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Abstract

Satellite observations of snow-covered regions in the microwave range have the potential to retrieve essential climate variables such as snow height. This requires a precise understanding of how microwave scattering is linked to snow microstructural properties (density, grain size, grain shape and arrangement). This link has so far relied on empirical adjustments of the theories, precluding the development of robust retrieval algorithms. Here we solve this problem by introducing a new microstructural parameter able to consistently predict scattering. This “microwave grain size” is demonstrated to be proportional to the measurable optical grain size and to a new factor describing the chord length dispersion in the microstructure, a geometrical property known as polydispersity. By assuming that the polydispersity depends on the snow grain type only, we retrieve its value for rounded and faceted grains by optimization of microwave satellite observations in 18 Antarctic sites, and for depth hoar in 86 Canadian sites using ground-based observations. The value for the convex grains (0.6) compares favorably to the polydispersity calculated from 3D micro-computed tomography images for alpine grains, while values for depth hoar show wider variations (1.2–1.9) and are larger in Canada than in the Alps. Nevertheless, using one value for each grain type, the microwave observations in Antarctica and in Canada can be simulated from in-situ measurements with good accuracy with a fully physical model. These findings improve snow scattering modeling, enabling future more accurate uses of satellite observations in snow hydrological and meteorological applications.

Plain Language Summary

Satellites are unique tools to observe the snow cover, especially in vast remote areas. Space-borne microwave sensors provide information about snow thickness and other properties, but with large uncertainties due to a poor understanding of how microwaves interact with the snow grains. Additional uncertainties are related to the snow effective grain size, which is a crucial but loosely-defined quantity, difficult to precisely measure in the field. Here, we introduce the concept of “microwave grain size”. This quantity has a clear theoretical definition and can be estimated from the product of the measurable optical grain size and a factor called polydispersity. Over 104 sites in Antarctica and Canada, we test the hypothesis that the polydispersity only depends on snow grain type, an observable quantity. The results show excellent modeling performance and yield polydispersity estimates: small values are found for rounded and faceted grains and high values are for cup-shaped crystals known as depth hoar. We explain these differences by differing degrees of microstructural arrangements. This study paves the way toward an improved use of satellite microwave remote sensing in hydrological and meteorological applications.

1 Introduction

Snow is a random heterogeneous medium composed of ice, air and possibly water and impurities. All its physical properties depend not only on the properties of these constituent materials but also on their geometrical arrangement at the micrometer scale, the so called microstructure (Torquato & Haslach, 2002). This applies in particular to the electromagnetic properties that control the propagation of waves in snow, such as the scattering and absorption coefficients. Scattering in snow is caused by the dielectric contrast between air and ice, and its amplitude highly depends on the length scales of the microstructure. The “snow grain size” is an intuitive property commonly estimated in the field (Fierz et al., 2009). However, it is loosely defined from a geometrical point of view because snow crystals often have very complex shapes, leading to imprecise and subjective measurements. Moreover this single metric is insufficient to fully describe all the length scales. Finding a rigorous mathematical representation of the microstructure

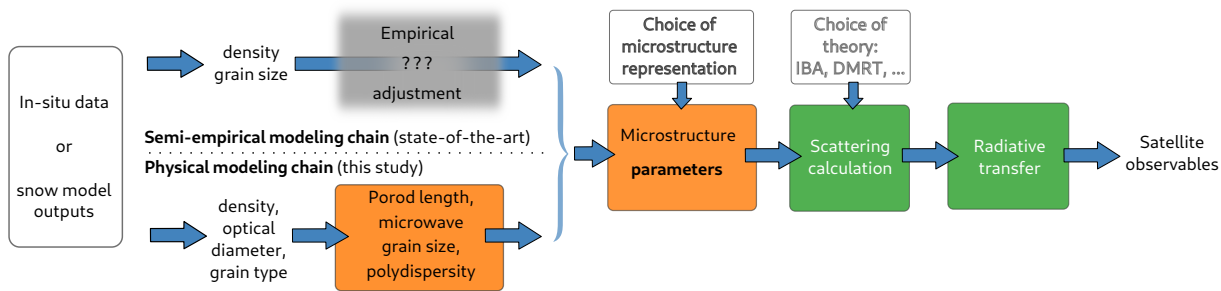


Figure 1. Two modeling chains for predicting satellite observables (brightness temperature, backscattering coefficient) from observed or modeled snowpack properties, 1) using empirical approach to adjust the microstructure parameters (top pathway) and 2) using the new physical pathway (bottom) presented in this study. The gray zone highlights where understanding is lacking, orange where understanding is progressing, green where established solutions provide sufficient accuracy.

73 and prescribing its length scales from actual measurements in the field or from snow evo-
 74 lution model outputs is the biggest problem to be solved for modeling interactions be-
 75 tween snow and electromagnetic waves. This step is crucial to ultimately predict satel-
 76 lite observables from snow measurable quantities (Fig. 1) and conversely for a more re-
 77 liable use of remote sensing to retrieve snow information for hydrological, meteorolog-
 78 ical and climate applications (Helmert et al., 2018; Hirahara et al., 2020). A paramount
 79 application is the retrieval of the snow height and snow water equivalent (snow mass on
 80 the ground), a major endeavour for snow hydrology (Rott et al., 2008; Lievens et al., 2019;
 81 Derksen et al., 2019; Pulliainen et al., 2020).

82 In the visible and infrared spectral range, this problem has been solved a long time
 83 ago by introducing the optical diameter d_{opt} . This stems from early modeling work where
 84 snow was described as a random collection of identical non-overlapping spherical parti-
 85 cles of ice (hard spheres) (Warren & Wiscombe, 1980). The sphere diameter, called the
 86 optical diameter or optical grain size (Wiscombe & Warren, 1980), proved to be a fruit-
 87 ful concept to predict scattering and absorption even when considering more complex
 88 geometrical shapes (Grenfell & Warren, 1999). For any shape, d_{opt} can be defined from
 89 the surface area of the ice/air interface S and ice volume V using $d_{\text{opt}} = 6S/V$. The
 90 relevance of the optical diameter comes from the fact that any medium exhibits a sim-
 91 ilar scattering behavior to hard spheres if the particles have the same S/V ratio (Grenfell
 92 & Warren, 1999), a property called S/V equivalence. This equivalence is not strict, as
 93 the scattering behavior has a small, residual dependence on particle shape (Picard, Ar-
 94 naud, et al., 2009). Hence, all modern snow optical radiative transfer models use the op-
 95 tical diameter to predict the albedo (Domine & Shepson, 2002). Measurements of the
 96 optical diameter and of the related metric called "specific surface area" defined as $\text{SSA} =$
 97 $6/\rho_{\text{ice}}d_{\text{opt}}$, has become considerably easier in the recent decades, with a variety of avail-
 98 able techniques based on adsorption of methane (Legagneux et al., 2002), high-resolution
 99 3D images obtained by micro-computed tomography (Kerbrat et al., 2008), or optical
 100 reflectance (Matzl & Schneebeli, 2006; Painter et al., 2007; Gallet et al., 2009; Arnaud
 101 et al., 2011). The latter is the most convenient technique in the field, and has enabled
 102 the collection of large data sets of SSA (Vargel et al., 2020).

103 In the microwave range, the problem stated above remains largely unsolved since
 104 a quantity equivalent to the optical diameter is hitherto missing. A first possible approach
 105 to represent snow is considering hard spheres (HS) with diameter d_{opt} , as in the visi-
 106 ble and infrared ranges. Results show however that this approach underestimates the

107 scattering amplitude in the microwave range (Brucker, Picard, et al., 2011). Artificially
 108 increasing the sphere diameters by a constant empirical factor was found to be an effective
 109 solution for local-scale studies (Brucker, Picard, et al., 2011; Picard et al., 2014) but
 110 requires specific regional adjustments (Roy et al., 2013; Vargel et al., 2020). This demon-
 111 strates the lack of geometrical insight which prevents generalization and application in
 112 robust satellite retrieval algorithms at large scales. A popular extension of the HS rep-
 113 resentation is the sticky hard sphere (SHS) model (Tsang et al., 1985; Macelloni et al.,
 114 2001; Picard et al., 2014) that aggregates the particles into clusters under the effect of
 115 an ad hoc attractive force (Fig. 2a and 2b). Because the clusters span a wider size than
 116 their constituent particles, the scattering amplitude is enhanced with respect to randomly
 117 positioned spheres. This highlights the prominent role of the inter-particle arrangement
 118 in the microwave range as opposed to the visible and infrared ranges where only the in-
 119 dividual particle size and shape matter. The strength of the attractive force can be ad-
 120 justed by a parameter called stickiness (Tsang et al., 1985). Despite being rigorously de-
 121 fined in terms of the pair potential between the particles, it is impossible to directly re-
 122 late this parameter to actual measurable microstructural properties.

