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Tectonics is a Hologram

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Tectonics is a hologram

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Abstract

A hologram is an image in which each area contains almost all the information about the entire system. It is a metaphor commonly used for complex systems in which the whole is bigger than the sum of the parts because of self-organization. And also the whole is smaller than the sum of the parts, since the collective organization limits the behavior of dynamic features. The tectonic evolution of the Earth is an emergent behavior of the lithosphere-mantle system, a witness of a program defined at the scale of rocks. Modeling the physics behind tectonics at a global scale became a reachable goal entering the 21st century. Geodynamicists developed numerical models of solid-state convection with yielding, and reproduced some fundamentals of planetary tectonics. In the past 15 years, several groups in the world have used these models to investigate how continents drift, seafloor spreads and

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plates evolve. These emergent characteristics tell that the whole is bigger than the sum of the parts. Slabs, plumes, ridges, plates are interdependent and constrain each other. The whole is smaller than the sum of the parts. In this context, searching for causality relationships between tectonic features seems vain. In this chapter, I consider this point of view to describe how convection models with yielding have changed and still change our views on how tectonics shape the Earth. I finally propose an outlook about this window that remains half-opened.

Keywords: Mantle convection, plate tectonics, numerical modelling, subduction, complex systems

1. Introduction

On every planet there is active tectonics, and there has been active tectonics. Even if it is extremely weak, macroscopic finite strain patterns shape their surface and draw their uniqueness. These structures are the way we identify them, traits of their identity and keys to their history. On Mercury, hundreds of kilometers long fault scarps cutting impact craters witness recent activity at planetary scale (Watters et al., 2016). Venus displays tectonic and volcanic structures, being rift systems (Olapa Chasma) interacting with volcanic eruptions (Idunn Mons), which have been identified as geologically recent (D’Incecco et al., 2020). The first months of seismic recording on Mars have evidenced tectonically active faults (Banerdt et al., 2020). Planetary-scale structures, as the Valles Marineris, show active deformation in the form of faults, landslides and possibly mud volcanoes (Kumar et al., 2019). Titan, Enceladus, Europa and Pluto each have their scars. And of course, the Earth

15 that we can observe more closely and connect the surface to its interior. Our
16 planet shows organized tectonics at a global scale, such that plate tecton-
17 ics has been developed as a successful framework Morgan (1968) to describe
18 its self-organization, linking together faults, volcanoes, deformed belts, sedi-
19 mentary deposits, remanent magnetism and even more, within a geometrical
20 and geodynamic logic. It can extend to climate and life. This extraordinary
21 organization of the surface of the Earth is the expression of a cascade of
22 interactions operating over a diversity of scales. At the microscopic scale,
23 relevant processes are the production of heat by unstable atomic nuclei, dif-
24 fusion of atoms and heat, chemical reactions, phase transformation in solids
25 and fluids, dislocation propagation in crystals and more. These small scale
26 processes interact with each other everywhere in the globe with a variety of
27 intensity and directions. They define a complex system, impossible to de-
28 scribe fully, highly non-linear, from which emerges a self-organization with
29 an apparent macroscopic behavior. Plate tectonics is a geometric theory to
30 describe this behavior.

31 In 1999, I started to work on the deep structure of the Earth's mantle
32 looking for relationships between dynamics and the geochemical diversity of
33 mantle-derived rocks (Coltice and Ricard, 1999). After studying the effects
34 of convective mixing on geochemical heterogeneity (Coltice and Schmalzl,
35 2006), I was convinced that the geochemical picture of the Earth was mostly
36 inherited from the early differentiation history of our planet, and I focused on
37 the formation of continents and the Earth's core (Labrosse et al., 2007; Mon-
38 teux et al., 2007; Rey and Coltice, 2008; Coltice et al., 2009, for example). I
39 enjoy fieldwork and observations so more and more, I looked at continents,

40 and started to make numerical models of convection with continental litho-
41 sphere (Coltice et al., 2007). The temperature dependence of viscosity was a
42 major factor for my conclusions. The problem was that a convection model
43 with temperature-dependent viscosity alone generates a sluggish surface at
44 best, and a one-plate planet at worst. Therefore, plate-like behavior was nec-
45 essary to model a mobile surface. I was lucky that Paul Tackley and Tobias
46 Rolf opened their door and allowed me to work with them. The first time
47 I watched movies of the models of Tobias, I was shocked. I could see plate
48 tectonics in operation. Dynamically. And in a second my mind switched.
49 I had the vision that these models would take us to a new description of
50 tectonics. I could see that the lithosphere dynamics and mantle convection
51 communities could reach their hands in terms of modeling and targets. Since
52 that time I have been working with these models and diving into complex
53 systems and self-organization. I could not consider anymore simple causal
54 links, but a systemic approach to the deformation over the whole globe.

55 Self-organization for tectonics is widely discussed, leading to abstract
56 models of interacting processes. For instance Miltenberger et al. (1993) have
57 created models where elasticity, threshold dynamics, and small-scale hetero-
58 geneity interact to produce relevant spatio-temporal patterns of faults. At
59 larger scale, Anderson (2002) proposed a vision where the interactions be-
60 tween stresses and temperature within the mantle-crust system generate the
61 emergent pattern of plate tectonics. The integrative work of Bird (2003)
62 was a step forward in the identification of self-organization at plate-scale by
63 describing a plate area distribution that can be modeled by a power law typ-
64 ical of fragmentation models (Sornette and Pisarenko, 2003). When there is

65 self-organization, it often becomes useless to separate a part from the whole.
66 For instance, questions like 'do plates make the mantle move, or does the
67 mantle make plates move ?' separate plates from the mantle, although they
68 form a single complex system. It can be rephrased: how can we link the
69 motion of the plates and the motion of the mantle? It is a holographic prin-
70 ciple (Morin, 1990): the parts are in whole, and information on the whole is
71 inscribed within the parts. Tectonics is then like a hologram: it is an image
72 in which each area contains almost the full information about the mantle-
73 lithosphere system. With such approach, causality becomes a fuzzy way to
74 tackle problems. For instance, the westward drift of the lithosphere is not
75 caused by the Pacific slabs. Westward drift and the Pacific slabs coexist at
76 the same instant and cannot be isolated to each other. They are the result
77 of a program, which could be compared to the genetic code in life sciences,
78 and a series of interactions at all scales. In this context, the whole is bigger
79 than the sum of the part: there is emergence. And also, the whole is smaller
80 than the parts: the system-specific constraints inhibit some behavior specific
81 to the parts.

82 In practice, geodynamicists studying the lithosphere design numerical
83 models with local equations and material properties defined at microscopic
84 scale. They look for the generation of macroscopic tectonic and magmatic
85 structures. Mantle convection models are a step forward in tackling the self-
86 organization question at a global scale, necessary to treat the lithosphere-
87 mantle system as a whole. The first models generating plate-tectonic like
88 patterns were published more than 20 years ago (Tackley, 1998; Trompert
89 and Hansen, 1998a; Moresi and Solomatov, 1998). The key to producing

90 plate-like behavior was to combine temperature-dependent viscosity and a
91 threshold model for the stress-strain rate relationship (Bercovici, 1993). Re-
92 cently, a similar step has been started for laboratory experiments using nat-
93 ural colloidal solutions (Davaille et al., 2017).

94 In this chapter, I will first explore some aspects of mantle convection
95 models producing tectonics to identify the program behind self-organization.
96 I will then focus on the emergent properties like subduction, remaining the
97 big fundamental for our planet. It leads to the relationships between the
98 parts and the whole. I finish with an outlook and final thoughts given the
99 promises of this approach.

