Title : Geochemical implication of Eu isotopic ratio in anorthosite: new evidence of Eu isotope fractionation during feldspar crystallization

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This manuscript has been submitted for publication in "Geosciences Journal". Please note that, despite having undergone peer-review, the manuscript has yet to be formally accepted for publication. Subsequent versions of this manuscript may have slightly different content. If accepted in "Geosciences Journal", the final version of this manuscript will be available via the 'Peer-reviewed Publication DOI' link on the right-hand side of this webpage. Please feel free to contact any of the authors; we welcome feedback.

1	Geochemical implication of Eu isotopic ratio in anorthosite: new
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23 ABSTRACT: Rare earth element geochemistry can provide critical information on the evolution of the crust-mantle system. Europium (Eu) exists in divalent and trivalent states, 24 and Eu^{2+} can substitute for Ca^{2+} during plagioclase feldspar crystallization in reducing 25 magmas. This leads to positive Eu anomaly in Ca-plagioclase-rich anorthosite derived from 26 the mantle and negative Eu anomalies in fractionated silica-rich crustal rocks. While many 27 studies have addressed Eu anomalies in REE data, especially in igneous rocs, almost none 28 have evaluated ratios of Eu's two stable isotopes (¹⁵¹Eu and ¹⁵³Eu) alongside Eu anomalies. 29 Here we report systematic variation of the Eu isotopic ratio (δ^{153} Eu) from igneous rocks 30 including anorthosite. This study detected a linear relationship between Eu anomalies and Eu 31 isotopic ratios. Rhyolites and highly fractionated granites exhibited large negative Eu 32 anomalies and negative δ^{153} Eu values while anorthosites exhibited large positive Eu 33 anomalies and positive δ^{153} Eu values. In the case of the highly fractionated igneous rocks 34 with negative Eu anomaly, the Eu isotope fractionation slope varied according to the degree 35 of magmatic differentiation for both extrusive and intrusive rocks. Our finding reveals that Eu 36 isotope fractionation in igneous rocks can provide new information related to magmatic 37 38 differentiation and plagioclase feldspar fractional crystallization including anorthosite formation. 39

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41 Key words: Eu isotope fractionation, Eu anomaly, magmatic differentiation, feldspar
42 crystallization, anorthosite

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45 1. INTRODUCTION

46

Rare earth elements (REEs) and their radiogenic isotopes (especially the ¹⁴⁷Sm-¹⁴³Nd and 47 ¹³⁸La-¹³⁸Ce system) have provided critical constraints on the geochemical evolution of Earth 48 and extra-terrestrial materials. Due to their similar chemical behavior and continuously 49 varying atomic masses, REEs normalized to standard materials provide valuable petrogenetic 50 51 information on magma evolution processes such as partial melting from the mantle-derived rocks, generalized crystallization from magma (Coryell et al., 1963; Fowler and Doig, 1983; 52 Masuda, 1962; Shearer and Papike, 1989; Weill and Drake, 1973) and fractional 53 crystallization (e.g., Bowen, 1928). 54

55 Most REEs exist in the trivalent (+3) state in natural systems, but Eu can exist in both divalent and trivalent states under magmatic redox conditions. Oxygen fugacity, temperature 56 and crystallization of minerals can induce Eu anomalies in igneous rocks (Burnham et al., 57 2015). Positive or negative Eu anomalies in igneous rocks are generally interpreted as 58 reflecting respective feldspar (particularly plagioclase) removal or accumulation during 59 magma evolution with the size of the anomaly interpreted as indicating the degree of 60 61 differentiation of the source magma (Fowler and Doig, 1983; Shearer and Papike, 1989; Weill 62 and Drake, 1973). For example, extremely large positive Eu anomaly in the anorthosite reflect concentration of Eu substitution in the Ca site of plagioclase feldspar as the Ca^{2+} site 63 in feldspar readily accepts Eu²⁺. However, highly fractionated granite and high-silica rhyolite 64 65 also exhibit large negative Eu anomalies.

The two stable isotopes of Eu, ¹⁵¹Eu (47.81%) and ¹⁵³Eu (52.19%), occur in relatively 66 similar natural abundances (Rosman and Taylor, 1998). Though 151 Eu decayed to 147 Pm by α 67 decay with the half-life $T_{1/2}=5\times10^{18}$ yr (Belli et al., 2007), ¹⁵¹Eu can be considered as a stable 68 isotope on Earth and solar system time scales. In addition, recently, Lee and Tanaka (2021a) 69 reported Eu isotope fractionation due to light Eu isotope enrichment (¹⁵¹Eu) in highly 70 fractionated granite and high-silica rhyolite with large negative Eu anomalies. The report 71 proposed that the heavier Eu isotope (¹⁵³Eu) enrichment in anorthosite exhibiting large Eu 72 73 positive anomalies may reflect Ca-feldspar crystallization processes. Recent reports of REE isotope fractionation addressing Ce, Nd, Sm, Dy, Er and Yb (Moynier et al., 2006; Nakada et 74 al., 2013; Shollenberger and Brebbecka, 2020; Hu et al., 2021) detected multiple isotope 75 76 fractionation effects among these elements except in cases where natural systems host only a single isotope. This may indicate existence of δREE based on the isotope fractionation of 10 77 REEs with multiple isotopes except four mono-isotope REEs 78

No report have addressed Eu anomalies and Eu isotopic ratios in anorthosite or other forms of gabbro or volcanic rocks such as andesite and trachyte. This work describes Eu isotope ratios and Eu anomalies among both intrusive and extrusive igneous rocks including anorthosites, gabbro, andesite and trachyte, and further interprets Eu isotopic ratios among basalt, rhyolite and granitoids.

The objective in this study were to evaluate Eu isotope fractionation in terms of Eu anomalies in various kinds of igneous rocks for potential use as a geochemical tracer or petrogenetic proxy.

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88 2. SAMPLES AND EXPERIMENTAL METHODS

89 **2.1.Samples**

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In order to establish a comprehensive picture of Eu isotopic variation in igneous rocks, a 91 92 total of 49 igneous rock samples were analyzed for Eu isotope ratios and REE abundances. Of these, 25 samples represented geochemical reference materials purchased from the U.S. 93 Geological Survey (USGS) and the Geological Survey of Japan (GSJ), while the other 94 95 samples consisted of anorthosites, granitoids and trachytes from Korea and Antarctica. The 25 geochemical reference materials in this study were as follows; Seven basalts (BCR2, 96 BHVO2 and BIR1a purchased from the USGS; JB1a, JB1b, JB2 and JB3 from the GSJ), four 97 andesites (AGV2 from USGS; JA1, JA2 and JA3 from GSJ), four rhyolites (RGM2 from 98 USGS; JR1, JR2 and JR3 from GSJ), one diabase (W-2a from USGS), one dolerite (DNC1a 99 100 from USGS), two gabbros (JGb1, JGb2 from GSJ), one syenite (STM2 from USGS), and five granites (G2 and GSP2 from USGS; JG1a, JG2 and JG3 from GSJ). The 24 rock samples 101 102 from Korea and Antarctica consisted of fourteen granites and five anorthosites from Korea, 103 and five trachytes from Antarctica. Absent a SRM trachyte, we used five trachytes from Antarctica (Table 1, see Appendices for more information). 104

105

106 "Table 1 is about here."

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108 2.2.Sample preparation procedures for determination Eu isotope ratio determination

109 Sample digestion procedures followed an approach in Lee et al (2016). Approximately 100-200 mg of each sample powder was dissolved in a 2:1 mixture of 2-4 mL of concentrated 110 111 HF (29M) and 1-2 mL of concentrated HNO₃ (16M) at ca. 160 °C for more than 72 hours in 15 mL Savillex vials. After the addition of 0.1-0.2 mL of concentrated HClO₄, the dissolved 112 sample solution was heated to dryness at ca. 180 °C for more than 1 day. The cakes were re-113 dissolved in a mixture of 1 mL concentrated HCl and 0.5 mL concentrated HNO₃ and then 114 dried ca. 160 °C for 1 day. Sample residues were re-dissolved in 10 ml 6 M HCl as a stock 115 solution. Of this, 0.5-1 ml was used to determine REE concentrations and the remainder was 116 used for Eu isotope ratio determination. 117

Prior to Eu purification, REE concentration were measured in each sample using inductively coupled plasma mass spectrometry (ICP-MS, NexION350, Perkin Elmer) at Korea Institute of Geoscience and Mineral Resources (KIGAM). Although REE concentration have been previously determined for the 25 geochemical reference rocks (USGS, GSJ) have previously been characterized, we reanalyzed their REE abundances for comparison and to establish analytical integrity. The analyzed REE data from the geochemical standard reference material (SRM) agreed within 5-10% (Table 1).

In this study, Eu was separated from the REE fraction using 0.12 M 2-hydroxyisobutyric acid (HIBA) with the pH adjusted to ~4.60 (Lee and Tanaka, 2019, 2021b). Nuryno et al. (1998) showed that the elution time and resolution during the REE separation depend on HIBA pH. The authors recommend pH 4.6 as optimal for REE separation by HIBA column chromatography. In addition, Lee and Tanaka (2021a, 2021b) showed that incomplete Eu purification by trace amounts of Gd can interfere with Eu isotopic ratio measurement to present pseudo-fractionation effect. Extraction procedures gave yields of 99.9 \pm 0.1% Eu except in case of highly fractionated, silica-rich igneous rocks such as granite (JG2, MA10)
and rhyolites (RGM2, JR series). Highly fractionated igneous rocks gave Eu yields of 98.5±
0.15%.

