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# Quantifying dyke-induced graben and dyke structure using 3D seismic reflection data

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## Abstract

During dyke intrusion, tensile stresses concentrated within the overlying rock may lead to the formation of normal faults. These faults typically form graben-bounding pairs that are sub-parallel to, and dip toward, the upper tip of their underlying dyke. Many studies use geometric properties extracted from the surface expression of such dyke-induced faults to estimate the geometry of subsurface dykes. These methods assume dyke-induced faults are planar and nucleate at the surface. However, recent investigations of the 3D structure of dyke-induced faults using seismic reflection data confirm they can be non-planar and have complex growth histories. Here, we use 3D seismic reflection surveys from offshore NW Australia to: (1) examine how the surface expression of dyke-induced faults relates to subsurface dyke geometry and depth; and (2) test whether subjective bias may influence the quantitative analyses of dyke-induced faults. We show displacement and dip vary across dyke-induced faults, supporting previous suggestions that faults nucleate between dyke upper tips and the free surface. We also find that depths of dyke upper tips predicted using graben width and area of loss measurements, whilst sensitive to fault dip variations and interpretation biases, are often similar to measured dyke depths. Both measured and

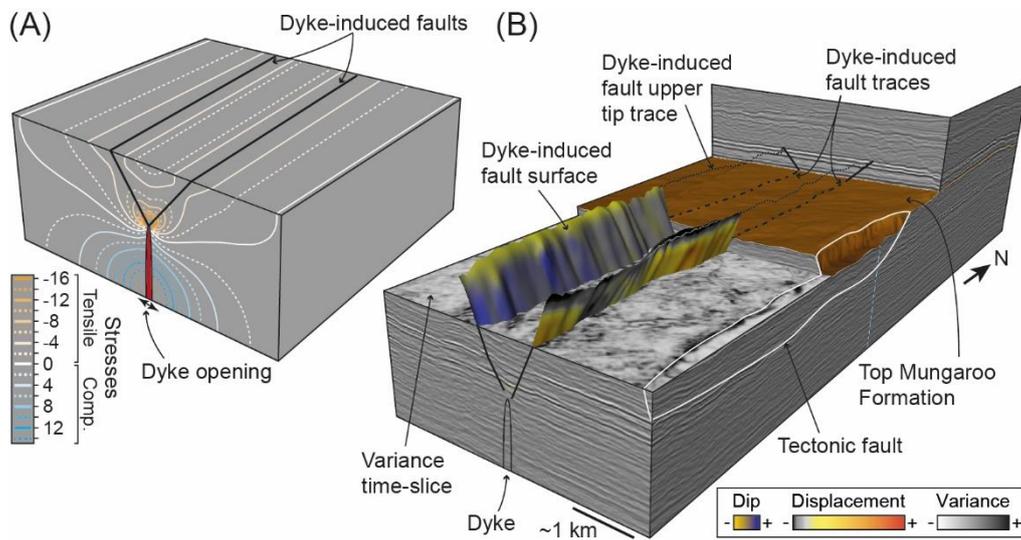
28 predicted dyke depths vary by several hundred metres along-strike, which we relate to the preservation of  
29 dyke heads, segmentation, and/or magma density changes. Overall, we show reflection seismology  
30 provides a better understanding of the 3D structure of dyke-induced faults and their relationship to the  
31 geometry and emplacement dynamics of their causal dykes.

32

### 33 1. Introduction

34 Igneous dykes facilitate magma transport and their intrusion generates both extensional and  
35 compressional stresses (Fig. 1A) (e.g., Rivalta et al., 2015, Rubin, 1993). Within the upper crust, dyke  
36 intrusion and opening is expected to concentrate tensile stresses above the dyke upper tip and in two  
37 elongate, sub-parallel zones at the free surface (Fig. 1A) (e.g., Koehn et al., 2019, Pollard et al., 1983,  
38 Rubin, 1992, Rubin & Pollard, 1988). Shear failure within this extensional stress field produces graben-  
39 bounding normal faults that strike sub-parallel to the underlying dyke and which extend from the dyke's  
40 upper tip to zones of maximum tensile stress at the surface (Fig. 1A) (e.g., Pollard et al., 1983, Trippanera  
41 et al., 2015a). In many active and ancient volcanic systems, on Earth and other planetary bodies, we can  
42 examine the surface expression of such dyke-induced faults and quantify their geometry, displacement,  
43 and kinematics (e.g., Perrin et al., 2022, Pollard et al., 1983, Rivas-Dorado et al., 2021, Trippanera et al.,  
44 2015a, Wilson & Head, 2002, Xu et al., 2016). We expect these fault properties to relate to the underlying  
45 dyke location, size, shape, and intrusion dynamics (e.g., Dumont et al., 2017, Rivas-Dorado et al., 2021,  
46 Trippanera et al., 2015a). Studying dyke-induced faults thus allows us to infer the structure and dynamics  
47 of subsurface dykes, which is crucial to understanding volcanic activity and the role of dyking in crustal  
48 extension (e.g., Dumont et al., 2016, Rivas-Dorado et al., 2021, Trippanera et al., 2015b, Wilson & Head,  
49 2002, Xu et al., 2016).

50



51

52 **Figure 1:** (A) Schematic 3D block diagram showing the modelled stress distribution around an opening  
 53 dyke, and the location of expected dyke-induced faults (based on Rubin, 1992). (B) 3D seismic reflection  
 54 data from the Chandon 3D seismic survey, offshore NW Australia, detailing the seismic expression of  
 55 dykes and dyke-induced faults (based on Magee & Jackson, 2020b). Fault dip and displacement  
 56 variations are shown on the opposing faults.

57

58 Several previous studies have used the width of dyke-induced graben and fault heaves to estimate  
 59 the depth to dyke upper tips and dyke thickness, respectively (e.g., Hjartardóttir et al., 2016, Perrin et al.,  
 60 2022, Rivas-Dorado et al., 2021, Rubin, 1992, Rubin & Pollard, 1988, Trippanera et al., 2015b). These  
 61 predictions assume that dyke-induced faults are planar and project down-dip to intersect at the dyke upper  
 62 tip (Fig. 1A), and that their cumulative heaves are equivalent to dyke thickness (e.g., Magee & Jackson,  
 63 2020b). Physical, numerical, and analytical models support these assumptions (e.g., Hardy, 2016, Koehn  
 64 et al., 2019, Mastin & Pollard, 1988, Pollard et al., 1983, Trippanera et al., 2015b), but it is difficult to test  
 65 their validity because: (1) we lack field exposures that reveal the 3D structure of dyke *and* dyke-induced  
 66 fault systems; and (2) models of ground movement related to active dyke-induced faulting are typically  
 67 non-unique (e.g., Wright et al., 2006).

68 Reflection seismology allows us to image entire dyke and dyke-induced fault systems in 3D at a  
 69 decametre-scale (Fig. 1B) (Bosworth et al., 2015, Magee & Jackson, 2020a, Magee & Jackson, 2020b). For  
 70 example, using seismic reflection data from offshore NW Australia, Magee and Jackson (2020b) show

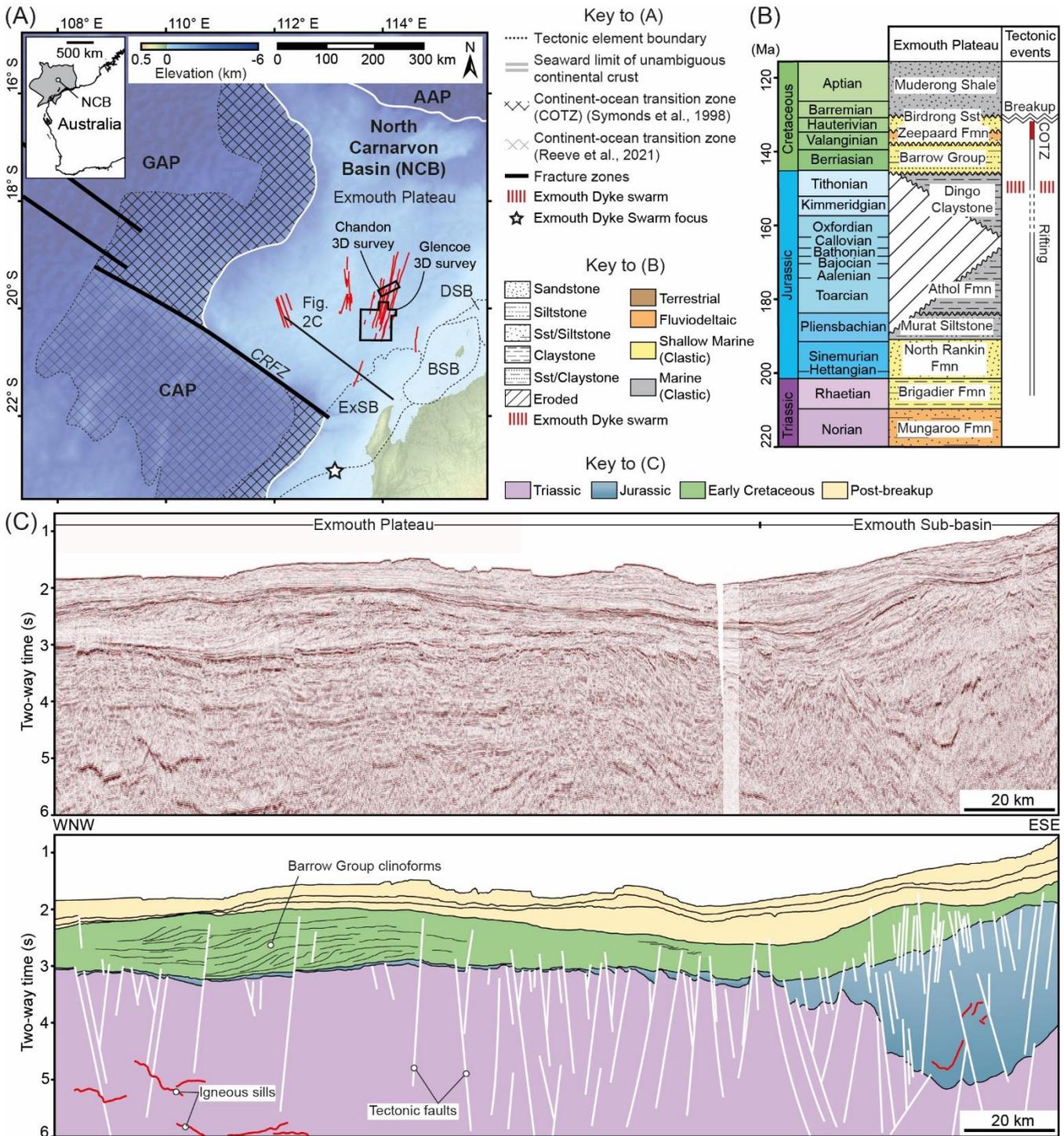
71 that fault displacement, heave, and dip varied laterally and vertically across two buried dyke-induced faults  
72 (Fig. 1B). Heterogeneity in fault dip indicates dyke-induced faults are not always planar, as is commonly  
73 assumed, and that their surficial heave may not fully reflect the extension the faults accommodate (Fig.  
74 1B) (Magee & Jackson, 2020b). These findings question the accuracy of dyke locations, sizes, and shapes  
75 estimated from the surface expression of dyke-induced faults (Magee & Jackson, 2020b). However,  
76 extracting quantitative data (e.g., dyke-induced fault properties) from seismic reflection data is subject to  
77 several objective and subjective sources of uncertainty (e.g., Alcalde et al., 2017a, Bond et al., 2007,  
78 Dimmen et al., 2022, Faleide et al., 2021, Wilson et al., 2019). We thus need to explore how these sources  
79 of uncertainty may affect interpretation of dyke-induced fault data extracted from seismic reflection  
80 volumes.

81 Here, we use two 3D seismic reflection datasets from offshore NW Australia to extend the study  
82 of Magee and Jackson (2020b). We specifically test: (1) if other dyke-induced fault pairs in the region  
83 show similar variations in displacement and dip across their surface (e.g., Fig. 1B) (Magee & Jackson,  
84 2020b); (2) the reliability of trigonometry and area of loss methods for estimating dyke upper tip depths,  
85 which we henceforth refer to as top-dyke depths, from dyke-induced graben properties (e.g., Pollard et al.,  
86 1983, Rivas-Dorado et al., 2021, Trippanera et al., 2015b); and (3) how sources of uncertainty affect  
87 variations in measured or calculated fault, graben, or dyke properties (e.g., Bond et al., 2007, Faleide et al.,  
88 2021). To achieve our aims, we compile and calculate fault property data from footwall and hanging wall  
89 cut-offs mapped by four individuals from the same faults. We present data for dyke-induced fault pairs  
90 above three dykes and show displacement and dip is variable across them all, consistent with the findings  
91 of Magee and Jackson (2020b). Although interpretation bias can introduce measurement errors,  
92 consistency between datasets produced by different individuals suggests we can, at least to a first order,  
93 relate the surface expression of dyke-induced faults to dyke geometry (cf. Magee & Jackson, 2020b).  
94 However, building confidence in estimated dyke parameters requires knowledge of how fault properties,  
95 particularly dip, change with depth. Unfortunately, this subsurface information is rarely available for active  
96 volcanic settings or other planetary bodies, but we suggest empirical data and relationships derived from  
97 reflection seismic data could help reduce uncertainty (Magee & Jackson, 2020b).

