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1 **Sensitivity, Accuracy and Limits of the Lightweight**
2 **Three-Component SmartSolo Geophone Sensor (5 Hz)**
3 **for Seismological Applications**

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6 **Key Points:**

- 7 • instrument tests
8 • nodal systems
9 • SmartSolo
10 • sensor performance

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11 **Abstract**

12 The use of Nodal systems based on autonomous geophone-based instruments entered the
13 field of Seismology only recently. These lightweight solutions revolutionized seismolog-
14 ical fieldwork through lightweight and wholistic instruments that are faster to deploy and
15 easier to handle. The IGU-16HR series of SmartSolo® is one example, but yet lacking
16 a thorough lab-based performance analysis. Here, we fill the knowledge gap, by perform-
17 ing a series of lab and field-based tests that focus on the sensors performance. The in-
18 vestigated parameters are the instruments transfer function, self-noise and overall per-
19 formance to classical seismometer-based instruments. In the real-world application we
20 show examples of H/V measurements of ambient vibrations in urban environments and
21 the performance ranges with teleseismic waveform recordings. Under lab conditions, the
22 nodal systems perform equally well as standard seismometers (e.g., Lennartz 3D/5s), even
23 in the frequency range down to 0.2Hz, way below their natural frequency. The restitu-
24 tion can be carried out correctly with manufacturer given transfer function. At least for
25 the vertical component, the instruments self-noise reaches the lower boundary of the global
26 minimum noise level, confirming the ability to properly record teleseismic phases down
27 to 0.1 Hz. In ambient noise studies the instrument limits are already reached at 0.8 Hz,
28 but still resolve the fundamental frequencies within the methods uncertainty ranges, based
29 on classical instrument data. These versatile and easy-to-use nodal systems are useful
30 and reliable for a wide range of seismological applications. In addition, their installation
31 is faster and reduced prices open the doors towards Large N installations and research
32 studies for groups that face limited financial budgets.

33 **Introduction**

34 Recent developments in seismological research have seen tremendous increases in
35 sheer size of data throughout the last decade (Quinteros et al., 2021; Arrowsmith et al.,
36 2022). This evolution has been accompanied by increasing computational power enabling
37 the processing of such large data-sets (Ahrens et al., 2011; Bozdog et al., 2014; MacCarthy
38 et al., 2020) and the introduction of Machine Learning techniques for seismological data
39 processing (Bergen et al., 2019; Kong et al., 2019; Arrowsmith et al., 2022). On the hard-
40 ware side, the introduction of low-cost geophone sensors (e.g., Raspberry Shake) often
41 in combination with wholistic software/hardware solutions enabled data recording in un-
42 precedented quantity of stations and for non-scientific audiences, for which the term "cit-

43 izen science” has been introduced (Chen et al., 2020; Subedi et al., 2020; De Plaen et
44 al., 2021; Calais et al., 2022). While the use of such low-cost instrument is limited and
45 cannot cover the full range of seismological methods (Anthony et al., 2019), integrated
46 nodal systems bear the potential to present a cost-efficient compromise of the performance
47 in between citizen instruments and classical seismological sensors.

48 Nodal systems are common practice in active seismic experiments for exploration
49 of hydrocarbon and other resources (Dean et al., 2018), in which numerous geophones
50 (mostly single component instruments) are regularly spaced over a site of interest record-
51 ing subsurface reflections of actively induced signals (e.g., explosive or sweep). Besides
52 extending to three-component instruments, latest developments in geophone sensors for
53 nodal installations saw major efforts in enhancing the level of autarky. To overcome is-
54 sues of power supply, communication and time accuracy in remote locations, integrated
55 nodal systems eliminate cable-based solutions and incorporate digitizer, data storage,
56 GPS and battery in a single acquisition unit (Dean & Sweeney, 2019). The first com-
57 mercially available node that also enabled continuous data recording was the Fairfield
58 ZLand node (A. T. Ringler et al., 2018). This instrument is also eligible to be used for
59 seismological research questions. With a fraction of the purchasing costs compared to
60 standard seismological acquisition systems, the installation of so-called Large N arrays
61 with 100s to 1000s of nodes became possible (Hand, 2014; Karplus & Schmandt, 2018;
62 Roux et al., 2018; Brenguier et al., 2015). One of the first installations of such kind was
63 realized in the Los Angeles basin with $\sim 13,000$ seismic stations covering an area of 16
64 x 16 km with three separated arrays and equidistant sensor spacing of 100m that enabled
65 unprecedented spatial sampling of wavefield and site-characteristics (Castellanos & Clay-
66 ton, 2021).

67 SmartSolo® recently released their IGU-16 series instruments. These geophone
68 instruments with a 5 Hz natural frequency are available as single (IGU-16 1C) or three-
69 component (IGU16-HR 3C) sensors and are equipped with 24 bits digitizers and GPS.
70 Batteries are modular and available as High Capacity Battery or Standard Capacity Bat-
71 tery Packs which, together with the sensor, eventually provides a single, closed casing
72 sensor. The total weight of the 3C (2.4 kg high capacity, 1.7 kg standard capacity bat-
73 tery) and size (10.3 x 9.5 x 18.7 cm) outperforms classical seismometer-digitizer set-ups.
74 During the installation of larger surveys, the operator profits from the reduced man-power
75 and time necessary. Due to the modular design of the nodes that allows the replacement

76 of their spike base with a tripod battery base, these sensors' potential use becomes in-
 77 dependent from the available surface structure in the survey area, i.e., urban environ-
 78 ments with a high degree of sealing. In the last years, the SmartSolo node series have
 79 been increasingly used for Large N installations in the field of passive seismology (e.g.
 80 Obermann et al., 2022; Chmiel et al., 2019).

