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Wavelet-based wavenumber spectral estimate of eddy kinetic energy: Idealized quasi-geostrophic flow

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Key Points:

- A wavelet-based spectral method to estimate eddy variability is described.
- Wavenumber spectra of eddies are estimated for a doubly-periodic quasi-geostrophic flow.
- The wavelet and Fourier approach agree well in their estimates of spectra and spectral flux.

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17 **Abstract**

18 A wavelet-based method is re-introduced in an oceanographic and spectral context to
 19 estimate wavenumber spectrum and spectral flux of kinetic energy and enstrophy. We
 20 apply this to a numerical simulation of idealized, doubly-periodic quasi-geostrophic flows,
 21 i.e. the flow is constrained by the Coriolis force and vertical stratification. The double
 22 periodicity allows for a straightforward Fourier analysis as the baseline method. Our wavelet
 23 spectra agree well with the canonical Fourier approach but with the additional strengths
 24 of negating the necessity for the data to be periodic and being able to extract local anisotropies
 25 in the flow. Caution is warranted, however, when computing higher-order quantities, such
 26 as spectral flux.

27 **Plain language summary**

28 Chaotic flows such as the ocean currents, atmospheric winds and turbulence in gen-
 29 eral are fundamentally impossible to analytically predict, namely, to formulate a math-
 30 ematical general solution. Nevertheless, the interest in describing such chaotic flows can
 31 be found in examples as old as Leonardo da Vinci’s sketch of turbulence. While we can-
 32 not obtain an analytical description of turbulence, we can extract statistical information
 33 from turbulent flows and a common descriptor has been the wavenumber spectrum. The
 34 spectrum reveals at each spatial scale, the level of variability the flow carries. Here, we
 35 re-introduce an alternative method in estimating the spectrum based on wavelet func-
 36 tions.

37 **1 Introduction**

38 Fundamental to the goal of properly modelling climate system dynamics is under-
 39 standing and quantifying how energy is both distributed, and ultimately transferred, across
 40 an extremely broad range of dynamically active space and time scales. In the atmospheric
 41 and ocean context, the most common means of quantifying the scale-dependent energy
 42 content of a chaotic, turbulent flow field is the energy spectrum (e.g. Taylor, 1938; Kol-
 43 mogorov, 1941; Charney, 1971; Nastrom & Gage, 1983; Yaglom, 2004) given by the Fourier
 44 transform of two-point (spatial or temporal) velocity correlations.

45 We will focus in this paper on spatial correlations since the behavior of wavenum-
 46 ber spectra are described by ‘inertial range’ theories predicting spectral slopes and cas-
 47 cades (Vallis, 2006). The standard Fourier approach has had great success in providing

48 us with spectral estimates of energy partition (e.g. Stammer, 1997; Xu & Fu, 2011, 2012;
 49 Uchida et al., 2017; Ajayi et al., 2021) and its straightforward mathematical formula-
 50 tion facilitates the spectral interpretation in the original context, namely statistically ho-
 51 mogeneous flows where Fourier decompositions are natural. Issues persist, however, in
 52 geophysical flows which are statistically inhomogeneous, anisotropic and non-stationary
 53 (Uchida, Jamet, et al., 2021). The assumption of homogeneity lies on the fact that a Fourier
 54 transform is a global operator over the entire space-time domain of interest. In other words,
 55 the Fourier description of the field conflates different regimes of an inhomogeneous flow.
 56 A notable example is in the separated Gulf Stream region where the energetics have been
 57 argued to be distinct from the gyre interior (Jamet et al., 2021).

58 With a growing acknowledgement of the shortcomings of the Fourier approach, there
 59 has been a recent effort in the geophysical sciences to re-examine the cross-scale ener-
 60 getics. Notable examples are: i) Aluie et al. (2018); Sadek and Aluie (2018); Schubert
 61 et al. (2020); Storer et al. (2022); Srinivasan et al. (2022) and Contreras et al. (2022) where
 62 they implement a spatial filter, ii) Lindborg (2015); Balwada et al. (2016, 2022); LaCasce
 63 (2016); Poje et al. (2017) and Pearson et al. (2020) where they use structure functions,
 64 iii) Jamet et al. (2020) where they employ the Green’s function, and iv) Uchida, Jamet,
 65 et al. (2021) where they use Empirical Orthogonal Functions all with the goal of exam-
 66 ining the KE spectra and cross-scale transfer in the wavenumber domain. The overall
 67 consistent picture is that at scales about $O(100\text{ km})$ where the oceanic motions are con-
 68 strained by the Earth’s rotation and vertical stratification, KE cascades upscale while
 69 KE on the scale of $O(10\text{ km})$ tend to cascade downscale due to a loss of balance with the
 70 two constraining forces. While all these approaches, including the Fourier method, can
 71 capture within limits the spatial anisotropy when examined on a two-dimensional (2D)
 72 wavenumber plane, they lose this information when reduced to one-dimensional (1D) spec-
 73 tral quantities.

74 Here, we use a wavelet-based technique which yields localized pseudo-Fourier 1D
 75 wavenumber spectra capable of capturing the local anisotropies in the flow (Daubechies,
 76 1992; Perrier et al., 1995). Wavelets emerged in the 1980s as a way to analyze time and
 77 space series in more local manner than was possible using Fourier techniques (e.g. Vasi-
 78 lyev & Paolucci, 1997; Doglioli et al., 2007; Alvera-Azcárate et al., 2007; Thomson & Emery,
 79 2014), although strong parallels and connections are to be found between the two meth-
 80 ods (Katul & Parlange, 1995; Torrence & Compo, 1998). We will argue the localized na-

81 ture of wavelets allows us to capture the inhomogeneity and anisotropy in the flow (Farge
 82 et al., 1992; Horbury et al., 2008). We then apply the wavelet approach to estimate the
 83 horizontal kinetic energy (KE) and enstrophy spectral flux from a doubly-periodic quasi-
 84 geostrophic (QG) flow, and to a flow subdomain where periodicity no longer applies. The
 85 comparisons illustrate some of the advantages of the wavelet approach.

86 The paper is organized as follows: We describe the QG model and provide an overview
 87 of the wavelet method in Section 2. Results are given in Section 3 where we compare our
 88 wavelet spectra to the canonical Fourier spectra. Conclusions are given in Section 4.

