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Extension of Methane Emission Rate Distribution for Permian Basin Oil and Gas Production Infrastructure by Aerial LiDAR

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Abstract

²
³ Aerial LiDAR measurements at 7474 oil and gas production facilities in the Per-
⁴ mian Basin yield a measured methane emission rate distribution extending to the de-
⁵ tection sensitivity of the method, 2 kg/h at 90% probability of detection. Emissions are
⁶ found at 38.3% of facilities scanned, a significantly higher proportion than reported in
⁷ lower-sensitivity campaigns. LiDAR measurements are analyzed in combination with
⁸ measurements of the heavy tail portion of the distribution (> 600 kg/h) obtained from
⁹ an airborne solar infrared imaging spectrometry campaign by Carbon Mapper (CM).
¹⁰ A joint distribution is found by fitting the aligned LiDAR and CM data. By compar-
¹¹ ing the aerial samples to the joint distribution, the practical detection sensitivity of
¹² the CM 2019 campaign is found to be 280 kg/h [256, 309] (95% confidence) at 50%
¹³ probability of detection for facility-sized emission sources. With respect to the joint
¹⁴ model distribution and its confidence interval, the LiDAR campaign is found to have
¹⁵ measured 103.6% [93.5%, 114.2%] of the total emission rate predicted by the model for
¹⁶ equipment-sized emission sources (~ 2 m diameter) with emission rates above 3 kg/h,

17 whereas the CM 2019 campaign is found to have measured 39.7% [34.6%, 45.1%] of the
18 same quantity for facility-sized sources (150 m diameter) above 10 kg/h. The analysis
19 is repeated with data from CM 2020-21 campaigns with similar results. The combined
20 distributions represent a more comprehensive view of the emission rate distribution in
21 the survey area, revealing the significance of previously underreported emission sources
22 at rates below the detection sensitivity of some emissions monitoring campaigns.

23 **Introduction**

24 Methane is a potent greenhouse gas with a warming potential 80 times greater than that of
25 CO₂ in a 20-year time frame.¹ Its current global emission rate is great enough to impact the
26 climate significantly, with a greater contribution to global temperature rise in the first ten
27 years after emission than CO₂ at current emission rates of both gases.² Consequently, mit-
28 igation of methane emissions is viewed as particularly important for meeting climate goals
29 within the next decade. Economic sectors including agriculture, waste disposal, and energy
30 are recognized as leading contributors to anthropogenic methane emissions, representing do-
31 mains where emissions can be most meaningfully mitigated. In the oil and natural gas (O&G)
32 industry, emissions arise from discrete infrastructure elements and associated processes that
33 can often be addressed with targeted intervention. Mitigation involves both the detection of
34 emission sources and follow-up with repair and/or upgrade of emitting equipment. Identi-
35 fying the most important emissions drivers and tracking the efficacy of mitigation efforts is
36 key to making emissions reductions effective and efficient.^{3,4}

37 Broadening the view of emissions from individual sources to a distribution of sources
38 provides large-scale context to set meaningful mitigation goals. Past characterization of
39 methane emission distributions has often relied on bottom-up estimates based on emission
40 factors, such as those used for the U.S. Environmental Protection Agency's Greenhouse
41 Gas Reporting Program and Greenhouse Gas Inventory. These estimates aim to identify
42 dominant emission sources at the component or equipment levels but have been shown to

43 misrepresent large-scale methane emissions distributions and the relative contribution of
44 different elements,^{3,5-8} with the greatest discrepancies existing in the production sector.⁹
45 In addition, emissions factors are meant to apply nationally, whereas emission intensities
46 in fact vary regionally and mitigation is performed locally.^{7,8} To more precisely account for
47 emissions and to inform mitigation efforts, measurement campaigns have been conducted to
48 obtain locally relevant empirical data within individual production basins throughout the
49 United States and Canada.¹⁰⁻¹⁵

50 Many recent research efforts have focused on the Permian Basin because of its sizable
51 share of U.S. O&G production, comprising 43% of domestic oil and 22% of natural gas
52 produced annually.¹⁶ Two studies on 2018/2019 Permian methane emissions both estimated
53 region-wide O&G methane intensity to be 3.7% of production,^{17,18} exceeding an estimated
54 national average of 2.3% for the full supply chain.⁷ More recent work provided methane
55 intensity estimates in the range of 5-6% in 2018, decreasing to 3-4% in 2020.¹⁹ Aerial mea-
56 surements conducted in 2019-21 coupled with simulated emission sources representing the
57 unmeasured part of the distribution provided a Permian Basin methane intensity estimate
58 of 5.29%.²⁰ With the exception of Ref. 20, these studies leveraged satellite observations for
59 inversion modeling and mass balance calculations, which are useful in benchmarking overall
60 emissions but lack the detection sensitivity or spatial resolution needed to identify individual
61 methane sources and understand their relation to infrastructure elements.

62 To provide a more specific account of emission sources, aerial campaigns conducted by
63 Carbon Mapper (CM) and a separate one reported by Chen et al. performed emissions
64 measurements in the Permian Basin using solar infrared imaging spectrometers. The first
65 CM campaign²¹ took place in 2019 and covered 55,000 km² in the Midland and Delaware
66 sub-basins located in Texas and New Mexico. Emission sources were localized and attributed
67 to individual facilities. Repeated sampling of the same sources was used to evaluate emis-
68 sion intermittency. Highly intermittent sources (0-25% observed persistence of facility-sized
69 sources) were responsible for 48% of all point source emissions in the sample. Further cam-

70 paigns were run in the Permian Basin in 2020-21²² in a spatial domain partially overlapping
71 with the 2019 campaign, with otherwise similar collection parameters.

72 The study by Chen et al.²³ was focused on the New Mexico Permian and encompassed
73 over 90% of wellheads in that region. Chen et al. compared the measured emission rate
74 distribution from their study to CM 2019 in an overlapping spatial region and found that
75 the CM 2019 campaign detected progressively fewer emission sources at rates below roughly
76 300 kg/h, while their own study observed similarly reduced detections below 100-150 kg/h.
77 Though the decline in detected emission sources suggests that the CM 2019 data under-
78 represent the actual emission sources present below the detection sensitivity, the heavy tail
79 portion of the data set can still valuably inform models of the emission rate distribution.

80 For the present work, CM data are combined and compared to compiled survey data from
81 Bridger Photonics' first generation Gas Mapping LiDAR (GML) sensor. Emission rates in
82 the range of 3-300 kg/h, which are underrepresented in the CM campaigns, are detected
83 by GML at their true frequency. In a complementary manner, the CM campaign data sets
84 offer extensive sampling of large-rate but infrequently-emitting sources. Detection data from
85 CM and GML campaigns are joined to obtain a comprehensive view of the emission rate
86 distribution in the survey region. Emission sources at rates observable by GML but not by
87 CM are seen to contribute most of the total rate for the whole distribution.

88 **Methods**

89 The Bridger Photonics Gas Mapping LiDARTM (GML) instrument is an aircraft-mounted
90 remote sensing device that maps methane concentration with coaligned dual LiDAR measure-
91 ments, geospatial data through a Global Navigation Satellite System, and aerial photography
92 to show plume shape, identify the source of the emission, and quantify the emission rate.
93 Coaligned range-finding and gas absorption lasers are spatially scanned in a conical pat-
94 tern below the aircraft. Return signal originating from ground-based backscatter is detected

95 at the sensor. Path-integrated gas concentration is measured using wavelength modulation
96 spectroscopy on the 1651 nm absorption line of methane. Flux rates are found from total
97 methane concentration integrated along the direction perpendicular to the gas flow direction,
98 multiplied by wind speed at the measured plume height. Details of the collection platform
99 have been described previously.²⁴

100 For the surveys used in this paper, the GML instrument was flown at an altitude of
101 206 m above ground level (AGL), with a measured detection sensitivity of 0.41 kg/h per
102 m/s wind speed at 90% probability of detection,²⁵ or nominally 2 kg/h at the average wind
103 speed of 4.9 m/s in Midland, TX.²⁶ Scan parameters are chosen so the distance between
104 LiDAR measurement points on the ground is at maximum 1 m.

105 The CM campaigns in this paper utilize two similar instruments called GAO and AVIRIS-
106 NG based on solar infrared spectroscopic imaging. The CM data offer extensive sampling
107 of the heavy tail of the distribution, but lower detection sensitivity and reduced spatial
108 source resolution compared to GML. The instruments were flown at altitudes of 4.5 km
109 and 8 km²¹ AGL. High flight altitudes like these offer greater coverage rates (land area per
110 time) but lower detection sensitivity. Performance of the CM instrument as a function of
111 altitude has been characterized in controlled releases²⁷ and modeled with a robust Bayesian
112 approach.²⁸ In this paper, campaign flights are grouped together regardless of altitude to
113 assess campaign-specific performance rather than instrument performance in general. The
114 campaign sensitivity is thus an average of the measurement sensitivity at the various condi-
115 tions observed in the campaign, including flight altitude.

116 Whereas the spatial pixel size increases with altitude (CM: 3-8 m for 3-8 km flight alti-
117 tude^{29,30}), it is important to distinguish between pixel size and source resolution, or spatial
118 area over which detected emissions are considered to come from the same source. An emis-
119 sion “source” in this paper means a set of synchronously or asynchronously detected plumes
120 falling into a defined aggregation area, whereas “emitter” means a source smaller than the
121 source resolution of the measurement system, inclusive of processing. In addition to limits

122 imposed by image resolution, CM employs a 150 m diameter aggregation area to define its
123 sources at roughly the size of a typical well pad.

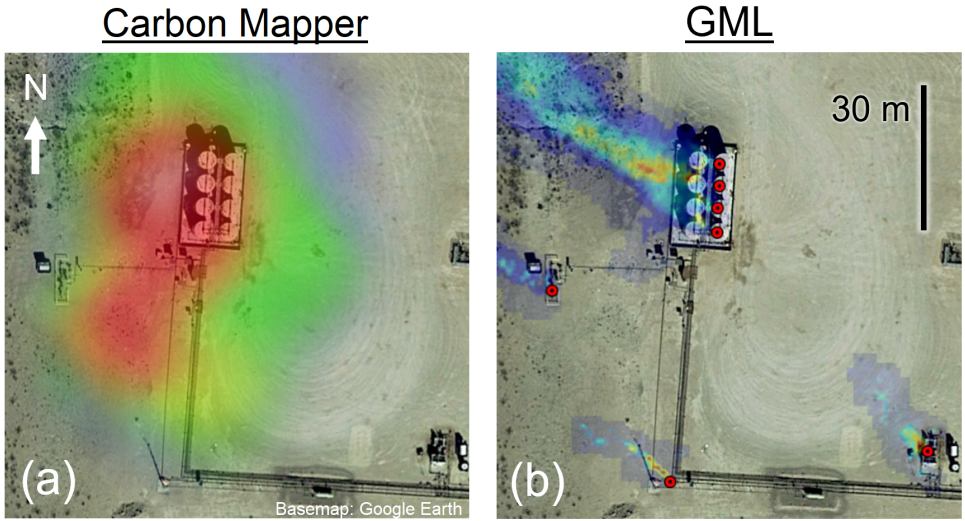


Figure 1: Example facility overlay with methane plume images from (a) CM 2019³¹ and (b) GML scans superimposed on satellite imagery. GML emission locations are marked by red dots. CM 2019 and GML plumes were observed on different dates; the GML image comprises plumes observed on multiple dates. GML identifies unique emission sources at an interval of 4-5 m on a tank battery (upper center).

124 To compare spatial characteristics of GML and CM emission sources, consider the exam-
125 ple plume imagery overlaid on satellite visible imagery shown in Fig. 1. The same facility
126 was observed by both AVIRIS-NG and GML on different dates. In (a), two possible gas
127 concentration peaks are not quite distinguishable, whereas in (b), multiple GML plumes are
128 visible. GML pins mark localization of point sources with a precision of 2 m, which roughly
129 corresponds to the size of typical production equipment. Asynchronous detections at these
130 locations must be localized to within 2 m to count toward the same emission source. For
131 CM, by contrast, all emitters within 150 m are aggregated to the same source. This tends to
132 increase source emission rates because multiple emitters are summed to obtain the reported
133 source rate. For comparison on an equal basis with CM, GML detections can also be aggre-
134 gated to 150 m diameter groups. Cases where spatial aggregation is performed are labeled
135 in the analysis. Further details of GML spatial aggregation are given in Sect. S1.