99 **2. Geological setting**

100 The Northern Carnarvon Basin, located offshore NW Australia, developed during several phases of rifting  
101 between Australia and Greater India in the Late Carboniferous-to-Early Cretaceous (Fig. 2) (e.g., Direen  
102 et al., 2008, Stagg et al., 2004, Tindale et al., 1998). The Exmouth Plateau is located in the south-west of  
103 the Northern Carnarvon Basin, and is a region of <10 km thick continental crust overlain by a <18 km  
104 thick sedimentary sequence (Fig. 2C) (e.g., Exon et al., 1992, Karner & Driscoll, 1999, Pryer et al., 2002).  
105 Rifting of the Exmouth Plateau began in the Late Triassic-to-Jurassic, forming an array of ~N-S striking,  
106 large (often >1 km throw) normal faults within pre-rift, fluvio-deltaic sedimentary rocks of the Mungaroo  
107 Formation (Figs 2B and C) (e.g., Bilal & McClay, 2022, Bilal et al., 2018, Stagg et al., 2004). The Exmouth  
108 Plateau was sediment-starved during this phase of rifting, thus contains a relatively condensed ( $\leq 100$  m  
109 thick), late Triassic-to-Early Jurassic marine succession (e.g., Figs 2B and C) (e.g., Exon et al., 1992, Karner  
110 & Driscoll, 1999). This latest Triassic-to-Early Jurassic strata is separated from the Late Jurassic, marine  
111 Dingo Claystone by the end Callovian regional unconformity (Fig. 2B) (e.g., Bilal et al., 2018, Tindale et  
112 al., 1998, Yang & Elders, 2016). Tectonic faulting reduced or ceased in the Late Jurassic across the North  
113 Carnarvon Basin, but renewed after formation of the Base Cretaceous unconformity (latest Tithonian;  
114 ~148 Ma) and during deposition of the Tithonian–Valanginian (~148–138 Ma), marine Barrow Group  
115 (Fig. 2B) (e.g., Gartrell et al., 2016, Paumard et al., 2018, Reeve et al., 2016). This renewed faulting  
116 produced N-S to NE-SW-striking, low-throw (<0.1 km) normal faults (e.g., Black et al., 2017). Continental  
117 break-up eventually occurred along the western and southern margins of the Exmouth Plateau in the Early  
118 Cretaceous (Valanginian–Hauterivian; ~135–130 Ma), followed by thermal subsidence and passive  
119 margin development (Figs 2B and C) (e.g., Direen et al., 2008, Reeve et al., 2021, Robb et al., 2005).

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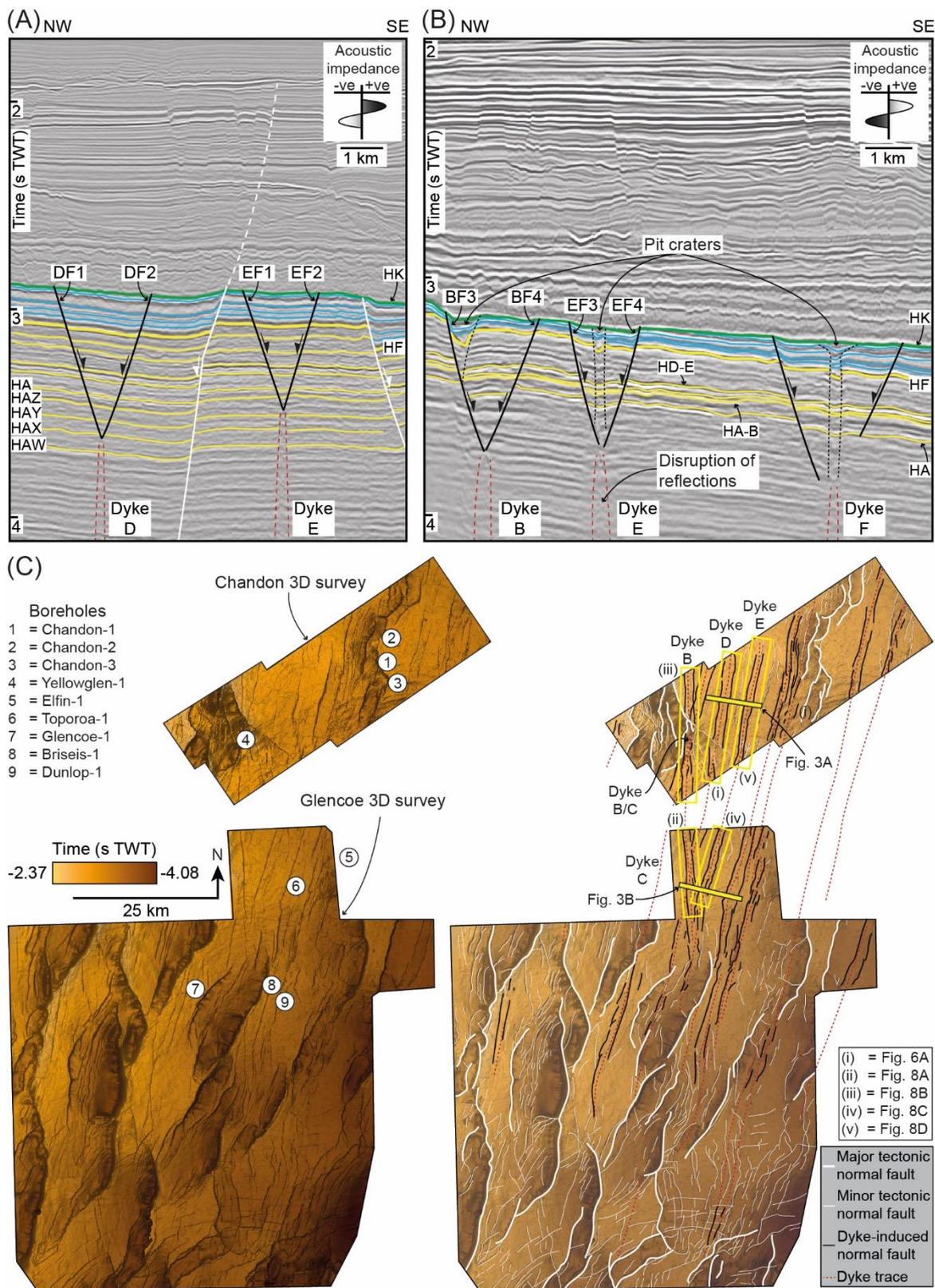
121  
 122 **Figure 2:** (A) Map of offshore NW Australia highlighting principal tectonic elements: ExSB = Exmouth  
 123 Sub-basin, BSB = Barrow Sub-basin, DSB = Dampier Sub-basin, CAP = Cuvier Abyssal Plain, GAP =  
 124 Gascoyne Abyssal Plain, AAP = Argo Abyssal Plain, CRFZ = Cape Range Fracture Zone. Elevation data  
 125 are based on the 2009 Australian Bathymetry and Topography grid (Geoscience Australia). The  
 126 Exmouth Dyke Swarm is also shown (Magee & Jackson, 2020a). (B) Tectono-stratigraphic column for  
 127 the Exmouth Plateau (based on Hocking et al., 1987, Longley et al., 2002, Magee & Jackson, 2020a,

128 Tindale et al., 1998). (C) Uninterpreted and interpreted 2D seismic line across the Exmouth Plateau and  
129 Exmouth Sub-basin showing the upper part of the sedimentary sequence (modified from Norcliffe et al.,  
130 2021).

### 131 **2.1. *The Exmouth Dyke Swarm and dyke-induced faults***

132 Seismic reflection data reveal a swarm of sub-vertical, low-amplitude zones that disrupt stratigraphic  
133 reflections within (and below) the Mungaroo Formation across the Exmouth Plateau (e.g., Figs 1B, 3A,  
134 and C) (Magee & Jackson, 2020a). Borehole data confirm these vertical zones of disruption correspond to  
135 dykes, each likely 10's of metres thick, belonging to the Exmouth Dyke Swarm (Magee & Jackson, 2020a).  
136 These dykes are 10's–100's km long and appear to radiate outwards from focal area within the Cuvier  
137 Margin, from which they likely propagated laterally northwards (e.g., Figs 2A) (Magee & Jackson, 2020a).  
138 A series of graben occur directly above and along the dykes, with the oppositely dipping faults intersecting  
139 at the dyke upper tips (e.g., Figs 1B and 3); these have been interpreted as dyke-induced faults (Magee &  
140 Jackson, 2020a, Magee & Jackson, 2020b). The dyke-induced faults offset siliciclastic Triassic-to-Jurassic  
141 strata, and terminate upwards at the Base Cretaceous unconformity, which is inferred to mark a syn-  
142 faulting free surface (i.e. the seabed; e.g., Figs 3A and B) (Magee & Jackson, 2020a, Magee & Jackson,  
143 2020b). Within the dyke-induced graben are numerous pit craters, which extend from dyke upper tips or  
144 dyke-induced fault planes up into the Upper Jurassic Dingo Claystone (e.g., Fig. 3B) (Magee & Jackson,  
145 2020a, Magee et al., 2022).

146



147

148 **Figure 3:** (A and B) Interpreted seismic sections showing dykes, dyke-induced faults, and stratigraphic  
 149 horizons (HK–HAW) mapped in the Chandon and Glencoe 3D surveys. Yellow horizons are Triassic,  
 150 blue horizons are Jurassic, and the green horizon marks the Base Cretaceous Unconformity (HK). Some  
 151 tectonic faults and pit craters are also highlighted (Magee & Jackson, 2020a). Line locations shown in

152 (C) and insets show how a downwards positive (+ve) or negative (-ve) change in acoustic impedance  
153 affects reflection colour. Uninterpreted sections shown in Supplementary Figure S2. (C) Uninterpreted  
154 and interpreted time-structure map of the Top Mungaroo Formation (horizon HF) in the Chandon and  
155 Glencoe 3D surveys, with tectonic and dyke-induced faults marked. Boreholes used in the study as well  
156 as underlying dyke traces are also shown (Magee & Jackson, 2020a).

157

### 158 3. Data and methods

#### 159 3.1. Data

160 We use the Chandon and Glencoe 3D seismic reflection surveys to analyse dyke-induced faults above  
161 three dykes (dykes B, D, and E; Fig. 3). Both seismic surveys are time-migrated, processed to zero-phase,  
162 and have bin size of 25 m; Chandon has a record length of 6 seconds two-way time (s TWT), whereas  
163 Glencoe extends down to 8 s TWT. The Chandon 3D survey has a SEG reverse polarity, whereby a positive  
164 reflection coefficient (indicating a downward increase in acoustic impedance) corresponds to a trough  
165 (black) reflection and a negative reflection coefficient (indicating downward decrease in acoustic  
166 impedance) is marked by a peak (white) reflection (Fig. 3A). In contrast, the Glencoe 3D survey has an  
167 SEG normal polarity, whereby a positive reflection coefficient corresponds to a peak (white) reflection  
168 and a negative reflection coefficient is marked by a trough (black) reflection (Fig. 3B). To constrain the  
169 age and lithology of mapped reflections, we tied four and five different boreholes to the Chandon and  
170 Glencoe 3D surveys, respectively (Fig. 3C). Checkshot data from these boreholes allowed us to establish  
171 time-depth relationships for the two seismic surveys (Supplementary Fig. S1; Supplementary Table S1),  
172 which we used to depth-convert measurements from TWT to metres. With these time-depth relationships  
173 and dominant frequencies of  $\sim 40\text{--}30$  Hz within the interval of interest, we estimate the limits of  
174 separability and visibility for both datasets to be  $\sim 20 \pm 4$  m and  $3 \pm 1$  m, respectively. The limit of  
175 separability defines the smallest vertical distance between two boundaries for them to be expressed in the  
176 data as two individually discrete reflections (e.g., Brown, 2011). At vertical separations less than the limit  
177 of separability, the reflected seismic wavelets overlap and undergo tuning associated with either  
178 constructive or destructive interference of the two wavelets; this results respectively in either a brightening

179 or dimming of the seismic amplitude. The peaks of the two reflections remain resolvable within the  
180 resultant merged wavelet down to a minimum vertical separation at the limit of visibility, at which point  
181 it is no longer possible to resolve the two component reflections as separate entities (e.g., Brown, 2011).  
182 For any boundaries separated by a vertical distance less than the limit of visibility, their reflection signal  
183 is obscured by seismic noise (Brown, 2011). The horizontal resolution within the interval interest, and  
184 given the time-depth and frequency ranges, would be equivalent to the bin size, which is ~25 m in both  
185 datasets.

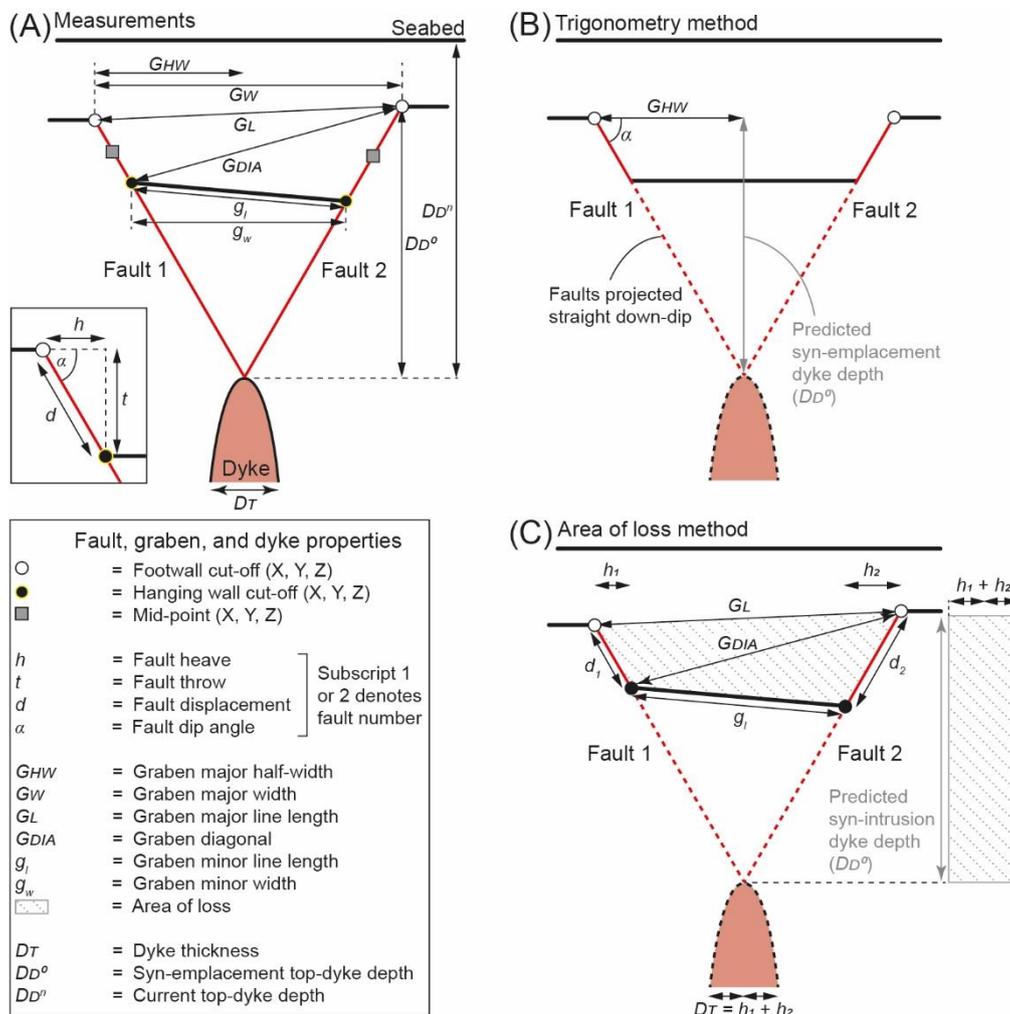
186

### 187 **3.2. Seismic interpretation**

188 To test the findings of Magee and Jackson (2020b), which focussed on the dyke-induced faults above  
189 Dyke E in the Chandon 3D survey, we extend their mapping of 11 seismic horizons (HA–HK; Fig. 3A) to  
190 areas above dykes B and D, which are imaged in the Chandon 3D survey. Of these horizons, HK  
191 corresponds to the Base Cretaceous unconformity and HF defines the Top Mungaroo Formation (Fig. 3A).  
192 The upper tip of some dykes we analyse are deeper than that of Dyke E, meaning their overlying dyke-  
193 induced faults extend to greater depths too; we thus also map several horizons below HA (i.e. HAZ–HAW;  
194 Fig. 3A). Where possible within the Glencoe 3D survey, we map the same HK–HA horizons above dykes  
195 B, D, and E (Fig. 3B). However, due to variations in data quality and reflection continuity within the  
196 Glencoe volume, we cannot map horizon HE and instead interpret a deeper reflection we name HD-E  
197 (Fig. 3B). In the Glencoe volume we also map an extra horizon, HA-B, between HA and HB (Fig. 3B)

