Tremor bursts reveal characteristics of slow-slip events that episodically load the shallow seismogenic zone of the San Andreas fault

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Strain accumulated on the deep extension of some faults are episodically released during transient slow-slip events which can subsequently load the shallow seismogenic region. At the San Andreas fault, the characteristics of slow-slip events are difficult to constrain geodetically due to their small deformation signal. Slow-slip events are often accompanied by coincident tremor bursts composed of many low-frequency earthquakes. Here we probabilistically estimate the spatiotemporal clustering properties of low-frequency earthquakes detected along the central San Andreas fault. We find that the tremor bursts follow a power-law spatial and temporal decay similar to earthquake after-shock sequences. The low-frequency earthquake clusters reveal that the underlying slow-slip events have two modes of rupture velocity. Compared to regular earthquakes, the slow-slip events have smaller stress drops and rupture velocities but follow similar magnitude-frequency, moment-area, and moment-duration scaling. Our results connect a broad spectrum of transient fault slips

#### with velocities spanning several orders of magnitude.

## Introduction

The establishment of continuous global positioning system (GPS) measurements led to the discovery of slow-slip events (SSEs) down dip of the seismogenically-locked region of the Nankai and Cascadia subduction zones (1–2). Subsequently, tectonic tremors that correlate spatially and temporally with SSEs were discovered (3–4). These long-duration tremors have been inferred to be the superposition of many low-frequency earthquakes (LFEs) that represent asperities that were repeatedly driven to failure by surrounding aseismic slip (5). This interpretation that LFEs are markers of SSEs is supported by recent studies that managed to extract slow slip deformation signal by using the timing of increased LFE rate as a guide to stack GPS time series (6–7).

Current surface geodetic measurements have shown to only be able to detect SSEs above moment magnitude  $(M_w) \sim 6$  (8). Along the central San Andreas fault, tremors and LFEs have long been observed (9-10) but it was only recently that the deformation signature of  $M_w$  4.9 SSEs were detected after stacking 20 such events using the timing of increased LFE rate (7). Since these SSEs in the deeper part of the fault might be episodically loading the shallow region that last ruptured in the 1857 magnitude 7.9 Fort Tejon earthquake (11), it is important that we have a more comprehensive catalog of the SSEs. The bias towards only detecting the largest SSEs also limits our ability to robustly characterize their scaling properties. For instance, it is still debated whether the moment  $(M_0)$  of SSEs is proportional to their duration (T) (12) or follows the  $T^3$  scaling observed for earthquakes (13–14). A  $M_0 \propto T$  scaling relationship for SSEs could reflect fundamentally different underlying dynamics compared to regular earthquakes (12), or simply be the by-product of only cataloging the largest 'bounded' events with aspect ratios > 1 (15).

Previous studies have used tremors and LFEs to characterize SSEs invisible to geodetic measurements. Along the central San Andreas fault, Shelly (10) cataloged 88 LFE families (Fig. 1) that represent groups of LFEs with similar waveforms and hence similar source mechanisms and locations. Thomas et al. (16) clustered the LFEs of each individual family base on their recurrence intervals using empirically-derived separation timescales while Lengliné et al. (17) proposed a stochastic model that successfully reproduced the temporal-clustering behavior of the LFEs. However, these studies did not account for interactions between LFE families and hence could not directly estimate the spatial properties of the underlying SSEs from the LFE clusters, even though the occurrence patterns of different families have been shown to correlate (Fig. 1c) (18–19). Clustering of tremors and LFEs that takes into account their spatial relationships have been attempted to identify secondary slip fronts at subduction zones but relied on rather ad hoc definitions of what constitutes a cluster (20–21).

In this paper, we estimate the spatiotemporal clustering properties of LFEs detected along the central San Andreas fault (10) probabilistically with minimal model assumptions and a priori parameterization. We then use the extracted LFE clusters to estimate the area, average slip, moment, duration, stress drop, and rupture velocity of the underlying SSEs and explore their scaling properties.

#### **Results**

#### **Clustering properties of LFEs**

We analyze the catalog of 88 LFE families that span  $\sim 150$  km along the central San Andreas fault (Fig. 1) (10) which includes more than 1 million events from 2001 to 2016. We limit our analysis to between 2006 and 2016 to minimize the impact of the 2004  $M_w$ 6 Parkfield earthquake. This leaves us with a catalog of  $\sim 750,000$  LFEs. The most populous family has  $\sim 22,000$  events while the least populous family has  $\sim 2,500$  events over this time period. We

model the LFE rate at time t as

$$\lambda(t) = \sum_{x=1}^{D} \left( \mu^x + \sum_{y=1}^{D} \sum_{t_j^y < t} K^{xy} g(t - t_j^y) \right), \tag{1}$$