136 Further parameters, including data selection, measurement time frame, survey area, and

137 scan repetitions were considered in compiling the data sets. Details of the data compilation
138 are given in Sect. S2. Statistical tests on the heavy tail part of the distribution are run to
139 check the assumption that GML and CM data sets sample the same distribution (Sect. S3).

140 **Results**

141 After alignment, we combine the CM and GML data sets to obtain a joint model of the emis-
142 sion rate distribution. We describe results in terms of the detection density and cumulative
143 emission rate distribution. The detection sensitivity of the CM campaign is quantified by
144 comparing the CM detection density to the joint model function, and the share of the cumu-
145 lative emission rate measured by each campaign (scaled by sample size) is also inferred by
146 comparing to the model. We first run the analysis on facility-sized sources (150 m diameter)
147 and then repeat the process on single-emitter sources. This highlights differences between
148 the distributions due to spatial aggregation to facility-sized sources. Results from the CM
149 2020-21 campaigns are also shown.

150 **Facility-sized aggregated (150 m) emission sources**

151 As a first step to joint analysis, we establish a comparison domain supported by both the
152 GML and CM samples. Sensitivity limits associated with each sample determine that emis-
153 sion sources below the full detection limit (FDL) will be detected with diminishing probability
154 as the emission rate decreases. The probability of detection (POD) for a given source can be
155 characterized rigorously as a function of emission rate, wind speed, and flight altitude using
156 controlled release data in a robust Bayesian formalism.²⁸ In this work we take a simple ap-
157 proach to restrict the emission rate domain to rates above the greater FDL of both samples.
158 For the GML and CM samples, the limiting FDL is set by CM. We choose $x_L = 600$ kg/h
159 as the effective FDL of the CM measurements. Accuracy of the declared FDL is not critical
160 as long as it is large enough to avoid introducing observations at significantly reduced POD.

161 All distributions are presented in the single-scan equivalent form described in Sect. S2.4.1,
 162 which can be understood as a distribution on a characteristic emission rate from a single
 163 observation of a given source, subject to detection sensitivity limits of the measurement cam-
 164 paign. The characteristic emission rate approximates an instantaneous source emission rate
 165 that would be observed in a single overflight; spatial aggregation and multiple overflights
 166 effectively sum and average the rate across observations of the source.

167 Before fitting the data to obtain a joint distribution, we run a preliminary check on
 168 the sample distributions using a hypothesis test based on the Kolmogorov-Smirnov (K-S)
 169 statistic. The test is meant to show whether the GML and CM samples differ significantly
 170 above the CM FDL, that is, in the heavy tail portion of the distribution. The outcome of the
 171 test does not oppose the assumption that GML and CM samples follow the same distribution
 172 (Sect. S3).

173 We next create a model of the distribution that represents both samples. The model
 174 density function is taken to follow a generalized lognormal distribution,

$$p(x) \propto \exp\left(-\left|\frac{x-x_0}{b}\right|^m\right), \quad (1)$$

175 where x is the base-10 logarithm of the emission rate and x_0 , b , and m are fit parameters
 176 where $b > 0$ and $m > 0$. Integration over the range $x_L \leq x < \infty$ yields the survival function

$$S(x) = \frac{1 - \operatorname{sgn}(x - x_0)\Gamma\left[\left(|x - x_0|/b\right)^m, \frac{1}{m}\right]}{1 - \operatorname{sgn}(x_L - x_0)\Gamma\left[\left(|x_L - x_0|/b\right)^m, \frac{1}{m}\right]}, \quad (2)$$

177 which has been adapted to the integration range and direction so that $\lim_{x \rightarrow \infty} S(x) = 0$
 178 and $S(x_L) = 1$. In the special case where $m = 2$, the generalized lognormal distribution is
 179 simply lognormal. In either case, the fit parameters are jointly optimized using maximum
 180 likelihood estimation (MLE). With the joint likelihood function given in Sect. S4, the samples
 181 can be fit jointly below their respective FDLs. A nominal value of $x_L = 3$ kg/h is chosen
 182 for the GML FDL for equipment-sized sources, consistent with the sensitivity of 2 kg/h

183 (90% POD) mentioned in the Methods section. For facility-sized sources, this is increased to
184 $x_L = 10$ kg/h to avoid underestimating the FDL, since spatial aggregation increases source
185 emission rates. The MLE fitting process accounts for differences in sample size so the source
186 densities are compared without requiring normalization based on survey size or number of
187 overflights.

188 Several candidate fits are considered. The joint fit is compared to single-sample fits
189 using lognormal and generalized lognormal forms for the density function. The purpose is to
190 confirm that the joint fit better represents the two samples and to choose a model function
191 that more accurately represents the two samples, particularly with respect to the “tailedness”
192 of the distribution determined by m in Eq. 1. After obtaining fit parameters, the candidate
193 models are assessed for relative likelihood of information loss using the Akaike information
194 criterion (AIC).³² The AIC comparison shows that the joint lognormal fit is optimal for 150 m
195 emission sources (GML with CM 2019 or CM 2020-21), whereas a generalized lognormal
196 model is preferred for equipment-sized emitters (GML with CM 2019; $m = 1.619$). Details
197 of the AIC analysis are given in Sect. S5. Best fit parameters for all candidate models
198 are shown in Table S1. Those from the optimal candidate provide the current best known
199 representation of the distribution based on the GML and CM data.

200 With fit parameters obtained from the joint likelihood analysis, the resulting density
201 function is shown in Fig. 2. Survey detections are binned by emission rate and the entire
202 sample is scaled to a reference total of 1000 detected sources above the CM FDL. Error bars
203 are placed at $\pm 2p_{\text{bin}}/\sqrt{n_{\text{bin}}}$, where p_{bin} is the density value of the bin and n_{bin} is the count
204 of emission sources in the bin. Confidence bounds for the model fit are calculated using
205 the likelihood ratio (LR) method at 5% rejection. The bounds consist of the most extreme
206 value of the distribution function at every emission rate among the locus of solutions on the
207 rejection contour. Fit agreement and scale factors are described in Sect. S6.

208 Though the three traces in Fig. 2 (CM 2019, GML, and model) agree above the CM FDL,
209 CM detection density diminishes rapidly at emission rates below the full detection limit. By

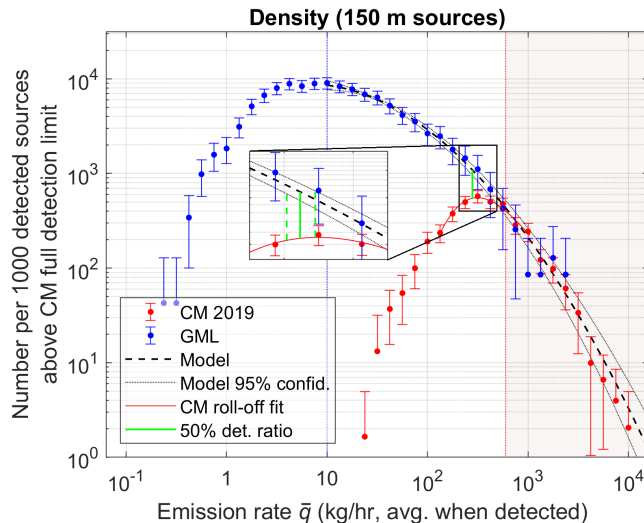


Figure 2: CM 2019 and GML emission source density as a function of emission rate, where sources have a 150 m diameter aggregation area. Zoomed in view near the sensitivity limit (inset) shows the 50% detection ratio with respect to model function and its confidence bounds. Model function follows Eq. 1 with $m = 2$, $x_0 = 0.797$, $b = 1.140$.

210 comparing the model to the CM detection distribution around the roll-off region using an
 211 error-weighted cubic polynomial fit of the binned data, the 50% detection ratio is placed at
 212 280 [256, 309] kg/h, where the confidence interval (CI) is found by comparing to the 95% CI
 213 of the model fit, neglecting error in the cubic polynomial estimating the roll-off, as shown in
 214 the inset. This resulting sensitivity is considerably higher than the detection limit quoted by
 215 Cusworth et al. at 10-20 kg/h but is consistent with a previous estimate of the sensitivity
 216 in the range 100-300 kg/h.²³ Without compensation, reduced POD leads to a significant
 217 underrepresentation of emission sources below the sensitivity. For example, comparing the
 218 detection density of CM to GML binned data at 100 kg/h shows that emission sources at this
 219 rate are in fact 14 times more common than the CM data would suggest. The CM campaign
 220 can be expected to underestimate both the fraction of facilities with emissions and the total
 221 emission rate for the facilities surveyed because of its sensitivity limit. In fact, the CM 2019
 222 campaign detected emissions at 1.48% of well sites whereas the GML detected emissions at
 223 38.3% of facilities (see Sect. S1).

224 Controlled release measurements could confirm the sensitivity findings reported in this

225 work. Alignment of flight altitude and on-the-ground properties of the sources observed in
226 the campaign, such as ground cover, would need to be considered. Moreover, the effects of
227 spatial aggregation would need to be accounted for to achieve the same measure of “realized”
228 detection sensitivity for the source definition used in the campaign.

229 **Equipment-sized emission sources**

230 Although the above results for 150 m sources show that emission rates less than ~ 300 kg/h
231 are underrepresented in the CM detection density, a further increase in density of lower
232 emission rates occurs when sources are resolved to equipment size scale (~ 2 m). Facility-
233 aggregated emission rates tend to be higher than equipment rates because co-located emitters
234 on a site count toward the same emission source. Equipment-sized source resolution tends to
235 be more practical for both bottom-up emissions modeling and identification for leak detection
236 and repair.

237 To obtain the equipment-scale emission distribution, GML detections are considered in
238 their native resolution (~ 2 m) and not aggregated to 150 m. Since CM sources are not
239 reported at finer resolution, we instead manually filter them based on associated plume
240 imagery³¹ to include only sources with a single point emission (see details in Sect. S7).

241 The detection density for equipment-sized sources is shown in Fig. 3 with facility (150 m)
242 detection density traces from Fig. 2 reproduced for comparison. At mid-range emission rates
243 (~ 3 -300 kg/h), density is significantly higher for equipment-sized sources than for facility-
244 sized sources. For example, comparing the two GML traces at 10 kg/h shows that equipment
245 sources at this emission rate are observed eight times more frequently than 150 m ones.
246 Single-emitter filters applied to the CM data set do not appear to distort the distribution
247 appreciably above the CM FDL. For other CM distributions (CM 2019 single emitters, CM
248 2020-21 150 m sources) CM detection sensitivity is assessed in a similar manner (Sect. S8).

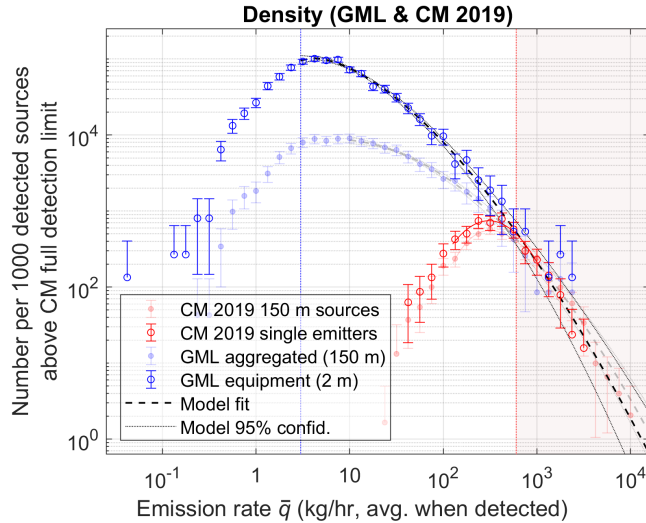


Figure 3: Density of detected emission sources with both 150 m aggregation diameter and no aggregation (single emitters), plotted together on the same axes. Model function for single emitter distribution follows Eq. 1 with $m = 1.619$, $x_0 = 0.629$, $b = 0.770$.