198 Using transects oriented perpendicular to fault strike and dyke trend, we measure the hanging wall  
199 and footwall cut-off pairs for each horizon, where they intersect the studied dyke-induced faults, and the  
200 upper tips of underlying dykes (e.g., Fig. 4A). Four of the authors independently mapped fault cut-offs and  
201 dyke tips along the same selected graben (Craig Magee = CM; Victoria Love = VL; Karima Fayez = KF;  
202 Billy Andrews = BA); this approach allows us to assess the impact of interpretation bias on our findings.  
203 Transect spacing is 100 m for BA and 125 m for CM, VL, and KF.

204



205

206 **Figure 4:** (A) Schematic diagram depicting the different dyke-induced fault and graben properties  
 207 calculated from the X, Y, and Z co-ordinates of mapped fault-cut-offs. (B) Sketch showing how fault dips  
 208 and graben half-widths can be used to project faults down-dip and estimate dyke upper tip depths. (C)  
 209 Sketch showing the graben area of loss and cumulative fault heave ( $h_1 + h_2$ ) can be used to define a  
 210 rectangle, the length of which can be considered equivalent to the dyke upper tip depth (modified from  
 211 Rivas-Dorado et al., 2021).

212

### 213 3.3. Measurements and calculations

214 From the coordinates of each interpreted fault cut-off pair, including those mapped by Magee and Jackson  
 215 (2020b), we calculate fault throw ( $t$ ) and heave ( $h$ ) (Fig. 4A). These throw and heave calculations allow  
 216 us to estimate fault dip ( $\alpha$ ) and displacement ( $d$ ), assuming the slip vector is dip-parallel (Fig. 4A) (Magee  
 217 & Jackson, 2020b). We project displacement data onto fault surfaces at X, Y, and Z coordinates marking  
 218 the mid-point between each paired footwall and hanging wall cut-off (Fig. 4A); a convergent interpolation

219 gridding algorithm was used to linearly extrapolate between data points and preserve any trends. For fault  
220 cut-offs mapped on faults along the same horizon and transect, we combine fault property measurements  
221 to derive cumulative throw, heave, and displacement values, as well as the average fault dip ( $\alpha_{av}$ ). Along  
222 each transect, we also use fault cut-off coordinates to calculate (Fig. 4A): (1) the horizontal graben width  
223 ( $G_W$ ) and half-width ( $G_{HW}$ ) between footwall cut-offs; (2) the line length ( $G_L$ ) between footwall cut-offs,  
224 which unlike  $G_W$  or  $G_{HW}$  accounts for differences in cut-off elevation; (3) the horizontal graben width ( $g_w$ )  
225 between hanging wall cut-offs; (4) the line length ( $g$ ) between the hanging wall cut-offs; and (5) the  
226 diagonal line length ( $G_{DIA}$ ) between one hanging wall cut-off and the opposing footwall cut-off.

227 We apply two methods to predict the current ( $D_D^n$ ) and syn-emplacement ( $D_D^0$ ) top-dyke depths  
228 from calculated graben width properties (Fig. 4). We first use trigonometry to estimate top-dyke depths  
229 from graben half-widths ( $G_{HW}$ ) and fault dips ( $\alpha$ ) measured at any stratigraphic level, such as the syn-  
230 emplacement free surface (Fig. 4B) (e.g., Pollard et al., 1983, Trippanera et al., 2015b). This trigonometric  
231 method assumes faults are planar and project straight down-dip (Fig. 4B) (Magee & Jackson, 2020b).  
232 Most studies assume  $\alpha$  is constant for both faults and, based on measurements and/or regional  
233 information, is  $\sim 70\text{--}60^\circ$  (e.g., Rubín & Pollard, 1988, Trippanera et al., 2015a). We present results where  
234 we consider that both faults have dips equivalent to either: (1) the average of the two fault dips ( $\alpha_{av}$ ); (2)  
235  $60^\circ$ , as this is thought typical of normal faults within an Andersonian framework (Anderson, 1951); and  
236 (3)  $45^\circ$ , which is similar to the dip of Late Jurassic-to-Early Cretaceous tectonic normal faults in the region  
237 (e.g., Magee et al., 2016). We henceforth refer to these three method variants as trig $\alpha_{av}$ , trig60, and trig45,  
238 respectively.

239 In addition to the trigonometry method for predicting top-dyke depths, we use an area of loss  
240 method modified from Rivas-Dorado et al. (2021). This method calculates the area between the four fault  
241 cut-offs on each transect, which typically define an irregular quadrilateral shape, and creates a rectangle  
242 of the same area (Fig. 4C) (Rivas-Dorado et al., 2021). By setting the width of this derived rectangle to  
243 equal the cumulative fault heave, which is assumed equivalent to dyke thickness ( $D_T$ ), the resulting length  
244 of the rectangle is taken as the vertical distance between the syn-emplacement surface and dyke upper tip  
245 (Fig. 4C); i.e. the graben area of loss is assumed to equal the area gained by dyke-driven extension (Rivas-

246 Dorado et al., 2021). Our methodology for finding the area of loss accounts for all cut-offs occurring at  
247 different elevations/depths, as opposed to the hanging wall cut-offs being assumed to have the same  
248 elevation (Rivas-Dorado et al., 2021).

249

### 250 **3.4. *Limitations and errors***

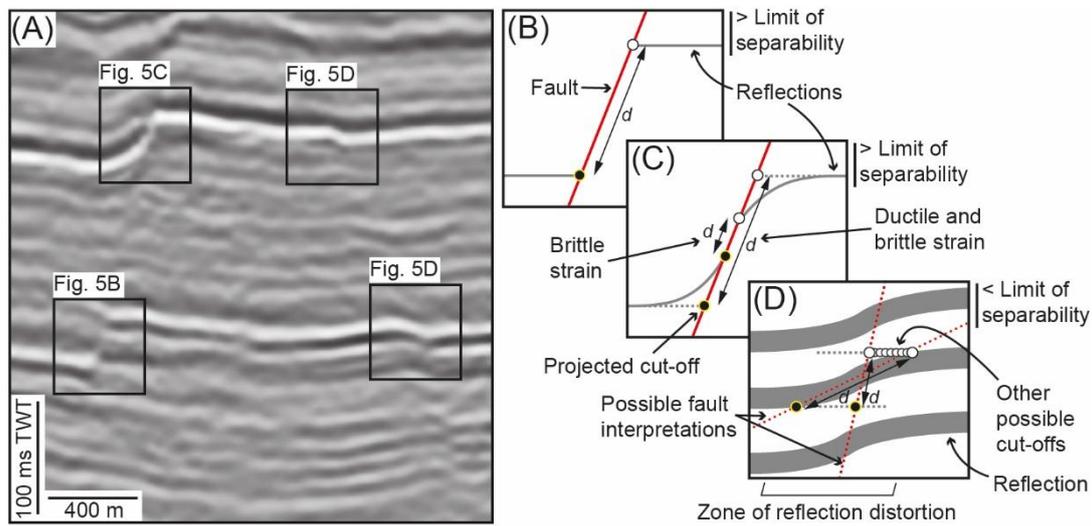
251 Although seismic reflection data provide unique insight into the 3D structure of dykes and dyke-induced  
252 faults, interpreting and quantitatively assessing these features, which often have displacements of <20 m,  
253 is affected by several objective and subjective uncertainties involved in seismic interpretation (e.g.,  
254 Dimmen et al., 2022, Magee & Jackson, 2020b, Wilson et al., 2019). Objective uncertainty, which can often  
255 be quantified, includes limitations related to resolution and quality of seismic imaging, as well as seismic  
256 velocities used for depth conversion (Faleide et al., 2021). For example, the seismic velocities we use are  
257 taken from borehole data, meaning they do not capture possible variations in seismic velocity across the  
258 study area, away from the borehole locations (Magee & Jackson, 2020b). However, because the time-  
259 depth relationships of the boreholes for each 3D survey we use are similar (Supplementary Fig. 1), we  
260 adopt the conservative view that calculated seismic velocities could vary by up to  $\pm 10\%$  (Magee & Jackson,  
261 2020a, Magee & Jackson, 2020b).

262 It is also important to consider that the dykes and dyke-induced faults in the Exmouth Plateau  
263 formed during the Late Jurassic, and have since been buried by several kilometres of sedimentary strata  
264 (Magee & Jackson, 2020a). Most sedimentary rocks compact as they are buried, reducing stratal  
265 thicknesses and rotating any pre-existing, inclined fractures to shallower angles (Allen & Allen, 2013).  
266 Burial-related compaction can thus reduce fault throw, potentially by up to 15% in sand-dominated or  
267 mixed sand-shale lithologies (Taylor et al., 2008), meaning our calculated fault dips, displacements, and  
268 predicted top-dyke depths are minimum estimates. However, we note that all dykes and dyke-induced  
269 faults occur at similar depths ( $\sim 3\text{--}4$  km) under a similar overburden thickness, so we suggest compaction,  
270 and compaction-related modification of the primary geometries, can be considered to have been constant  
271 across the study area (Magee & Jackson, 2020a, Magee & Jackson, 2020b). Although our measurements  
272 and calculations may not therefore reflect the absolute syn-emplacement 3D structure of the dyke and

273 dyke-induced faults, their relative values will be comparable, i.e. the current patterns of displacement, dip,  
274 or predicted dyke upper tip depth distribution will be the same as when dyking occurred in the Late  
275 Jurassic.

276           Subjective uncertainties and biases relate to those introduced by the person undertaking the  
277 interpretation and are generally more difficult to quantify (e.g., Alcalde et al., 2017a, Bond et al., 2007,  
278 Faleide et al., 2021, Wilson et al., 2019). For example, mapping footwall and hanging wall cut-offs of low  
279 offset faults, or faults with a high proportion of continuous deformation, is particularly prone to  
280 interpretation bias (Schaaf & Bond, 2019). Where fault displacement is substantially greater than the limit  
281 of separability and horizontal resolution, reflections are often clearly offset and discrete footwall and  
282 hanging wall cut-offs can be identified (Figs 5A and B) (Dimmen et al., 2022). In these cases, any fault-  
283 related continuous deformation (i.e. local rotation of bedding and thus reflections adjacent to faults;  
284 Delogkos et al., 2017) can be accounted for by projecting the regional dip of horizons onto the fault to  
285 define cut-off positions (Figs 5A and C) (e.g., Mansfield & Cartwright, 1996). In contrast, where fault  
286 displacement is close to or below the limit of separability and horizontal resolution, reflections are  
287 continuous but appear locally distorted (Figs 5A and D) (e.g., Dimmen et al., 2022, Faleide et al., 2021).  
288 Here, it is the interpreter's decision, which can be biased by experience (e.g., Bond et al., 2007), as to where  
289 to place the footwall and hanging wall cut-offs (Fig. 5D). For example, an interpreted portion of the fault  
290 will have a shallower dip and larger displacement (and heave) if cut-offs are mapped at the inflection  
291 points of a distorted reflection, as opposed to projecting the cut-offs to a mid-point (Fig. 5D). Any  
292 projection of cut-offs also influences measured graben width properties. Given limitations of data  
293 resolution and imaging quality, there is no unique, or 'correct', interpretation of fault cut-off location (e.g.,  
294 Alcalde et al., 2017a, Dimmen et al., 2022, Faleide et al., 2021).

295



296

297 **Figure 5:** (A) Uninterpreted seismic section from the Chandon 3D survey showing different reflection  
 298 configurations across two faults. (B-D) Schematic diagrams describing how and where fault footwall  
 299 and hanging wall cut-offs may be interpreted for the different reflection-fault interactions in (A) (e.g.,  
 300 Dimmen et al., 2022).

301

302 Previous studies have suggested interpretation bias in quantifying dyke-induced fault properties  
 303 can be conservatively accounted for by applying  $\pm 5\%$  measurement errors (Magee & Jackson, 2020b). To  
 304 preliminarily explore the effect of interpretation bias, and test these prior assumptions, we present data  
 305 based on independent mapping of fault cut-offs by four individuals (CM = Craig Magee; VL = Victoria  
 306 Love; KF = Karima Fayez; BA = Billy Andrews). Each interpreter mapped fault cut-offs along the Top  
 307 Mungaroo Formation for one or several dyke-induced fault pairs, using different transects and transect  
 308 spacings relative to others. We also conducted a repeat experiment whereby CM mapped dyke-induced  
 309 fault cut-offs above Dyke D along the Top Mungaroo Formation on the same transects on two occasions  
 310 (CM1 and CM2), separated by  $\sim 1$  year. To compare the results from these interpreter datasets, we use F-  
 311 tests to statistically determine the probability that calculated fault properties or predicted dyke upper tip  
 312 depths are not significantly different. If the calculated F value for the two or three datasets being compared  
 313 is less than a critical amount ( $F_{crit}$ ), the null hypothesis that the datasets may be considered equal is  
 314 accepted.