where D is the number of LFE families,  $\mu^x$  is the uniform background rate of family x,  $K^{xy}$  encodes the excitation strength i.e. the number of events in family x 'excited' by an event in family y on average, g is the normalized time-dependent excitation kernel, and  $t_j^y$  is the occurrence time of LFE j from family y. Therefore, we are only assuming that (i) the LFE clustering behavior is linear i.e. the contribution of different LFEs can be added up, (ii) a mean-field response to the occurrence of an LFE i.e. two LFEs from the same family are modeled similarly, and (iii) the excitation kernel g(t) does not vary between different families, though we obtain similar results when relaxing this assumption (see Materials and Methods). Similar models have been used to characterize the clustering properties of regular earthquakes, infectious diseases, and crime (22). We discretize g(t) as piece-wise constant and adopt an Expectation-Maximization approach (23–24) to estimate the parameters  $\mu^x$ ,  $K^{xy}$ , and g(t) (see Materials and Methods). We used a synthetic catalog to verify the algorithm's ability to correctly estimate the model parameters (see Materials and Methods).

g(t) characterizes the temporal-clustering property of the LFEs. We find that g decays with time, with the decay well-approximated by a power law up to 10 days, but with a plateau between 0.02 and 0.2 days (Fig. 2a). We fit the power law decay in the range  $[2 \cdot 10^{-4} - 2 \cdot 10^{-2}]$  days as well as between 0.2 and 10 days separately, and obtain  $g \propto t^{-1.8}$  for both time ranges. The shape of g(t) and the power-law exponent that we obtain are similar to what Lengliné et al. (17) obtained from applying a comparable stochastic model to one LFE family at a time, as well as the stacked inter-event-time density of the different LFE families (Fig. S1). Therefore, the power-law decay of g, which is similar to the temporal evolution of earthquake aftershock sequences, is a robust feature of the LFE catalog. Lengliné et al. (17) concluded that this feature

is unlikely to have arose from direct triggering of LFEs by the stress change due to a preceding LFE since the number of excited LFEs is not correlated with the amplitude of the 'mainshock' LFE in their analysis. Therefore, the power-law decay of g likely reflects triggering of LFEs by changes in stress or loading rate due to the underlying SSE. The plateau of g between 0.02 and 0.2 days (30 minutes and 5 hours) is rather perplexing but might indicate that there are two clustering timescales. When we allow g(t) to vary between LFE families (see Materials and Methods, Fig. S2a), we find that g(t) has such a plateau only for LFE families that have short-duration bursts occurring within long-duration bursts i.e. a trimodal interevent time distribution (Fig. S3) (16), similar to the secondary slip fronts observed within SSEs at the Cascadia and Nankai subduction zones (20, 25).

 $K^{xy}$  characterizes the excitation strength between LFE families and thus the spatial-clustering property of the LFEs. We find that a family that was previously suggested to be isolated (Fig. 1) (26) has the second smallest inter-family K value, lending confidence that K indeed captures the interaction between LFE families. On average, K decays with inter-family distance (Fig. 2b), even though inter-family distance is not present in our model (Eq. 1). The decay of K with distance is consistent with previous observations that the occurrence pattern of nearby LFE families are correlated (18–19). For along-strike distance, the decay of K is well-approximated by a power law up to 16 km, above which the value of K saturates possibly due to the values being too small to be resolvable (Fig. 2b). We fit the power-law decay between 1 and 16 km and obtain  $K \propto d^{-2.8}$ . For along-dip distance, we obtain  $K \propto d^{-2.5}$ . The power-law exponent of  $K \sim 3$  is similar to the expected earthquake aftershock density decay with distance from the mainshock if aftershocks are triggered by static stress change. The fast decay of  $K \sim 15$  km when grouping the LFE families based on the similarity of their long-timescale occurrence patterns. Inter-family excitation in the along-strike direction is about 10 times stronger com-

pared to the along-dip direction (Fig. 2b), which suggests that the underlying SSEs tend to propagate along-strike. However, we find that the decay of K is similar for both southeast (the excited family is located southeast of the exciting family) and northwest excitation (Fig. 2b), even though previous studies have suggested that earthquakes along the Parkfield segment of the San Andreas fault preferentially rupture to the southeast (e.g. 27) potentially due to the fault being a bimaterial interface (28).

