Discovery of a 15-kilometer-wide impact crater candidate in western Saudi Arabia

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Abdulrahman F. Toonsi*

Abstract

There are few reports on impact craters in Saudi Arabia. This paper aims to fill a gap in the literature in this area. Using satellite imaging, a circular structure with a diameter of approximately 15 kilometers was detected on the edge of Harrat Rahat volcanic field in western Saudi Arabia. Fieldwork showed that the circular structure is a newly-discovered impact crater candidate, this theory is aided by the presence of shock metamorphic features such as planar fractures in quartz grains and impact breccia. This impact crater candidate could be one of the largest exposed and preserved impact craters in Saudi Arabia.

Introduction

Impact structures are geological structures caused by extraterrestrial impacts such as asteroids and comets. When an impact occurs, a large mass of rock is vaporized and is subjected to large amounts of heat and pressure. The impact also breaks the rock beneath it and forces the broken shards together to form impact breccia. In order to prove impact events, certain geological features must be present. These include: shocked quartz (quartz with planar features), shatter cones and impact breccia. There are many terrestrial impact craters discovered and much more yet to be discovered (French, Bevan, and Koeberl, 2009, 123-70).

Methods

This paper describes the discovery of a large previously unknown impact crater candidate in western Saudi Arabia, on the eastern edge of Harat Rahat, which is a well-known volcanic field in the area. The search for terrestrial impact sites was conducted by using Google Earth satellite imagery. After the circular feature was identified, geological maps created by the Saudi Geological Survey were reviewed to check the general geology of the area. After that, the site was visited and rock and samples were taken from 5 different locations around the crater rim. See satellite images in figures 1-3.



Figure 1. The image above shows sample locations from different areas in the crate rim



Figure 2. The image above shows a comparison of the impact crater candidate(left) described in this paper compared to the Al Wahba volcanic crater(right). It highlights the difference in size and rules out the possibility that the crater on the left is of volcanic origin.



Figure 3. The image above shows the study area where the crater is located

Sand samples were taken from the deposits of seasonal streams coming from the crater rim in order to locate quartz grains with planar features. Microscopic images of sand samples were taken with an optical microscope with a magnification power of 40X to identify quartz grain with planar features (PFs).

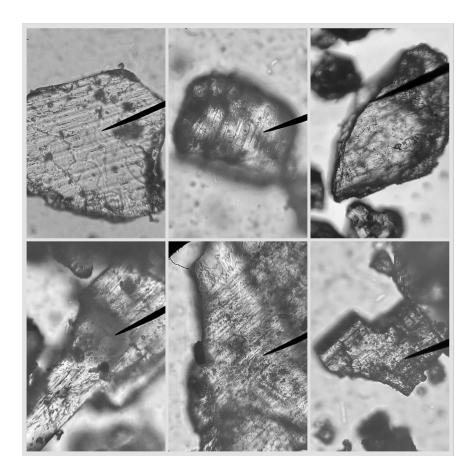


Figure 4: Planar features in 6 grains of quartz which is evidence of shock metamorphism, these samples were taken from a small gully originating from the crater rim.

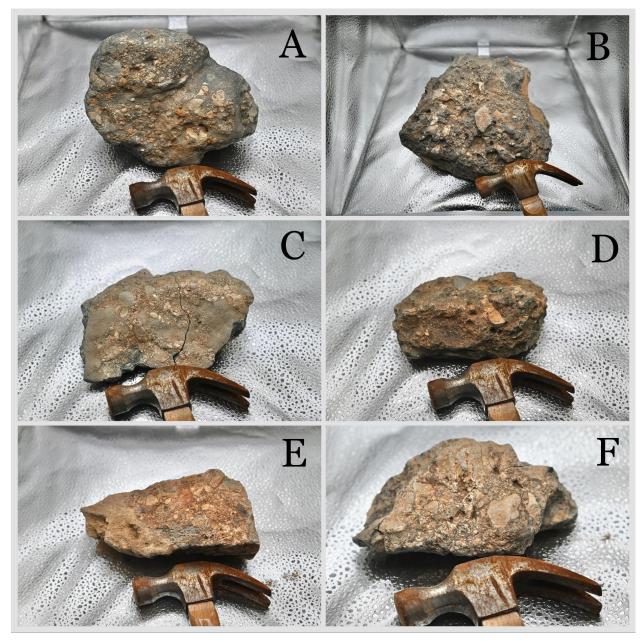


Figure 5: (A): Polymict breccia found in the south eastern part of the crater rim at the coordinates: 22 49'28"N 40 49'25"E. (B): Polymict breccia found in the western part of the crater rim at the coordinates: 22 50'00"N 4042'30"E. (C): Polymict breccia found in the eastern part of the crater rim at the coordinates: 22 52'26"N 40 50'16"E. (D): Polymict breccia found in the western part of the crater rim at the coordinates: 22 50'00"N 4042'30"E. (E): Polymict breccia found in the western part of the crater rim at the coordinates: 22 50'00"N 4042'30"E. (E): Polymict breccia found in the western part of the crater rim at the coordinates: 22 50,42"N 40 42'18"E. (F): Polymict breccia found in the western part of the crater rim at the coordinates: 22 50'00"N 4042'30"E.

