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1 Sea ice concentration satellite retrievals influenced by
2 surface changes due to warm air intrusions: A case study
3 from the MOSAiC expedition

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21 **Abstract**

22 Warm air intrusions over Arctic sea ice can rapidly change the snow and ice sur-
23 face conditions and can alter sea ice concentration (SIC) estimates derived from
24 satellite-based microwave radiometry without altering the true SIC. Here we focus
25 on two warm moist air intrusions that produced surface glazing during the Mul-
26 tidisciplinary drifting Observatory for the Study of Arctic Climate (MOSAiC)
27 expedition that reached the research vessel *Polarstern* in mid-April 2020. After
28 the events, we observe increased SIC deviations between different satellite prod-
29 ucts, including climate data records, and especially an underestimation of SIC
30 for algorithms based on polarization difference. To examine the causes of this
31 underestimation, we use the extensive MOSAiC snow and ice measurements to

32 computationally model the brightness temperatures of the surface on a local scale.
33 We further investigate the brightness temperatures observed by ground-based ra-
34 diometers at frequencies 6.9 GHz, 19 GHz and 89 GHz.

35 We show that the drop in the retrieved sea ice concentration of some satellite
36 products can be attributed to large-scale surface glazing, i. e., the formation of a
37 thin ice crust at the top of the snowpack, caused by the warming events. Another
38 mechanism affecting satellite products which are mainly based on gradient ratios
39 of brightness temperatures, is the interplay of the changed temperature gradient
40 in the snow and snow metamorphism. From the two analyzed climate data record
41 products, one is less affected by the warming events. The low frequency channels
42 at 6.9 GHz were less sensitive to these snow surface changes, which could be
43 exploited in future retrievals of sea ice concentration.

44 1. Introduction

45 The frozen blanket of the Arctic Ocean, the Arctic sea ice, controls fluxes of heat,
46 moisture and momentum between ocean and atmosphere. Arctic sea ice is also
47 a habitat for marine organisms. As the Arctic is warming rapidly over the last
48 decades, this integral component of the cryosphere is strongly affected, which has
49 a multitude of consequences both locally and outside of the Arctic (Semmler et al.,
50 2012; Fox-Kemper et al., 2021, e.g.). During the last decades, the summer Arctic
51 sea ice extent declined by about -13% (Perovich et al., 2017; Meier and Stroeve,
52 2022).

53 Observations from satellite microwave radiometers provide a more than 40-
54 year-long time series of Arctic sea ice area (Spreen and Kern, 2017), giving main
55 insights for sea ice research and providing findings for climate research (Meredith
56 et al., 2019). Retrievals of sea ice concentration (i. e., the percentage of an area
57 covered by sea ice) using passive microwave sensors take advantage of the strong
58 microwave emission contrast of ice and ocean, i.e., high for ice and often sev-
59 eral tens of Kelvin lower for ocean. Retrieval algorithms are either (i) based on
60 polarization difference, (ii) combine different frequencies, or (iii) use both, differ-
61 ent polarizations and frequencies. Overviews and inter-comparisons of different
62 retrieval algorithms are for example given in Ivanova et al. (2015) and Andersen
63 et al. (2006, 2007). In this case study, involving two moist and warm air intrusions
64 in April 2020, we investigate the performance of common sea ice concentration
65 retrievals of type (i) and (iii).

66 Moist air intrusions transporting water vapor poleward play an important role
67 in the Arctic climate system. They increase the downward longwave radiation flux

68 and the skin temperature and thus contribute to Arctic warming in winter (Woods
69 et al., 2013; Hao et al., 2019). There is evidence for an increase in the frequency of
70 extreme warming events, atmospheric rivers and cyclones in the central Arctic in
71 winter, related to an increase in meridional heat and moisture transport (Graham
72 et al., 2017; Valkonen et al., 2021; Rinke et al., 2017; Hao et al., 2019; Henderson
73 et al., 2021; Woods and Caballero, 2016; Zhang et al., 2023). We will refer to
74 these events as "warm air intrusions" in the following acknowledging that for spe-
75 cific events they can be different, e. g., in terms of moisture. Warm air intrusions
76 and associated wind and temperature changes can alter the sea ice concentration
77 by ice advection, breaking the ice and opening leads, or, to a lesser degree, by
78 melting the ice (mainly in connection with upwelling of warmer, e. g., Atlantic,
79 water along the ice margins but also by direct melting of the ice surface). How-
80 ever, warm air intrusions can also significantly change the atmosphere and the sur-
81 face in ways that alter satellite-measured microwave brightness temperatures (TB)
82 without changing the ice concentration (Liu and Curry, 2003). These changes can
83 cause spurious changes in sea ice concentration products (Tonboe et al., 2003)
84 based on brightness temperatures, i. e., they can cause wrong ice concentration re-
85 trievals in some cases. One possible effect on the snow surface is surface glazing.
86 By glazing we mean the formation of a thin ice layer on top of the snow due to
87 melt or precipitation (rain on snow (Stroeve et al., 2022, e.g.)) or other mecha-
88 nisms, e. g. winds, as observed for example in Antarctica (Scambos et al., 2012).
89 Onstott et al. (1987) found that a crust reduces emissivity at 37 and 94 GHz sig-
90 nificantly because of scattering within this layer. Smith (1996) and Comiso et al.
91 (1997) conjectured that ice layers in the snow can be a reason for an underestima-
92 tion of ice concentration referring to Mätzler et al. (1984). Mätzler et al. (1984)
93 showed that ice layers introduce interfaces with different refractive indices, ef-
94 fecting especially the horizontally-polarized brightness temperatures close to the
95 Brewster angle as described by the Fresnel equations and therefore alter the po-
96 larization difference. Rees et al. (2010) also observed this effect due to ice lenses
97 on snow on land in the Arctic.

98 To study and increase the understanding of the various processes that lead to
99 the strong recent changes in the Arctic climate, the Multidisciplinary drifting Ob-
100 servatory for the Study of Arctic Climate (MOSAIC) (Nicolaus et al., 2022; Shupe
101 et al., 2022a; Rabe et al., 2022) expedition was conducted for a full year from Oc-
102 tober 2019 to September 2020. The research icebreaker *R/V Polarstern* (Alfred-
103 Wegener-Institut Helmholtz-Zentrum für Polar- und Meeresforschung, 2017) was
104 moored to a sea ice floe and drifted with it. During the campaign ship-based,
105 ground-based, and airborne measurements of the ocean, sea ice, atmosphere, bio-

106 geochemistry and ecosystem in the vicinity of the ship were collected. The area
107 within about 2 km of *Polarstern* — named the ‘Central Observatory’ (CO)— was
108 intensively studied.

109 Warm air intrusions in April 2020

110 After a long period of cold winter conditions, two warm and moist air intrusions in
111 April 2020 dramatically warmed the CO (Shupe et al., 2022b). During the warm
112 air intrusions, air temperatures increased by up to 30 °C at the MOSAiC site, get-
113 ting close to and even above the melting point. The atmospheric events included
114 record-breaking total water vapor (Rinke et al., 2021) and high liquid water path,
115 increased wind speeds, precipitation, as well as changes in the aerosol regime
116 (Dada et al., 2022) and surface snow metamorphism. Before, during, and after the
117 warm air intrusion the actual SIC in the vicinity of MOSAiC was high ($> 95\%$).
118 Single leads opened during the events but nothing major in comparison to the peri-
119 ods before and after. This is confirmed by optical (MODIS) and radar (Sentinel-1)
120 satellite data as well as by observations from the expedition participants and by
121 helicopter-borne thermal infrared imagery (Thielke et al., 2022). The latter gives a
122 value for lead fraction, i. e., fraction of open water and thin (< 30 cm) young ice,
123 of the order of 1.5 % over the Central Observatory on April 23. Still the warm
124 air intrusion events affected satellite products of SIC. In conjunction with the
125 warming events, most satellite products showed a (wrong) decrease in SIC and
126 inter-product variability increased.

127 Microwave emission from open water depends mainly on surface tempera-
128 ture and surface roughness related to wave and foam formation. The microwave
129 emission of the snow-ice system, on the other hand, depends on snow and sea ice
130 properties such as density, temperature, salinity, stratification and microstructure.
131 Warm air intrusions can strongly impact these parameters and can thus change the
132 emission of the snow-ice system. Therefore, in order to understand their impact
133 on the sea ice concentration retrievals, we analyze the detailed snow and ice ob-
134 servations taken during MOSAiC (Nicolaus et al., 2022) and simulate the impact
135 of the April warm air intrusions on the microwave emission of the sea ice. This is
136 done on a floe-wide scale.

137 We then ‘zoom in’ further to adopt a local perspective and investigate the
138 ground-based radiometer observations taken at the Remote Sensing Site during
139 MOSAiC.

140 Of particular interest on all scales are responses of brightness temperatures to
141 large-scale surface glazing, which was observed at the MOSAiC CO.

142 In this paper, we study differences between several satellite ice concentration

143 products related to the April 2020 warm air intrusions. To explain the differences,
 144 we investigate the effect of these events on microwave brightness temperatures.
 145 We present the suite of such observations at *Polarstern* by satellite and ground-
 146 based radiometers on the ice floe measuring at the same frequencies (6.9 GHz,
 147 19 GHz and 89 GHz). We use in-situ snow and ice observations and microwave
 148 emission modelling to explore the impacts of glazing and snow metamorphism
 149 on brightness temperatures and, consequently, on SIC retrievals. The results are
 150 structured by the different scales that we are observing at, from a satellite view
 151 (Section 3.1) via a floe-wide perspective (Section 3.3) to a specific on-ice site
 152 (Section 3.4).

153 2. Data

154 2.1. Sea ice concentration: satellite products

155 We compare the sea ice concentration (SIC) around *Polarstern* based on different
 156 algorithms developed for satellite passive microwave remote sensing using differ-
 157 ent frequencies and polarization combinations. The datasets used are described in
 158 more detail in the following subsections. Table 1 provides an overview including
 159 the frequency channels that are used to compute SIC and the grid spacing. All
 160 products are available daily. The collocation procedure is the same for all prod-
 161 ucts: In order to account for drift we use *Polarstern*'s position resampled to hourly
 162 values and then choose the closest grid point in the satellite product for each hour.
 163 We then average over the whole day.

164 2.1.1. ASI Sea Ice algorithm

The ASI algorithm exploits the high spatial resolution of near 90 GHz channels
 and was initially developed for SSM/I sensors (Svendsen et al., 1987; Kaleschke
 et al., 2001). It was later adapted for the AMSR-E and AMSR2 sensors (Spren
 et al., 2008; Melsheimer, 2019). The polarization difference (PD)

$$\text{PD} = \text{TB V} - \text{TB H}, \quad (1)$$

165 where V denotes vertical and H horizontal polarization, at 89 GHz over open
 166 ocean is larger than over sea ice. This is used by the algorithm to distin-
 167 guish between these two surface types. The sea ice concentration is retrieved
 168 by a third-order polynomial of PD where the coefficients are determined by
 169 the tie points, i. e., typical values of PD over water (P_0) and consolidated ice,
 170 i. e., 100% ice concentration (P_1). To correct for weather influences over open

171 ocean, weather filters are applied. Here, we use the dataset operationally avail-
172 able on a 6.25 km grid at <https://seaice.uni-bremen.de> and <https://meereisportal.de>.
173

174 2.1.2. NASA Team Algorithm

175 The NASA Team algorithm (Cavalieri et al., 1984, 1997) uses vertically and
176 horizontally-polarized brightness temperature channels to calculate the polariza-
177 tion ratio, $PR = PD/(TB V + TB H)$, of 19.35 GHz, called PR(19) in the fol-
178 lowing, and the spectral gradient ratio $GR = (TB 37V - TB 19V)/(TB 37V +$
179 $TB 19V)$ between TB 19.35V and TB 37V, called GR(37/19) in the following.
180 These two ratios are then compared in a scatter plot where they form clusters.
181 These clusters can be identified as being correspondent to three surface types
182 (first-year ice, multiyear ice and ice-free ocean) and for each type three constant
183 tie points are determined (for each frequency channel). Values between the tie
184 points then represent mixtures of surface types. Weather filters are applied addi-
185 tionally. We use the NASA TEAM SIC operational product provided as part of
186 the NOAA/NSIDC Climate Data Record of Passive Microwave Sea Ice Concen-
187 tration, Version 4 (Meier et al., 2021).