#### **Estimating properties of underlying SSEs**

Now that we have estimated the parameters  $\mu$ , K, and q which govern the LFE rate at any given time, for every LFE i, we can calculate the probability that it is a background event and the probability that it was 'excited' by each preceding LFE j (see Materials and Methods). Using the probabilities associated with each LFE, we perform stochastic clustering (29) of the events to isolate individual LFE bursts. Each cluster includes one background event and the events it directly and indirectly 'excites'. We interpret this background event to mark the initiation of an SSE, while the excited events reflect the subsequent evolution of the SSE. Stochastic clustering is not unique and each catalog that it produces represents a sampling of the underlying structure. Here we present a catalog from one iteration, which includes 16,327 LFE clusters that involved ≥ 2 LFE families, but have verified that the presented statistics remain stable between different iterations as expected. For each of these clusters, we calculate their along-strike extent (L) and depth extent (W). We then infer the underlying SSE to have a rupture area A = LW (Fig. S4). The time difference between the first and last LFE in a cluster is taken as the SSE duration (T). We then estimate the average rupture velocity  $V_r = \frac{L}{T}$  which assumes unilateral propagation. For a rupture that starts at the middle and propagates bilaterally,  $V_r = \frac{L}{2T}$ , so our estimated rupture velocity is a proxy up to a factor of 2. While LFEs are often interpreted as asperities that are repeatedly driven to failure by surrounding aseismic slip (5), our estimating these SSE

properties using LFE clusters only requires that they are spatially and temporally coincident (7) and is valid regardless of the physical mechanism behind their coincidence.

Thomas et al. (16) have shown that LFEs on the San Andreas fault can be used to meter slip. The long-term slip rate of the 150-km-long creeping segment of the San Andreas fault is approximately 34 mm/yr (e.g. 30). For each LFE family x, we estimate the slip per LFE  $(d_x)$  by multiplying the long-term slip rate by the catalog duration (10 years) before dividing by the total number of LFEs in that family. For each LFE cluster, we can then estimate the total slip of each LFE family x by multiplying  $d_x$  by the number of LFEs of that family that is part of the cluster. We assume the total slip of each LFE family represents a point sample of the slip distribution over the rupture area of the underlying SSE i.e. there is slip throughout the area delineated by each LFE cluster but we only have an estimate of the slip amount where the LFE families are located. For each LFE cluster that involved  $\geq 2$  LFE families, we then take the mean slip of all the activated LFE families as the average slip  $(\overline{D})$  of the underlying SSE (Fig. S4) and calculate seismic moment  $M_o = \mu A \overline{D}$ , taking the shear modulus  $\mu = 30 \text{GPa}$ . While we are assuming that there are no aseismic slip on the LFE patches, if aseismic slip is a constant factor relative to seismic slip, it will just scale our  $M_o$  estimates by a constant factor and would not affect its inferred scaling relationship with other SSE properties. We calculate moment magnitude  $M_w=\frac{2}{3}(\log M_o-9.1)$  and find the 20 largest SSEs have a mean  $M_w$  of 5.1. This is just slightly larger than the geodetically-determined average  $M_{\rm w}$  of 4.9 for the SSEs associated with the 20 largest bursts of 10 LFE families on the NW Parkfield segment of the San Andreas fault (7), suggesting that our derived SSE moments are reasonable.

#### SSE scaling properties

Earthquakes follow a power-law magnitude-frequency distribution given as  $log_{10}(N) = a - bM$ , where N is the number of earthquakes of magnitude  $\geq M$ , and a and b are constants (31). We find

that the SSEs also approximately follow a power-law magnitude-frequency distribution (Fig. 3a). We obtain a b value of  $1.61 \pm 0.04$  (32-33) when taking a magnitude of completeness  $M_c = 3.9$ . Non-volcanic tremors that are approximately spatially-coincident with our LFE families in the southeast have been suggested to also follow a power-law magnitude-frequency distribution with a b value of 2.50 (34). The same study found that regular earthquakes on the creeping section of the San Andreas fault, which are on the shallower portion of the fault above our LFE families in the northwest, have 0.8 < b < 2.0. LFEs and SSEs at the Cascadia subduction zone have also been suggested to follow a power-law magnitude-frequency distribution, with the LFEs having b values  $\geq 5$  (35) while the SSEs have b values between 0.8-1.0 (14, 36). The differences in b values might reflect differences in stress condition (37) and fault roughness (38).