89 2 Theory and technique

90 We describe the configuration of our quasi-geostrophic (QG) model and provide
 91 an overview of the wavelet method.

92 2.1 Description of the quasi-geostrophic simulation

93 We consider a stochastically forced two-layer QG flow in a doubly periodic f plane
 94 domain (i.e. $\beta = f_y = 0$) under rigid-lid and flat bottom conditions. Solutions to the
 95 QG potential vorticity (PV) equation

$$96 \quad q_{j_t} + J(\psi_j, q_j) = -r_b \nabla^2 \psi_j \delta_{j,2} + \mathcal{Q} \quad (1)$$

98 are computed using the pseudo-spectral `pyqg` model (Abernathey et al., 2022), where
 99 $\delta_{i,j}$ is the usual Kronecker delta function and layer numbers are denoted $j = 1, 2$. The
 100 linear bottom drag coefficient is $r_b = 5.787 \times 10^{-7} \text{ s}^{-1}$. The PV in each layer are

$$101 \quad q_1 = \nabla^2 \psi_1 + F_1(\psi_2 - \psi_1), \quad (2)$$

$$102 \quad q_2 = \nabla^2 \psi_2 + F_2(\psi_1 - \psi_2). \quad (3)$$

104 The vortex stretching coefficients are $F_1 = \frac{(2\pi/R_d)^2}{1+\delta}$, $F_2 = \delta F_1$ where the internal Rossby
 105 deformation radius was prescribed as $R_d = 100 \text{ km}$. Each layer thickness is $(H_1, H_2) =$
 106 $(500, 2000) \text{ m}$ respectively, giving $\delta = H_1/H_2 = 0.25$. The square domain size is $L_0 =$
 107 1000 km with the spatial resolution of $\sim 2 \text{ km}$ (512×512 grid points). In order to pre-
 108 vent the system from equilibrating to the well-known single pair of positive and nega-
 109 tive vortices (Vallis, 2006), a vertically uniform forcing was introduced as

$$110 \quad \mathcal{Q} = A_q w(t, y, x), \quad (4)$$

111 where $A_q = 10^{-15} \text{ s}^{-2}$ is the amplitude and $w(t, y, x)$ is white noise in space-time with
 112 zero mean and $O(1)$ amplitude per layer (Fig. 1). The quantity $w(t, y, x)$ was computed
 113 by taking the inverse Fourier transform of a ring in wavenumber space

$$\hat{w}(t, k^y, k^x) = \begin{cases} a(t, k^y, k^x) + ib(t, k^y, k^x), & \text{if } (R_d + \delta_R)^{-1} < \sqrt{k^x^2 + k^y^2} < (R_d - \delta_R)^{-1}, \\ 0, & \text{otherwise} \end{cases}, \quad (5)$$

114
 115 where k^x and k^y are the zonal and meridional wavenumbers respectively, a and b are Gaus-
 116 sian random variables in the horizontal wavenumber space with zero mean and standard
 117 deviation of unity, and $\delta_R = 5 \text{ km}$. After taking the inverse Fourier transform, the hor-
 118 izontal spatial mean is removed and then divided by the maximum value in the horizon-
 119 tal dimension. In other words, the model is stochastically forced at scales about the Rossby
 120 radius uncorrelated in time. No background PV was prescribed. The model was spun
 121 up for 10 years from a state of rest, at which point area averaged energy had equilibrated
 122 (not shown), and then run for another 20 years with outputs saved every 10 days as in-
 123 stantaneous snapshots. The timeseries of the kinetic energy (KE; $K = (\mathbf{u} \cdot \mathbf{u})/2$) and
 124 potential energy (PE; $2(\psi_1 - \psi_2)^2/R_d^2$) for the 20 years of output are given in Fig. 2,
 which mirror each other and roughly show an equipartition.

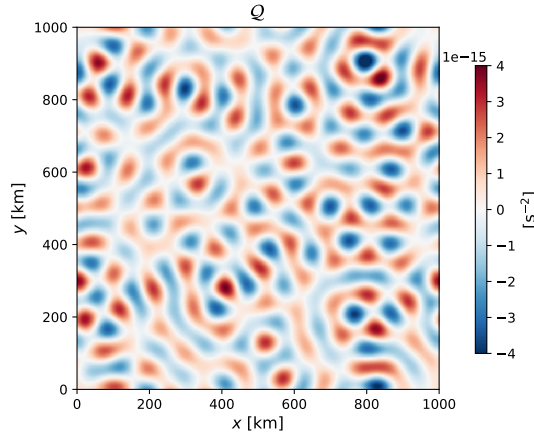


Figure 1. Example of how the vertically uniform stochastic forcing Q looks like for an arbitrary time step.

125

126 In this simple configuration, it is expected the flow will be both homogeneous and
 127 isotropic in the horizontal dimensions. Further, classical theory predicts the existence
 128 of an inverse cascade of KE and hence a $-5/3$ power law at scales larger than the forc-

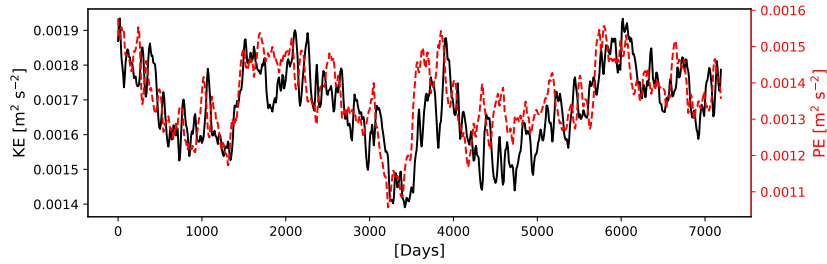


Figure 2. Vertically and domain averaged kinetic energy (KE; black solid) and (available) potential energy (PE; red dashed).

129 ing scale, and a forward cascade of enstrophy and hence -3 power law at smaller scales
 130 that are above the viscous dissipation scale (Vallis, 2006). In this sense, we ‘know’ what
 131 the answer should be and can use the results to test the efficacy of the wavelet trans-
 132 form. The double periodicity also allows for a straightforward comparison between the
 133 wavelet and Fourier approach as no windowing of the data is necessary in applying the
 134 transforms. We exhibit the top- and bottom-layer PV at the last time step of the model’s
 10th year in Fig. 3.

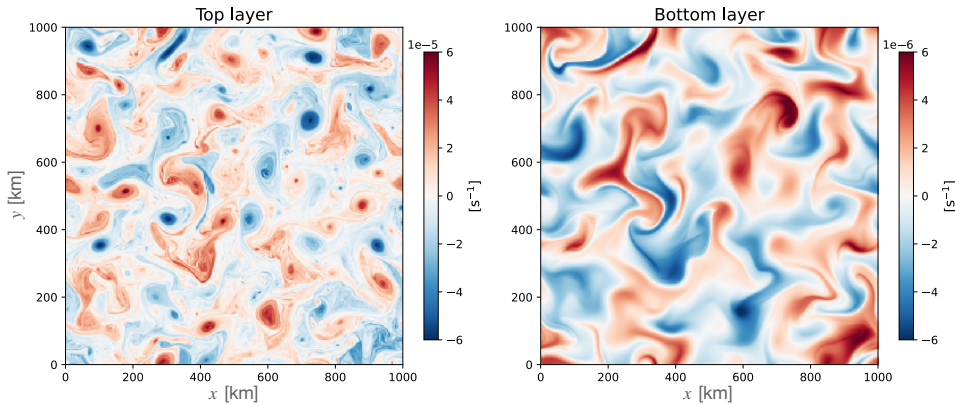


Figure 3. The PV at the last time step of the 10th simulated year in the top and bottom layer. Note the order of magnitude difference in the two panels.

