

Wavelet-based wavenumber spectral estimate of eddy kinetic energy: Idealized quasi-geostrophic flow

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Key Points:

- A wavelet-based spectral method to estimate eddy variability is described.
- Wavenumber spectra of eddies are estimated for a doubly-periodic quasi-geostrophic flow.
- The wavelet and Fourier approach agree well in their estimates of spectra and spectral flux.

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17 **Abstract**

18 A wavelet-based method is re-introduced in an oceanographic and spectral context to
19 estimate wavenumber spectrum and spectral flux of kinetic energy and enstrophy. We
20 apply this to a numerical simulation of idealized, doubly-periodic quasi-geostrophic flows,
21 i.e. the flow is constrained by the Coriolis force and vertical stratification. The double
22 periodicity allows for a straightforward Fourier analysis as the baseline method. Our wavelet
23 spectra agree well with the canonical Fourier approach but with the additional strengths
24 of negating the necessity for the data to be periodic and being able to extract local anisotropies
25 in the flow. Caution is warranted, however, when computing higher-order quantities, such
26 as spectral flux.

27 **Plain language summary**

28 Chaotic flows such as the ocean currents, atmospheric winds and turbulence in gen-
29 eral are fundamentally impossible to analytically predict, namely, to formulate a math-
30 ematical general solution. Nevertheless, the interest in describing such chaotic flows can
31 be found in examples as old as Leonardo da Vinci’s sketch of turbulence. While we can-
32 not obtain an analytical description of turbulence, we can extract statistical information
33 from turbulent flows and a common descriptor has been the wavenumber spectrum. The
34 spectrum reveals at each spatial scale, the level of variability the flow carries. Here, we
35 re-introduce an alternative method in estimating the spectrum based on wavelet func-
36 tions.

37 **1 Introduction**

38 Fundamental to the goal of properly modelling climate system dynamics is under-
39 standing and quantifying how energy is both distributed, and ultimately transferred, across
40 an extremely broad range of dynamically active space and time scales. In the atmospheric
41 and ocean context, the most common means of quantifying the scale-dependent energy
42 content of a chaotic, turbulent flow field is the energy spectrum (e.g. Taylor, 1938; Kol-
43 mogorov, 1941; Charney, 1971; Nastrom & Gage, 1983; Yaglom, 2004) given by the Fourier
44 transform of two-point (spatial or temporal) velocity correlations.

45 We will focus in this paper on spatial correlations since the behavior of wavenum-
46 ber spectra are described by ‘inertial range’ theories predicting spectral slopes and cas-
47 cades (Vallis, 2006). The standard Fourier approach has had great success in providing

48 us with spectral estimates of energy partition (e.g. Stammer, 1997; Xu & Fu, 2011, 2012;
 49 Uchida et al., 2017; Ajayi et al., 2021) and its straightforward mathematical formula-
 50 tion facilitates the spectral interpretation in the original context, namely statistically ho-
 51 mogeneous flows where Fourier decompositions are natural. Issues persist, however, in
 52 geophysical flows which are statistically inhomogeneous, anisotropic and non-stationary
 53 (Uchida, Jamet, et al., 2021). The assumption of homogeneity lies on the fact that a Fourier
 54 transform is a global operator over the entire space-time domain of interest. In other words,
 55 the Fourier description of the field conflates different regimes of an inhomogeneous flow.
 56 A notable example is in the separated Gulf Stream region where the energetics have been
 57 argued to be distinct from the gyre interior (Jamet et al., 2021).

58 With a growing acknowledgement of the shortcomings of the Fourier approach, there
 59 has been a recent effort in the geophysical sciences to re-examine the cross-scale ener-
 60 getics. Notable examples are: i) Aluie et al. (2018); Sadek and Aluie (2018); Schubert
 61 et al. (2020); Storer et al. (2022); Srinivasan et al. (2022) and Contreras et al. (2022) where
 62 they implement a spatial filter, ii) Lindborg (2015); Balwada et al. (2016, 2022); LaCasce
 63 (2016); Poje et al. (2017) and Pearson et al. (2020) where they use structure functions,
 64 iii) Jamet et al. (2020) where they employ the Green’s function, and iv) Uchida, Jamet,
 65 et al. (2021) where they use Empirical Orthogonal Functions all with the goal of exam-
 66 ining the KE spectra and cross-scale transfer in the wavenumber domain. The overall
 67 consistent picture is that at scales about $O(100\text{ km})$ where the oceanic motions are con-
 68 strained by the Earth’s rotation and vertical stratification, KE cascades upscale while
 69 KE on the scale of $O(10\text{ km})$ tend to cascade downscale due to a loss of balance with the
 70 two constraining forces. While all these approaches, including the Fourier method, can
 71 capture within limits the spatial anisotropy when examined on a two-dimensional (2D)
 72 wavenumber plane, they lose this information when reduced to one-dimensional (1D) spec-
 73 tral quantities.

74 Here, we use a wavelet-based technique which yields localized pseudo-Fourier 1D
 75 wavenumber spectra capable of capturing the local anisotropies in the flow (Daubechies,
 76 1992; Perrier et al., 1995). Wavelets emerged in the 1980s as a way to analyze time and
 77 space series in more local manner than was possible using Fourier techniques (e.g. Vasi-
 78 lyev & Paolucci, 1997; Doglioli et al., 2007; Alvera-Azcárate et al., 2007; Thomson & Emery,
 79 2014), although strong parallels and connections are to be found between the two meth-
 80 ods (Katul & Parlange, 1995; Torrence & Compo, 1998). We will argue the localized na-

81 ture of wavelets allows us to capture the inhomogeneity and anisotropy in the flow (Farge
 82 et al., 1992; Horbury et al., 2008). We then apply the wavelet approach to estimate the
 83 horizontal kinetic energy (KE) and enstrophy spectral flux from a doubly-periodic quasi-
 84 geostrophic (QG) flow, and to a flow subdomain where periodicity no longer applies. The
 85 comparisons illustrate some of the advantages of the wavelet approach.

86 The paper is organized as follows: We describe the QG model and provide an overview
 87 of the wavelet method in Section 2. Results are given in Section 3 where we compare our
 88 wavelet spectra to the canonical Fourier spectra. Conclusions are given in Section 4.

