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1                   **Kinematic and rheological controls on rift-related fault evolution**

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11                   **Key Points:**

- 12                   • Numerical models are qualitatively and quantitatively comparable to observational fault  
13                   networks
- 14                   • The scaling distribution of fault strain is not entirely power-law as there is a characteristic  
15                   upper-bound limit
- 16                   • Off-fault (i.e. non-extracted) deformation is significant in all models, accommodating  
17                   from 25-45% of extension depending on rift parameters.
- 18

19 **Abstract**

20 Continental extension is primarily accommodated by the evolution of normal fault  
21 networks. Rifts are shaped by complex tectonic processes and it has historically been difficult to  
22 determine the key rift controls using only observations from natural rifts. Here, we use 3D  
23 thermo-mechanical, high-resolution (<650 m) forward models of continental extension to  
24 investigate how fault network patterns vary as a function of key rift parameters, including  
25 extension rate, the magnitude of strain weakening, and the distribution and magnitude of initial  
26 crustal damage. We quantitatively compare modelled fault networks with observations of fault  
27 patterns in natural rift, finding key similarities in their along-strike variability and scaling  
28 distributions. We show that fault-accommodated strain summed across the entire 160 x 160 km  
29 study area increases linearly with time. We find that large faults do not abide by power-law  
30 scaling as they are limited by an upper finite characteristic,  $\omega_0$ . Fault weakening, and the spatial  
31 distribution of initial plastic strain blocks, exert a key control on fault characteristics. We show  
32 that off-fault (i.e. non-fault extracted) deformation accounts for 30-70% of the total extensional  
33 strain, depending on the rift parameters. As fault population statistics produce distinct  
34 characteristics for our investigated rift parameters, further numerical and observational data may  
35 enable the future reconstruction of key rifting parameters through observational data alone.

36 **1 Introduction**

37 Rifts are globally widespread, yet display a wide variety of structural styles (Fig. 1) as a  
38 result of crustal rheology (Buck, 1991; Ebinger et al., 1999), rift obliquity (Corti, 2008), the  
39 presence of pervasive lithological and mechanical heterogeneities (Fletcher and Hallet, 1983;  
40 Morley 1999), extensional velocity (Buck, 1991; Huisman and Beaumont, 2003), thermo-  
41 isostasy and loading (Kooi et al., 1992), crustal melting and magmatism (Rey et al., 2009; Corti  
42 et al., 2003), and syn-extension erosion and sedimentation (Burov 1997). While there is a global  
43 understanding of what parameters influence rift evolution and the final structure of rifted  
44 margins, the interaction and feedback of such intrinsically complex processes is still poorly  
45 constrained (Peron-Pinvidic et al., 2019). As rifts are characterised by the structural development  
46 of normal fault populations, the underlying rift controls may potentially be coded into the  
47 kinematics and geometry of rift-related normal fault networks.

48 Normal faults produce extensive networks of discontinuities at the surface and  
49 subsurface, at length scales ranging from millimetres to hundreds of kilometres (e.g. Willmese  
50 1996; Schultz and Fossen, 2008). With increasing brittle strain, faults nucleate within the crust  
51 and lengthen and gain displacement to eventually form a through-going fault system (Cowie et  
52 al., 2000). Fault systems enable the trapping of carbon dioxide (Kaldi et al., 2013) and  
53 hydrocarbons (Sorkhabi and Tsuji, 2005). The discontinuities within complex normal fault  
54 populations are particularly favourable for geothermal energy production (Faulds et al., 2015;  
55 Coolbaugh et al., 2006). Our understanding of how fault populations are characterised and how  
56 they evolve through time is thus crucial to the identification of potential energy storage and  
57 waste sites.

58 Faults have been recognised to exhibit scaling properties that describe the size  
59 distributions of fault length, displacement and the spatial distribution of the fault network (Cowie  
60 et al., 1995, Fig. 2). These properties provide insights into the underlying physical mechanism of  
61 fault nucleation and growth (Cowie 1998). Field-data of faults commonly observe a power-law  
62 distribution, i.e. they are scale invariant (Kakimi 1980; Scholz and Cowie, 1990; Walsh et al.,

63 1991; Jackson and Sanderson, 1992; Marrett and Allmendinger, 1992; Villedien et al., 1995),  
64 similar to the Gutenberg-Richter relationship observed for earthquakes (Gutenberg and Richter,  
65 1956; Kanamori and Anderson, 1975). However, numerical (Ackermann et al., 2001), physical  
66 (Spyropoulos et al., 1999) and observational (Cowie et al., 1993; Gupta and Scholz, 2000; Soliva  
67 and Schultz, 2008) studies of fault length distributions suggest that faults may scale  
68 exponentially in higher strain settings (Fig. 2). This has implications for the forecasting of the  
69 size, and density of sub-resolution faults, which typically uses a power-law distribution for  
70 extrapolation (e.g. Marrett and Allmendinger, 1991; Torabi et al., 2011). Sub-seismic faulting  
71 has a profound effect on (CO<sub>2</sub> or hydrocarbon) reservoir performance, as it may enhance or  
72 reduce reservoir permeability and overall flow properties (e.g. Antonellini and Aydin, 1994;  
73 Damsleth et al., 1998; Ferrill et al., 2009).

74 It has been difficult to deduce the true scaling nature of fault networks, as data are usually  
75 obtained from mapping which contains an insufficient range of fault numbers (Cladouhos and  
76 Marrett, 1996; Scholz 1997). Often fault observations span only 1 order of magnitude in terms of  
77 fault scale (Bonnet et al., 2001) and the limited spatial extent incurs sampling biases (Pickering  
78 et al., 1995). The finite size of datasets leads to underestimation of small faults due to limited  
79 resolution (truncation) and undersampling of the largest faults (censoring), which overall alter  
80 the appearance of the distribution (Fig. 2) and may lead to erroneous conclusions of the style and  
81 factors controlling fault patterns. Previous statistical strain analyses are often presented in 1D  
82 (through borehole images and well logs) or 2D (fault lengths in outcrop or summing heaves  
83 across cross sectional transects; Walsh et al., 1991; Marrett and Allmendinger, 1991; Cowie et  
84 al., 1995; Gupta and Scholz, 2000; Torabi et al., 2011). It remains a challenge to extrapolate  
85 results obtained from 1D and 2D datasets to 3D systems (Bonnet et al., 2001).

86 In addition to the lack of spatially extensive datasets in active and ancient rifts, there are  
87 even fewer temporally constrained datasets. Earthquakes only provide a short snapshot in time  
88 relative to the long geological timescales in which they accumulate. Few faults are associated  
89 with age-constrained growth strata that record fault activity over geological timescales (e.g.  
90 Meyer et al., 2002). The lack of kinematic constraint has led to different models of how  
91 individual faults grow (e.g. Walsh et al., 2002; Rotevatn et al., 2019; Pan et al., 2022a).  
92 Furthermore, it is unclear how other forms of deformation (i.e. diffuse, ductile or aseismic slip)  
93 are accommodated during extension. Limitations in spatial and temporal resolution from  
94 observational data means there is little understanding of how fault networks evolve  
95 quantitatively, and even less understanding of how underlying rift dynamics may govern the  
96 geometry and kinematics of natural rift fault networks.

97 Here, we utilise high resolution (<650 m), 3D thermo-mechanical forward models of  
98 lithospheric deformation to investigate the effect of well known, first-order rift controls, such as  
99 extension rate and rheology, on the structural expression of normal fault network evolution. The  
100 along-strike variability revealed by the 3D models enable key fault attributes, such as the length,  
101 active and cumulative strain, and strike, to be extracted across a range of spatial scales  
102 throughout their evolution. We characterise the scaling distribution of fault accommodated strain  
103 through time, and investigate how different model parameters affect the overall geometry of the  
104 fault network in 3D. We compare differences between measured crustal strain in the models with  
105 the summation of strain accommodated by normal faults, resulting in off-deformation percentage  
106 estimates for each model. Finally, we compare the modelled fault statistics with those observable

107 in the Earth's crust, allowing us to better understand and potentially infer first-order controls on  
 108 rift dynamics, based on observed fault patterns.

109

## 110 **2 Methods**

### 111 2.1. Model description and setup

112 The open-source dynamics code ASPECT (Kronbichler et al., 2012; Heister et al., 2017)  
 113 is used to model 3D continental extension. The thermomechanical model solves the  
 114 incompressible Boussinesq approximation of momentum, mass and energy equations, combined  
 115 with advection-diffusion equations. For a full description of the governing equations, see  
 116 Naliboff et al. (2020) and Pan et al. (2022b). The governing equations are solved on a 3D  
 117 gridded domain spanning 500 x 500 x 100 km (X, Y and Z respectively). The grids are 5 km  
 118 resolution at the sides and base of the model, and we use an adaptive mesh refinement to  
 119 successively reduce this down to 625 m over a 180 x 180 x 20 km region in the centre of the  
 120 model. Velocities are prescribed to the left and right of the model. Inflow at the model base  
 121 balances the outflow, and the top of the model is a free surface (Rose et al., 2017).

122 The model contains three compositional layers: the upper crust (0-20 km), the lower crust  
 123 (20-40 km), and the mantle lithosphere (40-100 km). Viscous flow laws for dislocation creep are  
 124 wet quartzite (Gleason and Tullis, 1995), wet anorthite (Rybacki et al., 2006) and dry olivine  
 125 (Hirth and Kohlstedt, 2003) containing background densities of 2700, 2800 and 3300 kg m<sup>-3</sup>. The  
 126 temperature distribution follows a characteristic conductive geotherm for the continental  
 127 lithosphere (Chapman, 1986) where the temperature at the base of each layer is 633, 893 and  
 128 1613 K (for parameters and assumptions used to solve for the conductive profile, see  
 129 Supplementary Table 1).

130 The compositional layers deform through a combination of nonlinear viscous flow  
 131 (following dislocation creep) and brittle plastic deformation (Glerum et al., 2018). Brittle  
 132 deformation follows a Drucker Prager yield criterion:

133

$$134 \quad \sigma'_{II} = \frac{6C\cos\phi + 2P\sin\phi}{\sqrt{(3 + \sin\phi)}}$$

135

136 Where the initial friction angle ( $\phi$ ) and cohesion ( $C$ ) are 30 and 20 MPa, and they linearly  
 137 weaken as a function of finite plastic strain (see Fig. 3). Following Duretz et al. (2020) the  
 138 Drucker Prager yield criterion is modified further to include a plastic (viscous) damper, which  
 139 acts as a stabilization term for the shear band width and produces mesh-independent results  
 140 provides a sufficient resolution to resolve the damper viscosity (1e21 Pa s here).

141 Instead of a single zone of initial weakness, a distribution consisting of binary strong and  
 142 weak blocks of plastic strain are prescribed in the upper crust, following Pan et al. (2022b). The  
 143 statistically randomised method geologically mimics variations in the initial strain field that may  
 144 reflect the mechanical properties of the rifting plate (e.g. Scholz and Contreras, 1998), the  
 145 presence of the pervasive and/or discrete fabrics (e.g. Versfelt and Rosendahl, 1989) and stronger  
 146 regions such as cratons (e.g. Dunbar and Sawyer, 1989; Versfelt and Rosendahl, 1989; Tommasi

147 and Vauchez, 2001; Ziegler and Cloetingh, 2004; Buitter and Torsvik, 2014). Overall the initial  
 148 strain field promotes the development of a complex fault network displaying the along-strike  
 149 variability in strike and displacement observed in natural rifts.