123 A second approach to describe snow is considering a two-phase porous random medium,
 124 which is more general than assuming a particle collection. Such a medium is completely
 125 defined by the indicator function $I(r_0)$, which takes the value 1 when ice is present at
 126 the position r_0 in 3D space and 0 otherwise (Fig. 3a). This binary representation has
 127 been successfully used to interpret data recorded with the small-angle scattering (SAS)
 128 of neutron or X rays techniques allowing the investigation of the structure of many mat-
 129 erials (Schmidt, 1991; Svergun & Koch, 2003). The main reason for this success is that
 130 the scattering amplitude is closely related to the auto-covariance function $\gamma(r)$ of the in-
 131 dicator function (Porod, 1951; Torquato & Haslach, 2002). The auto-covariance mea-
 132 sures how rapidly the indicator translated by some increasing distance r , $I(r_0+r)$, be-
 133 comes different from the original indicator $I(r_0)$ (Fig. 2d and 3b). r is called range or
 134 lag. The auto-covariance captures a combination of both size and arrangement of the
 135 medium structures. A fast decrease of the covariance at the origin ($r \approx 0$) indicates a
 136 medium with small structures (Fig. 2a and d), while a slow decrease indicates large struc-
 137 tures (Fig. 2c and d). Because of the importance of this decrease rate, the Porod length
 138 was introduced early in the development of the SAS technique (Porod, 1951). It is de-
 139 fined as the inverse of the decrease rate at the origin $l_p = -(\gamma'(0))^{-1}$. An important
 140 consequence established by (Debye et al., 1957) is that l_p is mathematically and unequiv-
 141 ocaly related to the optical diameter d_{opt} and density ρ for any porous medium with
 142 a smooth interface, $l_p = \frac{2}{3}(1 - \rho/\rho_{\text{ice}})d_{\text{opt}}$ (Mätzler, 2002). An equivalent expression
 143 is obtained as a function of SSA, $l_p = 4(1 - \rho/\rho_{\text{ice}})/SSA\rho_{\text{ice}}$. This general result
 144 is highly relevant for snow since SSA and density are measurable quantities. Objective val-
 145 ues of l_p can thereby be obtained easily. However, this is insufficient to predict microwave
 146 scattering because the long range behavior of the auto-covariance is important but un-
 147 constrained by l_p . For instance the clustering of spheres, as obtained with the SHS model,
 148 results in a covariance function with a more positive tail (Fig. 2b and d) than when the
 149 spheres are fully randomly positioned (Fig. 2a and d), resulting in stronger scattering
 150 for the same sphere size.

151 For an isotropic medium with randomly distributed and varying shapes, so-called
 152 Debye medium (Debye et al., 1957), the auto-covariance function has a decreasing ex-
 153 ponential form $\gamma(r) = \exp(-r/l_c)$. l_c is called the correlation length and is equal to
 154 the Porod length in this particular case. For this reason, these two lengths have been
 155 often confusingly called "correlation length" in the snow literature (Mätzler, 2002; Royer
 156 et al., 2017). Given that l_p , or likewise l_c , can be derived from measurements, this mi-
 157 crostructure representation has become popular to compute snow scattering (Mätzler
 158 & Wiesmann, 1999). Unfortunately, values of l_c derived from measurements in this way
 159 tend to overestimate scattering in many cases. Empirical scaling factors α have been there-
 160 fore introduced to adjust $l_c = \alpha l_p$ (Mätzler, 2002), which we refer to here as the scaled

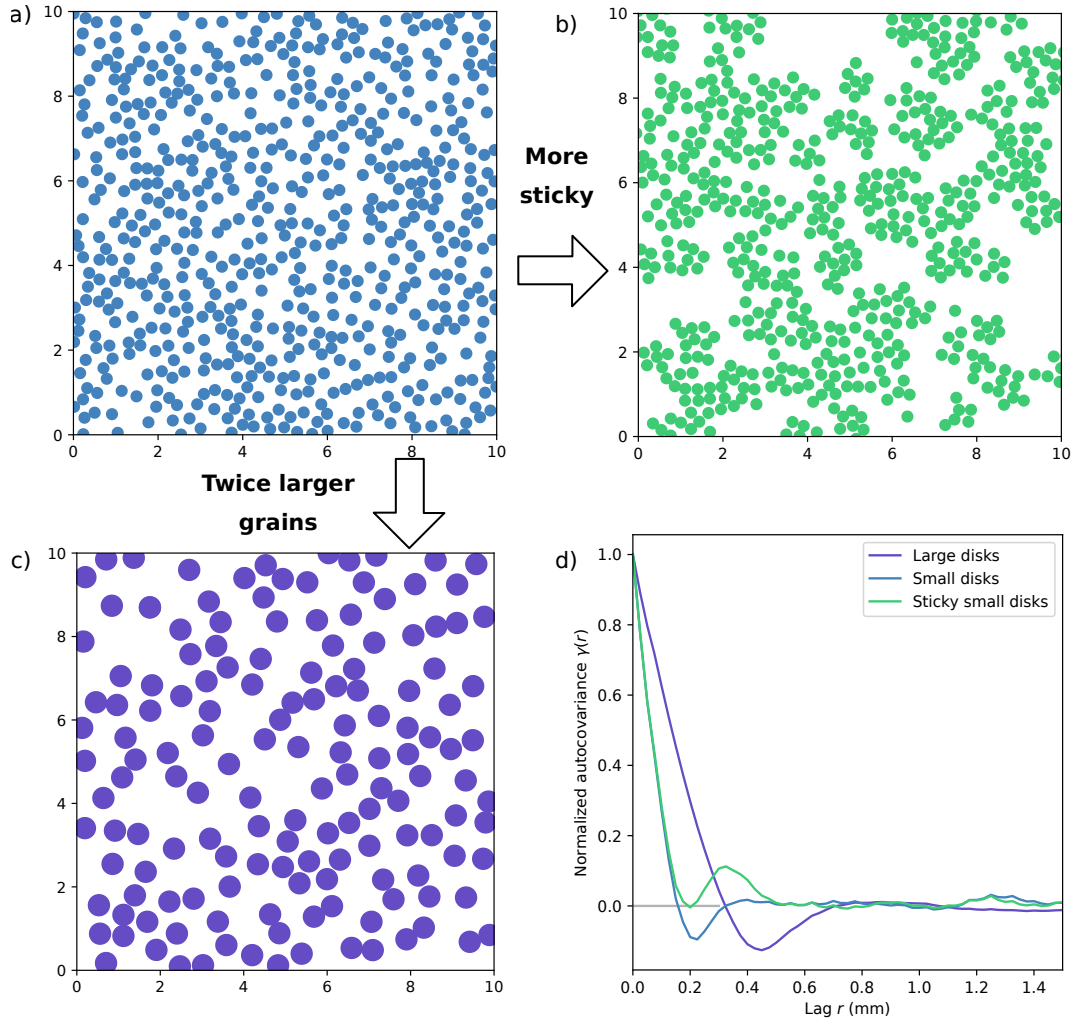


Figure 2. Comparisons of different microstructures composed of non-overlapping disks (in 2D for illustration purpose, (a, b and c)) and the resulting normalized auto-covariance function (d). The shape of the auto-covariance (d) informs about 1) the size of the particles, as it decreases more sharply from the origin small disks (a and b) than for large ones (c), and 2) has a longer or more positive tail for sticky disks (b) than for more randomly positioned disks (a). Both increasing size and stickiness enhances scattering.

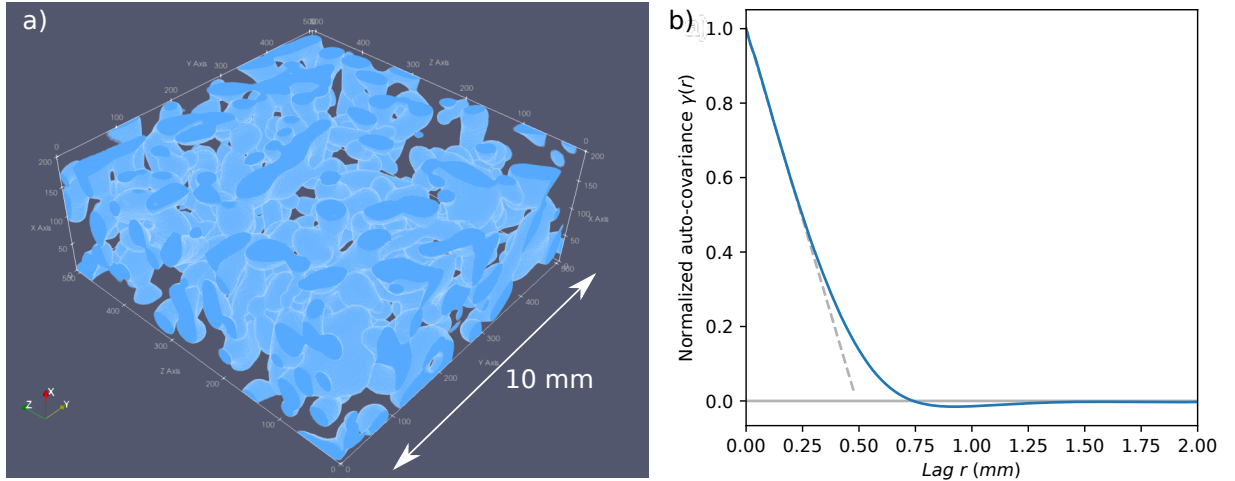


Figure 3. 3D Microstructure of a snow sample composed of alpine rounded grains. a) segmented micro-CT image representing ice (blue) and air (void). b) Normalized auto-covariance of the sample.

161 exponential microstructure (sEXP hereinafter). However this microstructure lacks any
 162 physical justification because the scaling factor has to be adjusted from region to region
 163 (Vargel et al., 2020), in full similarity to the empirical adjustments required for the SHS
 164 model outlined above. Other forms of auto-covariance function have been introduced for
 165 snow, in particular some based on cut-levelled Gaussian Random Fields (Ding et al., 2010;
 166 Sandells et al., 2021). However, they also introduce parameters that are difficult to re-
 167 late to snow physical properties.