Most igneous rocks analyzed in this study yielded high purity Eu fractions of about 40-700 ng/mL. These Eu fractions can be used at concentration of 10 ng/mL or higher when measuring Eu isotope ratios by MC-ICP-MS after dilution with 4~7 mL of 2% HNO₃ (Table 2). Given the high recovery, most Eu fractions in this study showed only negligible Gd impurities. However, because highly fractionated igneous rocks may contain Gd, we reperformed HIBA column chromatography after drying and resuspension of initial Eu fraction.

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142 2.3. Mass spectrometry for determination of Eu isotope ratio

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Recently, Lee and Tanaka (2019, 2021b) developed a highly precise and accurate 144 method for Eu isotopic ratio determination using a Sm spike as an internal standard in 145 combination with standard-sample-standard bracketing mass bias correction (C-SSBIN). Lee 146 et al. (2019) found that Eu isotopic ratio determined by C-SSBIN using Sm internal 147 standardization in NIST3117a gave relatively accurate and precise estimates for Eu 148 concentration down to 5ng/mL. Lee et al. (2021a, 2021b) describe optimal conditions for Eu 149 isotope ratio measurement by MC-ICP-MS wherein ¹⁵⁴Gd exerts less than 0.1% on ¹⁵⁴Sm. 150 We, accordingly, minimized isobaric interference using Sm from commercial ultrapure 151 Sm₂O₃ (Alfa Aesar) as a spike for normalizing Eu isotopic measurements. We also monitored 152 tail effects from both Gd and Sm. The separated Eu fraction showed no traces of Sm and only 153

data associated with a ratio of 154 Gd/ 154 Sm of 0.001 (0.1%) or less were further interpreted for degree of Eu isotope fractionation.

Eu isotopic ratios were measured in low-resolution mode with the Ni or Pt X-sampling 156 and Ni X-skimmer cones using multicollector inductively coupled plasma mass spectrometry 157 (MC-ICP-MS; Neptune Plus, Thermo Fisher Scientific Ltd.) in static mode with nine Faraday 158 cups (amplifier resistor: 10¹¹ ohm) at KIGAM. The instrument was tuned to achieve high 159 sensitivity while maintaining flattened, square peaks and stable signals enough to ensure 160 accurate measurements. The gain on each Faraday cup was monitored daily to ensure 161 normalization of its efficiency. The isotopes ¹⁴⁷Sm(L4), ¹⁴⁹Sm(L3), ¹⁵⁰Sm(L2), ¹⁵¹Eu(L1), 162 ¹⁵²Sm(C), ¹⁵³Eu(H1), ¹⁵⁴Sm(H2), ¹⁵⁵Gd(H3), and ¹⁵⁷Gd(H4) were monitored simultaneously 163 using nine Faraday cups for Sm normalization and Gd interference correction by the Gd 164 matrix (Lee and Tanaka, 2021b, Table 1). These parameters gave ¹⁵¹Eu and ¹⁵⁴Sm sensitivities 165 of 80-100mV/ppb and 45-50mV/ppb, respectively. Polyatomic isotopes by Eu, Sm and Gd 166 oxides did not appear during MC-ICP-MS analysis. 167

168

169 "Table 2 is about here."

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Samples subjected to Eu isotope measurement by MC-ICP-MS was suspended in 2%
HNO₃ prepared from 60% ultrapure HNO₃ (Merck, Darmstadt, Germany) and DIW (Milli-Q
system, Millipore, Milford, USA). We used a diluted solution of NIST 3117a (10,000 µg/mL,
Lot No. 120705) as an in-house standard for Eu isotope ratios. Data acquisition consisted of 1
block of 50 cycles with an integration time of 4.194 seconds and a sample aspiration rate of

80-100 µL/min. Peak centering was performed at the beginning of each analysis. A 250 s of
washout time was used between sample measurements. Blanks were checked during, before,
and after each sample measurement. After washing, ¹⁵¹Eu, ¹⁵⁴Sm and ¹⁵⁵Gd in the blank
solutions gave peaks of less than 0.15 mV, 0.11 mV and 0.02mV, respectively. This gives
concentrations of about 6 pg/mL for each REE in acids used during column chromatography.
Blank corrections subtracted each procedural blank from each REE measurement.

Operating conditions and data acquisition parameters including cup configuration followed those given in Lee and Tanaka (2021a, 2021b). We used ¹⁴⁷Sm/¹⁴⁹Sm (1.0868, Dubois et al., 1992) rather than ¹⁵⁰Sm/¹⁵⁴Sm to normalize Eu during Eu isotopic ratio determination. This gave more accurate and precise Eu isotope fractionation estimates for geological rocks subject to incomplete Eu purification and ¹⁵⁴Gd interference during ¹⁵⁰Sm/¹⁵⁴Sm normalization (Lee and Tanaka, 2021a, 2021b).

Eu isotope fractionation is represented in standard δ -notation in per mil relative to the 188 δ^{153} Eu follows: NIST3117a Eu standard solution as = 1,000 189 × $[(^{153}Eu/^{151}Eu_{sample})/(^{153}Eu/^{151}Eu_{NIST3117a}) - 1]$. Lee and Tanaka (2019, 2021c) report that 190 NIST3117a Eu standard reagent analyzed by MC-ICP-MS using Sm internal standard does 191 not exhibit any Eu isotope fractionation regardless of Sm isotope pair used for normalization. 192

193

194 **3. RESULTS**

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Table 3 lists REE abundances for USGS and GSJ SRMs analyzed in this study. Table 4
lists REE concentrations measured in igneous rocks from Korea and Antarctica including

198	anorthosites.
199	
200	"Tables 3 and 4 are about here."
201	
202	Figures 1 and 2 show chondrite-normalized REE patterns for SRMs and the Korea and
203	Antarctic rock samples. Figure 1a-c shows REE patterns from extrusive rocks such as basalt,
204	andesite and rhyolite, while Fig. 1d-f shows REE patterns from intrusive rocks such as
205	gabbro and granitoids.
206	
207	"Figures 1 and 2 are about here."
208	
209	Chondrite-normalized REE patterns from various kinds of igneous rocks (Figs. 1 and 2)
210	clearly show variation in the magnitude of Eu anomaly attributable to feldspar crystallization
211	during magmatic differentiation even though the samples are not cogenetic. The anorthosites
212	in Fig. 2a, in particular, exhibit a strikingly large Eu positive anomaly.
213	Table 5 lists Eu isotope ratios and the magnitude of the Eu anomaly from various kinds
214	of the igneous rocks. Table 3 also divides the samples into SRM and local igneous rocks for
215	comparison. Positive δ^{153} Eu values for the anorthosite indicated enrichment in the heavier Eu
216	(^{153}Eu) relative to the lighter Eu isotope (^{151}Eu) .
217	

218 "Table 5 is about here."

219

220 **4. DISCUSSION**

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This study sought to the relationship between Eu isotope fractionation and the 222 magnitude of Eu anomaly produced from feldspar crystallization during magmatic 223 differentiation. Figure 3 compares the magnitudes of Eu anomalies with δ^{153} Eu values 224 measured from the 49 samples to show Eu isotopic variation with a proxy for magma 225 differentiation. Three geochemical trends emerged from this comparison, First, δ^{153} Eu values 226 of highly fractionated granites and rhyolites attended large negative Eu anomalies, whereas 227 the anorthosites with large positive Eu anomalies showed relatively large positive δ^{153} Eu 228 values. This contrast indicates that the highly fractionated igneous rocks derived from felsic 229 magma in an upper crustal environment were enriched in the lighter Eu isotope (¹⁵¹Eu), 230 whereas Ca-plagioclase-rich anorthosites, which derive from the mafic magma in a lower 231 crustal environment, were enriched in the heavier Eu isotope (153 Eu). Second, δ^{153} Eu value 232 varied systematically with magnitude of the Eu anomaly. Third, intrusive rocks (red symbols 233 in Fig. 3) showed different slopes in anomaly vs. isotopic values relative to those of extrusive 234 rocks (green and blue symbols in Fig. 3). 235

236 "Figure 3 is about here."