249 Cumulative emission rate distribution

250 The density function weighted by emission rate can be integrated to yield the cumulative
 251 emission rate distributions shown in Fig. 4. For measured samples, the cumulative sum is
 252 given by Eq. S2 (single-scan equivalent). Results for CM 2019 (150 m sources and single
 253 emitter sources) are shown in this section; cumulative emission rates for CM 2020-21 are
 254 shown in Sect. S9.

255 Expected error due to sample variation is shown in the plot. Error bounds show the 2.5
 256 and 97.5 percentiles of the sample variation for an equivalently sized data set with the same
 257 number of detections above the corresponding FDL, assuming the best-fit model represents
 258 the “true” distribution. They are found by running a Monte Carlo simulation of random
 259 sets of detections drawn from the model density function (see Sect. S10). Sample error from
 260 sources with emission rates below each FDL is neglected, as is instrument quantification error.
 261 Sample variation in the heavy tail is responsible for much of the sample error along the entire
 262 trace. The relatively infrequent emitters in this emission rate range have a disproportionately
 263 large impact on cumulative emissions.

264 The fractional total emission rate measured in each survey can be found by comparing

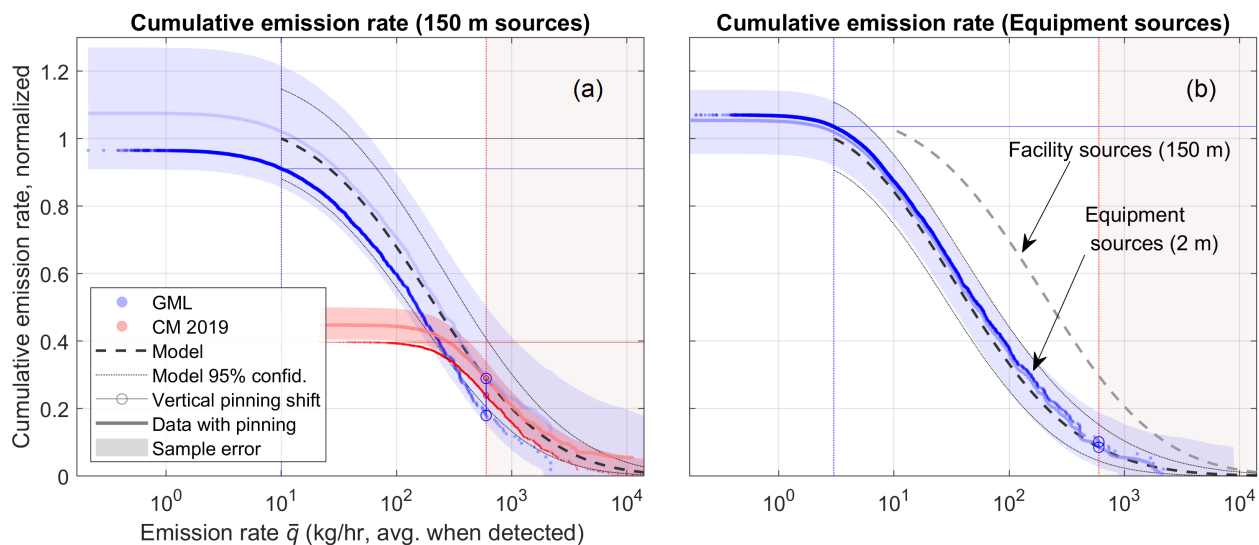


Figure 4: Cumulative emission rate distribution of (a) 150 m aggregated emission sources (GML and CM 2019) and (b) equipment sources (GML only). Error bounds (shaded regions) describe predicted sample variation. Model distribution from (a) is reproduced in (b) for comparison. All traces are normalized to equivalent campaign scale (spatial area, number of overflights). In (b), the “facility sources” model function is multiplied by the ratio of the quantity $c(10 \text{ kg/h})/c(0)$ for each sample (see Eq. S2, single-scan equivalent) so that cumulative emissions are comparable between traces. Vertically shifted copies of survey data pinned to the value of the model distribution at the CM FDL guide the eye to suggest the shape of the measured distribution supposing sample error above the CM FDL were suppressed.

265 the cumulative emission rate of each sample to the model function. Since the model does
266 not extend below the GML FDL due to the onset of sub-unity POD, we read cumulative
267 rates at this threshold (10 kg/h for 150 m sources, 3 kg/h for equipment sources) rather
268 than at the top of the curves. Comparing each sample to the model and its 95% CI, GML
269 is estimated to have measured 91.1% [79.4%, 103.4%] of the total emission rate predicted by
270 the model for 150 m sources above 10 kg/h and CM is estimated to have measured 39.7%
271 [34.6%, 45.1%]. For equipment-scale sources, GML measured 103.6% [93.5%, 114.2%] of the
272 cumulative emission rate above 3 kg/h. Measured fractions in excess of 100% signify the
273 measured distribution exceeding the joint model distribution over a portion of the model CI,
274 which can be expected from finite sample size and model fit uncertainty.

275 In the case of 150 m sources, both CM 2019 and GML appear to have undermeasured
276 the heavy tail compared to the model distribution. This can be seen from the cumulative
277 emission rates falling below the model in Fig. 4a. In fact, the measured distribution lies
278 outside the estimated sample error, which could be explained by a departure from lognormal
279 behavior above source emission rates of $10^{3.4}$ kg/h (see Sect. S6). The lower than expected
280 cumulative emission rate is consistent with a sharp drop in the measured CM 2019 survival
281 function at $10^{3.5}$ kg/h as shown in the inset of Fig. S8. By comparing the CM measured
282 distribution to the model function in Fig. 4a, it can be seen that most of the CM-model error
283 is indeed inherited from emission sources above $10^{3.5}$ kg/h. This suggests that either the
284 lognormal distribution does not describe the true emission rate distribution above $10^{3.5}$ kg/h
285 despite working well below it, or that the anomaly in CM 2019 data at $10^{3.5}$ kg/h might be
286 explained by undersampling, systematic error, or failure of invariance assumptions mentioned
287 in Sect. S2. In any case, further measurements of the distribution would more clearly resolve
288 this part of the heavy tail and explain the discrepancy.

289 Without POD compensation, missed emissions below the detection sensitivity of a given
290 campaign raise the apparent threshold responsible for a given share of the total emission rate.
291 For example, according to CM 2019 data alone, 90% of the total emission rate is contributed

292 by facilities with rates above 249 kg/h, whereas in the GML distribution the 90% facility
293 rate is 16.9 kg/h. In fact, the true 90% threshold will be even lower because emission sources
294 below GML detection sensitivity are underrepresented in the GML data set.

295 In addition, spatial aggregation of emission sources shifts (and reshapes) the entire curve
296 to larger emission rates. Comparing the model curves in Fig. 4b shows that the 150 m source
297 curve is shifted to the right of the equipment-level source curve by roughly a factor of 3-5
298 over most of the domain. The 90% threshold for the total detected emission rate shifts from
299 16.9 kg/h to 6.0 kg/h on the GML data traces. The shift toward smaller emission rates
300 can be significant, meaning that measured distributions and sensitivity thresholds should be
301 interpreted at the specified spatial aggregation level, and not directly compared at different
302 spatial aggregation levels.

303 Comparing the CM and GML distributions shows that the total methane emission rate
304 from O&G production infrastructure in the survey region is significantly greater than pre-
305 viously reported, with GML measuring 2.3 times that measured in the CM 2019 campaigns
306 (and also 2.3 times that in the CM 2020-21 campaigns; see Sect. S9). If observation of
307 emissions is viewed as an ergodic process, then the cumulative emission rate distributions
308 shown in Figs. 4 and S11 may be seen as representative of total average emission rate for
309 production infrastructure in the survey region. In this case, proportions of the total emission
310 rate from the plots can be compared to measurements of regional methane flux. Based on
311 top-down inversion from Tropospheric Monitoring Instrument (TROPOMI) measurements
312 of regional methane flux, Cusworth et al. estimated that sources measured in the CM 2019
313 survey represent 59% (CM 2020-21: 49%, where fractions for each of the three campaigns
314 are weighted by survey area) of the total methane emission rate in the survey region.²² This
315 estimate is somewhat higher than the proportion of cumulative emission rate measured by
316 CM 2019 compared to GML ($1/2.3 = 43\%$), suggesting that the emission rate inferred from
317 TROPOMI in Ref. 22 underestimates total emissions by 37% (i.e. $59\%/43\% - 1$) if GML
318 and CM campaign data are representative of the same emissions process.

319 In summary, GML detection data extends the measured emission rate distribution for
320 Permian Basin O&G production infrastructure beyond CM sensitivity limits by roughly
321 two orders of magnitude. In joint analysis, intensive sampling of the heavy tail by CM is
322 complemented by GML's higher detection sensitivity. In the region surveyed, facility-sized
323 emission sources with rates below the CM campaign detection sensitivity (280 kg/h at 50%
324 POD) contribute 67% of the total emission rate from sources with rates above 10 kg/h. The
325 density of these sources and their constituent equipment-size emission sources (at rates above
326 3 kg/h), was measured without POD degradation by GML. According to the GML sample
327 without POD correction, 90% of the total cumulative emission rate measured originates from
328 equipment-sized sources with rates larger than 6.0 kg/h. This threshold rate would become
329 even smaller if sources below 3 kg/h were measured at their true density rather than at POD
330 < 1 . Emissions monitoring campaigns require both high sensitivity and intensive sampling
331 to accurately capture the emissions distribution.

332 **Acknowledgement**

333 Bridger Photonics thanks the Advanced Research Program Agency – Energy (ARPA-E)
334 MONITOR program and the Montana Board of Research and Commercialization Technology
335 (MBRCT) for supporting the original development of GML hardware and analytics.

336 **Supporting Information Available**

337 The following data files will be available on the journal website at time of publication:

- 338 • GML methane emission source detection data at 150 m spatial aggregation (*.csv).
- 339 • GML methane emission source detection data at 2 m spatial aggregation (*.csv).
- 340 • Single-emitter classification of CM 2019 plume images (*.csv).

- Geospatial polygons bounding GML data set (*.kml).

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Supporting Information

Extension of Methane Emission Rate Distribution for Permian Basin Oil and Gas Production Infrastructure by Aerial LiDAR

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19 **S1. Spatial aggregation to 150 m sources**

20 Spatial analysis of GML data is performed by first assigning an emission origin point, or
21 “location,” to each detected plume. Detections observed at different times are associated
22 with the same location if they are co-located within 2 m. Emission locations were spatially
23 aggregated to 150 m sources by a clustering algorithm that iterates through a list of GML
24 locations to build a temporary table of locations within 150 m to all other locations in
25 the current cluster. After all unclustered locations in the list have been compared to the
26 temporary cluster (sequentially, in fixed arbitrary order), those in the temporary cluster are
27 removed from the waiting list. New clusters are formed in this way until no locations remain
28 in the waiting list.

29 GML detections can also be aggregated to “facilities” described by polygons enclosing site
30 assets. Facility polygons represent the boundaries around actual groups of surface infras-
31 tructure and are usually defined by the facility pad footprint. Polygons can be provided by
32 operators based on site data or generated from aerial photography, in which case the polygon
33 is drawn either manually or by an artificial intelligence model. A mix of AI-generated and
34 manually defined polygons was used in the data set in this work. A polygon is defined for
35 every facility on a GML flight path regardless of whether an emission is actually detected.