100 **2. The program of plate-like tectonic emergence in convection mod-** 101 **els: pseudo-plasticity**

102 *2.1. Context*

103 Tectonics has been a driver of mantle convection concepts. Holmes (1931)
104 built his model towards a description of the mechanics of continental drift.
105 With the advent of seafloor spreading, Turcotte and Oxburgh (1967) intro-
106 duced a semi-analytical model of mantle convection providing a framework
107 to the heat-flow-age relationship in the ocean floor. Plate tectonics imposed
108 a new concept of rigid blocks moving in pieces (Morgan, 1968), which was
109 followed by two types of modeling. First, to push the initiative on man-
110 tle convection, McKenzie et al. (1974) conducted the first time-dependent
111 numerical models of convection. Alongside, Richter and McKenzie (1978)
112 studied arising forces in viscous flow models (convection without the time-
113 dependent heat conservation equation) with plates, inspired by the approach

114 of Forsyth and Uyeda (1975). From these roots, time-dependent convection
115 models diversified, following the development of computers and parallel su-
116 percomputers, and viscous flow models as well. The first family of models
117 reached a limit by the end of the 1990's: they were useful to study global
118 properties (average surface heat flow, velocity etc...), physics of the deep man-
119 tle and geochemistry varying parameterizations but by no means they could
120 generate predictions for geological observations. The second family of mod-
121 els started with rooted limits: they could help characterize the Earth today,
122 but they could not do so for a fully dynamical perspective (time-dependent).
123 Therefore, at the end of the 1990's geoscientists took two different paths:
124 (1) working on the non-linear rheology of rocks (Bercovici, 1993) and nu-
125 merical methods (Tackley, 1998; Trompert and Hansen, 1998a; Moresi and
126 Solomatov, 1998) to build numerical models of mantle convection generat-
127 ing self-consistently plate-like behavior, (2) working on convection models
128 with imposed plates at the surface (Gable et al., 1991; Bunge et al., 1998;
129 Monnereau and Quéré, 2001; McNamara and Zhong, 2005).

130 By definition, imposing plates means imposing rigidity at the surface
131 and precluding modeling tectonics. I, therefore, focus on the models tak-
132 ing the rheology path, pointing the reader to review articles for more com-
133 prehensive work on both families of models (Lowman, 2011; Coltice et al.,
134 2017a). In 1998, three numerical modeling studies were published with con-
135 vection models generating plate-like behavior (Tackley, 1998; Trompert and
136 Hansen, 1998a; Moresi and Solomatov, 1998). These models integrated the
137 work of Bercovici (1993), formulating the hypothesis to keep the viscous ap-
138 proach to the problem, meaning that plate-like behavior is a viscous response.

139 The author tests the power of a variety of viscosity laws to generate plate-
140 like characters, being localization of the deformation and equipartitioning of
141 poloidal/toroidal kinetic energy (Olson and Bercovici, 1991). After disquali-
142 fying power laws, Dave Bercovici points to pseudo-plastic laws that succeed
143 in generating localized low viscosity pseudo-plate boundaries circling very
144 viscous pseudo-plates, and producing substantial toroidal flow. This seminal
145 work created the foundations of the convection models producing plate-like
146 tectonics since.

To start the process, the three publications cited above introduced a
yield stress parameterization, leading to an expression of the viscosity at the
strength limit σ_y being

$$\sigma_y = \mu_y \dot{\varepsilon}_{II},$$

147 with μ_y the viscosity at the pseudo-plastic limit and $\dot{\varepsilon}_{II}$ the second invariant
148 of the strain rate. This formulation is combined with a temperature depen-
149 dence of viscosity, making it high when cold, and low when hot. It launched
150 the race to identify the conditions of existence of plate-tectonics. At low
151 yield stress a convective flow will produce a soft and deformed surface, at
152 high yield stress a stagnant lid forms, and in between is the land of a range
153 of plate-like regimes. The issue lies in the quest for yield stress values, most
154 models pointing to below 150 MPa (O’neill and Lenardic, 2007; Richards
155 et al., 2001; van Heck and Tackley, 2008; Rolf and Tackley, 2011; Arnould
156 et al., 2018). Solomatov (2004) noted that the lower end values for the
157 yield stress (10-30 MPa) are similar to the ”stress drop during earthquakes
158 (Kanamori, 1994), and the stresses which produce the observed trench to-
159 pography (Zhong and Gurnis, 1994)”. The yield stress values have been a

160 new parameter to expand our vision of super-Earths, setting the question of
161 the existence of plate-tectonics at a central position (Valencia et al., 2007;
162 O’neill and Lenardic, 2007; van Heck and Tackley, 2011). However, subtleties
163 in initial conditions and melting history may be more critical than the yield
164 stress value itself (Seales and Lenardic, 2021; Lourenço et al., 2018).

165 Obtaining a range of yield stress values for plate-like behavior was a
166 fundamental step, but also one leading to confusion. The value obtained
167 in models is mostly taken out of its context to be compared directly to
168 experiments and observations without awareness. As a matter of fact, what
169 is called a yield stress in mechanics of solids has substantially higher values
170 for mantle rocks than what is needed in geodynamic models (Demouchy et al.,
171 2013; Zhong and Watts, 2013; Hansen et al., 2019). This has been a criticism
172 of convection models with yielding since. However, such quick comparison is
173 not appropriate for several reasons:

- 174 • evaluation of dynamical models rely on force (or time) ratios being
175 dimensionless numbers. One specific dimensional value in a model can
176 be compared to a natural value only if all other dimensional model
177 values are appropriate for the Earth, which is impossible to know and
178 achieve.
- 179 • In pseudo-plasticity the ”pseudo” word matters. It is a viscous model
180 parameterizing a rheology that is not only viscous. Therefore, compar-
181 ing the stress/strength envelopes of viscous models with stress/strength
182 envelopes for the Earth’s lithosphere is not appropriate. The latter
183 combines elasticity, brittle failure, plasticity and viscosity, showing dif-
184 ferent domains for the relevant rheological behavior (it is more like a

185 phase diagram). The former is just a curve for a given viscosity law.
186 This said, comparison with awareness is relevant. It means that one has to
187 find a physical domain in which both approaches can meet. It is also relevant
188 to work with dimensionless numbers in convection models. I will propose an
189 approach in section 3.

190 *2.2. Without and with pseudo-plasticity*

191 Pseudo-plasticity impacts convection in all its dimensions. The most ob-
192 vious is kinematics at the surface. As shown in Fig.1, the kinematics of
193 convection without pseudo-plasticity can be fully described by a flow diverg-
194 ing from a plume and converging to downwellings that can be either circular
195 or slightly planar when a viscosity jump is imposed (Bunge et al., 1997).
196 The theoretical ridge, trench and transform structure of plate tectonics is
197 not produced. The apparent long wavelength of thermal heterogeneity in
198 the mantle, imaged by seismic tomography (Su and Dziewonski, 1991) and
199 consistent with the geoid Cazenave et al. (1989), does not emerge from ev-
200 ery convection model. Without pseudo-plasticity, three parameterizations
201 produce long-wavelength thermal heterogeneity: a viscous jump within the
202 mantle (Bunge et al., 1997; Zhong et al., 2007), the presence of continen-
203 tal rafts at the surface (Gurnis, 1988; Guillou and Jaupart, 1995) and the
204 presence of rigid plates (Gurnis and Zhong, 1991). In terms of flow beneath
205 the thermal boundary layer, convection models without lateral viscosity con-
206 trasts do not generate mantle drag, i.e. velocity at the surface is always
207 faster in its direction than below the boundary layer (Coltice et al., 2019),
208 and toroidal motion does not exist (Chandrasekhar, 1961). Regions of diver-
209 gence correspond exactly to upwellings.

