Title : Geochemical implication of Eu isotope ratio in the anorthosite: a new evidence of Eu isotope fractionation by feldspar crystallization ".

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1	Geochemical implication of Eu isotope ratio in the anorthosite: a new
2	evidence of Eu isotope fractionation by feldspar crystallization
3	
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13	Rare earth element geochemistry is important for understanding the evolution of the crust-
14	mantle system. Europium (Eu) exists in divalent and trivalent states, and Eu <sup>2+</sup> can be substituted
15	for Ca <sup>2+</sup> during plagioclase feldspar crystallization in reducing magmas. This leads to positive
16	Eu anomaly in Ca-plagioclase-rich anorthosite derived from the mantle and negative Eu
17	anomalies in fractionated silica-rich crustal rocks. But while Eu anomalies are well known, Eu
18	has only two stable isotopes ( <sup>151</sup> Eu and <sup>153</sup> Eu), Eu isotope ratios have not been compared with
19	Eu anomalies in igneous rocks. Here we report a systematic variation of the Eu isotope ratio
20	$(\delta^{153/151}Eu)$ from igneous rocks including anorthosite. This study finds a linear relationship

22 fractionated granites having large negative Eu anomalies and negative  $\delta^{153/151}$ Eu values but

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between Eu anomalies and Eu isotope ratios in igneous rocks, with rhyolites and highly

anorthosites having large positive Eu anomalies and positive  $\delta^{153/151}$ Eu values. Particularly, in the area of the highly fractionated igneous rocks with negative Eu anomaly, the Eu isotope fractionation proceeds with different slope according to the degree of magmatic differentiation in extrusive (volcanic) and intrusive (plutonic) rocks. Our finding reveals that Eu isotope fractionation in igneous rocks will provide new information related to magmatic differentiation and plagioclase feldspar fractional crystallization including anorthosite formation in the crustmantle.

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Key words: Eu isotope fractionation, Eu anomaly, magmatic differentiation, feldspar
crystallization, anorthosite

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### **INTRODUCTION**

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38 Rare earth elements (REEs) and their radiogenic isotope geochemistry (especially the <sup>147</sup>Sm-<sup>143</sup>Nd and <sup>138</sup>La-<sup>138</sup>Ce system) have provided abundant information for interpreting the 39 geochemical evolution of Earth and extra-terrestrial materials as a result of their similar 40 chemical behavior and continuously varying atomic masses of REEs. In particular, the 41 42 geochemistry of chondrite-normalized REEs provides valuable petrogenetic information 43 during the magma evolution processes such as partial melting from the mantle-derived rocks 44 or crystallization from the magma (Coryell et al., 1963; Fowler and Doig, 1983; Masuda, 1962; Shearer and Papike, 1989; Weill and Drake, 1973). Most REEs have a stable (+3) state in 45 natural systems; however, Eu can exist in both divalent and trivalent state under magmatic 46 redox conditions, which indicates that the behavior of Eu during magmatic differentiation 47 depends on the oxygen fugacity and crystallization of minerals (Burnham et al., 2015). 48 49 Fractional crystallization is considered the dominant mechanism of magmatic differentiation and isolates crystallized-minerals from magma step by step (Bowen, 1928). Positive or 50 negative Eu anomalies from the igneous rocks are produced by feldspar (particularly 51 plagioclase) fractional crystallization with removal or accumulation, respectively, of 52 plagioclase during magma evolution and have been interpreted as indicating the degree of 53 differentiation of the source magma (Fowler and Doig, 1983; Shearer and Papike, 1989; Weill 54 55 and Drake, 1973). For example, extremely large positive Eu anomaly in the anorthosite is due to concentration of Eu due to be substituted into the Ca site in plagioclase feldspar because the 56 Ca<sup>2+</sup> site in feldspar readily accepts Eu<sup>2+</sup>. However, highly fractionated granite and high-silica 57 rhyolite shows extremely large negative Eu anomalies. 58

Eu has only two isotopes, <sup>151</sup>Eu (47.81%) and <sup>153</sup>Eu (52.19%) (Rossman and Taylor, 1998). Though <sup>151</sup>Eu decayed to <sup>147</sup>Pm by  $\alpha$  decay with the half-life T<sub>1/2</sub>=5×10<sup>18</sup>yr (Belli et al., 2007), <sup>151</sup>Eu can be considered as a stable isotope in earth and solar system. In addition, recently, Lee and Tanaka (2021a) reported Eu isotope fractionation due to light Eu isotope enrichment (<sup>151</sup>Eu) in highly fractionated granite and high-silica rhyolite with large negative Eu anomalies. The authors proposed that the heavier Eu isotope (<sup>153</sup>Eu) might be enriched in anorthosite with large Eu positive anomaly due to Ca-feldspar crystallization.

At present, there have been no report on Eu anomaly and Eu isotope ratio in anorthosite including gabbro as well as the volcanic rocks such as andesite and trachyte. Therefore, here, we report the Eu isotope ratio and Eu anomaly among plutonic (intrusive) rocks such as anorthosites, gabbro and volcanic (extrusive) rocks such as andesite and trachyte, and compare the data of Eu isotope ratio from the igneous rocks such as basalt, rhyolite and granitoids.

The objective in this article is to find a possibility of a new tracer for studying the relationship between Eu anomaly in the chondrite-normalized REE pattern and the Eu isotope fractionation by comparing the magnitude of Eu anomaly and the degree of Eu isotope fractionation in various kinds of igneous rocks such as the extrusive rocks and intrusive rocks including anorthosite.

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## SAMPLES AND EXPERIMENTAL METHODS

- 78 Samples
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80 In order to measure Eu isotope ratio of various kinds of igneous rocks, 49 igneous rock samples were used for Eu isotope ratio and REE abundance determination, of which 25 samples 81 were geochemical reference materials purchased from the United States Geological Survey 82 (USGS) and the Geological Survey of Japan (GSJ), while the others were anorthosites, 83 84 granitoids and trachytes from Korea and Antarctica. The 25 geochemical reference materials in this study were as follows; Seven basalts (BCR2, BHVO2 and BIR1a purchased from the 85 USGS; JB1a, JB1b, JB2 and JB3 from the GSJ), four andesites (AGV2 from USGS; JA1, JA2 86 87 and JA3 from GSJ), four rhyolites (RGM2 from USGS; JR1, JR2 and JR3 from GSJ), one diabase (W-2a from USGS), one dolerite (DNC1a from USGS), two gabbros (JGb1, JGb2 from 88 GSJ), one syenite (STM2 from USGS), and five granites (G2 and GSP2 from USGS; JG1a, 89 90 JG2 and JG3 from GSJ). The 24 rock samples from Korea and Antarctica are as follows; fourteen granites and five anorthosites from Korea, and five trachytes from Antarctica. Because 91 there was no SRM trachyte, we collected five trachytes from Antarctica. 92

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# 94 Sample digestion and determination of REE concentrations

The sample digestion procedure was based on the approach of Lee et al (2016). 95 Approximately 100~200 mg of each sample powder was decomposed by a 2:1 mixture of 2~4 96 97 mL of concentrated HF (29M) and 1~2 mL of concentrated HNO<sub>3</sub> (16M) at ca. 160 °C for more than 72 hours in 15 mL Savillex vial. After the addition of 0.1~0.2 mL of concentrated HClO<sub>4</sub>, 98 the decomposed sample solution was heated to dryness at ca. 180 °C for more than 1 day. The 99 100 cakes were re-dissolved by a mixture of 1 mL of concentrated HCl and 0.5 mL of concentrated HNO3 at ca. 160 °C for 1 day. Each sample solution was dried again, and diluted in 10 ml of 6 101 M HCl as a stock solution. Of this 10 ml stock solution, 0.5~1 ml stock solution was used to 102

determine the rare earth element (REE) concentrations, and the remainder was used fordetermination of Eu isotope ratio.