81 So far, a comprehensive study identifying the capabilities of geophone-based node
 82 sensors for seismological purposes has only been performed for the Fairfield ZLand sen-
 83 sors (A. T. Ringler et al., 2018), but is yet unavailable for the SmartSolo sensors. In this
 84 study we evaluate the SmartSolo instruments characteristics, performance and limits in
 85 order to justify their use in a variety of seismological applications. In a set of lab-based
 86 experiments we identify the sensors' transfer function, control the manufacturer's given
 87 poles and zeros, check the self-noise level, and compare the sensors with well-calibrated
 88 seismometers. After that, we show the performance of the sensors during field installa-
 89 tions with two examples focusing on teleseismic waveforms and ambient seismic noise
 90 measurements.

91 **Instrument tests**

92 **Instrument response derived from coherent waveforms**

93 In the recording of ground shaking, a seismic sensor acts as a filter in the sense of
 94 a linear, time-invariant system (LTI) (Scherbaum, 2006) when translating it into elec-
 95 tric voltages as an output signal. This alternation from input to output signal is repre-
 96 sented through the system's frequency response function or the transfer function. The
 97 quantitative description of the LTI then allows us to restore the original input signal by
 98 applying signal restitution to the obtained waveforms without further knowledge of the
 99 physical processes going on inside the filter (Scherbaum, 2006). The transfer function
 100 is then characterized by the complex poles and zeros.

101 Havskov & Alguacil (2015) have shown that it is possible to estimate the transfer
 102 function of a seismometer by using the natural vibrations of the ground as a shaking ta-
 103 ble recorded with two closely installed sensors. For the SmartSolo sensors, the output
 104 signal is expected to be contaminated by instrument noise and thus, we applied the cross-
 105 spectrum method (Eq. 1). In this method, the output of seismometer 1 is the input of
 106 seismometer 2 as a linear system that presents a transfer function in the form of:

$$T_2(\omega) = T_1(\omega) \frac{P_{21}(\omega)}{P_{11}(\omega)} \quad (1)$$

107 with P_{21} as the cross-spectrum between the outputs of both sensors and P_{11} as the au-
 108 topower spectrum of the output of sensor 1. Under the assumption the instrument re-
 109 sponse (as poles and zeros) given by the manufacturer is correct for sensor 1, we can es-
 110 timate the unknown response of sensor 2. This is repeated for all instrument pairs.

111 The estimation of the instrument response parameters represented by its poles and
 112 zeros is a non-linear operation. Therefore, T_2 is identified through the optimization of
 113 the poles and zeros and fitting the theoretical response function to the observed trans-
 114 fer function presented in equation 1. The misfit function of the optimization is repre-
 115 sented by the complex L2-norm.

116 In order to obtain highly correlated ground motions, 24 3C nodes have been closely
 117 co-located (in a so called ‘huddle’, Fig. 1c) in a regular grid of 1m x 1m overall exten-
 118 sion close to the Uccle permanent station of the Belgian seismic network (international
 119 code BE.UCC, Royal Observatory of Belgium, 1985). The location within Brussels as-
 120 sured a high noise level. During the recording period a teleseismic earthquake could be
 121 recorded (M7.3, Japan, GEOFON Data Centre, 1993) that further guarantees strong cor-
 122 relation of the obtained wavefield.

123 The resulting poles for the instrument response estimation strongly converge to-
 124 wards the values given by the manufacturer ($-22.2111-22.2178i$, $-22.2111+22.2178i$),
 125 with half of the estimated transfer functions obtaining misfits below 5%. The weighted
 126 mean for poles below this misfit threshold differs by $-0.0559 \mp 0.0552i$ from the man-
 127 ufacturer given values. Considering only the results with misfits below 2%, the poles dif-
 128 fer by $-0.0162 \mp 0.0158i$. Stronger misfit of the resulting transfer functions are foremost
 129 proportional to intersensor distances as the higher frequency sections of the recorded noise
 130 spectra de-correlate with increasing distance. This result could be reproduced for the
 131 horizontal components as well, with an overall greater spread of high misfit poles and
 132 zeros. This is likely due to the higher self-noise of the horizontal components (section
 133 2.2) that leads to less coherent waveforms, as they show lower signal-to-noise ratio of the
 134 teleseismic phases and are more affected by tilt of the sensor that reduces the overall sen-
 135 sitivity. However, the limitation to results with misfits below 5% or 2% leads to the same

136 range of differences between manufacturer and estimated poles and zeros as determined
 137 for the vertical component.

poles	$-22.2111 - 22.2178i$	$-22.2111 + 22.2178i$
zeros	$0i$	$0i$
sensitivity (@0 gain)	$76.7e3 \frac{mV}{\frac{m}{s}}$	
digitizer gain	3355.4428	

Table 1. Instrument Response for a SmartSolo IGU-16HR-3C node represented by Poles and Zeros.