150 In this paper, the plastic strain in the reference model (Model A) consists of strong and  
 151 weak values of 0.5 and 1.5 spatially over 2.5 km<sup>2</sup> blocks, with the friction angle and cohesion  
 152 weakening by a factor of 2 over the plastic strain interval. Relative to the reference model, we  
 153 investigate the effect of an increased spatial wavelength of 5 km blocks (Model E; Fig 3),  
 154 weakening the friction angle and cohesion by a factor of 4 (Model D) and reducing the contrast  
 155 between blocks to 0.5 and 0.6 (i.e. faster brittle weakening; Model C). We also investigate  
 156 extension rates of 1.25 and 5 mm/yr (see Fig. 3); these correspond to a range comparable to those  
 157 characterising slow active rifts (e.g. East Africa, Basin and Range) using GPS data (e.g. Argus  
 158 and Helflin 1995).

159

## 160 2.2. Fault network extraction

161 Fault analysis is sampled on a horizontal plane located 5 km beneath the initial model  
 162 surface as the full extent of the fault network is captured at this level (Fig. 5). Fault statistics are  
 163 extracted using an image-processing workflow from Pan et al. (2022b), which derives the strain  
 164 profile gradient from the active deformation field. This approach effectively defines a fault as a  
 165 region of clearly localised active slip, and enables the extraction of discrete fault segments within  
 166 a complex, interacting fault network (Fig 5c). Each fault label contains a unique accessible  
 167 index, and each label acts as a mask in order to extract active and finite strain, as well as to  
 168 compute geometries such as fault length and strike. In Pan et al. (2022b), fault labels are thinned  
 169 so that they are 1 element width along the fault in order to derive the cumulative euclidean  
 170 distance (i.e. fault length). In contrast, here the thinned labels are later enlarged (morphologically  
 171 dilated using a structuring element - see Supplementary Material for full details), in order to  
 172 capture the full extent of shear zone deformation, which typically span 2-3 grid elements (Fig.  
 173 5c-d). The total strain occupied and extracted by each fault label is therefore equivalent to the  
 174 geometric moment,  $M_g$  (see next section 2.3).

175

## 176 2.3. Fault strain summation

177 Incremental strain is often calculated in earthquake seismology using the Kostrov  
 178 summation method (Kostrov 1974, Molnar 1983). Similarly, strain on faults over geological  
 179 timescales can be determined using the geometric moment,  $M_g$ :

180

$$M_g = u_{ave}n$$

181 where  $u_{ave}$  is the average displacement and  $n$  is the fault surface area. The geometric  
 182 moment is directly related to the seismic moment (which is the geometric moment divided by the  
 183 shear modulus; Scholz 1989) allowing for comparison of faults active over geological timescales  
 184 with the Gutenberg-Richter power-law population of characteristic earthquake populations  
 185 (Gutenberg and Richter, 1944; Kanamari and Anderson 1975; Scholz and Cowie 1990).  
 186 Summation of the geometric moment divided by the area of regional interest ( $V$ ) allows for an  
 187 estimation of strain,  $\epsilon$  (Marrett and Allmendinger 1990; Molnar 1983):

188

$$\varepsilon = \sum_{k=1}^n \frac{M_g^k}{V}$$

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The above equation results in the total strain accommodated by the fault network (i.e. Fig. 5e). In addition to the strain occupied by our extracted faults (i.e. Fig. 5d), we also sum the total crustal strain,  $\varepsilon$  (i.e. Fig. 5a) and active strain (i.e. Fig. 5b) over the central 180 x 180 km high-resolution zone. As a result, the difference between the total crustal strain and fault strain equates to off-fault deformation (see Supplementary Figure 2).

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195

#### 2.4. Fault scaling distributions

196

197

The equation of a power law fault population distribution (e.g. Kakimi, 1980; Villemin and Sunwoo, 1987; Childs et al., 1990; Scholz and Cowie, 1990; Walsh et al., 1991) follows:

198

$$N(> \omega) = A \omega^{-\alpha}$$

199

200

201

202

where  $\omega$  is the measure of size (e.g. length, displacement or geometric moment),  $N$  is the cumulative number of values  $\geq \omega$ ,  $A$  is a constant and  $\alpha$  is the exponent. A log transformation of this equation gives a linear relationship between  $\log N$  and  $\log \omega$ , with the slope gradient representing the power law component,  $\alpha$  (Fig. 2).

203

204

An exponential scaling distribution better describes fault length distributions (Cowie et al., 1993; Ackermann et al., 2001; Spyropoulos et al., 2002) and follows:

205

$$N(> \omega) = A \exp\left(\frac{-\omega}{\omega_0}\right)$$

206

207

where the exponential law incorporates a characteristic scale  $\omega_0$ , which may reflect a physical length in the system such as layer thickness (Cowie, 1998).

208

209

210

The gamma distribution is more commonly used in earthquake statistics (Davy 1993) but it also characterises fault trace lengths in the Gulf of Corinth rift (Michas et al., 2015). The gamma law follows:

211

$$N(> \omega) = A \omega^{-\alpha} \exp_q\left(\frac{-\omega}{\omega_0}\right)$$

212

213

The gamma law distribution combines a power law from equation 4 with an exponential tail from equation 5 (Fig. 2).

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We used curve fitting and applied a non-linear least squares fit to find the optimal set of parameters for the defined function (equations 4-6). We use the geometric moment (see Section 2.3) in order to analyse the scaling distribution of strain, where the initial input variables for  $A$ ,  $\alpha$ ,  $q$  and  $\omega_0$  are the fault number, 1, 0, and the mean fault strain, respectively (following a similar methodology from Cowie et al., 2012). Only the best fitted distribution is shown in figures for clarity, however the parameters for other distributions are shown in the Supplementary Fig. 1. We assume that the smallest 30 faults are affected by truncation as the data points characteristically flatten off at this point in all models. The smallest 30 faults are therefore discarded for curve fitting, and the remaining number of faults is sufficiently large across several orders of magnitude (between 200-500 faults depending on the model) to provide a robust statistical analysis. We further discuss the implications of sampling bias in the discussion.

225

226 **3 Results**

## 227 3.1. Modelled fault network evolution

228 The general behaviour of the modelled fault network and their corresponding statistics  
229 are first described using reference Model A. All models show that fault network template is  
230 established from the initial onset of extension (Supplementary Video 1). In the first resolvable  
231 time increment (at 0.1 Myrs), the reference model A consists of a diffuse network of deformation  
232 across the model domain (Supplementary Video 1). From 0.1 - 0.5 Myrs, active deformation  
233 shows a reduction in element width as shear zones actively localise on the fault template - we  
234 attribute this behaviour as the full establishment of fault pattern lengths. After 1 Myrs of  
235 extension, a number of fault maximas appear randomly distributed throughout the model, and  
236 strain predominantly accumulates on the initial maxima points (Supplementary Video 1). The  
237 evolution of all models demonstrate similar general behaviour of widespread strain localisation,  
238 from an initially distributed deformation field, to a localised, large normal fault array (e.g. Cowie  
239 et al., 2000; Gawthorpe et al., 2003). While all investigated models reveal that fault patterns are  
240 established early and strain localises with extension, the fault pattern template and style of  
241 subsequent strain deformation exhibit considerable variability across the investigated rift  
242 parameters (Fig. 4).

243 The cumulative frequency of fault strain of reference model A shows a curved  
244 distribution similar to many other natural observations of fault scaling (e.g. Meyer et al., 2002;  
245 Soliva and Schultz, 2008). Power-law, exponential scaling, and gamma distributions all fit the  
246 observational data to  $R^2 \geq 0.99$ , although the high  $R^2$  value is attributed to the logarithmic nature  
247 of the dataset. We find that the gamma has the highest  $R^2$ , and critically, best matches the largest  
248 faults in the model (Supplementary Figure 1; Fig. 6a). Fig. 6a shows that up to 0.5 Myrs  
249 extension, the total number of faults increases as the strain increases - we correlate this to an  
250 'initiation' phase where deformation quickly localises and the lengths of the fault network  
251 pattern are fully established. For the remainder of extension, the fault distribution profile moves  
252 towards the right as faults increasingly accommodate strain, and the total number of faults stays  
253 relatively constant (Fig. 6a). The parameters of best fit show that  $\omega_0$  increases with extension,  
254 and  $q$  ranges between 1-3.  $\alpha$  stays relatively constant throughout time, ranging between 0.1-0.3  
255 (Fig. 6a, Supplementary Video 2).

256 While the central portion of the fault distribution progressively accommodates strain,  
257 reflected by stable increases in  $\alpha$ , faults accommodating the highest strain are variable  
258 throughout extension, where the magnitude of highest strain periodically fluctuates from the  $q$ -  
259 gamma fit trendline (Fig. 6a; Supplementary Video 2). We note that the periodic fluctuations of  
260 the fault distribution are sometimes initially marked by a break in slope on the log plots (e.g. see  
261 1 and 3 Myrs distribution on Fig. 6a) resulting in the appearance of a prominent tail. The  
262 fluctuations above (i.e. strain greater than) the trendline correspond to continuous active slip  
263 events, where the summed strain along large fault lengths is high. Conversely, deviations below  
264 the trendline suggest that the strain intermittently accommodated by faulting is relatively  
265 distributed throughout the crust (rather than accommodated by a few large faults). The total  
266 summed strain accumulation, accommodated by the extracted fault network, increases



267 approximately linearly through time, however there are small temporal fluctuations which reflect  
268 the oscillations between distributed and localised deformation (Fig. 6b).

269

## 270 3.2. Investigated rift model parameters

### 271 3.2.1. Extension rate

272 Models deforming at constant rates of 1.25 and 5 mm/yr (Models A and B; Fig. 3) allow  
273 for the investigation of extension rate on fault network evolution. The models show that higher  
274 extension rates (5 mm/yr) correspond to an overall higher magnitude of plastic strain  
275 deformation over increasingly diffuse regions (Fig. 7a and b). Our results also show that an  
276 increase in extension rates do not affect the overall spatial pattern of finite strain (Fig. 7a and b).  
277 For example, fault maximas and localisation occur on the same fault systems no matter the  
278 extension rate, although to some extent the fault strain magnitudes and shear width of  
279 deformation are variable across the investigated models (Fig. 7a and b).

280 Similar to the other investigated parameters, fault strain distributions for all models  
281 investigating extension rates abide by a q-gamma distribution ( $R^2 > 0.99$ ; Fig. 7). The  
282 distribution appears broadly similar across the investigated models, whereby greater extension  
283 rates accommodate higher strain throughout extension, so plot further towards the right in Fig. 7.  
284 For both models,  $\omega_0$  increases with extension, but overall magnitudes are higher for the 5 mm/yr  
285 model - at the final 5 Myrs timestep shown,  $\omega_0$  is  $2.7 \times 10^6$ , 3 times higher than that of the 1.25  
286 mm/yr model at  $8.7 \times 10^5$  (Fig. 7). For the 2.5 mm/yr model, alpha steepens from 0.21 to 0.07  
287 from 1 to 2 Myrs, respectively, before increasing for the remainder of extension (Fig. 7a). For the  
288 5 mm/yr model, alpha increases (i.e. the central portion of the slope shallows) with extension,  
289 and reaches a relatively higher value of 0.32 by 5 Myrs extension (in contrast to 0.2 for the 1.25  
290 mm/yr model).