135

136 2.2 Spectral Considerations

137 For the reasons outlined in the introduction, we depart from the classical Fourier
 138 approach to compute wavenumber spectra, but do note the utility of that wavenumber

139 spectrum emerges largely from Parseval's equality

$$140 \quad \int_{\mathbf{x}} K(\mathbf{x}) d\mathbf{x} = \int_{\mathbf{k}} \widehat{E}_K(\mathbf{k}) d\mathbf{k}, \quad (6)$$

141 where $\mathbf{x} = (x, y)$, $\mathbf{k} = (k^x, k^y)$ (e.g. Scott & Wang, 2005; Capet et al., 2008; Uchida
142 et al., 2017). The Fourier energy spectrum is given by $2\widehat{E}_K(\mathbf{k}) = \hat{\mathbf{u}}^* \cdot \hat{\mathbf{u}}$ where the Fourier
143 transform of the velocity is denoted by the hat ($\hat{\mathbf{u}}$) and the superscript * denotes the com-
144 plex conjugate. This equivalence of the area integrated KE to the wavenumber integrated
145 Fourier spectrum motivates the latter's interpretation as the KE density in the wavenum-
146 ber domain.

147 We base our spectral analysis on wavelet decompositions, rather than Fourier trans-
148 forms, as the space-time locality of wavelets does not require the data to be periodic. Given
149 a function dependent on two spatial dimensions, $f(\mathbf{x})$, its continuous wavelet transform
150 is given by (Daubechies, 1992; Torrence & Compo, 1998)

$$151 \quad \tilde{f}(s, \phi, \boldsymbol{\gamma}) = \int_{\Omega} f(\mathbf{x}) \frac{1}{s} \xi^*(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x} - \boldsymbol{\gamma}}{s} \right)) d\mathbf{x}, \quad (7)$$

152 where the integration is taken over the whole domain of interest Ω and \mathbf{R}^{-1} is the in-
153 verse of the rotation matrix

$$154 \quad \mathbf{R}^{-1} = \begin{pmatrix} \cos(\phi) & \sin(\phi) \\ -\sin(\phi) & \cos(\phi) \end{pmatrix}, \quad (8)$$

155 for rotation through an angle ϕ relative to the x axis. The quantity s is referred to as
156 the 'scale', $\boldsymbol{\gamma} (\in \mathbf{R}^2)$ are the two-dimensional coordinates of interest, $\xi(\mathbf{x})$ is the so-called
157 'mother' wavelet and $\xi(\mathbf{R}^{-1} \cdot (\mathbf{x} - \boldsymbol{\gamma})/s)$ in (7) are the daughter wavelets. The quan-
158 tities \tilde{f} are called the wavelet coefficients. Note that the field of wavelet coefficients is
159 a filtered version of the original data.

160 Subject to the 'admissibility condition' $C_{\Xi} < \infty$, the original function f can be
161 reconstructed from the wavelet coefficients (Daubechies, 1992; Torrence & Compo, 1998)

$$162 \quad f(\mathbf{x}) = \frac{1}{C_{\Xi}} \int_{\boldsymbol{\gamma}} \int_s \int_{\phi} \tilde{f}(s, \phi, \boldsymbol{\gamma}) \frac{1}{s^4} \xi(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x} - \boldsymbol{\gamma}}{s} \right)) d\phi ds d\boldsymbol{\gamma}. \quad (9)$$

163 If $\hat{\Xi}(\mathbf{k})$ is the Fourier transform of the mother wavelet, then

$$164 \quad C_{\Xi} = \int_{\mathbf{k}} \frac{\hat{\Xi}^* \hat{\Xi}}{\mathbf{k} \cdot \mathbf{k}} d\mathbf{k}. \quad (10)$$

165 The so-called 'admissibility condition' implies that the mother wavelet defines a well-behaved
166 class of wavelet transforms. Many functions satisfy (10) provided they have zero mean

$$167 \quad \int_{\mathbf{x}} \xi(\mathbf{x}) d\mathbf{x} = 0. \quad (11)$$

168 For current purposes, we will employ the so-called Morlet wavelet (Morlet et al., 1982;
 169 Gabor, 1946), i.e.

$$170 \quad \xi(\mathbf{x}) = (e^{-2\pi i \mathbf{k}_0 \cdot \mathbf{x}} - c_0) e^{-\frac{\mathbf{x} \cdot \mathbf{x}}{2x_0^2}}, \quad (12)$$

171 where c_0 is a constant included to ensure that (11) is met. The central wavenumber \mathbf{k}_0
 172 is taken to be $\mathbf{k}_0 = (k_0, 0)$ and the quantity x_0 is a reference length scale, here taken
 173 to be the Rossby radius ($x_0 = 100$ km), viz. the central length scale of the mother wavelet.
 174 We will choose $k_0 = 1/x_0$, in which case the constant c_0 is quite small and generally
 175 ignored (i.e. $c_0 = 0$), a convention adopted in this paper. Plots of (12) are found in Fig. 4.
 176 Note that the Morlet mother wavelet consists of a wave of wavelength x_0 inside a Gaus-
 177 sian envelope of decay scale $\sqrt{2}x_0$. Thus for $s = 1$ and $\phi = 0$, the wavelet coefficient
 178 produced by this transformation comments on the presence of the wavenumber $\mathbf{k}_0 =$
 179 $(k_0, 0)$ at location γ in the original data. Increasing the rotation angle ϕ and filtering
 180 returns information about the presence of the same wavelength at angle $-\phi$. Finally al-
 181 lowing s to vary modifies the filter so that the primary wavenumber of the filter is $k =$
 182 $1/(sx_0)$. The Morlet wavelet coefficient can thus be thought of as a ‘local’ Fourier trans-
 183 form at wavenumber $\mathbf{k}_0^\top \cdot \mathbf{R}^{-1}(\phi)/s$, where the superscript τ denotes a transpose. We
 184 note that Morlet wavelets are not orthonormal. However, this does not hinder our re-
 185 sults as we are focused on continuous wavelet transforms.

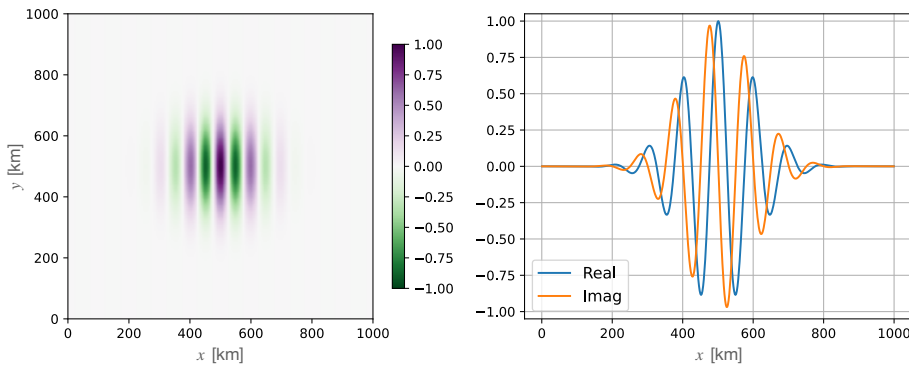


Figure 4. Structure of the mother Morlet wavelet (12) for $c_0 = 0$. A contour plot of the real part of the mother Morlet wavelet is shown in the left panel. Zonal transects of the real and imaginary parts at $y = 500$ km appear in the right panel. The reference lengthscale is $x_0 = 100$ km.