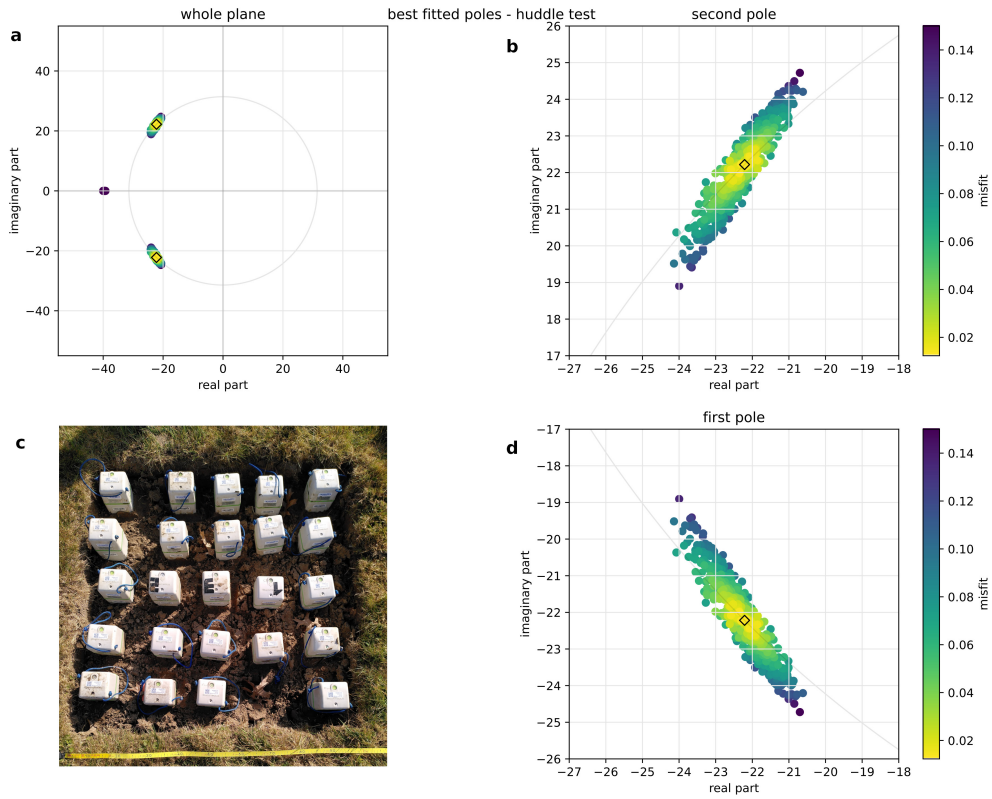


Figure 1. The resulting poles of the SmartSolo nodes huddle test, color-coded by misfit from the manufacturer’s values, shown by the diamond marker. a) result shown over the whole complex plane that has been defined as the solution space in the inversion. b) and d) close-up view of the two poles. c) 24 nodes co-located during the huddle test. Note the slightly imperfect installation, contributing negatively to the misfit values.

138 **Instrument self-noise and long-term noise stability**

139 The experiment set-up presented in figure 1c of 24 co-located SmartSolo 3C instru-
 140 ments allowed us to apply the three instrument approach of Sleeman et al. (2006) to iden-
 141 tify the instrument’s self noise based on a common, coherent input data. Here, we rely
 142 on the analysis of actual ground motion recordings during the self-noise test. As the sen-
 143 sor and digitizer are located within the same casing, we cannot measure their self-noise
 144 independently and the full recording system combining both sensor and digitizer is an-
 145 alyzed. For the most part, the self noise of digitizers lies up to 20dB below the self noise
 146 of the sensors A. T. Ringler et al. (2014) and thus we assume that the obtained noise
 147 spectra will reflect only the sensor’s self noise.

148 The comparison was performed for each instrument (i) using the two closest neigh-
 149 boring sensors of the grid (j, k). Similar to the huddle test, the use of the cross-spectrum
 150 (P_{ji}, P_{ik} , etc.) between the sensors eliminates the sensor’s transfer functions and noise
 151 cross-spectra. The systems self-noise autospectrum (N_{ii}) then can be expressed solely
 152 through power- and cross-spectra of the obtained output of the three sensors (i, j, k)
 153 under the assumption of a common recording input as follows:

$$N_{ii} = P_{ii} - P_{ji} \cdot \frac{P_{ik}}{P_{jk}} \quad (2)$$

154 In order to retain comparability of the experiment outcome of A. T. Ringler et al.
 155 (2018) in which the authors performed a lab test for the Fairfield nodes on a shaking ta-
 156 ble and comparison with broadband sensors, we apply the same Fourier transformation
 157 parameters, prior downsampling (decimate from 250 to 50 Hz), and moving average to
 158 smooth the resulting spectra. The input data is a 1-hr period at a Thursday night (2022-
 159 03-17 01:30:00 UTC) in order to minimize the environmental noise close to the BE.UCCS
 160 station (lat 50.797, lon 4.36) in an open field as the spike at the bases could not be re-
 161 moved. In order to reduce errors propagating from transfer functions uncertainties, the
 162 input waveforms have been restituted (A. Ringler et al., 2011).

163 *This following paragraph has been updated in version 3 and was faulty in the pre-*
 164 *vious versions.*

165 The resulting self-noise increases linearly between 0.7 to 15 Hz from around -140
 166 dB to -130 dB (figure 2), which lies around the usual background noise of silent hard rock