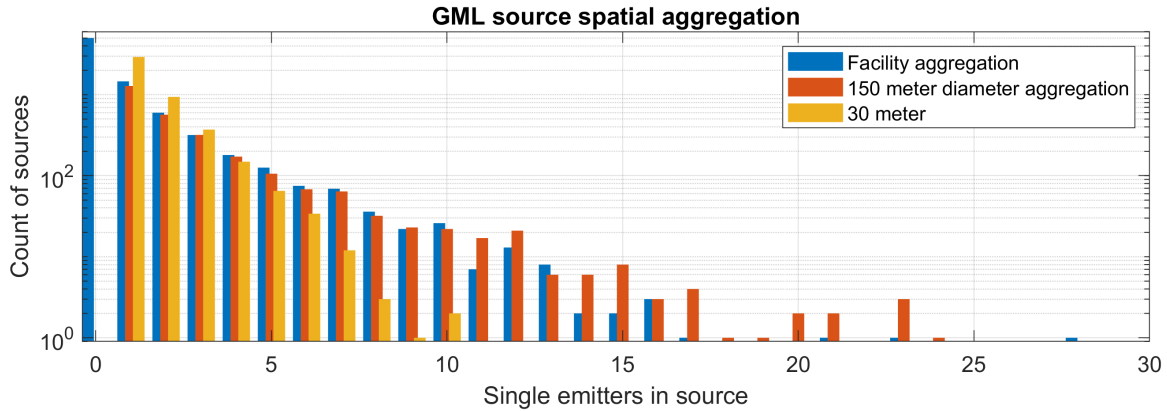


Figure S1: Histogram counting spatially aggregated sources by number of emitters in aggregation area.

36 Comparing the number of GML locations in each facility to the number per 150 m di-
 37 ameter source shows a near correspondence between the two aggregation styles (Fig. S1),
 38 supporting the use of 150 m aggregation to represent facility-sized sources. A smaller ag-
 39 gregation area (30 meters) displays a steeper roll-off in number of detection locations per
 40 source. For GML, the proportion of facilities with at least one detection was found to be
 41 38.3% (or 32.9% when considering only first overflights in the 15-minute scan window de-
 42 scribed in Sect. S2.4.2), much higher than the reported 1.48% rate^{S1} of well sites in the
 43 CM 2019 campaign. This difference may be explained by differences in detection sensitivity
 44 described in the analysis.

45 S2. Data preparation and alignment

46 Before jointly analyzing emission detection data from diverse sources, several aspects of data
 47 collection and data set composition must be considered. In this section we aim to address
 48 some important aspects of data alignment, specifically (1) compilation of GML survey data
 49 and types of emitters included, (2) spatial overlap of CM and GML surveys, (3) temporal
 50 overlap of CM and GML surveys, and (4) influence of scan repetitions on reported emission
 51 rates. In preparation for this work, efforts were made to directly align the data sets as
 52 much as possible. Where data sets do not align, limited assumptions of spatial or temporal

53 invariance are used to set up the analysis. These are pointed out where relevant in the details
54 of the data preparation process below.

55 **S2.1 Sample composition and emitter types included**

56 The GML sample is compiled from sets of anonymized survey data collected under contracts
57 for client O&G operators. Survey sets for compilation were chosen for geographical and tem-
58 poral overlap with CM data without considering analysis results. Whereas CM campaigns
59 blanketed entire geographic areas, GML surveys were targeted to client facilities. Clients
60 were given advance notice of when scans would occur (typ. accuracy ± 2 -3 days). The sample
61 is comprised of scans of sites belonging to 28 individual operators.

62 Sites included in the GML sample were in the O&G production sector and do not include
63 midstream/distribution infrastructure. Types of infrastructure included in the GML sample
64 consist of wells, separators, tanks, compressors, flares, vapor recovery units, generators, and
65 facility piping. Equipment types were not identified at every facility scanned, though this
66 capability is currently under development. In this work, CM data have been filtered to
67 exclude detections from O&G pipelines unless otherwise marked. In the CM 2019 data
68 set,^{S2} measurements with all *source type* tags were included except for “pipeline” and “NA.”
69 For CM 2020-21,^{S3} the accepted tags were “tank,” “well,” “compressor,” “processing,” and
70 “refinery.” Exclusion of pipelines seems to have a negligible effect on the shape of the CM
71 distribution, as shown in Sect. S11.

72 False positive detections can occur in GML detection data, but practically only near the
73 GML detection limit. For emission rates more than a factor of two above the GML detection
74 limit the likelihood of false positives is vanishingly small. GML uses a physics model of
75 the LiDAR measurement noise processes (shot noise, photodetector noise, speckle noise) to
76 estimate the noise on each methane concentration LiDAR measurement based on received
77 light levels. During processing of GML data the signal to noise ratio for each measurement
78 is used in a statistical algorithm to detect regions of elevated methane concentration. The

79 detected regions of elevated concentration are then submitted to emitter analysis, which only
80 assigns an emission if a hot spot in both detection confidence and concentration is detected
81 at the upwind end of the detected plume.

82 S2.2 Spatial overlap of CM and GML samples

83 GML and CM 2019 samples were restricted to the GAO coverage polygons in the Delaware
84 and Midland Basins provided in Ref. S1. Geography is shown in Fig. S2. Restriction to the
85 GAO polygons excludes 29 out of 1756 detected facility sources in CM 2019. GML detection
86 locations occupy a subregion of both GAO polygons. We assume that the complementary
87 area in the GAO polygon does not significantly affect the emission rate distribution.

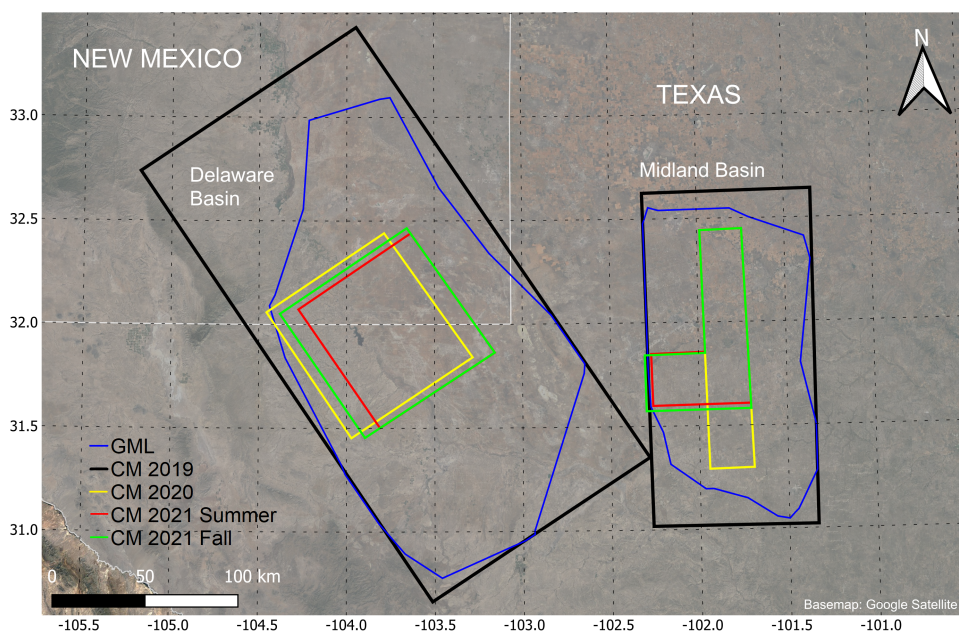


Figure S2: Geospatial domain of GML and CM samples. GML polygons contain all detected emission locations and include a random buffer so vertices do not correspond to detections. CM 2019 polygons reported in Ref. S1 represent areas surveyed at least once by GAO. CM 2020-21 polygons are approximate based on detection coordinates; details reported in Ref. S4.

88 Though the CM 2020-21 coverage areas^{S4} intersect the 2019 GAO polygons, they do not
89 cover the entire area of the 2019 polygons and contain a small amount of additional area
90 outside them. We do not explicitly align GML and CM 2020-21 survey areas in this work,

91 but rather assume that the emission rate distribution is roughly spatially invariant among
 92 these areas. We use the same GML data set for joint analysis with CM 2019 and CM 2020-
 93 21. We apply no geographic filters to CM 2020-21 other than to select the campaigns that
 94 took place in the Permian Basin (*source ID* markers “F,” “E,” and “J” in the published data
 95 set^{S3}).

96 S2.3 Temporal overlap of CM and GML samples

97 A timeline of plume detections in the GML and CM measurement campaigns is shown in
 98 Fig. S3. GML scans were performed between Jan 2020 and Feb 2022, whereas the CM
 99 campaigns took place in Sept-Nov 2019 (CM 2019) and Jul 2020-Nov 2021 (CM 2020-21).
 100 Analysis in the Results section assumes stationarity in the shape of the emission rate distri-
 101 bution with time (i.e. does not change with choice of time origin). However, stationarity of
 102 the scale of the distribution is not required. The joint analysis computes separate likelihoods
 103 for each data set and scales the density and cumulative emissions traces to the total density
 104 above the CM full detection limit.

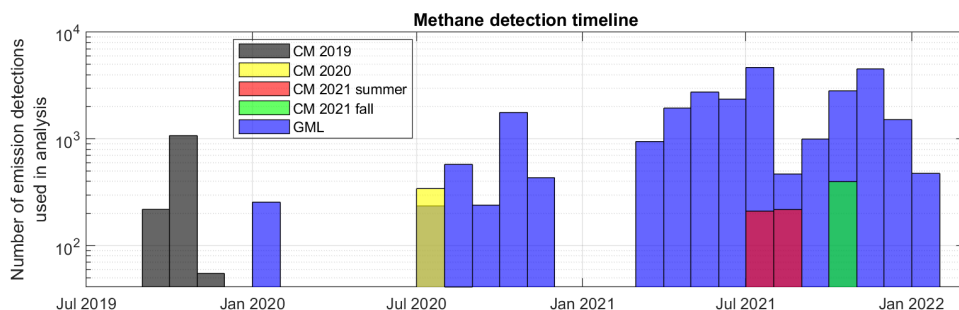


Figure S3: Plume detection counts versus time for CM and GML measurement campaigns.

105 S2.4 Scan repetitions

106 CM and GML campaigns were conducted with different approaches to scan repetitions.
 107 Number of scans per emission source is shown in Fig. S4. Fewer scans were performed per
 108 150 m source with GML (median: 2 scans) in comparison to CM (2019 median: 6 scans,

109 2020-21: 4 scans). In CM campaigns, repeated scans over a given source were performed
 110 independently of previous results. No minimum number of scans was used to filter the data
 111 sets for this work. In GML surveys, repeated scans were performed only on locations where
 112 an emission was detected in the first scan. This means that emissions measured by GML
 113 were effectively found in just one scan, and repeat measurements were not independent.
 114 Most emission sources in the CM campaign had multiple opportunities to be detected, so a
 115 greater fraction will have been detected. To address these issues, we describe two solutions
 116 below: how to express the distributions in a form that enables direct comparison, and how
 117 GML observations are handled considering overflight repetitions and conditionality.

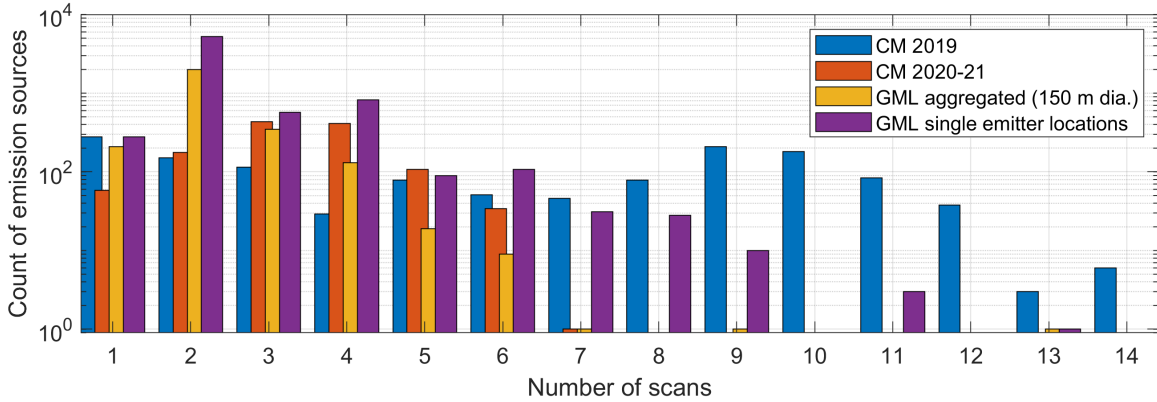


Figure S4: Histogram counting emission sources by number of scan repetitions.