The age of impact was estimated by looking for basalt fragments within the polymict breccia, no basalt was found within the breccia implying that the impact occurred before the formation of the basalt from Harat Rahat which has been dated to the Neogene.

Results

The crater candidate described in this study is located in western Saudi Arabia, approximately 215 kilometers from Jeddah with the coordinates of 22°52'28"N 40° 46' 16"E. It lies on the eastern edge of Harrat Rahat.

It measures approximately 15 km in diameter. The crater rim height was relatively similar at different measuring points and was approximately 50-70 meters higher than the surrounding plains, as measured by satellite imagery.

The rim was almost entirely composed of impact breccia containing crystalline rocks (figure 5). Impact breccia was found in 5 different regions of the crater rim, with the farthest distance between two sample points being 15 km from each other. Ten grains of shocked quartz candidates were found upon microscopic examination of sand samples. Images in figure 4 demonstrate these findings.

Discussion

This paper describes the discovery of a previously unknown impact crater candidate in western Saudi Arabia. This may be the largest exposed and preserved impact crater reported in Saudi Arabia as of yet.

Several features of this crater resemble those of other craters found on earth. Like other confirmed terrestrial impact craters, this crater hosts impact breccia, quartz with planar features, and a circular crater rim.

The map created by the Saudi Geological survey indicates that this area contains conglomerate rocks, which includes breccia. This further confirms the nature of the breccia/rocks reported here. Due to the sheer amount of breccia and with this great distance, the chances of this breccia forming in traditional ways such as stream deposits and alluvial fan are greatly reduced, leaving the most probable cause of formation being extraterrestrial impact.

The impact breccia did not contain any fragments of basalt from Harrat Rahat which implies that the asteroid or comet impacted before Harrat Rahat's formation which puts the age of the crater to be >10 million years. No uranium-lead dating has been done to confirm the age of the crater due to lack of equipment but a rough estimate can be made. Due to the large amounts of erosion the crater has faced and the thick deposits of sand that has covered the center of the crater, it is estimated that the crater formed either during the proterozoic or paleozoic eras, and unlikely to have happened in the cenozoic.

The crater could have been made by:

1: 1.1-kilometer in diameter asteroid made of dense rock traveling at 17 km/s which is the average speed of asteroid collisions with earth (figure 6).

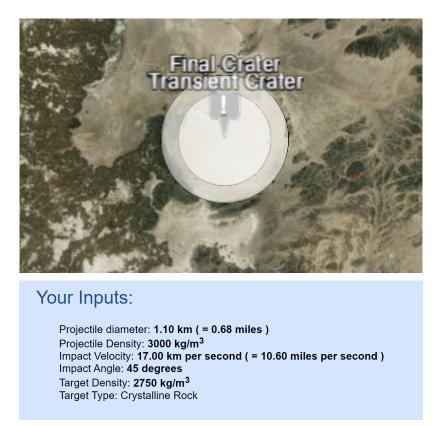
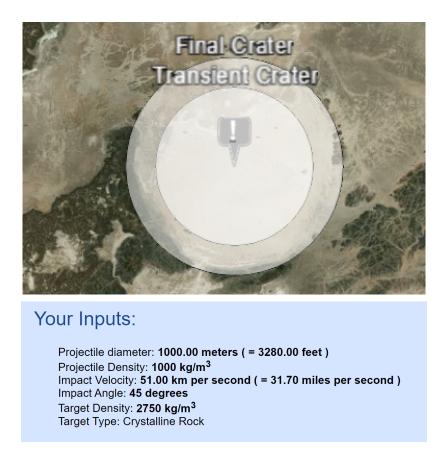


Figure 6: Figures 6 is a simulation conducted using the Impact Crater Map (Imperial College London) made by Gareth Collins, Robert Marcus, H.J Melosh.

2: 1-kilometer in diameter comet made of porous rock or ice traveling at 51 km/s which is the average speed of comet impacts (figure 7).



Figures 7: Figures 7 is a simulation conducted using the Impact Crater Map (Imperial College London) made by Gareth Collins, Robert Marcus, H.J Melosh.

Both of these scenarios would create a very similar impact size and structure to what has been described in this study, this is because the asteroid has a dense composition and slower speed but the comet has a low density but makes up for this with a much faster speed. The impact crater also appears to have a slight ellipse with the crater being 15.2 kilometers east to west and 16.3 kilometers from north to south. Whatever did cause the impact, the implications would be similar.

If the impact happened in the Phanerozoic then it could be the cause of a minor regional extinction. Future research could be aimed at determining an accurate age of impact, determining the age of impact could find a correlation between a possible regional extinction event. Future research could also look for meteoritic fragments to determine what type of asteroid or determine if the impactor was a comet.

Regarding the naming of the impact crater, As the discoverer, I propose the name "Toonsi Crater" for further discussion and evaluation by the scientific community.

References

(French, Bevan M., and Christian Koeberl. "The Convincing Identification of Terrestrial Meteorite Impact Structures: What Works, What Doesn't, and Why." *Earth-Science Reviews* 98, no. 1-2 (2010): 123–70. https://doi.org/10.1016/j.earscirev.2009.10.009).