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Spatio-temporal variability of small-scale leads based on helicopter winter sea ice surface temperatures

Linda Thielke<sup>1</sup>, Gunnar Spreen<sup>1\*</sup>, Marcus Huntemann<sup>1</sup>, Dmitrii Murashkin<sup>2,1</sup>

<sup>1</sup>University of Bremen, Institute of Environmental Physics, Bremen, 28359, Germany

<sup>2</sup>German Aerospace Center (DLR), Remote Sensing Technology Institute (IMF), Bremen, Germany

\*Corresponding author: [gunnar.spreen@uni-bremen.de](mailto:gunnar.spreen@uni-bremen.de)

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3 Linda Thielke<sup>1</sup>, Gunnar Spreen<sup>1\*</sup>, Marcus Huntemann<sup>1</sup>, Dmitrii Murashkin<sup>2,1</sup>

4 <sup>1</sup>University of Bremen, Institute of Environmental Physics, Bremen, 28359, Germany

5 <sup>2</sup>German Aerospace Center (DLR), Remote Sensing Technology Institute (IMF), Bremen,  
6 Germany

7 \*Corresponding author: [gunnar.spreen@uni-bremen.de](mailto:gunnar.spreen@uni-bremen.de)

8 **Abstract**

9 Surface temperature is crucial in studying the Arctic climate, particularly  
10 during winter. We examine 1 m resolution surface temperature maps of 35  
11 helicopter flights between 02 October 2019 and 23 April 2020, recorded during  
12 the Multidisciplinary drifting Observatory for the Study of Arctic Climate  
13 (MOSAIC). The seasonal cycle of the average surface temperature spans from  
14 265.6 K on 02 October 2019 to 231.8 K on 28 January 2020. The surface  
15 temperature is affected by atmospheric changes and also varies across scales.  
16 Furthermore, we concentrate on leads in sea ice because they allow for greater heat  
17 exchange between ocean and atmosphere than thick, snow-covered ice. Leads,  
18 which appear considerably warmer than sea ice, are classified by a temperature  
19 threshold. The local scale (5–10 km) lead area fraction varies between 0% and  
20 4% with a higher variability than on a regional scale (20–40 km), where leads  
21 cover a more stable fraction of 0-1% until mid-January when it increases to 4%.  
22 The variability in the lead area is caused by sea ice dynamics (opening and  
23 closing of leads), as well as thermodynamics with ice growth (lead closing).  
24 To understand better the ice rheology throughout the winter, we identify lead  
25 orientation distributions. We find that the orientation varies between different  
26 flights but the distribution mostly shows one prominent orientation peak. Thus,  
27 we are not able to determine predominant intersection angles, which would need  
28 two modes in the orientation distribution. The lead width distribution follows a  
29 power law with a negative exponent of 2.63, which agrees with literature values,  
30 proves the comparability to other datasets, and extends the existing relationship  
31 to the smaller scales, as observed here. The appearance of many more small leads

32 compared to wider leads is important since they only occur on the sub-footprint  
33 scale of thermal infrared satellite data. Sub-satellite-footprint lead statistics are  
34 essential for Arctic-climate investigations because the ocean-atmosphere heat  
35 exchange does not scale linearly with lead area fraction and is larger for smaller  
36 leads.

## 37 1. Introduction

38 This study presents the spatio-temporal evolution of the Arctic sea ice surface  
39 temperature and lead area fraction, as well as the lead width and intersection  
40 angle. In this analysis, we refer to fractures in the sea ice cover like cracks and  
41 leads (>50 m width according to the definition of the World Meteorological  
42 Organization (WMO), (WMO, 2014)), jointly as "leads". The helicopter-borne  
43 surface temperature measurements were taken as part of the Multidisciplinary  
44 drifting Observatory for the Study of Arctic Climate (MOSAiC) in the central  
45 Arctic (Shupe et al., 2020). The MOSAiC expedition (Sep 2019–Oct 2020)  
46 allowed us to collect in-situ measurements from the central Arctic over a whole  
47 seasonal cycle for different aspects of the Arctic system (Nicolaus et al., 2022;  
48 Rabe et al., 2022; Shupe et al., 2022). Our measurement program was part of the  
49 sea ice and remote sensing teams (Nicolaus et al., 2022), which conducted a large  
50 collection of data from sea ice physics, on-ice remote sensing, over albedo, to  
51 snow properties. The analysis is based on data from 35 helicopter survey flights  
52 between October 2019 to April 2020, recorded with an infrared camera over the  
53 same ice floe and surrounding regions along the Transpolar Drift.

54 The investigation of sea ice processes is crucial for studying climate warming,  
55 which is especially strong in the high latitudes (Arctic Amplification) (Serreze  
56 and Barry, 2011; Wendisch et al., 2017; Dai et al., 2019; Masson-Delmotte et al.,  
57 2021). The warming is even stronger in winter than in summer, related to the  
58 feedbacks of infrared (IR) radiation in winter and ice-albedo during summer  
59 (Bintanja and Van Der Linden, 2013). Sea ice becomes significantly thinner  
60 (Meredith et al., 2019; Masson-Delmotte et al., 2021) with an average reduction of  
61 2 m from the period 1958-1976 (submarine record) to the current altimeter period  
62 with strongest thinning during the ICESat period (2003-2008) (Kwok, 2018). With  
63 the decline in annual sea ice minimum extent in late summer, also the multiyear ice  
64 area has strongly decreased (Kwok, 2018). The thinner ice makes the sea ice more  
65 susceptible to wind and ocean current forcing, resulting in higher ice drift speeds  
66 (Spren et al., 2011; Kwok et al., 2013). Rampal et al. (2009) hypothesizes that  
67 thinner sea ice has less mechanical strength, allowing easier breaking of the sea

68 ice. The changing sea ice conditions influence the heat exchange between ocean  
69 and atmosphere, which is important for the whole Earth's Climate System and not  
70 only the Arctic regions (Serreze et al., 2009; Meredith et al., 2019). Leads and thin  
71 ice are much warmer than the surrounding sea ice and snow and thus heat loss is  
72 more than a magnitude larger in leads compared to the surrounding ice (Maykut,  
73 1982). Therefore, a better understanding of the interaction between ocean, sea ice,  
74 and atmosphere is essential. The high resolution lead data presented here have the  
75 potential for evaluation of the sub-footprint scale information of satellite remote  
76 sensing products.

77 Leads, with open water or thin ice cover, have high variability in time and  
78 space (Yu and Rothrock, 1996; Willmes and Heinemann, 2015). Therefore, it  
79 is important to monitor their conditions throughout the year. In Arctic pack ice  
80 during winter, a lead area fraction (open water and thin ice combined) of less  
81 than 10% can be expected (Yu and Rothrock, 1996), while in the Central Arctic  
82 lead area fractions are typically even lower (Wang et al., 2016). Willmes and  
83 Heinemann (2016) (2003–2015; satellite) and Wang et al. (2016) (1985–2014;  
84 model) could not find a trend in the lead area fraction, and Wang et al. (2016)  
85 found that the winds mainly determine the inter-annual variability in lead area  
86 fraction. However, a precise determination of the lead area fraction is crucial.  
87 Lüpkes et al. (2008) showed that a slight change in the high sea ice concentration  
88 (SIC) range, e.g., by the opening of leads, affects the near-surface air temperature.  
89 According to their study, a change of 1% in SIC could cause an air temperature  
90 change of up to 3.5 K. Small reductions in SIC, mostly induced by leads, have  
91 a (non-linear) more efficient effect on the heat exchange between the ocean and  
92 atmosphere than when a closed sea ice cover is present (Maykut, 1978). Therefore  
93 small changes in winter sea ice concentration, i.e., changes in lead area fraction  
94 are necessary to monitor. For example, if an increase of surface temperature on  
95 regional scale (e.g. satellite footprint) is (i) caused by many small leads the heat  
96 flux is stronger affected than if it is (ii) caused by a thinner but closed ice cover.

97 We first describe our helicopter measurement program and explain the  
98 principles of thermal sea ice observation. In the next part, we describe  
99 the spatio-temporal variability of our high resolution surface temperatures.  
100 Afterwards, we describe the lead classification based on the temperature  
101 difference. In the next section, we analyse the temporal variability of the surface  
102 temperature and lead area fraction on different spatial scales and present a case  
103 study of the November 2019 storm event. The last part focuses on the lead  
104 properties, i.e., width and orientation, derived after segmenting the classified  
105 leads.

## 106 2. Data

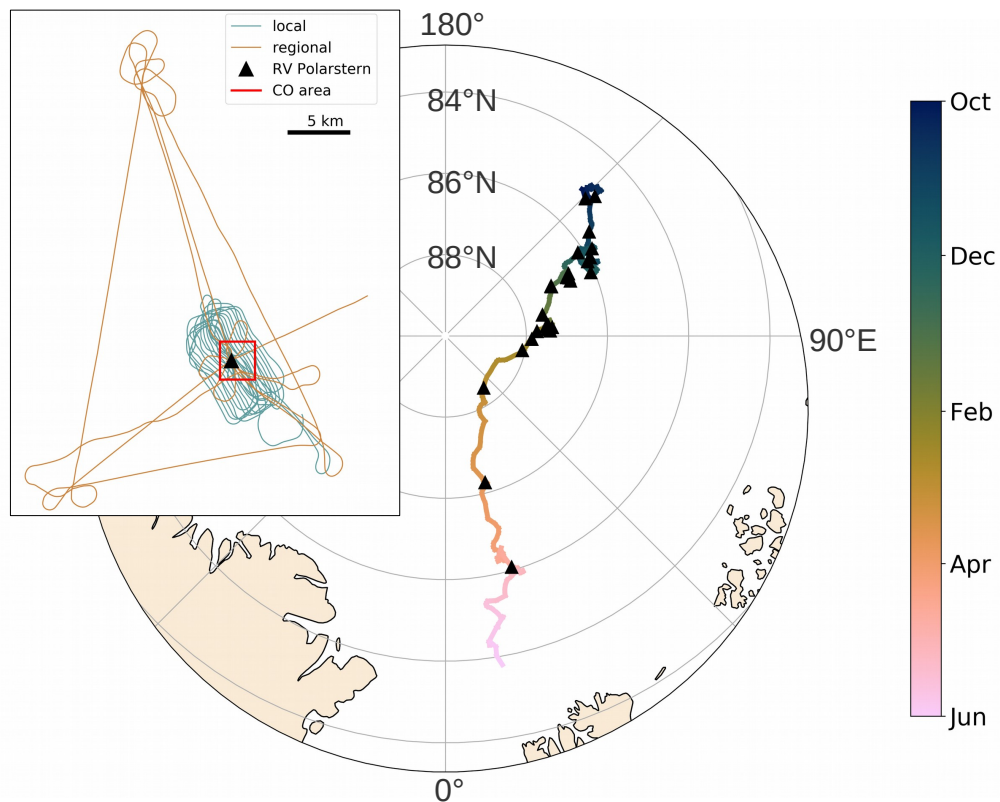
### 107 2.1. *Thermal sea ice observation*

108 With an infrared camera (InfraTec VarioCAM HD) installed, 35 helicopter flights  
109 were performed on a roughly weekly basis between 02 October 2019 and 23  
110 April 2020 from RV Polarstern (Alfred-Wegener-Institut Helmholtz-Zentrum für  
111 Polar- und Meeresforschung, 2017) (Figure 1). The set of flights consists of four  
112 main flight patterns: (i) Central Observatory (CO) (local), (ii) L-site triangles  
113 (regional), (iii) L-site grids (other), and (iv) event-related (other), like mapping  
114 particular leads. Detail about the surface temperature maps (Thielke et al., 2022)  
115 and pre-processing are presented in Thielke et al. (2022). To our knowledge  
116 regional scale sea ice infrared imaging has not yet been analysed and published  
117 before, such as done in the scope of this study.

118 Our measurements with helicopter-borne thermal infrared (TIR) imaging  
119 provide temperatures of the sea ice surface with a high spatial resolution of  
120 1 m which is substantially higher than TIR satellites, like MODIS, that have a  
121 resolution of about 1 km. Nevertheless, satellites are the primary tool for the  
122 Arctic sea ice state observations (Spren and Kern, 2017; Fox-Kemper et al.,  
123 2021). Compared to pan-Arctic coverage from satellites, we can provide with our  
124 helicopter data restricted area coverage from a local 5 km scale to a regional 40 km  
125 scale. Investigating the small-scale variability is important to better understand the  
126 representation of sea ice properties in models and satellite retrievals on a sub-grid  
127 scale (Vihma et al., 2014). Thus, this data is valuable for evaluating models and  
128 satellite retrievals (Ivanova et al., 2016).

