

Mineral dust aerosol impacts on global climate and climate change

Jasper F. Kok^{1,†}, Trude Storelvmo², Vlassis A. Karydis³, Adeyemi A. Adebisi⁴, Natalie M. Mahowald⁵, Amato T. Evan⁶, Cenlin He⁷ and Danny M. Leung¹

¹*Department of Atmospheric and Oceanic Sciences, University of California - Los Angeles, Los Angeles, CA, USA*

²*Department of Geoscience, University of Oslo, Oslo, Norway*

³*Institute for Energy and Climate Research: Troposphere, Forschungszentrum Jülich GmbH, Jülich, Germany*

⁴*Department of Life and Environmental Sciences, University of California - Merced, Merced, USA*

⁵*Department of Earth and Atmospheric Sciences, Cornell University, Ithaca, NY, USA*

⁶*Scripps Institution of Oceanography, University of California, San Diego, CA, USA*

⁷*Research Applications Laboratory, National Center for Atmospheric Research, Boulder, CO, USA*

[†]*Corresponding author: jfkok@ucla.edu*

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Abstract

Mineral dust aerosols impact Earth's energy budget through interactions with radiation, clouds, atmospheric chemistry, the cryosphere and biogeochemistry. In this review, we summarize these interactions and assess the resulting impacts of dust, and of changes in dust, on global climate and climate change. We find that the total effect of these interactions on Earth's global energy budget—the dust effective radiative effect—is $-0.2 \pm 0.5 \text{ Wm}^{-2}$ (90% confidence interval). Compared to pre-industrial times, global dust mass loading is $55 \pm 30\%$ higher in the modern climate, leading to changes in the Earth's energy budget. Indeed, this increase in dust has produced a global mean effective radiative forcing of $-0.07 \pm 0.18 \text{ Wm}^{-2}$. Current climate models and climate assessments do not represent the historical increase in dust and thus omit the resulting radiative forcing, biasing climate change projections and assessments of climate sensitivity. Climate model simulations of future changes in dust diverge widely and are very uncertain. Further work is thus needed to constrain the radiative effects of dust on climate and to improve the representation of dust in climate models.

Key points

1. The direct radiative effect due to dust interactions with radiation is $-0.15 \pm 0.35 \text{ Wm}^{-2}$ and accounts for a large fraction of the dust effective radiative effect and its uncertainty.
2. Dust interactions with clouds, atmospheric chemistry, the cryosphere, and biogeochemistry also contribute considerably to the uncertainty in the dust effective radiative effect, in part because of a lack of observational constraints, which are urgently needed.
3. Dust mass loading generated by all major source regions has increased since pre-industrial times, namely by 46 (2-102) % for North African dust, by 74 ± 37 % for Asian dust, and by 27 (-14 to 88) % for Southern Hemisphere dust.
4. It is more likely that dust net cools than that it net warms global climate.
5. The historical increase in dust has likely somewhat counteracted greenhouse warming.
6. Careful simulations with coupled climate models that reproduce the historical dust increase are needed to better constrain dust radiative forcing.

Introduction

Mineral dust aerosols are small rock-derived particles with diameter $D < \sim 100 \mu\text{m}$ that are suspended in the atmosphere^{1,2}. Most dust is produced by the ballistic impacts of wind-driven sand grains on sparsely vegetated and dry soils³, which ejects and fragments aggregates of soil particles^{1,4}. Owing to these mechanical impacts, dust is a relatively coarse aerosol, with most of its mass contained in the coarse ($D > 2.5 \mu\text{m}$) and super coarse ($D > 10 \mu\text{m}$) modes⁵.

Dust is produced in copious amounts in the world's deserts, loading the atmosphere with ~ 26 million tonnes of dust, which accounts for a large majority of the atmosphere's aerosol burden by mass^{6,7}. The Sahara Desert and the Sahel contribute $\sim 50\%$ of global dust emissions ($\sim 2100 \text{ Tg/yr}$) and mass loading ($\sim 13 \text{ Tg}$), the Asian deserts $\sim 40\%$ ($\sim 2000 \text{ Tg/yr}$ and $\sim 10 \text{ Tg}$), and the North American and Southern Hemisphere deserts and high latitude regions another $\sim 10\%$ ($\sim 500 \text{ Tg/yr}$ and $\sim 3 \text{ Tg}$) (**Fig. 1**)^{8,9}. Although much of the dust is deposited close to source regions, a substantial fraction is transported for thousands of kilometres. For example, plumes of African dust regularly travel across the tropical North Atlantic, reaching the southwestern United States and the Amazon Basin¹⁰.

The abundance and long-range transport of dust cause it to impact climate through various mechanisms. During transport, dust scatters and absorbs solar shortwave (SW) and terrestrial longwave (LW) radiation^{6,11}, modifies cloud properties through seeding cloud droplets and ice crystals^{12,13}, mixes with other aerosols¹⁴, and serves as a sink for radiatively important atmospheric trace gases¹⁴⁻¹⁷. Upon deposition, dust darkens snow and ice packs^{18,19}, and stimulates ecosystem productivity and CO_2 drawdown through the delivery of iron and phosphorus²⁰. Because some of these mechanisms cool whereas others warm^{6,14,21}, it is unclear whether dust exerts a net cooling or a net warming effect on global climate. Because measurements of dust deposition suggest that dust has increased since the pre-industrial era^{22,23}, this uncertainty in the sign and magnitude of dust radiative effects means it is unknown whether dust changes have enhanced or opposed anthropogenic warming.

In this review, we examine the impacts of dust, and of changes in dust, on global climate and climate change. We first summarize the various mechanisms through which dust impacts Earth's radiation budget, and assess the radiative effect produced by each mechanism. We then constrain the increase in dust loading since pre-industrial times and assess the radiative perturbation produced by this historical increase in dust. We also discuss the radiative perturbation due to possible future changes in dust and end with recommendations for future research priorities.

Mechanisms by which dust impacts climate

Dust can perturb Earth's energy balance via various mechanisms. In each case, a radiative effect arises, defined as the imbalance between incoming net solar radiation and outgoing infrared radiation at the top-of-atmosphere (TOA) resulting from an atmospheric constituent (in this case, dust)²⁴. These effects can be either instantaneous, such as scattering and absorbing SW and LW radiation, or an adjustment, such as altering cloud cover²⁵.

We calculate the radiative effect due to mechanism i in the modern climate, r_i (Wm^{-2}), as the change in Earth's energy balance, Δf_i (Wm^{-2}), produced per change in global dust mass loading from modern levels, ΔL_i (Tg), multiplied by the global modern dust loading, L (Tg). That is,

$$r_i \equiv \frac{\Delta f_i}{\Delta L_i} L. \quad (1)$$

The sum of all radiative effects then equals the effective radiative effect of dust, R (Wm^{-2}), which includes both instantaneous radiative effects and adjustments^{25,27},

$$R = \sum_i r_i. \quad (2)$$

Eqs. (1) and (2) define the dust effective radiative effect in such a way that it can be used to obtain the radiative perturbation, ΔF , due to a change in dust loading, ΔL_m , from its value in the modern climate,

$$\Delta F = R \frac{\Delta L_m}{L}. \quad (3)$$

We then define the effective radiative forcing of dust due to the change $\Delta L_{p \rightarrow m}$ in dust mass loading from pre-industrial to modern times as

$$\Delta F_{p \rightarrow m} = R \frac{\Delta L_{p \rightarrow m}}{L}. \quad (4)$$

Our use here of the term radiative forcing deviates slightly from previous work^{25,27} in which it denotes radiative perturbations that are entirely from anthropogenic forcing agents. However, because dust is a natural aerosol affected by climate changes and human land use changes, a radiative perturbation due to a historical change in dust can be partially due to both human land use changes (a forcing) and natural and anthropogenic climate changes (a feedback). Because these two contributions are difficult to disentangle, we refer to the entire radiative perturbation due to the historical change in dust as the **dust** effective radiative forcing.

Radiative effects from dust arise through interactions with radiation, atmospheric chemistry, clouds, the cryosphere, and biogeochemistry (**Fig. 2**). Each of these mechanisms are now discussed.

Interactions with radiation

Perhaps the best understood mechanism by which dust impacts climate is through the dust direct radiative effect (DRE), the perturbation of Earth's energy balance by scattering and absorption of radiation (**Fig. 2a**). Since dust spans a wide range of sizes, from $\sim 0.1 - 100 \mu\text{m}^{29}$, it interacts with both SW (centered around 550 nm wavelength) and LW (centered around 10 μm wavelength) radiation^{30,31}.

The sign and magnitude of the dust DRE depend on the balance between these interactions. For instance, scattering of SW radiation cools the climate while absorption of SW radiation warms, with an overall net cooling^{6,32}. In contrast, scattering and absorption of LW radiation both warm the climate since both decrease the transparency of the atmosphere to terrestrial LW radiation³³. Thus, the balance between cooling from SW scattering, and warming from SW absorption and LW scattering and absorption, dictate the dust DRE.

For SW radiation, the balance between scattering and absorption is influenced by dust particle size. Because absorption increases more strongly with particle size than scattering^{34,35}, the single-scattering albedo (SSA; the ratio of scattered radiation to total extinguished radiation) decreases with particle size. Indeed, submicron dust has an SSA close to 1, whereas supermicron dust absorbs a substantial fraction of extinguished radiation, exhibiting SSAs of ~ 0.95 at $D = 2 \mu\text{m}$, ~ 0.80 at $D = 10 \mu\text{m}$, and even lower for super coarse dust^{30,36}.

However, the exact SSA of dust aerosols depends on their complex refractive index, determined by particle mineralogy³¹. Absorption increases approximately linearly with iron oxide content, which is primarily provided by hematite and goethite³⁷. Dust optical properties can also be affected by mixing with other aerosols, especially black carbon³⁸. Observations suggest that this possible mixing has limited impact on the optical properties of most African dust^{39,40}, but could substantially affect those of East Asian dust⁴¹.

Although dust particle size and mineralogy determine the balance between SW scattering and absorption, the efficiency with which both processes perturb the TOA radiative flux depends on the albedo of the underlying surface. Indeed, the cooling effect of SW scattering is enhanced if the dust is situated above dark (low albedo) surfaces like the ocean and forests that would otherwise absorb most of the radiation⁴². Conversely, the warming effect of SW absorption is enhanced if the dust is situated above clouds or above high albedo land surfaces like snow, ice, and deserts that would otherwise scatter most of the radiation back to space^{35,43}.

Dust microphysical properties and mineralogy also influence the extinction of LW radiation. For example, because of its longer wavelengths, LW radiation is extinguished primarily by coarse dust^{30,33,44}. The sensitivity of LW extinction to mineralogy is less important than for SW interactions owing to the smaller variability in LW optical properties between minerals, and because LW scattering and absorption both warm the planet^{31,45,46}.

The efficiency with which dust extinction of LW radiation perturbs the TOA radiative flux also depends on the atmosphere's transparency to LW radiation and the elevation of the dust layer. Indeed, the TOA flux is only substantially impacted if the atmosphere is at least somewhat transparent to LW radiation, as is the case in the absence of clouds in the $\sim 8 - 13 \mu\text{m}$ 'atmospheric window' wavelength range^{33,47}. Furthermore, because LW emission depends on temperature, the LW warming depends on the temperature difference between the dust layer and the source of the LW radiation - usually the surface or clouds below the dust layer. In addition, the atmosphere's transparency to LW radiation decreases with the concentration of water vapor

and thus increases with height. As such, dust warming by LW extinction increases approximately linearly with the height of the dust layer^{33,42,47,48}.

Although the processes by which dust interacts with SW and LW radiation are relatively well understood, the resulting radiative effects are poorly constrained. For dust interactions with SW radiation, central estimates of SW DRE are -0.40 Wm^{-2} (-0.10 to -0.70 Wm^{-2} , 90% confidence interval)^{6,32,49-53} (**Fig. 3**); these estimates are determined using less absorbing optical properties and a coarser dust size distribution, consistent with experimental constraints^{4,6,29,37,40,54-56}. The wide range reflects substantial uncertainties in the dust size distribution⁶ and dust optical properties⁴⁶. For dust interactions with LW radiation, best estimates of LW DRE are $+0.25 \text{ Wm}^{-2}$ with a range of $+0.10$ to $+0.40 \text{ Wm}^{-2}$ (**Fig. 3**)^{6,32,46,49-52}; these estimates use realistic optical properties⁴⁵ and size distributions that are consistent with satellite constraints on the LW direct radiative effect^{48,57}. The range reflects uncertainties in dust LW optical properties⁴⁵, the height of dust plumes^{58,59}, the dust size distribution and the contribution of super coarse dust^{36,49,54}, and the effect of LW scattering by dust, which is neglected in climate models^{32,33} and is sometimes accounted for using a simple correction factor^{6,32,46,52}.

As a result of the uncertainties and opposing SW and LW DRE, it is unclear whether the dust DRE exerts a net cooling or warming effect. Combining the SW and LW DRE yields a net dust DRE of $-0.15 \pm 0.35 \text{ Wm}^{-2}$, consistent with other calculations^{6,32,46,49-51} (**Fig. 3**). As such, the dust DRE could either slightly warm or substantially cool the planet, or it could have little net impact. We assign medium confidence to this assessment because of the large body of research and availability of satellite-based constraints.

