non-  
167 functioning instruments, or instruments where the secondary microseism peak was not visible  
168 (*e.g.*, anomalously high noise floor). This results in average spectra for vertical ( $Z$ ), horizontal ( $H1$ ,  
169  $H2$ , or collectively  $H$ ), and pressure ( $P$ ) components at each BBOBS, as well as average cross-  
170 component coherence, admittance, and phase functions.

171 Lastly, we use the computed transfer functions to estimate average tilt- and compliance-corrected  
172  $Z$  spectra for each BBOBS (Crawford & Webb, 2000). In our discussion, we use  $Z$ -corrected to  
173 refer to the  $Z$  component where both tilt and compliance noise have been removed. To compute  
174 the  $Z$ -corrected spectra, all four components of a BBOBS must pass the aforementioned quality  
175 control procedures. After quality controls, we are left with data from 551 BBOBS with waveforms  
176 archived in the IRIS DMC that had at least one component, and 404 that have all four components.  
177 This includes instruments from 18 seismic experiments, with deployment years ranging from 2005  
178 – 2019 (Table S1). In calculating the transfer functions, we also calculate the cross spectral  
179 properties: the coherence, admittance, and phase (Bell et al., 2015). Of these, coherence is the  
180 simplest to interpret since it varies between zero and one and does not reflect discrepancies in  
181 instrument gain or polarity. High coherence between the vertical and pressure components at long  
182 periods indicates the presence of compliance noise, and high coherence between the vertical and  
183 horizontal components indicates tilt noise (Crawford & Webb, 2000). High vertical-pressure  
184 coherence also occurs within the microseism band, particularly near the secondary microseism.  
185 We analyze the coherences in order to constrain variability in the properties of tilt and compliance  
186 noise on the BBOBS.

187 The approach of Crawford & Webb (2000) scales the transfer function with the coherence; that is,  
188 a larger noise removal in the data will occur in locations with higher coherence  
189 values. Traditionally, tilt- and compliance- corrections are only calculated at frequencies  $< 0.1$   
190 Hz; another approach is to only calculate corrections in frequency ranges where the coherence  
191 between components is above a cutoff value (Bell et al., 2015; Tian & Ritzwoller, 2017). High  
192 coherences are also often observed in the secondary microseism band, and this transfer function  
193 correction approach has been extended to these higher frequencies with success (*e.g.*, Bowden et  
194 al., 2016). Since our investigations only rely on a systematic estimate of noise reduction after  
195 corrections, we calculate the corrections following Crawford & Webb (2000) across the entire  
196 frequency band, rather than exclude lower-coherence frequency ranges altogether. This may lead  
197 to a slightly higher estimate of noise reduction at frequencies with low coherence, but the effect  
198 should be minimal. In addition, users of BBOBS data should be cautious of potential distortion of  
199 time series or amplitude errors when applying corrections systematically over the entire frequency  
200 band. A more targeted removal approach may be required for some use cases.

### 201 **3.3 Metadata Compilation**

202 To compare the noise with instrument and deployment properties, we compile metadata  
203 information for all stations in our dataset. Instrument design was obtained from the IRIS DMC,  
204 and verified through review of individual cruise reports. Instruments are designed by one of the  
205 three centers that comprised OBSIP – the Lamont Doherty Earth Observatory (LDEO), Scripps  
206 Institute of Oceanography (SIO), and Woods Hole Oceanographic Institution (WHOI). We classify  
207 each instrument among eight unique designs according to differences in the seismometer, pressure  
208 gauge, or the overall package in which the instruments are housed. We do not attempt to distinguish  
209 between upgrades in datalogger versions within a given instrument design. For all designs, we

210 utilize the instrument responses archived with the data at the IRIS DMC. Our categorizations  
211 mirror those given in the OBSIP Final Report (Aderhold et al, 2019); however, we additionally  
212 distinguish between LDEO broadband instruments that were deployed with either a DPG or an  
213 APG. We also include information about the geographic environment of the instruments, which  
214 includes ocean basins and marginal seas as defined by the International Hydrographic  
215 Organization (1953). The deployments were predominantly located in the Pacific Ocean and its  
216 marginal seas, but also include the Atlantic Ocean and Lake Malawi. Details for our instrument  
217 design categorization are given in Table 1, and the frequency distributions of these parameters are  
218 shown in Figure S1. We also record the experiment in which each BBOBS was deployed.

219 We query the water depth of each station from the IRIS DMC; these values are reported by cruise  
220 logs, typically from sonar readings at the deployment location or through acoustic ranging (Russell  
221 et al., 2019), and are typically accurate to  $\sim 10$ m. Where possible, we determine the following  
222 geographic properties for each instrument using global compilations: the distance to the nearest  
223 land, the distance to the nearest tectonic plate boundary, the age of the underlying oceanic crust,  
224 the sediment thickness beneath each BBOBS, and the mean annual surface current velocity. To  
225 determine the distance to coastline, we calculate the distance to the nearest major landmass from  
226 each station using the dataset of Lee et al. (2018). This parameterization ignores small islands, as  
227 we are most interested in the relationship between noise sources that may arise from ocean-shelf  
228 interactions. We calculate the distance to the nearest plate boundary using the compilation of Bird  
229 (2003).

230 We estimate crustal age using the 2-arc-minute resolution seafloor age map from Müller et al.  
231 (2008), which is primarily constrained by prominent marine magnetic anomalies. We use the  
232 GlobSed model (Straume et al., 2019), a 5-arc-minute total sediment thickness grid for the world's

233 oceans and marginal seas, to estimate sediment thickness. We compute mean annual surface  
234 current velocity estimates using the near-surface velocity climatology data from the Global Drifter  
235 Program (Laurindo et al., 2017). For all three datasets, we estimate the variable at the BBOBS by  
236 extracting the geographic grid point that overlaps with the site location. If this did not exist, no  
237 value was assigned. In total, we compile and examine 11 metadata parameters at each BBOBS:  
238 “*Experiment*”, “*Instrument Design*”, “*Seismometer*”, “*Pressure Gauge*”, “*Environment*”,  
239 “*Water Depth*”, “*Distance from Land*”, “*Distance to Plate Boundary*”, “*Surface Current*”,  
240 “*Sediment Thickness*”, and “*Crustal Age*” (Table S2). The distributions for these parameters are  
241 shown in Figure S2.

242 There are limitations to our sampling of metadata in this analysis. Many investigated parameters  
243 are not evenly distributed. For example, the maximum value for distance to the coastline is 4020  
244 km, but > 75% of stations have values less than 1000 km. Some parameters do not have available  
245 values for all stations. For example, oceanic crustal age estimates do not exist for stations located  
246 on the continental shelves, on submerged Zealandia continental crust, and for lacustrine stations.

### 247 **3.4 Spectral Angle Calculation and Analysis**

248 A primary goal of this study is to determine the properties (*i.e.*, metadata) of a BBOBS that  
249 determine its noise characteristics. To this end, we divide the dataset of station spectra into  
250 subgroups defined by metadata parameter(s), and then quantify the similarity of spectra within  
251 each subgroup. If a certain parameter is highly determinative of noise, then spectra within each  
252 subgroup defined by that parameter should be similar to each other, but quite distinct from spectra  
253 in the other subgroups.

254 We use the “spectral angle” to quantify the (dis)similarity of spectra. This metric accounts for  
 255 differences in the shape, but not absolute amplitudes, of stations’ spectra (*e.g.*, Sohn & Rebello,  
 256 2002; Wan et al., 2002). We are primarily interested in variations in sources of noise (*e.g.*, changes  
 257 in the frequency distribution and extrema of different noise peaks and troughs), and not in an  
 258 average noise-level metric. The spectral angle is better suited to this than other metrics we tested  
 259 (*e.g.*, the Euclidean distance) that are overly sensitive to absolute amplitude. Spectral angle is also  
 260 diagnostic of differences in noise floor between instruments, due to differences in curvature of the  
 261 spectra between high noise peaks.

262 For a given pair of spectra,  $s_i$  and  $s_j$ , the spectral angle is computed in log-frequency space as

$$\theta_{i,j} = \cos^{-1} \left( \frac{s_i \cdot s_j}{|s_i| |s_j|} \right), \quad 1$$

263 We assign a penalty to each individual spectrum, defined as the root-mean-square of its spectral  
 264 angle with all other spectra in its subgroup:

$$p_i = \sqrt{\frac{1}{n-1} \sum_{j \neq i} \theta_{i,j}^2}. \quad 2$$

265 where  $n$  is the number of spectra in the subgroup. We then calculate the summed penalty for each  
 266 subgroup, describing mean spectral similarity, as the sum of the individual stations’ penalties:

$$P = \sum_i^n p_i. \quad 3$$

267 Finally, the overall penalty function for a given subgrouping scheme is the sum across all  
 268 subgroups’ penalties.

269 Since the effect of noise is expected to differ between the  $Z$ ,  $H1$  and  $H2$ , and  $Z$ -corrected  
270 components, we examine each of these independently. For consistency, we only analyze the 404  
271 instruments with four components that passed quality control. Since the BBOBSs are randomly  
272 oriented on the seafloor, we treat the  $H1$  and  $H2$  components as two representations of the  
273 horizontal noise, giving 808 spectra for this  $H$  component.

274 To start, we calculate three total penalties for the entire group of BBOBSs described above (for  $Z$ ,  
275  $Z$ -corrected, and  $H$  components). This yields a baseline measure of spectral dissimilarity amongst  
276 all stations in the dataset. We then systematically divide the dataset into subgroups of stations  
277 defined by each metadata parameter. For example, we use the “seismometer” metadata parameter  
278 to construct three subgroups of noise spectra defined by the parameter’s three categorical  
279 subdivisions: T-240s, CMG-3Ts, and T-Compacts. We calculate a penalty for each of the  $k$   
280 subgroups,  $P_k$ , as above, and a total penalty,  $P$ , as the summation of the three subgroup penalties.  
281 This is done for each seismic component  $Z$ ,  $H$ , and  $Z$ -corrected, in turn. In general,  $N$  in each  
282 subgroup varies with our choice of metadata parameter, as discussed below. Subgroups with zero  
283 or one station are excluded from the penalty calculation. The larger the reduction in overall penalty  
284 function when stations are subdivided according to a given metadata parameter, the more closely  
285 linked that parameter is to noise spectral shape.

286 The metadata parameters we use to subdivide the spectra fall into one of three types: (1)  
287 categorical, (2) numerical, (3) incomplete-numerical. For categorical parameters, we use one  
288 subgroup for each category. For numerical parameters, we utilize two subgroups, separated by a  
289 cutoff value. We determine the optimal threshold value by grid searching to obtain the cutoff that  
290 yields the two most internally similar subgroups. Finally, for the numerical variables that lack  
291 some data entries (*i.e.*, semi-numerical; Figure S2b, d,f), we place stations lacking numerical

292 values into a separate subgroup, and use the grid-search approach for numerical parameters for the  
293 remaining stations, yielding three subgroups.

294 Since multiple parameters influence the noise spectra, the subgrouping scheme uses a hierarchical  
295 framework. First, we perform the above analysis for each individual metadata parameter. We refer  
296 to such single-parameter subgroups as a “1-layer” analysis. For parameters that result in relatively  
297 high levels of penalty reduction, we then test the effect of producing additional subgroups by  
298 repeating this procedure two times, resulting in a “3-layer” analysis. In all cases, we evaluate  
299 success by computing the penalty reduction value, which compares the summed penalty to the  
300 baseline penalty.

301

## 302 **4. Results**

### 303 **4.1 Average Noise Spectra**

304 We present average power spectra for each BBOBS seismometer component in Figure 3. The  $Z$   
305 component data are, on average, between the New High and Low Noise Model (NHNM and  
306 NLNM; Peterson, 1993). As expected, the  $H$  components have higher values, above or near the  
307 NHNM at all frequencies. At frequencies lower than  $\sim 0.1$  Hz, the  $H$  components are on average  
308  $\sim 20 - 35$  dB noisier than the vertical components, likely due to bottom-current noise (Webb, 1998).  
309 The difference between components is less pronounced at shorter periods. Both the secondary and  
310 primary microseism are observed as clear peaks at  $\sim 0.14$  and  $\sim 0.07$  Hz, respectively, where the  
311 secondary peak is on average higher than the primary. However, the greatest variability between  
312 the spectra is observed in the primary microseism band on both the  $Z$  and  $H$  components. At  
313 frequencies  $< 0.05$  Hz, the infragravity signal manifests as an additional peak on a subset of the

314 vertical spectra. This peak is not observed on the horizontal components since it is drowned out  
315 by the stronger bottom-current noise. We also examined the pressure spectra; however, significant  
316 variability between experiments suggests a possible instrumentation or gain error for subsets of  
317 BBOBSs. Pressure gauge response functions can be prone to calibration error, although the cited  
318 variability is typically less than the order of magnitude observed here (Yang et al., 2012; Doran et  
319 al., 2019). At least one of the apparently anomalous pressure spectra is related to an error in the  
320 AACSE data that has since been reported and resolved in the IRIS DMC (Figure S5, S6). We note,  
321 however, that gain errors do not affect our ability to perform compliance removal, or interpret  
322 coherence or phase information between the  $Z$  and  $P$  components.

323 As predicted, for the  $Z$ -corrected components, we observe a reduction in noise across all  
324 frequencies after the transfer function corrections were applied. On average, the corrected spectra  
325 are  $\sim 5$ -10 dB quieter than the original, but reductions as large as  $\sim 40$  dB are observed. Maximal  
326 noise reduction is observed at  $\sim 0.01$ ,  $0.07$ , and  $0.14$  Hz, corresponding to tilt and compliance, the  
327 primary microseism, and the secondary microseism, respectively (Figure 3d). We test the order of  
328 corrections, comparing the final spectra when compliance is removed after tilt noise versus when  
329 tilt noise is removed after compliance noise. To first order, no difference is observed, and for the  
330 remaining analyses,  $Z$ -corrected spectra are calculated by first removing tilt and then compliance  
331 noise. The spectra for the seismic and pressure components grouped by experiment are shown on  
332 Figures S3-S7.

## 333 **4.2 Average Coherences**

334 For all BBOBS, we present the coherences between each horizontal and the vertical component,  
335  $H1-Z$  and  $H2-Z$ , and the coherence between the pressure and vertical components,  $P-Z$  (Figure 4a-

336 c). On the *H1-Z* and *H2-Z* pairs, we clearly observe high coherence values on a subset of the  
337 instruments at frequencies  $< 0.1$  Hz with no clear dependence on water depth. This is consistent  
338 with tilt noise on the *Z* component. We observe high *P-Z* coherence with a water-depth-dependent  
339 high frequency limit that agrees with the predicted cutoff for infragravity waves. This is consistent  
340 with compliance noise on the *Z* component. We also observe a region of high *P-Z* coherence at  
341 frequencies at and just above  $\sim 0.14$  Hz, consistent with the secondary microseism. Another, more  
342 moderate, peak observed at  $\sim 0.07$  Hz is consistent with the primary microseism. We recalculate  
343 the *H1-Z* and *H2-Z* coherences after compliance noise removal, and the *P-Z* coherence after tilt  
344 noise removal from the *Z* component (Figure 4d-f). Since tilt is typically assumed to be the larger  
345 noise source, we expect its removal should result in a more visible compliance signal, and an  
346 increase in *P-Z* coherence. As anticipated, we observe that the *P-Z* coherence tends to increase  
347 after the removal of tilt noise at frequencies below the infragravity cutoff. However, we also  
348 observe an increase in the *H1-Z* and *H2-Z* coherences for some instruments when we first remove  
349 the compliance noise. This suggests that the two noise sources may have similar amplitudes at  
350 some stations, in contrast to the assumption that tilt noise is typically a much larger noise source  
351 (Bell et al., 2015). This is discussed further in Section 5.1.

352 Coherences between horizontal and pressure component pairs are typically not investigated for  
353 BBOBS noise characterization and removal, as incoherence is predicted (Crawford & Webb,  
354 2000). We mostly observed *H1-P* and *H2-P* incoherence in our compiled dataset, with the  
355 exception of the shallowest stations, where coherences were  $> 0.5$  near 0.1 Hz for both these  
356 component pairs (Figure S8). Tilt and compliance corrections for the *Z* component can still be used  
357 in this frequency band at these shallow water instruments (An et al., 2020; Webb & Crawford,  
358 2010), as long as this coherence is accounted for.

### 359 **4.3 Determinants of Station Noise**

360 Our systematic calculation of the (dis)similarity between station noise spectra after subdividing  
361 stations by each metadata parameter yields quantitative estimates (in terms of “penalty”, the  
362 measure of dissimilarity) of the relative importance of these features in controlling noise  
363 properties. A higher penalty reduction suggests that a given parameter is a better predictor of  
364 spectral characteristics. As a baseline, the mean penalty per trace for the *Z*, *Z-corrected*, and *H*  
365 components is 4.94, 4.27, and 4.46, respectively. We report “penalty reduction” as a percent  
366 deviation from these values.