186 From the properties of wavelets, it is possible to show they satisfy a generalized Par-
 187 seval’s equality (cf. Appendix A; Daubechies, 1992; Torrence & Compo, 1998; Chen &

188 Chu, 2017), namely

$$189 \int_{\mathbf{x}} f(\mathbf{x})g(\mathbf{x})d\mathbf{x} = \frac{1}{C_{\Xi}} \int_{\phi} \int_s \int_{\gamma} \frac{\tilde{f}\tilde{g}^*}{s^3} d\gamma ds d\phi. \quad (13)$$

190 Note, if $f = g$, then the variance in f is captured via

$$191 \int_{\mathbf{x}} f^2(\mathbf{x})d\mathbf{x} = \frac{1}{C_{\Xi}} \int_{\phi} \int_s \int_{\gamma} \frac{\tilde{f}^*\tilde{f}}{s^3} d\gamma ds d\phi, \quad (14)$$

192 which identifies the quantity

$$193 \tilde{E}_S(\gamma, \phi, s) = \frac{1}{C_{\Xi}} \frac{\tilde{f}^*\tilde{f}}{s^3}, \quad (15)$$

194 as the energy density of f in wavelet space s and direction ϕ . In other words, Eq. (15)
195 gives a spectral energy estimate for f that belongs to location γ .

196 At this point, the scale factor in (15), s , is non-dimensional. It is more traditional
197 in fluid mechanics to discuss energy spectra in terms of wavenumber. As pointed out above,
198 the effective wavenumber associated with s is $k = 1/(sx_0) = 1/s_0$, where the quan-
199 tity s_0 has units of length. One can transform (14) from s to s_0 space as

$$200 \int_{\mathbf{x}} f^2(\mathbf{x})d\mathbf{x} = \frac{1}{C_{\Xi}} \int_{\phi} \int_{s_0} \int_{\gamma} \frac{\tilde{f}^*\tilde{f}}{s_0^3} x_0^2 d\gamma ds_0 d\phi, \quad (16)$$

201 and finally to wavenumber, $k = 1/s_0$, space, ending with

$$202 \int_{\mathbf{x}} f^2(\mathbf{x})d\mathbf{x} = \frac{1}{C_{\Xi}} \int_{\phi} \int_k \int_{\gamma} \tilde{f}^*\tilde{f} x_0^2 k d\gamma dk d\phi. \quad (17)$$

203 If we now produce wavelet coefficients for the stream function and PV from time
204 step n of our simulation, and manipulate them appropriately, we obtain

$$205 \tilde{E}_K^n(\gamma, \phi, k) = \frac{1}{C_{\Xi}} \mathcal{R} \left[(-\tilde{\psi}^*)\tilde{q} \right] x_0^2 k, \quad (18)$$

$$206 \tilde{Z}_K^n(\gamma, \phi, k) = \frac{1}{C_{\Xi}} \mathcal{R} \left[\frac{\tilde{q}^*\tilde{q}}{2} \right] x_0^2 k, \quad (19)$$

207 where $\mathcal{R}[\cdot]$ is the real part of the quantity \cdot , as a measure of energy and enstrophy den-
208 sity in wavelet transform space (cf. Vallis, 2006; Uchida, Deremble, & Penduff, 2021).
209 Each value of \tilde{E}_K^n and \tilde{Z}_K^n is a random number (as they are associated with each real-
210 ization of random eddies). Ensemble averaging those values where the members are snap-
211 shots at intervals of 30 days, returns an estimate of the energy spectrum as a function
212 of wavenumber k in direction ϕ . The interval of 30 days ensures temporal decorrelation
213 between the density estimates. The spatial locality of the mother wavelet permits the
214 interpretation of $\tilde{E}_K(\gamma, \phi, k) = \overline{\tilde{E}_K^n(\gamma, \phi, k)}$ as the local energy spectrum at location
215 γ . The same argument applies for enstrophy.
216

217 3 Results

218 We have opted for this work to calculate the wavelet coefficients explicitly, rather
 219 than by the frequently used Fourier transform method, in view of our eventual interest
 220 in applications to realistic aperiodic and inhomogeneous settings, such as the North At-
 221 lantic basin. The wavelet transform appropriate to the angle ϕ was taken between $[0, -\pi)$
 222 with the azimuthal resolution of $\pi/12$ radian ($= 15^\circ$). The sum of the product of the
 223 wavelet and the data spatially integrated is the wavelet coefficient at the location γ . In
 224 what follows, we consider the quasi two-dimensional flow in the top layer ($j = 1$).

225 3.1 Spectra over the entire domain

226 We examine and intercompare the wavelet and Fourier wavenumber spectra and
 227 spectral flux over the entire domain in this section. As the simulated domain is doubly
 228 periodic and on a uniform grid, it is an ideal case for the Fourier method; no window-
 229 ing nor spatial interpolation are applied prior to taking the transform. Although one of
 230 the strengths of the wavelet approach is in negating the necessity of periodicity, we have
 231 chosen such an idealized configuration to test the wavelet method against the Fourier
 232 method where the latter would provide the “true” spectra.

233 While the scaling factor s provides flexibility in defining the wavelet wavenumber,
 234 as opposed to the Fourier approach where, to employ Fast Fourier Transform (FFT) al-
 235 gorithms, the resolution is constrained to $1/L$ with $L (= 1000 \text{ km})$ being the domain size,
 236 we start by computing the wavelet spectra at the center location $\gamma = \gamma_0 = (y_c, x_c) =$
 237 $(500, 500) \text{ km}$ and use the same wavenumbers as the Fourier spectra (k_F). We see from
 238 Fig. 5 that the agreement between the Fourier and wavelet method is excellent (red solid
 239 and black dashed curves respectively) for both the energy and enstrophy spectra at scales
 240 above the dissipation scale.