188 2.1.3. NSIDC Climate Data Record

189 The National Snow and Ice Data Center (NSIDC) provides sea ice concentration
190 estimates as a Climate Data Record (CDR) starting in 1978 (Meier et al., 2021).
191 Here, SIC is computed both by NASA Team (see above) and the NASA Bootstrap
192 algorithm (Comiso, 1986; Comiso et al., 2017). The Bootstrap algorithm is based
193 on relationships of TB combinations of 19V and 37V, and 37V and 37H. Clus-
194 ters of pure surface types are determined in brightness temperature scatter plots
195 of these combinations. Tie points are derived daily based on these clusters. Ad-
196 ditionally, weather filters are applied. Then, the higher concentration value from
197 the two algorithms is chosen for each grid cell. We use the NSIDC CDR oper-
198 ational product provided by the NOAA/NSIDC Climate Data Record of Passive
199 Microwave Sea Ice Concentration, Version 4 (Meier et al., 2021).

200 2.1.4. OSI SAF Climate Data Record

201 OSI SAF Global Sea Ice Concentration interim climate data record (OSI-SAF
202 iCDR), release 2, provides daily sea ice concentration, starting in 2016 and us-
203 ing data from the SSMIS sensors from NOAA CLASS (EUMETSAT Ocean and
204 Sea Ice Satellite Application Facility., 2017). This sea ice concentration data set,
205 OSI-430-b, is based on a dynamic algorithm (Lavergne et al., 2016, 2019), gener-

Table 1: Summary of sea ice concentration products used.

Algorithm	Frequencies	Grid spacing
ASI	PD 89	6.25 km
NSIDC NASA Team	19.35V, 19.35H, 37V	25 km
NSIDC CDR	19.35V, 19.35H, 37V, 37H	25 km
OSI-SAF iCDR	19.35V, 37V, 37H	25 km

206 alizing the Bristol algorithm (Smith, 1996). Brightness temperatures at 19V, 37V
 207 and 37H span a 3-D space. Within this space, clusters or shapes close to lines for
 208 closed-ice and water are existent. The algorithm then projects the brightness tem-
 209 perature data on an optimized plane. This projection is found using daily updated
 210 training data sets, one for fully ice-covered and one for open water areas. The unit
 211 vector of this plane is found by principal component analysis (direction of high-
 212 est variance in ice concentration) and is then rotated to maximize accuracy. The
 213 final sea ice concentration is then calculated by a weighted linear combination of
 214 SIC computed from an algorithm dynamically tuned to perform better over open
 215 water and one dynamically tuned to perform better over high-concentration ice
 216 conditions, both applied in the respectively optimized planes in TB space. Addi-
 217 tionally, the TBs are corrected using operational analysis and forecast data from
 218 the European Centre for Medium-range Weather Forecasts (ECMWF) to account
 219 for atmospheric influences.

220 2.2. *Satellite remote sensing*

221 We investigate brightness temperatures measured by the microwave scanning radi-
 222 ometer AMSR2 on the GCOM-W1 spacecraft from the Japan Aerospace Ex-
 223 ploration Agency (JAXA), launched May 18, 2012. AMSR2 orbits the Earth at
 224 an altitude of 700 km in a near-polar, sun-synchronous sub-recurrent orbit with
 225 a swath width of 1450 km. The dual-polarized sensor has instantaneous fields of
 226 view (IFOV) ranging from 62 km \times 35 km at 6.9 GHz to 5 km \times 3 km at 89 GHz.
 227 We use the swath data, both ascending and descending, of the Level 1R product
 228 matched to the resolution of 6.9 GHz (Maeda et al., 2016), corresponding to an
 229 IFOV of 62 km \times 35 km. This product matches antenna patterns so that the bright-
 230 ness temperatures for all frequencies have the same field of view, facilitating easier
 231 comparisons between different frequencies. For every overflight of *Polarstern* we
 232 chose the measurement closest to the vessel’s hourly position. There are between
 233 five and seven overflights per day. For clarity, we show daily averaged values of

234 brightness temperatures at 6.9 GHz, 18.7 GHz, 36.5 GHz and 89 GHz in Figure 4.

235 2.3. *Ground-based remote sensing*

236 We focus on two microwave radiometers deployed during MOSAiC observing the
237 surface at frequencies ranging from 6.9 GHz to 89 GHz, similar to AMSR2. De-
238 tails are provided in the supplement Section 6.4. It is to note that due to increased
239 snow accumulation in front of the instruments after the warm air intrusions, the
240 snow was much deeper around the ground-based radiometers compared to the
241 surrounding MOSAiC floe. Especially during and after the second warm air intru-
242 sion, deep snow drifts formed in the field of view (FoV) of some radiometers (see
243 supplemental Section 6.4).

244 2.4. *Snow data*

245 Detailed snow measurements were performed during MOSAiC. In this study, we
246 analyze 132 SnowMicroPen (SMP) profiles (Macfarlane et al., 2021) taken be-
247 tween April 08 and April 27 to support our interpretation of the observed satel-
248 lite signals. From the raw SMP observations (penetration resistance) snow density
249 and specific surface area (SSA) can be estimated using empirical models (Proksch
250 et al., 2015; King et al., 2020). From density and SSA, the exponential correlation
251 length can be estimated, a parameter describing the microstructure of the snow
252 which is used in common snow microwave emission models (Tonboe et al., 2006,
253 e.g.,).

254 2.5. *Supporting data*

255 2.5.1. *Met Tower temperature*

256 We use the 2 m air temperature recorded from the 10 m meteorological mast in-
257 stalled on the floe (Cox et al., 2021).

258 2.5.2. *Precipitation*

259 To illustrate the timing of precipitation, we use data from the *Vaisala* Present
260 Weather Detector 22 (PWD22) precipitation gauge, an optical device that was in-
261 stalled on the deck of *Polarstern* and operated by the US Department of Energy
262 Atmospheric Radiation Measurement (ARM) program (Shi, 2019). Here, we use
263 1 minute mean precipitation rates (Figure 3). This product was also used as ref-
264 erence product in an inter-comparison of different snow precipitation sensors by
265 Wagner et al. (2022).

266 2.5.3. Total water vapor from radiosondes

267 The Level 2 dataset of balloon-borne radiosondes from the MOSAiC expedition
268 (Maturilli et al., 2021) was used to calculate total water vapor (TWV) from the
269 measured temperature, pressure and relative humidity profiles from the *Polarstern*
270 helicopter deck (at about 10 m height) to about 30 km altitude using the formula
271 for vapor pressure over liquid water below 0°C by Hyland and Wexler (1983) as
272 done in Walbröl et al. (2022). During the warm air intrusions, the radiosondes
273 were launched more often, up to seven times a day, while during the other periods
274 they were launched four times a day.

275 2.5.4. Liquid water path from HATPRO radiometer

276 We use liquid water path (LWP) retrieved from the ground-based low frequency
277 HATPRO (Humidity and Temperature Profiler) microwave radiometer operated
278 onboard the research vessel *Polarstern* during MOSAiC as input parameter to
279 model the atmosphere. The retrieved LWP is based on the retrieval algorithm as
280 described in Nomokonova et al. (2019). The radiometer has a temporal resolution
281 of 1 second. More information on this data can be found in Walbröl et al. (2022).

282 2.5.5. Reanalysis ERA5

283 From ECMWF fifth generation reanalysis ERA5 (Hersbach et al., 2020) we use
284 long- and shortwave radiation from the grid cell closest to *Polarstern* for the
285 SNOWPACK model simulations (Section 2.4).

286 2.5.6. Terrestrial laser scanner TLS

287 Supporting information about the snow surface topography was derived from
288 terrestrial laser scan (TLS) data taken on April 17 and on April 22 (Clemens-
289 Sewall et al., 2022c). The TLS uses a scanning, 1550 nm laser, to generate a
290 three-dimensional point cloud of the snow and ice surface at cm-scale resolu-
291 tion. See Deems et al. (2013) for a review of TLS applications to snow depth
292 measurements. Wind-blown snow particles were filtered out of TLS data using
293 the FlakeOut method (Clemens-Sewall et al., 2022a). From the measured topog-
294 raphy and its changes, we can deduce the changing snow thicknesses and effective
295 incidence angles (i.e., the incident angle of the tilted surface with respect to the
296 radiometer) within the footprints of the radiometers. The TLS data also includes
297 the backscatter reflectance of the surface at 1550 nm. Glazed areas are identifiable
298 in this data, because surface glazing reduces the backscatter reflectance (glazing
299 increases forward scattering and absorption).

300 3. Results

301 In the following we describe the temporal development of the retrieved sea ice
302 concentration and satellite-measured microwave brightness temperatures, first on
303 a large scale and then locally around *Polarstern*.

304 For the local analysis, we further describe the floe snow distribution by us-
305 ing the SMP measurements as model input to analyze the evolution of brightness
306 temperatures. In a second step we change the perspective to an even smaller scale
307 and study the data obtained by the ground-based radiometers. We then discuss the
308 integration of the observations from the different scales. Finally, this allows us to
309 come up with an interpretation of the satellite signal and the resulting differences
310 in sea ice concentration estimates.

311 Figure 1 shows the temporal evolution of total water vapor from ERA5 over
312 four days. The first intrusion, reaching the ship around April 16, originated in the
313 Russian Arctic while the second one around April 19 was approaching from the
314 North Atlantic. This illustrates the large area exposed to the warm air intrusions.

315 3.1. *Satellite perspective: sea ice concentration and brightness* 316 *temperatures*

317 The two warm air intrusions were large scale events (Figure 1) and thus also vis-
318 ible at large scale in the satellite data. This becomes evident when looking at
319 spatial maps of SIC (Figure 2). Here, we show the mean SIC for four consequent
320 days prior and after the event. Note that the SIC from ASI (first row) has a much
321 higher spatial resolution (6.25 km grid spacing compared to 25 km for all other
322 products). The last column shows the difference between the two time periods.
323 Blue colors denote a reduction of SIC after the events. In all products except for
324 the NSIDC CDR, decreases of SIC are visible in the Central Arctic to different
325 extents (black oval in the last column of Figure 2) as well as in the marginal ice
326 zone. The strongest effect is observed for the ASI product, followed by NASA
327 TEAM. Deviations between different products have increased after the events for
328 all products. The reduction in SIC in the Central Arctic is larger than suggested
329 by independent observations (from the ship and radar satellites) and likely caused
330 by the warm air intrusions as will be discussed in the following.