We find that the SSE moment and rupture area approximately follows the  $M_o \propto A^{1.5}$  scaling of regular earthquakes (Fig. 3b). Since we estimated  $M_o$  from A and  $\overline{D}$ , this is a direct result of  $\overline{D} \propto \sqrt{A}$  (Fig. S5a). There is deviation from this relationship at approximately  $M_o < 10^{13.5}$  (Fig. 3b), probably due to the poorer rupture area estimates for these smaller events since their length and width are comparable to the LFE location uncertainties of 1-2 km (I0). For bins of  $M_o > 10^{13.5}$  N·m, we calculate the median rupture area. Fitting these median values, we obtain  $M_o \propto A^{1.5}$  (Fig. 3b). This is consistent with our SSEs being mostly 'unbounded' events (I5) where both the length and width are still growing as the rupture propagates (Fig. S5b), although these events show a greater tendency to rupture in the along-strike direction (Fig. 2b) with the median SSE aspect ratio  $\frac{L}{W} = 2$ . For the largest events, W eventually saturates while L and  $\overline{D}$  continue to grow (Fig. S5b). This maximum W of  $\sim 13$  km (Fig. S5b) could reflect the maximum channel width for SSEs at the central San Andreas fault as determined by rheological change with depth, or simply a coincidence of asperities' distribution that resulted in the LFEs being located in a narrow 13-km band (Fig. 1b) even though SSEs can rupture beyond this zone.

The SSE stress drops are mostly within the range of 1-10 kPa (Fig. 3b), with a median value of 6 kPa, based on a circular crack model (39). This is a few orders of magnitude smaller than the stress drops of regular earthquakes on the San Andreas fault which are estimated to be on the order of 0.1 - 100 MPa (e.g. 40). The SSEs have a median stress drop of 2 kPa when assuming strike-slip faulting on a rectangular fault instead (41). Our results are comparable to the estimated stress drop of  $\sim$ 10 kPa from spectral analysis of two LFE families on the San Andreas fault (42) and is consistent with the low stress drops on the order of 10 kPa typically obtained for slow earthquake phenomena (e.g. 12, 14, 21, 35).

The logarithm of the SSE duration has a bimodal distribution (Fig. 3c) with a clear separation at approximately  $10^{3.5}$  seconds (84 minutes), which is in the time range when q has a plateau (Fig. 2a). This translates to a bimodal distribution in the rupture velocity, with the two modes at 6 and 708 km/day. These velocities are within the range of tremor migration velocities previously observed at the San Andreas fault (18) as well as at the Nankai (5) and Cascadia subduction zones (21, 43). We use  $T = 10^{3.5} s$  as a boundary to separate the SSEs into two populations. For the shorter-duration population, we calculate the median T for  $M_o$  bins in the range  $[10^{11} - 10^{15}]$  Nm. For the longer-duration population, we calculate the median T for  $M_o$ bins in the range  $[10^{12.5}-10^{16.5}]$  Nm. Fitting these median values, we obtain  $M_o \propto T^{3.1}$  and  $M_o \propto T^{2.8}$  for the shorter and longer-duration SSE populations respectively (Fig. 3d). Recently, Michel et al. (14) suggested that 64  $M_w$  5.3-6.8 SSEs cataloged by inverting GPS measurements at the Cascadia subduction zone likely follows a  $M_o \propto T^3$  relationship (their best-fit value is  $M_o \propto T^5$ ), while Frank and Brodsky (13) observed a similar scaling relationship for SSEs at the Mexican subduction zone when calibrating LFE amplitudes to GPS-measured moment rate of previously recorded SSEs to estimate the moment of smaller SSEs. The fact that three different methods applied to three different regions obtained similar results suggests that it is likely a universal property that SSEs follow a  $M_o \propto T^3$  scaling like regular earthquakes (39) instead of

the  $M_o \propto T$  scaling proposed for slow earthquakes (12), when the events are 'unbounded' (15). The previously apparent  $M_o \propto T$  scaling is likely an artifact of documenting only the largest slow events with different rupture velocities (Fig. 4).

#### **Conclusions**

We have shown that LFEs can be used to catalog the large number of SSEs that are episodically loading the shallow seismogenic zone. We find that the LFEs' clustering properties are similar to earthquake aftershock sequences. The underlying SSEs have two modes of rupture velocity with stress drops and rupture velocities a few orders of magnitude smaller than regular earthquakes. However, the SSEs follow similar magnitude-frequency, moment-area, and moment-duration scaling as regular earthquakes, suggesting that transient fault slips with velocities spanning many orders of magnitude may be governed by the same driving mechanism. Our observations provide important constraints on the relationships between LFEs, SSEs, and regular earthquakes.