Before Eu purification, we analyzed REE concentration of the sample using inductively coupled plasma mass spectrometry (ICP-MS, NexION350, Perkin Elmer) at KIGAM. Although Eu anomalies of the 25 geochemical reference rocks (USGS, GSJ) have previously been characterized, we also reanalyzed their REE abundances for comparison. The analyzed REE data from the geochemical reference sample powders (SRM) agreed with the recommended values within 5~10%.

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## 112 Experimental procedures for determination of Eu isotope ratio

Recently, Lee and Tanaka (2019, 2021b) developed a method for determining highly precise and accurate Eu isotope ratio using Sm spike as an internal standard in combination with standard-sample-standard bracketing mass bias correction (C-SSBIN). In addition, Lee and Tanaka (2021a, 2021b) also showed that incomplete Eu purification from the geological material lead to change in the Eu isotope ratio, that is, pseudo-fractionation of Eu isotope ratio.

In this study, Eu was separated from the obtained REE fraction using 0.12 M 2hydroxyisobutyric acid (HIBA) with the pH adjusted to ~4.60 (Lee and Tanaka, 2019, 2021b). To minimize isobaric interference because we used Sm which was prepared from ultrapure Sm<sub>2</sub>O<sub>3</sub> produced by Alfa Aesar as a spike for normalization of Eu isotopes, we always checked for tailing of both Gd and Sm.

Eu isotope ratios were measured using multicollector inductively coupled plasma mass spectrometry (MC-ICP-MS; Neptune Plus, Thermo Fisher Scientific Ltd.) in static mode with 125 nine Faraday cups at KIGAM. The instrument was tuned to achieve high sensitivity while maintaining flattened square peaks and stable signals enough to ensure accurate measurements. 126 The gain on each Faraday cup was monitored daily to ensure normalization of its efficiency. 127 Sample dilution for Eu isotope measurement by MC-ICP-MS was performed with 2% HNO<sub>3</sub> 128 which was prepared from 60% ultrapure HNO<sub>3</sub> (Merck, Darmstadt, Germany) and DIW (Milli-129 Q system, Millipore, Milford, USA). We used a diluted solution of NIST 3117a (10,000 µg/mL, 130 Lot No. 120705) as an in-house standard solution for comparison of the Eu isotope ratios. 131 The isotopes  ${}^{147}Sm(L4)$ ,  ${}^{149}Sm(L3)$ ,  ${}^{150}Sm(L2)$ ,  ${}^{151}Eu(L1)$ ,  ${}^{152}Sm(C)$ ,  ${}^{153}Eu(H1)$ , 132 <sup>154</sup>Sm(H2), <sup>155</sup>Gd(H3), and <sup>157</sup>Gd(H4) were monitored simultaneously using nine Faraday cups 133 for Sm normalization and Gd interference correction by the Gd matrix (Lee and Tanaka, 2021b). 134 135 Data acquisition consisted of 1 block of 50 cycles with an integration time of 4.194 seconds and a sample aspiration rate of 80-100  $\mu$ L/min. Peak centering was performed at the beginning 136 of each analysis and 250 s of washout time was used between sample measurements. Blanks 137 were checked during, before, and after each sample measurement. Operating conditions and 138 data acquisition parameters, including cup configuration, are the same as Lee and Tanaka 139 (2021a, 2021b). In determination of Eu isotope ratio, we used <sup>147</sup>Sm/<sup>149</sup>Sm (1.0868, Dubois et 140 al., 1992) for normalization to obtain an optimum value of Eu isotope fractionation from the 141 natural materials because of isobar matrix problem by <sup>154</sup>Gd due to incomplete separation from 142 the geological rock during <sup>150</sup>Sm/<sup>154</sup>Sm normalization (Lee and Tanaka, 2021a, 2021b). Eu 143 isotope fractionation is represented in standard  $\delta$ -notation in per mil relative to the NIST3117a 144 Eu standard solution as follows:  $\delta^{153/151}Eu = 1,000 \times [(^{153}Eu/^{151}Eu_{sample})/(^{153}Eu/^{151}Eu_{NIST3117a})$ 145 146 -1].

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#### RESULTS

149 Rare Earth Element concentration in the extrusive rocks and intrusive rocks

In order to validate the accuracy of the REE data in this study, we determined REE concentrations of 25 geochemical standard reference materials (SRM) produced by USGS and GSJ. REE abundances of the geochemical reference materials from the United States of Geological Survey (USGS) and Geological Survey of Japan (GSJ) measured in this study are reported in Supplementary Table S1. In addition, REE concentrations of the local igneous rocks in Korea and Antarctica including anorthosite are presented in Table 1.



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Fig. 1. Chondrite (McDonough and Sun, 1995)-normalized REE pattern of standard reference materials (SRMs)
of USGS and GSJ. (a), (b) and (c) are SRMs for volcanic (extrusive) rocks whereas (d), (e) and (f) are SRMs for
intrusive rocks such as gabbro, diabase and granitoids. In this paper, we classified dolerite and diabase as plutonic
rocks rather than volcanic rocks. The REE abundances of all SRMs were re-measured from this study (see Table
S1). REE patterns by solid black dots were drawn by recommended values for each SRM from USGS and GSJ.

163 Figures 1 and 2 are chondrite-normalized REE patterns for SRMs and Korea and Antarctic igneous rocks, respectively. In Fig. 1, REE patterns from the volcanic rocks such as basalt, 164 andesite and rhyolite were drawn in Fig. 1a~1c whereas those from the plutonic rocks such as 165 gabbro and granitoids were drawn in Fig. 1d~1f. The chondrite-normalized REE patterns from 166 various kinds of igneous rocks in Figs. 1 and 2 clearly show variation of the magnitude of Eu 167 anomaly due to feldspar crystallization during magmatic differentiation even though they are 168 not cogenetic igneous rocks. Particularly, the anorthosites in Fig. 2a has strikingly large Eu 169 positive anomaly. 170



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Fig. 2. Chondrite (McDonough and Sun, 1995)-normalized REE pattern of Korean and Antarctic igneous rocks.
Except (b) trachyte, the others are all collected from Korea. Trachyte was collected from Antarctica. The REE abundances of Korean anorthosites and Anrarctic trachytes were re-measured from this study. REE patterns of Koean granites are from Lee et al. (2004, 2006, 2008, 2013).

Eu isotope ratio and magnitude of Eu anomaly from various kinds of the igneous rocks in this study are presented in Table 2. For comparison, in Table 2, we divided the samples into SRM and local igneous rocks. The <sup>153/151</sup>Eu value of anorthosite all are positive, indicating that heavier Eu isotope (<sup>153</sup>Eu) was enriched compared to lighter Eu isotope (<sup>151</sup>Eu).