331 MOSAiC measurements allow us to study the effect of these intrusions lo-
332 cally: total water vapor (TWV), liquid water path (LWP) and 2 m air temperature
333 at *Polarstern* are shown in Figure 3, bottom row. For both warm air intrusions,
334 the rising temperatures (up to the freezing point for the second) coincide with in-
335 creased amounts of TWV (up to 13.4 mm) and LWP (up to around 0.47 mm). For

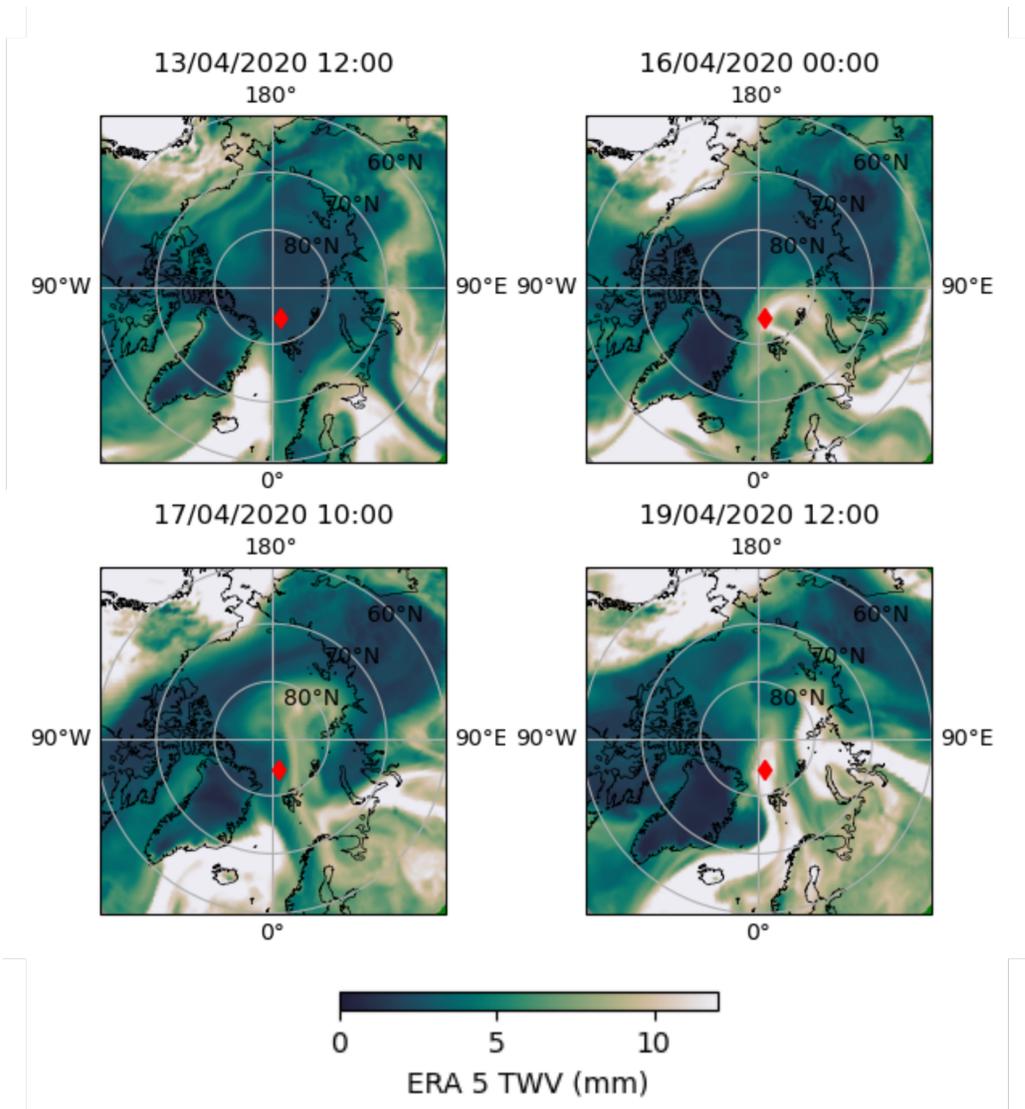


Figure 1: Total columnar water vapor (TWV, hourly values) from the reanalysis ERA5 for one day prior to the events and three days during the two warm air intrusions in April. The red diamond indicates the position of *Polarstern* at the day. The bright colors indicate high values of TWV.

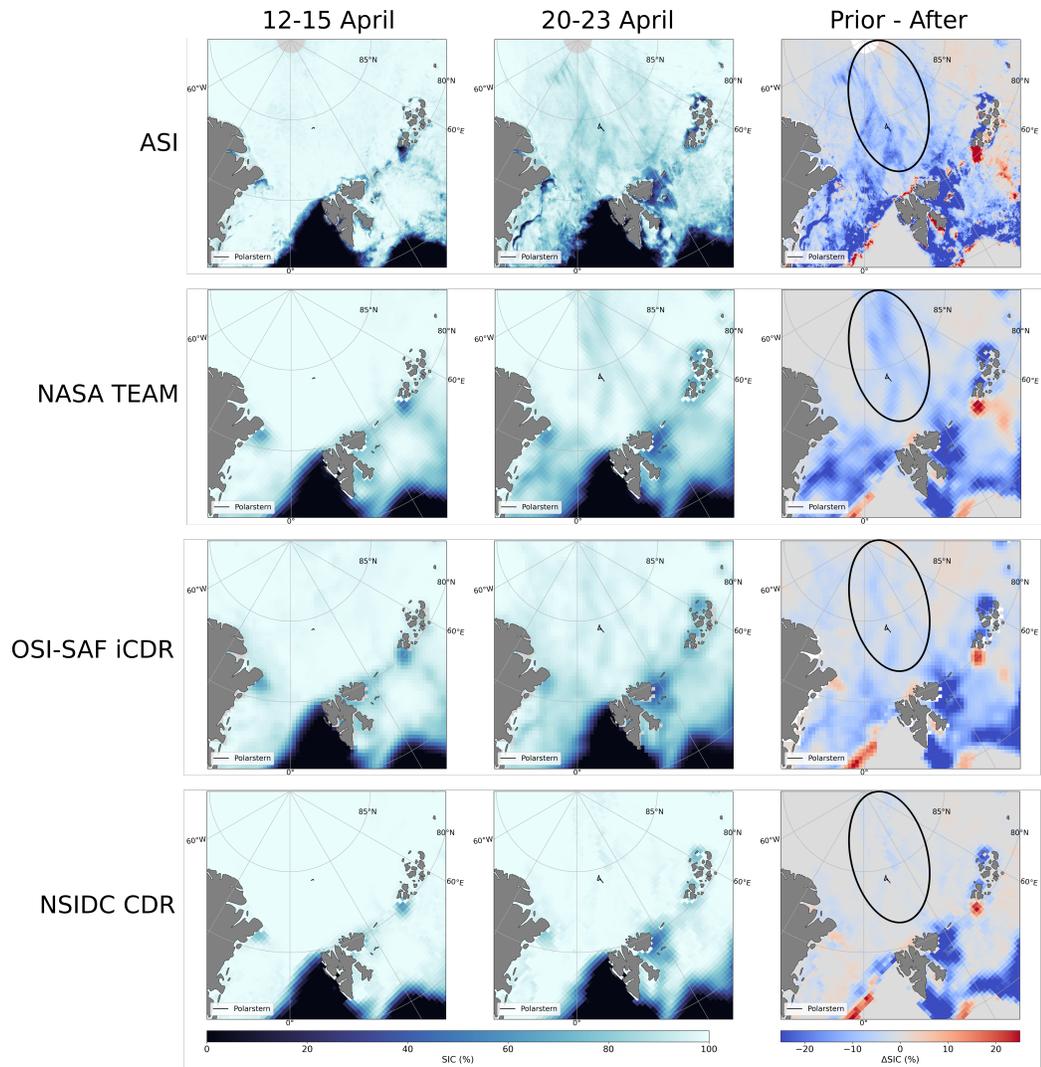


Figure 2: Sea ice concentration (SIC) from different satellite retrievals as four-day averages prior (first column) and after the intrusions (second column). The third column shows the difference between the second and first column, so the change in average SIC in the four days after the intrusions compared to the four days prior. The black oval in the last column marks a region where the different products deviate from each other after the events. *Polarstern*'s drift track is shown by a black line.

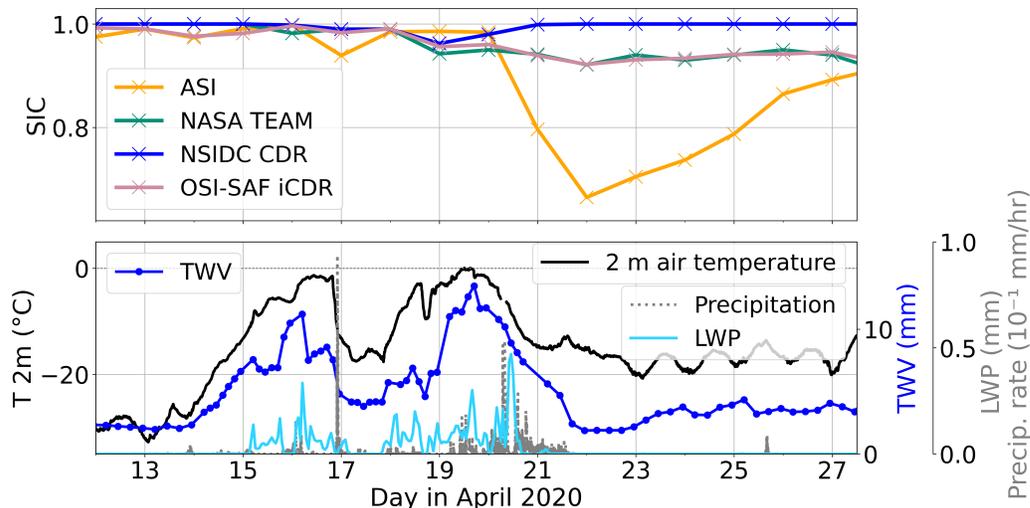


Figure 3: Sea ice concentration (SIC) and meteorological conditions during the events. Top row: SIC from four different operational products collocated to *Polarstern*. Bottom row: 2 m air temperature (black) from the Met Tower, total water vapor (TWV, dark blue) from radiosondes, liquid water path (LWP, light blue) from the HATPRO microwave radiometer (resampled to hourly values) and 1 min precipitation rates (grey) from the precipitation gauge (PWD22) installed on deck.

336 the ASI algorithm based on 89 GHz we observe a drop in SIC between the two
 337 warm air intrusions corresponding to the clear sky day on April 17 and a strong
 338 decline after the second event (Figure 3, top row). The SIC from NASA Team and
 339 OSI-SAF iCDR shows less variability during the intrusions but both decrease to
 340 92 % on April 22 and do not recover thereafter. The NSIDC CDR, on the other
 341 hand, stays at 100 % SIC. This algorithm includes both NASA Team and the Boot-
 342 strap algorithm, see Section 2.1.3, with the latter compensating for the decrease
 343 observed in SIC from NASA Team as discussed later. Before the events, all algo-
 344 rithms showed high sea ice concentration around 100 %. Similar to the large scale
 345 view (Figure 2), the spread between different products increased after the events.
 346 Using the NSIDC CDR data as reference, we observe SIC differences of around
 347 8 % for OSI-SAF iCDR and for NASA TEAM and up to 34 % for ASI. Although it
 348 has been observed by Tjernström et al. (2015) that warm air advections can cause
 349 rapid ice melt, and that dynamic effects can decrease SIC up to 3% in the high sea
 350 ice concentration domain. (Aue et al., 2022; Schreiber and Serreze, 2020) here