291

### 292 3.2.3. Rheology

293 Our results indicate that changes to crustal rheology exert the most significant control on  
294 fault patterns, both visually due to along-strike variations in strike, strain, length and density, and  
295 statistically due to differences in their scaling distributions (Fig. 9). Most notably, models with  
296 relatively weak faults (Models C and D) contain less faults, thus strain is localised onto fewer,  
297 larger structures (Fig. 8b and d). The cumulative strain profiles of these two weak models reflect  
298 their localised behaviour, where the gradient of strain distribution trends shallower relative to the  
299 reference model (Fig. 9). The q-gamma fitted parameters show that highly localised fault  
300 populations such as Models C and D both exhibit lower  $\alpha$  values of c. 0.03 and 0.09 and thus  
301 shallower trends for the central portion, and contain higher values of  $\omega_0$ . Although the  
302 cumulative strain distributions are comparable between the two weak models (Model C and  
303 Model D), their fault template are significantly different whereby Model D exhibits significantly  
304 more pronounced along-strike variability (Fig. 8d).

305 Investigations into the spatial wavelengths within the initial distributions of plastic strain  
306 (Model A:  $2.5 \text{ km}^2$  or Model E:  $5 \text{ km}^2$ ) here reveal that strain scaling distributions appear  
307 comparable as both exhibit relatively higher  $\alpha$  values (0.27 and 0.2, respectively) and lower  $\omega_0$   
308 values (Fig. 8a and e). However, the faults in Model E overall accommodate less strain  
309 magnitude in their scaling distribution (Fig. 8e), and the visual style of the pattern of deformation

310 differs significantly between the two models, with larger initial blocks producing more  
 311 continuous, sinuous fault patterns with less along-strike variation (Fig. 8c).

312

### 313 3.3. Fault-accommodated strain through time

314 Fig. 9 shows the total fault accommodated strain through time across all models, and a  
 315 snapshot of their fault strain distribution at 5 Myrs. For the cumulative distributions in Fig. 9a,  
 316 the geometric moment observed from the NW Shelf of Australia (Meyer et al., 2002; Pan et al.,  
 317 2022a) are shown. Fault strain from Pan et al. (2022a) is summed and divided by the studied area  
 318 size (1200 km<sup>2</sup>) for a dimensionless comparison of strain. Similarly from Bell et al. (2011), 2D  
 319 cross-sectional strain across three transects across the Corinth Rift are divided by the rift width.  
 320 Due to a range of uncertainty on the strain summation and age of rifting, we plot upper and lower  
 321 bound estimates (see Supplementary Material for values).

322 Overall the 5 rift models reveal that summed fault strain increases linearly through time  
 323 (Fig. 9b) when at constant extension rates. The rate at which summed fault strain increases is  
 324 variable, and lower rates correlate to lower rates of extension (Fig. 9b). Within models of the  
 325 same 5 mm/yr extension rate (Models A, C, D, E), higher rates of strain accumulation occur in  
 326 models characterised by relatively weak crust and/or faults (i.e. Models C and D). Overall we  
 327 find that the range of fault strain accumulation lies within the observational fault data (Fig. 10b).

328 Fig. 9a reveals that the strain scaling distribution of faults all exhibit the typical curvature  
 329 characterised by observational datasets. Fault populations are best fit by a q-gamma distribution  
 330 but exhibit variable fitted parameters, particularly q and  $\omega_0$ . The lower 1.25 mm/yr extension rate  
 331 model plots closest to the left, consistent with findings that it exhibits lower strain magnitude in  
 332 comparison to the other models (Fig. 9a). In contrast, Models C and D with weaker strain  
 333 mechanisms plot furthest to the right and thus their faults accommodate the highest strain  
 334 magnitudes. Noticeably, these models are characterised by lower, shallower power law values (c.  
 335 0.11-0.15) and higher  $\omega_0$  values of c.  $6 \times 10^6$  (Fig. 9a).

336

### 337 3.4. Off-fault deformation

338 The summed difference between extracted fault accommodated strain and total crustal  
 339 strain (e.g. Fig. 6) may account for diffuse deformation, which we term as off-fault deformation  
 340 (OFD). Supplementary Fig. 2b shows how the OFD of Model A changes through time. OFD%  
 341 rapidly decreases within the first 0.8 Myrs, which we correlate to fault organisation and initiation  
 342 (Supplementary Fig. 2b). For the remainder of time, OFD% *progressively* reduces,  
 343 corresponding to progressive strain localisation onto faults (Supplementary Fig. 2b). A similar  
 344 model behaviour to the reference model is depicted in Fig. 10, where the majority of the scatter  
 345 and outliers in the plot is due to the initiation phase in the first 1 Myrs. An exception is the model  
 346 (B) with a slow extension rate (1.25 mm/yr), which shows a reverse trend of fault localisation  
 347 from the onset of extension (Fig. 10). On average, Fig. 10 shows that OFD% accounts for  
 348 approximately 40% in the models, ranging from 25% to 45% depending on the investigated rift  
 349 parameters. We find that the model (C) with an increased rate (10x) of fault weakening contains  
 350 the lowest proportion (25%), followed by the model (D) with a large magnitude (4x) of fault  
 351 weakening - given the weakened mechanisms this corresponds to the high occurrence of strain  
 352 localisation observed (Section 3.2.3).

353

## 354 **4 Discussion**

### 355 *4.1. Fault strain comparison between numerical models and natural observations*

356 The extraction of fault geometry within an entire fault population has enabled us to  
357 demonstrate that modelled faults are *quantitatively* comparable to fault geometries observed in  
358 rifts (Fig. 9). Modelled fault populations exhibit a curved distribution similar to size distributions  
359 observations in the literature for length (e.g. Scholz et al., 1993; Yielding et al., 1996; Odling,  
360 1997), displacement (Marrett and Allmendinger, 1992; Knott et al., 1996) and geometric moment  
361 (e.g. Meyer et al., 2002; Bailey et al., 2005; Fig. 9a). The total fault strain of the models also  
362 quantitatively plot within the range of strain derived for observational data (Fig. 9b). Models  
363 which experienced 1.25 mm/yr rates of extension lie closest to the observational range for the  
364 Exmouth Plateau - taken at face value, this predicts that ancient extension rates are relatively low  
365 based on its magnitude strain through time.

366 Although our results have demonstrated that fault strain is largely comparable between  
367 models and observations, we find that direct rift inversions to recover rift parameters using  
368 summed strain measurements (i.e. Fig. 9b) are still not achievable due to uncertainties in i)  
369 measured observational strain, which is subject to interpretative biases and limited spatial  
370 resolution (e.g. Bonnet et al., 2001); and ii) overall rift duration, as the structure of active and  
371 ancient rifts observed provide only a snapshot in time, and there is often uncertainty on the age  
372 of the oldest syn-rift sediments to date the rift age as these deep sediments are rarely drilled.  
373 Fault backstripping may enable further data points of strain-time, however the earliest resolvable  
374 growth strata preserved in ancient systems rarely precedes 5 Myrs of rifting, therefore additional  
375 modelling of longer timescales may be needed for comparison.

376

### 377 *4.2. Normal fault growth and rift evolution*

378 The spatial and temporal scales covered by numerical modelling give insight on the  
379 earliest stages of fault growth, which are not easily constrained using seismic reflection and  
380 borehole data (Jackson et al., 2017). Here, all models establish their finite strain patterns from  
381 the nearly onset of extension, i.e. faults abide by the ‘constant-length’ model (Stage 1; Fig. 11)  
382 consistent with Walsh et al. (2002). The rapid establishment of fault patterns within the first  
383 <100 kyrs is attributed to the initial randomisation of the crustal (plastic) strain field. Our results  
384 demonstrate that the weaknesses in the initial plastic distribution control the loci for which strain  
385 maxima form (Stage 2; Fig. 11) and which eventually become the future sites of displacement  
386 accumulation (Stage 3; Fig. 11), similar to the findings of Cowie (1995). The behaviour of strain  
387 localisation onto through-going faults is consistent with Sornette et al. (1990), Cowie (1998) and  
388 Gupta et al. (1998), supporting that it is a fundamental characteristic of fault network evolution.  
389 While all models undergo strain localisation, models with larger (i.e. weaker) brittle strength  
390 contrasts between initial binary blocks or a faster rate of brittle weakening (C and D) trend more  
391 exponentially from the onset of extension (e.g. Fig. 8). The investigations of initial strain  
392 distribution in the models highlight the importance of crustal strength variations, which may be  
393 distributed randomly, or contain spatial significance i.e. pre-existing structures (e.g. Duffy et al.,  
394 2015), shear zones (e.g. Phillips et al., 2016) and terrane structures (e.g. Daly et al., 2014). We  
395 suggest that the relative and/or bulk strength from structural inheritances play an important role

396 in structural development and may be further investigated through numerical modelling, where  
397 the initial strain field can be statistically quantified.

398

#### 399 4.3. Fault strain scaling distribution evolution

400 Despite the assumption that faults abide by a power law scaling, we find that the  
401 distribution of the modelled faults is best described by a q-gamma distribution (a power law  
402 across moderate sized faults, with an exponential tail for the largest faults within a population),  
403 consistent with earthquake and seismic hazard assessments (Davy, 1993; Kagan 1997; Sornette  
404 and Sornette, 1999; Main 2000) and a study of a naturally occurring fault network in the Gulf of  
405 Corinth (Michas et al., 2015). Our results show that faults are not entirely fractal; instead, the  
406 distribution curve becomes steeper due to an upper bounding characteristic scale (e.g. Fig. 6).  
407 We argue that the large bound limit of the fault distribution in our results is real and is not due to  
408 censorship as the modelled area (180 x 180 km) is rift scale and is sufficiently large to capture  
409 the biggest strain-accommodating faults. Furthermore, the position at which the fall-off occurs is  
410 significantly greater than previous deductions of censorship effects, and occurs across all  
411 investigated models of different strain magnitudes. We speculate that the upper limit of fault  
412 distributions is related to similar upper bounding limits that define observational D-L profiles -  
413 faults cannot physically exhibit strain beyond a certain limit due to the thickness and strength of  
414 the crust (e.g. Cowie and Scholz, 1992).

415 Across all investigated model parameters,  $\alpha$  remains relatively consistent between 0-0.3  
416 and is comparable to similar scaling distributions of geometric moment from the NW Shelf of  
417 Australia ( $\alpha = 0.4-0.5$ ; Fig. 9). When fitted with a powerlaw distribution, the scaling exponent is  
418 c. 0.7 (consistent with existing literature which ranges between 0.3 - 2 (Villemin and Sunwoo  
419 1987; Scholz and Cowie 1990; Turcotte et al., 2007; Marrett and Allmendinger 1992; Cladouhos  
420 and Marrett 1995; Yielding et al., 1996; Knott et al., 1996; Schlische et al., 1996). Previous  
421 numerical modelling (Cowie et al., 1993, 1994) and outcrop data (Wojtal 1986, 1994, 1996)  
422 suggest that the power-law scaling exponent decreases through time, due to the progressive  
423 concentration of strain onto large faults. While our models clearly exhibit progressive strain  
424 localisation, and the fault distributions here are visually comparable to those in aforementioned  
425 studies, our study correlates the behaviour of fault distribution profiles to the upper-bound  
426 exponential scaling characteristic,  $\omega_0$ , which increases with time (Fig. 6, 9 and 11).