168 In summary, the snow-microwave community faces a two-fold problem: first, it is
 169 not yet established which representation of snow microstructure (sEXP, SHS, or another)
 170 is optimal for electromagnetic scattering calculations, and second, the parameters (di-
 171 ameter, stickiness, correlation length) to run the scattering models are often not mea-
 172 surable and need to be adjusted empirically. They are also usually specific to a single
 173 microstructure representation, which causes incompatibility and confusion on how to re-
 174 late them to one another and to measurable quantities.

175 We propose an alternative approach to solve this problem without relying on em-
 176 pirical adjustments (Fig. 1, bottom pathway). Following the concept of the optical di-
 177 ameter, we introduce a new metric called the microwave grain size (l_{MW}) which opti-
 178 mally predicts scattering (Sect. 2). In this paper, we then show that:

179 1) The microwave grain size is a unifying parameter because common microstruc-
 180 ture representations for snow (sEXP, SHS and others) can be reformulated with this new
 181 parameter, the density and the Porod length. Moreover, for a given microwave grain size
 182 value, the scattering amplitude at low frequencies becomes almost independent of the
 183 microstructure representation.

184 2) The microwave grain size is the product of the Porod length and a new intro-
 185 duced factor K that describes how widely the length scales vary in the microstructure.
 186 The so-called polydispersity K carries information on the shape of the particles and their
 187 relative arrangement. Furthermore, we find here that K is fairly constant ($\approx 15\%$ vari-
 188 ations) for a wide range of convex-grained snows, thus providing a physical way to es-
 189 timate l_{MW} from l_{p} and thus from easily measurable quantities in the field (SSA and
 190 density).

191 3) Taking K as depending on the type of snow grains (among classes in a univer-
 192 sally accepted classification, (Fierz et al., 2009)) only is an efficient way to predict mi-
 193 crowave observation from satellites which is comprehensively demonstrated for 104 sites
 194 in Antarctica and in Canada (Sec. 4).

195 This work hence provides a better understanding of the snow microstructure and
 196 a robust way to predict microwave scattering from measurable or observable snow prop-
 197 erties using traceable physical relationships. These new formulations are implemented
 198 in the open-source Snow Microwave Radiative Model. These findings and this model will
 199 help to rigorously link snowpack evolution model outputs to microwave emission and backscat-
 200 ter models inputs, which open great perspectives to improve the retrieval of crucial vari-
 201 ables, such as the snow height or snow water equivalent, using remote sensing.

202 2 Background

203 2.1 The microwave grain size

204 A natural definition of the microwave grain size follows from the Born approxima-
 205 tion (Born, 1926; Mätzler, 1998; Ding et al., 2010), used in several scientific domains (Porod,
 206 1951; Teubner & Strey, 1987; Gille, 2000), and relating the scattering amplitude σ_s to
 207 the auto-covariance function $\gamma(r)$ of a porous isotropic medium:

$$208 \sigma_s = k^4 \varepsilon(k, \rho) \tilde{\gamma}(k) \quad (1)$$

209 Here $k = 2\pi f/c$ is the wavenumber (f is the wave frequency and c is the speed of light)
 210 and ε is an electromagnetic term depending on the density ρ and the wavenumber. It
 211 only shows small variations in the case of dry snow in the frequency range 1–100 GHz
 212 (Löwe & Picard, 2015) and is not a source of uncertainties. The microstructure infor-
 213 mation is carried by $\tilde{\gamma}(k)$, the 3D Fourier transform of the isotropic auto-covariance func-
 214 tion $\gamma(r)$. We introduce the microwave grain size l_{MW} by noting that $\tilde{\gamma}(k)$ has the di-
 215 mension of a cubic power of a length. Taking the static limit ($k = 0$), a simple defini-
 216 tion follows:

$$217 l_{\text{MW}} = \tilde{\gamma}(0)^{1/3} = \left(\frac{1}{2} \int_0^\infty \gamma(r) r^2 dr \right)^{1/3} \quad (2)$$

218 As a direct consequence of this definition, the scattering amplitude is exactly propor-
 219 tional to the cubic power of l_{MW} in the static regime, and Eq. 2 remains a good approx-
 220 imation in the low frequency limit (1– 85 GHz for most snows, as shown in the follow-
 221 ing), as long as the k^4 term dominates the frequency variations over the electromagnetic
 222 and microstructure terms. This implies that knowing l_{MW} solves the problem of the scat-
 223 tering calculation. Nevertheless, no method exists to obtain or measure l_{MW} for snow
 224 yet. In the following we show that this parameter has very relevant properties and we
 225 devise a method for estimating it from measurable quantities.

226 2.2 The unifying role of the microwave grain size

227 We first show how the microwave grain size is related to the specific parameters
 228 of some commonly-used microstructure representations. For this, we use the definition
 229 (Eq. 2) with either the integration of the real space auto-covariance function $\gamma(r)$ or the
 230 Fourier space expression at the origin $\tilde{\gamma}(0)$. The details of the calculations are reported
 231 in supporting Text S1 and Table S1, and we briefly summarize the final results here. For
 232 the scaled exponential representation it is trivial to show by integration of the exponen-
 233 tial function (2) that:

$$234 l_{\text{MW},sEXP} = l_c = \alpha l_p \quad (3)$$

235 The microwave grain size would thus coincide with the widely used "exponential" cor-
 236 relation length (Mätzler, 2002) if the auto-covariance were strictly exponential. It also
 237 appears proportional to the Porod length. Similarly for the SHS microstructure, the mi-
 238 crowave grain size can be related to the sphere radius, density and stickiness. Since l_p

is a function of radius and density, the microwave grain size can also be related to the Porod length as follows:

$$l_{MW,SHS} = K_{SHS}(\tau, \rho)l_p \quad (4)$$

where the function K_{SHS} only depends on the stickiness τ and the snow density ρ (Table S1). This expression shows a clear and natural separation between the size (carried by l_p) and the effect of packing of the spheres (carried by K_{SHS}). Another widely used auto-covariance function was proposed by Teubner and Strey (TS representation hereinafter) for microemulsions of oil in water (Teubner & Strey, 1987). It has been previously introduced for snow (Löwe et al., 2011; Picard et al., 2018; Sandells et al., 2021). Here again a simple relationship can be obtained for TS:

$$l_{MW,TS} = K_{TS}(q)l_p \quad (5)$$

where we use the dimensionless parameter $q = l_p/d_{TS}$ to define the TS representation instead of the repeat distance d_{TS} as suggested by (Ruland, 2010). In contrast with these examples, not all the microstructure representations have closed analytical form yet. This is the case of the Gaussian Random Field microstructures as defined in (Ding et al., 2010) or (Sandells et al., 2021).

These three examples highlight that l_{MW} can be computed for different analytical forms of the auto-covariance function and related to the specific parameters of these forms. It is not guaranteed that an analytical expression of l_{MW} always exists for any microstructure representation, since the integration of $\gamma(r)$ may not be carried out in closed form. However, when they exist, these relationships make the different microstructure representations comparable. For instance in a study using the SHS representation, l_{MW} can be calculated from the radius, density and stickiness, and can then be compared to another study using the scaled exponential representation where l_{MW} is simply related to the scaling coefficient and the Porod length. This provides a way to re-evaluate past studies.

Noting the central role of the microwave grain size, we went a step further and inverted these relationships to obtain all specific parameters of the considered microstructure representations (sEXP, SHS and TS) as a function of the triplet microwave grain size l_{MW} , Porod length l_p and snow density ρ only (the equations are reported in supporting Text S2 and Table S2). The fact that such a common set of variables exists is an important and new result because it provides a unified way to parametrize different microstructure representations. Furthermore, given the definition of the microwave grain size, it is guaranteed that different microstructure representations predict the same scattering amplitude in the low frequency limit when the same microwave grain size value is used as input. Only at higher frequencies some differences between the microstructure representations may appear for a given microwave grain size, but the re-parameterization in terms of common triplet remains nevertheless effective and relevant. This result renders the choice of the best snow representation a secondary problem, and conversely implies that measuring the microwave grain size or deriving its value from measurable quantities become the primary task to be solved in order to predict snow scattering in the microwave range.

2.3 The microwave grain size from measurable quantities

To obtain the microwave grain size from measurable quantities, it is necessary to reveal its fundamental link to geometrical characteristics of the microstructure. We established here a relationship between the microwave grain size and the chord length distribution (CLD), independent of the particular choice of the auto-covariance functional form.