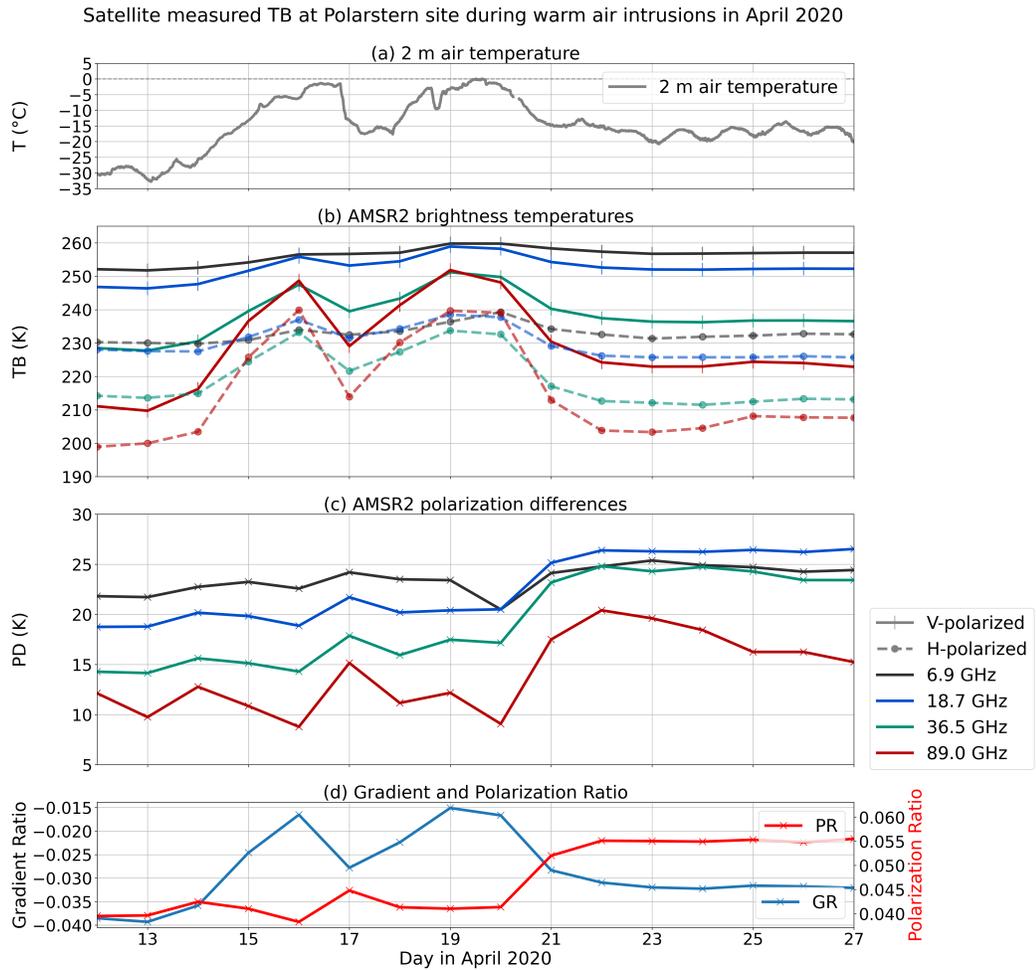


Figure 4: Effect of warm air intrusion on satellite-measured brightness temperatures. Top row shows the air temperature at 2 m (a). Collocated satellite measurement of TB (b) and PD (c) around *Polarstern* (daily averages). (d) Gradient ratio of 36.5 GHz and 18.7 GHz and polarization ratio of 18.7 GHz as defined in Section 2.1.2

351 the low sea ice concentration values at MOSAiC retrieved by some algorithms are
352 underestimating the actual sea ice concentration.

353 In our field observations, we did not observe a dramatic decrease in ice con-
354 centration at this time. This is also confirmed by SIC derived from satellite thermal
355 infrared data (MODIS instrument, only at clear sky, not shown), and such a signif-
356 icant drop in SIC can also not be seen in optical (MODIS) and radar (Sentinel-1)
357 satellite data (not shown). To understand this underestimation of sea ice concen-
358 tration, we analyze daily satellite AMSR2 brightness temperature data, shown in
359 Figure 4 (b). The signature of the warming events is clearly visible in the bright-
360 ness temperature time series at these frequencies. The higher frequencies (i. e.,
361 36.5 GHz and 89 GHz) follow the air temperature evolution in Figure 4 (a) more
362 closely. Especially ASI and NASA TEAM make use of polarization differences
363 PD and polarisation ratios PR, respectively. All frequencies show an increased
364 PD after the events as can be seen in the third row (c) of Figure 4. Again, the
365 higher frequencies show a larger increase. The PD increase on April 22 com-
366 pared to the mean of April 10 - April 13 ranges from a few Kelvin (≈ 3 K) for
367 6.9 GHz up to around 9 K and 11 K for 89 GHz and 36.5 GHz, respectively. A
368 smaller increase in PD at 89 GHz of a few Kelvin (≈ 4 K) is already visible around
369 April 17 between the two warm air intrusions. During that day, temperatures
370 dropped and clear sky conditions connected with a high longwave radiation loss
371 prevailed (Rinke et al., 2021). Images from the Panomax webcam onboard the ship
372 (<https://www.mosaic-panorama.org/>) reveal a lead opening close to
373 the ship on that day. Thus, the related decrease of SIC in the high-resolution ASI
374 product on April 17 is possibly real. During the second warm air intrusion, PD
375 for 89 GHz initially decreases again coinciding with rising TWP and LWP values.
376 Only after the event the rise in PD can be observed at all frequencies but strongest
377 at the higher frequencies. Figure 4 (d) illustrates that both the gradient ratio as
378 well as the polarization ratio show higher values after the intrusions (after April
379 20). The gradient ratio shows even higher values during the intrusions. While the
380 polarization ratio can be considered independent of physical temperature, this is
381 not necessarily the case for the gradient ratio due to different penetration depths
382 (and temperature-dependent permittivities) at 18.7 GHz and 36.5 GHz. Comiso
383 et al. (1997) argue that the effect should be small unless the snow cover emits
384 a sizable fraction of the measured TB. Assuming this is the case for 36.5 GHz
385 we attribute the strong response to rising and decreasing physical temperatures
386 to the strong change in temperature gradients in the snow cover as also depicted
387 later in Figure 5. For sea ice concentration estimates based on polarization dif-
388 ferences/gradient ratios, these changes result in (too) low retrieved SIC values as

389 shown in Figure 2 and Figure 3. It is not as straightforward to directly relate the
390 effect of these changes on SIC from OSI-SAF iCDR due to the complex retrieval
391 method. We note however that this algorithm is based on the three-dimensional
392 diagram in TB space of 19V, 37V and 37H, so strong changes in these frequencies
393 relative to each other will affect the retrieval. The OSI-SAF iCDR product pro-
394 vides two uncertainty estimates (algorithm uncertainty and "representativeness"
395 uncertainty, i. e. uncertainty due to resampling and mismatch of footprints at dif-
396 ferent channels (Lavergne et al., 2019)). Here, the sum of these two uncertainties
397 (gives as one standard deviation) increases from below 2% before the events to up
398 to 5.6% on April 21 (not shown) for the collocated data shown in Figure 3, i. e.,
399 the uncertainty estimates correctly identify a potential problem in the retrieved
400 SIC.

401 3.1.1. Atmospheric influence

402 In the few days after the intrusions, the temperature and TWV were on average
403 higher than before, while LWP is as low as during the first two weeks of April,
404 see Figure 3. Previous studies (Andersen et al., 2007; Oelke, 1997) demonstrated
405 that atmospheric events also can increase retrieved SIC and not only decrease like
406 in our case. Emissions from water vapor or liquid water path contributing to the
407 satellite signal are in general not polarized (Ulaby et al., 2013). This decreases PD,
408 which causes an increase in retrieved SIC for algorithms relying on PD (like ASI).
409 This is not observed here: we see an increase in PD after the warm air intrusion
410 events.

411 Scattered radiation, e. g., by ice particles in clouds, may have a polarized com-
412 ponent. However, observations by Troitsky et al. (2003) report values of PD and
413 duration of periods with polarization differences that are too small to explain the
414 development of PD we observe here. Also, compared to the surface emissions, the
415 contribution of the atmosphere is small at the low frequencies 6.9 and 19 GHz and
416 is thus in any case not the most likely candidate for the PD increase. Other factors
417 have to explain the increase in PD. We thus concentrate on explanations related to
418 changed surface emission.

419 3.2. *Snow accumulation and metamorphism*

420 Wagner et al. (2022) describe a significant snowfall event from April 16 – April
421 21 accompanying the warm air intrusions. Snowfall events could have had an in-
422 creasing effect on PD due to atmospheric scattering, but we observe the increase
423 *after* the snowfall. A detailed analysis suggests that much of the fresh snowfall

424 during this event may have been lost into leads (Clemens-Sewall et al., 2022b).
425 Even if there was snow accumulation on level ice, we would expect the fresher
426 snow (less dense, refractive index more similar to air) to decrease the PD. For the
427 large-scale area, this snowfall cannot explain the microwave signal. It is impor-
428 tant to keep in mind that the surface conditions at the Remote Sensing Site on
429 the MOSAiC floe were not representative for the larger area. Mainly because the
430 instruments themselves posed obstacles and caused artificial snow accumulation.
431 Thus, for the ground-based radiometers the snowfall is relevant in the interpreta-
432 tion (see Section 3.4).

433 For understanding and modeling the observed microwave emission, tempera-
434 ture profiles of the snow are important. Here we use temperature profiles from sim-
435 ulations with the SNOWPACK model. (Lehning et al., 2002a; Bartelt and Lehn-
436 ing, 2002; Lehning et al., 2002b; Wever et al., 2019) The model was initialized
437 with a snow pit from April 08 and driven with MET tower 2 m air temperature
438 and ERA-5 reanalysis data (Hersbach et al., 2020). For simulating the brightness
439 temperature of the snow/ice system of the MOSAiC floe, density and correlation
440 length from the SMP profiles are used (see section 6.1) together with the tem-
441 perature profiles from the SNOWPACK simulations. To match the varying snow
442 height of the SMP profiles, simulations with a snowpack of 10 to 30 cm in 5 cm
443 steps were performed. In these simulations, snowfall was omitted. We compared
444 the simulated temperature profiles with snow pit measurements (not shown). The
445 shape of the temperature profiles was well predicted from the SNOWPACK sim-
446 ulations, although we note a slight negative bias ($< -2^{\circ}\text{C}$). Figure 5 shows the
447 simulated snow temperature for a 20 cm deep snowpack. Overall, the snow tem-
448 perature increased by more than 10°C . After the events, the snow temperature
449 remains higher at the lower part of the snowpack for several days. These data
450 serves as input for the microwave emission modeling presented in Section 3.3.

451 The changes of the snow microstructure caused by the warm air intrusions
452 are evident when one looks at the SMP profiles shown in Figure 6. It shows the
453 distribution of the SMP profile averages for density (left), SSA (middle) and ex-
454 ponential correlation length (right). The data is split into before (April 08 - April
455 15) and after (April 20 - April 27) the warming events. A shift towards lower den-
456 sity and SSA and higher correlation length is visible in the data suggesting that
457 the warm air intrusions led to snow metamorphism. The strong and even inverted
458 temperature gradient effects the migration of water vapor (deposition and subli-
459 mation) in the snow and we would expect larger snow structures resulting in lower
460 SSA, which is indeed visible in the data. The changes in density are mainly in the
461 upper layers of the snow, possibly related to fresh snow.

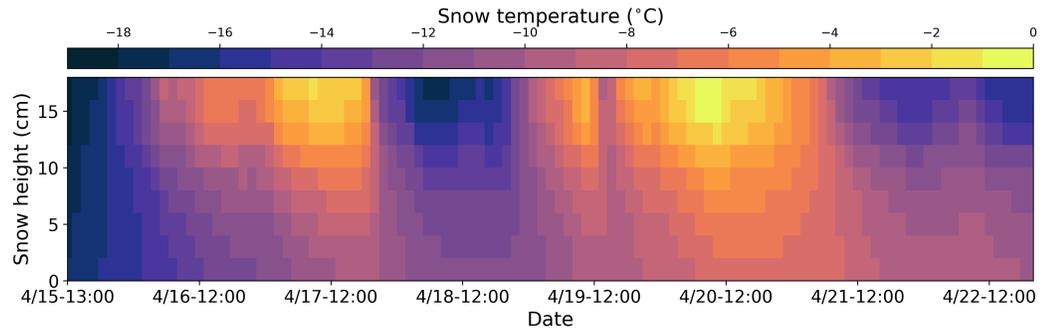


Figure 5: Snow temperature profile from a SNOWPACK simulation initialized with 20 cm of snow.

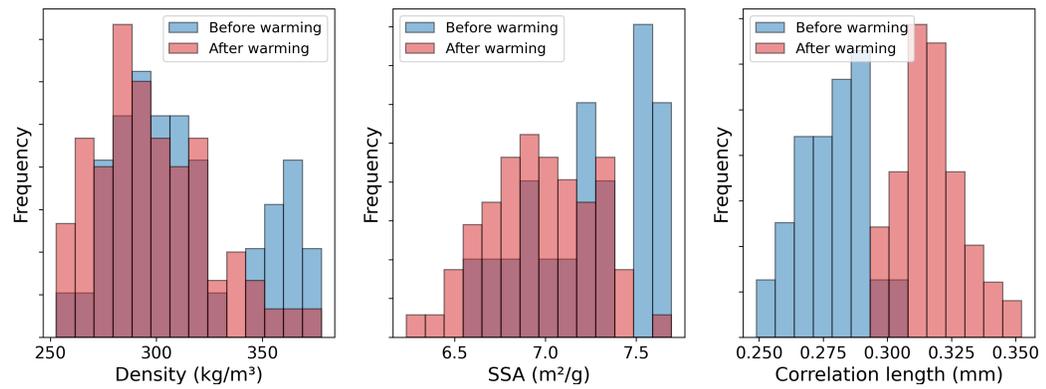


Figure 6: Histogram of the density, SSA and exponential correlation length of the 132 SMP profiles.