210 With pseudo-plasticity, the planform of convection involves sheet-like
211 downwellings with properties consistent with subduction on Earth (Crameri
212 and Tackley, 2014; Coltice et al., 2019) and cylindrical plumes (Arnould
213 et al., 2020). The surface pattern is a network of divergent, convergent and
214 possibly transform boundaries (Langemeyer et al., 2021). These boundaries
215 are sharp or diffuse, making them sometimes hard to automatically identify
216 (Mallard et al., 2017). In such models, large viscosity gradients come into
217 play. A viscosity value imposes its flow scale. With variable viscosity, a diver-
218 sity of flow wavelengths interact (Arnould et al., 2018). Viscosity gradients
219 allow decoupling and, therefore, the interior flow and the surface can differ
220 strongly, especially above upwellings. The dominant wavelength of convec-
221 tion is directly connected to the value of the yield stress. In the plate-like
222 regime, the wavelength is long, being degree 2 in a sphere (van Heck and
223 Tackley, 2008; Yoshida, 2008; Rolf et al., 2012; Mallard et al., 2016). The
224 surface expression of the long-wavelength flow pattern is the distribution of
225 large plates of similar sizes, comparable to those of the North American plate
226 and up to that of the Pacific plate (Mallard et al., 2016). The presence of a
227 viscosity jump is not a prerequisite for the long wavelength. A stiff surface
228 produces long wavelength as expected. The viscosity gradients at the base
229 of plates are strong, 5 to 7 orders of magnitude over tens of kilometers in
230 numerical models. They generate small-scale flow beneath old seafloor as
231 instabilities (Coltice et al., 2017c), but also short and intermediate scales
232 around sinking slabs (Király et al., 2017; Arnould et al., 2018). Toroidal
233 flow represents the motion of lower viscosity (and generally faster) material
234 around more viscous (and generally slower) material. The decoupling caused

235 by viscosity gradients produces contrasts in kinematics between the surface
236 and the shallow mantle. The plates smaller than the larger ones are produced
237 by fragmentation forced by subduction tractions (Mallard et al., 2016), hence
238 ridges are most of the time not the places of mantle upwellings but passive
239 structures. Therefore, with pseudo-plasticity, surface patterns do not reflect
240 deep structures, contrarily to convection models without strain localization.
241 This observation is consistent with the fact that mantle drag becomes a force
242 to be considered in such models (Coltice et al., 2019).

243 *2.3. On temperature-dependent viscosity*

244 Temperature dependent viscosity is a fundamental ingredient in generat-
245 ing plate-like behavior: it produces the lid structure, as long as the viscosity
246 contrast over the thermal boundary layer is larger than 10^4 (King, 2009),
247 the slab strength impulsing toroidal flow around it, and decoupling between
248 hot upwellings and plates. The dependence of viscosity upon temperature
249 is extreme in rocks. In the cold boundary layer, viscosity values are so high
250 that it ultimately means that viscosity is not the relevant rheological ap-
251 proximation to use. Instead, elasticity and brittle behavior would be more
252 appropriate. Viscosity drops quickly towards the base of the upper thermal
253 boundary layer. The viscosity contrast between the colder and hotter region
254 of the mantle, or within a model is a fundamental parameter to examine.
255 The magnitude is fundamental, but also its gradients are to be scrutinize
256 with a lot of care (King, 2009; Stein and Hansen, 2013).

257 Hence tackling the numerical problem to obtain stable and reliable so-
258 lutions is a challenge. To take that challenge, the first step is to develop
259 appropriate numerical methods (Yoshida and Ogawa, 2004; Choblet et al.,

260 2007; Tackley, 2008; May and Moresi, 2008; Kronbichler et al., 2012), and the
 261 second step is to choose parameterizations that can be reliably solved. In the
 262 recent literature, there are two types of temperature-dependent parameteri-
 263 zations for plate-like behavior models. The first one is based on controlling
 264 the viscosity contrast between the values of temperature being temperature
 265 at the surface T_s and temperature at the base $T_s+\Delta T$, ΔT being the tem-
 266 perature jump of reference:

$$\mu \propto \exp\left(\frac{A}{(T - T_s)/\Delta T + 1} - \frac{A}{2}\right), \quad (1)$$

267 in which T is the temperature variable and A the factor that controls the
 268 maximum viscosity contrast in the model. This type of parameterization is
 269 ideal when working with dimensionless numbers, which makes sense since
 270 reaching an Earth-like regime with these models is a big challenge in terms
 271 of computation. The viscosity dependence relies on relative temperature
 272 (being $T - T_s$). Trompert and Hansen (1998b); Tackley (1998); van Heck and
 273 Tackley (2008); Yoshida (2008); Foley and Becker (2009); Rolf and Tackley
 274 (2011); Mallard et al. (2016); Langemeyer et al. (2021) use that type of
 275 parameterization to study convection with plate-like behavior.

276 The second type is built on an experimental parameterization:

$$\mu \propto \exp\left(\frac{E_a}{RT} - \frac{E_a}{RT_0}\right), \quad (2)$$

277 in which T_0 is a reference temperature which determines the viscosity of
 278 reference in the model, and E_a the activation energy. This law explicitly uses
 279 the absolute value of temperature. This makes it less useful when exploring
 280 dimensionless numbers. However, the viscosity gradients between hot and
 281 cold are stronger than in equation (1), which is the basis of modelling plates:

282 close to rigid pieces over soft flowing mantle. Cramereri and Tackley (2014,
283 2016); Zhang and O'Neill (2016); Coltice et al. (2017a); Coltice and Shephard
284 (2018); Arnould et al. (2019); Coltice et al. (2019); Arnould et al. (2020) use
285 that type of parameterization in 3D models.

286 These two laws produce plate-like behavior when combined with a yield
287 stress formulation. The mean emergent properties (tectonic regimes, heat
288 flow, root mean square velocity) are weakly dependent on the temperature-
289 dependent formulation (Stein and Hansen, 2013). However, the tectonic
290 and convective structures are slightly different, especially when it comes to
291 trenches and subduction. The comparison of published models shows that
292 subduction with law 2 initiates from small 3D structures and develops arcuate
293 trenches that most of the time retreat (Cramereri and Tackley, 2014; Coltice
294 et al., 2019), while they are linear and stable features with law 1 (Rolf et al.,
295 2012; Mallard et al., 2016). For instance, in published work and from my
296 experience with both viscosity laws, plume induced subduction only exists
297 using law 2, highlighting the role of viscosity contrasts at the base of the
298 thermal boundary layer, around slabs and plumes. The viscosity laws are still
299 to explore and should not be neglected. As I will suggest in subsection 3.4,
300 they can drive different mechanisms for subduction initiation and plume-
301 lithosphere interaction.

302 **3. The whole is bigger than the sum of the parts parts. The whole** 303 **is smaller than the sum of the parts**

304 Solving the local equations of 3D convection with a pseudo-plastic approx-
305 imation leads, with the adequate parameters, to emergent tectonic features.

306 As for any complex system, there is no way to predict the existence of these
307 features from the local equations. Self-organization is an intrinsic property
308 of these models. I describe here some of the emergent features that could be
309 relevant for our planet today.