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## 316 4. Results

### 317 4.1. *Structural framework*

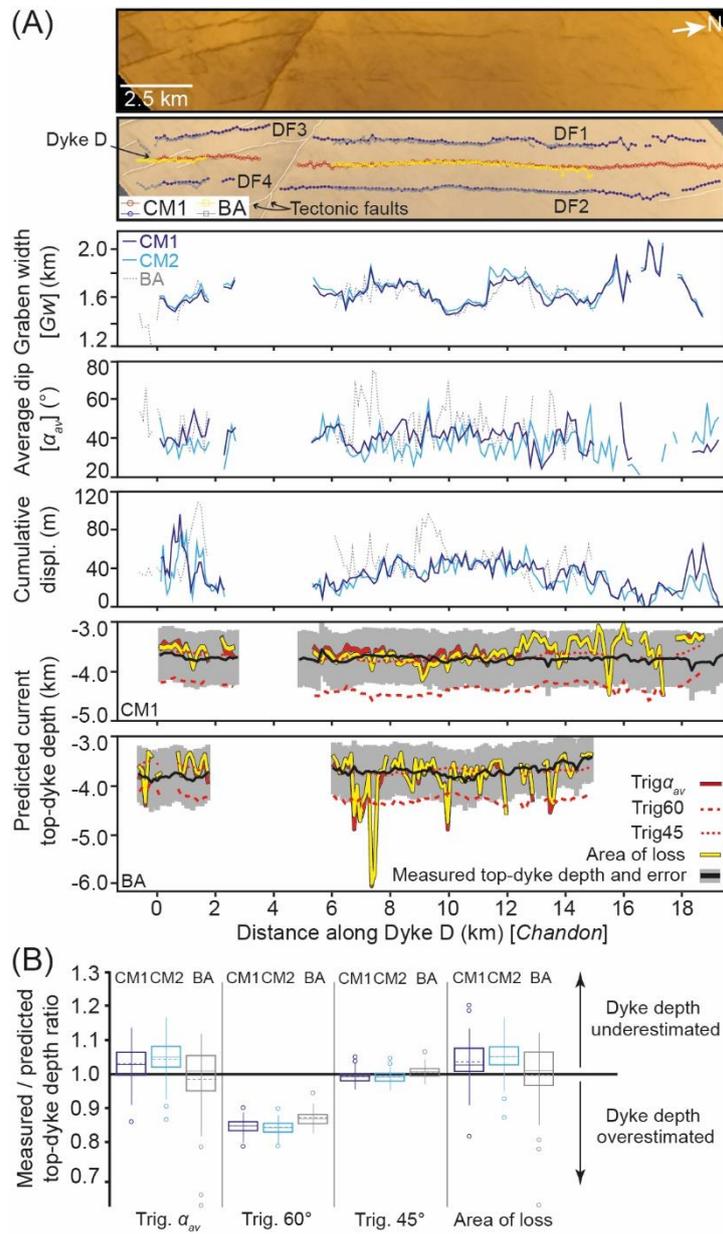
318 For all dykes, their associated dyke-induced faults are sub-parallel to the dyke strike (Fig. 3C). Dykes D  
319 and E are sub-parallel and strike  $\sim 012^\circ$  (Fig. 3C). Dyke B and its dyke-induced faults trend  $\sim 002^\circ$ ,  
320 intersecting other dykes (i.e. C, D, and E) and associated faults, respectively (Fig. 3C). Where these dykes  
321 and faults intersect, they can be difficult to identify and assign (Fig. 3C). Dyke B and its dyke-induced  
322 faults also cross-cut or are cross-cut by a major W-dipping,  $\sim$ N-S striking, tectonic normal fault, and several  
323 smaller associated tectonic faults within the Chandon 3D survey (Fig. 3C). Within the Chandon 3D survey,  
324 dyke-induced faults above Dyke D are also cross-cut, and possibly offset by, a minor S-dipping,  $\sim$ NW-SE  
325 striking, tectonic normal fault (Fig. 3C). Elsewhere in the study area, there is little connectivity between  
326 the studied dyke-induced faults and tectonic faults (Fig. 3C).

327

### 328 4.2. *Dyke D and its overlying dyke-induced faults*

329 To illustrate our research methodology, we first present quantitative data from the dyke-induced faults,  
330 which we separate into DF1-DF2 and DF3-DF4 pairs, above Dyke D in the Chandon 3D survey (Fig. 6A).  
331 We start by analysing graben width, average fault dip, and cumulative fault displacement along DF1-DF2  
332 and DF3-DF4 at the Top Mungaroo Formation (horizon HF) (Fig. 6). To calculate these data, CM and  
333 BA interpret footwall and hanging wall cut-offs on different seismic sections spaced 125 m or 100 m apart,  
334 respectively (Fig. 6A); two cut-off sets are mapped by CM (i.e. CM1 and CM2; Fig. 6). These data allow  
335 us to predict Dyke D top-dyke depths, which we compare to top-dyke depths measured from the seismic  
336 reflection data. Finally, we demonstrate how graben width measurements collected along different  
337 stratigraphic horizon improve top-dyke depth predictions.

338



339

340 **Figure 6:** (A) Uninterpreted and interpreted time-structure map of the Top Mungaroo Formation  
 341 (horizon HF) above Dyke D in the Chandon 3D survey (see Fig. 3C for key and location). Footwall cut-  
 342 off and dyke upper tip depth locations mapped by CM1 and BA are shown. Plots compare graben width,  
 343 average fault dip, and cumulative displacement (displ.) calculated from CM1, CM2, and BA cut-off  
 344 mapping. Also shown are plots comparing measured top-dyke depths with those predicted by CM1 and  
 345 BA using different methods (trig $\alpha_{av}$ , trig60, trig 45, and area of loss). Error bars are not shown because  
 346 we compare calculations subject to the same uncertainties and to improve clarity; we make an exception  
 347 for the measured top-dyke depths, for which we include a  $\pm 15\%$  error envelope, to enable a broader  
 348 comparison to predicted top-dyke depths. (B) Box-and-whisker plots comparing measured and predicted  
 349 top-dyke depth ratios calculated by different methods and using fault cut-off datasets mapped by

350 different interpreters (CM1, CM2, and BA). The boxes describe the interquartile range and median of  
351 each dataset, with the whiskers marking lower and upper extremes of the data. We also show the mean  
352 (dashed line) and outliers in these data.

353

354 Above Dyke D along the Top Mungaroo Formation, graben widths are  $\sim 1.20$ – $2.05$  km, average  
355 fault dips are  $\sim 20$ – $85^\circ$ , and cumulative displacements are  $\sim 5$ – $97$  m (Fig. 6; Supplementary Table S2).  
356 Dyke-induced fault lengths mapped by CM and BA differ, with CM interpreting DF1-DF2  $\sim 3$  km further  
357 northwards than BA (Fig. 6A). Where fault cut-offs interpreted by CM and BA spatially overlap, graben  
358 widths are similar, displaying systematic increases and decreases along-strike at the kilometre-scale (Fig.  
359 6A). The average fault dips and cumulative displacement calculated by CM1, CM2, and BA are more  
360 variable (Fig. 6A; Supplementary Table S2). Differences between the CM1 and CM2 datasets, which use  
361 the same transects, are due to small offsets of  $\lesssim 100$  m in cut-off positions (average is  $\sim 18$  m;  
362 Supplementary Table S3) but statistical F-tests reveal their variance is insignificant (i.e.  $F < F_{\text{crit}}$ ; Table 1);  
363 i.e. the CM1 and CM2 datasets may be considered equal. Conversely, F-tests demonstrate that variances  
364 between the CM1, CM2, and BA measurements of graben width, average fault dip, and cumulative  
365 displacement are significant (i.e.  $F > F_{\text{crit}}$ ; Table 1); i.e. the null hypothesis can thus be rejected and the CM  
366 and BA datasets considered unequal.

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**Table 1: Statistical F-tests comparing CM1, CM2, and BA data**

Property	Interpreter	Mean (m)	Variance	CM1 and CM2 comparison		CM1, CM2, and BA comparison	
				F	F <sub>crit</sub>	F	F <sub>crit</sub>
Horizontal graben width [ $G_w$ ]	CM1	1658	011843				
	CM2	1679	011904	1.01	1.35	14.58	3.02
	BA	1602	010577				
Average fault dip [ $\alpha_{av}$ ]	CM1	0040	000038				
	CM2	0037	000052	1.36	1.37	35.78	3.02
	BA	0046	000102				
Cumulative displacement	CM1	0036	000167				
	CM2	0034	000225	1.35	1.35	31.87	3.02
	BA	0050	000365				
Trig $\alpha_{av}$ estimated top- dyke depth	CM1	3556	029921				
	CM2	3501	030500	1.02	1.37	26.45	3.02
	BA	3778	205831				
Trig60 estimated current top-dyke depth	CM1	4282	008973				
	CM2	4300	009292	1.04	1.35	08.98	3.02
	BA	4246	008696				
Trig45 estimated current top-dyke depth	CM1	3675	003510				
	CM2	3686	003649	1.04	1.35	05.23	3.02
	BA	3660	003217				
Area of loss estimated current top-dyke-depth	CM1	3547	040773				
	CM2	3488	027126	1.50	1.35	21.90	3.02
	BA	3735	197886				

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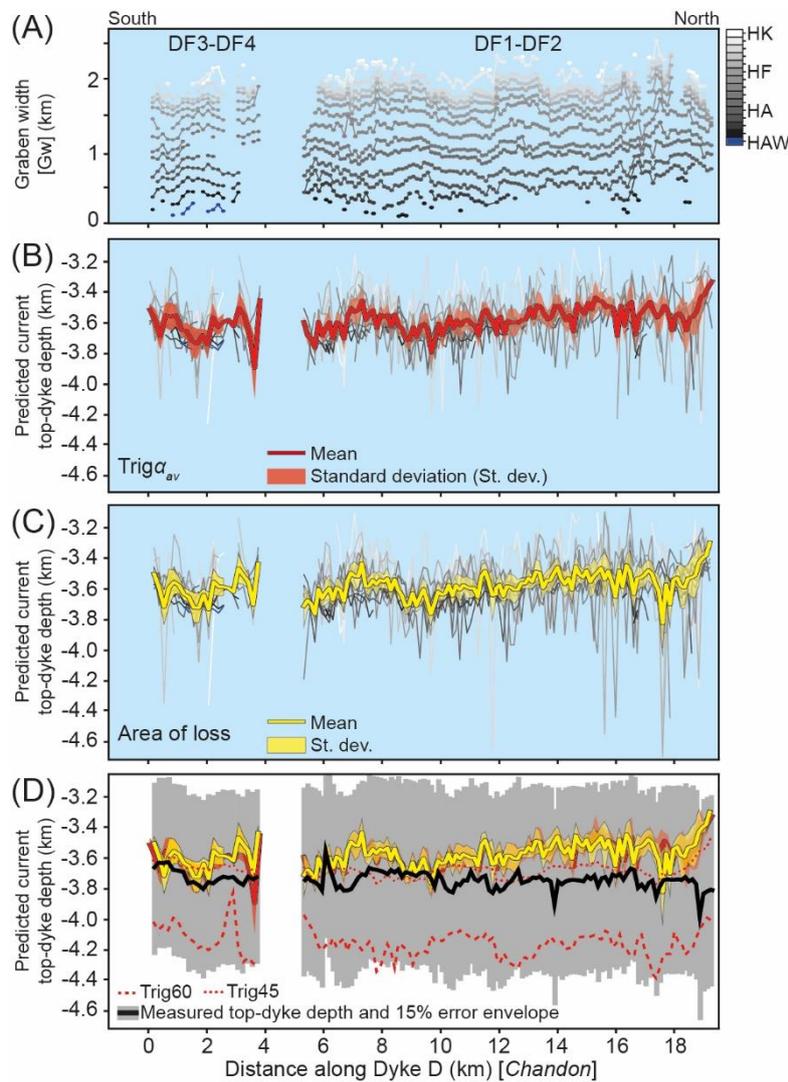
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Measured top-dyke depths for Dyke D in the Chandon 3D survey range from ~3.5–3.8 km beneath current sea-level (Supplementary Table S2), which broadly correspond to emplacement depths of ~0.9–1.1 km beneath the syn-intrusion surface, Horizon HK (Fig. 6A). Along the graben, over lateral distances of ~125–250 m, there are some abrupt changes of up to ~200 m in measured top-dyke depth superimposed onto subtle (<100 m) increases and decreases over the kilometre-scale (Fig. 6A). Predictions of current top-dyke depths derived using the trig $\alpha_{av}$ , trig45, and area of loss methods are often within error of measured top-dyke depths for both CM and BA (Fig. 6; Supplementary Table S2). However, top-dyke depths predicted using the trig60 method appear to overestimate dyke depths (Fig. 6A). Statistical F-tests reveal the variance of trig $\alpha_{av}$ , trig60, and trig 45 calculations by CM1 and CM2 is insignificant (Table 1). Only for top-dyke depths predicted using the area of loss method is the variance significant,  $F (1.50) > F_{crit} (1.35)$ , indicating the CM1 and CM2 calculations are unequal (Table 1). F-tests demonstrate the CM and BA top-dyke depth estimates are unequal, regardless of the prediction method used (Table 1).