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238 The Sancheong anorthosites analyzed in this study are plagioclase accumulates consisting of more than 95% of plagioclase. The rock contains less than 5% quartz and less 239 240 than 1% other minerals such as hornblende, biotite and muscovite. More than 98 % of the chemical compositions also consisted of SiO₂, Al₂O₃, CaO and Na₂O (Jeong et al. 1991; 241 Kang et al., 1994) indicating plagioclase-rich rock. Plagioclase also exhibits an extremely 242 large positive Eu anomaly. The positive Eu anomaly in anorthosite likely reflects substitution 243 of Eu^{2+} for Ca^{2+} in plagioclase during differentiation of the anorthositic (primary) magma 244 either in the upper mantle or at the lower crust. 245

Ismail et al. (1998) showed that isotope effects in Eu^{2+}/Eu^{3+} exchange reaction can 246 occur in aqueous solutions during cation exchange chromatography. These results showed 247 enrichment in the heavier isotope 153 Eu is enriched in Eu²⁺, which forms part of the Eu²⁺/Eu³⁺ 248 electron exchange system. Park et al. (2004) reported that the plagioclase of the Hadong-249 Sancheng anorthosite is characterized by a wide range of δ^{18} O values between - 4.4 and 8.2 250 ‰. Such value range indicates that the source magma of anorthosites was under the reduced 251 environment. Dubinina and Borisov (2018) showed that the increasing of the CaO content in 252 magma leaded the decreasing of ¹⁸O in the melt with structural change of oxygen atoms in the 253 melt. In natural melts, δ^{18} O has a tendency to increase from mafic to felsic rocks (Garlick, 254 255 1966). Moreover, the systematic correlation between Eu isotope fractionation and the magnitudes of Eu anomaly from the fractionated igneous rocks and anorthosites indicates that 256 Eu isotope fractionation was closely related to magmatic differentiation processes such as 257 feldspar fractional crystallization. Therefore, though it is difficult to compare directly the 258 conditions in the magma with those in the aqueous solutions, the enrichment of heavier 259 isotope ¹⁵³Eu in the anorthosites with large positive Eu anomaly may be explained due to 260

isotope effects in Eu^{2+}/Eu^{3+} electron exchange system during Ca-plagioclase accumulation in the anorthositic magma.

263 Another interesting feature in Figure 3 is that the trend of Eu isotope fractionation between the extrusive volcanic rocks and intrusive plutonic rocks seems to be different. 264 Recently, Millet et al. (2016) reported that a correlation between enrichment in heavy Ti 265 isotopes and SiO₂ content resulted from the fractional crystallization of Ti-bearing oxides. 266 267 Yuan et al. (2022) also found that Zr isotope fractionation has different trend in different igneous system such as tholeiitic and alkaline series. These results indicate that isotope 268 269 fractionation in heavy metal elements such as Ti and Zr should have a close relationship with the differentiation process of magma series. Although more in-depth studies are needed in the 270 future, such different trend of the Eu isotope fractionation in intrusive and extrusive rocks 271 may provide an important information for understanding the magma evolution in crust-272 mantle system. 273

The behavior of Eu is known to be determined by temperature and oxygen fugacity 274 (Weill and Drake, 1973), and Eu anomalies are controlled by crystal chemistry and magmatic 275 oxidation potential (Philpotts and Schnetler, 1968; Philpotts, 1970). Dauphas et al. (2014) 276 showed that equilibrium iron isotope fractionation is controlled mainly by the redox and 277 278 structural conditions in magma and suggested that magmatic differentiation is the main driver of Fe isotope fractionation in felsic magmas. In addition, Dauphas et al. (2014) proposed that 279 stable isotopes from heterovalent elements, including Eu, may show isotopic variations in 280 281 bulk rocks controlled by the redox and structural conditions in the magma. It means a possibility that a slight difference of the $\delta^{153}\text{Eu}$ values in the plutonic rocks and volcanic 282 rocks may be due to the oxidation potential in the magma. Further study is needed to clarify 283

the relationship between Eu isotope fractionation and oxidation potential in an intrusivemagmatic system.

286 In addition, we may need to notice for the plot of anorthosites and gabbros in Fig. 3. Anorthosite and gabbro are representative gabbroic rocks solidified through fractional 287 crystallization from primary magma, and gabbro also has relatively large positive Eu 288 anomaly. However, the δ^{153} Eu values of the gabbro are negative compared to those of the 289 290 anorthosites. Anorthosite is a unique and enigmatic rock type in Earth system. Although gabbro and anorthosite are plutonic rocks and are included in the gabbroic rock groups, the 291 292 geochemistry and mineralogy of gabbro are equivalent to those of basalt. This similarity may indicate that the trend in Eu isotope fractionation between gabbros and anorthosites is 293 different. However, at present, the reason for the shift in Eu isotope ratio values between the 294 plutonic rocks and volcanic rocks as well as for trend in Eu isotope fractionation between 295 anorthosite and gabbro is uncertain and, therefore, needs further study. 296

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298 5. CONCLUSION

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We compared the magnitude of Eu anomaly in the chondrite-normalized REE pattern from various kinds of igneous rocks (the extrusive rocks and intrusive rocks including anorthosite) and their Eu isotope ratio from the fractionation. The anorthosites having large positive Eu anomalies show a geochemical characteristic of a heavier Eu isotope (¹⁵³Eu) enrichment (i.e., positive δ^{153} Eu value) whereas the rhyolites and highly fractionated granites having large negative Eu anomalies show a geochemical characteristic of a lighter Eu isotope

(¹⁵¹Eu) enrichment (i.e., negative δ^{153} Eu value). Particularly, our results clearly showed that 306 variation of the magnitude of Eu anomaly and Eu isotope fractionation in igneous rocks has 307 308 systematic correlation, suggesting that Eu isotope fractionation in igneous rocks should be produced by feldspar crystallization during magma evolution. In addition, the Eu isotope 309 fractionation in the highly fractionated volcanic and plutonic rocks was proceeded with 310 311 different trend, implying that the Eu isotope fractionation from the intrusive and extrusive magma in the crustal environment may occur under different mechanism or geochemical 312 environment. Anorthosite is a unique and enigmatic rock type in Earth system. Therefore, 313 geochemical characteristic of Eu isotope fractionation in igneous rocks including anorthosite 314 may provide a valuable information for solving the enigma of anorthosite formation in future. 315

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317 ACKNOWLEDEMENTS

This work was supported by the grants from the Principal Research Fund of the Korea Institute of Geoscience and Mineral Resources (GP2020-003, GP2021-006) and the National Research Foundation of Korea (NRF) grant funded by the Korea government (MSIT) (2020R1F1A1075924, NP2020-012) to S-G. Lee. This work was also partly supported by grants from Korea Polar Research Institute Project (PE21050) to M.J. Lee.

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Figure Cations

Fig. 1. Chondrite (McDonough and Sun, 1995)-normalized REE pattern of standard reference materials (SRMs) of USGS and GSJ. (a), (b) and (c) are SRMs for volcanic (extrusive) rocks whereas (d), (e) and (f) are SRMs for intrusive rocks such as gabbro, diabase and granitoids. In this paper, we classified dolerite and diabase as plutonic rocks rather than volcanic rocks. The REE abundances of all SRMs were re-measured from this study (see Table A2). REE patterns by solid black dots were drawn by recommended values for each SRM from USGS and GSJ.

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Fig. 2. Chondrite (McDonough and Sun, 1995)-normalized REE pattern of Korean and
Antarctic igneous rocks. Except (b) trachyte, the others are all collected from Korea. Trachyte
was collected from Antarctica. The REE abundances of Korean anorthosites and Antarctic
trachytes were re-measured from this study. REE patterns of Korean granites are from Lee et
al. (2004, 2006, 2008, 2013).

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Fig. 3. (a) Variation of Eu isotope ratio according to magnitudes of Eu anomalies from igneous rocks. The error bars represent uncertainties (2SD) of the average δ^{153} Eu values from some of igneous rock samples. Lines; solid gray means no Eu anomaly in the chondritenormalized REE pattern; black dotted means none of Eu isotope fractionation; red and blue dotted lines are correlation lines between magnitude of Eu anomaly and degree of Eu isotope fractionation (δ^{153} Eu) of the plutonic rocks and volcanic rocks, respectively. (b) is an enlarged diagram of the rectangle in (a), that is, the plots of igneous rocks with little Eu anomaly







Sample Name	Rock Type	Area	Location and	n and reference							
M17102801-3			74°21'16.37"S	164°44'49.81"E							
J13010105			74°21'16.32"S	164°44'40.62"E							
K16012306	Trachyte	Mt. Melbourne (Antarctica)	74°20'44.53"S	164°42'28.57"E							
M171110-03		· · · · ·	74°20'43.99"S	164°45'33.29"E							
J13010104			74°21'14.40"S	164°43'55.98"E							
SA20170309 1-1			35°27'23.0"N	127°49'38.4"E							
SA20170309 -2			35°27'19.5"N	127°49'33.7"E							
SA20170309-3	Anorthosite	Sancheong (Korea)	35°26'44.2"N	127°47'23.1"E							
SA20170309-6-2			35°27'15.8"N	127°48'42.4"E							
SA20170310 3-3			35°27'55.0"N	127°51'41.5"E							
MA2018											
WA2018		Muamsa-	Lee et al. (2010, 2013)								
MA10		Weolaksan									
WAWR12		(Korea)									
WAWR8											
SM4		Seokmodo	Loo at al	(2006)							
SM26	Granitoid	(Korea)	Lee et al.	(2000)							
ICH3		Icheon	Songjeong-dong, Icheon	Les et al. (2008)							
ICH11		(Korea)	Daecheon-ri, Icheon	Lee et al. (2008)							
C2320		Pohang	L ag at al	(2008a)							
C2980		(Korea)	Lee et al.	(2000a)							
TD13		Taedo	Les et 1 (2010)								
TD2B		(Korea)	Lee et al. (2018)								