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## DISCUSSION

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A major objective of this study is to examine the relationship between the magnitude of 185 Eu anomaly due to feldspar crystallization during magmatic differentiation and Eu isotope 186 differentiation. Therefore, we first compared the magnitudes of Eu anomalies with the 187  $\delta^{153/151}$ Eu values of the 49 samples to investigate the possibility that Eu isotope fractionation 188 had occurred during magma differentiation (Fig. 3). Figure 3 illustrates three geochemical 189 characteristics of Eu isotope fractionation as follows; 1) The  $\delta^{153/151}$ Eu values of the highly 190 fractionated granites and rhyolites with extremely large negative Eu anomaly are negative, 191 whereas the anorthosites with extremely large positive Eu anomaly show relatively large 192 positive  $\delta^{153/151}$ Eu values. This contrast indicates that the highly fractionated igneous rocks, 193 which are emplaced from felsic magma in an upper crustal environment, were enriched in the 194 lighter Eu isotope (<sup>151</sup>Eu), whereas Ca-plagioclase-rich anorthosite, which are derived from the 195 mafic magma in a lower crustal environment, were enriched in the heavier Eu isotope (<sup>153</sup>Eu). 196 2) The  $\delta^{153/151}$ Eu value in the igneous rock varies systematically with magnitude of the Eu 197 anomaly. 3) Another interesting feature of the Eu fractionation trends in the volcanic and 198

plutonic rocks is that intrusive rocks (red symbols in Fig. 3) and extrusive rocks (green andblue symbols in Fig. 3) are distributed with different slopes.



Fig. 3. Variation of Eu isotope ratio according to magnitudes of Eu anomalies from igneous rocks. The error bars represent uncertainties (2SD) of the average  $\delta^{153/151}$ Eu values from some of igneous rock samples.

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Ismail et al. (1998) employed cation exchange chromatography and concluded that an 206 isotope effect in the  $Eu^{2+}/Eu^{3+}$  exchange reaction may occur in aqueous solutions; specifically, 207 they found isotope effects in which the heavier isotope  $^{153}$ Eu is enriched in Eu<sup>2+</sup> in the 208 Eu<sup>2+</sup>/Eu<sup>3+</sup> electron exchange system. The positive Eu anomaly in anorthosite can easily be 209 explained by the substitution of  $Eu^{2+}$  for  $Ca^{2+}$  in plagioclase during differentiation of the 210 211 anorthositic (primary) magma either in the upper mantle or at the lower crust. Therefore, we can suggest that the enrichment of heavier isotope <sup>153</sup>Eu in the anorthosites with large positive 212 Eu anomaly should be explained due to isotope effects in  $Eu^{2+}/Eu^{3+}$  electron exchange system 213 214 during Ca-plagioclase accumulation in the anorthositic magma. Moreover, the systematic correlation between Eu isotope fractionation and the magnitudes of Eu anomaly from the fractionated igneous rocks and anorthosites indicates that Eu isotope fractionation was closely related to magmatic differentiation processes such as feldspar fractional crystallization.

Another interesting finding in Figure 3 is that the trend of Eu isotope fractionation 218 between the extrusive volcanic rocks and intrusive plutonic rocks seems to be different. 219 220 Although more in-depth studies are needed in the future, the existence of such different slope 221 can be interpreted as indicating the possibility that the Eu isotope fractionation in intrusive and 222 extrusive rocks occurred by different mechanism or geochemical environment. The behavior 223 of Eu is known to be determined by temperature and oxygen fugacity (Weill and Drake, 1973), and Philpotts and Schnetler (1968) and Philpotts (1970) suggested that Eu anomalies are 224 controlled by crystal chemistry and magmatic oxidation potential. Dauphas et al. (2014) 225 showed that equilibrium iron isotope fractionation is controlled mainly by the redox and 226 structural conditions in magma and suggested that magmatic differentiation is the main driver 227 228 of Fe isotope fractionation in felsic magmas. In addition, Dauphas et al. (2014) proposed that stable isotopes from heterovalent elements, including Eu, may show isotopic variations in bulk 229 rocks controlled by the redox and structural conditions in the magma. Therefore, we also may 230 be able to consider a possibility that a slight difference of the  $\delta^{153/151}$ Eu values in the plutonic 231 rocks and volcanic rocks may be due to the oxidation potential in the magma. Further study is 232 needed to clarify the relationship between Eu isotope fractionation and oxidation potential in 233 an intrusive magmatic system. 234

Besides Eu isotope fractionation, recently, several research groups reported isotope fractionation study of REEs like Ce, Nd, Sm, Dy, Er and Yb (Moynier et al., 2006; Nakada et al., 2013; Shollenberger and Brebbecka, 2020; Hu et al., 2021). This means that combined information using chondrite-normalized REE pattern, radiogenic isotope geochemistry and
δREE will provide new constraints for understanding more clearly the processes in our Earth
and Planetary system such as redox conditions in the mantle and/or crust, early earth formation,
and crust/mantle differentiation.

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# CONCLUSION

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245 We compared the magnitude of Eu anomaly in the chondrite-normalized REE pattern from various kinds of igneous rocks such as the extrusive rocks and intrusive rocks including 246 anorthosite.and their Eu isotope ratio from the fractionation. The anorthosites having large 247 positive Eu anomalies show a geochemical characteristic of a heavier Eu isotope (<sup>153</sup>Eu) 248 enrichment (i.e., positive  $\delta^{153/151}$ Eu value) whereas the rhyolites and highly fractionated 249 250 granites having large negative Eu anomalies show a geochemical characteristic of a lighter Eu isotope (<sup>151</sup>Eu) enrichment (i.e., negative  $\delta^{153/151}$ Eu value). Particularly, our results clearly 251 showed that variation of the magnitude of Eu anomaly and Eu isotope fractionation in igneous 252 253 rocks has systematic correlation, suggesting that Eu isotope fractionation in igneous rocks should be produced by feldspar crystallization during magma evolution. In addition, the Eu 254 isotope fractionation in the highly fractionated volcanic and plutonic rocks was proceeded with 255 different trend, implying that the Eu isotope fractionation from the intrusive and extrusive 256 magma in the crustal environment may occur under different mechanism or geochemical 257 258 environment.

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**Figure Cations** 

360 361

Fig. 1. Chondrite (McDonough and Sun, 1995)-normalized REE pattern of standard reference materials (SRMs) of USGS and GSJ. (a), (b) and (c) are SRMs for volcanic (extrusive) rocks whereas (d), (e) and (f) are SRMs for intrusive rocks such as gabbro, diabase and granitoids. In this paper, we classified dolerite and diabase as plutonic rocks rather than volcanic rocks. The REE abundances of all SRMs were re-measured from this study (see Table S1). REE patterns by solid black dots were drawn by recommended values for each SRM from USGS and GSJ.

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Fig. 2. Chondrite (McDonough and Sun, 1995)-normalized REE pattern of Korean and
Antarctic igneous rocks. Except (b) trachyte, the others are all collected from Korea. Trachyte
was collected from Antarctica. The REE abundances of Korean anorthosites and Antarctic
trachytes were re-measured from this study. REE patterns of Korean granites are from Lee et
al. (2004, 2006, 2008, 2013).

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Fig. 3. Variation of Eu isotope ratio according to magnitudes of Eu anomalies from igneous rocks. The error bars represent uncertainties (2SD) of the average  $\delta^{153/151}$ Eu values from some of igneous rock samples.