We also measure graben width and predict top-dyke depths using fault cut-offs for DF1-DF2 and DF3-DF4 mapped at various stratigraphic horizons (e.g., Fig. 7). As part of the CM1 analysis, we show graben width is variable along DF1-DF2 and DF3-DF4 at all stratigraphic levels and decreases with depth

393 (Fig. 7A) (data provided in Magee & Love, 2021). Top-dyke depths, relative to current sea-level, predicted  
 394 from these graben width measurements at different stratigraphic levels do vary, but together define a  
 395 mean profile with relatively low standard deviations of  $\lesssim 300$  m (e.g., Figs 7B and C). Comparing the mean  
 396 top-dyke depth estimates with those measured from the seismic reflection data reveals that: (1) predictions  
 397 derived from the trig $\alpha_{av}$ , trig45, and area of loss methods are broadly comparable to measured top-dyke  
 398 depths, although on average they underestimate top-dyke depths by  $\lesssim 170$  m; and (2) the trig60  
 399 predictions are now within the error of measured top-dyke depths, but overestimate measured values by  
 400  $\sim 420$  m (Fig. 7D).

401



402

403 **Figure 7:** (A) Graben width measurements for horizons HK–HAW above Dyke D in the Chandon 3D  
 404 survey. Error bars not shown for clarity. (B) Top-dyke depths predicted using graben widths and average  
 405 fault dips acquired from each horizon (HK–HAW) on each measurement transect. From the predicted

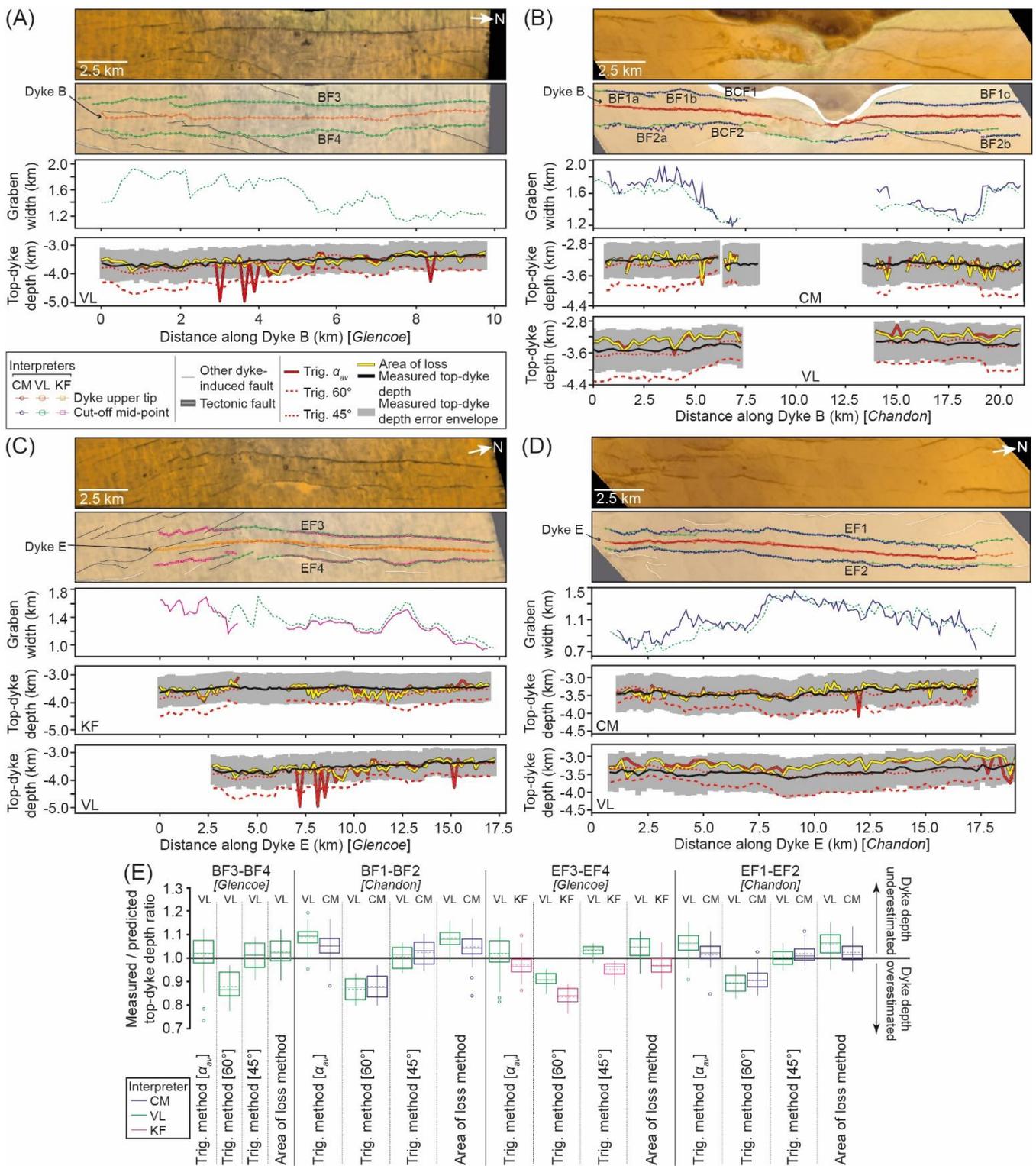
406 top-dyke depths for all horizons, we calculate their mean and standard deviation. (C) Top-dyke depths  
407 predicted using the area of loss method and fault cut-offs mapped for each horizon (HK–HAW) on each  
408 measurement transect. From the predicted top-dyke depths for all horizons, we calculate their mean and  
409 standard deviation. (D) Plot comparing measured top-dyke depths with mean predicted top-dyke depths  
410 estimated using different methods (trig $\alpha$ v, trig60, trig 45, and area of loss). Standard deviation  
411 envelopes are shown for the trig $\alpha$ v and area of loss methods.

412

### 413 ***4.3. Graben widths along the Top Mungaroo Formation and predicted top-dyke depths***

414 To further test how interpreter bias may influence quantitative dyke-induced fault analyses, we examine  
415 how graben width measurements and predicted top-dyke depths for dykes B and E vary between three  
416 interpreters (i.e. CM, VL, and KF; Fig. 8) (all data available in Magee, 2022, Magee & Love, 2021). For this  
417 analysis, we split the dyke-induced faults traces above Dyke B in the Chandon 3D survey into sections  
418 (Fig. 8B): (1) faults BF1a and BF1b both dip eastwards and overlap laterally by ~700 m towards their  
419 upper tips, but are otherwise hard-linked and occur opposite BF2a; (2) further north, the merging of faults  
420 above dykes B and C make it difficult to accurately map their continuation, so we define faults BCF1 and  
421 BCF2 separately; and (3) north of where dykes B and C link, we map dyke-induced faults BF1c and BF2a.  
422 The Dyke B dyke-induced faults analysed within the Glencoe 3D survey are labelled BF3-BF4 (Fig. 8A).  
423 The Dyke E dyke-induced faults in the Chandon 3D survey are named EF1-EF2 (Magee & Jackson, 2020b)  
424 and EF3-EF4 in the Glencoe 3D survey (Figs 8C and D).

425



426

427 **Figure 8:** (A-D) Uninterpreted and interpreted time-structure maps of the Top Mungaroo Formation  
 428 (horizon HF) above BF dyke B and E1 in the Chandon and Glencoe 3D surveys (see Fig. 3C for key and  
 429 locations). Footwall cut-off and dyke upper tip depth locations mapped by different interpreters are  
 430 shown. Plots compare graben width, average fault dip, and cumulative displacement (displ.) calculated  
 431 from different interpreters' cut-off mapping. Also shown are plots comparing measured top-dyke depths  
 432 with those predicted using different methods (trig $\alpha_{av}$ , trig60, trig 45, and area of loss). Error bars are

433 not shown because we compare calculations subject to the same uncertainties and to improve clarity; we  
434 make an exception for the measured top-dyke depths, for which we include a  $\pm 15\%$  error envelope, to  
435 enable a broader comparison to predicted top-dyke depths. (E) Box-and-whisker plots comparing  
436 measured and predicted top-dyke depths ratios calculated by different methods and using fault cut-off  
437 datasets mapped by different interpreters. The boxes describe the interquartile range and median of  
438 each dataset, with the whiskers marking lower and upper extremes of the data. We also show the mean  
439 (dashed line) and outliers in these data.

440

441 Across the Top Mungaroo Formation (Horizon HF), we observe both gradual changes along-strike  
442 in graben width (e.g., EF1-EF2; Fig. 8D), and abrupt changes where overlapping fault segments occur (e.g.,  
443 BF3-BF4; Fig. 8A). Most measurements of graben width at the Top Mungaroo Formation are similar (i.e.  
444  $<100$  m difference) but can vary by up to  $\sim 400$  m (Figs 8B-E). Comparing measured top-dyke depths and  
445 those predicted using the graben widths measured along the Top Mungaroo Formation above of dykes B  
446 and E reveals that (Fig. 8): (1) predicted top-dyke depths derived from the trig  $\alpha_{av}$ , trig45, and area of loss  
447 methods are broadly comparable to measured top-dyke depths; (2) top-dyke depth predictions calculated  
448 using the trig60 method overestimates measured top-dyke depths; (3) even though the four interpreters  
449 mapped fault cut-offs on different transects, there is generally good agreement in their predicted top-dyke  
450 depths, except for EF3-EF4; (4) top-dyke depths appear to increase and decrease by  $\sim 100$  m over several  
451 kilometres; and (5) abrupt and isolated changes in predicted top-dyke depths along individual graben  
452 occur when calculated using the trig  $\alpha_{av}$  and area of loss methods, but are absent when fault dip is assumed  
453 to be  $60^\circ$  or  $45^\circ$ . Statistical F-tests reveal that the variance between graben width measurements acquired  
454 by different interpreters above the same fault are not significant and may thus be considered equal (i.e.  
455  $F < F_{crit}$ ; Table 2). We also show that variance between top-dyke depth predictions for Dyke E in the  
456 Chandon 3D survey is insignificant may be considered equal, regardless of the interpreter or prediction  
457 method, but for the Glencoe 3D survey most interpreter predictions are significantly different from each  
458 other, except where the area of loss method was employed (Table 2). For Dyke B in the Chandon 3D

459 survey, the trig60 and trig45 top-dyke depth predictions calculated from CM and VL interpretations may  
 460 be considered equal, but the trig $\alpha_{av}$  and area of loss data unequal (Table 2).

461

**Table 2: Statistical F-tests comparing CM, VL, and KF data**

Dyke-induced graben		Property	Interpreter	Mean (m)	Variance	F	F <sub>crit</sub>
Dyke B, survey	Chandon 3D	Horizontal graben width [G <sub>w</sub> ]	CM	1561	35490	1.31	1.49
			VL	1489	27083		
		Trig $\alpha_{av}$ estimated top- dyke depth	CM	3386	28122	1.65	1.49
			VL	3183	16996		
		Trig60 estimated current top-dyke depth	CM	4044	47071	1.14	1.49
			VL	3988	41204		
		Trig45 estimated current top-dyke depth	CM	3473	23542	1.08	1.49
			VL	3443	21842		
Area of loss estimated current top-dyke-depth	CM	3403	32538	2.05	1.49		
	VL	3198	15896				
Dyke E, survey	Chandon 3D	Horizontal graben width [G <sub>w</sub> ]	CM	1120	30234	1.30	1.38
			VL	1102	39193		
		Trig $\alpha_{av}$ estimated top- dyke depth	CM	3379	17325	1.14	1.39
			VL	3205	19709		
		Trig60 estimated current top-dyke depth	CM	3818	23523	1.14	1.38
			VL	3803	26909		
		Trig45 estimated current top-dyke depth	CM	3408	09373	1.04	1.38
			VL	3400	09740		
Area of loss estimated current top-dyke-depth	CM	3389	17627	1.24	1.38		
	VL	3214	21813				
Dyke E, survey	Glencoe 3D	Horizontal graben width [G <sub>w</sub> ]	KF	1303	34003	1.22	1.48
			VL	1321	27901		
		Trig $\alpha_{av}$ estimated top- dyke depth	KF	3520	21619	3.58	1.44
			VL	3618	77428		
		Trig60 estimated current top-dyke depth	KF	4056	26940	1.87	1.48
			VL	4031	14413		
		Trig45 estimated current top-dyke depth	KF	3579	10988	2.17	1.48
			VL	3548	05053		
Area of loss estimated current top-dyke-depth	KF	3509	19878	1.01	1.44		
	VL	3535	19741				

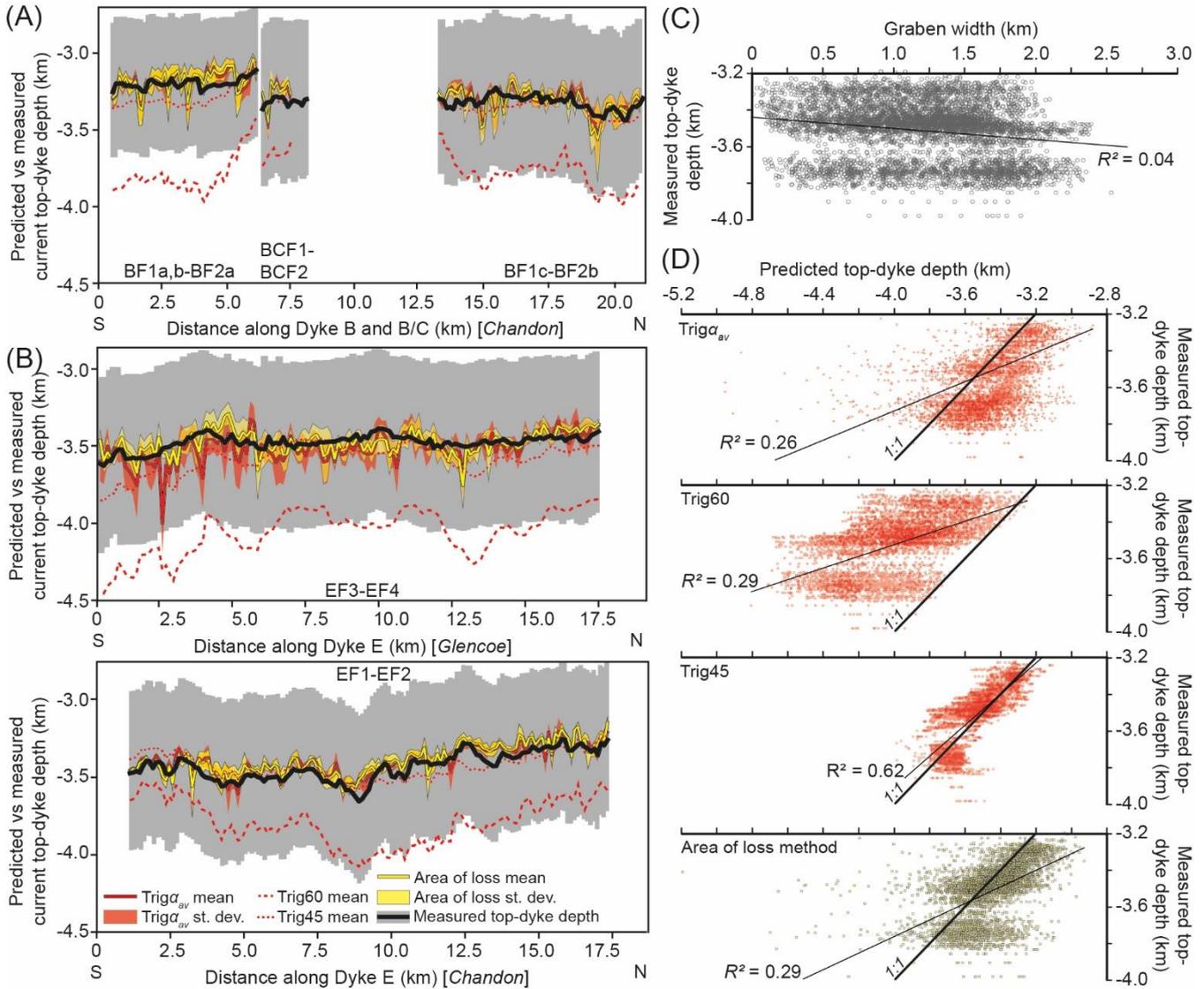
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#### 463 **4.4. Graben widths and predicted top-dyke depths in 3D**