310 *3.1. Continental drift*

311 Continental drift in convection models has been extensively studied since
312 Gurnis (1988), but mostly without plate-like behavior. This seminal work
313 confirmed the idea that continents and convection interact leading to a self-
314 organization that is characterized by long wavelength convection and moving
315 continents. The latter drift towards downwellings (being pressure lows in the
316 boundary layer). Once it lies above it, because it is buoyant and viscous, it
317 shuts down the downwelling and another one initiates elsewhere. This view
318 guided Zhong et al. (2007) to propose that supercontinent cycles (Nance
319 et al., 2014; Mitchell et al., 2021, for reviews on supercontinent cycles), are
320 organized by an assembly phase when mantle convection is dominated by a
321 single subduction system in one hemisphere attracting all continents, and a
322 dispersion phase when two new systems of subduction dominate away from
323 the supercontinent. Convection models with viscous continental rafts pro-
324 duced supercontinent cycles with relevant time scales (Trubitsyn and Rykov,
325 1995; Lowman and Jarvis, 1999; Phillips and Bunge, 2007, among others).
326 Would the fact that supercontinent cycles could be produced without ac-
327 counting for plates (with the exception of Lowman and Jarvis, 1999) and
328 strong slabs suggest that continental drift is independent of plate tectonics?

329 Adding plate-like behavior to models with continents show it is not the
330 case. A major conclusion in the previous convection studies is that the long

331 wavelength structure of the flow is the dominating feature for continental
332 drift and supercontinent cyclicity. In convection models without plate-like
333 tectonics, long wavelength is produced by the existence of continental rafts
334 (Gurnis, 1988; Guillou and Jaupart, 1995; Phillips and Coltice, 2010), or
335 by the viscosity jump within the mantle (Bunge et al., 1996; Zhong et al.,
336 2007). However, when plate-like behavior is accounted for the long wave-
337 length, the flow is generated primarily by the existence of strong plates (van
338 Heck and Tackley, 2008; Yoshida, 2008; Rolf et al., 2012, 2018; Mallard et al.,
339 2016) (long wavelengths already exist without viscous jump and continents).
340 Therefore, the initial reasoning could stand, but the physics changes. Irregu-
341 larities in supercontinent cycles become more connected to the non-linearity
342 of the rheological response in the lithosphere (Rolf et al., 2014, 2018) than
343 the hypothetical existence of strong plumes (Phillips and Bunge, 2007). The
344 details of continent aggregation and breakup also differ. For example, in
345 models without plate-like behavior, supercontinent formation always happen
346 by extroversion, meaning that following the breakup, continent pieces collide
347 on the exterior margins. However, this is not always the case on Earth. The
348 formation of Pangea happened by introversion, continent pieces having col-
349 lided on the interior margins (Murphy and Nance, 2003). Such behavior is
350 not often observed in models with plate-like behavior, but has been observed
351 in Coltice et al. (2019). More work is needed to evaluate if introversion hap-
352 pens for specific convection parameterisations. Also, in this numerical model,
353 continental collision happens synchronously with rifting in distant regions,
354 which is typical of the Earth (Himalayas growing synchronously with the
355 opening of the African rift).

356 *3.2. Seafloor spreading*

357 Within "ocean" basins (oceans do not explicitly exist in the models), the
358 numerical solutions of convection with yielding generate seafloor spreading
359 characteristics. The basins are fragmented into pieces that behave like plates,
360 90 % of active strain being restricted to 1-10 % of the surface depending on
361 model parameters. Fragmentation is represented by divergent boundaries,
362 sometimes with vorticity component, which have similarities with mid-ocean
363 ridges. They are long linear features with no clear connection with the un-
364 derlying velocity and temperature fields. They appear as passive structures
365 of fragmentation emerging to accommodate forces imposed by slabs, plumes
366 and viscous stresses between the boundary layer and the interior. The con-
367 nections between ridge systems correspond to triple junctions (Figure 1C). In
368 convection models with a viscous law 2, quadruple junctions also exist, but
369 only temporarily and related to plume-ridge configurations (Arnould et al.,
370 2020). Ridges are a fundamental example of the assertion "the whole is
371 smaller than the sum of the parts" because the fragmentation pattern in such
372 models is forced fully by the collective interactions over the whole surface.
373 It forms a network that adapts constantly to global changes in mechanical
374 stresses. Only the thermal dimension involves some memory, meaning some
375 sort of independence of a ridge from its surroundings. A ridge is a hotter
376 area, a thinner boundary layer. Hence, a locally weaker area. This acts as a
377 restoring force on a system, contributing to keep the ridge where it has been.

378 The fragmentation pattern varies with the yield stress value. The plate's
379 area distribution is a convenient way to explore the fragmentation pattern.
380 For the Earth, the distribution can be divided into several large plates of a

381 similar size (being of the order of the Pacific plate size for instance), and
382 a distribution of smaller plates that follow a power law (Bird, 2003). A
383 discussion about the possibility of using one single density function or 3
384 different ones (one for small plates, one for intermediate and one for large)
385 could be relevant (Sornette and Pisarenko, 2003; Vallianatos and Sammonds,
386 2010). Whatever the mathematical interpretation, 3D spherical convection
387 models with yielding generate such distributions innately (Mallard et al.,
388 2016). It is a typical emergent pattern of convection with such rheology. A
389 lower yield stress shifts the distribution to smaller and hence more numerous
390 plates, while a larger yield stress pushes the distribution towards a smaller
391 number of larger plates. Fragmentation is more localized in areas where
392 subduction is active, which represents the main driver of fragmentation.

393 Plumes influence the localization of ridges by weakening thermally the
394 boundary layer in their vicinity. Because they are several 100 K hotter than
395 their surroundings, plumes reduce the thickness of the boundary layer they
396 are in contact with, and over areas extending several 100 km. Therefore, the
397 impact of a plume close to a ridge can lead to a ridge jump (Fig. 2A), to
398 the localization of the fragmentation of a whole ocean basin (Fig.2B), or to
399 nothing at all. Sometimes, it is even the hot wake of a plume, which is like
400 a thermal scar in the boundary layer, that focuses the fragmentation of a
401 plate (Fig.2C). This type of tectonic event depends on too many interactions
402 (local and long-distance) to be predicted, unfortunately.

403 Textbooks often represents seafloor spreading with the typical symmet-
404 ric magnetic stripe patterns. They represent the distribution of ages on the
405 seafloor, from which so many properties are correlated: bathymetry, heat

406 flow, hydrothermal activity among others. The seafloor age-area distribu-
407 tion on Earth today is called triangular: the youngest seafloor dominate and
408 a very limited amount of >100 My seafloor exists. One can interpret that the
409 age of subduction is equally distributed. This is surprising for a convective
410 system for which a critical age determines, in principle, the onset of cold
411 instabilities. I will come back to it in subsection 3.4. Indeed, Labrosse and
412 Jaupart (2007) showed that simple convection models generate distributions
413 of seafloor ages that are rectangular: the area of a given age is constant
414 up to a critical age beyond which the area is 0. In a 3D spherical geome-
415 try, convection with yielding produce a variety of age-area distributions, and
416 when continents are added, the triangular one dominates (Coltice et al., 2012,
417 2013). Continental margins impose a geometrical forcing. They also generate
418 stronger time-dependence of convection, which translates into strong fluctua-
419 tions of the production of new seafloor. Both lead to triangular-like age-area
420 distributions. The interpretation is that the present-day age-area distribution
421 of the Earth represents "forced" subduction below continents and changes
422 in seafloor production. Hence, depending on the continental configuration
423 (Coltice et al., 2014) and timing, the age-area distribution evolves between
424 rectangular and exponential decay, going through triangular shapes.