#### 367 **4.3.1 1-Layer Analysis**

368 We first discuss results for our 1-layer analysis (Figure 5a). The largest penalty reduction is  
369 obtained when grouping stations by “*Experiment*” (a mean penalty reduction of 17.4%, when  
370 averaging over the *Z*, *H*, and *Z-corrected* components). Next is “*Instrument Design*”, which  
371 produced an average penalty reduction of 15.6%, and yields the largest penalty reduction for the  
372 *Z* component (19.8%). However, neither of these parameters directly illuminate the physical  
373 processes controlling seismic noise, as they strongly co-vary with other metadata. For example, 14  
374 (out of 18) experiments involve only one type of instrument design and seismometer (Table S1).  
375 Experiments typically occupy a small footprint (Figure 1), so intra-experiment variation in the  
376 environment is also limited. Similarly, “*Instrument Design*” co-varies with “*Seismometer*”,  
377 “*Water Depth*” (e.g., TRM designs are only deployed in < 1000 m), and “*Pressure Gauge*”.  
378 Nonetheless, the significant penalty reduction under these two parameters demonstrates that  
379 experiment and instrument parameters collectively have substantial impact on noise  
380 characteristics, reinforcing the need for careful deployment planning.

381 The “*Seismometer*” parameter has the next greatest influence on the noise spectra (Figure 5a),  
382 reducing the total penalty by ~10% for both *Z* and *H* components. For the *Z* component, the  
383 “*Seismometer*” subgroup spectra show different signatures of classic BBOBS noise (Figure 6).  
384 The CMG-3Ts display a power law (linear in log-log space) amplitude increase at frequencies  
385 below 0.1 Hz, characteristic of tilt noise. By contrast, T-240s and T-Compact subgroups have more  
386 spectral curvature and multiple inflection points at the same low frequencies, which is  
387 characteristic of compliance noise (Bell et al., 2015). The *H* spectra provide further insight (Figure  
388 6c). All three seismometer subgroups show bottom current noise (an ~18dB/Hz-decade linear  
389 increase in log-log space at low frequencies). However, the subgroups of *H* spectra are clearly  
390 distinguished by their noise notch relative to their secondary microseism peak. The CMG-3Ts have  
391 the least distinct noise notch (just ~15 dB below the secondary microseism), and are the noisiest  
392 instruments, on average, at long periods, especially for horizontals. T-Compacts have an  
393 intermediate noise notch (~20dB below the secondary microseism), partially influenced by the  
394 substantial primary microseism associated with shallow water shielded instruments. T-240s have  
395 a noise notch ~30dB below the secondary microseism, and have the quietest horizontals at long  
396 periods. The average spectra for the *Z-corrected* components for the CMG-3Ts and the T-  
397 Compacts are nearly identical; however, long period noise on the T-240s remains ~ 20 dB quieter  
398 (Figure 6e).

399 In the higher-frequency band (0.1-1 Hz) dominated by the secondary microseism, the T-240s peak  
400 ~10 dB lower than the other sensors, on all components. This observation is somewhat surprising,  
401 as the secondary microseism peak is ubiquitous in all ocean environments and does not vary  
402 dramatically even with depth (Figure 6). We have considered the possibility that incorrect  
403 instrument gain(s) may contribute to this apparent offset (Doran & Laske, 2019), and tested the

404 potential impact of this error on our results by adding a constant 10 dB to spectra that have an  
405 anomalously low secondary microseism peak. The overriding conclusion of our tests is that the  
406 quantitative subgroupings and associated interpretations are not sensitive to gain uncertainty of  
407 this magnitude, but this may slightly impact the absolute amplitude differences observed between  
408 different seismometers. We discuss these metadata uncertainties further below.

409 “*Water Depth*” has a similar level of influence on the noise spectra as the “*Seismometer*”  
410 parameter. These two categories are also parsimonious in their numbers of subgroups (2 and 3,  
411 respectively). “*Water Depth*” is particularly deterministic for *Z* component noise, with a penalty  
412 reduction of 14.1%, compared to 4.8% for *H*, and 4.0% for *Z-corrected* (Figure 5). Unlike for the  
413 “*Seismometer*” parameter, “*Water Depth*” subgroups were determined by grid search. The cutoff  
414 depths separating these subgroups are relatively shallow for all three components, between ~ 200  
415 - 500 m depth. This cutoff separates shallow versus deep noise environments, reflecting the  
416 distinctive signal of shallow water infragravity waves in the 0.04 - 0.1 Hz band, likely due to direct  
417 wave loading that overlaps with primary microseism frequencies, observed on all components  
418 (Webb & Crawford, 2010; An et al., 2021). Our depth resolution is limited by the depths at which  
419 BBOBS were deployed; thus, we report the maximum and minimum depths of the shallow and  
420 deep subgroups, respectively (Figure 6). While this shallow water signal is present on all  
421 components, compliance noise continues to influence the *Z* component at lower frequencies (<  
422 0.03 Hz), following the characteristic depth dependence of infragravity waves (Figure 4). This  
423 explains the greater penalty reduction for the *Z* component, relative to the *H* and *Z-corrected*  
424 components. A more detailed investigation of the grid search for the *Z* component indicates two  
425 depths where there are sharp changes in the penalty function: (1) the 200-500 m cutoff discussed  
426 above, and (2) an inflection at ~ 2600 m water depth (Figure S9). This may reflect the variable

427 frequencies of the compliance noise, and suggests broadly that categorization into shallow, mid-  
428 depths, and deep-water regimes is useful for predicting overall BBOBS noise levels for a  
429 deployment.

430 The remaining parameters are less useful singular determinants of noise characteristics, as they  
431 mostly have smaller penalty reductions ( $< 5\%$  for each component). The penalty reductions for the  
432  $Z$  component using “*Crustal Age*” and “*Distance from Land*” are somewhat higher; however,  
433 covariance with “*Water Depth*” likely explains this observation. Covariance amongst parameters  
434 is exemplified by the apparently perverse observation that “*Pressure Gauge*” has some apparent  
435 predictive power for the noise characteristics of the seismic components.

#### 436 **4.3.2 3-Layer Analysis**

437 For the 1-layer analysis, the power of any single parameter to predict noise characteristics is  
438 limited, with 6.5% of variation explained, on average. This low value indicates the multifactorial  
439 controls on BBOBS noise. Therefore, we expand our analysis up to three layers to determine which  
440 combinations of parameters yield subgroupings with the most similar spectral characteristics. This  
441 also helps us test which parameters (if any) have a secondary role in regulating noise variability.

442 We start with a new baseline 2-layer analysis. We group stations by “*Seismometer*” (three  
443 subgroups), and then by “*Water Depth*” (two subgroups), yielding a total of six subgroups (Figure  
444 7). These two parameters were chosen on the basis of their high penalty reduction (Section 4.3.1),  
445 their parsimonious subgroupings, their relative lack of covariance with simpler parameters, and their  
446 clear physical relationships with noise, facilitating interpretation. The 2-layer analysis yielded a  
447 20% average penalty reduction (horizontal lines in Figure 5b), with 27.4% for the  $Z$  component,  
448 19.2% for the  $H$  components, and 13.5% for  $Z$ -corrected. As above, the larger penalty reduction

449 for the  $Z$  component reflects its sensitivity to both compliance and tilt. This is further supported  
450 by the relatively low penalty reduction (*i.e.*, higher inter-group similarity) for the corrected vertical  
451 component ( $Z$ -corrected), which nominally has these effects removed. On the other hand, the fact  
452 that the  $Z$ -corrected component still has non-zero penalty reduction demonstrates that factors other  
453 than tilt and compliance influence noise characteristics, or that these corrections do not work  
454 perfectly. More sophisticated methods of noise removal, such as algorithms that account for  
455 temporal variability of the transfer functions, or iterative removal processes (Bell et al., 2015; Tian  
456 & Ritzwoller, 2017), might drive this “ $Z$ -corrected” penalty reduction down further.

457 We conducted independent grid searches for water depth cutoff values in each seismometer  
458 subgroup. The optimal depth cutoff for the T-Compact subgroup was between 354 - 430 m, very  
459 similar to the shallow cutoff depths discussed above (Figure 7). T-Compact seismometers are used  
460 in the majority of the shielded BBOBS instruments designed for shallow water deployments <  
461 1000 m. By contrast, the T-240 and CMG-3T subgroups have deeper apparent cutoffs, between  
462 2564 - 2687 m and 2785 - 2822 m, respectively. This is similar to the secondary mid-depth cutoff  
463 noted above. Since these subgroups do not include most of the shallow-water BBOBS, they do not  
464 include characteristically very shallow (< 500m depth) spectra, and so their intra-group grid  
465 searches find what we believe to be an important local minimum in penalty at ~2600m. Close  
466 investigation of the grid search results for the Trillium Compacts (Figure S9) reveals the same  
467 local inflection at ~2600 m cutoff depth.

468 Finally, we perform a 3-layer analysis. Each of the six subgroups from the 2-layer “*Seismometer*”  
469 and “*Water Depth*” analysis is further subdivided according to each of the remaining parameters  
470 (*e.g.*, Figure 7, Figures S10-12), and penalty reduction is measured. Aside from “*Experiment*” (see  
471 caveats below), “*Environment*” yields the highest 3-layer penalty reduction (25.8% mean

472 reduction, or 5.8% above the 2-layer baseline, for all components). This is likely explained by  
473 differences in the frequency distribution of the secondary microseism across ocean basins  
474 (Babcock et al., 1994; Yang et al., 2012). “*Distance from Land*” is the numerical parameter that  
475 provided the highest penalty reduction, approximately 5.5% above the 2-layer baseline for all  
476 components. While not included in the main analysis, we test the effects of including “*Water*  
477 *Depth*” again in the third layer. It results in similar penalty reductions as “*Distance from Land*”,  
478 which is likely a consequence of covariance between water depth and distance from land.

479 Other numerical parameters in the 3-layer analysis yield lower penalty reduction, but all improve  
480 upon the 2-layer baseline by  $> \sim 3\%$ . The fact that several, rather than any one, of these parameters  
481 control noise characteristics is highlighted by the particularly high penalty reduction (9.9% above  
482 the 2-layer baseline across components) for “*Experiment*” in the 3-layer analysis. As discussed  
483 above, because individual experiments do not usually span large portions of metadata space, this  
484 parameter effectively combines many other parameters. Thus, it functions as a heuristic for the  
485 extent to which station noise is determined by all the station metadata collectively. One way of  
486 looking at this is as a lower bound for the aspects of station noise that are deterministically based  
487 on instrument type and location, with the remainder of variability owing to random site  
488 characteristics and spatiotemporally varying sea state. A final point of note is that moving from  
489 the 1- to 3-layer hierarchy, “*Instrument Design*” switches from providing the second highest to  
490 the second lowest penalty reduction. Of course, instrument design is strongly related to  
491 seismometer type and - via shielding - water depth. However, this result indicates that having  
492 controlled for covariance with those factors, the design of the instrument is not itself highly  
493 impactful on noise characteristics.

#### 494 **4.3.3 Significance of Observations**

495 Finally, we consider the possibility that we observe penalty reduction simply by chance. We test  
496 the significance of the observed penalty reduction by computing the penalty reduction for 10,000  
497 random subgroups of spectra. In each iteration, we keep the same number of subgroups and number  
498 of stations in each subgroup as in the true groupings analyzed above. However, instead of assigning  
499 stations into each subgroup according to their metadata parameters, we assign them randomly.  
500 Taking the example of the 3-layer analysis of “*Seismometer*”, “*Water Depth*”, and “*Distance from*  
501 *Land*”, the random assignments yield a mean penalty per trace of 4.93 (the standard deviation is  
502 0.02, and 95% of the random iterations yield a value above 4.89). This is only a 0.88% penalty  
503 reduction from the baseline (of 4.94), compared with the 32.7% penalty reduction when the data  
504 are grouped according to real parameters (which is 88 standard deviations removed from the  
505 baseline). This analysis establishes the strong significance of the relationships between metadata  
506 and station noise characteristics spectra (Figure S13).

507

#### 508 **4.4 Frequency and Amplitude Variability**

509 While the previous analysis focuses on causes of inter-station variability in the noise spectra from  
510 0.001 - 1 Hz, most seismic applications of BBOBS data use a band-limited frequency range. We  
511 calculate mean noise levels in four commonly used frequency bands. We focus on the following  
512 bands: (1) 0.1 - 1 Hz, centered over the secondary microseism band; (2) 0.05 - 0.1 Hz, centered  
513 over the primary microseism band and traditional noise notch in BBOBS instruments, both of  
514 which are relevant for ambient-noise analyses (*e.g.*, Zha et al., 2013; Russell et al., 2019; Yang et  
515 al., 2020), teleseismic body-wave imaging (*e.g.*, Wolfe et al., 2009; Hawley et al., 2016; Bodmer  
516 et al., 2018; Eilon & Abers, 2017), scattered-wave imaging (*e.g.*, Leahy et al., 2010; Janiszewski  
517 & Abers, 2015; Rychert et al., 2018; Mark et al, 2021), and shear-wave splitting (*e.g.*, Collins et

518 al., 2012; Eilon et al., 2014; Bodmer et al., 2015; Lynner & Bodmer, 2017); (3) 0.01 - 0.05 Hz, the  
519 primary band for teleseismic long-period body- and surface-wave velocity and attenuation imaging  
520 (*e.g.*, Weeraratne et al., 2007; Laske et al., 2011; Jin et al., 2015; Cai et al., 2018; Janiszewski et  
521 al., 2019); (4) 0.005 - 0.01 Hz, of interest for very long-period surface wave (*e.g.*, Lin et al., 2016)  
522 and normal-mode (*e.g.*, Bécel et al., 2011) studies.

523 Based on the spectral angle analysis in the previous section, we plot noise levels as a function of  
524 water depth, parsed according to seismometer and instrument design (Figure 8). This analysis  
525 complements our spectral angle approach by comparing average absolute amplitudes in discrete  
526 frequency bands, rather than amplitude-agnostic spectral shape. Figure 8 illuminates several  
527 points that should be considered carefully during experiment design. First, there are no clear trends  
528 in noise level as a function of water depth, seismometer, or instrument type in the secondary  
529 microseism band (0.1 - 1 Hz). Most BBOBS tend to cluster near the NHNM (Peterson, 1993) in  
530 this range for both the vertical and horizontal components. Importantly, this is true even for the  
531 shallowest BBOBS, as this band is largely above the frequencies at which compliance noise is  
532 present. Some instruments display noise levels up to  $\sim 50$  dB quieter on average, but it is possible  
533 that instrument-gain uncertainty contributes to these outliers (Doran & Laske, 2019).

534 In the primary microseism band (0.05-0.1 Hz),  $Z$  and  $H$  component noise levels increase for  
535 shallow BBOBS, consistent with direct seafloor loading due to infragravity waves; this effect is  
536 reduced for the  $Z$ -corrected component, but relatively high noise levels at the shallowest depths  
537 persist even after corrections. For each component, shallow water instruments have the highest  
538 noise levels in this frequency range.

539 In the lower frequency bands ( $< 0.05$  Hz), the effect of compliance noise, in addition to tilt, is  
540 observed on the  $Z$  components, evidenced by the dependence on water depth. In contrast, the  $H$   
541 components as a whole do not show a clear dependence on water depth in these bands; this may  
542 reflect the effectiveness of instrument shielding in mitigating strong shallow seafloor currents,  
543 and/or the fact that seafloor currents are pervasive at all ocean depths. The CMG-3T seismometers  
544 show a stronger trend of decreasing noise levels with water depth relative to the other  
545 seismometers, indicating that this trend may depend on instrumentation; however, further analysis  
546 is needed to assess the significance of this observation. The compliance and tilt corrections are  
547 generally effective in these bands, and the  $Z$ -corrected noise levels are largely distributed between  
548 the NLNM and the NHNM.

549 On average, BBOBS containing T-240 seismometers have the quietest noise levels at all frequency  
550 bands, but the differences become more pronounced at lower frequencies. This observation holds  
551 even accounting for the possibility of gain errors in some T-240 deployments (Doran & Laske,  
552 2019). Importantly, this difference remains after tilt and compliance removal; that is, T-240s have,  
553 on average, the quietest  $Z$ -corrected components, with many deployments showing noise  
554 characteristics just above the NLNM. At frequencies lower than 0.1 Hz, BBOBS that contain a  
555 CMG-3T seismometer show higher noise levels than BBOBS with other sensors, particularly on  
556 the  $H$  components. The exaggeration of this effect at frequencies lower than 0.1 Hz is consistent  
557 with the presence of bottom current and tilt noise.

558

## 559 **5. Discussion**

560 Station metadata are strongly predictive of BBOBS noise characteristics. When stations are  
561 grouped by metadata parameters, there is substantially more similarity between spectra within  
562 those groups compared to the similarity averaged across the whole dataset (Figures 5 and 7). The  
563 water depth and seismometer type are the two most important factors that determine noise  
564 characteristics. The covariance between seismometer and instrument design complicates  
565 understanding the relative roles of these two parameters. However, consideration of these results  
566 may be useful during experiment design. For instance, if analysis relies upon 0.05 - 0.1 Hz period  
567 teleseismic *S-p* converted phases to evaluate mantle discontinuities, it is ill-advised to deploy in <  
568 500 m water depths, as these signals will likely be dominated by noise that persists after tilt and  
569 compliance corrections. Similarly, CMG-3Ts seem to be the most noise-prone seismometers  
570 across a range of environments. Here we further discuss sources of noise, implications for data  
571 quality and traditional noise corrections, limitations of our analysis, and potential next steps for  
572 the BBOBS community.