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Rock Type	Area	Sample Name	La (ppm)	Ce (ppm)	Pr (ppm)	Nd (ppm)	Sm (ppm)	Eu (ppm)	Gd (ppm)	Tb (ppm)	Dy (ppm)
		M17102801-3	78.42	190.5	18.12	68.07	13.17	3.26	11.29	1.80	10.81
	NG	J13010105	103.3	198.7	21.92	80.38	15.09	3.46	13.09	2.04	11.92
Trachyte	Mt. Melbourne	K16012306	87.78	164.8	18.74	70.02	13.35	3.76	12.51	1.68	9.89
	(Antarctica)	M171110-03	86.93	172.4	18.10	70.93	12.98	3.49	12.50	1.77	10.11
		J13010104	78.42	190.5	18.12	68.07	13.17	3.26	11.29	1.80	10.81
		SA20170309 1-1	1.64	3.25	0.36	1.41	0.26	0.87	0.23	0.03	0.17
		SA20170309 -2	1.73	3.18	0.36	1.43	0.27	0.99	0.28	0.04	0.24
Anorthosite	Sancheong (Korea)	SA20170309-3	1.65	3.12	0.35	1.47	0.31	0.66	0.33	0.05	0.27
		SA20170309-6-2	1.42	2.27	0.22	0.77	0.11	1.24	0.08	0.01	0.05
		SA20170310 3-3	1.41	2.45	0.28	1.17	0.22	1.04	0.26	0.04	0.23
	Muamsa- Weolaksan (Korea)	MA10	12.42	28.12	4.23	16.35	6.12	0.01	7.61	1.62	12.39
		WAWR12	20.97	54.60	5.67	21.40	6.56	0.13	8.13	1.59	11.35
		WAWR8	49.60	96.30	10.80	40.20	8.07	0.24	7.15	1.07	6.56
		MA2018	46.39	60.86	11.64	42.91	10.59	0.31	11.08	1.93	13.23
		WA2018	27.30	62.63	7.88	30.31	9.17	0.20	10.63	2.17	15.58
	Seokmodo	SM4	48.70	81.50	8.11	27.70	4.00	0.87	4.00	0.36	1.35
Granitoid	(Korea)	SM26	72.80	120.80	13.50	48.50	6.66	1.78	4.55	0.47	1.85
	Icheon	ICH3	41.20	83.00	10.32	40.28	7.29	1.54	5.42	0.64	3.24
	(Korea)	ICH11	54.50	98.70	11.10	42.10	7.34	1.12	5.75	0.71	2.92
	Pohang	C2320	16.15	28.50	3.03	11.35	2.20	0.59	2.18	0.35	2.30
	(Korea)	C2980	11.95	20.93	2.24	8.49	1.64	0.72	1.65	0.27	1.58
	Taedo	TD13	79.97	141.0	19.52	65.33	11.11	0.90	7.83	1.33	9.16
	(Korea)	TD2B	138.3	273.9	29.76	101.8	18.36	1.21	15.36	2.50	15.84

Ho (ppm)	Er (ppm)	Tm (ppm)	Yb (ppm)	Lu (ppm)	Reference
2.10	5.79	0.83	5.14	0.66	
2.26	6.25	0.88	5.57	0.83	
1.88	5.19	0.71	4.50	0.65	
1.99	5.58	0.73	4.88	0.72	
2.10	5.79	0.83	5.14	0.66	
0.03	0.08	0.01	0.06	0.01	
0.05	0.13	0.02	0.12	0.02	
0.05	0.15	0.02	0.12	0.02	This study
0.01	0.02	0.00	0.02	0.002	
0.05	0.14	0.02	0.13	0.02	
2.92	10.46	1.86	14.55	2.34	
2.41	7.57	1.21	8.54	1.25	
1.40	3.88	0.61	3.97	0.56	
2.66	8.95	1.42	9.67	1.43	
3.31	11.85	2.14	16.01	2.51	
0.20	0.44	0.06	0.37	0.05	Lee et al.
0.30	0.64	0.09	0.52	0.07	(2006)
0.57	1.37	0.21	1.33	0.18	Lee et al.
0.51	1.29	0.16	0.93	0.13	(2004)
0.50	1.42	0.28	1.46	0.25	Lee et al.
0.36	0.97	0.17	1.03	0.17	(2008)
2.05	6.54	1.06	6.96	1.07	This
3.23	9.31	1.39	8.73	1.27	study

## Table 2. Eu isotope ratio from SRM and local igneous rocks in this study

Rock type			В	Basalt (SR	.M)				Andesit	e (SRM)			Trach	yte (Anta	rctica)			Rhyolit	e (SRM)		SI	RM	Gabbro	(SRM)
Sample name	BCR2 <sup>1)</sup>	JB2	JB1a	JB1b	JB3	BHVO2	BIR1a	JA1	JA2	JA3	AGV2	M1710 2801-3	J1301 0105	K1601 2306	M1711 10-03	J1301 0104	JR2	JR1	RGM2	JR3	W2a	DNC1a	JGb2	JGb1
Eu/Eu*2)	0.94	0.96	0.99	0.94	0.97	1.07	1.13	0.93	0.95	0.80	0.99	0.24	0.16	0.94	1.01	0.25	0.07	0.15	0.56	0.06	0.98	1.08	3.38	1.27
$\delta^{153/151}Eu^{3)}$	-0.08	-0.02	-0.05	-0.06	-0.05	-0.02	0.09	-0.13	-0.09	-0.08	-0.03	-0.27	-0.41	-0.07	-0.05	-0.54	-0.37	-0.28	-0.10	-0.50	-0.03	0.04	0.08	-0.07
2SD	0.04 (n=12)	0.14 (n=10)	0.14 (n=8)	0.11 (n=2)	0.10 (n=14)	0.17 (n=4)	0.05 (n=7)	0.20 (n=7)	0.09 (n=7)	0.23 (n=7)	0.12 (n=3)	-	-	-	-	-	-	0.08 (n=2)	0.02 (n=3)	0.27 (n=4)	0.04 (n=4)	0.14 (n=6)	0.23 (n=11)	0.19 (n=7)
											Intrusiv	e (plutoni	c) rocks											
Rock type			Anorthosi	ite								Granitoids						Granitoids (SRM)						
Sample name	SA1703 09-1-1	SA1703 09-2	SA1703 09-3	SA1703 09-6-2	SA1703 10-3-3	MA10	MA <sup>4)</sup> 2018	WAWR 2018	WAWR 12	WAWR 8	SM4	SM26	ICH11	ІСН3	C2320	C2980	TD13	TD2B	STM2	JG1a	G2	JG3	GSP2	JG2
Eu/Eu*	37.71	12.28	6.29	12.98	9.77	0.01	0.09	0.06	0.06	0.09	0.66	0.93	0.51	0.72	0.81	1.32	0.30	0.21	1.01	0.48	0.79	0.89	0.35	0.03
$\delta^{153/151}Eu$	0.118	0.125	0.172	0.184	0.244	-0.28	-0.11	-0.07	-0.17	-0.16	0.05	0.00	-0.03	-0.07	-0.02	-0.02	-0.07	-0.18	0.00	-0.03	-0.10	-0.01	-0.07	-0.31

# Extrusive (volcanic) rocks

0.30 (n=6)