Chords are line segments intersecting an infinite line with the two phases of a porous medium (Fig. 4a) (Torquato & Haslach, 2002). The CLD of each phase is a statistical

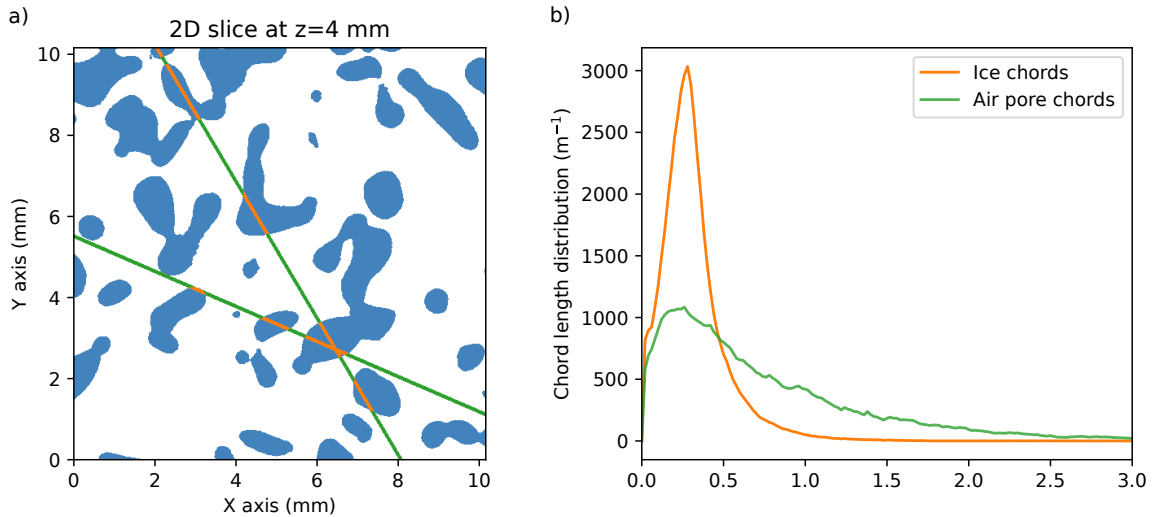


Figure 4. a) A 2D slice of the 3D Microstructure shown in Fig. 3a with examples of chords in the air (orange segments) and ice (green segments). b) The distribution of the lengths of all the chords in the 3D samples for the air and ice.

289 characteristics of the microstructure (Fig. 4b). Intuitively, the average chord length μ_1
 290 of the ice (and air) gives information on the size of the ice grains (and of the air pores
 291 respectively). It is related to l_p for any microstructure by: $\mu_1 = l_p/(1-\phi)$ where ϕ is
 292 the fractional volume of the ice phase, i.e. the ratio between snow density ρ and pure
 293 ice density $\rho_{ice}=917 \text{ kg m}^{-3}$ (Ruland, 2010). It can be therefore unequivocally estimated
 294 from the measurable SSA and density. The higher order moments ($\mu_i, i > 1$) of the CLD
 295 carry information on how dispersed the chord lengths are, that is, all higher order mo-
 296 ments have a small value only if all the chords have similar lengths, and a high value for
 297 complex shapes or a large range of sizes. This property is called polydispersity. The sec-
 298 ond moment μ_2 was previously used to parametrize a generalized version of the TS mi-
 299 crostructure model (Ruland, 2010). An important conclusion of this study was that the
 300 scattering amplitude increases not only as a function of size (the first order moment) but
 301 also with polydispersity (the higher order moments). We adapt and generalize here this
 302 idea to a wide class of microstructures. To this end we establish a general relationship
 303 linking the microwave grain size and the chord length moments. This is achieved in two
 304 steps (details of the calculation are given in the supporting text S3), first by relating the
 305 microwave grain size to the second derivative of the Laplace transform of the auto-covariance
 306 function $\hat{\gamma}(s)$:

$$307 \quad l_{\text{MW}} = \left(\frac{1}{2} \int_0^\infty r^2 \gamma(r) dr \right)^{1/3} = \left(\frac{1}{2} \hat{\gamma}''(0) \right)^{1/3} \quad (6)$$

308 and second by using an approximation for the Laplace transform of the CLD established
 309 in a previous study (Roberts & Torquato, 1999). This approximation assumes that the
 310 pore chord distribution is exponential, a class of microstructures known as the Boolean
 311 model (Bilodeau et al., 2007). In such a model, the solid phase is built up by randomly
 312 positioning a finite set of primary shapes (e.g. spheres, cubes, polyhedra, etc., or any com-
 313 bination of them) in space with possible overlap. In such a model, the pore CLD is ex-
 314 ponential if the primary shapes are all convex (Bourgeois & Lyman, 1997). We verified
 315 how close to an exponential is the pore CLD for 167 snow samples collected in the Alps
 316 from in-lab snow growth experiments (Fig. S1) and concluded that the Boolean model
 317 applies well to snow. Note that even with convex primary shapes, the resulting microstruc-

318 ture has concave parts, as it is common in snow (depth hoar, grain boundaries), because
 319 overlaps are allowed in the Boolean model.

320 The relationship obtained after applying these two steps links the microwave grain
 321 size to the first four moments of the ice CLD:

$$322 \quad l_{\text{MW}} = Kl_{\text{p}} \quad (7)$$

$$323 \quad K = \left(\frac{\mu_4}{24\mu_1^4} - \frac{\mu_2\mu_3}{6\mu_1^5}\phi + \frac{\mu_2^3}{8\mu_1^6}\phi^2 \right)^{1/3} (1-\phi)^{-2/3} \quad (8)$$

325 Here we only consider snows with $\phi < 0.5$ (density less than 468 kg m^{-3}) because for
 326 $\phi > 0.5$ it is recommended to swap air and ice (air primary shapes in an ice background)
 327 (Dierking et al., 2012). This relationship highlights the proportionality between the mi-
 328 crowave grain size and the Porod length through the factor K . We call this latter fac-
 329 tor the "microwave polydispersity" because it only involves ratios between the higher
 330 order moments and the first order moment of the CLD, and thus measure the chord poly-
 331 dispersity. As opposed to (Ruland, 2010) we demonstrate here that the second order mo-
 332 ment is insufficient to fully characterize the polydispersity as relevant to microwave scat-
 333 tering, the first four moments are all required.

334 Before studying this equation in its general form, it is instructive to consider the
 335 case $\phi = 0$, a medium with isolated grains and very low density even though it does
 336 not apply to snow. The microwave polydispersity of such a sparse medium writes:

$$337 \quad K_{\text{sparse}} = \left(\frac{\mu_4}{24\mu_1^4} \right)^{1/3} \quad (9)$$

338 It only depends on the first and fourth moments of the ice CLD which can be related
 339 to the volume V and surface area S of the particles using the Cauchy formula (Mazzolo
 340 et al., 2003), leading to:

$$341 \quad K_{\text{sparse}} = \frac{S}{8\pi^{1/3}V^{2/3}} \quad (10)$$

342 This equation offers a practical means to compute the microwave polydispersity for any
 343 geometrical particle with known surface area and volume (when the medium is sparse).
 344 Moreover, it gives an intuitive understanding of the polydispersity by noting that the
 345 ratio $\frac{S}{V^{2/3}}$ is related to the isoperimetric shape factor $f_1 = 6V/\pi^{1/2}S^{3/2}$, a common mea-
 346 sure of sphericity of particles (Redenbach et al., 2012). f_1 indeed takes its highest pos-
 347 sible value for spheres and decreases with the particle elongation. The microwave poly-
 348 dispersity K_{sparse} is proportional to $f_1^{-2/3}$, implying that spheres are the least efficient
 349 scatterers, and the scattering amplitude increases with elongation for a given Porod length.
 350 This result may explain why representing snow as non-overlapping ice spheres usually
 351 underestimates scattering and that large empirical scaling factors had to be used in the
 352 past to reconcile model simulations and observations (Brucker, Picard, et al., 2011; Roy
 353 et al., 2013; Picard et al., 2014). To conclude for sparse media, the microwave grain size
 354 can be interpreted as the product of an elongation indicator (K_{sparse}) and the particle
 355 size (l_{p}).

356 In the case of dense media such as snow, the polydispersity given by Eq. 8 involves
 357 two additional terms in ϕ and ϕ^2 , with a more complex combination of CLD moments.
 358 Furthermore, the second and third moments cannot be related to S and V only. Despite
 359 this complexity, the formulation provides several hints. First, it confirms the idea of (Ruland,
 360 2010) about the influence of the chord polydispersity on scattering. Second, it shows that
 361 the polydispersity tends to decrease with increasing density (the first order term in ϕ
 362 is negative and $(1-\phi)^{-2/3}$ decreases with increasing ϕ) at least for moderate densities
 363 ($\phi^2 \approx 0$). This implies that the microwave polydispersity recovers a well known and
 364 important effect in dense packings, where the scattering amplitude of packed particles
 365 is lower than the sum of individual particle scattering (Tsang & Kong, 2001). And last,

Eq. 8 allows us to estimate the polydispersity value from the CLD, which itself can be obtained from micro-CT imaging of real snow. This equation hence provides a means to obtain the polydispersity and then the microwave grain size from measurable quantities.

3 Materials and Methods

3.1 Micro-CT Data and Chord Length Distribution

The dataset used here to compute CLD was first presented in (Löwe et al., 2013). It comprises 167 snow samples scanned with X-ray tomography (micro-CT), producing 3D images at a resolution ranging from 5.1 and 10.7 μm . The samples are in fact of two categories: 37 of them are individual samples collected in the Alps while the remaining was obtained by sampling at different times from 6 in-lab snow maturation experiments. These experiments differ from each other by the imposed thermal gradient conditions, from isothermal to 100 K m^{-1} . All the samples were assigned to a snow type (depth hoar, rounded grains, faceted crystals, decomposing and fragmented precipitation particles, melt forms and precipitation particles) according to the international classification of seasonal snow on the ground (Fierz et al., 2009). The dataset is therefore quite heterogeneous and is not representative of any snow on Earth, but is adequate to illustrate the effect of snow types on polydispersity. The CLD of ice and air was extracted from each 3D image after binarization, by drawing lines in the vertical and two perpendicular horizontal directions as presented in (Krol & Löwe, 2016).