### 573 **5.1 Tilt and Compliance Noise**

574 We have already suggested that the strong link between noise characteristics, and the seismometer  
575 and water depth parameters is primarily driven by variations in the tilt and compliance effects,  
576 respectively. Here we investigate how coherence between components can illuminate the relative  
577 roles of these noise sources as a function of seismometer, water depth, and instrument design.  
578 Importantly, the coherence is insensitive to any gain errors. We also discuss how these may affect  
579 noise removal approaches.

580 Compliance noise is characterized by high average *P-Z* coherence from  $\sim 0.005$  Hz (due to our  
581 instrument response removal procedure) up to the theoretical infragravity frequency limit

582 (Crawford & Webb, 2000). Figure 4 shows very good agreement between the high-frequency limit  
583 of  $P$ - $Z$  coherence and the predicted cutoff frequency at water depths spanning the full range from  
584 0-6000 m. Unsurprisingly, water depth is a primary factor in determining a station's compliance  
585 noise signature. We estimate the presence of tilt noise from the maximum average coherence (in  
586 the range 0.005 - 0.035 Hz) between the  $H1$ - $Z$  and  $H2$ - $Z$  components. We follow the method of  
587 Bell et al. (2015), grid searching through horizontal component azimuths to find the orientation  
588 (theoretically the physical tilt direction) that gives the maximum coherence with the vertical ( $H_{\text{tilt}}$ ).

589 First, we observe that  $Z$ - $H_{\text{tilt}}$  coherence is higher on instruments with CMG-3T seismometers  
590 (Figure 9), consistent with their higher propensity for tilt noise. On average, the coherence is  $\sim$   
591 0.8, above the typical benchmark value used for tilt removal (Bell et al., 2015; Tian & Ritzwoller,  
592 2017). This higher tilt noise could arise from higher susceptibility of these instruments to current  
593 noise, and/or a tendency of these instruments to remain slightly out of level (*i.e.*, to have a  $Z$   
594 component that is not perfectly vertical). The analyses in Section 4.3 suggest that a combination  
595 of these effects may be important. While all seismometer types show similar low-frequency noise  
596 with a log-linear slope below 0.03 Hz on the  $H$  components, indicative of bottom current noise,  
597 amplitudes are systematically higher on CMG-3T seismometers (Figure 6), suggesting that these  
598 sensors are more strongly impacted directly by currents. In addition, only this seismometer shows  
599 this log-linear trend on the  $Z$  component (Figure 6), suggesting that it more commonly transfers  
600 current noise into  $Z$ -component tilt noise. Bell et al. (2015) also reported high tilt noise on these  
601 seismometers using data from just the first year of the Cascadia Initiative deployment, and  
602 suggested a tendency for their tilt direction to preferentially align with  $H1$ . Using our expanded  
603 dataset, we observe no systematic tilt direction (Figure 9). In contrast to this high tilt susceptibility,  
604 BBOBSs that use either the T-Compacts or T-240s have mean  $Z$ - $H_{\text{tilt}}$  coherences that are  $< 0.5$ ,

605 lower than typical benchmark values for useful tilt noise removal (Bell et al., 2015; Tian &  
606 Ritzwoller, 2017). The T-Compacts offer the next highest  $Z-H_{\text{tilt}}$  coherence after the CMG-3T  
607 seismometers; of these, average coherence for the TRM and AB shielded designs is particularly  
608 low, supporting the suggestion that shielding protects against horizontal noise contamination.  
609 However, unshielded BBOBSs using the T-240s also have comparably low values, suggesting that  
610 these seismometers may simply be less susceptible to tilt noise.

611 Tilt noise is typically assumed to be higher amplitude than compliance noise, but is not always  
612 present. Compliance noise is always present but may be masked by strong tilt noise (Crawford &  
613 Webb, 2000). It is therefore conventional to first remove tilt noise, which should lead to an increase  
614 in the  $P-Z$  coherence allowing for subsequent removal of the compliance noise. Our analysis  
615 (Figure 9) suggests this sequence of noise removal is particularly important for stations with CMG-  
616 3Ts. On the other hand, Tian & Ritzwoller (2017) find that both tilt and compliance noise interfere  
617 with each other (that is  $H1-Z$  and  $H2-Z$  coherence increases after compliance removal, and  $P-Z$   
618 coherence increases after tilt removal), consistent with relative similarity in their strength. They  
619 suggest that multiple iterations of corrections may be appropriate in such cases. To investigate this  
620 systematically, we compare the  $P-Z$  coherence before and after tilt correction (Figure 10a), and  
621  $H1-Z$  and  $H2-Z$  coherences before and after compliance correction (Figure 10b, c). For consistency  
622 between the two, we report the average coherence over the frequency range where compliance  
623 effects are present, which is inclusive of the range where tilt noise is expected for all stations.

624 As expected, the  $P-Z$  coherence increases for most instruments after tilt noise removal (Figure  
625 10a), validating conventional noise removal approaches (e.g., Wei et al., 2015; Accardo et al.,  
626 2017; Cai et al., 2018; Janiszewski et al., 2019). This test also reinforces the predominance of tilt  
627 effects on CMG-3T instruments: these seismometers have the largest gains in  $P-Z$  coherence, but

628 see essentially no change in compliance-corrected  $H1-Z$  or  $H2-Z$  coherences (Figure 10b). In  
629 contrast, for instruments with Trillium seismometers, we find increases in both the tilt-corrected  
630  $P-Z$  coherence and compliance-corrected  $H1-Z$  and  $H2-Z$  coherences. This suggests that at  
631 individual instruments either the two noise sources are similar in amplitude and interfere with one  
632 another, or that in some cases compliance removal may improve the ability to distinguish tilt noise  
633 on an instrument. A more detailed analysis of individual instruments is necessary to distinguish  
634 between these end member behaviors. In addition, recalculation of the tilt orientation after  
635 compliance removal, and testing of iterative noise removal methods may further help to determine  
636 properties and best practices related to the noise and its removal, but is beyond the scope of this  
637 study (Tian & Ritzwoller, 2017). Furthermore, whether this behavior remains stationary  
638 throughout the deployment of an instrument remains unclear.

639 Lastly, coherence actually decreases after noise removal at a subset of the TRM and AB  
640 instruments (Figure 10). These instruments have mostly high ( $> 0.5$ )  $P-Z$  coherences in the  
641 expected frequency range for compliance noise. We reiterate that these instruments also have high  
642  $H1-Z$  and  $H2-Z$  coherences in the primary microseism band (Figure 4); the lower ( $< 0.5$ .) values  
643 observed here stem from averaging over the entire compliance frequency band, which is wider  
644 than the microseism. For such a decrease in coherence to occur, the noise across all four  
645 components of the BBOBSs must be coherent, which is a property only observed on shallow water  
646 instruments. With the exception of one AR instrument that may be affected by an error, all  
647 instruments that have a decrease in coherence are deployed in less than 280 m water depth.

## 648 **5.2 Shallow Water Instruments**

649 Shallow water BBOBSs have demonstrably distinct noise characteristics (Figure 6; Webb &  
650 Crawford, 2010; An et al., 2021). Since this is one of the strongest defining characteristics of  
651 observed BBOBS noise, here we further investigate if these characteristics are present on all  
652 shallow water instruments. Given the set of water depths chosen as the optimal division between  
653 shallow and deep instrument noise characteristics (e.g., Section 4.3), we suggest  $< 500$  m depth as  
654 a conservative limit below which shallow water noise characteristics should be expected. These  
655 spectra typically contain a high amplitude peak on all components within the primary microseism  
656 band, extending to the predicted infragravity wave cutoff frequency. This peak is reduced, but not  
657 removed, by noise corrections (Figure 6). Within this band, pressure coherence with all seismic  
658 components of the BBOBSs is high (Figure 4; Figure S8) due to both vertical and horizontal  
659 loading of the seafloor from ocean waves directly above the instrument (Webb & Crawford, 2010).

660 Only three experiments in our dataset deployed instruments at depths  $< 500$  m: the Alaska-  
661 Aleutians Community Seismic Experiment (AACSE), located offshore Alaska; the Cascadia  
662 Initiative (CI), located offshore the northwest coast of the United States; and SEGMeNT, located  
663 in Lake Malawi in Africa. The former two share environmental similarities: the continental shelf  
664 adjacent to the Pacific Ocean basin. Their noise characteristics are also similar (Figure FS14); the  
665 majority of these shallow BBOBSs contain the expected high amplitude peak on vertical and  
666 horizontal components at  $\sim 0.07$  Hz. In contrast, this feature is much weaker at the Lake Malawi  
667 stations (Figure FS14), which record the primary ocean microseism in the far field. Lake Malawi  
668 stations instead manifest a strong noise peak at 0.3-1.6 Hz (Carchedi et al., 2022). This is likely  
669 due to differences in the characteristic wavelength of wind-driven waves in lacustrine versus  
670 oceanic settings. While microseisms are generated at lakes, they have distinctively higher  
671 frequencies than those generated in the oceans (Xu et al., 2017; Smalls et al., 2019) explaining the

672 strong 0.3-1.6 Hz peak (Carchedi et al., 2022). Lake infragravity waves are also present (Accardo  
673 et al, 2017), seen from 0.02-0.06 Hz (Figure S14).

674 Importantly, these differences may impact the application of different seismic analysis techniques  
675 on the data. For example, at Lake Malawi the separation of the ocean microseism from both the  
676 lake-generated microseism at higher frequency and lake infragravity waves at lower frequency  
677 allowed Accardo et al. (2017) to observe clear ambient noise cross-correlation signals between  
678 lake-bottom and land seismometers in the 0.04-0.125 Hz range, including at those instruments  
679 deployed at depths < 500 m. By contrast, ambient noise cross-correlations from shallow-water  
680 instruments in Cascadia had low signal-to-noise ratios at these frequencies (Janiszewski et al.,  
681 2019; Tian & Ritzwoller, 2017), due to local ocean-generated waves swamping the microseism  
682 signal.

### 683 **5.3 Limitations**

684 Although this study constitutes the largest systematic review of BBOBS noise characteristics  
685 conducted to date, there are important limitations to the dataset. Chief among these is that US  
686 BBOBS deployments using modern instrumentation have unevenly sampled large swaths of the  
687 metadata parameter space. For instance, there is more data from the Pacific Ocean than elsewhere,  
688 and a relative paucity of stations atop thick sediments or at great distance from coastlines (Figures  
689 S1, S2). A corollary to this uneven sampling is covariance in several station parameters, which  
690 makes it more challenging to tease apart the individual influences of, say, shielding versus shallow  
691 water on noise. Although we have attempted to pick apart the most important parameters  
692 controlling noise characteristics (Section 4.3), intrinsic covariation makes it impossible to separate  
693 parameters completely. This is most clearly seen from Figure 5 where the “*Experiment*” is the most

694 important parameter determining noise characteristics, simply because the small geographic  
695 footprint of most experiments means that other station parameters are alike within each  
696 experiment, and most experiments use a homogeneous instrument design. A small number of co-  
697 located pilot deployments of BBOBS with different seismometers or instrument designs in shallow  
698 and deep locations could test the robustness of the results presented here. We also suggest that co-  
699 deployments should be an essential aspect of testing new BBOBS designs whenever possible.

700 Our analysis is dependent on the accuracy of both data and metadata archived within the IRIS  
701 DMC, and one example of a co-located deployment suggests that errors may exist. The PLUME  
702 experiment (Doran & Laske, 2019) utilized an intermixed array of T-240 and CMG-3T sensors in  
703 relatively deep water. The individual seismometer spectra (Figures S3-S4) group into distinct  
704 clusters, with the T-240s offset to significantly lower power at all bands. Given the similar  
705 deployment environment for these instruments, the simplest interpretation for the offset is a gain  
706 error. Based on the secondary microseism peak (0.1-1 Hz), the CMG-3Ts are biased ~10 dB too  
707 high, or the T-240s are biased low. Doran & Laske (2019) analyzed this apparent bias and  
708 calculated station-specific gain corrections of x2 or x4 for the PLUME T-240 observations.

709 Here, we take a more general approach to specifically assess whether such issues could  
710 significantly impact our analyses. We estimate that the T-240 data are biased approximately 10 dB  
711 low. This is based on the observations that the average T-240 spectrum is lower than our full-  
712 dataset average in the secondary microseism band by approximately this amount (Figure 6), and  
713 that several older T-240 experiments have low noise levels relative to more recent experiments  
714 using those same instruments, including in the secondary microseism band (Figures S3-S4).

715 To test the impact of these gain uncertainties on our analysis, we collect all the T-240 spectra from  
716 these suspect experiments collected prior to 2011, and increase them by 10 dB. We then re-run  
717 the spectral-angle analysis, and compare the resulting groupings to those presented above (Section  
718 4.3). The dominant groupings are unchanged, as are the majority of the details of the spectral  
719 characteristics within each grouping. The weaker secondary microseism peak in the average T-  
720 240 spectrum is no longer present, but the T-240 spectrum at long periods remains lower than the  
721 other instruments, particularly for the *Z* and *Z-corrected* components. After this adjustment, the  
722 difference between the *H* component noise on the T-240 and T-Compact is minimal; however,  
723 both remain lower than the observed noise levels for the CMG-3Ts. This evaluation reassures us  
724 that our primary conclusions are robust in the face of metadata uncertainty of the scale suggested  
725 by Figures S3-S4.

726 Finally, the limited duration of standard OBS deployments ( $\leq 12$  months) means that our analysis  
727 is subject to the idiosyncrasies of experiment timing. As an example, the recording period for the  
728 HOBITSS experiment on the Hikurangi forearc largely overlapped the 2014-2016 El Niño event,  
729 confounding our ability to assess the normative noise characteristics of this particular margin. In  
730 this study of overall trends, we have chosen not to consider seasonal variability of noise, which  
731 can be substantial (*e.g.*, Grob *et al.*, 2011), and in addition we do not consider secular changes in  
732 noise with time (*cf.* Bell *et al.*, 2015). Further, this study uses only instruments from US-funded  
733 BBOBS deployments; many other designs exist that we have not included here. Their future  
734 inclusion would likely mitigate covariances between metadata parameters (particularly between  
735 seismometer and instrument design), and yield a wider geographic footprint.

736 As the marine geophysical community plans for long-term BBOBS observatories (Kohler *et al.*,  
737 2020), it would be worthwhile to invest resources in exploring the noise characteristics of these

738 under-sampled regions of metadata parameter space. The dataset presented in this study assists in  
739 framing noise domain gaps that future pilot experiments could fill.

740

## 741 **6. Conclusions**

742 We have computed representative noise spectra for 551 broadband BBOBS stations spanning 18  
743 experiments deployed between 2005 – 2019, including seismic components and pressure gauges.  
744 We also calculated cross-spectral properties (admittance, phase, coherence) that help reveal and  
745 quantify seismic noise induced by bottom currents and infragravity waves. The resultant dataset  
746 constitutes the most comprehensive sampling of noise characteristics at seafloor stations to date.  
747 Our analysis supplies a framework for BBOBS users to compare and assess the noise  
748 characteristics of individual datasets, better anticipate noise characteristics for newly acquired  
749 data, and provide a baseline catalog that will continue to grow in detail and utility as the marine  
750 geophysics community expands BBOBS sampling of the world’s diverse seafloor.

751 By grouping noise spectra based on metadata parameters, we demonstrate that there are significant  
752 systematics to BBOBS noise characteristics. The most important determinants of noise  
753 characteristics are the seismometer (which strongly covaries with instrument design), and the water  
754 depth at which it is deployed. Accounting for other factors, BBOBSs with CMG-3T seismometers  
755 seem to have higher low-frequency noise than average, and those with T-240 sensors have lower  
756 noise levels, particularly on the vertical components. CMG-3T instruments have higher tilt noise  
757 on the vertical components, most clearly seen at long periods, and overall, more noisy horizontals.  
758 Although noise is correlated with seismometer (and by extension instrument design) type, we find

759 no systematic orientation of the tilt noise, suggesting that none of the BBOBSs' engineering creates  
760 a bias in tilt direction.

761 We have shown, for the first time, that the theoretical depth-frequency limit for seafloor  
762 compliance is closely matched by the data spanning 0-6000 m in water depth. BBOBS deployed  
763 on continental shelves in shallow water (< 500 m) have systematically different noise properties,  
764 characterized in particular by higher noise in the primary microseism band on all four components.  
765 The exception is shallow water lake instruments, which have low noise in the global microseism  
766 band, and a unique  $\sim 0.4$ Hz peak. This and other departures from our main groupings will need to  
767 be reevaluated in the future as new datasets provide wider sampling of station properties.