89 2 Theory and technique

90 We describe the configuration of our quasi-geostrophic (QG) model and provide
 91 an overview of the wavelet method.

92 2.1 Description of the quasi-geostrophic simulation

93 We consider a stochastically forced two-layer QG flow in a doubly periodic f plane
 94 domain (i.e. $\beta = f_y = 0$) under rigid-lid and flat bottom conditions. Solutions to the
 95 QG potential vorticity (PV) equation

$$96 \quad q_{j_t} + J(\psi_j, q_j) = -r_b \nabla^2 \psi_j \delta_{j,2} + \mathcal{Q} \quad (1)$$

98 are computed using the pseudo-spectral `pyqg` model (Abernathey et al., 2022), where
 99 $\delta_{i,j}$ is the usual Kronecker delta function and layer numbers are denoted $j = 1, 2$. The
 100 linear bottom drag coefficient is $r_b = 5.787 \times 10^{-7} \text{ s}^{-1}$. The PV in each layer are

$$101 \quad q_1 = \nabla^2 \psi_1 + F_1(\psi_2 - \psi_1), \quad (2)$$

$$102 \quad q_2 = \nabla^2 \psi_2 + F_2(\psi_1 - \psi_2). \quad (3)$$

104 The vortex stretching coefficients are $F_1 = \frac{(2\pi/R_d)^2}{1+\delta}$, $F_2 = \delta F_1$ where the internal Rossby
 105 deformation radius was prescribed as $R_d = 100 \text{ km}$. Each layer thickness is $(H_1, H_2) =$
 106 $(500, 2000) \text{ m}$ respectively, giving $\delta = H_1/H_2 = 0.25$. The square domain size is $L_0 =$
 107 1000 km with the spatial resolution of $\sim 2 \text{ km}$ (512×512 grid points). In order to pre-
 108 vent the system from equilibrating to the well-known single pair of positive and nega-
 109 tive vortices (Vallis, 2006), a vertically uniform forcing was introduced as

$$110 \quad \mathcal{Q} = A_q w(t, y, x), \quad (4)$$

111 where $A_q = 10^{-15} \text{ s}^{-2}$ is the amplitude and $w(t, y, x)$ is white noise in space-time with
 112 zero mean and $O(1)$ amplitude per layer (Fig. 1). The quantity $w(t, y, x)$ was computed
 113 by taking the inverse Fourier transform of a ring in wavenumber space

$$\hat{w}(t, k^y, k^x) = \begin{cases} a(t, k^y, k^x) + ib(t, k^y, k^x), & \text{if } (R_d + \delta_R)^{-1} < \sqrt{k^x^2 + k^y^2} < (R_d - \delta_R)^{-1}, \\ 0, & \text{otherwise} \end{cases}, \quad (5)$$

114
 115 where k^x and k^y are the zonal and meridional wavenumbers respectively, a and b are Gaus-
 116 sian random variables in the horizontal wavenumber space with zero mean and standard
 117 deviation of unity, and $\delta_R = 5 \text{ km}$. After taking the inverse Fourier transform, the hor-
 118 izontal spatial mean is removed and then divided by the mean of the absolute values in
 119 the horizontal dimension in order to have the magnitude on the order of unity. In other
 120 words, the model is stochastically forced at scales about the Rossby radius uncorrelated
 121 in time. No background PV was prescribed. The model was spun up for 10 years from
 122 a state of rest, at which point area averaged energy had equilibrated (not shown), and
 123 then run for another 20 years with outputs saved every 10 days as instantaneous snap-
 124 shots. The timeseries of the kinetic energy (KE; $K = (\mathbf{u} \cdot \mathbf{u})/2$) and potential energy
 125 (PE; $2(\psi_1 - \psi_2)^2/R_d^2$) for the 20 years of output are given in Fig. 2, which mirror each
 other and roughly show an equipartition.

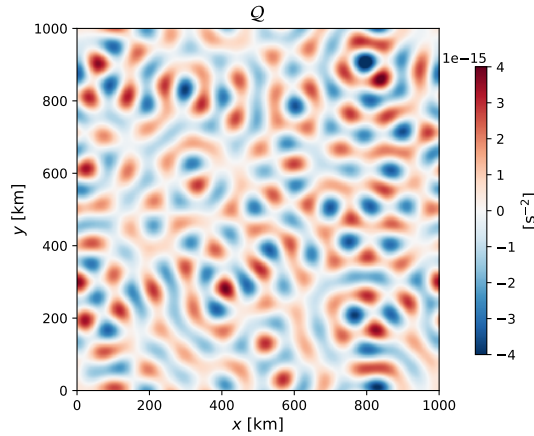


Figure 1. Example of how the vertically uniform stochastic forcing \mathcal{Q} looks like for an arbitrary time step.

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127 In this simple configuration, it is expected the flow will be both homogeneous and
 128 isotropic in the horizontal dimensions. Further, classical theory predicts the existence

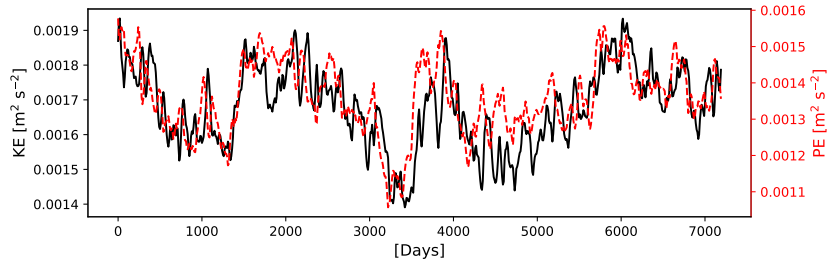


Figure 2. Vertically and domain averaged kinetic energy (KE; black solid) and (available) potential energy (PE; red dashed).

129 of an inverse cascade of KE and hence a $-5/3$ power law at scales larger than the forc-
 130 ing scale, and a forward cascade of enstrophy and hence -3 power law at smaller scales
 131 that are above the viscous dissipation scale (Vallis, 2006). In this sense, we ‘know’ what
 132 the answer should be and can use the results to test the efficacy of the wavelet trans-
 133 form. The double periodicity also allows for a straightforward comparison between the
 134 wavelet and Fourier approach as no windowing of the data is necessary in applying the
 135 transforms. We exhibit the top- and bottom-layer PV at the last time step of the model’s
 10th year in Fig. 3.

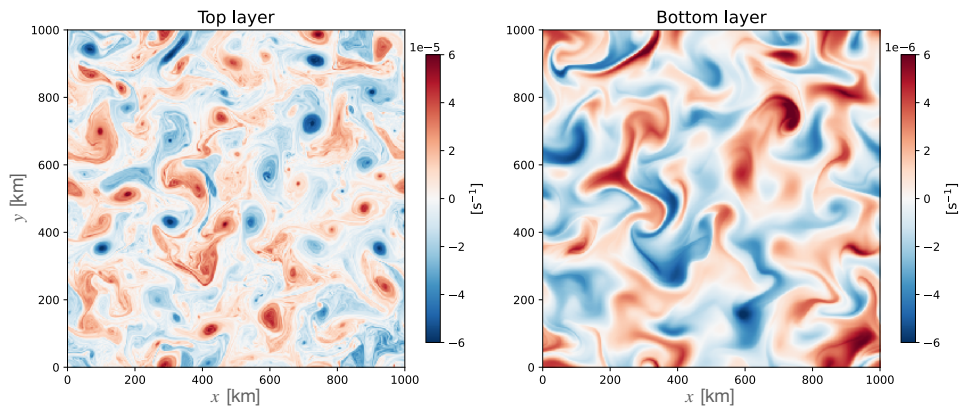


Figure 3. The PV at the last time step of the 10th simulated year in the top and bottom layer. Note the order of magnitude difference in the two panels.