Interactions with atmospheric chemistry

Dust affects atmospheric chemistry through numerous interactions with atmospheric trace gases and aerosols. Although freshly emitted mineral dust is considered insoluble, it is reactive towards trace acidic gases derived from anthropogenic pollutants and sea salt^{17,60}. Mineral dust particles collected throughout the world are notably associated with nitrate⁶¹⁻⁶³. Nitric acid interacts with the non-volatile mineral cations of dust, forming salts to maintain the charge balance in the aerosol phase⁶⁴. The uptake of such acidic vapors is very rapid due to their ability to react with carbonates and other minerals through simple acid-base chemistry⁶⁵. Over continents, such interactions of mineral cations with anthropogenic sulfuric acid causes the accumulation of substantial amounts of sulfate on dust surfaces⁶⁶. In contrast, over oceans, mineral cations are commonly associated with chloride derived from sea salt⁶⁷.

Mineral dust also provides surfaces for the adsorption of inorganic (notably SO_2 , NO_2 , and O_3) and organic trace gases¹⁷, affecting the optical properties, hygroscopicity and atmospheric residence time of both dust and anthropogenic aerosols. Therefore, dust particles provide a substantial sink for the direct removal of important atmospheric constituents like O_3 , affecting the oxidative capacity of the atmosphere and the ozone radiative forcing^{68,69}. Dust particles also provide reaction sites for the oxidation of SO_2 to sulfuric acid⁷⁰ and the formation of nitrous acid through heterogeneous reactions of NO_2 ⁷¹. However, such heterogeneous formation of salts occurs at a much slower pace than through the direct uptake of acidic vapours since acid anhydrides (for example SO_2) do not initially contain any acidic protons⁶⁵. Additionally, the high pH values found on the alkaline mineral particles can promote the formation of ammonium nitrate on their surface^{72,73}.

All these interactions of dust with atmospheric gases can transform the surface and even the bulk chemical composition of dust particles^{74,75}. This chemical processing of dust is highly dependent on both the gas phase composition and on the dust chemical composition^{64,76}, which depends on the mineralogy of the source soil⁷⁷.

The chemical ageing of dust due to these various reactions creates a soluble coating that increases the dust particle's hydrophilicity, which in turn affects the residence time of dust and its interactions with clouds. For example, the interaction of a calcite-containing dust particle with nitric acid converts the insoluble calcium carbonate to the highly hygroscopic calcium nitrate⁷⁸. The increased hygroscopicity of the chemically aged dust increases its water adsorption efficiency, making it grow more rapidly under humid conditions, thus causing it to form cloud droplets and extinguish radiation more efficiently. On the contrary, the increased water uptake by the large, aged dust particles can also deplete in-cloud supersaturation, thereby reducing the number of smaller anthropogenic particles that are activated and grow into cloud droplets¹³. Furthermore, chemical ageing of mineral dust can also reduce its ice nucleating ability⁸⁰.

These heterogeneous and multiphase reactions affect the atmospheric loading of both dust and non-dust aerosols. Nitrate formation associated with the mineral cations removes nitric acid from the gas phase, decreasing the formation of ammonium nitrate aerosols. Similarly, sulfate formation on dust decreases SO₂ abundance and thus the formation of sulfate aerosols. As such, dust can reduce the concentration of anthropogenic CCN both by adsorption of precursor gases and through coagulation with anthropogenic aerosols. Furthermore, the hygroscopic growth of aged dust can increase its scavenging and deposition rate, reducing its atmospheric residence time and loading^{15,81}. However, modelling results suggest that these effects can enhance the total accumulation mode dust burden through a reduced loss by coagulation with coarse dust particles¹⁵.

The physicochemical interactions of mineral dust with atmospheric composition can thus affect the direct and indirect radiative forcing of both dust and non-dust aerosols (Fig. 2b). These effects can be both negative or positive, depending on the region and the prevailing impacts of dust on the atmospheric aerosol loading and composition^{15,16}. A net cooling effect of -0.05 Wm^{-2} has been calculated for the effect of dust on the total aerosol DRE¹⁵, driven mostly by an enhanced burden of the accumulation mode dust aerosols and decreased absorption of SW radiation due to the modified aerosol composition of mineral dust. However, observations of dust during transport across the Atlantic and Mediterranean oceans indicate that the size distribution of dust with diameters less than $5 \mu\text{m}$ remains remarkably constant and that optical properties do not change appreciably^{39,40,55}. Moreover, a critical effect of heterogeneous chemistry on dust surfaces is to reduce the atmospheric loading of anthropogenic aerosols, thereby decreasing their direct radiative cooling, resulting in a net warming of $+0.12$ to $+0.20 \text{ Wm}^{-2}$ ^{16,82,83}. Overall, the impact of dust interactions with atmospheric chemistry on the aerosol DRE is highly uncertain. The resulting radiative effect is assessed at $0.10 \pm 0.15 \text{ Wm}^{-2}$ to encompass the possibility of the slight cooling of -0.05 Wm^{-2} ¹⁵ as a lower bound and the larger warming found by others^{16,82,83} as an upper bound. We assign very low confidence to this assessment.

Interactions with clouds

Dust particles influence clouds via multiple interactions, including changing the thermodynamic environment by absorbing SW and LW radiation and serving as cloud condensation nuclei (CCN) and ice-nucleating particles (INPs). Radiative perturbations produced by dust effects on

warm clouds, mixed-phase clouds, cirrus (ice) clouds, and by semi-direct effects are discussed next.

Dust indirect effects on warm clouds

There are three main pathways through which dust particles can affect warm clouds: first, by increasing the concentration of CCN, as laboratory studies have shown that various types of (unprocessed) mineral dusts possess a modest ability to act as CCN^{84,85}, which is further enhanced by atmospheric processing (ageing) of dust⁸⁶; second, by reducing the concentration of non-dust CCN through coagulation and adsorption of precursor gasses; and third by acting as giant CCN, which can form cloud droplets at relatively low supersaturation and thus deplete water vapor to such an extent that overall cloud droplet formation is suppressed. The second and third pathways are both thought to reduce cloud droplet number concentrations (CDNC), and thus reduce cloud albedo and shorten cloud lifetimes.

Although some modelling results found that the pathways that decrease CDNC dominate (Fig. 2c, lower branch), amounting to a decrease of as much as -11% in the global mean CDNC¹³, most other modelling has found that dust slightly increases the global mean CDNC abundance. These latter results thus suggest that the effect of dust acting as CCN dominates (Fig. 2c, upper branch), albeit with large differences in the magnitude of the dust-induced CDNC contribution^{21,87,88}. As a CDNC increase is expected to increase cloud albedo and extend cloud lifetime through well-established mechanisms²⁸, a dust-induced increase in CDNC would have a net cooling effect (Fig. 2c, upper branch). For example, global simulations with the CAM5 model²¹ resulted in a ~1% increase in CDNC for a 3-fold increase in dust emissions, and this CDNC increase in turn produced a negative forcing of -0.01 Wm^{-2} . Such an effect is indeed supported by estimates based on satellite observations⁸⁹.

Although past work thus reached contradictory conclusions regarding the net global effects of dust on warm clouds, there is broad agreement that the sign and magnitude of the dust contribution to CDNC is highly heterogeneous in both space and time^{13,21}. Given the relatively sparse research and disagreement on the sign of the global mean CDNC contribution from dust, we assess the corresponding perturbation to Earth's TOA radiation budget through changes to liquid clouds to likely be negative but close to zero, with an uncertainty range of -0.10 to $+0.10 \text{ Wm}^{-2}$. This assessment is based on scaling the estimates of CCN/CDNC changes^{13,21,87,88} with the forcing estimate per change in CCN/CDNC cited above²¹, and has low confidence.

Dust indirect effects on mixed-phase clouds

Although the ability of dust to act as CCN is somewhat ambiguous, their ice-nucleating ability is undisputed^{90,91}. A wide variety of dust particles have been investigated in the laboratory and found to be efficient INPs both in the immersion mode (freezing cloud droplets from within) and in the deposition mode (nucleating ice through vapor depositing onto them, possibly triggered by freezing of condensed water in particle pores⁹²). The former is the ice formation mechanism thought to be of greatest relevance for mixed-phase clouds (MPCs). These are clouds with temperatures between approximately -38°C and 0°C that can consist of either supercooled liquid droplets, ice crystals, or a mixture. MPCs are generally optically thick and efficiently reflect incoming SW radiation (a cooling effect). Their optical thickness also allows them to absorb virtually all outgoing LW radiation, reducing the amount of LW radiation emitted to space (a warming effect). The former (SW) effect has been found to dominate in the global mean⁹³. In an

INP-limited (“pristine”) environment, MPCs will be optically thick and usually have liquid cloud tops⁹⁴, with only small amounts of ice residing in the cloud interior or below cloud base as ice crystals rapidly grow and sediment out (Fig. 2d, left schematic). In a dust-enriched environment, MPCs will be partly or completely glaciated, depending on the dust abundance and INP efficiency. This cloud glaciation results in an overall reduction of cloud albedo and thus a positive (warming) radiative effect (Fig. 2d, right schematic). An increase in dust loading, and thus INPs, therefore likely produce a warming effect on climate by reducing the cooling effect of MPCs (Fig. 2d).

Modelling results on the effects of dust on MPCs generally agree qualitatively, but differ quantitatively. Global simulations with the E3SM model⁹⁵ found that dust effects on mixed-phase clouds perturb the TOA radiation budget by +0.05 to +0.26 Wm⁻². This perturbation arises from a reduction in cloud liquid water and a corresponding increase in cloud ice (Fig. 2d). For comparison, global simulations with the CAM5 model estimated that going from a very pristine state with only 10% of current dust emissions to present-day dust emissions induced a perturbation of only 0.01 to 0.10 Wm⁻² through dust-INP effects on mixed-phase clouds, again by shifting cloud phase in favor of more ice²¹. However, as the atmospheric dust loading change in the latter study is smaller than in the former, these estimates are broadly consistent with each other. These modelling results are further supported by satellite observations that found that dust-enriched environments tend to have MPCs with a larger proportion of ice than their counterparts in largely dust-free environments^{96,97}. Thus, a perturbation to the TOA radiation budget of approximately 0.10 Wm⁻² due to dust effects on MPCs is supported, but with a relatively large assessed uncertainty range of 0 to 0.20 Wm⁻² and low confidence, owing to the limited body of research.

Dust indirect effects on cirrus clouds

The dominant role of dust particles in cirrus cloud formation worldwide is supported by in situ measurements, satellite observations, and numerical modelling^{98,99}. Cirrus clouds are pure ice clouds residing in the upper troposphere at temperatures below approximately -38°C. These clouds have a net warming effect on climate by reducing emission of LW radiation to space more effectively than they reflect SW radiation¹⁰⁰. Cirrus clouds can form by two different mechanisms: homogeneous freezing, in which small solution droplets freeze spontaneously, and heterogeneous freezing, in which ice crystals form on INPs⁹¹. The latter mechanism requires only modest supersaturation but can only occur when sufficient INPs are present and typically results in low concentrations of large ice crystals. The former mechanism requires much higher supersaturation but does not rely on the presence of INPs and typically results in high concentrations of small ice crystals¹². The transition from homogeneous to heterogeneous freezing has been estimated to occur for INP concentrations between 10 and 100 L⁻¹.¹⁰¹

The impact of dust on cirrus clouds is thus highly dependent on whether non-dust INPs are present (Fig. 2e). In conditions that favor heterogeneous freezing (high INP concentration), additional dust INPs would add ice crystals and reduce their size, while in conditions that favor homogeneous freezing (low INP concentration), additional dust could reduce the number of ice crystals and increase their size by shifting nucleation from occurring homogeneously to occurring heterogeneously. The former scenario would make cirrus clouds optically thicker and extend their lifetimes, while the latter scenario would do the opposite.

The perturbation of the TOA radiation budget would naturally be opposite in the two scenarios, and at present it is unclear which one dominates globally. Thus, although global modelling results of dust impacts on cirrus clouds have in the past produced net radiative perturbations of opposite sign¹², this difference does not signify a complete lack of process understanding, but rather indicates different assessments of which cirrus formation mechanism dominates in the absence of dust.

Research that incorporated up-to-date laboratory results of ice nucleation on dust particles^{92,102,103} generally find an optical thinning of cirrus clouds due to dust (Fig. 2e, top schematic). This thinning yields large opposing perturbations to both LW and SW radiation at the TOA, but the LW effect tends to dominate, producing a net negative (cooling) perturbation^{21,104}. The corresponding overall radiative effect was estimated at -0.4 Wm^{-2} using global simulations with the CAM5 model¹⁰⁴, whereas simulations for a more moderate dust change (going from 10% to 100% of present emissions) with a modified version of the same model²¹ found a range from -0.32 to $+0.05 \text{ Wm}^{-2}$. We therefore assess the perturbation of the TOA radiation budget due to dust effects on cirrus clouds to -0.20 Wm^{-2} , with a 90% confidence interval of -0.40 to $+0.10 \text{ Wm}^{-2}$. This range encompasses the strongest reported cooling effects¹⁰⁴ as a lower bound and the possibility of a slight warming as an upper bound. We assign low confidence to this assessment due to the limited body of research.