768 We found that grouping by experiment yielded the highest similarity of spectra, indicating that the  
769 combination of station parameters (similar instrumentation, geographic footprint, etc.)  
770 deterministically controls overall BBOBS noise. This holds promise for informed experiment  
771 planning; overall noise properties are station contingent, but largely predictable. Despite this, we  
772 recognize that our analysis is incomplete, limited by uneven global sampling, and covariance  
773 between important metadata parameters. Key future work may include systematic analysis of  
774 seasonal and other temporal variability, expansion of the dataset to include additional instrument  
775 designs and deployment locations, including non-US-funded deployments, buried or cabled  
776 instrumentation, and testing the effects of iterative noise removal procedures.

777

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787

788 **Author Contribution Statement:** HAJ designed and coordinated the study, carried out the  
789 processing and calculation of spectral properties for all instruments, organized all data for analysis.  
790 ZE assisted in study design and coordination of the writing. JR aided in processing of YOUNG  
791 ORCA data. BB, ZE, HAJ, and SC designed and carried out the spectral angle analysis. HAJ, JR,  
792 ZE, JG, SM, WH each contributed to the collection of metadata parameters for the stations. All  
793 authors participated in interpretation and manuscript preparation.

794

795 **Data Availability Statement:** The seismic and pressure time series data are available for  
796 download through the IRIS Data Management Center (<http://www.iris.edu>). The assembled station  
797 metadata table, and calculated spectral properties are available in the online supplementary material.  
798 Additionally, the calculated noise spectra and cross spectra, and metadata table will archived at Dryad,  
799 currently available upon request. The ATaCR package used to process the data is available here:  
800 <https://github.com/helenjanisz/ATaCR>.

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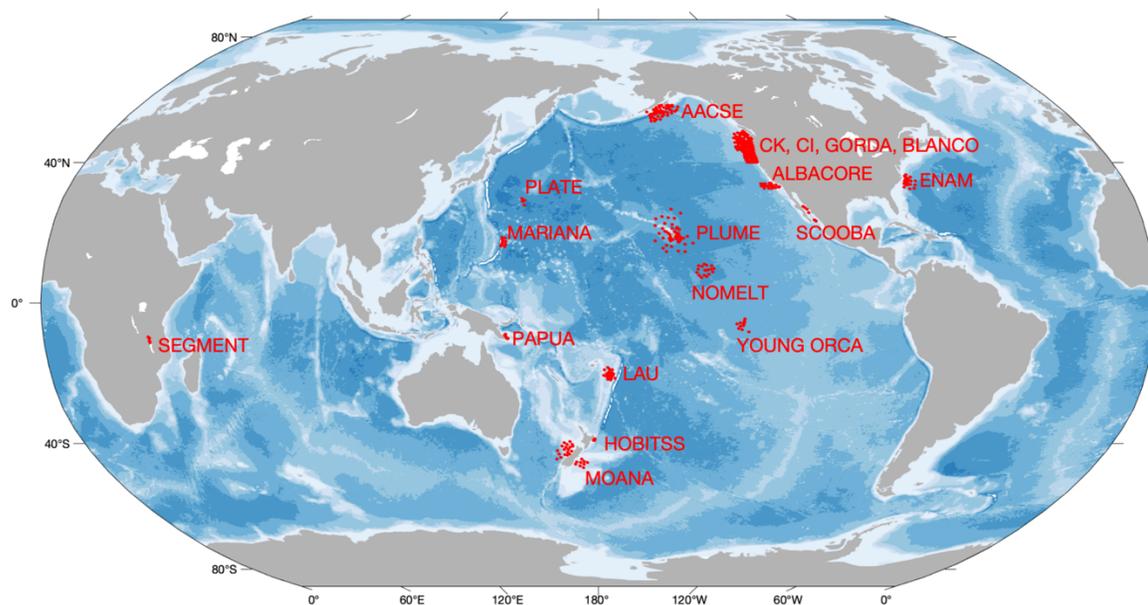
805 **Figures and Tables**

806 **Table 1:** Information related to different BBOBS instrument types included in this study,  
 807 including defined abbreviations to distinguish them.

Abbreviation	Design Institution	Seismometer	Pressure Gauge	Shielding	Instrument Name
AB	SIO	T-Compact	DPG	Syntactic Foam	Abalone
B2	SIO	T-240	DPG	None	SIO Unshielded Broadband
BA	LDEO	T-Compact	APG	None	LDEO Unshielded APG Broadband
BD	LDEO	T-Compact	DPG	None	LDEO Unshielded DPG Broadband
TRM	LDEO	T-Compact	APG	Steel Plates	LDEO Trawl Resistant Mount OBS
AR	WHOI	T-Compact	DPG	None	WHOI ARRA
BG	WHOI	CMG-3T	DPG	None	WHOI BBOBS
KE	WHOI	CMG-3T	DPG	None	WHOI KECK <sup>1</sup>

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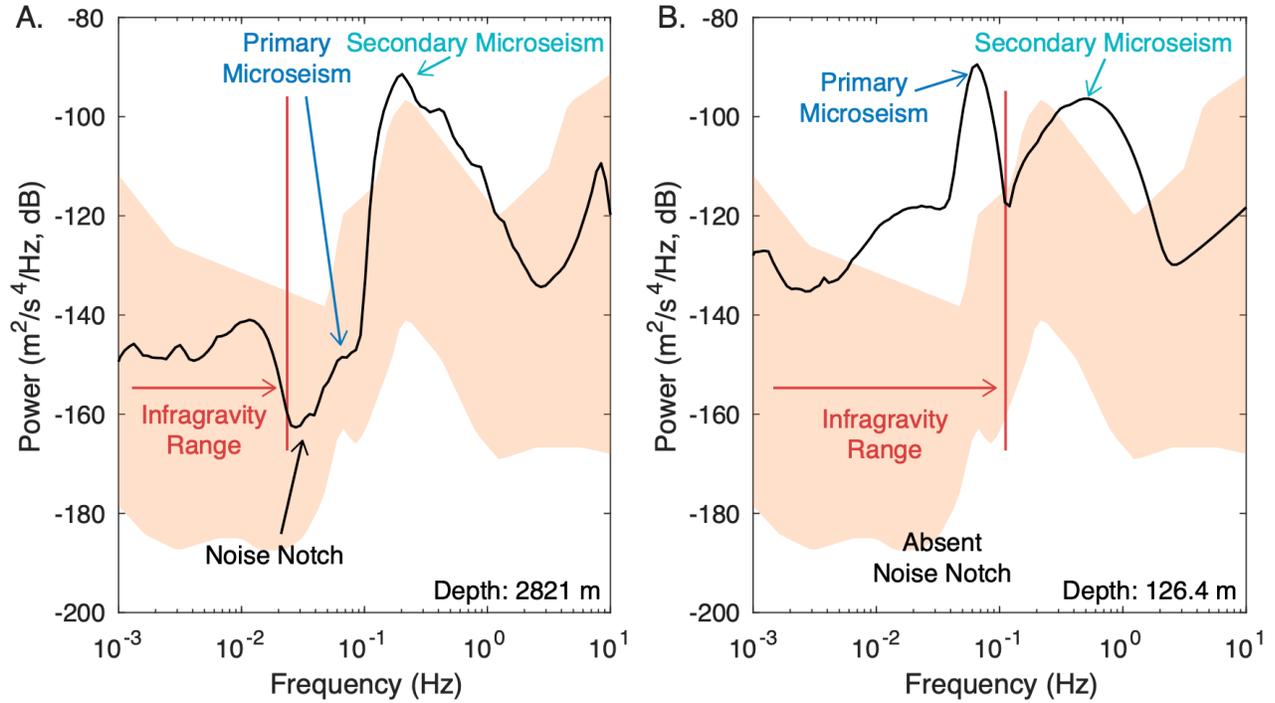
809 Note: 1. The WHOI KECK also includes a strong-motion accelerometer distinguishing it from the  
 810 WHOI BBOBS.



811

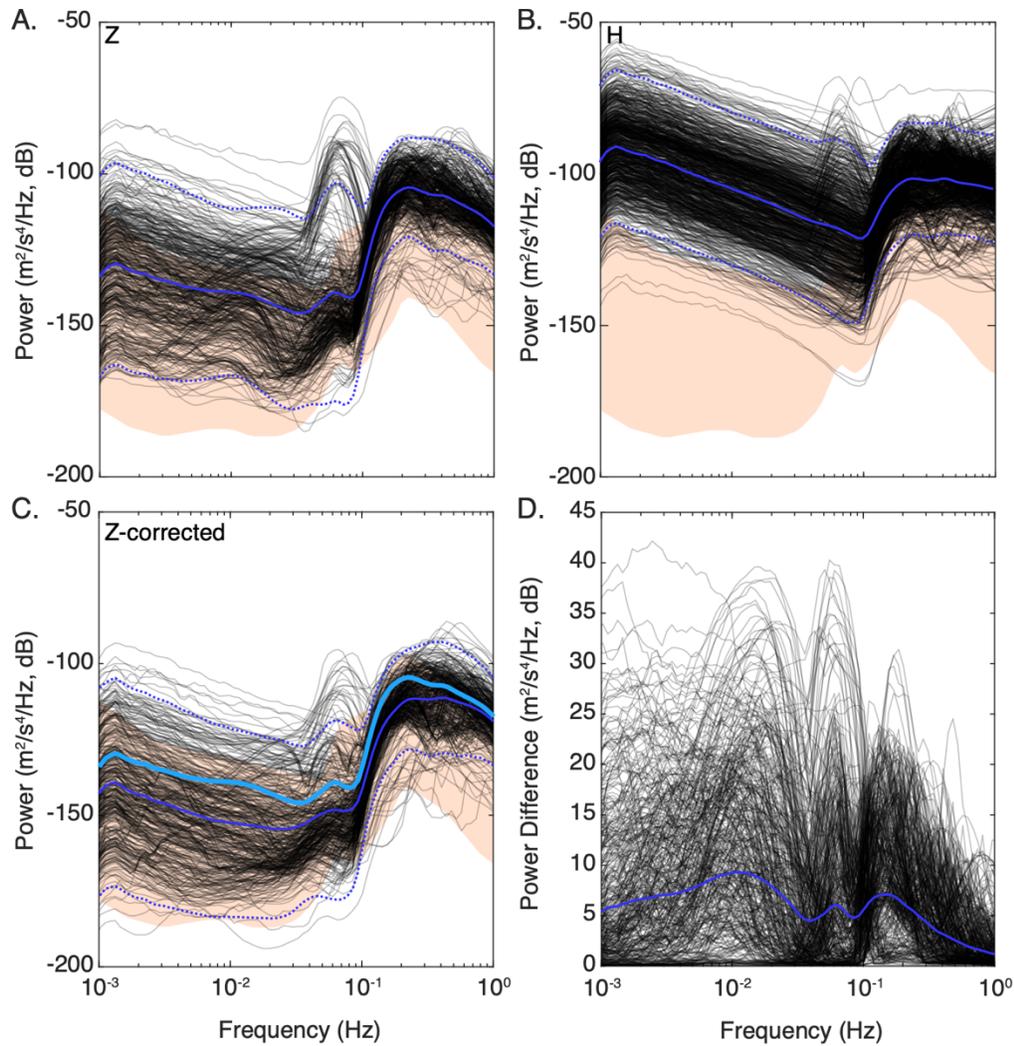
812 **Figure 1:** Map of analyzed BBOBSs (red circles). Details corresponding to each deployment are  
813 given in Table S1. Made using M\_Map (Pawlowicz, 2020).

814



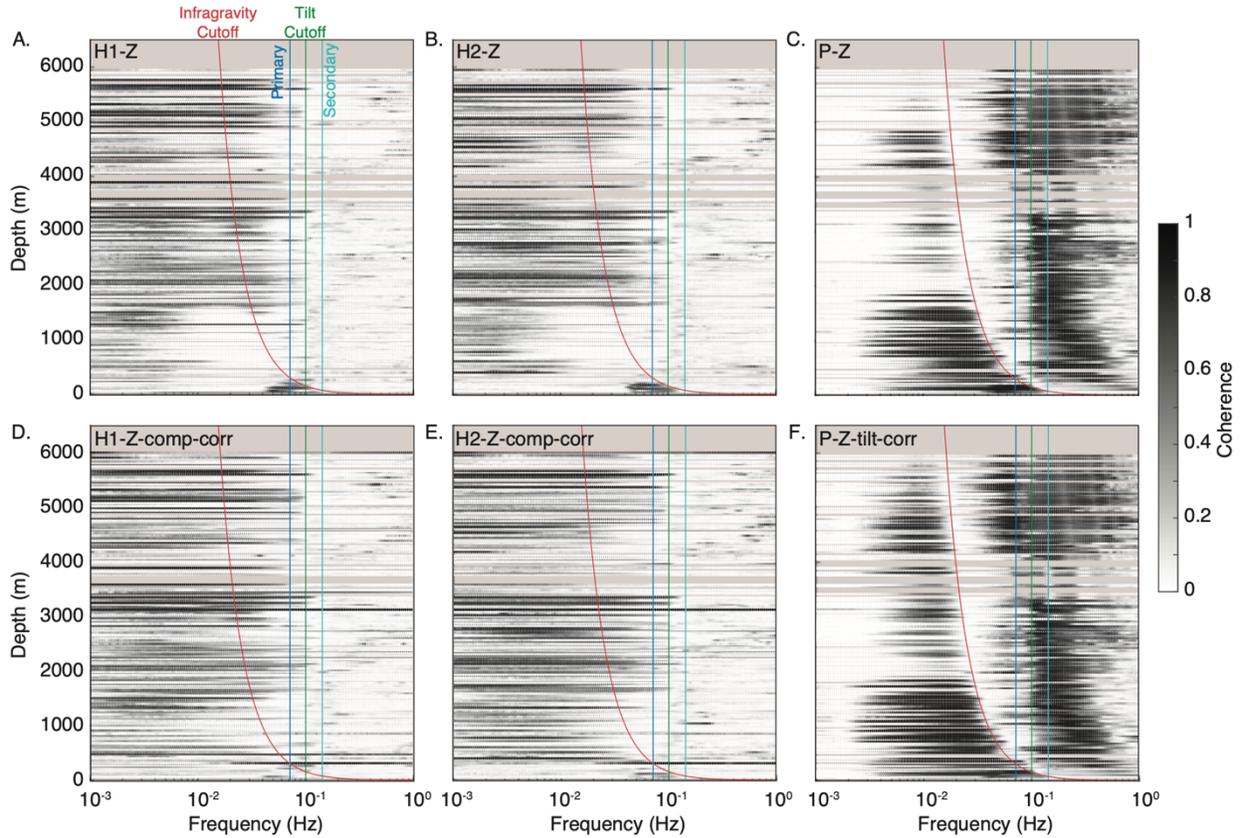
815

816 **Figure 2:** Examples of noise spectra for the Z component of a BBOBS for (a) deep-water station  
817 J36A and (b) shallow-water station M08A, both from the Cascadia Initiative. The primary, and  
818 secondary microseism peaks, the noise notch, and the infragravity band are labeled. The Peterson  
819 (1993) high and low noise ranges are shown as the orange shaded area.



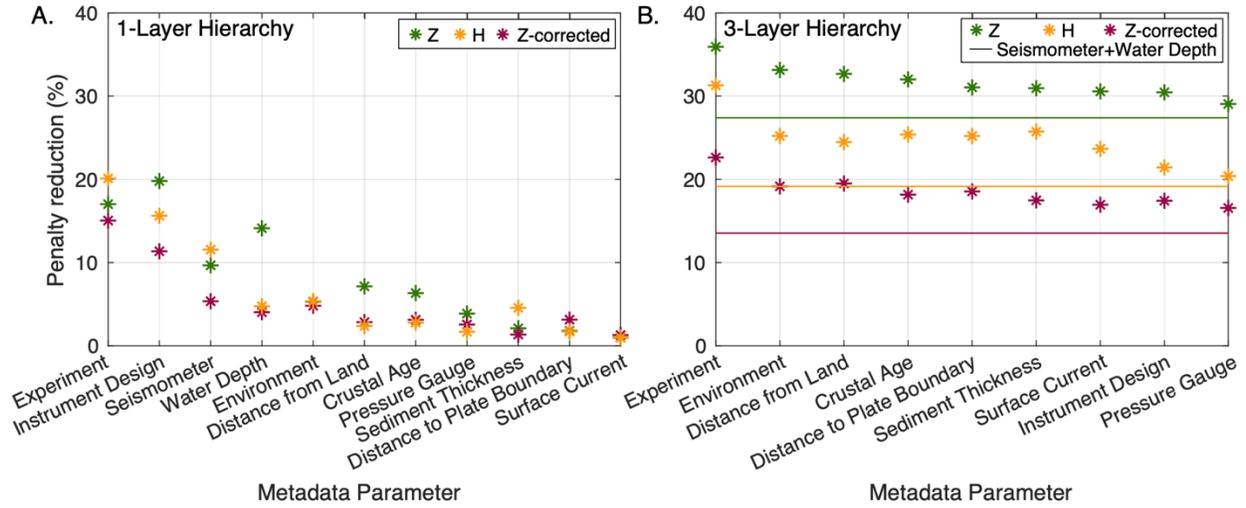
820

821 **Figure 3:** Power spectra for (a)  $Z$ , (b)  $H$ , and (c)  $Z$ -corrected for all individual stations. Solid and  
 822 dashed dark blue lines indicate mean and  $2\text{-}\sigma$  standard deviations, respectively. Solid light blue  
 823 line on (c) is the uncorrected  $Z$  mean for comparison. The Peterson (1993) high and low noise  
 824 ranges are shown as the orange shaded area. (d) Difference between the  $Z$  and  $Z$ -corrected spectra;  
 825 positive values indicate lower values for the corrected dataset. The dark blue line indicates the  
 826 mean difference.