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2.2 Spectral Considerations

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For the reasons outlined in the introduction, we depart from the classical Fourier approach to compute wavenumber spectra, but do note the utility of that wavenumber spectrum emerges largely from Parseval's equality

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$$\int_{\mathbf{x}} K(\mathbf{x}) d\mathbf{x} = \int_{\mathbf{k}} \widehat{E}_K(\mathbf{k}) d\mathbf{k}, \quad (6)$$

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where $\mathbf{x} = (x, y)$, $\mathbf{k} = (k^x, k^y)$ (e.g. Scott & Wang, 2005; Capet et al., 2008; Uchida et al., 2017). The Fourier energy spectrum is given by $2\widehat{E}_K(\mathbf{k}) = \hat{\mathbf{u}}^* \cdot \hat{\mathbf{u}}$ where the Fourier transform of the velocity is denoted by the hat ($\hat{\mathbf{u}}$) and the superscript * denotes the complex conjugate. This equivalence of the area integrated KE to the wavenumber integrated Fourier spectrum motivates the latter's interpretation as the KE density in the wavenumber domain.

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We base our spectral analysis on wavelet decompositions, rather than Fourier transforms, as the space-time locality of wavelets does not require the data to be periodic. Given a function dependent on two spatial dimensions, $f(\mathbf{x})$, its continuous wavelet transform is given by (Daubechies, 1992; Torrence & Compo, 1998)

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$$\tilde{f}(s, \phi, \gamma) = \int_{\Omega} f(\mathbf{x}) \frac{1}{s} \xi^* \left(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x} - \gamma}{s} \right) \right) d\mathbf{x}, \quad (7)$$

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154

where the integration is taken over the whole domain of interest Ω and \mathbf{R}^{-1} is the inverse of the rotation matrix

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$$\mathbf{R}^{-1} = \begin{pmatrix} \cos(\phi) & \sin(\phi) \\ -\sin(\phi) & \cos(\phi) \end{pmatrix}, \quad (8)$$

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for rotation through an angle ϕ relative to the x axis. The quantity s is referred to as the 'scale', $\gamma \in \mathbf{R}^2$ are the two-dimensional coordinates of interest, $\xi(\mathbf{x})$ is the so-called 'mother' wavelet and $\xi(\mathbf{R}^{-1} \cdot (\mathbf{x} - \gamma) / s)$ in (7) are the daughter wavelets. The quantities \tilde{f} are called the wavelet coefficients. Note that the field of wavelet coefficients is a filtered version of the original data.

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Subject to the 'admissibility condition' $C_{\Xi} < \infty$, the original function f can be reconstructed from the wavelet coefficients (Daubechies, 1992; Torrence & Compo, 1998)

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$$f(\mathbf{x}) = \frac{1}{C_{\Xi}} \int_{\gamma} \int_s \int_{\phi} \tilde{f}(s, \phi, \gamma) \frac{1}{s^4} \xi \left(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x} - \gamma}{s} \right) \right) d\phi ds d\gamma. \quad (9)$$

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If $\widehat{\Xi}(\mathbf{k})$ is the Fourier transform of the mother wavelet, then

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$$C_{\Xi} = \int_{\mathbf{k}} \frac{\widehat{\Xi}^* \widehat{\Xi}}{\mathbf{k} \cdot \mathbf{k}} d\mathbf{k}. \quad (10)$$

166 The so-called ‘admissibility condition’ implies that the mother wavelet defines a well-behaved
 167 class of wavelet transforms. Many functions satisfy (10) provided they have zero mean

$$168 \int_{\mathbf{x}} \xi(\mathbf{x}) d\mathbf{x} = 0. \quad (11)$$

169 For current purposes, we will employ the so-called Morlet wavelet (Morlet et al., 1982;
 170 Gabor, 1946), i.e.

$$171 \xi(\mathbf{x}) = (e^{-2\pi i \mathbf{k}_0 \cdot \mathbf{x}} - c_0) e^{-\frac{\mathbf{x} \cdot \mathbf{x}}{2x_0^2}}, \quad (12)$$

172 where c_0 is a constant included to ensure that (11) is met. The central wavenumber \mathbf{k}_0
 173 is taken to be $\mathbf{k}_0 = (k_0, 0)$ and the quantity x_0 is a reference length scale, here taken
 174 to be the Rossby radius ($x_0 = 100$ km), viz. the central length scale of the mother wavelet.
 175 We will choose $k_0 = 1/x_0$, in which case the constant c_0 is quite small and generally
 176 ignored (i.e. $c_0 = 0$), a convention adopted in this paper. Plots of (12) are found in Fig. 4.
 177 Note that the Morlet mother wavelet consists of a wave of wavelength x_0 inside a Gaus-
 178 sian envelope of decay scale $\sqrt{2}x_0$. Thus for $s = 1$ and $\phi = 0$, the wavelet coefficient
 179 produced by this transformation comments on the presence of the wavenumber $\mathbf{k}_0 =$
 180 $(k_0, 0)$ at location γ in the original data. Increasing the rotation angle ϕ and filtering
 181 returns information about the presence of the same wavelength at angle $-\phi$. Finally al-
 182 lowing s to vary modifies the filter so that the primary wavenumber of the filter is $k =$
 183 $1/(sx_0)$. The Morlet wavelet coefficient can thus be thought of as a ‘local’ Fourier trans-
 184 form at wavenumber $\mathbf{k}_0^\top \cdot \mathbf{R}^{-1}(\phi)/s$, where the superscript τ denotes a transpose. We
 185 note that Morlet wavelets are not orthonormal. However, this does not hinder our re-
 186 sults as we are focused on continuous wavelet transforms.

187 From the properties of wavelets, it is possible to show they satisfy a generalized Par-
 188 seval’s equality (cf. Appendix A; Daubechies, 1992; Torrence & Compo, 1998; Chen &
 189 Chu, 2017), namely

$$190 \int_{\mathbf{x}} f(\mathbf{x})g(\mathbf{x})d\mathbf{x} = \frac{1}{C_{\Xi}} \int_{\phi} \int_s \int_{\gamma} \frac{\tilde{f}\tilde{g}^*}{s^3} d\gamma ds d\phi. \quad (13)$$

191 Note, if $f = g$, then the variance in f is captured via

$$192 \int_{\mathbf{x}} f^2(\mathbf{x})d\mathbf{x} = \frac{1}{C_{\Xi}} \int_{\phi} \int_s \int_{\gamma} \frac{\tilde{f}^*\tilde{f}}{s^3} d\gamma ds d\phi, \quad (14)$$

193 which identifies the quantity

$$194 \tilde{E}_S(\gamma, \phi, s) = \frac{1}{C_{\Xi}} \frac{\tilde{f}^*\tilde{f}}{s^3}, \quad (15)$$

195 as the energy density of f in wavelet space s and direction ϕ . In other words, Eq. (15)
 196 gives a spectral energy estimate for f that belongs to location γ .