Dust semi-direct effects on clouds

Absorption of radiation by mineral dust can modify the temperature profile¹⁰⁵, which can change atmospheric stability, the moisture profile, and secondary circulations, all of which can alter cloud distributions¹⁰⁶⁻¹⁰⁸. These processes, known as aerosol semi-direct effects (SDEs)^{109,110}, were broadly described in the IPCC's Sixth Assessment Report as atmospheric adjustments to instantaneous aerosol direct radiative effects without considering effects due to changes in surface temperature^{25,28}. Because dust accounts for about a third of shortwave absorption by all aerosols, the contribution of dust to SDEs is crucial to accurately quantifying the overall dust effective radiative forcing^{111,112}.

The magnitude of the dust SDE, and whether it results in a positive (warming) or a negative (cooling) radiative effect, depends primarily on two factors: the relative position of the dust and cloud layers within the atmospheric column and the amount of radiation absorbed by the dust layer^{106,107}. In turn, radiation absorption by dust depends on dust loading and microphysical properties, including dust mineralogical composition, shape, and size distribution^{42,43,45,113}.

Understanding of the pathways through which dust semi-directly impacts different cloud regimes follows that of SDEs produced by other absorbing aerosols, like black carbon^{106,114}. For low-altitude clouds, the pathways for dust SDEs can be categorized into cases where the dust layer is above, within or near, and below the cloud layer (Fig. 2f). When dust is located above boundary-layer clouds, local heating by the above-cloud dust can stabilize the boundary layer by strengthening its capping inversion, causing an increased build-up of moisture in the boundary layer. This increased moisture increases the cloud cover, which results in a negative SDE (left schematic of Fig. 2f)^{107,115}. Conversely, when dust is located within or near boundary-layer clouds, the local heating could result in reduction of relative humidity, which could evaporate the cloud and result in a positive SDE (middle schematic of Fig. 2f)^{107,116}. Finally, when dust is located below boundary-layer clouds, the local heating may enhance convergence and available

moisture, increasing cloud cover and resulting in a negative SDE (right schematic of Fig. 2f)^{117,118}.

Radiation absorption by dust can also generate SDEs for mid and high-altitude clouds. These SDEs involve the compensation between the warming effect produced by dust absorption, which tends to decrease cloud cover, and an increase in moisture convergence, which tends to increase cloud cover¹⁰⁶. Although the effect of the enhanced moisture convergence can overwhelm the warming effect, resulting in increased globally averaged high-altitude cloud cover during the summer, the overall annual-mean dust SDE is to decrease the high cloud cover^{106,119,120}.

This understanding of dust SDEs assumes that dust, like other absorbing aerosols such as black carbon, warms the atmospheric layer in which they are present¹²¹. This assumption is based on evidence that dust radiative warming due to SW absorption dominates over dust radiative cooling due to LW emission^{122,123}. However, past research likely underestimated the amount of coarse dust, which emits LW radiation more strongly than fine dust^{49,56}. Because accounting for the observed abundance coarse dust particles could produce substantial LW radiative cooling of the atmosphere^{32,36,124}, the understanding of the different pathways through which dust can semi-directly impact clouds remains incomplete.

Because of the uncertainties in the various pathways by which dust absorption semi-directly influences cloud cover (Fig. 2f), a global observational estimate of dust SDE is not currently available. Instead, observationally based assessments have focused on dust-dominated regions^{107,116,125}. For example, satellite observations show that annual dust SDE is negative ($-1.2 \pm 1.4 \text{ Wm}^{-2}$) over the North Atlantic Ocean¹⁰⁷. Since estimates of dust SDE show strong spatial variability and because dust SDE is driven by different dominant mechanisms for different cloud regimes over the ocean than over land¹⁰⁶, scaling such observationally based regional dust SDE estimates to global values is difficult. In addition, accurate retrievals of dust microphysical properties, including dust optical properties and size distribution, are lacking from global-scale satellite and ground-based platforms¹¹², making it difficult to obtain global estimates of dust SDE.

In the absence of global observational estimates, climate models simulations have reported a net positive global annual mean dust SDE¹²⁶. These estimates vary by over an order of magnitude, between 0.01 and 0.16 Wm^{-2} , and depend on the climate model used^{21,126,127}. These positive SDE estimates are consistent with an overall decrease in cloud cover in these simulations. Although model estimates of dust SDE and cloud changes are thus relatively consistent with each other, they could be biased because of unaccounted for uncertainties in dust absorption properties, the vertical distributions of dust and clouds, an underestimate of LW radiative cooling by coarse dust, and the parameterization of cloud processes^{56,58,111}. Therefore, based on the above model simulations, the dust SDE is estimated at $0.07 \pm 0.07 \text{ Wm}^{-2}$, but with low confidence due to these possible biases and limited research.

Interactions with the cryosphere

Dust interactions with the cryosphere impact climate by altering cryospheric conditions via dust direct and indirect radiative effects (Figs. 2a-f) and by darkening snow and ice surfaces after deposition (Fig. 2g), which leads to a positive surface radiative effect (Fig. 2g). This dust-induced snow albedo effect accelerates snow and glacier melting^{18,128,129}, which triggers a strong, positive surface albedo feedback on the climate system¹³⁰. The dust-induced snow albedo effect is influenced by many factors, including dust concentration in snow^{131,132}, dust optical properties

as determined by its size distribution and chemical composition^{132,133}, dust-snow mixing state^{134,135}, snow grain size and shape¹³⁴, snowpack properties^{136,137}, and illumination conditions^{132,138}.

Observations indicate strong heterogeneity in dust concentrations in snow/ice. Along with different snowpack and atmospheric conditions, this variability in dust concentrations leads to large variations in the dust-induced snow albedo reduction and the associated surface radiative effects. For instance, the springtime dust-induced snow albedo effect is estimated to be less than 0.5 Wm^{-2} for the Arctic^{139,140}, up to 5 Wm^{-2} for remote mid-latitude snowpacks (such as the Tibetan Plateau)^{139,141}, and about $10\text{--}50 \text{ Wm}^{-2}$ over polluted mid-latitude snowpacks (such as the U. S. Rocky Mountains)^{18,128}. In some extremely polluted mid-latitude mountains, the local instantaneous snow albedo effect can be as high as $100\text{--}300 \text{ Wm}^{-2}$ ^{142,143}. The dust-induced snow albedo effect is typically larger in aged snow than in fresh snow¹⁸, because of the stronger light penetration and hence larger light absorption by dust in aged snow. Most research has focused on a few cryospheric hotspots in the Northern Hemisphere (the Rocky Mountains, Tibetan Plateau, European Alps, and the Arctic) during spring, when the dust-induced snow albedo effect is more prominent and often reaches its annual maximum.

There are only limited estimates of the global annual mean dust-induced snow albedo effect, with a central estimate of $+0.013 \text{ Wm}^{-2}$ and a 90% confidence interval of $0.007\text{--}0.03 \text{ Wm}^{-2}$ ^{19,131,144}. Although the snow albedo radiative effect is smaller than most other dust radiative effects, it can still be more substantial regionally, particularly over polluted mid-latitude snowpacks¹⁸.

Estimates of the dust-induced snow albedo effect are still associated with large uncertainties due to complicated and poorly constrained dust-snowpack-radiation interactions. Variations in the poorly constrained dust-snow mixing state, snow grain shape, dust size distribution and dust chemical composition can cause up to a factor of two uncertainty in the dust-induced snow albedo effect^{134,135}. Moreover, the limited knowledge of dust evolution within the snowpack - for instance due to dust scavenging by melting water and dust enrichment at the snowpack surface - also adds to the uncertainty of the estimated snow albedo effect. Owing to the potential nonlinearity in dust-snow-radiation interactions and dust wet deposition, the dust-induced snow albedo effect may not increase linearly with dust concentration in the atmosphere or snowpack. Considering these uncertainties and the limited research, we assign low confidence to our estimate of the dust-induced snow albedo effect.

Interactions with biogeochemistry

Dust can influence ocean and land biogeochemistry, both directly through the addition of nutrients and pollutants to ecosystems, as well as indirectly through modifying precipitation, temperature, and radiation²⁰. Atmospheric deposition of dust onto oceans provides iron, a limiting nutrient in high nutrient low chlorophyll (HNLC) regions^{145,146}. In addition, nitrogen fixing organisms in the ocean have higher iron requirements, thereby linking iron deposition to the oceanic nitrogen cycle^{147,148}. Although initial research suggested that atmospheric deposition was the dominant source of new iron^{145,149}, other ocean sources also have a substantial role in the iron cycle¹⁵⁰⁻¹⁵². Overall, atmospheric inputs of iron to the ocean modulate ecosystem productivity and carbon sequestration on the timescale of decades^{146,153}.

, The soluble fraction of the iron is the most important for dust particles sinking through the ocean mixed layer. The deposition of soluble iron has increased since pre-industrial times, both

because of the historical increase of dust over this time period and because of an increase in iron solubilization during transport due to increased anthropogenic pollution^{22,154,155}. Additionally, some other important sources of soluble iron have also increased, including from wildfires and anthropogenic combustion^{156,157}. The resulting alleviation of iron limitation has increased ecosystem productivity, which in turn has reduced the atmospheric concentration of carbon dioxide and its radiative forcing (Fig. 2h).

Several ocean biogeochemical models include iron and its coupling to the nitrogen cycle and can therefore estimate the reduction of CO₂ concentrations due to the alleviation of iron limitation^{158,159}. These models suggest that the increased deposition of soluble iron over the 20th century resulted in the uptake of ~4 ppm of CO₂, producing a radiative perturbation of $-0.07 \pm 0.07 \text{ Wm}^{-2}$ ^{20,160}. Because approximately half of this increase in soluble iron was estimated to be due to a simulated ~40% increase in dust over the 20th century, these results imply a radiative effect due to dust-biogeochemistry interactions of $-0.12 \pm 0.12 \text{ Wm}^{-2}$ (Eq. 1). Confidence in this assessment is very low, as it is based on only one study. Note that the radiative effect due to dust-biogeochemistry interactions differ from that due to other interactions in that its effect increases over time. Consequently, the radiative perturbation that it produces depends on the timescale.

Dust also contains phosphorus, a limiting nutrient in some tropical forests and grasslands^{161,162}, as well as in some ocean ecosystems^{146,163}. For example, phosphorus from long-range transported North African dust may help maintain the productivity of the Amazon rainforest¹⁶⁴. However, because inputs from atmospheric deposition of desert dust are thought to be important in the Amazon on millennial time scales¹⁶⁵ any contribution of changes in this phosphorus input probably produces a negligible contribution to dust radiative forcing since pre-industrial times. Dust also serves as a ballast, enhancing the downward transfer of organic material within the ocean, but there is not yet a quantitative estimate of the impacts in terms of productivity or carbon uptake feedback from this process^{170,171}. In addition, desert dust could include elements that can be toxic to ocean or land ecosystems, such as Cu, although current estimates suggest that this effect is not important to Earth's radiation budget¹⁷².

The dust effective radiative effect

To determine the climatic impact of past and future changes in atmospheric dust, it is critical to assess the dust effective radiative effect R (Eq. 2), which equals the sum of the various radiative effects generated by dust (Fig. 3). Many of these radiative effects oppose one another, resulting in a median estimate of $R = -0.2 \text{ Wm}^{-2}$, with a wide 90% confidence interval of -0.7 to $+0.3 \text{ Wm}^{-2}$. (Note that we neglected some rapid adjustments in assessing R , such as responses by water vapor and the lapse rate to dust direct radiative effects, but these adjustments are likely small¹⁷³.) As such, the net effect of dust on Earth's global radiation budget could be negligible, a substantial net cooling, or a small net warming.

On regional scales and for different seasons, the dust effective radiative effect can differ substantially from its global and annual mean in Figure 3. This regional and seasonal variability occurs because the various radiative effects are sensitive to the spatiotemporal variability in dust concentration, microphysical properties (mineralogy and size distribution), and environmental conditions (surface albedo and cloud cover). For instance, dust over reflective deserts likely produces substantial warming because of the high dust concentration, coarse size distribution¹²⁴, and because reflective surfaces reduce cooling produced by SW scattering and enhance warming

produced by SW absorption^{35,174}. Similarly, dust likely produces net warming over snow and ice-covered regions because the high surface albedo enhances warming produced by dust absorption of SW radiation and because dust deposition decreases the surface albedo^{42,131}. In contrast, dust over oceans usually produces cooling because dust is finer further from source regions and because the ocean albedo is only ~ 0.1 in the visible spectrum¹⁷⁵. To determine the climate impacts of dust, it is thus critical not only to constrain the global mean dust effective radiative effect but also to constrain its spatiotemporal pattern.

Dust radiative forcing

Because dust produces a potentially large effective radiative effect, a change in atmospheric dust loading since pre-industrial times could have produced a substantial effective radiative forcing. Dust loading could have changed due to both climate change and widespread human land use changes (Box 1). Knowledge of the change in dust loading from pre-industrial to modern times depends largely on dust deposition records that resolve both the modern and the pre-industrial climate. Many of these deposition records show increases in dust deposition between modern and pre-industrial times, sometimes by a factor of ~ 4 ^{23,169,176,177}.

Dust reconstruction

We reconstructed the evolution of the global dust mass loading since pre-industrial times by combining 19 dust deposition records^{23,169,176,177} with constraints on the source regions providing the deposition flux to each deposition core^{8,44} (see Supplementary Information). This dust reconstruction used a bootstrap resampling method to propagate uncertainties in both the experimental deposition records and the constraints on source region-resolved deposition fluxes to each deposition site; nonetheless, errors should be interpreted as a lower bound.