118 S2.4.1 Transformation to single-scan equivalent

119 GML and CM distributions are expressed in a “single-scan equivalent” form for alignment.
 120 We adopt the notation and terminology of Cusworth et al.^{S1} for the persistence-adjusted
 121 emission rate $q = f\bar{q}$, where f is the observed persistence $f = M/N$, with M as the number
 122 of non-zero unique detections and N as the number of scans, and \bar{q} is the mean of all non-zero
 123 measured emission rates

$$\bar{q} = \frac{1}{M} \sum_{i=1}^M q_i, \quad (\text{S1})$$

124 where q_i is a non-zero unique emission rate measurement. For a given source emitting
 125 intermittently at a single rate, \bar{q} should be consistent across number of measurement scans,

126 which aids in comparing measurements with different numbers of scans.

127 To plot emission density and cumulative emission rate on a \bar{q} axis requires further adjust-
 128 ment using the persistence. Consider a point on the detection density function, or rather,
 129 a single point in a discrete series representing detection frequency, as shown in Fig. S5a.
 130 The persistence adjusted detection frequency (blue), where the emission rate is q , is acces-
 131 sible only from a repetitive sample set and not from a single scan, since the persistence f is
 132 needed to obtain q . The detection frequency can be replotted at \bar{q} , which effectively removes
 133 the persistence from the emission rate. This results in an effectively higher emission rate
 134 (green dashed) which overrepresents the density at this emission rate. To obtain a correctly
 135 weighted frequency for summation, or density for integration, the prevalence of the source
 136 must be reduced by the persistence (red). Where the density or frequency function contains
 137 many points, the remapping of q to \bar{q} and the persistence weighting f applies to all points
 138 on the curve.

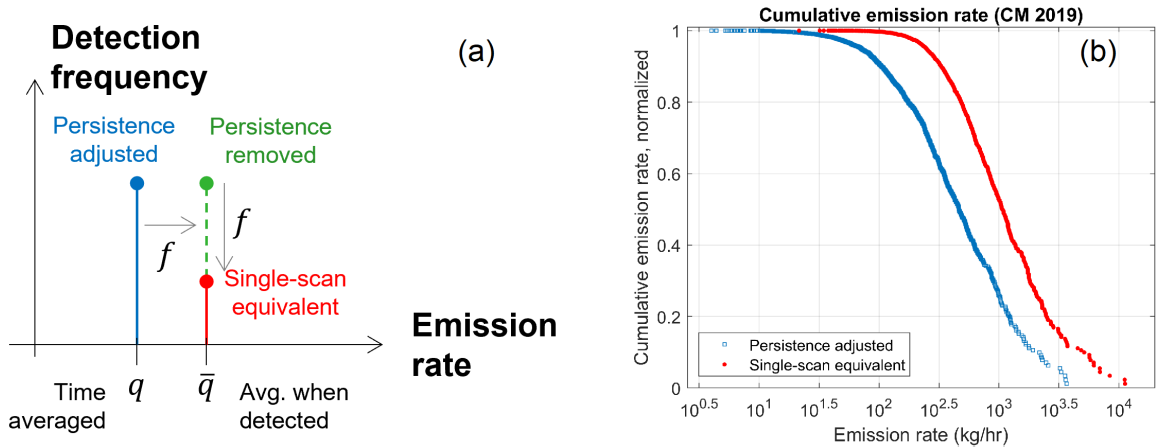


Figure S5: Transformation of emission distribution using the observed persistence f , for (a) a discrete point in a measurement series and (b) the cumulative emission rate distribution (CM 2019, 150 m sources). In both plots, emission rate on the x -axis means either q or \bar{q} as indicated. Distributions shown in (b) follow Eq. S2, reaching the same total emission rate without independent normalization.

139 Next, consider the implications for the cumulative emission rate distribution. For a finite

140 set of measurements, the cumulative emission rate is computed using the discrete sum

$$c(x) = \frac{1}{\sum q} \cdot \begin{cases} \sum_{q \geq x} q & \text{(persistence adjusted)} \\ \sum_{\bar{q} \geq x} f\bar{q} & \text{(single-scan equivalent)} \end{cases}, \quad (\text{S2})$$

141 where x is the emission rate. The result of Eq. S2 applied to the CM 2019 sample used
 142 for analysis is shown in Fig. S5b. Different forms of the sum yield the same total emission
 143 rate since each source contributes the same argument to the sum ($q = f\bar{q}$). Effectively the
 144 contributions have been reordered and replotted on the x -axis according to the correspond-
 145 ing value of q or \bar{q} . As a result, the single-scan equivalent distribution is a reshaped and
 146 horizontally shifted version of the persistence-adjusted distribution.

147 Although this treatment conveniently transforms distributions for comparison regardless
 148 of number of scans, the resulting distributions are not exact. Noting that the observed
 149 persistence f is an observation of an event with probability equal to the actual source per-
 150 sistence times the probability of detection (POD), some distortion of the distribution can be
 151 expected where sources with $\text{POD} < 1$ from below the FDL are shifted above it. Whereas
 152 this affects multi-scan data sets like CM, single-scan data sets (which GML approximates)
 153 would not be affected.

154 S2.4.2 GML detections

155 GML observations of a given emission source come at three different levels: overflight, loca-
 156 tion scan, and aggregated source scan. A location scan is comprised of one or more aerial
 157 passes (“overflights”) of an emission source seen at GML source resolution (~ 2 m). The
 158 first measurement out of all overflights within a 15-minute time window, inclusive of mea-
 159 surements with zero and non-zero emission rates, is selected to represent the emission rate
 160 for the scan. Scan measurements are then converted to a persistence-adjusted rate q and
 161 associated observed persistence f for the location. These are used to find the “average when

162 detected” rate \bar{q} in the same way as for the CM data, using Eq. S1.

163 For spatially aggregated sources (150 m), emission rates are found by adding the persis-
164 tence adjusted emission rates for each location in the source, and dividing by a composite
165 persistence value for the source,

$$f_{\text{agg}} = \frac{\sum_i q_{\text{loc},i} f_{\text{loc},i}}{\sum_i q_{\text{loc},i}}, \quad (\text{S3})$$

166 where $q_{\text{loc},i}$ is the persistence adjusted emission rate and $f_{\text{loc},i}$ is the observed persistence,
167 where both correspond to the i^{th} location in the source. In other words, the aggregated
168 source persistence is an average of the observed location persistence values, weighted by the
169 persistence-adjusted location emission rates. The average emission rate for the source, when
170 detected, is then calculated as $\bar{q} = \sum_i q_{\text{loc},i} / f_{\text{agg}}$.

171 S3. Statistical test on CM and GML distributions

172 A two-sample Kolmogorov-Smirnov (K-S) test^{S5} is used to check for differences between
173 the tails of the GML and CM samples. This is a non-parametric test with a standard null
174 hypothesis (no statistically significant difference between the samples). Here the test is
175 performed on the survival function for a single-scan equivalent sample,

$$S(x) = \frac{\sum_{\bar{q} \geq x}^{\infty} f(\bar{q})}{\sum_{\bar{q} = x_L}^{\infty} f(\bar{q})}, \quad (\text{S4})$$

176 where x is the emission rate and $x \geq x_L$, with x_L as the lower bound of a range of interest,
177 and f is the observed persistence for a given measurement with emission rate \bar{q} . The sum is
178 represented as a stepwise function for the K-S test. As mentioned in the Results section, we
179 choose $x_L = 600 \text{ kg h}^{-1}$ as the effective full detection limit of CM measurements.

180 The survival function of the GML and CM 2019 samples are plotted in Fig. S6. The

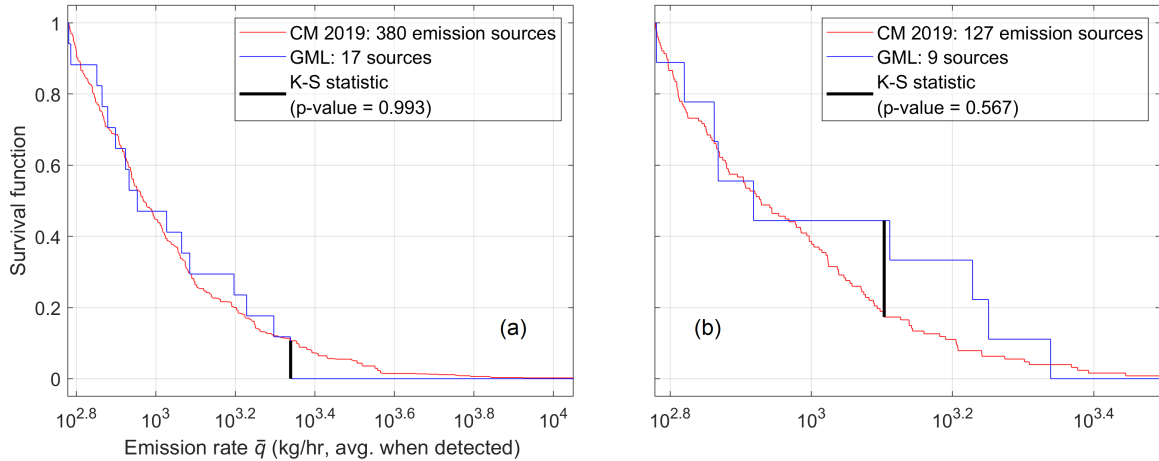


Figure S6: Survival function of single-scan equivalent CM 2019 and GML source detections above 600 kg h^{-1} where sources are defined by a (a) 150 m aggregation diameter and (b) single emitter. Kolmogorov-Smirnov (K-S) statistic and associated p -value are shown.

181 Kolmogorov-Smirnov (K-S) statistic shows the maximum absolute residual between the two
 182 sample distributions. The number of measurements in the GML sample is small in this
 183 range. In both cases the associated p -values are high and do not indicate rejection of the
 184 null hypothesis.

185 The K-S test is also used to check the measured emission distribution for the CM 2020-21
 186 campaigns against GML. In 2020-21, CM conducted three campaigns around the Midland
 187 and Delaware sub-basins (2020 summer, 2021 summer, 2021 fall). Each campaign is smaller
 188 than CM 2019 in number of detections and number of overflights (see Fig. S1). Spatial
 189 overlap among these campaigns is partial; overlap with CM 2019 is also partial.^{S4} For the
 190 analysis in this paper, no controls for spatial overlap were used, under the assumption that
 191 the shape of the emission distribution is spatially invariant over the CM 2019 and CM 2020-21
 192 domains. The GML data set is unchanged whether comparing to CM 2019 or CM 2020-21.