464 To better evaluate the validity of top-dyke depth prediction methods, and extend the 3D analysis of DF1-  
 465 DF2 and DF3-DF4 (Fig. 7), we measured graben width across all mapped horizons above Dyke B in the  
 466 Chandon 3D survey and Dyke E in the Chandon and Glencoe 3D surveys (all data provided in Magee,  
 467 2022, Magee & Love, 2021). From these graben width measurements, we estimate the top-dyke depth  
 468 beneath each horizon (e.g., Figs 7 and 9). The measured top-dyke depths beneath current sea-level range  
 469 from ~3.3–3.7 km and ~3.3–3.6 km for dykes B and E, respectively (Figs 9A and B). Along the graben,

470 over lateral distances of  $\sim 125\text{--}250$  m, there are some abrupt changes of up to  $\sim 100$  m in measured top-  
 471 dyke depth superimposed onto broad increases and decreases of up to  $\sim 300$  m over the kilometre-scale  
 472 (Figs 9A and B). By compiling data for all stratigraphic horizons, we show there is no correlation ( $R^2=0.04$ )  
 473 between measured graben widths and measured top-dyke depths (Fig. 9C).

474



475

476 **Figure 9:** (A-B) Plots comparing measured top-dyke depths with mean predicted top-dyke depths  
 477 estimated from all horizons using different methods ( $\text{trig}\alpha_{av}$ , trig60, trig 45, and area of loss) for dykes  
 478 B and E in the Chandon and Glencoe 3D surveys. Standard deviation envelopes are shown for the  
 479  $\text{trig}\alpha_{av}$  and area of loss methods, and the error envelope (grey) of the measured top-dyke depths is  
 480  $\pm 15\%$ . (C) Measured graben widths from all horizons compared to underlying top-dyke depths. (D) All

481 measured top-dyke depths compared to predicted top-dyke depths estimated from the trig $\alpha_{av}$ , trig60,  
482 trig 45, and area of loss methods used.

483

484 There is often scatter between the predicted top-dyke depths from different horizons, but mean  
485 values display relatively low ( $\lesssim 200$  m) standard deviations (Figs 9A and B). Like our analysis of predicted  
486 top-dyke depths predicted from graben widths along the Top Mungaroo Formation, we show predictions  
487 derived from the trig $\alpha_{av}$ , trig45, and area of loss methods appear broadly comparable to measured top-  
488 dyke depths (Figs 9A and B). In contrast, top-dyke depths predicted using the trig60 method appear to  
489 overestimate dyke depths (Figs 9A and B). If we compare all top-dyke depth predictions, and not just  
490 derived means, those calculated using the trig $\alpha_{av}$  and area of loss methods are positively but weakly  
491 ( $R^2=0.26$  and  $0.29$ ) correlated to measured top-dyke depths; these methods have Root-Mean Square  
492 Errors (RSME) of  $\sim 180$  m and seemingly tend to underestimate top-dyke depths (Fig. 9D). The trig60  
493 method is similarly positively but weakly ( $R^2=0.29$ ) correlated to measured top-dyke depths, but  
494 overestimates top-dyke depths and has an RSME of  $\sim 500$  m (Fig. 9D). Top-dyke depth predictions derived  
495 from the trig45 method display a moderate ( $R^2=0.62$ ), positive correlation with the measured top-dyke  
496 depths, and a RSME of  $\sim 100$  m (Fig. 9D).

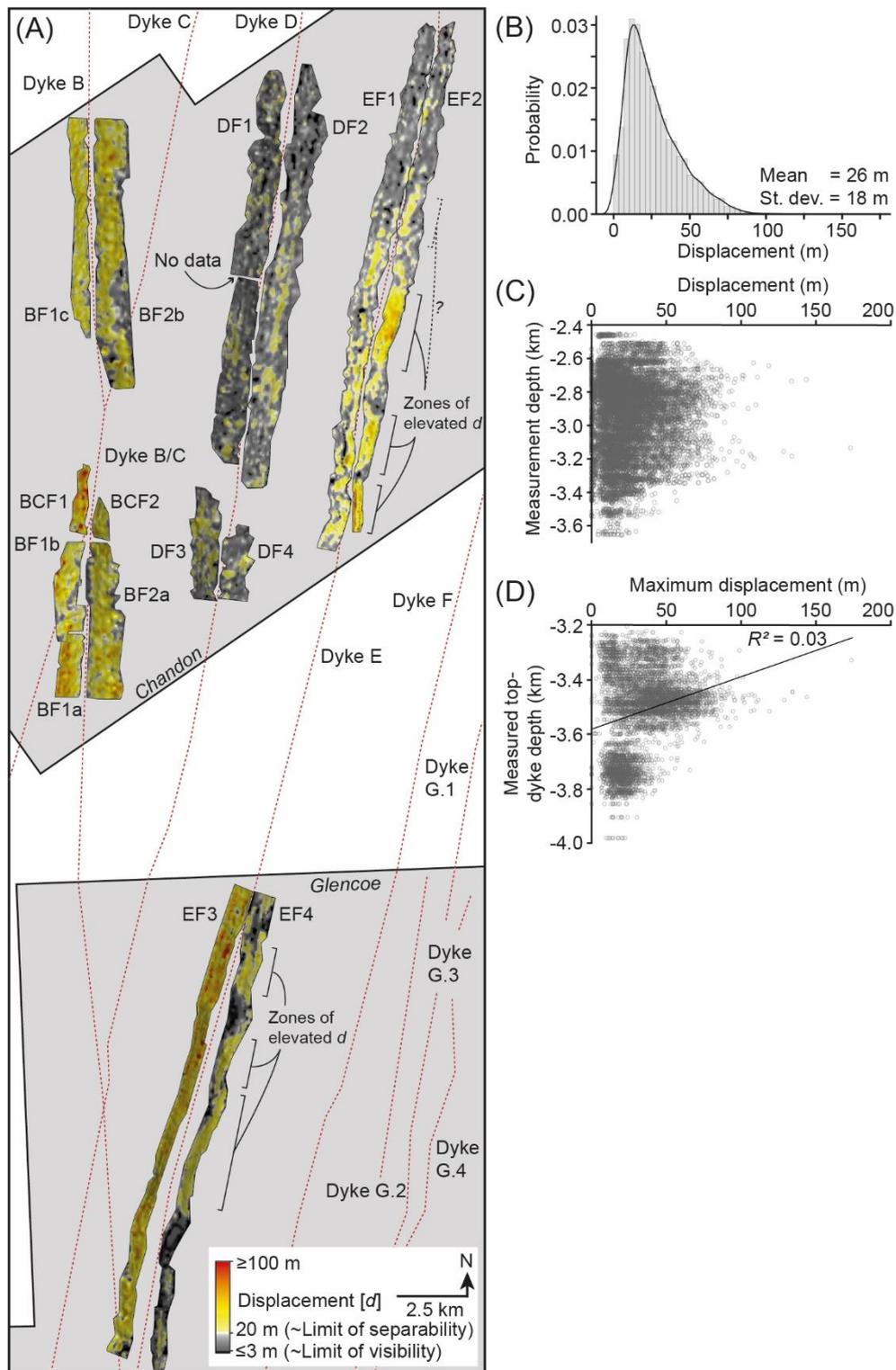
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#### 498 **4.5. Fault displacement and dip in 3D**

499 We show how displacement varies along-strike and down-dip of all studied dyke-induced faults (Figs 10A)  
500 (all data provided in Magee, 2022, Magee & Love, 2021). Across these dyke-induced faults, displacement  
501 has a right-skew distribution but is  $\sim 26$  m on average with a standard deviation of  $\sim 18$  m (Fig. 10B).  
502 Displacement maxima rarely occur at fault upper or lower tips, with displacement typically higher towards  
503 fault centres (Figs 10A and C). Across EF1 and EF2 in the Chandon 3D survey, displacement gradually  
504 decreases northwards and is consistently greatest (up to  $\sim 101 \pm 15$  m) on the W-dipping fault (EF2) where  
505 three, possibly four, zones of locally elevated displacement can be recognised (Fig. 10A) (see also Magee  
506 & Jackson, 2020b). There are no such zones of elevated displacement along EF1 (Fig. 10A). In the Glencoe  
507 Figure 11: (A) Fault dip ( $\alpha$ ) maps (see Magee & Jackson, 2020b for method). Dyke traces also shown

508 (Magee & Jackson, 2020a). Fault dip maps may contain interpolation errors derived from our choice  
509 gridding algorithms, but we consider these negligible given the high density of measurement locations  
510 across both faults. (B) Probability density function plot of measured average fault dips ( $\alpha_{av}$ ), showing  
511 they have a normal distribution. (C) Fault dip compared to the depth of the hanging wall cut-off used in  
512 its calculation. 3D survey, it is the E-dipping fault, EF3, which consistently has the greatest displacement  
513 (up to  $\sim 134 \pm 18$  m) (Fig. 10A). Yet EF3 shows no clear zones of elevated displacement, whereas EF4 can  
514 be sub-divided into at least three zones (Fig. 10A). Where zones of elevated displacement occur along EF2  
515 and EF4, they broadly appear to correlate with areas of higher fault dip (Fig. 10). For the dyke-induced  
516 faults above dykes B and D, displacement gradually decreases northwards but there are no obvious zones  
517 of locally elevated displacement (Fig. 10A). The dyke-induced faults above Dyke D have lower  
518 displacement (up to  $\sim 65 \pm 10$  m) than those above dykes B (up to  $\sim 173 \pm 26$  m) and E (Fig. 10A). There  
519 is not statistical correlation between measured top-dyke depths and maximum displacements measured  
520 on each horizon at each transect, but there is a hint that dyke-induced faults accrue more displacement  
521 when dykes reach shallower levels (Fig. 10D).

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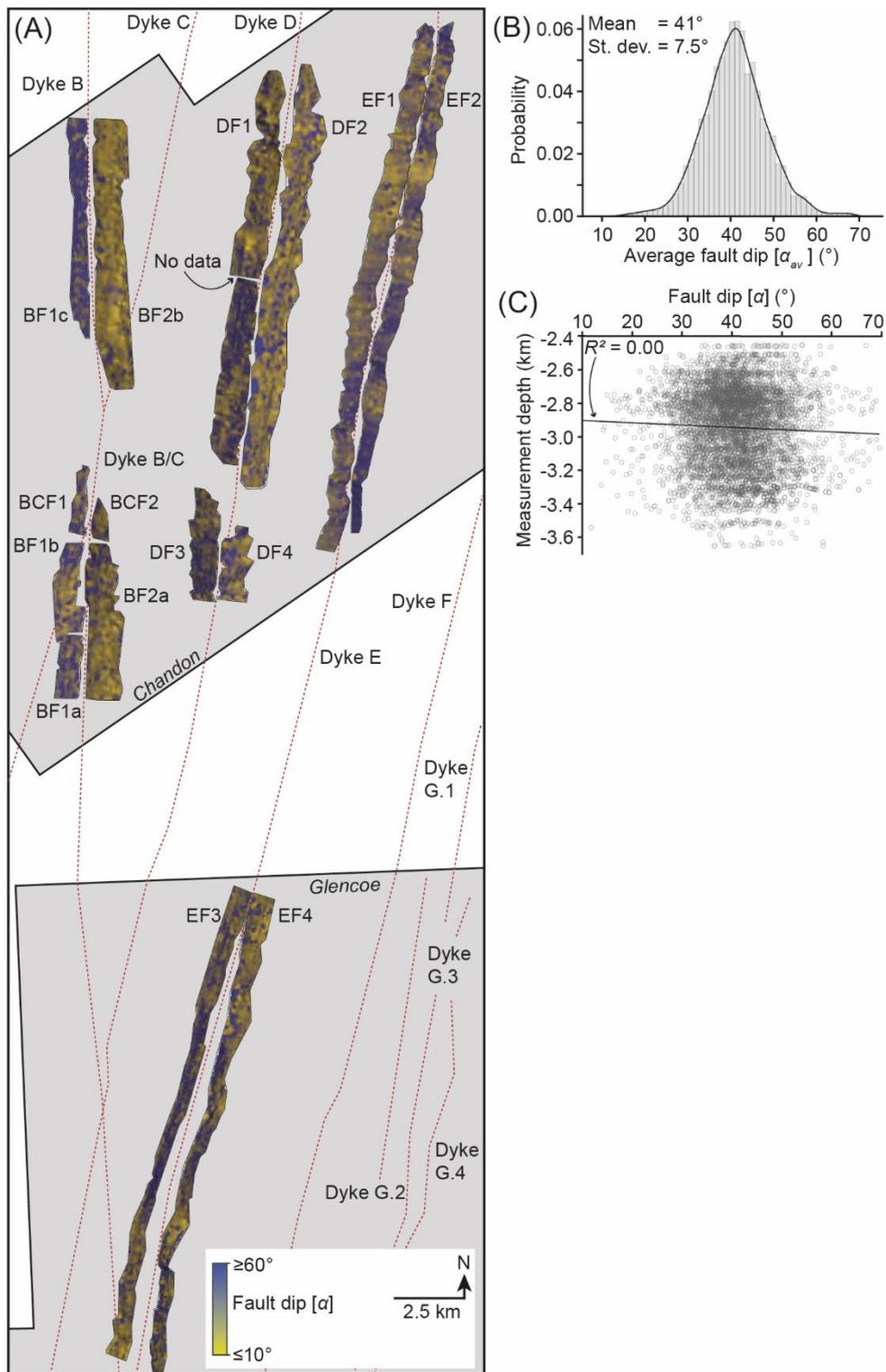
524 **Figure 10:** (A) Fault displacement maps (see Magee & Jackson, 2020b for method). Dyke traces also  
 525 shown (Magee & Jackson, 2020a). Fault displacement maps may contain interpolation errors derived  
 526 from our choice gridding algorithms, but we consider these negligible given the high density of  
 527 measurement locations across both faults. (B) Probability density function plot of measured fault  
 528 displacements, showing they have a right-skewed distribution. (C) Fault displacement compared to the  
 529 depth of the hanging wall cut-off used in its calculation. (D) Maximum displacement of the two

530 opposing dyke-induced faults on each measurement transect at each horizon, compared to the measured  
531 underlying top-dyke depth.