427 Our results demonstrate that while the majority of the fault network distribution steadily  
428 accumulates displacement through time (as the plot moves towards the right with consistent  $\alpha$ ),  
429 the tail-end of the distribution (i.e. strain accommodated on the largest faults) is more transient  
430 (Fig. 6). We correlate the transient behaviour of large faults to the complexity of the fault  
431 network and interaction of fault segments due to competing stress redistribution (e.g. Cartwright  
432 et al., 1995; Dawers and Anders, 1995). Here, the extracted faults exhibit long lengths when they  
433 interact, producing strike-continuous slip events, and subsequently strain is redistributed along  
434 the entire length of the fault system (across relay ramps) resulting in a high fault strain value. We  
435 find that the occurrence of a prominent tail in log plots (e.g. at 2 Myrs on Fig. 6; see also Fig. 11)  
436 may be reflective of linkage events for large faults, consistent with initial discussions from  
437 Wojtal et al. (1996). The break in slope on log-log plots and prominent tail is apparent in fault  
438 size distribution plots from experimental models (e.g. Ackermann et al., 2001), and field studies  
439 on Earth (e.g. Casey et al., 2003; Soliva and Schultz, 2008; Gudmundsson et al., 2013) and other

440 planetary bodies (e.g. Vallianatos et al., 2016). While some studies discuss the presence of a  
441 prominent incipient tail (e.g. Walsh et al., 2003; Gudmundsson et al., 2013; Soliva and Schultz,  
442 2008), its relevance is not fully characterised or understood, and we encourage future work to  
443 investigate its significance.

444 Our attempt to characterise scaling laws have highlighted how sampling protocols and  
445 fitting procedures can and have led to different outcomes (i.e. exponential, power law or a  
446 combination of both), particularly due to the effect of sampling biases. Specifically, the cut-off  
447 point used to account for truncation is highly subjective and often assumes a power law scaling  
448 distribution *a priori* (Supplementary Fig. 3). Historically, geoscientists perform a log-log  
449 transform of  $\omega$  and  $N(>\omega)$  (equation 4 in Methods; see Fig. 2 for log-log transform) and either fit  
450 by eye or use a least-squares estimation of the resulting straight line after truncating their data  
451 (Pickering et al., 1995). Such methods are not statistically validated and can produce inaccurate  
452 estimates (Clark et al., 1995; Bonnet et al., 2001; Clauset et al., 2009). We suggest that the  
453 inconsistency of such procedures have led to different scaling conclusions. For example,  
454 analogue modelling from Ackermann et al. (2002) found that size distributions changed from  
455 powerlaw to exponential. Similarly, observational data from Gupta and Scholz (2000) found that  
456 size distributions are power law for low strain, and exponential for high strain settings. In  
457 contrast, for fault distributions that appear similarly curved, Cowie (1995), Bonnet (1997) and  
458 Soliva and Schultz (2008) describe a transition from distributed to localised faulting as an  
459 initially exponential distribution that evolves into a power law scaling distribution.

460 Overall, the characterisation of fault strain distributions has provided insight on how the  
461 fault population evolves with time, the characteristics of which are not easily conveyed by  
462 conventional displacement-length plots. Given changes in scaling exponent are regarded as  
463 precursors to large earthquakes (Smith, 1981) and, for volcanic areas, precursors to eruptions  
464 (Gresta and Patane, 1983), we suggest that future work should provide a comprehensive,  
465 statistically robust investigation on fault scaling distributions that will enable a better  
466 understanding of fault evolution and its implication to related geohazard.

467

#### 468 4.4. Off-fault deformation

469 Our results demonstrate that OFD accounts for at least 25 to 45% of total crustal strain,  
470 depending on the investigated rift parameters (Fig. 10). These results are consistent with the  
471 presence of so-called ‘hidden’ deformation, initially proposed by Kautz and Sclater (1988),  
472 where faults accommodated 40-50% of the known extension in clay analogue models, and 70-  
473 80% in sand analogue models. Our results lie within the range of  $40\% \pm 23\%$  OFD determined  
474 for the East Californian Shear Zone (Herbert et al., 2014) and within the range of OFD%  
475 calculated using ground-penetrating radar and palaeoseismic trench data, where drag folding and  
476 fault block rotations accommodated 46% of extension in the Taupo Rift (McClymont et al.,  
477 2010).

478 At the crustal scale, Marrett and Allmendinger (1992) and Walsh et al (1991) find that as  
479 much as 40% of the total extension is accounted for by small-scale faulting in seismically  
480 imaged extensional basins of the North Sea. In contrast, Scholz and Cowie (1990) argue that  
481 small-scale faulting accounted for less than 10% of the total strain budget, and Cowie (1995)  
482 reconcile both arguments by proposing that the relative importance of small faults decreases as  
483 the total strain increases due to strain localisation. As our results show that a significant

484 proportion of strain is accommodated off-fault, such that deformation may be distributed in the  
 485 form of fault drag, rotated fault blocks and/or pure bulk thinning of the crust, we suggest small-  
 486 scale faulting (estimated through the extrapolation of fault size distributions) is not sufficient to  
 487 account for total strain, and is only partially responsible for the discrepancy between summed  
 488 fault strain and extension predicted by thermal subsidence (e.g. Marsden et al., 1990). In  
 489 addition, our results show that OFD% progressively decreases to account for strain localisation,  
 490 however we find that the decrease accounts for <10%, and overall OFD% may still remain  
 491 relatively high by the end of model time (Fig. 10). This suggests that OFD% in the crust is  
 492 inherent to fault evolution during continental extension.

493

## 494 **Conclusions**

- 495 1. We demonstrate that high-resolution 3D numerical models of continental extension  
 496 statistically produce geologically realistic faulting patterns.
- 497 2. Numerical models show that fault patterns are rapidly established from the onset of  
 498 extension (within <100 ka) and thus abide by the ‘constant-length’ fault growth model.
- 499 3. We find that distribution of initial strain and rate of fault weakening exert the strongest  
 500 control on fault pattern variability- this highlights the significant role of pre-existing  
 501 crustal fabric in controlling rift geometry
- 502 4. The scaling distribution of strain is not power law across all scales; instead all modelled  
 503 fault populations are subject to an upper bound characteristic scale.
- 504 5. The amount of strain accommodated by faults is quantified, and we find that diffuse, off-  
 505 fault deformation accounts for 25-45% of total crustal strain, the extent of which depends  
 506 on rift parameters.
- 507 6. A robust statistical analysis of fault scaling distributions with consideration of sampling  
 508 biases across scales is needed to fully deduce fault scaling laws.
- 509 7. The characterisation of strain distributions combined with a better understanding of  
 510 OFD% may enable the future inversion of rift controls through observations alone.

511

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517

## 518 **Open Research**

519 The model simulations were run with ASPECT version 2.5.0-pre (main, e911bed). This  
 520 specific version can be obtained by cloning the ASPECT repository (git clone  
 521 <https://github.com/geodynamics/aspect/>) then within the repository checking out the specific  
 522 commit (git checkout e911bed). ASPECT was compiled with deal.II 9.4.0, Trilinos 12.8.1, and

523 p4est 2.3.2. dealII and the aforementioned libraries were compiled with the candi installation  
524 package (<https://github.com/dealii/candi>, branch deal.II-9.4).

525 The ASPECT model data has been interpolated as 2D numpy files and are included in  
526 the supplementary files. From this, an image-processing based workflow is applied on the  
527 interpolated numpy files to automatically extract the faults – the python script is provided in the  
528 supplementary files. The script outputs a spreadsheet table with geometrical properties of each  
529 extracted fault, and a jupyter notebook in the supplementary file provides the code to reproduce  
530 the figures directly from the spreadsheet. The supplementary files can be found at  
531 <https://figshare.com/s/dcc8ac81f0bd0fa50a0f>.

532

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776

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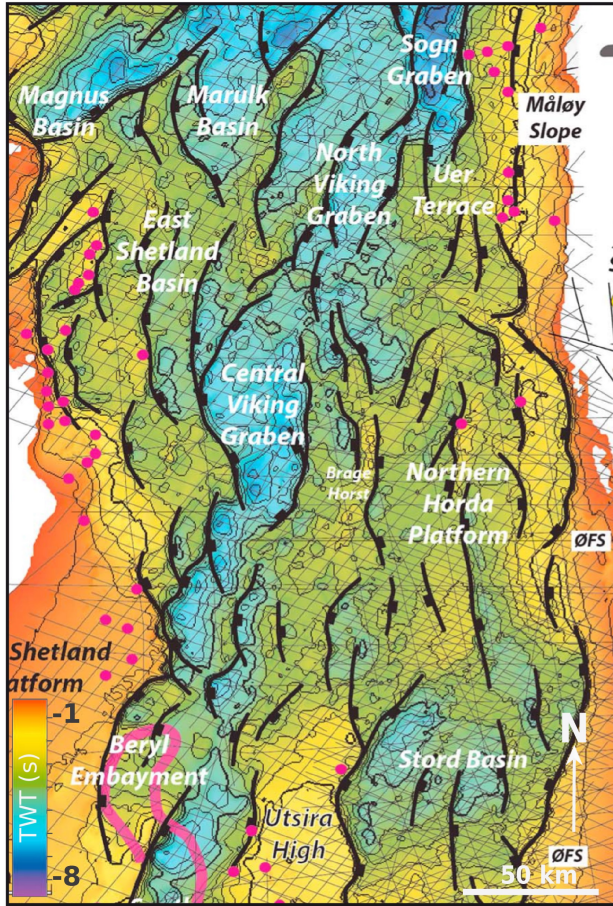
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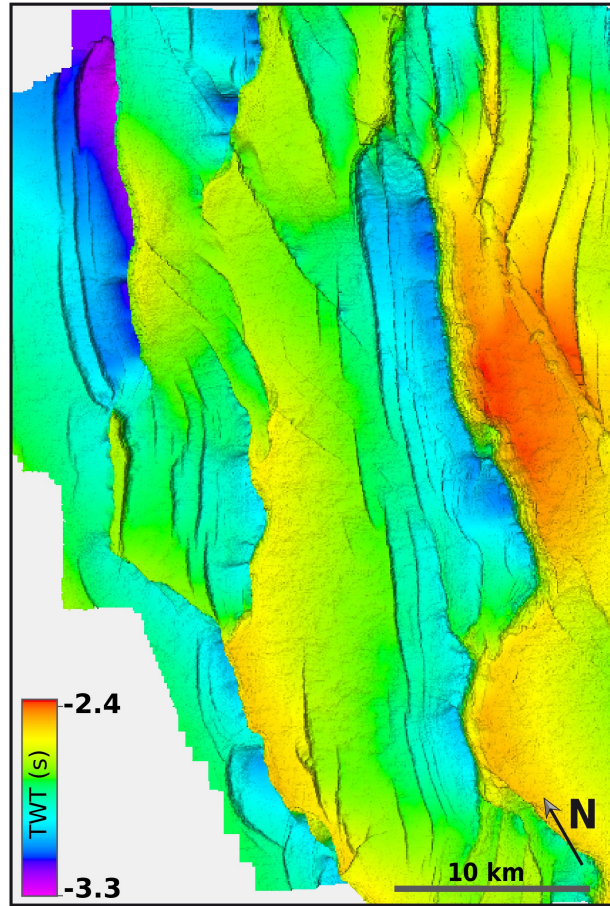
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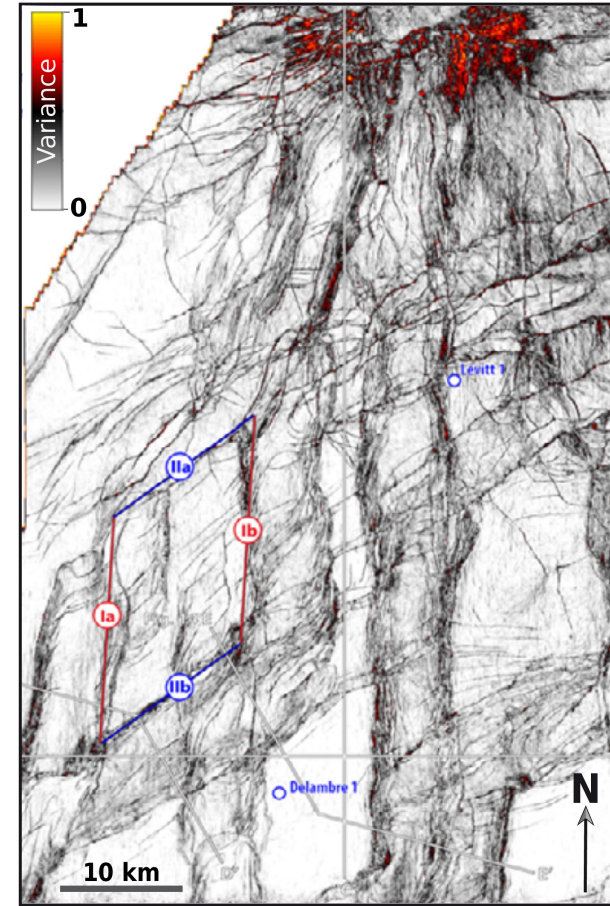
Northern North Sea  
(Fazlihani et al., 2017)