129 The TIR temperature can distinguish open water and thin ice from thick ice,  
130 particularly for thin ice thickness of less than 1 m (Shokr and Sinha, 2015).  
131 Open water rarely exists in winter because the freezing starts directly after a lead  
132 opening. Therefore, we expect to capture mainly thin ice and only small open  
133 water areas with significantly warmer surface temperatures. Open water and thin  
134 ice areas influence the Arctic heat budget by allowing increased heat exchange  
135 between the ocean and the atmosphere. Above 1 m ice thickness, heat flux changes  
136 are minimal and have minor relevance for the Arctic heat budget (Maykut, 1982).  
137 Maykut (1982) found that in winter, the heat contribution from thin ice in leads is  
138 similar to the open water area and even larger than the dominating thick ice area.

139 The radiation in the TIR spectral region has a very small penetration depth  
140 on a sub-millimeters scale in snow, and ice (Shokr and Sinha, 2015, pages  
141 272,294). As a result, the TIR brightness temperature provides a measurement of  
142 the upper surface of snow or sea ice. Thus, the recorded temperature is expected



**Figure 1. Helicopter flight locations and flight patterns**

The colored track shows the drift of RV Polarstern from October 2019 until June 2020. The black triangles represent the location of the 35 helicopter flights. Additionally, as inlay on the left, we show a typical local (turquoise) and regional (orange) flight pattern with Polarstern as the center (black triangle). The red box marks the CO area (according to Figure S1, in Supplemental material).

143 to be influenced by atmospheric changes through the radiation balance at the  
144 snow/ice-air interface. Clouds strongly influence the surface temperature (Vihma  
145 and Pirazzini, 2005), i.e., they reduce the radiative cooling (Wang et al., 2001).  
146 Our flights were performed only during calm and clear weather conditions. Thus,  
147 we can neglect a dependence on changing cloud cover. However, the changing air  
148 temperature still plays a role, which needs to be taken into account (Thielke et al.,  
149 2022).

## 150 2.2. *Meteorological context*

151 How representative are our results from the MOSAiC winter in terms of surface  
152 temperature and lead area fraction in context with the meteorological condition?  
153 The meteorological conditions are discussed in Rinke et al. (2021) based on  
154 the ERA5 reanalysis data between 1979 and 2019. There were mostly typical  
155 meteorological conditions present during MOSAiC, although some unusual  
156 events happened during our observation period and before the expedition. Summer  
157 2019, before the MOSAiC expedition started, was very warm and had unusually  
158 long low sea ice extent as well as thinner ice (Rinke et al., 2021; Krumpfen et al.,  
159 2020). During the expedition, unusual conditions occurred during the following  
160 periods, all according to Rinke et al. (2021):

- 161 • Unusual cold at the beginning of November 2019 and March 2020
- 162 • Warming events in mid-November, beginning of December, mid-February,  
163 and mid-April
- 164 • Unusual positive Arctic Oscillation with associated fast sea ice drift in  
165 spring 2020 (Krumpfen et al., 2021; Dethloff et al., 2022)
- 166 • Anomalous low pressure January to April 2020 associated with more  
167 frequent storm events during winter and spring (relatively low cyclone  
168 counts for October 2019 – January 2020)

## 169 2.3. *Supporting data*

170 We use atmospheric data from the 12 m meteorological mast on the MOSAiC ice  
171 floe, i.e., 2 m air temperature, measured with a Vaisala HMT 330, and 10 m wind  
172 speed, measured with a Metek uSonic-3 cage Cox et al. (2021). These supporting  
173 measurements were measured at the location of Met City in the CO (Details see  
174 Shupe et al. (2022)).

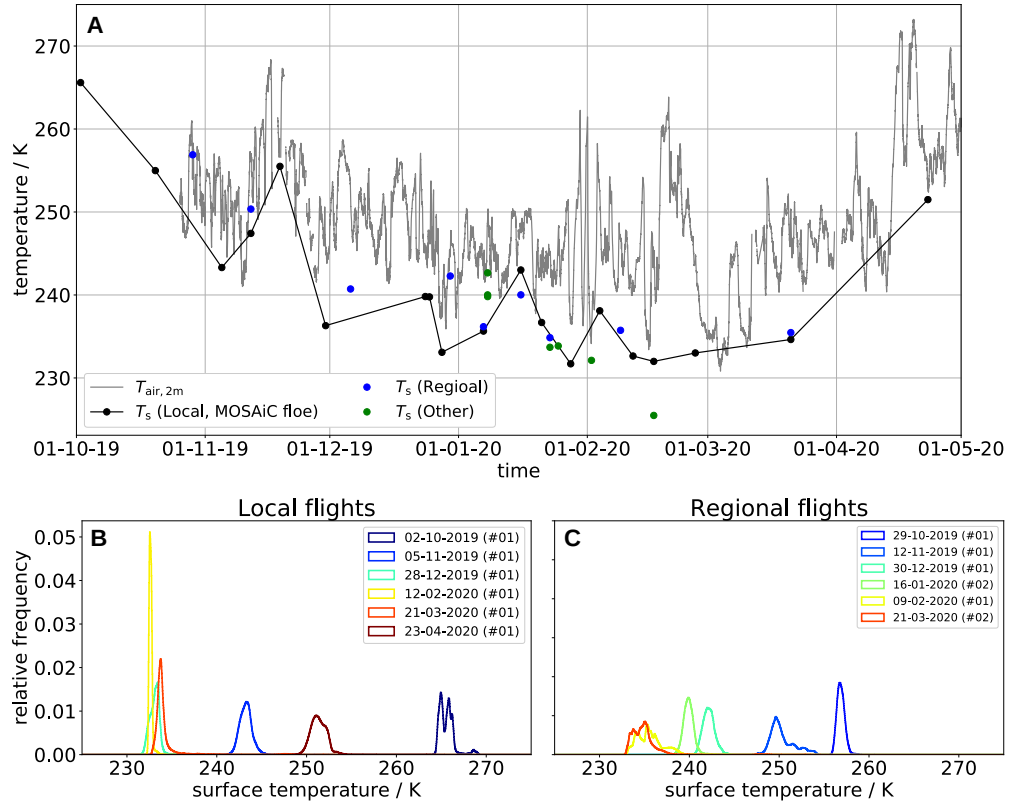
### 175 **3. Surface temperature variability**

176 This study focuses on the gridded time-fixed helicopter surface temperature  
177 maps (Thielke et al., 2022), which will be referred to as surface temperature  
178 for simplicity. Based on the series of 35 helicopter flights with the TIR camera  
179 during the MOSAiC winter, we present and discuss the spatio-temporal surface  
180 temperature variability. We present and discuss the complete seasonal cycle of the  
181 MOSAiC ice floe and its surrounding for the whole winter 2019/2020. There are  
182 unprocessed data from nine more flights available (for detailed explanation see  
183 Thielke et al. (2022)).

#### 184 *3.1. Temporal variability*

185 Ice surface temperature varies even on short timescales, i.e., within the flight  
186 duration of 90 minutes. This effect, however, is largely corrected in our data  
187 set. See Thielke et al. (2022) for how the corrected and time-fixed surface  
188 temperature maps are calculated. Here, in the following, we discuss the surface  
189 temperature seasonal variability. The average surface temperature decreased from  
190 02 October 2019 at 265.6 K until it reaches its minimum with 231.8 K on  
191 28 January 2020 (Figure 2 A). Later in the winter season, the average surface  
192 temperature increased to 251.4 K until the latest flight on 23 April 2020, while  
193 at that time, the 2 m air temperature was already about 20 K higher and close to  
194 the freezing point. However, the temporal evolution of the surface temperature  
195 is comparable to that of the 2 m air temperature. This is consistent with our  
196 expectation that, due to the shallow penetration depth (micrometer range) of  
197 electromagnetic waves in the thermal infrared region, air temperature will have  
198 a substantial influence on our surface temperature observations. Also, Vihma  
199 and Pirazzini (2005) highlight the importance of the surface temperature and  
200 coupling to the atmosphere. At the same time, the heterogeneity of the surface  
201 temperature in ice-covered regions can also influence the atmosphere. But as long  
202 as the surface is frozen, the surface temperatures stay well below the freezing  
203 point. The surface temperature can be cooler during clear sky conditions due  
204 to radiative cooling. A prominent interruption in the cooling happened at the  
205 beginning of the winter in mid-November due to a substantial increase in the  
206 surface and air temperature caused by a storm event (Rinke et al., 2021). More  
207 warming events (subsection 2.2) are reflected in the surface temperature record.  
208 From mid-February onward, the frequency of flights was reduced, so we cannot  
209 reflect all single atmospheric events. However, we can show the warming of the  
210 surface temperature towards spring.





**Figure 2. Evolution of MOSAiC surface temperatures from 35 helicopter flights.**

(A) Temporal evolution of the average surface temperature throughout winter 2019/2020 from 02 October 2019 to 23 April 2020. Black indicates the local flights covering the Central Observatory (CO). They are connected to show the temporal evolution of the primary MOSAiC observation area. The regional flights, repeatedly visiting the L-Sites in the MOSAiC distributed network, are displayed in blue, whereas green shows additional flights not falling in one of these two categories. The grey line represents the 2 m air temperature measured at the floe in Met City. In the lower panel, a selection of surface temperature distributions is shown for different dates in the winter for (B) the local and (C) the regional flights. The colors continue from blue (begin of the winter) to red (end of the winter).

### 211 3.2. *Spatial variability*

212 In January 2020, there is a high density of flights, which allows us to illustrate the  
213 variability between different scales nicely or even on a short timescale for similar  
214 spatial surveys (Figure 2 A). On 07 and 16 January 2020, a local (black) and a  
215 regional (blue) flight were conducted on the same day (the regional flight is in  
216 both cases about 3 hours later). On 07 January, the average surface temperature of  
217 the different flights is similar, with an increase of only 0.6 K for the regional  
218 flight, which corresponds to an increase of 0.8 K in the 2 m air temperature.  
219 On 16 January 2020, the difference is larger with a decrease of 2.7 K while  
220 the 2 m air temperature only decreased by about 0.9 K. Thus, changes in the  
221 spatial surface temperature variability, either on the local or regional scale, has to  
222 explain the increasing temperature difference within nine days between the local  
223 and regional scale (flights had similar flight patterns). A likely candidate would  
224 have been changing lead fraction but actually a change of number of leads does  
225 not seem to be the reason for the higher temperature difference on 16 January: On  
226 07 January there is actually a higher lead area fraction for the colder local flight  
227 (2.02% vs. 0.23% for the regional flight, see Table 1). While on 16 January 2020  
228 the colder regional flight had a higher lead area fraction (1.37% vs. 0% on local  
229 scale; see Table 1). Thus, likely changes in other ice types with different thermal  
230 properties, snowfall, or snow redistribution should have caused these changes in  
231 spatial surface temperature variability.

232 The spread of the surface temperature varies from flight to flight. It is  
233 illustrated with the exemplary selection of six surface temperature distributions  
234 for local and regional scales throughout the winter season (Figure 2 B, C). Mostly,  
235 the distributions are wider for regional flights (right) than for local flights (left)  
236 because they include a larger variety of surface types due to the wider spatial  
237 extent. The major peak represents the predominant surface type, snow-covered  
238 thick ice, in all cases. The surface temperatures of this thick ice are more  
239 similar to the 2 m air temperature because of the reduced heat flux from the  
240 ocean through the thick ice and snow (Shokr and Sinha, 2015). The warm tail  
241 shows the presence of leads, but its peak is often too small (only visible in the  
242 log-scale), and the different thin ice thicknesses in leads of different ages widen  
243 the lead temperature distribution. In the local flights, we can find a transition from  
244 a wider distribution at the beginning of the winter season (blue) to a narrow  
245 distribution in mid-winter (yellow), and back to a wider distribution towards  
246 the end of the winter season (red). The more narrow distributions indicate the  
247 prevalence of the thick, consolidated ice due to cold and constant conditions. A

248 wider distribution indicates the presence of several ice classes and spread towards  
249 warmer temperatures.