425 *3.3. Transform zones*

426 Transform faults were a key feature to demonstrate the occurrence of
427 seafloor spreading (Wilson, 1965) and build a plate tectonic theory (Mor-
428 gan, 1968). Their existence is not for granted. Indeed, a plate boundary
429 showing pure rotational behavior is a very special configuration. Boundaries
430 combining divergence with rotation could be the rule. It actually is in most

431 models of convection with yielding, in which transform shear zones (faults
432 do not exist in viscous models) are replaced by ridges with strong rotational
433 components. Overall, the toroidal component of the velocity, which corre-
434 sponds to vorticity and potentially transform motion, is comparable between
435 convection models with plate-like behavior and the Earth (van Heck and
436 Tackley, 2008). But the tectonic features expressing it are somewhat dif-
437 ferent. Therefore two major questions arise: what physical parameters are
438 needed to allow pure transform motion to emerge ; what does the transform
439 offset expresses in terms of regional and global dynamical evolution?

440 Transform motion has been observed in 3D spherical models of convection
441 with yielding (Coltice et al., 2017b, 2019), but as a feature which is not as
442 dominant as on Earth today, except at specific times during a calculation.
443 Exploring the parameter space of these models is extremely difficult: com-
444 putations require weeks to months, the non-linearities imply that changing
445 one parameter redefines the whole flow regime. The number of parameters
446 to vary is larger than 10. However, global and regional lithospheric models
447 already define the scope of the generation of transform motion. Langemeyer
448 et al. (2021) reproduced large and distinct transform offsets in a series of
449 3D spherical models of convection with yielding. When the yield stress is
450 constant with depth and close to the yield stress that sets the transition to
451 the stagnant lid regime, transform motion emerges within elongated ridge
452 systems. In these models, the offsets develops with time and grow, starting
453 from a corrugated ridge system. The models that produce these transform
454 motions only develop a small number of very large plates, which differs from
455 the Earth plate size distribution. This study shows that transform motion

456 emerges in global models when plates are stiff (the yield stress value is large).
457 When they are not, transform motion exists and develops, but quickly de-
458 rive into transpressive-transpressive motion. Weakening is an issue, because
459 if a slight divergence onsets in the weak areas, it grows at the expense of
460 transforms, which seem to require some degree of strength. Fig. 3 shows
461 the transient evolution of a transform offset I obtained in a calculation I
462 conducted for parameter exploration, which has similarities with those of
463 Langemeyer et al. (2021), since plates are large and not many. The offsets
464 develop through time and are not set from the start of the model.

465 Regional models of the lithosphere that are pulled from the sides can
466 provide complementary views on transform systems. Gerya (2010) proposes
467 that strain weakening is a major ingredient to nucleate and develop transform
468 segments, which acts as instabilities in asymmetric ridges. 3D lithospheric
469 models including grain size reduction initiate transforms when reaching ex-
470 treme weakening (Schierjott et al., 2020). However, the results of global
471 models and regional models seem to conflict at first sight. The missing con-
472 nections come from: (1) weakening mechanisms have not been explored in
473 3D spherical models of convection yet, but attention is brought up and work
474 is in progress (Fuchs and Becker, 2019; Rolf and Arnould, 2021) (modeling
475 weakening is a big issue in terms of numerical methods, especially at a global
476 scale (Duretz et al., 2020)) and (2) regional models show transform onset
477 over $<3-4$ Myrs while global models have shown that the difficulty is to keep
478 these structures stable for long periods if they are weak.

479 *3.4. Subduction*

480 Subduction appears as the essence of tectonics on Earth. When looking
481 for the onset of plate tectonics, scientists search for clues of subduction within
482 the oldest crustal minerals (Turner et al., 2020, for instance). In classical
483 convection models, without strongly varying viscosity, the surface boundary
484 layer is unstable and drips form and sink into the interior. Making the
485 connection with subduction is another story since other characteristics are
486 required. Among them, subduction on Earth is one-sided, with sheet-like
487 structures dipping with a variety of angles, migrating and deforming the
488 upper plate (Schellart, 2007; Husson, 2012).

489 *3.4.1. Downwellings or subduction?*

490 In 3D spherical models of convection with yielding, convergent boundaries
491 are linear features, mostly arcuate depending on their width, and down-
492 wellings are sheet-like structures. Therefore, they have a lot in common with
493 subduction zones on Earth. However, most models display two-sided sub-
494 duction, with both sides of the boundary layer sinking into the interior. As
495 a result, the cold downwellings are mostly vertical until they reach a rheo-
496 logical/phase boundary where their geometry can change (Yanagisawa et al.,
497 2010). One-sided subduction is the extreme configuration of asymmetric
498 subduction (Gerya et al., 2008). It is observed in convection models with
499 yielding when a thin weak layer is continuously formed at the surface of the
500 models (Crameri et al., 2012; Coltice et al., 2017b, 2019) and reinforced by
501 the incorporation of a free-surface (Crameri et al., 2012; Crameri and Tack-
502 ley, 2015). It can also be observed in models with strong viscosity gradients,
503 which typically employ the viscosity law 2 (Coltice et al., 2017a; Arnould

504 et al., 2018; Rodriguez et al., 2021; Coltice et al., 2019).

505 For the models with asymmetric to one-sided subduction, the more nar-
506 row the trench, the more arcuate and the faster their retreat, as predicted
507 by regional models (Stegman et al., 2006). However, no systematic study
508 has been made yet. Subduction is the singular dynamic feature for convec-
509 tion with plate-like tectonics. In a global model, no subducting slab remains
510 isolated from the information in the whole system. It is typically a feature
511 influenced by the systemic constraints, the collective interactions. They cer-
512 tainly inhibit some of the behaviors that can be studied in models and
513 experiments involving the sinking of one or two isolated slabs into a box. For
514 instance, the type of regime diagram found in state-of-the-art slab models
515 (Garel et al., 2014) could be explored in the a large-scale context. For now,
516 looking for situations in which the whole is smaller than the sum of the parts
517 has not yet been done.

518 Another point to raise is the termination of trenches. While geological
519 and seismological observation could favor the tearing of slabs on their sides
520 (Jolivet et al., 2015), convection models display slab sides that are rolling-
521 up (Fig. 4: Line 2). That difference may suggest again the importance of
522 weakening mechanisms to complete the picture of an Earth-like tectonics in
523 global models.

524 Slabs produced by convection models are very stiff in the upper mantle,
525 generating strong toroidal motion around them, and buckle and fold within
526 a more viscous lower mantle where they reside for long times and sink slowly.
527 Therefore, the shape of slabs contrasts between upper and lower mantle struc-
528 ture, being well defined and in continuity with the surface in the former, and

529 diffuse and folded around in the latter. Part of it comes from the viscosity
530 jump, as in Strak and Schellart (2021) for instance, part of it comes from the
531 existence of other slabs, plumes and dynamic features.

532 *3.4.2. Onset of subduction*

533 On Earth, most subduction zones initiated in the past 65 My (Gurnis
534 et al., 2004) within oceans, and are fully active at present. The detection
535 of incipient subduction is a very difficult task because it requires it exists
536 today and we have enough observations to document it. Some examples are
537 proposed (Duarte et al., 2021), but research is ongoing for definitive data
538 to validate such hypothesis. In the literature, the variety of subduction ini-
539 tiation mechanisms comes from combinations of observations and regional
540 modeling in which the initial configuration is imposed (Stern and Gerya,
541 2018, for a review). Again, the collective interactions are fundamental here
542 to inhibit ranges of possibilities that can be found in models where onset
543 of subduction is studied in isolation. Indeed, reviewing subduction initia-
544 tion points to the impact of external contributors to the process on Earth
545 (Lallemand and Arcay, 2021).