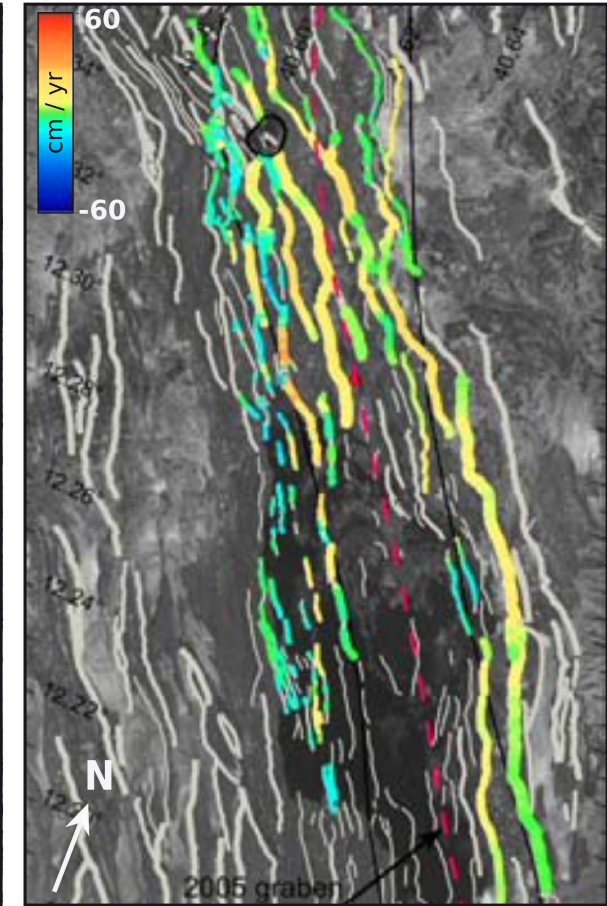
Exmouth Plateau, NW  
Australia (Pan et al., 2021)



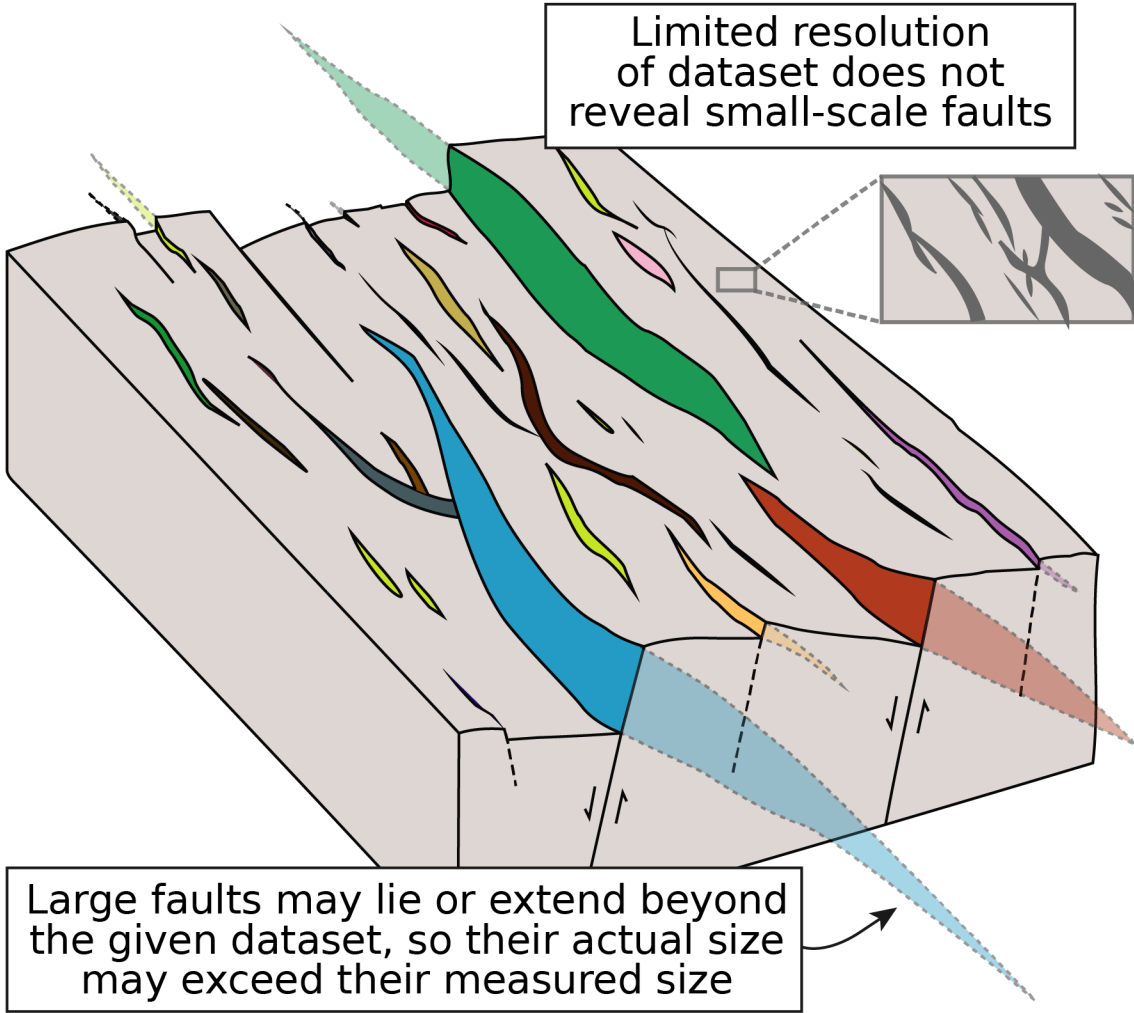
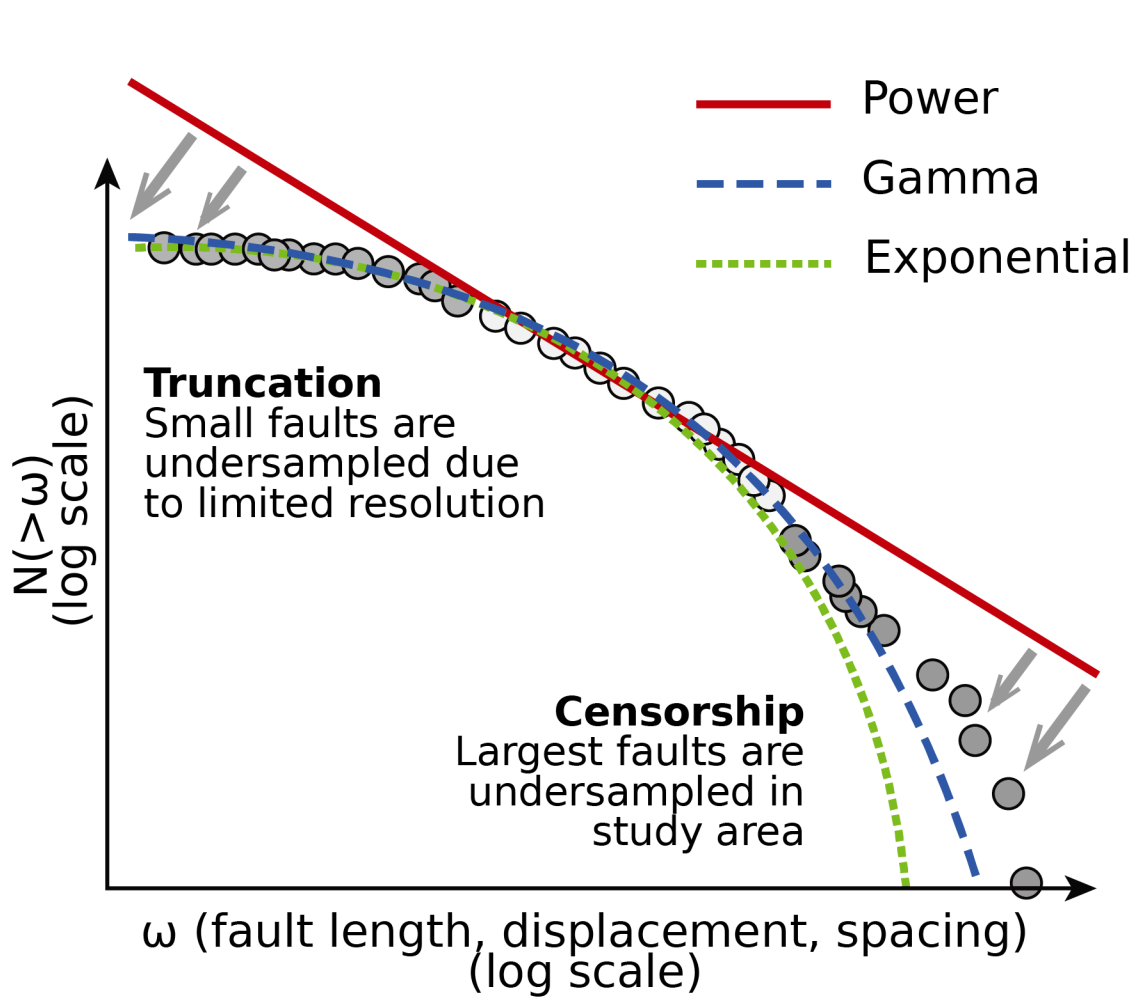
Beagle Sub-basin, NW  
Australia (McCormack et al., 2018)