193 Fig. S7 shows the distributions and K-S test results. For the CM 2020 campaign, a devia-
 194 tion around $10^{2.9} \text{ kg h}^{-1}$ is responsible for a slightly low p -value of 0.166. When grouped with
 195 the other campaign data, however, the CM 2020 deviation no longer causes the maximum
 196 difference in sample distributions (comparing Fig. S7a and Fig. S7d). For analysis in the
 197 rest of this paper, all three CM 2020-21 campaigns were merged into one data set as shown

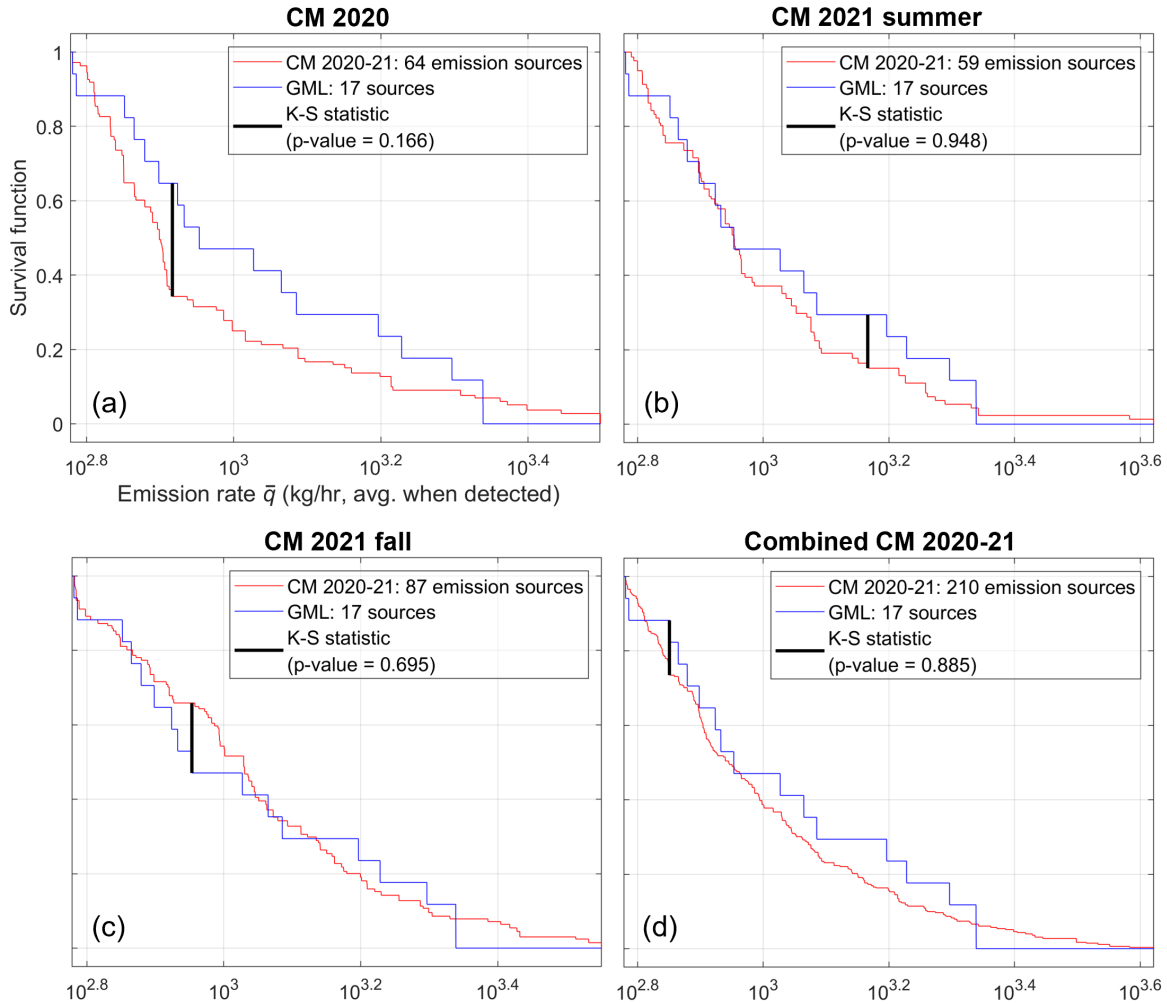


Figure S7: Survival function of CM 2020-21 and GML single-scan equivalent source detections (150 m aggregation diameter) for emission rates above 600 kg h^{-1} for campaigns taking place in (a) 2020 summer, (b) 2021 summer, (c) 2021 fall, and (d) all 2020-21 campaigns together. Kolmogorov-Smirnov (K-S) statistic and associated p -value for each case are indicated.

198 in Fig. S7d.

199 S4. Likelihood function

200 The likelihood function $L(\theta)$, where θ is the vector of fit parameters, is based on the density
 201 function from Eq. 1 normalized to the integration range $x_{L,i} \leq x < \infty$,

$$p_i(x) = \frac{m}{bd_i\Gamma(1/m)} \exp\left(-\left|\frac{(x-x_0)}{b}\right|^m\right), \quad (\text{S5})$$

202 where $x_{L,i}$ is the FDL, $d_i = 1 - \text{sgn}(x_{L,i} - x_0)\Gamma[|(x_{L,i} - x_0)/b|^m, 1/m]$, and the subscript i
 203 has been added to denote the sample (i.e. GML or CM). Using the standard form for the
 204 likelihood function, $L(\theta) = \prod_{j=1}^n p(X_j|\theta)$, where X_j are the observed emission rates in the
 205 sample, we obtain the log likelihood function for the i^{th} sample,

$$LL_i(\theta) = \sum_{j=1}^n \left[\ln(f_{i,j}/\bar{f}_i) + \ln\left(\frac{m}{bd_i\Gamma(1/m)}\right) - \left|\frac{X_{i,j} - x_0}{b}\right|^m \right], \quad (\text{S6})$$

206 where \bar{f}_i is the mean persistence of the sample in the limited domain ($x \geq x_{L,i}$). The term
 207 $f_{i,j}/\bar{f}_i$ performs the persistence weighting (vertical part of the density transformation) from
 208 persistence-adjusted to single-scan equivalent described in Sect. S2.4.1 while maintaining the
 209 property that $\int_{x_L}^{\infty} p(x)dx = 1$.

210 For joint fits, because the samples are independent, we take the product of likelihoods to
 211 obtain the joint log likelihood function

$$LL(\theta) = \sum_i LL_i(\theta), \quad (\text{S7})$$

212 where $i = 1, 2$.

213 S5. Akaike information criterion (AIC) analysis

214 AIC analysis was performed on lognormal and generalized lognormal fits to single-sample
215 data sets and joint data sets. Results are shown in Table S1. As seen by the location of
216 AIC minima under single-sample fits where the fit is tested with the same sample in the “1”
217 rows, joint fits do not provide the best representation of each single sample. They instead
218 reduce the joint likelihood of the two independent samples taken together, as seen by the
219 location of joint relative AIC minima under joint fit columns in “2” rows. Models fit to the
220 CM distribution alone tend to have very low values of joint relative likelihood of information
221 loss (see “3” rows), suggesting that models fit to the CM samples alone are not predictive
222 of the entire distribution through the range over which GML is assessed ($\geq 3 \text{ kg h}^{-1}$ or
223 $\geq 10 \text{ kg h}^{-1}$). In addition to the lognormal and generalized lognormal functions shown,
224 log-logistic (with an extra parameter for horizontal shift), Fréchet, Gumbel, and power law
225 model functions were tested but were not optimal in any case.

Table S1: Akaike information criterion (AIC) analysis for different data sets: (a) CM 2019 survey data, 150 m emission sources; (b) CM 2019 survey data, single emitter (or equipment-sized) emission sources; (c) CM 2020-21 survey data, 150 m emission sources. Each table displays (1) AIC values obtained from the likelihood function (parameter values x_0, b, m), (2) relative AIC values $AIC_{rel,j} = AIC_{i,j} - AIC_{min,j}$, where (i, j) signify (sample, fit), and (3) relative joint likelihood of information loss minimization, where the joint likelihood is taken as the product of likelihoods corresponding to each sample, i.e. $\exp[-\sum_j AIC_{rel,j}/2]$.

CM 2019, 150 m sources	(1)	<table border="1"> <thead> <tr> <th>Fit → ↓ Test</th> <th>CM Lognorm.</th> <th>GML Lognorm.</th> <th>Joint Lognorm.</th> <th>CM Gen.Logn.</th> <th>GML Gen.Logn.</th> <th>Joint Gen.Logn.</th> <th>Min.</th> </tr> </thead> <tbody> <tr> <td>CM</td> <td>54.0</td> <td>57.8</td> <td>55.4</td> <td>55.9</td> <td>75.8</td> <td>57.4</td> <td>54.0</td> </tr> <tr> <td>GML</td> <td>1873</td> <td>1080</td> <td>1081</td> <td>1112</td> <td>1081</td> <td>1083</td> <td>1080</td> </tr> <tr> <td>(Params)</td> <td>(1.840, 0.880)</td> <td>(0.880, 1.088)</td> <td>(0.797, 1.140)</td> <td>(-4.443, 6.057, 6.224)</td> <td>(-0.350, 2.167, 3.232)</td> <td>(0.882, 1.072, 1.929)</td> <td></td> </tr> </tbody> </table>	Fit → ↓ Test	CM Lognorm.	GML Lognorm.	Joint Lognorm.	CM Gen.Logn.	GML Gen.Logn.	Joint Gen.Logn.	Min.	CM	54.0	57.8	55.4	55.9	75.8	57.4	54.0	GML	1873	1080	1081	1112	1081	1083	1080	(Params)	(1.840, 0.880)	(0.880, 1.088)	(0.797, 1.140)	(-4.443, 6.057, 6.224)	(-0.350, 2.167, 3.232)	(0.882, 1.072, 1.929)	
	Fit → ↓ Test	CM Lognorm.	GML Lognorm.	Joint Lognorm.	CM Gen.Logn.	GML Gen.Logn.	Joint Gen.Logn.	Min.																										
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	(2)	<table border="1"> <tbody> <tr> <td>CM</td> <td>0</td> <td>3.8</td> <td>1.4</td> <td>1.9</td> <td>21.9</td> <td>3.4</td> <td></td> </tr> <tr> <td>GML</td> <td>793</td> <td>0</td> <td>0.6</td> <td>32.0</td> <td>1.4</td> <td>2.6</td> <td></td> </tr> <tr> <td>Sum</td> <td>793</td> <td>3.8</td> <td>2.1</td> <td>33.9</td> <td>23.3</td> <td>6.0</td> <td>2.1</td> </tr> </tbody> </table>	CM	0	3.8	1.4	1.9	21.9	3.4		GML	793	0	0.6	32.0	1.4	2.6		Sum	793	3.8	2.1	33.9	23.3	6.0	2.1								
	CM	0	3.8	1.4	1.9	21.9	3.4																											
	GML	793	0	0.6	32.0	1.4	2.6																											
	Sum	793	3.8	2.1	33.9	23.3	6.0	2.1																										
	(3)	<table border="1"> <tbody> <tr> <td>0</td> <td>0.41</td> <td>1</td> <td>0</td> <td>0</td> <td>0.14</td> <td></td> </tr> </tbody> </table>	0	0.41	1	0	0	0.14																										
0	0.41	1	0	0	0.14																													
CM 2019, single emitter sources	(1)	<table border="1"> <thead> <tr> <th>Fit → ↓ Test</th> <th>CM Lognorm.</th> <th>GML Lognorm.</th> <th>Joint Lognorm.</th> <th>CM Gen.Logn.</th> <th>GML Gen.Logn.</th> <th>Joint Gen.Logn.</th> <th>Min.</th> </tr> </thead> <tbody> <tr> <td>CM</td> <td>-166</td> <td>-164</td> <td>-164</td> <td>-164</td> <td>-154</td> <td>-158</td> <td>-166</td> </tr> <tr> <td>GML</td> <td>67852</td> <td>3499</td> <td>3499</td> <td>6414</td> <td>3484</td> <td>3486</td> <td>3484</td> </tr> <tr> <td>(Params)</td> <td>(2.346, 0.541)</td> <td>(0.303, 1.056)</td> <td>(0.311, 1.052)</td> <td>(-6.901, 9.355, 16.52)</td> <td>(0.654, 0.731, 1.531)</td> <td>(0.629, 0.770, 1.619)</td> <td></td> </tr> </tbody> </table>	Fit → ↓ Test	CM Lognorm.	GML Lognorm.	Joint Lognorm.	CM Gen.Logn.	GML Gen.Logn.	Joint Gen.Logn.	Min.	CM	-166	-164	-164	-164	-154	-158	-166	GML	67852	3499	3499	6414	3484	3486	3484	(Params)	(2.346, 0.541)	(0.303, 1.056)	(0.311, 1.052)	(-6.901, 9.355, 16.52)	(0.654, 0.731, 1.531)	(0.629, 0.770, 1.619)	
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	CM	0	1.8	1.7	1.4	12.1	7.8																											
	GML	64336	14.1	14.2	2930	0	1.8																											
	Sum	64336	15.9	15.9	2931	12.1	9.7	9.7																										
	(3)	<table border="1"> <tbody> <tr> <td>0</td> <td>0.04</td> <td>0.04</td> <td>0</td> <td>0.30</td> <td>1</td> <td></td> </tr> </tbody> </table>	0	0.04	0.04	0	0.30	1																										
0	0.04	0.04	0	0.30	1																													
CM 2020-21, 150 m sources	(1)	<table border="1"> <thead> <tr> <th>Fit → ↓ Test</th> <th>CM Lognorm.</th> <th>GML Lognorm.</th> <th>Joint Lognorm.</th> <th>CM Gen.Logn.</th> <th>GML Gen.Logn.</th> <th>Joint Gen.Logn.</th> <th>Min.</th> </tr> </thead> <tbody> <tr> <td>CM</td> <td>-137</td> <td>-129</td> <td>-136</td> <td>-136</td> <td>-129</td> <td>-130</td> <td>-136</td> </tr> <tr> <td>GML</td> <td>14235</td> <td>1080</td> <td>1081</td> <td>1625</td> <td>1082</td> <td>1082</td> <td>1080</td> </tr> <tr> <td>(Params)</td> <td>(2.650, 0.500)</td> <td>(0.880, 1.088)</td> <td>(0.928, 1.059)</td> <td>(-4.713, 7.623, 17.04)</td> <td>(-0.350, 2.167, 3.232)</td> <td>(-1.957, 3.599, 4.494)</td> <td></td> </tr> </tbody> </table>	Fit → ↓ Test	CM Lognorm.	GML Lognorm.	Joint Lognorm.	CM Gen.Logn.	GML Gen.Logn.	Joint Gen.Logn.	Min.	CM	-137	-129	-136	-136	-129	-130	-136	GML	14235	1080	1081	1625	1082	1082	1080	(Params)	(2.650, 0.500)	(0.880, 1.088)	(0.928, 1.059)	(-4.713, 7.623, 17.04)	(-0.350, 2.167, 3.232)	(-1.957, 3.599, 4.494)	
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	(3)	<table border="1"> <tbody> <tr> <td>0</td> <td>0.81</td> <td>1</td> <td>0</td> <td>0.39</td> <td>0.50</td> <td></td> </tr> </tbody> </table>	0	0.81	1	0	0.39	0.50																										
0	0.81	1	0	0.39	0.50																													