Rock type			Anorthosi	te			Granitoids												Granitoids (SRM)						
Sample name	SA1703 09-1-1	SA1703 09-2	SA1703 09-3	SA1703 09-6-2	SA1703 10-3-3	MA10	MA <sup>4)</sup> 2018	WAWR 2018	WAWR 12	WAWR 8	SM4	SM26	ICH11	ІСН3	C2320	C2980	TD13	TD2B	STM2	JG1a	G2	JG3	GSP2		
Eu/Eu*	37.71	12.28	6.29	12.98	9.77	0.01	0.09	0.06	0.06	0.09	0.66	0.93	0.51	0.72	0.81	1.32	0.30	0.21	1.01	0.48	0.79	0.89	0.35		
$\delta^{153/151}Eu$	0.118	0.125	0.172	0.184	0.244	-0.28	-0.11	-0.07	-0.17	-0.16	0.05	0.00	-0.03	-0.07	-0.02	-0.02	-0.07	-0.18	0.00	-0.03	-0.10	-0.01	-0.07		
2SD	-	0.01 (n=2)	0.05 (n=3)	0.03 (n=3)	-	0.40 (n=2)	-	-	-	0.31 (n=4)	-	-	-	-	-	-	-	-	0.06 (n=5)	0.06 (n=13)	0.19 (n=4)	0.04 (n=10)	0.05 (n=17)		

1) The samples of Bold character are SRMs of USGS and GSJ.

2) The magnitude of Eu anomaly is defined as s the ratio  $Eu_N/Eu^*$  where  $Eu^*$  is SQRT(Sm<sub>N</sub> x Gd<sub>N</sub>).

3) Eu isotope ratio normalized by <sup>147</sup>Sm-<sup>149</sup>Sm isotope pair (Lee and Tanaka, 2021a, 2021b).

4) Eu isotope data (Bold Italic numbers) are from Lee and Tanaka (2021a)







Table S1. Concentrations of Rare earth element of standard reference materials (SRMs) measured in this study