250 In the time series, we could already show the close connection to the  
251 atmospheric state, represented by the 2 m air temperature. Additionally, we look  
252 at the dependence of the surface temperature standard deviation, as a measure  
253 of spatial variability, (leads were excluded to have a comparable basis of the  
254 thick ice) on the 10 m wind speed (Figure S2, in Supplemental material) . We  
255 expect a lower surface temperature standard deviation for higher wind speeds  
256 caused by an increased exchange between the surface and atmosphere. We find  
257 a correlation of  $-0.38$  between the surface temperature standard deviation and  
258 10 m wind speed around the target time of the flight with the significance of a  
259 p-value of 0.04. The relationship with wind speed supports our assumptions that  
260 increasing wind speeds reduce surface temperature variability and can explain the  
261 greater sensible heat exchange due to faster air mass exchange. Because leads are  
262 not taken into account here, a potentially more dynamic ice pack that would result  
263 in more leads (and thus greater temperatures) can be ruled out. There is no relation  
264 between the average surface temperatures and the standard deviation. Although a  
265 low-temperature regime can create more compact and consolidated ice (uniform  
266 temperature), even under cold conditions, deformation can cause variability in ice  
267 classes with warmer surface temperatures, which increases the variation.

#### 268 **4. Lead classification**

269 The leads are classified based on a one-dimensional, temperature-only approach,  
270 i.e., leads are characterized by a specific surface temperature range, defined  
271 by the temperature distribution of the corresponding flight. We apply a binary  
272 classification, discriminating between sea ice (snow-covered thicker ice) and  
273 leads, which are mostly covered with thin ice due to the fast freezing of the  
274 ocean surface under cold winter conditions. The classification is applied to the  
275 surface temperature grids. The pre-processing described in [Thielke et al. \(2022\)](#)  
276 avoids the influence of changes in surface temperature during one flight. The  
277 thermal distinction for surface types with larger ice thickness is getting weaker  
278 due to low heat transfer through the ice ([Maykut, 1978](#)). Thus, the discrimination  
279 of leads from thick ice is easier due to large temperature differences, and we  
280 do not aim to discriminate the thick ice classes further (e.g., in first-year and  
281 second-year ice). We apply dynamic thresholds for different flights (Table S1,  
282 in Supplemental material and [subsection 4.1](#)) in order to establish the same  
283 classes defined with different surface temperature distributions in different flights.

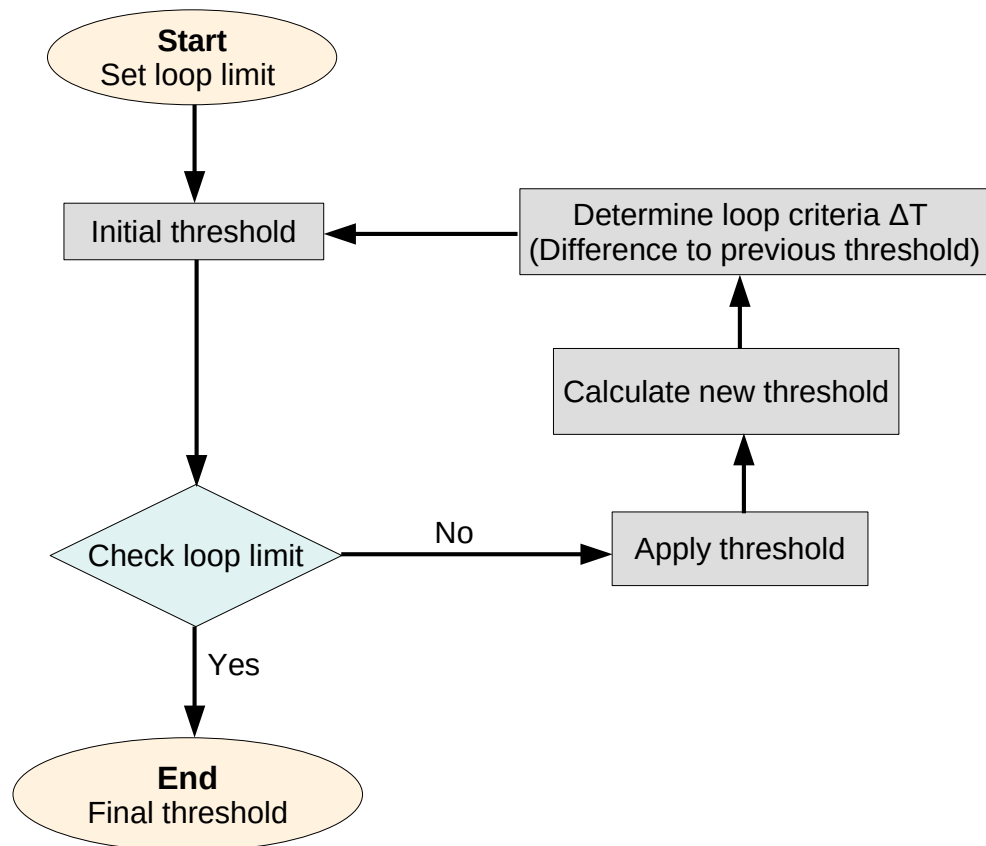
284 The dynamic threshold is required because the surface temperature is strongly  
285 connected to the ambient air temperature, which is strongly variable with time  
286 (see [subsection 3.1](#)).

#### 287 *4.1. Iterative threshold selection*

288 We apply the iterative threshold selection from [Ridler et al. \(1978\)](#) to the  
289 two-dimensional temperature arrays. The temperature distribution has its major  
290 mode towards the colder part of the distribution and a smaller secondary mode  
291 along the tail towards the warmest temperatures (caused, e.g., by leads). The  
292 initial threshold is the middle range of the temperature distribution (average of  
293 the minimum and maximum value). Due to the long tail towards the warmer  
294 temperatures, it is ensured that the initial threshold is on the warmer side of the  
295 major mode so that the iteration can converge towards a minimum between the  
296 two modes. Starting from the initial threshold, the threshold is adjusted iteratively  
297 based on the new "lead" mask (defined by the threshold of each iteration) for the  
298 temperature array until it reaches the final temperature threshold. The updated  
299 threshold is calculated from the mean of the "lead" (all values larger than the  
300 threshold) and "sea ice" (all values smaller than the threshold) temperatures  
301 based on the current threshold. The iteration stop criterion is achieved when  
302 the temperature threshold change between one iteration to the next is within  
303 the tolerance of 0.02 K, which corresponds to the precision of the IR camera.  
304 The main steps are shown in the flowchart in [Figure 3](#), and the Python3 code  
305 is shared in part A of the Supplemental material. For three flights (20191224\_01,  
306 20191225\_01, 20200108\_01), the threshold did not converge to a reasonable value;  
307 therefore, the tolerance had to be increased to 0.8 K. For the cases with this  
308 larger allowed tolerance in the difference, the obtained threshold still results in a  
309 reasonable lead classification, i.e., can be confirmed by the manual classification  
310 (see below in [subsection 4.2](#)).

##### 311 **4.1.1. Classification example 20 October 2019**

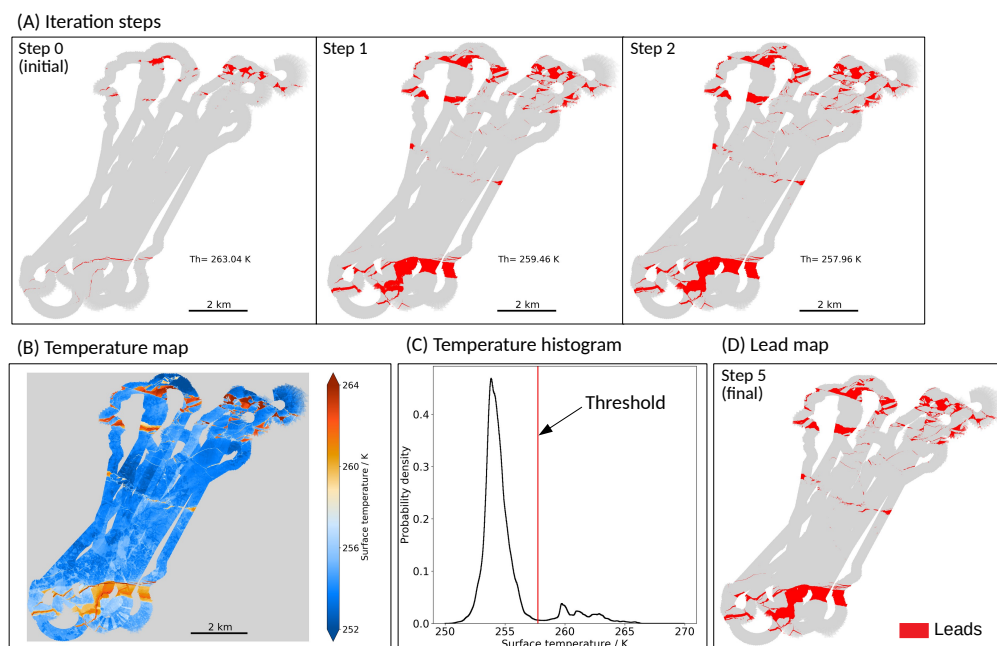
312 The threshold-based lead classification is shown for the flight on 20 October  
313 2019 at the beginning of the winter season in [Figure 4](#). We show steps 0 (initial  
314 threshold) to 2 for the temperature threshold iteration (A). With step 2, the result  
315 is already close to the final result (step 5) shown in (D). The surface temperature  
316 (B) is dominated by low temperatures (blue), associated with snow-covered thick  
317 ice. The warm surface temperatures (red domain) are referred to as leads. The  
318 binary classification map (D) resulting in "sea ice" (gray) or "leads" (red) is based



**Figure 3. Flowchart for the processes of the iterative threshold selection.**

Main processing steps for the iterative threshold selection to determine a temperature threshold for lead classification.

319 on the iterative temperature threshold applied to the temperature distribution (C),  
 320 showing the two classes as two main temperature regimes. The lead area fraction  
 321 for this case on 20 October 2019 is the highest in our time series, with close to  
 322 10% for the full area covered by the helicopter flight.



**Figure 4. Lead classification example for the flight on 20 October 2019.**

(A) Results from steps 0, 1, and 2 for the iterative threshold selection. (D) Final (step 5) binary lead classification based on (B) the gridded surface temperature maps. (C) The temperature distribution of (B). The red vertical line represents the found iterative temperature threshold to discriminate between "lead" and "no lead" surfaces.

#### 323 4.2. Evaluation using manual thresholds

324 We use manual thresholds as a comparison for the reliability of the iterative  
 325 method. The authors performed the manual threshold selection based on the  
 326 minima in the distribution and visual approval of the classified map compared  
 327 to the surface temperature map. The manual selection is a rather arbitrary and  
 328 conservative choice but it can be used for the evaluation of the automatic, iterative  
 329 classification method. The thresholds are determined for each flight individually

330 because the surface temperature values and their distributions change from flight  
331 to flight. The manual selection was, in some cases, more conservative, i.e., it  
332 has warmer thresholds but does not significantly influence the resulting lead area  
333 fractions (Table S1, in Supplementary material). Although the manual temperature  
334 threshold is 1.97 K higher than the one from the automatic and reproducible  
335 method (subsection 4.1), its derived lead area fraction is only 0.1% lower than for  
336 the iterative method. Thus, the threshold difference does not significantly affect  
337 the lead area fraction because it is in the minimum of the surface temperature  
338 distribution. The small difference demonstrates that the automatic method aligns  
339 well with the manually defined thresholds and the resulting lead classification. We  
340 chose results from the automatic method for the following discussion because it  
341 is reproducible and can also be applied to further flights.