#### **Materials and Methods**

## **Modeling LFE rate**

For the most generalized model, the LFE rate at time t can be modeled as

$$\lambda(t) = \sum_{x=1}^{D} \left( \mu^x + \sum_{y=1}^{D} \sum_{t_j^y < t} g^{xy} (t - t_j^y) \right), \tag{1}$$

where D is the number of LFE families,  $\mu^x$  is the uniform background rate of LFE family x,  $g^{xy}$  is the time-dependent excitation kernel that encodes the influence of an LFE from family y on the future rate of family x, and  $t_j^y$  is the occurrence time of LFE j from family y. g(t) can be discretized as piece-wise constant:

$$g_m = g(T_m < t < T_{m+1}) (2)$$

where  $T_m$  are the discretization times and  $m \in [1:M]$  with  $T_1 = 0$  and  $T_M = 10$  days. For M = 20 and D = 88, we would have to solve for  $\sim 155,000$  parameters. Therefore, we simplify the model by assuming that g(t) does not vary between different families such that

$$\lambda(t) = \sum_{x=1}^{D} \left( \mu^x + \sum_{y=1}^{D} \sum_{t_j^y < t} K^{xy} g(t - t_j^y) \right), \tag{3}$$

where  $K^{xy}$  encodes the excitation strength i.e. the number of events in family x excited by an event in family y on average, and g(t) is normalized such that it represents a probability density function:

$$\sum_{m=1}^{M} g_m \delta t_m = 1,\tag{4}$$

where  $\delta t_m$  is the discretization time step of  $g_m$ , i.e.,  $\delta t_m = T_{m+1} - T_m$ . With this model, we only have to solve for  $\sim 8,000$  parameters.  $\sum_{x,y} K^{xy}$  has to be < D for the model to be stable;  $\sum_{x,y} K^{xy} > D$  would result in an infinite number of events within a finite time period, i.e.,  $\lambda$  growing to infinity. We adopt an Expectation-Maximization approach to estimate the parameters  $\mu^x$ ,  $K^{xy}$ , and  $g_m$  (23–24). The Expectation step involves computing the probability that event i from family x was excited by event y from family y:

$$p_{ij}^{xy} = \frac{K^{xy}g(t_i - t_j)}{\mu^x + \sum_{z=1}^D \sum_{t_i^z < t_i} K^{xz}g(t_i - t_l^z)},$$
 (5)

where  $\{t_l^z < t_i\}$  are all events occurring before  $t_i$ , and the probability that event i from family x was a background event (not excited by any previous events):

$$p_{i0}^{x} = \frac{\mu^{x}}{\mu^{x} + \sum_{z=1}^{D} \sum_{t_{i}^{z} < t_{i}} K^{xz} g(t_{i} - t_{i}^{z})},$$
(6)

such that

$$p_{i0}^x + \sum_{j,y} p_{ij}^{xy} = 1. (7)$$

The log-likelihood function associated with Eq. 3 is thus

$$L = \sum_{x=1}^{D} \left[ \sum_{i=1}^{N} p_{i0}^{x} \log \mu^{x} - \mu^{x} T + \sum_{y=1}^{D} \sum_{m=1}^{M} \left[ \sum_{i,j \in A_{m}^{xy}} p_{ij}^{xy} \log K^{xy} g_{m} - n_{y} K^{xy} g_{m} \delta t_{m} \right] \right], \quad (8)$$

where T is the duration of the time series,  $A_m^{xy}$  is the set of pairs of events such that  $(t_i^x - t_j^y) \in (T_{m+1} - T_m)$ , and  $n_y$  is the number of events from family y.

The Maximization step then involves updating the background rates as

$$\frac{\partial L}{\partial \mu^x} = 0 \quad \Leftrightarrow \quad \mu^x = \frac{1}{T} \sum_{i=1}^N p_{i0}^x, \tag{9}$$

the excitation kernel as

$$\frac{\partial L}{\partial g_m} = 0 \quad \Leftrightarrow \quad g_m = \frac{\sum_{x=1}^D \sum_{y=1}^D \sum_{i,j \in A_m^{xy}} p_{ij}^{xy}}{\delta t_m \sum_{x=1}^D \sum_{y=1}^D n_y K^{xy}},$$
(10)

and the excitation strengths as

$$\frac{\partial L}{\partial K^{xy}} = 0 \quad \Leftrightarrow \quad K^{xy} = \frac{\sum_{i,j} p_{ij}^{xy}}{n_y \sum_{m=1}^{M} g_m \delta t_m} = \frac{\sum_{i,j} p_{ij}^{xy}}{n_y} \tag{11}$$

as a result of normalizing g following Eq. 4. We start with initial guesses of  $\mu^x$  (uniform values of 1 for all families),  $g_m$  (random values between 0 and 1), and  $K^{xy}$  (random values between 0 and 1). We then iterate through the Expectation-Maximization steps until the difference of the estimated parameters between two successive iterations is smaller than a certain threshold. We obtain  $\sum_{x,y} K^{xy} = 81.6$ . We generate a 10-year synthetic catalog with the g,  $\mu$ ,

and  $K^{xy}$  that we obtained for the San Andreas LFE catalog, with  $K^{xy}$  first multiplied by 0.8 to test if we can resolve parameters from a catalog with weaker inter-event excitation. We find that we can correctly estimate the model parameters from the synthetic catalog using the proposed Expectation-Maximization algorithm (Fig. S6).