Dabbahu-Manda-Hararu,  
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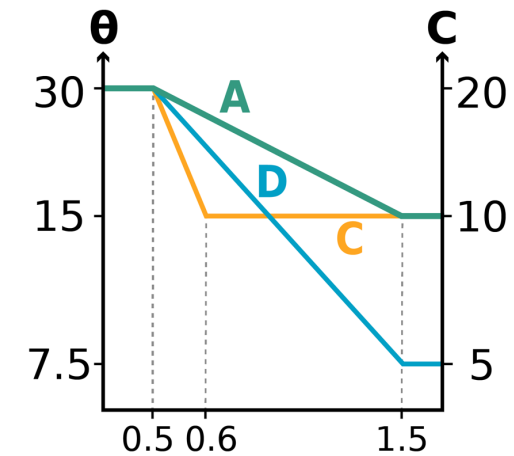
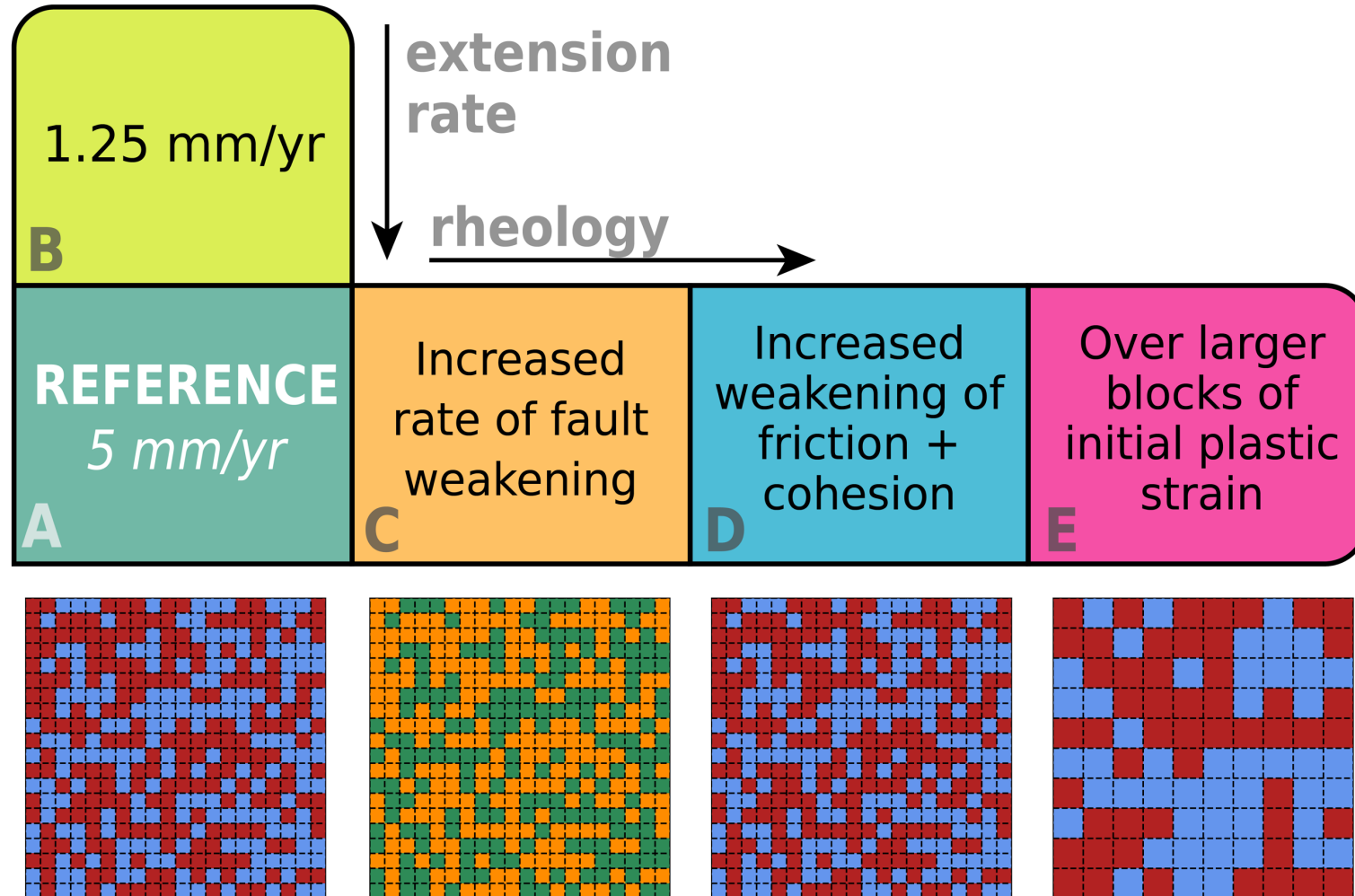


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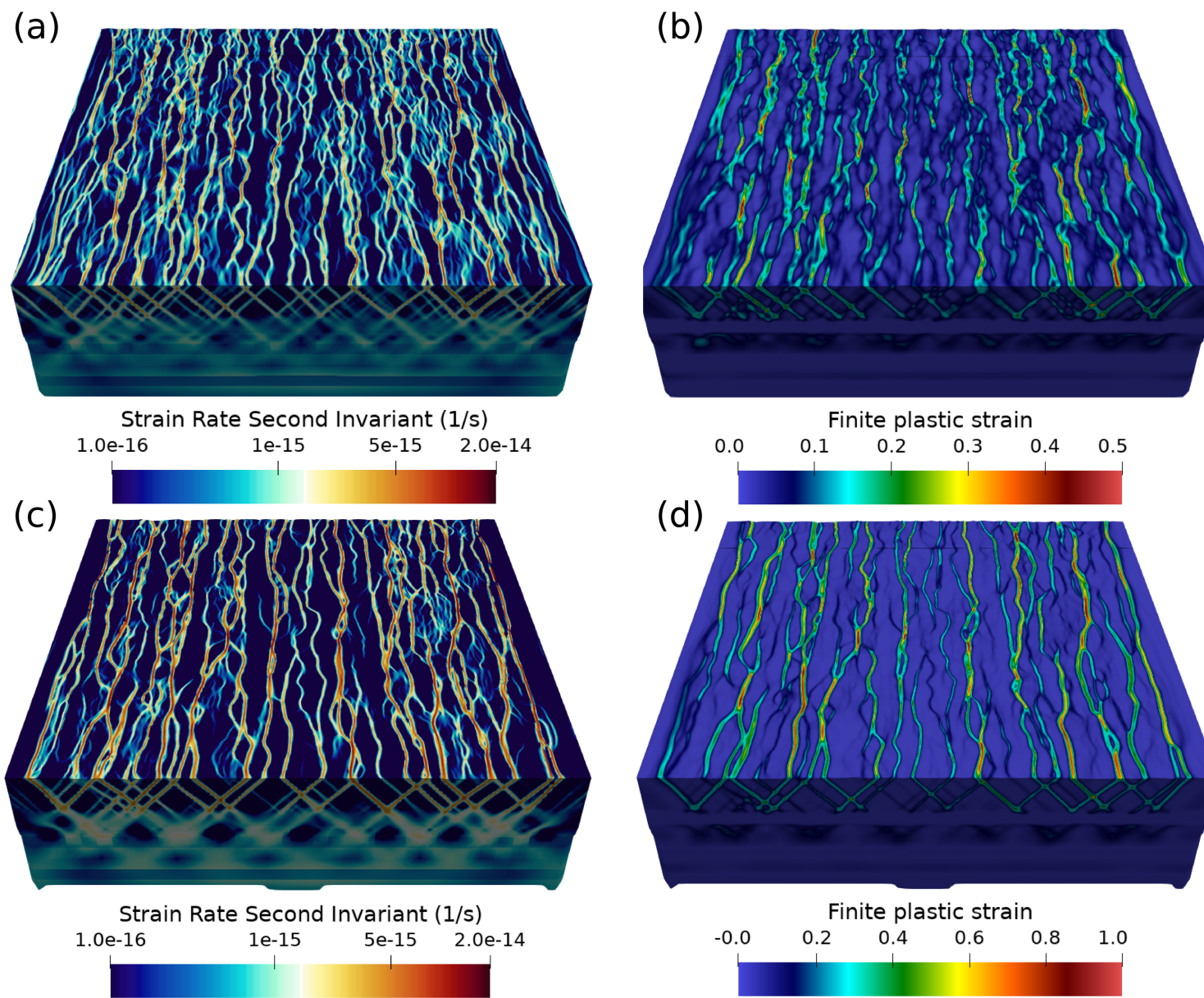