226 S6. Model fit and scaling

227 Results from the fit optimization for CM 2019 (150 m sources) are shown in Fig. S8. Measured
 228 data in each survival function are plotted according to Eqn. S4, which scales each sample to
 229 $S(x_{L,i}) = 1$ at the respective FDL, $x_{L,i}$, where i denotes the sample. The model function is
 230 correspondingly normalized using integrals over the density $p_i(x)$ given by Eqn. S5.

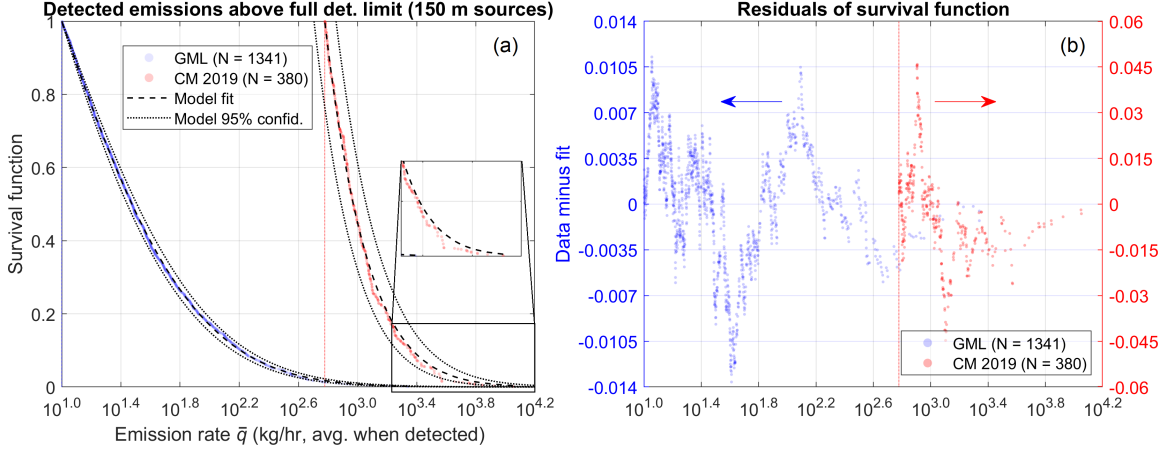


Figure S8: Joint model fitting of 150 m aggregated emission sources showing (a) survival function and (b) fit residuals. Inset: zoomed-in view of largest CM emission rates.

231 Residuals for both traces show that the survival function crosses the model multiple
 232 times without a strong bias toward the positive or negative values. However, CM residuals
 233 are negative for emission rates above roughly $10^{3.4}$ kg h⁻¹. This does not strongly impact the
 234 density function fit, but it does influence sample agreement with the model for cumulative
 235 emissions (i.e. the integral of the density function weighted by emission rate).

236 When plotted as density functions as in Fig. 2, traces are scaled to a common reference.
 237 The factor $1000/n_{\text{CM}}(x > x_{L,\text{CM}})$ is used to scale the CM data to 1000 detections above the
 238 CM FDL, where n is the number of detected sources in the specified range. The GML series
 239 is scaled by the factor

$$\frac{1000}{n_{\text{GML}}(x > x_{L,\text{GML}})} \frac{1}{\int_{x_{L,\text{CM}}}^{\infty} p_{\text{GML}}(x)}, \quad (\text{S8})$$

240 where “CM” or “GML” fill in the subscript i in Eqn. S5. The right-hand term of Eqn. S8
 241 rescales the number of detected sources above the GML FDL by the ratio of the survival

242 functions to each FDL, where both terms in the ratio are evaluated at the CM FDL (that is,
 243 recognizing the numerator as $1 = \int_{x_{L,CM}}^{\infty} p_{CM}(x)$). These scale factors assume that the size
 244 of both samples is sufficiently large above the respective FDL and that sample error in the
 245 number of detected sources is negligible. Likewise, the model function is scaled by the factor
 246 $1000 / \int_{x_{L,CM}}^{\infty} p_{GML}(x)$ but with no assumptions about sample size.

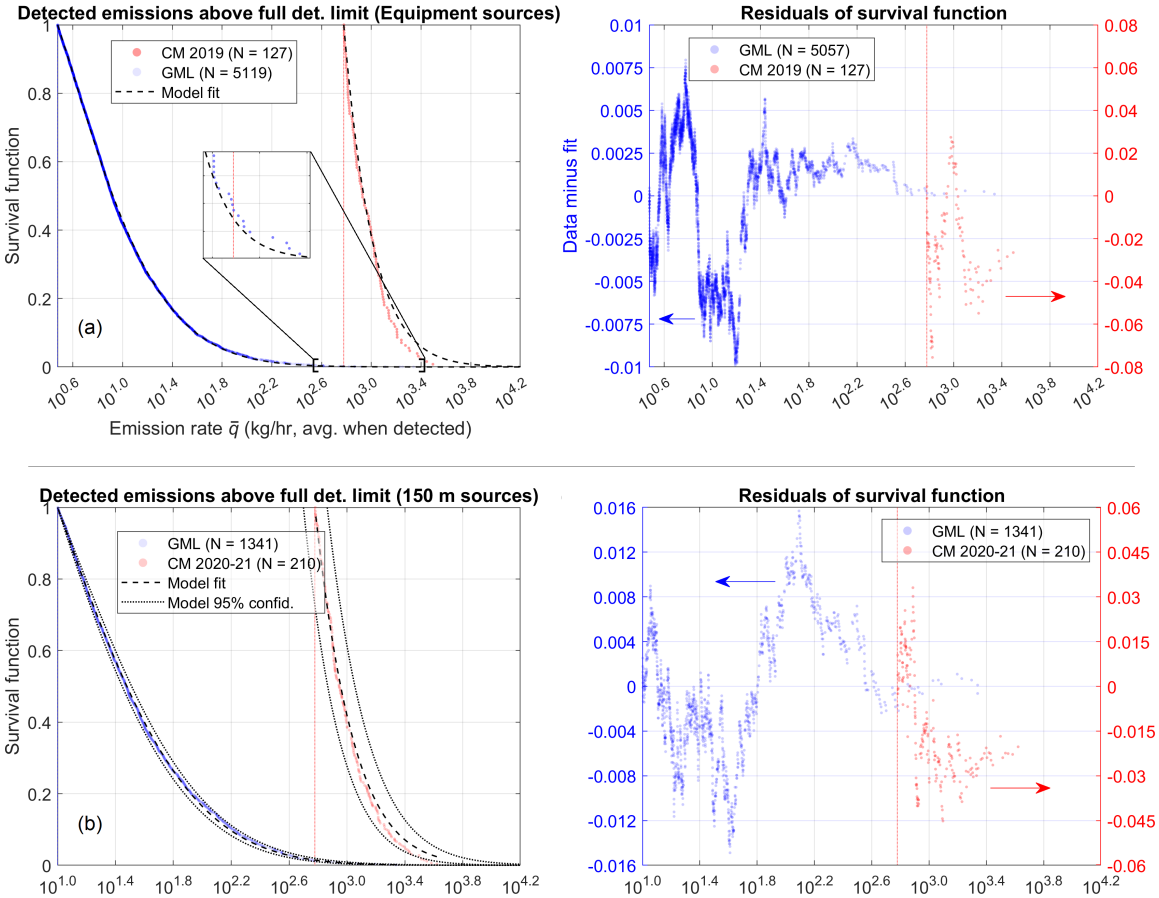


Figure S9: Survival function and residuals showing joint model fitting of GML to (a) CM 2019 equipment-scale sources and (b) CM 2020-21 facility-scale sources (150 m diameter). Inset (a): zoomed-in view of largest GML emission rates.

247 Fit results in terms of the survival function are shown for CM 2019 equipment-scale
 248 sources and CM 2020-21 150 m sources in Fig. S9. In cases where the CM residuals tend
 249 to be negative but GML residuals tend to be positive, the fit is located in between the two
 250 samples. To some extent, the model disagrees with CM due to the GML measurements in
 251 these cases. Other possible reasons for the fit to be above the CM measurement distribution

252 in the heavy tail include (1) the model functional form or parameter values do not adequately
253 represent the rapid decline in sources in the heavy tail, or (2) the heavy tails measured by
254 CM are reshaped relative to GML by other factors such as quantification bias, such as that
255 reported in Ref. S6. Assuming that the CM and GML measured distributions are in fact
256 aligned, apparent differences may be explained by the heavy tail of the distribution rolling
257 off faster than the model fits above emission rates of roughly $10^{3.4}$ kg h⁻¹.

258 **S7. Equipment-scale emission source filtering**

259 Human analysts classified CM 2019 plume images in a binary fashion as either “single emitter”
260 or “multiple or unclear.” Detections classified as “single emitter” were selected for analysis.
261 Scan data were cut to include only the first scan at each source, making the data set effectively
262 single-scan ($f = 1$). This filter changed the number of CM 2019 sources after other filters
263 (removing pipelines, restricting to GAO survey polygons) from 1348 at 150 m to 645 single
264 emitters. For GML, skipping aggregation increased the number of GML sources from 2727
265 to 7176, though the number of GML sources above the CM FDL shrank from 17 to 9. Spatial
266 aggregation significantly affects the CM distribution, whereas O&G pipeline sources do not
267 (Section S11). Fit residuals display similar behavior to those from 150 m sources (Sect. S6).

268 **S8. Density plots (CM 2019 single emitters, CM 2020-21)**

269 Detected emission density from analysis with CM 2019 single emitter and CM 2020-21 sam-
270 ples is shown in Fig. S10. As shown in Fig. S10a, the CM 2019 single emitter sensitivity
271 at 50% POD is seen to be 321 [277, 382] kg h⁻¹. This overlaps with the confidence interval
272 of the CM 2019 detection sensitivity at 150 m aggregation. Further details of the single
273 emitter distribution in contrast to the 150 m distribution are described around Fig. 3. For
274 the CM 2020-21 sample, Fig. S10b shows the density with CM 2019 traces reproduced for
275 comparison. The sensitivity at 50% POD is 252 [227, 282] kg h⁻¹, which suggests a possi-

276 ble improvement over the CM 2019 campaign sensitivity (possibly from flight altitude; CM
 277 2020-21 campaigns were flown at 4.5 km AGL only). The CM 2020-21 sample is scaled to
 278 the CM 2019 sample using the ratio of the GML scale values given in Eqn. S8 from analysis
 279 with both CM samples, a value of 0.833. This ensures that the CM traces are scaled to one
 280 another such that the GML traces from both analyses coincide exactly. In other words, the
 281 CM traces are both scaled to 1000 total CM 2019 detections above the CM FDL using the
 282 GML distribution as a common reference.