546 Little attention has been addressed to subduction initiation in convec-
547 tion models with yielding, except Cramer and Tackley (2016) and Ulvrova
548 et al. (2019), the latter being in 2D. Subduction initiation is a result of self-
549 organization. There is no weak zone placed initially to favor it in these mod-
550 els. Therefore, they appear where a dense instability can develop and where
551 the local stress exceeds the yield stress. Cramer and Tackley (2016) ob-
552 served that the impact of hot plumes favor subduction initiation by thinning
553 locally the boundary layer (weakening) and generating a lateral buoyancy

554 contrast. Modeling of a free surface helps because the lateral pressure gra-
555 dients caused by topography changes can be accounted in a better way than
556 free slip boundary conditions. The existence of continental margins can focus
557 a fraction of subduction initiation (Rolf and Tackley, 2011; Ulvrova et al.,
558 2019). This can happen because continents are lighter, imposing a buoyancy
559 gradient at the ocean-continent boundary (Nikolaeva et al., 2010; Lévy and
560 Jaupart, 2012; Rey et al., 2014).

561 Digging into 3D spherical models of convection with plate-like behavior
562 shows that subduction initiation and growth differs depending on the type
563 of temperature dependence of the viscosity discussed before. The models
564 with law 1 initiate subduction by boundary layer yielding that can operate
565 at the scale of an ocean: subduction can initiate has a shear zone extending
566 over thousands of kilometers (Fig. 4: Line 1). It never happens that way in
567 models with law 2 (both Crameri and Tackley, 2016 and Ulvrová et al. 2019
568 uses that law). Subduction initiation starts on a very confined region and
569 develops by extending the trench on both sides, in a growing arc fashion. In
570 Coltice et al. (2019), subduction initiates where a local buoyancy contrast
571 can grow, like a boundary between two basins of differing ages (Fig. 4: Line
572 2), or under the effect of a plume impinging on relatively young seafloor
573 (Rodriguez et al., 2021).

The onset of subduction in convection models with yielding brings to
the table the balance between stresses and buoyancy. The growth of a cold
instability from the boundary layer happens for a critical nondimensional
number in convection models without yielding, being a local critical Rayleigh

number Ra_{cr} :

$$Ra_{cr} = \frac{\rho_0 g \alpha \Delta T \delta_{cr}^3}{\kappa \mu},$$

574 ρ_0 being the reference density, g the gravity acceleration, α the thermal ex-
575 pansion, ΔT the temperature contrast over the layer, δ_{cr} the critical thickness
576 of the layer, κ the thermal diffusivity and μ the viscosity of the layer. There-
577 fore, the fundamental value is the thickness of the layer, which is mostly
578 determined by its thermal age. Without yielding, subduction happens when
579 the boundary layer is locally old enough, thick enough.

When yielding comes into the story, two critical nondimensional numbers come into play, because the growth of the instability at the boundary layer scale (not small scale convection) imposes that the local stress exceeds the yield stress. Therefore, the first nondimensional number is the ratio between the yield stress and buoyancy:

$$S_b = \frac{\sigma_y}{\rho_0 g \alpha \Delta T \delta},$$

δ being here the local thickness of the layer. And the second nondimensional number is the ratio between the yield stress and viscous stresses:

$$S_t = \frac{\sigma_y \delta}{\mu \kappa},$$

580 which corresponds to the nondimensional yield stress often expressed in pub-
581 lications. The ratio of S_t/S_b is by definition the Rayleigh number. Therefore,
582 to characterize the system, one needs either the Rayleigh number and one
583 among S_b and S_t , or just S_b and S_t .

584 Because of a new condition on the stress to reach, the local Ra_{cr} can be
585 exceeded. The lithosphere is unstable but strong enough not to yield and

586 sink. Therefore small scale convection at the base of the boundary layer initi-
587 ates, buffering the thickness of the boundary layer for seafloor ages exceeding
588 the critical age. If not, the boundary layer can become even more thick and
589 hence old modeled seafloor becomes extremely stiff because of the tempera-
590 ture dependence of the viscosity. As a consequence, initiating subduction on
591 old seafloor is difficult because the local S_t is very large. Solomatov (2004)
592 proposed that small scale convection could initiate subduction in convection
593 models with yielding, bridging a theory and 2D numerical models using a
594 Frank-Kamenetskii law for the temperature-dependence of the viscosity. 3D
595 models using law 2 do not produce subduction initiation from small-scale
596 convection. A reason is certainly related to the fact that small-scale insta-
597 bilities have temperature contrasts proportional to the inverse of the local
598 viscosity contrast ($\Delta\mu/\mu$) (Davaille and Jaupart, 1994). In the models using
599 law 2, the viscosity gradients are so large at the base of the boundary layer
600 that the buoyancy of small scale instabilities is not strong enough to gen-
601 erate yielding. Therefore, stronger buoyancy effect are required as plumes,
602 continent-ocean boundaries or boundaries between basins of contrasting ages
603 (transform motion, or a scarred plate boundary).

604 The published models of convection with yielding in 3D are not yet able
605 to take into account some relevant physics for subduction initiation on Earth:
606 weakening, viscous anisotropy or melting among others. For now, they can
607 only account for the feedbacks between instantaneous stresses and buoyancy.

608 4. Outlook

609 20 years of modeling mantle convection with plate-like behavior have
610 opened the gates between lithospheric modeling and convection modeling.
611 The tools are now similar for many research groups: Stokes solvers with
612 variable material properties and rheologies that are projected in the viscous
613 space, pseudo-plasticity being one of them. Parallel supercomputers pro-
614 vide the computational power to determine the numerical solutions in 3D.
615 These models, as any model, are imperfect. Therefore, it is useful to carry
616 experiments and compute other kinds of numerical models (Morra et al.,
617 2007; Bonnardot et al., 2008; Combes et al., 2012; Gerardi et al., 2019, for
618 instance) to identify the areas of progress and increase awareness on their
619 use. The convergence of modeling approaches between communities that
620 tackle tectonics from 2 sides is very promising. Communities already gather
621 in common science meetings. However, it will still take time for both ap-
622 proaches to reach their prime: mantle convection models are requested to
623 account for the essential rheological behaviors identified by microphysics and
624 lithospheric modeling (Burov and Watts, 2006; Bercovici and Ricard, 2014,
625 among others), and lithosphere models are requested to initiate the regional
626 models with temperature and chemical fields consistent with self-organized
627 convection.

628 Yet, convection models with yielding can be used to study the underlying
629 mechanisms of plate tectonics in a way geological observations come into
630 play (Mallard et al., 2016; Rolf and Pesonen, 2018; Rolf et al., 2018; Ulvrova
631 et al., 2019; Coltice et al., 2019; Rodriguez et al., 2021; Langemeyer et al.,
632 2021). How plate boundaries initiate, evolve and cease, how the tectonic

633 jigsaw reorganizes, how continents break and mountains build in details, and
634 so many other questions. These models provide a systemic approach to core
635 questions that have remained unanswered since the birth of plate tectonic
636 theory.

637 I focus then on 4 major points to deal with:

638 • **Rheology vs. initial conditions.** The answer to a tectonic problem
639 is often searched within rheological parameters, especially in regional
640 models. It is natural because these models specify thermal initial con-
641 ditions built on assumptions. However, the consistency between rheol-
642 ogy and initial conditions is harshly and rarely met. Ideally, the initial
643 set up with the physical parameters used should correspond to a solu-
644 tion, obtained from solving the equations, being a self-organized state.
645 For instance, the initial conditions (boundary layer temperature dis-
646 tribution, temperature at depth) for a subduction model is supposed
647 to be built on a convection model with the chosen parameterization.
648 The forces applied on a lithospheric model are supposed to be consis-
649 tent with what a 3D model that self-organizes would produce. Tak-
650 ing Earth-like conditions is not consistent with a choice of parameters
651 which, by essence, cannot match exactly and comprehensively with the
652 Earth (material properties but also physics). As a consequence, varying
653 the rheology of the model tells that what is observed comes from feed-
654 backs between the initial conditions and the parameterization. These
655 are non-linear feedbacks. The question is then: does the result express
656 consequences of uncertainties or relevant mechanisms? And in a com-
657 plex system, it is very hard to separate both possibilities. Additionally,

658 the whole is smaller than the parts, which leads to the question: does
659 the observed behavior express when interactions exist?