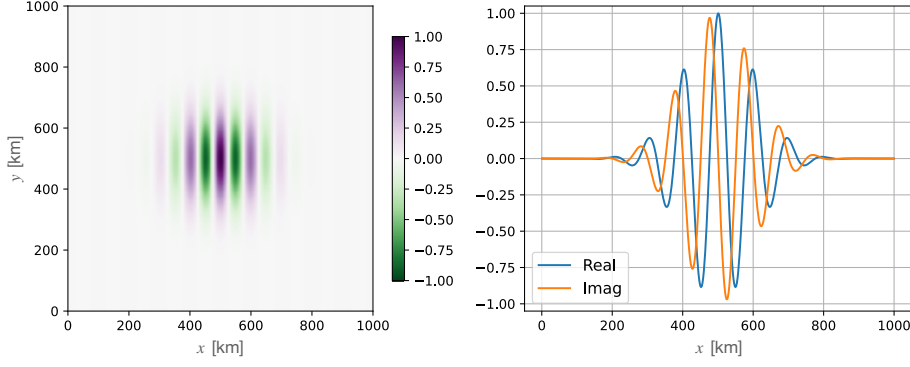


Figure 4. Structure of the mother Morlet wavelet (12) for $c_0 = 0$. A contour plot of the real part of the mother Morlet wavelet is shown in the left panel. Zonal transects of the real and imaginary parts at $y = 500$ km appear in the right panel. The reference lengthscale is $x_0 = 100$ km.

197 At this point, the scale factor in (15), s , is non-dimensional. It is more traditional
 198 in fluid mechanics to discuss energy spectra in terms of wavenumber. As pointed out above,
 199 the effective wavenumber associated with s is $k = 1/(sx_0) = 1/s_0$, where the quan-
 200 tity s_0 has units of length. One can transform (14) from s to s_0 space as

$$201 \int_{\mathbf{x}} f^2(\mathbf{x}) d\mathbf{x} = \frac{1}{C_{\Xi}} \int_{\phi} \int_{s_0} \int_{\gamma} \frac{\tilde{f}^* \tilde{f}}{s_0^3} x_0^2 d\gamma ds_0 d\phi, \quad (16)$$

202 and finally to wavenumber, $k = 1/s_0$, space, ending with

$$203 \int_{\mathbf{x}} f^2(\mathbf{x}) d\mathbf{x} = \frac{1}{C_{\Xi}} \int_{\phi} \int_k \int_{\gamma} \tilde{f}^* \tilde{f} x_0^2 k d\gamma dk d\phi. \quad (17)$$

204 If we now produce wavelet coefficients for the stream function and PV from time
 205 step n of our simulation, and manipulate them appropriately, we obtain

$$206 \tilde{E}_K^n(\gamma, \phi, k) = \frac{1}{C_{\Xi}} \mathcal{R} \left[(-\tilde{\psi}^*) \tilde{q} \right] x_0^2 k, \quad (18)$$

$$207 \tilde{Z}_K^n(\gamma, \phi, k) = \frac{1}{C_{\Xi}} \mathcal{R} \left[\frac{\tilde{q}^* \tilde{q}}{2} \right] x_0^2 k, \quad (19)$$

209 where $\mathcal{R}[\cdot]$ is the real part of the quantity \cdot , as a measure of energy and enstrophy den-
 210 sity in wavelet transform space (cf. Vallis, 2006; Uchida, Deremble, & Penduff, 2021).
 211 Each value of \tilde{E}_K^n and \tilde{Z}_K^n is a random number (as they are associated with each real-
 212 ization of random eddies). Ensemble averaging those values where the members are snap-
 213 shots at intervals of 30 days, returns an estimate of the energy spectrum as a function
 214 of wavenumber k in direction ϕ . The interval of 30 days ensures temporal decorrelation

215 between the density estimates. The spatial locality of the mother wavelet permits the
 216 interpretation of $\tilde{E}_K(\gamma, \phi, k) = \overline{\tilde{E}_K^n(\gamma, \phi, k)}$ as the local energy spectrum at location
 217 γ . The same argument applies for enstrophy.

218 **3 Results**

219 We have opted for this work to calculate the wavelet coefficients explicitly, rather
 220 than by the frequently used Fourier transform method, in view of our eventual interest
 221 in applications to realistic aperiodic and inhomogeneous settings, such as the North At-
 222 lantic basin. The wavelet transform appropriate to the angle ϕ was taken between $[0, -\pi)$
 223 with the azimuthal resolution of $\pi/12$ radian ($= 15^\circ$). The sum of the product of the
 224 wavelet and the data spatially integrated is the wavelet coefficient at the location γ . In
 225 what follows, we consider the quasi two-dimensional flow in the top layer ($j = 1$).

226 **3.1 Spectra over the entire domain**

227 We examine and intercompare the wavelet and Fourier wavenumber spectra and
 228 spectral flux over the entire domain in this section. As the simulated domain is doubly
 229 periodic and on a uniform grid, it is an ideal case for the Fourier method; no window-
 230 ing nor spatial interpolation are applied prior to taking the transform. Although one of
 231 the strengths of the wavelet approach is in negating the necessity of periodicity, we have
 232 chosen such an idealized configuration to test the wavelet method against the Fourier
 233 method where the latter would provide the “true” spectra.

234 While the scaling factor s provides flexibility in defining the wavelet wavenumber,
 235 as opposed to the Fourier approach where, to employ Fast Fourier Transform (FFT) al-
 236 gorithms, the resolution is constrained to $1/L$ with $L (= 1000 \text{ km})$ being the domain size,
 237 we start by computing the wavelet spectra at the center location $\gamma = \gamma_0 = (y_c, x_c) =$
 238 $(500, 500) \text{ km}$ and use the same wavenumbers as the Fourier spectra (k_F). We see from
 239 Fig. 5 that the agreement between the Fourier and wavelet method is excellent (red solid
 240 and black dashed curves respectively) for both the energy and enstrophy spectra at scales
 241 above the dissipation scale.