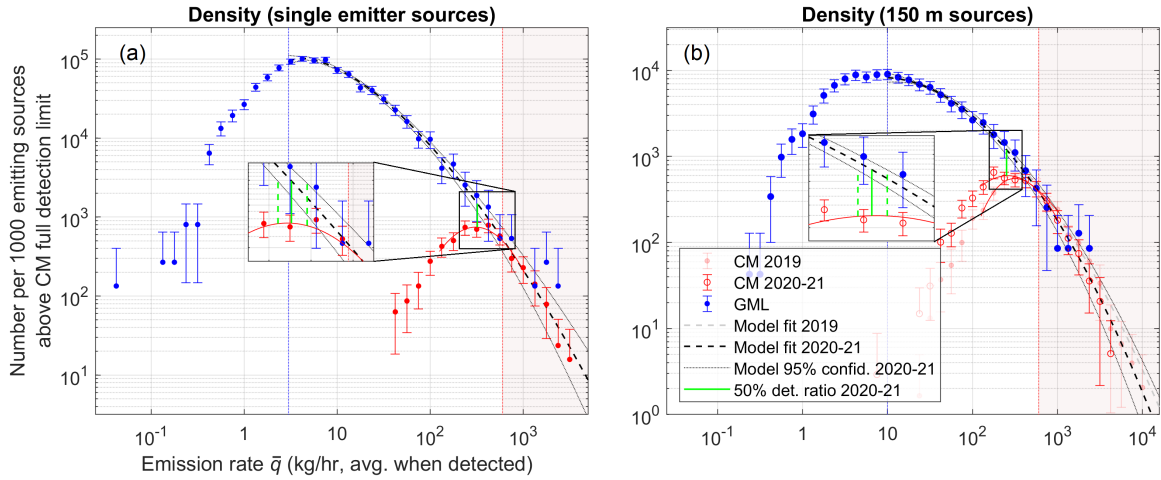


Figure S10: Emission source density from joint analysis with GML for (a) CM 2019 single emitters and (b) CM 2020-21 with 150 m sources. Zoomed in view near the CM sensitivity (insets) shows the 50% detection ratio with respect to model function, along with confidence bounds (dashed green). Model function for CM 2020-21 (150 m) distribution follows Eq. 1 with $m = 2$, $x_0 = 0.928$, $b = 1.059$.

283 S9. Cumulative emission rate distribution (CM 2020-21)

284 The CM 2020-21 cumulative emission rate distribution is displayed in Fig. S11. CM 2019
 285 measured data and model function are reproduced in the plot for comparison. Both CM data
 286 sets were analyzed jointly with the GML data set. By comparing the measured distributions
 287 at 10 kg h^{-1} to the model function and its confidence bounds at this emission rate, we find
 288 that CM 2020-21 measured 43.4% [37.8%, 49.2%] of the total cumulative emission rate from
 289 150 m sources above 10 kg h^{-1} , whereas GML measured 98.2% [85.5%, 111.3%]. These

290 results are similar to those obtained with CM 2019 data, suggesting consistency between the
 291 CM 2019 and 2020-21 campaigns.

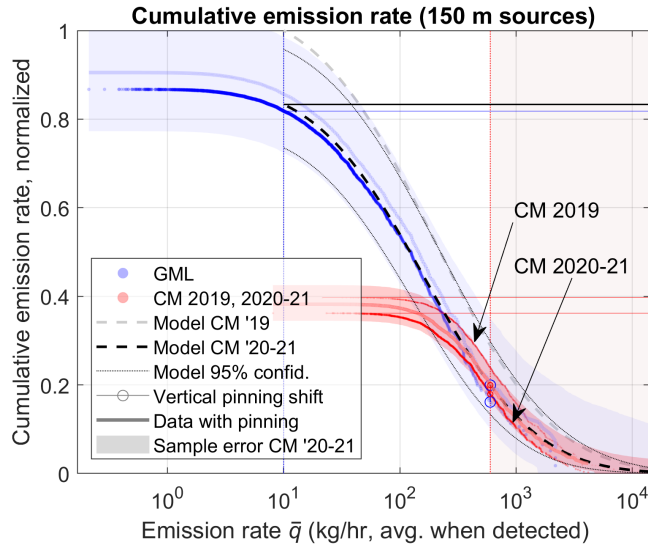


Figure S11: Cumulative emission rate distribution of GML and CM 2020-21 measurements with 150 m diameter aggregated emission sources. Joint GML/CM 2019 model and CM 2019 measured distribution are reproduced for comparison. Distributions from CM 2020-21 joint analysis are scaled to those from the CM 2019 analysis. All traces are normalized to equivalent campaign scale (spatial area, number of overflights). Vertically shifted copies of measured data pinned to the value of the model distribution at the CM FDL guide the eye to suggest the shape of the measured distribution supposing sample error above the CM FDL were suppressed.

292 Relative scaling of the CM 2020-21 density function to CM 2019 results in the different
 293 cumulative emission rate totals shown in the plot. This was performed as described in
 294 Sect. S8, where GML density was used as a reference. GML to CM scaling is implied directly
 295 from the joint fit without any ad hoc parameters. The ratio of measured totals between the
 296 two campaigns (CM 2020-21/CM 2019) is 91% when scaled to one another accounting for
 297 sample size and number of overflights, using GML as a reference.

298 S10. Monte Carlo estimation of sample error

299 A Monte Carlo algorithm is used to obtain percentile ranges on the cumulative emission rate
 300 as a function of source emission rate. New samples are synthesized from the joint best-fit

301 density function at emission rates above the respective FDL. Size of synthesized samples
302 matches the number of detected sources above the FDL in the measured samples. Since
303 the density function expresses the single-scan equivalent, the number of overflights is one
304 for each synthesized detection. For each of $n_{MC} = 10,000$ Monte Carlo trials, a vector of
305 $n(x < x_{L,i})$ random numbers uniformly distributed on the interval $(0, 1)$ is generated. The
306 random numbers are input as arguments to the inverse of the survival function on the domain
307 above the FDL to generate source emission rates. Cumulative emission rate versus source
308 emission rate is calculated from each Monte Carlo trial. Percentiles are found from the set
309 of synthesized Monte Carlo trials on a grid of source emission rates.

310 Simulated sample error supports the emission rate domain down to each respective FDL.
311 Sample error below the FDL is represented by assuming the same cumulative emission rate
312 increase as the measured sample, with no additional error contributed by samples below the
313 FDL.

314 **S11. Exclusion of pipelines from CM data set**

315 Exclusion of O&G pipeline sources in the CM 2019 sample produces negligible change in the
316 survival function. By comparison, the effect of filtering the data to single emitter sources
317 changes the distribution significantly. Fig. S12 shows the survival function for “with pipeline”
318 and “without pipeline” filters for 150 m and single emitter aggregation styles. Moderate p -
319 values, and hence no statistically significant difference, are seen between the distributions
320 including or excluding pipelines within each aggregation style. However, the p -value for
321 a comparison across aggregation styles is outside the 95% confidence interval ($p < 0.05$),
322 indicating that those distributions differ significantly.

323 Despite the lack of significant change in the shape of the CM 2019 distribution with the
324 inclusion or exclusion of pipelines, small differences around the detection roll-off (300 kg h^{-1})
325 lead to slightly different estimates of the detection sensitivity. Density functions for both

326 aggregation styles with and without pipelines are shown in Fig. S13. Comparing the model
 327 bounds at 95% confidence to respective cubic polynomial roll-off fits yields detection sen-
 328 sitivity intervals of 233-279 (with pipelines) and 256-309 kg h^{-1} (no pipelines) for 150 m
 329 sources, and 258-356 and 277-382 kg h^{-1} , respectively, for single emitter sources. Since these
 330 intervals overlap significantly, the detection sensitivity roll-off can be considered as weakly
 331 dependent on both types of data filter.

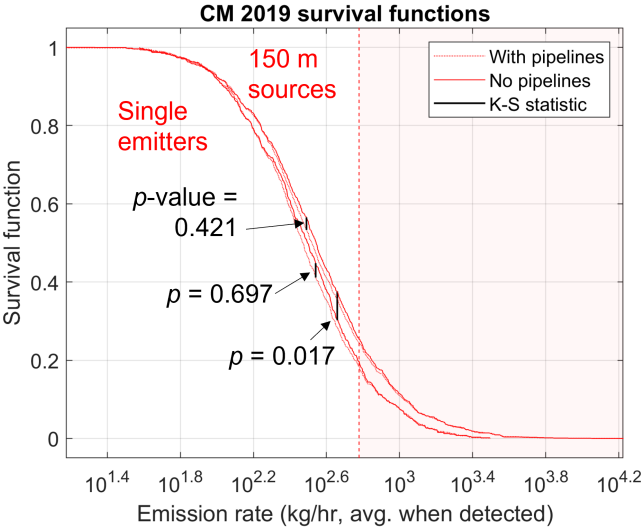


Figure S12: Comparison of CM 2019 survival functions over the range of emission rates in the sample. Data are filtered to either include or exclude O&G pipeline emission sources at both 150 m aggregation and single emitter sources. p -values are indicated for “with pipeline” and “without pipeline” comparisons within each source type and a “no pipeline” comparison across the two source types.

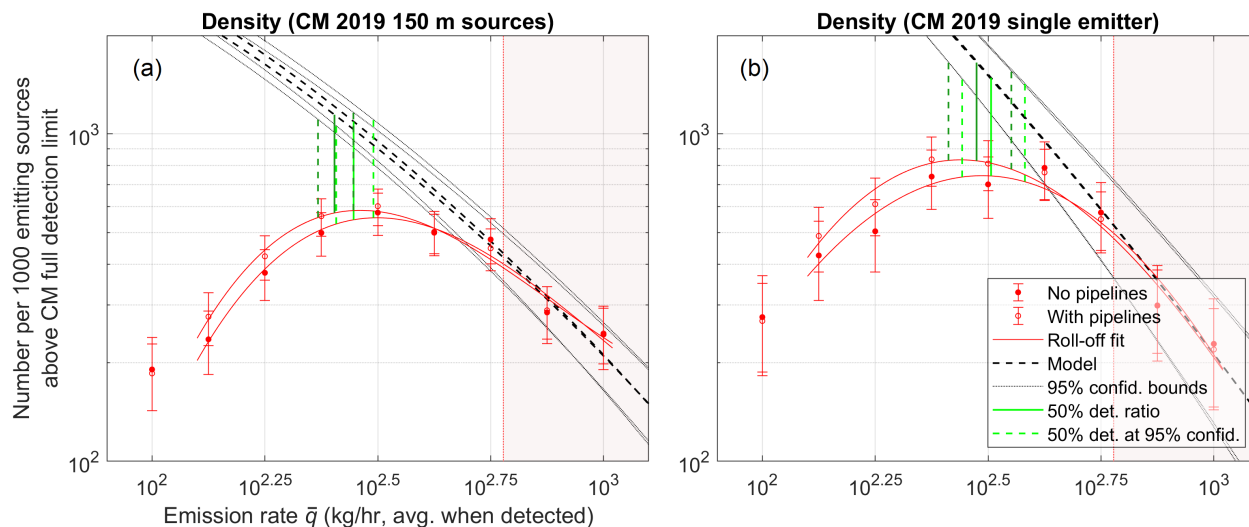


Figure S13: CM 2019 detected emission density plots showing influence of O&G pipeline sources on emission distribution around detection roll-off for 150 m aggregated sources (a) and single emitter sources (b). Model functions are reproduced from joint analysis with GML.

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