Table 1. General information for the rock samples used in this study

Tuble 2. Operating conditions for the	
Instrument settings	
RF power (W)	1200
Plasma Ar gas flow rate (L/min)	16
Auxiliary Ar gas flow rate (L/min)	1
Ar carrier gas flow rate (L/min)	1.04
Sample cone	X Nickel, 1.1 mm orifice (wet plazma) Jet Nickel, 1.1 mm orifice (dry plazma)
Skimmer cone	X-cone; nickel, 0.8 mm orifice X-cone; platinum, 0.8 mm orifice
Sample uptake rate	$\delta U \sim 100 \ \mu L/11111$
Sample uptake Time	150 s
Wash time	250 s (Cetac ASX-110 automatic sampler), 180 s (Aridus II)
Lens settings	Optimized for maximum analyte signal intensity,
	□flat-topped peaks and stability
Data acquisition parameters	
Scan type	Static measurements
Cup configuration	147 Sm(L4) ^{, 149} Sm(L3), 150 Sm(L2), 151 Eu (L1), 152 Sm (C), 153 Eu (H1), 154 Sm (H2), 155 Gd (H3), 157 Gd (H4)
Zoom optics	Focus quad: 6 V and dispersion quad: 0 V
Sourcitie-it	80~100 mV/ ppb for 151 Eu, 45 ~ 50 mV/ ppb for 154 Sm,
Sensitivity	20 ± 2 mV/ ppb for 155 Gd
Integration time	4.19 sec
Number of integrations	1
Cycles/blocks or runs/passes	60

Table 2. Operating conditions for the Neptune Plus MC-ICP-MS

Table 3. Concentrations of Rare earth element of standard reference materials (SRMs) measured in this study