242 We also show in Fig. 5 a case where we arbitrarily increase the wavelet wavenum-
 243 ber resolution at scales larger than 50 km where the inverse cascade is expected (black
 244 dashed curve); we take $s_0 = [2\Delta x, \dots, 5x_0]$ monotonically spaced with 30 increments,

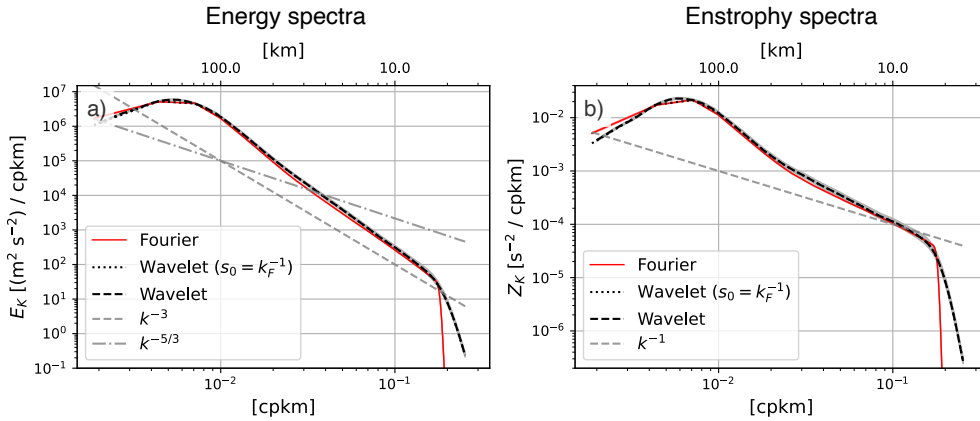


Figure 5. The isotropic (azimuthally-integrated) energy and enstrophy wavenumber spectra of the top layer (a,b). For the wavelet approach, spectra at $\gamma = \gamma_0$ where the wavenumbers are identical to the Fourier wavenumbers ($s_0 = k_F^{-1}$; black dotted) and where the wavenumber resolution is increased at scales larger than 50 km (black dashed) are given. The wavenumber are shown in the lower x axes and corresponding lengthscale in the upper axes. The colored shadings indicate the 95% boot-strapped confidence interval and are shown for the Fourier spectra (red solid) and wavelet spectra with increased wavenumber resolution (black dashed) although the intervals are narrower than the curves themselves.

245 which is trimmed for scales smaller than 50 km, and concatenate this with the Fourier
 246 length scales below 50 km. Features at the lowest wavenumbers (i.e. largest spatial scales)
 247 are better captured compared to the red solid and black dotted curves in Fig. 5 where
 248 the Fourier wavenumber resolution is low. This is beneficial as the scales of interest in
 249 the oceanographic context are often length scales about and larger than the Rossby ra-
 250 dius, associated with mesoscale eddies (Chelton et al., 1998, 2011). The enstrophy spec-
 251 tra are slightly steeper than k^{-1} at scales below the Rossby radius (Fig. 5b), which is
 252 consistent with the KE spectral slope also being steeper than -3 . We attribute the steeper
 253 slope to the excessive PV variance introduced by the stochastic forcing cascading down-
 254 scale (cf. Fig. 6b) and the sporadic emergence of coherent structures (e.g. Fig. 3 left panel;
 255 Benzi et al., 1988; Maltrud & Vallis, 1991). While the spectral slopes do not match ex-
 256 actly to what is expected from the inertial range theory, it is known that the spectral
 257 slopes are sensitive to the model configuration of forcing and dissipation (Maltrud & Val-
 258 lis, 1991), and this does not diminish the agreement between the Fourier and wavelet spec-
 259 tral estimates.

260 Using the wavelet transformation, we can also diagnose the KE and enstrophy spec-
 261 tral flux as

$$262 \quad \tilde{\varepsilon}_K(\gamma, \phi, k) = -\frac{1}{C_\Psi} \int_{k>\kappa} \mathcal{R} \left[\overline{\tilde{u}^*(\mathbf{u} \cdot \nabla u)} + \overline{\tilde{v}^*(\mathbf{u} \cdot \nabla v)} \right] x_0^2 \kappa \, d\kappa, \quad (20)$$

$$263 \quad \tilde{\eta}_K(\gamma, \phi, k) = -\frac{1}{C_\Psi} \int_{k>\kappa} \mathcal{R} \left[\overline{\tilde{q}^*(\mathbf{u} \cdot \nabla q)} \right] x_0^2 \kappa \, d\kappa, \quad (21)$$

264 where negative values imply an inverse cascade towards larger scales and positive val-
 265 ues a forward cascade towards smaller scales (Arbic et al., 2013; Khatri et al., 2018).

267 Comparisons of the spectral fluxes computed using wavelets at a single point (black),
 268 standard Fourier spectra (red) and spatial averages of point-wise wavelets (blue) are shown
 269 in Fig. 6a,b. All approaches clearly indicate a broad forward enstrophy cascade range
 270 at scales smaller than the forcing scale/Rossby radius. Similarly, there is general agree-
 271 ment on the existence of an inverse energy cascade in the limited range of scales larger
 272 than the forcing scale. The lower panels in Fig. 6 show the azimuthally-integrated spec-
 273 tral transfers, i.e. the integrand of (20) and (21).

274 In contrast to calculations of the spectra themselves shown in Fig. 5, the spectral
 275 fluxes computed from wavelet data taken at a single spatial point differ significantly from
 276 the global Fourier estimates. As described below, the wavelet spectral flux estimates are
 277 highly sensitive to the amount of spatial and temporal averaging employed, despite the
 278 homogeneity and statistical stationarity of the flow field. This sensitivity arises because
 279 the flux is the transfer cumulatively integrated from the largest wavenumbers towards
 280 smaller wavenumbers (i.e. (20)) so values at high wavenumbers can have a substantial
 281 effect on the flux at low wavenumbers.

282 The 95% boot-strapped confidence intervals, computed by randomly re-sampling
 283 spectral quantities 9999 times, are shown by shading in Fig. 6. In all cases, single-point
 284 wavelet flux and transfer estimates are highly uncertain, while Fourier estimates are not.

285 We argue this dependency on averaging is associated with the fact that the wavelet
 286 estimate of the spectral transfer only incorporates spatially local information while the
 287 Fourier approach effectively yields a domain-averaged estimate. Namely, the global two-
 288 point correlation function, stemming from the assumption of homogeneity in the Fourier
 289 approach, acts as a spatial averaging operator (cf. Uchida, Jamet, et al., 2021). For this
 290 setting, this assumption is valid, hence the superior performance in flux estimation of
 291 the Fourier approach. Note, however, that the transfer estimates emerging from the wavelet

292 approach, while noisy, do largely agree with those of the Fourier approach. It is in the
 293 integration of the transfers where initial noise in the estimates can result in an erroneous
 294 outcome (compare black and blue curves in Fig. 6d).

295 The expectation is that if we were to take the explicit wavelet transform at every
 296 single grid point, the spatial average of the wavelet spectral flux would converge to the
 297 Fourier approach. We examined this by estimating the wavelet spectral flux and trans-
 298 fer at every five grid points in the diagonal direction (i.e. every ~ 14 km) up to 125 grid
 299 points apart (~ 280 km) from the center point (101 locations in total along $y - y_c =$
 300 $\pm(x - x_c)$). The spatial average of them shown as blue curves in Fig. 6 all come closer
 301 to the Fourier estimate than the black curves. Comparisons of the domain averaged wavelet
 302 estimates to those derived via standard Fourier approach, both in their mean and con-
 303 fidence intervals, significantly improve when averaged over 101 locations ($\langle \tilde{\varepsilon}_K \rangle, \langle \tilde{\eta}_K \rangle$ where
 304 $\langle \cdot \rangle$ is the averaging operator over 101 locations; blue curves in Fig. 6b,d). The Fourier
 305 and wavelet spectral transfer and flux also no longer differ at the 95% confidence inter-
 306 val.