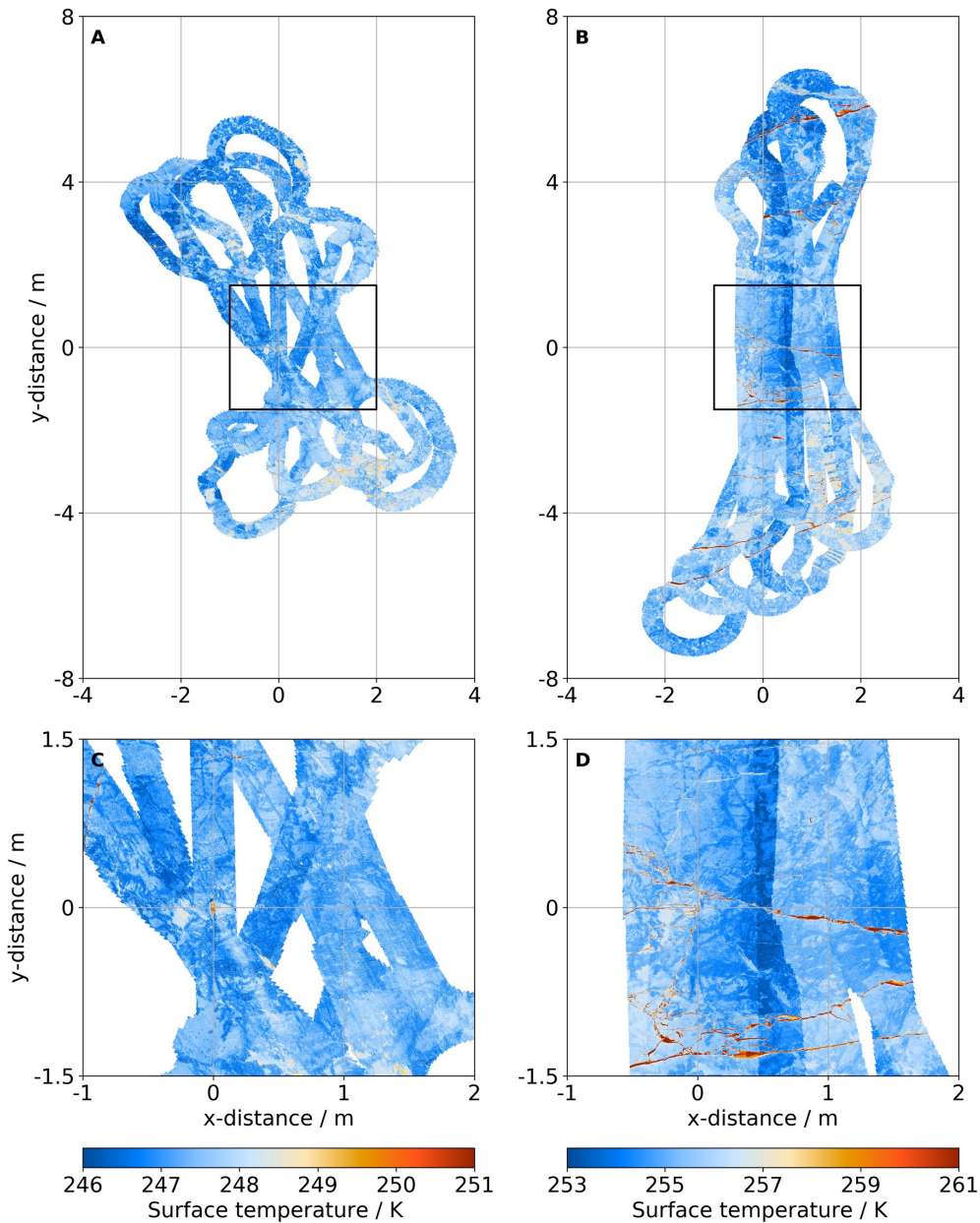
## 342 5. Winter lead area fraction

### 343 5.1. Lead formation during the November storm event

344 Here, we discuss the storm event, which happened from 16 to 20 November 2019  
345 (Rinke et al., 2021). This event had a major influence on the MOSAiC Central  
346 Observatory (CO) due to several leads appearing across the measurement sites.  
347 It had a significant relevance for several measurements (Nicolaus et al., 2022;  
348 Shupe et al., 2022) and was, e.g., influencing the snow transport as discussed in  
349 Nandan et al. (2022). We conducted one flight before (12 November 2019) and  
350 one flight after the storm (19 November 2019) and compare both flights directly  
351 in Figure 5. This storm event with high wind speeds was associated with warm  
352 air advection. It resulted in increased ice dynamics, which caused the break-up of  
353 the sea ice along various fracture lines, which are visible in Figure 5 B and D.  
354 The surface temperatures on 19 November 2019 after the storm are overall higher  
355 than before the storm event (mind the different temperature scales for (A)/(C) and  
356 (B)/(D)). Before the storm, there are a few narrow cracks in the outer areas of  
357 the flight pattern (Figure 5 (A)), but no prominent cracks in the vicinity of RV  
358 Polarstern (Figure 5 (C)). The surface temperature map of 19 November 2019  
359 includes warm linear structures throughout the CO area and beyond (Figure 5 (B,  
360 D)), which causes an increase of lead area fraction from 0.07% (Figure 5 (C)) to  
361 1.73% (Figure 5 (D)) and therewith to a higher surface temperature variability.

### 362 5.2. Spatio-temporal variability

363 We discuss the spatio-temporal variability of the lead area fraction, which was  
364 calculated based on our lead classification (section 4). The local lead area fraction



**Figure 5. Surface temperature maps before and after the November storm event.**

Comparison of two time-fixed surface temperature maps displayed in relative coordinates before the storm event (A, C) on 12 November 2019, and after the storm event (B, D) on 19 November 2019 with their respective CO area around RV Polarstern (0,0). Please note the different temperature ranges, adjusted to allow clearer visibility of the spatial variability of the temperature. For the flight on 19 November 2019, an increased area of warmer temperatures (reddish) is prominent.



365 (orange line in [Figure 6](#)) is constrained to the CO area of  $3 \times 3$  km. In the CO area  
366 (shown in [Figure S1](#) in Supplemental material), always the same area around RV  
367 Polarstern is covered and makes the lead area fraction better comparable during  
368 the winter season as for the entire local flights. In most cases, the CO area lead area  
369 fraction is close to the one for the complete local flights (not shown); however,  
370 there is a large difference in lead area fraction on 20 October 2019 (1.3% for  
371 the CO area; 10.2% for the full local flight) because the majority of the detected  
372 leads are outside of the CO area. The data coverage of the CO area is sufficient  
373 for all flights with more than 50% and, except for the flights in November, even  
374 with more than 75% ([Figure S3](#), in Supplemental material). The lead area fraction  
375 within the CO area shows high variability between 0% and 4%, but no trend can  
376 be seen in the temporal evolution ([Figure 6](#), orange line). For the November storm  
377 event ([subsection 5.1](#)), there is an increase in lead area fraction for the CO area  
378 from close to 0% on 12 November 2019 to 1.7% on 19 November 2019. We find  
379 for the regional scale, there is a steady low lead area fraction between 0% and  
380 1% until mid-January. Towards the end of the winter season (21 March 2020), the  
381 lead area fraction increases to up to 4%. The increase in lead area fraction might  
382 be related to the increased number of storm events between January and April,  
383 compared to previous months ([subsection 2.2](#)).

384 The higher lead area fraction observed in March in our data aligns with  
385 the temporal lead area evolution discussed for the regional scale in [Krumpen](#)  
386 [et al. \(2021\)](#). However, they observe a distinct higher lead area fraction of up  
387 to 20%, derived from MODIS TIR satellite data. The regional lead area fraction  
388 evolution also agrees well with the regional Satellite Synthetic Aperture Radar  
389 (SAR) derived time series from [Guo et al. \(2022\)](#), which is in the similar range of  
390 0-4% as ours and starts to increase only in March. [Kortum et al. \(2022\)](#) performed  
391 an ice classification based on SAR satellite data during winter on the same scale  
392 as our CO area. Our leads should be represented by the sum of open water and  
393 young ice classes of this study, where daily data are available (higher temporal  
394 resolution than our data). Their daily data generally show a comparable to our  
395 lead area fraction below 5% in mid-winter but exceed this value on a few days  
396 to up to 15% around 23 November 2019. Additionally, during March and April,  
397 their lead area fraction is for a longer time on a higher level of up to 10%. The  
398 high values at the beginning of the winter in [Kortum et al. \(2022\)](#) might be caused  
399 by the characteristics of the method based on SAR data. The same method from  
400 [Kortum et al. \(2022\)](#) was applied on the regional scale. The regional lead area  
401 fraction has a peak in mid-November of 6%, stays below 4% and even lower  
402 during mid-winter, and increases in the second half of March to 7% (Karl Kortum,

403 personal communication on 08 December 2022). While we do not capture their  
404 peak in November and the absolute values differ with a few percentage points,  
405 also the regional evolution aligns with our time series. Thus our results align  
406 reasonably well (considering different temporal and spatial sampling) with the  
407 two SAR-based studies, while for the MODIS TIR based study only the temporal  
408 evolution agrees but the absolute values are different (much higher in the MODIS  
409 lead time series).

410 Within our data, we see scale-dependent differences in the lead area fraction  
411 with less variability on the regional scale than on the local CO scale but no trend  
412 in the local scale, while the regional lead area fraction increases throughout the  
413 winter (Figure 6). Nevertheless, the overall magnitude is similar. Thus the CO area  
414 is representative of the measurement sites in the CO of the MOSAiC expedition  
415 but the temporal development does not necessarily represent the lead area fraction  
416 on a larger scale. Nonetheless, the local data are helpful for a better understanding  
417 of the condition at and around the MOSAiC floe, particularly in connection to  
418 other in-situ measurements. Our lead area fractions (0-4%) are comparable to  
419 other previous studies of Marcq and Weiss (2012) with 1–2%, and Lindsay and  
420 Rothrock (1995) with 2–3%; both these winter lead area fractions were also  
421 derived in the central Arctic. Generally, the lead area fraction for the MOSAiC  
422 winter seems to align with the climatological mean and might be influenced by  
423 the changing location due to the MOSAiC drift (Krumpen et al., 2021). Yet,  
424 comparing different lead area fraction retrievals remains challenging because of  
425 different definitions of leads (e.g. open leads vs. leads covered by thin ice or even  
426 frost flowers) with other methods used on different scales (von Albedyll et al.,  
427 2022).

### 428 5.2.1. Relation to wind speed

429 We are interested in the connection between lead area fraction and wind speed  
430 because wind events can cause increased ice dynamics and, therefore, possibly  
431 more leads. Thus, we compare our lead area fraction to the 10 m wind speed.  
432 We use the 7-day running mean for the wind speed to find prominent high wind  
433 regimes rather than short-term fluctuations because we can not represent these  
434 fluctuations with the limited temporal frequency of the helicopter flights.

435 We do not find significant correlation between wind speed and lead area  
436 fraction on the local and regional scale. However, we see sometimes a relation,  
437 such as for the increase of the lead area fraction (Figure 6) during the November  
438 storm event (subsection 5.1). We have to note that our flights are only snapshots  
439 of a specific time with a weekly to biweekly frequency, while leads can open and

440 close within hours. However, in most cases, they prevail for several days (if not  
441 closed by another ice dynamic event) until, eventually, the ice thickness and snow  
442 accumulation within them gets too thick to be discernible from the surrounding  
443 ice in TIR imagery.

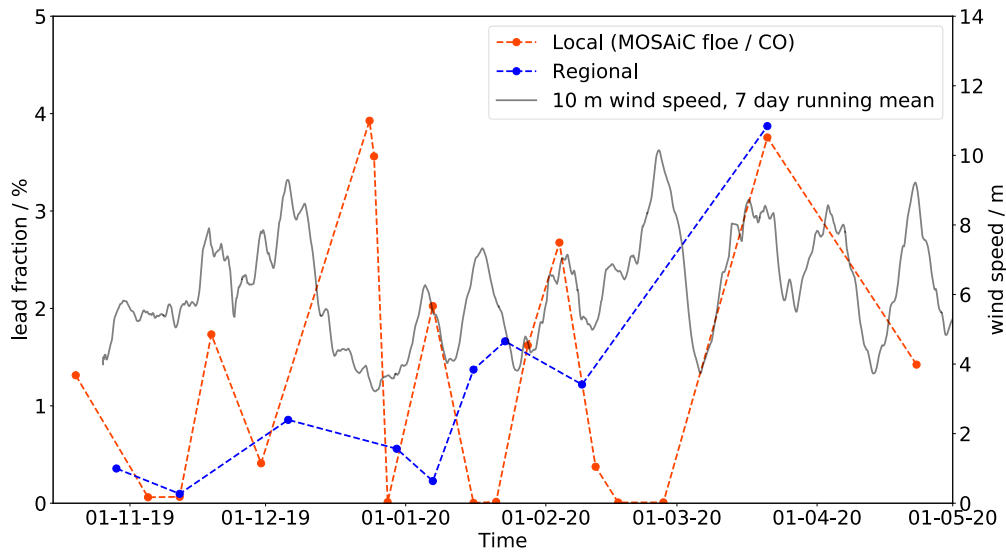
444 We highlight two cases in our time series: 1) the highest local lead area fraction  
445 within the CO area of 4% end of December occurs after a high wind regime that  
446 lasted several days, and 2) basically no presence of leads in mid-January during  
447 an increased wind speed regime while the regional fraction increases. Especially  
448 the high lead area fraction variability for the local CO area illustrates that local  
449 changes are rather random, not always representing large scale changes (see  
450 different temporal development between local and regional lead area fraction).  
451 Nevertheless, the local lead area fraction is valuable in combination with other  
452 interdisciplinary measurements obtained during MOSAiC and valuable for getting  
453 a better process understanding.