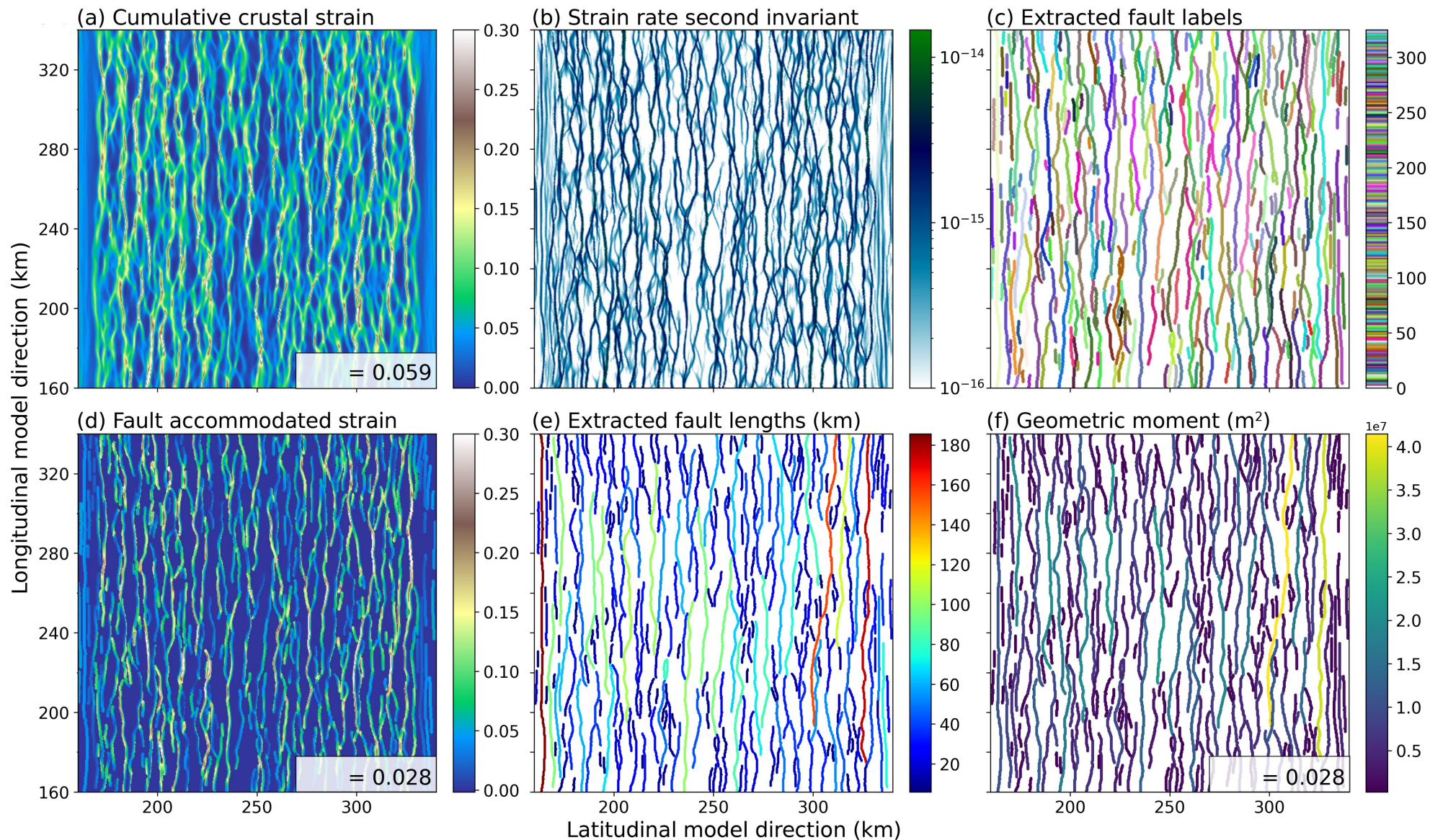
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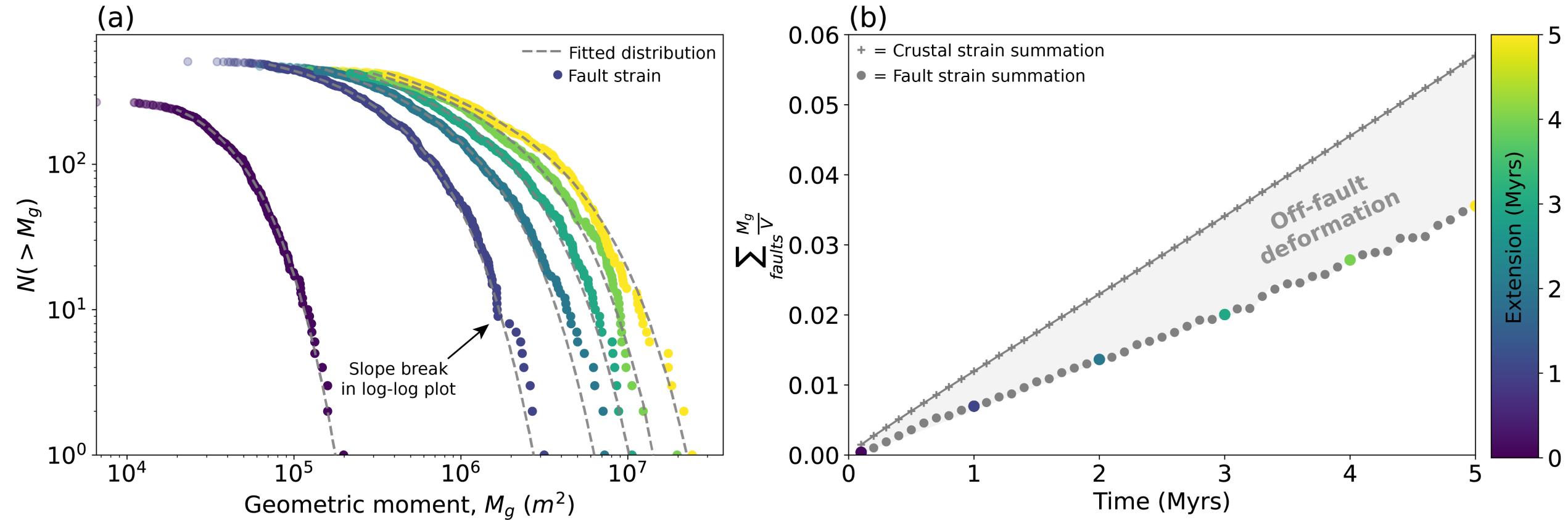
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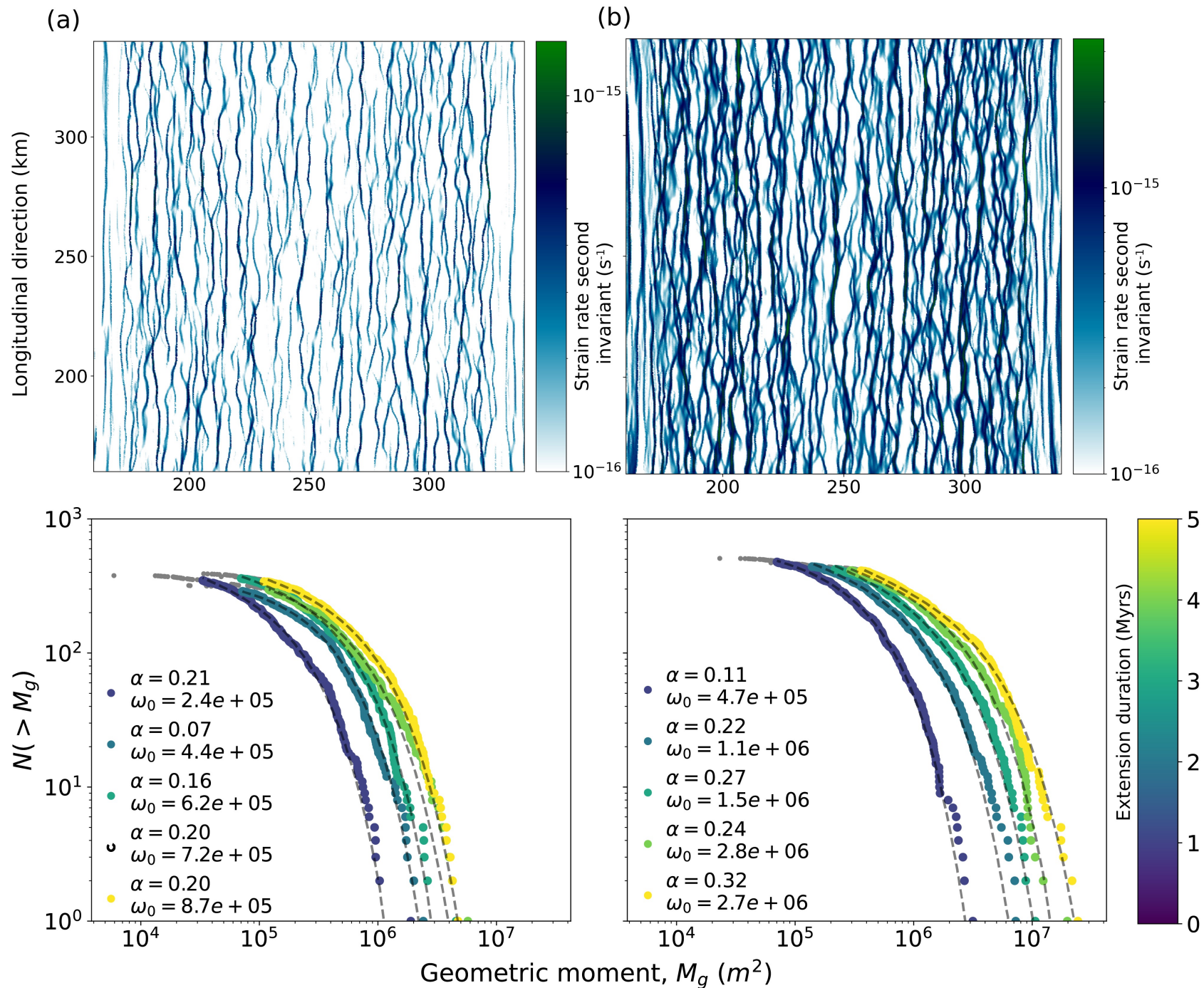
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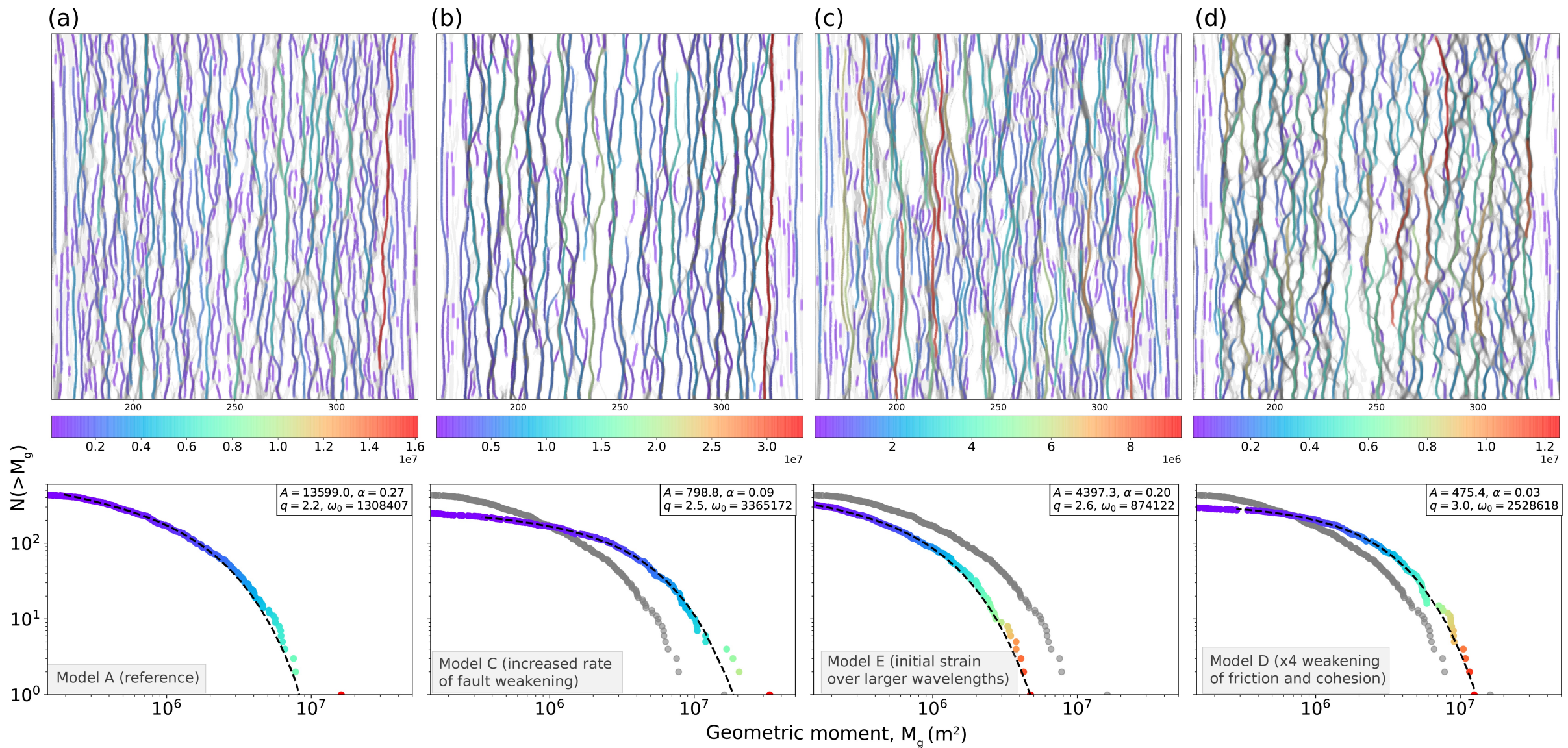
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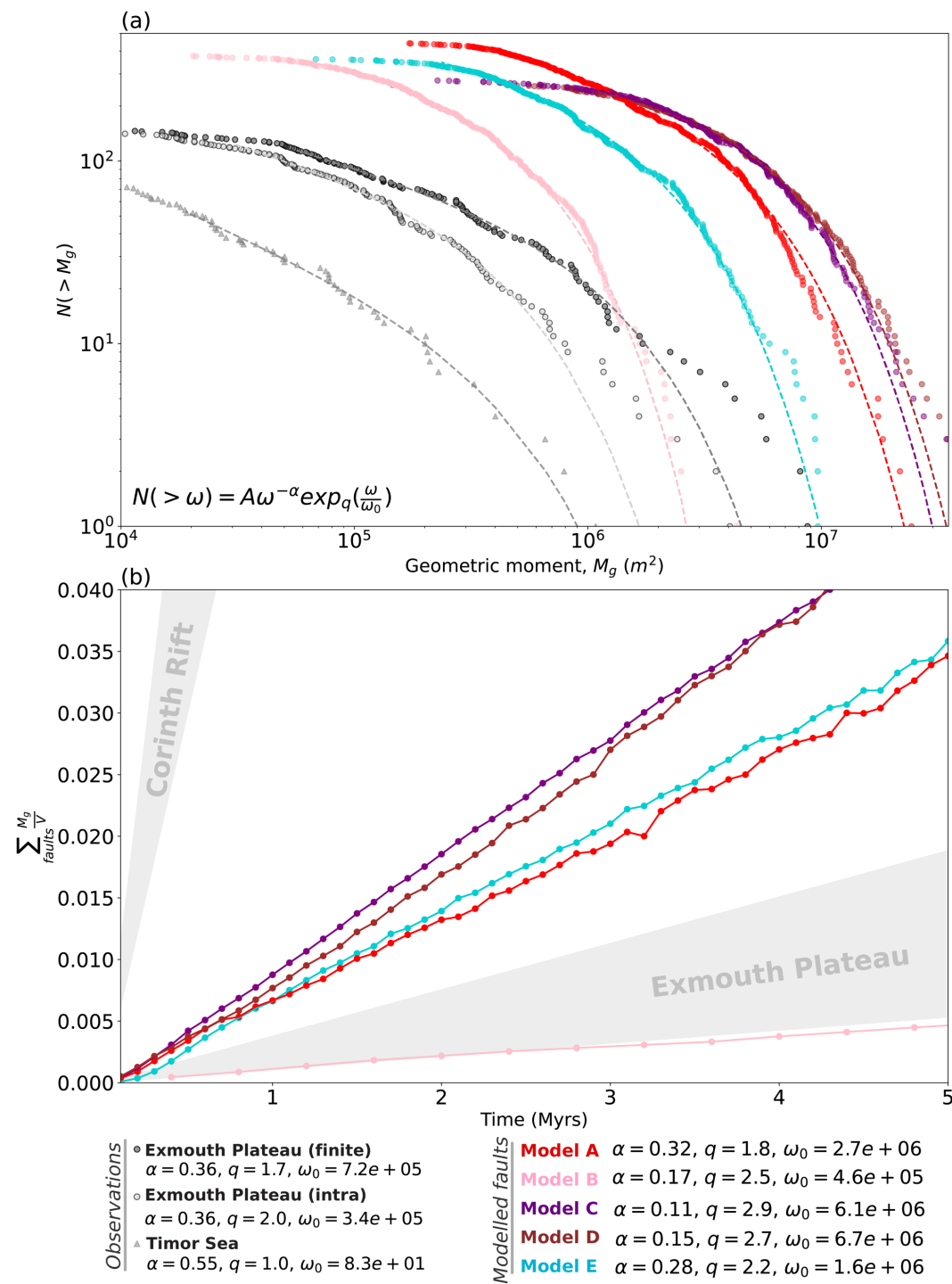
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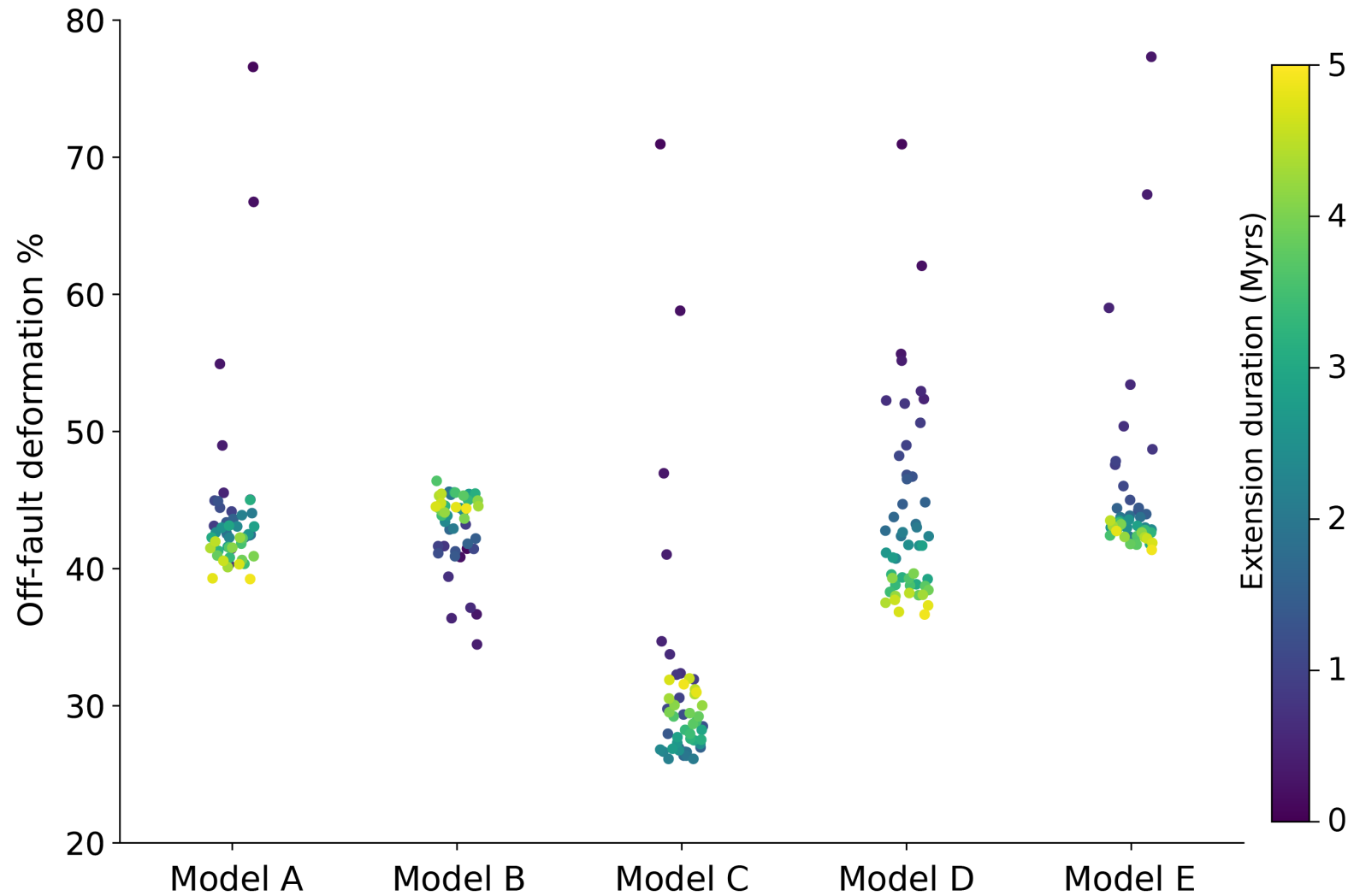
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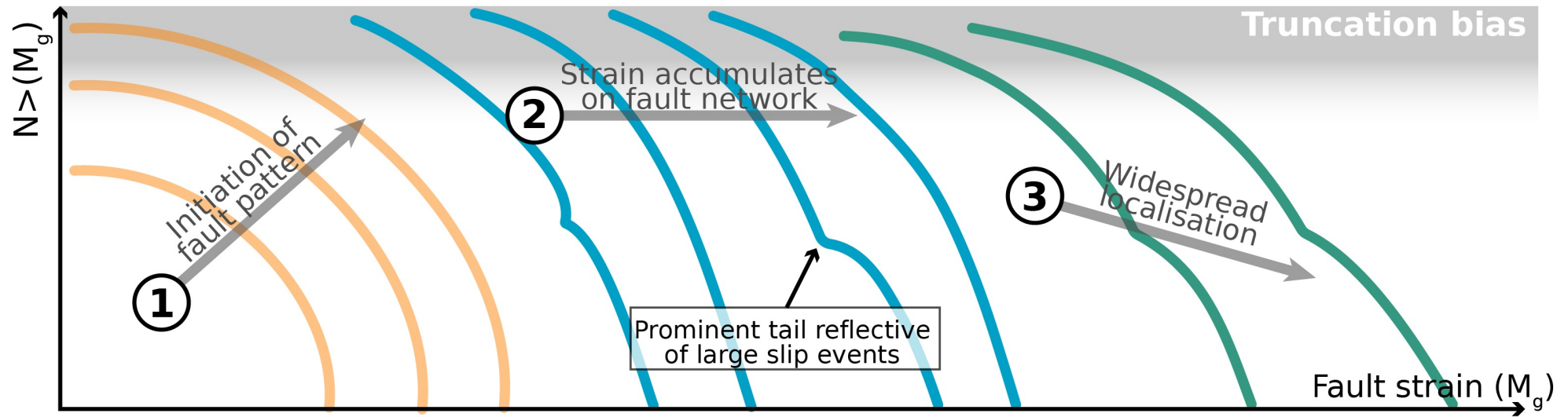