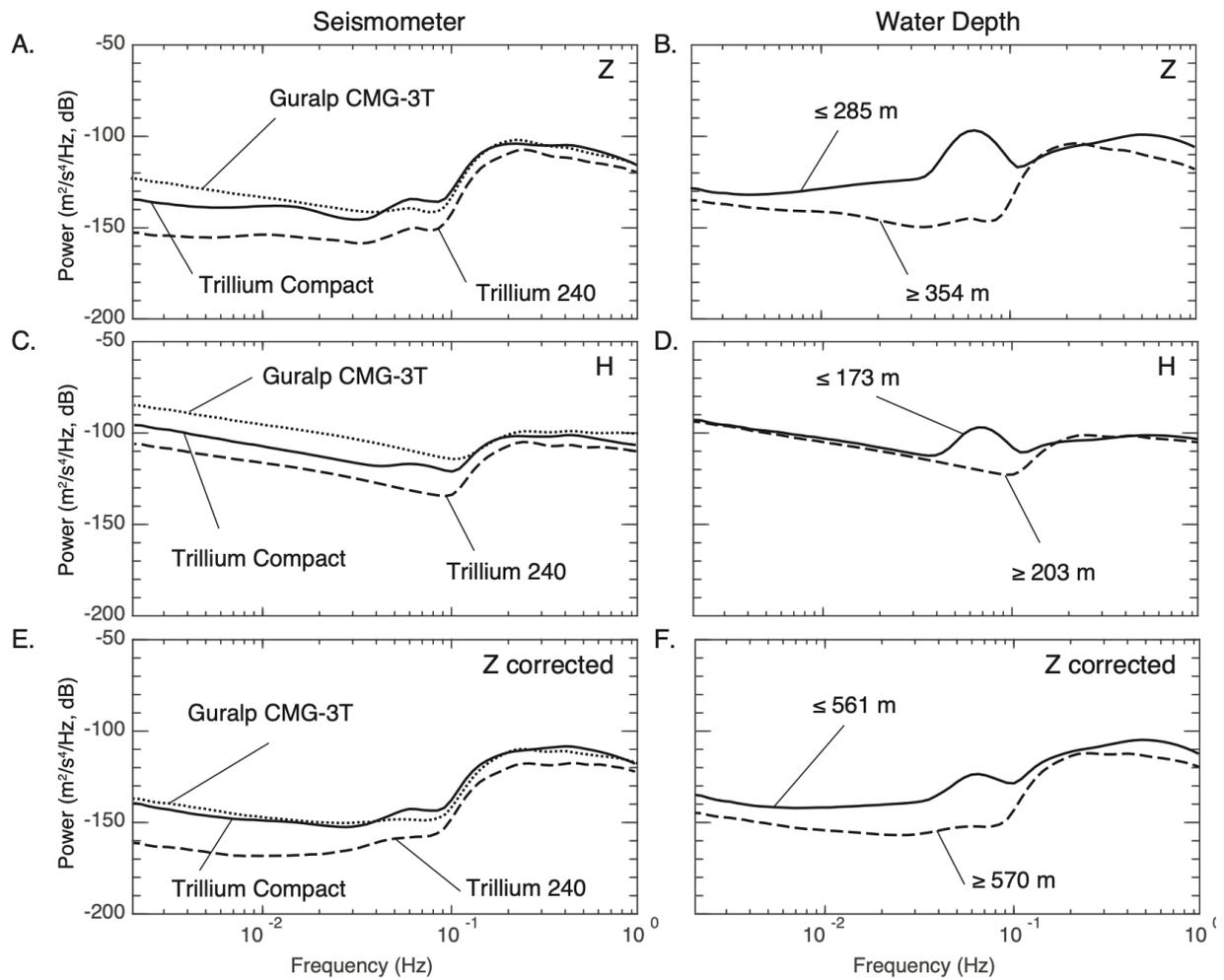
827

828 **Figure 4:** Coherences of the *H1*, *H2*, and *P* with the *Z* for each BBOBS compared with the water  
 829 depth of the instrument. The red line indicates the predicted infragravity cutoff frequency ( $f$ ) as a  
 830 function of water depth ( $d$ ), using the equation  $f = \sqrt{\frac{g}{2\pi d}}$  (Bell et al., 2015); the green, dark blue,  
 831 and light blue lines indicate the tilt cutoff (0.1 Hz), primary (0.07 Hz), and secondary (0.14 Hz)  
 832 microseism peaks respectively. (a-c) Coherences calculated with the *Z* component; (d-e) *H1-Z* and  
 833 *H2-Z* coherences with the *Z-compliance-corrected* component; (f) *P-Z* coherence with the *Z-tilt-*  
 834 *corrected* component.



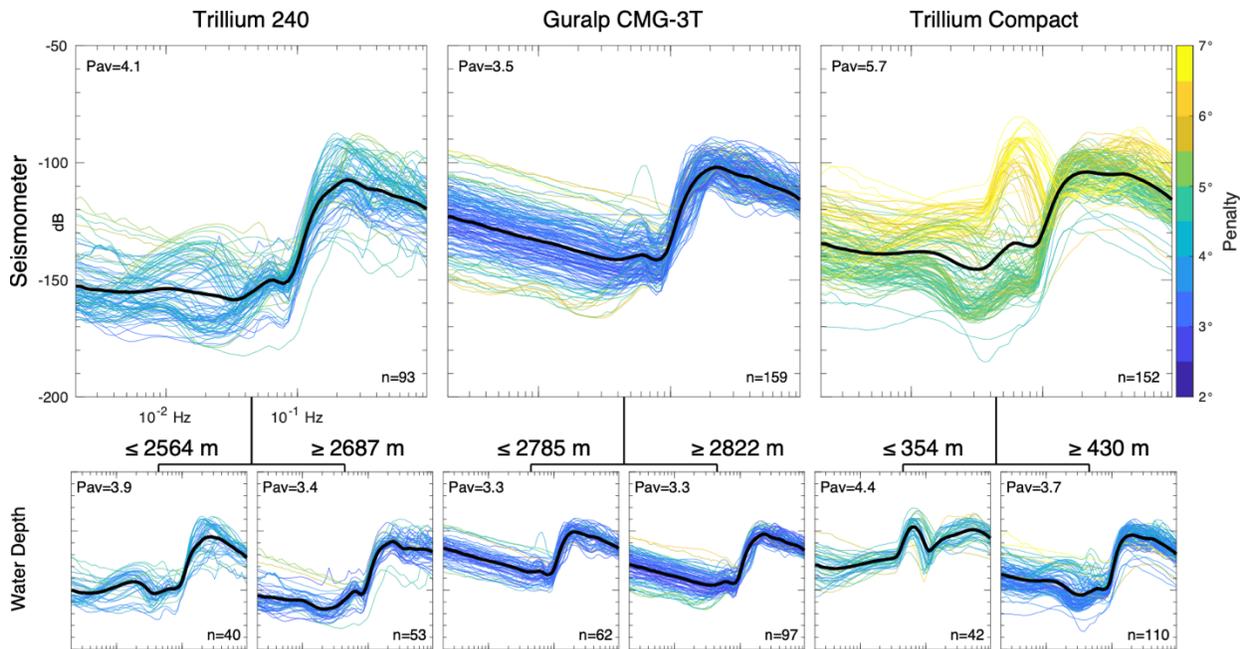
835

836 **Figure 5:** Percentage penalty reduction in spectral angle for  $Z$ ,  $H$ , and  $Z$ -corrected for each  
 837 metadata parameter subdivision. Larger reductions indicate more similarity within the final  
 838 subgroups. Parameters are sorted from left to right in descending order of their average penalty  
 839 reductions. (a) Results for each metadata parameter (1-layer analysis). (b) Results after  
 840 subgrouping the BBOBS by “*Seismometer*”, then “*Water Depth*”, and then the labeled metadata  
 841 parameter (3-layer analysis). The 2-layer penalty reductions for “*Seismometer*” and “*Water*  
 842 *Depth*” are shown by the solid lines.



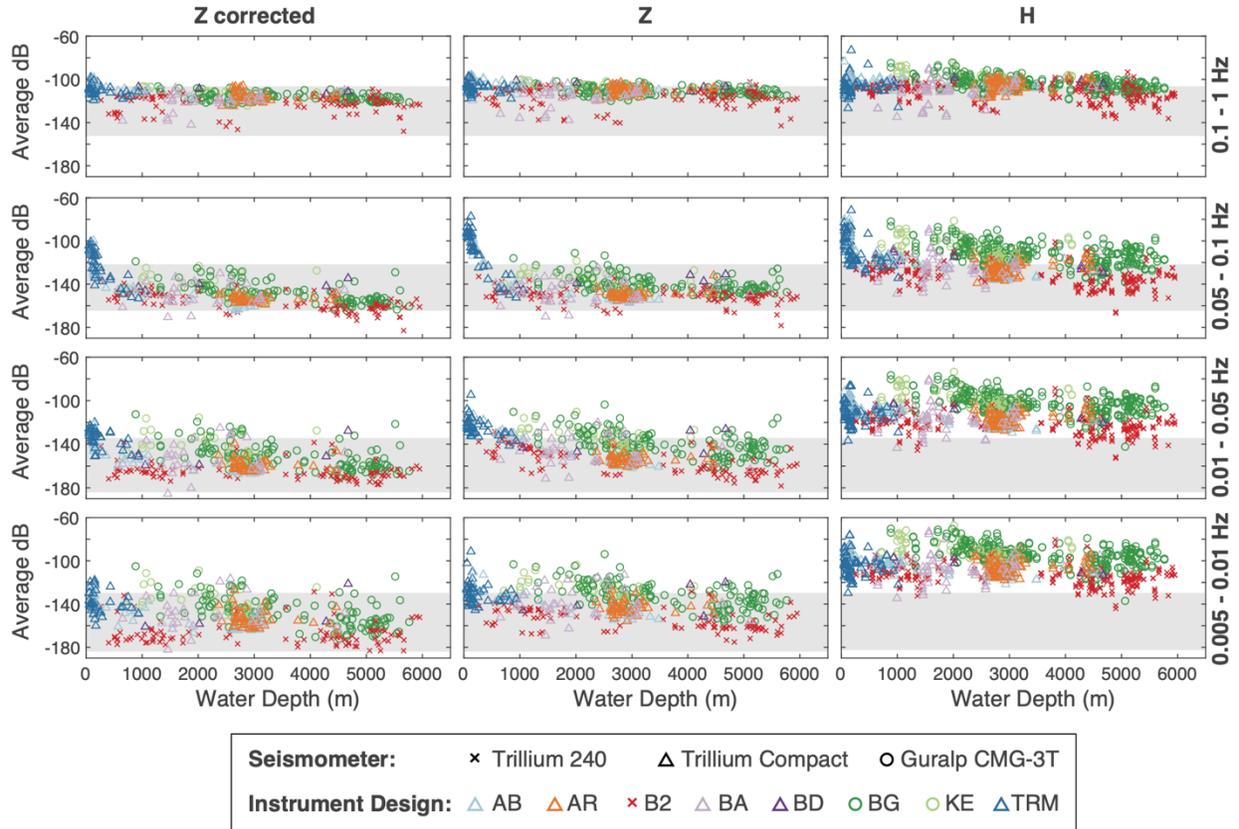
843

844 **Figure 6:** Average spectra calculated from the resultant metadata subgroups based on  
 845 “*Seismometer*” and “*Water Depth*”. (a) The average *Z* spectra for the three “*Seismometer*”  
 846 subgroups: CMG-3T, T-Compact, and T-240. (b) The average *Z* spectra for the two “*Water*  
 847 *Depth*” subgroups (grid search determined cutoff depths indicated). (c) Same as (a), but for the *H*  
 848 components. (d) Same as (b), but for the *H* components. (e) Same as (a), but for the *Z-corrected*  
 849 component. (f) Same as (b), but for the *Z-corrected* component.



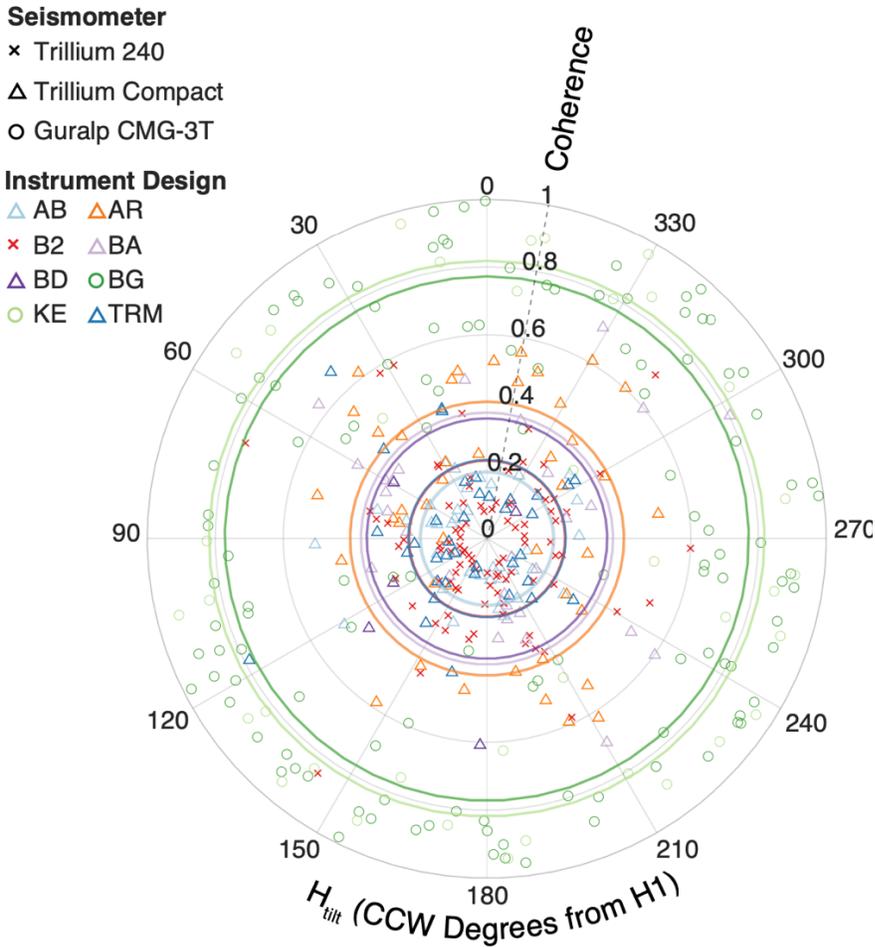
850

851 **Figure 7:** Example of subgrouping by spectral similarity, showing the 2-layer analysis for  $Z$   
 852 spectra. These were first subgrouped by “*Seismometer*”, then subgrouped by “*Water Depth*”. The  
 853 seismometer types and threshold depths are indicated above each plotted subgroup. For each  
 854 subgroup, the average spectrum is plotted in black. Individual spectra are colored according to  
 855 their average spectral angle (*i.e.*, penalty) from the other spectra in that subgroup. The same  
 856 vertical and horizontal scale is used for all plots. The number of spectra ( $n$ ) and the average penalty  
 857 ( $P_{av}$ ) for each subgroup is given in the corresponding plots.



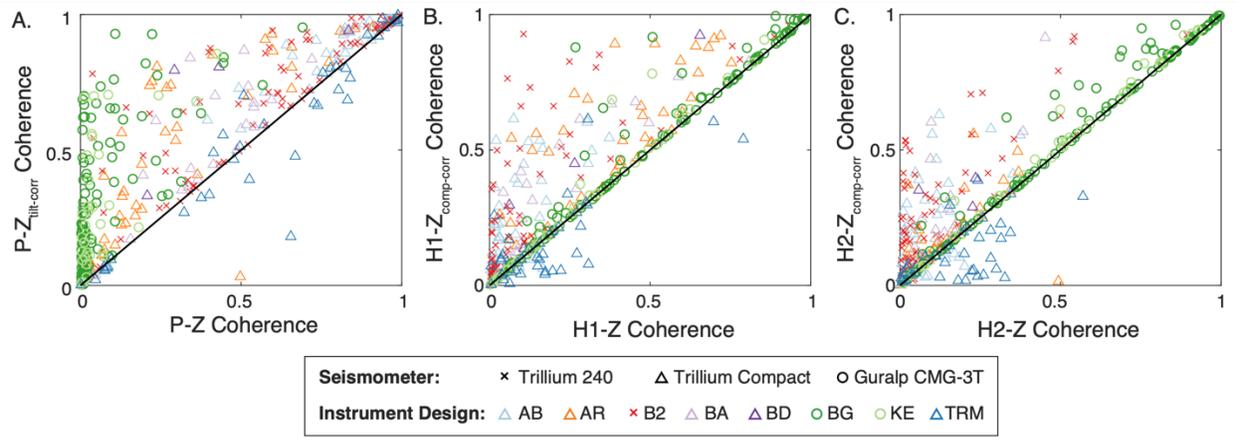
858

859 **Figure 8:** Average power for the  $Z$ ,  $H$ , and  $Z$ -corrected components plotted as a function of water  
 860 depth for each BBOBS in four frequency bands (ranges shown on right). Symbols indicate the  
 861 seismometer; colors indicate the instrument design (see Table 1 for more details). Gray shading  
 862 indicates the average Peterson (1993) high and low noise model range in each frequency band.