532

533 We show dip varies along-strike and down-dip of all studied dyke-induced faults, although there  
534 are no systematic changes in its distribution (Figs 11A). Average fault dips, calculated from the dips of  
535 both opposing dyke-induced faults on each transect at each horizon, are normally distributed with a mean  
536 of  $41^\circ$  and standard deviation of  $7.5^\circ$  (Fig. 11B). There is no systematic change in average fault dip with  
537 depth (Fig. 11C).

538



539

540 **Figure 11:** (A) Fault dip ( $\alpha$ ) maps (see Magee & Jackson, 2020b for method). Dyke traces also shown  
 541 (Magee & Jackson, 2020a). Fault dip maps may contain interpolation errors derived from our choice  
 542 gridding algorithms, but we consider these negligible given the high density of measurement locations  
 543 across both faults. (B) Probability density function plot of measured average fault dips ( $\alpha_{av}$ ), showing  
 544 they have a normal distribution. (C) Fault dip compared to the depth of the hanging wall cut-off used in  
 545 its calculation.

## 547 5. Discussion

## 548 5.1. Interpretation bias

549 The studied dyke-induced faults have relatively low displacements (typically <75 m, with an average of  
550 ~26 m; Fig. 10B). Given the vertical ( $\sim 20 \pm 4$  m) and horizontal ( $\sim 25$  m) resolution limits of the seismic  
551 reflection data used, these low displacements mean many portions of the dyke-induced faults are not  
552 imaged as clear offsets in reflections (e.g., Fig. 5i). Instead, the dyke-induced faults often correspond to  
553 areas where seismic reflections are subtly distorted and appear folded (e.g., Figs 3A, B, and 5iii). In such  
554 a situation, a dyke-induced fault may extend through the distorted reflections, but the interpreter has to  
555 decide whether the fault intersects the middle of the rotated reflection limb, either of its inflection points,  
556 or elsewhere (e.g., Fig. 5iii) (e.g., Faleide et al., 2021). It is also plausible that the distorted reflections  
557 correspond to folded strata, perhaps generated by fault propagation folding above a buried fault tip, and  
558 there is no fault present at that stratigraphic level at all. These uncertainties, associated with mapping fault  
559 cut-offs, affect all quantitative fault measurements (e.g., displacement and graben width) (e.g., Alcalde et  
560 al., 2017a, Faleide et al., 2021). In their quantitative dyke-induced fault study, Magee and Jackson (2020b)  
561 assumed this interpreter bias could be accounted for by considering measured or calculated fault  
562 properties had errors of  $\pm 5\%$ .

563 In our study, where the same fault cut-off sets were mapped by the same interpreter, but at a  
564 different time (i.e. CM1 and CM2), variations in average dip and cumulative displacement occur due to  
565 supposedly coincident fault cut-offs being mapped up to  $\sim 100$  m along-strike ( $\sim 18$  m on average)  
566 (Supplementary Table S3). These distances have the same order of magnitude as fault heave and throw,  
567 and therefore significantly affect calculated dips and displacements. Conversely, the  $\lesssim 100$  m differences  
568 in fault cut-off position are much less than, and thus have little impact on graben widths ( $>1500$  m;  
569 Supplementary Table S2). Despite these differences in fault cut-off position, our statistical F-test confirm  
570 little variance between fault properties extracted from the two compared datasets (i.e. CM1 and CM2),  
571 indicating they can be considered equal (Table 1). Additionally, top-dyke depths predicted from both  
572 datasets using trigonometry methods are statistically equal; however, those predicted using the area of loss

573 method are unequal (Table 1). Our data thus suggest that although an individual interpreter may  
574 introduce bias into their fault cut-off mapping, their findings are generally consistent. Such ‘internal  
575 consistency’ from geologists has been observed for several data types, including the characterisation of  
576 faults and fractures (Andrews et al., 2019, Shipton et al., 2020) and models built from seismic reflection  
577 datasets (Alcalde & Bond, 2022); this is likely a result of the mental model(s) of the interpreter (i.e. the  
578 simplified internal representation of the process or problem being assessed) (e.g., Gibson et al., 2016,  
579 Shipley & Tikoff, 2016). These mental models are influenced by the cogitative style and experience of the  
580 interpreter (Bond et al., 2007, Shipley & Tikoff, 2016), as well as the purpose and time constraints  
581 associated with data collection and quality control (e.g., Shipton et al., 2020).

582         In our study, where fault properties and predicted top-dyke depths are derived from fault cut-offs  
583 mapped by *different* interpreters, their values are broadly similar and the same profile patterns along the  
584 graben length are reproduced (Figs 6 and 8). Our statistical F-test shows that, in some instances, fault  
585 properties and predicted top-dyke depths can be considered equal, but not always (Tables 1 and 2). There  
586 are many potential reasons why data derived from different interpreters is variable, including: (1) seismic  
587 reflection data quality (Alcalde et al., 2017a, Alcalde et al., 2017b, Faleide et al., 2021); (2) interpreter  
588 experience (Bond et al., 2007, Bond et al., 2015); (3) applied methods of mapping (e.g., were seismic  
589 attributes used to constrain interpretations?) (Rankey & Mitchell, 2003); (4) the vertical exaggeration and  
590 scale used during fault picking (Alcalde et al., 2019); and (5) the time available to collect and quality check  
591 the dataset (e.g., Bond et al., 2007, Bond et al., 2015, Faleide et al., 2021, Macrae et al., 2016). Previous  
592 work has shown that variance between interpreters can be reduced where a pre-interpretation picking  
593 strategy was implemented, training material was used by all interpreters prior to picking, and/or an  
594 element of group working or training is employed (see Alcalde & Bond, 2022 and references therein).  
595 Overall, our data suggest that the arbitrary 5% error applied by Magee and Jackson (2020b) may be  
596 insufficient to describe extracted fault properties, with errors of  $\gg 5\%$  observed. Therefore, until a more  
597 detailed parametric study can be carried out, we suggest that arbitrary but conservative errors of 10% are  
598 used, particularly where fault displacement is close to or below the separation limit of the dataset.

599

## 600 5.1. *Predicting top-dyke depths from dyke-induced fault surface expressions*

601 Our data enable us to examine how predicted top-dyke depths, estimated using several methods, compare  
602 to measured depths. For example, we test the predictive power of the trigonometry method using an  
603 average of the measured faults dips ( $\text{trig}\alpha_{av}$ ), which varies on each transect, compared to where fault dip  
604 is assumed constant (at  $60^\circ$  or  $45^\circ$ ). We show that predictions of top-dyke depths made using graben  
605 width, fault dip, and/or displacement measurements are typically within error of measured top-dyke  
606 depths (Figs 6-9). Despite these similarities, predictions made using the  $\text{trig}\alpha_{av}$  and area of loss methods  
607 typically underestimate top-dyke depths and display only a weak, positive relationship to measured top-  
608 dyke depths ( $R^2 = 0.26$  and  $0.29$ , respectively) (Fig. 9D). These discrepancies and weak correlation  
609 between  $\text{trig}\alpha_{av}$  and area of loss method predictions and measured top-dyke depths relate to the  
610 incorporation of measured, as opposed to assumed, fault dip data. Fault dip measurements are susceptible  
611 to interpretation bias and vary across the fault plane over short (100–200 m) length-scales (Fig. 11A).  
612 Instead, projection of a fault with such variations in dip may be best achieved by using a constant dip  
613 representative of the entire fault. Although assuming dyke-induced faults dip consistently at  $60^\circ$  leads to  
614 overestimates of top-dyke depths, our data do show that assuming a consistent dip of  $45^\circ$  provides a good  
615 fit between predicted and measured top-dyke depths, likely because it is similar to the mean ( $41^\circ$ ) of all  
616 fault dip measurements (Figs 9D and 11B). Our data support previous studies that use constant fault dips  
617 to predict top-dyke depths (e.g., Hjartardóttir et al., 2016), implying we can relate the surface expression  
618 of dyke-induced faults, at least to some extent, to underlying dyke geometry (cf. Magee & Jackson, 2020b).  
619 However, we note that without 3D imaging of dyke-induced fault planes, we cannot ascertain whether the  
620 average of fault dips measured at a single surface (e.g., a planetary surface), or an assumed fault dip (e.g.,  
621  $60^\circ$  or  $45^\circ$ ), are representative of its dip variations with depth. For example, if fault dips continuously  
622 decrease with depth, such that the faults are concave-upwards, assuming fault dip is constant will result  
623 in overestimating top-dyke depths (Magee & Jackson, 2020b).

624

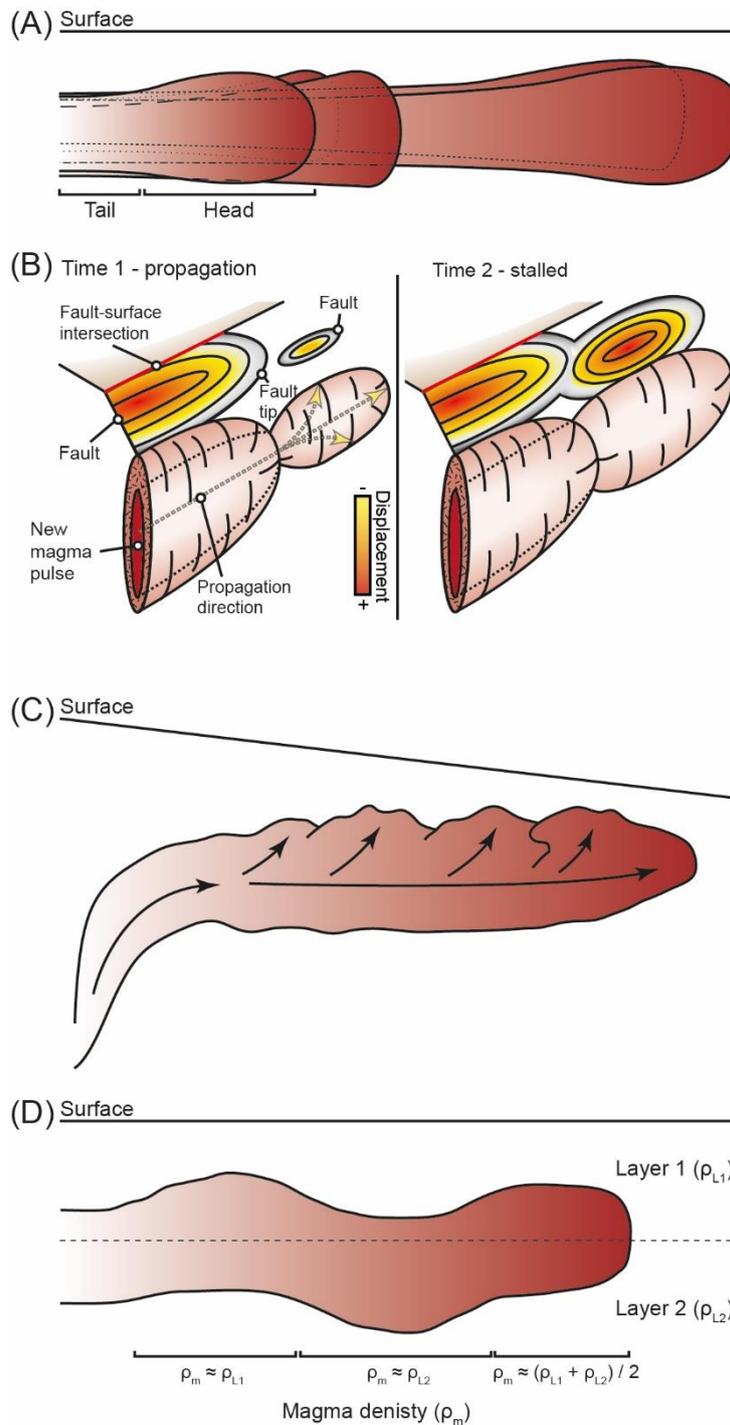
### 625 5.1.1. *Top-dyke depths*

626 Dyke upper tips are currently located  $\sim 3.3\text{--}3.8$  km below sea-level (Figs 6-9), and extended upwards to  
627 depths of at least  $\sim 1$  km when they were intruded, assuming horizon HK marks the syn-emplacement free  
628 surface (Magee & Jackson, 2020a, Magee & Jackson, 2020b). Both our measurements and predictions  
629 suggest top-dyke depths vary along-strike by several hundred metres, with changes either being abrupt or  
630 gradual over several kilometres (Figs 6-9). Abrupt changes in along-strike top-dyke depths are within  
631 error, and so they could be an artefact of our interpretation (Figs 6-9). However, dykes B and E, and to a  
632 lesser extent Dyke D, show systematic increases and decreases in upper tip depth of up to  $\sim 300$  m, with  
633 wavelengths of several kilometres (Figs 6-9); similar fluctuations in top-dyke tip depth, albeit on length  
634 scales of 10's of kilometres, have been shown for other members of the Exmouth Dyke Swarm (Magee &  
635 Jackson, 2020a) and for dykes in Elysium Fossae, Mars (Rivas-Dorado et al., 2021). Kilometre-scale, along-  
636 strike variations in top-dyke depth have been observed in recent dyking episodes (e.g., Ágústsdóttir et al.,  
637 2016, Tripanera et al., 2019, Xu et al., 2016), and may be a common feature of dyke swarms (e.g., Magee  
638 & Jackson, 2020a, Rivas-Dorado et al., 2021).