241 We also show in Fig. 5 a case where we arbitrarily increase the wavelet wavenum-
 242 ber resolution at scales larger than 50 km where the inverse cascade is expected (black
 243 dashed curve); we take $s_0 = [2\Delta x, \dots, 5x_0]$ monotonically spaced with 30 increments,
 244 which is trimmed for scales smaller than 50 km, and concatenate this with the Fourier
 245 length scales below 50 km. Features at the lowest wavenumbers (i.e. largest spatial scales)
 246 are better captured compared to the red solid and black dashed curves in Fig. 5 where
 247 the Fourier wavenumber resolution is low. This is beneficial as the scales of interest in

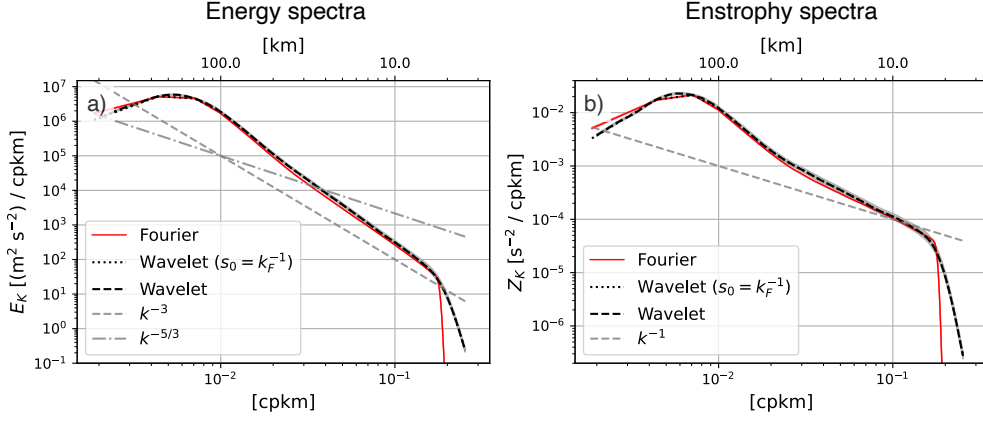


Figure 5. The isotropic (azimuthally-integrated) energy and enstrophy wavenumber spectra of the top layer (a,b). For the wavelet approach, spectra at $\gamma = \gamma_0$ where the wavenumbers are identical to the Fourier wavenumbers ($s_0 = k_F^{-1}$; black dotted) and where the wavenumber resolution is increased at scales larger than 50 km (black dashed) are given. The wavenumber are shown in the lower x axes and corresponding lengthscale in the upper axes. The colored shadings indicate the 95% boot-strapped confidence interval and are shown for the Fourier spectra (red solid) and wavelet spectra with increased wavenumber resolution (black dashed) although the intervals are narrower than the curves themselves.

248 the oceanographic context are often length scales about and larger than the Rossby ra-
 249 dius, associated with mesoscale eddies (Chelton et al., 1998, 2011). The enstrophy spec-
 250 tra are slightly steeper than k^{-1} at scales below the Rossby radius (Fig. 5b), and is con-
 251 sistent with the KE spectral slope also being steeper than -3 . We attribute the steeper
 252 slope to the excessive PV variance introduced by the stochastic forcing cascading down-
 253 scale (cf. Fig. 6b) and the sporadic emergence of coherent structures (e.g. Fig. 3 left panel;
 254 Benzi et al., 1988; Maltrud & Vallis, 1991). While the spectral slopes do not match ex-
 255 actly to what is expected from the inertial range theory, it is known that the spectral
 256 slopes are sensitive to the model configuration of forcing and dissipation (Maltrud & Val-
 257 lis, 1991), and this does not diminish the agreement between the Fourier and wavelet spec-
 258 tral estimates.

259 Using the wavelet transformation, we can also diagnose the KE and enstrophy spec-
 260 tral flux as

$$261 \quad \tilde{\varepsilon}_K(\gamma, \phi, k) = -\frac{1}{C_\Psi} \int_{k>\kappa} \mathcal{R} \left[\overline{\tilde{u}^*(\mathbf{u} \cdot \nabla u)} + \overline{\tilde{v}^*(\mathbf{u} \cdot \nabla v)} \right] x_0^2 \kappa \, d\kappa, \quad (20)$$

$$262 \quad \tilde{\eta}_K(\gamma, \phi, k) = -\frac{1}{C_\Psi} \int_{k>\kappa} \mathcal{R} \left[\overline{\tilde{q}^*(\mathbf{u} \cdot \nabla q)} \right] x_0^2 \kappa \, d\kappa, \quad (21)$$

264 where negative values imply an inverse cascade towards larger scales and positive val-
 265 ues a forward cascade towards smaller scales (Arbic et al., 2013; Khatri et al., 2018).

266 Comparisons of the spectral fluxes computed using wavelets at a single point (black),
 267 standard Fourier spectra (red) and spatial averages of point-wise wavelets (blue) are shown
 268 in Fig. 6a,b. All approaches clearly indicate a broad forward enstrophy cascade range
 269 at scales smaller than the forcing scale/Rossby radius. Similarly, there is general agree-
 270 ment on the existence of an inverse energy cascade in the limited range of scales larger
 271 than the forcing scale. The lower panels in Fig. 6 show the azimuthally-integrated spec-
 272 tral transfers, i.e. the integrand of (20) and (21).

273 In contrast to calculations of the spectra themselves shown in Fig. 5, the spectral
 274 fluxes computed from wavelet data taken at a single spatial point differ significantly from
 275 the global Fourier estimates. As described below, the wavelet spectral flux estimates are
 276 highly sensitive to the amount of spatial and temporal averaging employed, despite the
 277 homogeneity and statistical stationarity of the flow field. This sensitivity arises because
 278 the flux is the transfer cumulatively integrated from the largest wavenumbers towards
 279 smaller wavenumbers (i.e. (20)) so values at high wavenumbers can have a substantial
 280 effect on the flux at low wavenumbers.

281 The 95% boot-strapped confidence intervals, computed by randomly re-sampling
 282 spectral quantities 9999 times, are shown by shading in Fig. 6. In all cases, single-point
 283 wavelet flux and transfer estimates are highly uncertain, while Fourier estimates are not.

284 We argue this dependency on averaging is associated with the fact that the wavelet
 285 estimate of the spectral transfer only incorporates spatially local information while the
 286 Fourier approach effectively yields a domain-averaged estimate. Namely, the global two-
 287 point correlation function, stemming from the assumption of homogeneity in the Fourier
 288 approach, acts as a spatial averaging operator (cf. Uchida, Jamet, et al., 2021). For this
 289 setting, this assumption is valid, hence the superior performance in flux estimation of
 290 the Fourier approach. Note, however, that the transfer estimates emerging from the wavelet

291 approach, while noisy, do largely agree with those of the Fourier approach. It is in the
 292 integration of the transfers where initial noise in the estimates can result in an erroneous
 293 outcome (compare black and blue curves in Fig. 6d).

294 The expectation is that if we were to take the explicit wavelet transform at every
 295 single grid point, the spatial average of the wavelet spectral flux would converge to the
 296 Fourier approach. We examined this by estimating the wavelet spectral flux and trans-
 297 fer at every five grid points in the diagonal direction (i.e. every ~ 14 km) up to 125 grid
 298 points apart (~ 280 km) from the center point (101 locations in total along $y - y_c =$
 299 $\pm(x - x_c)$). The spatial average of them shown as blue curves in Fig. 6 all come closer
 300 to the Fourier estimate than the black curves. Comparison of the domain averaged wavelet
 301 estimates to those derived via standard Fourier approach, both in their mean and con-
 302 fidence intervals, significantly improve when averaged over 101 locations ($\langle \tilde{\varepsilon}_K \rangle, \langle \tilde{\eta}_K \rangle$ where
 303 $\langle \cdot \rangle$ is the averaging operator over 101 locations; blue curves in Fig. 6b,d). The Fourier
 304 and wavelet spectral transfer and flux also no longer differ at the 95% confidence inter-
 305 val.