Туре	Sample	La	Ce	Pr	Nd	Sm	Eu	Gd	Tb	Dy	Но	Er	Tm	Yb	Lu	Eu/Eu*1)	Ref
	BCR2 <sup>2)</sup>	25.0	53.0	6.80	28.0	6.70	2.00	6.80	1.07	6.41	1.33	3.66	0.54	3.50	0.51	0.90	USGS
		$\begin{array}{c} 25.0\pm0.9\\ (n=\!4,1\sigma_m) \end{array}$	$53.7 \pm 1.1$ (n=4, 1 $\sigma_m$ )	$\begin{array}{c} \textbf{6.73} \pm \textbf{0.16} \\ \textbf{(n=4, 1}\sigma_{m}) \end{array}$	$\begin{array}{c} 28.5\pm0.6\\ (n{=}4,1\sigma_m) \end{array}$	$\begin{array}{c} 6.58 \pm 0.15 \\ (n{=}4,1\sigma_m) \end{array}$	$\begin{array}{c} 2.05\pm0.04\\ (n{=}4,1\sigma_m) \end{array}$	$6.76 \pm 0.17$ (n=4, 1 $\sigma_m$ )	$1.03 \pm 0.03$ (n=4, 1 $\sigma_{m}$ )	$6.48 \pm 0.13$ (n=4, 1 $\sigma_{m}$ )	$\begin{array}{c} 1.32\pm0.03\\ (n{=}4,1\sigma_m) \end{array}$	$\begin{array}{c} 3.77 \pm 0.09 \\ (n{=}4, 1\sigma_m) \end{array}$	$\begin{array}{c} 0.52\pm0.01\\ (n{=}4,1\sigma_m) \end{array}$	$3.41 \pm 0.05$ (n=4, 1 $\sigma_m$ )	$\begin{array}{l} 0.51 \pm 0.01 \\ (n{=}4,1\sigma_m) \end{array}$	0.94	this study
-	BHVO2	15.00	38.00	-	25.00	6.20	2.05	6.30	0.90	-	1.04	-	-	2.00	0.28	1.00	USGS
		$14.9 \pm 0.44$ (n=3, 1 $\sigma_{\rm m}$ )	$36.2 \pm 1.80$ (n=3, 1 $\sigma_{\rm m}$ )	$5.17 \pm 0.18$ (n=3, 1 $\sigma_{\rm m}$ )	$23.2 \pm 1.22$ (n=3, 1 $\sigma_{\rm m}$ )	$5.76 \pm 0.16$ (n=3, 1 $\sigma_{\rm m}$ )	$2.05 \pm 0.10$ (n=3, 1 $\sigma_{\rm m}$ )	$5.90 \pm 0.34$ (n=3, 1 $\sigma_{\rm m}$ )	$0.97 \pm 0.07$ (n=3, 1 $\sigma_{\rm m}$ )	$4.89 \pm 0.36$ (n=3, 1 $\sigma_{\rm m}$ )	$0.97 \pm 0.03$ (n=3, 1 $\sigma_{\rm m}$ )	$2.39 \pm 0.17$ (n=3, 1 $\sigma_{\rm m}$ )	$0.26 \pm 0.02$ (n=3, 1 $\sigma_{\rm m}$ )	$1.66 \pm 0.12$ (n=3, 1 $\sigma_{\rm m}$ )	$0.25 \pm 0.02$ (n=3, 1 $\sigma_{\rm m}$ )	1.07	this study
-	BIR1a	0.63	1.90	-	2.50	1.10	0.55	1.80	-	4.00	-	-	-	1.70	0.26	1.13	USGS
		$0.64 \pm 0.05$ (n=3, 1 $\sigma$ )	$1.99 \pm 0.10$ (n=3.1 $\sigma$ )	$0.40 \pm 0.05$	$2.49 \pm 0.08$ (n=3.1 $\sigma$ )	$1.12 \pm 0.05$ (n=3.1 $\sigma$ )	$0.53 \pm 0.02$ (n=3.1 $\sigma$ )	$1.80 \pm 0.23$ (n=3.1 $\sigma$ )	$0.36 \pm 0.03$	$2.61 \pm 0.20$ (n=3.1 $\sigma$ )	$0.57 \pm 0.05$	$1.72 \pm 0.17$ (n=3, 1 $\sigma$ )	$0.34 \pm 0.03$	$1.94 \pm 0.05$ (n=3.1 $\sigma$ )	$0.26 \pm 0.01$ (n=3, 1 $\sigma$ )	1.13	this study
-	ID1o	(1 5, 10 <sub>m</sub> )	(I 5, 10 <sub>m</sub> )	(II 5, 10 <sub>m</sub> )	(II 5, 10 <sub>m</sub> )	(II 5, 10 <sub>m</sub> )	(II 3, 10 <sub>m</sub> )	(II 5, 10 <sub>m</sub> )	(I 5, 10 <sub>m</sub> )	(I 5, 10 <sub>m</sub> )	(I 5, 10 <sub>m</sub> )	(II 5, 10 <sub>m</sub> )	(II 5, 10 <sub>m</sub> )	$(115, 10_m)$	(II 5, 10 <sub>m</sub> )	0.01	CSI
basalt	JDIA	37.00 22.7 + 0.8	746 + 0.7	7.50 6.47 ± 0.19	20.00	5.07	1.40	4.07	0.69	<b>3.99</b>	0.70 + 0.02	2.10	0.35	<b>2.10</b>	0.35	0.91	GSJ
-		$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$24.7 \pm 0.0$ (n=6, 1 $\sigma_{\rm m}$ )	$4.72 \pm 0.11$ (n=6, 1 $\sigma_{\rm m}$ )	$1.49 \pm 0.03$ (n=6, 1 $\sigma_{\rm m}$ )	$4.43 \pm 0.10$ (n=6, 1 $\sigma_{\rm m}$ )	$(n=6, 1\sigma_m)$	$3.76 \pm 0.16$ (n=6, 1 $\sigma_{\rm m}$ )	$(n=6, 1\sigma_m)$	$2.12 \pm 0.05$ (n=6, 1 $\sigma_{\rm m}$ )	$(n=6, 1\sigma_m)$	$1.91 \pm 0.03$ (n=6, 1 $\sigma_{\rm m}$ )	$(n=6, 1\sigma_m)$	0.99	this study
	JB1b	41.20	71.80	7.73	27.10	5.17	1.59	4.38	0.69	3.73	0.67	1.97	0.31	2.10	0.31	1.02	GSJ
-		40.20	70.10	7.46	26.90	4.93	1.59	5.35	0.77	4.14	0.79	2.22	0.33	2.01	0.31	0.94	this study
	JB2	2.35	6.76	1.01	6.63	2.31	0.86	3.28	0.60	3.73	0.75	2.60	0.41	2.62	0.40	0.95	GSJ
		$2.18 \pm 0.8$ (n=8, 1 $\sigma_{\rm m}$ )	$6.39 \pm 0.14$ (n=8, 1 $\sigma_{\rm m}$ )	$1.11 \pm 0.05$ (n=8, 1 $\sigma_{m}$ )	$\begin{array}{l} 6.26 \pm 0.17 \\ (n{=}8,1\sigma_m) \end{array}$	$2.22 \pm 0.04$ (n=8, 1 $\sigma_{m}$ )	$0.83 \pm 0.03$ (n=8, 1 $\sigma_{m}$ )	$\begin{array}{l} 3.09 \pm 0.07 \\ (n{=}8,1\sigma_m) \end{array}$	$0.56 \pm 0.01$ (n=8, 1 $\sigma_{m}$ )	$3.77 \pm 0.09$ (n=8, 1 $\sigma_{m}$ )	$0.81 \pm 0.04$ (n=8, 1 $\sigma_{\rm m}$ )	$2.51 \pm 0.07$ (n=8, 1 $\sigma_m$ )	$\begin{array}{l} 0.37 \pm 0.01 \\ (n{=}8,1\sigma_m) \end{array}$	$\begin{array}{l} 2.42 \pm 0.06 \\ (n=8, 1\sigma_m) \end{array}$	$0.37 \pm 0.01$ (n=8, 1 $\sigma_{m}$ )	0.96	this study
-	JB3	8.81	21.50	3.11	15.60	4.27	1.32	4.67	0.73	4.54	0.80	2.49	0.42	2.55	0.39	0.90	GSJ
		$8.82 \pm 0.70$ (n=3, 1 $\sigma_{m}$ )	$\begin{array}{c} 22.0\pm2.5\\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 3.45 \pm 0.35 \\ (n{=}3,1\sigma_m) \end{array}$	$16.0 \pm 1.78$ (n=3, 1 $\sigma_{m}$ )	$4.31 \pm 0.35$ (n=3, 1 $\sigma_{m}$ )	$1.41 \pm 0.16$ (n=3, 1 $\sigma_{m}$ )	$\begin{array}{l} 4.55\pm0.56\\ (n{=}3,1\sigma_m) \end{array}$	$0.78 \pm 0.07$ (n=3, 1 $\sigma_{m}$ )	$4.64 \pm 0.73$ (n=3, 1 $\sigma_{m}$ )	$1.01 \pm 0.12$ (n=3, 1 $\sigma_{m}$ )	$2.71 \pm 0.44$ (n=3, 1 $\sigma_{m}$ )	$\begin{array}{c} 0.42 \pm 0.05 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 2.57 \pm 0.22 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 0.39 \pm 0.05 \\ (n{=}3,1\sigma_m) \end{array}$	0.97	this study