The atmospheric loading of dust with a volume-equivalent diameter less than $20 \mu\text{m}$ has increased from $19 \pm 6 \text{ Tg}$ in the pre-industrial period (defined here as 1841-1860) to $29 \pm 8 \text{ Tg}$ in the modern climate (1981-2000). As such, global dust mass loading has increased by $55 \pm 30 \%$ (Fig. 4a). Although substantial, this increase is less than the doubling of dust suggested by previous research^{22,23}. A large contributor to this increase has been Asian dust, which has increased by $74 \pm 37 \%$ from $8 \pm 3 \text{ Tg}$ in pre-industrial times to $13 \pm 5 \text{ Tg}$ in modern times (Fig. 4c). North African dust has increased less, from $10 \pm 4 \text{ Tg}$ in the pre-industrial period to $14 \pm 4 \text{ Tg}$ in the modern climate, representing a 46 (2-102) % increase. Both African and Asian dust mass loading peaked in the 1980s and then decreased substantially, consistent with changes observed from long-term dust concentration measurements, visibility records, and satellite observations¹⁷⁸⁻¹⁸⁴. Dust has likely also increased in the Southern Hemisphere, from $1.2 (0.6-2.2) \text{ Tg}$ to $1.6 (0.8-2.2) \text{ Tg}$, representing a 27 (-14 to 88) % increase (Fig. 4d). Satellite observations suggest that global dust mass loading has been relatively stable since the year 2000, the end point of the analysis, with some notable regional trends, such as in Central and East Asia¹⁸⁵.

This large historical increase in dust mass loading is inadequately accounted for in current climate models and climate assessments. In fact, twelve climate models with prognostic dust cycles in the Coupled Model Intercomparison Project phase 6 (CMIP6) model ensemble^{186,187} show little change in dust mass loading since pre-industrial times (Fig. 5). This failure of models to reproduce the historical dust increase could be due to several reasons (Box 1). If the dust increase has been largely driven by natural and anthropogenic climate changes, then the model failure could be either due to an inaccurate representation of these changes in models or because

modelled dust emissions are not sufficiently sensitive to changes in climate. This latter possibility is suggested by the common use in climate models of empirical dust source functions to parameterize the spatial distribution of dust emissions^{188,189}. Because dust source functions are static, they mask physical links between changeable surface properties and dust emissions. As such, their use can cause models to underestimate the sensitivity of dust emissions to changes in climate¹⁹⁰. Conversely, if the dust increase has been largely driven by human land use changes (Box 1), as suggested by research indicating that approximately a quarter of current dust emissions originate from regions heavily impacted by human land use¹⁹¹ (Box 1), then the model failure to reproduce the dust increase could be caused by an underestimation of land use and land cover changes in drylands and the resulting increases in dust emissions.

Radiative forcing due to dust increase

The historical increase in dust loading could have produced a substantial radiative forcing. Combining $R = -0.2 \pm 0.5 \text{ Wm}^{-2}$ with the $55 \pm 30\%$ historical dust loading increase yields a dust effective radiative forcing from 1841 to 2000 of $\Delta F_{p \rightarrow m} = -0.07 \pm 0.18 \text{ Wm}^{-2}$. Dust radiative forcing could thus either have substantially contributed to, or slightly opposed, the total aerosol effective radiative forcing of -1.1 (-1.7 to -0.4) Wm^{-2} for the period of 1750 to 2019²⁸.

Note that the calculations of R and $\Delta F_{p \rightarrow m}$ are subject to important limitations. First, these calculations assume that radiative effects increase linearly with aerosol loading^{6,192} (Eqs. 3 and 4). However, the increase of radiative effects with aerosol loading is usually less-than-linear, especially for interactions with clouds and biogeochemistry^{20,21,26}. Moreover, the radiative effects of dust vary in space, such that $\Delta F_{p \rightarrow m}$ depends on the spatial pattern of dust increases, which the simple calculation here does not account for. For instance, Asian dust likely has an outsize impact on Northern Hemisphere cirrus clouds⁹⁸ and high latitude dust emissions are likely important in controlling the glaciation of mixed-phase clouds^{193,194} but are not explicitly included in the dust reconstruction. Careful simulations with coupled climate models that reproduce the historical dust increase are thus needed to better constrain dust radiative forcing.

Because current climate models do not reproduce the historical dust increase, these models omit the potentially important radiative forcing due to increased dust interactions with radiation, clouds, atmospheric chemistry and the cryosphere. (Note that changes in CO_2 and other greenhouse gases due to dust interactions with biogeochemistry are inherently included in climate model runs forced by observed greenhouse gas concentrations.) Dust radiative forcing was thus not accounted for in constraints on the total aerosol effective radiative forcing in the IPCC Sixth Assessment Report²⁸. Because constraints on climate sensitivity depend strongly on the aerosol radiative forcing since pre-industrial times¹⁹⁵, the failure by models and climate assessments to account for the historical increase in dust could thus have biased constraints on climate sensitivity and projections of future climate changes¹⁹⁶.

Future changes in dust radiative forcing

Future changes in dust radiative forcing are likely to be dominated by changes in atmospheric dust loading, which in turn will be determined by several factors. One important factor will be future changes in soil moisture since drier soils are more susceptible to aeolian erosion because of reduced soil cohesive forces and less vegetation^{1,2}. In CMIP5 and CMIP6 models, changes in precipitation are the main driver of soil moisture changes, yet there is a wide degree of

divergence in model projections of precipitation¹⁹⁷. Models do consistently show that as the planet warms the evaporative demand over land increases¹⁹⁸, which by itself would reduce soil moisture. However, the effects of reduced soil moisture may be countered by CO₂ fertilization, which reduces plant water losses. This could reduce dust emissions by driving an expansion of vegetation into arid regions¹⁹⁹, although the magnitude of this effect is uncertain²⁰⁰. Terrestrial stilling, the observed downward trend in surface wind speeds over land surfaces²⁰¹, could also affect dust emissions, with models suggesting a future reduction in Northern Hemisphere winds²⁰². However, changes in atmospheric circulation patterns thought to impact surface wind speeds over dust producing regions may be more important¹⁸⁰. Another consequence of planetary warming is an increase in precipitation variability²⁰³, and thus extreme rainfall events²⁰⁴, potentially increasing future sediment supply—and aeolian erosion—via alluvial and fluvial recharge²⁰⁵. Finally, future climate and land use changes could drive a decline in biological soil crusts that reduce dust emissions, which is a mechanism for increasing emissions that is not accounted for in current models²⁰⁶.

Model estimates of future changes in dust are sensitive to methodology²⁰⁷ and span the range of an increase in dust due to increasing aridity²⁰⁸ to a decrease due to CO₂ fertilization driving an expansion of vegetation into arid regions^{199,209}. Starting with CMIP5, simulations from some models included either prescribed dust emissions or fully interactive dust. However, both regional^{189,210-212} and global²¹³ analyses of these models found that the dust mean state had substantial biases, that CMIP5 models did not reproduce historical dust variability, and that modelled dust emissions were insufficiently sensitive to changes in surface conditions. An analysis of dust changes over land in RCP 8.5 simulations, for which CO₂ emissions continue unabated throughout the 21st century, showed no secular trends in global dust²¹³. An analysis of CMIP6 simulations demonstrated that many of these previously identified model deficiencies also exist in these newer climate model simulations (Fig. 5) and that the inter-model differences in dust are also growing relative to earlier CMIP efforts, suggesting that as model complexity increases so does model divergence in future projections of dust²¹⁴.

Given the inability of models to reproduce historical dust changes and the large spread in model projections of future dust change, it is not surprising that estimates of the change in dust radiative forcing per degree planetary warming, the so-called dust-climate feedback (units Wm⁻²K⁻¹), is similarly uncertain. An analysis of the output from 6 CMIP6 models that participated in an aerosol intercomparison project found that these models differed in the sign of the dust-climate feedback¹⁸⁷, with a multimodel mean feedback of 0.0026 ± 0.0048 Wm⁻²K⁻¹. Other research has speculated that a key driver of the model inconsistencies was the simulation of surface winds¹⁸⁷, which in turn may be related to the relatively coarse resolution of a typical climate model²¹⁵. These results from CMIP6 are consistent with earlier research that estimated a multimodel mean feedback for CMIP5 models that was not statistically different from zero²¹⁶. However, using a dust emission scheme that responded more realistically to changes in climate²¹⁷ enhanced the dust climate feedback due to changes in the dust direct radiative effect by an order of magnitude, yielding a range of -0.04 to +0.02 Wm⁻²K⁻¹. On a regional scale, the dust climate feedback close to source regions is likely an additional order of magnitude larger²¹⁶.

Given the lack of confidence in model projections of future changes in the dust burden, and the substantial uncertainties in dust direct and indirect radiative effects, there is a low degree of confidence in the ability of models to predict future changes in the dust radiative forcing.

Summary & Future Perspectives

We assessed the global mean effective radiative effect of dust in the modern climate at $R = -0.2 \pm 0.5 \text{ Wm}^{-2}$ (Fig. 3). Despite the considerable uncertainty in the sign and magnitude of R , which arises from the numerous uncertain and sometimes opposing mechanisms through which dust impacts climate, it is more likely that dust cools than that it warms global climate. We further found that global dust loading in the modern climate is $55 \pm 30\%$ higher than it was in pre-industrial times (Fig. 4), which has exerted a global mean effective radiative forcing of $\Delta F_{p \rightarrow m} = -0.07 \pm 0.18 \text{ Wm}^{-2}$. The historical increase in dust has thus likely somewhat counteracted greenhouse warming.

Current climate models fail to capture the historical increase in dust loading (Fig. 5) and thus inadequately account for dust radiative forcing, which could have caused biases in assessments of climate sensitivity and projections of future climate changes^{195,196}. Substantial additional research is thus needed both to better constrain R and $\Delta F_{p \rightarrow m}$ and to enable climate models to reproduce the historical increase in dust.

The dust direct radiative effect (DRE) contributes most to the uncertainty in R and $\Delta F_{p \rightarrow m}$ (Fig. 3). Future research should focus on reducing its uncertainty by better constraining dust optical properties through *in situ* and remote sensing observations. For instance, the information on soil mineralogy to be provided by NASA's 2022 Earth Surface Mineral Dust Source Investigation (EMIT) mission could help constrain dust optical properties²¹⁸. Additionally, models likely greatly underestimate the atmospheric concentration of super coarse dust^{49,55,56,124}, which warms by absorbing SW and LW radiation. This should be addressed by obtaining more measurements of emitted and transported dust that extend to the difficult-to-measure super coarse dust size range^{29,36,54,55}, and by developing improved parameterizations of super coarse dust emission²²⁰ and deposition and implementing those in climate models.

Another priority for future research should be better constraining the radiative effects of dust due to interactions with clouds, anthropogenic aerosols, and biogeochemistry, which together contribute the remaining uncertainty in R (Fig. 3). Because of a dearth of observational constraints, our assessment of these radiative effects was mostly based on modelling studies. However, models struggle to correctly account for interactions of dust with clouds and anthropogenic aerosols, in part because of the mismatch in scales between the small scales at which the relevant processes occur and the large scales of climate model grid boxes^{38,219,221}. As such, there is an urgent need for comprehensive *in situ* and satellite observations to constrain these interactions^{12,193}. For instance, more satellite and *in situ* observations of cirrus interactions with dust and other INPs^{98,99} could elucidate the relative importance of homogeneous and heterogeneous nucleation of ice crystals, which determines the sign of the radiative effect of dust interactions with cirrus (Fig. 2e)¹². Furthermore, dust radiative effects due to interactions with clouds could be better constrained with future model simulations at a sufficiently high (kilometre-scale²²¹) resolution to resolve the critical sub-grid scale turbulence and cloud processes that currently must be parameterized in models²¹⁹. Finally, constraining radiative effects due to dust interactions with biogeochemistry requires an improved characterization of dust composition and how this evolves during transport, as well as accurate knowledge of which land and ocean regions are nutrient limited¹⁵³.

We also recommend that the community conducts multi-model experiments to obtain more robust estimates of the various dust radiative effects and of R and $\Delta F_{p \rightarrow m}$. These experiments

should also investigate the uncertainty in radiative effects that result from model differences in dust optical properties, size distribution, model resolution, meteorology, the spatiotemporal distribution of dust emission fluxes, and parameterizations for dust deposition and dust interactions with clouds, radiation, atmospheric chemistry, the cryosphere, biogeochemistry, and other aerosols. Such multi-model experiments could be done in the context of the Aerosol Model Intercomparison project (AeroCom), which has previously performed multi-model experiments for anthropogenic aerosols^{25,222}.

Future research should also prioritize addressing the failure of models to reproduce the historical increase in dust (Fig. 5). Doing so requires an improved understanding of the factors driving changes in the atmospheric dust loading since pre-industrial times, including the relative roles of changes in land use, wind speed, soil properties, sediment supply, and vegetation cover^{180,223}. Additionally, new observations and modelling are needed to clarify the meteorological processes that generate the high wind speeds that produce dust, such as cold pool outflows from moist convection^{215,224,225}, and to improve the representation of those processes in climate models. Finally, more physically based dust emission schemes need to be developed and implemented into climate models. These schemes should explicitly account for dust emissions from high latitudes, which have an outsize effect on climate through interactions with clouds^{193,194}. Furthermore, dust emission schemes should avoid using empirical dust source functions as these do not respond to changes in climate; instead, emission schemes should use process understanding to account for the dependence of the spatiotemporal pattern and mineralogical composition of dust emissions on wind, soil properties, sediment supply, and vegetation coverage^{190,226,227}. A challenge will be to achieve this without making these schemes too sensitive to parameters such as soil moisture that non-linearly increase dust emissions^{1,2} and that have considerable variability in climate models¹⁹⁷. These fundamental improvements in dust emission schemes are also needed for meaningful predictions of future changes in dust and for more accurate predictions of dust impacts on regional climate.