Туре	Sample	La	Ce	Pr	Nd	Sm	Eu	Gd	Tb	Dy	Но	Er	Tm	Yb	Lu	Eu/Eu*1)	Ref
	BCR2 ²⁾	25.0	53.0	6.80	28.0	6.70	2.00	6.80	1.07	6.41	1.33	3.66	0.54	3.50	0.51	0.90	USGS
		$\begin{array}{c} 25.0\pm0.9\\ (n{=}4,1\sigma_m) \end{array}$	53.7 ± 1.1 (n=4, 1 σ_{m})	6.73 ± 0.16 (n=4, 1 σ_{m})	$\begin{array}{c} 28.5\pm0.6\\ (n{=}4,1\sigma_m) \end{array}$	$\begin{array}{l} \textbf{6.58} \pm \textbf{0.15} \\ \textbf{(n=4, 1}\sigma_{m}) \end{array}$	$\begin{array}{l} 2.05\pm0.04\\ (n{=}4,1\sigma_m) \end{array}$	$\begin{array}{l} \textbf{6.76} \pm \textbf{0.17} \\ \textbf{(n=4, 1}\sigma_{m}) \end{array}$	$\begin{array}{c} 1.03 \pm 0.03 \\ (n{=}4, 1\sigma_m) \end{array}$	$\begin{array}{l} \textbf{6.48} \pm \textbf{0.13} \\ \textbf{(n=4, 1}\sigma_{m}) \end{array}$	$\begin{array}{c} 1.32\pm0.03\\ (n{=}4,1\sigma_m) \end{array}$	$\begin{array}{l} 3.77 \pm 0.09 \\ (n{=}4, 1\sigma_m) \end{array}$	$\begin{array}{l} 0.52 \pm 0.01 \\ (n{=}4,1\sigma_m) \end{array}$	3.41 ± 0.05 (n=4, 1 σ_m)	0.51 ± 0.01 (n=4, 1 σ_{m})	0.94	this study
	BHVO2	15.00	38.00	-	25.00	6.20	2.05	6.30	0.90	-	1.04	-	-	2.00	0.28	1.00	USGS
		$\begin{array}{c} 14.9\pm0.44\\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 36.2 \pm 1.80 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 5.17\pm0.18\\(n{=}3,1\sigma_m)\end{array}$	$\begin{array}{c} 23.2\pm1.22\\(n{=}3,1\sigma_m)\end{array}$	$\begin{array}{c} 5.76\pm0.16\\(n{=}3,1\sigma_m)\end{array}$	$\begin{array}{c} 2.05\pm0.10\\(n{=}3,1\sigma_m)\end{array}$	$\begin{array}{c} 5.90 \pm 0.34 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 0.97 \pm 0.07 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{l} 4.89 \pm 0.36 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 0.97 \pm 0.03 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 2.39 \pm 0.17 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 0.26 \pm 0.02 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 1.66 \pm 0.12 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{l} 0.25\pm0.02\\ (n{=}3,1\sigma_m) \end{array}$	1.07	this study
	BIR1a	0.63	1.90	-	2.50	1.10	0.55	1.80	-	4.00	-	-	-	1.70	0.26	1.13	USGS
		$\begin{array}{c} 0.64 \pm 0.05 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 1.99 \pm 0.10 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 0.40 \pm 0.05 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 2.49 \pm 0.08 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 1.12 \pm 0.05 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 0.53 \pm 0.02 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 1.80 \pm 0.23 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 0.36 \pm 0.03 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 2.61 \pm 0.20 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 0.57 \pm 0.05 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 1.72\pm0.17\\(n{=}3,1\sigma_m)\end{array}$	$\begin{array}{c} 0.34 \pm 0.03 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 1.94 \pm 0.05 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{l} 0.26 \pm 0.01 \\ (n{=}3,1\sigma_m) \end{array}$	1.13	this study
basalt	JB1a	37.60	65.90	7.30	26.00	5.07	1.46	4.67	0.69	3.99	0.71	2.18	0.33	2.10	0.33	0.91	GSJ
Dasan		$\begin{array}{c} 33.7\pm0.8\\(n{=}6,1\sigma_m)\end{array}$	$\begin{array}{l} 74.6\pm9.7\\ (n{=}6,1\sigma_m) \end{array}$	$\begin{array}{c} \textbf{6.47} \pm \textbf{0.18} \\ \textbf{(n=6, 1}\sigma_{m}) \end{array}$	$\begin{array}{c} 24.7\pm0.6\\ (n{=}6,1\sigma_m) \end{array}$	4.72 ± 0.11 (n=6, 1 σ_m)	$\begin{array}{c} 1.49 \pm 0.03 \\ (n=\!6,1\sigma_m) \end{array}$	$\begin{array}{l} 4.43 \pm 0.10 \\ (n{=}6, 1\sigma_m) \end{array}$	$\begin{array}{l} 0.66 \pm 0.01 \\ (n{=}6,1\sigma_m) \end{array}$	3.76 ± 0.16 (n=6, 1 σ_m)	$\begin{array}{c} 0.70 \pm 0.02 \\ (n{=}6,1\sigma_m) \end{array}$	$\begin{array}{c} 2.12\pm0.05\\ (n{=}6,1\sigma_m) \end{array}$	$\begin{array}{c} 0.29 \pm 0.01 \\ (n{=}6, 1\sigma_m) \end{array}$	1.91 ± 0.05 (n=6, 1 σ_{m})	$\begin{array}{l} 0.28 \pm 0.01 \\ (n{=}6,1\sigma_m) \end{array}$	0.99	this study
	JB1b	41.20	71.80	7.73	27.10	5.17	1.59	4.38	0.69	3.73	0.67	1.97	0.31	2.10	0.31	1.02	GSJ
		40.20	70.10	7.46	26.90	4.93	1.59	5.35	0.77	4.14	0.79	2.22	0.33	2.01	0.31	0.94	this study
	JB2	2.35	6.76	1.01	6.63	2.31	0.86	3.28	0.60	3.73	0.75	2.60	0.41	2.62	0.40	0.95	GSJ
		$\begin{array}{c} 2.18\pm0.8\\ (n{=}8,1\sigma_m) \end{array}$	$\begin{array}{c} 6.39 \pm 0.14 \\ (n{=}8,1\sigma_m) \end{array}$	1.11 ± 0.05 (n=8, 1 σ_{m})	$\begin{array}{l} 6.26 \pm 0.17 \\ (n{=}8,1\sigma_m) \end{array}$	$\begin{array}{c} 2.22 \pm 0.04 \\ (n{=}8,1\sigma_m) \end{array}$	$\begin{array}{c} 0.83 \pm 0.03 \\ (n=\!8, 1\sigma_m) \end{array}$	$\begin{array}{l} 3.09\pm0.07\\(n{=}8,1\sigma_m)\end{array}$	$\begin{array}{l} 0.56\pm0.01\\ (n{=}8,1\sigma_m) \end{array}$	$\begin{array}{c} 3.77 \pm 0.09 \\ (n{=}8,1\sigma_m) \end{array}$	0.81 ± 0.04 (n=8, 1 σ_{m})	$\begin{array}{c} 2.51 \pm 0.07 \\ (n{=}8, 1\sigma_m) \end{array}$	$\begin{array}{c} 0.37 \pm 0.01 \\ (n{=}8,1\sigma_m) \end{array}$	$\begin{array}{l} 2.42\pm0.06\\ (n{=}8,1\sigma_m) \end{array}$	$\begin{array}{l} 0.37\pm 0.01\\ (n{=}8,1\sigma_m) \end{array}$	0.96	this study
	JB3	8.81	21.50	3.11	15.60	4.27	1.32	4.67	0.73	4.54	0.80	2.49	0.42	2.55	0.39	0.90	GSJ
		$\begin{array}{c} 8.82 \pm 0.70 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 22.0\pm2.5\\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 3.45 \pm 0.35 \\ (n=3, 1\sigma_m) \end{array}$	16.0 ± 1.78 (n=3, 1 σ_{m})	$\begin{array}{l} 4.31 \pm 0.35 \\ (n{=}3,1\sigma_m) \end{array}$	1.41 ± 0.16 (n=3, 1 σ_m)	$\begin{array}{l} 4.55\pm0.56\\(n{=}3,1\sigma_m)\end{array}$	$\begin{array}{c} 0.78 \pm 0.07 \\ (n{=}3,1\sigma_m) \end{array}$	4.64 ± 0.73 (n=3, 1 σ_m)	1.01 ± 0.12 (n=3, 1 σ_{m})	2.71 ± 0.44 (n=3, 1 σ_{m})	$\begin{array}{c} 0.42 \pm 0.05 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{c} 2.57 \pm 0.22 \\ (n{=}3,1\sigma_m) \end{array}$	$\begin{array}{l} 0.39\pm0.05\\ (n{=}3,1\sigma_m) \end{array}$	0.97	this study
	JR1	19.70	47.20	5.58	23.30	6.03	0.30	5.06	1.01	5.69	1.11	3.61	0.67	4.55	0.71	0.17	GSJ
		17.35	43.83	5.58	22.52	5.37	0.27	5.48	0.94	6.09	1.25	4.25	0.70	4.94	0.76	0.15	this study
	JR2	16.30	38.80	4.75	20.40	5.63	0.14	5.83	1.10	6.63	1.39	4.36	0.74	5.33	0.88	0.07	GSJ
		13.72	37.58	4.53	17.77	4.95	0.09	5.29	0.98	6.39	1.41	4.42	0.74	5.08	0.81	0.05	this study
Rhyolite	JR3	179	327	33.1	107	21.3	0.53	19.7	4.29	21.5	4.70	14.0		20.3	2.80	0.06	GSJ
		167 ± 5 (n=4, 1 $\sigma_{\rm m}$)	309 ± 9 (n=4, 1 $\sigma_{\rm m}$)	30.9 ± 1.5 (n=4, 1 $\sigma_{\rm m}$)	101 ± 3 (n=4, 1 $\sigma_{\rm m}$)	20.1 ± 0.7 (n=4, 1 $\sigma_{\rm m}$)	0.41 ± 0.02 (n=4, 1 $\sigma_{\rm m}$)	20.2 ± 0.72 (n=4, 1 σ_{m})	3.99 ± 0.25 (n=4, 1 $\sigma_{\rm m}$)	27.1 ± 1.3 (n=4, 1 $\sigma_{\rm m}$)	6.00 ± 0.33 (n=4, 1 $\sigma_{\rm m}$)	18.3 ± 0.7 (n=4, 1 $\sigma_{\rm m}$)	2.96 ± 0.18 (n=4, 1 $\sigma_{\rm m}$)	19.0 ± 0.8 (n=4, 1 $\sigma_{\rm m}$)	2.71 ± 0.07 (n=4, 1 $\sigma_{\rm m}$)	0.06	this study
	RGM2	25.00	48.00	5.00	20.00	4.00	0.70	3.60	0.60	3.30	0.80	2.20			0.40	0.56	USGS
		22.7 ± 1.8 (n=2, 1 σ_{m})	47.2 ± 1.0 (n=2, 1 $\sigma_{\rm m}$)	5.27 ± 0.31 (n=2, 1 σ_{m})	19.4 ± 0.90 (n=2, 1 σ_{m})	3.91 ± 0.48 (n=2, 1 σ_{m})	0.70 ± 0.03 (n=2, 1 σ_{m})	3.81 ± 0.31 (n=2, 1 σ_m)	$\begin{array}{c} 0.57 \pm 0.02 \\ (n=2, 1\sigma_m) \end{array}$	3.60 ± 0.19 (n=2, 1 σ_{m})	0.73 ± 0.06 (n=2, 1 σ_{m})	2.33 ± 0.09 (n=2, 1 σ_m)	0.36 ± 0.00 (n=2, 1 σ_{m})	0.52 ± 0.09 (n=2, 1 σ_{m})	0.38 ± 0.01 (n=2, 1 $\sigma_{\rm m}$)	0.56	this study
	JA1	5.24	13.30	1.71	10.90	3.52	1.20	4.36	0.75	4.55	0.95	3.04	0.47	3.03	0.47	0.93	GSJ
		$\begin{array}{l} 4.69 \pm 0.06 \\ (n=\!6, 1\sigma_{m}) \end{array}$	11.8 ± 1.52 (n=6, 1 σ_{m})	2.01 ± 0.04 (n=6, 1 $\sigma_{\rm m}$)	10.7 ± 0.2 (n=6, 1 σ_{m})	3.30 ± 0.03 (n=6, 1 σ_{m})	1.16 ± 0.02 (n=6, 1 σ_{m})	4.31 ± 0.05 (n=6, 1 σ_{m})	$\begin{array}{l} 0.72\pm0.01\\ (n=\!6,1\sigma_m) \end{array}$	4.65 ± 0.12 (n=6, 1 σ_m)	0.94 ± 0.01 (n=6, 1 σ_{m})	3.04 ± 0.02 (n=6, 1 σ_{m})	0.44 ± 0.00 (n=6, 1 σ_{m})	2.91 ± 0.02 (n=6, 1 σ_m)	0.44 ± 0.00 (n=6, 1 $\sigma_{\rm m}$)	0.93	this study
	JA2	15.80	32.70	3.84	13.90	3.11	0.93	3.06	0.44	2.80	0.50	1.48	0.28	1.62	0.27	0.92	GSJ
Andesite —		$\begin{array}{c} 14.5\pm0.25\\(n=\!6,1\sigma_{m})\end{array}$	28.8 ± 3.9 (n=6, 1 $\sigma_{\rm m}$)	3.48 ± 0.05 (n=6, 1 σ_{m})	13.8 ± 0.2 (n=6, 1 σ_{m})	2.93 ± 0.04 (n=6, 1 σ_{m})	0.92 ± 0.02 (n=6, 1 σ_{m})	$\begin{array}{l} 2.99 \pm 0.03 \\ (n=\!6, 1\sigma_m) \end{array}$	$\begin{array}{l} 0.46 \pm 0.00 \\ (n=\!6, 1\sigma_m) \end{array}$	2.78 ± 0.04 (n=6, 1 σ_{m})	$\begin{array}{l} 0.54\pm0.01\\ (n=\!6,1\sigma_{m}) \end{array}$	1.69 ± 0.04 (n=6, 1 σ_{m})	$\begin{array}{l} 0.25 \pm 0.00 \\ (n=\!6, 1\sigma_{m}) \end{array}$	$\begin{array}{c} 1.59 \pm 0.03 \\ (n=\!6, 1\sigma_m) \end{array}$	0.24 ± 0.01 (n=6, 1 σ_{m})	0.95	this study
	JA3	9.33	22.80	2.40	12.30	3.05	0.82	2.96	0.52	3.01	0.51	1.57	0.28	2.16	0.27	0.83	GSJ
		$\begin{array}{l} 8.91 \pm 0.25 \\ (n{=}6, 1\sigma_{m}) \end{array}$	21.0 ± 3.9 (n=6, 1 σ_m)	$\begin{array}{l} 2.74 \pm 0.05 \\ (n=\!6,1\sigma_{m}) \end{array}$	12.2 ± 0.2 (n=6, 1 σ_m)	3.02 ± 0.04 (n=6, 1 σ_{m})	$\begin{array}{l} 0.82 \pm 0.02 \\ (n=\!6,1\sigma_{m}) \end{array}$	$\begin{array}{l} 3.23 \pm 0.03 \\ (n=\!6, 1\sigma_m) \end{array}$	0.51 ± 0.00 (n=6, 1 σ_{m})	3.21 ± 0.04 (n=6, 1 σ_m)	$\begin{array}{l} 0.65 \pm 0.01 \\ (n=\!6, 1\sigma_{m}) \end{array}$	1.99 ± 0.04 (n=6, 1 σ_{m})	0.29 ± 0.00 (n=6, 1 σ_{m})	1.94 ± 0.03 (n=6, 1 σ_{m})	0.29 ± 0.01 (n=6, 1 σ_{m})	0.80	this study
	AGV2	38.00	68.00	8.30	30.00	5.70	1.54	4.69	0.64	3.60	0.71	1.79	0.26	1.60	0.25	0.91	USGS
		$\begin{array}{c} 29.3\pm0.4\\ (n{=}6,1\sigma_m) \end{array}$	$\begin{array}{c} 74.6 \pm 10.5 \\ (n=\!6, 1\sigma_m) \end{array}$	$\begin{array}{c} \textbf{6.26} \pm \textbf{0.08} \\ \textbf{(n=6, 1\sigma_m)} \end{array}$	$\begin{array}{c} 24.2 \pm 0.25 \\ (n{=}6, 1\sigma_{m}) \end{array}$	$\begin{array}{c} 4.39 \pm 0.05 \\ (n=\!6, 1\sigma_m) \end{array}$	1.41 ± 0.16 (n=6, 1 σ_{m})	$\begin{array}{c} 4.55\pm0.56\\(n{=}6,1\sigma_m)\end{array}$	$\begin{array}{c} 0.78 \pm 0.07 \\ (n{=}6,1\sigma_m) \end{array}$	$\begin{array}{c} 4.64 \pm 0.73 \\ (n{=}6, 1\sigma_m) \end{array}$	1.01 ± 0.12 (n=6, 1 σ_{m})	271 ± 0.44 (n=6, 1 σ_m)	$\begin{array}{c} 0.42 \pm 0.05 \\ (n=\!6, 1\sigma_m) \end{array}$	$\begin{array}{c} 2.57 \pm 0.22 \\ (n{=}6, 1\sigma_{m}) \end{array}$	$\begin{array}{l} 0.39 \pm 0.05 \\ (n{=}6,1\sigma_m) \end{array}$	0.96	this study
	G2	89.0	160	18.0	55.0	7.20	1.40	4.30	0.48	2.40	0.40	0.92	0.18	0.48	0.11	0.78	USGS
		96.1 \pm 12.1 (n=2, 1 σ_{m})	160 ± 2 (n=2, 1 σ_{m})	16.1 ± 1.1 (n=2, 1 σ_{m})	53.0 ± 1.3 (n=2, 1 σ_m)	$\begin{array}{c} 7.18 \pm 0.06 \\ (n=2, 1\sigma_m) \end{array}$	$\begin{array}{l} 1.55\pm0.09\\ (n{=}2,1\sigma_m) \end{array}$	$\begin{array}{l} 4.07\pm0.17\\(n{=}2,1\sigma_m)\end{array}$	$\begin{array}{l} 0.47 \pm 0.01 \\ (n=2, 1\sigma_m) \end{array}$	2.13 ± 0.16 (n=2, 1 σ_m)	$\begin{array}{c} 0.32\pm0.04\\ (n{=}2,1\sigma_m) \end{array}$	$\begin{array}{c} 0.89 \pm 0.02 \\ (n{=}2,1\sigma_m) \end{array}$	$\begin{array}{c} 0.10 \pm 0.05 \\ (n{=}2,1\sigma_m) \end{array}$	0.66 ± 0.11 (n=2, 1 σ_{m})	0.07 ± 0.02 (n=2, 1 σ_{m})	0.86	this study