3.2 Snow In-situ Measurements

In-situ measurements were collected in Antarctica and Canada to compute microwave grain size and perform the microwave simulations. In Antarctica, snow properties were measured at 18 sites (Table S3) over a large range of latitudes during three scientific traverses, namely Vanish (2011-2012), ASUMA (2016-2017) and EAIIST (2019-2020). Additional measurements were taken in 2011 at Dome C (Picard et al., 2014). A relatively similar protocol was applied at every site. A borehole was drilled up to a depth of typically 8 m (4.1–17.9 m). The extracted core was sliced in ≈ 10 cm long pieces. Snow density was obtained by measuring the diameter, height and mass of each cylindrical slice. If a slice was not cylindrical, the height was recorded, and the density was set to that of the nearest cylindrical slice. The Specific Surface Area (SSA) profile was measured by short-wave infrared reflectometry using the Posssum and Asssap instruments (Arnaud et al., 2011; Libois et al., 2015). On ASUMA and EAIIST, Asssap was used to record the SSA profile along each extracted snow core of 50–100 cm length. The profiles were then assembled and the small gaps between each core were filled by linear interpolation. The profile resolution is about 1 cm. At Dome C and at sites S2b and S4 on Vanish, Asssap was used to take a single record for every 10 cm slice in a cold chamber in France (Picard et al., 2014). At point S2 on Vanish, Posssum (Arnaud et al., 2011) was directly used in the borehole to record the full profile at 1 cm resolution. This profile is however short, only 4.9 m. Both instruments, Posssum and Asssap are based on the same principle and have been inter-calibrated many times. Their accuracy was estimated to 15% against independent SSA measurements (Arnaud et al., 2011). Since the measurements from a single borehole were used for each simulation that was then compared to satellite observations representative of a 12.5 km (or 25 km) wide pixel, it is expected that the intra-pixel spatial variability is a large source of uncertainties. This prevents performing a very precise site-by-site comparison between simulations and observations. The annual mean temperature at each site was measured with a Pt100 sensor at 10 or 20 m depths, after 24h stabilization. The complete temperature profile was not recorded because of its changing nature. For this reason, the simulations are conducted with an uniform temperature equal to the temperature measured at 10 or 20 m depth, and they are

416 compared with the annual average brightness temperature. The density and SSA pro-
 417 files are also considered independent of time. This approximation is valid because due
 418 to the cold conditions and the low accumulation, the main rapid changes only occur in
 419 the topmost ≈ 20 cm of the snowpack, the remaining being stable over years on the Antarc-
 420 tic Plateau. This is applied to the low frequencies (10, 19 and 37 GHz) where the snow-
 421 pack portion contributing to emitted signal is larger than about one meter depth, and
 422 this is mathematically justified by the quasi-linearity of the temperature dependence in
 423 the heat equation and the radiative transfer equation in snow (Picard, Brucker, et al.,
 424 2009). In contrast at 89 GHz, because the radiation is emitted by the topmost ≈ 20 cm
 425 of the snowpack, and the snow properties were measured in summer, the simulations use
 426 an uniform temperature equal to the mean December-January 2 m air temperature ex-
 427 tracted from the ERA5 reanalysis. The results are compared with the average bright-
 428 ness temperature over the same months.

429 In Canada, 86 sites (Table S4) were sampled over a large latitudinal range. The
 430 density, SSA and temperature profiles were measured in snowpits down to the ground
 431 as detailed in (Vargel et al., 2020). The density was measured with a density cutter and
 432 a scale. The SSA was measured on samples extracted from the pits using the IRIS in-
 433 strument (Montpetit et al., 2012) based on short-wave infrared reflectometry as Asssap
 434 and Posssum. The profiles of temperature and the soil temperature were recorded for
 435 each pit.

436 3.3 Microwave Simulations

437 The Snow Microwave Radiative Model (SMRT) (Picard et al., 2018) is used to con-
 438 duct the simulations of microwave thermal emission. The model represents the snowpack
 439 as a stack of horizontal layers specified with the in-situ properties as follows. The lay-
 440 ering is directly derived from the density profile. In Antarctica, because some profiles
 441 are too short (e.g. S2 on Vanish) with respect to the microwave penetration depth at
 442 the lowest frequency (10 GHz), the modeled snowpack is extended down to 30 m depth
 443 by repeating the lower meter of the measured profile. The SSA which is usually sampled
 444 with a higher resolution than density is averaged for each density layer. The Porod length
 445 l_p is then deduced from density and SSA. In Canada, with a snow height rarely exceed-
 446 ing 1.5 m, the soil is a significant contributor to the microwave signal at low frequencies
 447 (10 GHz and 19 GHz) and certainly plays a small role at 37 GHz as well. Unfortunately
 448 the soil characteristics relevant to microwave simulations (soil permittivity, surface rough-
 449 ness, ...) are in general difficult to measure, and were not available here. This problem
 450 was solved by optimization of the soil parameters by (Vargel et al., 2020) using the ob-
 451 servations at low frequencies. We have taken here the soil parameters of that study with-
 452 out any further adjustment.

453 To explore the role of microstructure representation, SHS, sEXP and TS are con-
 454 sidered for most simulations. The original version of TS is limited to $K < 1$ (Teubner
 455 & Strey, 1987) but has been extended proposed by (Ruland, 2010). The latter is imple-
 456 mented in SMRT as detailed in the supporting Text S2.

457 The other settings of SMRT are common to previous studies (Picard et al., 2018;
 458 Vargel et al., 2020). In short, the Improve Born Approximation (IBA) is used to com-
 459 pute the scattering and absorption coefficients in each layer and the Discrete Ordinate
 460 method (DORT) solves the radiative transfer equation for the whole snowpack account
 461 for multiple scattering between the layers. The outputs for each site are the brightness
 462 temperature at four frequencies and at horizontal and vertical polarizations.

3.4 Microwave Observations

Microwave observations were compared to the model simulations in order to retrieve the polydispersity and assess the simulation performance. In Antarctica, the microwave brightness temperature observations at 10, 19, 37 and 89 GHz were recorded by the Advanced Microwave Scanning Radiometer 2 (AMSR2) sensor onboard Japan’s Global Change Observation Mission 1st - Water “SHIZUKU” (GCOM-W1) satellite. We extracted the observations at the nearest pixel of each site (Table S3) from the National Snow and Ice center (NSIDC) AMSR-E/AMSR2 Unified Level 3 daily product, version 2. The product has a resolution of 25 km at 10 GHz and 12.5 km at the higher frequencies. The observations were averaged over the period 2013–2019. The typical brightness temperature accuracy is ± 1.5 K.

In Canada, the observations at 86 sites (Table S4) were obtained with ground-based radiometers operating at the same frequencies as AMSR2 (Vargel et al., 2020) though not all the frequencies and polarizations were observed at all sites due to instrumental failure or availability. The accuracy is typically 2 K. These data mainly differ from satellite data by the small field of view of the sensor which is at the meter scale, and is coincident with the snow properties measurements.

4 Results

4.1 Polydispersity of Snow Samples

Fig. 5 shows the microwave polydispersity K for the 167 samples taken in the field in the Alps or from in-lab snow growth experiments and imaged by micro-CT. The graph distinguishes two categories of snows as a function of the grain shape, with convex grains on the one hand and depth hoar on the other hand. Convex grains include rounded and faceted grains (typical of alpine dry snow) and melt forms (occurring during melt). There are grouped together because their respective mean polydispersity is 0.72 ± 0.084 (1σ , $n=53$) for rounded grains, 0.71 ± 0.073 ($n=33$) for faceted grains, 0.68 ± 0.028 ($n=5$) for melt forms, showing no significant differences (pair-wise Welch’s t-test, $p > 0.05$). Depth hoar, also known as cups because of their hollow shape, features higher values 0.85 ± 0.081 ($n=62$) than the other grains, with a significantly different mean ($p \ll 0.05$). Meanwhile, we note that our values compare well with values (0.8 – 1.2) obtained for the Boreal Finnish depth hoar in a recent investigation (Leinss et al., 2020) where the empirical scaling factor of the exponential function was determined using micro-CT images (according to our equation 3 the polydispersity K is equal to this empirical factor α).

These first results obtained with micro-CT images show that K spans a relatively narrow range 0.71 ± 0.078 (1σ , $n=91$) for rounded grains, faceted grains and melt forms, if compared to the ≈ 10 -fold potential range of variation of l_p . This result suggests that when micro-CT measurements are not available, running microwave simulations with a constant value of K for these convex grain shapes could be sufficient. The next section tests this hypothesis.