660 • **Memory.** Localization with memory is a step to make for 3D spherical
661 convection models. For now, yielding in these models mean instantane-
662 ous yielding and healing. Although, there is already a strong effect
663 of thermal and chemical memory. Indeed, 100 My is a typical time
664 scale for a thermal anomaly to die out, and even more for a chemical
665 anomaly (a continental block for instance). A plume, a ridge, or a
666 subduction imposes a thermal anomaly which in turn produces a rhe-
667 ological anomaly that persists for a long time. Wether memory driven
668 by grain size reduction and growth dominates over thermal/chemical
669 memory is a relevant question and depends on the time and space scales
670 of the process studied. Coltice and Shephard (2018) and Bello et al.
671 (2015) highlighted this issue: differences in initial conditions produce
672 quick differences in terms of tectonics and thermal state, that are larger
673 or comparable to differences in rheological parameters. Also, buoyancy
674 is the driving force of the system, rheology is a filter to its expression.
675 In a system that self-organizes, thermal anomalies develop on their own
676 depending on the full organization of the flow, not necessarily where
677 a weak zone exists (Fuchs and Becker, 2019). A classical property in
678 a complex system is that although emergence implies that the whole
679 is more than the sum of the pieces, the sum of the pieces is also more
680 than the whole. Indeed, the interactions between the pieces are con-
681 straining the global structures. Therefore, while memory (being a local
682 property) in an isolated regional model is dominant and can express,

683 the interactions between dynamic regions of a global model (producing
684 self-organization) may inhibit that effect in the final structuration.

685 • **Automatic analysis of tectonics.** Analyzing the tectonics of 3D
686 spherical models with fine resolution that generate billions of years of
687 geological histories is a challenge. Tectonics is in the details of the
688 structure. On Earth, plates are a first way to map global structures,
689 before zooming into them and their boundaries. It is primordial that
690 computers can identify automatically discrete structures in 3D, or at
691 least give some help. Computer vision and machine learning strategies
692 are promising and will give a boost to studying tectonics in 3D global
693 models (Bremer et al., 2010; Mallard et al., 2017; Duclaux et al., 2020;
694 Wrona and Brune, 2021; Tierny et al., 2017).

695 • **Inverse methods.** 3D spherical models of convection with plate-like
696 behavior can already be predictive enough to be used for inverse mod-
697 eling (Coltice and Shephard, 2018; Coltice et al., 2019), depending of
698 course on what the inverse problem is. Using these models to inverse
699 geological and geophysical observations, without making assumption on
700 the tectonic history, would open a way to make reconstructions with-
701 out the plate tectonic assumptions. Data assimilation tests on sim-
702 ple models are promising (Bocher et al., 2016; Li et al., 2017; Bocher
703 et al., 2018), but there is still a step to make to push these inverse
704 models towards 3D sphericity and time-dependence. The leap forward
705 that oceanography has made for circulation reconstructions could be
706 in reach in the next years, with the promise to bring together tectonic

707 and deep flow reconstructions.

708 **5. Final thoughts**

709 Models of global convection with yielding generate emergent tectonics at
710 the surface within a range of parameters. The patterns observed within the
711 top cold boundary layer rarely simply mirror the flow structure in the weaker
712 hot mantle beneath it. The models which have been developed in the past
713 20 years are computationally demanding in 3D, but provide unique ways to
714 investigate how deformation at the surface of planets evolve, how lithospheric
715 shear zones of any kind initiate, develop, cease, interact with each other and
716 deep flow. For now, these models have barely been employed to study the
717 physics at play for emerging tectonic features like ridges, transform, subduc-
718 tion, collision, continent cycles... But the bridge with lithospheric models
719 is close to be built, pushing the modeling towards a synergistic approach.
720 In convection models with plate-like behavior, the systemic complexity pre-
721 cludes simple causal links, such as "subduction triggers a rift", "this plume
722 breaks this continent", "plate velocity accelerates thanks to this subduction
723 initiation", because a full set of interactions at a variety of scales generate a
724 single self-organized state, that can substantially alter if a small perturbation
725 is added. It is easy to end up in the chicken or egg question, and difficult to
726 work looking for simplicity or Occam's razor principles. An illustration of the
727 causality blindness is given by Lorenz (1972) when discussing the butterfly
728 effect in complex chaotic systems:

729 "1. If a single flap of a butterfly's wing can be instrumental in generating
730 a tornado, so all the previous and subsequent flaps of its wings, as can the

731 flaps of the wings of the millions of other butterflies, not to mention the
732 activities of innumerable more powerful creatures, including our own species.

733 2. If a flap of a butterfly’s wing can be instrumental in generating a
734 tornado, it can equally well be instrumental in preventing a tornado.”

735 Change tornado to subduction, and the flap of a butterfly’s wing to an
736 earthquake and it would fit the convection-tectonics system. Even though
737 looking for causal links between slabs, plumes, plates and other dynamic
738 features seems vain, convection models with yielding provide a full description
739 of the fields involved (temperature, velocity, rheology), helping to investigate
740 systematics, and how tectonics rely on the physical parameters being the
741 underlying program of the emerging diversity. The physics of these models
742 is relevant to the Earth, but also to the planets of our solar system, allowing
743 to investigate the intrinsic properties leading to their unique identity. The
744 hologram metaphor suggests tectonic structures contain information about
745 the whole history of planetary interiors. Geodynamic models are a source of
746 light to reveal it.

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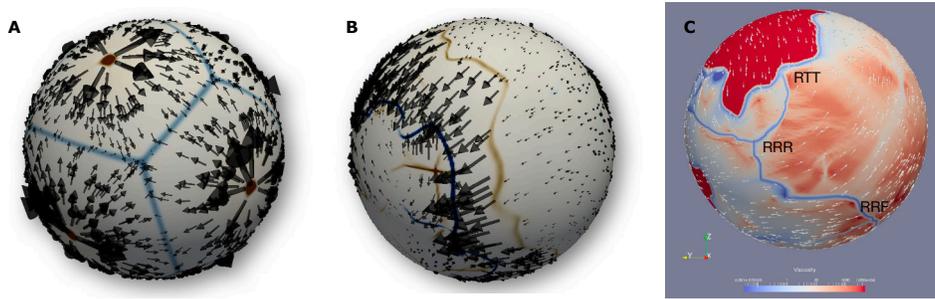


Figure 1: A: Surface velocities, diverging area in blue and convergent area in red in a convection model with viscosity varying only radially (model "Radial viscosity - basal heating" in Coltice et al., 2019). B: Surface velocities, diverging area in blue and convergent area in red in a convection model with viscosity varying only radially (model "Reference model" in Coltice et al., 2019). C: 3D snapshot of the viscosity at 10km depth on the hemisphere of a convection model, showing triple junctions (RRR=ridge-ridge-ridge, RTT=ridge-trench-trench and RRF=ridge-ridge-transform). Arrows show the direction of the velocity at the surface. The convection model has the same parameterization as that of Coltice et al. (2017), but with continental rafts initially positioned as Pangea. Thermal initial conditions are obtained running the code fixing the tracers until a statistical state is reached.