## 454 **6. Lead properties**

### 455 *6.1. Lead segmentation*

456 We apply a segmentation algorithm to the lead map to define the properties of  
457 single leads, i.e, width and orientation. The segmentation is performed according  
458 to the watershed segmentation (Najman and Schmitt, 1994). Next, a set of object  
459 lead properties (width and orientation from enclosing ellipse, its area, orientation,  
460 and major axis) is derived based on (Burger and Burge, 2009) with 'scikit-image'  
461 library for Python.

462 In Figure 7, we illustrate the object properties for two example lead segments.  
463 The warmer temperatures on the left (yellow) are classified as lead, consistent  
464 with the red areas on the right that indicate the lead areas. We retrieve the lead  
465 properties width and orientation (calculated from ellipse parameters), assuming  
466 the lead properties are representative, even if the ellipse does not cover the full  
467 lead due to the limited spatial coverage of our data or if the lead is interrupted.  
468 We therefore can only determine width but not length of the leads. The key  
469 lead parameters are the classified area (red) in the enclosing rectangle (dashed  
470 line in Figure 7), minor and major axis, as well as the orientation of the ellipse  
471 (pointing in the direction of the major axis). The zero line for the orientation  
472 is the north-south axis (all our surface temperature maps are oriented North  
473 along the y-axis). The ellipse defined for the lead segment is not representing  
474 the real length, but can be seen as a stable approximation for an object of arbitrary  
475 shape. Generally, the ellipses of close-by leads can overlap, which is required



**Figure 6. Evolution of the lead area fraction on different scales.**

Temporal evolution of the lead area fraction throughout winter 2019/2020 from 02 October 2019 to 23 April 2020. The orange points show the lead area fraction for the CO area. The blue points illustrate the lead area fraction for the regional flights, visiting the L-Sites. The grey line shows the 10 m wind speed averaged to a 7-day running mean. Please note that there might be a minor influence by in and out coming support vessels, which could slightly increase the lead area fraction by breaking the ice.

**Table 1. Lead area fraction values from [Figure 6](#).**

Lead area fraction values for the local and regional scale in %. The same as the displayed values of the data shown in [Figure 6](#).

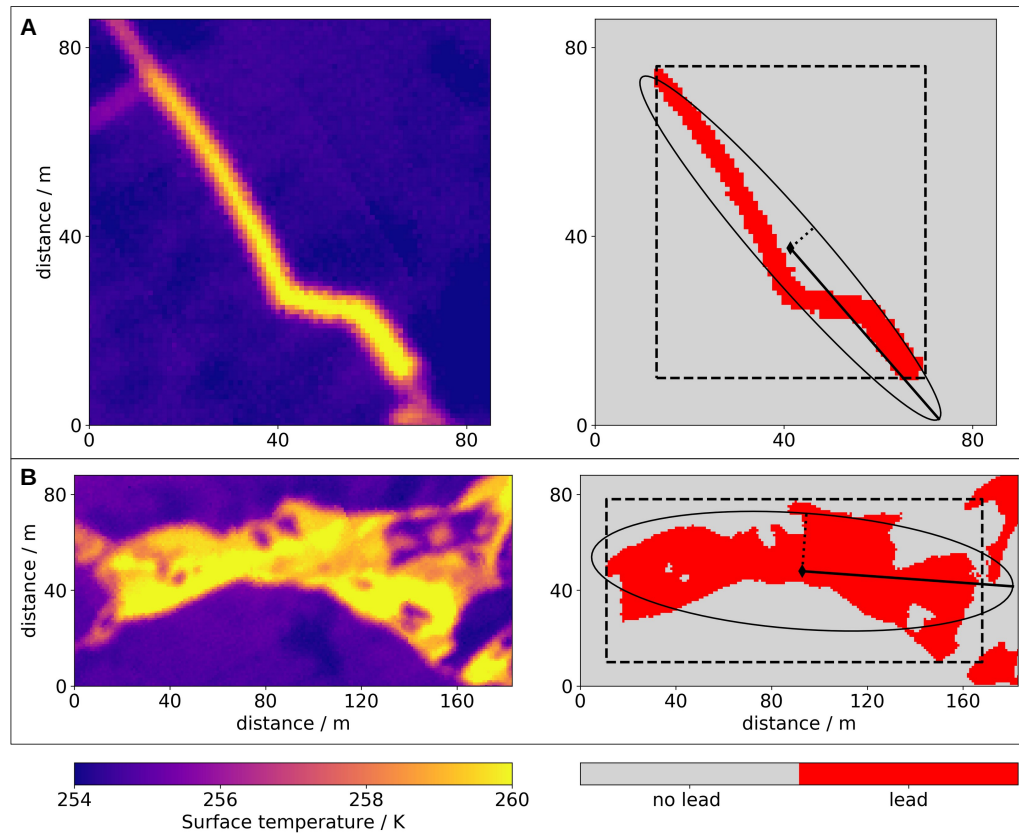
Date	Local fraction / %	Regional fraction / %
2019-10-20	1.31	-
2019-10-29	-	0.36
2019-11-05	0.06	-
2019-11-12	0.07	0.10
2019-11-19	1.73	-
2019-11-30	0.41	-
2019-12-06	-	0.86
2019-12-24	3.93	-
2019-12-25	3.56	-
2019-12-28	0.01	-
2019-12-30	-	0.56
2020-01-07	2.02	0.23
2020-01-16	0.00	1.37
2020-01-21	0.01	-
2020-01-23	-	1.66
2020-01-28	1.62	-
2020-02-04	2.68	-
2020-02-09	-	1.22
2020-02-12	0.37	-
2020-02-17	0.01	-
2020-02-27	0.01	-
2020-03-21	3.76	3.87
2020-04-23	1.42	-

476 to calculate the lead properties individually, even though the two lead segments  
477 are not overlapping (Figure S4, in Supplemental material). We must deal with  
478 some artificial effects, such as the map's edge or shifts inside the map caused by  
479 small offsets in the geolocation of different helicopter overflights. Shifts or gaps  
480 could cause an artificial break of a lead into more segments, whereas it would  
481 have been only a single lead. Also, due to ice drift direction changes (i.e., shear),  
482 which can cause real breaks and gaps in the classified leads, the segments can  
483 represent a subset of a lead. The segments of the subsets of leads will result in an  
484 overestimation of the total number of leads. However, it is not expected to impact  
485 our results for lead width and orientation (we do not analyse the number of leads).  
486 Therefore, we assume that the segmentation is representative of our purpose of  
487 an overall statistical analysis of lead width and orientation. Width and orientation  
488 may also be critical parameters for evaluating the turbulent heat flux from leads  
489 (Tschudi et al., 2002). On the one hand, the efficiency of the heat transfer is  
490 dependent on the orientation relative to the wind direction (e.g., Tetzlaff et al.,  
491 2015). On the other hand, the heat transfer is more efficient for narrow leads,  
492 which makes the transfer dependent on the lead width distribution (Marcq and  
493 Weiss, 2012).

## 494 6.2. *Lead orientations*

495 A good understanding of lead orientation is crucial because they represent the ice  
496 dynamics of the sea ice (Lindsay and Rothrock, 1995). Ringeisen et al. (2019)  
497 emphasize the lack of knowledge of lead orientation at the floe scale because  
498 of missing high resolution observations. Here, the MOSAiC observations like  
499 ours can contribute new data. Better knowledge of small scale leads is also  
500 crucial for a good representation of ice rheology in sea ice models (Hutter et al.,  
501 2018; Ringeisen et al., 2021). The orientation of leads shows the effect of ice  
502 dynamics in sea ice, influences it, and is connected to prevailing regional wind  
503 and ocean current (Lindsay and Rothrock, 1995). In the long term, leads have a  
504 non-random orientation during the Arctic winter, mainly influenced by coastlines  
505 and atmospheric and oceanic currents (e.g., Bröhan and Kaleschke, 2014). In  
506 general, lead features, including width and orientation, are similar across a large  
507 range of scales, including the smallest scales (Schulson, 2004).

508 We here look at the lead orientations of nine local flights (full coverage, i.e. not  
509 restricted to the CO area), which have in the CO area a lead area fraction of  $\geq 1\%$ .  
510 This ensures a sufficient presence of leads to perform a statistical analysis of lead  
511 orientations. We decided not to connect the single lead segments which might be



**Figure 7. Lead segmentation to derive lead width and orientation properties.**

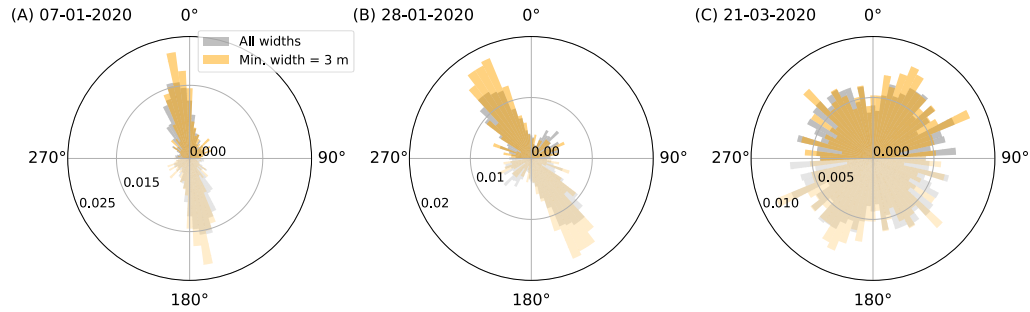
Two lead segments from the lead classification result of the flight on 20 October 2019 with the temperature map on the left and the lead classification including the ellipse geometry on the right. The ellipse and their major axis (solid) and minor axis (dotted) are shown. The dashed rectangle marks the area from which the classified area in red is determined. (A) Shows a narrow lead with a mean lead width of 3 m and an orientation (of the major axis) of  $-41^\circ$ . (B) Shows a wider and slightly scattered lead. It has a 26 m mean width and  $-86^\circ$  as orientation.

512 split within one lead (subsection 6.1) because we look at statistical distributions of  
513 lead properties and do not distinguish single leads. Furthermore, we bin our data  
514 in  $5^\circ$  steps.