4.2 Retrieval of the polydispersity from microwave observations and in-situ data

We use the microwave grain size l_{MW} to predict snowpack microwave emission first in Antarctica and second in Canada. The in-situ measurements provide the profiles of SSA and density, from which l_p can be deduced without approximation. Since no coincident micro-CT measurements were taken, we deduce l_{MW} by assuming that the polydispersity K is constant (but unknown) for the rounded and faceted grains, the prevailing snow grain types on the Antarctic plateau. These measurements and derivatives, along with an assumption on the microstructure representation, are sufficient to fully prescribe

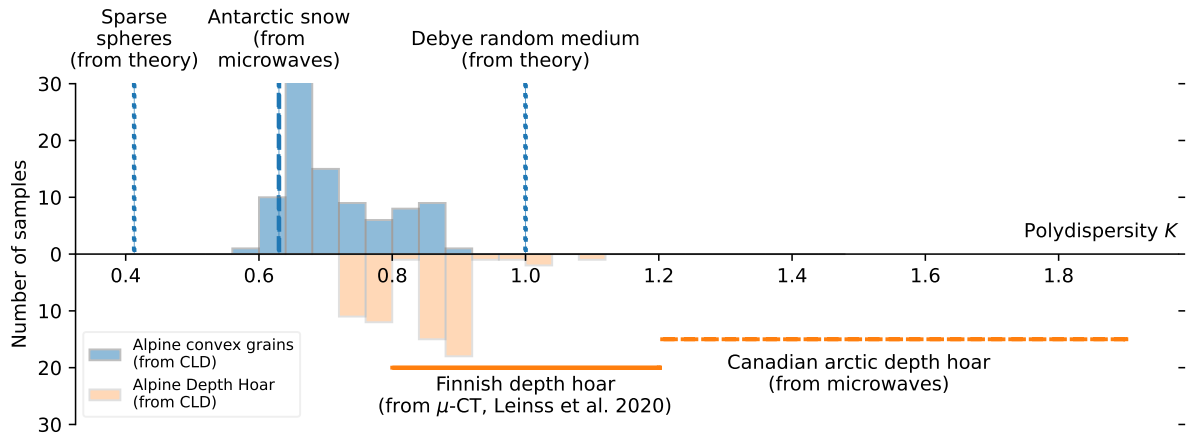


Figure 5. Polydispersity K of convex grains (blue) and depth hoar (orange), i) calculated for 167 snow samples using micro-computed images (blue and orange vertical bars) ii) obtained in this study from theory (vertical dotted lines) or microwave retrieval (blue vertical and orange horizontal dashed lines), iii) and derived from (Leinss et al., 2020) which use micro-computed images (horizontal solid line). The vertical lines are used for values determined with a reasonable accuracy, while the horizontal lines are used when a wider range of values is determined.

511 the microstructure in every snow layer. We performed the SMRT simulations for three
 512 different microstructure representations (sEXP, SHS, and TS) parameterized with the
 513 unifying triplet $l_{\text{MW}} = Kl_p, l_p,$ and ρ . For each microstructure representation, the
 514 optimal K value was determined by minimizing the root mean square error (RMSE) cal-
 515 culated between the simulated and observed brightness temperatures at 19 and 37 GHz
 516 and at vertical polarization (Fig. 6a). We then test the simulations with the optimal K
 517 on a wider set of frequencies (10, 19, 37, 89 GHz) and at both vertical and horizontal po-
 518 larizations (Fig. 7 and S2).

519 The RMSE calculated at two frequencies and vertical polarization features a clear
 520 minimum, as a function of K , of 5.8 K, 5.7 K and 6.2 K for sEXP, SHS and TS respec-
 521 tively (Fig. 6a). When the simulations with the optimal K are run at the four frequen-
 522 cies and two polarizations, the average RMSE is 11.4 K, 11.3 K, 11.79 K for sEXP, SHS
 523 and TS respectively (Fig. 7). Both results show small differences in performance between
 524 the microstructure representations. This reflects past findings where different microstruc-
 525 ture representations have been used with equal success (Royer et al., 2017; Vargel et al.,
 526 2020). This is an expected outcome of the microwave grain size definition as discussed
 527 above. Split per frequency, the RMSE is the lowest at 37 GHz, and increases at 19, 10
 528 and 89 GHz (Fig. 4a). We attribute these variations mainly to the in-situ measurement
 529 uncertainties, and the difference of scale between the in-situ and the satellite measure-
 530 ments. At 10 GHz and 19 GHz, the microwave emanates from the surface to about 15-
 531 20 m and 5-10 m depth respectively, whereas the measurements were taken up to only
 532 ≈ 8 m on average (Table S3). Even though we extended the simulated snowpack down-
 533 ward (Sec. 3.2), this is approximate and may explain part of the uncertainties in the re-
 534 sults at the two lowest frequencies. Conversely, at 89 GHz, the microwaves emanate from
 535 the top 20 cm of the snowpack. This zone was sampled for all the cores but with a ver-
 536 tical resolution of 10 cm that is too coarse for this high frequency. The frequency 37 GHz
 537 is optimal given our experimental sampling, with waves mainly coming from the upper-
 538 most first meter, where accurate and detailed measurements were taken at all sites. Re-
 539 garding polarization, the performance is better in vertical polarization (blue in Fig. 6)

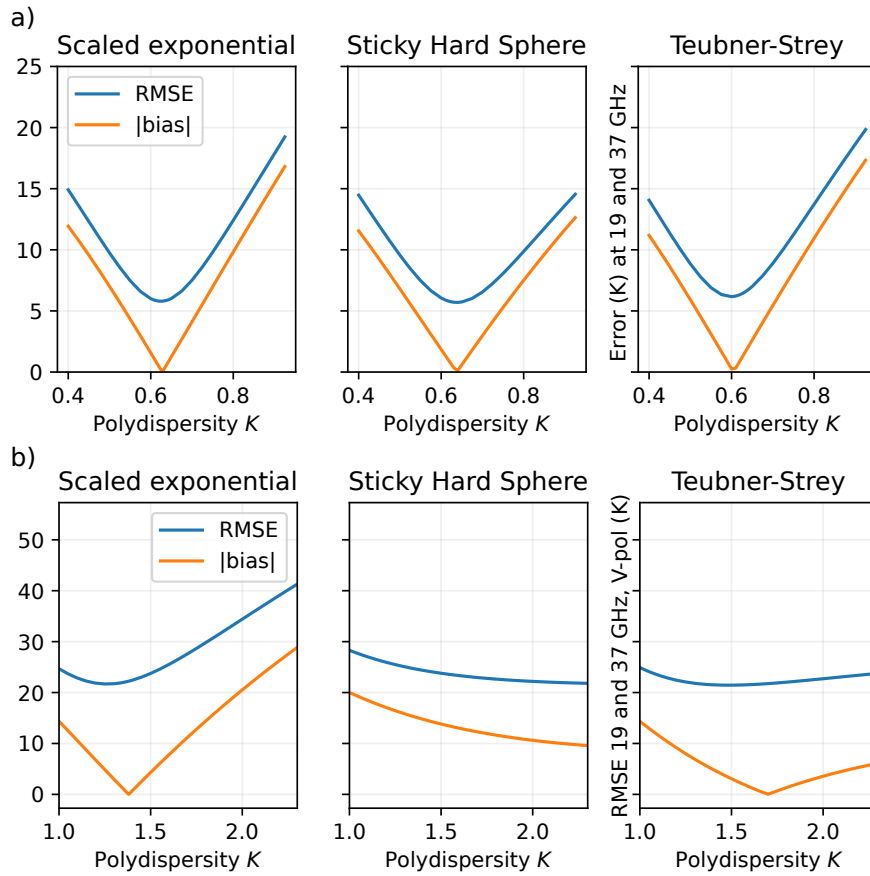


Figure 6. RMSE and bias between simulations and observations calculated at 19 and 37 GHz, at vertical polarization, as a function of the polydispersity value a) applied to all Antarctic sites over the whole profiles and b) applied to all Canadian sites for the depth hoar layer.

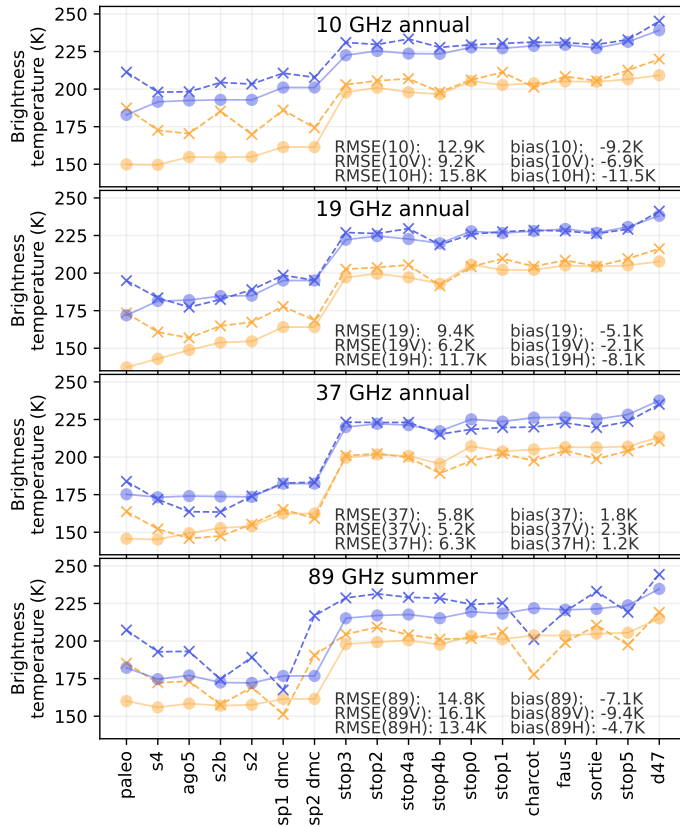


Figure 7. Observed (cross) and simulated (circle) brightness temperatures at four frequencies and vertical (blue) and horizontal (yellow) polarizations at 18 sites in Antarctica (sorted from the inner plateau to the coast, Table S3) using sticky hard spheres and the optimal polydispersity of 0.64.