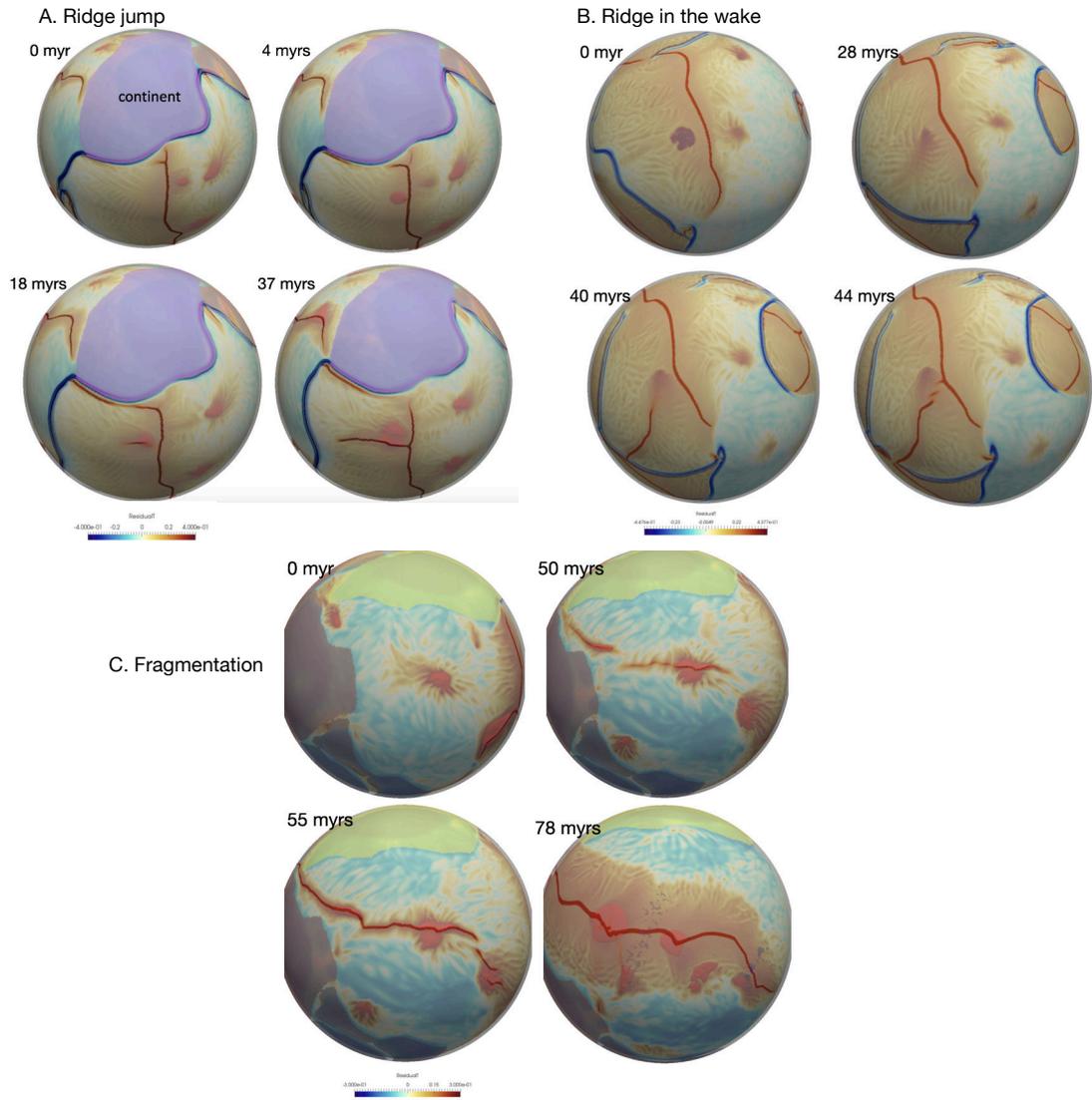


Figure 2: 3D snapshot of plate fragmentation styles involving plumes. Convection model corresponds to running the model of Coltice and Shephard (2018) without imposing plate motions. Color is the non-dimensional residual temperature (temperature difference with the lateral mean) at depth 50km. Red isotherm of non-dimensional temperature 0.9 highlight plume position (except in B). Continental rafts are shaded areas (except in B). A. corresponds to a ridge switching direction under the influence of a plume. B. corresponds to fragmentation occurring within the wake of the plume. C. corresponds to the fragmentation of a large ocean basin.

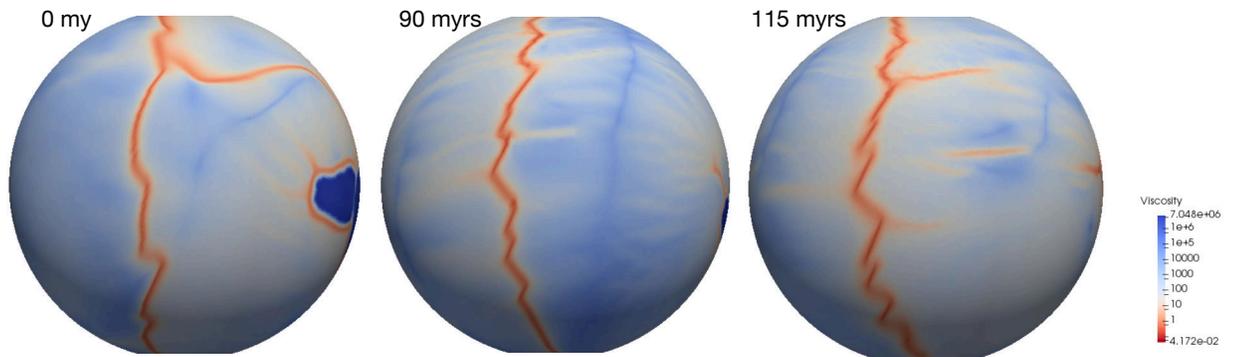


Figure 3: 3D snapshots of the development of transform-like shear zones in a convection model built on Earth's continent model in Coltice et al. (2017) but with weaker continental lithosphere on their edges. Weaker means viscosity and yield stress are both 10 times lower than that of the mantle. These edges get easily recycled and stretched within the convective system and can reach the surface again.

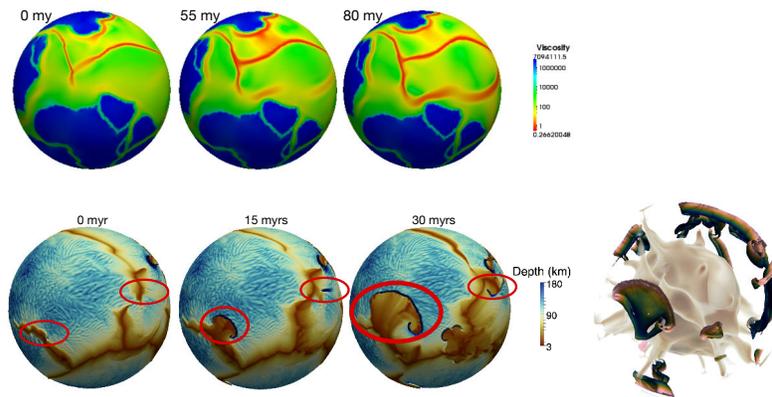


Figure 4: Line 1: 3D snapshots of model PLC1 in Bello et al. (2014) showing the initiation of a linear subduction zone close to the equator. Line 2: 3D snapshots of the fine resolution model in Coltice et al. (2019), showing initiation of subduction. In the southern hemisphere, subduction initiates around a plume. In the East side of the northern hemisphere, a subduction initiates close to a ridge at a smooth boundary between to basin of slightly different ages. On the left: Corresponding interior structure of the 30 Myr snapshot. Hot isotherm shows the presence of plumes. Cold isotherm is colored by depth and represents the sinking slabs in the upper mantle of the model. The sides of the slabs in such model are rolling up.