	GSP2	180	410	51.0	200	27.0	2.30	12.0	1.40	6.10	1.00	2.20	0.30	1.60	0.23	0.39	USGS
		176 ± 19	453 ± 30	53.1 ± 4.4	197 ± 15	25.6 ± 1.5	2.30 ± 0.11	12.1 ± 0.6	1.29 ± 0.06	5.82 ± 0.27	0.94 ± 0.03	2.34 ± 0.09	0.27 ± 0.01	1.61 ± 0.03	0.22 ± 0.01	0.40	this study.
		(n=4, $1\sigma_{m}$)	(n=4, $1\sigma_m$)	(n=4, 1 σ_{m})	(n=4, $1\sigma_m$)	(n=4, $1\sigma_m$)	(n=4, $1\sigma_m$)	(n=4, $1\sigma_m$)	(n=4, $1\sigma_m$)	(n=4, $1\sigma_m$)	(n=4, $1\sigma_{m}$)	(n=4, $1\sigma_m$)	$(n=4, 1\sigma_m)$	(n=4, 1 _s)	(n=4, $1\sigma_m$)	0.40	uns study
	JG1a	21.30	45.00	5.63	20.40	4.53	0.70	4.10	0.81	4.44	0.82	2.57	0.38	2.70	0.44	0.50	GSJ
Granite		20.9 ± 0.69	45.0 ± 1.9	5.02 ± 0.11	19.2 ± 0.5	4.32 ± 0.11	0.68 ± 0.02	4.21 ± 0.16	0.72 ± 0.03	4.70 ± 0.24	0.96 ± 0.06	2.90 ± 0.17	0.44 ± 0.03	2.98 ± 0.22	0.43 ± 0.03	0.40	4 1
		(n=5, $1\sigma_{m}$)	(n=5, $1\sigma_m$)	(n=5, $1\sigma_m$)	$(n=5, 1\sigma_m)$	$(n=5, 1\sigma_m)$	(n=5, $1\sigma_m$)	(n=5, 1 _s)	(n=5, 1 _m)	$(n=5, 1\sigma_m)$	$(n=5, 1\sigma_m)$	(n=5, 1 σ_m)	$(n=5, 1\sigma_m)$	(n=5, $1\sigma_m$)	(n=5, $1\sigma_{m}$)	0.48	this study
	JG2	19.90	48.30	6.20	26.40	7.78	0.10	8.01	1.62	10.50	1.67	6.04	1.16	6.85	1.22	0.04	GSJ
		19.0 ± 0.72	49.6 ± 1.3	6.00 ± 0.26	24.4 ± 0.75	7.97 ± 0.22	0.08 ± 0.02	9.63 ± 0.47	1.87 ± 0.13	12.4 ± 0.8	2.69 ± 0.17	8.01 ± 0.38	1.26 ± 0.07	8.11 ± 0.39	1.21 ± 0.07	0.02	4 1
		(n=4, 1 $\sigma_{\rm m}$)	(n=4, 1 σ_{m})	(n=4, 1 _m)	(n=4, $1\sigma_m$)	(n=4, $1\sigma_{m}$)	(n=4, $1\sigma_m$)	(n=4, 1 σ_{m})	(n=4, $1\sigma_m$)	(n=4, 1 _s)	(n=4, 1 σ_{m})	(n=4, 1 σ_{m})	(n=4, 1 σ_{m})	(n=4, $1\sigma_m$)	(n=4, 1 $\sigma_{\rm m}$)	0.03	this study
	JG3	20.60	40.30	4.70	17.20	3.39	0.90	2.92	0.46	2.59	0.38	1.52	0.24	1.77	0.26	0.87	GSJ
		19.7 ± 0.23	37.2 ± 4.8	4.47 ± 0.07	16.9 ± 0.03	3.20 ± 0.03	0.91 ± 0.02	2.95 ± 0.02	0.44 ± 0.01	2.64 ± 0.03	0.50 ± 0.00	1.58 ± 0.01	0.23 ± 0.00	1.60 ± 0.02	0.24 ± 0.00	0.00	
		(n=6, 1 $\sigma_{\rm m}$)	(n=6, 1 σ_m)	(n=6, 1 _m)	(n=6, 1 _m)	(n=6, 1 σ_{m})	(n=6, 1 _m)	(n=6, 1 \sigma_m)	(n=6, 1 _m)	(n=6, 1 σ_m)	(n=6, 1 σ_m)	(n=6, 1 σ_m)	(n=6, 1 σ_m)	(n=6, 1 _m)	(n=6, 1 $\sigma_{\rm m}$)	0.90	this study
	W2a	10.00	23.00	-	13.00	3.30	1.00	-	0.63	3.60	0.76	2.50	0.38	2.10	0.33	-	USGS
diabase		10.0 ± 0.07	20.6 ± 2.8	2.86 ± 0.02	12.7 ± 0.06	3.19 ± 0.04	1.12 ± 0.02	3.76 ± 0.05	0.60 ± 0.01	3.74 ± 0.04	0.72 ± 0.01	2.23 ± 0.01	0.32 ± 0.00	2.02 ± 0.02	0.30 ± 0.01	0.00	4 1
		(n=6, 1 $\sigma_{\rm m}$)	(n=6, 1 _m)	(n=6, 1 _m)	(n=6, 1 \sigma_m)	(n=6, 1 σ_m)	(n=6, 1 _m)	(n=6, 1 _m)	(n=6, 1 _m)	(n=6, 1 _m)	(n=6, 1 _m)	(n=6, 1 _m)	(n=6, 1 _m)	(n=6, 1 \sigma_m)	(n=6, 1 $\sigma_{\rm m}$)	0.98	this study
	DNC1a	3.60		1.20	5.20	1.41	0.59	2.00	0.42	3.00	0.62	1.70	0.33	2.00	0.27	1.07	USGS
dolerite		4.20 ± 0.39	10.3 ± 0.9	1.37 ± 0.12	6.31 ± 0.57	1.80 ± 0.17	0.77 ± 0.08	2.61 ± 0.25	0.48 ± 0.04	3.40 ± 0.25	0.74 ± 0.03	2.47 ± 0.22	0.37 ± 0.03	2.44 ± 0.20	0.38 ± 0.03	1.00	41
		(n=4, $1\sigma_{m}$)	$(n=4, 1\sigma_m)$	(n=4, $1\sigma_m$)	(n=4, 1 σ_{m})	(n=4, 1 σ_{m})	(n=4, $1\sigma_m$)	(n=4, 1 σ_{m})	(n=4, $1\sigma_{m}$)	(n=4, $1\sigma_{m}$)	1.08	this study					
	STM2	154	256	25.0	81.0	12.0	3.45	8.00	1.38	8.01	1.55	4.40	0.55	4.20	0.60	1.07	USGS
syenite		153 ± 5	258 ± 17	26.8 ± 0.9	79.5 ± 7	12.5 ± 0.3	3.72 ± 0.08	12.5 ± 0.1	1.30 ± 0.28	7.90 ± 0.11	1.54 ± 0.07	4.25 ± 0.05	0.66 ± 0.04	0.69 ± 0.13	0.06 ± 0.00	1.01	4 1
		$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	(n=2, $1\sigma_m$)	(n=2, $1\sigma_m$)	(n=2, $1\sigma_m$)	(n=2, $1\sigma_m$)	$(n=2, 1\sigma_m)$	(n=2, $1\sigma_m$)	$(n=2, 1\sigma_m)$	1.01	this study					
	JGb1	3.60	8.17	1.13	5.47	1.49	0.62	1.61	0.29	1.56	0.33	1.04	0.16	1.06	0.15	1.22	GSJ
		3.40 ± 0.09	7.55 ± 1.2	1.12 ± 0.03	5.29 ± 0.14	1.40 ± 0.04	0.61 ± 0.01	1.52 ± 0.19	0.27 ± 0.01	1.70 ± 0.12	0.33 ± 0.03	1.00 ± 0.07	0.14 ± 0.01	0.92 ± 0.04	0.13 ± 0.01	1.27	this study.
achbro		$(n=5, 1\sigma_m)$	$(n=5, 1\sigma_m)$	(n=5, $1\sigma_m$)	(n=5, $1\sigma_m$)	$(n=5, 1\sigma_m)$	(n=5, $1\sigma_m$)	$(n=5, 1\sigma_m)$	(n=5, $1\sigma_m$)	$(n=5, 1\sigma_m)$	1.27	this study					
gabbio	JGb2	1.50	3.00	0.39	1.80	0.51	0.59	0.48	0.15	0.60	0.15	0.36	0.06	0.39	0.06	3.63	GSJ
		1.28 ± 0.13	2.27 ± 0.42	0.37 ± 0.01	1.74 ± 0.04	0.49 ± 0.01	0.59 ± 0.00	0.58 ± 0.06	0.10 ± 0.03	0.63 ± 0.02	0.12 ± 0.01	0.40 ± 0.02	0.06 ± 0.00	0.39 ± 0.00	0.06 ± 0.00	2 20	this study.
		$(n=2, 1\sigma_m)$	(n=2, $1\sigma_{m}$)	(n=2, $1\sigma_m$)	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	(n=2, 1 _m)	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	$(n=2, 1\sigma_m)$	(n=2, $1\sigma_{m}$)	(n=2, 1 _m)	(n=2, $1\sigma_{m}$)	$(n=2, 1\sigma_m)$	3.38	unis study