Since it has been observed that different LFE families have different interevent time distribution (16), we explored having one g(t) per family such that

$$\lambda(t) = \sum_{x=1}^{D} \left( \mu^x + \sum_{y=1}^{D} \sum_{t_j^y < t} K^{xy} g^x (t - t_j^y) \right), \tag{12}$$

where  $g^x$  is the time-dependent excitation kernel that encodes the influence of an LFE on the future rate of family x. With this model, we have to solve for  $\sim 9,500$  parameters and obtain  $\sum_{x,y} K^{xy} = 83.3$ . While  $g^x(t)$ 's shape varies depending on the LFE family's inter-event-time distribution (Fig. S3), we obtain similar scaling properties (Fig. S2b and S7) as when using Eq. 3, except that we now obtain  $M_o \propto T^{4.2}$  for the shorter-duration SSE population (Fig. S7d).

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# **Figures**

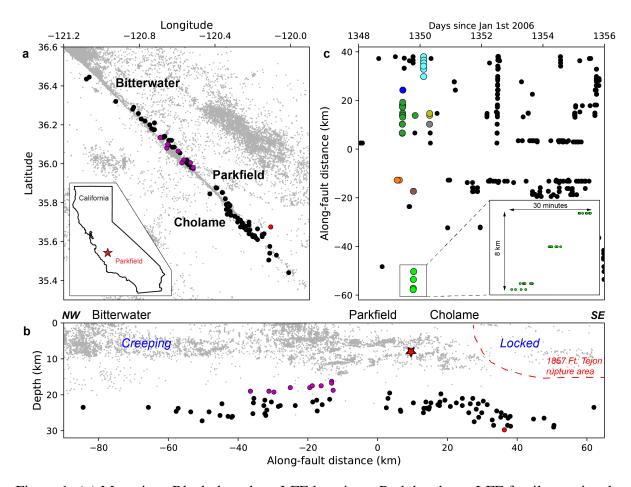


Figure 1: (a) Map view. Black dots show LFE locations. Red dot shows LFE family previously suggested to be isolated (26). Magenta dots show 10 LFE families that Rousset et al. (7) successfully used as guide to stack GPS time series and extract slow-slip deformation signal. Grey dots show earthquakes from 1984 to 2011 (44). (b) Along-fault cross-section. Red star shows epicenter of  $M_w$  6.0 Parkfield earthquake. (c) Space-time plot of the LFEs. Black dots show all events. Colored dots show example clusters - events of the same color belong to the same cluster.

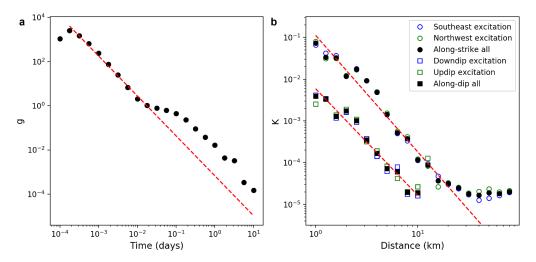


Figure 2: (a) g as a function of time. Dashed red line shows least-squares fit giving  $g \propto t^{-1.8}$ . (b) K as a function of distance. Since the absolute K value is affected by the LFE family size, we account for this effect by multiplying  $K^{xy}$  by  $\frac{n_y}{n_x}$ , where  $n_x$  is the number of events from the excited family x and x0 is the number of events from the exciting family x0, before calculating the mean x0 value for different inter-family distance bins. Dashed red lines show least-squares fits giving x1 giving x2 distances respectively.