	JR1	19.70	47.20	5.58	23.30	6.03	0.30	5.06	1.01	5.69	1.11	3.61	0.67	4.55	0.71	0.17	GSJ
		17.35	43.83	5.58	22.52	5.37	0.27	5.48	0.94	6.09	1.25	4.25	0.70	4.94	0.76	0.15	this study
-	JR2	16.30	38.80	4.75	20.40	5.63	0.14	5.83	1.10	6.63	1.39	4.36	0.74	5.33	0.88	0.07	GSJ
		13.72	37.58	4.53	17.77	4.95	0.09	5.29	0.98	6.39	1.41	4.42	0.74	5.08	0.81	0.05	this study
	JR3	179	327	33.1	107	21.3	0.53	19.7	4.29	21.5	4.70	14.0		20.3	2.80	0.06	GSJ
Rnyonte		$167 \pm 5$ (n=4, 1 $\sigma_{\rm m}$ )	$309 \pm 9$ (n=4, 1 $\sigma_{\rm m}$ )	$30.9 \pm 1.5$ (n=4, 1 $\sigma_{\rm m}$ )	$101 \pm 3$ (n=4, 1 $\sigma_{\rm m}$ )	$20.1 \pm 0.7$ (n=4, 1 $\sigma_{\rm m}$ )	$0.41 \pm 0.02$ (n=4, 1 $\sigma_{\rm m}$ )	$20.2 \pm 0.72$ (n=4, 1 $\sigma_{\rm m}$ )	$3.99 \pm 0.25$ (n=4, 1 $\sigma_{\rm m}$ )	$27.1 \pm 1.3$ (n=4, 1 $\sigma_{\rm m}$ )	$6.00 \pm 0.33$ (n=4, 1 $\sigma_{\rm m}$ )	$18.3 \pm 0.7$ (n=4, 1 $\sigma_{\rm m}$ )	$2.96 \pm 0.18$ (n=4, 1 $\sigma_{\rm m}$ )	$19.0 \pm 0.8$ (n=4, 1 $\sigma_{\rm m}$ )	$2.71 \pm 0.07$ (n=4, 1 $\sigma_{\rm m}$ )	0.06	this study
-	RGM2	25.00	48.00	5.00	20.00	4.00	0.70	3.60	0.60	3.30	0.80	2.20			0.40	0.56	USGS
		$22.7 \pm 1.8$ (n=2 1 $\sigma$ )	$47.2 \pm 1.0$ (n=2 1 $\sigma$ )	$5.27 \pm 0.31$ (n=2 1 $\sigma$ )	$19.4 \pm 0.90$ (n=2 1 $\sigma$ )	$3.91 \pm 0.48$ (n=2 1 $\sigma$ )	$0.70 \pm 0.03$ (n=2 1 $\sigma$ )	$3.81 \pm 0.31$ (n=2 1 $\sigma$ )	$0.57 \pm 0.02$ (n=2 1 $\sigma$ )	$3.60 \pm 0.19$ (n=2 1 $\sigma$ )	$0.73 \pm 0.06$ (n=2 1 $\sigma$ )	$2.33 \pm 0.09$ (n=2 1 $\sigma$ )	$0.36 \pm 0.00$ (n=2 1 $\sigma$ )	$0.52 \pm 0.09$ (n=2 1 $\sigma$ )	$0.38 \pm 0.01$ (n=2 1 $\sigma$ )	0.56	this study
	TA 1	(II 2, 10m)	13 30	(I 2, 10m)	(ii 2, 10 <sub>m</sub> )	(ii 2, 10m)	(II 2, 10m)	(II 2, 10m)	0.75	(II 2, 10m)	0.05	(ii 2, 10m) 3.04	0.47	(ii 2, 10 <sub>m</sub> )	0.47	0.03	CSI
	JAI	<b>5.24</b> 4 69 ± 0.06	13.50 $11.8 \pm 1.52$	1.71	10.90 $10.7 \pm 0.2$	3.32	1.20 $1.16 \pm 0.02$	4.30	0.73 = 0.73	$4.55 \pm 0.12$	0.93	3.04	0.47	2.03	0.47	0.75	GOJ
-		$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	0.93	this study
	JA2	15.80	32.70	3.84	13.90	3.11	0.93	3.06	0.44	2.80	0.50	1.48	0.28	1.62	0.27	0.92	GSJ
Andesite -		$\begin{array}{c} 14.5\pm0.25\\ (n=\!6,1\sigma_m) \end{array}$	$28.8 \pm 3.9$ (n=6, 1 $\sigma_{\rm m}$ )	$3.48 \pm 0.05$ (n=6, 1 $\sigma_{m}$ )	$13.8 \pm 0.2$ (n=6, 1 $\sigma_{m}$ )	$2.93 \pm 0.04$ (n=6, 1 $\sigma_{\rm m}$ )	$0.92 \pm 0.02$ (n=6, 1 $\sigma_{\rm m}$ )	$2.99 \pm 0.03$ (n=6, 1 $\sigma_{\rm m}$ )	$0.46 \pm 0.00$ (n=6, 1 $\sigma_{\rm m}$ )	$2.78 \pm 0.04$ (n=6, 1 $\sigma_{\rm m}$ )	$0.54 \pm 0.01$ (n=6, 1 $\sigma_{m}$ )	$1.69 \pm 0.04$ (n=6, 1 $\sigma_{\rm m}$ )	$0.25 \pm 0.00$ (n=6, 1 $\sigma_{m}$ )	$1.59 \pm 0.03$ (n=6, 1 $\sigma_{m}$ )	$0.24 \pm 0.01$ (n=6, 1 $\sigma_{\rm m}$ )	0.95	this study
Andesne	JA3	9.33	22.80	2.40	12.30	3.05	0.82	2.96	0.52	3.01	0.51	1.57	0.28	2.16	0.27	0.83	GSJ
		$8.91 \pm 0.25$ (n=6, 1 $\sigma_{\rm m}$ )	$21.0 \pm 3.9$ (n=6, 1 $\sigma_{\rm m}$ )	$2.74 \pm 0.05$ (n=6, 1 $\sigma_{\rm m}$ )	$12.2 \pm 0.2$ (n=6, 1 $\sigma_{\rm m}$ )	$3.02 \pm 0.04$ (n=6, 1 $\sigma_{\rm m}$ )	$0.82 \pm 0.02$ (n=6, 1 $\sigma_{\rm m}$ )	$3.23 \pm 0.03$ (n=6, 1 $\sigma_{\rm m}$ )	$0.51 \pm 0.00$ (n=6, 1 $\sigma_{\rm m}$ )	$3.21 \pm 0.04$ (n=6, 1 $\sigma_{\rm m}$ )	$0.65 \pm 0.01$ (n=6, 1 $\sigma_{\rm m}$ )	$1.99 \pm 0.04$ (n=6, 1 $\sigma_{\rm m}$ )	$0.29 \pm 0.00$ (n=6, 1 $\sigma_{\rm m}$ )	$1.94 \pm 0.03$ (n=6, 1 $\sigma_{\rm m}$ )	$0.29 \pm 0.01$ (n=6, 1 $\sigma_{\rm m}$ )	0.80	this study
-	AGV2	38.00	68.00	8.30	30.00	5.70	1.54	4.69	0.64	3.60	0.71	1.79	0.26	1.60	0.25	0.91	USGS
		$29.3 \pm 0.4$ (n=6, 1 $\sigma$ .)	$74.6 \pm 10.5$ (n=6, 1 $\sigma$ .)	$6.26 \pm 0.08$ (n=6, 1 $\sigma$ .)	$24.2 \pm 0.25$ (n=6, 1 $\sigma$ .)	$4.39 \pm 0.05$ (n=6, 1 $\sigma$ .)	$1.41 \pm 0.16$ (n=6, 1 $\sigma$ .)	$4.55 \pm 0.56$ (n=6.15.)	$0.78 \pm 0.07$ (n=6, 1 $\sigma$ .)	$4.64 \pm 0.73$ (n=6, 1 $\sigma$ .)	$1.01 \pm 0.12$ (n=6, 1 $\sigma$ .)	$271 \pm 0.44$ (n=6, 1 $\sigma$ .)	$0.42 \pm 0.05$ (n=6, 1 $\sigma$ )	$2.57 \pm 0.22$ (n=6, 1 $\sigma$ .)	$0.39 \pm 0.05$ (n=6, 1 $\sigma$ .)	0.96	this study
	G2	80 N	160	18 A	55 N	7 20	1 40	<u>( 0, 10m</u> )	0 48	2 40	0.40	( 0, 10m)	0.18	0.48	0 11	0.78	USCS
	02	96 1 + 12 1	160 + 2	16.1 + 1.1	53.0 + 1.3	7 18 + 0.06	1.70 1.55 + 0.00	4 07 + 0 17	0.47 + 0.01	2 13 + 0 16	0.32 + 0.04	0.92 0.89 + 0.02	0.10	0. <b>0</b>	$0.07 \pm 0.02$	0.70	0000
		$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	0.86	this study