¹⁾ The magnitude of Eu anomaly is defined as s the ratio Eu_N/Eu* where Eu* is SQRT(Sm_N x Gd_N). The magnitude was calculated based on the values of Sm, Eu and Gd from the reference.

²⁾ The bold number of this raw are recommended value by United States of Geological Survey (USGS) or Geological Survey of Japan (GSJ)

Rock Type	Area	Sample Name	La (ppm)	Ce (ppm)	Pr (ppm)	Nd (ppm)	Sm (ppm)	Eu (ppm)	Gd (ppm)	Tb (ppm)	Dy (ppm)	Ho (ppm)	Er (ppm)	Tm (ppm)	Yb (ppm)	Lu (ppm)	Eu/Eu*	Reference
		M17102801-3	78.42	190.5	18.12	68.07	13.17	3.26	11.29	1.80	10.81	2.10	5.79	0.83	5.14	0.66	0.81	
Trachyte		J13010105	103.3	198.7	21.92	80.38	15.09	3.46	13.09	2.04	11.92	2.26	6.25	0.88	5.57	0.83	0.75	
	Mt. Melbourne	K16012306	87.78	164.8	18.74	70.02	13.35	3.76	12.51	1.68	9.89	1.88	5.19	0.71	4.50	0.65	0.89	
	(Antarctica)	M171110-03	86.93	172.4	18.10	70.93	12.98	3.49	12.50	1.77	10.11	1.99	5.58	0.73	4.88	0.72	0.84	
		J13010104	78.42	190.5	18.12	68.07	13.17	3.26	11.29	1.80	10.81	2.10	5.79	0.83	5.14	0.66	0.81	
		SA20170309 1-1	1.64	3.25	0.36	1.41	0.26	0.87	0.23	0.03	0.17	0.03	0.08	0.01	0.06	0.01	10.83	
		SA20170309 -2	1.73	3.18	0.36	1.43	0.27	0.99	0.28	0.04	0.24	0.05	0.13	0.02	0.12	0.02	11.08	This
	Sancheong (Korea)	SA20170309-3	1.65	3.12	0.35	1.47	0.31	0.66	0.33	0.05	0.27	0.05	0.15	0.02	0.12	0.02	6.29	study
	(110104)	SA20170309-6-2	1.42	2.27	0.22	0.77	0.11	1.24	0.08	0.01	0.05	0.01	0.02	0.00	0.02	0.002	38.53	
		SA20170310 3-3	1.41	2.45	0.28	1.17	0.22	1.04	0.26	0.04	0.23	0.05	0.14	0.02	0.13	0.02	13.09	
		MA10	12.42	28.12	4.23	16.35	6.12	0.01	7.61	1.62	12.39	2.92	10.46	1.86	14.55	2.34	0.01	
	Muamsa- Weolaksan (Korea)	WAWR12	20.97	54.60	5.67	21.40	6.56	0.13	8.13	1.59	11.35	2.41	7.57	1.21	8.54	1.25	0.05	
		MA2018	46.39	60.86	11.64	42.91	10.59	0.31	11.08	1.93	13.23	2.66	8.95	1.42	9.67	1.43	0.09	
		WA2018	27.30	62.63	7.88	30.31	9.17	0.20	10.63	2.17	15.58	3.31	11.85	2.14	16.01	2.51	0.06	
		WAWR8	49.60	96.30	10.80	40.20	8.07	0.24	7.15	1.07	6.56	1.40	3.88	0.61	3.97	0.56	0.10	Lee et al. (2013)
	Seokmodo	SM4	48.70	81.50	8.11	27.70	4.00	0.87	4.00	0.36	1.35	0.20	0.44	0.06	0.37	0.05	0.66	Lee et al.
Granitoid	(Korea)	SM26	72.80	120.80	13.50	48.50	6.66	1.78	4.55	0.47	1.85	0.30	0.64	0.09	0.52	0.07	0.99	(2006)
	Icheon	ICH3	41.20	83.00	10.32	40.28	7.29	1.54	5.42	0.64	3.24	0.57	1.37	0.21	1.33	0.18	0.75	Lee et al.
	(Korea)	ICH11	54.50	98.70	11.10	42.10	7.34	1.12	5.75	0.71	2.92	0.51	1.29	0.16	0.93	0.13	0.53	(2008a)
-	Pohang	C2320	16.15	28.50	3.03	11.35	2.20	0.59	2.18	0.35	2.30	0.50	1.42	0.28	1.46	0.25	0.81	Lee et al.
	(Korea)	C2980	11.95	20.93	2.24	8.49	1.64	0.72	1.65	0.27	1.58	0.36	0.97	0.17	1.03	0.17	1.33	(2008b)
	Taedo	TD13	79.97	141.0	19.52	65.33	11.11	0.90	7.83	1.33	9.16	2.05	6.54	1.06	6.96	1.07	0.30	This
	(Korea)	TD2B	138.3	273.9	29.76	101.8	18.36	1.21	15.36	2.50	15.84	3.23	9.31	1.39	8.73	1.27	0.22	study

Table 4 . Rare earth element concentrations of igneous rocks from Korea and Antarctica used in this study