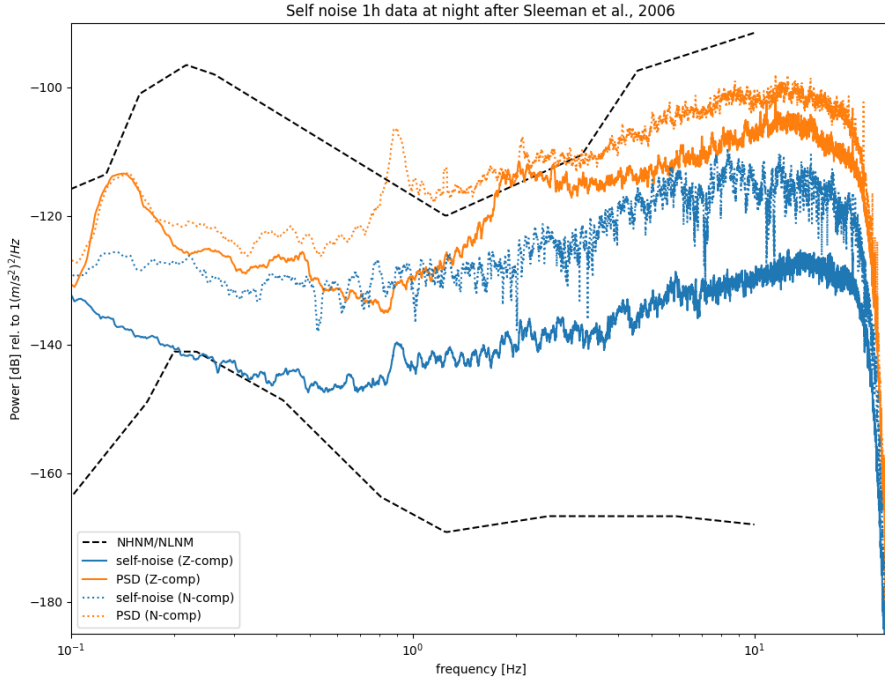


Figure 2. Outcome of the self-noise test following Sleeman et al. (2006). Blue curves is the estimated system self noise for a SmartSolo 3C sensor. Orange curves show the power spectrum of the recorded ground motion. The solid lines correspond to the vertical component and dashed colored lines show the equivalent for the North component. The black, dashed lines give the upper and lower global noise model bounds (Peterson, 1993).

167 stations in the Belgian network. For lower frequencies the self-noise is steadily increas-
 168 ing but remains around the NLNM until 0.2 Hz. For higher frequencies the self-noise is
 169 decreasing, due to the taper applied before the downsampling. The overall shape of the
 170 noise spectrum is comparable to the Fairfield nodes (A. T. Ringler et al., 2018), that was
 171 installed in the ASL underground vault and thus likely shows the true Fairfield self-noise.
 172 Above 1 Hz the SmartSolo node shows slightly higher noise-levels which corresponds to
 173 the nigher theoretical noise floor (3 - 16 dB). While for lower frequencies the SmartSolo
 174 node shows consistently lower self-noise that even reaches the NLNM levels (Peterson,
 175 1993) and outperforms the Fairfield instrument. This observation underlines a decent
 176 sensitivity of the SmartSolo nodes for a seismological purpose as a passive sensor for tem-
 177 poral installations for recording ambient seismic noise, detecting local tectonic and/or

178 induced earthquakes, investigating activity in geothermal fields, and they remain per-
 179 formant over a large frequency range from 5 s to their chosen Nyquist frequency (here
 180 25 Hz). The increase in noise level towards the higher periods probably makes them less
 181 suitable to investigate teleseisms, microseism and storms.

182 The horizontal components show on average a 15 dB higher noise level that are more
 183 sensitive to signal distortion due to tilt. The cross- sensor comparison of all 24 installed
 184 instruments in the Huddle test experiment shows no major distortions or anomalies for
 185 individual instruments. Only a few nodes show some irregular higher variability of the
 186 self-noise spectra. We suggest this might be related to the different level of coupling of
 187 each sensor that further introduces incoherencies in the recorded wavefields in the am-
 188 bient noise frequency range. Such incoherence of the input of the three sensors is then
 189 further propagated into the noise spectrum analysis. In future analyses of the Smart-
 190 Solo sensors this could be avoided by using a shaking table instead of relying on coher-
 191 ent waveform recordings. The three first generation instruments (indicated in figure 1c
 192 with the letters 1, 2, 3) installed alongside the newest generations show the same out-
 193 come.

194 Comparison with well-calibrated seismometers

	Güralp DM24 + 3ESP	CityShark II + Lennartz 3D	SmartSolo
natural/corner frequency	30 s	5s	5 Hz
sampling frequency	100 Hz	250 Hz	250 Hz
downsampled frequency	50 Hz	50 Hz	50 Hz

Table 2. Overview of seismometer and digitizer combinations with sampling specifications.

195 In a lab-based instrument test, the SmartSolo sensors were compared with well-
 196 calibrated, standard seismometers. The SmartSolo nodes were co-located with (*i*) the
 197 surface sensor of the Uccle station (network station code: BE.UCCS), that consists of
 198 a Güralp DM24/3ESP instrument, and (*ii*) with a Lennartz 3D sensor connected to a
 199 CityShark digitizer (Chatelain et al., 2000) for comparison to a standard instrument that
 200 is used for ambient seismic noise measurements (figure 3). During this experiment, we
 201 also investigated if the modular use of different base set-ups of the Smartsolo nodes al-