863

864 **Figure 9:** The tilt orientation ( $H_{tilt}$ ), measured as a function of degrees counterclockwise from  $H1$ ,  
 865 and the corresponding  $H_{tilt}$ - $Z$  coherence for each BBOBS. Symbols and colors indicate  
 866 seismometer and instrument design. Colored lines show the average  $H_{tilt}$ - $Z$  coherence for each  
 867 instrument design.



868

869 **Figure 10:** Comparison of coherences before and after tilt or compliance corrections. Symbols  
 870 and colors indicate seismometer and instrument type; symbols that plot above the black line  
 871 indicate an increase in coherence after corrections, below the line indicate a decrease in coherence,  
 872 and along the line indicate no change. (a) Comparison of the  $P$ - $Z$  and  $P$ - $Z$ -tilt-corrected  
 873 coherences. (b) Comparison of the  $H1$ - $Z$  and  $H1$ - $Z$ -compliance-corrected coherences. (c)  
 874 Comparison of the  $H2$ - $Z$  and  $H2$ - $Z$ -compliance-corrected coherences.

875

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- 1124

1125 Supplementary material for

1126 **Broadband Ocean Bottom Seismometer Noise Properties**

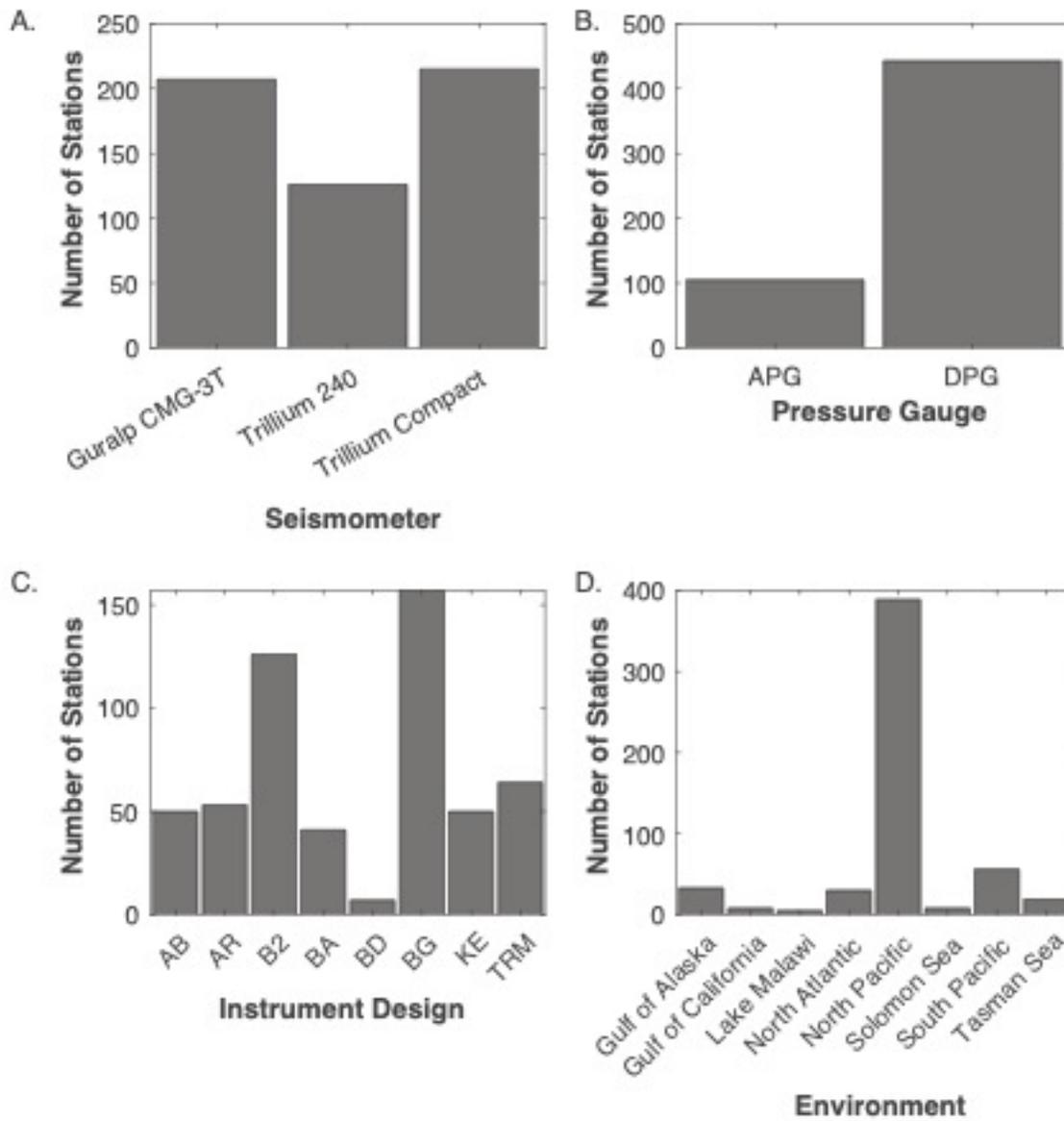
1127 Helen A. Janiszewski<sup>\*</sup>, Z. Eilon<sup>2</sup>, J.B. Russell<sup>3</sup>, B. Brunsvik<sup>2</sup>, J.B. Gaherty<sup>4</sup>, S.G. Mosher<sup>5</sup>, W.B. Hawley<sup>6</sup>,  
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1131 Science, University of Santa Barbara, Santa Barbara, CA; 3. Dept. of Earth, Environmental and  
1132 Planetary Sciences, Brown University, Providence, RI; 4. School of Earth & Sustainability, Northern  
1133 Arizona University, Flagstaff, AZ; 5. Department of Earth and Environmental Sciences, University of  
1134 Ottawa, Ottawa, Canada; 6. Lamont-Doherty Earth Observatory of Columbia University, Palisades,  
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1136 **Contents:**

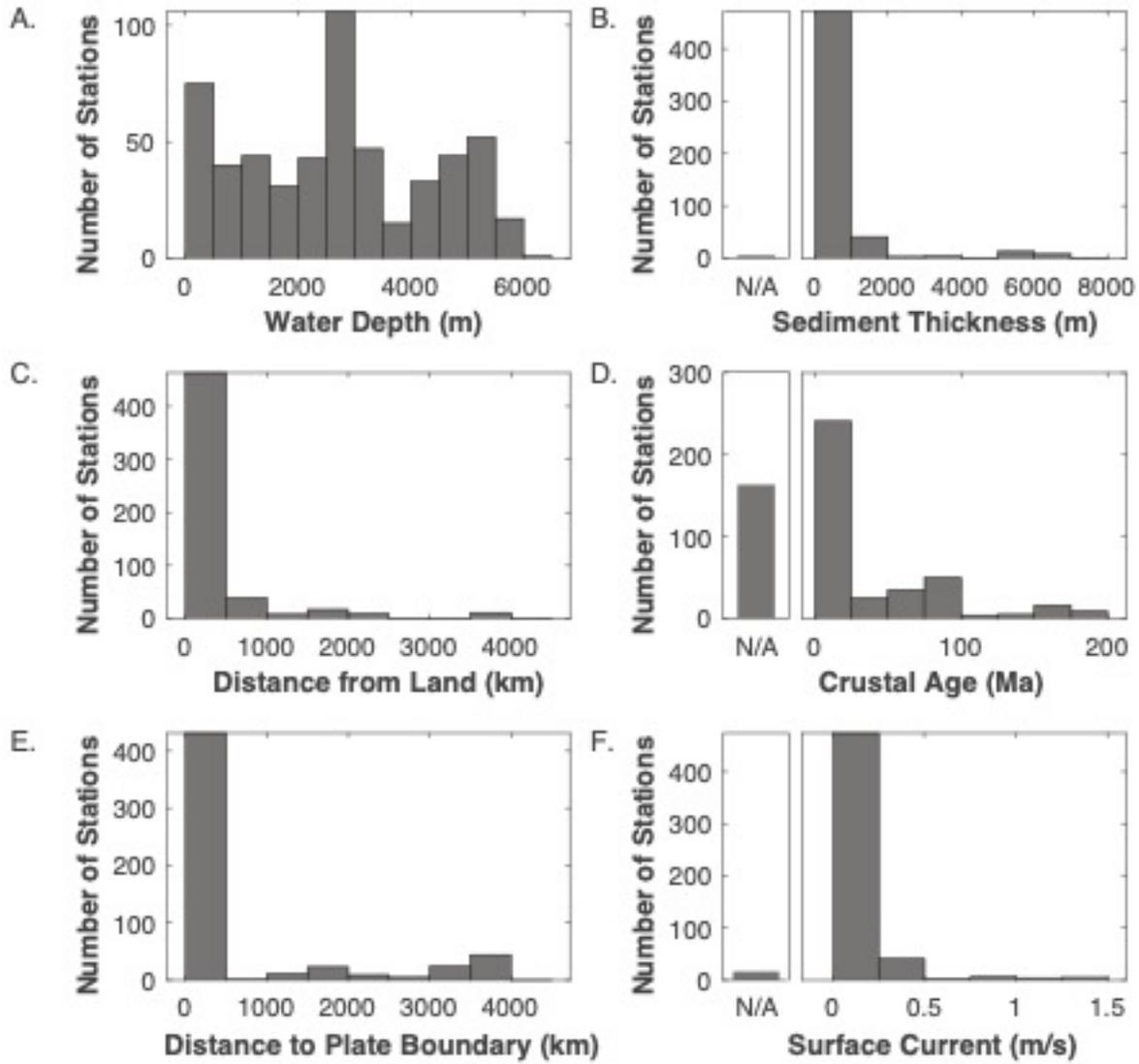
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1138 Figure S1: Deployment instrumentation and location histograms.  
1139 Figure S2: Deployment environmental properties histograms.  
1140 Figure S3: Vertical component noise spectra sorted by experiment.  
1141 Figure S4: Horizontal component noise spectra sorted by experiment.  
1142 Figure S5: Absolute pressure gauge component noise spectra sorted by experiment.  
1143 Figure S6: Differential pressure gauge component noise spectra sorted by experiment.  
1144 Figure S7: Corrected vertical component noise spectra sorted by experiment.  
1145 Figure S8: Horizontal-pressure coherences as a function of water depth.  
1146 Figure S9: Results of spectral angle grouping water-depth grid search.  
1147 Figure S10: Example of 3-layer spectral angle results for the vertical component.  
1148 Figure S11: Example of 3-layer spectral angle results for the horizontal component.  
1149 Figure S12: Example of 3-layer spectral angle results for the vertical corrected component.  
1150 Figure S13: Significance tests for spectral angle analysis.  
1151 Figure S14: Comparison of lake- and shelf-deployed shallow water vertical spectra.  
1152 Table S1: Information about the experiments that were used for our analysis. Includes appropriate  
1153 citation information for datasets, as well as the network names under which the data is catalogs in  
1154 the IRIS DMC.  
1155 Table S2: Assembled metadata parameter table for all examined instruments. A description of the  
1156 source of the data, including references, is provided in the main text. This table also indicates if a  
1157 particular channel was flagged as “good” after quality control procedures, where a value of “1”  
1158 indicates the spectra was accepted and “0” indicates it was discarded.  
1159 Data Access: Calculated noise spectra and cross spectra, and metadata table will be archived at  
1160 Dryad, doi available upon request, will be made public and included with publication.

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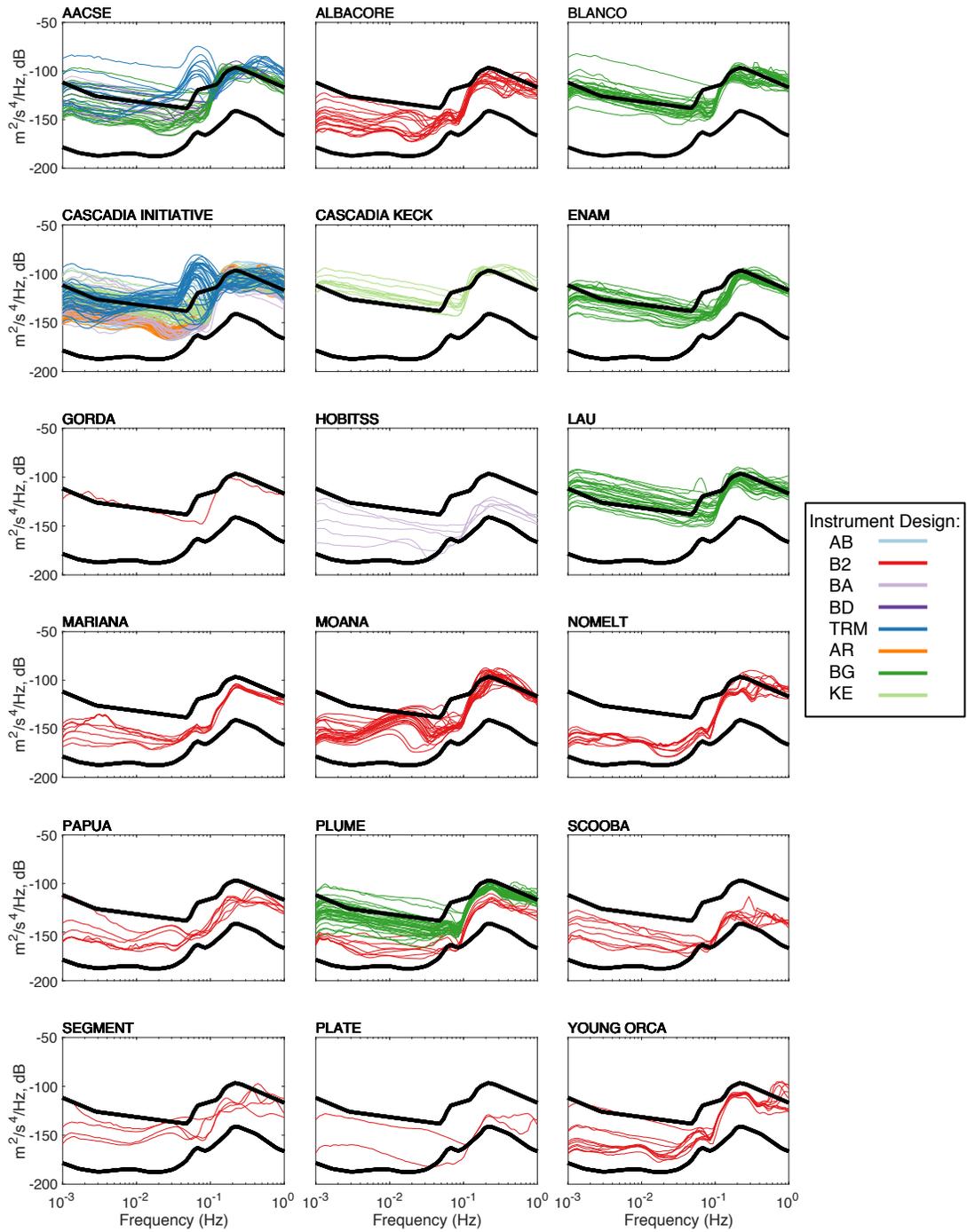
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1167 **Figure S1:** Distributions of categorical station variables. (a) Types of seismometers, (b) types of  
 1168 pressure gauges, (c) BBOBS designs, and (d) deployment environment included in our analysis.