515 Comparing the nine flights between 20 October 2019 and 23 April 2020, we  
516 see a high temporal variability in the lead orientation distribution (see examples  
517 in Figure 8), also shown for passive microwave based analysis in Bröhan and  
518 Kaleschke (2014). We find prevailing orientations of  $-80^\circ$ ,  $-10^\circ$ ,  $30^\circ$ , or  $60^\circ$  (Table  
519 S2, in Supplemental material). In the following part, we focus on three examples  
520 from 07 January 2020, 28 January 2020, and 21 March 2020 (Figure 8). We  
521 identify modes of the lead orientation probability distributions of the orientation  
522 from  $-90$  to  $90^\circ$ , binned in  $5^\circ$  steps (Figure S6, in Supplemental material). We  
523 constrain our data to elongated ellipse shapes with an axis ratio (major/minor)  
524 of at least two. Additionally, we compare the leads of all widths with leads of a  
525 width of more than 3 m which is consistent with the valid range of the power law.  
526 With the constraint of the axis ratio, the data are reduced to 89% of the full data  
527 set. With the minimum width of 3 m the data amount is reduced to 21% of the  
528 complete data set. Starting with the case in March (Figure 8 C), we cannot find a  
529 major peak in the distribution of all lead width. With only wider leads ( $\geq 3$  m),  
530 the distribution of orientation angles is modified to a preferred direction at  $35^\circ$ ,  
531 but still most orientations are present and not a clear prevailing orientation can  
532 be identified. Going backward in time to the end of January (Figure 8 B), we  
533 have one prominent orientation at  $-35^\circ$  (all leads) which is even more emphasised  
534 for leads with the minimum width of 3 m (slightly shifted to  $-30^\circ$ ). There is a  
535 second minor peak at  $40^\circ$ , but this is very small and does not allow us to infer  
536 any intersections between two main orientations. For the case on 07 January 2020  
537 (Figure 8 A) we identify one clear main direction of  $-10^\circ$ . We see a variation in  
538 the primary lead orientation throughout the winter but no prevailing orientations  
539 on longer time scales. For none of the nine investigated flight we can infer two  
540 main directions (bimodal distribution) from which we could infer an average  
541 intersection angle. Usual lead intersection angles from different studies, including  
542 satellite and laboratory measurements, would be  $30$ – $50^\circ$  (Hutter et al., 2022),  
543 also shown for a SAR data set from MOSAiC Ringeisen et al. (2022). There is a  
544 difference between using all data and the width restricted subset, but overall both  
545 show the same picture (Table S2, in Supplemental material). The variability might  
546 depend on the regional wind patterns that create direction-related fracture patterns.  
547 The investigation of the reasons for the variability in lead orientation is beyond  
548 the scope of this study, but is encouraged for future research. Different to many  
549 previous studies is that we are (a) far from land (which can introduce prevailing



550 lead orientation by restricting ice drift in one direction) and (b) following the  
 551 Lagrangian approach of the MOSAiC drift, which results in different locations of  
 552 the Arctic Ocean to be monitored.



**Figure 8. Orientation angles of leads for three example cases.**

Probability density distribution for the orientation angles of the flight from (A) 07 January 2020, (B) 28 January 2020, and (C) 21 March 2020, as polar histogram. The radius indicated the probability density, which is different for all three cases. Only lead segments with an axis ratio (major/minor)  $\geq 2$  are included. We discriminate between two cases: leads of all widths included (gray) and only leads with a minimum width of 3 m included (orange). The lead orientation have only a range of  $180^\circ$  but are valid in both directions, they are mirrored to the opposite direction (slightly transparent). The total number of lead segments used for the histograms ( $270^\circ$  to  $90^\circ$  only) are (all;  $\geq 3$  m): A=(1736; 500), B=(1326; 303); C=(1378; 464).

### 553 6.3. Lead width distribution

We discuss here the power law scaling of lead width (i.e. many more narrow leads than wide leads). [Equation 1](#) gives the relation between lead width and number of observed leads (as probability density) of a respective width assuming a power law relationship:

$$f(x) = ax^{-b}. \quad (1)$$

554 The parameter  $a$  is the scaling parameter (related to the number of measurements),  
 555 but not further analysed here. The parameter  $x$  is the variable lead width, and  
 556  $b$  is the power law exponent, determining the (negative) slope. Thus, a larger  
 557 power law exponent  $b$  results in a steeper (more negative) power law. The ratio  
 558 of the classified segment area (shown in red in [Figure 7](#)) and the major axis

length of the ellipse approximates the lead width. We detected in total 33855 lead segments in our classified maps for all 35 flights (but see explanation above why the number of segments should not be mistaken as the number of leads). The detected lead width varies between 1 m and 464 m. From the distribution of the lead widths, we perform a linear fit for Equation 1 (Figure 9 A) in the log-log space with logarithmic bins. We exclude leads smaller than 3 m width because they are too close to the spatial resolution of the data set to be fully resolved in the segmentation. This can be seen from the deviation from the power law below 3 m in Figure 9. This is confirmed by the stabilisation of the power law exponent for a minimum lead width of 3 m and larger (Figure S5, in Supplemental material). However, for a minimum lead width between 9 m and 26 m we see a slight increase of the power law exponent. We do not know the reason for the increase but our hypothesis is that the value is less reliable because of the strong decrease of number of observations available for the power law fit. Our power law is calculated up to the lead width of 336 m (largest logarithmic bin). The resulting exponent of  $b=2.63$  agrees with literature values at the upper end of the previously found exponent value range (2.0 to 2.6) (Wadhams, 1981; Wadhams et al., 1985; Marcq and Weiss, 2012; Wernecke and Kaleschke, 2015; Qu et al., 2019) and proves the compatibility with other datasets. From the stability of  $b$  in Figure S5 (in Supplemental material) we estimate the uncertainty of our  $b$  to be smaller than the range of the literature values of 2.0–2.6. The so far presented literature values of the power law exponents are summarized in Muchow et al. (2021). Lindsay and Rothrock (1995) determines a smaller exponent of  $1.6 \pm 0.18$  (less steep), which might differ because the power law is calculated to the lead width that is equal to the spatial resolution, while we see in our data that the power law is not valid anymore close to the spatial resolution and the slope between the bins has a smaller absolute value. In previous studies, the range of the power law exponent (dependent on the instrument and resolution) was determined starting between 20 m and 2 km lead width. Our study adds to the lower end of lead width with a range down to 3 m lead width and shows that the power law agrees with other methods and resolutions. As expressed by the power law, there are many more leads with small lead width, which were not resolved in previous studies. Also, our 3 m, the smallest resolvable lead width, likely is not the end of the lead width distribution. There are likely many cracks with a smaller width, which we do not resolve (but also can be important for, e.g., heat flux estimates). Our exponent is one of the largest (i.e., most negative) compared to the literature values. The other studies are also based on primarily winter data (Oct-Apr) but performed in different regions of the Arctic, which might influence the results due to different

597 characteristics of the ice rheology. The power law distribution tells us that there  
598 are many more narrow leads than wider leads, which emphasizes the importance  
599 of small-scale features. The area contribution of the smallest leads are: (i) 4%, for  
600 lead width  $< 3$  m, (ii) 64%, for lead width between 3 and 100 m, and (iii) 32%, for  
601 lead width  $> 100$  m.

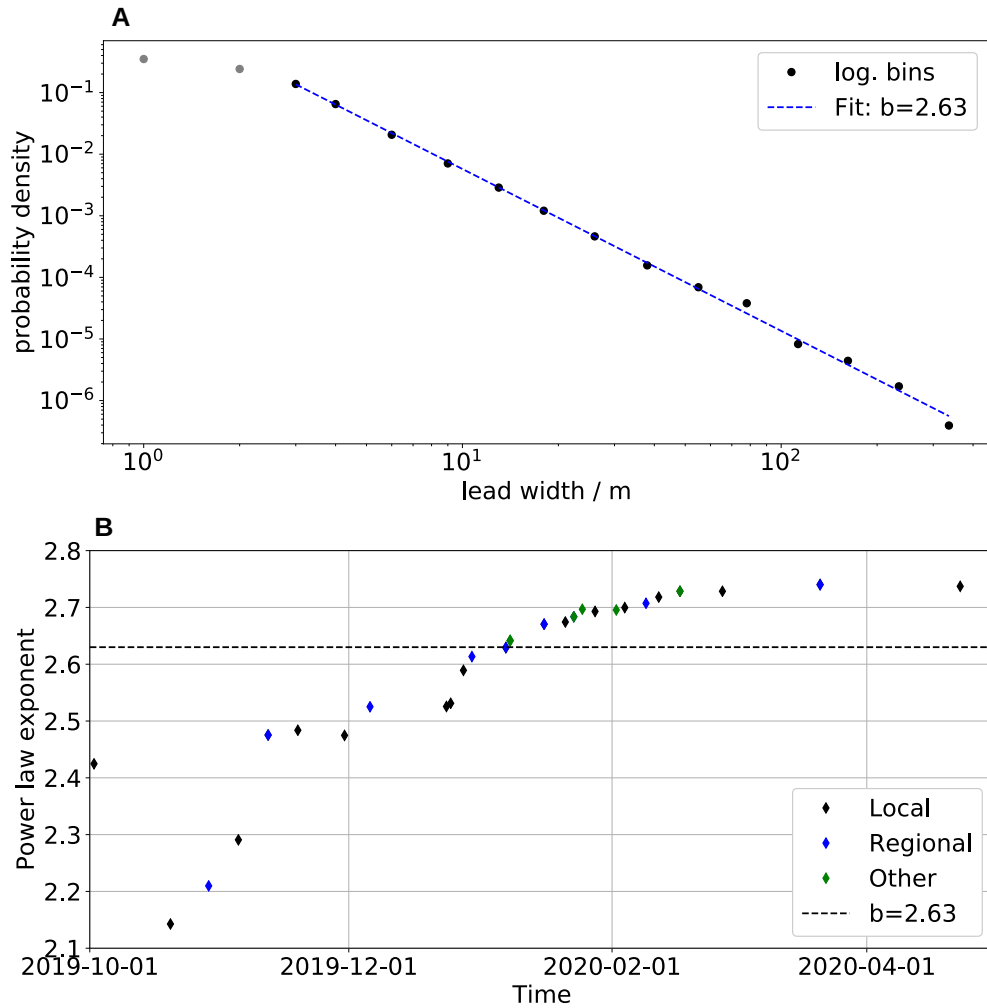
602 Additionally, we found a seasonal dependence of the power law exponent,  
603 with a tendency of an increasing power law exponent throughout the winter  
604 season (Figure 9 B). The seasonal increase in the power law exponent can  
605 also have a spatial component because of the drift into different regions during  
606 MOSAiC (Figure 1). The power law exponent drops from 2.42 to 2.14 at the  
607 start of the winter season in October (freeze-up time and consolidations of the  
608 ice north of the Laptev Sea). This is followed by a steady increase to 2.63 on  
609 07 January 2020 (Central Arctic). Following that, there is a further increase and  
610 then stabilisation around 2.74 in March and April (North of Svalbard). Mind  
611 that we are not covering the full melting and summer season, which again might  
612 introduce a change in the exponent. For the power law exponent, there is: (i) no  
613 scale dependence (no variation between local and regional flight, also on the same  
614 day; compare black and blue dots), and (ii) no clear effect on the exponent by a  
615 rapid change in lead area fraction (subsection 5.2) caused by ,e.g., the November  
616 storm event (Figure 9 B). An increasing exponent during winter time contrasts  
617 with the findings of Lindsay and Rothrock (1995) where the monthly average of  
618 the power law exponent for the central Arctic decreases from February to April  
619 and again decrease from October to December in the following season. We can  
620 only comment on several theories without providing a certain explanation why  
621 the power law exponent increases (relatively more narrow leads) throughout the  
622 winter. Three exponents in October and the beginning of November are lower (2.1  
623 to 2.3), probably because during the freeze-up phase the ice floes were still in  
624 rather free drift, which could more easily cause relatively wider leads (decrease of  
625 the exponent). The power law fit aligns for these three flights not as good as for the  
626 other flights, which is more prominent for smaller lead widths. In December and  
627 January, the exponent is increasing from 2.5 to 2.7, which may be related to a more  
628 consolidated and thicker ice pack far away from the coastlines (potential change of  
629 internal ice strength). This time was also characterized by less storms and lower  
630 wind speeds (Figure 6). The stabilization at the end of the winter could show  
631 the maximum power law exponent that can be reached during winter before it  
632 decreases towards the melt season with free drift conditions (which is not included  
633 in our dataset anymore).

634 The clear power law relationship for the lead width found here supports the

635 validity of our lead property data and that representative statistical conclusions  
636 can be obtained from it. Our results indicate that we miss many leads in satellite  
637 remote sensing products with coarser spatial resolutions. This could already be  
638 extrapolated from the found power-law in previous studies, but is here shown for  
639 the first time down to a lead width of 3 m. Our widest leads are still narrower  
640 than the resolution of about 1 km of thermal infrared satellites. Lead retrieval  
641 results vary (e.g., in frequency and how thick the ice in the lead can be) between  
642 different remote sensing approaches (von Albedyll et al., 2022). Thus, direct  
643 and absolute comparison of lead retrievals can be difficult for different products  
644 and resolutions. Nevertheless, the same physical principles (like the power law  
645 correspondence) are valid for different scales and resolutions (Wernecke and  
646 Kaleschke, 2015). We recommend that any lead width study should check if the  
647 power law scaling conditions are fulfilled to increase confidence in the validity of  
648 the obtained results.