306 We also exhibit the angular orientation of the spectral flux, which the wavelet ap-
 307 proach can extract via its dependence on the angle ϕ (Fig. 7). The flux shown in Fig. 6a,b
 308 as blue dashed curves are the azimuthal integration of angle-dependent fluxes exhibited
 309 in Fig. 7. As the simulated QG flow is configured to be isotropic, the anisotropy seen
 310 in the spectral flux are statistically insignificant within the 95% boot-strapped confidence
 311 interval; the KE flux exhibits an inverse cascade and enstrophy flux a forward cascade
 312 across all angles (Fig. 7b,d).

313 We are thus led to be cautious interpreting single-point wavelet spectral calcula-
 314 tions when applied to what might be termed higher order quantities, like spectral flux.
 315 However, we also point out this is a sword that cuts in both directions. The accuracy
 316 of the Fourier flux estimates depends strongly on their area-wide integrative effect in this
 317 homogeneous setting. Were the flow not homogeneous, the integrative character of the
 318 Fourier approach would obscure the meaning of the result.

319 **3.2 Spectra over a non-periodic subdomain**

320 We now examine the spectra taken over the subdomain given by $y = 200 - 800$ km
 321 and $x = 200 - 800$ km in anticipation of realistic data where periodicity is never satis-

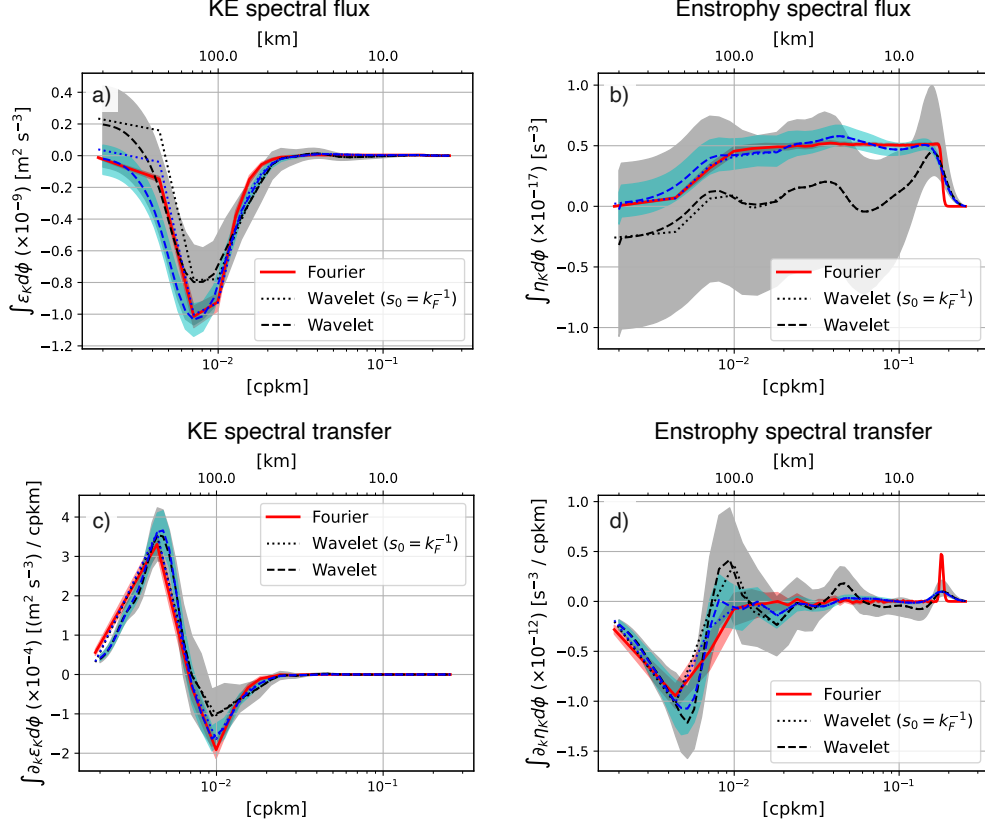


Figure 6. The isotropic (azimuthally integrated) KE and enstrophy wavenumber spectral flux (a,b) and transfer (c,d) respectively. The Fourier method is shown in red and the wavelet approach at $\gamma = \gamma_0$ with wavenumbers identical to the Fourier wavenumbers in dotted ($s_0 = k_F^{-1}$) and the case with increased wavenumber resolution at smaller wavenumbers in dashed curves respectively. The black curves show the wavelet flux and transfer at $\gamma = \gamma_0$, while the blue curves show them averaged over the 101 locations ($\langle \tilde{\varepsilon}_K \rangle, \langle \tilde{\eta}_K \rangle$). The colored shadings indicate the 95% bootstrapped confidence interval and are shown for the Fourier spectra (red solid) and wavelet spectra with increased wavenumber resolution (black and blue dashed curves). The wavenumber are shown in the lower x axes and corresponding lengthscale in the upper axes.

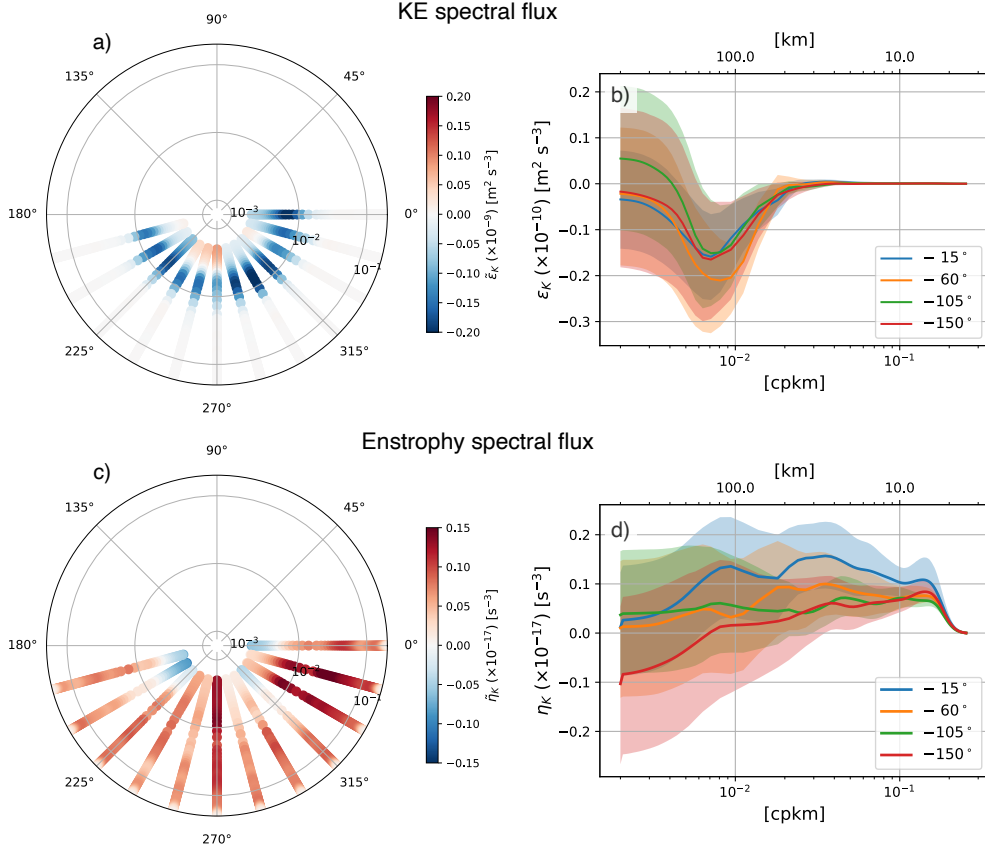


Figure 7. The angular dependence of the KE and enstrophy spectral flux from the wavelet approach plotted radially average over the 101 locations ($\langle \tilde{\varepsilon}_K(\phi, k) \rangle$, $\langle \tilde{\eta}_K(\phi, k) \rangle$); a,c). The radial axes are the wavenumbers in logarithmic scaling with the increased wavenumber resolution. The fluxes are symmetric about the origin so we only show for angles $[0, -\pi)$. The 95% boot-strapped confidence intervals are given for a four arbitrary angles (b,d).