639         There are several reasons why the upper tip depth of laterally propagating dykes may increase and  
640 decrease along-strike. For example, a head region may develop at the propagating edge of a dyke, which  
641 is taller and reaches a shallower level than the tail (Figs 12A and B) (e.g., Rivalta et al., 2015, Rivas-  
642 Dorado et al., 2021). If the seismic expression of the intrusions corresponds to multiple, adjacent dykes  
643 (Magee & Jackson, 2020a), differences in the distance each dyke propagated could cause top-dyke depths  
644 to vary beneath the graben (Fig. 12A) (Rivas-Dorado et al., 2021). Alternatively, propagation of a dyke  
645 (with or without a head) may have stalled, leading to its crystallisation and pressurisation until magma  
646 broke out from dyke 'nose' to form a new segment that grew laterally and vertically (Fig. 12B) (Healy et  
647 al., 2018, Magee & Jackson, 2020b). A transition to vertical or inclined magma flow towards dyke tops can  
648 also cause dyke upper tips to segment and locally propagate vertically, affecting dyke height (e.g., Poland  
649 et al., 2008). Such dyke segmentation can particularly occur when emplacement occurs beneath a volcano  
650 load (Fig. 12C) (e.g., Poland et al., 2008), as the sloping topography results in changes to the lithostatic  
651 stresses above and along underlying dykes (Urbani et al., 2017). Density layering in the host rock is  
652 another relevant factor that controls the vertical stability of dykes, and therefore their depths; stratified

653 host rocks are more likely to contain dykes with smaller vertical extents, which are stable under a wider  
654 range of conditions (Pollard & Townsend, 2018). Lateral changes in the vertical stratigraphic sequence  
655 may thus control how close to the free surface a dyke may reach. Finally, increases or decreases in magma  
656 density may cause dyke upper tips to become unstable and adjust to a new structural levels (Fig. 12D)  
657 (e.g., Townsend et al., 2017). Regardless of the driving mechanism, areas where top-dyke depths become  
658 shallower may focus magma flow and promote local upwards propagation, potentially leading to and  
659 explaining the restricted distribution of fissure eruptions along dyke lengths (e.g., Pansino et al., 2019,  
660 Woods et al., 2019).

661



662

663 **Figure 12:** Possible mechanisms for changing top-dyke depths along-strike. (A) Laterally propagating  
 664 dykes develop a head region that is taller and extends to shallower levels than the tail behind (modified  
 665 from Rivas-Dorado et al., 2021). (B) Cyclical stalling and propagation of the dyke leads to magma  
 666 breaking out from the dyke 'nose', where it is vertically restricted (Magee & Jackson, 2020b). (C) Under  
 667 a variable volcano load, the ascent and lateral propagation of a dyke results in segmentation of its upper  
 668 tip, which locally focuses magma flow and drives upward propagation (modified from Poland et al.,  
 669 2008). (D) Changes in magma density, relative to that of a host rock displaying a density stratification,

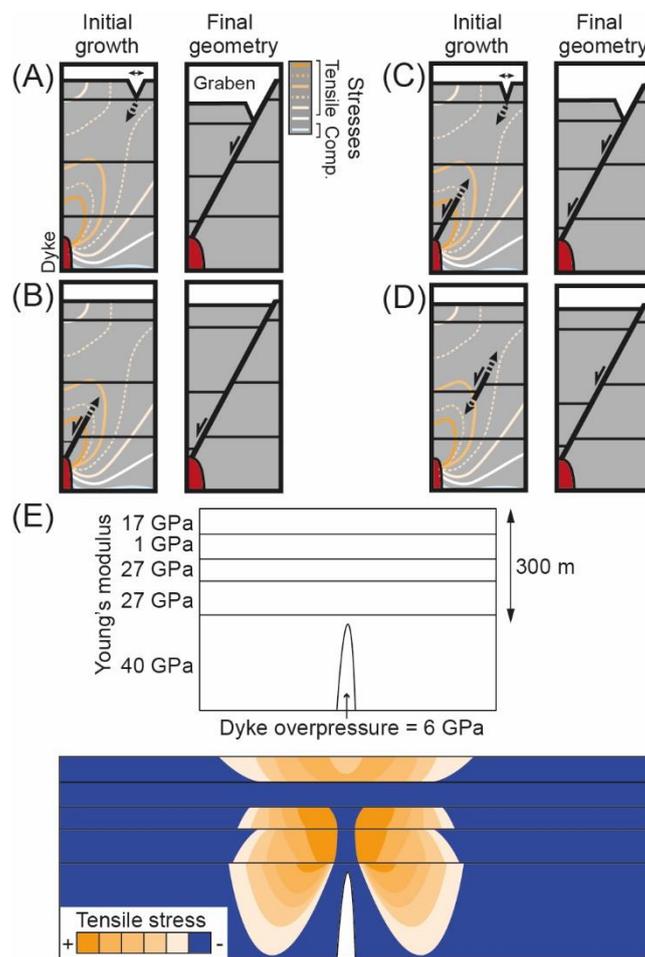
670 can cause the upper and lower tip positions of a propagating dyke to destabilize and transgress upwards  
671 or downwards (Townsend et al., 2017). Where the magma density is equal to the average of two host  
672 rock layers of different densities, the dyke upper and lower tips are equally spaced from the layer  
673 boundary (Townsend et al., 2017). If magma density changes (e.g., through vesiculation, crystallisation,  
674 and/or cooling) and becomes closer in density to one of the layer, it will preferentially intrude within  
675 that layer, causing the upper and lower tips to move (Townsend et al., 2017).

676

## 677 ***5.2. Dyke-induced fault growth and kinematics***

678 Field- and geophysical-based observations, coupled with physical, analytical, and numerical models  
679 suggest dyke-induced faults probably nucleate either: (1) near the surface, initiating as vertical tensile  
680 fractures that propagate downwards and develop into shear fractures (Fig. 13A) (e.g., Al Shehri &  
681 Gudmundsson, 2018, Trippanera et al., 2015a, Trippanera et al., 2015b, Von Hagke et al., 2019); (2) as  
682 shear fractures that grow upwards from dyke upper tips (Fig. 13B) (e.g., Koehn et al., 2019, Rubin, 1992,  
683 Xu et al., 2016); (3) at both the surface and dyke upper tip, linking as they propagate downwards and  
684 upwards, respectively (Fig. 13C) (Rowland et al., 2007, Tentler, 2005); or (4) between the surface and  
685 dyke tip, propagating both upwards and downwards (Fig. 13D) (Koehn et al., 2019, Magee & Jackson,  
686 2020b, Mastin & Pollard, 1988).

687



688

689 **Figure 13:** (A-D) Dyke-induced fault growth models (based on Koehn et al., 2019, Magee & Jackson,  
 690 2020b, Mastin & Pollard, 1988, Rubin & Pollard, 1988, Tentler, 2005, Trippanera et al., 2015b).

691 Horizontal stress patterns above an intruding dyke in a homogeneous elastic medium, showing tensile  
 692 stress is concentrated at the surface and above the dyke tip (redrawn from Rubin & Pollard, 1988). (E)

693 Numerical model of tensile stresses above a dyke arrested at 300 m below the surface, with an  
 694 overpressure of 6 GPa, in a layered medium (redrawn from Al Shehri & Gudmundsson, 2018; their Fig.  
 695 21). Tensile stresses are concentrated above the dyke tip and in two zones at the surface, but the  
 696 presence of layering disrupts its distribution; e.g., the weak layer, which has a Young's modulus of 1 GPa,  
 697 suppresses stress concentration (Al Shehri & Gudmundsson, 2018).

698

699 If we expect displacement to be greatest where faults nucleate (e.g., Pollard & Segall, 1987,  
 700 Trippanera et al., 2015b), our mapped displacement distributions suggest fault segments initially  
 701 developed between the dyke upper tip and coeval free surface (Figs 10A and C) (Koehn et al., 2019, Magee

702 & Jackson, 2020b, Mastin & Pollard, 1988). Yet dyke opening induces tensile stresses in the overburden,  
703 focused at the dykes upper tip and in two zones at the free surface, where modelling suggest fault  
704 nucleation most likely to occur (Fig. 1A) (e.g., Koehn et al., 2019, Pollard et al., 1983, Rubin, 1992, Rubin  
705 & Pollard, 1988). However, these models assume dyke intrusion occurs within a homogenous, elastic half-  
706 space (e.g., Koehn et al., 2019, Pollard et al., 1983, Rubin, 1992, Rubin & Pollard, 1988). Where mechanical  
707 layers with different physical properties are modelled above dykes, tensile stresses have been shown to  
708 become concentrated in relatively strong units away from the dyke tip or free surface (Fig. 13E) (Al Shehri  
709 & Gudmundsson, 2018). The dyke-induced faults we analyse primarily offset heterolithic fluvio-deltaic  
710 rocks (e.g., Bilal et al., 2018, Martin et al., 2018), as well as Jurassic sandstones and claystones (Figs 2B,  
711 3A and B) (Tindale et al., 1998). The layered sedimentary succession offset by the dyke-induced faults is  
712 thus likely mechanically heterogeneous, which could have affected stress distribution and fault nucleation  
713 during dyking (e.g., Al Shehri & Gudmundsson, 2018, Schöpfer et al., 2006). Alternatively, displacement  
714 may preferentially accrue on faults away from their nucleation site if post-emplacment dyke thickening  
715 instigated slip (Magee & Jackson, 2020b); in this scenario, displacement distribution cannot be used to  
716 reconstruct dyke-induced fault kinematics. Overall, we suggest it likely there are multiple controls on  
717 where dyke-induced faults nucleate, which could make it difficult to assess their relation to dyke  
718 emplacement and structure.

719

## 720 6. Conclusions

721 Dyke-induced faults are common in many volcanic settings on Earth and other planetary bodies (e.g.,  
722 Mars). Because we often have no or little access to the subsurface in these locations, relating the surface  
723 expression of dyke-induced faults to the underlying dyke geometry can provide key insight into the 3D  
724 structure of and processes active during dyke emplacement. However, deriving dyke geometry from the  
725 surface expression of dyke-induced faults is only feasible if the faults project straight down-dip to the dyke  
726 upper tip. Yet recent analysis of 3D seismic reflection imagery suggests dyke-induced faults are non-planar.  
727 We show that interpretation bias does introduce uncertainty into the quantitative analysis of dyke-induced  
728 faults using seismic reflection data, but that we can still use these data to confidently understand dyke and

729 dyke-induced fault structure in 3D. By quantitatively measuring dyke-induced fault and graben  
730 geometries, we demonstrate that: (1) predictions of top-dyke depths are typically within error of measured  
731 top-dyke depths, but most accurate when incorporating information on the average dip of entire faults;  
732 and (2) dyke-induced fault displacement and dip are variable along-strike and down-dip. Our work  
733 supports previous findings and relates dyke-induced faulting to the nucleation of isolated segments above  
734 propagating dyke segments, which linked as they grew in response to dyke stalling and thickening. Cyclical  
735 dyke propagation and stalling, or other processes leading to the emplacement of dyke segments, may also  
736 explain the hectometre-scale variation in dyke upper tip depths we observe along dyke strike. Overall, our  
737 work suggests that we can relate the surface expression of dyke-induced faults to subsurface dyke  
738 geometry, but only if we have some information on the 3D structure of the faults. Reflection seismology  
739 is a critical tool for developing our understanding of dyke-induced faulting.

740

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745 the provision of the seismic interpretation software.

746

#### 747 **Author contributions**

748 CM conceived and designed the work, acquired the data, contributed to data analysis and interpretation,  
749 and wrote the manuscript. Both VL and KF were involved in project design and data analysis, and proof-  
750 read the manuscript. BA, SR-D, CO, and EB were involved in data analysis and interpretation, and critically  
751 revised the manuscript.

752

#### 753 **Data availability statement**

754 Seismic reflection and borehole data used in this study are freely available from either the Geoscience  
755 Australia NOPIMS (<https://nopims.dmp.wa.gov.au/nopims>) data repository or that of the UK

756 National Geoscience Data Centre (NGDC; [https://www.bgs.ac.uk/geological-data/national-](https://www.bgs.ac.uk/geological-data/national-geoscience-data-centre/)  
757 [geoscience-data-centre/](https://www.bgs.ac.uk/geological-data/national-geoscience-data-centre/)). The NOPIMS data access centre allows “Wells” (i.e., for this study Chandon-  
758 1, Chandon-2, Chandon-3, Yellowglen 1, Briseis-1, Dunlop-1, Elfin-1, Glencoe-1, and Toporoa-1) and  
759 “Surveys” (i.e., for this study the Chandon 3D MSS and Glencoe 3D seismic reflection surveys) from  
760 offshore Australia to be searched for. From these search results, borehole data can be downloaded by  
761 highlighting the correct borehole and using “view details for selected rows”; due to the file size of 3D  
762 seismic reflection segy data, few can be downloaded but all can be added to a basket and requested from  
763 Geoscience Australia. These data can also be downloaded from the NGDC under the title “3D seismic  
764 reflection surveys (Chandon and Glencoe) and borehole data from offshore NW Australia.” All  
765 measurements and calculations acquired during this research, as well as mapped seismic horizons, are  
766 provided as supplementary information or hosted by the NGDC (Magee, 2022, Magee & Love, 2021).

767

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