## 649 7. Conclusions

650 On a local (5–10 km) and regional (20–40 km) scale, we investigate the variation  
651 of the surface temperatures in time and space and derived lead properties. Along  
652 the MOSAiC drift during the winter season, we use high resolution surface  
653 temperature maps obtained from helicopter flights to examine small-scale lead  
654 properties. First, we investigated the surface temperatures and found: (i) its  
655 temporal variability is influenced by meteorological changes, such as warm air  
656 intrusions, often associated with high wind speeds, and (ii) its spatial variability  
657 over thick ice decreases as wind speed increases. For each flight, we retrieve  
658 classified lead maps and lead area fractions, based on a lead classification applied  
659 to the surface temperature maps using a dynamic temperature threshold. We see  
660 a high variability of the local lead area fraction and the influence of events, like  
661 the November 2019 storm. On a regional scale, there is a more stable lead area  
662 fraction evolution between 0% and 1% (until January), followed by an increase  
663 to 4% (March). This evolution agrees well with other MOSAiC studies on the  
664 regional scale. From the classified lead maps, we additionally determine lead  
665 width and orientation for all lead segments of every flight. This reveals three main  
666 findings: 1) the lead width distribution follows a power law (Equation 1) with  
667 an exponent of  $b=2.63$  (narrow leads dominate wide leads), which is consistent  
668 with previous research, 2) the power law exponent increases in the course of  
669 the winter, 3) small-scale leads typically have one primary orientation. However,  
670 that orientation changes between the flights and throughout the winter season and



**Figure 9. Lead width distribution with the power law fit for all and single flights.**

(A) The logarithmic frequencies of the lead widths of all 35 flights combined, also binned logarithmic, are represented as black points. The blue dashed line shows the negative power law fit exponent  $b=2.63$ . The power law fit is constrained to the lead width  $\geq 3$  m. (B) Time series of the power law exponent for all 35 flights between 02 October 2019 and 23 April 2020; in black for local flight, in blue for regional flights, and in green for other flight types. The horizontal line marks the exponent of all flights (2.63) from (A).

671 no overall prevailing orientation is found. The abundance of small scale leads  
 672 emphasises the necessity to understand their physical processes better, where  
 673 our high spatial resolution data can help. However, those narrow leads are not  
 674 individually included in the current thermal infrared satellite data of about 1 km  
 675 resolution (e.g., MODIS). We suggest a representation of the smallest leads on  
 676 the satellite sub-footprint scale because the heat transfer is not linear with surface  
 677 temperature. In fact, the heat exchange is larger for leads within thick sea ice  
 678 compared to larger areas of uniformed thinner sea ice with the same average  
 679 surface temperature. Such parameterizations could also improve model simulation  
 680 for considering small scale lead processes.

681 Beyond this study, we plan to perform a one-to-one comparison of the  
 682 high-resolution helicopter-borne data with thermal infrared satellite data, e.g.,  
 683 MODIS ice surface temperatures. The aim is to assess how well the lead's  
 684 impact on the atmosphere is represented in the satellite retrieval. Additionally,  
 685 comparisons with the deformation rate from buoy grids on different scales or  
 686 inter-comparison with helicopter-borne topography data can be used to understand  
 687 the MOSAiC lead characteristics better.

## 688 **References**

- 689 Alfred-Wegener-Institut Helmholtz-Zentrum für Polar- und Meeresforschung.  
 690 2017. Polar Research and Supply Vessel POLARSTERN operated by  
 691 the Alfred-Wegener-Institute. *Journal of large-scale research facilities* **3**.  
 692 doi:<http://dx.doi.org/10.17815/jlsrf-3-163>.
- 693 Bintanja R, Van Der Linden EC. 2013. The changing seasonal climate in the  
 694 Arctic. *Scientific Reports* **3**: 1–8. ISSN 20452322. doi:<https://doi.org/10.1038/srep01556>.
- 696 Bröhan D, Kaleschke L. 2014. A nine-year climatology of arctic sea ice lead  
 697 orientation and frequency from AMSR-E. *Remote Sensing* **6**(2): 1451–1475.  
 698 doi:[10.3390/rs6021451](https://doi.org/10.3390/rs6021451).
- 699 Burger W, Burge MJ. 2009. *Undergraduate Topics in Computer Science. Principles of Digital Image Processing. Core Algorithms*. Springer-Verlag  
 700 London. doi:<https://doi.org/10.1007/978-1-84800-195-4>.
- 702 Cox C, Gallagher M, Shupe M, Persson O, Solomon A, et al. 2021. 10-meter (m)  
 703 meteorological flux tower measurements (Level 1 Raw), Multidisciplinary  
 704 Drifting Observatory for the Study of Arctic Climate (MOSAiC), central  
 705 Arctic, October 2019 - September 2020 [dataset]. *Arctic Data Center*  
 706 <https://doi.org/10.18739/A2VM42Z5F>.

- 707 Dai A, Luo D, Song M, Liu J. 2019. Arctic amplification is caused by sea-ice  
708 loss under increasing CO<sub>2</sub>. *Nature communications* **10**(1): 1–13. doi:<https://doi.org/10.1038/s41467-018-07954-9>.  
709
- 710 Dethloff K, Maslowski W, Hendricks S, Lee YJ, Goessling HF, et al. 2022.  
711 Arctic sea ice anomalies during the MOSAiC winter 2019/20. *The*  
712 *Cryosphere* **16**(3): 981–1005. doi:10.5194/tc-16-981-2022. <https://tc.copernicus.org/articles/16/981/2022/>.  
713
- 714 Fox-Kemper B, et al. 2021. The Physical Science Basis. Contribution of Working  
715 Group I to the Sixth Assessment Report of the Intergovernmental Panel on  
716 Climate Change: Ocean, Cryosphere and Sea Level Change. *Climate change*  
717 .
- 718 Guo W, Itkin P, Singha S, Paul Doulgeris A, Johansson M, et al. 2022. Sea ice  
719 classification of TerraSAR-X ScanSAR images for the MOSAiC expedition  
720 incorporating per-class incidence angle dependency of image texture. *The*  
721 *Cryosphere Discussions* **2022**: 1–29. doi:<https://doi.org/10.5194/tc-2022-86>.  
722
- 723 Hutter N, Bouchat A, Dupont F, Dukhovskoy D, Koldunov N, et al. 2022. Sea  
724 Ice Rheology Experiment (SIREx): 2. Evaluating Linear Kinematic Features  
725 in High-Resolution Sea Ice Simulations. *Journal of Geophysical Research: Oceans* **127**(4). doi:<https://doi.org/10.1029/2021JC017666>.  
726
- 727 Hutter N, Losch M, Menemenlis D. 2018. Scaling Properties of Arctic Sea Ice  
728 Deformation in a High-Resolution Viscous-Plastic Sea Ice Model and in  
729 Satellite Observations. *Journal of Geophysical Research: Oceans* **123**(1):  
730 672–687. doi:<https://doi.org/10.1002/2017JC013119>.  
731
- 732 Ivanova N, Rampal P, Bouillon S. 2016. Error assessment of satellite-derived  
733 lead fraction in the Arctic. *Cryosphere* **10**(2): 585–595. doi:<https://doi.org/10.5194/tc-10-585-2016>.  
734
- 735 Kortum K, Singha S, Spreen G. 2022. Robust Multiseasonal Ice Classification  
736 From High-Resolution X-Band SAR. *IEEE Transactions on Geoscience and*  
737 *Remote Sensing* **60**: 1–12. doi:<https://doi.org/10.1109/TGRS.2022.3144731>.  
738
- 739 Krumpfen T, Birrien F, Kauker F, Rackow T, Albedyll Lv, et al. 2020. The  
740 MOSAiC ice floe: sediment-laden survivor from the Siberian shelf. *The*  
741 *Cryosphere* **14**(7): 2173–2187. doi:<https://doi.org/10.5194/tc-14-2173-2020>.  
742
- 743 Krumpfen T, von Albedyll L, Goessling HF, Hendricks S, Juhls B, et al. 2021.  
744 MOSAiC drift expedition from October 2019 to July 2020: Sea ice conditions  
from space and comparison with previous years. *The Cryosphere* **15**(8):  
3897–3920. doi:<https://doi.org/10.5194/tc-15-3897-2021>.
- 745 Kwok R. 2018. Arctic sea ice thickness, volume, and multiyear ice coverage:  
Losses and coupled variability (1958-2018). *Environmental Research Letters*

- 745 **13**(10). doi:<https://doi.org/10.1088/1748-9326/aae3ec>.
- 746 Kwok R, Spreen G, Pang S. 2013. Arctic sea ice circulation and drift speed:  
747 Decadal trends and ocean currents. *Journal of Geophysical Research: Oceans*  
748 **118**(5): 2408–2425. doi:<https://doi.org/10.1002/jgrc.20191>.
- 749 Lindsay R, Rothrock D. 1995. Arctic sea ice leads from advanced very high  
750 resolution radiometer images. *Journal of Geophysical Research: Oceans*  
751 **100**(C3): 4533–4544. doi:<https://doi.org/10.1029/94JC02393>.
- 752 Lüpkes C, Vihma T, Bimbaum G, Wacker U. 2008. Influence of leads in sea ice  
753 on the temperature of the atmospheric boundary layer during polar night.  
754 *Geophysical Research Letters* **35**(3): 2–6. ISSN 00948276. doi:<https://doi.org/10.1029/2007GL032461>.
- 756 Marcq S, Weiss J. 2012. Influence of sea ice lead-width distribution on turbulent  
757 heat transfer between the ocean and the atmosphere. *Cryosphere* **6**(1):  
758 143–156. doi:<https://doi.org/10.5194/tc-6-143-2012>.
- 759 Masson-Delmotte V, Zhai P, Pirani A, Connors S, Péan C, et al. 2021. IPCC, 2021:  
760 Climate Change 2021: The Physical Science Basis. Contribution of Working  
761 Group I to the Sixth Assessment Report of the Intergovernmental Panel on  
762 Climate Change.
- 763 Maykut GA. 1978. Energy exchange over young sea ice in the central Arctic.  
764 *Journal of Geophysical Research: Oceans* **83**(C7): 3646–3658. doi:<https://doi.org/10.1029/JC083iC07p03646>.
- 766 Maykut GA. 1982. Large-Scale Heat Exchange and Ice Production in the Central  
767 Arctic. *Journal of Geophysical Research* **87**(C10): 7971–7984. doi:<https://doi.org/10.1029/JC087iC10p07971>.
- 769 Meredith M, Sommerkorn M, Cassotta S, Derksen C, Ekaykin A, et al. 2019. Polar  
770 Regions. Chapter 3, IPCC Special Report on the Ocean and Cryosphere in a  
771 Changing Climate. *IPPC* .
- 772 Muchow M, Schmitt AU, Kaleschke L. 2021. A lead-width distribution for  
773 Antarctic sea ice: a case study for the Weddell Sea with high-resolution  
774 Sentinel-2 images. *The Cryosphere* **15**(9): 4527–4537. doi:<https://doi.org/10.5194/tc-15-4527-2021>.
- 776 Najman L, Schmitt M. 1994. Watershed of a continuous function. *Signal*  
777 *Processing* **38**(1): 99–112. doi:[https://doi.org/10.1016/0165-1684\(94\)90059-0](https://doi.org/10.1016/0165-1684(94)90059-0).
- 779 Nandan V, Willatt R, Mallett R, Stroeve J, Geldsetzer T, et al. 2022. Wind  
780 Transport of Snow Impacts Ka- and Ku-band Radar Signatures on Arctic  
781 Sea Ice. *The Cryosphere Discussions* **2022**: 1–38. doi:10.5194/tc-2022-116.
- 782 Nicolaus M, Perovich DK, Spreen G, Granskog MA, von Albedyll L, et al. 2022.