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Competing interests

The authors declare no competing interests.

Author contributions

J.F.K. led the review, performed the dust reconstruction, wrote the Supplementary material, prepared the figures, and compiled the paper. T.S. and A.A.A. contributed the section on clouds and figures 2c-f. V.A.K. contributed the section on atmospheric chemistry and figure 2b. N.M.M. contributed the section on biogeochemistry and a draft of figure 2h. C.H. contributed the section on the cryosphere and figure 2g. A.T.E. contributed the section on future dust changes. D.M.L. contributed to the CMIP6 results in figure 5. All authors contributed to the manuscript preparation, discussion and writing.

Supplementary information

Supplementary information is available for this paper at <https://doi.org/10.1038/s415XX-XXX-XXXX-X>

Data availability

The dust reconstruction data shown in Figure 4 are available at [link to be added upon article acceptance].

Figure legends

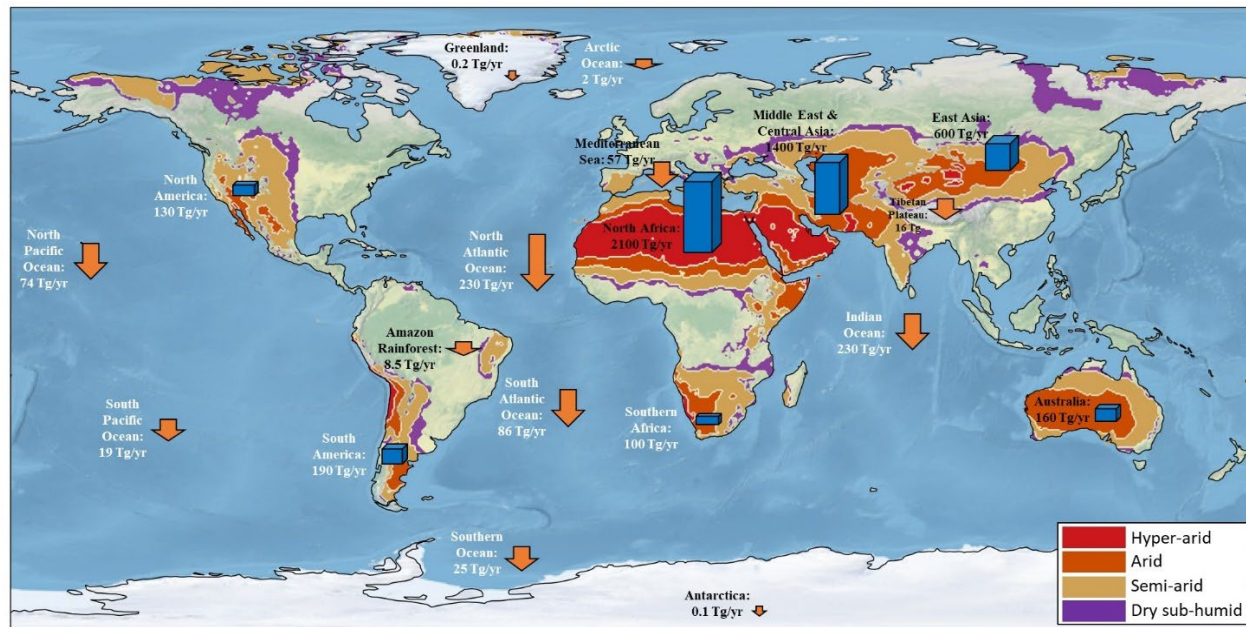
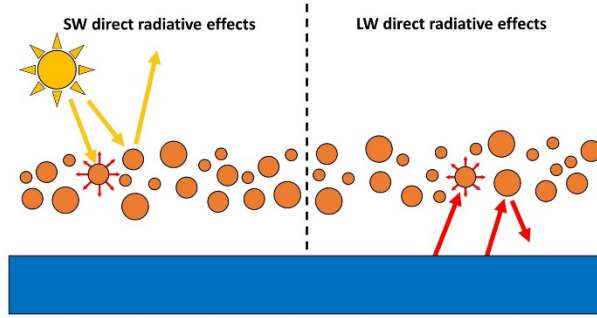


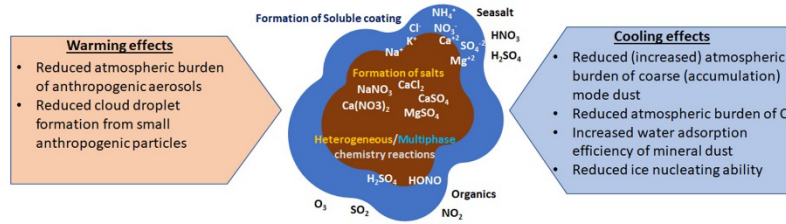
Figure 1. Main sources and sinks of dust in the global dust cycle. Emission fluxes (blue bars) from the world's main dust source regions and deposition fluxes (orange arrows) to regions where dust can impact surface albedo or biogeochemistry. Fluxes are for dust with geometric (volume-equivalent) diameter up to 20 μm and are based on constraints for 2004-2008⁴⁴; emissions from high latitude regions are not included. Shading represents dryland classification based on the aridity index: hyper-arid regions ($\text{AI} <$

0.05; red shading), arid regions ($0.05 < AI < 0.20$; orange shading), semi-arid regions ($0.20 < AI < 0.50$; light brown shading), and dry sub-humid regions ($0.50 < AI < 0.65$; green shading)²²⁸. Most dust is emitted from drylands in North Africa and Asia, which are collectively known as the “dust belt”²²⁹.

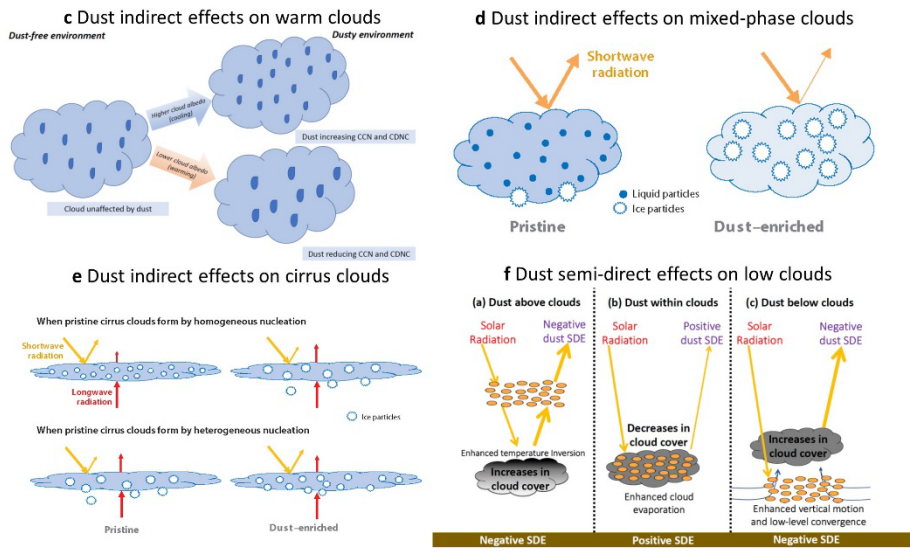
a Dust interactions with radiation



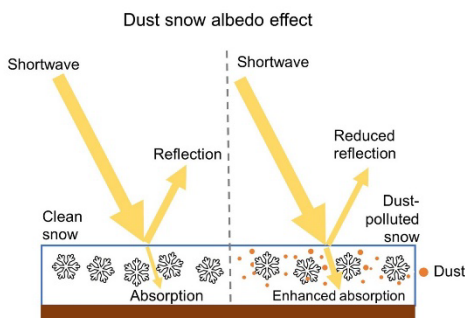
b Dust interactions with atmospheric chemistry



Dust interactions with clouds



g Dust interactions with the cryosphere



h Dust interactions with biogeochemistry

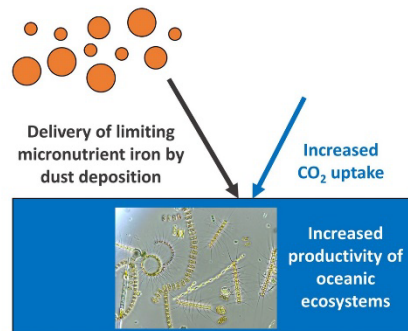


Figure 2. Mechanisms through which dust impacts climate. **a** | dust direct interactions with shortwave (SW) and longwave (LW) radiation. **b** | dust interactions with atmospheric chemistry and the induced perturbations to the radiative fluxes at the top-of-atmosphere exerting a warming (left) or cooling (right) effect on global climate. The brown core represents the freshly emitted insoluble dust particle and the surrounding blue area represents the acquired soluble coating through interactions with atmospheric chemistry. **c** | dust indirect effects on warm clouds occur by dust increasing cloud albedo through adding to CCN and increasing CDNC (upper branch) and by dust decreasing cloud albedo by reducing non-dust CCN through enhanced particle coagulation and adsorption of precursor gases and by dust giant CCN reducing in-cloud supersaturation (lower branch). **d** | dust indirect effects on mixed-phase clouds (MPCs), illustrated by MPC formation in pristine (left) and dust-enriched (right) environments. **e** | dust indirect effects on cirrus clouds, separated by the dominant ice crystal formation mechanism in the absence of dust. **f** | dust semi-direct effects on low clouds due to local heating generated by dust absorption, separated by location of dust relative to clouds. **g** | radiative effects of dust deposited on snow and ice, illustrated by snow reflectivity without (left) and with (right) dust deposited onto the snowpack. **h** | effect of dust on CO₂ concentrations due to interactions with ocean biogeochemistry. Yellow arrows represent SW radiation and red arrows represent LW radiation.

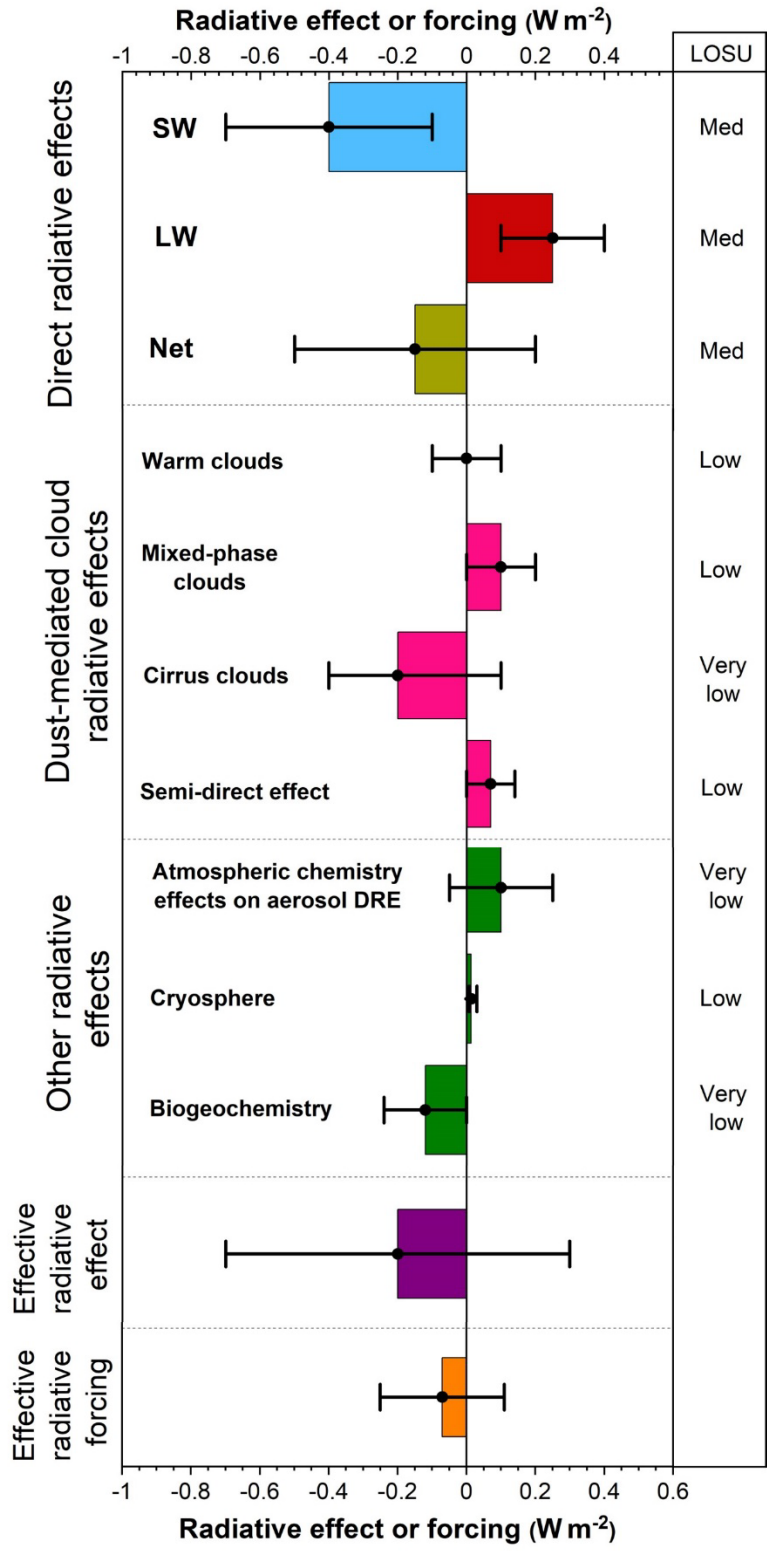


Figure 3. The global mean effective radiative effect and radiative forcing of dust at the top-of-atmosphere. Perturbations to Earth’s radiation budget by dust through direct radiative effects, dust-mediated cloud radiative effects, and various other radiative effects. The sum of all radiative effects equals the dust effective radiative effect R (Eq. 2) and the portion of that dust effective radiative effect

that is due to the increase in dust since pre-industrial times is the effective radiative forcing $\Delta F_{p \rightarrow m}$ (Eq. 4). Error bars denote the 90% confidence range. The column on the right denotes the level of scientific understanding (LOSU), or confidence in the assessment of each radiative effect, following past practice²⁷. The global mean dust effective radiative effect and radiative forcing of dust are uncertain in sign and magnitude, but are more likely to cool than to warm the climate.