462 The change in increased correlation length affects the scattering strength of the
463 snow: in general higher correlation lengths lead to stronger scattering. On the other
464 hand, higher snow temperatures (Figure 5) increase emissions. To understand how
465 these surface changes affect brightness temperatures in more detail we therefore
466 model the microwave emissions in the following.

467 3.3. *Floe perspective: microwave emission modeling*

468 We further adopt the local perspective by using a statistical approach in modelling
469 the brightness temperatures. If we assume that the SMP measurements are repre-
470 sentative of the Central Observatory (of the order of ≈ 2 km) in terms of statistical
471 distributions of the measured quantities, we should be able to simulate the effect
472 of the warm air intrusions on the microwave signature. A comparison to satel-
473 lite data is only meaningful on a qualitative level due to the different scales and
474 atmospheric effects.

475 3.3.1. Brightness temperatures from modelling

476 We show the modelled polarization ratio of 18.7 GHz (PR(19)), the gradient ratio
477 of 36.5 GHz and 18.7 GHz at vertical polarization (GR(37/19)) and the polariza-
478 tion difference at 89 GHz (PD89) in Figure 7. The model output for the individual
479 frequencies is shown in the supplement (Figure S2). The blue histograms corre-
480 spond to input profiles from before (April 08 - April 15), while the red histograms
481 show modelled brightness temperatures using profiles from after the warming
482 events (April 21 - April 27). In general, the model output shows a clear increase
483 of brightness temperatures for all frequencies (at both polarizations) except for
484 36.5 GHz, where the situation is reversed (see Section 6.2). Qualitatively, an in-
485 crease is also seen in the satellite data, but not for the horizontally-polarized TB
486 at 18.7 GHz and 36.5 GHz.

487 When looking at the ratios that are used in many sea ice concentration algo-
488 rithms, the polarization ratio at 18.7 GHz remains mostly unchanged if no surface
489 glazing is modeled (Figure 7, upper row). GR(37/19) increases noticeable com-
490 pared to the values prior to the warming events and PD at 89 GHz shows only a
491 slight increase. From a satellite perspective we observe an increase in all three
492 quantities, see Figure 8, which is different from what we observe in the model
493 without glaze ice layer. Understanding the pronounced rise in satellite-measured
494 PD, which causes the strong drop in sea ice concentration from the ASI algorithm,
495 is key for potentially improving such PD based SIC algorithms. As discussed ear-
496 lier, the change in PD89 cannot be explained by variability in downwelling radia-

497 tion due to cloud cover and thus must be due to changes of the snow surface which
498 is however not captured by the quantities derived from SMP profiles (density, SSA
499 and correlation length).

500 3.3.2. Model experiment floe: simulation of a glaze ice layer

501 Visual observations during the expedition and also the TLS reflectance data (see
502 supplemental material, Figure S3 and Figure S4) suggest the development of a
503 glaze ice layer, which formed in some spots of the ice floe during the first, and
504 almost everywhere during the second warming event. Glaze ice layers at the top
505 of the snowpack can have a strong impact on the microwave emission of the snow
506 (Grenfell and Putkonen, 2008; Mätzler et al., 1984; Rees et al., 2010; Smith, 1996,
507 e.g.). Studies on the effect of such ice layers at the surface of the snowpack have
508 shown that, close to the Brewster angle, they usually have a minor impact on
509 vertically-polarized TB, but strongly influence horizontally-polarized TB (Rees
510 et al., 2010, e.g.) due to the high dielectric contrast between the snow and the
511 ice layer. Thus, algorithms utilizing polarization differences or ratios (e. g., ASI
512 and NASA-Team) will be influenced by the presence of such layers. How strong a
513 certain frequency is impacted depends generally on the thickness of the ice layer
514 (Montpetit et al., 2013, e.g.). When we include such an ice layer in the SMP-
515 based modelling, the modelled data (Figure 7 bottom row) shows relative changes
516 (increase in PR(19) and PD89) that are comparable to the ones observed from
517 satellite, see Figure 8. GR(37/19) is hardly effected since it is based on vertically-
518 polarized brightness temperatures.

519 3.4. *Site perspective: ground-based radiometers*

520 We now adopt the site perspective and investigate the ground-based radiometer
521 measurements.

522 3.4.1. Brightness temperatures from ground-based radiometers

523 The brightness temperatures of the ground-based radiometers between 15 and 20
524 April 2020 are summarized in Figure 9. The data is smoothed applying an one
525 hour running mean. The analyzed period can be divided into four phases. The first
526 phase is the period from April 15 to April 17. In this period the first warm air in-
527 trusion hit the MOSAiC site. Temperatures are rising from -12°C to -2°C at 2 m
528 height. Increased wind speed and (wet) snowfall led to changes of the snow cover.
529 This is evident at, e. g., 6.9 GHz horizontal polarization, fluctuations indicate here
530 that surface properties in the FoV of the radiometer changed and that there might

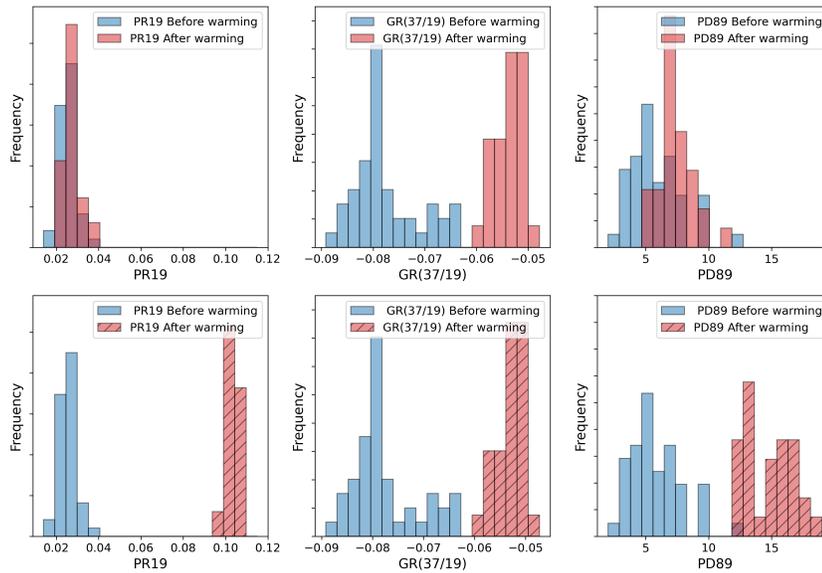


Figure 7: Histogram of simulated PR(19), GR(37/19) and PD89 based on 84 the SMP profiles. The data is split into the periods from April 08 - April 15 (blue) and April 21 - April 27 (red). In the bottom panel, a thin ice layer (2 mm) was added on top of the snow (red, hatched) to simulate the effect of surface glazing after the second warming wave.

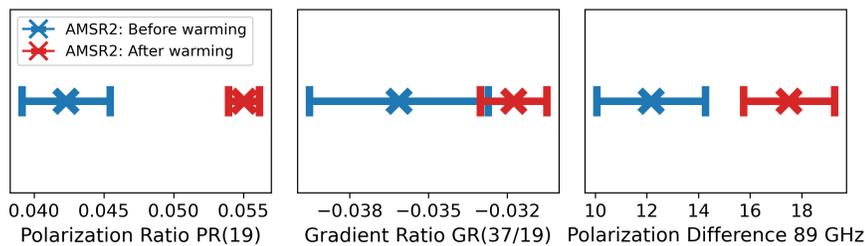


Figure 8: Mean and standard deviation of PR(19), GR(37/19) and PD89 derived from AMSR2 brightness temperatures collocated to *Polarstern* from Figure 4. The data is split into the periods from April 08 - April 15 (blue) and April 21 - April 27 (red) matching the modeled periods shown in Figure 7.

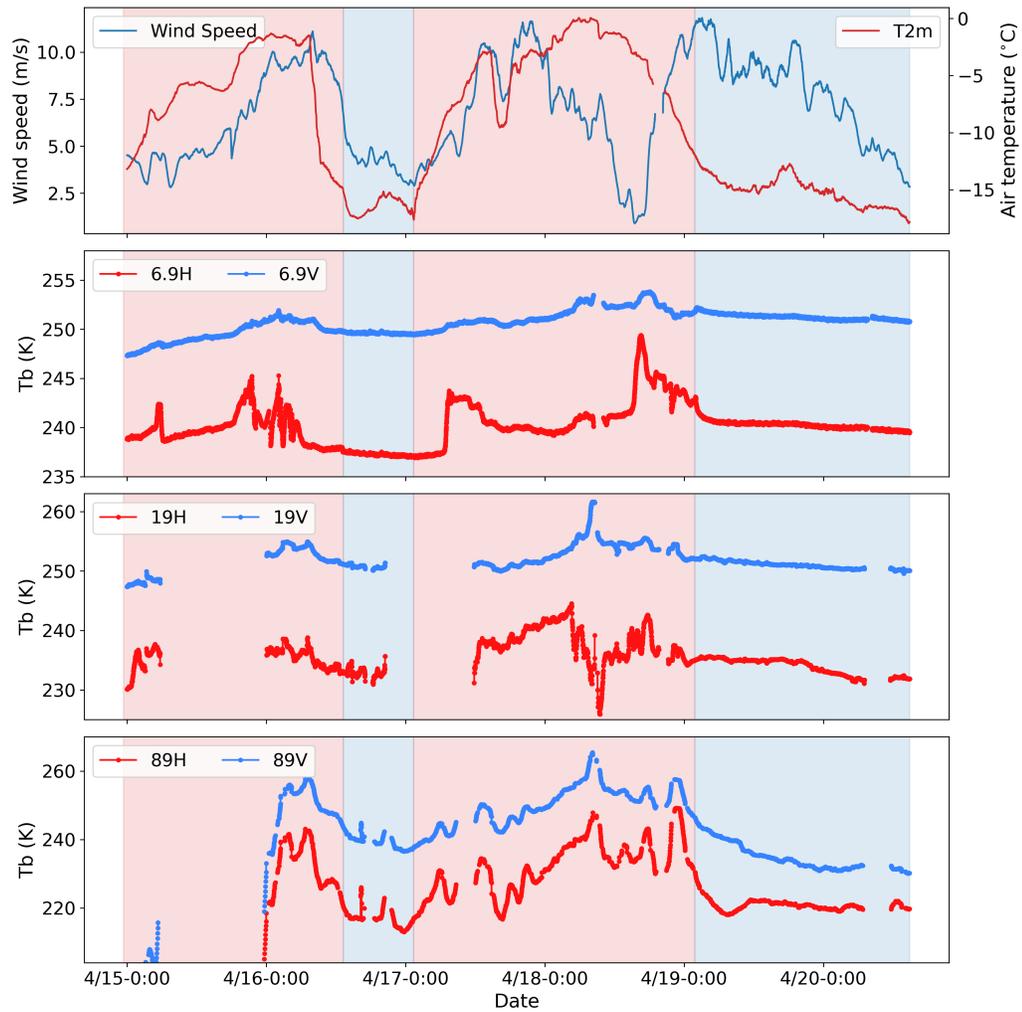


Figure 9: Top: 2 m air temperature and 10 m wind speed during the warm air intrusions. Rows 2 – 4: Observed brightness temperatures for vertical (V) and horizontal (H) polarizations for 6.9 GHz, 19 GHz, and 89 GHz, respectively. Background shading indicate the four different phases for cloudy (red) and clear sky (blue) conditions.

531 be already some low amount of liquid water building in the uppermost layer of
532 the snowpack. While the average brightness temperature at lower frequencies (6.9
533 to 19 GHz) is only slightly increasing, a strong increase can be found at 89 GHz,
534 indicating that mainly the temperature of the upper snow layer changed (89 GHz
535 has the lowest penetration depth), which is consistent with Figure 5. Strong short-
536 term fluctuations at 89 GHz are captured well in the simulations and are likely due
537 to fluctuations of reflected downwelling radiation caused by increased cloud cover
538 and higher LWP.