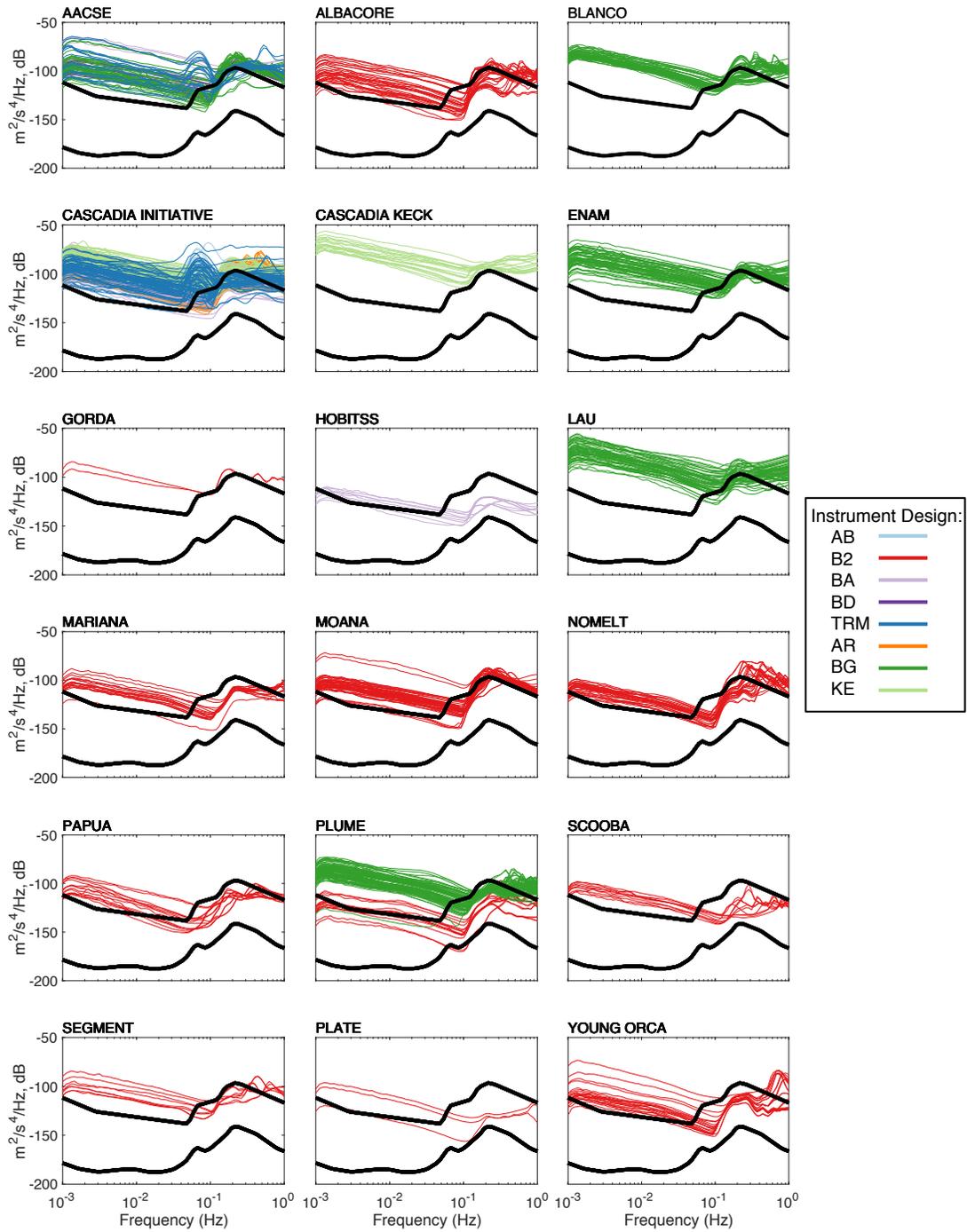
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1170 **Figure S2:** Distributions of numerical station variables. (a) Water depth, (b) sediment thickness,  
1171 (c) distance from major coastline, (d) crustal age, (e) distance to nearest plate boundary, (f) surface  
1172 current.



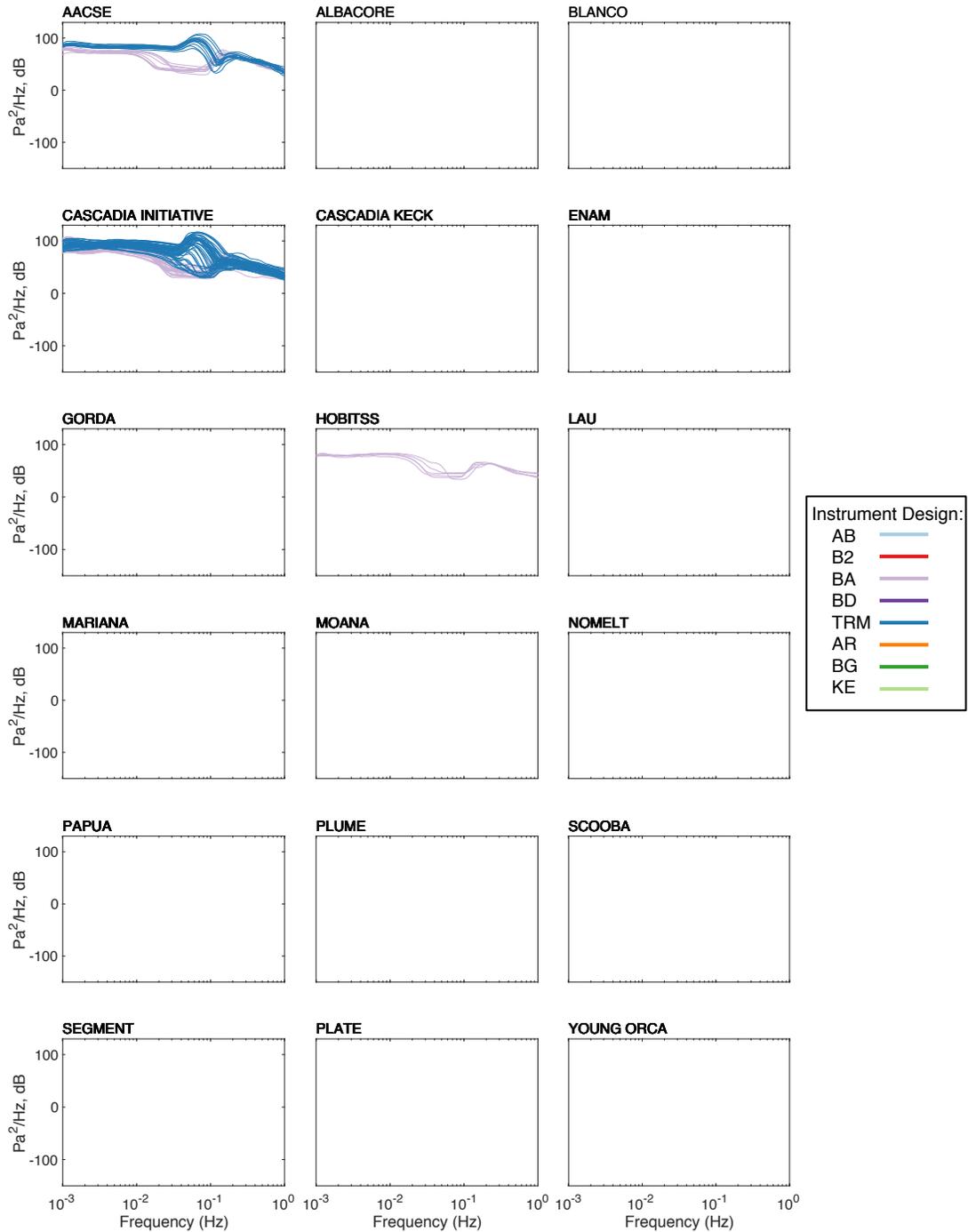
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1174 **Figure S3:** Power spectra for the Z component for each experiment (Table S1). Spectra are  
 1175 colored by instrument design; abbreviations are defined in Table 1.

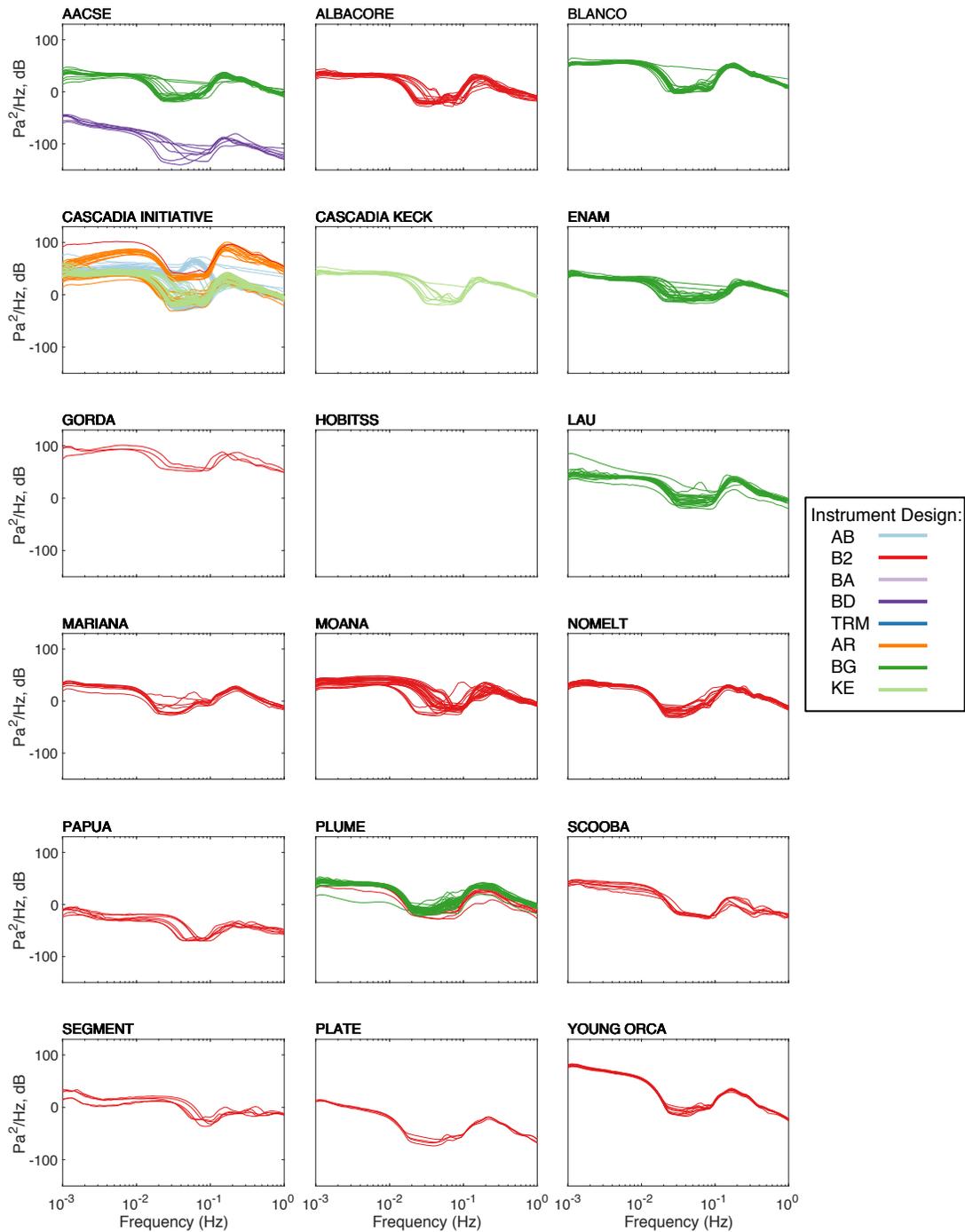


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**Figure S4:** Power spectra for the  $H$  components for each experiment (Table S1). Spectra are colored by instrument design; abbreviations are defined in Table 1.

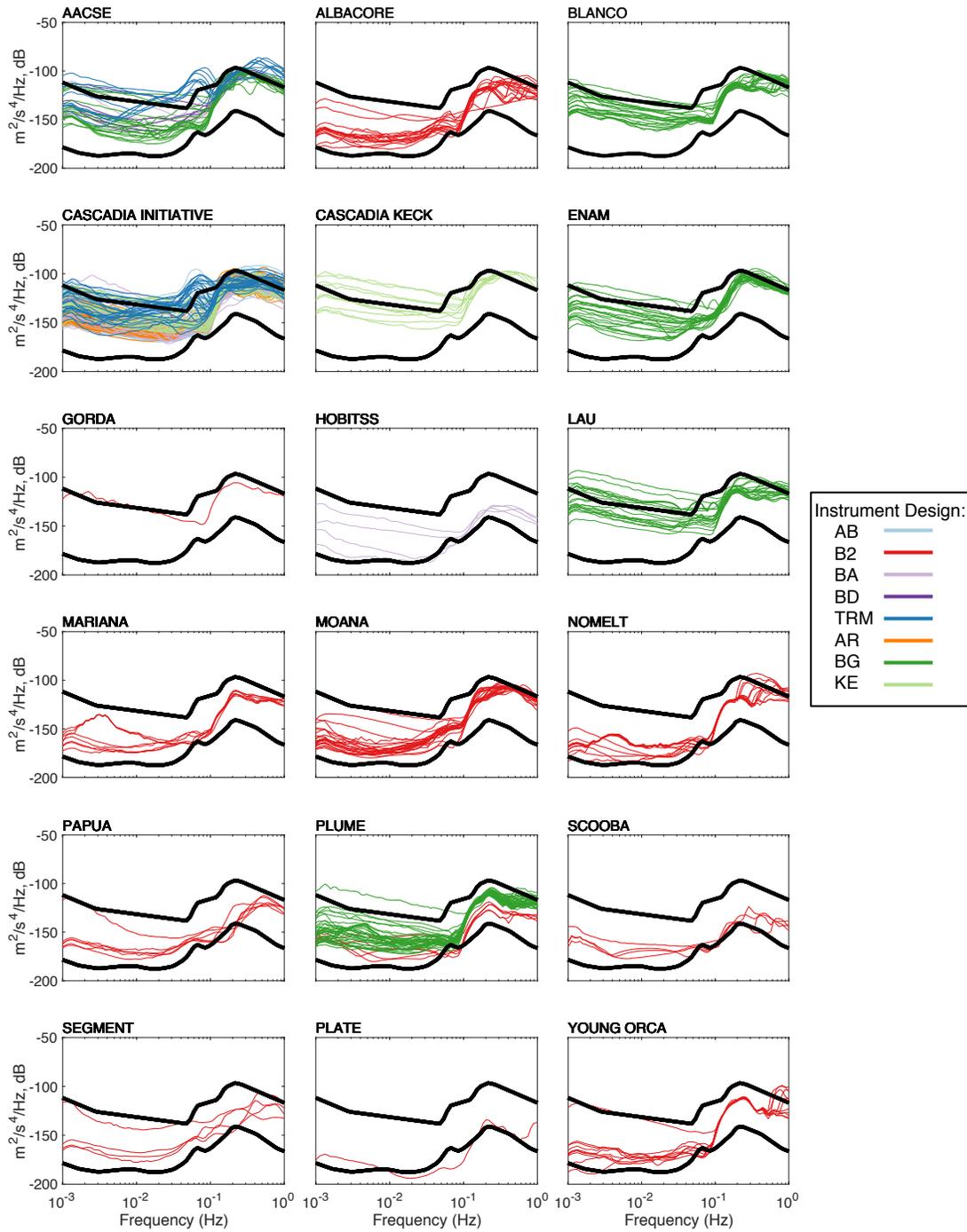


1180  
 1181 **Figure S5:** Power spectra for the APGs for each experiment (Table S1). Spectra are colored by  
 1182 instrument design; abbreviations are defined in Table 1. Experiments where APGs were not  
 1183 deployed do not show any data. The response has not been removed from the instruments.



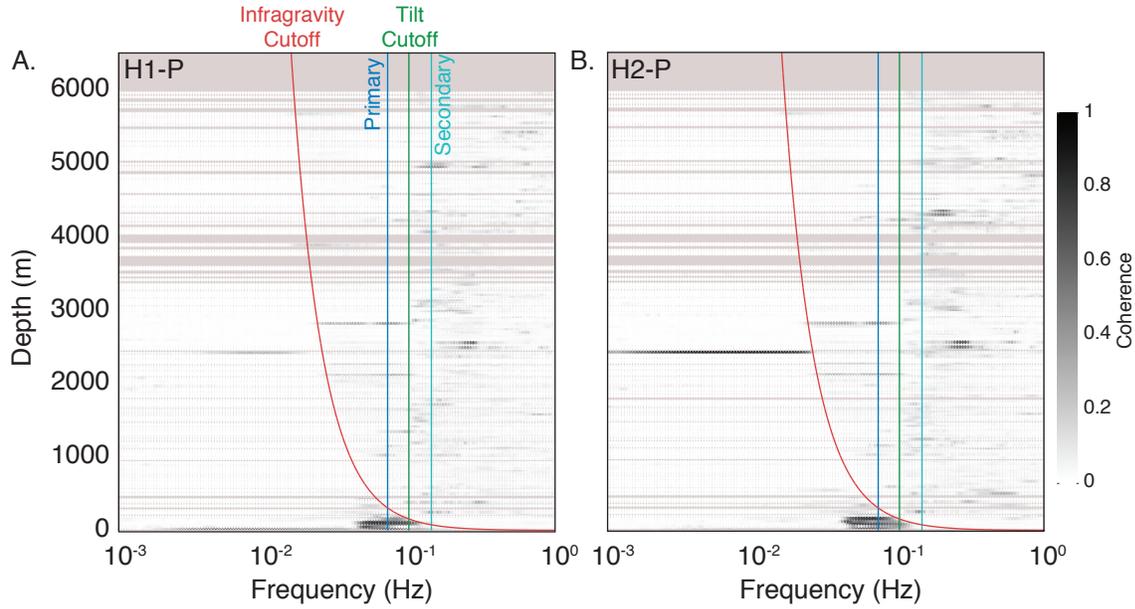
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**Figure S6:** Power spectra for the DPGs for each experiment (Table S1). Spectra are colored by instrument design; abbreviations are defined in Table 1. Response is removed, and data are processed identically for all experiments. Several of the experiments have instruments or subsets of instruments where the amplitudes of the DPG spectra are significantly outside the normal amplitude range. This is likely due to errors in instrument calibration, since it is systematically observed across different experiment or instrument type subsets.