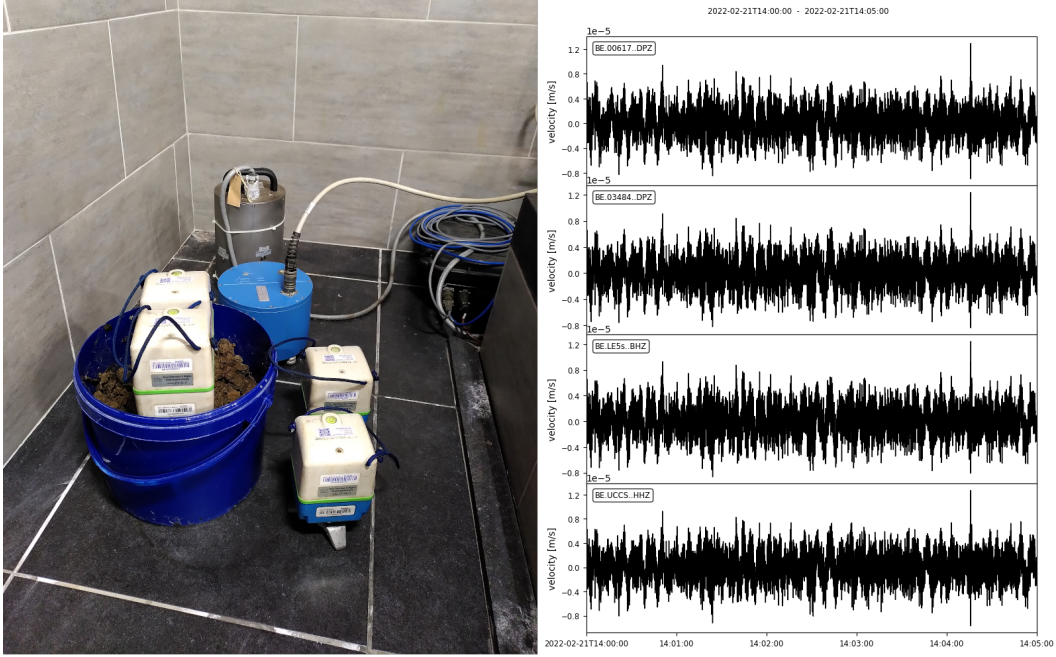


Figure 3. Co-location test of Smartsolo sensors with well-calibrated seismometers. Left: Two nodes each with different base set-ups either on a tripod or with a central spike in a sand-filled bucket, Lennartz LE3D/5s (blue instrument) connect to a Cityshark and Uccle surface station BE.UCCS (gray instrument in the back, Güralp CMG 3ESP). Right: restituted waveforms of all four kinds of sensors: From top to bottom: 1) node in bucket, 2) node on tripod, 3) LE3D/5s, 4) permanent sensor.

202 ters the recorded noise field. Two different set-ups were tested: (*i*) a 3C node connected
 203 to the High Capacity Battery Pack (gray) on a central spike installed in a sand-filled bucket
 204 and (*ii*) a 3C node connected to the Standard Battery Pack (blue) on a steel tripod base
 205 installed on the floor of the cave next to the listed seismometers above (figure 3).

206 To compare the obtained waveforms in the time and spectral domain, we first re-
 207 moved the instrument responses of all sensors (table 1). The restituted waveforms of all
 208 four sensor types are highly congruent in obtained ground velocity amplitudes and time
 209 accuracy. This congruence demonstrates the accuracy of the nodes' poles and zeros iden-
 210 tified during the huddle test (see section above). In order to quantify the waveform sim-
 211 ilarity, we computed the coherence of all instrument combinations as the normalized cross-
 212 spectra (figure 4).

213 In comparison to the well-calibrated Güralp instrument, the node sensor installed
 214 on a tripod has the highest overall coherence with nearly perfect similarity from 20 Hz
 215 down to 10 s, way below its natural frequency (figure 4). A small deviation is present
 216 between 0.85 and 1.05 Hz that is more evident for the nodes with a central spike, but
 217 the waveform similarity always exceeds a 0.9 coherence. The decreasing coherence above
 218 20 Hz for the SmartSolo sensors appears to be a filter artifact that propagates from the
 219 different decimation applied to the waveforms to result in a common sampling frequency
 220 (table 2). For the CityShark with Lennartz instrument the ~ 1 Hz coherence drop is
 221 absent but above 4 Hz the waveform similarity to all other sensors in this test is steadily
 222 decreasing and falls below 0.9 at around 13 Hz. Due to the absence of a lowpass filter
 223 close to the Nyquist frequency of the raw data, we presume the existence of an analog
 224 filter in the CityShark digitizer with a cut-off that starts around 13 Hz and is not included
 225 in the instrument’s transfer function.

226 The instrument comparison in the spectral domain is visualized in figure 4 and was
 227 obtained by dividing the power spectra of all instruments with all other instruments co-
 228 located during the experiment. Similar to the waveform similarity, we obtain flat spec-
 229 tral divisions at the ratio of 1 from $\lesssim 10$ Hz down to less than 10s. Here, the similarity
 230 deviation around ~ 1 Hz of the SmartSolo sensors becomes evident again and also is
 231 more pronounced for the sensors with a central spike installed in a sand-filled bucket.
 232 However, this effect can only be observed for the vertical components and is absent for
 233 the horizontal components. In the low frequency range below 0.2 Hz (remind that the
 234 node’s natural frequency is 5 Hz) the horizontal spectra of the SmartSolo sensors devi-
 235 ate stronger from the spectrum obtained with the well-calibrated instrument as it can
 236 be observed for the vertical components. The deviation from the well-calibrated instru-
 237 ment is even larger for sensors that were installed with a spike in the sand-filled bucket
 238 and thus, results from the fact that the nodes were not fully buried and resulting in poorer
 239 leveling in comparison to the tripod based nodes.

240 **Real-world observations (applications, sensitivity)**

241 **Teleseismic arrivals**

242 In the previous chapter, it was shown that the waveforms obtained with the Smart-
 243 Solo nodes bear the potential to recover ground motion far below their own natural fre-