539 The second phase, between April 17 and April 18 (first blue shading in Fig-
540 ure 9) is marked by a rapid cooling below -15°C , clear sky and calm conditions.
541 Brightness temperatures at all frequencies are stable during this phase with a slight
542 cooling at 19 GHz and a strong drop in brightness temperature at 89 GHz.

543 The third phase marks the second, stronger warm air intrusion where the 2 m
544 air temperature reached 0°C . Strong changes for horizontally-polarized TB at all
545 frequencies indicate changes in the snow surface due to snow fall and drift as
546 well as accumulation and possible formation of liquid water. However, during
547 this period, increased snow accumulation around the instruments complicate the
548 interpretation of the data at the remote sensing site as the exact timings of snow
549 dune formation are unknown.

550 Of special interest is the fourth phase, the period right after the second warm-
551 ing event, where a strong decline in brightness temperatures is observed at higher
552 frequencies. At 89 GHz, TB H drops by around 30°C within 6.5 hours. For vertical
553 polarization, this decrease is less pronounced.

554 Consequently, PD at 89 GHz shows a strong increase during this period, which
555 is consistent with the increase observed by the satellites (Figure 4). At the lower
556 frequencies, this drop in TB H is less pronounced. In contrast to the satellite ob-
557 servations, the PD 89 of the ground-based observations recovers after less than
558 one day and thus much faster than the satellite measurements. A likely explana-
559 tion is the accumulation of snow in front of the ground-based radiometers, which
560 does not occur on the satellite scale. This already highlights the need of auxiliary
561 data when using ground-based measurements to interpret satellite data due to the
562 local snow conditions. Unlike the snow drifts that accumulated in front of the in-
563 struments, most of the level ice on the floe-scale did not experience accumulation
564 during this event.

565 3.4.2. Model experiment on-ice site: simulation of a glaze ice layer

566 Similar to the simulations of the SMP profiles (Section 3.3.2), we perform an
567 experiment with simulating the effect of a glaze ice layer. In the model, the glaze

568 ice layer is approximated by a thin, radiometrically flat ice layer (≤ 1.6 mm) at
569 top of the snowpack. The setup is chosen such that the model can reproduce the
570 observed PD changes from the ground-based radiometer at 6.9 GHz and 89 GHz
571 during phase 4, i. e., the clear sky phase after the second warm air intrusion (right
572 blue shading in Figure 9). As discussed earlier, during this phase, the drop in PD
573 at 89 GHz cannot be explained by cloud forcing and is most likely due to the
574 formation of a glaze ice layer in the FoV of the radiometer.

575 For the experiment we simulate an ice layer that starts developing at April 20
576 at around 13:00 and grows up to 1.6 mm until 17:00. After that we allow a layer
577 of new snow to accumulate on top of this ice layer. We know of the increased
578 snow accumulation in front of the instruments due to snow drift formation as
579 wind speed was high during this period (Figure 9), however the exact timing of
580 new snow accumulations remains a reasonable assumption. The reflectance data
581 from the TLS scan on April 22 shows that a glaze ice layer is not visible anymore
582 in the FoV of the radiometers by this date. Figure 10 shows the observed and
583 simulated brightness temperatures between April 20 10:00 and April 21 10:00.

584 With this set-up it is possible to reproduce the observed increase of polariza-
585 tion difference at 89 GHz indicating that indeed the formation of a glaze ice layer
586 most likely led to the strong increase in observed PD 89.

587 In summary, the temporal development of microwave brightness temperatures
588 measured by the on-ice radiometers and especially their polarization difference
589 can be explained if a thin glaze ice layer is added in the microwave emission
590 model. Such a glaze layer actually was observed in the field. The effect of the
591 glaze ice layer is larger at higher frequencies and mainly affect the polarization
592 difference.

593 4. Discussion

594 The effects of the warm air intrusions on sea ice concentration retrievals are man-
595 ifold, but we believe we have identified dominant mechanisms for the case study
596 presented here. First, the changed temperature gradients in the snow and snow
597 metamorphism (increase in correlation length). This influences SIC estimates that
598 rely on the gradient ratio $GR(37/19)$. The increase in $GR(37/19)$ is due to the
599 changed snow temperature (gradient) and is not fully compensated for by snow
600 metamorphism (which decreases $GR(37/19)$). Second, we attribute the strong in-
601 crease of satellite-measured PD after the warm air events to an ice glazing of the
602 snow surface that was present on a large scale. Such glazing was observed in the
603 field and model results suggest that it can explain the observed satellite microwave

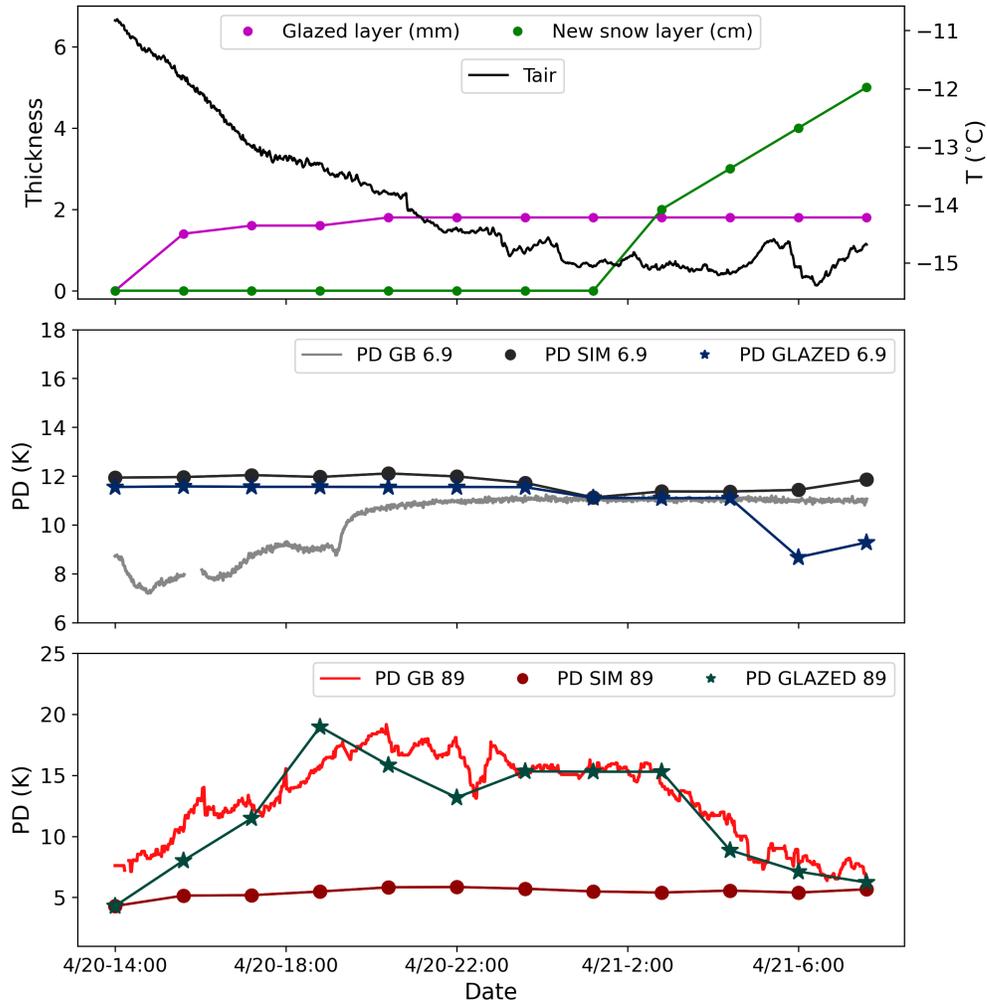


Figure 10: Model experiment adding a glaze ice layer on top of the snowpack covering parts of phase 4 after the second warm air intrusion (2nd blue shading in Figure 9). The top panel shows 2 m air temperature and the glazed and snow layer setup of the experiment. Panel two and three show the observed (solid) and simulated brightness temperatures without (dotted) and with (stars) the glazed layer for PD6.9 (middle) and PD89 (bottom).

604 brightness temperatures. The higher PD causes a decrease in SIC in some sea ice
605 concentration products that we investigate in this study (significantly for the al-
606 gorithm that mainly rely on PD89, as ASI does, or PR(19) as NASA TEAM, but
607 also visible for algorithms like OSI-SAF that use the TB space spanned by both
608 polarizations at 36.5 GHz as part of their retrieval; see Section 2.1).

609 We note a stronger response of the microwave emissions to the warming events
610 at higher frequencies that have smaller penetration depths into the snow. A thin
611 ice crust also effects these frequencies more while the wavelengths of the lower
612 frequencies are large compared to the thickness of the crust. SIC retrievals us-
613 ing vertical polarization or ones that are based on lower frequencies like 6.9 GHz
614 are less affected. For example, the dual algorithm approach of the NSIDC-CDR
615 largely mitigates the glazing impact in this case, as the Bootstrap algorithm gives
616 a high value for SIC. While this is an advantage for the case presented here, this
617 approach might also overestimate sea ice concentration in other situations. For
618 example, the North East Water polynya at the Greenland coast that opened af-
619 ter the events (Figure 2, middle column, first three rows), is hardly visible in the
620 NSIDC-CDR product.

621 Hypothesizing that such glazing events should increase in a warming Arctic,
622 as not only warm air intrusions but also rain on snow events can cause them, the
623 choice of low frequency / vertical polarization algorithms could be a more robust
624 choice for the future. In the study by Rees et al. (2010) about an ice crust on snow
625 on land 6.9 GHz showed the least response to the ice crust as well. We note that,
626 different to our case, PD at 19 GHz was affected most in their study, which is
627 likely due to a different thickness of the ice crust and different snow conditions.

628 Modeling snow and ice microwave emission remains a challenge especially at
629 satellite-footprint scale, even if a large amount of observational data is available.
630 Partly this is due to the local heterogeneity making it difficult, e. g., to match ra-
631 diometer observations and ground-based observations. For example, we did not
632 only observe differences in the snow cover between the ground-based radiometer
633 footprints at the same site but also within one footprint (see supplemental Fig-
634 ure S1). We overcome this problem by using the vast amount of snow profiles,
635 thanks to the SMP measurements, in combination with modelling, providing a sta-
636 tistical description of expected surface brightness temperatures (Figure 7). Still,
637 if the radiometers measure snow conditions that are not representative of the sur-
638 roundings (due to snow accumulation), comparisons to these TB measurements
639 are impeded. The statistical description might be more comparable to a satellite
640 observation in case of negligible atmospheric effects. However we cannot be cer-
641 tain that the SMP data are representative of a satellite footprint of the order of tens

642 of kilometres. Also, certain parameters, such as surface roughness that are impor-
643 tant for emission modeling of satellite observations, are not directly available. Co-
644 herence effects, that depend on frequency and layer thickness, are another source
645 of uncertainty. Nevertheless, the temporal evolution of the ground-based radiome-
646 ter measurements can be reproduced by a microwave emission model if the glaze
647 ice layer is included. Similarly, including a glaze layer allows us to qualitatively
648 model relative changes as observed from space using the SMP measurements as
649 input. This is one of the few existing cases where a false change in satellite SIC
650 can be fully explained by the observed surface changes from ground-based mea-
651 surements (snow and ice physics and radiometers). We are confident that adopting
652 the three perspectives from the different scales (satellite, floe, site) allowed for a
653 plausible interpretation of the observations.

654 5. Conclusion

655 Arctic amplification, i. e., the more rapid and stronger increase of temperatures
656 in the Arctic compared to low latitudes (Serreze and Barry, 2011; Screen and
657 Simmonds, 2010; Comiso and Hall, 2014; Wendisch et al., 2017, 2023), can lead
658 to an increased occurrence of warm air intrusions (Graham et al., 2017).