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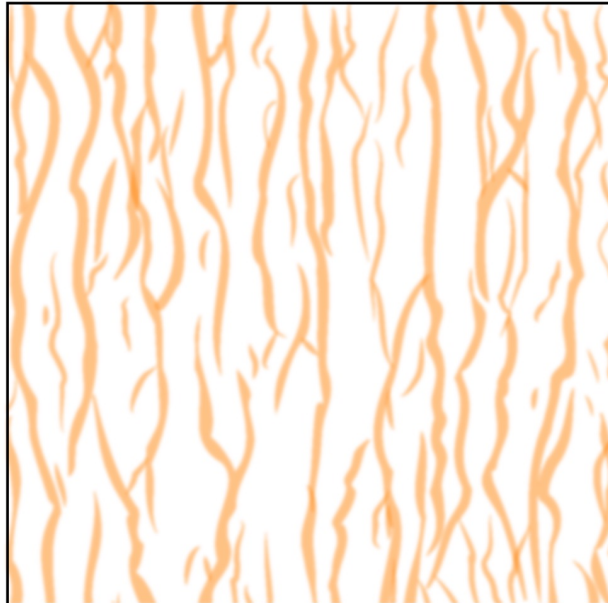




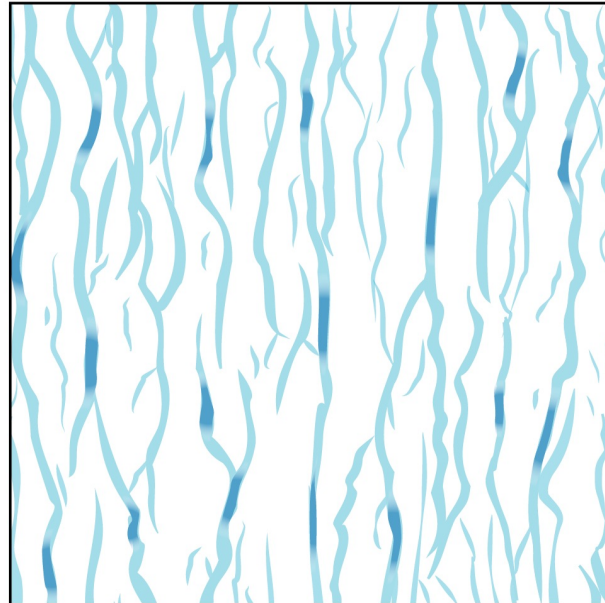
**Fig 11.** Schematic evolution of fault distribution and network evolution



**1)** Diffuse fault pattern forms, entire network rapidly localises



**2)** Strain accumulation with randomly distributed maxima



**3)** Permanent localisation as largest faults accommodate strain