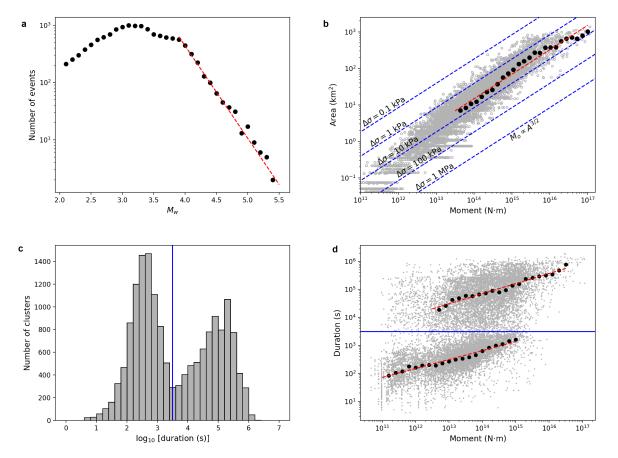


Figure 3: (a) Non-cumulative magnitude-frequency distribution of the SSEs. Black circles show the number of SSEs in different  $M_w$  bins. Dashed red line shows maximum-likelihood fit giving a b value of 1.61. (b) Relationship between seismic moment  $(M_o)$  and rupture area (A) of the SSEs. Black circles show the median A for different  $M_o$  bins. Dashed red line shows least-squares fit giving  $M_o \propto A^{1.5}$ . Dashed blue lines show the  $M_o \propto A^{1.5}$  theoretical scaling relationships for circular cracks with constant stress drop (39). (c) Histogram of the SSE duration. Vertical blue line marks the local minimum used to separate the two populations of SSEs. (d) Relationship between seismic moment  $(M_o)$  and duration (T) of the SSEs. Horizontal blue line is the same duration boundary as in (c) that is used to separate the SSEs into two populations. For each SSE population, black circles show the median T for different  $M_o$  bins. Dashed red lines show least-squares fits giving  $M_o \propto T^{3.1}$  and  $M_o \propto T^{2.8}$  for the shorter and longer-duration SSE populations respectively.

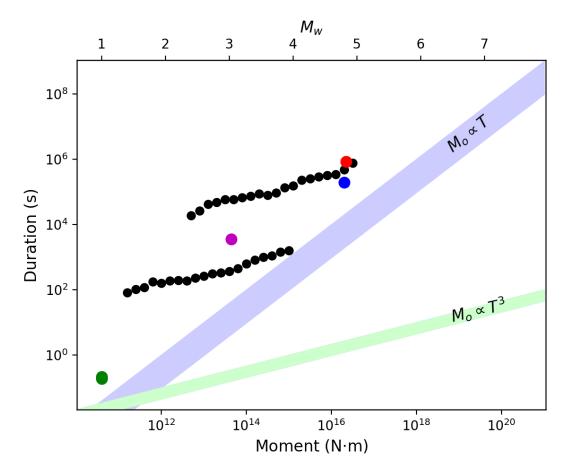


Figure 4: Comparison between slow events on the San Andreas fault and proposed moment-duration scaling for regular (green bar) and slow (blue bar) earthquakes (12). Green circles: two LFE families (42); Magenta circle: episodic creep events (45); Blue circle: slow earthquake (46); Red circle: average of 20 SSEs (7); Black circles: SSEs in this study (same as in Fig. 3d).

## **Supplementary Materials**

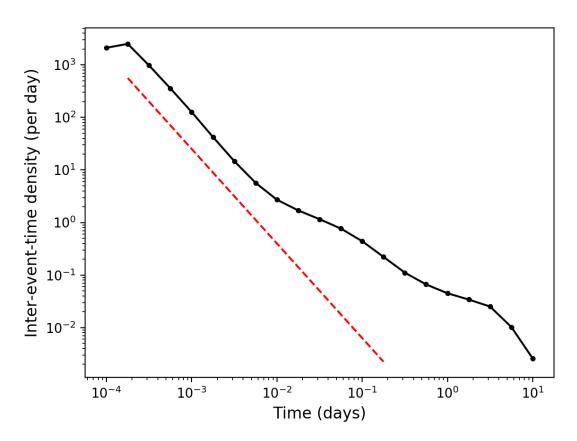


Figure S1: Stacked inter-event time distribution of all LFE families. Dashed red line has slope of -1.8.

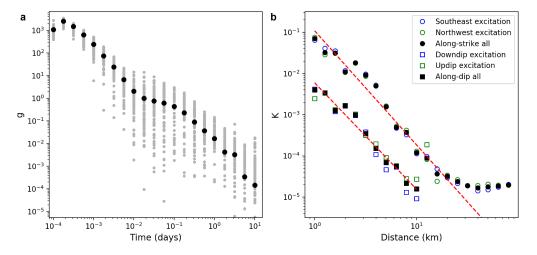


Figure S2: Obtained g and K when g is allowed to vary for different LFE families (see Materials and Methods) (a) g as a function of time. Grey dots show g for different LFE families, while black dots show the same g as in Fig. 2a for comparison. (b) K as a function of distance. Dashed red lines show least-squares fits giving  $K \propto d^{-2.8}$  and  $K \propto d^{-2.6}$  for along-strike and along-dip distances respectively.