659 As shown in this study, they can affect brightness temperatures measured by
660 satellite microwave radiometers and cause uncertainties in the derived sea ice con-
661 centration products. In our case, the warm air intrusions led to a large-scale spuri-
662 ous strong decrease in SIC and an increase in the deviations between different SIC
663 products. These deviations lasted for several days. Only one climate data record
664 was little affected. It, however, has the tendency to always produce high SIC val-
665 ues, which was the correct solution in this case.

666 We offer an interpretation of the signals based on observations from the MO-
667 SAiC expedition and by microwave emission modeling, taking into account atmo-
668 spheric effects as well as surface snow metamorphism. As an explanation for the
669 changed microwave emissions during and after the warming events, we propose
670 the formation of a large-scale glaze ice layer, which is persistent even days after
671 the warm air intrusion.

672 Many recent sea ice concentration studies, (Lu et al., 2022, e. g.), focus on the
673 influence of the atmosphere on ice concentration. In our case, the surface changes
674 are highly relevant and should be included in future evaluations of sea ice con-
675 centration retrieval algorithms. Especially, near 90 GHz algorithms were thought
676 to be less sensitive to these surface effects like ice lenses within the snow because
677 of their small penetration depth compared to the lower frequencies (Ivanova et al.,

678 2015). However, as shown in this study, certain surface effects like glazing can
679 have a strong influence on these algorithms. Identifying similar events, their scale
680 in time and space and their frequency of occurrences can provide additional in-
681 sights to quantify whether this effect is significant on longer temporal and spatial
682 scales and for climate data records. This is also interesting for past sea ice concen-
683 tration datasets and inter-comparison studies of sea ice concentration algorithms.
684 In the future, due to a projected increase in such warm air intrusions, the rele-
685 vance of their effects on sea ice climate records from satellites and the distinction
686 between actual influence of climate warming on sea ice, as described by, e. g.,
687 Merkouriadi et al. (2020), and retrieval uncertainties will become more important.
688 Regarding the estimation of product uncertainties this needs consideration as well,
689 possibly by using dynamic uncertainty estimates, i. e., uncertainties that are nei-
690 ther constant nor only dependent on SIC. A good example is the OSI-SAF iCDR
691 product that provides such estimates which show an increased uncertainty after
692 the events.

693 The detailed MOSAiC measurements allowed for a sophisticated emission
694 model setup to simulate the ground-based radiometer data that enabled us to in-
695 terpret the observed satellite changes. When considering ice concentration algo-
696 rithms based on physical forward models, however, it is not yet feasible to fully
697 account for the variability of surface emissions due to the high variability and
698 amount of needed input parameters as well as model physics limitations.

699 Instead, multi-frequency methods exploiting the synergy of the robustness
700 of 6.9 GHz and the high spatial resolution of 89 GHz are a promising approach
701 for future retrievals. Upcoming satellite missions like the Copernicus Imaging
702 Microwave Radiometer (CIMR) (Donlon, 2020) will provide measurements at
703 6.9 GHz at a much higher spatial resolution (around 15 km) than current satellite
704 sensors, which makes it well suited for sea ice concentration retrievals at higher
705 spatial resolution (5 km at 37 GHz) and higher accuracy (using 6.9 GHz) than what
706 is available today.

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1188 Contributions

1189 Janna E. Rückert and Philip Rostosky contributed equally to this work. First de-
 1190 scription and interpretation of event: LK. Conception and design of this study:
 1191 JR, PR, MH, GS. Acquisition of data: MH, GS, LK, AM, RN, JS. Analysis and
 1192 interpretation of data: JR, PR, MH, DCS, KE, LK, JL, AM, AW, GS. Drafting,
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1218 Competing interests

1219 The authors have declared that no competing interests exist.

1220 6. Supplemental material

1221 6.1. *SI - MEMLS_ice initialization*

1222 MEMLS_ice is a microwave emission model for layered snow and ice to simu-
1223 late the upwelling brightness temperatures of a snow-ice system in the microwave
1224 regime, based on its dielectric properties. The model needs several snow and ice
1225 properties as input data. For the snow, most of these properties like density, expo-
1226 nential correlation length or temperature can be obtained from SnowMicroPen and
1227 snow pit samples, while some parameters like snow wetness or interface rough-
1228 ness have to be assumed. The roughness is modelled by the modification of the
1229 interface reflectivity following Choudhury et al. (1979). In principle, a glazing
1230 layer can lead to coherent reflections in the microwave regime. However, as we
1231 have not enough information about the thickness and uniformity of that ice layer,
1232 we do not calculate coherent reflections in our modelling of the glazed layer. Note
1233 that coherence is in principle implemented in MEMLS but deactivated for our
1234 purposes. During the warm air intrusion events, one ice core sample was taken at
1235 April 17 in the vicinity (i.e., $\mathcal{O}100$ m) of the remote sensing site. We use this ice
1236 core as first guess for the sea ice properties and then slightly varied the salinity
1237 and density profiles to obtain brightness temperatures at 6.9 GHz that are similar

1238 to the ground-based radiometer observations before the warm air intrusion. For
1239 all layers, the deviations from the ice core samples are less than 1 ppt for salinity
1240 and less than 50 Kgm^{-3} for the density. Also note that the Central Observatory
1241 was on second-year ice (with a low amount of salinity in the upper layers) at the
1242 beginning of the expedition in 2019 but that there was new ice formation during
1243 the winter. On the satellite footprint scale a mixture of second- (dominating) and
1244 first-year ice prevailed. Except for the ice temperature, for which a linear fit is
1245 used between the snow/ice and ice/ocean interface temperatures, the ice profile
1246 properties are kept constant for the whole time series. Since the warm air intru-
1247 sions were short-term events and the temperature at the snow-ice interface only
1248 slightly changed, this is a valid assumption. Table S1 shows the initialization pro-
1249 file for April 08 07:40 based on the SMP profile S49M2235 (Macfarlane et al.,
1250 2021) for 6.9 GHz. At the microwaves frequencies investigated here, only the up-
1251 per ice layer is relevant due to the high absorption in the ice. Therefore, only the
1252 upper 40 cm of the ice are shown (bottom five rows in Table S1). The model was
1253 initialized with 2 m thick ice. Thus there is no influence from the ocean below.
1254 The downwelling sky temperature is calculated with the radiative transfer model
1255 PAMTRA (Mech et al., 2020) using the HATPRO TWV and LWP data as well as
1256 profiles from radiosondes launched during MOSAiC.
1257 Note that the snow height from the SMP profiles was usually lower than the snow
1258 in the field of view of the radiometers due to enhanced snow accumulation. The
1259 TLS scan data provides accurate relative differences of snow height. In order
1260 to obtain absolute values, we use the measurement of the maximum snow dune
1261 height with a ruler stick on April 22, namely 90 cm as reference point. The ac-
1262 quired snow depths are in agreement with other manual ruler stick measurements
1263 of snow height performed close to the radiometer footprints. Table S2 shows the
1264 snow height at the different sensors from TLS scans from April 17 and April 22.
1265 Figure S1 shows the surface elevation at April 22 based on a TLS scan. In ad-
1266 dition, the FoV of the different frequencies of the ground-based radiometers is
1267 shown. High snow dunes were forming during the second warm air intrusion in
1268 front of the radiometers.

1269 6.2. *S2 - SMP-based modelling*

1270 Figure S2 shows the simulated brightness temperatures at different frequencies for
1271 the time period before the warming events (April 08 to April 15, blue) and after
1272 the warming events (April 21 - April 27, red). For the simulations, the SMP pro-
1273 files were averaged to layers of at least 2 cm thickness. To maintain the layering

Table S1: Initialization profile for the MEMLS_ice simulation for April 08 07:40 for 6.9 GHz. Shown are the parameters: layer thickness LT (cm), Temperature T (K), Density ρ (Kg m^{-3}), exponential correlation length pc (mm), wetness W, salinity S (ppt), interface roughness R (mm). Parameters used for fitting the simulations are shown in *italics*. The downwelling sky temperature at 6.9 GHz is 6.1 K

Depth	LT	T	ρ	pc	W	S	R
0	2.1	250.1	269.8	0.292	0	0	0
2.1	3.3	250.9	374.6	0.344	0	0	0
5.4	4.0	251.9	311.5	0.298	0	0	0
9.4	2.8	252.5	299.0	0.271	0	0	0
12.2	2.0	252.8	287.7	0.271	0	0	0
14.2	3.0	253.0	287.7	0.271	0	0	0
17.2	2.1	252.9	312.8	0.299	0	0	0
19.3	3.0	252.8	264.6	0.203	0	0	0
22.3	8	253.0	<i>880</i>	<i>0.450</i>	0	<i>3.08</i>	<i>2.20</i>
30.3	8	253.5	<i>880</i>	<i>0.450</i>	0	<i>2.93</i>	0
38.3	8	253.2	<i>880</i>	<i>0.450</i>	0	<i>2.80</i>	0
46.3	8	253.4	<i>880</i>	<i>0.450</i>	0	<i>2.66</i>	0
54.3	8	253.6	<i>880</i>	<i>0.450</i>	0	<i>2.50</i>	0

Table S2: Average snow height (cm) in the FoV of the different channels (GHz) at April 17 and April 22

Date	6.9	10.7	19	89
April 17	48	46	45	44
April 22	48	53	46	71

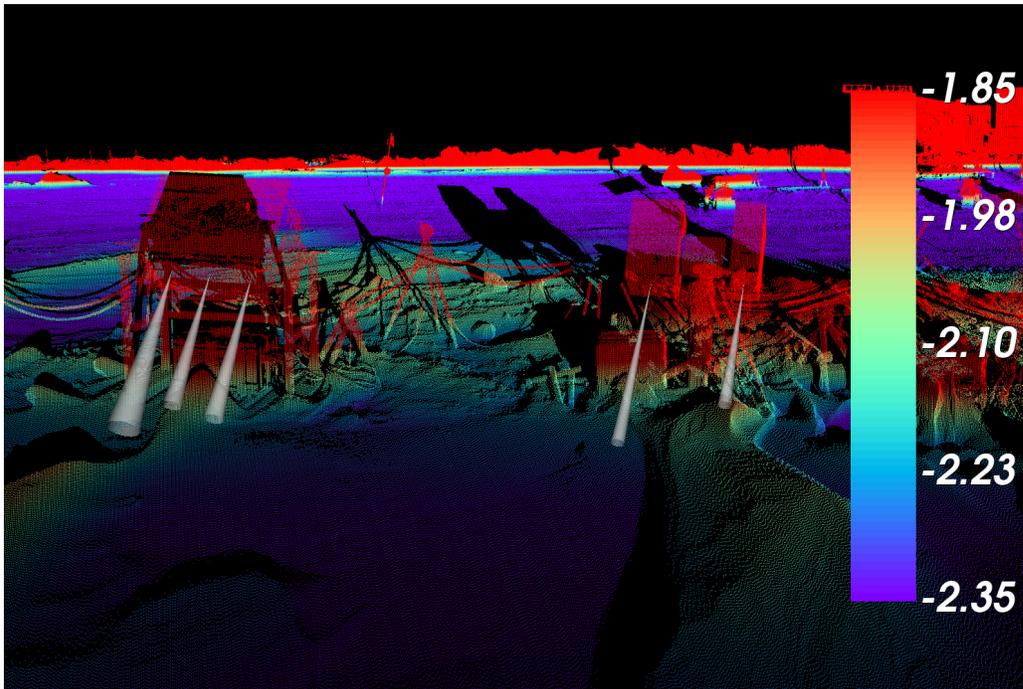


Figure S1: TLS surface height with respect to the height of the laser scanner (2.5 m) from 22 April. The cones indicate the viewing direction if the different sensors. From left to right: HUTRAD frequencies (GHz): 10.7, 18.7 (not used), 6.9, SBR frequencies (GHz): 19, 89.

1274 in the snowpack, snow type classifications provided by Kaltenborn et al. (2022)
 1275 were used for the initial step. Strong changes in simulated brightness temperatures
 1276 are visible at all frequencies. At 6.9 GHz and 19 GHz, brightness temperatures are
 1277 higher after the warming events caused by a generally higher snowpack tempera-
 1278 ture. In contrast, at 37 GHz, increased scattering due to snow metamorphism leads
 1279 to a decrease of brightness temperatures. Note that no glazed layer was used in
 1280 these simulations.