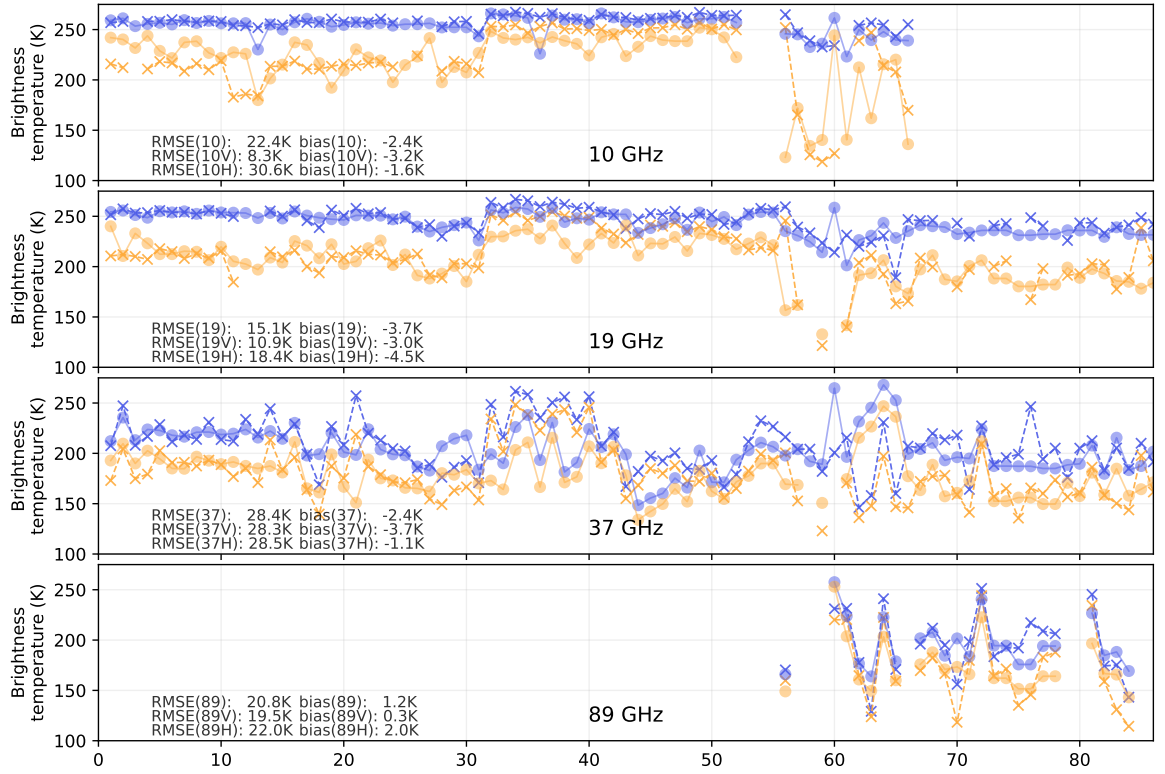


Figure 8. Observed (cross) and simulated (circle) brightness temperatures at four frequencies and vertical (blue) and horizontal (yellow) polarizations at 86 sites in Canada using the Teubner and Strey microstructure. Sites are listed in Table S4.

540 than in horizontal polarization (yellow), which is a classical result (e.g. (Durand et al.,
 541 2008; Wójcik et al., 2008)) explained by the insensitivity of the vertical polarization to
 542 the snowpack density layering. Overall the modeling errors are of the same order as in
 543 other studies where optimizations were applied (e.g. (Picard et al., 2014; Macelloni et
 544 al., 2001)), and the model shows excellent skills to reproduce the latitudinal gradient show-
 545 ing brightness temperatures increasing from the plateau to the coast. We conclude that
 546 assuming a unique constant K for rounded and faceted grains is suitable to predict the
 547 microwave signal in Antarctica, given the uncertainties in the in-situ measurements and
 548 the difference of scale between the satellite and in-situ observations.

549 The optimal polydispersity value is 0.63, 0.64, 0.60 for sEXP, SHS and TS respec-
 550 tively (Fig. 4a). These values fall in the lower range of K obtained from micro-CT on
 551 the alpine rounded and faceted grains (Fig. 5). This result is remarkable because the two
 552 estimates are fully independent, providing for the first time a link between the microwave-
 553 optimized scaling factor and its microstructural origin. Furthermore, the three optimal
 554 K values for the different microstructures are close to one another (within 7%) which
 555 comes from the unifying character of the microwave grain size. While this Antarctic dataset
 556 provides a first confirmation that a constant K is suitable for microwave simulations, the
 557 variety of grain types is limited, only rounded and faceted grains are present on the Antarc-
 558 tic Plateau.

559 We further test our hypothesis on the Canadian dataset where highly metamor-
 560 phized snow is omnipresent as depth hoar. The typical eastern Canadian Arctic snow-
 561 pack consists of an upper part of rounded or faceted grains overlying a bottom part of
 562 depth hoar. For the upper part, we make and test the hypothesis that the optimal K
 563 value obtained in Antarctica also applies in the Canadian environments. In contrast, to
 564 account for the particular scattering efficiency of the depth hoar in the lower part, we
 565 consider a specific value for depth hoar (K_{DH}). This value is obtained by optimization
 566 as done previously, by minimizing the difference between simulations and observations
 567 at 19 and 37 GHz in vertical polarization.

568 The RMSE calculated at two frequencies and vertical polarization shows a min-
 569 imum, with values of 22.7 K, 20.7 K and 21.6 K K respectively for sEXP, SHS and TS (Fig.
 570 6a). These three values are of the same order. However for SHS, the minimum is not marked
 571 and the bias never reaches 0 K, mainly due to a systematic overestimation of the sim-
 572 ulated brightness temperature at 37 GHz. The simulations also become numerically un-
 573 stable (diagonalisation error in the DORT solver in SMRT, (Picard et al., 2018)) for large
 574 polydispersity at 89 GHz preventing the exploration of polydispersity values above 2.3.
 575 SHS appears to be unable to cope with high polydispersity and to produce strong enough
 576 scattering for a given sphere size. This shows the limit of the sphere model even with
 577 highly clustered particles. We conclude that SHS is unsuitable for depth hoar, in line with
 578 past studies (Löwe & Picard, 2015; Vargel et al., 2020). The results in brightness tem-
 579 perature (Fig. 8 and S2) again show the good skills of the model. The simulations at
 580 89 GHz, showing virtually no bias (e.g. 1.2 K for TS, p-value of 0.7), are very insight-
 581 ful because only the upper layer contributes at this high frequency and these simulations
 582 are therefore independent of the K_{DH} optimisation. This confirms that the polydis-
 583 persity estimated in Antarctica and used without adjustment in the upper layer here ap-
 584 plies well to the rounded and faceted grains in Canada. In contrast, the lower frequen-
 585 cies are sensitive to the depth hoar layer, and do depend on the optimized K_{DH} value.
 586 The optimal value only weakly depends on the microstructure representation choice, 1.25
 587 and 1.5 for sEXP and TS respectively. However, the determination is relatively impre-
 588 cise as shown by the wide minima in Fig. 6b. This is particularly true for TS, the RMSE
 589 changes by less than 1 K over the range 1.2–1.9. If instead of the RMSE minimum, we
 590 consider a null bias as an optimization criterion, we would obtain optimal polydispers-
 591 ity of 1.4 and 1.7 for sEXP and TS respectively. Despite these uncertainties, we con-
 592 clude that the optimal K_{DH} is certainly above 1, which is significantly higher than the

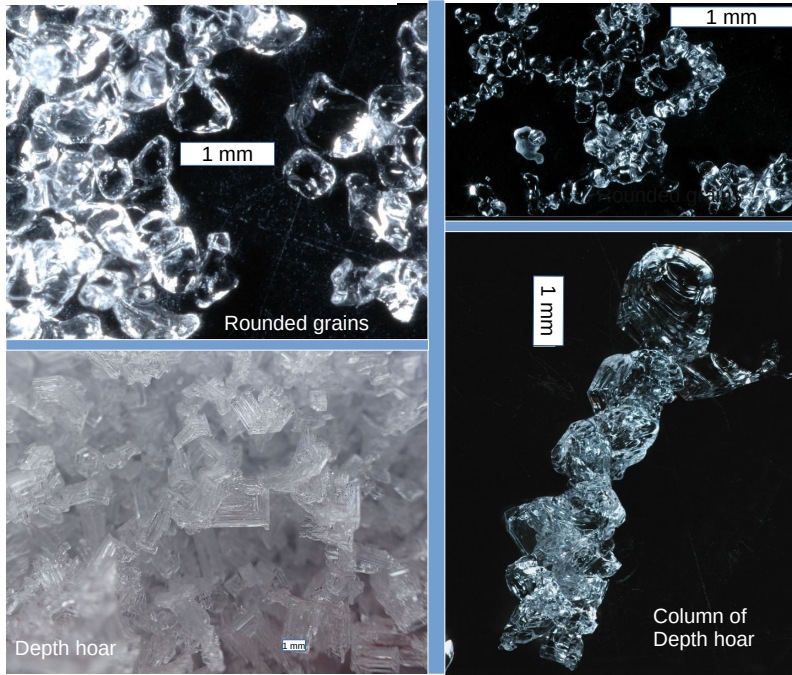


Figure 9. Macrophotography of rounded grains and depth hoar. The white bar indicates the 1 mm scale.