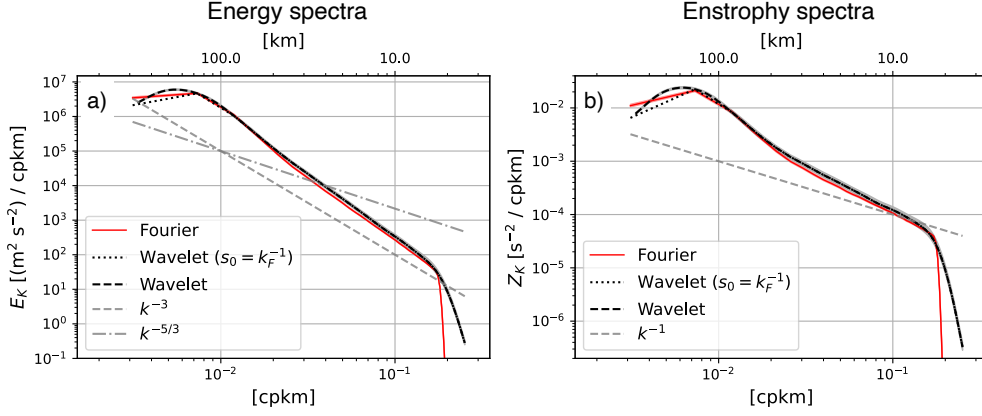


Figure 8. Same as Fig. 5 but for the subdomain of $y = 200\text{-}800$ km and $x = 200\text{-}800$ km.

The confidence intervals are again narrower than the curves themselves.

322 fied. As the data are no longer periodic, the Fourier approach requires the data to be
 323 windowed. This will highlight the strength of the locality in the wavelet approach where
 324 windowing of the data is unnecessary. Prior to taking the Fourier transforms, we applied
 325 Hann windows (Arbic et al., 2013; Uchida et al., 2017; Uchida, Jamet, et al., 2021) and
 326 then corrected for their amplitude. Comparing Figs. 5 and 8, we see that the spectral
 327 estimates are still robust. The low resolution at lower wavenumbers from the Fourier method
 328 makes it difficult to detect the spectral shape at scales above ~ 100 km due to the do-
 329 main size being small; there are only two wavenumber points at scales larger than the
 330 Rossby radius (red curves in Fig. 8). The wavelet approach, on the other hand through
 331 its flexibility in s , still captures a smooth spectral estimate; the scaling for the wavelet
 332 approach was adjusted to $s_0 = [2\Delta x, \dots, 3x_0]$ over 30 monotonic increments in order to
 333 account for the smaller domain and then replaced by the Fourier wavenumber at scales
 334 smaller than 50 km.

335 Regarding the spectral transfer, the Fourier approach is significantly affected by
 336 the tapering at the lowest wavenumbers (red curves in Figs. 6c,d and 9c,d) but the wavelet
 337 approach is still able to capture the change in sign in its curvature (blue dashed curves
 338 in Figs. 6c,d and 9c,d). The enstrophy spectral flux and transfer tend to be particularly
 339 sensitive to the local nature of wavelet transforms. Similar to Fig. 6, we see a significant
 340 improvement in the wavelet estimates both in their mean and confidence intervals when
 341 averaged over 101 locations, particularly for enstrophy ($\langle \tilde{\eta}_K \rangle$; blue curves in Fig. 9b,d).

342

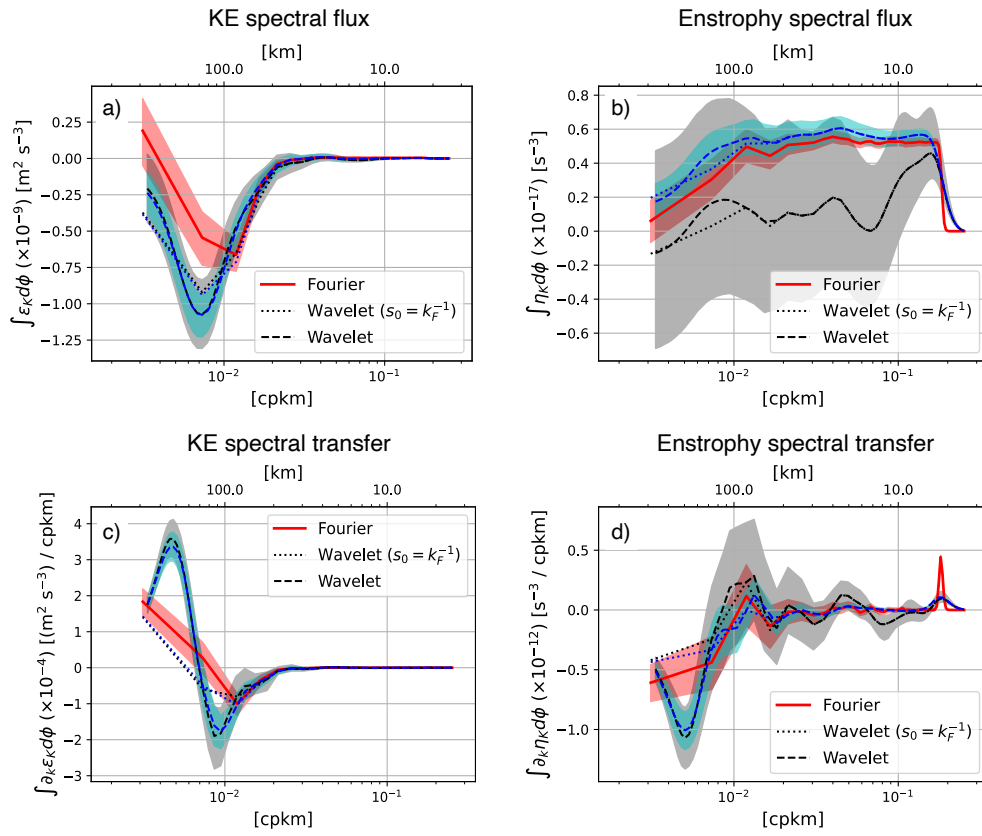


Figure 9. Same as Fig. 6 but for the subdomain of $y = 200\text{-}800$ km and $x = 200\text{-}800$ km.