307 We also exhibit the angular orientation of the spectral flux, which the wavelet ap-
 308 proach can extract via its dependence on the angle ϕ (Fig. 7). The flux shown in Fig. 6a,b
 309 as blue dashed curves are the azimuthal integration of angle-dependent fluxes exhibited
 310 in Fig. 7. As the simulated QG flow is configured to be isotropic, the anisotropy seen
 311 in the spectral flux are statistically insignificant within the 95% boot-strapped confidence
 312 interval; the KE flux exhibits an inverse cascade and enstrophy flux a forward cascade
 313 across all angles (Fig. 7b,d).

314 We are thus led to be cautious interpreting single-point wavelet spectral calcula-
 315 tions when applied to what might be termed higher order quantities, like spectral flux.
 316 However, we also point out this is a sword that cuts in both directions. The accuracy
 317 of the Fourier flux estimates depends strongly on their area-wide integrative effect in this
 318 homogeneous setting. Were the flow not homogeneous, the integrative character of the
 319 Fourier approach would obscure the meaning of the result.

320 **3.2 Spectra over a non-periodic subdomain**

321 We now examine the spectra taken over the subdomain given by $y = 200 - 800$ km
 322 and $x = 200 - 800$ km in anticipation of realistic data where periodicity is never satis-

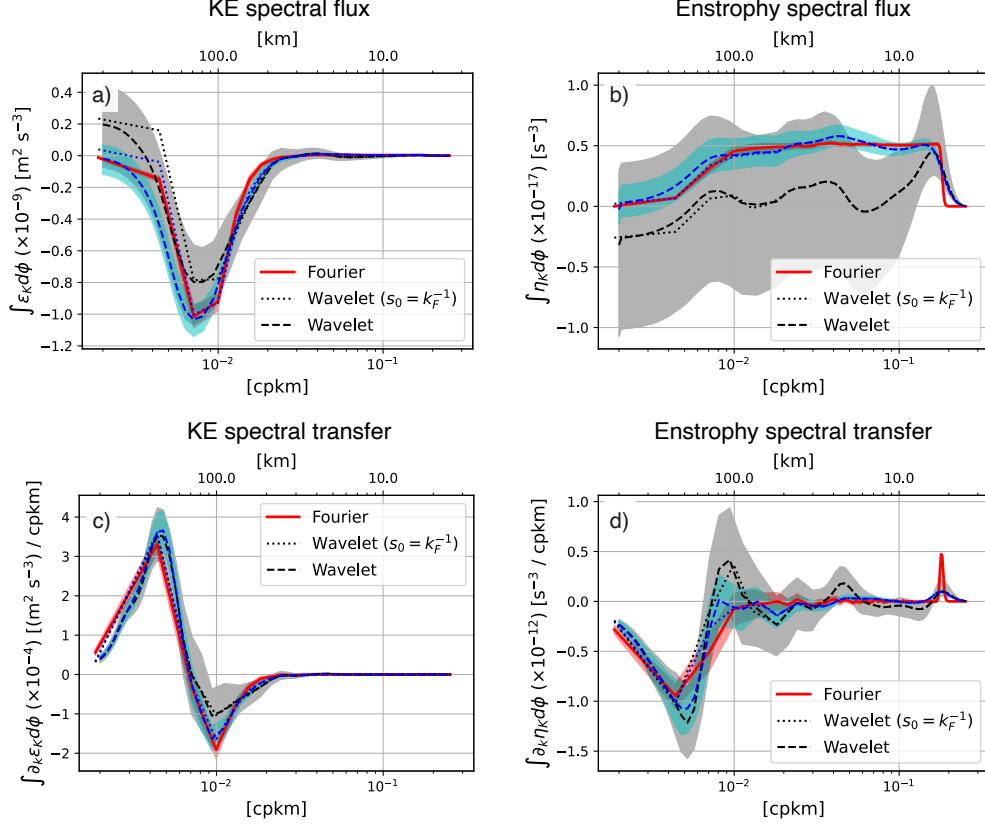


Figure 6. The isotropic (azimuthally integrated) KE and enstrophy wavenumber spectral flux (a,b) and transfer (c,d) respectively. The Fourier method is shown in red and the wavelet approach at $\gamma = \gamma_0$ with wavenumbers identical to the Fourier wavenumbers in dotted ($s_0 = k_F^{-1}$) and the case with increased wavenumber resolution at smaller wavenumbers in dashed curves respectively. The black curves show the wavelet flux and transfer at $\gamma = \gamma_0$, while the blue curves show them averaged over the 101 locations ($\langle \tilde{\varepsilon}_K \rangle, \langle \tilde{\eta}_K \rangle$). The colored shadings indicate the 95% bootstrapped confidence interval and are shown for the Fourier spectra (red solid) and wavelet spectra with increased wavenumber resolution (black and blue dashed curves). The wavenumber is shown in the lower x axes and corresponding lengthscale in the upper axes.