1281 6.3. S3 - Glazed layer observations

1282 Figure S3 shows the TLS reflectance for April 17 and April 22. Glazed areas have
 1283 low TLS reflectance (dark purple colors) in the near-IR due to increased absorp-
 1284 tion and specular reflections. On April 17, only a few patchy glazed areas are

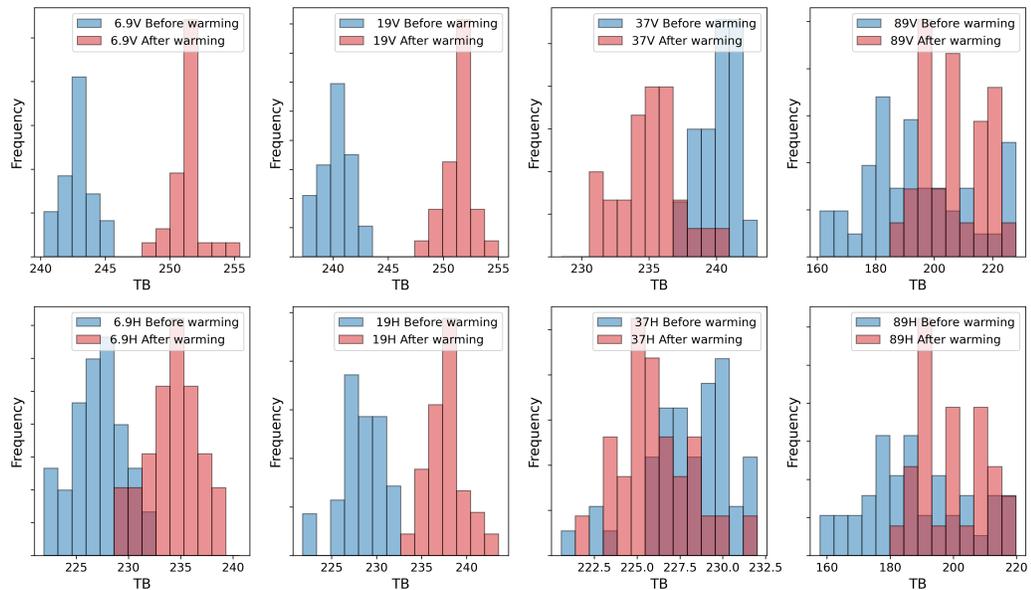


Figure S2: Simulated brightness temperature from SMP profiles. The data is split into the periods from April 08 - April 15 (blue) and April 21 - April 27 (red).

1285 present (e. g., the dark purple patch in the upper right) and none are within the in-
 1286 struments' footprints. On April 22, widespread glazing was observed on unaltered
 1287 snow surfaces (e. g., upper portion of figure). However, the snow drifts that formed
 1288 in the instruments' footprints were not glazed. The glazing was also evident from
 1289 visual observations. Figure S4 shows the glazed snow surfaces observed on April
 1290 22 at the remote sensing site (Figure S4, left). The snow in front of the radiome-
 1291 ters is freshly accumulated. Also on the panorama camera mounted on *Polarstern*,
 1292 the glazing was visible at April 22 (Figure S4, right).

1293 6.4. S4 - Ground-based radiometer observations

1294 We focus on two microwave radiometers deployed during MOSAiC observing
 1295 the surface at frequencies ranging from 6.9 GHz to 89 GHz, similar to AMSR2.
 1296 Details are provided in Table S3. They were installed next to each other on the
 1297 ice facing the same area but the footprints do not overlap (see Figure S1). The
 1298 low-frequency system of HUTRAD (Helsinki University of Technology Radiome-
 1299 ter) (Hallikainen et al., 1996; Colliander et al., 2007; Lemmetyinen et al., 2009)
 1300 measures at three frequencies at 6.825, 10.65 and 18.7 GHz, at two orthogonal

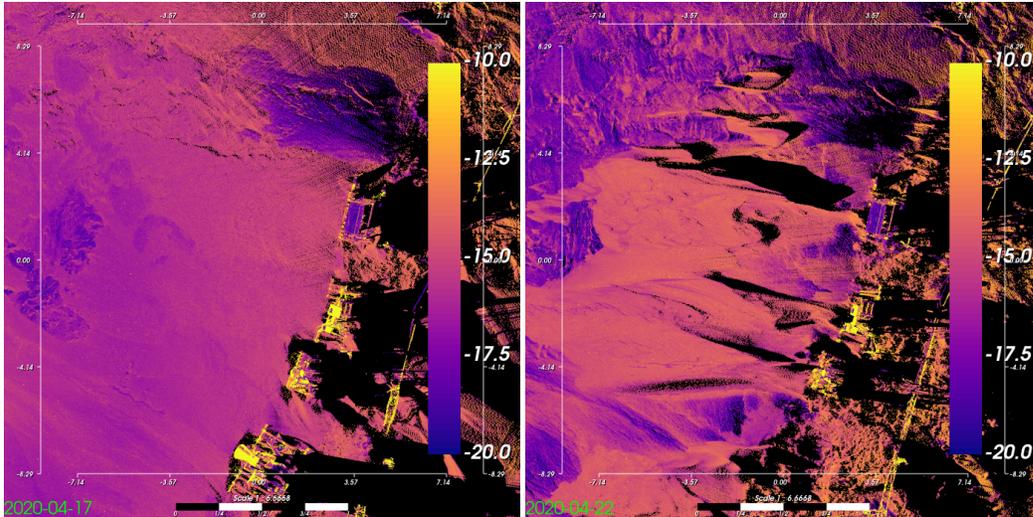


Figure S3: Glazed snow surfaces observed in Terrestrial Laser Scanning (TLS) reflectance near the HUTRAD and SBR instruments on April 17 (left) and April 22 (right). Scale bars are in meters.



Figure S4: Left: Image from April 22 at the remote sensing site (photo: Lars Kaleschke). Right: panorama image from Panomax webcam onboard *Polarstern* (<https://www.mosaic-panorama.org/>) from April 13 and April 22.

1301 polarizations H and V. During MOSAiC, HUTRAD had a sampling rate of 1 s,
1302 the dataset used here was averaged to 1 minute temporal resolution in order to de-
1303 crease fluctuations. HUTRAD was observing at a fixed incident angle of 45° . With
1304 an instrument height of 0.8 m the largest footprint size (at vertically-polarized
1305 6.9 GHz) is $30\text{ cm} \times 42\text{ cm}$. The radiometer was deployed at a fixed location dur-
1306 ing the whole period investigated. The uncertainty (i.e., $2 \times$ standard deviation)
1307 of the calibrated brightness temperature is estimated around 3 K. It is to note that
1308 the internal temperature control of the 18.7 GHz receiver failed early on in the
1309 MOSAiC campaign, which made measurements unstable and unusable for our
1310 investigation.

1311 The surfaced based radiometer from University of Manitoba (SBR) radiome-
1312 ter observes at 19, 37 and 89 GHz (Radiometrics, 2004). Radiometric properties
1313 are provided in Table S3, for more detailed specifications see Table 1 in Derk-
1314 sen et al. (2012). The SBR was measuring in scanning-mode with incident angle
1315 varying from 40° to 70° in 5° steps. The sampling rate varies between 1 s and 5 s.
1316 For this study, we only use the measurements at 55° , since most data was col-
1317 lected at this angle. The footprint size is about $20\text{ cm} \times 35\text{ cm}$ (assuming 1.1 m
1318 instrument height). Similar to HUTRAD the SBR data is resampled to 1 minute
1319 temporal resolution. The uncertainty of the observations at 89 GHz is similar to
1320 the lower frequencies (6 K for horizontally- and 5 K for vertically-polarized TB).
1321 Unfortunately, during leg 3, the 37 GHz receiver was not functioning.

1322 As the 18.7 GHz horizontally-polarized channel from HUTRAD was unstable
1323 during our period of interest, we show the values measured by SBR which mea-
1324 sures at 19 GHz. During the first warming event some values from SBR (between
1325 two calibrations) had to be removed because they show unrealistic fluctuations.

1326 Corrupted scans and data points with clearly non-physical values were filtered
1327 out manually for Figure 9.

1328 Figure S1 illustrates the formation of snow dunes in front of the radiometers
1329 that further increase the local (intra- and inter-footprint) variability of the snow
1330 conditions.

1331 Data accessibility statement

1332 The AMSR2 satellite data is available via Japan Aerospace Exploration
1333 Agency's G-Portal <https://gportal.jaxa.jp/gpr/?lang=en>. The
1334 NOAA/NSIDC Climate Data Record of Passive Microwave Sea Ice Concen-
1335 tration, Version 4 is available via [https://nsidc.org/data/g02202/](https://nsidc.org/data/g02202/versions/4)
1336 [versions/4](https://nsidc.org/data/g02202/versions/4). The OSI-430-b SIC product data is available via <https://>

Table S3: HUTRAD and SBR radiometric properties: Shown are frequency (GHz), polarization P, radiometric bandwidth BW (MHz) and sensitivity (K), incidence angle θ (degree), field of view FoV given by the half power beam width (degree) and the FoV on the ground (cm) based on an instrument height of 0.8 m (HUTRAD) and 1.1 m (SBR).

Frequency (GHz)	HUTRAD				SBR	
	6.9		10.7		19	89
Polarization	H	V	H	V	H/V	H/V
Bandwidth (MHz)	310	310	120	120	1000	4000
Sensitivity (K)	0.09	0.11	0.24	0.22	0.04	0.08
θ ($^{\circ}$)	45	45	45	45	55	55
Field of View ($^{\circ}$)	11.2	14.8	6.6	9.1	6	5.88
Field of View (cm)	22×32	30×42	13×19	18×26	20×35	20×35

1337 `osi-saf.eumetsat.int/` (copyright 2021 EUMETSAT). The radiometer,
 1338 snow, temperature, laser scan and radiosonde data used in this manuscript was
 1339 produced as part of the international Multidisciplinary drifting Observatory for
 1340 the Study of the Arctic Climate (MOSAiC) with the tag MOSAiC20192020
 1341 and the Project ID: AWI PS122 00. The HUTRAD radiometer data is avail-
 1342 able via PANGAEA <https://doi.org/10.1594/PANGAEA.954608>.
 1343 The radiosonde data is available via PANGAEA [https://doi.org/10.](https://doi.org/10.1594/PANGAEA.928656)
 1344 [1594/PANGAEA.928656](https://doi.org/10.1594/PANGAEA.928656). The HATPRO LWP data product is available via
 1345 PANGAEA <https://doi.org/10.1594/PANGAEA.941389>. The near-
 1346 surface temperature measurements are available from the Arctic Data Center
 1347 (Cox et al., 2021). The Present Weather Detector 22 precipitation gauge data is
 1348 accessible via ARM: [https://adc.arm.gov/discovery/#/results/](https://adc.arm.gov/discovery/#/results/instrument_code::pwd)
 1349 [instrument_code::pwd](https://adc.arm.gov/discovery/#/results/instrument_code::pwd), last access: 07 November 2022. The SBR radiome-
 1350 ter is available via PANGAEA [https://doi.pangaea.de/10.1594/](https://doi.pangaea.de/10.1594/PANGAEA.956108)
 1351 [PANGAEA.956108](https://doi.pangaea.de/10.1594/PANGAEA.956108) (dataset in review). The TLS data is available via the Arc-
 1352 tic Data Center (Clemens-Sewall, 2021). ERA5 data are made available by
 1353 the Copernicus Climate Change Service (C3S) at [https://cds.climate.](https://cds.climate.copernicus.eu/cdsapp#!/home)
 1354 [copernicus.eu/cdsapp#!/home](https://cds.climate.copernicus.eu/cdsapp#!/home). C3S (2017): ERA5. Fifth generation of
 1355 European Centre for Medium-range Weather Forecasts atmospheric reanalyses
 1356 of the global climate. Copernicus Climate Change Service Climate Data Store,
 1357 2017–2020.