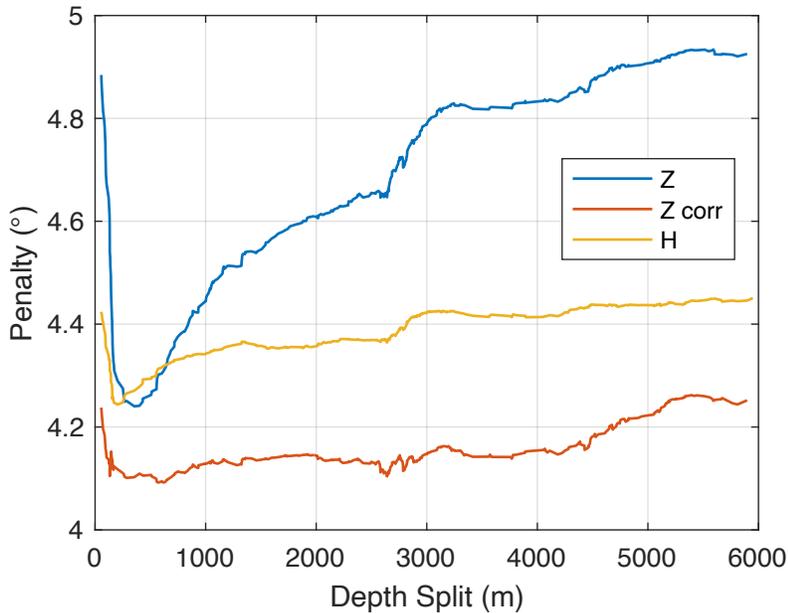


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**Figure S7:** Power spectra for the *Z-corrected* component for each experiment (Table S1). Spectra are colored by instrument design; abbreviations are defined in Table 1.

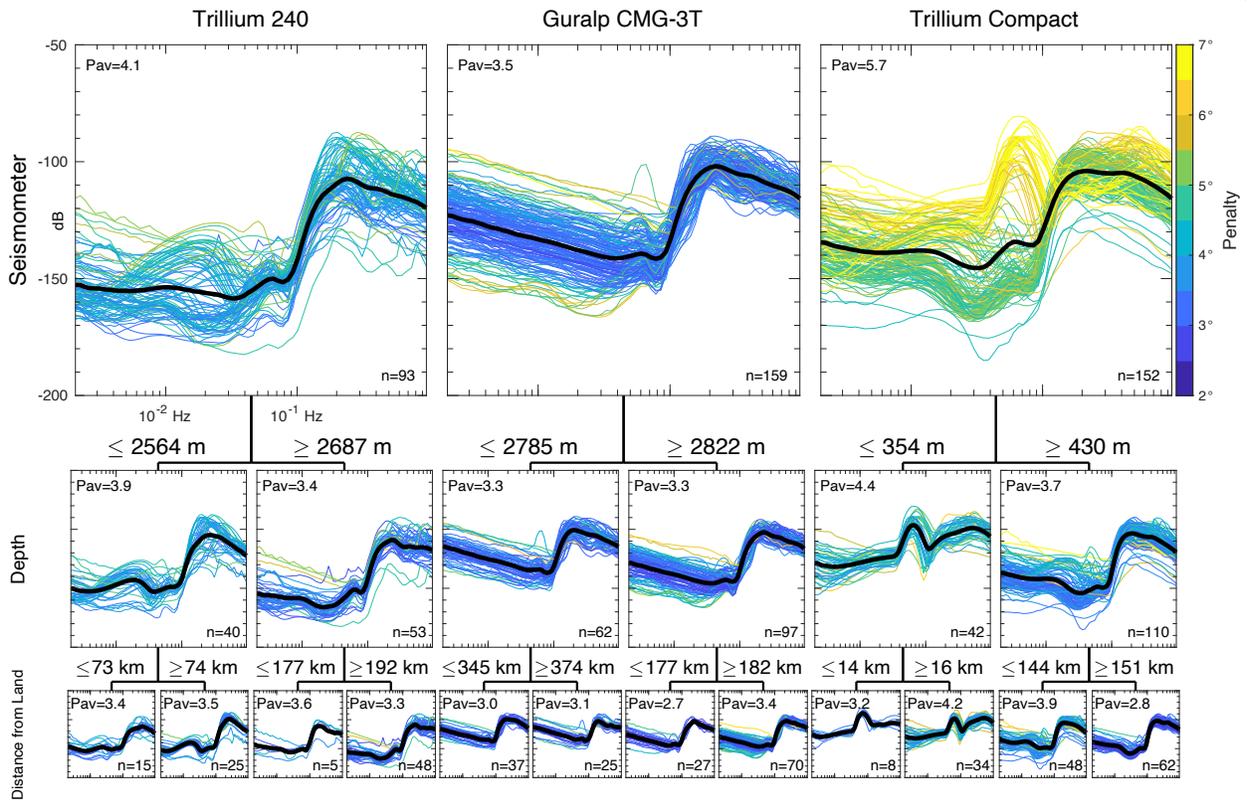


1195 **Figure S8:** Coherence of the (a) *H1* and (b) *H2* components with the *P* component for each  
 1196 BBOBS compared with the water depth of the instrument. The red line indicates the predicted  
 1197 infragravity cutoff frequency, using  $f = \sqrt{\frac{g}{2\pi d}}$  (Bell et al., 2015); the green, dark blue, and light  
 1198 blue lines indicate the tilt cutoff (0.1 Hz), primary (0.07 Hz), and secondary (0.14 Hz)  
 1199 microseism peaks respectively.  
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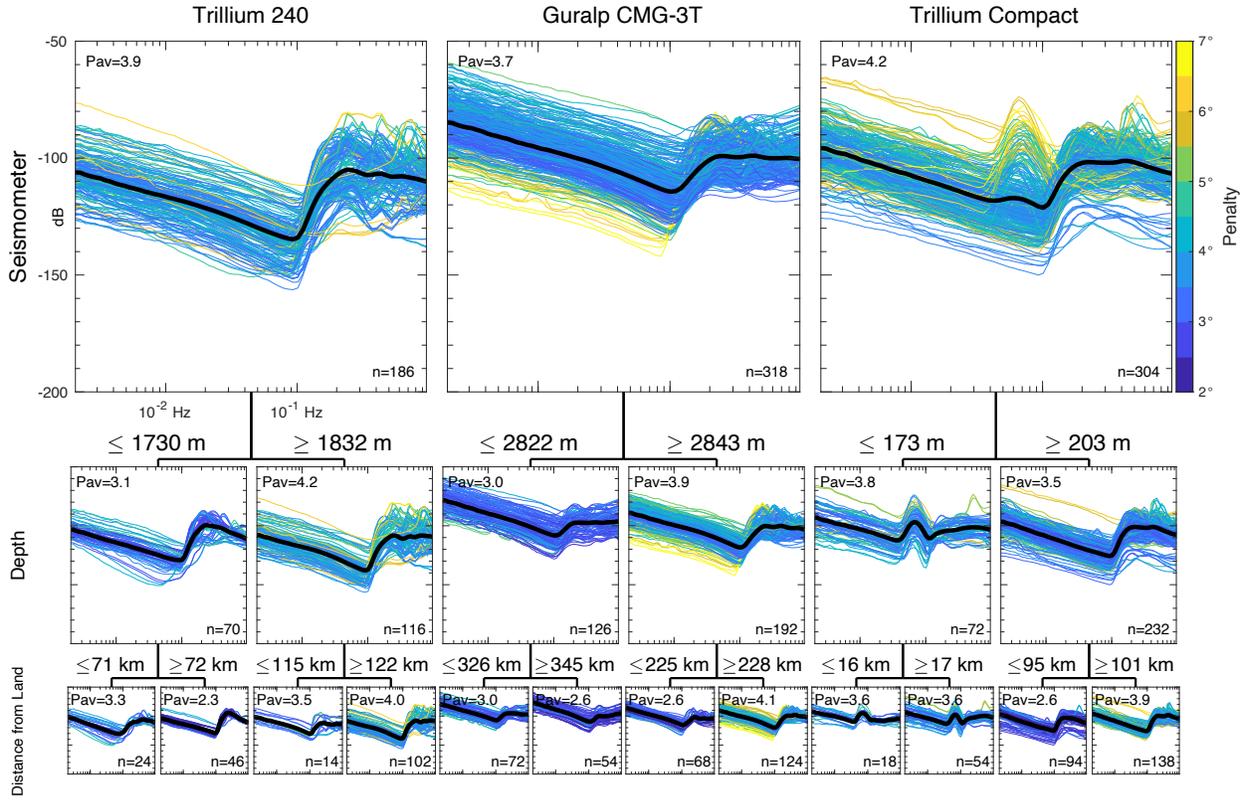
1202 **Figure S9:** Examples of the grid search procedure used to find the best water depth value to split  
 1203 spectra into two subgroups for the *Z*, *H*, and *Z-corrected* components. The average penalty is  
 1204 reported as a function of water depth split.  
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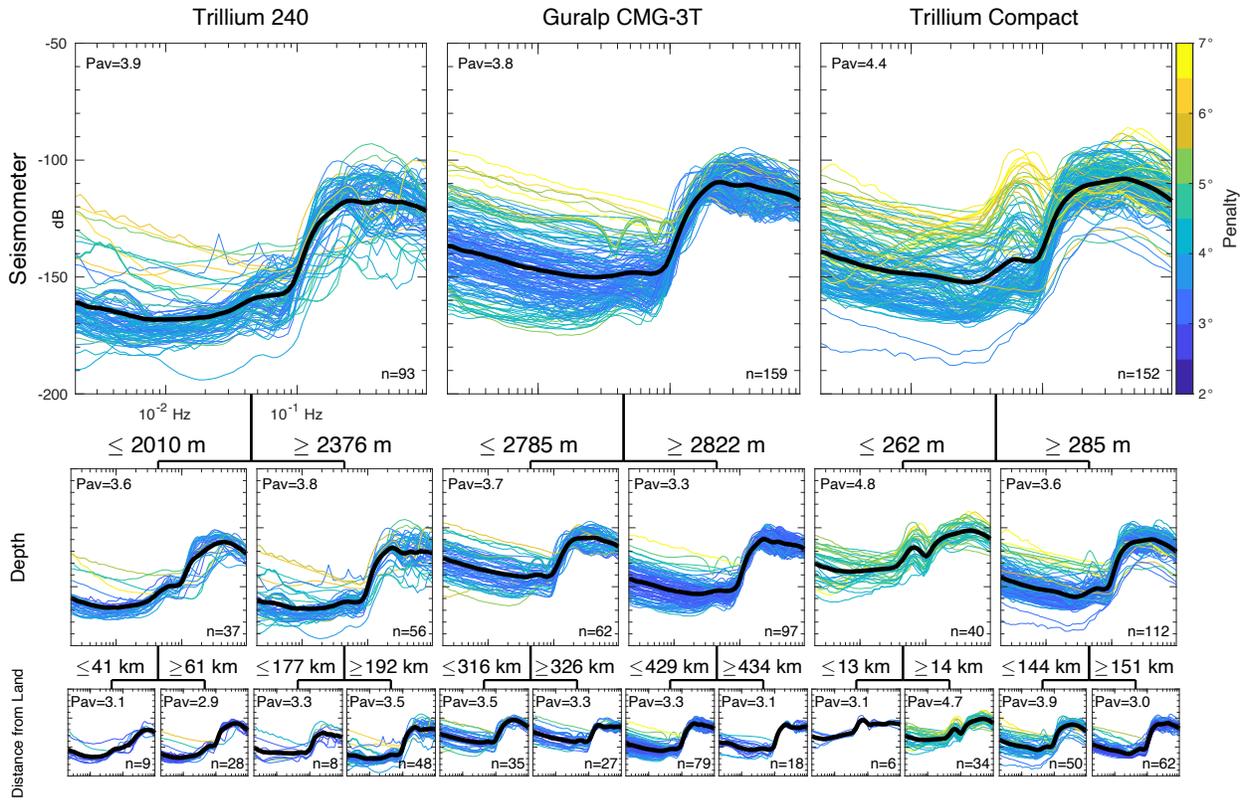
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**Figure S10:** Vertical component 3-layer hierarchy example. Same as Figure 7, but showing an example where the third layer subgroup of spectra is based on “Distance from Land” (bottom row).



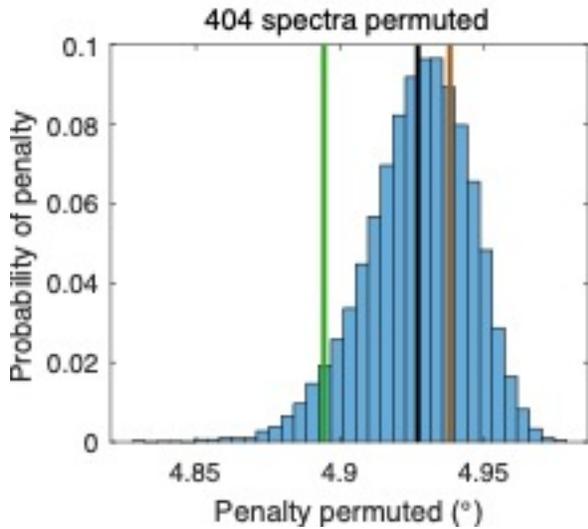
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**Figure S11:** Same as Figure S10, but for the *H* components.



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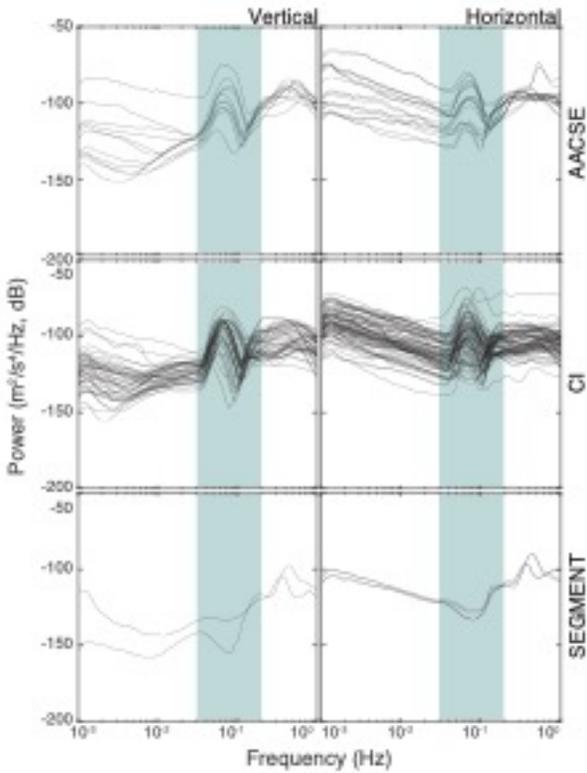
**Figure S12:** Same as Figure S10, but for the *Z*-corrected component.



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1221 **Figure S13:** Results of test of distribution of penalties after randomly permuting spectra between  
 1222 clusters for the 3-layer analysis on the *Z* component grouped for “*Seismometer*”, “*Water*  
 1223 *Depth*”, and “*Distance from Land*”. Averaged results of the random permutations (black),  
 1224 compared with the baseline penalty (brown). The green line shows the penalty above which 95%  
 1225 of perturbations resulted in.

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1229 **Figure S14:** Comparison of vertical and horizontal spectra from all BBOBS deployed in water  
1230 depths < 500 m, which include instruments from the AACSE, CI, and SEGMeNT experiments  
1231 (Table S1). Blue shaded region extends over the primary microseism band, and is meant to  
1232 emphasize the difference in spectra between instruments deployed on the continental shelf (e.g.  
1233 AACSE and CI) as opposed to in a lake environment (e.g. SEGMeNT).

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1245 **Table S1- Experiment List**

Experiment Name	Network	Instrument Types	Years	Citation
SCOOBA	ZL	B2	2005-06	Sumy et al., 2013
PLUME	YS	BG, B2	2005-07	Wolfe et al., 2009
MOANA	ZU	B2	2009-10	Collins & Sheehan, 2009
PLATE	Z6	B2	2009-10	Takeo et al., 2014
LAU	YL	BG	2009-10	Zha et al., 2014
ALBACORE	2D	B2	2010-11	Lin et al., 2015
CASCADIA KECK	7A	KE	2010-11	Toomey et al., 2014
PAPUA	ZN	B2	2010-11	Abers & Gaherty, 2010
CASCADIA INITIATIVE	7D	TRM, BA, KE, AR, AB	2011-15	Toomey et al., 2014
NOMELT	ZA	B2	2011-12	Lin et al., 2016
MARIANA	XF	B2	2012-13	Wiens, 2012
BLANCO	X9	BG	2012-13	Nabelek & Braunmiller, 2012
GORDA	Z5	B2, AB, BA	2013-15	Nabelek & Braunmiller, 2013
ENAM	YO	BG	2014-15	Gaherty, 2014
HOBITSS	YH	BA	2014-15	Wallace et al., 2014
SEGMeNT	YQ	B2	2015	Gaherty et al., 2013
AACSE	XO	TRM, BA, BD, BG, KE	2018-19	Abers et al., 2018
YOUNG ORCA	XE	B2	2018-19	Eilon et al., 2021

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