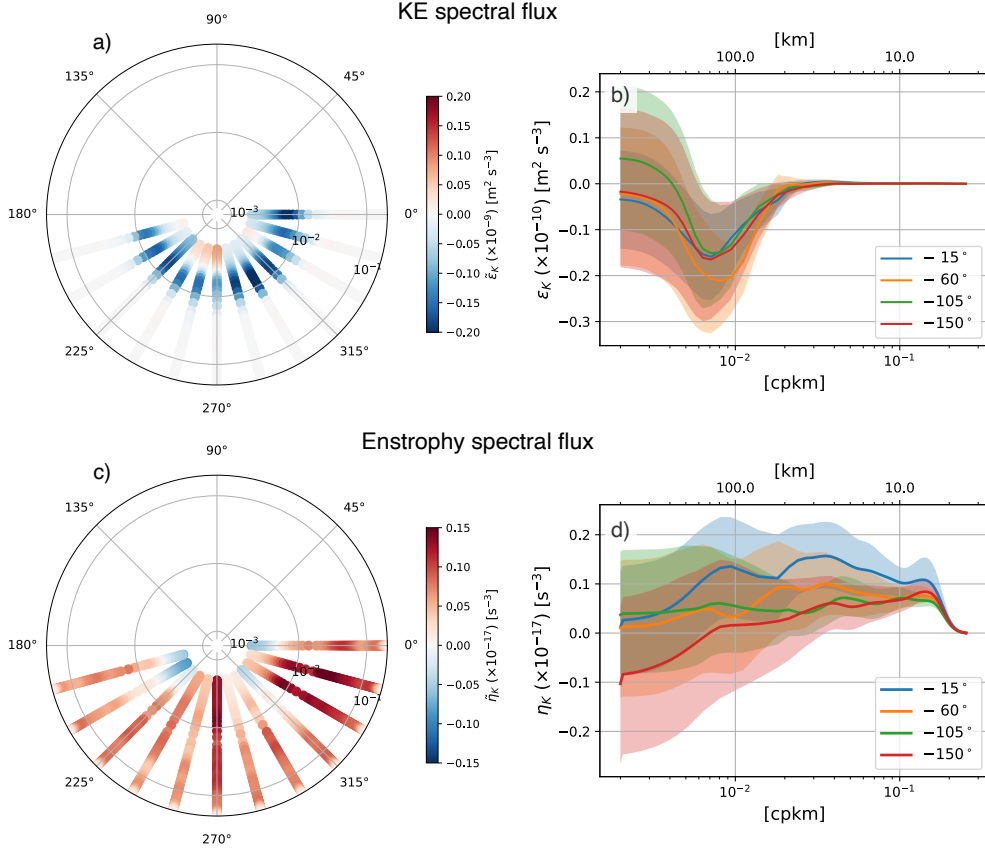


Figure 7. The angular dependence of the KE and enstrophy spectral flux from the wavelet approach plotted radially averaged over the 101 locations ($\langle \tilde{\varepsilon}_K(\phi, k) \rangle$, $\langle \tilde{\eta}_K(\phi, k) \rangle$; a,c). The radial axes are the wavenumbers in logarithmic scaling with the increased wavenumber resolution. The fluxes are symmetric about the origin so we only show for angles $[0, -\pi)$. The 95% boot-strapped confidence intervals are given for four arbitrary angles (b,d).

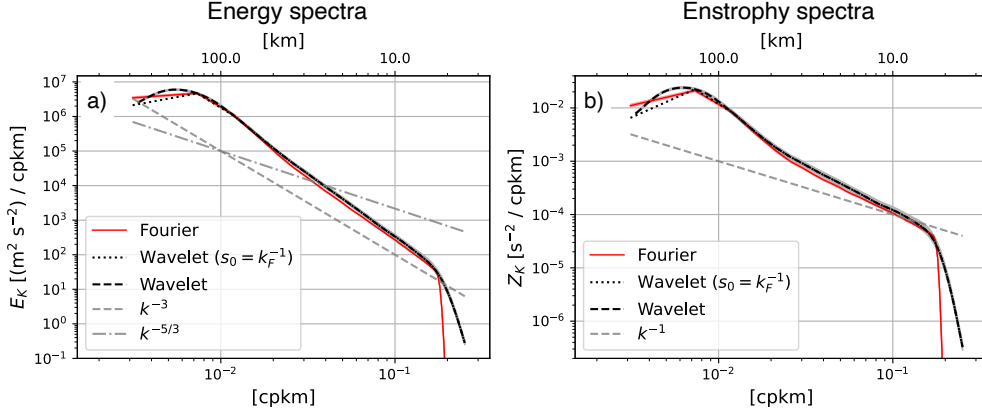


Figure 8. Same as Fig. 5 but for the subdomain of $y = 200\text{-}800$ km and $x = 200\text{-}800$ km. The confidence intervals are again narrower than the curves themselves.

323 fied. As the data are no longer periodic, the Fourier approach requires the data to be
 324 windowed. This will highlight the strength of the locality in the wavelet approach where
 325 windowing of the data is unnecessary. Prior to taking the Fourier transforms, we applied
 326 Hann windows (Arbic et al., 2013; Uchida et al., 2017; Uchida, Jamet, et al., 2021) and
 327 then corrected for their amplitude. Comparing Figs. 5 and 8, we see that the spectral
 328 estimates are still robust. The low resolution at lower wavenumbers from the Fourier method
 329 makes it difficult to detect the spectral shape at scales above ~ 100 km due to the do-
 330 main size being small; there are only two wavenumber points at scales larger than the
 331 Rossby radius (red curves in Fig. 8). The wavelet approach, on the other hand through
 332 its flexibility in s , still captures a smooth spectral estimate; the scaling for the wavelet
 333 approach was adjusted to $s_0 = [2\Delta x, \dots, 3x_0]$ over 30 monotonic increments in order to
 334 account for the smaller domain and then replaced by the Fourier wavenumber at scales
 335 smaller than 50 km.

336 Regarding the spectral transfer, the Fourier approach is significantly affected by
 337 the tapering at the lowest wavenumbers (red curves in Figs. 6c,d and 9c,d) but the wavelet
 338 approach is still able to capture the change in sign in its curvature (blue dashed curves
 339 in Figs. 6c,d and 9c,d). The enstrophy spectral flux and transfer tend to be particularly
 340 sensitive to the local nature of wavelet transforms. Similar to Fig. 6, we see a significant
 341 improvement in the wavelet estimates both in their mean and confidence intervals when
 342 averaged over 101 locations, particularly for enstrophy ($\langle\tilde{\eta}_K\rangle$; blue curves in Fig. 9b,d).
 343

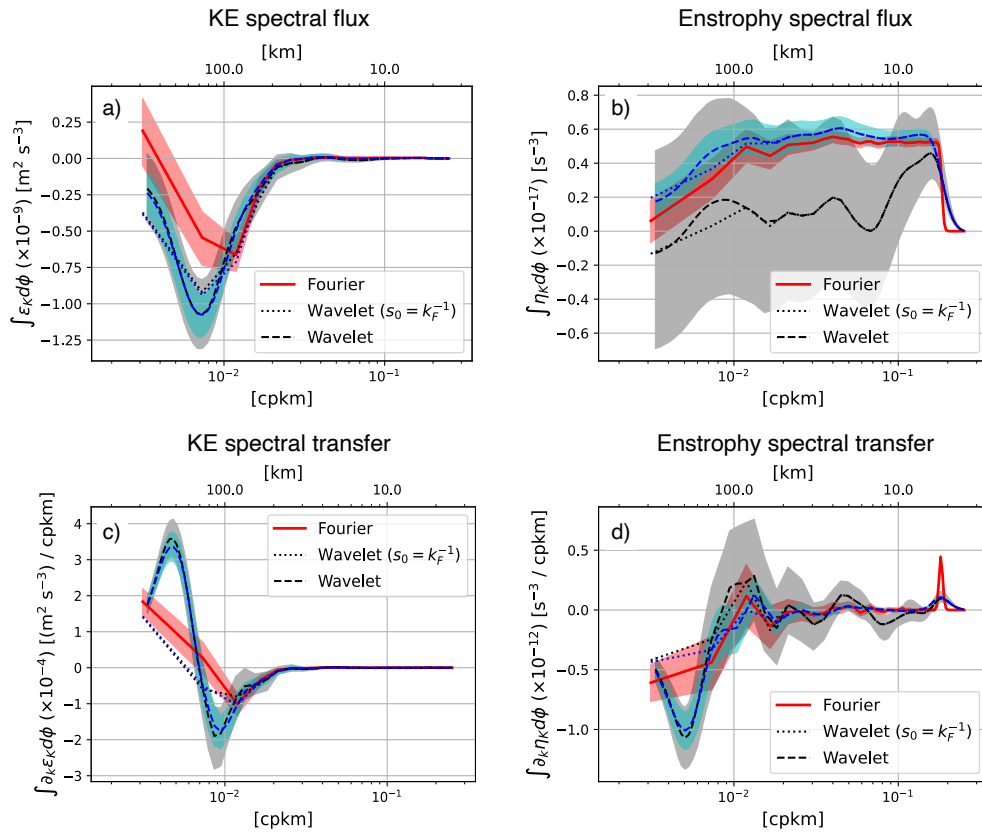


Figure 9. Same as Fig. 6 but for the subdomain of $y = 200\text{-}800$ km and $x = 200\text{-}800$ km.