593 polydispersity of rounded and faceted grains. It is also significantly higher than the alpine
 594 depth hoar polydispersity estimated from micro-CT. This suggests that scattering by the
 595 depth hoar in Canada is much stronger than that in the Alps (for a given l_p). Obser-
 596 vations indicate that the structure of hoar is indeed different between these regions (Domine
 597 et al., 2016; Satyawali & Schneebeli, 2010). The eastern Canadian Arctic depth hoar is
 598 often of centimeter size and is more developed due to the very strong vertical temper-
 599 ature gradient prevailing during the entire winter season. The high polydispersity could
 600 thus be explained by the large ratio between the micrometer scales (the steps and the
 601 thin walls of the depth hoar crystals) and the centimeter size of the crystals or even the
 602 long range organisation between the crystals as in columnar depth hoar (Fig. 9). In the
 603 Alps, depth hoar is often tinier and less structured because the thermal gradients are weaker
 604 and operate over a shorter period (mostly the beginning of the snow season) which jus-
 605 tifies a smaller polydispersity.

606 5 Discussion and Conclusion

607 This study establishes a fully tractable chain of physical links to conduct simula-
 608 tions of microwave scattering from measurable snow physical properties (Fig. 1). For each
 609 snow layer, density and specific surface area (SSA) or optical diameter d_{opt} provide the
 610 Porod length l_p which is then converted to the microwave grain size l_{MW} by multipli-
 611 cation with the microwave polydispersity K . We showed that assigning a constant value
 612 to K depending on the traditional grain shape leads to satisfactory simulations. An op-
 613 timal value of ≈ 0.6 for rounded, faceted and melt forms and 1.2-1.9 depending on the
 614 microstructure representation for depth hoar in the eastern Canadian Arctic was obtained.
 615 The confidence in these values is relatively high for the former group, composed of con-
 616 vex grains, because we obtained a similar estimate with two independent methods (CLD
 617 direct calculation and microwave retrieval). However, this comparison was not performed
 618 at the same site, because of the lack of coincident micro-CT and microwave observations.

619 For depth hoar, the value is more uncertain, but it is certainly much larger than that
 620 of the rounded and faceted crystals, which can be understood by the morphological dif-
 621 ferences (Fig. 9) and the possibly wider range of structure in depth hoar (e.g. soft depth
 622 hoar, indurated depth hoar from wind slabs, indurated depth hoar from melt freeze lay-
 623 ers, columnar depth hoar) (Domine et al., 2018).

624 An immediate application of this new tractable chain is to perform microwave sim-
 625 ulations with the outputs from state-of-the-art snowpack models such as CROCUS (Vionnet
 626 et al., 2012) and SNOWPACK (Lehning et al., 1999). These models predict snow evo-
 627 lution from timeseries of meteorological conditions. As their outputs include all the vari-
 628 ables required by our chain (density, SSA and traditional grain shape), it becomes ob-
 629 solete to rely on empirical coefficients (Brucker, Royer, et al., 2011). This achievement
 630 should increase the interest in microwave satellite observations to assess or constrain the
 631 snowpack models in the future. Our findings open new perspectives in large scale sim-
 632 ulations of microwave signatures (Pulliainen et al., 2020) and in data assimilation of mi-
 633 crowave observations in snow hydrological models (Durand & Margulis, 2006).

634 In the future, instead of relying on the traditional grain shape to infer the poly-
 635 dispersity value, direct and more precise values could be obtained. For in-situ surveys,
 636 polydispersity can be obtained from snow samples imaged by micro-CT, although this
 637 involves significant work. From a modeling point of view, CROCUS and SNOWPACK
 638 already have a "sphericity" prognostic variable to represent grain shape. Unfortunately
 639 the "sphericity" definition established three decades ago (Brun et al., 1989) is not com-
 640 patible with the isoperimetric shape factor which we demonstrated to be equivalent to
 641 the microwave polydispersity in sparse media. More work is needed to relate these quan-
 642 tities. An even more advanced and promising avenue is the future snow evolution mod-
 643 els that are expected to describe metamorphism laws more closely to the microstructure
 644 (Leinss et al., 2020). A model able to predict the evolution of the auto-covariance func-
 645 tion or of the CLD from fundamental thermodynamic principles would indeed enable seam-
 646 less predictions of the microwave polydispersity.

647 However, there are still some important unsolved issues. From a theoretical point
 648 of view, a better understanding of the peculiar geometrical features of the microstruc-
 649 ture controlling the CLD is needed. Although our results for isolated convex grains are
 650 simple and intuitive (the polydispersity K is a measure of grain sphericity), the situa-
 651 tion for dense media seems more complex. The equation (Eq. 8) established to estimate
 652 the polydispersity as a function of the chord length moments gives a practical way to
 653 compute K from micro-CT, but does not reveal exactly which geometrical features of
 654 dense media control the polydispersity. The long-range order in the medium, character-
 655 izing how grains are arranged relatively to each other, is known to influence K (Chen
 656 et al., 1990; Ruland, 2010) but investigations on the order in snow microstructure is lack-
 657 ing. It will also be important to determine whether the polydispersity can be assumed
 658 to be constant for depth hoar crystals grown in different conditions. The range of depth
 659 hoar polydispersity estimated in the present study is about 1.2–1.9 (50% variation) from
 660 microwave and even larger when including the calculation from micro-CT. It is certainly
 661 the largest source of l_{MW} uncertainties considering that l_p can be derived from mea-
 662 surements of SSA, with 15% uncertainties, and density with 10% uncertainties. This large
 663 range is not a surprise according to our field experience. Depth hoar is certainly the snow
 664 type with the largest visual variations in crystal size, shape and order across the world.
 665 Columnar depth hoar features the largest and most organized crystals (Fig. 9) and is
 666 expected to yield high polydispersity, while some depth hoar found at the bottom of the
 667 alpine snowpack is often small and random. Further investigation on depth hoar with
 668 micro-CT is required. This study also assumes an isotropic medium from the very be-
 669 ginning although snow geometrical properties are known to be different in the vertical
 670 and horizontal directions (Krol & Löwe, 2016). A possible approach may be to consider
 671 a different microwave grain size for each Cartesian direction (Leinss et al., 2020).

672 Introducing the microwave grain size l_{MW} and the polydispersity K provides a way
 673 to relate the different microstructure models but it does not solve the problem of choos-
 674 ing the most adequate microstructure representations for snow. All the representations
 675 reach approximately the same RMSE after optimization of the polydispersity, except SHS
 676 in the case of depth hoar. In light of the results, the scaled exponential may seem at-
 677 tractive because of its simplicity and efficiency, and the scaling factor introduced empir-
 678 ically in the past by fitting exponential curves to measured auto-covariance functions (Mätzler,
 679 2002; Krol & Löwe, 2016) or by microwave optimization (Royer et al., 2017). It appears
 680 to correspond to the polydispersity K , but it is not strictly similar as fitting an expo-
 681 nential curve to the auto-covariance function may differ from integrating this function
 682 with Eq. 6. The scaling factor of 0.75 (on average) established by (Mätzler, 2002) is close
 683 to our estimates of K for alpine snows, and the increasing trend from fresh snow and de-
 684 composing particles to faceted grains and to depth hoar (Supporting Table S5) has sim-
 685 ilarities with our findings (Fig 5). Despite these great advantages, the sEXP does not
 686 respect the required mathematical properties of an auto-covariance function at the origin (Torquato
 687 & Haslach, 2002). The impact of this inconsistency on microwave scattering is negligi-
 688 ble because the long-range behavior of the auto-covariance function matters most (Eq.
 689 6). However, it is important in the optical range (Krol & Löwe, 2016) making sEXP un-
 690 suitable for a unified treatment of snow in the optical and microwave ranges. The SHS
 691 representation has a fully valid auto-covariance function and yields the best performance
 692 in Antarctica, but clearly fails to represent depth hoar. Teubner–Strey seems adequate
 693 for any type of grains, even though performance in Antarctica is slightly reduced com-
 694 pared to SHS. To discriminate the representation performances, future work should in-
 695 vestigate the snow microwave response at higher frequencies (e.g. 150 GHz available on
 696 the Microwave Humidity Sounder) where the microstructure details play a more promi-
 697 nent role. This requires higher resolution measurements of snow properties than what
 698 collected so far.

699 The theory developed in this study is of interest beyond snow and microwaves. It
 700 clarifies how the microstructure of a porous medium controls wave scattering, when the
 701 wavelength is larger than the grain size, independently of the constituent materials and
 702 of the wave nature. The microwave grain size and the polydispersity K as defined here
 703 are new and general metrics useful to investigate a variety of media and waves. Conversely,
 704 we expect that the tied theoretical links will help to transfer new knowledge from the
 705 materials science to snow scattering and ultimately contribute to more efficient remote
 706 sensing applications.

707 Open Research

708 All microwave and snow data needed to evaluate the conclusions are available in
 709 the repository <https://doi.org/10.18709/perscido.2022.05.ds367> (Picard, Löwe, et al., 2022).
 710 The version of the SMRT model code with unified microstructure is available from
 711 <https://doi.org/10.5281/zenodo.6518996> (Picard, Sandells, & Löwe, 2022). The code to
 712 produce the main result figures (Fig 5–8) is made available from
 713 <https://doi.org/10.5281/zenodo.6519037> (Picard, Sandells, & Löwe, 2022).

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726 Conflict of Interest

727 The authors declare no conflicts of interest relevant to this study.

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