343 4 Conclusions and discussion

344 In this study, we have described and documented a wavelet-based technique for spec-
 345 tral analyses in an oceanographic context. The wavelet approach employed here, through
 346 its dependence on a scale parameter s , returns effectively a one-dimensional (1D) spec-
 347 tral estimate, and its incorporating of two-dimensional data allows for information re-
 348 garding local anisotropies through its angular dependency ϕ (Fig. 7).

349 We have demonstrated its utility by applying it to a doubly-periodic, two-layer, quasi-
 350 geostrophic (QG) simulation. The flow analyzed in this study is highly idealized being
 351 spatially isotropic and homogeneous in the horizontal dimensions. The idealized setting,
 352 however, is expected to yield known spectral cascades, so it can be used as a test bed
 353 for the wavelet approach. The agreement between the wavelet and Fourier approach, par-
 354 ticularly for the spectra (Figs. 5 and 8), encourages the usage of wavelets with its ad-
 355 ditional strengths of being able to capture the local features of the flow. While numer-
 356 ically efficient algorithms exist to take the wavelet transform (coined as Fast Wavelet Trans-
 357 forms; e.g. Beylkin et al., 1991), they face the same conundrum as FFTs requiring: i)
 358 periodic boundary conditions, and ii) filling in missing data points. We have, therefore,
 359 taken the approach of explicitly computing the wavelet transform (7), which negates the
 360 two necessities and will benefit realistic settings such as the North Atlantic basin. The
 361 robustness of the spectra is comforting, but we also emphasize the need for caution when
 362 computing higher-order spectral quantities, like spectral fluxes, which involve spatial deriva-
 363 tives. The disagreement arises from the local nature of wavelets; the Fourier method in-
 364 corporates spatially global information and hence can be thought as a spatial average
 365 of spectral estimates. This is evident from the fact that upon spatially averaging the wavelet
 366 spectral transfer over multiple locations, the confidence interval improved and its mean
 367 converged towards the Fourier estimate.

368 Our work is complementary to recent work by Aluie et al. (2018); Sadek and Aluie
 369 (2018); Schubert et al. (2020); Storer et al. (2022) and Contreras et al. (2022) where they
 370 use a spatial filter to examine the KE spectra and cross-scale transfer, Lindborg (2015);
 371 Balwada et al. (2016, 2022); LaCasce (2016); Poje et al. (2017) and Pearson et al. (2020)
 372 where they implement structure functions, Jamet et al. (2020) where they employ the
 373 Green’s function, and Uchida, Jamet, et al. (2021) where they use Empirical Orthogo-
 374 nal Functions. Barkan et al. (2021) and Srinivasan et al. (2022) apply the filtering method

375 in both the spatiotemporal dimensions. Liang and Anderson (2007); Liang (2016) and
376 Yang et al. (2021) are also interesting attempts in implementing a multiscale window trans-
377 form to examine the energy exchange across spatiotemporal scales by decomposing the
378 flow with a set of orthogonal windows. Here, we have documented the wavelet-based cross-
379 scale energetics in the spectral context. While the form of the Parseval’s equality will
380 slightly change, namely in the power of scaling s , the wavelet method can also be extended
381 to estimating frequency-wavenumber spectra (e.g. Torres et al., 2018; Uchida et al., 2019);
382 this will allow us to decompose the balanced and unbalanced motions in non-periodic
383 settings.

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396 **Data availability statement**

397 The wavelet transforms were taken using the `xwavelet` Python package (Uchida
398 & Dewar, 2022) and Fourier transforms using the `xrft` Python package (Uchida et al.,
399 2022). The `pyqg` model is available through Github (Abernathey et al., 2022, [https://](https://github.com/roxyboy/pyqg/tree/ring)
400 github.com/roxyboy/pyqg/tree/ring). Jupyter notebooks used to run the `pyqg` sim-
401 ulation and conduct analyses are available via Github ([https://github.com/roxyboy/](https://github.com/roxyboy/QG-wavelets)
402 [QG-wavelets](https://github.com/roxyboy/QG-wavelets); a DOI will be added upon acceptance of the manuscript).

Appendix A Parseval's equality

In this appendix, we review the Parseval's equality for two-dimensional wavelet transforms (Daubechies, 1992; Torrence & Compo, 1998; Mallat, 1999; Chen & Chu, 2017),

i.e.

$$\int f g d\mathbf{x} = \frac{1}{C_{\Xi}} \iiint \tilde{f} \tilde{g}^* \frac{1}{s^3} d\gamma ds d\phi. \quad (\text{A1})$$

Using (9), the right-hand side can be expanded as

$$\begin{aligned} \iiint \tilde{f} \tilde{g}^* \frac{1}{s^3} d\gamma ds d\phi &= \iiint \frac{1}{s^5} \int_{\mathbf{x}} f(\mathbf{x}) \xi^*(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x} - \gamma}{s}\right)) d\mathbf{x} \int_{\boldsymbol{\chi}} g^*(\boldsymbol{\chi}) \xi(\mathbf{R}^{-1} \cdot \left(\frac{\boldsymbol{\chi} - \gamma}{s}\right)) d\boldsymbol{\chi} d\gamma ds d\phi \\ &= \int_{\mathbf{x}} \int_{\boldsymbol{\chi}} f g^* \iiint \frac{1}{s^5} \xi^*(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x} - \gamma}{s}\right)) \xi(\mathbf{R}^{-1} \cdot \left(\frac{\boldsymbol{\chi} - \gamma}{s}\right)) d\gamma ds d\phi d\mathbf{x} d\boldsymbol{\chi}. \end{aligned} \quad (\text{A2})$$

Now, consider the wavelet transform of the Dirac delta function

$$\begin{aligned} \tilde{\delta} &= \int_{\mathbf{x}} \delta(\mathbf{x} - \mathbf{x}') \frac{1}{s} \xi^*(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x} - \gamma}{s}\right)) d\mathbf{x} \\ &= \frac{1}{s} \xi^*(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x}' - \gamma}{s}\right)). \end{aligned} \quad (\text{A3})$$

Hence, the inverse wavelet transform becomes

$$\begin{aligned} \delta &= \frac{1}{C_{\Xi}} \iiint \tilde{\delta} \frac{1}{s^4} \xi(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x} - \gamma}{s}\right)) d\phi ds d\gamma \\ &= \frac{1}{C_{\Xi}} \iiint \frac{1}{s^5} \xi^*(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x}' - \gamma}{s}\right)) \xi(\mathbf{R}^{-1} \cdot \left(\frac{\mathbf{x} - \gamma}{s}\right)) d\phi ds d\gamma. \end{aligned} \quad (\text{A4})$$

Plugging (A4) into (A2) yields

$$\iiint \tilde{f} \tilde{g}^* \frac{1}{s^3} d\gamma ds d\phi = C_{\Xi} \int_{\mathbf{x}} f g^* d\mathbf{x}, \quad (\text{A5})$$

and we obtain (A1).

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