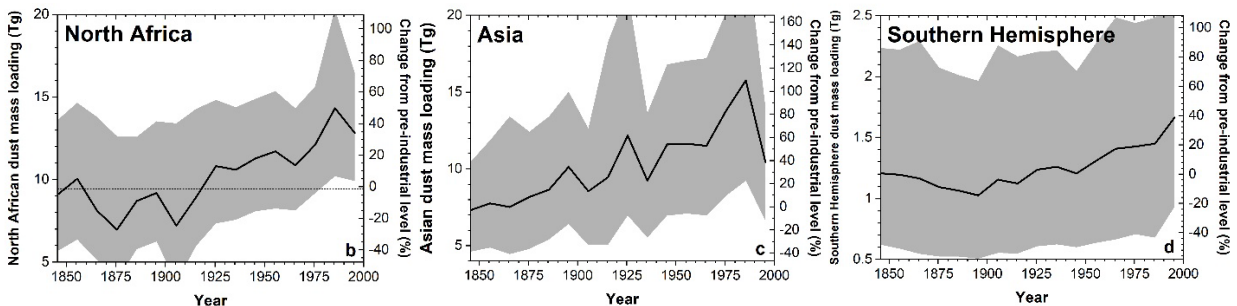
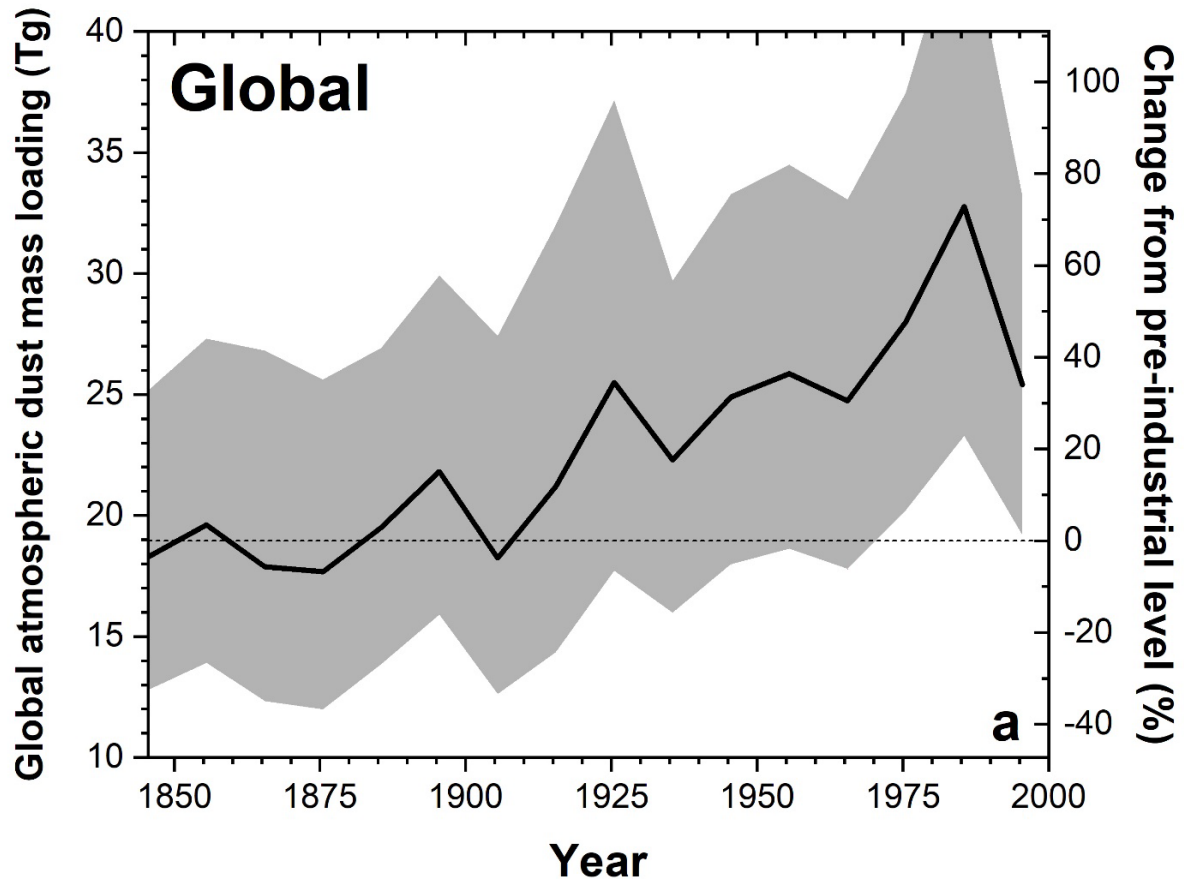


Figure 4. Atmospheric dust mass loading changes since pre-industrial times. a| Reconstructed globally integrated dust mass loading. **b|** as in a, but for loading contributed by dust from North Africa. **c|** as in a, but for loading contributed by dust from Asia. **d|** as in a, but for loading contributed by dust from the Southern Hemisphere. The solid line denotes the median dust loading estimate, the shading the 90% confidence range, and the dotted line the average pre-industrial (1841-1860) dust loading. Dust loadings were obtained by combining 19 records of dust deposition with constraints on the spatially resolved dust deposition fluxes produced by the world's main dust source regions^{8,44}; see Supplement for details. Dust has increased in all three regions, translating to a 55 ± 30 % rise in global dust mass loading in modern times (1981-2000) compared to pre-industrial.

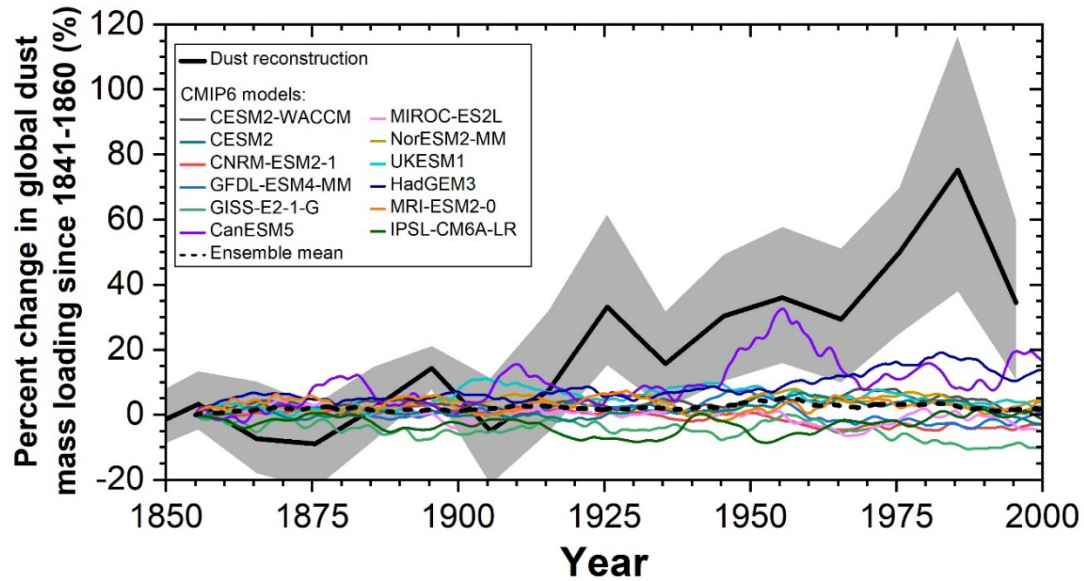


Figure 5. Climate model representations of historical changes in dust loading. Changes in global dust loading relative to the period 1841-1860 obtained from the dust reconstruction (solid black line) and simulated by 12 CMIP6 climate models with prognostic dust aerosol cycles²¹⁴ (thin colored lines). CMIP6 data are 10-year running means from historical runs¹⁸⁶. Grey shading denotes the 90% confidence interval for the dust reconstruction. All models and the ensemble mean (dashed black line) fail to reproduce the large historical increase in dust loading.

Box 1. Drivers of the historical increase in dust loading. The large historical increase in dust observed in deposition records and the reconstruction of dust mass loading (Fig. 4) can be either due to human land use changes or due to natural and anthropogenic changes in climate²³⁰.

The observational record shows that dust is highly sensitive to climate. Indeed, dust records in some regions show a variation of a factor of ~2-4 due to climate variability over the 20th century^{178,180,183} and dust has increased by a factor of ~2-4 in transitions between interglacial and glacial periods^{231,232}. As such, changes in aridity, vegetation cover, and wind speed due to natural climate variability could have driven (part of) the long-term increase in dust loading, as has been suggested for North Africa^{178,180,233}. In addition, anthropogenic changes to climate and atmospheric composition could also have affected dust loading, both by increasing aridity and by higher CO₂ concentrations fertilizing plants at desert margins²³⁴, with the net effect on desert extent and dust emissions still unclear²³⁵.

Human land use changes could also have increased dust emissions. The Industrial Revolution and the rise of industrialized agriculture have resulted in a dramatic increase in the area of land used by humans: the fraction of the ice-free land area used for agriculture has quadrupled from ~9% in 1850 to ~35% in 2000²³⁶. This large-scale conversion of wildlands to agricultural land has included many semi-arid and arid regions (Fig. 1), for which human land use changes can result in dramatic increases in dust emission²³⁷⁻²³⁹. Additionally, anthropogenic changes in water management that result in the drying of inland bodies of water might also have substantially increased dust emissions, such as has occurred for Owen's Lake in California in the early 20th century²⁴⁰ and more recently for the Aral Sea in Central Asia^{241,242}.

Modelling has been unable to determine whether the historical increase in dust, which models have been unable to reproduce²² (Fig. 5), has been primarily driven by climate or land use changes. Indeed, past research has diverged on the fraction of the global dust burden in the current climate emitted from anthropogenically disturbed sources, with results ranging from as little as 0% to as much as 50%^{199,207,230,243-245}. Similarly, modelling results on effects of changes in climate and CO₂ concentrations on dust loading also differ, with results varying between a decrease of -20% and an increase of +60% in dust loading^{199,230,235}.

Although large uncertainties thus remain in how climate and land use changes have contributed to the historical increase in dust loading (Fig. 4), two observational findings suggest that anthropogenic land use change has been a key driver of the long-term increase in dust loading (Fig. 4). First, the timing of increases in dust deposition in various deposition records appears to coincide with the rise of industrialized agriculture in source regions²³. And second, satellite observations suggest that ~25% of modern dust emissions originate from regions heavily impacted by human land use¹⁹¹. This finding implies that human land use changes have increased dust mass loading by ~33% since pre-industrial times, which accounts for the majority of the $55 \pm 30\%$ increase in dust mass loading since pre-industrial times (Fig. 4). Moreover, satellite observations indicate that the fraction of dust emitted from anthropogenically disturbed surfaces is substantially higher for Asian than for North African source regions, which is qualitatively consistent with the finding of a larger historical increase of Asian than of North African dust (Figs. 4b-d). Nonetheless, substantial additional work is needed to determine the exact causes of the historical increase in dust for each of the world's main dust source regions.

Supplementary Methods

We reconstructed atmospheric dust loading between the years 1841-2000 by combining a compilation of dust deposition records with constraints on the deposition flux produced by each of nine major source regions to each deposition site. These constraints on source-resolved deposition fluxes were obtained from the Dust Constraints from joint Observational-Modelling-Experimental analysis (DustCOMM) data set. This data set includes uncertainties on deposition fluxes that comparisons against independent data suggest are realistic^{1,2}.