	GSP2	180	410	51.0	200	27.0	2.30	12.0	1.40	6.10	1.00	2.20	0.30	1.60	0.23	0.39	USGS
		$176 \pm 19$	$453\pm30$	$53.1\pm4.4$	$197\pm15$	$25.6\pm1.5$	$2.30\pm0.11$	$12.1\pm0.6$	$1.29\pm0.06$	$5.82\pm0.27$	$0.94\pm0.03$	$2.34\pm0.09$	$0.27\pm0.01$	$1.61\pm0.03$	$0.22\pm0.01$	0.40	this study
		$(n=4, 1\sigma_m)$	(n=4, 1 $\sigma_{m}$ )	(n=4, 1 $\sigma_m$ )	(n=4, 1 $\sigma_m$ )	(n=4, $1\sigma_m$ )	(n=4, 1 $\sigma_m$ )	(n=4, 1 $\sigma_{m}$ )	(n=4, 1 $\sigma_{m}$ )	(n=4, 1 $\sigma_{m}$ )	(n=4, 1 $\sigma_{m}$ )	(n=4, $1\sigma_m$ )	(n=4, 1 $\sigma_m$ )	(n=4, 1 $\sigma_m$ )	(n=4, $1\sigma_{m}$ )	0.40	uns study
	JG1a	21.30	45.00	5.63	20.40	4.53	0.70	4.10	0.81	4.44	0.82	2.57	0.38	2.70	0.44	0.50	GSJ
Granite		$20.9\pm0.69$	$45.0\pm1.9$	$5.02\pm0.11$	$19.2\pm0.5$	$4.32\pm0.11$	$0.68\pm0.02$	$4.21\pm0.16$	$0.72\pm0.03$	$4.70\pm0.24$	$0.96\pm0.06$	$2.90\pm0.17$	$0.44\pm0.03$	$2.98 \pm 0.22$	$0.43\pm0.03$	0.48	this study
		$(n=5, 1\sigma_m)$	$(n=5, 1\sigma_m)$	(n=5, $1\sigma_m$ )	(n=5, 1 $\sigma_m$ )	$(n=5, 1\sigma_m)$	(n=5, 1 $\sigma_{m}$ )	$(n=5, 1\sigma_m)$	$(n=5, 1\sigma_m)$	$(n=5, 1\sigma_m)$	$(n=5, 1\sigma_m)$	$(n=5, 1\sigma_m)$	$(n=5, 1\sigma_m)$	(n=5, 1 $\sigma_{m}$ )	$(n=5, 1\sigma_m)$	0.40	uns study
	JG2	19.90	48.30	6.20	26.40	7.78	0.10	8.01	1.62	10.50	1.67	6.04	1.16	6.85	1.22	0.04	GSJ
		$19.0\pm0.72$	$49.6\pm1.3$	$6.00\pm0.26$	$24.4\pm0.75$	$7.97 \pm 0.22$	$0.08\pm0.02$	$9.63\pm0.47$	$1.87\pm0.13$	$12.4\pm0.8$	$2.69\pm0.17$	$8.01\pm0.38$	$1.26\pm0.07$	$8.11\pm0.39$	$1.21\pm0.07$	0.03	this study
		$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	(n=4, $1\sigma_{m}$ )	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	(n=4, $1\sigma_{m}$ )	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	0.05	uns study
	JG3	20.60	40.30	4.70	17.20	3.39	0.90	2.92	0.46	2.59	0.38	1.52	0.24	1.77	0.26	0.87	GSJ
		$19.7\pm0.23$	$37.2\pm4.8$	$4.47\pm0.07$	$16.9\pm0.03$	$3.20\pm0.03$	$0.91\pm0.02$	$2.95\pm0.02$	$0.44\pm0.01$	$2.64\pm0.03$	$0.50\pm0.00$	$1.58\pm0.01$	$0.23\pm0.00$	$1.60\pm0.02$	$0.24\pm0.00$	0.90	this study
		$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	(n=6, 1 $\sigma_{\rm m}$ )	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	(n=6, 1 $\sigma_{\rm m}$ )	$(n=6, 1\sigma_m)$	0.90	uns study						
	W2a	10.00	23.00	-	13.00	3.30	1.00	-	0.63	3.60	0.76	2.50	0.38	2.10	0.33	-	USGS
diabase		$10.0\pm0.07$	$20.6\pm2.8$	$2.86\pm0.02$	$12.7\pm0.06$	$3.19\pm0.04$	$1.12\pm0.02$	$3.76\pm0.05$	$0.60\pm0.01$	$3.74\pm0.04$	$0.72\pm0.01$	$2.23\pm0.01$	$0.32\pm0.00$	$2.02\pm0.02$	$0.30\pm0.01$	0.98	this study
		$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	(n=6, 1 $\sigma_{\rm m}$ )	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	(n=6, 1 $\sigma_{\rm m}$ )	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$	$(n=6, 1\sigma_m)$		uns study
	DNC1a	3.60		1.20	5.20	1.41	0.59	2.00	0.42	3.00	0.62	1.70	0.33	2.00	0.27	1.07	USGS
dolerite		$4.20\pm0.39$	$10.3\pm0.9$	$1.37\pm0.12$	$6.31\pm0.57$	$1.80\pm0.17$	$0.77\pm0.08$	$2.61\pm0.25$	$0.48\pm0.04$	$3.40\pm0.25$	$0.74\pm0.03$	$2.47\pm0.22$	$0.37\pm0.03$	$2.44\pm0.20$	$0.38\pm0.03$	1.08	this study
		$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	$(n=4, 1\sigma_m)$	1.00	uns study
	STM2	154	256	25.0	81.0	12.0	3.45	8.00	1.38	8.01	1.55	4.40	0.55	4.20	0.60	1.07	USGS
syenite		$153 \pm 5$	258 ± 17	$26.8\pm0.9$	79.5 ± 7	$12.5 \pm 0.3$	$3.72 \pm 0.08$	$12.5 \pm 0.1$	$1.30 \pm 0.28$	$7.90 \pm 0.11$	$1.54\pm0.07$	$4.25\pm0.05$	$0.66 \pm 0.04$	$0.69 \pm 0.13$	$0.06 \pm 0.00$	1.01	this study
		$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, l\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, l\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	1101	
	JGb1	3.60	8.17	1.13	5.47	1.49	0.62	1.61	0.29	1.56	0.33	1.04	0.16	1.06	0.15	1.22	GSJ
		$3.40 \pm 0.09$	$7.55 \pm 1.2$	$1.12 \pm 0.03$	$5.29 \pm 0.14$	$1.40 \pm 0.04$	$0.61 \pm 0.01$	$1.52 \pm 0.19$	$0.27 \pm 0.01$	$1.70\pm0.12$	$0.33\pm0.03$	$1.00\pm0.07$	$0.14 \pm 0.01$	$0.92 \pm 0.04$	$0.13 \pm 0.01$	1.27	this study
gabbro —		$(n=5, 1\sigma_m)$	$(n=5, 1\sigma_m)$	$(n=5, l\sigma_m)$	$(n=5, l\sigma_m)$	$(n=5, l\sigma_m)$	$(n=5, 1\sigma_m)$	$(n=5, 1\sigma_m)$	$(n=5, l\sigma_m)$	$(n=5, 1\sigma_m)$	$(n=5, l\sigma_m)$	$(n=5, 1\sigma_m)$	$(n=5, l\sigma_m)$	$(n=5, 1\sigma_m)$	$(n=5, 1\sigma_m)$	1127	
0	JGb2	1.50	3.00	0.39	1.80	0.51	0.59	0.48	0.15	0.60	0.15	0.36	0.06	0.39	0.06	3.63	GSJ
		$1.28 \pm 0.13$	$2.27 \pm 0.42$	$0.37 \pm 0.01$	$1.74 \pm 0.04$	$0.49 \pm 0.01$	$0.59 \pm 0.00$	$0.58 \pm 0.06$	$0.10 \pm 0.03$	$0.63 \pm 0.02$	$0.12 \pm 0.01$	$0.40 \pm 0.02$	$0.06 \pm 0.00$	$0.39 \pm 0.00$	$0.06 \pm 0.00$	3.38	this study
		$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	2.00	into occurry

<sup>1)</sup> The magnitude of Eu anomaly is defined as s the ratio Eu<sub>N</sub>/Eu\* where Eu\* is SQRT(Sm<sub>N</sub> x Gd<sub>N</sub>). The magnitude was calculated based on the values of Sm, Eu and Gd from the reference.

<sup>2)</sup> The bold number of this raw are recommended value by United States of Geological Survey (USGS) or Geological Survey of Japan (GSJ)