Table 5. Eu isotope ratio from SRM and local igneous rocks in this study

									Ext	rusive (vo	lcanic) ro	ocks									In	trusive (pl	utonic) roo	cks
Rock type	Basalt (SRM) Andesite (SRM)												Trach	nyte (Anta	arctica)			Rhyolit	e (SRM)		Diabase	Dolerite	Gabbro	o (SRM)
Sample name	BCR2 ¹⁾	JB2	JB1a	JB2	JB3	BHVO2	BIR1a	JA1	JA2	JA3	AGV2	M1710 2801-3	J1301 0105	K1601 2306	M1711 10-03	J1301 0104	JR2	JR1	RGM2	JR3	W2a	DNC1a	JGb1	JGb2
$(Eu/Eu^*)_N^{(2)}$	0.94	0.96	0.99	0.96	0.97	1.07	1.13	0.93	0.95	0.80	0.99	0.81	0.75	0.89	0.84	0.81	0.07	0.15	0.56	0.06	0.98	1.08	1.27	3.38
Eu concentration range (ng/mL) in the solution during MC-ICP-MS operation ³⁾	20~100	10~80	20~100	20~80	20~100	20~100	20~50	50~200	20~100	20~60	5~30	80	60	70	50	50	5~15	15~30	10~40	10~20	20~40	10~40	20~40	10~40
$\delta^{153/151} Eu^{4)}$	-0.08 ⁵⁾	-0.02	-0.05	-0.02	-0.05	-0.02	0.09	-0.13	-0.09	-0.08	-0.03	0.02	0.00	-0.10	-0.13	-0.26	-0.37	-0.28	-0.10	-0.50	-0.04	0.04	-0.07	-0.08
2SD	0.04 (n=12)	0.14 (n=10)	0.14 (n=8)	0.14 (n=10)	0.10 (n=14)	0.17 (n=4)	0.05 (n=7)	0.20 (n=7)	0.09 (n=7)	0.23 (n=7)	0.12 (n=3)	-	-	-	-	-	-	0.08 (n=2)	0.02 (n=3)	0.27 (n=4)	0.08 (n=5)	0.14 (n=6)	0.19 (n=7)	0.23 (n=11)
¹⁵⁴ Gd [*] / ¹⁵⁴ Sm ⁶⁾ (%, 2SD)	0.03± 0.07	0.00± 0.01	0.01± 0.02	0.00± 0.01	0.01± 0.04	0.04± 0.09	0.02± 0.03	0.00± 0.01	0.01± 0.02	0.03± 0.06	0.00± 0.00	0.09	0.08	0.06	0.03	0.03	0.02	$\begin{array}{c} 0.00 \pm \\ 0.00 \end{array}$	0.03± 0.02	$\begin{array}{c} 0.00 \pm \\ 0.00 \end{array}$	0.04± 0.07	0.01± 0.03	0.02± 0.05	0.03± 0.05
											Intrusi	ve (pluton	iic) rocks											
Rock type		1	Anorthosi	te								Granitoid	s								Granito	ids (SRM)		
Sample name	SA1703 09-1-1	SA1703 09-2	SA1703 09-3	SA1703 09-6-2	SA1703 10-3-3	MA10	MA ⁴⁾ 2018	WAWR 2018	WAWR 12	WAWR 8	SM4	SM26	ICH11	ІСН3	C2320	C2980	TD13	TD2B	STM2	JG1a	G2	JG3	GSP2	JG2
$(Eu/Eu^{*})_{N}^{2)}$	10.7	11.1	6.3	38.5	13.1	0.01	0.09	0.06	0.05	0.10	0.66	0.99	0.51	0.75	0.81	1.33	0.30	0.22	1.01	0.48	0.79	0.89	0.35	0.03
Eu concentration range (ng/mL) in the solution during MC-ICP-MS operation ³⁾	20	30~60	20~80	50~200	60	2~5	10	10	20	15~40	45	60	30	40	35	40	20	0	40~150	10~20	10~80	10~40	25~70	5~40
$\delta^{153/151} Eu^{4)}$	0.12	0.12	0.17	0.18	0.24	-0.28 ⁵⁾	-0.11	-0.07	-0.17	-0.16	0.05	0.00	-0.03	-0.07	-0.02	-0.02	-0.07	-0.18	0.00	-0.03	-0.10	-0.01	-0.07	-0.31
2SD	-	0.01 (n=2)	0.05 (n=3)	0.03 (n=3)	-	0.40 (n=2)	-	-	-	0.31 (n=4)	-	-	-	-	-	-	-	-	0.06 (n=5)	0.06 (n=13)	0.19 (n=4)	0.04 (n=10)	0.05 (n=17)	0.30 (n=6)
¹⁵⁴ Gd ^{*/154} Sm ⁶⁾ (%, 2SD)	0.00	0.00± 0.00	0.00± 0.00	0.00± 0.00	0.00	0.01± 0.04.	0.00	0.00	0.10	0.05± 0.06	0.06	0.01	0.01	0.01	0.04	0.02	0.02	0.00	0.00± 0.00	0.02± 0.05	0.01± 0.03	0.01± 0.04	0.04± 0.07	0.00± 0.00

1) The samples of Bold character are SRMs of USGS and GSJ.

2) The magnitude of Eu anomaly is defined as s the ratio $(Eu/Eu^*)_N$ where Eu* is SQRT $(Sm_N \times Gd_N)$.

3) approximately range value (ng/mL) of Eu concentration in the sample solution diluted by 2% HNO₃ for MC-ICP-MS determination

4) Eu isotope ratio normalized by ¹⁴⁷Sm-¹⁴⁹Sm isotope pair (Lee and Tanaka, 2021a, 2021b).

5) Eu isotope data of bold Italic numbers re-calculated with data of Lee and Tanaka (2021a)

6) degree of ¹⁵⁴Gd intensity effect for ¹⁵⁴Sm (%): ¹⁵⁴Gd^{*} (calculated intensity) = measured intensity of ¹⁵⁵Gd × 2.18 % (abundance of ¹⁵⁴Gd)/14.8 % (abundance of ¹⁵⁵Gd) (Rosman and Taylor, 1998)

Supplementary Material for

Geochemical implication of Eu isotopic ratio in anorthosite: new evidence of Eu isotope fractionation during feldspar crystallization

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Samples

This study was initiated under the assumption that Eu isotope fractionation is associated with Eu anomaly during the evolution of the igneous rocks. Therefore, when selecting samples, except for anorthosite, silica concentration of igneous rocks and variation of the magnitude of their Eu anomaly were considered.

Except 25 geochemical reference materials purchased from the United States Geological Survey (USGS) and the Geological Survey of Japan (GSJ), anorthosites and granitoids from Korea and trachytes from Antarctica were collected for this study. The 25 geochemical reference materials in this study were as follows; Seven basalts (BCR2, BHVO2 and BIR1a purchased from the USGS; JB1a, JB1b, JB2 and JB3 from the GSJ), four andesites (AGV2 from USGS; JA1, JA2 and JA3 from GSJ), four rhyolites (RGM2 from USGS; JR1, JR2 and JR3 from GSJ), one diabase (W-2a from USGS), one dolerite (DNC1a from USGS), two gabbros (JGb1, JGb2 from GSJ), one syenite (STM2 from USGS), and five granites (G2 and GSP2 from USGS; JG1a, JG2 and JG3 from GSJ).

Twenty four local rock samples from Korea and Antarctica are as follows; fourteen granites and five anorthosites from Korea, and five trachytes from Antarctica. The trachytes were selected because I did not obtain SRMs.

1. Anorthosites (Sancheong-Hadong, SA)

The Sancheong-Hadong anorthosites occur in the southern Yeongnam Massif, South Korea. The anorthosites are known to be differentiated from mantle-derived parental magma during the Paleoproterozoic period (~1875 – 1863Ma) (Kwon and Jeong, 1990; Lee et al., 2014, 2017). This period is correspond to emplacement age of Proterozoic massif-type anorthosites in the world (Ashwal, 2017). The majority of primary plagioclase (An56-68) in the anorthosite was replaced by aggregates of smaller recrystallized grains (An46-54) (Lee et al., 2014). In order to measure Eu isotope ratio and REE abundances from anorthosites, we collected twenty seven rock samples from the outcrops in the Sancheong-Hadong area.

2. Muamsa (MA)-Weolaksan (WA) granites

The Muamsa-Weolaksan granites occur in the Hwanggangri district located in the NE margin of the metamorphic area of the Okcheon Belt, South Korea, which are highly fractionated Cretaceous granites with A-type geochemical characteristics. The Weolaksan granite is a batholith type whereas the Muamsa granite is narrow stock type. The granites are medium- to coarse-grained biotite granites, and were evolved from the granitic magma derived from the same source material (Lee et al., 2010, 2013, 2018).

We collected five granite samples from the outcrops in the Muamsa-Weolaksan area for determination of Eu isotope ratio and REE abundance.

3. Pohang (C) granites

The Pohang granites were collected from the core samples in the Pohang area covered with Tertiary sedimentary rocks and Cretaceous volcanic at southeastern part in the Gyeongsang Basin, South Korea. The Pohang granites were emplaced at Cretaceous age and show geochemical features that they were derived from mantle source material (Lee et al., 2008b).

We collected two granitic core samples for determination of Eu isotope ratio and REE abundance.

4. Seokmodo (SM) granite

Seokmodo is a small island located on the west coast of Korea and consists of the Jurassic biotite granitoids. Seokmodo is also known as the border between South Korea and North Korea (Lee et al., 2006).

We collected two granitoids for Eu isotope and REE abundance measurement.

5. Icheon (ICH) granite

The Icheon Jurassic granite occurs in the area which is located in the central Korean Peninsula. The northeastern part of these granite body consists of the Precambrian biotite gneiss. We selected two granite sample powders which was reported by Lee et al. (2008a) for Eu isotope and REE abundance measurement.

6. Hataedo (TD) granite

Hataedo is a small island, which is located in the southwestern sea (Yellow Sea) near to the Korean Peninsula. Hataedo granite occurs in the Hatedo and is Proterozoic alkali feldspar granite (Lee et al., 2021). We collected two granites by Lee et al. (2021) for Eu isotope and REE abundance measurement.

7. Antarctic volcanic rocks

We analyzed volcanic rocks from the Mt. Melbourne volcanic field (MMVF) in the Melbourne Province, Antarctica. They represent one of the most extensive volcanic provinces in the world, comparable to the alkali volcanic rocks in the East African rift systems. The McMurdo Volcanic Group is divided into three provinces: Hallett, Melbourne, and Erebus. The Melbourne province includes the three large stratovolcanoes of Mt. Overlord, The Pleiades and Mt. Melbourne which are composed of a wide range of intermediate and felsic alkali differentiates evolved from alkali basaltic rocks. The Mt. Melbourne (74.35° S; 164.70° E) is a 2732m high alkaline stratovolcano composed of scoria cones, domes, lava flows and various pyroclastic deposits (Wörner and Viereck, 1989; Giordano et al., 2012).

In this study, we collected five trachytes from Mt. Melbourne, and analysed their Eu isotope ratios and REE abundances.

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