Our approach is an extension of the methodology used in Mahowald et al.³, who also used a compilation of dust deposition records to reconstruct dust loading. We improved upon the approach in Mahowald et al.³ in a number of ways. First, we used almost double the number of dust deposition records (19 versus 11), including several recently published records that constrain dust from North Africa and East Asia⁴⁻⁹. These two source regions together account for ~60% of global dust loading² but most deposition records constraining these two regions were not available in Mahowald et al.³. Second, we used explicit constraints on the contributions of the different source regions to the deposition flux at each site, allowing a statistical approach that finds the most likely evolution of dust loading over time for each of nine major dust source regions^{1,2}. Finally, we used a bootstrap approach with resampling to propagate errors in both deposition flux records and in the DustCOMM source region-resolved deposition fluxes^{10,11}.

Below, we first describe the compilation of dust deposition records, after which we describe the methodology for reconstructing dust loading.

Compilation of dust deposition records

To construct our compilation of dust deposition records, we built on the compilation of 25 deposition records obtained by Hooper and Marx¹². Each of these records span at least 150 years of the pre-industrial to modern climate, ends after 1975, and has a time resolution of at least one data point per 10 years. Some of these records are based on measurements of dust particles in a sedimentary matrix (e.g., ice) and some are based on proxies for dust, such as the concentration of crustal elements like Mg and Ca. Measurements of the concentration of dust or a dust proxy are then converted to a dust deposition flux by calculating the mass accumulation rate of the sedimentary matrix using an age model that links the depth of the sediment to its age (e.g., Ref. 13). We use a subset of 13 of these deposition records, all of which Hooper and Marx¹² identified as being dominated by long-range transported dust, rather than by locally generated dust (also see discussion in Mahowald et al.³). Note that these 13 records exclude three dust records from Antarctic ice cores that were used in Hooper and Marx¹² but that were affected by local dust emissions (WAIS¹⁴ and Mount Erebus¹⁵) or that are unpublished data sets (Siple Dome). Because these 13 records included only one record dominated by North African dust⁶, we added four more published deposition records dominated by North African dust, namely records from Barbados⁴, the Swiss-Italian Alps⁵, Corsica⁷, and Cape Verde¹⁶. Additionally, we added deposition records from the northern Antarctic Peninsula¹⁷ and the Colorado mountains¹⁸, both of which were used by Mahowald et al.³. All deposition records used in this study are listed in Table S1.

Data from all deposition records were processed as follows. First, the data from each deposition record were linearly interpolated to obtain annual values in the period 1841-2000, after which all fluxes were normalized to the value of the median dust flux in the period 1991-2000. If the

record was missing data, either at the start or end of the 1841-2000 period, then these data points were filled in by taking the median of the 10 years most proximal to the missing data. We then took the median dust flux for each decade to reduce the substantial noise and year-to-year variability contained in some deposition records, which might be due to various factors unrelated to changes in dust emitted from large source regions (see Refs. ^{2,19,20} and further discussion below).

An exception on the data processing procedure above was made for the Cape Verde record of Evan & Mukhopadhyay¹⁶. This data set is unusual in that its accuracy was confirmed by correlation with satellite-observed aerosol optical depth data, but this data set also only spans the period 1955 – 2000. To fill in this record in the period 1841-1954, we used the mean of the normalized fluxes for the four other records that are dominated by North African dust, namely from Mauritania Canyon⁶, the Bahamas⁴, the Swiss-Italian Alps⁵, and Corsica⁷. We then normalized the mean of those four records such that the dust flux over the period 1955-1974 equaled that of the Evan & Mukhopadhyay¹⁶ record.

Reconstruction of dust loading since pre-industrial times

We used the compilation of dust deposition records to reconstruct the historical changes in dust loading generated by each of nine source regions. To do so, we used the DustCOMM data constraints on the relative contribution of each source region to each deposition core site in the modern climate (Fig. S1). That is, given the compilation of deposition flux timeseries, we obtained the evolution of normalized dust loading per source region that minimizes the disagreement between the deposition flux timeseries and the reconstructed dust loading. Specifically, we obtained the time evolution of dust loading generated by each source region by finding, for each decade, the values that minimizes the cost function¹

$$\chi(t)^2 = \sum_{j=1}^{N_{\text{dep}}} \left[\sum_{i=1}^{N_{\text{sr}}} \lambda_i(t) \hat{f}_{i,j}^{\text{cc}} - \tilde{\beta}_j(t) \right]^2, \quad (\text{S1})$$

where $N_{\text{dep}} = 19$ is the number of deposition flux timeseries in the compilation (Table S1) and $N_{\text{sr}} = 9$ is the number of source regions, which are western North Africa, eastern North Africa, the Southern Sahara and Sahel, the Middle East & Central Asia, East Asia, North America, Australia, South America, and Southern Africa (see Fig. S1 and Ref. ¹). Together, these 9 source regions account for the vast majority of natural dust emissions¹. However, emissions from high latitudes ($\sim 2\text{-}3\%$ of global emissions²¹) and from anthropogenic activities like industrial processes and vehicular traffic on dirt roads are not included²². Furthermore, $\hat{f}_{i,j}^{\text{cc}}$ is the DustCOMM constraint on source region i 's fractional contribution to deposition site j in the current climate, and $\tilde{\beta}_j(t)$ is the measured deposition flux at site j for decade t , normalized by the median dust deposition flux in the period 1991-2000. Finally, $\lambda_i(t)$ is the normalized globally integrated dust loading generated by source region i ,

$$\lambda_i(t) = \frac{L_i(t)}{L_i^{\text{cc}}}, \quad (\text{S2})$$

where $L_i(t)$ is the dust loading generated by source region i in decade t and L_i^{cc} is the DustCOMM constraint on the dust loading generated by source region i in the modern climate (see Table 1 in Ref.²). Because the DustCOMM constraints were obtained for the years 2004-

2008, here we need to assume that dust loading in the period 1991-2000 equals that in the period 2004-2008, which introduces additional error in our results. Furthermore, Eqs. (S1) and (S2) assume that a source region's deposition flux scales linearly with its globally integrated dust loading.

Note that the solution for Eq. (S1) is over-constrained, with 19 deposition cores and only nine source regions, which helps reduce errors on $\lambda_i(t)$. Nonetheless, because deposition fluxes from the Western North Africa, Eastern North Africa, and Sahel source regions tend to be correlated, and because of a dearth of deposition records dominated by dust from the Sahel and Eastern North Africa, we cannot separately constrain the Western North Africa ($i = 1$), Eastern North Africa ($i = 2$), and Sahel ($i = 3$) source regions. Therefore, we grouped these three source regions into a single source region similar to the North Africa source region used in Mahowald et al.³, meaning that we forced $\lambda_1(t) = \lambda_2(t) = \lambda_3(t)$ when solving (S1). Furthermore, we keep the Southern Africa source region ($i = 9$) constant (i.e., $\lambda_9(t) = 1$) because none of the deposition records are dominated by dust from this relatively minor source region ($\sim 1\text{-}2\%$ of global dust loading²), making its temporal evolution under-constrained.

Measurements of deposition fluxes and DustCOMM constraints on the relative contribution of each source region both carry substantial uncertainties^{2,23-25}. We propagated these uncertainties into $\lambda_i(y)$ and thus into the reconstruction of each source region's globally integrated loading using a bootstrap procedure^{10,11} in which we resample the deposition sites. Specifically, we repeat the following steps (also see Ref.¹) a large number of times ($\sim 1,000$):

1. We obtain a realization of the 3D global dust cycle, including the source region-resolved deposition flux to each deposition site, per the procedure detailed in Kok et al.¹ (see especially section 2.4 in Ref. ¹). From this realization, we obtain $\hat{f}_{i,j}^{cc}$ and L_i^{cc} .
2. We randomly draw, with replacement, 19 cores from our compilation of 19 cores. As such, for each iteration, some cores will be used not at all, some only once, and some more than once. This resampling procedure propagates uncertainty due to errors in the dust deposition records and yields $\tilde{\beta}_j(y)$.
3. We use equation S1 to obtain $\lambda_i(t)$, the globally integrated normalized dust loading for each source region. To prevent unrealistic solutions from propagating into our results, for instance for when a random drawing of deposition cores in step 2 results in one or more source regions being poorly constrained, we only use an iteration's result if the change in the dust loading due to each source region is less than a factor of five (i.e., $\min\left(\frac{L_i(t)}{L_i^{cc}}\right) > 0.2$ and $\max\left(\frac{L_i(t)}{L_i^{cc}}\right) < 5$).
4. We use equation S2 to calculate $L_i(t)$ and the time evolution of the globally integrated dust loading as $L_{\text{glb}}(t) = \sum_{i=1}^{N_{\text{sr}}} L_i(t)$.

This approach yields a large number of realizations that represent the probability distribution of $L_i(t)$ and $L_{\text{glb}}(t)$. We report the median and one standard error range of these results in Figures 4 and 5 in the main text. Additionally, we show the resulting comparisons against each of the 19 deposition cores in Fig. S2 and the reconstructed loading for each source region in Fig. S3. We also compare the dust reconstruction against long-term dust surface concentration measurements made since the 1970s at Barbados and Miami in Fig. S4. The dust reconstruction is in excellent

agreement with the Barbados record (Fig. S4a) and captures the trend in the Miami record, although it underestimates the magnitude of the peak dust loading in the 1980s there (Fig. S4b).

As discussed in more detail in Mahowald et al.³, using deposition cores to reconstruct global dust loading requires a number of critical assumptions that could introduce large errors in the results, not all of which are quantified by our error propagation procedure. First, our methodology assumes that the spatial pattern of the dust deposition flux per source region has not changed since pre-industrial times. This is likely inaccurate, as the historical increase in dust loading suggests that new source regions have been activated over time, either by climate or human land use changes. However, the effect of this limitation on our results might be limited because variability of simulated dust emissions in even the modern climate is very large (e.g., Figure 2 in Ref.²⁶) and DustCOMM draws on simulations from six different global models. Because satellite analyses indicate that ~25% of present day emissions are generated from surfaces with substantial human land use²⁷ it thus appears unlikely that the difference in the spatial pattern between dust emissions in the modern and pre-industrial period will be large compared to the substantial diversity in the spatial pattern of emissions that is already accounted for by DustCOMM. Second, measurements and simulations of dust deposition fluxes are subject to large experimental and representation errors^{23,26,28}, which could cause further errors in the reconstruction of global dust loading. Our approach of using relative changes in dust deposition over time at each deposition site addresses this concern somewhat as errors in trends of dust deposition fluxes are usually substantially lower than in absolute values⁴. Third, trends in dust fluxes at deposition sites might not reflect trends in dust emissions but instead in dust transport pathways¹⁹ and in the precipitation that drives wet deposition²⁹. However, ice core records indicate that rates of sea-salt aerosol deposition have remained approximately constant since pre-industrial times, which indicates that wet deposition rates have not changed substantially³⁰. The effect of possible changes in dust transport pathways are difficult to quantify and it is possible that a trend in deposition data at a given site is due to long-term changes in dust transport pathways rather than changes in emissions. This concern about changes in transport causing spurious trends in deposition data is ameliorated somewhat by our use of deposition data from 19 deposition sites (Table 1). That is, because the bias caused by changes in transport pathways will differ for each site it constitutes a random error, the effect of which is thus reduced with more data. Also note that this random error is inherently propagated into our results by the bootstrap procedure with resampling. Fourth, trends in deposition fluxes will not be reflective of trends in major source regions if dust from a deposition site is dominated by local dust or by dust from a non-representative subsection of a major source region²⁹. Errors in the classification of which deposition records are reflective of long-range transported dust from major source regions^{3,12} could thus cause errors in the dust reconstruction. Also note that more deposition records could facilitate a substantially more accurate dust reconstruction because it could allow the use of smaller source regions in solving Eq. (S1); for instance, more deposition records could allow the temporal evolution of the western North Africa, eastern North Africa, and Southern Sahara and Sahel to be solved without combining these source regions. Considering these potentially substantial sources of uncertainty, the error estimates on our reconstructed global dust loading should be considered a lower bound.

Supplementary Figures and Tables

Table S1. Summary of compilation of dust deposition records.