- 783 Overview of the MOSAiC expedition: Snow and sea ice. *Elem Sci Anth*  
784 **10**(1): 000046. doi:<https://doi.org/10.1525/elementa.2021.000046>.
- 785 Nixdorf U, Dethloff K, Rex M, Shupe M, Sommerfeld A, et al. 2021.  
786 MOSAiC extended acknowledgement. *Zenodo* [https://doi.org/10.  
787 5281/zenodo.5541624](https://doi.org/10.5281/zenodo.5541624).
- 788 Qu M, Pang X, Zhao X, Zhang J, Ji Q, et al. 2019. Estimation of turbulent heat flux  
789 over leads using satellite thermal images. *The Cryosphere* **13**(6): 1565–1582.  
790 doi:<https://doi.org/10.5194/tc-13-1565-2019>.
- 791 Rabe B, Heuzé C, Regnery J, Aksenov Y, Allerholt J, et al. 2022. Overview of the  
792 MOSAiC expedition: Physical oceanography. *Elem Sci Anth* **10**(1): 00062.  
793 doi:<https://doi.org/10.1525/elementa.2021.00062>.
- 794 Rampal P, Weiss J, Marsan D. 2009. Positive trend in the mean speed and  
795 deformation rate of Arctic sea ice, 1979–2007. *Journal of Geophysical*  
796 *Research: Oceans* **114**(5). doi:<https://doi.org/10.1029/2008JC005066>.
- 797 Ridler T, Calvard S, et al. 1978. Picture thresholding using an iterative selection  
798 method. *IEEE trans syst Man Cybern* **8**(8): 630–632. doi:[https://doi.org/10.  
799 1109/TSMC.1978.4310039](https://doi.org/10.1109/TSMC.1978.4310039).
- 800 Ringeisen D, Hutter N, von Albedyll L. 2022. Deformation lines in Arctic sea ice:  
801 intersection angles distribution and mechanical properties. *EGUsphere* pp.  
802 1–22. doi:10.5194/egusphere-2022-1481.
- 803 Ringeisen D, Losch M, Bruno Tremblay L, Hutter N. 2019. Simulating  
804 intersection angles between conjugate faults in sea ice with different  
805 viscous-plastic rheologies. *Cryosphere* **13**(4): 1167–1186. doi:[https://doi.  
806 org/10.5194/tc-13-1167-2019](https://doi.org/10.5194/tc-13-1167-2019).
- 807 Ringeisen D, Tremblay LB, Losch M. 2021. Non-normal flow rules affect fracture  
808 angles in sea ice viscous-plastic rheologies. *Cryosphere* **15**(6): 2873–2888.  
809 doi:<https://doi.org/10.5194/tc-15-2873-2021>.
- 810 Rinke A, Cassano JJ, Cassano EN, Jaiser R, Handorf D. 2021.  
811 Meteorological conditions during the MOSAiC expedition.  
812 *Elementa: Science of the Anthropocene* **9**(1). doi:<https://doi.org/10.1525/elementa.2021.00023>.  
813 [https://online.  
814 ucpress.edu/elementa/article/9/1/00023/118092/  
815 Meteorological-conditions-during-the-MOSAiC](https://online.ucpress.edu/elementa/article/9/1/00023/118092/Meteorological-conditions-during-the-MOSAiC).
- 816 Schulson EM. 2004. Compressive shear faults within arctic sea ice: Fracture on  
817 scales large and small. *Journal of Geophysical Research C: Oceans* **109**(7).  
818 doi:<https://doi.org/10.1029/2003JC002108>.
- 819 Serreze MC, Barrett AP, Stroeve JC, Kindig DN, Holland MM. 2009. The  
820 emergence of surface-based Arctic amplification. *Cryosphere* **3**(1): 11–19.

- 821 ISSN 19940424. doi:<https://doi.org/10.5194/tc-3-11-2009>.
- 822 Serreze MC, Barry RG. 2011. Processes and impacts of Arctic amplification:  
823 A research synthesis. *Global and Planetary Change* **77**(1): 85–96. ISSN  
824 0921-8181. doi:<https://doi.org/10.1016/j.gloplacha.2011.03.004>.
- 825 Shokr M, Sinha N. 2015. *Sea ice: physics and remote sensing*. John Wiley & Sons.
- 826 Shupe MD, Rex M, Blomquist B, Persson POG, Schmale J, et al. 2022. Overview  
827 of the MOSAiC expedition: Atmosphere. *Elem Sci Anth* **10**(1): 00060.  
828 doi:<https://doi.org/10.1525/elementa.2021.00060>.
- 829 Shupe MD, Rex M, Dethloff K, Damm E, Fong A, et al. 2020. Arctic Report Card  
830 2020: The MOSAiC Expedition: A Year Drifting with the Arctic Sea Ice.  
831 NOAA doi:<https://doi.org/10.25923/9g3v-xh92>.
- 832 Spreen G, Kern S. 2017. Methods of satellite remote sensing of sea ice, in *Sea ice*  
833 / edited by David N. Thomas. John Wiley & Sons: Chap. 9.
- 834 Spreen G, Kwok R, Menemenlis D. 2011. Trends in Arctic sea ice drift and  
835 role of wind forcing: 1992-2009. *Geophysical Research Letters* **38**(19): 1–6.  
836 doi:<https://doi.org/10.1029/2011GL048970>.
- 837 Tetzlaff A, Lüpkes C, Hartmann J. 2015. Aircraft-based observations of  
838 atmospheric boundary-layer modification over Arctic leads. *Quarterly*  
839 *Journal of the Royal Meteorological Society* **141**(692): 2839–2856. doi:<https://doi.org/10.1002/qj.2568>.
- 841 Thielke L, Huntemann M, Hendricks S, Jutila A, Ricker R, et al.  
842 2022. Helicopter-borne thermal infrared sea ice surface temperature  
843 maps with 1 m resolution during the MOSAiC expedition, NetCDF  
844 format, version 3 [dataset]. PANGAEA [https://doi.org/10.1594/](https://doi.org/10.1594/PANGAEA.950683)  
845 [PANGAEA.950683](https://doi.org/10.1594/PANGAEA.950683).
- 846 Thielke L, Huntemann M, Hendricks S, Jutila A, Ricker R, et al. 2022. Sea ice  
847 surface temperatures from helicopter-borne thermal infrared imaging during  
848 the MOSAiC expedition. *Scientific Data* **9**(1): 1–16. doi:[https://doi.org/10.](https://doi.org/10.1038/s41597-022-01461-9)  
849 [1038/s41597-022-01461-9](https://doi.org/10.1038/s41597-022-01461-9).
- 850 Thielke L, Huntemann M, Spreen G. 2022. Lead classification maps from  
851 helicopter-borne surface temperatures with 1 m resolution during the  
852 MOSAiC expedition [dataset]. PANGAEA [https://doi.org/10.](https://doi.org/10.1594/PANGAEA.951568)  
853 [1594/PANGAEA.951568](https://doi.org/10.1594/PANGAEA.951568).
- 854 Tschudi MA, Curry JA, Maslanik JA. 2002. Characterization of springtime leads  
855 in the Beaufort/Chukchi seas from airborne and satellite observations during  
856 FIRE/SHEBA. *Journal of Geophysical Research: Oceans* **107**(10). doi:[https://doi.org/10.](https://doi.org/10.1029/2000jc000541)  
857 [1029/2000jc000541](https://doi.org/10.1029/2000jc000541).
- 858 Vihma T, Pirazzini R. 2005. On the factors controlling the snow

- 859 surface and 2-m air temperatures over the Arctic sea ice in winter.  
860 *Boundary-Layer Meteorology* **117**(1): 73–90. doi:[https://doi.org/10.1007/  
861 s10546-004-5938-7](https://doi.org/10.1007/s10546-004-5938-7).
- 862 Vihma T, Pirazzini R, Fer I, Renfrew IA, Sedlar J, et al. 2014. Advances in  
863 understanding and parameterization of small-scale physical processes in the  
864 marine Arctic climate system: a review. *Atmospheric Chemistry and Physics*  
865 **14**(17): 9403–9450. doi:<https://doi.org/10.5194/acp-14-9403-2014>.
- 866 von Albedyll L, et al. 2022. Lead fractions from SAR-derived sea ice divergence  
867 during MOSAiC. *In preparation* .
- 868 Wadhams P. 1981. Sea-ice topography of the Arctic Ocean in the region 70 W to  
869 25 E. *Philosophical Transactions of the Royal Society of London Series A,  
870 Mathematical and Physical Sciences* **302**(1464): 45–85. doi:[https://doi.org/  
871 10.1098/rsta.1981.0157](https://doi.org/10.1098/rsta.1981.0157).
- 872 Wadhams P, McLaren AS, Weintraub R. 1985. Ice thickness distribution  
873 in Davis Strait in February from submarine sonar profiles. *Journal of  
874 Geophysical Research: Oceans* **90**(C1): 1069–1077. doi:[https://doi.org/10.  
875 1029/JC090iC01p01069](https://doi.org/10.1029/JC090iC01p01069).
- 876 Wang Q, Danilov S, Jung T, Kaleschke L, Wernecke A. 2016. Sea ice  
877 leads in the Arctic Ocean: Model assessment, interannual variability and  
878 trends. *Geophysical Research Letters* **43**(13): 7019–7027. ISSN 19448007.  
879 doi:<https://doi.org/10.1002/2016GL068696>.
- 880 Wang S, Wang Q, Jordan RE, Persson PO. 2001. Interactions among longwave  
881 radiation of clouds, turbulence, and snow surface temperature in the  
882 Arctic: A model sensitivity study. *Journal of Geophysical Research  
883 Atmospheres* **106**(D14): 15323–15333. ISSN 01480227. doi:[https://doi.org/  
884 10.1029/2000JD900358](https://doi.org/10.1029/2000JD900358).
- 885 Wendisch M, Brückner M, Burrows J, Crewell S, Dethloff K, et al. 2017.  
886 Understanding causes and effects of rapid warming in the Arctic. *Eos* **98**.  
887 doi:<http://dx.doi.org/10.1029/2017EO064803>.
- 888 Wernecke A, Kaleschke L. 2015. Lead detection in Arctic sea ice from CryoSat-2:  
889 Quality assessment, lead area fraction and width distribution. *Cryosphere*  
890 **9**(5): 1955–1968. doi:<https://doi.org/10.5194/tc-9-1955-2015>.
- 891 Willmes S, Heinemann G. 2015. Pan-arctic lead detection from MODIS thermal  
892 infrared imagery. *Annals of Glaciology* **56**(69): 29–37. ISSN 02603055.  
893 doi:<https://doi.org/10.3189/2015AoG69A615>.
- 894 Willmes S, Heinemann G. 2016. Sea-ice wintertime lead frequencies and regional  
895 characteristics in the Arctic, 2003–2015. *Remote Sensing* **8**(1): 2003–2015.  
896 ISSN 20724292. doi:<https://doi.org/10.3390/rs8010004>.

- 897 WMO. 2014. SEA ICE NOMENCLATURE. [https://library.wmo.int/index.php?lvl=notice\\_display&id=6772,](https://library.wmo.int/index.php?lvl=notice_display&id=6772)  
898 [https://library.wmo.int/index.php?lvl=notice\\_display&id=6772,](https://library.wmo.int/index.php?lvl=notice_display&id=6772)  
899 [lastaccessFebruary16,2022.](https://library.wmo.int/index.php?lvl=notice_display&id=6772)
- 900 Yu Y, Rothrock DA. 1996. Thin ice thickness from satellite thermal imagery.  
901 *Journal of Geophysical Research C: Oceans* **101**(C11): 25753–25766.  
902 doi:<https://doi.org/10.1029/96JC02242>.

### 903 **Contributions**

- 904 Contributed to conception and design: LT, GS, MH  
905 Contributed to acquisition of data: LT, GS, MH  
906 Contributed to analysis and interpretation of data: all authors  
907 Drafted the article: LT  
908 Revised the article: all authors  
909 Approved the submitted version for publication: all authors  
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### 926 **Competing interests**

927 The authors have no competing interests, as defined by Elementa, that might be  
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929 **Data accessibility statement**

- 930 • Helicopter-borne surface temperature maps, 1 m resolution: (Thielke et al.,  
931 2022).
- 932 • Lead classification maps, 1 m resolution: (Thielke et al., 2022)
- 933 • Atmospheric in-situ data: Cox et al. (2021) [updated version used]