344 4 Conclusions and discussion

345 In this study, we have described and documented a wavelet-based technique for spec-
 346 tral analyses in an oceanographic context. The wavelet approach employed here, through
 347 its dependence on a scale parameter s , returns effectively a one-dimensional (1D) spec-
 348 tral estimate, and its incorporating of two-dimensional data allows for information re-
 349 garding local anisotropies through its angular dependency ϕ (Fig. 7).

350 We have demonstrated its utility by applying it to a doubly-periodic, two-layer, quasi-
 351 geostrophic (QG) simulation. The flow analyzed in this study is highly idealized being
 352 spatially isotropic and homogeneous in the horizontal dimensions. The idealized setting,
 353 however, is expected to yield known spectral cascades, so it can be used as a test bed
 354 for the wavelet approach. The agreement between the wavelet and Fourier approach, par-
 355 ticularly for the spectra (Figs. 5 and 8), encourages the usage of wavelets with its ad-
 356 ditional strengths of being able to capture the local features of the flow. While numer-
 357 ically efficient algorithms exist to take the wavelet transform (coined as Fast Wavelet Trans-
 358 forms; e.g. Beylkin et al., 1991), they face the same conundrum as FFTs requiring: i)
 359 periodic boundary conditions, and ii) filling in missing data points. We have, therefore,
 360 taken the approach of explicitly computing the wavelet transform (7), which negates the
 361 two necessities and will benefit realistic settings such as the North Atlantic basin. The
 362 robustness of the spectra is comforting, but we also emphasize the need for caution when
 363 computing higher-order spectral quantities, like spectral fluxes, which involve spatial deriva-
 364 tives. The disagreement arises from the local nature of wavelets; the Fourier method in-
 365 corporates spatially global information and hence can be thought as a spatial average
 366 of spectral estimates. This is evident from the fact that upon spatially averaging the wavelet
 367 spectral transfer over multiple locations, the confidence interval improved and its mean
 368 converged towards the Fourier estimate.

369 Our work is complementary to a growing list of literature on spectral methods al-
 370 ternative to the Fourier approach: Aluie et al. (2018); Sadek and Aluie (2018); Schubert
 371 et al. (2020); Storer et al. (2022) and Contreras et al. (2022) where they use a spatial
 372 filter to examine the KE spectra and cross-scale transfer, Lindborg (2015); Balwada et
 373 al. (2016, 2022); LaCasce (2016); Poje et al. (2017) and Pearson et al. (2020) where they
 374 implement structure functions, Jamet et al. (2020) where they employ the Green’s func-
 375 tion, and Uchida, Jamet, et al. (2021) where they use Empirical Orthogonal Functions.

376 Barkan et al. (2021) and Srinivasan et al. (2022) apply the filtering method in both the
377 spatiotemporal dimensions. Liang and Anderson (2007); Liang (2016) and Yang et al.
378 (2021) are also interesting attempts in implementing a multiscale window transform to
379 examine the energy exchange across spatiotemporal scales by decomposing the flow with
380 a set of orthogonal windows. Here, we have documented the wavelet-based cross-scale
381 energetics in the spectral context. While the form of the Parseval’s equality will slightly
382 change, namely in the power of scaling s , the wavelet method can also be extended to
383 estimating frequency-wavenumber spectra (e.g. Torres et al., 2018; Uchida et al., 2019);
384 this will allow us to decompose the balanced and unbalanced motions in non-periodic
385 settings.

386 **Open Research**

387 The wavelet transforms were taken using the `xwavelet` Python package (Uchida
388 & Dewar, 2022) and Fourier transforms using the `xrft` Python package (Uchida et al.,
389 2022). The `pyqg` model is available through Github (Abernathy et al., 2022). Jupyter
390 notebooks used to run the `pyqg` simulation and conduct analyses are available via Github
391 (Uchida, 2023).

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Appendix A Parseval's equality

In this appendix, we review the Parseval's equality for two-dimensional wavelet transforms (Daubechies, 1992; Torrence & Compo, 1998; Mallat, 1999; Chen & Chu, 2017),

i.e.

$$\int f g d\mathbf{x} = \frac{1}{C_{\Xi}} \iiint \tilde{f} \tilde{g}^* \frac{1}{s^3} d\gamma ds d\phi. \quad (\text{A1})$$

Using (9), the right-hand side can be expanded as

$$\begin{aligned} \iiint \tilde{f} \tilde{g}^* \frac{1}{s^3} d\gamma ds d\phi &= \iiint \frac{1}{s^5} \int_{\mathbf{x}} f(\mathbf{x}) \xi^*(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x} - \gamma}{s}\right)) d\mathbf{x} \int_{\chi} g^*(\chi) \xi(\mathbf{R}^{-1} \cdot \left(\frac{\chi - \gamma}{s}\right)) d\chi d\gamma ds d\phi \\ &= \int_{\mathbf{x}} \int_{\chi} f g^* \iiint \frac{1}{s^5} \xi^*(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x} - \gamma}{s}\right)) \xi(\mathbf{R}^{-1} \cdot \left(\frac{\chi - \gamma}{s}\right)) d\gamma ds d\phi d\mathbf{x} d\chi. \end{aligned} \quad (\text{A2})$$

Now, consider the wavelet transform of the Dirac delta function

$$\begin{aligned} \tilde{\delta} &= \int_{\mathbf{x}} \delta(\mathbf{x} - \mathbf{x}') \frac{1}{s} \xi^*(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x} - \gamma}{s}\right)) d\mathbf{x} \\ &= \frac{1}{s} \xi^*(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x}' - \gamma}{s}\right)). \end{aligned} \quad (\text{A3})$$

Hence, the inverse wavelet transform becomes

$$\begin{aligned} \delta &= \frac{1}{C_{\Xi}} \iiint \tilde{\delta} \frac{1}{s^4} \xi(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x} - \gamma}{s}\right)) d\phi ds d\gamma \\ &= \frac{1}{C_{\Xi}} \iiint \frac{1}{s^5} \xi^*(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x}' - \gamma}{s}\right)) \xi(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x} - \gamma}{s}\right)) d\phi ds d\gamma. \end{aligned} \quad (\text{A4})$$

Plugging (A4) into (A2) yields

$$\iiint \tilde{f} \tilde{g}^* \frac{1}{s^3} d\gamma ds d\phi = C_{\Xi} \int_{\mathbf{x}} f g^* d\mathbf{x}, \quad (\text{A5})$$

and we obtain (A1).

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