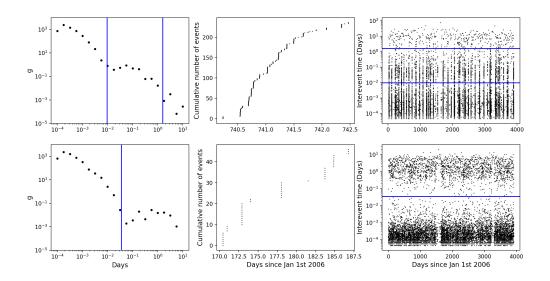


Figure S3: (**Top**) LFE family which has a trimodal interevent time distribution (16 Fig. 4). (**Bottom**) LFE family which has a bimodal interevent time distribution (16 Fig. 3). (**Left**) g as a function of time. Blue lines show separation timescales determined by Thomas et al. (16) base on the interevent time distribution. (**Middle**) Cumulative number of events with time. (**Right**) Interevent time versus days. Blue lines show separation timescales determined by Thomas et al. (16).

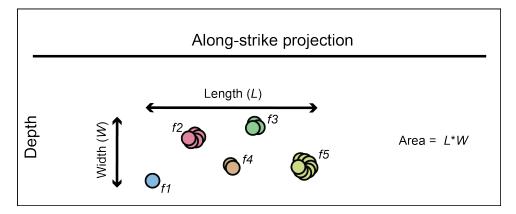


Figure S4: Illustration showing how SSE properties are estimated from an LFE cluster that involves 5 families. Average slip  $\overline{D} = \frac{1}{5}(d_{f1} \cdot n_{f1} + d_{f2} \cdot n_{f2} + d_{f3} \cdot n_{f3} + d_{f4} \cdot n_{f4} + d_{f5} \cdot n_{f5})$ , where  $d_{f1}$  is the slip per LFE from family f1 (see main text) and  $n_{f1}$  is the number of LFE from family f1 that is part of the cluster  $(n_{f1} = 1 \text{ in this example})$ .

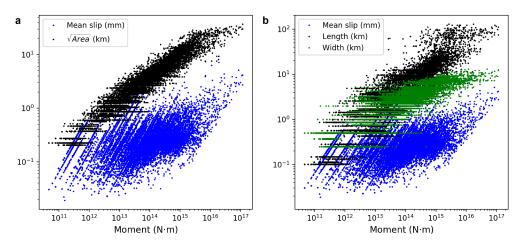


Figure S5: Relationship between length, width, area, mean slip, and moment.

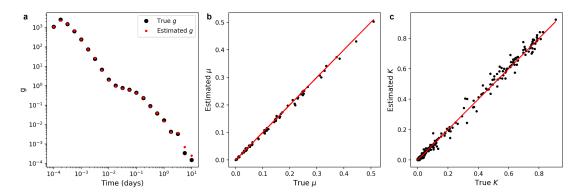


Figure S6: Comparison between true and estimated g,  $\mu$ , and K from synthetic catalog. Red lines have slope of 1.

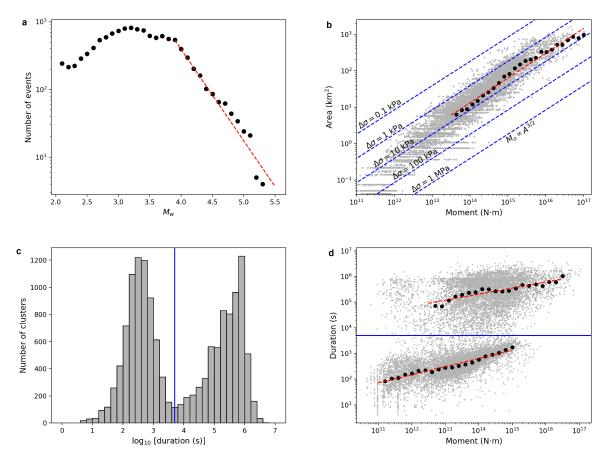


Figure S7: Obtained SSE scaling properties when g is allowed to vary for different LFE families (see Materials and Methods). (a) Non-cumulative magnitude-frequency distribution of the SSEs. Dashed red line shows maximum-likelihood fit giving a b value of  $1.35 \pm 0.03$ . (b) Relationship between seismic moment  $(M_o)$  and rupture area (A) of the SSEs. Black circles show the median A for different  $M_o$  bins. Dashed red line shows least-squares fit giving  $M_o \propto A^{1.5}$ . Dashed blue lines show the  $M_o \propto A^{1.5}$  theoretical scaling relationships for circular cracks with constant stress drop (39). (c) Histogram of the SSE duration. Vertical blue line marks the local minimum  $(10^{3.7} \text{ seconds})$  used to separate the two populations of SSEs. (d) Relationship between seismic moment  $(M_o)$  and duration (T) of the SSEs. Horizontal blue line is the same duration boundary as in (c) that is used to separate the SSEs into two populations. For each SSE population, black circles show the median T for different  $M_o$  bins. Dashed red lines show least-squares fits giving  $M_o \propto T^{4.2}$  and  $M_o \propto T^{3.2}$  for the shorter and longer-duration SSE populations respectively.