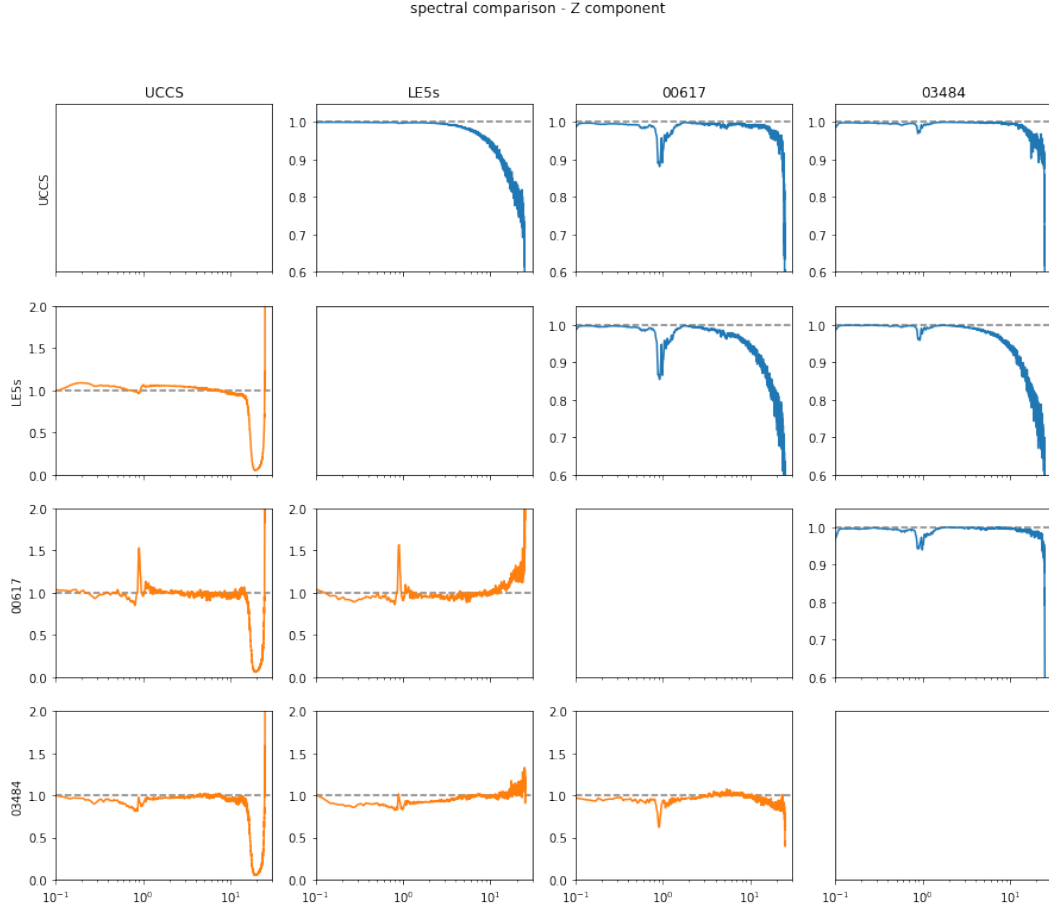


Figure 4. Waveform similarity between different types of sensors. Upper part (blue curves) shows the coherence amplitudes between the different sensors. Lower part (orange curves) gives the spectral divisions of all sensor combinations.

244 quency. During two longer term SmartSolo array installations in 2020 and 2022 around
 245 the BE.UCCS station, two teleseismic events in Kermadec (Mw 7.4, June 18, 2020) and
 246 Japan (Mw 7.3, March 16, 2022) respectively occurred during the surveys. To compare
 247 the node’s performance with BE.UCCS, waveforms were first restituted to velocity and
 248 then bandpass filtered between 20 s and 3 s (figure 5).

249 The waveforms of the vertical component of a single Smartsolo sensor perfectly match
 250 the waveforms obtained with a Güralp instrument, with only slightly higher amplitudes
 251 for BE.UCCS. For both waveforms, the first arrivals of the body wave phases could be
 252 identified on a single vertical component for both events (PKIKP for Kermadec at 162°
 253 distance and PP for Japan at 84° distance). The surface waves of the 2022 Mw

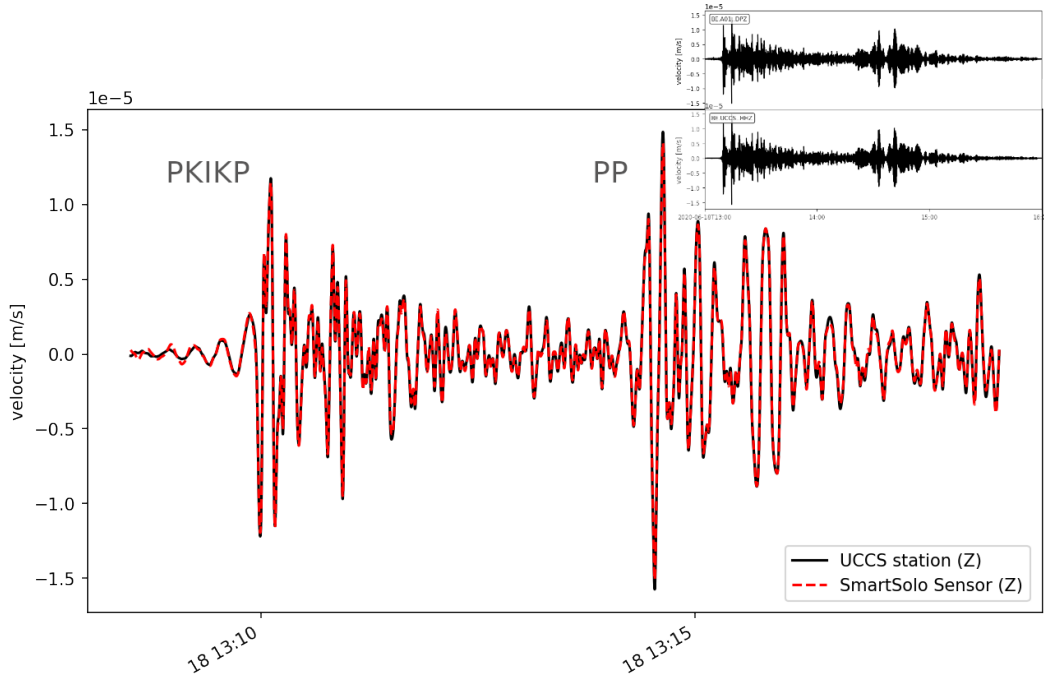


Figure 5. Teleseismic PKIKP and PP phases of the Mw 7.4 Kermadec earthquake (June 18, 2020) recorded with the vertical components at the BE.UCCS station (black) and with a single Smartsolo sensor (red). Both waveforms have been restituted to velocity and filtered between 20 s and 3 s. The inlet shows the full length (3 hours) of the teleseismic earthquake recorded with the SmartSolo Node (top trace) and the Güralp sensor (lower trace).