Dust deposition site	Region	Type of record	Latitude	Longitude	Associated source(s)*	Record length	Approximate age resolution	Reference
NEGIS	Greenland	Ice core	75.62°	-35.96°	East Asia, North Africa, North America	1607 – 2005	1 year	Vallelonga et al. (2014) ³¹
GISP2	Greenland	Ice core	72.6°	-38.5°	East Asia, North Africa, North America	665 – 1990	2-3 years (irregular)	Zielinski & Mershon (1997) ³²
20D	Greenland	Ice core	65.02°	-44.87°	East Asia, North Africa, North America	1767 – 1984	1 year	Kang et al. (2003) ³³
Mount Logan	Yukon, Canada	Ice core	60.58°	-140.5°	East Asia	1000 – 1998	1 year	Osterberg et al. (2008) ⁹
Motianling	Great Hinggan Range, China	Peat core	47.35°	120.65°	East Asia	1860 – 2007	3-30 years (irregular)	Bao et al. (2012) ⁸
Point d'Escuminac	New Brunswick, Canada	Peat core	46.92°	-65.5°	North America	1550 – 1996	4-25 years (irregular)	Kylander et al. (2009) ³⁴
San Juan Lakes	Colorado, USA	Lake sediment core	38°	-108°	North America	1855 - 2005	30 years	Neff et al. (2008) ¹⁸
Dasuopo	Mount Xixiabangma, Himalaya	Ice core	28.38°	85.72°	Middle East / Central Asia	1450 – 1996	1 year	Thompson et al. (2000) ³⁵
Mount Everest #1	East Rongbuk Glacier, Himalaya	Ice core	28.03°	86.96°	Middle East / Central Asia	1690 - 2002	1 year	Kaspari et al. (2009) ³⁶
Mount Everest #2	East Rongbuk Glacier, Himalaya	Ice core	27.98°	86.92°	Middle East / Central Asia	1843 – 1997	1 year	Kang et al. (2003) ³³
Mauritania Canyon	Atlantic Ocean	Ocean sediment core	16.83°	-16.73°	North Africa	1180 BC – 2005	1-2 years	Mulitza et al. (2010) ⁶
Sal Island	Cape Verde	Coral core	16.76°	-22.89°	North Africa	1955 – 2008	1 year	Evan & Mukhopadhyay (2010) ¹⁶
Andros Island	Bahamas	Coral core	25°	-78°	North Africa	1065 – 2011	10-30 years (irregular)	Hayes et al. (2017) ⁴
Colle Gnifetti	Italian Alps	Ice core	45.94°	7.88°	North Africa	1780 – 2006	1 year	Clifford et al. (2019) ⁵
Lake Bastani†	Corsica	Lake sediment core	42.07°	9.13°	North Africa	1220 BC - 2013	3-4 years	Sabatier et al. (2020) ⁷ ; Pierre Sabatier, personal communication (2022)
Snowy Mountains	NSW, Australia	Peat core	-36.46°	148.3°	Australia	500 – 2005	2 – 10 years (irregular)	Marx et al. (2014) ³⁷
Law Dome	Budd Coast, East Antarctica	Ice core	-66.77°	112.8°	South America, Australia	1480 – 1996	1 year	Souney et al. (2002) ³⁸
Siple Station	Ellsworth land, West Antarctica	Ice core	-75.92°	-84.25°	South America, Australia	1417 – 1983	1 year	Mosley-Thompson et al. (1990) ³⁹

*Associated sources are based on Fig. S1.

†The data of Sabatier et al.⁷ were further corrected for sediment matrix effects through centered-log-ratio (CLR) transformation (Sabatier, personal communication, 2022).

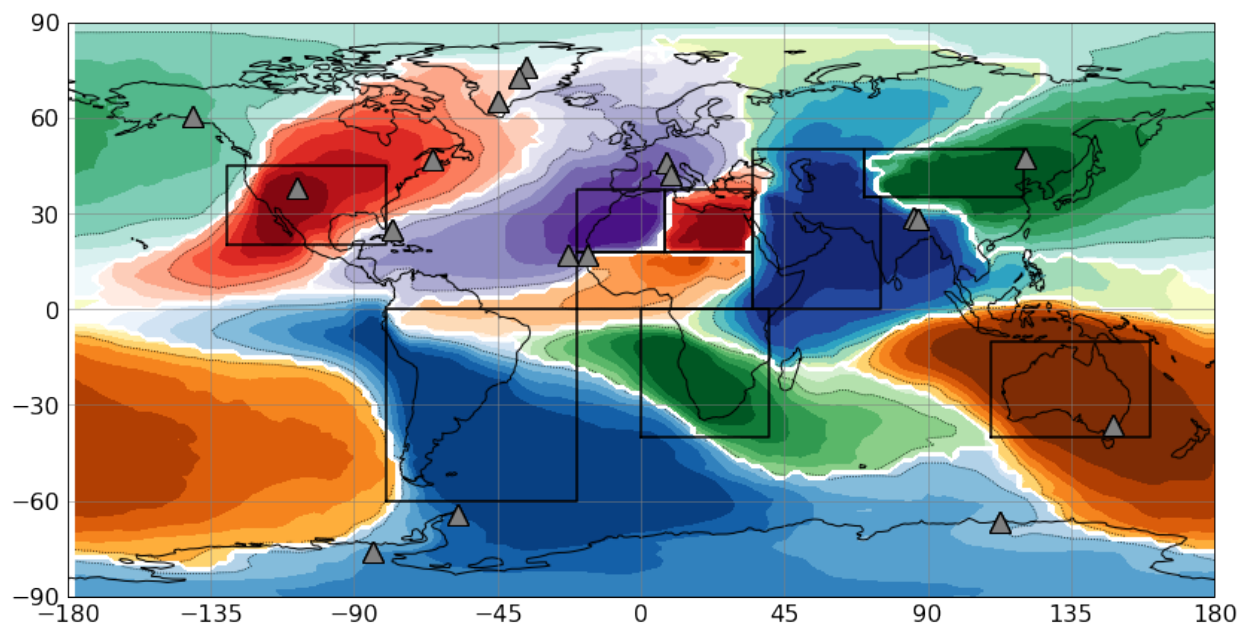


Figure S1. Map of deposition record sites and dominant dust source regions. Different colors represent the fraction of the dust deposition flux supplied by different dominant source regions, with shading decreasing in increments of 0.1 from a maximum of 1. The black boxes denote the nine major source regions (western North Africa, eastern North Africa, the Southern Sahara and Sahel, the Middle East & Central Asia, East Asia, North America, Australia, South America, Southern Africa). Gray triangles denote the locations of the 19 dust deposition records used to reconstruct dust loading since pre-industrial times. After Ref. ².

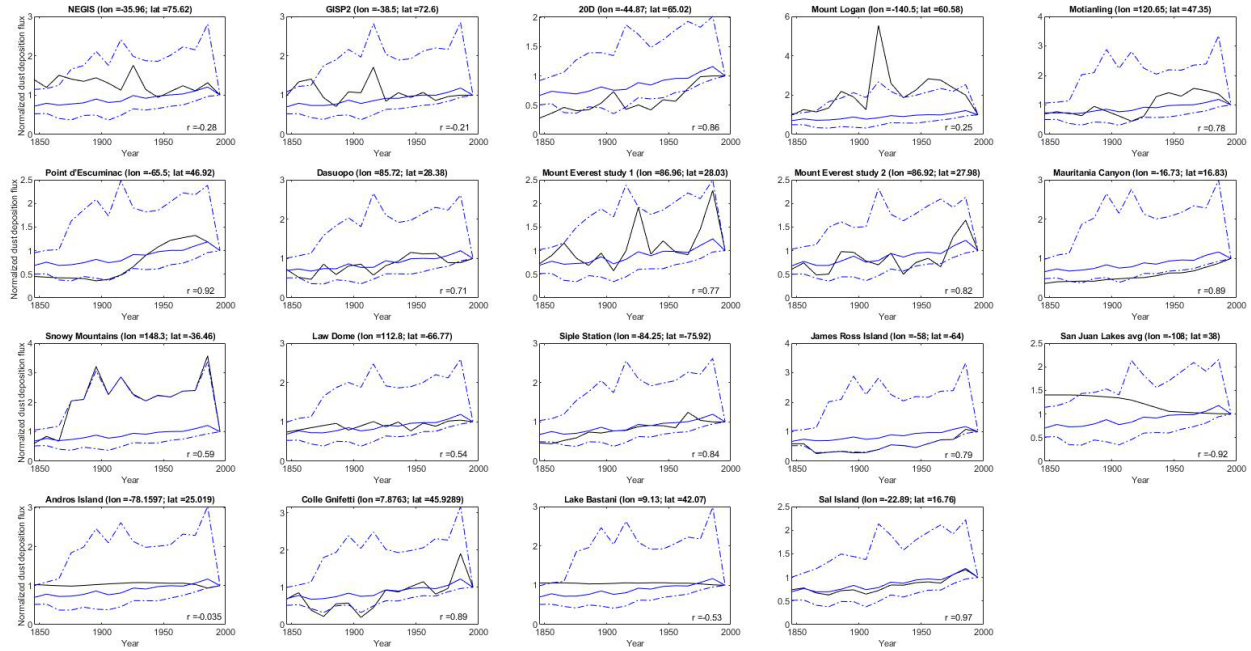


Figure S2. Comparisons between dust deposition records and reconstructed dust deposition fluxes. Shown are results for the 19 deposition records in our compilation (Table S1). The solid black line represents the measured normalized deposition fluxes (or deposition flux proxy) after data processing (see text), the solid blue line the median reconstructed normalized deposition flux, and the dash-dotted blue lines the one standard error interval on the reconstructed normalized deposition flux. Also listed is the correlation coefficient between the measured and median reconstructed normalized deposition fluxes.

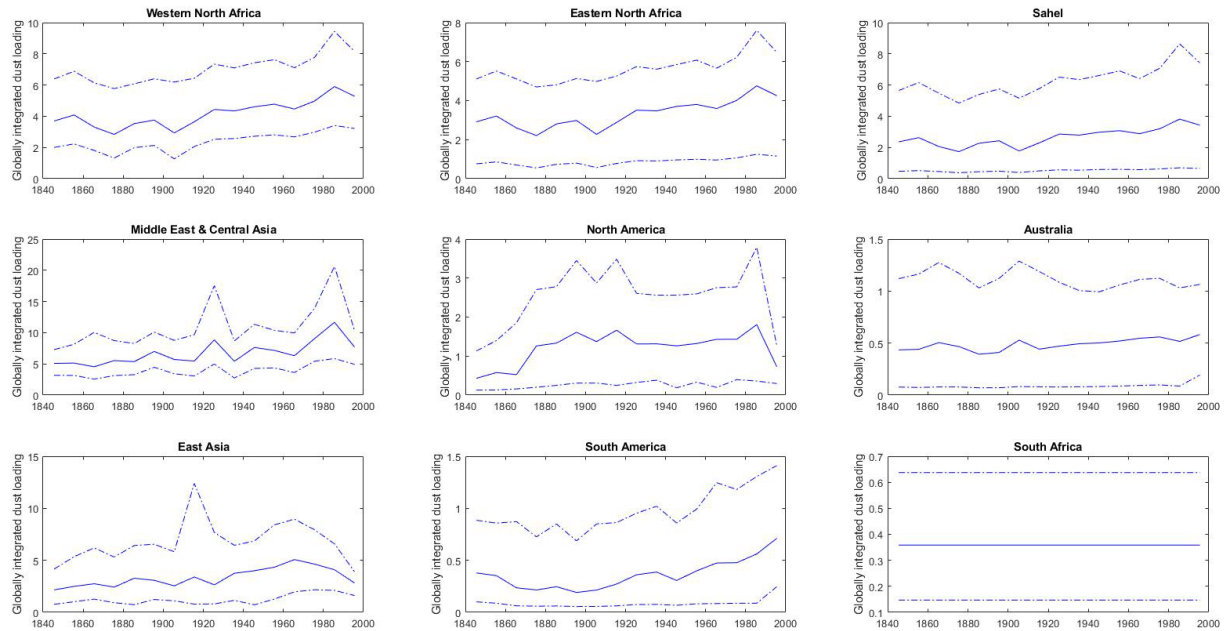


Figure S3. Reconstructed globally integrated dust loading for each source region. The solid line represents the median estimate and the dash-dotted lines the one standard error interval.

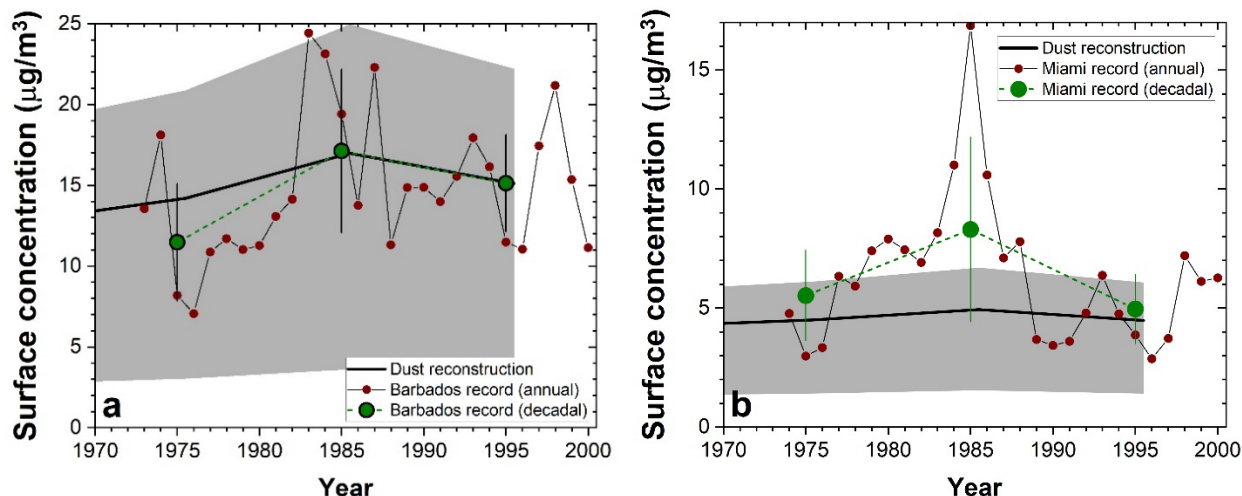


Figure S4. Comparison of dust reconstruction against long-term dust surface concentration measurements. **a**| Comparison against measurements at Barbados. **b**| Comparison against measurements at Miami. Small brown circles denote the measured annual average surface concentration, which was calculated from previously reported monthly average values⁴⁰. Large green circle denote decadal averages of the annual values for a more direct comparison against the dust reconstruction. Error bars on the decadal values denote the standard deviation of the annual values. The solid black line denotes the median estimate of the reconstructed dust surface concentration and the shading denotes the 90% confidence range.

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