254 7.3 Japan earthquake could only be retrieved when lowering the bandpass filter down
 255 to 100 s, due to their lower dominant frequencies.

256 The horizontal components have a much reduced sensitivity in the very long pe-
 257 riod range. Thus for the Mw 7.4 Kermadec event, the earthquake can only be identified
 258 by stacking the waveforms of at least 20 nodes. It is important to mention that the sta-
 259 tion BE.UCCS around which the tests were performed is located in the city of Brussels
 260 and possesses one of the highest seismic noise levels in the whole BE network (Lecocq
 261 et al., 2020). In contrast, for the Mw 7.3 Japan earthquake, the first arrival S-phases and
 262 the surface waves can already be identified on a single horizontal sensor, but at the same
 263 cost as described before for the vertical component.

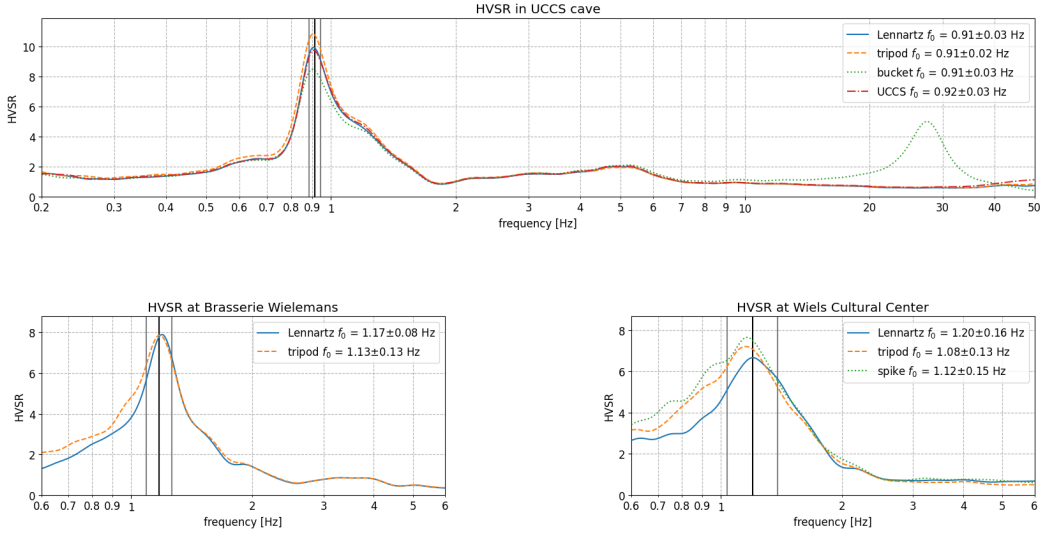


Figure 6. Instrument comparison through H/V analysis at three locations in Brussels. Waveforms have been restituted before the processing. HVSr graph from recordings a) at the location of UCC surface sensor (50.7973N, 4.3605E) from the sensor comparison lab test (fig. 3), with the blue solid line for LE3D-5s with Cityshark, orange dashed line for SmartSolo node on tripod base, green dotted line for node with spike in a sandfilled bucket and red dot-dashed line for Guralp permanent sensor. b) Location of the former Wielemans Brewery (50.8261N, 4.32646E) with LE3D-5s and Smartsolo sensors ~ 10 cm apart. c) Location at the Wiels Cultural Center (50.82453N, 4.3259E) with LE3D-5s, node on tripod and node with spike digged into a grass field. Intersensor distance 5 - 10 m.

264 Ambient noise application

265 The preceding lab-based tests infer a suitable frequency range that justifies to use
 266 the SmartSolo sensors for passive measurements of ambient seismic noise in the frequency
 267 range of 0.2 to 25 Hz. As an example of an ambient noise application, we show three ex-
 268 amples of Horizontal-to-Vertical Spectral Ratio (HVSr) (Nakamura, 1989; Molnar et al.,
 269 2022) surveys in Brussels, Belgium. In the framework of a shallow geothermal feasibil-
 270 ity study, we prospected several sites in Brussels with non-invasive ambient noise obser-
 271 vations prior to drilling. For the region of Brussels (Belgium), a conversion law exists
 272 to estimate the depth to bedrock from fundamental resonance frequency (f_0) values, de-
 273 rived from Horizontal-to-Vertical Spectral Ratio (HVSr) analysis of ambient noise mea-
 274 surements co-located with well logs (Van Noten et al., 2022). In the Brussels capital re-

275 gion, the main acoustic impedance contrast corresponds to a fundamental frequency range
 276 between 0.6 and 1.6 Hz (Van Noten et al., 2022).

277 At first, the data of the ideal case-study of the instrument comparison test (fig. 3)
 278 have been analyzed using Geopsy (Wathelet et al., 2020). The location of the perma-
 279 nent station (BE.UCCS) within the cave of the Royal Observatory of Belgium (ROB)
 280 reassures constant environmental conditions (e.g., stable temperature, no insulation). The
 281 circular street present around the ROB also provides sufficient distance and azimuthal
 282 coverage of anthropogenic noise sources. In this case study, the same time windows (120s
 283 long) have been used for all co-located instruments and the horizontal components have
 284 been averaged when computing the HVSR spectra. The HVSR spectra for all instruments
 285 (fig. 6a) are congruent for the most parts. The fundamental frequency f_0 can be repro-
 286 duced by all sensors within 50% of the given uncertainty range (given as one standard
 287 deviation in Geopsy). For the HVSR amplitude at the f_0 , the Lennartz seismometer (con-
 288 nected to the City-shark) is comparable to the UCCS, Gralp permanent sensor. The
 289 node on the tripod gives a 10% higher and the node in the bucket a 15% lower ampli-
 290 tude value. For frequencies above 13 Hz, the HVSR curve computed from the nodal in-
 291 strument in the bucket deviates strongly and contains a second peak at 27.5 Hz. We pre-
 292 sume this peak is related either to a bad coupling of the spike base in the sand-filled bucket,
 293 or to an impedance contrast between the bucket and the tiling floor with its concrete base.
 294 Considering the lower HVSR amplitude at f_0 , the former is more likely as tilting shows
 295 stronger negative effects of the horizontal than the vertical components according to the
 296 manufacturer.

297 Under real-world conditions we present two examples from the southern part of Brus-
 298 sels. At the first location (fig. 6b), located next to the former *Brasserie Wielemans*, the
 299 Lennartz 5s with Cityshark and a SmartSolo node with a tripod base have been placed
 300 on a sidewalk, 10 cm apart. The street presents high traffic amounts, including public
 301 busses, streetcars and pedestrians passing next to the sensors. The second location, lo-
 302 cated 200m away from the first case (*Wiels Cultural Center*, fig. 6c) consists of a park-
 303 ing spot next to the same street and a community garden. That allowed us to install the
 304 Lennartz and tripod node next to a node with a spike for comparison. Distances between
 305 the sensors lie between 5 and 10 m. The f_0 values obtained at both locations are the same
 306 on average and given the uncertainties. This is expected as both locations show no el-
 307 evation difference and have a similar geologic subsurface structure located in the Senne-

308 valley. Above 1.3 Hz the HVSR curves for each location are congruent. The shape of the
 309 HV peak for the *Brasserie Wielemans* is sharper and presents smaller uncertainties. The
 310 *Wiels* location demonstrates larger variations in f_0 as well as the corresponding ampli-
 311 tude.

312 The HVSR curves deviate strongly for frequencies below their f_0 peaks. The in-
 313 spection of all analyzed time windows for the spectral analysis in Geopsy reveals much
 314 stronger fluctuation of the individual HVSR curves below 0.8 Hz. This leads to the wider
 315 average HVSR peak around f_0 in this frequency range. Below the S-wave resonance fre-
 316 quency the wavefield is dominated by body-waves and nearby sources (Lunedei & Malis-
 317 chewsky, 2015). In a densed urban area, as shown in the two examples, it might be im-
 318 possible to decouple the instrument from the noise generating infrastructure (i.e., side-
 319 walk of a heavily used street). In addition, noise-receiver distances are short and noise
 320 sources are non-stationary. Here, we recommend careful selection of investigation sites
 321 with longer recording periods and multiple locations. The impact of the noise instabil-
 322 ity below 0.8 Hz could be reduced through restitution of the raw waveform data. This
 323 step limits the amplitude differences of the mean HVSR curves as presented in figure 6.

324 The main distinction between the SmartSolo nodes and the classical seismometer-
 325 digitizer set-up became obvious in the handling of the hardware during the survey. The
 326 integrated node sensors outperform classical instruments in size, weight and usability.
 327 In the same time needed for one trained surveyor to install a seismometer-digitizer set-
 328 up, a single surveyor can transport and install up to 4 nodes. The use of multiple instru-
 329 ments might introduce some redundancy, but allows to capture potential lateral varia-
 330 tions over short distances. Especially in urban contexts, additional sensors assure suc-
 331 cessful data recordings in cases of unwanted noise sources (e.g., traffic, pumps, etc.), un-
 332 known subsurface cavities (e.g., channels, sewers) or bad coupling (see above).

333 Conclusions

334 With three different “lab-based” tests using coherent ground motion recordings,
 335 we demonstrated the high performance of the 3C SmartSolo sensors (IGU-16HR-3C).
 336 The manufacturer given values for the transfer function could be reproduced in the so-
 337 called huddle-test and were used to accurately reconstitute the instrument responses. Their
 338 overall self-noise resides around the global minimum noise level (Peterson, 1993) over a

339 wide frequency range, through which they become versatile and useful for a wide range
340 of seismological applications, such as seismotectonics in local and regional distances, noise
341 tomography, ambient noise studies and applied geophysics. In direct comparison to stan-
342 dard instruments in use for decades for seismological surveys, the nodes show at least
343 the same performance levels, even beyond their natural frequency, while having the ad-
344 vantage of highly reduced purchasing costs, weight, and installation and dismantling time.
345 This study endorses the use of SmartSolo nodes as low-budget alternatives, either for
346 Large N installations or for research groups that have limited financial resources to per-
347 form seismotectonic or ambient noise studies using more expensive but higher-quality
348 seismic sensors.

349 **Data and Resources**

350 All seismic waveforms processed for this study have been obtained at the Royal Obser-
351 vatory of Belgium and are available alongside with the publicly available python codes
352 at <https://gitlab-as.oma.be/martinz/smartsolo-nodes-paper>.

353 **Declaration of Competing Interests**

354 The authors declare no competing interests.

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